

Title: Gravitational Collapse in SD

Date: Jun 25, 2015 09:45 AM

URL: <http://pirsa.org/15060026>

Abstract:



SAPIENZA
UNIVERSITÀ DI ROMA

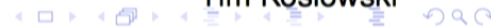


SD@Convergence Workshop

Gravitational Collapse in SD

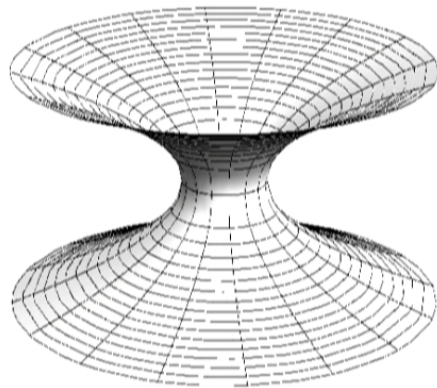
Andrea Napoletano
Sapienza Università di Roma

In collaboration with:
Flavio Mercati
Henrique Gomes
Tim Koslowski



Isotropic Wormhole Solution

$$ds^2 = \left(\frac{1 - \frac{\alpha}{4r}}{1 + \frac{\alpha}{4r}} \right)^2 dt^2 + \left(1 + \frac{\alpha}{4r} \right)^4 (dr^2 + r^2 d\Omega_2)$$



- ▶ No Singularity
- ▶ Symmetry under inversion $r \rightarrow \frac{\alpha^2}{16r}$
- ▶ Two Schwarzschild exteriors glued together at the throat

THE ISOTROPIC SOLUTION IS ETERNAL

"A Birkhoff theorem for Shape Dynamics" H. Gomes arXiv:1305.0310



The Thin Shell Model

WHAT?

Dynamical creation of the isotropic solution through gravitational collapse

HOW?

Spherically symmetric infinitely thin shell of dust



The Thin Shell Model: Starting Point

Metric tensor and conjugate momentum

$$g_{ij} = \begin{pmatrix} \mu^2 & 0 & 0 \\ 0 & \sigma & 0 \\ 0 & 0 & \sigma \sin(\theta)^2 \end{pmatrix} \quad p^{ij} = \sin(\theta) \begin{pmatrix} \frac{f}{\mu} & 0 & 0 \\ 0 & \frac{1}{2}s & 0 \\ 0 & 0 & \frac{1}{2}s \sin(\theta)^{-2} \end{pmatrix}$$

Hamiltonian, diffeo and conformal constraints

$$\frac{1}{2\sigma\mu^2} [2f\sigma\mu^2s - f^2\mu^3 + \mu(\sigma')^2 + 4\sigma\mu^3 + 4\sigma\mu'\sigma' - 4\sigma\mu\sigma''] = \delta(r - R) \sqrt{\frac{p^2}{\mu^2} + M^2}$$
$$\mu f' - \frac{1}{2}s\sigma' = -\frac{P}{2}\delta(r - R)$$
$$\mu f + s\sigma = 0$$



Solution to Constraints

Solutions

- ▶ $f = p^{rr} \mu = \frac{A}{\sqrt{\sigma}}$
- ▶ $\mu^2 = g_{rr} = \frac{(\sigma')^2}{4\alpha\sqrt{\sigma} + 4\sigma + \frac{A^2}{\sigma}}$

Integration Constants

- ▶ A_{in}, A_{out}
- ▶ $\alpha_{in}, \alpha_{out}$

4 spatial integration constants
2 inside the shell, 2 outside the shell

Metric EoM and LFE

Metric Tensor

$$\dot{g}_{ij} = \left(\frac{2f\mu^2 N}{\sigma} + 2\mu\xi\mu' + 2\mu^2\xi' \right) \delta^r_i \delta^r_j + \left(\frac{\sigma s N}{\mu} + \xi\sigma' \right) (\delta^\theta_i \delta^\theta_j + \delta^\phi_i \delta^\phi_j \sin^2 \theta)$$

Conjugate Momentum

$$\begin{aligned} \dot{p}^{ij} = & -\frac{\sin \theta}{4\sigma\mu^3} \left[5f^2 N\mu^2 + 4\sigma(N'\sigma' - \mu^2\xi f') + N(-4\sigma\mu^2 + (\sigma')^2) + 4f\sigma\mu(\xi\mu' + \mu\xi') \right] \delta^i_r \delta^j_r \\ & + \frac{\sin \theta}{4\sigma^2\mu^2} \left[(f^2 N\mu^3 + N\mu(\sigma')^2 + 2\sigma^2(2N'\mu' + \mu^2(\xi s' + s\xi') - 2\mu N'')) \right. \\ & \left. - 2\sigma(-N\sigma'\mu' + \mu(N'\sigma' + N\sigma'')) \right] (\delta^i_\theta \delta^j_\theta + \delta^i_\phi \delta^j_\phi \sin^{-2} \theta) \\ & + \sin \theta \delta^i_r \delta^j_r \frac{p^2/\mu^4}{2\sqrt{p^2/\mu^2 + M^2}} N(R)\delta(r - R) \end{aligned}$$

Lapse Fixing Equation

$$\begin{aligned} \frac{1}{4\sigma\mu^2} \left[6f^2 N\mu^3 + 4\sigma \left(-\mu(2N'\sigma' + N\sigma'') + N\sigma'\mu' + N\mu^3 \right) + N\mu(\sigma')^2 \right. \\ \left. + \sigma^2 \left(\mu(3Ns^2 - 8N'') + 8N'\mu' \right) \right] + \frac{p^2/\mu^2}{2\sqrt{p^2/\mu^2 + M^2}} N\delta(r - R) = 0 \end{aligned}$$



Solutions to EoM

Lapse function

$$N = \frac{\sigma'}{2\mu\sqrt{\sigma}} \left(c_1 + c_2 \int_{r_a}^r dy \frac{\mu^3(y)}{(\sigma'(y))^2} \right)$$

Shift vector

$$\xi = \frac{\dot{\sigma}}{\sigma'} + \frac{A}{\sigma^{\frac{1}{2}}\sigma'} N(r)$$

Metric tensor and conjugate momentum

$$\dot{\alpha} = 0 \quad c_2 = -2\dot{A}$$



Assumptions and jump conditions

Integration Constants

$$A_{\text{in}} \ A_{\text{out}} \ \alpha_{\text{in}} \ \alpha_{\text{out}} \ C_{1\text{in}} \ C_{1\text{out}}$$

Assumptions

- ▶ Continuity of the metric tensor
- ▶ No mass inside the shell
- ▶ Asymptotic flatness at infinity
- ▶ Continuity of the Lapse
- ▶ $N = 1$ at infinity

Jump Conditions

- ▶ Diffeo constraint
- ▶ Hamiltonian constraint
- ▶ LFE
- ▶ EoM for p_{ij}

$$A_{\text{in}} \ A_{\text{out}} \ \alpha_{\text{out}} \ C_{1\text{out}} \ \alpha_{\text{in}} = 0$$



Assumptions and jump conditions

Integration Constants

$$A_{\text{in}} \ A_{\text{out}} \ \alpha_{\text{in}} \ \alpha_{\text{out}} \ C_{1\text{in}} \ C_{1\text{out}}$$

Assumptions

- ▶ Continuity of the metric tensor
- ▶ No mass inside the shell
- ▶ Asymptotic flatness at infinity
- ▶ Continuity of the Lapse
- ▶ $N = 1$ at infinity

Jump Conditions

- ▶ Diffeo constraint
- ▶ Hamiltonian constraint
- ▶ LFE
- ▶ EoM for p_{ij}

$$A_{\text{in}} \ A_{\text{out}} \ \alpha_{\text{out}} \ C_{1\text{out}} \ \alpha_{\text{in}} = 0$$



Reduced Phase Space

On-Shell Relation

$$(\alpha_{\text{out}} + \rho)A_{\text{in}}^2 + \rho A_{\text{out}}^2 + \left(\frac{M^2}{16\rho} - \alpha_{\text{out}} - 2\rho\right)A_{\text{in}}A_{\text{out}} - \rho^3(\alpha_{\text{out}})^2 - \frac{M^2\rho^2}{8}\left(\frac{M^2}{32\rho} - \alpha_{\text{out}} - 2\rho\right) = 0$$

Independent Variables

- ▶ $\rho = \sigma(R) = g_{\theta\theta}(R)$ Area of the Shell
- ▶ $A_{\text{in}} A_{\text{out}}$ Related to the Momentum P
- ▶ α_{out} Related to the ADM mass

UNDERDETERMINED SYSTEM

The on-shell relation and the condition $\alpha_{\text{out}} = \text{const.}$ define a 2-dimensional manifold - not a one dimensional curve

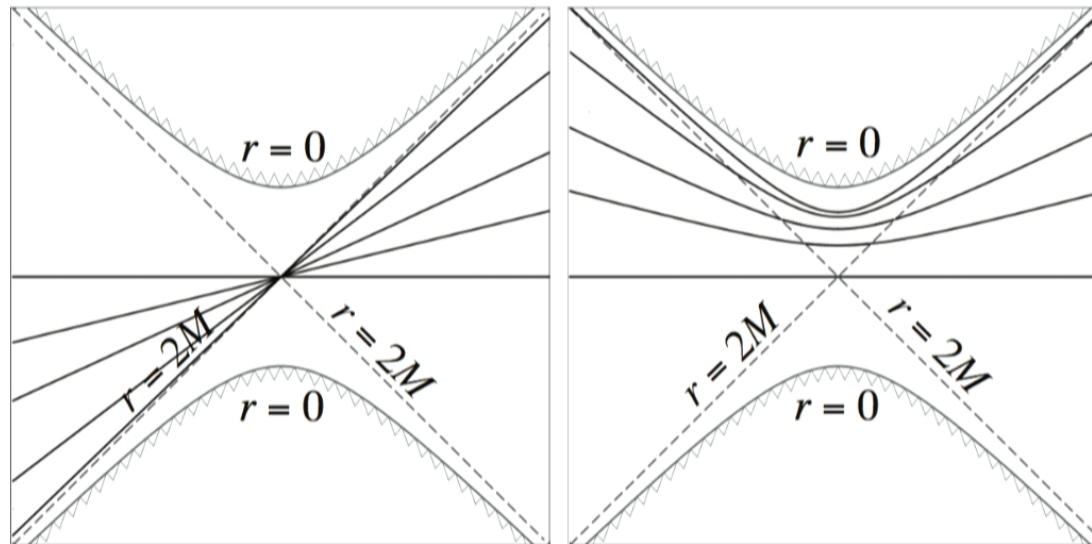


What is the under determination?

- ▶ Spherically symmetric thin shell
- ▶ Conformal constraint $g_{ij}p^{ij} = 0$
- ▶ General result: Birkhoff Theorem

Schwarzschild space-time in maximal foliation

MAXIMAL FOLIATION OF SCHWARZCHILD IS NOT UNIQUE



"3+1 Formalism in General Relativity" E.ourgoulhon 

GR point of view

Isotropic static foliation $A_{\text{out}} = 0$

$$ds^2 = \left(\frac{1 - \frac{\alpha}{4r}}{1 + \frac{\alpha}{4r}} \right)^2 dt^2 + \left(1 + \frac{\alpha}{4r} \right)^4 (dr^2 + r^2 d\Omega_2)$$

C foliation $C \leftrightarrow A_{\text{out}}$

$$ds^2 = f(C(\tau))\dot{C}(\tau)^2 d\tau^2 + g(C(\tau))\dot{C}(\tau) d\tau dr + \frac{1}{1 + \frac{\alpha}{y} + \frac{C(\tau)^2}{4y^2}} dy^2 + y^2 d\Omega_2^2.$$

A 4-dimensional diffeomorphism connects the two solutions
THERE IS NO AMBIGUITY IN GR



Why then the ambiguity in SD?

- ▶ SD rests on a preferred notion of simultaneity
- ▶ SD symmetries are 3d diffeomorphisms and 3d conformal transformations
- ▶ The 2 maximal foliations of Schwarzschild are not connected by a 3d conformodiffeo

What fixes A_{out} ?

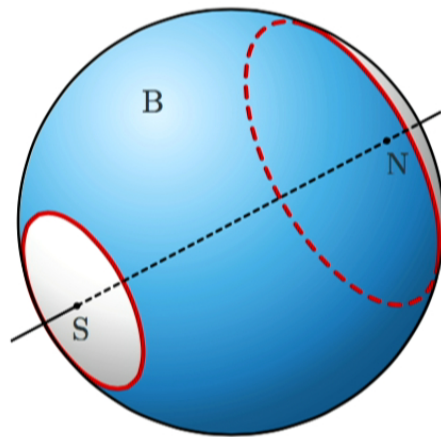
The rest of the universe

Twin Shell Model

SD is naturally defined in a compact universe

A spherically symmetric compact universe with S^3 topology is the simplest model

$$ds^2 = d\psi^2 + \sin^2 \psi (d\theta^2 + \sin^2 \theta d\phi^2)$$



It requires at least two thin shell to not have any mass at the origin



The twin shell model: Starting point

Ansatz

$$g_{ij} = \text{diag} \{ \mu^2, \sigma, \sigma \sin^2 \theta \} \quad p^{ij} = \text{diag} \left\{ \frac{f}{\mu}, \frac{1}{2} s, \frac{1}{2} s \sin^{-2} \theta \right\} \sin \theta \quad \xi^i = \{ \xi, 0, 0 \}$$

Hamiltonian Constraint

$$\int d\theta d\phi \left[\frac{p_{ij}^T p_{ij}^T}{\sqrt{g}} - \frac{1}{6} (\langle p \rangle^2 - 12\Lambda) \sqrt{g} - \sqrt{g} R \right] + 4\pi \sum_{a \in \{S, N\}} \delta(\psi - \Psi_a) \sqrt{g^{rr} P_a^2 + M_a^2} = 0$$

Diffeo constraint

$$-2 \int d\theta d\phi \nabla_j p^j_i - 4\pi \delta^{\psi}_i \sum_{a \in \{S, N\}} \delta(\psi - \Psi_a) P_a = 0$$

Conformal constraint: CMC Foliation

$$g_{ij} p^{ij} - \sqrt{g} \langle p \rangle = 0$$



Solutions to EoM and OS Relation

Diffeo and Hamiltonian constraints

$$f = \frac{1}{3} \langle p \rangle \sigma + \frac{A}{\sigma^2} \quad \mu^2 = \frac{(\sigma')^2}{\left(\frac{2}{3} A \langle p \rangle + 4\alpha\right) \sigma^2 + 4\sigma + \frac{A^2}{\sigma} + \frac{1}{9} (\langle p \rangle^2 - 12\Lambda) \sigma^2}$$

We get two on shell relations

$$\frac{M_S^4}{16\rho_S^2} - M_S^2 \left(\frac{A_B A_S}{\rho_S^4} + 4 - T \rho_S^2 \right) - 4(A_B - A_S)^2 \left(\frac{16}{\rho_S^2} - 4T \right) - \left(\frac{4}{\rho_S^3} A_S (A_S - A_B) + \frac{M_S^2}{2\rho_S} \right) X + X^2 = 0$$

$$\frac{M_N^4}{16\rho_N^2} - M_N^2 \left(\frac{A_B A_N}{\rho_N^4} + 4 - T \rho_N^2 \right) - 4(A_B - A_N)^2 \left(\frac{16}{\rho_N^2} - 4T \right) - \left(\frac{4}{\rho_N^3} A_N (A_N - A_B) + \frac{M_N^2}{2\rho_N} \right) X + X^2 = 0$$

$$\rho^2 = \sigma(\psi) \quad X = \left(\frac{2}{3} A_B \langle p \rangle + 4\alpha_B \right) \quad T = \frac{1}{9} (12\Lambda - \langle p \rangle^2)$$



Twin Shell DoF

- ▶ $\rho_S \rho_N$
 - ▶ $A_S A_N A_B$
 - ▶ $\alpha_S \alpha_N \alpha_B$
 - ▶ $\langle p \rangle V$
- ▶ 2 on shell relations
 - ▶ $\alpha_S = 0 \alpha_N = 0; \alpha_B = \text{const}$

4 matter DoF (2 for each shell)
2 scale DoF (volume and momentum)

THE SYSTEM IS FULLY DETERMINED

Towards Asymptotic Flatness

- ▶ Take the limit $A_N \rightarrow \infty$
- ▶ Set to 0 each of the coefficients of A_N
- ▶ 3 equations that can be solved in terms of ρ_n α_B A_B

2 possible solutions

$$\begin{aligned} \rho_N &= \sqrt{\frac{4 - \sqrt{16 - M_N^2 T}}{2T}}, & \alpha_B &= -\sqrt{\frac{4 - \sqrt{16 - M_N^2 T}}{2T}} \left(\frac{1}{8} \sqrt{16 - M_N^2 T} + \frac{1}{2} \right), & A_B &= 0 \\ \rho_N &= \sqrt{\frac{\sqrt{16 - M_N^2 T} + 4}{2T}}, & \alpha_B &= \left(\frac{1}{8} \sqrt{16 - M_N^2 T} - \frac{1}{2} \right) \sqrt{\frac{\sqrt{16 - M_N^2 T} + 4}{2T}}, & A_B &= 0 \end{aligned}$$

Asymptotic flatness becomes a good approximation for $T \rightarrow 0$

$$V \propto \frac{1}{T} \Leftrightarrow \lim_{T \rightarrow 0} V = \infty$$



Towards Asymptotic Flatness

- ▶ Take the limit $A_N \rightarrow \infty$
- ▶ Set to 0 each of the coefficients of A_N
- ▶ 3 equations that can be solved in terms of ρ_n α_B A_B

2 possible solutions

$$\begin{aligned} \rho_N &= \sqrt{\frac{4 - \sqrt{16 - M_N^2 T}}{2T}}, & \alpha_B &= -\sqrt{\frac{4 - \sqrt{16 - M_N^2 T}}{2T}} \left(\frac{1}{8} \sqrt{16 - M_N^2 T} + \frac{1}{2} \right), & A_B &= 0 \\ \rho_N &= \sqrt{\frac{\sqrt{16 - M_N^2 T} + 4}{2T}}, & \alpha_B &= \left(\frac{1}{8} \sqrt{16 - M_N^2 T} - \frac{1}{2} \right) \sqrt{\frac{\sqrt{16 - M_N^2 T} + 4}{2T}}, & A_B &= 0 \end{aligned}$$

Asymptotic flatness becomes a good approximation for $T \rightarrow 0$

$$V \propto \frac{1}{T} \Leftrightarrow \lim_{T \rightarrow 0} V = \infty$$



Deriving $A_{\text{out}} = 0$

Late time limit $T \rightarrow 0$

$$\rho_N \rightarrow \frac{M_N}{4}$$

$$\alpha_B \rightarrow -\frac{M_N}{4}$$

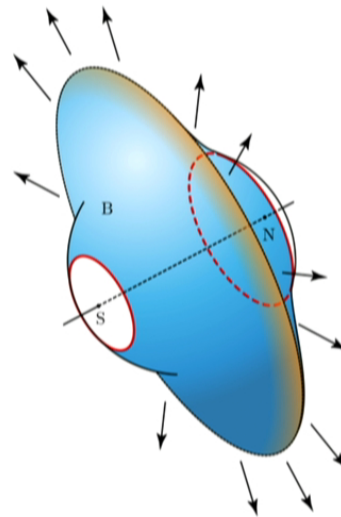
$$A_B \rightarrow 0$$

$$\rho_N \sim \frac{2}{\sqrt{T}} \rightarrow \infty$$

$$\alpha_B \sim -\frac{M_N^2}{32} \sqrt{T} \rightarrow 0^-$$

$$A_B \rightarrow 0$$

$$\rho_N \rightarrow \frac{M_N}{4} \quad A_{\text{out}} = A_B = 0 \quad \alpha_{\text{out}} = \alpha_B = -\frac{M_N}{4}$$



Deriving $A_{\text{out}} = 0$

Late time limit $T \rightarrow 0$

$$\rho_N \rightarrow \frac{M_N}{4}$$

$$\alpha_B \rightarrow -\frac{M_N}{4}$$

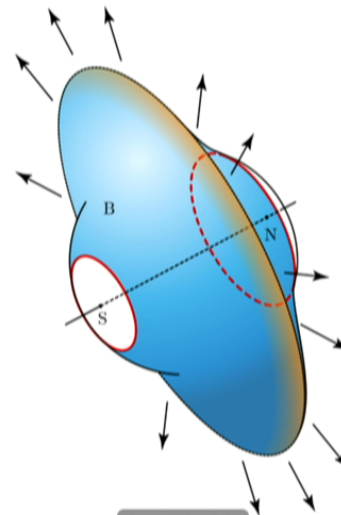
$$A_B \rightarrow 0$$

$$\rho_N \sim \frac{2}{\sqrt{T}} \rightarrow \infty$$

$$\alpha_B \sim -\frac{M_N^2}{32} \sqrt{T} \rightarrow 0^-$$

$$A_B \rightarrow 0$$

$$\rho_N \rightarrow \frac{M_N}{4} \quad A_{\text{out}} = A_B = 0 \quad \alpha_{\text{out}} = \alpha_B = -\frac{M_N}{4}$$



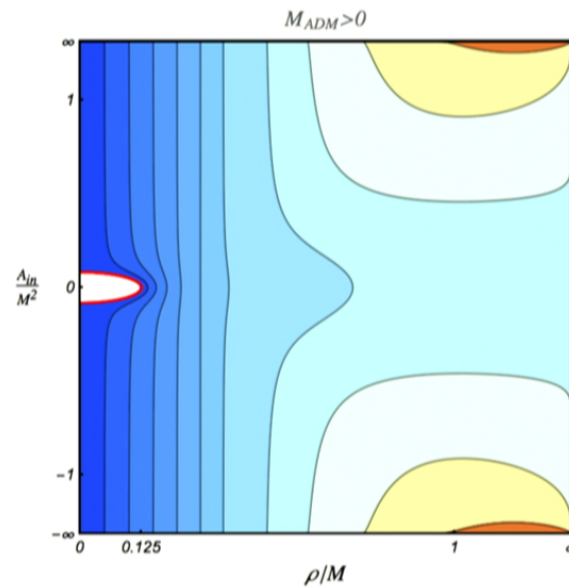
Page 18 of 25



Reduced phase space $A_{\text{out}} = 0$

On Shell Relation for the Thin Shell

$$(\alpha_{\text{out}} + \rho)A_{\text{in}}^2 - \rho^3(\alpha_{\text{out}})^2 - \frac{M^2\rho^2}{8} \left(\frac{M^2}{32\rho} - \alpha_{\text{out}} - 2\rho \right) = 0$$



A_{out} diverges when the dynamics freezes or the shell escapes

Thin Shell Symplectic Reduction

Symplectic potential

$$\theta = \int dr d\theta d\phi p^i \delta g_{ij} + 4\pi P \delta R$$
$$- 8\pi \left(\delta A_{\text{in}} \int_0^R dr \frac{\mu}{\sqrt{\sigma}} + \delta A_{\text{out}} \int_R^\infty dr \frac{\mu}{\sqrt{\sigma}} \right) + 4\pi P \delta R$$

Symplectic form

$$\omega = \delta\theta = -8\pi \left(\delta A_{\text{in}} \frac{\mu(R)}{\sqrt{\sigma(R)}} - \delta A_{\text{out}} \frac{\mu(R)}{\sqrt{\sigma(R)}} \right) \wedge \delta R,$$

$$\boxed{\omega = 4\pi \delta P \wedge \delta R}$$

P and R are the conjugate dynamical variables



ADM Mass

Definition

$$M_{\text{ADM}} = \frac{1}{16\pi} \lim_{r \rightarrow \infty} \int d\theta d\phi \left(\bar{\nabla}^j g_{rj} - \bar{\nabla}_r \bar{g}^{ij} g_{ij} \right) r^2 \sin \theta$$

$$M_{\text{ADM}} = \frac{1}{2} \lim_{r \rightarrow \infty} \left(r \mu^2 - \sigma' + \frac{1}{r} \sigma \right)$$

Assumption: Asymptotic flatness $\lim_{r \rightarrow \infty} \sigma = r^2$

$$M_{\text{ADM}} = -\frac{\alpha_{\text{out}}}{2}$$

α_{out} plays the role of the ADM Hamiltonian



Thin Shell $A_{\text{out}} = 0$: Isotropic Gauge

Isotropic gauge condition & $A_{\text{out}} = 0$

$$\frac{\mu^2 r}{\sigma(r)} = \frac{1}{r^2} \Rightarrow \sigma(r) = r^2 \left(1 - \frac{\alpha_{\text{out}}}{4r}\right)^4$$

EoM in terms of R and P : $\alpha_{\text{out}}(R, P)$

- ▶ On shell relation

$$(\alpha_{\text{out}} + \rho)A_{\text{in}}^2 - \rho^3(\alpha_{\text{out}})^2 - \frac{M^2 \rho^2}{8} \left(\frac{M^2}{32\rho} - \alpha_{\text{out}} - 2\rho \right) = 0$$

- ▶ Isotropic relation

$$\rho^2 = R^2 \left(1 - \frac{\alpha_{\text{out}}}{4R}\right)^4$$

- ▶ Diffeo jump condition

$$A_{\text{in}} - \frac{RP}{2} = 0$$

The solution of this system $\alpha_{\text{out}} = \alpha_{\text{out}}(R, P)$



Thin Shell EoM

$$(\dots)d\alpha_{\text{out}} + (\dots)dA_{\text{in}} + (\dots)d\rho + (\dots)dR + (\dots)dP = 0$$

$$(\dots)d\alpha_{\text{out}} + (\dots)dA_{\text{in}} + (\dots)d\rho + (\dots)dR + (\dots)dP = 0$$

$$(\dots)d\alpha_{\text{out}} + (\dots)dA_{\text{in}} + (\dots)d\rho + (\dots)dR + (\dots)dP = 0$$

$$d\alpha_{\text{out}} = \left(\frac{\partial \alpha_{\text{out}}}{\partial R} \Big|_{A_{\text{in}} = \frac{R^2 P}{2}, \rho = R \left(1 - \frac{\alpha_{\text{out}}}{4R}\right)^2} \right) dR + \left(\frac{\partial \alpha_{\text{out}}}{\partial P} \Big|_{A_{\text{in}} = \frac{R^2 P}{2}, \rho = R \left(1 - \frac{\alpha_{\text{out}}}{4R}\right)^2} \right) dP$$

$$dA_{\text{in}} = \left(\frac{\partial A_{\text{in}}}{\partial R} \Big|_{A_{\text{in}} = \frac{R^2 P}{2}, \rho = R \left(1 - \frac{\alpha_{\text{out}}}{4R}\right)^2} \right) dR + \left(\frac{\partial A_{\text{in}}}{\partial P} \Big|_{A_{\text{in}} = \frac{R^2 P}{2}, \rho = R \left(1 - \frac{\alpha_{\text{out}}}{4R}\right)^2} \right) dP$$

$$d\rho = \left(\frac{\partial \rho}{\partial R} \Big|_{A_{\text{in}} = \frac{R^2 P}{2}, \rho = R \left(1 - \frac{\alpha_{\text{out}}}{4R}\right)^2} \right) dR + \left(\frac{\partial \rho}{\partial P} \Big|_{A_{\text{in}} = \frac{R^2 P}{2}, \rho = R \left(1 - \frac{\alpha_{\text{out}}}{4R}\right)^2} \right) dP$$

$$\dot{R} = -\frac{1}{2} \frac{d\alpha_{\text{out}}}{dP}(R, P, \alpha_{\text{out}}) \quad \dot{P} = +\frac{1}{2} \frac{d\alpha_{\text{out}}}{dR}(R, P, \alpha_{\text{out}})$$

Thin Shell Numerical Simulations

