

Title: Quantum theory and spacetime: allies, not enemies

Date: May 12, 2015 09:50 AM

URL: <http://pirsa.org/15050078>

Abstract: It has become conventional wisdom to say that quantum theory and gravitational physics are conceptually so different, if not incompatible, that it is very hard to unify them. However, in the talk I will argue that the operational view of (quantum) information theory adds a very different twist to this picture: quite on the contrary, quantum theory and space-time are highly fine-tuned to fit to each other.

After a recap of ideas by von Weizsäcker, Wootters, and Popescu and Rohrlich, I will show how uncertainty relations, the number of degrees of freedom of the Bloch ball, and the existence of entangled states and possibly the Tsirelson bound can be understood from space-time geometry alone. Conversely, I will show how the 3+1 Lorentz group of space-time can be derived from a purely informational communication scenario of two observers that describe local quantum physics in different Hilbert space bases (joint work with Philipp Hoehn).

Folk wisdom:

Quantum theory and general relativity **hard to unify**
because they are so different / incompatible:



Folk wisdom:

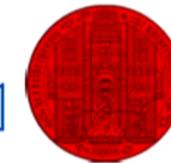
Quantum theory and general relativity **hard to unify**
because they are so different / incompatible:

QM

Probabilistic, usually
fixed background, ...

GR

Deterministic, background-
independent, ...



Folk wisdom:

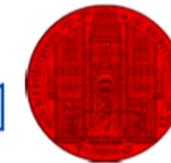
Quantum theory and general relativity **hard to unify**
because they are so different / incompatible:

QM

Probabilistic, usually
fixed background, ...

GR

Deterministic
non-renormalizable
background,
...
\$



Folk wisdom:

Quantum theory and general relativity **hard to unify**
because they are so different / incompatible:

QM

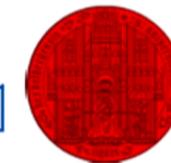
Probabilistic, usually
fixed background, ...

GR

Deterministic
non-renormalizable
background, ...



This talk: quantum information adds a different
facet to this picture...



This talk



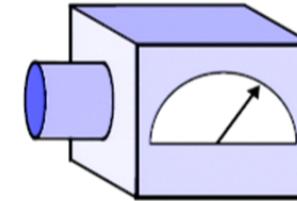
QT space-time

They mutually constrain (part of) each other's structure.
They are "strange allies".

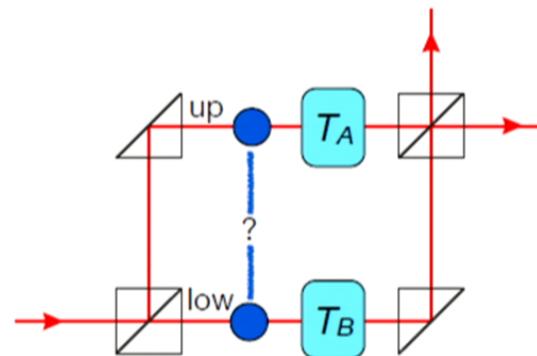


Outline

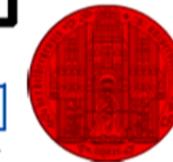
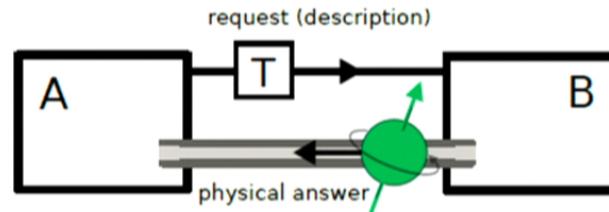
1. A glance beyond quantum theory



2. Relativity of simultaneity
on an interferometer



3. SO(3,1) and quantum communication



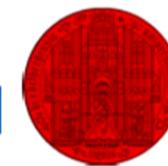
1. A glance beyond QT

Goal: argue counterfactually. "What if **XY** was different?"



QT

space-time



1. A glance beyond QT

Goal: argue counterfactually. "What if **XY** was different?"



QT

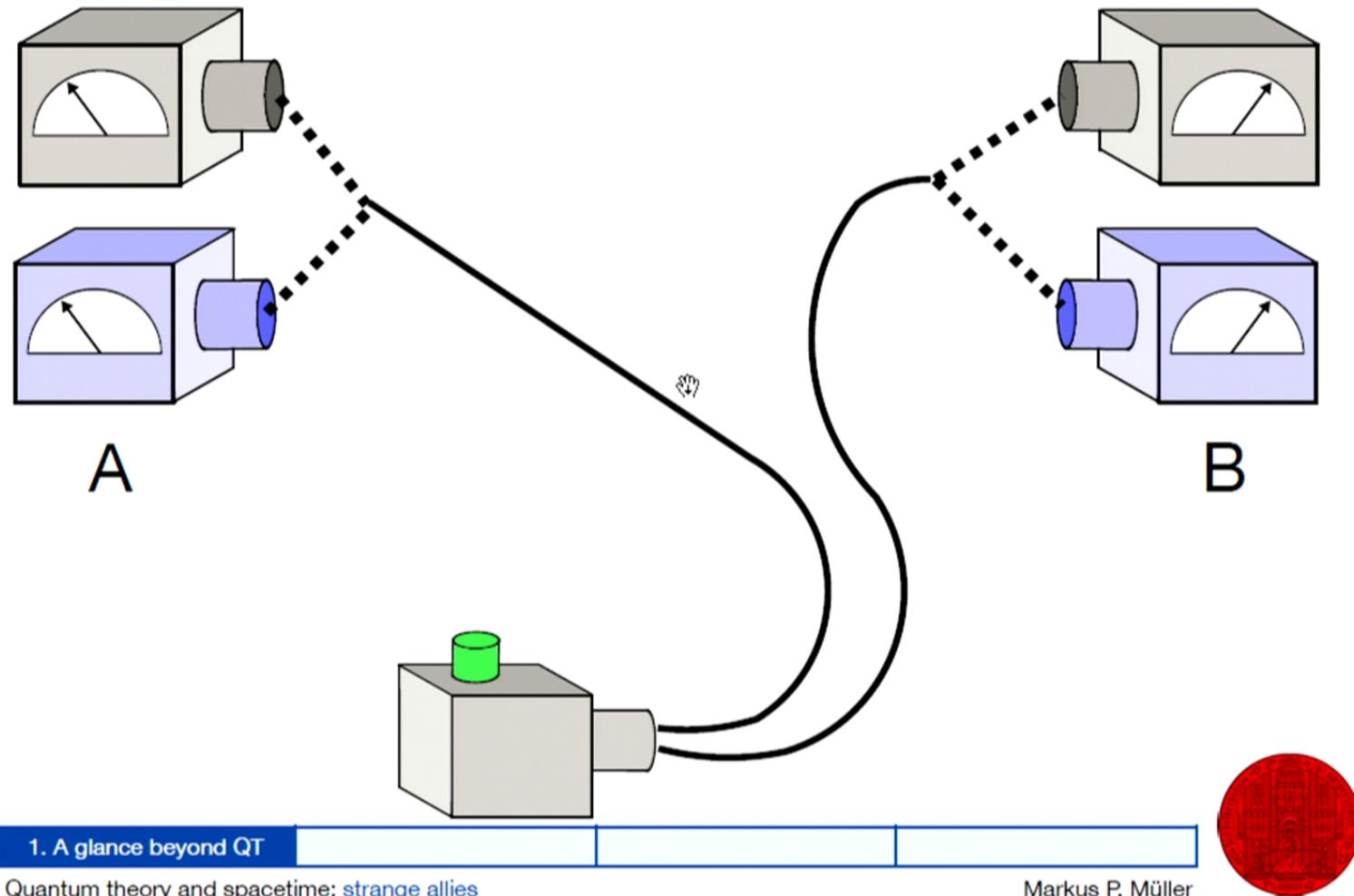
space-time

?

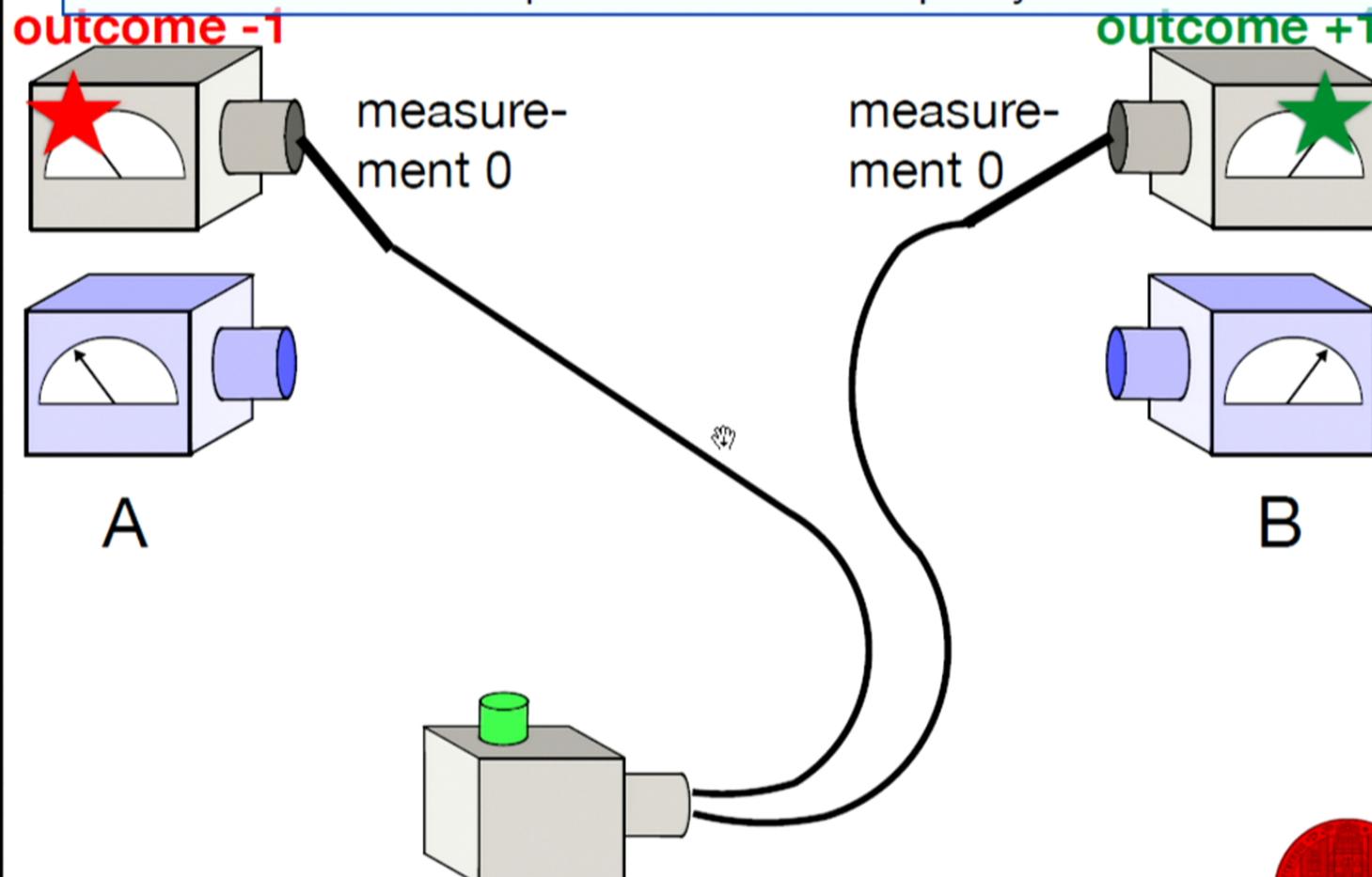
Different dimensionality,
signature, torsion,
discrete graph etc.



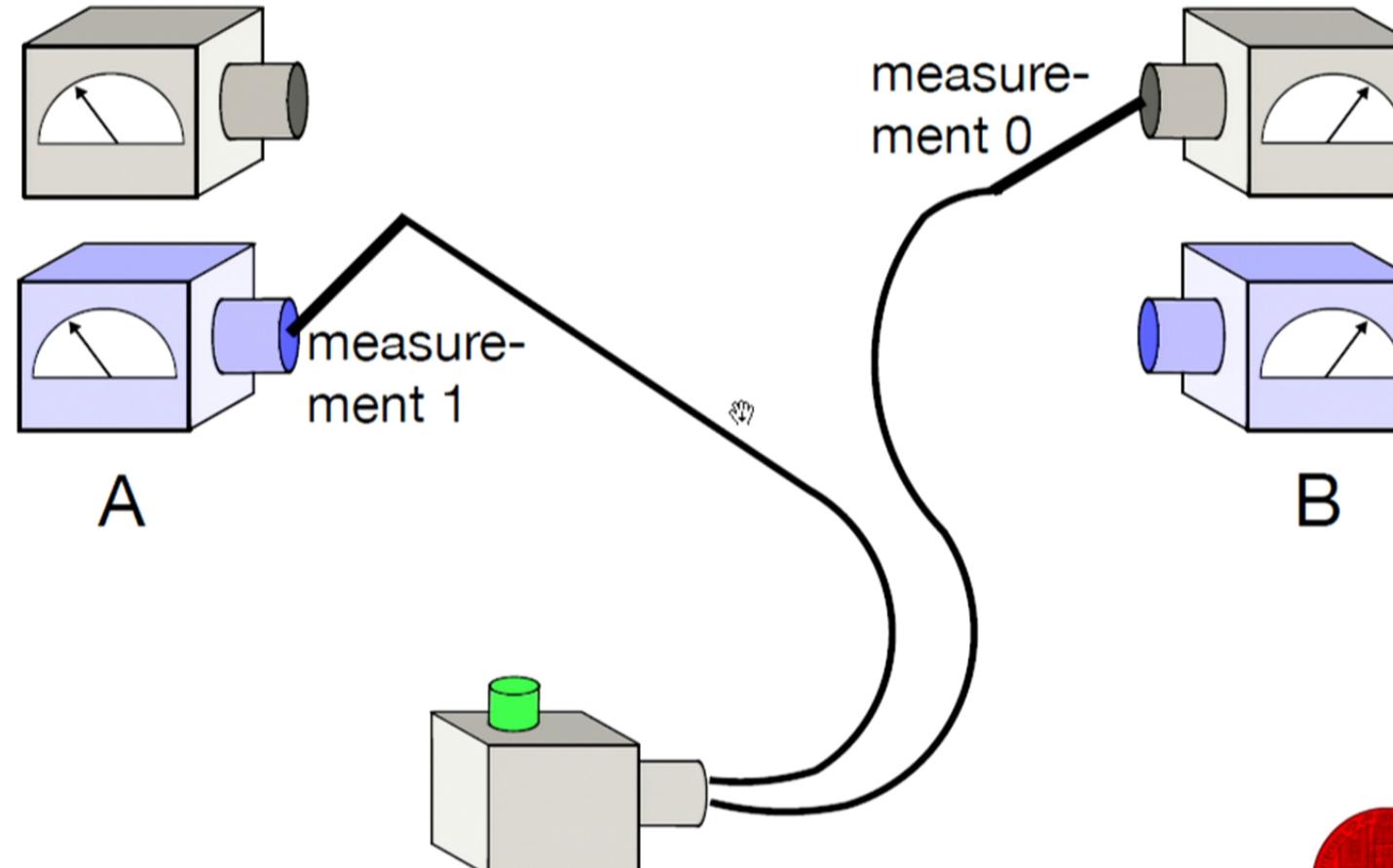
Example: Bell-CHSH inequality



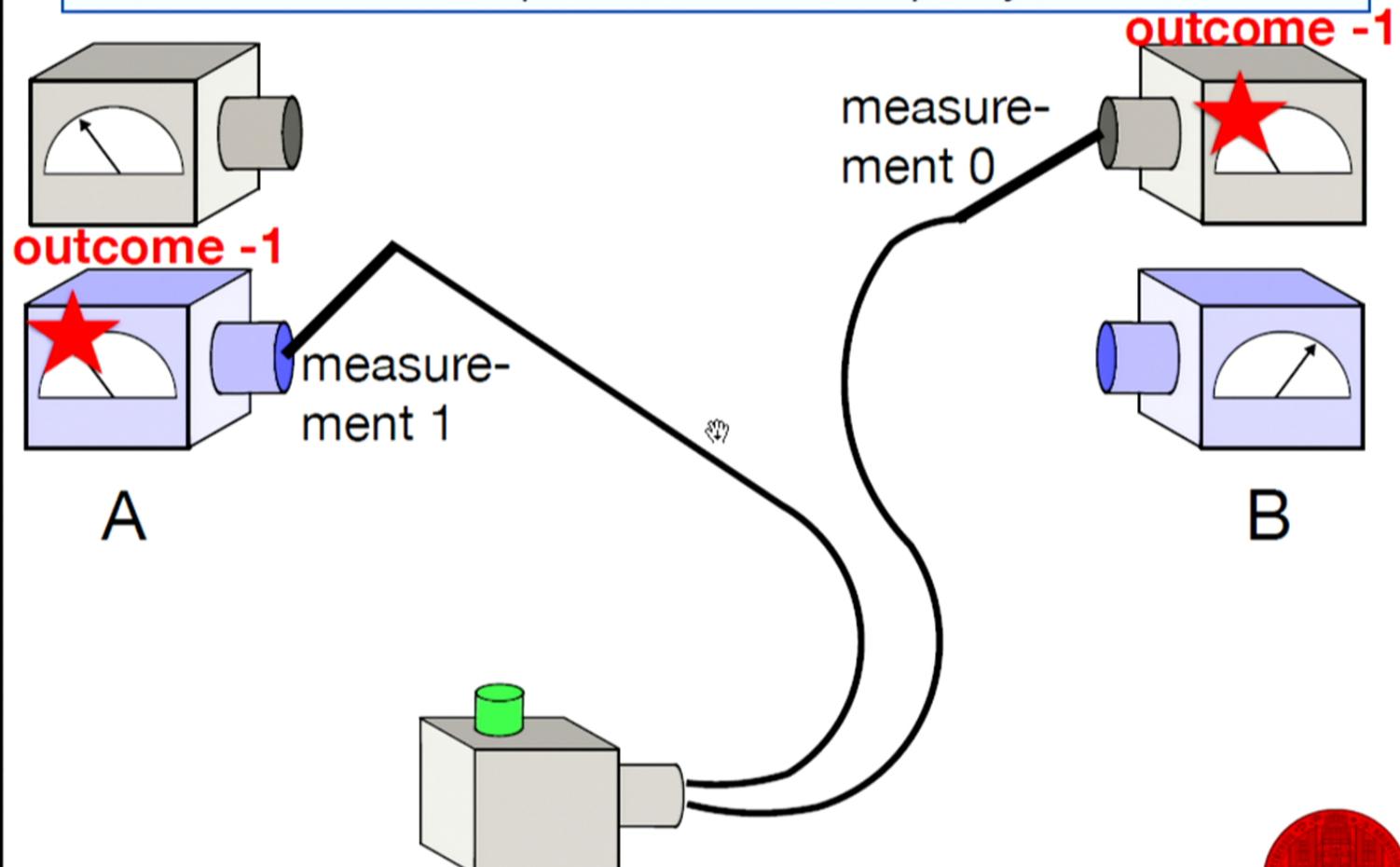
Example: Bell-CHSH inequality



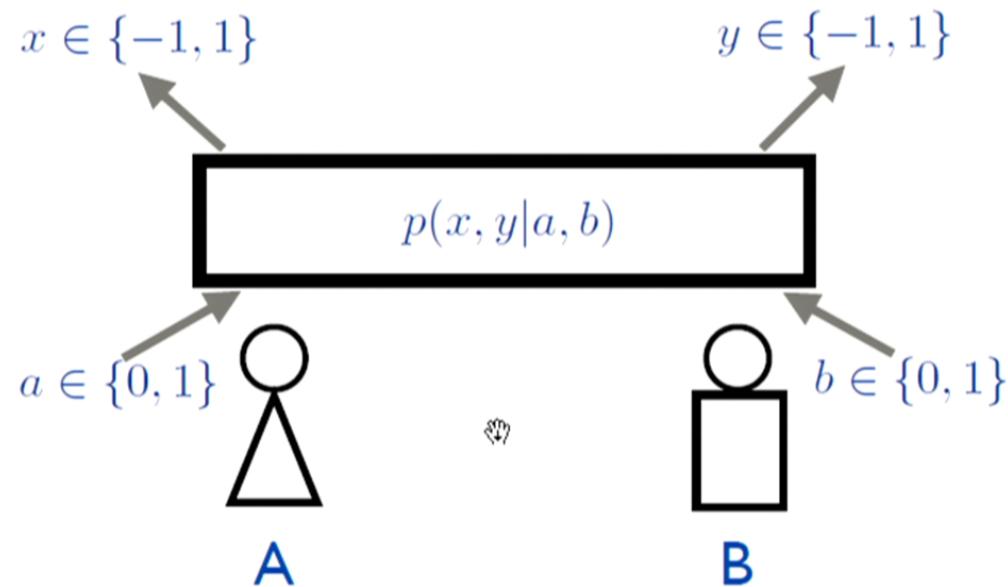
Example: Bell-CHSH inequality



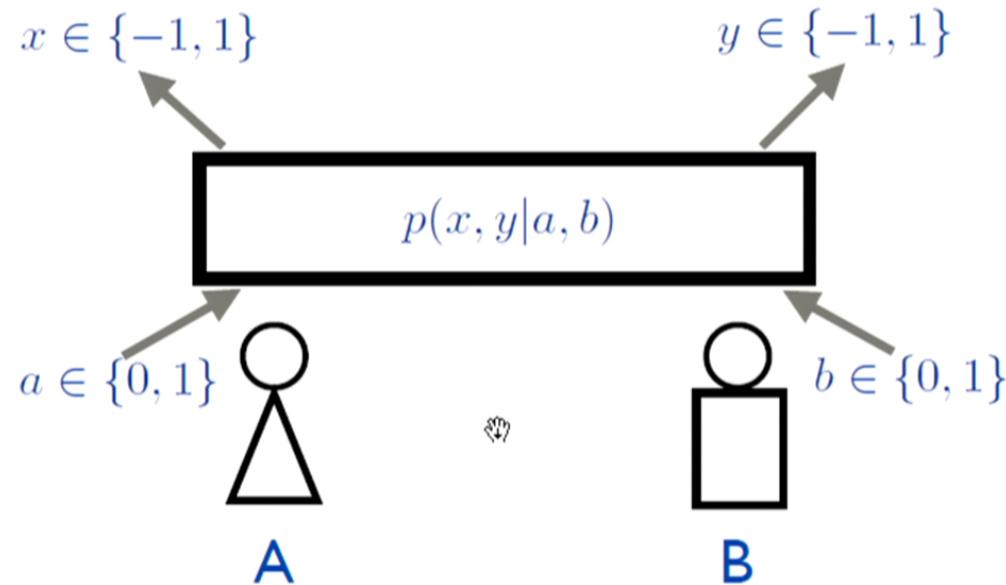
Example: Bell-CHSH inequality



Example: Bell-CHSH inequality



Example: Bell-CHSH inequality

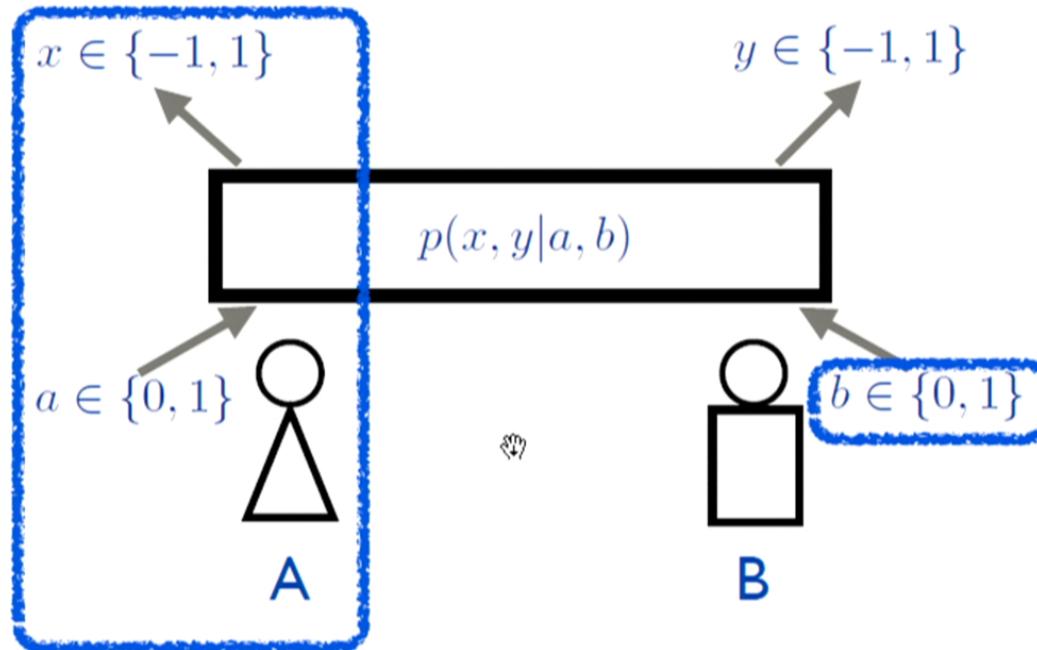


Quantum physics: $p(x, y|a, b) = \langle \psi | \hat{P}_a^x \otimes \hat{P}_b^y | \psi \rangle$.

Classical physics: $p(x, y|a, b)$ conditional prob.

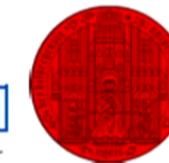


Example: Bell-CHSH inequality

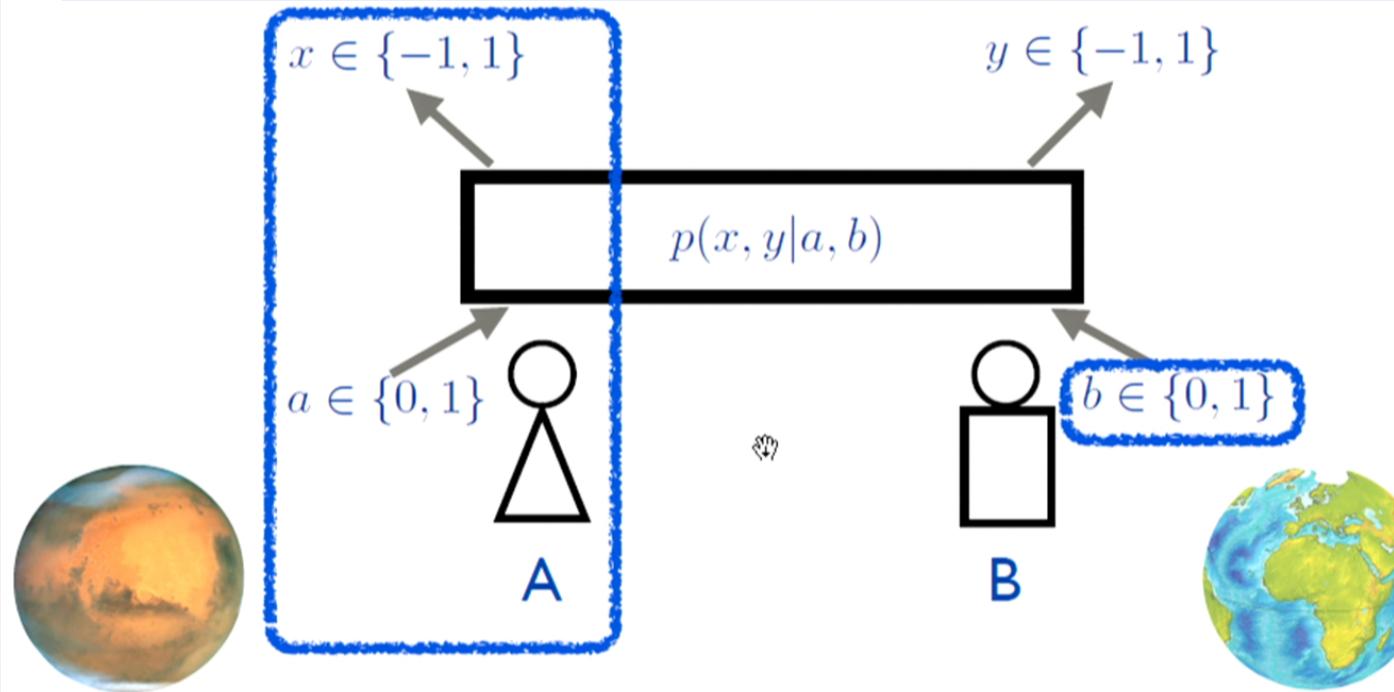


Classical and quantum physics satisfy **no-signalling**:

$p(x|a)$ does not depend on b (and vice versa).

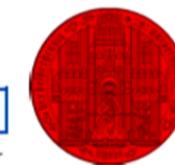


Example: Bell-CHSH inequality



Classical and quantum physics satisfy **no-signalling**:

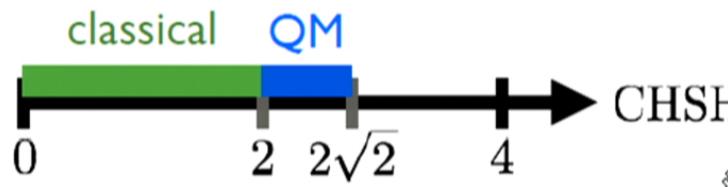
$p(x|a)$ does not depend on b (and vice versa).



Example: Bell-CHSH inequality

Classical probability distributions satisfy Bell inequality:

$$\text{CHSH} := |C_{00} + C_{01} + C_{10} - C_{11}| \leq 2 \quad \text{where } C_{ab} := \mathbb{E}(x \cdot y|a, b).$$



Quantum: Bell inequality violation.

$$\text{CHSH} \leq 2\sqrt{2}.$$

S. Popescu and D. Rohrlich, Found. Phys. 24, 379 (1994).

In 1994, Popescu and Rohrlich asked:

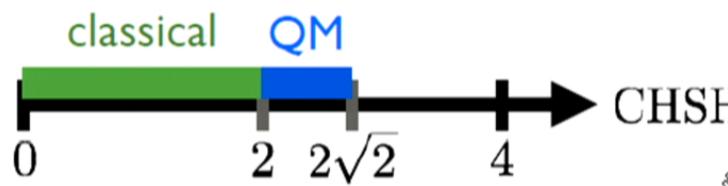
Are quantum correlations the most general correlations that are non-local, but still fit into relativistic spacetime?



Example: Bell-CHSH inequality

Classical probability distributions satisfy Bell inequality:

$$\text{CHSH} := |C_{00} + C_{01} + C_{10} - C_{11}| \leq 2 \quad \text{where } C_{ab} := \mathbb{E}(x \cdot y|a, b).$$



Quantum: Bell inequality violation.

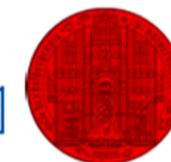
$$\text{CHSH} \leq 2\sqrt{2}.$$

S. Popescu and D. Rohrlich, Found. Phys. 24, 379 (1994).

In 1994, Popescu and Rohrlich asked:

Are quantum correlations the most general correlations that are non-local, but still fit into relativistic spacetime?

Answer: **No.** There are ("PR box") correlations that are non-signalling, but not possible in quantum theory.



Example: Bell-CHSH inequality

Classical probability distributions satisfy Bell inequality:

$$\text{CHSH} := |C_{00} + C_{01} + C_{10} - C_{11}| \leq 2 \quad \text{where } C_{ab} := \mathbb{E}(x \cdot y|a, b).$$



PR-box correlations: $p(x, y|a, b)$

$$p(+1, +1|a, b) = p(-1, -1|a, b) = \frac{1}{2} \quad \text{if } (a, b) \in \{(0, 0), (0, 1), (1, 0)\}$$

$$p(+1, -1|1, 1) = p(-1, +1|1, 1) = \frac{1}{2}$$



Example: Bell-CHSH inequality

Classical probability distributions satisfy Bell inequality:

$$\text{CHSH} := |C_{00} + C_{01} + C_{10} - C_{11}| \leq 2 \quad \text{where } C_{ab} := \mathbb{E}(x \cdot y|a, b).$$



PR-box correlations: $p(x, y|a, b)$ are non-signalling, and have

$$p(+1, +1|a, b) = p(-1, -1|a, b) = \frac{1}{2} \quad \text{CHSH=4.}$$

if $(a, b) \in \{(0, 0), (0, 1), (1, 0)\}$

$$p(+1, -1|1, 1) = p(-1, +1|1, 1) = \frac{1}{2}$$



Example: Bell-CHSH inequality

Classical probability distributions satisfy Bell inequality:

$$\text{CHSH} := |C_{00} + C_{01} + C_{10} - C_{11}| \leq 2 \quad \text{where } C_{ab} := \mathbb{E}(x \cdot y|a, b).$$



PR-box correlations: $p(x, y|a, b)$ are non-signalling, and have

$$p(+1, +1|a, b) = p(-1, -1|a, b) = \frac{1}{2} \quad \text{CHSH=4.}$$

if $(a, b) \in \{(0, 0), (0, 1), (1, 0)\}$

$$p(+1, -1|1, 1) = p(-1, +1|1, 1) = \frac{1}{2}$$

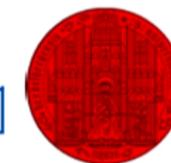


Conclusion part 1



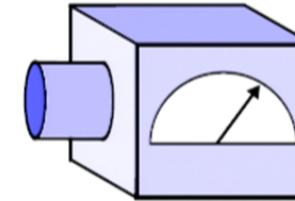
QT ← space-time

Space-time structure enforces no-signalling,
but does not fully determine all of QT

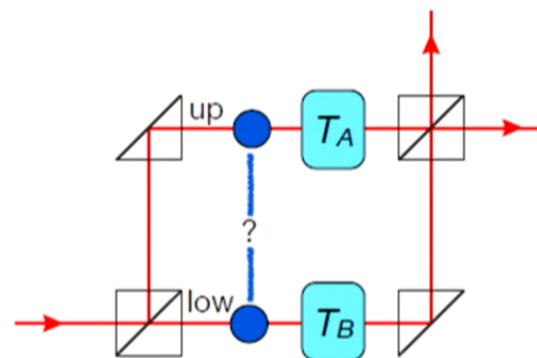


Outline

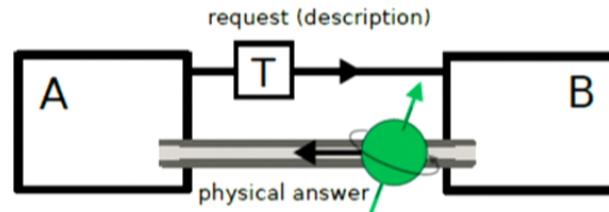
1. A glance beyond quantum theory



2. Relativity of simultaneity
on an interferometer



3. SO(3,1) and quantum communication



1. A glance beyond QT

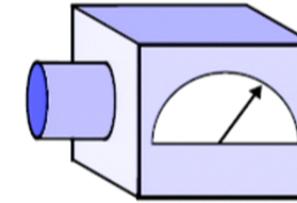
Quantum theory and spacetime: strange allies

Markus P. Müller

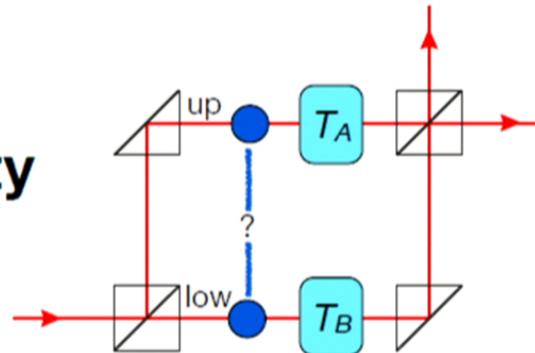


Outline

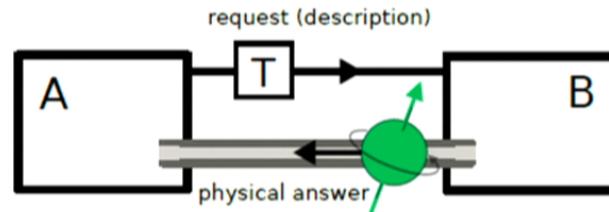
1. A glance beyond quantum theory



**2. Relativity of simultaneity
on an interferometer**



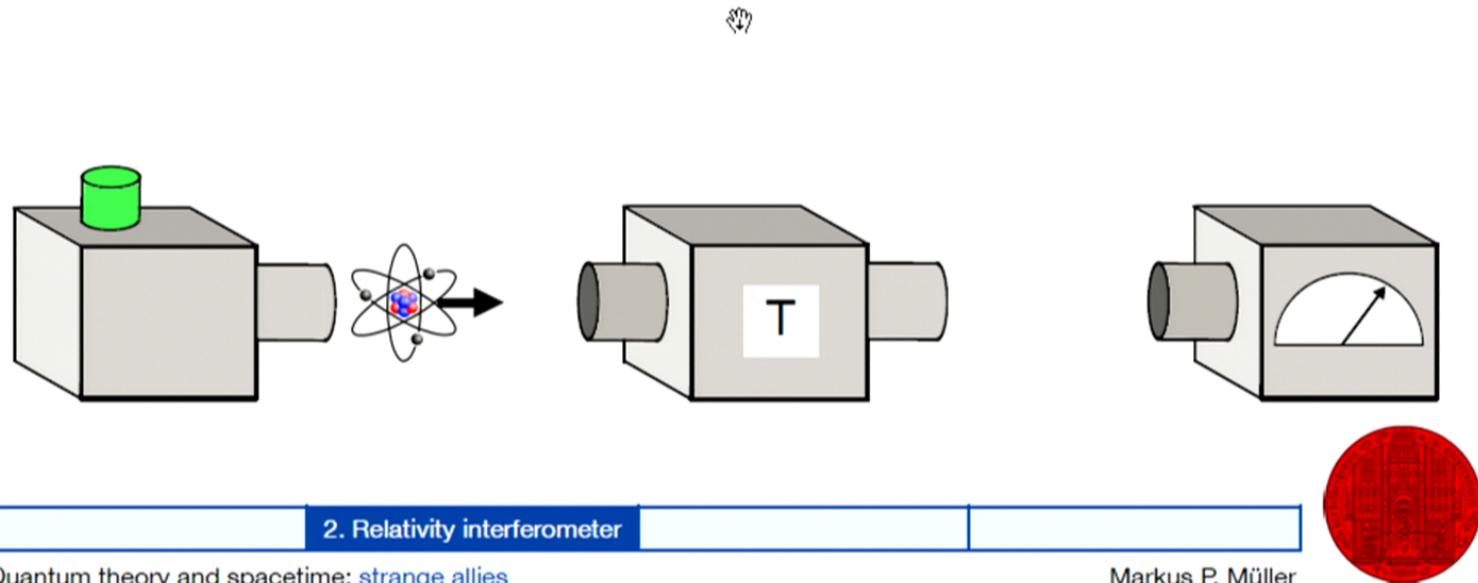
3. SO(3,1) and quantum communication



General state spaces

Preparation, transformation, measurement

- On every push of button, the preparation device outputs a physical system in some **state**.

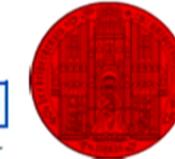
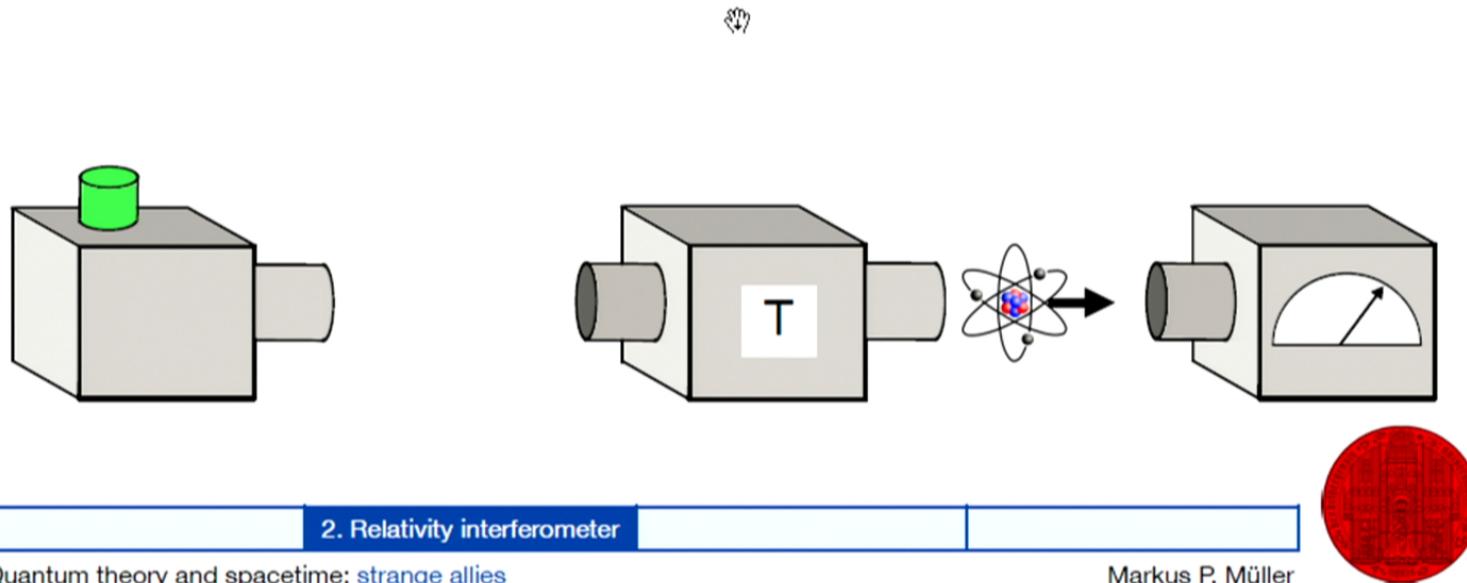


2. Relativity interferometer

General state spaces

Preparation, transformation, measurement

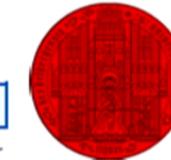
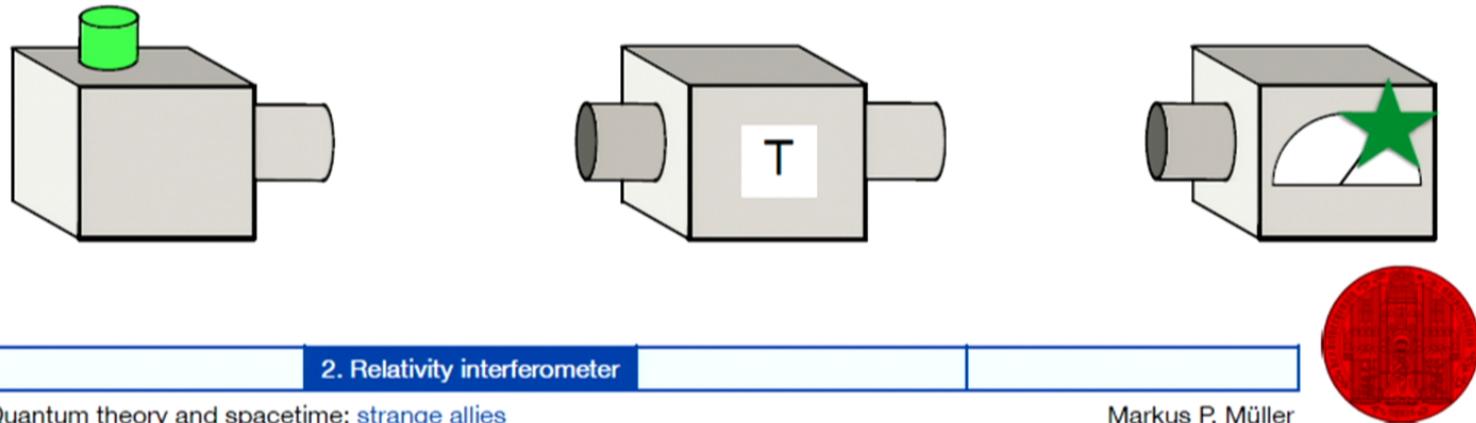
- On every push of button, the preparation device outputs a physical system in some **state**.
- The transformation device modifies the state. Convex-linearity: statistical mixtures are respected.



General state spaces

Preparation, transformation, measurement

- On every push of button, the preparation device outputs a physical system in some **state**.
- The transformation device modifies the state. Convex-linearity: statistical mixtures are respected.
- One of many possible **measurements** is performed. A definite classical outcome is obtained, with prob. **determined by state**.



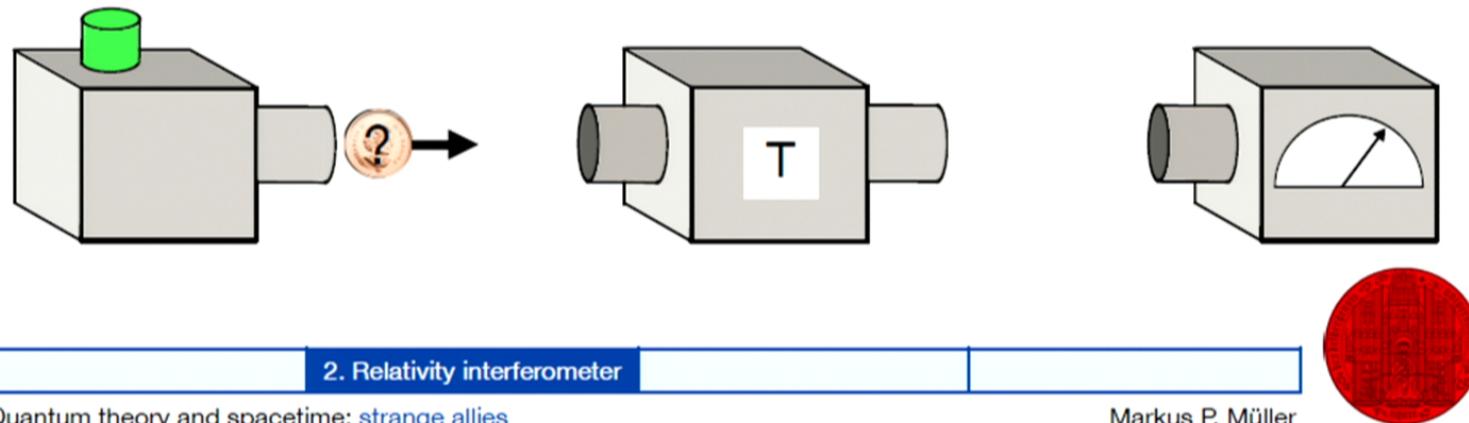
General state spaces

Example: classical coin toss.



- The preparation device prepares a physical system in a state ω . Here

$$\omega = \begin{pmatrix} \text{Prob(heads)} \\ \text{Prob(tails)} \end{pmatrix} = \begin{pmatrix} p \\ 1-p \end{pmatrix}.$$



2. Relativity interferometer

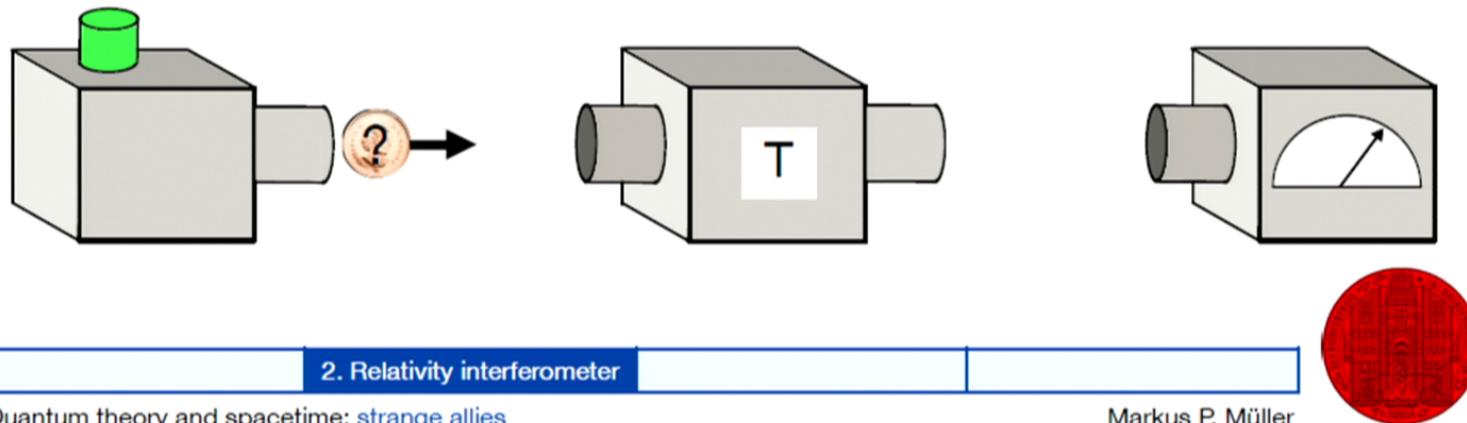
General state spaces

Example: classical coin toss.



- The preparation device prepares a physical system in a state ω . Here

$$\omega = \begin{pmatrix} \text{Prob(heads)} \\ \text{Prob(tails)} \end{pmatrix} = \begin{pmatrix} p \\ 1-p \end{pmatrix}.$$



2. Relativity interferometer

General state spaces

Example: classical coin toss.

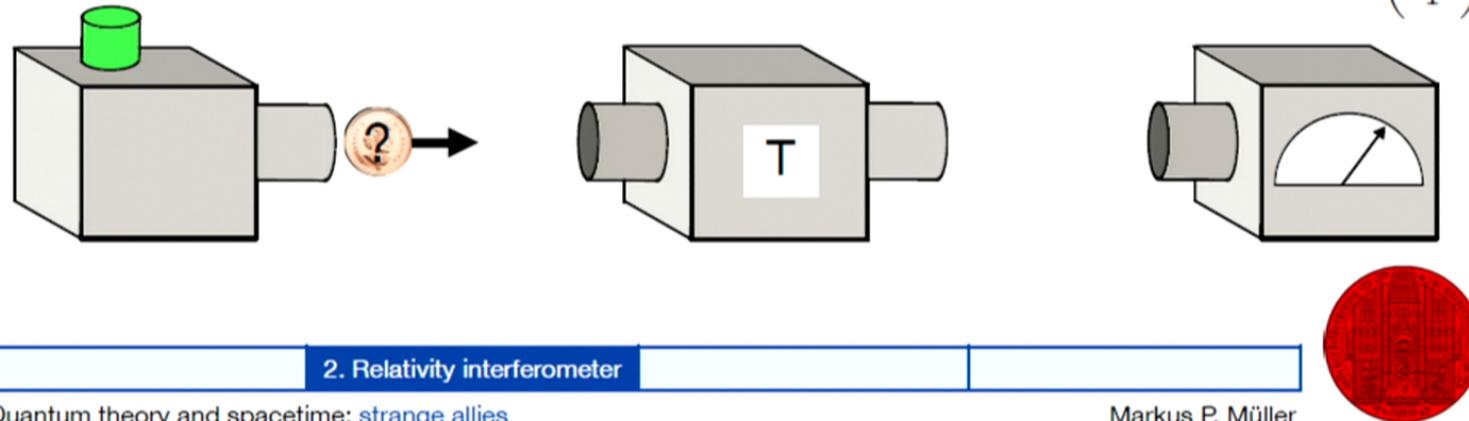


- The preparation device prepares a physical system in a state ω . Here

$$\omega = \begin{pmatrix} \text{Prob(heads)} \\ \text{Prob(tails)} \end{pmatrix} = \begin{pmatrix} p \\ 1-p \end{pmatrix}.$$

$$\begin{array}{c} \left(\begin{array}{c} 1 \\ 0 \end{array} \right) \\ \downarrow \\ \text{blue dot} \\ \left(\begin{array}{c} 1/2 \\ 1/2 \end{array} \right) \\ \downarrow \\ \left(\begin{array}{c} 0 \\ 1 \end{array} \right) \end{array}$$

State space Ω : the set of all possible states



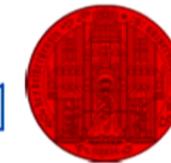
2. Relativity interferometer

General state spaces

Example: quantum spin-1/2 particle.



ψ



2. Relativity interferometer

Quantum theory and spacetime: strange allies

Markus P. Müller

General state spaces

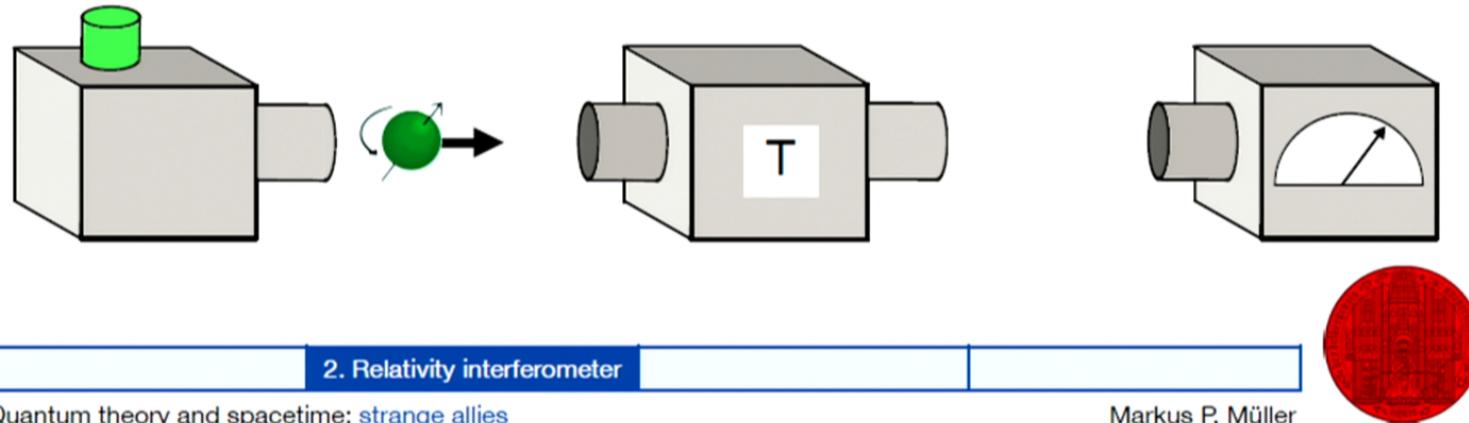
Example: quantum spin-1/2 particle.



- The preparation device prepares a spin-1/2 particle in quantum state ω .

$$\alpha|\uparrow\rangle + \beta|\downarrow\rangle$$

More generally: ω is 2x2 density matrix.

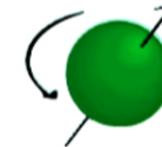


2. Relativity interferometer



General state spaces

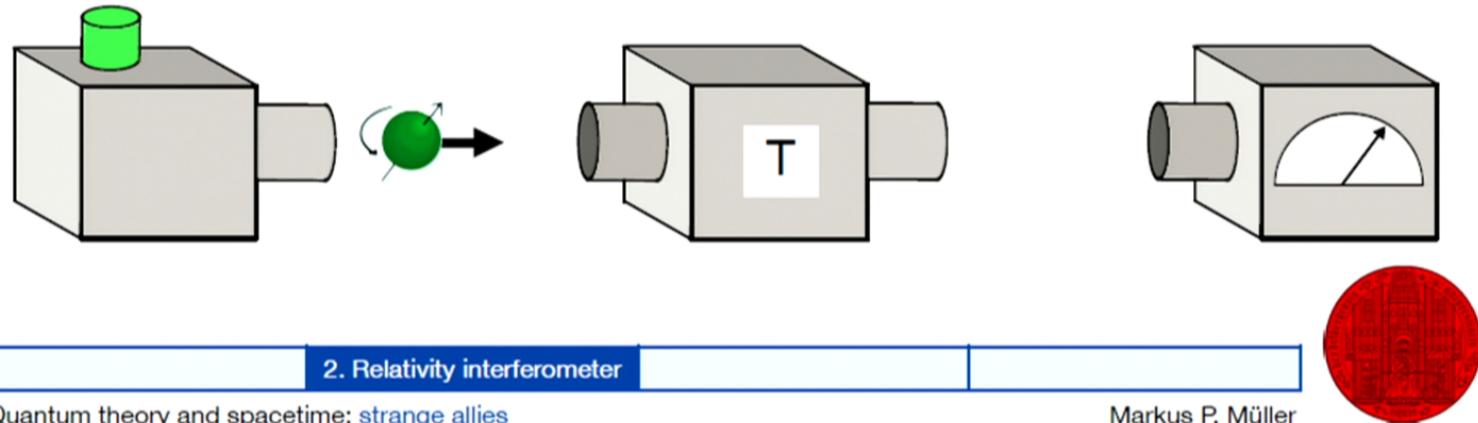
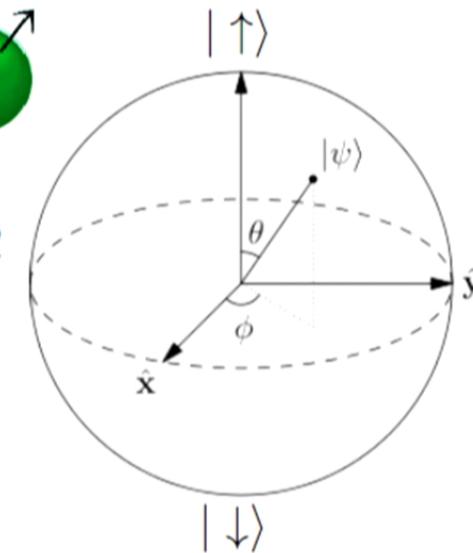
Example: quantum spin-1/2 particle.



- The preparation device prepares a spin-1/2 particle in quantum state ω .

$$\cos \frac{\theta}{2} | \uparrow \rangle + e^{i\phi} \sin \frac{\theta}{2} | \downarrow \rangle$$

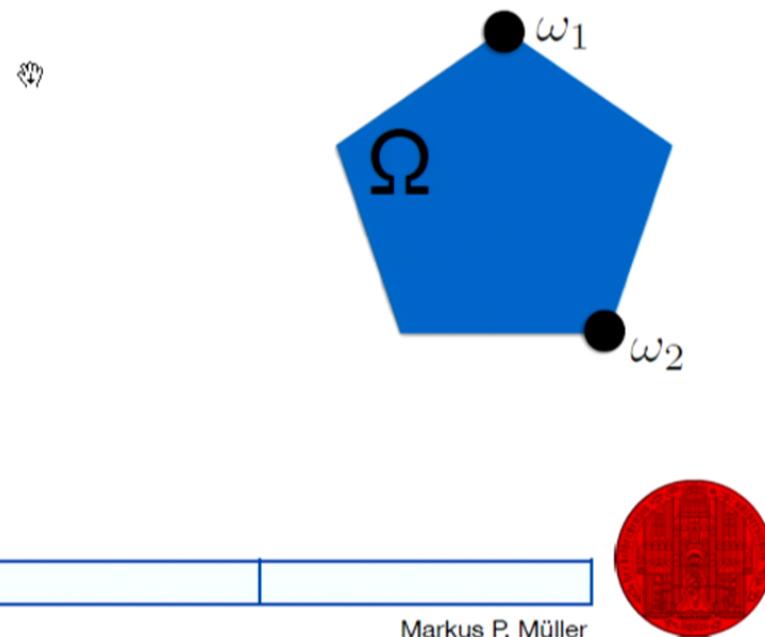
More generally: ω is 2x2 density matrix.



2. Relativity interferometer

General state spaces

The **set of all possible states** of a given physical system is called the state space Ω .

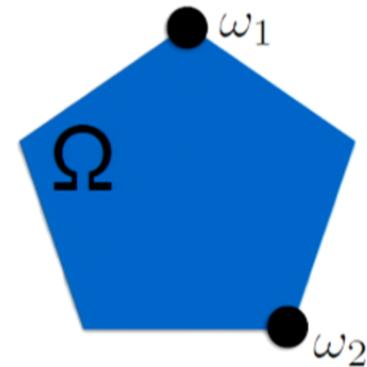
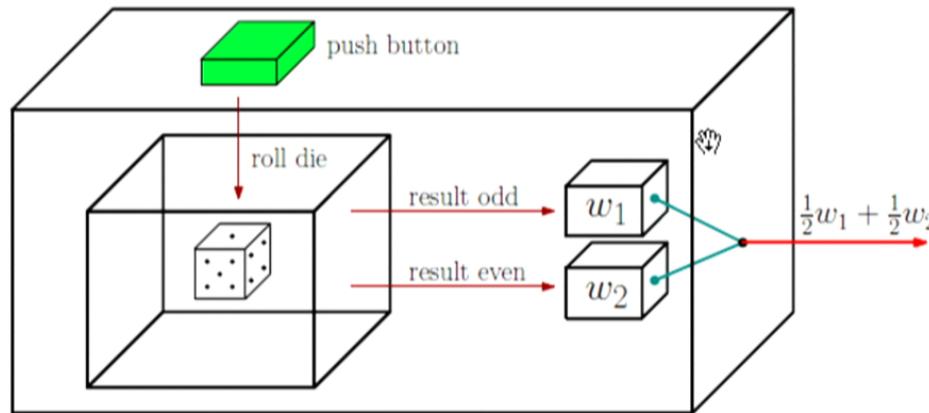


General state spaces

The **set of all possible states** of a given physical system is called the state space Ω .

Preparation of **statistical mixtures**: $\omega = \lambda\omega_1 + (1 - \lambda)\omega_2$

$$(0 \leq \lambda \leq 1)$$



General state spaces

The **set of all possible states** of a given physical system is called the state space Ω .

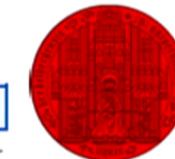
Preparation of **statistical mixtures**: $\omega = \lambda\omega_1 + (1 - \lambda)\omega_2$
 $(0 \leq \lambda \leq 1)$

QT: Ω_N = set of $N \times N$ density matrices

CPT: Ω_N = set of prob. distributions
 (p_1, \dots, p_N) .



Thus Ω is a **convex set**.



General state spaces



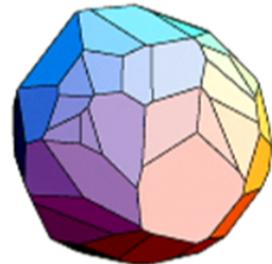
classical
bit



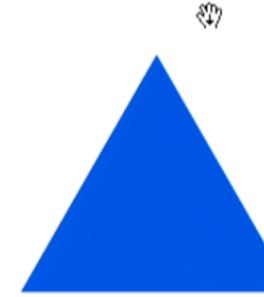
quantum
bit



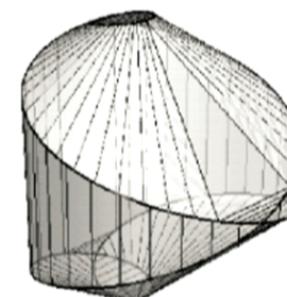
"gbit"



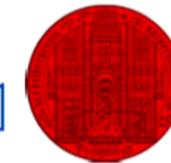
Arbitrary convex
state space



Classical "trit"
(3-level-system)

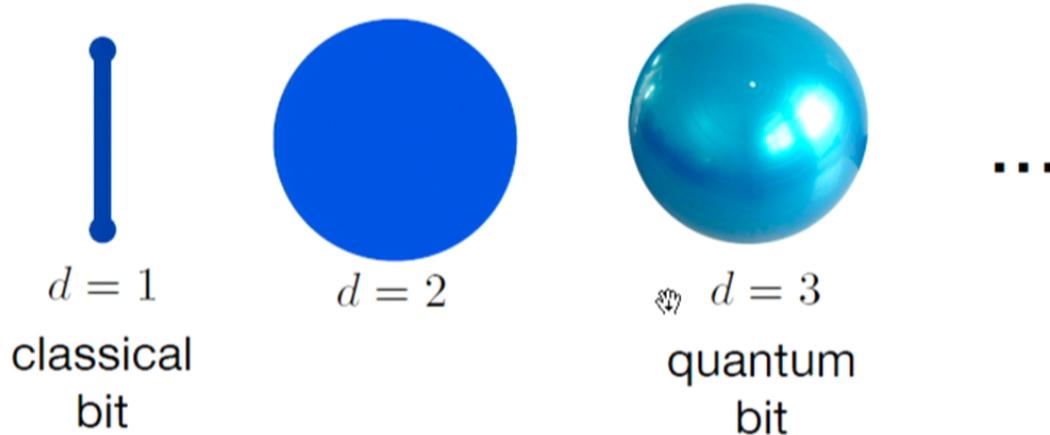


Quantum "trit".
Complicated, 8D!



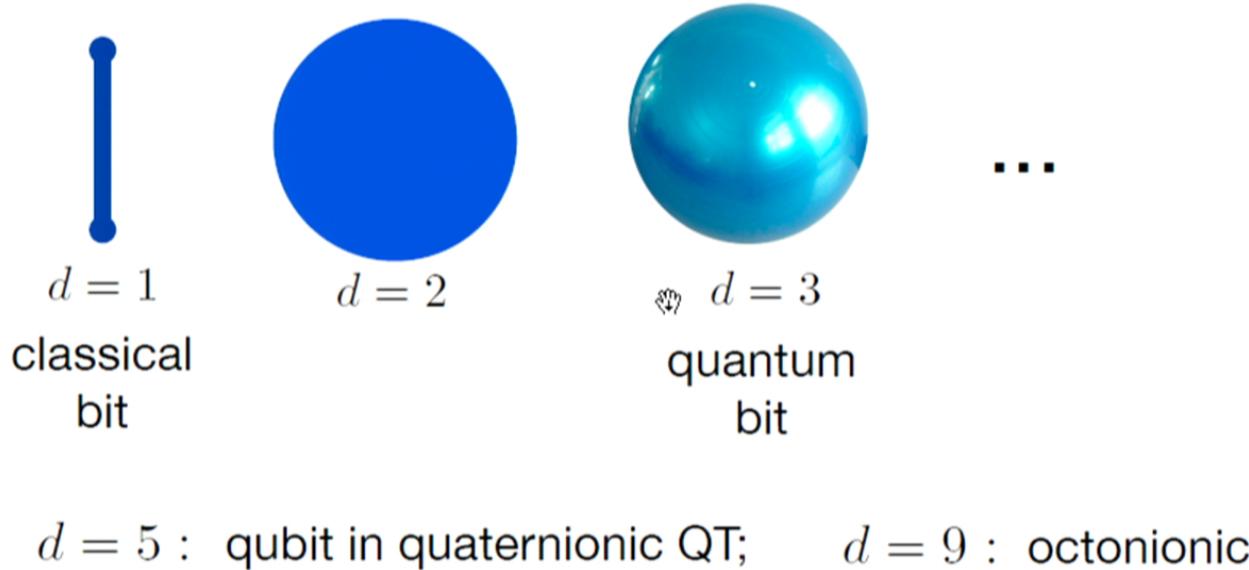
Relativistic constraints on the state space

Two-level state spaces (“bits”) are naturally ball state spaces:



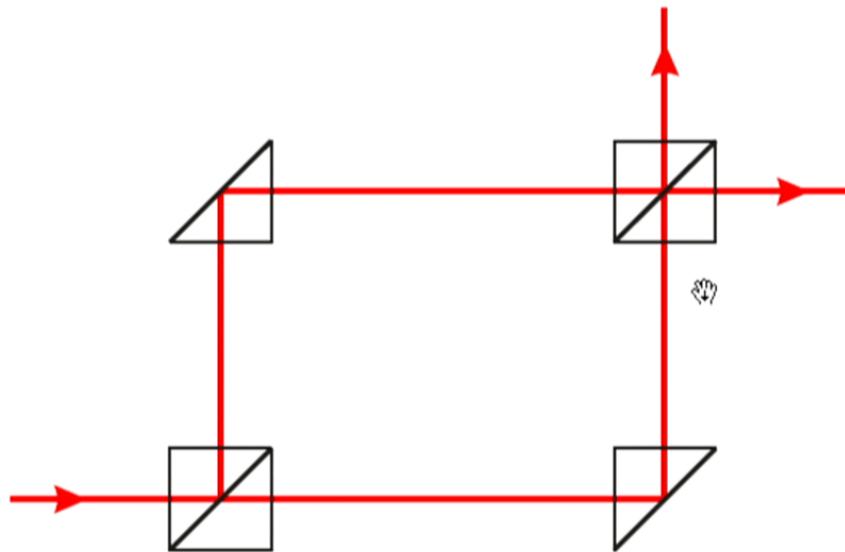
Relativistic constraints on the state space

Two-level state spaces (“bits”) are naturally ball state spaces:



Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112



2. Relativity interferometer

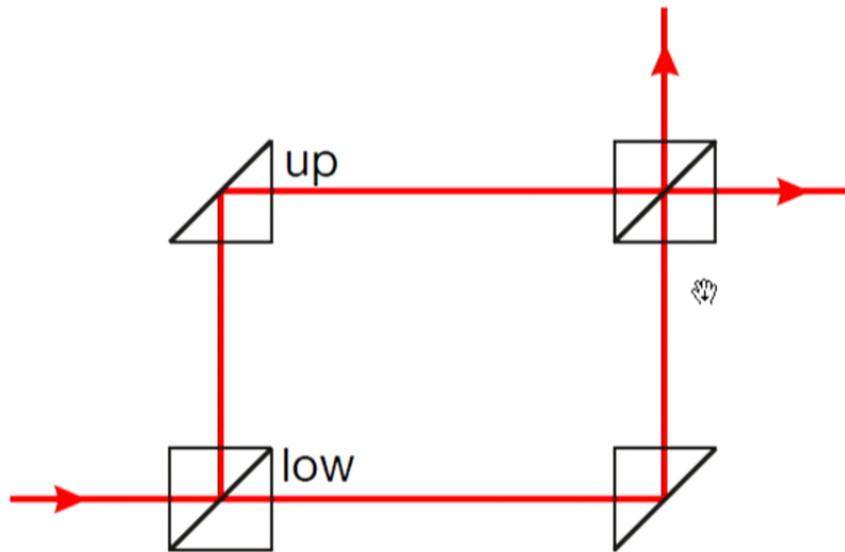
Quantum theory and spacetime: strange allies

Markus P. Müller



Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112



2. Relativity interferometer

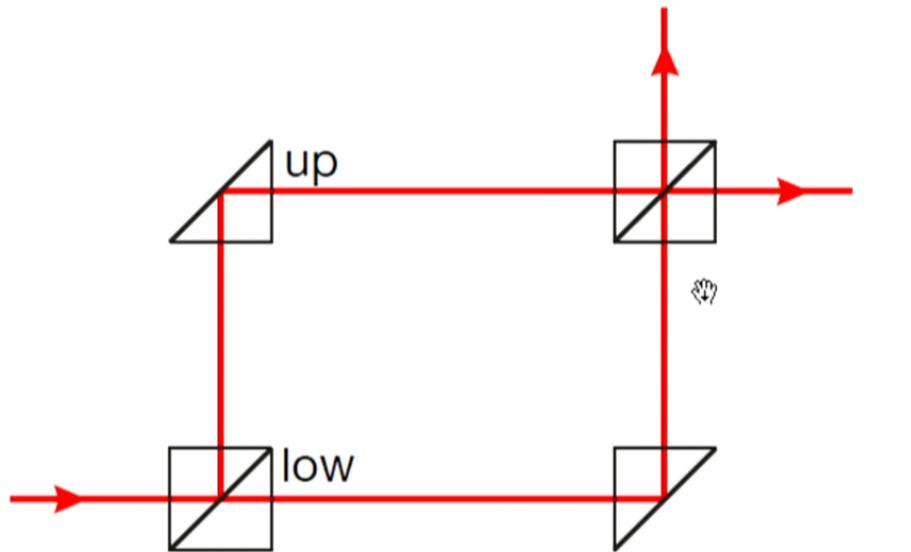
Quantum theory and spacetime: strange allies

Markus P. Müller

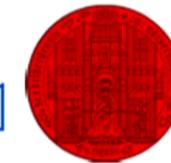


Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112

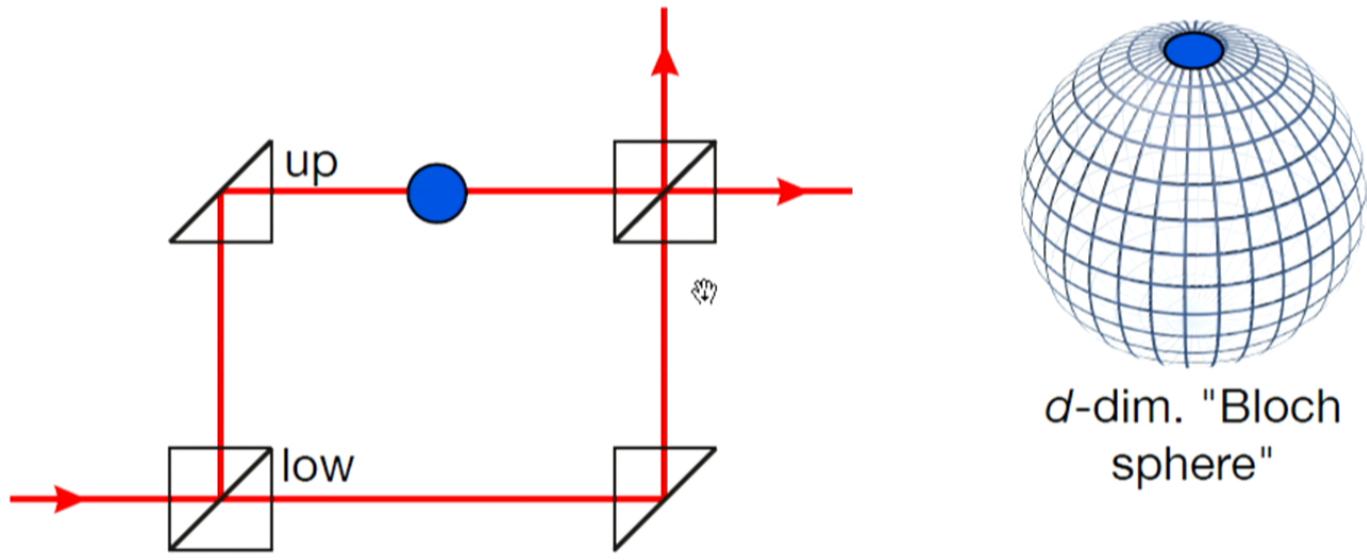


d -dim. "Bloch sphere"



Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112

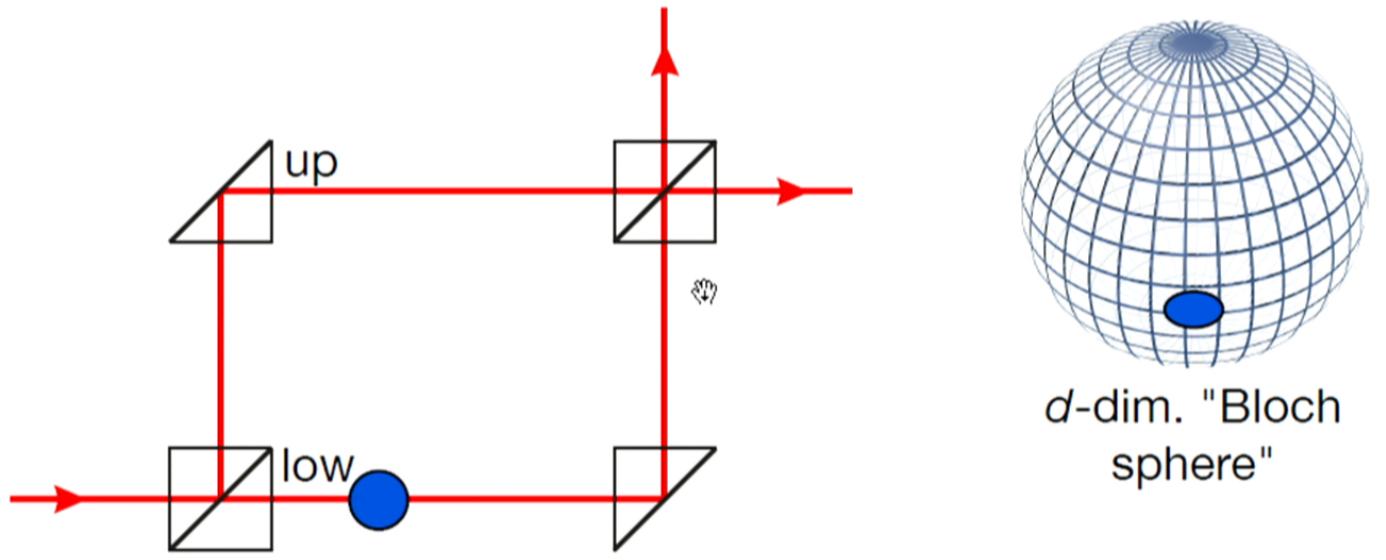


North-pole state: particle definitely in upper branch.



Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112

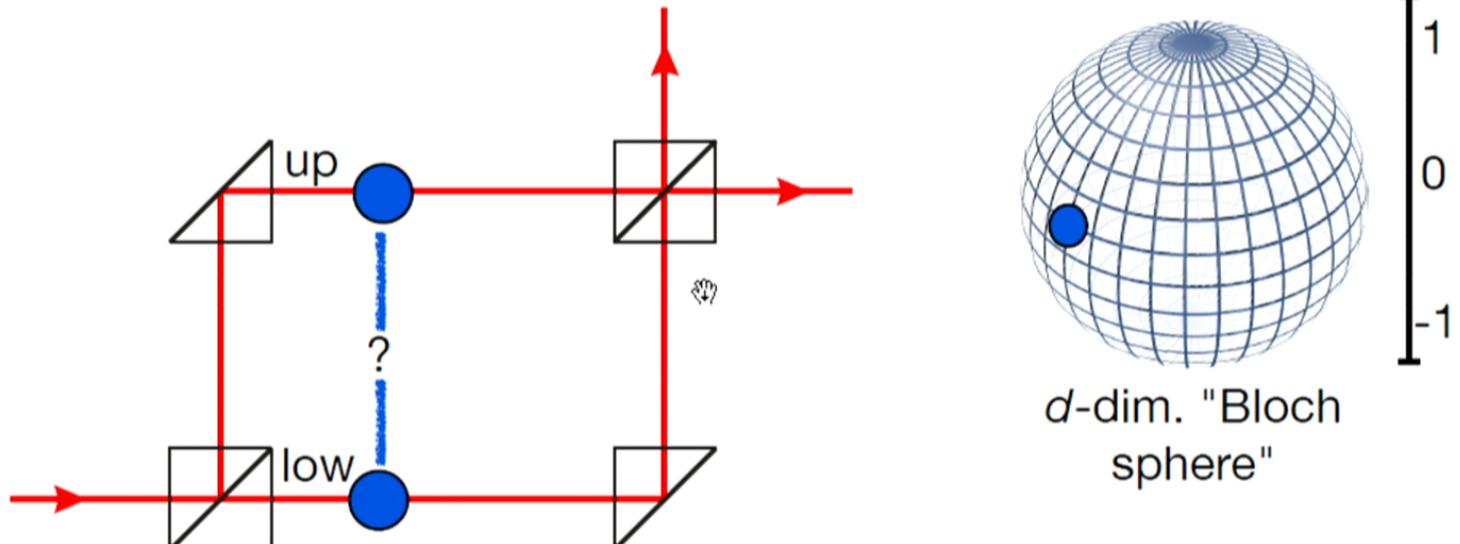


South-pole state: particle definitely in lower branch.



Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112



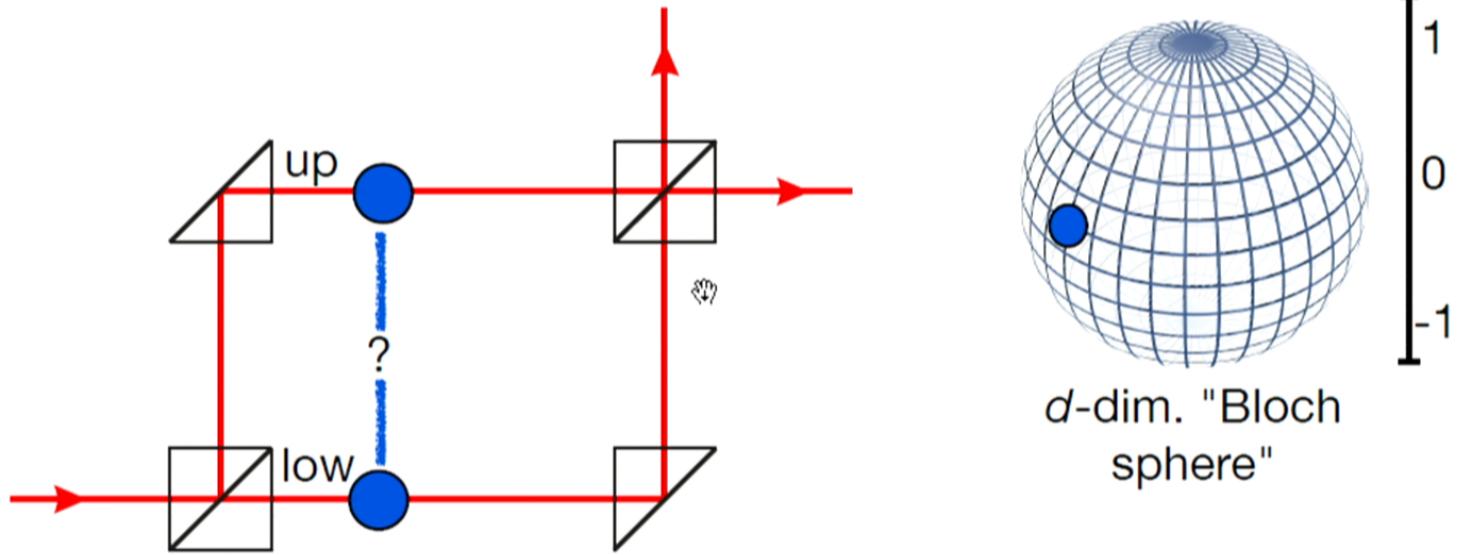
State on equator $z=0$: probability 1/2 for each.

	2. Relativity interferometer	
--	------------------------------	--



Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112



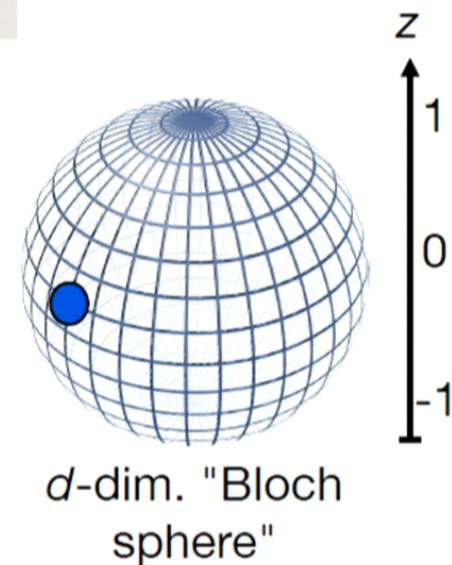
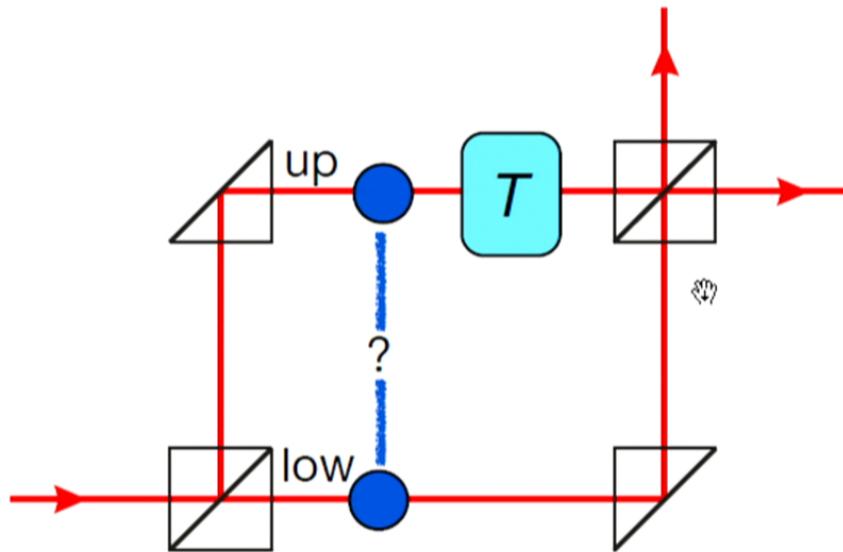
State on equator $z=0$: probability 1/2 for each.

$$p(\text{up}) = \frac{1}{2}(z + 1)$$



Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112



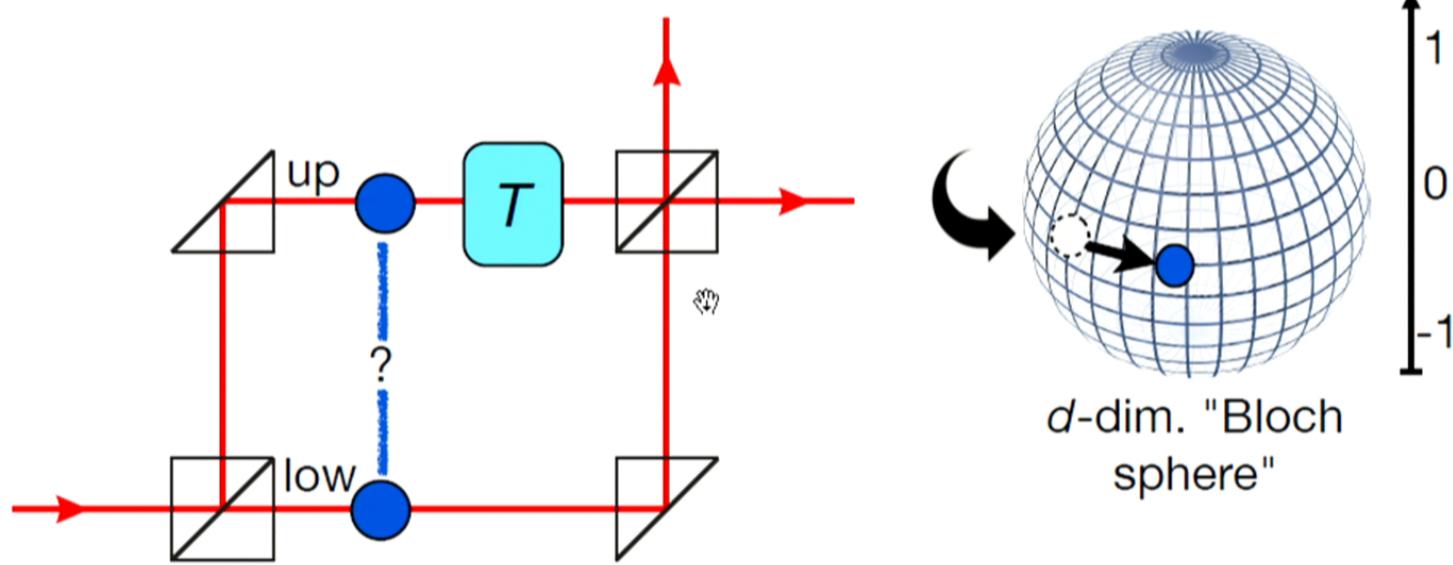
d -dim. "Bloch sphere"

What transformations T can we perform locally in one arm...
... without any information loss?



Relativistic constraints on the state space

A. Garner, MM, O. Dahlsten, arXiv:1412.7112

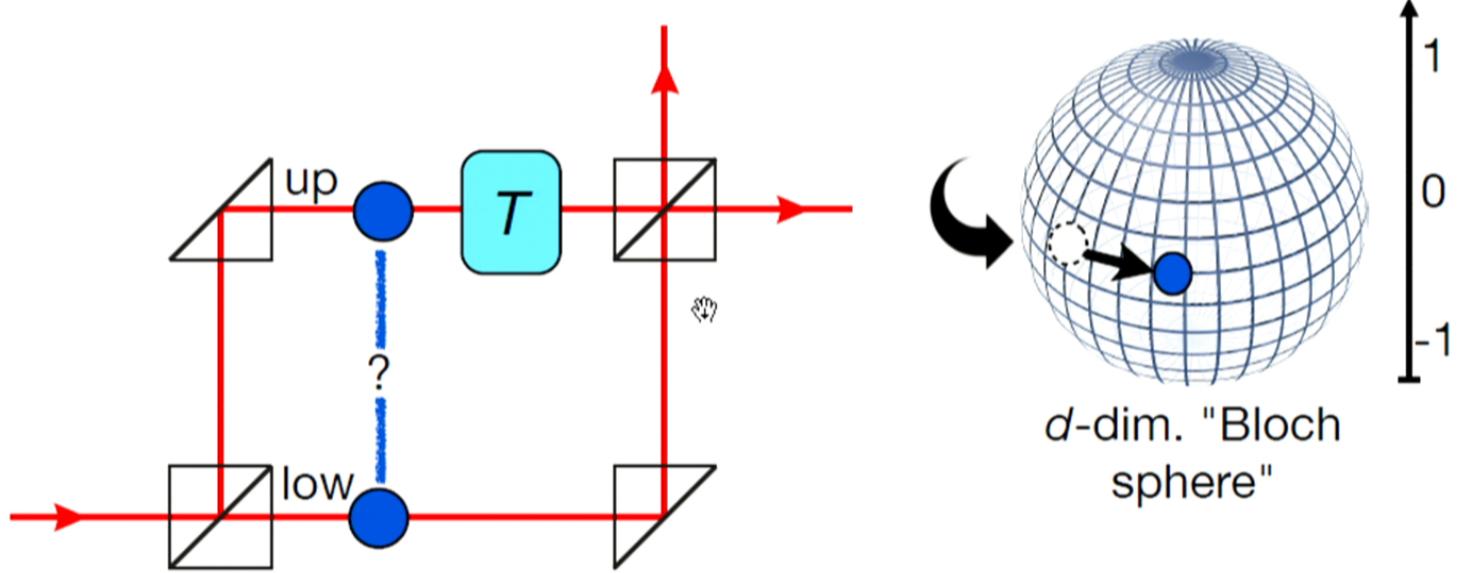


T must be a **rotation** of the Bloch ball (reversible+linear)...
... and must preserve $p(\text{up})$, i.e. **preserve the z-axis**.



Relativistic constraints on the state space

Assumption: $\mathcal{G}_{\text{up}} = \mathcal{G}_{\text{low}} \simeq \text{SO}(d - 1)$.

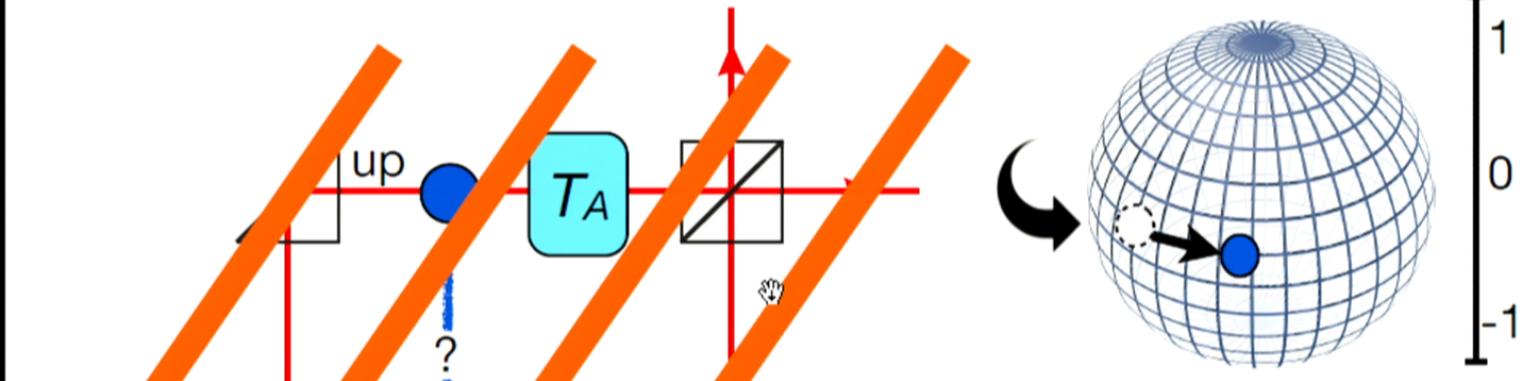


T must be a **rotation** of the Bloch ball (reversible+linear)...
... and must preserve $p(\text{up})$, i.e. **preserve the z-axis**.



Relativistic constraints on the state space

Assumption: $\mathcal{G}_{\text{up}} = \mathcal{G}_{\text{low}} \simeq \text{SO}(d - 1)$.



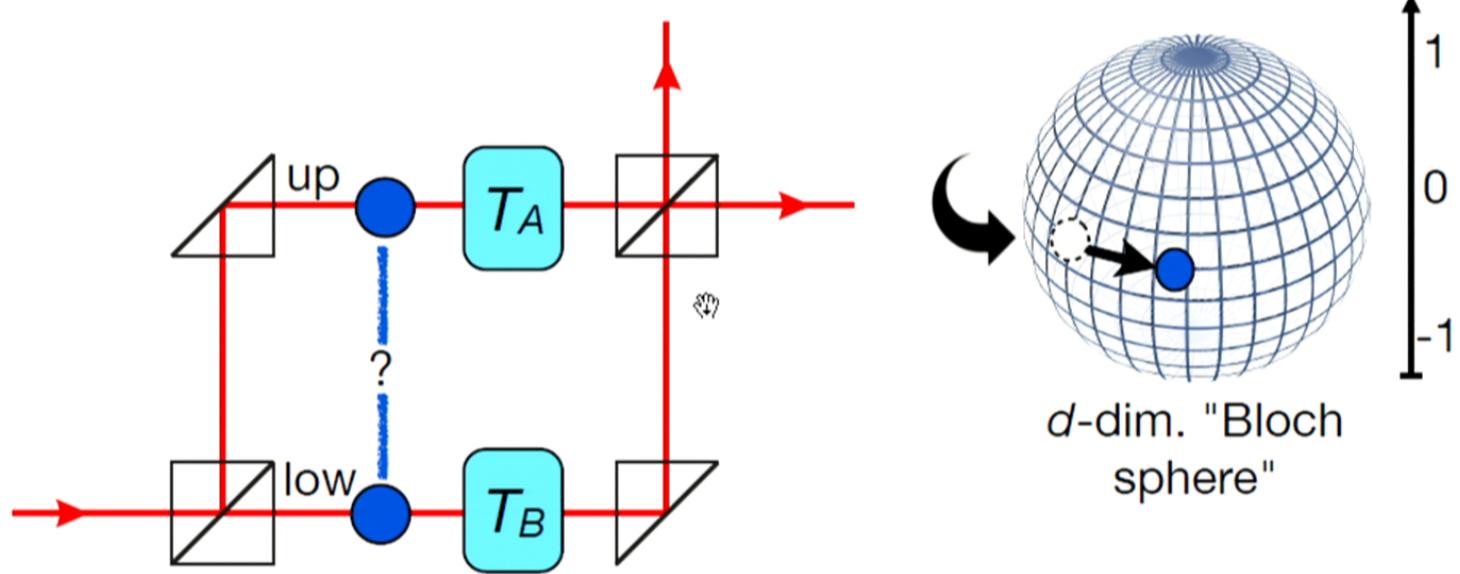
Relativity: there is one frame of reference in which T_A happens first, and then T_B ...

	2. Relativity interferometer
--	------------------------------



Relativistic constraints on the state space

Assumption: $\mathcal{G}_{\text{up}} = \mathcal{G}_{\text{low}} \simeq \text{SO}(d - 1)$.



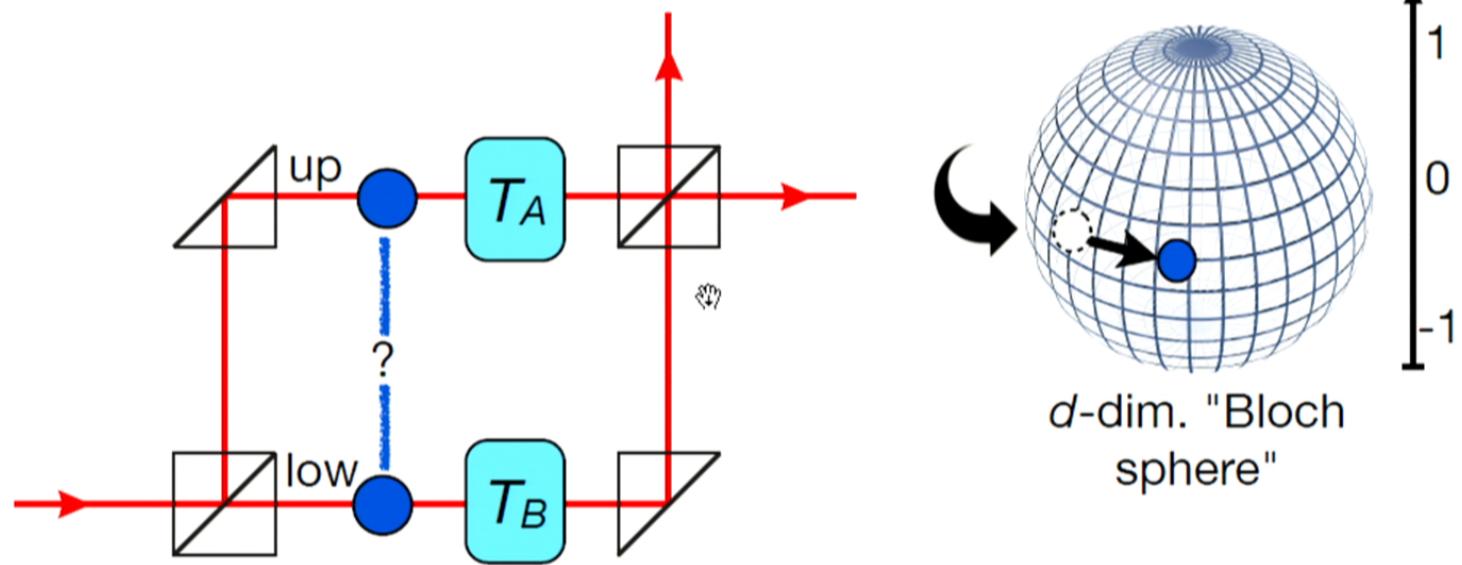
Detector click statistics is Lorentz-invariant

$$\Rightarrow T_A T_B = T_B T_A \text{ for all } T_A, T_B \in \text{SO}(d - 1).$$



Relativistic constraints on the state space

$\Rightarrow d \leq 3$ (In fact, $d=3$, otherwise these transformations are all trivial.)



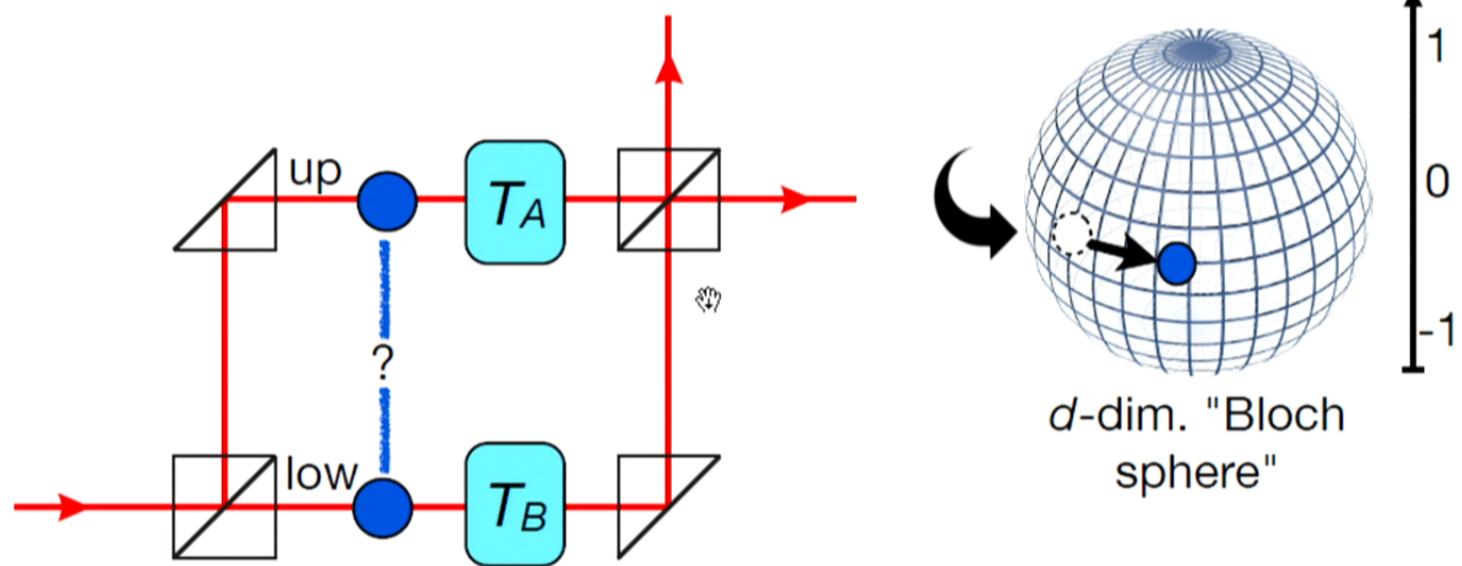
Detector click statistics is Lorentz-invariant

$$\Rightarrow T_A T_B = T_B T_A \text{ for all } T_A, T_B \in \mathrm{SO}(d-1).$$



Relativistic constraints on the state space

$\Rightarrow d \leq 3$ (In fact, $d=3$, otherwise these transformations are all trivial.)



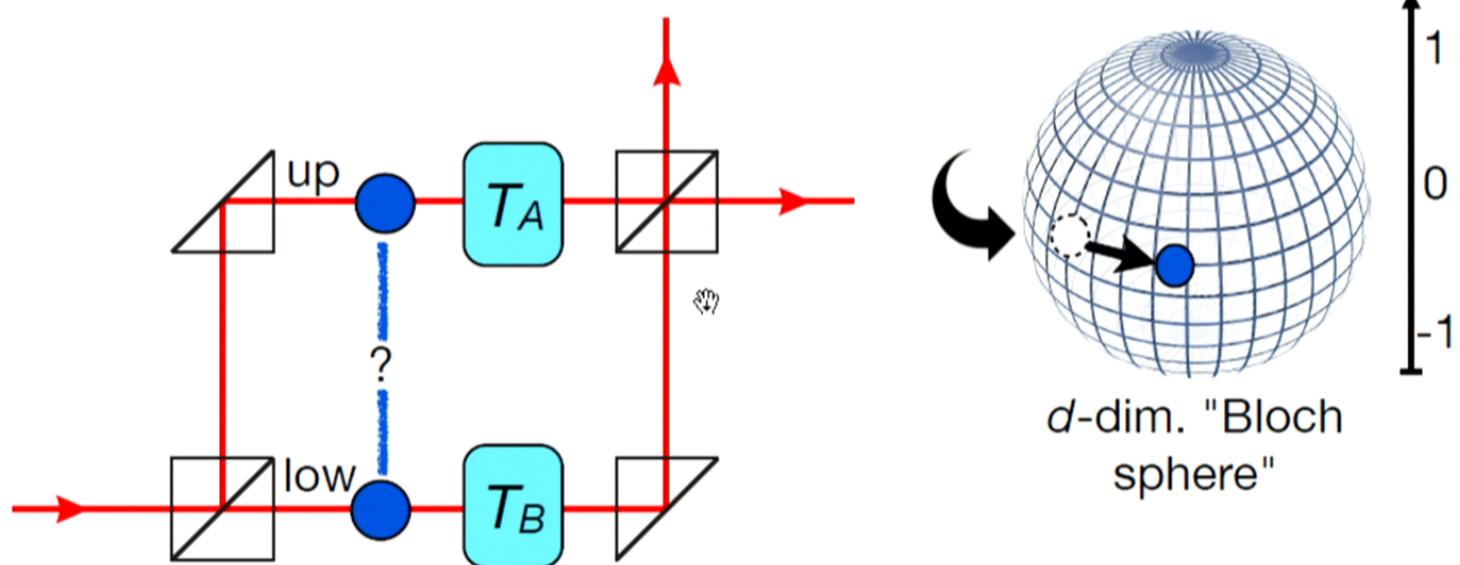
Detector click statistics is Lorentz-invariant

$$\Rightarrow T_A T_B = T_B T_A \text{ for all } T_A, T_B \in \mathrm{SO}(d-1).$$



Relativistic constraints on the state space

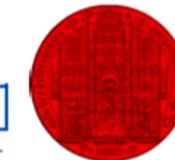
Weaker assumption: \mathcal{G}_{up} and \mathcal{G}_{low} isomorphic



2. Relativity interferometer

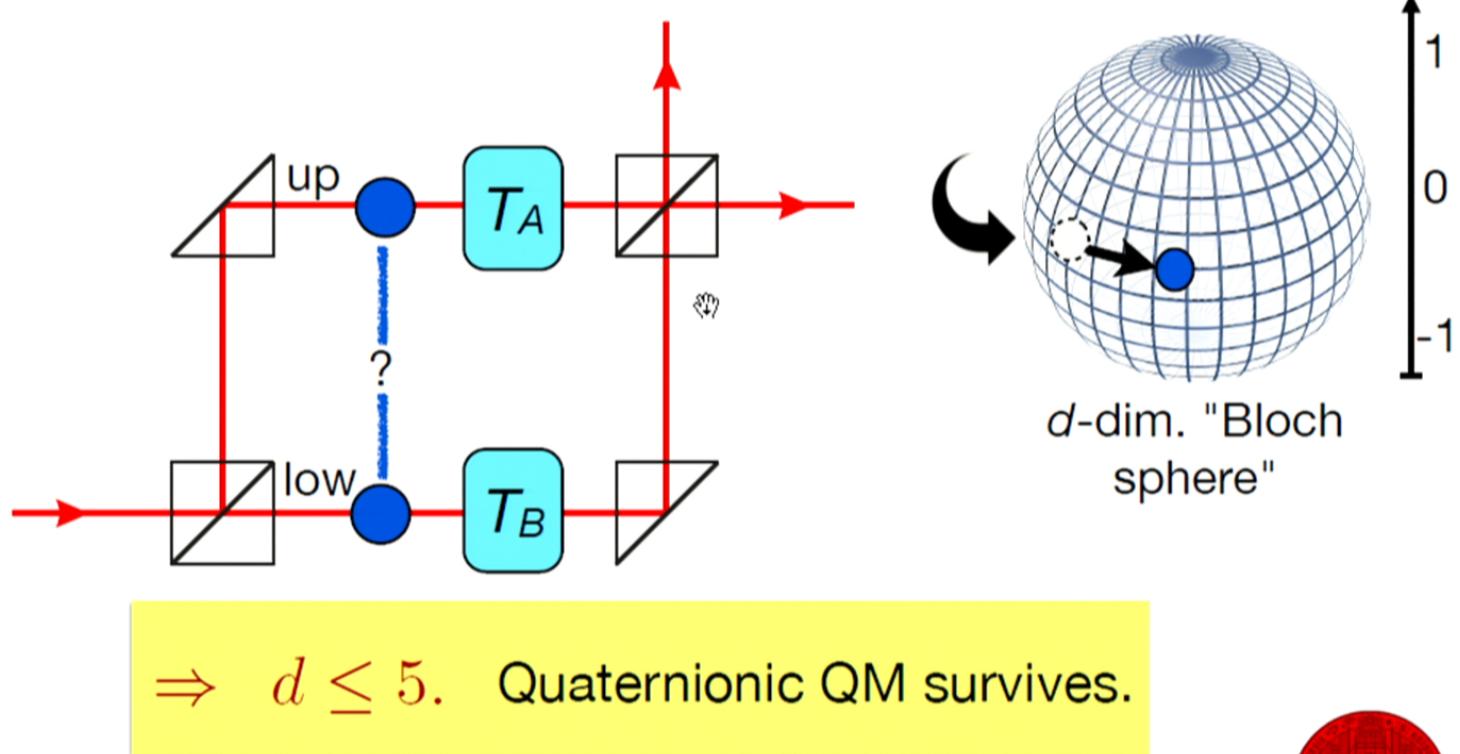
Quantum theory and spacetime: strange allies

Markus P. Müller



Relativistic constraints on the state space

Weaker assumption: \mathcal{G}_{up} and \mathcal{G}_{low} isomorphic

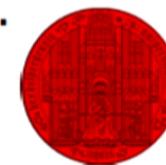


Conclusion part 2



QT ← space-time

Space-time structure **constrains QT's state space**:
upper bound on # of incompatible qubit observables.



	2. Relativity interferometer	
--	------------------------------	--

Conclusion part 2

Consequences for actual interference experiments:

PHYSICAL REVIEW LETTERS

VOLUME 42

12 MARCH 1979

NUMBER 11

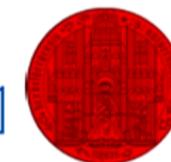


Proposed Test for Complex versus Quaternion Quantum Theory

Asher Peres

Department of Physics, Technion-Israel Institute of Technology, Haifa, Israel
(Received 7 December 1978)

If scattering amplitudes are ordinary complex numbers (not quaternions) then there is a universal algebraic relationship between the six coherent cross sections of any three scatterers (taken singly and pairwise). A violation of this relationship would indicate either that scattering amplitudes are quaternions, or that the superposition principle fails. Some experimental tests are proposed, involving neutron diffraction by crystals made of three different isotopes, neutron interferometry, and K_S -meson regeneration.



2. Relativity interferometer

Quantum theory and spacetime: strange allies

Markus P. Müller

Conclusion part 2

Consequences for actual interference experiments:

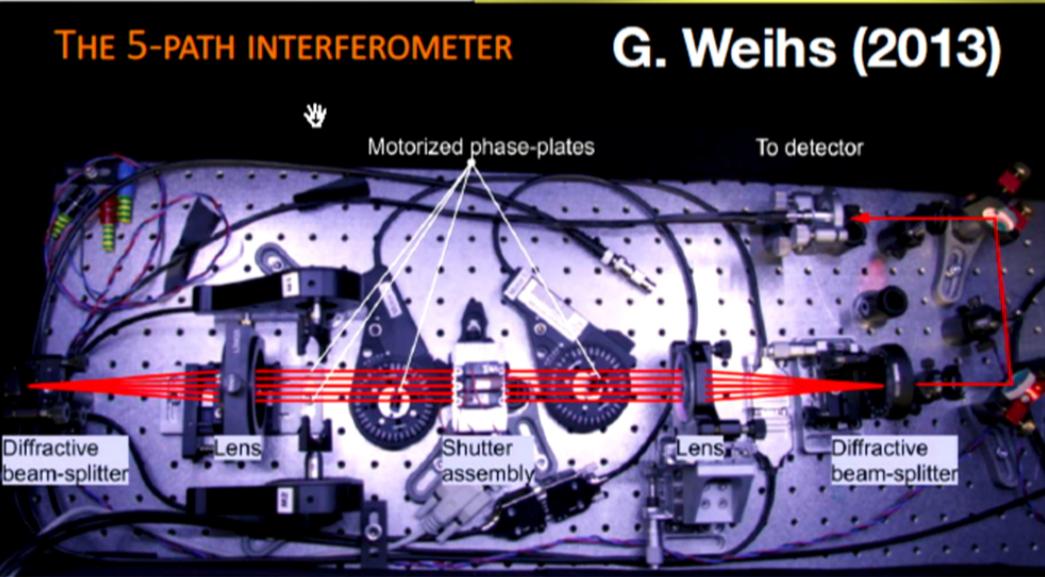
PHYSICAL REVIEW LETTERS

VOLUME 42

Proposed Test for

Department of Physics

If scattering amplitudes a universal algebraic relation between scatterers (taken singly and together) fails, either that scattering amplitude fails. Some experimental results made of three different is



2. Relativity interferometer

Quantum theory and spacetime: strange allies

Markus P. Müller



Could it also go the other way?



QT  space-time

		3. SO(3,1)	
--	--	------------	--

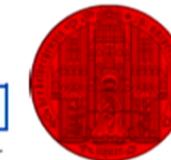
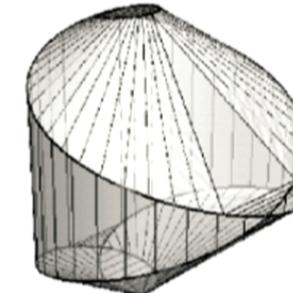


Hints for QT → space-time

Spin: representation of rotation group $SO(d)$

reversible transformation G_R of state space

↑↓ 1:1
spatial rotation R

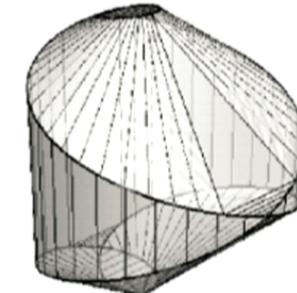


Hints for QT → space-time

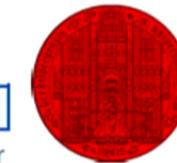
Spin: representation of rotation group $SO(d)$

reversible transformation G_R of state space

↑↓ 1:1
spatial rotation R



Impossible for most state spaces!

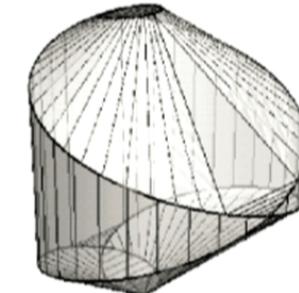


Hints for QT → space-time

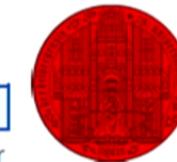
Spin: representation of rotation group $SO(d)$

reversible transformation G_R of state space

↑↓ 1:1
spatial rotation R



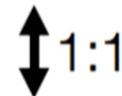
Impossible for most state spaces!



Hints for QT → space-time

Spin: representation of rotation group $SO(d)$

linear structure of statistical mixtures



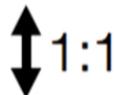
linear structure of tangent space



Hints for QT → space-time

Spin: representation of rotation group $\text{SO}(d)$

linear structure of statistical mixtures



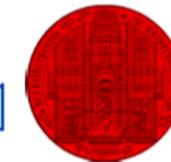
linear structure of tangent space



→ only possible for standard qubit and $d=3$.

But the 3D-Bloch ball can be derived from
information-theoretic postulates!

(Cf. talk by Philipp Hoehn; Lucien Hardy 2001; Dakic, Brukner 2009;
Masanes, Müller 2011; Chribella, D'Ariano, Perinotto 2011, uvm.)



	3. $\text{SO}(3,1)$	
--	---------------------	--

Hints for QT → space-time

Recall Philipp's talk:

GR: has informational essence (e.g., horizons)

- standard: spacetime \Rightarrow causal structure \Rightarrow information flow
- alternative: information flow \Rightarrow causal structure \Rightarrow spacetime (up to scale)
[Hawking, Geroch,...]



Hints for QT → space-time

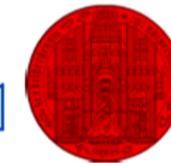
Recall Philipp's talk:

GR: has informational essence (e.g., horizons)

- standard: spacetime \Rightarrow causal structure \Rightarrow information flow
- alternative: information flow \Rightarrow causal structure \Rightarrow spacetime (up to scale)
[Hawking, Geroch,...]



Can we infer (local) space-time symmetry from informational relations between observers?



The general setup

Philipp Hoehn and MM, arXiv:1412.8462

Related (but different): B. Dakic and C. Brukner, arXiv:1307.3984 (!!)
MM and LI. Masanes, New J. Phys. **15**, 053040 (2013)

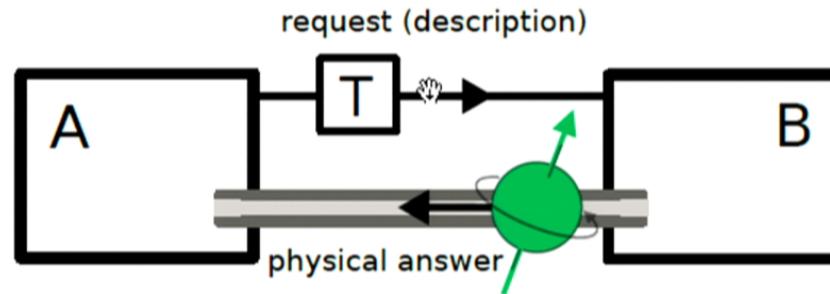


The general setup

Philipp Hoehn and MM, arXiv:1412.8462

Related (but different): B. Dakic and C. Brukner, arXiv:1307.3984 (!!)

MM and LI. Masanes, New J. Phys. **15**, 053040 (2013)

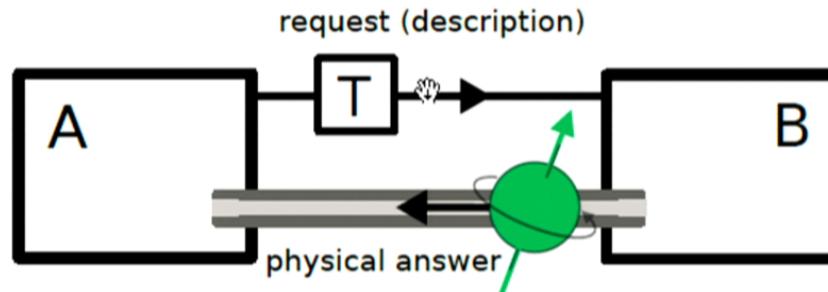


The general setup

Philipp Hoehn and MM, arXiv:1412.8462

Related (but different): B. Dakic and C. Brukner, arXiv:1307.3984 (!!)

MM and LI. Masanes, New J. Phys. **15**, 053040 (2013)



Two distant observers (Alice, Bob) who have never met.

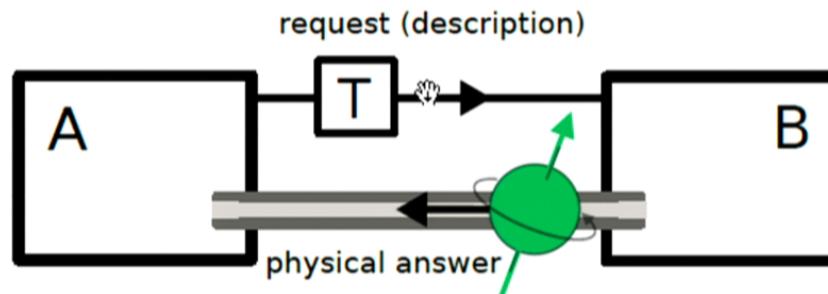
Alice (on phone): "Please send the following object: [description]"

Bob: "Ok!", and sends some physical object back.



The general setup

Problem: A+B have never met
→ lack of common reference frame



Two distant observers (Alice, Bob) who have never met.

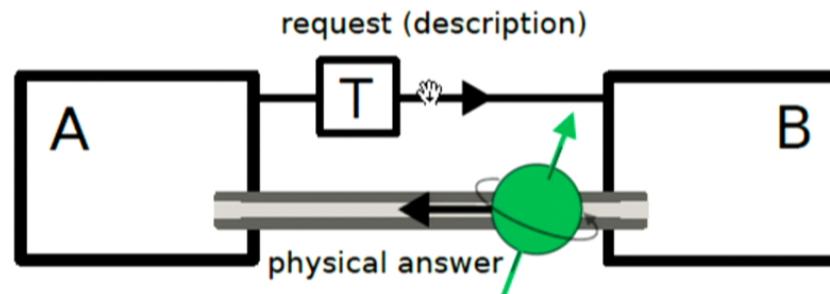
Alice (on phone): "Please send the following object: [description]"
Bob: "Ok!", and sends some physical object back.



The general setup

Problem: A+B have never met
→ lack of common reference frame
→ need correcting transformation $T = \varphi_B \circ \varphi_A^{-1}$.

Bob's description
of physical objects



Two distant observers (Alice, Bob) who have never met.

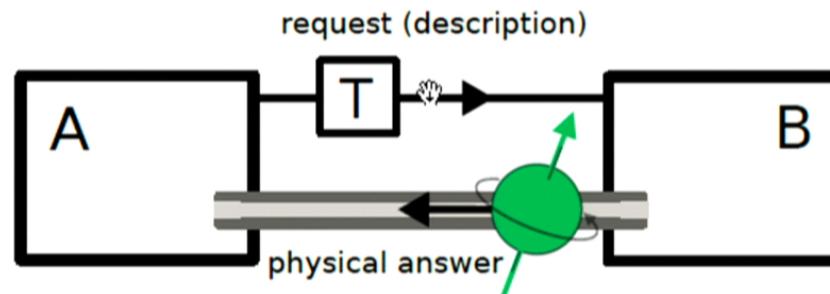
Alice (on phone): "Please send the following object: [description]"
Bob: "Ok!", and sends some physical object back.



The general setup

Problem: A+B have never met.
→ lack of common reference frame
→ need correcting transformation $T = \varphi_B \circ \varphi_A^{-1}$.

Alice's description
of physical objects

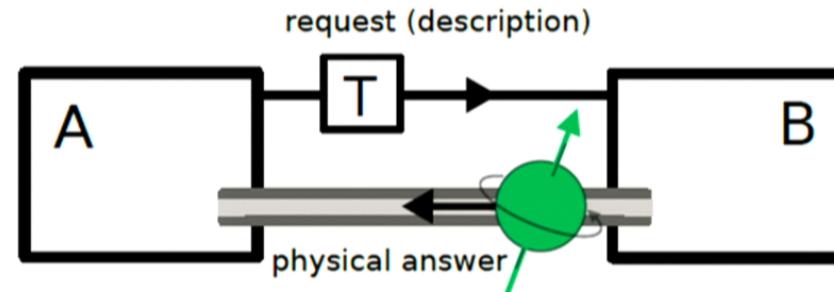


Two distant observers (Alice, Bob) who have never met.

Alice (on phone): "Please send the following object: [description]"
Bob: "Ok!", and sends some physical object back.



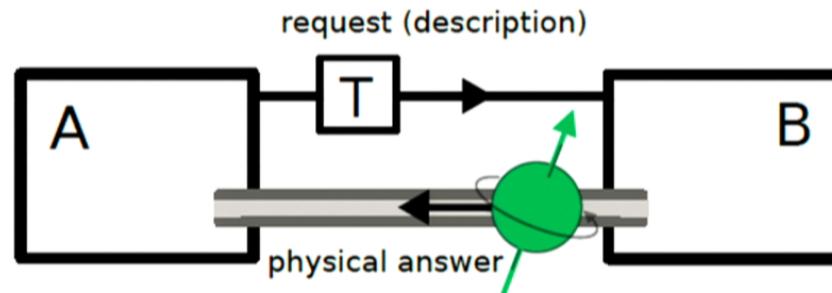
The general setup



Example: assume classical mechanics;
physical objects= billiard balls with specified v .



The general setup

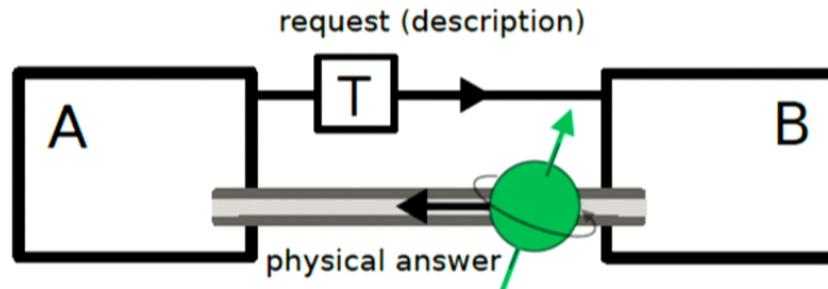


Example: assume classical mechanics;
physical objects= billiard balls with specified v .

A and B could choose arbitrary continuous encodings φ_A, φ_B .
 φ : physical velocities $\rightarrow \mathbb{R}^3$.



The general setup



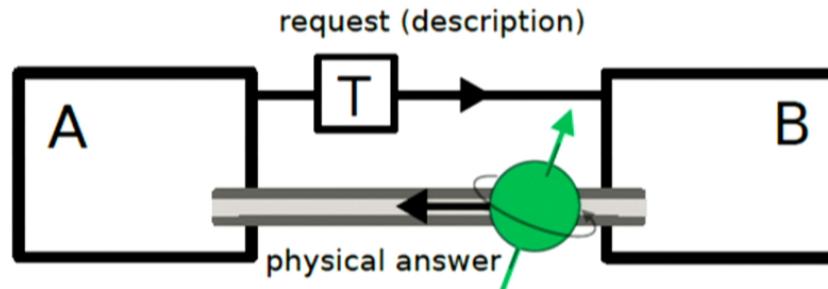
Example: assume classical mechanics;
physical objects= billiard balls with specified v .

A and B could choose arbitrary continuous encodings φ_A, φ_B .
 φ : physical velocities $\rightarrow \mathbb{R}^3$.

Then there would always be an element $T \in \mathcal{G}$ that does the job,
 $\mathcal{G} = \{\text{homeomorphisms of } \mathbb{R}^3\}$.



The general setup



Example: assume classical mechanics;
physical objects= billiard balls with specified v .

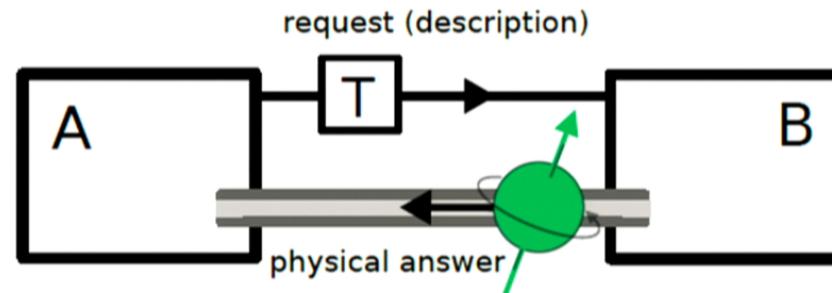
A and B could choose inertial frame encodings φ_A, φ_B .

Then $T \in \mathcal{G}'$ is sufficient, where

$$\mathcal{G}' = \text{Aff}(3, \mathbb{R}), \text{ the affine group.}$$



The general setup



Example: assume classical mechanics;
physical objects= billiard balls with specified v .

A and B could choose inertial frame encodings φ_A, φ_B .

Then $T \in \mathcal{G}'$ is sufficient, where

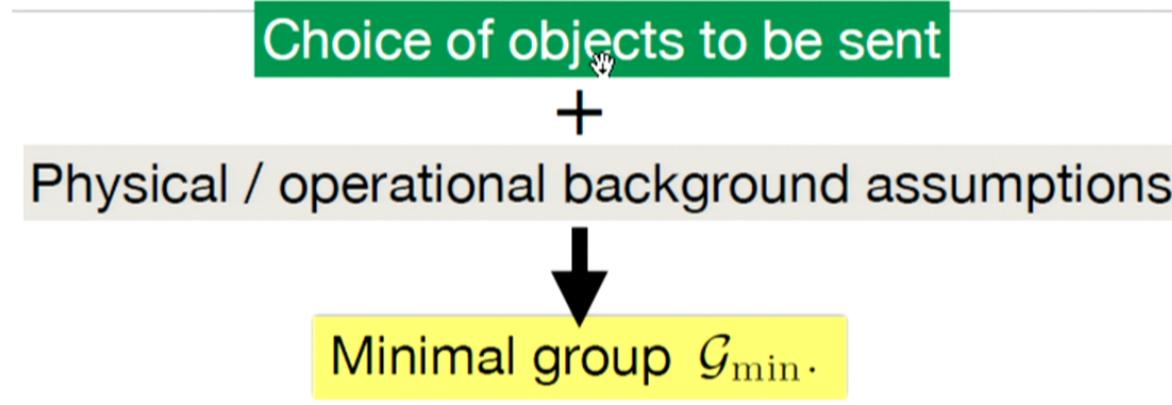
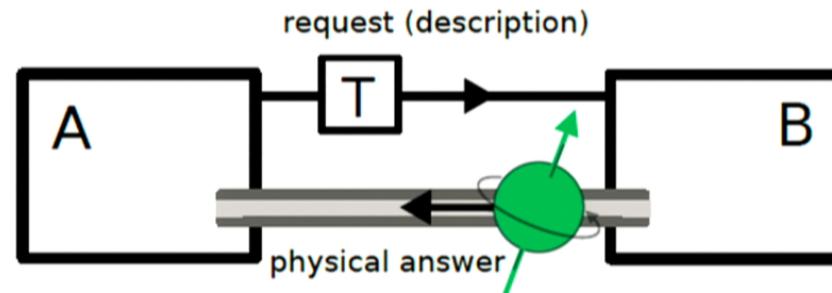
$$\mathcal{G}' = \text{Aff}(3, \mathbb{R}), \text{ the affine group.}$$

Then $\mathcal{G}' \subsetneq \mathcal{G}$, so this is "better" than before.

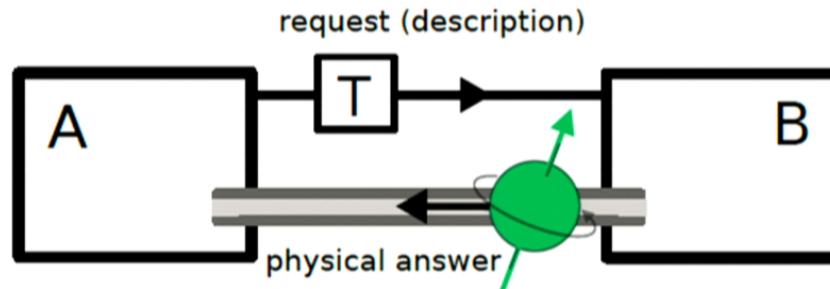
They cannot do better: $\mathcal{G}_{\min} = \text{Aff}(3, \mathbb{R})$.



The general setup



The general setup



Choice of objects to be sent

+

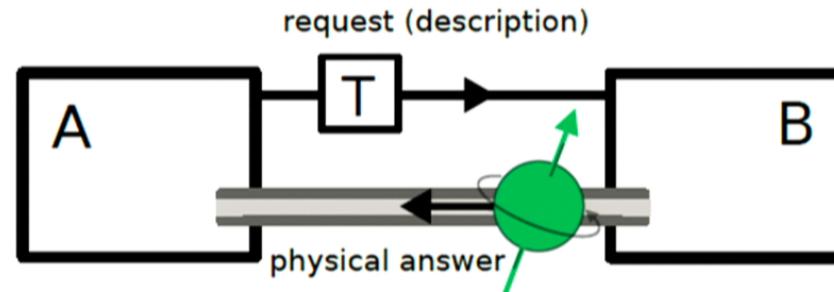
Physical / operational background assumptions



Minimal group \mathcal{G}_{\min} .

Theorem: If \mathcal{G}_{\min} exists then it is unique (as an abstract group).

Sending finite-dimensional quantum states



Choice of objects to be sent

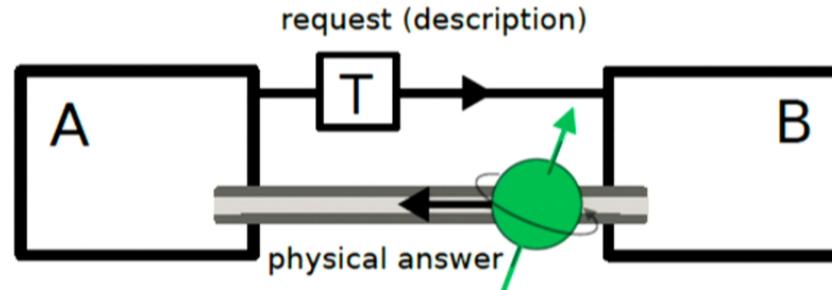
+

Physical / operational background assumptions



Minimal group \mathcal{G}_{\min} .

Sending finite-dimensional quantum states



Quantum states $\rho_{n \times n}$ for some $n \in N$

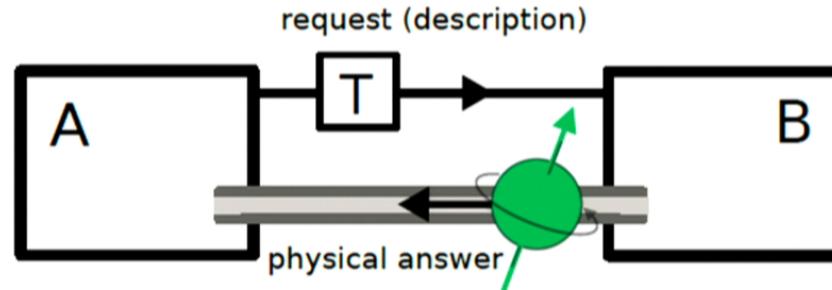
+

No (helpful) additional assumptions at all



Minimal group \mathcal{G}_{\min} .

Sending finite-dimensional quantum states



Quantum states $\rho_{n \times n}$ for some $n \in \mathbb{N}$

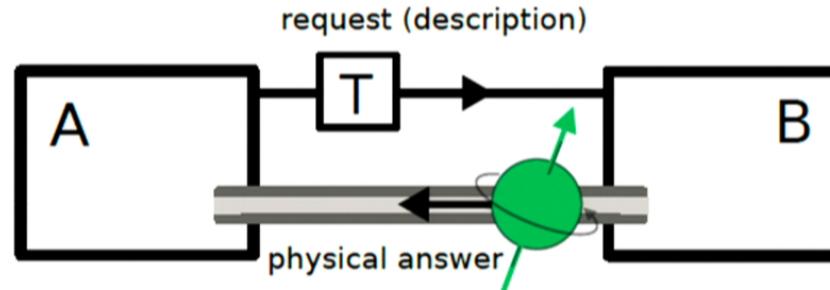
+

No (helpful) additional assumptions at all



Huge minimal group $\mathcal{G}_{\min} = \left\{ \bigotimes_{n \in \mathbb{N}} \bigotimes_{\mathcal{H}_n} (U_{\mathcal{H}_n} \bullet U_{\mathcal{H}_n}^{-1}) \right\},$
one unitary / antiunitary U for each Hilbert space.

Sending finite-dimensional quantum states



In our universe, we have additional structure:

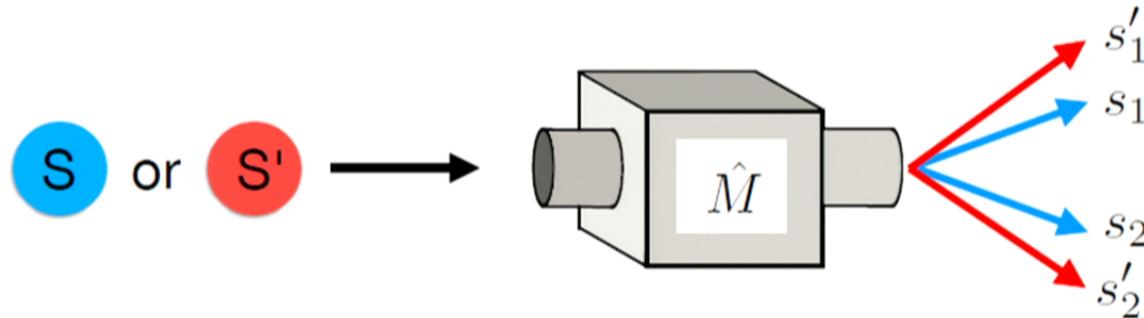
Can measure "the same" observable on different Hilbert spaces.

Does this help to get a smaller \mathcal{G}_{\min} ?

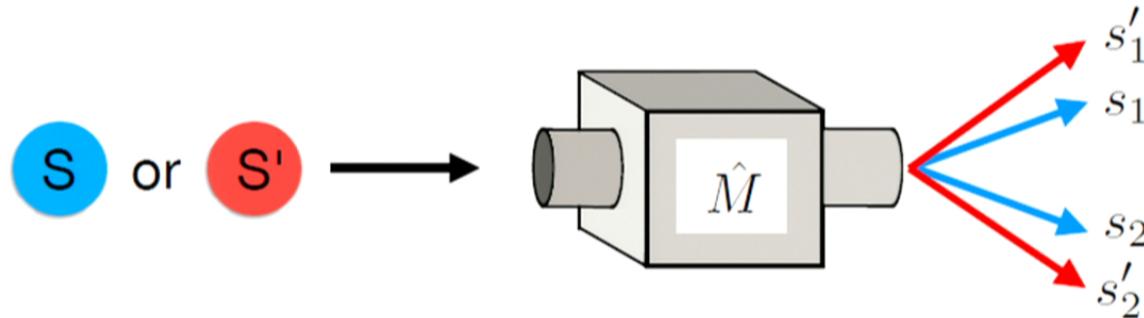
Yes!



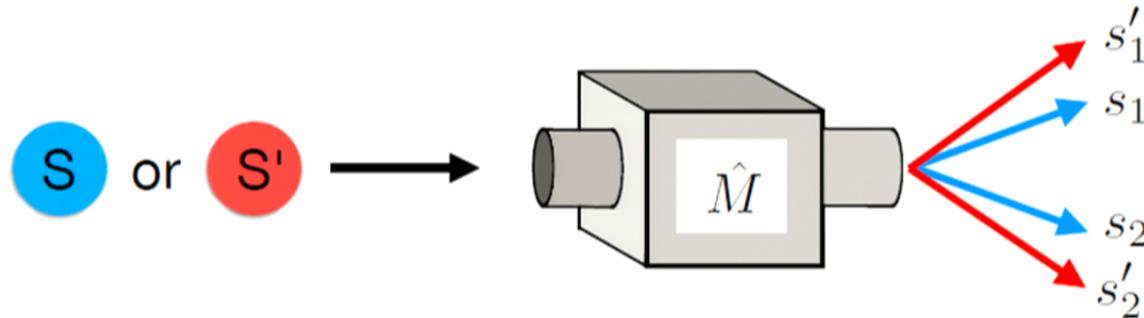
Same observable on different quantum systems S



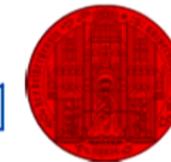
Same observable on different quantum systems S



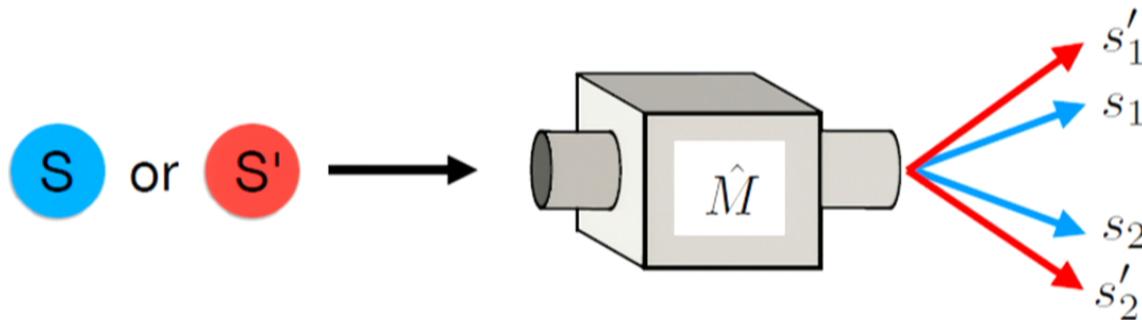
Same observable on different quantum systems S



Physical quantity \hat{M} is called (S, S') -co-measurable.



Same observable on different quantum systems S



Physical quantity \hat{M} is called (S, S') -co-measurable.

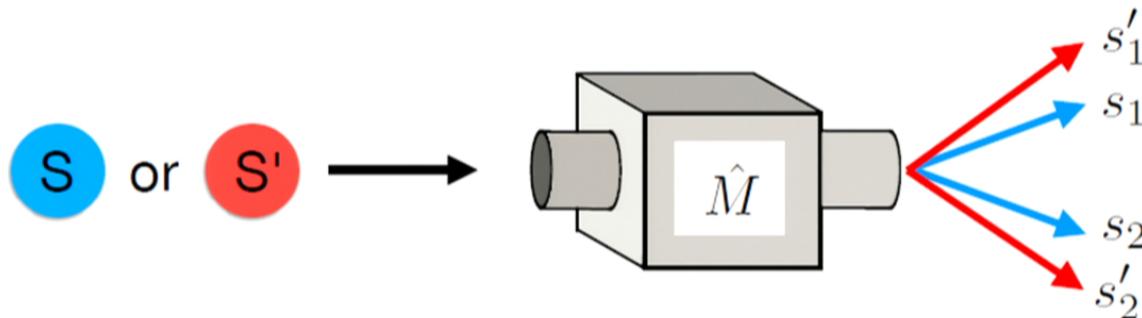
Example: S =spin-1/2 qubit, S' =spin-1 qutrit

$$\hat{M} = S_z, \text{ spin in } z\text{-direction}$$

$$\hat{M}(S) = \begin{pmatrix} 1/2 & 0 \\ 0 & -1/2 \end{pmatrix}, \quad \hat{M}(S') = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & -1 \end{pmatrix}.$$



Same observable on different quantum systems S



Physical quantity \hat{M} is called (S, S') -co-measurable.

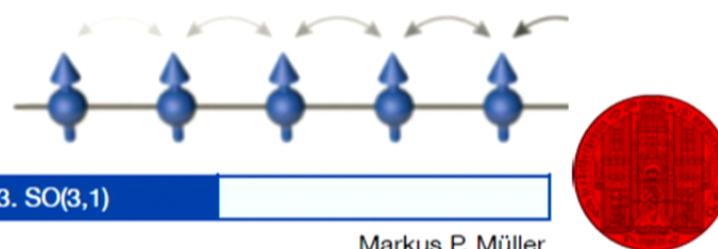
Example: S =spin-1/2 qubit, S' =spin-1 qutrit

$$\hat{M} = S_z, \text{ spin in } z\text{-direction}$$

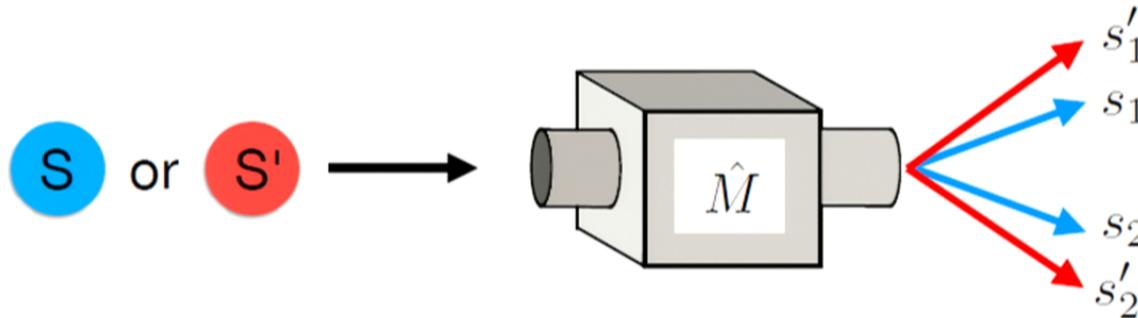
Possible origin: "natural interaction" between S and S' ?

\hat{M} conserved quantity?

Build measurement device
from many S -systems?



Same observable on different quantum systems S



Physical quantity \hat{M} is called (S, S') -co-measurable.

Example: S =spin-1/2 qubit, S' =spin-1 qutrit

$$\hat{M} = S_z, \text{ spin in } z\text{-direction}$$

Possible origin: "natural interaction" between S and S' ?

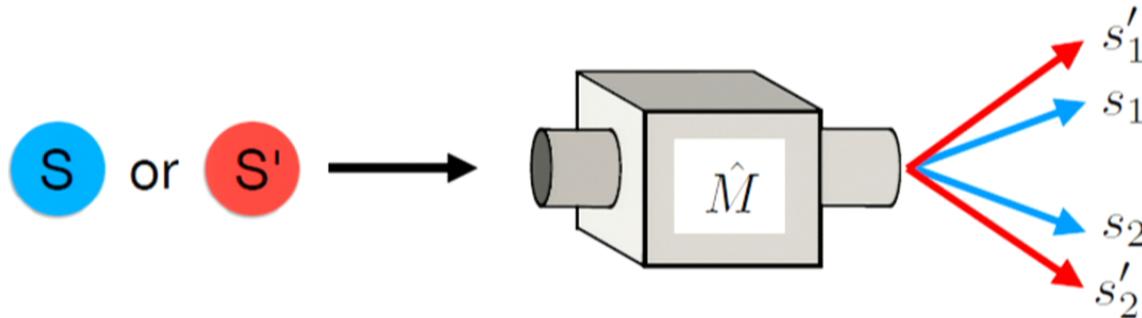
\hat{M} conserved due to

Build from any S -systems?

Doesn't matter for us.



Same observable on different quantum systems S



Physical quantity \hat{M} is called (S, S') -co-measurable.

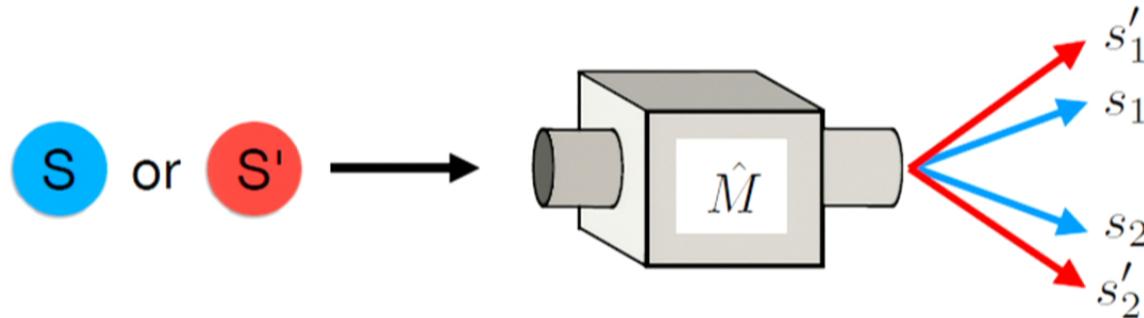
Set of co-measurable observables: consistency conditions.

$$\hat{M}_1 = \hat{S}_z; \quad \hat{M}_2 = \begin{cases} \hat{S}_z & \text{if spin-1/2 sent in} \\ \hat{S}_x & \text{otherwise} \end{cases}$$

cannot both be co-measurable.

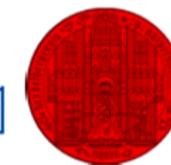


Same observable on different quantum systems S

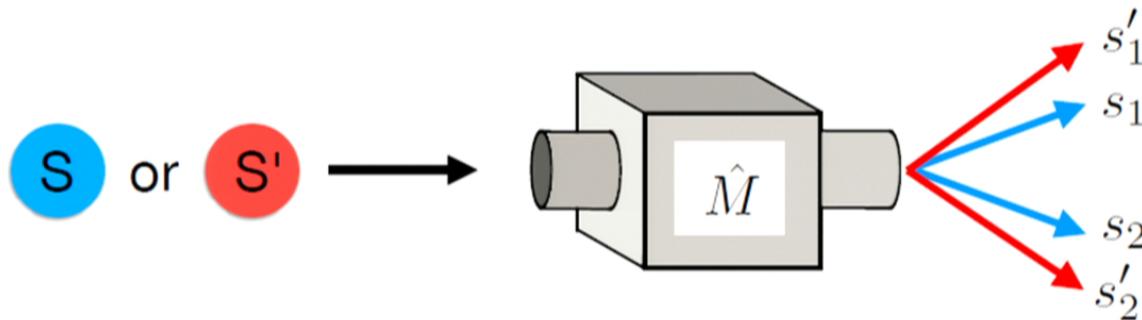


Consistency conditions for (S, S') -co-measurable obs.:

	3. SO(3,1)	
--	------------	--



Same observable on different quantum systems S

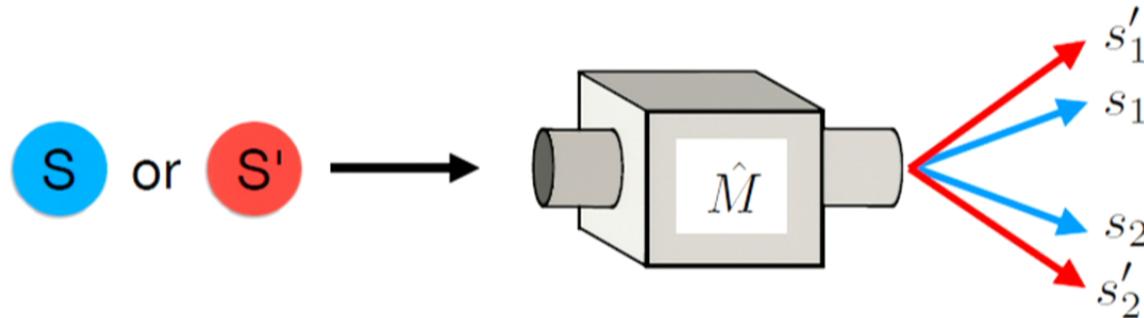


Consistency conditions for (S, S') -co-measurable obs.:

- $\hat{M}_n(S) \xrightarrow{n \rightarrow \infty} \hat{M}(S) \Rightarrow \hat{M}_n(S') \xrightarrow{n \rightarrow \infty} \hat{M}(S')$.
- $\hat{M}_1(S) \leq \hat{M}_2(S) \Rightarrow \hat{M}_1(S') \leq \hat{M}_2(S')$.
- $\hat{M}(S) = 0 \Rightarrow \hat{M}(S') = 0$.



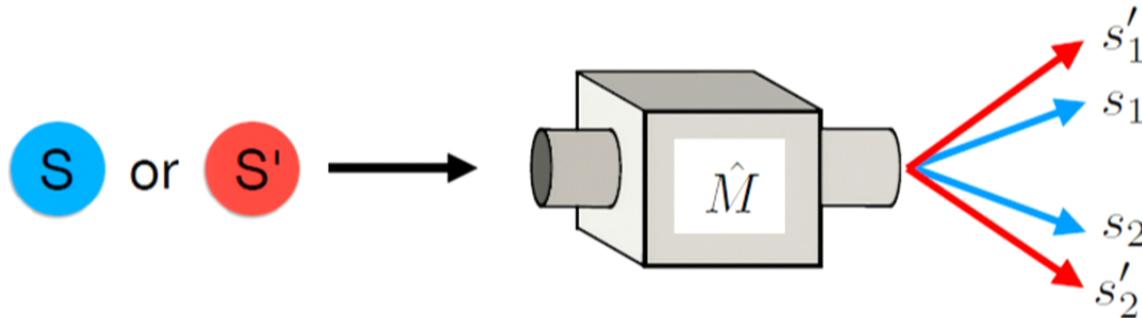
Same observable on different quantum systems S



Call S and S' equivalent if all their observables are (S, S') -co-measurable. Equivalence classes: \mathbf{S} .



Same observable on different quantum systems S

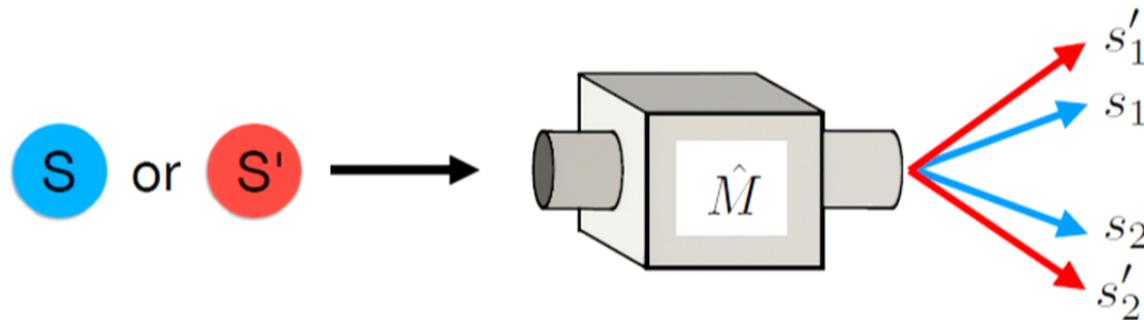


Consistency conditions for (S, S') -co-measurable obs.:

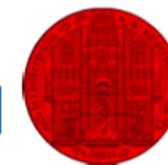
- $\hat{M}_n(S) \xrightarrow{n \rightarrow \infty} \hat{M}(S) \Rightarrow \hat{M}_n(S') \xrightarrow{n \rightarrow \infty} \hat{M}(S')$.
- $\hat{M}_1(S) \leq \hat{M}_2(S) \Rightarrow \hat{M}_1(S') \leq \hat{M}_2(S')$.
- $\hat{M}(S) = 0 \Rightarrow \hat{M}(S') = 0$.



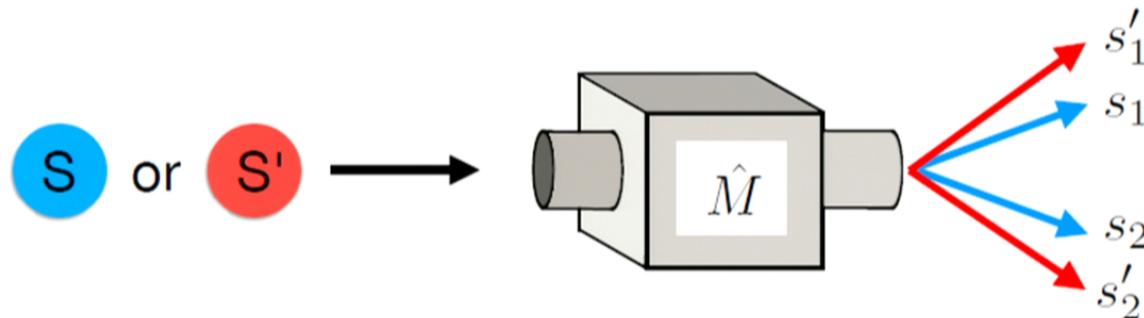
Same observable on different quantum systems S



Call S and S' equivalent if all their observables are (S, S') -co-measurable. Equivalence classes: \mathbf{S} .



Same observable on different quantum systems S



Call S and S' equivalent if all their observables are (S, S') -co-measurable. Equivalence classes: \mathbf{S} .

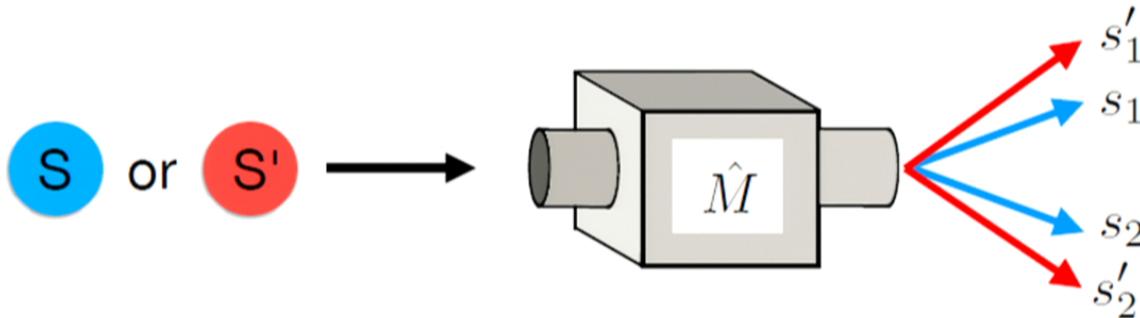
Theorem: Then $\dim S = \dim S'$, and $\hat{M}(S) = X \hat{M}(S')^T X^\dagger$.

invertible

possibly transpose



Same observable on different quantum systems S



Call S and S' equivalent if all their observables are (S, S') -co-measurable. Equivalence classes: \mathbf{S} .

Theorem: Then $\dim S = \dim S'$, and $\hat{M}(S) = X \hat{M}(S') X^\dagger$.

- Ex.:**
- $S = S'$ up to choice of basis ($\Rightarrow X$ unitary),
 - S and S' spins of electrons with different momenta.



Interaction graph

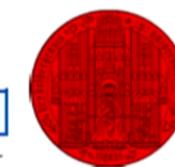
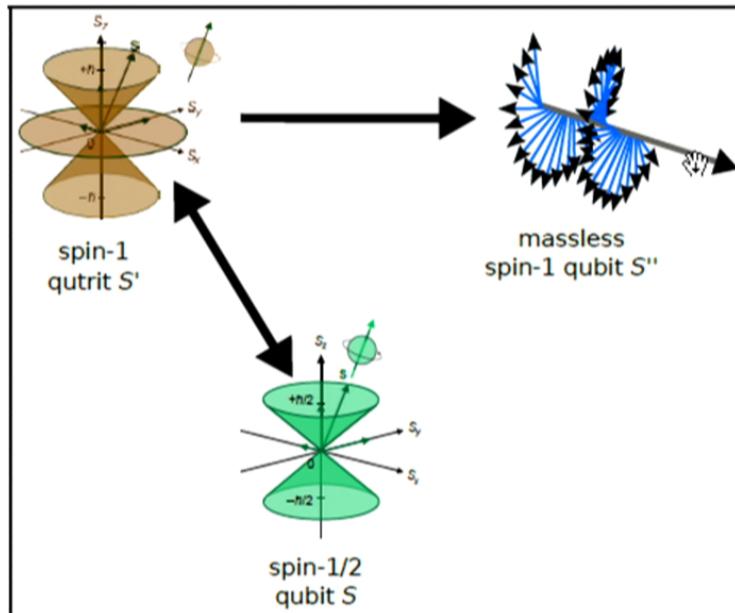
Draw an arrow from \mathbf{S} to \mathbf{S}' if the set of $(\mathbf{S}, \mathbf{S}')$ -co-measurable observables is [tomographically complete](#) on \mathbf{S}' .



Interaction graph

Draw an arrow from \mathbf{S} to \mathbf{S}' if the set of $(\mathbf{S}, \mathbf{S}')$ -co-measurable observables is **tomographically complete** on \mathbf{S}' .

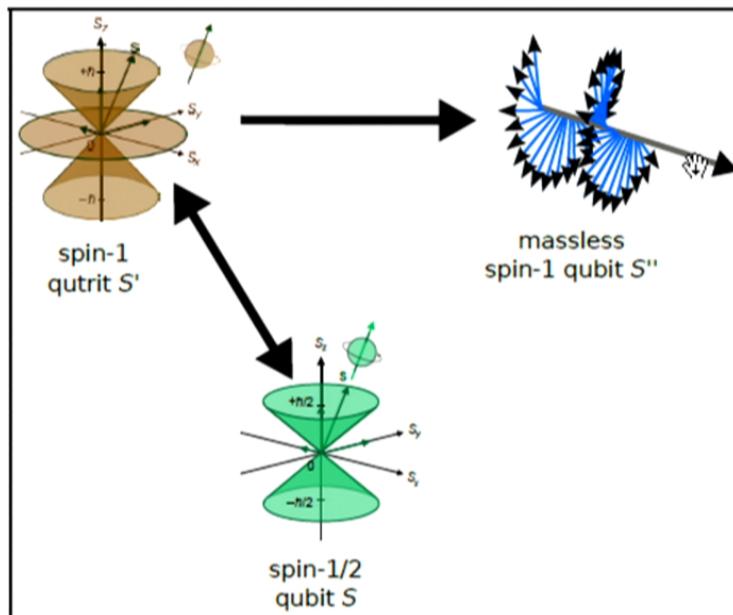
In our specific physical world:



Interaction graph

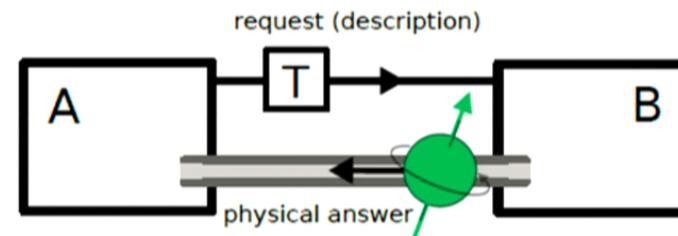
Draw an arrow from \mathbf{S} to \mathbf{S}' if the set of $(\mathbf{S}, \mathbf{S}')$ -co-measurable observables is **tomographically complete** on \mathbf{S}' .

In our specific physical world:



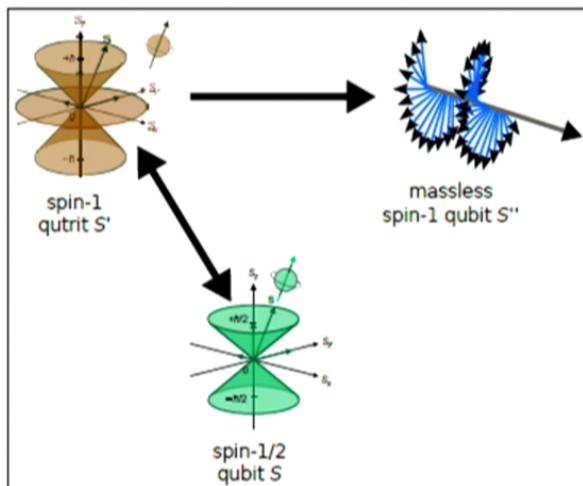
Interpretation:

Alice and Bob can lift their agreement on description of \mathbf{S} to description of \mathbf{S}' .



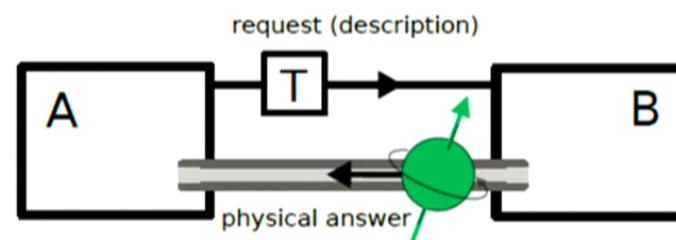
Interaction graph

A: Dear Bob, please send me the qutrit state $|\psi\rangle = (1, 0, -1)/\sqrt{2}!$



Interpretation:

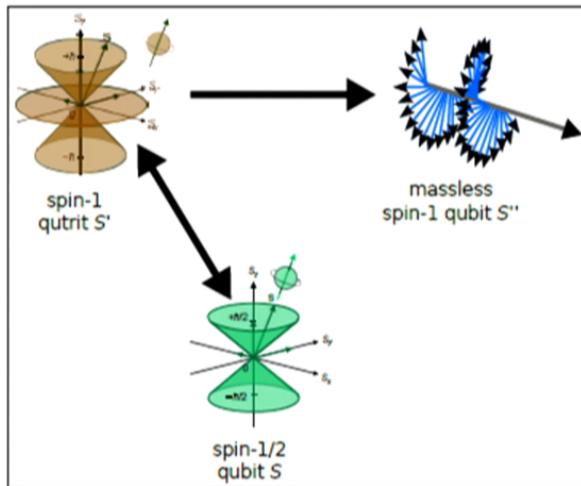
Alice and Bob can lift their agreement on description of **S** to description of **S'**.



Interaction graph

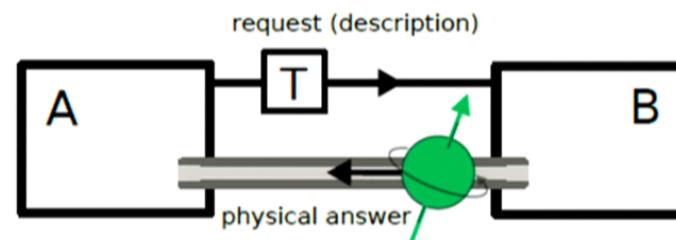
A: Dear Bob, please send me the qutrit state $|\psi\rangle = (1, 0, -1)/\sqrt{2}!$

B: Don't know how to do this... haven't set an \mathbf{S}' - reference frame!



Interpretation:

Alice and Bob can lift their agreement on description of \mathbf{S} to description of \mathbf{S}' .

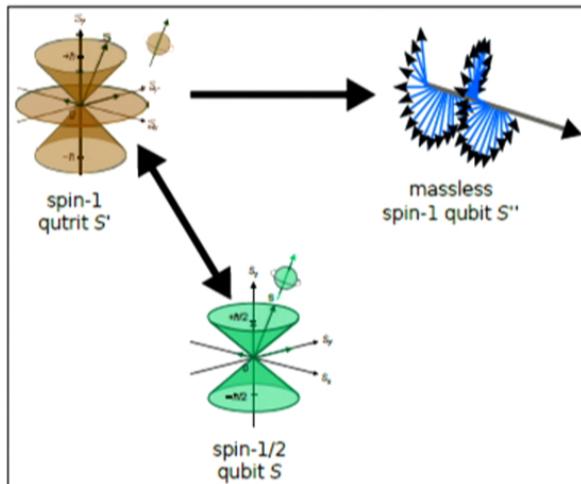


Interaction graph

A: Dear Bob, please send me the qutrit state $|\psi\rangle = (1, 0, -1)/\sqrt{2}!$

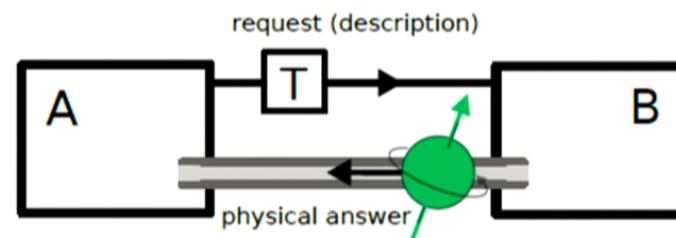
B: Don't know how to do this... haven't set an \mathbf{S}' - reference frame!

A: But we have for \mathbf{S} ! So please build the measurement devices with $\hat{M}_i(\mathbf{S}) = [\text{descriptions}]$. Then I mean the state $|\psi\rangle$ with the following outcome probabilities in these device: $[\text{numbers}]$.

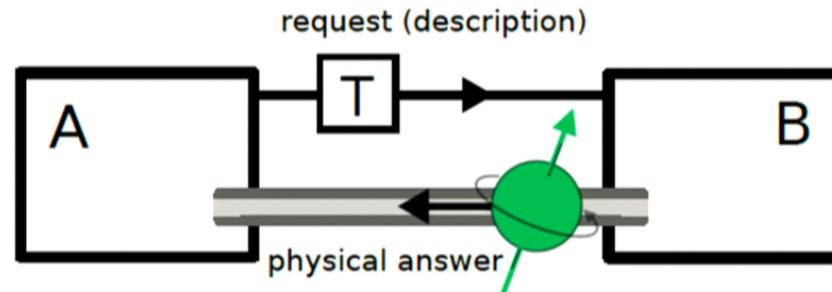


Interpretation:

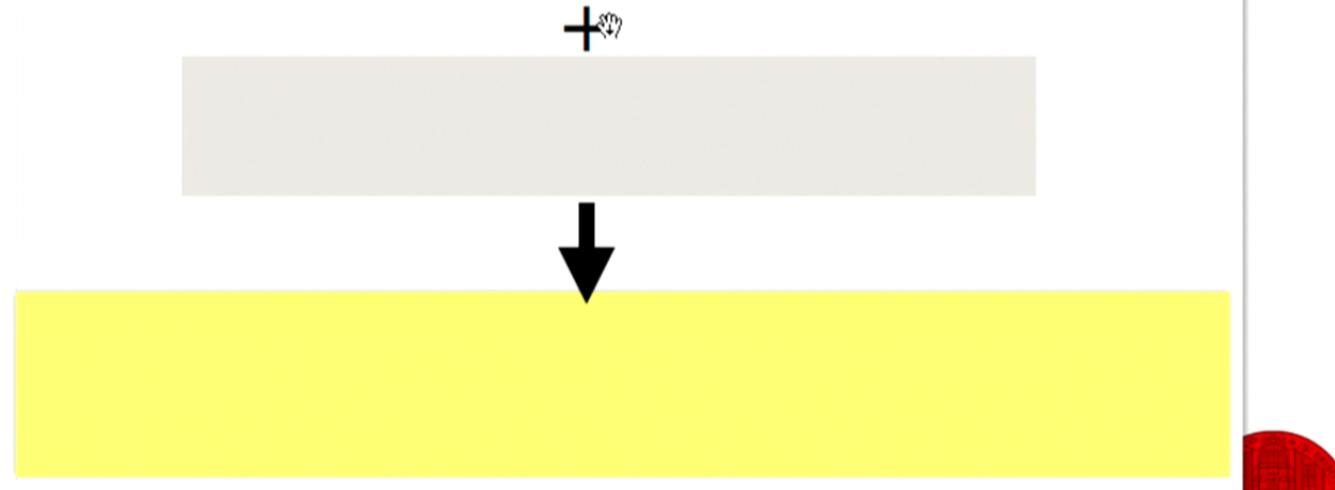
Alice and Bob can lift their agreement on description of \mathbf{S} to description of \mathbf{S}' .



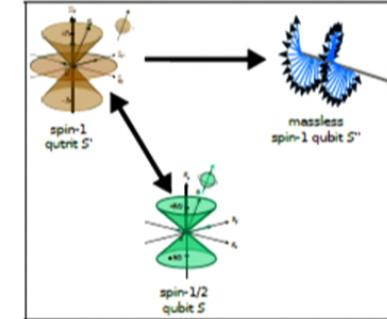
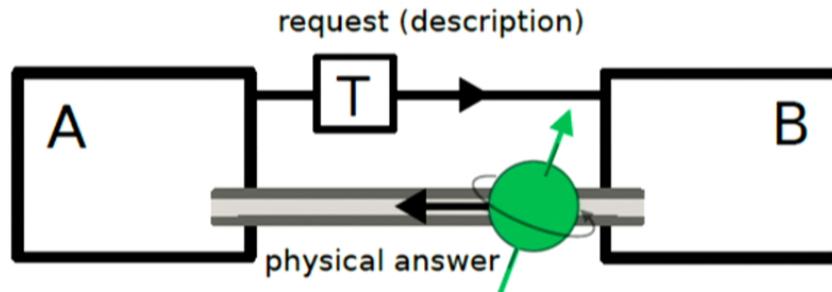
The communication task revisited



Quantum states ρ on any system S



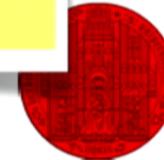
The communication task revisited



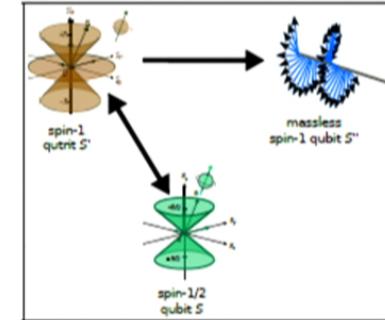
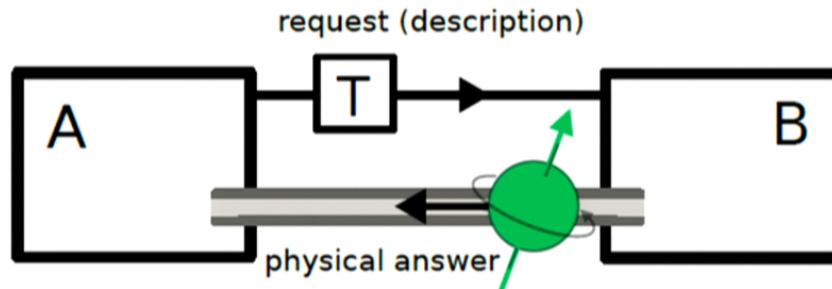
Quantum states ρ on any system S



Lots of co-measurable observables:
interaction graph is *connected*



The communication task revisited



Quantum states ρ on any system S



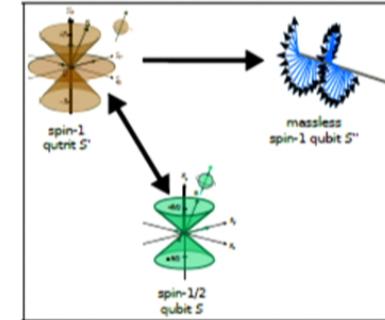
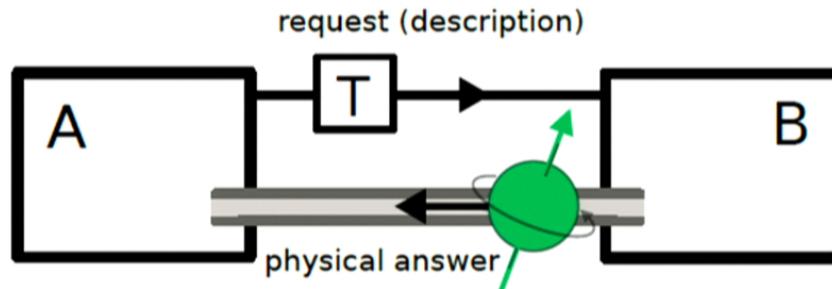
Lots of co-measurable observables:
interaction graph is *connected*



\mathcal{G}_{\min} is the symmetry group of the
smallest-dimensional root.



The communication task revisited



Quantum states ρ on any system S



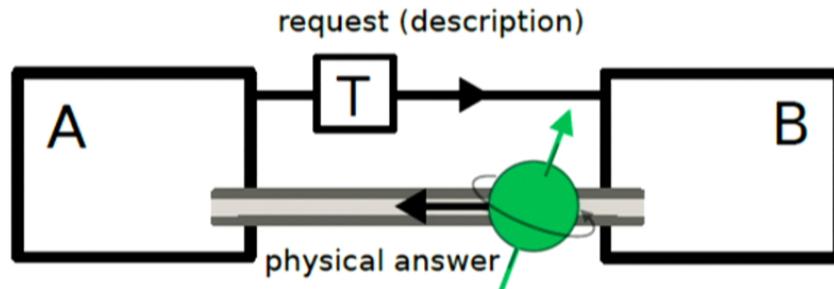
Minimality: interaction graph connected,
and there is a **qubit root**



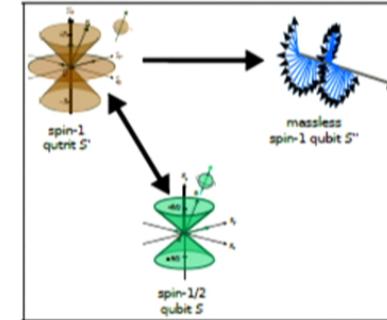
$$\begin{aligned}\mathcal{G}_{\min} &= \left\{ \hat{M} \mapsto X \hat{M} X^\dagger \text{ or } X \hat{M}^T X^\dagger \mid X \in \mathrm{GL}(2, \mathbb{C}) \right\} \\ &\simeq \mathbb{R}^+ \times \mathrm{O}^+(3, 1).\end{aligned}$$



Representations of the Lorentz group



$$\mathcal{G}_{\min} = \mathbb{R}^+ \times O^+(3, 1).$$



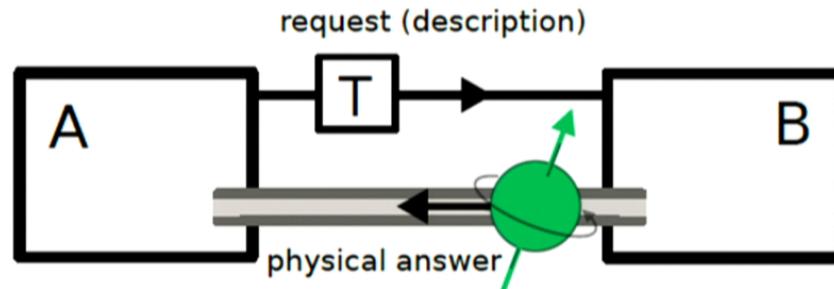
Theorem: Let \mathbf{S} be the "root qubit", and \mathbf{S}' any system.

- If all \mathbf{S} -observables are $(\mathbf{S}, \mathbf{S}')$ -co-measurable, then \mathbf{S}' carries a projective rep. of $SO^+(3, 1)$.

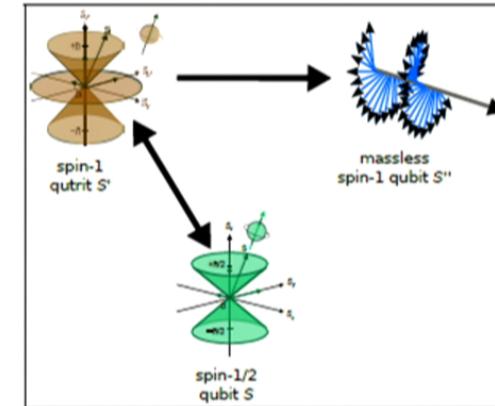
	3. $SO(3,1)$	
--	--------------	--



Representations of the Lorentz group



$$\mathcal{G}_{\min} = \mathbb{R}^+ \times O^+(3, 1).$$



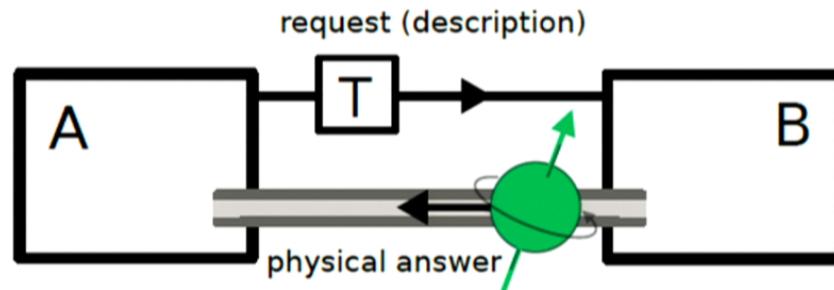
Theorem: Let \mathbf{S} be the "root qubit", and \mathbf{S}' any system.

- If all \mathbf{S} -observables are $(\mathbf{S}, \mathbf{S}')$ -co-measurable, then \mathbf{S}' carries a projective rep. of $SO^+(3, 1)$.
- All other \mathbf{S}' carry a proj. rep. of the subgroup that preserves the $(\mathbf{S}, \mathbf{S}')$ -co-measurable observables ("Wigner little groups").

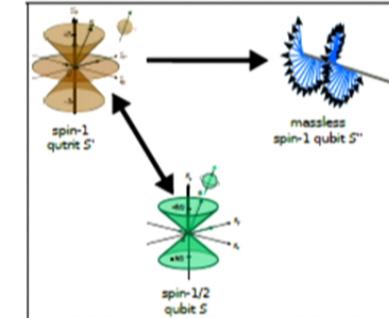
	3. $SO(3,1)$	
--	--------------	--



Interpretation of \mathcal{G}_{\min}



$$\mathcal{G}_{\min} = \mathbb{R}^+ \times O^+(3, 1).$$



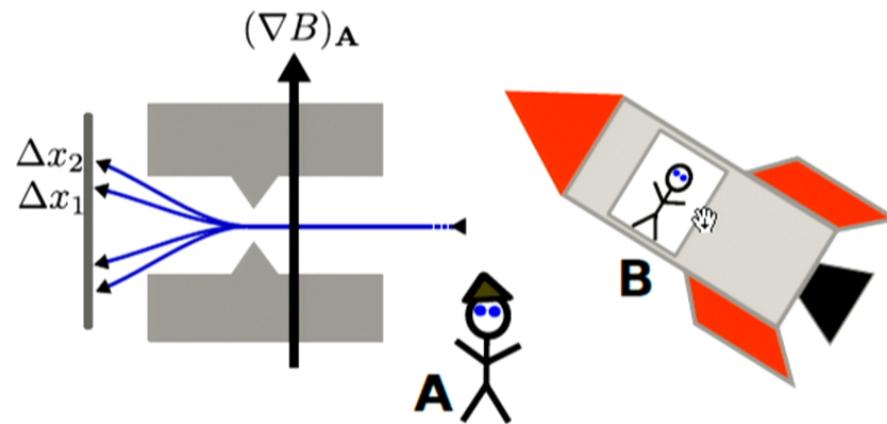
Interpretation of \mathbb{R}^+ -factor: Alice and Bob choose different units.



Relativistic Stern-Gerlach measurements

M. C. Palmer, M. Takahashi and H. F. Westman, *WKB analysis of relativistic Stern-Gerlach measurements*, Annals of Physics **336**, 505–516 (2013); [arXiv:1208.6434].

P. Hoehn and MM, in preparation



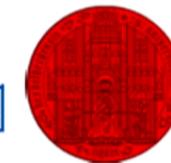
momentum p

$$\mathcal{H}_p \simeq \mathbb{C}^2$$

$S_\Lambda : \mathcal{H}_p \rightarrow \mathcal{H}_{\Lambda p}$ isometry

$$\hat{M}_B = S_\Lambda \hat{M}_A S_\Lambda^\dagger$$

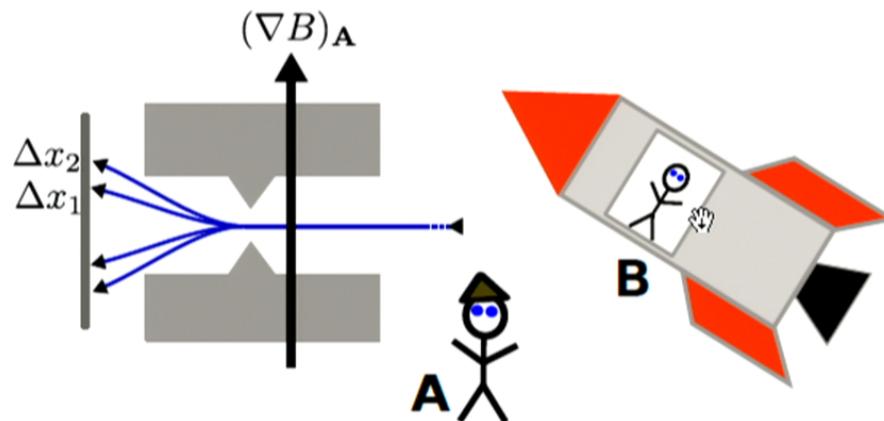
	3. SO(3,1)	
--	------------	--



Relativistic Stern-Gerlach measurements

M. C. Palmer, M. Takahashi and H. F. Westman, *WKB analysis of relativistic Stern-Gerlach measurements*, Annals of Physics **336**, 505–516 (2013); [arXiv:1208.6434].

P. Hoehn and MM, in preparation



momentum p

$$\mathcal{H}_p \simeq \mathbb{C}^2$$

$S_\Lambda : \mathcal{H}_p \rightarrow \mathcal{H}_{\Lambda p}$ isometry

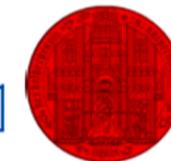
$$\hat{M}_B = S_\Lambda \hat{M}_A S_\Lambda^\dagger$$

Observed acceleration
of wave packets:

$$G^{A,B} = (G_0^{A,B}, \vec{G}^{A,B})$$

Eigenvalues of \hat{M} : $G_0^{A,B} \pm |\vec{G}^{A,B}|$

Rest frame of particle $G_0^{A,B} = 0$, but boosted observer sees
asymmetric deflections \rightarrow two different eigenvalues



Reversal of standard viewpoint

Standard point of view:

space-time symmetries → representations,
operational properties

Here:



Operational properties → representations,
space-time (?) symmetries

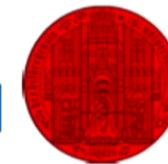


Conclusions



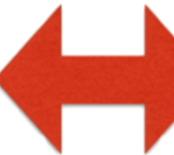
QT ←→ space-time

Interesting insights into architecture of physics; quantum gravity?
Common link: information theory / operationalism.



Conclusions



QT  space-time

Interesting insights into architecture of physics; quantum gravity?
Common link: information theory / operationalism.

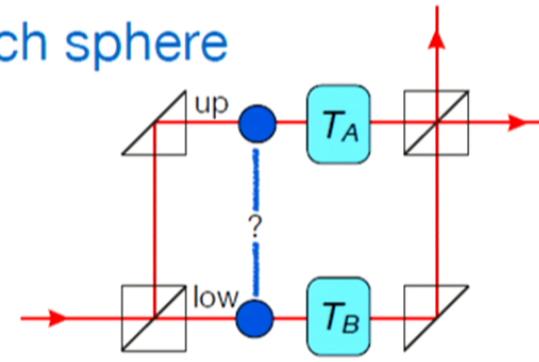
Lots of related work (Q-reference frames, QIT \leftrightarrow QG, ...)



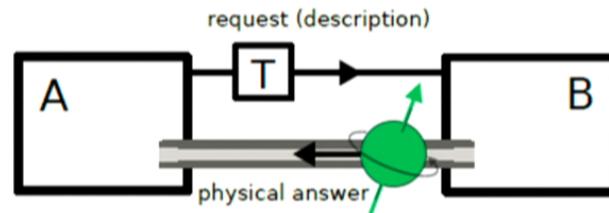
Conclusions

- Relativity of simultaneity on interferometer implies $d \leq 3$ / $d \leq 5$ for the Bloch sphere

Andy Garner, MM, Oscar Dahlsten,
arXiv:1412.7112



- Quantum communication task with operational background assumptions (satisfied in our world) yields the Lorentz transformations (and rep's)



Philipp Hoehn, MM,
arXiv:1412.8462.

