

Title: Expansion of Lie algebras and accidental symmetries in Lovelock theories

Date: Apr 28, 2015 02:00 PM

URL: <http://pirsa.org/15040071>

Abstract: <p>Main properties of generalized contraction methods of Lie algebras, known also as expansion methods, are briefly introduced. Between some of their physical applications, one might study the nature of solutions in theories constructed with those expanded algebras. In particular, as we are interested in solutions that could be relevant in the context of AdS/CFT and Holographic Superconductors, we would like to study the holographic QFT dual to Chern-Simons gravity for an expansion of AdS algebra. As a first step, we studied charged static spherically symmetric BH solutions of a CS theory for the most simple extension of AdS symmetry:  $AdS_{\tilde{A}}-U(1)$ . It is shown that in this kind of higher dimensional gravity, degeneracy in some sectors of the space of solutions can appear. In fact, arbitrary functions remain undetermined after the field equations are imposed. This is related to an increase in local symmetries and it is shown that the knowledge of these "accidental symmetries" can help to formulate a simple criterion that avoids unwanted degenerate ansätze. Finally, main properties of Pure Lovelock gravity are presented and some issues about black hole solutions this theory are also discussed.</p>

# Expansion of Lie algebras and accidental symmetries in Lovelock theories

Nelson Merino

Pontificia Universidad Católica de Valparaíso

April 2015



# Contents

- ① Introduction
- ② Expansion methods
- ③ Accidental symmetries
  - Action, field equations and some solutions
  - Torsion and degeneracy
  - Accidental symmetries
- ④ Conclusions

# Introduction

We use first order Einstein-Cartan formulation of gravity

$$(e_{\mu}^a, \omega_{\mu}^{ab}; g_{\mu\nu} = \eta_{ab} e_{\mu}^a e_{\nu}^b; R^{ab} = (d\omega + \omega^2)^{ab}, T^a = D e^a).$$

For example, in  $D = 4$  Einstein-Cartán theory

$$\mathcal{L}_{EC}(e, \omega) = \kappa \left( \epsilon_{abcd} R^{ab} e^c e^d - \frac{\Lambda}{6} e^a e^b e^c e^d \right); \quad \kappa = \frac{1}{32\pi G},$$

is equivalent to Einstein-Hilbert theory

$$(g_{\mu\nu}, \Gamma_{\mu\nu}^{\lambda}(g), \text{ under assumption } T_{\mu\nu}^{\lambda} = 0, \nabla^{\Gamma} g_{\mu\nu} = 0)$$

# Introduction

We use first order Einstein-Cartan formulation of gravity

$$(e_{\mu}^a, \omega_{\mu}^{ab}; g_{\mu\nu} = \eta_{ab} e_{\mu}^a e_{\nu}^b; R^{ab} = (d\omega + \omega^2)^{ab}, T^a = D e^a).$$

For example, in  $D = 4$  Einstein-Cartán theory

$$\mathcal{L}_{EC}(e, \omega) = \kappa \left( \epsilon_{abcd} R^{ab} e^c e^d - \frac{\Lambda}{6} e^a e^b e^c e^d \right); \quad \kappa = \frac{1}{32\pi G},$$

is equivalent to Einstein-Hilbert theory

$$(g_{\mu\nu}, \Gamma_{\mu\nu}^{\lambda}(g), \text{ under assumption } T_{\mu\nu}^{\lambda} = 0, \nabla^{\Gamma} g_{\mu\nu} = 0)$$

$$\mathcal{L}_{EH}(g, \Gamma(g)) = -\kappa \sqrt{-g} (R - 2\Lambda).$$

Generalization to  $D$ -dimensions: Lanczos-Lovelock (L-L) gravity

$$\mathcal{L} = \sum_{p=0}^{\lfloor \frac{D}{2} \rfloor} \alpha_p \mathcal{L}^{(p)},$$

$$\mathcal{L}^{(p)} = \epsilon_{a_1 \dots a_D} R^{a_1 a_2} \dots R^{a_{2p-1} a_{2p}} e^{a_{2p+1}} \dots e^{a_D}.$$

Requiring the theory to have the maximum possible number of degrees of freedom, leads to a special choosing<sup>1</sup> of  $\alpha_p$  which in even  $D$  gives to **Born-Infeld (BI)** gravity; while in **odd**  $D$  it gives **AdS-Chern-Simons (CS)** gravity, described by

---

<sup>1</sup>R. Troncoso, J. Zanelli, Class. Quantum Grav. 17 (2000) 4451.

<sup>2</sup>Dadich et al, arXiv:1006.0337, 1201.4994.

Requiring the theory to have the maximum possible number of degrees of freedom, leads to a special choosing<sup>1</sup> of  $\alpha_p$  which in even  $D$  gives to **Born-Infeld (BI)** gravity; while in **odd**  $D$  it gives **AdS-Chern-Simons (CS)** gravity, described by

$$\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = d\mathcal{L}_{\text{CS}}(\mathbf{A}),$$

---

<sup>1</sup>R. Troncoso, J. Zanelli, Class. Quantum Grav. 17 (2000) 4451.

<sup>2</sup>Dadich et al, arXiv:1006.0337, 1201.4994.

Requiring the theory to have the maximum possible number of degrees of freedom, leads to a special choosing<sup>1</sup> of  $\alpha_p$  which in even  $D$  gives to **Born-Infeld (BI)** gravity; while in **odd**  $D$  it gives **AdS-Chern-Simons (CS)** gravity, described by

$$\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = d\mathcal{L}_{\text{CS}}(\mathbf{A}), \quad \text{with } \mathbf{A} = A^A T_A = \frac{1}{2} \omega^{ab} J_{ab} + \frac{1}{\ell} e^a P_a,$$

---

<sup>1</sup>R. Troncoso, J. Zanelli, Class. Quantum Grav. 17 (2000) 4451.

<sup>2</sup>Dadich et al, arXiv:1006.0337, 1201.4994.

Requiring the theory to have the maximum possible number of degrees of freedom, leads to a special choosing<sup>1</sup> of  $\alpha_p$  which in even  $D$  gives to **Born-Infeld (BI)** gravity; while in **odd**  $D$  it gives **AdS-Chern-Simons (CS)** gravity, described by

$$\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = d\mathcal{L}_{\text{CS}}(\mathbf{A}), \quad \text{with } \mathbf{A} = A^A T_A = \frac{1}{2} \omega^{ab} J_{ab} + \frac{1}{\ell} e^a P_a,$$

$$\langle \{J_{ab}, P_a\} \rangle = \mathfrak{so}(D-1, 2),$$

---

<sup>1</sup>R. Troncoso, J. Zanelli, Class. Quantum Grav. 17 (2000) 4451.

<sup>2</sup>Dadich et al, arXiv:1006.0337, 1201.4994.



Requiring the theory to have the maximum possible number of degrees of freedom, leads to a special choosing<sup>1</sup> of  $\alpha_p$  which in even  $D$  gives to **Born-Infeld (BI)** gravity; while in **odd**  $D$  it gives **AdS-Chern-Simons (CS)** gravity, described by

$$\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = d\mathcal{L}_{\text{CS}}(\mathbf{A}), \quad \text{with } \mathbf{A} = A^A T_A = \frac{1}{2} \omega^{ab} J_{ab} + \frac{1}{\ell} e^a P_a,$$

$$\langle \{J_{ab}, P_a\} \rangle = \mathfrak{so}(D-1, 2), \quad \mathbf{F} = d\mathbf{A} + \mathbf{A} \wedge \mathbf{A} = \frac{1}{2} \left( R^{ab} + \frac{1}{\ell^2} e^a e^b \right) J_{ab} + \frac{1}{\ell} T^a P_a$$

---

<sup>1</sup>R. Troncoso, J. Zanelli, *Class. Quantum Grav.* 17 (2000) 4451.

<sup>2</sup>Dadich et al, arXiv:1006.0337, 1201.4994.



# Introduction

We use first order Einstein-Cartan formulation of gravity

$$(e_{\mu}^a, \omega_{\mu}^{ab}; g_{\mu\nu} = \eta_{ab}e_{\mu}^ae_{\nu}^b; R^{ab} = (d\omega + \omega^2)^{ab}, T^a = De^a).$$

For example, in  $D = 4$  Einstein-Cartán theory

$$\mathcal{L}_{EC}(e, \omega) = \kappa \left( \epsilon_{abcd} R^{ab} e^c e^d - \frac{\Lambda}{6} e^a e^b e^c e^d \right); \quad \kappa = \frac{1}{32\pi G},$$

is equivalent to Einstein-Hilbert theory

$$(g_{\mu\nu}, \Gamma_{\mu\nu}^{\lambda}(g), \text{ under assumption } T_{\mu\nu}^{\lambda} = 0, \nabla^{\Gamma} g_{\mu\nu} = 0)$$

$$\mathcal{L}_{EH}(g, \Gamma(g)) = -\kappa \sqrt{-g} (R - 2\Lambda).$$

Generalization to  $D$ -dimensions: Lanczos-Lovelock (L-L) gravity

$$\mathcal{L} = \sum_{p=0}^{\lfloor \frac{D}{2} \rfloor} \alpha_p \mathcal{L}^{(p)},$$

$$\mathcal{L}^{(p)} = \epsilon_{a_1 \dots a_D} R^{a_1 a_2} \dots R^{a_{2p-1} a_{2p}} e^{a_{2p+1}} \dots e^{a_D}.$$

Requiring the theory to have the maximum possible number of degrees of freedom, leads to a special choosing<sup>1</sup> of  $\alpha_p$  which in even  $D$  gives to **Born-Infeld (BI)** gravity; while in **odd**  $D$  it gives **AdS-Chern-Simons (CS)** gravity, described by

$$\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = d\mathcal{L}_{\text{CS}}(\mathbf{A}), \quad \text{with } \mathbf{A} = A^A T_A = \frac{1}{2} \omega^{ab} J_{ab} + \frac{1}{\ell} e^a P_a,$$

$$\langle \{J_{ab}, P_a\} \rangle = \mathfrak{so}(D-1, 2), \quad \mathbf{F} = d\mathbf{A} + \mathbf{A} \wedge \mathbf{A} = \frac{1}{2} \left( R^{ab} + \frac{1}{\ell^2} e^a e^b \right) J_{ab} + \frac{1}{\ell} T^a P_a$$

---


$$\text{Ex. in } D = 5 : \mathcal{L}_{\text{CS}}(\mathbf{A}) = k \epsilon_{abcde} \left( \frac{1}{5!^5} e^a e^b e^c e^d e^e + \frac{2}{3!^3} R^{ab} e^c e^d e^e + \frac{1}{!} R^{ab} R^{cd} e^e \right).$$


---

<sup>1</sup>R. Troncoso, J. Zanelli, *Class. Quantum Grav.* 17 (2000) 4451.

<sup>2</sup>Dadich et al, arXiv:1006.0337, 1201.4994.

Requiring the theory to have the maximum possible number of degrees of freedom, leads to a special choosing<sup>1</sup> of  $\alpha_p$  which in even  $D$  gives to **Born-Infeld (BI)** gravity; while in odd  $D$  it gives **AdS-Chern-Simons (CS)** gravity, described by

$$\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = d\mathcal{L}_{\text{CS}}(\mathbf{A}), \quad \text{with } \mathbf{A} = A^A T_A = \frac{1}{2} \omega^{ab} J_{ab} + \frac{1}{\ell} e^a P_a,$$

$$\langle \{J_{ab}, P_a\} \rangle = \mathfrak{so}(D-1, 2), \quad \mathbf{F} = d\mathbf{A} + \mathbf{A} \wedge \mathbf{A} = \frac{1}{2} \left( R^{ab} + \frac{1}{\ell^2} e^a e^b \right) J_{ab} + \frac{1}{\ell} T^a P_a$$

---


$$\text{Ex. in } D = 5 : \mathcal{L}_{\text{CS}}(\mathbf{A}) = k \epsilon_{abcde} \left( \frac{1}{5!^5} e^a e^b e^c e^d e^e + \frac{2}{3!^3} R^{ab} e^c e^d e^e + \frac{1}{!} R^{ab} R^{cd} e^e \right).$$


---

Another interesting family is **Pure Lovelock** theory<sup>2</sup>, which have only  $\alpha_0$  and  $\alpha_N$  ( $N = [(D-1)/2]$ ) non-vanishing.

---

<sup>1</sup>R. Troncoso, J. Zanelli, *Class. Quantum Grav.* 17 (2000) 4451.

<sup>2</sup>Dadich et al, arXiv:1006.0337, 1201.4994.

$$R \dots R e' + e \dots e$$

$$R \dots R e e + e \dots e$$



$\alpha_1 R \dots R e + \alpha_2 e \dots e$  } odd  
 $\alpha_2 R \dots R e e + \alpha_1 e \dots e$  } even

We will see here that some L-L theories might have **degenerated sectors** in the space of solutions.

We will see here that some L-L theories might have **degenerated sectors** in the space of solutions.

For example, consider (in the torsionless sector:  $T^a = De^a = 0$ ) the following ansatz,

$$ds^2 = -f^2(r)dt^2 + \frac{dr^2}{f^2(r)} + r^2 d\Omega_{D-2}^2 \quad ; \quad \text{with } f^2 = 1 - \psi(r)r^2 .$$

We will see here that some L-L theories might have **degenerated sectors** in the space of solutions.

For example, consider (in the torsionless sector:  $T^a = De^a = 0$ ) the following ansatz,

$$ds^2 = -f^2(r)dt^2 + \frac{dr^2}{f^2(r)} + r^2 d\Omega_{D-2}^2 \quad ; \quad \text{with } f^2 = 1 - \psi(r)r^2 .$$

L-L field equations reduces to,

$$\mathcal{F}(\psi(r)) = \sum_p \tilde{\alpha}_p \psi^p = \frac{\mu}{r^{D-1}} ,$$

where  $\mu$  is the mass parameter and  $\tilde{\alpha}_p = \alpha_p 2\kappa (D-2p)(D-2)!$ .



When coupling constants are such that  $F(\psi)$  has a degenerate zero, the field equations leaves  $g_{tt}$  arbitrary (and were called "geometrically free" solutions by J. T. Wheeler).

---

<sup>3</sup>Bañados et al (1994), Crisostomo et al (2000).

When coupling constants are such that  $F(\psi)$  has a degenerate zero, the field equations leaves  $g_{tt}$  arbitrary (and were called "geometrically free" solutions by J. T. Wheeler).

There are two ways to reduce arbitrariness of the theory:

---

<sup>3</sup>Bañados et al (1994), Crisostomo et al (2000).

When coupling constants are such that  $F(\psi)$  has a degenerate zero, the field equations leaves  $g_{tt}$  arbitrary (and were called "geometrically free" solutions by J. T. Wheeler).

There are two ways to reduce arbitrariness of the theory:

- To assume fully degeneracy of that polynomial, i.e. that there is a unique vacua

$$(\psi - \beta)^N = 0,$$

which amounts to choose  $\alpha_p$  as in the BI or CS case until some order  $k \leq [D/2]$ .

---

<sup>3</sup>Bañados et al (1994), Crisostomo et al (2000).

When coupling constants are such that  $F(\psi)$  has a degenerate zero, the field equations leaves  $g_{tt}$  arbitrary (and were called "geometrically free" solutions by J. T. Wheeler).

There are two ways to reduce arbitrariness of the theory:

- To assume fully degeneracy of that polynomial, i.e. that there is a unique vacua

$$(\psi - \beta)^N = 0,$$

which amounts to choose  $\alpha_p$  as in the BI or CS case until some order  $k \leq [D/2]$ .

They are known as dimensionally continued BHs<sup>3</sup>,

---

<sup>3</sup>Bañados et al (1994), Crisostomo et al (2000).

When coupling constants are such that  $F(\psi)$  has a degenerate zero, the field equations leaves  $g_{tt}$  arbitrary (and were called "geometrically free" solutions by J. T. Wheeler).

There are two ways to reduce arbitrariness of the theory:

- To assume full degeneracy of that polynomial, i.e. that there is a unique vacua

$$(\psi - \beta)^N = 0,$$

which amounts to choose  $\alpha_p$  as in the BI or CS case until some order  $k \leq [D/2]$ .

They are known as dimensionally continued BHs<sup>3</sup>,

$$f^2 = 1 - \beta r^2 - r^2 \left( \frac{\mu}{r^{D-1}} \right)^{\frac{1}{N}} \xrightarrow{\text{large } r} f^2 = 1 - ar^2 - \frac{b}{r^{D-3}}, \text{ where } \begin{cases} a \sim \Lambda \\ b \sim \mu \end{cases}$$

Schwarzschild dS/AdS BH

<sup>3</sup>Bañados et al (1994), Crisostomo et al (2000).

- The other way is to restrict to non vanishing coefficients  $\alpha_0$  and  $\alpha_N$  (with  $N = [(D - 1) / 2]$ ), known as **Pure Lovelock** case<sup>4</sup>:

---

<sup>4</sup>Dadich et al, arXiv:1006.0337, 1201.4994.

- The other way is to restrict to non vanishing coefficients  $\alpha_0$  and  $\alpha_N$  (with  $N = [(D - 1) / 2]$ ), known as **Pure Lovelock** case<sup>4</sup>:

$$\psi^N - \alpha = 0,$$

---

<sup>4</sup>Dadich et al, arXiv:1006.0337, 1201.4994.

- The other way is to restrict to non vanishing coefficients  $\alpha_0$  and  $\alpha_N$  (with  $N = [(D - 1) / 2]$ ), known as **Pure Lovelock** case<sup>4</sup>:

$$\psi^N - \alpha = 0,$$

$$f^2 = 1 - r^2 \left( \alpha + \frac{\mu}{r^{D-1}} \right)^{\frac{1}{N}},$$

---

<sup>4</sup>Dadich et al, arXiv:1006.0337, 1201.4994.



- The other way is to restrict to non vanishing coefficients  $\alpha_0$  and  $\alpha_N$  (with  $N = [(D - 1) / 2]$ ), known as **Pure Lovelock** case<sup>4</sup>:

$$\psi^N - \alpha = 0,$$

$$f^2 = 1 - r^2 \left( \alpha + \frac{\mu}{r^{D-1}} \right)^{\frac{1}{N}},$$

→ dS/AdS Schwarzschild BH  $\left( f^2 = 1 - ar^2 - \frac{b}{r^{D-3}} \right),$   
large  $r$

→ dimensionally continued BH.  
 $r \rightarrow r_h$


---

<sup>4</sup>Dadich et al, arXiv:1006.0337, 1201.4994.

This work also deals with [expansion methods](#) of Lie algebras<sup>5</sup> and gravity theories constructed with them (for which BH and a cosmological solutions have already been studied<sup>6</sup>).

---

<sup>5</sup>Hatsuda-Sakaguchi (2003); de Azcarraga, Izquierdo, Picon, Varela (2003); Izaurieta, Rodriguez, Salgado (2006).


<sup>6</sup>Quinzacara *et al*, arXiv:1401.1797; Crisóstomo *et al*, arXiv:1401.2128. 

This work also deals with [expansion methods](#) of Lie algebras<sup>5</sup> and gravity theories constructed with them (for which BH and a cosmological solutions have already been studied<sup>6</sup>).

With O. Miskovic (PUCV, Chile) → to explore solutions of those theories that could be relevant in AdS/CFT context, because:

---

<sup>5</sup>Hatsuda-Sakaguchi (2003); de Azcarraga, Izquierdo, Picon, Varela (2003); Izaurieta, Rodriguez, Salgado (2006).

<sup>6</sup>Quinzacara *et al*, arXiv:1401.1797; Crisóstomo *et al*, arXiv:1401.2128. 

$L_{CS}(A)$

$$A = A^p \underbrace{T_p}_G$$



$L_{CS}(A)$

$$A = A^p \underbrace{T_a}$$

Ads  $\rightarrow$  G

$$L_{CS}(A, G) \rightarrow A \quad EN$$

$$A = A^p \underbrace{T_a}$$

$$A \rightarrow G$$




This work also deals with [expansion methods](#) of Lie algebras<sup>5</sup> and gravity theories constructed with them (for which BH and a cosmological solutions have already been studied<sup>6</sup>).

With O. Miskovic (PUCV, Chile) → to explore solutions of those theories that could be relevant in AdS/CFT context, because:

- Four-dimensional holographic QFT at finite T is dual to five-dimensional BH in AAdS gravity
- Non-perturbative phenomena and phase transitions in QFT can be analyzed holographically

---

<sup>5</sup>Hatsuda-Sakaguchi (2003); de Azcarraga, Izquierdo, Picon, Varela (2003); Izaurieta, Rodriguez, Salgado (2006).

<sup>6</sup>Quinzacara *et al*, arXiv:1401.1797; Crisóstomo *et al*, arXiv:1401.2128. 


This work also deals with [expansion methods](#) of Lie algebras<sup>5</sup> and gravity theories constructed with them (for which BH and a cosmological solutions have already been studied<sup>6</sup>).

With O. Miskovic (PUCV, Chile) → to explore solutions of those theories that could be relevant in AdS/CFT context, because:

- Four-dimensional holographic QFT at finite T is dual to five-dimensional BH in AAdS gravity
- Non-perturbative phenomena and phase transitions in QFT can be analyzed holographically
- There are some unconventional superconductors (discovered '78 and 86') whose theoretical description so far is based on holography [e.g., arXiv:1308.2976 for a recent review].

---

<sup>5</sup>Hatsuda-Sakaguchi (2003); de Azcarraga, Izquierdo, Picon, Varela (2003); Izaurieta, Rodriguez, Salgado (2006).

<sup>6</sup>Quinzacara *et al*, arXiv:1401.1797; Crisóstomo *et al*, arXiv:1401.2128. 



- Field content: Gravitational field + Electromagnetic field + Symmetry breaking matter fields

- Field content: Gravitational field + Electromagnetic field + Symmetry breaking matter fields

---

**Ex. 1:** scalar field  $\Phi$  in AdS gravity is related to the order parameter  $\mathcal{O}$  in QFT

- Field content: Gravitational field + Electromagnetic field + Symmetry breaking matter fields

---

**Ex. 1:** scalar field  $\Phi$  in AdS gravity is related to the order parameter  $\mathcal{O}$  in QFT

**Ex. 2:** Gauss-Bonnet holographic superconductor + Maxwell field + Minimally coupled scalar field [e.g., arXiv:1009.1991]

GB:  $\alpha_0 e^{-e} + \alpha_1 R e e e + \alpha_2 R R e e$

$L_{CS}(\dots)$

$A = A$

Ads  $\rightarrow$



GB:  $\frac{1}{2}e - e + \frac{1}{2}Reee + \frac{1}{2}RRR$

$L_{CS}($

$A = A$

Ans  $\rightarrow$



- Field content: Gravitational field + Electromagnetic field + Symmetry breaking matter fields

---

**Ex. 1:** scalar field  $\Phi$  in AdS gravity is related to the order parameter  $\mathcal{O}$  in QFT

**Ex. 2:** Gauss-Bonnet holographic superconductor + Maxwell field + Minimally coupled scalar field [e.g., arXiv:1009.1991]  
GB coupling  $\alpha$  does not admit the Chern-Simons limit  $\alpha \rightarrow \alpha_{CS}$ .

---

**Our proposal:** *Studying holographic QFT dual to Chern-Simons gravity for an expansion of AdS algebra*

- Field content: Gravitational field + Electromagnetic field + Symmetry breaking matter fields

---

**Ex. 1:** scalar field  $\Phi$  in AdS gravity is related to the order parameter  $\mathcal{O}$  in QFT

**Ex. 2:** Gauss-Bonnet holographic superconductor + Maxwell field + Minimally coupled scalar field [e.g., arXiv:1009.1991]  
GB coupling  $\alpha$  does not admit the Chern-Simons limit  $\alpha \rightarrow \alpha_{CS}$ .

---

**Our proposal:** *Studying holographic QFT dual to Chern-Simons gravity for an expansion of AdS algebra*

As a first step in arXiv:1406.3096 - also in collaboration with G. Giribet (UBA, Argentina) and J. Zanelli (CECs, Chile) - we studied static spherically symmetric BH solutions of a CS theory for the most simple extension of AdS symmetry:  $\mathfrak{so}(4,2) \times U(1)$ .

*Objective for the first part:*

- To briefly describe main properties of the expansion methods.

---

<sup>7</sup>Charge is needed for the holographic application.



*Objective for the first part:*

- To briefly describe main properties of the expansion methods.

*Second part:*

- To show that charged<sup>7</sup> solutions in CS theory for  $so(4,2) \times U(1)$  requires to have non-vanishing torsion.

---

<sup>7</sup>Charge is needed for the holographic application.

*Objective for the first part:*

- To briefly describe main properties of the expansion methods.

*Second part:*

- To show that charged<sup>7</sup> solutions in CS theory for  $so(4,2) \times U(1)$  requires to have non-vanishing torsion.
- To see that degenerated sectors appear in the space of solutions and explain how this is related to the presence of additional local symmetries (called "accidental").

---

<sup>7</sup>Charge is needed for the holographic application.



*Objective for the first part:*

- To briefly describe main properties of the expansion methods.

*Second part:*

- To show that charged<sup>7</sup> solutions in CS theory for  $so(4,2) \times U(1)$  requires to have non-vanishing torsion.
- To see that degenerated sectors appear in the space of solutions and explain how this is related to the presence of additional local symmetries (called "accidental").
- To explain how that knowledge was useful to identify a **new physical solution** ; and then why it should be useful in the next step, when considering CS theory based on expanded algebras.

---

<sup>7</sup>Charge is needed for the holographic application.

# Contents

- 1 Introduction
- 2 Expansion methods**
- 3 Accidental symmetries
  - Action, field equations and some solutions
  - Torsion and degeneracy
  - Accidental symmetries
- 4 Conclusions

## Expansion methods, brief introduction

Expansion methods are examples of procedures giving **non-trivial relations** between Lie algebras and groups.

## Expansion methods, brief introduction

Expansion methods are examples of procedures giving **non-trivial relations** between Lie algebras and groups.

They allow to understand **interrelations** between physical theories.

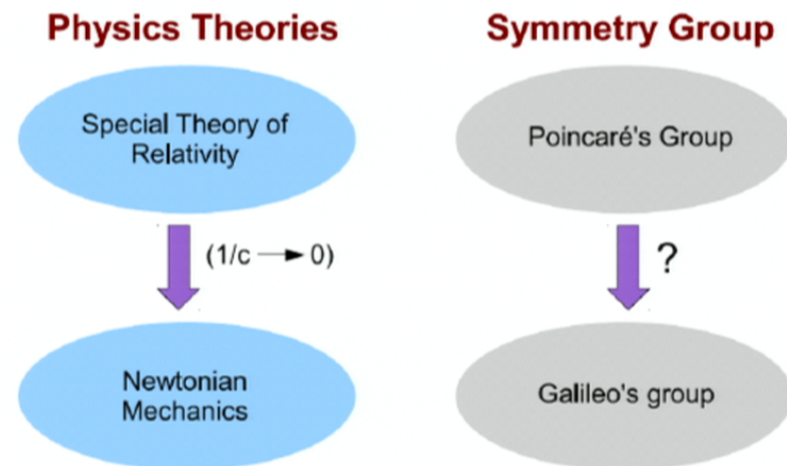
The first example, proposed by I.E. Segal in 1951,

## Expansion methods, brief introduction

Expansion methods are examples of procedures giving **non-trivial relations** between Lie algebras and groups.

They allow to understand **interrelations** between physical theories.

The first example, proposed by I.E. Segal in 1951,



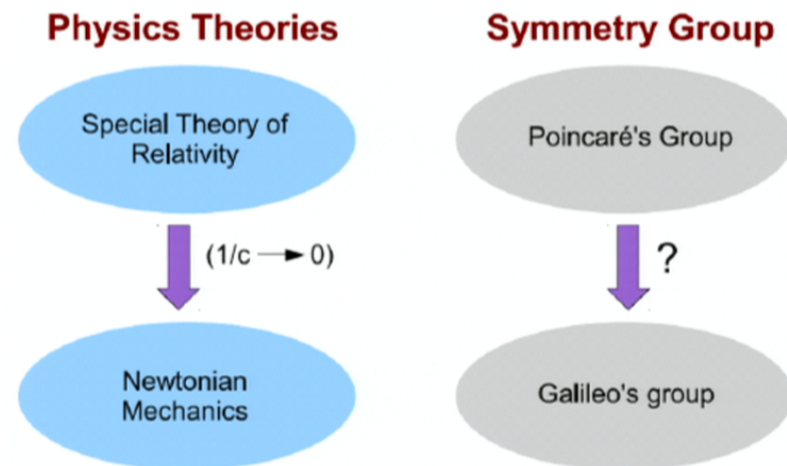


## Expansion methods, brief introduction

Expansion methods are examples of procedures giving **non-trivial relations** between Lie algebras and groups.

They allow to understand **interrelations** between physical theories.

The first example, proposed by I.E. Segal in 1951,



### *Inönü-Wigner contraction of $\mathcal{G}$ :*

It is made with respect to a subalgebra  $\mathcal{L}$ , first rescaling elements of the coset  $\mathcal{G}/\mathcal{L}$  by some parameter and then performing a non trivial limit.

*Inönü-Wigner contraction of  $\mathcal{G}$  :*

It is made with respect to a subalgebra  $\mathcal{L}$ , first rescaling elements of the coset  $\mathcal{G}/\mathcal{L}$  by some parameter and then performing a non trivial limit.

*Example:  $\mathfrak{iso}(3,1)$  from  $\mathfrak{so}(3,2)$*

### *Inönü-Wigner contraction of $\mathcal{G}$ :*

It is made with respect to a subalgebra  $\mathcal{L}$ , first rescaling elements of the coset  $\mathcal{G}/\mathcal{L}$  by some parameter and then performing a non trivial limit.

*Example:  $\mathfrak{iso}(3, 1)$  from  $\mathfrak{so}(3, 2)$*

$$[M_{ab}, M_{cd}] = \eta_{ac}M_{bd} - \eta_{bc}M_{ad} - \eta_{ad}M_{bc} + \eta_{bd}M_{ac}, \quad a, b, \dots = 1, \dots, 5$$

Rescaling  $M_{5\mu} = RP_{\mu}$ , with  $\mu, \nu, \dots = 1, \dots, 4$  leads to

### *Inönü-Wigner contraction of $\mathcal{G}$ :*

It is made with respect to a subalgebra  $\mathcal{L}$ , first rescaling elements of the coset  $\mathcal{G}/\mathcal{L}$  by some parameter and then performing a non trivial limit.

*Example:  $\mathfrak{iso}(3, 1)$  from  $\mathfrak{so}(3, 2)$*

$$[M_{ab}, M_{cd}] = \eta_{ac}M_{bd} - \eta_{bc}M_{ad} - \eta_{ad}M_{bc} + \eta_{bd}M_{ac}, \quad a, b, \dots = 1, \dots, 5$$

Rescaling  $M_{5\mu} = RP_{\mu}$ , with  $\mu, \nu, \dots = 1, \dots, 4$  leads to

$$[P_{\mu}, P_{\nu}] = \frac{1}{R^2} \eta_{55} M_{\mu\nu}$$



Different mechanisms known as contractions, deformations and extensions appeared in literature.

*Brief history:*

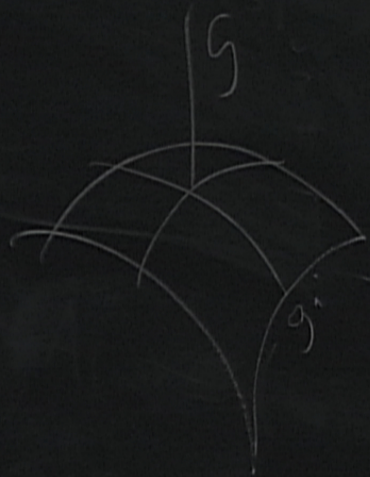
Different mechanisms known as contractions, deformations and extensions appeared in literature.

*Brief history:*

- 1953: Inönü-Wigner contractions
- 2000 - 2003: Generalized contractions, Weimar Woods
- 2003: Expansion method, Hatsuda, Sakaguchi  
[arXiv:hep-th/0106114] & de Azcarraga, Izquierdo, Picon, Varela  
[arXiv: hep-th/0212347]

$$GB := \alpha_e \cdot e + \alpha_{Re} e + \alpha_{RR} e$$

$$L_{CS}(A, G) \rightarrow A \text{ EN}$$



$$A = A^e T_a$$

$$A_{DS} \rightarrow G$$

Different mechanisms known as contractions, deformations and extensions appeared in literature.

*Brief history:*

- 1953: Inönü-Wigner contractions
- 2000 - 2003: Generalized contractions, Weimar Woods
- 2003: Expansion method, Hatsuda, Sakaguchi  
[arXiv:hep-th/0106114] & de Azcarraga, Izquierdo, Picon, Varela  
[arXiv: hep-th/0212347]
- 2006: S-expansion method, Izaurieta, Rodriguez, Salgado  
[arXiv:heo-th/0606215]

*Advantages and some applications of the S-expansion method:*

- Direct **supersymmetric** extension
- It provides non trivial **invariant tensors** for the expanded algebra (different from the (super)symmetrized trace)
- That was useful to construct<sup>8</sup> a CS action for the  $M$ -algebra (which was obtained as an expansion of  $\mathfrak{osp}(32|1)$ )

---

<sup>8</sup>F. Izaurieta, *et al*, *arXiv: 0606225*; see also de Azcárraga *et al*, *arXiv:0212347*

<sup>9</sup>F. Izaurieta *et al*, *arXiv:0903.4712*

<sup>10</sup>Edelstein *et al*, *arXiv:0605174*; Izaurieta *et al*, *arXiv:0905.2187*; P.K. Concha, *et al*, *arXiv:1309.0062*, *1402.0023*, *arXiv:1405.7078*.



*Advantages and some applications of the S-expansion method:*

- Direct **supersymmetric** extension
- It provides non trivial **invariant tensors** for the expanded algebra (different from the (super)symmetrized trace)
- That was useful to construct<sup>8</sup> a CS action for the  $M$ -algebra (which was obtained as an expansion of  $\mathfrak{osp}(32|1)$ )
- A **dual formulation**<sup>9</sup> of S-expansion method permits to perform the expansion at the level of the Lagrangian.

---

<sup>8</sup>F. Izaurieta, *et al*, *arXiv: 0606225*; see also de Azcárraga *et al*, *arXiv:0212347*

<sup>9</sup>F. Izaurieta *et al*, *arXiv:0903.4712*

<sup>10</sup>Edelstein *et al*, *arXiv:0605174*; Izaurieta *et al*, *arXiv:0905.2187*; P.K. Concha, *et al*, *arXiv:1309.0062*, *1402.0023*, *arXiv:1405.7078*.

*Advantages and some applications of the S-expansion method:*

- Direct **supersymmetric** extension
- It provides non trivial **invariant tensors** for the expanded algebra (different from the (super)symmetrized trace)
- That was useful to construct<sup>8</sup> a CS action for the  $M$ -algebra (which was obtained as an expansion of  $\mathfrak{osp}(32|1)$ )
- A **dual formulation**<sup>9</sup> of S-expansion method permits to perform the expansion at the level of the Lagrangian.
- By constructing lagrangians invariant under expansions of  $\mathfrak{so}(D-1,2)$ , it has been found standard General Relativity (SGR) in **even** and **odd** dimensions as a special limit of a **BI** and **CS** Lagrangian respectively<sup>10</sup>.

<sup>8</sup>F. Izaurieta, *et al*, *arXiv: 0606225*; see also de Azcárraga *et al*, *arXiv:0212347*

<sup>9</sup>F. Izaurieta *et al*, *arXiv:0903.4712*

<sup>10</sup>Edelstein *et al*, *arXiv:0605174*; Izaurieta *et al*, *arXiv:0905.2187*; P.K. Concha, *et al*, *arXiv:1309.0062*, *1402.0023*, *arXiv:1405.7078*.

All generalized contractions, mentioned before, can be reproduced in the frame of the S-expansion method by using one of the semigroups  $S_E^{(N)} = \{\lambda_0, \dots, \lambda_{N+1}\}$  defined by:

$$\begin{aligned}\lambda_\alpha \lambda_\beta &= \lambda_{\alpha+\beta}, \text{ if } \alpha + \beta < N + 1 \\ \lambda_\alpha \lambda_\beta &= \lambda_{N+1}, \text{ if } \alpha + \beta \geq N + 1\end{aligned}$$

Expansions with **other semigroups** can generate algebras that cannot be reached, nor by any contraction neither by an expansion of de Azcarraga et al.

All generalized contractions, mentioned before, can be reproduced in the frame of the S-expansion method by using one of the semigroups  $S_E^{(N)} = \{\lambda_0, \dots, \lambda_{N+1}\}$  defined by:

$$\begin{aligned}\lambda_\alpha \lambda_\beta &= \lambda_{\alpha+\beta}, \text{ if } \alpha + \beta < N + 1 \\ \lambda_\alpha \lambda_\beta &= \lambda_{N+1}, \text{ if } \alpha + \beta \geq N + 1\end{aligned}$$

Expansions with **other semigroups** can generate algebras that cannot be reached, nor by any contraction neither by an expansion of de Azcarraga et al.

As many physical applications have been appeared by using this methods, we wanted to answer the following question:

All generalized contractions, mentioned before, can be reproduced in the frame of the S-expansion method by using one of the semigroups  $S_E^{(N)} = \{\lambda_0, \dots, \lambda_{N+1}\}$  defined by:

$$\begin{aligned}\lambda_\alpha \lambda_\beta &= \lambda_{\alpha+\beta}, \text{ if } \alpha + \beta < N + 1 \\ \lambda_\alpha \lambda_\beta &= \lambda_{N+1}, \text{ if } \alpha + \beta \geq N + 1\end{aligned}$$

Expansions with **other semigroups** can generate algebras that cannot be reached, nor by any contraction neither by an expansion of de Azcarraga et al.

As many physical applications have been appeared by using this methods, we wanted to answer the following question:

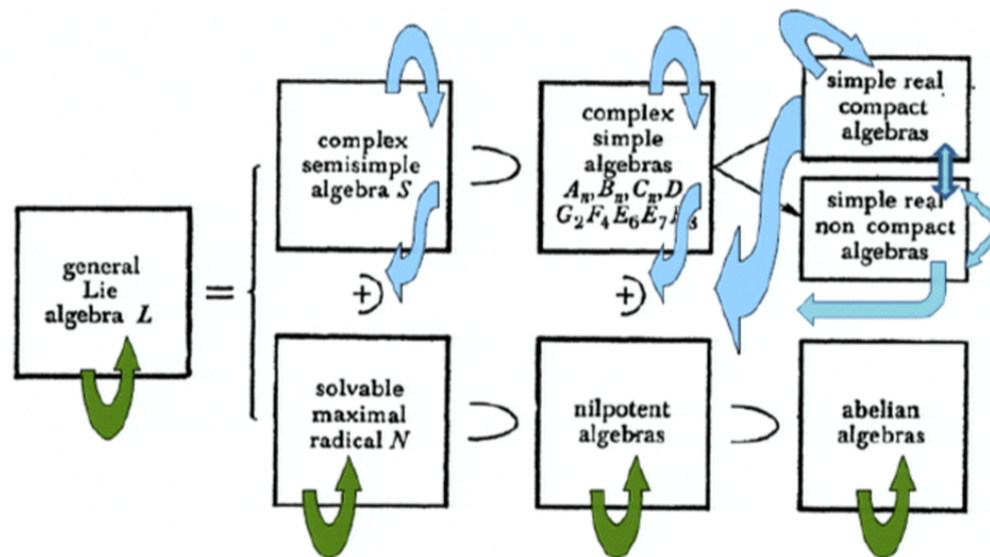
Given two Lie algebras, can they be related by an S-expansion?



To answer that, one should consider the complete family of abelian semigroups, i.e., to take into account the history of their classification:

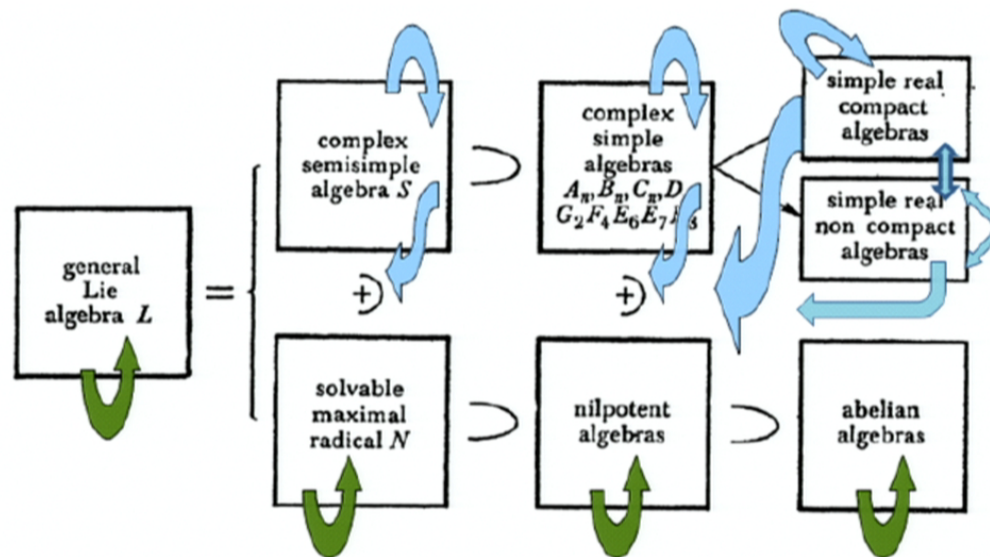
order	$Q = \#$ semigroups	
1	1	
2	4	
3	18	
4	126	[Forsythe '54]
5	1,160	[Motzkin, Selfridge '55]
6	15,973	[Plemmons '66]
7	836,021	[Jurgensen, Wick '76]
8	1,843,120,128	[Satoh, Yama, Tokizawa '94]
9	<b>52,989,400,714,478</b>	[Distler, Kelsey, Mitchell '09]

In particular, we have found<sup>11</sup> some criteria, related to the preservation of some properties of the Lie algebra, that allows to answer that question:



<sup>11</sup>L. Andrianopoli, N. Merino, F. Nadal, M. Trigiante, General properties of the S-expansion method, *arXiv:1308.4832*

In particular, we have found<sup>11</sup> some criteria, related to the preservation of some properties of the Lie algebra, that allows to answer that question:



<sup>11</sup>L. Andrianopoli, N. Merino, F. Nadal, M. Trigiante, General properties of the S-expansion method, *arXiv:1308.4832*

We have also implemented computer programs to perform expansion with any semigroup (up to order 6 ) that allow us:

- to study all **resonant decompositions** of a semigroup,
- to establish the isomorphism of an **arbitrary semigroup** with one of those classified in the literature and

As an example, we have applied the procedure<sup>12</sup> in the context of Bianchi algebras:

**Principal Idea:** can be related 2 and 3-dimensional isometries?

Considering the two set of algebras:  $[X_1, X_2] = 0$  and  
 $[X_1, X_2] = X_1$

Group	Algebra
→ type I	$[X_1, X_2] = [X_1, X_3] = [X_2, X_3] = 0$
→ type II	$[X_1, X_2] = [X_1, X_3] = 0, [X_2, X_3] = X_1$
→ type III	$[X_1, X_2] = [X_2, X_3] = 0, [X_1, X_3] = X_1$
type IV	$[X_1, X_2] = 0, [X_1, X_3] = X_1, [X_2, X_3] = X_1 + X_2$
→ type V	$[X_1, X_2] = 0, [X_1, X_3] = X_1, [X_2, X_3] = X_2$
type VI	$[X_1, X_2] = 0, [X_1, X_3] = X_1, [X_2, X_3] = hX_2,$ where $h \neq 0, 1$
type VII <sub>1</sub>	$[X_1, X_2] = 0, [X_1, X_3] = X_2, [X_2, X_3] = -X_1$
type VII <sub>2</sub>	$[X_1, X_2] = 0, [X_1, X_3] = X_2, [X_2, X_3] = -X_1 + hX_2,$ where $h \neq 0 (0 < h < 2)$ .
type VIII	$[X_1, X_2] = X_1, [X_1, X_3] = 2X_2, [X_2, X_3] = X_3$
type IX	$[X_1, X_2] = X_3, [X_2, X_3] = X_1, [X_3, X_1] = X_2$

<sup>12</sup>R. Caroca, I. Kondrashuk, N. Merino and F. Nadal, *arXiv:1104.3541*



As an example, we have applied the procedure<sup>12</sup> in the context of Bianchi algebras:

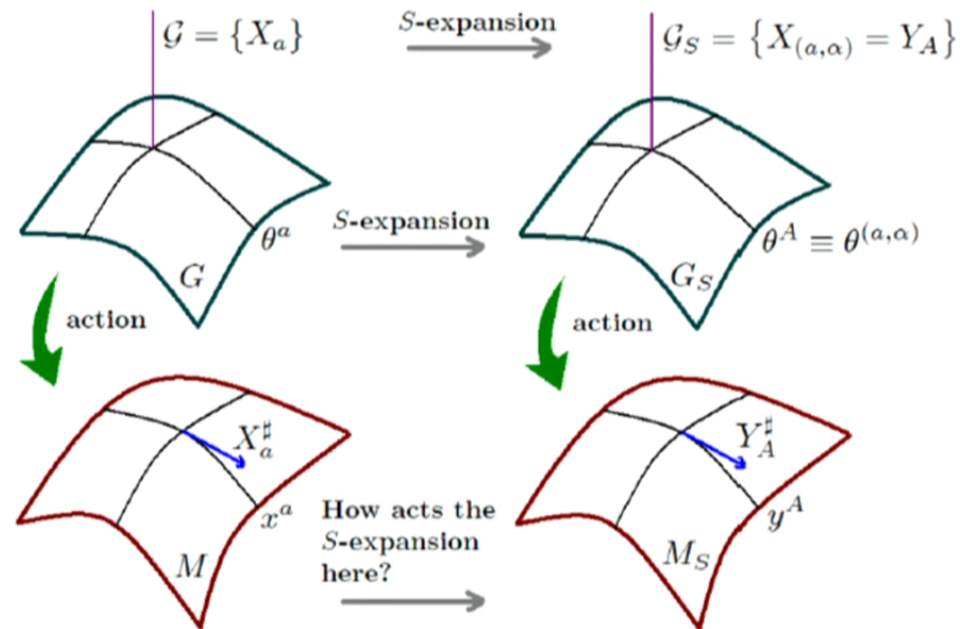
**Principal Idea:** can be related 2 and 3-dimensional isometries?

Considering the two set of algebras:  $[X_1, X_2] = 0$  and  
 $[X_1, X_2] = X_1$

Group	Algebra
→ type I	$[X_1, X_2] = [X_1, X_3] = [X_2, X_3] = 0$
→ type II	$[X_1, X_2] = [X_1, X_3] = 0, [X_2, X_3] = X_1$
→ type III	$[X_1, X_2] = [X_2, X_3] = 0, [X_1, X_3] = X_1$
type IV	$[X_1, X_2] = 0, [X_1, X_3] = X_1, [X_2, X_3] = X_1 + X_2$
→ type V	$[X_1, X_2] = 0, [X_1, X_3] = X_1, [X_2, X_3] = X_2$
type VI	$[X_1, X_2] = 0, [X_1, X_3] = X_1, [X_2, X_3] = hX_2,$ where $h \neq 0, 1$
type VII <sub>1</sub>	$[X_1, X_2] = 0, [X_1, X_3] = X_2, [X_2, X_3] = -X_1$
type VII <sub>2</sub>	$[X_1, X_2] = 0, [X_1, X_3] = X_2, [X_2, X_3] = -X_1 + hX_2,$ where $h \neq 0$ ( $0 < h < 2$ ).
type VIII	$[X_1, X_2] = X_1, [X_1, X_3] = 2X_2, [X_2, X_3] = X_3$
type IX	$[X_1, X_2] = X_3, [X_2, X_3] = X_1, [X_3, X_1] = X_2$

<sup>12</sup>R. Caroca, I. Kondrashuk, N. Merino and F. Nadal, *arXiv:1104.3541*

In particular, we currently working<sup>13</sup> in finding how the mechanism works in the case of Lie group transformations.



<sup>13</sup>Discussions with M. Calderón, I. Kondrashuk and M. Trigiante

As the metric could represent a solution of either GR or Cosmology, we expect this extension to be useful to **relate solutions** in different kind of theories.

In fact, a BH<sup>14</sup> and a cosmological<sup>15</sup> solution have already been studied for a CS theory based on expanded algebras.

However, an important issue related to **degeneracy** in some sectors of the space of solutions in higher curvature theories must be taken into account. **This is the subject of the next part of the talk.**

---

<sup>14</sup>Quinzacara *et al*, arXiv:1401.1797

<sup>15</sup>Crisóstomo *et al*, arXiv:1401.2128

# Action, field equations and some solutions

The Lagrangian is a CS density constructed from  $\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = dL_{\text{CS}}(\mathbf{A})$ ,

## Action, field equations and some solutions

The Lagrangian is a CS density constructed from  $\langle \mathbf{F} \wedge \dots \wedge \mathbf{F} \rangle_g = dL_{\text{CS}}(\mathbf{A})$ ,

$$\mathbf{A} = A^A \mathbf{T}_A = \frac{1}{2} \omega^{ab} \mathbf{J}_{ab} + \frac{1}{\ell} e^a \mathbf{P}_a + A \mathbf{T}_1, \quad T_A \in \mathfrak{so}(4,2) \times \mathfrak{u}(1),$$

$$\mathbf{F} = \frac{1}{2} \left( R^{ab} + \frac{1}{\ell^2} e^a e^b \right) \mathbf{J}_{ab} + \frac{1}{\ell} T^a \mathbf{P}_a + F \mathbf{T}_1, \quad F = dA,$$

so we have,



## Action, field equations and some solutions

The Lagrangian is a CS density constructed from  $\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = dL_{\text{CS}}(\mathbf{A})$ ,

$$\mathbf{A} = A^A \mathbf{T}_A = \frac{1}{2} \omega^{ab} \mathbf{J}_{ab} + \frac{1}{\ell} e^a \mathbf{P}_a + A \mathbf{T}_1, \quad T_A \in \mathfrak{so}(4, 2) \times \mathfrak{u}(1),$$

$$\mathbf{F} = \frac{1}{2} \left( R^{ab} + \frac{1}{\ell^2} e^a e^b \right) \mathbf{J}_{ab} + \frac{1}{\ell} T^a \mathbf{P}_a + F \mathbf{T}_1, \quad F = dA,$$

so we have,

$$\begin{aligned} I_{\text{CS}}[\mathbf{A}] &= \int_M \mathcal{L}_{\text{CS}}(\mathbf{A}) = \frac{i}{3} \int_M \left\langle \mathbf{A} \mathbf{F}^2 - \frac{1}{2} \mathbf{F} \mathbf{A}^3 + \frac{1}{10} \mathbf{A}^5 \right\rangle \\ &= \int_M \left[ \mathcal{L}_{\text{AdS}}(e, \omega) + \mathcal{L}_{\text{U}(1)}(A) + \mathcal{L}_{\text{int}}(e, \omega, A) \right], \end{aligned}$$

## Action, field equations and some solutions

The Lagrangian is a CS density constructed from  $\langle \mathbf{F} \wedge \cdots \wedge \mathbf{F} \rangle_g = dL_{\text{CS}}(\mathbf{A})$ ,

$$\mathbf{A} = A^A \mathbf{T}_A = \frac{1}{2} \omega^{ab} \mathbf{J}_{ab} + \frac{1}{\ell} e^a \mathbf{P}_a + A \mathbf{T}_1, \quad T_A \in \mathfrak{so}(4, 2) \times \mathfrak{u}(1),$$

$$\mathbf{F} = \frac{1}{2} \left( R^{ab} + \frac{1}{\ell^2} e^a e^b \right) \mathbf{J}_{ab} + \frac{1}{\ell} T^a \mathbf{P}_a + F \mathbf{T}_1, \quad F = dA,$$

so we have,

*uvucue -*

$$\begin{aligned} I_{\text{CS}}[\mathbf{A}] &= \int_M \mathcal{L}_{\text{CS}}(\mathbf{A}) = \frac{\alpha}{3} \int_M \left\langle \mathbf{A} \mathbf{F}^2 - \frac{1}{2} \mathbf{F} \mathbf{A}^3 + \frac{1}{10} \mathbf{A}^5 \right\rangle \\ &= \int_M \left[ \mathcal{L}_{\text{AdS}}(e, \omega) + \mathcal{L}_{U(1)}(A) + \mathcal{L}_{\text{int}}(e, \omega, A) \right], \end{aligned}$$

$$\mathcal{L}_{\text{AdS}}(e, \omega) = \frac{k}{4\ell} \epsilon_{abcde} \left( R^{ab} R^{cd} + \frac{2}{3\ell^2} R^{ab} e^c e^d + \frac{1}{5\ell^4} e^a e^b e^c e^d \right) e^e,$$

$$\mathcal{L}_{U(1)}(A) = \beta A F^2,$$

$$\mathcal{L}_{\text{int}} = \frac{\alpha}{2} \left[ R^{ab} R_{ab} + \frac{2}{\ell^2} \left( R^{ab} e_a e_b - T^a T_a \right) \right] A.$$

Field equations are

$$\mathbf{A} = A^{\mu} \mathbf{I}_A = \frac{1}{2} \omega^{ab} \mathbf{J}_{ab} + \frac{1}{\ell} e^a \mathbf{T}_a + A \mathbf{I}_1, \quad \mathbf{I}_A \in \mathfrak{so}(4, 2) \times \mathfrak{u}(1),$$

$$\mathbf{F} = \frac{1}{2} \left( R^{ab} + \frac{1}{\ell^2} e^a e^b \right) \mathbf{J}_{ab} + \frac{1}{\ell} T^a \mathbf{P}_a + F \mathbf{T}_1, \quad F = dA,$$

ive,

$$\begin{aligned} I_{\text{CS}}[\mathbf{A}] &= \int_M \mathcal{L}_{\text{CS}}(\mathbf{A}) = \frac{i}{3} \int_M \left\langle \mathbf{A} \mathbf{F}^2 - \frac{1}{2} \mathbf{F} \mathbf{A}^3 + \frac{1}{10} \mathbf{A}^5 \right\rangle \\ &= \int_M \left[ \mathcal{L}_{\text{AdS}}(e, \omega) + \mathcal{L}_{\text{U}(1)}(A) + \mathcal{L}_{\text{int}}(e, \omega, A) \right], \end{aligned}$$

In the local coordinates  $x^\mu = (t, r, x^m)$  (with  $x^m = x, y, z$ ), the ansatz is:

$$ds^2 = -f^2(r)dt^2 + \frac{dr^2}{f^2(r)} + r^2 \delta_{mn} dx^m dx^n ,$$

In the local coordinates  $x^\mu = (t, r, x^m)$  (with  $x^m = x, y, z$ ), the ansatz is:

$$ds^2 = -f^2(r)dt^2 + \frac{dr^2}{f^2(r)} + r^2 \delta_{mn} dx^m dx^n,$$

and has seven Killing vectors:

$$\tilde{\zeta} = \zeta^\mu \partial_\mu = c \partial_t + a_m \epsilon^{mk}{}_n x^n \partial_k + b_m \partial_m.$$



In the local coordinates  $x^\mu = (t, r, x^m)$  (with  $x^m = x, y, z$ ), the ansatz is:

$$ds^2 = -f^2(r)dt^2 + \frac{dr^2}{f^2(r)} + r^2 \delta_{mn} dx^m dx^n,$$

and has seven Killing vectors:

$$\tilde{\zeta} = \tilde{\zeta}^\mu \partial_\mu = c \partial_t + a_m \epsilon^{mk} x^n \partial_k + b_m \partial_m.$$

In order to use [Riemann-Cartan formalism](#), we split group indices as  $a = (0, 1, i)$  (with  $i = 2, 3, 4$ ), so

$$e^0 = f(r) dt, \quad e^1 = \frac{dr}{f(r)}, \quad e^i = r \delta_m^i dx^m := r dx^i.$$

Demanding the gauge field  $F = dA$  and torsion to have the same isometries ( $\mathfrak{L}_\zeta F = 0$  and  $\mathfrak{L}_\zeta T^\lambda_{\mu\nu} = 0$ ) leads to,

Demanding the gauge field  $F = dA$  and torsion to have the same isometries ( $\mathfrak{L}_\xi F = 0$  and  $\mathfrak{L}_\xi T^\lambda_{\mu\nu} = 0$ ) leads to,

$$A = A_t(r) dt + A_r(r) dr ,$$

$$T^0 = -\frac{\chi_t}{f} dt \wedge dr , \quad T^1 = f \chi_r dt \wedge dr ,$$

$$T^i = \frac{1}{r} (\psi_t dt + \psi_r dr) \wedge dx^i + \frac{\phi}{2r} \delta^{ik} \epsilon_{knm} dx^n \wedge dx^m .$$

Demanding the gauge field  $F = dA$  and torsion to have the same isometries ( $\mathfrak{L}_\xi F = 0$  and  $\mathfrak{L}_\xi T^\lambda_{\mu\nu} = 0$ ) leads to,

$$A = A_t(r) dt + A_r(r) dr ,$$

$$T^0 = -\frac{\chi_t}{f} dt \wedge dr , \quad T^1 = f \chi_r dt \wedge dr ,$$

$$T^i = \frac{1}{r} (\psi_t dt + \psi_r dr) \wedge dx^i + \frac{\phi}{2r} \delta^{ik} \epsilon_{knm} dx^n \wedge dx^m .$$

**: functions to solve:**

Demanding the gauge field  $F = dA$  and torsion to have the same isometries ( $\mathfrak{L}_\xi F = 0$  and  $\mathfrak{L}_\xi T^\lambda_{\mu\nu} = 0$ ) leads to,

$$A = A_t(r) dt + A_r(r) dr ,$$

$$T^0 = -\frac{\chi_t}{f} dt \wedge dr , \quad T^1 = f \chi_r dt \wedge dr ,$$

$$T^i = \frac{1}{r} (\psi_t dt + \psi_r dr) \wedge dx^i + \frac{\phi}{2r} \delta^{ik} \epsilon_{knm} dx^n \wedge dx^m .$$

**Eight functions to solve:**

$$f(r), A_t(r), A_r(r), \phi(r), \chi_r(r), \chi_t(r), \psi_r(r), \psi_t(r) .$$



Field equations are

$$\begin{aligned} \delta e^a : \quad 0 &= \frac{k}{4} \epsilon_{abcde} F^{bc} F^{de} - \frac{2\alpha}{\ell} T_a F, \\ \delta \omega^{ab} : \quad 0 &= \frac{k}{\ell} \epsilon_{abcde} F^{cd} T^e + 2\alpha F_{ab} F, \\ \delta A : \quad 0 &= FF + \frac{\alpha}{2} R^{ab} R_{ab} - \frac{\alpha}{\ell^2} d(T^a e_a). \end{aligned}$$

with  $F^{ab} = R^{ab} + \frac{1}{\ell^2} e^a e^b$ .

We look for exact spherically symmetric **charged BH solutions** to these field equations.

### Field equations reduces to four independent ones:

Demanding the gauge field  $F = dA$  and torsion to have the same isometries ( $\mathfrak{L}_\xi F = 0$  and  $\mathfrak{L}_\xi T^\lambda_{\mu\nu} = 0$ ) leads to,

$$A = A_t(r) dt + A_r(r) dr ,$$

$$T^0 = -\frac{\chi_t}{f} dt \wedge dr , \quad T^1 = f \chi_r dt \wedge dr ,$$

$$T^i = \frac{1}{r} (\psi_t dt + \psi_r dr) \wedge dx^i + \frac{\phi}{2r} \delta^{ik} \epsilon_{knm} dx^n \wedge dx^m .$$

### Eight functions to solve:

$$f(r), A_t(r), A_r(r), \phi(r), \chi_t(r), \chi_r(r), \psi_t(r), \psi_r(r) .$$

Field equations reduces to four independent ones:

$$0 = \left( -\frac{\psi_t^2}{f^2} + f^2 (\psi_r - r)^2 + \frac{\phi^2}{4r^2} - \frac{r^4}{\ell^2} \right) \psi_r + \frac{\phi}{r} \left( \frac{r}{2} \phi' - \phi \right) ,$$

$$0 = \left( -\frac{\psi_t^2}{f^2} + f^2 (\psi_r - r)^2 + \frac{\phi^2}{4r^2} - \frac{r^4}{\ell^2} \right) \psi_t ,$$

$$0 = \frac{Cl\alpha}{k} r^2 f E_{tr} - r f \chi_r \psi_t + f^2 \eta - r f^2 \eta' \\ - f \frac{r^3}{\ell^2} + r^2 f \chi_t + r \chi_t \eta - r^2 f^2 f' - r f f' \eta ,$$

$$0 = \eta \chi_t - f \chi_r \psi_t - r^2 f \chi_t'^2 f f'^2 + r^2 f^2 f'' - \frac{r^2}{\ell^2} f .$$

A solution with axial torsion ( $\phi(r) \neq 0$ ) was considered by Canfora et al [arXiv:0707.1056], however, is uncharged.

### Case with non-vanishing $\psi_r$ and $\phi$

$$f^2(r) = \frac{r^2}{\ell^2} + br - \mu,$$

$$\phi(r) = 2Cr^2,$$

$$\psi_r = r \frac{\sqrt{r^2 + \ell^2 br - \ell^2 \mu} - \varepsilon_\psi \sqrt{r^2 - \ell^2 C^2}}{\sqrt{r^2 + \ell^2 br - \ell^2 \mu}},$$

$$A_t = \Phi - \frac{k}{Cl\alpha} \left[ \frac{r^2}{\ell^2} + \frac{br}{2} - \sqrt{\left( \frac{r^2}{\ell^2} + br - \mu \right) \left( \frac{r^2}{\ell^2} - C^2 \right)} \right],$$

$$A_r = 0.$$

This  $f(r)$  represents the five-dimensional analogue of the hairy BH solution considered in conformal gravity and massive gravity in three dimensions by Oliva et al [arXiv:0905.1510, 0905.1545].

The parameter  $b$  can be regarded as a gravitational hair. For some range of the parameters  $\mu$  and  $b$ , the solution represents a topological BH (or black brane).

The parameter  $b$  can be regarded as a gravitational hair. For some range of the parameters  $\mu$  and  $b$ , the solution represents a topological BH (or black brane).

Study of the corresponding horizons, asymptotic behavior, calculation of the mass (on the curve  $\ell b = \pm 2\sqrt{C^2 - \mu}$ , where electric field vanishes), Hawking temperature and entropy of this black branes was also made in arXiv:1406.3096.



# Torsion and degeneracy

By adding non-vanishing  $\psi_t$  one gets

$$\begin{aligned}
 f^2 &= \frac{r^2}{\ell^2} + br - \mu + \theta, & \phi &= 2Cr^2, \\
 A_t &= \Phi - \frac{k}{C\ell\alpha} \left( rff' + \frac{f\eta}{r} \right), & \psi_r &= r + \frac{\eta}{f}, \\
 \chi_r &= \frac{r^2\theta''}{2\psi_t}, & \psi_t &= \varepsilon\psi f \sqrt{\eta^2 + C^2r^2 - \frac{r^4}{\ell^2}},
 \end{aligned}$$

## Torsion and degeneracy

By adding non-vanishing  $\psi_t$  one gets

$$\begin{aligned}
 f^2 &= \frac{r^2}{\ell^2} + br - \mu + \theta, & \phi &= 2Cr^2, \\
 A_t &= \Phi - \frac{k}{C\ell\alpha} \left( rff' + \frac{f\eta}{r} \right), & \psi_r &= r + \frac{\eta}{f}, \\
 \chi_r &= \frac{r^2\theta''}{2\psi_t}, & \psi_t &= \varepsilon\psi f \sqrt{\eta^2 + C^2r^2 - \frac{r^4}{\ell^2}},
 \end{aligned}$$

where  $\theta(r)$  is an **arbitrary function**.

## General solution

$$f^2 = \frac{r^2}{\ell^2} + br - \mu + \theta,$$

$$A_t = \Phi - \frac{k}{Cl\alpha} \left[ \frac{r^2}{\ell^2} + \frac{br}{2} + \frac{r\theta'_t}{\theta' - \theta'_r + \theta_t} \left( \frac{r^2}{\ell^2} + br - \mu + \theta \right) + \frac{r(\theta'_r - \theta_t)}{2} \right],$$

$$A_r = 0,$$

$$\phi = 2Cr^2,$$

$$\psi_t = \varepsilon_\psi \varepsilon_f r \sqrt{\frac{r^2}{\ell^2} + br - \mu + \theta} \sqrt{\left( \frac{r^2}{\ell^2} + br - \mu + \theta \right) \left( \frac{r\theta'_t}{\theta' - \theta'_r + \theta_t} \right)^2 + C^2 - \frac{r^2}{\ell^2}},$$

$$\psi_r = r \left( 1 + \frac{r\theta'_t}{\theta' - \theta'_r + \theta_t} \right),$$

$$\chi_t = \frac{\theta' - \theta'_r + \theta_t}{2},$$

$$\chi_r = \frac{\varepsilon_\psi \varepsilon_f r \theta''_r}{2 \sqrt{\frac{r^2}{\ell^2} + br - \mu + \theta} \sqrt{\left( \frac{r^2}{\ell^2} + br - \mu + \theta \right) \left( \frac{r\theta'_t}{\theta' - \theta'_r + \theta_t} \right)^2 + C^2 - \frac{r^2}{\ell^2}}},$$

where  $\theta_t(r)$ ,  $\theta_r(r)$ , and  $\theta(r)$  are **arbitrary functions**. 

In CS (super)gravity theories, the appearance of arbitrary functions arise from degeneracies in the symplectic structure on certain sectors of phase space<sup>16</sup>.

In those sectors the system **acquires** extra gauge symmetry and **looses** dynamical degrees of freedom.

---

<sup>16</sup>Bañados, Garay, Henneaux / arXiv:hep-th/9506187, 9605159.

<sup>17</sup>Saavedra-Troncoso-Zanelli (2001).

In CS (super)gravity theories, the appearance of arbitrary functions arise from degeneracies in the symplectic structure on certain sectors of phase space<sup>16</sup>.

In those sectors the system **acquires** extra gauge symmetry and **looses** dynamical degrees of freedom.

However it is not an exclusive feature of CS theory since, as mention at the begining, arbitrary functions (called “**geometrically free solutions**” by J. T. Wheeler) are known to exist in some sectors of general Lovelock gravity.

---

<sup>16</sup>Bañados, Garay, Henneaux / arXiv:hep-th/9506187, 9605159.

<sup>17</sup>Saavedra-Troncoso-Zanelli (2001).

In CS (super)gravity theories, the appearance of arbitrary functions arise from degeneracies in the symplectic structure on certain sectors of phase space<sup>16</sup>.

In those sectors the system **acquires** extra gauge symmetry and **looses** dynamical degrees of freedom.

However it is not an exclusive feature of CS theory since, as mention at the begining, arbitrary functions (called “**geometrically free solutions**” by J. T. Wheeler) are known to exist in some sectors of general Lovelock gravity.

Furthermore, this behavior also exist in many mechanical systems<sup>17</sup>.

---

<sup>16</sup>Bañados, Garay, Henneaux / arXiv:hep-th/9506187, 9605159.

<sup>17</sup>Saavedra-Troncoso-Zanelli (2001).



General Lovelock theory has a pathological structure of its phase space because of the **non-invertible relation** between the metric and its conjugate momentum<sup>18</sup>.

This introduces an indeterminacy in the dynamical evolution and leads to degenerate dynamics.

Metrics with undetermined components were reported in higher-dimensional theories in the **torsionless** case as well, e.g., in Einstein-Gauss-Bonnet (EGB) AdS gravity when the transverse section of the metric is maximally symmetric<sup>19</sup>.

---

<sup>18</sup>Teitelboim-Zanelli (1987).

<sup>19</sup>Zegers, gr-qc/0505016.

<sup>20</sup>Oliva, arXiv:1210.4123.

General Lovelock theory has a pathological structure of its phase space because of the **non-invertible relation** between the metric and its conjugate momentum<sup>18</sup>.

This introduces an indeterminacy in the dynamical evolution and leads to degenerate dynamics.

Metrics with undetermined components were reported in higher-dimensional theories in the **torsionless** case as well, e.g., in Einstein-Gauss-Bonnet (EGB) AdS gravity when the transverse section of the metric is maximally symmetric<sup>19</sup>.

In case  $f(t, r)$ , there are still branches with undetermined components in CS theories<sup>20</sup>.

---

<sup>18</sup>Teitelboim-Zanelli (1987).

<sup>19</sup>Zegers, gr-qc/0505016.

<sup>20</sup>Oliva, arXiv:1210.4123.

It has been argued that the arbitrariness in the metric that appear in five-dimensional CS AdS gravity **can be removed** by:

- changing the cosmological constant, so that CS gravity becomes effectively EGB gravity<sup>21</sup>.
- gauge-fixing<sup>22</sup>, however a solution obtained in this way is still degenerate, i.e., the gauge-fixing hides the original arbitrariness in the metric.

---

<sup>21</sup>Bañados, hep-th/0310160.

<sup>22</sup>Aros-Contreras, gr-qc/0601135.

# Accidental symmetries

The presence of three arbitrary functions in the general solution is consequence of a local symmetry.

This symmetry cannot be a restriction of the gauge transformation  $A' = g^{-1}(A + d)g$  that preserves the form of the spherically symmetric ansatz  $A$ .

# Accidental symmetries

The presence of three arbitrary functions in the general solution is consequence of a local symmetry.

This symmetry cannot be a restriction of the gauge transformation  $A' = g^{-1}(A + d)g$  that preserves the form of the spherically symmetric ansatz  $A$ .

We have shown that the infinitesimal gauge transformations that preserve this ansatz are necessarily rigid ( $g = \text{Const}$ ).

# Accidental symmetries

The presence of three arbitrary functions in the general solution is consequence of a local symmetry.

This symmetry cannot be a restriction of the gauge transformation  $A' = g^{-1}(A + d)g$  that preserves the form of the spherically symmetric ansatz  $A$ .

We have shown that the infinitesimal gauge transformations that preserve this ansatz are necessarily rigid ( $g = \text{Const}$ ).

Thus, residual gauge symmetries of this kind cannot explain this.



# Accidental symmetries

→ Our background is not generic and it possesses additional local symmetries (different from  $\Lambda$  and  $\zeta$ ) called "accidental" because they happen to exist **only in certain sectors**.

# Accidental symmetries

→ Our background is not generic and it possesses additional local symmetries (different from  $\Lambda$  and  $\zeta$ ) called "accidental" because they happen to exist **only in certain sectors**.

In fact, we proved that field equations (in the branch where all spherically symmetric torsion components are switched on) **are insensitive** to the infinitesimal changes

$$\delta\theta = 2\sigma(r),$$

$$\delta\theta_t = 2 \int dr \tau(r),$$

$$\delta\theta_r = -2 \int dr \rho(r) + 2 \int dr \int_0^r ds \tau(s) + 2\sigma(r),$$

# Accidental symmetries

which induces the following local transformations on the fields

$$\begin{aligned}\delta f &= \frac{\sigma}{f}, \\ \delta A_t &= -\frac{k}{Cl\alpha} \left[ r\sigma' + \frac{2\eta}{rf} \sigma + \frac{rf^2}{\chi_t} \tau - \left( r + \frac{f\eta}{r\chi_t} \right) \rho \right], \\ \delta \psi_r &= \frac{r^2}{\chi_t} \tau - \frac{\eta}{f\chi_t} \rho, \\ \delta \psi_t &= \left( \frac{\psi_t}{f^2} + \frac{\eta^2}{\psi_t} \right) \sigma + \frac{f^2\eta}{\psi_t\chi_t} \left( r^2f\tau - \eta\rho \right), \\ \delta \chi_r &= \frac{r^2}{\psi_t} \sigma'' - \chi_r \left( \frac{1}{f^2} + \frac{\eta^2}{\psi_t^2} \right) \sigma + \frac{r^2}{\psi_t} \left( 1 - \frac{f^3\eta\chi_r}{\psi_t\chi_t} \right) \tau + \frac{f^2\eta^2\chi_r}{\psi_t^2\chi_t} \rho - \frac{r^2}{\psi_t} \rho', \\ \delta \chi_t &= \rho,\end{aligned}$$

with local parameters  $\sigma(r)$ ,  $\tau(r)$  and  $\rho(r)$ .

# Accidental symmetries

The transformations are Abelian because  $[\delta_1, \delta_2] = 0$  upon acting on any field.

# Accidental symmetries

The transformations are Abelian because  $[\delta_1, \delta_2] = 0$  upon acting on any field.


This new unexpected on-shell symmetry  $U(1) \times U(1) \times U(1)$  cannot be a Cartan subgroup of  $SO(2, 4) \times U(1)$  because we already showed that there are no residual gauge symmetries.

In fact, we found that on this sector there is only **one degree of freedom** and that Hamiltonian is (off-shell) invariant under 4-parameter local symmetry that on-shell reduces to the 3-parameter transformations  $(\delta f, \delta A_t, \delta \psi_r, \delta \psi_t, \delta \chi_r, \delta \chi_t)$ .

## Conclusion and future directions

Degeneracy in the space of solutions may appear in different families of L-L gravity in cases **with** and **without torsion**.

---

<sup>23</sup>Bañados-Garay-Henneaux, *arXiv:hep-th/9506187*, 9605159. 



## Conclusion and future directions

Degeneracy in the space of solutions may appear in different families of L-L gravity in cases **with** and **without torsion**.

We studied charged BH in CS  $\text{AdS} \times U(1)$  gravity and saw there exist degenerated branches in the static spherically symmetric sector.

**In contrast with a generic** CS AdS gravity with a  $U(1)$  field<sup>23</sup>, that possesses maximal number of degrees of freedom (14 in this case), we found that **there is only one** dynamically propagating mode in the static symmetric sector of phase space.

---

<sup>23</sup>Bañados-Garay-Henneaux, *arXiv:hep-th/9506187, 9605159*. 

## Conclusion and future directions

Indeed, **missing** degrees of freedom were related to an **increase** in local symmetries and thus our example provide **an explicit realization of a non-generic CS gravity**.

In particular this shows that the knowledge of these "accidental symmetries" can help to formulate a simple criterion that avoids unwanted degenerate ansatze.

## Conclusion and future directions

Indeed, **missing** degrees of freedom were related to an **increase** in local symmetries and thus our example provide **an explicit realization of a non-generic CS gravity**.

In particular this shows that the knowledge of these "accidental symmetries" can help to formulate a simple criterion that avoids unwanted degenerate ansatze.

This way, we identified two interesting solutions: the axial torsion one - already known in the literature -, and a **new 2-components torsion solution**.

## Conclusion and future directions

This issue of accidental symmetries should then be taken into account when studying charged solutions in CS theory based on expanded algebras, which is the next step to find an application in the Holographic context.



$G \times U(1)$

$\alpha_N R \dots R e^{i\alpha_0} \dots e$

$\alpha_N R \dots R e^{i\alpha_0} \dots e$

schritt  $\Theta(r)$

odd

even