Title: Exact holographic mapping, tensor networks and space-time geometry

Date: Feb 27, 2015 11:00 AM

URL: http://pirsa.org/15020134

Abstract: Holographic duality is a duality between gravitational systems and non-gravitational systems. In this talk, I will propose a different approach for understanding holographic duality named as the exact holographic mapping. The key idea of this approach can be summarized by two points: 1) The bulk theory and boundary theory are related by a unitary mapping in the Hilbert space. 2) Space-time geometry is determined by the structure of correlations and quantum entanglement in a quantum state. When applied to lattice systems, the holographic mapping is defined by a unitary tensor network. For free fermion boundary theories, I will discuss how different bulk geometries are obtained as dual theories of different boundary states. A particularly interesting case is the AdS black hole geometry and the interpretation of the interior of a black hole. We will also discuss dual geometries of topological states of matter.

Pirsa: 15020134 Page 1/38

#### Outline

- Background and motivation
- The definition of exact holographic mapping (EHM) and the free fermion example
- Thermodynamics and black hole geometry
- Optimal disentanglers and the interior of blackhole
- Holographic dual of topological states
- Discussion and summary

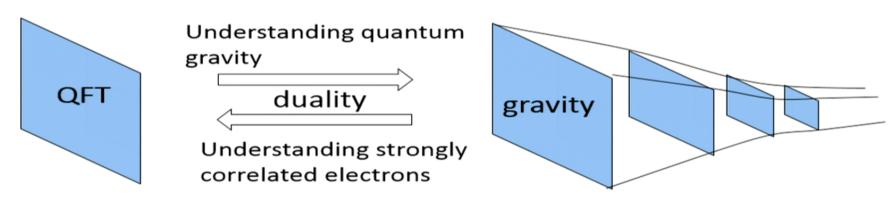
Ref: XLQ, arXiv:1309.6282 (2013)

**Unpublished works** 

- 1) Michael Zaletel, XLQ
- Yingfei Gu, Chinghua Lee, Gil Y. Cho, Xueda Wen, Shinsei Ryu, XLQ

Pirsa: 15020134 Page 2/38

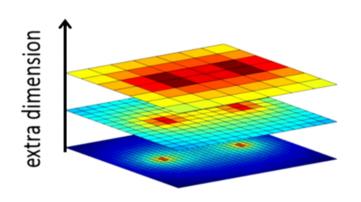
#### Holographic duality



- Holographic duality, or AdS/CFT (Maldacena '97, Witten '98, Gubser, Klebanov & Polyakov '98)

(E. Akhmedov, '98, Heemskerk & Polchinski '10)

 Application to condensed matter (For a review, see S. Sachdev, Annual Review of Condensed Matter Physics 3, 9 (2012))



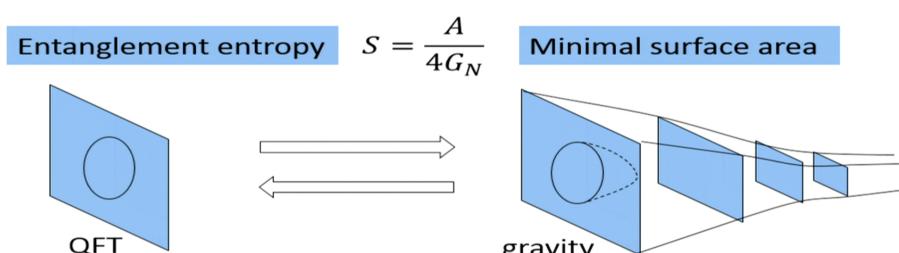
Pirsa: 15020134 Page 3/38

#### Holographic duality

- Holographic duality also relates space-time geometry with quantum entanglement.
- Bekenstein-Hawking formula

Black hole entropy 
$$S = \frac{A}{4G_N}$$
 Black hole area

Ryu-Takayanagi formula (S. Ryu & T. Takayanagi PRL '06)

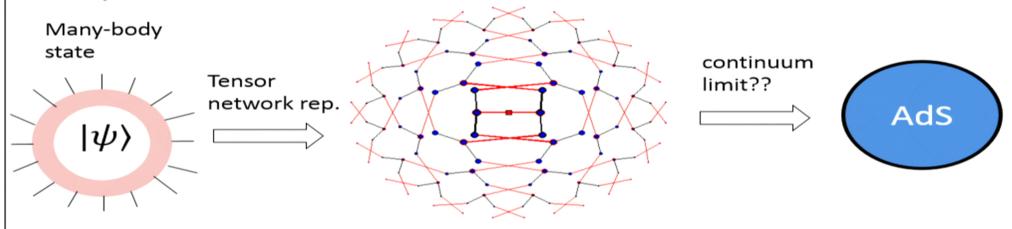


gravity

Pirsa: 15020134 Page 4/38

#### Tensor networks

- Multiscale Entanglement Renormalization Ansatz (MERA). A unitary tensor network describing critical states (Vidal '07)
- MERA has been proposed to be related to AdS/CFT (Swingle '10, Evenbly&Vidal '11, Haegeman et al '11, Nozaki et al '12, Hartman&Maldacena '13)
- How this relation exactly works remain an open question



Pirsa: 15020134 Page 5/38

# Step back and think about two questions about quantum space-time geometry

- Classical space-time: Riemann manifold consisting of points, and the distance between points given by geodesics.
- Quantum space-time geometry: How to define points in the geometry, and the distance between points?
- Points: A set of points corresponds to a direct product decomposition of the Hilbert space  $\mathbb{H} = \prod_{\mathbf{x}}^{\otimes} \mathbb{H}_{\mathbf{x}}$ .
- Distance: Distance between points shall be measured by physical correlation functions. <u>More</u> correlated sites are closer to each other.

Pirsa: 15020134 Page 6/38

#### How to define distance between points?

- A choice of points is only meaningful if physical correlations are local.  $E_{k}$
- Example: Massive free fermion

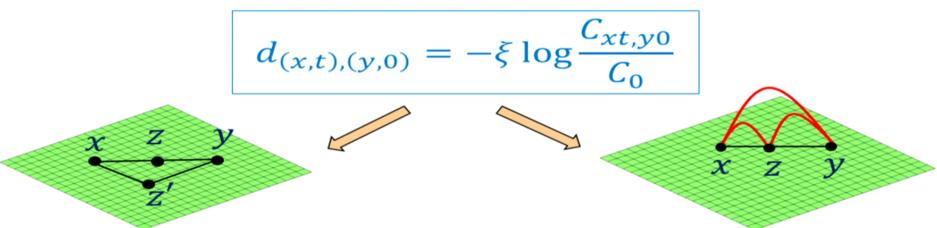


• Real space basis x. Imaginary time Correlation function  $C_{xt,y0} = \langle G | c_{xt}^+ c_{y0} | G \rangle \propto C_0 e^{-d_{(x,t),(y,0)}/\xi}$ . Natural distance definition

$$d_{(x,t),(y,0)} = -\xi \log \frac{C_{xt,y0}}{C_0}$$

- For all massive state, this definition works and recovers the classical space-time geometry.
- Mutual information  $I_{xy} = S_x + S_y S_{xy}$  can be used as an unbiased measure of equal-time correlation. (Wolf et al '08)

# From gapped states to critical states



# Gapped states Geodesic distance restored Triangle inequality holds

$$d_{xz'} + d_{z'y} \ge d_{xy}$$

#### **Critical states**

$$C_{xy} \propto |x - y|^{-2\Delta}$$

$$d_{xy} \propto \log |x - y|$$

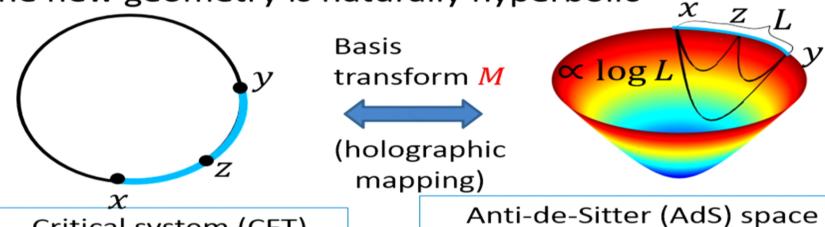
$$d_{xz'} + d_{z'y} > d_{xy}$$

 A new basis and new geometry is required for critical system

#### Relation to holographic duality

- Lesson from the discussion so far:
- A new basis may be defined for critical systems, which leads to a new geometry.

The new geometry is naturally hyperbolic



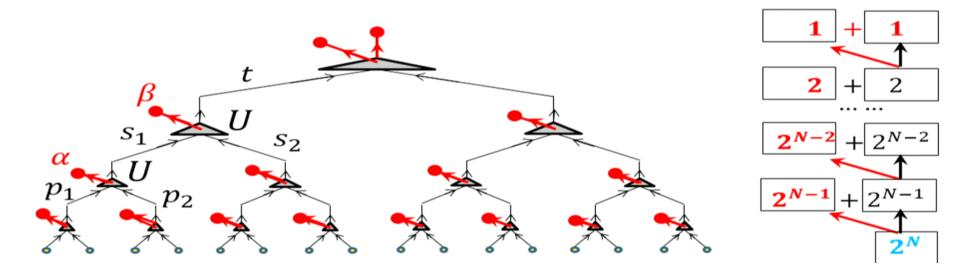
Critical system (CFT)

$$C_{xy} \propto |x - y|^{-2\Delta}$$
  
Scaling dimensions

Anti-de-Sitter (AdS) space  $C_{xy} \propto e^{-\frac{d_{xy}}{\xi}}$   $d_{xy} \propto \log|x-y|$  Mass of bulk fields

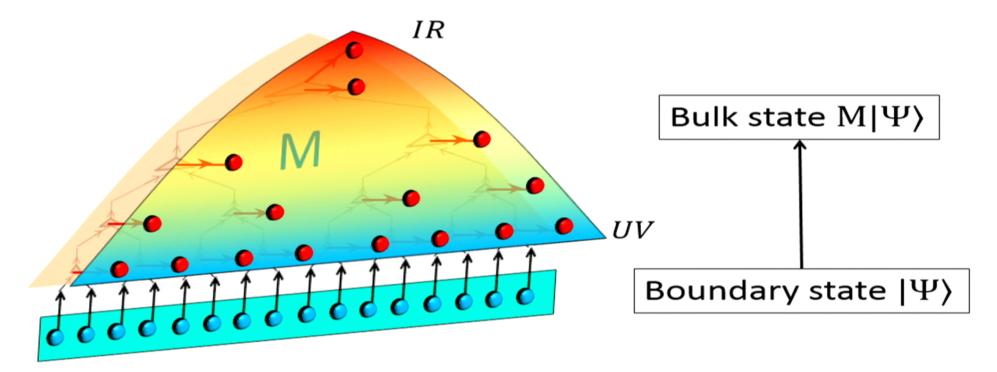
# Definition of the exact holographic mapping (EHM)

- 1. Starting from 2<sup>N</sup> sites, U maps each two neighboring sites to a "low energy" site and a "high energy" site
   2. Repeat step 1 the 2<sup>N-1</sup> low energy sites.
- Unitary mapping  $M=\prod_{\mathrm{network}}U$  maps boundary (2<sup>N</sup> sites) to bulk (  $1+1+2+\cdots+2^{N-1}=2^N$  sites)
- A modification of MERA



Pirsa: 15020134 Page 10/38

#### Definition of the exact holographic mapping

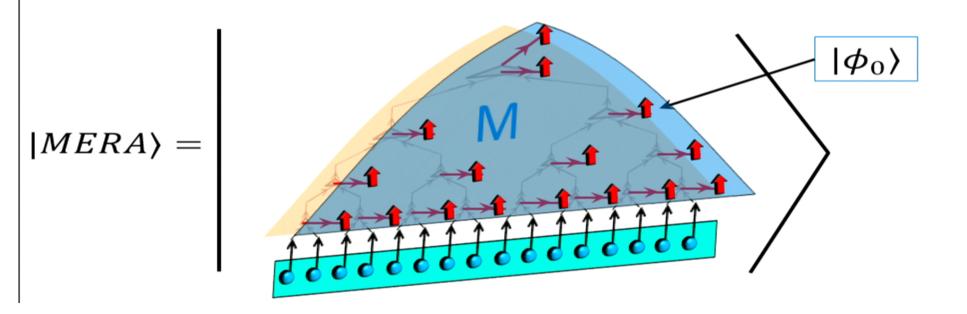


- A "lossless" generalization of real space RG.
- Degrees of freedom at different energy scales are all kept, and they can entangle with each other.

Pirsa: 15020134 Page 11/38

#### Relation of EHM and MERA

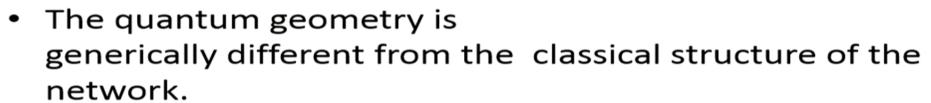
- A MERA state corresponds to acting the reverse EHM of a direct product bulk state  $|MERA\rangle=M^{-1}\prod^{\bigotimes}|\phi_0
  angle$
- The goal of EHM is to allow bulk states to entangle and use that to probe the geometry.

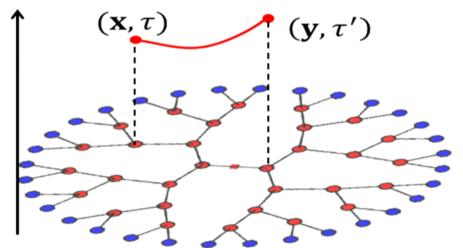


Pirsa: 15020134 Page 12/38

# Exact holographic mapping: bulk geometry

- Bulk space-time geometry determined by bulk correlation functions  $d_{(x,\tau),(y,\tau')} = -\xi \log \frac{\langle o_{(x,\tau)} o_{(y,\tau')} \rangle}{c_0}$ . ( $\tau$  is the imaginary time)
- Mutual information can be used for equal time distance.
   (A realization of ER=EPR (Maldacena&Susskind))
- Different choices of mapping
   U correspond to different
   choices of "spatial slices".





#### Exact holographic mapping: free fermion

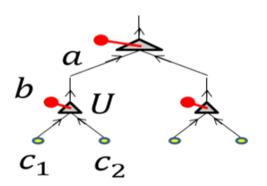
- 1+1d lattice Dirac fermion  $H = \sum_k c_k^+ (\sin k \, \sigma_x + (m+1-\cos k)\sigma_v) c_k$
- Unitary mapping

• 
$$\binom{c_1}{c_2} \rightarrow \binom{a}{b} = \frac{1}{\sqrt{2}} \binom{c_1 + c_2}{c_1 - c_2}$$

In many-body Hilbert space

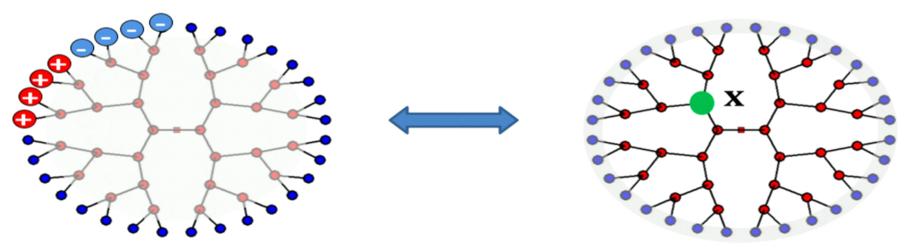
$$U = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1/\sqrt{2} & 1/\sqrt{2} & 0 \\ 0 & -1/\sqrt{2} & 1/\sqrt{2} & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix},$$

$$|00\rangle |10\rangle |01\rangle |11\rangle$$



# Exact holographic mapping: free fermion

Basis transformation between boundary and bulk

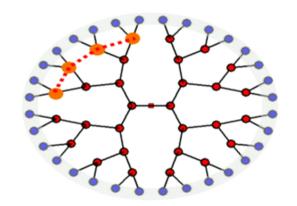


"Haar wavelet" wavefunctions at the boundary  $\phi_{\mathbf{x}}(i)$ 

Local basis in the bulk

$$b_{\mathbf{x}} = \sum_{i} \phi_{\mathbf{x}}^{*} (i) c_{i}$$

- T = m = 0
- Spatial direction: compare the distance obtained from  $d_{xy}=-\xi\lograc{I_{xy}}{I_0}$  to the AdS space distance

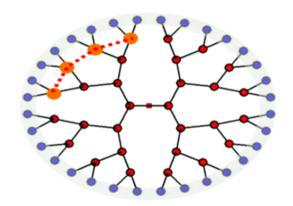


 Embedding of the bulk sites in AdS is determined by scale invariance (special for the critical state)

$$(j,n) \to (\rho,\theta), j = 1,2,..., 2^{N-n}$$
  
 $\rho = 2^{N-n} \frac{1}{2\pi}, \theta = \left(j - \frac{1}{2} + \frac{1}{2^{n+1}}\right) \frac{1}{\rho},$ 

• Angle direction  $d_{(x,n)(y,n)}/\xi \simeq \frac{2R}{\xi} \log \frac{|x-y|}{R}$ 

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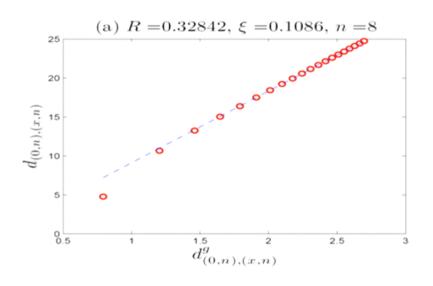
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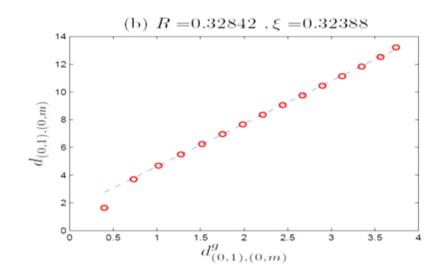
Fitting of the angle direction distance

→ AdS radius

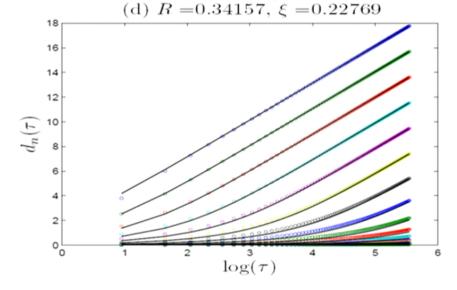
$$R = (64\pi^2 \log 4)^{-\frac{1}{6}} \simeq 0.33, \xi = R/3,$$
(Ching Hua Lee &XLQ)

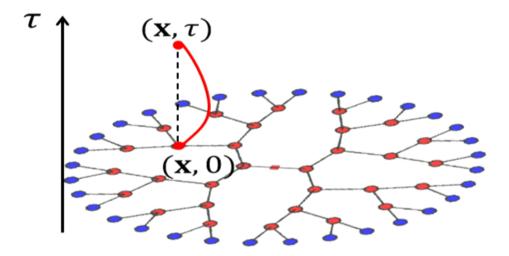
• Radial direction  $d_{(x,n),(x,1)}$ , qualitatively agree





- Time-direction: Imaginary time direction distance defined by  $d_{(x,\tau)(x,0)} = -\xi \log \langle Tb_{x\sigma}(\tau)b_{x\sigma}^+(0) \rangle$
- Compared with  $AdS_3$  formula  $d_{(x,\tau),(x,0)}=R \operatorname{acosh}\left[\left(\frac{\rho^2}{R^2}+1\right) \operatorname{cosh}\frac{\tau}{R}-\frac{\rho^2}{R^2}\right]$
- R can be independently fitted  $R \simeq 0.34$



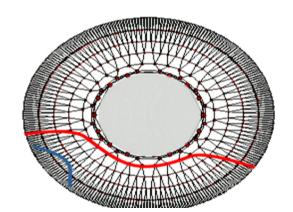


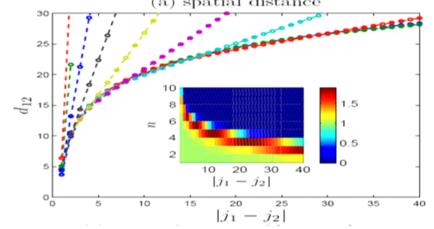
# Finite temperature and black hole

- T>0, m=0. Geometry is modified for non-critical systems, even if the same mapping is chosen.
- Spatial direction:  $d_{(x,n),(y,n)}$  Cross-over from AdS ( $\propto \log |x-y|$ ) to Euclidean ( $\propto |x-y|$ ).
- IR limit  $n \to N$ , Stretched horizon region.

• angle-direction distance  $d_{0n,xn}\simeq (1-2\pi T)^{2^{n+1}x}\Rightarrow$  Area saturates to a finite value  $2\pi \rho\simeq 4\pi T\cdot 2^N$  (Ching Hua

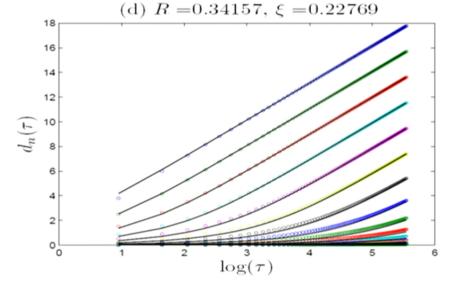
Lee &XLQ)

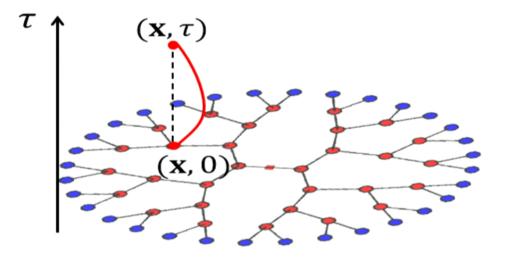




Pirsa: 15020134 Page 20/38

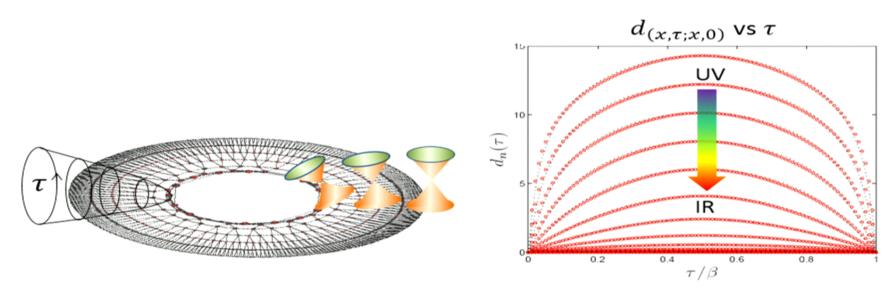
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- R can be independently fitted  $R \simeq 0.34$





#### Finite temperature and black hole

- Time direction: In IR region, fermion bandwidth $\ll T$
- The time dependence of correlation function exponentially slows down.  $d_{(x,\tau),(x,0)} o 0$  in IR
- General reason: reduced density matrix of the stretched horizon region  $\rho_{IR} \propto I$  Trivial time evolution.



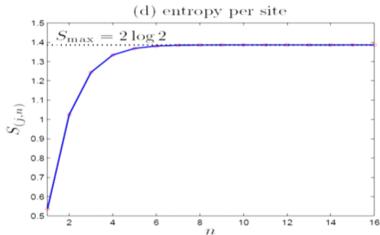
Pirsa: 15020134 Page 22/38

#### Finite temperature and black hole

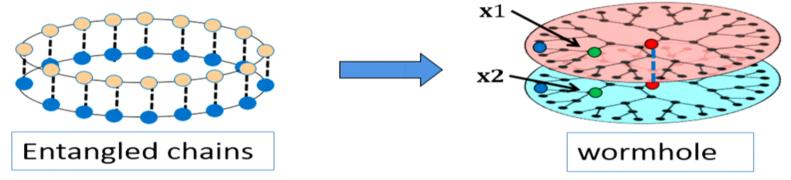
The observed behavior agrees with BTZ blackhole metric

• 
$$ds^2 = (\rho^2 - b^2)d\tau^2 + R^2 \frac{d\rho^2}{\rho^2 - b^2} + \rho^2 d\theta^2$$

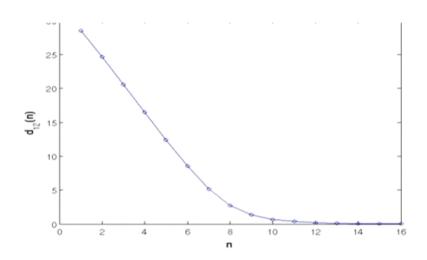
- Near the horizon  $\rho \to b$ , the time direction becomes infinitely short. We observe this behavior in the correlation functions, because the states in IR are in the infinite temperature limit.
- Each bulk-site in the IR carries maximal entanglement entropy.
- $\log D = \log 4$ .
- The entropy in this region shall be considered as the black-hole entropy



#### Einstein-Rosen bridge



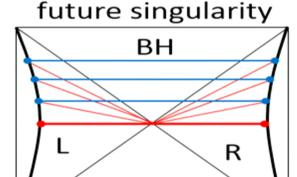
- There are different black hole geometries.
- A pure state of a two-leg ladder system in a thermal double state  $|\Psi\rangle=\sum_n e^{-\beta E_n/2}\,|\bar{n}\rangle_L|n\rangle_R$  is dual to a worm hole (Maldacena '01)
- Time evolution of two-sided blackhole (Hartman&Maldacena '13)



Pirsa: 15020134 Page 24/38

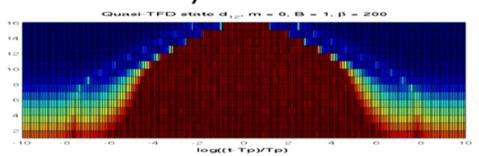
# Quantum quench on the wormhole geometry

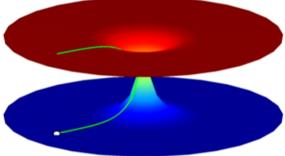
- The blackhole interior can be indirectly probed by the distance between outside points (Hartman & Maldacena JHEP '13)
- Time evolution by  $H_R H_L$  gives  $|\Psi(t)\rangle = \sum_n e^{-\left(2it + \frac{\beta}{2}\right)E_n} |\bar{n}\rangle_L |n\rangle_R$ , and changes the entanglement structure between the two sides.



WH

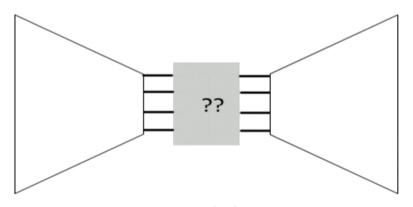
• In free theory, the worm hole stretches and returns in time  $t\sim L/2$ 

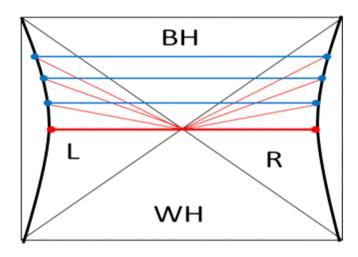




#### Tensor network point of view

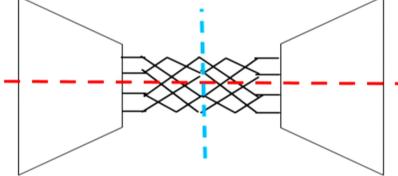
Outside frame (red)





 Hartman-Maldacena tensor network. Does it correspond to the minimal surface (blue)?

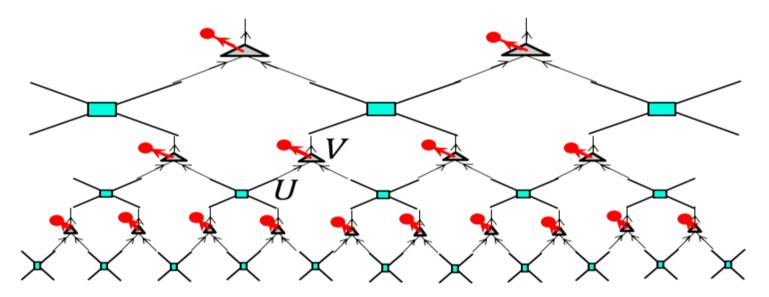
 It looks more like the green slice. The scrambling of time evolution is removed. Throat size never shrinks.



Pirsa: 15020134 Page 26/38

#### More general EHM tensor networks

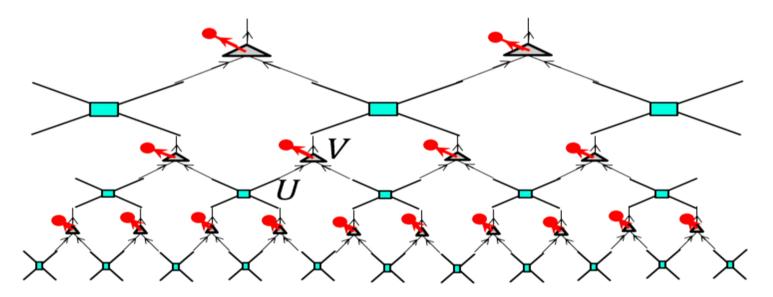
- In general, the suitable EHM should depend on the physical state.
- In the same way as MERA, we can modify EHM by introducing disentanglers



Pirsa: 15020134 Page 27/38

#### More general EHM tensor networks

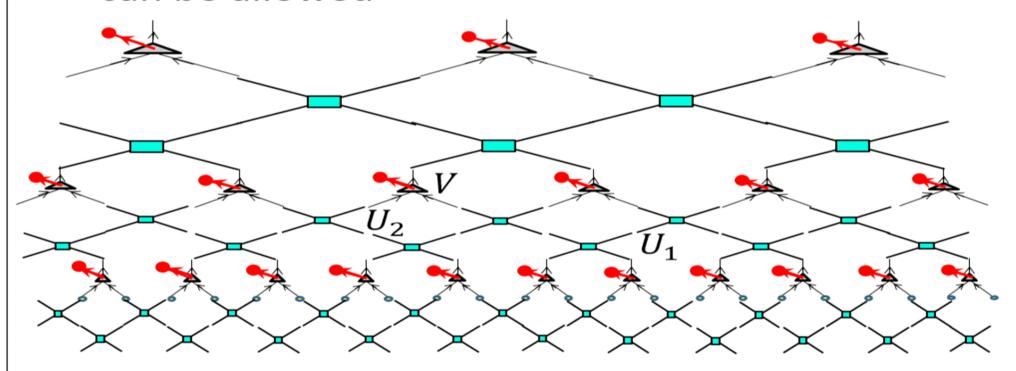
- In general, the suitable EHM should depend on the physical state.
- In the same way as MERA, we can modify EHM by introducing disentanglers



Pirsa: 15020134 Page 28/38

# More general EHM tensor networks

 More generally, multilayers of disentanglers can be allowed



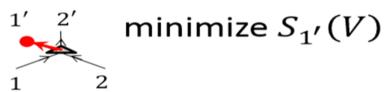
Pirsa: 15020134 Page 29/38

#### Criteria of optimal EHM tensor networks

- Although there is probably no unique choice of EHM which is the "best" choice, there shall be one family of tensor networks that are "optimal", which means the remaining entanglement of the bulk states is most short-ranged compared to other EHM with the same structure.
- Criteria:
- For disentanglers



For coarse-graining



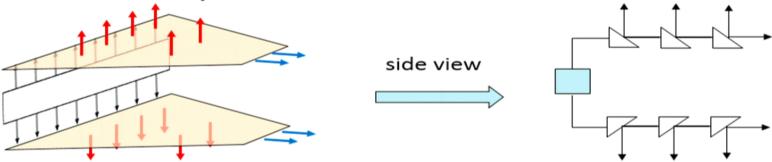
"Distill" the entanglement to fewer IR qubits

Possibly related criteria in Evenbly&Vidal PRB '10

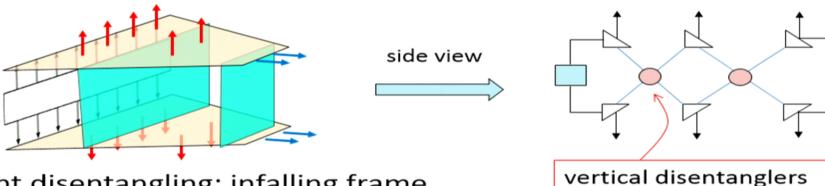
Pirsa: 15020134 Page 30/38

#### Application of the optimal EHM to the thermal double state

Two different disentangling scheme can be applied to the doubled system.



Separate disentangling: outside frame

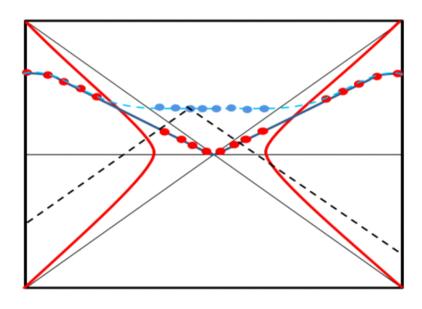


Joint disentangling: infalling frame

Pirsa: 15020134 Page 31/38

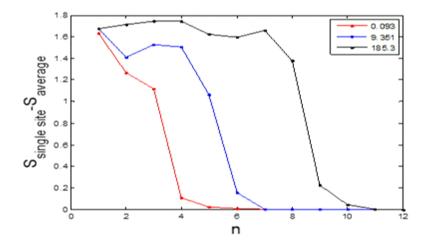
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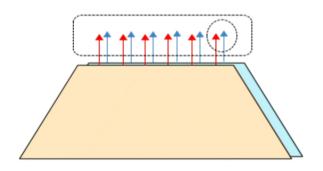
- For the quenched thermal double state  $|\Psi(t)\rangle=\sum_n e^{-rac{eta E_n}{2}-2iE_nt}|n_L\rangle|n_R\rangle$  the separate disentangling stops working at energy scale  $E\!\sim\!T=1/eta$ .
- Outside frame (red dots): separate left and right bulk regions. Left and right are not smoothly connected
- Infalling frame (blue dots): states in IR depends on both sides of the boundary.
- Blue states are further disentangled compared to red, to define a smooth geometry.

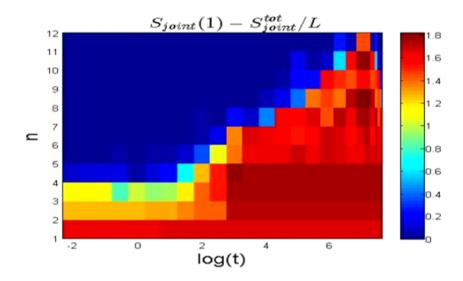


#### Numerical results

- Carrying the disentangling procedure until the IR states are direct products  $ho_{IR} \simeq \prod_i 
  ho_i$
- This factorization can be judged by the value of  $\sum S(\rho_i) S(\rho_{IR})$



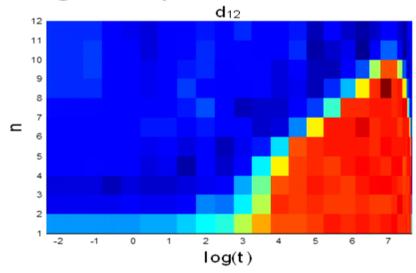


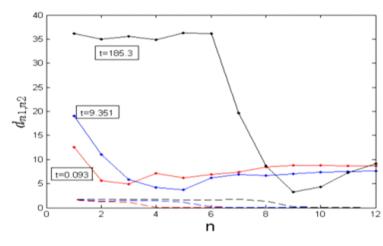


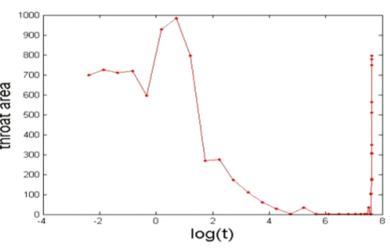
Pirsa: 15020134 Page 33/38

#### Numerical results

- Distance between bulk sites
- The throat remains smooth, with a shrinking area (and expand again at  $t \sim L/2$ )
- Question: Can we see the singularity?



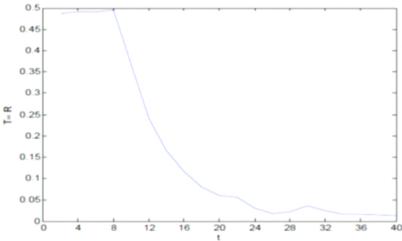


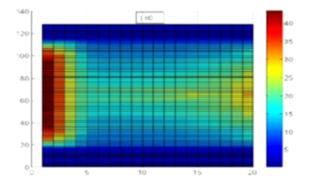


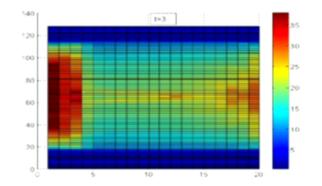
Pirsa: 15020134 Page 34/38

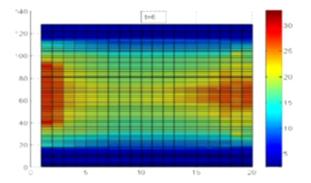
Topological edge states and black hole evaporation

- Temperature vs time
- Difference from semiclassical blackhole: specific heat is still positive



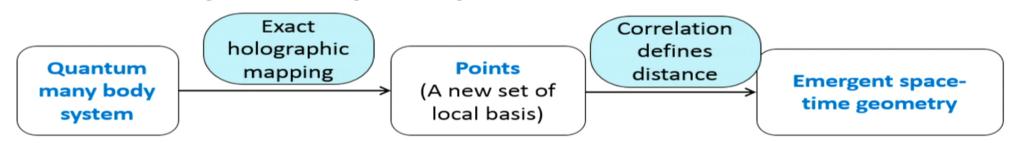






Pirsa: 15020134 Page 35/38

#### Summary and open questions

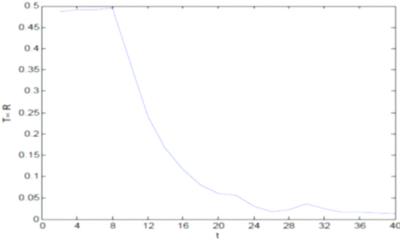


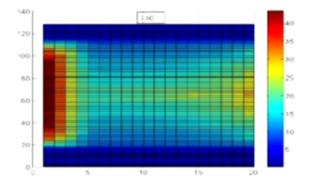
- Relation between quantum entanglement and space-time geometry.
- A quantum model for blackholes.
- A possible new approach to strongly correlated condensed matter systems.
- Open questions:
  - Application to physical strongly correlated problems.
  - ♦ The black-hole information paradox.
  - Or How to describe a "quantum" geometry, beyond the concepts of points and distance?

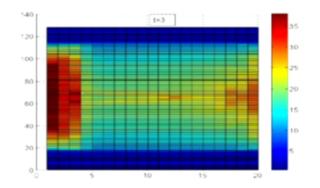
Pirsa: 15020134 Page 36/38

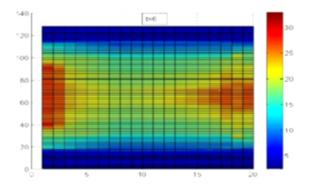
Topological edge states and black hole evaporation

- Temperature vs time
- Difference from semiclassical blackhole: specific heat is still positive



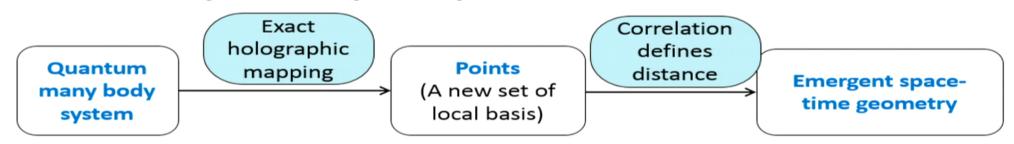






Pirsa: 15020134 Page 37/38

#### Summary and open questions



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Pirsa: 15020134 Page 38/38