

Title: Gravitational Physics-6

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Abstract:

## Lecture 6 Gravity & field theory

In field theory we use a Lagrangian formulation, on the manifold we need to be able to integrate.

$$d^4x$$

However  $\sqrt{|g|} d^4x$  is inv

To vary this volume element

use  $\det M = \exp \text{tr} \log M$

$$\Rightarrow \delta(\det M) = \det M \text{tr}(M^{-1} \delta M)$$

However  $\sqrt{|g|} d^4x$  is invariant

any this volume element

use  $\det M = \exp \operatorname{tr} \log M$

$$\Rightarrow \delta(\det M) = \det M \operatorname{tr}(M^{-1} \delta M)$$

$$\Rightarrow \delta \sqrt{|g|} = \frac{1}{2} \frac{1}{\sqrt{|g|}} g^{ab} \delta g_{ab}$$

int

ment

$\text{tr} \log M$

$\text{tr}(M^{-1} \delta M)$

$$\delta g_{ab} = -\frac{\sqrt{-g}}{2} g_{ab} \delta g^{ab}$$

e.g. Massless scalar  $\mathcal{L}_\phi = \frac{1}{2}(\partial\phi)^2$

$$S_\phi = \int d^4x \sqrt{g}$$

inv

ment

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e.g. Massless scalar  $\mathcal{L}_\phi = \frac{1}{2}(\partial\phi)^2$

$$S_\phi = \int d^4x \sqrt{g} \frac{1}{2}(\partial\phi)^2$$

$$\delta S_\phi = \int d^4x \sqrt{g} \partial_\mu \phi \partial_\nu \delta\phi g^{\mu\nu}$$

int

ent

$r \log M$

$M \delta(M^{-1} \delta M)$

$$\delta g_{ab} = -\frac{\sqrt{-g}}{2} g_{ab} \delta g^{ab}$$

e.g. Massless scalar  $\mathcal{L}_\phi = \frac{1}{2}(\partial\phi)^2$ .

$$S_\phi = \int d^4x \sqrt{g} \frac{1}{2}(\partial\phi)^2 = \int d^4x \frac{\sqrt{g}}{2} \phi_{,\mu} \phi_{,\nu} g^{\mu\nu}$$

$$\delta S_\phi = \int d^4x \sqrt{g} \partial_\mu \phi \partial_\nu \delta\phi g^{\mu\nu} + \frac{1}{2} \sqrt{g} (\phi_{,\mu} \phi_{,\nu} - \frac{1}{2}(\partial\phi)^2 g_{\mu\nu}) \delta g^{\mu\nu}$$

invt  
 ment  
 $\text{tr} \log M$   
 $M \text{tr}(M^{-1} \delta M)$   
 $\delta g_{ab} = -\frac{\sqrt{-g}}{2} g_{ab} \delta g^{ab}$

e.g. Massless scalar  $\mathcal{L}_\phi = \frac{1}{2}(\partial\phi)^2$ .

$$S_\phi = \int d^4x \sqrt{g} \frac{1}{2}(\partial\phi)^2 = \int d^4x \frac{\sqrt{g}}{2} \phi_{,\mu} \phi_{,\nu} g^{\mu\nu}$$

$$\delta S_\phi = \int d^4x \sqrt{g} \partial_\mu \phi \partial_\nu \delta\phi g^{\mu\nu} + \frac{1}{2} \sqrt{g} (\phi_{,\mu} \phi_{,\nu} - \frac{1}{2}(\partial\phi)^2 g_{\mu\nu}) \delta g^{\mu\nu}$$

$$= \int d^3x \sqrt{g} \frac{\delta S}{\delta \phi}$$

int

ment

$\text{tr} \log M$

$M^{-1} \delta M$

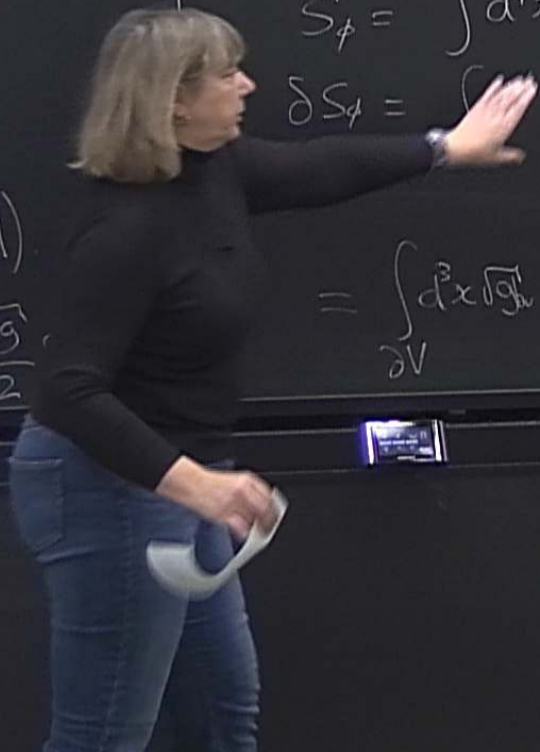
$\delta g_{ab} = -\frac{\sqrt{g}}{2}$

e.g. Massless scalar  $\mathcal{L}_\phi = \frac{1}{2}(\partial\phi)^2$ .

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$$\delta S_\phi = \int d^4x \sqrt{g} \partial_\mu \phi \partial_\nu \delta\phi g^{\mu\nu} + \frac{1}{2} \sqrt{g} (\phi_{,\mu} \phi_{,\nu} - \frac{1}{2}(\partial\phi)^2 g_{\mu\nu}) \delta g^{\mu\nu}$$

$$= \int d^3x \sqrt{g} (n^\mu \partial_\mu \phi) + \int d^4x \sqrt{g} \left[ -\frac{1}{\sqrt{g}} \partial_\mu (\sqrt{g} g^{\mu\nu} \phi) \right]$$



invt  
 ment  
 tr log M  
 tr(M^{-1} \delta M)  
 $\delta g_{ab} = -\frac{\sqrt{g}}{2} g_{ab} \delta g^{ab}$

e.g. Massless scalar  $\mathcal{L}_\phi = \frac{1}{2}(\partial\phi)^2$

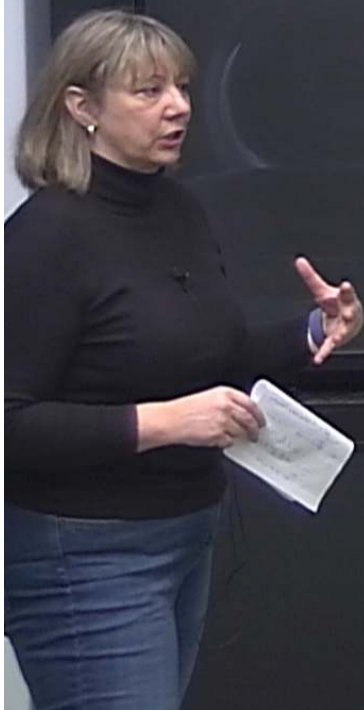
$$S_\phi = \int d^4x \sqrt{g} \frac{1}{2}(\partial\phi)^2 = \int d^4x \sqrt{g} \frac{1}{2} \phi_{,\mu} \phi_{,\nu} g^{\mu\nu}$$

$$\delta S_\phi = \int d^4x \sqrt{g} \partial_\mu \phi \partial_\nu \delta\phi g^{\mu\nu} + \frac{1}{2} \sqrt{g} (\phi_{,\mu} \phi_{,\nu} - \frac{1}{2}(\partial\phi)^2 g_{\mu\nu}) \delta g^{\mu\nu}$$

$$= \int d^3x \sqrt{g} (n^\mu \partial_\mu \phi) + \int d^4x \sqrt{g} \left[ \frac{-1}{\sqrt{g}} \partial_\mu (\sqrt{g} \delta\phi) \right] + \frac{1}{2} [\partial_\mu \phi \partial_\nu \phi - \frac{1}{2}(\partial\phi)^2 g_{\mu\nu}] \delta g^{\mu\nu}$$

$$\frac{\delta S_\phi}{\delta \phi} =$$

For gravitational action, need scalar



For gravitational action, need scalar

$$\delta R = \delta R_{\mu\nu} g^{\mu\nu} + R_{\mu\nu} \delta g^{\mu\nu}$$

$$\delta R_{\mu\nu} = \delta \Gamma^{\lambda}_{\mu\nu;\lambda} - \delta \Gamma^{\lambda}_{\mu\nu;\lambda}$$

Hence:

$$\delta R_{ab} = \frac{1}{2} \left[ \nabla_c \nabla_a \delta g^c_b + \nabla_b \nabla_a \delta g^c_a - \nabla_c \nabla^c \delta g_{ab} - \nabla_b \nabla_a \delta g^c_c \right]$$

change to  $\delta g^{-1}$

$$\delta R_{\mu\nu} g^{\mu\nu} = -\nabla_\mu \nabla_\nu \delta g^{\mu\nu} + g_{\mu\nu} \delta g^{\mu\nu}$$

This is a total deriv

Lecture 6 Gravity & field theory

In field theory we use a Lagrangian formulation on the manifold  $M$ .

$$\delta R_{\mu\nu} = \delta \Gamma^{\lambda}_{\mu\nu;\lambda} - \delta \Gamma^{\lambda}_{\mu\lambda;\nu}$$

$$= \delta \Gamma^{\lambda}_{\mu\nu;\lambda} - \delta \Gamma^{\lambda}_{\mu\lambda;\nu}$$

NC

← TENSORIAL

$$\delta \Gamma^{\mu}_{\nu\lambda} = \frac{1}{2} g^{\mu\sigma} (\delta g_{\sigma\nu;\lambda} + \delta g_{\sigma\lambda;\nu} - \delta g_{\nu\lambda;\sigma})$$

hence

$$\delta \Gamma^a_{bc} = \frac{1}{2} [\nabla_b \delta g^a_c + \nabla_c \delta g^a_b - \nabla^a \delta g_{bc}]$$

$$\nabla_b \delta g^a_c$$

$$g_{\mu\nu} \delta g^{\mu\nu}$$

⚠ This is a total deriv, but, the boundary term we get on integration is

$$\int d^3x n_{\mu} \nabla_{\nu} \delta g^{\mu\nu} - n^{\mu} \nabla_{\mu} \delta g$$

So we need to set normal derivs of  $\delta g$  to zero.

We therefore write

-

Combine with interacting scalar:

$$L_\phi = \frac{1}{2}(\partial\phi)^2 - V(\phi)$$

$\hookrightarrow \frac{\lambda}{2}(\phi^2 - \eta^2)^2$

$$S_{\text{TOT}} = S_{\text{EH}} + S_\phi$$

$V$  has two vacua  $\phi = \pm\eta$



$\phi$  - e.o.m.

$$\square\phi + 2\lambda(\phi^2 - \eta^2)\phi = 0$$

Ask that  $\phi \rightarrow \pm\eta$  as  
 $z \rightarrow \pm\infty$  in Minkowski  
in absence of G.R.

$$\phi = \phi(z)$$

Then  $\phi = \eta \tanh[\sqrt{\lambda}\eta(z-z_0)]$

$$\phi' = \sqrt{\lambda}\eta^2 \text{sech}^2[\dots]$$

$\eta^2$



Combine with interacting scalar

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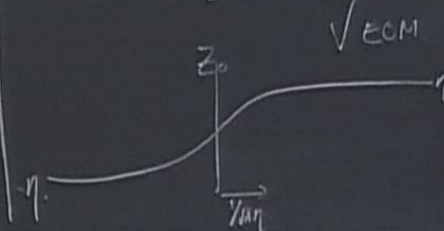
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Then  $\phi = \eta \tanh[\sqrt{\lambda}\eta(z-z_0)]$

$$\phi' = \sqrt{\lambda}\eta^2 \text{sech}^2[\dots]$$

$$\phi'' = -2\lambda\eta^3 \text{sech}^2[\dots] \tanh[\dots]$$

$$= -2\lambda\eta^2 (1 - \phi^2/\eta^2)\phi$$



1th interacting scalar

$$\phi^2 - V(\phi) \rightarrow \frac{\lambda}{2} (\phi^2 - \eta^2)^2$$

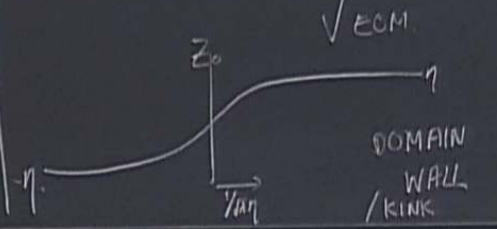
two  $\phi = \pm \eta$

$\phi$ - e.o.m.  $\square \phi + 2\lambda(\phi^2 - \eta^2)\phi = 0$

Ask that  $\phi \rightarrow \pm \eta$  as  $z \rightarrow \pm \infty$  in Minkowski in absence of G.R.

$$\phi = \phi(z)$$

Then  $\phi = \eta \tanh[\sqrt{\lambda} \eta (z - z_0)]$   
 $\phi' = \sqrt{\lambda} \eta^2 \text{sech}^2[\dots]$   
 $\phi'' = -2\lambda \eta^3 \text{sech}^2[\dots] \tanh[\dots]$   
 $= -2\lambda \eta^2 (1 - \phi^2/\eta^2) \phi$



Topologically stable



th interacting scalar:

$$\phi^2 - V(\phi) \rightarrow \frac{\lambda}{2} (\phi^2 - \eta^2)^2$$

$S \neq$

two vacua  $\phi = \pm \eta$



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Ask that  $\phi \rightarrow \pm \eta$  as  $z \rightarrow \pm \infty$  in Minkowski in absence of G.R.

$$\phi = \phi(z)$$

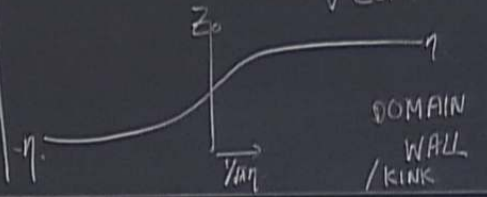
Then  $\phi = \eta \tanh[\sqrt{\lambda} \eta (z - z_0)]$

$$\phi' = \sqrt{\lambda} \eta^2 \text{sech}^2[\dots]$$

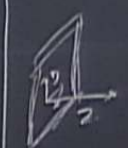
$$\phi'' = -2\lambda \eta^3 \text{sech}^2[\dots] \tanh[\dots]$$

$$= -2\lambda \eta^2 (1 - \phi^2/\eta^2) \phi$$

$\sqrt{\text{EOM}}$



Topologically stable



$$T_{\mu\nu} = \phi_{,\mu} \phi_{,\nu} - (\phi_{,\alpha} \phi^{,\alpha}) \delta_{\mu\nu}$$

$$= \dots \delta_{\mu\nu} + \dots$$



$$T_0^0 = T_x^x = T_y^y = 2V, \quad T_z^z = 0$$

Mostly a vacuum, but wall has  
 $\infty$  area  $\int T_0^0 dz = \frac{4}{3} \sqrt{\lambda} \eta^3$ .

+ GRAVITY  $T_0^0$  etc looks like  $\Lambda$   
in 3D space at  $Z_0$ .

$$\text{Try } ds^2 = A^2(z) \gamma_{\mu\nu} dx^\mu dx^\nu - dz^2.$$

WARNING indicates  $t, x \dots$  not 4D!

$$= T^y_y = 2V, \quad T^z_z = 0$$

vacuum, but wall has

$$\int P_0 dz = \frac{4}{3} \sqrt{\lambda} \eta^3$$

$T^0_0$  etc looks like  
pressure at  $z_0$ .

$$\text{Try } ds^2 = A^2(z) \gamma_{\mu\nu} dx^\mu dx^\nu - dz^2$$

WARNING  $\mu$  indicates  $t, x$  ... not 4D!

Cartan:  $\gamma_{\mu\nu}$  arb (D-1) diml metric

$$\underline{\omega}^z = dz \quad \underline{\omega}^a = A(z) \omega^a_{\mu} e^{\mu}_a dx^\mu$$

$$d\underline{\omega}^z = 0; \quad d\underline{\omega}^a = \frac{A'}{A} \underline{\omega}^z \wedge \omega^a_{\mu} - \Theta^a_{\mu\nu} \omega^{\mu}_a \omega^{\nu}_b$$

$$\Rightarrow \Theta^a_{\hat{b}} = \Theta^a_{\mu} \hat{b}^{\mu} \quad \& \quad \Theta^a_{\hat{z}} = \frac{A'}{A} \underline{\omega}^a$$

Toret looks like  $\Lambda$   
 space at  $z_0$ .

$$\underline{\omega}^z = dz \quad \underline{\omega} = \omega^a e_a$$

$$d\underline{\omega}^z = 0; \quad d\underline{\omega}^a = \frac{A'}{A} \underline{\omega}^z \wedge \underline{\omega}^a - \Theta_0^a \wedge \underline{\omega}^b$$

$$R^{\hat{a}}_{\hat{z}} = \frac{A''}{A} \underline{\omega}^z \wedge \underline{\omega}^a - A' \Theta_0^a \wedge \underline{\omega}^z + \Theta_0^a \wedge \frac{A'}{A} \underline{\omega}^b$$

$$a = \frac{1}{A^2} R_0^a \hat{z} \hat{z} a + \left(\frac{A'}{A}\right)^2 (\delta_{z\hat{z}} \eta_{\hat{z}a} - \delta_{\hat{z}a} \eta_{z\hat{z}})$$

$$\hat{z} \hat{z} \hat{z} = -\frac{A''}{A} \delta^{\hat{z}a} \hat{z}$$

$$\lambda \tau = \frac{1}{A^2} R_0^{\mu\nu} \lambda \tau + \left(\frac{A'}{A}\right)^2 (\delta_{\lambda\tau}^{\mu\nu} - \delta_{\tau\lambda}^{\mu\nu})$$

$$v \tau = \frac{A''}{A}$$

Hence  $R^z_z = (D-1) \frac{A''}{A}$

$$R^{\mu\nu}_{\mu\nu} = \frac{1}{A^2} R_0^{\mu\nu} + (D-2) \left(\frac{A'}{A}\right)^2 \delta_{\mu\nu} + \frac{A''}{A} \delta^{\mu\nu}$$

where  $D$  is dim of  $ds^2$   
 For us,  $D=4$ , let  $\kappa/l^2$  be curv  
 $d\gamma$

$$G^z_z = 3 \left( \frac{\kappa}{l^2 A^2} - \frac{A''}{A} \right) = 8\pi G (V - \frac{1}{2} \phi^4)$$



$\phi_{\text{kink}}$ 

$$A = 1 - \epsilon \left( \frac{2}{3} \log \cosh z - \frac{1}{6} \operatorname{sech}^2 z + \frac{2}{3} \ln 2 \right) \sim \left| \frac{z}{l} \right|$$

⚠ This is a total deriv, but, the boundary term we get on integration is

$$\int d^3x n_\mu \nabla_\nu \delta g^{\mu\nu} - n^\mu \nabla_\mu \delta g$$

So we need to set normal derivatives of  $\delta g$  to zero.

We therefore write

$$S_{\text{EH}} = -\frac{1}{16\pi G} \int d^4x \sqrt{g} R$$

$$\& \frac{\delta S_{\text{EH}}}{\delta g^{\mu\nu}} = -\frac{1}{16\pi G} (R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu})$$