

Title: Macroscopic Quantum Objects

Date: Dec 18, 2014 01:00 PM

URL: <http://pirsa.org/14120018>

Abstract: <p>With recent advancement of experimental physics, macroscopic objects, which are typically well-described by classical physics, can now be isolated so well from their environment, that their quantum uncertainties can be studied quantitatively. In the research field called "optomechanics", mechanical motions of masses from picograms to kilograms are being prepared into nearly pure quantum states, and observed at time scales ranging from nanoseconds to milliseconds. In practice, optomechanics experiments have been constructed to measure extremely weak classical forces, e.g., due to gravitational waves, acting on macroscopic test objects. In this case, experiments must be designed in such a way that quantum uncertainties of the test objects are avoided as much as possible --- often by employing the quantum correlations between the state of light and the motion of the test object, which can build up during the measurement process.</p>

<p>Optomechanics experiments can also be used to search for possible deviations from standard quantum mechanics when macroscopic objects are involved. In this case, experiments are designed to highlight as much as possible the quantum-state evolution of the macroscopic objects. It is hoped that these macroscopic quantum mechanics experiments will either lead our way toward new physics, or at least put experimental constraints on how standard quantum mechanics might be modified.</p>



**LIGO
Scientific
Collaboration**

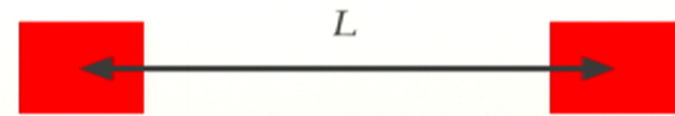
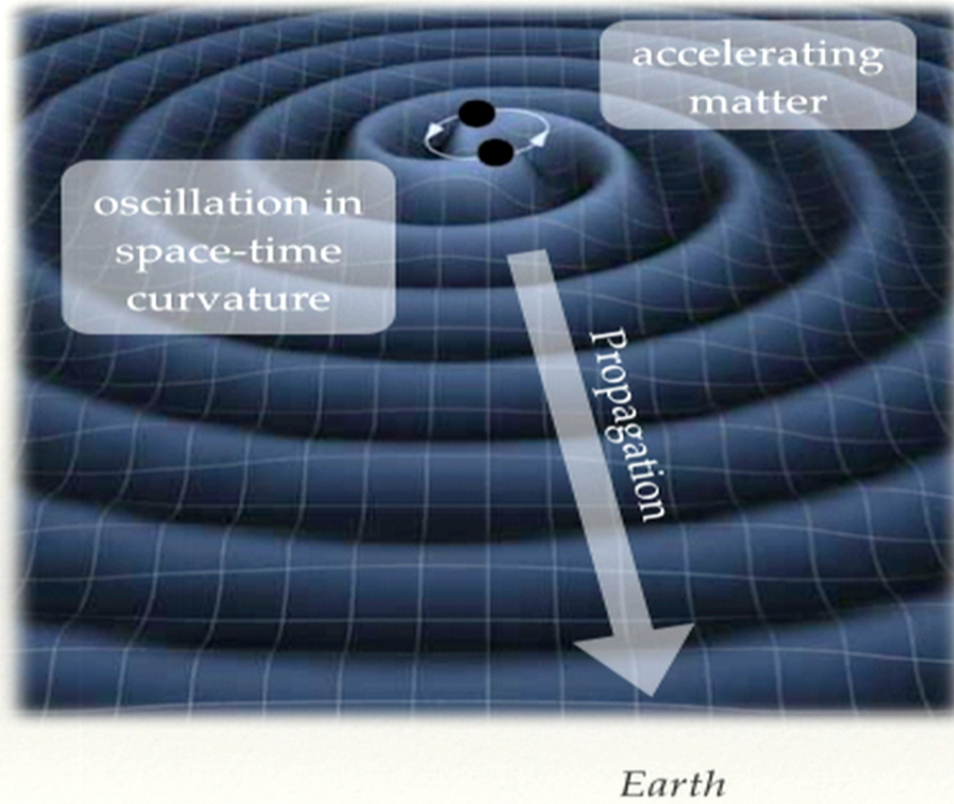
Macroscopic Quantum Objects

Yanbei Chen
California Institute of Technology



GORDON AND BETTY
MOORE
FOUNDATION

Gravitational Waves

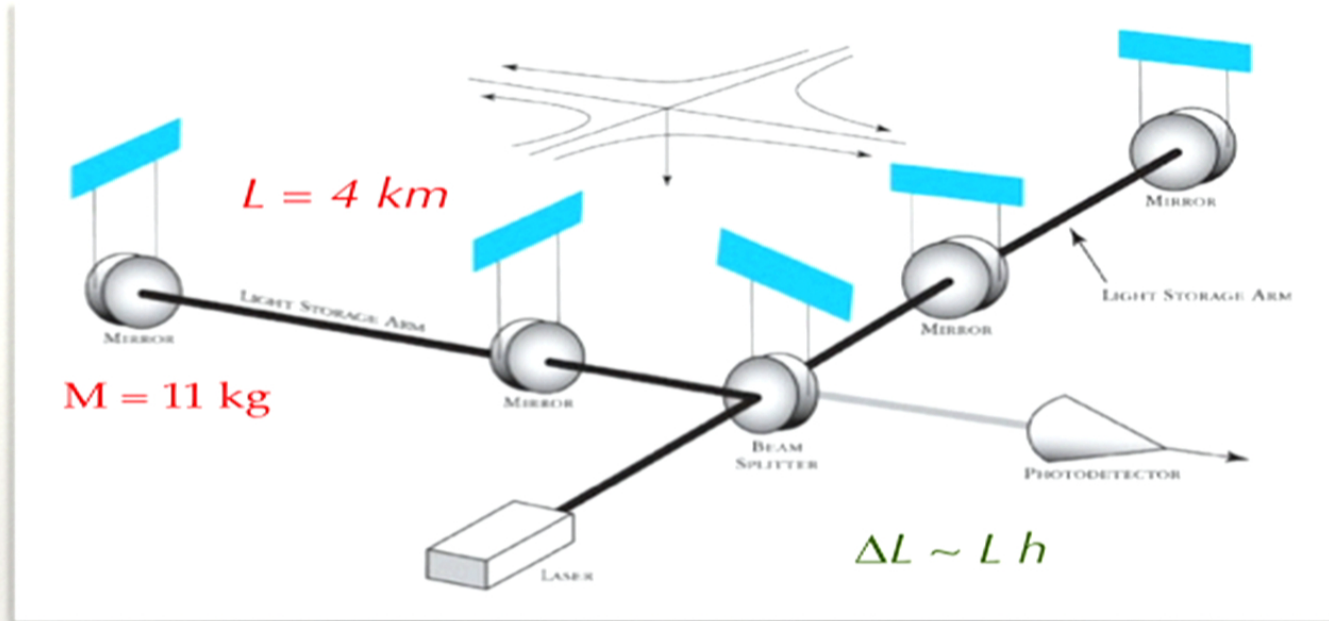


motions of free test masses
separated by $L \ll \lambda$

$$\Delta L \sim L h$$

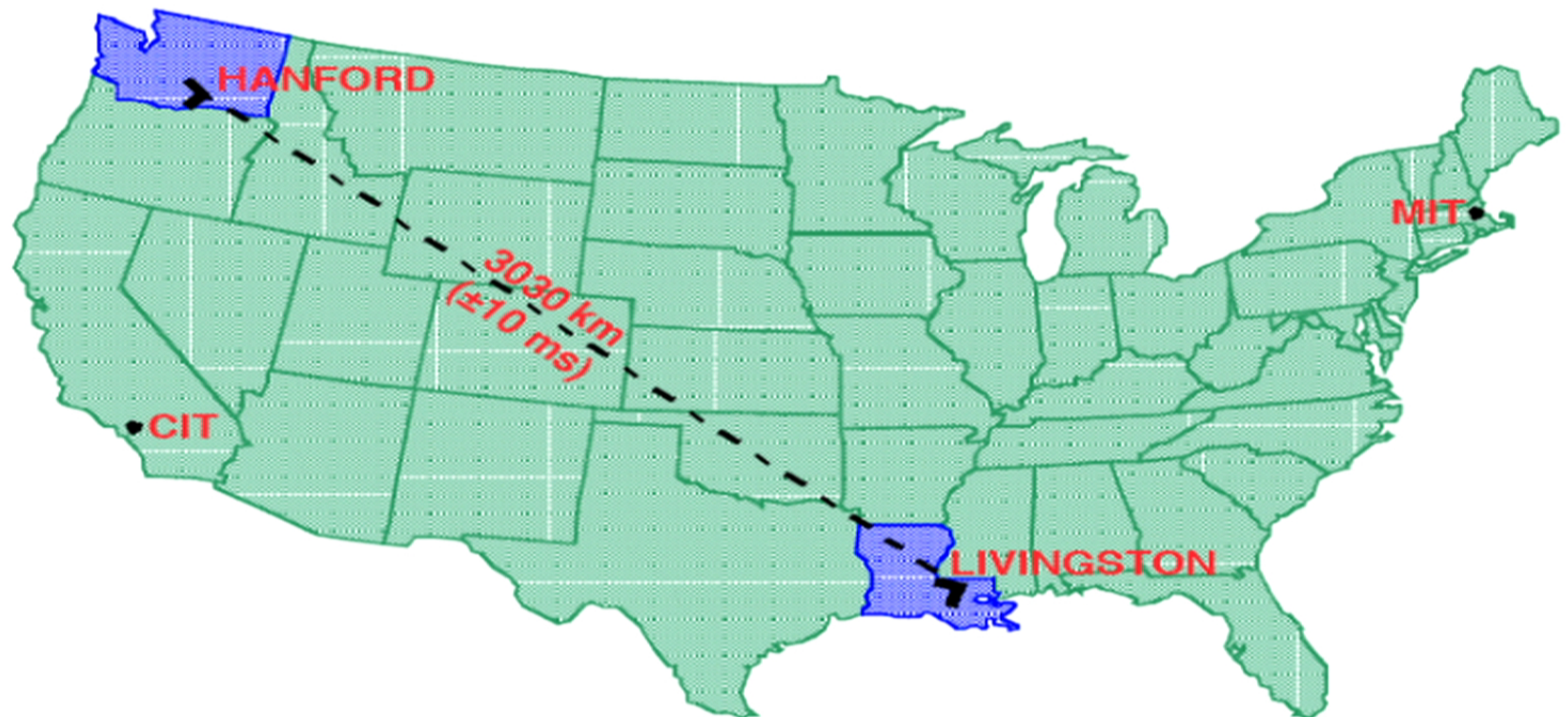
Laser Interferometer Gravitational Wave Detectors

A detector from the Laser Interferometer Gravitational-wave Observatory (LIGO)



- Michelson Interferometer, with 4 km arms and kg-scale mirrors
- In 2007, $\Delta h \sim 10^{-21}$, $\Delta x \sim 10^{-18} \text{ m}$ at 100 Hz (entire range 10 Hz -- 10 kHz)
 - Each photon resolves $\lambda/F \sim 10^{-8} \text{ m}$
 - 10^{20} photons injected within 0.01 second
- "Heisenberg Uncertainty": " $\Delta x \cdot \Delta p \sim 100 \hbar$ " ($\Delta p \sim m\Omega \cdot \Delta x$)
- 2nd generation detectors (~2015): " $\Delta x \cdot \Delta p \sim \hbar/2$ ", 3rd generation, " $\Delta x \cdot \Delta p < \hbar/2$ "

Laser Interferometer Gravitational-wave Observatory (LIGO)



4

Laser Interferometer Gravitational-wave Observatory (LIGO)



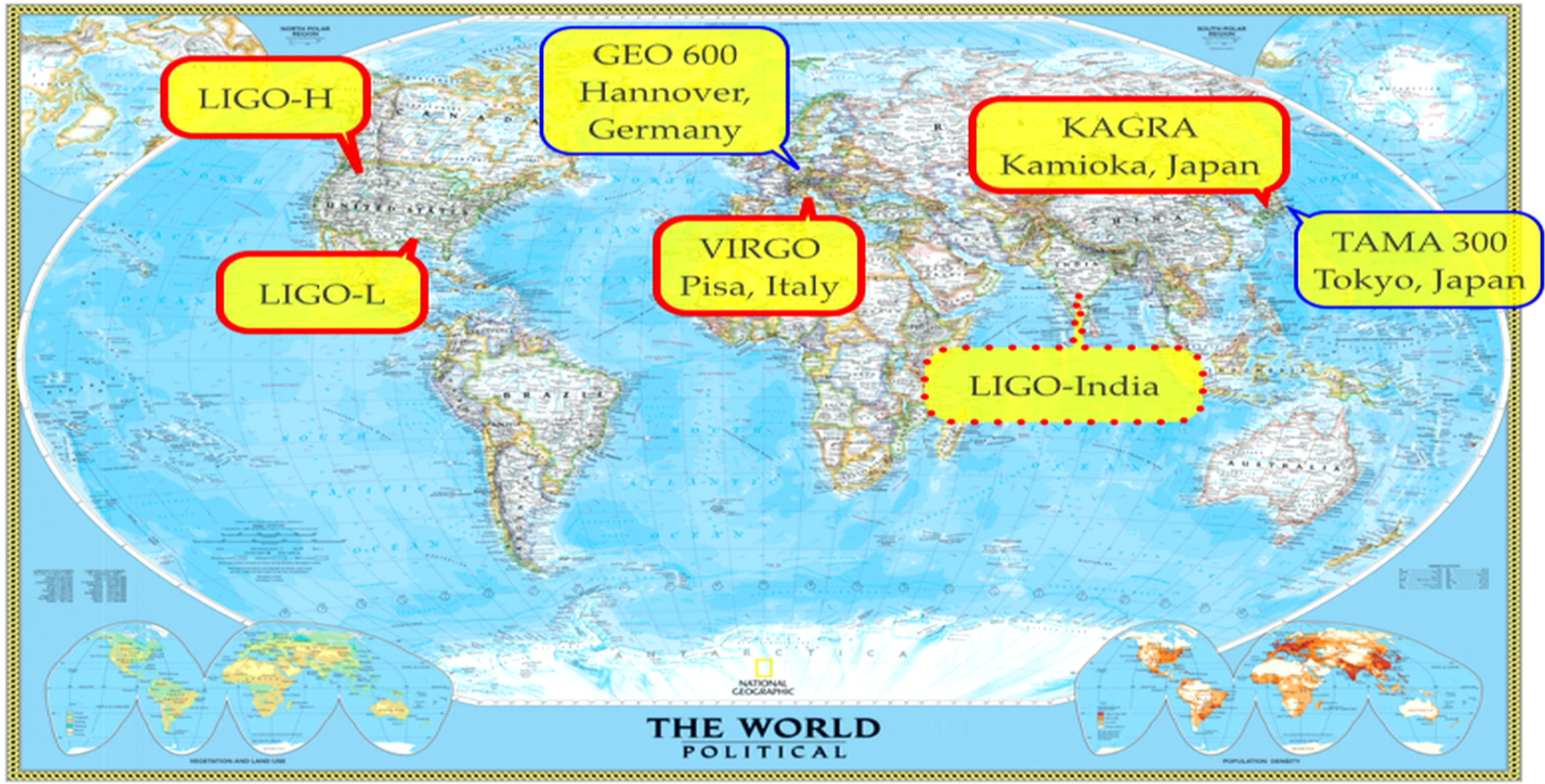
LIGO Hanford, WA site



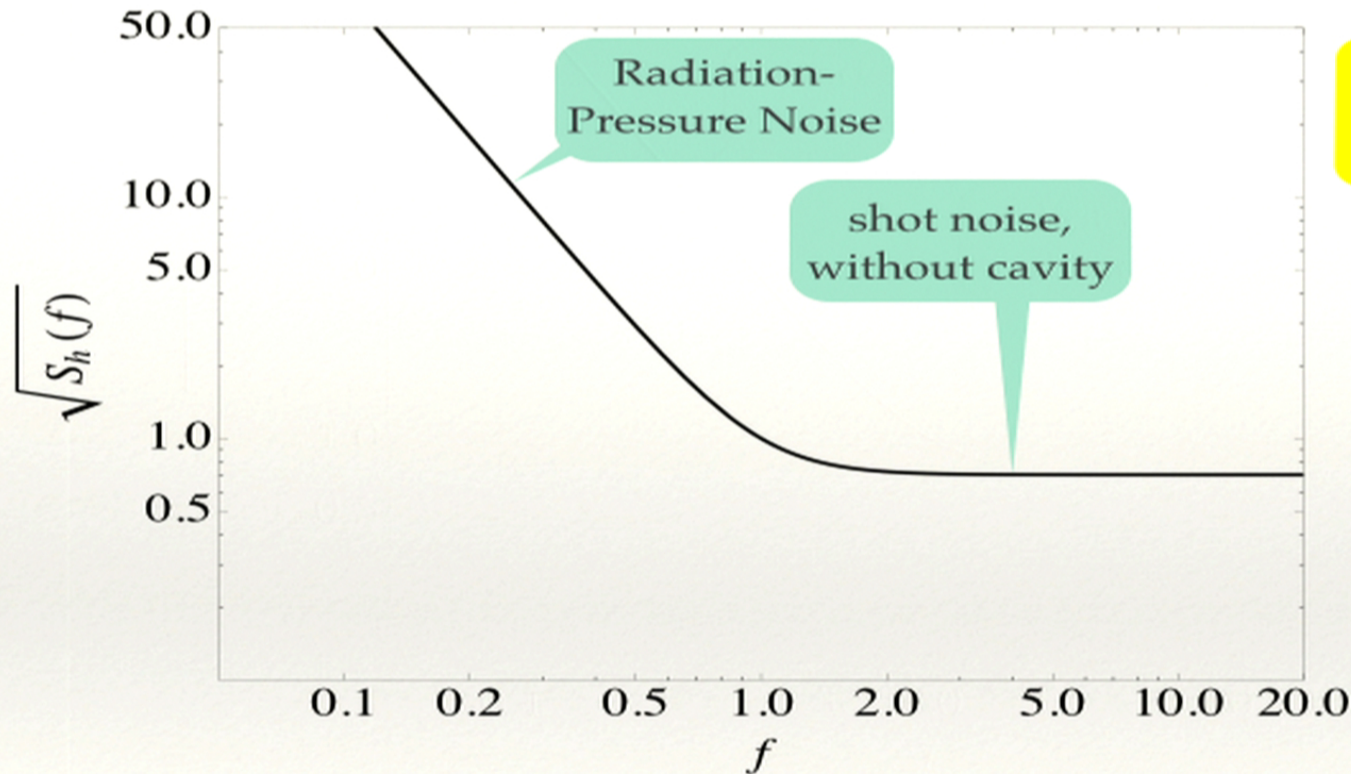
LIGO Livingston, LA site

4

A Global Network



Standard Quantum Limit

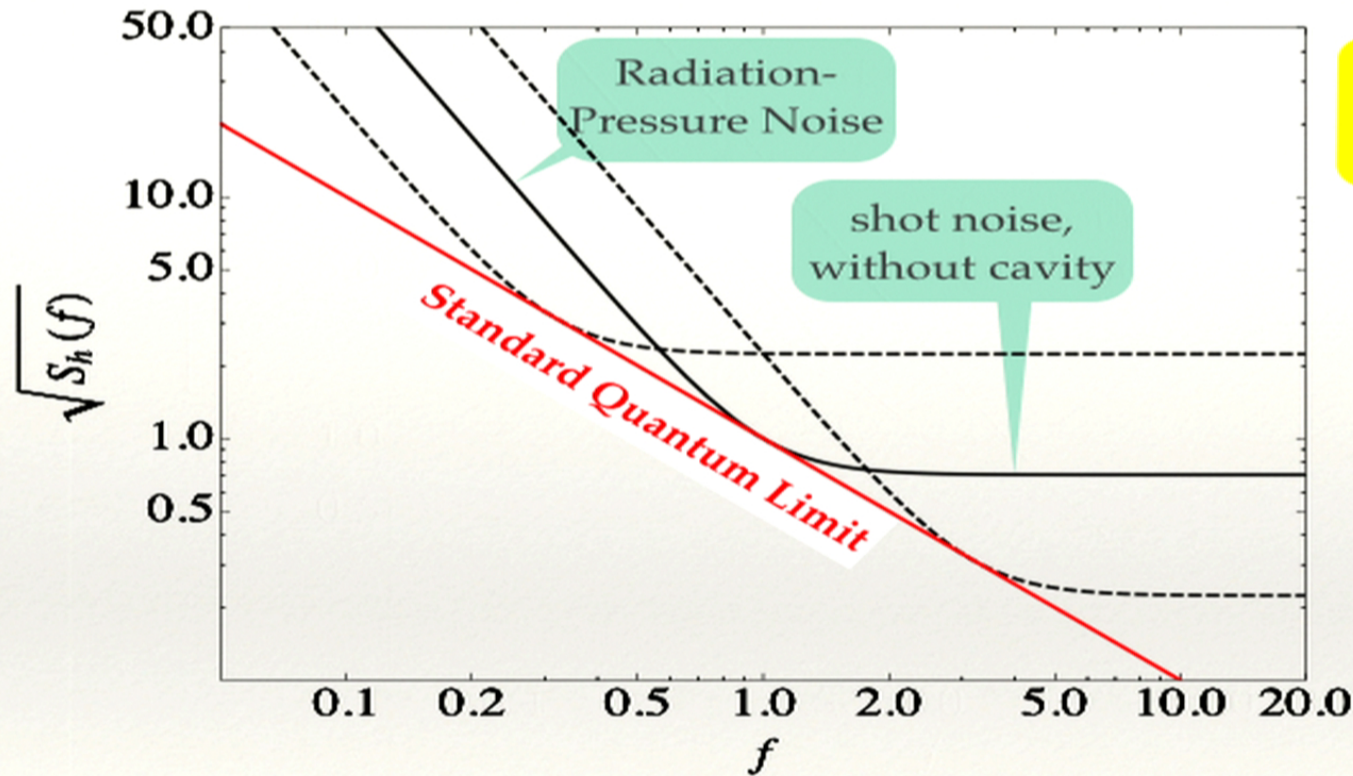


$S_h(f)$
Noise Spectral Density
(noise power / frequency)

$\sqrt{f \cdot S_h(f)}$
r.m.s. noise
for broadband signal.

- ❖ Shot Noise $\sqrt{S_{shot}} \sim 1 / \sqrt{I}$
- ❖ Radiation-Pressure Noise $\sqrt{S_{rad\ press}} \sim \sqrt{I} / (M\Omega^2)$

Standard Quantum Limit



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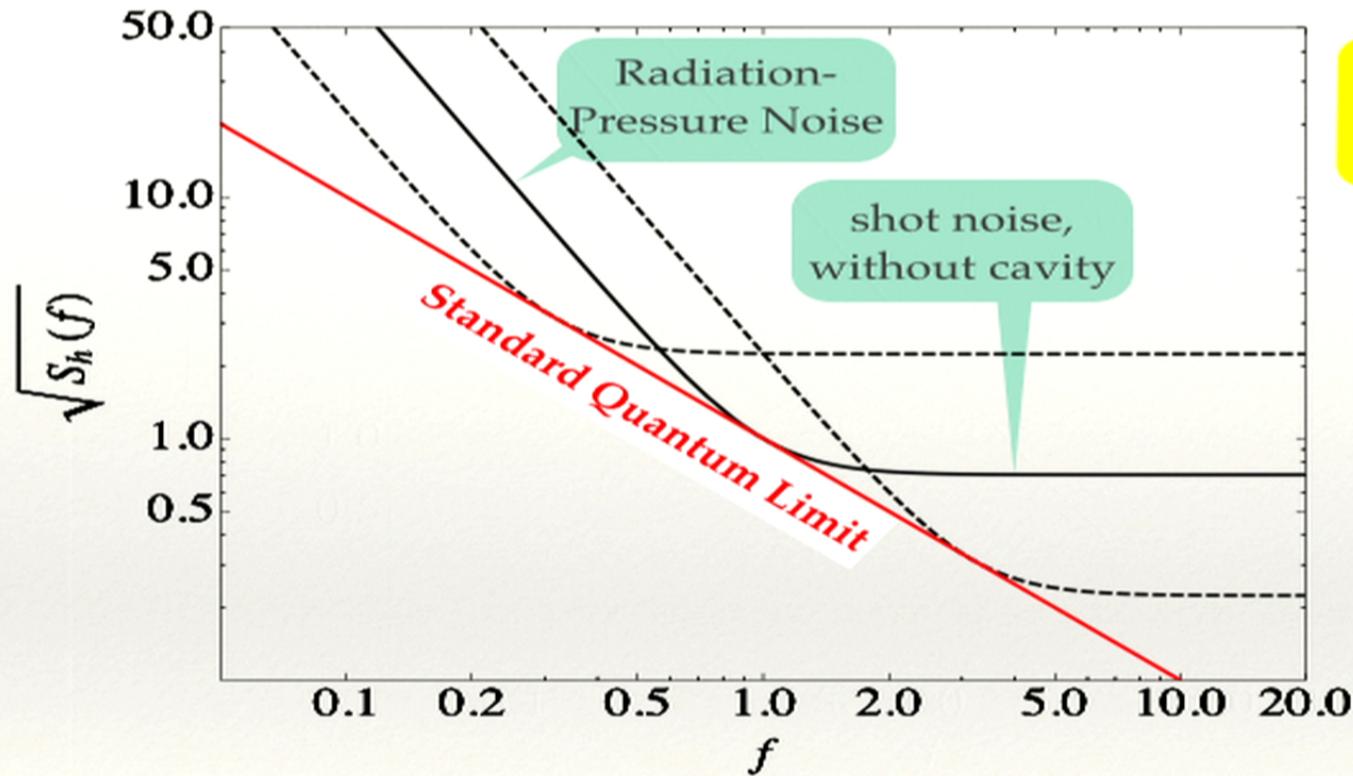
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Standard Quantum
Limit
$$\sqrt{S_h^{SQL}} = \sqrt{\frac{8\hbar}{M\Omega^2 L^2}}$$

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Trade-off Between Shot Noise and Radiation-Pressure Noise: Standard Quantum Limit

Standard Quantum Limit



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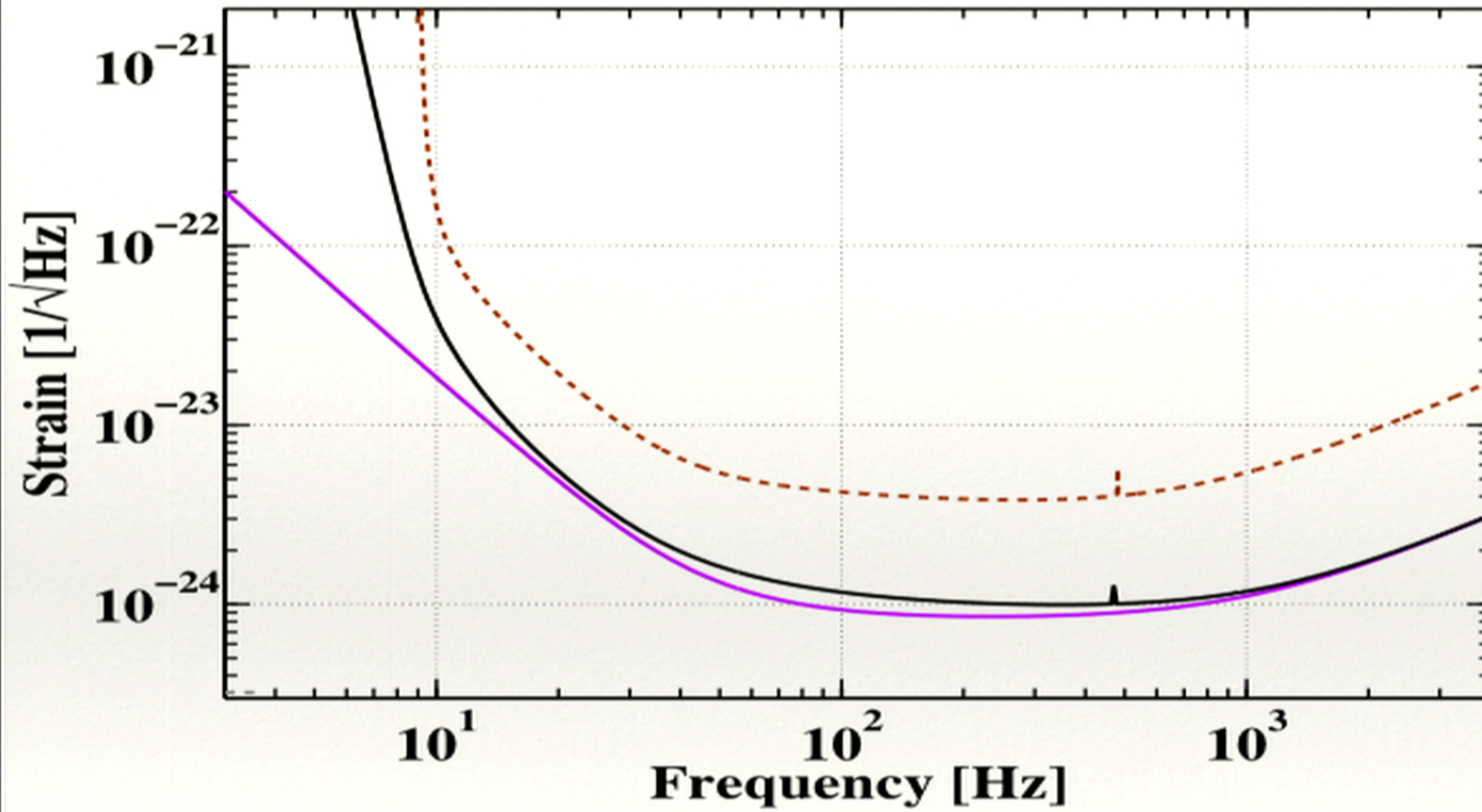
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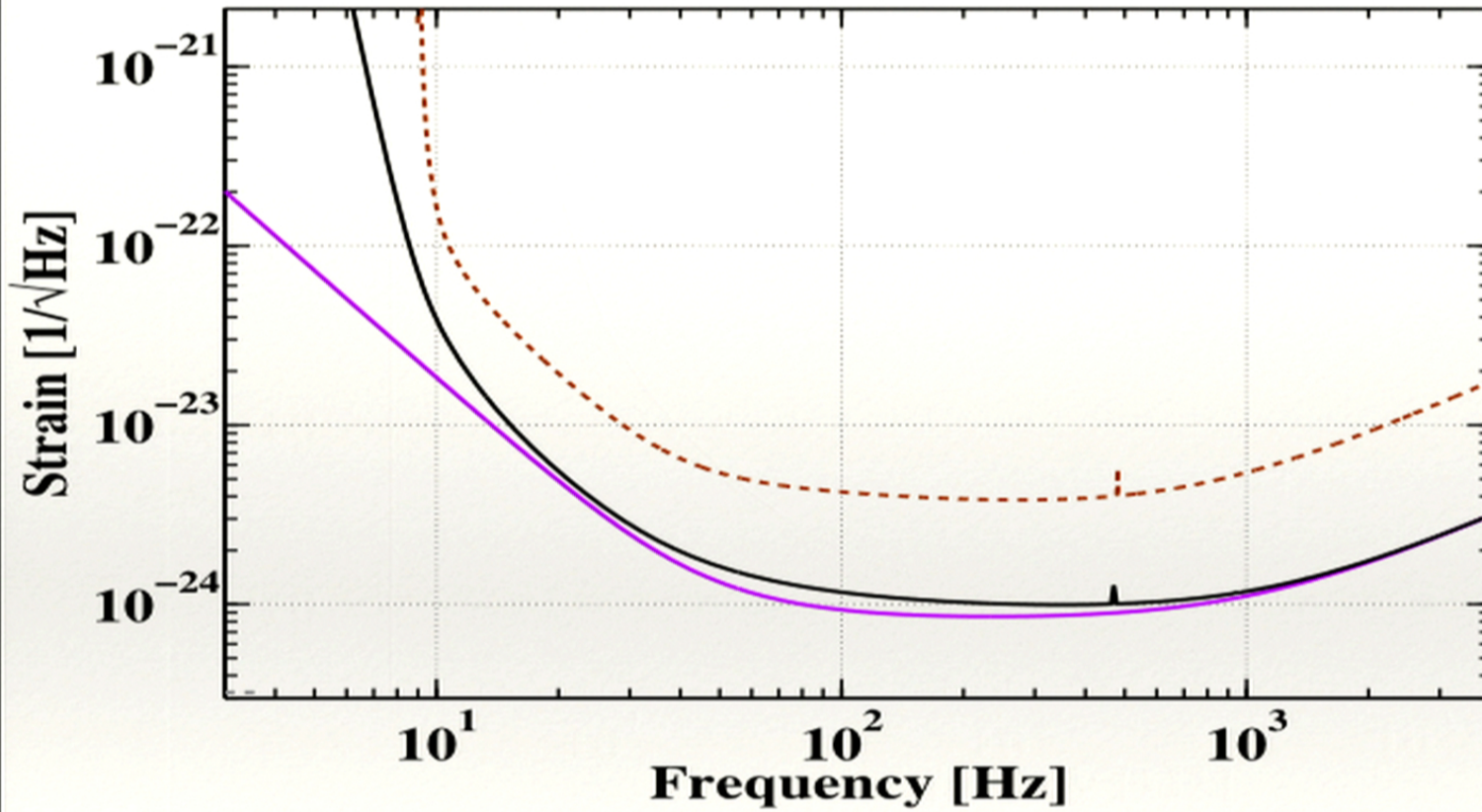
at SQL:
"Δx · Δp = ħ/2"

Upgrading Advanced LIGO Detectors



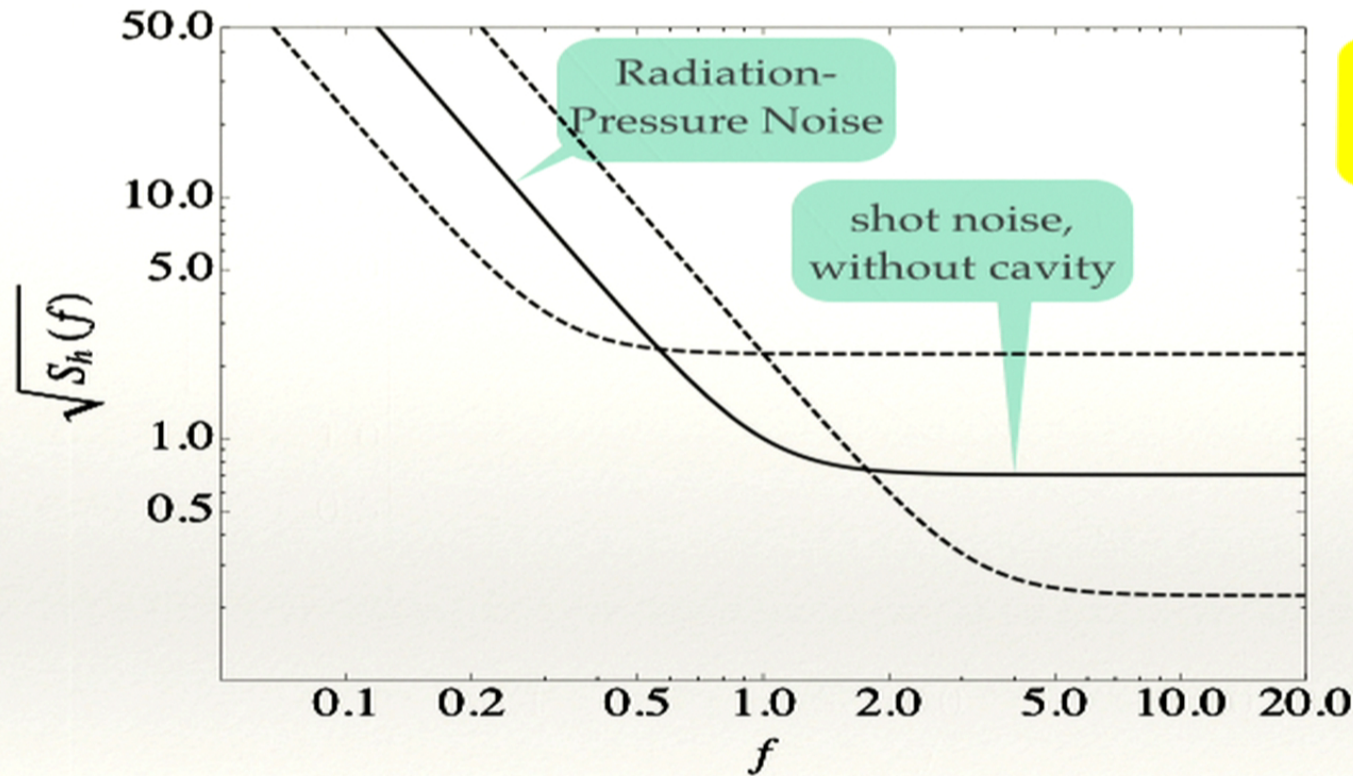
*figure from R. Adhikari et al.,
Instrumentation Science White Paper
LIGO-T1300433*

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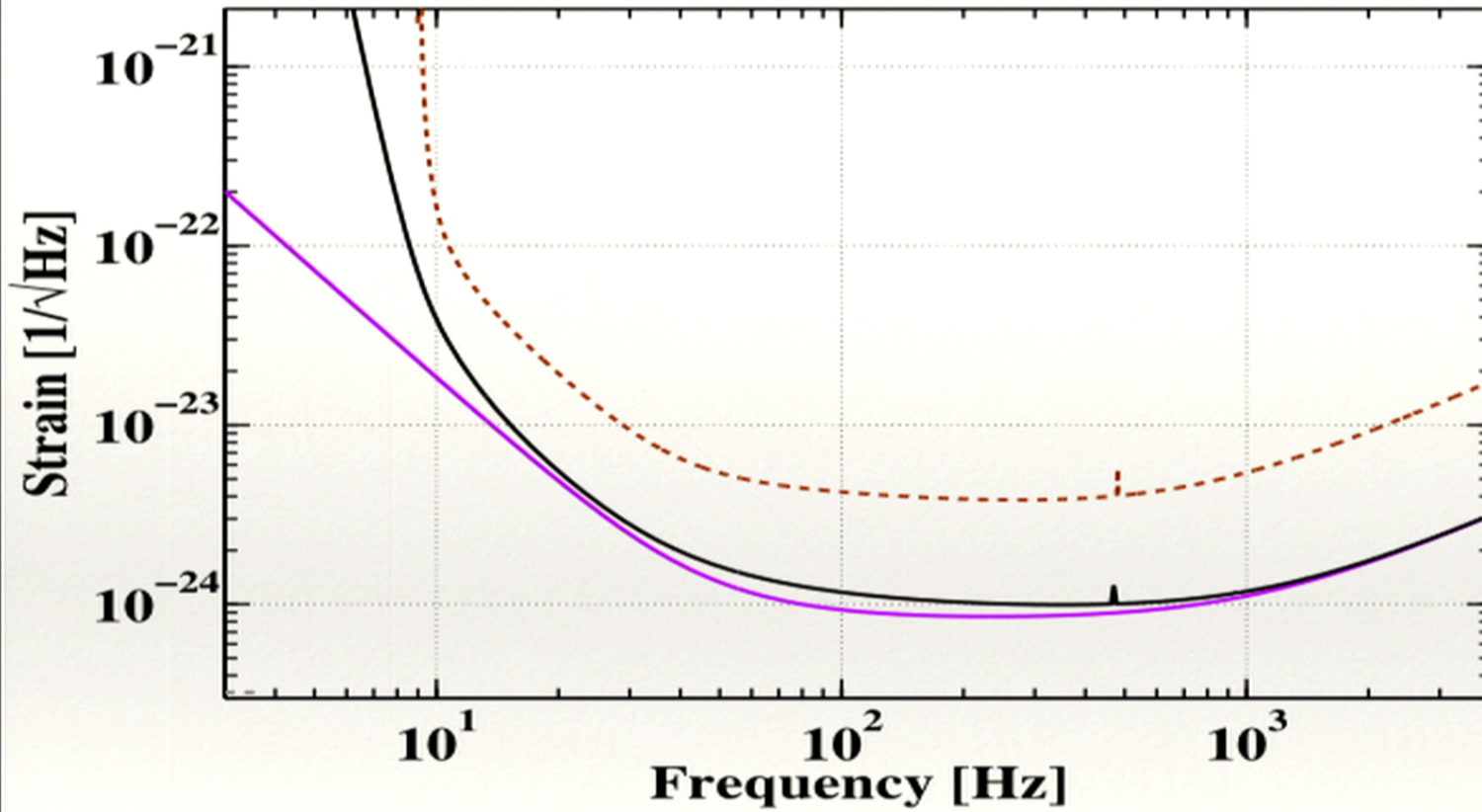


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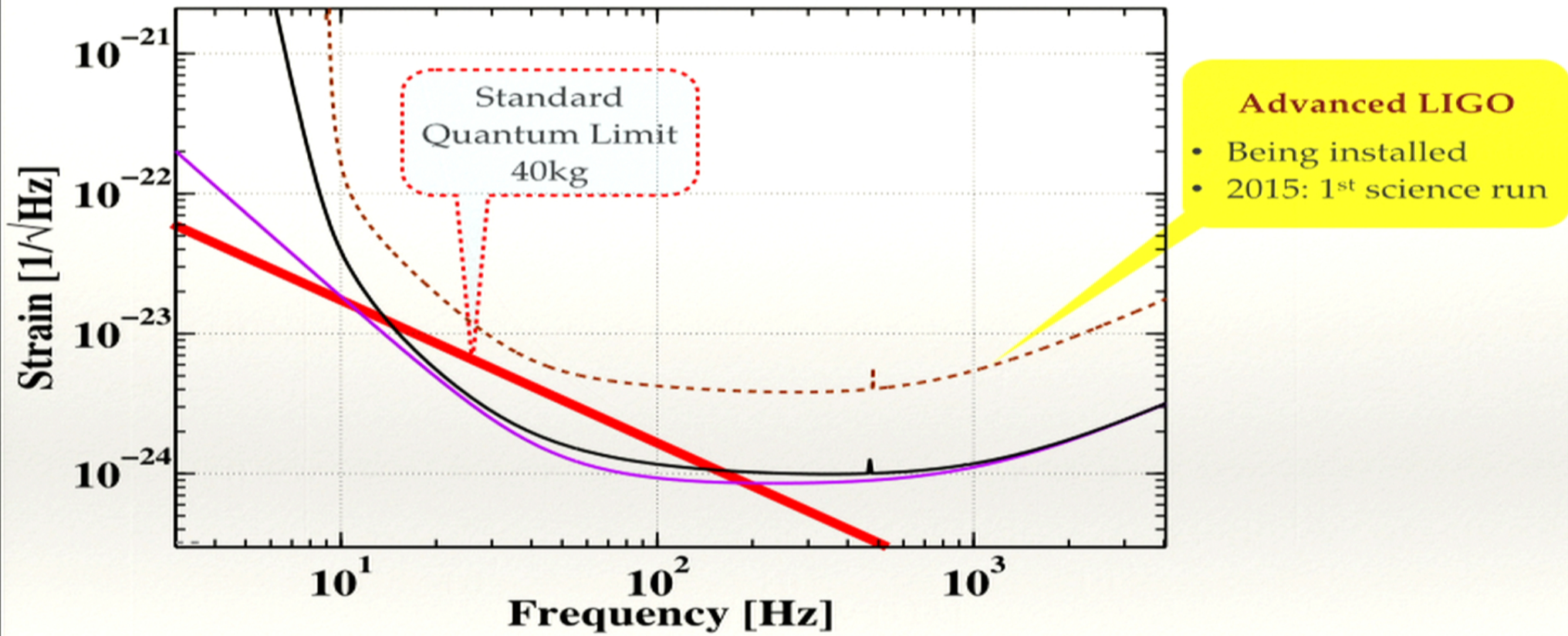
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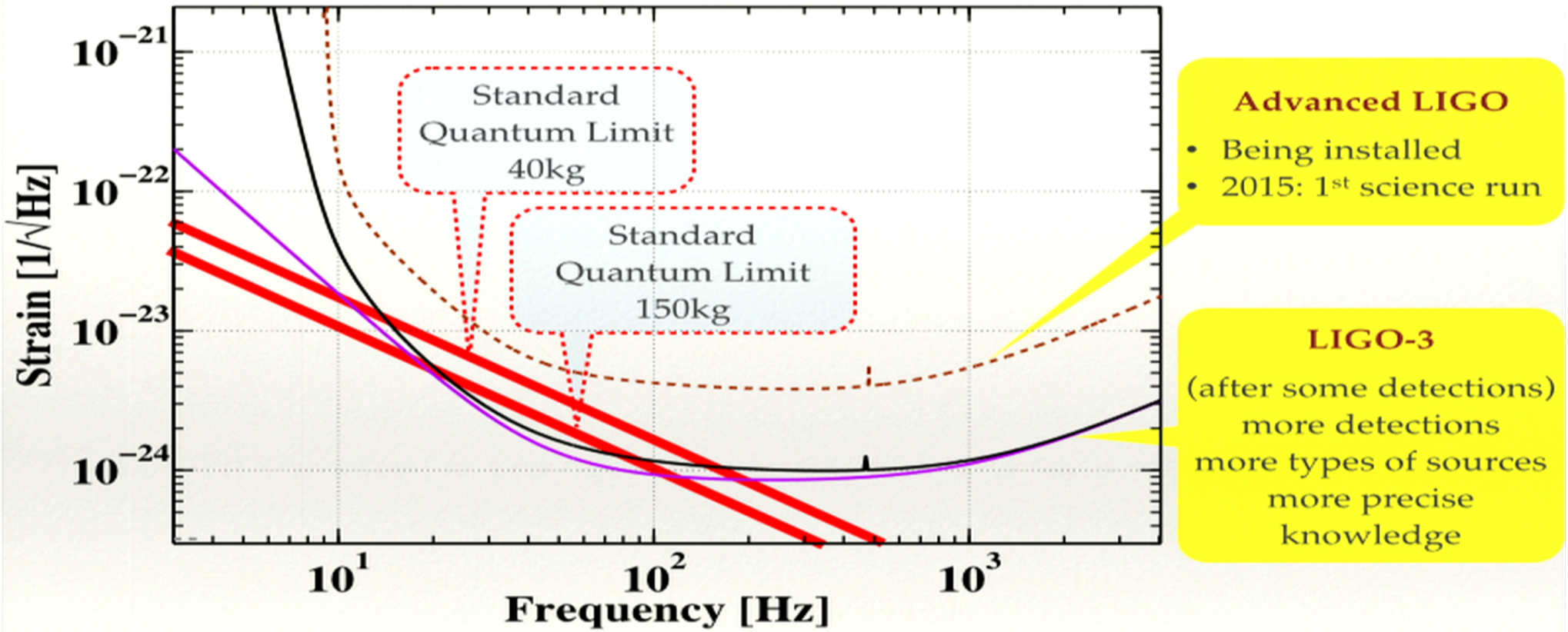
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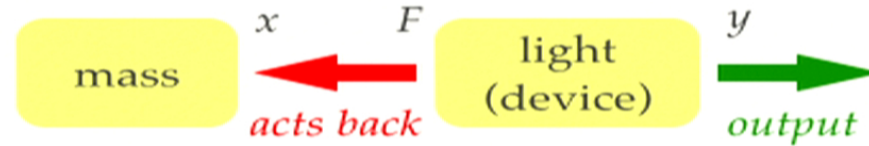


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Origin of the Standard Quantum Limit

test-mass device coupling

$$V_I = -\alpha x \cdot F$$



- Output field contains:

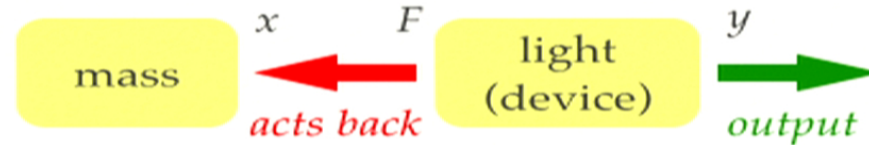
$$y = Z + \alpha \left[x_{\text{zero-point}} + x_{\text{signal}} - \alpha \frac{F}{M\Omega^2} \right]$$



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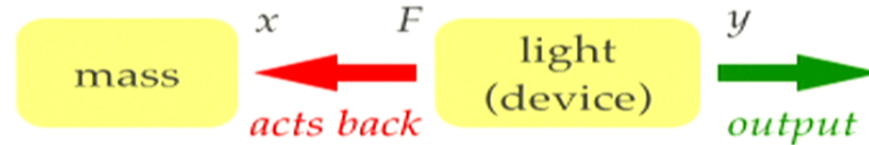
- Heisenberg Uncertainty Principle: $S_{ZZ}S_{FF} - |S_{ZF}|^2 \geq \hbar^2$

- Total noise spectrum: $S_{xx}^{sig} = \frac{S_{ZZ}}{\alpha^2} + \alpha^2 \frac{S_{FF}}{M^2\Omega^4} - 2\text{Re} \frac{S_{ZF}}{M\Omega^2} + S_{xx}^{\text{zero-point}}$

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sensing
noise

noise
from ZPF

signal

back-action
noise

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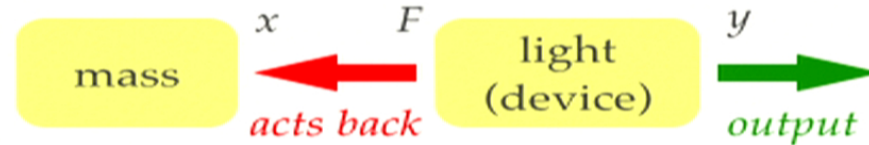
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$$\text{IF } S_{ZF}=0: \frac{S_{ZZ}}{\alpha^2} + \alpha^2 \frac{S_{FF}}{M^2\Omega^4} \geq 2\sqrt{\frac{S_{ZZ}S_{FF}}{M^2\Omega^4}} \geq \frac{2\hbar}{M\Omega^2} \equiv S_{xx}^{\text{SQL}}$$

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correlation between sensing and back-action noise

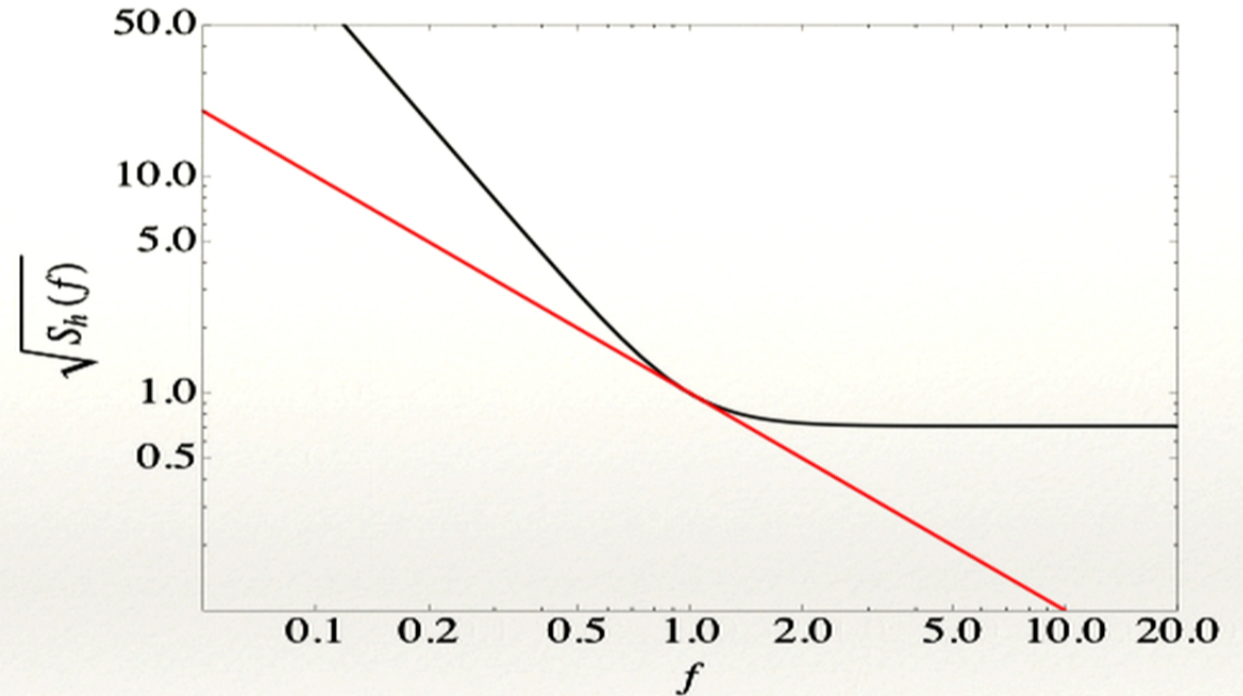
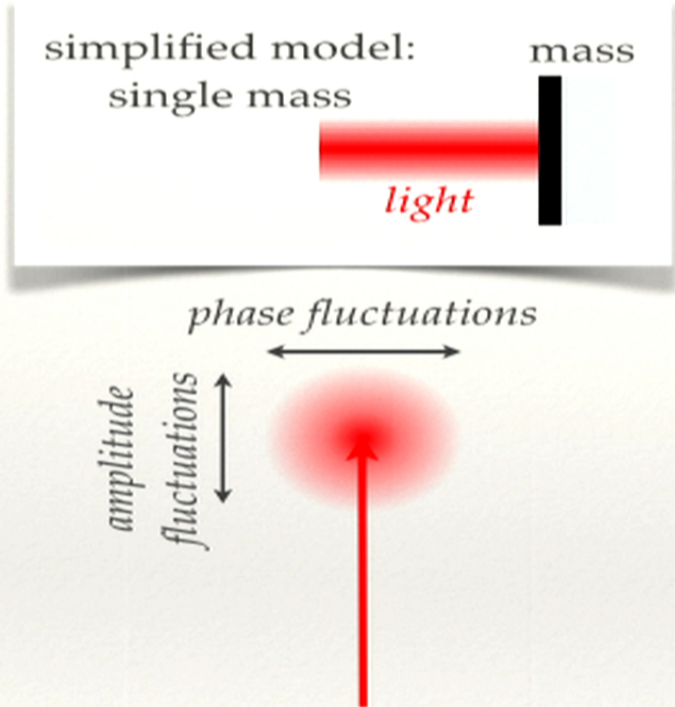
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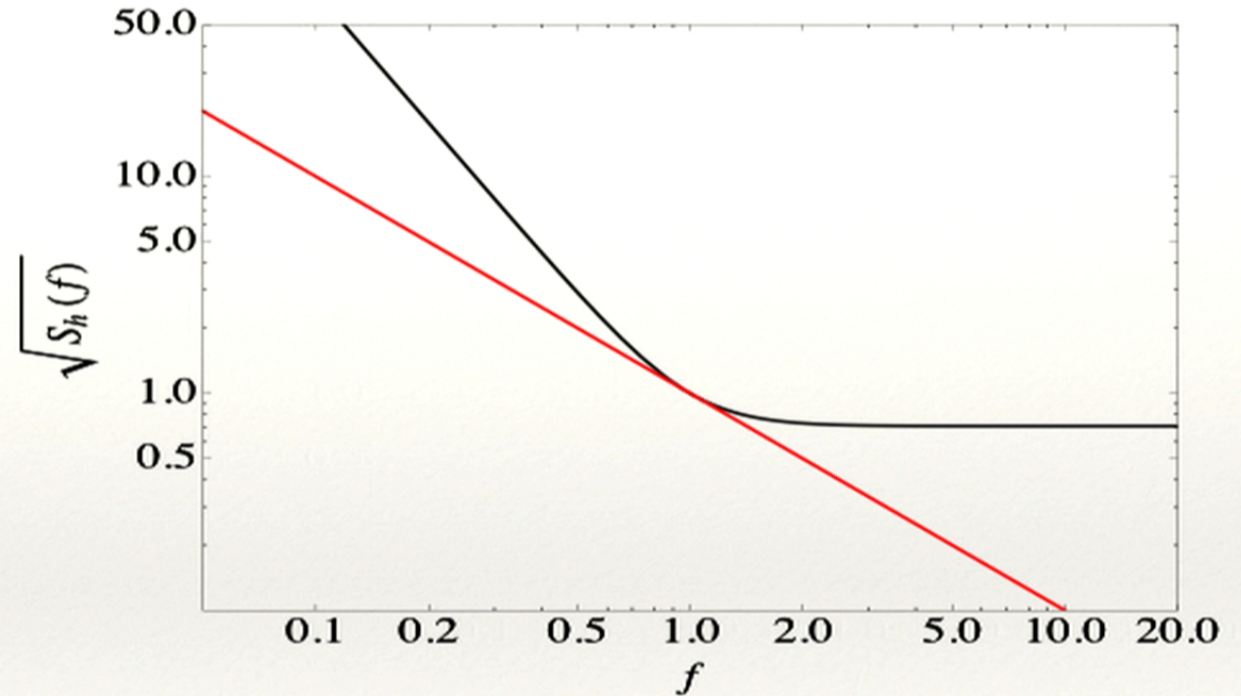
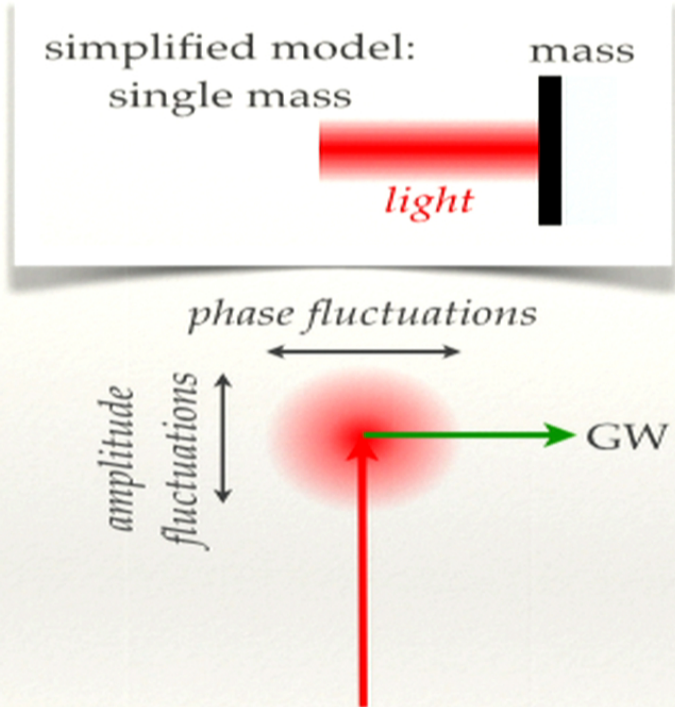
Building the Necessary Correlation

[Caves, Unruh, 1990s; H.J. Kimble et al., 2001]



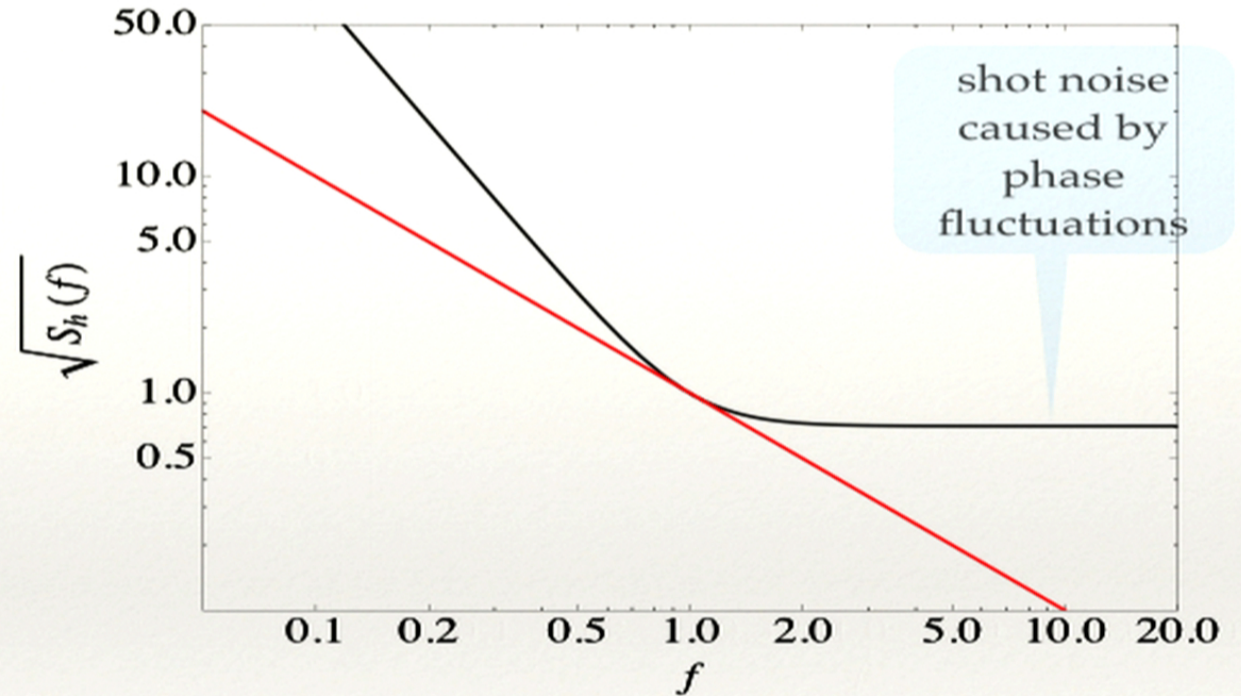
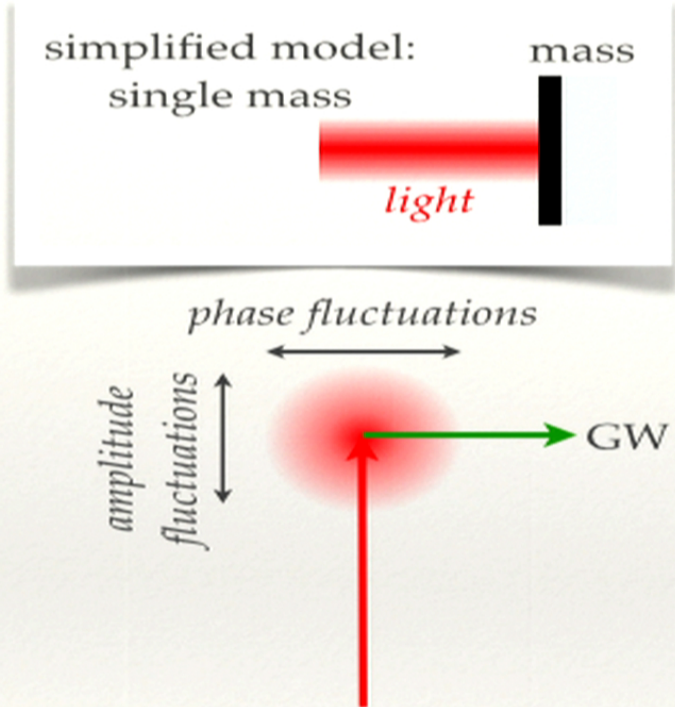
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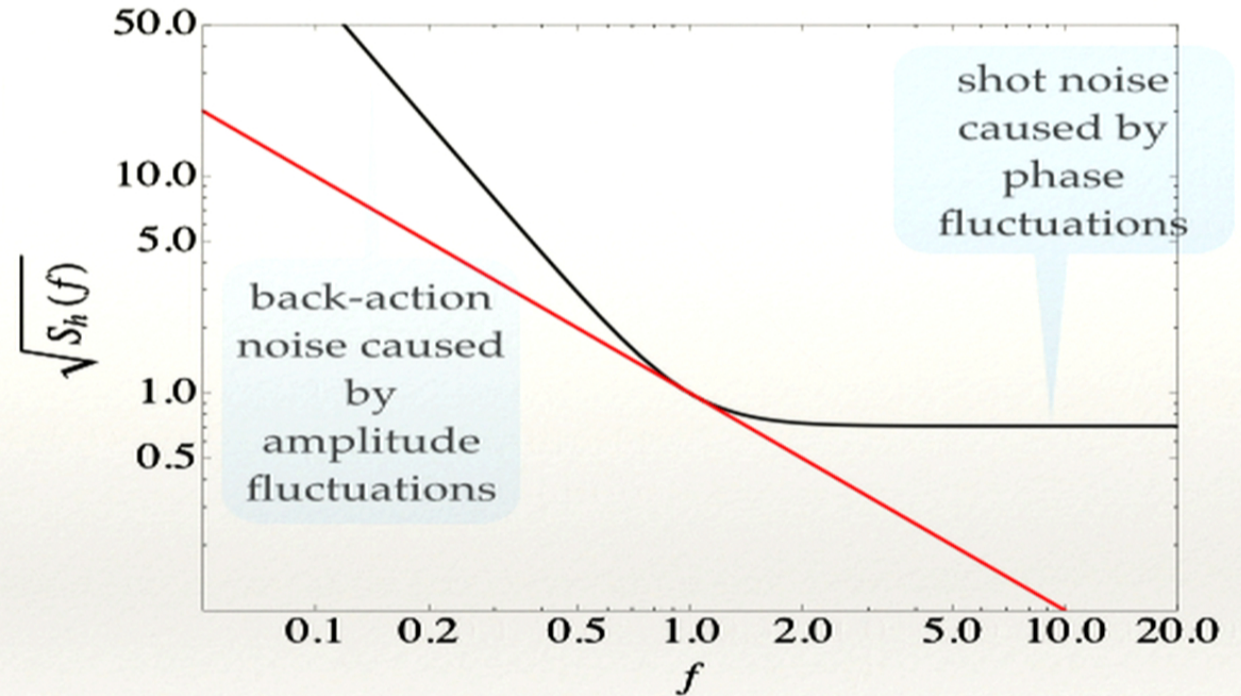
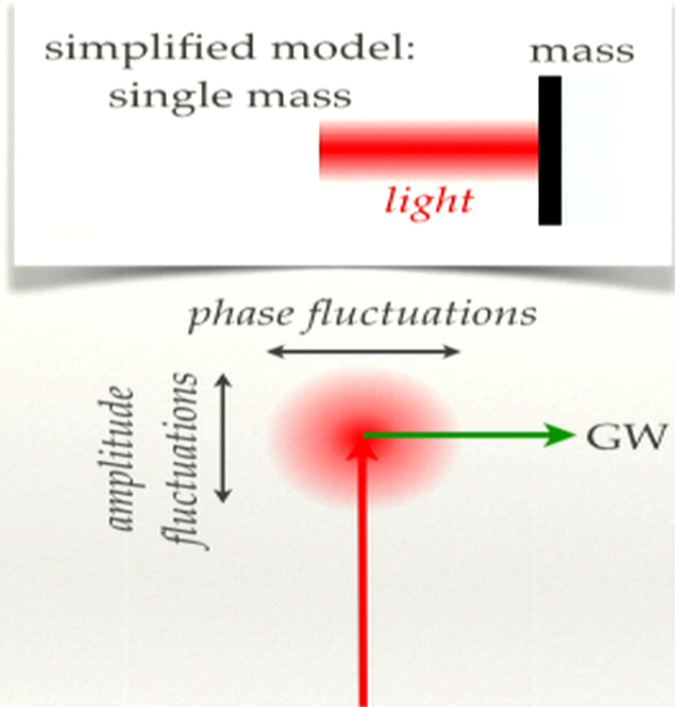
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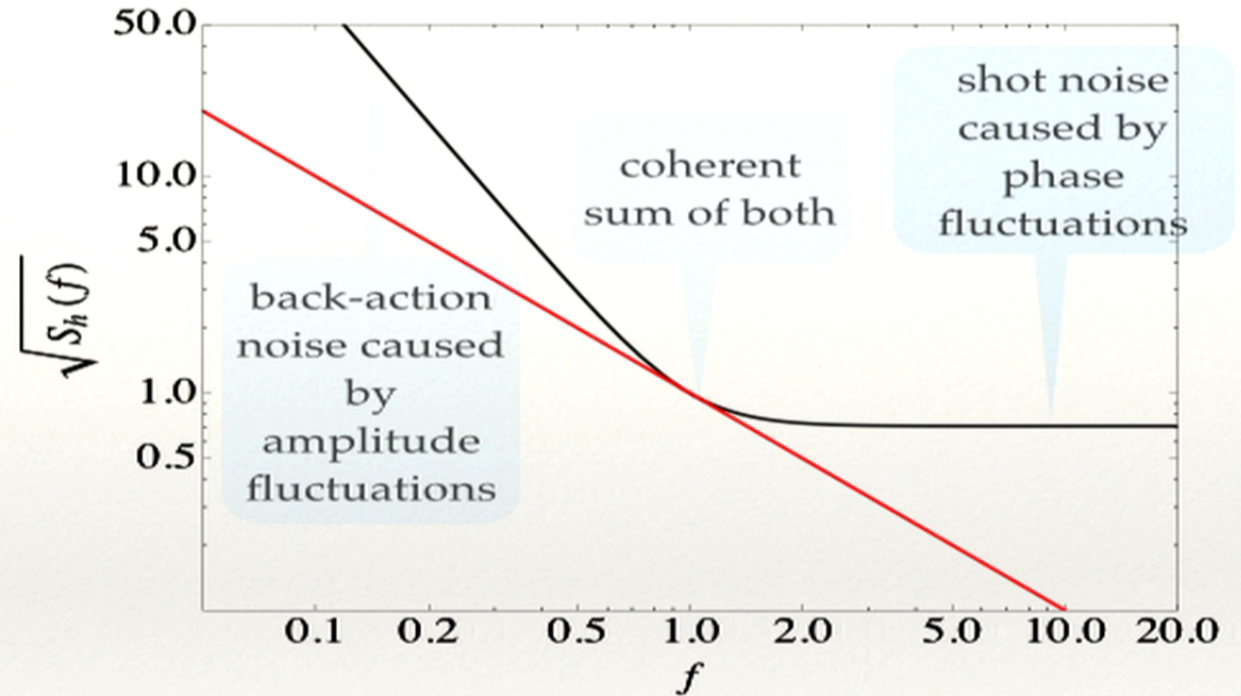
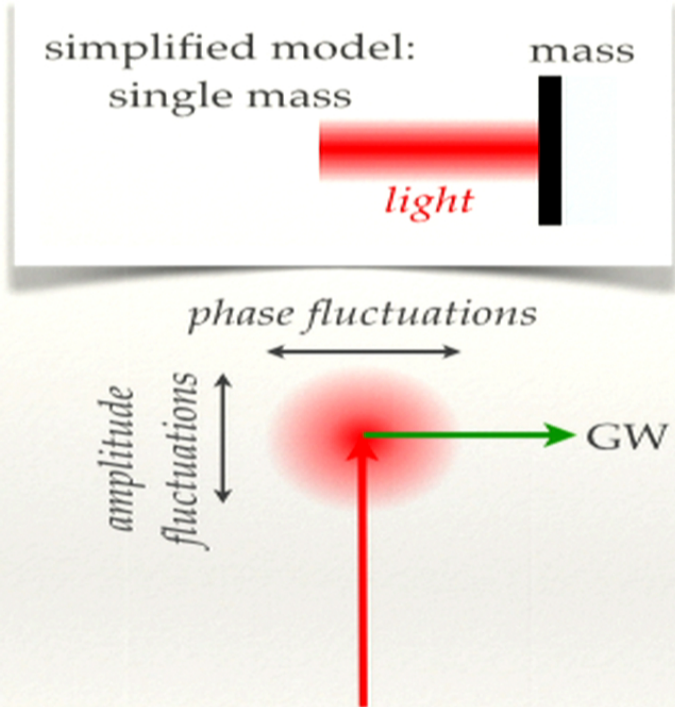
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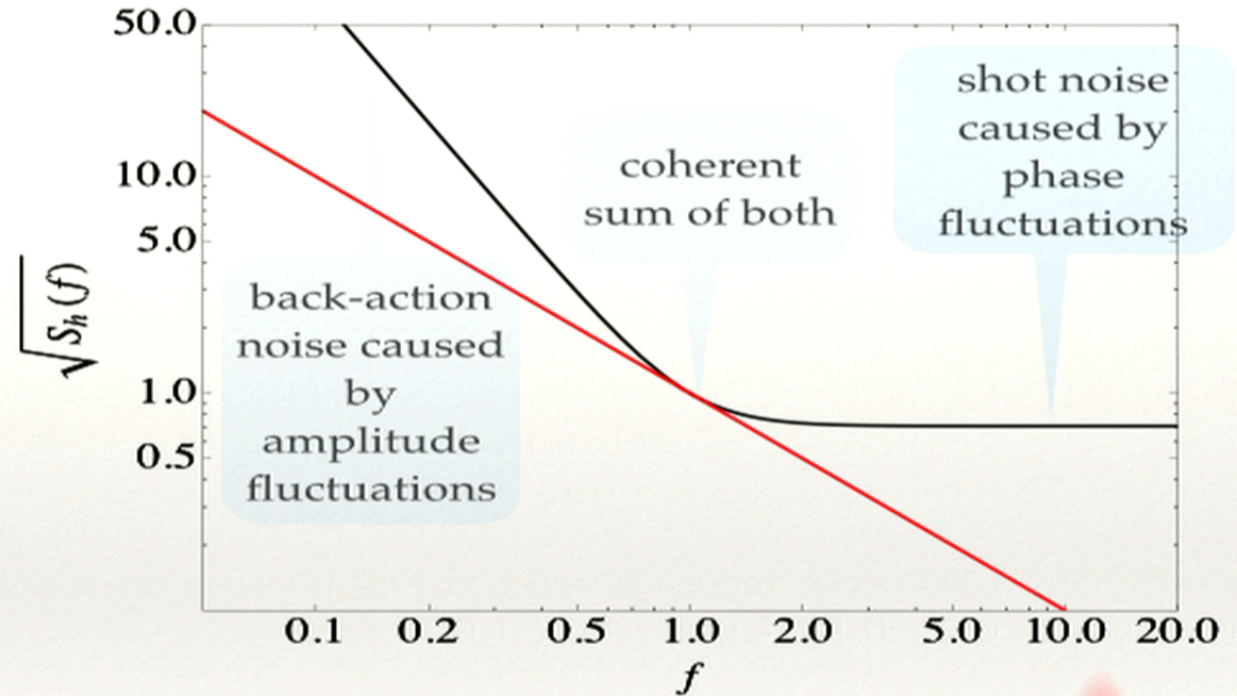
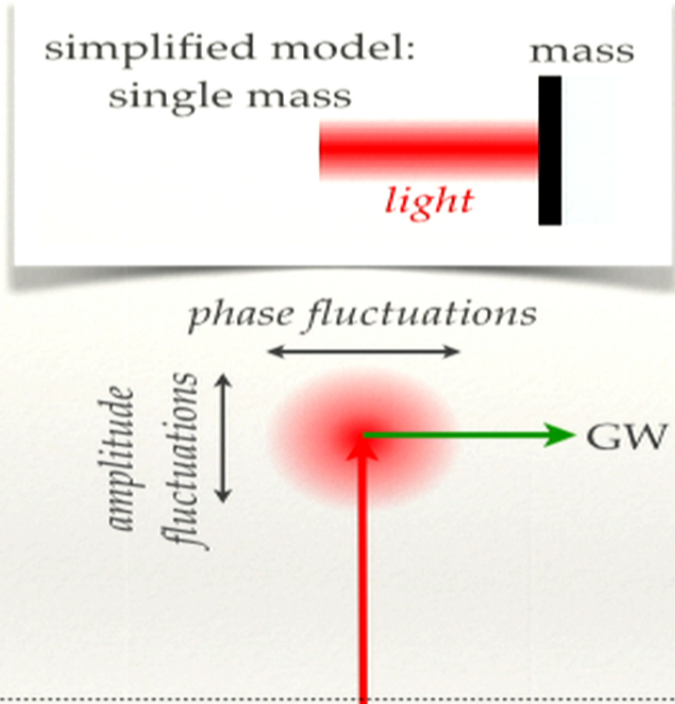
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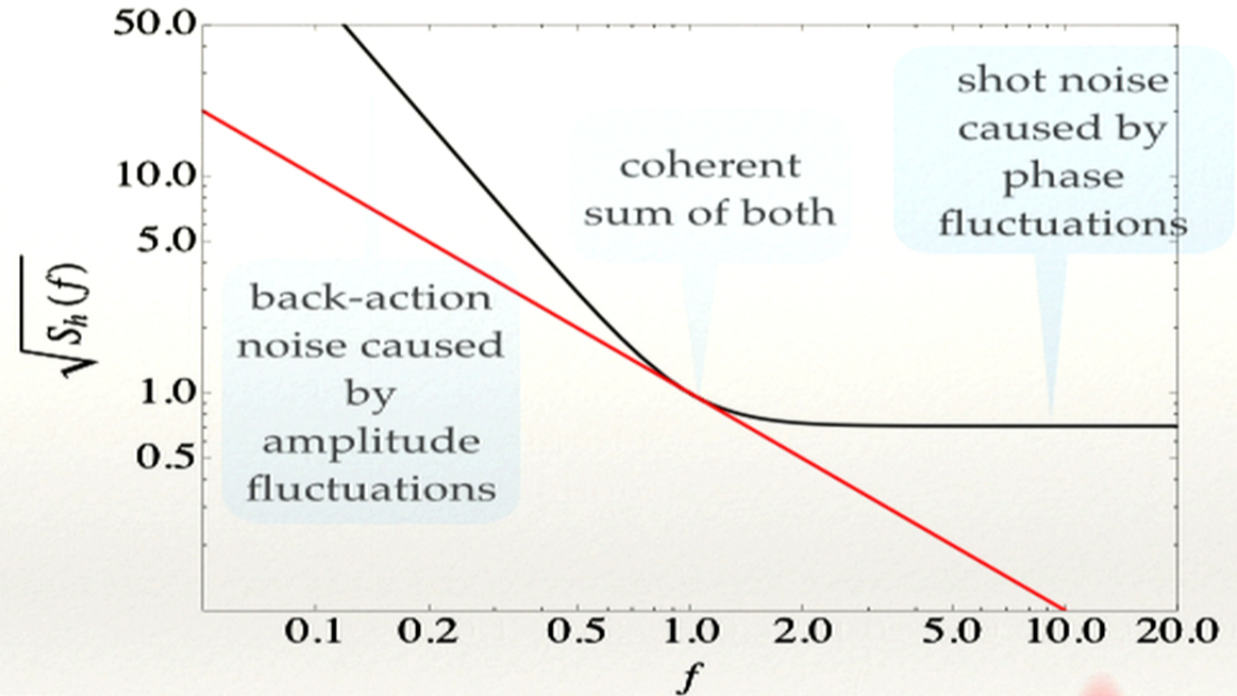
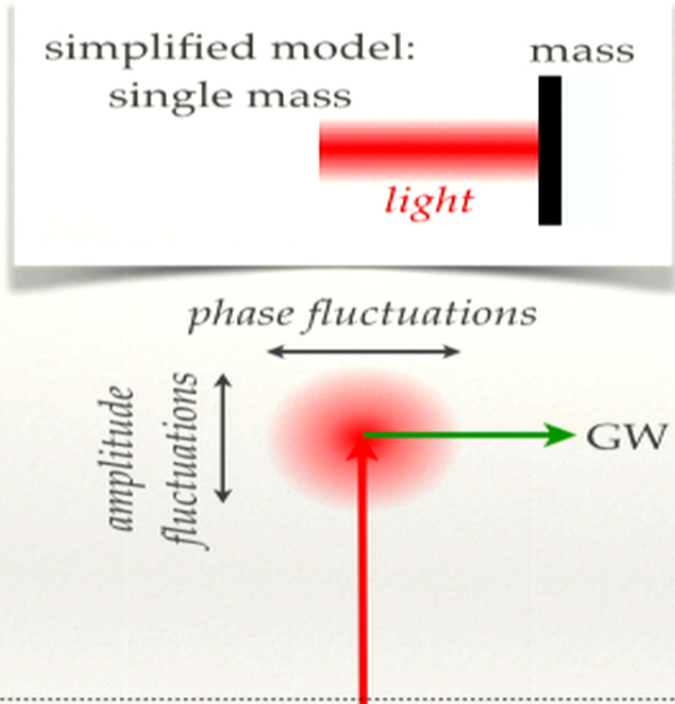
EM field can be prepared into "squeezed states"

Wu, Xiao & Kimble, 1987
R. Schnabel, N. Mavalvala, D.E. McClelland & P. K. Lam, 2010

phase squeezed

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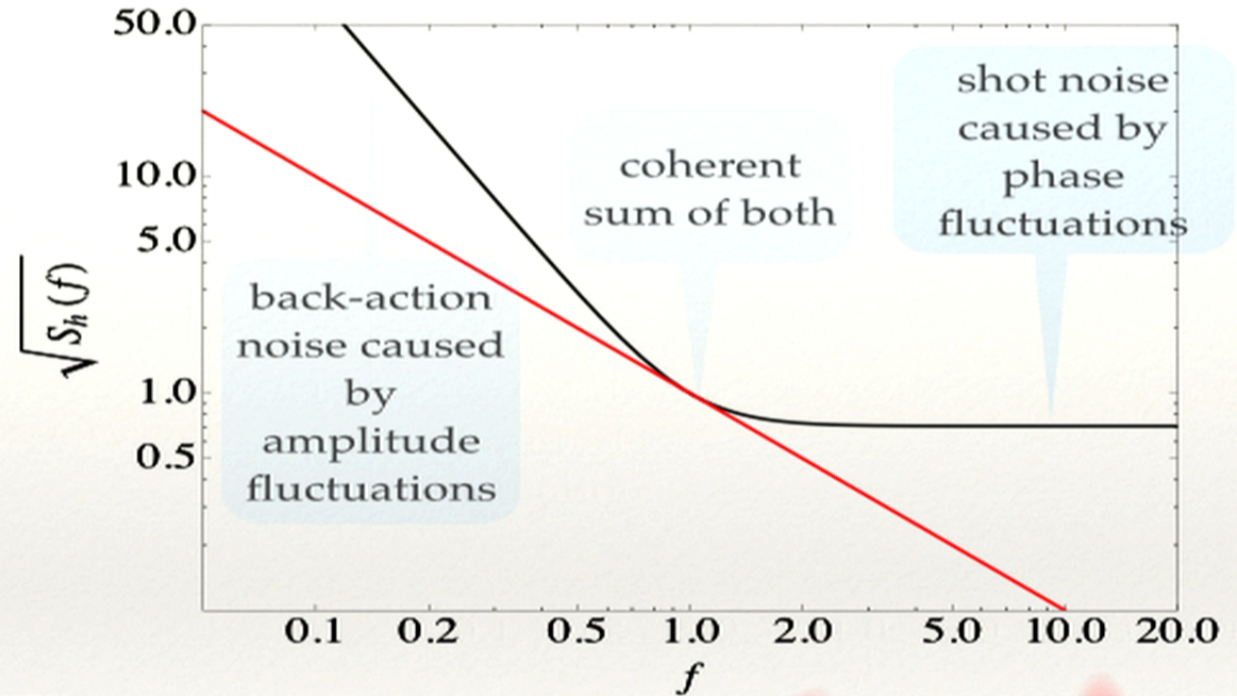
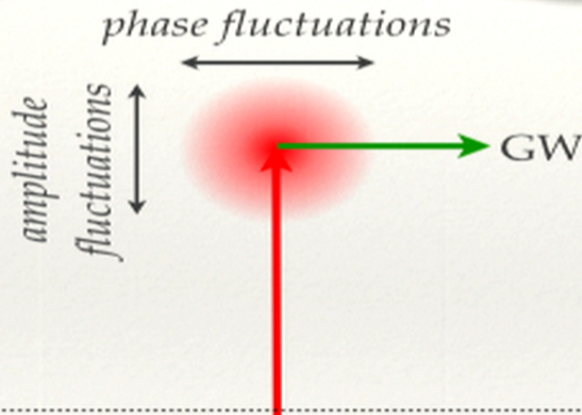
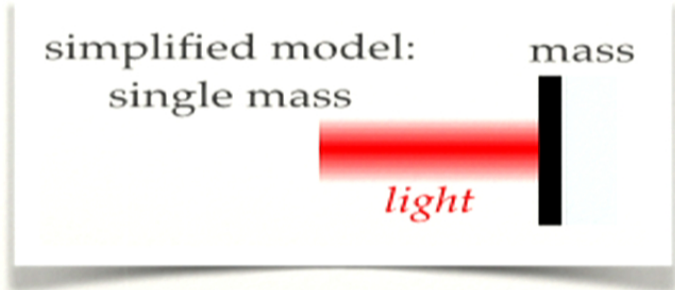


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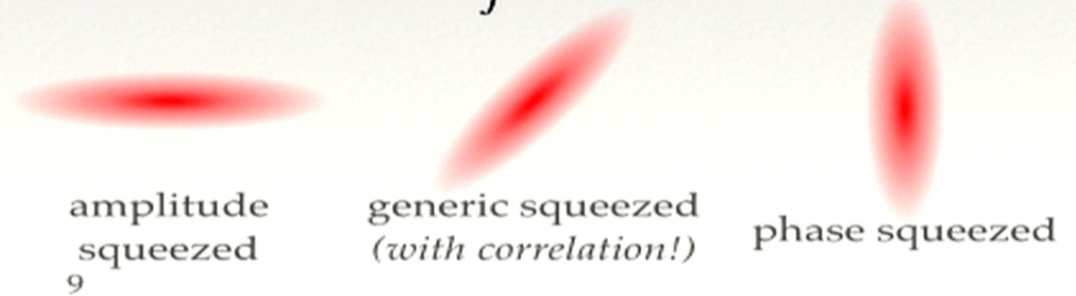
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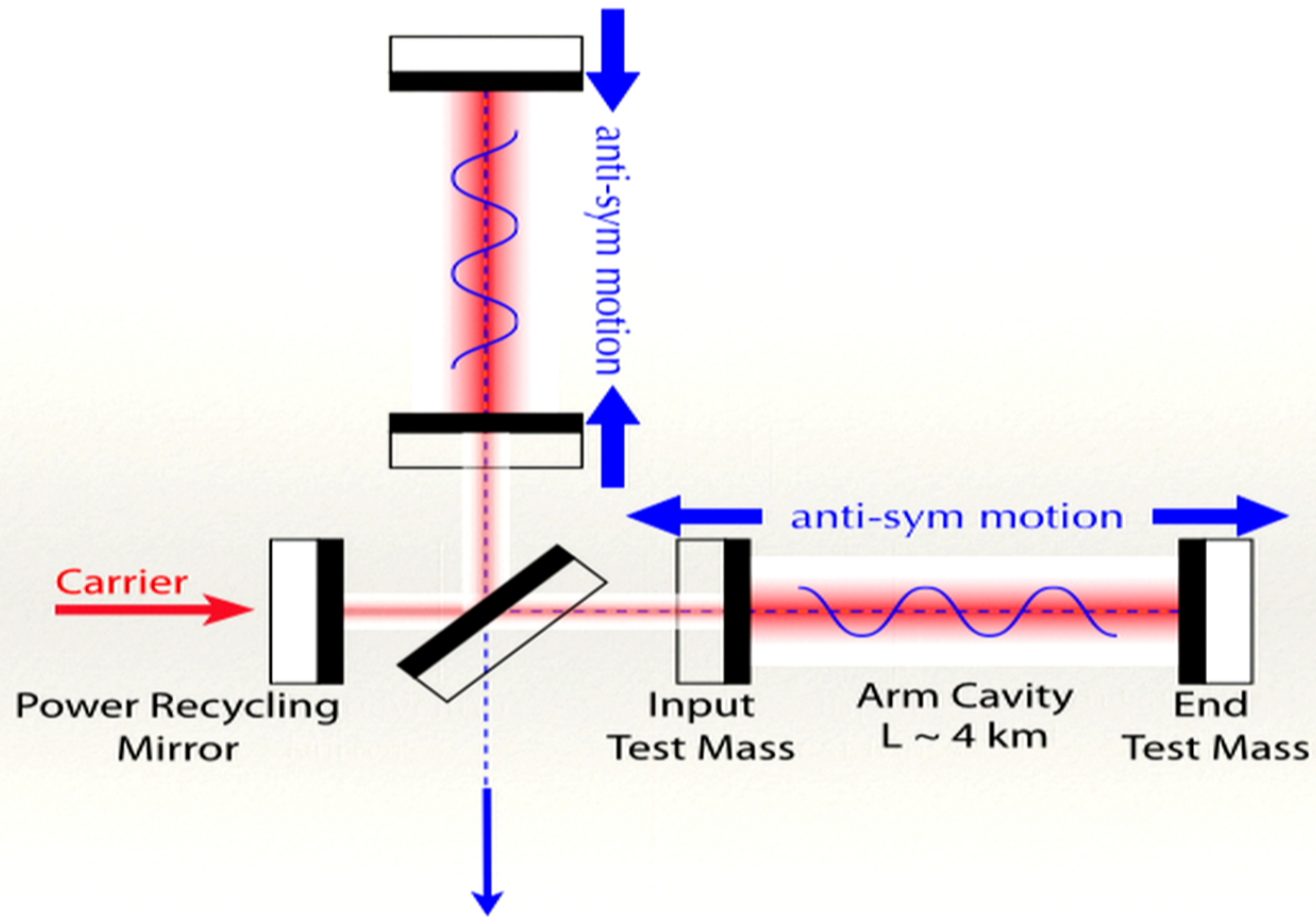


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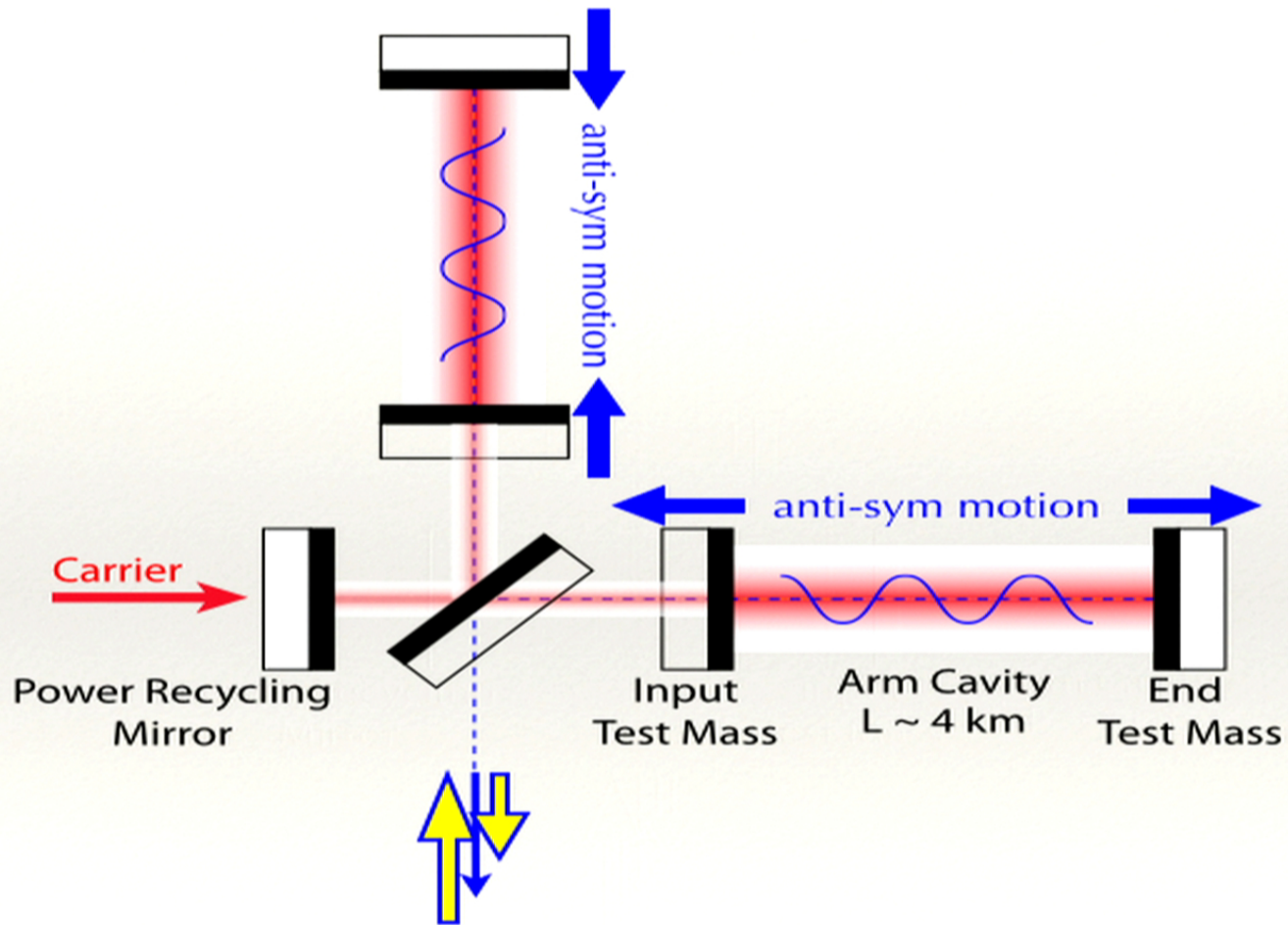


Injection of Squeezed Vacuum

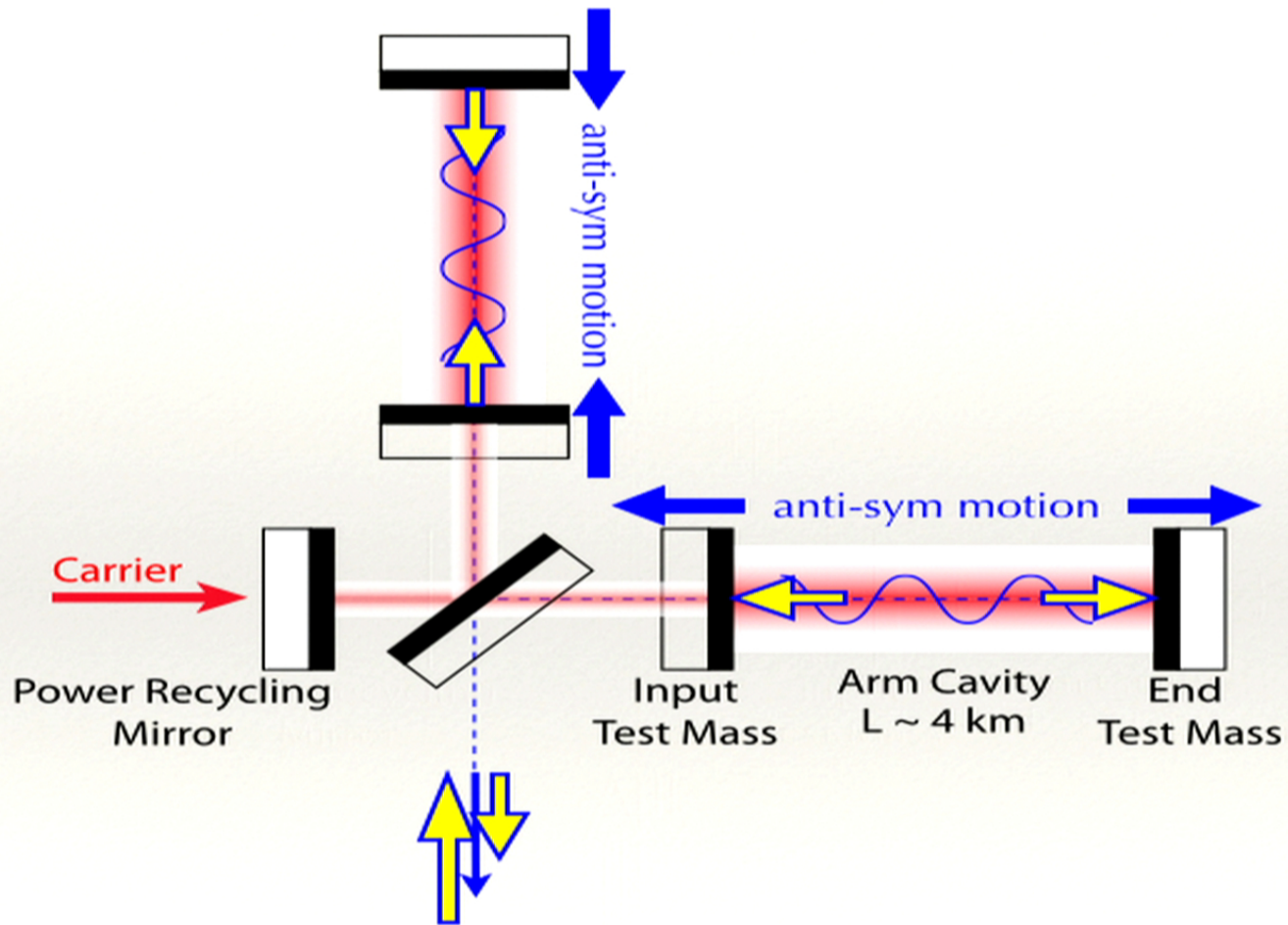


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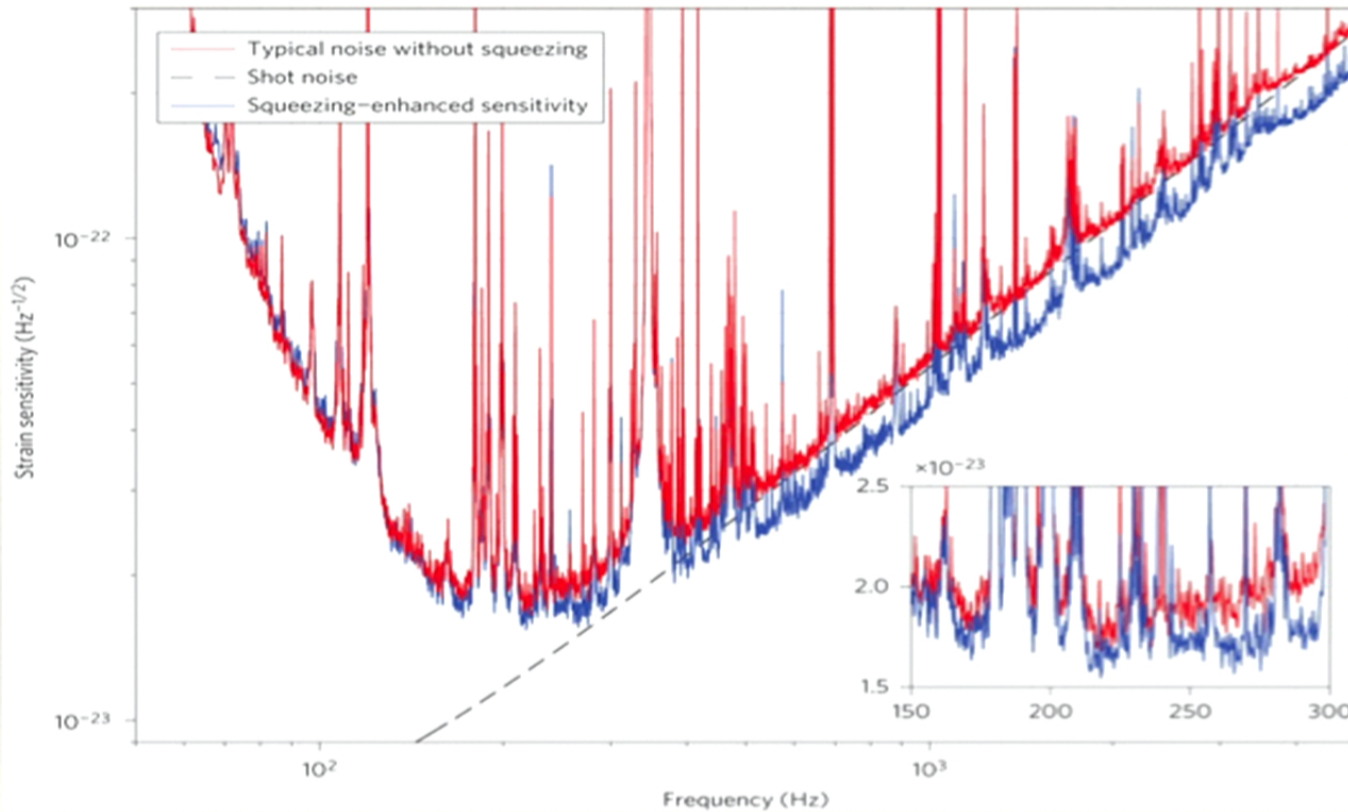


Injection of Squeezed Vacuum



10

Squeezing in LIGO



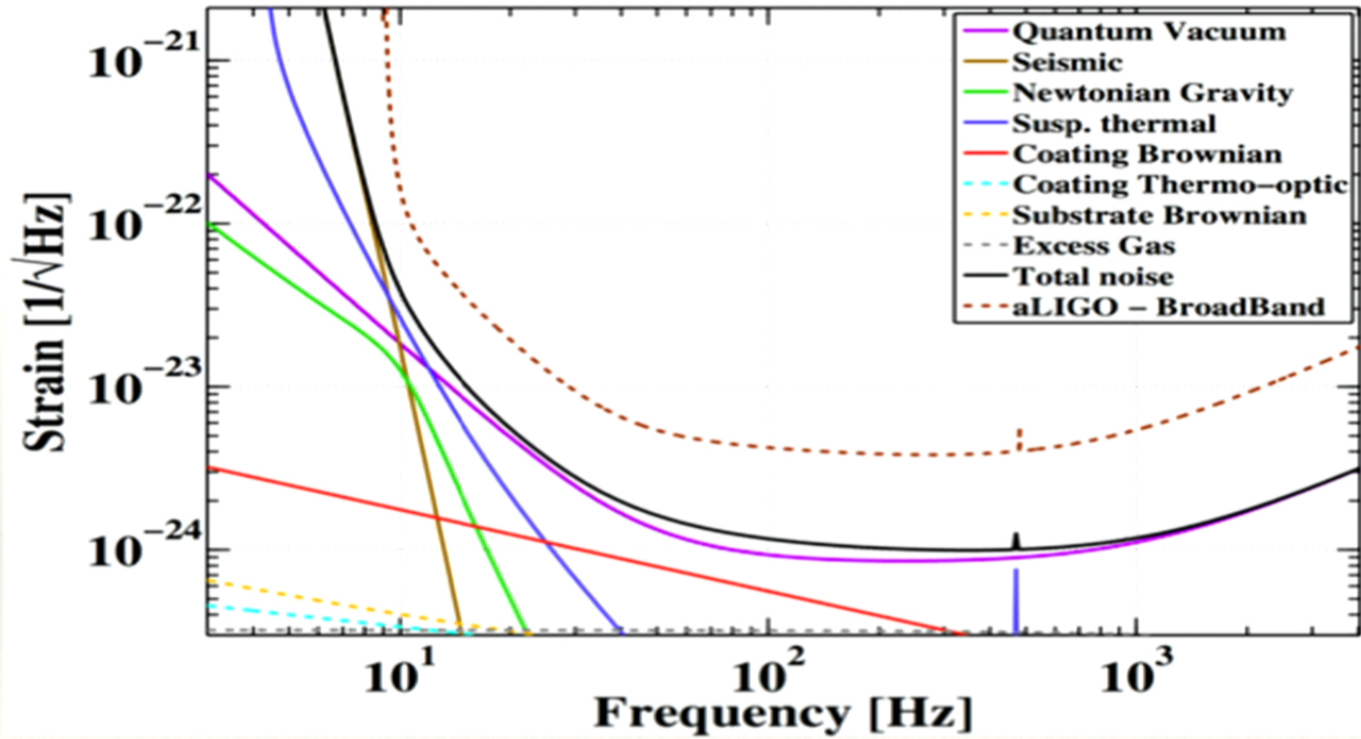
limited by optical losses

- initial LIGO
 - ~2dB squeezing
- Advanced LIGO
 - permits ~6 dB
- LIGO-3
 - frequency-dependent
 - requires “filter cavities”

Phase Squeezing in LIGO

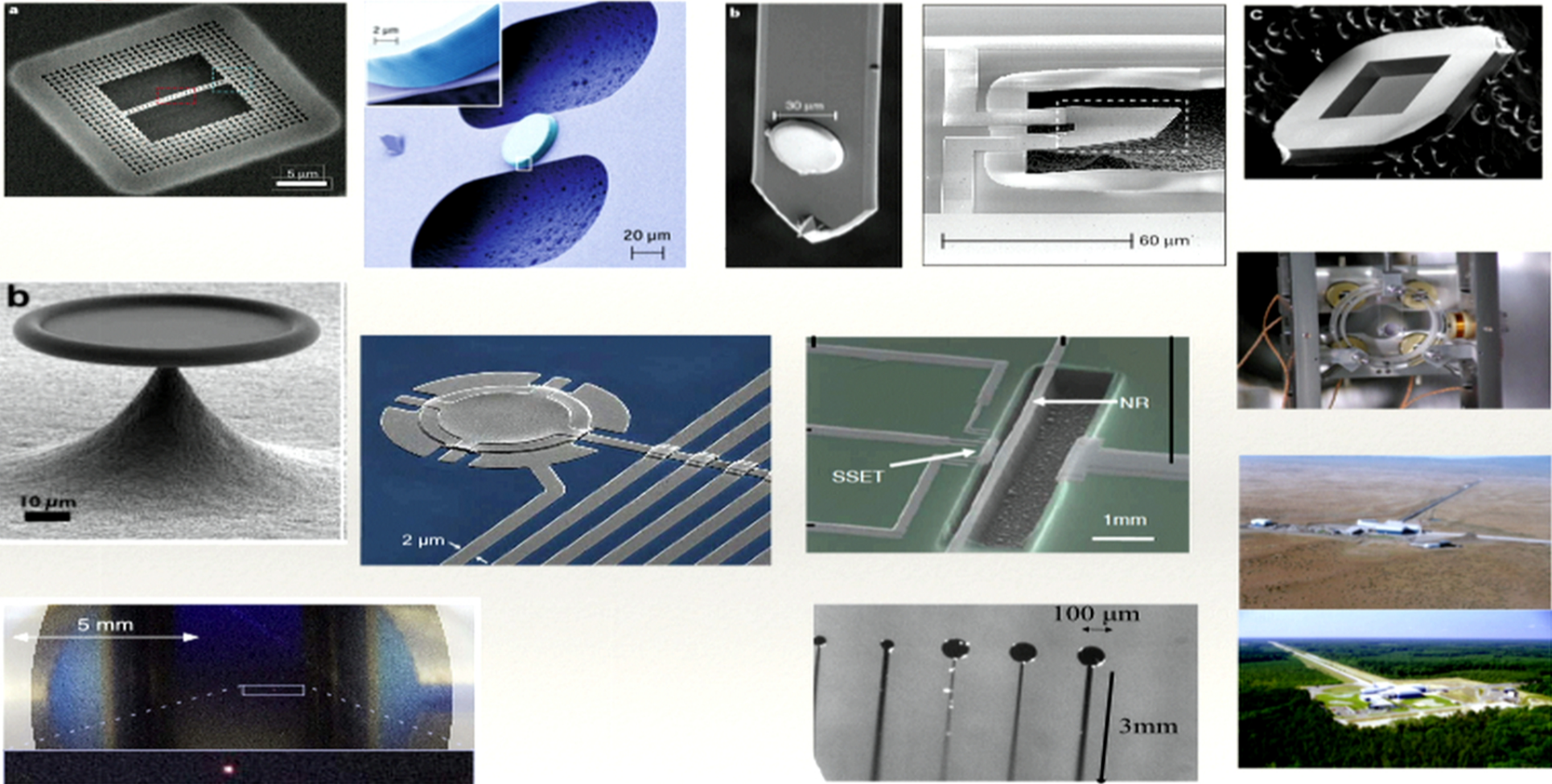
[LIGO Scientific Collaboration, 2013]

Other sources of noise ...



R&D in LIGO makes test masses "more and more quantum"

Optomechanics/Electromechanics at Different Scales



Macroscopic Quantum Mechanics

Demonstration of (Macroscopic) Quantum Mechanics

- ❖ Prepare & verify quantum states (e.g., entanglement) of macroscopic objects
- ❖ Prepare quantum states without classical counterparts
- ❖ Quantum state of macroscopic objects can be reconstructed through tomography.

[review articles by Aspelmeyer, Kippenberg and Marquardt, 2013; Meystre, 2013; Y Chen, 2013]

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“Testing Quantum Mechanics”?

Testing Macroscopic Quantum Mechanics

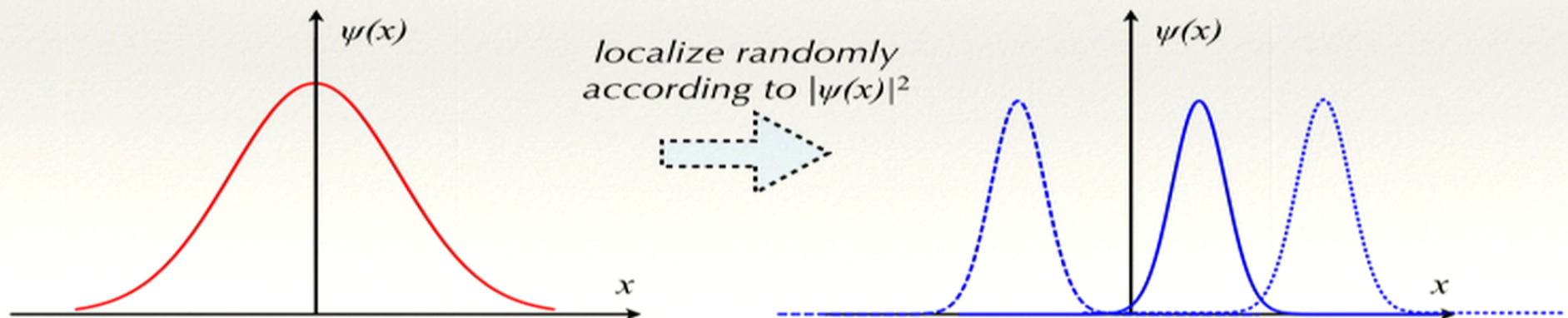
- ❖ Does standard quantum mechanics apply to a macroscopic test mass? (*Why not?*)

Testing Macroscopic Quantum Mechanics

- ❖ Does standard quantum mechanics apply to a macroscopic test mass? (*Why not?*)
- ❖ **Motivation 1: “Intrinsic Randomness of Quantum Mechanics”**
 - ❖ Macroscopic world is classical, while microscopic world is quantum.
 - ❖ “Quantum Measurement” is necessary to convert from quantum to classical.
 - ❖ “Quantum Measurement gives random answers.
- ❖ **“Collapse Models”**
 - ❖ Conversion from quantum to classical happens spontaneously, but in an unknown manner.
 - ❖ A. Bassi et al., Rev. Mod. Phys., 2013; O. Romero-Isart, PRA 2011; O. Romero-Isart et al., PRL 2011.

Testing Macroscopic Quantum Mechanics

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Collapse Models

- ❖ Collapse in x will cause disturbance in p .
- ❖ Diffusion in phase space, same as a *stochastic force*.
- ❖ Example: "CSL" models



*each atom feels a random force due to the collapses
the total force scales with mass*

Nimmrichter, Hornberger and Hammerer (2014)

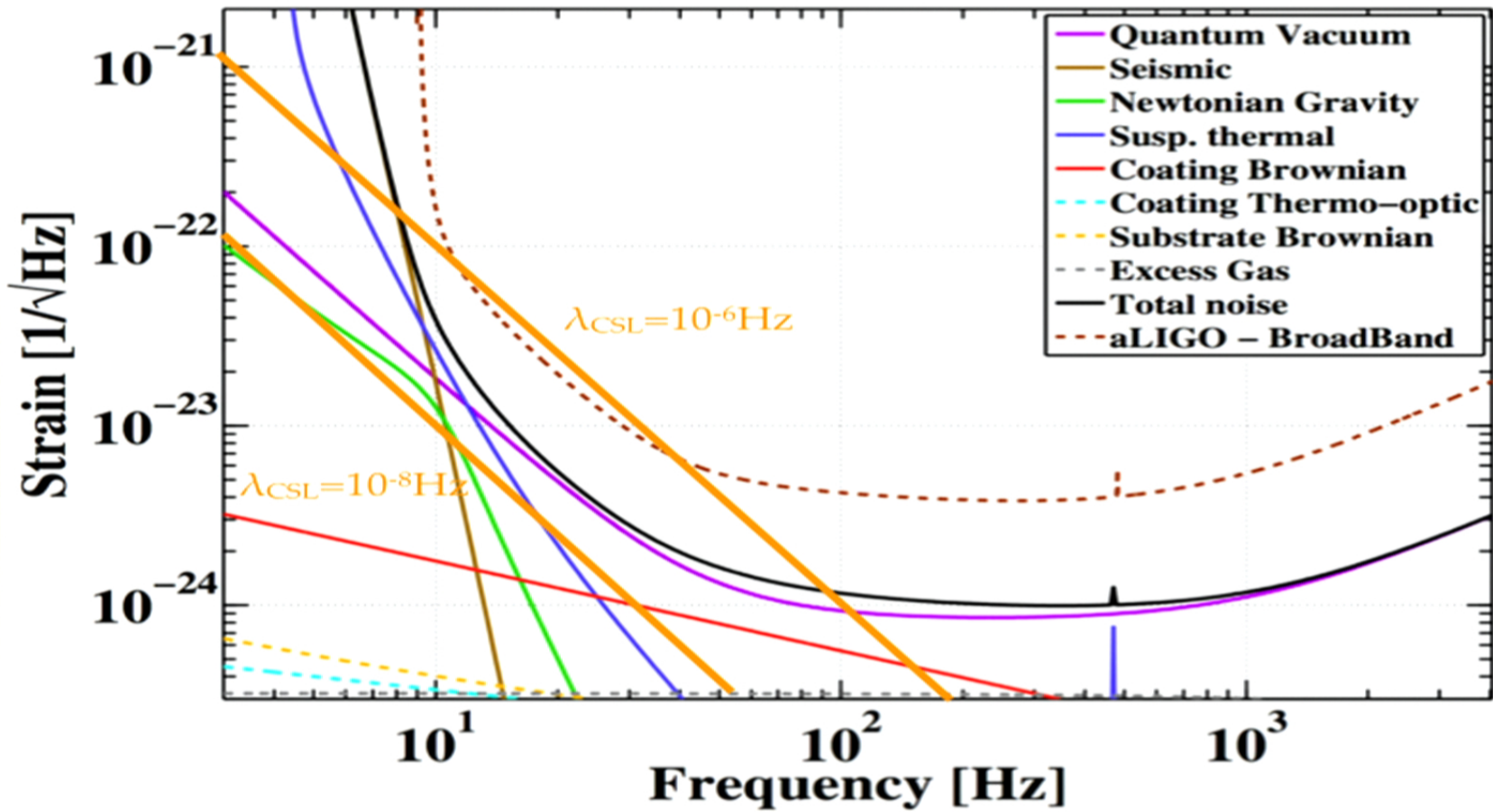
$$S_f^{\text{CSL}} \propto m^{2/3}, \quad S_f^{\text{Gravity}} \propto m$$

$$S_f^{\text{SQL}} = 2\hbar m \omega^2$$

$$S_f^{\text{SQL}} / S_f^{\text{CSL}} \propto m^{1/3} \omega^2, \quad S_f^{\text{SQL}} / S_f^{\text{Gravity}} \propto \omega^2.$$

This means low-frequency oscillators are more promising in testing these models.

Collapse Models with LIGO



Collapse Models

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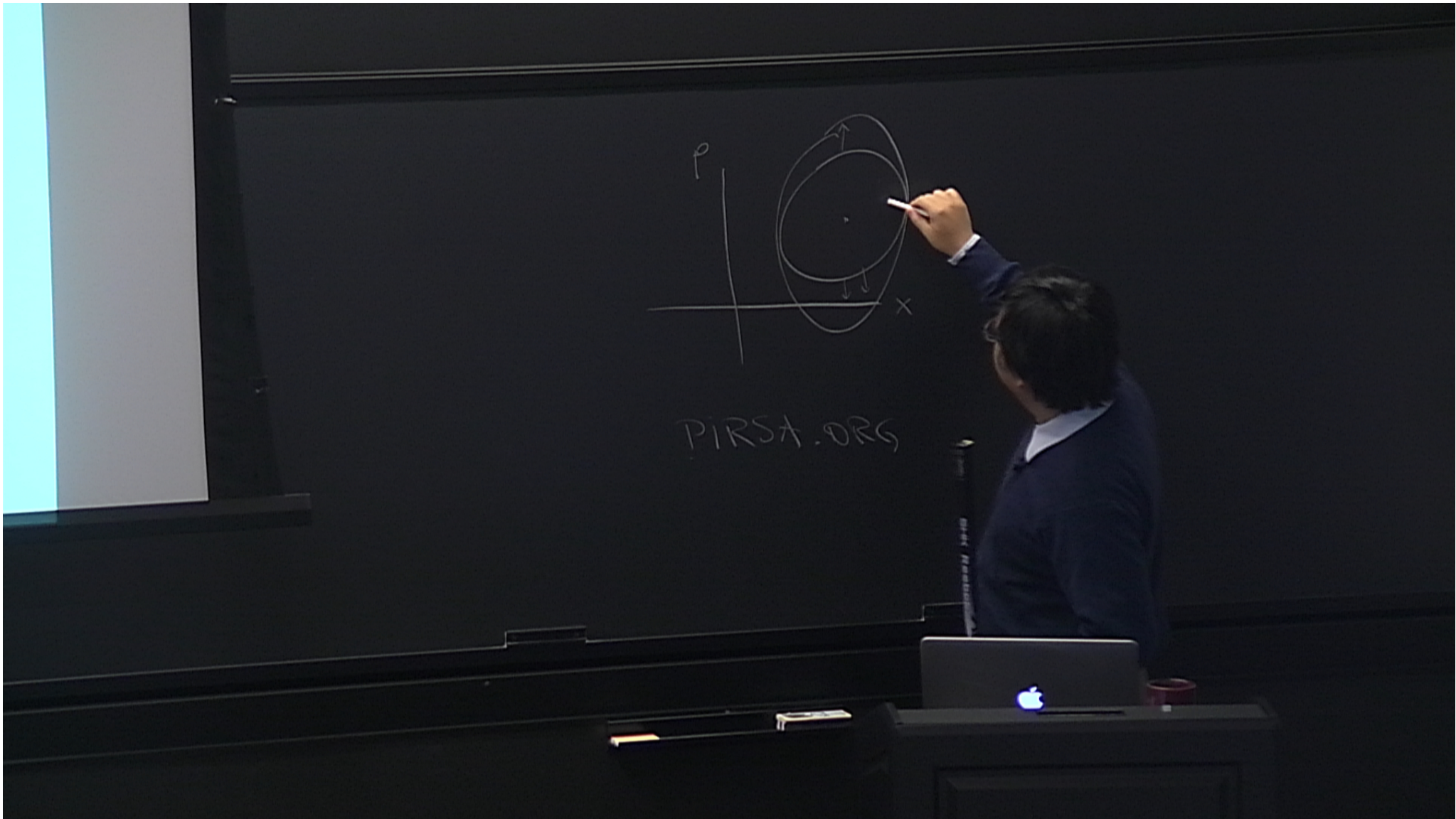
Testing Macroscopic Quantum Mechanics

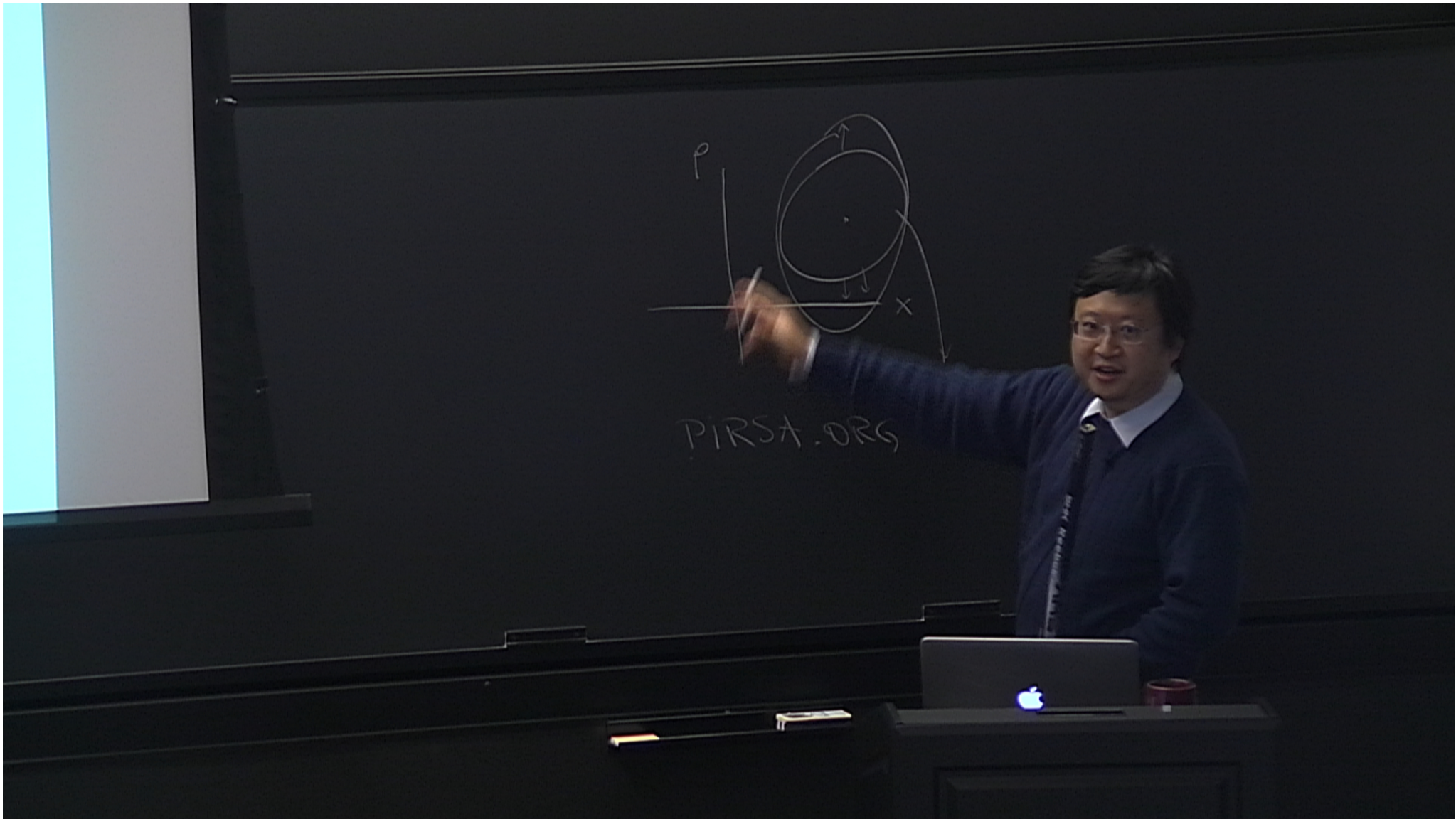
- ❖ Does standard quantum mechanics apply to a macroscopic test mass? (*Why not?*)
- ❖ **Motivation 2: how to confirm that gravity is quantum.**



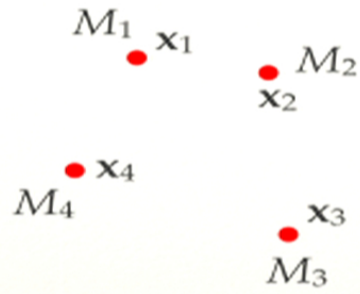
If you believe in quantum mechanics up to any level then you have to believe in gravitational quantization in order to describe this experiment."

R.P. Feynman, 1957





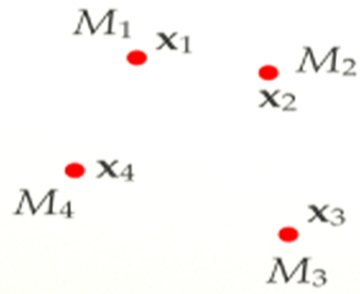
Physics of “Self-Gravitating” Quantum Systems



$$\hat{\phi}(\mathbf{x}) = - \sum G \frac{M_j}{|\mathbf{x} - \hat{\mathbf{x}}_j|}$$
$$\hat{V} = \sum_i -\frac{1}{2} M_j \hat{\phi}(\hat{\mathbf{x}}_j) = - \sum_{i < k} \frac{GM_j M_k}{|\hat{\mathbf{x}}_j - \hat{\mathbf{x}}_k|} + (\text{Self Energy})$$

This potential term appears in the Schrödinger Equation

Physics of “Self-Gravitating” Quantum Systems

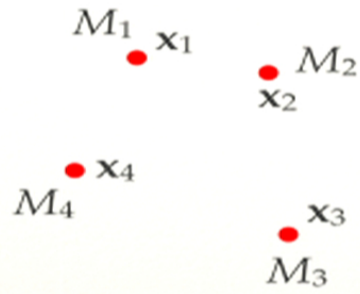


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By contrast, what are the alternatives?

1. Gravity is just classical, ϕ takes expectation value. [Møller 1962, Rosenfeld 1963, Kibble 1976, Diosi 1981; H. Yang et al., 2013]
2. Gravity is “classical” through continuous measurement, causing error in quantum-information transfer and decoherence. [Kafri & Taylor, 2013; Kafri, Taylor & Milburn, 2014; Diosi, 1987]

Classical Gravity: Schrödinger-Newton Equation

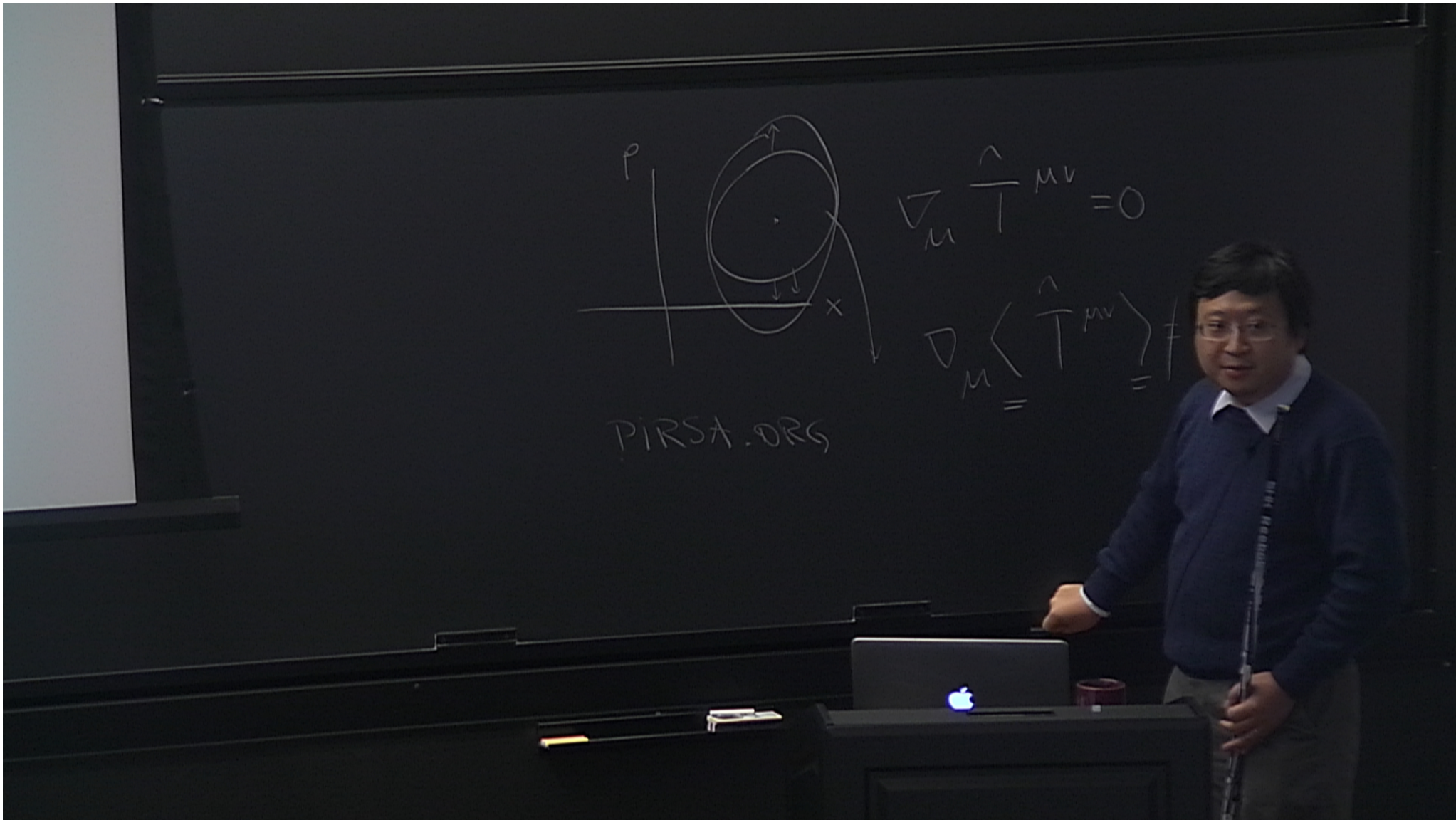
$$\nabla^2 \phi = 4\pi G \langle \hat{\rho} \rangle \Rightarrow \phi(\mathbf{x}) = - \int d^3 \mathbf{y} \frac{G \langle \hat{\rho}(\mathbf{y}) \rangle}{|\mathbf{x} - \mathbf{y}|}$$

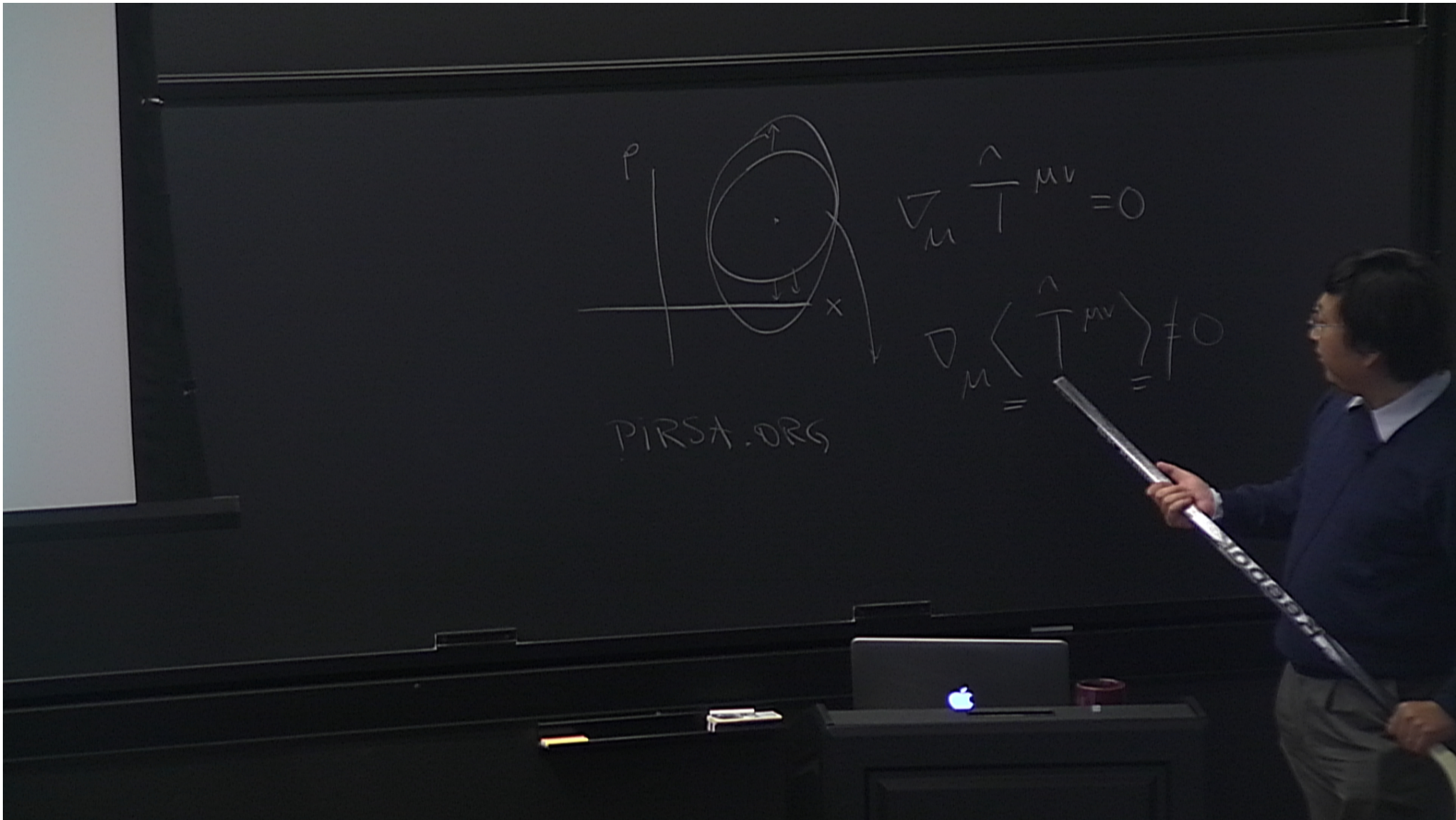
$$i\hbar \partial_t \psi(\mathbf{x}_1, \dots, \mathbf{x}_n) = \hat{H}_0 \psi(\mathbf{x}_1, \dots, \mathbf{x}_n) - \frac{1}{2} \sum_j M_j \phi(\mathbf{x}_j) \psi(\mathbf{x}_1, \dots, \mathbf{x}_n)$$

[Møller 1962, Rosenfeld 1963; Kibble 1976; Diosi 1981; Carlip 2008; Guilini 2012; H. Yang et al., 2013]

Features of the Schrödinger-Newton Equation

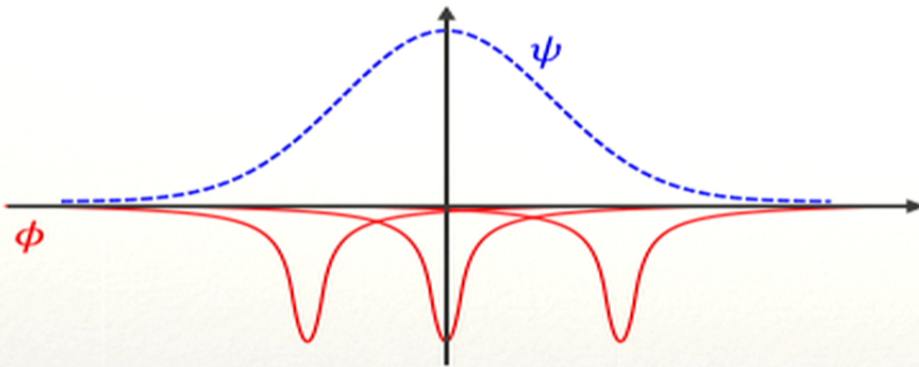
1. Nonlinear Quantum Mechanics
2. Recovers classical dynamics.
3. Absolutely No Quantum Information Transfer Between Objects.
4. Center of Mass separates out, satisfies a separate equation.





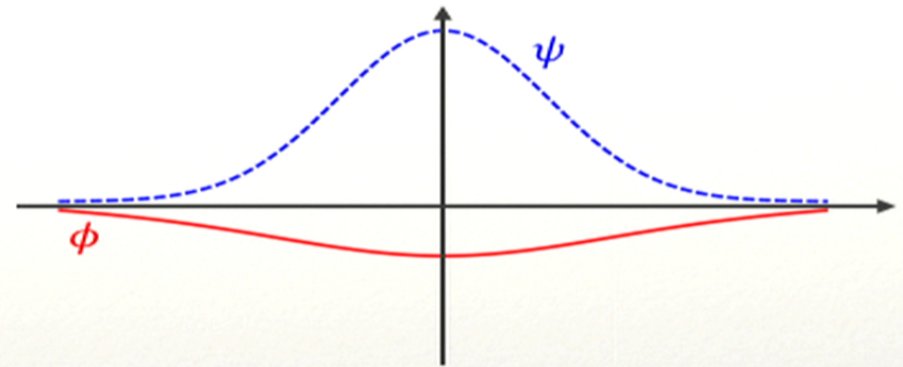
Self Interaction in Classical Gravity

Quantum "Self Gravity"



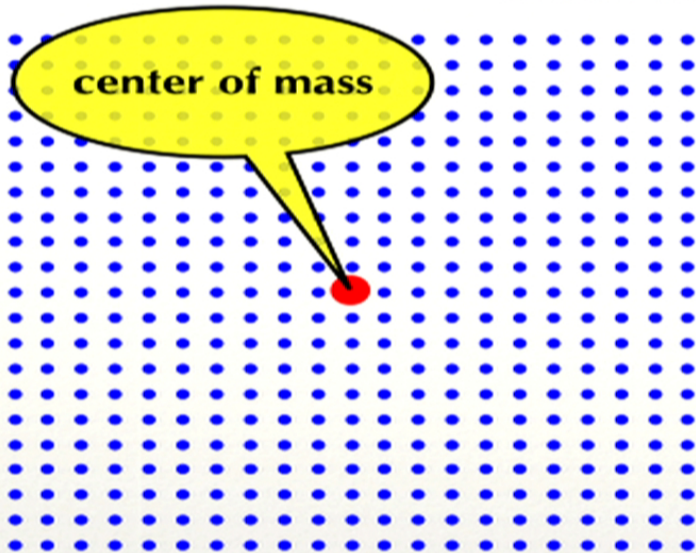
*particle carries own gravity field
gravity field entangled with particle
back action negligible*

Classical "Self Gravity"

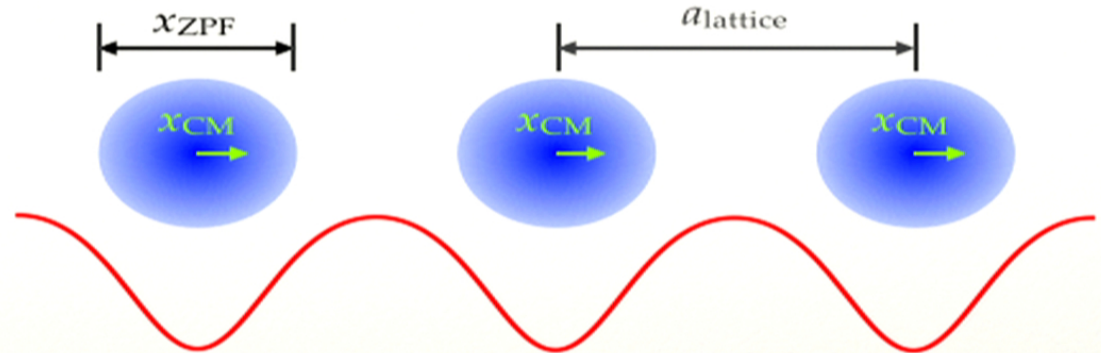


*unique classical field
wave packet attracted by its own potential*

Self Interaction for Crystals



a **macroscopic crystal** made up from atoms



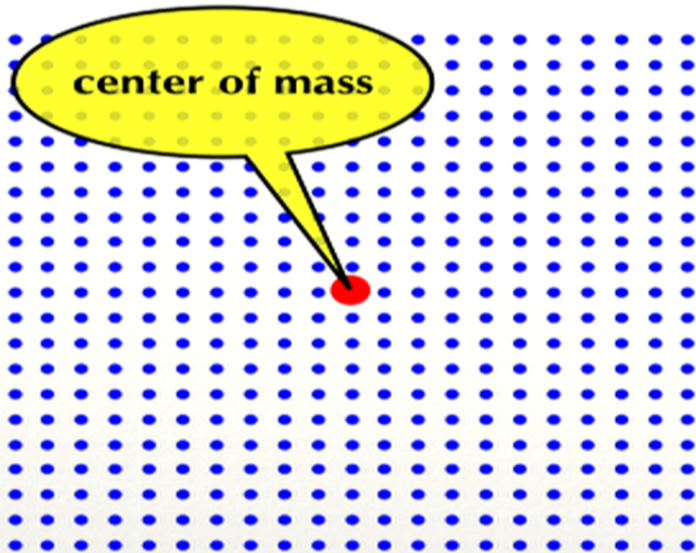
$$x_{\text{ZPF}} \sim \sqrt{\frac{\hbar}{m\omega_{\text{Debye}}}} \sim 10^{-12} \text{ m} \ll a_{\text{lattice}} \sim 10^{-10} \text{ m}$$

$$x_{\text{CM}} \ll x_{\text{ZPF}}$$

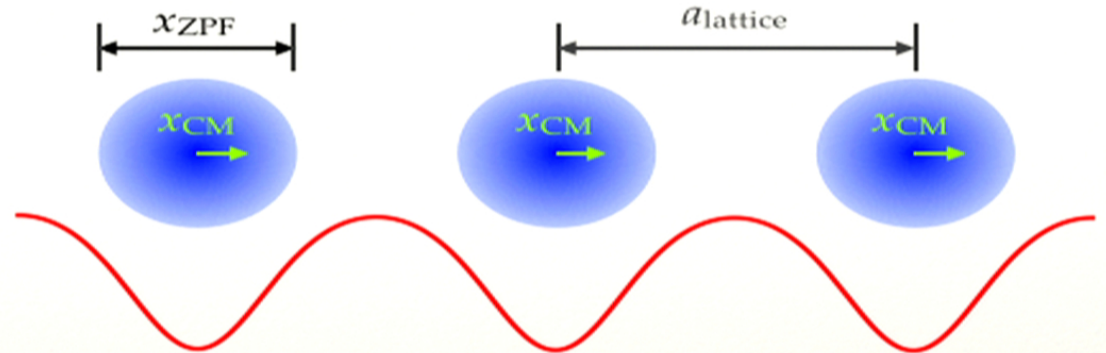
$$i\hbar \frac{\partial \Psi_{\text{CM}}}{\partial t} = \left[-\frac{\hbar^2 \nabla^2}{2M} + \frac{1}{2} M \omega_{\text{CM}}^2 x^2 + \frac{1}{2} M \omega_{\text{SN}}^2 (x - \langle x \rangle)^2 \right] \Psi_{\text{CM}}$$

$$\omega_{\text{SN}}^2 = \frac{Gm}{12\sqrt{\pi}x_{\text{ZPF}}^3} \gg \omega_g^2 \quad \omega_{\text{SN}}^{\text{Si}} = 4 \times 10^{-2} \text{ s}^{-1} \approx 100 \omega_g^{\text{Si}}$$

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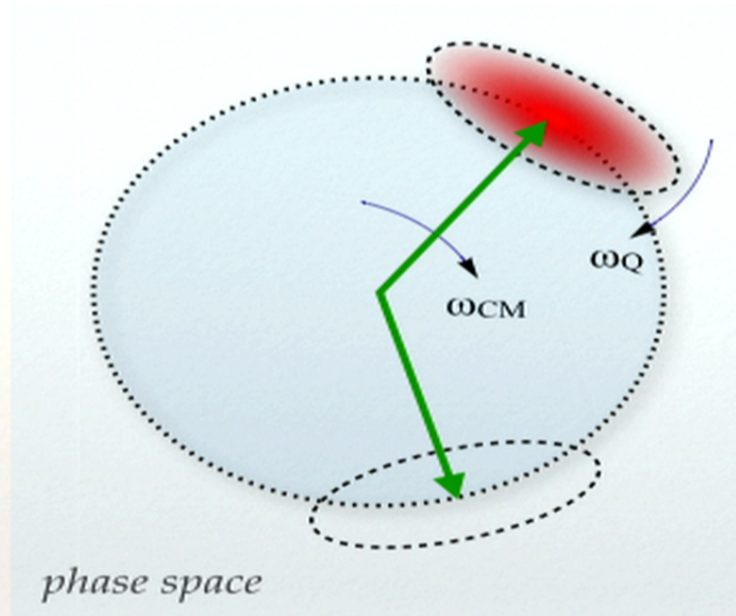
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$$\omega^{\text{Nb}} = 0.1 \text{ s}^{-1}, \quad \omega_{\text{SN}}^{\text{W}} = 0.25 \text{ s}^{-1}$$

Classical Gravity: Macroscopic Test Mass

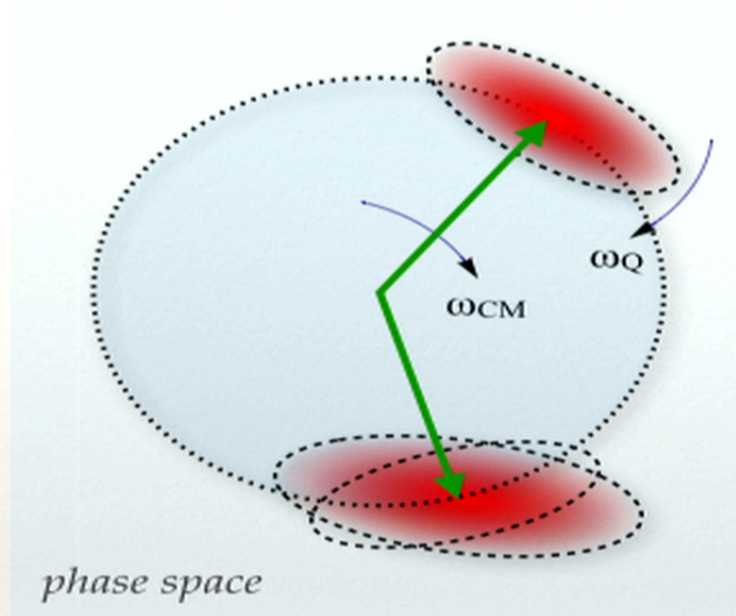


$$\dot{\hat{x}} = \hat{p} / M$$

$$\dot{\hat{p}} = -M\omega_{\text{CM}}^2 \hat{x} - M\omega_{\text{SN}}^2 (\hat{x} - \langle \hat{x} \rangle)$$

$\omega_c / (2\pi)$ (Hz)	100	10	1	0.1
Q^{Si}	3×10^8	3×10^6	3×10^4	300
Q^{Nb}	4×10^7	4×10^5	4×10^3	40
Q^{W}	6×10^6	6×10^4	6×10^2	6

Classical Gravity: Macroscopic Test Mass

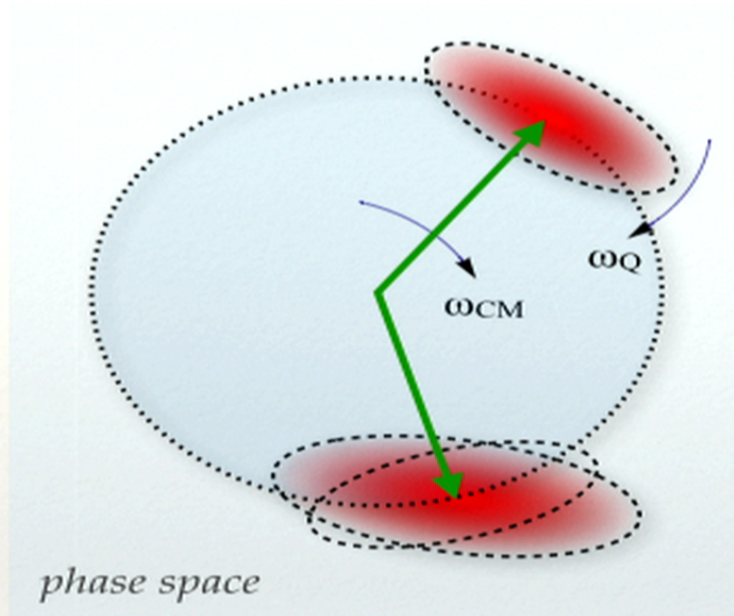


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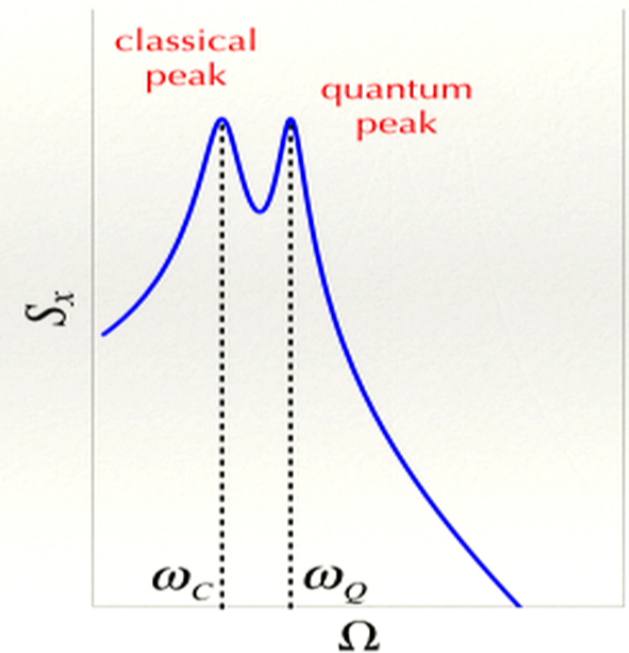
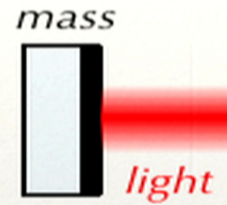


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Quantum noise ellipse rotate at a different frequency:

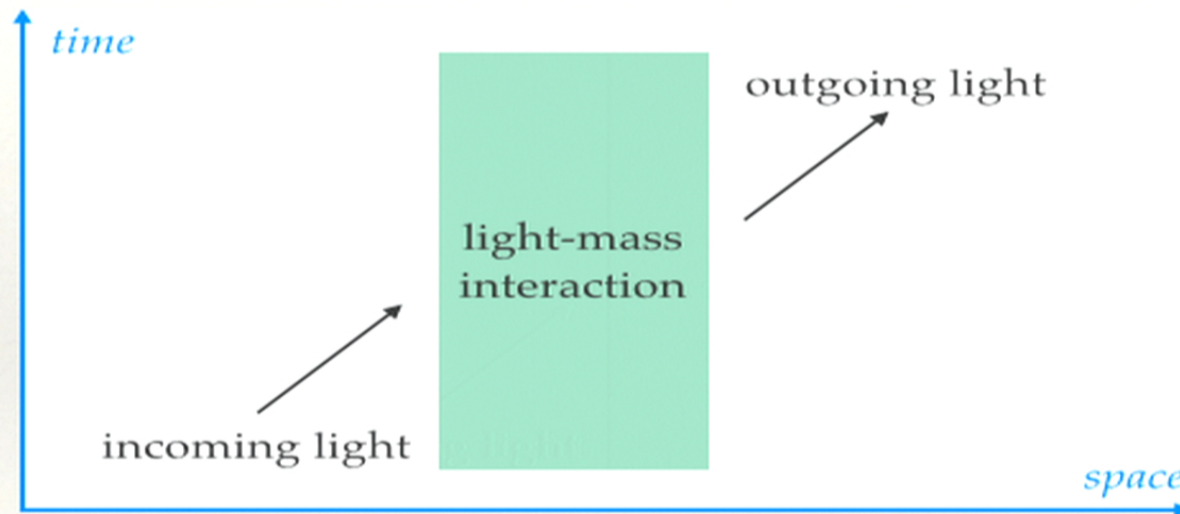
$$\omega_{\text{Q}}^2 = \omega_{\text{CM}}^2 + \omega_{\text{SN}}^2$$



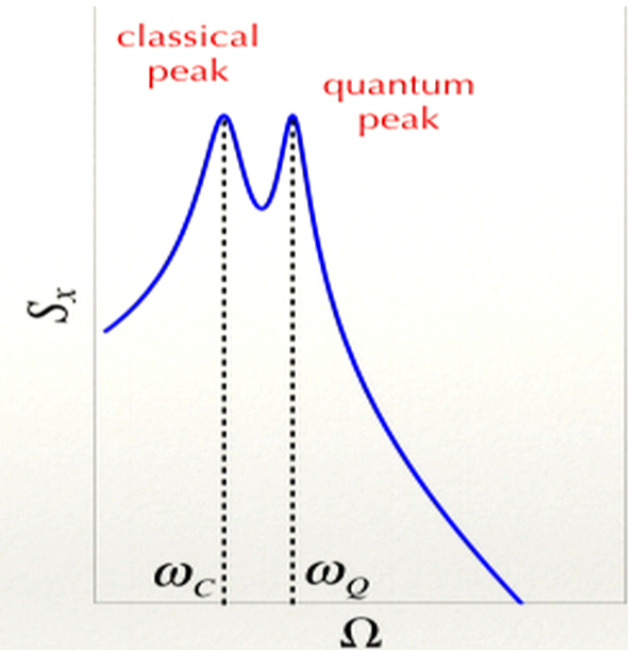
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Q^{W}	6×10^6	6×10^4	6×10^2	6

“Fundamental Flaw” of Classical Gravity

1. Interpretation of **quantum-state reduction** is a long-standing issue.
2. According to standard (linear quantum mechanics), wavefunction collapse can take place at any “moment”.



3. Unlike other situations, nonlinear quantum mechanics actually *depends on interpretation of quantum mechanics*: **Different features at ω_Q will be predicted using other interpretations.**



Summary

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1. An important ingredient of LIGO noise is due to **Heisenberg Uncertainty** of **kg-scale test masses**
 - Standard Quantum Limit represents the trade-off between **sensing** and **back-action noise**
 - techniques have been developed to circumvent SQL by altering optical configurations, often requiring **correlations between sensing and back-action noise**.
2. Various proposals for modifying quantum mechanics for macroscopic objects can be tested
 - none of the individual theories look correct/ fully-motivated on its own
 - the effects they induce (excess noise, separation in frequencies) can be tested
 - even some *gravity* effects seems within reach [*as well as "nonlinear QM", Bialynicki-Birula and Mycielski 1970s , Weinberg 1980s, Gähler, Klein & Zeilinger, 1981]*]