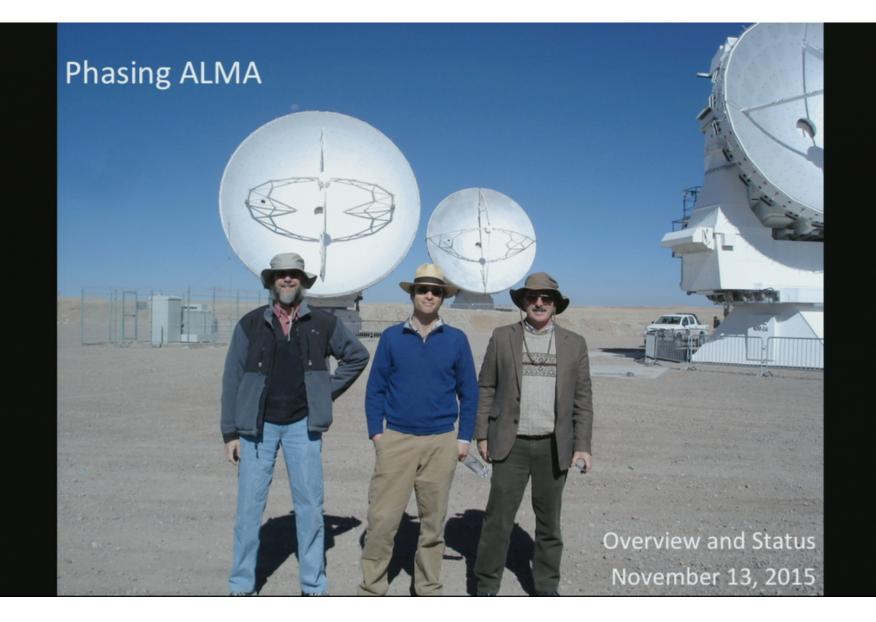
Title: Phasing ALMA

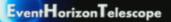
Date: Nov 13, 2014 03:45 PM

URL: http://pirsa.org/14110107

Abstract: We are now two years into development of hardware and software for phasing ALMA, with emphasis on Band 6 (1.2 mm) but applicability to bands 3 and 7. Central to the effort is software to continuously calculate and apply a phasing solution, and the hardware interface cards (PIC) and other upgrades to the correlator that apply the solution, format the data and extract the correlated data stream. Upstream from the correlator is a hydrogen maser, which is the new ALMA time standard. Downstream from the correlator are four high speed Mark 6 data recorders collectively capable of capturing 64 Gb/s, and an optical fiber link to bring the signal from the AOS, where the antennas and correlator are located, to the OFS where the recorders are located. All the hardware is currently installed and functional. A formal acceptance review for the maser was held this past week, and acceptance of the rest of the hardware will be reviewed in December. While some debugging of software is ongoing, the basic software elements are in place and (more or less) functional. Formal commissioning activities are planned to begin in January



Thanks to...

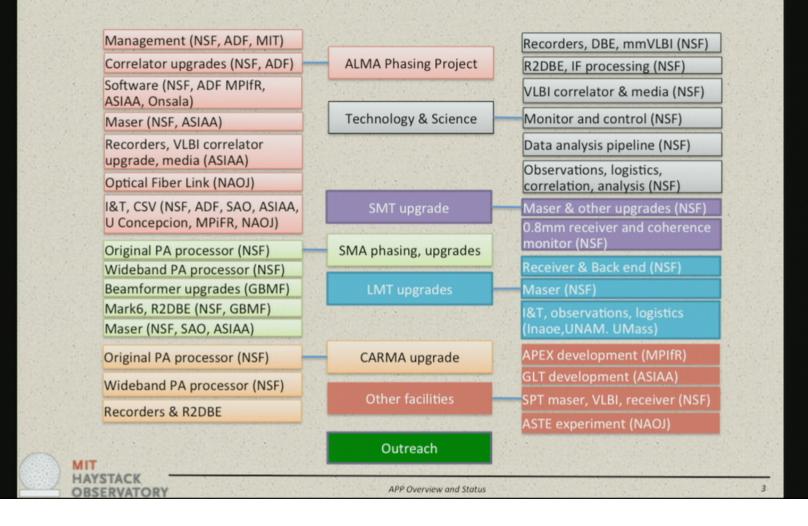


- Participating organizations below
- Sponsoring organizations
 - NSF Major Research Instrumentation program
 - ALMA NA Development Fund
 - Cost sharing partners
- ASTRON, Onsala, ESO, Univ. of Grenoble, IRAM, Board members, and especially JAO for making staff available to assist APP



EHT development (overview)

EventHorizonTelescope



Essential design features

EventHorizonTelescope

- Correlator upgrades are in parallel with the existing hardware/ firmware and do not affect the normal data path
- Interferometric ALMA data products are available while the VLBI data are being captured.
- Only relatively "minor" software enhancements to the ALMA code base are required.
- All of the phasing corrections are performed in the correlator and supporting computer system with no front-end or Local Oscillator modifications required.
- The phasing system includes a commercial hydrogen maser to provide a high stability, spectrally pure frequency standard to ALMA for routine operations as well as VLBI.

APP Overview and Status

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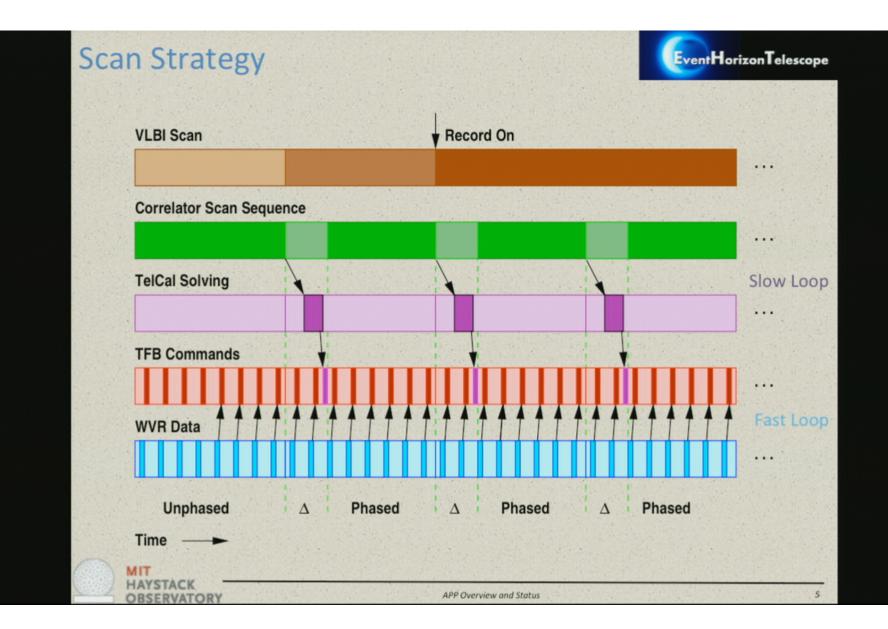
Essential design features

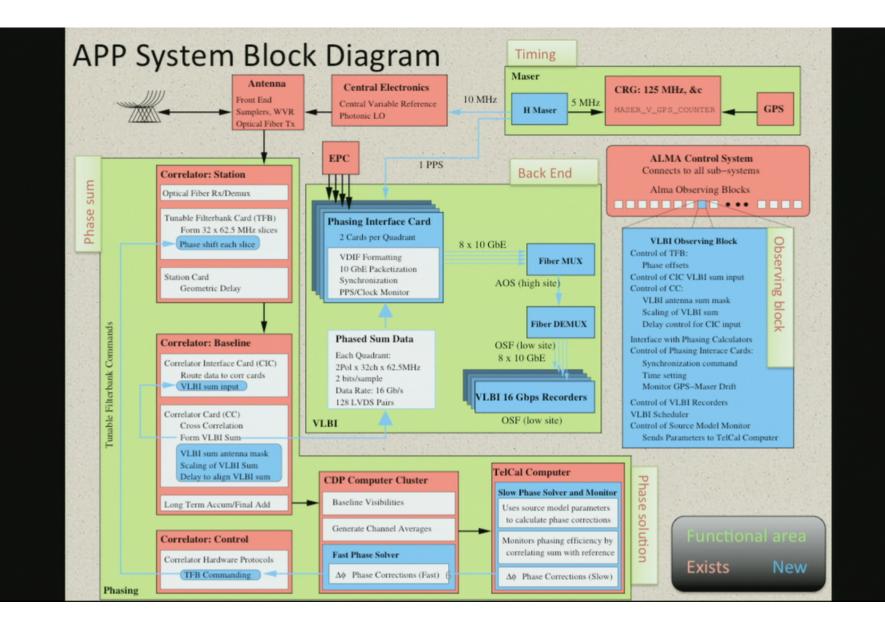
EventHorizonTelescope

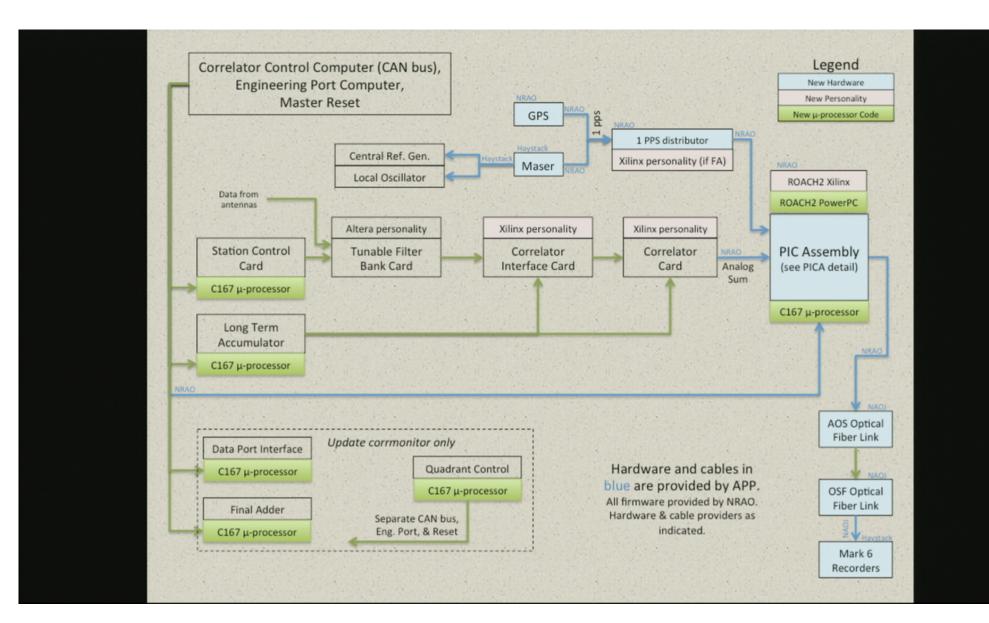
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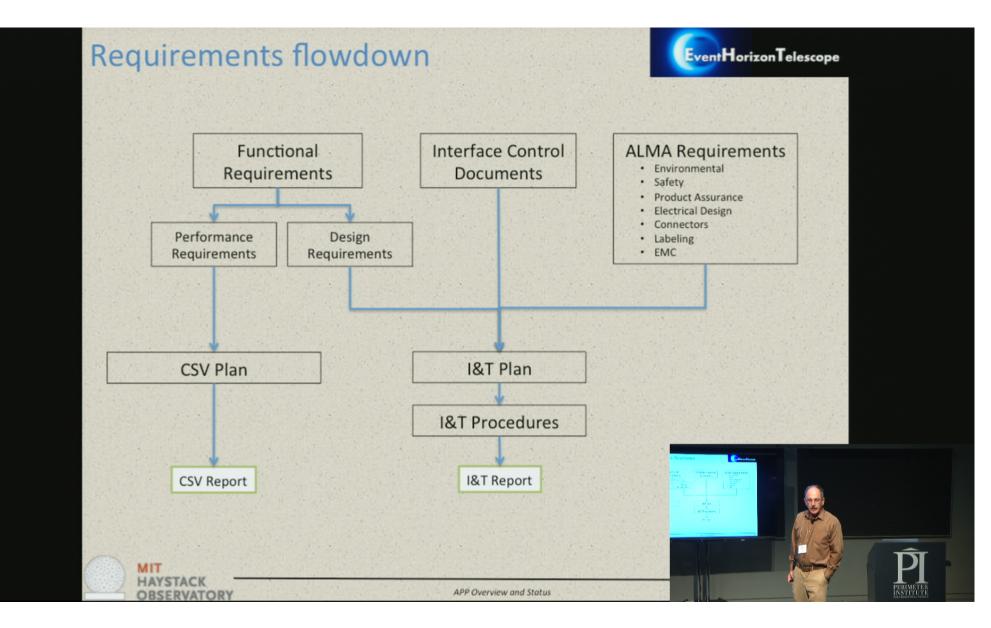
APP Overview and Status

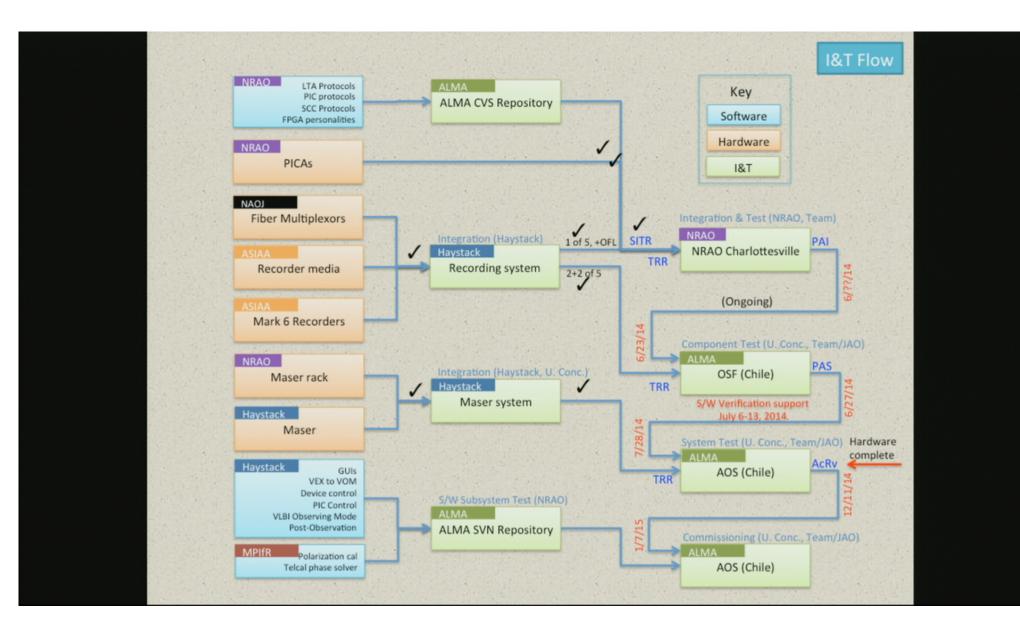
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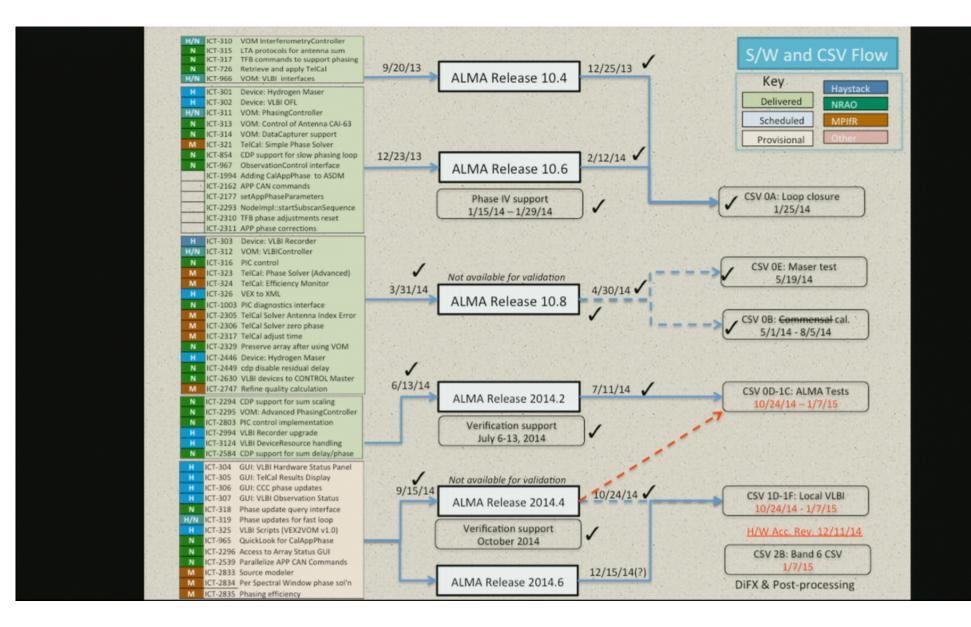












Hardware is complete and installed!

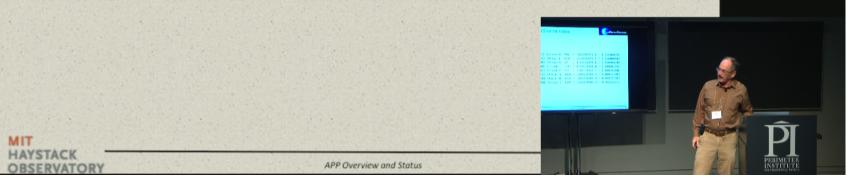


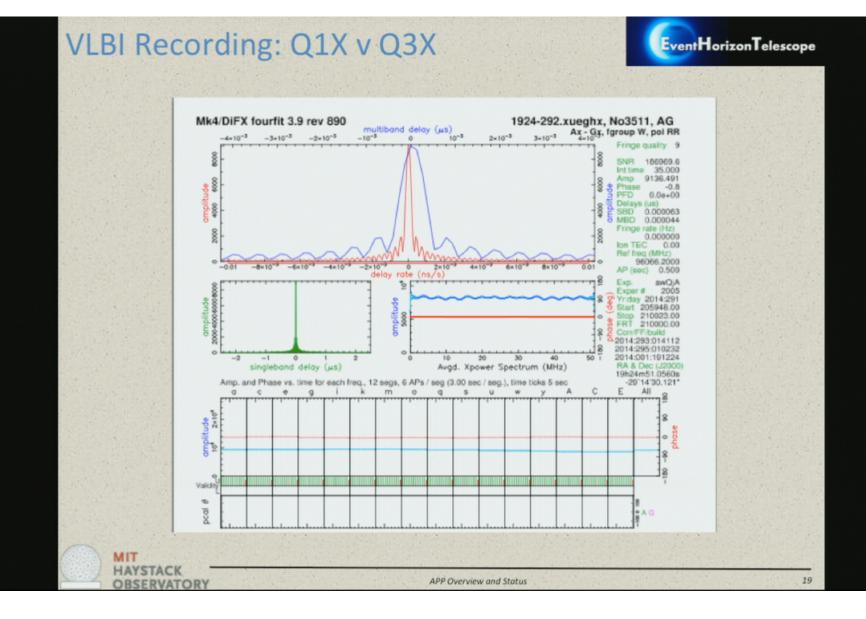


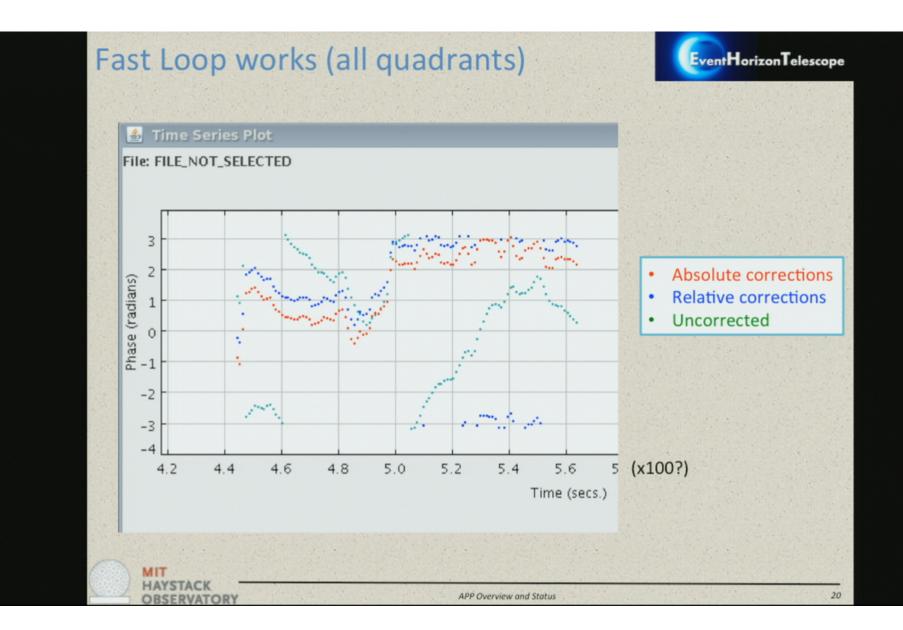
Recorder/OFLS error rates

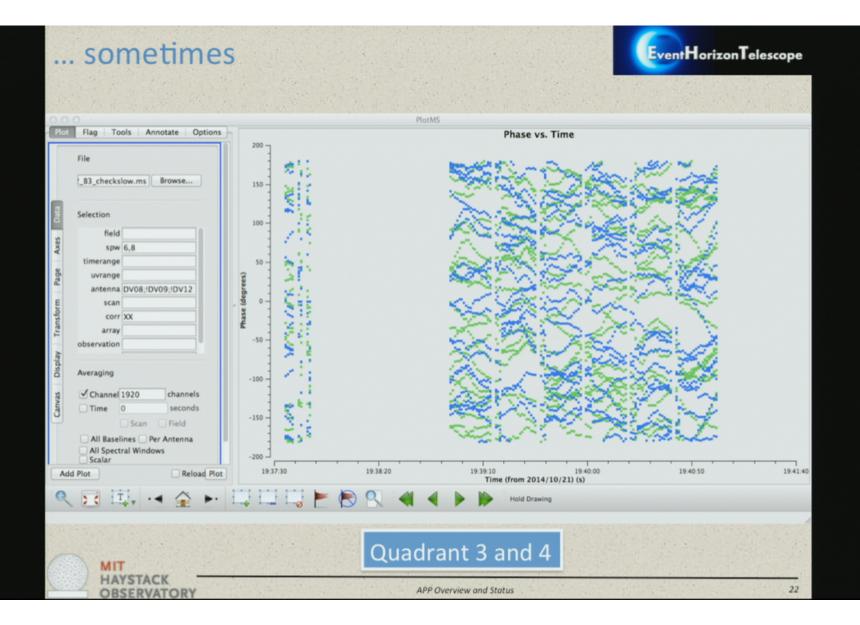
EventHorizonTelescope

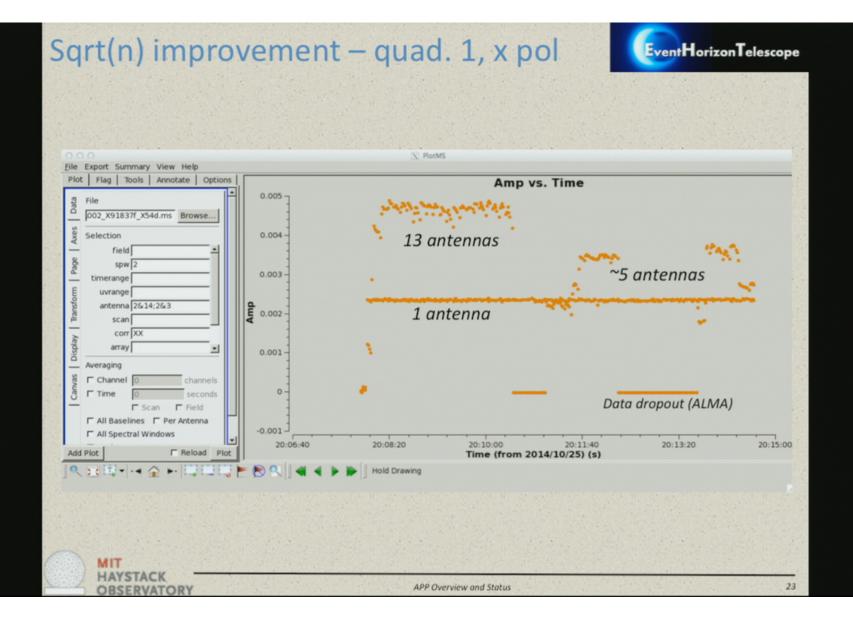
	Mark6-4005	Stream-0:	909	1	1286538975.5	=	0.000000706	
1	Mark6-4005	Stream-1:	1220	1	1286538975.5	=	0.00000948	
	Mark6-4006	Stream-0:	62	1	1286538978.0	=	0.00000048	
11	Mark6-4006	Stream-1:	1712	1	1286538978.0	=	0.000001330	
-	Mark6-4007	Stream-0:	574	1	1286539004.5	=	0.000000446	
4	Mark6-4007	Stream-1:	1885	1	1286539004.5	=	0.000001465	
-	Mark6-4008	Stream-0:	1420	1	1286539006.0	=	0.000001103	
-	Mark6-4008	Stream-1:	1587	1	1286539006.0	=	0.000001233	











Outstanding issues

EventHorizonTelescope

 Phasing only demonstrated successfully for Quadrant 1, x polarization. Indexing bug?

APP Overview and Status

- Phasing efficiency isn't as good as it should be, by comparison with off-line analysis of phasing data.
- Didn't succeed with Band 6 (weather?)

HAYSTAC

Phase 0 (Pre-Commissioning)

EventHorizonTelescope

Objectives:

- Test VOM block execution during real observations
- Insure that APP operation does not negatively impact standard ALMA operations
- Verify that data products meet specifications for phasing stability, efficiency (via commensal observing)
 - Characterize polarization properties (via dedicated observations)

Proposed Tests:

MIT

- Test 0A: Observations of calibrator sources (non-VLBI mode)
- Test OB: Commensal calibrator source observations (non-VLBI mode)
- Test OC*: Polarization characterization with the ALMA Beacon
 - Test OD: Polarization characterization using observations of astronomical sources

APP Overview and Status

Does not require PIC, maser, OFL, or recorders:

*Contingent upon availability/suitability of Beacon for testing

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Phase 1 ("Local VLBI")

EventHorizonTelescope

Objectives:

Exercise the APP in VLBI and pseudo-VLBI observing modes through "local VLBI"

Demonstrate readiness of APP to participate in a global VLBI network

Proposed Tests:

- Test 1A: Stepped scans with closure phase tests
 - Test 1B*: "Single-dish" spectroscopy of astronomical line source
- Test 1C: Inter-quadrant VLBI
- Test 1D*: VLBI between APP and OSF antenna (fringe test)
- Test 1E*: VLBI between APP and OSF antenna (stepped scans)

APP Overview and Status

Test 1F: VLBI between APP and APEX (fringe test)

*Desirable, not mandatory

HAYSTAC

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Phase 2 (Global VLBI)

EventHorizonTelescope

Objectives:

MIT

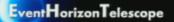
- Utilize APP in full-fledged global VLBI experiments
- Demonstrate the functionality of the fully-integrated APP for science
- Proposed Tests:
 - Test 2A: Global VLBI (3 mm = Band 3)
 - Test 2B: Global VLBI (1.3 mm = Band 6)

As per agreement, observing targets will include neither Sgr A nor M87

APP Overview and Status

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Anticipated VLBI Observations



- Jan 2015 (Commissioning)
 - Local VLBI w/ APEX and/or OSF
 - Possible 1.3 mm global VLBI with SPT
 - Probably just one ALMA dish, test of VLBI (not phasing)
 - Pending hardware acceptance
- March 2015
- May 2015 (Commissioning)
 - Global 3 mm VLBI (band 3)
 - w/ VLBA, others TBD
- March 2016

MIT

- Science observations, subject to approval
- Band 6, 16-32 Gb/s
- Participants TBD

APP Overview and Status

