

Title: Search for new physics in atoms: cosmic PNC and variation of alpha

Date: Jun 18, 2014 05:00 PM

URL: <http://pirsa.org/14060030>

Abstract: We consider pseudo-scalar and pseudo-vector interaction of atomic electrons with hypothetical dark matter particles (e.g., axions). These interactions lead to oscillating atomic parity non-conserving (PNC) amplitudes and/or oscillating electric dipole moments (EDM). In static
 limit for PNC, existing atomic PNC experiments can be used to constrain time component of the pseudo-vector field.

 Possible variation of fundamental constants is suggested by theories unifying gravity with other interactions. Evidence of the space/time variation of the fine structure constant alpha is found in
 the quasar absorption spectra. Optical transitions in highly charged ions can be used as sensitive tools for studying time variation of alpha in laboratory.

Cosmic PNC

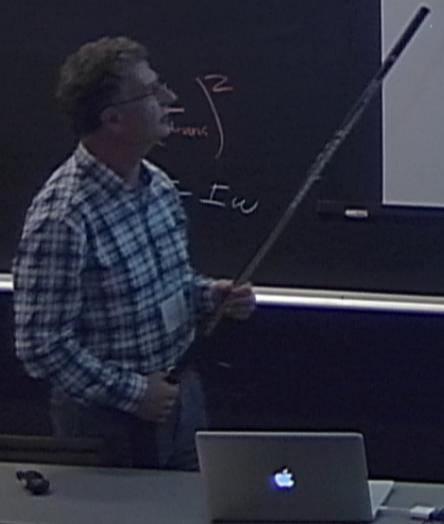
Assuming dark matter breaks parity we consider pseudoscalar and pseudovector interactions.

Pseudoscalar field (e.g., axions):

$$L^{PS} = \kappa \hbar (\partial_\mu \varphi) \bar{\psi} \gamma^\mu \gamma^5 \psi - i k m_f c^2 \phi \bar{\psi} \gamma^5 \psi$$

Pseudovector field:

$$L^{PV} = b_\mu \bar{\psi} \gamma^\mu \gamma^5 \psi$$



Lead to oscillating PNC and EDM:

$$E_{PNC}^{PS} = (\kappa + \frac{k}{2})\hbar\omega_q \sin(\omega_q t)K_{PNC}$$

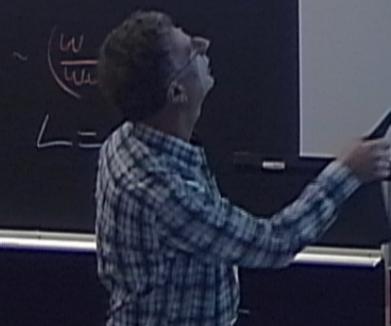
$$E_{EDM}^{PS} = -(2\kappa + k)\hbar^2\omega_q^2 \cos(\omega_q t)K_{EDM}$$

$$E_{PNC}^{PV} = b_0 K_{PNC}$$

$$E_{EDM}^{PV} = -2ib_0\hbar\omega_b \cos(\omega_b t)K_{EDM}$$

No static EDM for PS or PV!

No static PNC for PS field.



K_{PNC} and K_{EDM} are electron structure factors to be found in atomic calculations

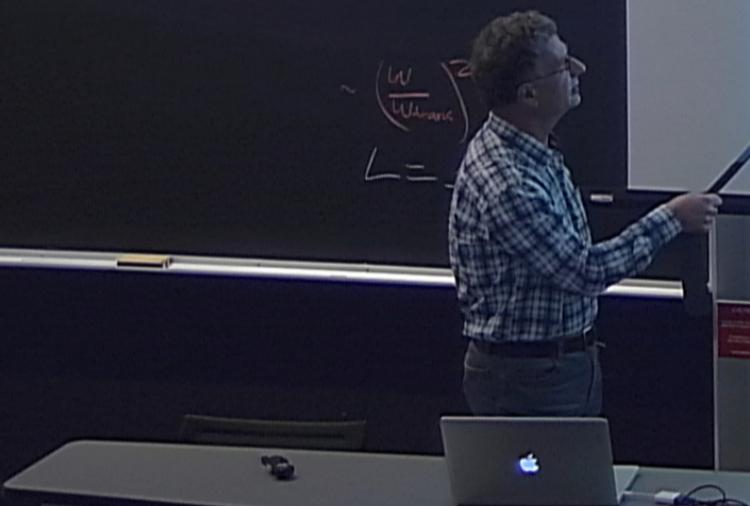
$$K_{PNC} = \sum_n \left[\frac{\langle b|d|n\rangle\langle n|\gamma^5|a\rangle}{E_a - E_n} + \frac{\langle b|\gamma^5|n\rangle\langle n|d|a\rangle}{E_b - E_n} \right] \quad (\text{Similar to PNC due to } Q_W \text{ or AM.})$$

$$K_{EDM} = \sum_n \frac{\langle a|d|n\rangle\langle n|\gamma^5|a\rangle}{(E_a - E_n)^2} \quad (\text{Extra square in the denominator compared to EDM due to SM or eEDM})$$

Both PS and PV interactions are reduced to matrix elements of γ^5

In non-relativistic limit $\langle a|\gamma^5|b\rangle \propto (E_a - E_b)\langle a|r|b\rangle$

Which means that $K_{EDM} \propto \alpha_0$ and $K_{PNC} \equiv 0$



PV contribution to atomic PNC

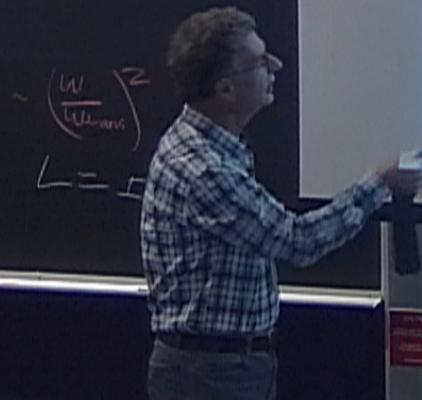
In case of $b_0(t) = b_0 = \text{const}$

limits on b_0 can be found from existing PNC measurements.
We consider Cs, Tl and Dy.

Interpretation requires relativistic calculation since in
non-relativistic limit

$$\langle a | \gamma^5 | b \rangle \propto (E_a - E_b) \langle a | r | b \rangle \quad \text{It is 0 if } E_a = E_b \text{ e.g. in Dy}$$

$$\text{and } K_{PNC} = \sum_n \left[\frac{\langle b | d | n \rangle \langle n | \gamma^5 | a \rangle}{E_a - E_n} + \frac{\langle b | \gamma^5 | n \rangle \langle n | d | a \rangle}{E_b - E_n} \right] = C - C = 0$$



In single-electron case: $\langle a | \gamma^5 | b \rangle = i \int (f_b g_a - g_b f_a) dr$

Similar to standard PNC: $\langle a | \gamma^5 \rho | b \rangle = i \int (f_b g_a - g_b f_a) \rho_N(r) dr$

α -expansion:

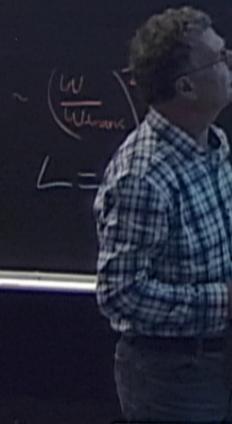
$$\langle a | \gamma^5 | b \rangle =$$

$$\alpha(E_b - E_a) \langle a | \sigma \cdot \mathbf{r} | b \rangle + \quad \sim \alpha - \text{gives zero}$$

$$2 \langle a | \gamma^5 (1 + \sigma \cdot \mathbf{L}) | b \rangle \quad \sim \alpha^3 - \text{gives non-zero}$$

In single-electron case (Cs,Fr,Tl):

$$\mathcal{E}_{\text{PNC}}^{i \rightarrow f} = \alpha^2 (\kappa_i + \kappa_f) \langle f | \gamma^5 \mathbf{d} | i \rangle$$



In single-electron case: $\langle a | \gamma^5 | b \rangle = i \int (f_b g_a - g_b f_a) dr$

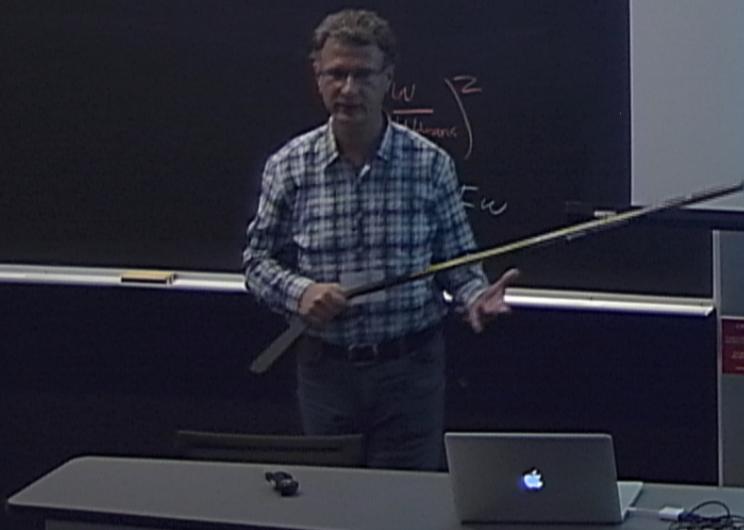
Similar to standard PNC: $\langle a | \gamma^5 \rho | b \rangle = i \int (f_b g_a - g_b f_a) \rho_N(r) dr$

α -expansion:

$$\begin{aligned} \langle a | \gamma^5 | b \rangle &= \\ \alpha(E_b - E_a) \langle a | \sigma \cdot \mathbf{r} | b \rangle + &\quad \sim \alpha - \text{gives zero} \\ 2 \langle a | \gamma^5 (1 + \sigma \cdot \mathbf{L}) | b \rangle &\quad \sim \alpha^3 - \text{gives non-zero} \end{aligned}$$

In single-electron case (Cs,Fr,Tl):

$$E_{\text{PNC}}^{i \rightarrow f} = \alpha^2 (\kappa_i + \kappa_f) \langle f | \gamma^5 \mathbf{d} | i \rangle$$



We have good agreement between γ^5 and α^3 terms for Cs and Ti; we use α^3 term for Dy.

The limit on β_0 is found by

$$b_0 < \frac{|A_{\text{expt}} - A_{\text{theor}}| + \sigma_{\text{expt}} + \sigma_{\text{theor}}}{K_{PNC}}$$

Atom	A_{expt} (10^{-11} a.u.)	A_{theor} (10^{-11} a.u.)	K_{PNC} (10^{-7} a.u.)	b_0 (10^{-7} a.u.)	b_0 (10^{-15} GeV)
Cs	0.8428(26)	0.8353(42)	2.5	6	15
Ti	25.6(2)	24.8(7)	2.1	800	2200
Dy	2.3(3.0) Hz*	4(4) Hz*	0.08	3	7

* $1\text{Hz} = 1.52 \times 10^{-16}$ a.u.

Cs: Wood *et al.*, Science 275, 1759 (1997); Dzuba *et al.*, PRL 109, 203003 (2012);

Ti: Vetter *et al.*, PRL 74, 2658 (1995); Dzuba *et al.*, JPB 20, 3297 (1987);

Dy: Nguyen *et al.*, PRA 56, 3453 (1997); Dzuba and Flambaum, PRA 81, 052515 (2010).

We have good agreement between γ^5 and α^3 terms
for Cs and Ti; we use α^3 term for Dy.

The limit on β_0 is found by

$$b_0 < \frac{|A_{\text{expt}} - A_{\text{theor}}| + \sigma_{\text{expt}} + \sigma_{\text{theor}}}{K_{PNC}}$$

Atom	A_{expt} (10^{-11} a.u.)	A_{theor} (10^{-11} a.u.)	K_{PNC} (10^{-7} a.u.)	b_0 (10^{-7} a.u.)	b_0 (10^{-15} GeV)
Cs	0.8428(26)	0.8353(42)	2.5	6	15
Tl	21.6(2)	24.8(7)	2.1	800	2200
Dy	2.3(3.0) Hz*	4(4) Hz*	0.08	3	7

1 GeV = 1.52×10^{-16} a.u.

Cs: Wood *et al.*, Science 275, 1759 (1997); Dzuba *et al.*, PRL 109, 203003 (2012);

Tl: Vetter *et al.*, PRL 74, 2658 (1995); Dzuba *et al.*, JPB 20, 3297 (1987);

Dy: Nguyen *et al.*, PRA 56, 3453 (1997); Dzuba and Flambaum, PRA 81, 052515 (2010).

We have good agreement between γ^5 and α^3 terms for Cs and Tl; we use α^3 term for Dy.

The limit on β_0 is found by

$$b_0 < \frac{|A_{\text{expt}} - A_{\text{theor}}| + \sigma_{\text{expt}} + \sigma_{\text{theor}}}{K_{PNC}}$$

Atom	A_{expt} (10^{-11} a.u.)	A_{theor} (10^{-11} a.u.)	K_{PNC} (10^{-7} a.u.)	b_0 (10^{-7} a.u.)	b_0 (10^{-15} GeV)
Cs	0.8428(26)	0.8353(42)	2.5	6	15
Tl	25.6(2)	24.8(7)	2.1	800	2200
Dy	2.3(3.0) Hz*	4(4) Hz*	0.08	3	7

* 1Hz = 1.52×10^{-16} a.u.

Cs: Wood *et al.*, Science 275, 1759 (1997); Dzuba *et al.*, PRL 109, 203003 (2012);

Tl: Vetter *et al.*, PRL 74, 2658 (1995); Dzuba *et al.*, JPB 20, 3297 (1987);

Dy: Nguyen *et al.*, PRA 56, 3453 (1997); Dzuba and Flambaum, PRA 81, 052515 (2010).

PV interaction of dark matter with nuclear protons and neutrons gives additional contribution κ_b to nuclear anapole moment (AM)

$$\kappa_p = \frac{2\sqrt{2}\hbar\pi\alpha\mu\langle r^2 \rangle}{a_0^3 m_p c} b_0^N \quad H_{AM} = \frac{G_F K}{\sqrt{2}} \frac{\alpha \cdot I}{I} (\kappa_a + \kappa_b) \rho_N(r)$$

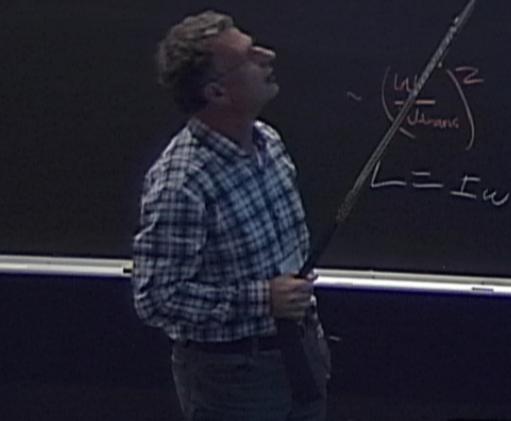
AM moment measurements in Cs and Tl can be used to put limits on b_0^P

Atom	Exp κ_a	Theor κ_a	b_0^P (a.u.)
Cs	0.364(62)	0.15 – 0.23	1.1
Tl	-0.22(30)	0.10 – 0.24	3.1

Cs: Wood et al, Science 275, 1759 (1997); Flambaum and Murray, PRC 56, 1641 (1997);
Dmitriev and Telitsin, Nucl. Phys. A 613, 237 (1987); Haxon et al PRL 86, 5287 (2001);
Tl: Vetter et al, PRL 74, 2658 (1995); Khriplovich, Phys. Lett. A 197, 316 (1995).

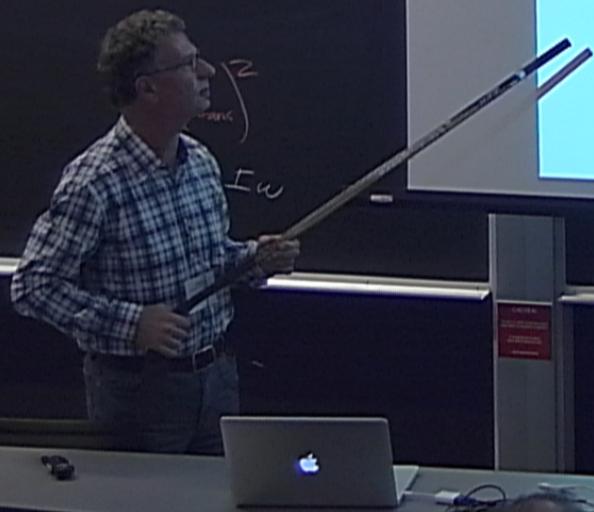
Summary

- PS and PV interaction of dark matter with atomic electrons or nucleons leads to oscillating PNC or EDM
- There is no static EDM
- There is no static PNC in case of PS interaction.
- There is static PNC due to time component of the PV field.
- Interpretation of EDM measurements is simple and based on atomic polarizabilities.
- PNC amplitudes are suppressed and interpretation requires relativistic expansion of the operator.
- Existing PNC measurements can be used to put limits on the time component of PV interaction of dark matter with atomic electrons or protons.



Summary

- PS and PV interaction of dark matter with atomic electrons or nucleons leads to oscillating PNC or EDM
- There is no static EDM
- There is no static PNC in case of PS interaction.
- There is static PNC due to time component of the PV field.
- Interpretation of EDM measurements is simple and based on atomic polarizabilities.
- PNC amplitudes are suppressed and interpretation requires relativistic expansion of the operator.
- Existing PNC measurements can be used to put limits on the time component of PV interaction of dark matter with atomic electrons or protons.



Variation of alpha

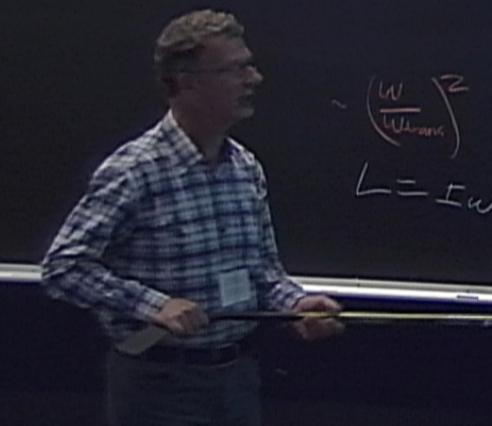
Possible variation of fundamental constants is suggested by theories unifying gravity with other interactions.

Fitting of the quasar absorption spectra data suggests spatial variation of alpha (Webb et al PRL 107, 191101 (2011)):

$$\frac{\delta\alpha}{\alpha_0} = (1.10 \pm 0.25) \times 10^{-6} r \cos\psi \text{ Glyr}^{-1} \text{(Australalian or alpha dipole).}$$

Earth movements in the framework of α -dipole leads to time variation of α in laboratory (Berengut and Flambaum EPL 97, 20006 (2012)):

$$\frac{\partial\alpha}{\partial t} \frac{1}{\alpha} = [1.35 \times 10^{-18} \cos\psi + 1.4 \times 10^{-20} \cos\omega t] \text{yr}^{-1} \approx 10^{-19} \text{yr}^{-1}$$



Variation of alpha

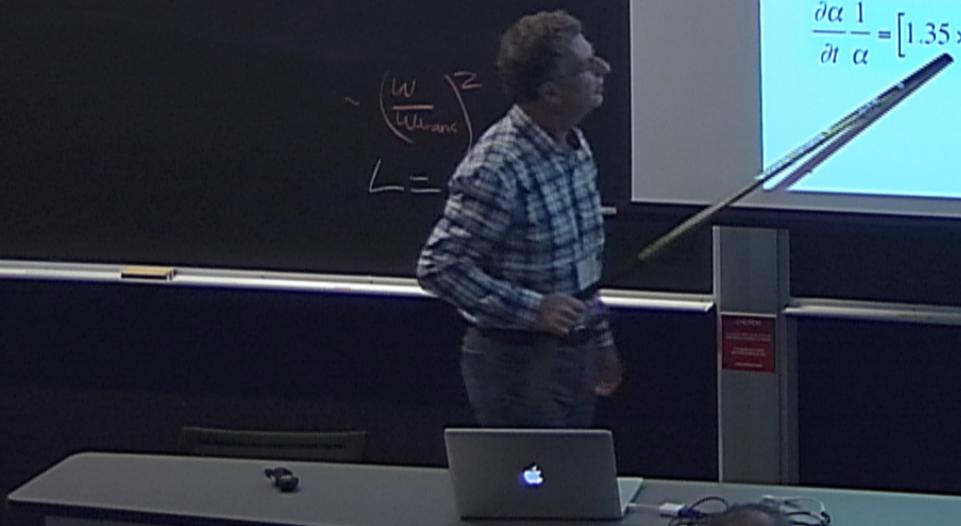
Possible variation of fundamental constants is suggested by theories unifying gravity with other interactions.

Fitting of the quasar absorption spectra data suggests spatial variation of alpha (Webb et al PRL 107, 191101 (2011)):

$$\frac{\delta\alpha}{\alpha_0} = (1.10 \pm 0.25) \times 10^{-6} r \cos\psi \text{ Glyr}^{-1} \text{(Australalian or alpha dipole).}$$

Earth movements in the framework of α -dipole leads to time variation of α in laboratory (Berengut and Flambaum EPL 97, 20006 (2012)):

$$\frac{\partial\alpha}{\partial t} \frac{1}{\alpha} = [1.35 \times 10^{-18} \cos\psi + 1.4 \times 10^{-20} \cos\omega t] \text{yr}^{-1} \approx 10^{-19} \text{yr}^{-1}$$



Sr and Yb clocks are not sensitive to variation of alpha

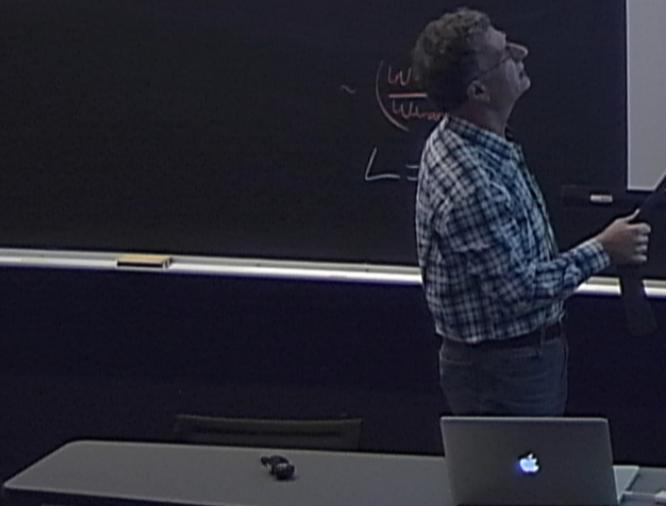
$$\frac{\partial \omega}{\partial t} \frac{1}{\omega} = K \frac{\partial \alpha}{\partial t} \frac{1}{\alpha} \quad K=0.31 \text{ for Yb and } K=0.062 \text{ For Sr}$$

We need $K \gg 1$!

$$\omega = \omega_0 + q \left[\left(\alpha / \alpha_0 \right)^2 - 1 \right] \quad K = 2q / \omega_0$$

Atom	State 1	State 2	K
Yb II	4f ¹⁴ 6s	4f ¹³ 6s ²	-6
Cs II	5d ¹⁰ 6s	5d ⁹ 6s ²	-3
Te	5p ⁴ 3P ₁	5p ⁴ 3P ₀	106
Dy	4f ¹⁰ 5d6s	4f ⁹ 5d ² 6s	$\sim 10^8$

$$\omega_0 \approx 0$$



Sr and Yb clocks are not sensitive to variation of alpha

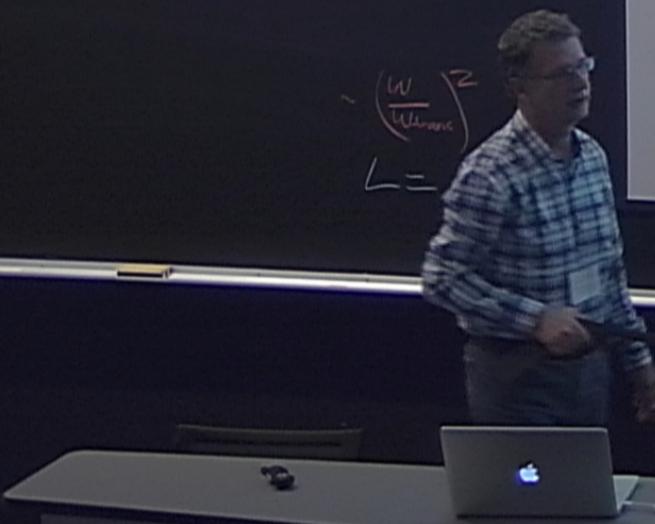
$$\frac{\partial \omega}{\partial t} \frac{1}{\omega} = K \frac{\partial \alpha}{\partial t} \frac{1}{\alpha} \quad K=0.31 \text{ for Yb and } K=0.062 \text{ For Sr}$$

We need $K \gg 1$!

$$\omega = \omega_0 + q \left[\left(\alpha / \alpha_0 \right)^2 - 1 \right] \quad K = 2q / \omega_0$$

Atom/ion	State 1	State 2	K
Yb II	4f ¹⁴ 6s	4f ¹³ 6s ²	-6
Hg II	5d ¹⁰ 6s	5d ⁹ 6s ²	-3
Te	5p ⁴ ³ P ₁	5p ⁴ ³ P ₀	106
Dy	4f ¹⁰ 5d6s	4f ⁹ 5d ² 6s	$\sim 10^8$

$$\omega_0 = 0$$



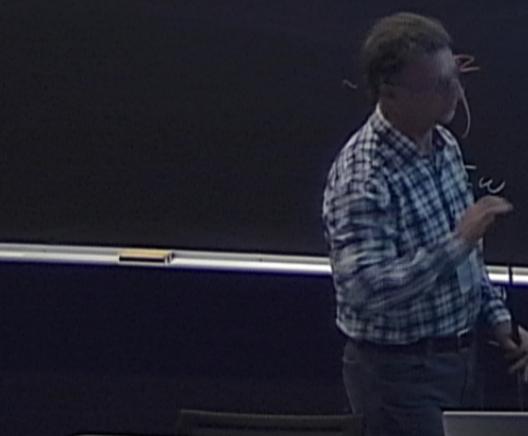
$$\frac{\partial \alpha}{\partial t} \frac{1}{\alpha} = (-5.8 \pm 6.9) \times 10^{-17} \text{ yr}^{-1} \quad \text{Dy, Leefer et al, PRL 111, 060801 (2013)}$$

$K = 2q/\omega_0$ We want to keep ω_0 optical and look for large q .

$$\Delta E = \frac{E}{v} (Z\alpha)^2 \left(\frac{1}{j+1/2} - C(Z, j, l) \right) \quad q = \Delta E_a - \Delta E_b$$

Large q can be found in s-f or p-f transitions of highly charged ions.

Problem: such transitions are usually not optical.



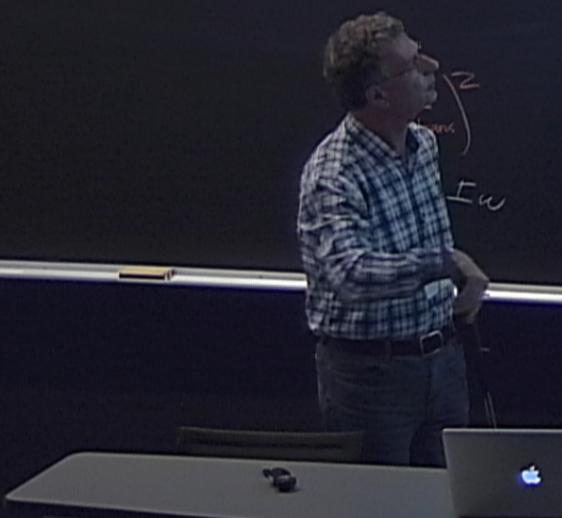
$$\frac{\partial \alpha}{\partial t} \frac{1}{\alpha} = (-5.8 \pm 6.9) \times 10^{-17} \text{ yr}^{-1} \quad \text{Dy, Leefer et al, PRL 111, 060801 (2013)}$$

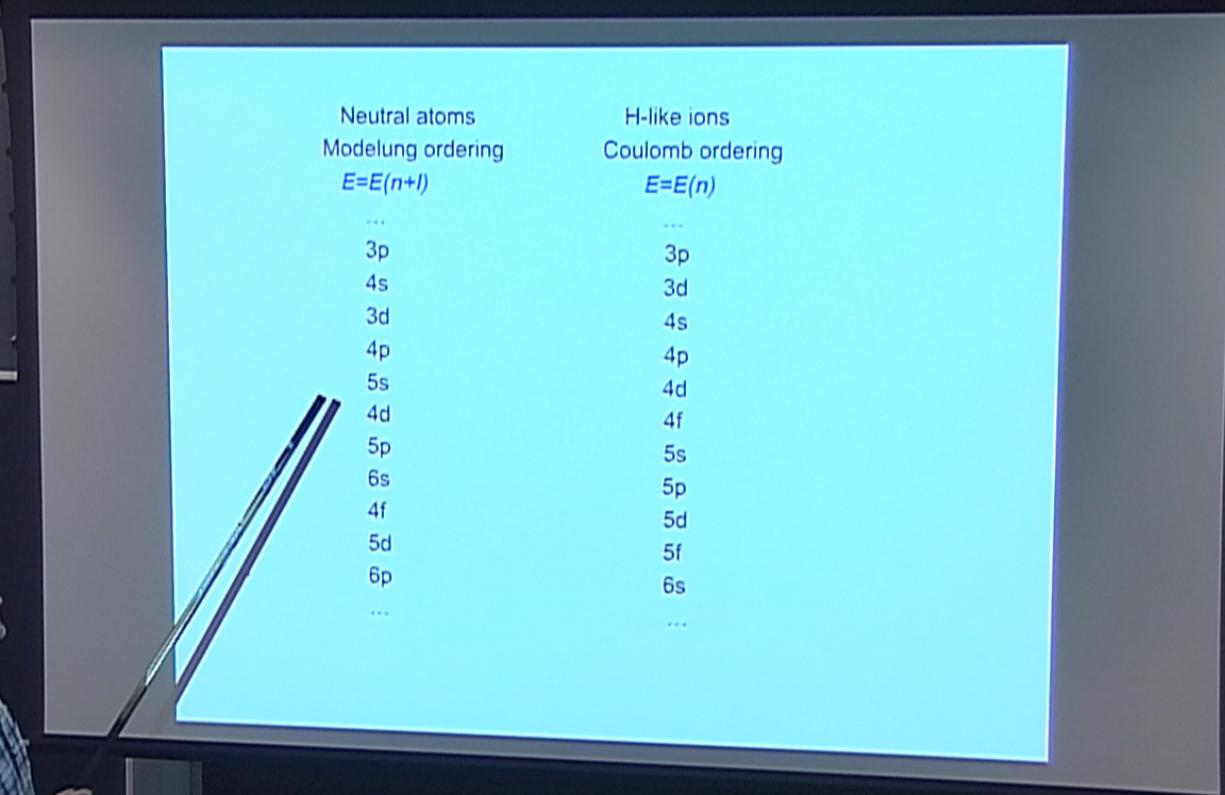
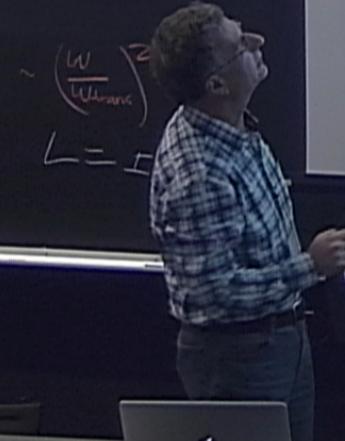
$K = 2q/\omega_0$ We want to keep ω_0 optical and look for large q .

$$\Delta E = \frac{E}{v} (Z\alpha)^2 \left(\frac{1}{j+1/2} - C(Z, j, l) \right) \quad q = \Delta E_a - \Delta E_b$$

Large q can be found in s-f or p-f transitions of highly charged ions.

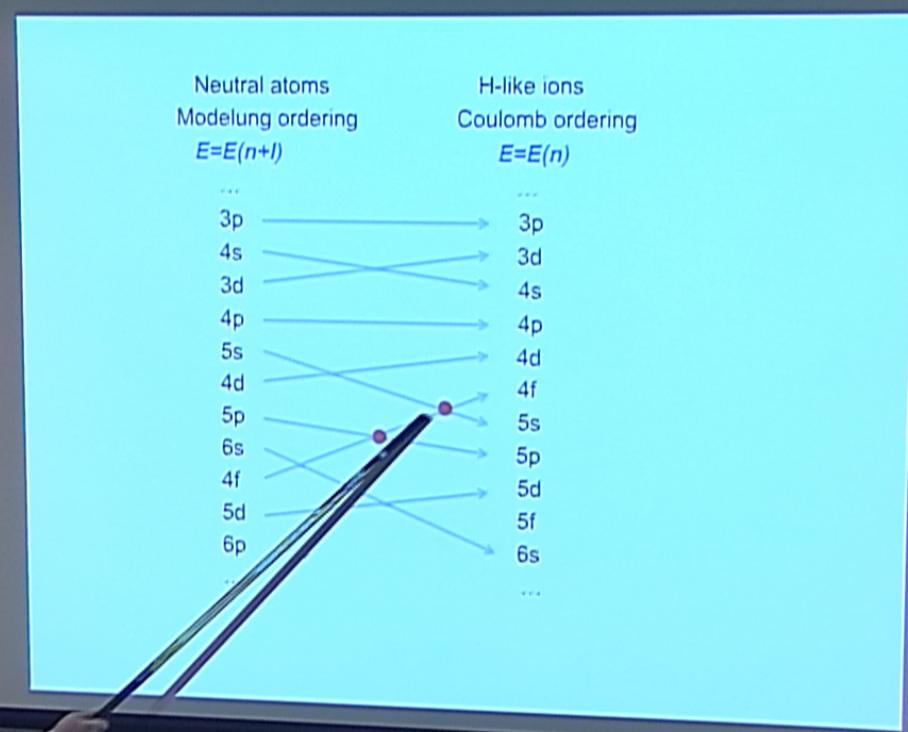
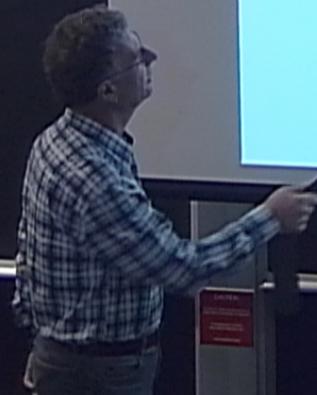
Problem: such transitions are usually not optical.





$$\sim \left(\frac{U}{U_{\text{screen}}} \right)^2$$

$$L = L_{\text{sc}}$$



There many such ions, e.g., for the 4s – 5f transitions:

Ion	Ground state	Clock state
Eu^{13+}	$5s^2 4f^2$	$5s 4f^3$
Gd^{13+}	$5s^2 4f^3$	$5s 4f^4$
Gd^{14+}	$5s^2 4f^4$	$5s 4f^5$
Dy^{13+}	$5s^2 4f^5$	$5s 4f^6$
Eu^{14+}	$5s^2 4f^6$	$5s 4f^7$
Tm^{14+}	$5s^2 4f^7$	$5s 4f^8$
Yb^{14+}	$5s^2 4f^8$	$5s 4f^9$
Lu^{14+}	$5s^2 4f^9$	$5s 4f^{10}$
...

Complicated electron structure makes analysis difficult.

Extra criterion: simple electron structure (one, two or three valence electrons).



Example:

Sm¹³⁺: $E(5s^2 4f^2 F_{7/2}) = -21.4297$ a.u.,
 $E(5s 4f^2 4H_{9/2}) = -21.3375$ a.u., $\Delta = 0.4\%$.

Theoretical uncertainty should be <<0.4% !

We use the SD+CI method which treats core-valence and valence-valence correlations to all orders

(Satonova et al, PRA 80, 012516 (2009)).

Ion	State	Expt	Calc
Nd ¹³⁺	4f _{5/2}	55870	55706
	4f _{7/2}	60300	60134
Sm ¹⁵⁺	4f _{7/2}	6555	6444
	5s _{1/2}	60384	60517

Accuracy for intervals ~ 1%

Criteria:

- Optical transition between ground and a metastable state.
- Sensitive to variation of alpha (5s-4f or 5p-4f).
- Simple electron structure
- Stable isotopes

Ion	Ground state	Clock state	Energy (cm ⁻¹)	K
Nd ¹³⁺	5s _{1/2}	4f _{7/2}	55706	3.7
Sm ¹⁵⁺	4f _{5/2}	5s _{1/2}	60517	-4.4
Ce ⁹⁺	5p _{1/2}	4f _{5/2}	54683	2.3
Pr ¹⁰⁺	5p _{1/2}	4f _{5/2}	3702	40
Nd ¹¹⁺	4f _{5/2}	5p _{1/2}	53684	-3.2
Nd ¹²⁺	5s ² ¹ S ₀	5s4f ³ F ₂	79469	2.6
Sm ¹⁴⁺	4f ² ³ H ₄	5s4f ³ F ₂	2172	-118
Sm ¹³⁺	5s ² 4f ² F _{5/2}	4f ₂ 5s ³ H _{7/2}	20254	12
Pr ⁹⁺	5p ² ³ P ₀	5p4f ³ G ₃	20216	4.2
Nd ¹⁰⁺	4f ² J=4	5p4f J=3	1564	-104

There are 10 ions with few transitions in each ion

Example: Sm¹⁴⁺ vs Pr^{10+:} $\frac{\partial \omega}{\partial t} \frac{1}{\omega} \sim -160 \frac{\partial \alpha}{\partial t} \frac{1}{\alpha}$

$$\sim \left(\frac{W}{W_{\text{trans}}} \right)^2$$

$$L = L_w$$



Another possibility: Nuclear clock.

^{229}Th , transition between isomeric nuclear states ($\Delta E \sim 7.6$ eV)

$$\frac{\delta\omega}{\omega} \sim 10^{-19} \quad (\text{Campbell et al, PRL 108, 120809 (2012)}).$$

$$\frac{\delta\omega}{\omega} = \frac{\Delta V_C}{\omega} \frac{\delta\alpha}{\alpha} \quad \frac{\Delta V_C}{\omega} \sim 10^2 + 10^4 \quad \begin{matrix} \text{Nuclear theory cannot give} \\ \text{accurate value of } \Delta V_C! \end{matrix}$$

What cannot be calculated can be measured!

(Berengut et al, PRL 102, 210801 (2009)).

$$\frac{\Delta V_C}{(\text{MeV})} \approx -506 \frac{\Delta \langle r^2 \rangle}{\langle r^2 \rangle} + 23 \frac{\Delta Q_0}{Q_0}$$

$\Delta \langle r^2 \rangle$ - from isomeric frequency shift measurements

ΔQ_0 - from hfs measurements

