

Title: Probing Gravity and Small Forces with Torsion Balances

Date: Jun 17, 2014 01:30 PM

URL: <http://pirsa.org/14060018>

Abstract: The EotWash group at the University of Washington has developed a set of torsion balance instruments to probe the properties of gravity and to search for new weak forces. Current efforts focus on improved tests of the principle of equivalence, the inverse square law at short distances, and spin-coupled interactions. These experiments and prospects for the future will be discussed.

Probing Gravity and Small Forces with Torsion Balances

Blayne Heckel
University of Washington

the Eöt-Wash group:

Faculty

E.G. Adelberger
J. Gundlach
S. Fleischer
B. H.

Staff

E. Swanson

Postdocs

Andreas Kraft
K. Venkateswara

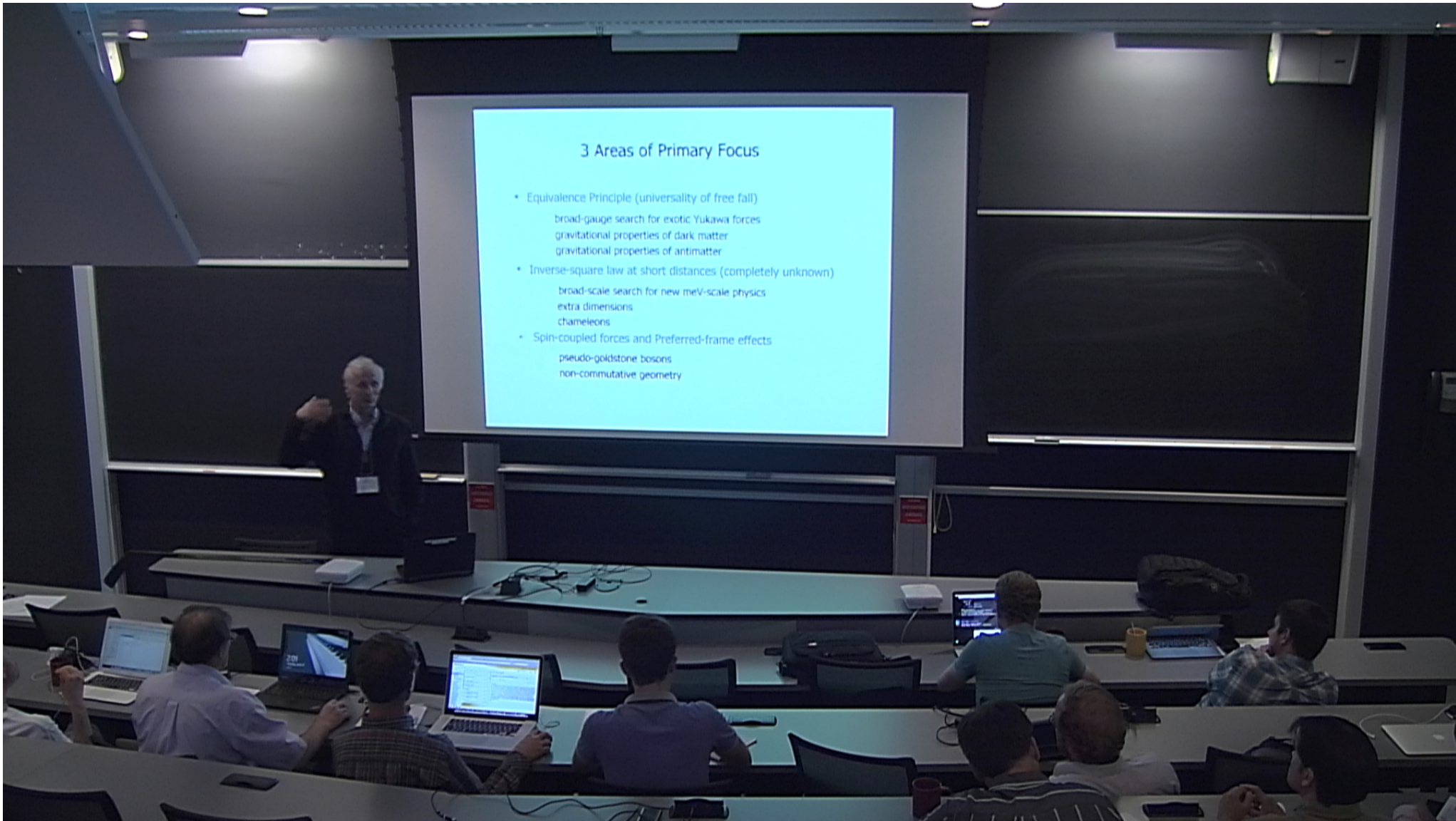
Students

C. Hagedorn
J. Lee
E. Shaw
W. Terrano
M. Turner
T. Wagner

Primary support from NSF Grant PHY0969199 with supplements from the DOE Office of Science
and to a lesser extent NASA

3 Areas of Primary Focus

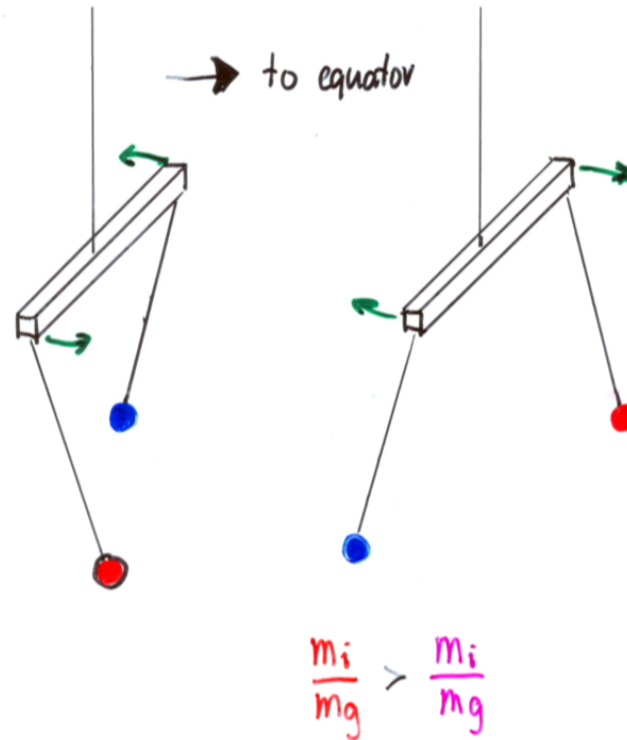
- Equivalence Principle (universality of free fall)
 - broad-gauge search for exotic Yukawa forces
 - gravitational properties of dark matter
 - gravitational properties of antimatter
- Inverse-square law at short distances (completely unknown)
 - broad-scale search for new meV-scale physics
 - extra dimensions
 - chameleons
- Spin-coupled forces and Preferred-frame effects
 - pseudo-goldstone bosons
 - non-commutative geometry



3 Areas of Primary Focus

- Equivalence Principle (universality of free fall)
 - broad-gauge search for exotic Yukawa forces
 - gravitational properties of dark matter
 - gravitational properties of antimatter
- Inverse-square law at short distances (completely unknown)
 - broad-scale search for new meV-scale physics
 - extra dimensions
 - chameleons
- Spin-coupled forces and Preferred-frame effects
 - pseudo-goldstone bosons
 - non-commutative geometry

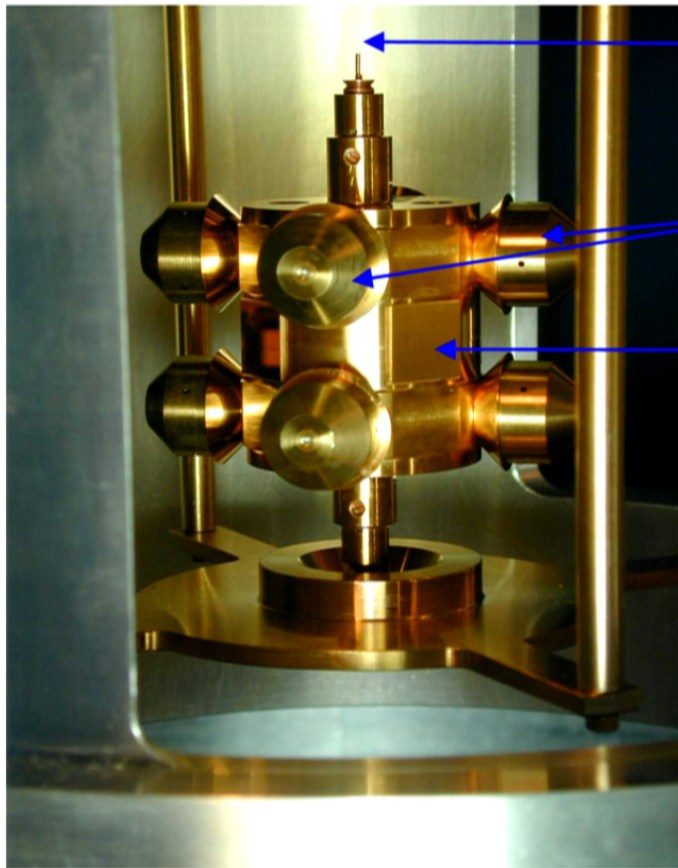
Testing the WEP by watching things fall sideways



Balance only twists if force vectors are not parallel
Down is not a unique direction if the EP is violated
or if the gravity field is not uniform

torsion pendulum of our recent EP test

T. A. Wagner et al., Class. Quant. Grav. 29, 184002 (2012)



20 μm diameter tungsten fiber

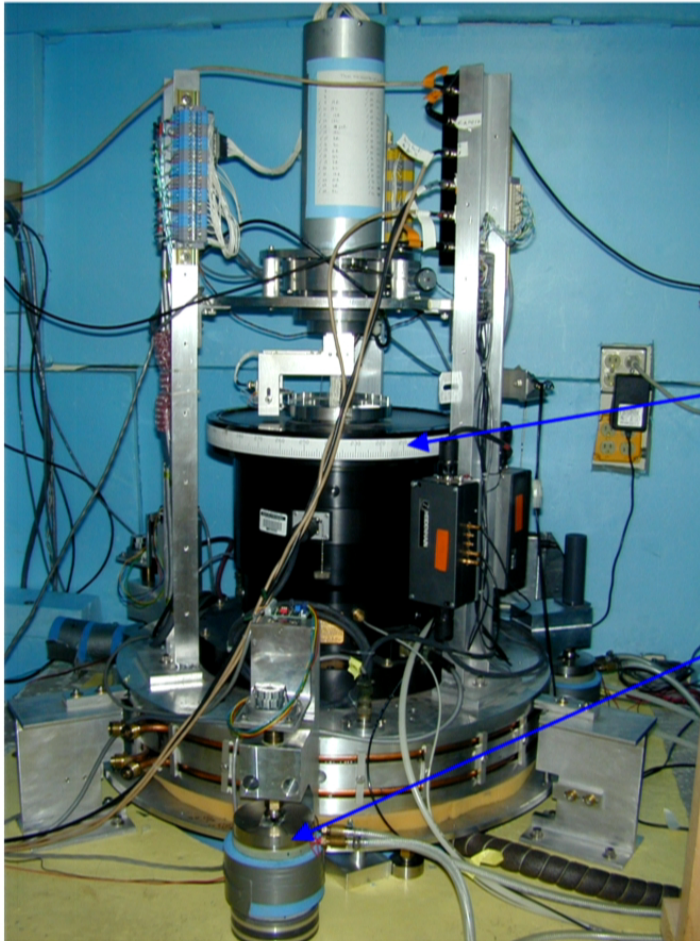
eight 4.84 g test bodies
(4 Be & 4 Ti) or (4 Be & 4 Al)

4 mirrors for measuring
pendulum twist

symmetrical design
suppresses false effects
from gravity gradients, etc.

free osc freq:	1.261 mHz
quality factor:	4000
machining tolerance:	5 μm
total mass :	70 g

Eöt-Wash torsion balance hangs from turntable that rotates at about 0.833 mHz



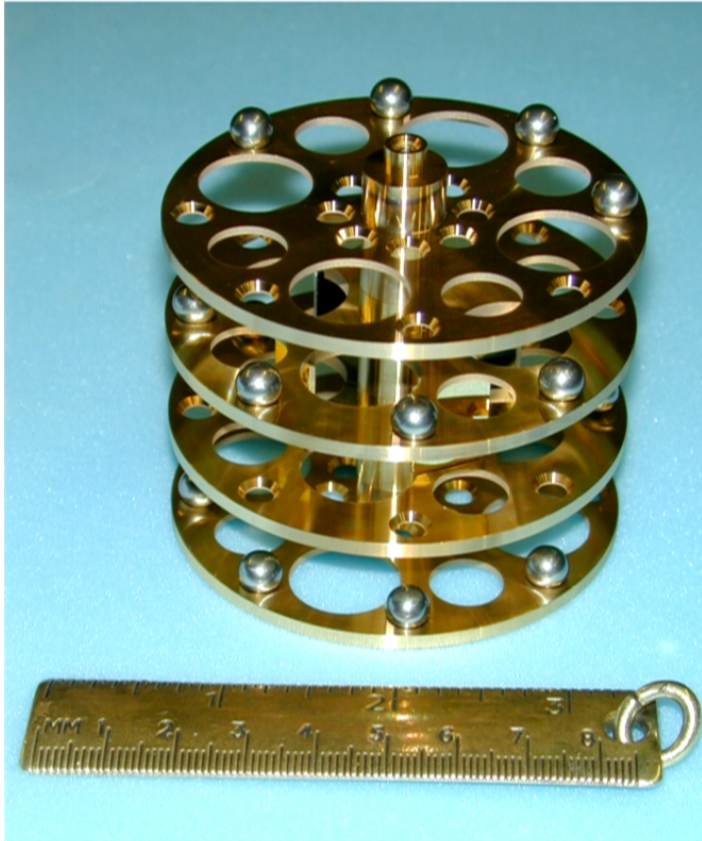
turntable requirements:

- 1) constant rotation rate
- 2) rotation axis must be along the suspension fiber

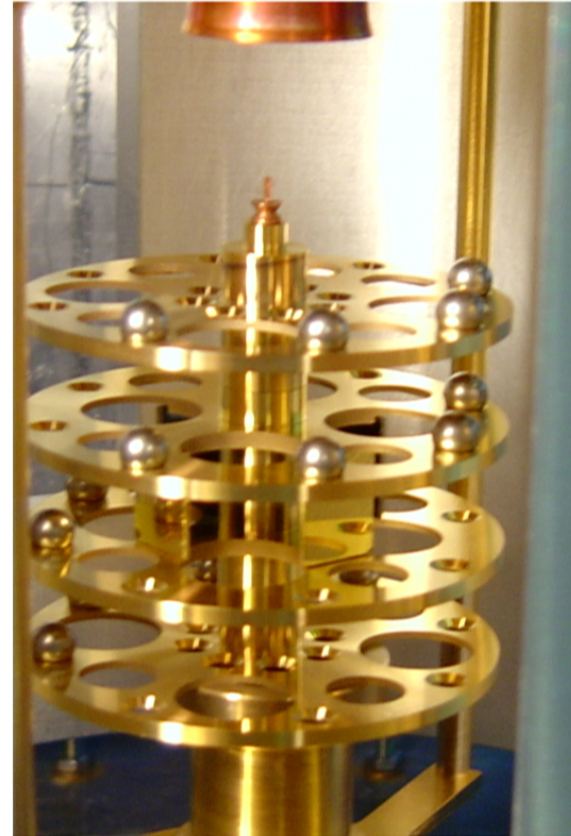
air-bearing turntable

thermal expansion feet
feedback to keep turntable
rotation axis level

gravity-gradiometer pendulums

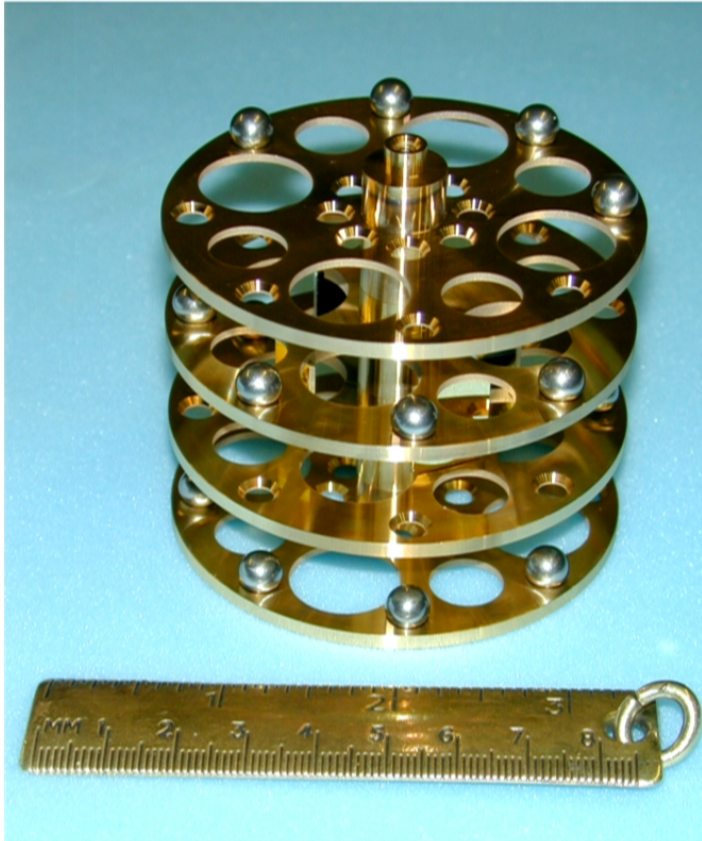


q_{41} configuration on a table

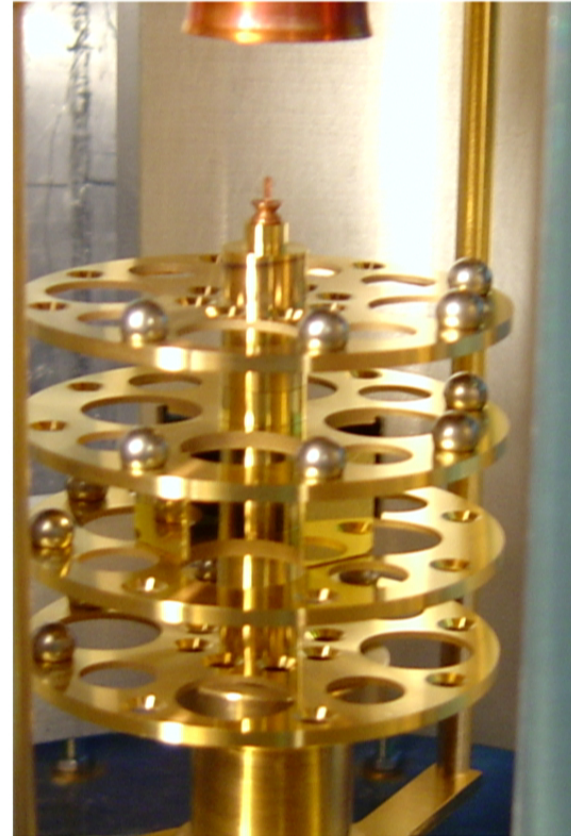


q_{21} configuration installed

gravity-gradiometer pendulums

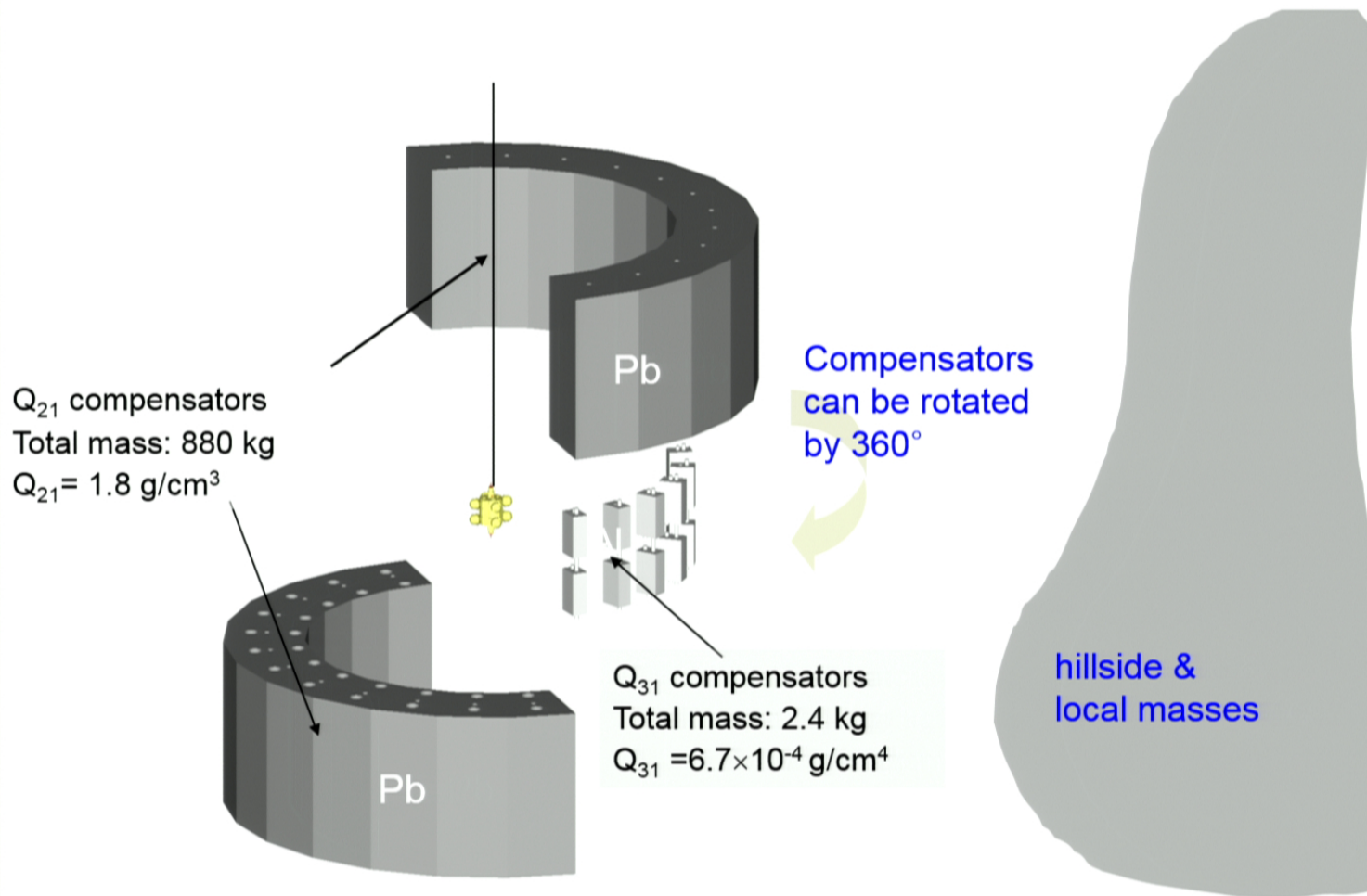


q_{41} configuration on a table



q_{21} configuration installed

gravity-gradient compensation



results with 1σ uncertainties

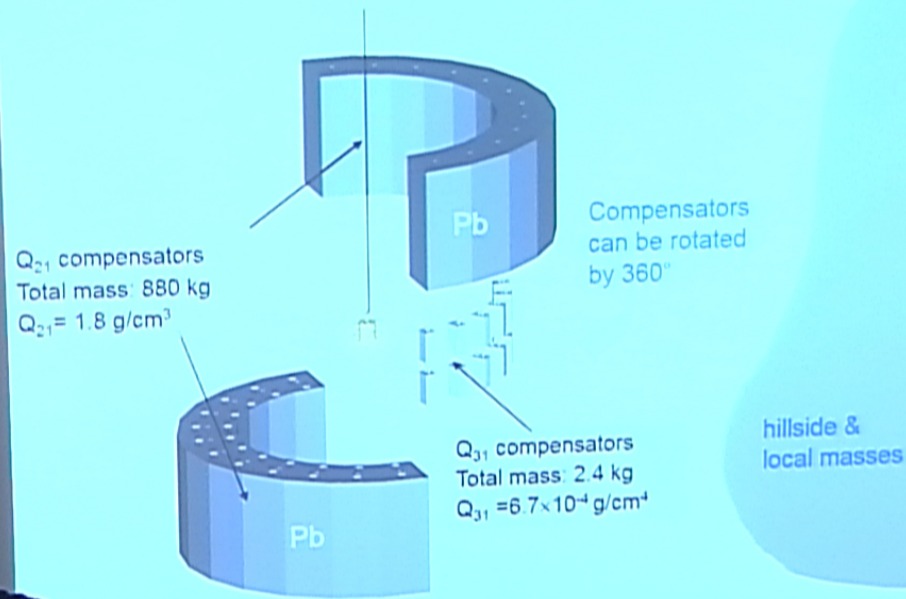
		Be-Ti	Be-Al
Δa_N	$(10^{-15} \text{ m s}^{-2})$	0.6 ± 3.1	-1.2 ± 2.2
Δa_W	$(10^{-15} \text{ m s}^{-2})$	-2.5 ± 3.5	0.2 ± 2.4
Δa_\odot	$(10^{-15} \text{ m s}^{-2})$	-1.8 ± 2.8	-3.1 ± 2.4
Δa_g	$(10^{-15} \text{ m s}^{-2})$	-2.1 ± 3.1	-1.2 ± 2.6
η_\oplus	(10^{-13})	0.3 ± 1.8	-0.7 ± 1.3
η_\odot	(10^{-13})	-3.1 ± 4.7	-5.2 ± 4.0
η_{DM}	(10^{-5})	-4.2 ± 6.2	-2.4 ± 5.2

Table 2. Error budget for the lab-fixed Be-Ti differential accelerations. Corrections were applied for gravitational gradients and tilt, only upper limits were obtained on the magnetic and temperature effects. All uncertainties are 1σ .

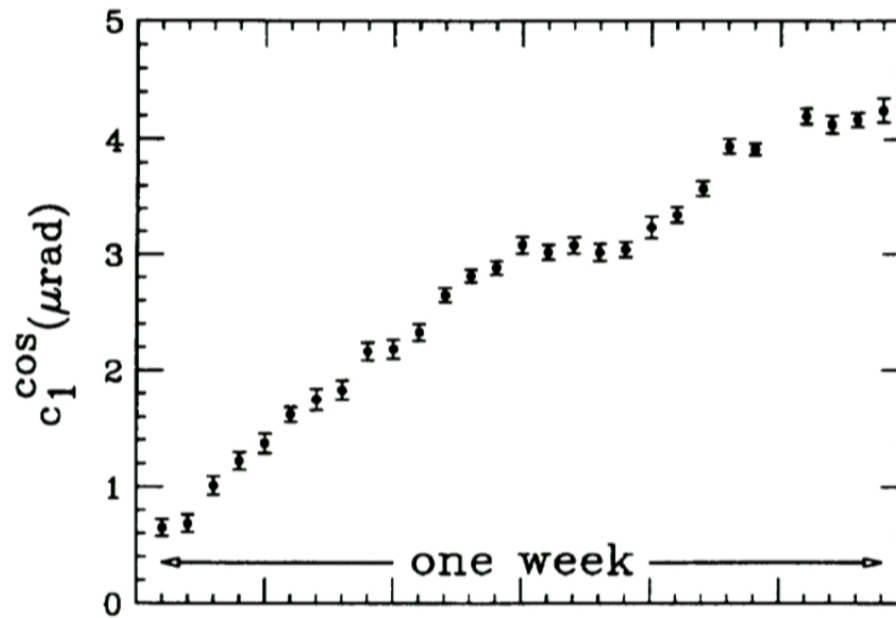
Uncertainty source	$\Delta a_{N,Be-Ti} (10^{-15} \text{ m s}^{-2})$	$\Delta a_{W,Be-Ti} (10^{-15} \text{ m s}^{-2})$
Statistical	3.3 ± 2.5	-2.4 ± 2.4
Gravity gradients	1.6 ± 0.2	0.3 ± 1.7
Tilt	1.2 ± 0.6	-0.2 ± 0.7
Magnetic	0 ± 0.3	0 ± 0.3
Temperature gradients	0 ± 1.7	0 ± 1.7

PhD project of Todd Wagner

gravity-gradient compensation



limitations on gradient cancellation



these data were taken in early November



results with 1σ uncertainties

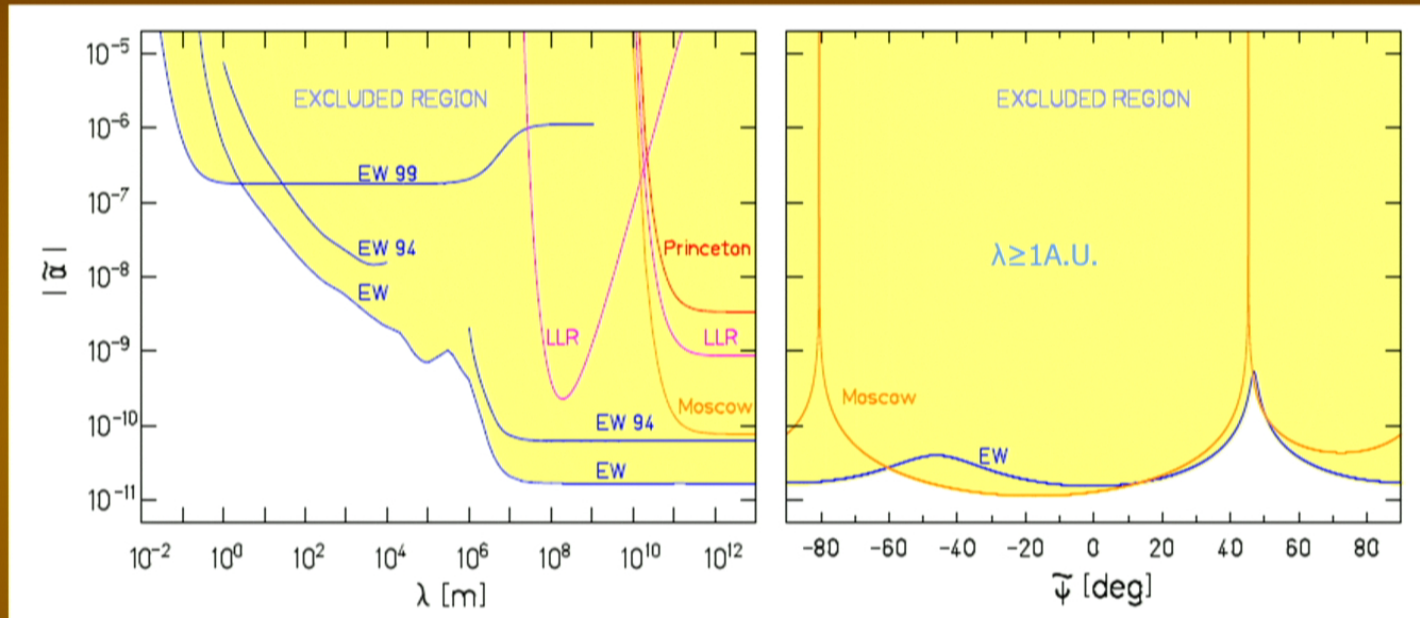
		Be-Ti	Be-Al
Δa_N	$(10^{-15} \text{ m s}^{-2})$	0.6 ± 3.1	-1.2 ± 2.2
Δa_W	$(10^{-15} \text{ m s}^{-2})$	-2.5 ± 3.5	0.2 ± 2.4
Δa_\odot	$(10^{-15} \text{ m s}^{-2})$	-1.8 ± 2.8	-3.1 ± 2.4
Δa_g	$(10^{-15} \text{ m s}^{-2})$	-2.1 ± 3.1	-1.2 ± 2.6
η_\oplus	(10^{-13})	0.3 ± 1.8	-0.7 ± 1.3
η_\odot	(10^{-13})	-3.1 ± 4.7	-5.2 ± 4.0
η_{DM}	(10^{-5})	-4.2 ± 6.2	-2.4 ± 5.2

Table 2. Error budget for the lab-fixed Be-Ti differential accelerations. Corrections were applied for gravitational gradients and tilt, only upper limits were obtained on the magnetic and temperature effects. All uncertainties are 1σ .

Uncertainty source	$\Delta a_{N,Be-Ti} (10^{-15} \text{ m s}^{-2})$	$\Delta a_{W,Be-Ti} (10^{-15} \text{ m s}^{-2})$
Statistical	3.3 ± 2.5	-2.4 ± 2.4
Gravity gradients	1.6 ± 0.2	0.3 ± 1.7
Tilt	1.2 ± 0.6	-0.2 ± 0.7
Magnetic	0 ± 0.3	0 ± 0.3
Temperature gradients	0 ± 1.7	0 ± 1.7

PhD project of Todd Wagner

95% confidence level exclusion plot for interactions coupled to B-L



Yukawa attractor integral based on:

$0.5m < \lambda < 5m$

$1m < \lambda < 50km$

$5km < \lambda < 1000km$

$1000km < \lambda < 10000km$

lab building and its major contents

topography

USGS subsurface density model

PREM earth model

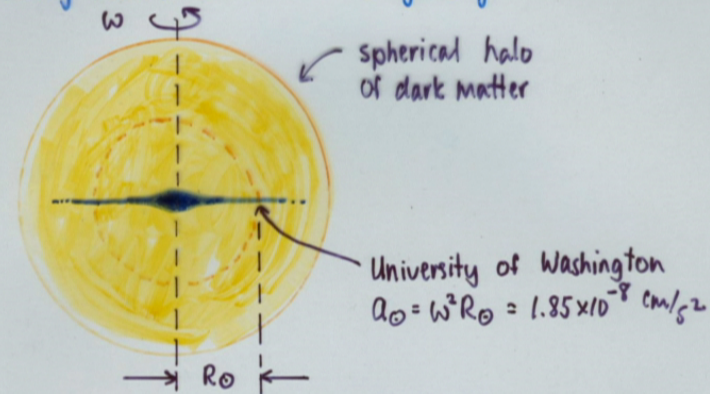
T. A. Wagner et al., Class. Quant. Grav. 29, 184002 (2012)

Is gravity the only long-range force between dark and luminous matter?

Could there be a long-range scalar interaction that couples dark-matter & standard-model particles?

OUR EXPERIMENTAL STRATEGY G.W. STUBBS

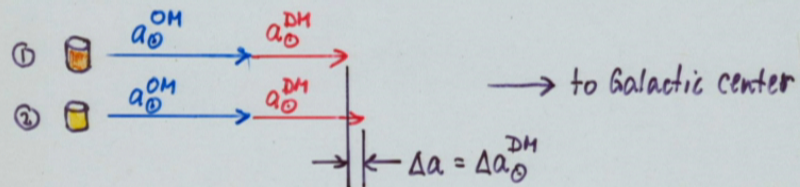
check universality of free fall for different materials falling toward center of our galaxy.



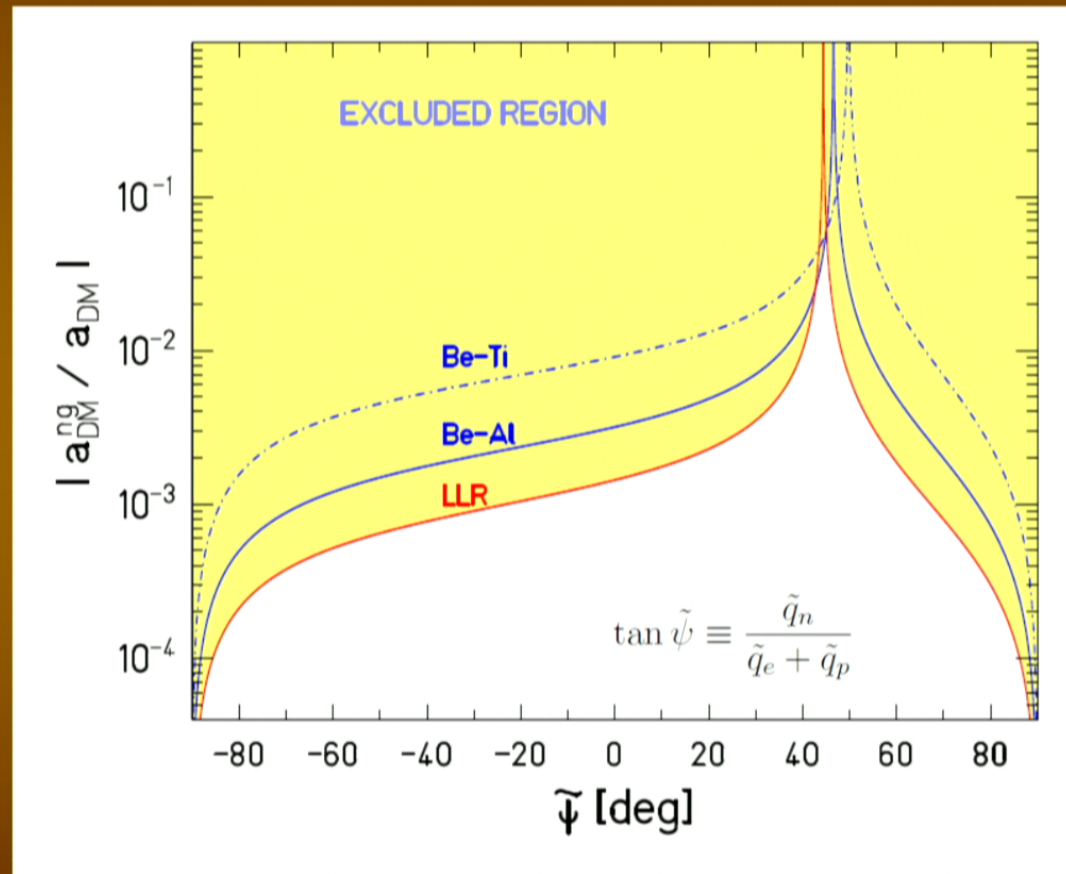
although 90% of galaxy mass is thought to be DM much of it lies outside R_{\odot} , so

$$a_{\odot}^{\text{DM}} = 25\text{-}30\% a_{\odot} \Rightarrow a_{\odot}^{\text{DM}} \approx 5 \times 10^{-9} \text{ cm/s}^2$$

We can make interesting statement about non-grav. component of a_{\odot}^{DM} if we can detect differential accels. with a sensitivity of $10^{-3} a_{\odot}^{\text{DM}} \approx 5 \times 10^{-12} \text{ cm/s}^2$



95% confidence limits on non-gravitational acceleration of hydrogen by galactic dark matter



at most 6% of the acceleration can be non-gravitational

Current status of our EP tests

- Hydrogen rich test bodies (CH_2)
- Continuous monitoring of Q_{21} gravitational gradient field
- Higher precision tilt sensors
- Quartz torsion fiber (higher Q than tungsten)

We anticipate a factor of 3 improvement in sensitivity with test body pairs with larger “charge” differences.

motivations for sub-millimeter tests of the inverse-square law

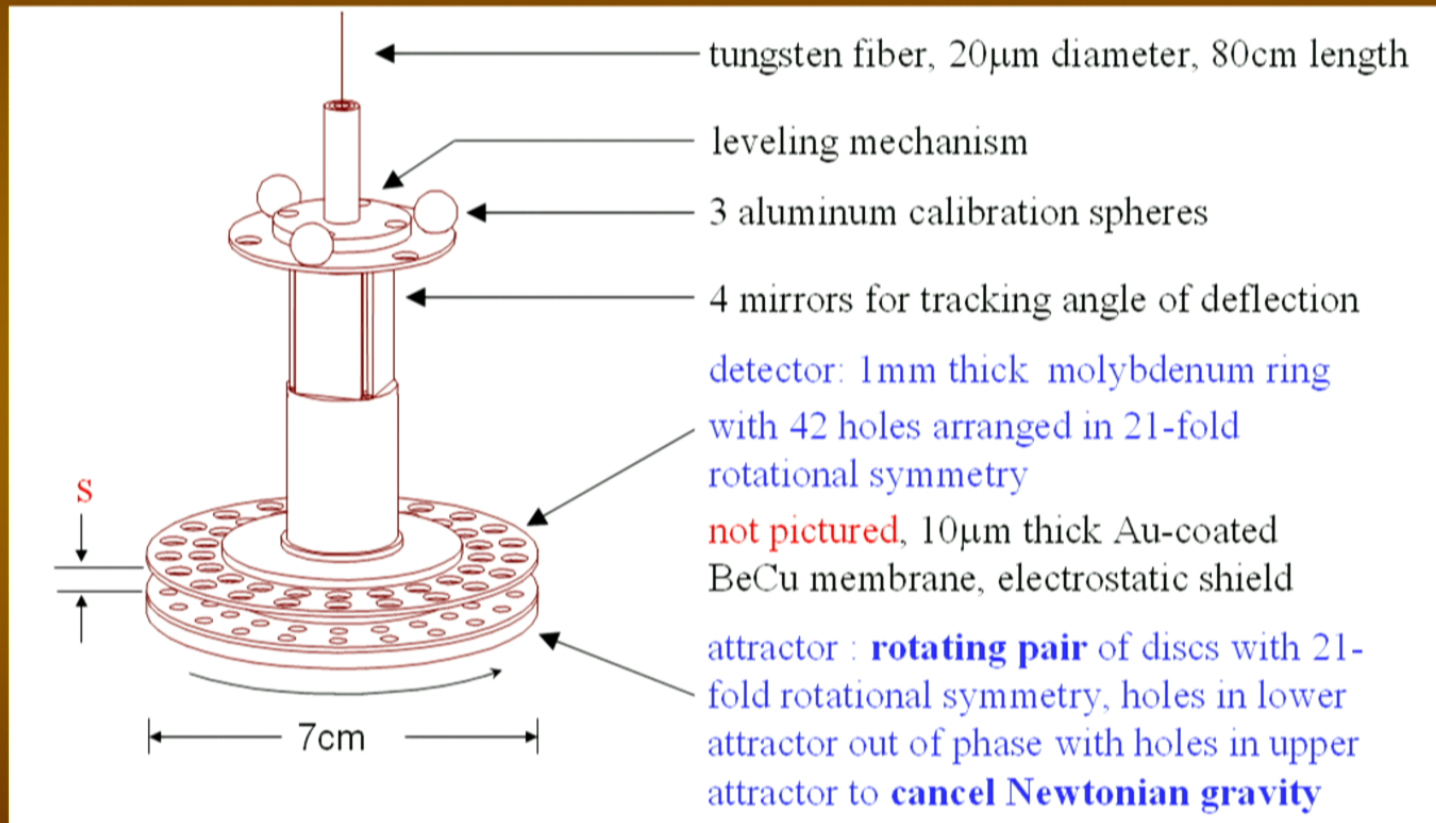
- explore an untested regime
- probe the dark-energy length scale

$$\rho_d \approx 3.8 \text{ keV/cm}^3$$

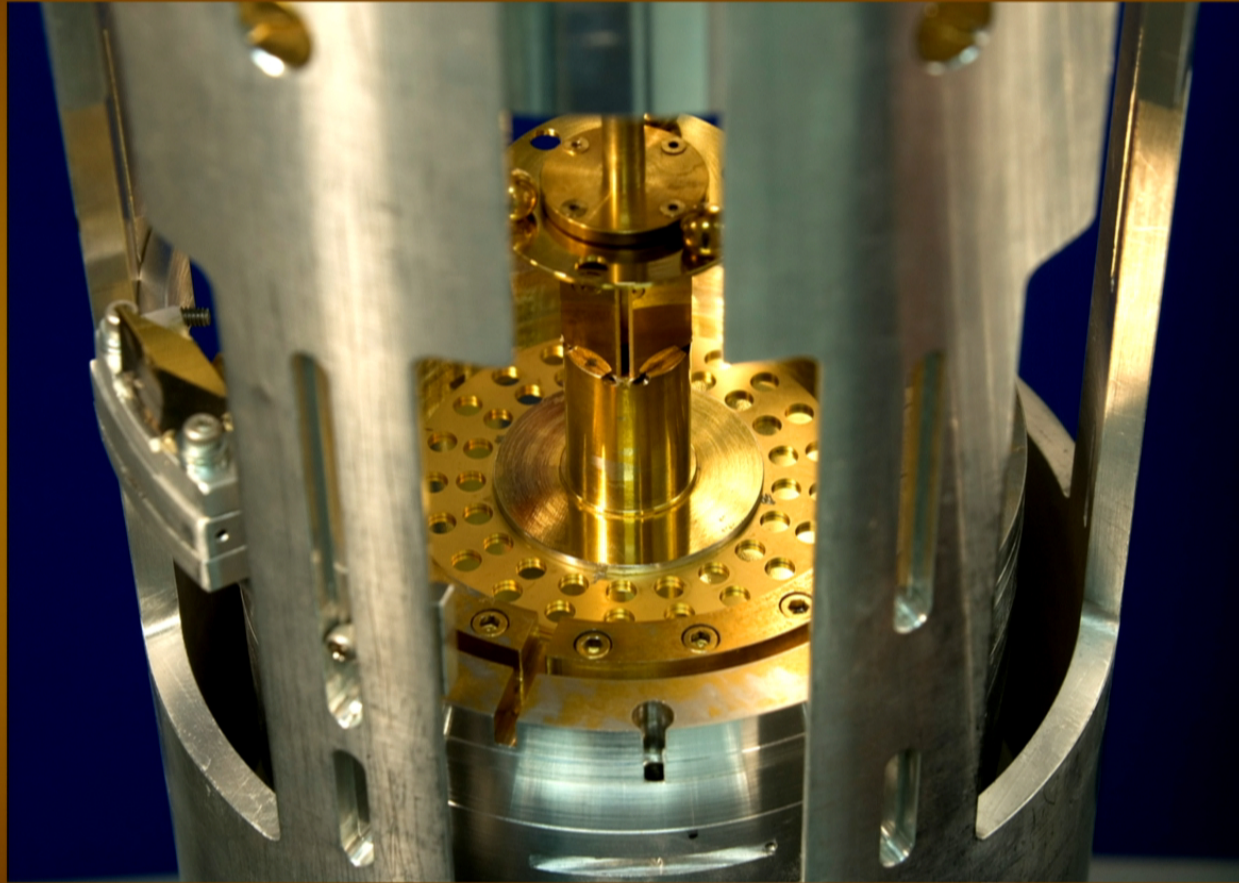
$$\lambda_d = \sqrt[4]{\hbar c / \rho_d} \approx 85 \text{ } \mu\text{m}$$

- search for proposed new phenomena
 - large extra dimensions: why is gravity so weak?
 - chameleons: what happened to the stringy scalars?

the 42-hole ISL pendulum



D.J. Kapner et al., PRL 98, 021101(2007)

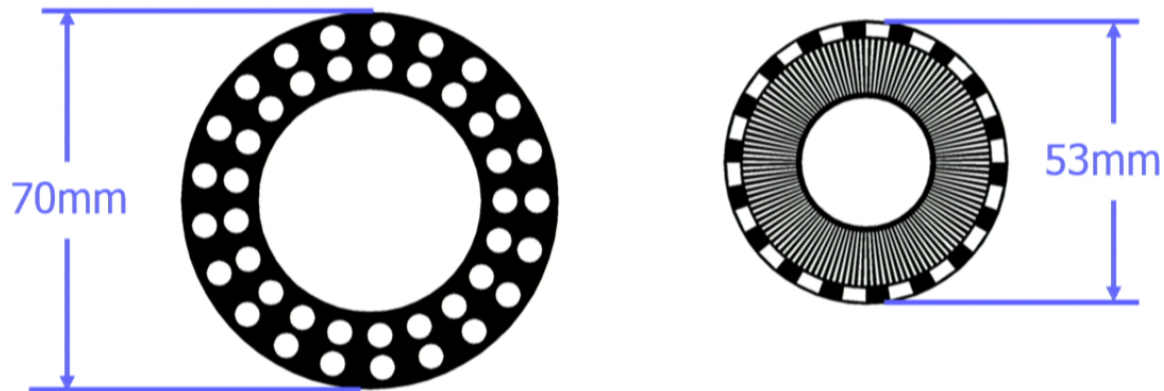


Mary Levin photo

our next-generation short-range instrument

$$N_Y = \frac{\partial E_Y}{\partial \phi} \approx 2\pi\alpha G\lambda^3 \rho_D \rho_A \frac{\Delta A}{\Delta \phi} \exp\left(-\frac{s}{\lambda}\right)$$

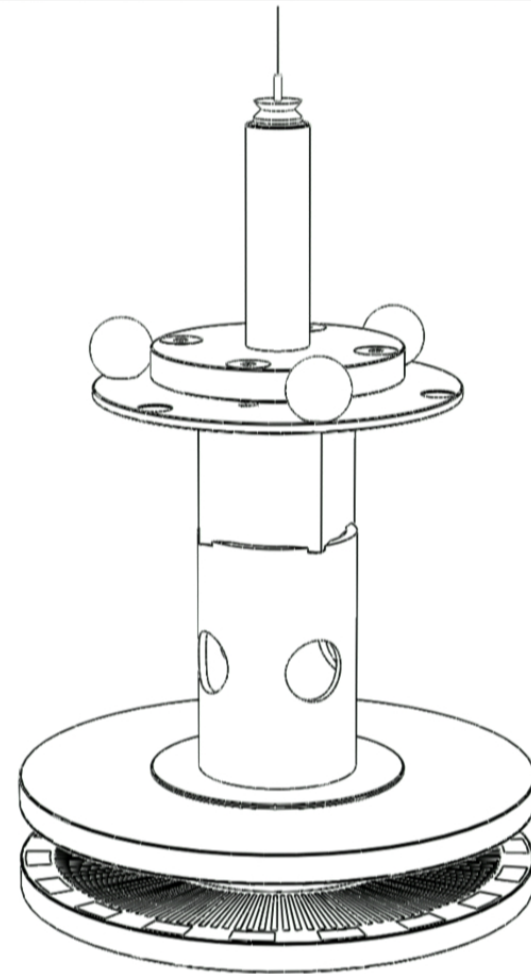
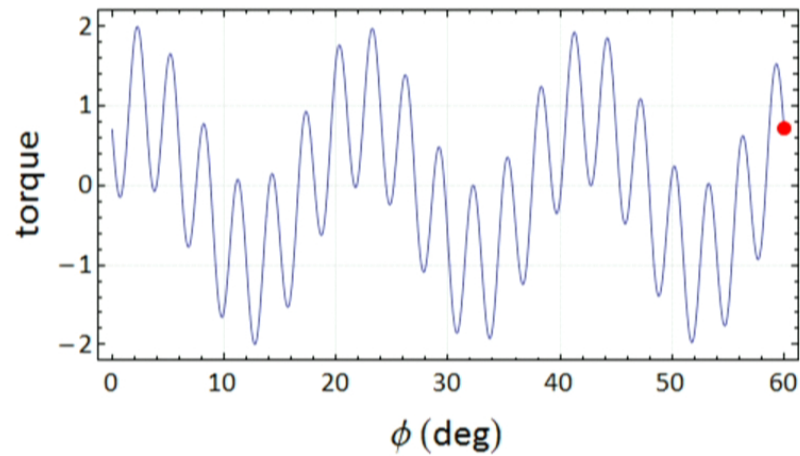
	<u>Kapner et al.</u>	<u>Cook et al.</u>
symmetry:	21	120 & 18
material:	molybdenum (10.3 g/cm ³)	tungsten (19.3 g/cm ³)
thickness:	1 mm	0.05 mm
attractor:	2 pieces	1 piece



::

Cook et al.'s Experiment

::

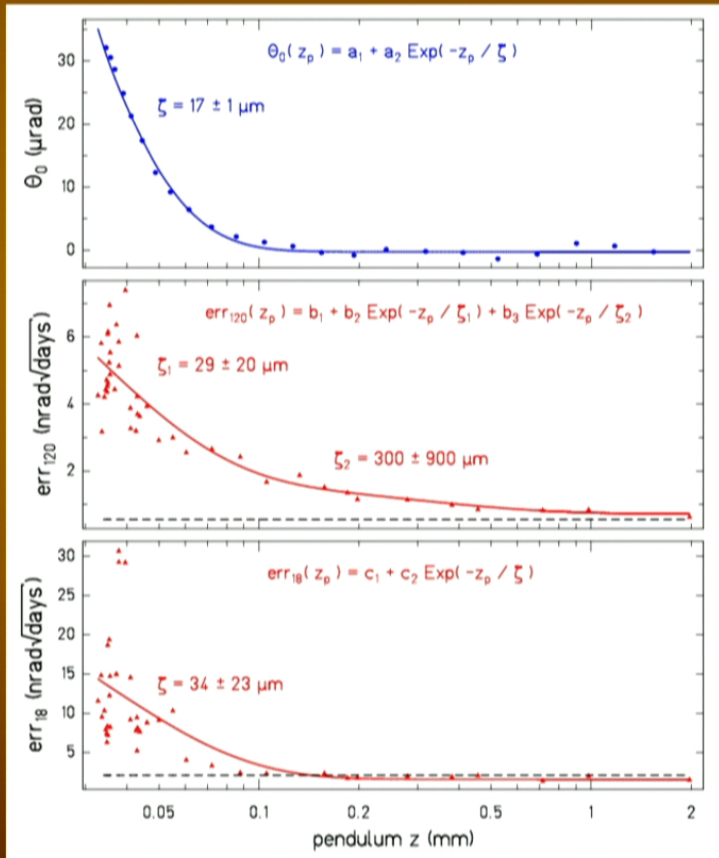


simulation is speeded up by
factor of ≈ 1000

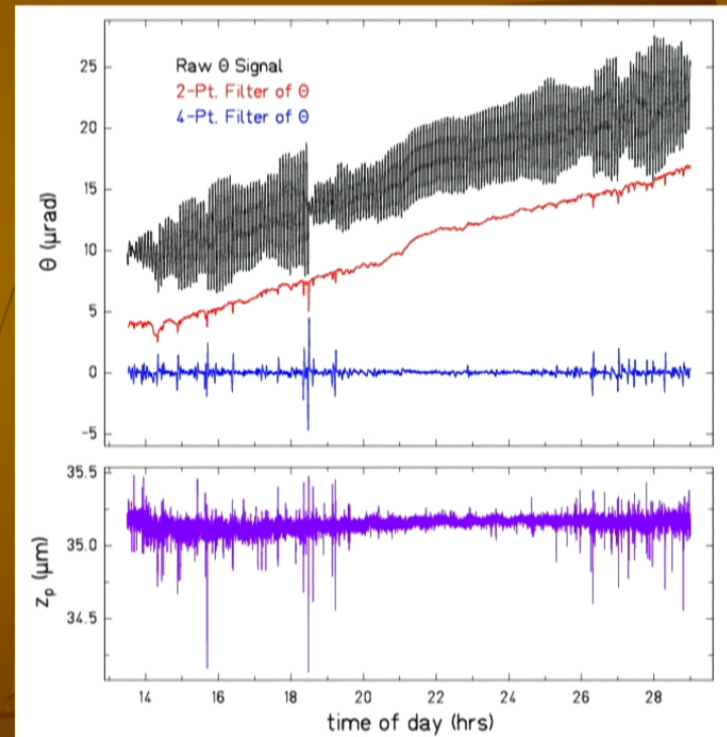
Ted Cook | tedcook@gmail.com | www.npl.washington.edu/eotwash

patch fields

patch field potential minimum
not aligned with fiber minimum



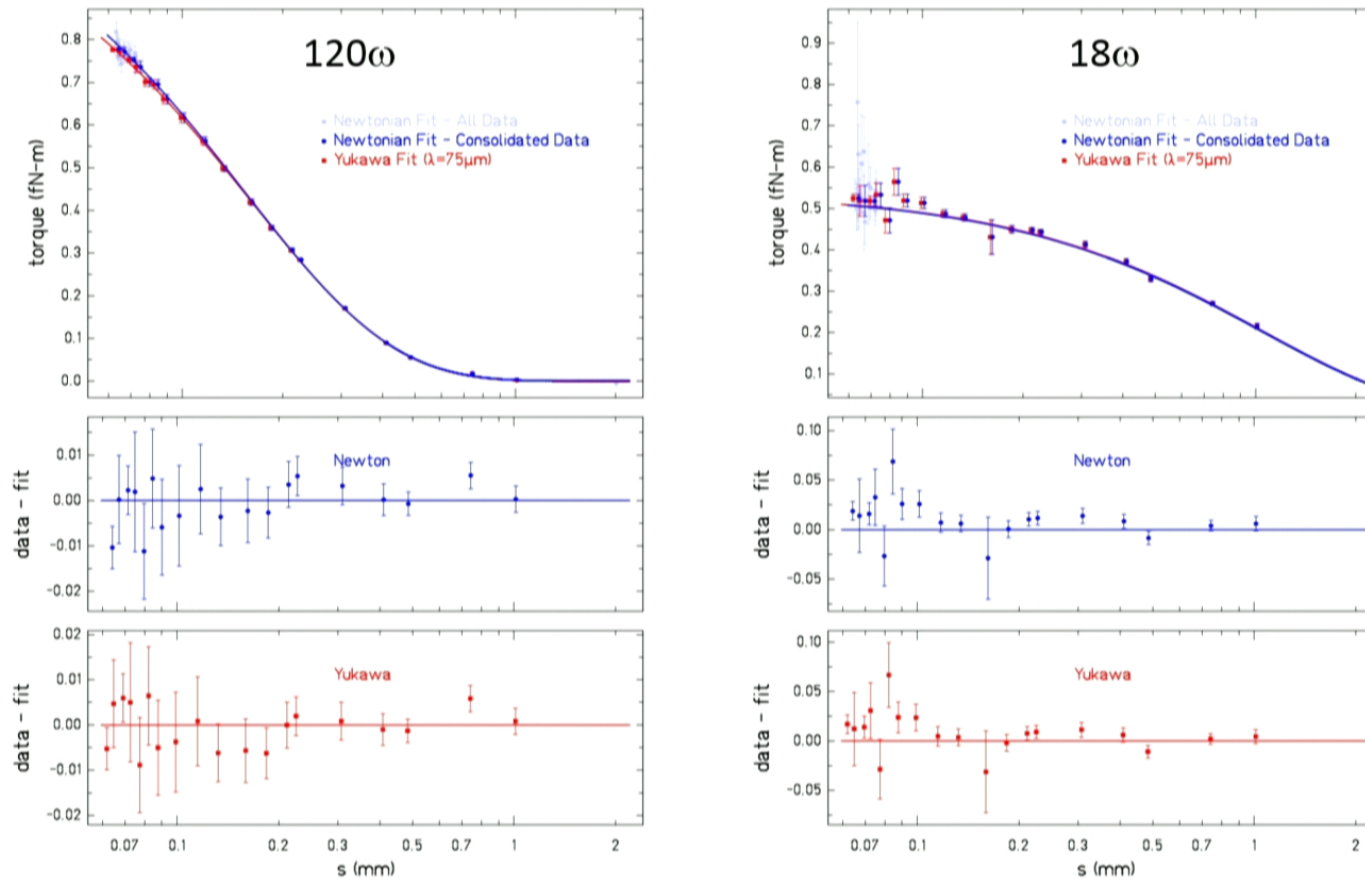
vibrations



(attractor not turning)

Data Fit

$$\lambda = 75 \mu\text{m}; \alpha = -0.16 \pm 0.05$$

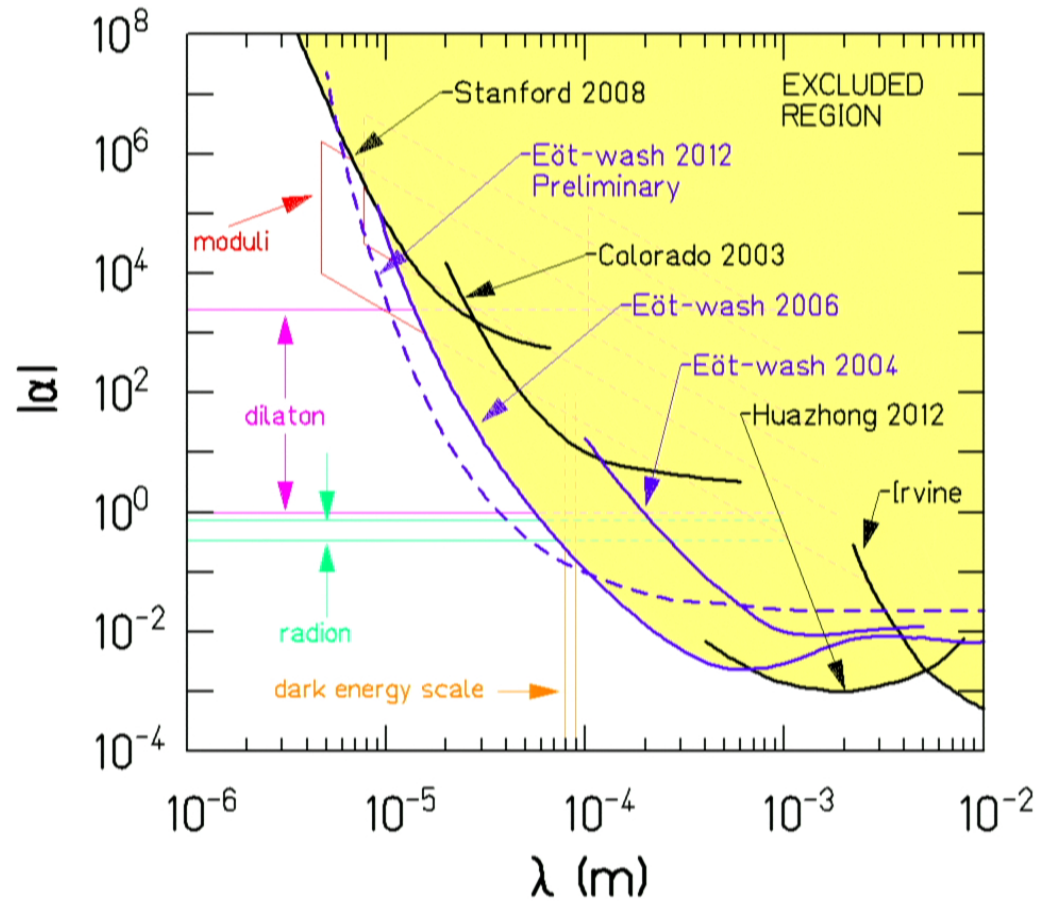


Ted Cook | tedcook@gmail.com | www.npl.washington.edu/eotwash

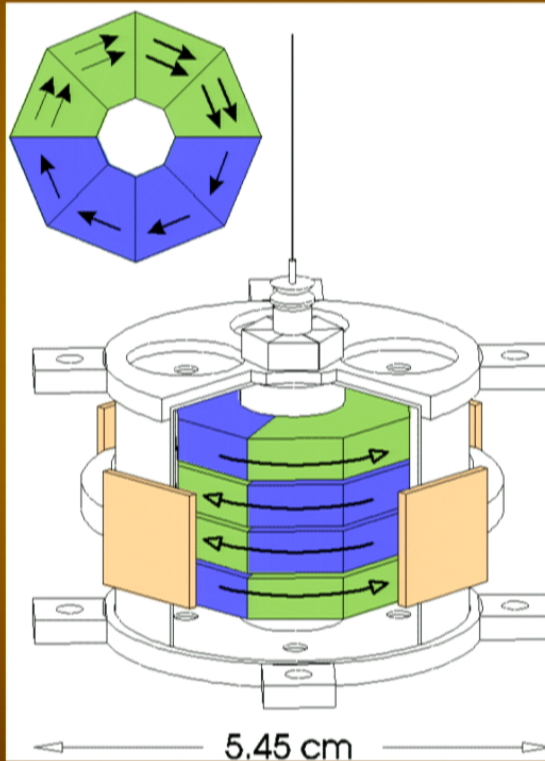
Cook's preliminary 95% C.L. results

order of magnitude higher sensitivity below 40 μm :

We hope to do significantly better in the an improved iteration of Cook's device

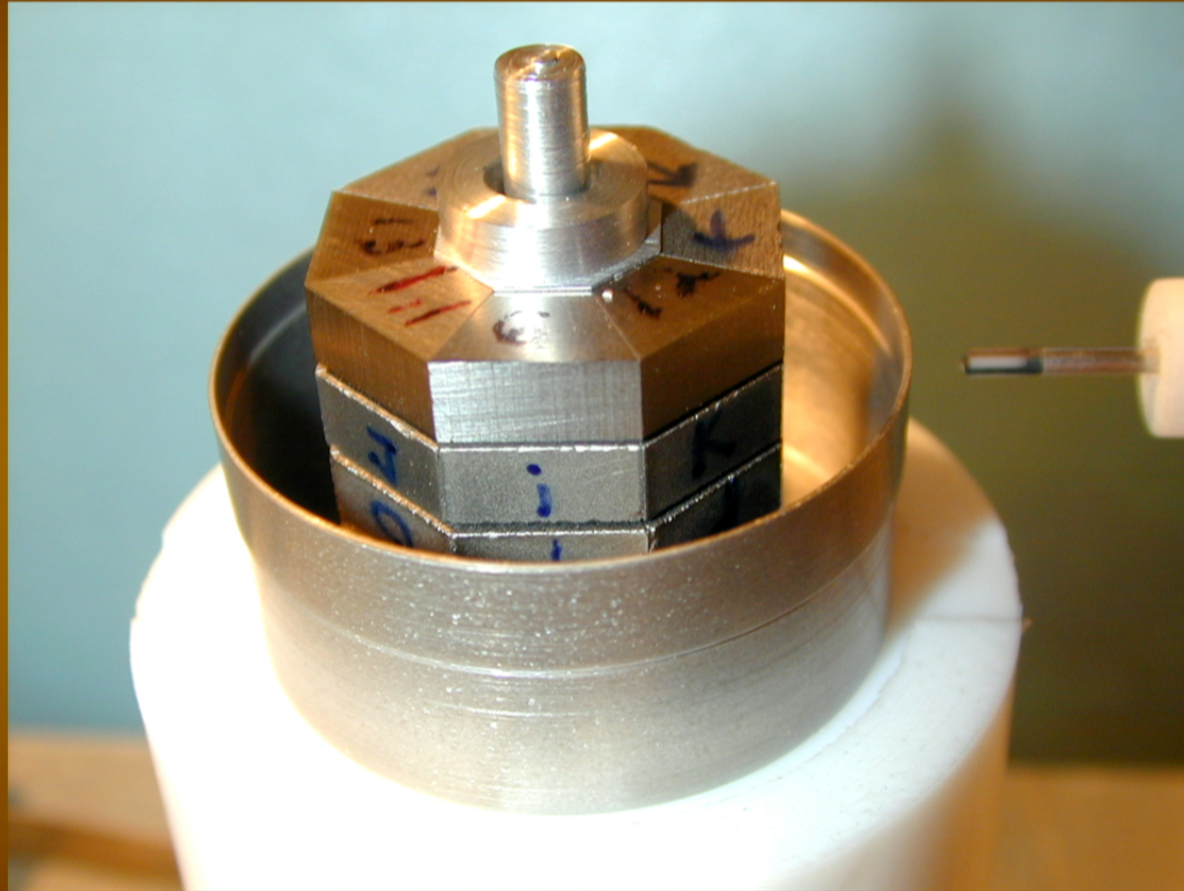


the Eöt-Wash spin pendulum



- 9.8×10^{22} polarized electrons
- negligible mass asymmetry
- negligible composition asymmetry
- flux of B confined within magnets
- negligible external B field
- Alnico: all B comes from electron spin: spins point opposite to B
- SmCo_5 : Sm 3^+ ion has spin pointing along total B and its spin B field is nearly canceled by its orbital B field--so B of SmCo_5 comes almost entirely from the Co's electron spins
- therefore the spins of Alnico and Co cancel and pendulum's net spin comes from the Sm and $J = \uparrow \downarrow S$

measuring the spin pendulum's stray B field



B inside = 9.6 ± 0.2 kG

B outside \approx few mG

Our sensitivity

- our upper limit on the energy required to invert an electron spin about an arbitrary axis fixed in inertial space is $\sim 10^{-21}$ eV
- this is comparable to the electrostatic energy of two electrons separated by ~ 10 astronomical units

constraint on non-commutative geometry

If electrons are point-like up to $\Lambda = 1 \text{ TeV}$, this corresponds to a minimum observable area

$$|\theta^{\mu\nu}| \leq 6 \times 10^{-58} \text{ m}^2$$

$6 \times 10^{-58} \text{ m}^2 \sim (10^6 L_p)^2$
where L_p is the Planck Length = $\sqrt{\hbar G/c^3} = 1.6 \times 10^{-35} \text{ m}$

or $\sim (10^3 L_U)^2$
where L_U is the GUT scale = $\hbar c / 10^{16} \text{ GeV}$

Recent limits on short range spin-spin coupling

$$V_{ee}(\vec{r}) = g_P^2 \frac{\hbar^2}{16\pi m_e^2 c^2} (\vec{\sigma}_1 \cdot \vec{\nabla}_1)(\vec{\sigma}_2 \cdot \vec{\nabla}_2) \frac{e^{-r/\lambda}}{r}$$

