

Title: Introduction, Scope and goals of the workshop, Open questions in the field

Date: Jun 16, 2014 09:10 AM

URL: <http://pirsa.org/14050149>

Abstract:

New Ideas in Low-Energy Tests of Fundamental Physics

General Thoughts

Derek F. Jackson Kimball
California State University

Origins...

Origins...

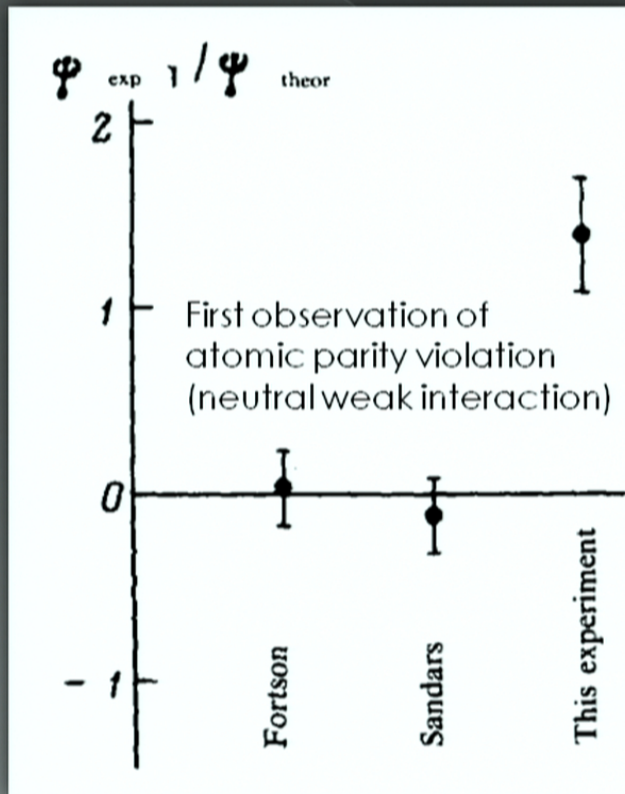


Max Zolotorev



Eugene Commins

Origins...



Barkov & Zolotorev, JETP Lett. **27**, 357 (1978).

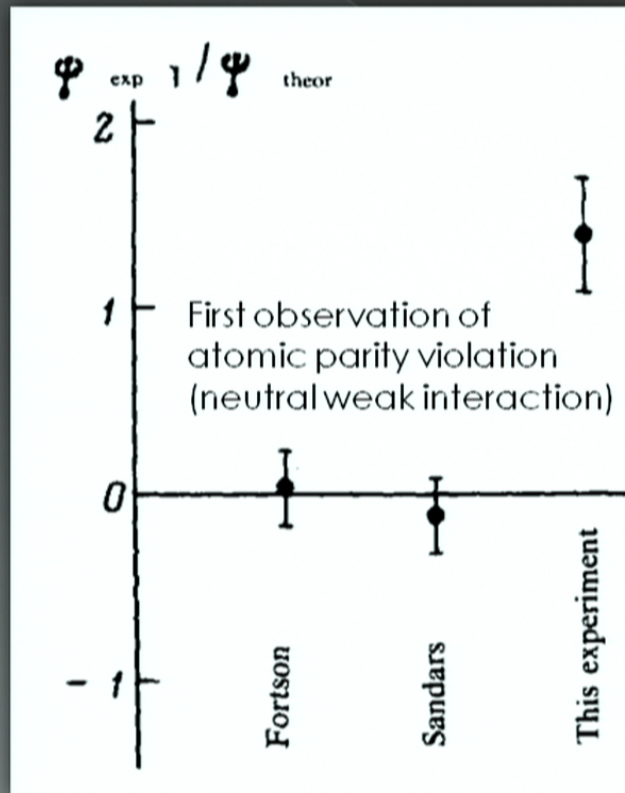


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Max Zolotorev

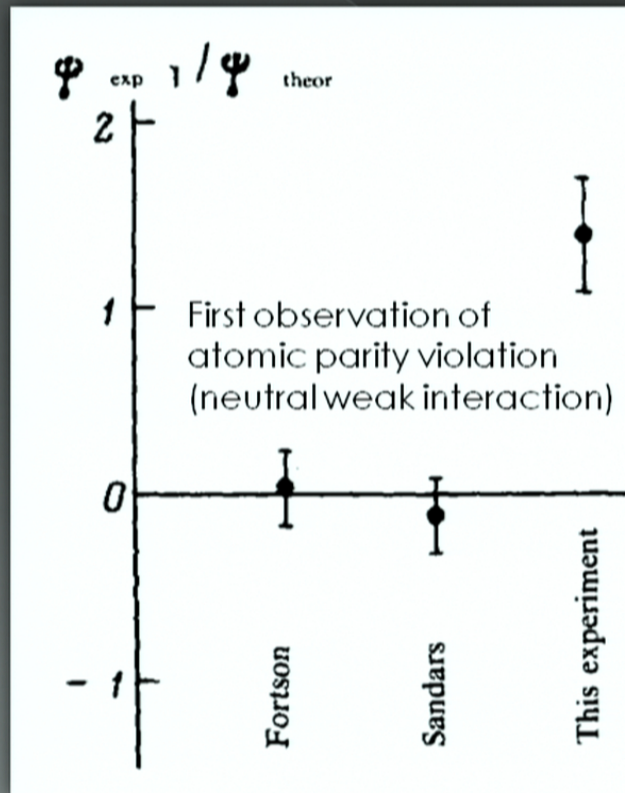


Eugene Commins



Dmitry Budker

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Dmitry Budker

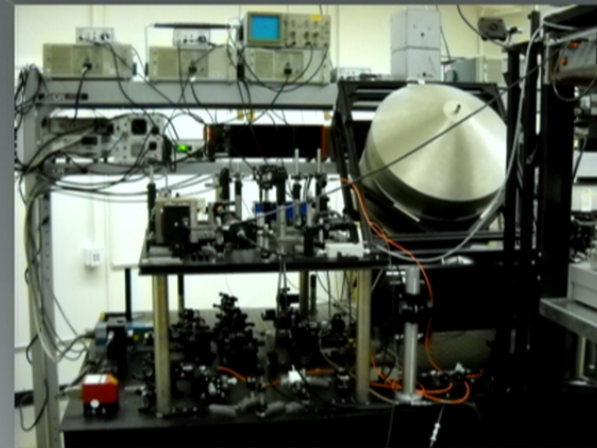


me

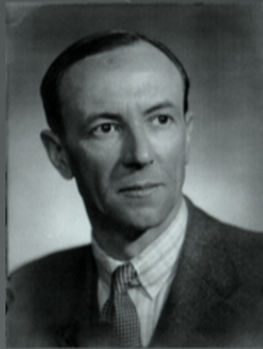
Interface between theorists and experimentalists

Interface between theorists and experimentalists

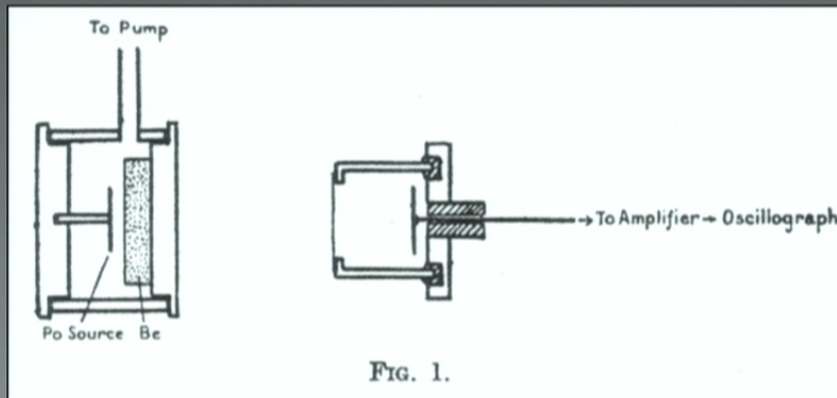
$$\mathcal{L}_\phi = |\partial_\mu \phi|^2 - V(\phi);$$
$$V(\phi) = \frac{\lambda}{s_0^{2N-4}} |2^{N/2} \phi^N - s_0^N|^2,$$



The state of particle physics?



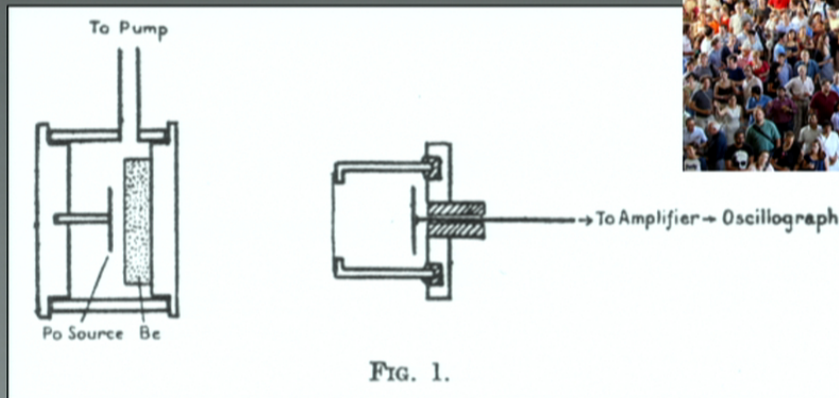
James Chadwick
discovered neutron 1932



The state of particle physics?



James Chadwick
discovered neutron 1932



CERN team
discovered Higgs 2013

Confusion is preface to
discovery!



Motivated by great mysteries

- ◉ Missing CP-violation for baryogenesis,
- ◉ Strong CP problem,
- ◉ Dark matter,
- ◉ Dark energy,
- ◉ etc.

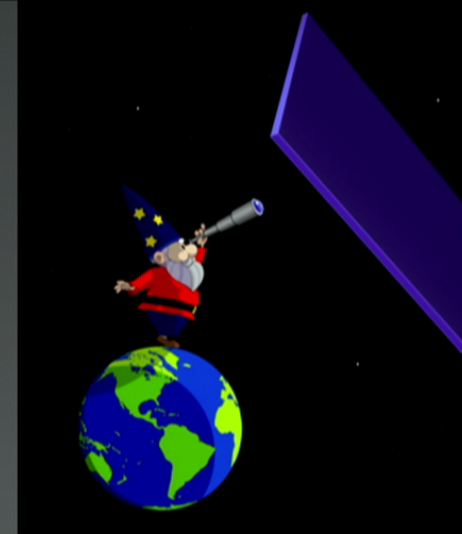
Finite set of parameters to explore with atomic experiments?

- Elementary fermions have only monopole (charge/mass) and dipole (spin) moments.
- Elementary particles can be in relative motion.
- Quantum field theory, with general assumptions, allows only a finite number of possible spin-0 and spin-1 potentials.

A new direction in low-energy tests of fundamental physics...

- ◉ Instead of searching for local or universal signatures of new physics in atomic systems, search for transient signatures of astrophysical origin...
- ◉ But where to search, how to search, and what to search for? Motivation and connections between different measurements come from theory.

Global Network of Optical Magnetometers to search for Exotic Physics (GNOME)



PRL **110**, 021803 (2013)

PHYSICAL REVIEW LETTERS

week ending
11 JANUARY 2013

Detecting Domain Walls of Axionlike Models Using Terrestrial Experiments

M. Pospelov,^{1,2} S. Pustelny,^{3,4,*} M. P. Ledbetter,⁴ D. F. Jackson Kimball,⁵ W. Gawlik,³ and D. Budker^{4,6}

¹*Department of Physics and Astronomy, University of Victoria, Victoria, British Columbia V8P 1A1, Canada*

²*Perimeter Institute for Theoretical Physics, Waterloo, Ontario N2J 2W9, Canada*

³*Institute of Physics, Jagiellonian University, Reymonta 4, 30-059 Kraków, Poland*

⁴*Department of Physics, University of California at Berkeley, Berkeley, California 94720-7300, USA*

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⁶*Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA*

(Received 23 July 2012; revised manuscript received 4 November 2012; published 11 January 2013)

Cosmic Axion Spin Precession Experiment (CASPER)



PHYSICAL REVIEW X **4**, 021030 (2014)

Proposal for a Cosmic Axion Spin Precession Experiment (CASPER)

Dmitry Budker,^{1,5} Peter W. Graham,² Micah Ledbetter,³ Surjeet Rajendran,² and Alexander O. Sushkov⁴

¹*Department of Physics, University of California, Berkeley, California 94720, USA
and Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA*

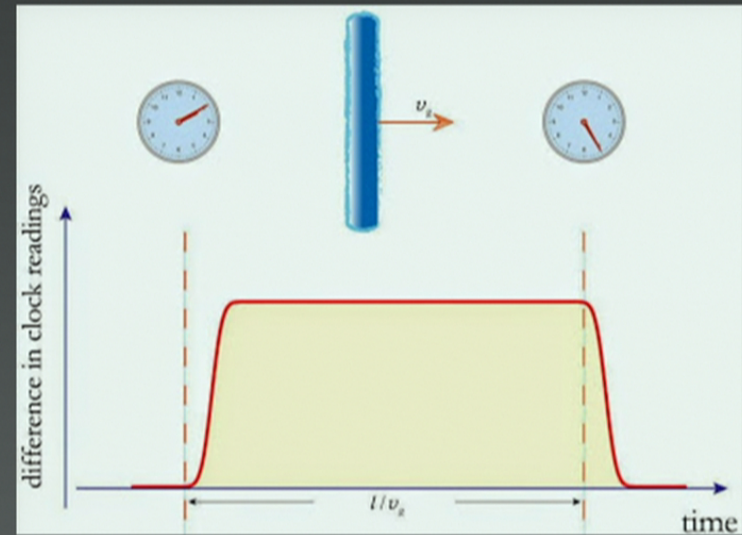
²*Department of Physics, Stanford Institute for Theoretical Physics, Stanford University,
Stanford, California 94305, USA*

³*AOSense, 767 North Mary Avenue, Sunnyvale, California 94085-2909, USA*

⁴*Department of Physics and Department of Chemistry and Chemical Biology, Harvard University,
Cambridge, Massachusetts 02138, USA*

⁵*Helmholtz Institute Mainz, Johannes Gutenberg University, 55099 Mainz, Germany
(Received 9 July 2013; published 19 May 2014)*

Network of atomic clocks



Hunting for topological dark matter with atomic clocks

A. Derevianko

Department of Physics, University of Nevada, Reno, NV 89557, USA

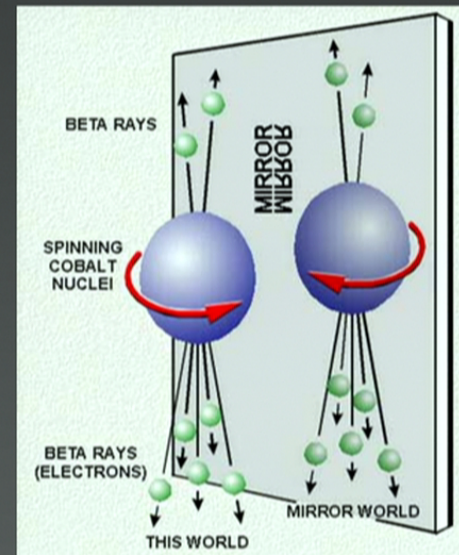
M. Pospelov

*Department of Physics and Astronomy, University of Victoria, Victoria, BC V8P 1A1, Canada and
Perimeter Institute for Theoretical Physics, Waterloo, ON N2J 2W9, Canada*

(Dated: November 7, 2013)

Derevianko and Pospelov, arXiv:1311.1244 (2013).

Astrophysically induced parity violating effects



Axion-induced effects in atoms, molecules and nuclei: parity non-conservation, anapole moments, electric dipole moments, and spin-gravity and spin-axion momentum couplings

Y. V. Stadnik and V. V. Flambaum
School of Physics, University of New South Wales, Sydney 2052, Australia
(Dated: December 24, 2013)

Stadnik and Flambaum, arXiv:1312.6667 (2013).

Transient violations of physical laws?

- Broadly speaking, it seems possible that many physical laws could be subject to transient, small violations from some unknown astrophysical effects, and violations could easily have escaped detection!
- But given difficulty of experiments and our finite resources, what are the most promising laws to test? What are the most promising systems with which to perform tests?

OPEN QUESTIONS



Dmitry Budker

**Helmholtz Institute Mainz
Johannes Gutenberg U.**

**UC Berkeley Physics
NSD LBNL**

Perimeter Institute, June 2014

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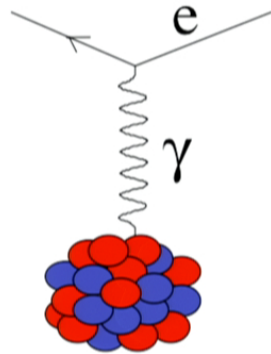
Perimeter Institute, June 2014

THE PLAN ...

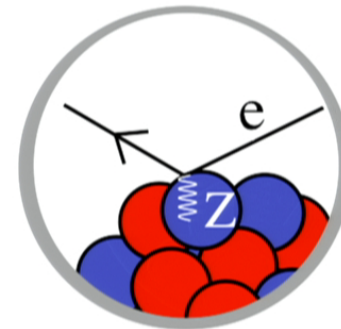
- Gravity tests with α ?
- Does gravity respect parity ?
- Cosmic parity violation ?
- Antimatter tests beyond \bar{g} ?
- Open GNOME & CASPEr Qns

ATOMIC PARITY VIOLATION

- Main Source: Z exchange



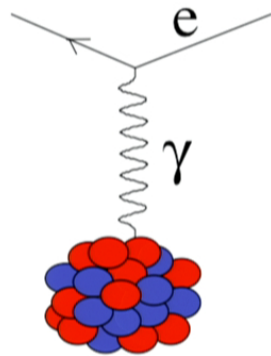
Electromagnetic
interaction
(conserves parity)



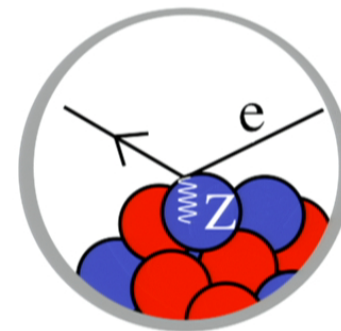
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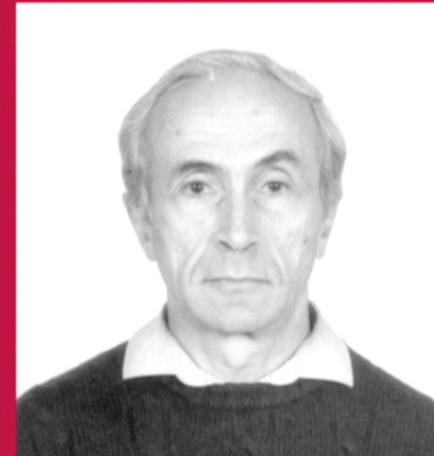
- P-odd, T-even correlation $\vec{\sigma} \cdot \vec{p}$

Dy

- ▣ Ideal APV amplifier? Dzuba, Flambaum, Khriplovich (1986)
 - Fully degenerate opposite-parity levels
 - Large Z^3 ($Z=66$)

Dy

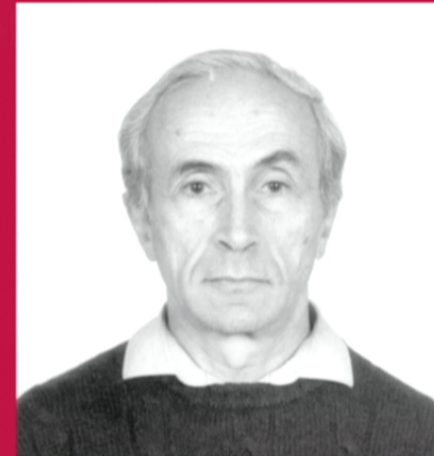
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Prof. Iosif B. Khriplovich

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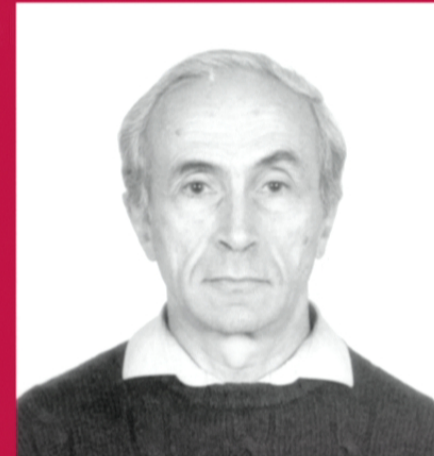
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 - Large Z^3 ($Z=66$)
- ▣ Also
 - Many stable isotopes: $A=164-156$
 - Two $I=5/2$ isotopes (anapole)
- Degenerate levels require quantum juggling to measure APV



Prof. Iosif B. Khriplovich

The parity violation experiment in Dy

PHYSICAL REVIEW A

VOLUME 56, NUMBER 5

NOVEMBER 1997

Search for parity nonconservation in atomic dysprosium

A. T. Nguyen,¹ D. Budker,^{1,2} D. DeMille,^{1,*} and M. Zolotarev³

¹*Physics Department, University of California, Berkeley, California 94720-7300*

²*Nuclear Science Division, E. O. Lawrence Berkeley National Laboratory, Berkeley, California 94720*

³*Center for Beam Physics, E. O. Lawrence Berkeley National Laboratory, Berkeley, California 94720*

(Received 2 June 1997)

Results of a search for parity nonconservation (PNC) in a pair of nearly degenerate opposite-parity states in atomic dysprosium are reported. The sensitivity to PNC mixing is enhanced in this system by the small energy separation between these levels, which can be crossed by applying an external magnetic field. The metastable odd-parity sublevel of the nearly crossed pair is first populated. A rapidly oscillating electric field is applied to mix this level with its even-parity partner. By observing time-resolved quantum beats between these sublevels, we look for interference between the Stark-induced mixing and the much smaller PNC mixing. To guard against possible systematic effects, reversals of the signs of the electric field, the magnetic field, and the decrossing of the sublevels are employed. We report a value of $|H_w| = |2.3 \pm 2.9 \text{ (statistical)} \pm 0.7 \text{ (systematic)}|$ Hz for the magnitude of the weak-interaction matrix element. A detailed discussion is given of the apparatus, data analysis, and systematic effects. [S1050-2947(97)02111-2]



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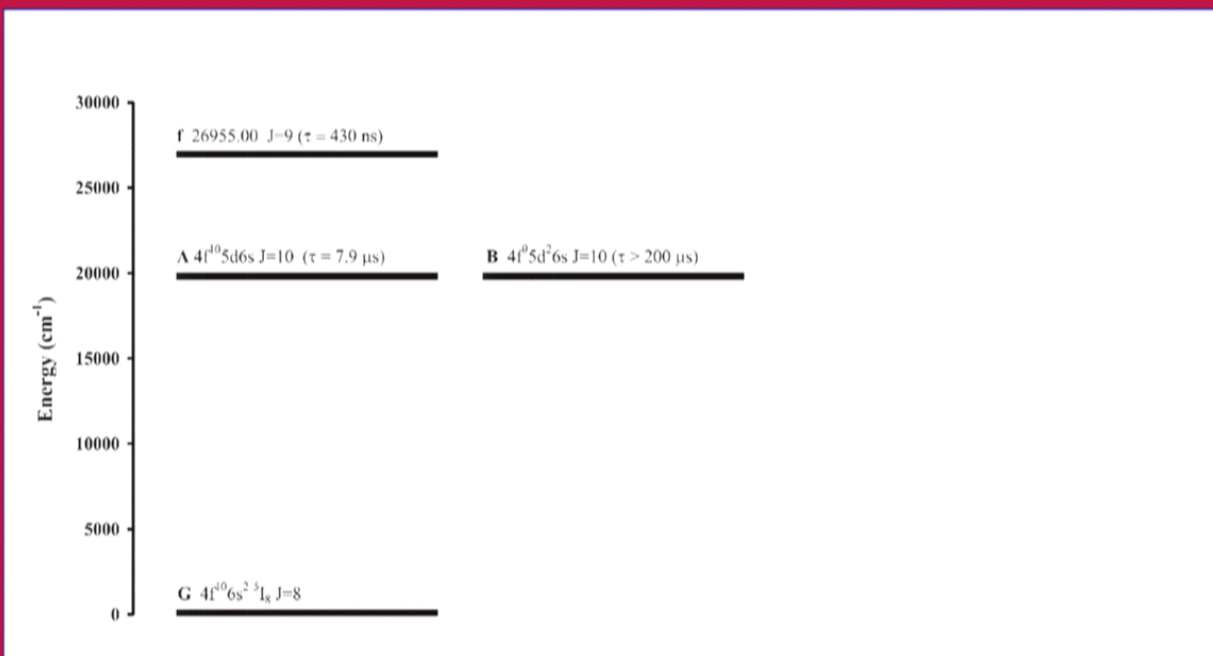
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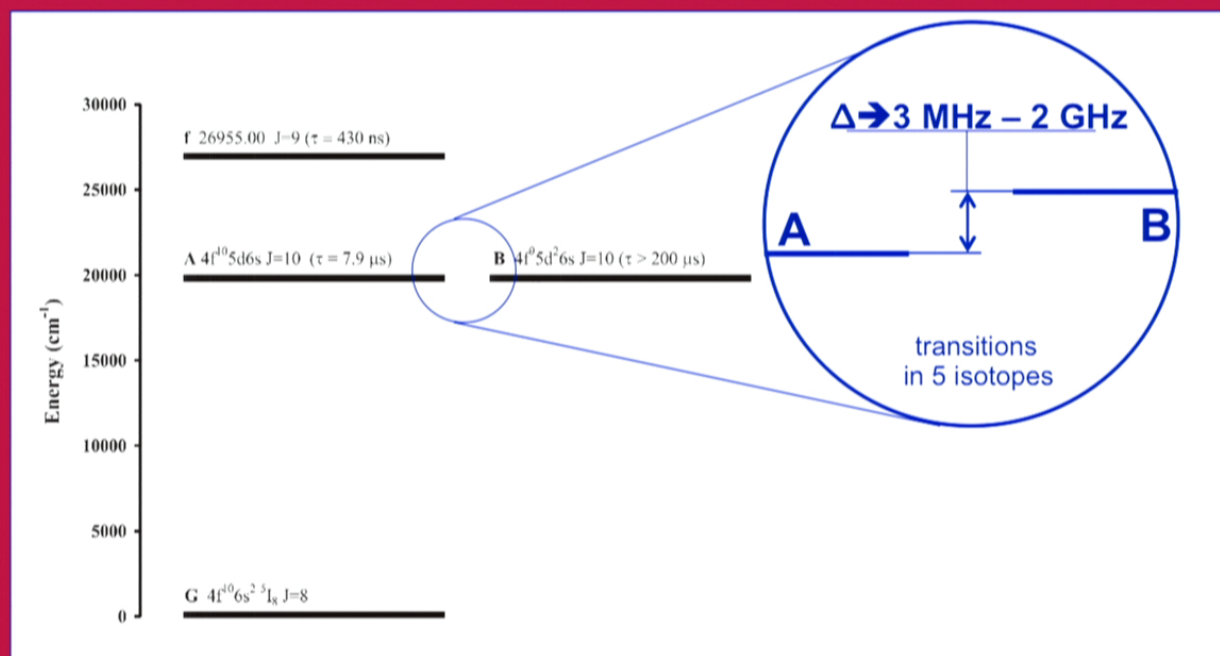
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Search in Atomic Dy



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Sensitivity

- Transition linewidth γ determined by lifetime of A
($\tau = 7.9 \mu\text{s}$) and transit broadening: $\gamma \sim 40 \text{ kHz}$
- Counting rate $\sim 10^9 \text{ s}^{-1}$
- 10 mHz sensitivity in 6 hours of averaging

$$\delta\nu \sim \frac{\gamma}{\sqrt{N}} \sim \frac{1.3}{\sqrt{T_{avg}}} \text{ Hz}$$

$$\dot{\alpha}/\alpha \sim 5 \times 10^{-18} \text{ yr}^{-1}$$

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fractional frequency
stability

Easy

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fractional frequency
stability

Easy

$$\frac{\delta\nu}{\gamma} \sim 3 \times 10^{-7}$$

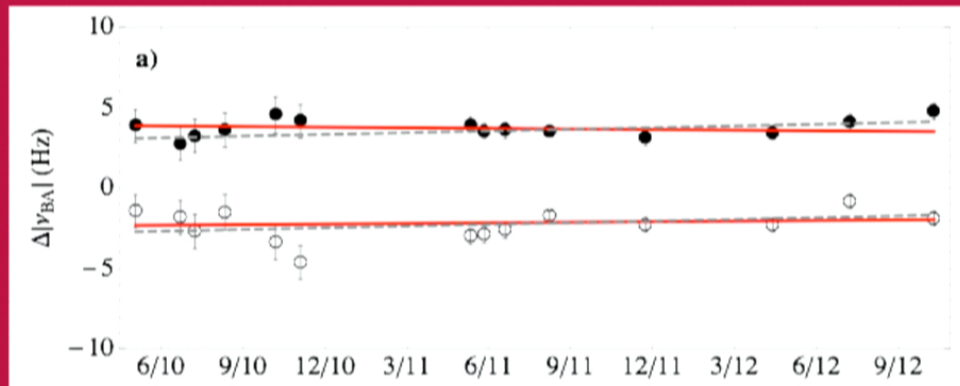
Difficult

“splitting the line”



New Limits on Variation of the Fine-Structure Constant Using Atomic Dysprosium

N. Leefer,¹ C. T. M. Weber,² A. Cingöz,³ J. R. Torgerson,⁴ and D. Budker^{1,5}



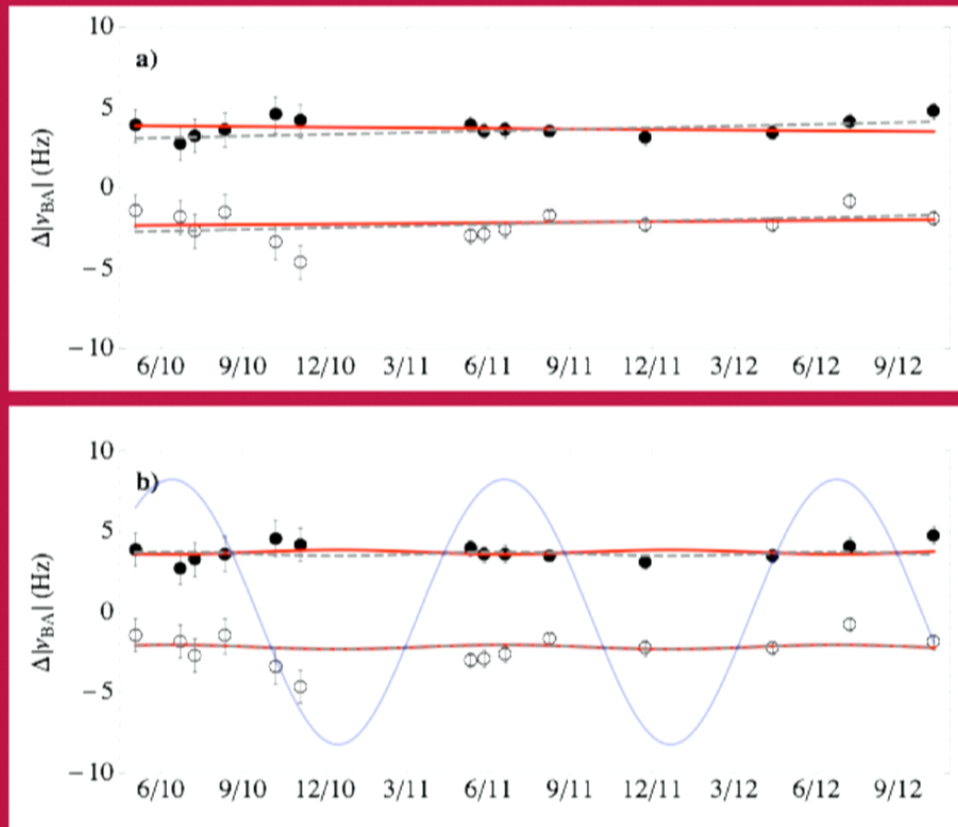
Isotope Comparison

$$\dot{\alpha}/\alpha = (-5.8 \pm 6.9) \times 10^{-17}$$



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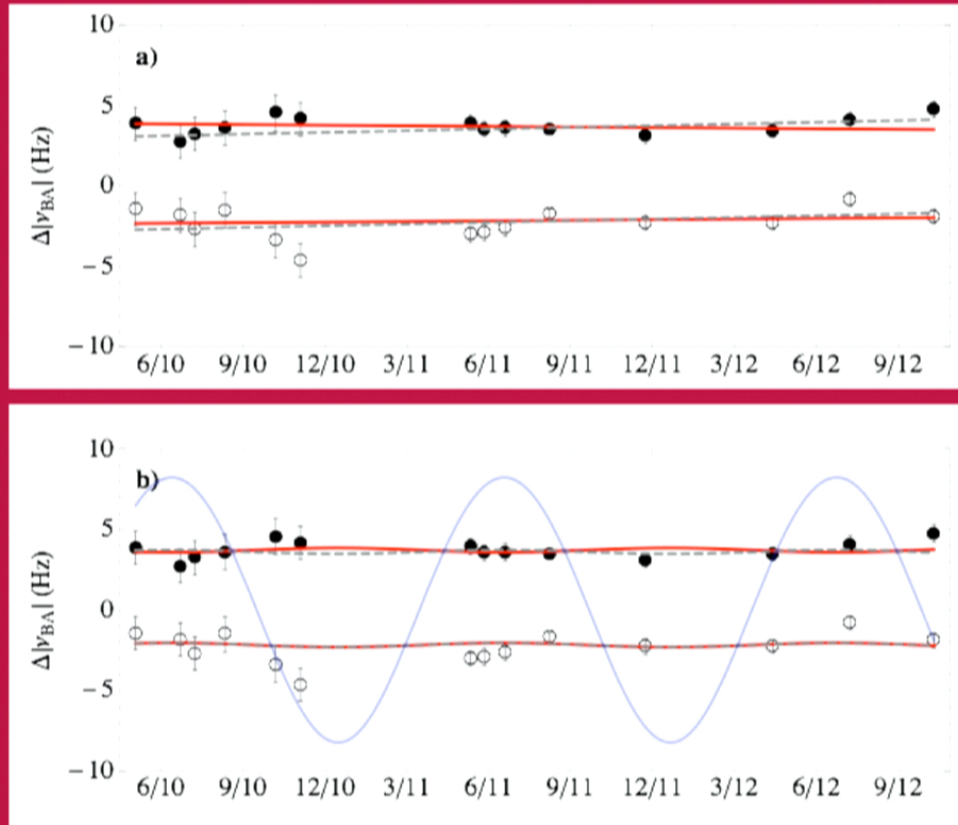
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Systematics Limited!



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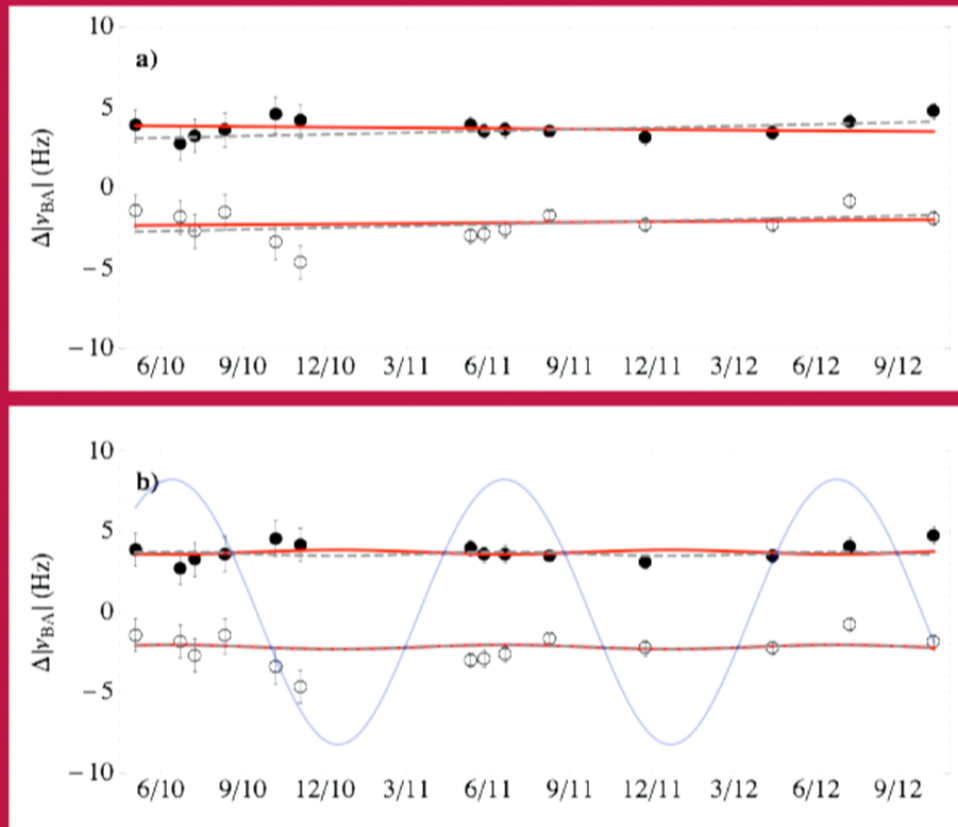
$$\frac{\Delta\alpha}{\alpha} = k_{\alpha} \frac{\Delta U}{c^2}$$

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The effect of gravitational field gradient on atomic energies

A. Gerhardus,* N. Leefer,[†] M. Hohensee, and D. Budker
Physics Department, University of California, Berkeley 94720-7300, USA

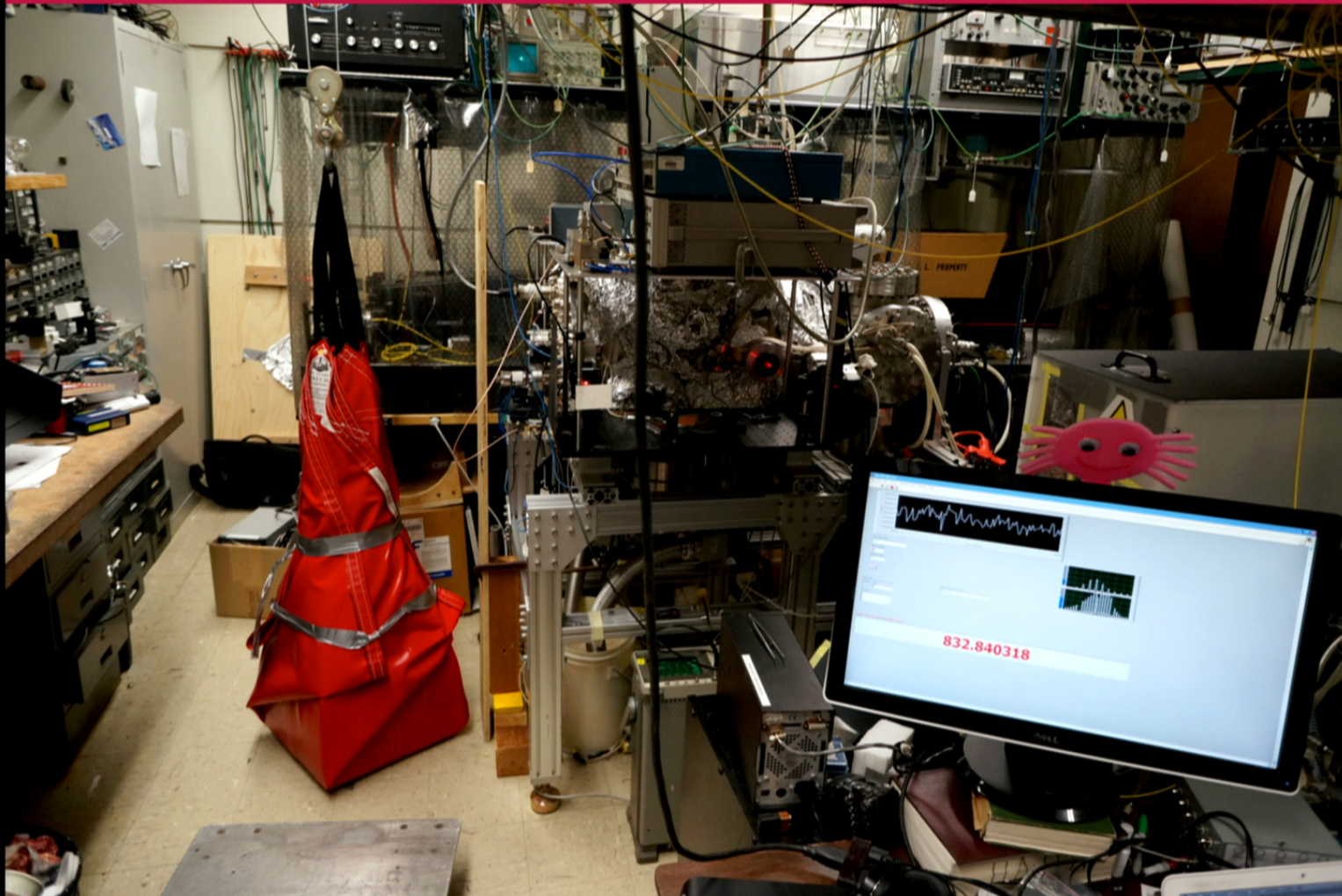
P. Graham
Department of Physics, Stanford University, Stanford 94305, USA
(Dated: November 24, 2013)

$$\alpha = \frac{e^2}{4\pi} (1 + D^{ij} g_{ij})$$

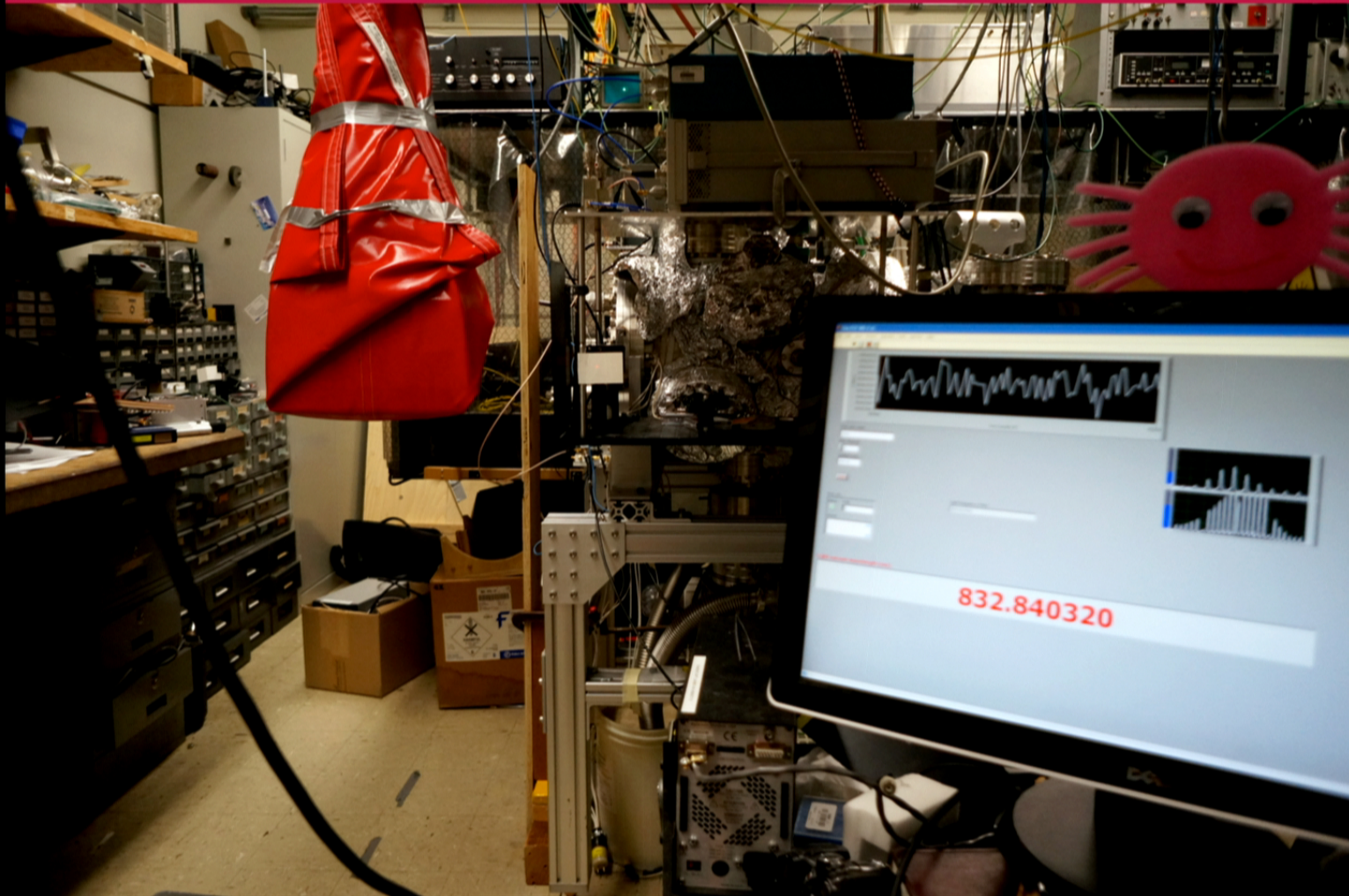
Background
tensor field

Gravitational-field
gradient

Search for background tensor field



Search for background tensor field



Search for background tensor field

- Move the 300 kg mass
- $\rightarrow D^{ij} < 10^{21} \text{ eV}^{-2}$ (exptl limits)
- Can also wait for earth to rotate w.r.t. cosmic gradient

Q: Does such search make sense ?

$$\alpha = \frac{e^2}{4\pi} (1 + D^{ij} g_{ij})$$

Background
tensor field

Gravitational-field
gradient

- Is parity conserved by gravitation ?

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(GDMs are T-odd,...) → centrifuge (EEP)

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- Can we test gravity via APV ?

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- Probably not...

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- ... but can look for exotic cosmic fields

DM, DE, ...

ONE THING LEADS TO ANOTHER

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DM, DE, ...

Limits on \mathcal{P} -odd interactions of cosmic fields with electrons, protons and neutronsB. M. Roberts,^{1,*} Y. V. Stadnik,^{1,†} V. A. Dzuba,¹ V. V. Flambaum,^{1,2} N. Leeper,³ and D. Budker^{3,4,5}¹*School of Physics, University of New South Wales, Sydney 2052, Australia*²*New Zealand Institute for Advanced Study, Massey University, Auckland 0745, New Zealand*³*Helmholtz Institute Mainz, Johannes Gutenberg University, 55099 Mainz, Germany*⁴*Department of Physics, University of California at Berkeley, Berkeley, CA 94720-7300, USA*⁵*Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA*

(Dated: June 13, 2014)

We propose methods for extracting limits on the strength of \mathcal{P} -odd interactions of pseudoscalar and pseudovector cosmic fields with electrons, protons and neutrons. Candidates for such fields are dark matter (including axions) and dark energy, as well as several more exotic sources described by standard-model extensions. Calculations of parity nonconserving amplitudes and atomic electric dipole moments induced by these fields are performed for H, Li, Na, K, Rb, Cs, Ba⁺, Tl, Dy, Fr, and Ra⁺. From these calculations and existing measurements in Dy and Cs, we constrain the parity-violating interaction of a static pseudovector cosmic field at 9×10^{-15} GeV for the electron, and 3.1×10^{-8} GeV for the proton.

Limits on \mathcal{P} -odd interactions of cosmic fields with electrons, protons and neutronsB. M. Roberts,^{1,*} Y. V. Stadnik,^{1,†} V. A. Dzuba,¹ V. V. Flambaum,^{1,2} N. Leefer,³ and D. Budker^{3,4,5}¹*School of Physics, University of New South Wales, Sydney 2052, Australia*²*New Zealand Institute for Advanced Study, Massey University, Auckland 0745, New Zealand*³*Helmholtz Institute Mainz, Johannes Gutenberg University, 55099 Mainz, Germany*⁴*Department of Physics, University of California at Berkeley, Berkeley, CA 94720-7300, USA*⁵*Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA*

(Dated: June 13, 2014)

We propose methods for extracting limits on the strength of \mathcal{P} -odd interactions of pseudoscalar and pseudovector cosmic fields with electrons, protons and neutrons. Candidates for such fields are dark matter (including axions) and dark energy, as well as several more exotic sources described by standard-model extensions. Calculations of parity nonconserving amplitudes and atomic electric dipole moments induced by these fields are performed for H, Li, Na, K, Rb, Cs, Ba⁺, Tl, Dy, Fr, and Ra⁺. From these calculations and existing measurements in Dy and Cs, we constrain the parity-violating interaction of a static pseudovector cosmic field at 9×10^{-15} GeV for the electron, and 3.1×10^{-8} GeV for the proton.

TABLE II. Limits (1σ) on the interaction strengths of a PV cosmic field with electrons (b_0^e) and protons (b_0^p) in GeV.

| | PNC quantity | $ b_0^e $ | $ b_0^p $ |
|----|-------------------------------------|-----------------------|----------------------|
| Cs | $E_{\text{PNC}}(6s-7s)$ | 1.5×10^{-14} | 3.1×10^{-8} |
| Tl | $E_{\text{PNC}}(6p_{1/2}-6p_{3/2})$ | 2.2×10^{-12} | 8.5×10^{-8} |
| Dy | $\langle A \hat{h} B\rangle$ | 7×10^{-15} | |

Q: How do cosmic PNC searches
fare compared to other searches ?

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Q: So how do we test parity in gravity?

Search for Exotic Physics with Antiatoms ?



A New Era...

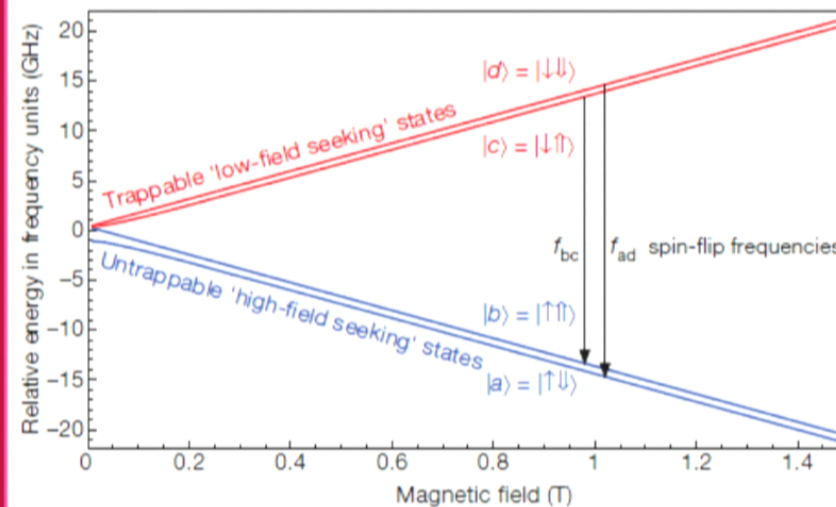
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Resonant quantum transitions in trapped antihydrogen atoms

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WHAT ELSE IN ADDITION TO FALLING UP AND 1S-2S (CPT)?

WHAT ELSE IN ADDITION TO FALLING UP AND 1S-2S (CPT)?



It difficult to come up with a scenario that will have exotic physics with antimatter that would not “leak” into the ordinary-matter sector without “fine tuning”



CPT and Lorentz Invariance

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CPT Violation Implies Violation of Lorentz Invariance

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(Received 28 January 2002; published 18 November 2002)

A interacting theory that violates *CPT* invariance necessarily violates Lorentz invariance. On the other hand, *CPT* invariance is not sufficient for out-of-cone Lorentz invariance. Theories that violate *CPT* by having different particle and antiparticle masses must be nonlocal.

~~CPT~~ → ~~LI~~

CPT and Lorentz Invariance

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CPT violation does not lead to violation of Lorentz invariance and vice versa

Masud Chaichian^{a,*}, Alexander D. Dolgov^{b,c,d}, Victor A. Novikov^d, Anca Tureanu^a

~~CPT~~ \nrightarrow ~~LI~~

CPT and Lorentz Invariance

CPT, Lorentz invariance, mass differences,
and charge non-conservation

A.D. Dolgov, V.A. Novikov

April 26, 2012

arxiv.org/abs/1204.5612v1

$$\begin{array}{c} \cancel{\text{CPT}} \not\rightarrow \cancel{\text{LI}} \\ \text{LI} \leftrightarrow m = \overline{m} \\ \text{(even with } \cancel{\text{CPT}} \text{)} \end{array}$$

Search for Exotic Physics with Antiatoms ?

Personal opinion:

pay no attention to what theorists (or
astronomers) say, and measure what is
important

Fun stuff to do with (anti)atomic clocks...

- ▣ Dependence on gravitational potential and gradients
- ▣ Anomalous spin-mass and spin-spin couplings
- ▣ Noise correlations?
- ▣ Temporal variation

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- ▣ Temporal variation

But many (most?) specific models are already constrained by experiments with ordinary matter, positrons, and antiprotons!

Finally...

Why stop at
antihydrogen?

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ANTI ENERGY SECRETARY

Q: What tests should be done with antimatter beyond \bar{g} ?



GNOME & CASPEr Open Qns

GNOME & CASPEr Open Qns

- ▣ Axion “g” factors for ^{129}Xe , ^3He , ^{87}Rb , ...
- ▣ Optimal experimental strategy for “GNOME now”
- ▣ “GNOME-AC” ?