

Title: Effective theories of vortex lines

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Abstract: Vortex lines are a distinctive feature of superfluids and are characterized by a very peculiar dynamics. In this talk, I will first discuss the behavior of vortex lines in a non-relativistic superfluids in the incompressible limit. I will then introduce an effective theory of vortex lines coupled to sound which applies to relativistic superfluids. I will conclude by briefly discussing the similarities between the effective theory for vortex lines and non-relativistic General Relativity.

Effective Theories of Vortex Lines

Riccardo Penco
Columbia University

with Steve Gubser, Bart Horn, Alberto Nicolis — *to appear*
Low Energy Challenges for High Energy Physicists
Perimeter Institute, May 26-30, 2014

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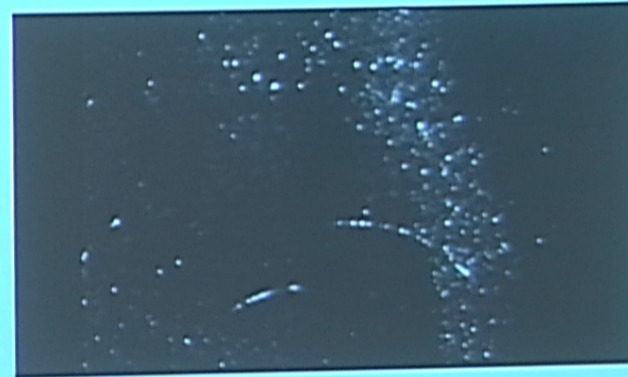
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Why Vortex Lines?

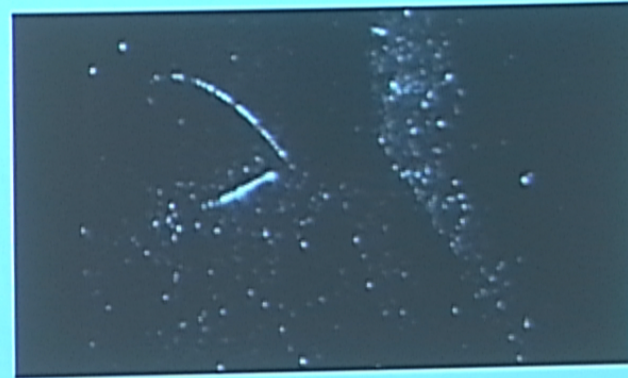


They exist in superfluids and ordinary fluids.
They don't behave like strings.

Video credit: E. Fonda, D. P. Meichle, N. T. Ouellette, S. Hormoz, K. R. Sreenivasan, D. P. Lathrop

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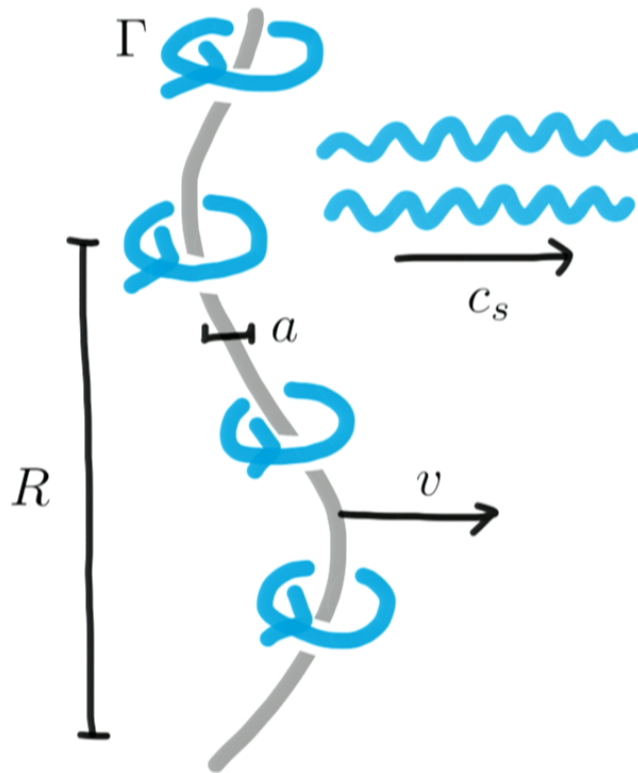
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Basic Quantities



Vortex lines: $\nabla \times \vec{v} \neq 0$

Circulation: $v(r) \sim \Gamma/r$

Regime of effective theory:

$$L \gg a \quad T \gg \frac{a}{c_s}$$

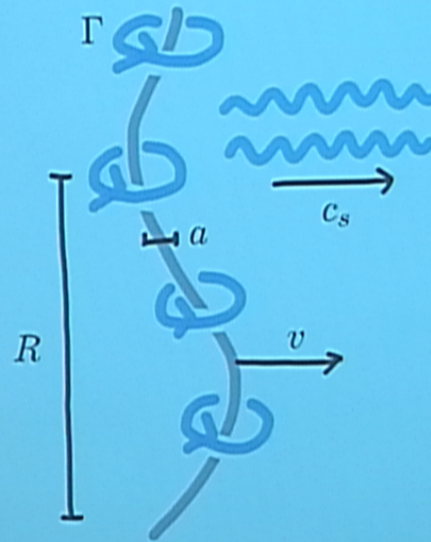
Non-relativistic motion:

$$v \sim \Gamma/R \ll \Gamma/a \lesssim c_s$$

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Superfluids

Superfluid = spontaneously broken U(1) at finite density

$$\mathcal{L} = F(X) \quad X = \partial_\mu \phi \partial^\mu \phi \quad \phi = t + \varphi$$



$$X = -1 - 2\dot{\varphi} - \dot{\varphi}^2 + (\nabla\varphi)^2$$

$$\mathcal{L} = a_0 + a_1(X + 1) + a_2(X + 1)^2 + \dots \quad a_{n+1} \sim \frac{a_n}{c_s^2}$$

NR limit: $\vec{v} = \nabla\varphi \longrightarrow \nabla \times \vec{v} = 0$

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Sen 2002

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$$\langle \psi \rangle = \int d^d k \sum_{\omega} \frac{1}{k^2 + \omega^2}$$

Introducing Vortex Lines

$$\mathcal{L} = a_0 + a_1(X+1) + a_2(X+1)^2 + \dots \quad a_{n+1} \sim \frac{a_n}{c_s^2}$$

1. Take incompressible, non-relativistic limit: $c_s, c \rightarrow \infty$

2. Introduce external sources to redefine velocity:

$$\nabla\varphi \rightarrow \nabla\varphi + \vec{S} \quad \nabla \times \vec{S} = \Gamma \int d\sigma \partial_\sigma \vec{X}(t, \sigma) \delta(\vec{x} - \vec{X}(t, \sigma))$$

3. Integrate out φ

$$\langle \psi(0) \rangle = \frac{k}{X^4} \quad k = \frac{1}{16}$$

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Magnetostatic analogy

(incompressible limit)

The hydrophoton couples to vortex lines
like
the vector potential couples to lines of current
(in Coulomb gauge)

magnetostatic	incompr. hydro
current \vec{J}	vorticity $\vec{\omega}$
vector potential \vec{A}	hydrophoton \vec{A}
magnetic field \vec{B}	velocity field \vec{v}

$$\langle \vec{J}(\vec{r}) \rangle = \frac{k}{X^4} \quad k = \frac{1}{16}$$
$$\langle \phi^2 \rangle = \int \frac{1}{\epsilon + \mu + \nu}$$

Questions

1. What is the origin of the hydrophoton?
2. What about local interactions with sound?
3. What about the log in the dispersion relation?

$$\langle \psi(0) \rangle = \frac{k}{X^4} \quad k = \frac{1}{16}$$
$$\langle \psi^2 \rangle = \int dk \sum_{\omega} \frac{1}{k^2}$$

$$\langle \phi^2 \rangle = \int d^4k \sum_n \frac{1}{k^2 + \omega_n^2 + \nu}$$

$(k^{-1} h)$

“Fun” Fact

Kelvin waves = Goldstone bosons

unbroken = P_t, P_z, J_z , broken = P_x, P_y, J_x, J_y

Coset Construction \longrightarrow



\longleftarrow Lorentz, gauge & reparametrization invariance

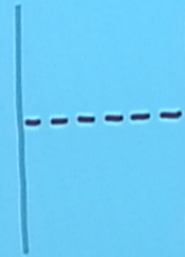
$$S_{2D} = \int d\tau d\sigma \left\{ A_{\mu\nu} \frac{dX^\mu}{d\tau} \frac{dX^\nu}{d\sigma} + \sqrt{-k} F \left(\frac{\sqrt{-h}}{\sqrt{-k}}, Y \right) \right\}$$

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Interactions

between vortex lines




Non-relativistic,
incompressible

$$V \sim \Gamma^2 \int d\sigma d\sigma' \frac{\partial_\sigma \vec{X} \cdot \partial_{\sigma'} \vec{X}'}{|\vec{X} - \vec{X}'|}$$


Classical RG Running

The tension T can be fixed by demanding that the self-energy of a static vortex line is finite.

$$\frac{dE}{d\sigma} = \Gamma^2 \int d\sigma' \frac{\partial_\sigma \vec{X} \cdot \partial_{\sigma'} \vec{X}'}{|\vec{X} - \vec{X}'|} + T |\partial_\sigma \vec{X}|$$



$$\sim \Gamma^2 \log(L/a)$$

 $T(\mu) \sim \Gamma^2 \log(\mu a) + \text{finite}$

Lund & Regge, 1976

Classical RG Running

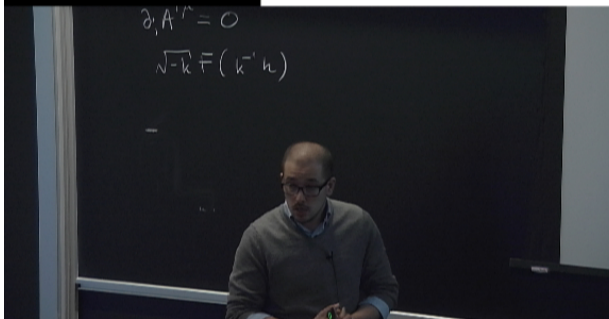
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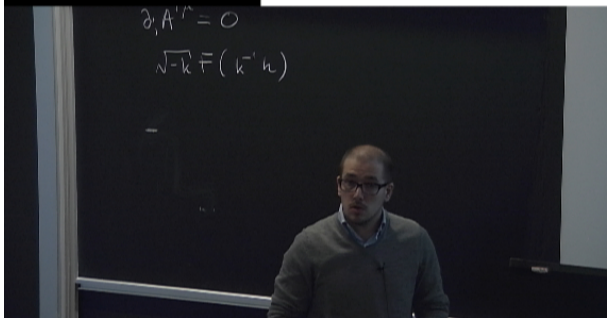
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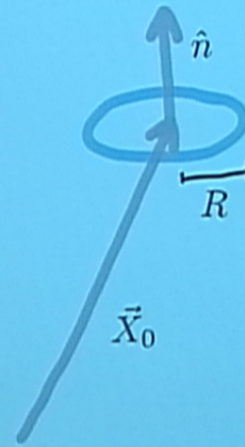
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$$k^{-1} h$$

Vortex Rings



Lagrangian:

$$\mathcal{L} = -\Gamma R^2 \pi \hat{n} \cdot \partial_t \vec{X}_0 - \frac{\Gamma^2}{2} R \log(R\mu)$$

Equations of motion:

$$\partial_t (R^2 \hat{n}) = 0$$

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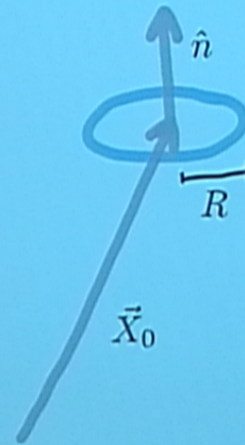
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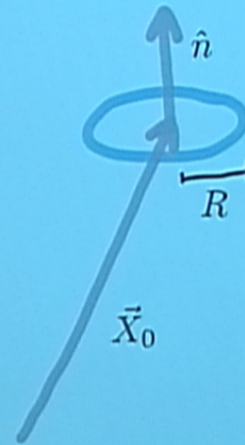
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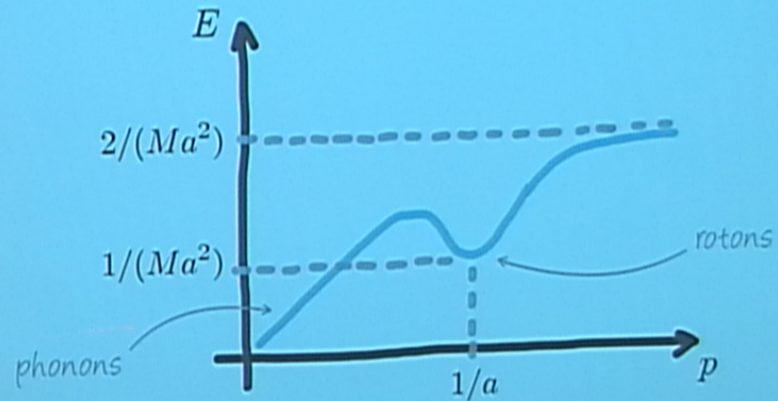
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Rotons in He4

(work in progress)



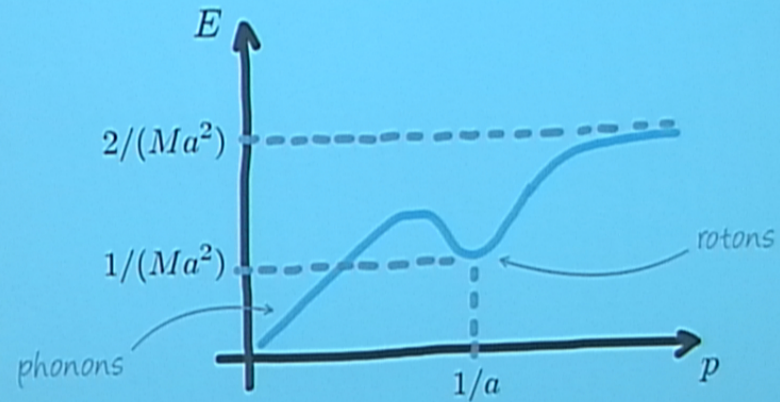
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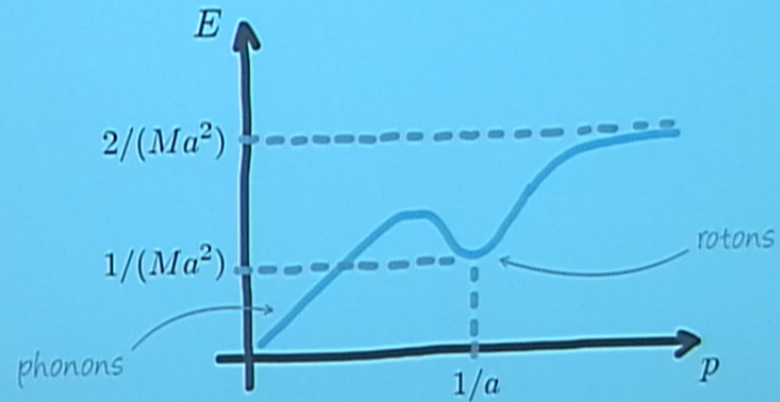
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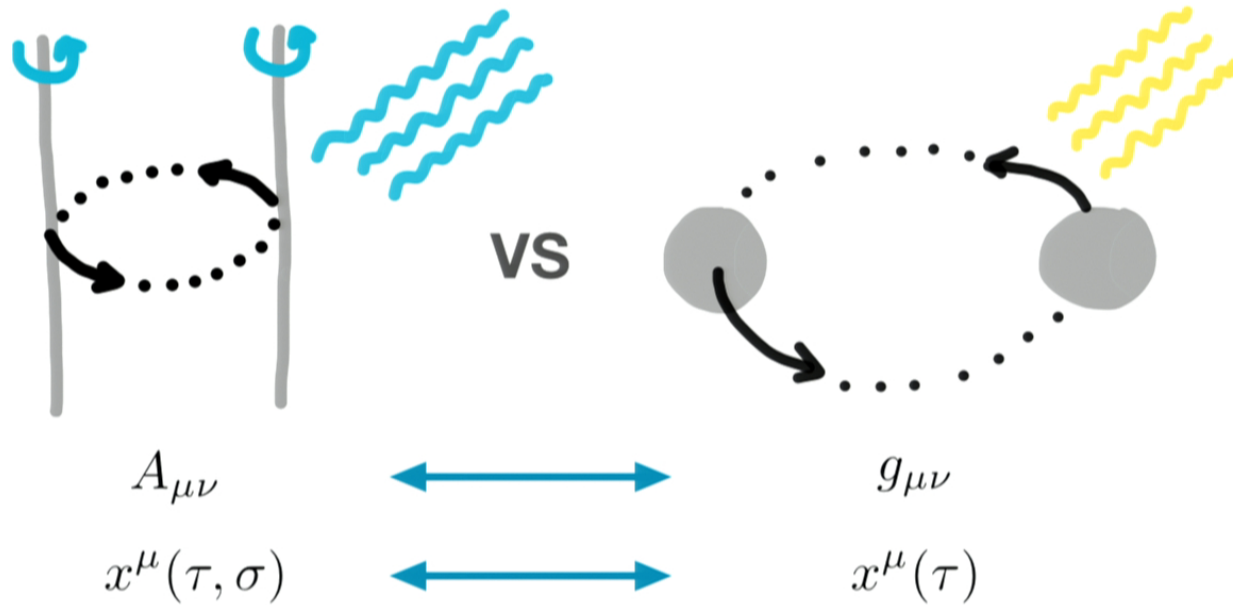
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Vortex Lines vs Binaries

(outlook)

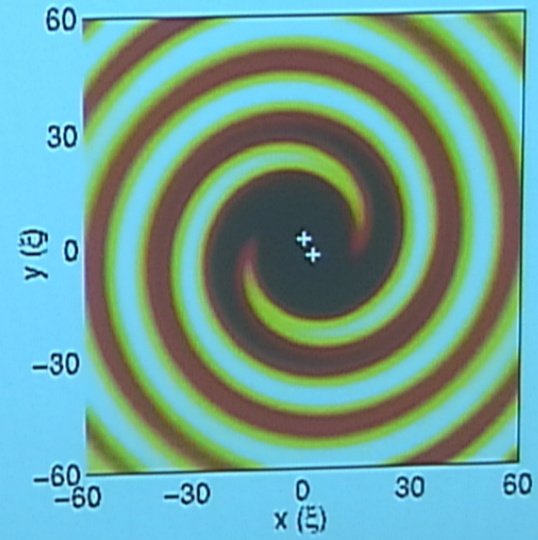


Endlich & Nicolis, 2013

Golberger & Rothstein, 2006

$$= \int dk \sum_n \frac{1}{k^2 + \omega^2 + \mu}$$

Numerical Simulations



Video credit: Carlo Barenghi

$$= \int dk \sum_{\omega} \frac{1}{k^2 + \omega^2 + \nu}$$

Velocity Power Counting

(outlook)

Working in Coulomb gauge with $c_s \sim 1$

$$\vec{B} \equiv \vec{B}_{pot} + \vec{B}_{rad}$$

$$\partial_i \vec{B}_{pot} \sim \frac{1}{r} \vec{B}_{pot}$$

$$\partial_t \vec{B}_{pot} \sim \frac{v}{r} \vec{B}_{pot}$$

$$\partial_t \vec{B}_{rad} \sim \partial_i \vec{B}_{rad} \sim \frac{v}{r} \vec{B}_{rad}$$

$$e^{iS[\vec{B}_{rad}, X^\mu]} = \int \mathcal{D}\vec{A} \mathcal{D}\vec{B}_{pot} e^{iS[\vec{A}, B_{pot} + \vec{B}_{rad}, X^\mu]}$$

“post-Newtonian” expansion for vortex lines!

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