

Title: A Fourth-Order Solution-Adaptive CENO Scheme for Space-Physics Flows on Three-Dimensional Multi-Block Cubed-Sphere Grids

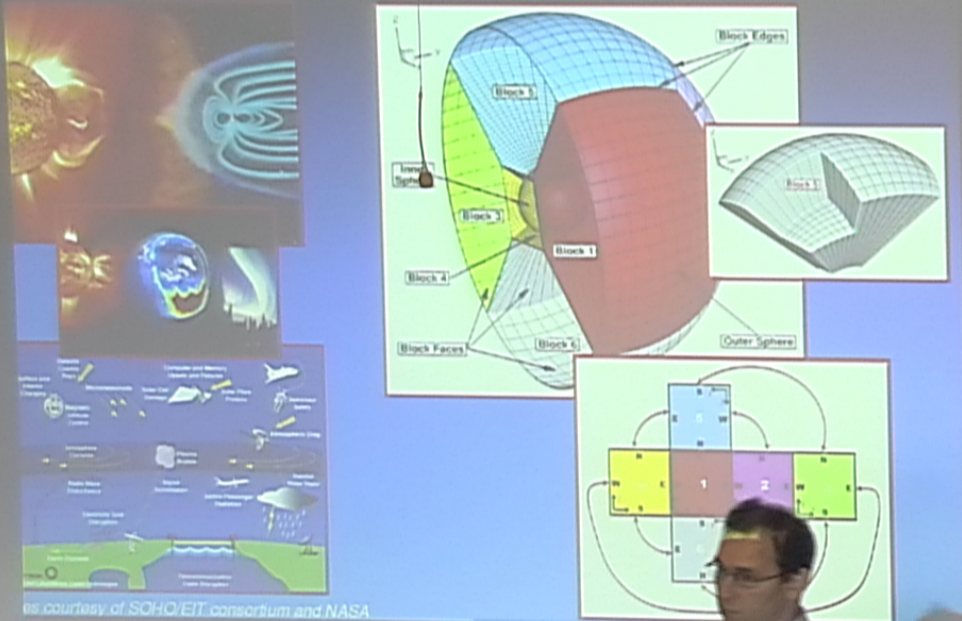
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Abstract: Accurate efficient and scalable computational methods are highly desirable for large-scale scientific computing applications especially for problems exhibiting spatial and temporal multi-resolution scales non-trivial geometries and complex boundary conditions (BSc). For global magnetohydrodynamics (MHD) modelling of space-physics problems high-performance approaches could significantly reduce the grid requirements to achieve targeted solution accuracies thereby enabling more affordable yet accurate predictions of space-plasma flows. Key challenges encountered relate to providing solenoidal magnetic fields accurate discretizations on spherical domains capturing of MHD shocks and implementing accurate BCs. This talk gives an overview of a fourth-order finite-volume discretization procedure in combination with a parallel solution-adaptive algorithm for the computation of MHD space plasmas on cubed-sphere grids. Numerical results to demonstrate the accuracy and capability of the multidimensional high-order solution-adaptive cubed-sphere computational framework are presented.

# Main Application Driver: 3D Space-Physics Flows

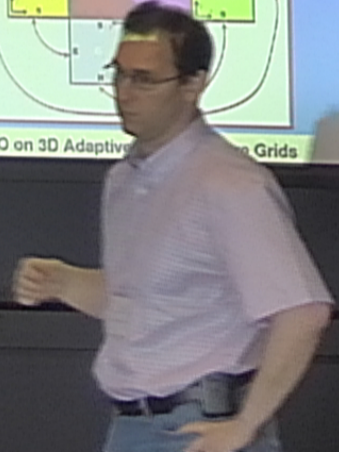
Unstructured-Sphere Discretization ( Sadourny, 1972; Ronchi *et. al.*, 1996)



as courtesy of SOHO/EIT consortium and NASA

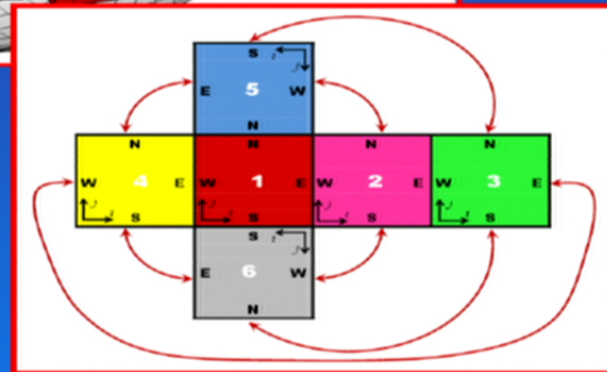
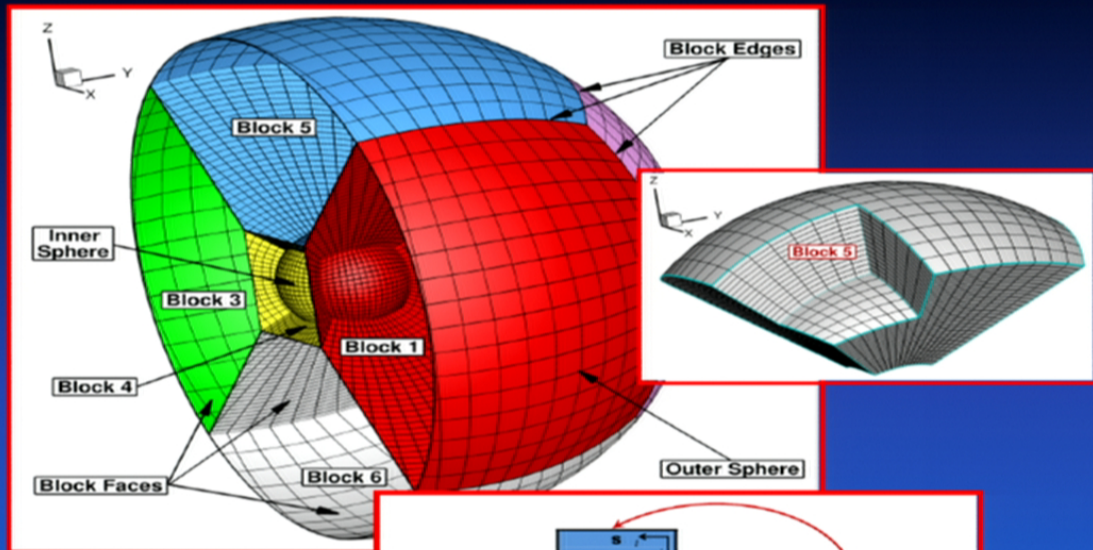
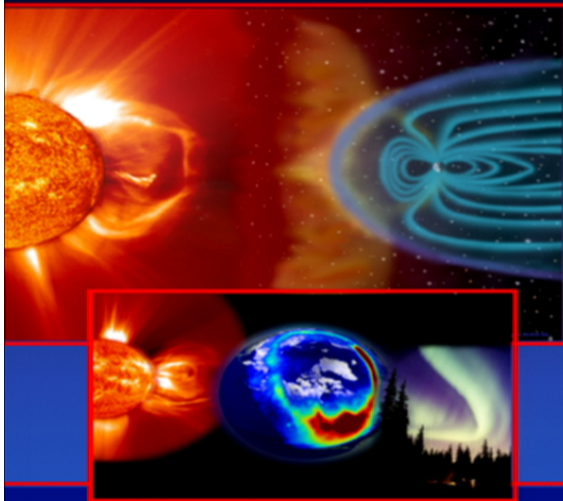
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4th-Order CENO on 3D Adaptive Grids



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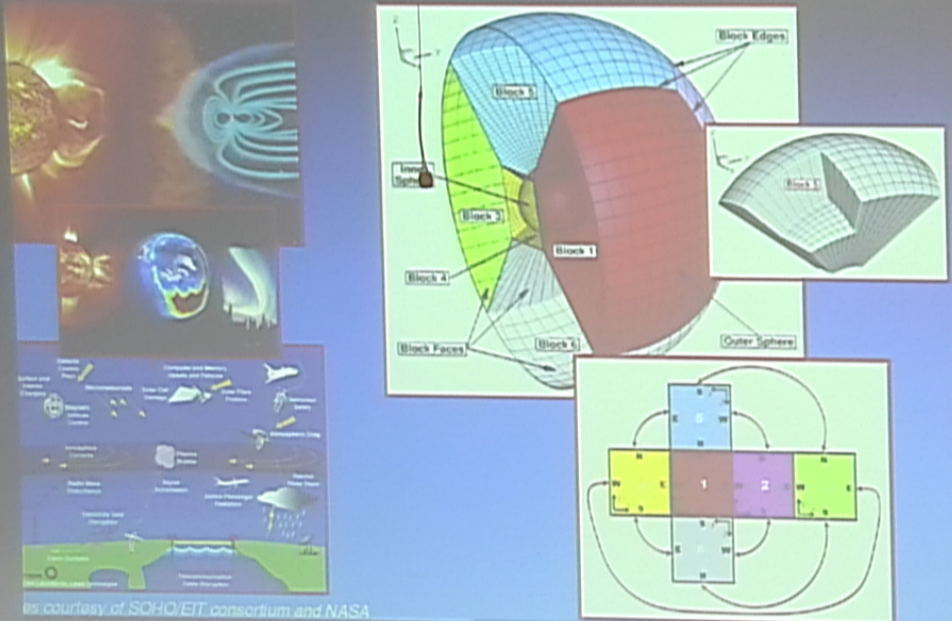
Cubed-Sphere Discretization ( Sadourny, 1972; Ronchi *et. al.*, 1996)



Images courtesy of SOHO/EIT consortium and NASA

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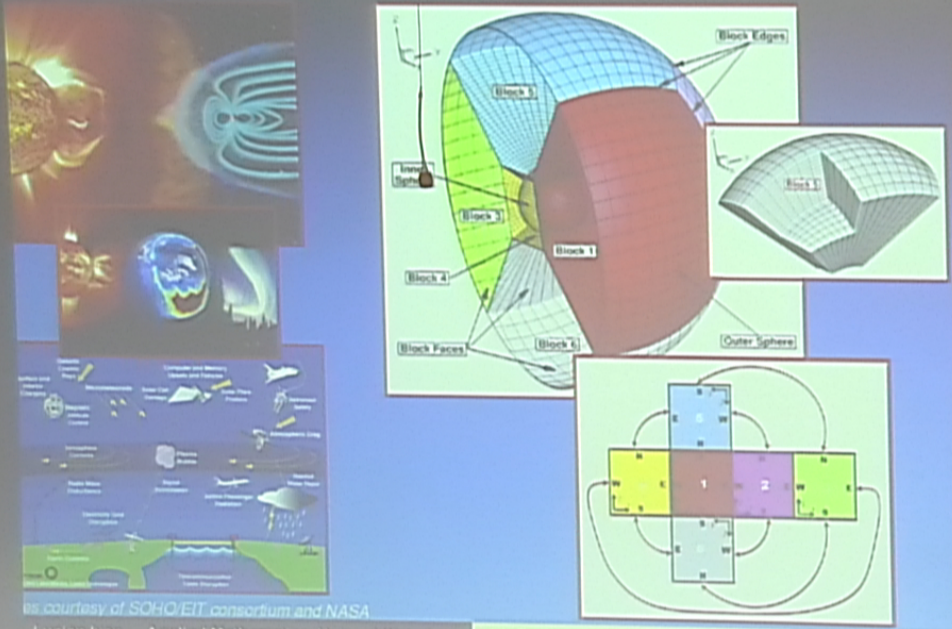
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Cubed-Sphere Discretization ( Sadourny, 1972; Ronchi *et. al.*, 1996)



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# Magnetohydrodynamics (MHD) Model

Approaches to Deal with the Divergence Constraint Condition,  $\nabla \cdot \vec{B} = 0$

## Powell Source Term (Powell *et al.*, 1999)

$$\mathbf{S}_{\text{num}} = -\nabla \cdot \vec{B} [ 0, \vec{B}, \vec{V}, \vec{V} \cdot \vec{B} ]^T$$

- 8-wave MHD system that is symmetric and Galilean invariant  
 $\lambda_{1,2} = v_x \pm c_{fx}$ ,  $\lambda_{3,4} = v_x \pm c_{Ax}$ ,  $\lambda_{5,6} = v_x \pm c_{sx}$ ,  $\lambda_{7,8} = v_x$
- Numerical error in  $\nabla \cdot \vec{B}$  is convected out of the domain by  $\lambda_8 = v_x$

## Divergence Correction Technique: Generalized Lagrange Multiplier (GLM)-MHD (Dedner *et al.*, 2002)

$$\frac{\partial \vec{B}}{\partial t} + \nabla \cdot (\vec{V} \vec{B} - \vec{B} \vec{V}) + \nabla \psi = 0$$
$$\frac{\partial \psi}{\partial t} + c_h^2 \nabla \cdot \vec{B} = -\frac{c_h^2}{c_p^2} \psi$$

- Solve an extra transport equation for the GLM,  $\psi$
- $\lambda_{8,9} = \pm c_h$ , the largest eigenvalue in the domain

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# Parallel AMR Finite-Volume Framework

## Second-order Finite-Volume Formulation

### Semi-Discrete Integral Form for Hexahedral Cell (i,j,k)

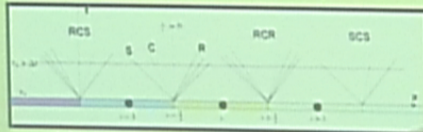
$$\frac{d\bar{U}_{ijk}}{dt} = -\frac{1}{V_{ijk}} \sum_{m=1}^{N_f} (\vec{F}_H \cdot \vec{n} A)_{i,j,k,m} + (S)_{ijk} = \mathbf{R}_{ijk}(\mathbf{U})$$

### Key Elements of Numerical Scheme

- Second-order spatial discretization
  - Multi-dimensional limited linear least-squares reconstruction (Barth 1993; Barth-Jespersen 1993; Venkatakrishnan 1993)

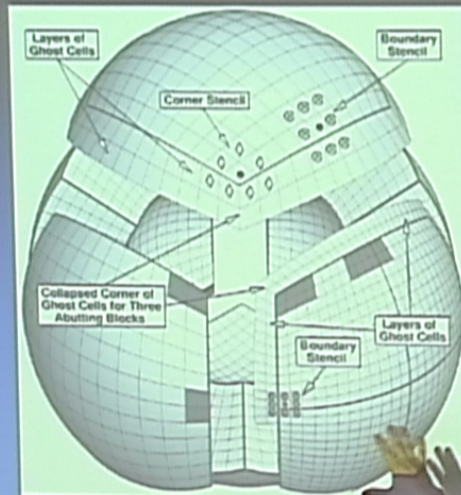
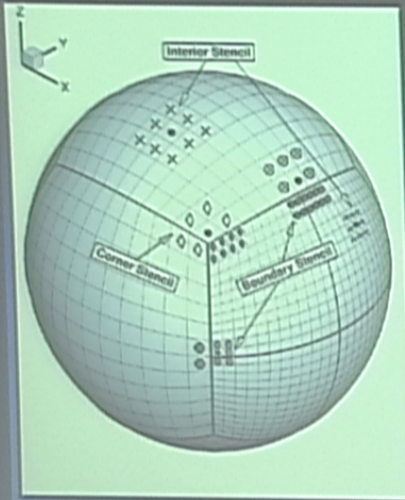
$$\mathbf{W}_{ijk}(\vec{x}) = \bar{\mathbf{W}}_{ijk} + \Phi_{ijk} \vec{\nabla} \bar{\mathbf{W}} \cdot (\vec{x} - \vec{x}_{ijk})$$

- Upwinding flux evaluation (HLL 1983, Lax-Friedrichs 1996)



- Time-integration procedure
  - Multi-stage explicit time marching schemes

## Handling of Block Resolution Changes



### Solution Transfer at Block Interfaces

- Fine-to-Coarse: apply conservation principles
- Coarse-to-Fine: apply linear approximation

## D MHD CENO-GLM Formulation

Follow the 2D Approach: A. Susanto *et al.*, "High-Order Central ENO Finite-Volume Scheme with Divergence Cleaning Approach for the Ideal MHD Equations", (*JCP*, 2013)

### Overview Idea of the High-Order MHD Algorithm

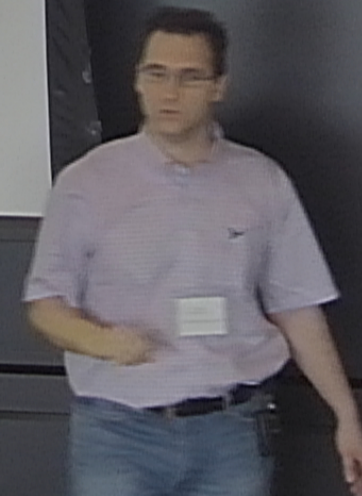
- Apply a high-order **CENO** approach (Ivan & Groth, 2007, 2011, 2014) to estimate accurately the residual
  - Combine an unlimited  $k$  exact reconstruction (Barth, 1993) with a monotonicity preserving limited linear ( $k=1$ ) scheme
  - A smoothness indicator triggers switching for non-smooth data
- Use **CENO + GLM-MHD** (Dedner *et al.*, 2002) to satisfy  $\nabla \cdot \vec{B} = 0$

### Accurate Evaluation of $R_{ijk}(\vec{U})$

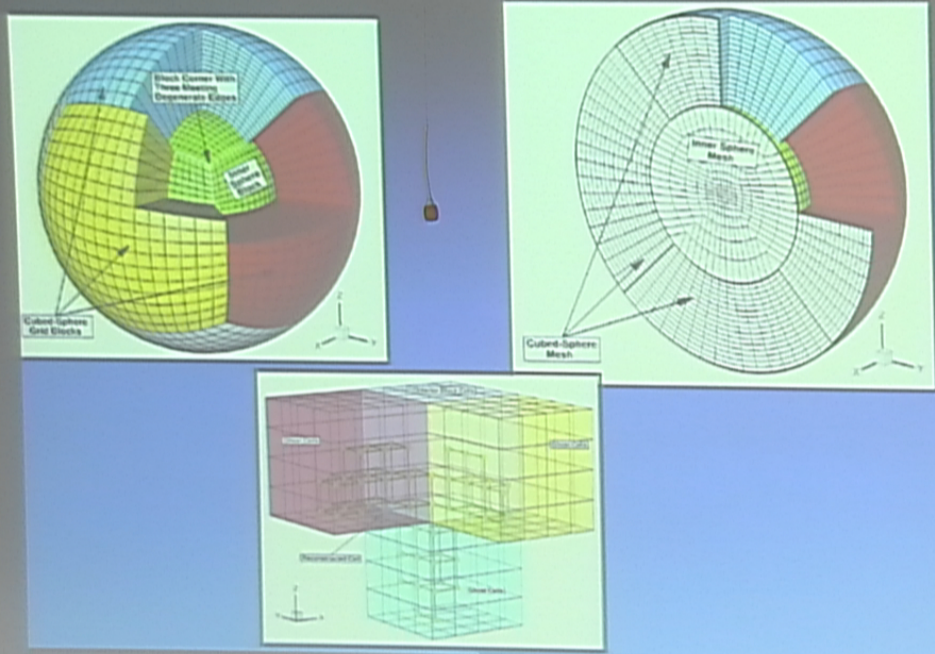
- More accurate calculation of flux and source term integrals

$$\frac{1}{V_{ijk}} \oint_{\partial V} \vec{F}_H \cdot \vec{n} da = \frac{1}{V_{ijk}} \sum_{l=1}^{N_f} \sum_{m=1}^{N_G} (\omega \vec{F}_H \cdot \vec{n} \Delta a)_{i,j,k,l,m} \quad (N_G \geq 4)$$

$$\frac{1}{V_{ijk}} \iiint_V (S) dv = \frac{1}{V_{ijk}} \sum_{v=1}^{N_V} (\omega(S))_{i,j,k,v} \quad (N_V \geq 4)$$



# How to Include an Interior Domain? ( e.g., a Planet)



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# High-Order Numerical Results

rho-Density Vortex Problem in Solid-Core Cubed-Sphere Grid

$$\rho = 9, \vec{V} = (1, 1, 2), \vec{X}_i = (-2, -2.5, -3), \vec{X}_f = (1, 0.5, 3), t = 3$$

