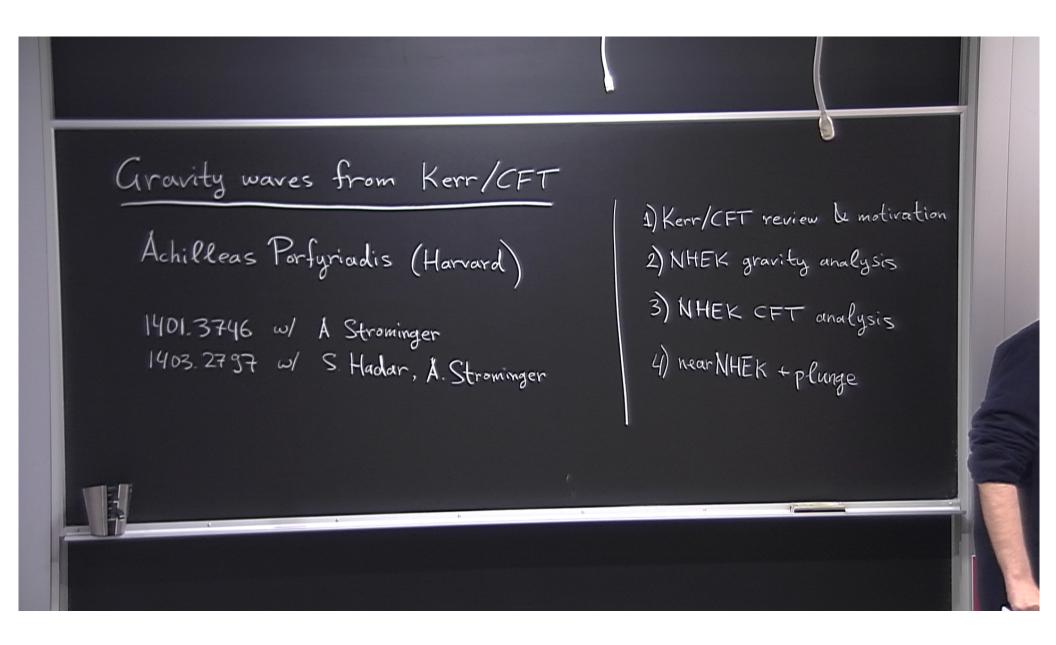
Title: Gravity waves from Kerr/CFT

Date: May 15, 2014 01:00 PM

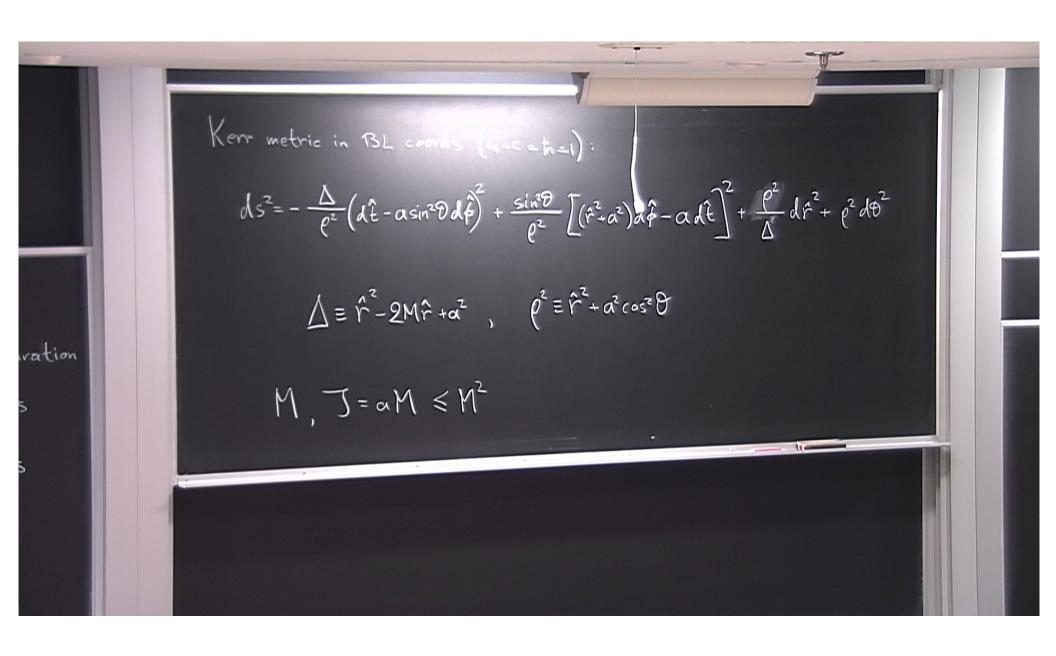
URL: http://pirsa.org/14050009

Abstract: Astronomical observation suggests the existence of near-extreme Kerr black holes whose horizons spin at nearly the speed of light. Properties of diffeomorphisms imply that the dynamics of the high-redshift near-horizon region of near-extreme Kerr, which includes the innermost-stable-circular-orbit (ISCO), is governed by an infinite-dimensional emergent conformal symmetry. This symmetry may be exploited to analytically, rather than numerically, compute a variety of potentially observable processes. In this talk I will show how we compute and study the conformal transformation properties of the gravitational radiation emitted by an orbiting massive object in the large-redshift near-horizon region. I will also use conformal symmetry of the near-horizon region to compute the gravitational radiation produced during the plunge phase following the object's crossing of the ISCO.

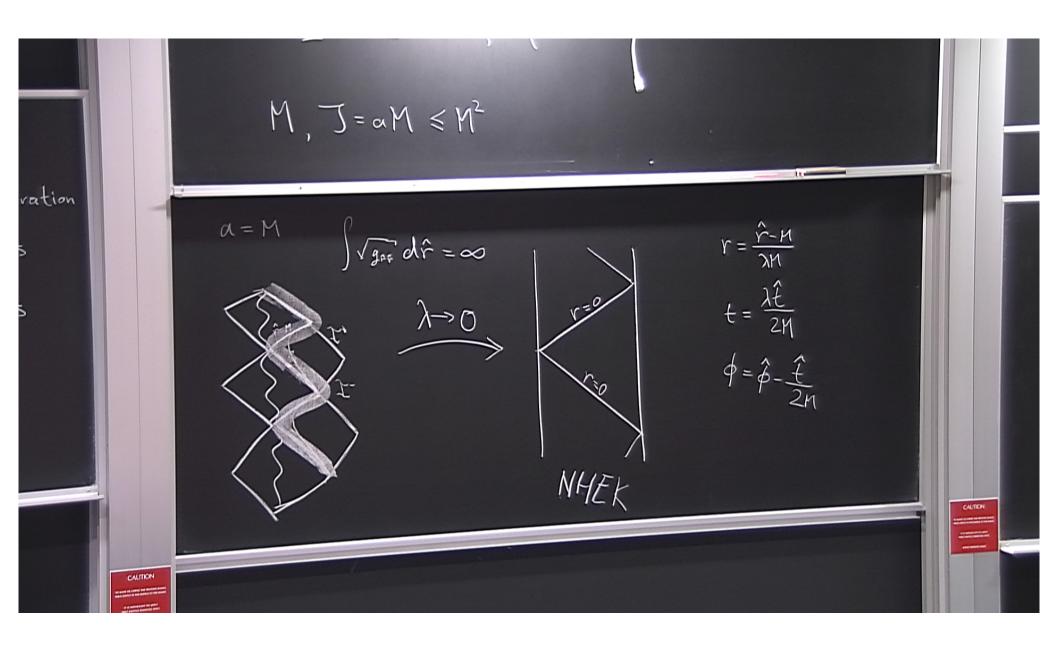
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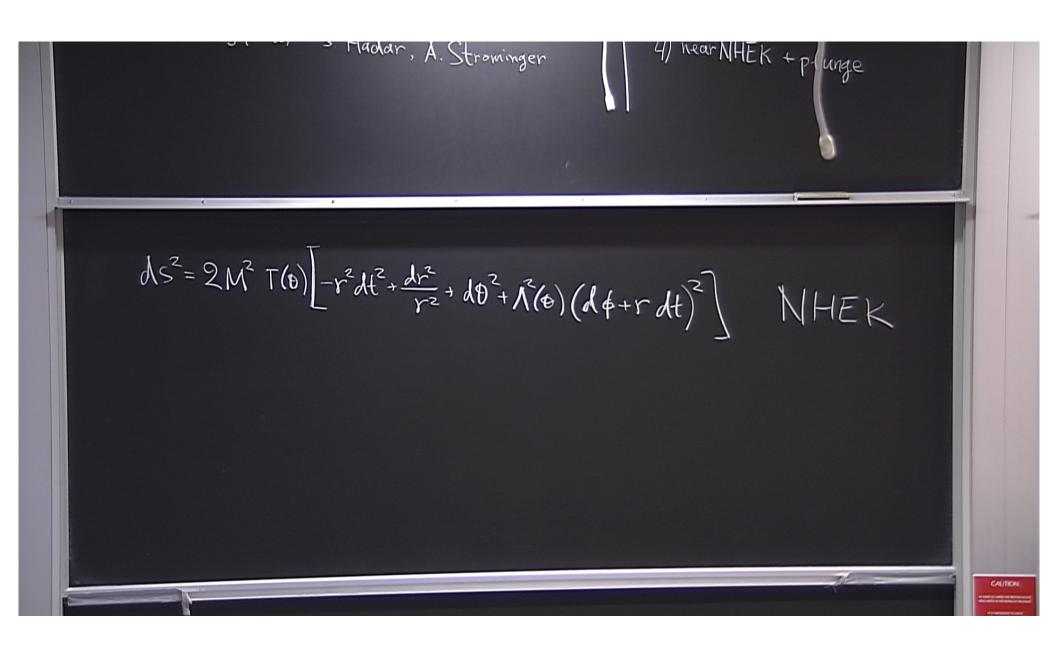
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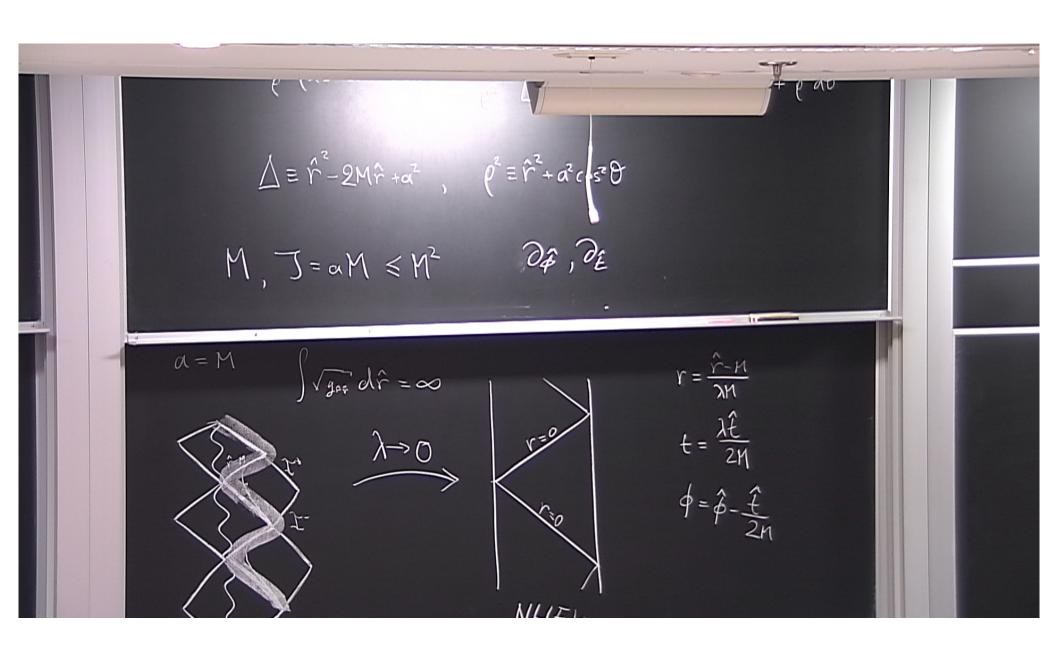
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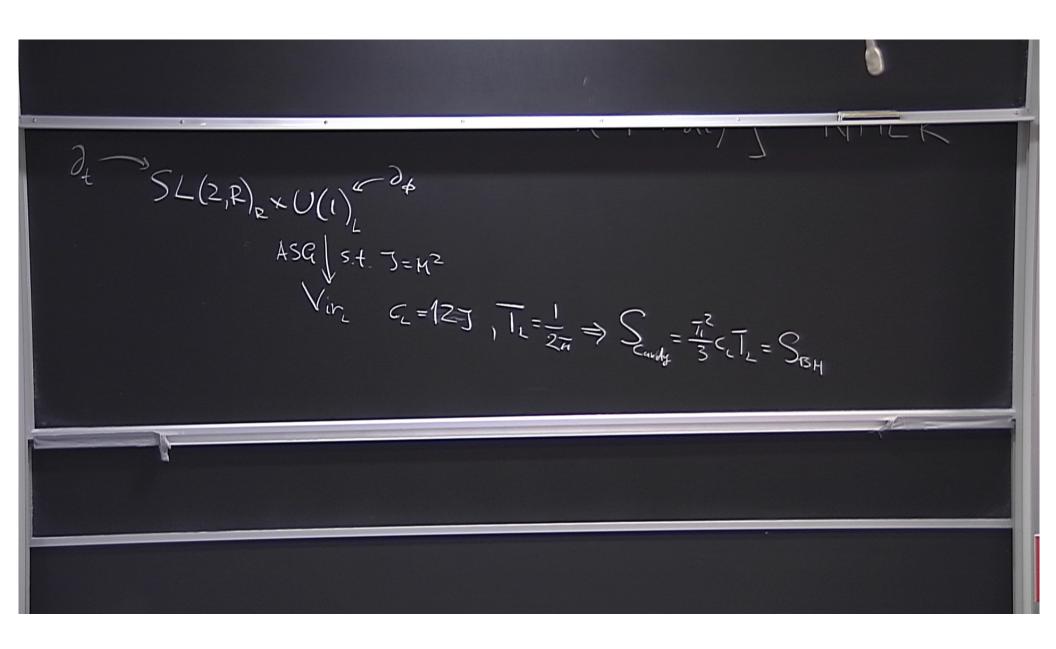
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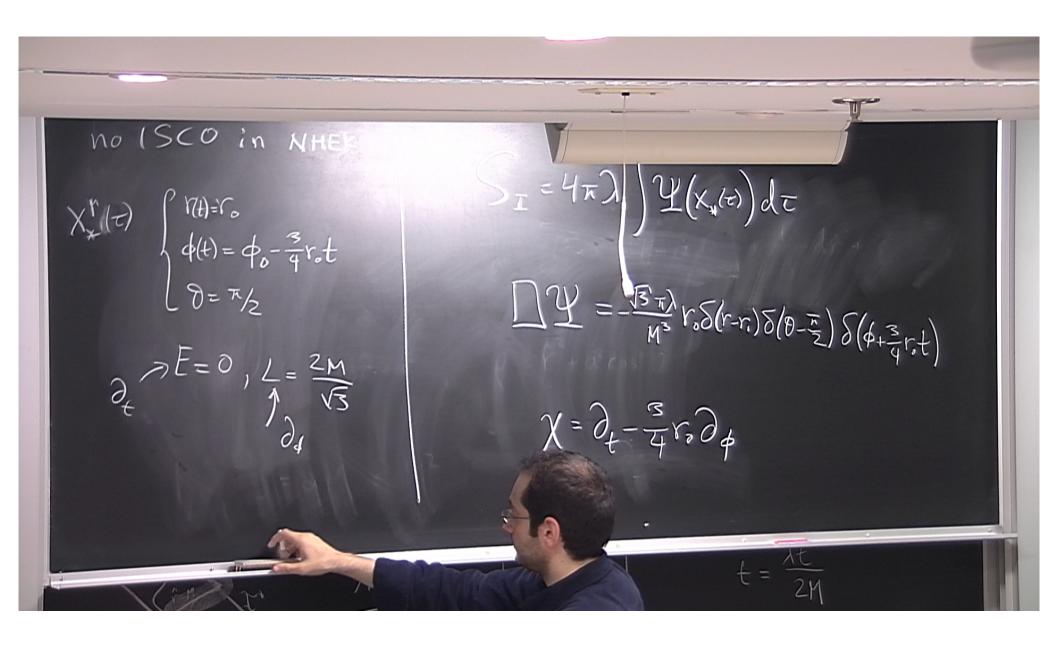
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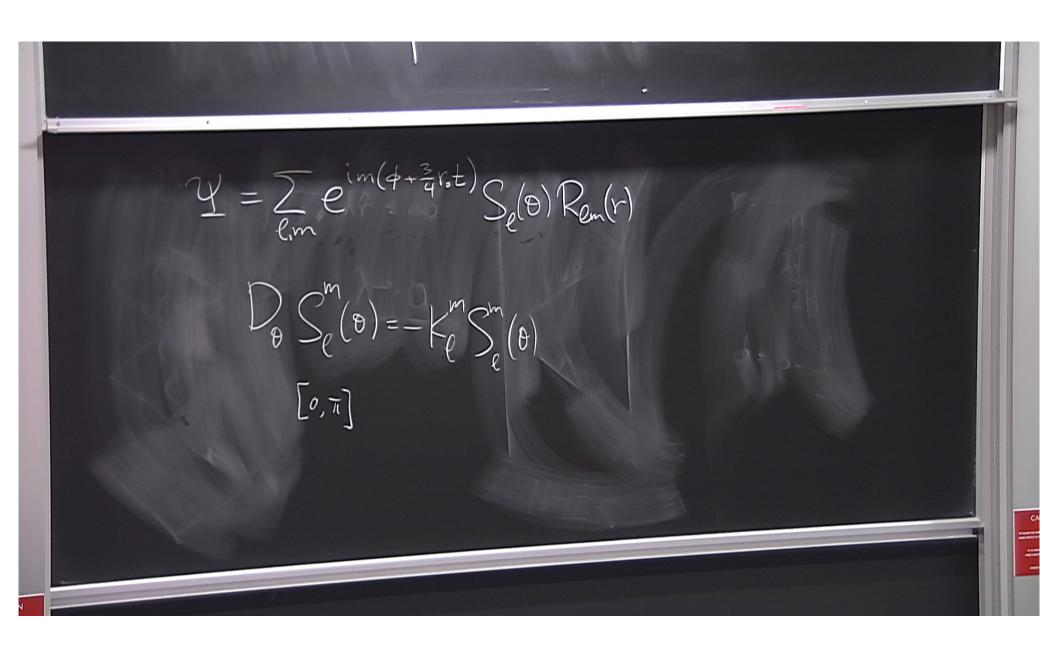
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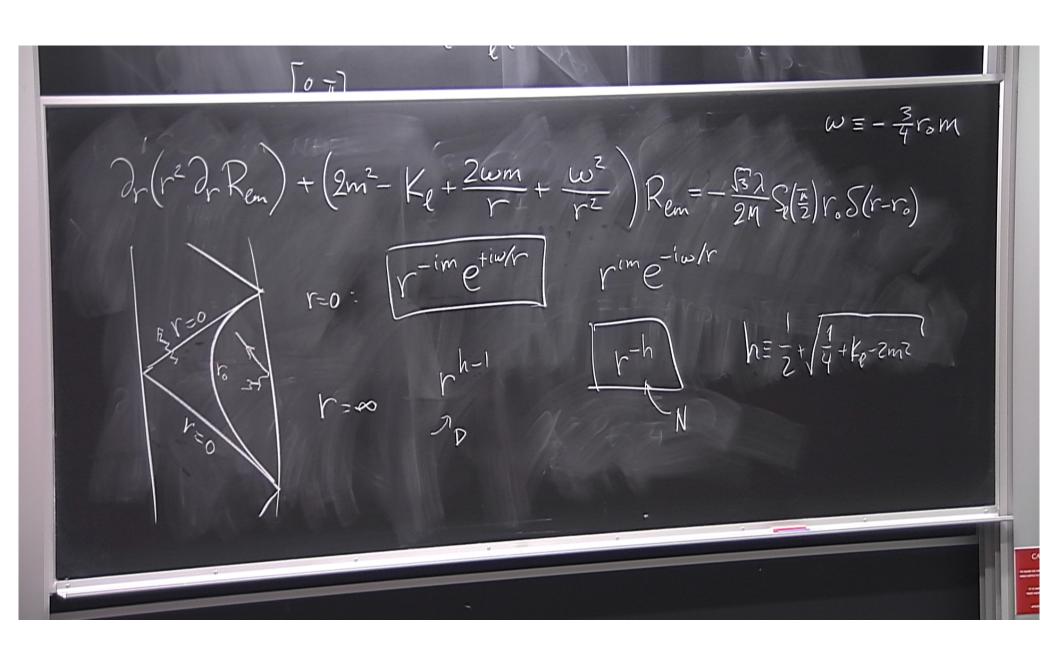
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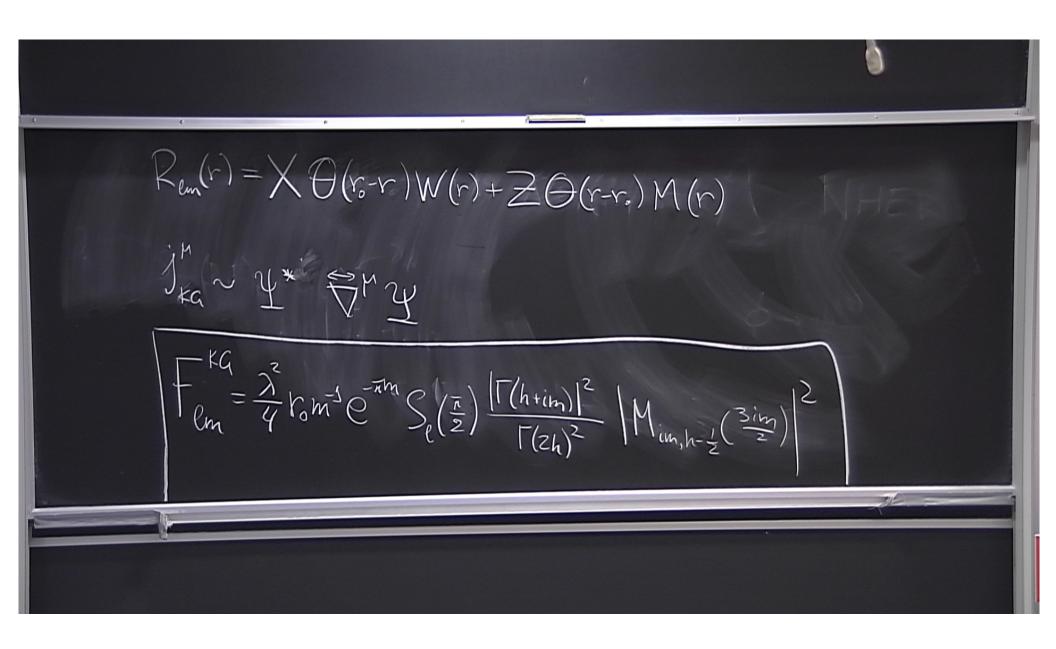
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$$\partial_{r}(r^{2}\partial_{r}R_{em}) + \left(2m^{2} - K_{\ell} + \frac{2\omega m}{r} + \frac{\omega^{2}}{r^{2}}\right)R_{em} = -\frac{53}{24}S(\frac{\pi}{2})r_{o}S(r-r_{o})$$

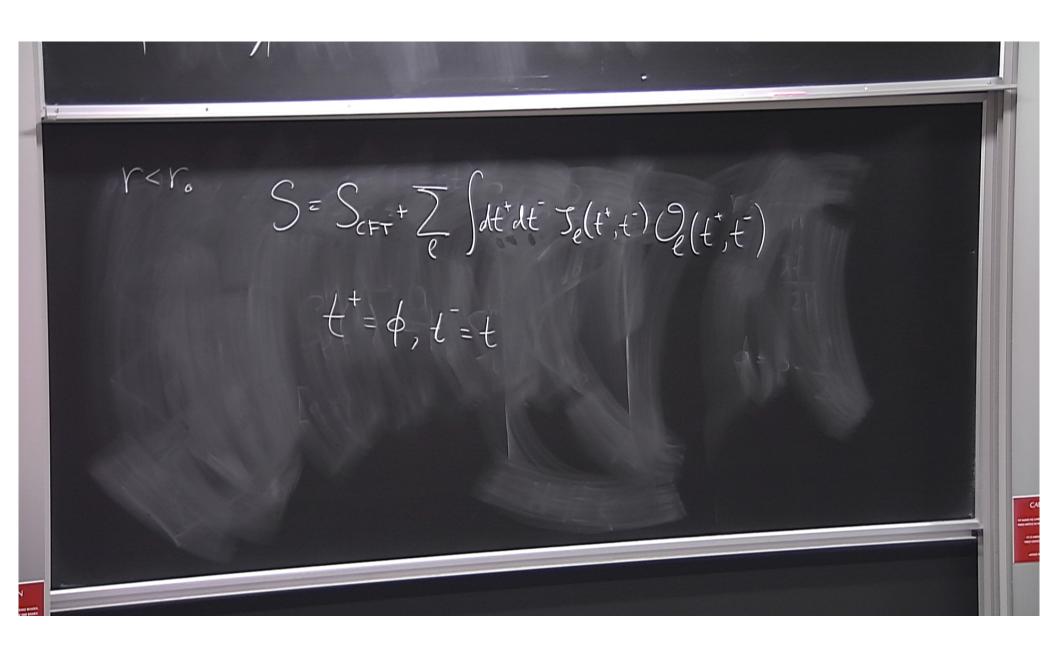
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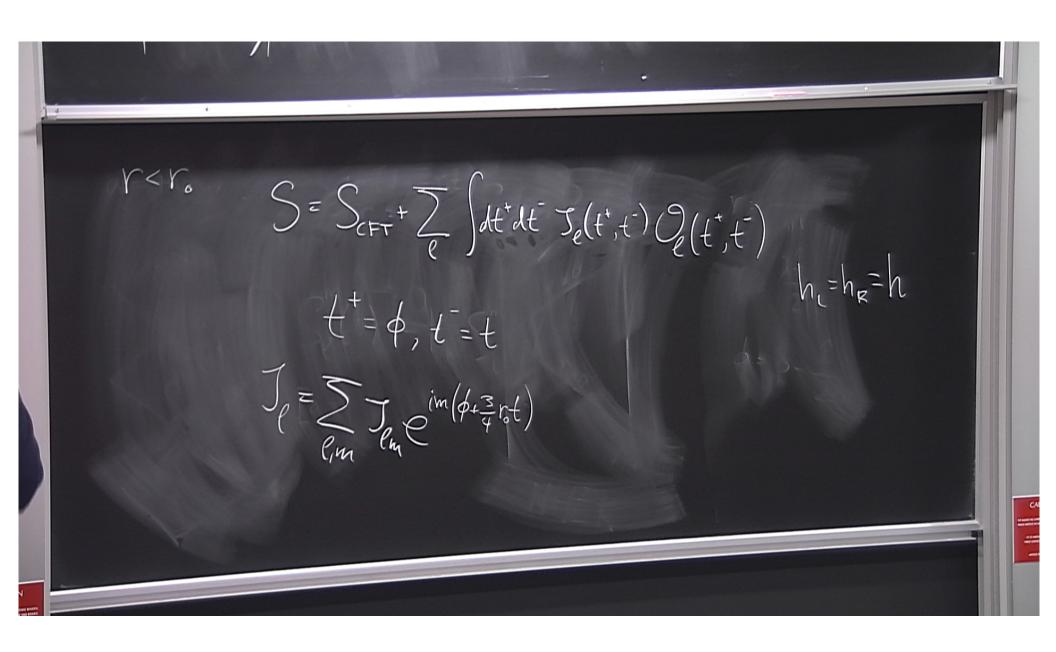
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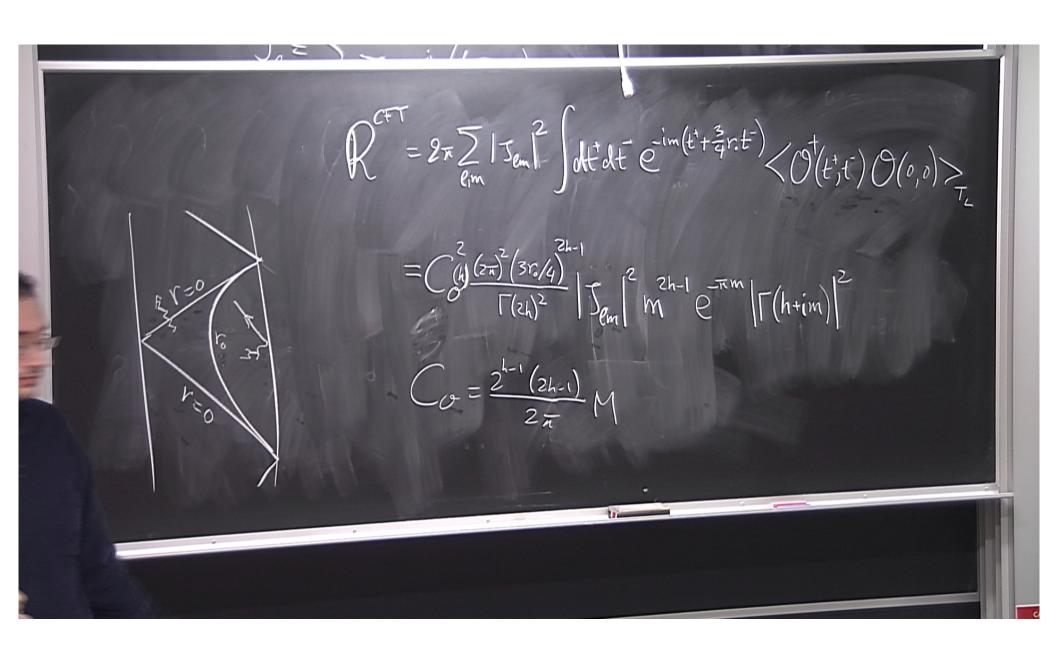
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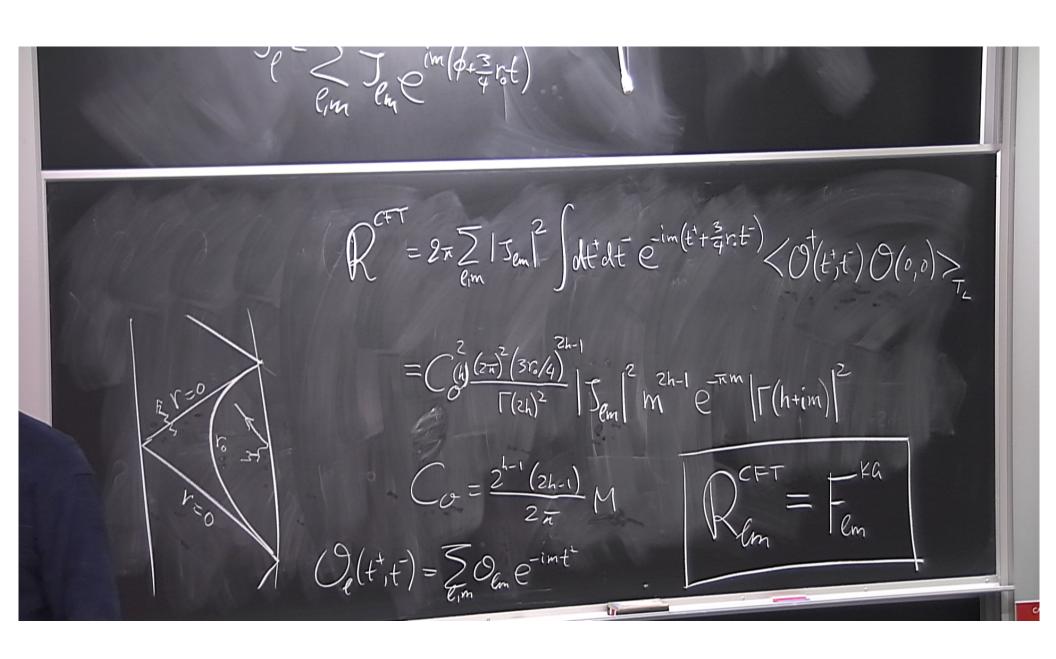
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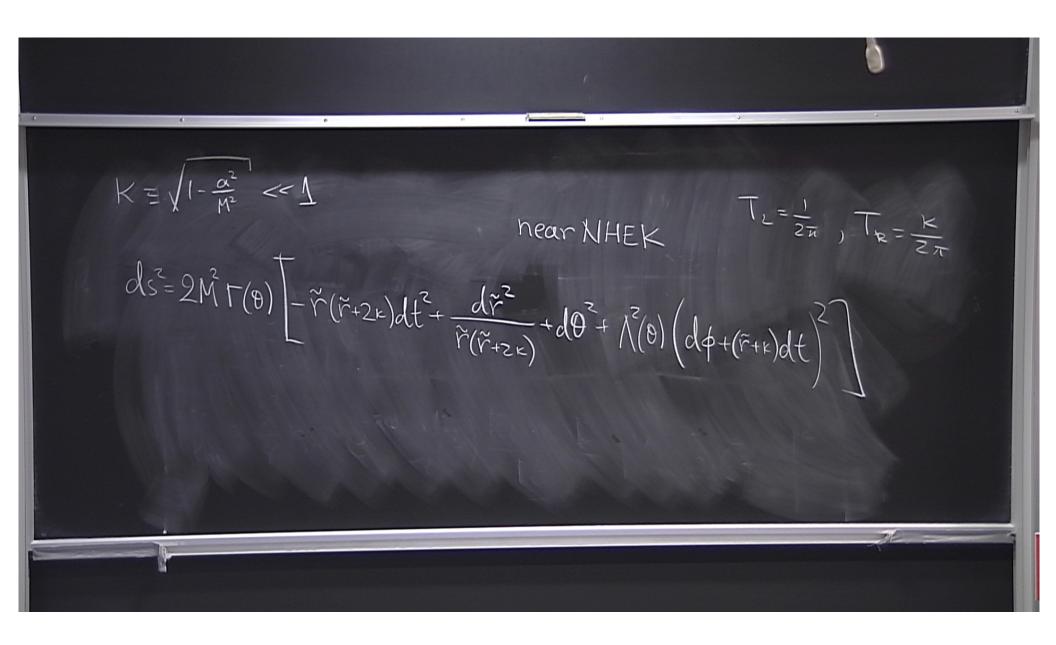
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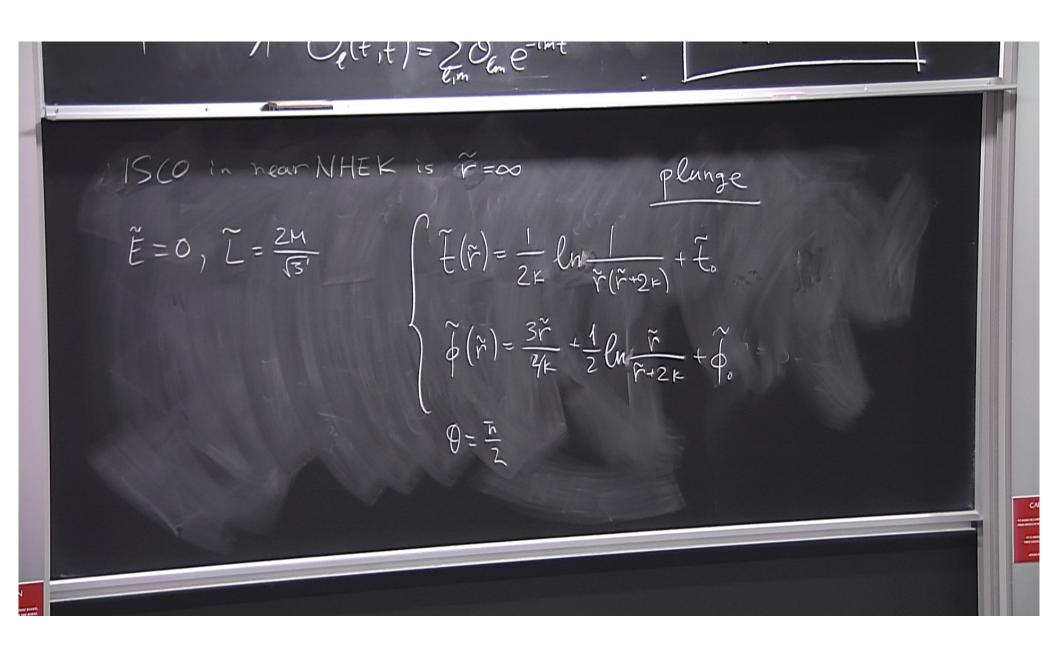
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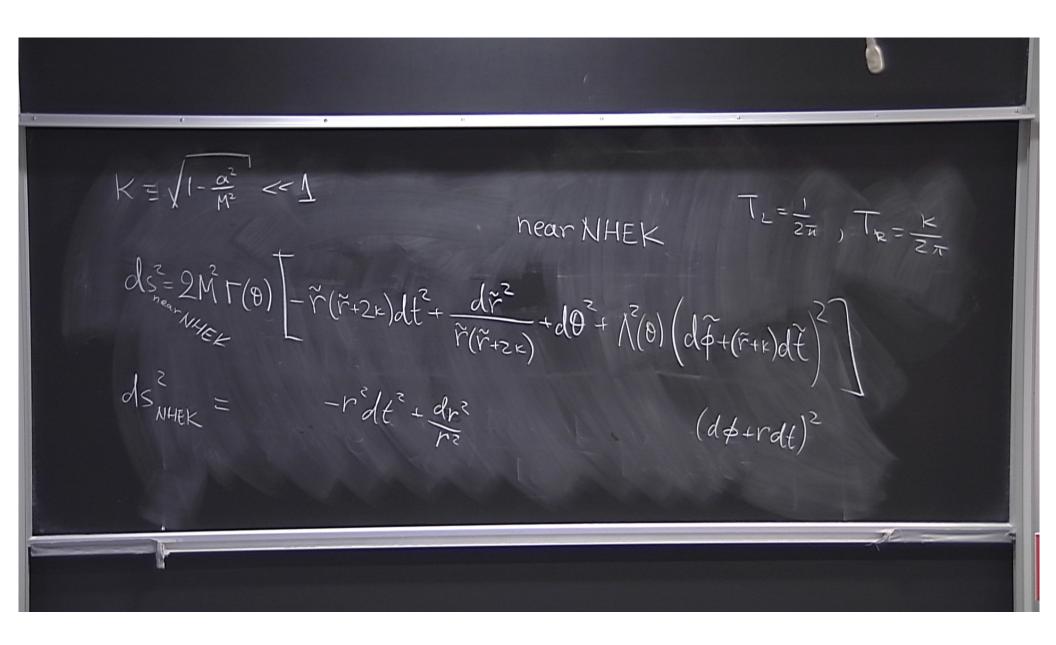
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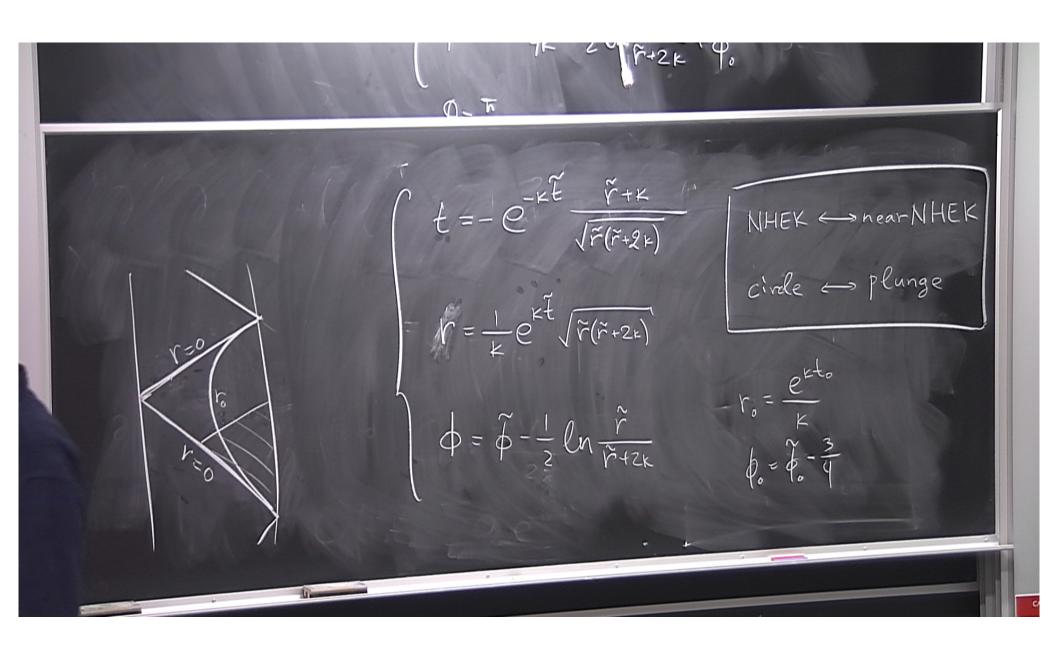
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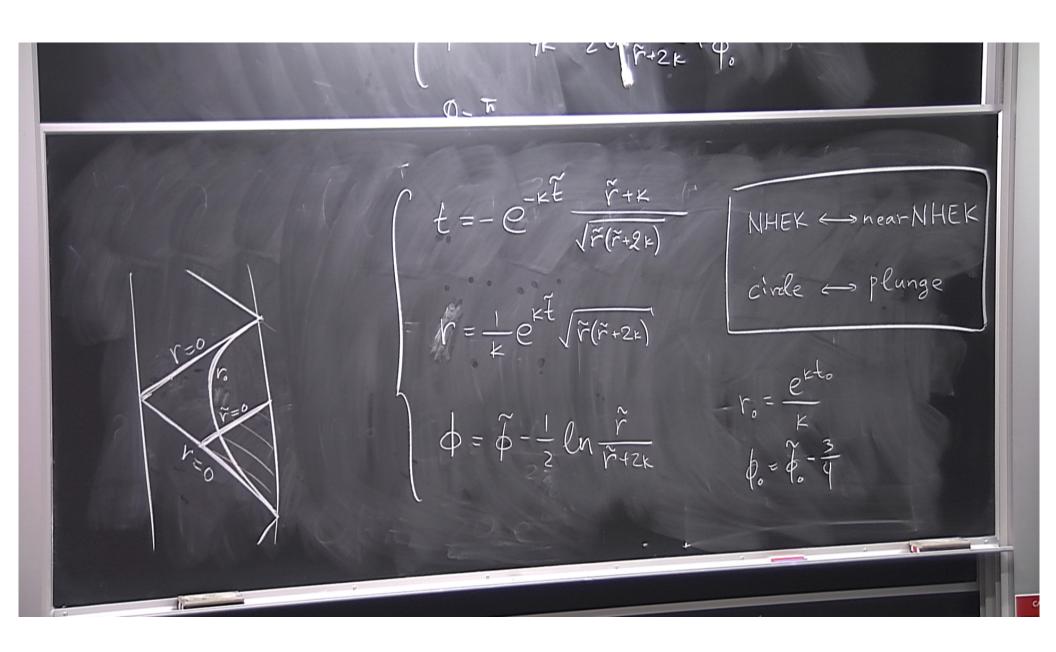
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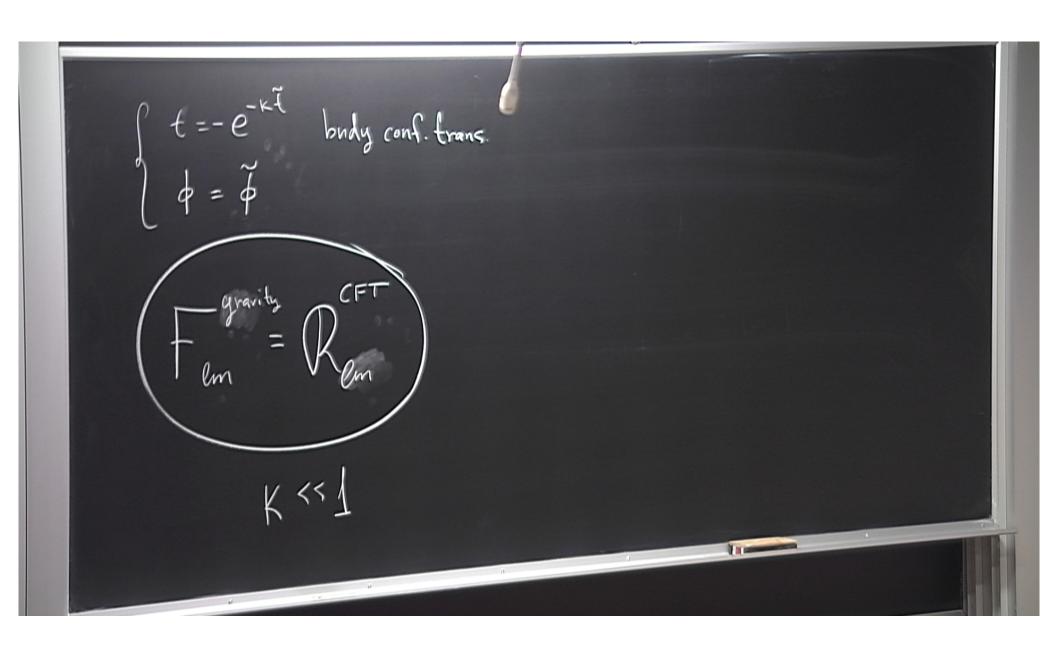
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