

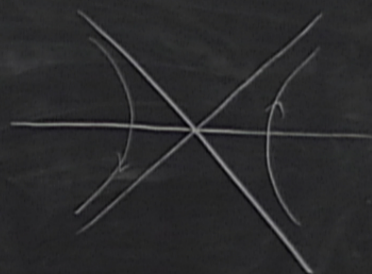
Title: Firewall evasion by quantum gravity

Date: Feb 13, 2014 02:30 PM

URL: <http://pirsa.org/14020136>

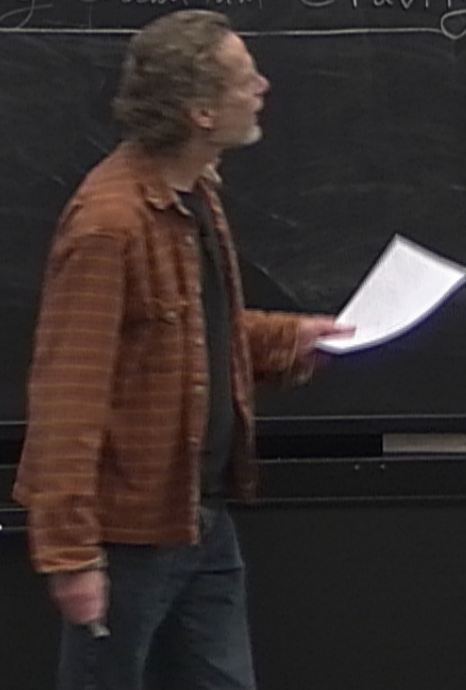
Abstract: It has been argued that if black hole evaporation is a unitary quantum process, then a black hole horizon must be cloaked by a "firewall", i.e. a highly excited state of local quantum fields. This reasoning is based on factorizing the Hilbert space into interior and exterior degrees of freedom. Such factorization ignores the Wheeler-deWitt constraint equation, which arises from the diffeomorphism invariance of quantum gravity. I will argue that this constraint evades the firewall.

Firewall Erasion by Quantum Gravity

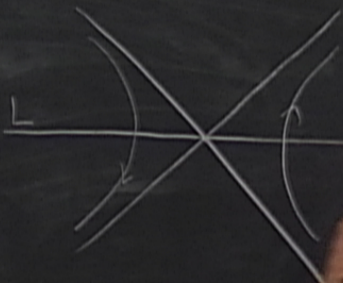


Minkowski

$|0\rangle$: vacuum of any
Stable Q F



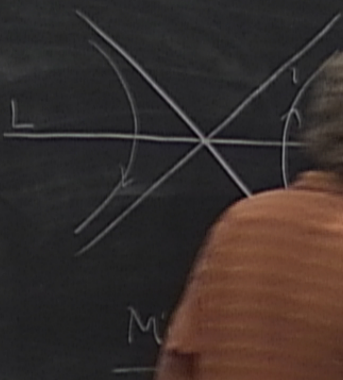
Transition by Quantum Gravity



Minkowski

$|0\rangle$: vacuum of any Lorentz-inv
Stable QFT
 $\text{Tr} |0\rangle\langle 0| = e^{\frac{-2\pi H_{\text{horiz}}}{\hbar}}$

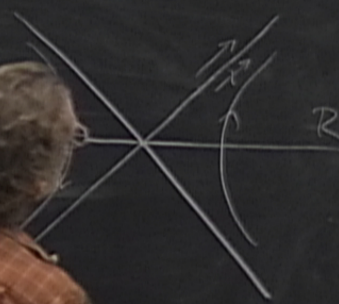
Transition by Quantum Gravity



$|0\rangle$: vacuum of any Lorentz-inv
Stable QFT

$$\text{Tr}_L |0\rangle\langle 0| = e^{\frac{-2\pi H_{\text{horiz}}}{\hbar}}$$

Analysis by Quantum Gravity



Penrose

$|0\rangle$: vacuum of any Lorentz-inv
Stable QFT

$$\text{Tr}(|0\rangle\langle 0|) = e^{\frac{-2\pi H_{\text{horiz}}}{\hbar}}$$

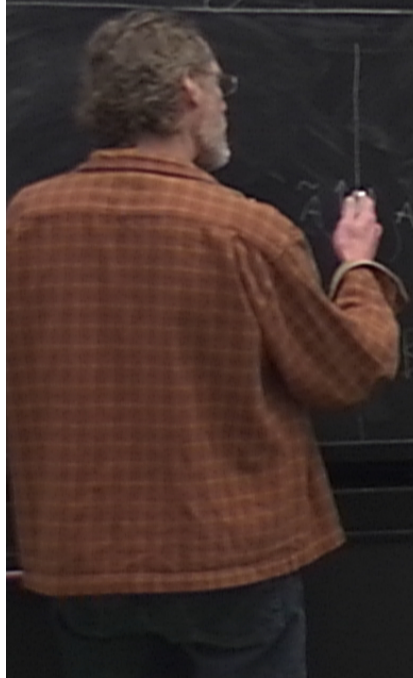
Transition by Quantum Gravity

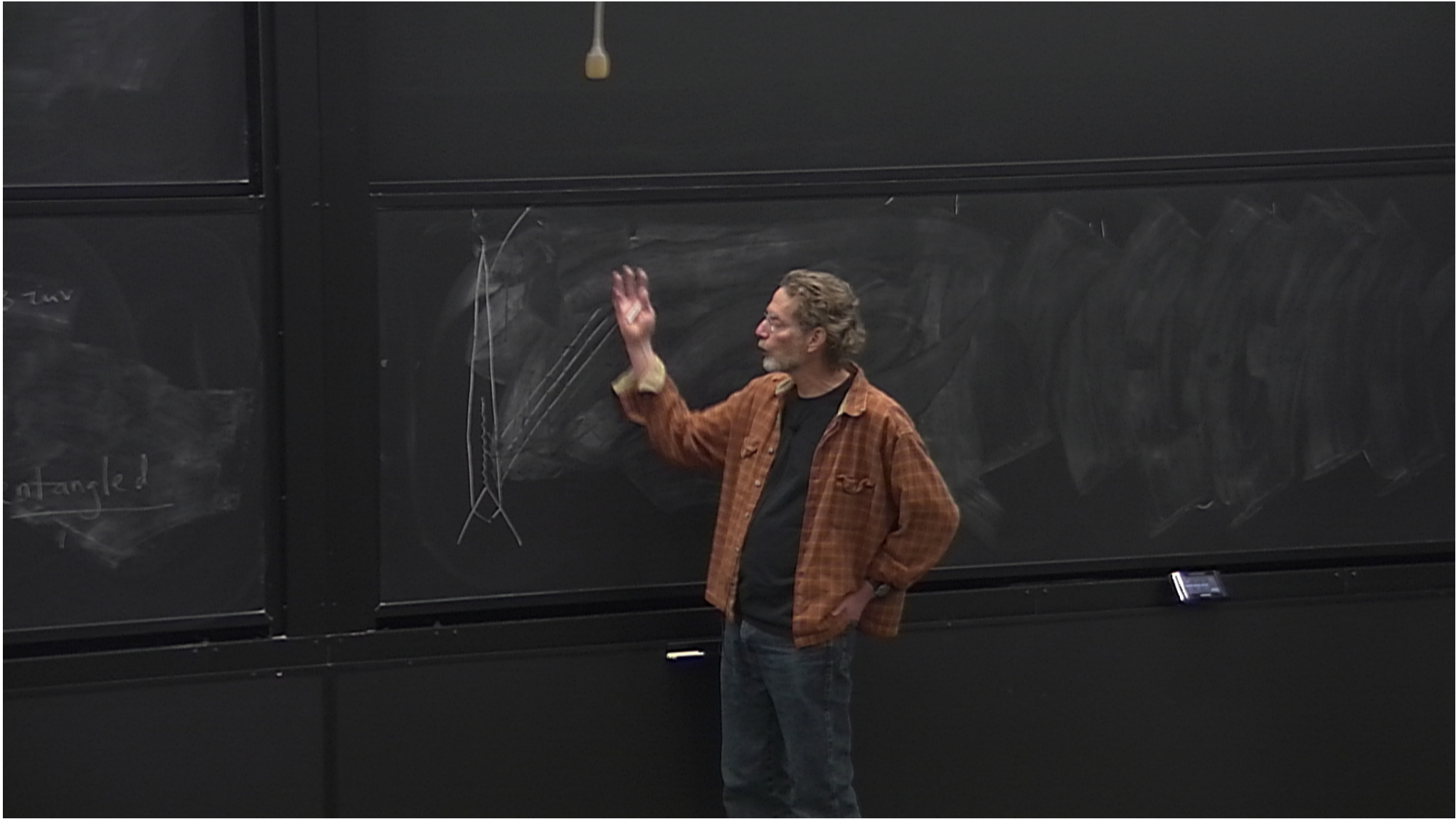


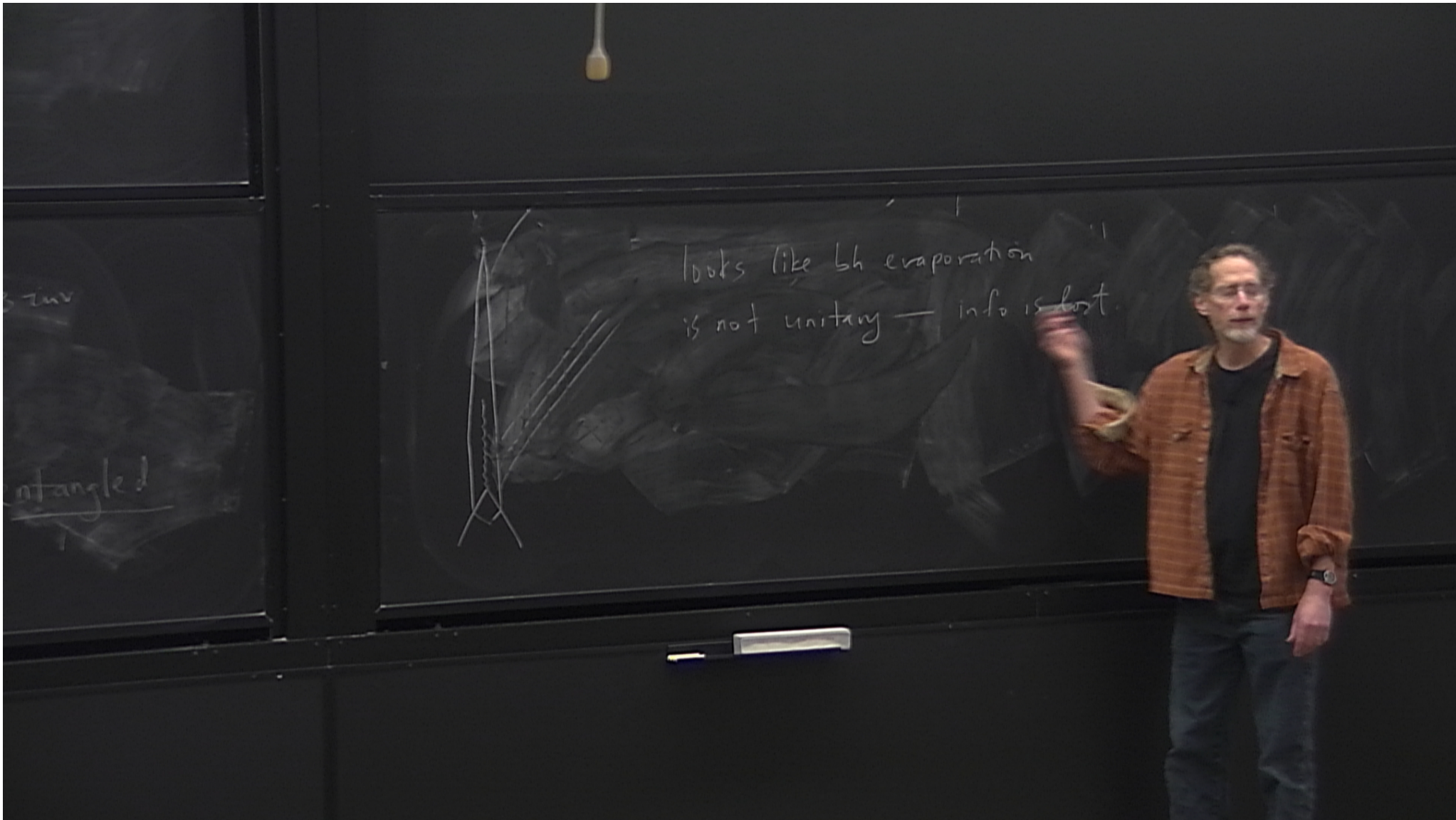
$|0\rangle$: vacuum of any Lorentz-inv
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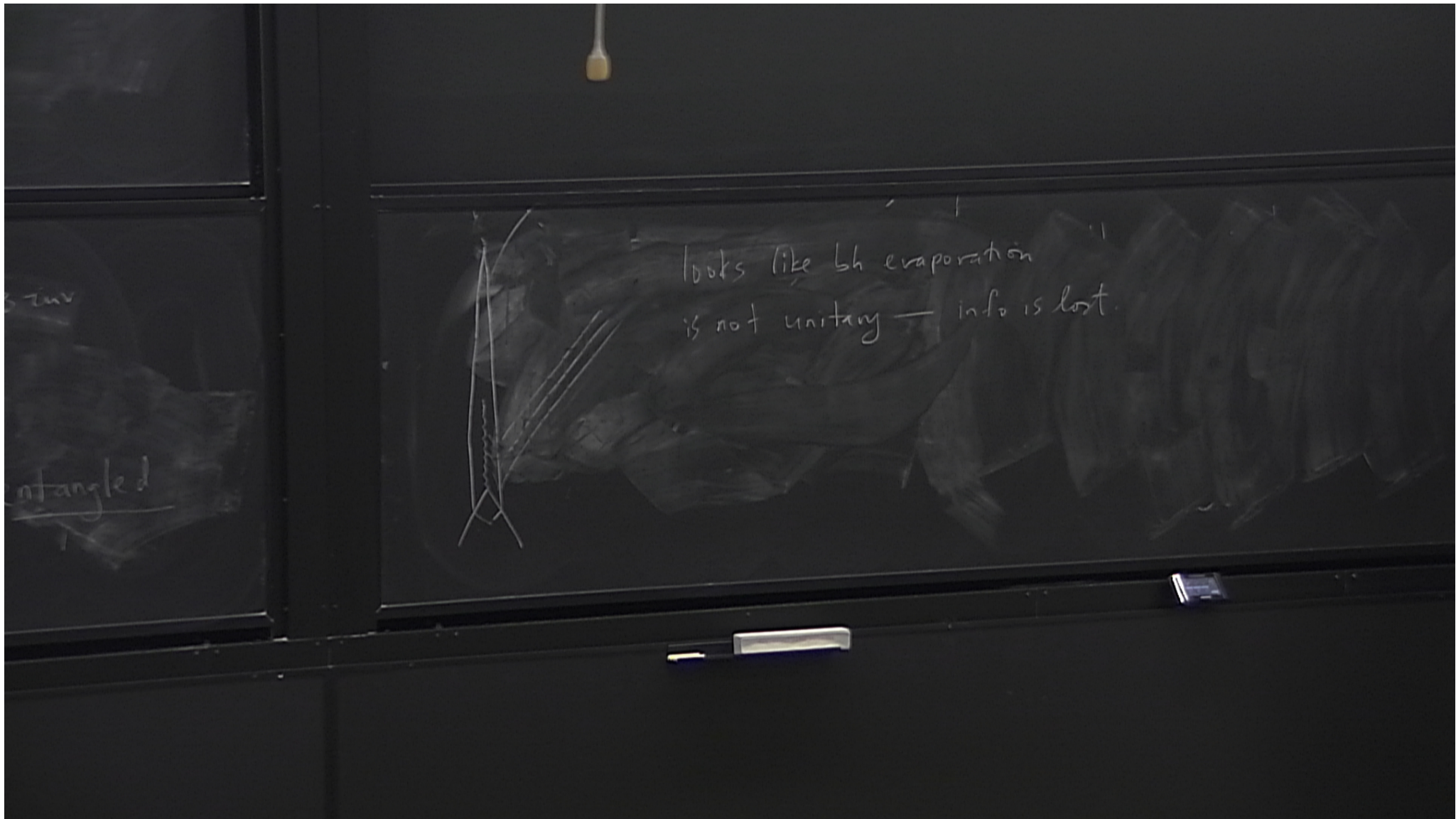
$$\text{Tr}_L |0\rangle\langle 0| = e^{\frac{-2\pi H_{\text{horiz}}}{\hbar}}$$

$$\Psi_{A\tilde{A}} \sim \sum_n e^{-\pi n \omega} |n\rangle_A |n\rangle_{\tilde{A}} \quad \text{entangled}$$





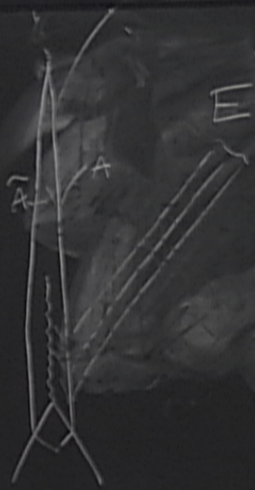
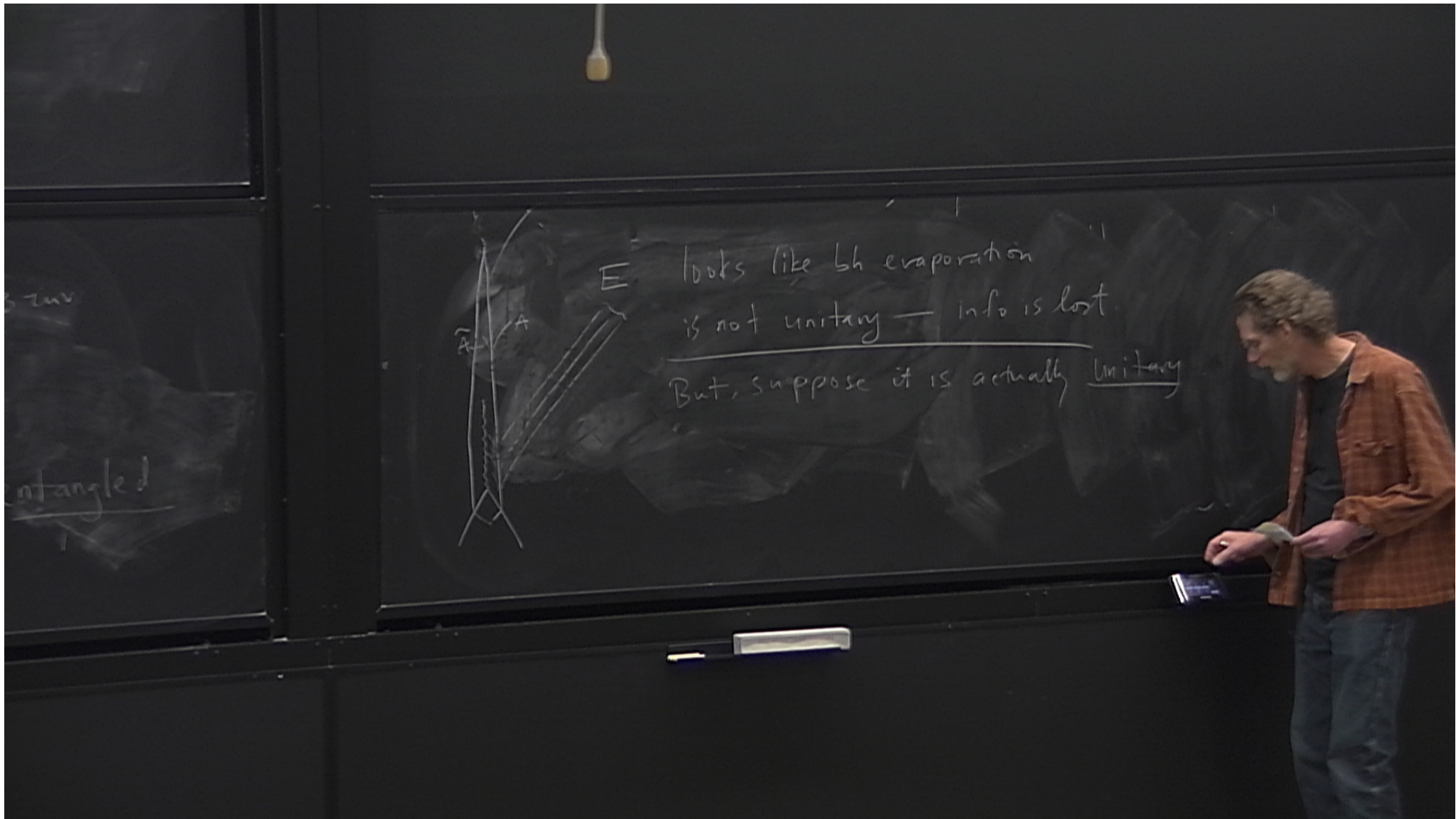




looks like bh evaporation
is not unitary — info is lost.

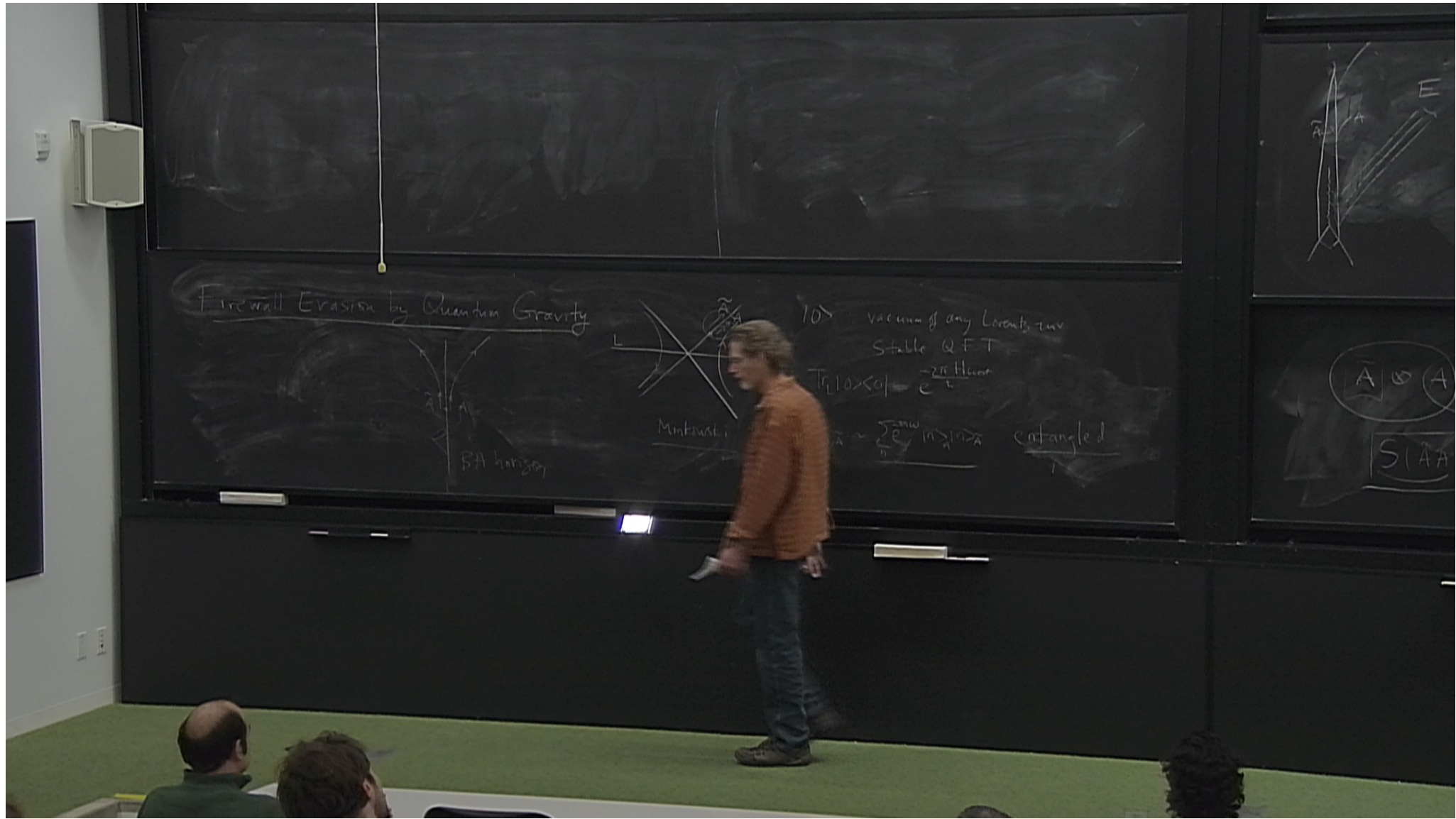


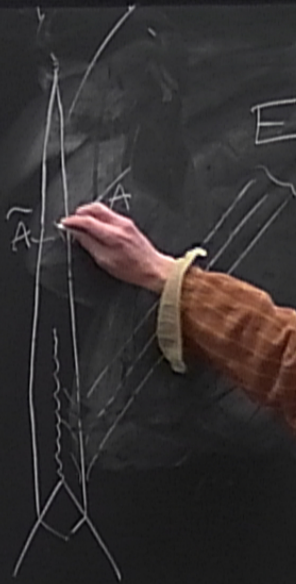
entangled



looks like bh evaporation
is not unitary — info is lost.
But, suppose it is actually unitary

entangled





E looks like lh evaporation
is not unitary — info is lost.

+, suppose it is actually unitary



E looks like bh evaporation
is not unitary — info is lost.

But, suppose it is actually unitary



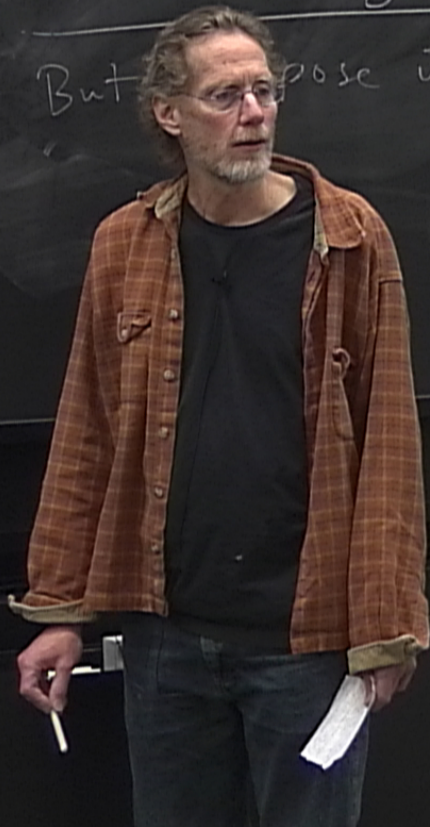
E looks like bh evaporation
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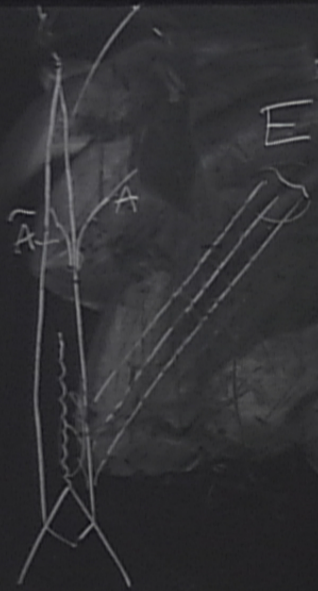
But, use it is actually unitary



E looks like bh evaporation
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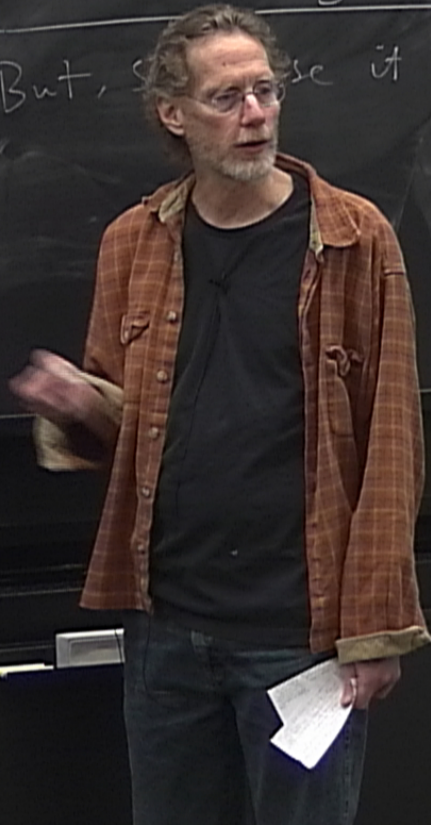
But suppose it is actually unitary





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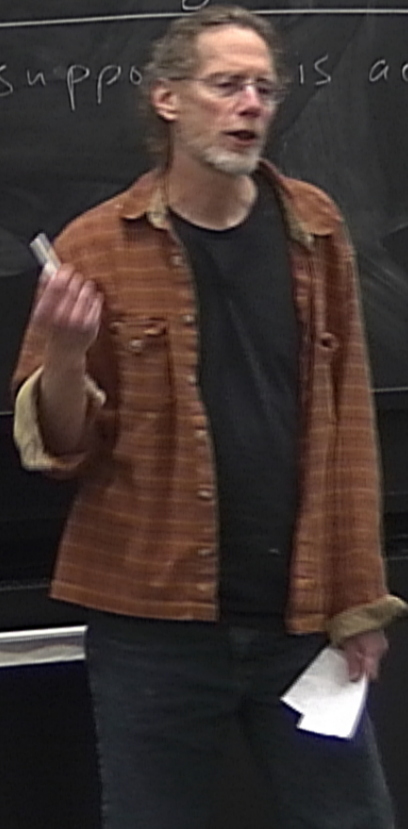
But, since it is actually unitary





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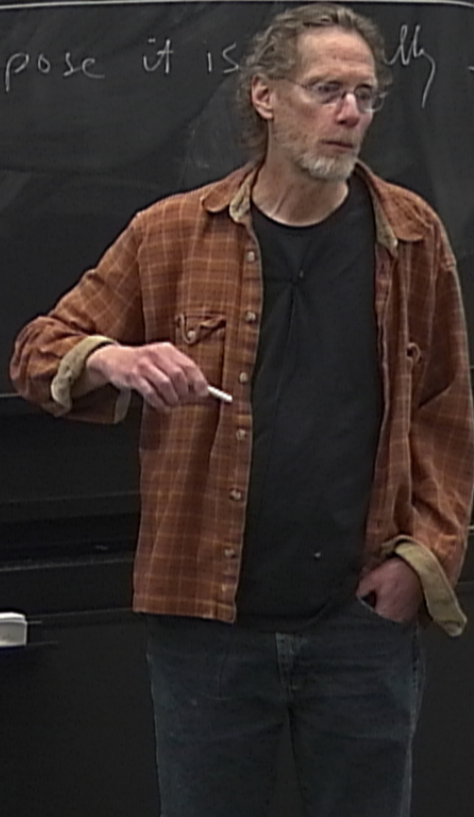
But, suppo is actually unitary





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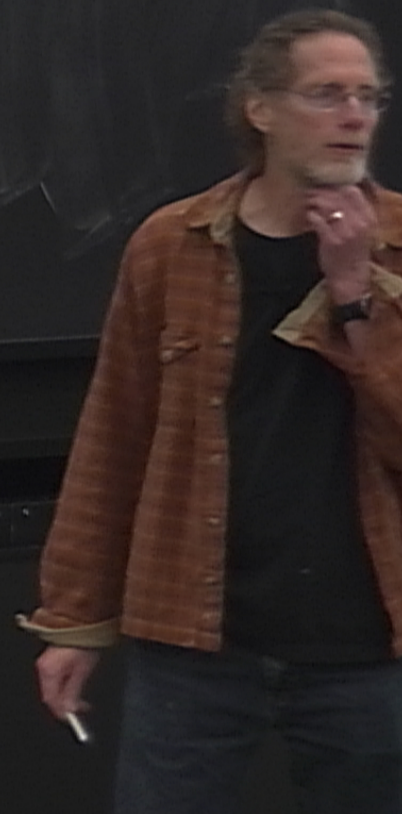
But, suppose it is unitary





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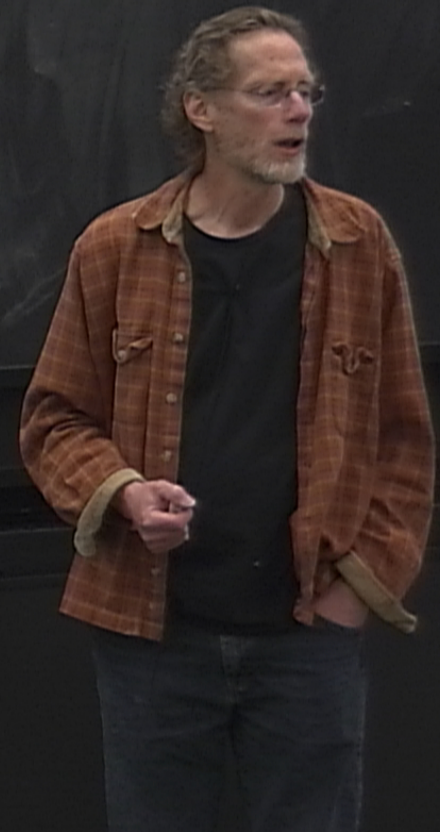
But, suppose it is actually unitary





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But, suppose it is actually unitary



BH horizon

Marolf (2009) 044010
PRD 79

BH horizon

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PRD 79.

Asym. flat & Asym. AdS.

BH horizon

Mandl (2009) 044010
PRD 79.

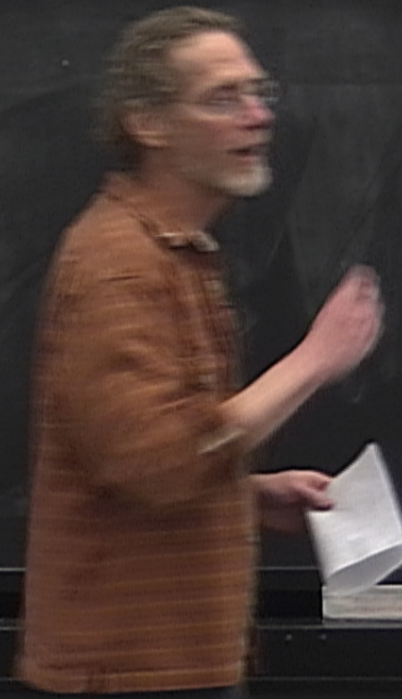
Asym. flat & Asym. AdS.

BH horizon

Marolf (2009) 044010
PRD 79.

Asym. flat & Asym. AdS.

in diffeo invariant theory



BH horizon

4010

in diffeo invariant theory

$$\hat{C}_m = G_m^0 - T_m^0 \quad ; \text{ constraints, generate diffeos.}$$

AdS.

H

free invariant theory

$$C_\mu = G_\mu^0 - T_\mu^0 \quad ; \quad \text{constraints, generate diffeos.}$$

$$H = \int_\Sigma N^\mu C_\mu + H_{\partial\Sigma}$$

physical observable

→ evolve by

$$[C_\mu, C_\nu] = 0$$

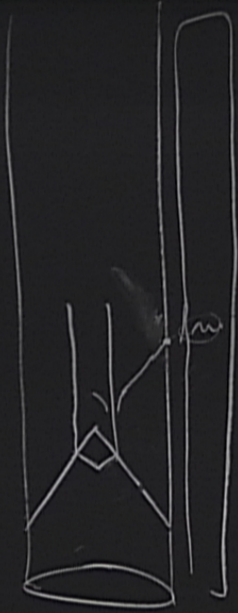
$$[H_{\partial\Sigma}, C_\mu] = 0$$

free invariant theory

$$C_\mu = G_\mu^0 - T_\mu^0 \quad \text{constraints, } g \text{ differs.}$$

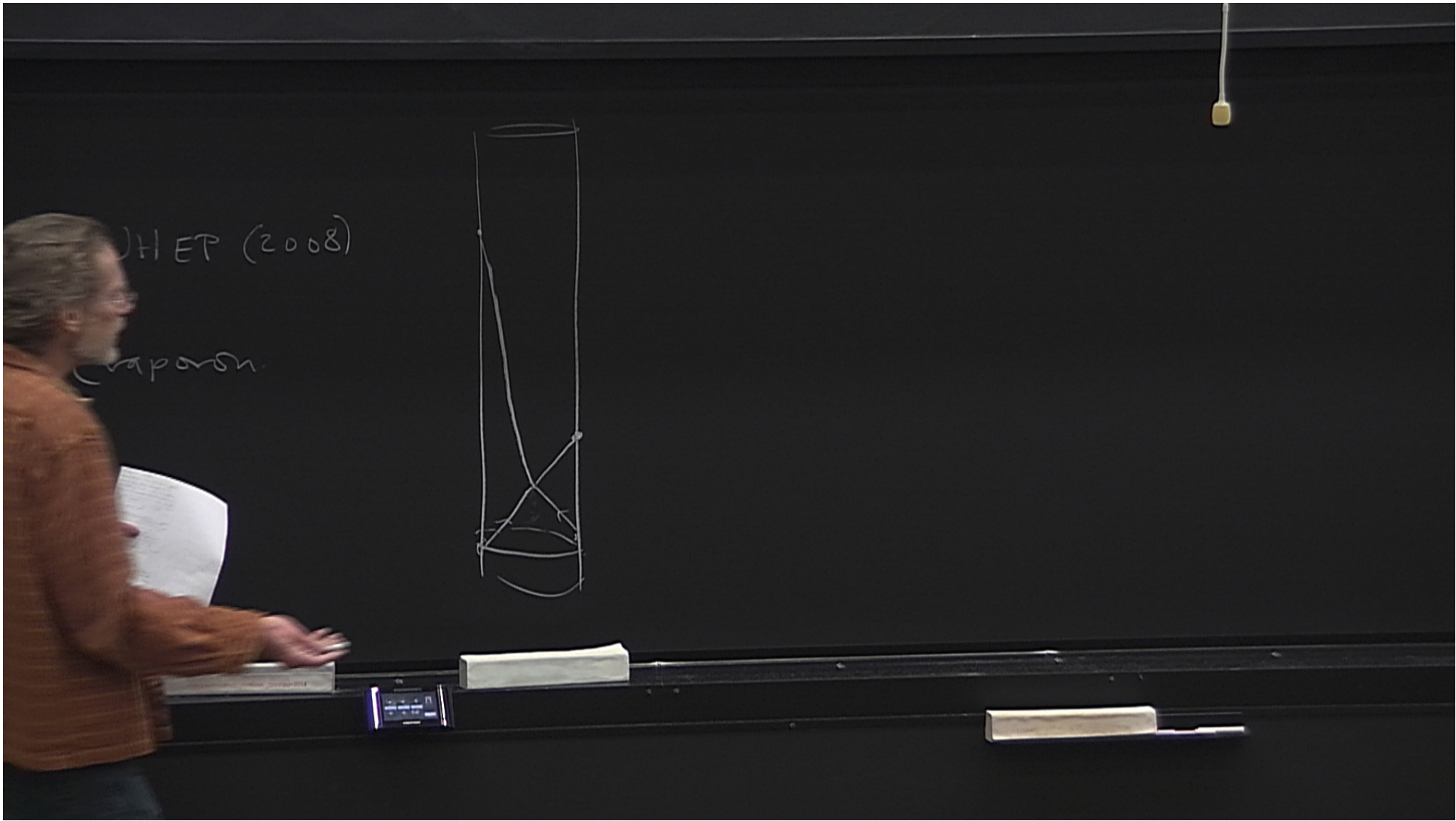
$$H = \int_\Sigma N^\mu C_\mu + H_{\partial\Sigma}$$

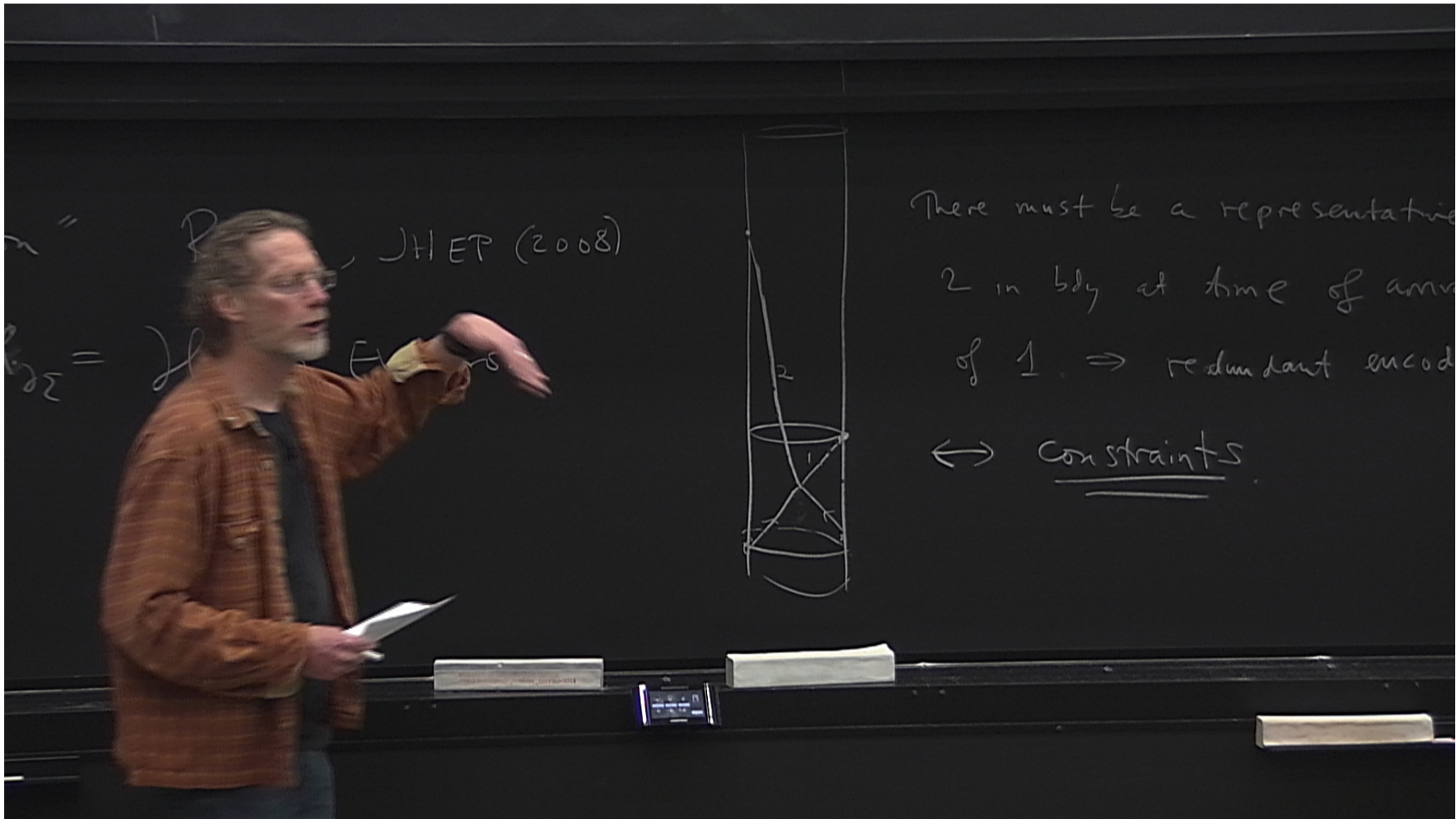
physical observables commute w C_μ . $[\mathcal{O}, C_\mu] = 0$
evolve by bdy term. $[H, \mathcal{O}] = [H_{\partial\Sigma}, \mathcal{O}]$



"Evaporon" Rocha, JHEP (2008)

$$\text{let } \mathcal{H}_{\Sigma\Sigma} = \mathcal{H}_{\text{AdS}} \otimes \text{Evaporon}$$

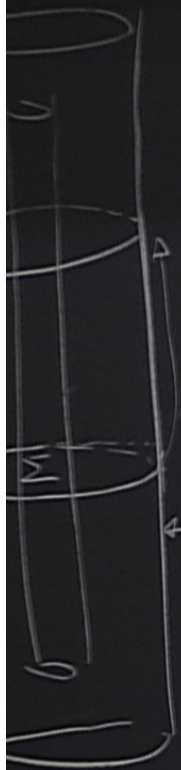




P , JHEP (2008)
 $\mathcal{H} = \mathcal{H} \oplus \mathcal{H}_s$



There must be a representative
2 in bdy at time of arrival
of 1 \Rightarrow redundant encoding
 \leftrightarrow Constraints



$$\begin{array}{c}
 \nearrow \mathcal{g}^+ \\
 \searrow \mathcal{g}^-
 \end{array}
 \xrightarrow{i^0}
 \left(\tilde{A} \otimes A \otimes E \right)$$

$A_{\infty}^{(t)}$: alg. of bdy observables

4 spin j systems. $\rightarrow \mathcal{H}_{\text{kin}} = j \otimes j \otimes j \otimes j$

Symm: Rotation inv.

Constraints: $\vec{J}_{\text{tot}} |4\rangle = 0$

observables: $[0, \vec{J}_{\text{tot}}] = 0$

4 Spin j systems.

$$\mathcal{H}_{kin} = (j \otimes j) \otimes (j \otimes j)$$

$$(0 \oplus 1 \oplus 2 \oplus \dots \oplus 2j) \otimes (0 \oplus 1 \oplus 2 \oplus \dots)$$

Symm. Projection inv.

Constram $\vec{J}_{tot} |4\rangle = 0$

obs $[0, \vec{J}_{tot}] = 0$

$$\Pi_{\text{single}} \mathcal{H}_{kin} = \mathcal{H}_{\text{phys}}$$

4 spin j systems

$$\mathcal{H}_{kin} = (j \otimes j) \otimes (j \otimes j)$$

$$(0 \oplus 1 \oplus 2 \oplus \dots \oplus 2j) \otimes (0 \oplus 1 \oplus 2 \oplus \dots)$$

Symm: Rotation inv.

$$\Pi_{\text{singlet}} \mathcal{H}_{kin} = \mathcal{H}_{\text{phys}}$$

Constraints: $\vec{J}_{tot} |\psi\rangle = 0$

$$0 \oplus 0_1 \oplus 0_2 \oplus \dots \oplus 0_{2j} \oplus \text{NON SINGLET}$$

observables: $[0, \vec{J}_{tot}] = 0$