

Title: Sequestering the Standard Model Vacuum Energy

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Abstract: We propose a very simple reformulation of General Relativity, which completely sequesters from gravity {\it all} of the vacuum energy from a matter sector, including all loop corrections and renders all contributions from phase transitions automatically small. The idea is to make the dimensional parameters in the matter sector functionals of the 4-volume element of the universe. For them to be nonzero, the universe should be finite in spacetime. If this matter is the Standard Model of particle physics, our mechanism prevents any of its vacuum energy, classical or quantum, from sourcing the curvature of the universe. The mechanism is consistent with the large hierarchy between the Planck scale, electroweak scale and curvature scale, and early universe cosmology, including inflation. Consequences of our proposal are that the vacuum curvature of an old and large universe is not zero, but very small, that $w_{DE} \rightarrow 1$ is a transient, and that the universe will collapse in the future.

Q Padilla

hep-th/1309.6562

$$S = \int d^4x \sqrt{g} \left(\frac{M_p^2}{2} R - \mathcal{L}(g^{\mu\nu}, \psi, \phi) \right)$$

$$V_{\text{vac}} = \langle 0 | \mathcal{H} | 0 \rangle = 0 + \text{circle with slash}$$

CAUTION

DO NOT TOUCH THE CONTENTS OF THIS BOX

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$$S = \int d^4x \sqrt{g} \left(\frac{M_{\text{Pl}}^2}{2} R - \mathcal{L}(\psi, \psi^\dagger) \right)$$

$$V_{\text{vac}} = \langle 0 | \mathcal{H} | 0 \rangle = 0 + \underbrace{\quad}_{M_{\text{cv}}}$$

CAUTION

DO NOT TOUCH THE BOARD OR THE BOARDER

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$$S = \int d^4x \sqrt{g} \left(\frac{M_p^2}{2} R - \Lambda - Z(g^{\mu\nu}, \psi, \phi) \right)$$

$$\Lambda + V_{vac} = \Lambda \langle 0 | \mathcal{L} | 0 \rangle = \text{[diagrams]} + 1$$

$\frac{1}{\epsilon} \left(c_1 M_{\mu\nu}^{\otimes 4} + c_2 M_{\mu\nu}^{\otimes 2} \right) + \text{[triangle diagram]} \in$

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$$S = \int d^4x \sqrt{g} \left(\frac{M_p^2}{2} R + \Lambda \right) Z(g, \psi, \phi)$$

$$\Lambda + V_{vac} = \Lambda \langle 0 | \infty | 0 \rangle = \text{diagrams} + 1$$

$\frac{1}{\Lambda} \left(c_1 M_{\mu\nu}^{\otimes 4} + c_2 M_{\mu\nu}^{\otimes 8} \right) + \text{triangle diagram} \in$

$$S = S_{\Lambda=0} +$$

$$\int \sqrt{g} \Lambda$$

$$\int \phi^4 \sqrt{g}$$

1

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$$S = \int d^4x \sqrt{-g} \left(\frac{M_{\text{Pl}}^2}{2} R - \Lambda - Z(g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi) \right)$$

$$\langle 0 | \mathcal{Z} | 0 \rangle = \Lambda_1 + \Lambda_2 + \Lambda_3$$

$$\frac{M_{\text{phys}}}{M_{\text{Pl}}} = \lambda \frac{m}{M_{\text{Pl}}} \quad \bar{g}_{\mu\nu} = \lambda^2 g_{\mu\nu}$$

$$S = \int d^4x \left[\sqrt{g} \left(\frac{M_{\text{Pl}}^2}{2} R - \Lambda \right) - \sqrt{g} \mathcal{L}(\bar{g}^{\mu\nu} \bar{\psi} \partial_\mu \bar{\psi}) \right. \\ \left. - \sqrt{g} \lambda^4 \mathcal{L}(\bar{g}^{\mu\nu} \psi \partial_\mu \psi) \right]$$

$$S = \int d^4x \sqrt{g} \left(\frac{M_p^2 R}{2} - \Lambda \right) - \sqrt{g} \mathcal{L}(g, \psi, \bar{\psi})$$

$$- \sqrt{g} \bar{\psi} \gamma^\mu \nabla_\mu \psi + \sqrt{g} \bar{\psi} \gamma^\mu \gamma^5 \nabla_\mu \psi$$

$$\langle 0 | Z | 0 \rangle = \lambda^4 \left(\bar{\Lambda}_1 + \bar{\Lambda}_2 + \bar{\Lambda}_3 \right)^{1,1,1}$$

$$\frac{M_{\text{phys}}}{M_{\text{Pl}}} = \lambda \frac{m}{M_{\text{Pl}}}$$

$$\bar{g}_{\mu\nu} = \lambda^2 g_{\mu\nu}$$

$$S = \int d^4x \left[\sqrt{g} \left(\frac{M_{\text{Pl}}^2}{2} R - \Lambda \right) - \sqrt{g} \bar{\psi} \gamma^\mu \partial_\mu \psi - \sqrt{g} \lambda^4 \bar{\psi} \gamma^\mu \partial_\mu \psi \right]$$

$$\langle 0 | Z | 0 \rangle = \lambda^4 \left(\bar{\Lambda}_1 + \bar{\Lambda}_2 + \bar{\Lambda}_3 \right) \gamma^\mu \partial_\mu \psi$$

$$S = \int d^4x \sqrt{g} \left(\frac{M_p^2}{2} R - \Lambda - \frac{\lambda^4}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi \right)$$

$$\int d^4x \sqrt{g} = \frac{\sigma^4}{M^4 \lambda^4}$$

$$+ \sigma \left(\frac{\Lambda}{M^4 \lambda^4} \right)$$

$$S = \int d^4x \left(\frac{M_p^2}{2} R - \Lambda - \frac{1}{2} \lambda^2 (g^{\mu\nu} - \eta^{\mu\nu}) \right)$$

$$\int d^4x \sqrt{g} = \frac{\sigma}{M^4 \lambda^4}$$

$$\int d^4x \sqrt{g} T^{\mu\nu} = \frac{4 \Lambda \sigma}{\lambda^4 M^4}$$

$$+ \sigma \left(\frac{\Lambda}{M^4 \lambda^4} \right)$$

$$S = \int d^4x \sqrt{g} \left(\frac{M_p^2}{2} R - \Lambda - \frac{1}{2} \lambda^2 (g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi) \right)$$

$$\int d^4x \sqrt{g} = \frac{\sigma}{M^4 \lambda^4}$$

$$\int d^4x \sqrt{g} T^{\mu}_{\mu} = \frac{4 \Lambda \sigma}{M^4 \lambda^4} \Rightarrow \Lambda = \frac{1}{4} \frac{\int d^4x \sqrt{g} T^{\mu}_{\mu}}{\int d^4x \sqrt{g}} = \frac{1}{4} \frac{\int d^4x \sqrt{g} T^{\mu}_{\mu}}{\int d^4x \sqrt{g}}$$

$$S = \int d^4x \sqrt{g} \left(\frac{M_p^2}{2} R - \Lambda - \frac{1}{2} \lambda^2 (g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi) \right)$$

$$\int d^4x \sqrt{g} = \frac{\sigma}{M^4 \lambda^4}$$

$$\int d^4x \sqrt{g} T^{\mu}_{\mu} = \frac{4 \Lambda \sigma}{M^4 \lambda^4} \Rightarrow \Lambda = \frac{1}{4} \frac{\int d^4x \sqrt{g} T^{\mu}_{\mu}}{\int d^4x \sqrt{g}} = \frac{1}{4} \langle T^{\mu}_{\mu} \rangle$$

$$\langle 0 | \mathcal{Z} | 0 \rangle = \Lambda_1 + \Lambda_2 + \Lambda_3$$

$$M_b^2 G_v^M = T_v^M - \frac{1}{4} \langle T_v^M \rangle \delta_v^M$$

$$S = \int d^4x \left[\frac{1}{2} (g_{\mu\nu} \partial^\mu \phi \partial^\nu \phi - \Lambda) \right] - \frac{1}{4} (g_{\mu\nu} \partial^\mu \phi \partial^\nu \phi - \Lambda)$$

$$\langle 0 | \mathcal{Z} | 0 \rangle = \Lambda_1 + \Lambda_2 + \Lambda_3$$

$$\mathcal{L} = \frac{\mathcal{L}_0 + \langle 0|V|0 \rangle + \mathcal{L}_{loc}}{}$$

$$\langle \mathcal{L}_0 + \langle 0|V|0 \rangle \rangle = \mathcal{L}_0 + \langle 0|V|0 \rangle$$

$$T_{44} = \frac{\delta}{\delta g_{\mu\nu}} \int^{\text{Matter}} \mathcal{L}_{loc}$$

$$g_{\mu\nu} \rightarrow \frac{1}{\Omega^2} g_{\mu\nu}$$

$$\lambda \rightarrow \Omega \lambda$$

$$\Lambda \rightarrow \Omega^4 \Lambda$$

$$\Lambda \rightarrow \Lambda - \alpha$$

$$\mathcal{L} \rightarrow \mathcal{L} + \alpha \lambda^4$$

CAUTION

BEWARE OF CATCHES AND TRICKS

SEE THE INSTRUCTIONS FOR MORE

ADDITIONAL INFORMATION

$$g_{\mu\nu} \rightarrow \frac{1}{\Omega^2} g_{\mu\nu}$$

$$\lambda \rightarrow \Omega \lambda$$

$$\Lambda \rightarrow \Omega^4 \Lambda$$

$$\delta S \rightarrow \delta' \frac{\alpha}{M^4}$$

$$\Lambda \rightarrow \Lambda - \alpha$$

$$\mathcal{L} \rightarrow \mathcal{L} + \alpha \lambda^4$$

$$\alpha \propto \left(\frac{m_{\text{phys}}}{M_{\text{Pl}}}\right)^4$$

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$$S = \int d^4x \left(\frac{M_p^2}{2} R - \Lambda - \frac{\lambda^4}{2} g^{\mu\nu} g_{\mu\nu} \right)$$

$$\int d^4x \sqrt{-g} T^{\mu}_{\mu} \sim \frac{V_B}{\lambda^2} \int dt a^3 \rho$$

$$S = \int d^4x \left(\frac{M_p^2}{2} R - \Lambda - \frac{1}{\lambda^2} \mathcal{L}(g^{\mu\nu}, \psi, \partial\psi) \right)$$

$|g| \leq 1$

$$\int d^4x \sqrt{-g} T^{\mu}_{\mu} \sim$$

$$\frac{2}{1+w} \frac{V}{R}$$

$$\int dt a^3 \int_0^{\frac{2a}{Hw}} \frac{2a}{Hw}$$

$$\int dt \frac{1}{(t-t_{end})^2} \sim \frac{1}{(t-t_{end})^2}$$

$$w \leq -1$$

$$\langle \tau_m^M \rangle \sim M_{Pl}^2 H_{age}^2$$

$$\leq M_{Pl}^2 H_0^2$$

$$\langle \tau_m^M \rangle \sim M_{Pl}^2 H_{age}^2$$

$$\underbrace{NR^2} \leq \underbrace{M_{Pl}^2 H_0^2}$$