

Title: 13/14 PSI - Cosmology Review - Lecture 1

Date: Jan 27, 2014 11:30 AM

URL: <http://pirsa.org/14010065>

Abstract:

Why Cosmo?

1) "Standard Model"

Review

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Office 220

Why Cosmo?

1) "Standard Model"

Review  
Boyle  
, Office 330

## Why Cosmo?

1) "Standard Model"

2)  $\vec{M}$  Meet Max Sym, FRW, BH

3,4) FRW, kinematics + dynamics  
"ΛCDM"

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BBN, CMB, Dark M

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BBN, CMB, Dark Matter, Dark Energy, Matter, Anti-Matter

Week 3:

Inflation, QFT in Curved spacetime

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Inflation, QFT in Curved Spacetime  
Unruh Effect, BH Thermo,  
Cosmological Perts.

Initial Conditions

# Why Cosmo?

1) "Standard Model"

2)  $\vec{T}$  Meet Max Sym.

3,4) FRW, kinematics +  
" $\Lambda$ CDM"

Week 2:

BBN, CMB, Dark Matter, Dark Energy,  $\frac{\text{Matter}}{\text{Anti-Matter}}$

Week 3:

Inflation, QFT in Curved Spacetime  
Unruh Effect, BH Thermo,  
Cosmological Perts.

Initial Conditions, Hartle + Hawking

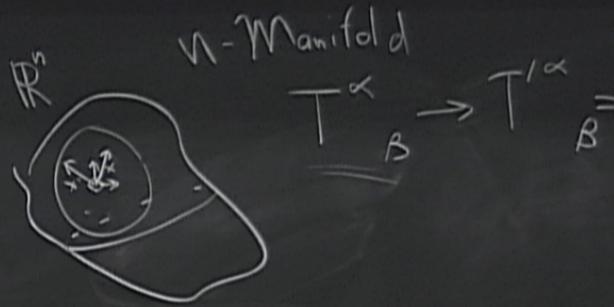
GR, SM

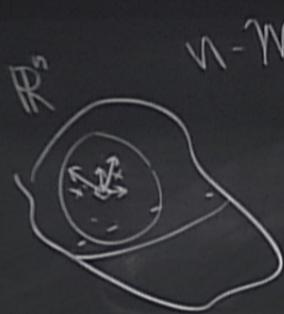
GR → DG

- 1) Manifold, Tensors
- 2) Connection, Covariant Deriv.
- 3)  $g_{\mu\nu}, \dots$

$n$ -Man

- 1) Manifold, Tensors
- 2) Connection, Covariant Deriv.
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N-Manifold

$$T^{\alpha}_{\beta} \rightarrow T'^{\alpha}_{\beta} = \frac{\partial x'^{\alpha}}{\partial x^{\gamma}} \frac{\partial x^{\delta}}{\partial x'^{\beta}} T^{\gamma}_{\delta}$$

$$\partial_M T^{\alpha}_{\beta} \rightarrow \partial'_M T'^{\alpha}_{\beta} \rightarrow \nabla_M T^{\alpha}_{\beta} = \partial_M T^{\alpha}_{\beta} + \Gamma^{\alpha}_{m\delta} T^{\delta}_{\beta} -$$

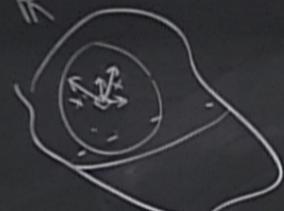
$\partial_M = \frac{\partial}{\partial x^M}$

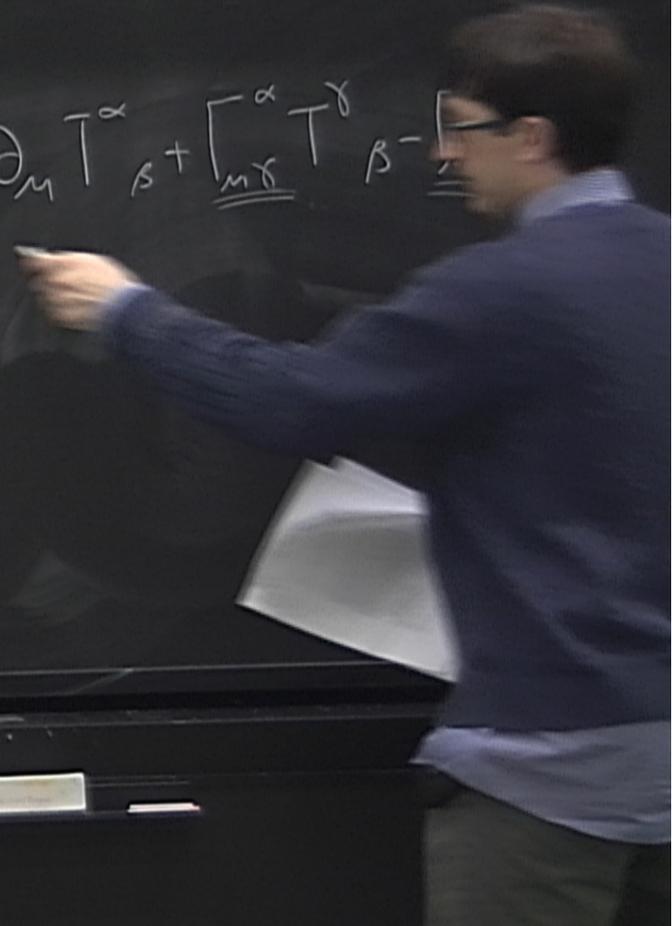
$\mathbb{R}^n$   $N$ -Manifold

$$T^\alpha_B \rightarrow T'^\alpha_B = \frac{\partial x'^\alpha}{\partial x^\gamma} \frac{\partial x^\delta}{\partial x'^\beta} T^\gamma_\delta$$

$$\partial_M T^\alpha_B \rightarrow \partial'_M T'^\alpha_B \rightarrow \nabla_M T^\alpha_B = \partial_M T^\alpha_B + \Gamma^{\alpha}_{\mu\delta} T^\delta_B - \Gamma^{\beta}_{\mu\delta} T^\alpha_B$$

$\partial_M = \frac{\partial}{\partial x^M}$





Manifold

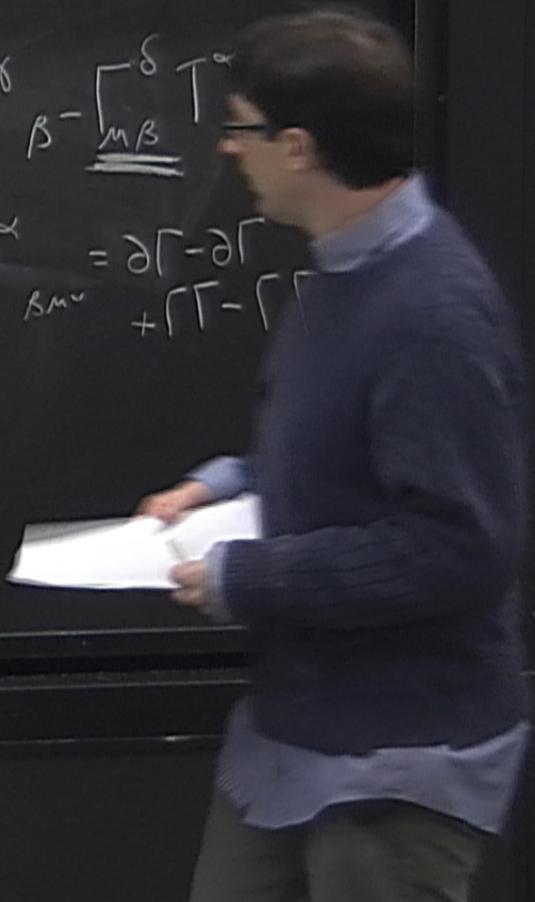
$$T^{\alpha}_{\beta} \rightarrow T'^{\alpha}_{\beta} = \frac{\partial x'^{\alpha}}{\partial x^{\gamma}} \frac{\partial x^{\delta}}{\partial x'^{\beta}} T^{\gamma}_{\delta}$$

$$\partial_M T^{\alpha}_{\beta} \rightarrow \partial'_M T'^{\alpha}_{\beta} \rightarrow \nabla_M T^{\alpha}_{\beta} = \partial_M T^{\alpha}_{\beta} + \Gamma^{\alpha}_{M\gamma} T^{\gamma}_{\beta} - \Gamma^{\delta}_{M\beta} T^{\alpha}_{\delta}$$

$$\partial_M = \frac{\partial}{\partial x^M}$$

$$(\nabla_M \nabla_N - \nabla_N \nabla_M) V^{\alpha} = R^{\alpha}_{\beta\mu\nu} V^{\beta}$$

$$R^{\alpha}_{\beta\mu\nu} = \partial_{\mu} \Gamma^{\alpha}_{\nu\beta} - \partial_{\nu} \Gamma^{\alpha}_{\mu\beta} + \Gamma^{\alpha}_{\mu\gamma} \Gamma^{\gamma}_{\nu\beta} - \Gamma^{\alpha}_{\nu\gamma} \Gamma^{\gamma}_{\mu\beta}$$



DG

1) Manifold, Tensors

2) Connection, Covariant Deriv.

3)  $g_{\mu\nu}, \dots$

$$x^\alpha(\tau), u^\alpha = \frac{dx^\alpha}{d\tau}$$



$n$ -Manifold

$$T^{\alpha}_{\beta} \rightarrow T'^{\alpha}_{\beta} = \frac{\partial x'^{\alpha}}{\partial x^{\gamma}} \frac{\partial x^{\delta}}{\partial x'^{\beta}} T^{\gamma}_{\delta}$$

$$\partial_M T^{\alpha}_{\beta} \rightarrow \partial'_M T'^{\alpha}_{\beta}$$

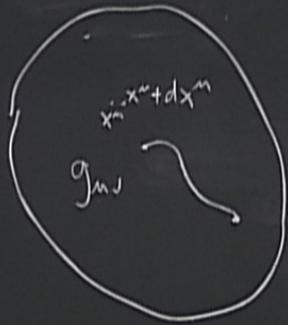
$$u^{\alpha} \nabla_{\alpha} T^{\beta}_{\gamma} = 0$$

$$\partial_M = \frac{\partial}{\partial x^M}$$

$$u^{\alpha} \nabla_{\alpha} u^{\mu} = 0$$

geodesic

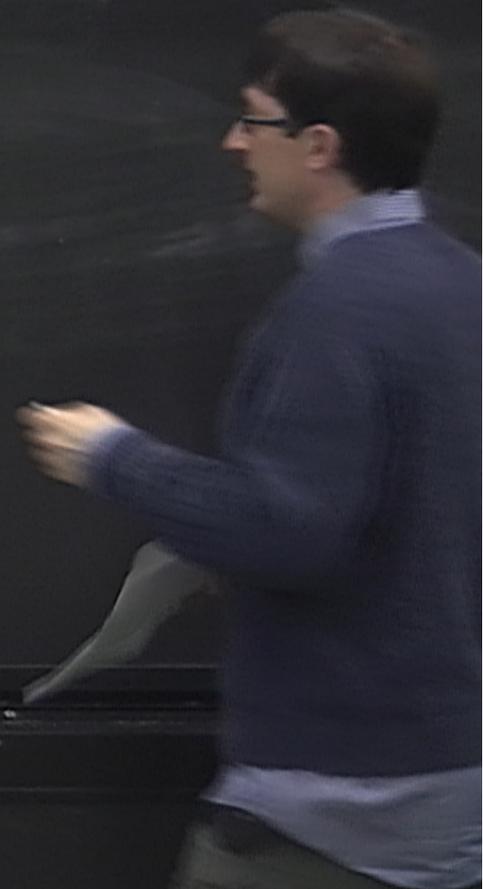
$$\left( \nabla_M \right)$$

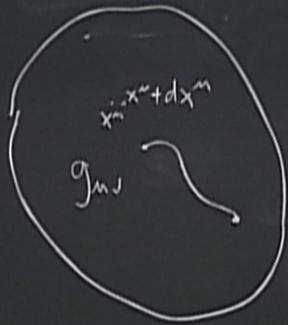


$$ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu$$

$\Gamma_{\mu\beta}^\alpha$ ,  $\underline{g_{\mu\nu}}$   
Levi-Civita

$$\Gamma_{\beta\delta}^\alpha = \Gamma_{\delta\beta}^\alpha$$





$$ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu$$

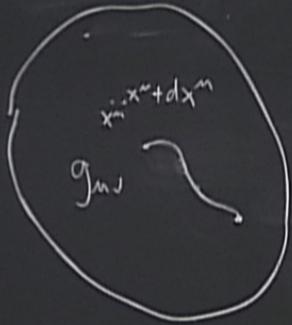
$$\Gamma_{m\beta}^\alpha, \underline{g_{\mu\nu}}$$

Levi-Civita

$$\Gamma_{\beta\delta}^\alpha = \Gamma_{\delta\beta}^\alpha$$

$$\nabla_\alpha g_{\mu\nu} = 0$$

$$\Gamma_{\beta\gamma}^\alpha = \frac{1}{2} g^{\alpha\delta} (\dots)$$



$$ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu$$

$$\Gamma_{\mu\beta}^\alpha, \underline{g_{\mu\nu}}$$

Levi-Civita

$$\Gamma_{\beta\delta}^\alpha = \Gamma_{\delta\beta}^\alpha$$

$$\nabla_\alpha g_{\mu\nu} = 0$$

$$\Gamma_{\beta\gamma}^\alpha = \frac{1}{2} g^{\alpha\delta} (\partial_\beta g_{\delta\gamma} + \partial_\gamma g_{\delta\beta} - \partial_\delta g_{\beta\gamma})$$

$$ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu$$

$$L = \frac{1}{2} g_{\mu\nu}(x) \dot{x}^\mu \dot{x}^\nu$$

$$\dot{x}^\mu = \frac{dx^\mu}{d\tau}$$

$$g_{\mu\nu}(x) dx^\mu dx^\nu$$

$\Gamma_{\mu\beta}^\alpha$ ,  $g_{\mu\nu}$   
Levi-Civita

$$\Gamma_{\beta\delta}^\alpha = \Gamma_{\delta\beta}^\alpha$$

$$\nabla_\alpha g_{\mu\nu} = 0$$

$$\Gamma_{\beta\delta}^\alpha = \frac{1}{2} g^{\alpha\delta} \left( g_{\beta\delta,\gamma} + g_{\gamma\delta,\beta} - g_{\beta\gamma,\delta} \right)$$

$$L(q, \dot{q}) \rightarrow \frac{\partial L}{\partial q} = \left( \frac{\partial L}{\partial q} \right)$$

$$ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu$$

$$L = \frac{1}{2} g_{\mu\nu}(x) \dot{x}^\mu \dot{x}^\nu$$

$$\dot{x}^\mu = \frac{dx^\mu}{d\tau}$$

GR

$$S = \int \sqrt{-g} \mathcal{L} d^4x$$

$$\nabla_m V^\alpha = \left( \partial_m V^\alpha + \Gamma_{m\beta}^\alpha V^\beta \right)$$

$$V^\alpha \rightarrow V'^\alpha = \frac{\partial x'^\alpha}{\partial x^\beta} V^\beta$$

$$\Phi^a \rightarrow \Phi'^a = U(x)^a_b \Phi^b$$
$$D_m \Phi^a = \partial_m \Phi^a + A_{mb}^a \Phi^b$$

GR

$$S = \int \sqrt{-g} \mathcal{L} d^4x$$

$$\nabla_m V^\alpha = (\partial_m V^\alpha + \Gamma_{\mu\beta}^\alpha V^\beta)$$

$$V^\alpha \rightarrow V'^\alpha = \frac{\partial x'^\alpha}{\partial x^\beta} V^\beta$$

$$(\nabla_m \nabla_\nu - \nabla_\nu \nabla_m) V^\alpha = \boxed{R^\alpha_{\beta\mu\nu}} V^\beta$$

$$\mathcal{L}_{GR} = \left( \begin{array}{c} -2\Lambda + g^{\beta\nu} R_{\beta\alpha\nu} \\ \uparrow \\ R \end{array} \right)$$

$16\pi G_N$

$\propto RR$

$$\Phi^a \rightarrow \Phi'^a = U(x)^a_b \Phi^b$$

$$D_\mu \Phi^a = \partial_\mu \Phi^a + A_{\mu b}^a \Phi^b$$

$$D_{[\mu} \Phi^a = i \boxed{F^a_{b\mu\nu}} \Phi^b$$

$F_{\mu\nu}$

	SU(3)	SU(2)	U(1)
$(u_L) = q_L$	3	2	+1/6
$u_R^i$	3	1	+2/3
$d_R^i$	3	1	-1/3
$(\nu_L) = l_L$	1	2	-1/2
$\nu_R^i$	1	1	0
$e_R^i$	1	1	-1
$h$	1	2	+1/2

$$[A_m] = 1$$

$$[\psi] = \frac{3}{2}$$

$$[h] = 1$$

$$\text{Renorm } [\ ] \leq 4$$

