

Title: General Relativity for Cosmology - Lecture 7

Date: Oct 15, 2013 04:00 PM

URL: <http://pirsa.org/13100003>

Abstract:

Recall: Physical motivation for the "Metric Tensor"

□ In Minkowski space, in inertial and cartesian coordinates:

4-dim spacetime distance!

$$[\text{distance}(x, \hat{x})]^2 = -(x^0 - \hat{x}^0)^2 + (x^1 - \hat{x}^1)^2 + (x^2 - \hat{x}^2)^2 + (x^3 - \hat{x}^3)^2$$

↑
indep. of choice
of inertial cds.

$$= \eta_{\mu\nu} (x^\mu - \hat{x}^\mu) (x^\nu - \hat{x}^\nu)$$

with $\eta_{\mu\nu} = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix}$

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□ In Minkowski space, in an arbitrary coordinate system:

$[g_{\mu\nu}(x) (x^\mu - \hat{x}^\mu) (x^\nu - \hat{x}^\nu)]^2$

Recall: Physical motivation for the "Metric Tensor"

□ In Minkowski space, in inertial and cartesian coordinates:

$$\begin{aligned} \left[\overset{\substack{\text{4-dim spa.-time} \\ \text{distance!}}}{\text{distance}}(x, \hat{x}) \right]^2 &= -(x^0 - \hat{x}^0)^2 + (x^1 - \hat{x}^1)^2 + (x^2 - \hat{x}^2)^2 + (x^3 - \hat{x}^3)^2 \\ &= \eta_{\mu\nu} (x^\mu - \hat{x}^\mu) (x^\nu - \hat{x}^\nu) \end{aligned}$$

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$$\text{with } \eta_{\mu\nu} = \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

□ In Minkowski space, in an arbitrary coordinate system:

$$\left[\text{distance}(x, \hat{x}) \right]^2 = g_{\mu\nu}(x) (x^\mu - \hat{x}^\mu) (x^\nu - \hat{x}^\nu) + \mathcal{O}^3$$

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(e.g. polar cds, or accelerated cds) with $g_{\mu\nu}(x) \neq \eta_{\mu\nu}$

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Generalization to curved spacetime

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Generalization to curved space-time historically

□ Generalization to curved space-time, historically:

Allow even such $g_{\mu\nu}(x)$ which in no coordinate system obey:

$$g_{\mu\nu}(x) = \eta_{\mu\nu} \text{ for all } x \in M$$

$\Rightarrow g_{\mu\nu}(x)$ is not simply $\eta_{\mu\nu}$ in noninertial coordinates

\Rightarrow Such $g_{\mu\nu}(x)$ take us beyond special relativity!

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Require $g_{\mu\nu}$ to be such that

for each $x \in M$ there exists a coordinate

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system so that at least at x :

$$g_{\mu\nu}(x) = \eta_{\mu\nu} \left(\begin{array}{l} \text{i.e., locally, special relativity holds} \\ \text{dist}(x, \tilde{x})^2 = \eta_{\mu\nu}(x - \tilde{x})^\mu(x - \tilde{x})^\nu + o^3 \\ \text{to lowest nontrivial order.} \end{array} \right)$$

Recall: Math. definition of the metric tensor:

□ g is covariant tensor of rank $(0,2)$

(because n is in special relativity)

□ Enforce equivalence principle:

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e.g. $\theta^\mu(x) = dx^\mu$

□ Thus, if some cotangent vectors $\theta^\mu(x)$
form a basis at each point x , then

g is of the form:

$$g(x) = g_{\mu\nu}(x) \theta^\mu(x) \otimes \theta^\nu(x)$$

↑ recall: $g_{\mu\nu}(x) = g_{\nu\mu}(x)$ and $g_{\mu\nu}$ is invertible (since nondegenerate)

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\rightsquigarrow Modern view of the equivalence principle:

Recall: We asked that for each point $p \in M$ there is a coordinate

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3 / 23

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obeys: $g_{\mu\nu}(p) = \eta_{\mu\nu}$ (in general only at p)

Modern formulation of the equivalence principle:

Independently of any choice of coordinate system:

There are choices of dual

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4 / 23

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4 / 23

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Independently of any choice of coordinate system:

There are choices of dual bases $\{\theta^{\mu}(x)\}, \{\partial_{\nu}(x)\}$ of $T_x^*(M), T_x(M)$

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so that: $g_{\mu\nu}(x) = g(e_\mu(x), e_\nu(x)) = \eta_{\mu\nu} \quad \forall x \in M$

Now, knowing distances through $g_{\mu\nu}$, what else follows?

- Distances yield volumes, namely $g_{\mu\nu}(x)$ induces an $\Omega(x)$.
- g, g^{-1} yield duality of covariance and contravariance.

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□ $*$ yields $(,)$ making the Λ_p Hilbert spaces.
for Riemannian manifolds

□ g yields co-derivative $\delta: \Lambda_p \rightarrow \Lambda_{p-1}$

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5 / 23

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5 / 23

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Namely:

□ Assume, as always, that M is oriented.

□ Consider a positive chart.

(i.e. has positive $\det(\text{Jacobian})$ with given atlas)

Then:

$$:= |\det(g_{i,j}(x))|$$

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□ $\Omega := \sqrt{|g|} dx^1 \wedge dx^2 \wedge \dots \wedge dx^n$ is a well-defined volume form.

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Proof: □ Nonzero for all $p \in M$?

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Then:

$$\tilde{g}_{ij}(\tilde{x}(x)) = g_{rs}(x) \frac{\partial x^r}{\partial \tilde{x}^i} \frac{\partial x^s}{\partial \tilde{x}^j} \quad \text{because covariant}$$

i.e., as matrices:

$$\tilde{g} = \left(\frac{\partial x}{\partial \tilde{x}} \right)^T g \left(\frac{\partial x}{\partial \tilde{x}} \right) \quad \text{now take determinant:}$$

$$\Rightarrow |\tilde{g}| = \left| \frac{\partial x}{\partial \tilde{x}} \right|^2 |g| \quad \text{i.e. } |\tilde{g}|^{1/2} = \left| \frac{\partial x}{\partial \tilde{x}} \right| |g|^{1/2}$$

$$\text{Also: } d\tilde{x}^1 \wedge \dots \wedge d\tilde{x}^n = \det \left(\frac{\partial \tilde{x}}{\partial x} \right) dx^1 \wedge \dots \wedge dx^n$$

7

7/23

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Notation: (Ω is an n -form. What are its coefficients, as a covariant $(0, m)$ tensor?)

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$$\Rightarrow |\tilde{g}|^{1/2} d\tilde{x}^1 \wedge \dots \wedge d\tilde{x}^m = \underbrace{\left| \frac{\partial \tilde{x}}{\partial x} \right| \left| \frac{\partial x}{\partial \tilde{x}} \right|}_1 |g|^{1/2} dx^1 \wedge \dots \wedge dx^m \quad \checkmark$$

Notation: (Ω is an n -form. What are its coefficients, as a covariant $(0, m)$ tensor?)

□ Define:

$g_{ij}(\tilde{x}(x)) = g_{rs}(x) \frac{\partial x^r}{\partial \tilde{x}^i} \frac{\partial x^s}{\partial \tilde{x}^j}$ because covariant

i.e., as matrices:

$\tilde{g} = \left(\frac{\partial x}{\partial \tilde{x}}\right)^T g \left(\frac{\partial x}{\partial \tilde{x}}\right)$ now take determinant:

$\Rightarrow |\tilde{g}| = \left|\frac{\partial x}{\partial \tilde{x}}\right|^2 |g|$ i.e. $|\tilde{g}|^{1/2} = \left|\frac{\partial x}{\partial \tilde{x}}\right| |g|^{1/2}$

Also: $d\tilde{x}^1 \wedge \dots \wedge d\tilde{x}^n = \det\left(\frac{\partial \tilde{x}}{\partial x}\right) dx^1 \wedge \dots \wedge dx^n$

$\Rightarrow |\tilde{g}|^{1/2} d\tilde{x}^1 \wedge \dots \wedge d\tilde{x}^n = \frac{|\tilde{g}|^{1/2}}{\left|\frac{\partial \tilde{x}}{\partial x}\right|} \frac{1}{|g|^{1/2}} |g|^{1/2} dx^1 \wedge \dots \wedge dx^n$ ✓

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$$\epsilon_{i_1, \dots, i_n} := \begin{cases} +1 & \text{if } (i_1, \dots, i_n) \text{ is even permutation of } (1, 2, \dots, n) \\ -1 & \text{if } (i_1, \dots, i_n) \text{ is odd permutation of } (1, 2, \dots, n) \\ 0 & \text{else} \end{cases}$$

unlike in SRT, ϵ_{\dots} is not canonical, because Ω is: →

□ Then, Ω also reads:

$$\begin{aligned} \Omega &= \sqrt{|g|} dx^1 \wedge \dots \wedge dx^n && (n\text{-form}) \\ &= \sqrt{|g|} \underbrace{\epsilon_{i_1, \dots, i_n}}_{=: \Omega_{i_1, \dots, i_n}} dx^{i_1} \otimes dx^{i_2} \otimes \dots \otimes dx^{i_n} \\ &= \Omega_{i_1, \dots, i_n} dx^{i_1} \otimes \dots \otimes dx^{i_n} && (\text{covariant tensor}) \end{aligned}$$

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if ϵ :

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8 / 23

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Q: Other use of g ?

A. A metric tensor g is used to define the inner product of two vectors u and v in a vector space V as $\langle u, v \rangle = g(u, v)$. This inner product is used to define the length of a vector u as $|u| = \sqrt{g(u, u)}$ and the angle between two vectors u and v as $\cos \theta = \frac{g(u, v)}{|u| |v|}$.

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Q: Other use of g

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Why? a) \square should be non-directional 2nd derivative, but $d^2 = 0$.

b.) need e.g. $\square: \Lambda^p \rightarrow \Lambda^p$ for Klein Gordon, i.e., need degree of forms conserved by \square .

Strategy:

A) Use g for a covariant \leftrightarrow contravariant tensors relation

B) Define a map "Hodge" $*$: $\Lambda_r \rightarrow \Lambda_{n-r}$

C) Define the "Cocodervative": $\delta: \Lambda_r \rightarrow \Lambda_{r-1}$

D) Define "Laplacian/d'Alembertian": $\square = d\delta + \delta d$

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A) Covariant \leftrightarrow contravariant tensors equivalence through g :

\square $g(x)$ can be used as a map: by evaluation of one tensor factor:

$$g(x): T_x(M)' \rightarrow T_x(M),$$

$$g(x): \xi^i(x) e_i \otimes \omega_j \rightarrow g_{\mu\nu}(x) \theta^\mu(x) \otimes \theta^\nu(x) (\xi^i e_i \otimes \omega_j)$$

$$= \underbrace{g_{\mu\nu}(x)}_{\in \mathbb{F}_x(M)} \underbrace{\xi^i(x)}_{\in T_x(M)} \underbrace{\theta^\nu(x)}_{\in T_x(M)} \in T_x(M),$$

$$\theta^\nu(e_i) = \delta_i^\nu$$

\Rightarrow For the coefficient

functions we have: $g: \xi^i(x) \rightarrow \omega_j(x) = g_{ij}(x) \xi^i(x)$ (relative to bases θ^i, e_j)

\square Conversely, g^{-1} acts as:

$$g^{-1}(x): T_x(M) \rightarrow T_x(M)'$$

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$$g'(x): T_x(M)_1 \rightarrow T_x(M)'$$

$$g'(x): \omega_p(x) \rightarrow \xi(x) = g'^{\nu\sigma}(x) \omega_\sigma(x)$$

In this way, g, g' can lower or raise any

tensor index, e.g.: $g: t^{ij}_k \rightarrow t_i^j{}_k = g_{is} t^{is}{}_k$

and: $g': \tau^{ij}_k \rightarrow \tau^{i'j'}_k = g'^{i'k} \tau^{ij}_k$

B) The Hodge $*$ map: $\Lambda_p \rightarrow \Lambda_{n-p}$

Recall:
 $\dim(\Lambda_p) = \binom{n}{p} = \dots$
 $= \dim(\dots)$

Idea: \square each $\nu \in \Lambda_p$ is a covariant tensor

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Concretely:

Consider any $v := \frac{1}{p!} v_{i_1, \dots, i_p} dx^{i_1} \wedge \dots \wedge dx^{i_p} \in \Lambda_p$

anything totally antisymmetric
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$$= v_{i_1 \dots i_p} dx^{i_1} \otimes \dots \otimes dx^{i_p}$$

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$n-p$ factors

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arrange maps out here \rightarrow \rightarrow components as a covariant vector

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$$\Omega(\tilde{v}) = \underbrace{\Omega_{i_1, \dots, i_p}}_{!!} \tilde{v}^{i_1, \dots, i_p} \underbrace{dx^{i_{p+1}} \otimes \dots \otimes dx^{i_n}}_{n-p \text{ factors}} \in \Lambda_{n-p}$$

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Proposition:

Assume $v \in \Lambda_p$. Then

$$d(n-p)+s$$

← E.g. $s=1$ for space-time

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$$**v = (-1)^{p(n-p)+s} v$$

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What is s ? The "signature" of g is $\text{sgn}(g) = (r, s)$, where in diagonal form: $g = \begin{pmatrix} 1 & & & \\ & \dots & & \\ & & 1 & \\ & & & -1 & \\ & & & & \dots \end{pmatrix}$

Use $*$ to turn $\Lambda(M)$ into an "Inner Product Space":

What is s ? The "signature" of g is $\text{sgn}(g) = (r, s)$, where in diagonal form: $g = \begin{pmatrix} 1 & & & \\ & \ddots & & \\ & & 1 & \\ & & & -1 & \\ & & & & \ddots \\ & & & & & -1 & \\ & & & & & & \ddots \end{pmatrix}$

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Definition:

The Hodge $*$ provides a "scalar" (or also called "inner") product for $\Lambda(M)$:

Exercises:

- 1.) Write (α, β) in coordinates
- 2.) Show that $(,)$ is always positive definite on Λ_0 , i.e., $(\alpha, \alpha) > 0 \forall \alpha \in \Lambda_0, \alpha \neq 0$.

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Recall: For any operator $A: D_A \subset \mathcal{X} \rightarrow \mathcal{X}$ (with D_A dense, i.e., $\overline{D_A} = \mathcal{X}$), its adjoint A^* is defined to have the domain

$$D_{A^*} := \left\{ v \in \mathcal{X} \mid \exists w \in \mathcal{X} \forall z \in D_A: \langle v, Az \rangle = \langle w, z \rangle \right\}$$

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$$(\delta \alpha, \beta) := -(\alpha, d\beta) \quad \forall \alpha \in D_\delta, \beta \in D_d$$

c) The Codifferential δ explicitly

Clearly: $\delta: \Lambda^k(M) \rightarrow \Lambda^{k-1}(M)$

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Properties: $\square \delta^2 = 0$

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"Field strength": $F_{\mu\nu}(x) := \begin{pmatrix} 0 & -E_1 & -E_2 & -E_3 \\ E_1 & 0 & B_3 & -B_2 \\ E_2 & -B_3 & 0 & B_1 \\ E_3 & B_2 & -B_1 & 0 \end{pmatrix}$, $F = F_{\mu\nu} dx^\mu \wedge dx^\nu$

↖ electric field
 ↘ magn. field
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□ F is assumed to be an exact 2-form, i.e.,

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$$dF = d^2A = 0 \quad !$$

→ One calls them "structure equations".

Remark:

The gauge principle of electrodynamics is the observation that, for any $w \in \Lambda_0$:

$$A \text{ and } \tilde{A} := A + dw$$

describe the same physics.

The Aharonov-Bohm effect and topological phases in general, can make A itself visible when Poincaré lemma doesn't apply.

They do because the (classically,) observable fields are only the E and B fields in \mathbb{F} and since $d^2 = 0$:

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D The Laplacian/d'Alembertian, Δ , \square :

□ Definition of the Laplacian:

$$\Delta := \delta d + d\delta$$

(Some authors define Δ as the negative of this)

□ Clear: $\Delta : \Lambda^p(M) \rightarrow \Lambda^p(M)$

□ If signature $s=1$: Then also called d'Alembertian and denoted $\square := d\delta + \delta d$.

□ Action on, e.g., $f \in \Lambda_0(M)$ in a chart: Exercise: verify

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$$\Delta f = \frac{1}{\sqrt{|g|}} \left(\sqrt{|g|} g^{\mu\nu} f_{,\mu} \right)_{,\nu}$$

$$= \left(-\frac{\partial^2}{\partial x^0{}^2} + \frac{\partial^2}{\partial x^1{}^2} + \frac{\partial^2}{\partial x^2{}^2} + \frac{\partial^2}{\partial x^3{}^2} \right) f$$

if $g = \eta$

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Properties of the d'Alembertian, \square , in the Hilbert space $\Lambda(M)$: , if Λ is a Hilbert space

* Defined:

$$\square : \Lambda_r(M) \rightarrow \Lambda_r(M)$$

$$\square : \psi \rightarrow (\delta d + d\delta)\psi$$

* In the Hilbert space $\Lambda(M)$:

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$$\square = \delta d + d\delta \text{ obeys } (d, \square\beta) = (\square d, \beta)$$

* \square is self-adjoint, $\square = \square^+$, for suitable boundary conditions, or if $\partial M = \emptyset$ and assuming $(,)$ is positive definite.

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on the manifold, space (α, β) .

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* Exercises: \square Verify $\square = \square^+$ formally, using only $\delta = -d^+$.
 \square Verify that $\square^* = *\square$, $\square d = d\square$, $\square \delta = \delta \square$.

Consequences of the self-adjointness of \square : - if \mathcal{H} is a Hilbert space

A) The operators Δ and \square can be diagonalized, with real spectrum.

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Consequences of the self-adjointness of \square : $- \square / \Gamma$ is a Hilbert space

A) The operators Δ and \square can be diagonalized, with real spectrum.

B) For Riemannian manifolds, $\text{spec}(\Delta) \subset [0, \infty)$.

C) For Riem. manifolds of finite volume: $\text{spec}(\Delta)$ is discrete.

Still the finite volume Riemannian case.
D) Then, $\text{spec}(\Delta)$ is carrying a lot of information about (M, g) !

Remark: There exists a related mathematical discipline, called "Spectral Geometry", combining differential
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of Riemannian manifolds, (M, g)

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Application: Klein-Gordon "action":

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$$\underbrace{g^{\mu\nu}}_{\in T^2} e_\mu \otimes e_\nu \quad \underbrace{(\phi_{,\alpha} \theta^\alpha)}_{\in T_1} \otimes \underbrace{(\phi_{,\beta} \theta^\beta)}_{\in T_1}$$

$$S[\phi] := \frac{1}{2} \int_M \underbrace{g^{\mu\nu}}_{\uparrow \Lambda_0} \underbrace{\phi_{,\mu}}_{\uparrow \Lambda_0} \underbrace{\phi_{,\nu}}_{\uparrow \Lambda_0} \underbrace{\Omega}_{\uparrow \Lambda_n}$$

↑
Klein Gordon
field $\phi \in F(M)$

(Recall special relativity:
 $S[\phi] = \int_{\mathbb{R}^4} \eta^{\mu\nu} \phi_{,\mu} \phi_{,\nu} d^4x$)

$$= \frac{1}{2} \int_M g^{\mu\nu}(x) \left(\frac{\partial}{\partial x^\mu} \phi \right) \left(\frac{\partial}{\partial x^\nu} \phi \right) \underbrace{\sqrt{|g(x)|}}_{\uparrow \Omega} d^m x$$

↑ next: integrate by parts!

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Obtain the Klein Gordon wave equation:

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$$S[\phi] := \frac{1}{2} \int_M \underbrace{g^{\mu\nu}}_{\uparrow \Lambda_\mu} \underbrace{\phi_{,\mu}}_{\uparrow \Lambda_\nu} \underbrace{\phi_{,\nu}}_{\uparrow \Lambda_\nu} \underbrace{\Omega}_{\uparrow \Lambda_\mu}$$

↑
Klein Gordon
field $\phi \in F(M)$

(Recall special relativity:)

$$S[\phi] = \int_{\mathbb{R}^4} \eta^{\mu\nu} \phi_{,\mu} \phi_{,\nu} d^4x$$

$$= \frac{1}{2} \int_M g^{\mu\nu}(x) \left(\frac{\partial}{\partial x^\mu} \phi \right) \left(\frac{\partial}{\partial x^\nu} \phi \right) \underbrace{\sqrt{|g(x)|}}_{\Omega} d^n x$$

↑ next: integrate by parts!

$$= \frac{1}{2} \int_M -\phi \underbrace{\frac{\partial}{\partial x^\nu} \left(\sqrt{|g|} g^{\mu\nu} \frac{\partial}{\partial x^\mu} \phi \right)}_{=\square\phi} \frac{1}{\sqrt{|g|}} \underbrace{\sqrt{|g|}}_{=\Omega} d^n x$$

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□ Recall: Euler Lagrange equation $\frac{\partial \mathcal{L}}{\partial \phi} = \frac{\partial}{\partial x^\mu} \frac{\partial \mathcal{L}}{\partial \phi_{,\mu}}$

□ Here: $\mathcal{L} = -\frac{1}{2} \phi \square \phi$ (the 0-form that we are integrating: $S = \int_M \mathcal{L} \Omega$)

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(with "mass": $\mathcal{L} = -\frac{1}{2} \phi (\square + m^2) \phi$)

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□ Higgs field (Gives all particles their mass. Found at LHC. Nobel to Higgs, Englert (Brout) in 2013)

□ Inflaton field (crucial ingredient in modern cosmology → see later)