Title: The impact of the Atmosphere on space-bound Vehicles

Date: Aug 16, 2013 10:10 AM

URL: http://pirsa.org/13080046

Abstract: <span>When rockets and space-bound vehicles are fired into space, they have to deal with turbulence, winds and waves in the Earth's atmosphere. Different height regimes cause different problems. In the lower regions (0 to 20 km altitude), winds and turbulence are major issues, whereas above 50 km altitude, large atmospheric waves have serious importance on safe passage. Similar issues arise on re-entry. In this talk, I will discuss the ways in which these atmospheric phenomena manifest themselves, and how they are best mitigated. One example of particular significance will be the Space Shuttle Columbia disaster, for which I consulted with NASA in respect to possible atmospheric influences. I will also describe some of the instrumentation used to measure these phenomena.

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# The impact of the Atmosphere on space-bound Vehicles

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CAM, Waterloo, 2013

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#### Atmospheric Physics – how does in fit?

Nobel Prize Winners in Physics and Chemistry:

#### Physics:

Rayleigh (1904), Appleton (1947), Braun and Marconi (1909), Van der Waals (1910), Compton and Wilson (1927), Raman (optics – 1930),

Chemistry: Arrhenius (1903- Greenhouse), Langmuir (1932), Debye and Huckel (1936), Herzberg (1971), Crutzen, Molina and Rowland (1995).

Peace Prize: IPCC and Gore, 2007.

Other notables: Lorenz (Chaos theory), Richardson (Turbulence), Fourier (Greenhouse), Rossby, Bjerknes, Prandtl, von Karman ...

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Studies in Atmospheric Physics requires background in.. Electromagnetism, Thermodynamics, Optics, Fluids, Nonlinear processes, gas kinetics, radar, signal-processing, Quantum Mechanics, scattering theory, computer modeling, chemistry...

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## $A \, typical \, weather \, forecast \, now a days \dots$



Forecast AQHI						
Mon 3 Jun	Tue 4 Jun	Wed 5 Jun	Thu 6 Jun	Fri 7 Jun	Sat 8 Jun	Sun 9 Jun
	<b>₩</b>	**		45	*	*
	20°C	19°C	19°C	18°C	30% 20°C	22°C
6	°C 8	PC 10			°C 12	°C

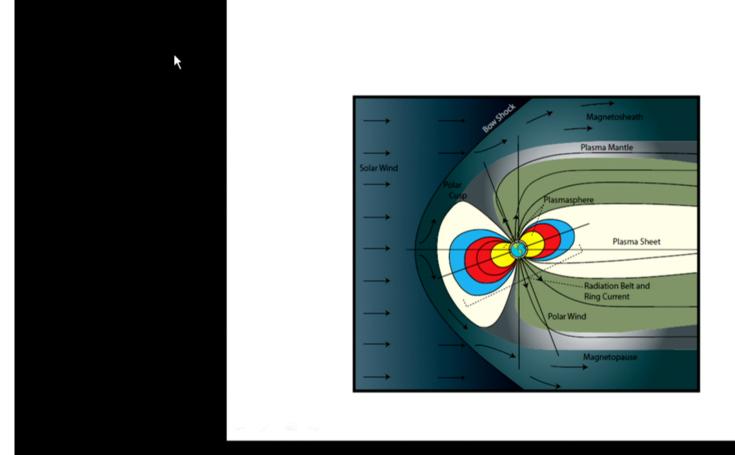
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An aviation weather forecast is not too much different, but might include wind maps, turbulence map. But what about in the future – what will a "weather forecast" look like in the days of space travel?

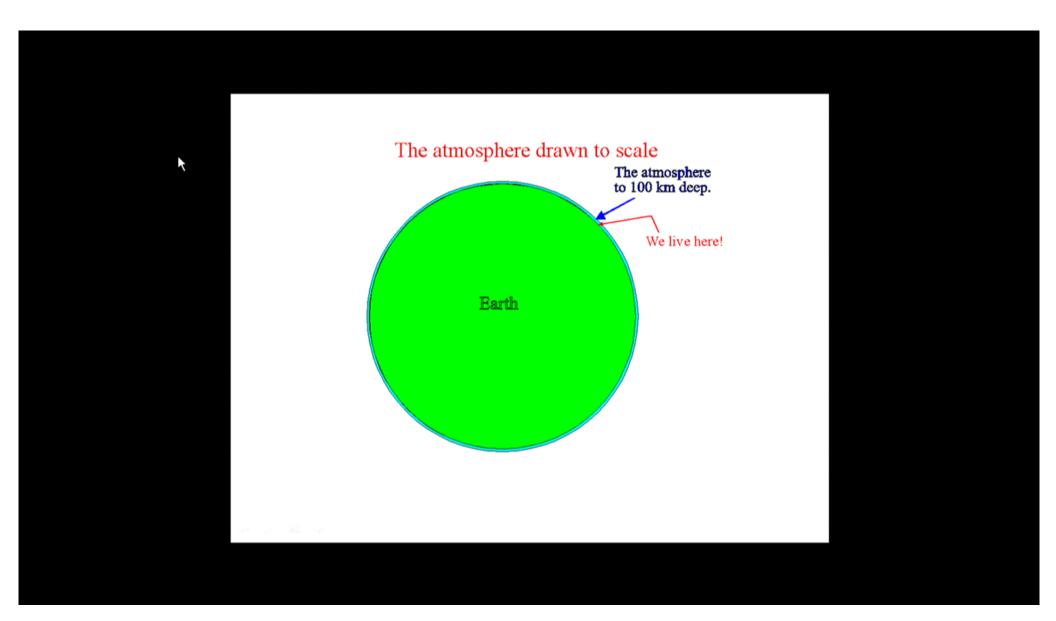
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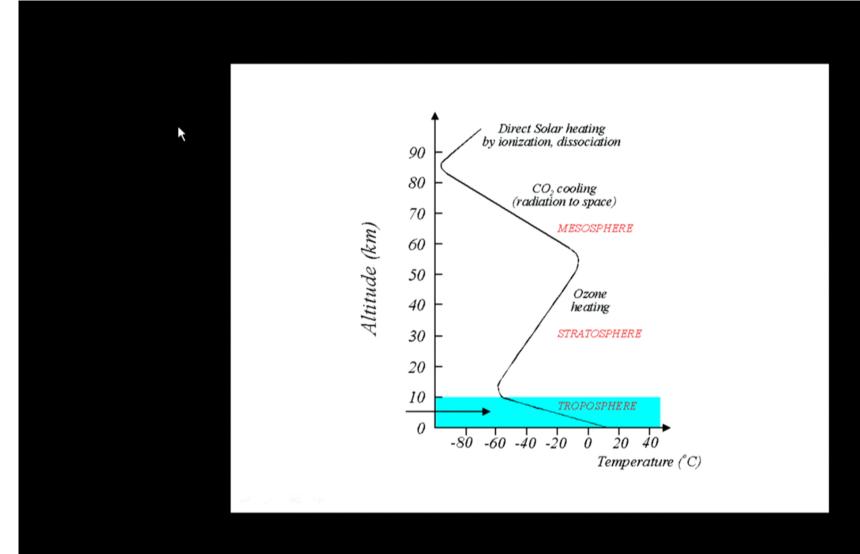
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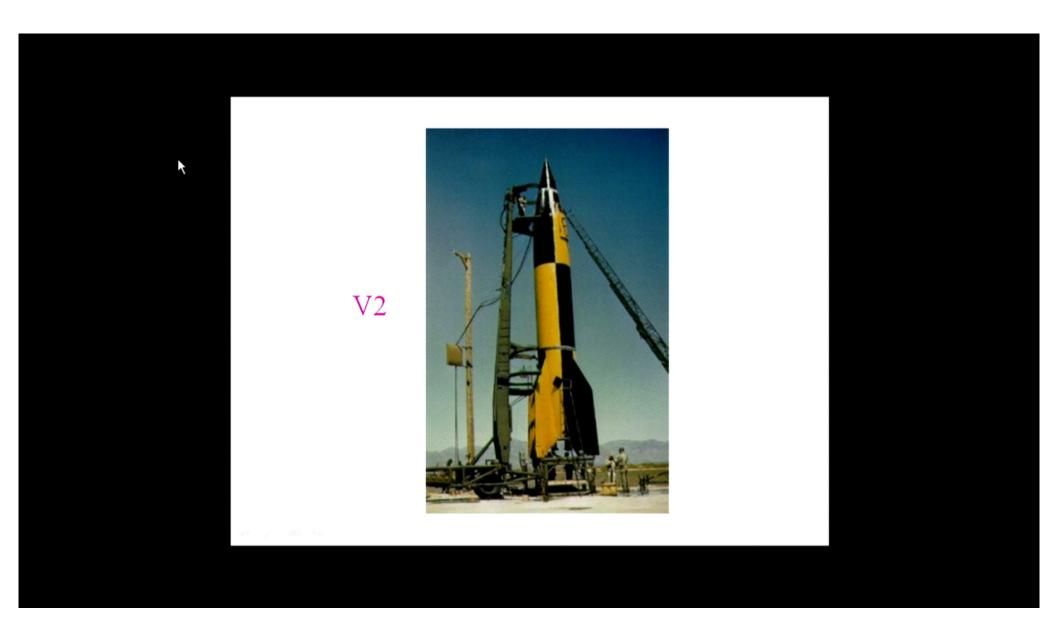


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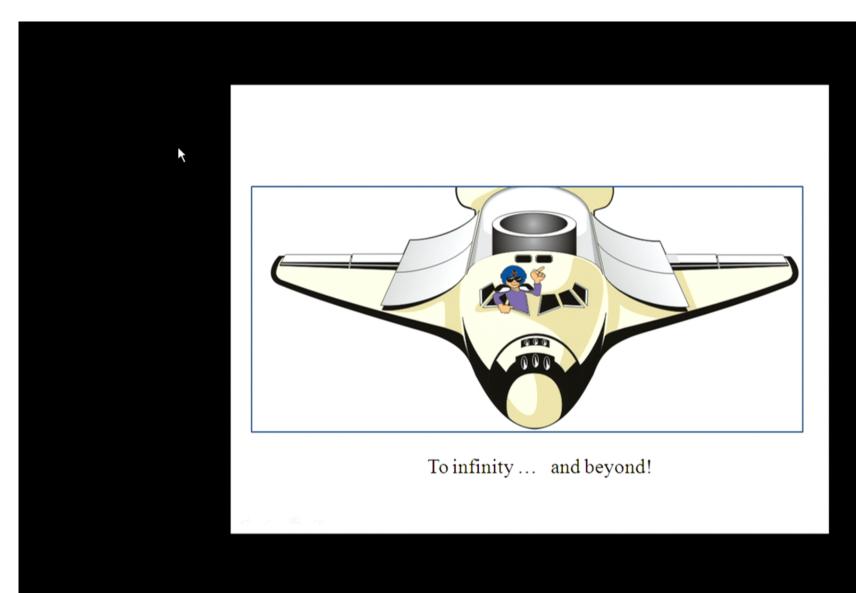


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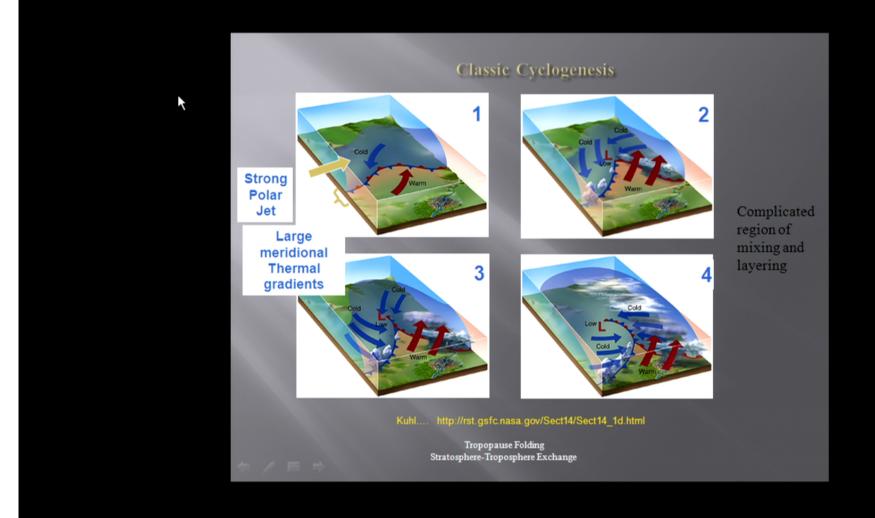
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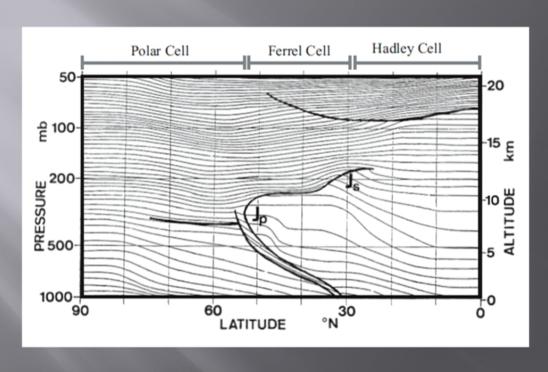
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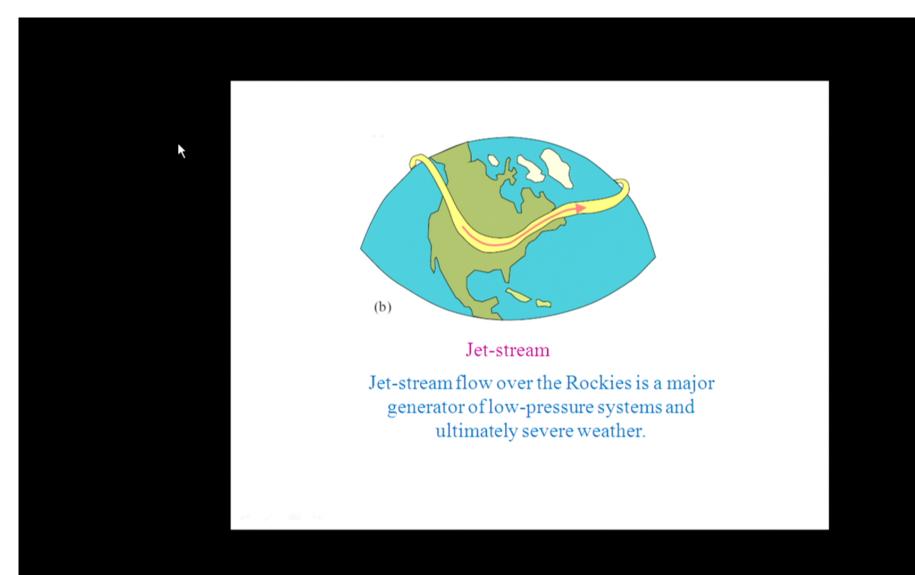


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 $Complicated \ latitudinal \ structure$ 

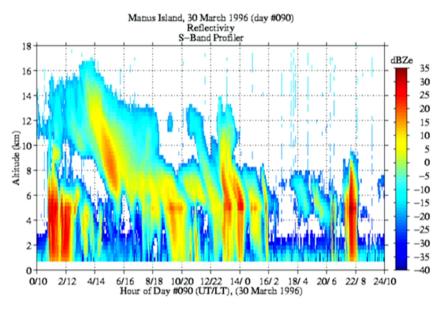
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University of Colorado, Boulder



Giant plumes of hot air rising into the stratosphere..

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## What does this picture tell us about the upper atmosphere?



Credit: Mel Wright



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## So... cold temperatures ... Triple point, sublimation curve..

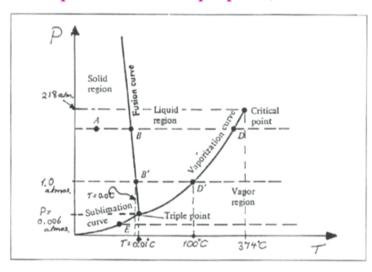


Figure 3.6: More realisitic phase diagram for water. Remember that for most substances the fusion curve points up and to the right - it only points to the left as water expands on freezing.

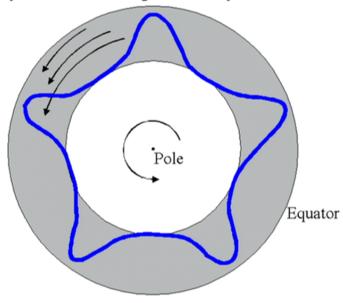
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So... cold temperatures, no water..just ice and vapour Stable temperature profile, (heating due to Ozone chemistry) – pretty boring, right? But something else enters the picture... ... WAVES....

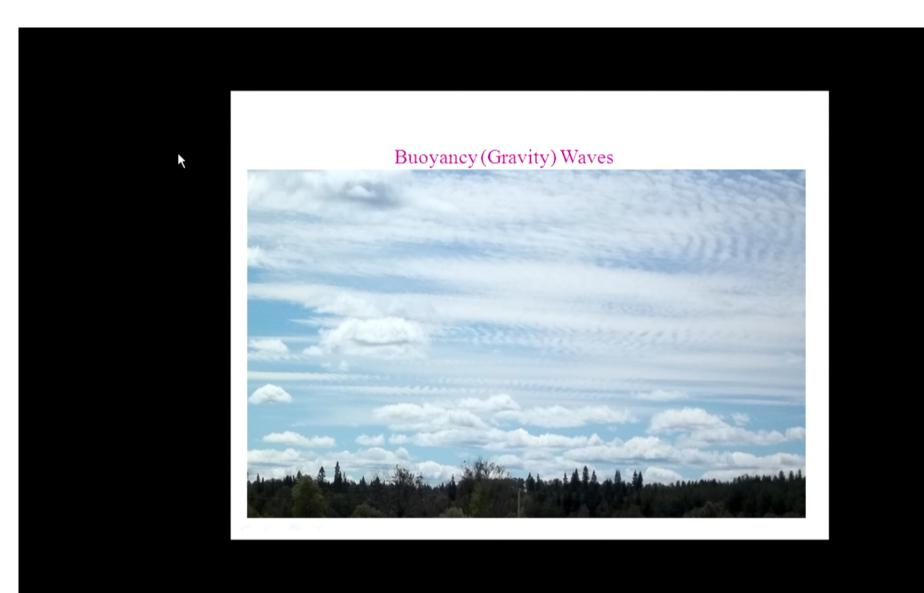
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### PLANETARY WAVES

Development of large-scale waves by differential heating and velocity shear.



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Waves – generated either in-situ or in the troposphere, and propagate upward.

#### Navier-Stokes

$$\frac{D\vec{u}}{Dt} = -\frac{1}{\rho} \vec{\nabla} P - 2\vec{\Omega} \times \vec{u} + \vec{g} + \nu \nabla^2 \vec{u}, \qquad (2.1)$$

$$\frac{D\rho}{Dt} = \frac{1}{c_s^2} \frac{DP}{Dt}$$
 (2.2)

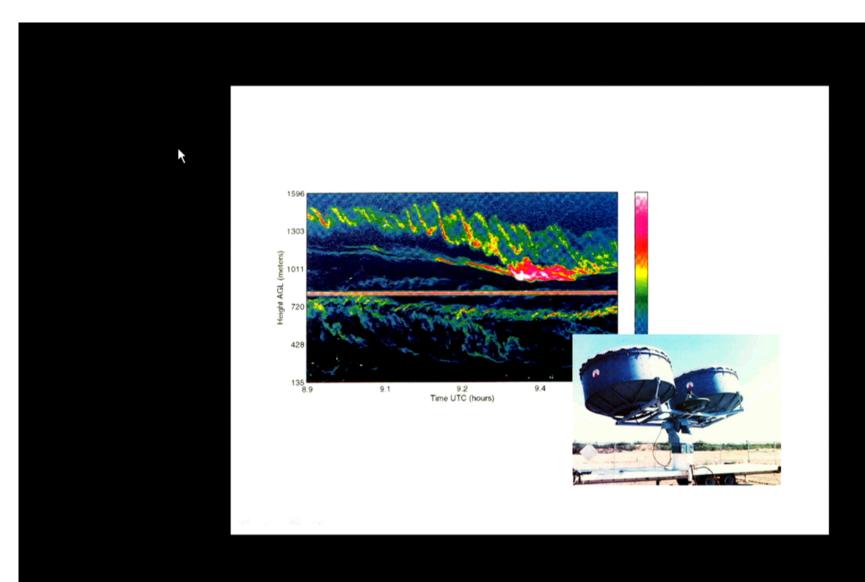
$$\frac{D\rho}{Dt} + \rho \vec{\nabla} \cdot \vec{u} = 0 \tag{2.3}$$

$$\frac{D\Theta}{Dt} = \frac{\kappa}{\rho} \nabla^2 \Theta \eqno(2.4)$$





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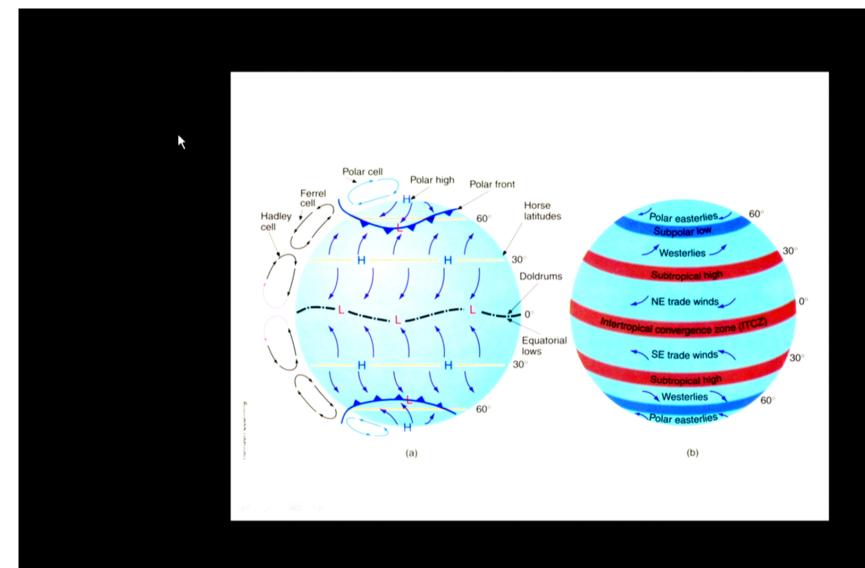


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In the stratosphere, two main impacts for aircraft: 1. Highly developed non-linearities in planetary-waves lead to development of subsequent radiated waves and turbulence Breakdown of gravity waves leads to development of clear-air-turbulence, which can be quite severe.

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Dynamically, even more important, since they impact the general circulation of the atmosphere. First, remind ourselves of the causes of the tropospheric circulation

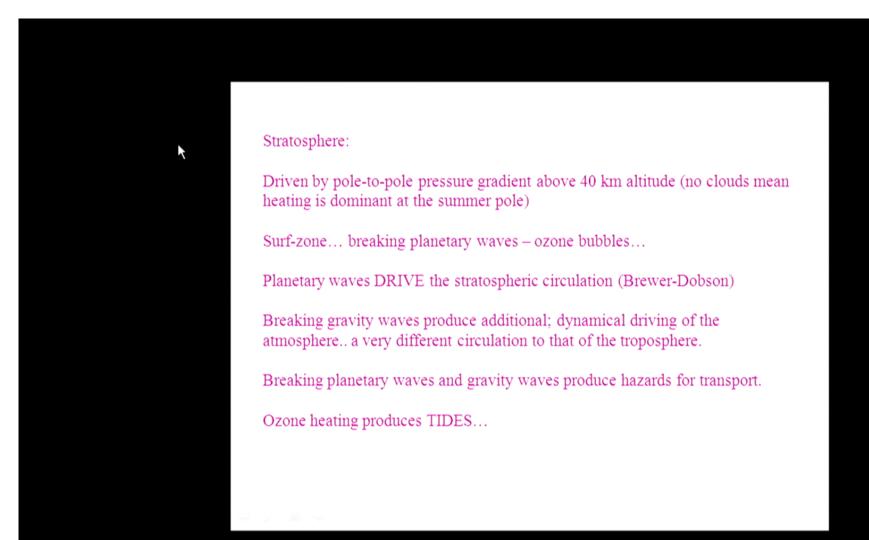


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- These effects, combined with other factors like ocean currents, land-sea differences, local orography, and water-related effects (latent heat transport, precipitation) are responsible for much of our "weather".
- Largely radiatively driven
- But what about the "upper" atmosphere?

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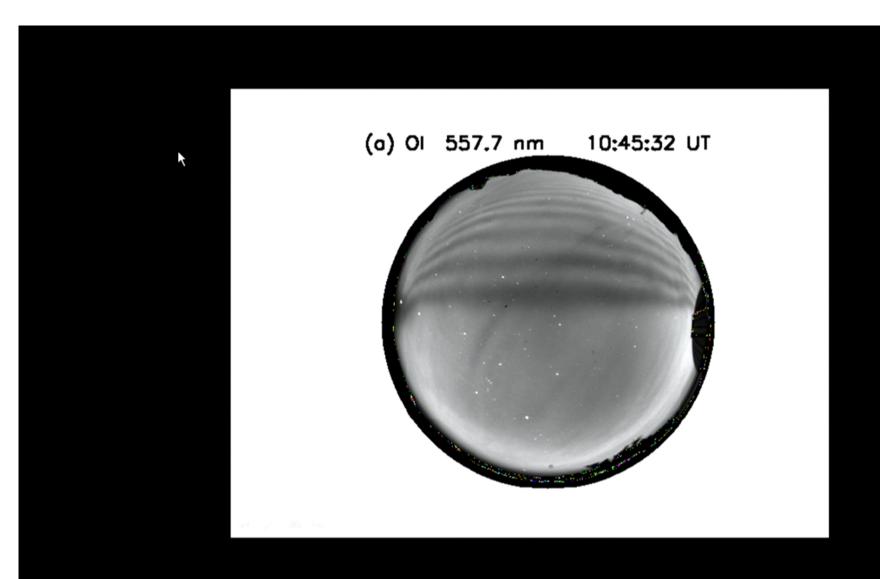
## Mesospheric Circulation

- Little to no water vapour
- Heating distribution differs
- Ozone heating, particle heating, IR radiation
- Dynamical and wave processes much more important
- Waves \*\*\*
- · No direct frictional drag due to ground
- coupling with troposphere and ionosphere

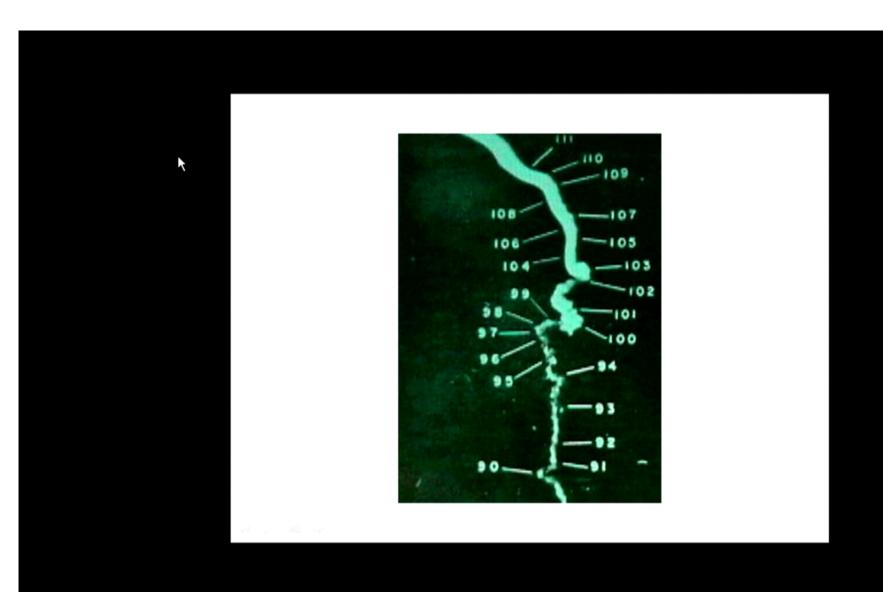
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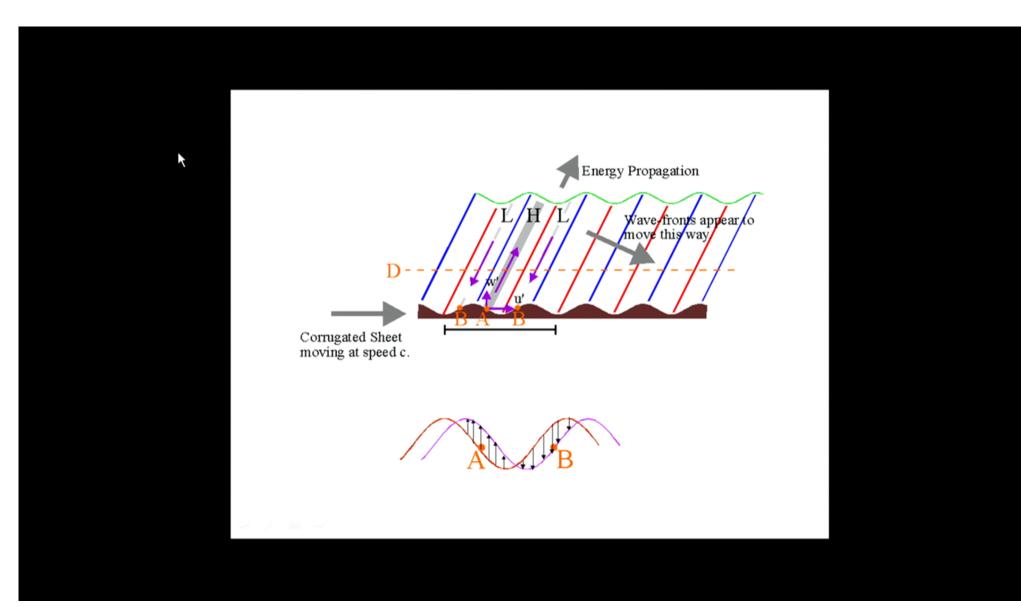
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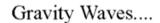
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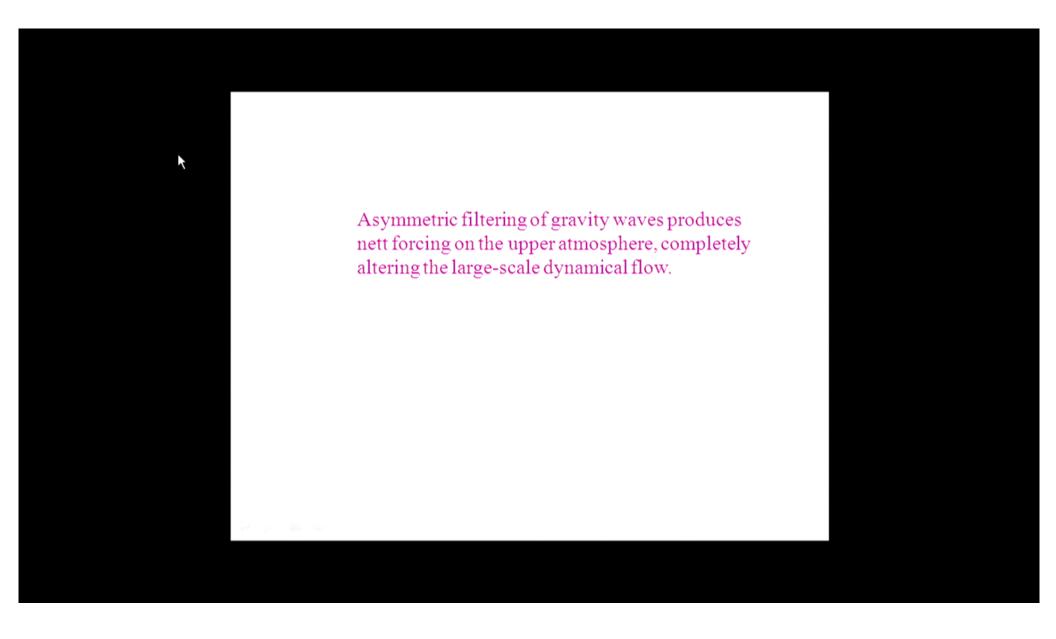


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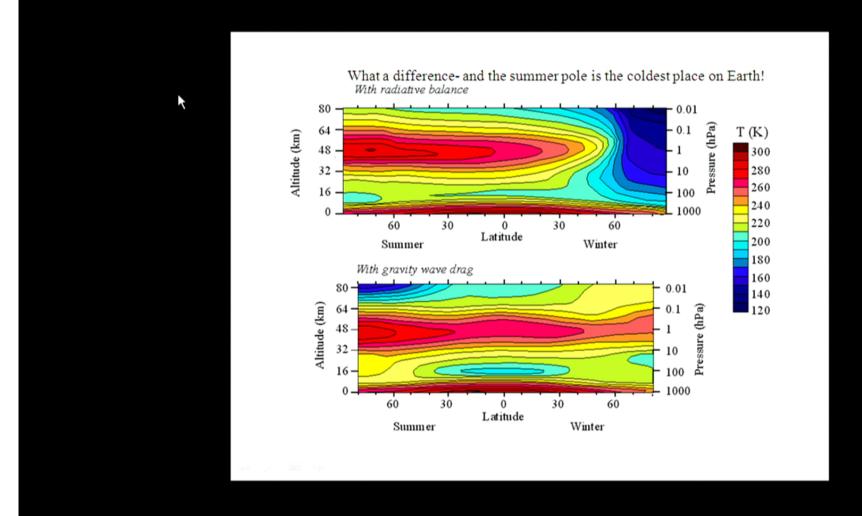


- 1. Grow exponentially with height in order to conserve kinetic energy per unit volume
  - 2. Can achieve velocity amplitudes of the order of tens of metres per second.
  - 3. Dissipate when they encounter a critical level (phase speed equals mean wind)
  - 4. Convectively break down when their amplitude u' exceeds |u-c|.
  - 5. Can also break down by other processes (slant-wise instability, K-H instability etc.)

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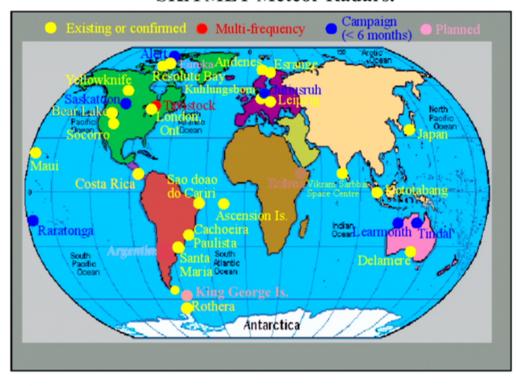


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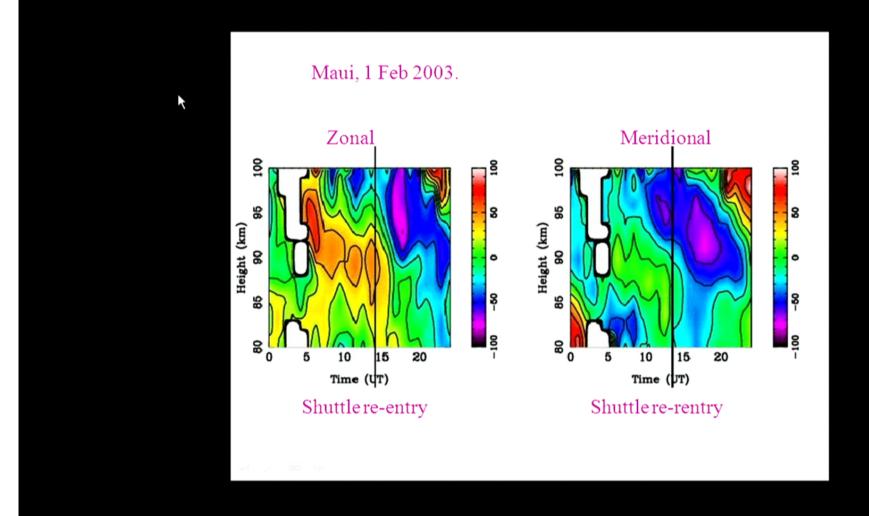


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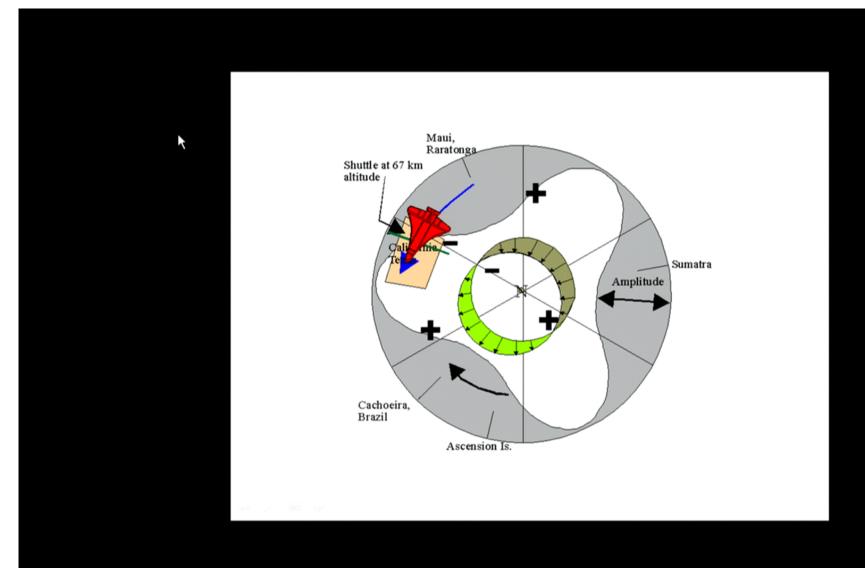
## **SKiYMET Meteor Radars.**



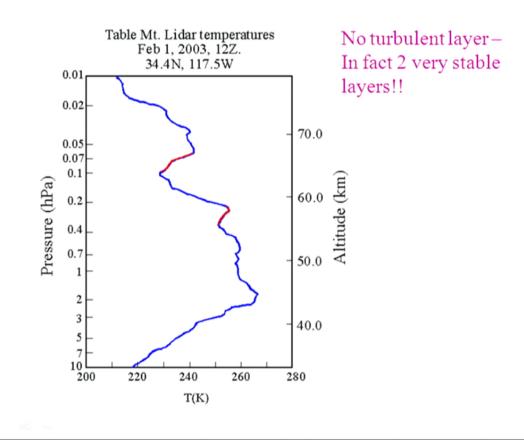
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We need to recognize that

- Total density  $\rho = \langle \rho \rangle + \rho'$
- Total velocity v = <v> + v'

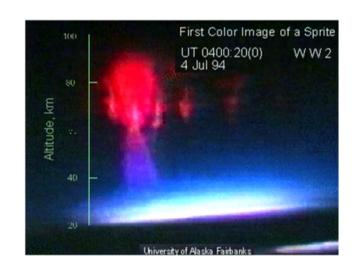
The degree of bufetting experienced by the space-craft is

$$\sigma_{\rm F} = 0.5 \, C_{\rm D} \, A \, \rho_0 \, u_0 . v_{\rm RMS} (2 + \beta u_0)$$

where  $v_{RMS}$  is the RMS along-track velocity,  $u_0$  is the shuttle speed,

and 
$$\beta \propto 1/\sqrt{(gT)} \sqrt{(dT/dz + \Gamma_a)}$$

The nett result is that at low speeds, (< Mach 1), buffeting is due to standard turbulence, BUT... at high speeds (~ Mach 18) the dominant source of buffeting is the spacecraft bouncing off gravity-wave density perturbations as it hurtles through the air!



Sprites and Elves – 30-95 km "lightning" and electromagnetic pulses (durations a few milliseconds)

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