

Title: Canonical Quantum Gravity - 3

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Abstract:

# A Double Dose of Abelian LQG

Casey Tomlin

with Adam Henderson, Alok Laddha, and Madhavan Varadarajan





## Quantum General Covariance

- Every theory with diff symmetry has constraints that generate the “hypersurface deformation” algebra

$$\{D[\vec{N}], D[\vec{M}]\} = D[\mathcal{L}_{\vec{N}}\vec{M}]$$

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- State-of-the-art LQG does not sufficiently capture this relation:
  - Algebra computed partially “on-shell” [Nikolai et al.]
  - Density weight responsible for trivial RHS [Lewandowski, Pullin et al.]
  - Ultralocality responsible for trivial LHS

## Smolin's Weak-Coupling Limit<sup>1</sup>

Euclidean, self-dual, first order action:

$$S[e, \omega] = \frac{1}{G_N} \int |e| e_i^\mu e_j^\nu R_{\mu\nu}{}^{IJ}[\omega], \quad \omega_\mu{}^{IJ} = \frac{1}{2} \epsilon^{IJ}{}_{KL} \omega_\mu{}^{KL}$$

Define  $A = G_N^{-1} \omega$ , take  $G_N \rightarrow 0$ , 3+1 split, get

$$S[A, E] = \int dt \left( \int_\Sigma d^3x E_i^a \dot{A}_a^i - G[\Lambda] - D[\vec{N}] - H[\underline{N}] \right)$$

where

$$G[\Lambda] = \int \Lambda^i \partial_a E_i^a \quad \text{—U(1)<sup>3</sup> Gauss}$$

$$D[\vec{N}] = \int E_i^a \mathcal{L}_{\vec{N}} A_a^i \quad \text{—diffeo}$$

$$H[\underline{N}] = \frac{1}{2} \int \underline{N} \epsilon^{ijk} E_i^a E_j^b F_{ab}^k[A] \quad \text{—Euclidean Hamiltonian with}$$

**Abelian** curvature  $F_{ab}^i := 2\partial_{[a} A_{b]}^i$

Subalgebra of  $D$  and  $H$  again generates the HD algebra

**Goal:** Quantize  $H$  such that  $[H, H] = D$  off-shell

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<sup>1</sup>CQG 9 883 1992

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$$\hat{F}_\delta^i = \frac{\text{tr}(h_{\square}\tau^i)}{\delta^2} + \frac{3i}{2\ell_{\text{P}}^2} \frac{\text{tr}(h_{\square} - \mathbf{1})}{\delta^2} \hat{E}_i(S_\delta)$$

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- What space supports the  $\delta \rightarrow 0$  limit?
  - Not  $\mathcal{H}_{\text{kin}}$ , but a well-chosen set of distributions (Lewandowski-Marolf habitat):

$$\hat{D}[\vec{N}] \Psi^f \sim \Psi^{\mathcal{L}_{\vec{N}} f}$$

- Apply these ideas to the  $U(1)^3$  Hamiltonian constraint

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## 3+1 Finite-Triangulation Hamiltonian

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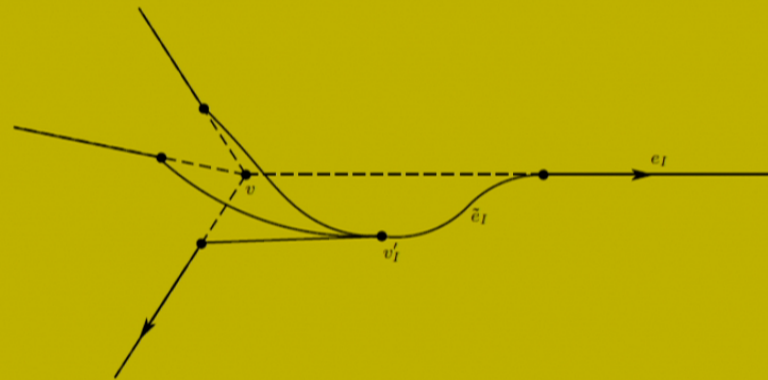
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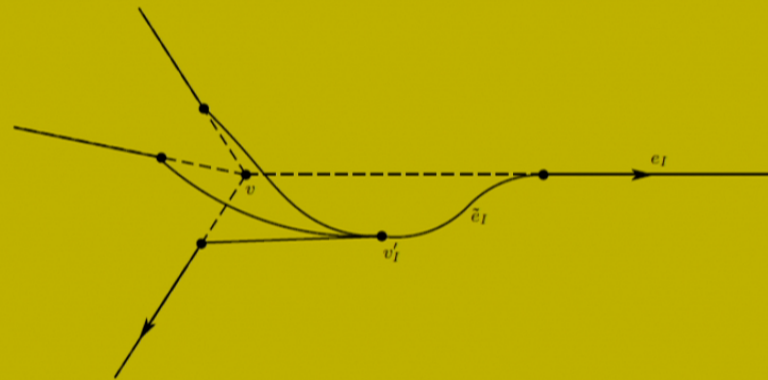
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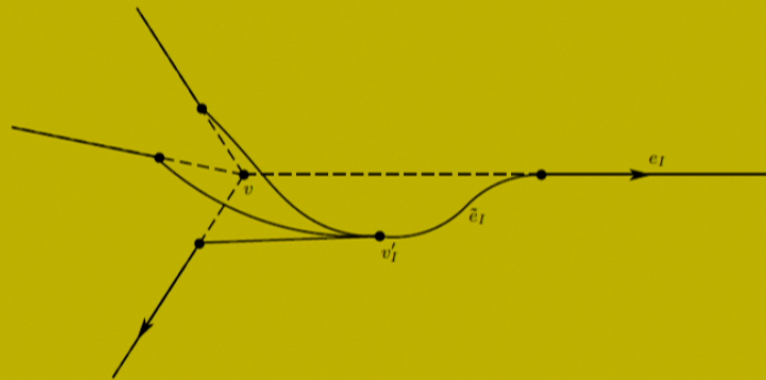
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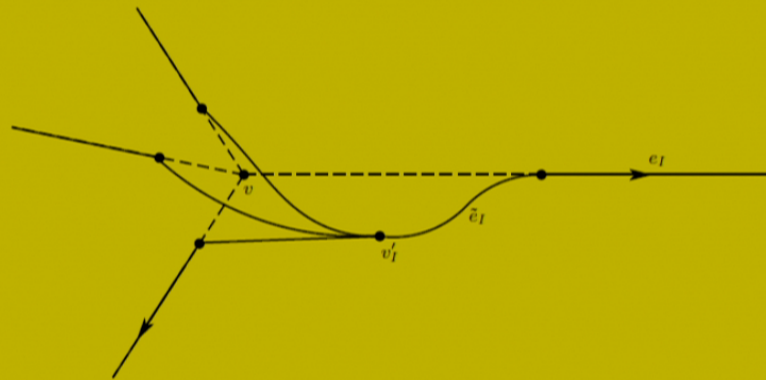
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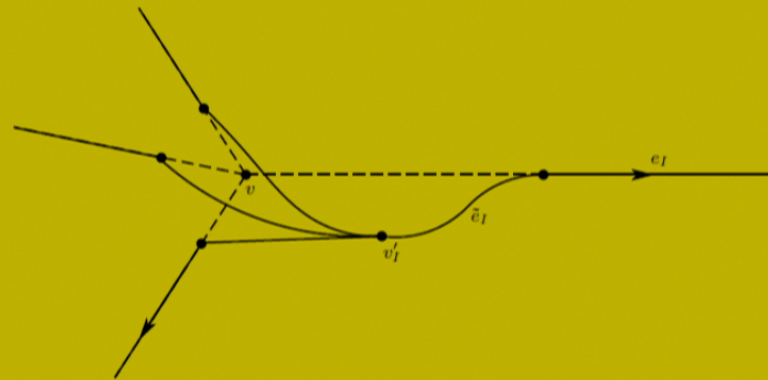
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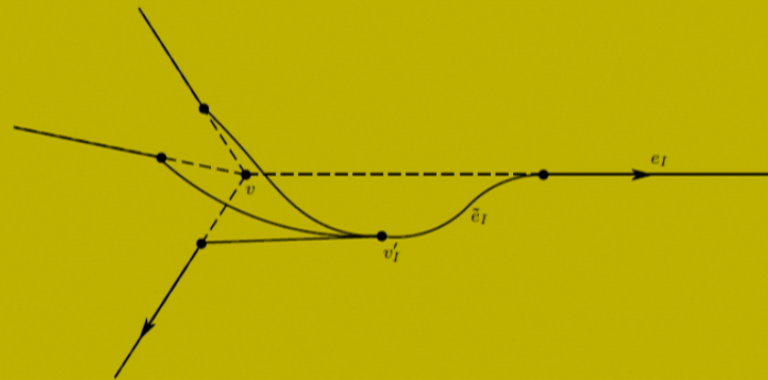
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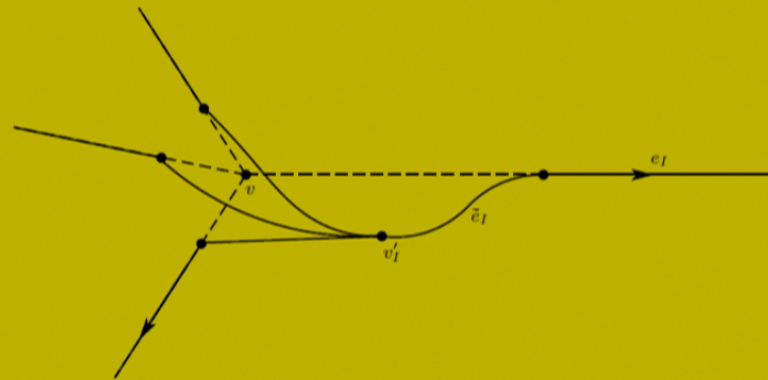
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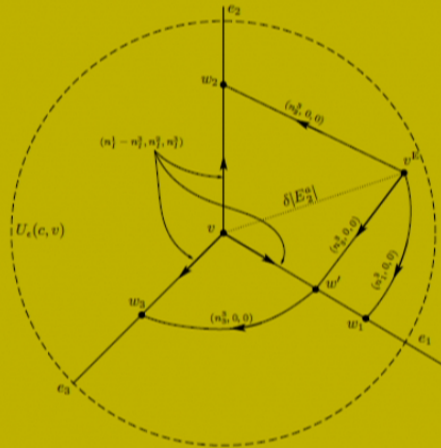
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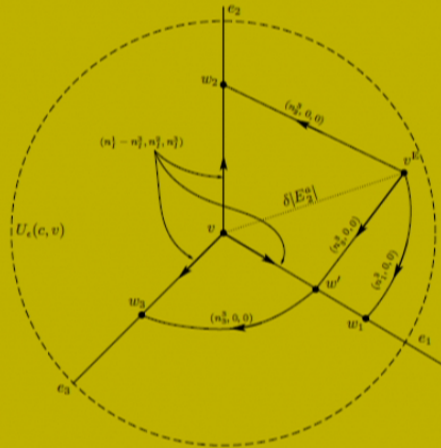
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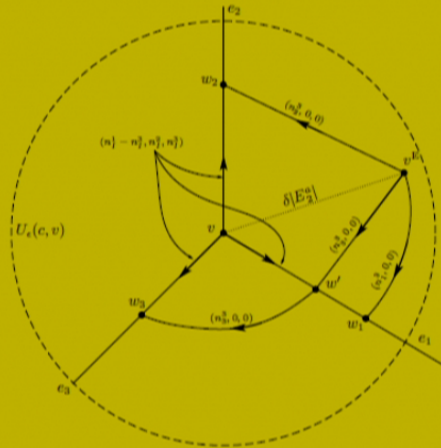
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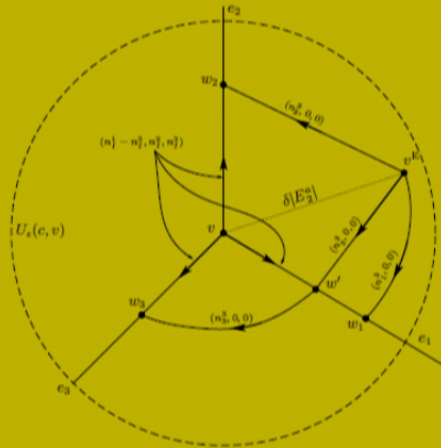
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Nontrivial action of 2<sup>nd</sup>  $\hat{H}$  on quantum shift components which gave the first deformation



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## Remarks

- ▶ Off-shell closure in a precise sense

$$\lim_{\delta, \delta' \rightarrow 0} (\Psi | [\hat{H}[N], \hat{H}[M]]_{\delta, \delta'} | c \rangle) = \lim_{\delta, \delta' \rightarrow 0} (\Psi | \hat{D}[\vec{\omega}]_{\delta, \delta'} | c \rangle)$$

and there is more than one way to do it.

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- ▶ Speculation: U(1)<sup>3</sup> E-representation has linear-in-momenta constraints. Could investigate in flux rep of Dittrich et al.

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## Why use Effective Equations?

- Correlation functions are calculated with an absolutely **generalized initial state**, as required for cosmology.
- Can avoid several technical difficulties like the exact structure of inner products on the Hilbert space, or the non-unique nature of self-adjoint extensions.
- Systematic way to realize **higher derivative corrections** in the equations of motion for a canonically quantized system.
- New perspective on known features of QFT, like renormalization, which may prove to be useful while quantizing with a dynamical background.
- Being **canonical**, applicable to certain models of LQG and LQC.



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## The New Variables

[M. Bojowald and A. Skirzewski, 2006]

- Define expectation values, with respect to some state, as:

$$\tilde{G}^{a,n} := \langle (\hat{p} - \langle \hat{p} \rangle)^a (\hat{q} - \langle \hat{q} \rangle)^{n-a} \rangle_{\text{Weyl}} \quad (2.1)$$

- Begin with a Hamiltonian operator:  $\hat{H} = \hat{H}(\hat{q}, \hat{p})$   
Take its expectation value with respect to the same state to define an **'effective' Quantum Hamiltonian**

$$\begin{aligned} H_Q := \langle \hat{H} \rangle &= \left\langle \hat{H}(\langle \hat{q} \rangle + (\hat{q} - \langle \hat{q} \rangle), \langle \hat{p} \rangle + (\hat{p} - \langle \hat{p} \rangle)) \right\rangle \\ &= \sum_{n=0}^{\infty} \sum_{a=0}^n \frac{1}{n!} \binom{n}{a} \frac{\partial^n H(q, p)}{\partial p^a \partial q^{n-a}} \tilde{G}^{a,n} \end{aligned} \quad (2.2)$$

- A point in this infinite dimensional space is completely specified by  $(\langle \hat{q} \rangle, \langle \hat{p} \rangle, \tilde{G}^{a,n})$



## The Equations of Motion

Let  $q := \langle \hat{q} \rangle$  and  $p := \langle \hat{p} \rangle$ .

The Hamilton's equations of motion gives us

$$\dot{q} = \{q, H_Q\} \quad (2.5)$$

$$\dot{p} = \{p, H_Q\} \quad (2.6)$$

$$\dot{\tilde{G}}^{a,n} = \{\tilde{G}^{a,n}, H_Q\} \quad (2.7)$$

Instead of solving the Schrödinger's partial differential equation, we have to solve this **infinite set of coupled ordinary differential equations**.

- The validity of the solutions to these equations of motion are subject to certain '**Uncertainty Relations**', imposed on the moments.



## The Effective Quantum Hamiltonian

The Hamiltonian for an oscillator with a perturbation term is

$$\hat{H} = \frac{1}{2m}\hat{p}^2 + \frac{1}{2}m\omega^2\hat{q}^2 + \hat{U}(\hat{q})$$

The corresponding 'effective' Quantum Hamiltonian is

$$H_Q = \frac{1}{2m}p^2 + \frac{1}{2}m\omega^2q^2 + U(q) + \frac{\hbar\omega}{2}(G^{0,2} + G^{2,2}) + \sum_n \frac{1}{n!}(\hbar/m\omega)^{n/2}U^{(n)}(q)G^{0,n} \quad (2.8)$$

where  $G^{a,n} = \hbar^{-n/2}(m\omega)^{n/2-a}\tilde{G}^{a,n}$  are now dimensionless quantities.



We need to make two approximations:

- Moments need to be solved **perturbatively** in  $(\frac{\hbar}{L})^{1/2}$ . Here  $L$  is some angular momentum scale provided by the perturbing potential.
- Need to make an **adiabatic approximation** for the moments where we assume they are slowly varying with time but the evolution of  $q$  and  $p$  are free. Derivatives with respect to time in equations of motion are rescaled as  $\frac{d}{dt} \rightarrow \lambda \frac{d}{dt}$ . In the end, we shall set  $\lambda = 1$ .

Thus, we can expand the moments as

$$G^{a,n} = \sum_c \sum_j G_{e,j}^{a,n} \left(\frac{\hbar}{L}\right)^{e/2} \lambda^j \quad (2.10)$$

At a given order in  $\sqrt{\frac{\hbar}{L}}$ , denoted by the index  $e$ , the adiabatic approximation gives

$$0 = \{G_{e,j}^{a,n}, H_0\} \quad (2.11)$$

to leading order, and

$$G_{e,j}^{a,n} = \{G_{e,j}^{a,n}, H_0\} \quad (2.12)$$

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to leading order, and

$$\dot{G}_{e,j}^{a,n} = \{G_{e,j+1}^{a,n}, H_Q\} \quad (2.12)$$

for higher orders.



## Equation of motion for $q$ up to $\hbar^{3/2}$ and fourth adiabatic order

We may now rewrite the equation of motion as:

$$\ddot{q} = -\omega^2 q - U'(q)/m \quad (2.16)$$

$$-\frac{\hbar}{2m^2\omega} U'''(q) [f(q, \dot{q}) + f_1(q, \dot{q})\ddot{q} + f_2(q)\ddot{q}^2 + f_3(q, \dot{q})\ddot{q} + f_4(q)\ddot{q}^2] + \mathcal{O}(\hbar^2)$$

where

$$f(q, \dot{q}) = \frac{1}{2} \left(1 + \frac{U''(q)}{m\omega^2}\right)^{-1/2} + \frac{U''''(q)\dot{q}^2}{16m\omega^4} \left(1 + \frac{U''(q)}{m\omega^2}\right)^{-5/2} - \frac{5(U''''(q))^2\dot{q}^2}{64m^2\omega^6} \left(1 + \frac{U''(q)}{m\omega^2}\right)^{-7/2}$$

$$- \frac{U''''''(q)\dot{q}^4}{64m\omega^6} \left(1 + \frac{U''(q)}{m\omega^2}\right)^{-7/2} + \frac{21(U''''(q))^2\dot{q}^4}{256m^2\omega^8} \left(1 + \frac{U''(q)}{m\omega^2}\right)^{-9/2}$$

$$+ \frac{7U''''''(q)U''''(q)\dot{q}^4}{64m^2\omega^8} \left(1 + \frac{U''(q)}{m\omega^2}\right)^{-9/2} - \frac{231U''''(q)(U''''(q))^2\dot{q}^4}{512m^3\omega^{10}} \left(1 + \frac{U''(q)}{m\omega^2}\right)^{-11/2}$$

$$+ \frac{1155(U''''(q))^4\dot{q}^4}{4096m^4\omega^{12}} \left(1 + \frac{U''(q)}{m\omega^2}\right)^{-13/2} \quad (2.17)$$



## CW Potential for a (0 + 1)-dimensional system

[S. Coleman and E. Weinberg, 1973]

For a given Lagrangian  $L(q, \dot{q}, t) = \frac{1}{2}m\dot{q}^2 - V(q)$ , with a vev defined by  $\langle 0 | q | 0 \rangle := q_0$ , the Effective Coleman-Weinberg potential is given by

$$V_{\text{eff}}(q) = V(q_0) + \frac{\hbar}{2\sqrt{m}} \int \frac{dk}{2\pi} \log \left( \frac{k^2 + V''(q_0)}{k^2} \right) + O(\hbar^2) \quad (2.22)$$

This integral is obviously convergent and it gives:

$$V_{\text{eff}}(q) = V(q_0) + \frac{\hbar}{2\sqrt{m}} \sqrt{V''(q_0)} + O(\hbar^2) \quad (2.23)$$



## The Setup

- Use the ‘in-in’ formalism to get equal-time correlation functions
- The ‘phi-fourth’ Hamiltonian

$$\hat{H} = \int d^3x \left[ \frac{\hat{\pi}^2(x)}{2} + \frac{m^2}{2} \hat{\phi}^2(x) + \frac{1}{2} (\nabla \hat{\phi}(x))^2 + \lambda \hat{\phi}^4(x) \right]$$

Define

$$G^{a,b}(x_1, \dots, x_a; y_1, \dots, y_b, t) := \left\langle (\hat{\pi}(x_1, t) - \langle \hat{\pi}(x_1, t) \rangle) \dots (\hat{\pi}(x_a, t) - \langle \hat{\pi}(x_a, t) \rangle) \times (\hat{\phi}(y_1, t) - \langle \hat{\phi}(y_1, t) \rangle) \dots (\hat{\phi}(y_b, t) - \langle \hat{\phi}(y_b, t) \rangle) \right\rangle_{\text{Weyl}} \quad (3.1)$$

$$\nabla_{x_i} \nabla_{y_j} \left[ G^{a,b}(x_1, \dots, x_a; y_1, \dots, y_b, t) \right] := \left\langle (\hat{\pi}(x_1, t) - \langle \hat{\pi}(x_1, t) \rangle) \dots \nabla_{x_i} (\hat{\pi}(x_i, t) - \langle \hat{\pi}(x_i, t) \rangle) \dots \times (\hat{\phi}(y_1, t) - \langle \hat{\phi}(y_1, t) \rangle) \dots \nabla_{y_j} (\hat{\phi}(y_j, t) - \langle \hat{\phi}(y_j, t) \rangle) \dots \right\rangle_{\text{Weyl}} \quad (3.2)$$







With  $\langle \hat{\pi}(x) \rangle := \pi(x)$  and  $\langle \hat{\phi}(x) \rangle := \phi(x)$ ,

$$\begin{aligned}
 H_Q = \frac{1}{2} \int d^3x & \left[ \pi^2(x) + G^{2,0}(x, x) + m^2 (\phi^2(x) + G^{0,2}(x, x)) \right. \\
 & + \nabla_x^2 (G^{0,2}(x, x)) + (\nabla \phi(x))^2 + 2\lambda \{ \phi^4(x) \\
 & \left. + 6\phi^2(x) G^{0,2}(x, x) + 4\phi(x) G^{0,3}(x, x, x) + G^{0,4}(x, x, x, x) \} \right] \quad (3.3)
 \end{aligned}$$

The (equal time) Poisson Algebra is defined as:

$$\{ \phi(x), \pi(y) \} := \frac{1}{i\hbar} \left\langle \left[ \hat{\phi}(x), \hat{\pi}(x) \right] \right\rangle = \delta^3(x - y) \quad (3.4)$$

The equations of motion are derived as:

$$\frac{d}{dt} [\mathcal{O}] := \{ H_Q, \mathcal{O} \} \quad (3.5)$$



## EOM (Higher Order Moments)

The general scheme for equations of higher order moments

$$\begin{aligned}
 \dot{G}^{0,n}(y_1, \dots, y_n, t) &\sim G^{1,n-1}(y_1, \dots, y_n, t) \\
 \dot{G}^{1,n-1}(y_1, \dots, y_n, t) &\sim G^{2,n-2}(y_1, \dots, y_n, t) + G^{0,n}(y_1, \dots, y_n, t) \\
 &\quad + \lambda G^{0,n+2}(y_1, \dots, y_n, t) \\
 &\quad \vdots \\
 \dot{G}^{n-1,1}(y_1, \dots, y_n, t) &\sim G^{n-2,2}(y_1, \dots, y_n, t) + G^{n,0}(y_1, \dots, y_n, t) \\
 &\quad + \lambda G^{n-2,4}(y_1, \dots, y_n, t) \\
 \dot{G}^{n,0}(y_1, \dots, y_n, t) &\sim G^{n-1,1}(y_1, \dots, y_n, t) + \lambda G^{n-1,3}(y_1, \dots, y_n, t) \quad (3.8)
 \end{aligned}$$



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 H_Q = \frac{1}{2} \int d^3x & \left[ \pi^2(x) + G^{2,0}(x, x) + m^2 (\phi^2(x) + G^{0,2}(x, x)) \right. \\
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## Solving these equations

- Expand the moments in powers of the coupling constant,  
$$G^{a,b} = \sum_e \bar{\lambda} G_e^{a,b}$$
- Solve for the moments in lower orders in  $\bar{\lambda}$ , **starting with the free field solutions.**
- Plug the (solved) lower order  $\bar{\lambda}$  moments, in the equations containing higher order in  $\bar{\lambda}$ .
- In this way, perturbatively solve for the moments, which shall give us the required **correlation functions.**



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## Cancellation of the tadpole term

For  $\phi^4$  theory,

$$\begin{aligned}\ddot{\phi}(y, t) = & -(m^2 - \nabla_y^2)\phi(y, t) \\ & + 4\lambda\phi^3(y, t) + 12\lambda\phi(y, t)G^{0,2}(y, y, t) \\ & + 4\lambda G^{0,3}(y, y, y, t)\end{aligned}\quad (3.9)$$

In this case,  $\phi(y, t) = 0$  is easily a solution up to any order since all odd moments (including  $G^{0,3}(y_1, y_2, y_3, t)$ ) are zero up to any order.

For  $\phi^3$  theory,

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In order for  $\phi(y, t) = 0$  to be a solution of this equation, we require an additional term (proportional to  $\phi$ ) in the Hamiltonian (or equivalently, Lagrangian) which will cancel off the  $G^{0,2}(y, y, t)$  up to whichever order we want.



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For a particular initial value of the moments, given by

$$\begin{aligned} G^{0,2}(y, z, 0) &= \frac{1}{2\pi^3} \int \frac{d^3k}{2\sqrt{k^2 + m^2}} e^{i\vec{k}\cdot(\vec{y}-\vec{z})} \quad \text{and} \\ \dot{G}^{0,2}(y, z, 0) &= 0 \end{aligned} \quad (3.13)$$

we reproduce the usual result from QFT, that is,

$$G^{0,2}(y, z, t) = \int \frac{d^3k}{2(2\pi^3)\sqrt{k^2 + m^2}} e^{i\vec{k}\cdot(\vec{y}-\vec{z})} \quad (3.14)$$

The unique factorization of  $\omega = \omega_y - \omega_z$  is why the two results (rightly) match up.

- The propagator has been calculated to agree up to one loop order with QFT.



## Important lessons and looking ahead

So, why Effective Equations?

- Using these canonical techniques for effective action, we **recover** the usual **QFT results** and also extend them, for instance, by including **more general states**.
- There is well defined **systematic way to derive the higher derivative corrections** while avoiding some technical difficulties.

Where are these useful?

- Currently being applied to certain models of **isotropic, homogeneous cosmology** and also to a **de Sitter background**.
- Current work is underway to include (perturbative) **quantum corrections** in the Scalar and Diffeomorphism constraints of **spherical LQG**, and see what effects they have on the hyperspace deformation algebra. In the high curvature regime, these might be of the same order as that of other non-perturbative corrections (like holonomy corrections), and hence they should be included for a full analysis.

# Born-Oppenheimer decomposition for non-commuting slow variables

Alexander Stottmeister

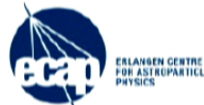
(work with T. Thiemann (forthcoming))

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July 25, 2013

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Waterloo, Canada



## Outline

- 1 Conceptual setup
- 2 A coherent state approach to the Born-Oppenheimer decomposition
- 3  $T^*U(1)$  &  $T^*SU(2)$  theories
  - Applications to LQG, LQC & WdW
- 4 Outlook & Final Remarks



## The Born-Oppenheimer decomposition in canonical LQG

The Born-Oppenheimer decomposition has a long tradition in quantum gravity (cf. [Kiefer, 2004]). It consists in a splitting of gravity-matter system  $\mathfrak{H}$  in **slow**  $\mathfrak{H}_S$  and **fast**  $\mathfrak{H}_F$  sectors.

$$\mathfrak{H} = \mathfrak{H}_S \otimes \mathfrak{H}_F \quad (1)$$

Approximation schemes in this setting are governed by one or more **adiabatic scales**, e.g.  $\frac{m_\Phi}{m_P}$ .

Its **direct application** to loop quantum gravity is **prevented by the noncommutativity of the fluxes**, if one intends to work in a representation admitting a **parametrisation by classical metrics** ( $q_{ab}$ ) in the **fast** sector, e.g. quantum fields on classical backgrounds (cf. [Giesel, Tambornino, Thiemann, 2009]).

It is important to note, that the Born-Oppenheimer decomposition is **not a semiclassical approximation scheme per se** (cf. [Teufel, 2003]), but we will **combine it with coherent state methods** (cf. [Faure & Zhilinskii, 2001], in the context of spin systems) enabling us to consider the semiclassical ( $\hbar \rightarrow 0$ ) and the adiabatic limit **simultaneously**. Furthermore, these methods allow us to surpass the difficulties imposed by the noncommutativity to some extent.

## Remarks

1. It is interesting to note that the passage from a theory of **quantum gravity** to **quantum field theory** on curved spacetimes can be interpreted as a **measurement problem in disguise**.

Recently, the possible need for a simultaneous consideration of the adiabatic and semiclassical limit w.r.t. the measurement problem has been pointed out (cf. [Landsman & Reuvers, 2013]).

2. The application of the framework to **deparametrised models** seems to be especially interesting (cf. [Giesel & Thiemann, 2012] for an overview).
3. From a mathematical point of view intriguing connections with **complex geometry**, **microlocal analysis** and **pseudo-differential calculus** arise (cf. [Teufel, 2003, Hörmander, 1983, Hörmander, 1985]).

## The (generalised) Born-Oppenheimer decomposition

(cf. [Chruściński & Jamiolkowski, 2004, Kiefer, 2004])

Let us shortly discuss the setup of the **traditional Born-Oppenheimer scheme**:

Consider a quantum mechanical system described by a triple

$$H, \mathfrak{H}, \mathfrak{A} = \{Q, P, q, p\}. \quad (2)$$

Assume a splitting

$$\mathfrak{H} = \mathfrak{H}_S \otimes \mathfrak{H}_F, \quad \mathfrak{A} = \mathfrak{A}_S \otimes \mathfrak{A}_F \quad (3)$$

and a decomposition

$$H = H_S(P, Q) \otimes \mathbb{1}_F + H_{S \otimes F}(Q, p, q). \quad (4)$$

If the subset  $\{Q\} \subset \mathfrak{A}_S$  is a **commutative subalgebra** ( $\cong C(\sigma(Q))$ ), one considers the **restriction** of  $H_{S \otimes F}$  to its joint spectral subspaces  $\mathfrak{H}_Q \cong \mathfrak{H}_F$

$$H_{S \otimes F}|_{\mathfrak{H}_Q} : \mathfrak{H}_Q \longrightarrow \mathfrak{H}_Q. \quad (5)$$

The **eigenvalue problem** of  $H$  is parametrised by the eigenvalue problems of the restrictions  $H_{S \otimes F}|_{\mathfrak{H}_Q}$ .

$$(\lambda_\Psi - E_{nF}(Q)) \Psi_{nF}(Q) = (H_S(P + A^F(Q), Q) \circ \Psi)_{nF}(Q) \quad (BOE) \quad (6)$$

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$$\mathfrak{H} = \mathfrak{H}_S \otimes \mathfrak{H}_F, \quad \mathfrak{A} = \mathfrak{A}_S \otimes \mathfrak{A}_F \quad (3)$$

and a decomposition

$$H = H_S(P, Q) \otimes \mathbb{1}_F + H_{S \otimes F}(Q, p, q). \quad (4)$$

If the subset  $\{Q\} \subset \mathfrak{A}_S$  is a **commutative subalgebra** ( $\cong C(\sigma(Q))$ ), one considers the **restriction** of  $H_{S \otimes F}$  to its joint spectral subspaces  $\mathfrak{H}_Q \cong \mathfrak{H}_F$

$$H_{S \otimes F}|_{\mathfrak{H}_Q} : \mathfrak{H}_Q \longrightarrow \mathfrak{H}_Q. \quad (5)$$

The **eigenvalue problem** of  $H$  is parametrised by the eigenvalue problems of the restrictions  $H_{S \otimes F}|_{\mathfrak{H}_Q}$ .

$$(\lambda_\Psi - E_{nF}(Q)) \Psi_{nF}(Q) = (H_S(P + A^F(Q), Q) \circ \Psi)_{nF}(Q) \quad (BOE) \quad (6)$$



## The (generalised) Born-Oppenheimer decomposition

(cf. [Chruściński & Jamiolkowski, 2004, Kiefer, 2004])

Let us shortly discuss the setup of the **traditional Born-Oppenheimer scheme**:

Consider a quantum mechanical system described by a triple

$$H, \mathfrak{H}, \mathfrak{A} = \{Q, P, q, p\}. \quad (2)$$

Assume a splitting

$$\mathfrak{H} = \mathfrak{H}_S \otimes \mathfrak{H}_F, \quad \mathfrak{A} = \mathfrak{A}_S \otimes \mathfrak{A}_F \quad (3)$$

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## Remarks

1. The object  $A^F = A^F(Q)$  is a (generalised) *Berry-Simon connection*.
2. The Born-Oppenheimer equation (6) is **exact**, but there are several approximation schemes, e.g. the **Born-Oppenheimer approximation** ( $A^F = 0$ ) or the **adiabatic or no-mixing approximation** ( $A^F$  preserves the eigenspaces of (5)).
3. **Effective Hamiltonians** for the slow variables arise, if we **restrict the slow dynamics** to spectral subspaces of  $H_{S \otimes F|S_Q}$ , i.e.

$$H_{EF}(P, Q)_{nm} := \langle n^F | H | m^F \rangle. \quad (7)$$

In the adiabatic approximation, we are led to the **quantum geometric forces** in  $\sigma(Q)$ .

4. Additional complications arise due to **spectral instabilities**, e.g. eigenvalue crossings or bifurcations.
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## The coherent state form of the Born-Oppenheimer decomposition I

Again, we consider a quantum mechanical system given in terms of a triple

$$H, \mathfrak{H}, \mathfrak{A} = \{A, A^*, q, p\}, \quad (8)$$

a splitting

$$\mathfrak{H} = \mathfrak{H}_S \otimes \mathfrak{H}_F, \quad \mathfrak{A} = \mathfrak{A}_S \otimes \mathfrak{A}_F \quad (9)$$

and a decomposition

$$H = H_S(A, A^*) \otimes \mathbb{1}_F + H_{S \otimes F}(A, A^*, q, p). \quad (10)$$

Furthermore, we require the existence of a **complete** set of **coherent states** in the **slow sector**, but there are **NO** commutativity assumptions

$$A|z\rangle = z|z\rangle, \quad \mathbb{1}_S = \int_{\Gamma_{\mathbb{C}}^S} d\mu(z, \bar{z}) |z\rangle\langle z|, \quad \langle z|O_S|z\rangle = 0 \Leftrightarrow O_S = 0. \quad (11)$$

This allows us to obtain a **diagonal form** for  $H_{S \otimes F}$  in  $\mathfrak{H}_S$

$$H_{S \otimes F} = \int_{\Gamma_{\mathbb{C}}^S} d\mu(z, \bar{z}) P_{H_{S \otimes F}}(z, \bar{z}) |z\rangle\langle z|. \quad (12)$$



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## The coherent state form of the Born-Oppenheimer decomposition II

In the following, we will use the spectral decomposition of the **upper symbol**

$$P_{H_{S \otimes F}}(z, \bar{z}) : \mathfrak{H}_F(z, \bar{z}) \longrightarrow \mathfrak{H}_F(z, \bar{z}) \quad (13)$$

as input for the eigenvalue problem of  $H$ .

$$(K \circ U((\lambda_\Psi - E_{n^F}))\Psi)_{n^F}(z, \bar{z}) = \left( H_S(z, \bar{z}, \partial^{A^F}, \bar{\partial}^{A^F}) \circ \Psi \right)_{n^F}(z, \bar{z}) \quad (CSBOE) \quad (14)$$

The **resolution of unity** of the coherent states system identifies a **fibre bundle structure**

$$\mathfrak{H}_S \otimes \mathfrak{H}_F \cong \mathcal{H}L^2(\Gamma_{\mathbb{C}}^S, d\mu, \mathfrak{H}_F) \quad (15)$$

and typically leads to additional **flatness constraints** on the solutions

$$\partial^{A^F}(\rho\Psi)_{n^F}(z, \bar{z}) = 0. \quad (16)$$



## Remarks

1. The **upper symbol** is generically **not unique**. It can be determined from the **lower symbol**  $\langle z | H_{S \otimes F} | z \rangle$  by **duality** (non trivial), and exists for a (strongly) dense set of operators (cf. [Simon, 1980, Klauder & Skagerstam, 1985]).
2. The solutions to the eigenvalue equation of  $H$  can be considered as **holomorphic, horizontal sections** of a  $\mathfrak{h}_F$ -bundle on the complexified (slow) phase space  $\Gamma_{\mathbb{C}}^S$ .
3. The main difference of this approach arises through the **integral operator**  $K \circ U$ , which is composed of the **coherent states kernel**  $K$  and the **bundle transition operator**  $U$ .
4. Similar to the traditional approach, we obtain a (generalised) **Berry-Simon connection**  $A^F = A^F(z, \bar{z})$ .
5. Completeness can be conveniently achieved in **complexifier framework** exploiting **holomorphicity** (cf. [Thiemann, 2006]). The associated covariant differential can be interpreted as **covariant Dolbeault operator**.
6. **Approximation schemes** similar to those of the traditional Born-Oppenheimer approach are **conceivable** in the **semiclassical limit**, but require a detailed **asymptotic expansion** of  $K \circ U$  and the upper symbol  $P_{H_{S \otimes F}} = P_{H_{S \otimes F}}(z, \bar{z})$  (heat kernel analysis, microlocal analysis).  
Essentially the same is required for effective Hamiltonians, spectral stability and gap conditions.

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Essentially the same is required for effective Hamiltonians, spectral stability and gap conditions.

## $T^*U(1)$ coherent states I

(cf. [Hall, 1994, Varadarajan, 2000, Thiemann & Winkler, 2001a, Hall & Mitchell, 2002], [Kowalski, Rembielinski, Papaloucas, 1996])

If we consider theories with phase space  $T^*U(1) \cong U(1)_{\mathbb{C}} \cong \mathbb{C}^*$  for the slow sector, the (slow) Hilbert space will be given by

$$\mathfrak{H}_S = L^2(\mathbb{R}_{\text{Bohr}}, d\mu_H) \cong \bigoplus_{\theta \in S^1} \mathfrak{H}_\theta. \quad (17)$$

Two sets of elementary operators are important for the construction

$$\{U, J \mid J^* = J, U^* = U^{-1}, [J, U] = U\}, \{X, X^* \mid X = e^{-\frac{1}{2}J^2} U e^{\frac{1}{2}J^2} = U e^{-J - \frac{1}{2}}\}. \quad (18)$$

The coherent state system is constructed by **heat kernel methods** and adapted to the decomposition of  $\mathfrak{H}_S$ :

$$X|\xi\rangle_{j_0} = \xi|\xi\rangle_{j_0}, \quad j_0 = \frac{\theta}{2\pi}, \quad |\xi\rangle_{j_0} = c_0^{j_0}(\xi) \sum_{j \in \mathbb{Z}} e^{-\frac{1}{2}j^2} (e^{j_0} \xi)^{-j} |j\rangle, \quad (19)$$
$$\mathbb{1}_{\mathfrak{H}_\theta} = \int_{\mathbb{C}^*} \frac{d\xi \wedge d\bar{\xi}}{4\pi^{\frac{3}{2}} i |\xi|^2} e^{-(\ln|\xi| + j_0)^2} |\xi\rangle_{j_0} \langle_{j_0} \xi|, \quad \xi = e^{-l+i\varphi}.$$



## $T^*SU(2)$ coherent states I

(cf. [Hall, 1994, Thiemann, 2001, Thiemann & Winkler, 2001a], [Bianchi, Magliaro, Perini, 2010])

Theories admitting a phase space  $T^*SU(2) \cong SU(2)_{\mathbb{C}} \cong SL_2(\mathbb{C})$  for the slow sector, will be formulated on the (slow) Hilbert space

$$\mathfrak{H}_S = L^2(SU(2), d\mu_H). \quad (23)$$

Similar to the  $U(1)$ -case, we need to sets of **elementary operators**

$$\{h, \vec{p} \mid [h, h] = 0, [\vec{p}, h] = -\frac{1}{2}\vec{\sigma}h, \vec{p} \times \vec{p} = i\vec{p}\}, \{A, A^* \mid A = e^{\frac{1}{2}\vec{p}^2} h e^{-\frac{1}{2}\vec{p}^2} = e^{-\frac{1}{2}(\vec{p}\cdot\vec{\sigma} - \frac{3}{4})} h\} \quad (24)$$

Again, the coherent state system is constructed by heat kernel methods:

$$A|g\rangle = g|g\rangle, |g\rangle = \sum_{j \in \frac{1}{2}\mathbb{N}_0} (2j+1) e^{-\frac{1}{2}j(j+1)} \sum_{m,n=-j}^j \pi_j(g)_{mn} |jmn\rangle \quad (25)$$

$$\mathbb{1}_{H_S} = \int_{SL_2(\mathbb{C})} d\mu(g, \bar{g}) |g\rangle\langle g|, g = e^{-\vec{l}\cdot\vec{\sigma}} h, h \in SU(2).$$

## Applications in quantum gravity

1. In a first attempt, the method has been applied to **WdW theory**, as it makes **computations simpler**.

A specific model that was considered is **FLRW-cosmology coupled to a scalar field** leading to the reduced action

$$\begin{aligned}
 S_{\text{red}} = & \int_{\mathbb{R}} \left[ p_a \dot{a} + \int_{\Sigma} d^3x p_{\Phi} \dot{\Phi} \right. & (29) \\
 & - N \left\{ \left[ -\frac{\kappa'}{4a} p_a^2 + \frac{1}{\kappa'} \left\{ \frac{\Lambda}{3} a^3 - a \right\} \right] \right. \\
 & \left. \left. - \left[ \int_{\Sigma} d^3x \left( \frac{\lambda'}{2\sqrt{a^3 \tilde{q}}} p_{\Phi}^2 + \frac{a^3}{\lambda'} \sqrt{\tilde{q}} \left\{ \frac{1}{2a^2} \tilde{q}^{ab} (\tilde{D}_a \Phi)(\tilde{D}_b \Phi) + V(\Phi) \right\} \right) \right] \right] \right\},
 \end{aligned}$$

which leads to the following **CSBOE**

$$\lambda_{\Psi} \Psi_{\{n_k\}}(z, \bar{z}) = H_G(\hat{z}, \hat{z}^*) \Psi_{\{n_k\}}(z, \bar{z}) + \int_{\mathbb{C}} \frac{d^2z}{\pi} E_{\{n_k\}}(z, \bar{z}) \langle z|z \rangle \Psi_{\{n_k\}}(z, \bar{z}), \quad (30)$$

$$E_{\{n_k\}}(z, \bar{z}) = \frac{a(z, \bar{z})^3}{2\lambda'} e^{-\frac{9}{2}\hbar} \oint d\mu(k) (k^2 + m^2 a(z, \bar{z})^2 e^{-\frac{16}{2}\hbar}) n_k.$$



## Applications to LQG, LQC & WdW

2. The formulas for **canonical LQG** on a single graph are under consideration (**truncated dynamics**).

Earlier work in this respect makes use of the lower symbol  
(cf. [Sahlmann & Thiemann, 2006a, Sahlmann & Thiemann, 2006b]).

An extension to infinite graphs along the lines of the infinite tensor product construction is conceivable, but probably needs a revision to make the semiclassical limit feasible

([Thiemann & Winkler, 2001b], [Sahlmann, Thiemann, Winkler, 2001]).

3. The  $U(1)$ -case applies to LQC as well as linearized gravity ( $U(1)^3$ ) or Maxwell theory in the parametrized field theory framework.

(cf. [Ashtekar & Singh, 2011, Varadarajan, 2000])

4. It would be interesting to consider the **gauge invariant sector**, but it is not strictly necessary, since the coherent state mainly serve as a technical tool in the formulation of the method [Bahr & Thiemann, 2009].

Although, it could be advantageous to go to the gauge invariant sector in the approximations schemes.

5. The **volume operator** in relation to coherent states will be crucial in this approach [Flori, 2009, Flori & Thiemann, 2008], as it enters in all **vacuum & matter coupling Hamiltonians** in an essential way.

## Outlook & Final Remarks

- The asymptotic analysis of  $K \circ U$  and the upper symbol  $P_{H_{S \otimes F}} = P_{H_{S \otimes F}}(z, \bar{z})$  is technically conceivable, but **more involved** than in the original BOE.
- The proposed method applies to **general Hamiltonian systems** that admit a **fast-slow decomposition**.  
Especially, in LQG it applies to **totally constrained** as well as **deparametrised models**.
- The (nonlinear) **Fock space structures** of the  $T^*U(1)$  &  $T^*SU(2)$  theories **differ** from those discussed in the (mathematical) literature (cf. [Hall, 2001]).  
Nevertheless, these constructions generalise to the case of **Lie groups of compact type**  $K$ , as well.
- Similarly, the **inversion formulas** developed for  $T^*U(1)$  &  $T^*SU(2)$  generalise to arbitrary  $T^*K$ , but the determination of the upper symbol of the **ground state projection**  $P_{|0\rangle}$  is **non trivial**.



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# The $SL(2,R)$ totally constrained model within the Uniform Discretizations quantization approach

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## Classical system

- 1) The system consists in two pairs  $(u_i, p_i)$  and  $(v_i, \pi_i)$ , with  $i = 1, 2$ , and three constraints

$$H_1 = \frac{1}{2}(p_1^2 + p_2^2 - v_1^2 - v_2^2), \quad H_2 = \frac{1}{2}(\pi_1^2 + \pi_2^2 - u_1^2 - u_2^2),$$
$$D = u_1 p_1 + u_2 p_2 - v_1 \pi_1 - v_2 \pi_2,$$

$$\{H_1, H_2\} = D, \quad \{H_1, D\} = -2H_1, \quad \{H_2, D\} = 2H_2.$$

- 2) The constants of motion  $-so(2, 2)$  Lie algebra–

$$O_{12} = u_1 p_2 - p_1 u_2, \quad O_{23} = u_2 v_1 - p_2 \pi_1, \quad O_{14} = u_1 v_2 - p_1 \pi_2,$$
$$O_{13} = u_1 v_1 - p_1 \pi_1, \quad O_{24} = u_2 v_2 - p_2 \pi_2, \quad O_{34} = \pi_1 v_2 - v_1 \pi_2.$$

## Classical system: observables

3) A more convenient choice  $so(2, 1) \times so(2, 1)$  algebra–

$$Q_1 = \frac{1}{2}(O_{23} + O_{14}), \quad Q_2 = \frac{1}{2}(-O_{13} + O_{24}), \quad Q_3 = \frac{1}{2}(O_{12} - O_{34}),$$

$$P_1 = \frac{1}{2}(O_{23} - O_{14}), \quad P_2 = \frac{1}{2}(-O_{13} - O_{24}), \quad P_3 = \frac{1}{2}(O_{12} + O_{34}).$$

$$\{Q_i, Q_j\} = \epsilon_{ij}^k Q_k, \quad \{P_i, P_j\} = \epsilon_{ij}^k P_k, \quad \{Q_i, P_j\} = 0.$$

4) Identities between observables and constraints

$$Q_1^2 + Q_2^2 - Q_3^2 = P_1^2 + P_2^2 - P_3^2 = \frac{1}{4}(D^2 + 4H_1H_2) =: \mathcal{C},$$

$$4Q_3P_3 = (u_1^2 + u_2^2)H_1 - (u_1p_1 + u_2p_2 + v_1\pi_1 + v_2\pi_2)D - (v_1^2 + v_2^2)H_2.$$

5) Solution space: four cones joined in the origin

a)  $P_i = 0$  and  $Q_3 \in \mathbb{R}$ , with  $Q_1^2 + Q_2^2 = Q_3^2$ ,

b)  $Q_i = 0$  and  $P_3 \in \mathbb{R}$ , with  $P_1^2 + P_2^2 = P_3^2$ ,

c)  $Q_i = 0$  and  $P_i = 0$ .



## Kinematical Hilbert space

1) Kinematical Hilbert space  $\mathcal{H}_{\text{kin}} = \mathcal{L}^2(\mathbb{R}^4)$  (and  $\hbar = 1$ ).

2) Operator representation

$$\begin{aligned}\hat{p}_i\psi(u, v) &= -i\partial_{u_i}\psi(u, v), & \hat{\pi}_i\psi(u, v) &= -i\partial_{v_i}\psi(u, v), \\ \hat{u}_i\psi(u, v) &= u_i\psi(u, v), & \hat{v}_i\psi(u, v) &= v_i\psi(u, v).\end{aligned}$$

3) Quantum constraints

$$\hat{H}_1 = -\frac{1}{2}(\partial_{u_1}^2 + \partial_{u_2}^2 + v_1^2 + v_2^2), \quad \hat{H}_2 = -\frac{1}{2}(\partial_{v_1}^2 + \partial_{v_2}^2 + u_1^2 + u_2^2),$$

$$\hat{D} = -i(u_1\partial_{u_1} + u_2\partial_{u_2} - v_1\partial_{v_1} - v_2\partial_{v_2}).$$

(factor ordering of  $\hat{D}$  yields anomaly free constraint algebra)

$$[\hat{H}_1, \hat{H}_2] = i\hat{D}, \quad [\hat{H}_1, \hat{D}] = -2i\hat{H}_1, \quad [\hat{H}_2, \hat{D}] = 2i\hat{H}_2.$$

## Uniform discretizations: quantum description

1) Simultaneous diagonalization of  $\hat{H}$ ,  $\hat{H}_-$ ,  $\hat{Q}_3$  and  $\hat{P}_3$  on  $\mathcal{H}_{\text{kin}}$ :

a)  $k = \sigma_p(\hat{H}_-) \in \mathbb{Z}$ , b)  $2q_3 = \sigma_p(\hat{Q}_3) \in \mathbb{Z}$ , c)  $2p_3 = \sigma_p(\hat{P}_3) \in \mathbb{Z}$ .

d) The continuous spectrum is

$$\sigma_c(\hat{H}) = \lambda_{\text{cont}} = \frac{1}{2} + \frac{1}{2}x^2 + k^2 > 0, \quad x \in [0, \infty).$$

Otherwise, if  $k > 0$  and  $|q_3 + p_3| - |q_3 - p_3| \geq 2$  or  $k < 0$  for  $|q_3 + p_3| - |q_3 - p_3| \leq 2$  the discrete counterpart:

$$\sigma_p(\hat{H}) = \lambda_{\text{discr}} = 2t(1 - t) + k^2,$$

with  $t = 1, 2, \dots, \frac{1}{2} \min(|k|, ||q_3 + p_3| - |q_3 - p_3||)$  for even  $k$ ,

and  $t = \frac{3}{2}, \frac{5}{2}, \dots, \frac{1}{2} \min(|k|, ||q_3 + p_3| - |q_3 - p_3||)$  for odd  $k$ .

## Uniform discretizations: modified observable algebra description

1) Let us define

$$\hat{t} = \frac{1}{2}\hat{I} + \sqrt{\frac{1}{4}\hat{I} - \hat{\mathcal{C}}_{\text{disc}}}, \quad \hat{\varepsilon}_q := \hat{I} - \delta_{|\hat{Q}_3|, \hat{t}}, \quad \hat{\varepsilon}_p := \hat{I} - \delta_{|\hat{P}_3|, \hat{t}}.$$

A more convenient family of observables is  $\tilde{Q}_\pm := \hat{\varepsilon}_q \hat{Q}_\pm \hat{\varepsilon}_q$ ,  
and  $\tilde{P}_\pm := \hat{\varepsilon}_p \hat{P}_\pm \hat{\varepsilon}_p$ ,

$$\tilde{Q}_\pm |q_3, p_3\rangle_{t,k} = \pm(1 - \delta_{|q_3|, t})(1 - \delta_{|q_3 \pm 1|, t}) \frac{-i}{\sqrt{2}} [q_3 \pm t] |q_3 \pm 1, p_3\rangle_{t,k},$$

$$\tilde{P}_\pm |q_3, p_3\rangle_{t,k} = \pm(1 - \delta_{|p_3|, t})(1 - \delta_{|p_3 \pm 1|, t}) \frac{-i}{\sqrt{2}} [p_3 \pm t] |q_3, p_3 \pm 1\rangle_{t,k},$$

2) The subspaces  $\{|q_3 = \pm t, p_3\rangle_{t,k}\}$  and  $\{|q_3, p_3 = \pm t\rangle_{t,k}\}$ , with  $q_3, p_3 \in (-\infty, -t] \cup [t, \infty)$  respectively, remain invariant.

## Master constraint programme

1)  $2\hat{\mathbf{M}} = \hat{H}_+^2 + \hat{H}_-^2 + \hat{D}^2$  has a minimum nonvanishing eigenvalue in  $\sigma_c$  ( $k = 0$  and  $\lambda_{\text{cont}} = 1/2$ ) and the operators  $\hat{Q}_3$  and  $\hat{P}_3$  are unbounded on that subspace.

2) The prescription in this case

$$\hat{\mathbf{M}}'' = \hat{\mathbf{M}} - \frac{1}{2}\hat{I} + \frac{1}{2}(\hat{Q}_3\hat{P}_3)^2$$

with the (on shell) observables  $\hat{Q}'_i := |\text{sgn}(\hat{Q}_3)|\hat{Q}_i|\text{sgn}(\hat{Q}_3)|$  and  $\hat{P}'_i := |\text{sgn}(\hat{P}_3)|\hat{P}_i|\text{sgn}(\hat{P}_3)|$ .

3) Physical Hilbert space

$$\{|\lambda_{\text{cont}} = 1/2, k = 0, q_3, p_3\rangle\} \quad \text{with} \quad q_3 = 0 \quad \text{or} \quad p_3 = 0$$



## Conclusions and outlook

- 1) Within the uniform discretizations approach (as well as in the MC) we provide a prescription for the quantization of an  $SL(2, \mathbb{R})$  model by:
  - a) Considering the whole infrared spectrum of the discrete Hamiltonian (discrete dynamics).
  - b) Together with a suitable choice of observable algebra.
- 2) The physical Hilbert space is a subspace of the kinematical one.
- 3) Study of the discrete (quantum) dynamics (comparison between evolving constants and conditional probabilities).

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# Constraint Lie algebra and true local Hamiltonian for the CGHS model

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## Why the CGHS model?

The Callan-Giddings-Harvey-Strominger model is 2D dilatonic ( $\Phi$ ) model which (coupled to matter field  $f$ ) reads

$$S_{\text{CGHS}} = \int d^2x \sqrt{-|g|} \left( \frac{1}{8} \Phi^2 R + \frac{1}{2} g^{ab} \partial_a \Phi \partial_b \Phi + \frac{1}{2} \Phi^2 \lambda^2 \right) - \int d^2x \sqrt{-|g|} g^{ab} \partial_a f \partial_b f$$

- It contains black hole solutions, FRW cosmological models, Hawking radiation etc.,



## What do we address?

We propose a classical formulation of the CGHS model where we have:

- **CGHS in Ashtekar-like variables:** very similar canonical transformation from a generic 2D action  $\implies$  CGHS/3+1 in Ashtekar-like variables.

## What do we address?

We propose a classical formulation of the CGHS model where we have:

- **CGHS in Ashtekar-like variables:** very similar canonical transformation from a generic 2D action  $\implies$  CGHS/3+1 in Ashtekar-like variables.
- **Possibility of traditional Dirac quantization:** the constraint algebra is a Lie algebra even in presence of matter,

## A generic 2D dilatonic Lagrangian

- An “almost” generic diffeomorphism invariant action yielding 2nd order differential equations for the metric  $g$  and a scalar (dilaton) field  $\Phi$  in 2D

$$L_g = \sqrt{-|g|} \left\{ Y(\Phi)R + \frac{1}{2}g^{ab}\partial_a\Phi\partial_b\Phi + V(\Phi) \right\},$$

$$L_m = -\sqrt{-|g|}W(\Phi)g^{ab}\partial_a f\partial_b f.$$

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- ▶  $Y(\Phi)$ ,  $V(\Phi)$  and  $W(\Phi)$  model specific functions of the dilaton field.
- ▶ Contains CGHS ( $\Phi$ =dilaton field), 3+1 spherically symmetric ( $ds^2 = g_{\mu\nu}dx^\mu dx^\nu + \Phi^2(d\theta^2 + \sin^2(\theta)d\phi^2)$ ), etc.



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Write the theory in tetrads

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Make a Legendre transformation to a generic Hamiltonian



Use specific  $Y(\Phi)$ ,  $V(\Phi)$  and  $W(\Phi)$



Make a **canonical transformation** to new variables for CGHS:

$$P_1 = 2 \cosh(\eta) E^\varphi, \quad P_2 = 2 \sinh(\eta) E^\varphi, \quad P_\omega = E^x,$$

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A **second class Hamiltonian**  $H(K_x, E^x, K_\varphi, E^\varphi, \Phi, P_\Phi, f, P_f)$   
with two **second class constraints**  $\alpha$  and  $\mu$ :  $\{\mu, \alpha\} \neq 0$ .

## The CGHS Hamiltonian in new variables

Write the theory in tetrads

### Similar transformations for the 3+1 case

This is much like what we do to get the 3+1 case from generic action:

$$P_1 = 2 \frac{\cosh(\eta)}{E^{x\frac{1}{4}}} E^\varphi, \quad P_2 = 2 \frac{\sinh(\eta)}{E^{x\frac{1}{4}}} E^\varphi, \quad P_\omega = \frac{E^x}{2}$$

but we get a first class system.

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## Second class procedure

### Similar transformations for the 3+1 case

Finally we get a total Hamiltonian density in new variables

$$H = N\mathcal{H} + N^1\mathcal{D}$$

with  $\mathcal{H}$  and  $\mathcal{D}$  being the Hamiltonian and diffeomorphism constraints and  $N$  and  $N^1$  lapse and shift.



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## Transforming the Hamiltonian constraint

Rescale lapse and shift

$$\bar{N}^1 = N^1 + \frac{NK_\varphi}{E^{x'}}, \quad \bar{N} = N \frac{E^\varphi E^x}{E^{x'}}$$

↓

The **total Hamiltonian** density will become

$$H_T = \bar{N} \left[ \overbrace{\partial_x \left( \frac{E^{x'^2}}{2E^\varphi{}^2 E^x} - 2E^x \lambda^2 - \frac{K_\varphi^2}{2E^x} \right) - \frac{f' P_f K_\varphi}{E^x E^\varphi} + \frac{E^{x'} (f'^2 + P_f^2)}{2E^\varphi{}^2 E^x}}^{\mathcal{H}} \right] + \bar{N}^1 \left[ \underbrace{-U_x E^{x'} + f' P_f + E^\varphi K'_\varphi}_D \right]$$

## Lie algebra of constraints

Now  $\mathcal{H}$  constraint has strongly Abelian algebra with itself in both vacuum and coupled-to-matter cases:

$$\begin{aligned} & \{\mathcal{H}(x), \mathcal{H}(y)\}_D = 0, \\ & \left\{ \mathcal{H}(x) \Big|_{f=0, P_f=0}, \mathcal{H}(y) \Big|_{f=0, P_f=0} \right\}_D = 0 \end{aligned}$$

chance of implementing back-reaction?

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Constraints algebra is now a Lie algebra,

$$\{\mathcal{D}, \mathcal{D}\} = \mathcal{D}$$

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and the same method works for vacuum 3+1 model.

## Possibility of resolution of singularity?

CGHS Hamiltonian in this formulation is very similar to the 3+1 model.



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Gambini-Pullin method of **eliminating the singularity** for 3+1 can be used here? (work under development).

The metric becomes an evolving constant operator  $\implies$  self-adjointness of this operator leads to removal of singularity.

## Deparametrizing and gauge fixing

- Deparametrizing the constraints by transformation to new canonical coordinates (Brown, Kuchar, Thiemann, Giesel, Gambini, Pullin, ...).
- **Step 1: fix the first gauge**  $\zeta_1 = E^x - h(x) \approx 0$  to get

$$\dot{\zeta}_1 = 0 \implies \bar{N}^1 = 0$$

so that

$$H_{PF} = \bar{N} \left[ \partial_x \left( \frac{1}{2} \frac{h'^2}{hE\varphi^2} - 2h\lambda^2 - \frac{1}{2} \frac{K_\varphi^2}{h} \right) - \frac{f'P_f K_\varphi}{hE\varphi} + \frac{1}{2} \frac{P_f^2 h'}{hE\varphi^2} + \frac{1}{2} \frac{h'f'^2}{hE\varphi^2} \right]$$

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## Deparametrizing and gauge fixing

- **Step 3:** find the **conjugate momentum**

$$P_X = -\frac{hh'}{\Omega(K_\varphi, X) (K_\varphi + \Omega(K_\varphi, X))}$$

- **Step 4:** From  $\mathcal{H} \approx 0$  find  $K_\varphi = K_\varphi(X, f, P_f)$ . Substituting this in above gives the new **total Hamiltonian**

$$H_{\text{tot}} = \bar{N} \left( P_X + \frac{hh'}{\Omega(X, f, P_f) (K_\varphi(X, f, P_f) + \Omega(X, f, P_f))} \right)$$

## The true dynamics

- The true local Hamiltonian gives the correct evolution: since  $f$  and  $P_f$  commute with  $P_X$ , we get

$$\dot{f} = \left\{ f, \int dx H_{\text{tot}} \right\}_D = \left\{ f, \int dx H_{\text{true}} \right\}_D,$$
$$\dot{P}_f = \left\{ P_f, \int dx H_{\text{tot}} \right\}_D = \left\{ P_f, \int dx H_{\text{true}} \right\}_D,$$

- It is a **local** Hamiltonian density:  $\bar{N}$  not an integral of canonical variables, the **Hamiltonian not an integral of an integral** (non-local), eqs. of motion local.



## Summary

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## Summary

- The CGHS can be written in a similar Ashtekar-like variables as in the 3+1, from a generic 2D dilatonic Lagrangian. In this formulation:
- The constraint algebra can be cast into a Lie algebra: possibility of completing the Dirac quantization / implementing back-reaction?
- Similarity to the 3+1 Hamiltonian: possibility to eliminate the singularity a la Gambini and Pullin?