

Title: Searching for Other Universes

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URL: <http://pirsa.org/13070029>

Abstract:

Searching for Other Universes

Matthew C. Johnson
York University and Perimeter Institute



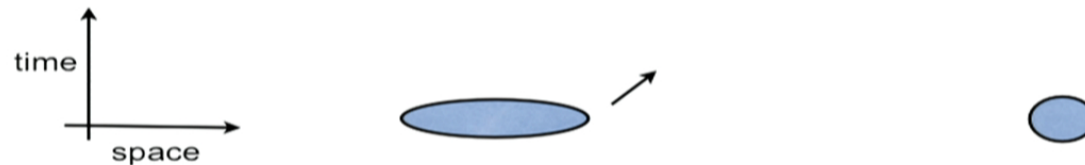


The Universe

Everything.

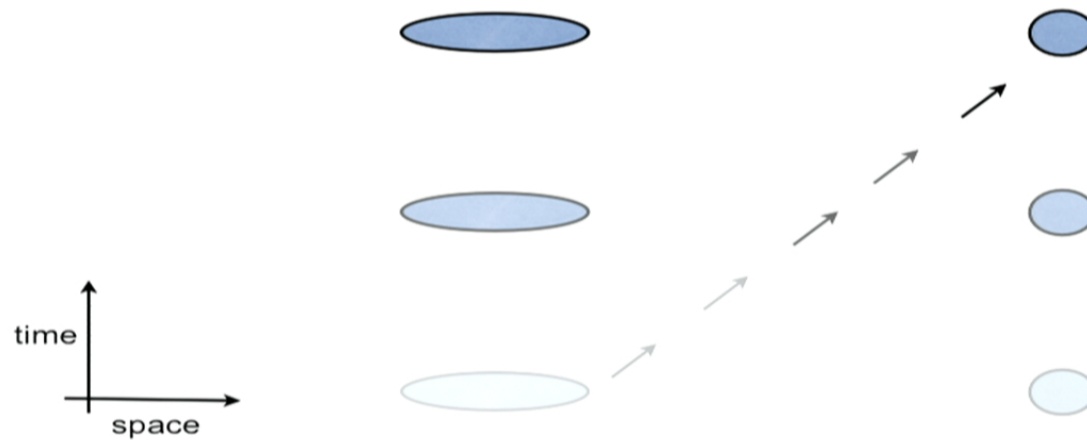
Far Out = Far Back in Time

Light has a constant speed.
Looking out is looking back in time.



Far Out = Far Back in Time

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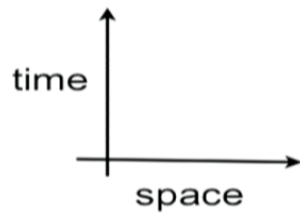


Cosmic Microwave Background (CMB) radiation

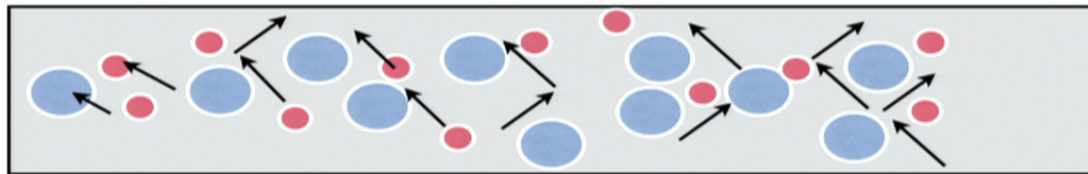
- The radiation released when neutral atoms were formed.

Cosmic Microwave Background (CMB) radiation

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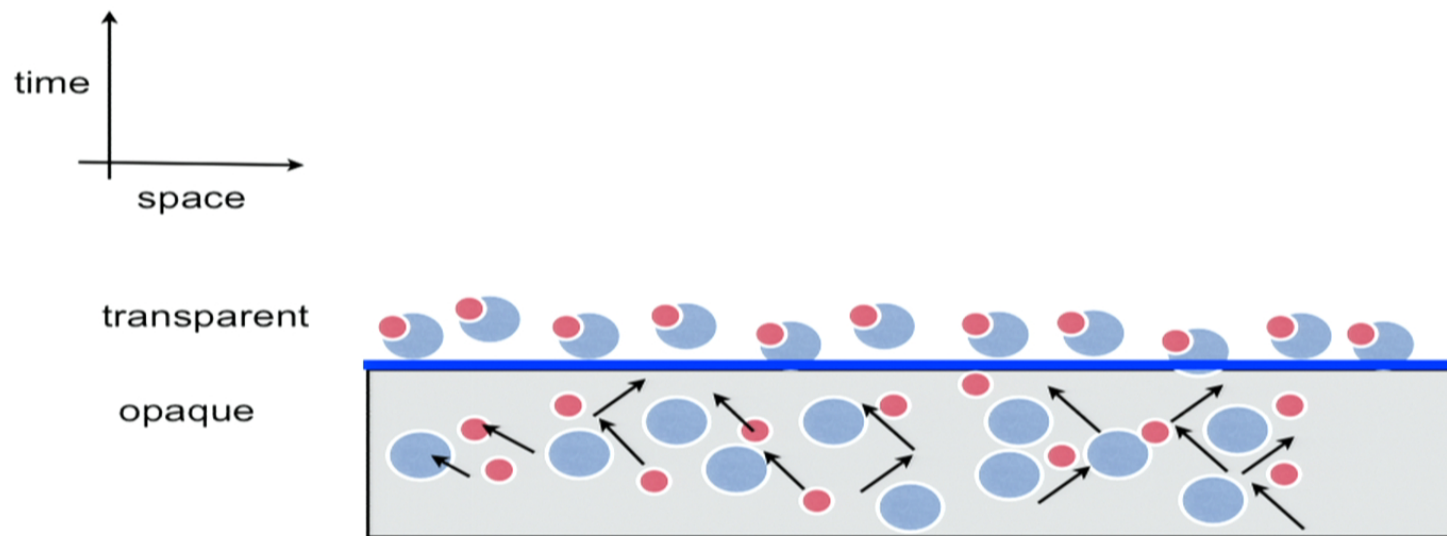


opaque



Cosmic Microwave Background (CMB) radiation

- The radiation released when neutral atoms were formed.



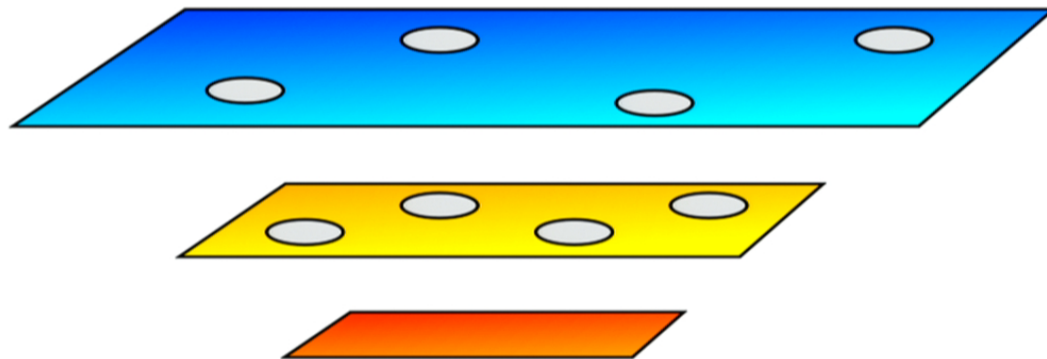
What is our best model?

- “Our whole universe was in a hot dense state, then nearly 14 billion years ago expansion started.....” (The big bang theory)



What is our best model?

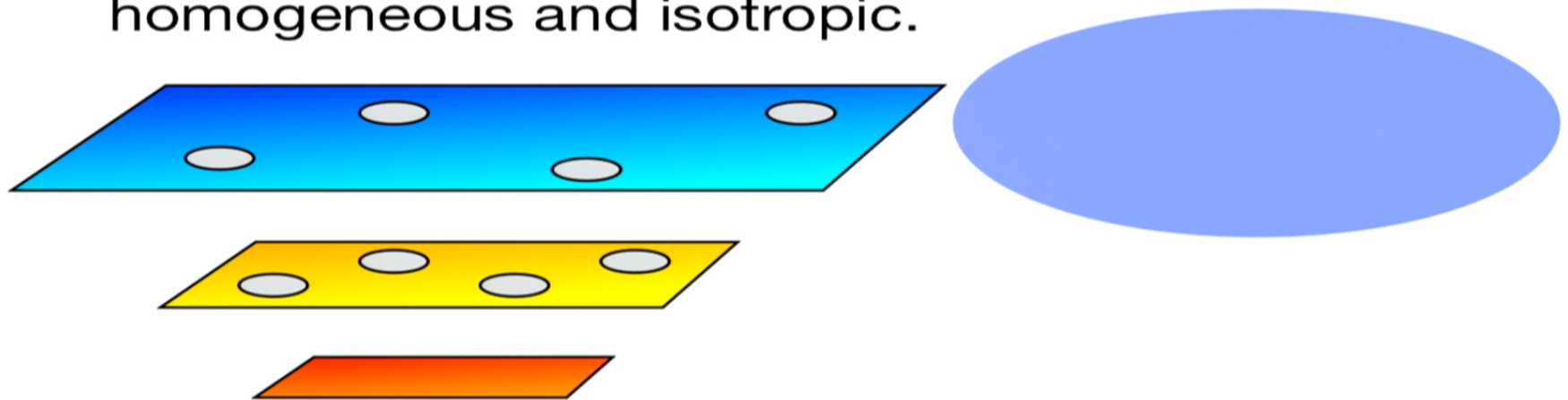
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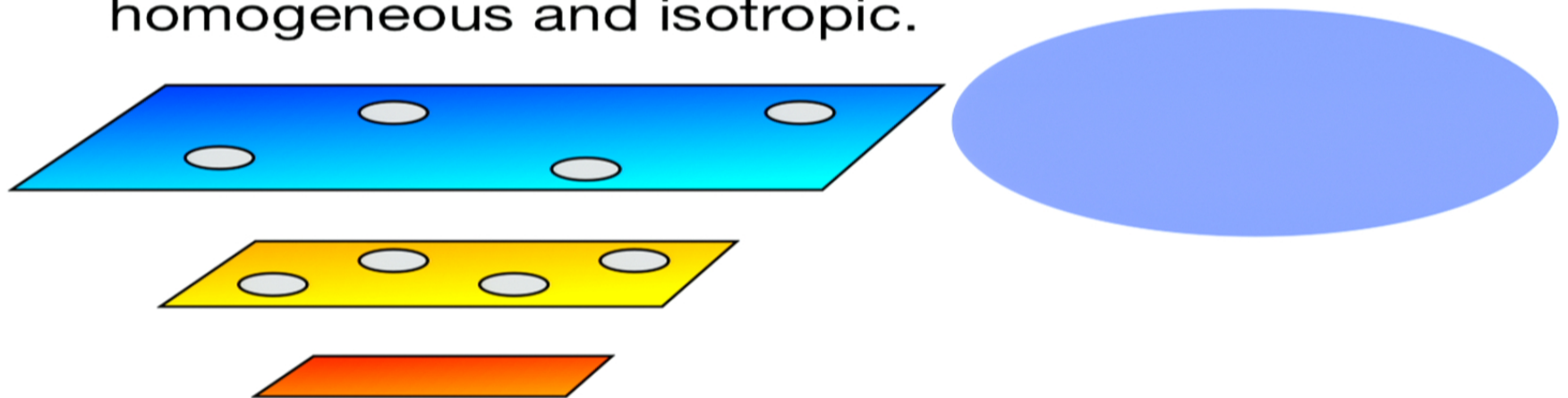
The Observable Universe is nearly homogeneous and isotropic.



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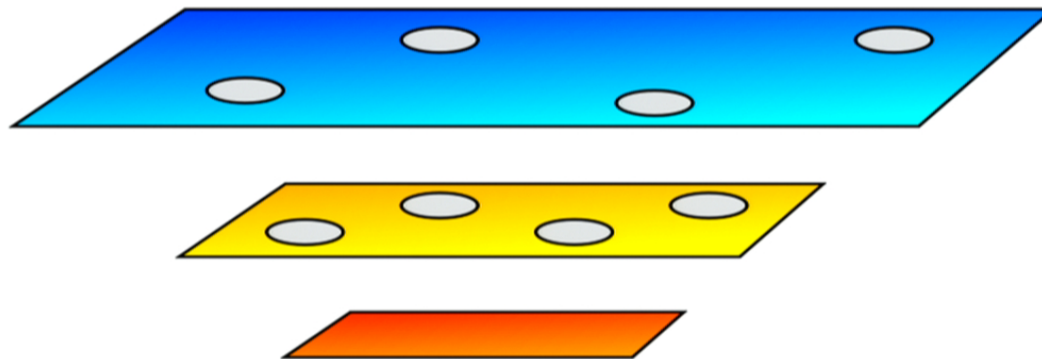
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What is our best model?

- “Our whole universe was in a hot dense state, then nearly 14 billion years ago expansion started.....” (The big bang theory)

The Observable Universe is statistically homogeneous and isotropic.



$1/1000000$

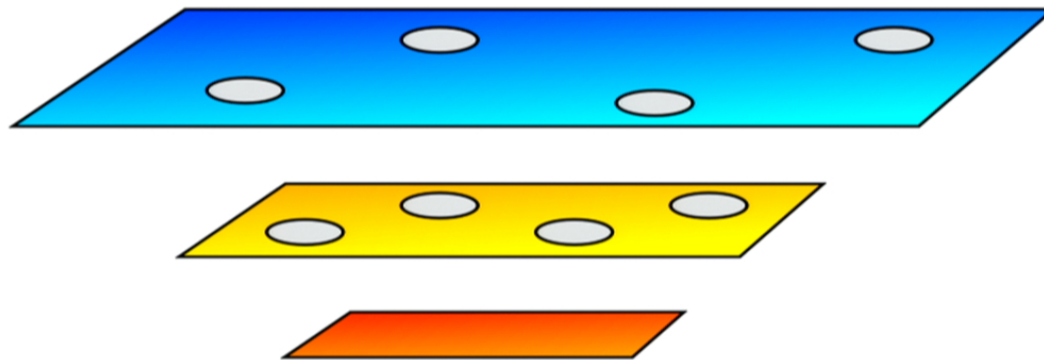
Nobel Prize 2006!

John Mather and George Smoot

What is our best model?

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The expansion of the Universe is accelerating!

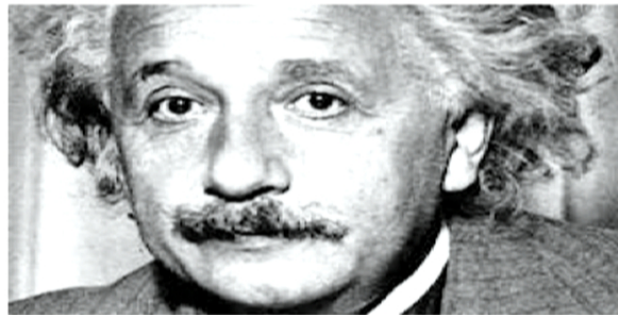


Nobel Prize 2011!

Adam Riess, Saul Perlmutter, Brian Schmidt

What is our best model?

- “Our whole universe was in a hot dense state, then nearly 14 billion years ago expansion started.....” (The big bang theory)
- The properties of expansion in a homogeneous and isotropic universe are determined by what the Universe contains.



General Relativity

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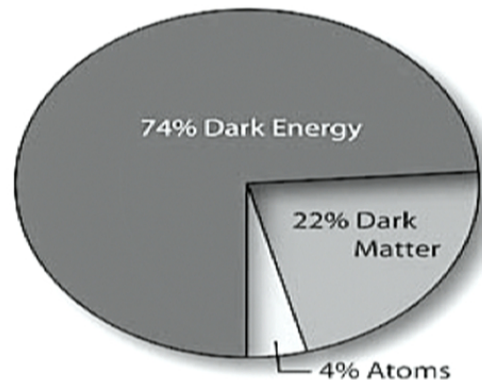
$$R_{\mu\nu} - \frac{1}{2}R g_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

Spacetime
curvature

Matter & energy

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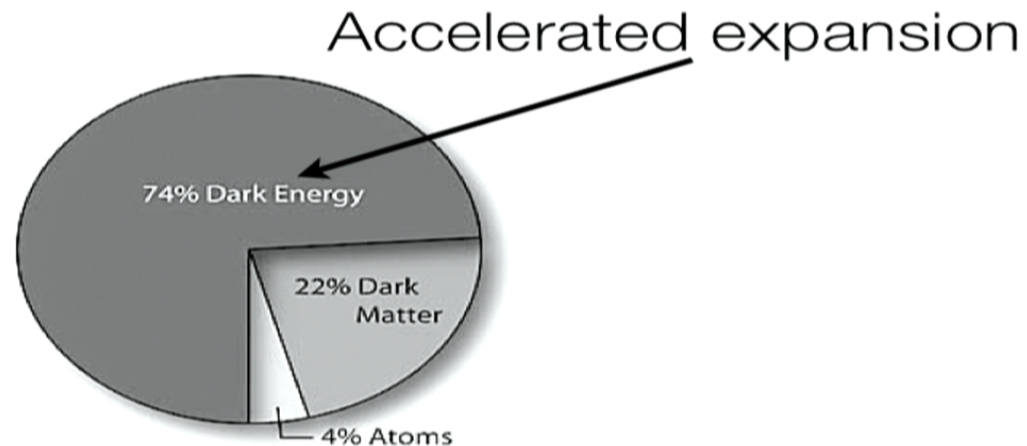
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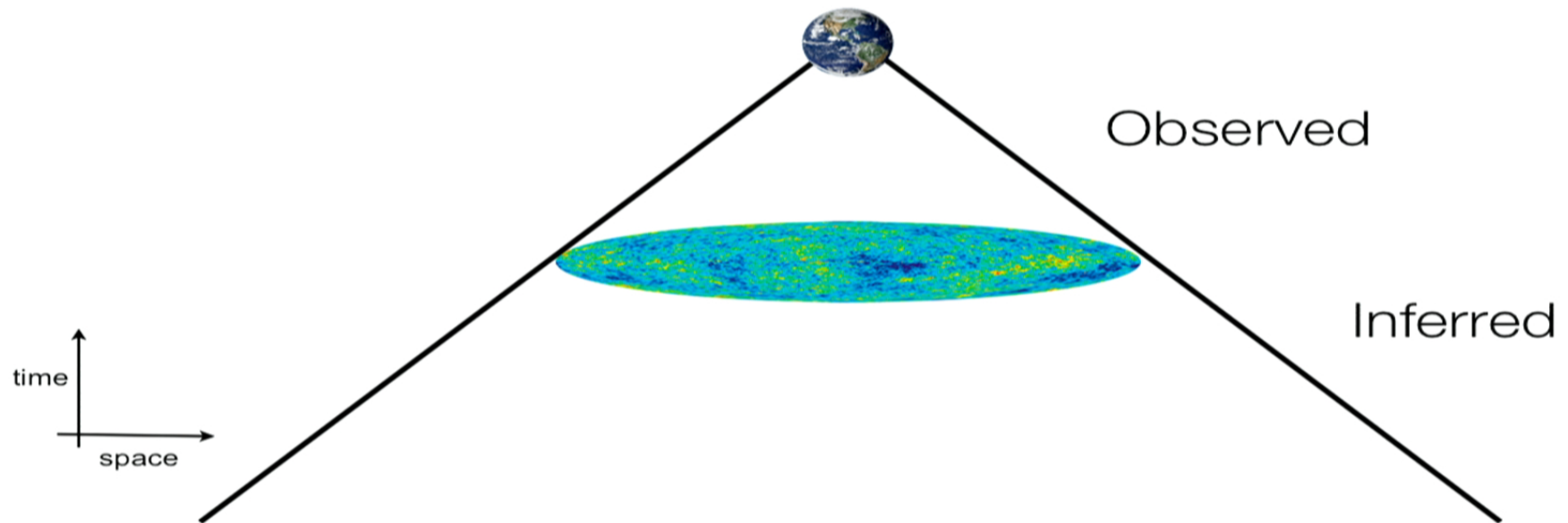
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Λ CDM

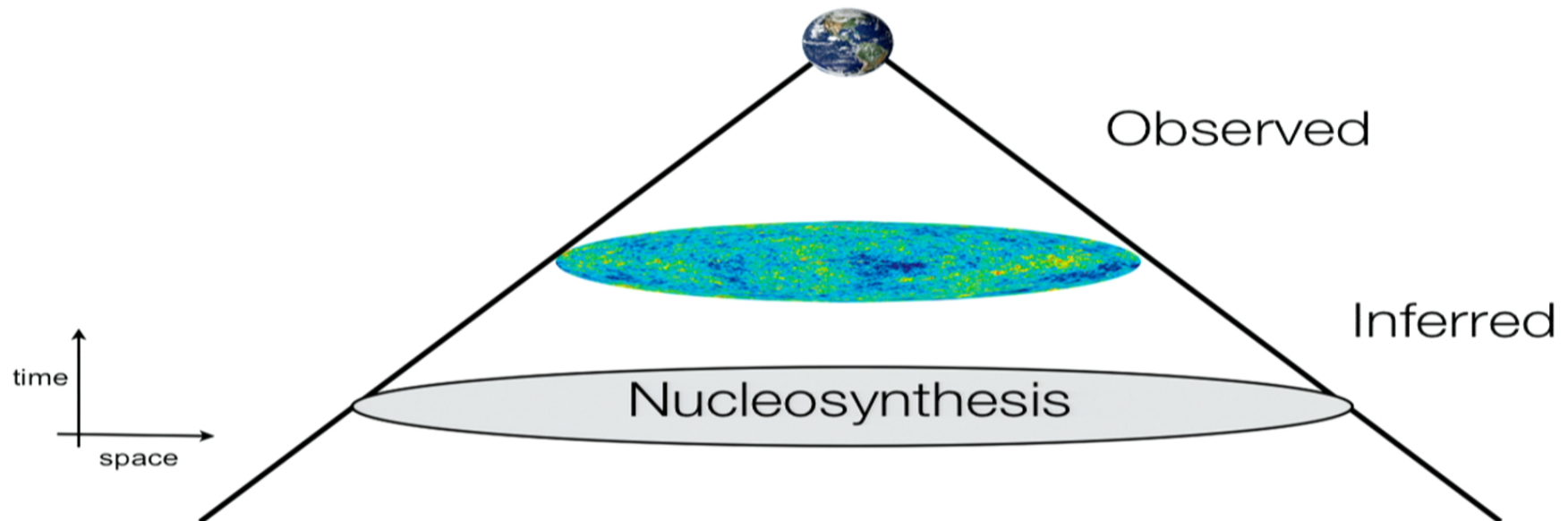
The Inferred Universe

Everything that we think exists based on theoretical models.



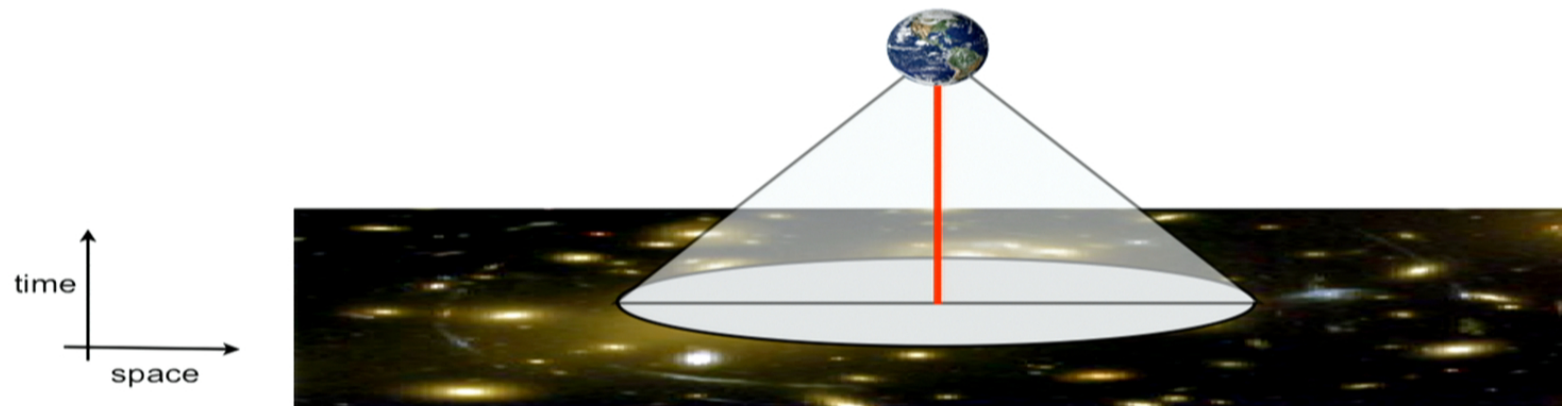
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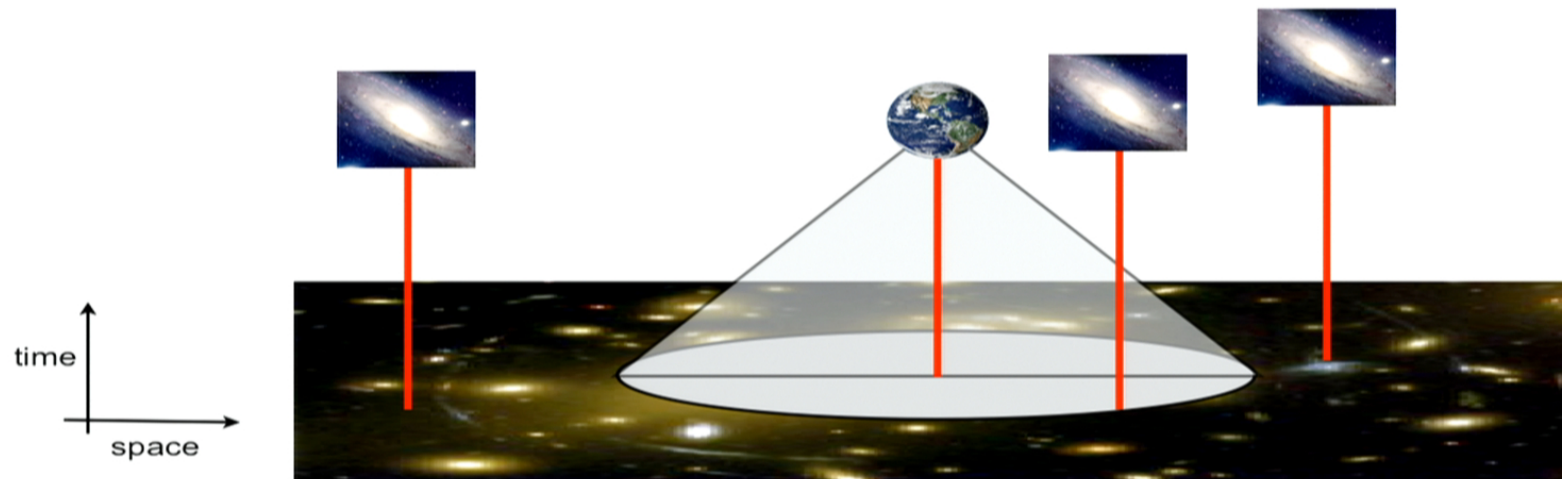
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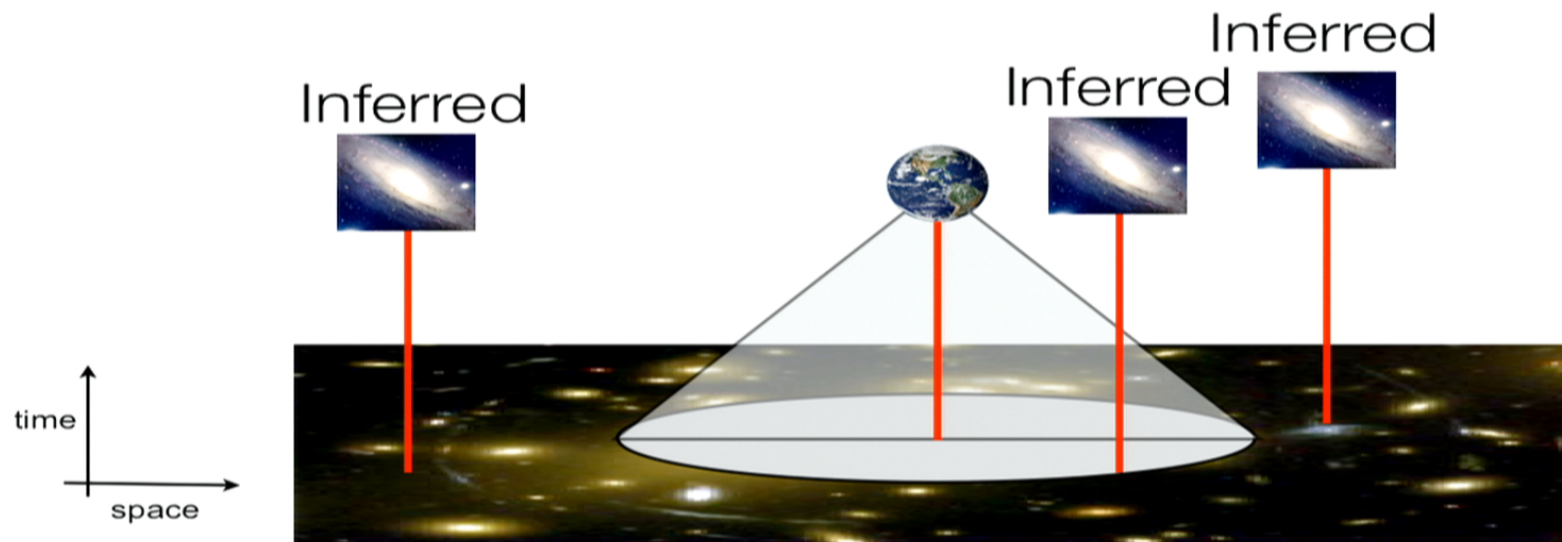
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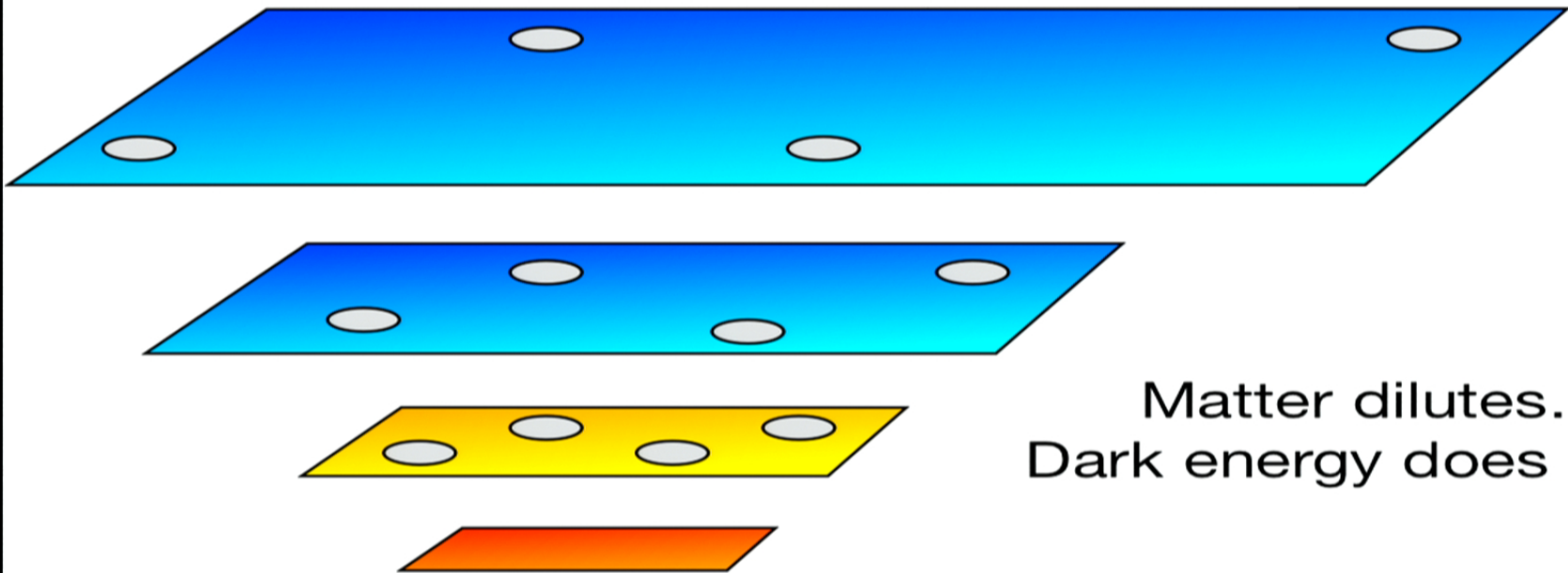


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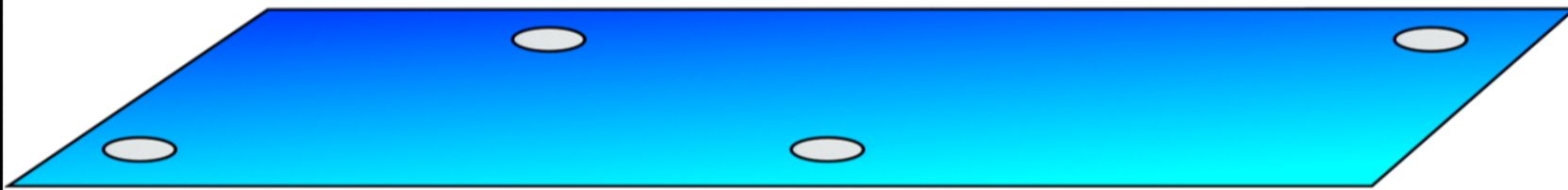


What is dark energy?

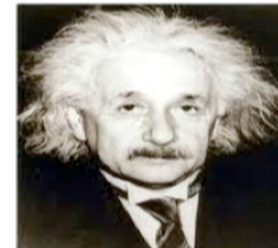


Matter dilutes.
Dark energy does not!

What is dark energy?



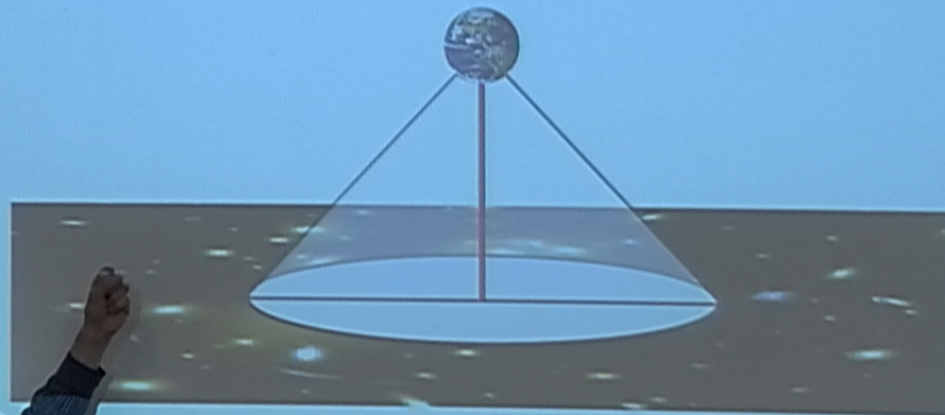
Cosmological Constant



The Inferred Universe

Everything that we think exists based on
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This is all we will ever see!



The Inferred Universe

Everything that we think exists based on what we observe and theoretical models.

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The Inferred Universe

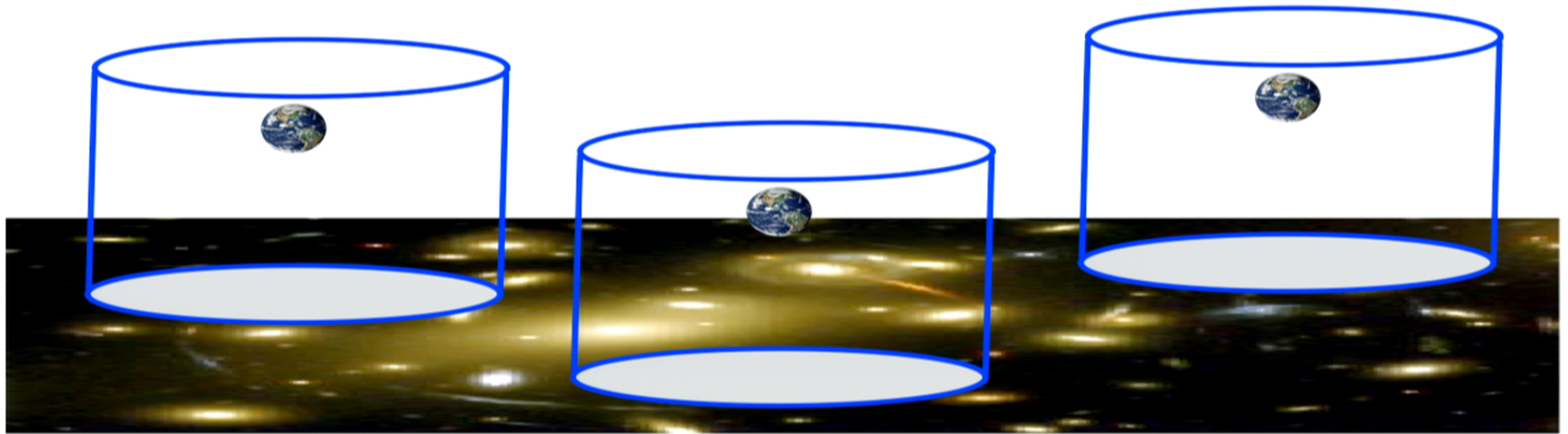
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The Multiverse

A set of Observable Universes.

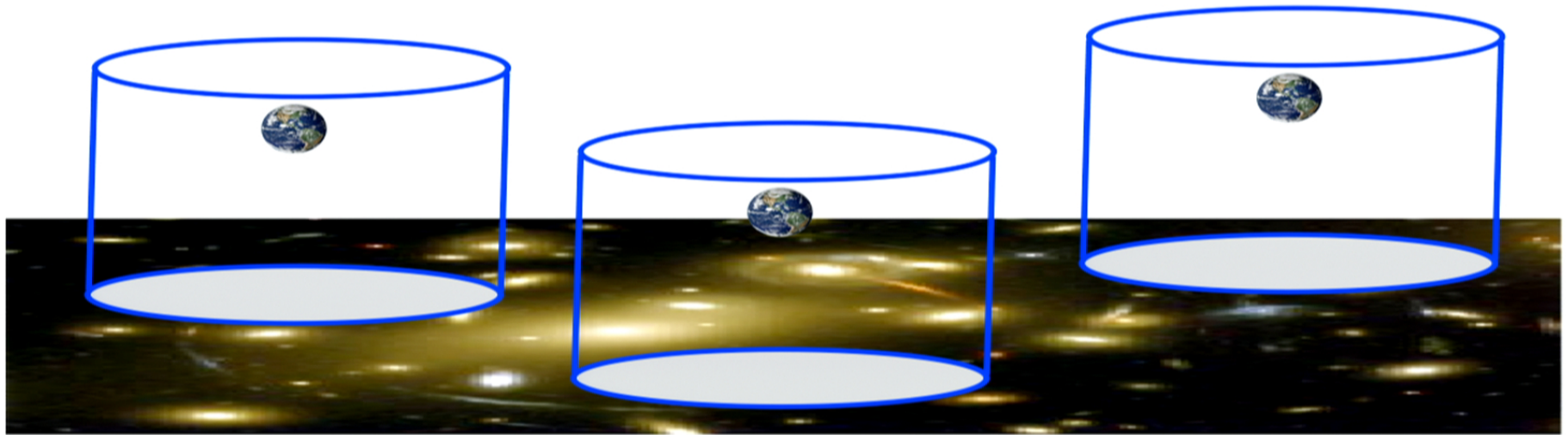


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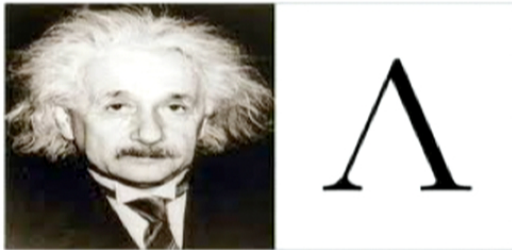
What can be different in the different Observable Universes?

Which one do we inhabit?



What is dark energy?

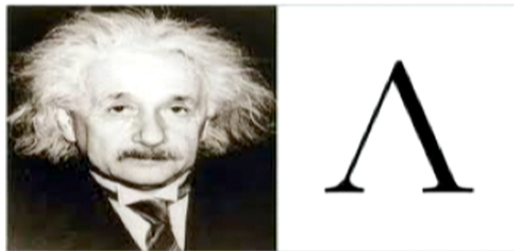
Cosmological Constant



Prediction from
Quantum Field Theory
(zero-point energy)

What is dark energy?

Cosmological Constant



Prediction from
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Nobel Prize 2011!

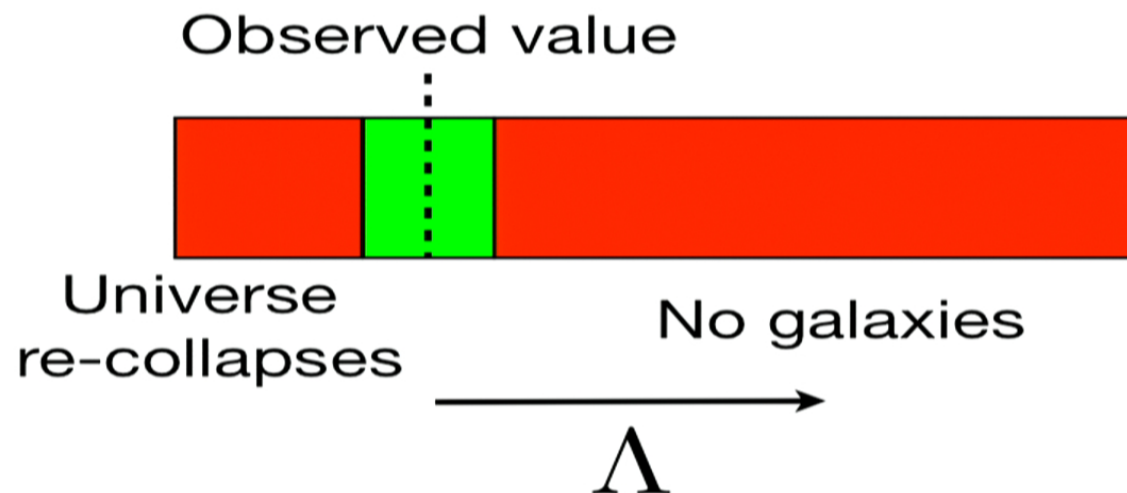
Adam Riess, Saul Perlmutter, Brian Schmidt

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Weinberg's Prediction of Λ



What if the cosmological constant varies between Observable Universes?



String Theory



(Graviton: particle associated with gravity)

String Theory



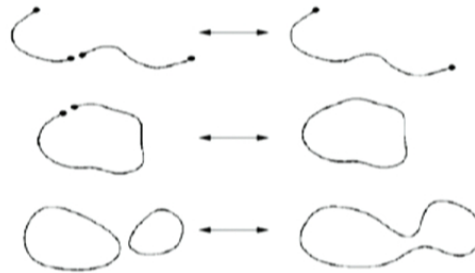
(Graviton: particle associated with gravity)



The theory of gravitons does not work!
(not a good quantum theory of gravity)

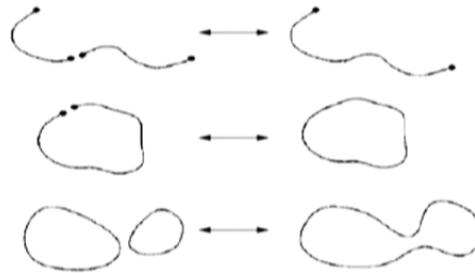


String Theory

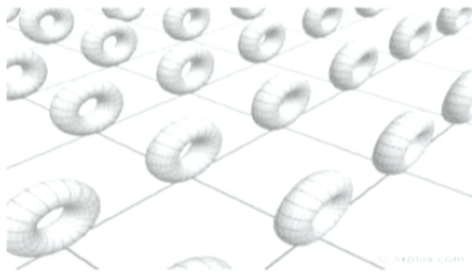


- String theory: A good theory of quantum gravity!
- Unifies all forces and fundamental particles!

String Theory

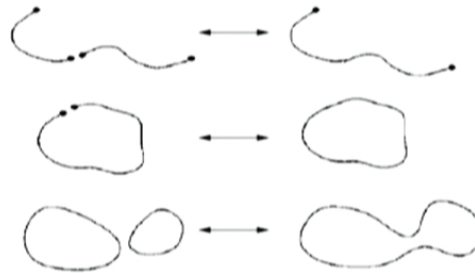


- String theory: A good theory of quantum gravity!
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- This only works if there are 9 dimensions of space!

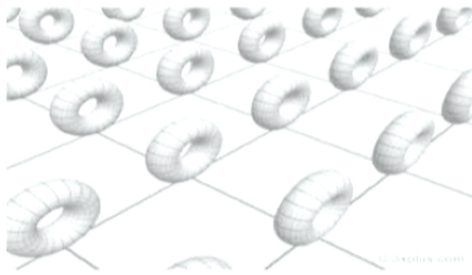


The solution: make the extra dimensions small!

String Theory



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The solution: make the extra dimensions small!

String Theory

- To keep the extra dimensions small, need to add energy.



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Λ : energy stored in the extra dimensions.

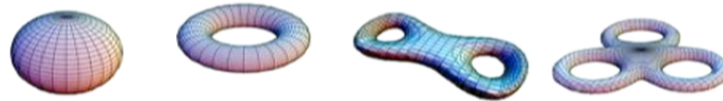
String Theory

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Λ : energy stored in the extra dimensions.

- The extra dimensions can assume many configurations:



Many possible values of Λ !

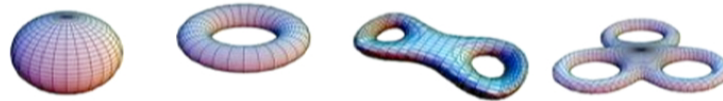
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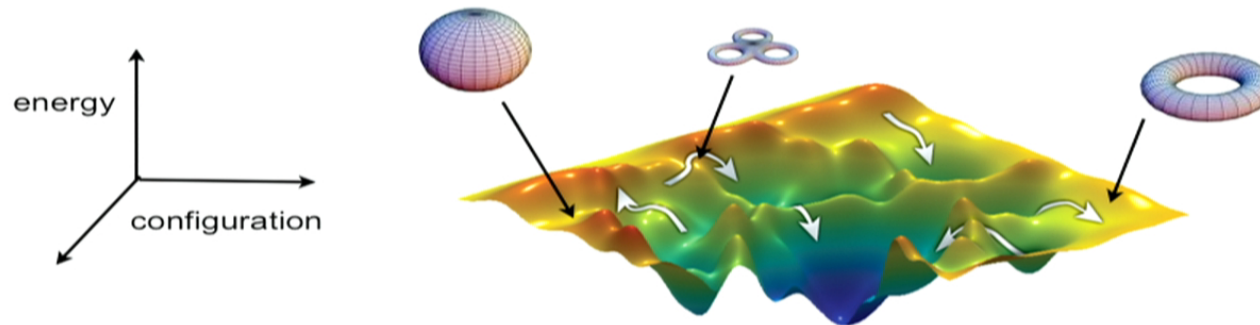
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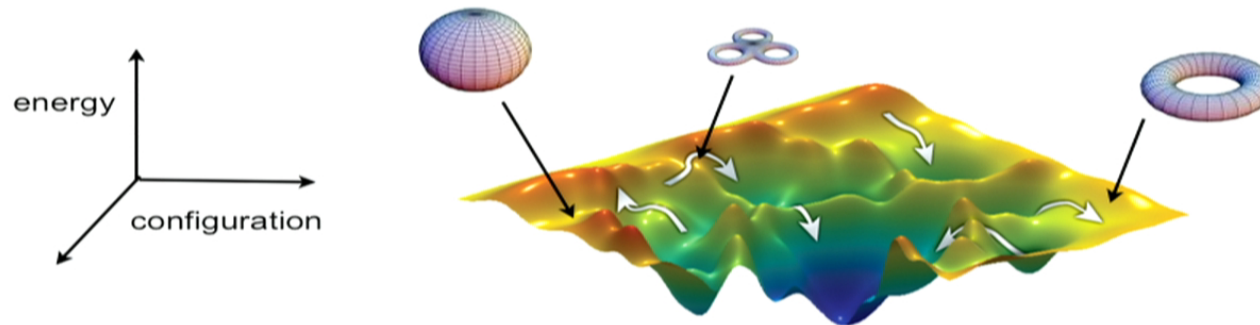
10^{500}

The String Theory Landscape



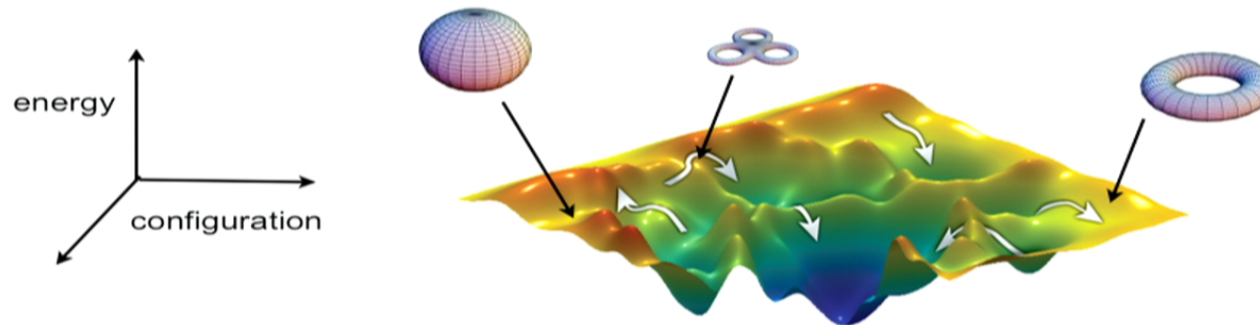
- Configurations can be deformed into one another.
- Can each of these configurations be physically realized?

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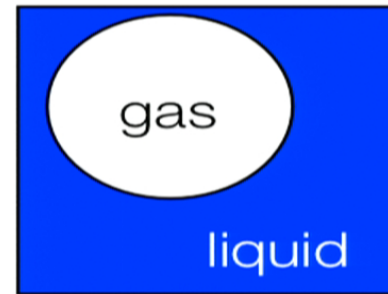
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Many possible values of Λ !

Cosmic Phase Transitions



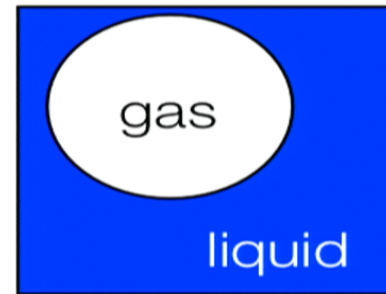
Liquid water goes to steam by the formation of bubbles.



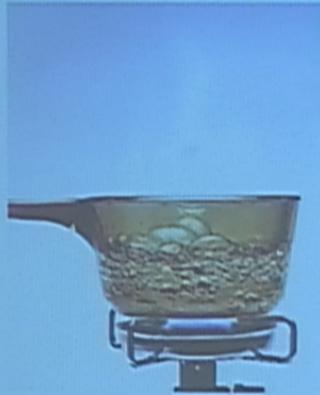
Cosmic Phase Transitions



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Cosmic Phase Transitions

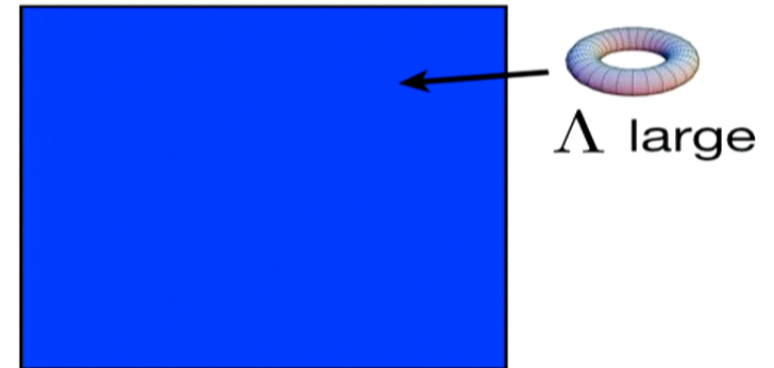
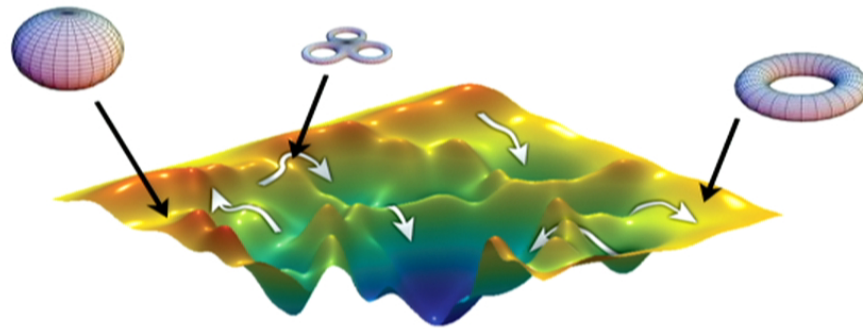


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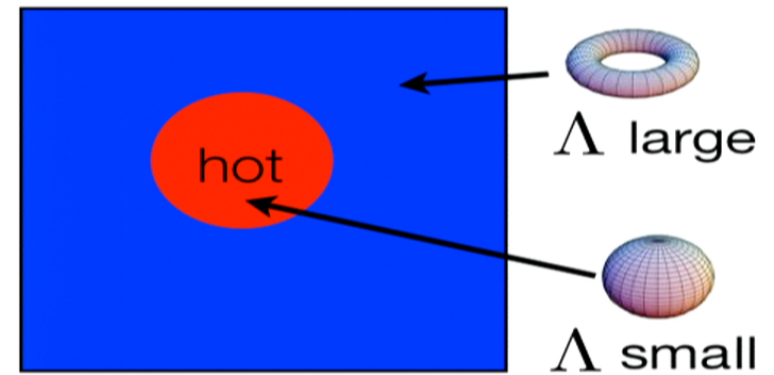
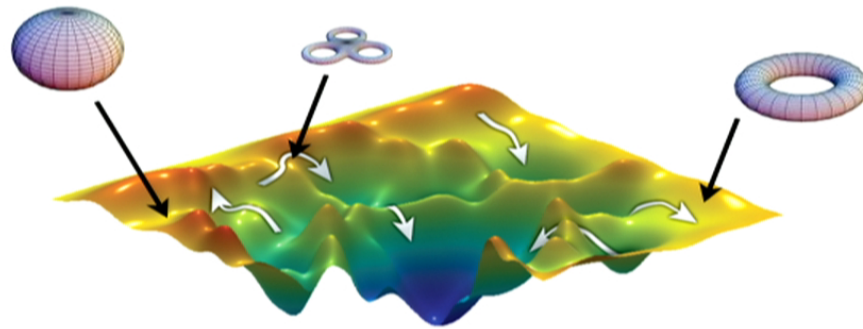
Cosmological phase transitions are possible.
Happen even at zero temperature!

Cosmic Phase Transitions



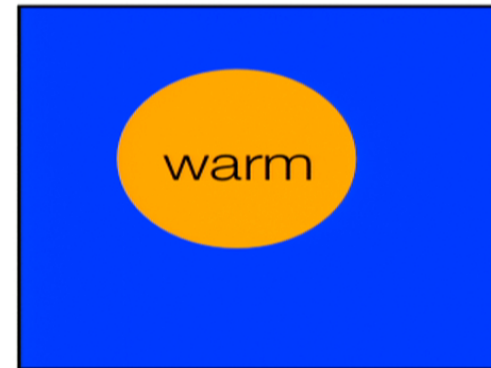
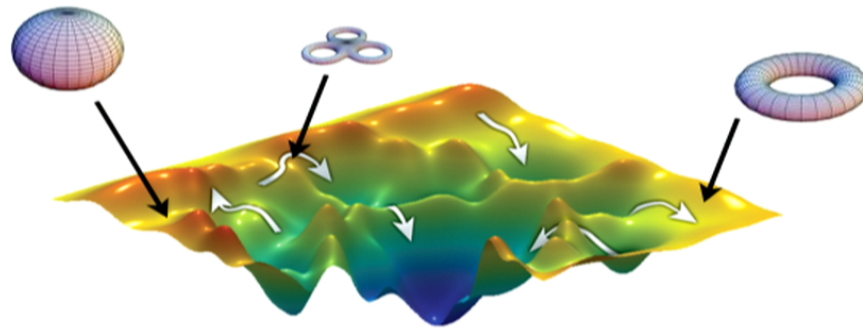
- The configuration of the extra dimensions can undergo phase transitions!
- This means Λ changes as well!

Cosmic Phase Transitions



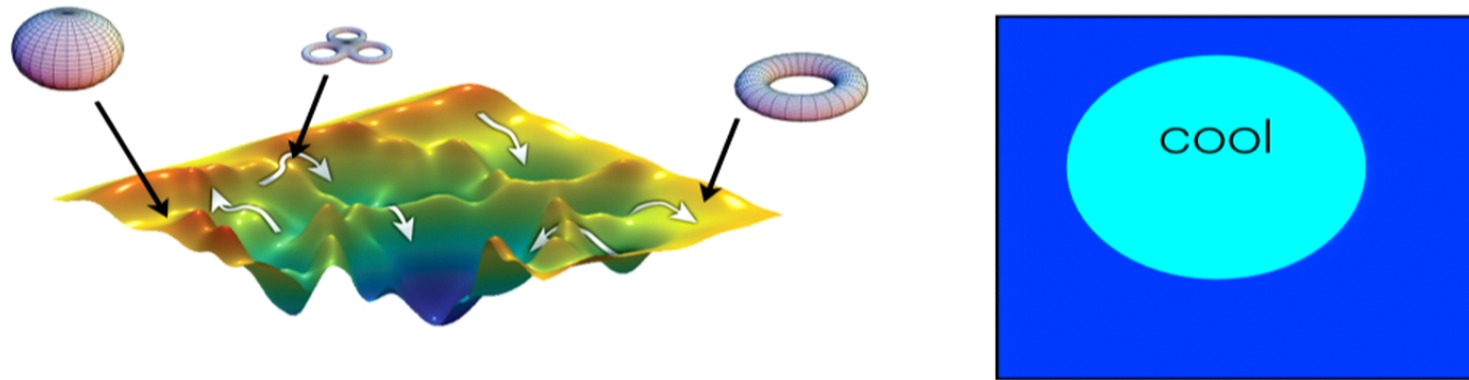
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Cosmic Phase Transitions



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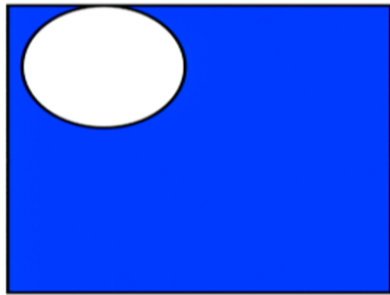
Cosmic Phase Transitions



- The configuration of the extra dimensions can undergo phase transitions!
- This means Λ changes as well!
- Our Observable Universe is contained in a bubble.

Cosmic Phase Transitions

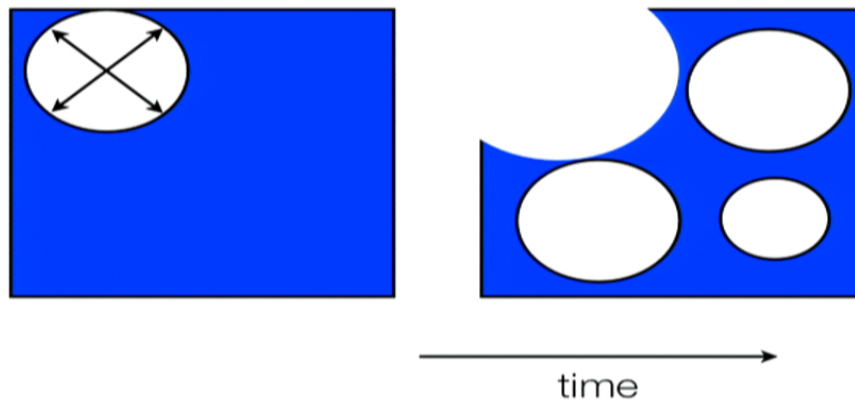
- In a static or decelerating universe:



time →

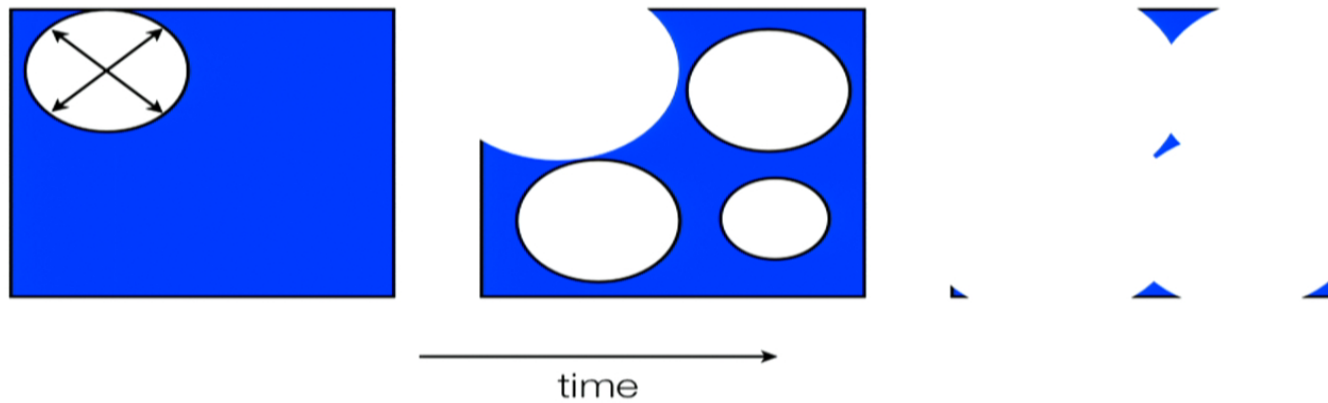
Cosmic Phase Transitions

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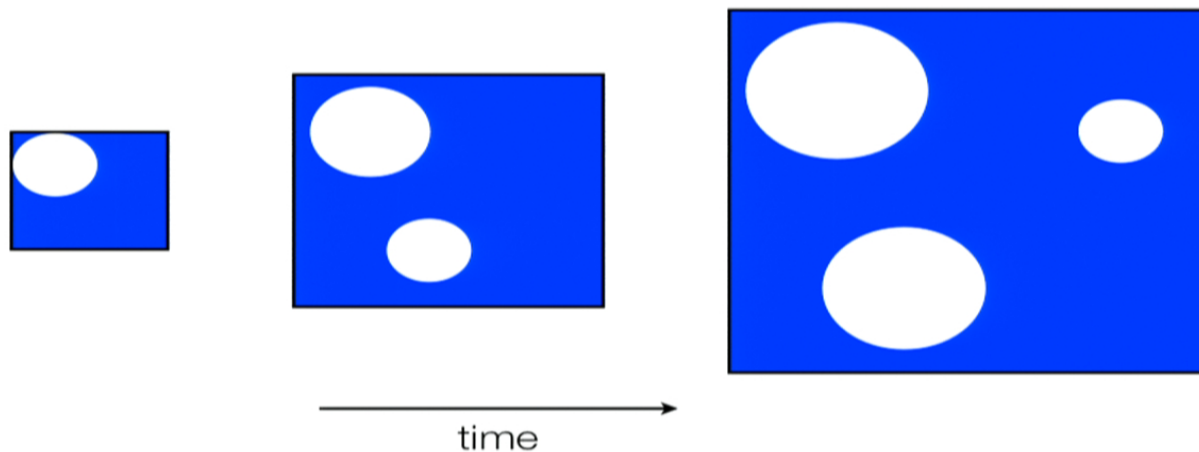
Cosmic Phase Transitions

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Cosmic Phase Transitions

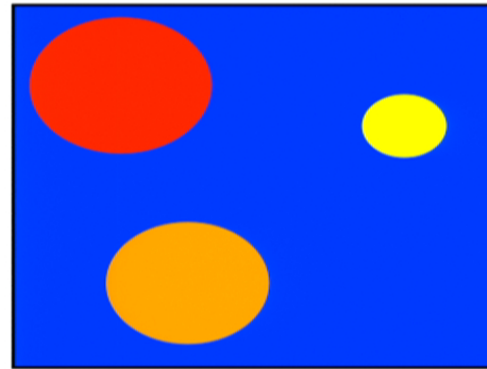
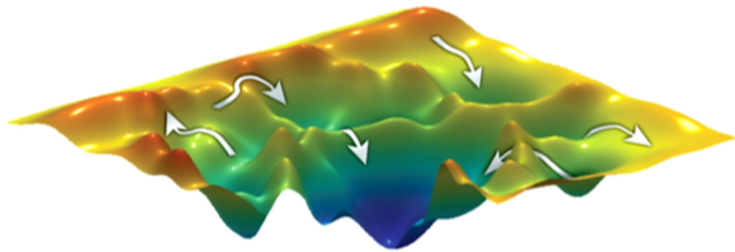
- In an accelerating universe:



Eternal Inflation

When the rate of bubble formation is lower than the rate of expansion, accelerated expansion doesn't end everywhere!

Eternal Inflation

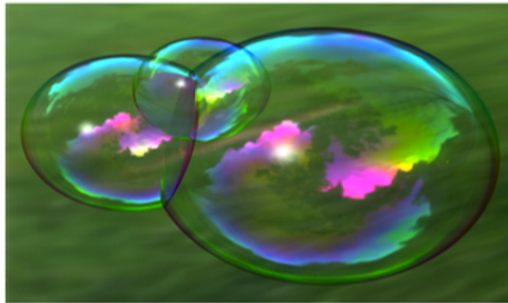


Many values of Λ are realized!

Observational Tests of Eternal Inflation

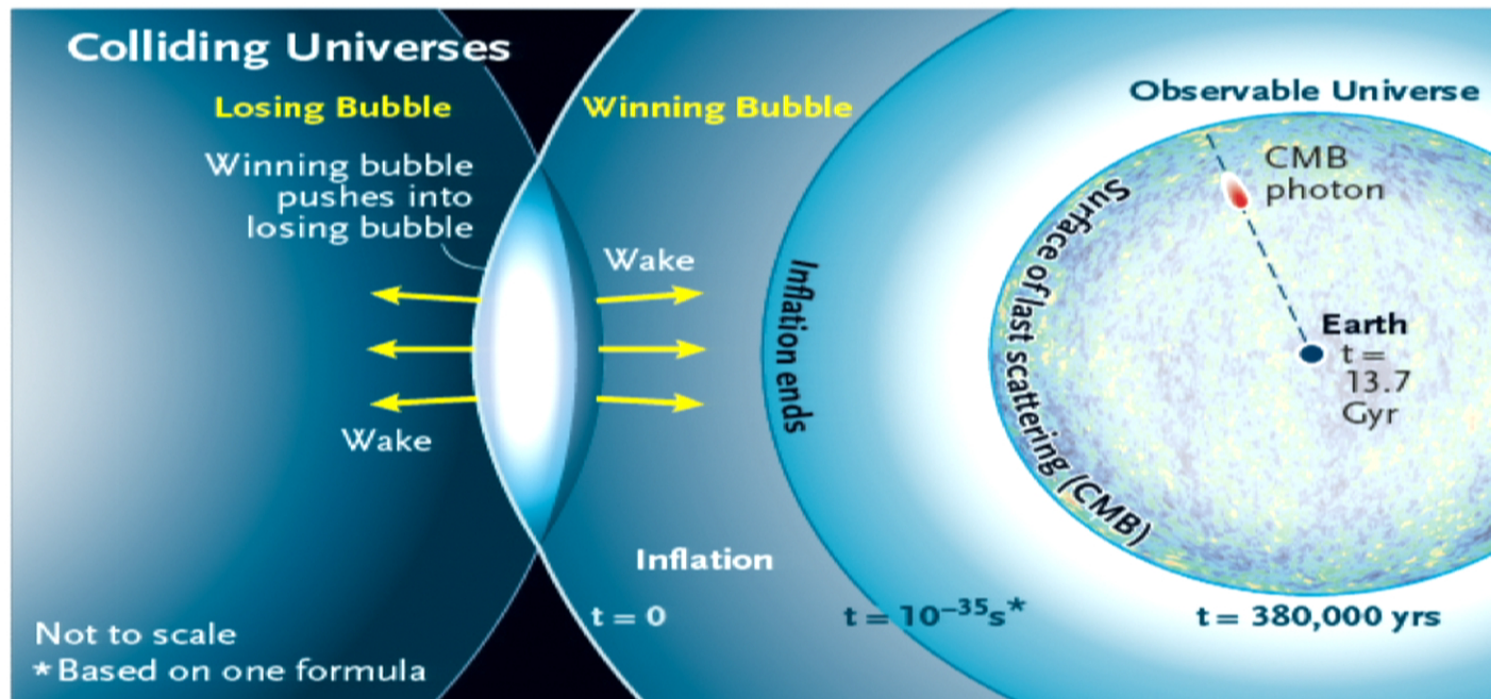
- But is eternal inflation experimentally verifiable?

Our bubble does not evolve in complete isolation....



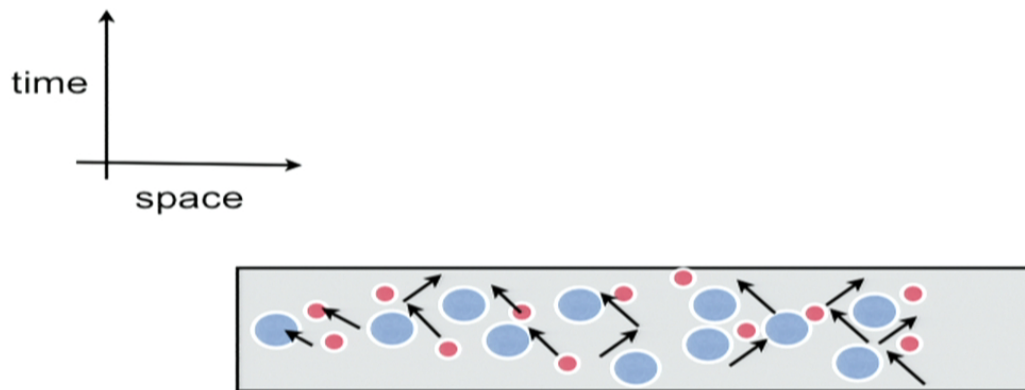
The collision of our bubble with others provides an observational test of the multiverse.

Observational Tests of Eternal Inflation



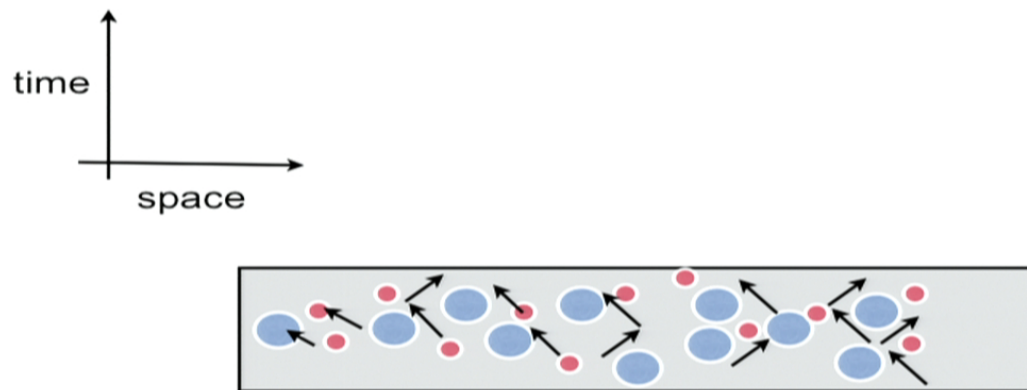
Cosmic Microwave Background (CMB) radiation

- Temperature anisotropies encode density perturbations.



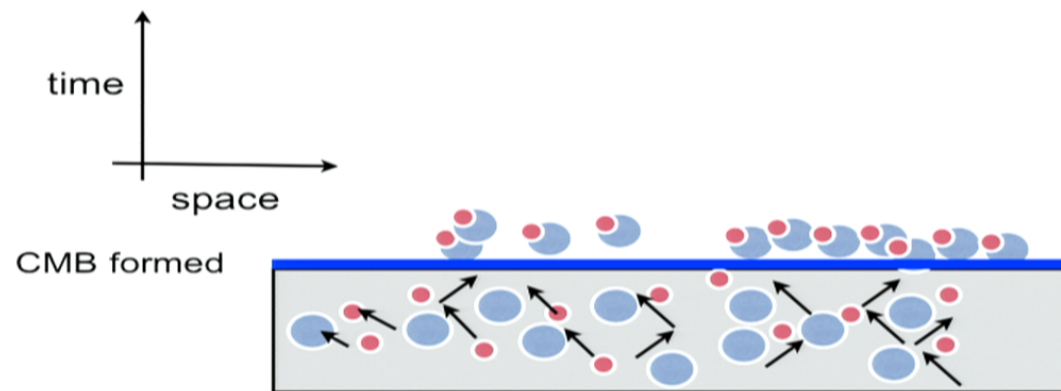
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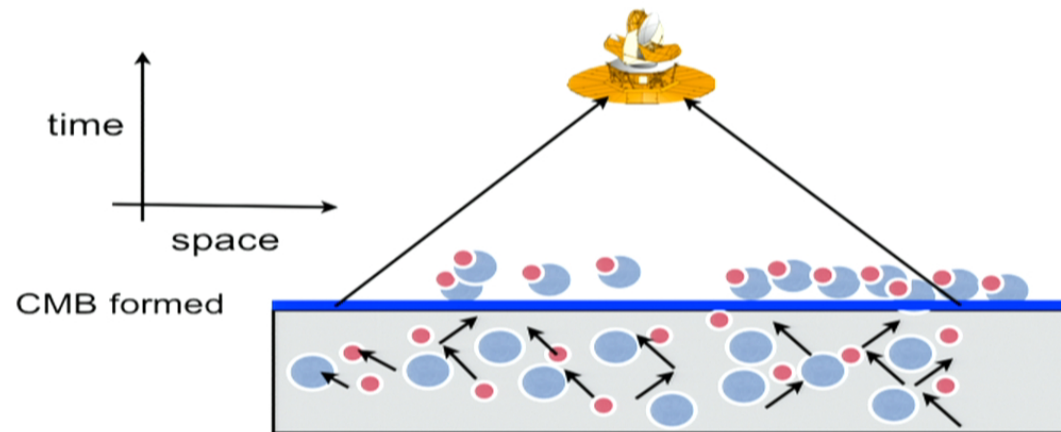
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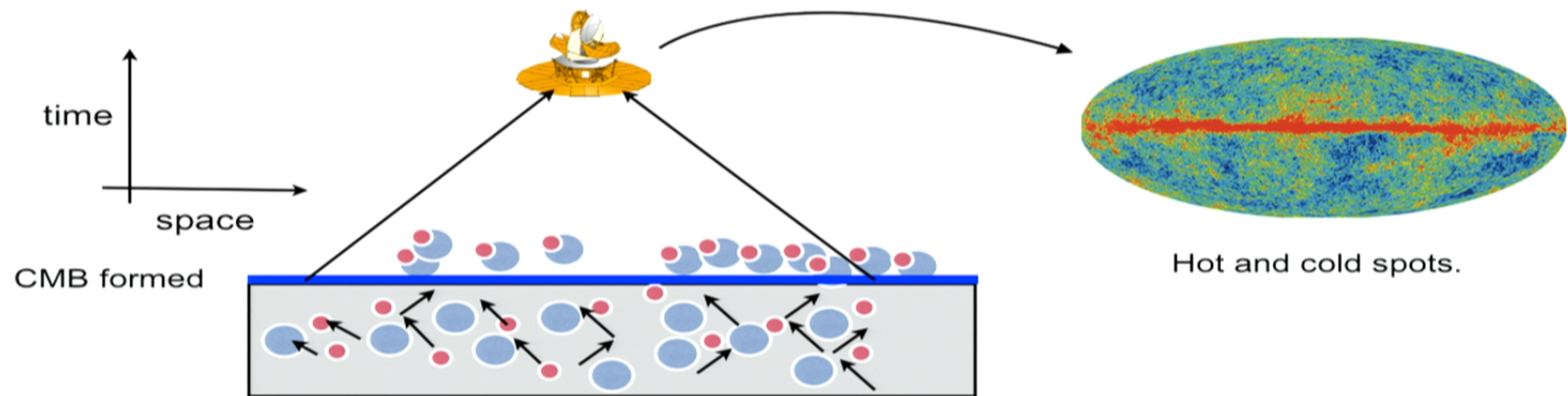
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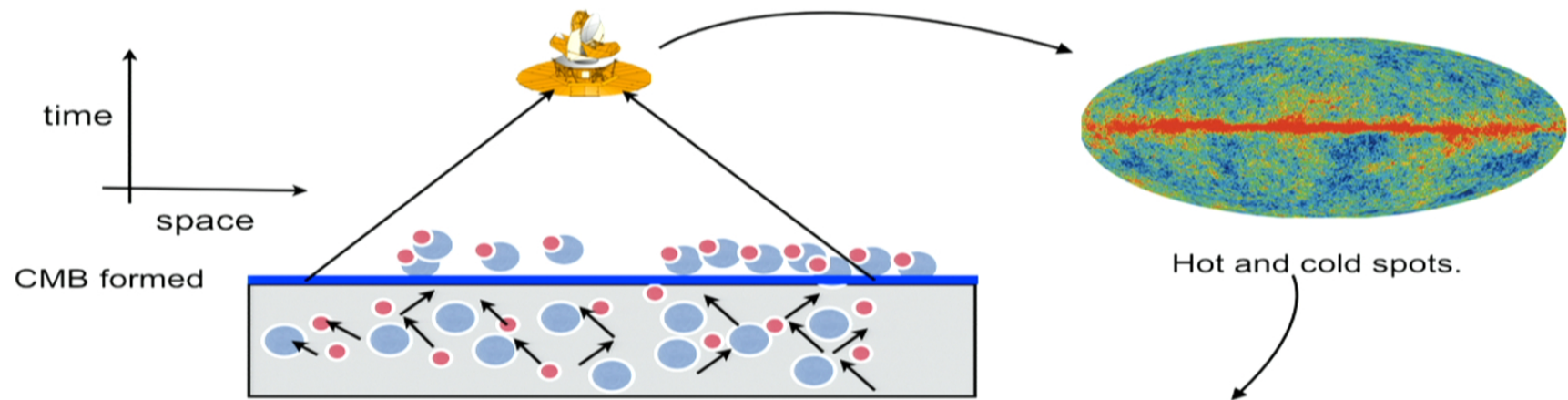
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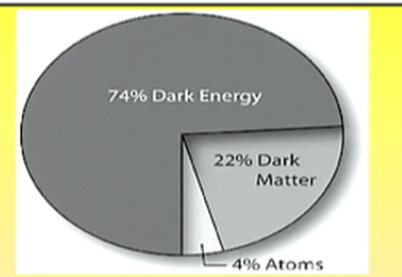


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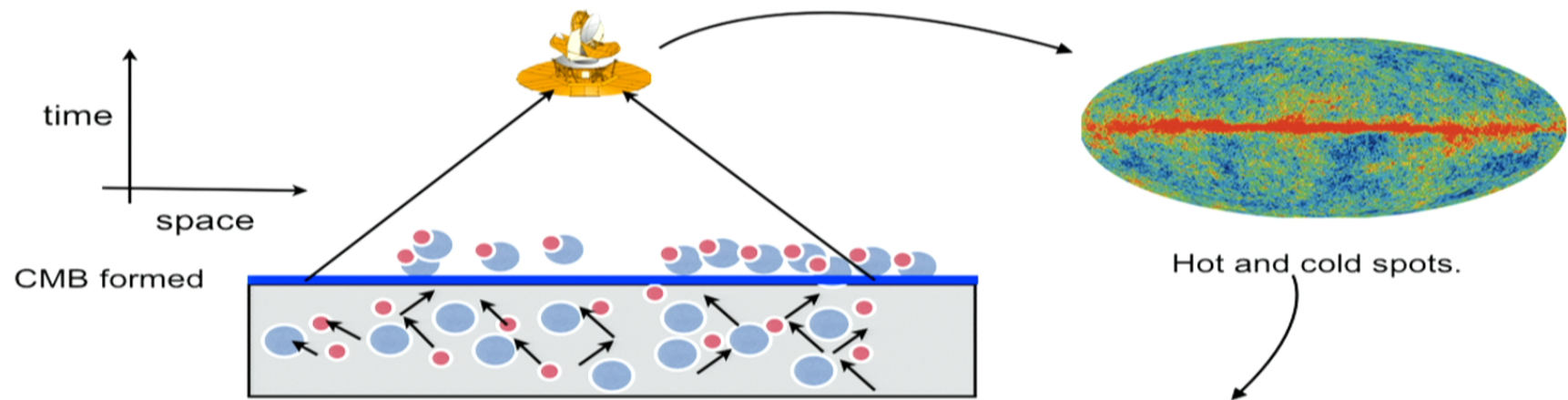


- We live in a nearly flat, accelerating universe, composed almost entirely of dark energy and dark matter.

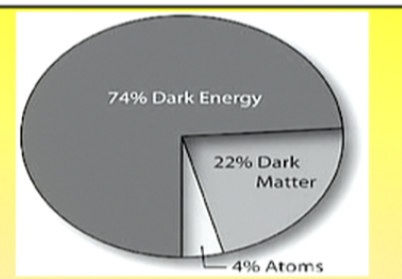


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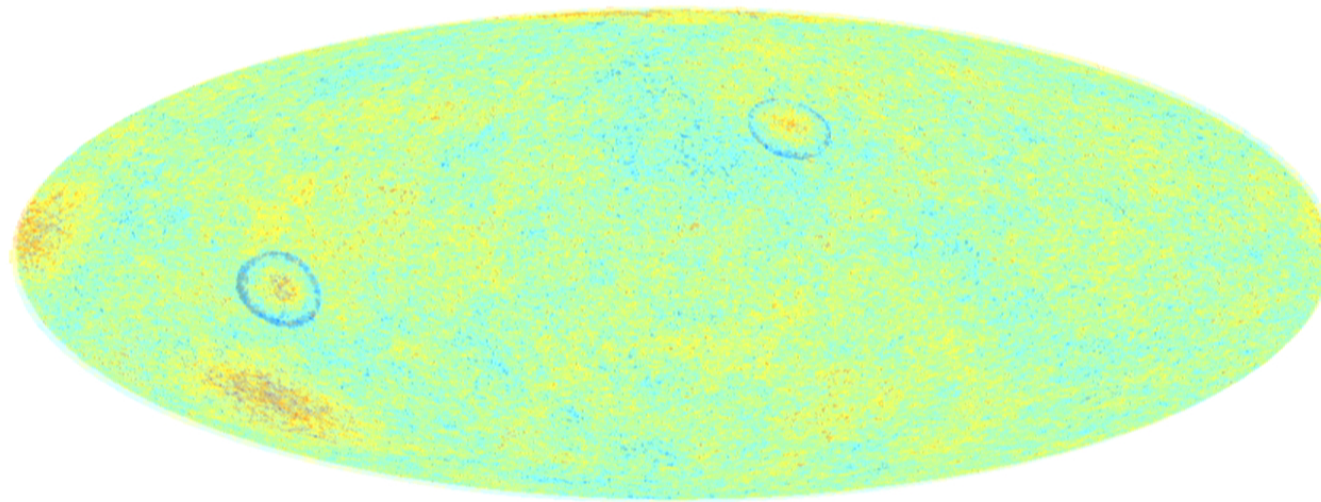


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Bubble Collisions

- Bubble collisions induce extra features in the CMB:

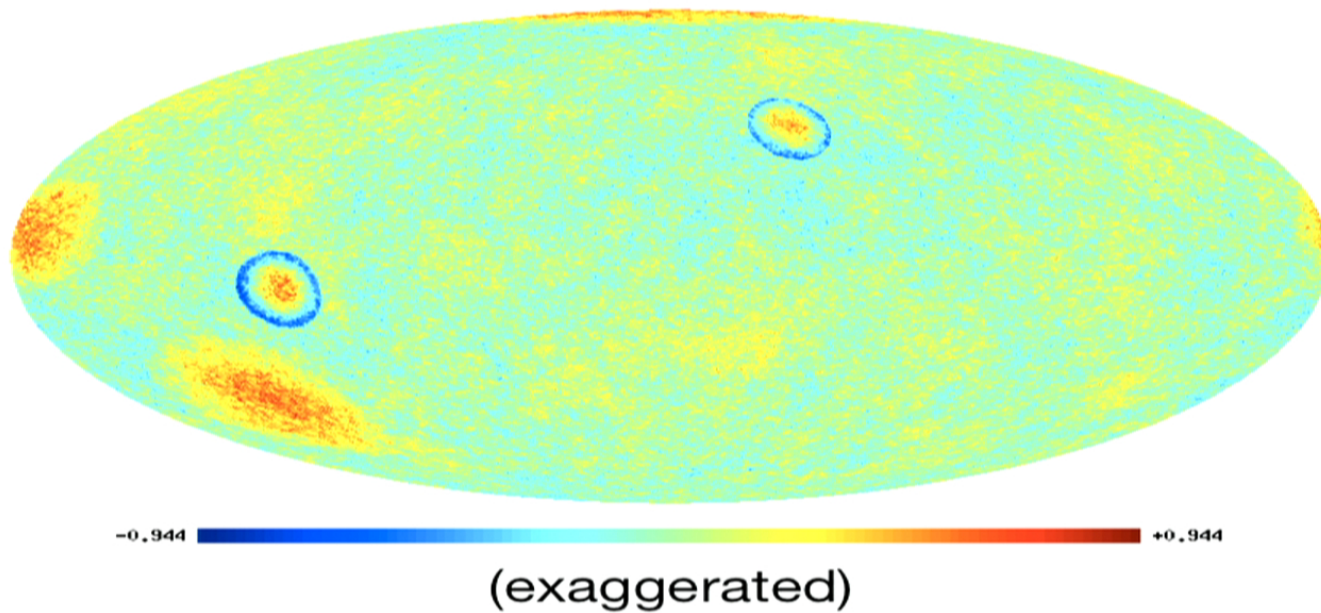


-0.944  +0.944

(exaggerated)

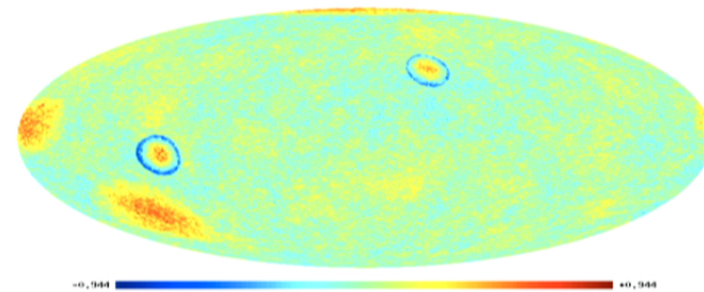
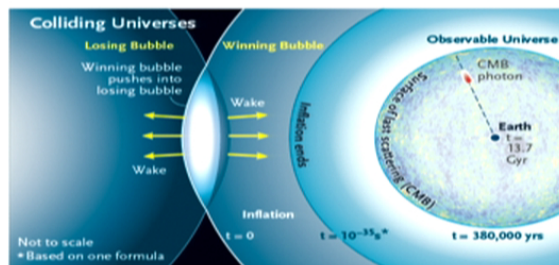
Bubble Collisions

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Bubble collisions model

- The model:



$$\bar{N}_s$$

expected number of collisions

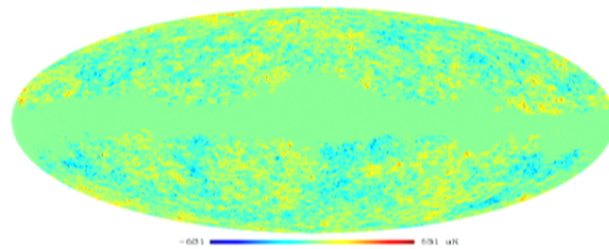
$$\mathbf{m}$$

parameters characterizing each collision

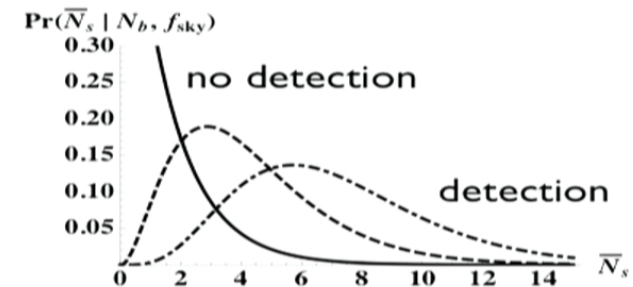
$$\text{Pr}(N_s, \mathbf{m})$$

How many of each type do I expect to find?

Searching for collisions



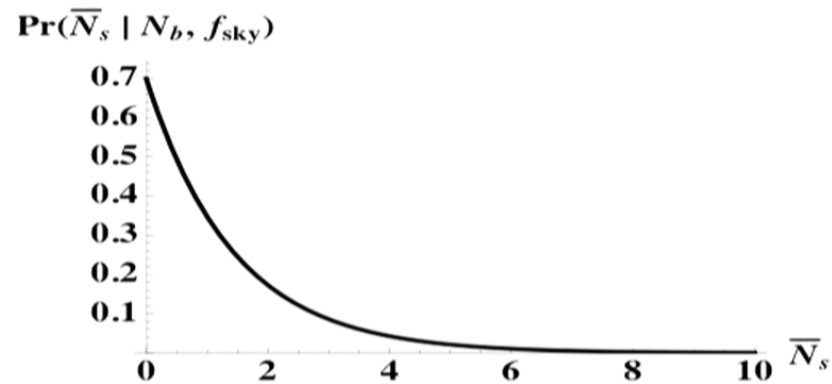
$$\Pr(\bar{N}_s | \mathbf{d})$$



- To calculate this, need to test for:
 - Arbitrary number of collisions
 - Arbitrary position on the sky
 - Arbitrary amplitude, shape, and size (lying within prior $\Pr(N_s, \mathbf{m})$)

Observational Tests of Eternal Inflation

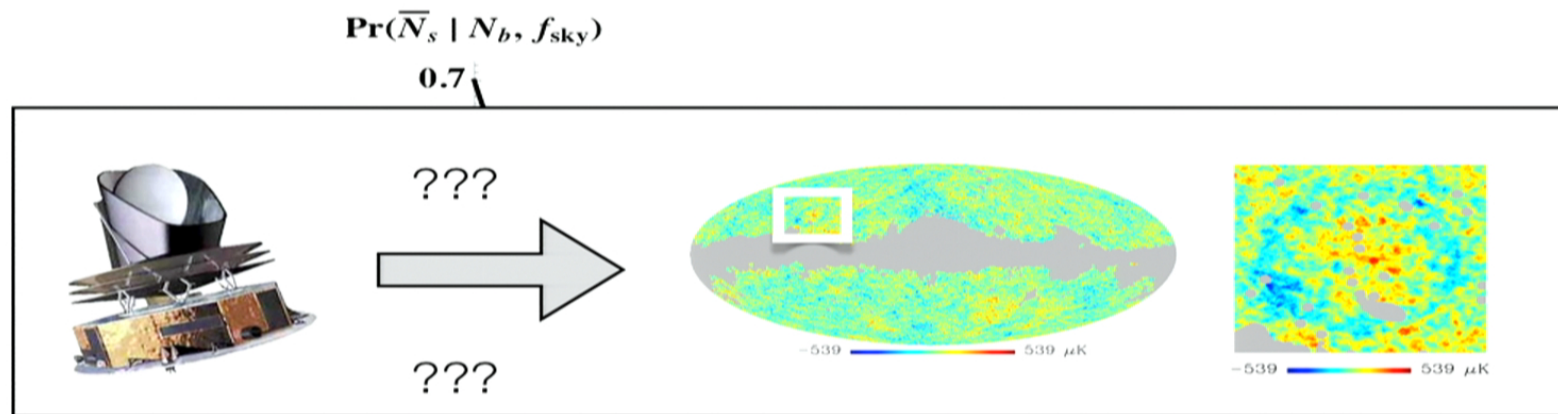
- Results using data from the WMAP satellite



$$\bar{N}_s < 1.6 \text{ at } 68\% \text{ CL}$$

Observational Tests of Eternal Inflation

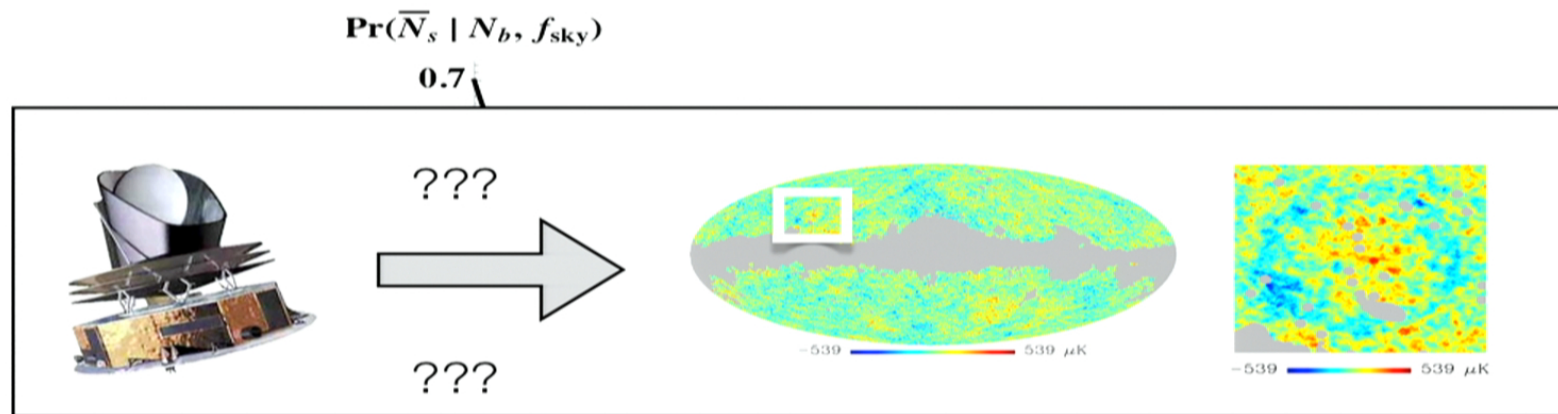
- Results using data from the WMAP satellite



$$\bar{N}_s < 1.6 \text{ at } 68\% \text{ CL}$$

Observational Tests of Eternal Inflation

- Results using data from the WMAP satellite



$$\bar{N}_s < 1.6 \text{ at } 68\% \text{ CL}$$

Conclusions

We may inhabit a Multiverse.
We may know soon!

