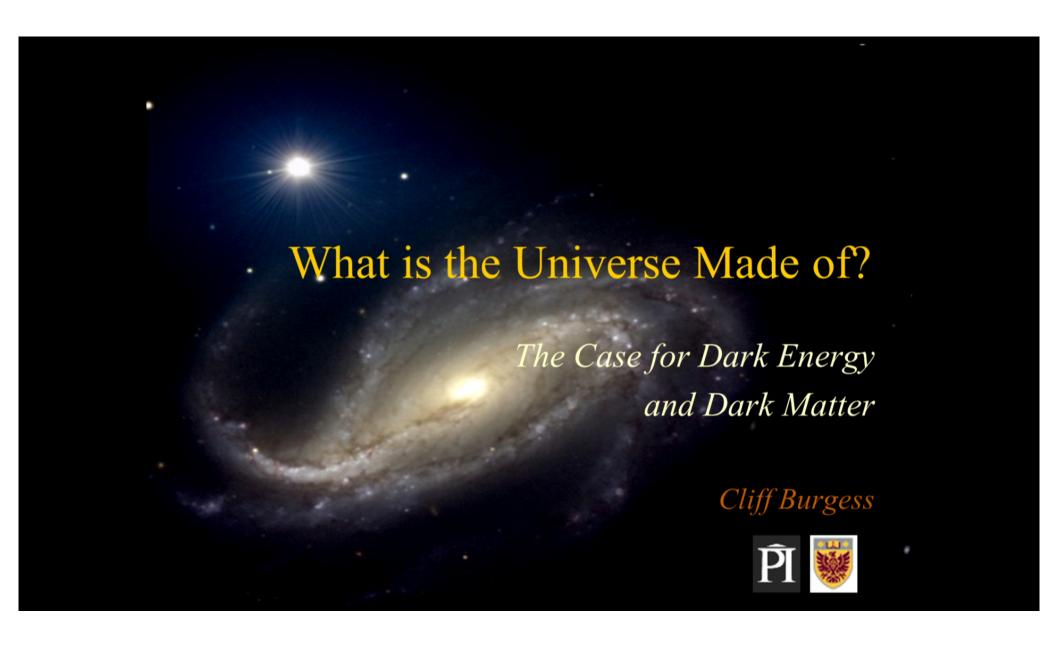
Title: What is the Universe Made Of? The Case for Dark Matter and Dark Energy

Date: Jul 13, 2012 07:30 PM

URL: http://pirsa.org/12070004

Abstract: For the first time in human thought it is now
br>possible to observationally determine how much matter is in the Universe as a
br>whole. These observations strongly support the "Concordance Model― of Hot Big
br>Bang Cosmology, and reinforce earlier indications that ordinary matter (atoms,
br>nuclei and electrons) make up at present at most 4% of the total of the
br>Universal energy density. The big surprise was that the rest consists of *two*
br>kinds of unknown forms of matter: the so-called Dark Matter and Dark Energy.
br>This talk summarizes for non-specialists the various lines of evidence for
br>their existence, and some of the theoretical ideas which have been proposed to
br>account for their properties

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1905 – A Big Year for Einstein

Photo-electric Effect

• "On a Heuristic Point of View concerning the Production and Transformation of Light." rcd Mar 18, pub Jun 9

Brownian Motion

• "On the Movement of Small Particles Suspended in Stationary Liquids Required by the Molecular-Kinetic Theory of Heat." rcd May 11, pub Jul 18

Special Relativity

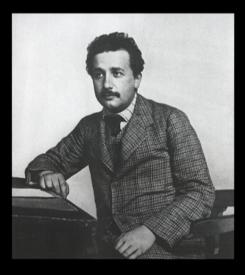
 "On the Electrodynamics of Moving Bodies." rcd Jun 30, pub 26 Sep

Size of Molecules*

• "A New Determination of Molecular Dimensions." rcd Aug 19, pub Feb 8

Mass-Energy Equivalence

• "Does the Inertia of a Body Depend upon Its Energy Content?" rcd Sep 27, pub Nov 21



Albert Einstein

* PhD Thesis and most cited

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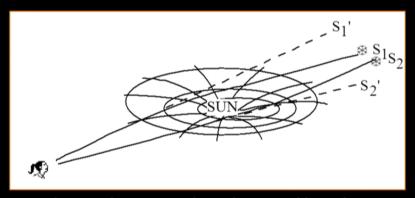
Outline

- The Hot Big Bang
 - Evidence for an Expanding Universe
 - Evidence for a Hot Big Bang
 - Evidence for Dark Matter and Dark Energy
- Contact With Fundamental Physics
 - · Dark Matter
 - Why Dark Energy is Harder
- Outlook

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Gravity, Matter & Geometry

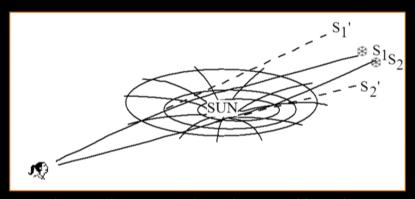


- According to Einstein gravity is really the response of space and time to the presence of matter.
 - The presence of energy curves space and time.
 - The curvature of space changes how objects move.

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Gravity, Matter & Geometry



- According to Einstein gravity is really the response of space and time to the presence of matter.
 - The presence of energy curves space and time.
 - The curvature of space changes how objects move.
- Knowing how matter is distributed over large scales tells us the shape and evolution of the Universe.

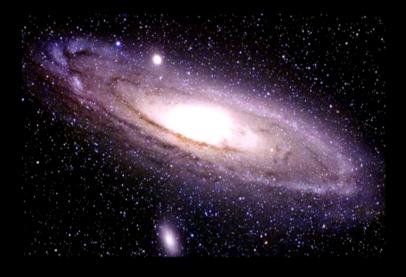
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Distribution of Visible Matter

Courtesv: Jason Ware

- Visible matter is found in space as stars, dust and gas.
- Most stars are found inside galaxies.
- Most galaxies are found inside clusters of galaxies.



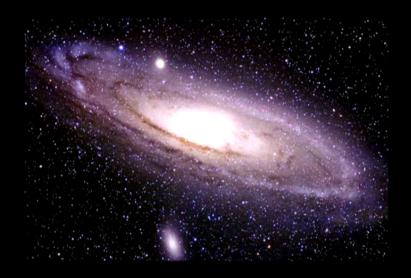
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Distribution of Visible Matter

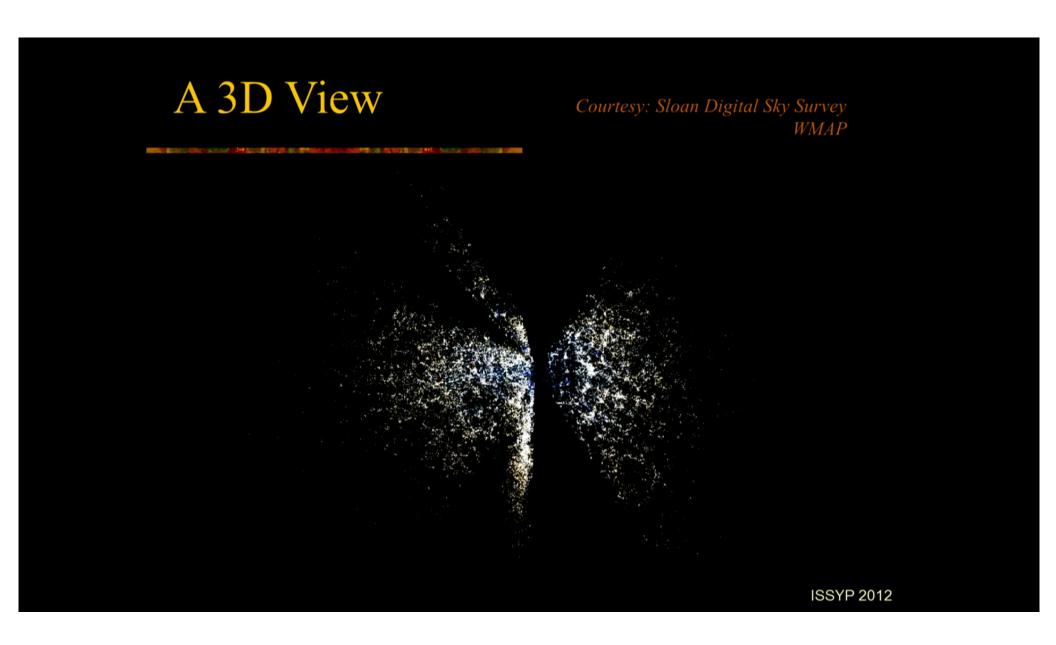
Courtesy: Jason Ware

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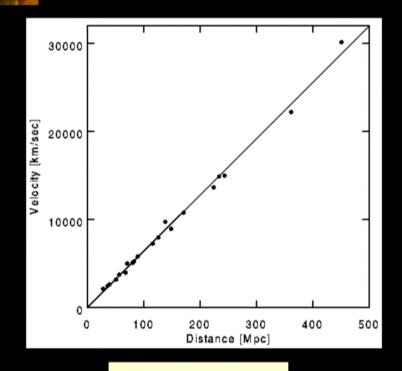
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- The sky is dark
- The Hubble Law
- The homogeneity and isotropy of the universe
- The slower decay of more distant supernovae

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Pirsa: 12070004 Page 10/62

- The sky is dark
- The Hubble Law
- The homogeneity and isotropy of the universe
- The slower decay of more distant supernovae



$$v = H_0 d$$

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Courtesy: Ned Wright's Cosmology Page

- The sky is dark
- The Hubble Law
- The homogeneity and isotropy of the universe
- The slower decay of more distant supernovae

In all directions galaxies recede from us with a speed proportional to their distance from us:

Implies all observers see the same law.

As expected if all started at the same point simultaneously.

$$v = H_0 d$$

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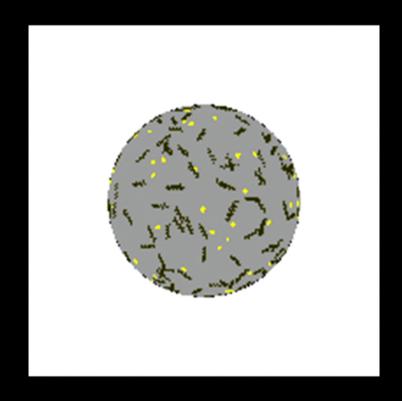
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- The sky is dark
- The Hubble Law
- The homogeneity and isotropy of the universe
- The slower decay of more distant supernovae



Courtesy: Ned Wright's Cosmology Page

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The Hot Big Bang

- Matter in an expanding container cools.
 - Extrapolating back in time suggests the Universe was once very small and very hot.
- The Hot Big Bang theory assumes the Universe was initially a small, hot soup of elementary particles.

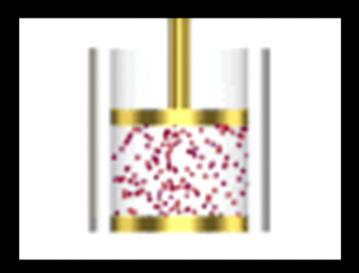


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How Would We Know?

- The Universe being hot *means* that the particles within it are randomly moving with higher speeds.
- At high enough temperatures particles get knocked apart into their constituents.

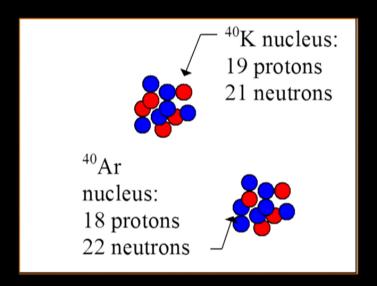


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How Would We Know?

- Atoms can survive once the Universe cools below 1000 degrees.
 - Charged electrons and nuclei combine to form neutral atoms.
- Nuclei can survive below 10 billion degrees.
 - Protons and neutrons combine to form nuclei.



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- Primordial element abundances
- The cosmic microwave background
- \bullet T_{CMB} vs distance

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- Primordial element abundances
- The cosmic microwave background
- T_{CMB} vs distance

When the universe cools to the point where nuclei can form (around 10¹⁰ K) then the relative abundances of the light elements can be calculated.

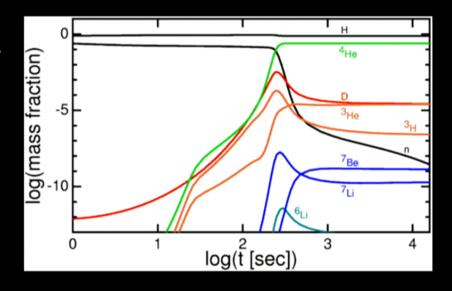
The predictions for the isotopes of H, He, Li and Be agree with the measured abundances.

Agreement is very sensitive to the total number of atoms present.

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- Primordial element abundances
- The cosmic microwave background
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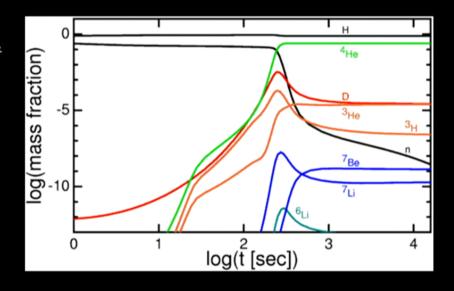
Burles, Nolette & Turner, 1999

Courtesy: Ned Wright's Cosmology Page

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- Primordial element abundances
- The cosmic microwave background
- T_{CMB} vs distance



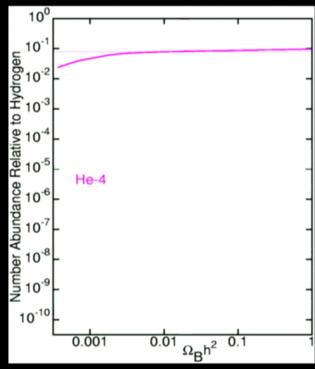
Burles, Nolette & Turner, 1999

Courtesy: Ned Wright's Cosmology Page

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- Primordial element abundances
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Total Mass Density of Atoms

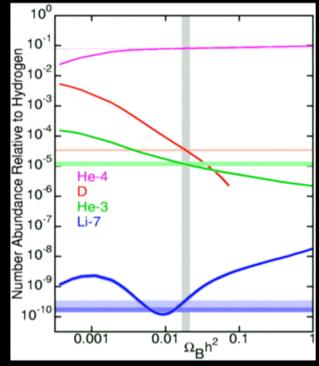
Burles, Nolette & Turner, 1999

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Courtesy: Ned Wright's Cosmology Pag

Pirsa: 12070004 Page 22/62

- Primordial element abundances
- The cosmic microwave background
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Total Mass Density of Atoms

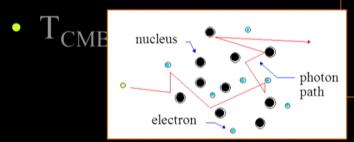
Burles, Nolette & Turner, 1999

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Courtesy: Ned Wright's Cosmology Pag

Pirsa: 12070004 Page 23/62

- Primordial element abundances
- The cosmic microwave background



When the universe cools enough for atoms to form (around 10³ K) it becomes transparent to light.

Above 1000 degrees atoms break up into electrons and nuclei.

Unlike atoms, electrons and nuclei carry electric charge and so scatter light very efficiently.

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- Primordial element abundances
- The cosmic microwave background
- T_{CMB} vs distance

When the universe cools enough for atoms to form (around 10³ K) it becomes transparent to light.

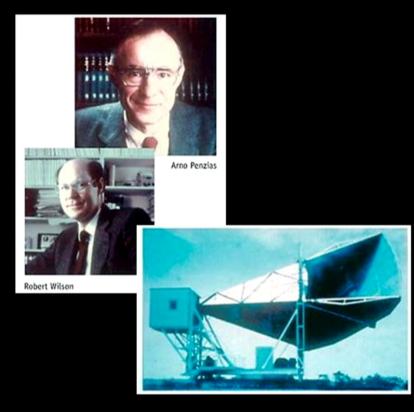
The light present at this time is still running around the universe.

This light is now seen as microwaves due to the intervening expansion of the universe.

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Pirsa: 12070004 Page 25/62

- Primordial element abundances
- The cosmic microwave background
- \bullet T_{CMB} vs distance

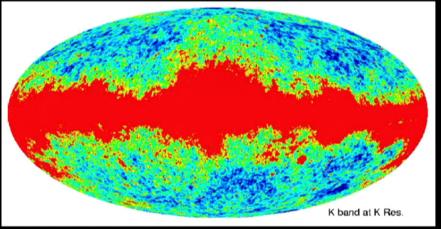


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WMAP

- Primordial element abundances
- The cosmic microwave background
- T_{CMB} vs distance



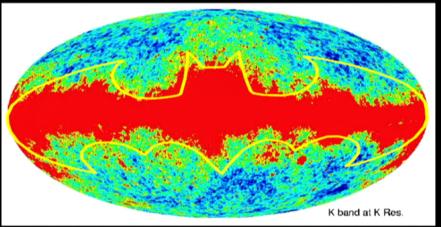
The temperature of the CMB has been measured as a function of direction.

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Batfit courtesy of Stephane Coutu

- Primordial element abundances
- The cosmic microwave background
- T_{CMB} vs distance



The temperature of the CMB has been measured as a function of direction.

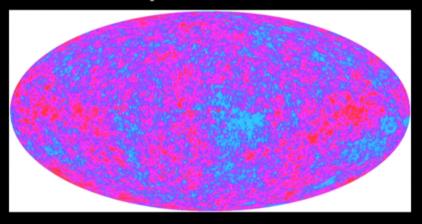
What is it telling us?

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- Primordial element abundances
- The cosmic microwave background
- T_{CMB} vs distance

CMB Temperature vs Direction



WMAP collaboration

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Outline

- The Hot Big Bang
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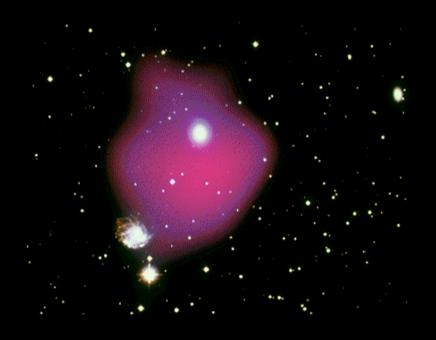
- Mass in galaxies
- Mass in clusters of galaxies
- Temperature fluctuations in the CMB
- Start of galaxy formation



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Pirsa: 12070004 Page 31/62

- Mass in galaxies
- Mass in clusters of galaxies
- Temperature fluctuations in the CMB
- Start of galaxy formation



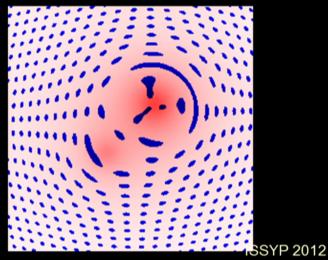
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Pirsa: 12070004 Page 32/62

- Mass in galaxies
- Mass in clusters of galaxies
- Temperature fluctuations in the CMB
- Start of galaxy formation

Courtesy: Ned Wright's Cosmology Page





Pirsa: 12070004 Page 33/62

- Mass in galaxies
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- Temperature fluctuations in the CMB
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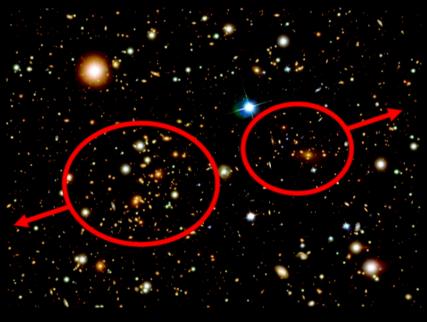
The Bullet Cluster: Separating
Dark and Visible Matter

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Colliding galaxy clusters

- Mass in galaxies
- Mass in clusters of galaxies
- Temperature fluctuations in the .
 CMB
- Start of galaxy formation



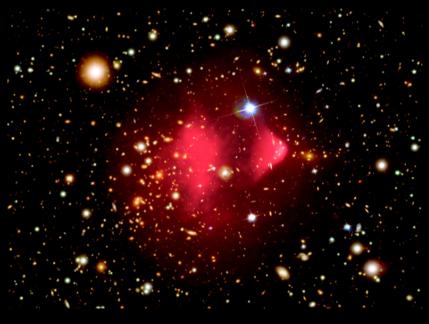
The Bullet Cluster: Separating Dark and Visible Matter

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Hot intra-cluster gas

- Mass in galaxies
- Mass in clusters of galaxies
- Temperature fluctuations in the CMB
- Start of galaxy formation



The Bullet Cluster: Separating
Dark and Visible Matter

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Mass measured by lensing

- Mass in galaxies
- Mass in clusters of galaxies
- Temperature fluctuations in the CMB
- Start of galaxy formation



The Bullet Cluster: Separating
Dark and Visible Matter

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- Mass in galaxies
- Mass in clusters of galaxies
- Temperature fluctuations in the CMB
- Start of galaxy formation

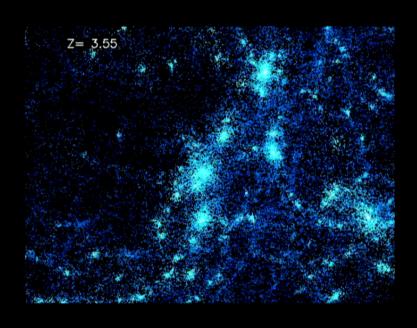


The Bullet Cluster: Separating
Dark and Visible Matter

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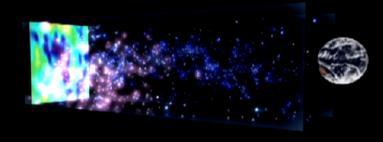
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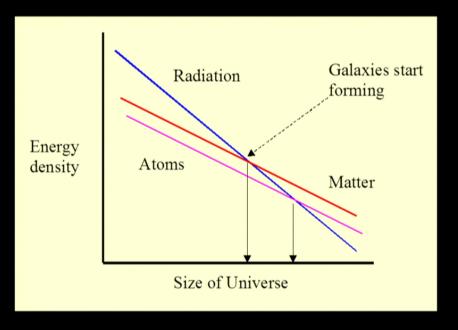
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- Mass in galaxies
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- Temperature fluctuations in the CMB
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- Brightness of very distant supernovae
- Flatness of the universe as a whole

Very distant objects should not precisely follow Hubble's Law because gravitational attraction should decelerate the universal expansion.

This can be tested by looking for deviations from Hubble's Law for very distant supernovae.

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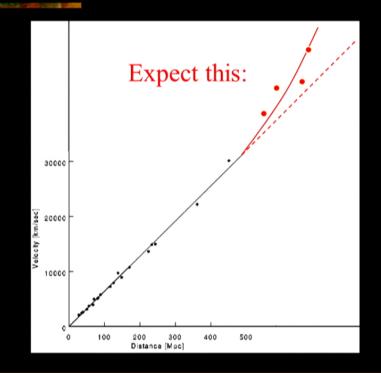
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- Brightness of very distant supernovae
- Flatness of the universe as a whole

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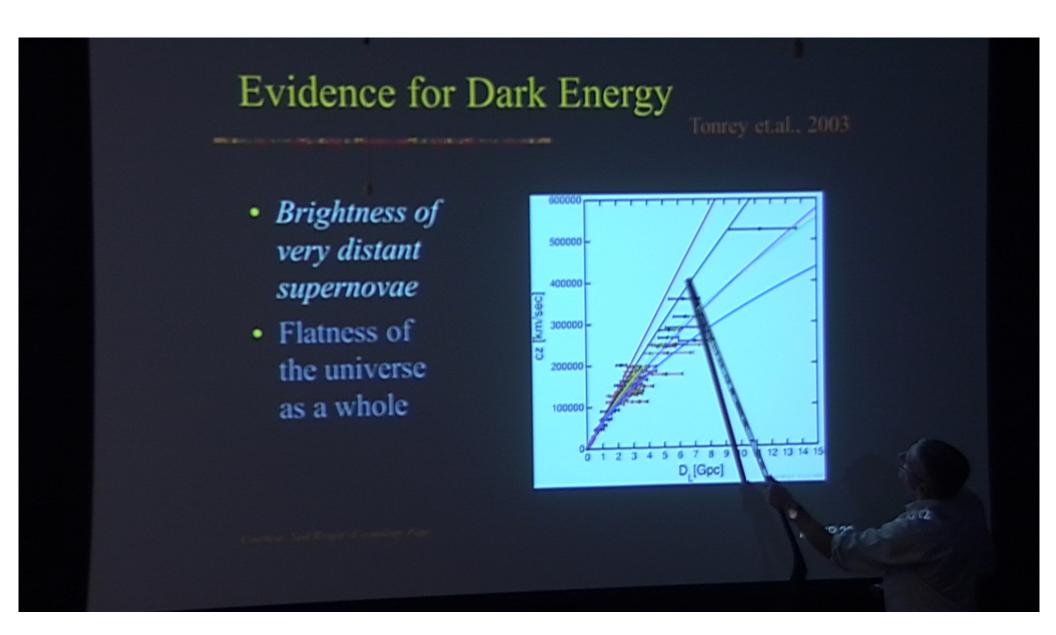
- Brightness of very distant supernovae
- Flatness of the universe as a whole



The universal expansion should be decelerating due to gravitational attraction

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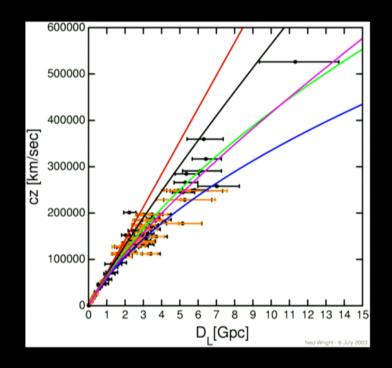
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Tonrey et.al., 2003

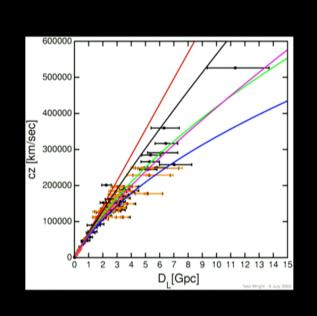
- Brightness of very distant supernovae
- Flatness of the universe as a whole

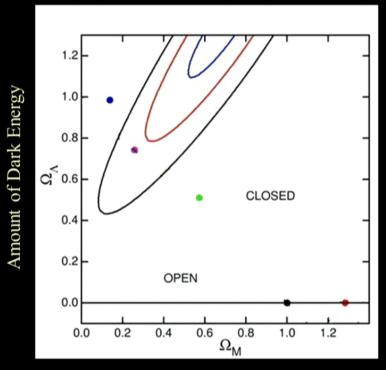


Courtesy: Ned Wright's Cosmology Page

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Tonrey et.al., 2003





Amount of Dark Matter

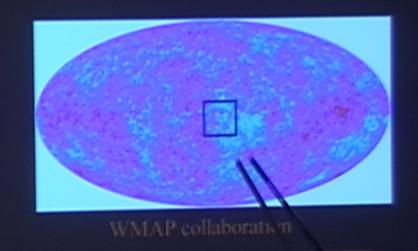
Courtesy: Ned Wright's Cosmology Page

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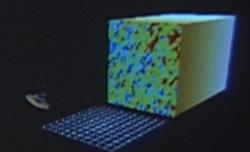
- Brightness of very distant supernovae
- Flatness of the universe as a whole



These are due to sound wa 'es in the primordial gas which emitted this light.

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- Brightness of very distant supernovae
- Flatness of the universe as a whole

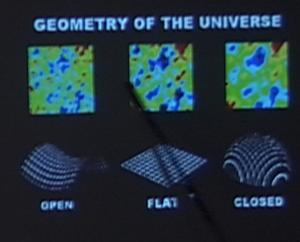


The CMB allows the inference of the properties of the later universe through which these photons pass SSYP 2012

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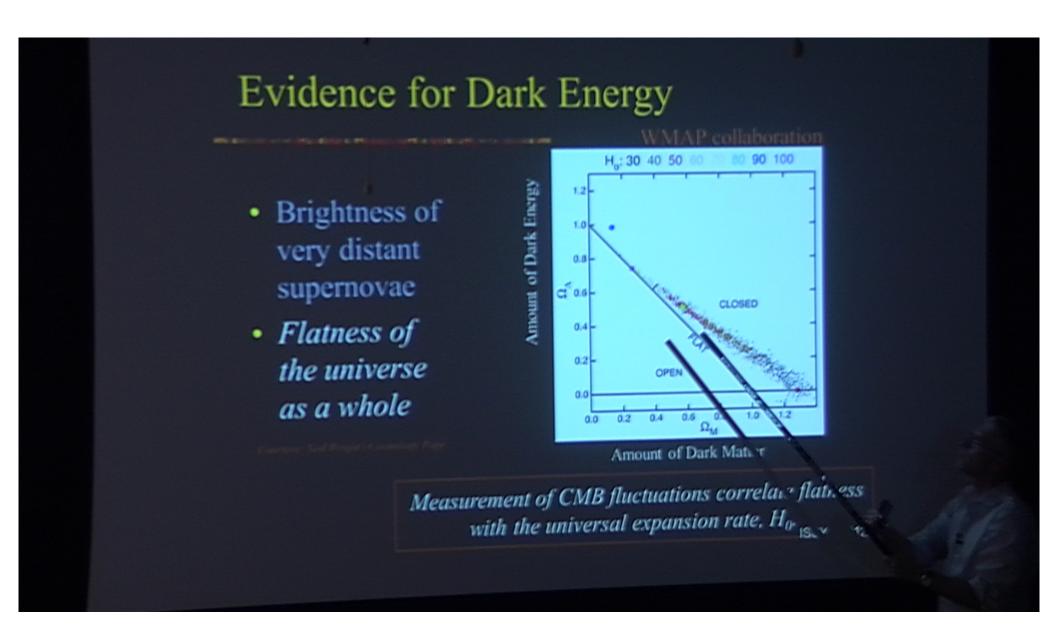


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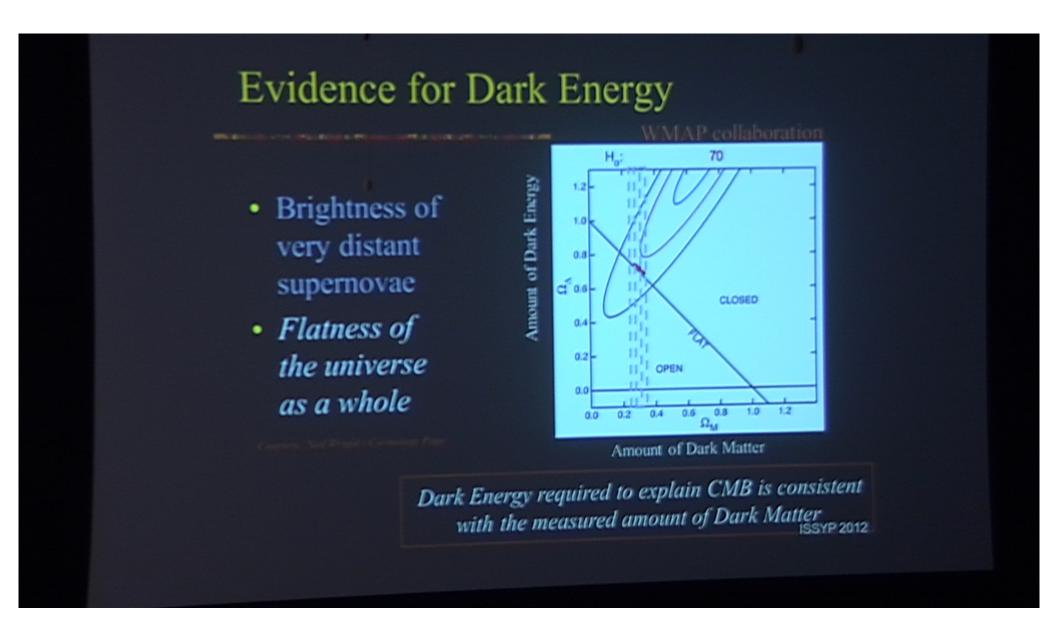


The CMB allows the inference of the properties of the later universe through which these photo: Later universe through which these photos.

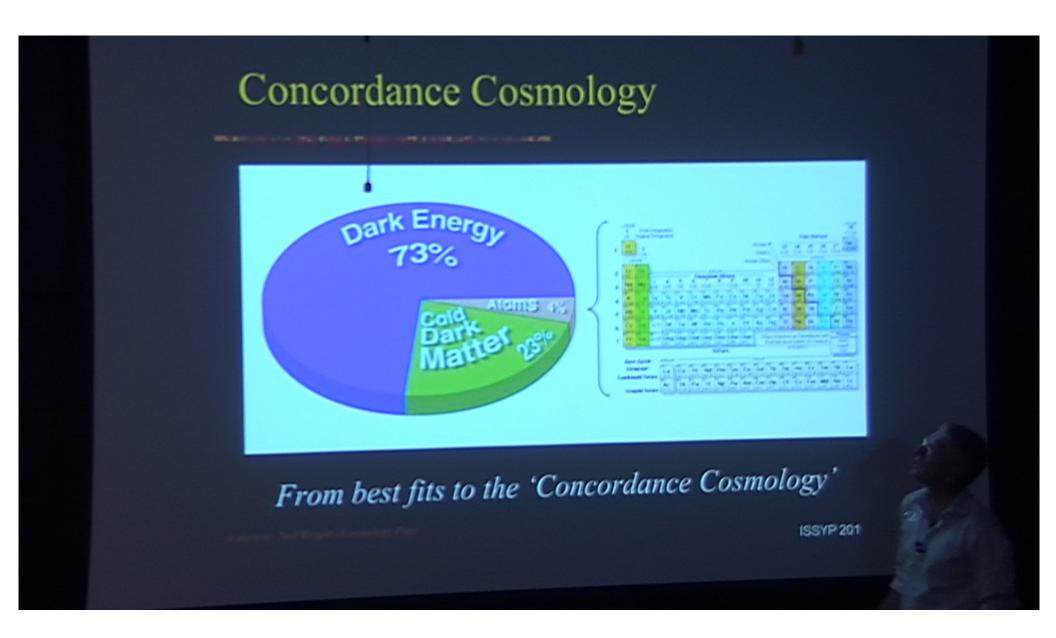
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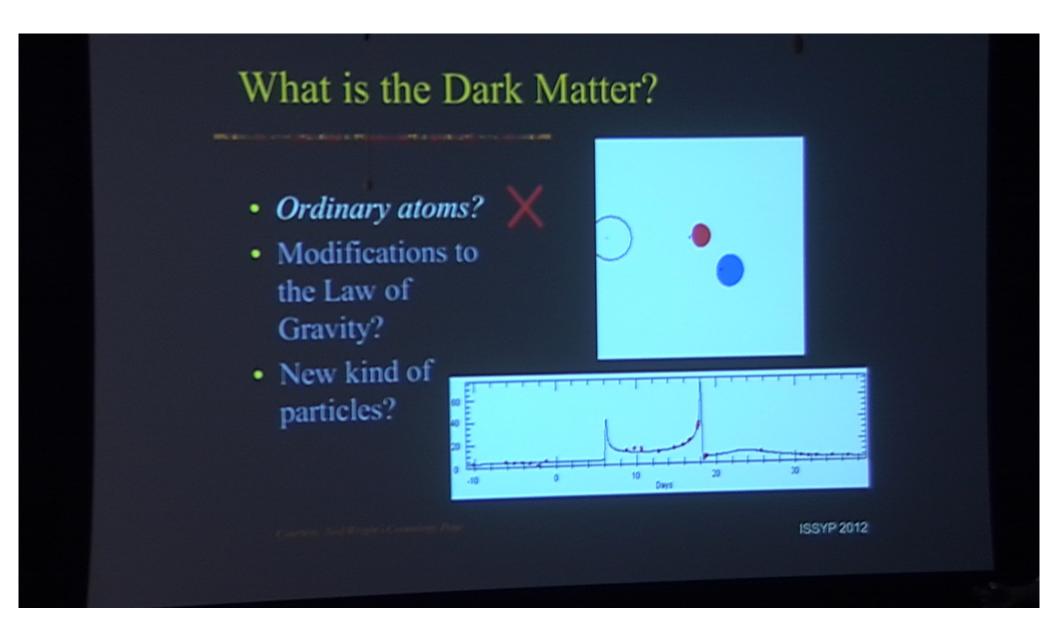
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What is the Dark Matter?

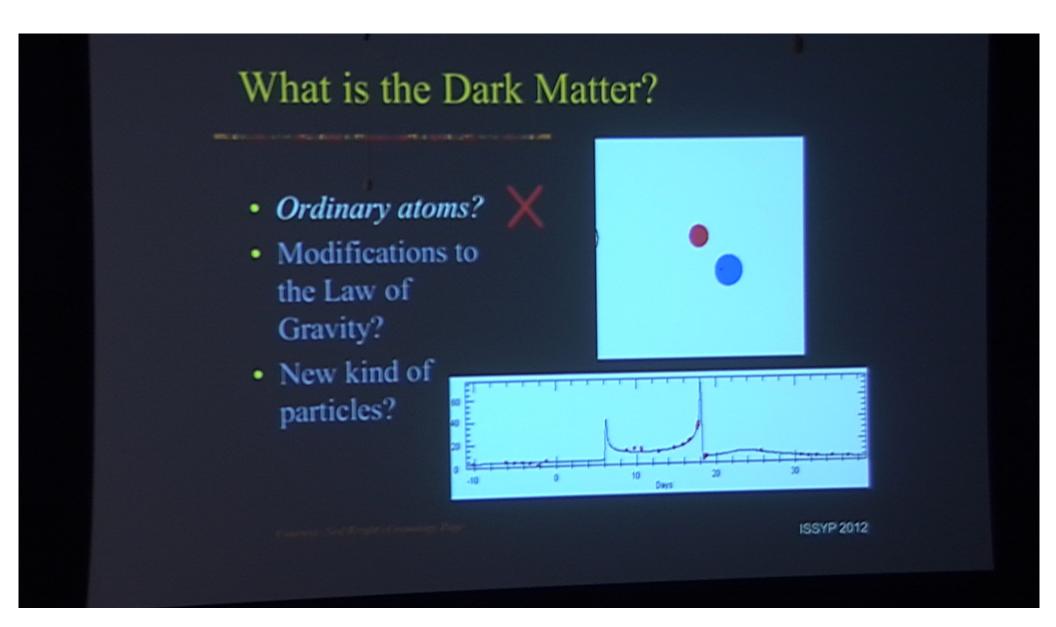
- Ordinary atoms?
- Modifications to the Law of Gravity?
- New kind of particles?

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Pirsa: 12070004 Page 55/62



Pirsa: 12070004 Page 56/62

What is the Dark Matter?

- Ordinary atoms?
- Modifications to the Law of Gravity?
- New kind of particles?

The devil is in the details:

No proposals yet succeed for galaxies and clusters and the CMB.

Very difficult to modify gravity at long distances without having problems with fundamental principles.

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What is the Dark Matter?

- Ordinary atoms?
- Modifications to the Law of Gravity?
- New kind of particles?

Weakly Interacting Massive Particles arise in most theories of microscopic physics.

Their residual cosmic abundance is naturally the right size to agree with the observed amount of Dark Matter.

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Outlook

- Cosmological observations are now redundantly testing the Hot Big Bang model.
 - Observations support the 'Concordance Cosmology'.

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Outlook

- Cosmological observations are now redundantly testing the Hot Big Bang model.
 - Observations support the 'Concordance Cosmology'.
- The concordance involves several lines of independent evidence for both Dark Matter and Dark Energy.
 - Neither can be dark forms of ordinary atoms.
- Dark Matter may be new kinds of elementary particles, but Dark Energy is more difficult to embed into a fundamental theory (although a few proposals do exist).
 - If so we'll know from a variety of observational tests.

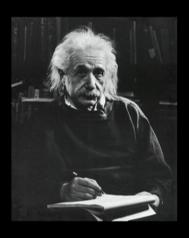
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Outlook

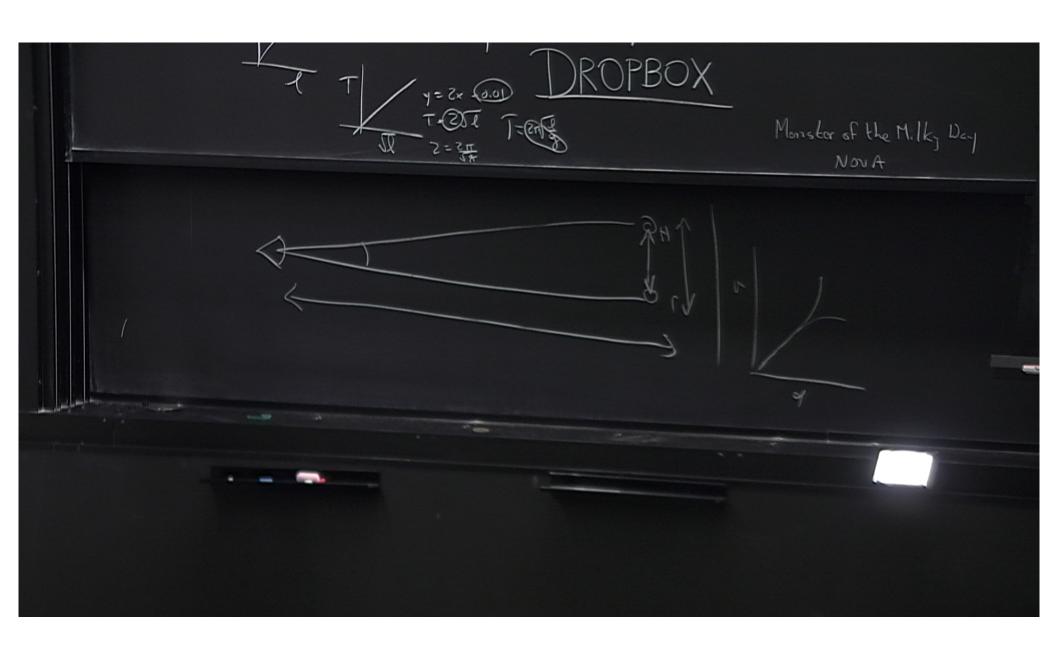
"If I had only known, I'd have become a locksmith."

Albert Einstein



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