

Title: Time and a Physical Hamiltonian for Quantum Gravity

Date: May 12, 2012 09:50 AM

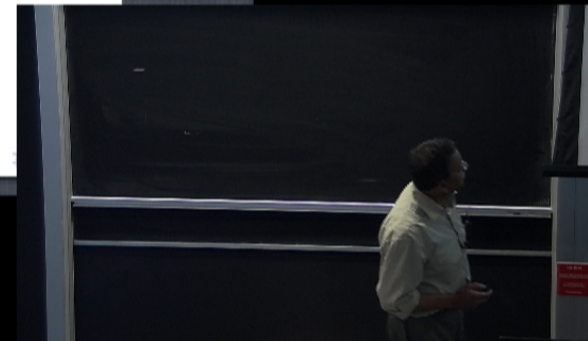
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Abstract: I will describe an approach to the problem of time that uses dust as a time variable. The canonical theory is such that there is a true Hamiltonian with spatial diffeomorphisms as the only gauge symmetry. This feature, and the form of the Hamiltonian, suggest a model for non-perturbative quantum gravity that is computationally accessible using the formalism of loop quantum gravity.

*Hey, raise those eyes for you do not see
the universe's mysteries.
Behind the green are games anew.
Despair not.*

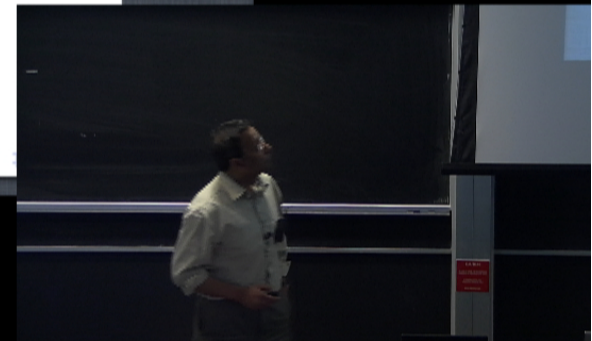
– Hafez of Shiraz (14th century)

1. Problems of time in QG and possible "solutions"
2. Canonical QG: state of the art
3. A model QG: classical theory and quantization
4. Summary and problems



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Navigation icons: back, forward, search, etc.



Problem of time in QG

3+1 ADM action

$$S = \int d^3x dt \left(\tilde{\pi}^{ab} \dot{q}_{ab} + P_\phi \dot{\phi} - N(\mathcal{H}_G + \mathcal{H}_M) - N^a(C_a^G + C_a^M) \right). \quad (1)$$

source of problem: Hamiltonian constraint

$$\mathcal{H} \equiv \mathcal{H}_G(q, \pi) + \mathcal{H}_M = 0. \quad (2)$$

“Dirac quantization”

$$\hat{\mathcal{H}}\Psi[q, \phi] = 0. \quad (3)$$

Not a TDSE – evolution is gauge.

What is the Hilbert space in which $\Psi[q, \phi]$ lives?

How is $\hat{\mathcal{H}}$ defined?

What is the physical Hilbert space?

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"Solutions" to the problem of time

Fix time gauge: (partially) reduced phase space quantization
(eg. Horava-Lifshitz, Shape dynamics)

Problems:

no canonical choice

square root or other complicated form of physical hamiltonian: hard to write operator

different choices \rightarrow unitarily inequivalent quantum theories.

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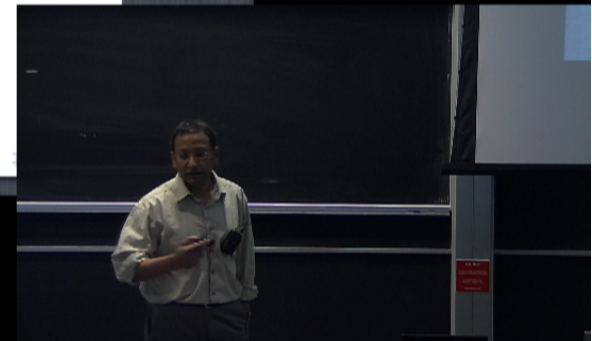
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Do not fix time gauge: Dirac Quantization

Emergent time: time arises in IR (WKB or other mechanism)

Relational time: correlations between (quantum) observables

Hamiltonian constraint is linear in one of the momenta \rightarrow functional (many fingered time) Schrodinger equation.

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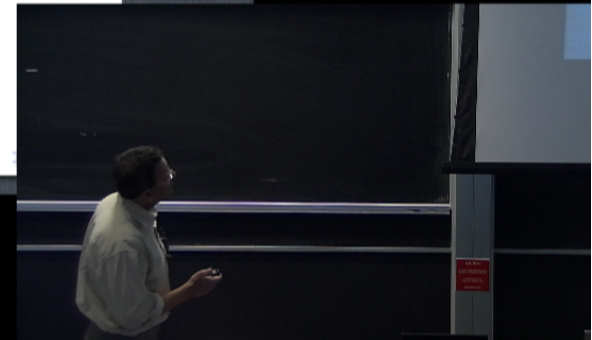


Non-perturbative QG: state of the Dirac quantization program

1. Define a suitable *kinematical Hilbert space* H_{kin}
2. Define constraints as operators on H_{kin} .
3. Check algebra of constraint operators
4. Find space of solution of the constraints: physical Hilbert space H_{phys} .
5. physical observables, semiclassical states \dots

State of the art in LQG: 1, 2 + some solutions of constraints that also satisfy

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The model

$$S = \frac{1}{4G} \int d^4x \sqrt{-g} R + S_M - \frac{1}{2} \int d^4x \sqrt{-g} M (g^{ab} \partial_a T \partial_b T + 1),$$

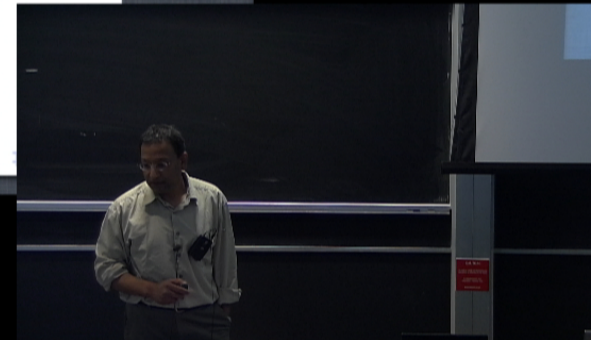
Fields g, M, T . S_M is any matter action.

With $U_a = \partial_a T$ the dust stress-energy tensor is

$$T^{ab} = M U^a U^b + (M/2) g^{ab} (g_{cd} U^c U^d + 1)$$

Special case of Brown-Kuchar action which has 4 dust fields.

Considered much before: in spherical symmetry \rightarrow
Tolman-Bondi-Lemaitre model.



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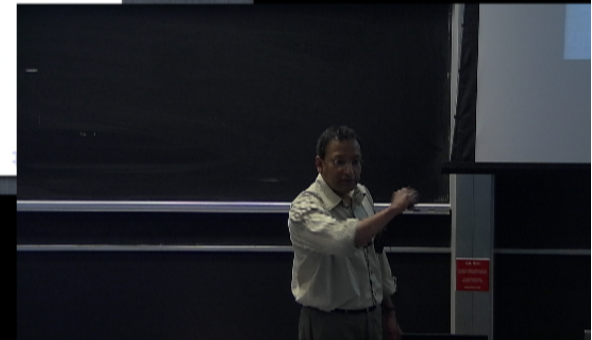
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Hamiltonian theory of dust

Substitute ADM metric

$$ds^2 = -N^2 dt^2 + q_{ab}(N^a dt + dx^a)(N^b dt + dx^b)$$

into dust action:

$$L_D = \frac{M\sqrt{q}}{2N} \left[(\dot{T} + N^a \partial_a T)^2 - N^2 (q^{ab} \partial_a T \partial_b T + 1) \right].$$

$$p_T = \frac{\partial L_D}{\partial \dot{T}} = \sqrt{q} \frac{M}{N} (\dot{T} + N^a \partial_a T)$$

$$S_D = \int dt d^3x \left[p_T \dot{T} - N \mathcal{H}_D - N^a C_a^D \right].$$

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– a square root Hamiltonian.

Final canonical theory: $(q, \pi), (T, p_T), (\phi, P_\phi)$ subject to constraints

$$\mathcal{H}_G + \mathcal{H}_D + \mathcal{H}_M = 0$$

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Time gauge fixing

Dirac quantization is difficult.

But the form of the dust Hamiltonian suggest a natural time gauge:

$$T = t.$$

This is second class with \mathcal{H} .

Preservation of gauge under time evolution gives

$$\dot{T} = \dot{t} = 1 = \{T, \int_{\Sigma} d^3x (N\mathcal{H} + N^a \mathcal{C}_a)\}|_{T=t}$$

This fixes $N = 1$ and leaves N^a arbitrary. Lapse and shift decouple!

$$-p_T = H_{phys} = \mathcal{H}_G(\pi, q) + \mathcal{H}_M(\phi, P_\phi)$$

- Physical hamiltonian is not a square root.
- It is independent of time.

Quantization

Basic idea

Use the Ashtekar-Barbero canonical variables for gravity: (A_a^i, E_i^a) .

Use LQG H_{kin} : spin network basis.

H_{kin} carries a representation of the spatial diffeomorphism group.
(cf. QFT Hilbert space carries a rep of Poincare group).

At least one construction of \mathcal{H}_G available on H_{kin} (Thiemann)

But in this theory \mathcal{H}_G is the physical Hamiltonian!

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H_{kin}

Connection triad formulation contains Gauss law constraint $D_a E^{ai} = 0$ in addition to $\mathcal{H} = 0$ and $\mathcal{C}_a = 0$.

Connection representation:

$$\psi[A] = f(U_{\gamma_1}^{j_1}(A), U_{\gamma_2}^{j_2}(A), \dots, U_{\gamma_n}^{j_n}(A))$$

γ_i : curves embedded in spatial slice; collection $\{\gamma_i\}$ is a graph Γ .

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Spin network states:

Tie up holonomies at points of intersection of edges using *intertwiners*

\mathcal{I} : eg. trivalent vertex

$$\psi[A] = [U_{\gamma_1}^{j_1}(A)]^{\alpha_1 \beta_1} [U_{\gamma_2}^{j_2}(A)]^{\alpha_2 \beta_2} [U_{\gamma_3}^{j_3}(A)]^{\alpha_3 \beta_3} \mathcal{I}_{\alpha_1 \alpha_2 \alpha_3}^{j_1 j_2 j_3}$$

SU(2) Haar measure provides inner product.

$$|\Gamma; j_1, j_2 \dots j_n; \mathcal{I}_1 \dots \mathcal{I}_m\rangle$$

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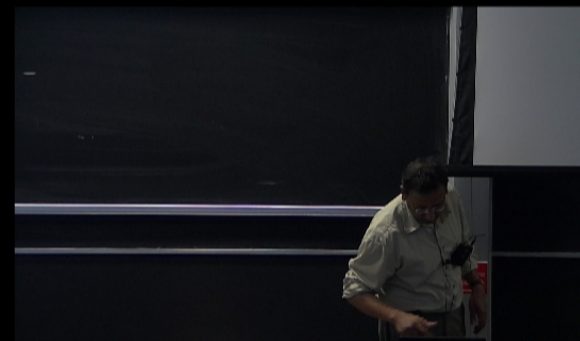
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But in this theory \mathcal{H}_G is the physical Hamiltonian!

Navigation icons: back, forward, search, etc.



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Hamiltonian operator

Physical Hamiltonian density is

$$\mathcal{H}_G = \frac{\gamma^2}{2\sqrt{\det E}} E_i^a E_j^b \left(\epsilon^{ij}{}_k F_{ab}^k + 2(1 - \gamma^2) K_{[a}^i K_{b]}^j \right)$$

A possible operator (following Thiemann):

$$\hat{H}^G = \sum_{v \in V(\Gamma)} \hat{\mathcal{H}}_v^G$$

The sum is over vertices of a graph.

$\hat{\mathcal{H}}_v^G$ is composed of:

- (i) the volume operator $\hat{V}(v)$
- (ii) the combination $\hat{h}_e[\hat{h}_e^{-1}, \hat{V}]$ of holonomies and volume.
- (iii) the holonomies $\hat{h}_{\square}(v)$ along the minimal closed loops based at v .

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Summary

- ▶ A model for non-perturbative QG: a completion of the LQG program.
- ▶ No gauge anomaly $[H(N), H(M)]$: hamiltonian constraint \rightarrow physical Hamiltonian.
- ▶ Physical Hilbert space, Hamiltonian operator, observables (area, volume, \dots), semiclassical states.
- ▶ Matter couplings immediate: just add any matter term to H_{phys} .

Navigation icons: back, forward, search, etc.

Open problems

Model provides a test bed for physical/technical problems.

- ▶ Other ways to write the Hamiltonian operator. \hat{H} on H_{diff} ?
- ▶ What do reduced models look like? eg. Spherical symmetry + scalar field (work in progress)
- ▶ What is a black hole in quantum gravity?
- ▶ Emergence of semiclassical gravity. Hawking radiation?
- ▶ Cosmology with dust time: FRW reduction, fluctuations etc.
- ▶ Lorentz invariance?

Curiosities

1. Another use of spin network Hilbert space:
SU(2) theory (VH, K. Kuchar '91)

$$S = \frac{1}{L^2} \int \text{Tr}(e \wedge e \wedge F(A)) \left(+ \Lambda \int \text{C.S.}(A) \right)$$

e_a^i SU(2) dreibein in 4d. A_a^i connection.

- power counting non-renormalizable.
- 3 local degrees of freedom.
- Hamiltonian constraint vanishes identically. Spatial diffeo + Gauss constraints only.
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2. "Dusty string"

$$S = \int d^2x \sqrt{h} (G^{AB} h^{ab} \partial_a X_A \partial_b X_B + M(h^{ab} \partial_a T \partial_b T + 1))$$

- not conformally invariant.
- $T = t$ gauge: physical Hamiltonian with $\text{Diff}(S^1)$ constraint.
- manifest target space Poincare invariance.
- gravitons?

3. "Little λ ": Can put this in by hand into the gravitational physical Hamiltonian in the dust gauge fixed action.

$$S^{GF} = \int d^3x dt \left[\pi^{ij} \dot{q}_{ij} - \frac{1}{\sqrt{q}} (\pi^{ij} \pi_{ij} - \lambda \pi^2) - \sqrt{q} V(q) \right]$$



