

Title: Explorations in Particle Theory - Lecture 1

Date: Apr 02, 2012 09:00 AM

URL: <http://pirsa.org/12040000>

Abstract:

Dark Matter

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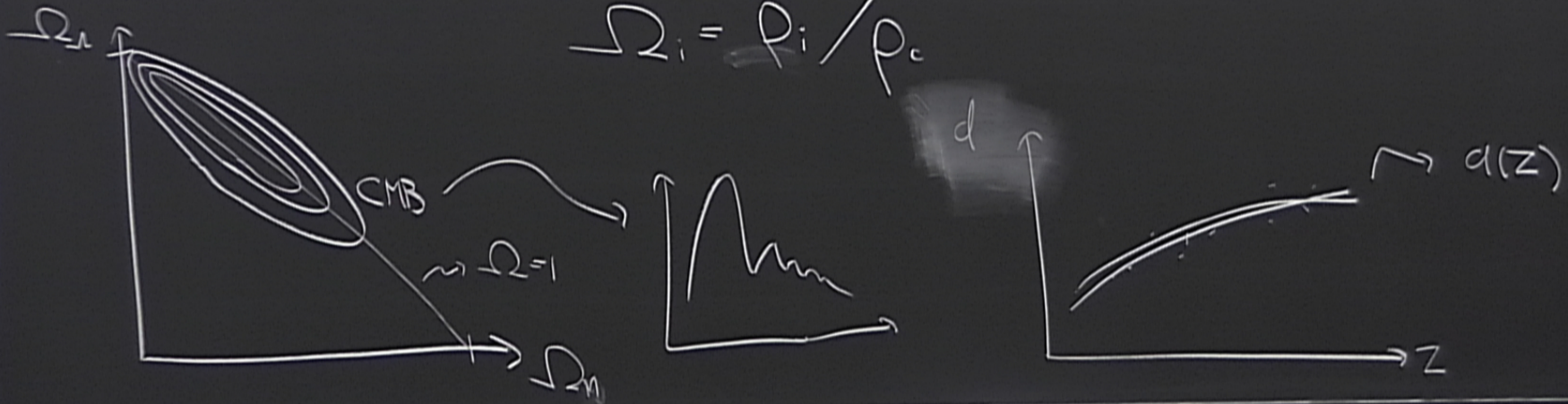
## Evidence for DM

$\Lambda$ CDM  $\rightarrow$  Dark Energy  $\oplus$  Dark Matter  $\oplus$  Baryons

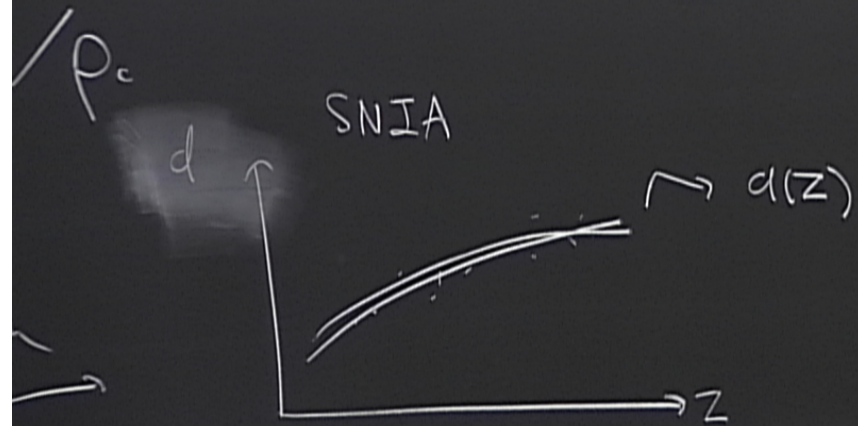
# Evidence for DM

$\Lambda$ CDM  $\rightarrow$  Dark Energy  $\oplus$  Dark Matter  $\oplus$  Baryons

$$\Omega_i = \rho_i / \rho_c$$



# Dark Matter ⊕ Baryons

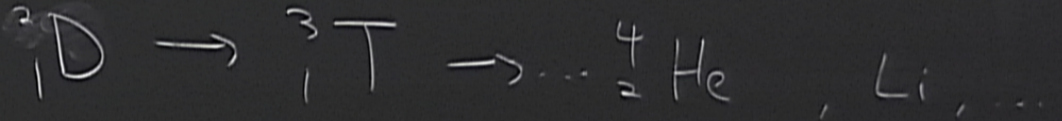
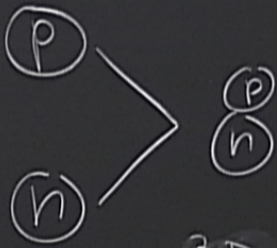


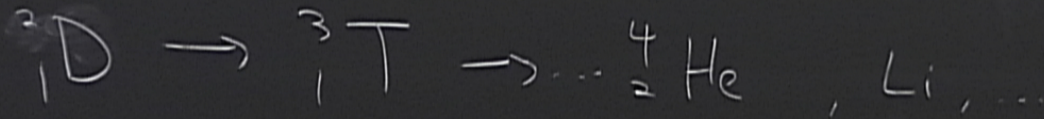
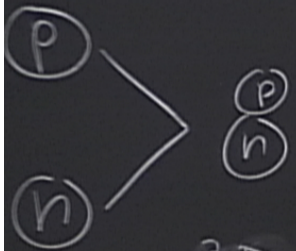
$$\Omega_\Lambda = 0.72$$

$$\Omega_M = 0.22$$

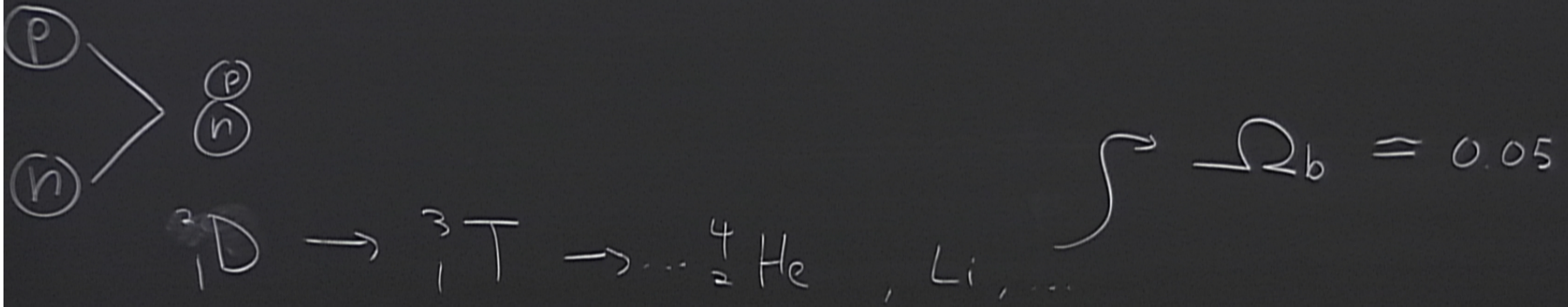
$$\Omega_b = 0.05$$

Nucleosynthesis:

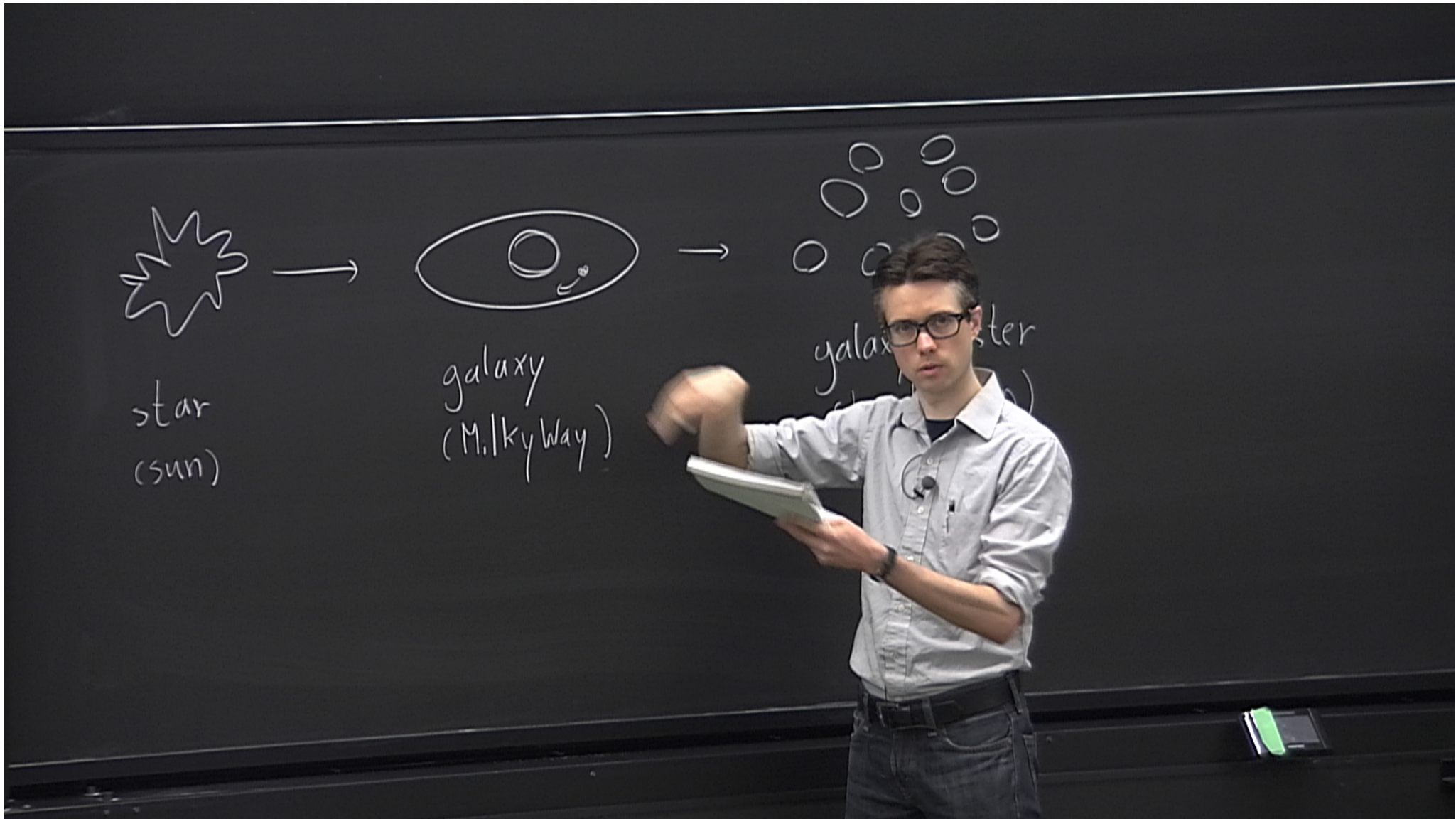




$$\Omega_b = 0.05$$

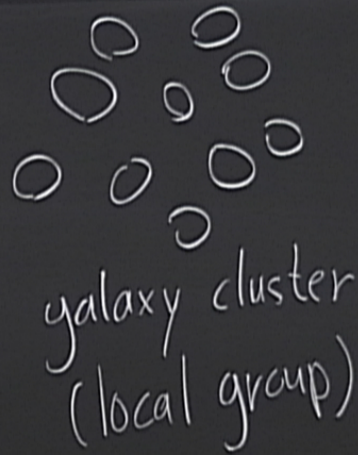








galaxy  
(Milky Way)

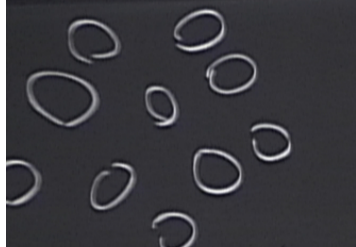


galaxy cluster  
(local group)



galaxy superclusters

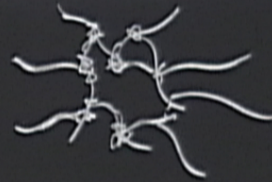




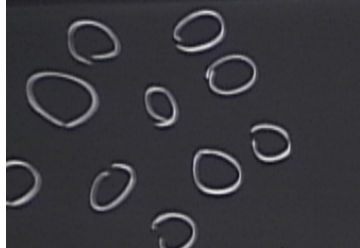
galaxy cluster  
(local group)



galaxy superclusters



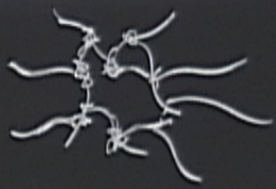
filaments and voids



galaxy cluster  
(local group)



galaxy superclusters



filaments and voids

→ ...  
galaxy superclusters

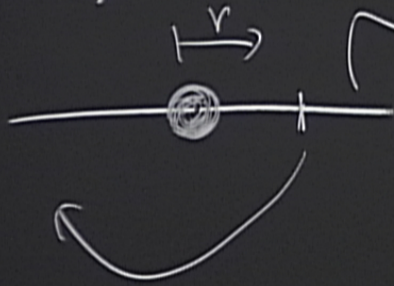


→  
filaments and voids

Hubble



# Galaxy DM

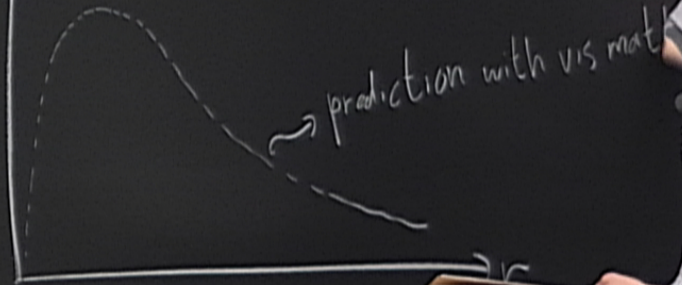


$v(r)$  = local rotation speed

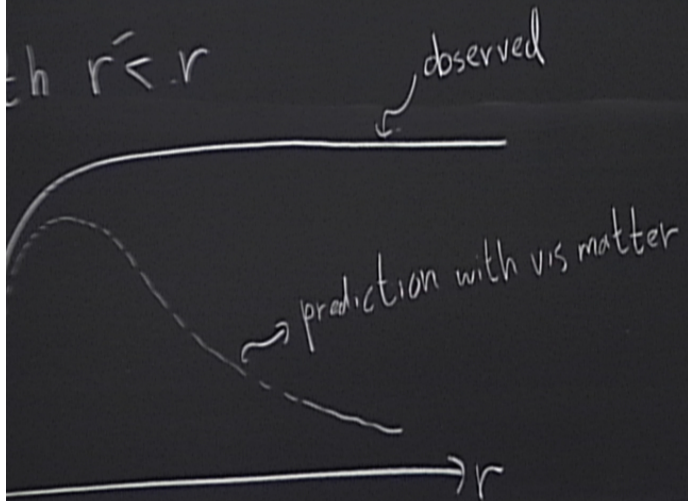
$M_{vis}(r)$  = total visible matter with  $r' < r$

$$m \frac{v^2}{r} = \frac{G M(r) \cdot m}{r}$$

$v(r)$



Assume spherical "DM halo"



Assume spherical "DM halo"

$v \sim \text{const}$  at large  $r$

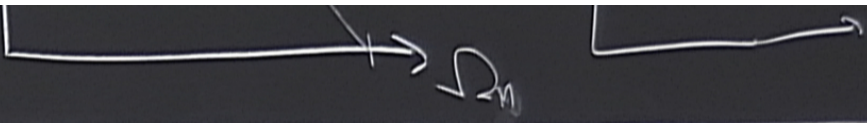
$$\rho(r) \sim r^{-2}$$

for  $r < r$  observed

prediction with vis matter

$r$

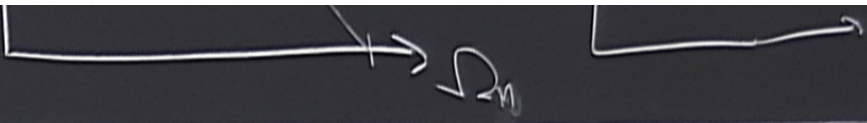




☺

# Galaxy Clusters

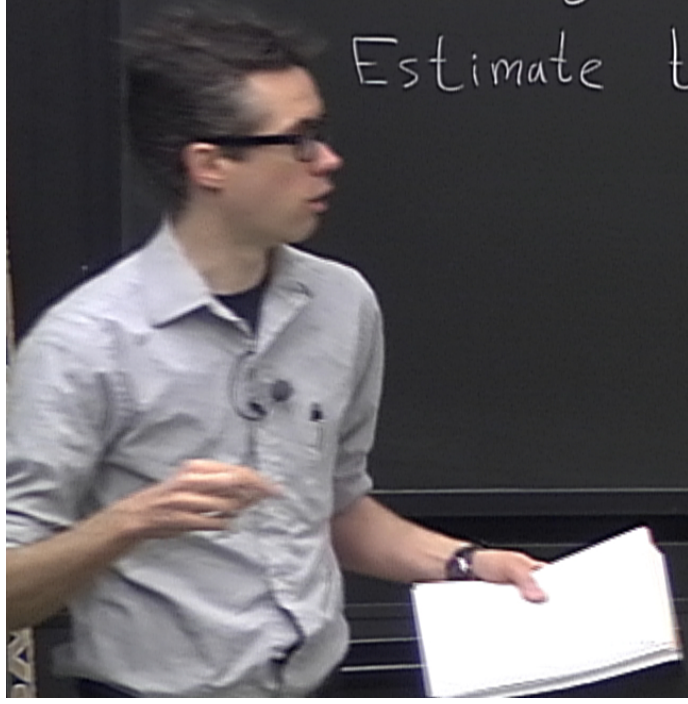
⊕



☺ Galaxy Clusters

↳ self-gravitating

Estimate total mass from luminosity.



☺ Galaxy Clusters

↳ self-gravitating

Estimate total mass from luminosity.

Virial Theorem:  $\langle T \rangle \sim \langle V \rangle$   
 $\frac{5}{2} N^2 \sim \frac{GM}{R}$

① Galaxy Clusters

↳ self-gravitating

Estimate total mass from luminosity

Virial Theorem:  $\langle T \rangle \sim \langle V \rangle$   
 $\frac{5}{2} N^2 \sim \frac{GM_{vis}}{R}$

Gravitational



# Gravitational Lensing

ss from lu

$\langle T \rangle$


$\frac{5}{N^2}$

$\frac{GM_{vis}}{R}$

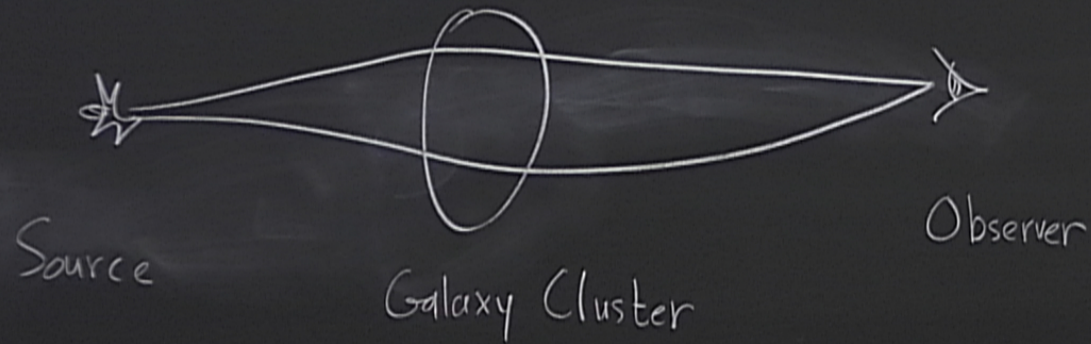


# Gravitational Lensing

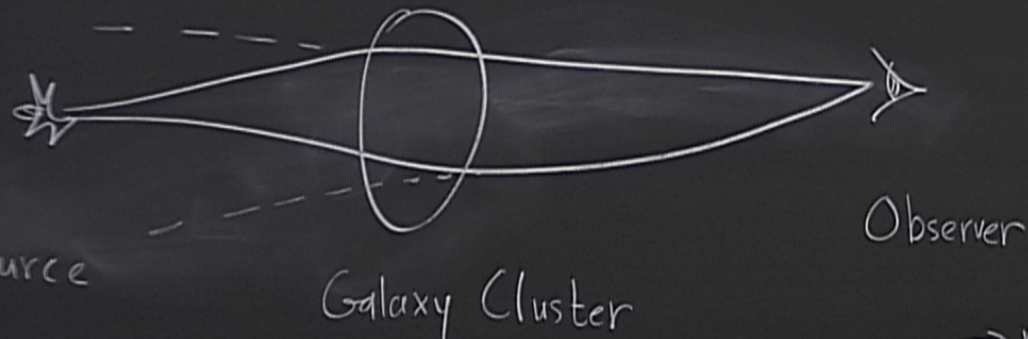
  
Source

  
Observer

# Gravitational Lensing



# Gravitational Lensing



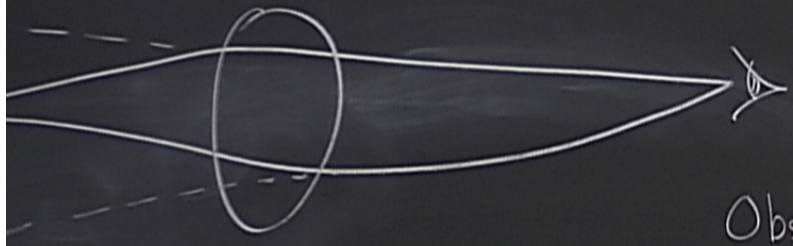
Strong Lensing: multiple source images

Weak Lensing: distorted image

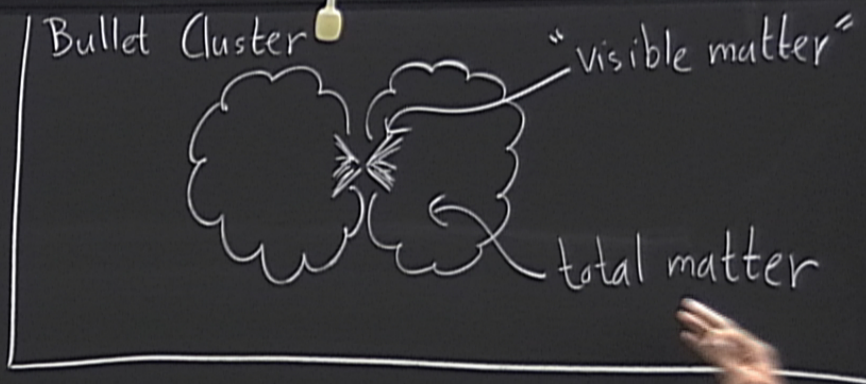
maps "see" much more matter than is luminous



Gravitational Lensing



Galaxy Cluster  
lensing. multiple source images  
lensing: distorted image



Observer

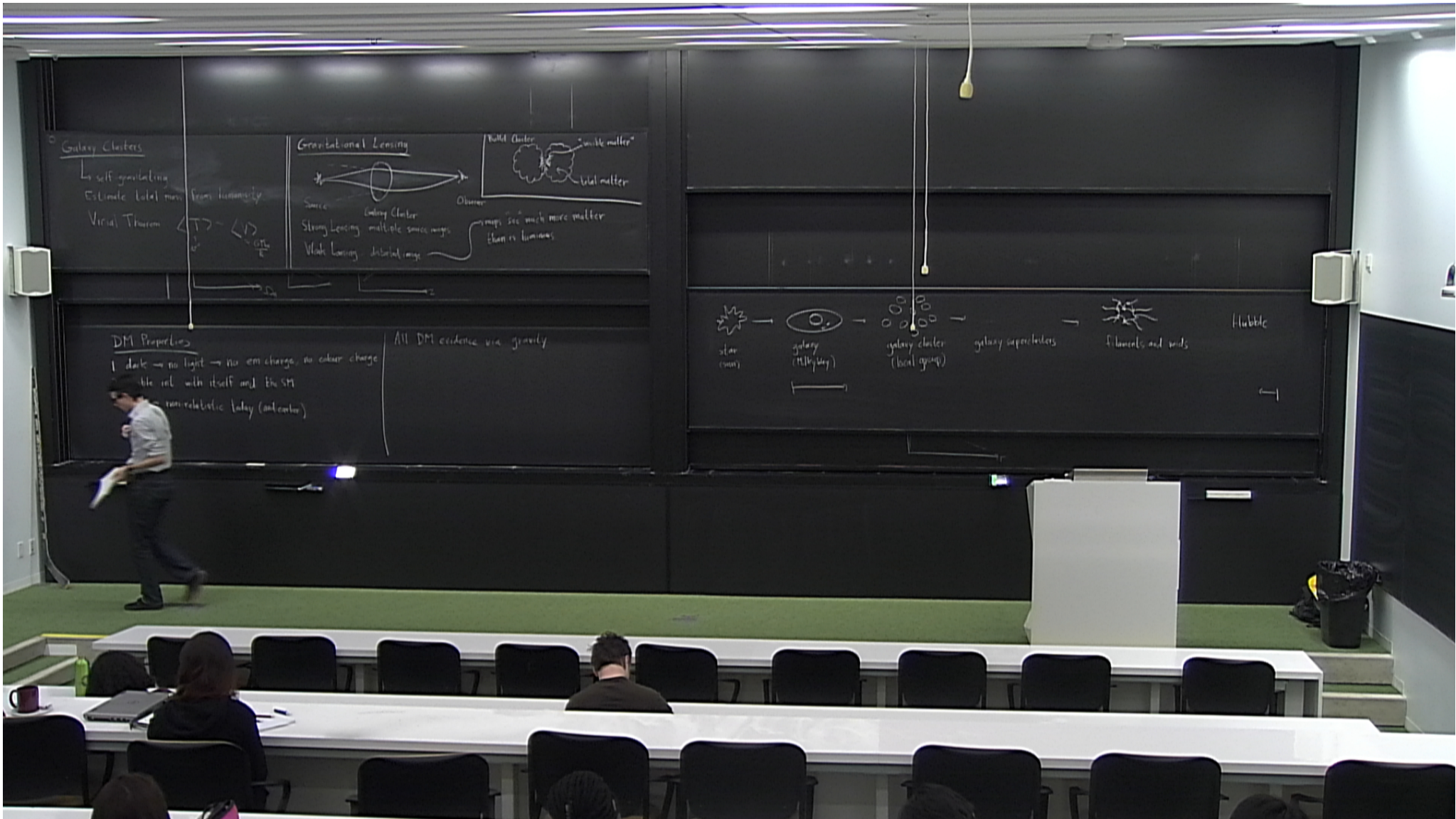
maps "see" much more matter than is luminous

## DM Properties

1. dark  $\rightarrow$  no light  $\rightarrow$  no em charge, no colour charge
2. feeble int. with itself and the SM

## DM Properties

1. dark  $\rightarrow$  no light  $\rightarrow$  no em charge, no colour charge
2. feeble int. with itself and the SM
3. cold  $\rightarrow$  non-relativistic today (and earlier)



All DM evidence via gravity.

or charge

All DM evidence via gravity.

Modified Gravity

or charge

All DM evidence via gravity.

or charge

Modified Gravity

All DM evidence via gravity.

Modified Gravity

or charge



All DM evidence via gravity.

Modified Gravity

Experiment

↳ non-gravitational evidence for DM.

# Outline

1 DM formation

↳ how do you <sup>get</sup>  $\rho_{DM}$  cosmologically?

DM is WIMP-like

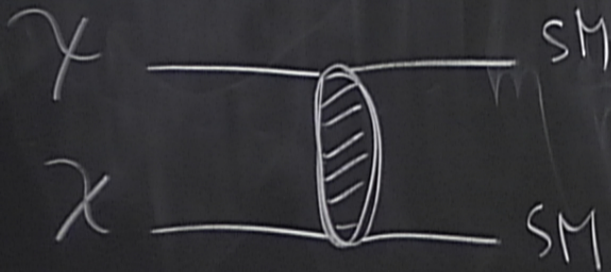
# Outline

1 DM formation

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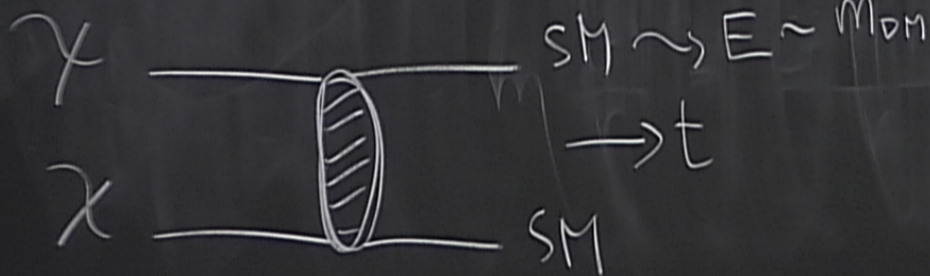
↳ DM is massive  
weak-ish int



# Outline

## I DM formation

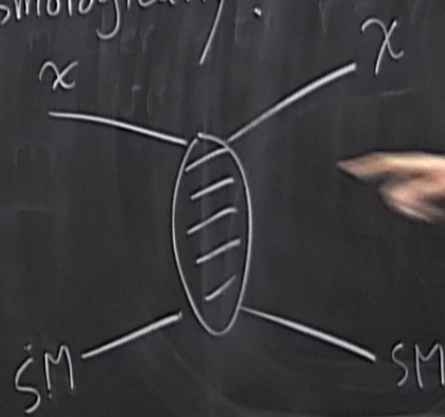
↳ how do you <sup>get</sup>  $\rho_{DM}$  cosmologically?



Indirect Detection

DM is WIMP-like

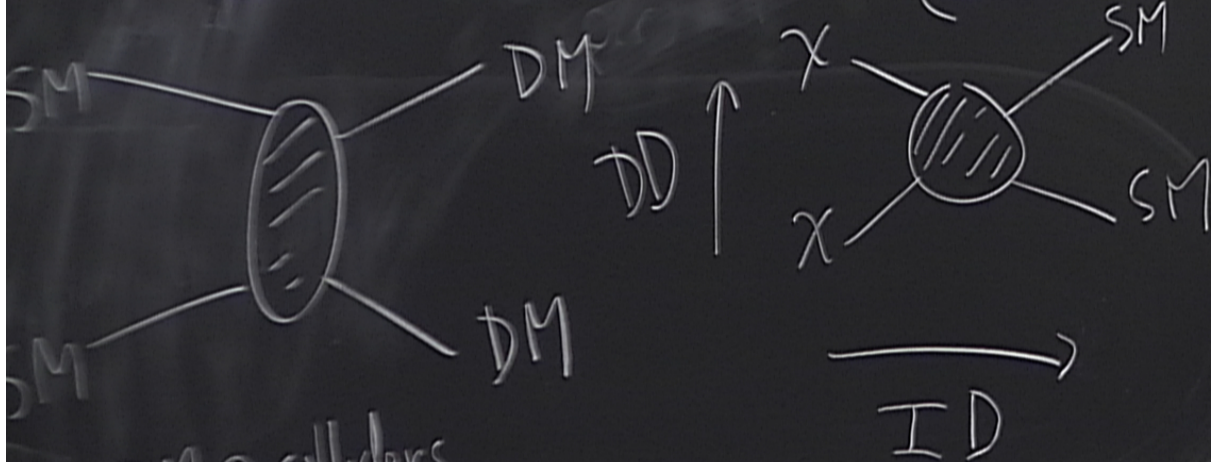
↳ DM is massive  
weak-ish



WIMP-like

↳ DM is massive

weak-ish interaction with the SM collider



3. Candidates

4. Axions

