

Title: Goldstone Fermion Dark Matter

Date: Jan 06, 2012 10:00 AM

URL: <http://pirsa.org/12010121>

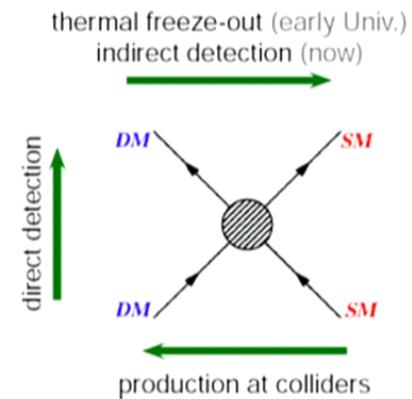
Abstract: We propose that the fermionic superpartner of a weak-scale Goldstone boson can be a natural WIMP candidate. The p-wave annihilation of this 'Goldstone fermion' into pairs of Goldstone bosons automatically generates the correct relic abundance, whereas the XENON100 direct detection bounds are evaded due to suppressed couplings to the Standard Model. Further, it is able to avoid indirect detection constraints because the relevant s-wave annihilations are small. The interactions of the Goldstone supermultiplet can induce non-standard Higgs decays and novel collider phenomenology.

## SUSY LSP as WIMP

- R-Parity  $\rightarrow$  Stability
- Weak Scale  $\rightarrow$  Naturalness

Natural WIMP

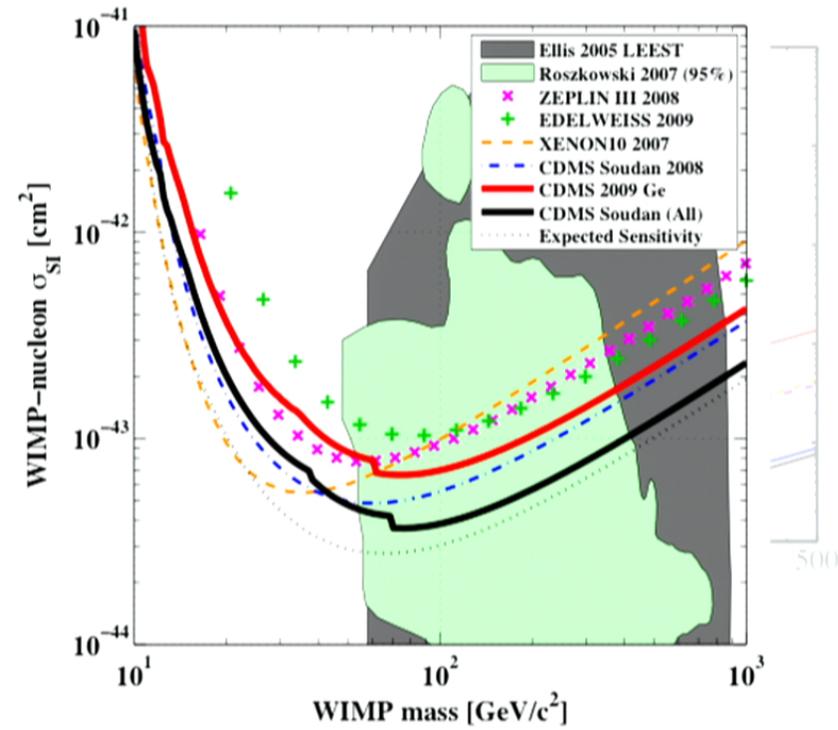
Miracle within orders of magnitude!



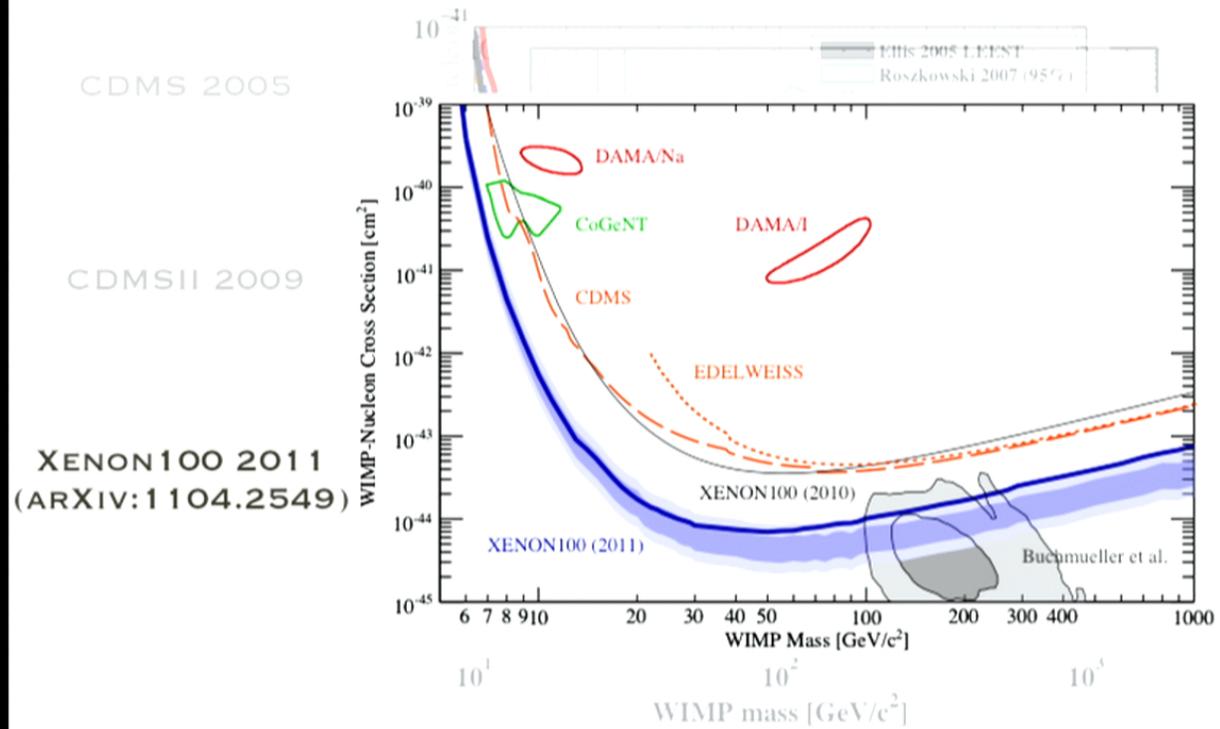
# Increased Tension w/ Direct Detection Experiment

CDMS 2005

CDMSII 2009

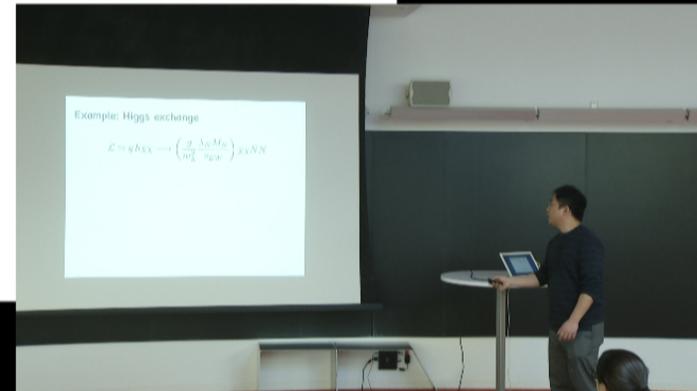


# Increased Tension w/ Direct Detection Experiment



## Example: Higgs exchange

$$\mathcal{L} \approx g h \bar{\chi} \chi \longrightarrow \left( \frac{g}{m_h^2} \frac{\lambda_N M_N}{v_{EW}} \right) \bar{\chi} \chi \bar{N} N$$



## Example: Higgs exchange

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$$\sigma_{SI} \sim \frac{4g^2}{\pi} \lambda_N^2 \frac{\mu^2 M_N^2}{m_h^4 v_{EW}^2} \approx \boxed{10^{-42} - 10^{-41} \text{ cm}^2}$$

**WAY ABOVE XENON100 BOUND**

$$\sigma_{SI} \sim 7.0 \times 10^{-45} \text{ cm}^2$$

## Additional Motivation

Interesting extension of MSSM that lead to "buried higgs" type models.

Bellazzini, Csaba, Falkowski, Weiler

- additional global symmetry broken at

$f \sim 500\text{GeV}$

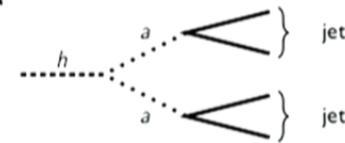
- PGB  $\sim O(10)\text{GeV}$ , couple derivatively to Higgs

$$\frac{1}{f^2} h^2 \partial_\mu a \partial^\mu a$$

- Light Higgs dominantly decay to jets, and

buried in QCD background; Higgs mass

possibly below LEP bound



Could the fermionic partner of PGB be the LSP?

## Goldstone multiplet in SUSY

- Add a global U(1) broken in the SUSY limit
- simplest example:  $W = S (N\bar{N} - f^2)$
- Goldstone Boson ( $\tilde{\text{axion}}$ )
  - Goldstone Fermion ( $\tilde{\text{axino}}$ )
  - S-Goldstone ( $\tilde{\text{saxion}}$ )

$$A = \frac{1}{\sqrt{2}}(s + ia) + \sqrt{2}\theta\chi + \theta^2 F$$

## Low energy effective theory

$$A = \frac{1}{\sqrt{2}}(s + ia) + \sqrt{2}\theta\chi + \theta^2 F$$

S-GOLDSTONE

GOLDSTONE BOSON

GOLDSTONE FERMION

## Low energy effective theory

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S-GOLDSTONE

GOLDSTONE BOSON

GOLDSTONE FERMION

$$A \longrightarrow A + i \text{const}$$

U(1)-SYMMETRY

INVARIANT:  $(A + A^\dagger)$

KAHLER POTENTIAL

$$K = K(A + A^\dagger)$$

NO SUPERPOTENTIAL

$$W = 0$$

## Simplest Example

$$W = S(\bar{N}N - f^2)$$

$$K = \bar{N}^\dagger \bar{N} + N^\dagger N$$

THEORY-UV

↓ integrate out heavy  
modes with mass  $\sim f$

$$K = \cosh(A + A^\dagger)$$

$$W = 0$$

THEORY-IR

$$N \sim f e^{A/f}$$

$$\bar{N} \sim f e^{-A/f}$$

FIELD CONTENT-IR

MORE GENERALLY

$$\Phi_i = f_i e^{q_i A/f}$$

$$f^2 = \sum_i q_i^2 f_i^2$$

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# Supersymmetric Non-linear sigma model

B. Zumino, Phys. Lett. B 87 (1979) 203.

$$\mathcal{L} = K''(s) \left( \frac{1}{2} \partial^\mu s \partial_\mu s + \frac{1}{2} \partial^\mu a \partial_\mu a + \frac{i}{2} \chi^\dagger \bar{\sigma}^\mu \partial_\mu \chi - \frac{i}{2} \partial_\mu \chi^\dagger \bar{\sigma}^\mu \chi \right) - \frac{1}{\sqrt{2}} K'''(s) (\chi^\dagger \bar{\sigma}^\mu \chi \partial_\mu a) + \frac{1}{4} \left( K''''(s) - \frac{K''''^2(s)}{K''(s)} \right) (\chi \chi) (\chi^\dagger \chi^\dagger)$$

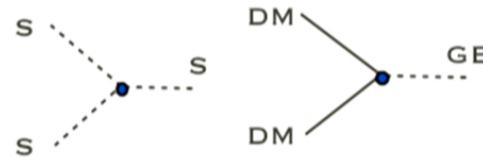
$$\frac{\partial^2 K}{\partial A \partial A^\dagger} = 1 + b_1 \frac{q}{f} (A + A^\dagger) + \dots$$



Goldstone-current interaction

$$\frac{1}{2\sqrt{2}} \left( b_1 \frac{1}{f} + o(s/f^2) \right) (\bar{\chi} \gamma^\mu \gamma^5 \chi) \partial_\mu a$$

$$b_1 = \frac{1}{qf^2} \sum_i q_i^3 f_i^2$$



4-fermion interaction

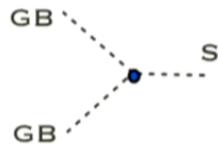
$$-\frac{1}{16f^2} (b_1^2 + \dots) [(\bar{\chi} \chi)^2 - (\bar{\chi} \gamma^5 \chi)^2]$$

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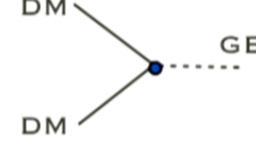
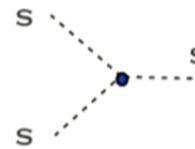
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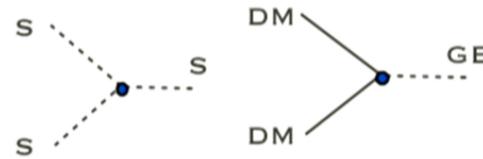
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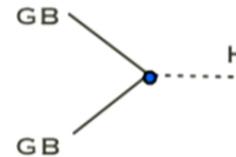
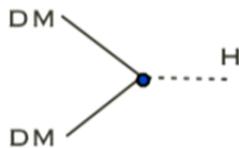
$$-\frac{1}{16f^2} (b_1^2 + \dots) [(\bar{\chi} \chi)^2 - (\bar{\chi} \gamma^5 \chi)^2]$$

## Scalar Mixing with Higgs

$$K = K(A + A^\dagger, \Phi_{SM})$$

$$K \subset \frac{1}{f}(A + A^\dagger)H_u H_d + \dots \rightarrow \frac{v}{f}(\partial s \partial h)$$

$$\left(1 + b_1 \frac{\sqrt{2}}{f}s + c_h \frac{v}{f^2}h\right) \left[\frac{1}{2}\partial^\mu s \partial_\mu s + \frac{1}{2}\partial^\mu a \partial_\mu a + \frac{i}{2}\bar{\chi}\gamma^\mu \partial_\mu \chi\right]$$



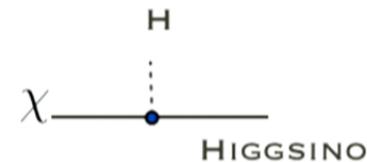
## Fermion kinetic mixing

$$\mathcal{L}_{\text{KM}} = \frac{i}{2f} \left[ \left( \chi^\dagger \bar{\sigma}^\mu \partial_\mu \tilde{H}_u - \partial_\mu \chi^\dagger \bar{\sigma}^\mu \tilde{H}_u \right) (c_1 H_d + \dots) + \text{h.c.} \right]$$

$$\rightarrow i\epsilon_u \chi^\dagger \bar{\sigma}^\mu \partial_\mu \tilde{H}_u^0 + i\epsilon_d \chi^\dagger \bar{\sigma}^\mu \partial_\mu \tilde{H}_d^0 + \text{h.c.}$$

$$\epsilon_{u,d} \sim v_{\text{EW}}/f$$

canonical normalizing kinetic term  
leads to a mixing of GF and Higgsino



small Higgsino component when  $\mu$  large

$$\epsilon_{u,d} m_\chi / \mu \sim v_{\text{EW}} m_\chi / f \mu.$$

In our scenario, we restrict us to  
this case, and ignore this mixing

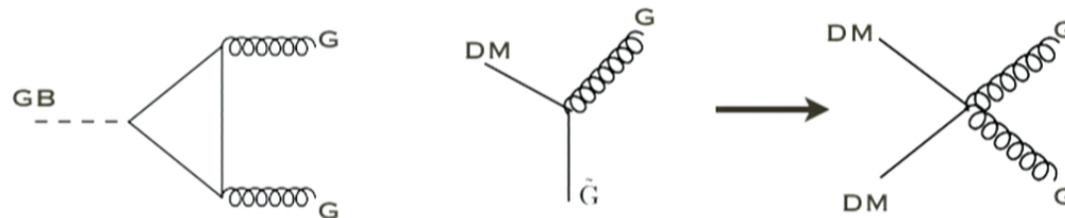
## Anomaly induced interactions

$$V_{an} = \frac{c_{an}}{\sqrt{2}f} a G_{\mu\nu}^a \tilde{G}_{\mu\nu}^a \quad W_{an} = -\frac{c_{an}}{f} A W^a W^a$$

$$\mathcal{L}_{anomaly} \subset \frac{c_{an}}{f\sqrt{2}} \left( a G_{\mu\nu}^a \tilde{G}_{\mu\nu}^a + i \bar{\chi} G_{\mu\nu}^a \sigma^{\mu\nu} \gamma^5 \lambda^a \right)$$

$$\mathcal{L}_y = ia \sum_{i=1}^{N_\Psi} y_i \bar{\Psi}_i \gamma^5 \Psi_i$$

If all fermions masses equal  $c_{an} = \frac{\alpha}{8\pi} q_\Psi N_\Psi$



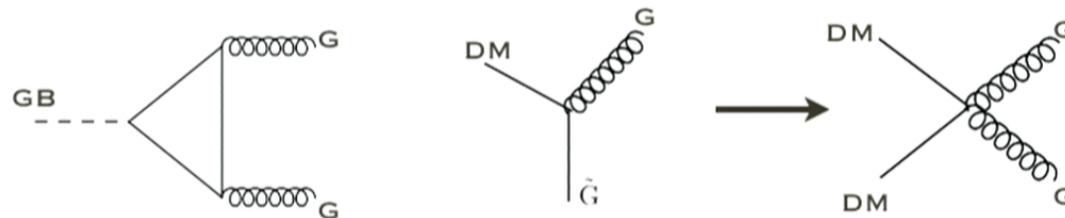
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## The Story So far ...

U(1)-broken spont.

$$W = S(\bar{N}N - f^2)$$

## Summary of interactions

$$W = S(\bar{N}N - f^2) + SH_uH_d + N\Phi_1\Phi_2$$

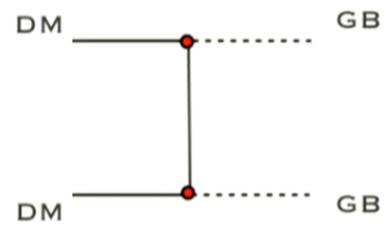
$$\frac{b_1}{f} \times \left( \begin{array}{c} \text{DM} \quad \text{GB} \quad \text{S} \quad \text{S} \quad \text{DM} \\ \diagdown \quad \text{---} \quad \diagup \quad \text{---} \quad \diagdown \\ \text{DM} \quad \text{GB} \quad \text{S} \quad \text{S} \quad \text{DM} \\ \diagup \quad \text{---} \quad \diagdown \quad \text{---} \quad \diagup \\ \text{DM} \quad \text{GB} \quad \text{S} \quad \text{S} \quad \text{DM} \end{array} \right)$$

$$\frac{v}{f^2} \times \left( \begin{array}{c} \text{DM} \quad \text{H} \\ \diagdown \quad \text{---} \\ \text{DM} \quad \text{H} \\ \diagup \\ \text{GB} \quad \text{H} \\ \diagdown \quad \text{---} \\ \text{GB} \quad \text{H} \\ \diagup \end{array} \right)$$

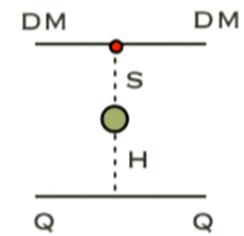
$$\frac{\alpha_s N}{8\pi f} \times \left( \begin{array}{c} \text{GB} \quad \text{G} \\ \diagdown \quad \text{---} \quad \diagup \\ \text{G} \end{array} \right) \quad \left( \frac{\alpha_s N}{8\pi f} \right)^2 \times \frac{1}{M_\lambda}$$

# General Feature

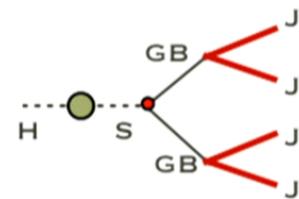
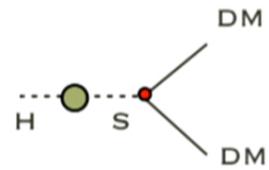
## DM-ANNIHILATION



## DIRECT DETECTION



## NON-STANDARD HIGGS DECAYS



## SUSY Breaking

- Until now entire A massless, but SUSY breaking can give mass to s and  $\chi$
- Add soft SUSY breaking, assume all globally symmetric

- Goldstone Fermion  $\chi = \sum_i \frac{q_i f_i}{f} \psi_i$

U(1) invariance of W  $\sum_j W_{ij} q_j f_j = -W_i q_i = -F_i q_i$   
 $\rightarrow m_\chi \approx q_i \langle F_i \rangle / f.$

Tamvakis-Wyler: If  $F_i$  vanish for all U(1) charged fields,  $\chi$  remain massless

## SUSY Breaking

- however SUSY breaking globally symmetric does NOT imply  $q_i F_i = 0$

- reason: mixing among F-terms  $F_X^\dagger F_i$

- assume spurion  $\langle X \rangle = x + F_X \theta^2$

- For example  $K \sim \frac{X^\dagger X}{M^2} \phi_i^\dagger \phi_i \longrightarrow \frac{x f_i}{M^2} F_X^* F_i$

$$F_i \rightarrow F_i^0 + \frac{x f_i}{M^2} F_X \longrightarrow m_\chi = F_i q_i / f \sim \frac{x}{M} m_i$$



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$\downarrow$   
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$\downarrow$   
 0

# Mass Spectrum

GOLDSTONE BOSON  $m_a = 0$  UP TO U(1)-BREAKING TERMS

GOLDSTONE FERMION

$m_\chi \ll m_i$  CAN BE ACHIEVED  $\rightarrow$  LSP  
DEPEND ON DETAILS OF SUSY BREAKING

S-GOLDSTONE  $m_s \approx m_i$

$$m_a < m_\chi \ll m_s$$

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$$m_a < m_\chi \ll m_s$$

## Parameters

Parameter	Description	Scan Range
$f$	Global symmetry breaking scale	500 GeV – 1.2 TeV
$m_\chi$	Goldstone fermion mass	50 – 150 GeV
$m_a$	Goldstone boson mass	8 GeV – $f/10$
$b_1$	$\chi\chi a$ coupling	[0, 2]
$c_{an}$	Anomaly coefficient	0.06
$c_h$	Higgs coupling	[-1, 1]
$\delta$	Explicit breaking $ia\bar{\chi}\gamma^5\chi$ coupling	3/2

$$\mathcal{L} \supset \left[ \frac{1}{2}(\partial a)^2 + \frac{1}{2}\bar{\chi}\not{\partial}\chi \right] c_h \frac{v}{f} h + \frac{b_1}{2\sqrt{2}f} (\bar{\chi}\gamma^\mu\gamma^5\chi) \partial_\mu a + \frac{c_{an}}{f\sqrt{2}} a G\tilde{G} + i\delta a\bar{\chi}\gamma^5\chi$$

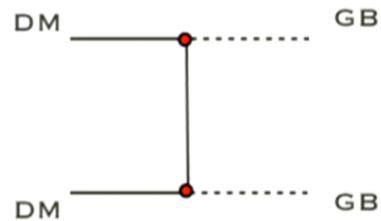
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## DM annihilation

DM-ANNIHILATION



$$\left( b_1 \frac{1}{2\sqrt{2}f} \right) (\bar{\chi} \gamma^\mu \gamma^5 \chi) \partial_\mu a = b_1 \frac{m_\chi}{f} (\bar{\chi} \gamma^5 \chi) a$$

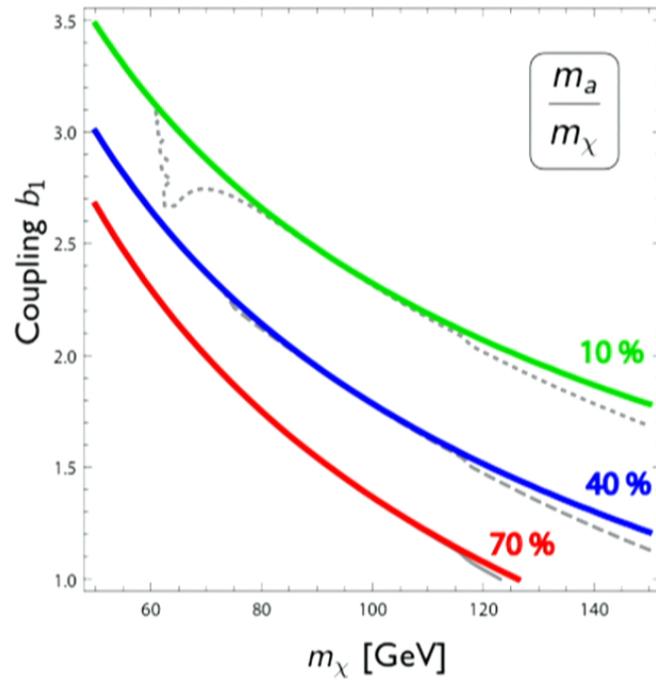
LEADING

$$\langle \sigma v \rangle = a + b \langle v^2 \rangle = \left( \frac{1}{8\pi} b_1^4 \frac{m_\chi^2}{f^4} \right) \frac{T_f}{m_\chi} \approx \text{pb}$$

P-WAVE

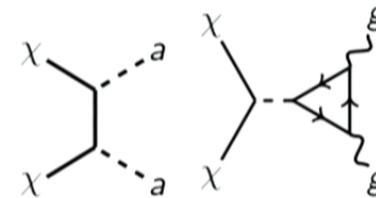
# DM annihilation

$$\Omega h^2 = 0.11$$



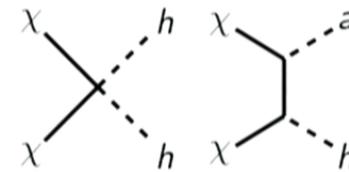
## Dominant contribution

Kähler, anomaly,  $U(1)$



## Subleading

Mixing with Higgs



## Negligible

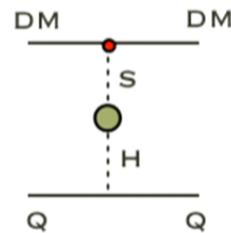
$\chi\chi \rightarrow s \rightarrow aa$ ,  $\chi\chi \xrightarrow{t,u} hh$

## Direct Detection

operator	coefficient	origin
$\bar{\chi}\chi G_{\mu\nu}^a G_{\mu\nu}^a$	$\frac{1}{f^2 M_\lambda}$	anomaly
$\bar{\chi}\chi \bar{q}q$	$\frac{y_q}{m_h^2} \times \left( \frac{m_\chi v_{EW}}{f^2} \right)$	mixing

## HIGGS EXCHANGE

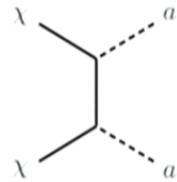
DIRECT DETECTION



$$\sigma_{SI} \approx \sigma_{SI}^{MSSM} \times \left( \frac{m_\chi v_{EW}}{f^2} \right)^2$$

$$10^{-46} \quad 10^{-42} \quad \times \quad 10^{-4}$$

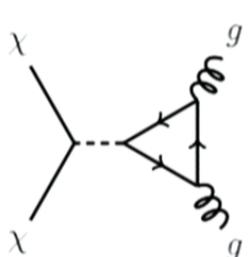
# Indirect Detection: FERMI



LEADING P-WAVE

$$\langle v\sigma \rangle_{\text{Th}} = 3 \times 10^{-26} \text{ cm}^3/\text{s}$$

TODAY SUPPRESSED BY VELOCITY  $v \simeq 10^{-3}$



SUBLEADING S-WAVE

$$0.2 - 0.3 \times \langle v\sigma \rangle_{\text{Th}} \times 10^{-3}$$

$$\frac{\Gamma(\gamma\gamma)}{\Gamma(gg)}$$

$\gamma$  - ray **LINES**  
(30-200 GEV)

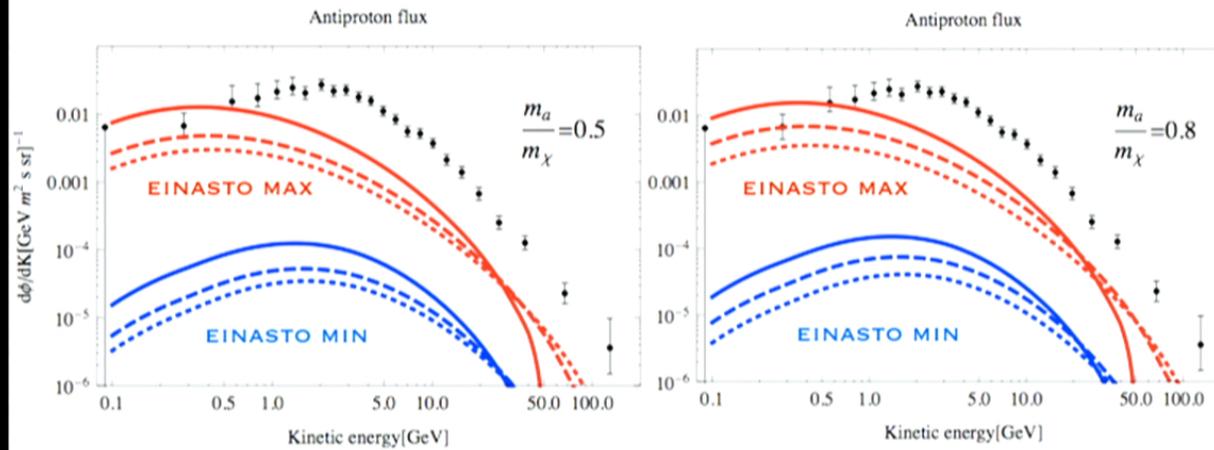
$$\mathcal{O}(10) \times \langle v\sigma \rangle_{\text{Th}}$$

$\gamma$  - ray **DIFFUSE**  
(20-100 GEV)

$$\chi\chi \rightarrow \pi^0, \dots$$

$\mathcal{O}(10)$  smaller

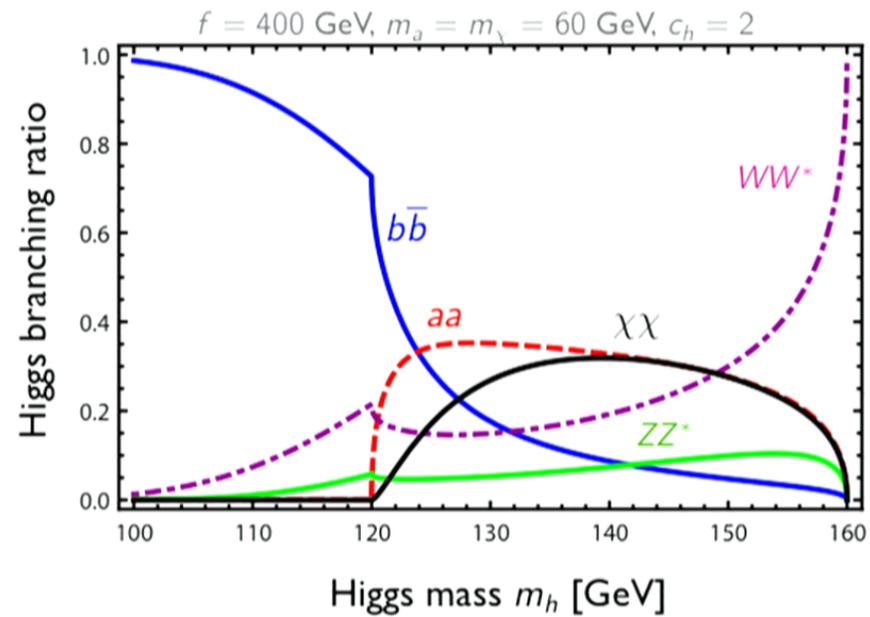
# Indirect Detection: PAMELA



WE USED THE PPC 4 DM ID: "POOR PARTICLE PHYSICIST COOKBOOK FOR DARK MATTER INDIRECT DETECTION"  
DEVELOPED BY M. CIRELLI ET AL. ARXIV 1012.4515

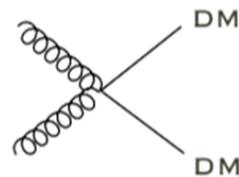
## Non-standard Higgs Decay

partial buried & invisible: suppressed SM channels, MET,  
Width < 1



# Goldstone Fermion @ LHC

Direct production through gluons  $\rightarrow$  ISR Monojets



operator	coefficient	origin	overall scale
$\bar{\chi}\chi G_{\mu\nu}^a G_{\mu\nu}^a$	$\frac{\alpha_s^2}{64\pi^2 f^2 M_\lambda}$	anomaly	$\frac{\alpha_s}{8M_*^3}$
$\bar{\chi}\chi \bar{q}q$	$\frac{y_q}{m_h^2} \times \left(\frac{m_\chi v_{EW}}{f^2}\right)$	mixing	$\frac{m_q}{2M_*^3}$

$gg \rightarrow a^* \rightarrow \chi\chi$  may be within  $5\sigma$  reach with  $100 \text{ fb}^{-1}$

1005.1286, 1005.3797, 1008.1783, 1103.0240, 1108.1196

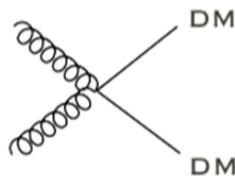
**Cascade decays**, LOSP  $\rightarrow \chi$ :  $\bar{\chi}G\lambda$  anomaly and  $\chi-\tilde{H}$  kinetic mixing

Decays typically prompt, a reconstruction is difficult for light masses.

Heavy fermions  $\Psi$  in anomaly may appear as “fourth generation” quarks

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## Conclusion

- minimal SUSY DM squeezed by direct detection bounds
- simple extension: add a  $U(1)$  and break it spontaneously in susy
  - like an axino at the TeV scale
  - GF could be the LSP and ordinary WIMP
  - annihilation rate dominated by
  - direct detection naturally below Xenon100 but not too far via Higgs exchange

## Scalar Mixing with Higgs

$$K = K(A + A^\dagger, \Phi_{SM})$$

$$K \subset \frac{1}{f}(A + A^\dagger)H_u H_d + \dots \rightarrow \frac{v}{f}(\partial s \partial h)$$

$$\left(1 + b_1 \frac{\sqrt{2}}{f} s + c_h \frac{v}{f^2} h\right) \left[ \frac{1}{2} \partial^\mu s \partial_\mu s + \frac{1}{2} \partial^\mu a \partial_\mu a + \frac{i}{2} \bar{\chi} \gamma^\mu \partial_\mu \chi \right]$$

