

Title: Fluctuations and Viscosity

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Abstract: Sound waves with long-distance propagation are both a consequence of hydrodynamics, and a danger to hydrodynamics' very existence, as they violate the assumption of local equilibration. In the talk, I will discuss what the thermally excited sound and shear waves do to viscosity. In 2+1 dimensions, the shear viscosity and the diffusion constant cease being independent transport coefficients. In 3+1 dimensions, the fluctuations render the second-order hydrodynamics invalid.



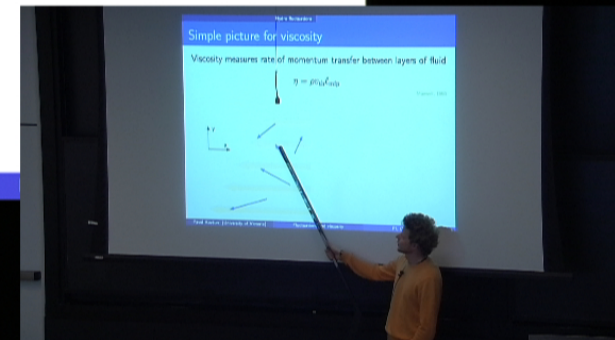
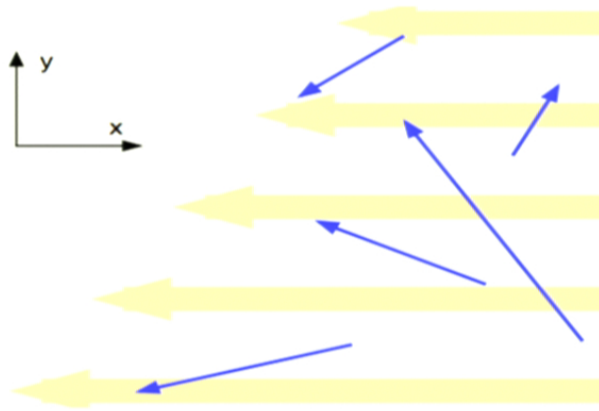


## Simple picture for viscosity

Viscosity measures rate of momentum transfer between layers of fluid

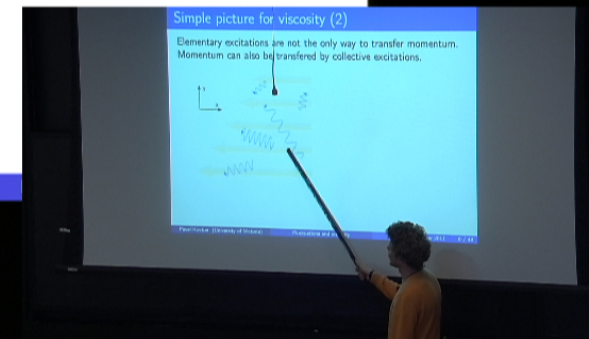
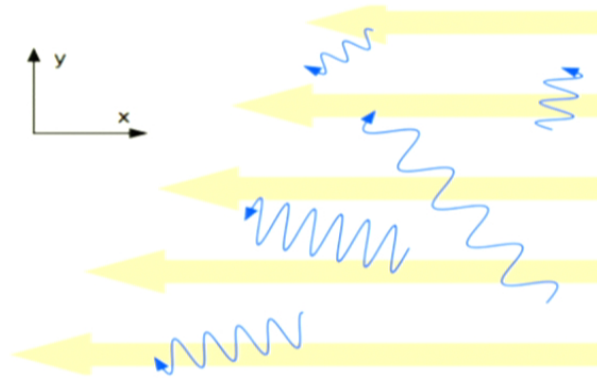
$$\eta = \rho v_{\text{th}} \ell_{\text{mfp}}$$

Maxwell, 1860



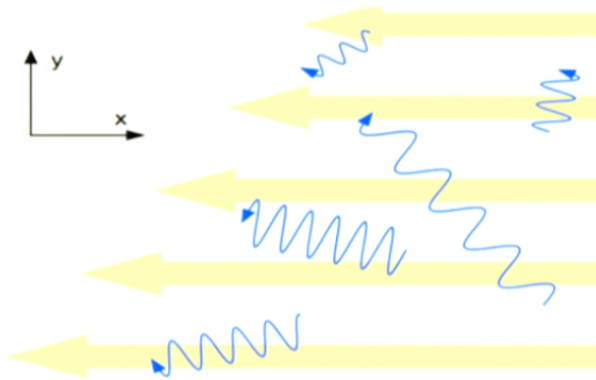
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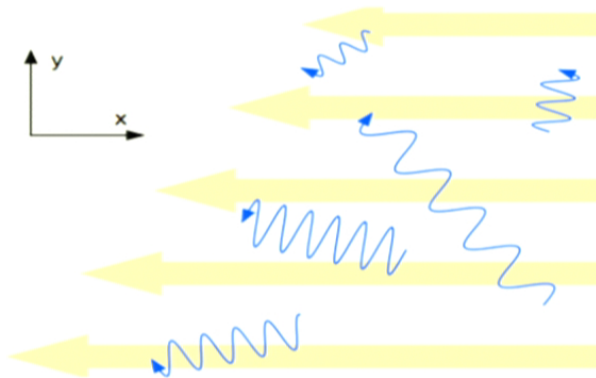
$$\ell_{\text{mfp}} \sim \frac{1}{\frac{\eta}{\epsilon + P} \mathbf{k}^2}$$

$$\eta_1 \sim \int^{k_{\text{max}}} d^3k \frac{T}{\frac{\eta_0}{\epsilon + P} \mathbf{k}^2} \sim \frac{k_{\text{max}} T^2}{\eta_0 / s}$$



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1. Total viscosity  $\eta_{\text{total}} = \eta_0 + \eta_1$  is bounded from below
2. This integral IR finite in  $d = 3+1$ , but IR divergent in  $d = 2+1$

Forster+Nelson+Stephen, 1977

## Interaction of hydro modes

- In hydro, there are no arbitrary “coupling constants” like  $g$
- Coefficients of non-linear terms are fixed by symmetry (Galilean or Lorentz) E.g.

$$J^\mu = nu^\mu + \nu^\mu, \quad T^{\mu\nu} = (\epsilon + P)u^\mu u^\nu + P\eta^{\mu\nu} + \tau^{\mu\nu}.$$

All transport coefs  $\eta, \zeta, \kappa$  are present already in linearized hydro

- Interaction of modes will change hydro correlation functions
- Was known since late 1960's – “mode-mode coupling”

## Long-time tails

Start with  $\mathbf{J} = -D\nabla n + n\mathbf{v}$ , take  $\mathbf{k} = 0$ . Schematically:

$$\begin{aligned}
 \langle \mathbf{J}(t)\mathbf{J}(0) \rangle &\supset \int d^d x \langle n(t, \mathbf{x})\mathbf{v}(t, \mathbf{x})n(0)\mathbf{v}(0) \rangle \\
 &= \int d^d x \langle n(t, \mathbf{x})n(0) \rangle \langle \mathbf{v}(t, \mathbf{x})\mathbf{v}(0) \rangle \\
 &\sim \int d^d k e^{-D\mathbf{k}^2 t} e^{-\gamma_\eta \mathbf{k}^2 t} \\
 &\sim \left[ \frac{1}{(D + \gamma_\eta)t} \right]^{d/2}
 \end{aligned}$$

See e.g. Arnold+Yaffe, PRD 1997  
(known since late 1960's)



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When FT, the convective contribution to  $S(\omega)$  is

$$\begin{aligned} S(\omega) &\sim \omega^{1/2}, & d = 3 \\ S(\omega) &\sim \ln(\omega), & d = 2 \end{aligned}$$

## Correction to Kubo formulas

Recall Kubo formula for the diffusion constant:

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This was derived in linear response. With the non-linear terms:

$$D^{\text{full}} = \lim_{\omega \rightarrow 0} (D + \text{const } \omega^{1/2}) , \quad d = 3$$

$$D^{\text{full}} = \lim_{\omega \rightarrow 0} (D + \text{const } \ln(\omega)) , \quad d = 2$$

Same applies to shear viscosity:

$$\eta^{\text{full}} = \lim_{\omega \rightarrow 0} (\eta + \text{const } \omega^{1/2}) , \quad d = 3$$

$$\eta^{\text{full}} = \lim_{\omega \rightarrow 0} (\eta + \text{const } \ln(\omega)) , \quad d = 2$$



## Comment

- In AdS/CFT, the  $\ln(\omega)$  correction is  $1/N^{3/2}$  suppressed
- Transport coefficients come out finite in 3 + 1 dimensional *classical* gravity
- Long-time tails come from quantum corrections to classical gravity
- This is an example where long-time limit does not commute with large-N limit

Kovtun+Yaffe, 2003  
Caron-Huot + Saremi, 2009

## Can do the same calculation in momentum space



One-loop diagram with sound and/or shear waves in the loop

$$S_{xy,xy}(\omega, \mathbf{k}=0) = (\epsilon + P)^2 \int \frac{d\omega'}{2\pi} \frac{d^3k}{(2\pi)^3} \left( \Delta_{xx}(\omega', \mathbf{k}) \Delta_{yy}(\omega - \omega', -\mathbf{k}) + \Delta_{xy}(\omega', \mathbf{k}) \Delta_{yx}(\omega - \omega', -\mathbf{k}) \right)$$

where  $\Delta_{ij} = \text{FT of } \langle u_i(x) u_j(0) \rangle$



## When the dust settles...

$$G_{xy,xy}^R(\omega \ll k_{\max}, k=0) = -i\omega\eta_0 - i\omega \frac{17T k_{\max}}{120\pi^2 \gamma \eta_0} + (1+i)\omega^{3/2} \frac{(7 + (3/2)^{3/2})T}{2^{3/2} \pi \gamma \eta_0} + O((k_{\max} \gamma \eta_0)^2, \omega^2)$$

PK-Moore-Romatschke, 2011

The contribution due to hydro fluctuations is suppressed at either small coupling, or large N

Pavel Kovtun (University of Victoria)

Fluctuations and viscosity

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## Implications for the shear viscosity

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- Estimate  $k_{\max}\gamma_{\eta_0} \sim 1/2$ , then

$$\eta_{\text{total}}/s \gtrsim 0.16\hbar$$

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- Current hydro simulations of QGP are blind to these effects because they simply solve the classical hydro equations and ignore the fluctuations

## Exactly the same happens for second-order relativistic hydro (Israel-Stewart)

In linearized second order hydro:

$$G_{xy,xy}^R(\omega, \mathbf{k}) = P - i\omega\eta + \left(\eta\tau_{\Pi} - \frac{\kappa}{2}\right)\omega^2 - \frac{\kappa}{2}\mathbf{k}^2 + \dots$$

Baier+Romatschke+Son+Starinets+Stephanov, 2007



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But this gets seriously modified by 1-loop hydro fluctuations,

$$G_{xy,xy}^R(\omega, \mathbf{k}=0) = P - i\omega\eta - \text{const } |\omega|^{3/2}(1 + i \text{sign}(\omega)) + \dots$$



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Blindly apply Kubo formula

$$\eta\tau_{\Pi} - \frac{\kappa}{2} = \lim_{\omega \rightarrow 0} \frac{1}{2} \frac{\partial^2}{\partial \omega^2} \text{Re} G_{xy,xy}^R(\omega, \mathbf{k}=0) \rightarrow \infty$$

This means  $\tau_{\Pi}$  does not exist

## Can we save second-order hydro?

- Can estimate when  $\omega^{3/2}$  term becomes comparable to  $\omega^2$  term
- 2nd-order hydro breaks down below some  $\omega_*$  depends on  $\eta_0/s$
- If  $\eta_0/s \sim 0.16$ , then  $\omega_* \sim T/20$ ,  
2nd-order hydro OK for heavy-ion collisions
- If  $\eta_0/s \sim 0.08$ , then  $\omega_* \sim 2.5T$ ,  
2nd order hydro makes no sense for heavy-ion collisions

PK+Moore+Romatschke, 2011

## Is there hope for hydrodynamics?

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Is 3+1 hydro meaningless beyond first derivatives?

Hydro is not meaningless.  
Rather, viscosity, conductivity etc become scale-dependent  
“running masses” in the low-energy effective hydro theory

## Langevin equation

Brownian particle:

$$m \frac{d^2 x}{dt^2} = -(6\pi\eta a) \frac{dx}{dt} + f(t),$$

$(6\pi\eta a)$  = friction coefficient (Stokes law)

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Noise properties:

$$\langle \xi(t) \rangle = 0, \quad \langle \xi(t) \xi(t') \rangle = C \delta(t - t').$$

$C$  determines the strength of the noise

## Correlation function of $q(t)$

- Take the Langevin equation  $\dot{q}(t) + \gamma q(t) = \xi(t)$
- Solve for  $q(t)$  in terms of  $\xi(t)$
- Find  $\langle q(t)q(t') \rangle$  by averaging over  $\xi(t)$
- When  $\gamma t, \gamma t' \gg 1$ , find

$$\langle q(t)q(t') \rangle = \frac{C}{2\gamma} e^{-\gamma|t-t'|}$$

- Fourier transform:

$$S(\omega) = \frac{C}{\omega^2 + \gamma^2}$$

## Noise strength

Recall  $\langle \xi(t)\xi(t') \rangle = C\delta(t - t')$

What determines the noise strength  $C$ ?

- Assume thermal equilibrium
- Demand that the correlation functions satisfy the FDT:

$$\text{Im } G_R(\omega) = \frac{\omega}{2T} S(\omega)$$

- To find  $G_R$ , introduce source (external force)

$$\delta q(t) = \int dt' G_R(t-t') \delta f(t')$$

- Langevin equation gives  $G_R(\omega) = \frac{i}{\omega + i\gamma}$
- Demand FDT:

$$C = 2T$$



## Path integral for Brownian particle

Let us now represent the Brownian motion as Quantum Mechanics  
(0+1 dimensional quantum field theory)

## Path integral for Brownian particle

**Step 1** Write Langevin equation as  $EoM \equiv (\dot{q} + \frac{\partial F}{\partial q} - \xi) = 0$

**Step 2** Gaussian noise:

$$\langle \dots \rangle = \int \mathcal{D}\xi e^{-W[\xi]}(\dots), \quad \text{where } W[\xi] = \frac{1}{2C} \int dt' \xi(t')^2.$$

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**Step 3** Recall  $\delta(f(x)) \sim \delta(x-x_0)$ , where  $x_0$  solves  $f(x_0) = 0$ . So

$$\int \mathcal{D}q J \delta(EoM) q(t_1) q(t_2) \dots = \underbrace{q_\xi(t_1)}_{\text{satisfy } EoM(q, \xi) = 0} \underbrace{q_\xi(t_2)}_{\text{satisfy } EoM(q, \xi) = 0} \dots$$



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**Step 4** Write  $\delta(EoM) = \int \mathcal{D}p e^{i \int p EoM}$ , do the integral over  $\xi(t)$ .

## Path integral for Brownian particle (2)

When the dust settles:

$$\langle q(t_1) \dots q(t_n) \rangle = \int \mathcal{D}q \mathcal{D}p J e^{iS[q,p]} q(t_1) \dots q(t_n)$$

where

$$S[q, p] = \int dt \left( p\dot{q} + p \frac{\partial F}{\partial q} + \frac{iC}{2} p^2 \right) .$$

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For the simple Langevin equation  $F(q) = \frac{1}{2}\gamma q^2$ ,

$$S(\omega) = \text{FT of } \langle q(t)q(t') \rangle = \frac{C}{\omega^2 + \gamma^2} ,$$

as expected.



## Bottomline:

In the stochastic model

$$\dot{q}(t) + \underbrace{\frac{\partial F(q)}{\partial q}}_{\text{relaxation term}} = \underbrace{\xi(t)}_{\text{noise term}}$$

correlation functions can be derived from field theory with

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## Functional integral for hydro

Correlation functions in linearized hydro:

$$\langle \epsilon(\mathbf{x}, t) \pi_k(\mathbf{x}', t') \dots \rangle = \int \mathcal{D}\epsilon \mathcal{D}\boldsymbol{\pi} \mathcal{D}\eta \mathcal{D}\boldsymbol{\lambda} e^{iS} \epsilon(\mathbf{x}, t) \pi_k(\mathbf{x}', t') \dots$$

$$S = \int_{t, \mathbf{x}} \left( \eta \left( \frac{\partial \epsilon}{\partial t} + \nabla \cdot \boldsymbol{\pi} \right) + \lambda_i \left( \frac{\partial \pi_i}{\partial t} + v_s^2 \partial_i \epsilon - M_{ij} \pi_j \right) - i\bar{\omega} T \lambda_i M_{ij} \lambda_j \right)$$

Can integrate out the auxiliary field  $\boldsymbol{\lambda}$ :

$$S_{\text{eff}}[\epsilon, \boldsymbol{\pi}] = \frac{1}{2} \int_{t, \mathbf{x}, \mathbf{x}'} E_i(t, \mathbf{x}) D_{ij}(\mathbf{x}, \mathbf{x}') E_j(t, \mathbf{x}')$$

where  $E_i \equiv \left( \frac{\partial \pi_i}{\partial t} + v_s^2 \partial_i \epsilon - M_{ij} \pi_j \right)$ , and  $M_{ij} D_{jk} = -\frac{1}{2\bar{\omega} T} \delta(\mathbf{x} - \mathbf{x}') \delta_{ik}$

Note the action  $S_{\text{eff}}[\epsilon, \boldsymbol{\pi}]$  is time-reversal invariant, as it should be

**This effective action produces the correct hydro response functions**



## Correlation functions

Once know  $S_{\pi_i \pi_j}(\omega, \mathbf{k})$ , the others follow from energy conservation:

$$\omega S_{\epsilon \pi_i}(\omega, \mathbf{k}) = k_l S_{\pi_l \pi_i}(\omega, \mathbf{k}),$$

$$\omega S_{\epsilon \epsilon}(\omega, \mathbf{k}) = k_l S_{\pi_l \epsilon}(\omega, \mathbf{k}).$$

Can read off correlation functions from the effective action  $S_{\text{eff}}[\epsilon, \boldsymbol{\pi}]$ :

$$S_{\pi_i \pi_j}(\omega, \mathbf{k}) = \left( \delta_{ij} - \frac{k_i k_j}{\mathbf{k}^2} \right) \frac{2\gamma_\eta \bar{\omega} T \mathbf{k}^2}{\omega^2 + (\gamma_\eta \mathbf{k}^2)^2} + \frac{k_i k_j}{\mathbf{k}^2} \frac{2\gamma_s \bar{\omega} T \mathbf{k}^2 \omega^2}{(\omega^2 - v_s^2 \mathbf{k}^2)^2 + (\gamma_s \mathbf{k}^2 \omega)^2}$$

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## Stochastic model for linearized hydro

$$\frac{\partial \epsilon}{\partial t} = -\nabla \cdot \boldsymbol{\pi},$$

$$\frac{\partial \pi_i}{\partial t} = -v_s^2 \partial_i \epsilon + M_{ij} \pi_j + \xi_i(\mathbf{x}, t).$$

Dissipative terms:

$$M_{ij} \equiv \gamma_\eta (\nabla^2 \delta_{ij} - \partial_i \partial_j) + \gamma_s \partial_i \partial_j$$

Noise correlations:

$$\langle \xi_i(\mathbf{x}, t) \xi_j(\mathbf{x}', t') \rangle = -2\bar{w} T M_{ij} \delta(\mathbf{x} - \mathbf{x}') \delta(t - t')$$

Note the same  $M_{ij}$  must appear both in the hydro equations, and in the noise correlations



## Bottomline

- We have an effective action for linearized relativistic hydro
- This effective action is **not** meant to reproduce the classical hydro equations
- Rather, it is to be used to construct the generating functional for the correlation functions of  $T^{0\mu}(x)$ 
  - i)* at low energies
  - ii)* in real time
  - iii)* near thermal equilibrium

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Now that we know how to construct the effective action for linearized hydro, can look at the full non-linear hydrodynamics

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Can integrate out the auxiliary field  $\boldsymbol{\lambda}$ :

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## A simple toy model

- Incompressible fluid: impose  $\nabla \cdot \boldsymbol{\pi} = 0$
- Momentum conservation:

Forster+Nelson+Stephen, 1977

$$\partial_t \pi_i = -\partial_j T_{ij} + \xi_i, \quad T_{ij} = P\delta_{ij} - \gamma_\eta(\partial_i \pi_j + \partial_j \pi_i) + \frac{\pi_i \pi_j}{\bar{w}}$$

- Current conservation:

$$\partial_t n = -\partial_i J_i + \theta, \quad J_i = -D\partial_i n + \frac{n\pi_i}{\bar{w}}$$

- Stochastic model:

$$\partial_t \pi_i = -\partial_i P + \gamma_\eta \nabla^2 \pi_i - \frac{(\boldsymbol{\pi} \cdot \nabla) \pi_i}{\bar{w}} + \xi_i,$$

$$\partial_t n = D \nabla^2 n - \frac{(\boldsymbol{\pi} \cdot \nabla) n}{\bar{w}} + \theta,$$

## Effective action for the toy model

$$S_{\text{eff}} = \int dt d^d x \left( \mathcal{L}^{(2)} + \mathcal{L}^{(int)} \right)$$

$$\begin{aligned} \mathcal{L}^{(2)} = & -\frac{\sigma}{2} \rho \nabla^2 \rho - \frac{\tilde{\sigma}}{2} \lambda_i \nabla^2 \lambda_i - i \rho (\partial_t n - D \nabla^2 n) - i \lambda_i (\partial_t \pi_i - \Gamma \nabla^2 \pi_i) \\ & + \bar{\psi}_i (\partial_t - \Gamma \nabla^2) \psi_i + \bar{\psi}_n (\partial_t - D \nabla^2) \psi_n, \end{aligned}$$

$$\begin{aligned} \mathcal{L}^{(int)} = & -\frac{i}{w} \rho \pi_i \partial_i n - \frac{i}{w} \lambda_i \pi_j \partial_j \pi_i \\ & + \frac{1}{w} \bar{\psi}_i \partial_k \pi_i \psi_k + \frac{1}{w} \bar{\psi}_i \pi_k \partial_k \psi_i + \frac{1}{w} \bar{\psi}_n \partial_i n \psi_i + \frac{1}{w} \bar{\psi}_n \pi_k \partial_k \psi_n, \end{aligned}$$

plus the constraints  $\partial_i \pi_i = 0$ ,  $\partial_i \lambda_i = 0$ ,  $\partial_i \bar{\psi}_i = 0$ ,  $\partial_i \psi_i = 0$ .

The constants are  $\sigma = 2TD\chi$ ,  $\tilde{\sigma} = 2T\Gamma w$ ,  $\Gamma = \eta/w$ .

## One-loop correlation functions in the toy model

As  $\mathbf{k} \rightarrow 0$ :

$$\langle T_{0i} T_{0j} \rangle = \frac{2T\eta\Gamma(\omega)\mathbf{k}^2}{\omega^2 + \left(\Gamma(\omega)\mathbf{k}^2\right)^2}, \quad \langle J_0 J_0 \rangle = \frac{2T\chi D(\omega)\mathbf{k}^2}{\omega^2 + \left(D(\omega)\mathbf{k}^2\right)^2}.$$

This looks like the familiar linear response functions, except  $D$  and  $\eta$  now depend on  $\omega$ .

In  $d=3$  dimensions:

$$\Gamma(\omega) = \Gamma - \frac{23}{30\pi s} \frac{\sqrt{|\omega|}}{(4\Gamma)^{3/2}}, \quad D(\omega) = D - \frac{1}{3\pi s} \frac{\sqrt{|\omega|}}{[2(\Gamma+D)]^{3/2}}.$$

Conventional Kubo formulas make sense:

$$D = \frac{1}{2T\chi} \lim_{\omega \rightarrow 0} \lim_{\mathbf{k} \rightarrow 0} \frac{\omega^2}{\mathbf{k}^2} G_{nn}(\omega, \mathbf{k})$$



## One-loop correlation functions in the toy model

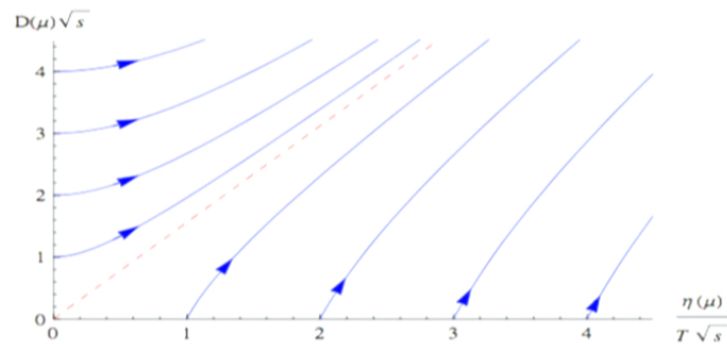
As  $\mathbf{k} \rightarrow 0$ :

$$\langle T_{0i} T_{0j} \rangle = \frac{2T\eta\Gamma(\omega)\mathbf{k}^2}{\omega^2 + \left(\Gamma(\omega)\mathbf{k}^2\right)^2}, \quad \langle J_0 J_0 \rangle = \frac{2T\chi D(\omega)\mathbf{k}^2}{\omega^2 + \left(D(\omega)\mathbf{k}^2\right)^2}.$$

This looks like the familiar linear response functions, except  $D$  and  $\eta$  now depend on  $\omega$ .

In  $d=2$  dimensions:

$$\Gamma(\omega) = \Gamma(\mu) + \frac{1}{32\pi s} \frac{1}{\Gamma(\mu)} \ln \frac{\mu}{\omega}, \quad D(\omega) = D(\mu) + \frac{1}{8\pi s} \frac{1}{\Gamma(\mu) + D(\mu)} \ln \frac{\mu}{\omega}$$

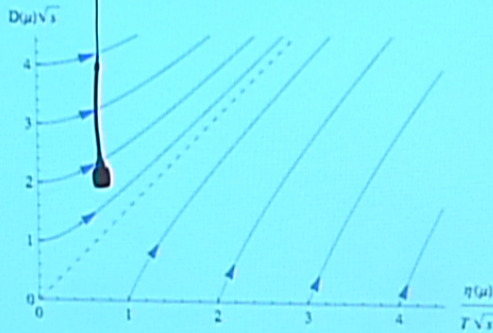
RG flow diagram in  $d=2$ 

In the extreme low-frequency limit  $\mu \rightarrow 0$ :

$$DT = \frac{\sqrt{17} - 1}{2} \frac{\eta}{s} \approx 1.56 \frac{\eta}{s}$$



## RG flow diagram in $d=2$



In the extreme low-frequency limit  $\mu \rightarrow 0$ :

$$DT = \frac{\sqrt{17} - 1}{2} \frac{\eta}{s} \approx 1.56 \frac{\eta}{s}$$

$D$  and  $\eta$  are not independent transport coefficients in extreme IR.



## Hydro fluctuations imply that

- $\frac{\eta}{s}$  is bounded from below in real-world QCD
- Second-order relativistic hydrodynamics strictly speaking does not exist
- However, 2nd order hydro still OK for heavy-ion collisions if  $\eta/s$  is sufficiently large
- Fluctuation effects disappear in the  $N \rightarrow \infty$  limit

## What I would like to understand

- I only showed the effective action for linearized hydro and the toy model. Can one find the covariant action for the full non-linear relativistic hydro? Work in progress with GM and PR!
- Effective action for hydro from AdS/CFT?
- Effective action for relativistic superfluids?
- How do transport coefficients in 2+1 dim flow at non-zero density?
- How do transport coefficients in 2+1 dim flow in external magnetic field?
- Other 2-nd order transport coefficients in relativistic hydro?