

Title: Dark Matter Halos in 6D

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Abstract: I will describe various efforts to understand the phase space structure of dark matter halos. Implications for dark matter detection experiments will also be discussed.

Dark Matter Halos in 6D

Unravelling Dark Matter
Perimeter Institute
September 2011

Larry Widrow
Queen's University

in collaboration with
David Stiff, Pascal Elahi, and Robert Thacker

Formation of dark matter halos within
the context of the standard CDM paradigm
is a relatively “clean” physics problem

The System:
collisionless particles
Newtonian gravity

Initial Conditions:
negligible velocity dispersion
linear, random perturbations
power spectrum given by “fundamental physics” (inflation)

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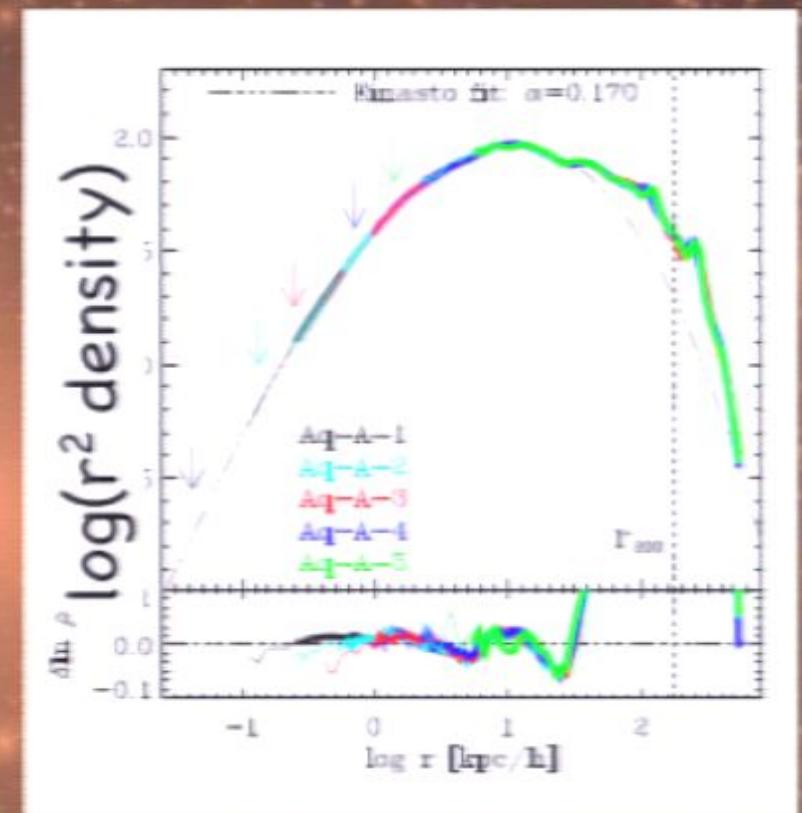
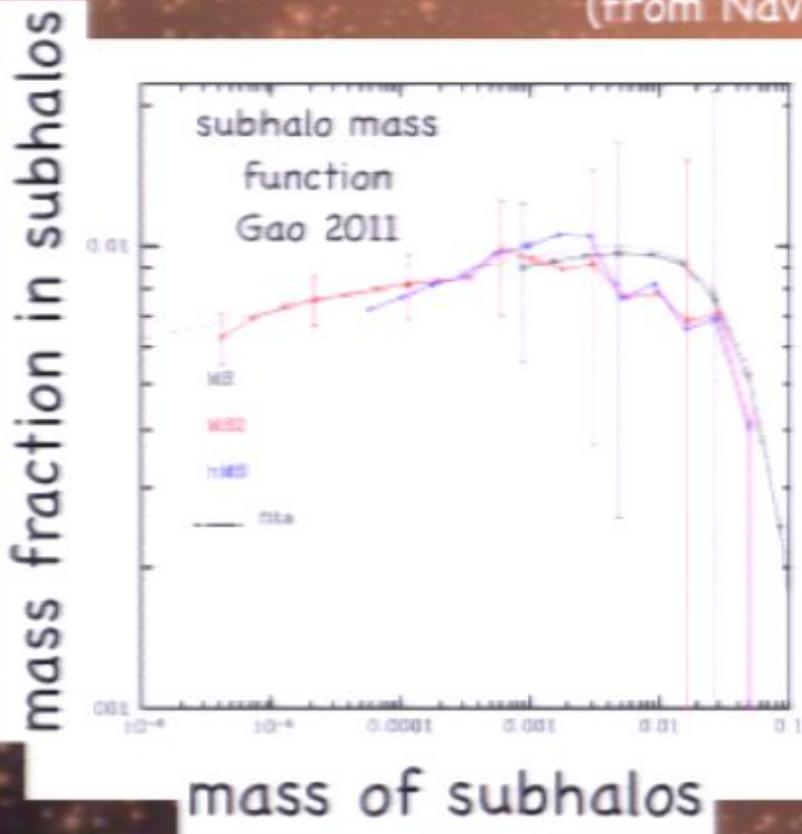
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dark halo phenomenology from simulations

spherically-averaged
density profile
(from Navarro, 2010)

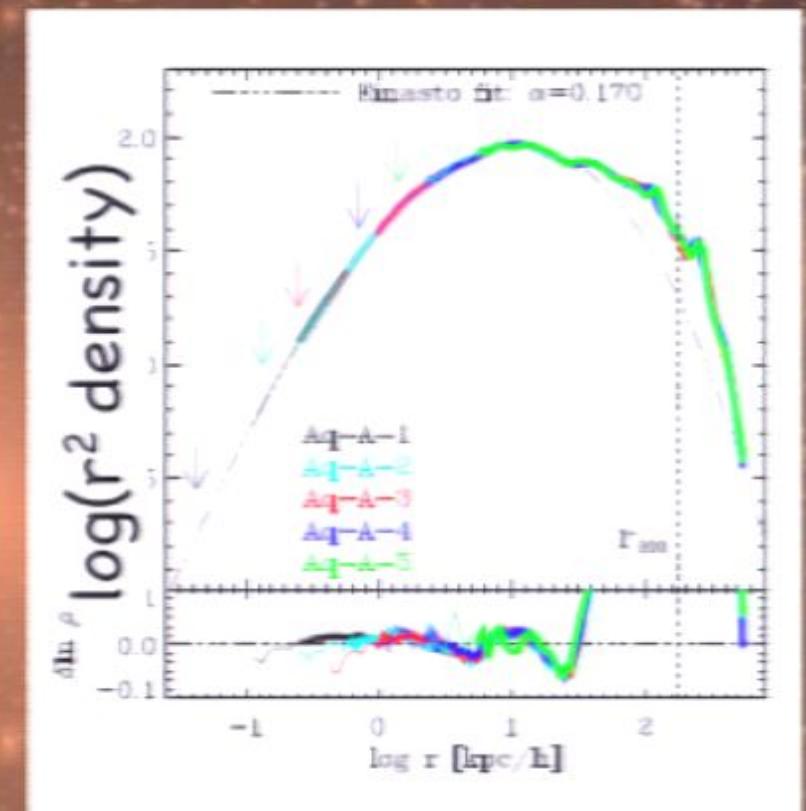
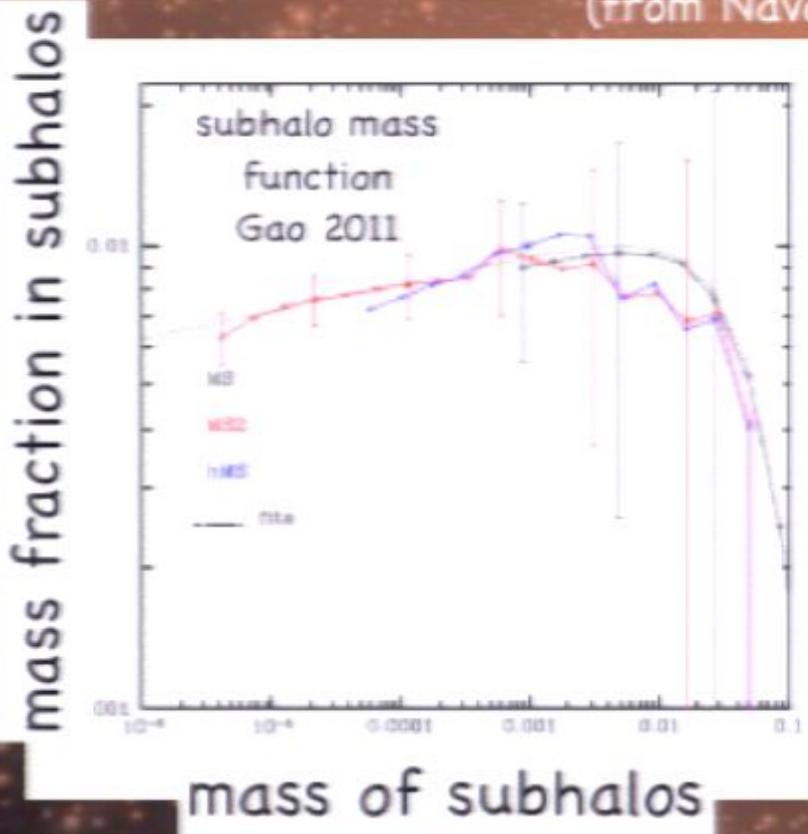


velocity anisotropy profile
pseudo-phase space profile
rotation
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figure from Diemand et al. 2006

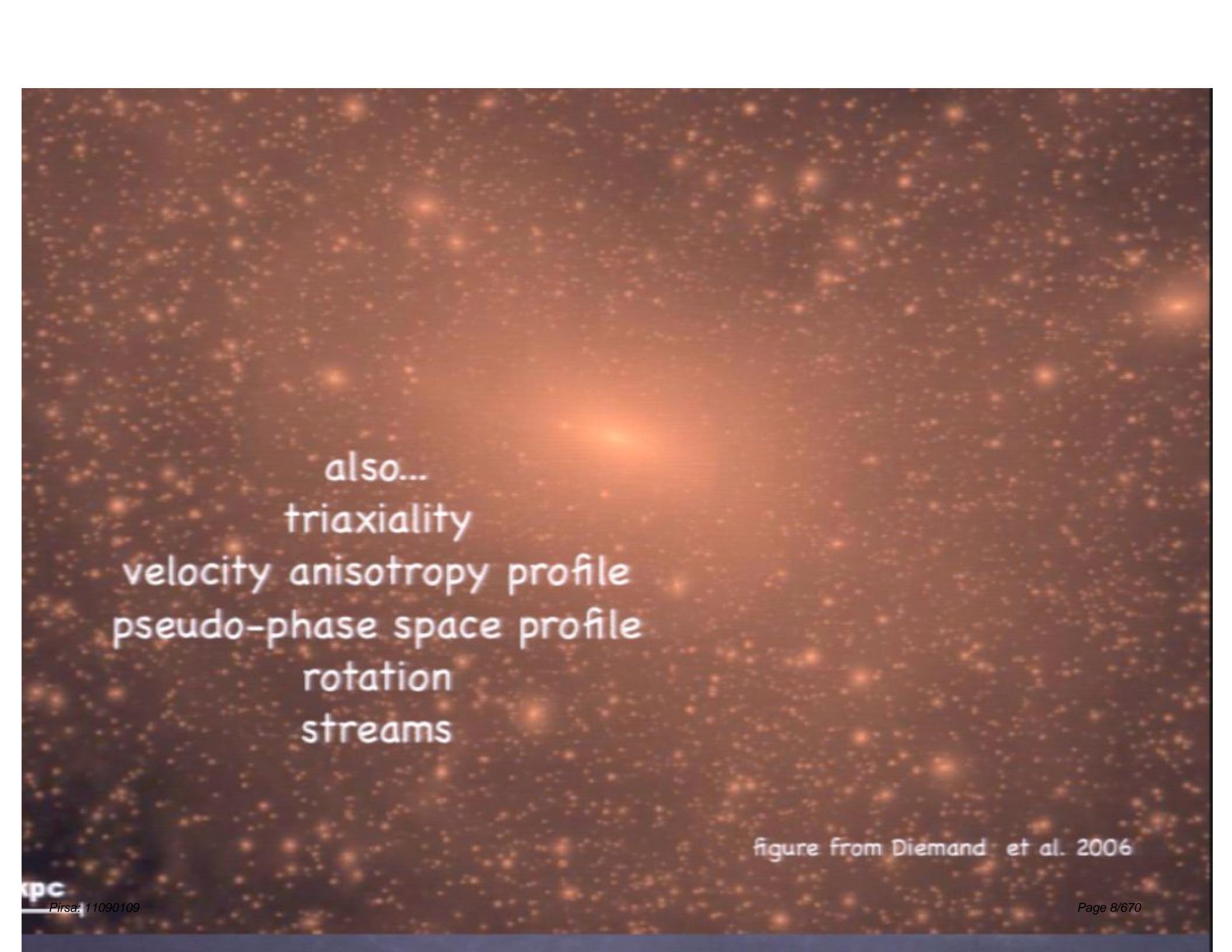
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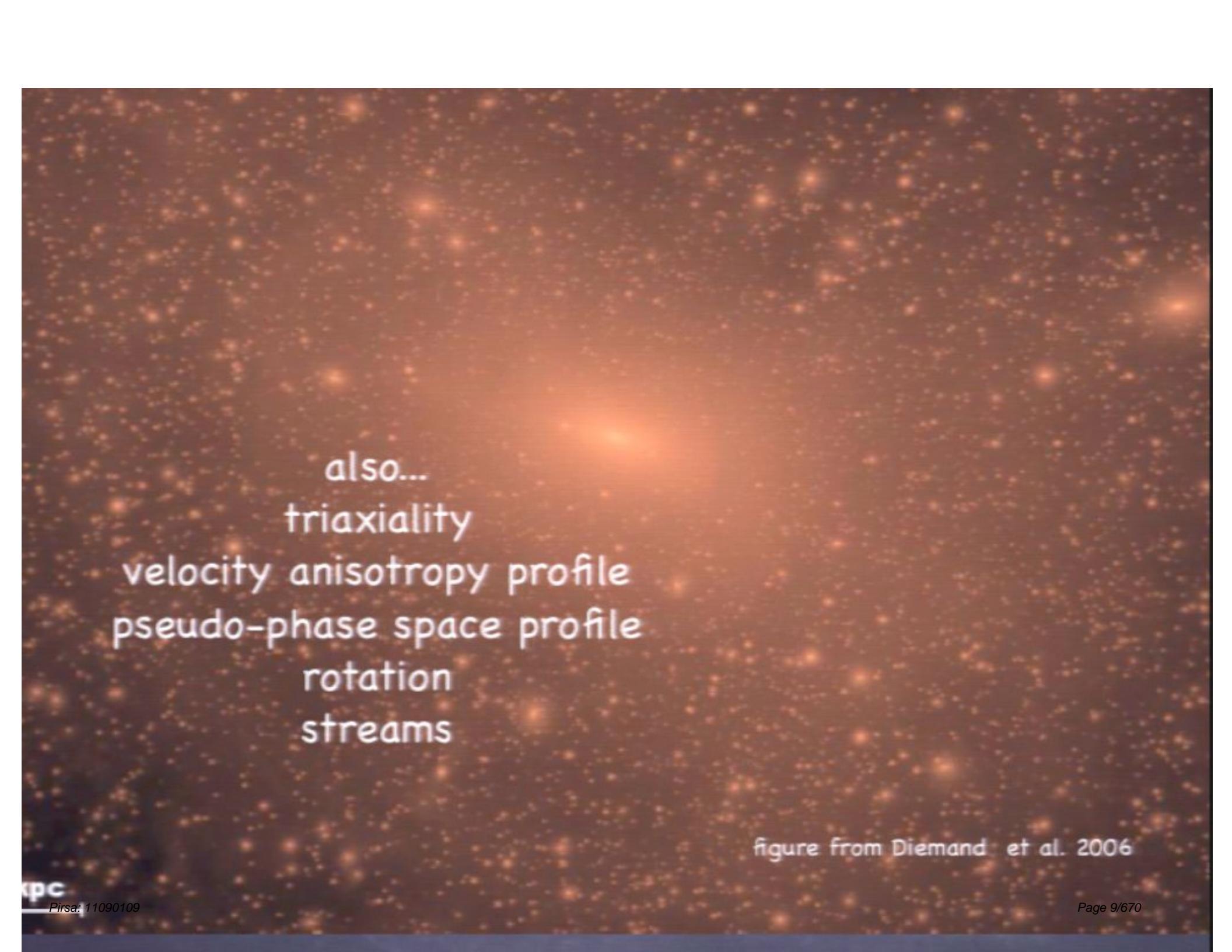
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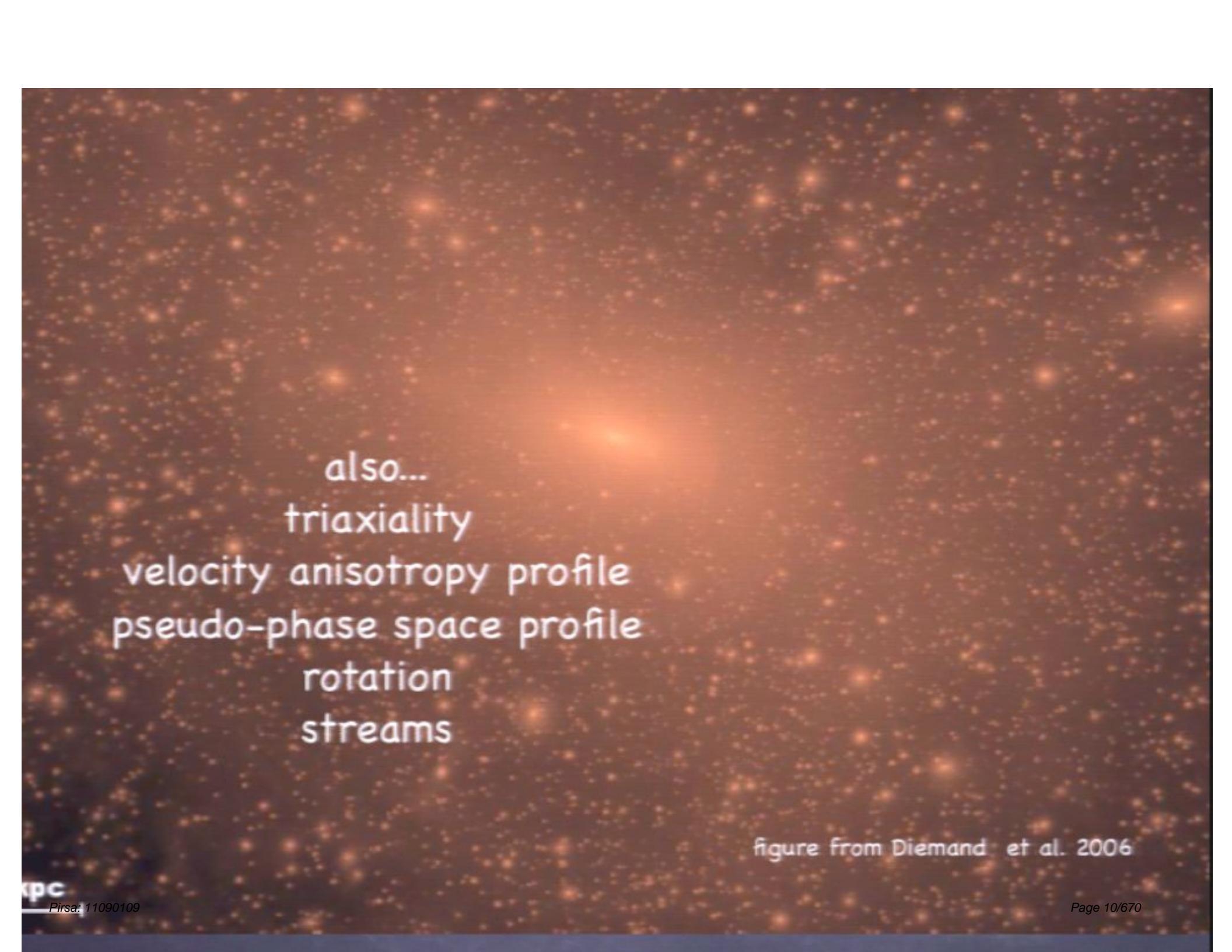
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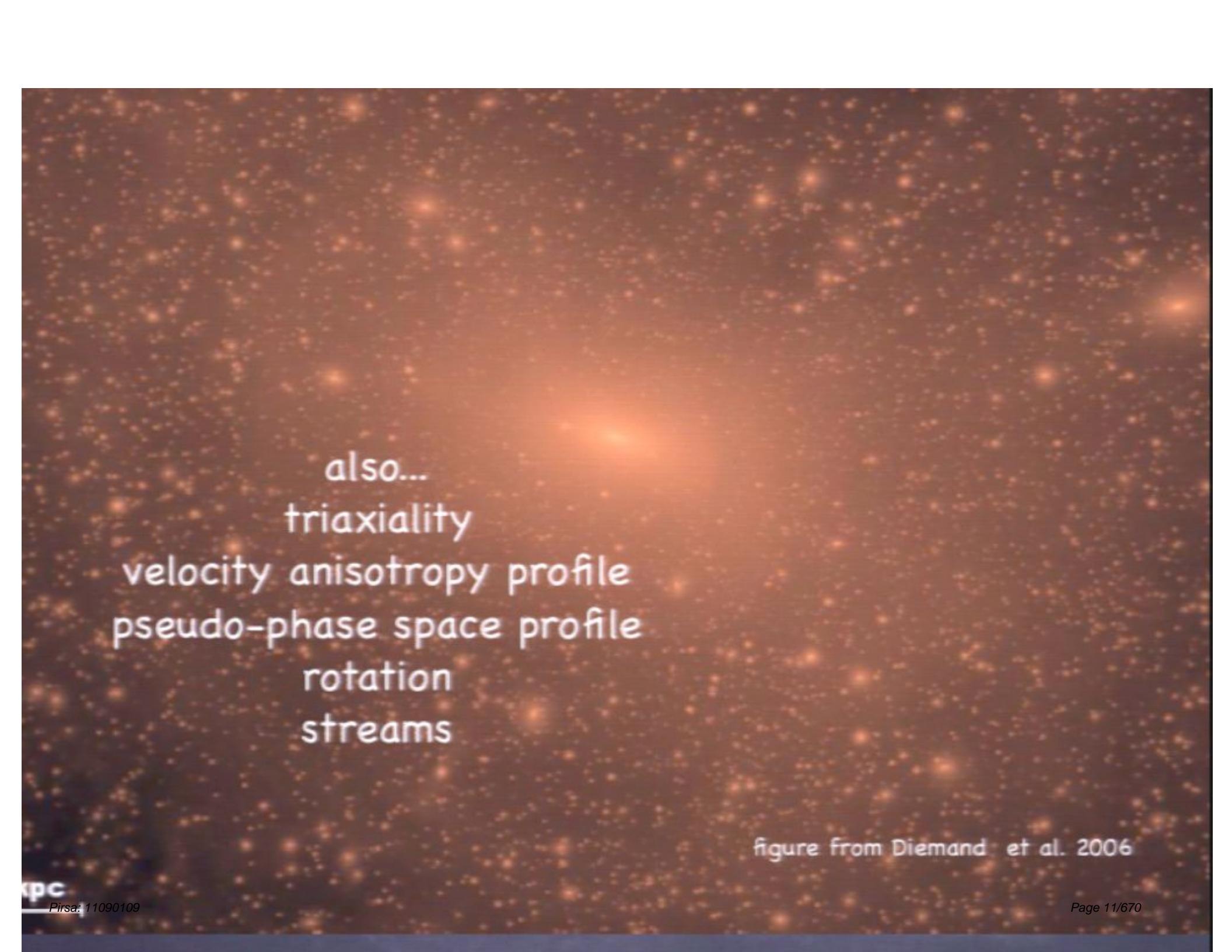
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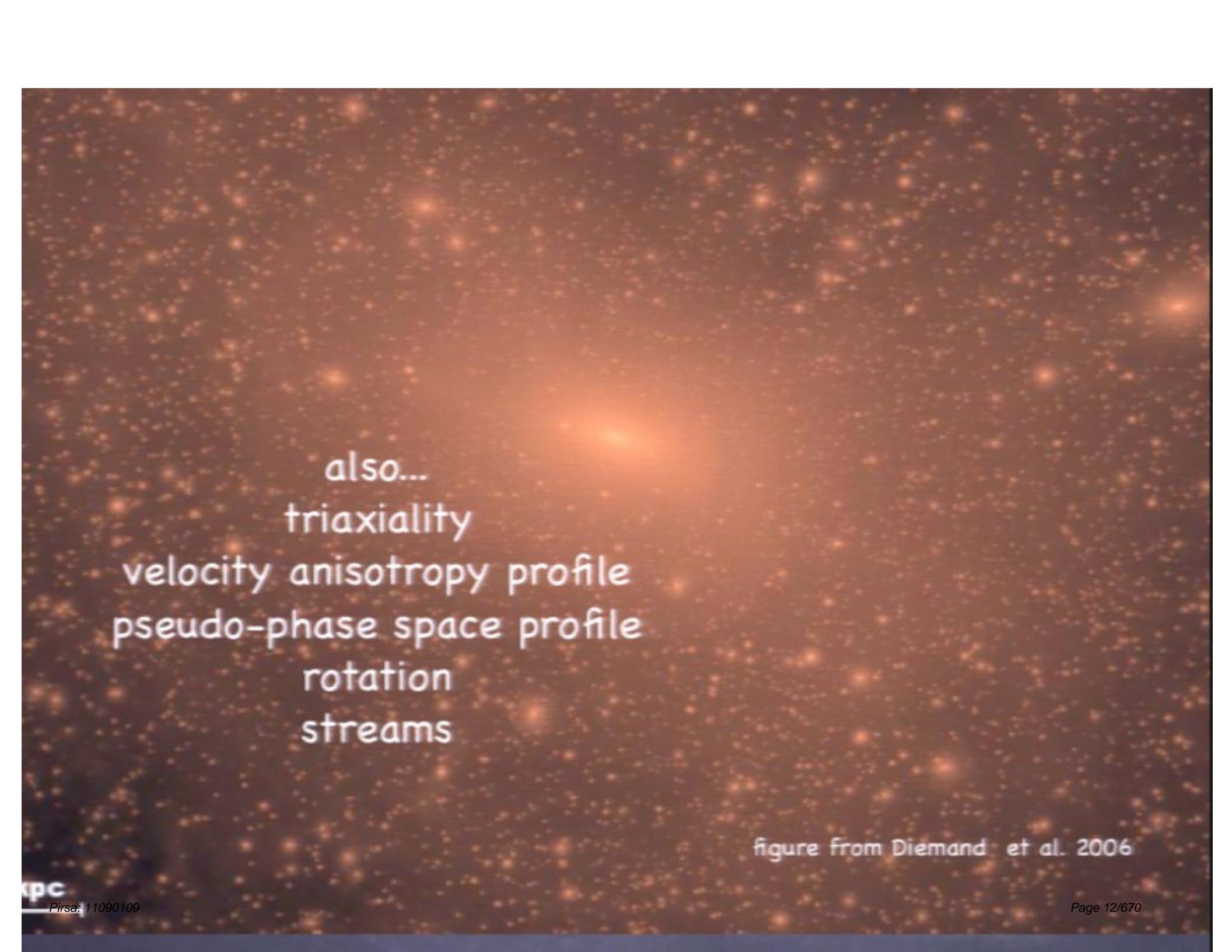
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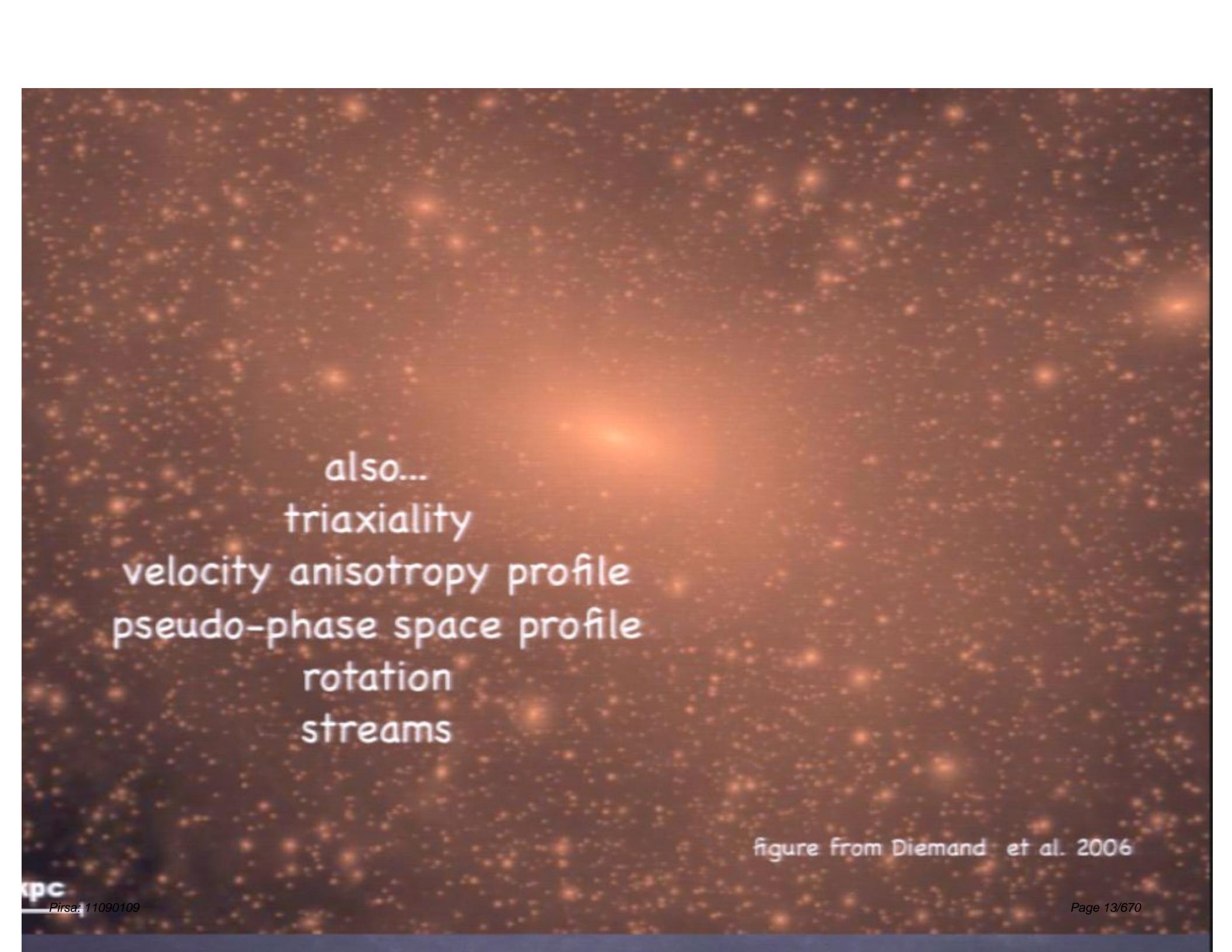
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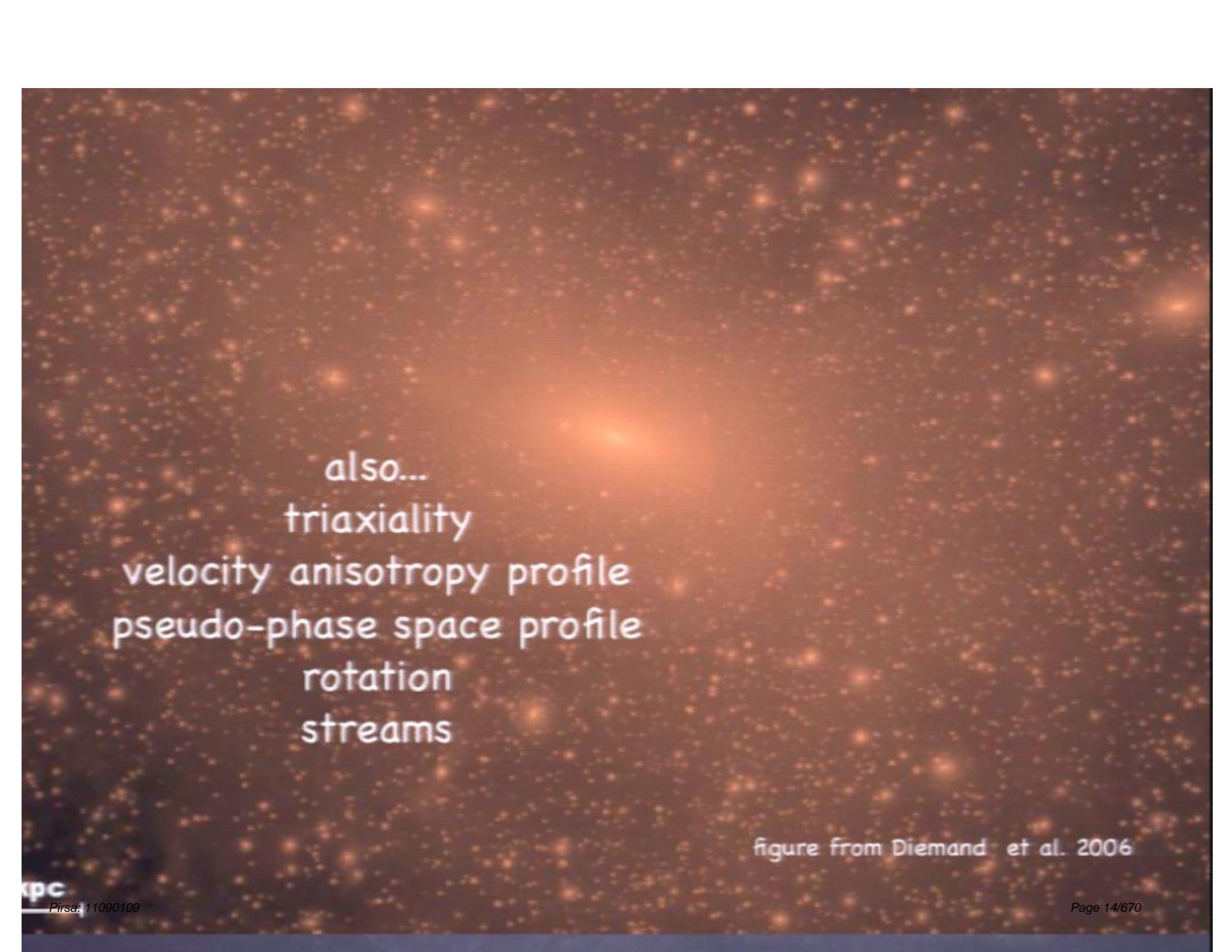
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In addition, there are hints of self-similarity between
halos of a different size



Moore et al. 1999

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Ultimate goal as dark matter theorists:
to develop a generative, theoretical
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build synthetic halos
indistinguishable from the halos
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Generative model →

push beyond resolution limits
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see also work by other conference participants Taylor, Afshordi, Vogelsberger, Kuhlen,

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$$f(\mathbf{x}, \mathbf{v}, t)$$

All dark matter detection experiments probe the DF, i.e., detector response involves some integral over the DF

$$\mathcal{R}(E, \dots) \sim \int f(\mathbf{x}, \mathbf{v}) I(E, \dots; \mathbf{x}, \mathbf{v}) d^3x d^3v$$

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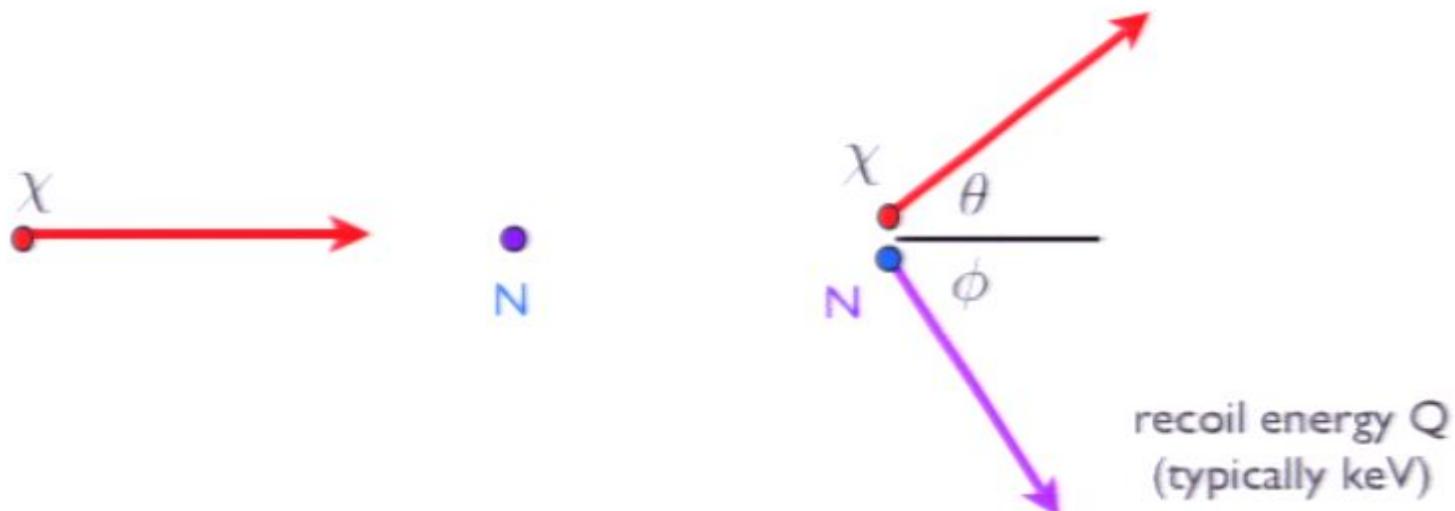
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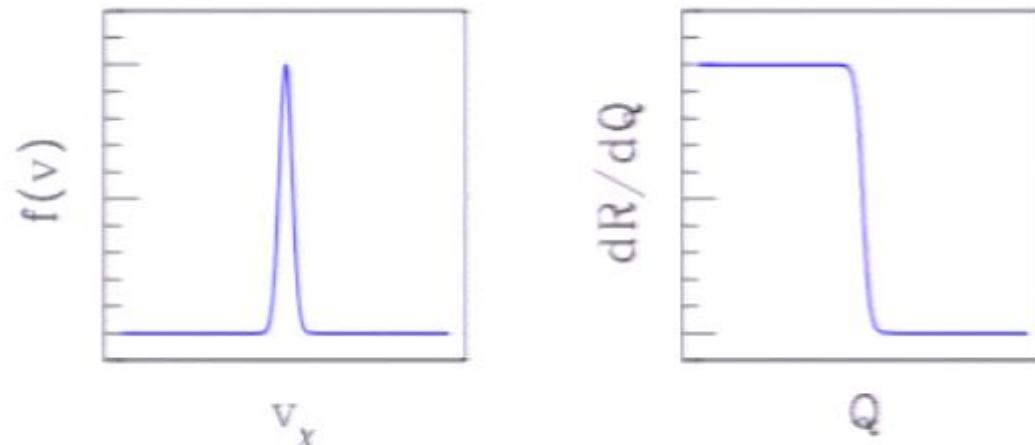
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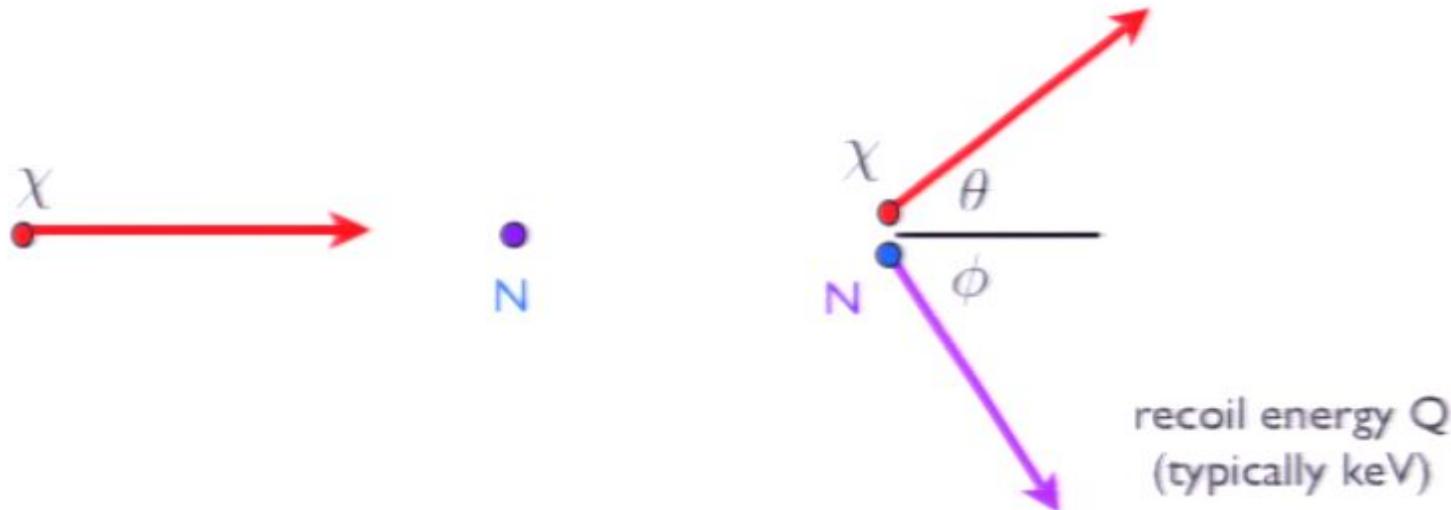
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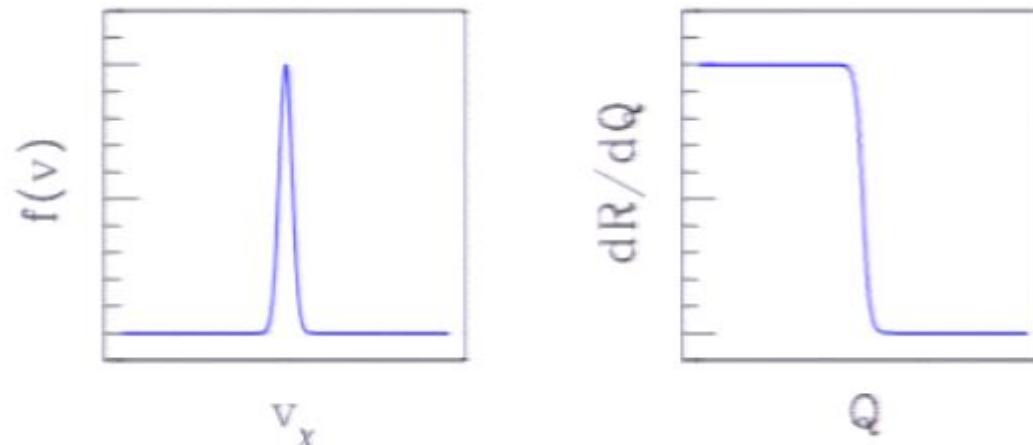
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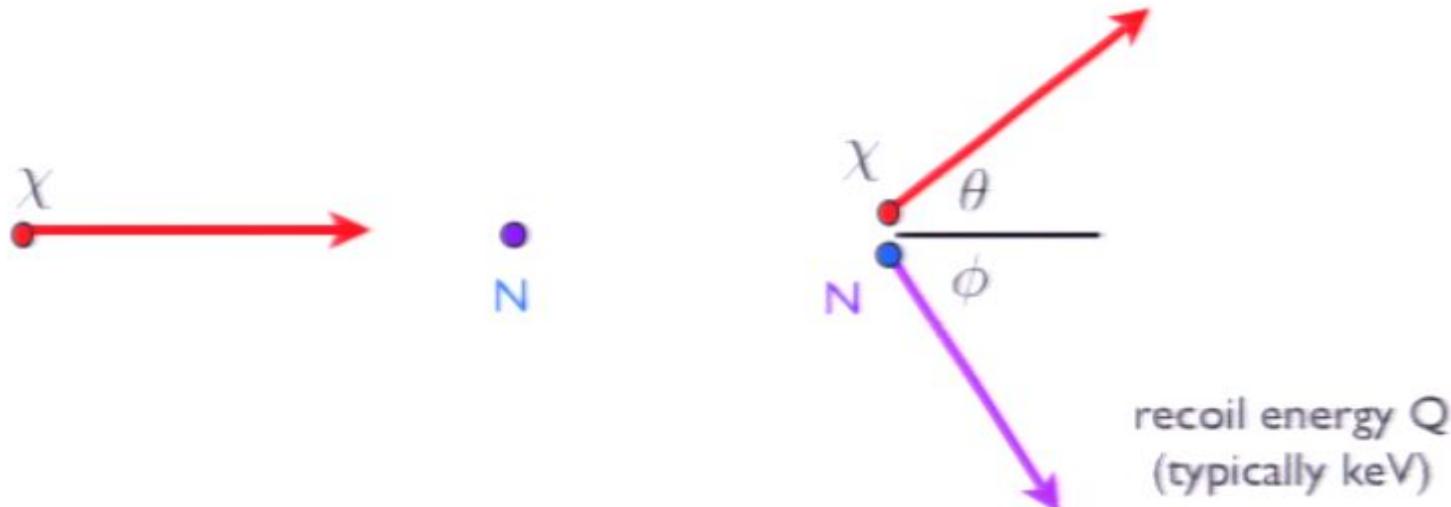




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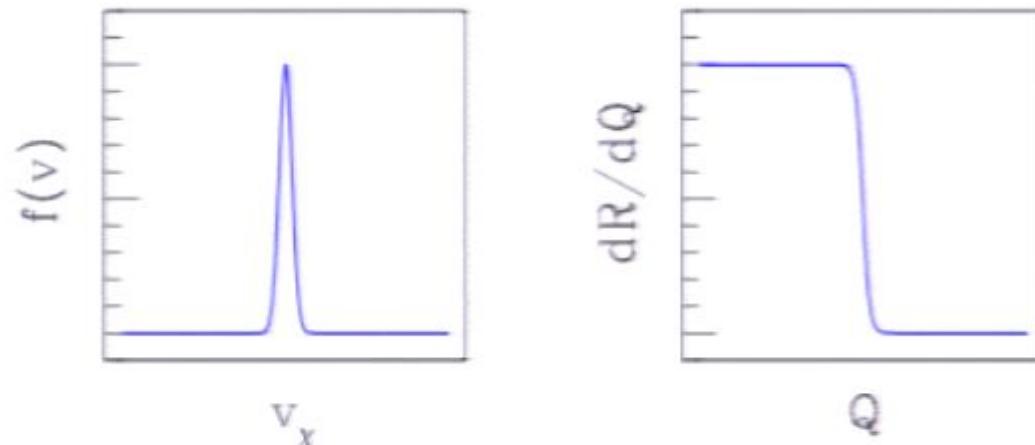
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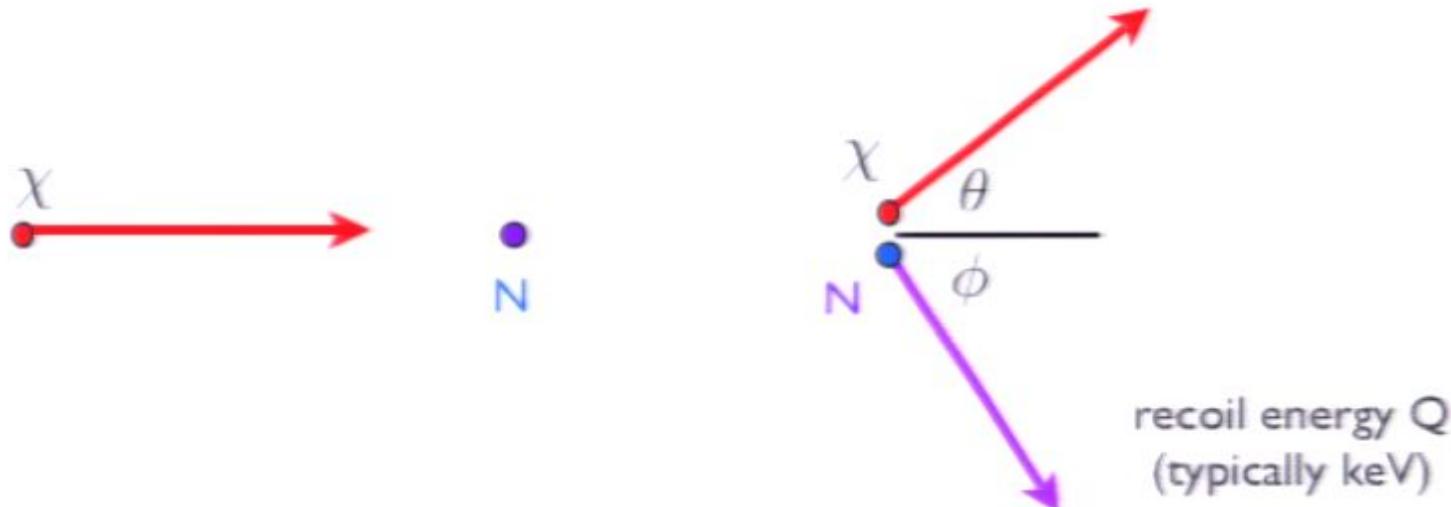




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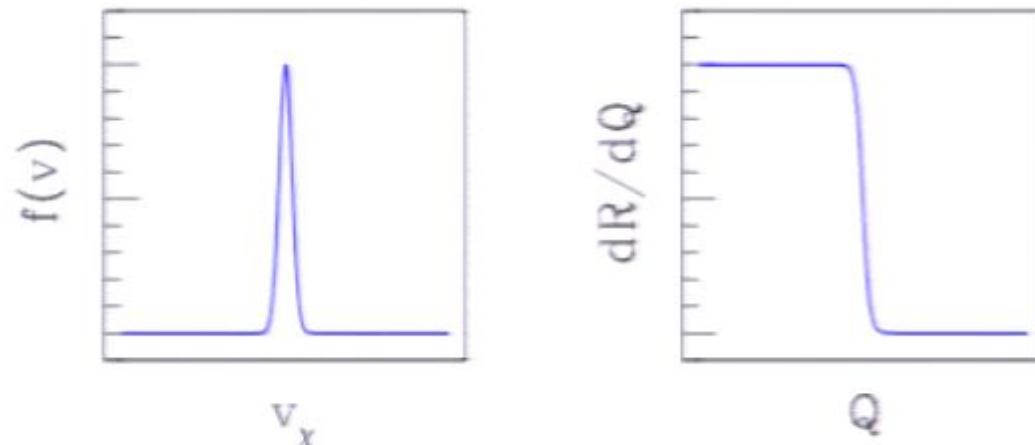
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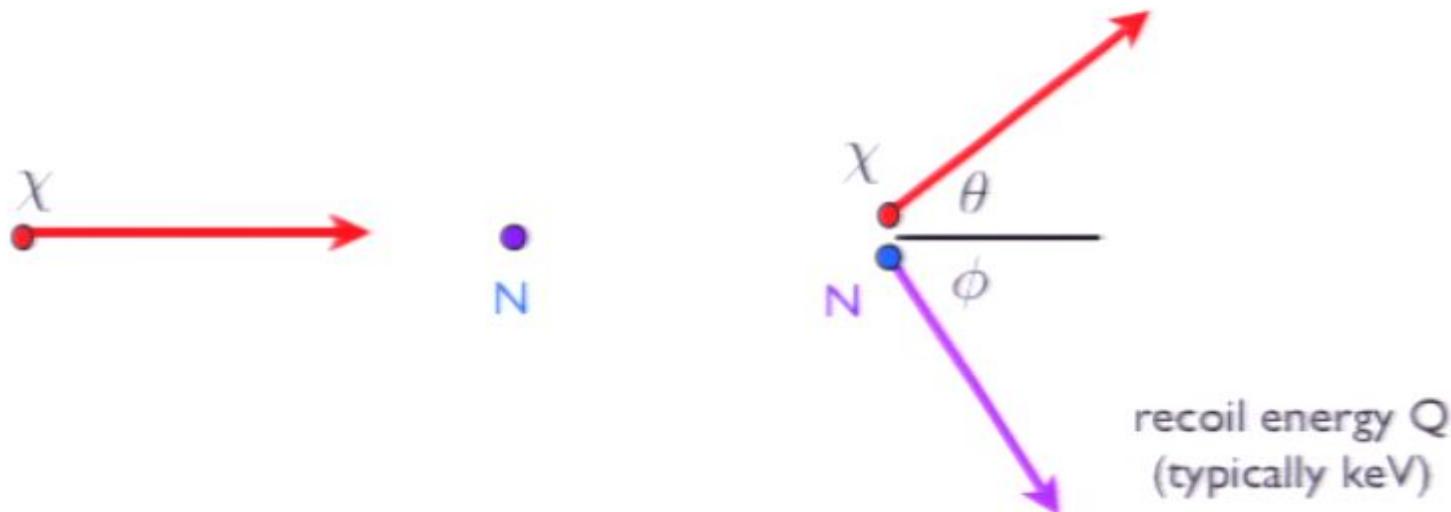




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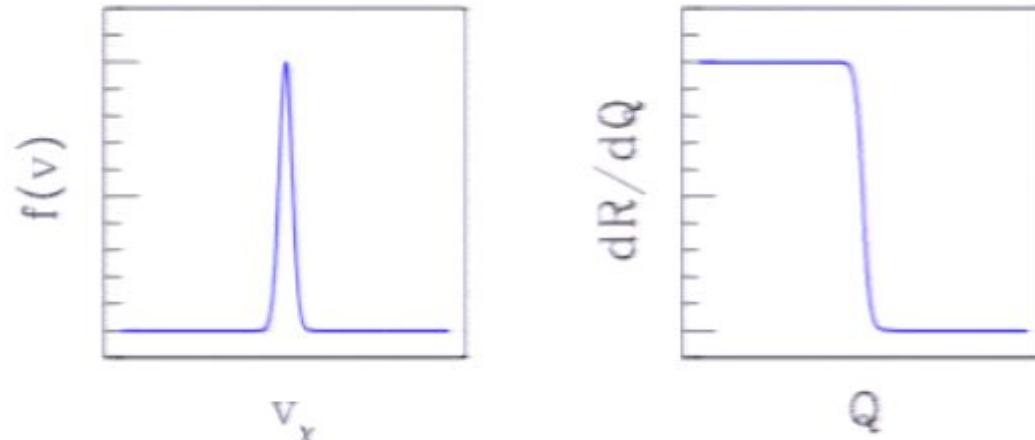
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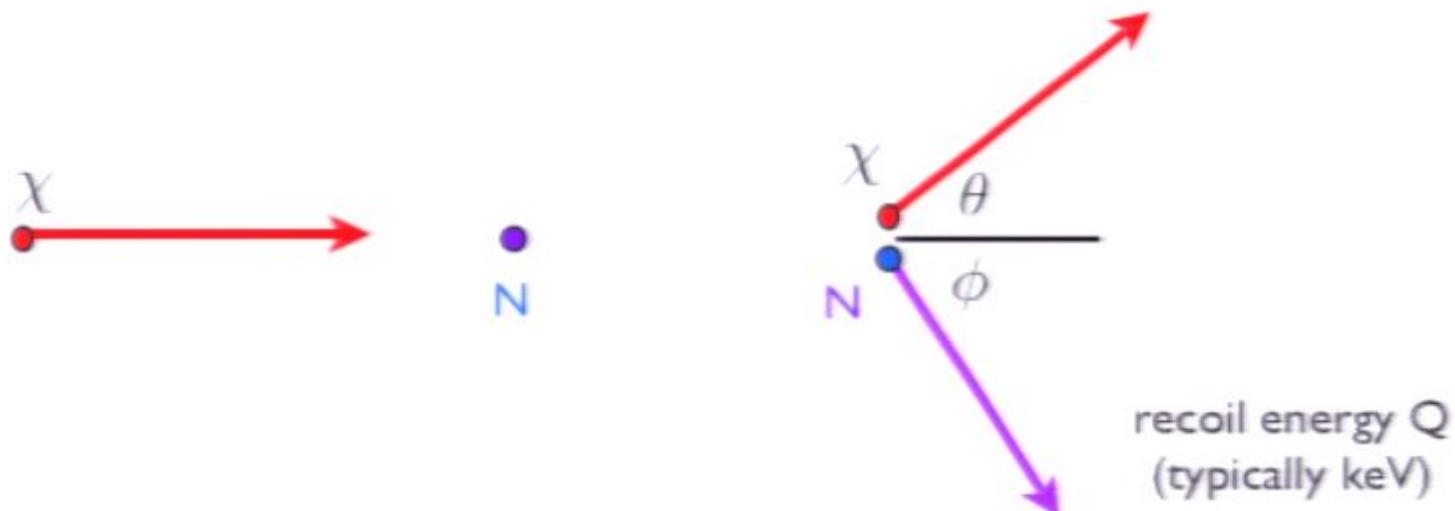




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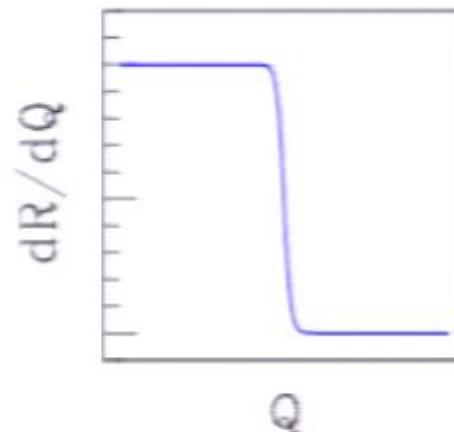
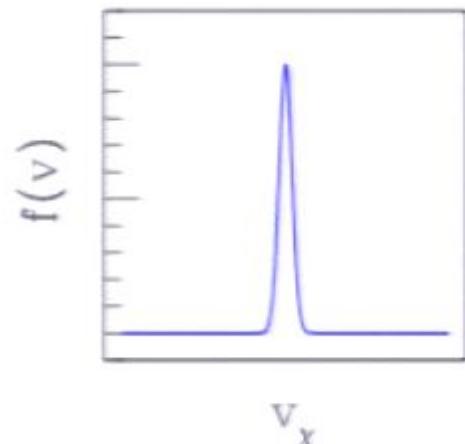
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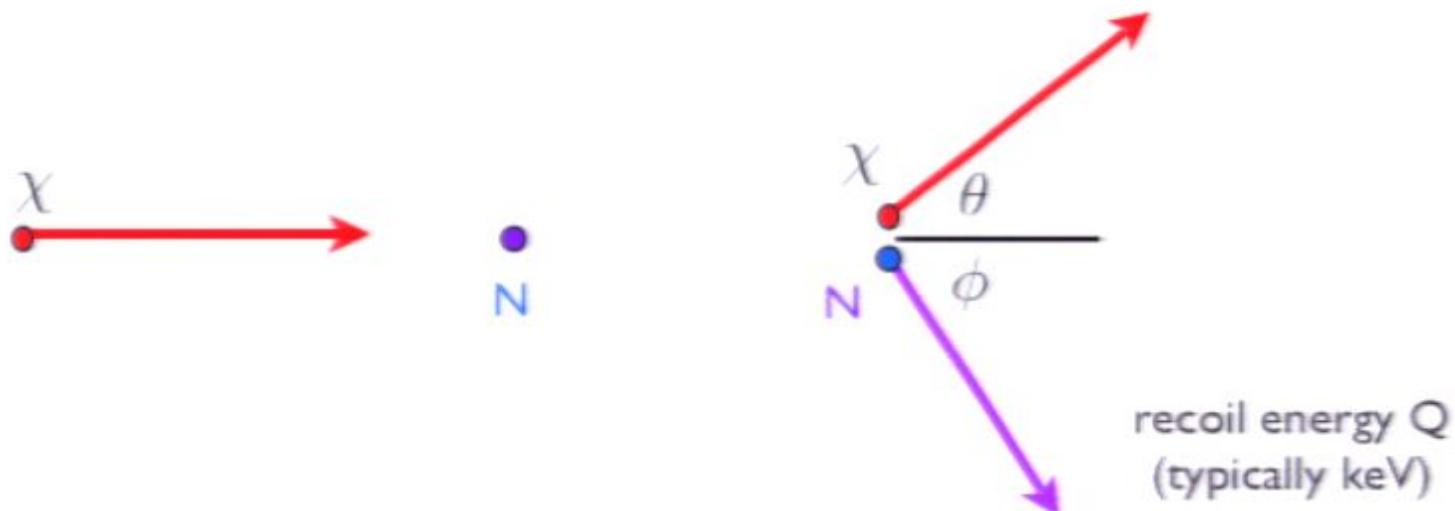




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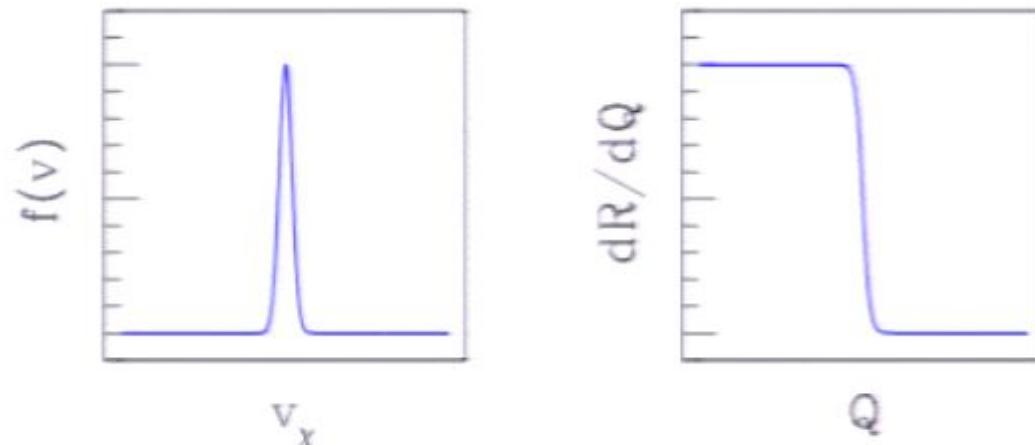
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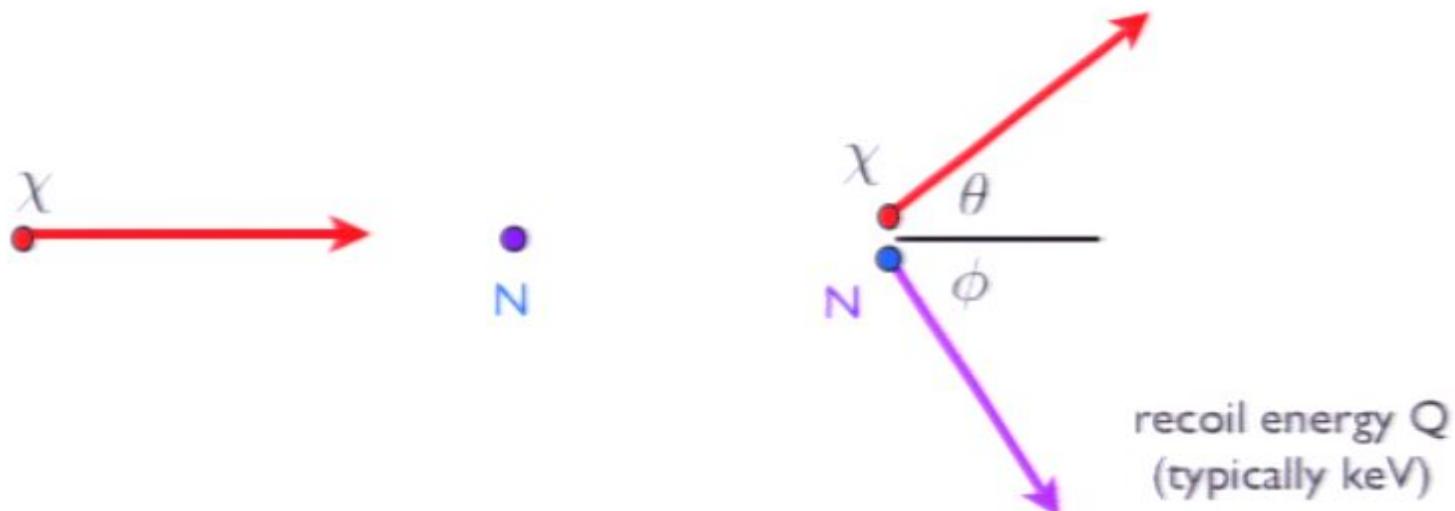




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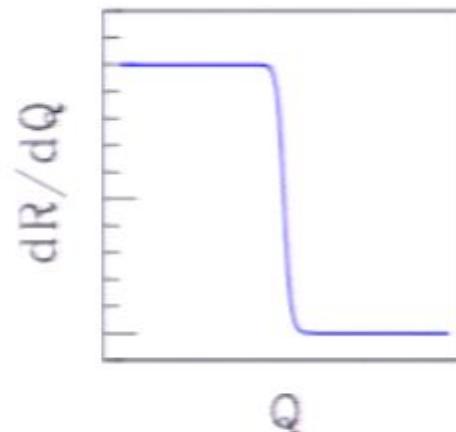
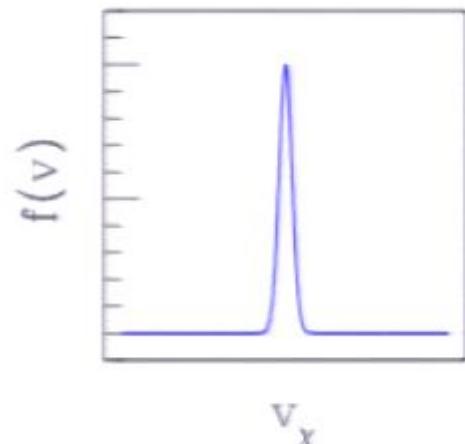
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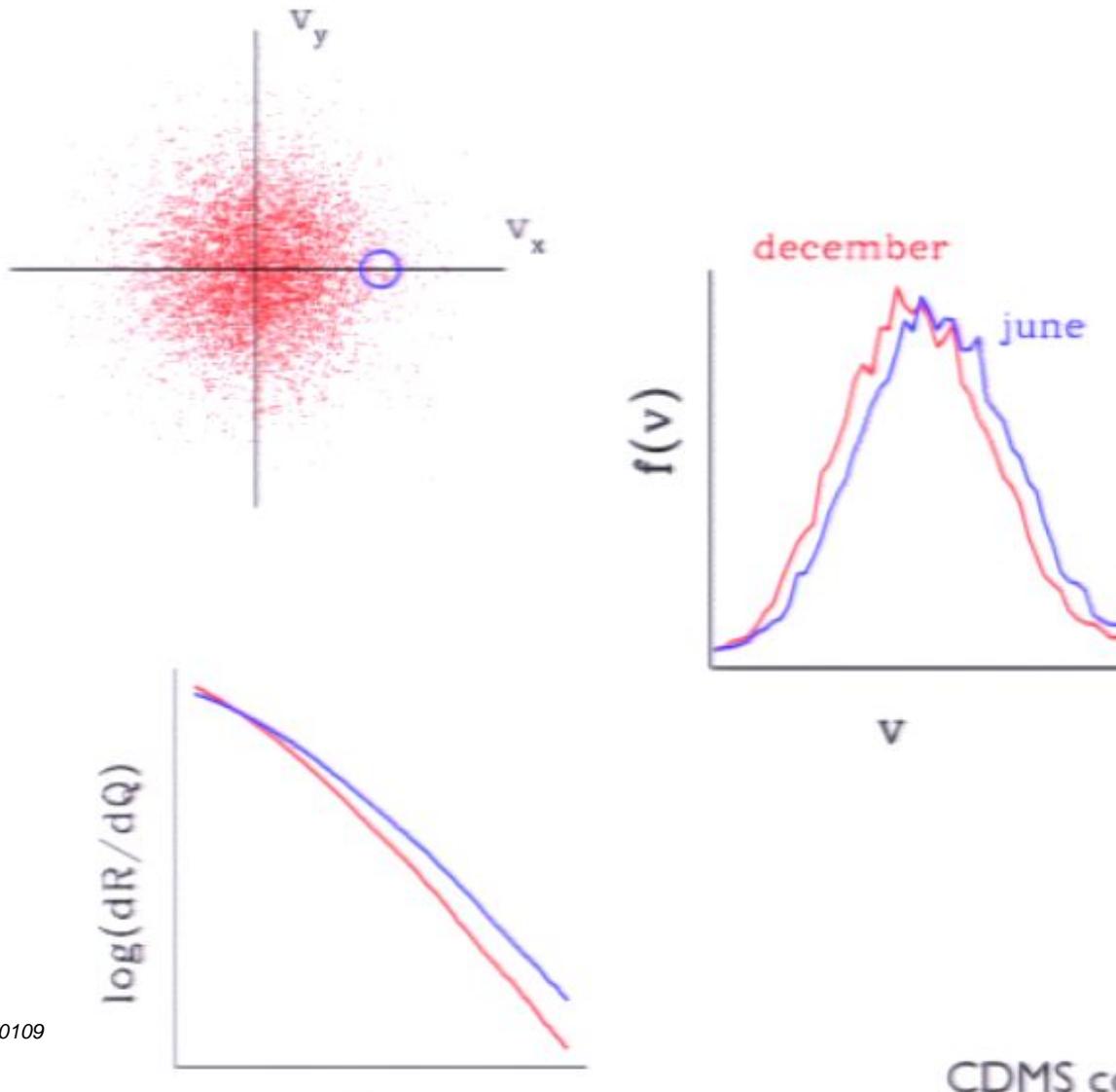


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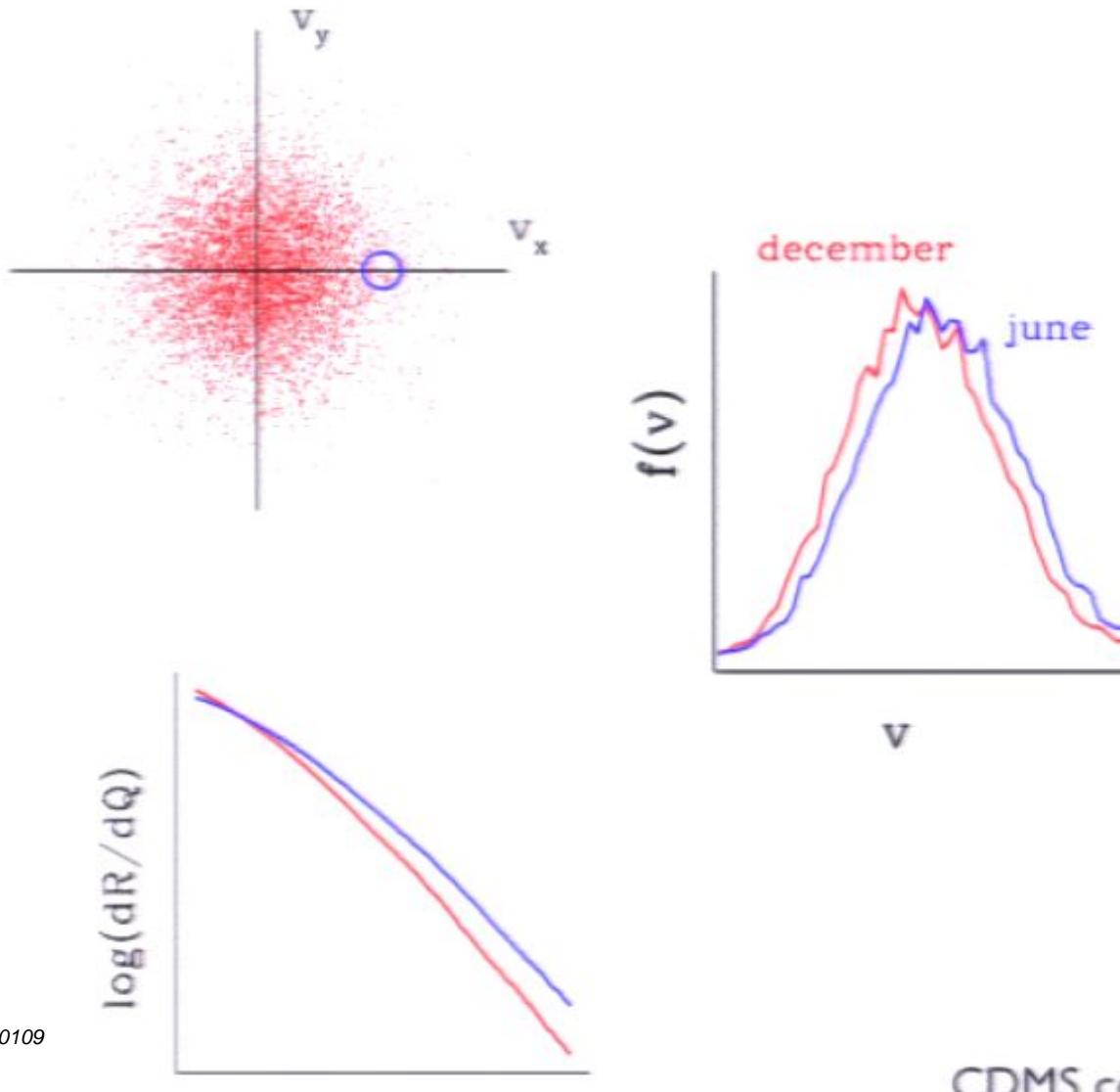
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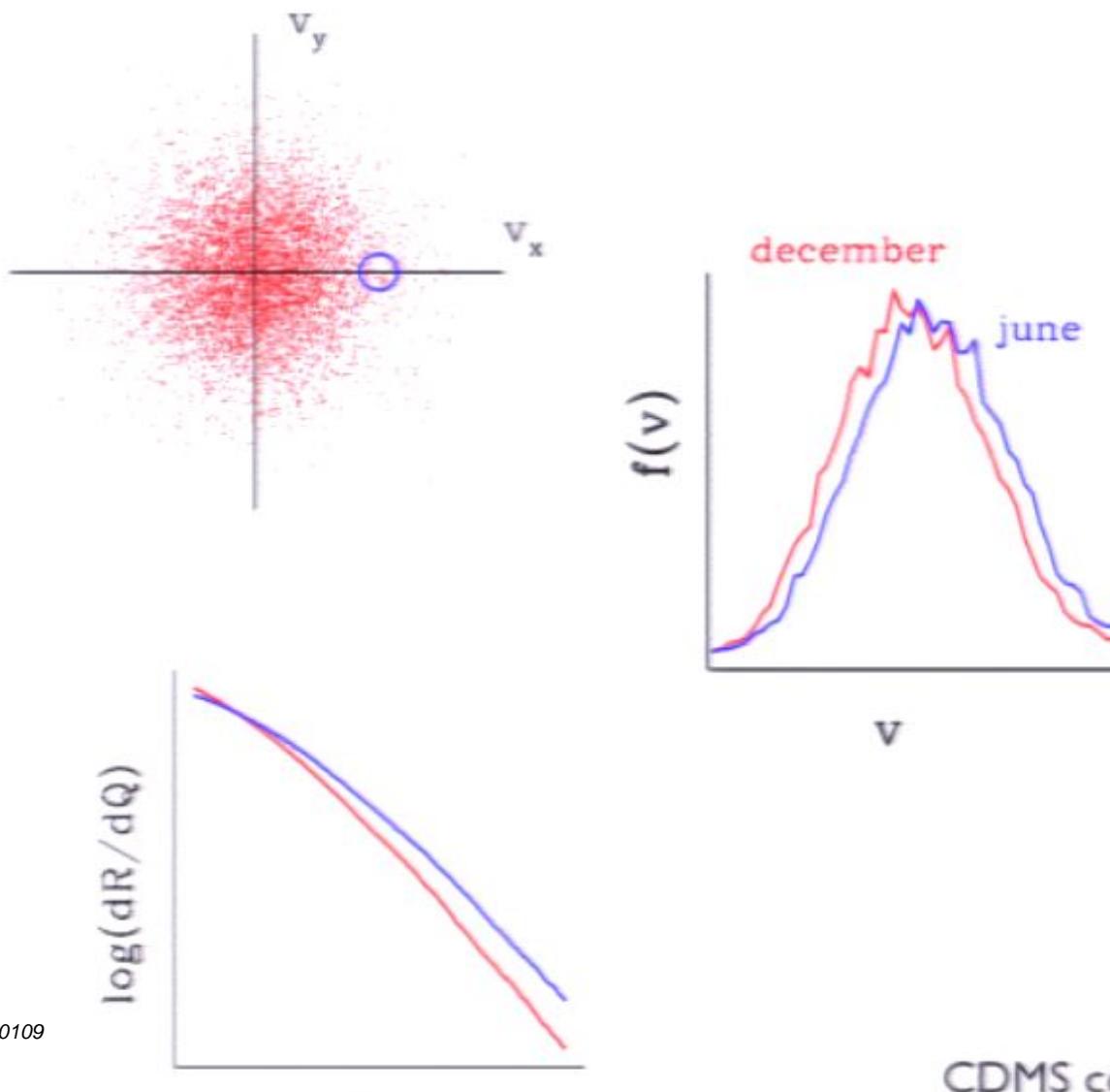
standard model used in DM detection experiments assumes a Maxwellian velocities in the Galactic rest frame



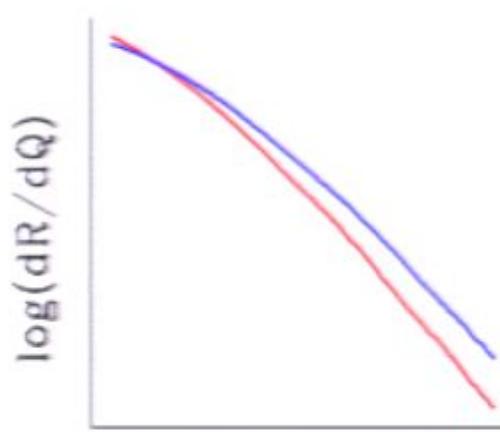
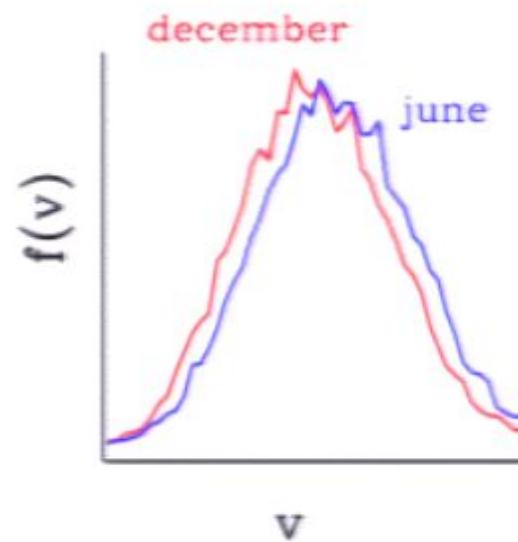
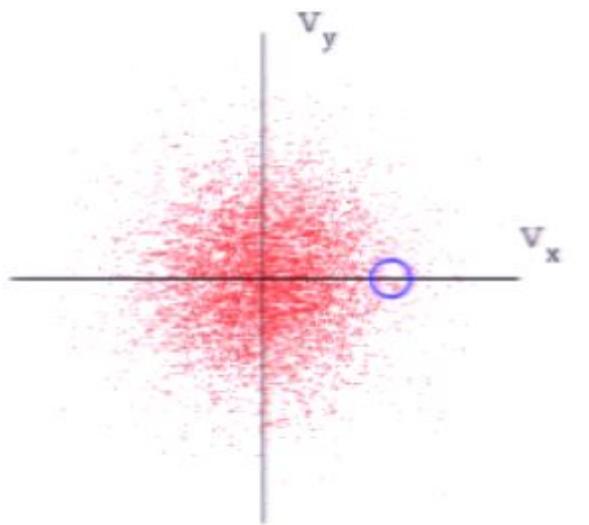
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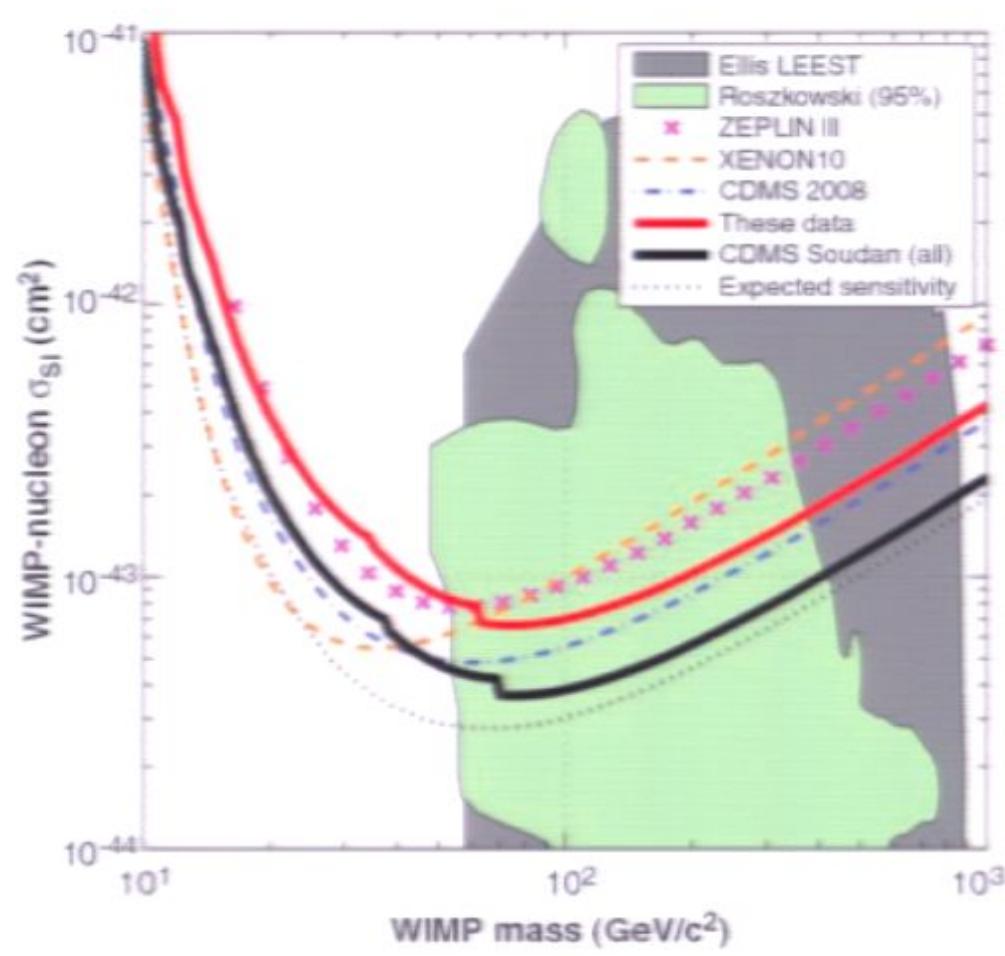
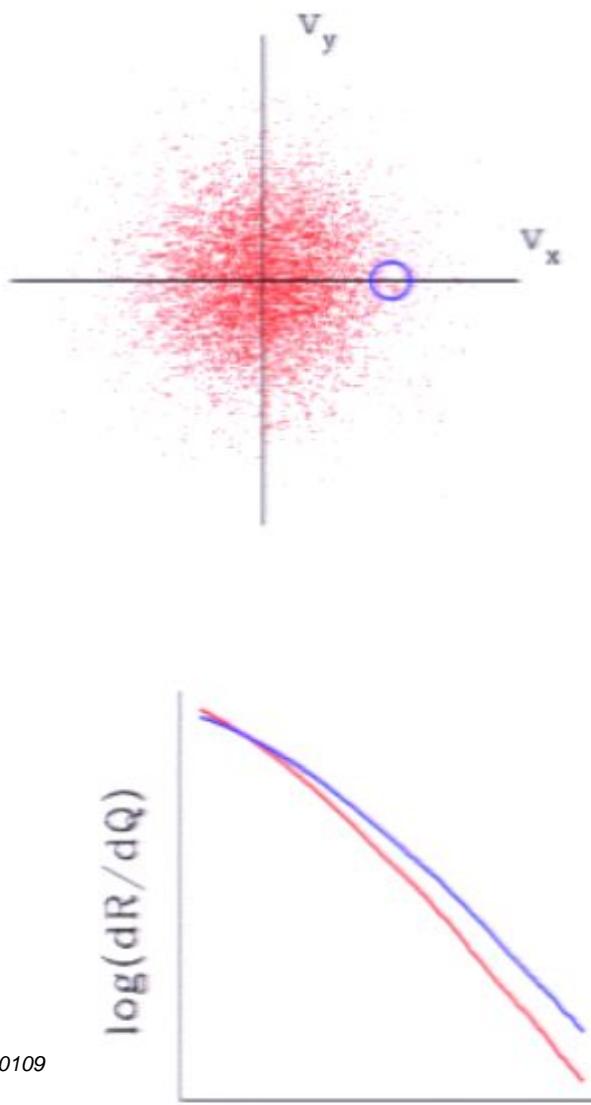
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Baby steps toward a generative model:

Given $\rho(r)$ and the assumptions of spherical symmetry and isotropic velocities, find $f(E)$ via the Eddington inversion formula

For “flattened” and rotating halos, consider $f(E, L_z)$

For velocity anisotropy, consider $f(E, L)$

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Challenges, I Limited resolution

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$$\sim 10^3 M_\odot$$

In typical WIMP models, the first objects to form in the Earth. Thus, we are 9 or more orders of magnitude short of doing fully realistic simulations of a MW-sized halo

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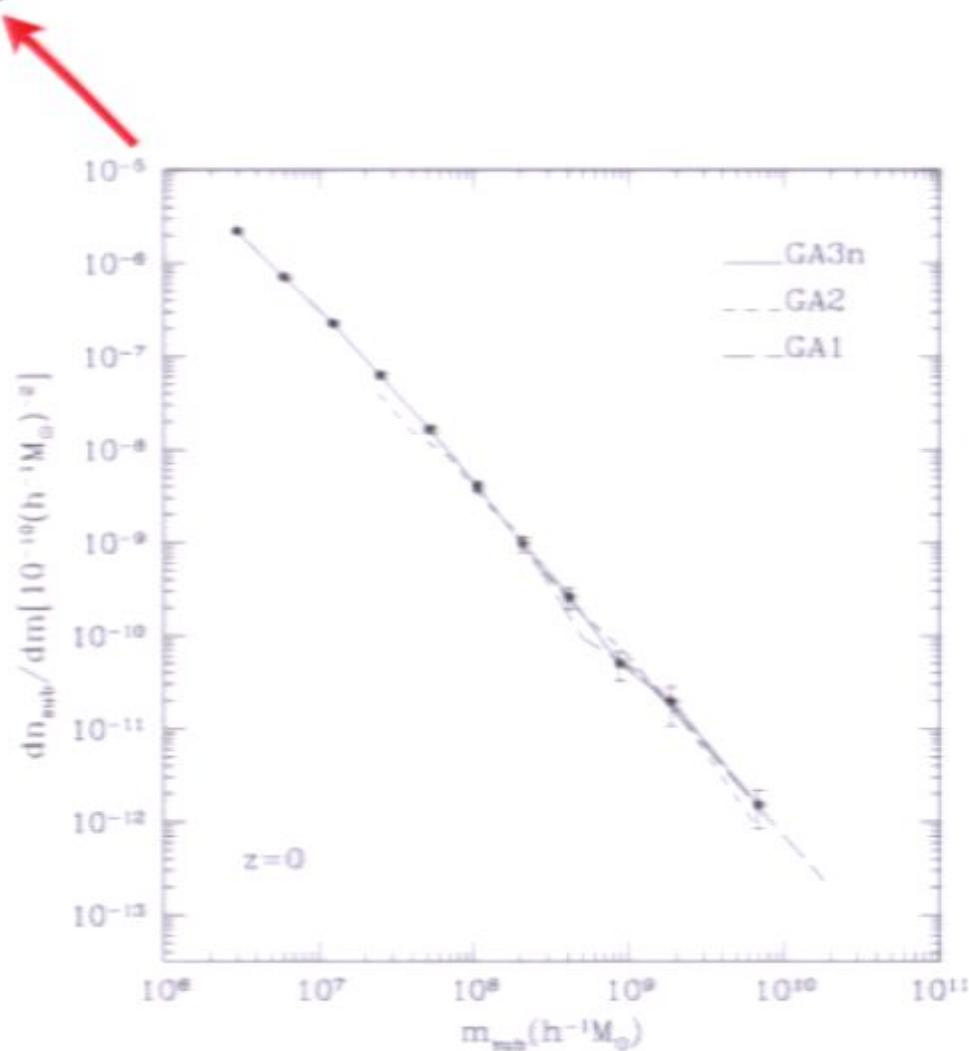
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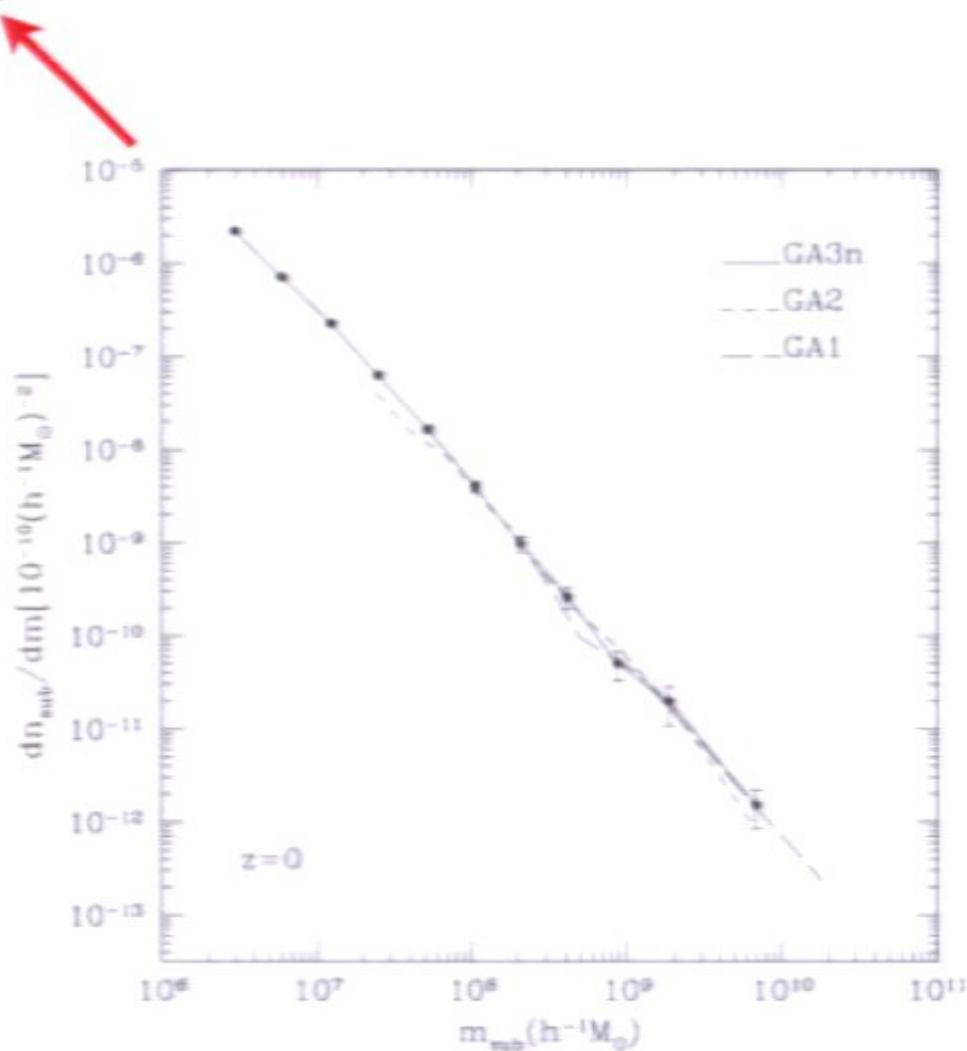
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subhalo mass function from
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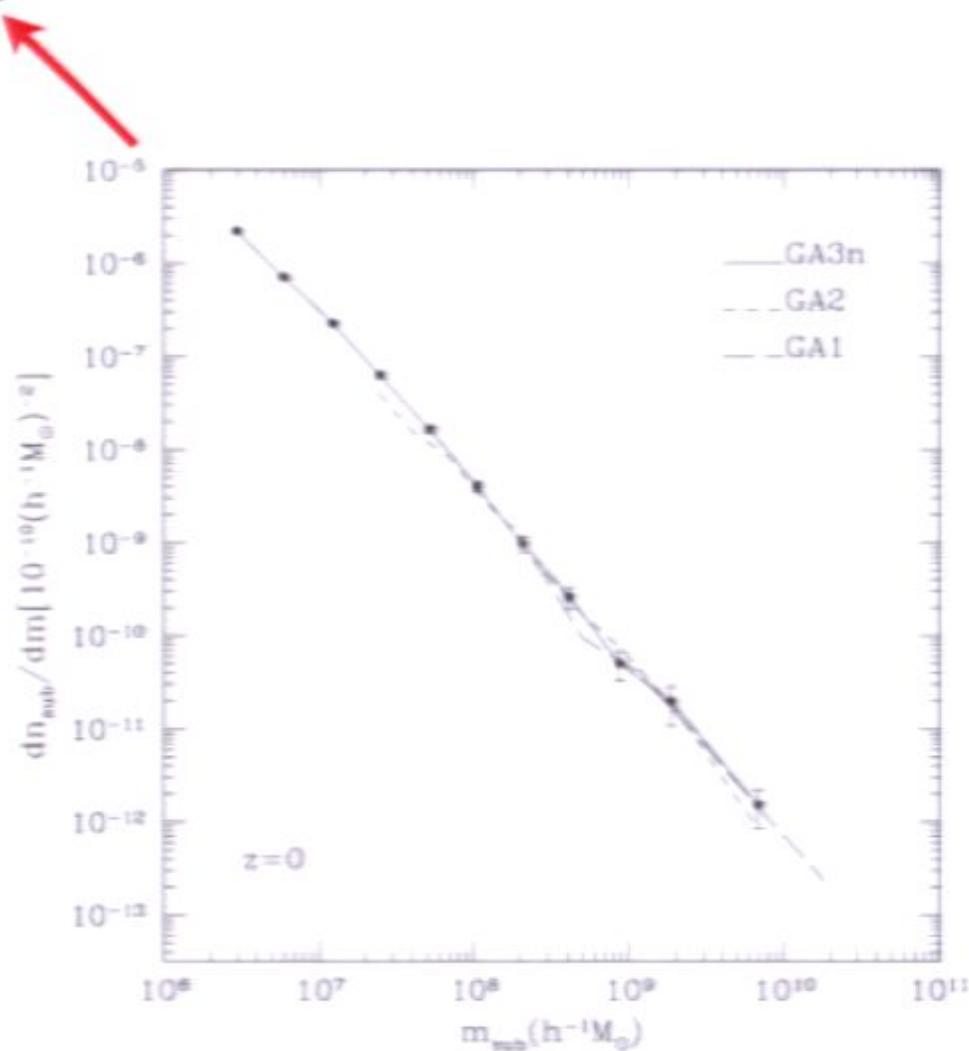
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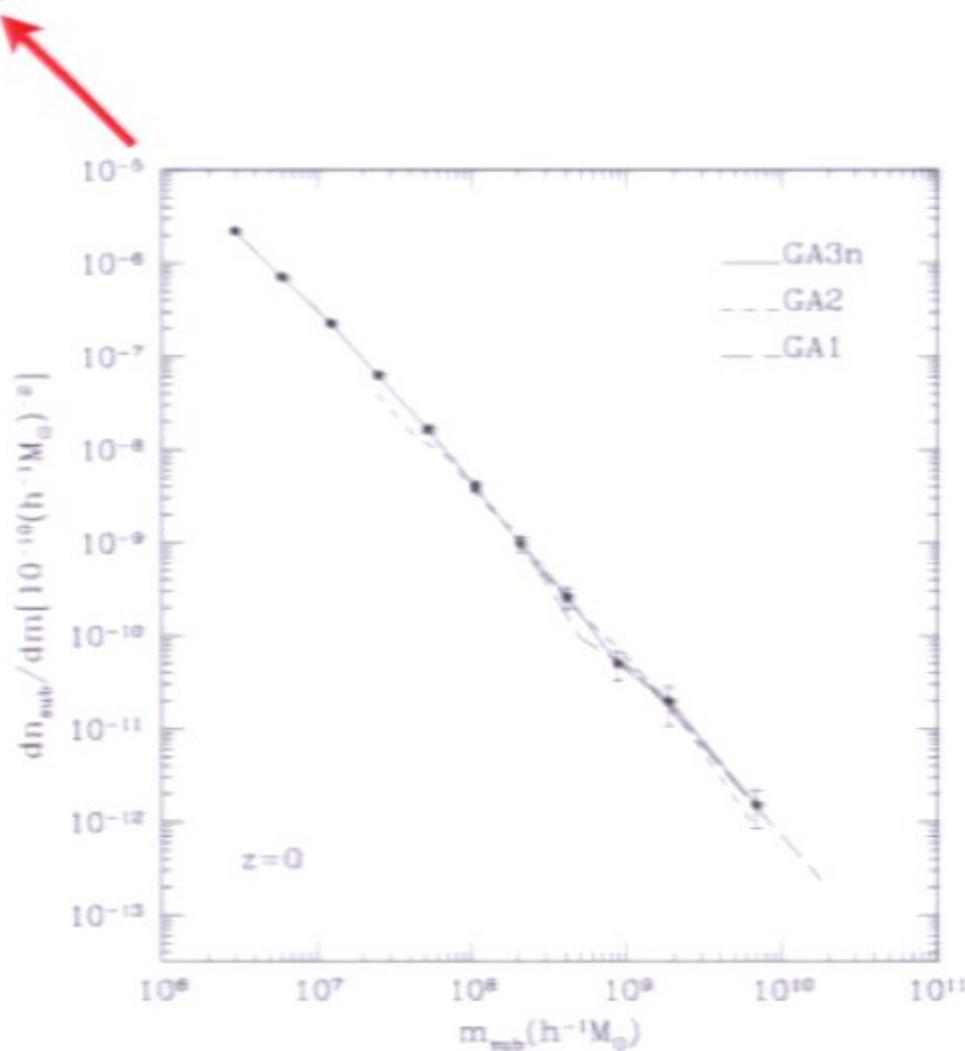
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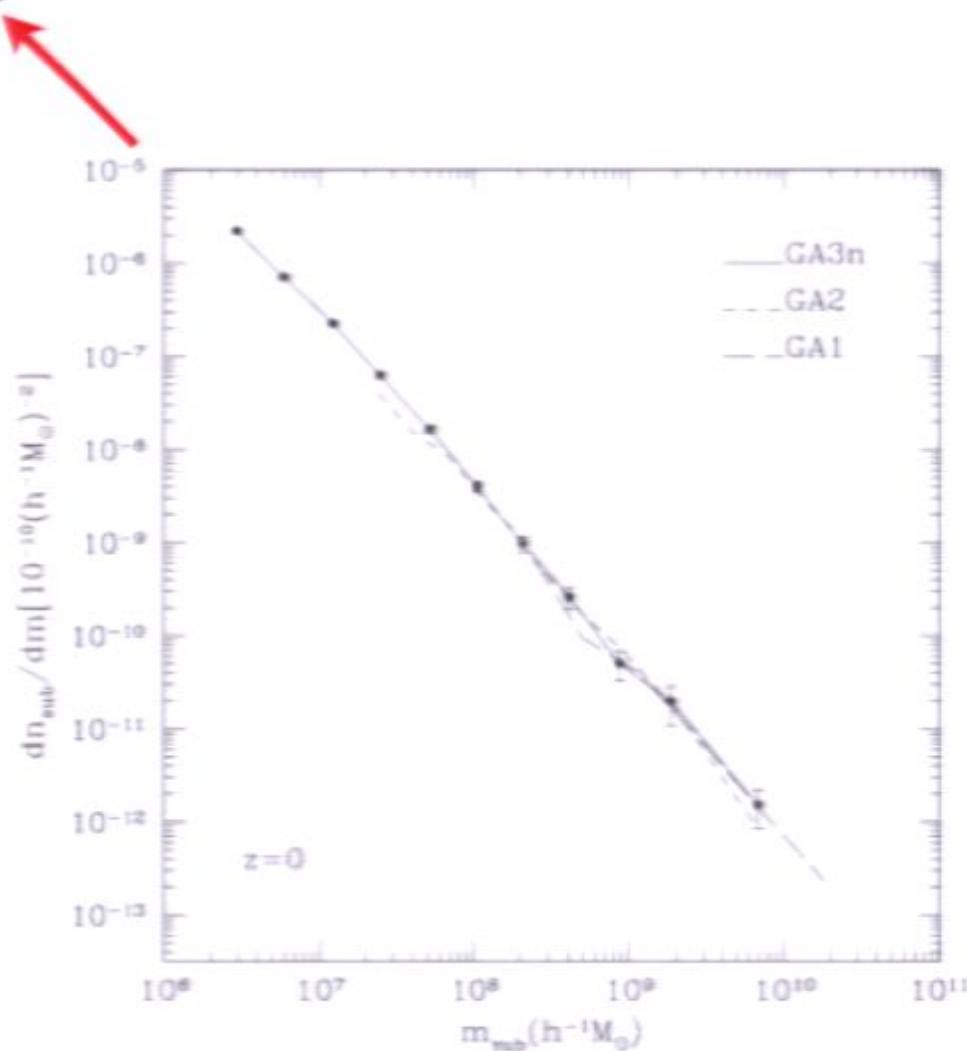
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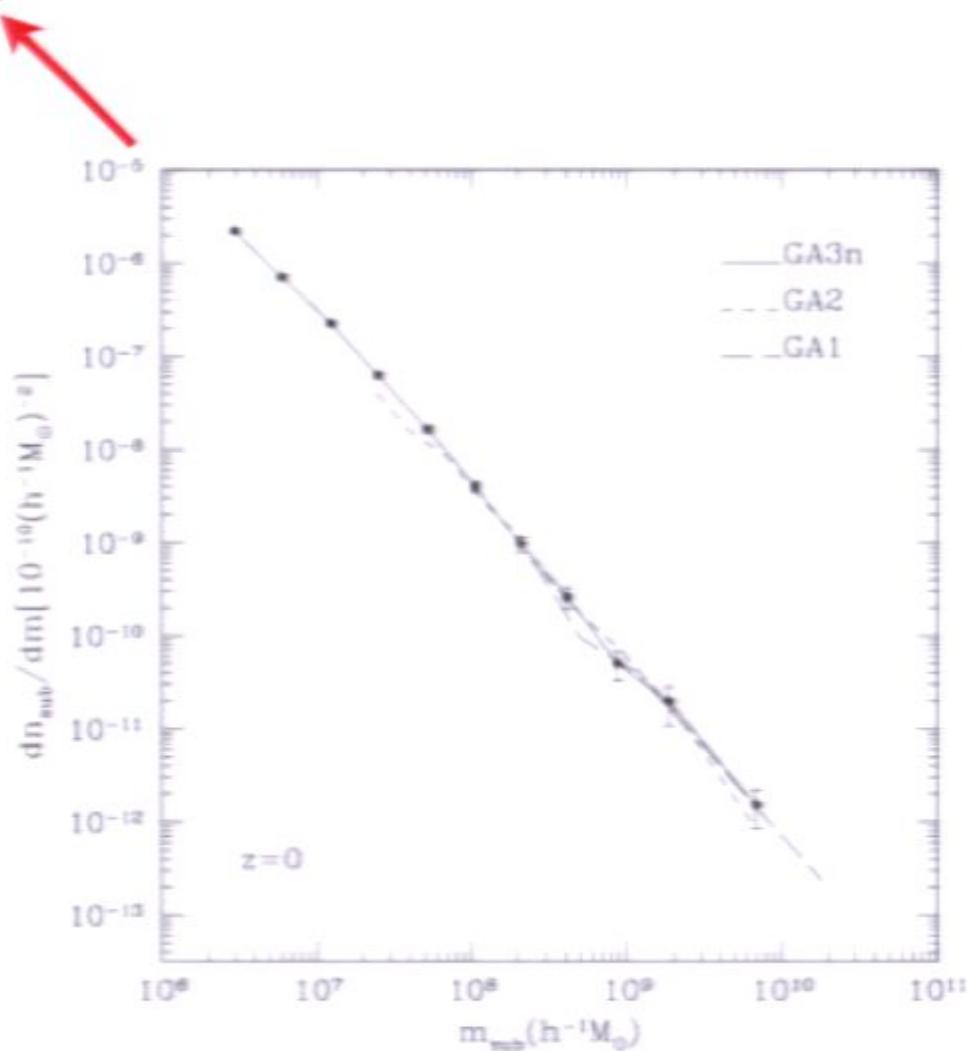
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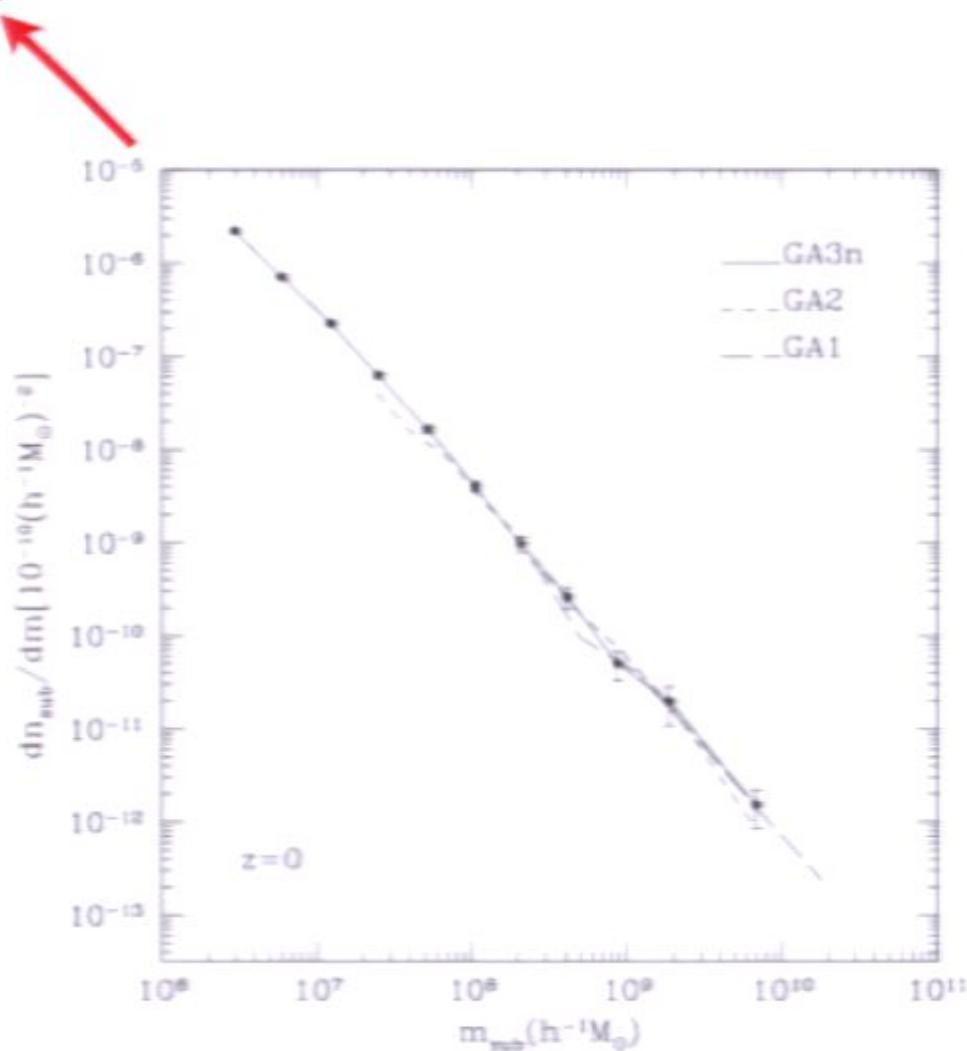
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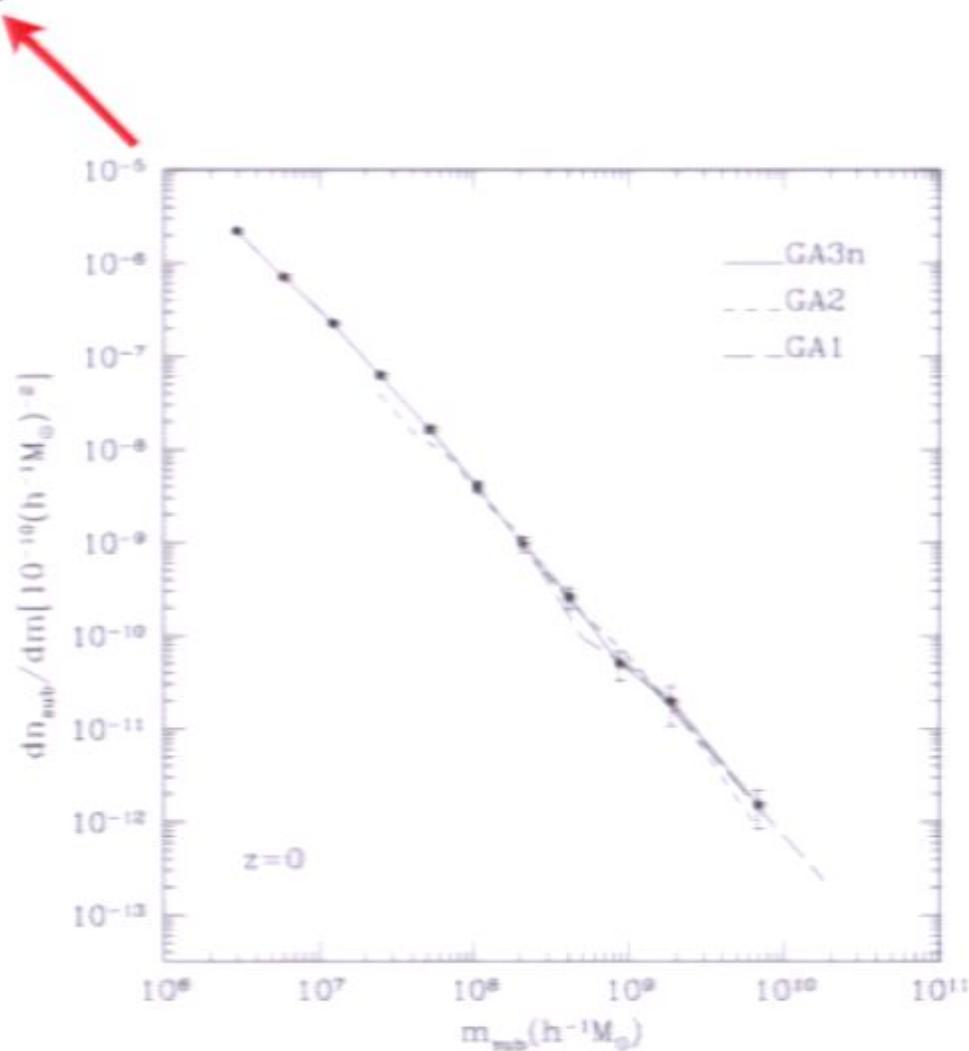
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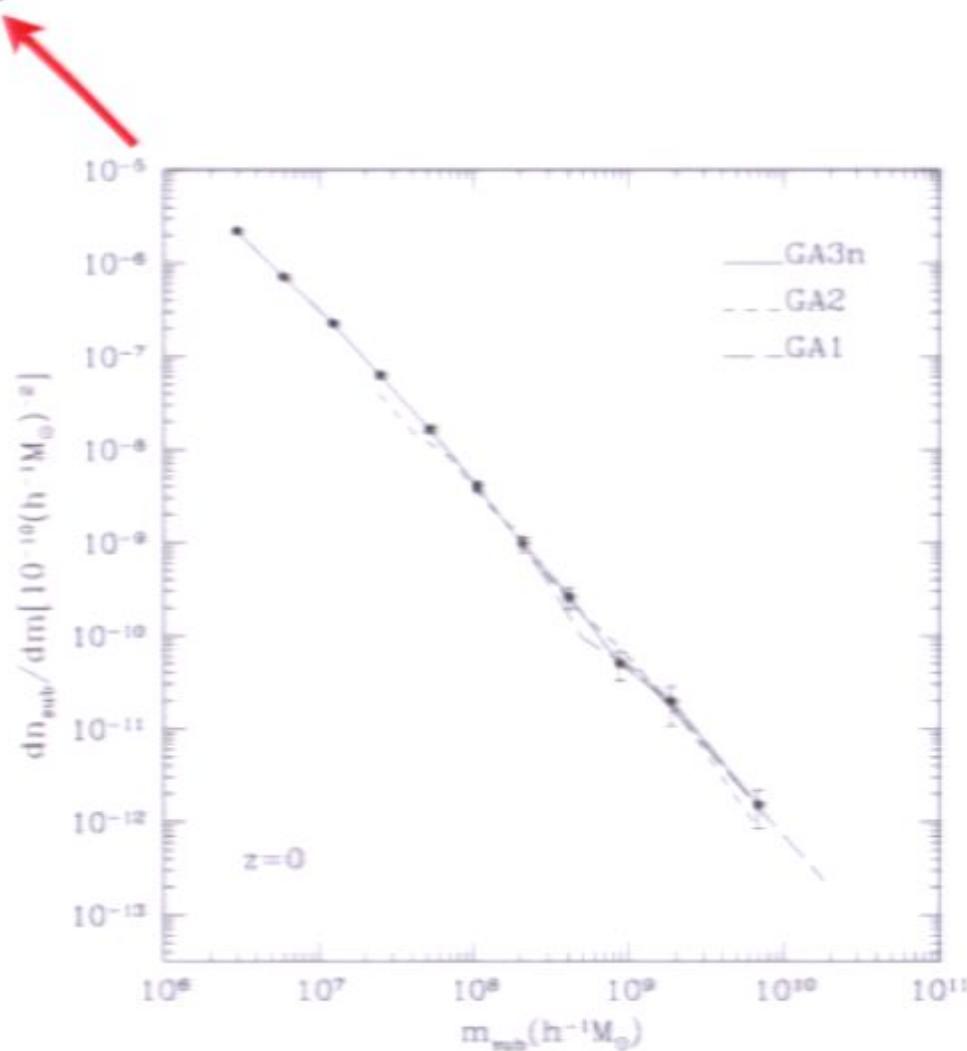
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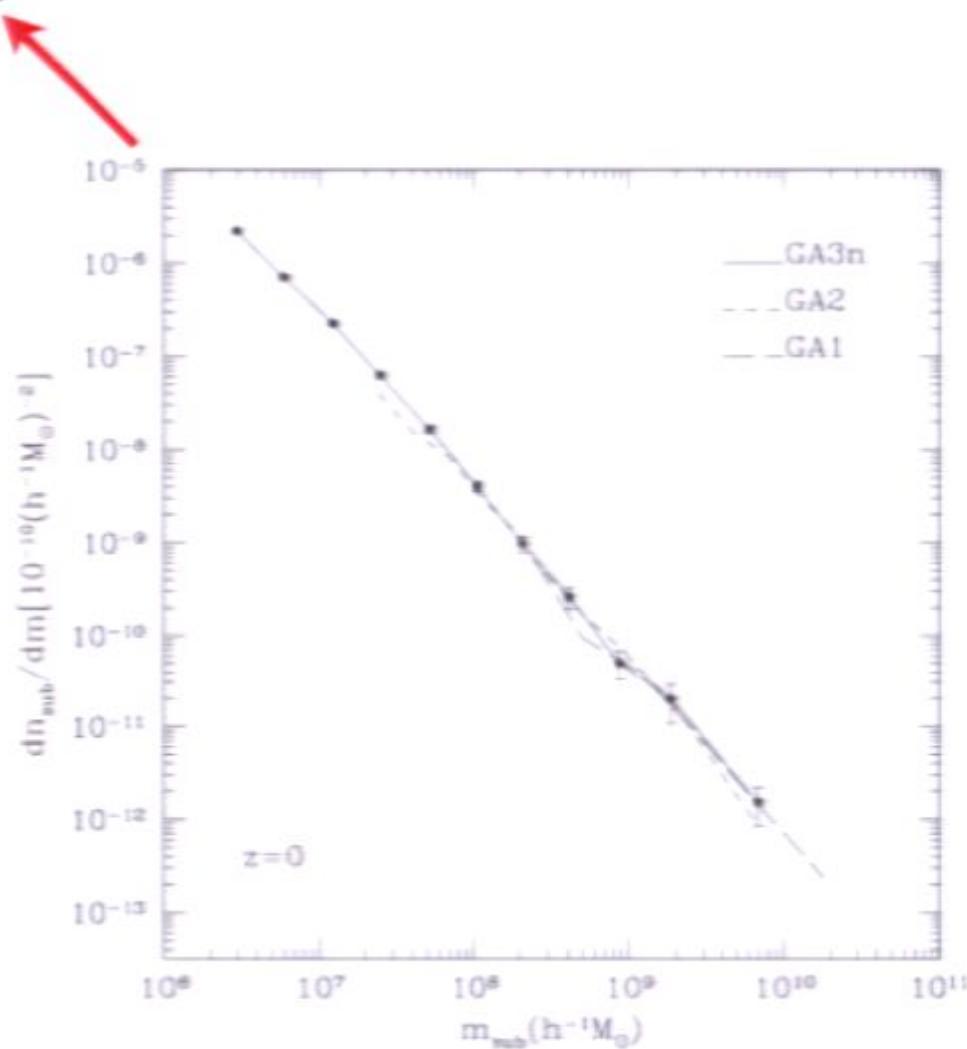
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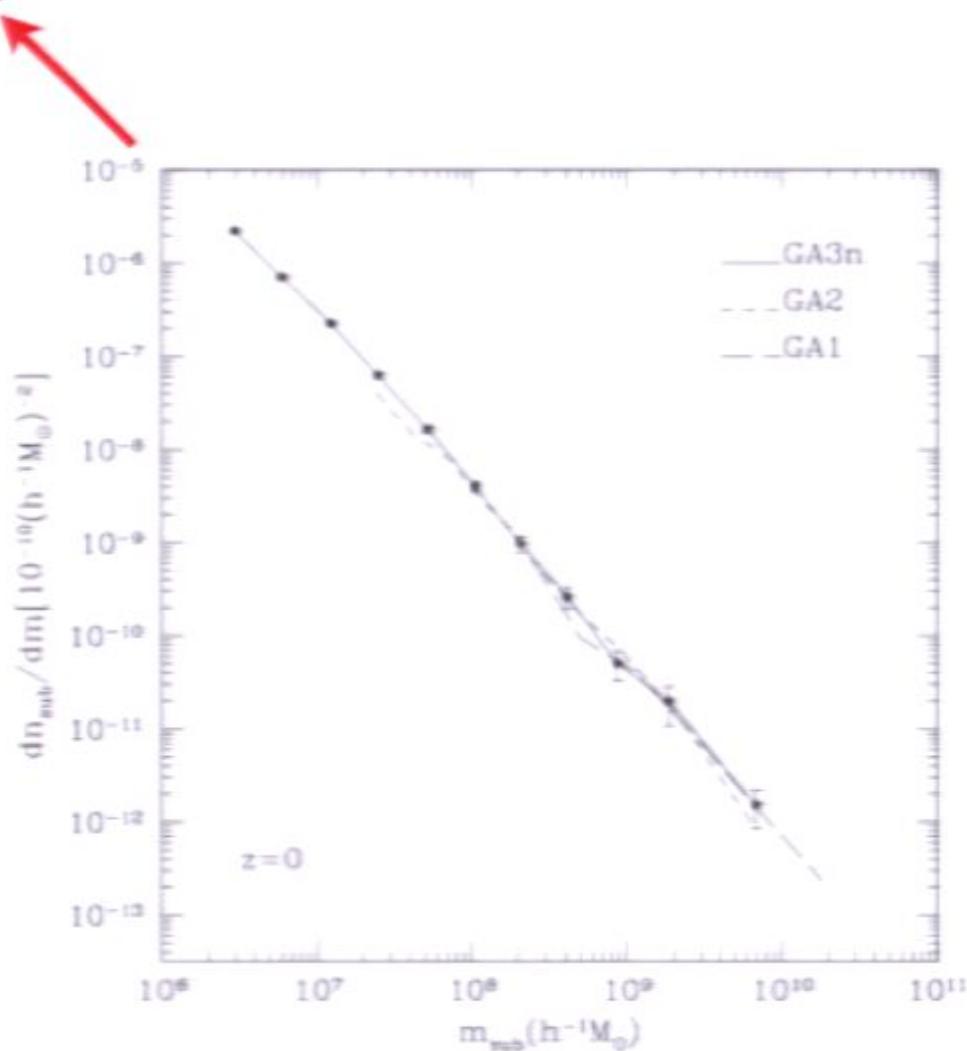
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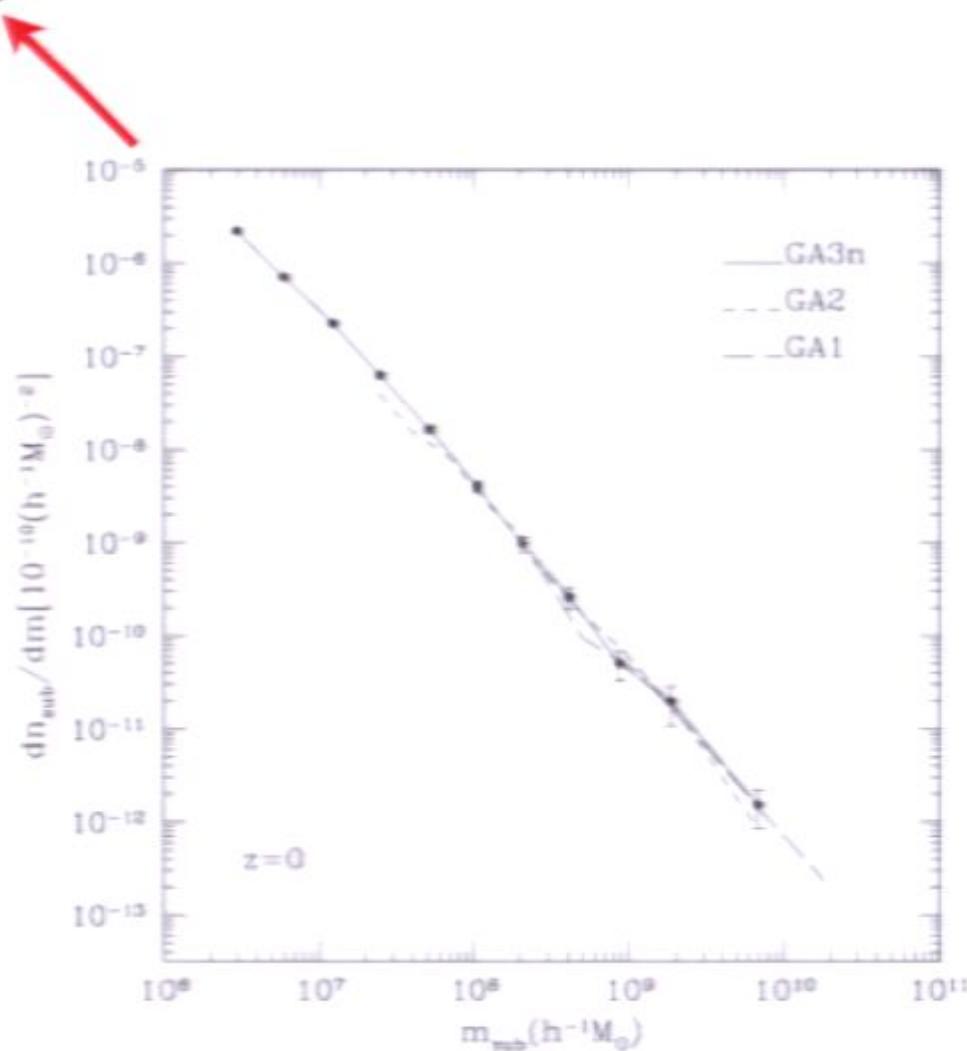
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Challenges, II

Pathological behavior of the CDM power spectrum at very small scales

For the first objects in the CDM hierarchy structures form over a wide range of scales nearly simultaneously.

Clustering no longer hierarchical and a significant dynamic range in simulations is difficult to achieve.

Other schemes to understand structure formation (e.g., perturbation theory) also have problems

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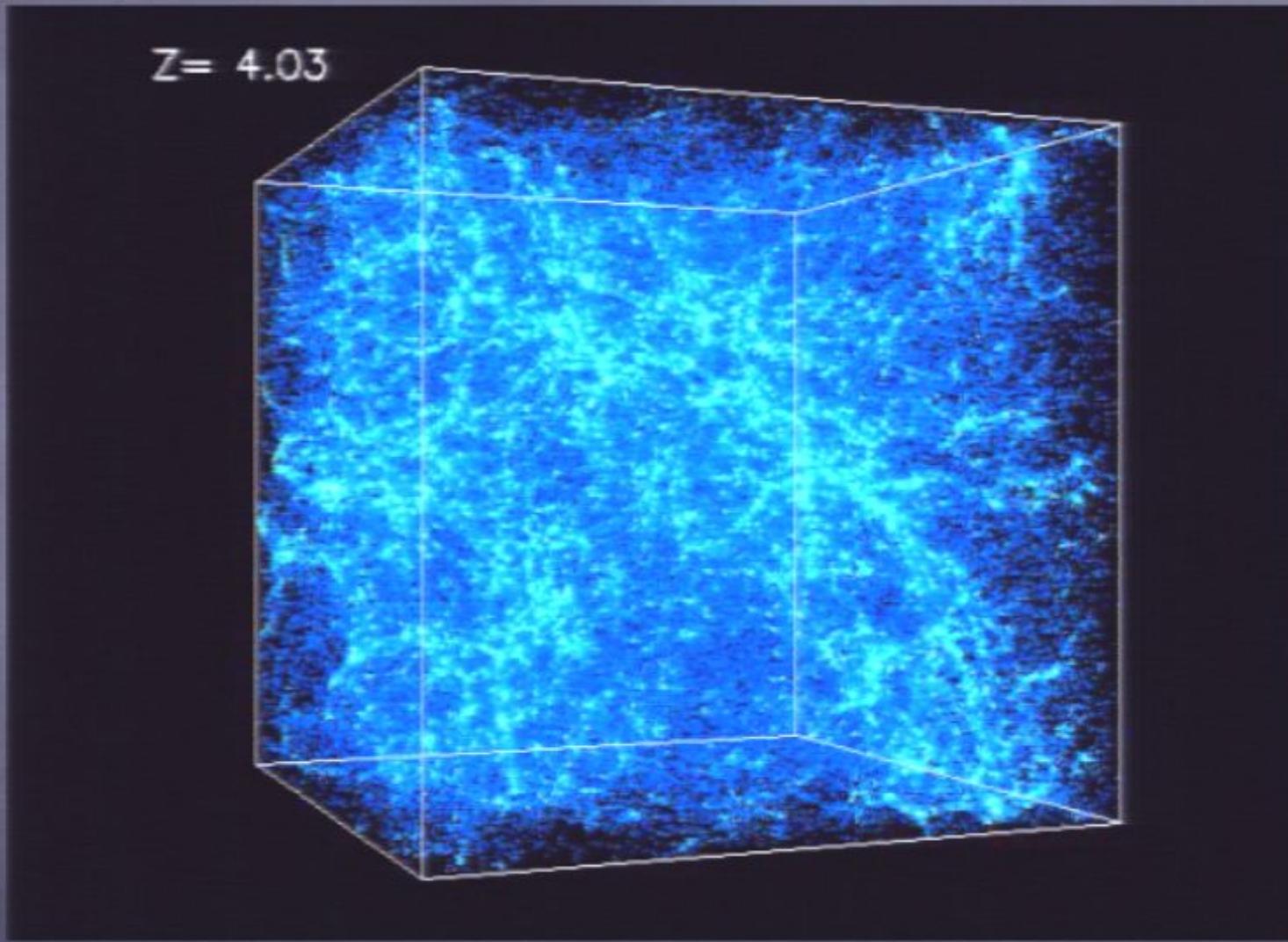
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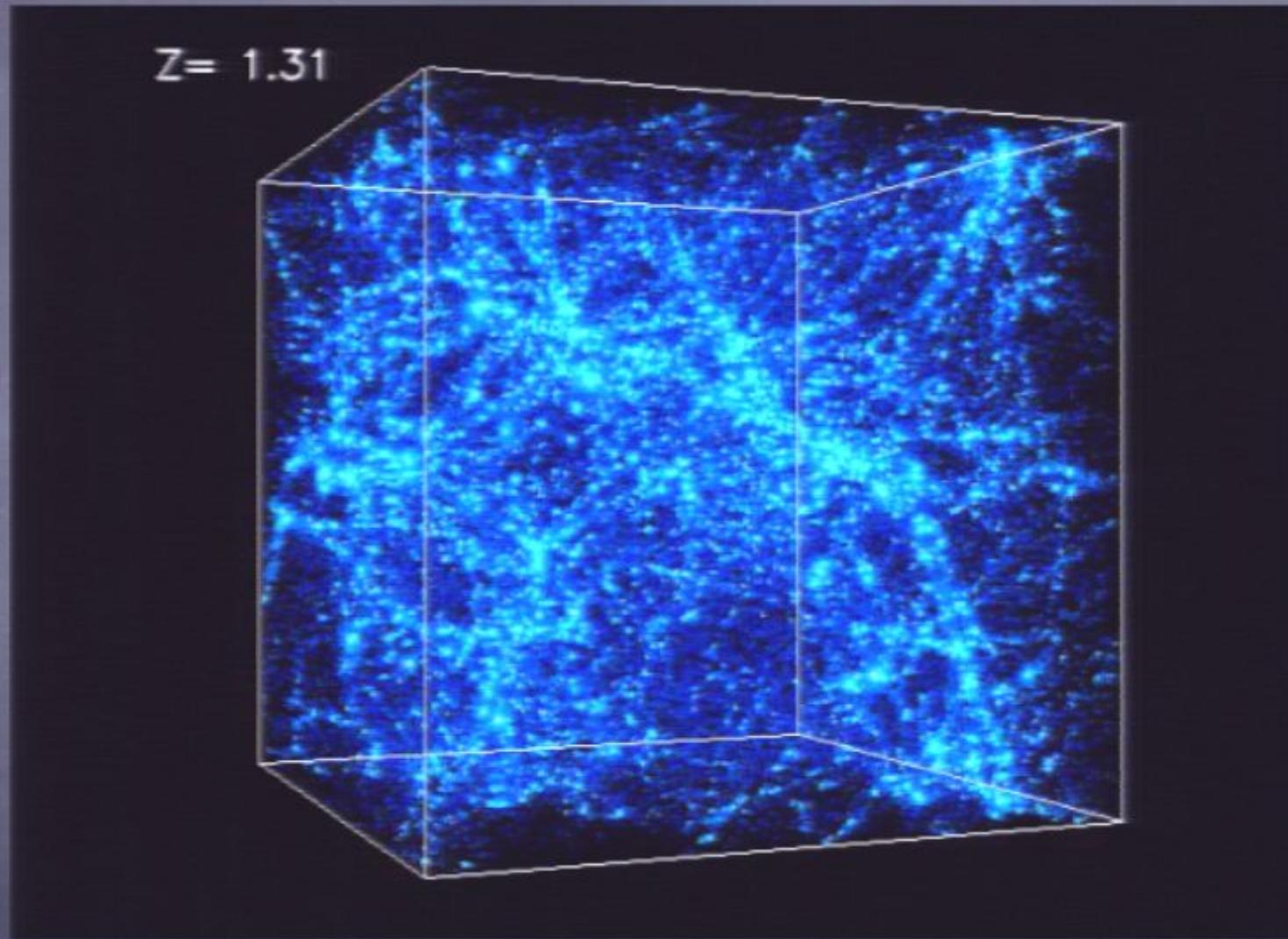
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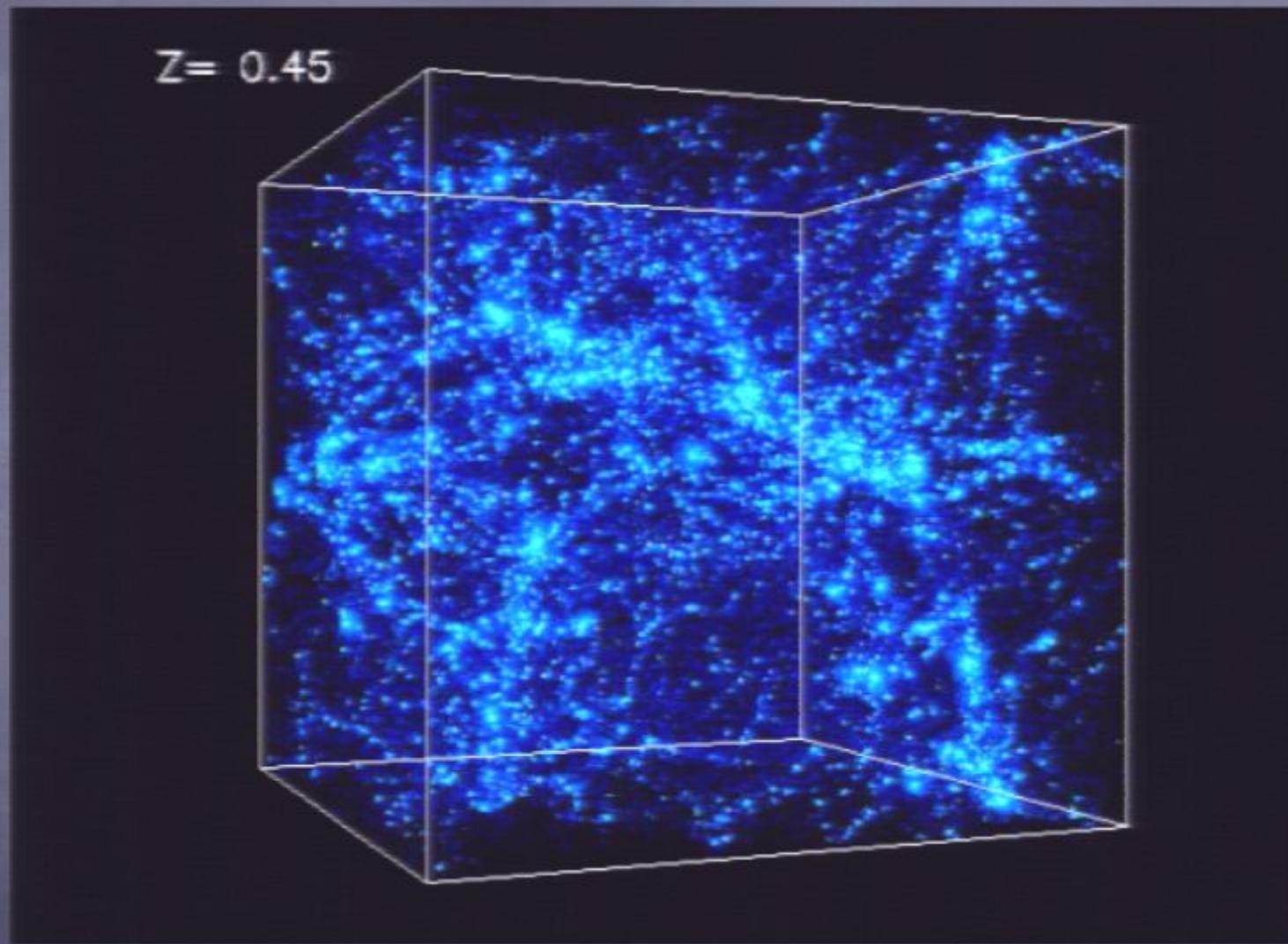
Other schemes to understand structure formation (e.g., perturbation theory) also have problems



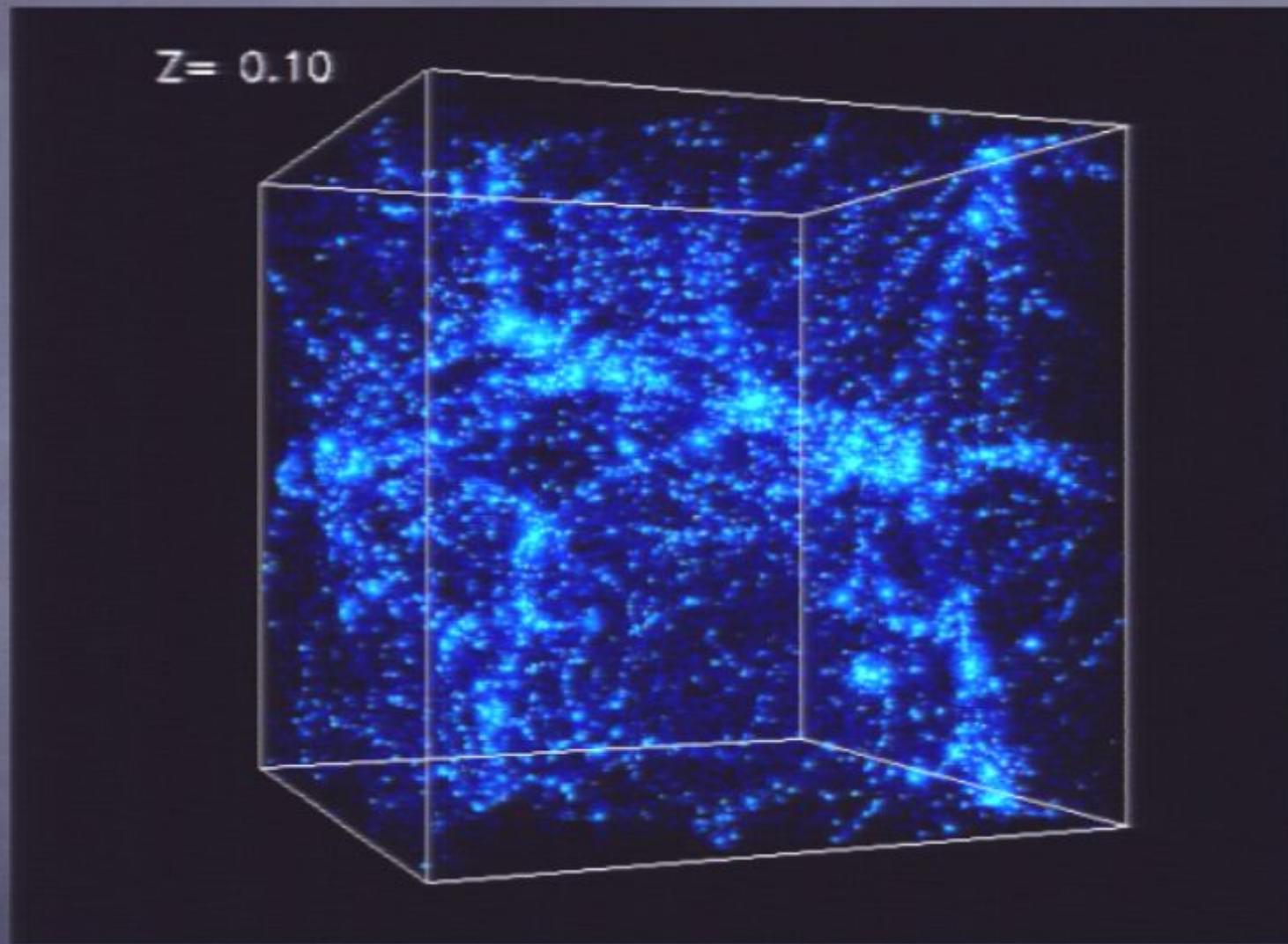
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Center for Cosmological Physics website



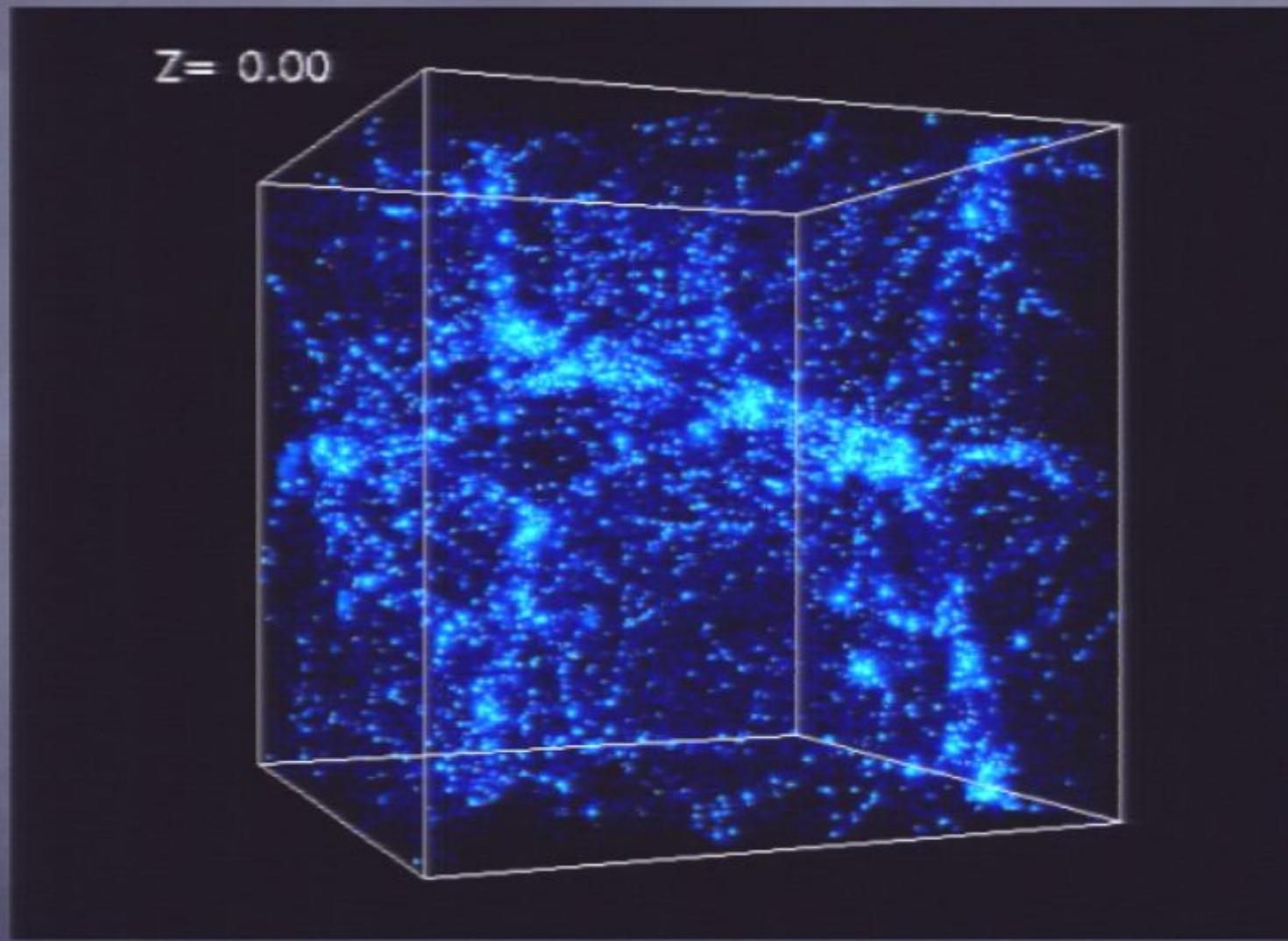
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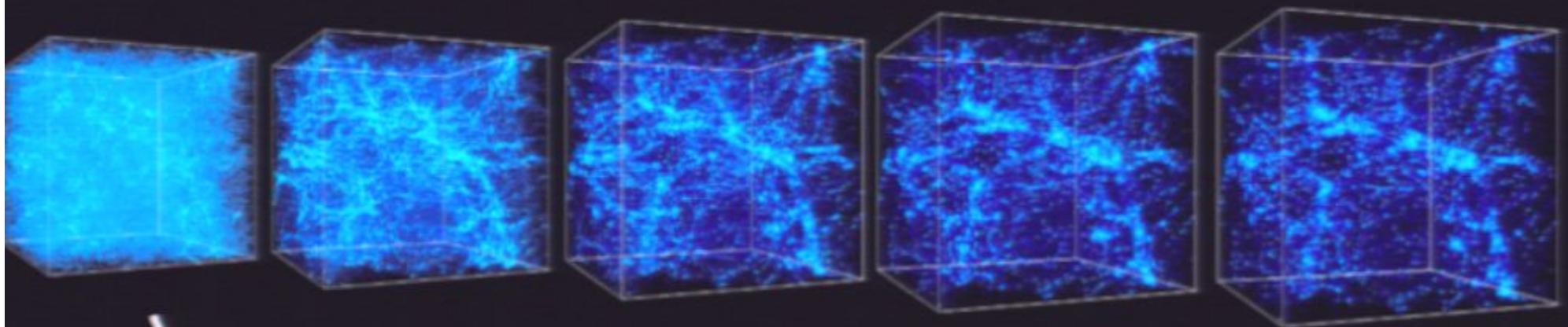
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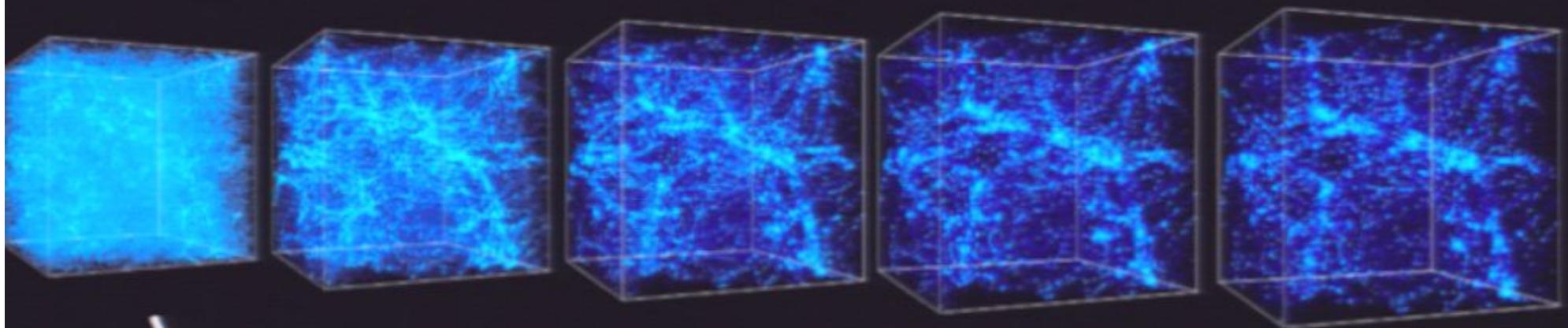


ICs:
mass fluctuation
associated with scale

$$\lambda \sim \frac{2\pi}{k}$$

power per logarithmic bin in k

$$k^3 \left(\frac{\delta \rho_k}{\rho} \right)^2 \equiv k^3 P(k) \equiv \Delta^2(k)$$



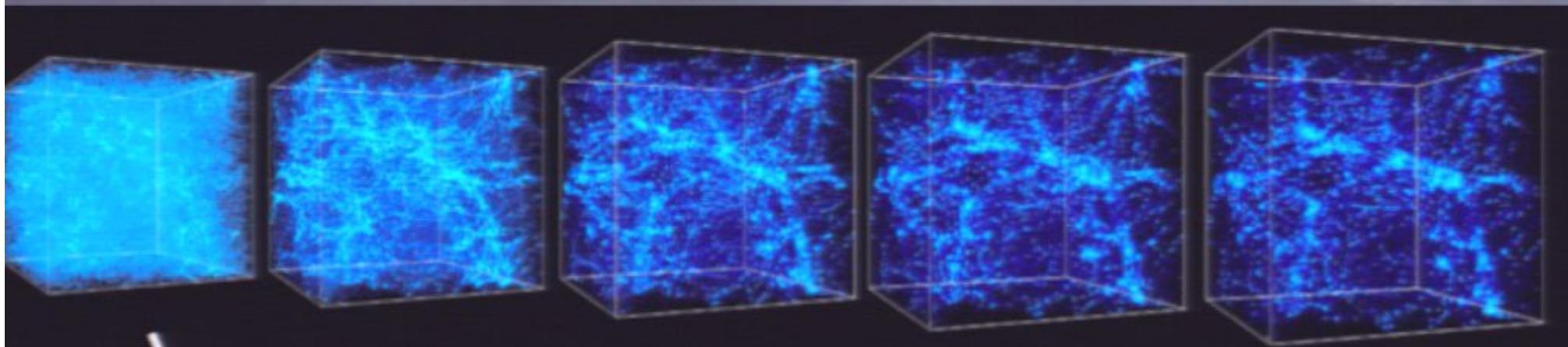
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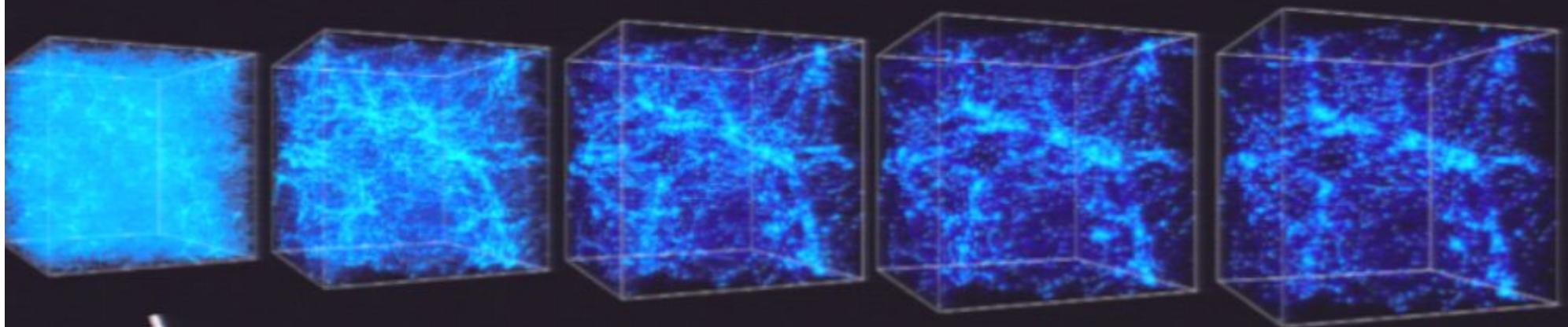
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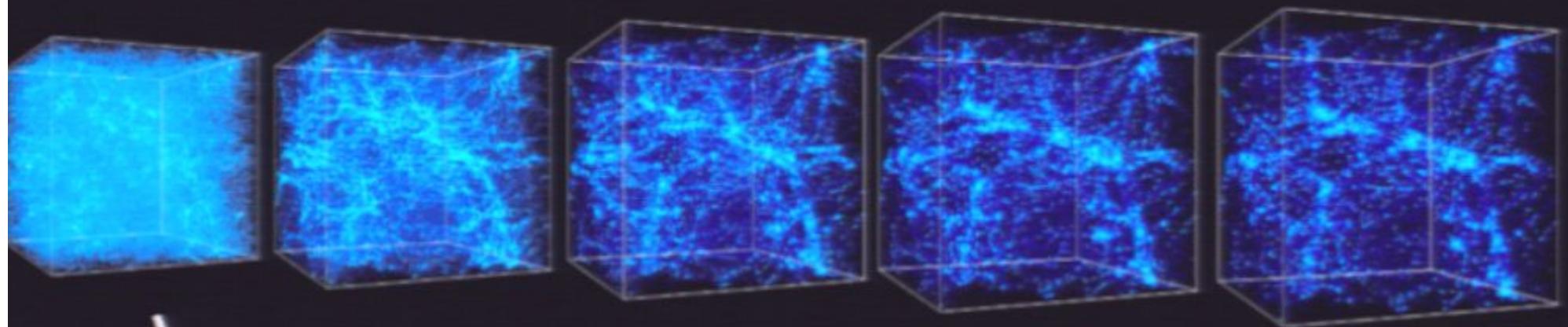
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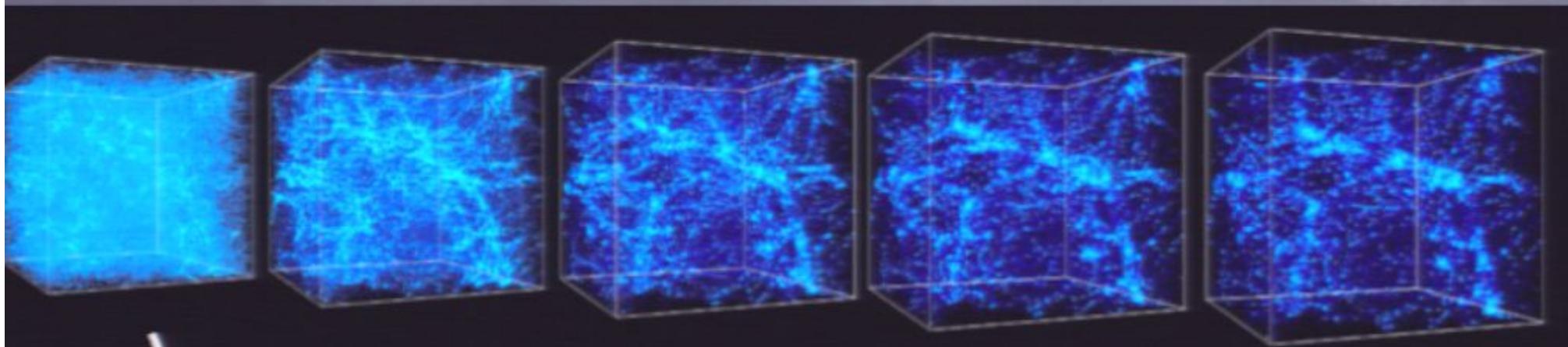
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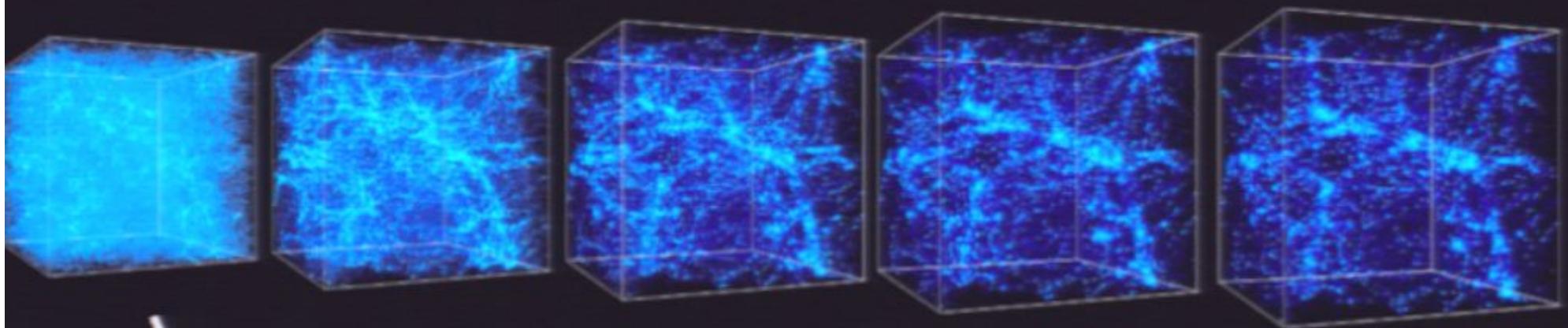


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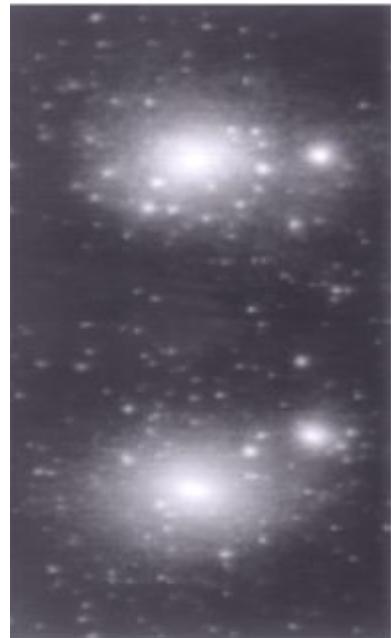
$$\begin{aligned}
\Delta^2(k) &= k^3 P(k) \\
&= k^3 P_{\text{primordial}}(k) T^2(k) \\
&\sim k^{n_{\text{eff}} + 3} \quad (\text{over small range in } k) \\
&\sim \text{const.} \quad (\text{i.e., } n_{\text{eff}} \simeq -3) \quad \text{for } k \rightarrow \infty
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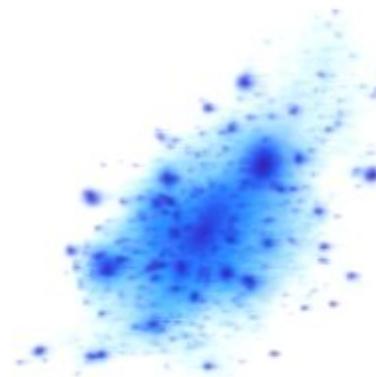
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Moore et al. 1999

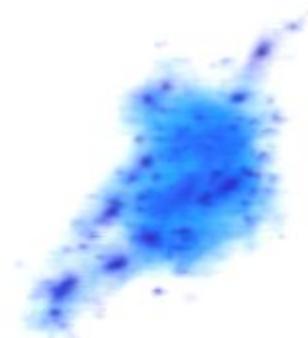


simulations with scale-free power-spectra $P(k) \propto k^n$

$$n = -1$$



$$n = -2.5$$



galaxy
 $n = -2.2$



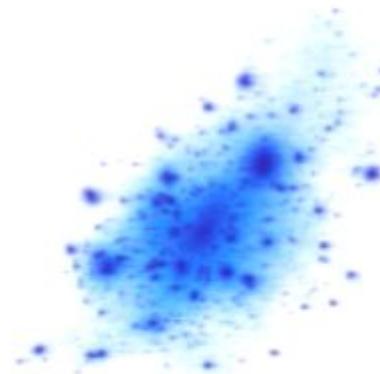
Moore et al. 1999



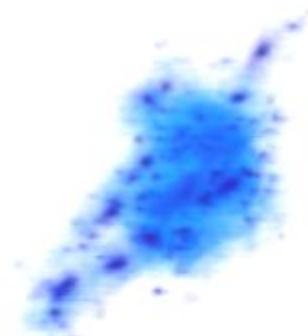
cluster
 $n = -1.8$

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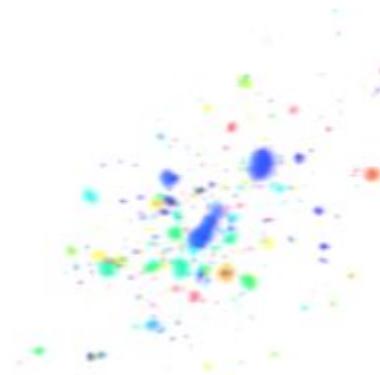
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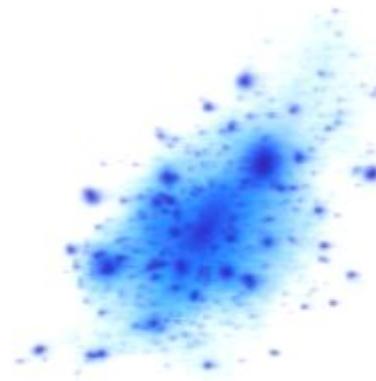
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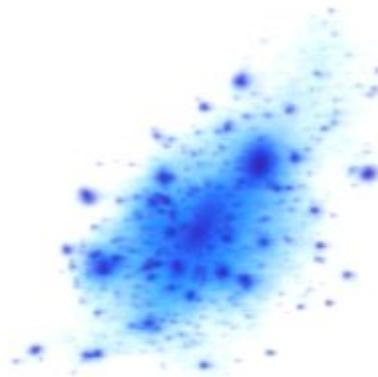
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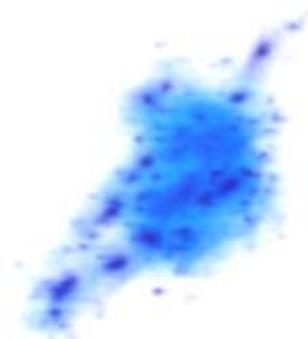
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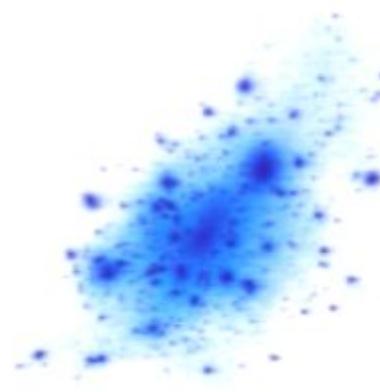
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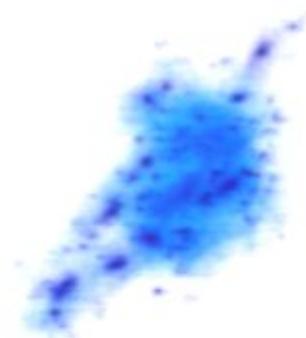
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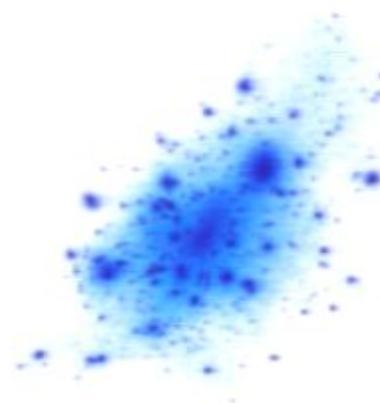
Moore et al. 1999



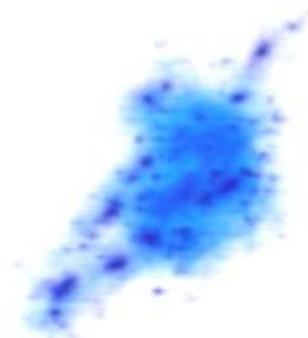
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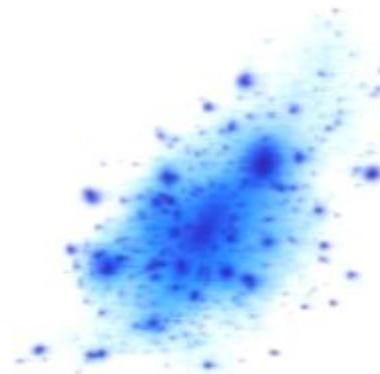
Moore et al. 1999



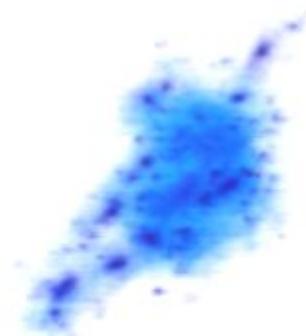
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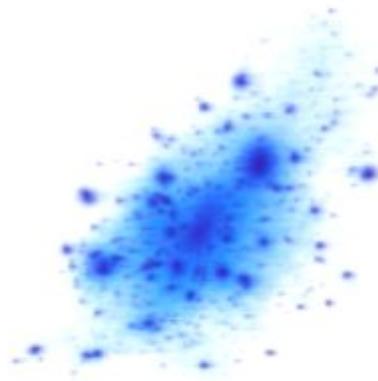
Moore et al. 1999



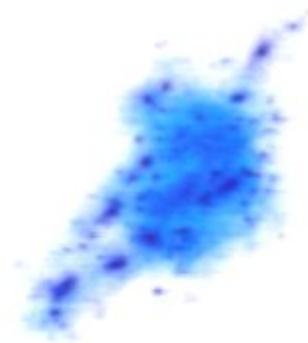
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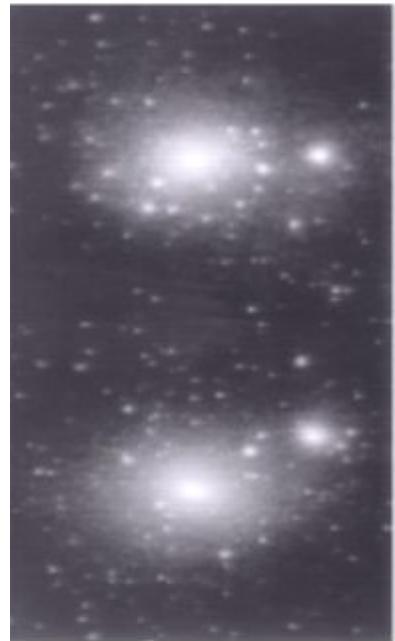
$n = -1$



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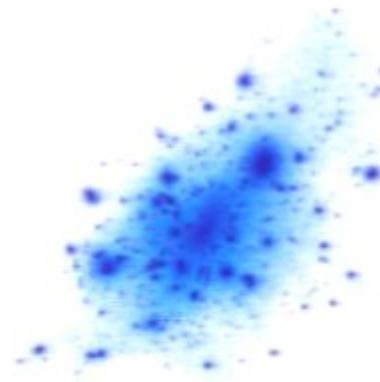
Moore et al. 1999



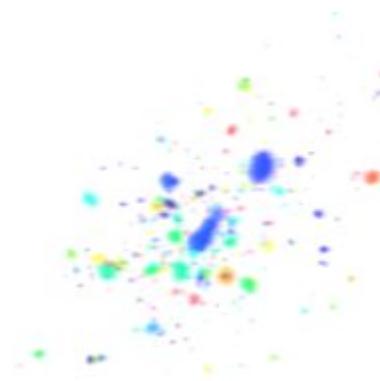
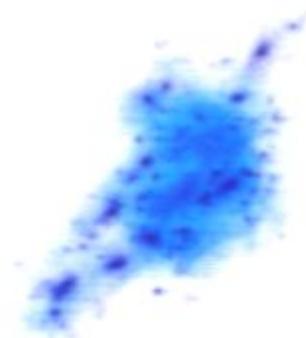
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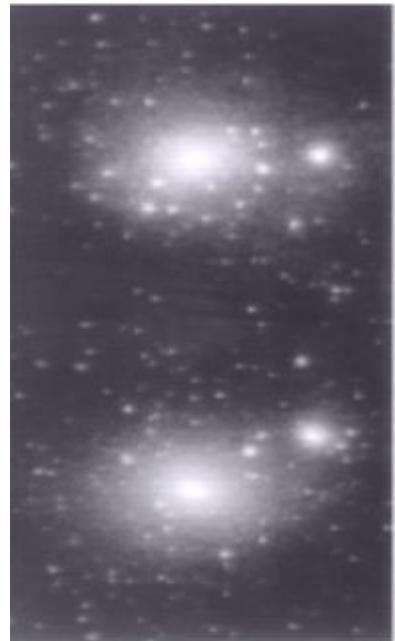
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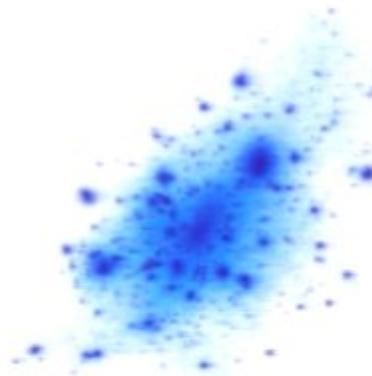
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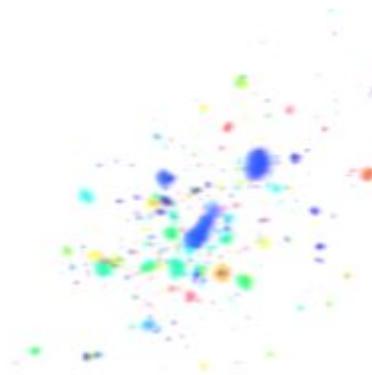
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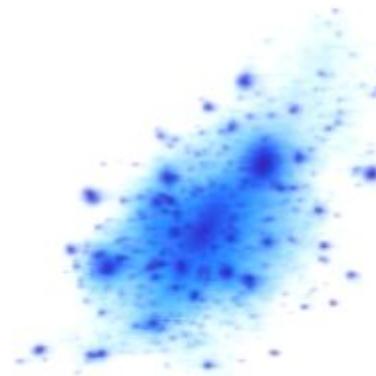
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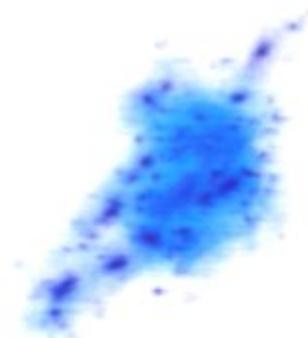
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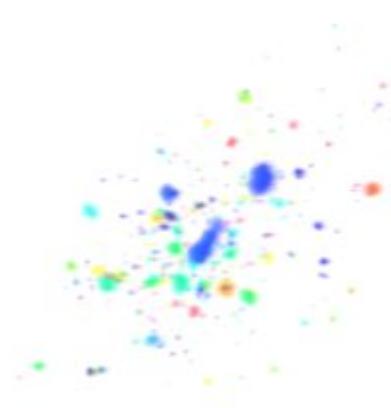
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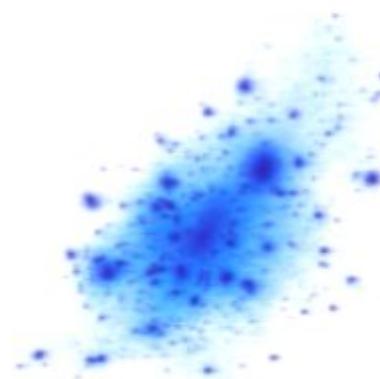
Moore et al. 1999



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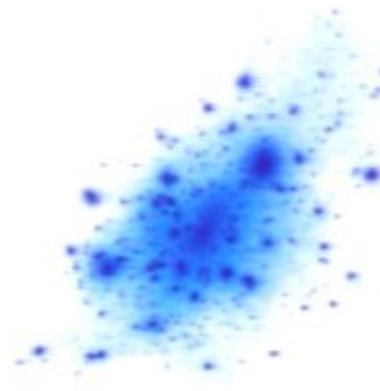
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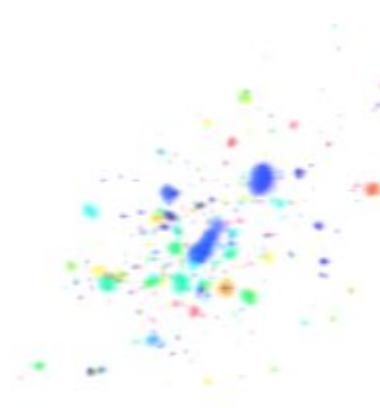
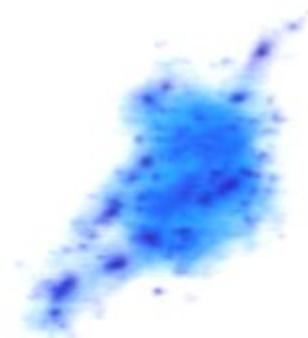
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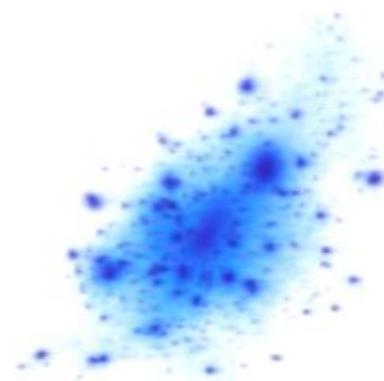
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Moore et al. 1999

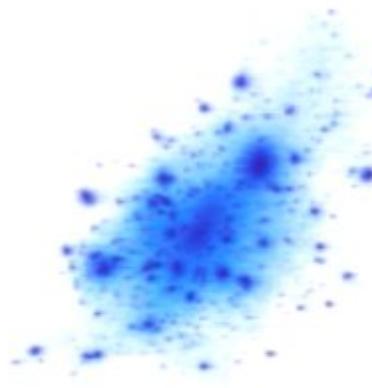


cluster
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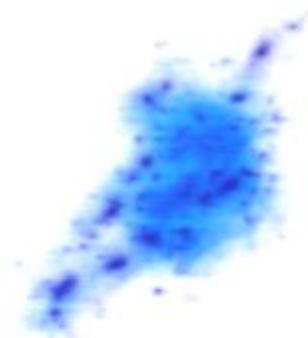
galaxy
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simulations with scale-free power-spectra $P(k) \propto k^n$

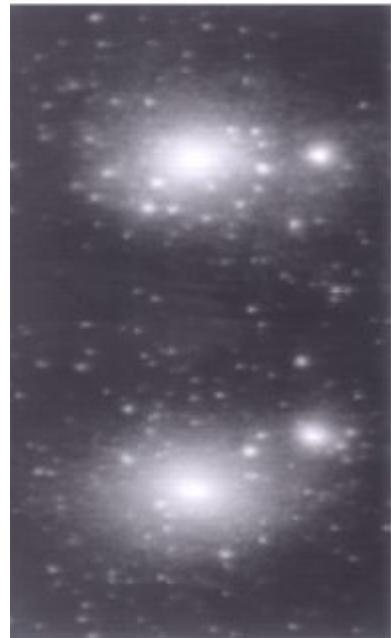
$$n = -1$$



$$n = -2.5$$



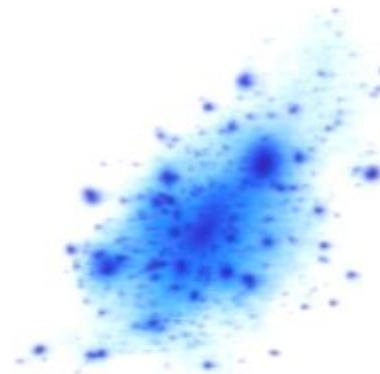
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cluster
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simulations with scale-free power-spectra $P(k) \propto k^n$

$n = -1$



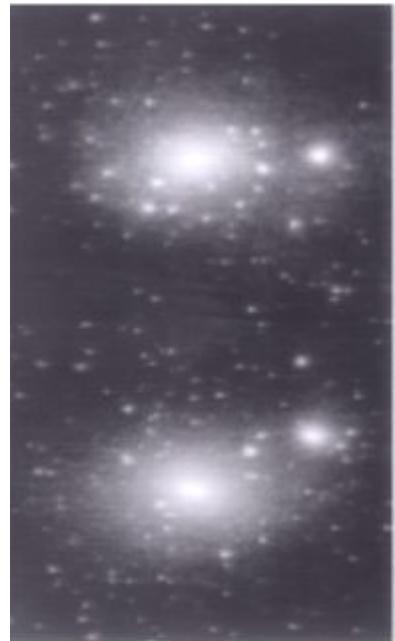
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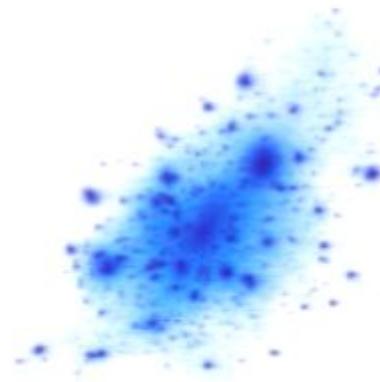
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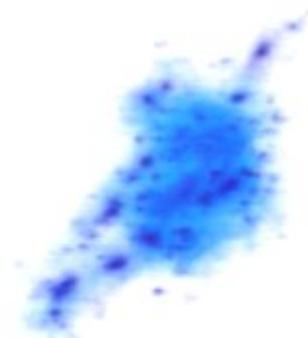
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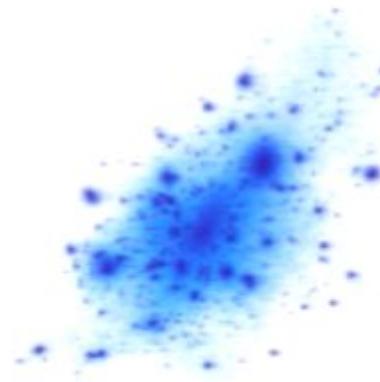
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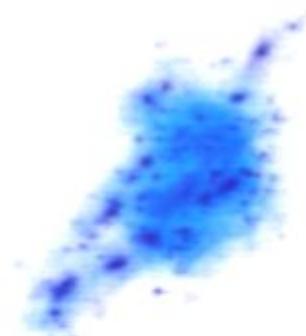
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EARLY SUPERSYMMETRIC COLD DARK MATTER SUBSTRUCTURE

JÜRGEN DIEMAND, MICHAEL KUHLEN, AND PIERO MADAU

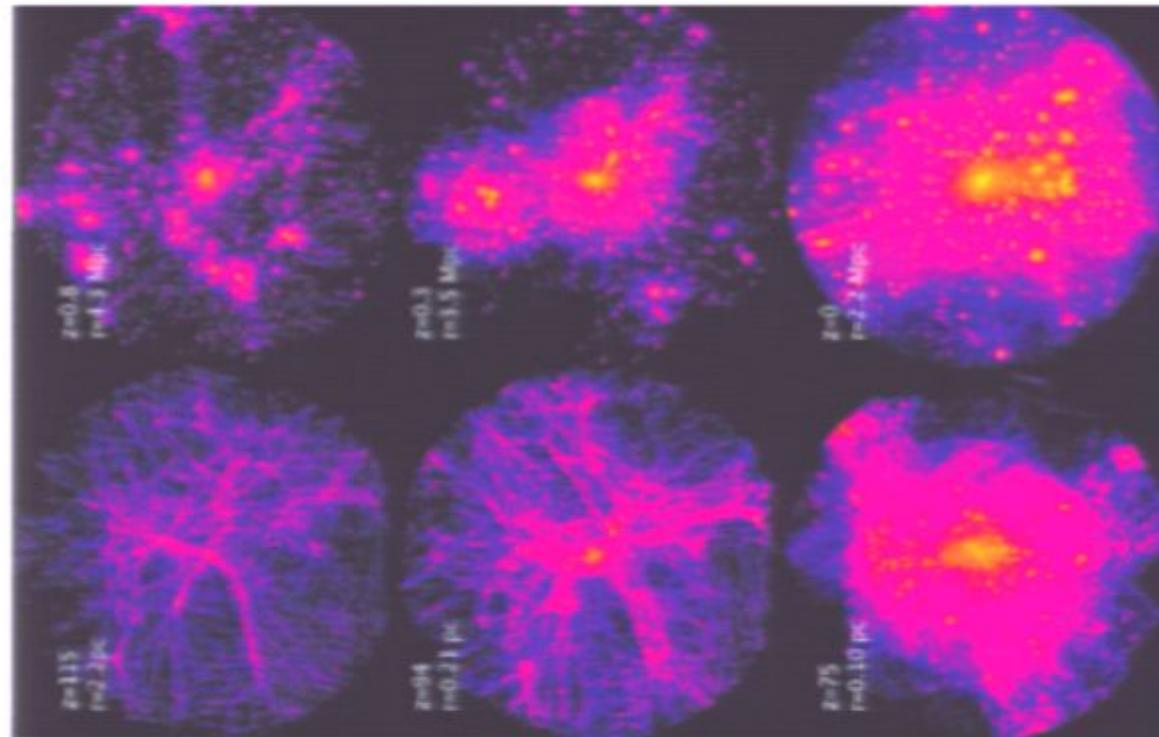
Department of Astronomy and Astrophysics, 477 Clark Kerr Hall, University of California, Santa Cruz, CA 95064; diemand@ucolick.org

Received 2006 February 26; accepted 2006 May 31

ABSTRACT

Earth-mass "microhalos" may be the first objects to virialize in the early universe. Their ability to survive the hierarchical clustering process as substructure in the larger halos that form subsequently has implications for dark matter detection experiments. We present a large N -body simulation of early substructure in a supersymmetric cold dark matter (SUSY-CDM) scenario characterized by an exponential cutoff in the power spectrum at $M_c = 10^{-6} M_\odot$. The simulation resolves a $0.014 M_\odot$ parent SUSY halo at $z = 75$ with 14 million particles. On these scales, the effective index of the power spectrum approaches -3 , and a range of mass scales collapses almost simultaneously.

galactic scale



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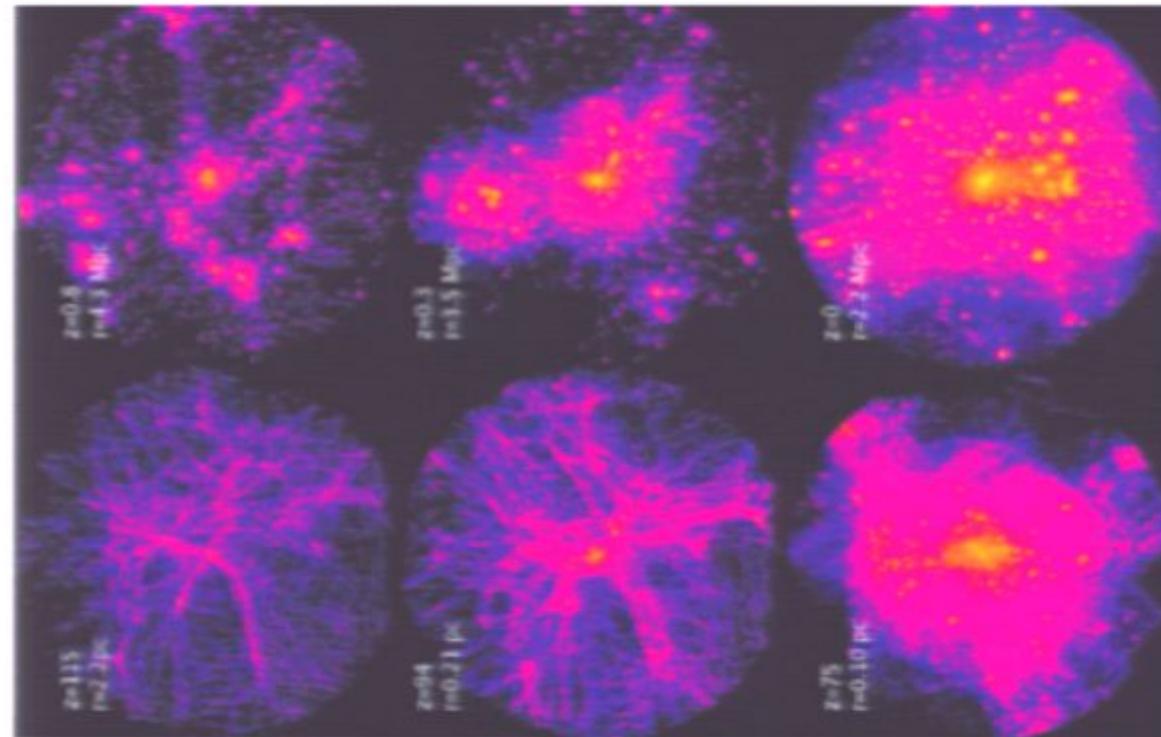
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galactic scale



time →

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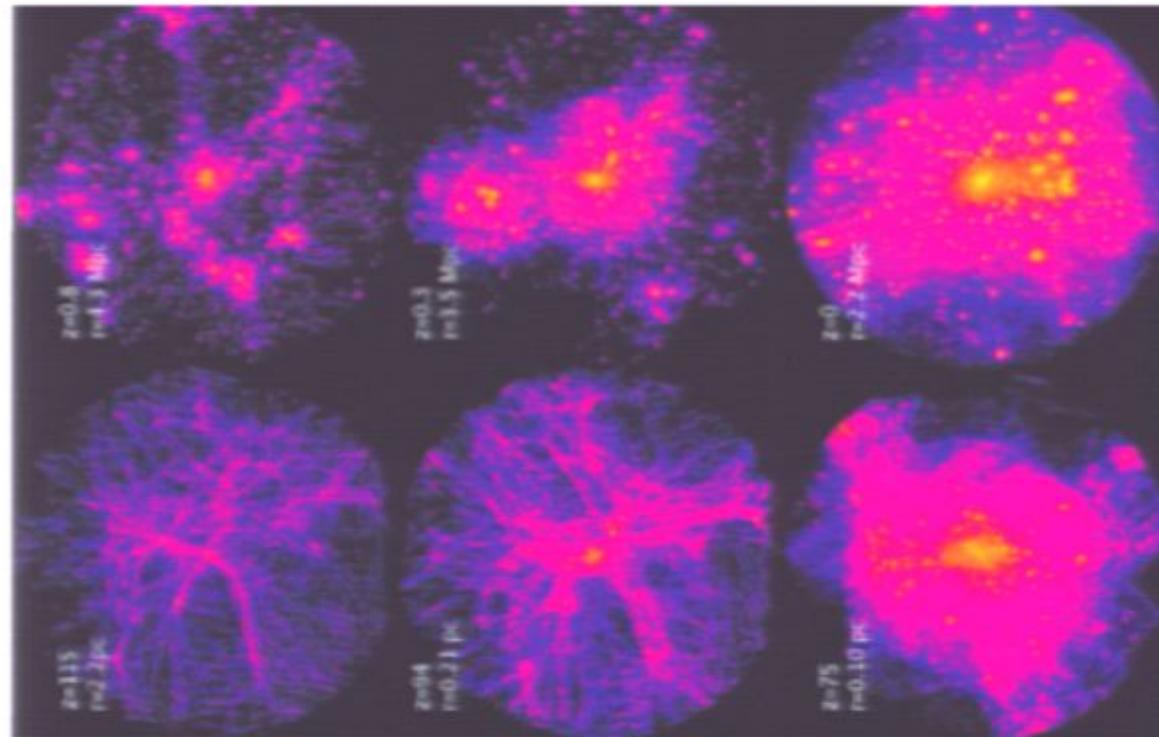
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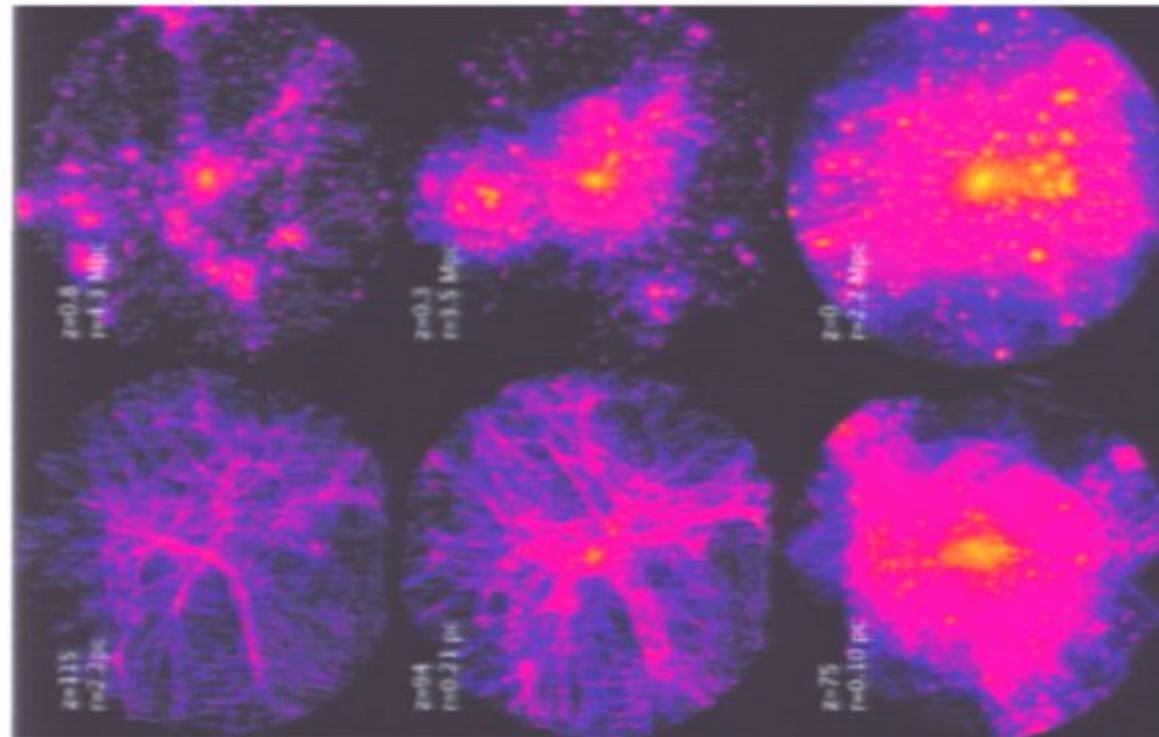
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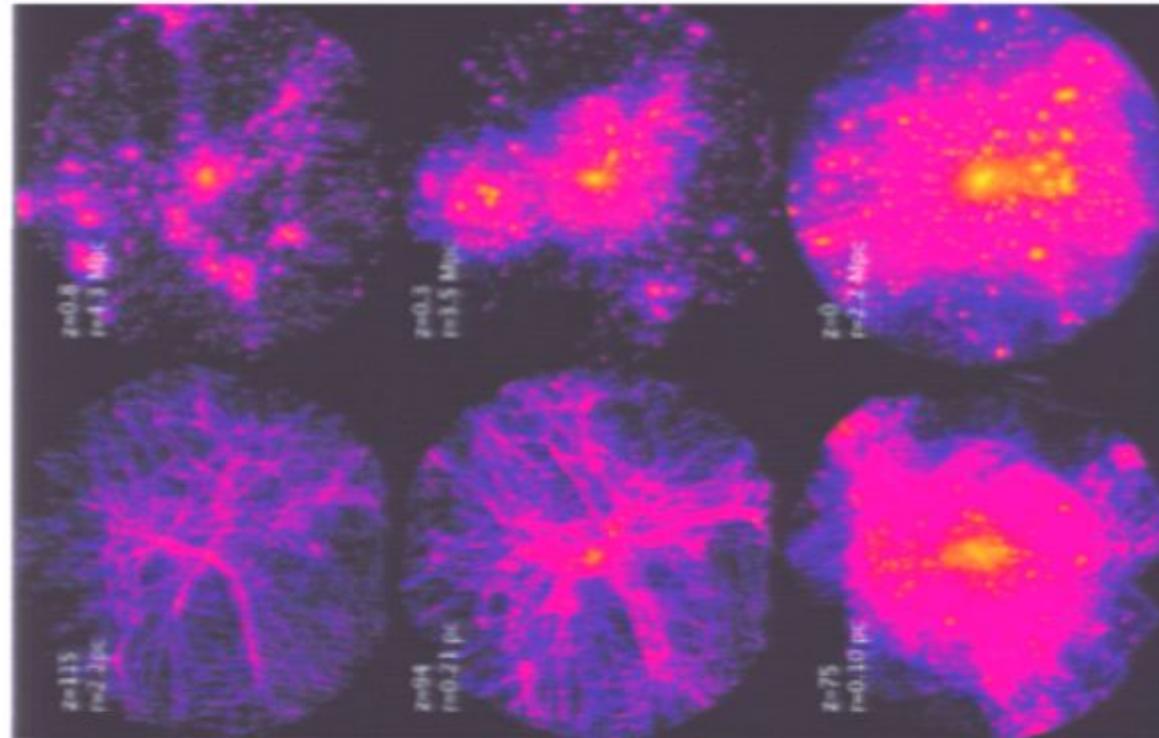
Department of Astronomy and Astrophysics, 477 Clark Kerr Hall, University of California, Santa Cruz, CA 95064; diemand@ucolick.org

Received 2006 February 26; accepted 2006 May 31

ABSTRACT

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galactic scale



time →

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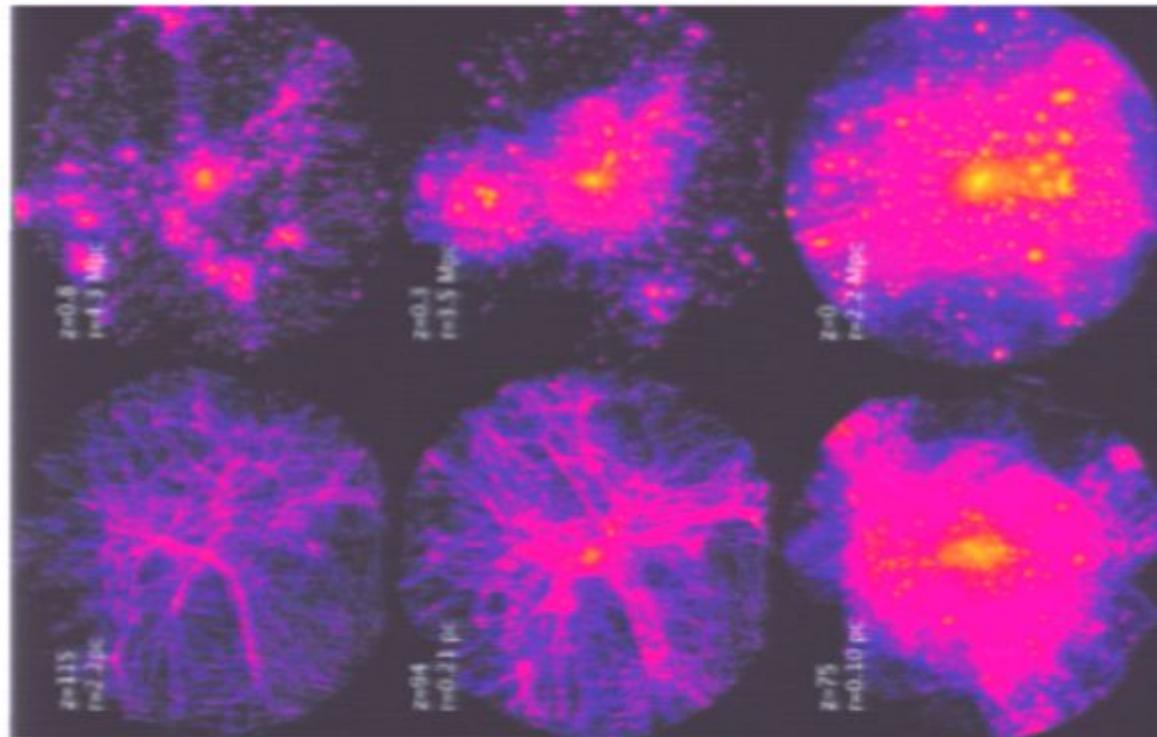
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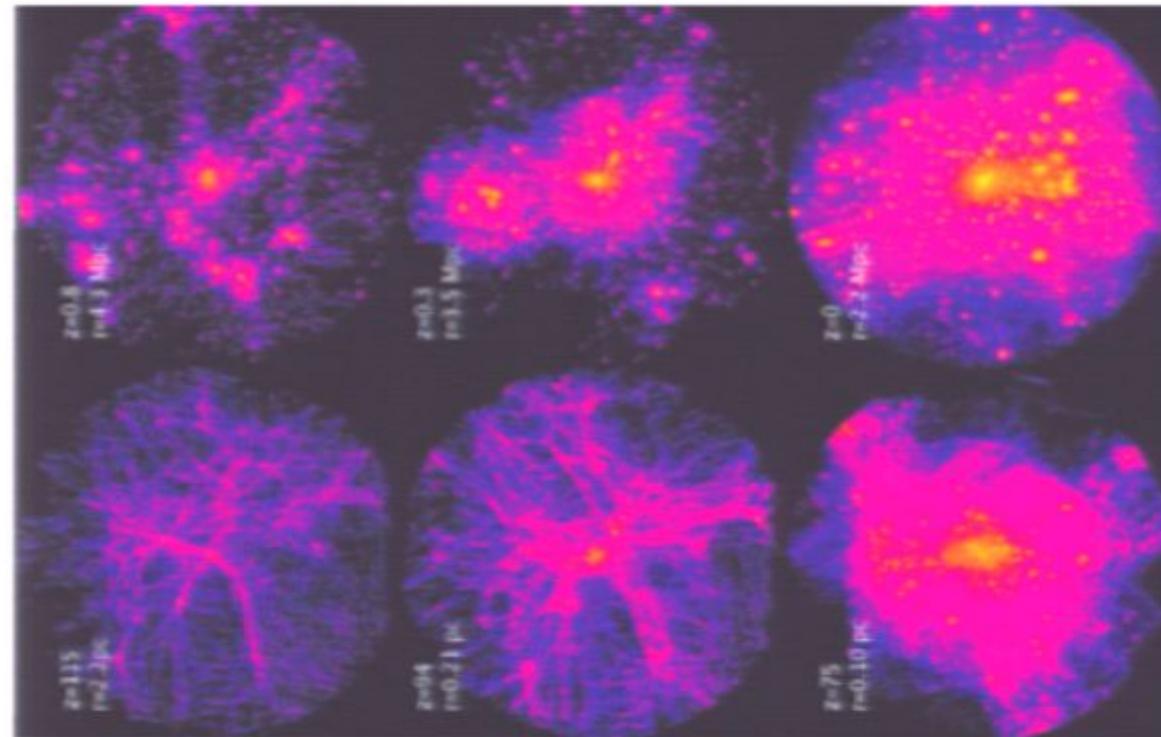
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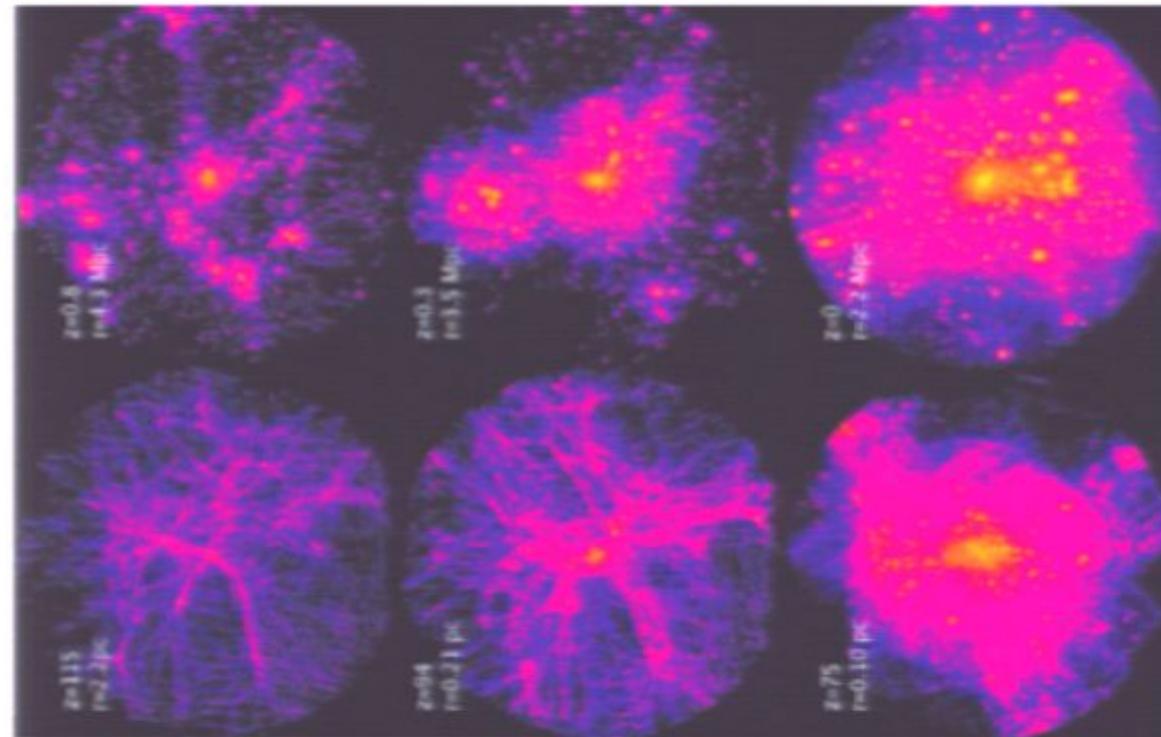
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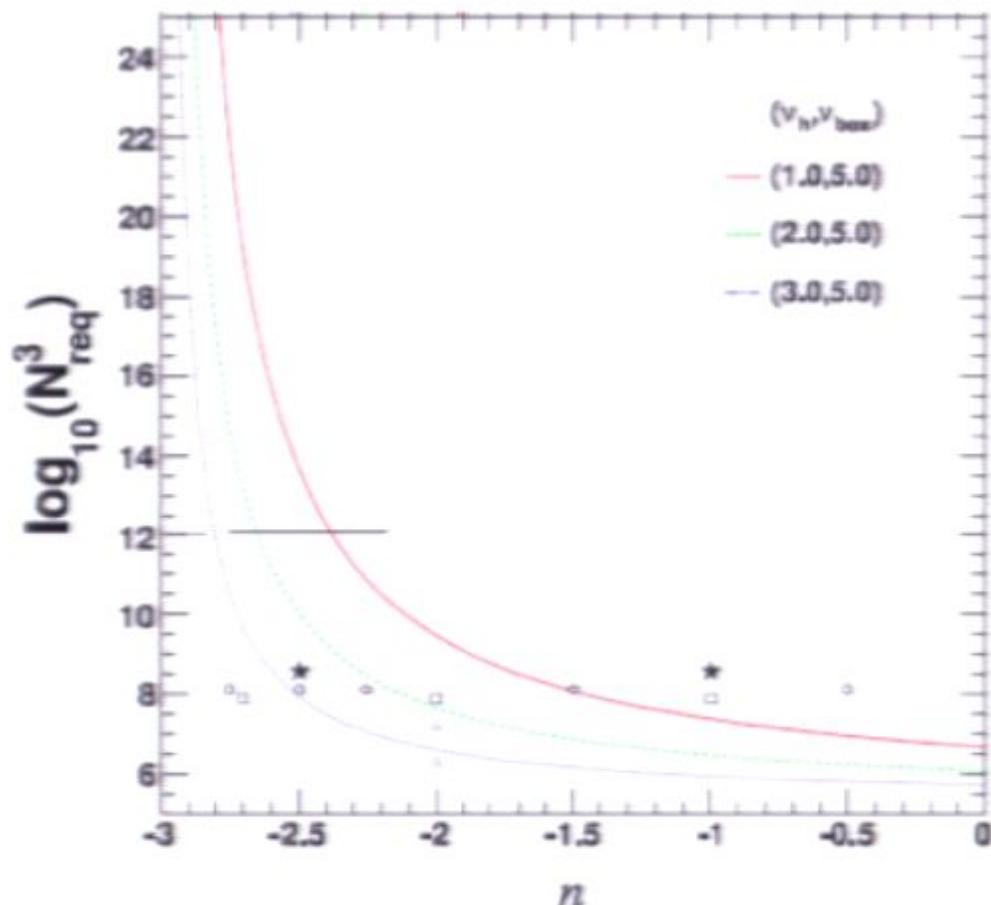
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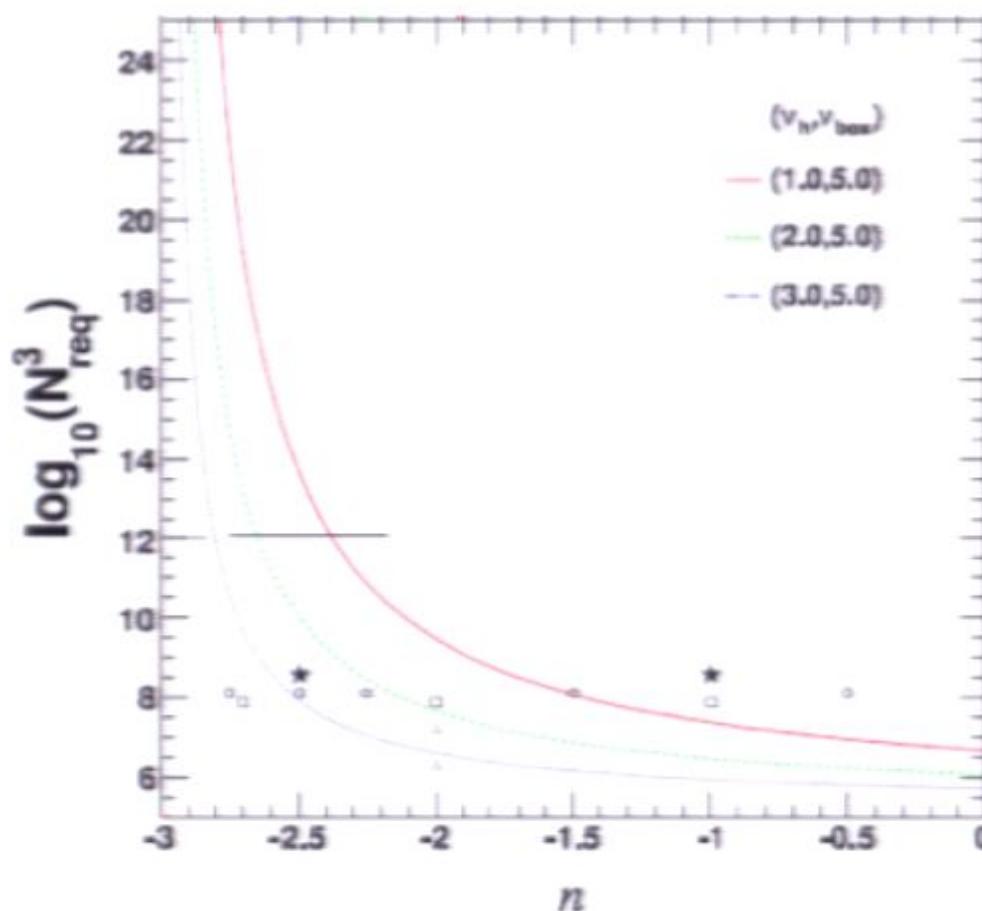


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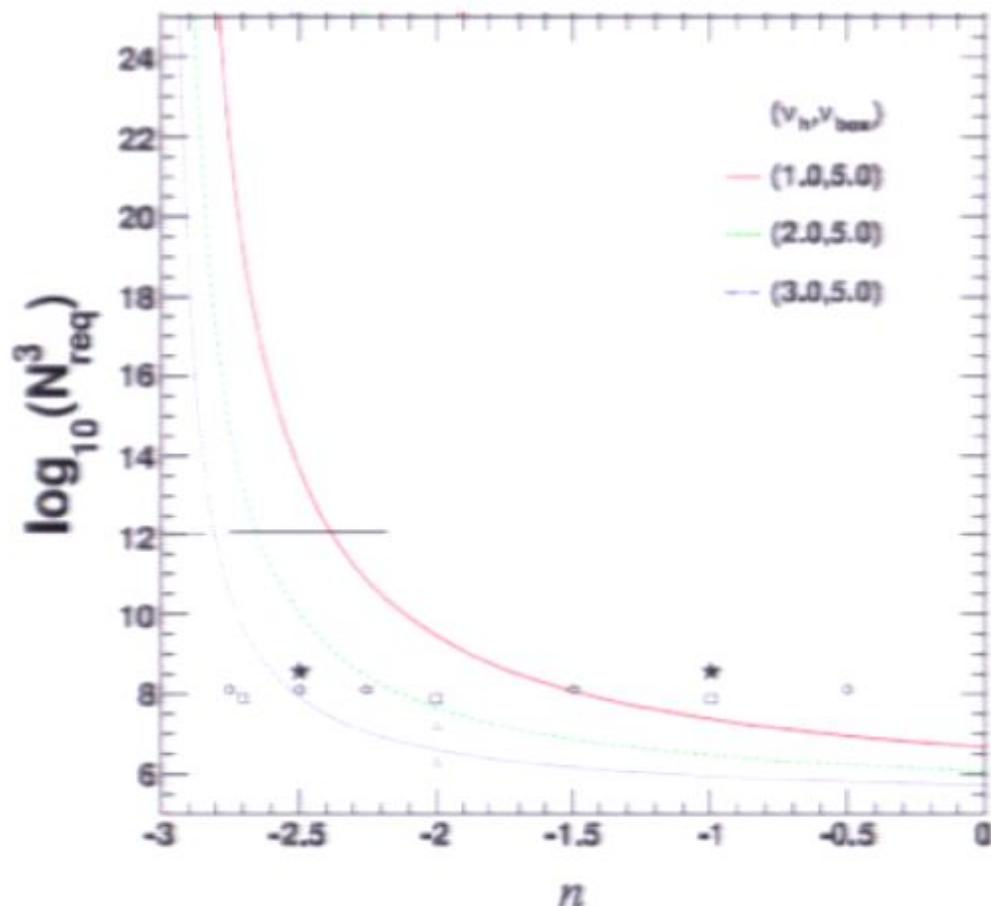
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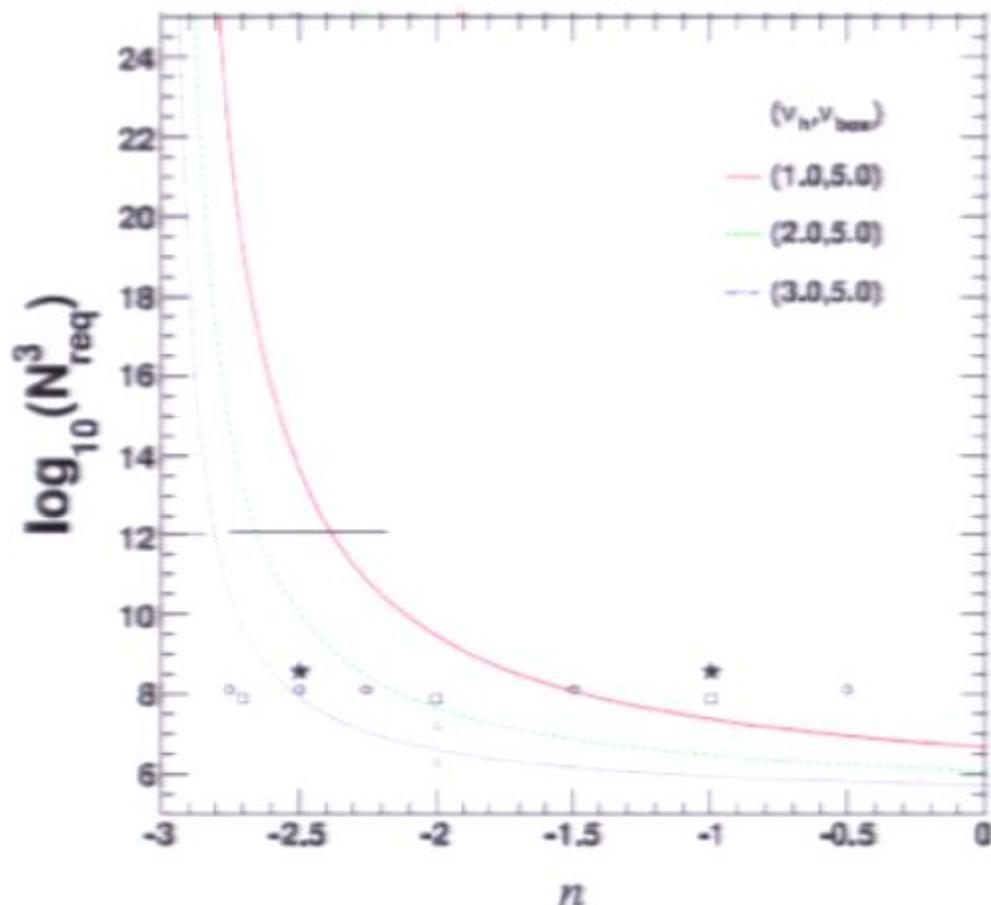
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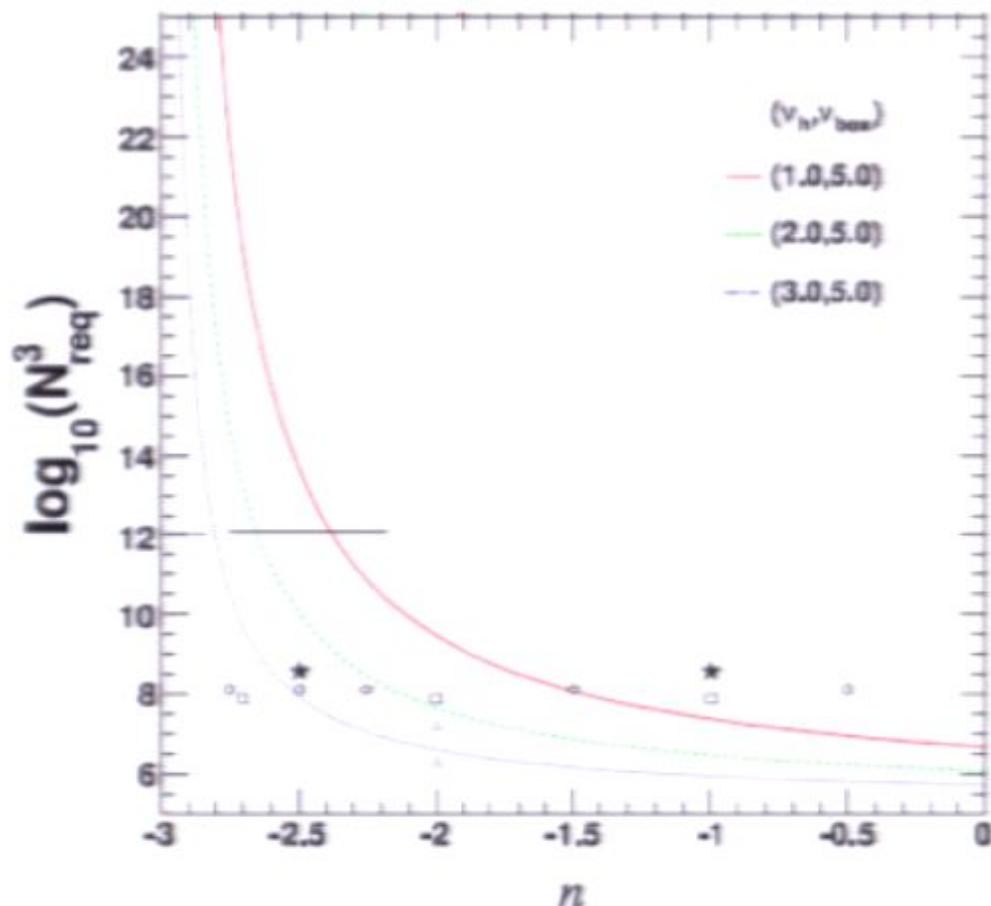
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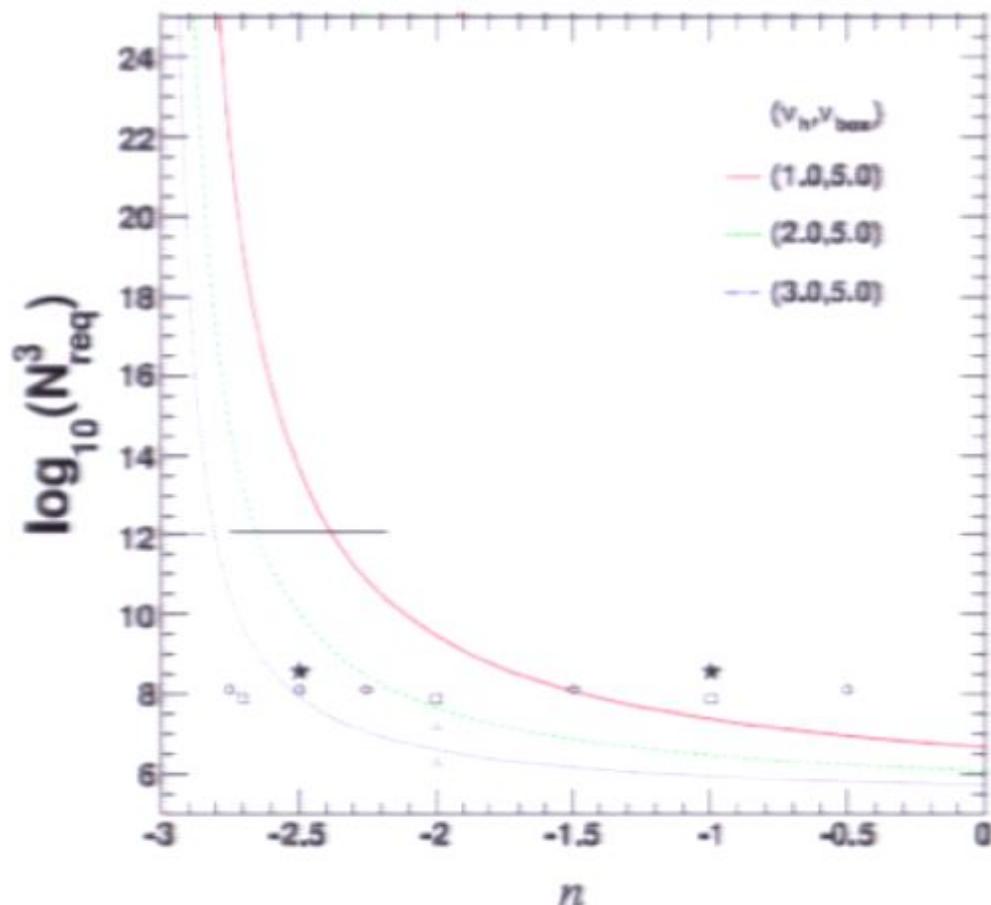
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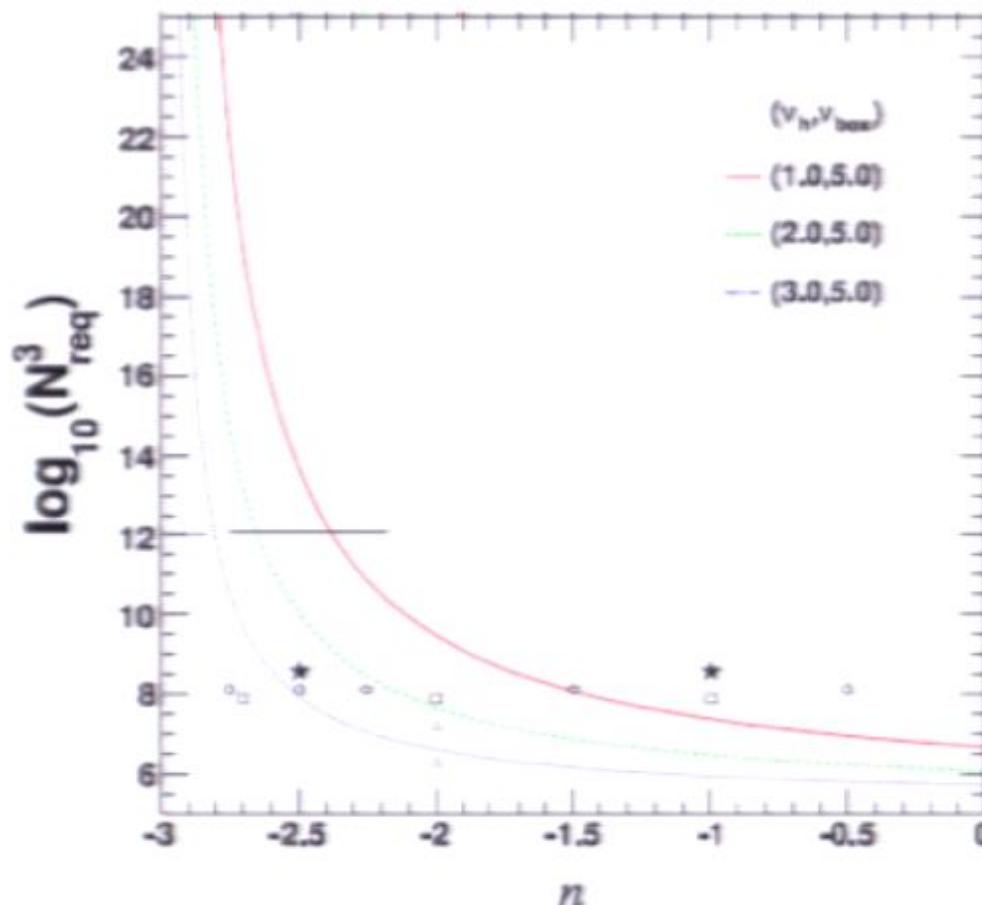
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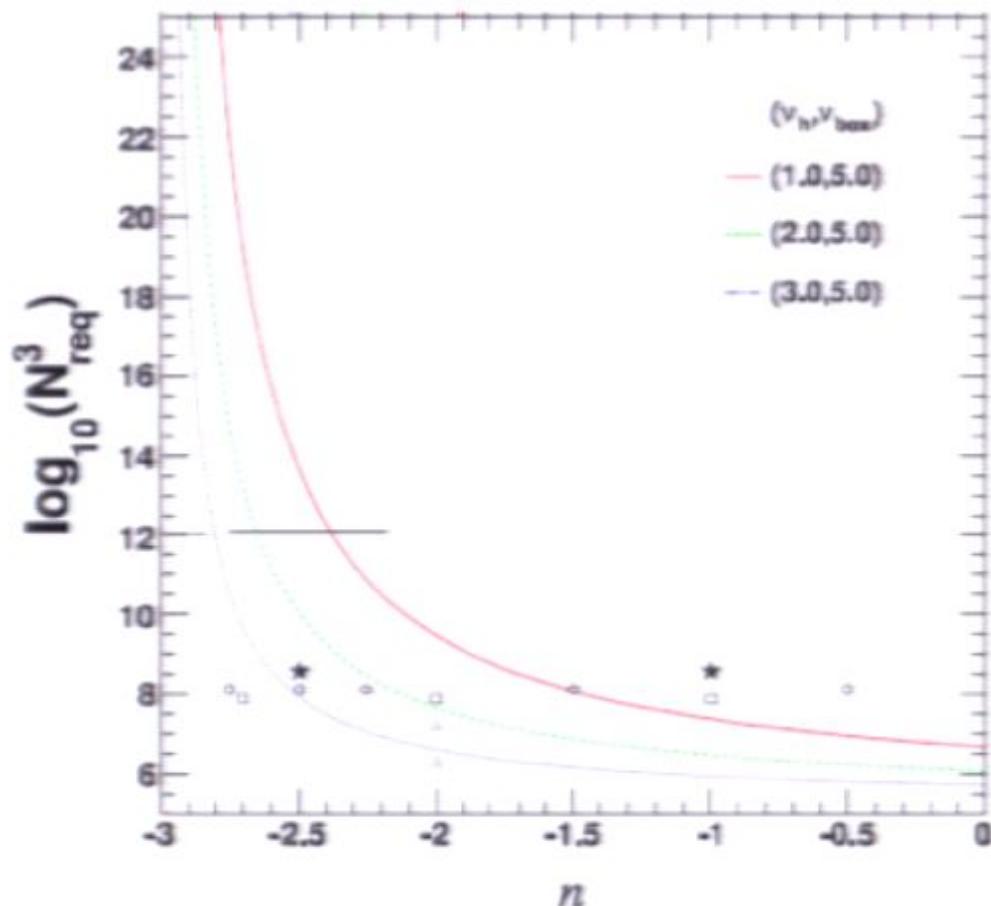
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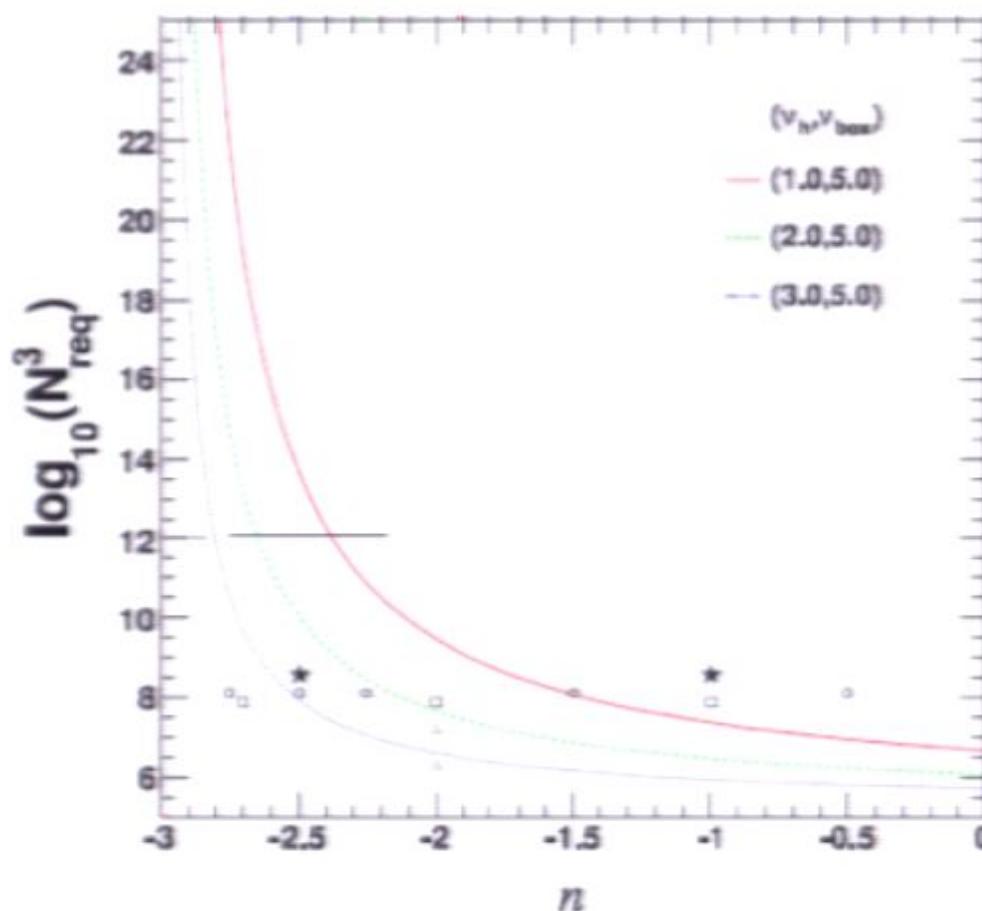
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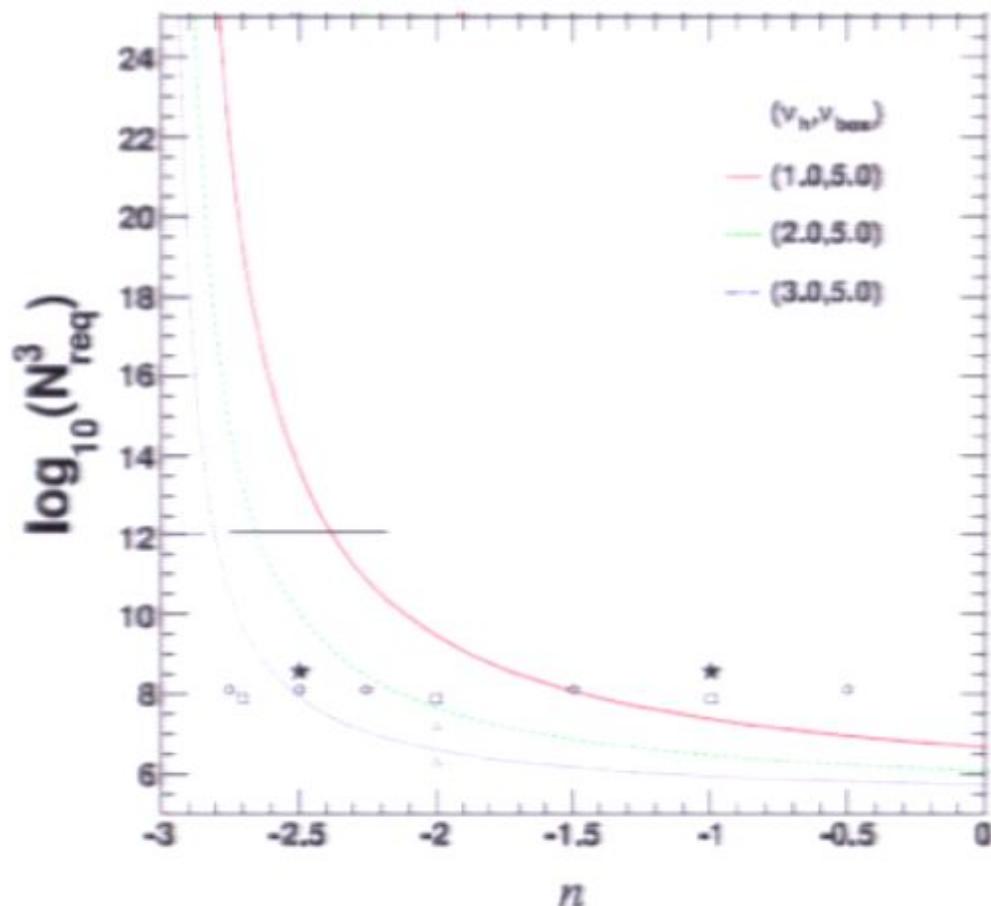
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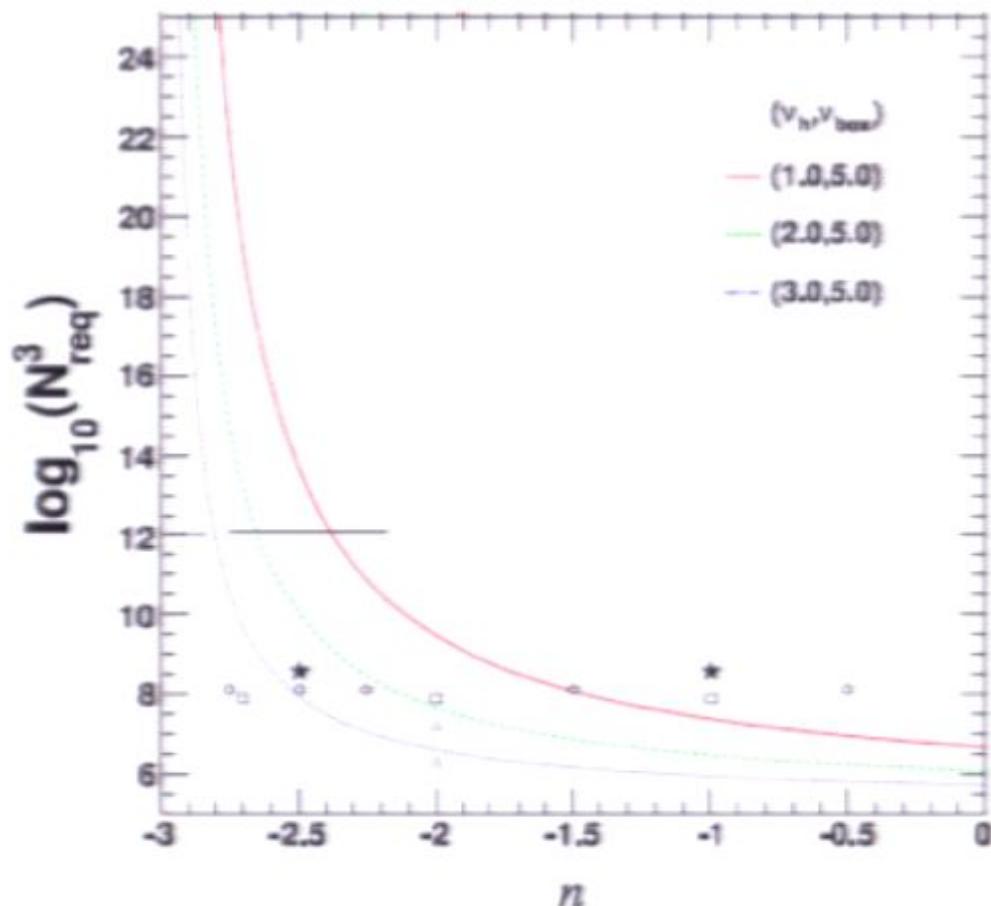
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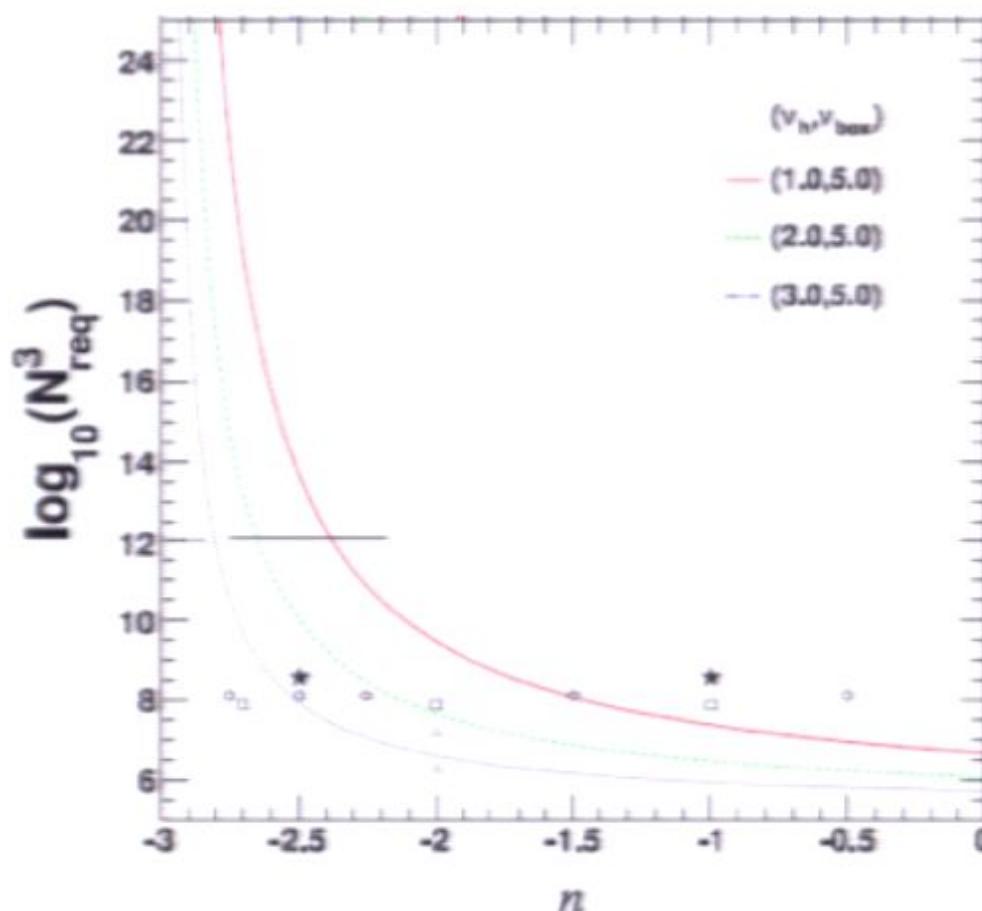
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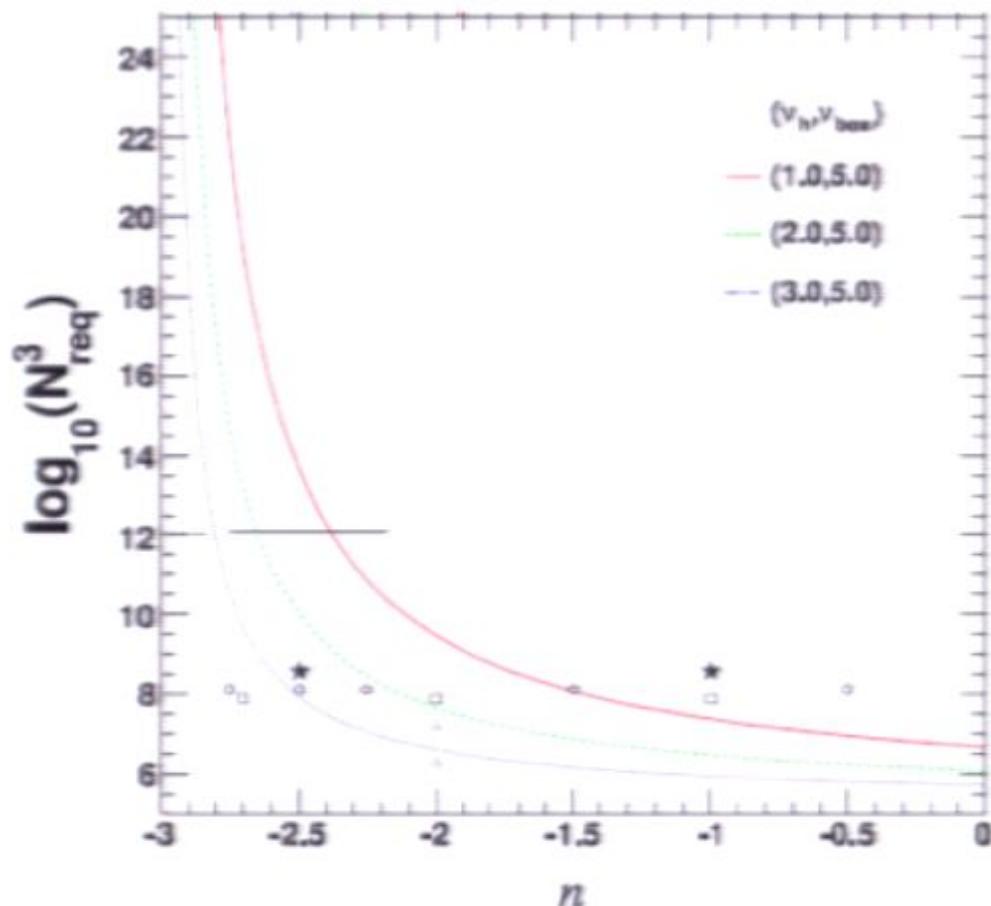
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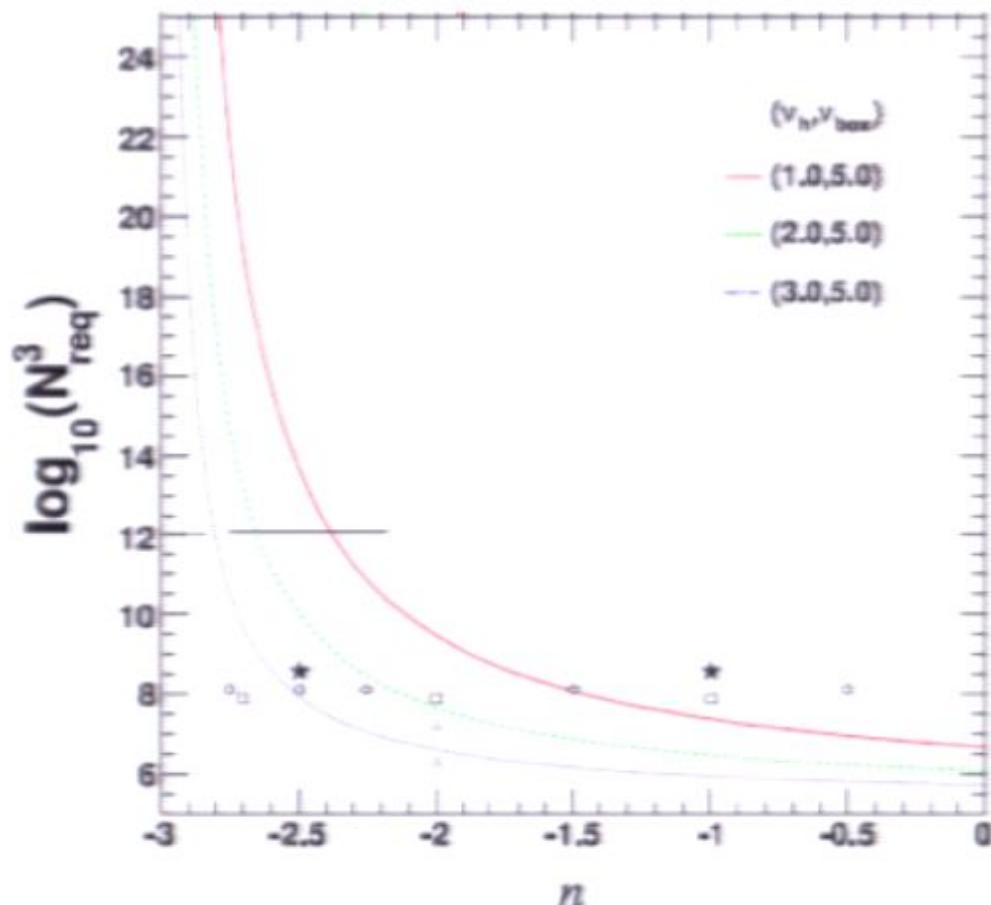
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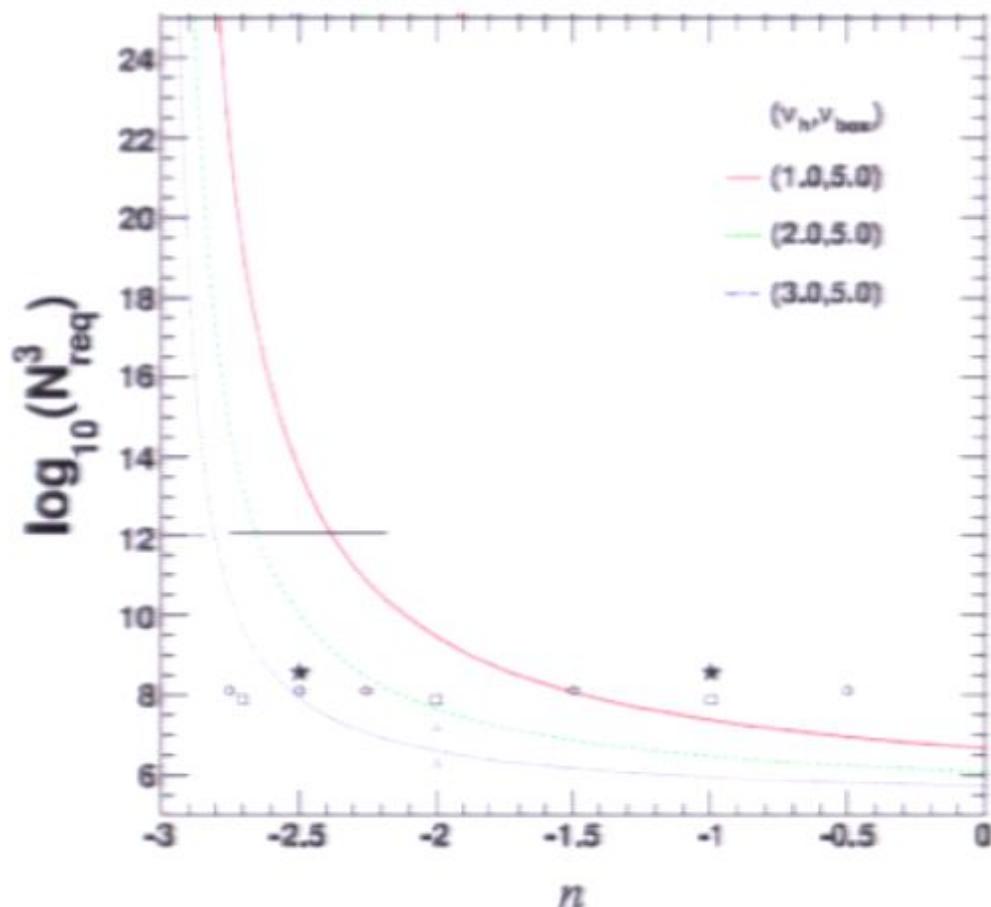
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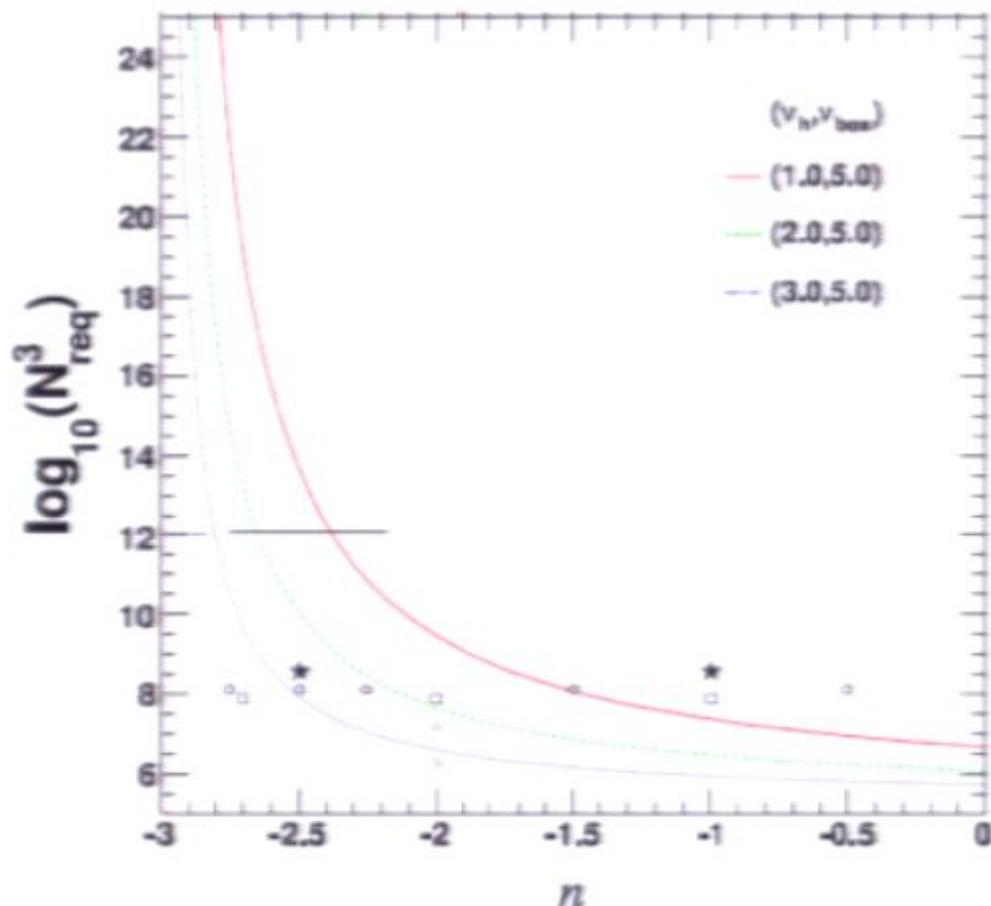
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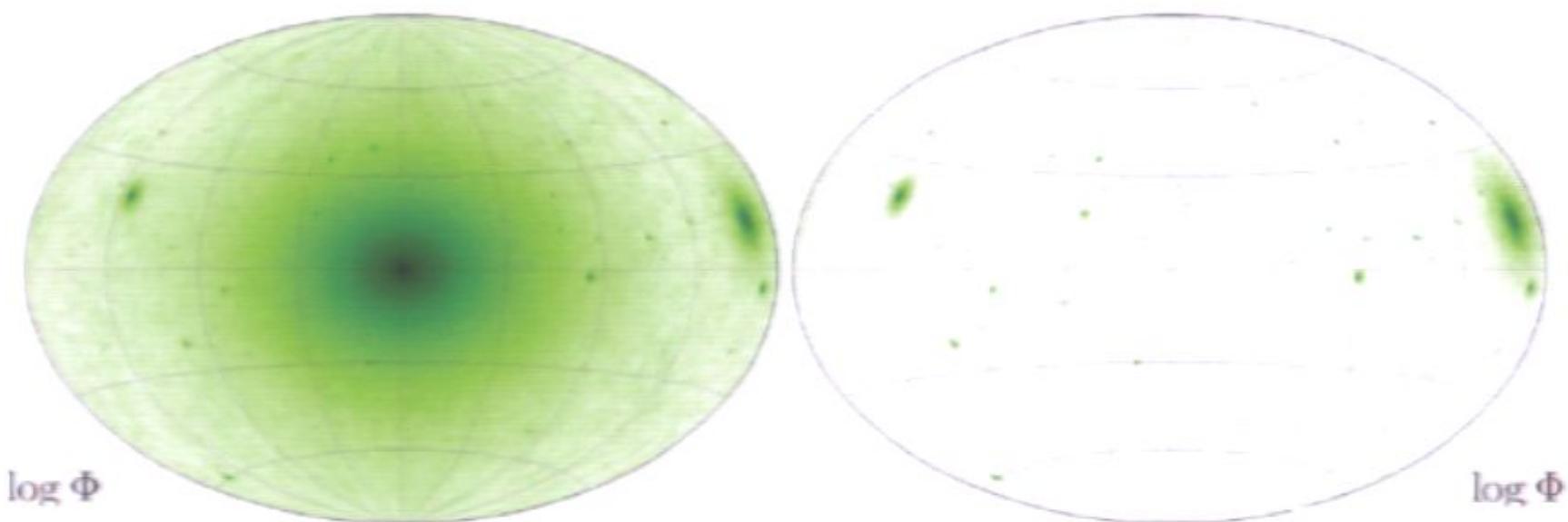
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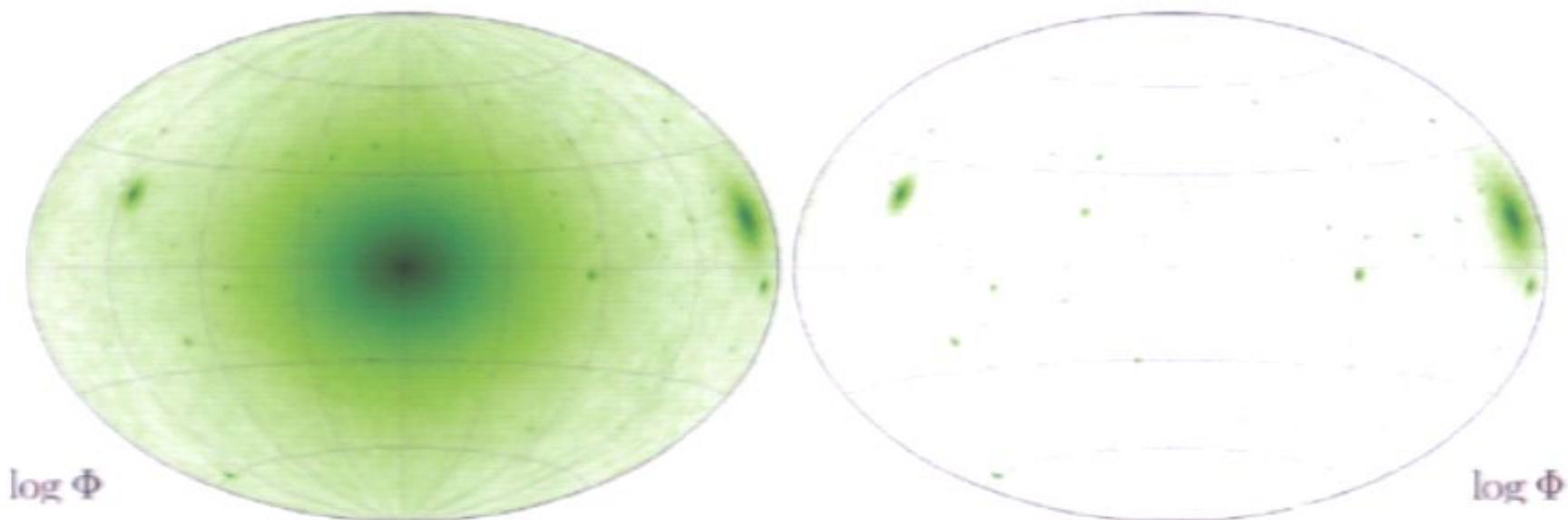
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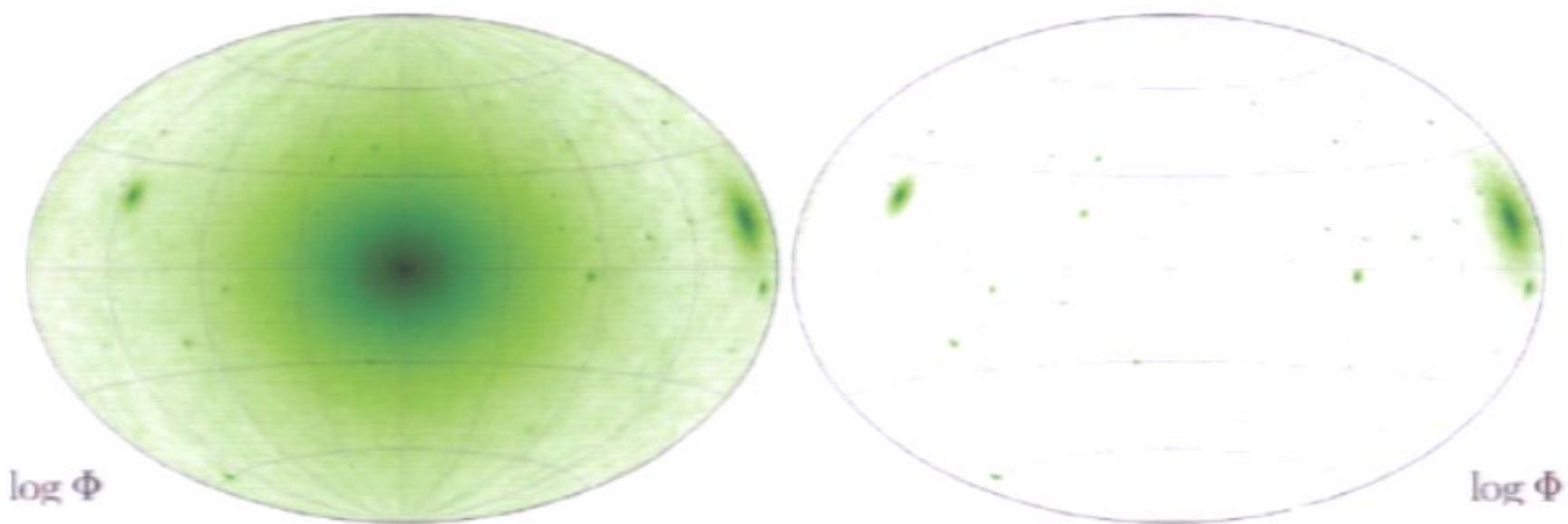
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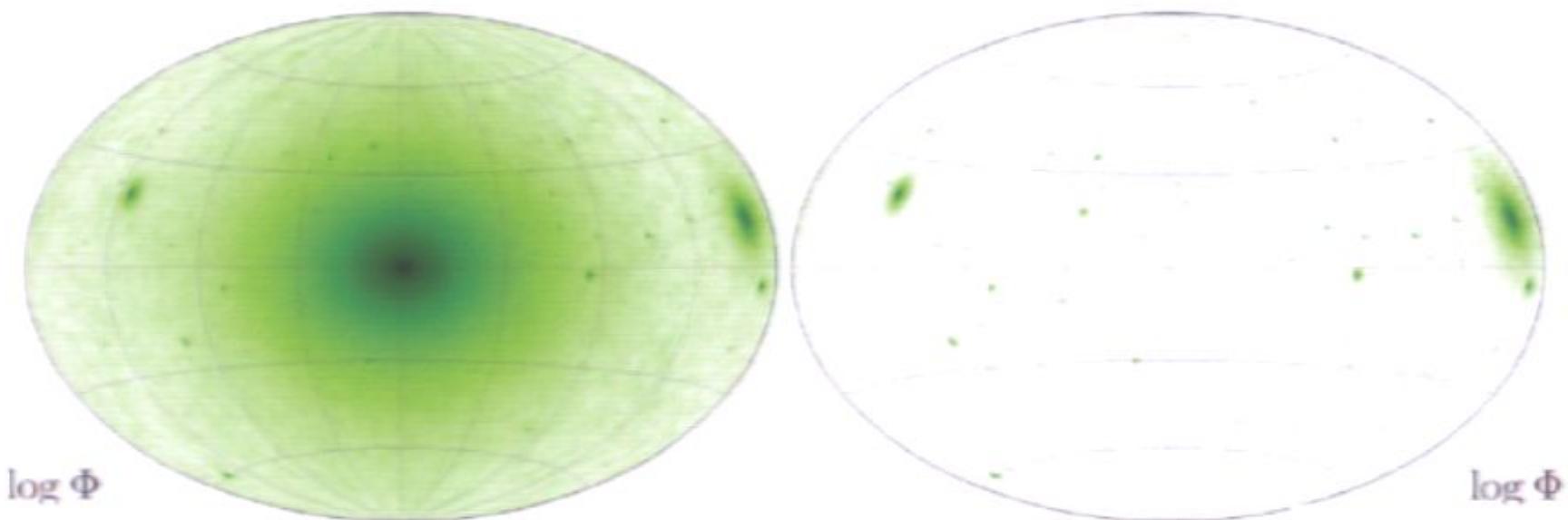
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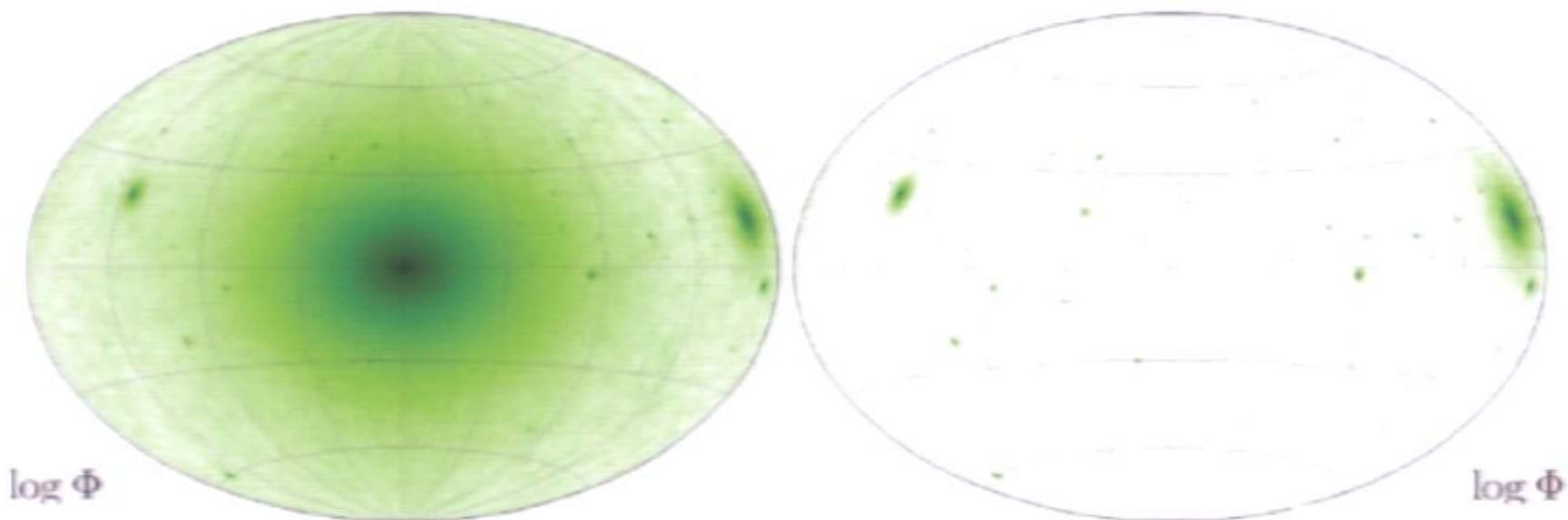
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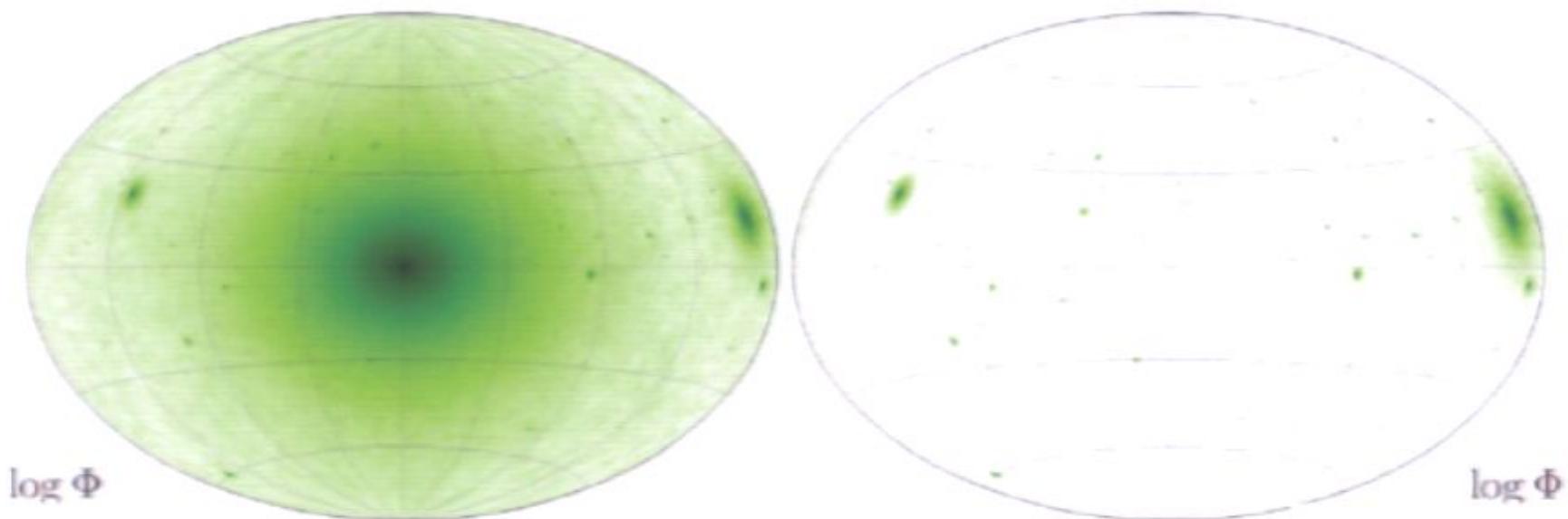
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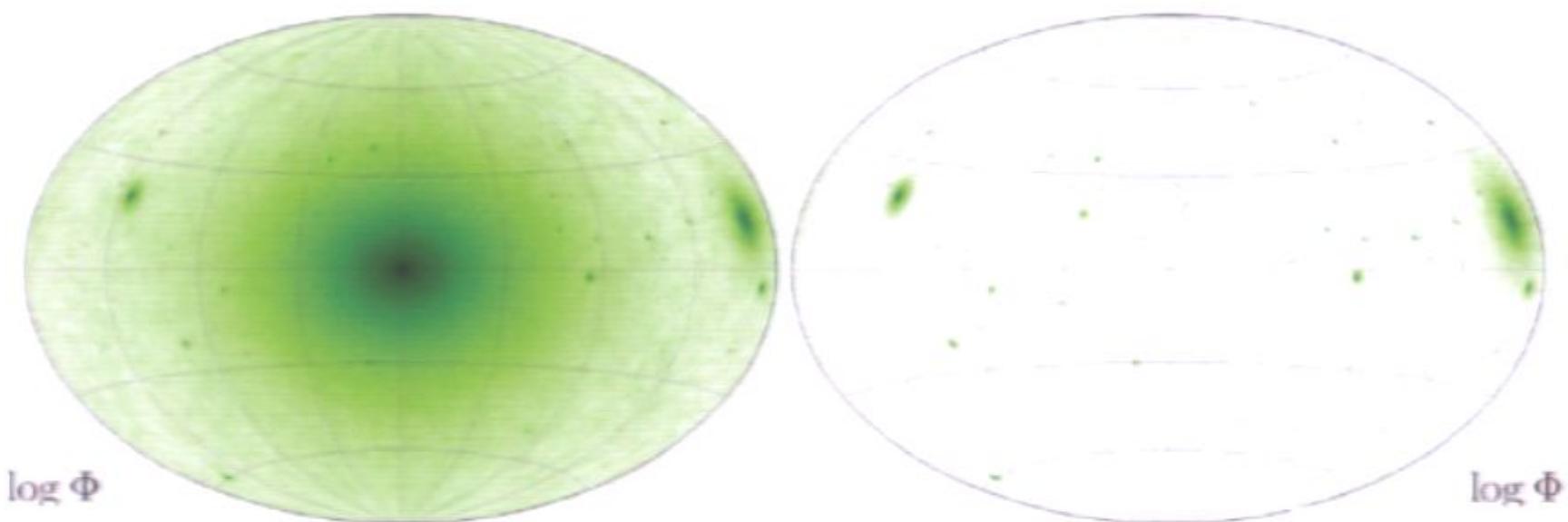
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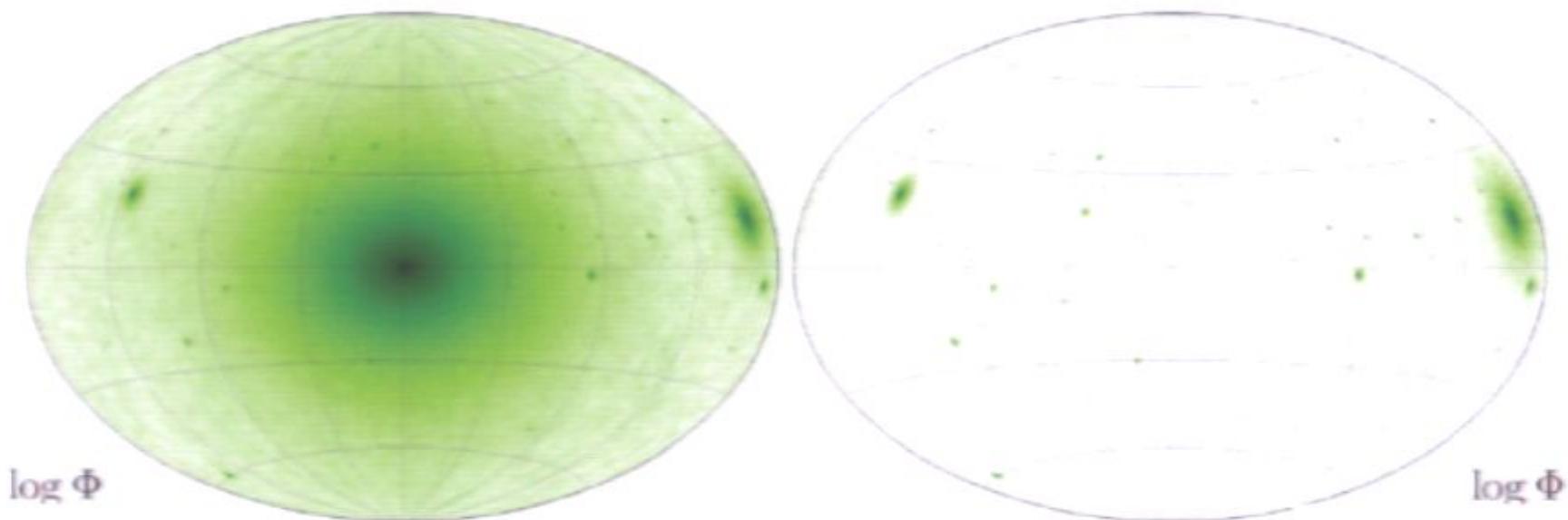
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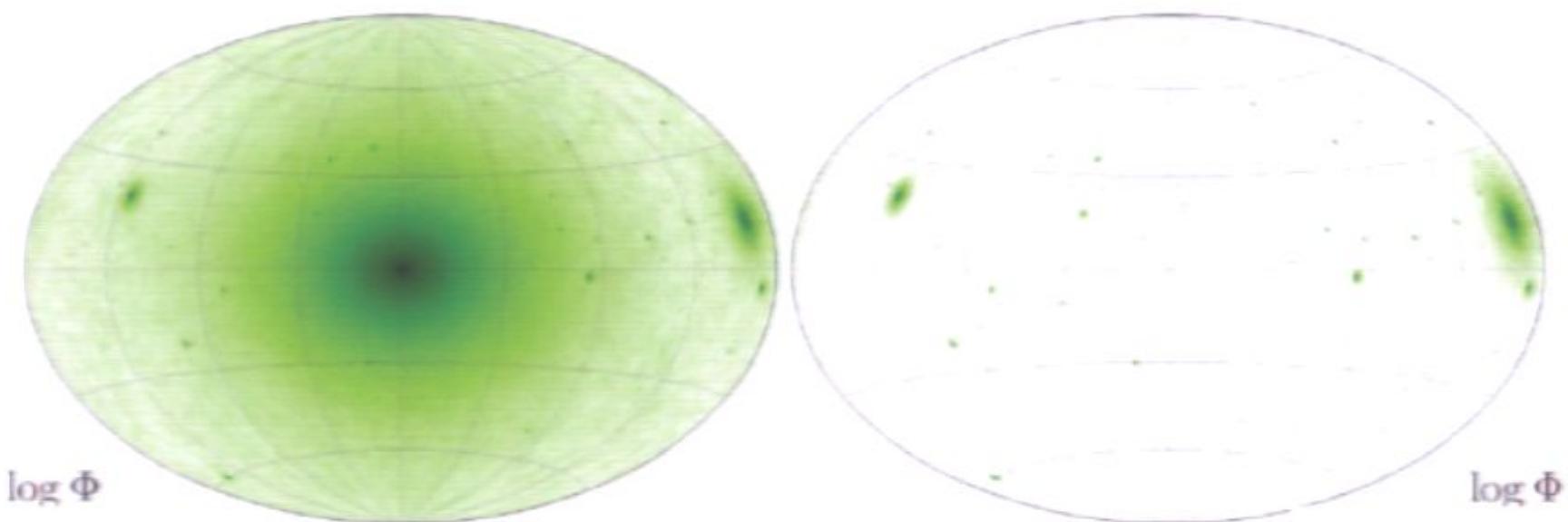
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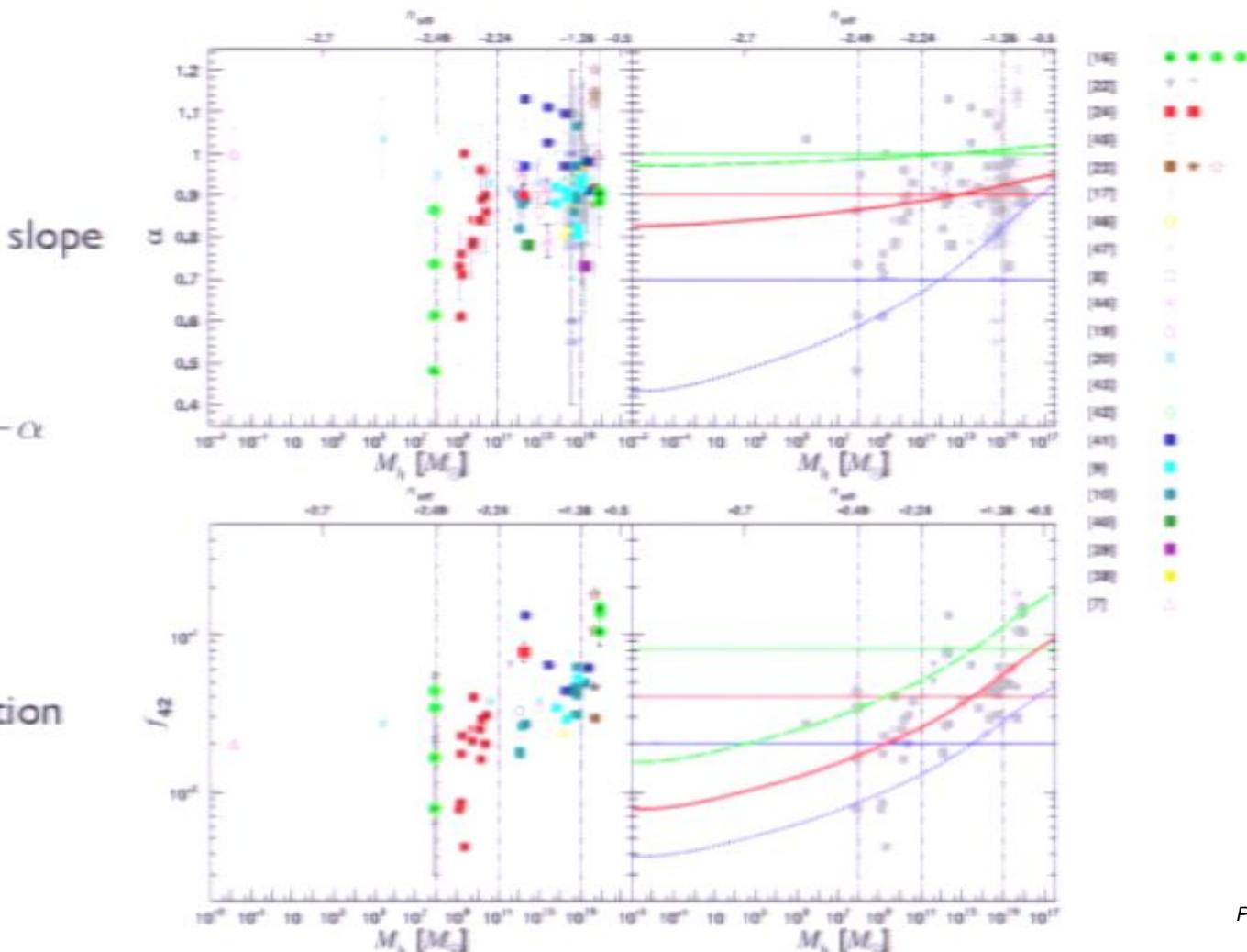


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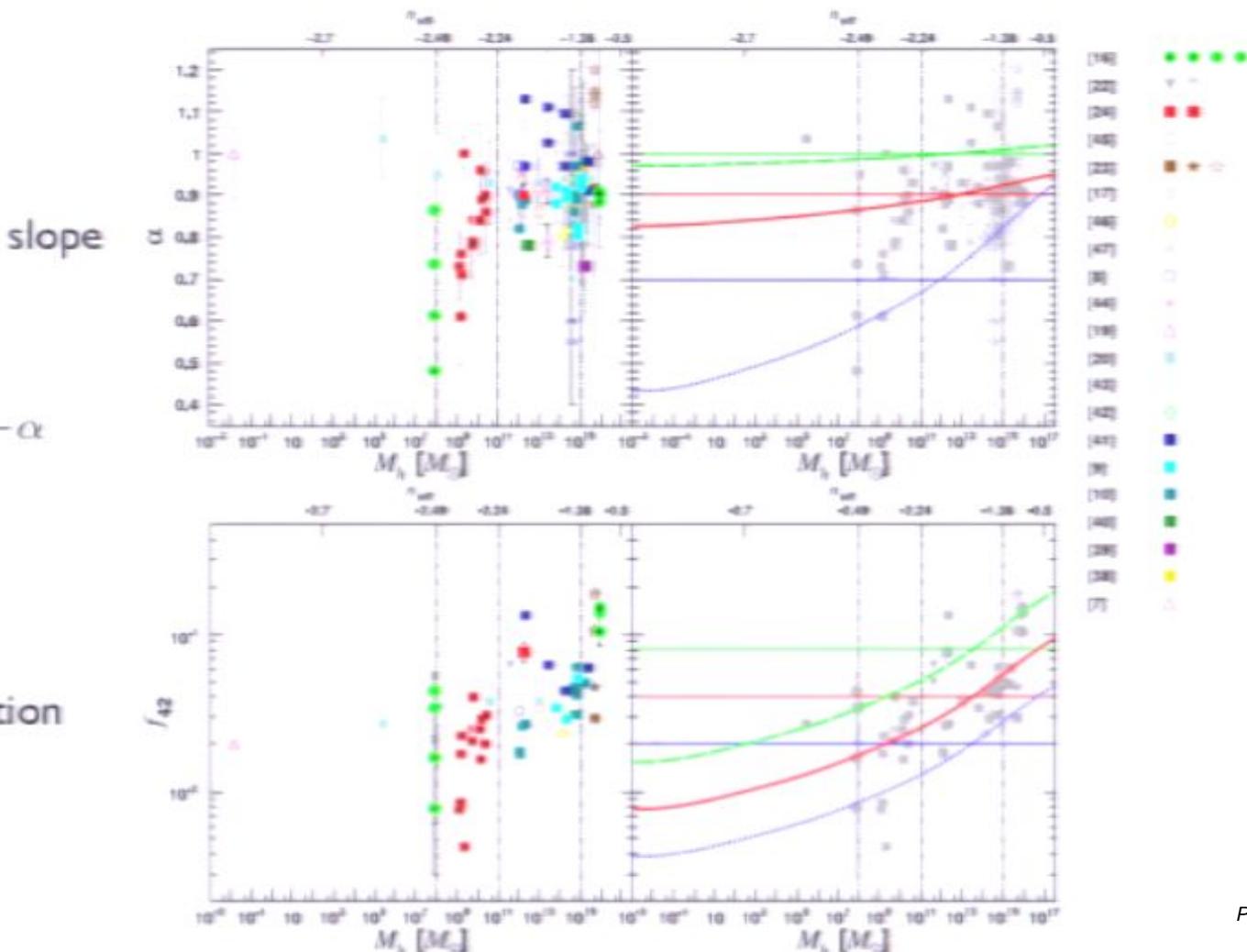
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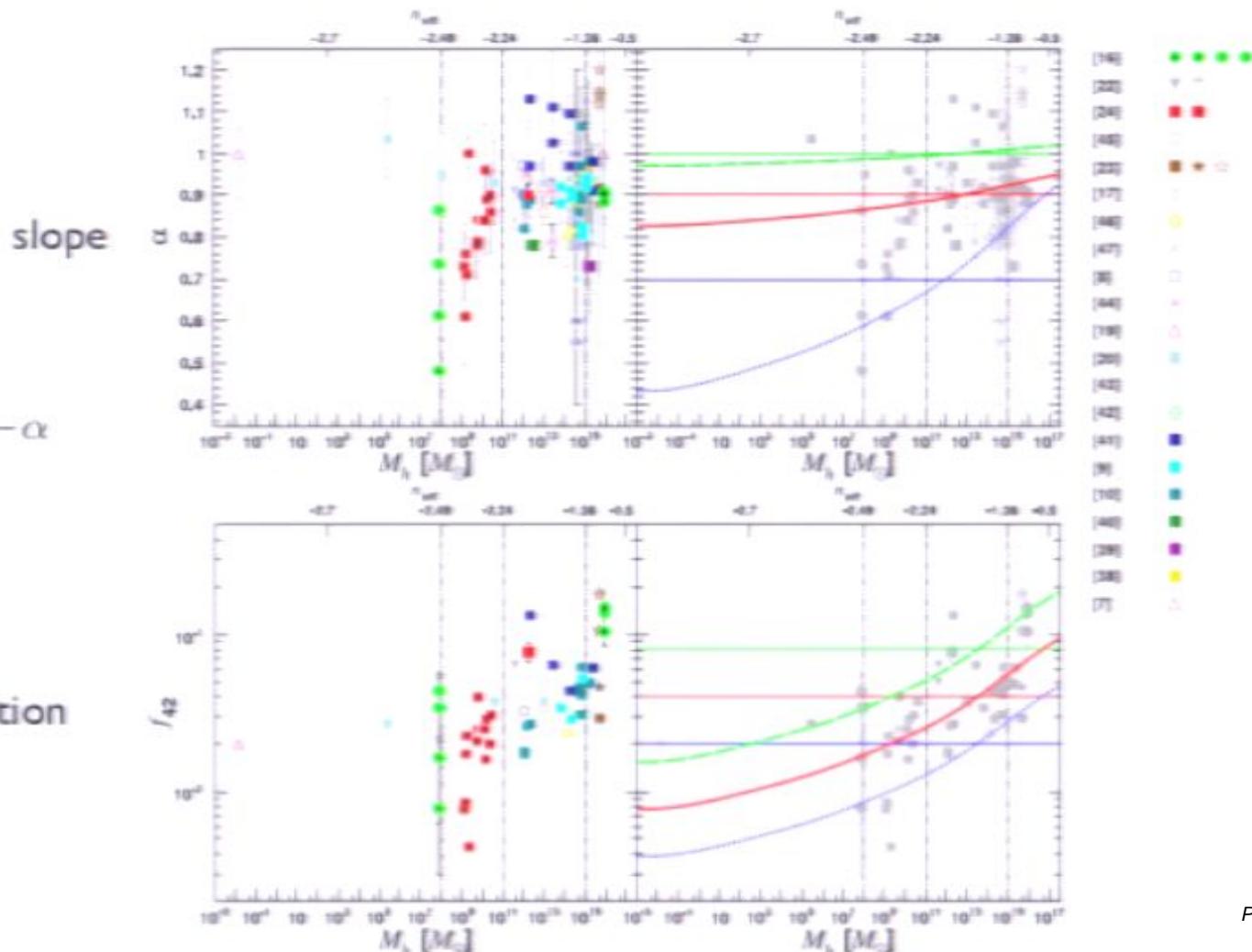
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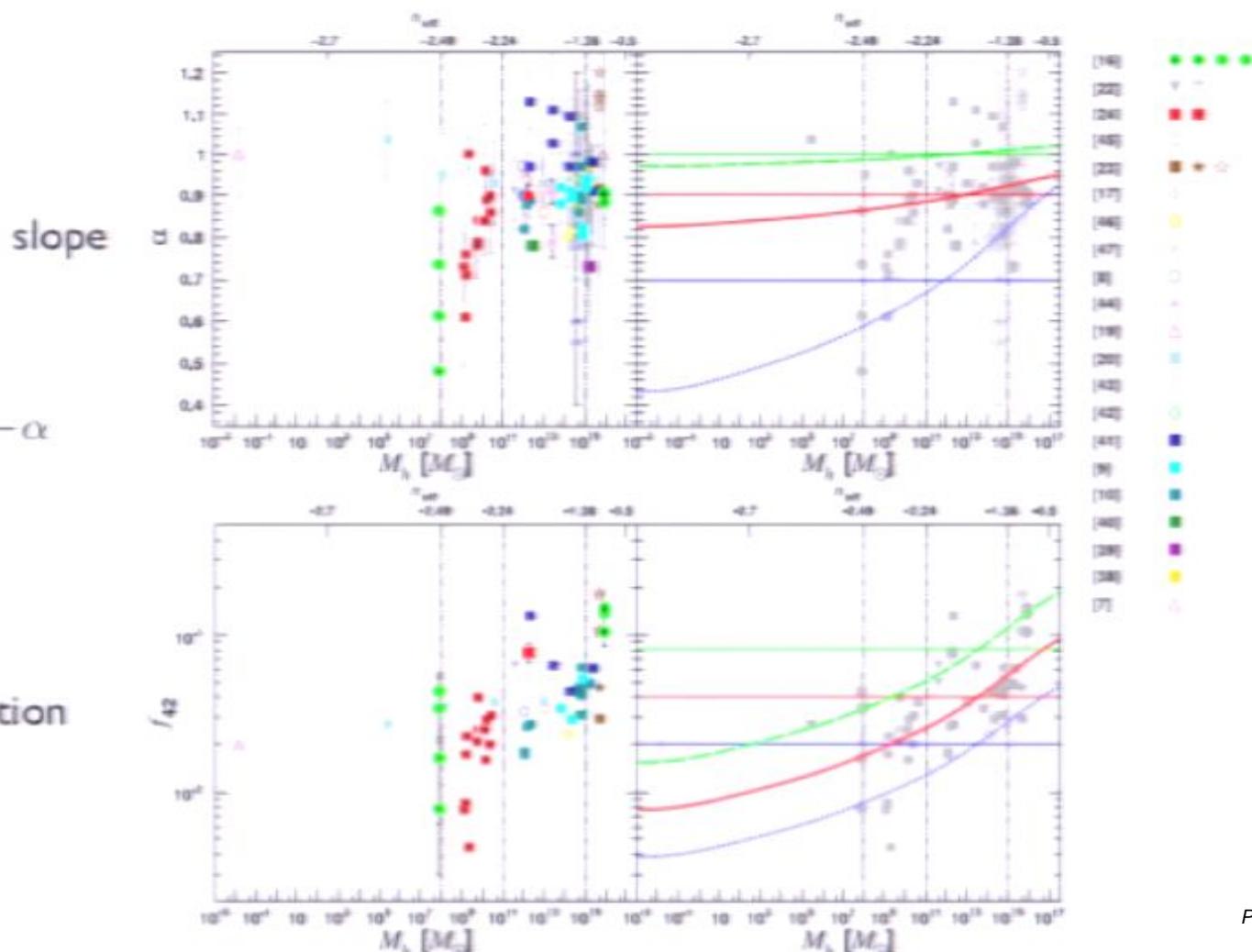
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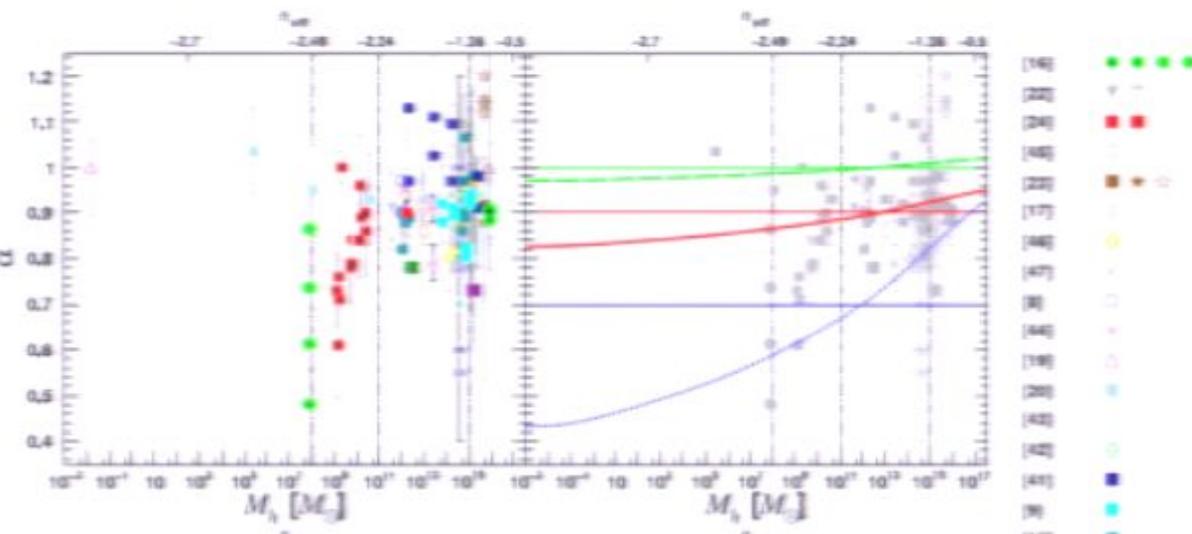


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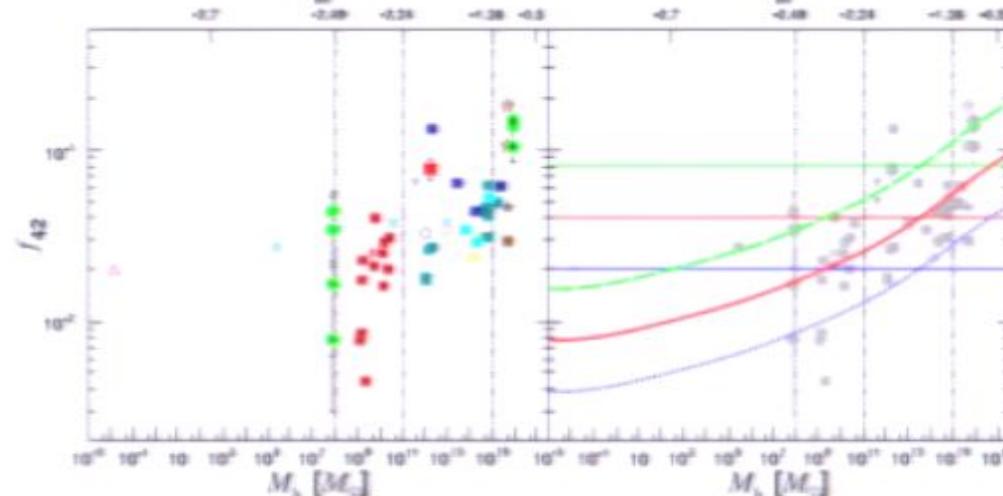
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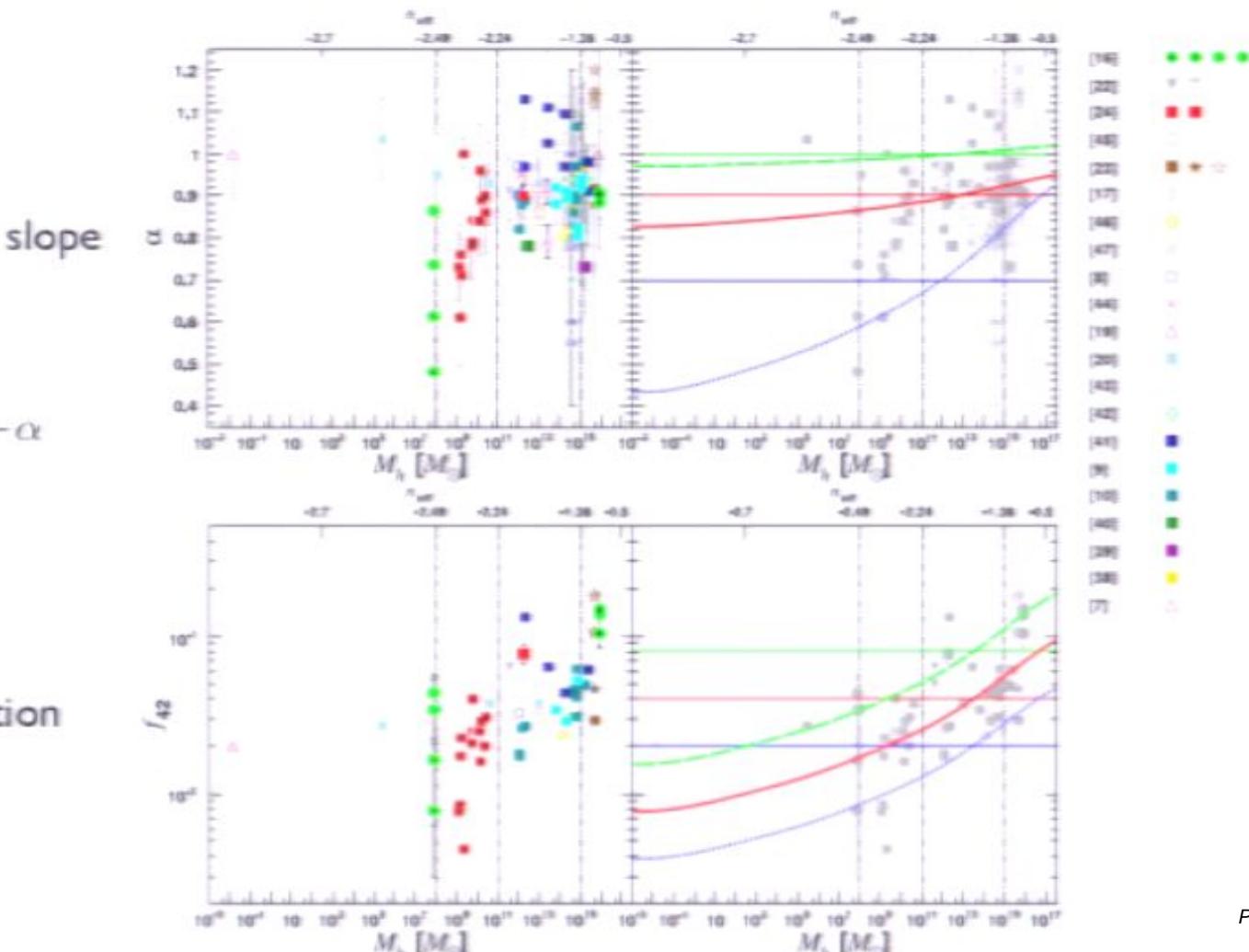


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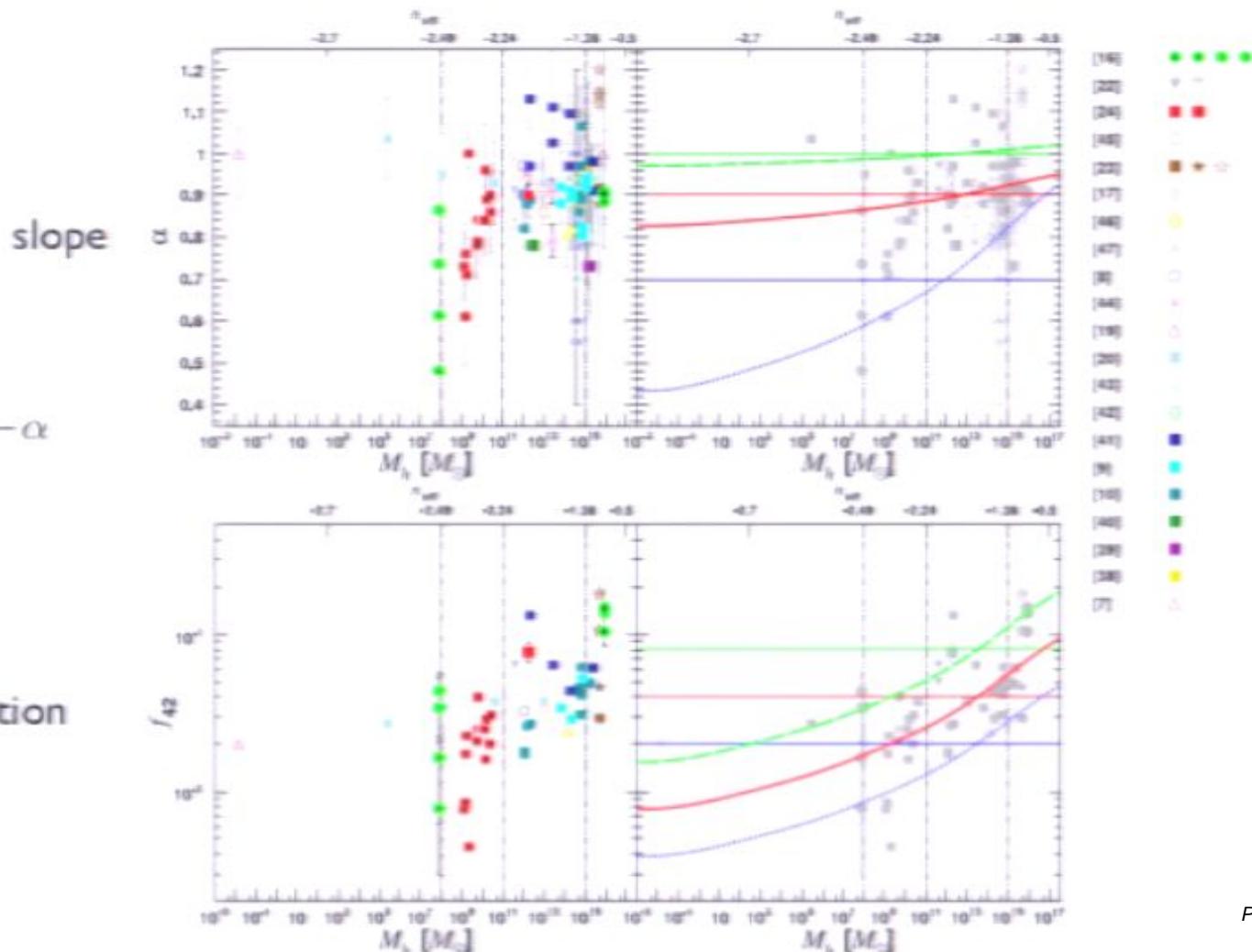
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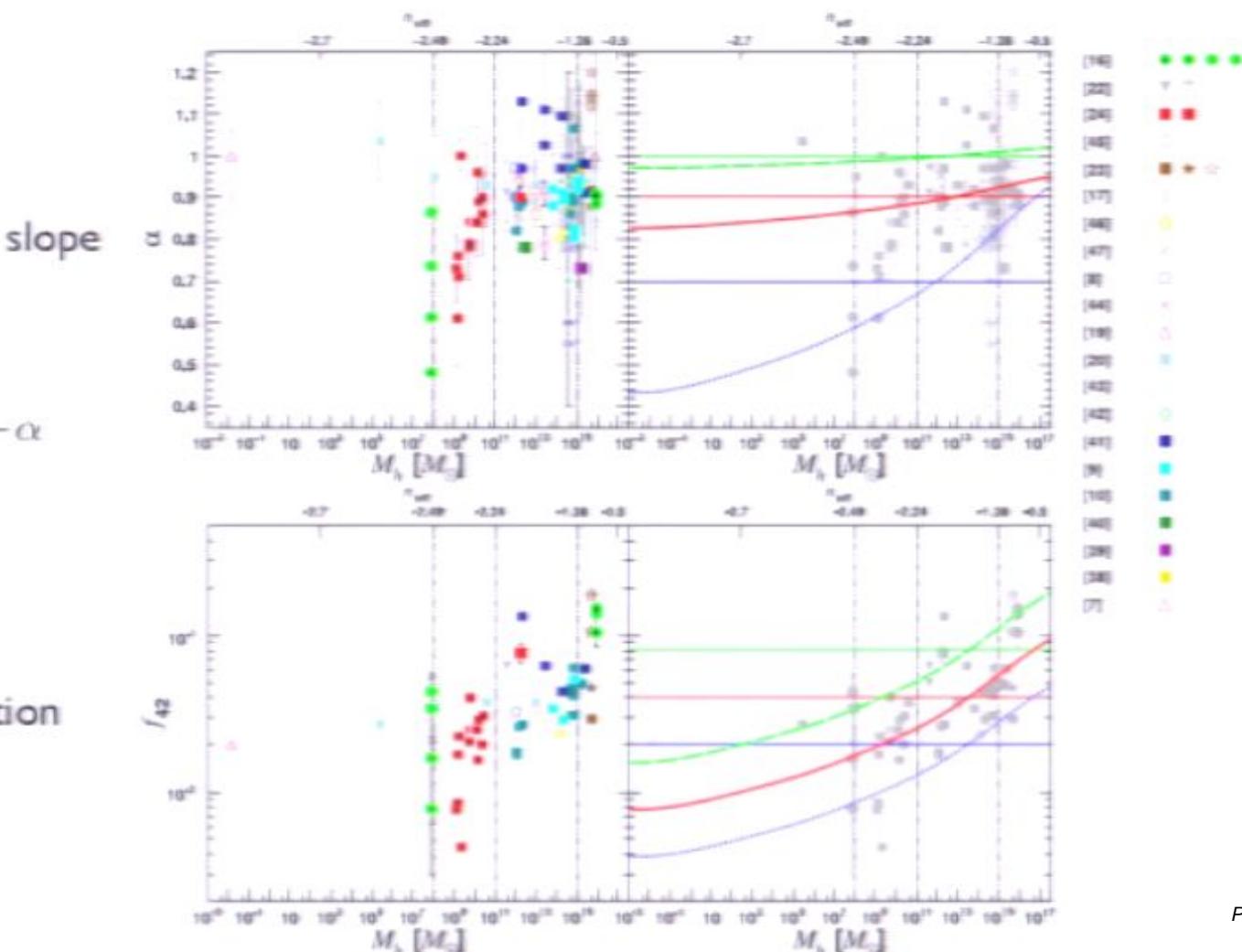
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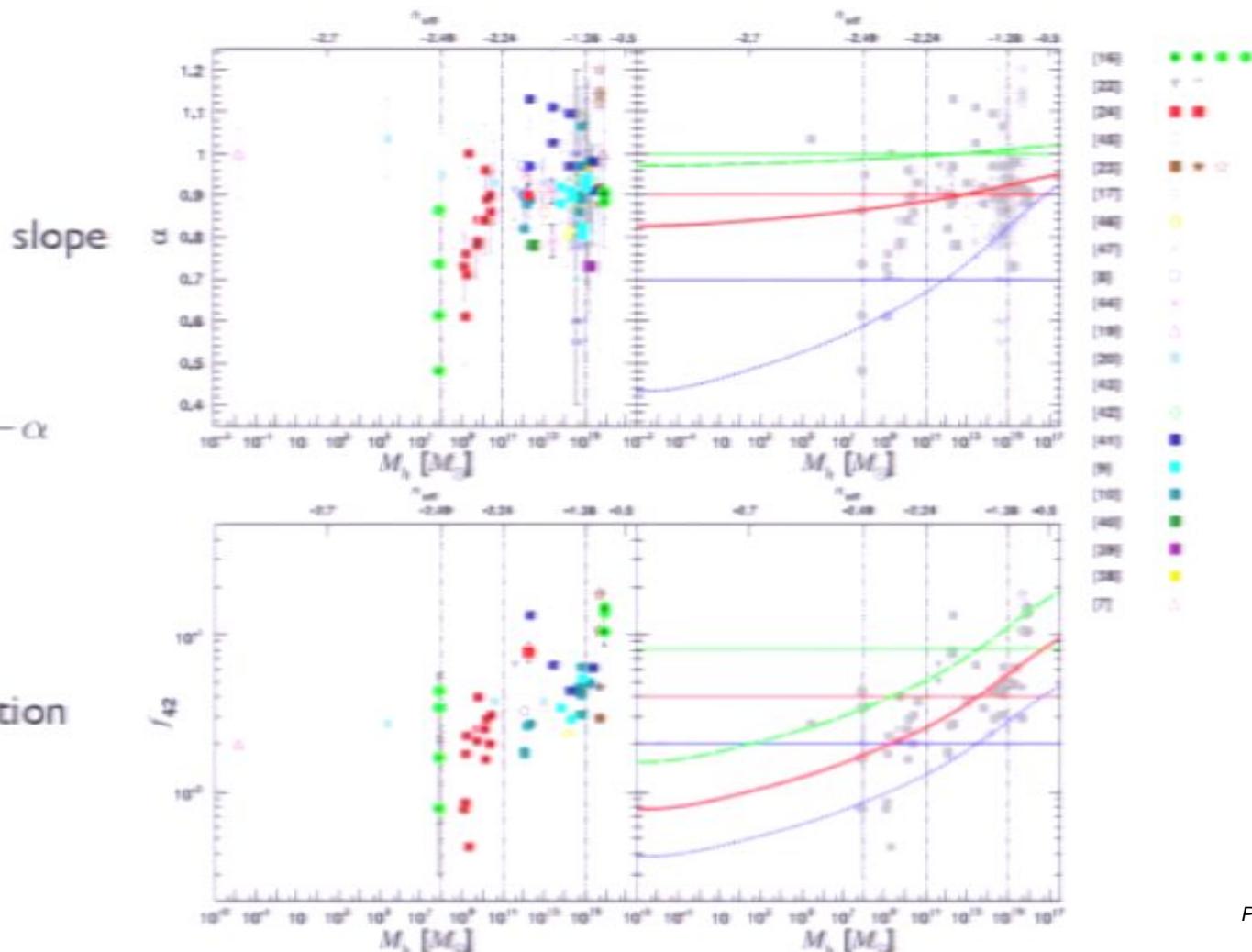
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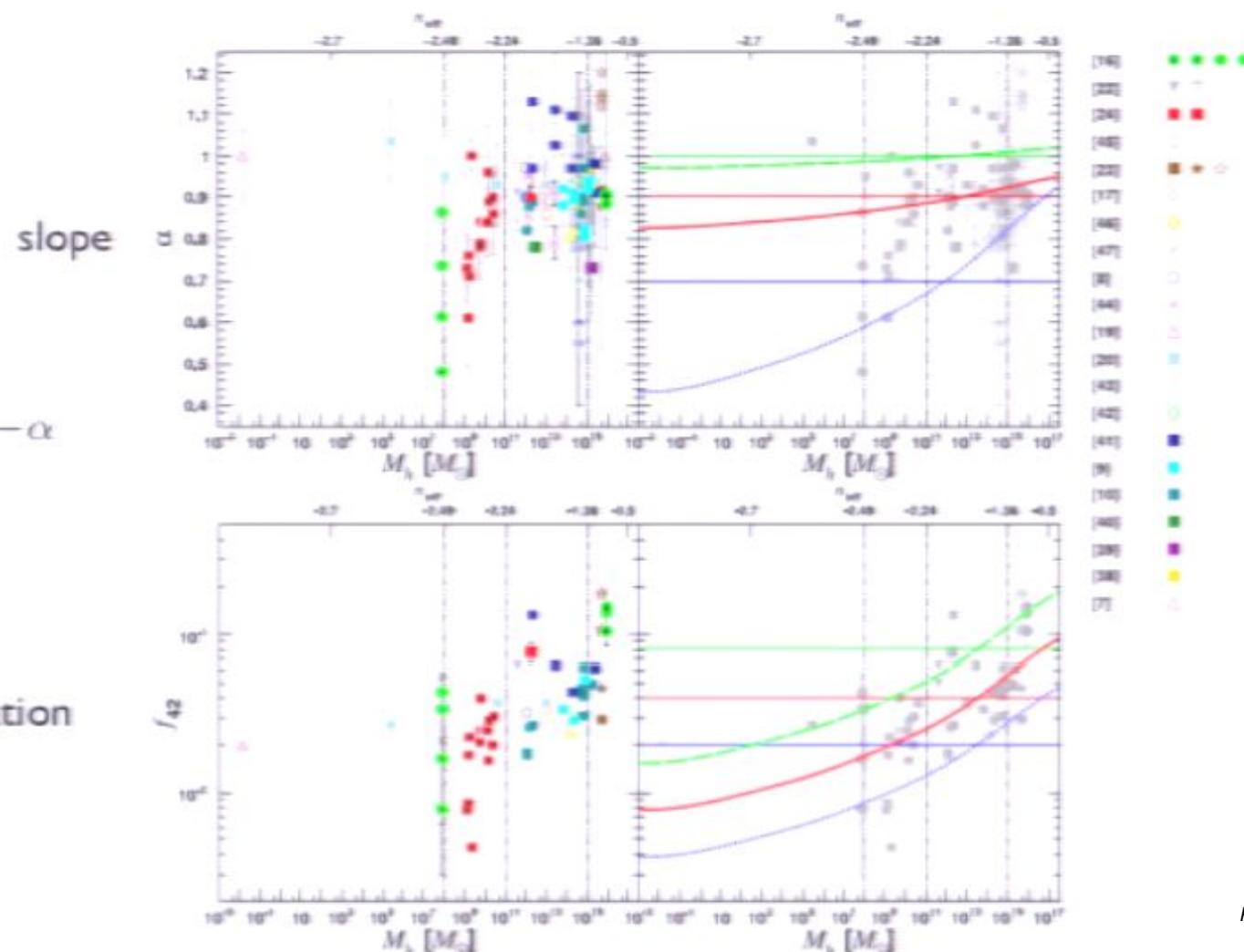
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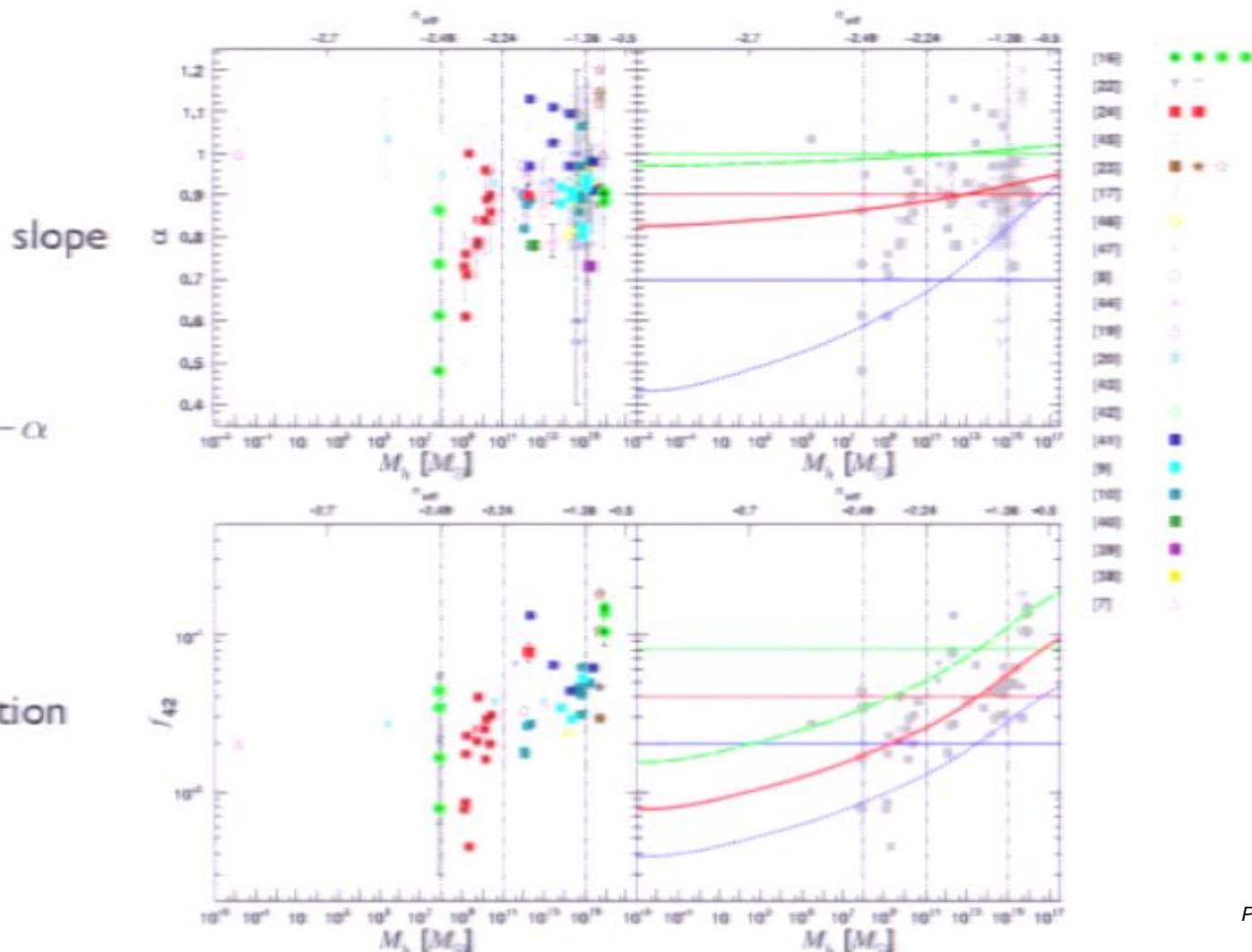
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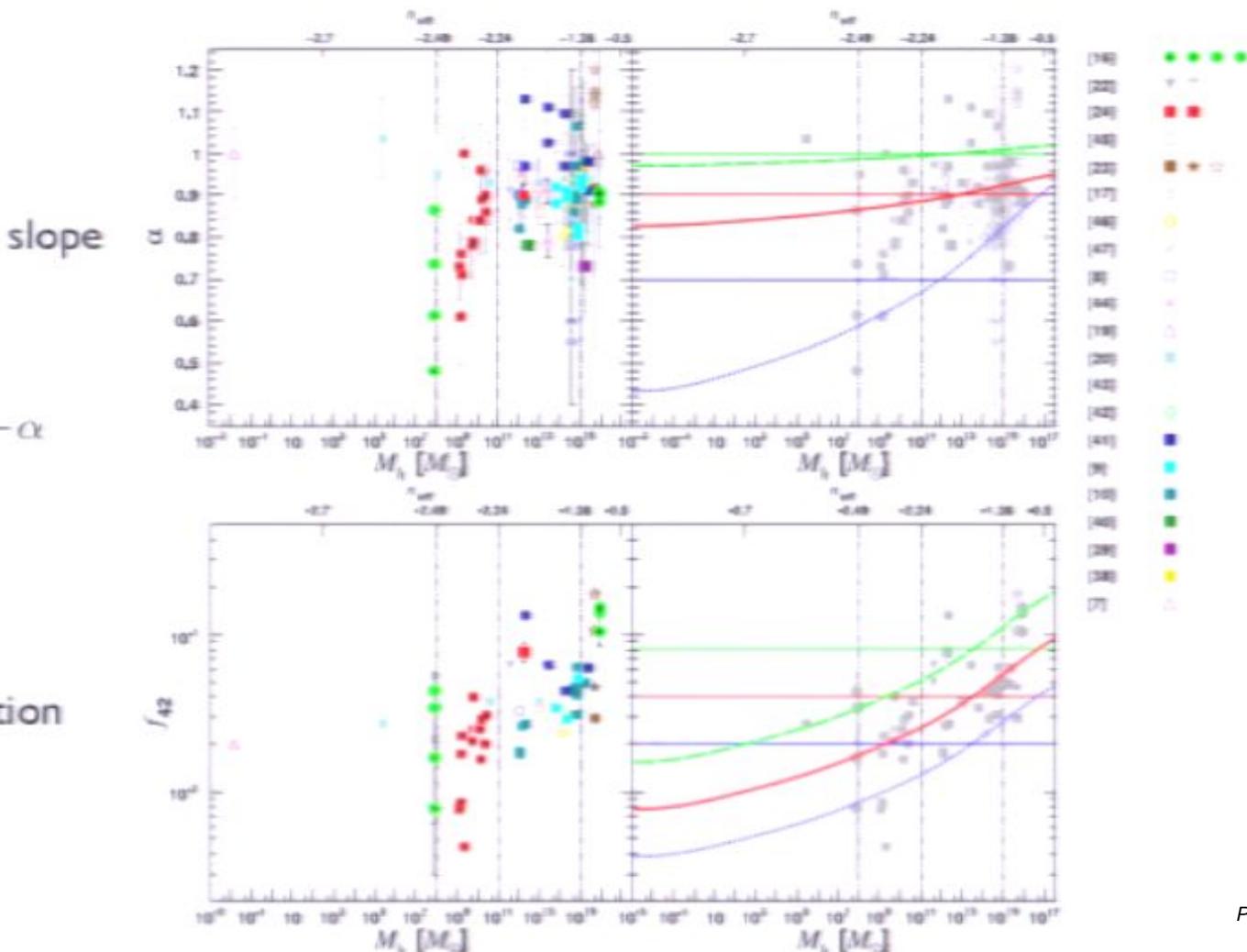
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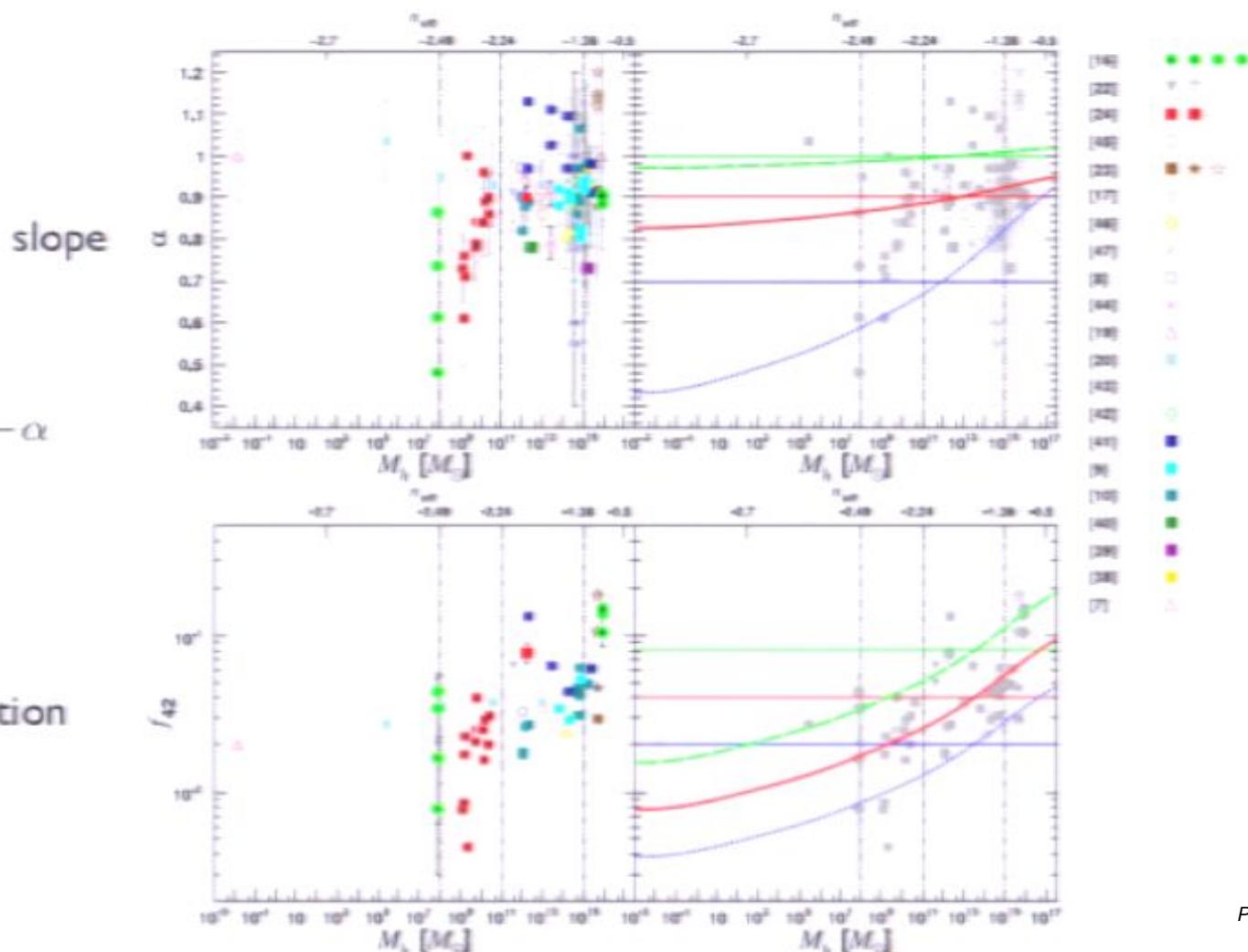
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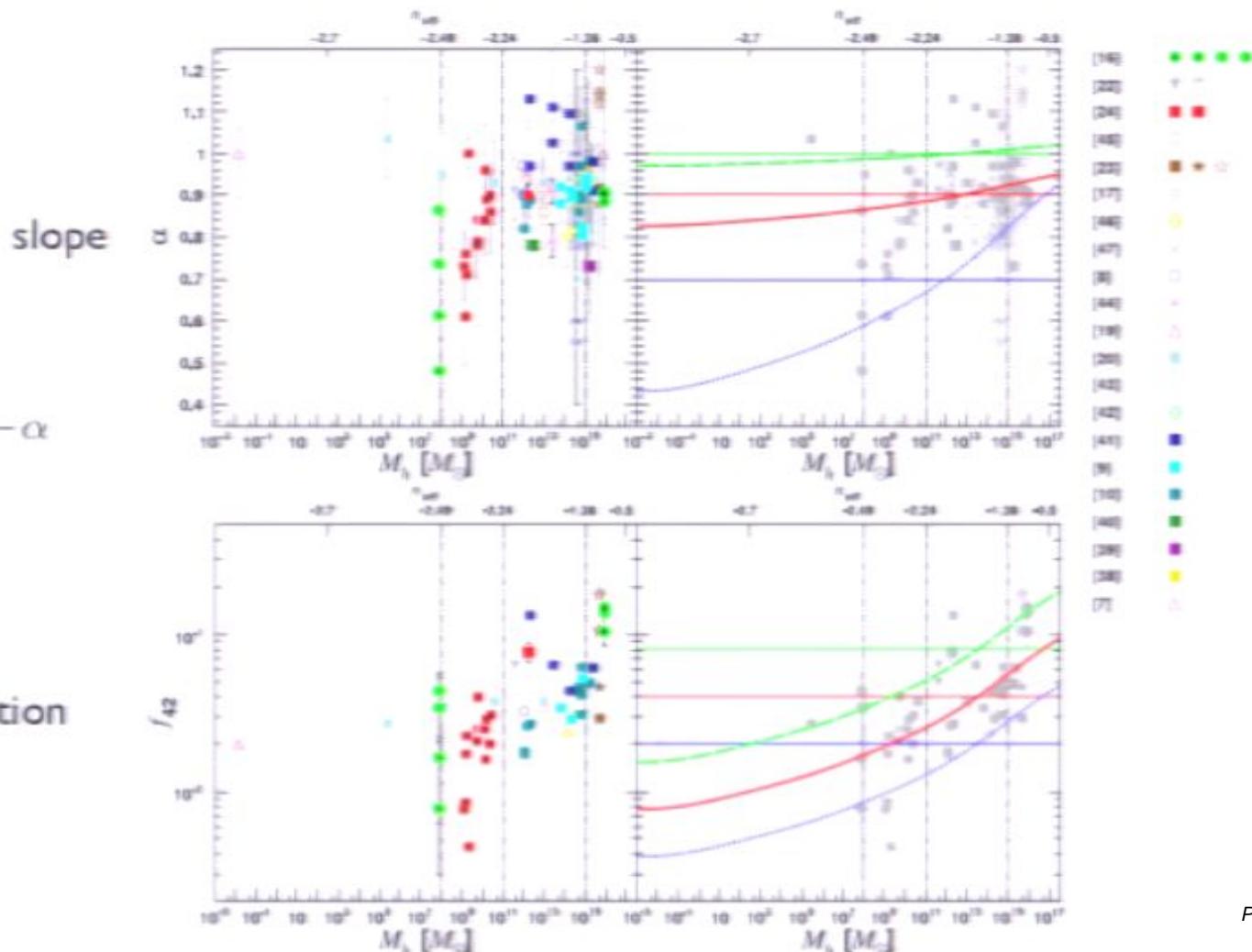
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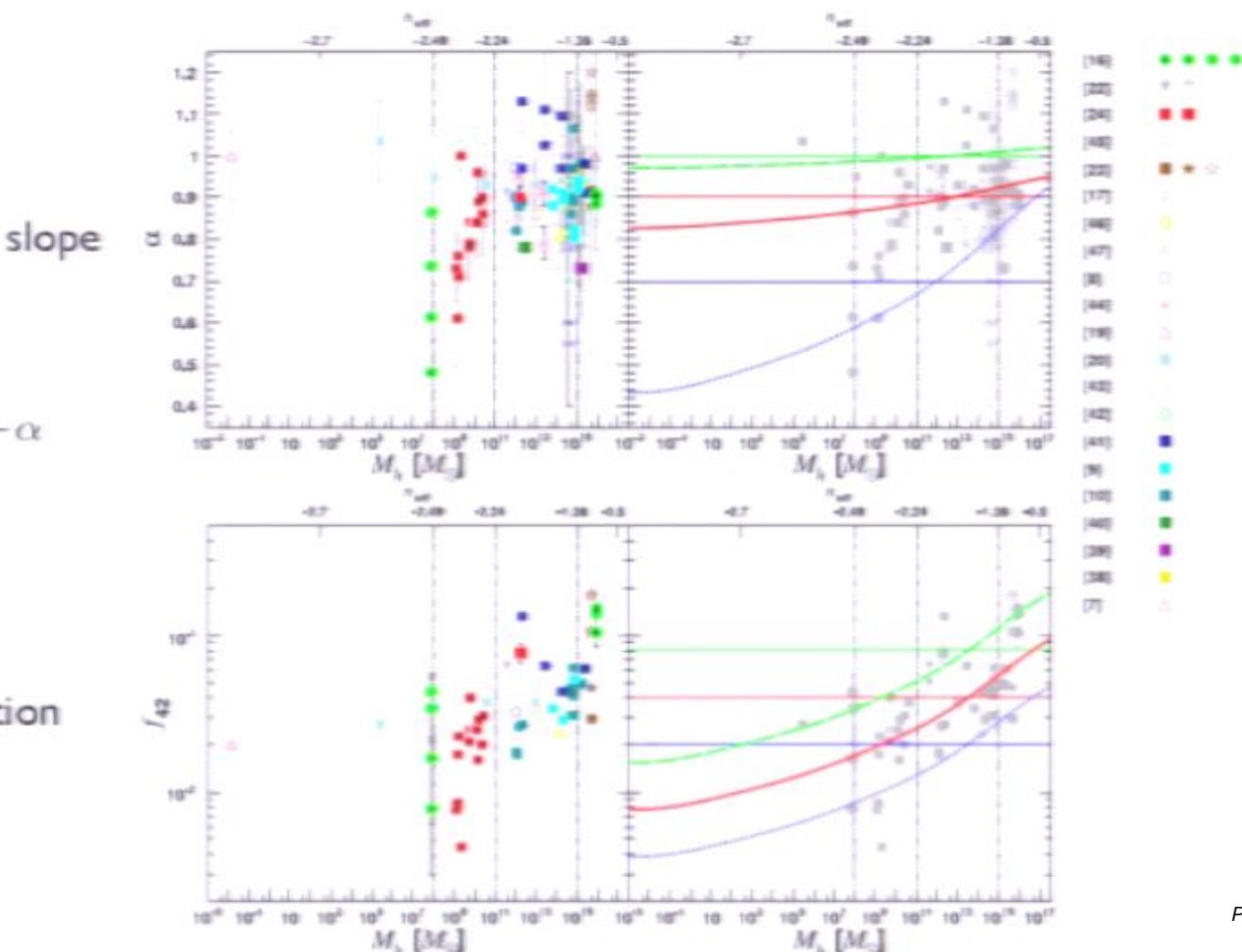
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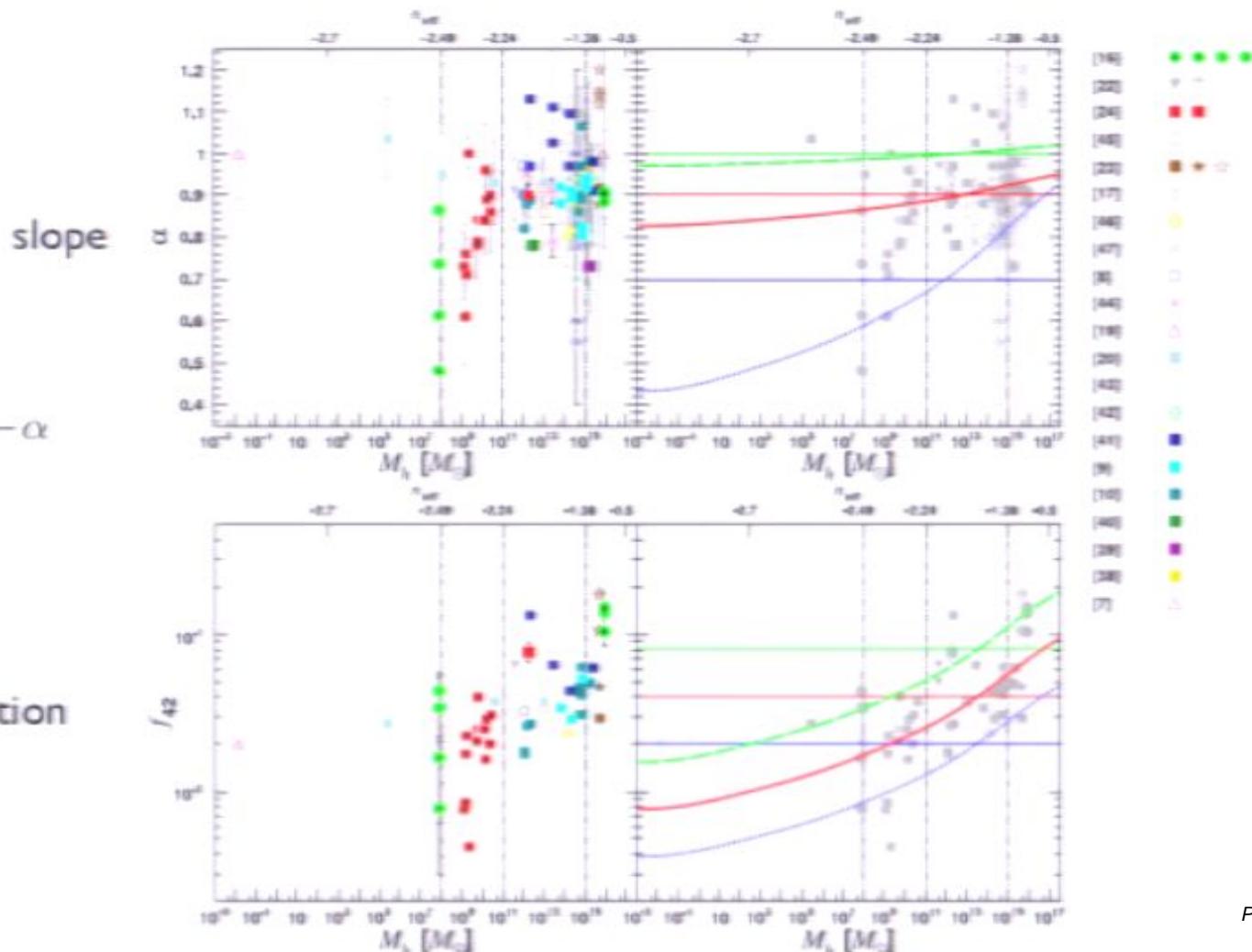
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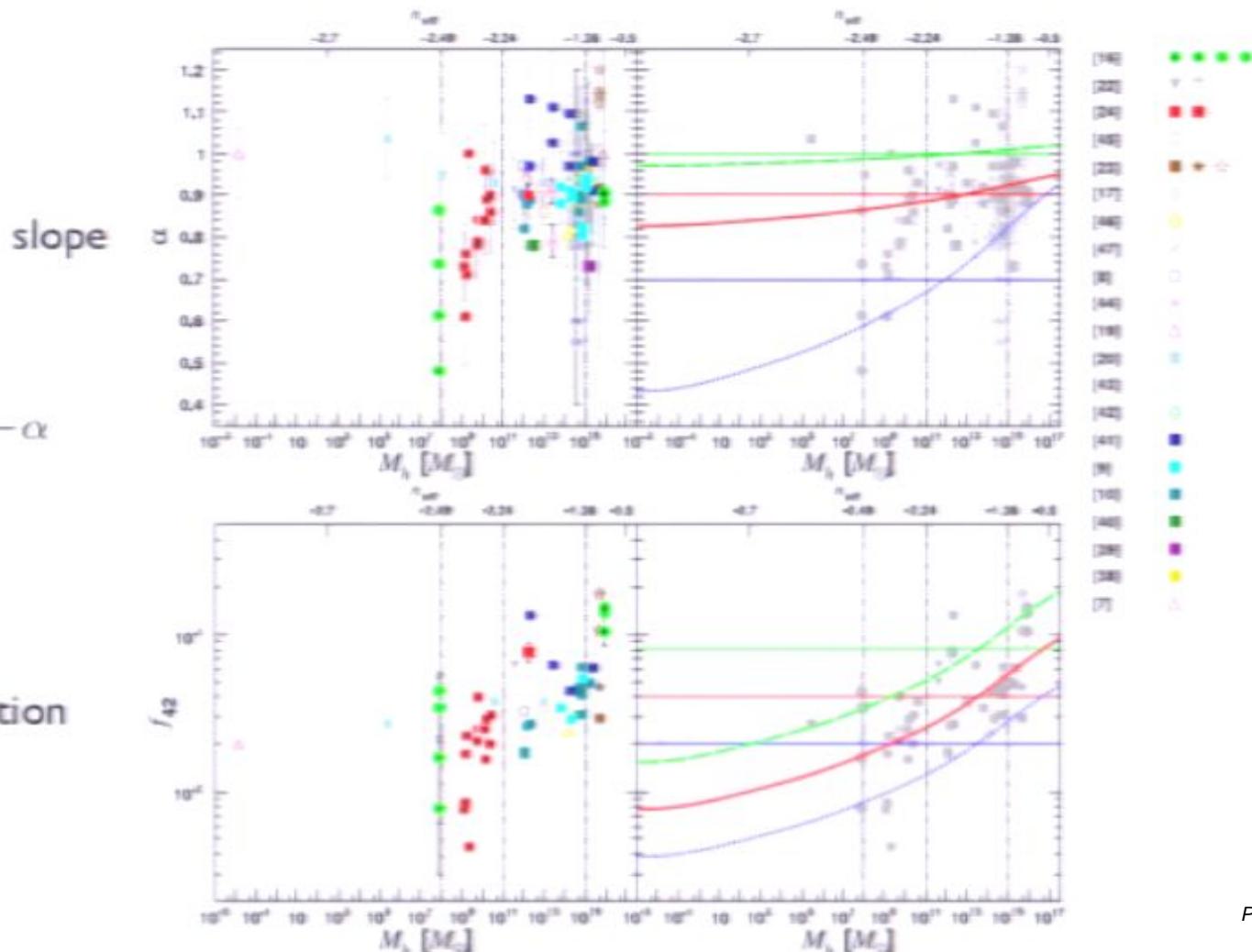
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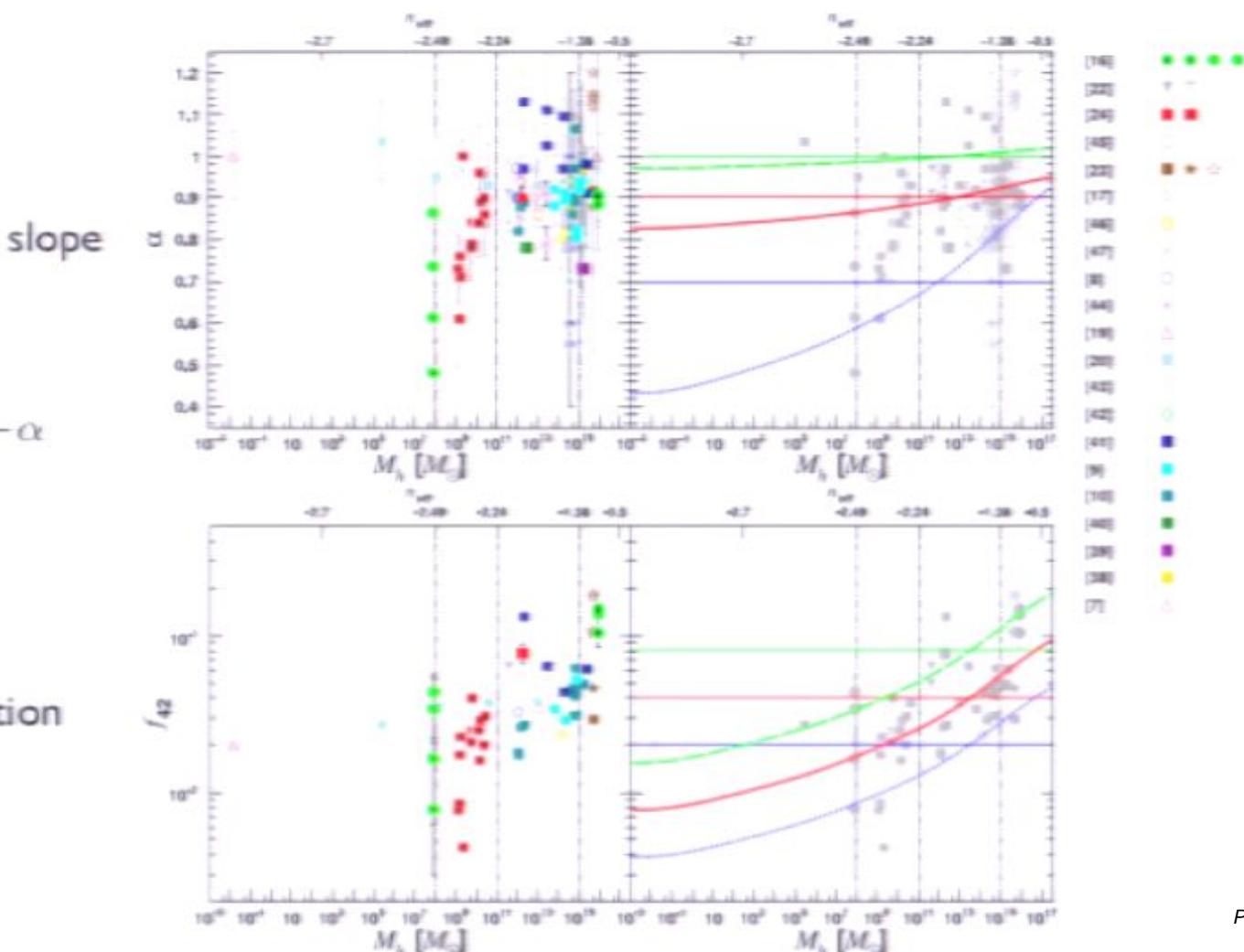
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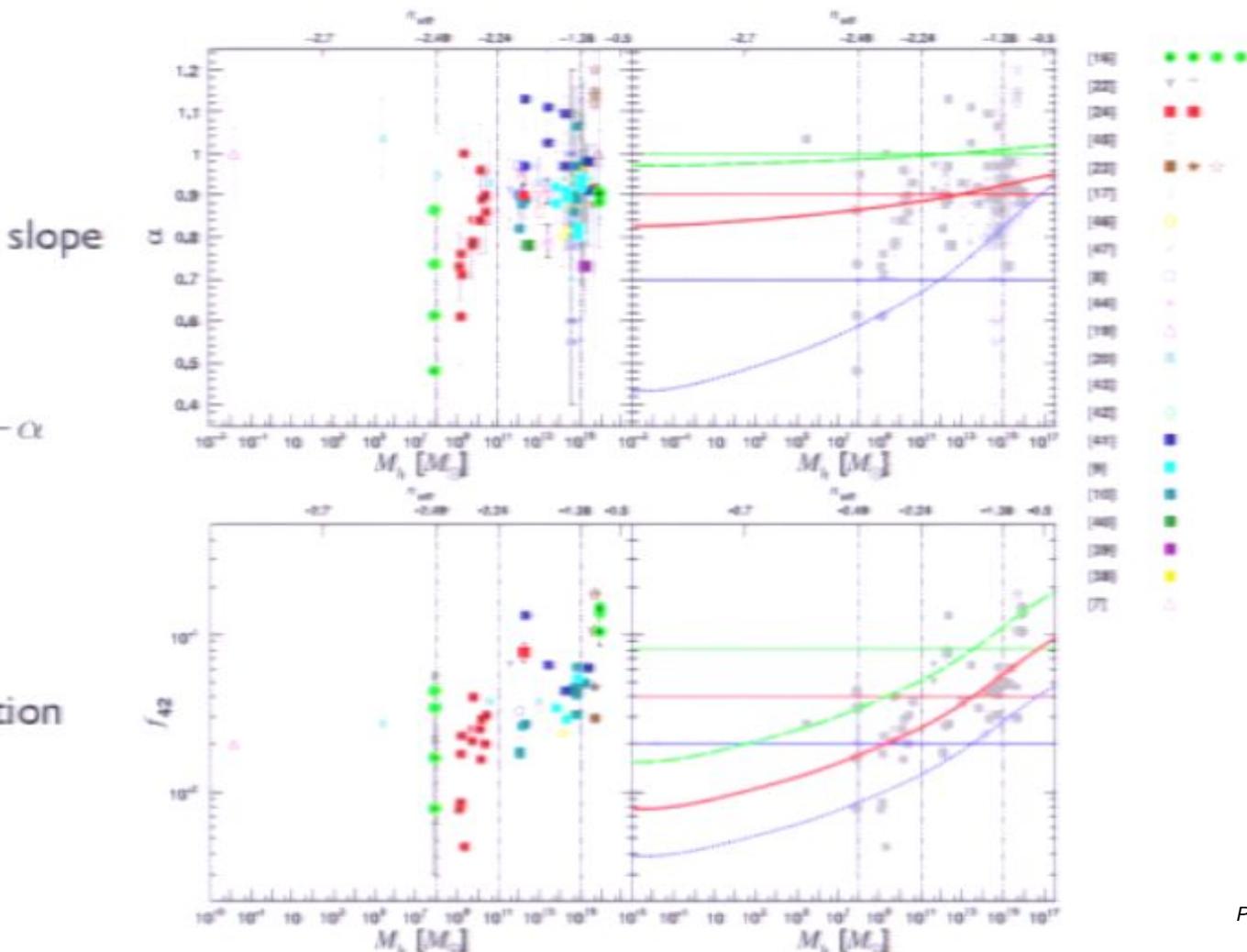
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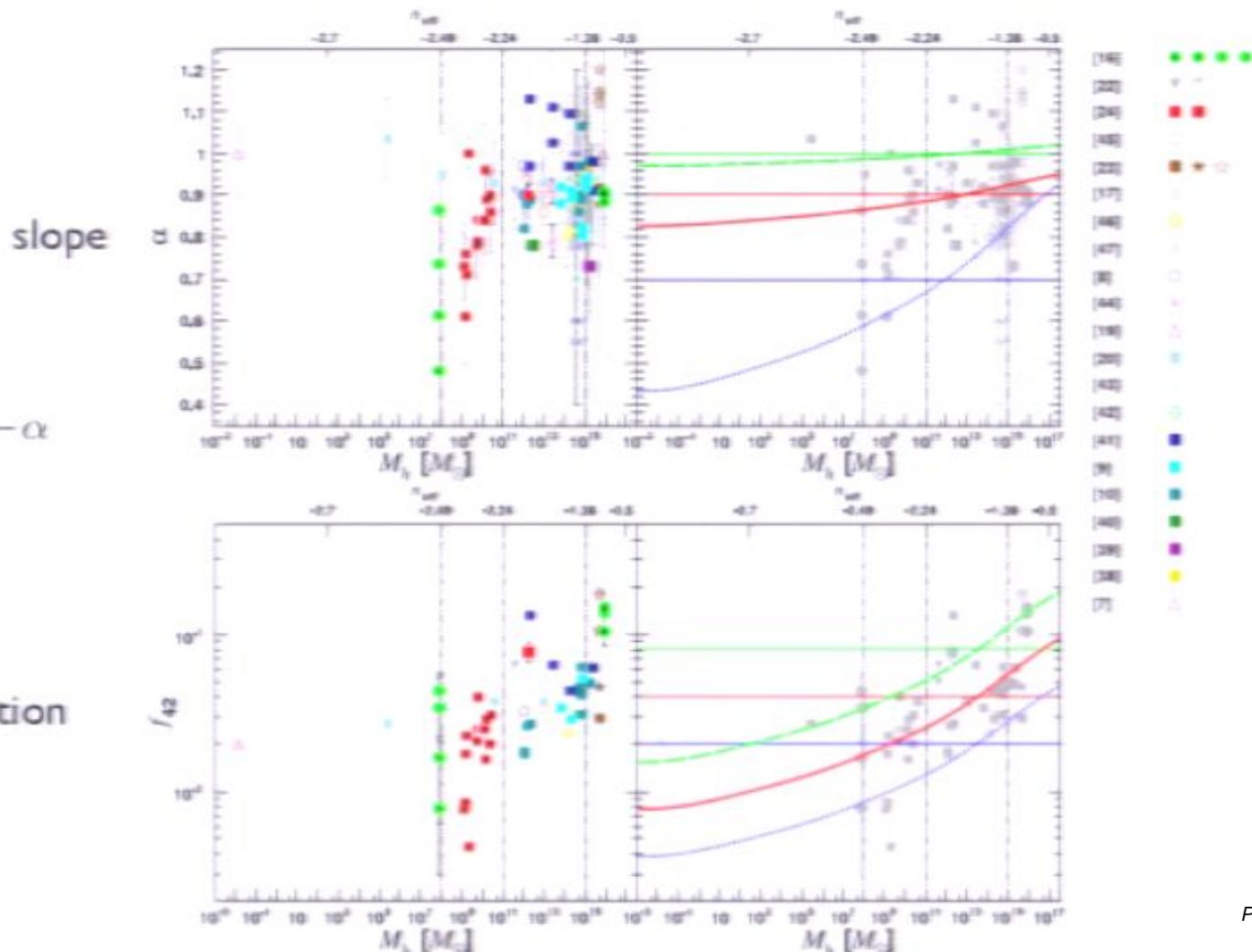
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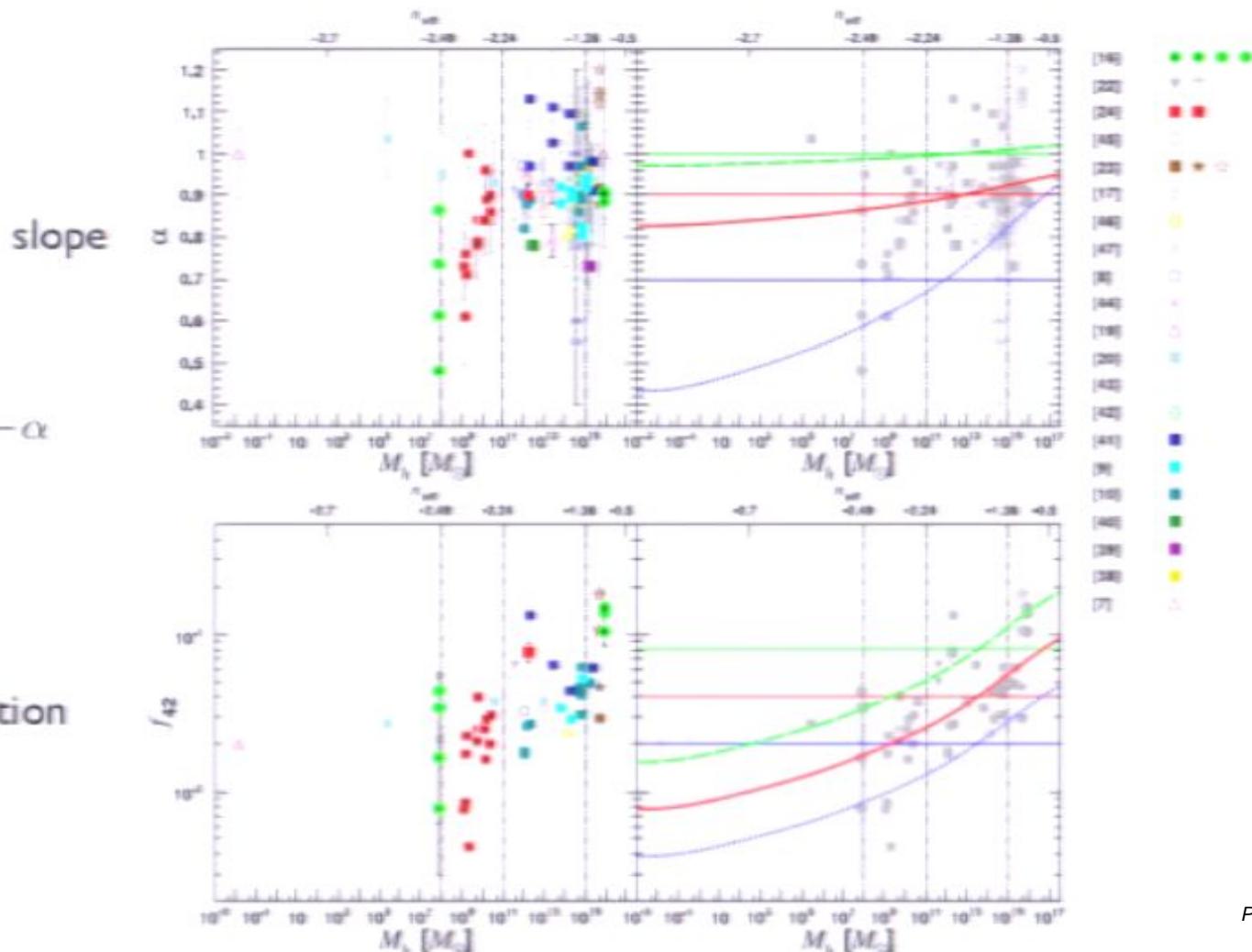
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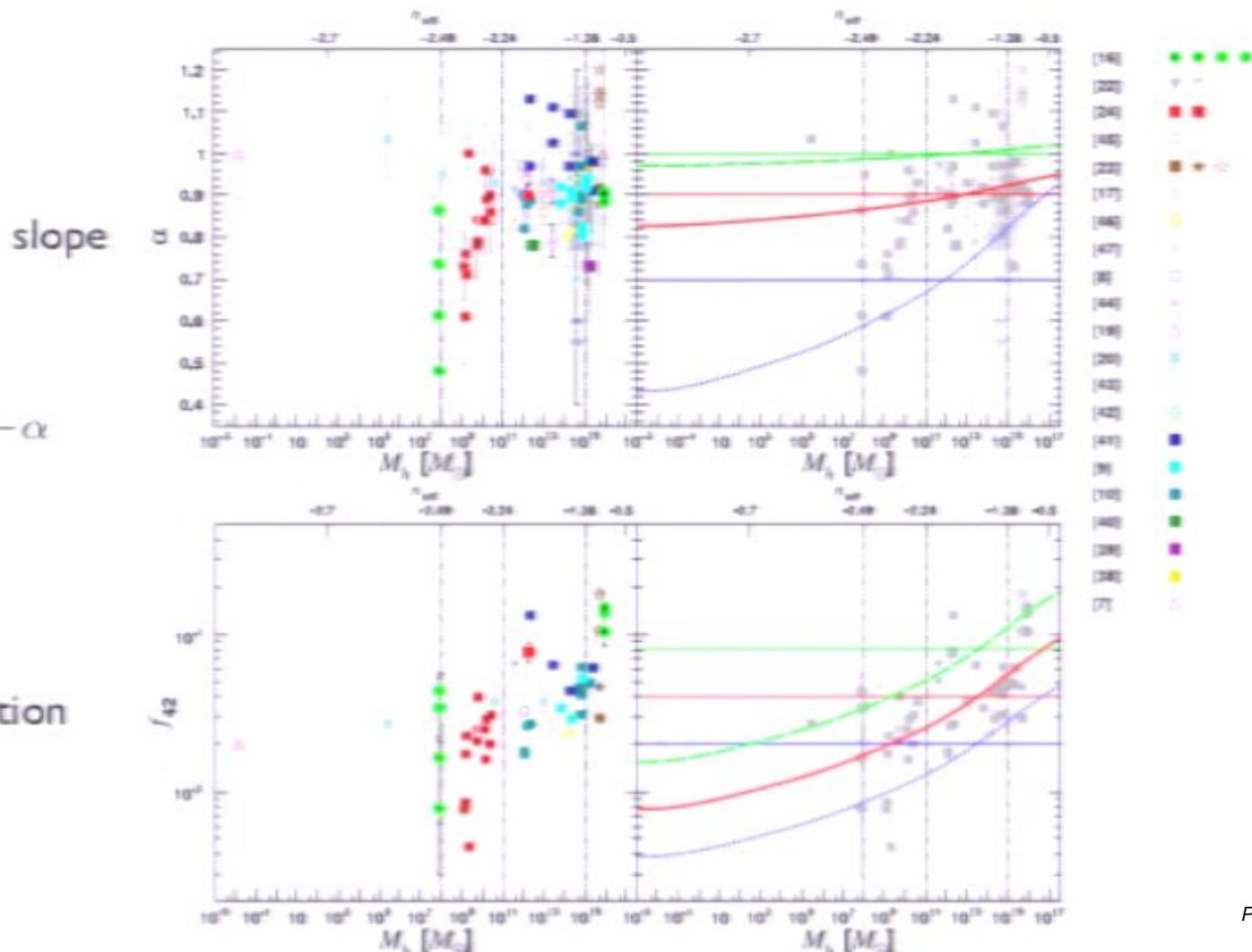
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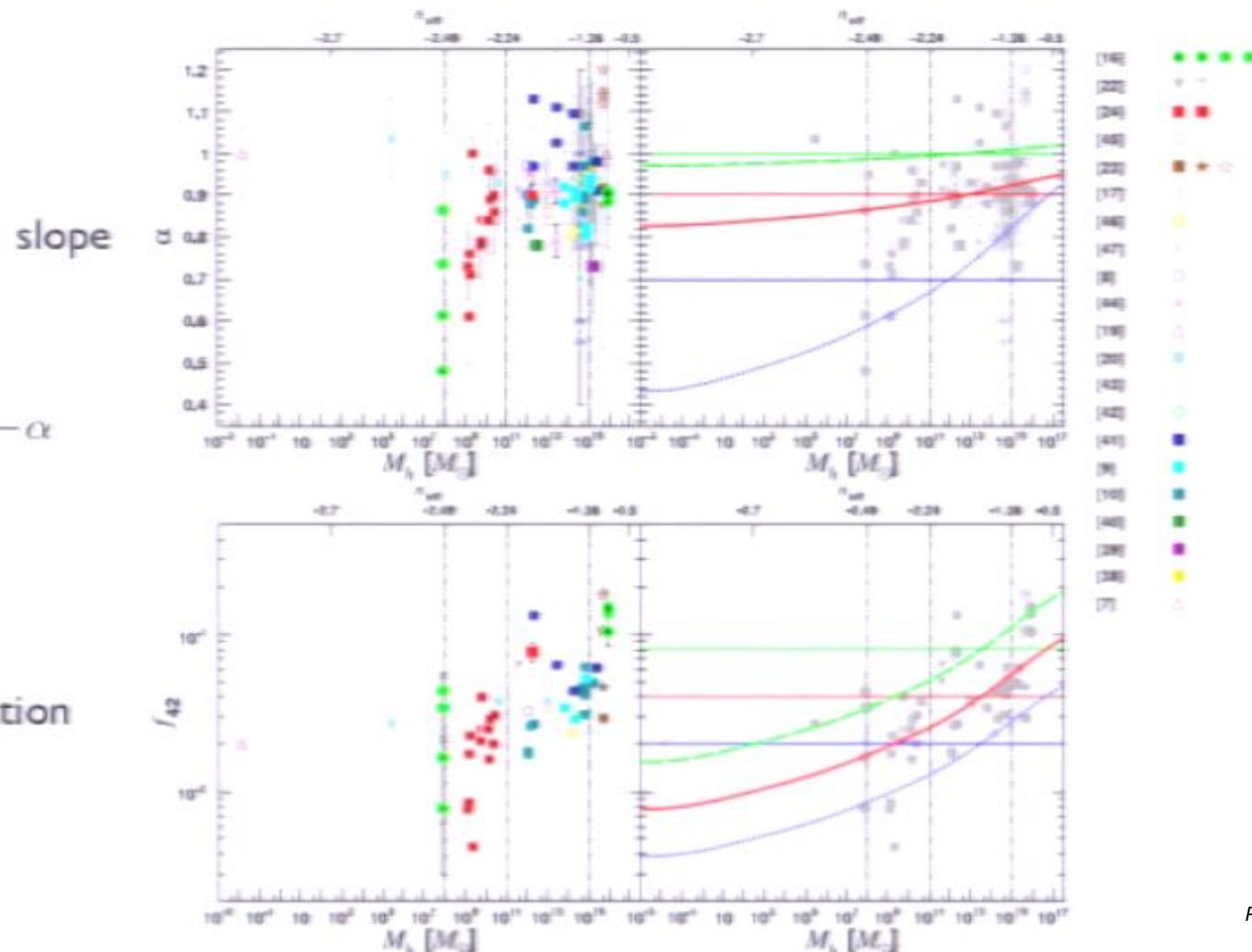
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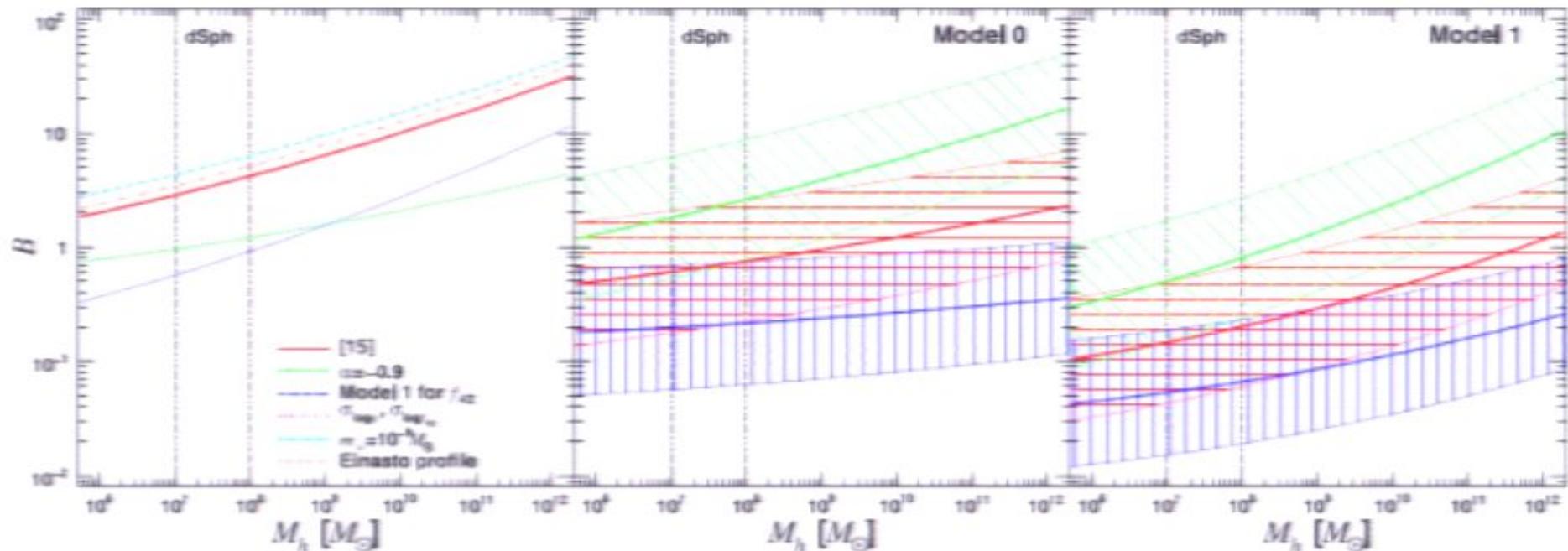
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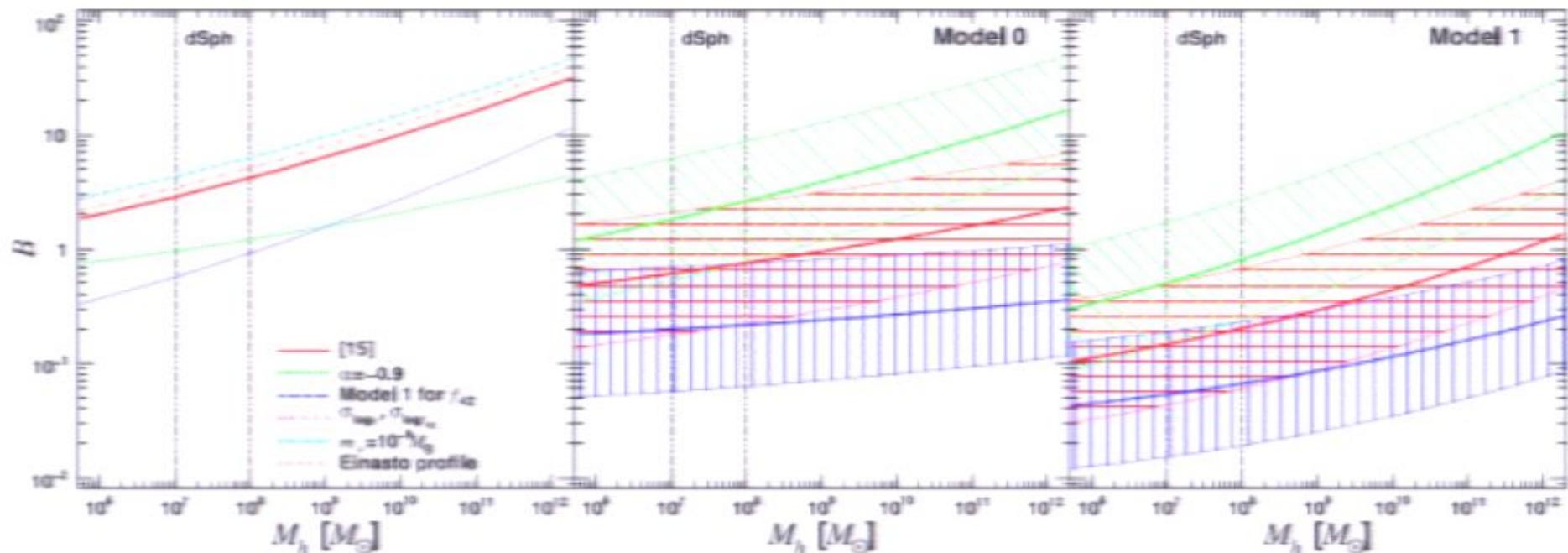
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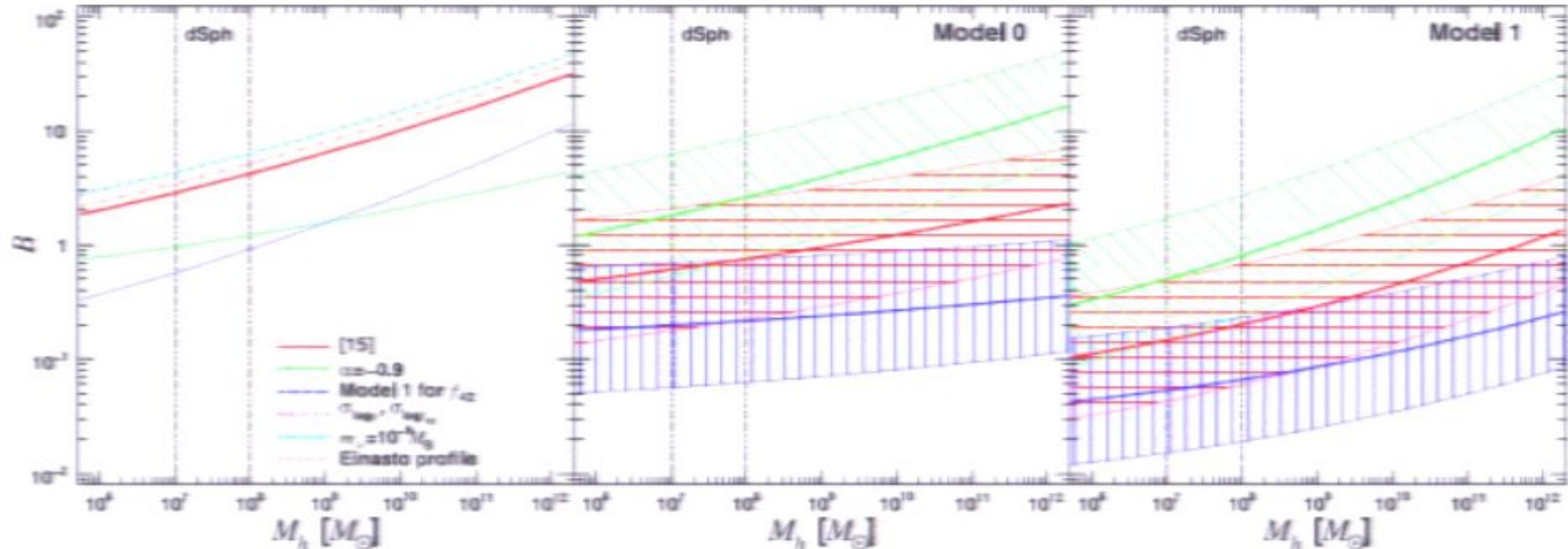
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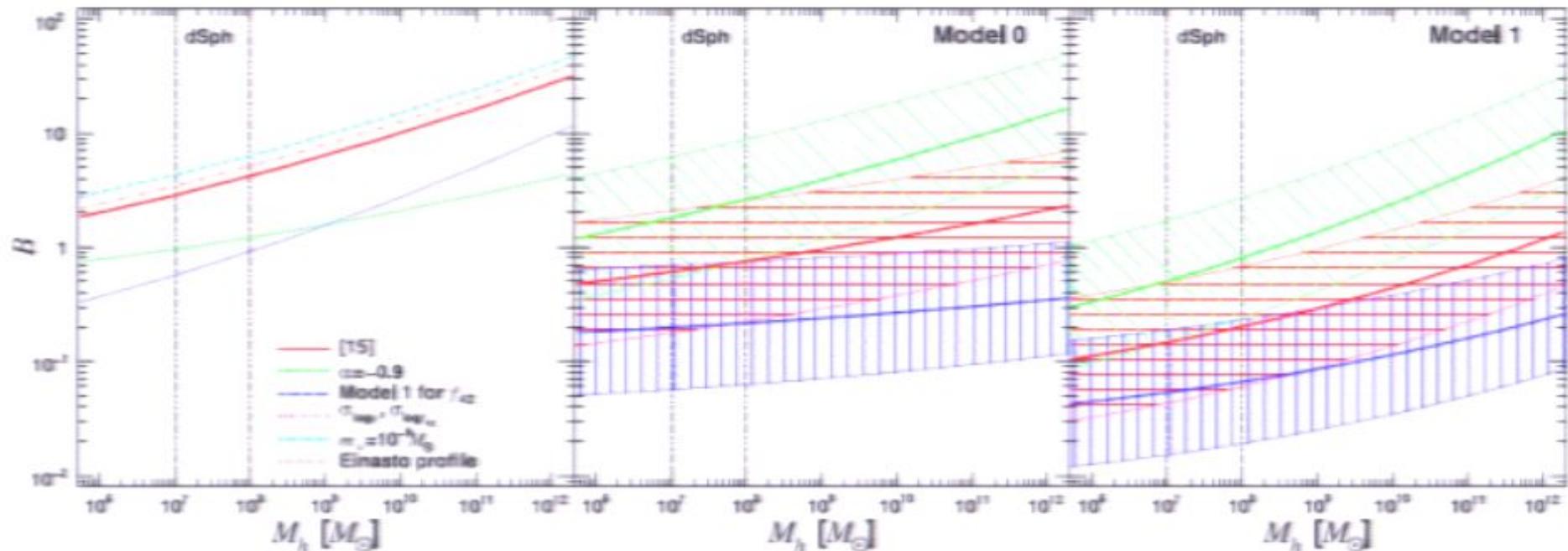
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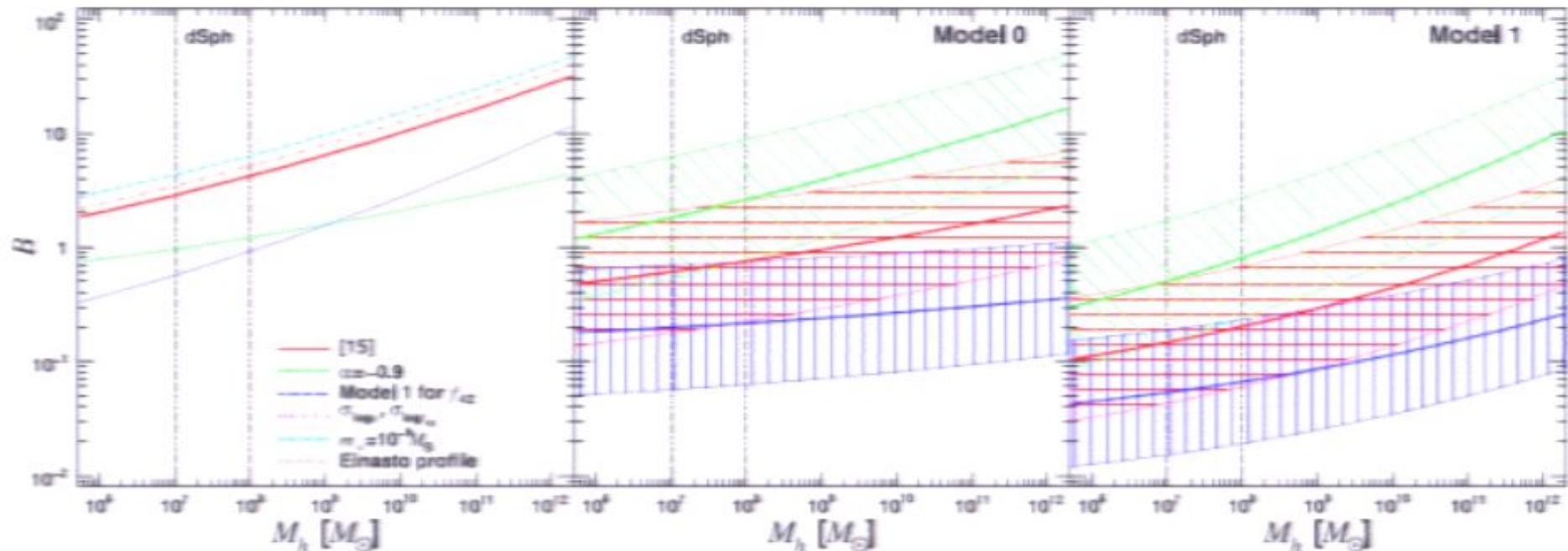
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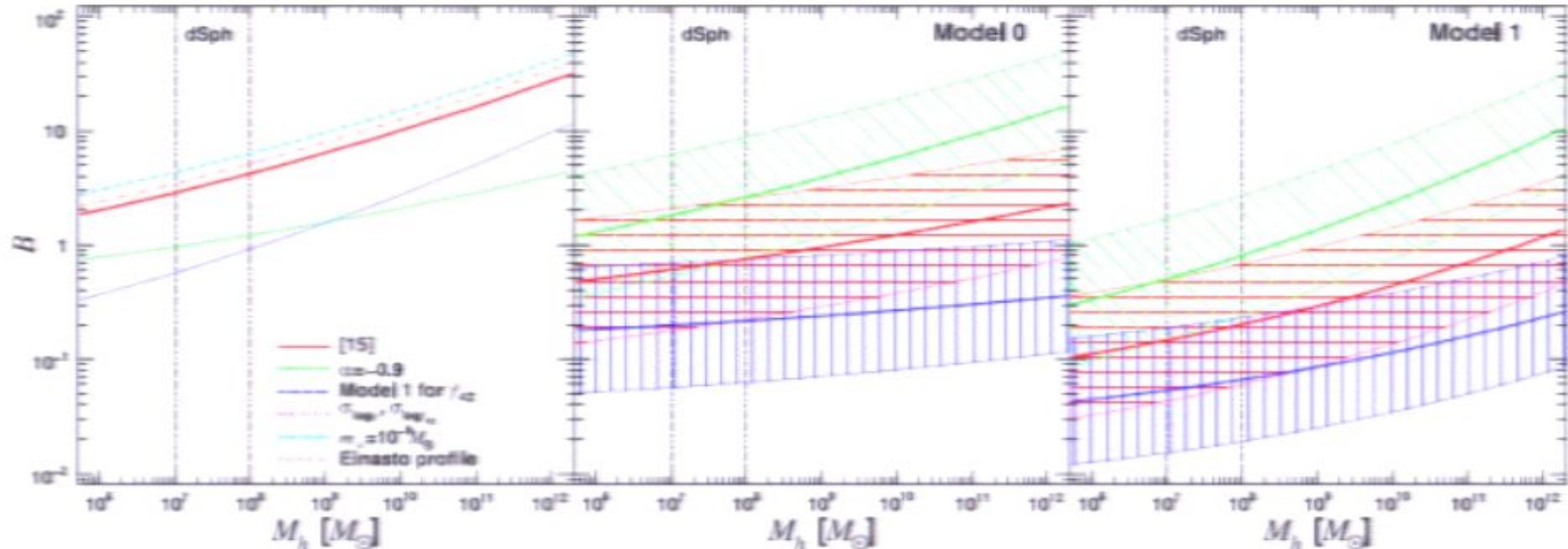
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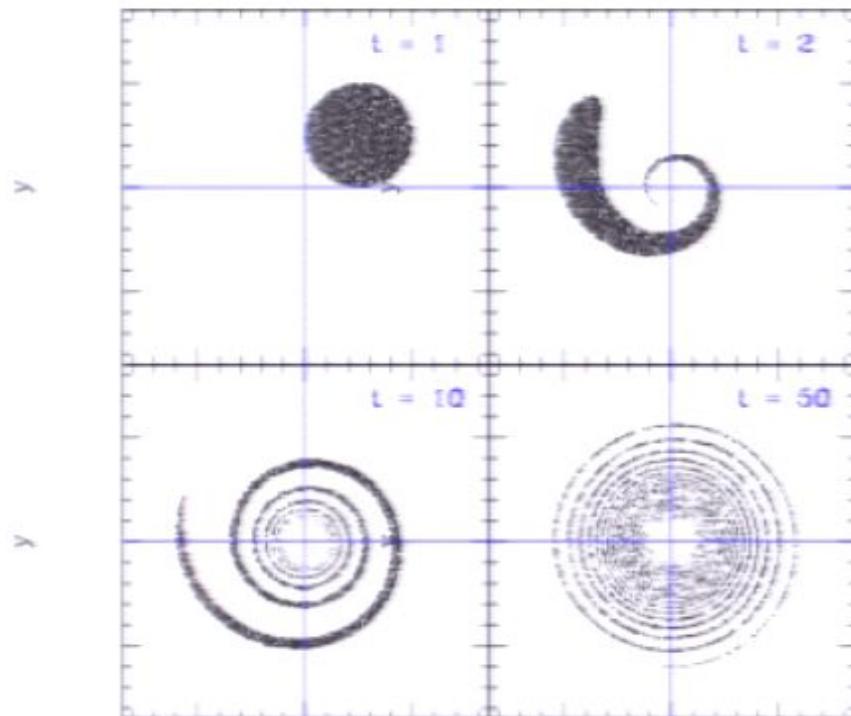
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Challenges, III

Liouville's theorem and the phase-space sheet

for collisionless system's, the phase space density along a particle's orbit is constant



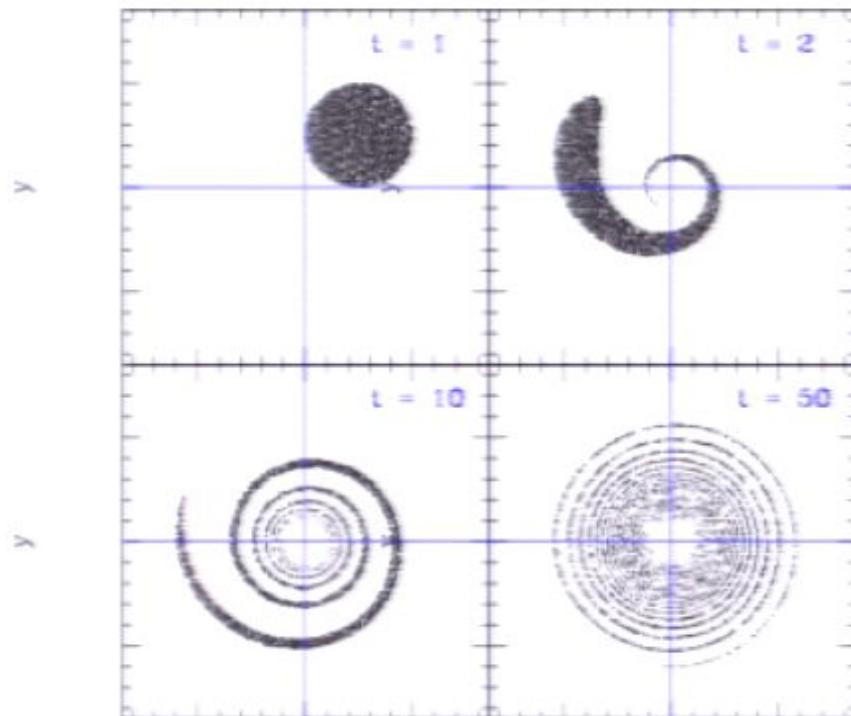
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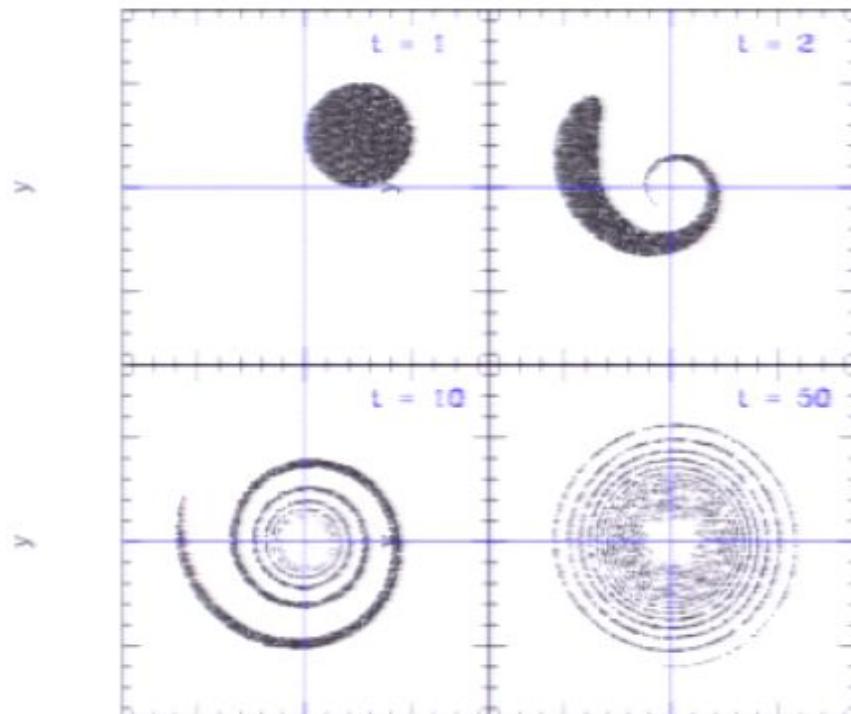
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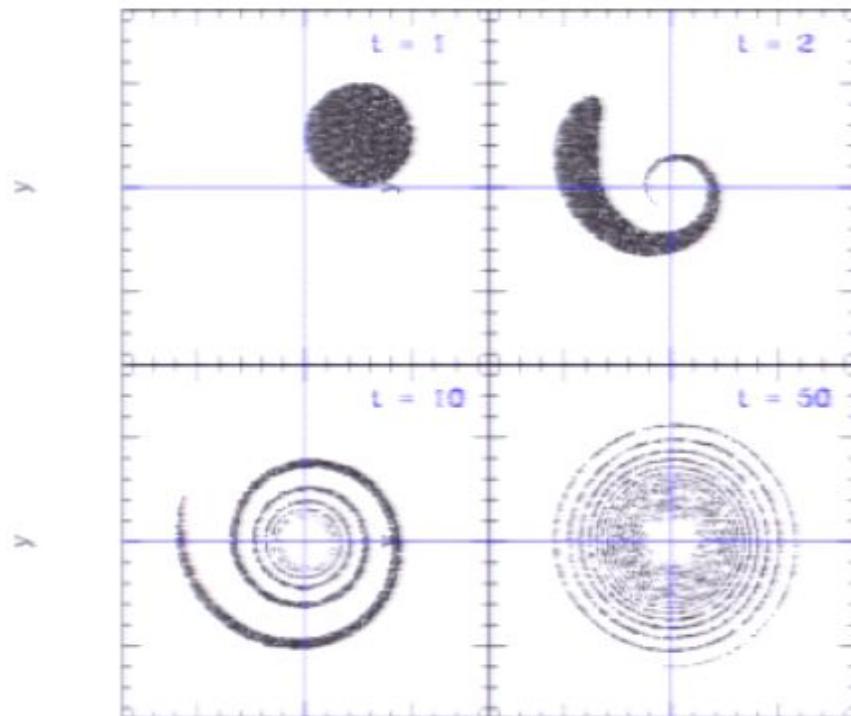
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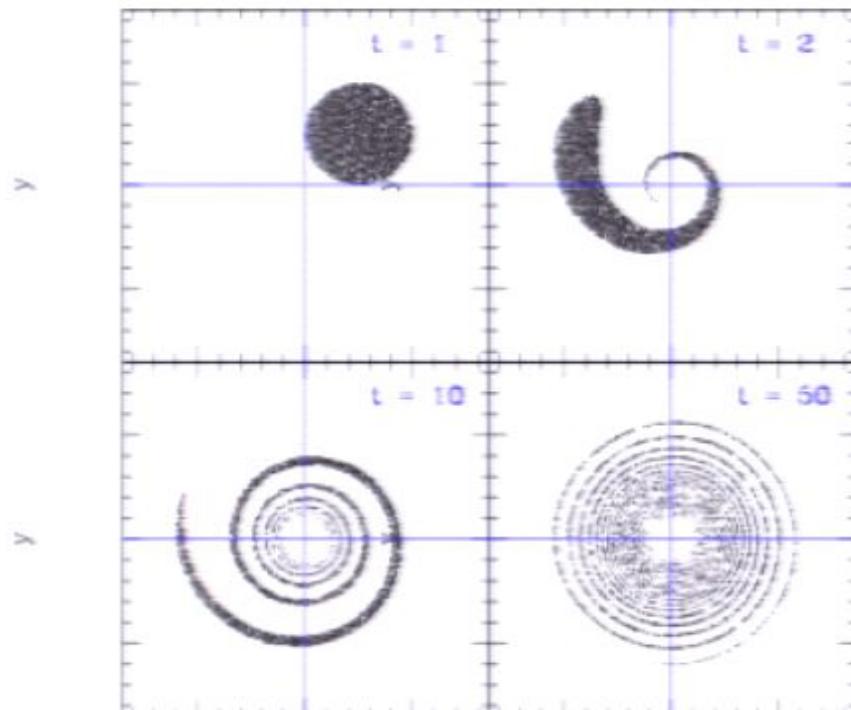
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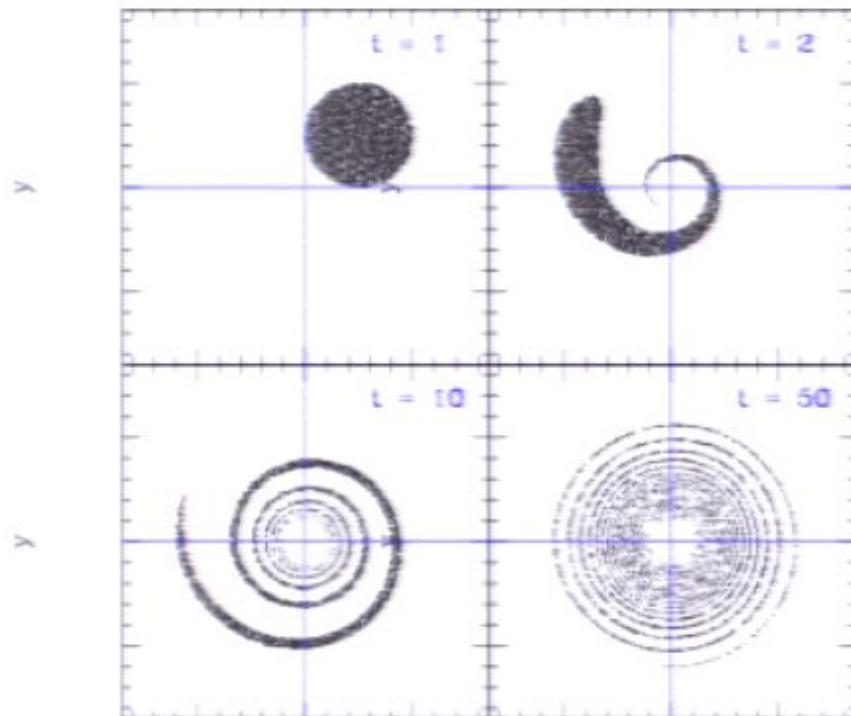
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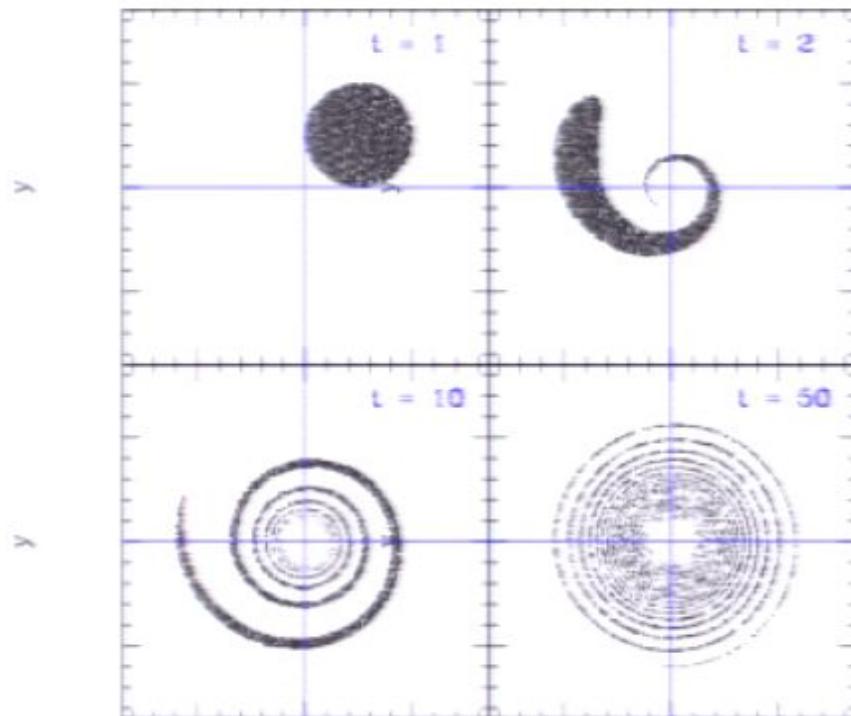
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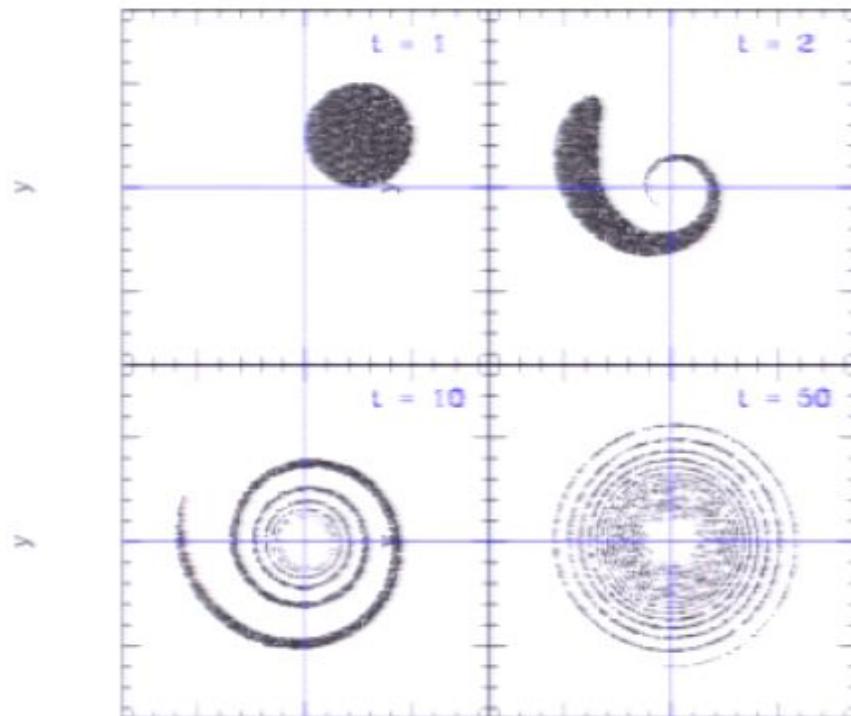
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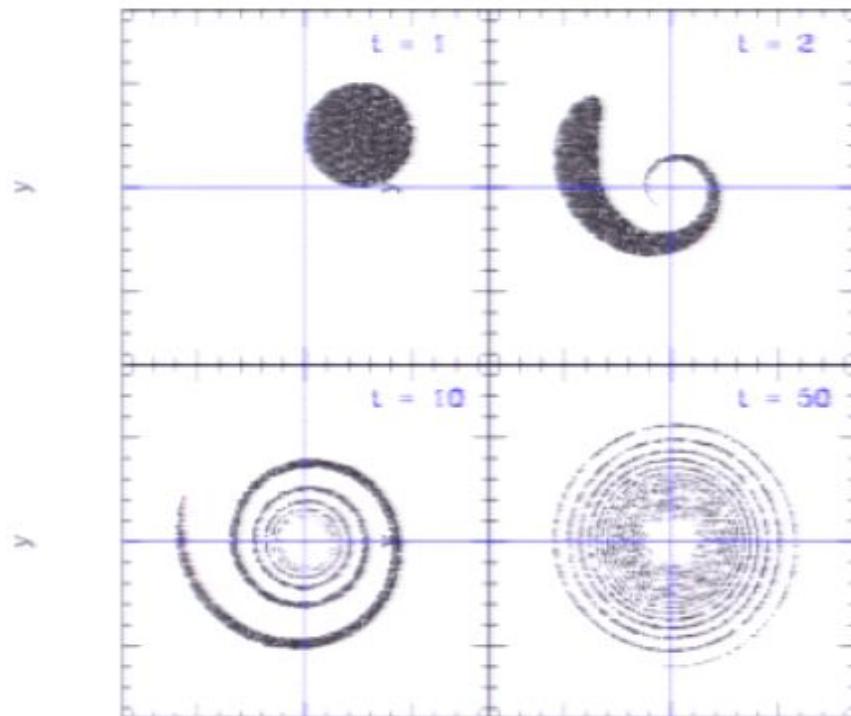
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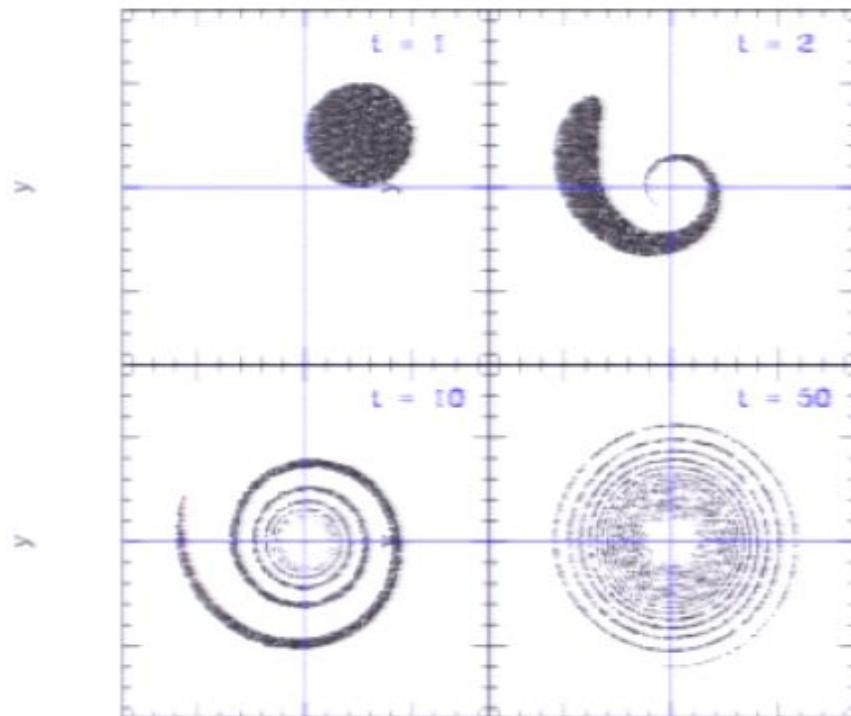
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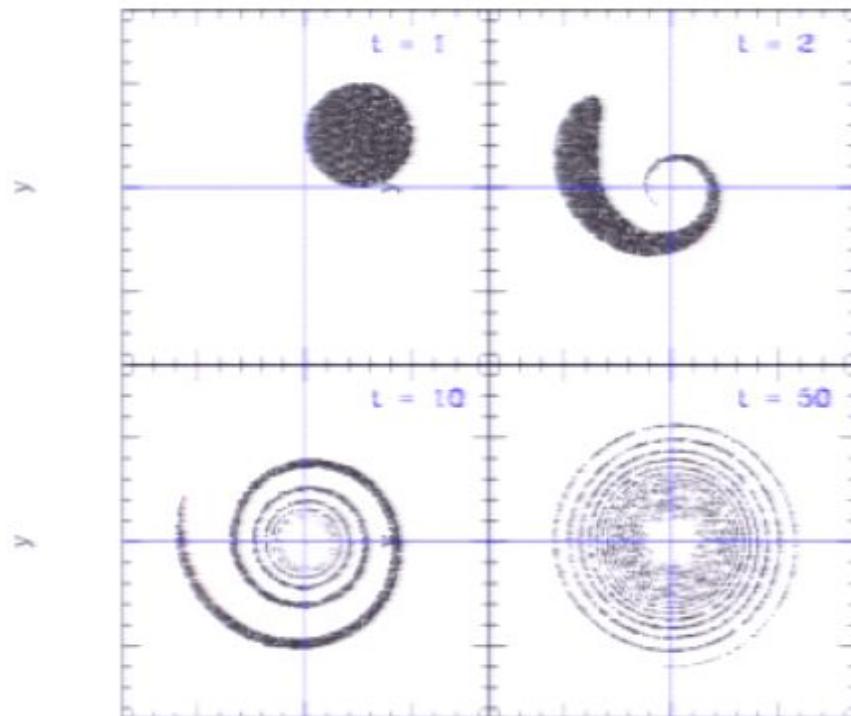
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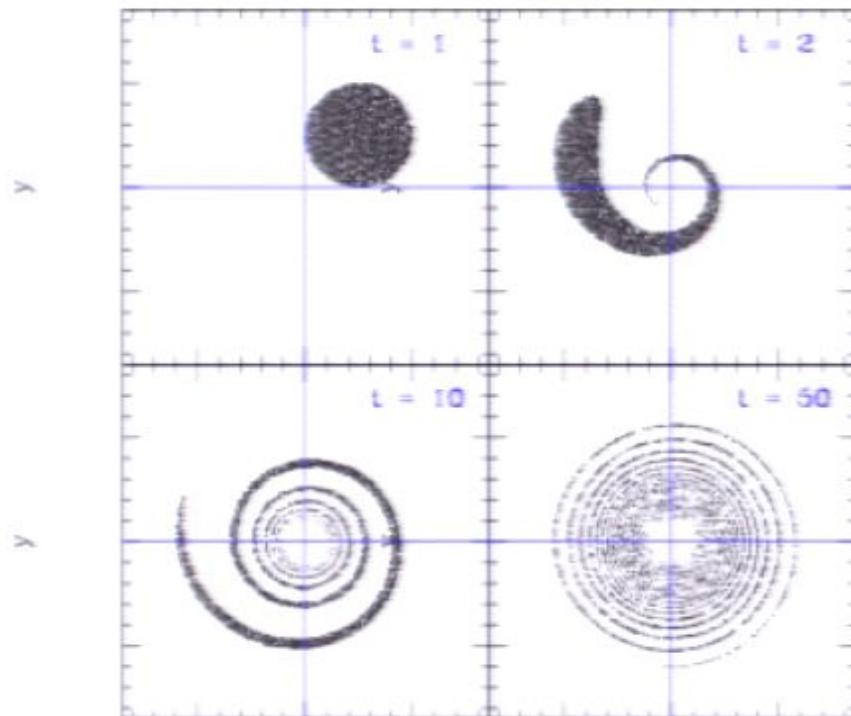
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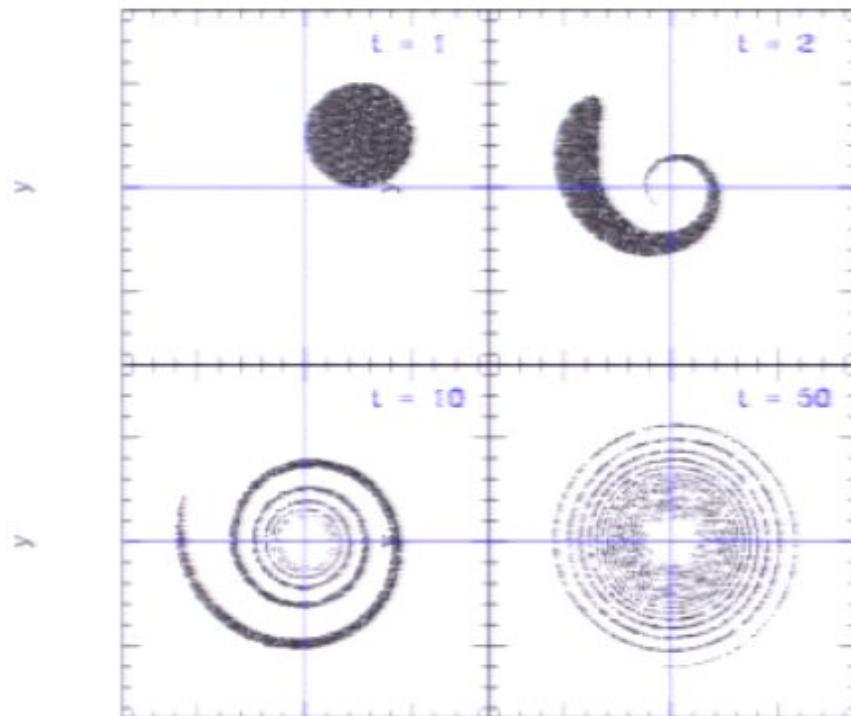
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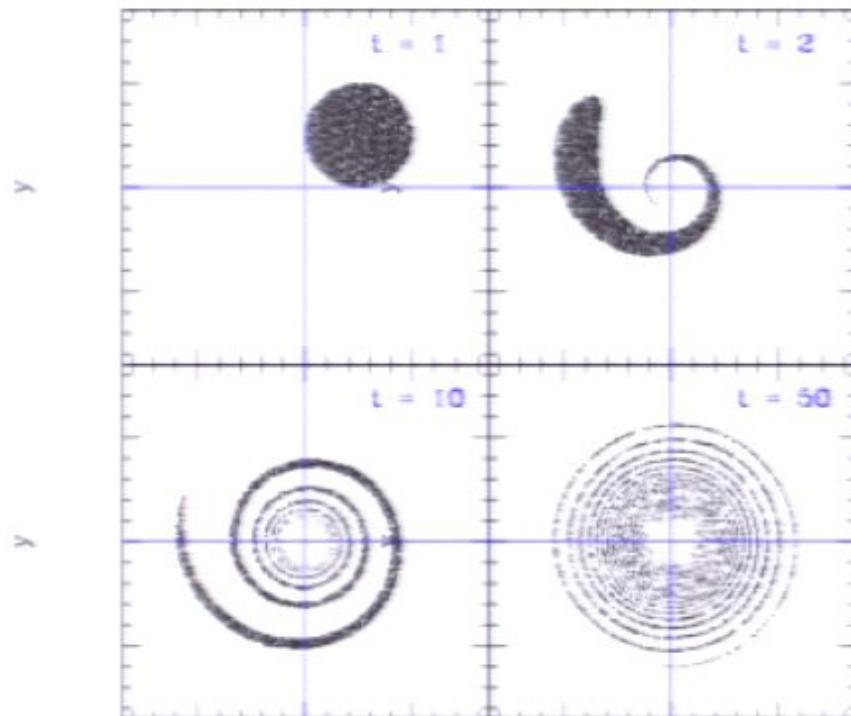
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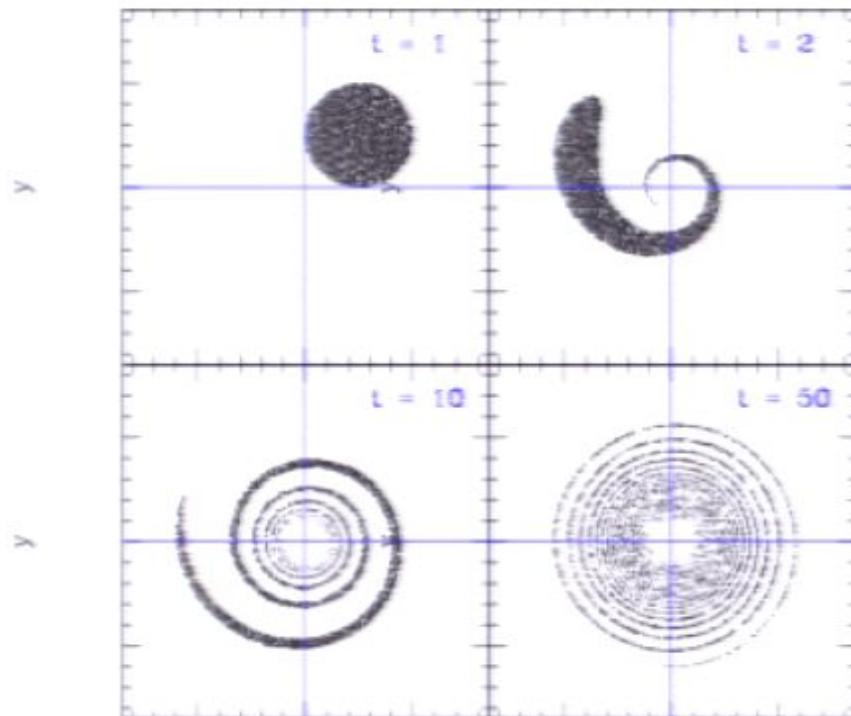
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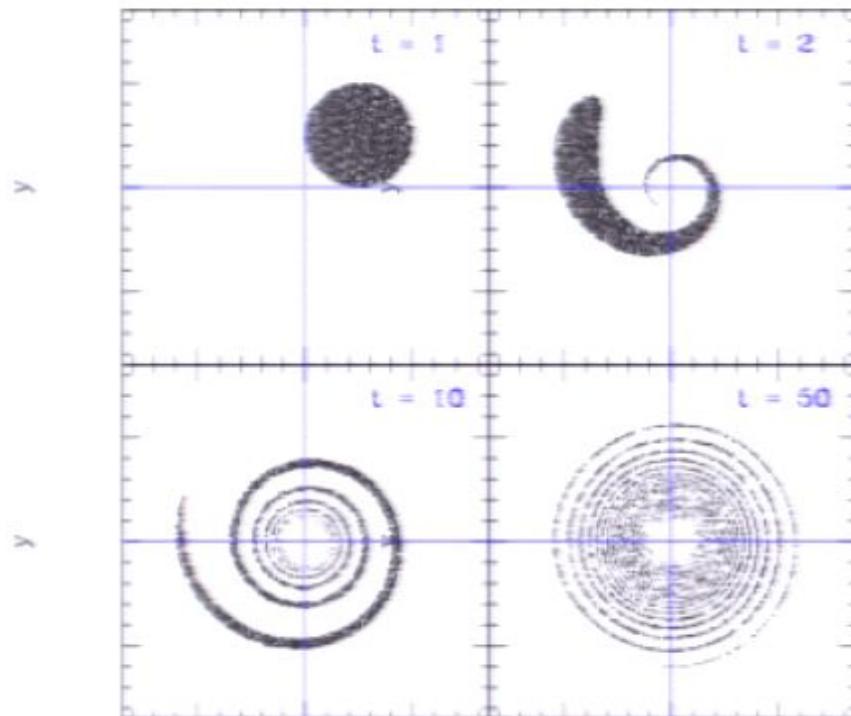
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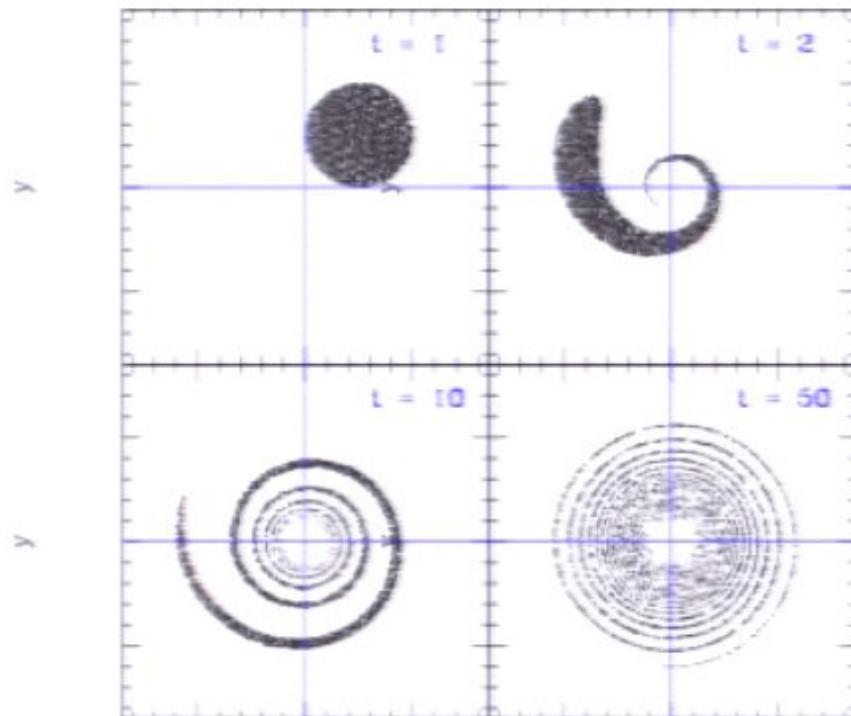
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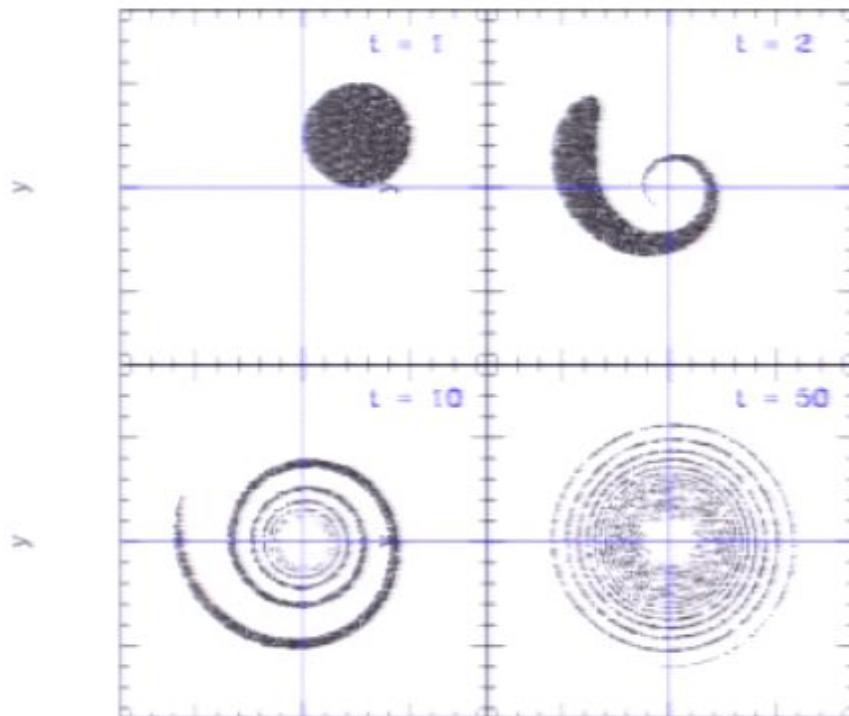
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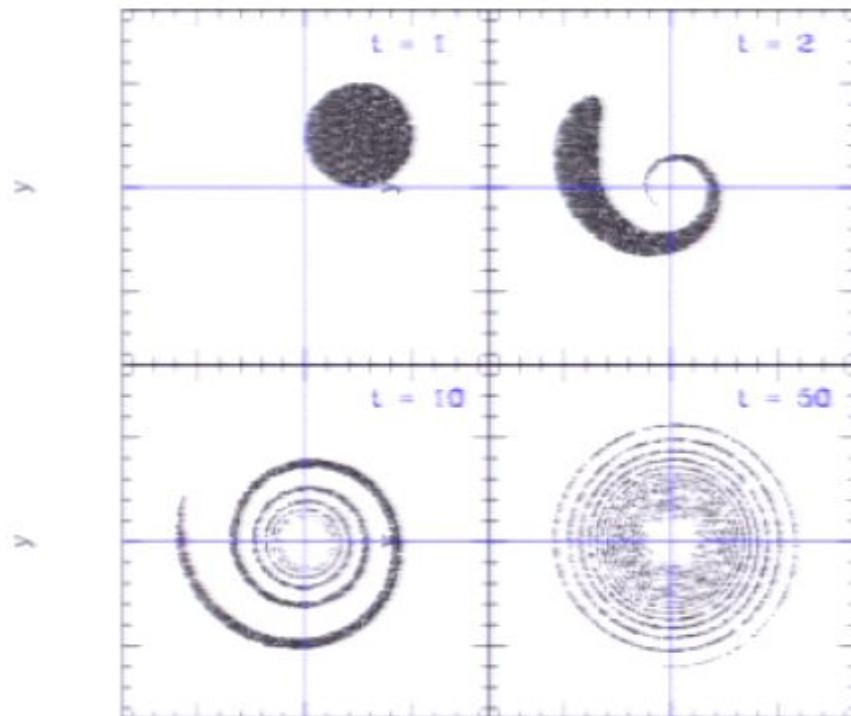
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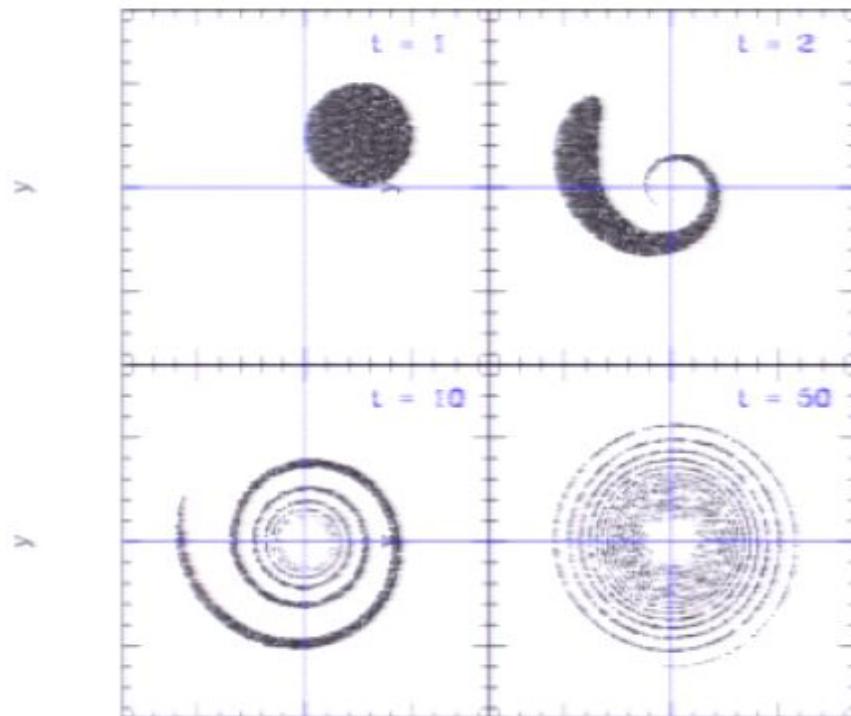
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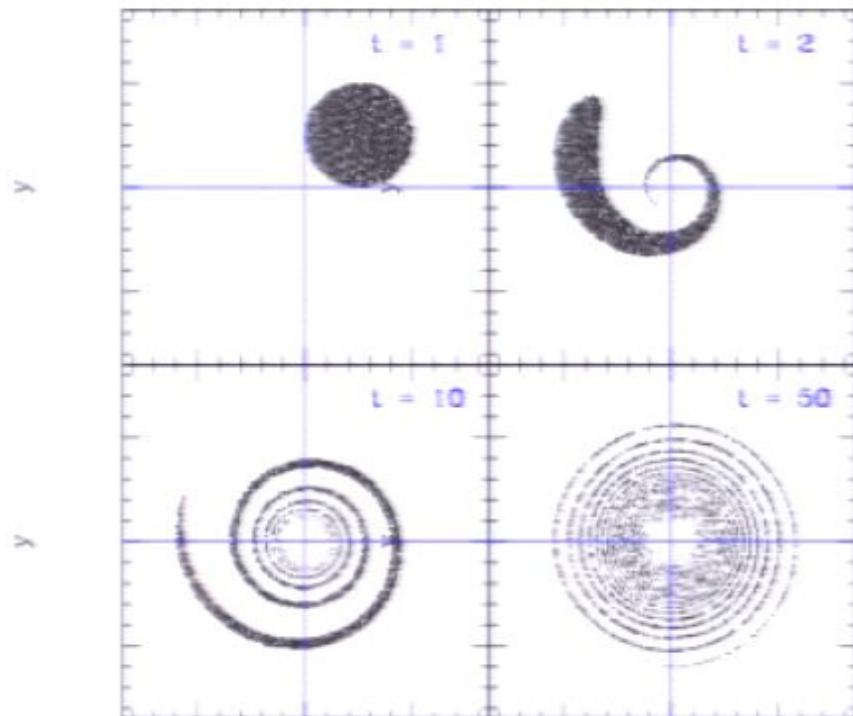
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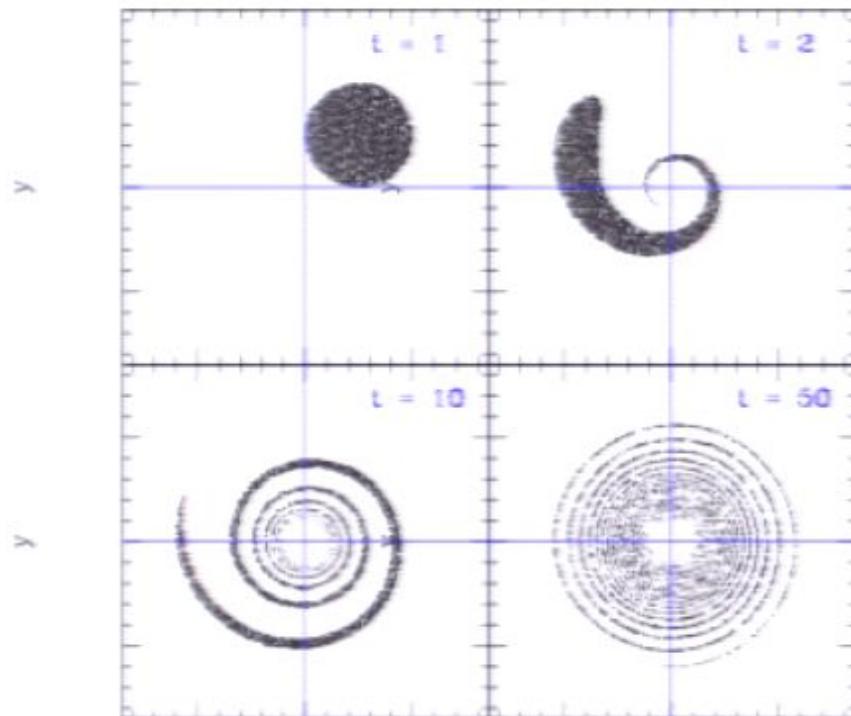
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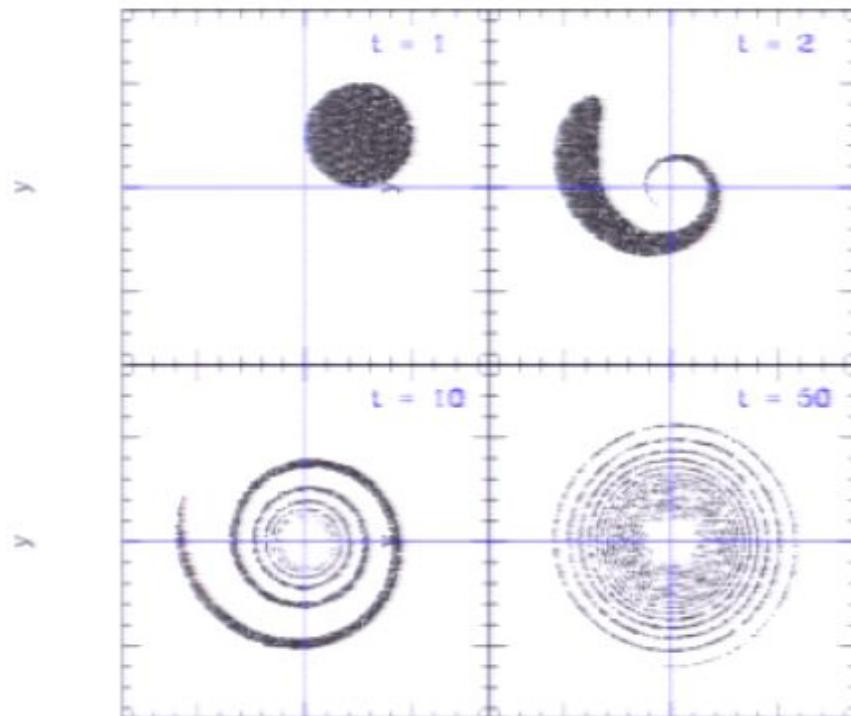
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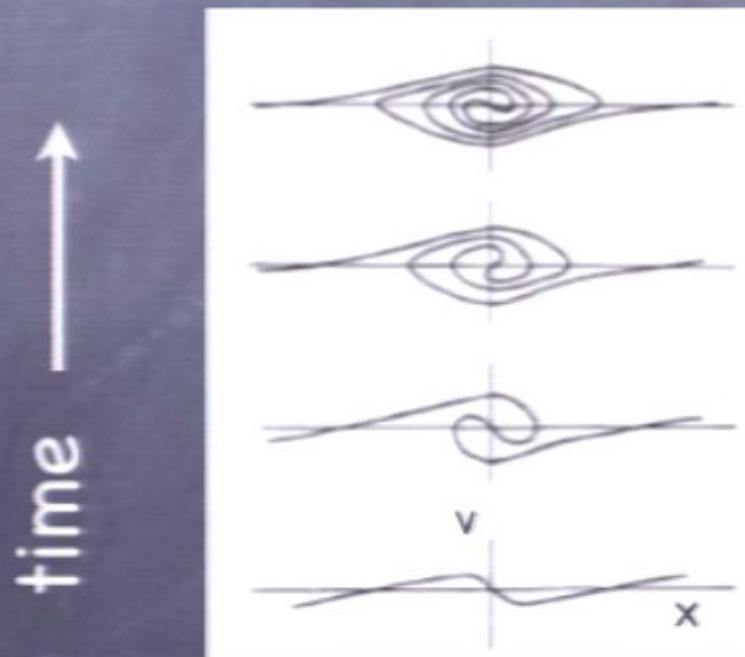
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^b Physics Department, University of Florida, Gainesville, FL 32611, USA

Received: 3 July 1992

We discuss the extent to which the phase-space distribution of cold dark matter particles is thermalized in a galactic halo and find that there are large deviations from a thermal distribution in that the highest energy particles have discrete values of velocity. The central values and intensities of the corresponding peaks provide detailed information on the history of the Galaxy. This information becomes immediately available if a signal is found in a cavity detector of galactic halo axions.



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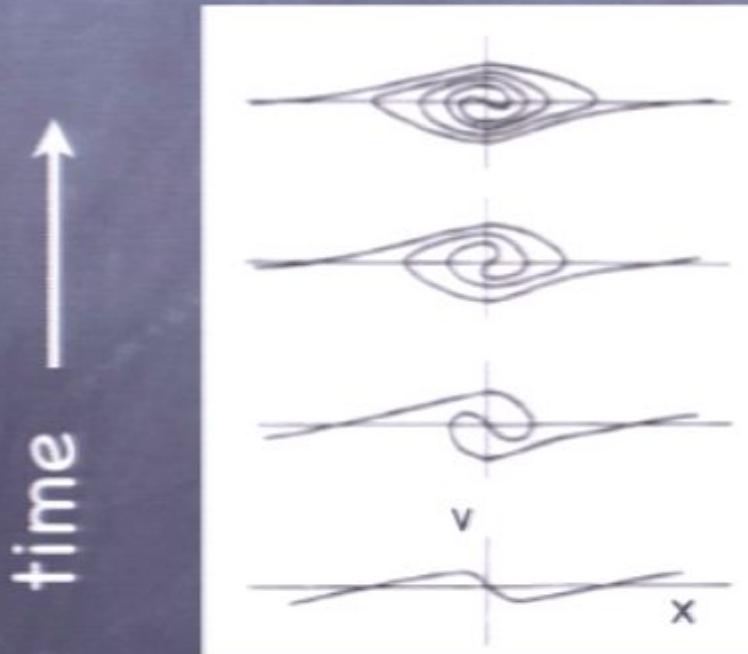
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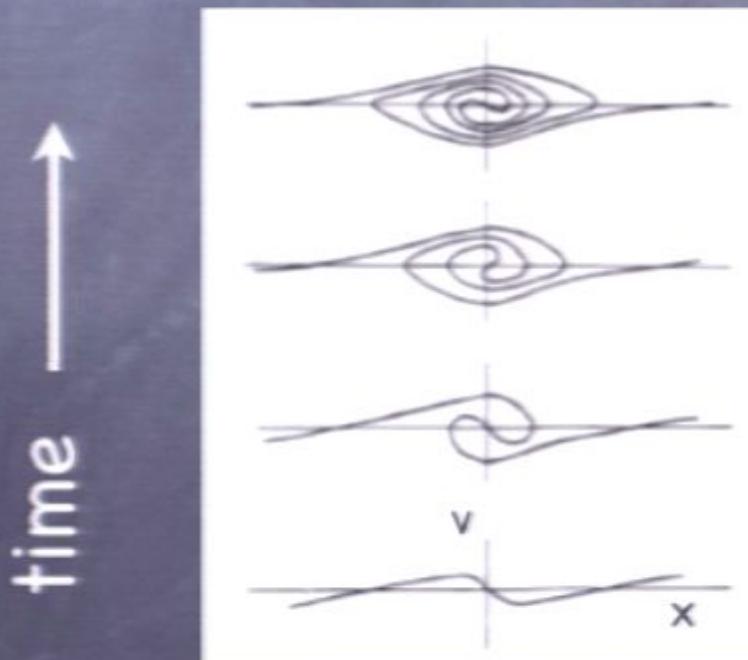
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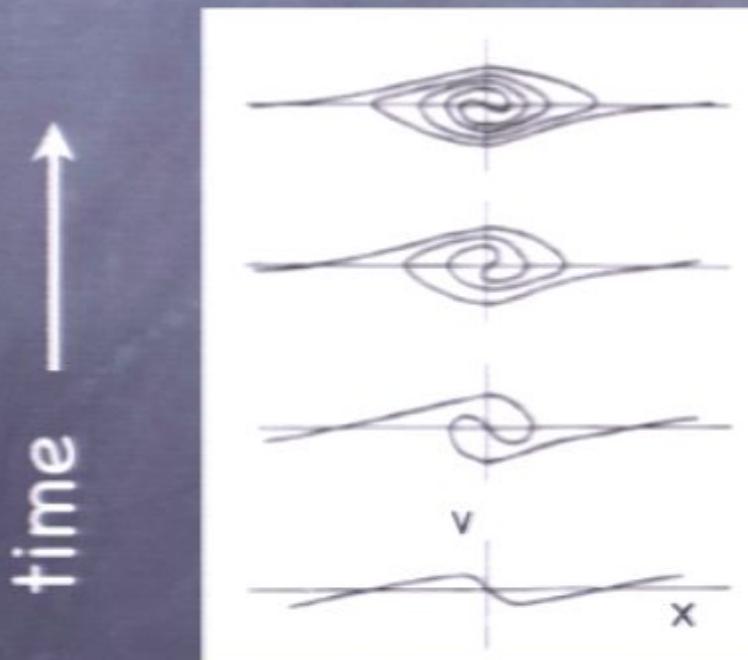
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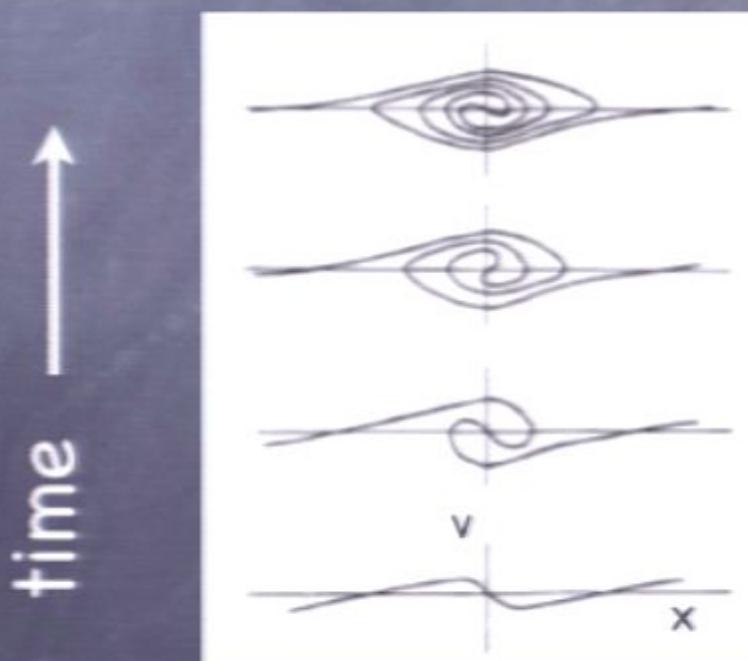
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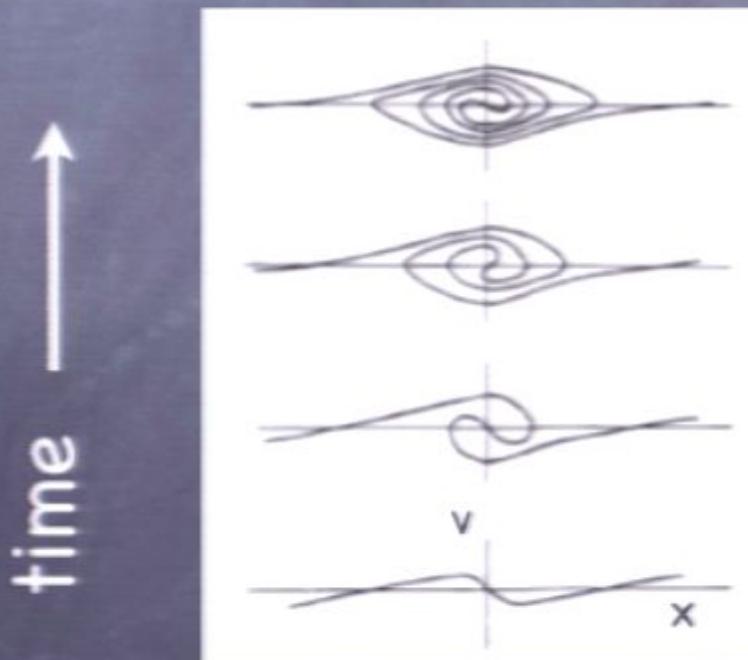
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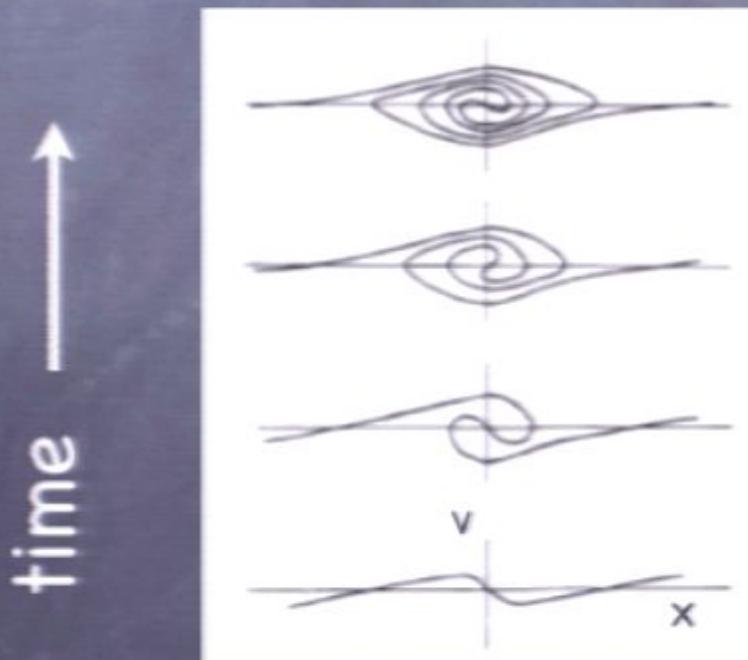
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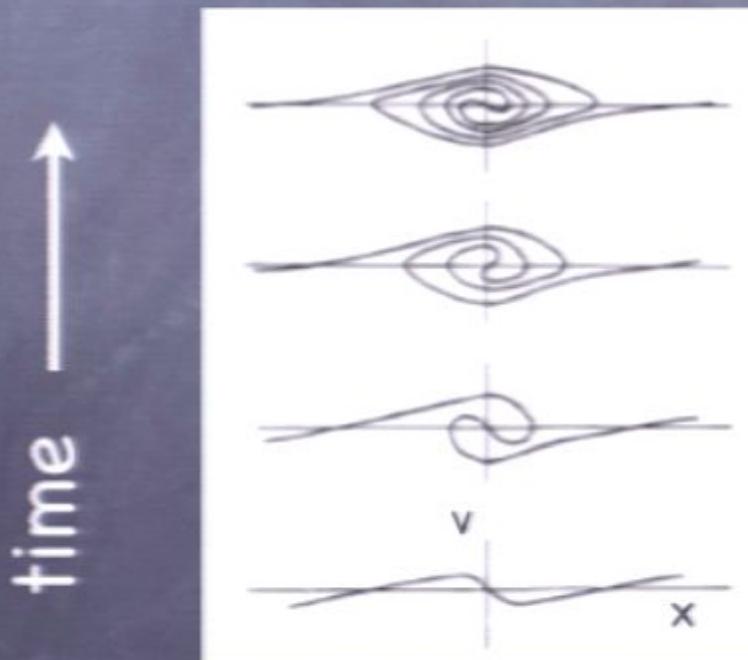
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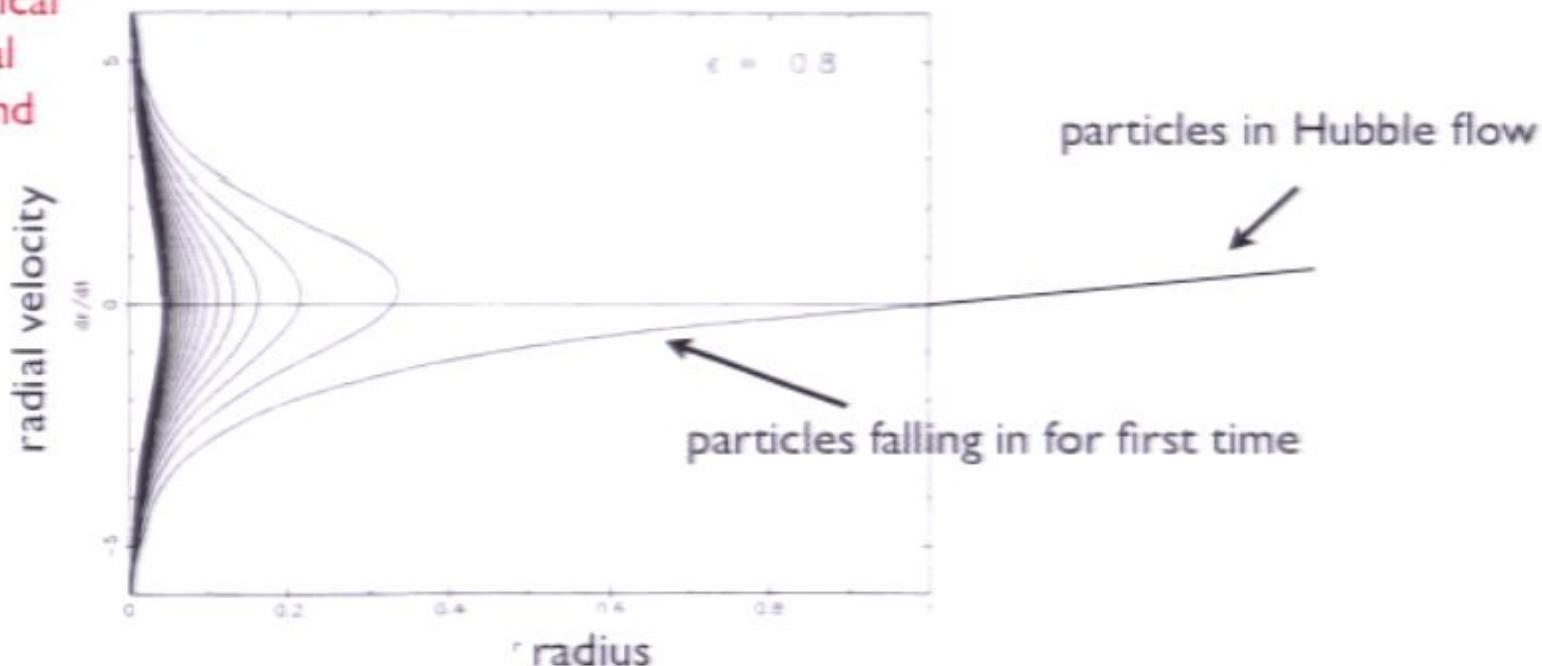
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each point on curve represents spherical shell with radial velocity dr/dt and radius $r(t)$



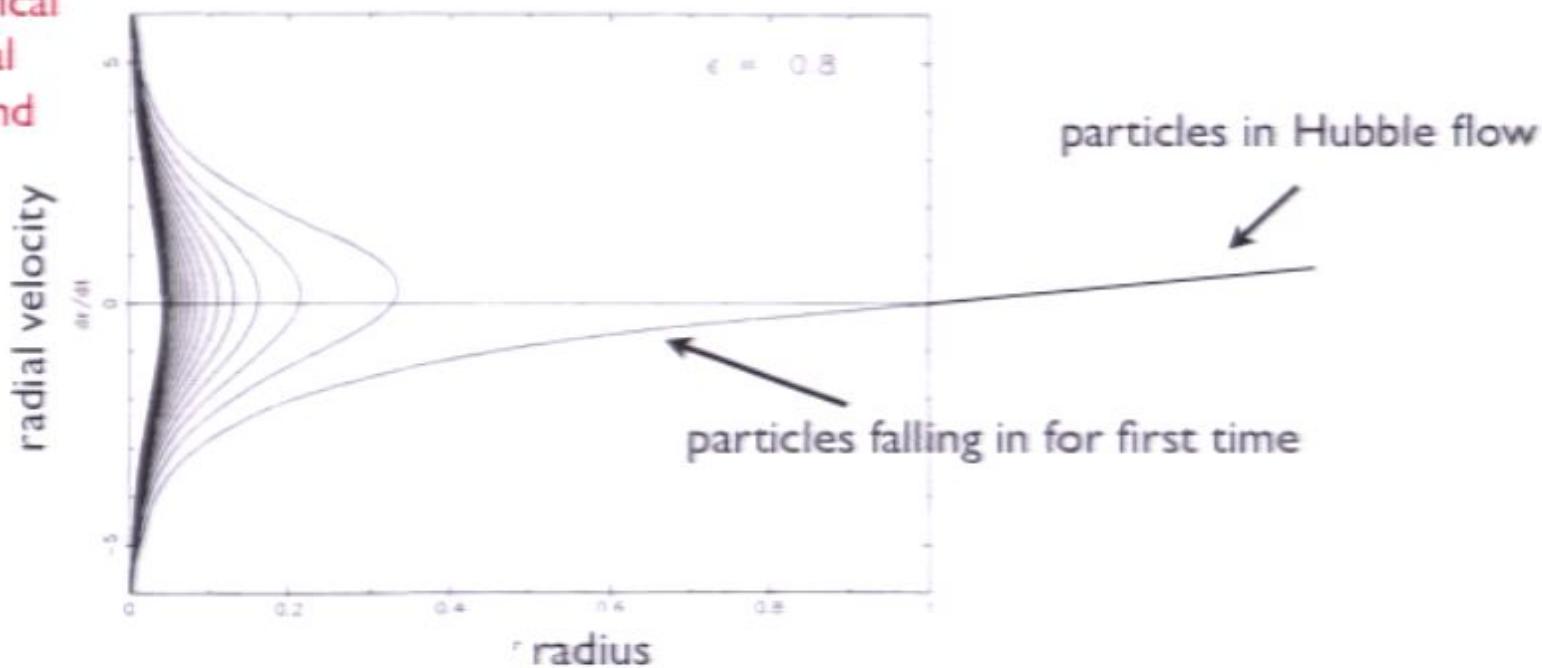
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Note that this solution

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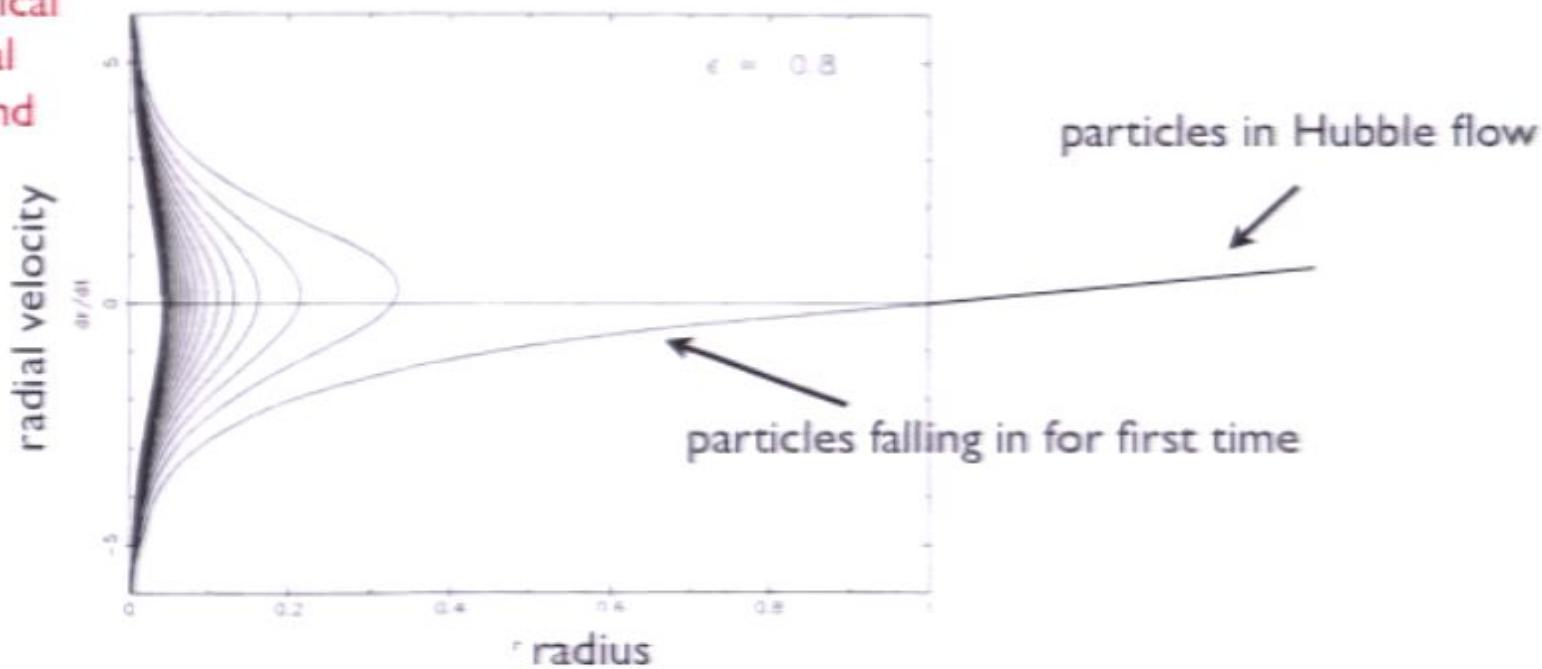
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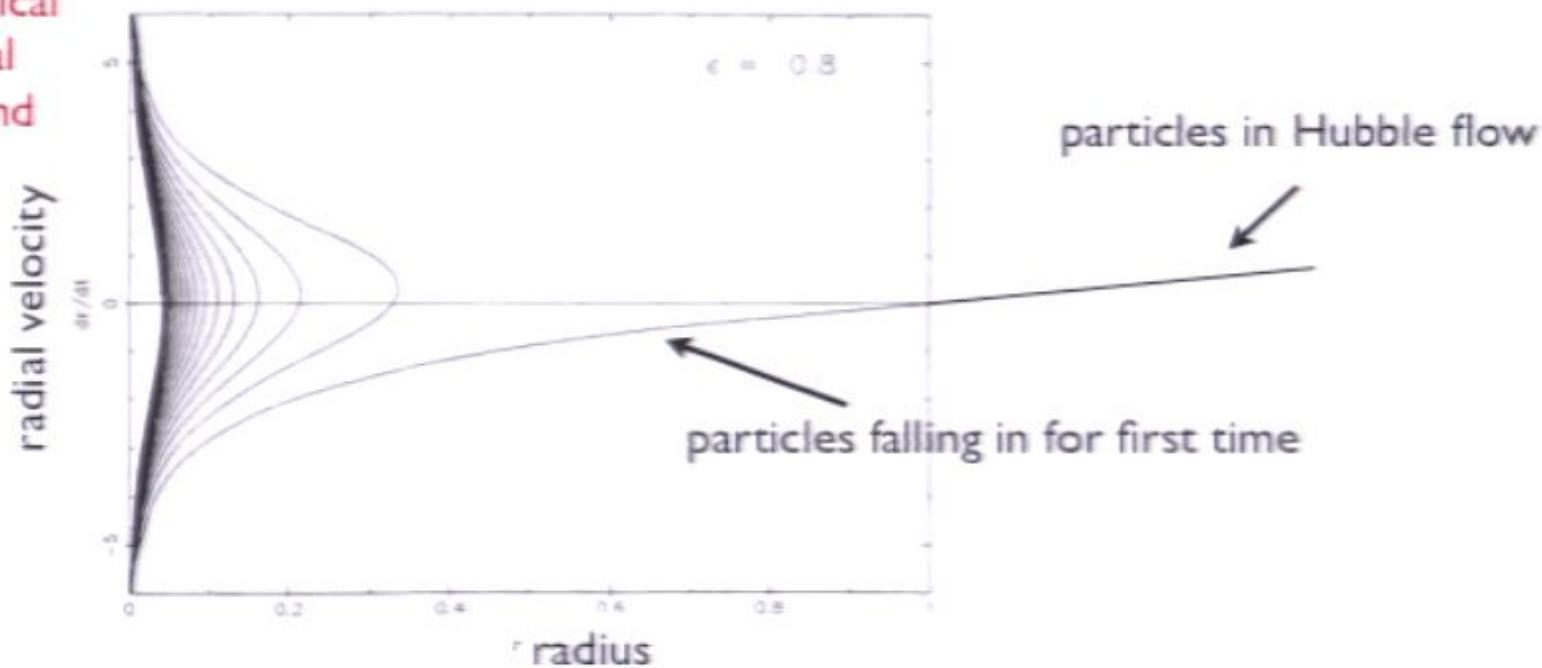
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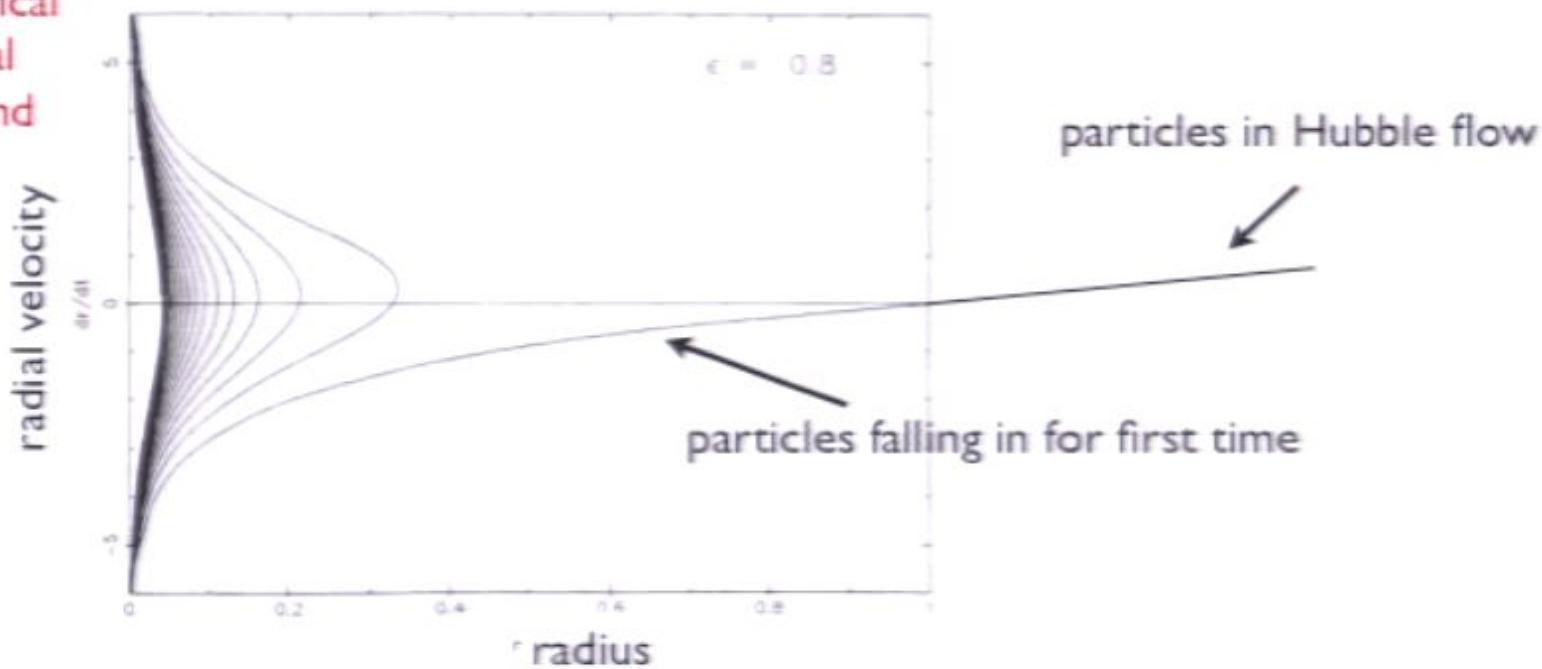
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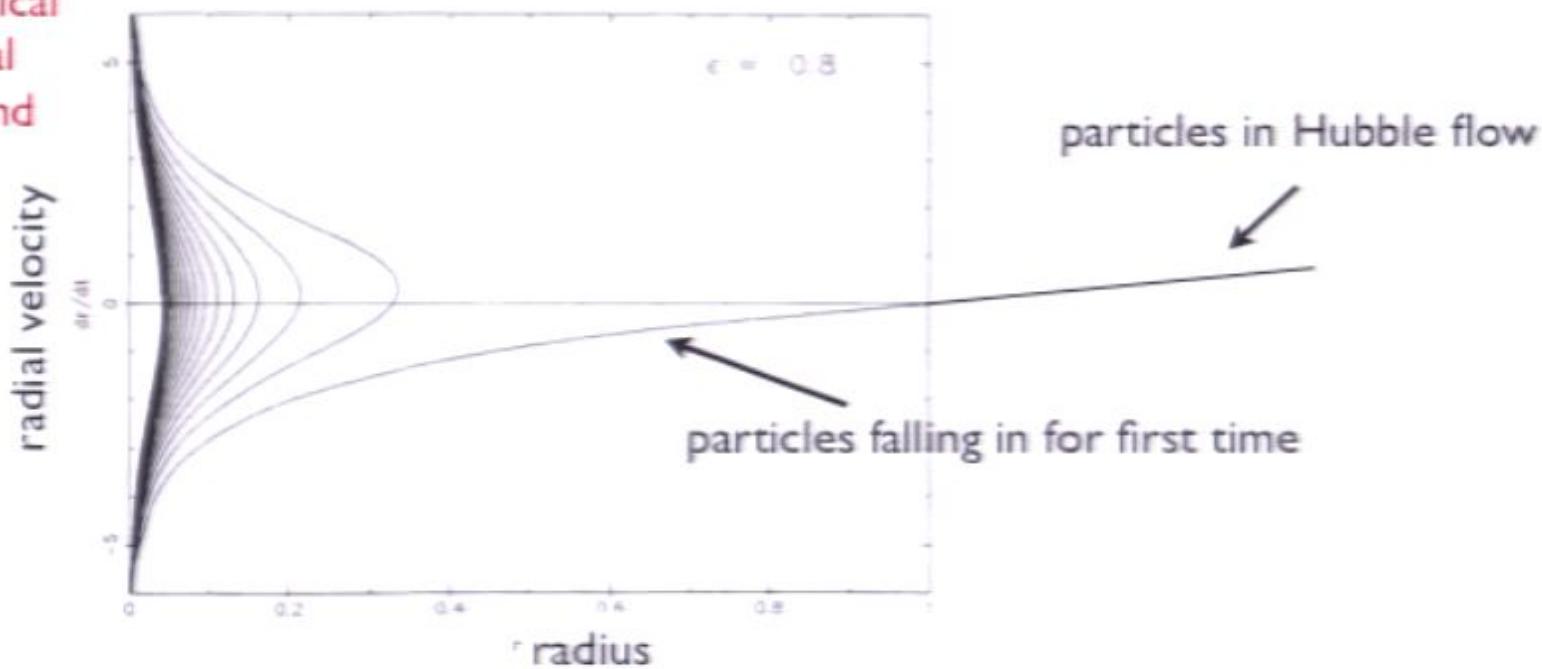
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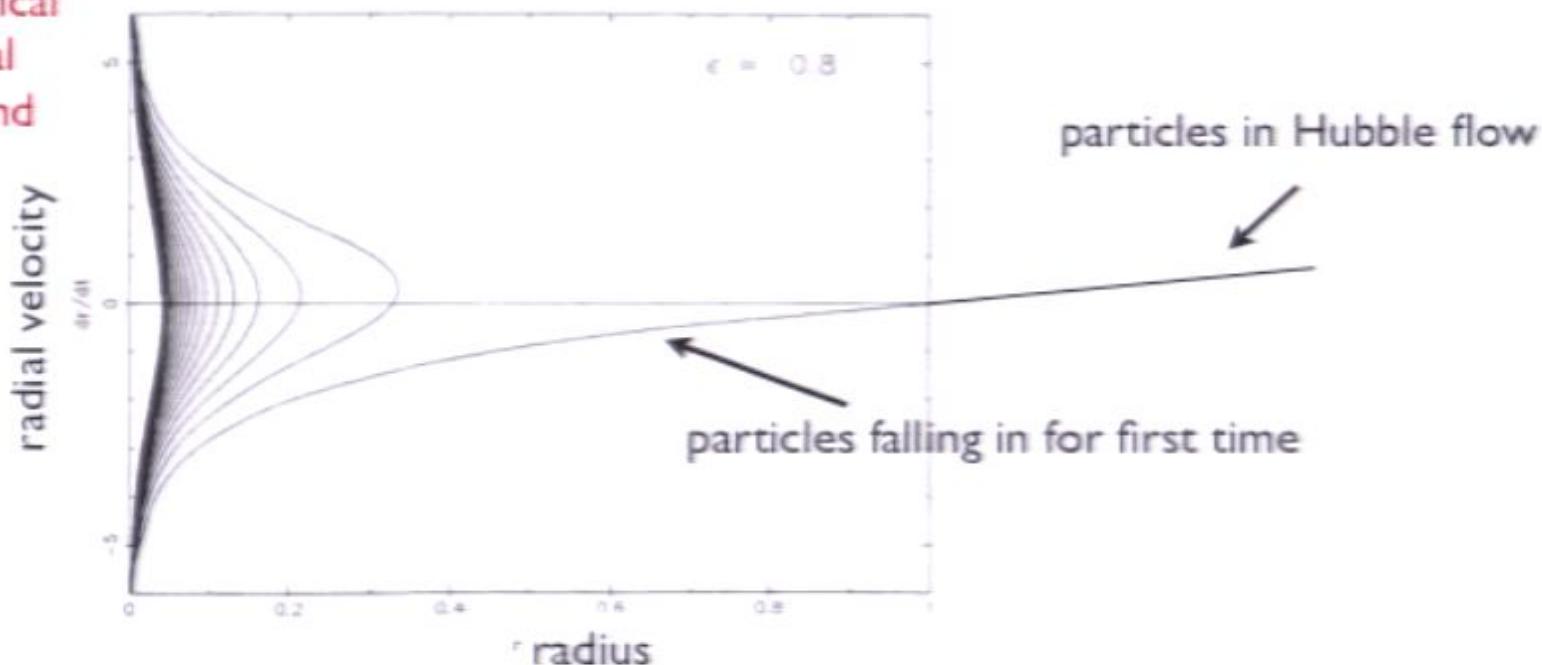
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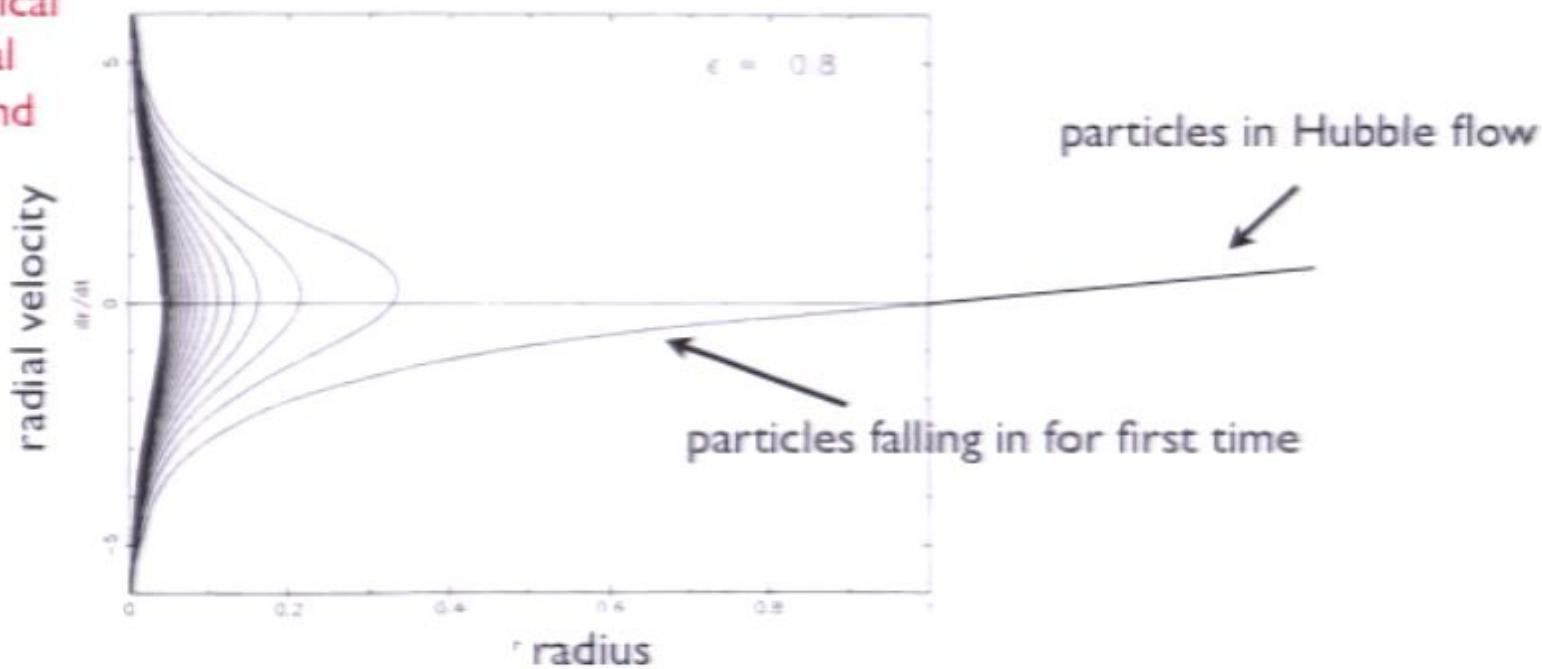
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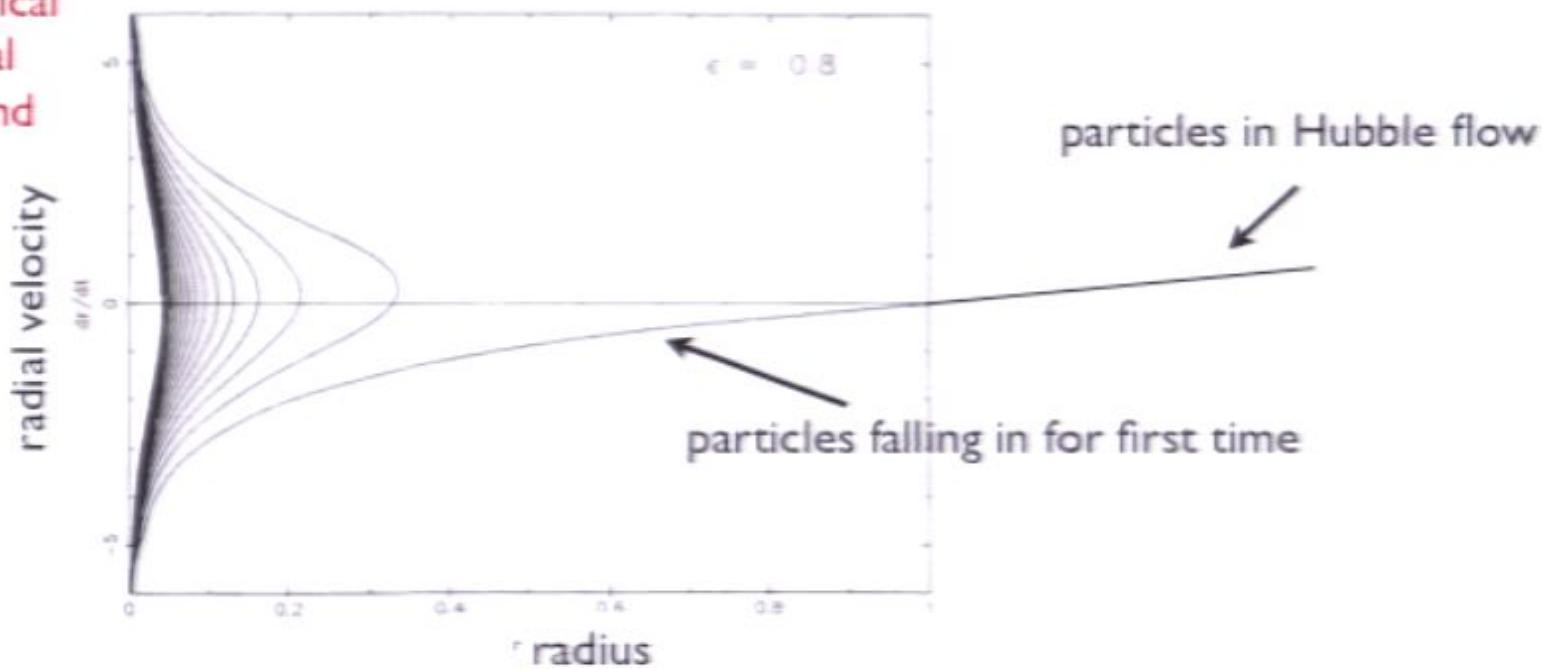
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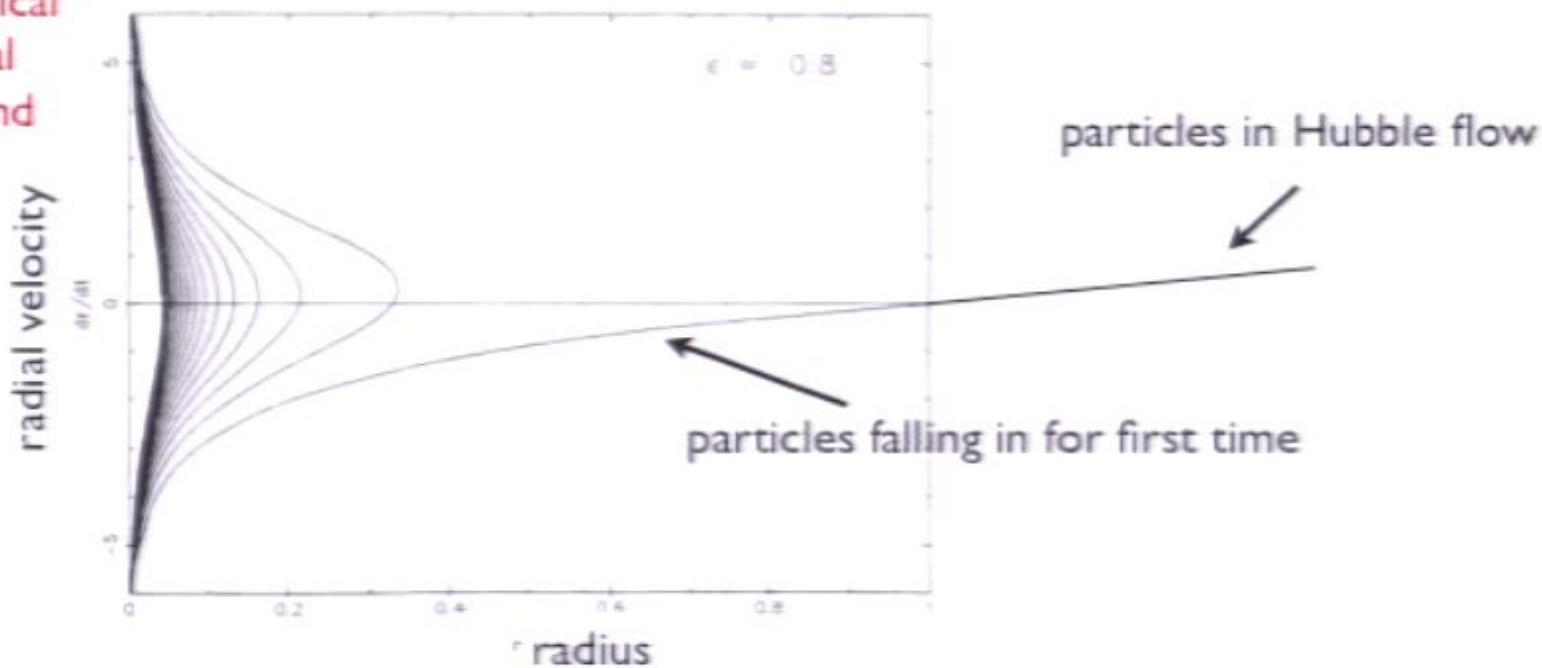
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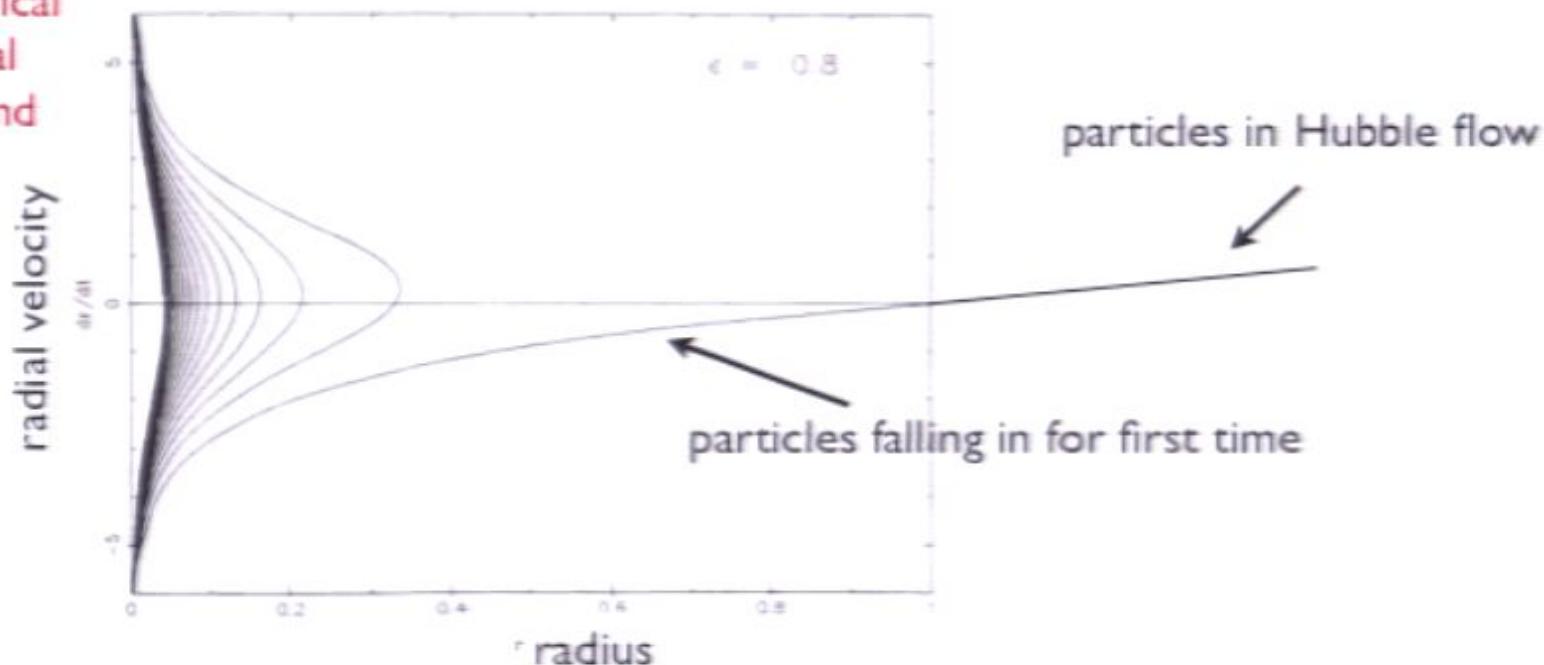
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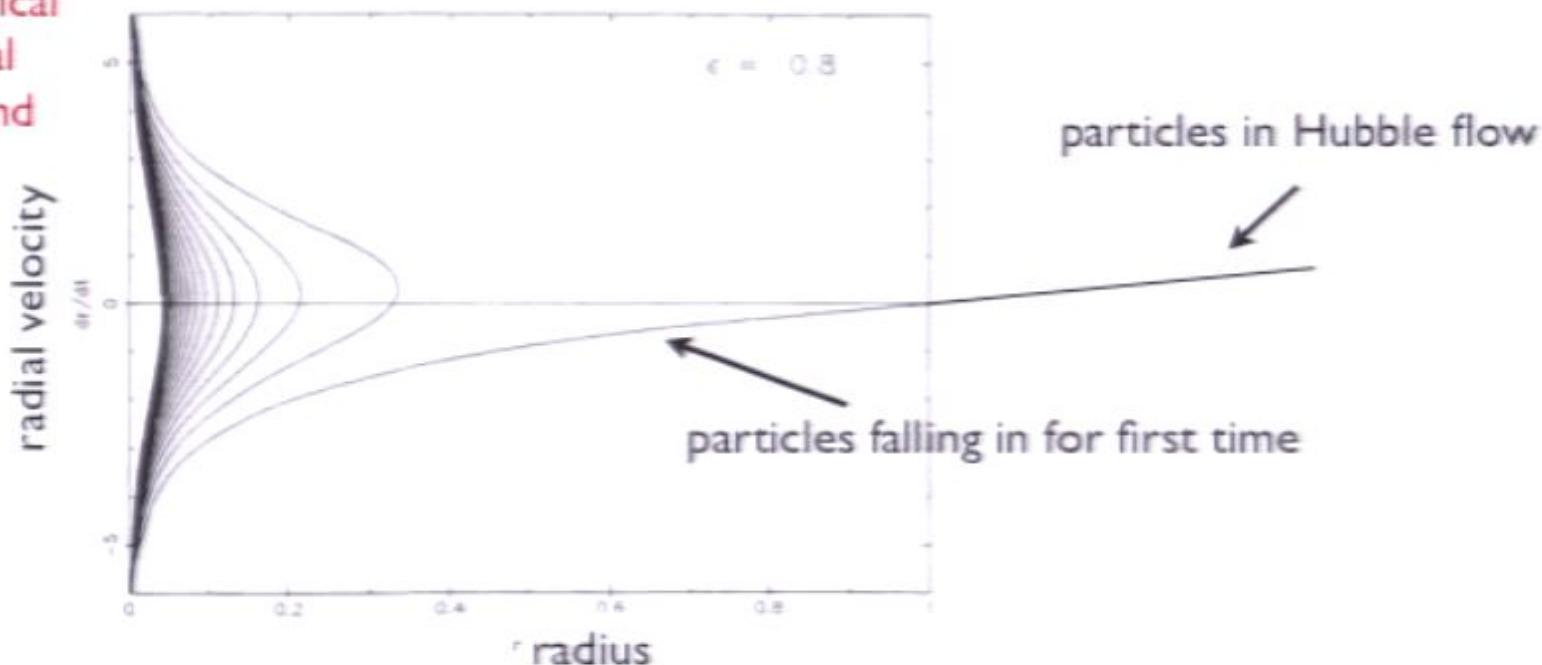
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represents spherical
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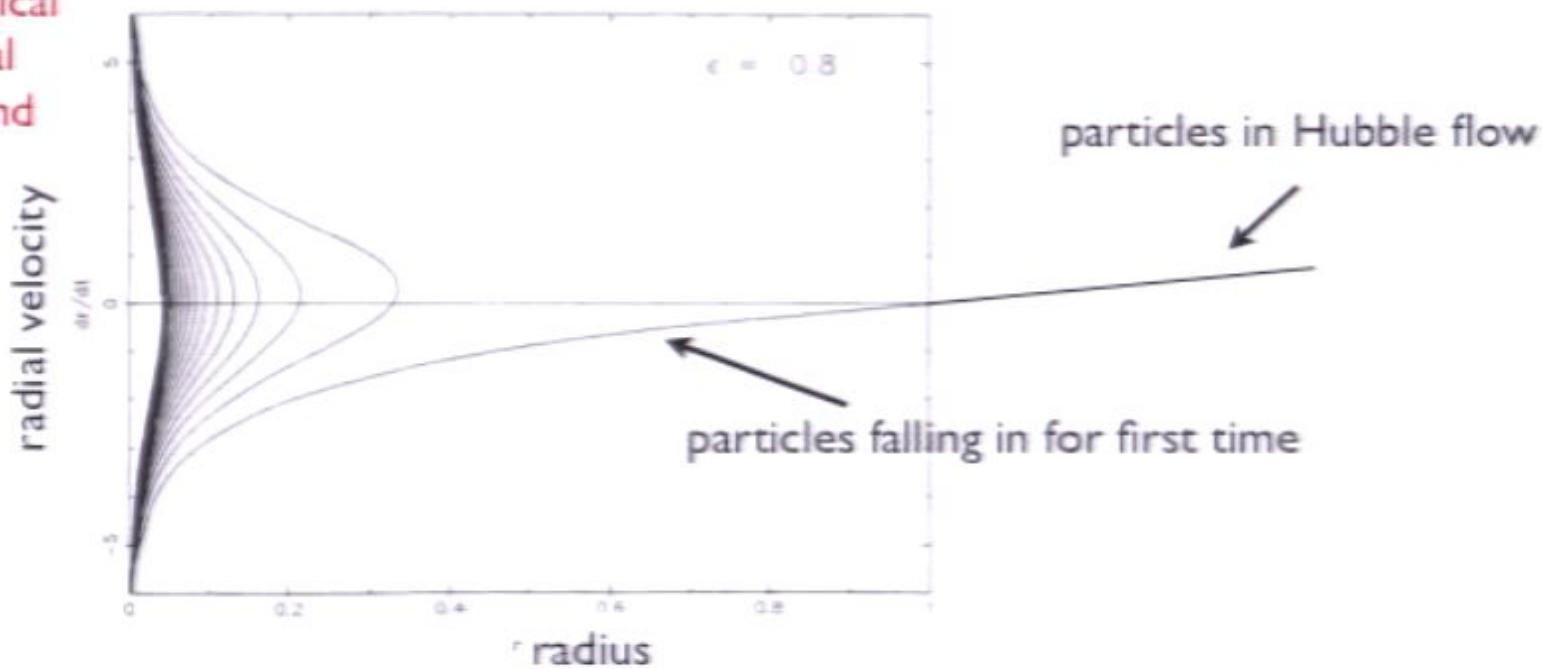
Fillmore & Goldreich, 1984 and Bertschinger, 1985

Note that this solution

- (a) satisfies Liouville's theorem (i.e., phase space DF remains a cold, 3D sheet in 6D phase space)
- (b) solution is time-reversible

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spherical collapse of a scale-free perturbation in an expanding Universe

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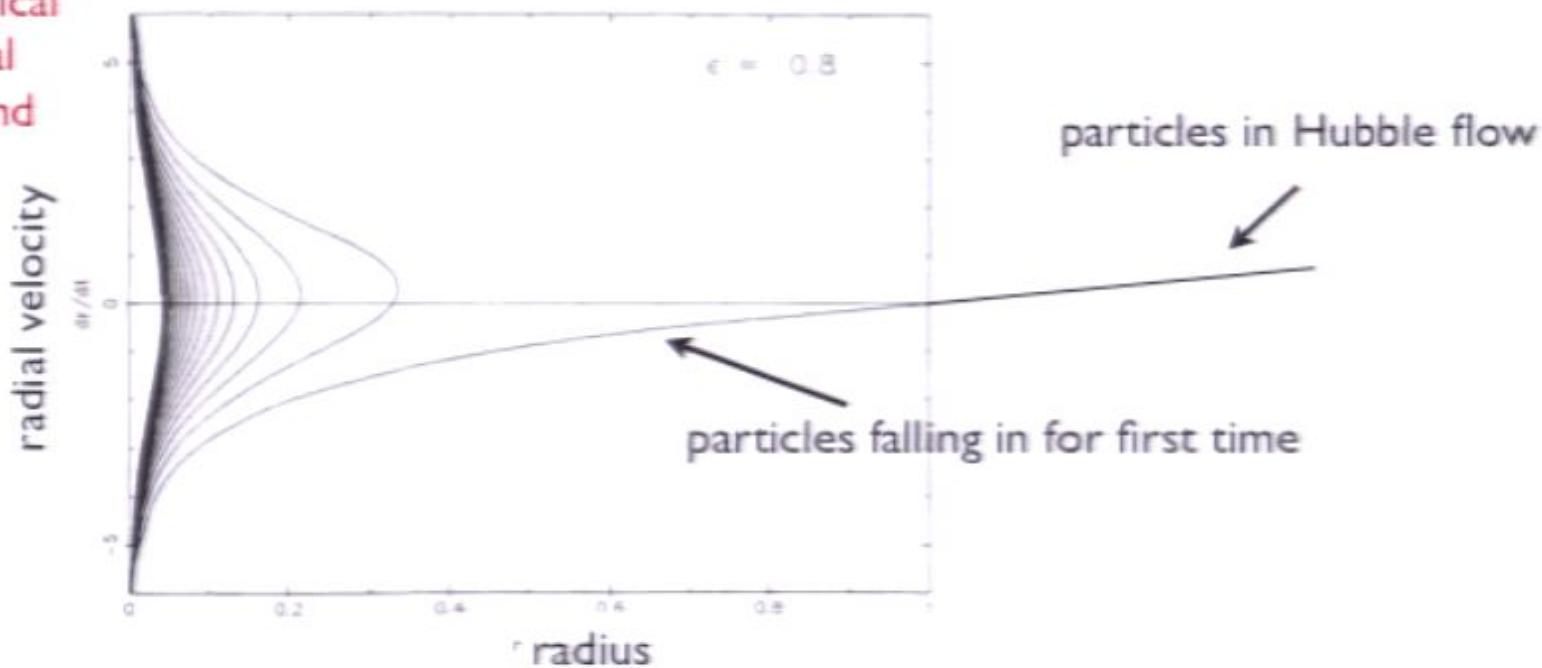
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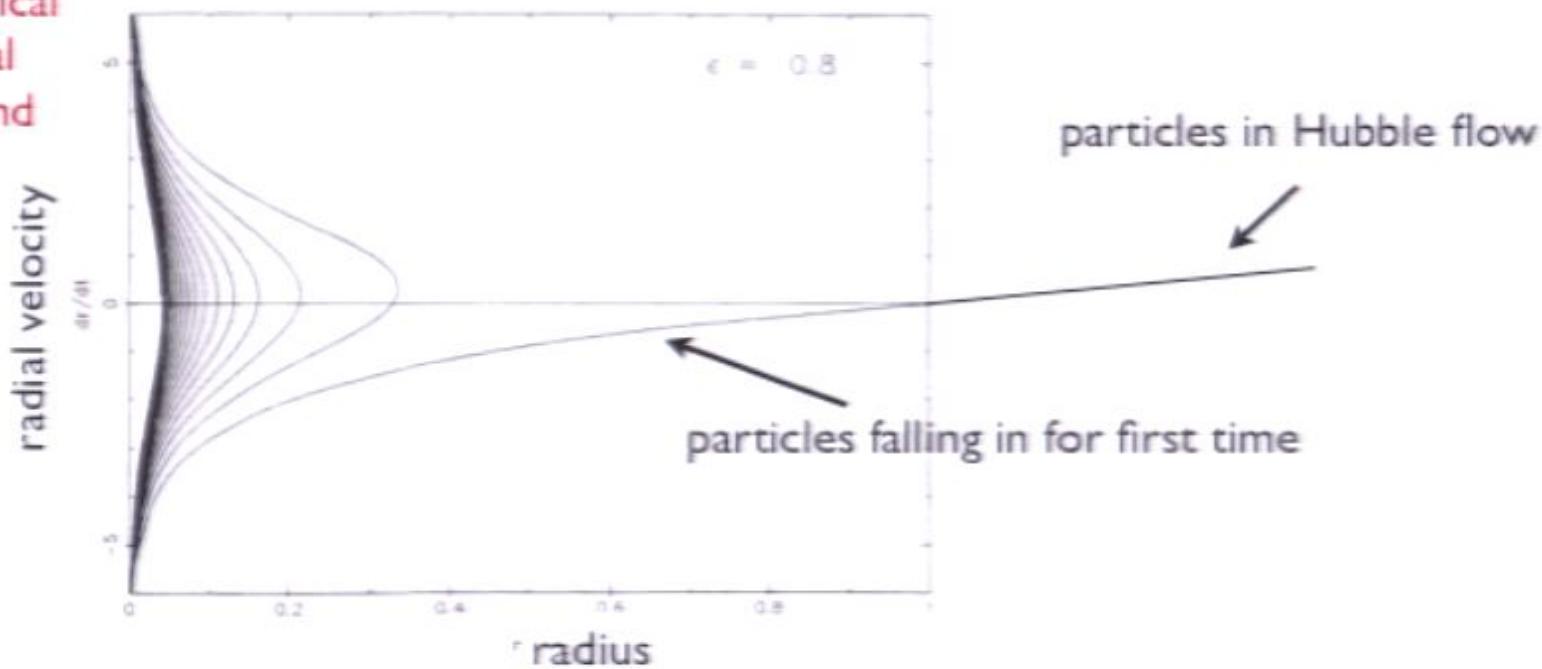
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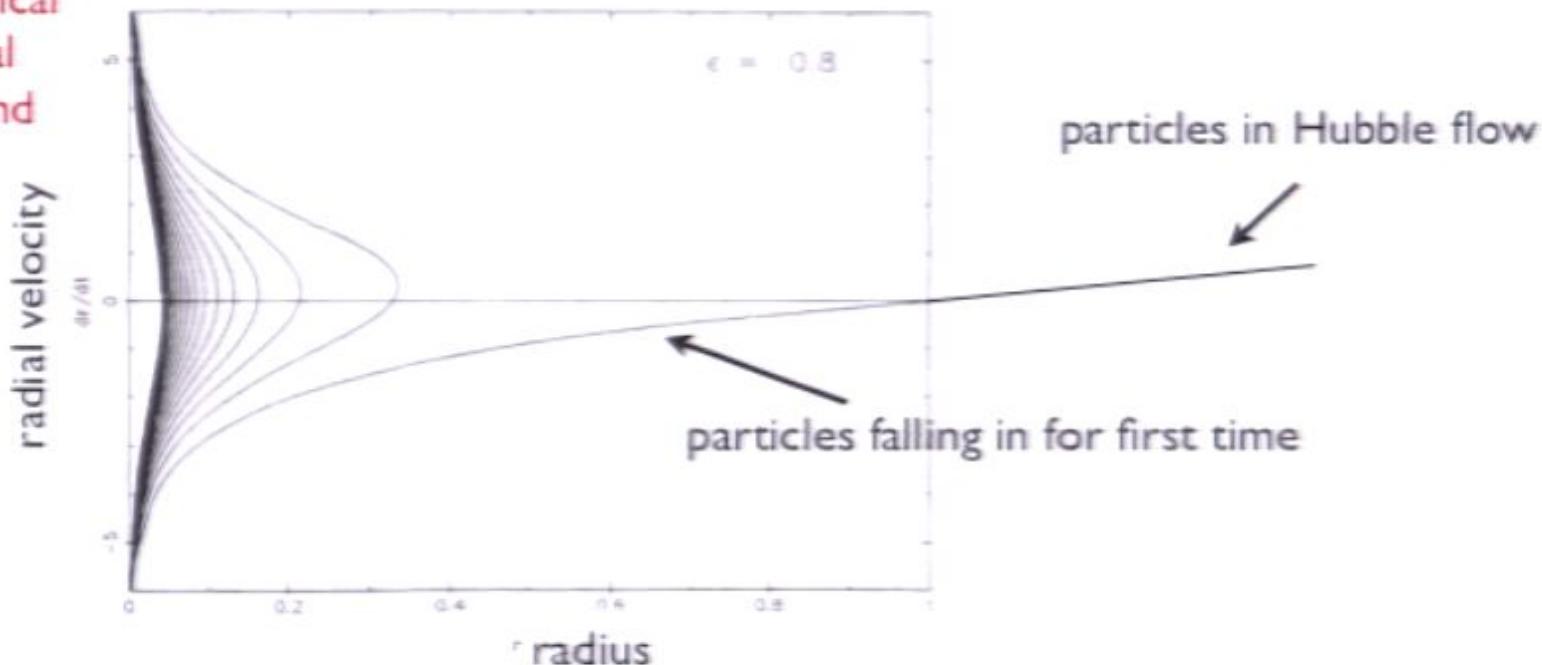
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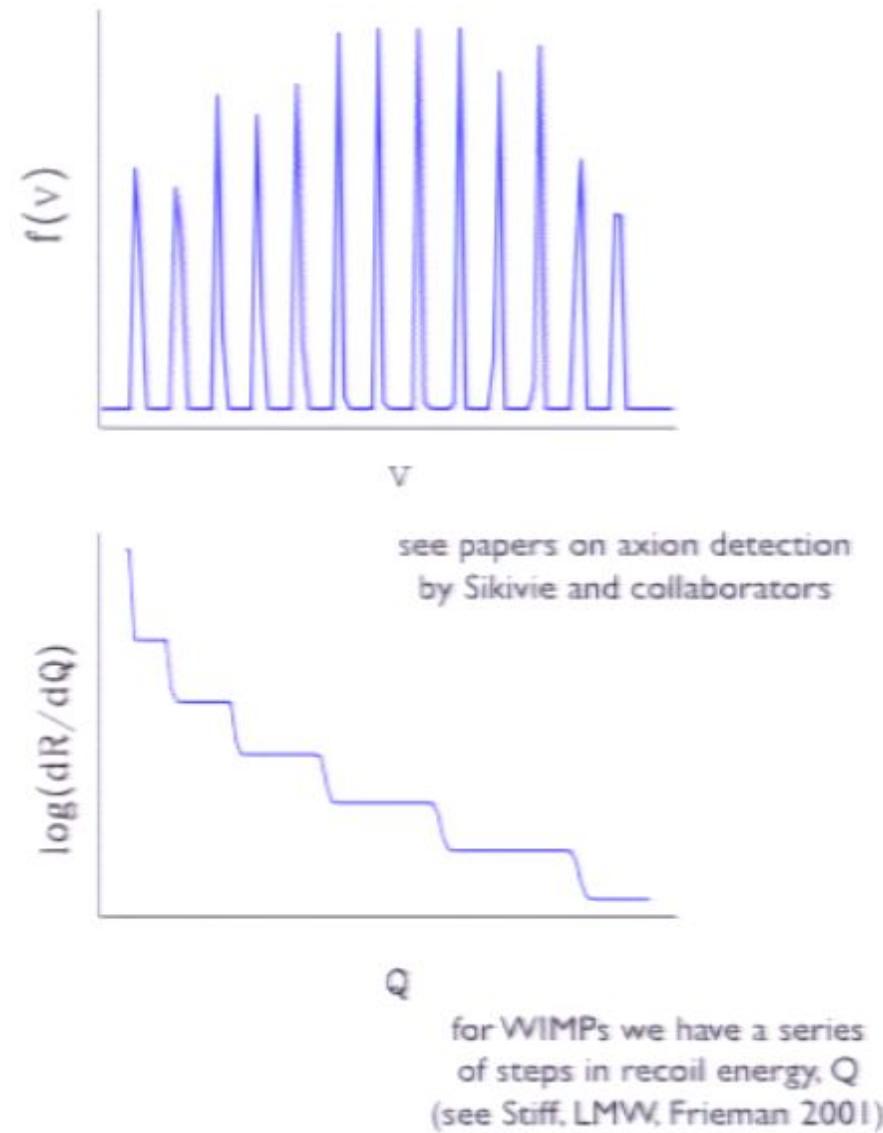
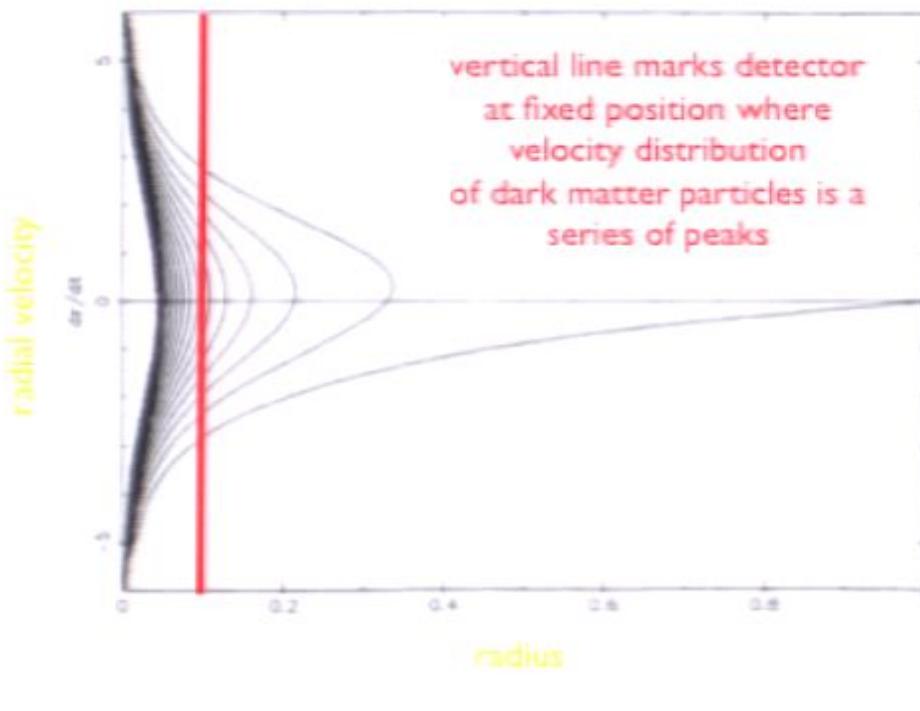
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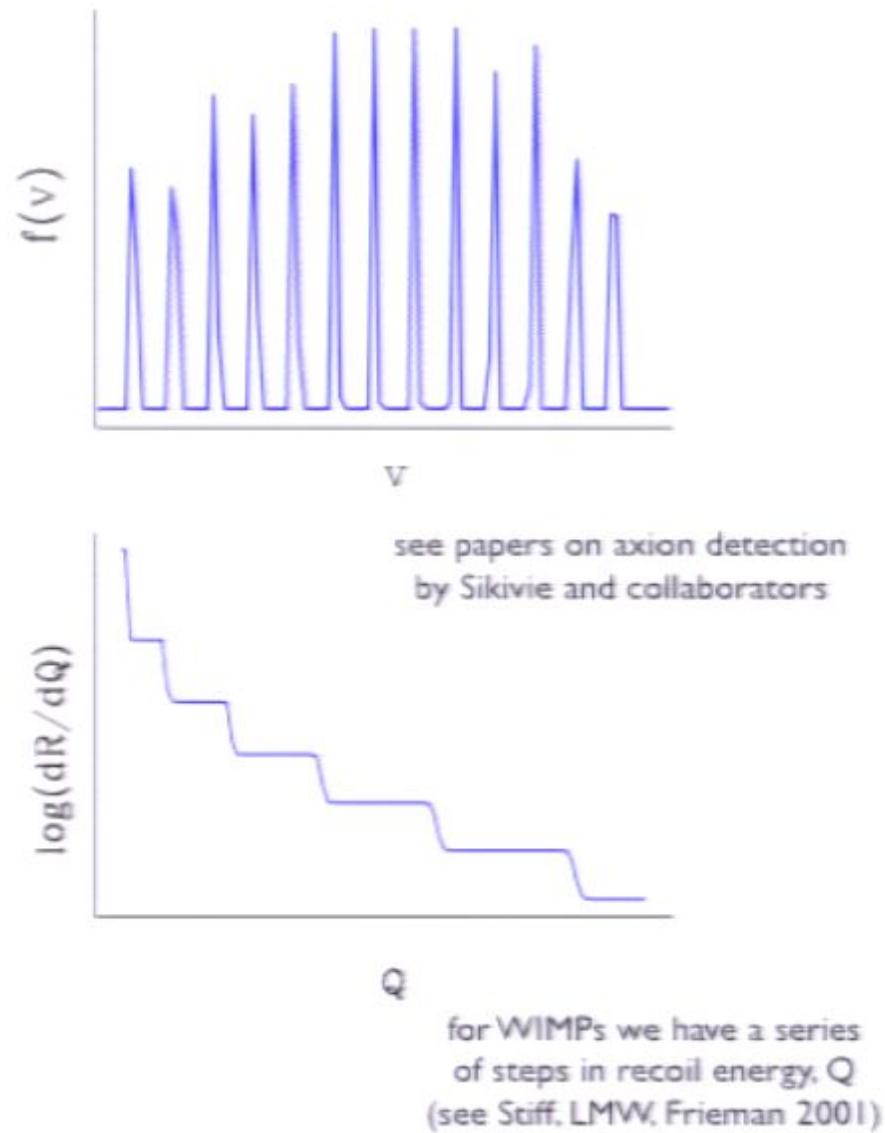
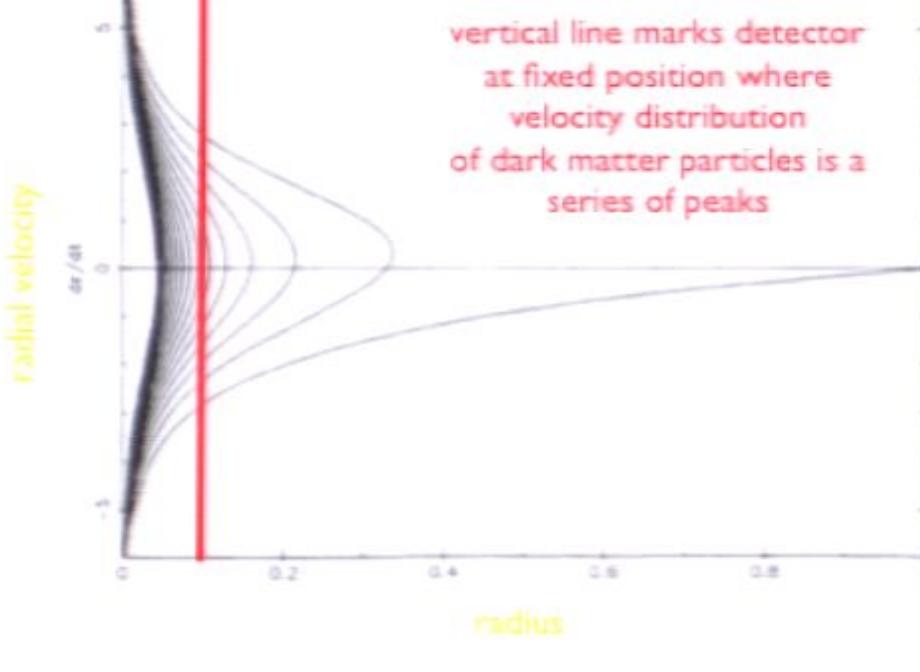


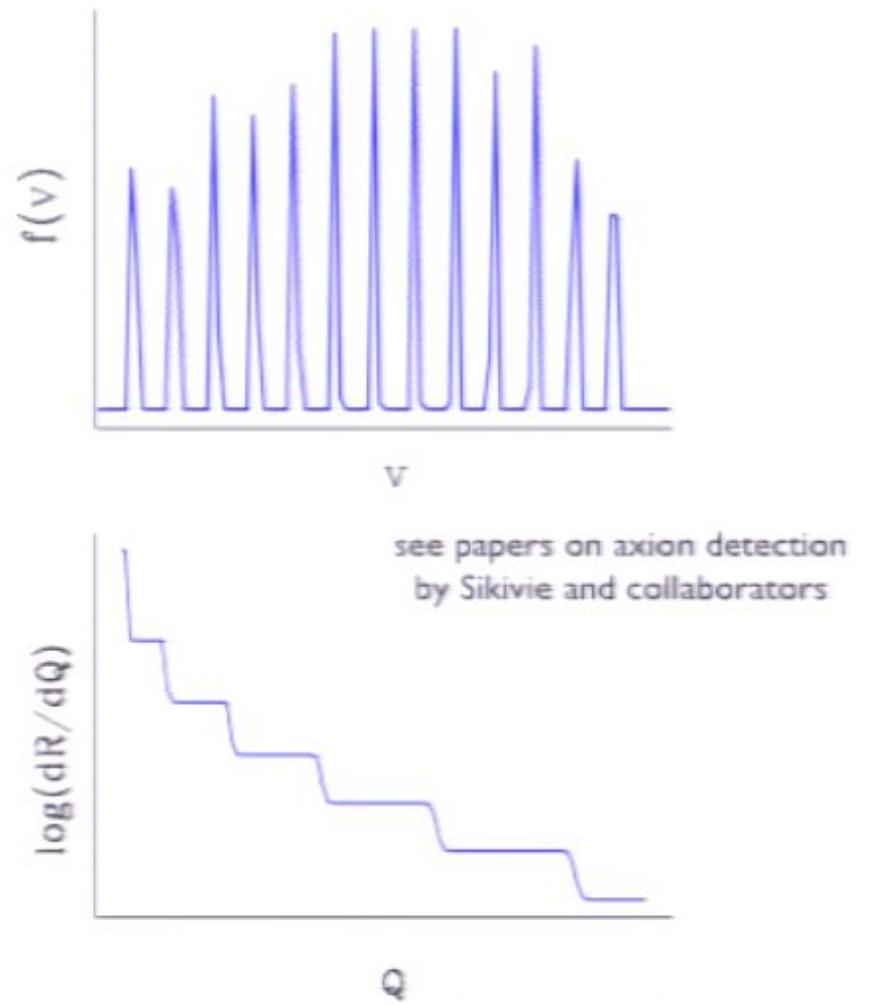
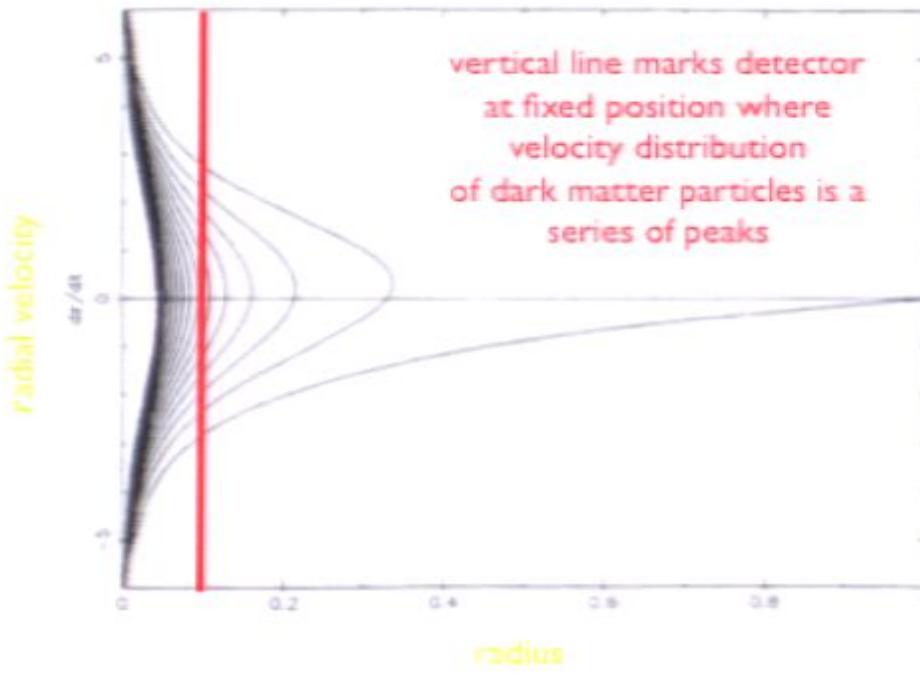
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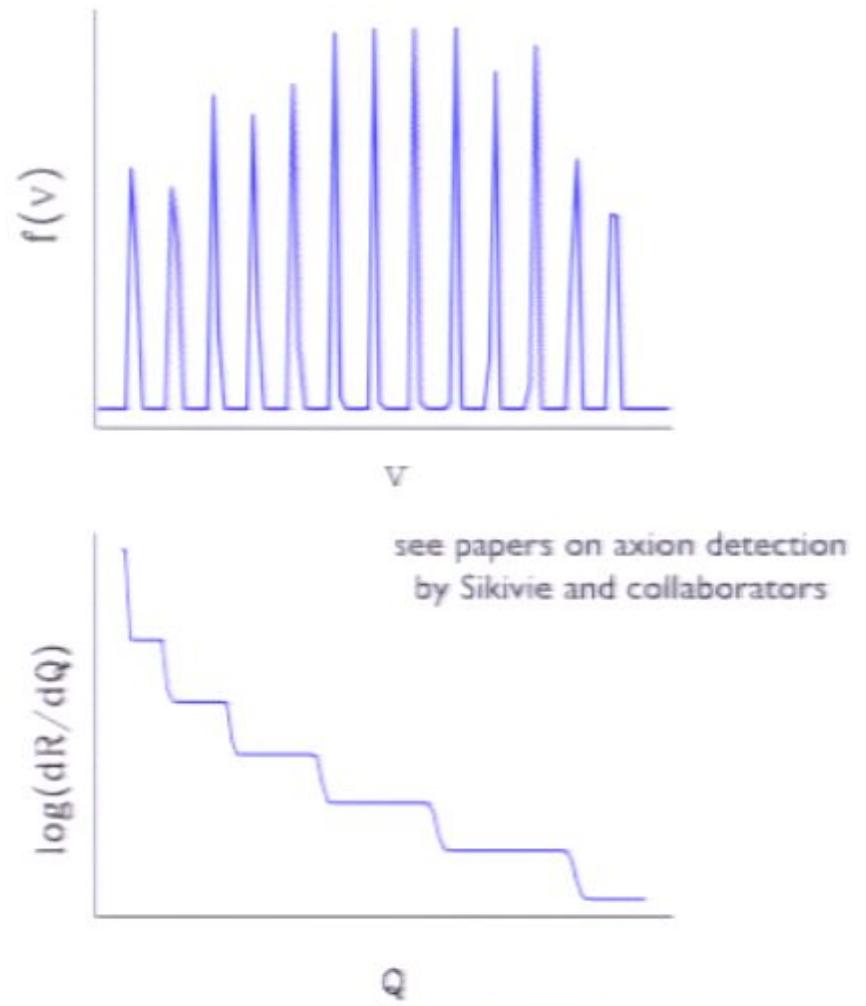
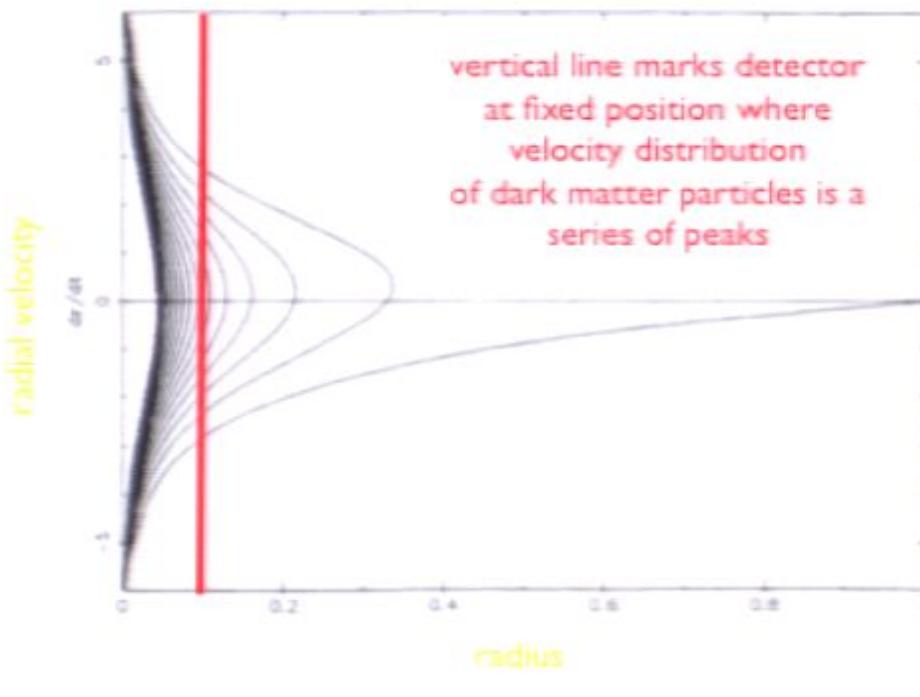
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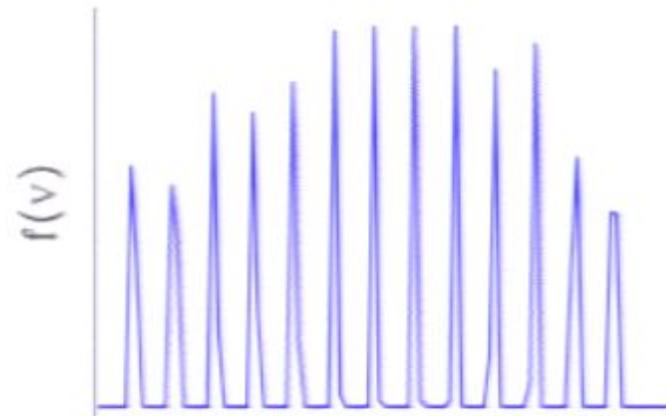
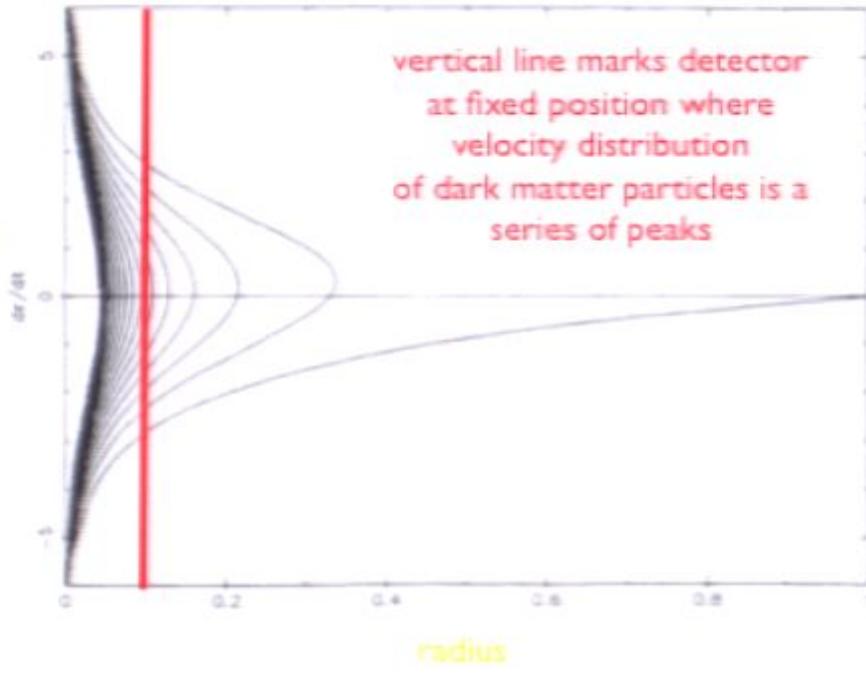


for WIMPs we have a series of steps in recoil energy, Q (see Stiff, LMW, Frieman 2001)

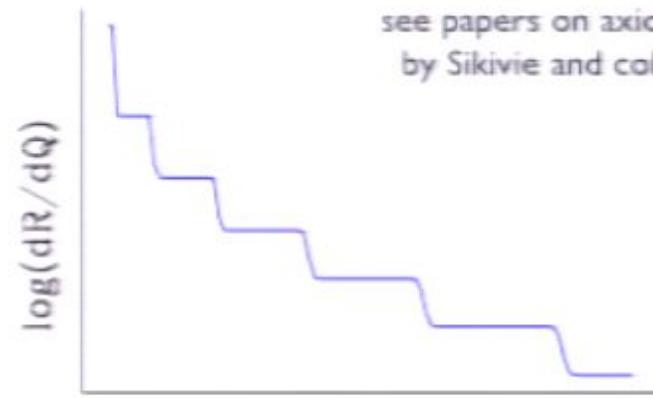


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radial velocity

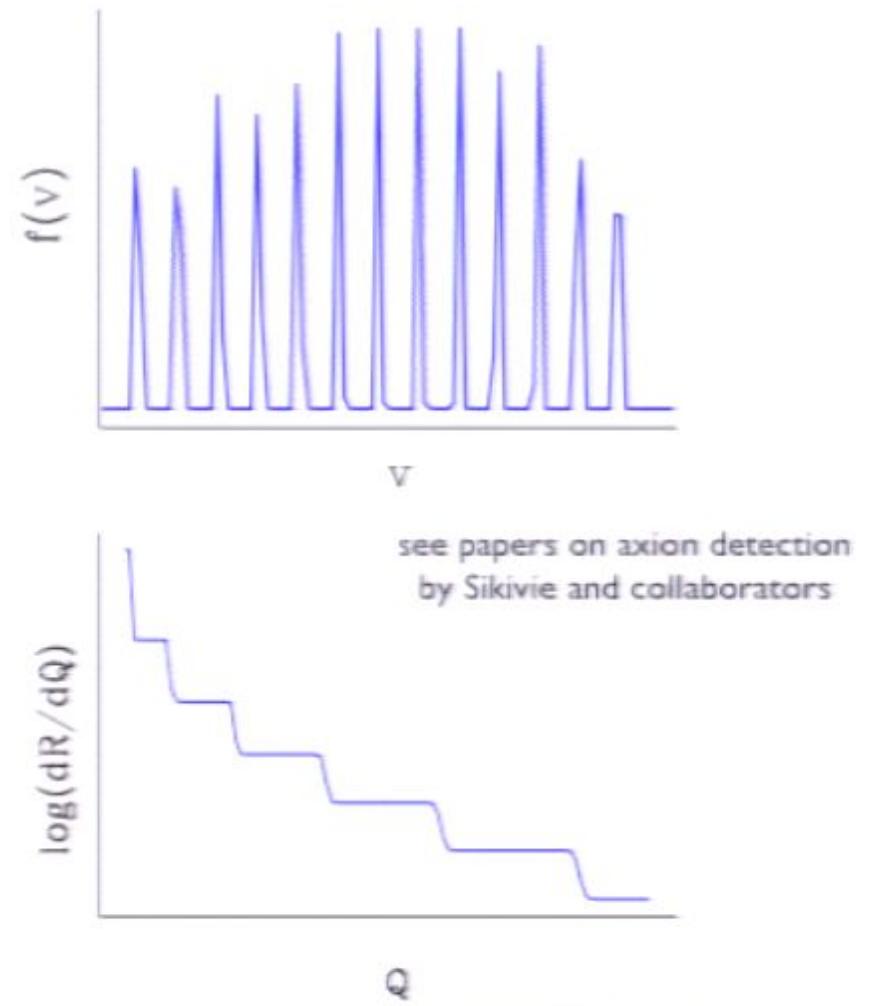
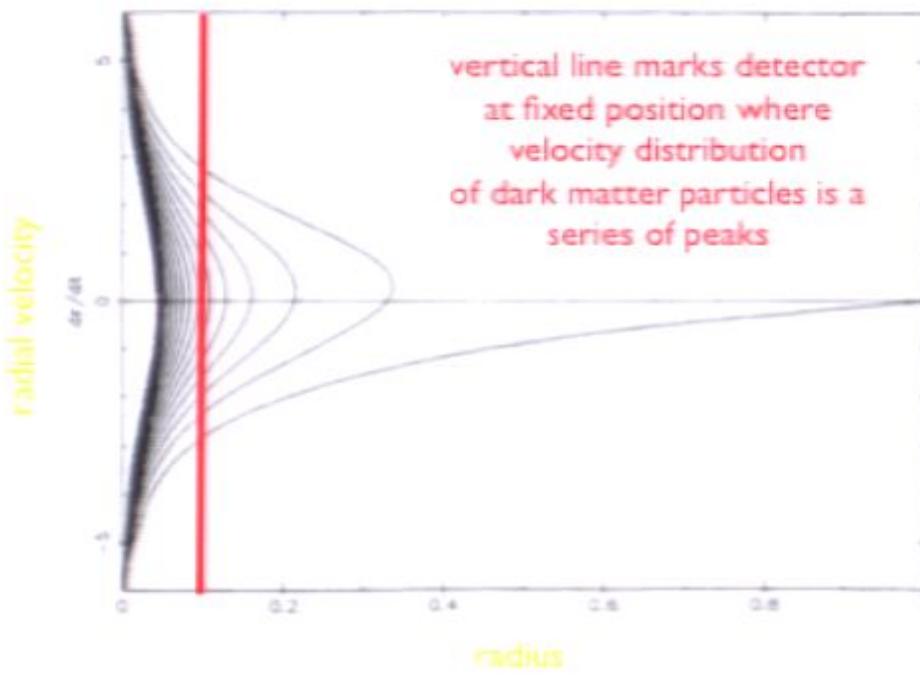


v



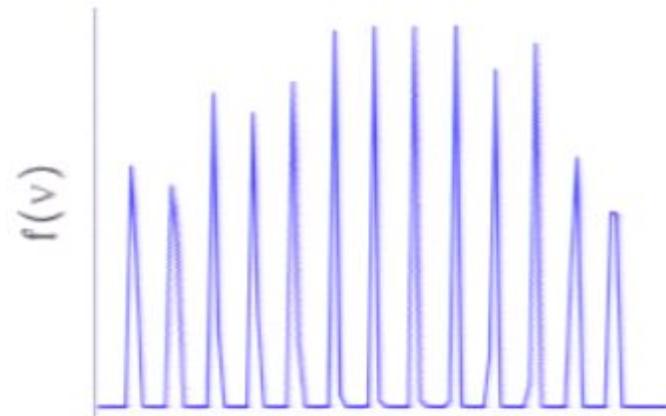
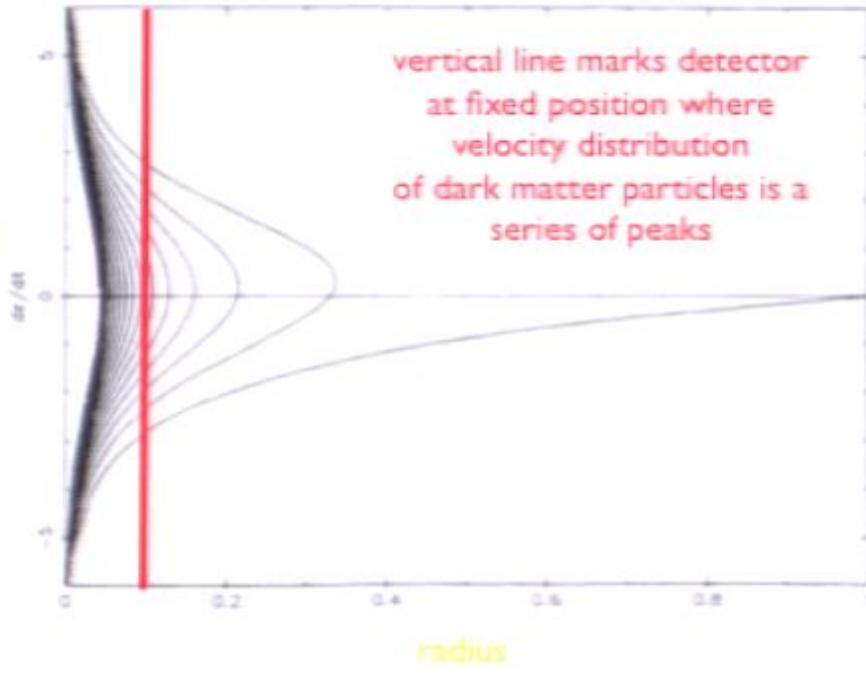
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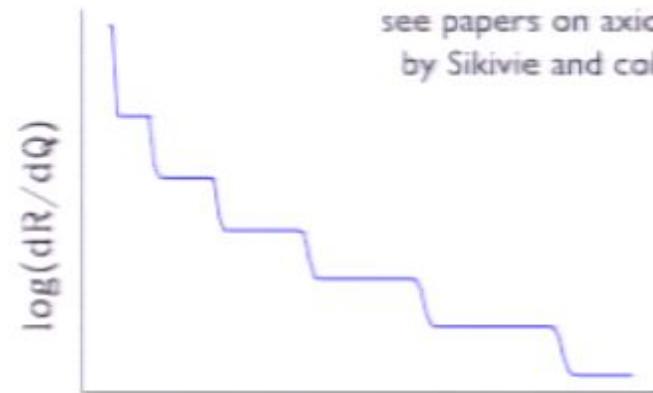


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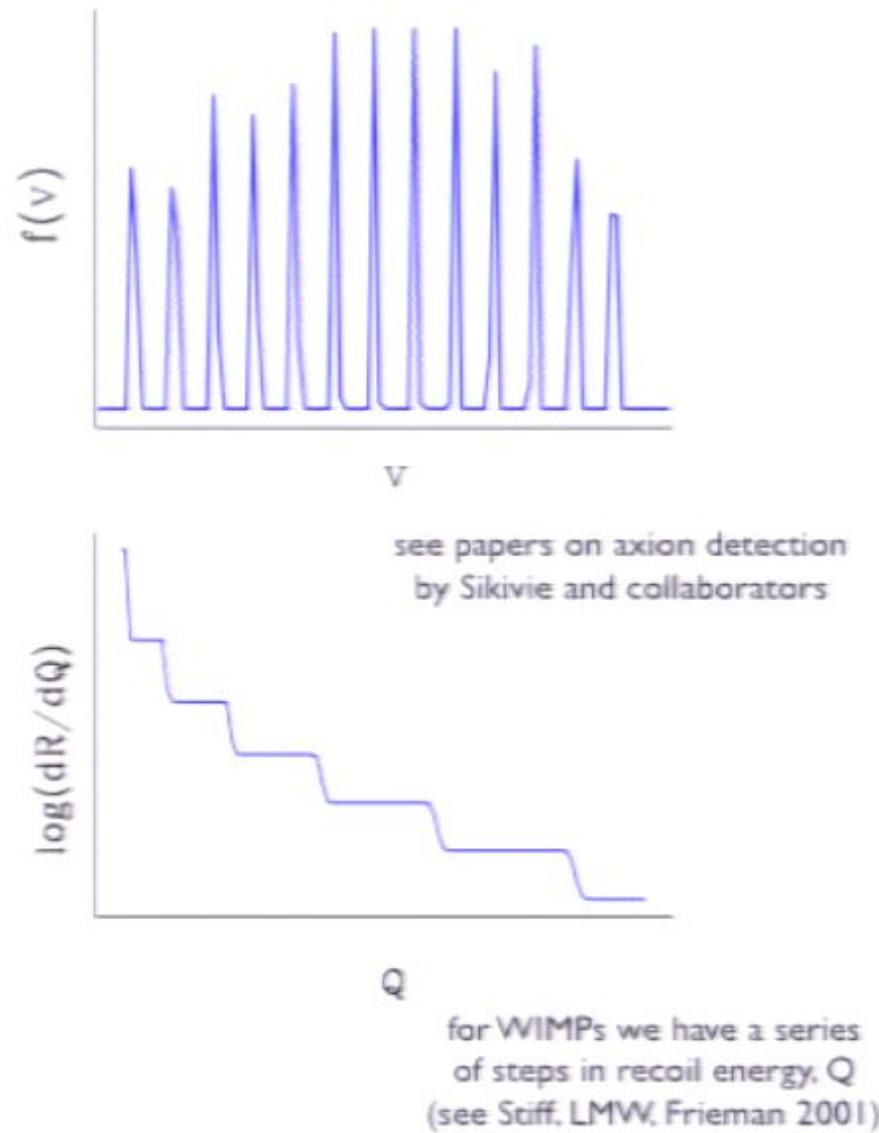
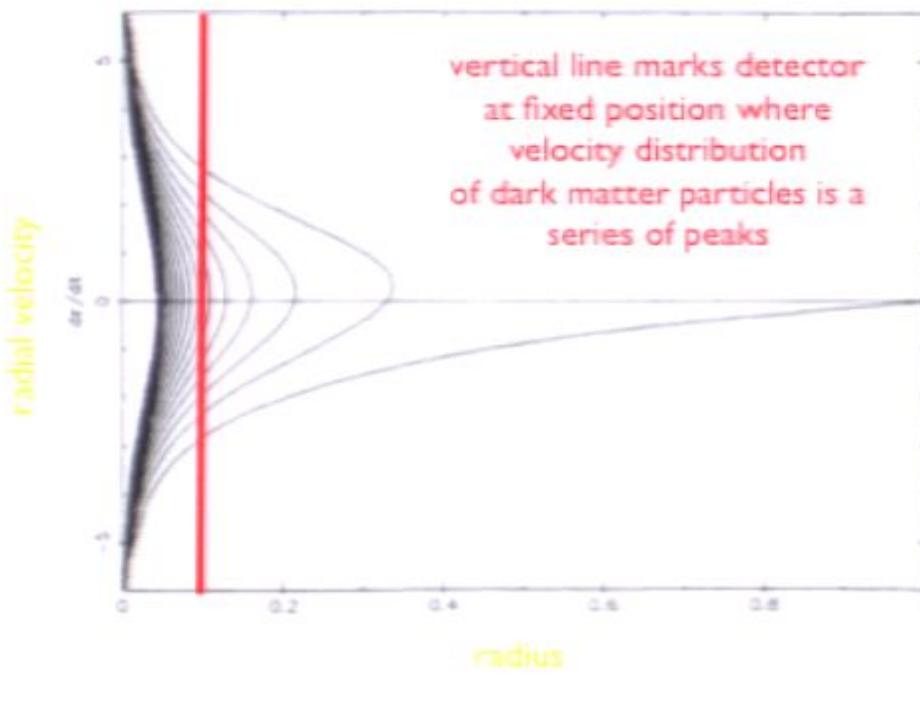


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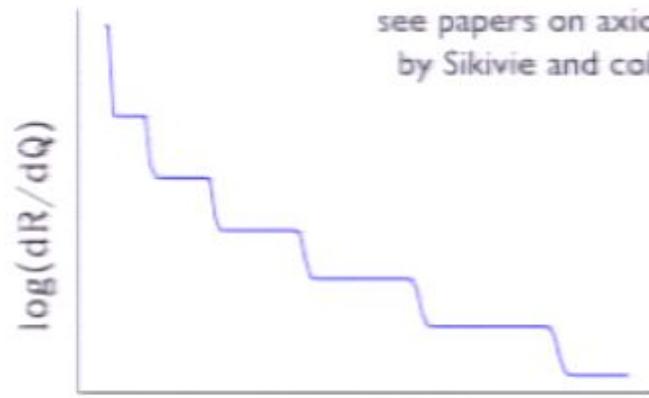
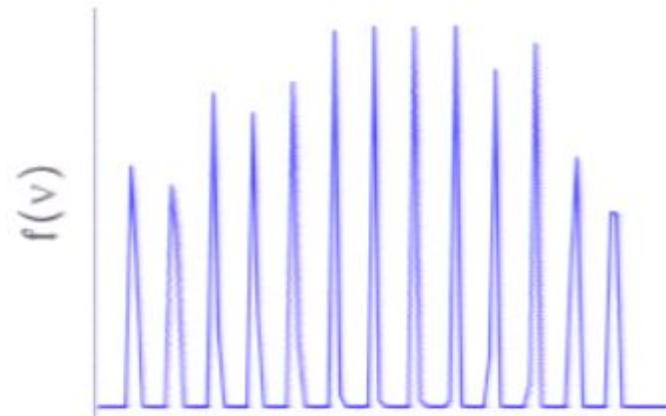
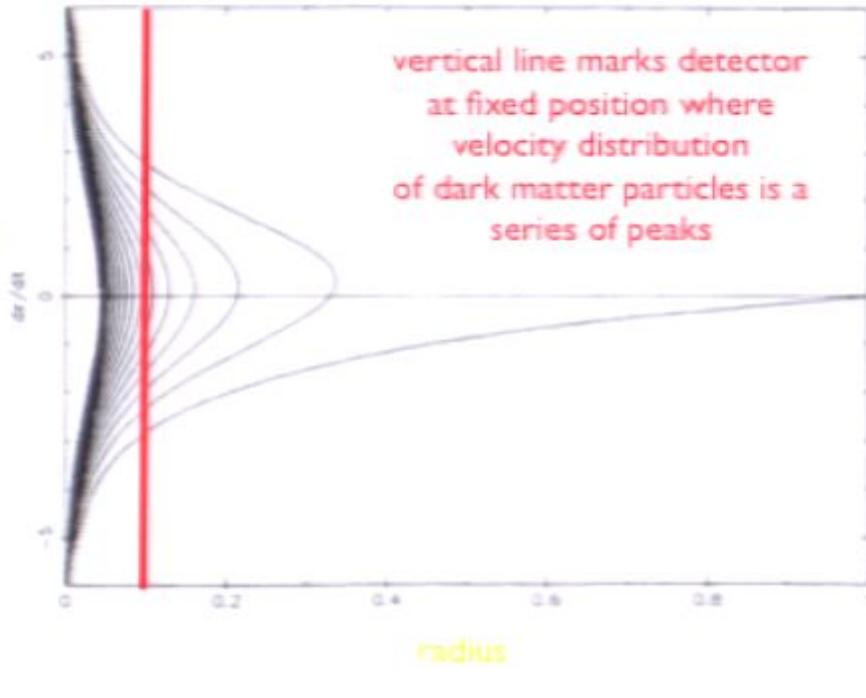


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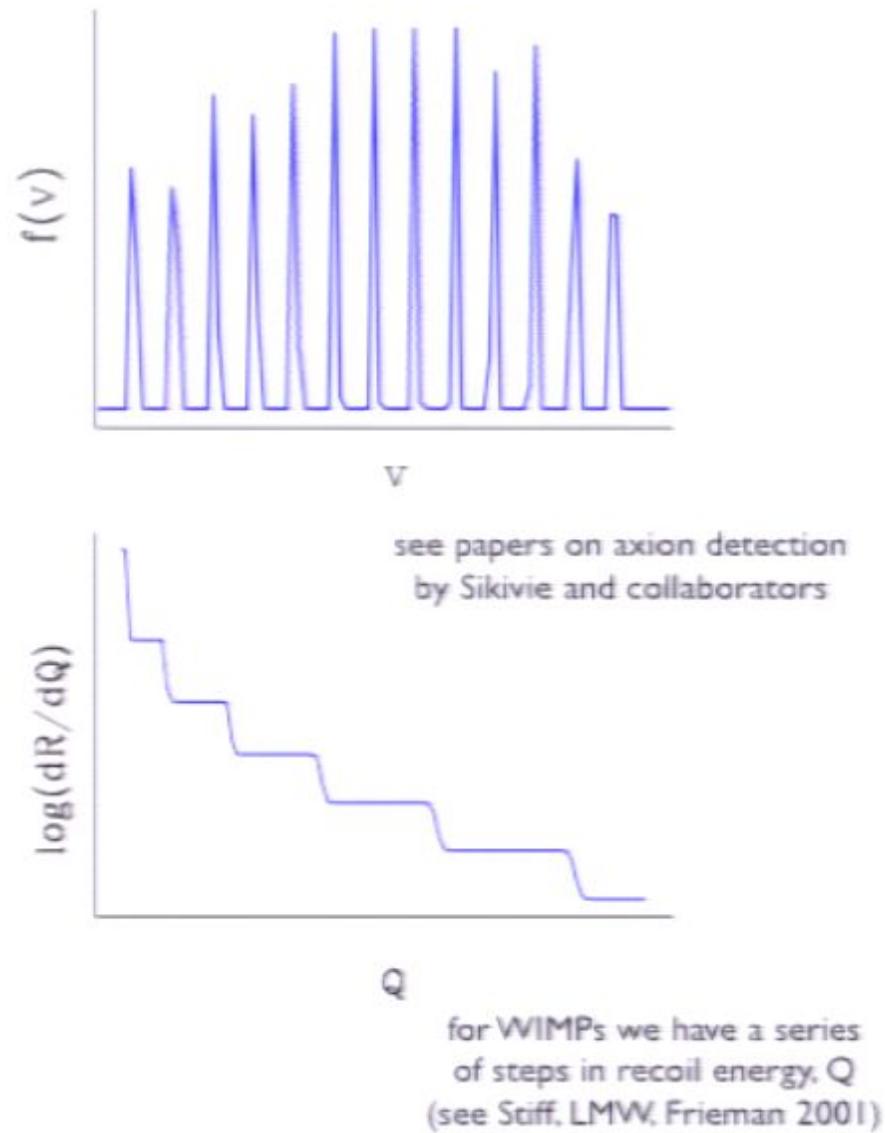
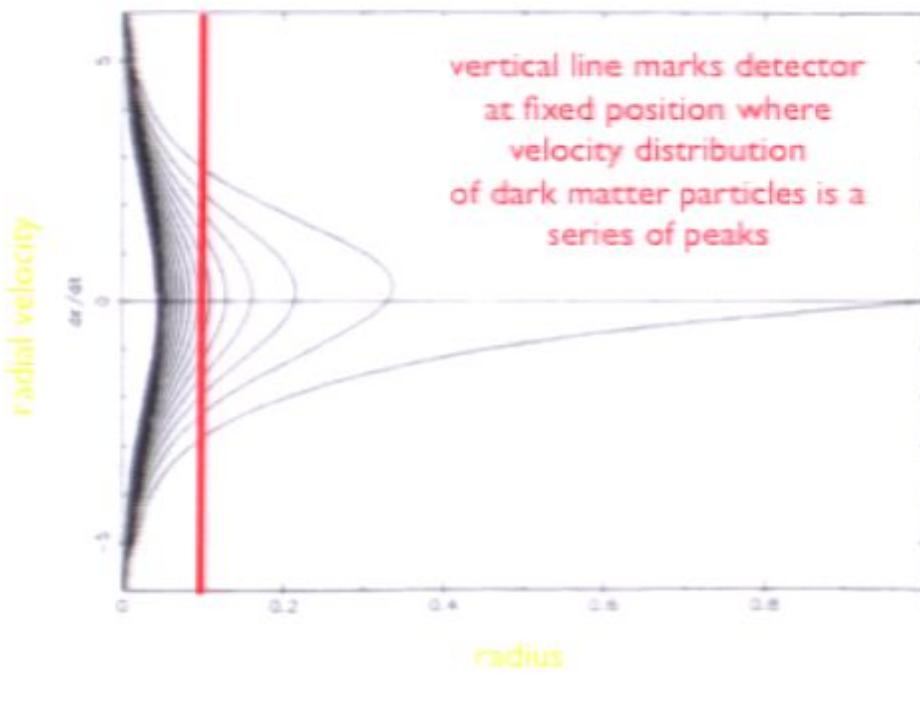


radial velocity

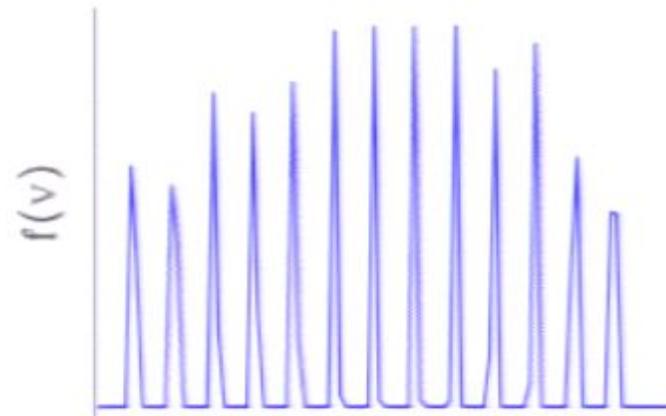
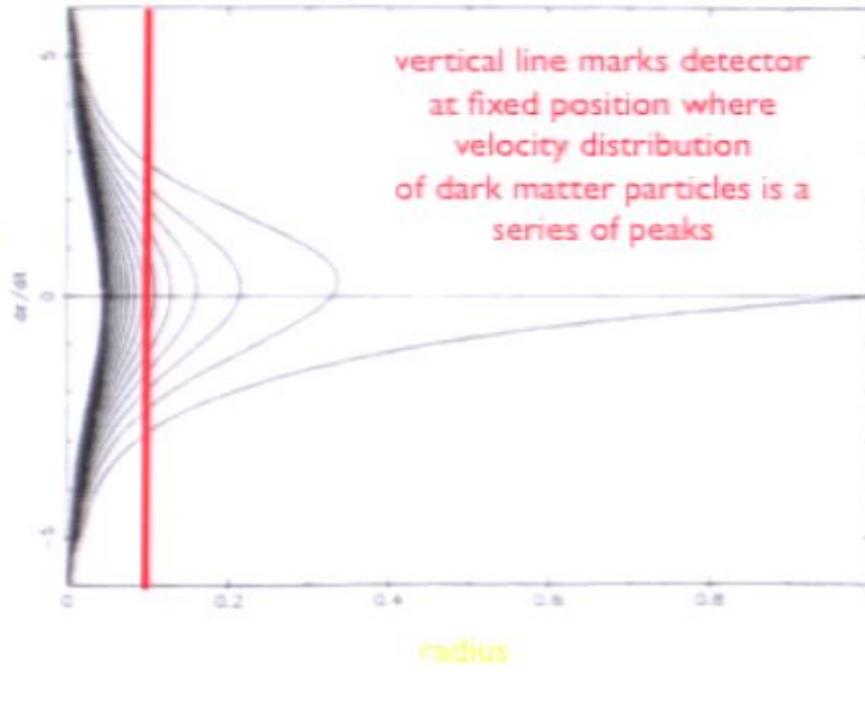


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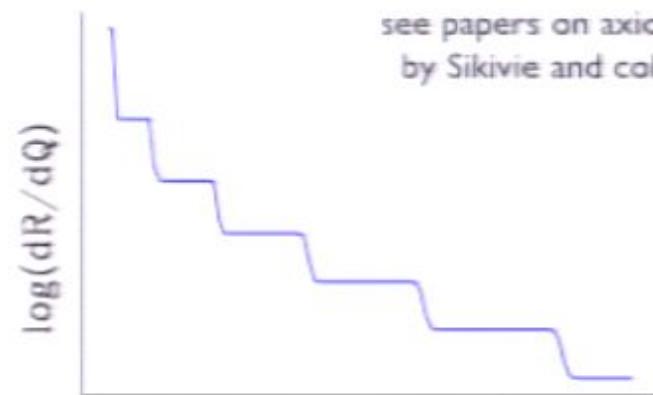
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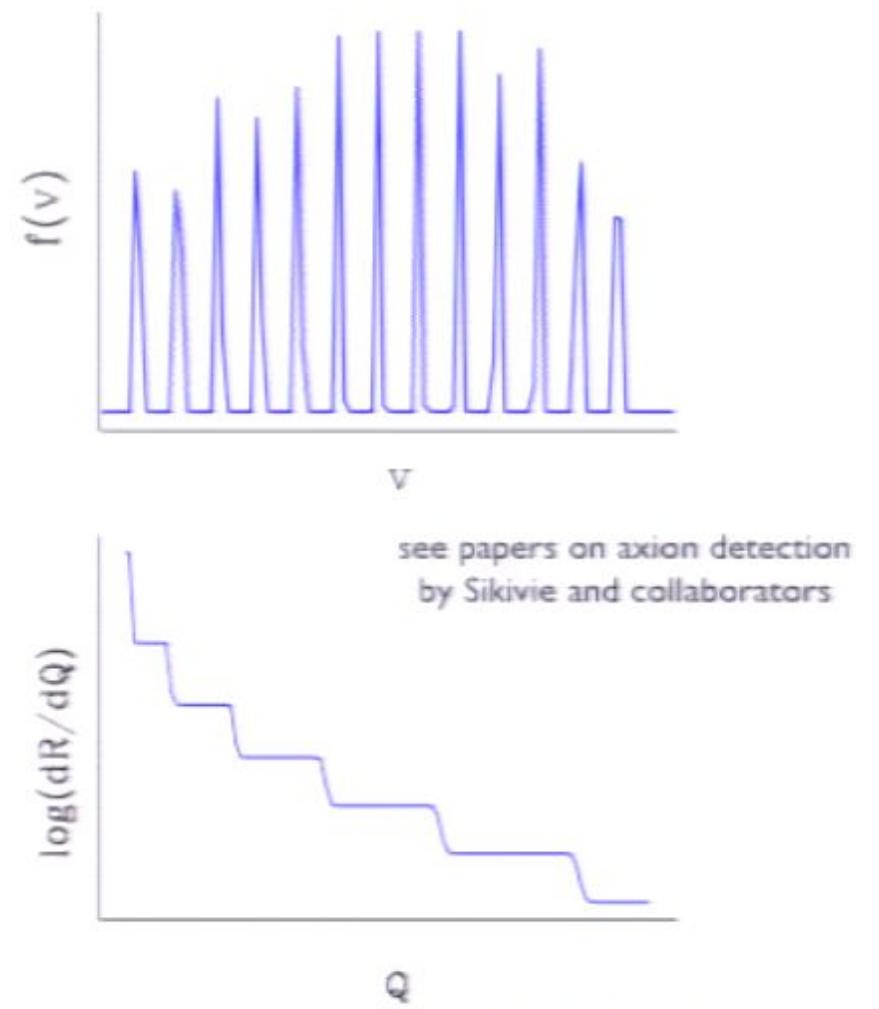
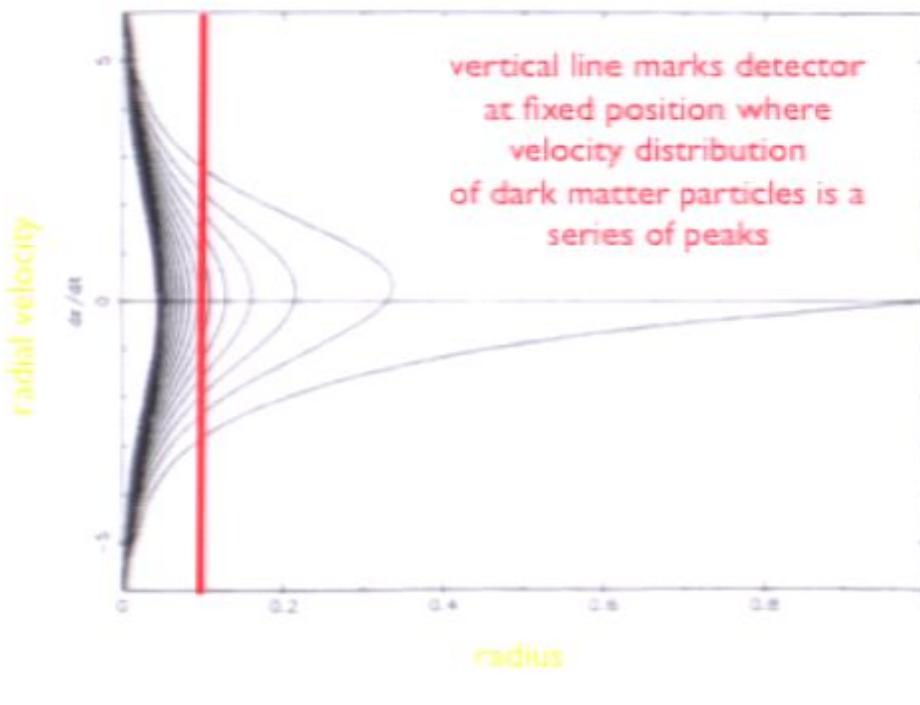


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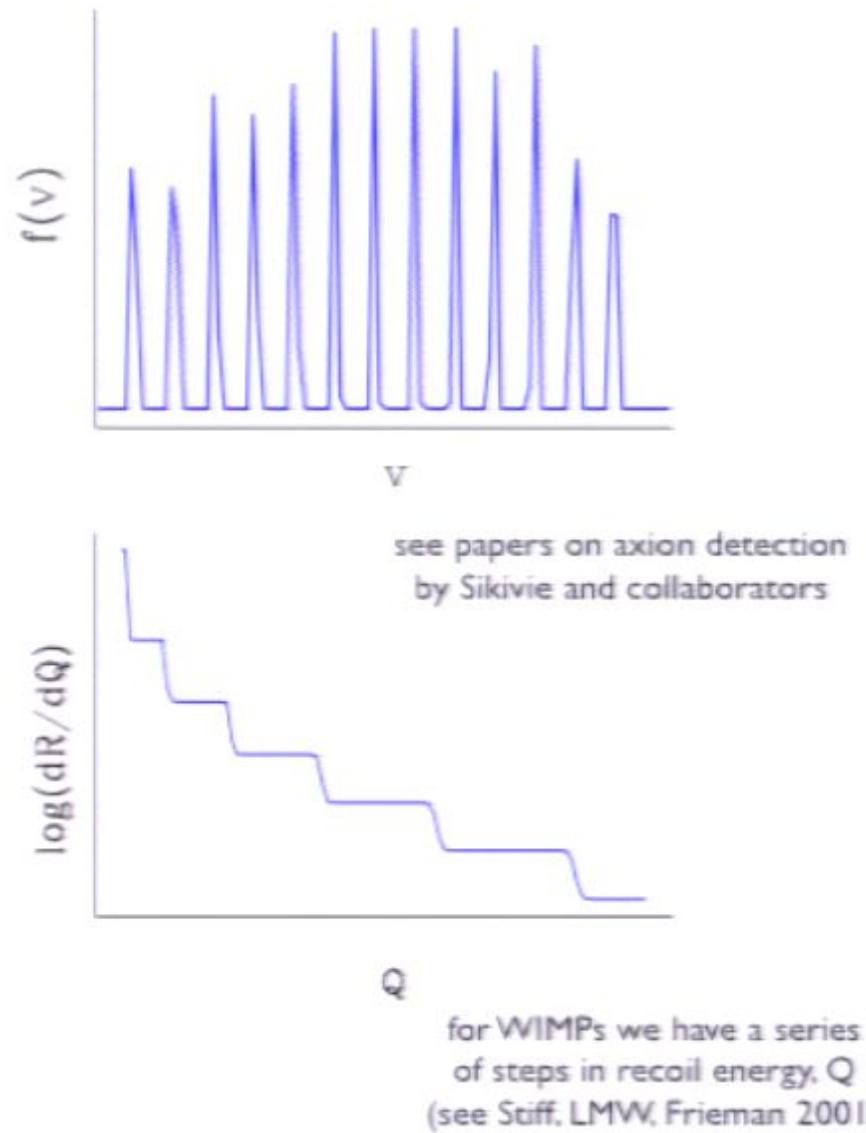
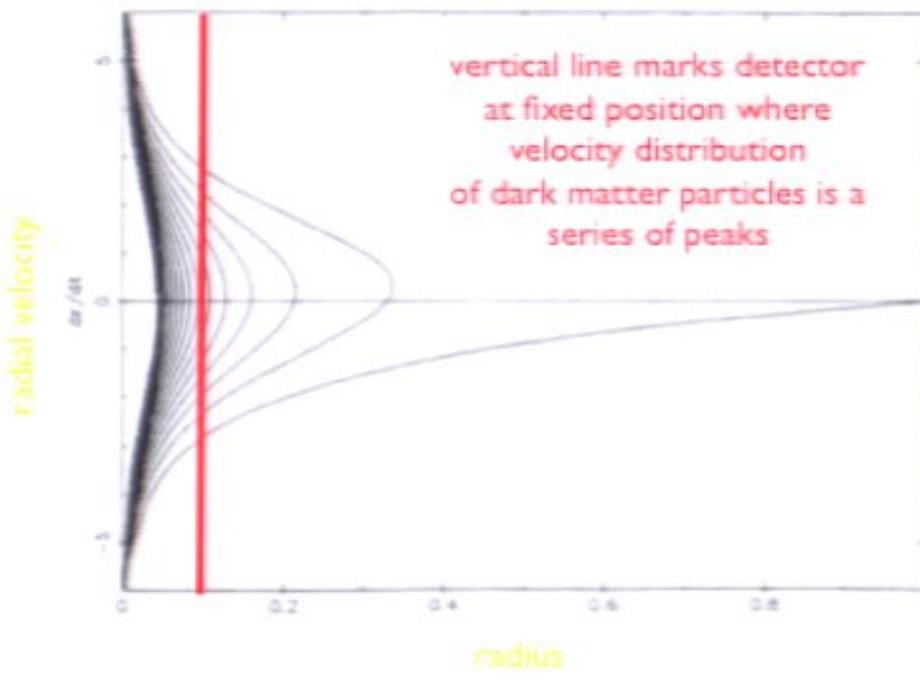


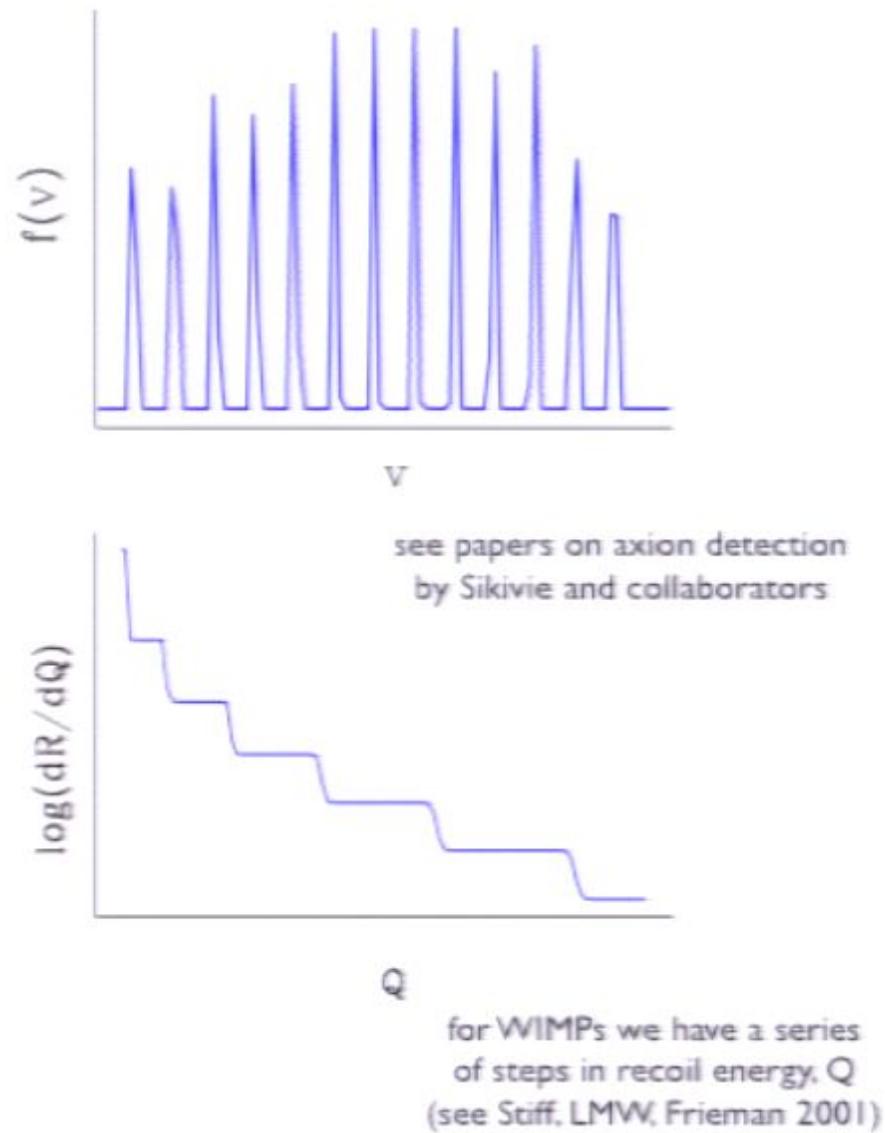
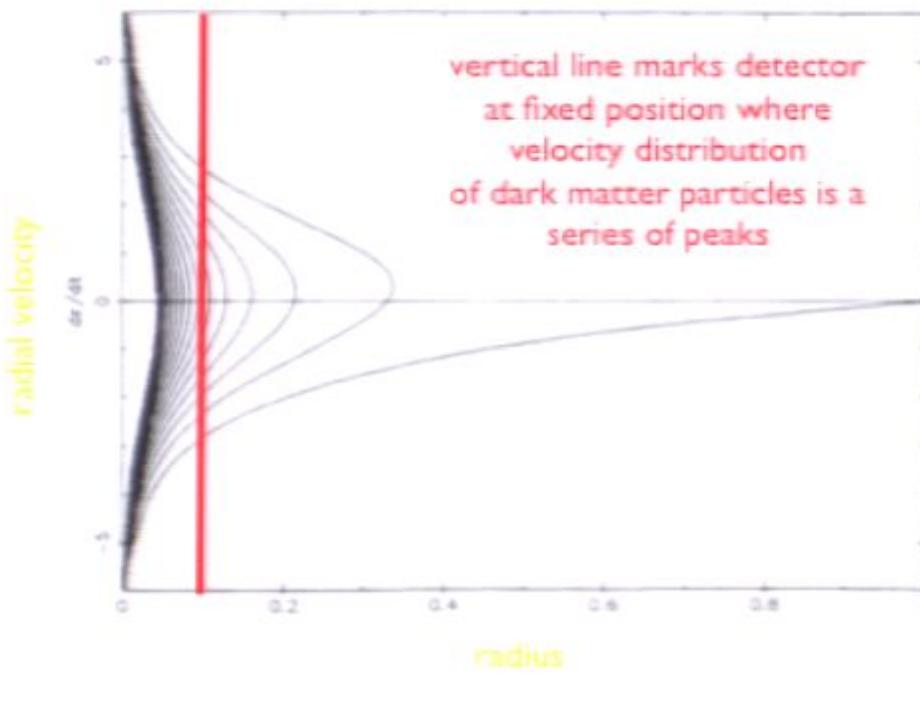
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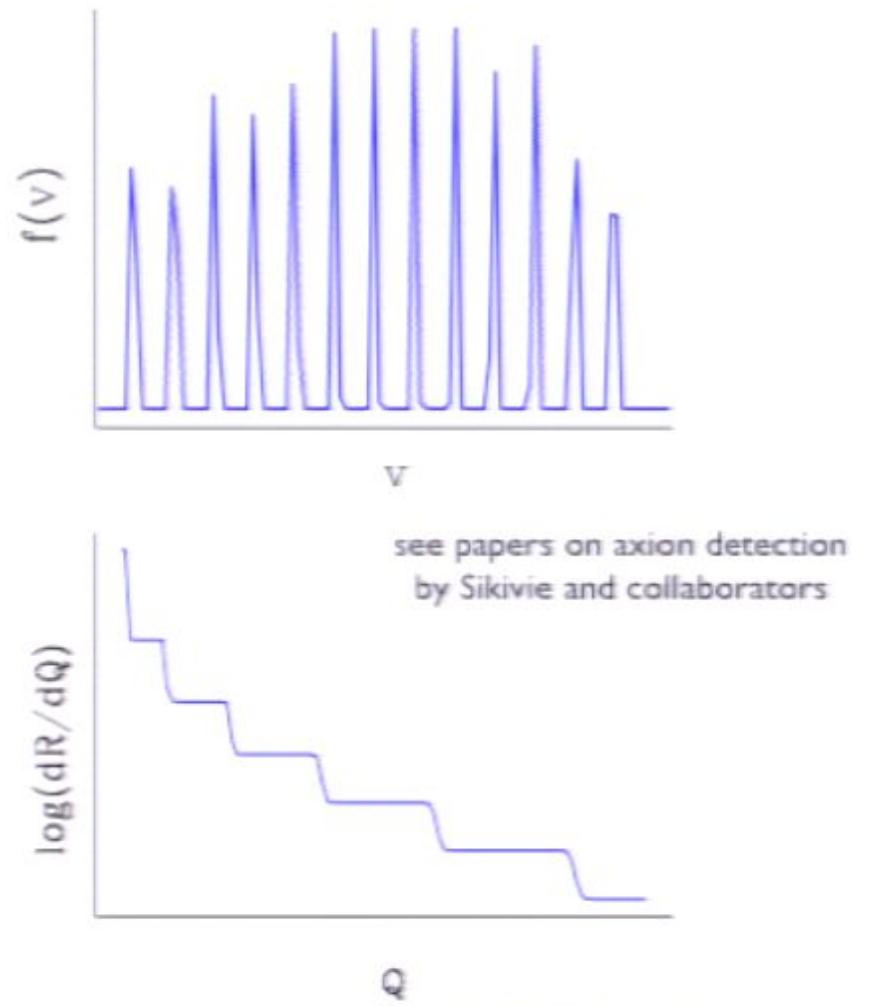
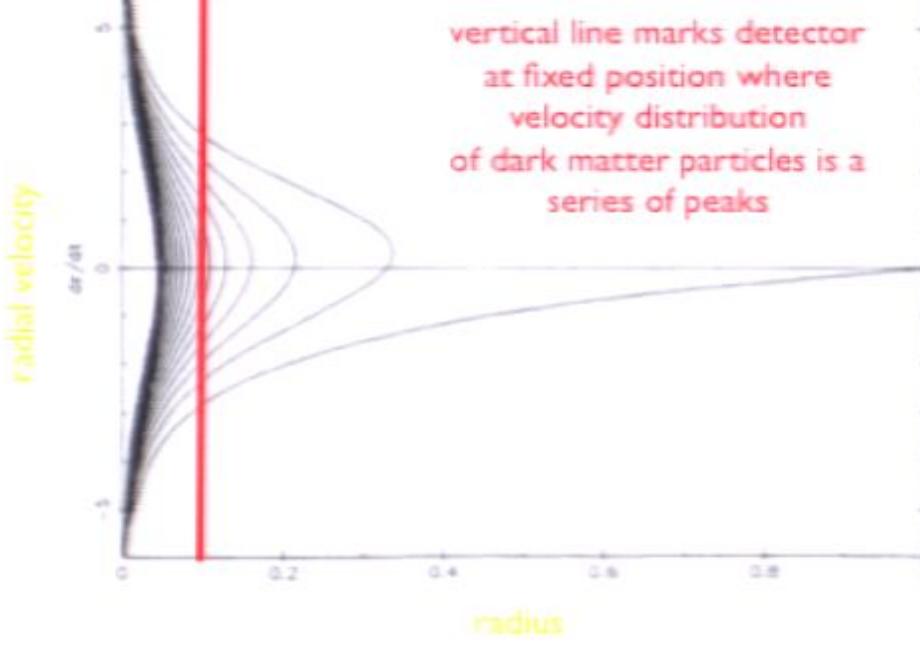
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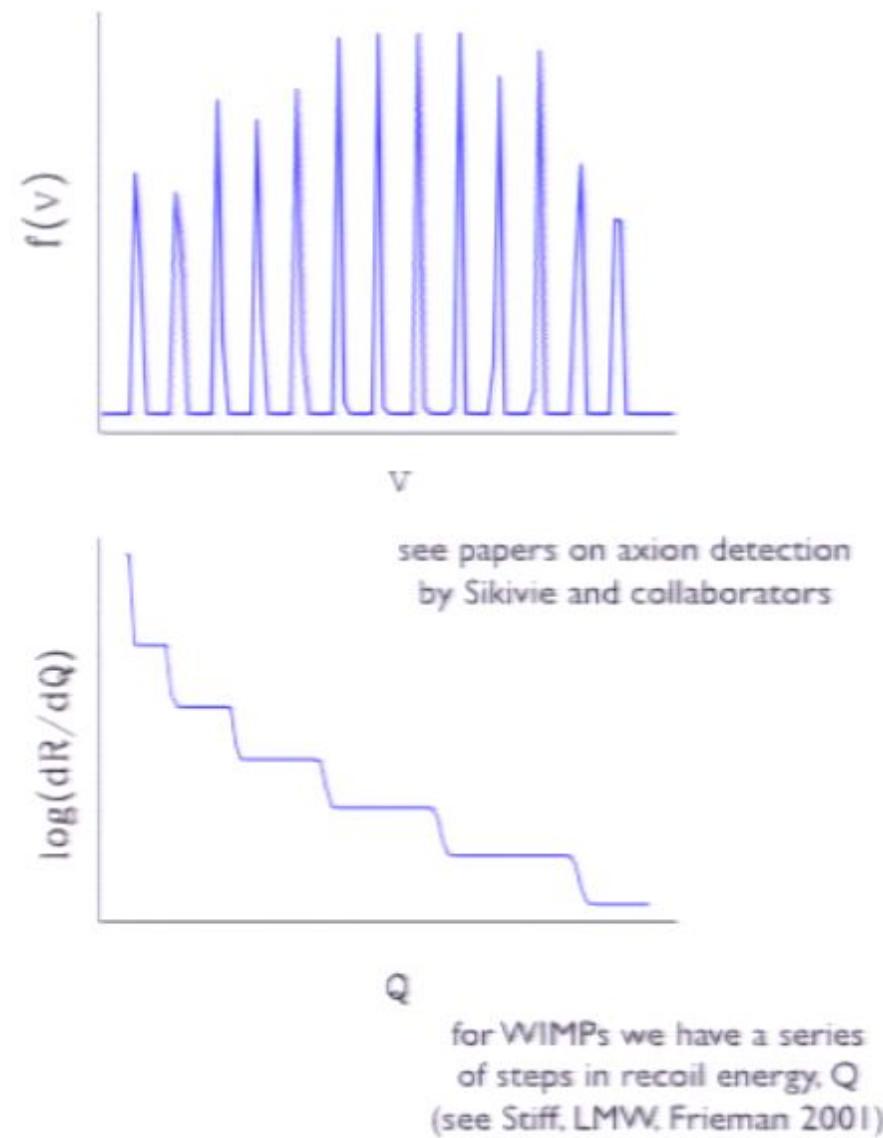
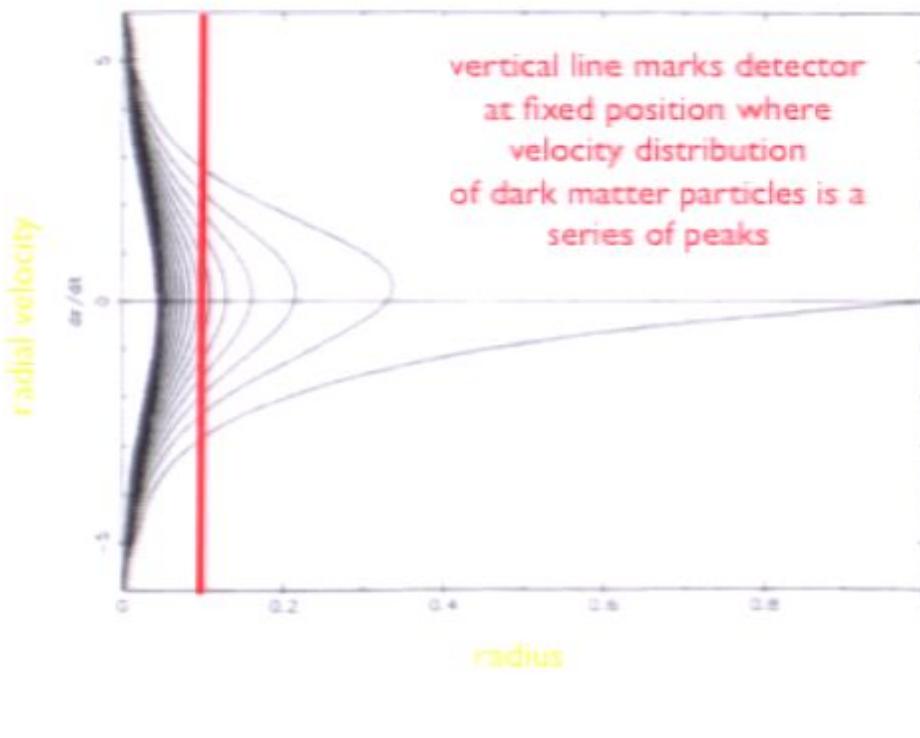
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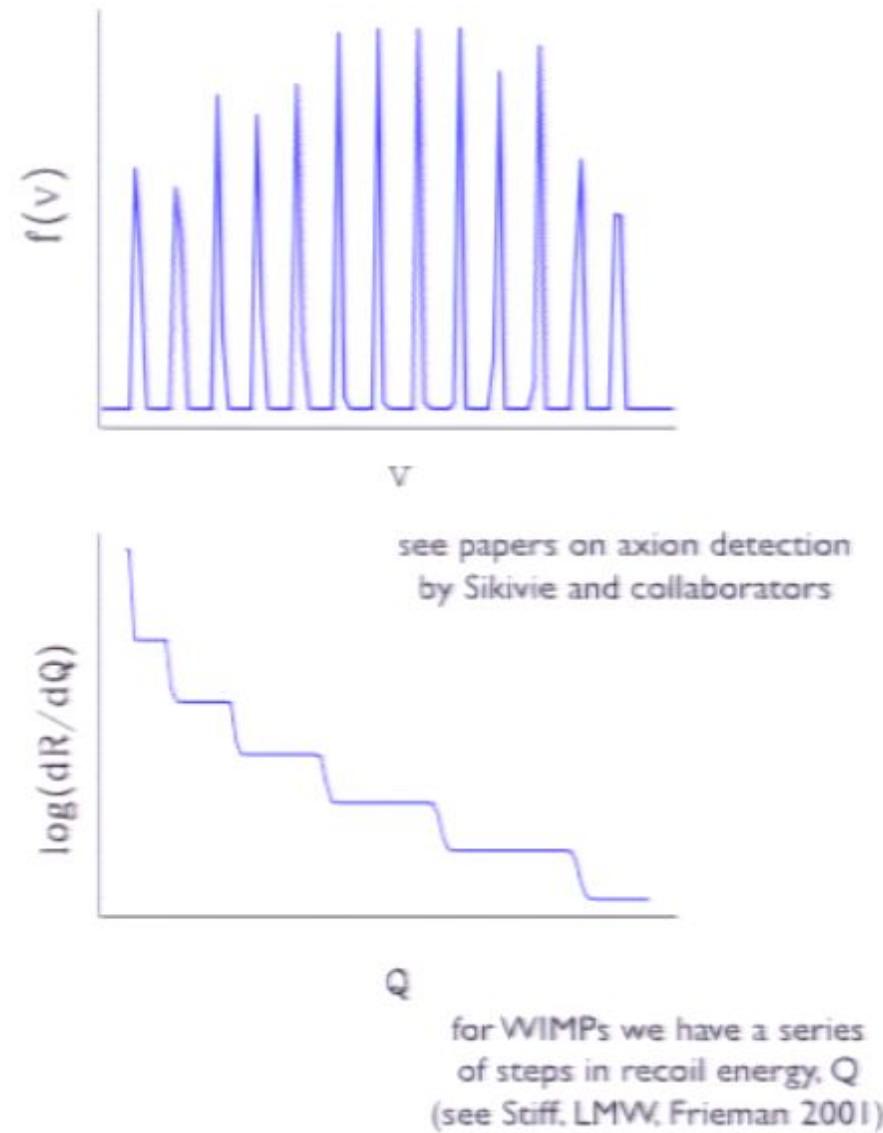
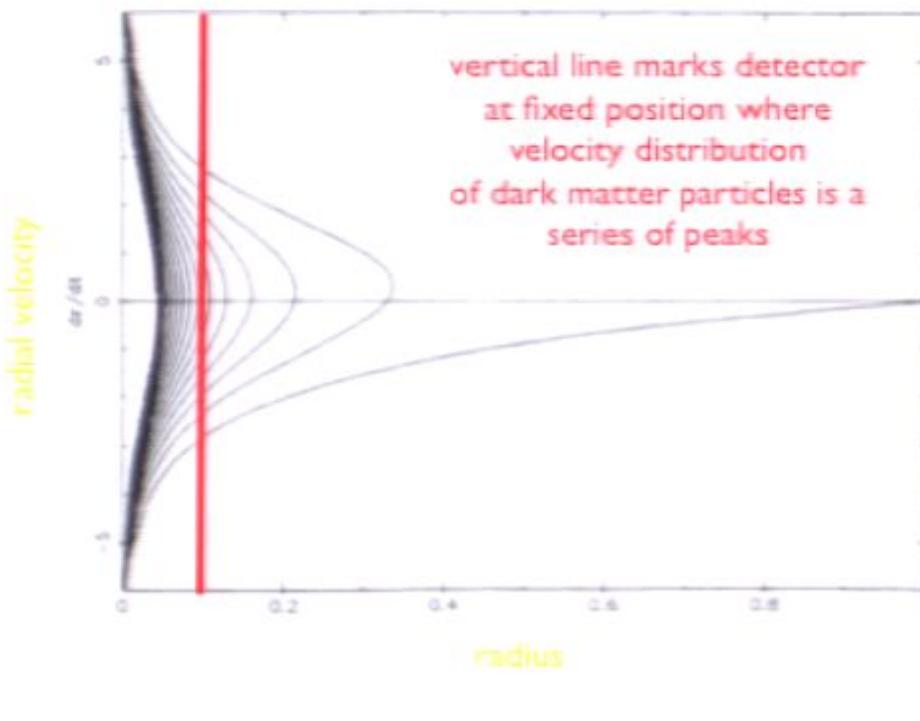




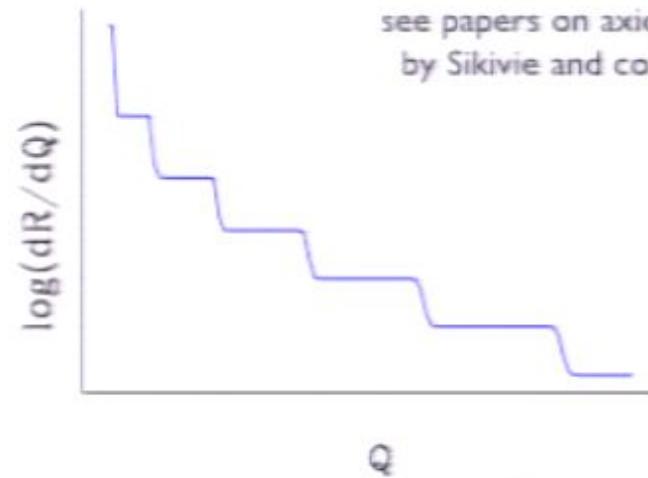
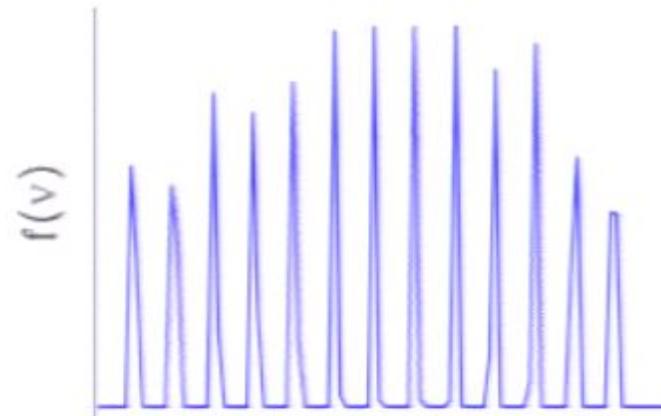
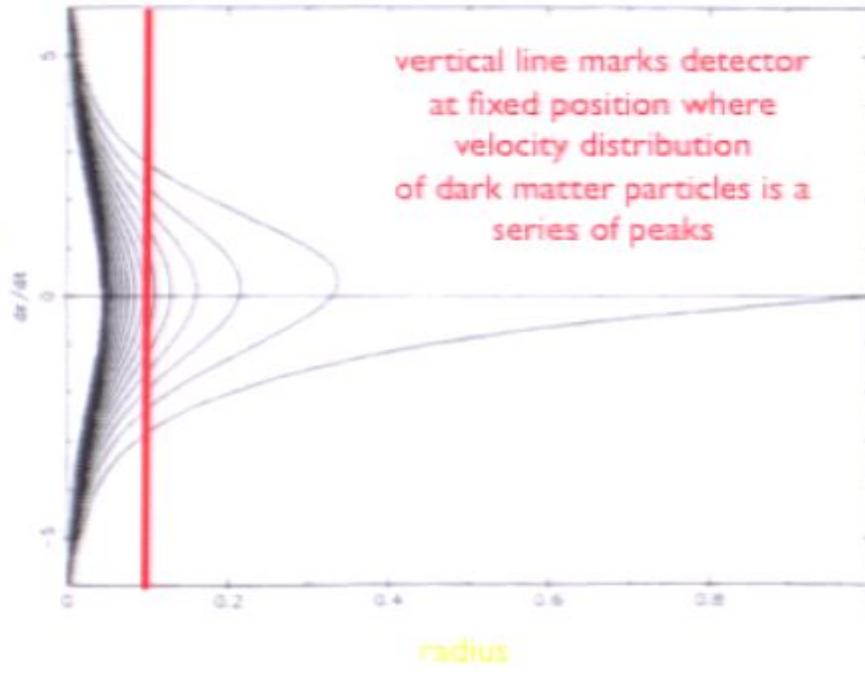


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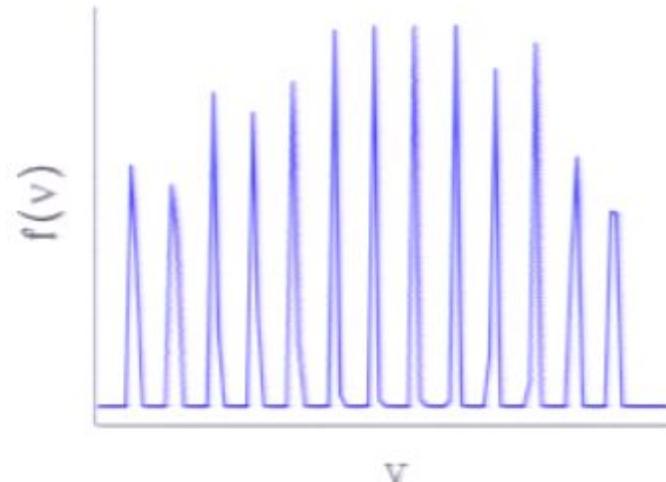
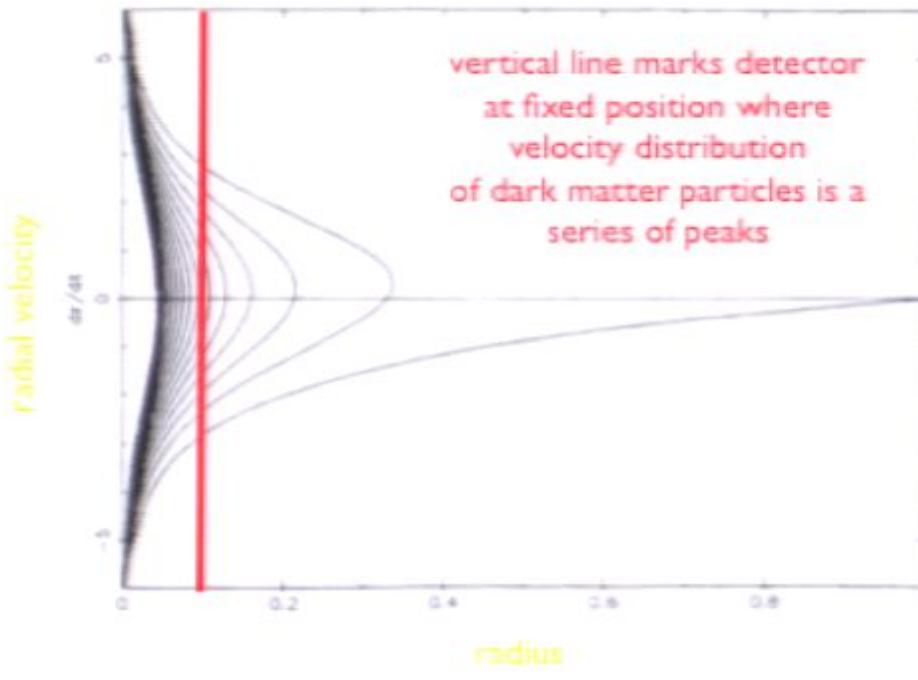




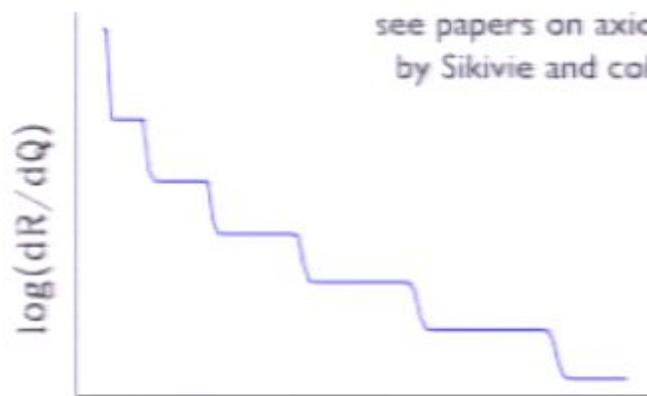
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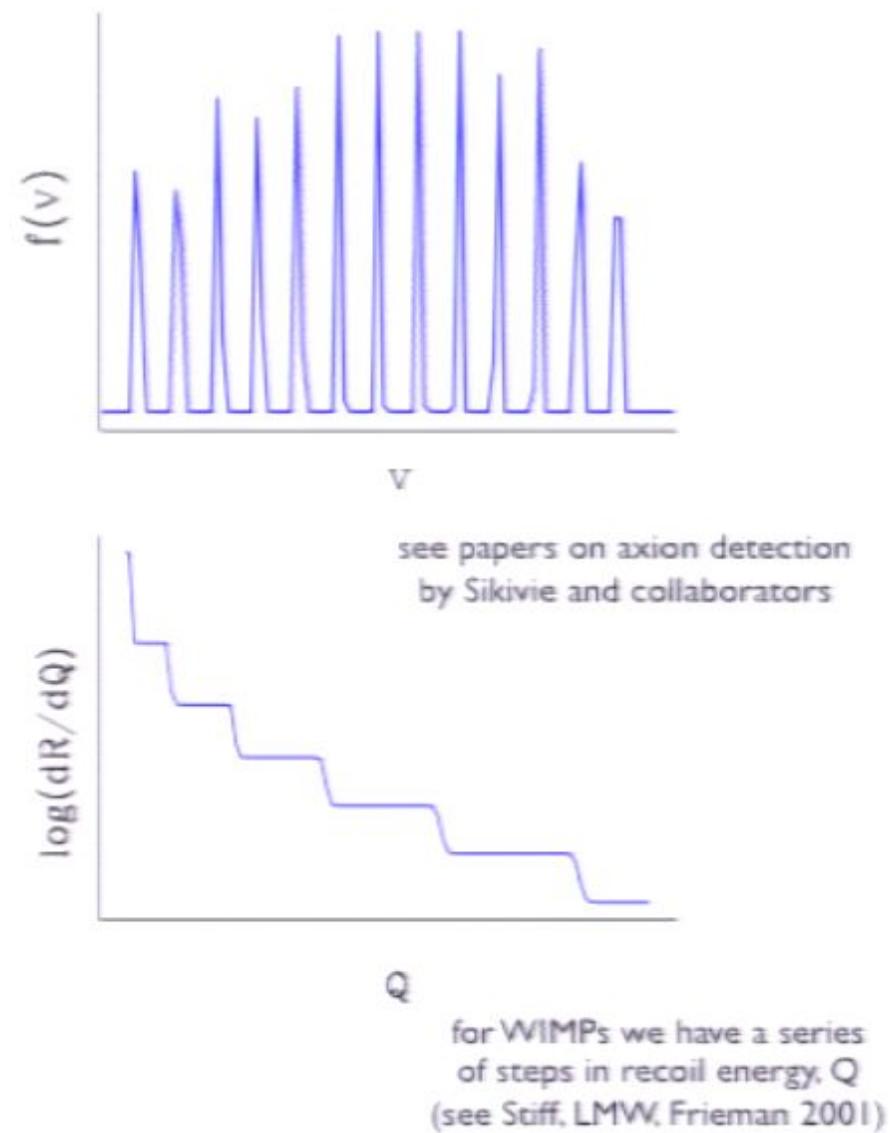
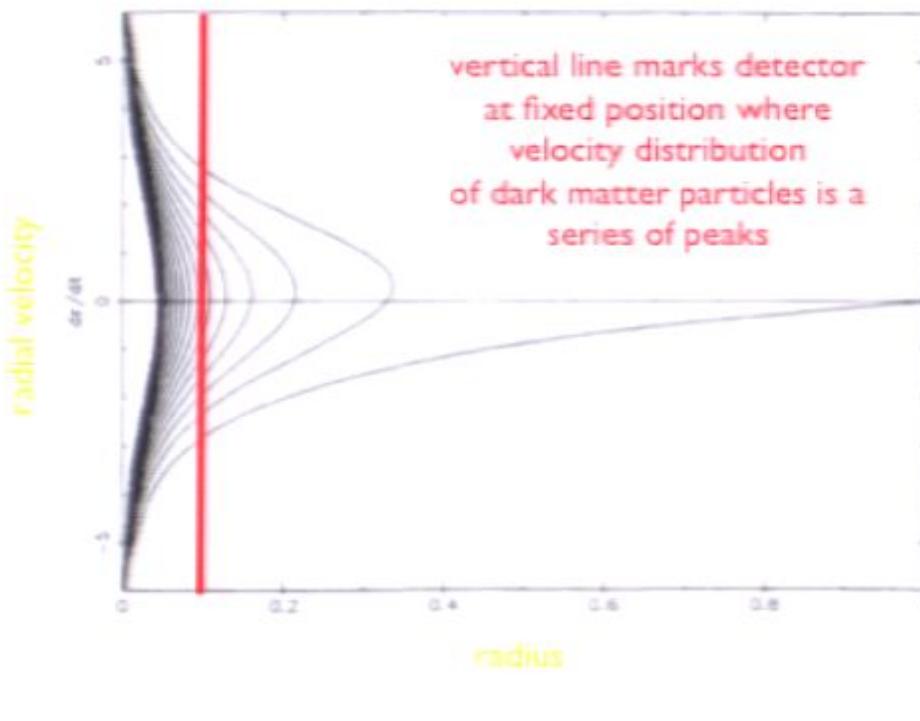


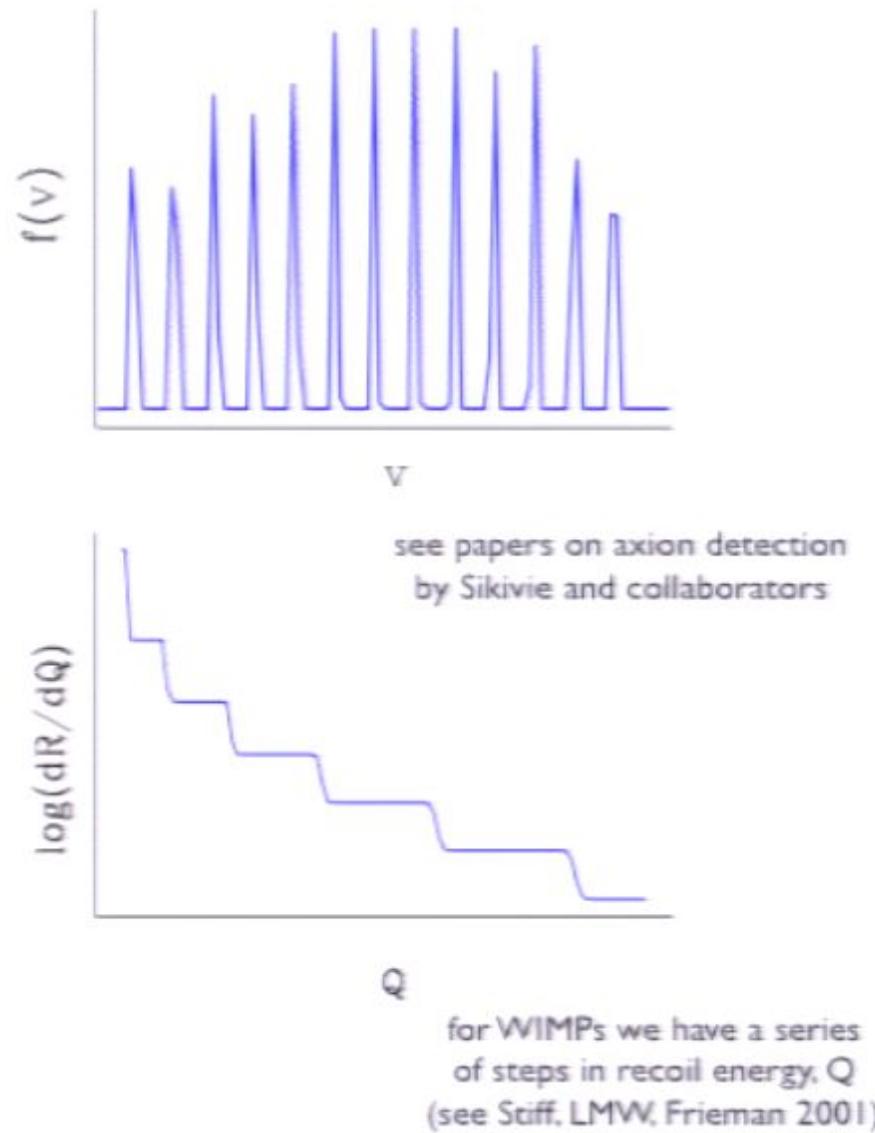
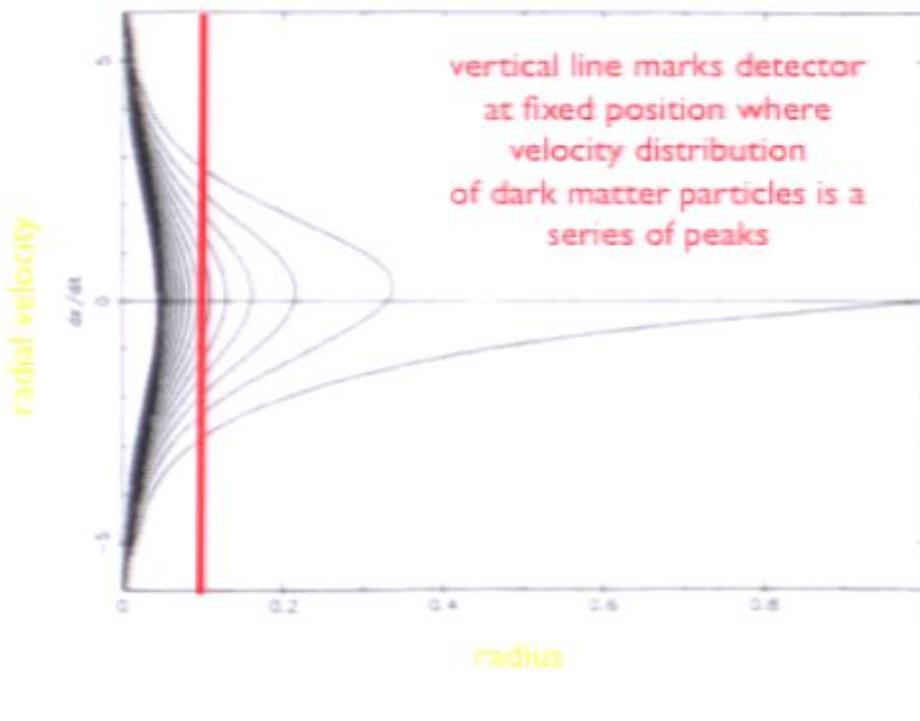
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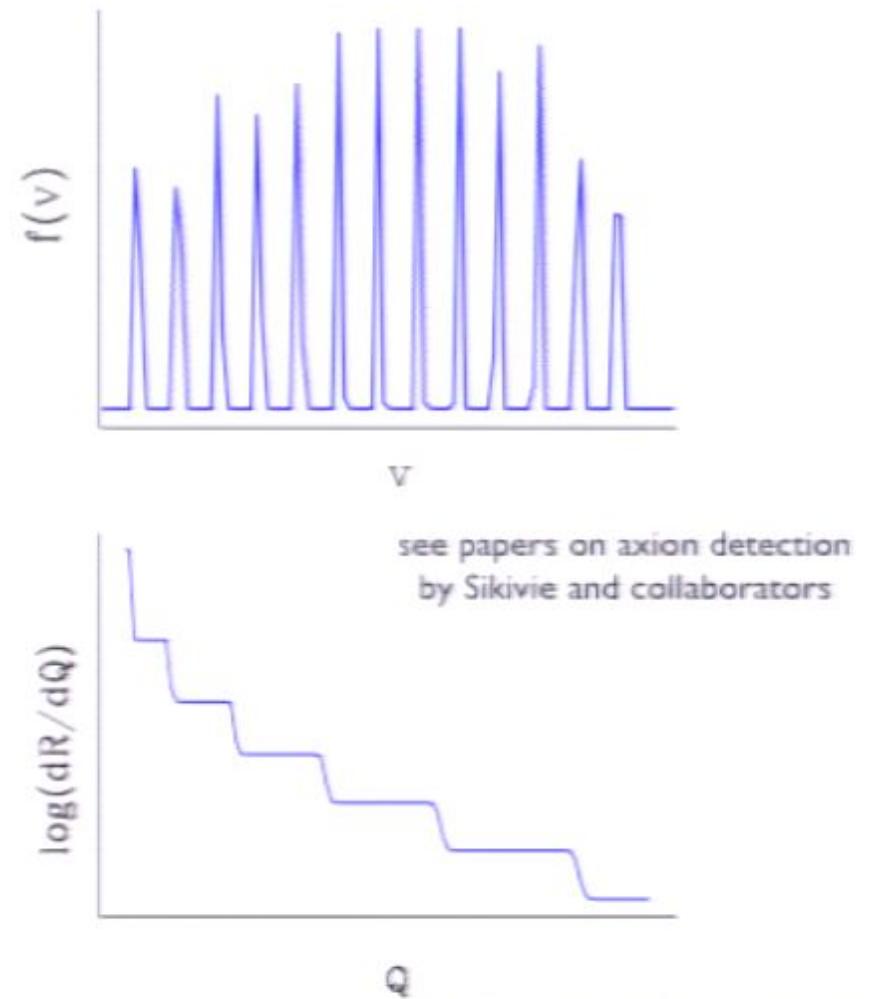
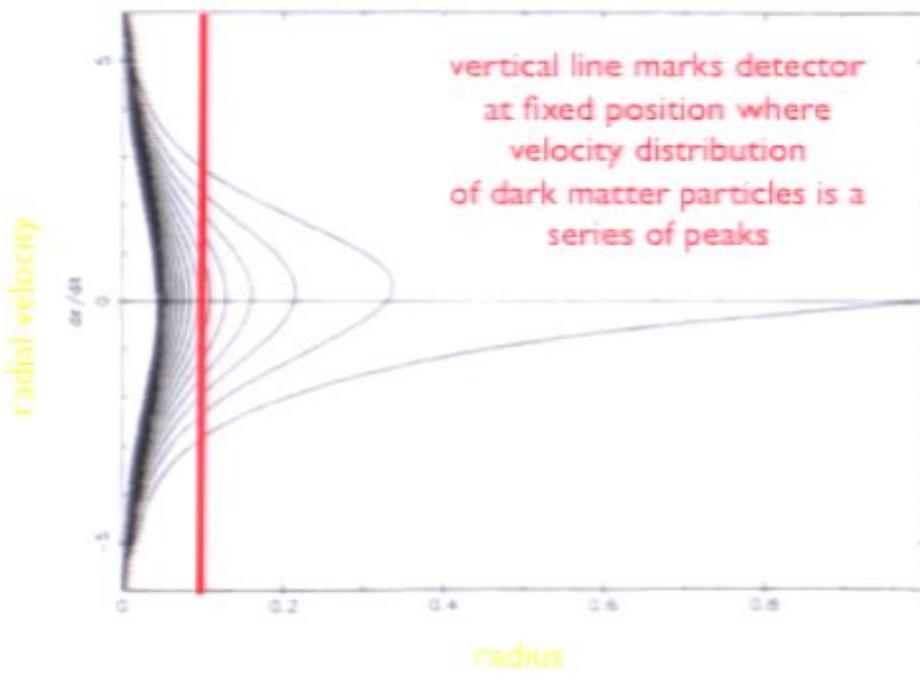


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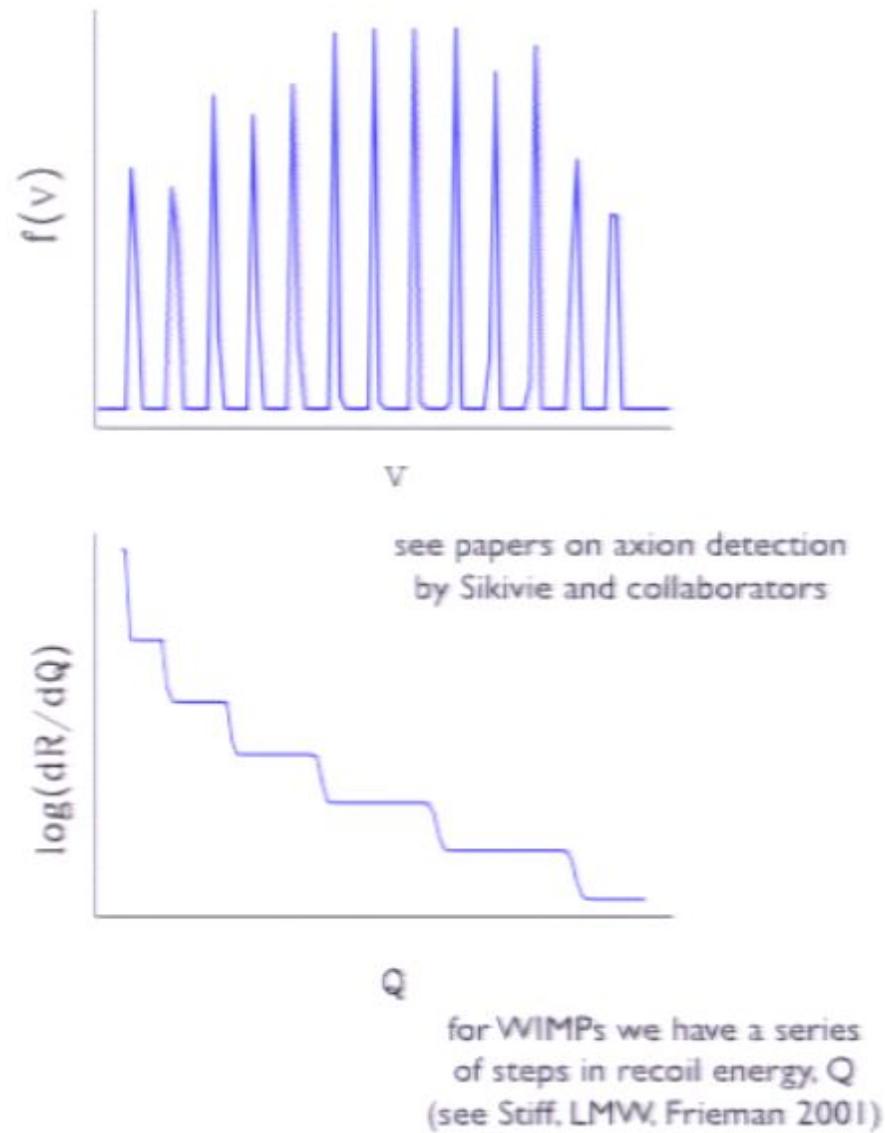
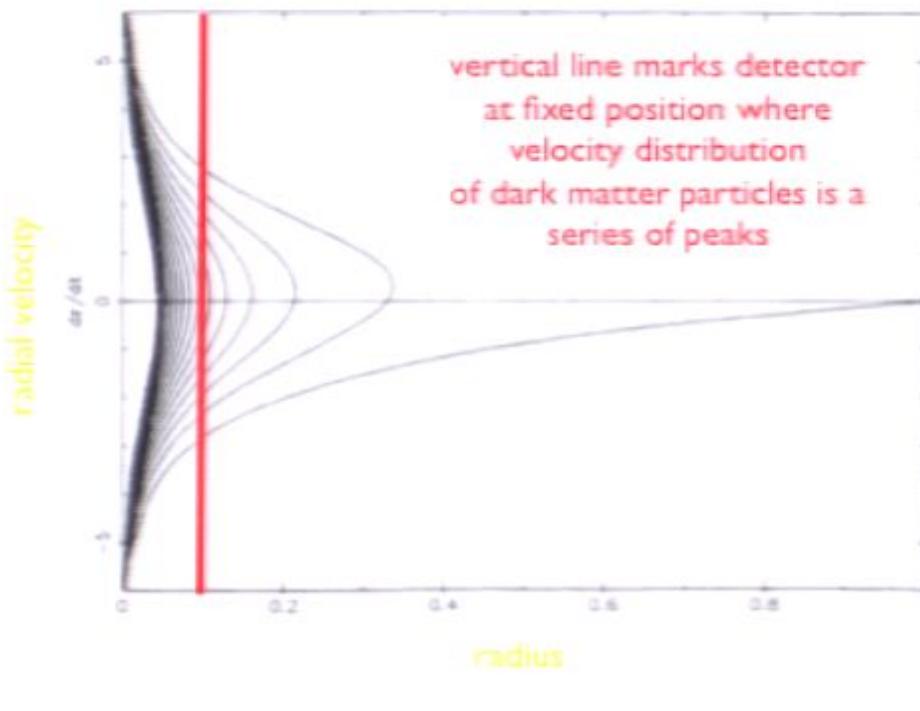
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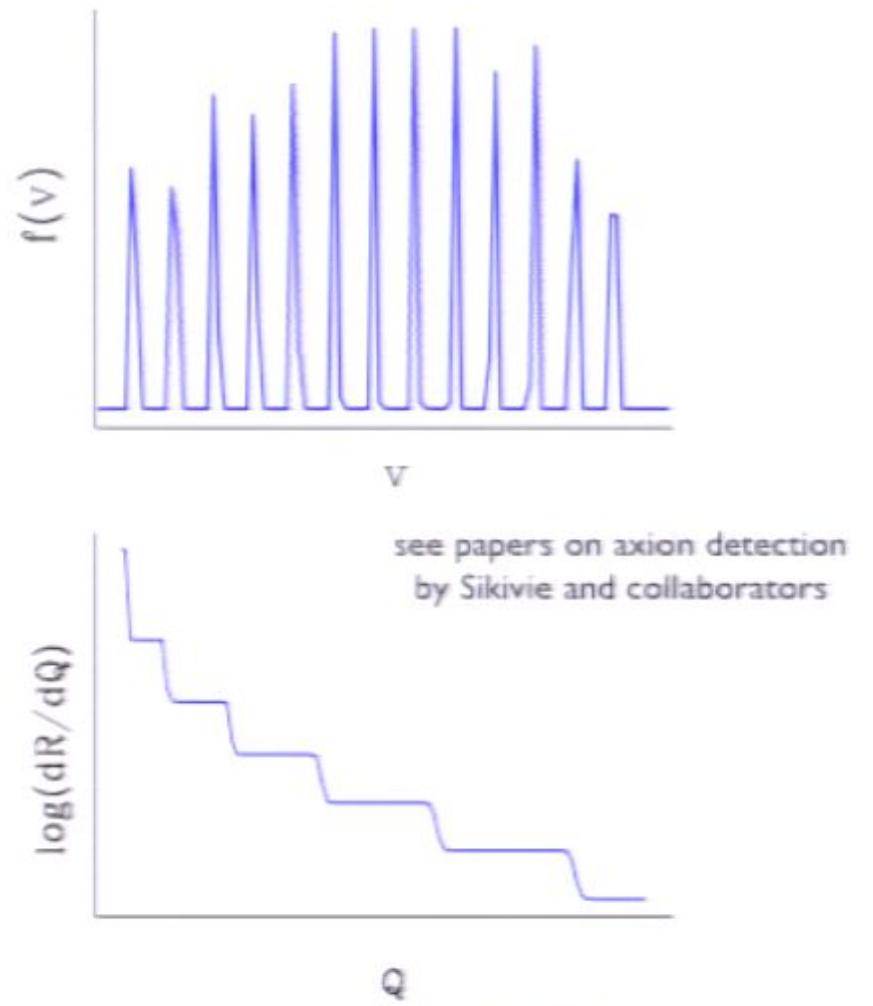
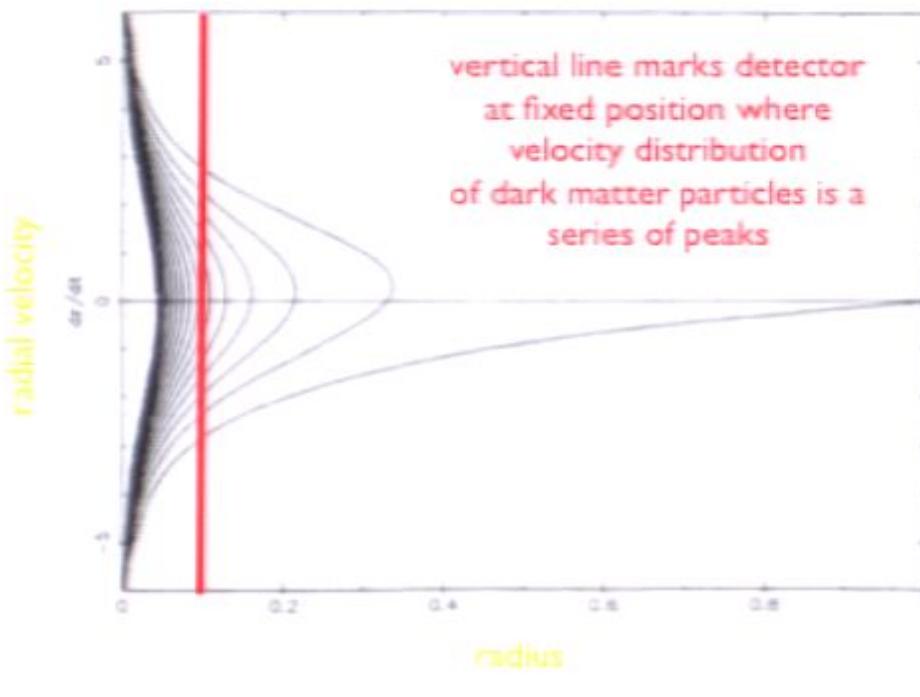




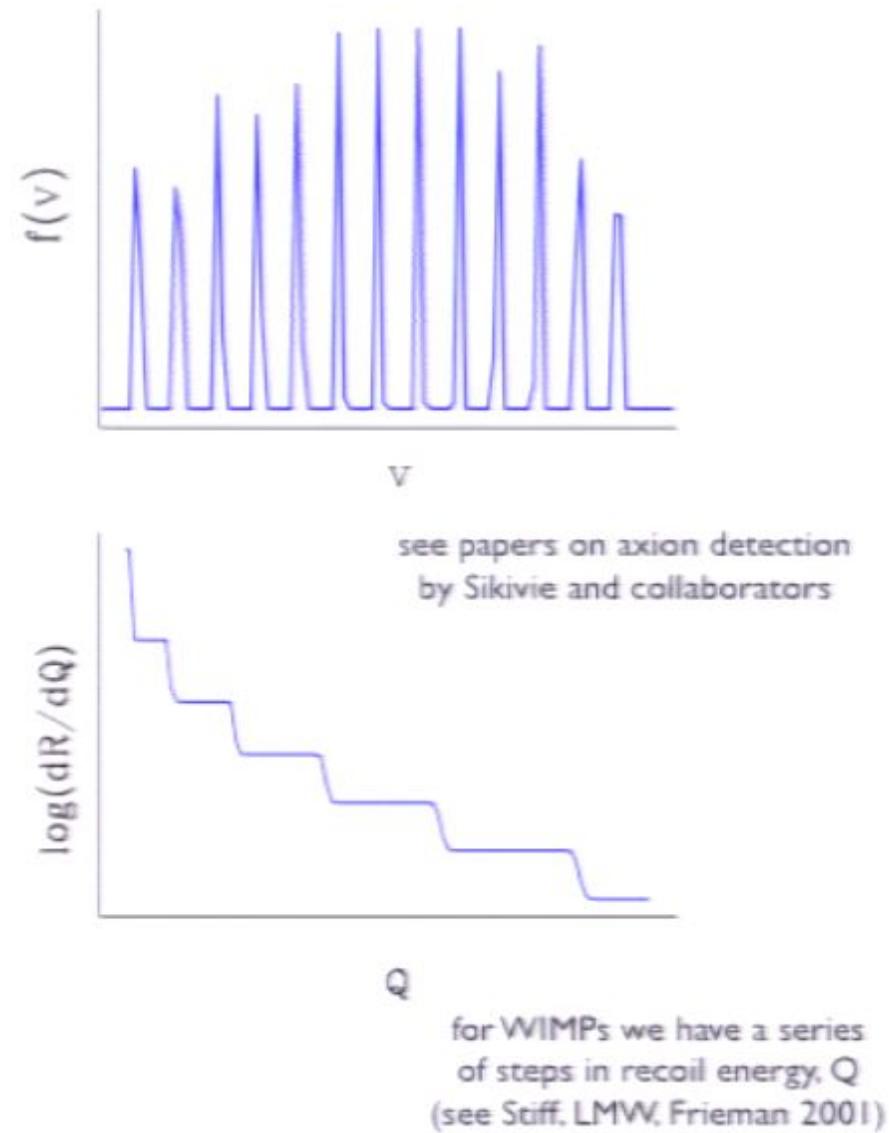
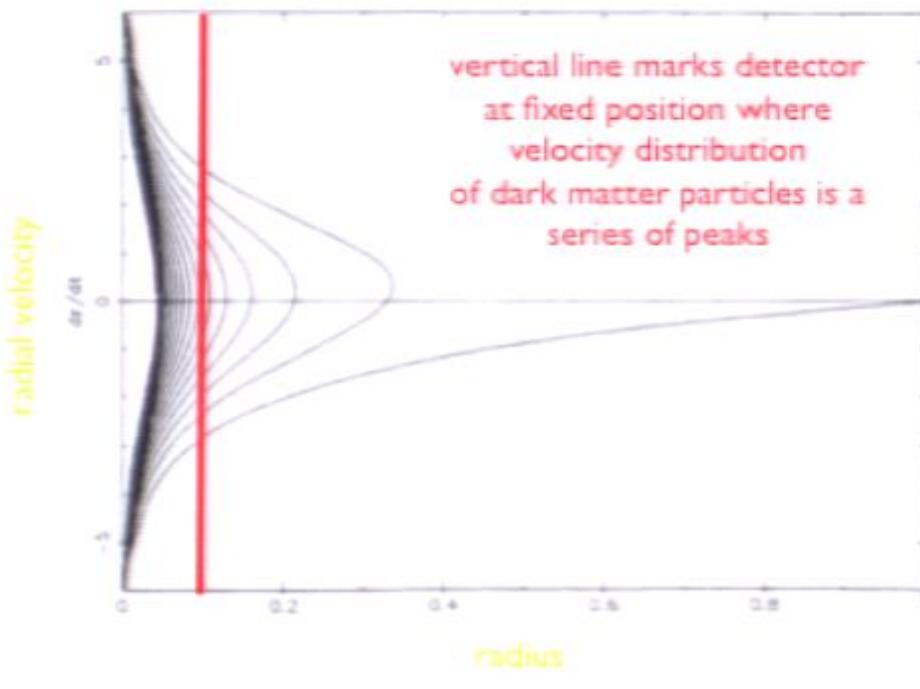


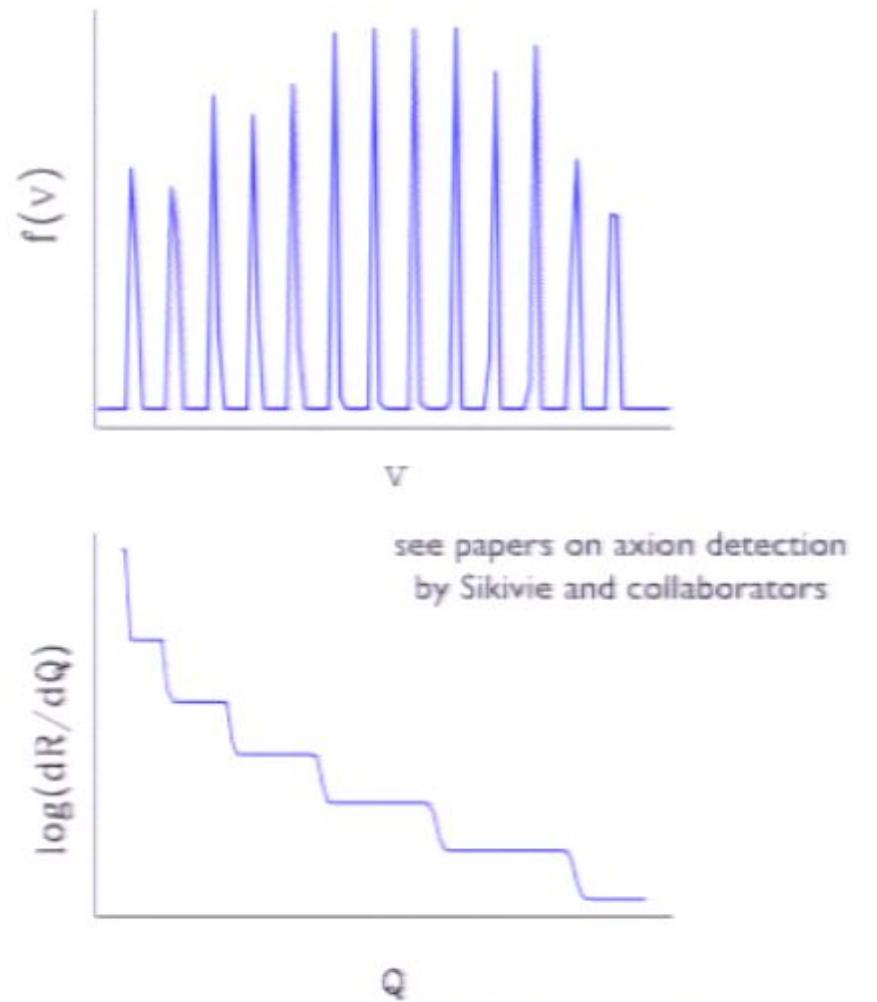
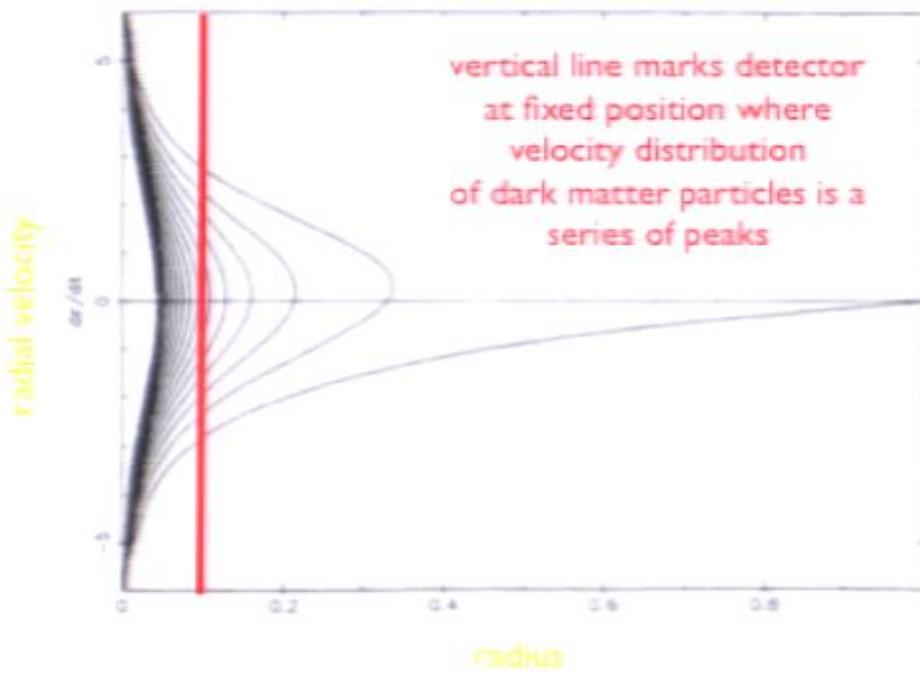
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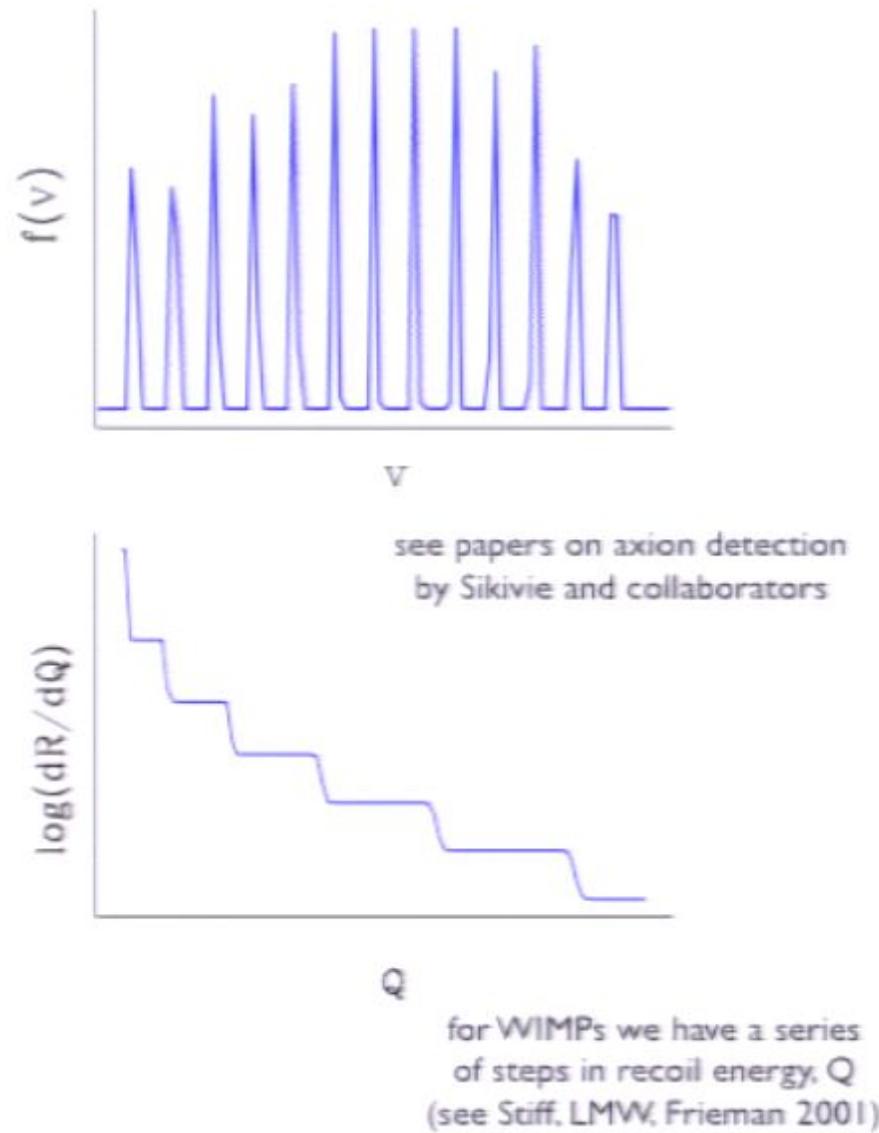
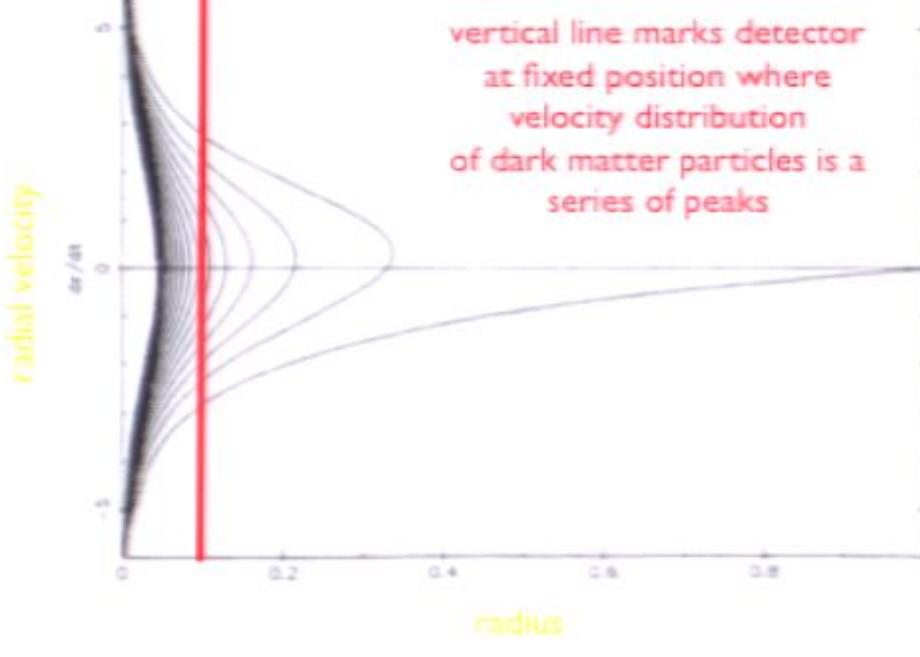


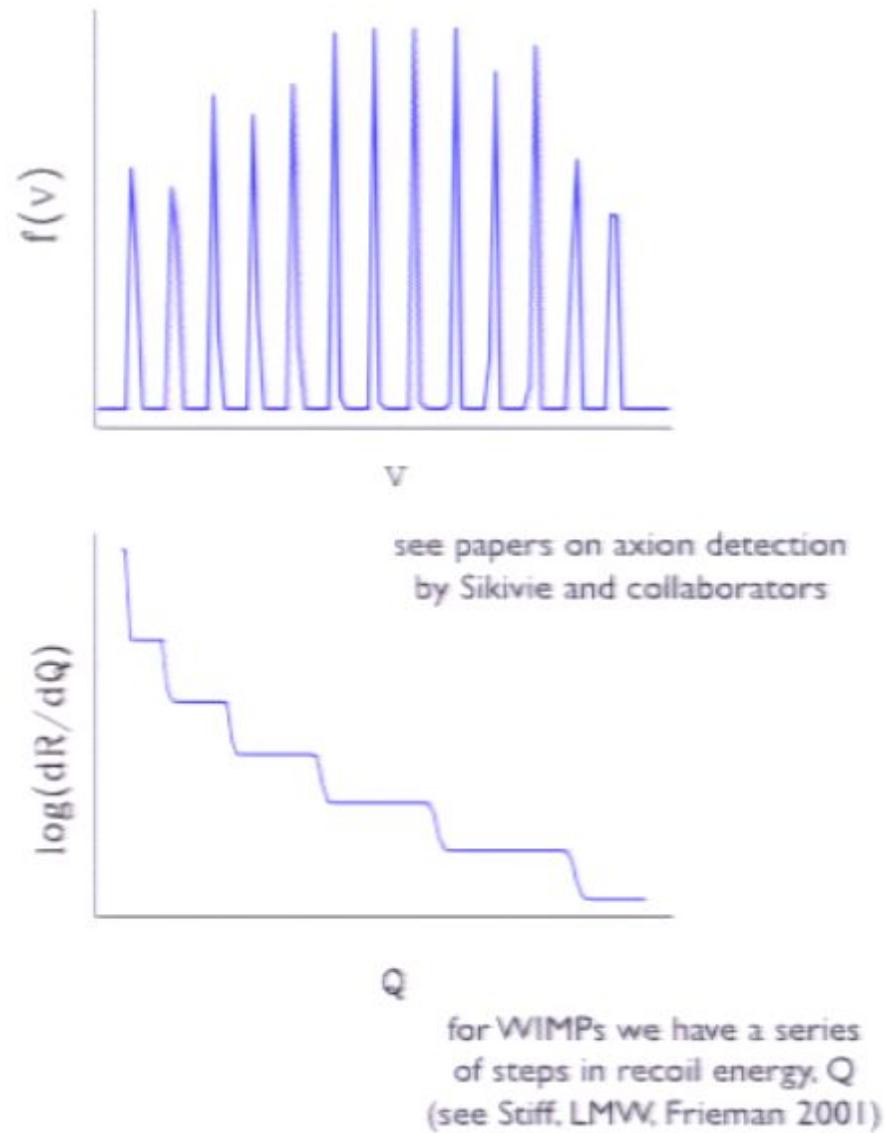
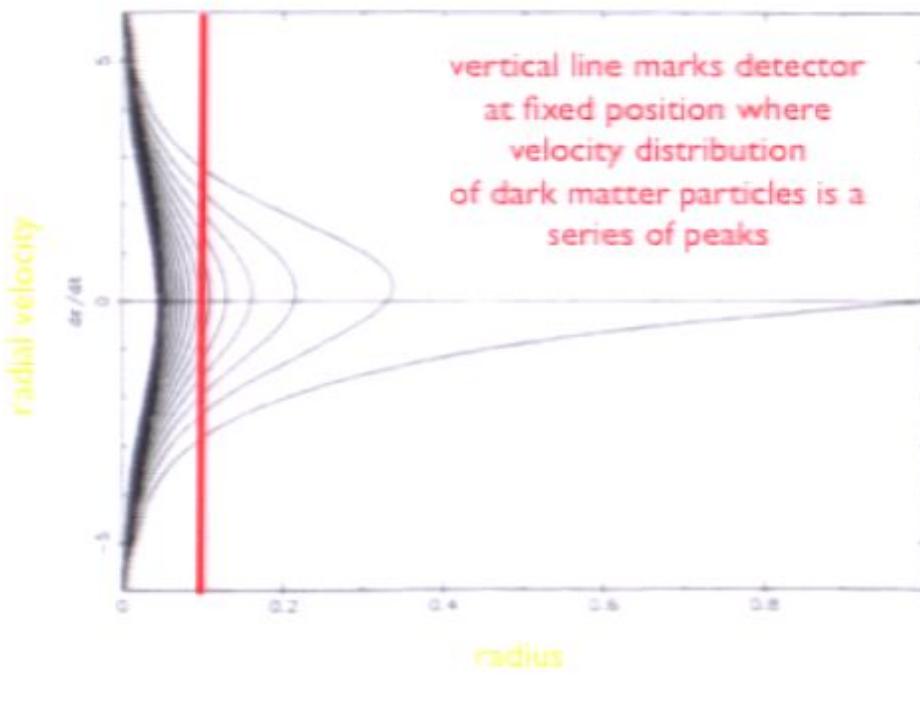
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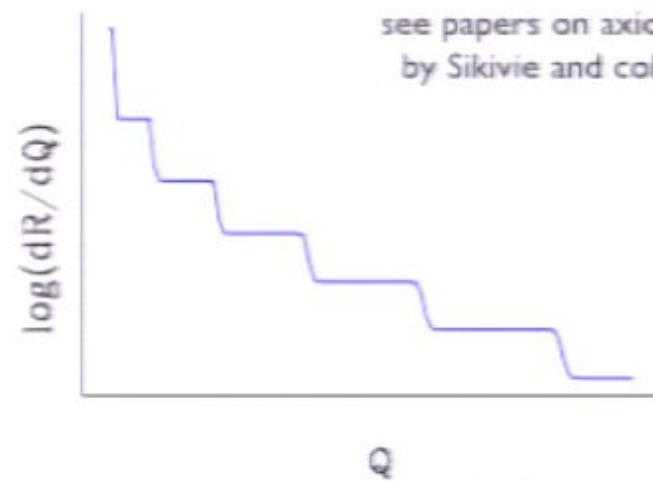
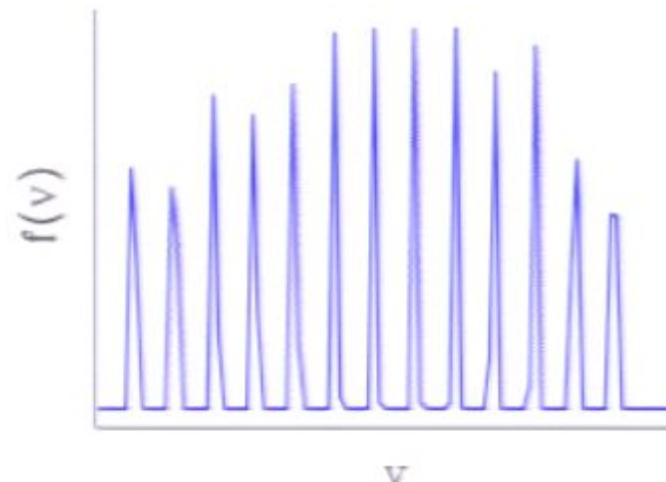
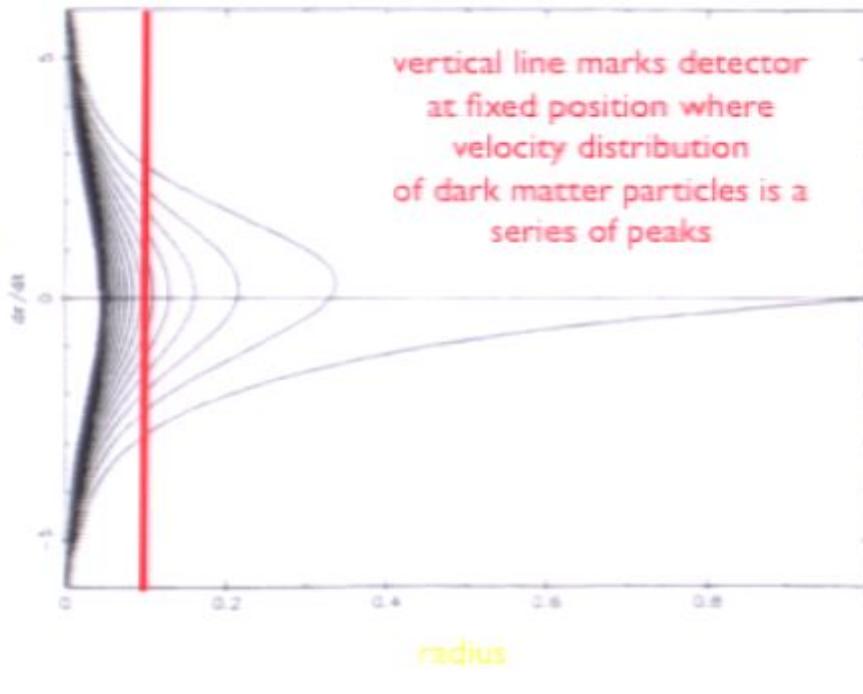


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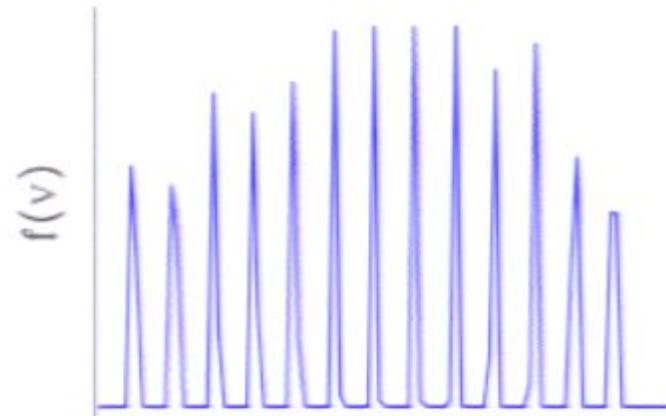
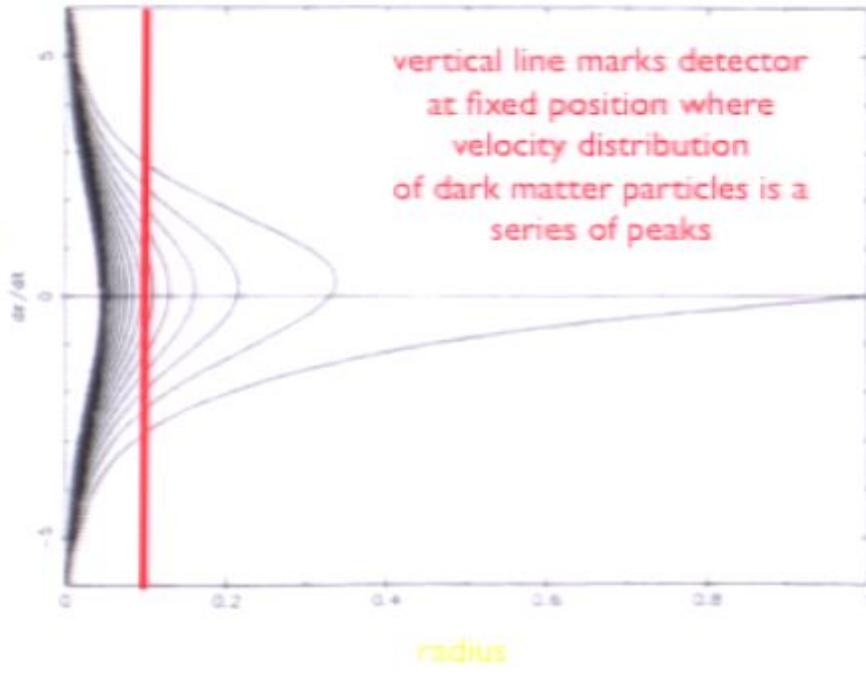


radial velocity

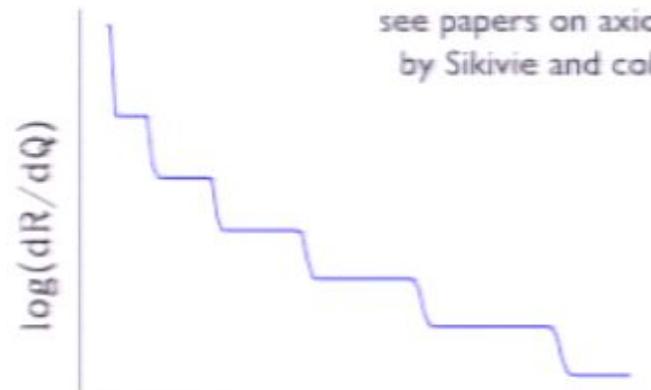


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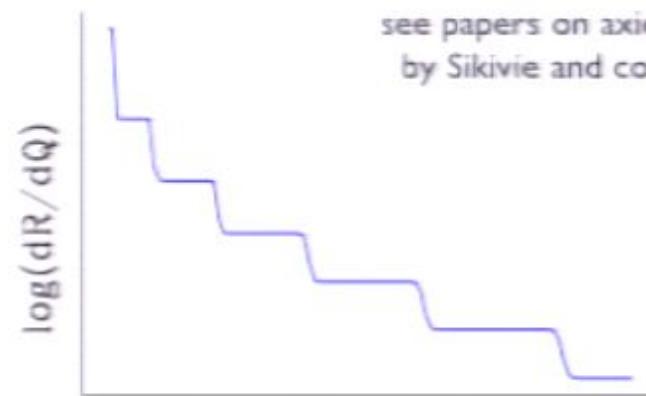
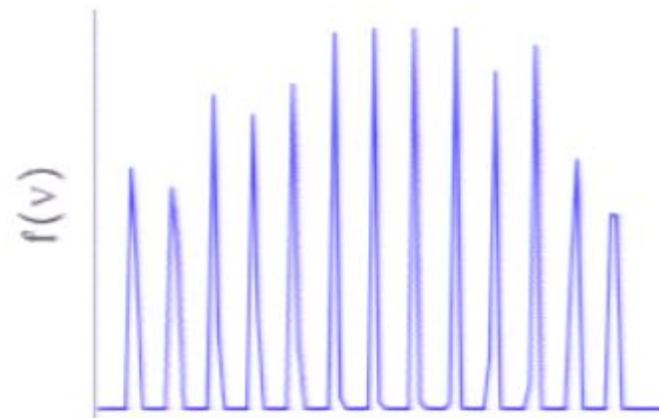
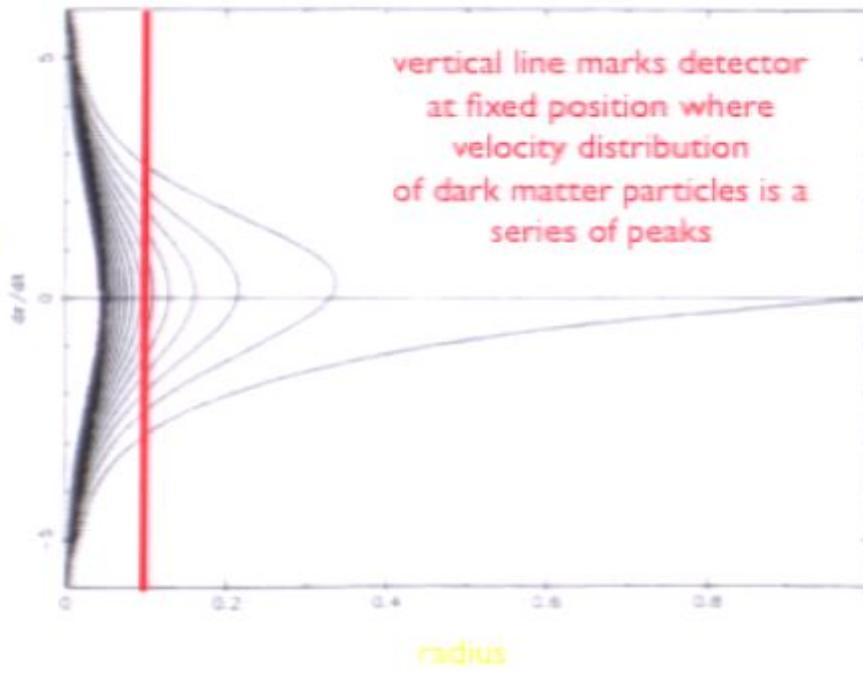
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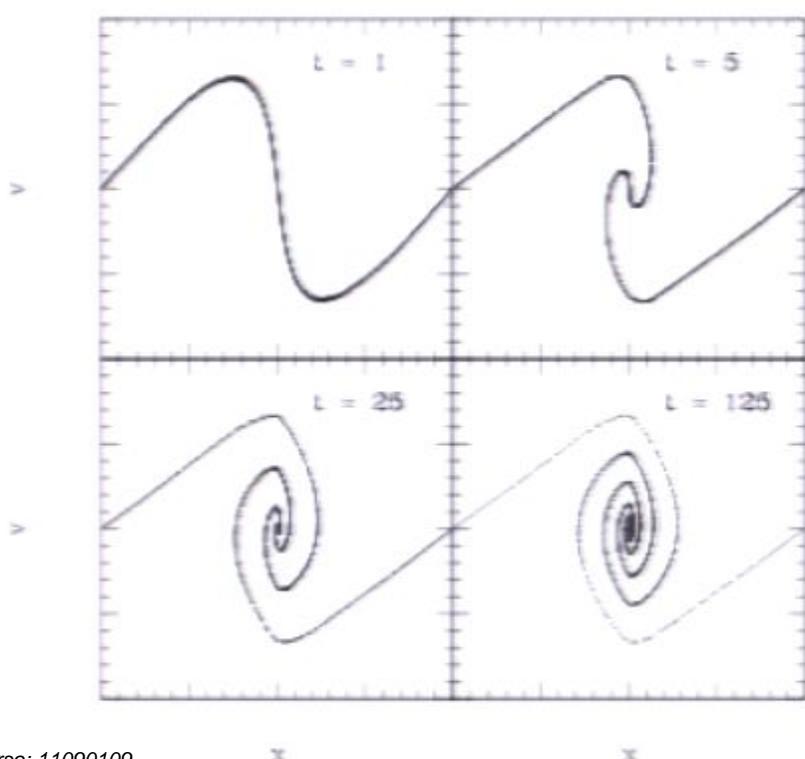
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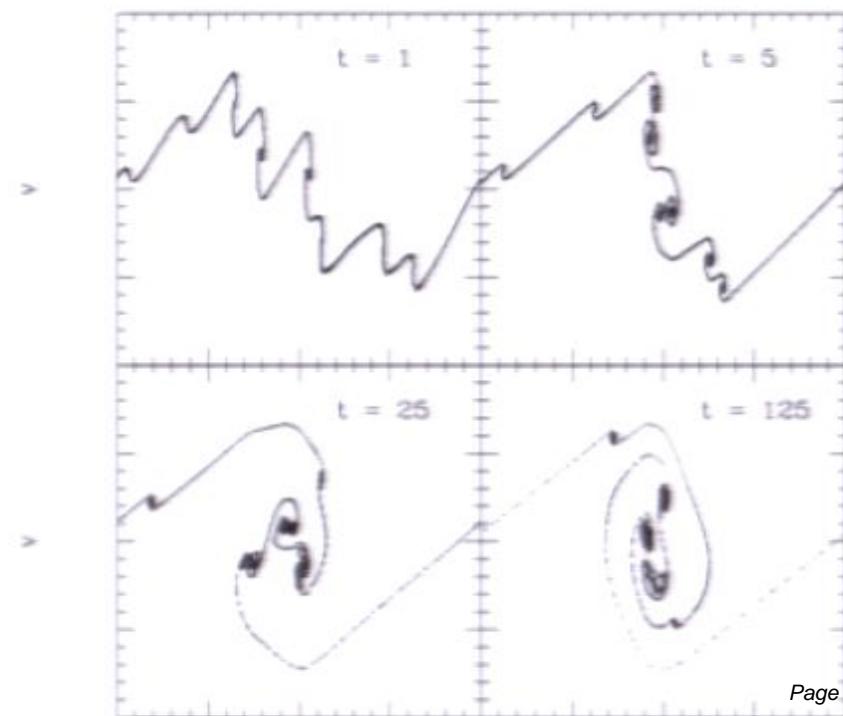
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In the standard model for structure formation on all scales and structure formation proceeds perturbations exist hierarchically

single mode -- planar collapse



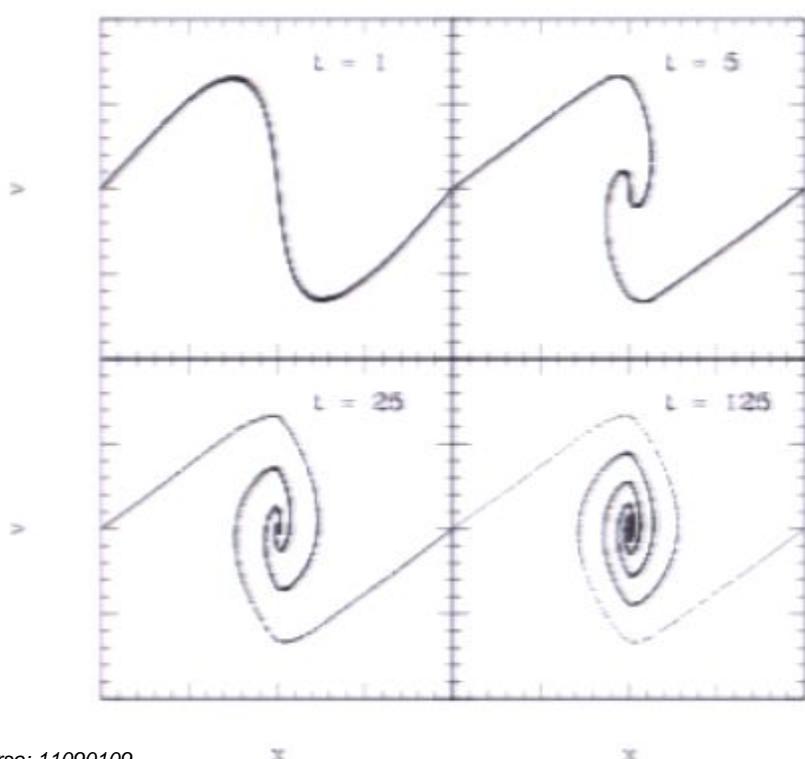
1D spectrum of modes



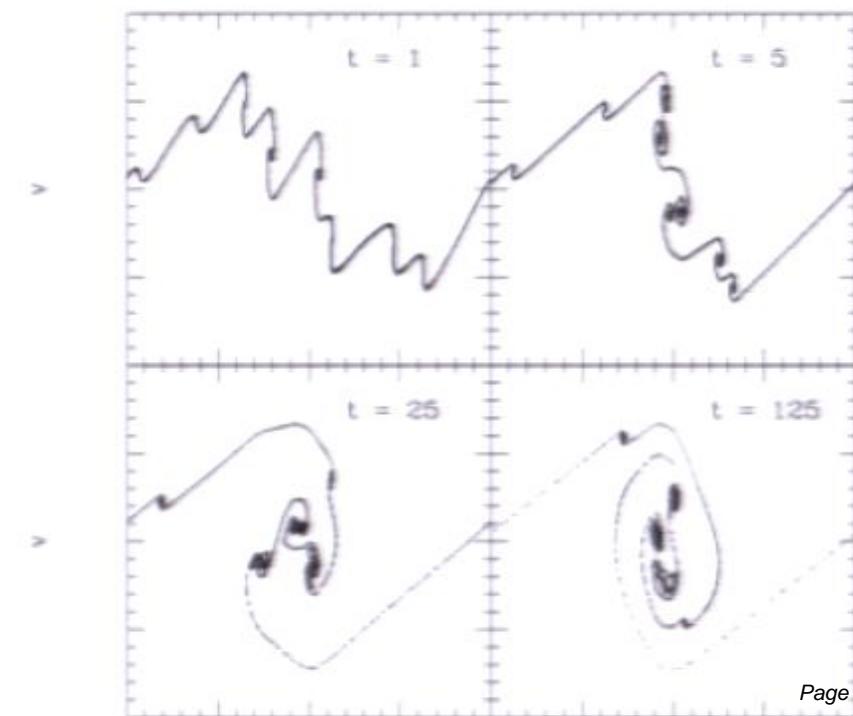
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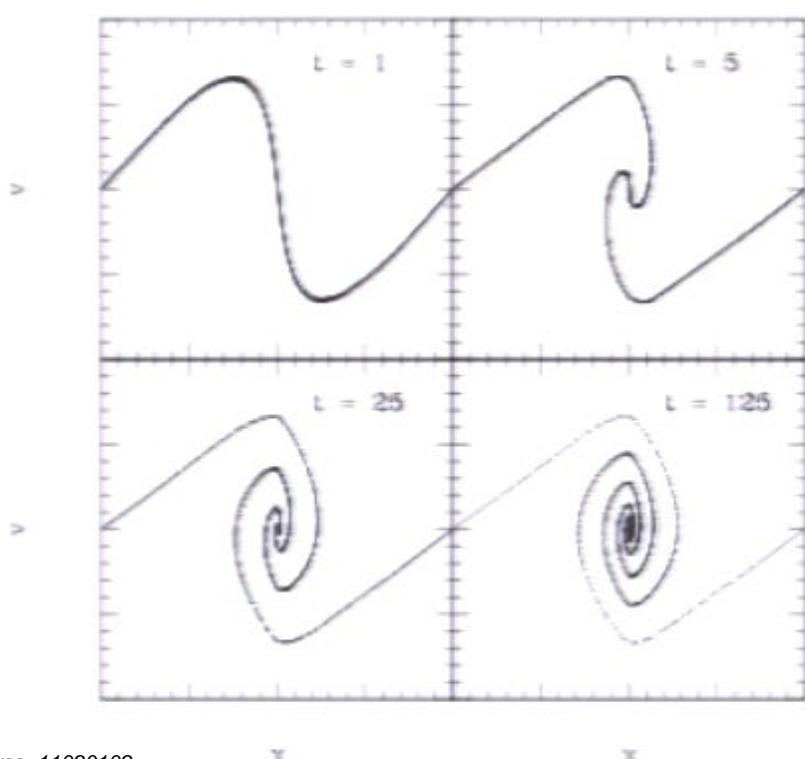
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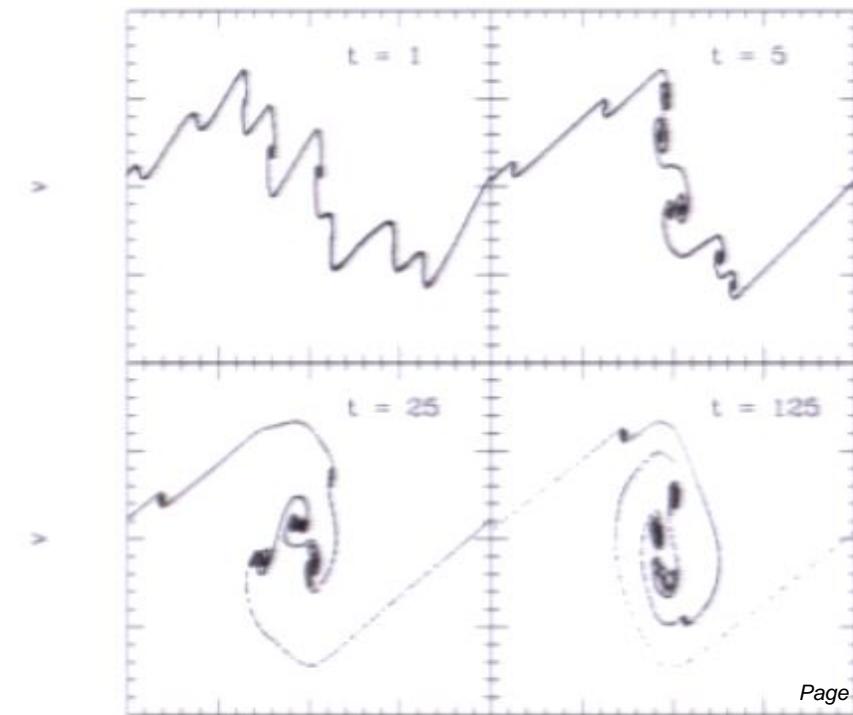
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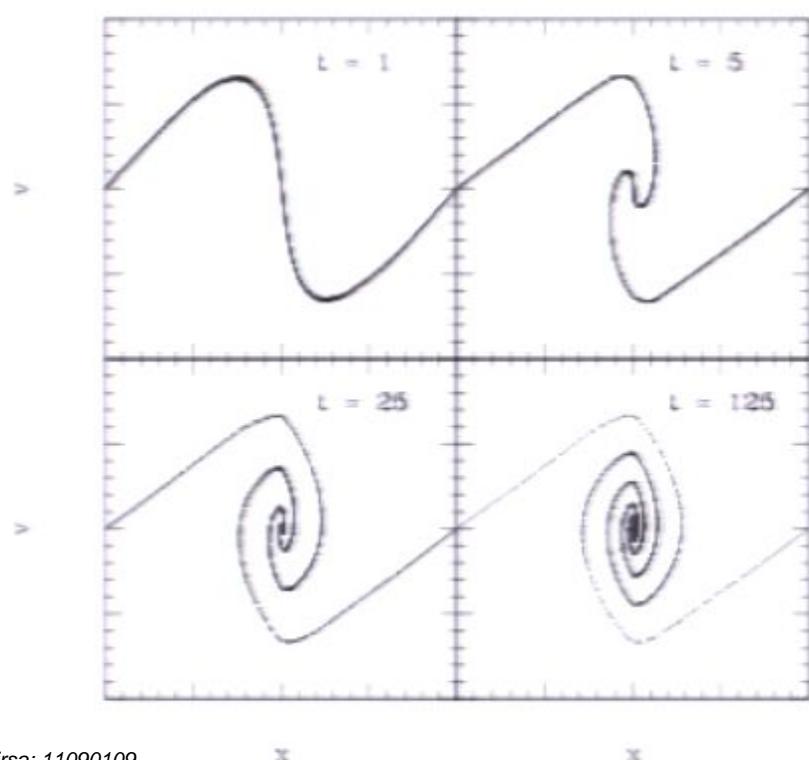
1D spectrum of modes



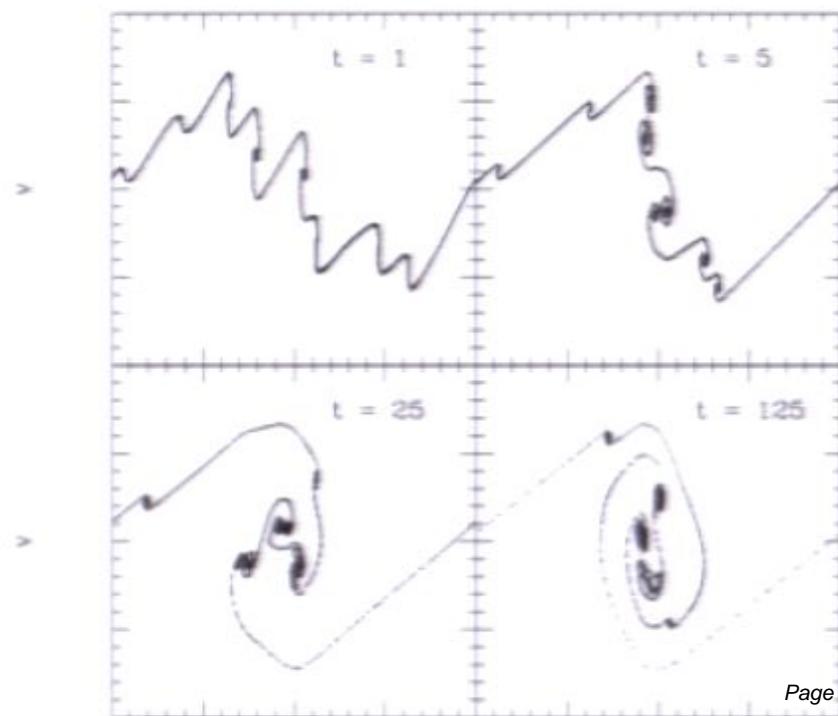
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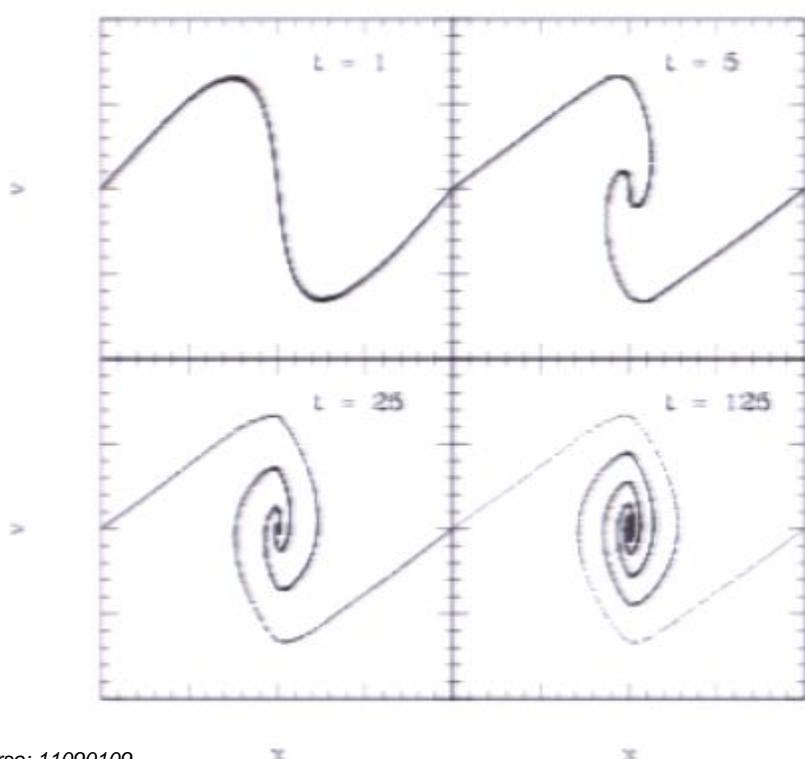
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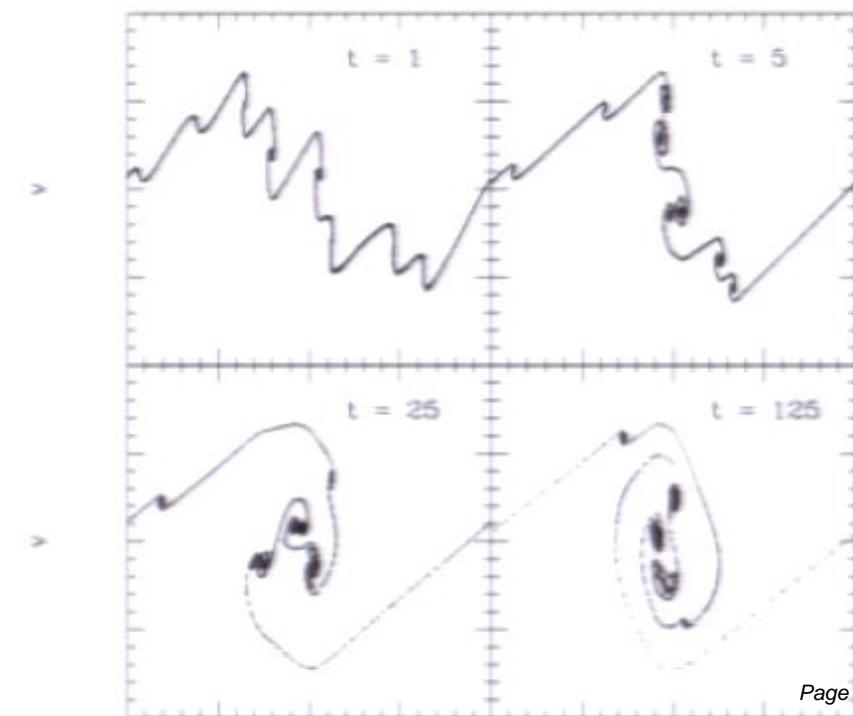
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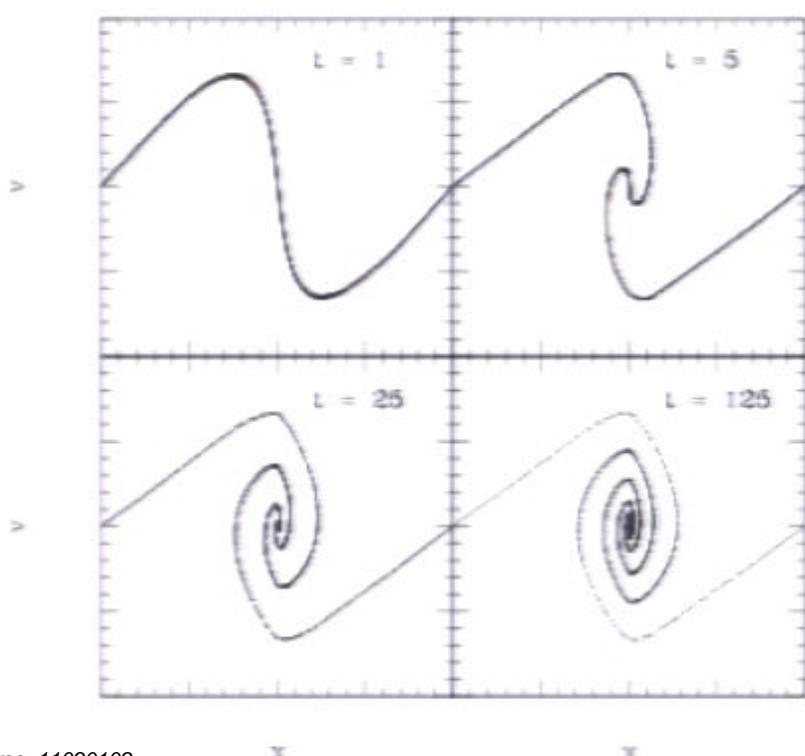
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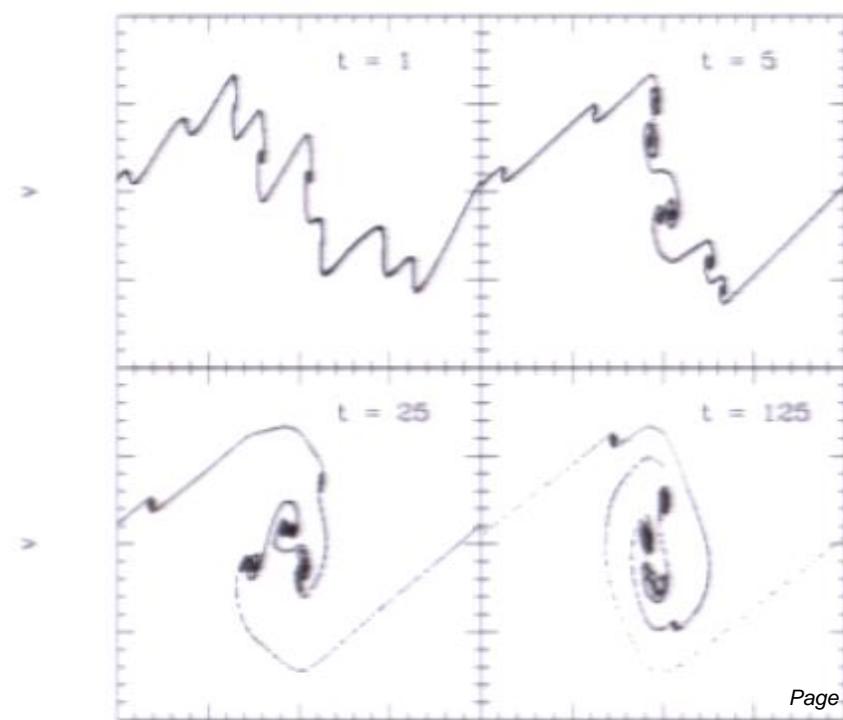
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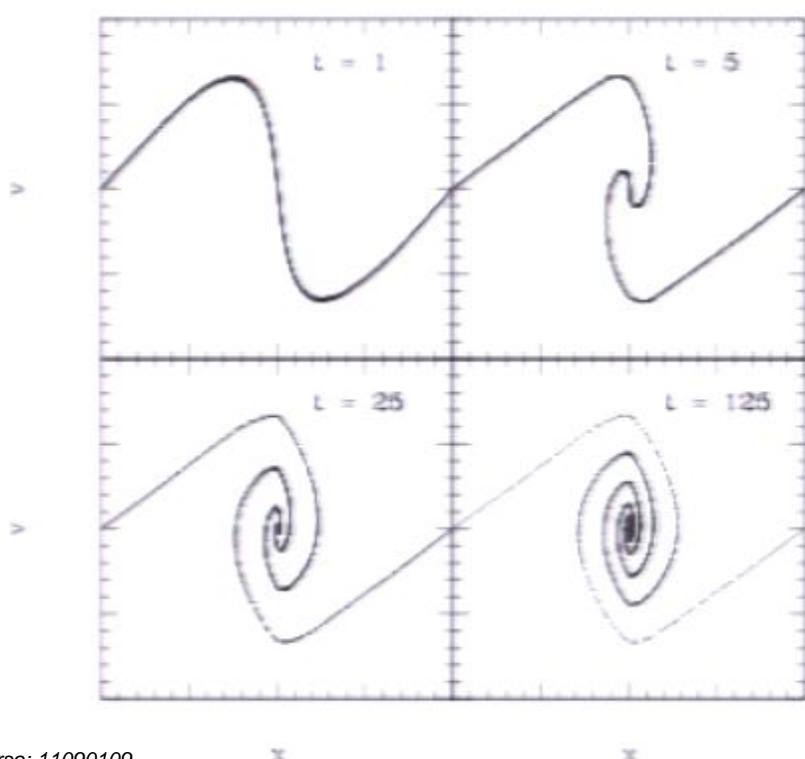
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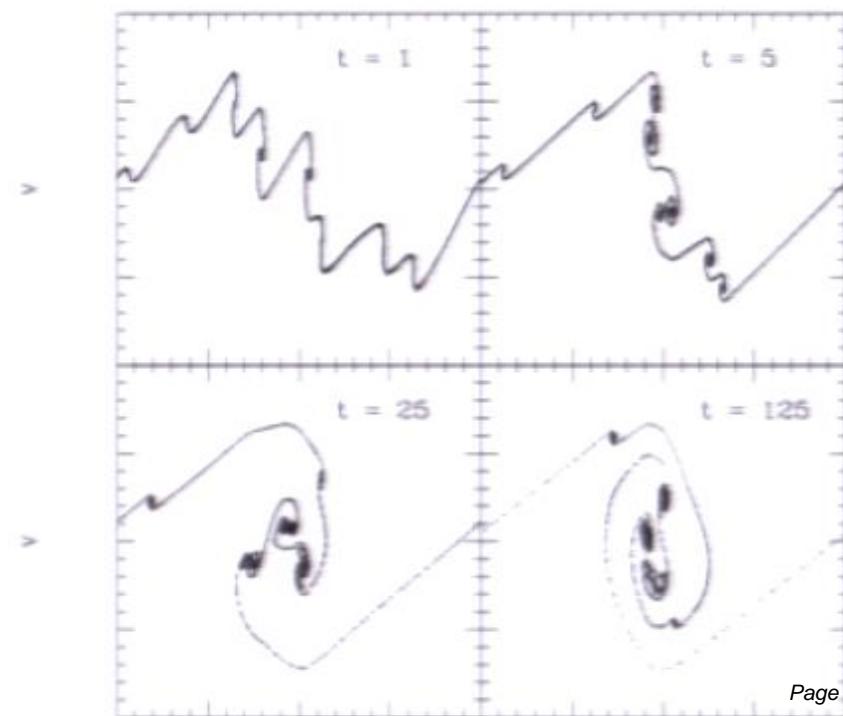
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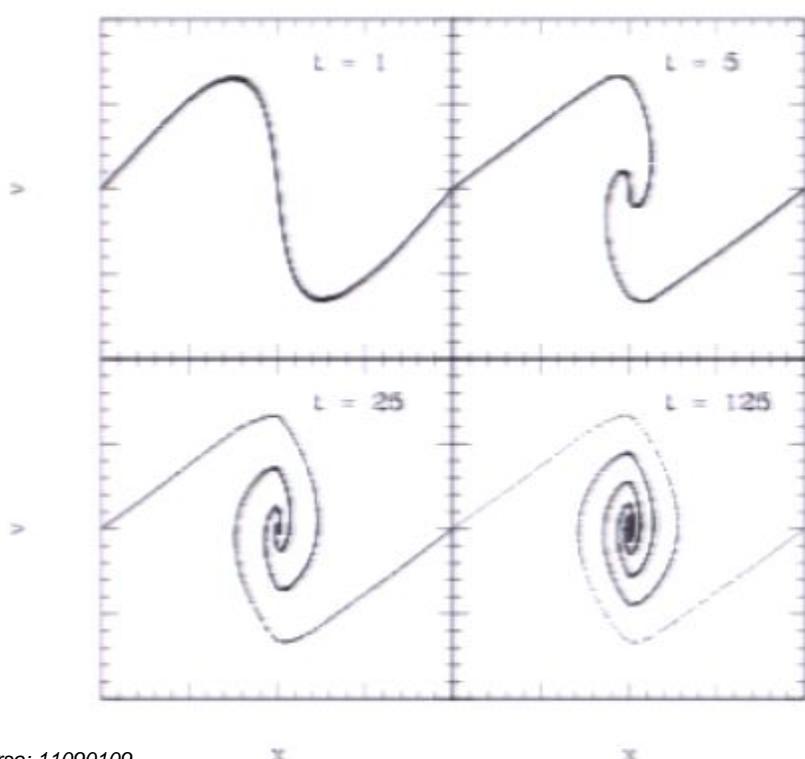
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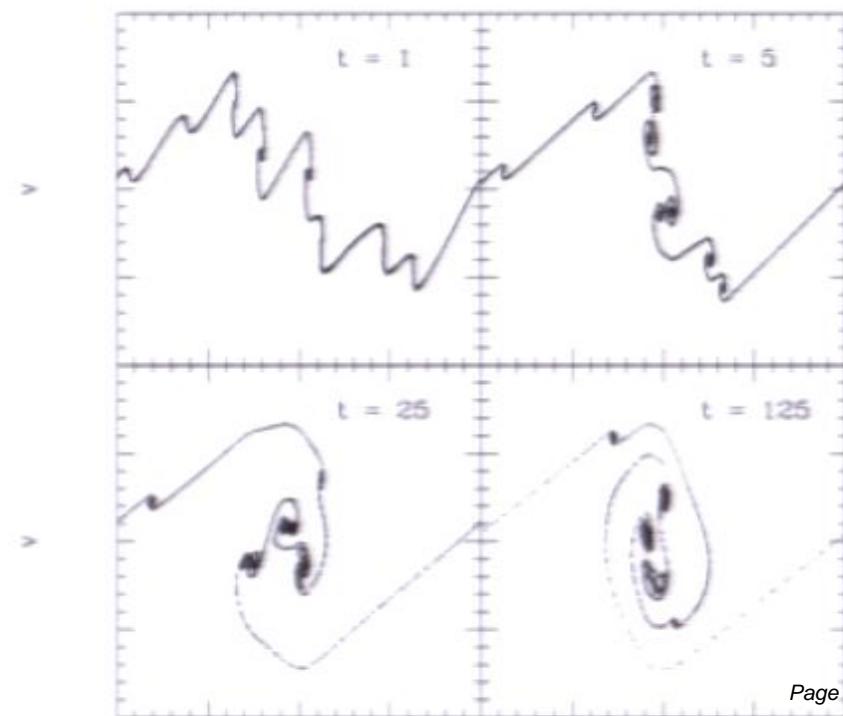
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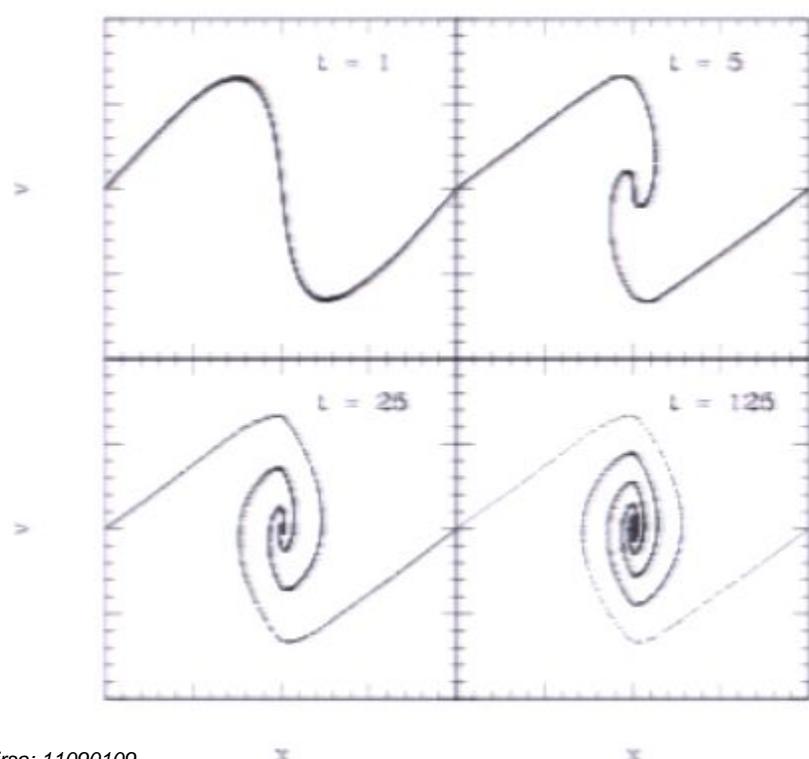
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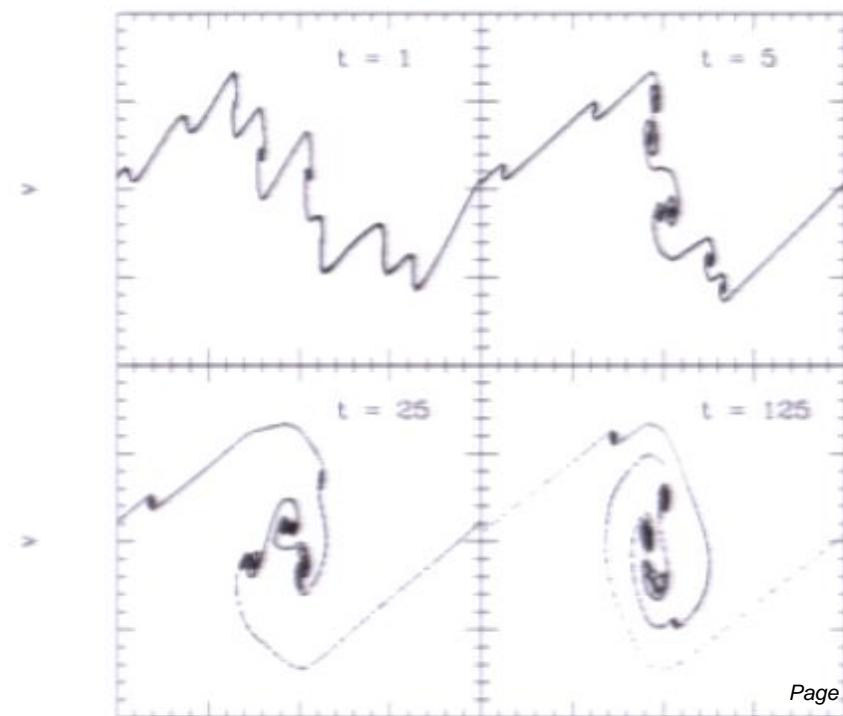
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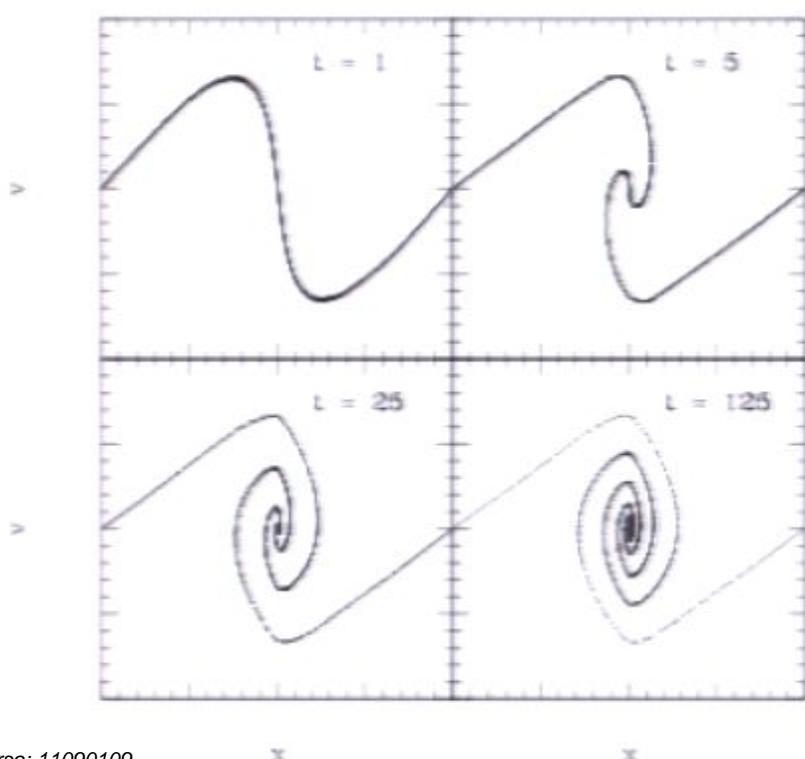
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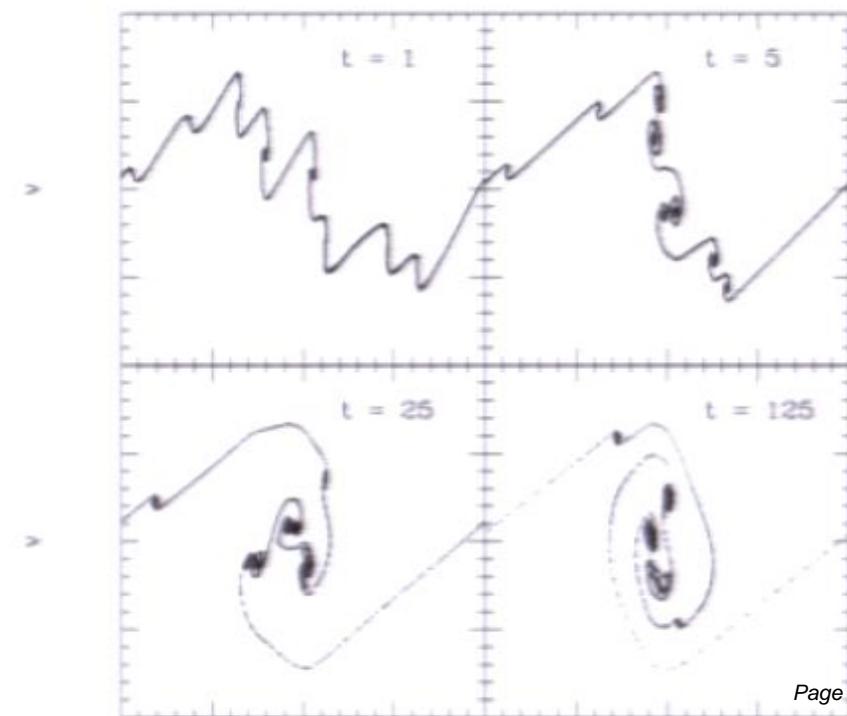
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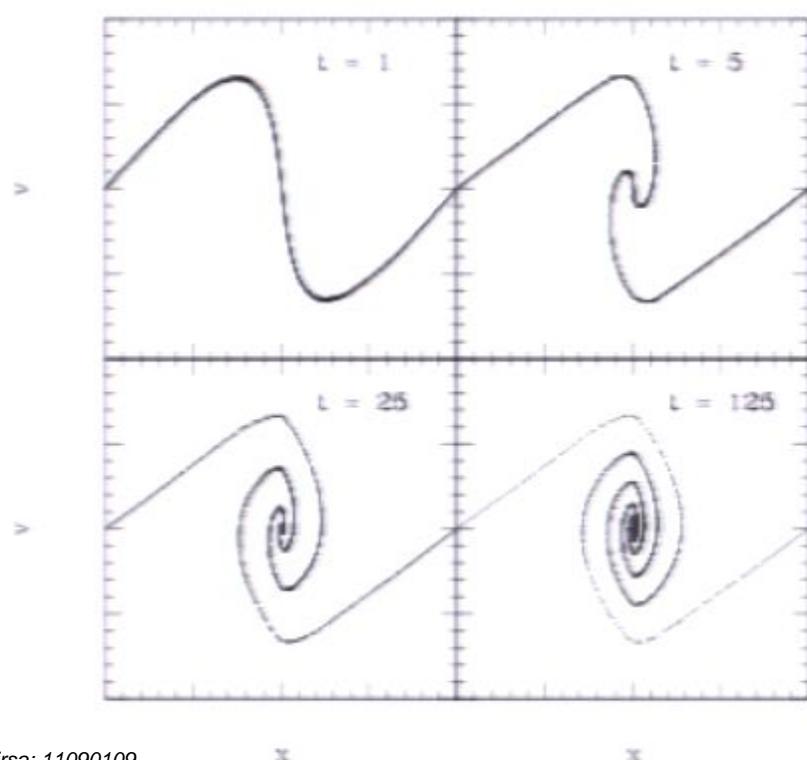
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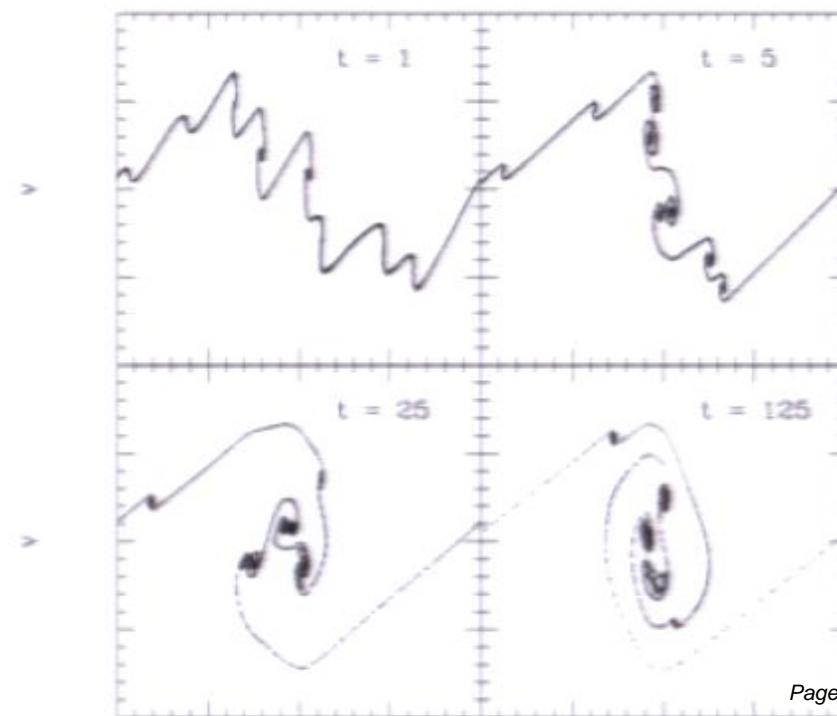
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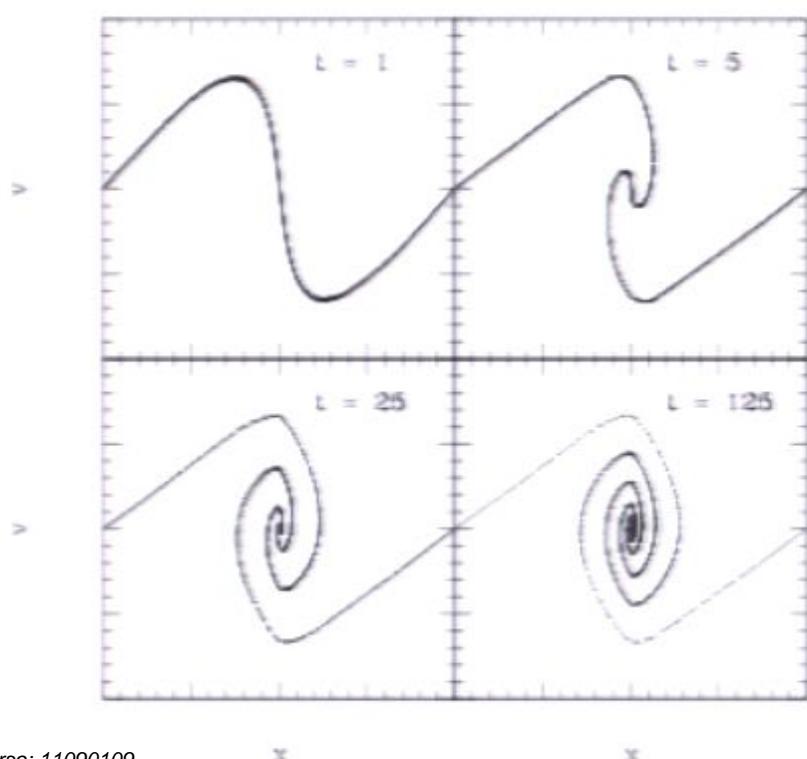
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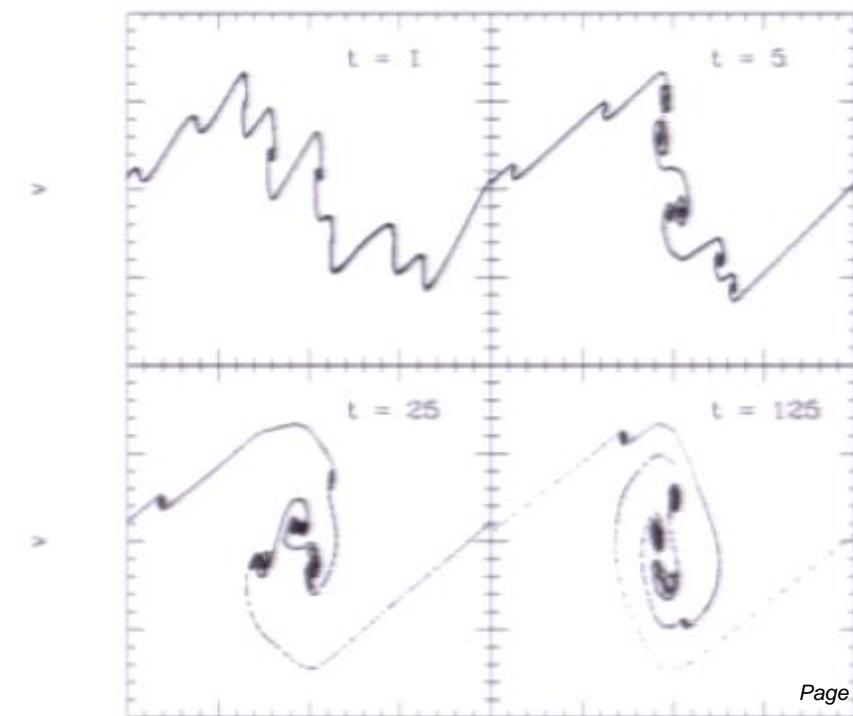
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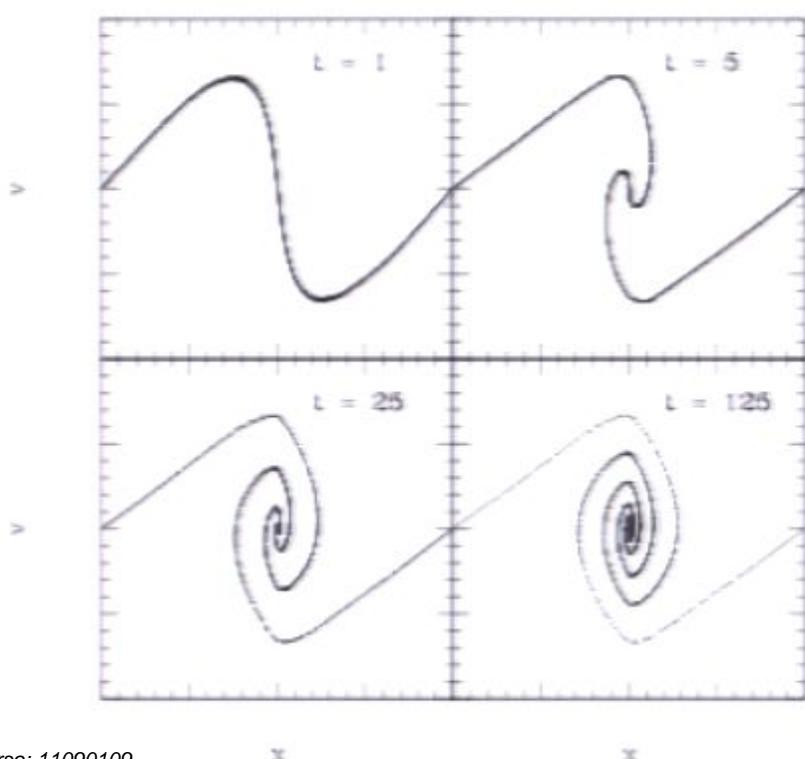
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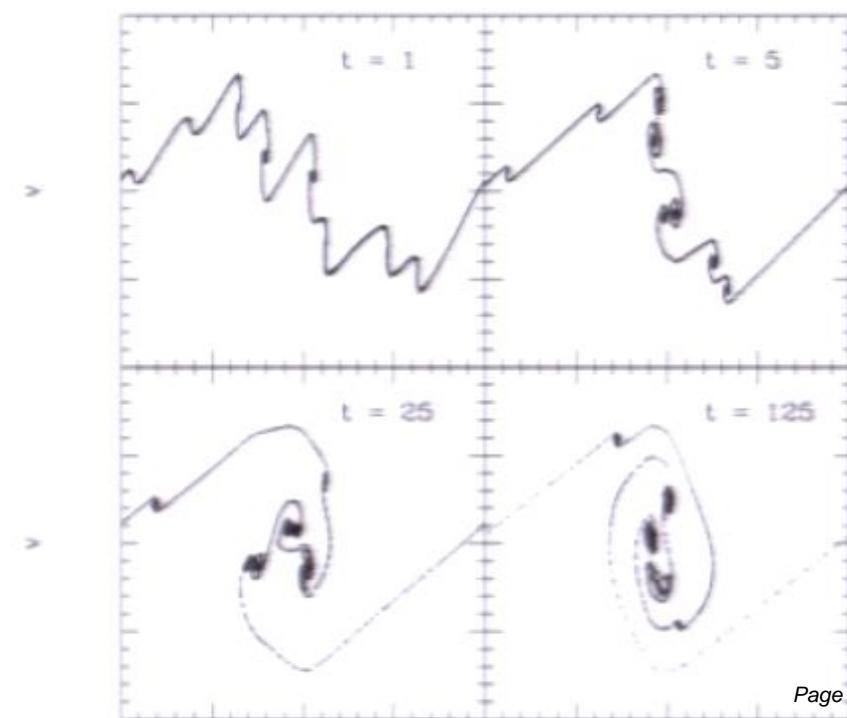
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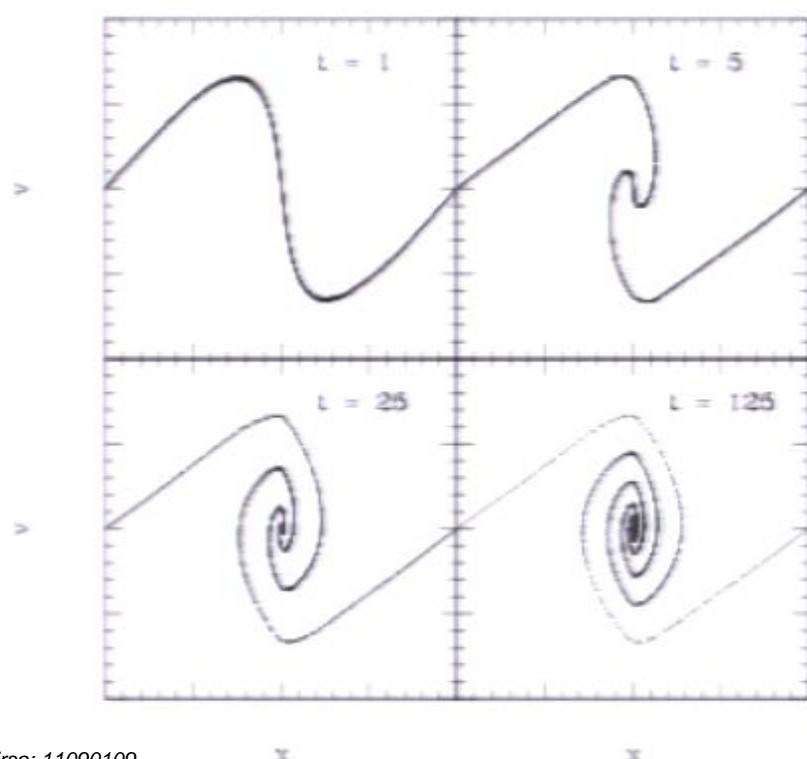
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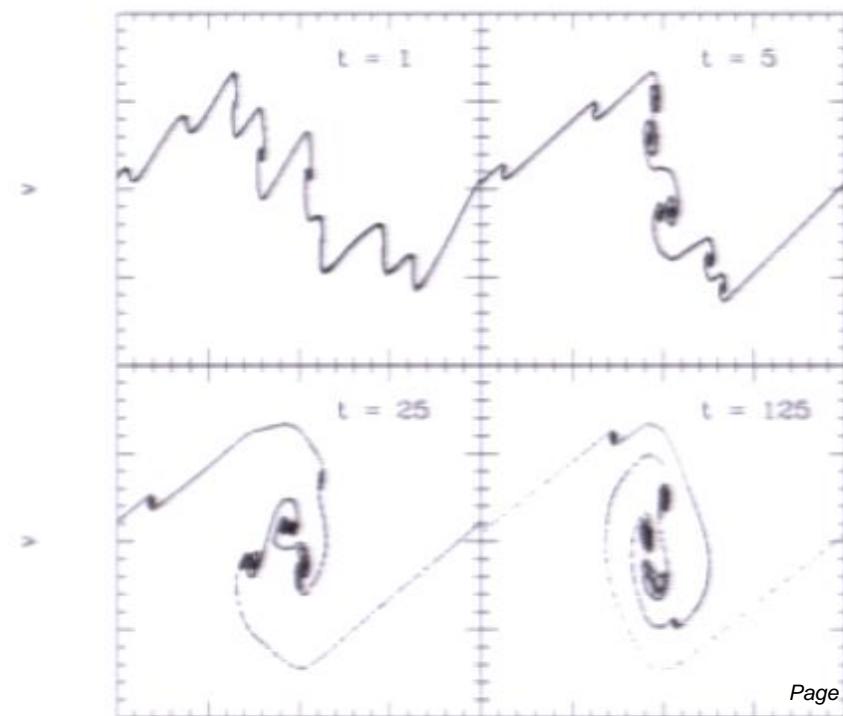
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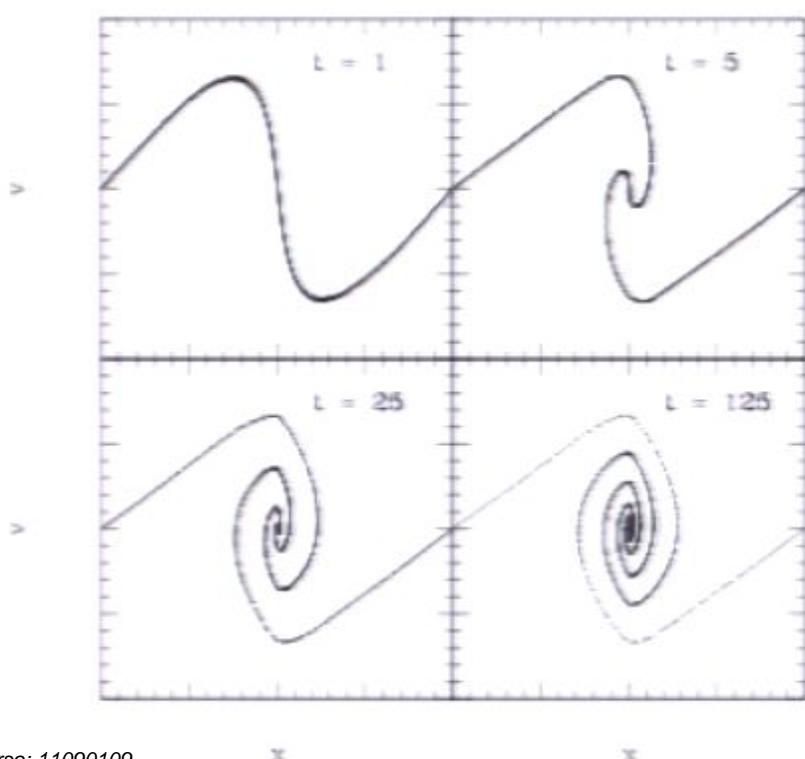
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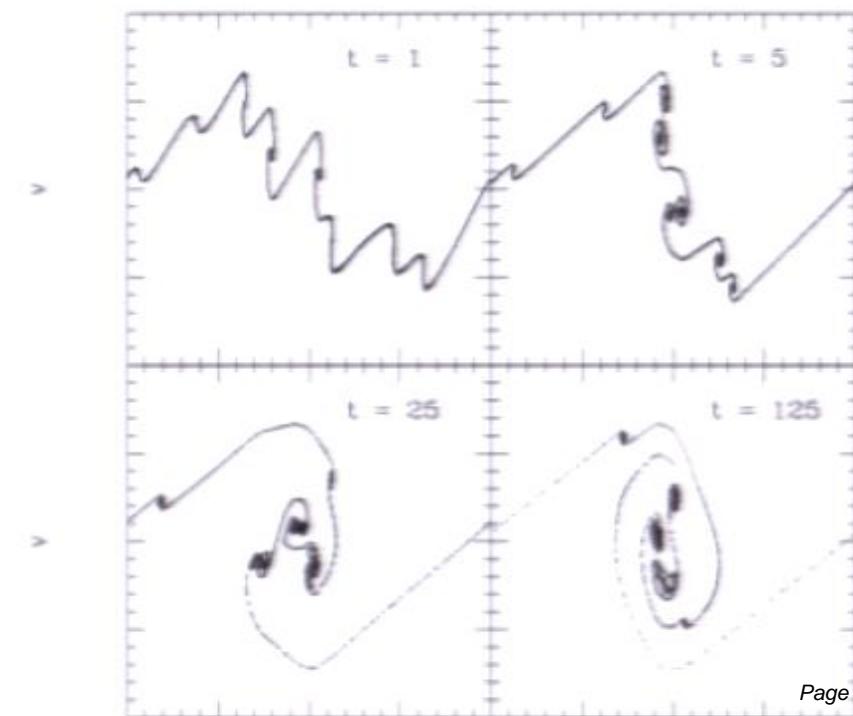
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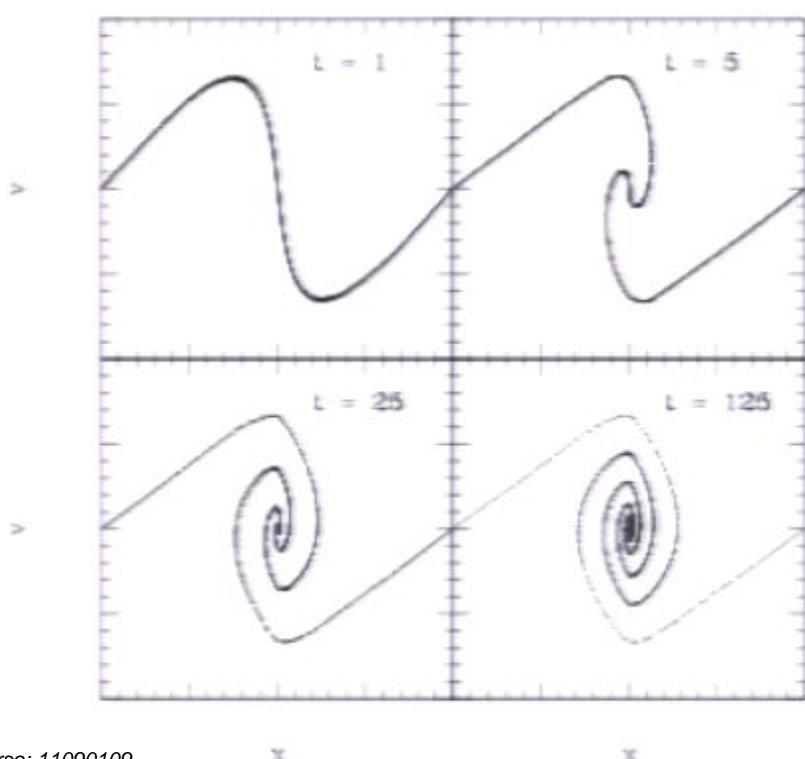
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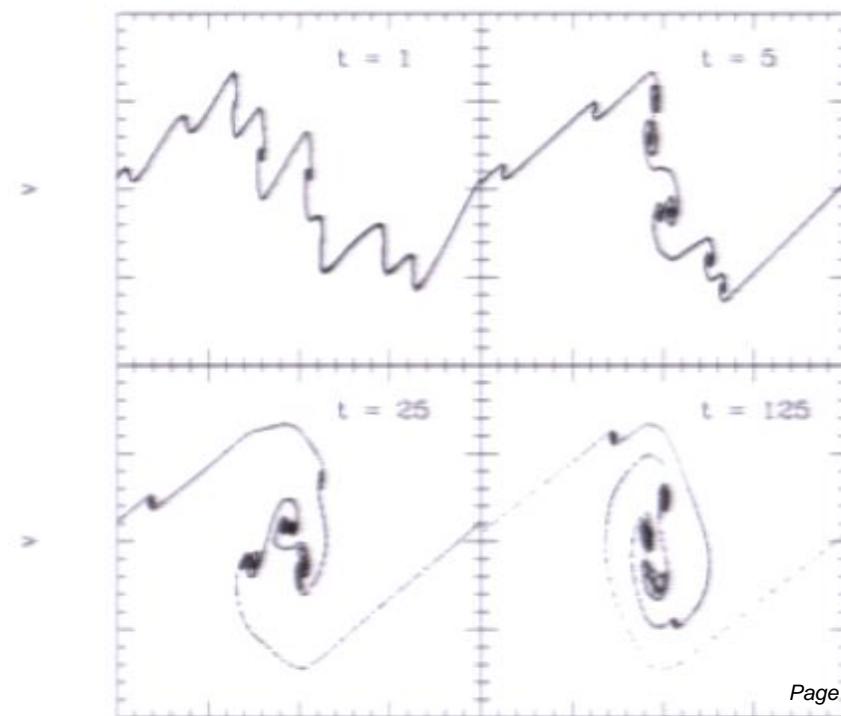
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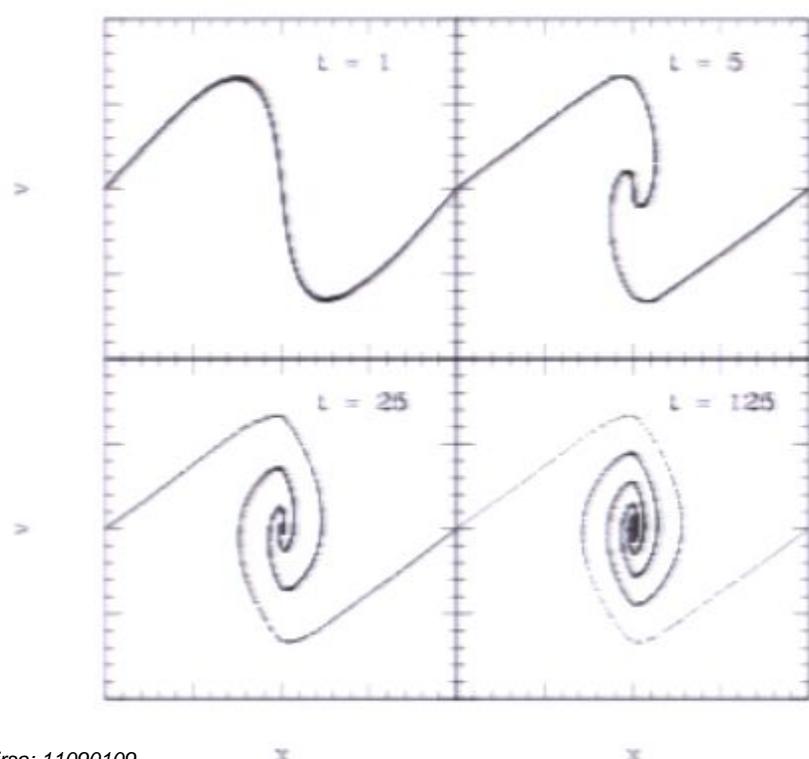
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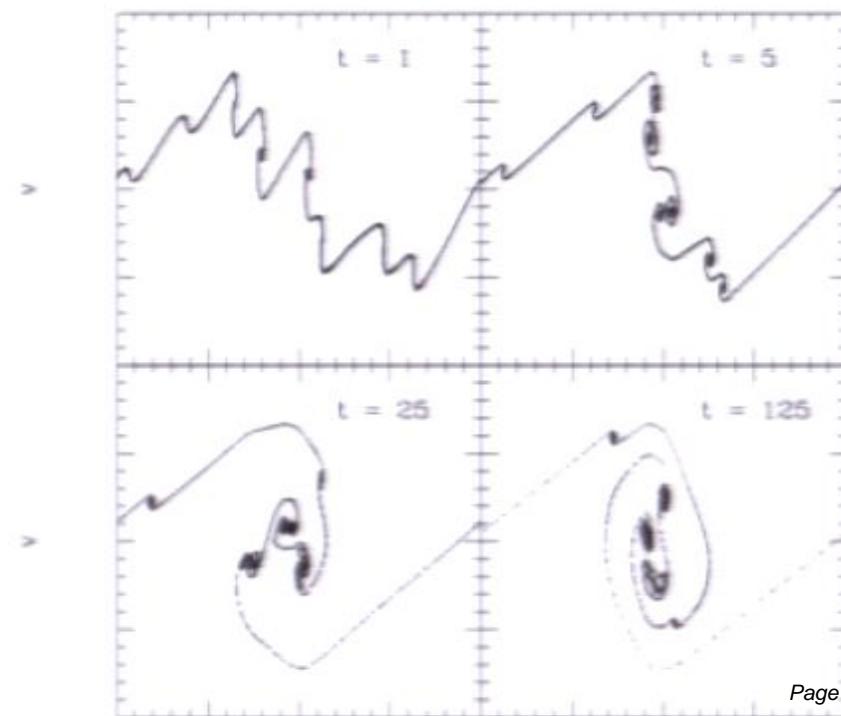
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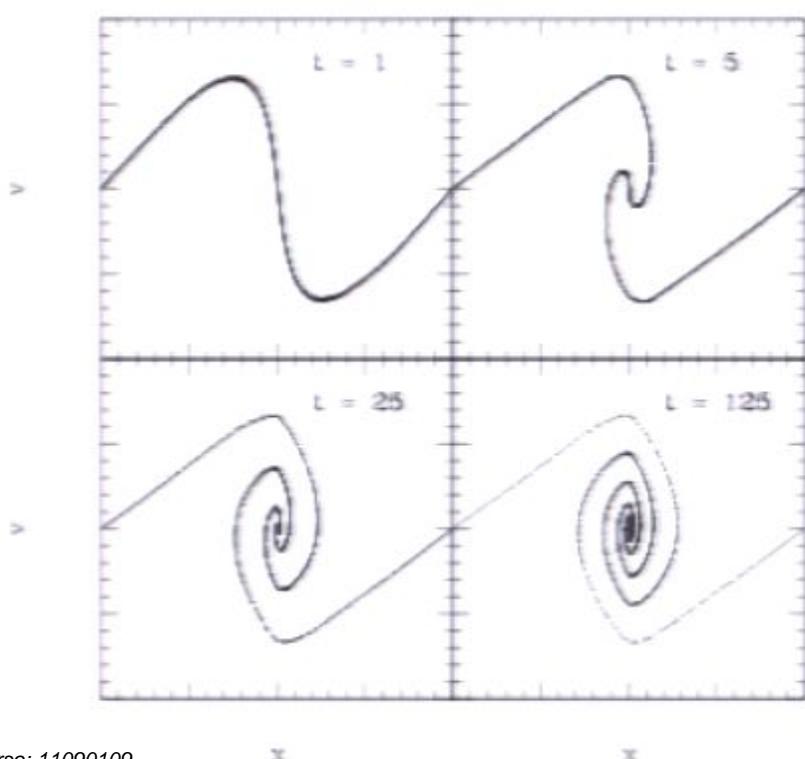
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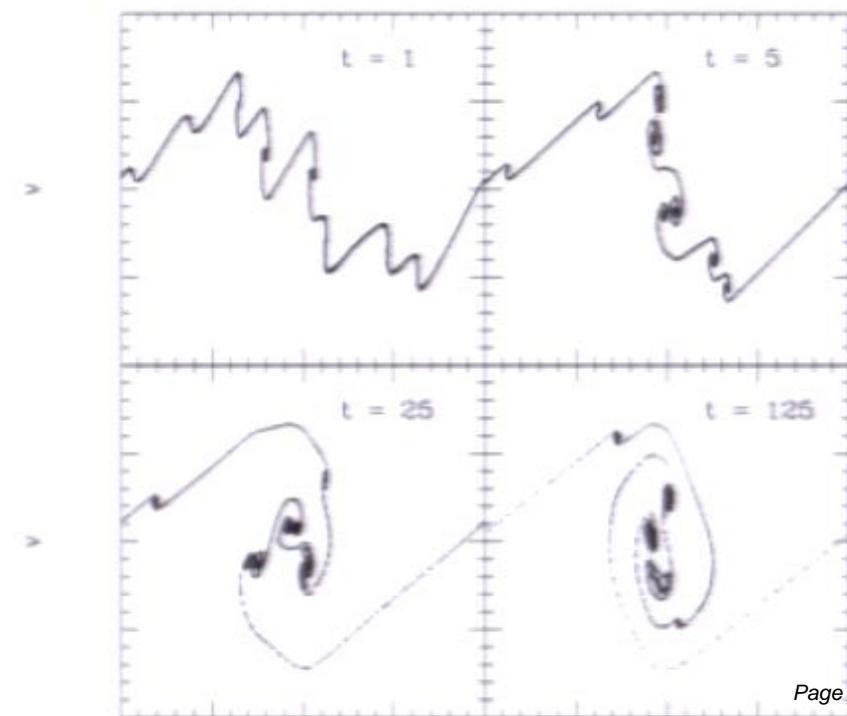
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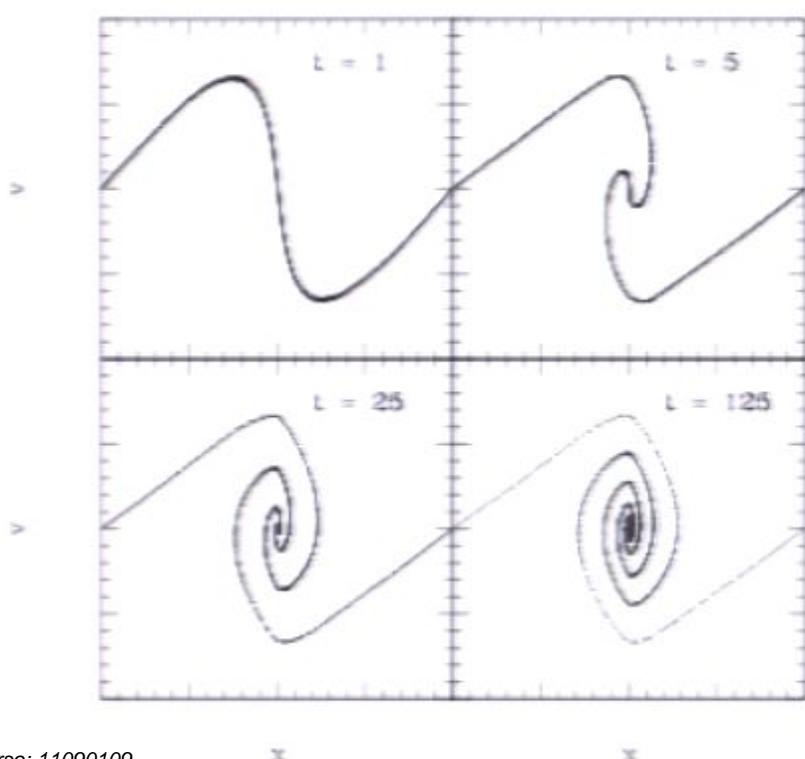
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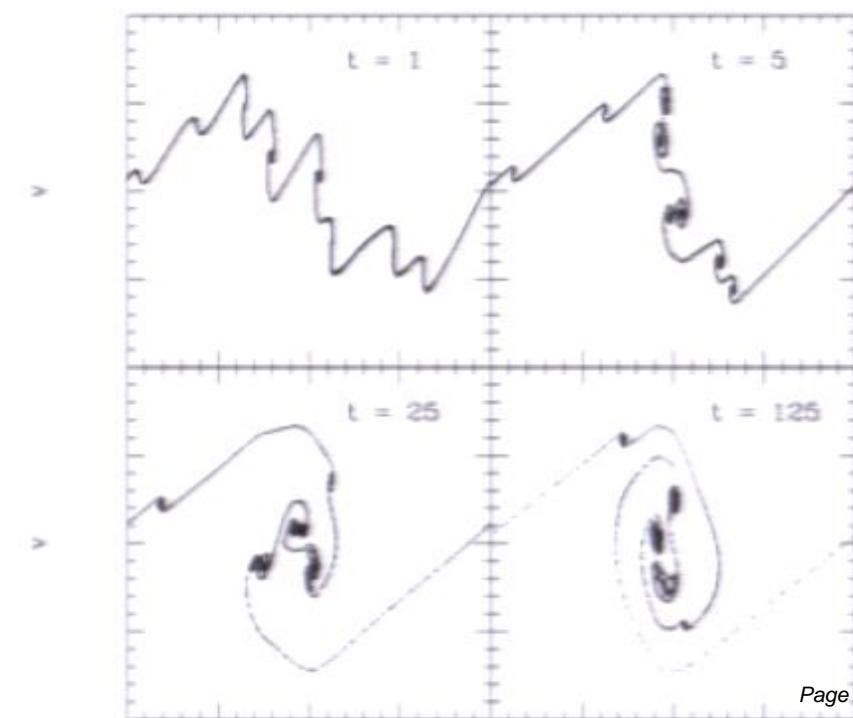
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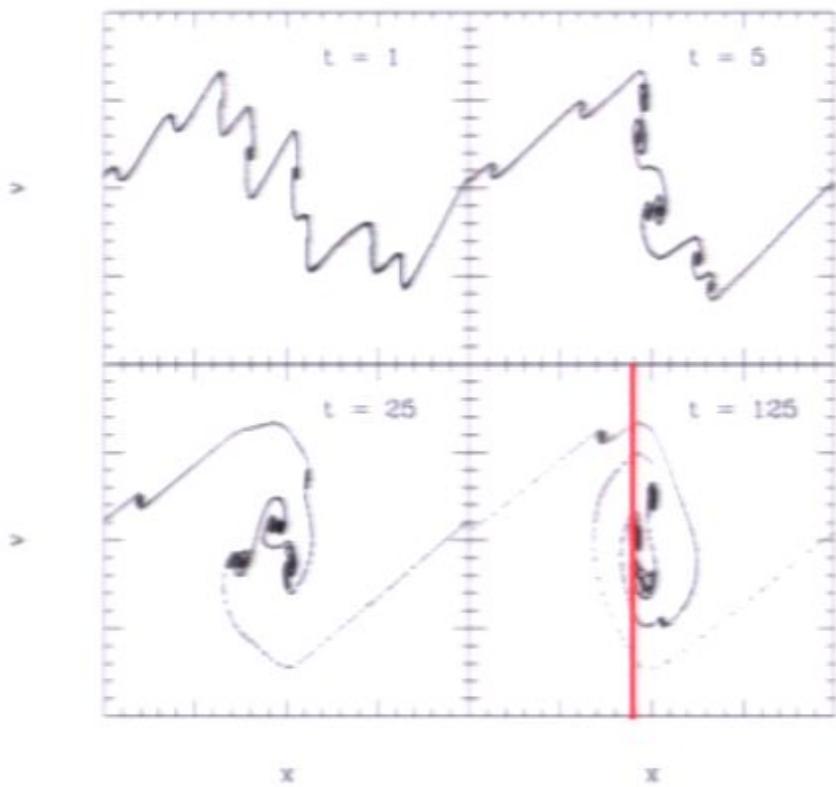
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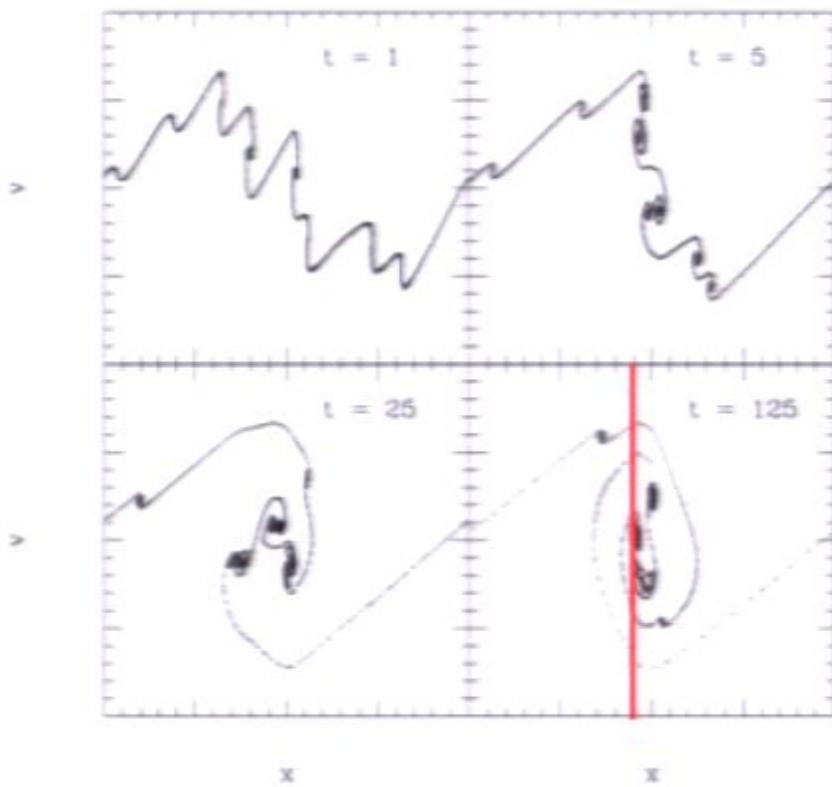
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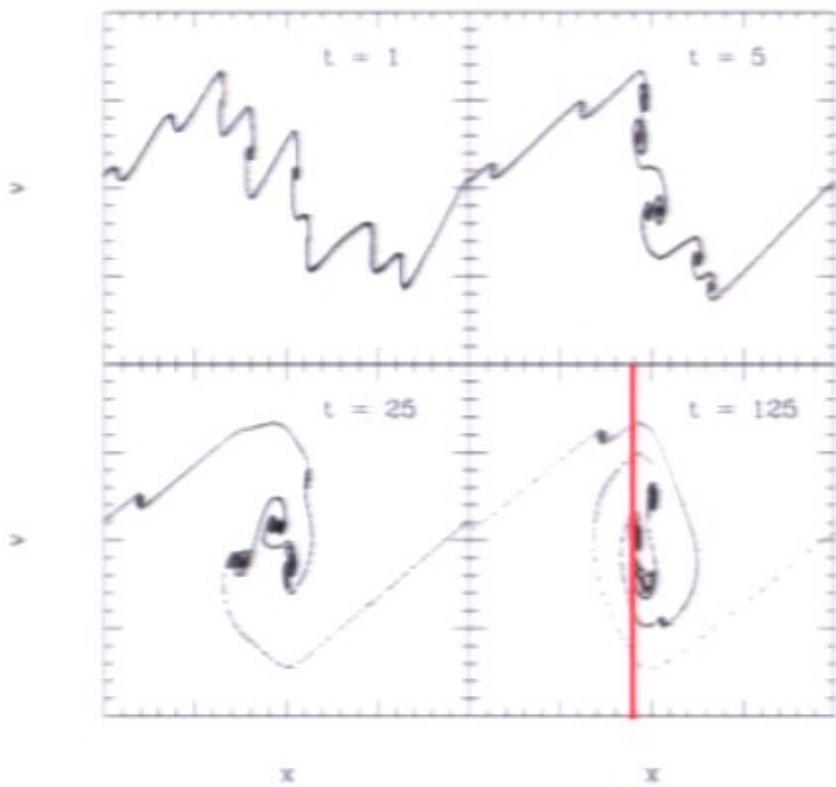
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“sees” particles with
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distribution (peaks and subpeaks)

ideal terrestrial detector occupies a 3D plane
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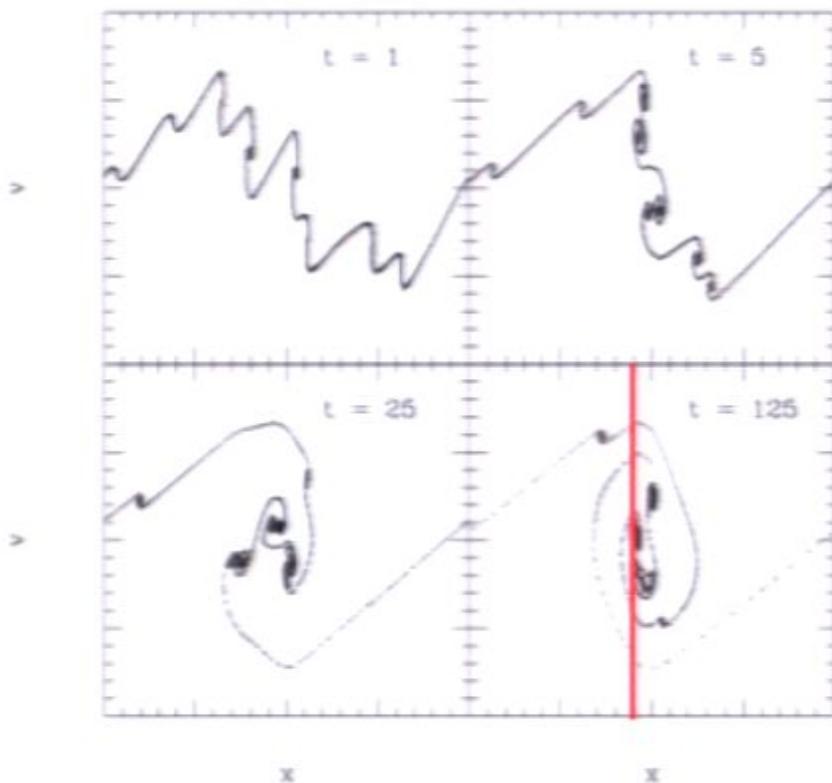
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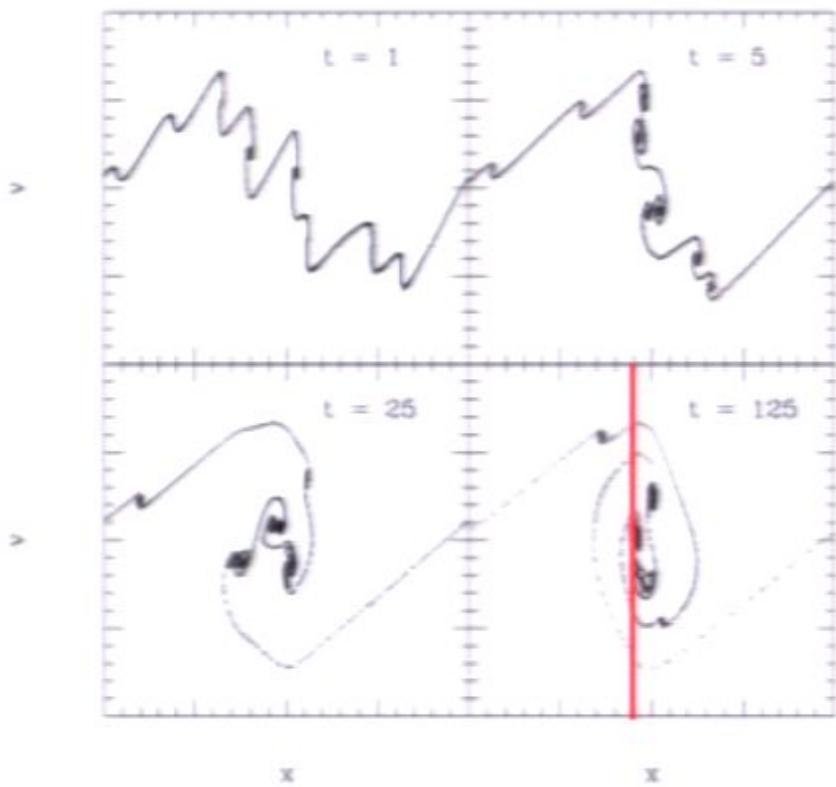
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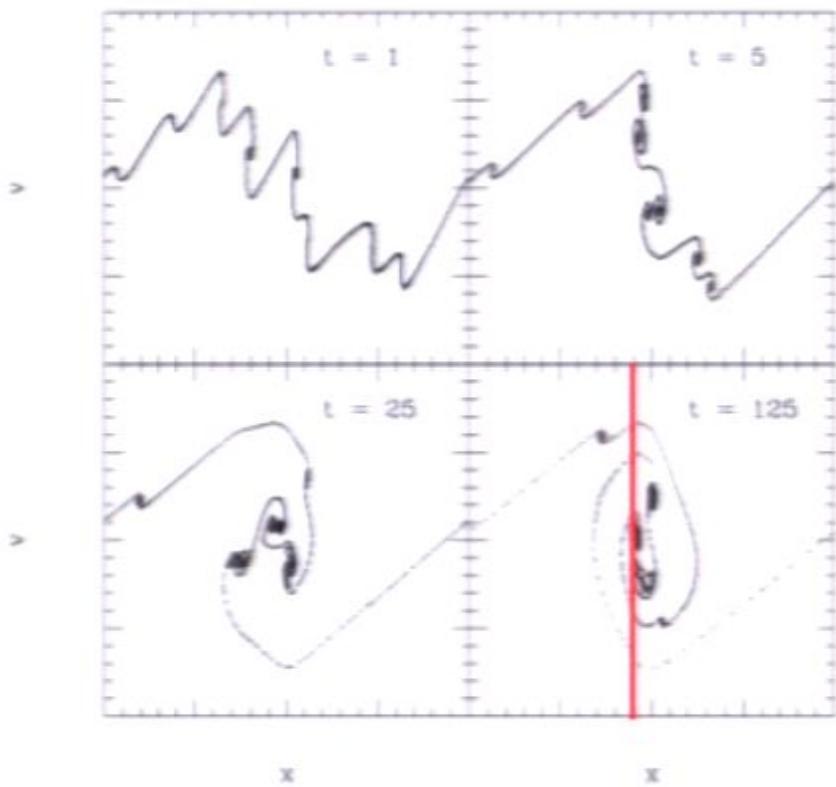
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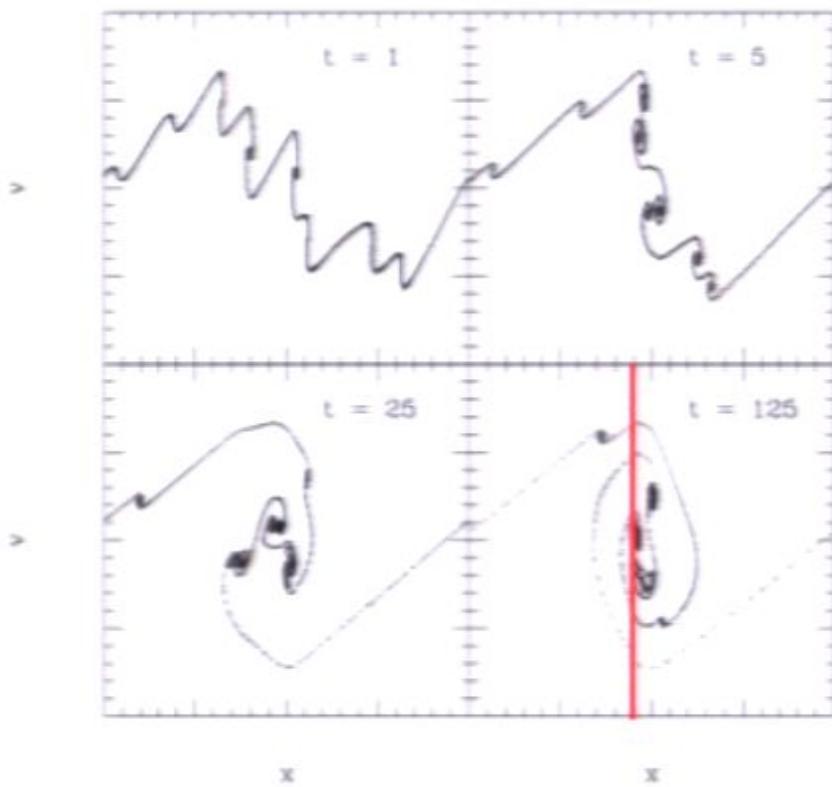
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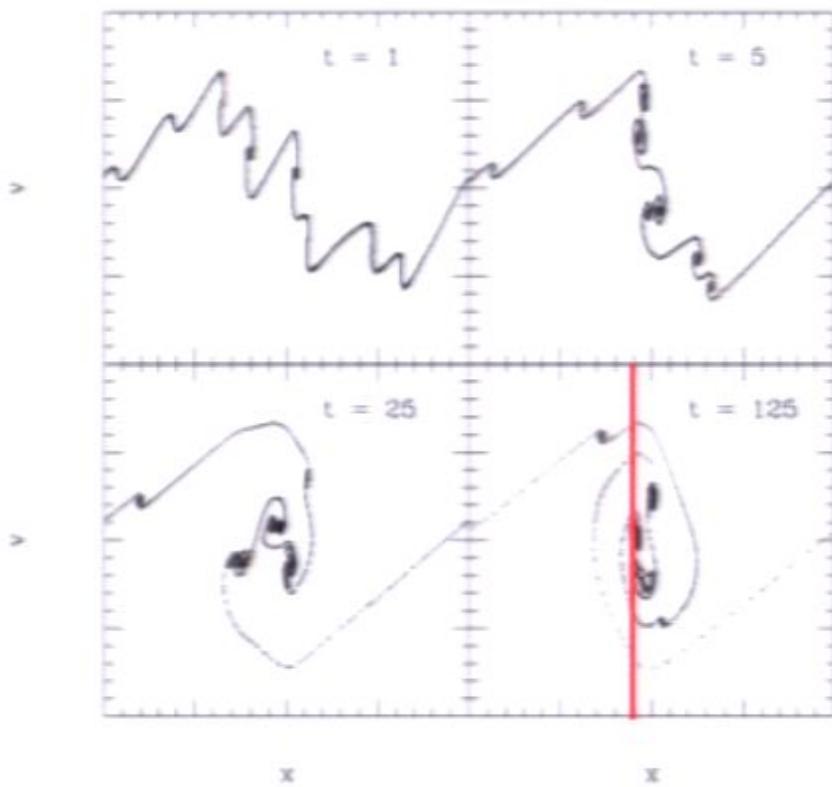
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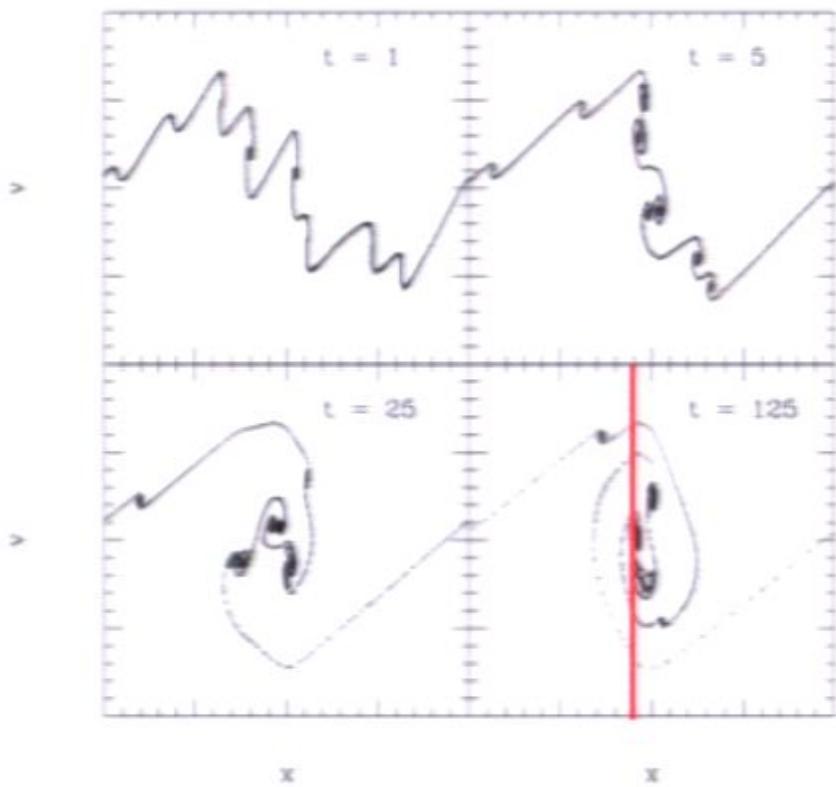
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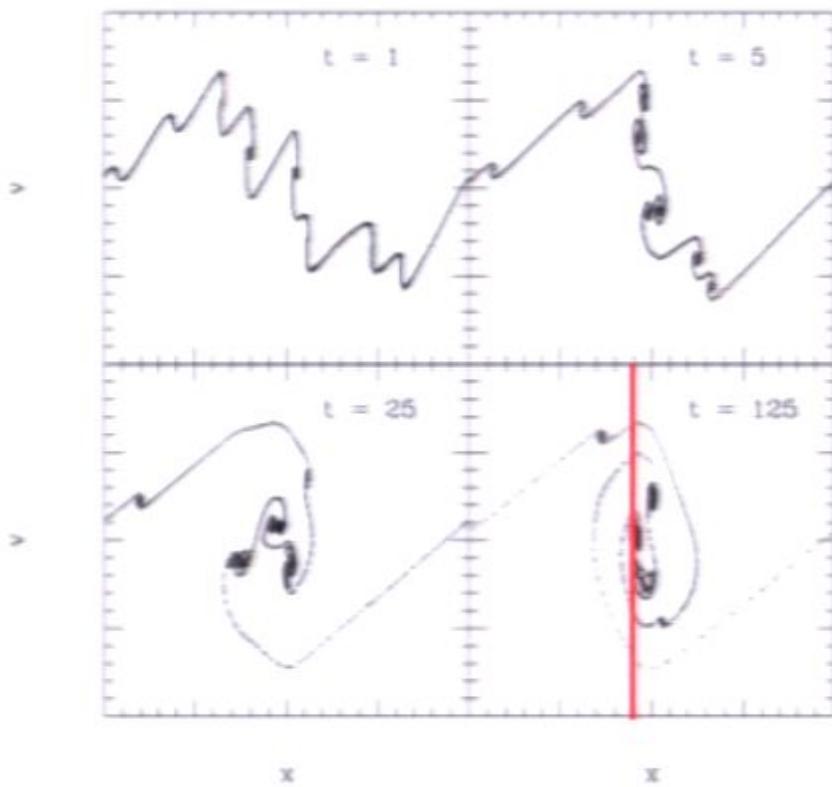
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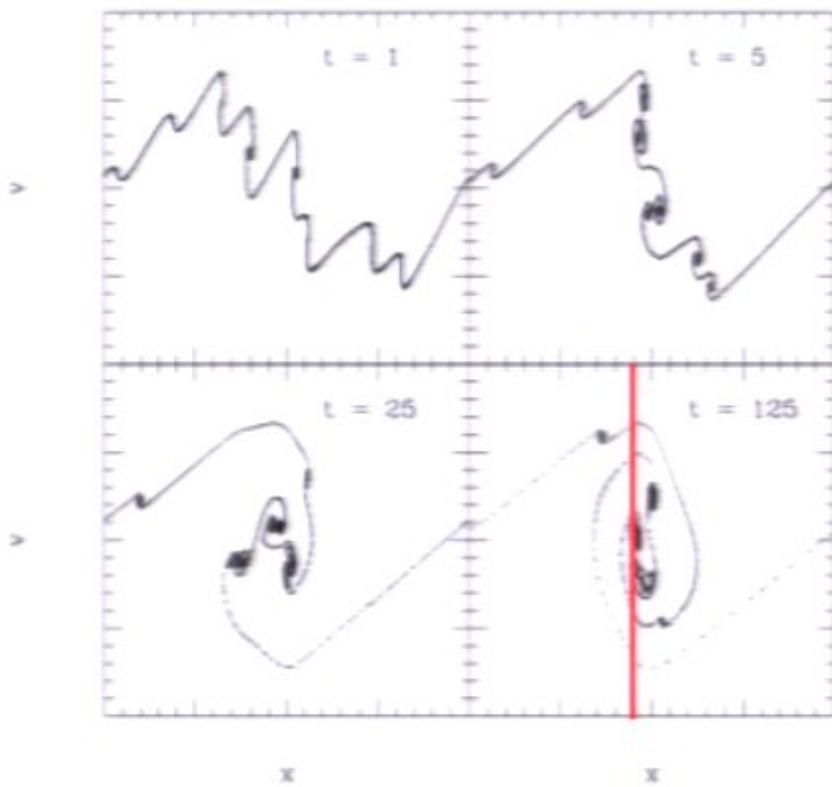
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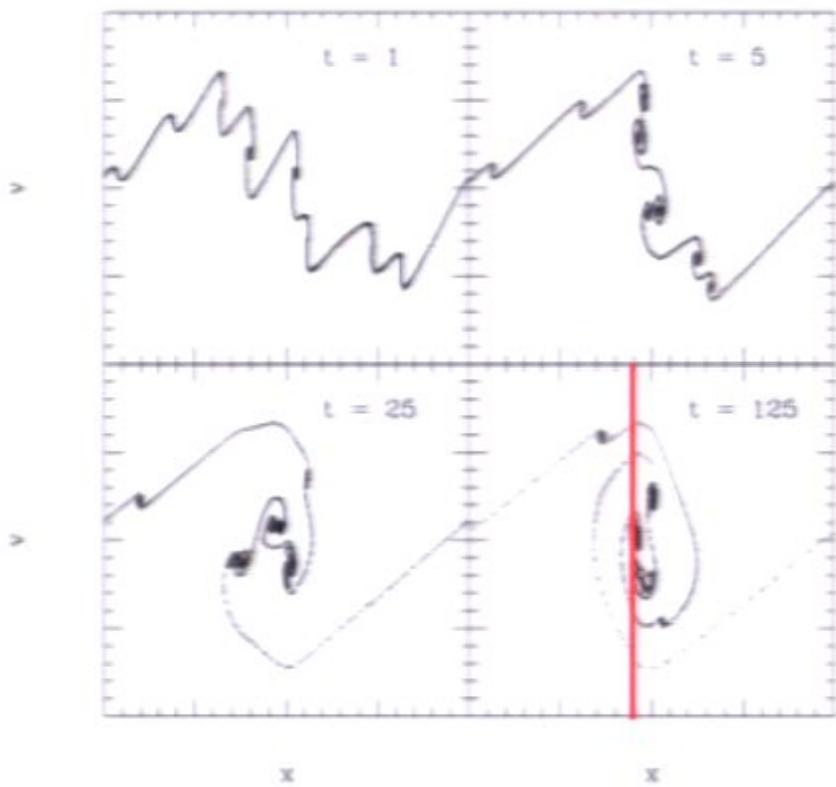
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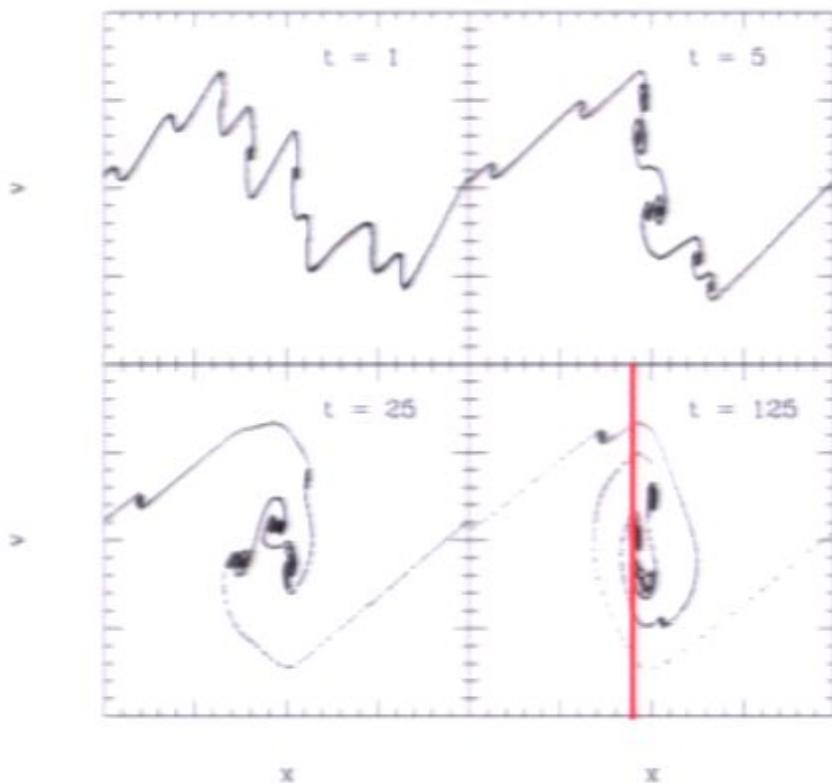
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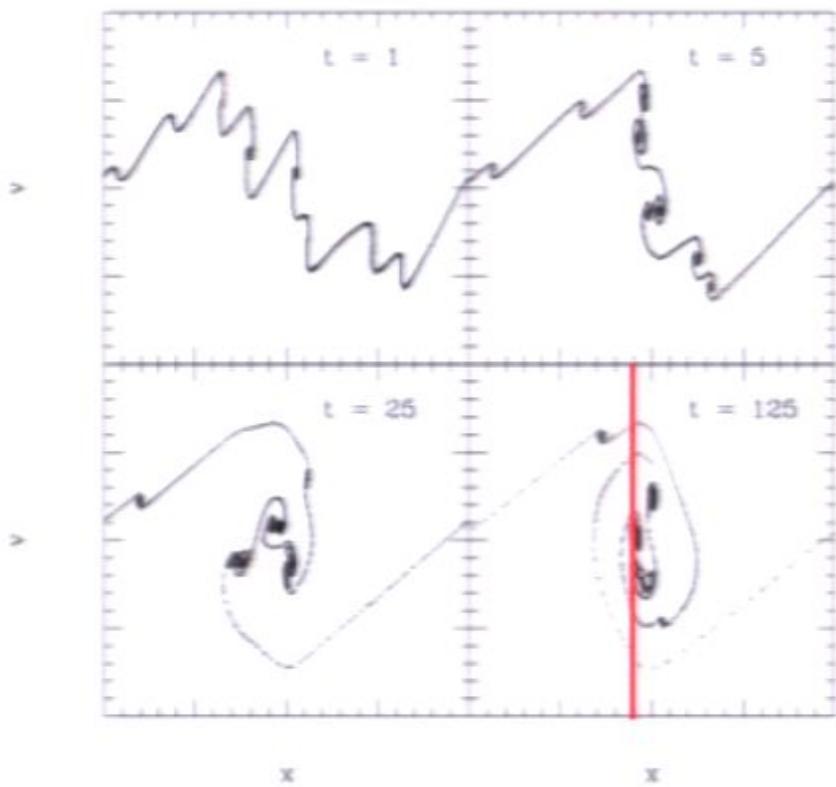
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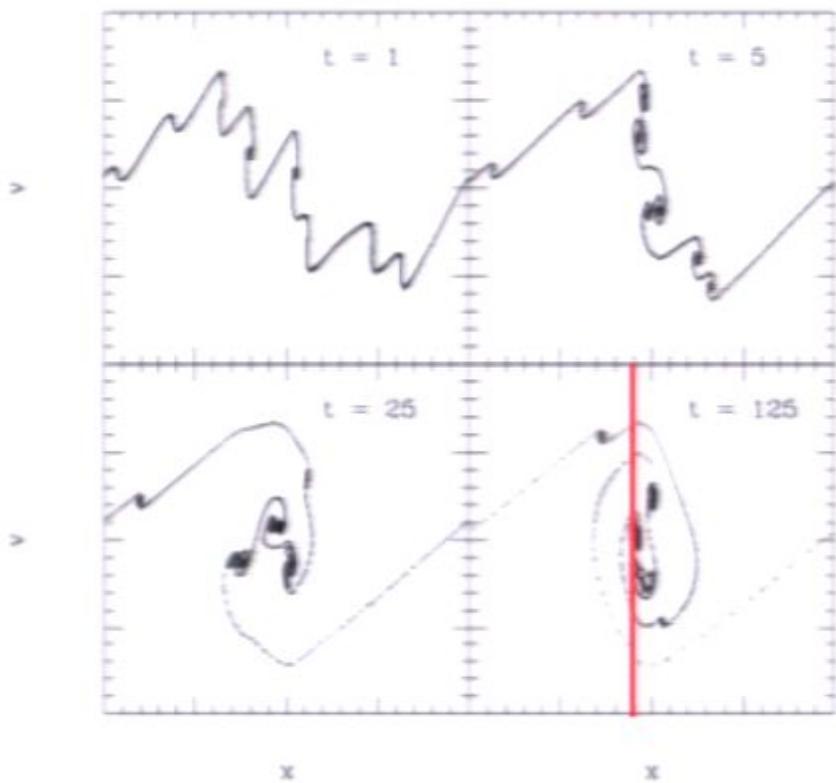
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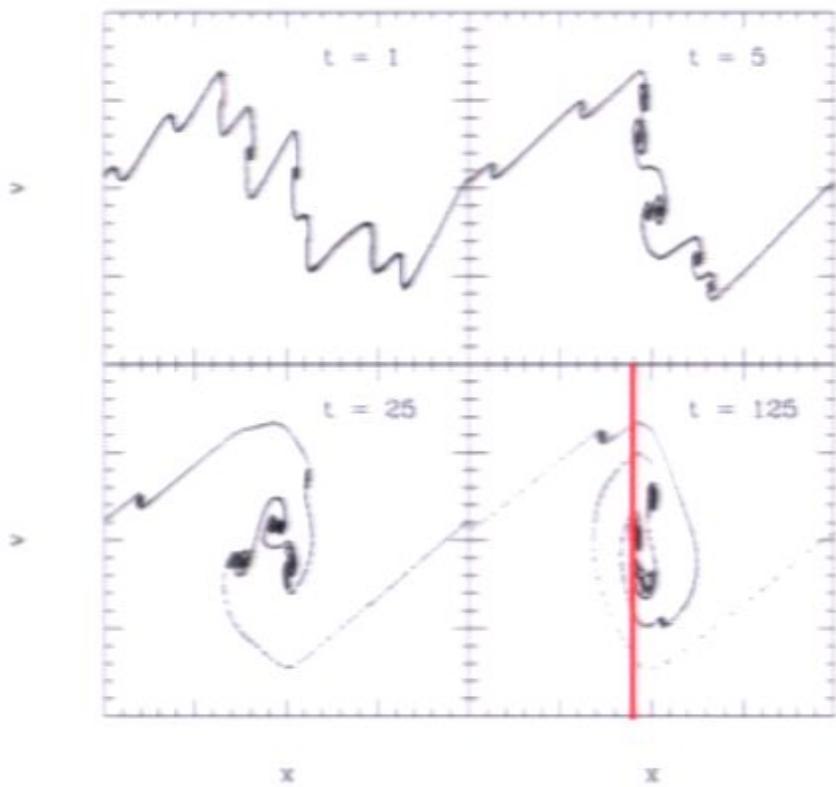
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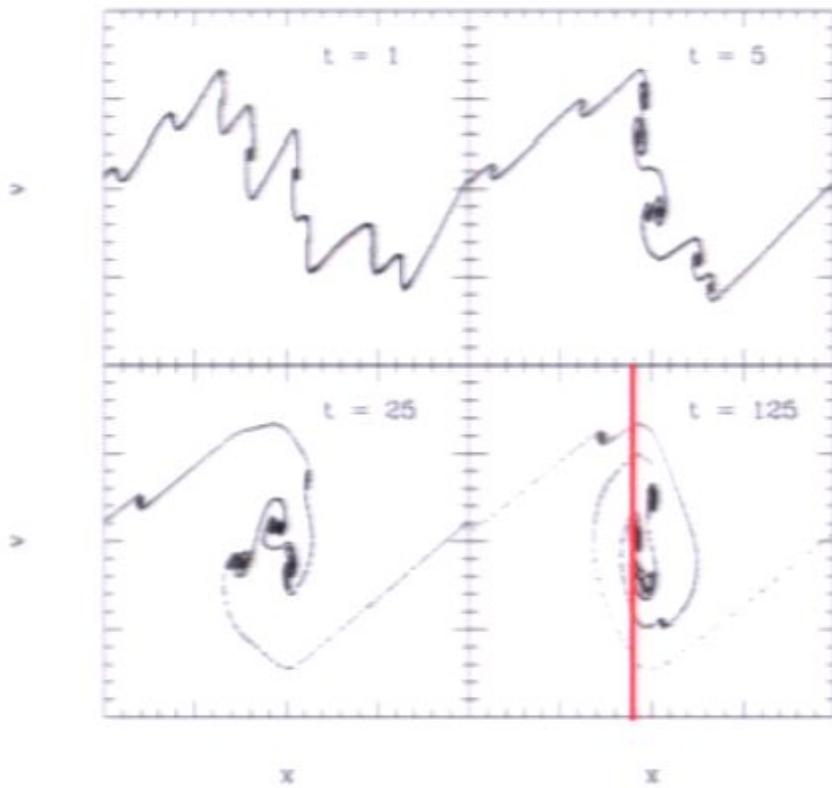
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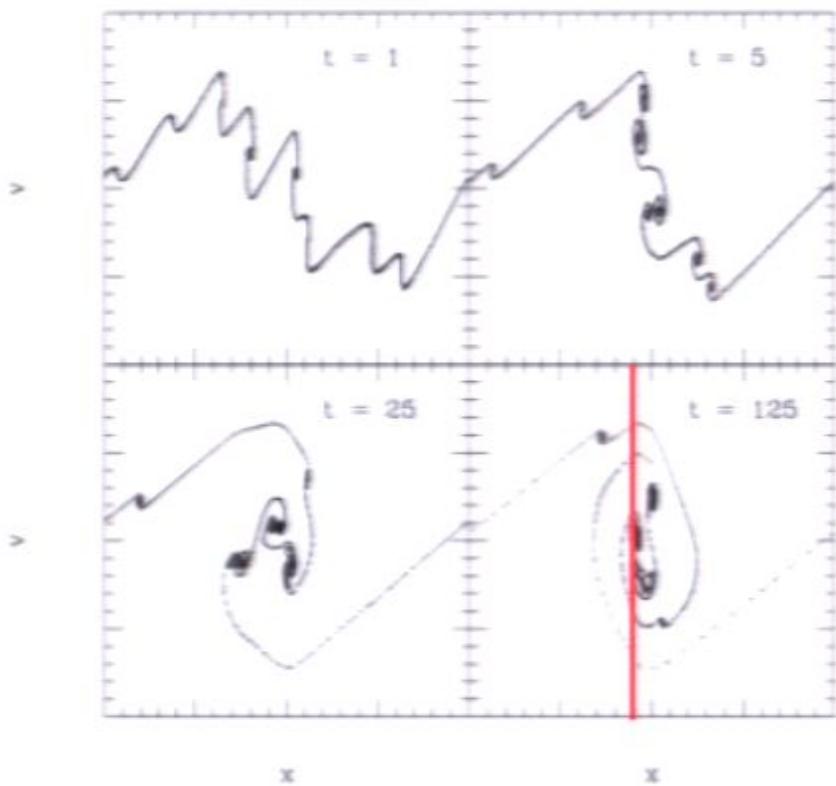
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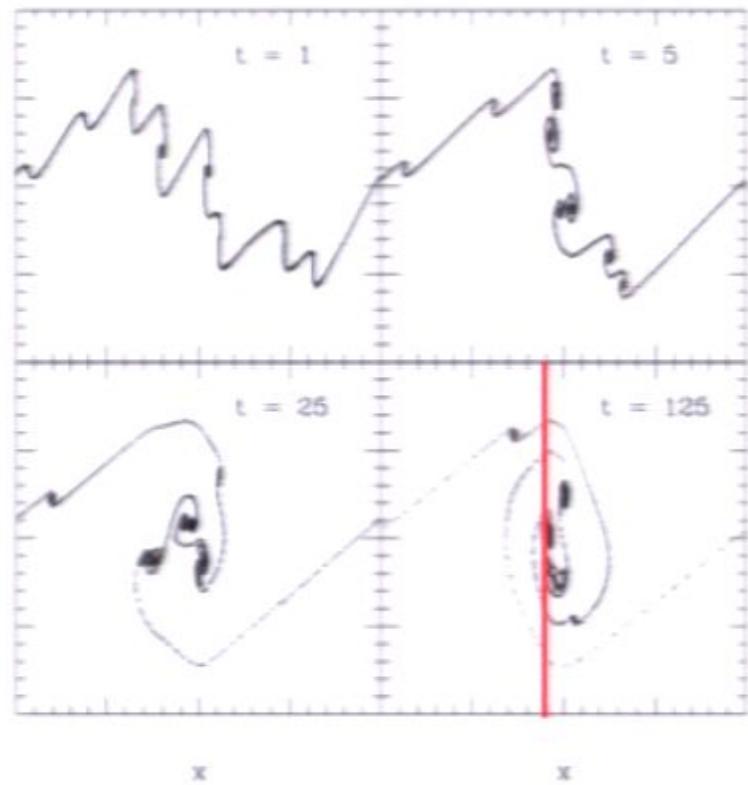
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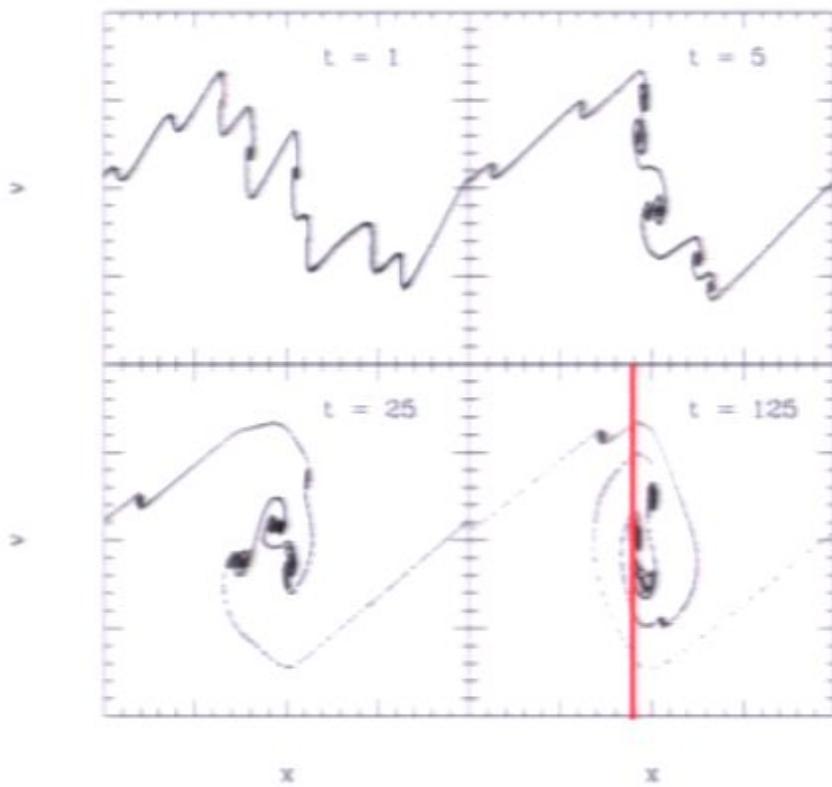
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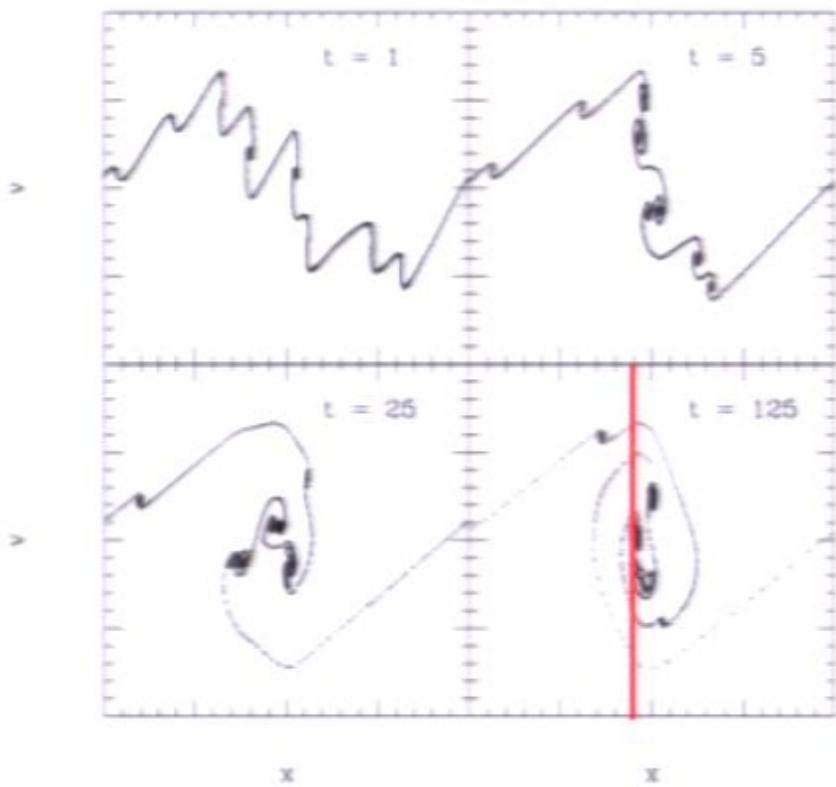
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Do N-body simulations correctly
evolve the phase space sheet?

If they did, can we properly visualize/study it?

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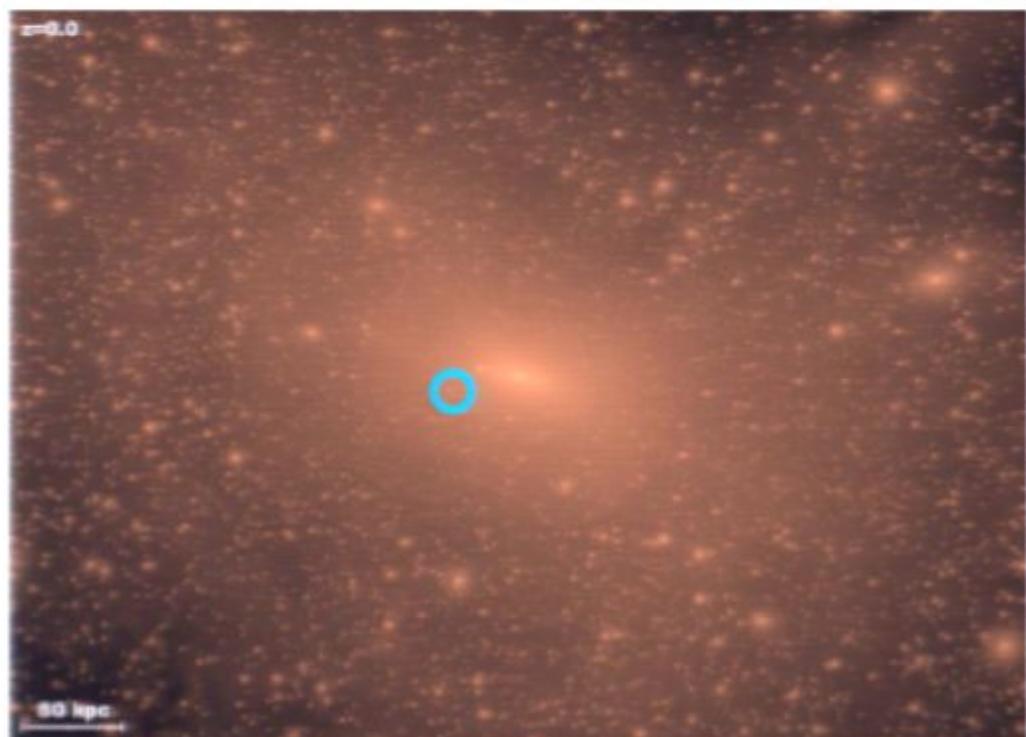
Use simulation to characterize

$$f(\mathbf{v}, \mathbf{x}_d)$$

see LMW & Dubinski 1998; Moore et al.
2001 and many others

Pick a region of the simulated
halo representative of the Earth.
Select particles in this region and
compute $f(\mathbf{v})$.

From $f(\mathbf{v})$, compute dR/dQ



“Uncertainty principle” for phase space:
Larger volumes give better velocity-space resolution
but worse spatial resolution, and vice versa

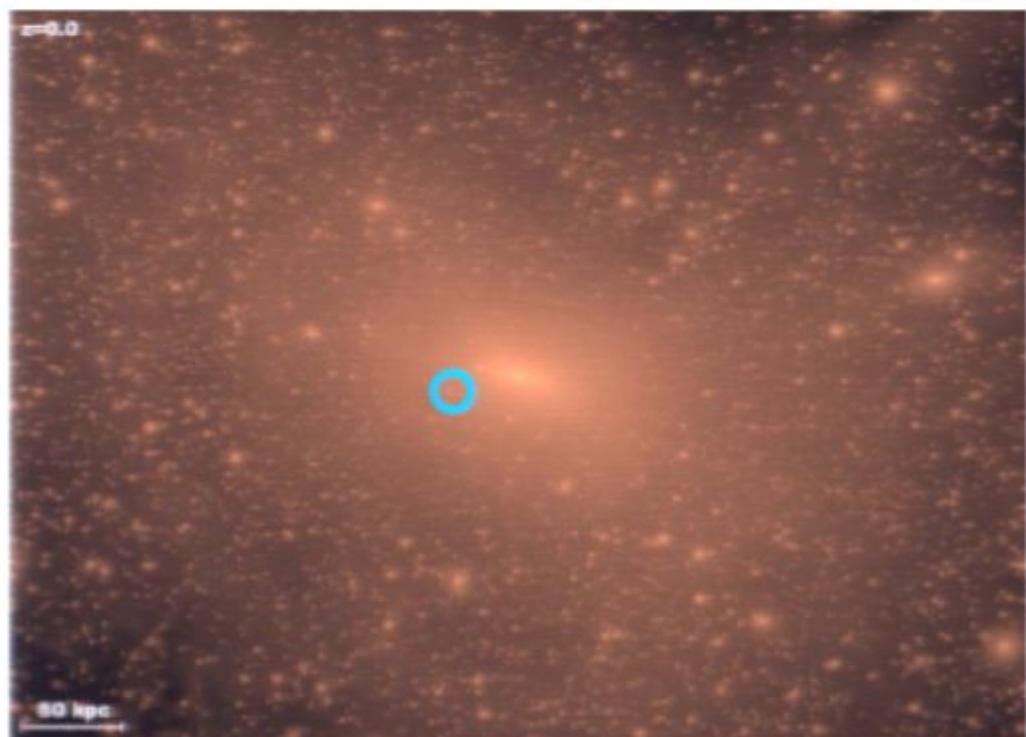
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Select particles in this region and
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From $f(\mathbf{v})$, compute dR/dQ



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Larger volumes give better velocity-space resolution
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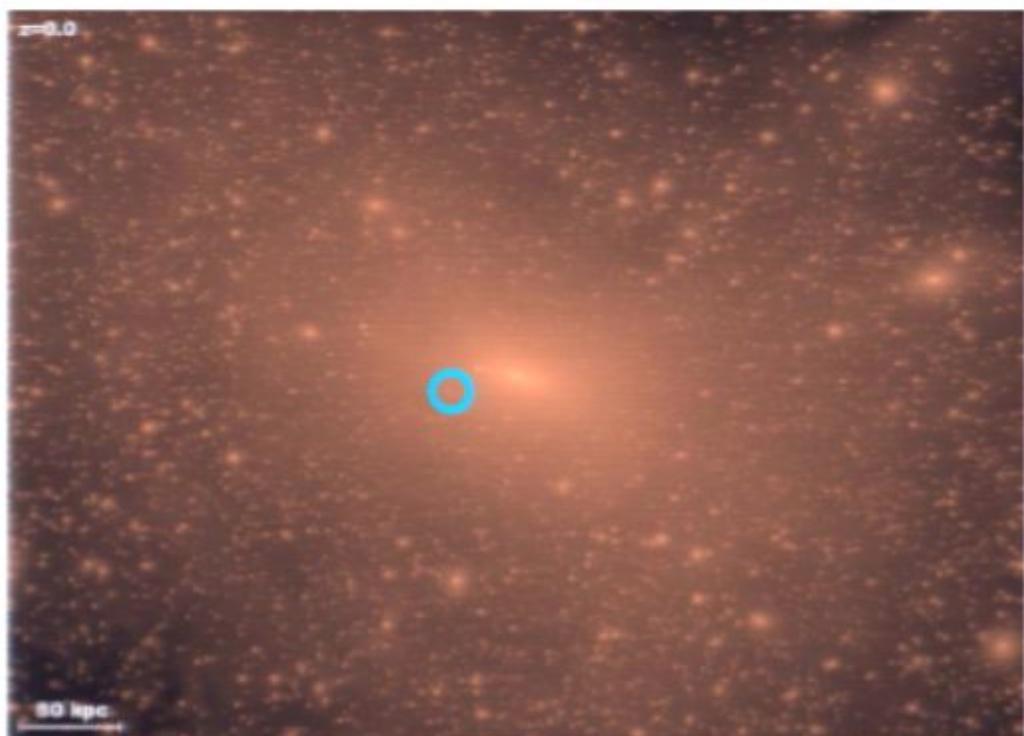
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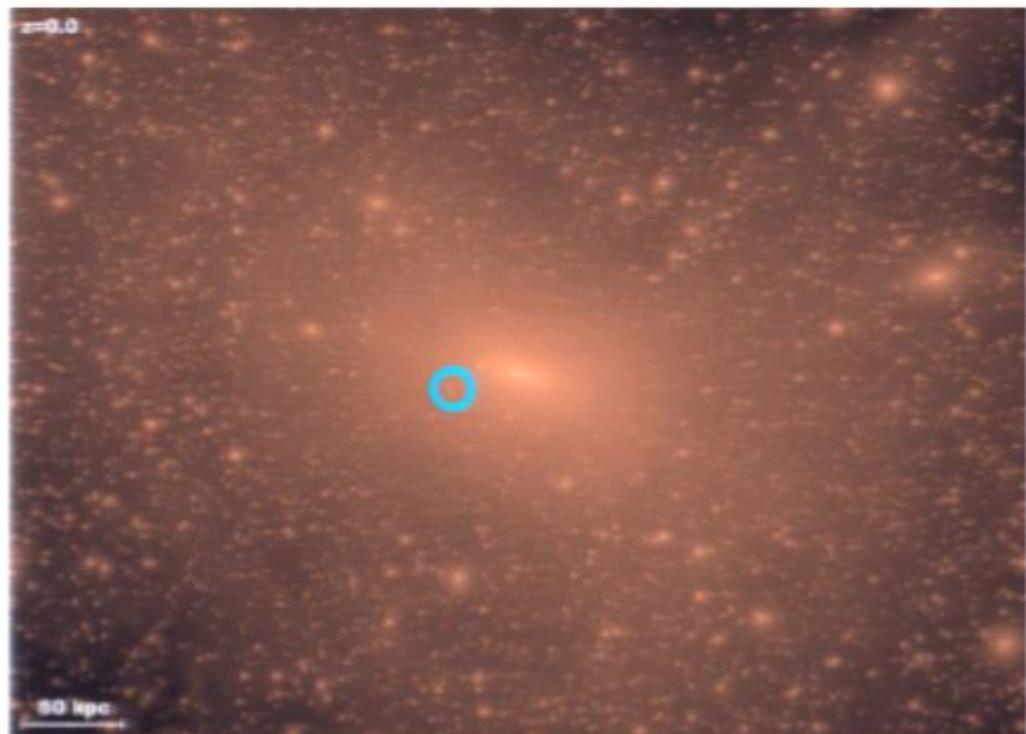
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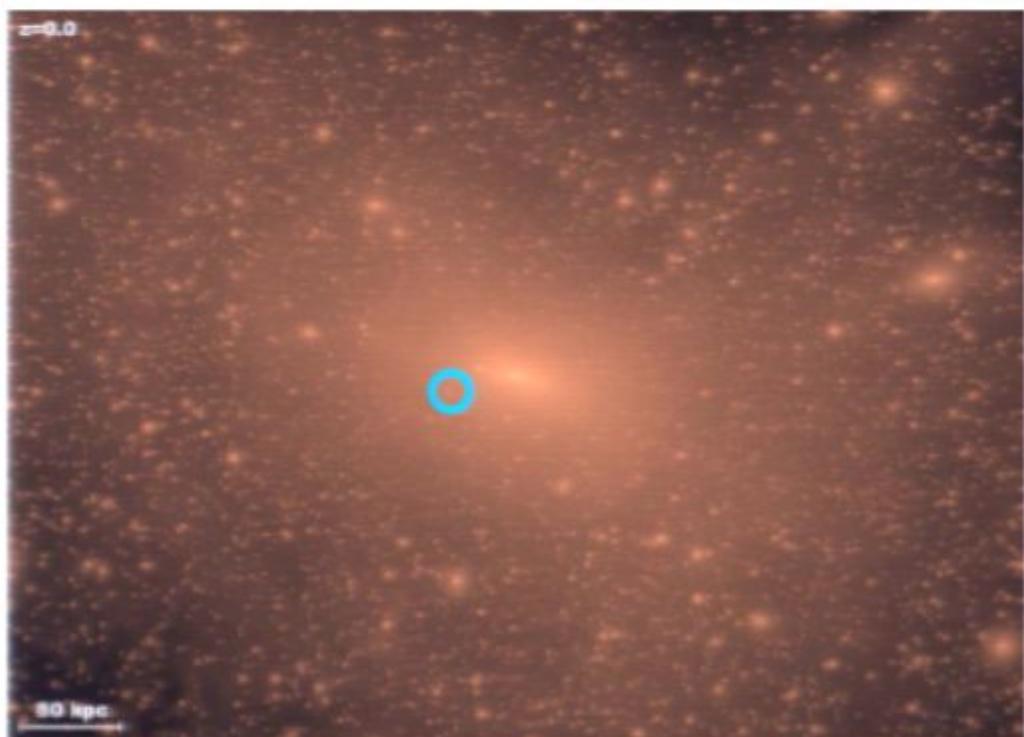
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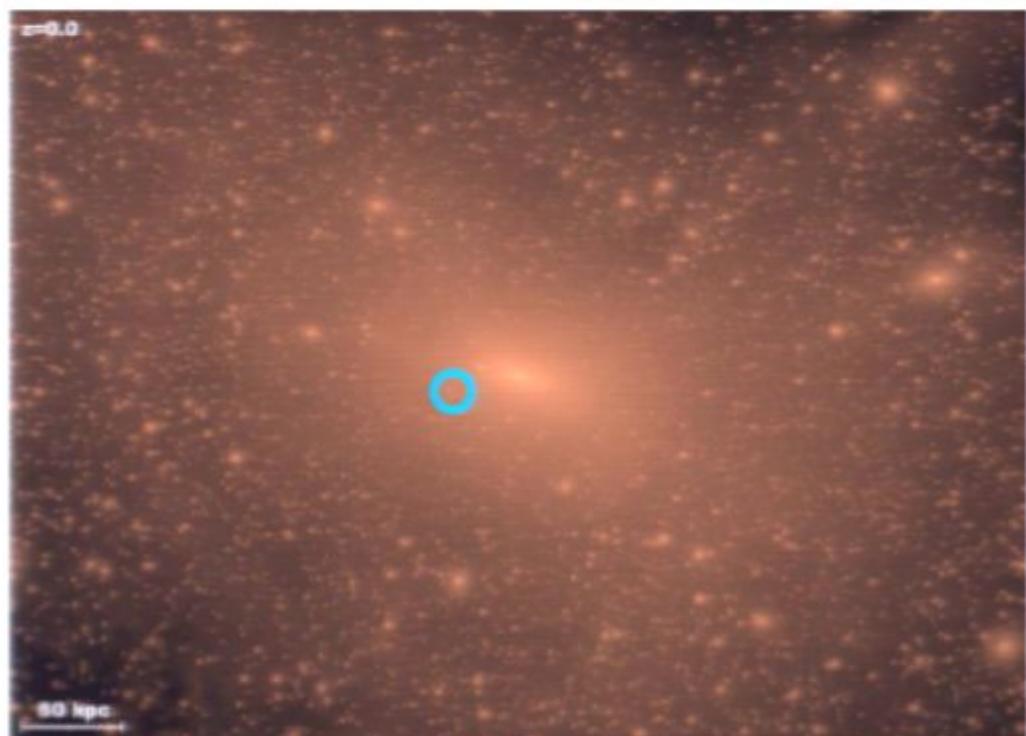
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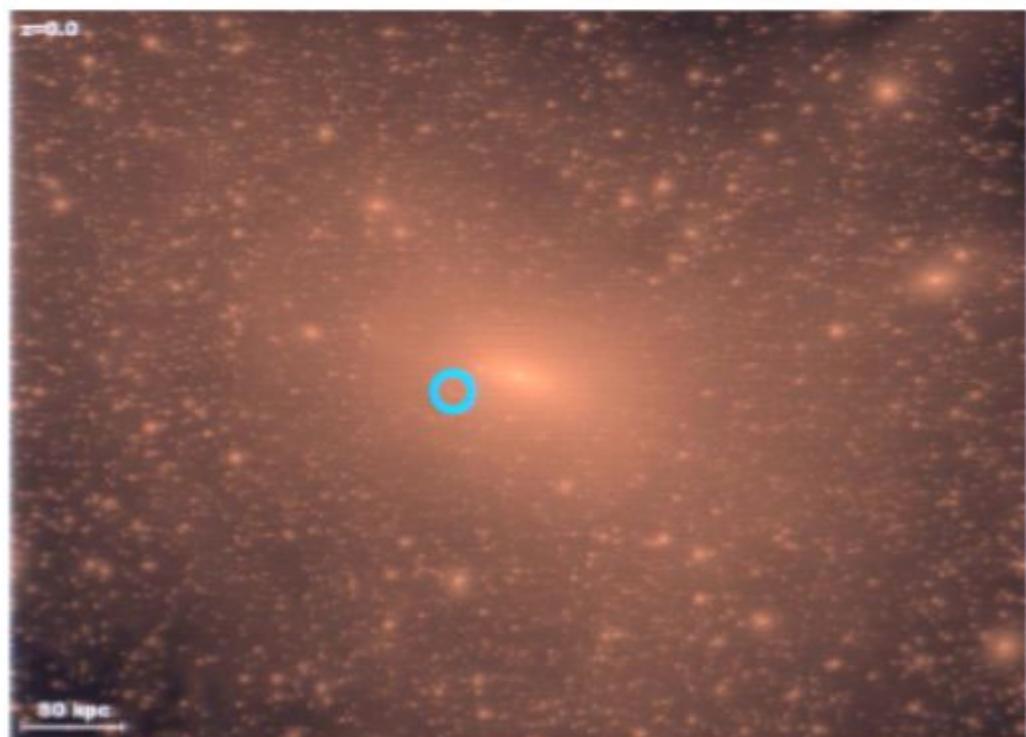
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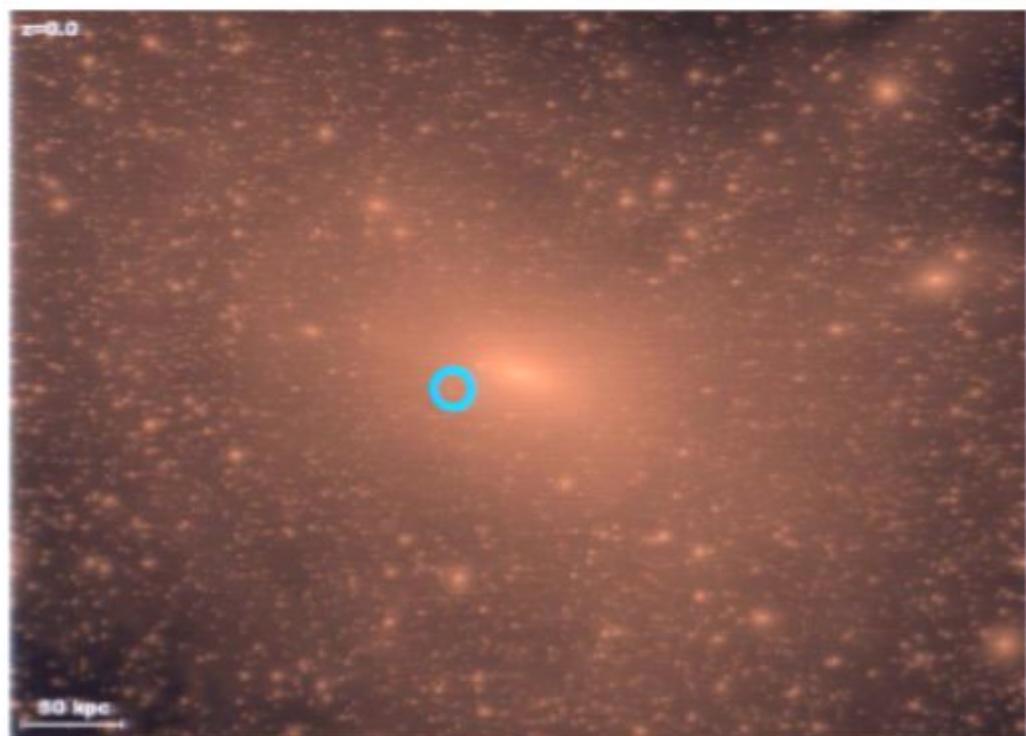
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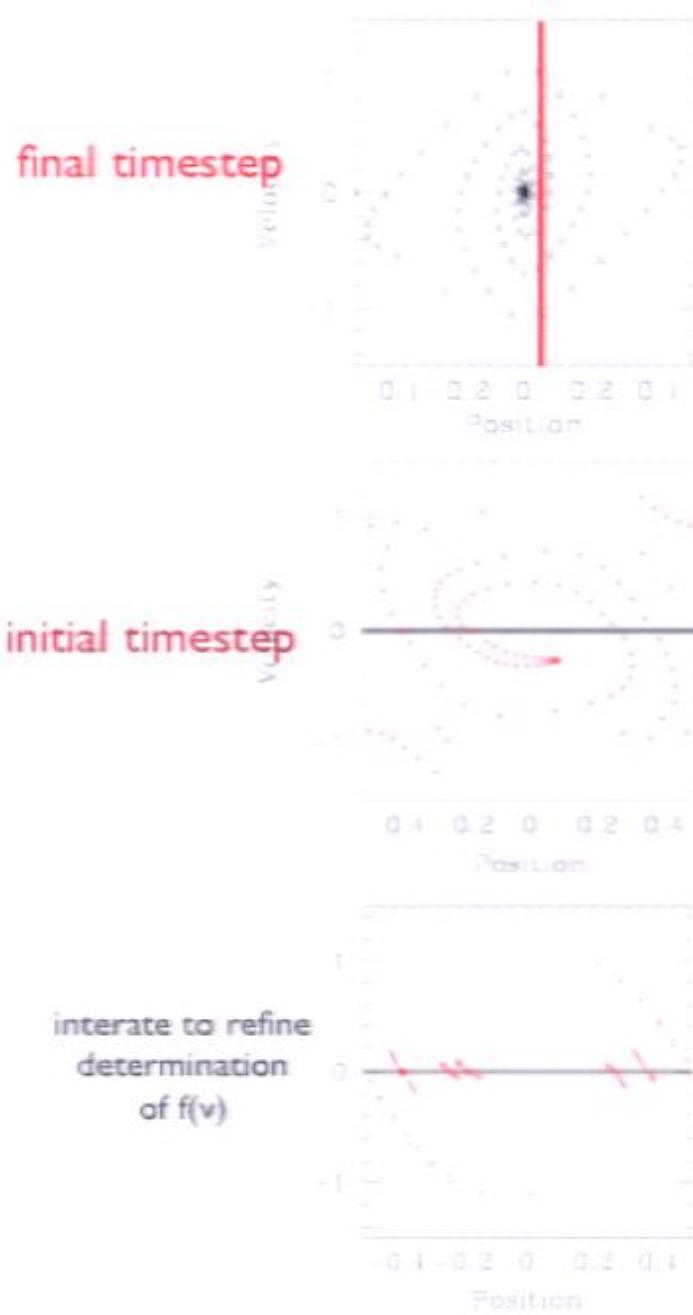
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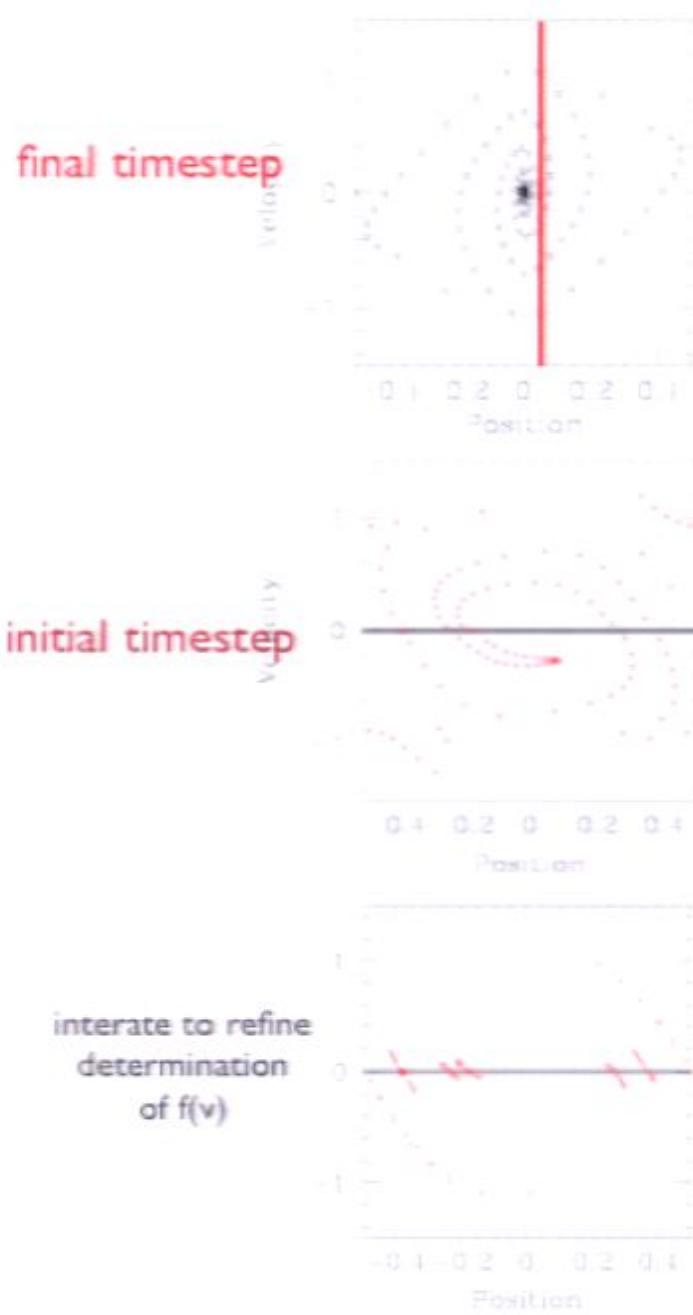
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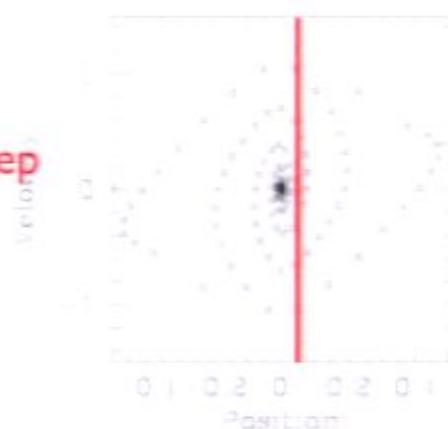
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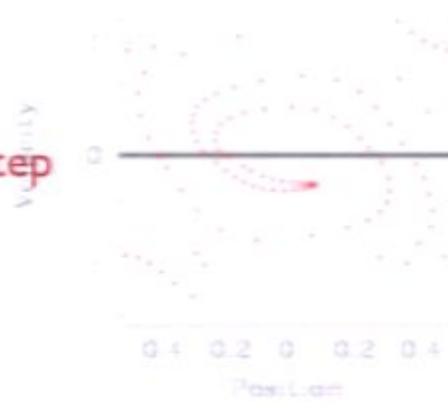
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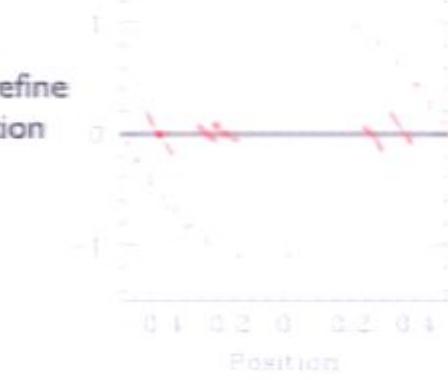
final timestep



initial timestep



iterate to refine determination of $f(v)$



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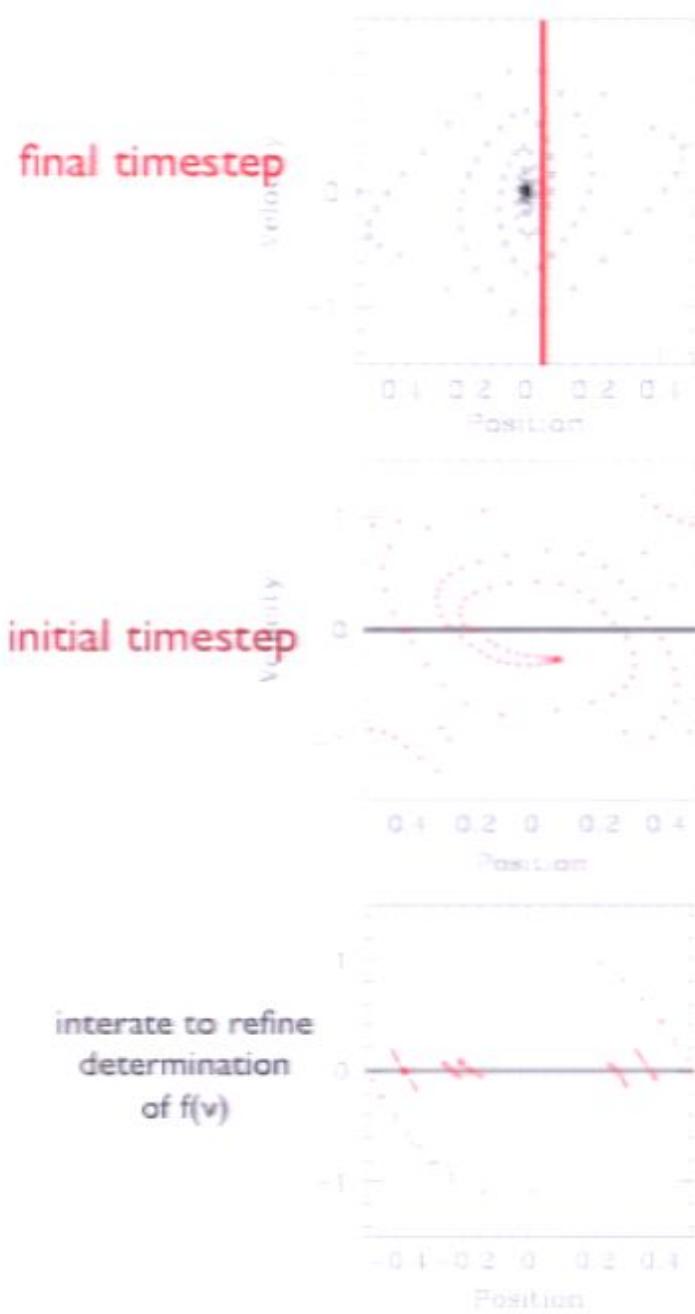
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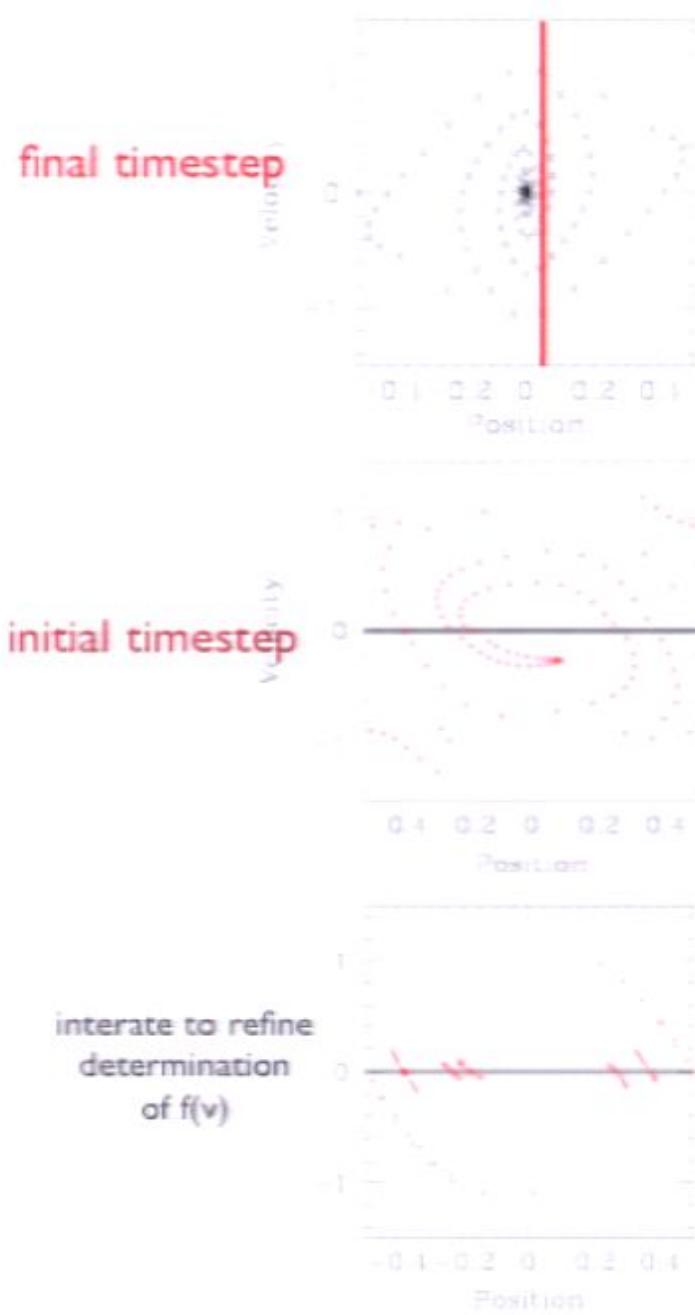
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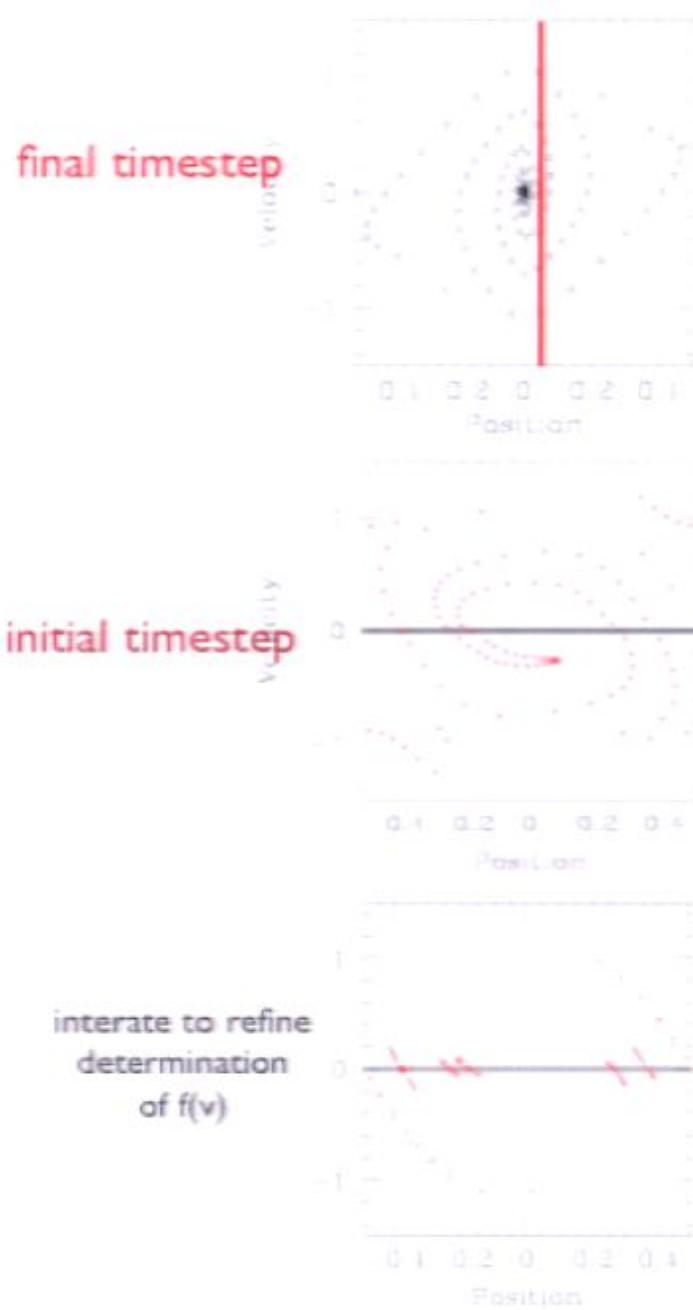
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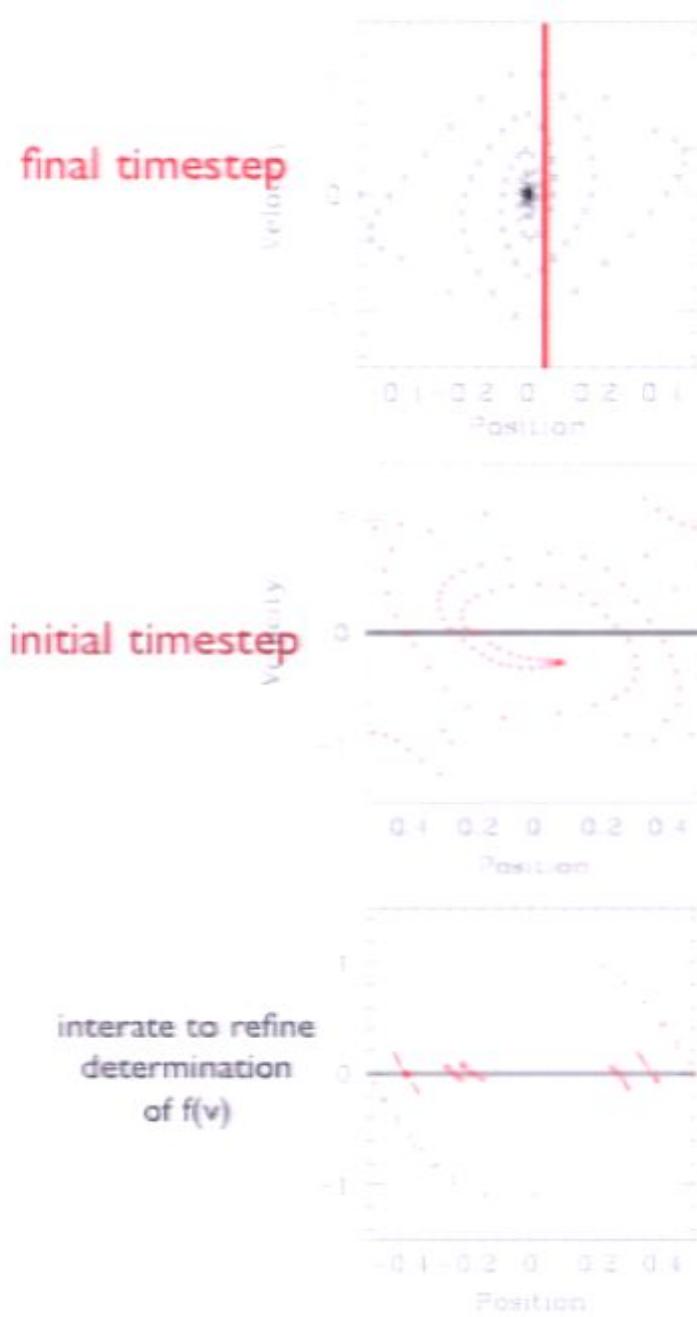
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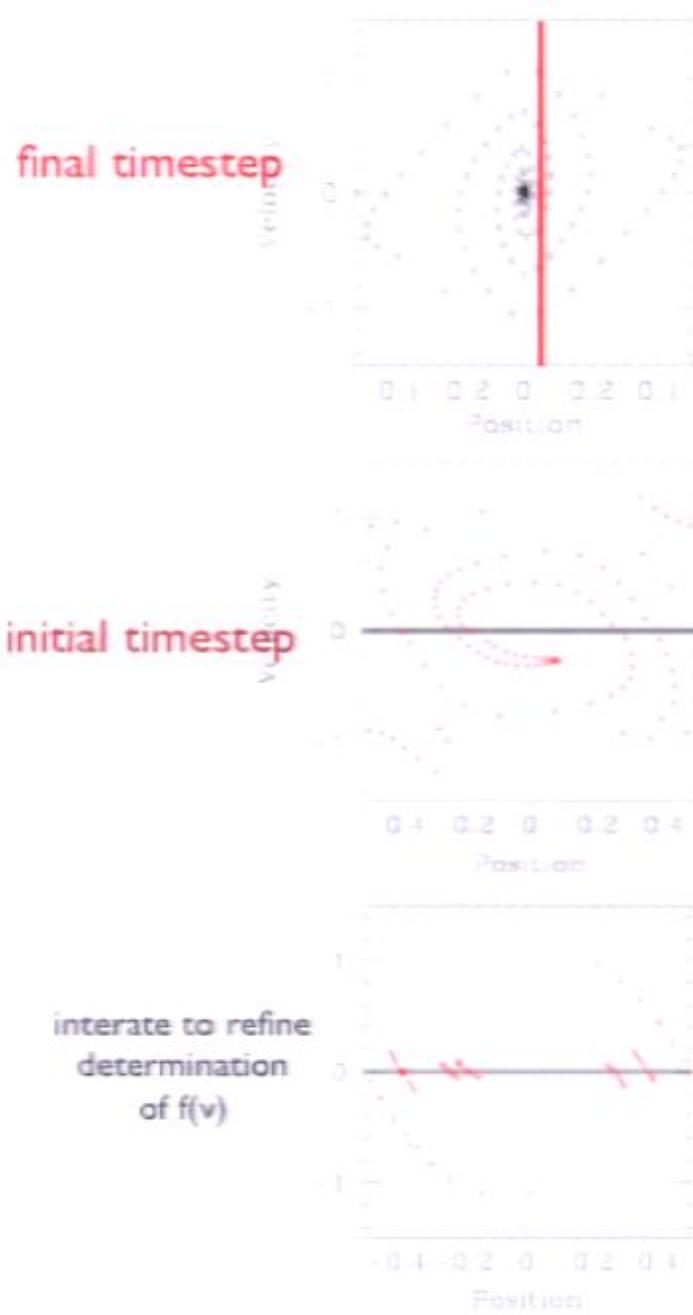
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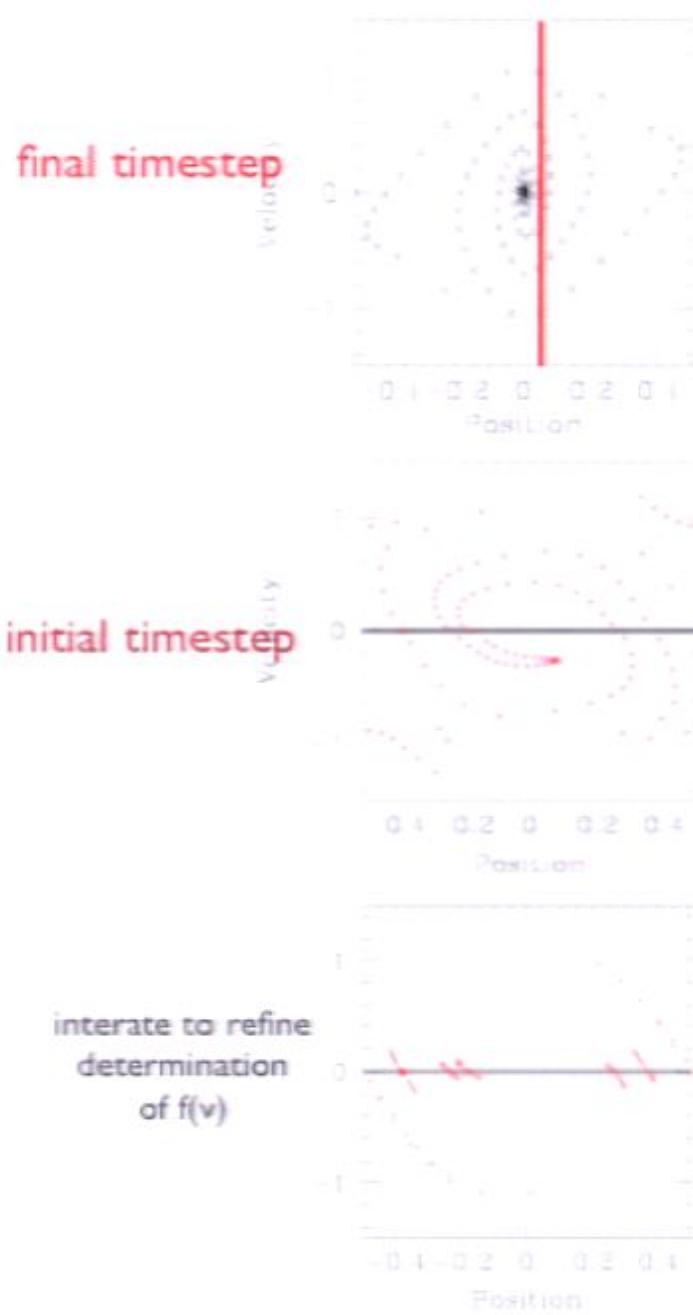
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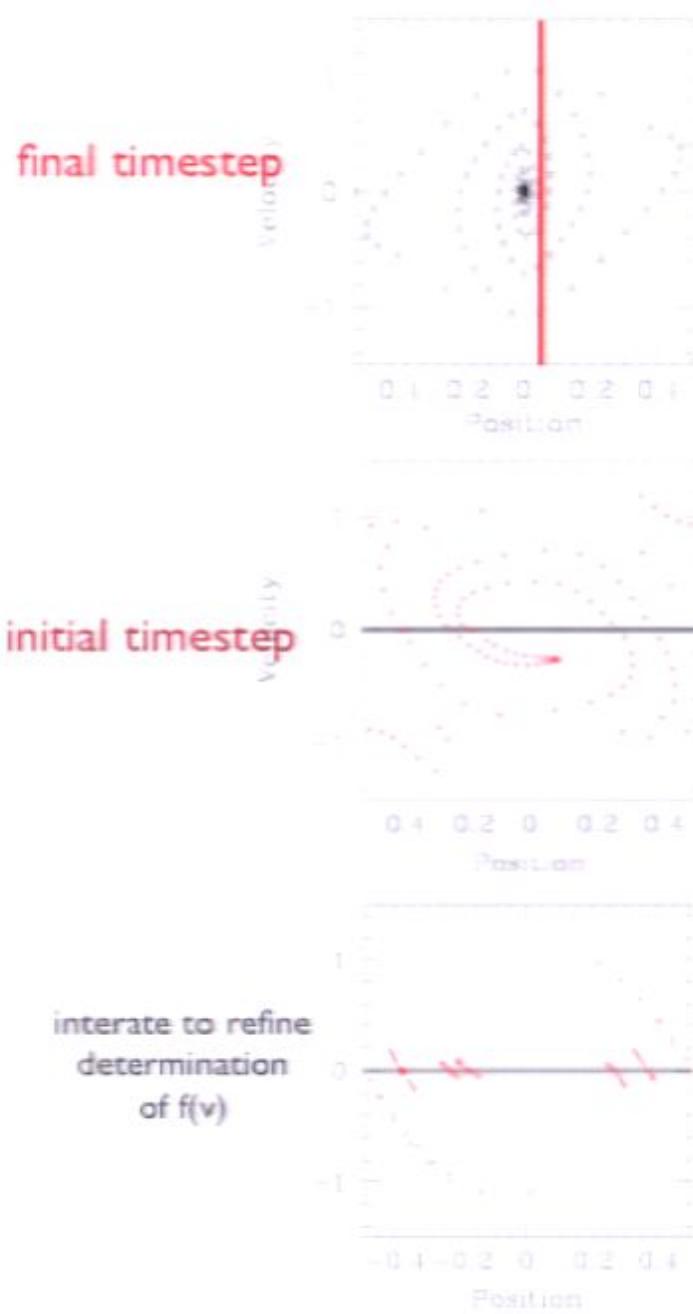
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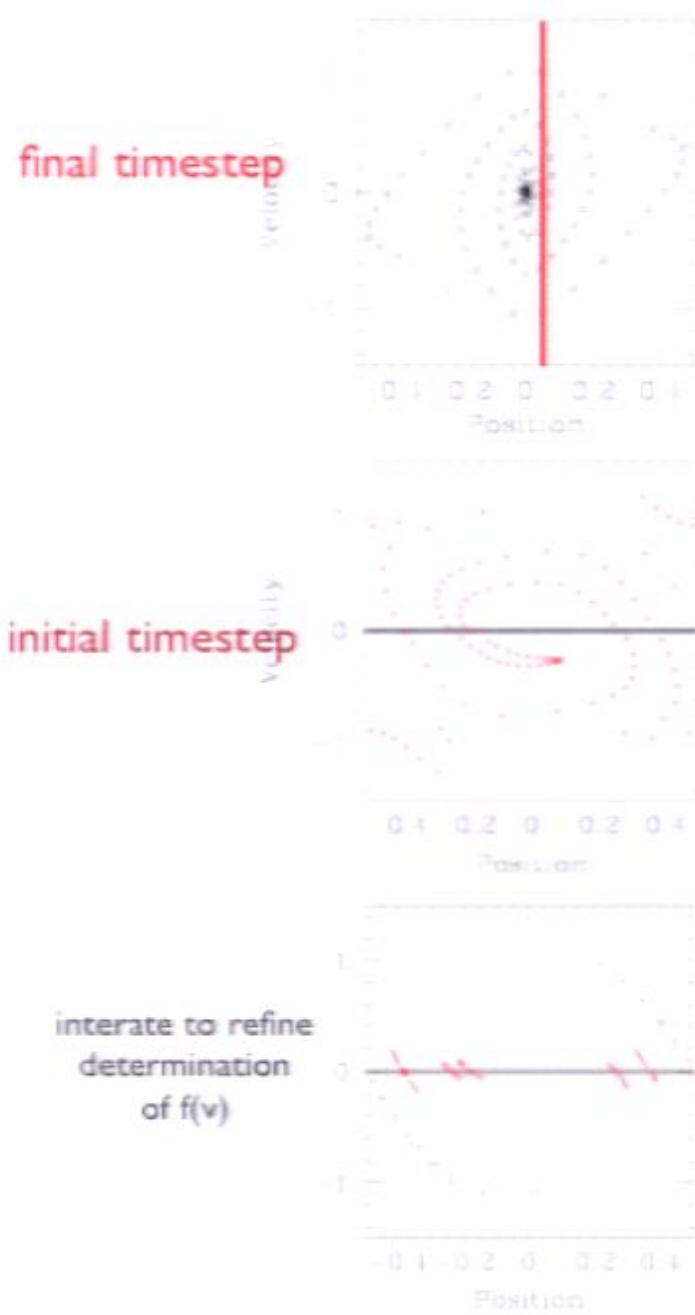
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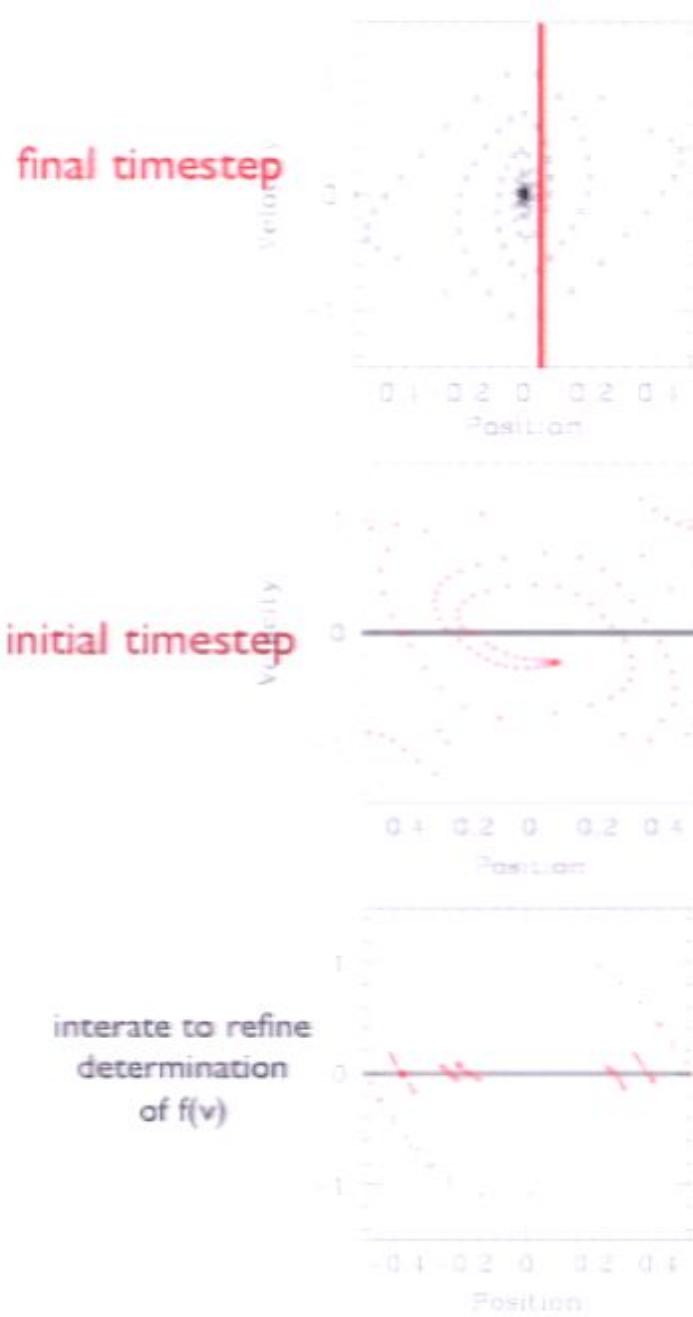
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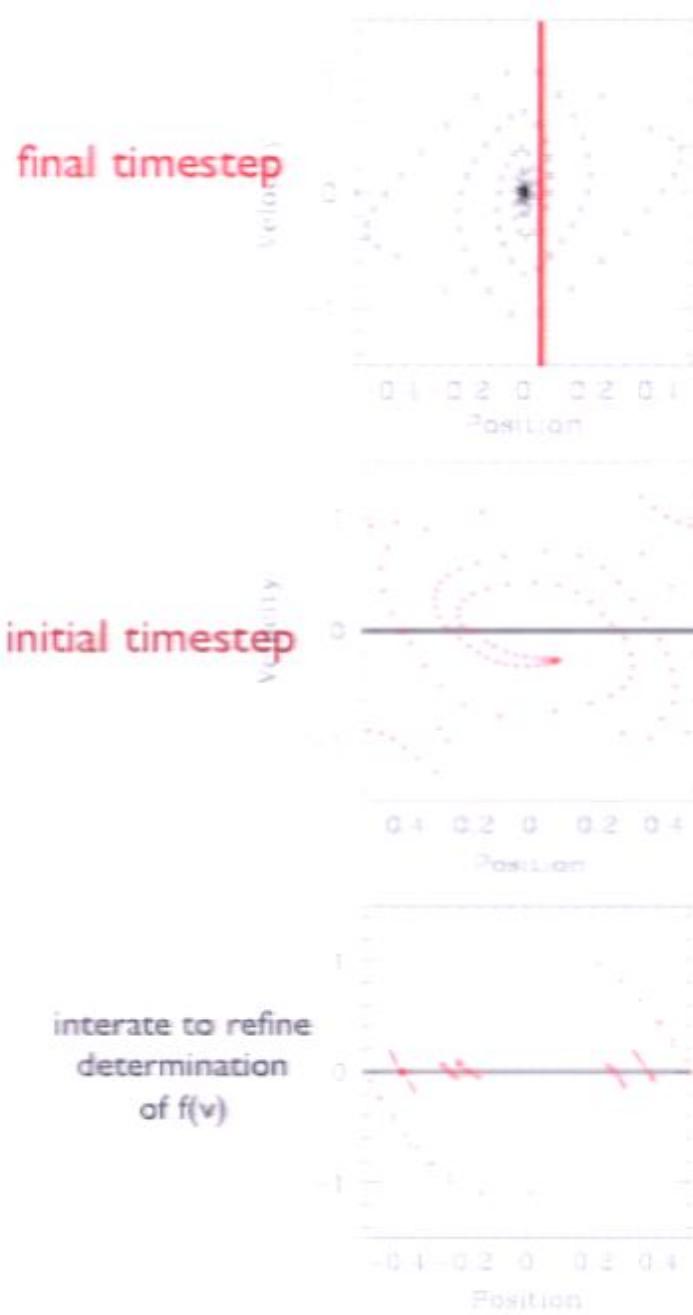
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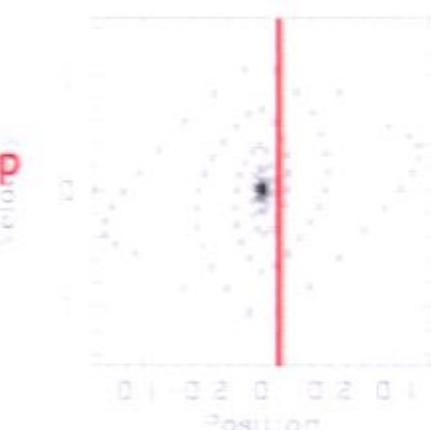
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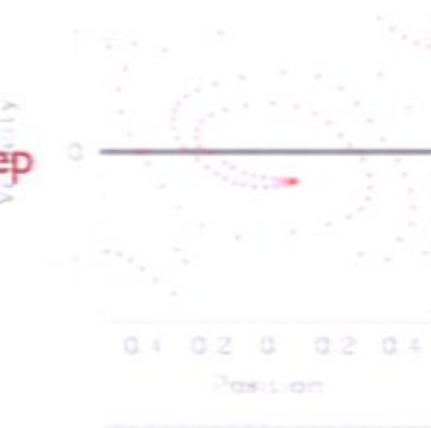
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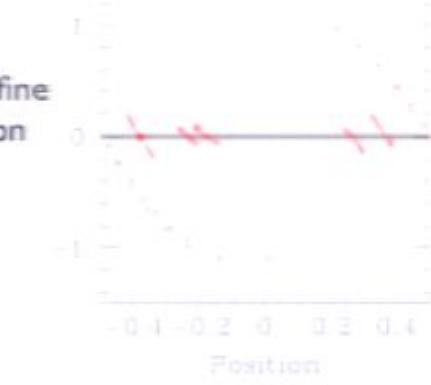
final timestep



initial timestep



iterate to refine determination of $f(v)$



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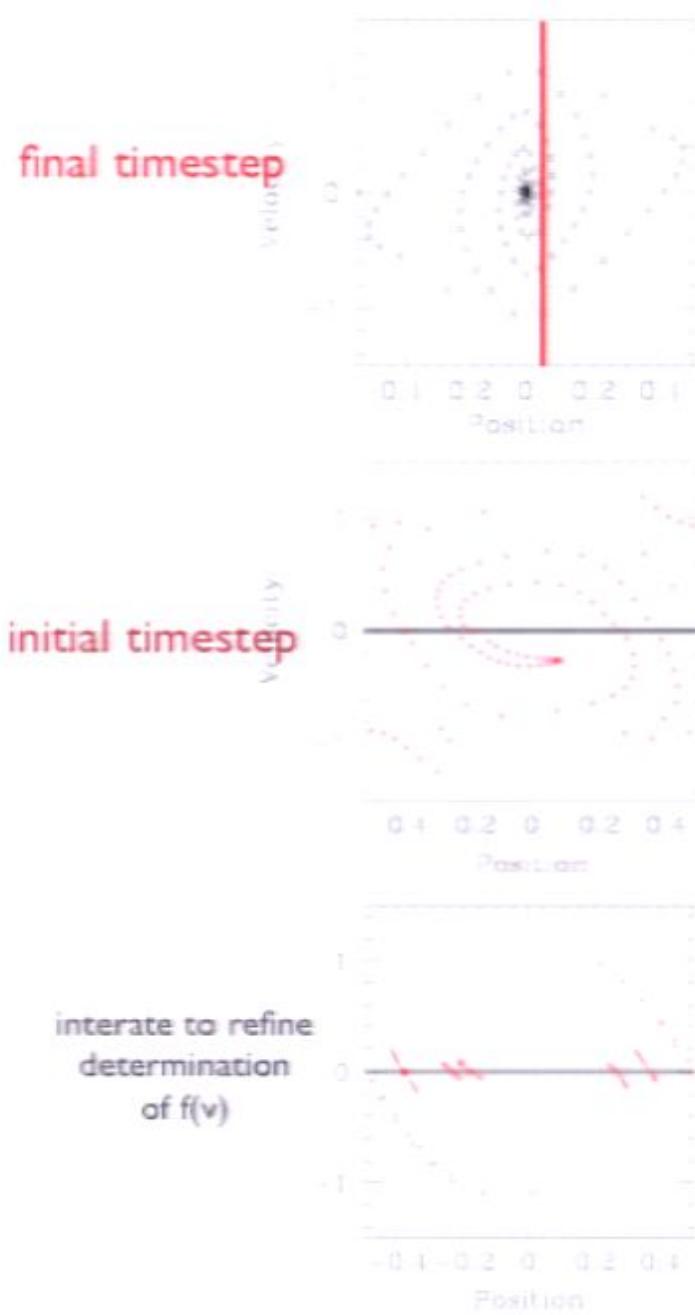
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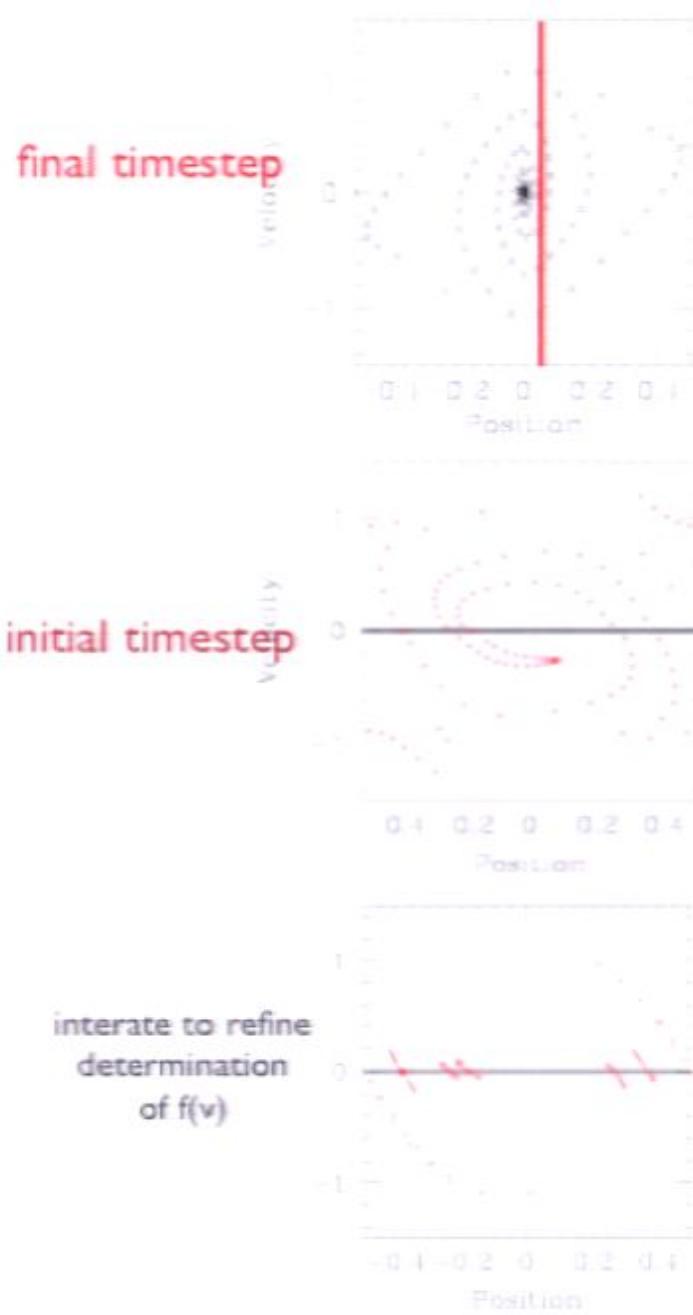
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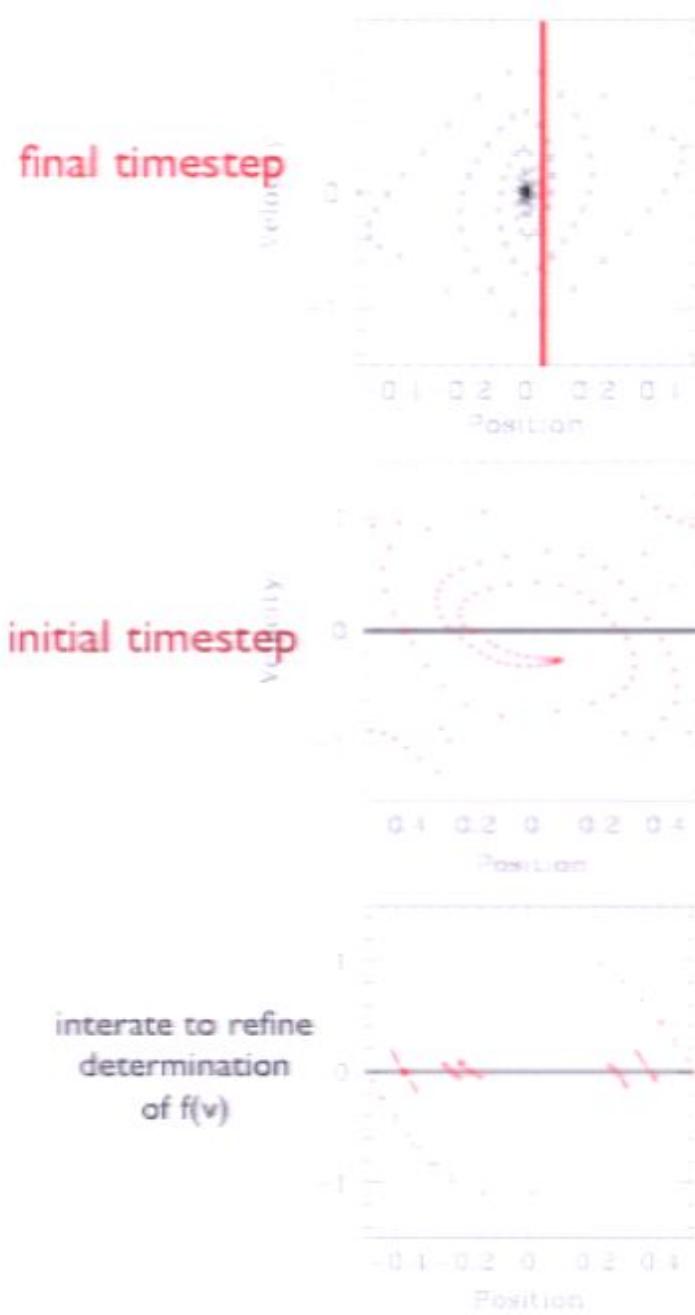
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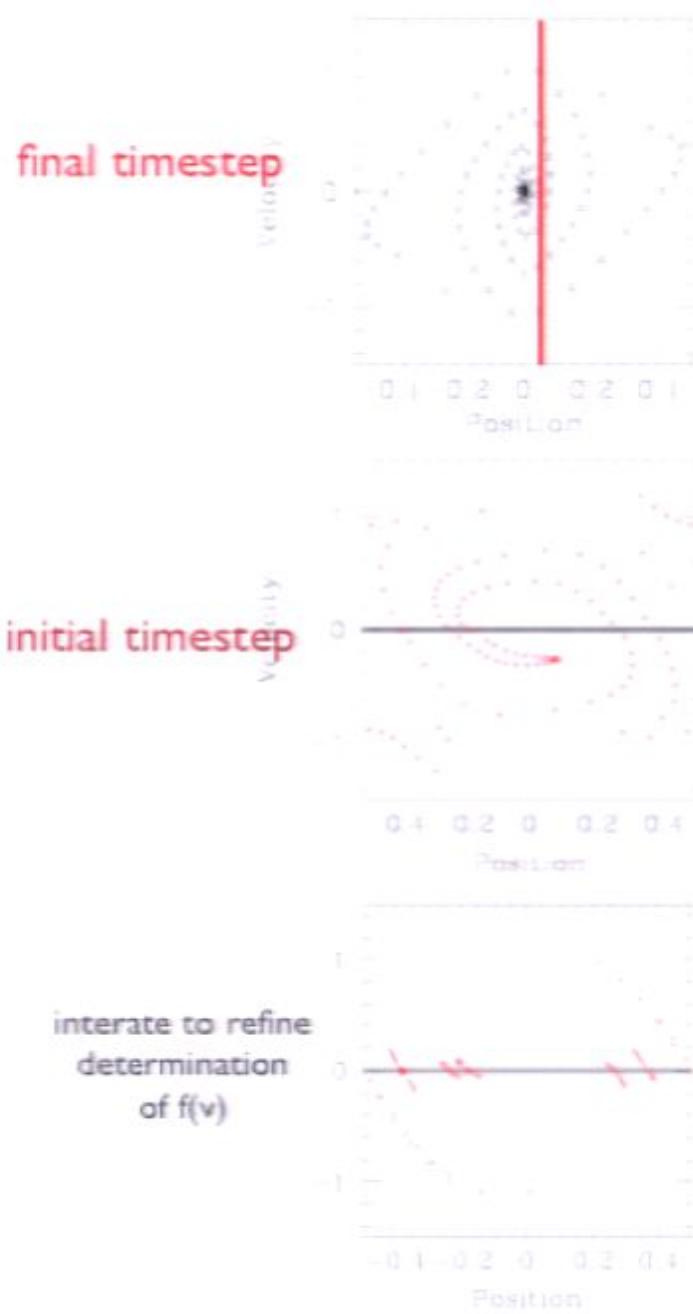
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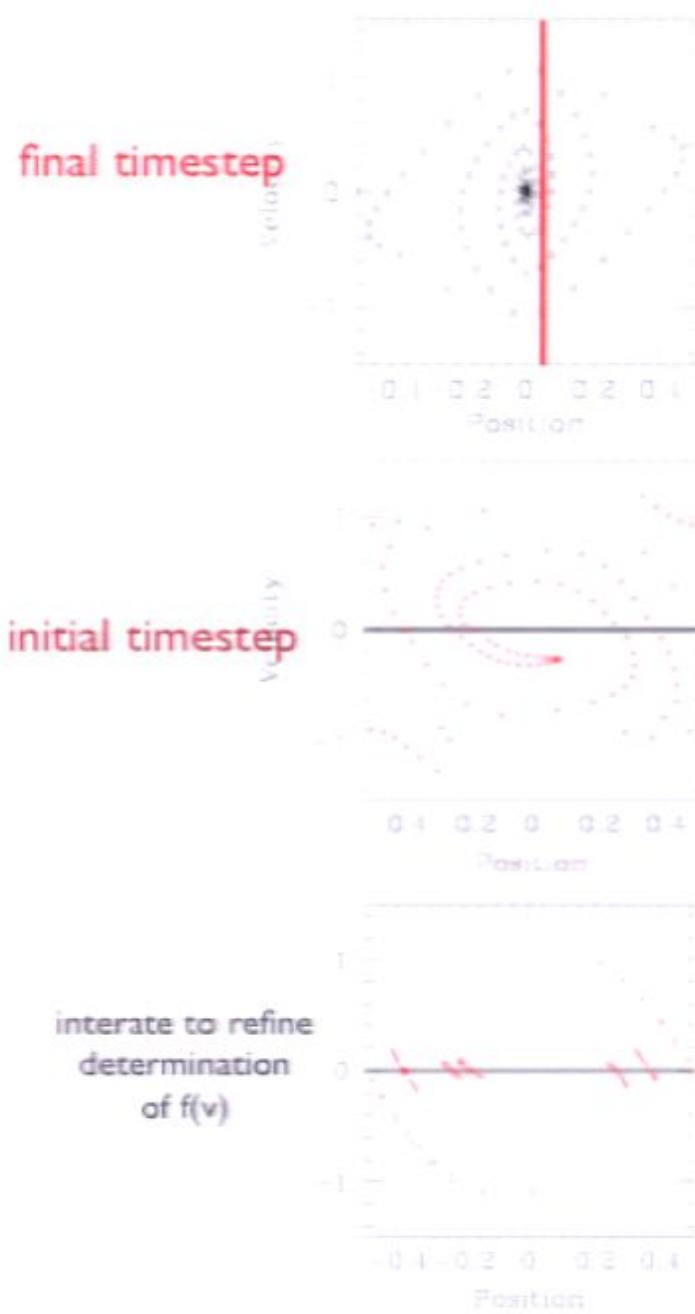
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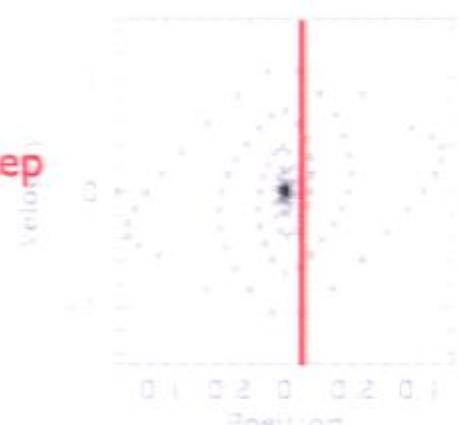
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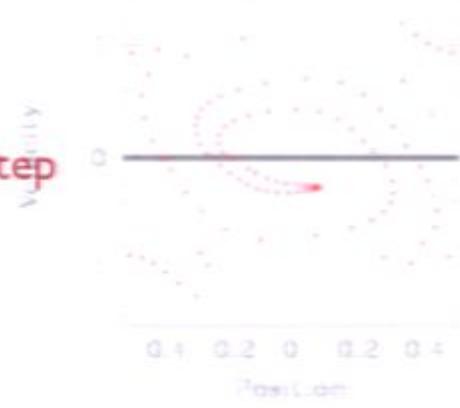
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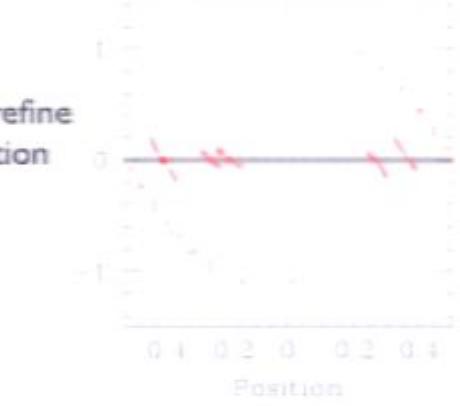
final timestep



initial timestep



iterate to refine determination of $f(v)$



Stiff & LMW (PRL, 2003)

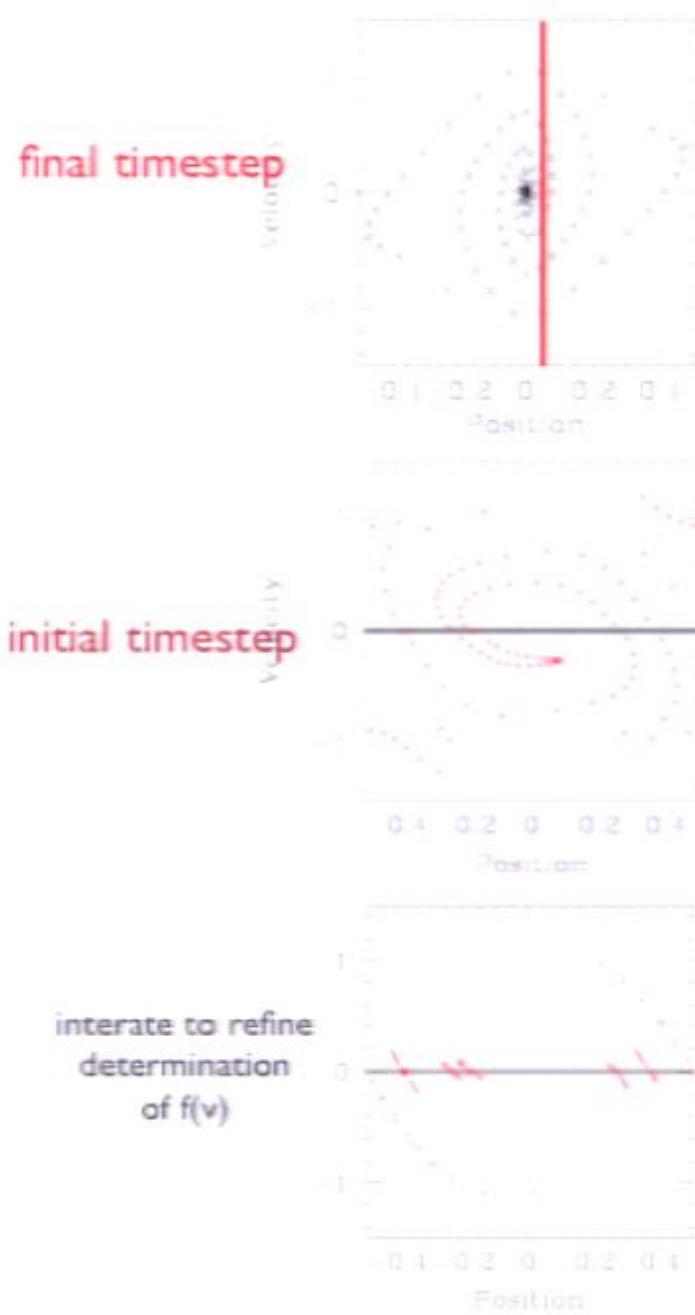
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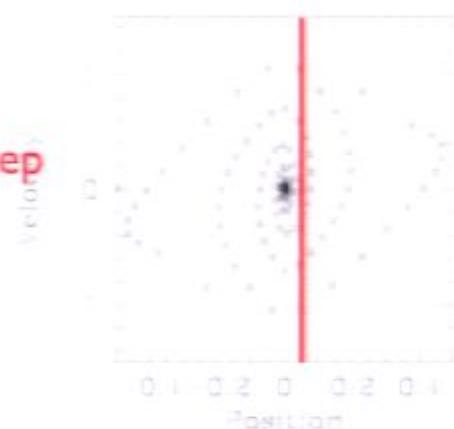
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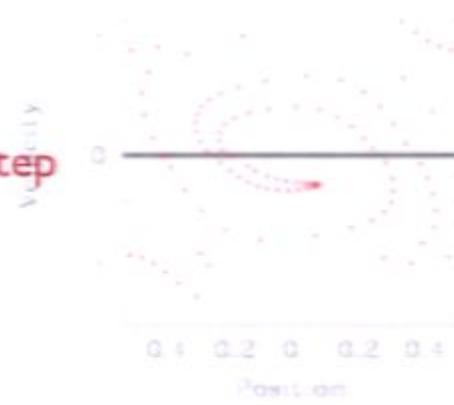
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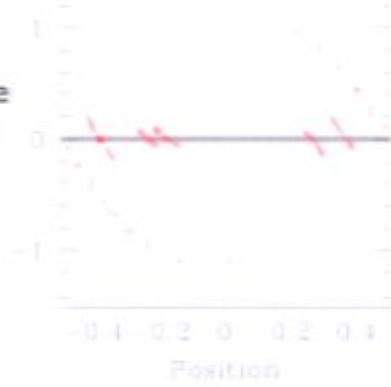
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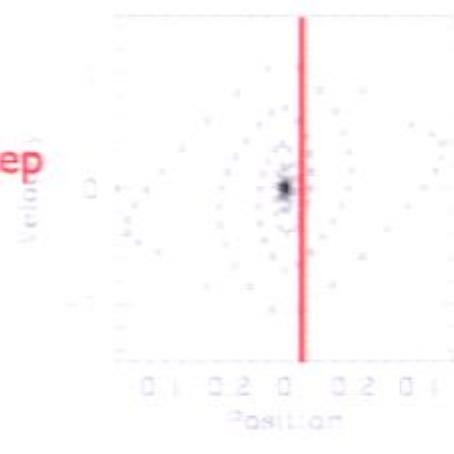
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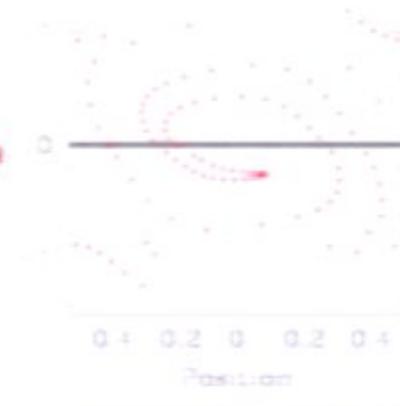
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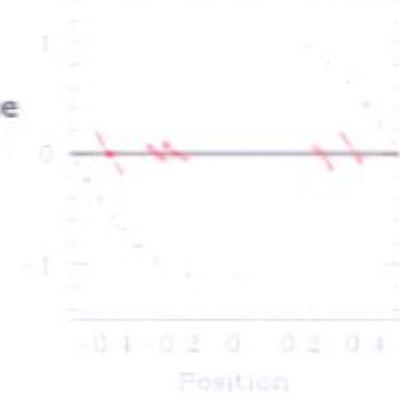
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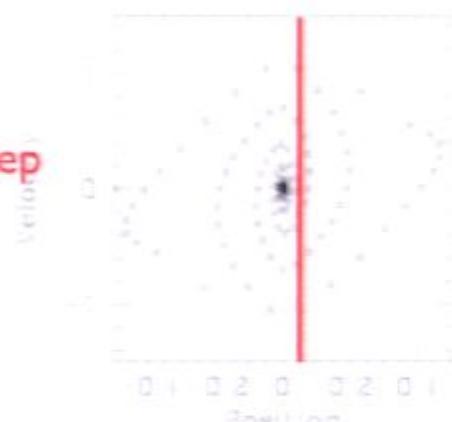
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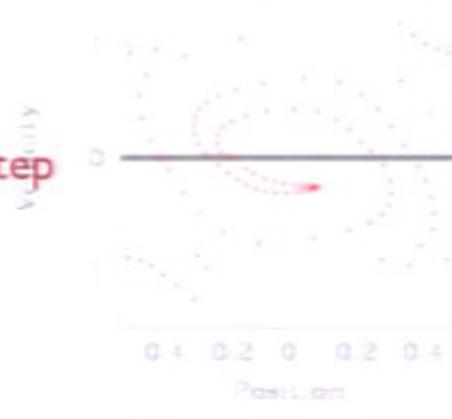
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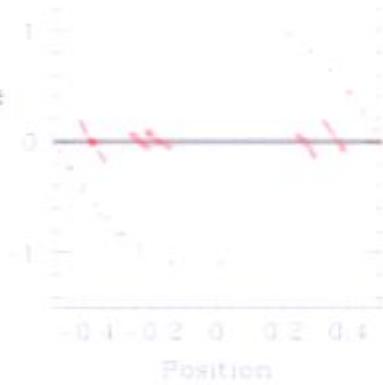
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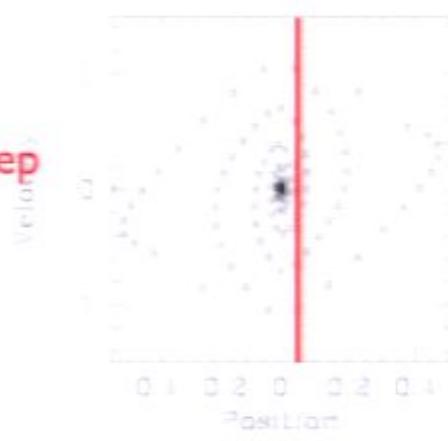
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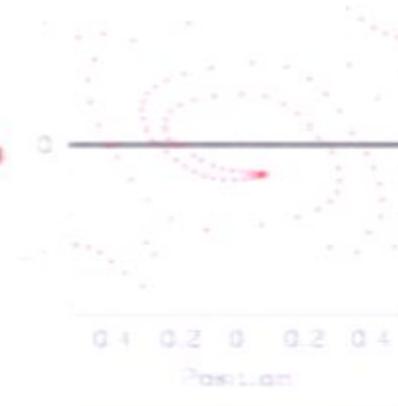
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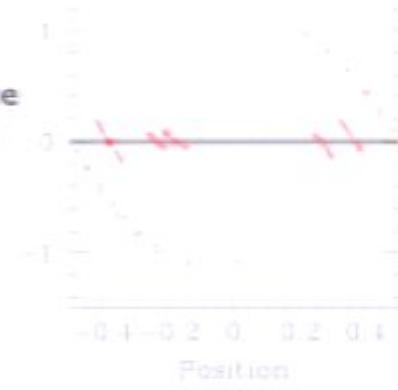
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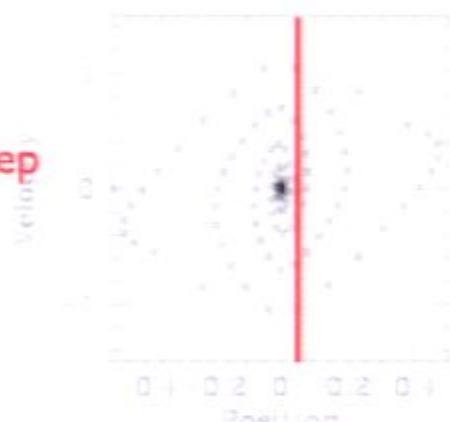
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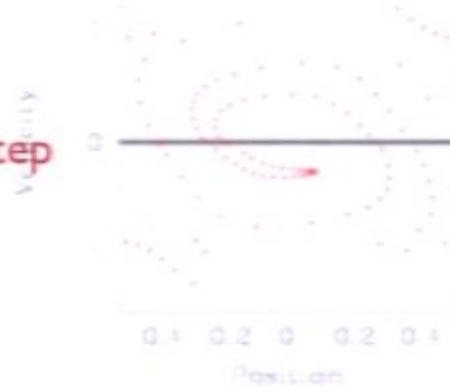
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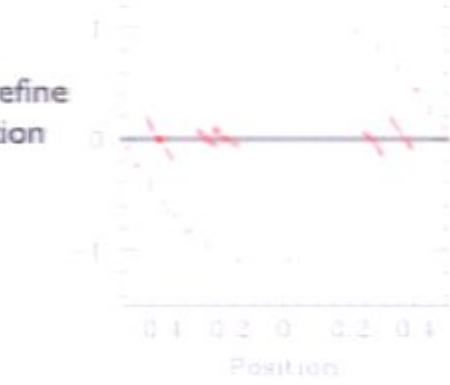
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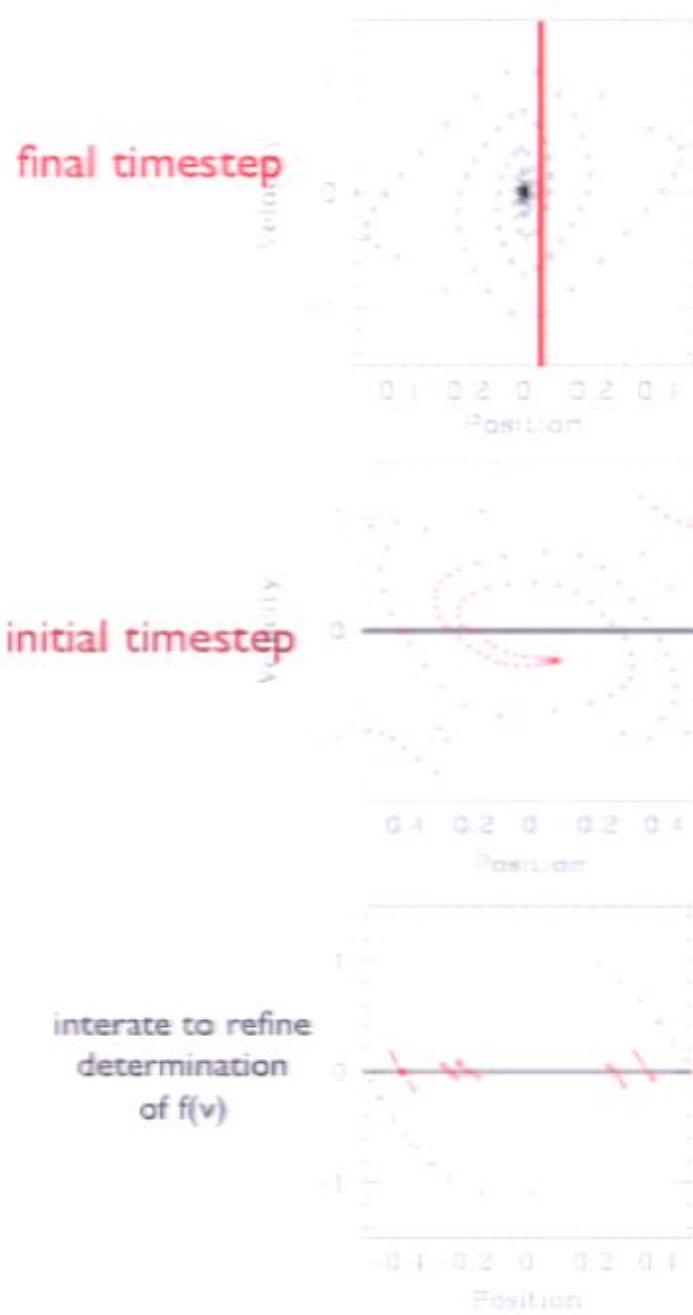
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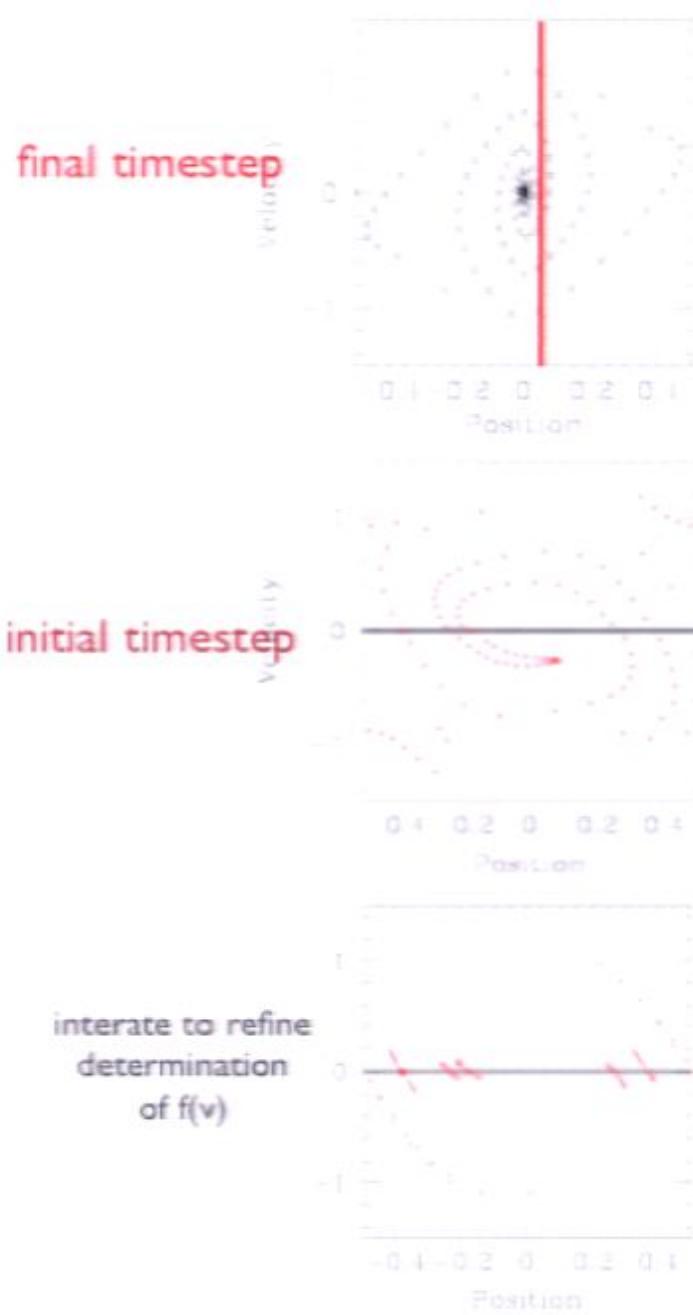
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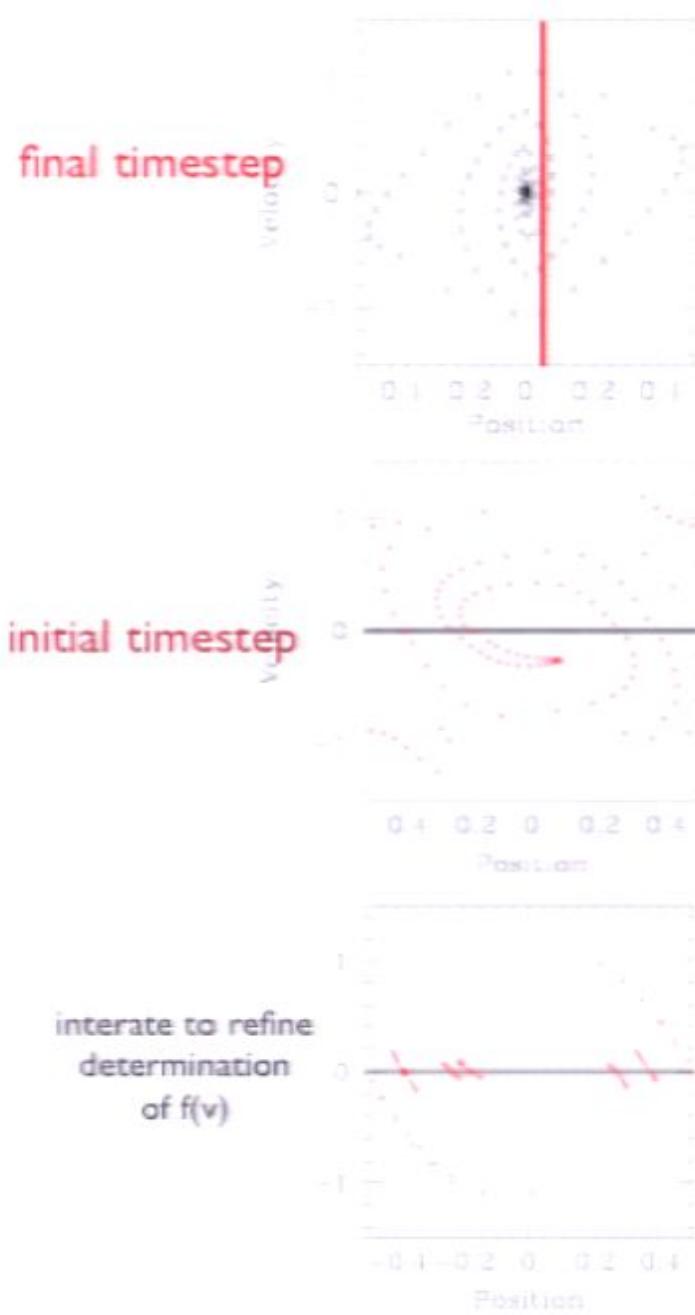
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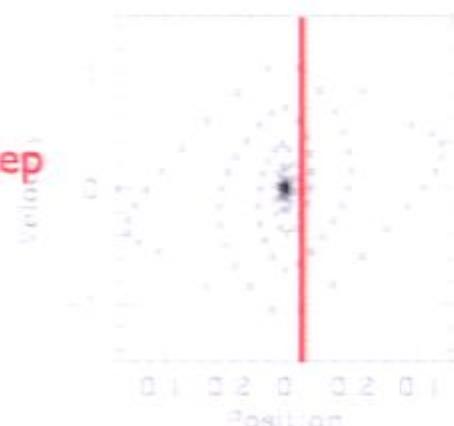
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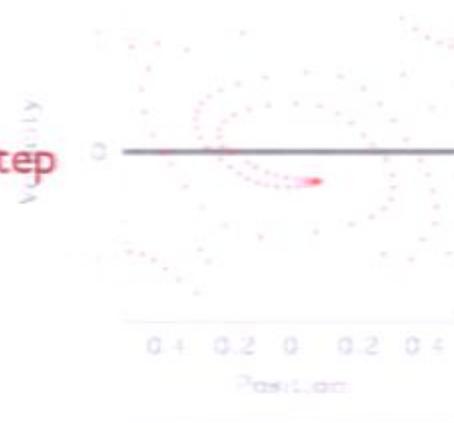
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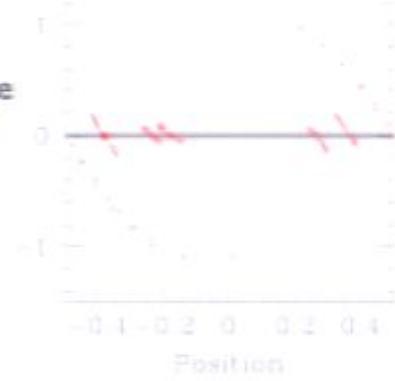
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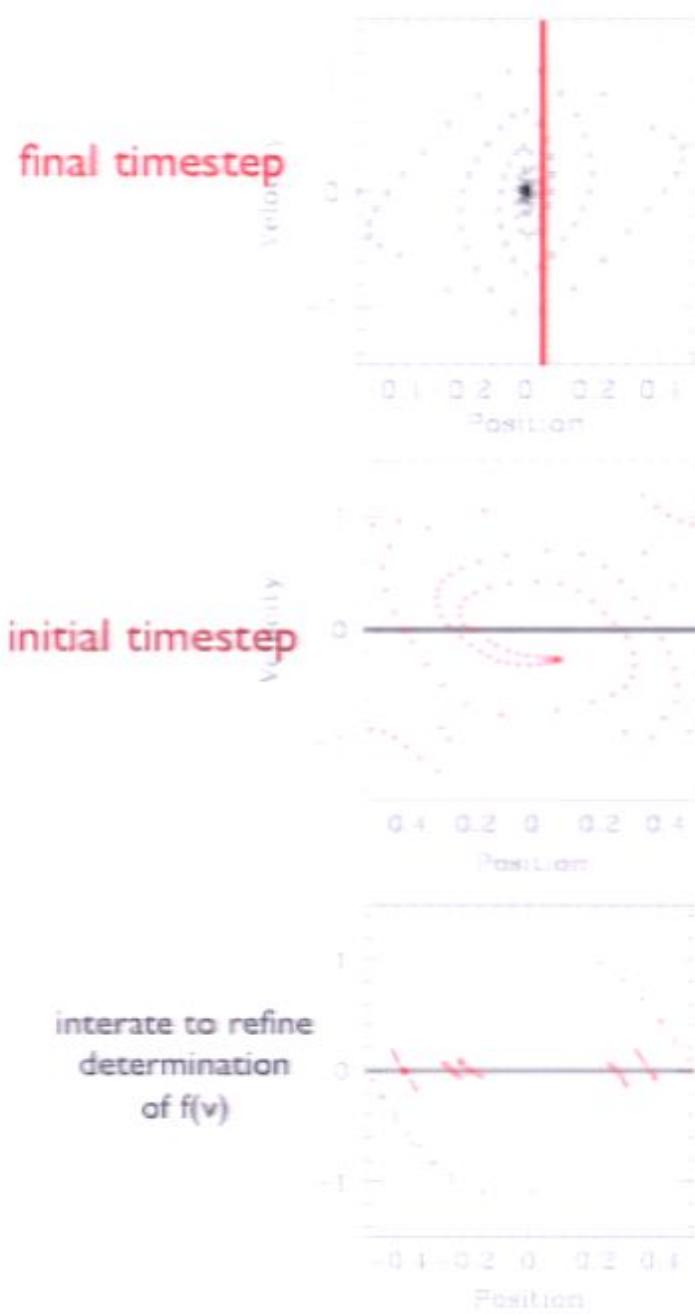
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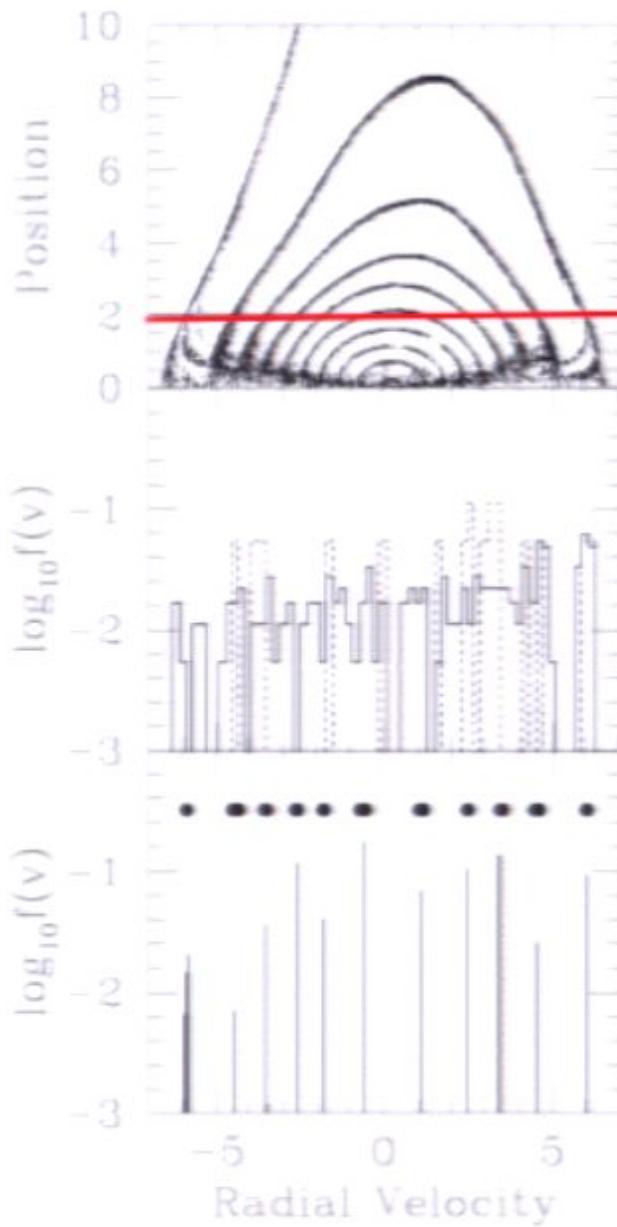
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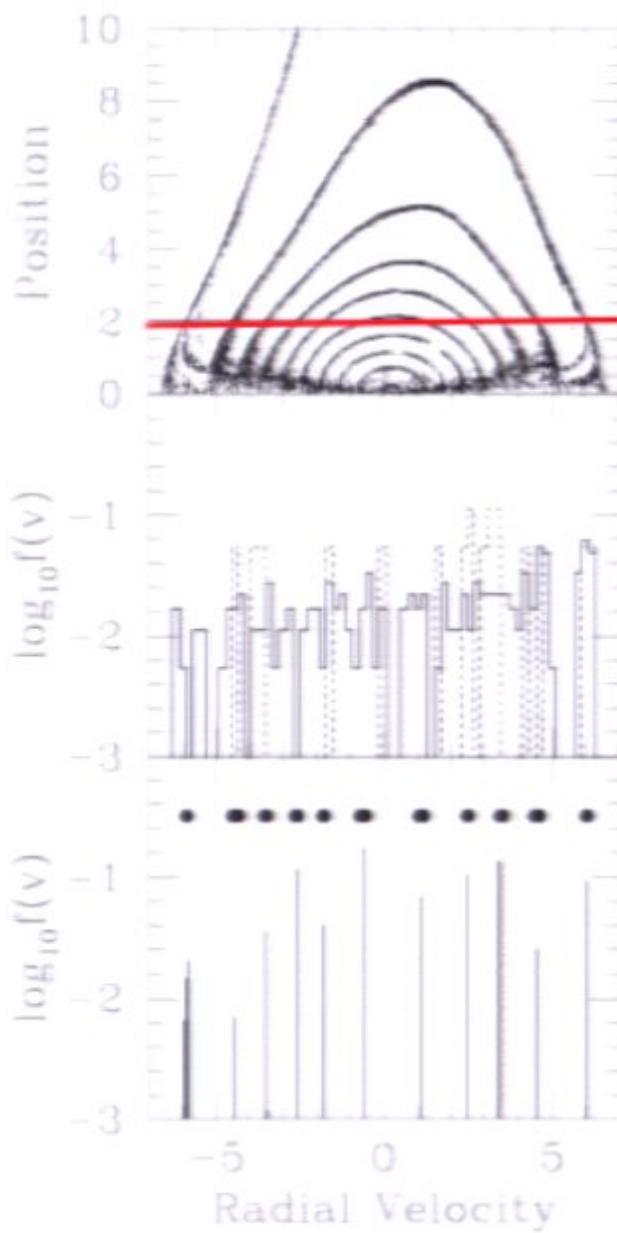


final frame for spherical collapse
from **high-res** shell-model calculation

red line marks position of detector

velocity distribution measured
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**too few particles to map out true
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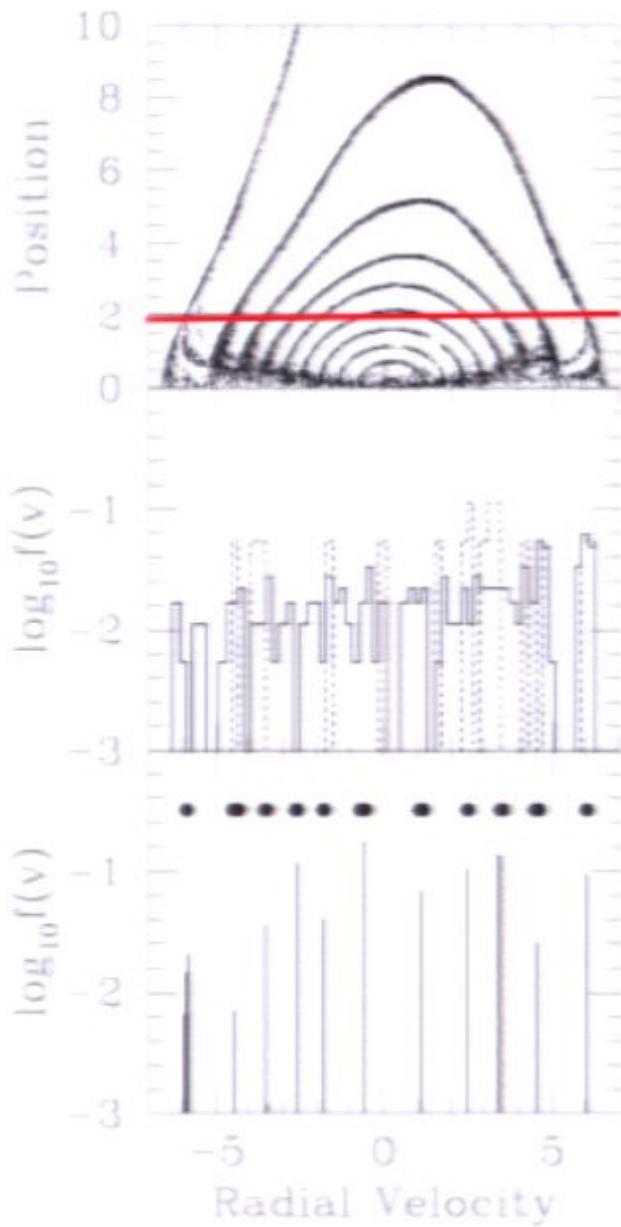


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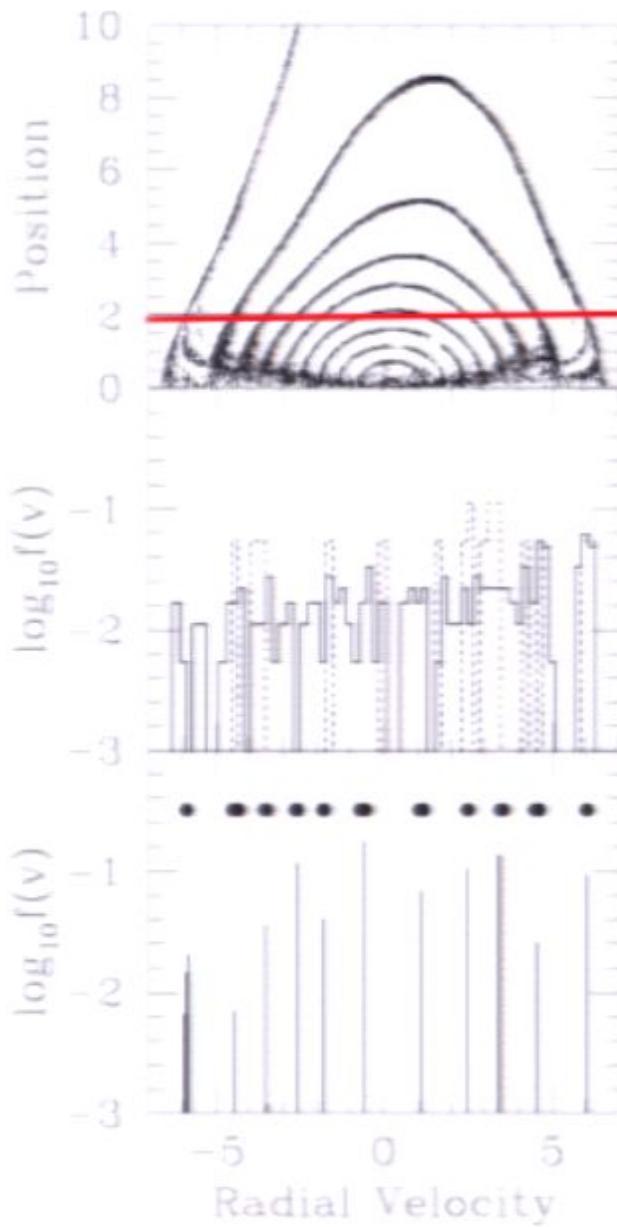


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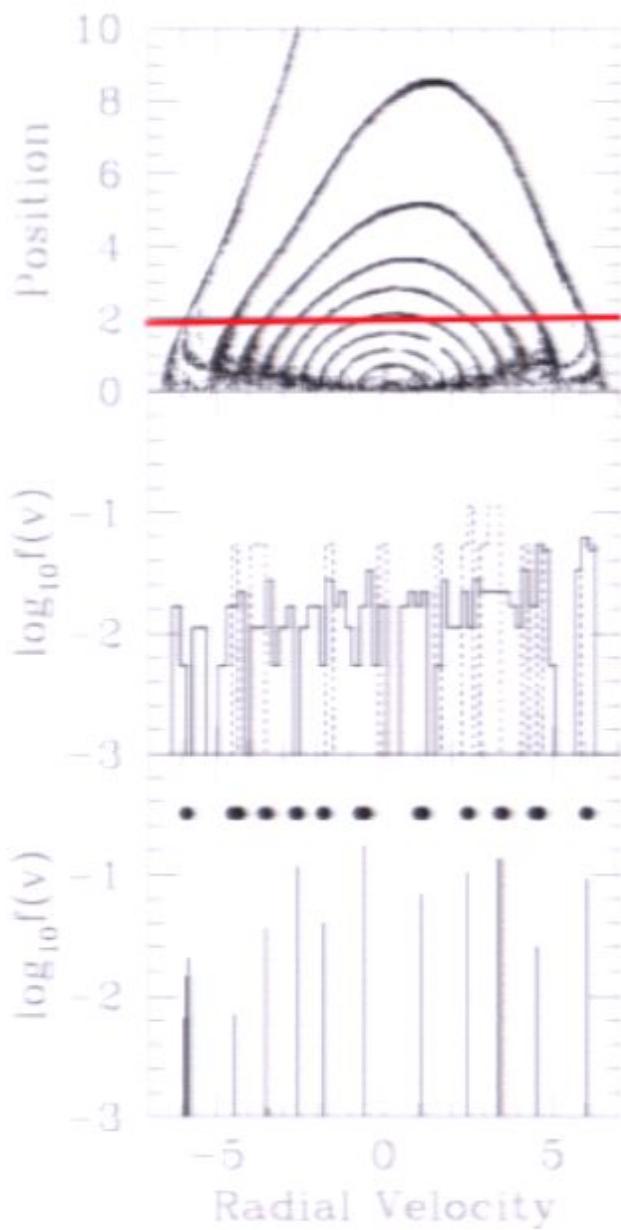


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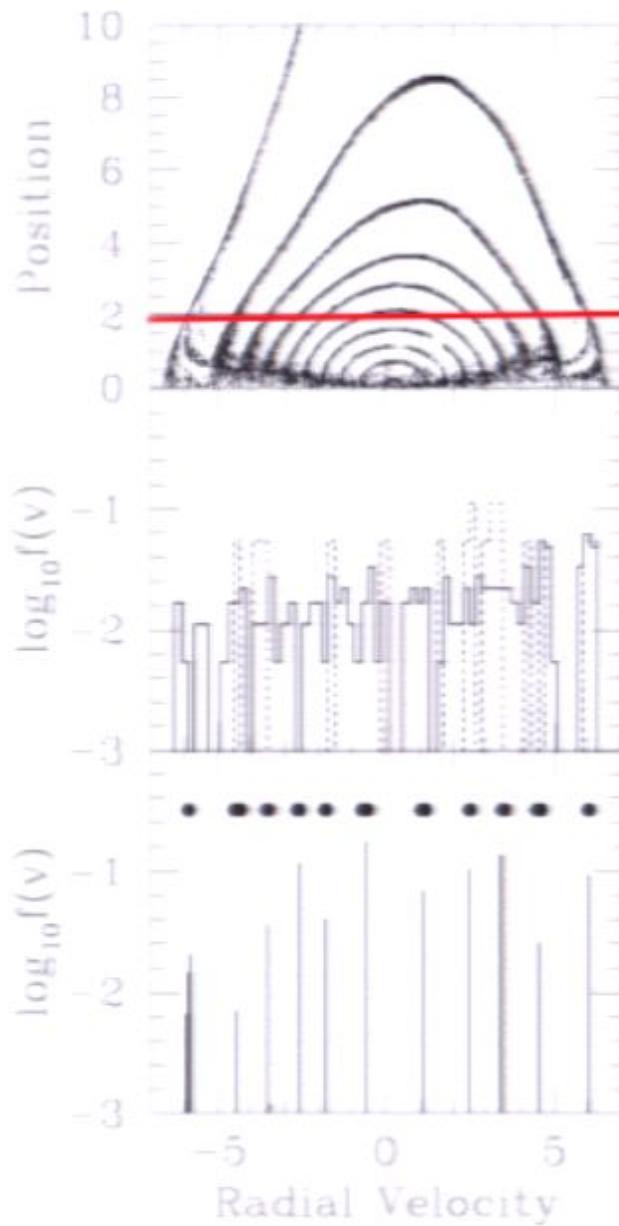


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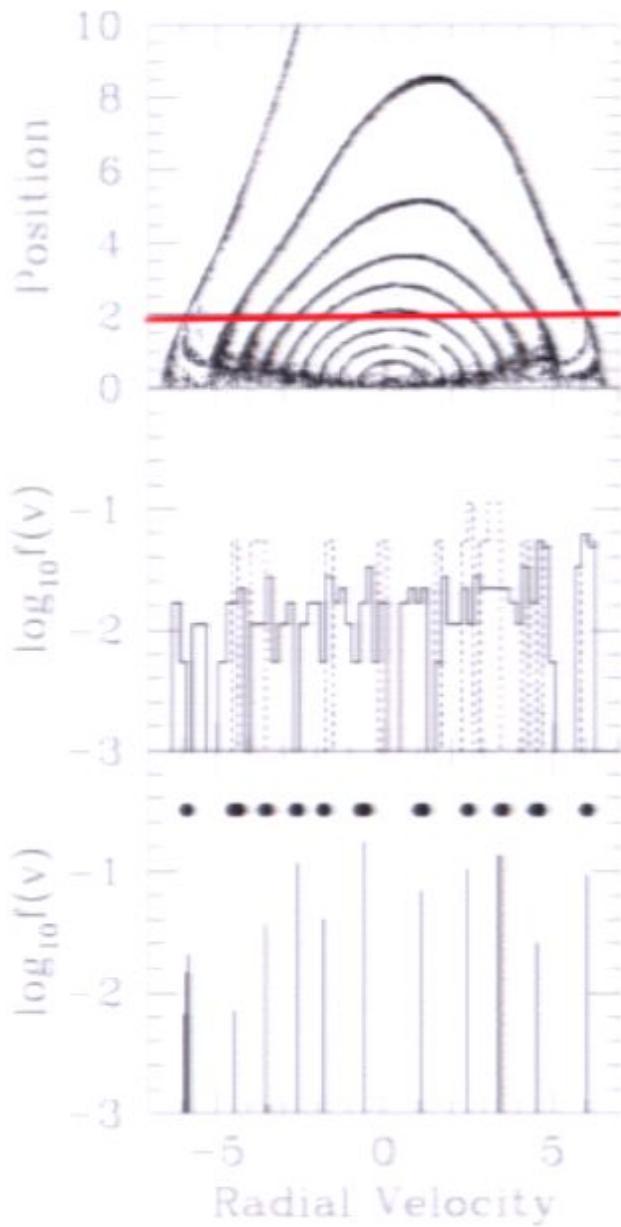


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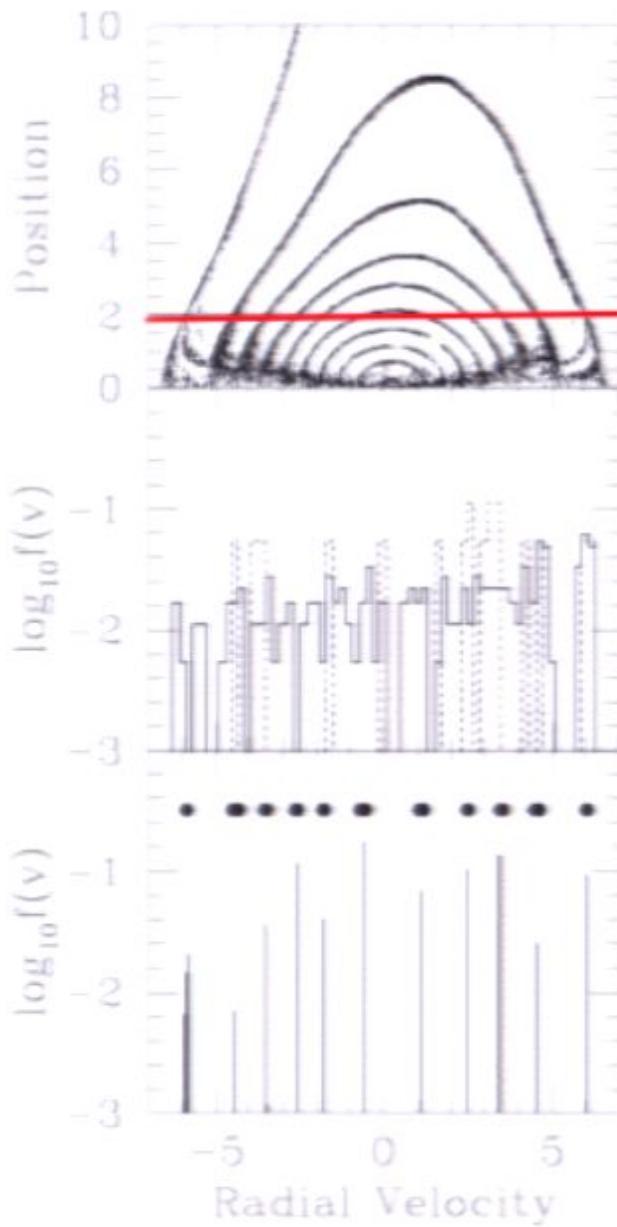


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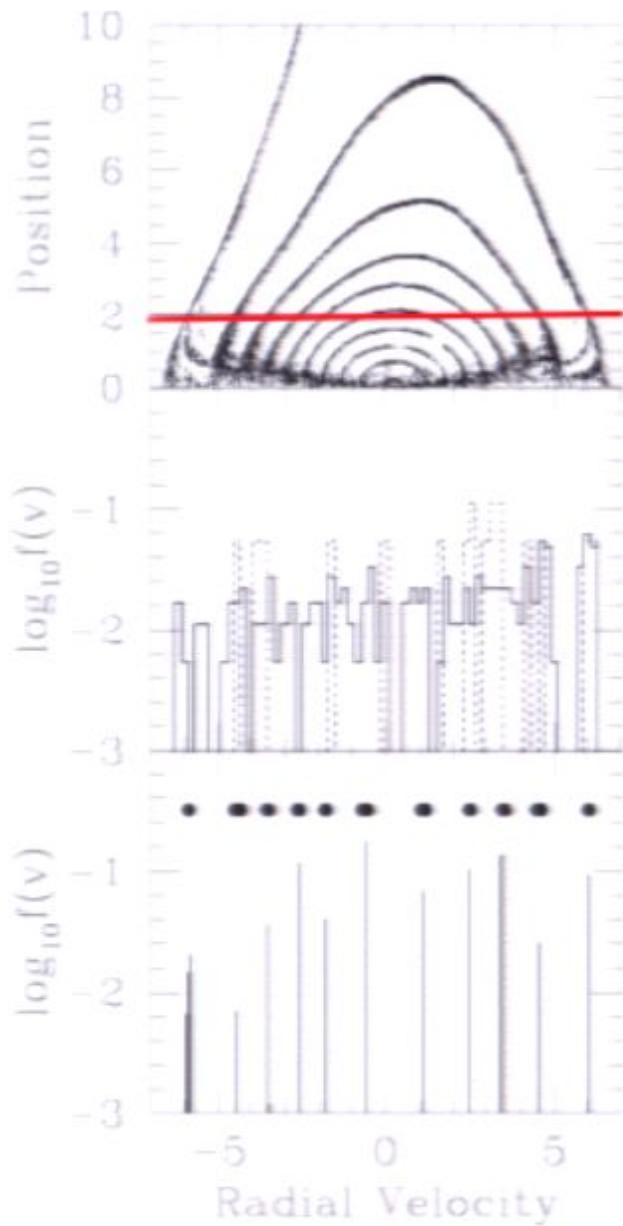


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Liouville's theorem must be satisfied.

Of course, cosmological N-body simulations
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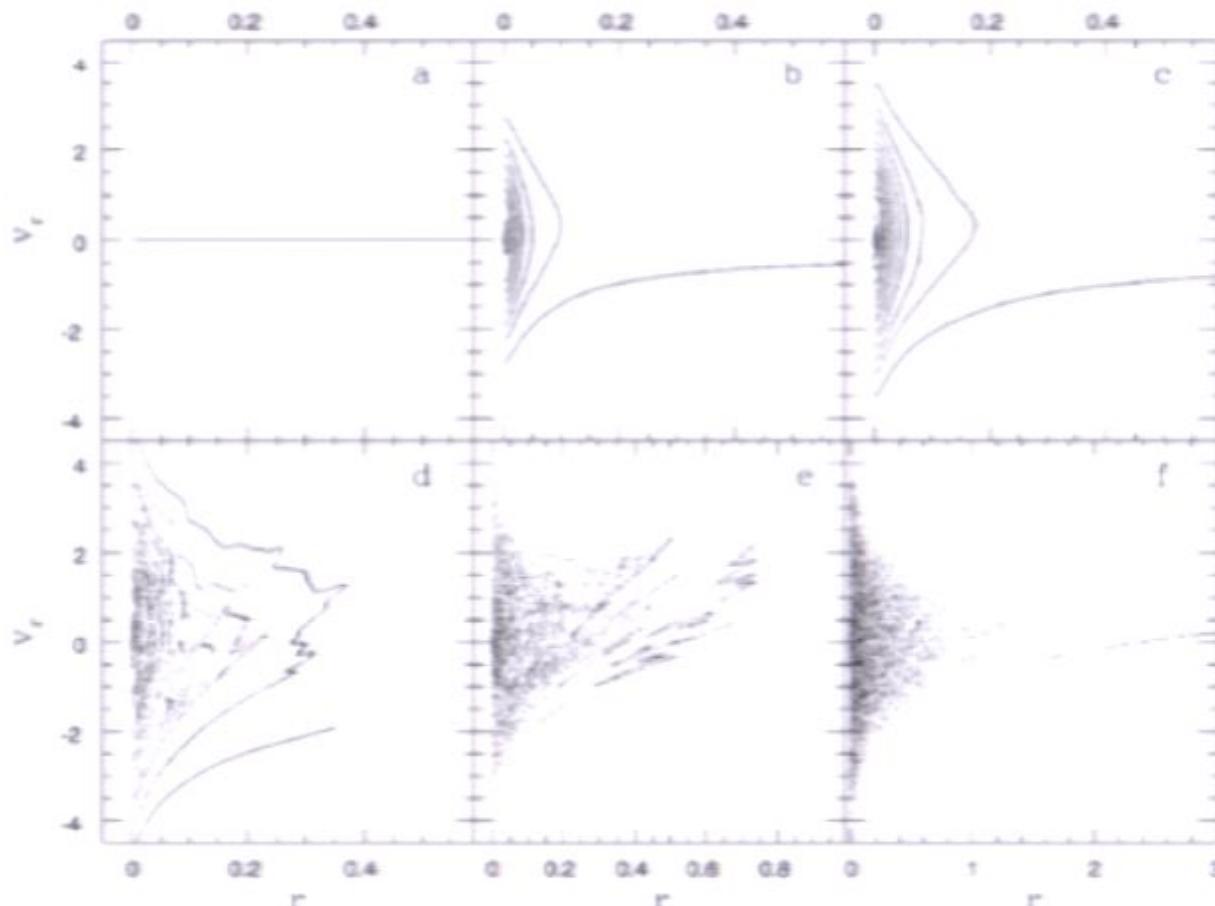
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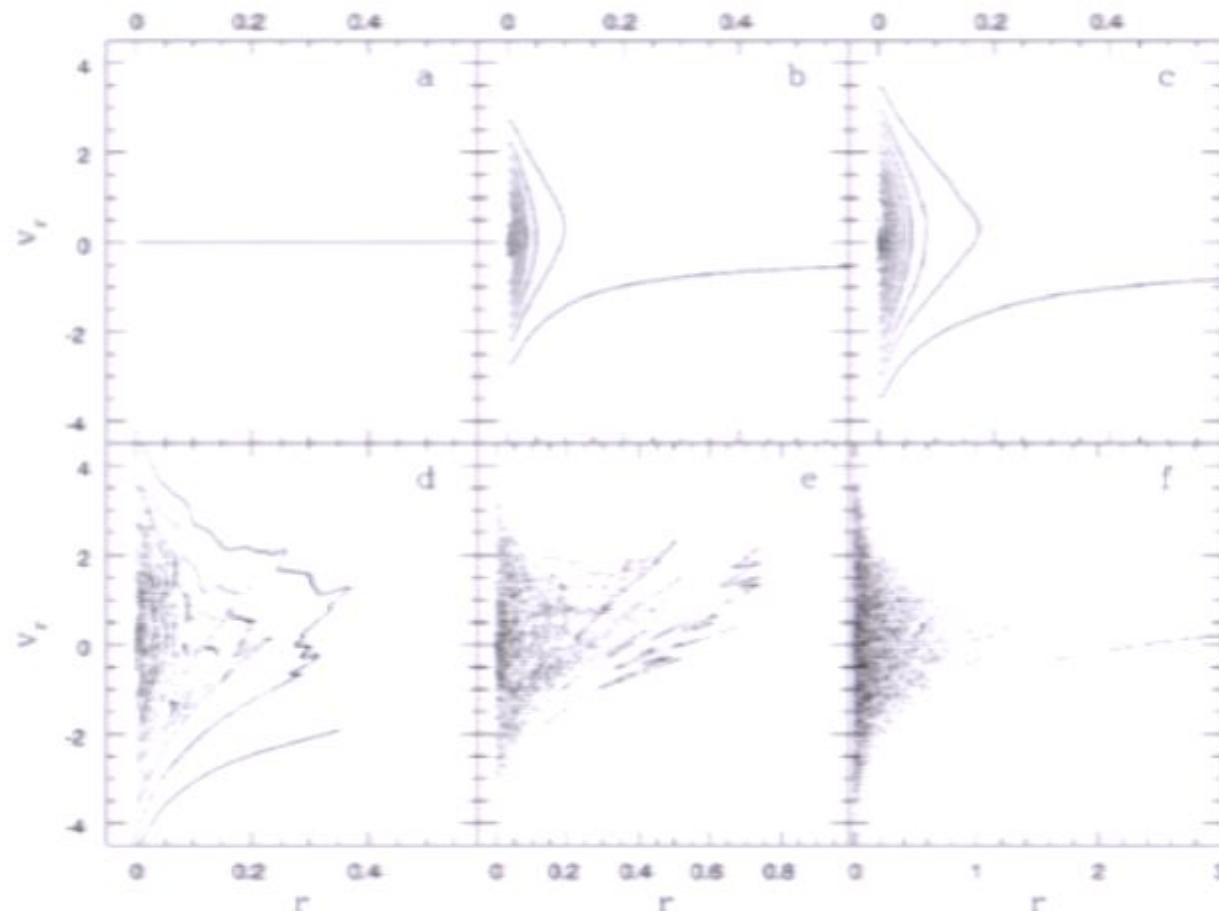
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FMG pick position of detector at final epoch and consider test-particles at this position with different velocities.

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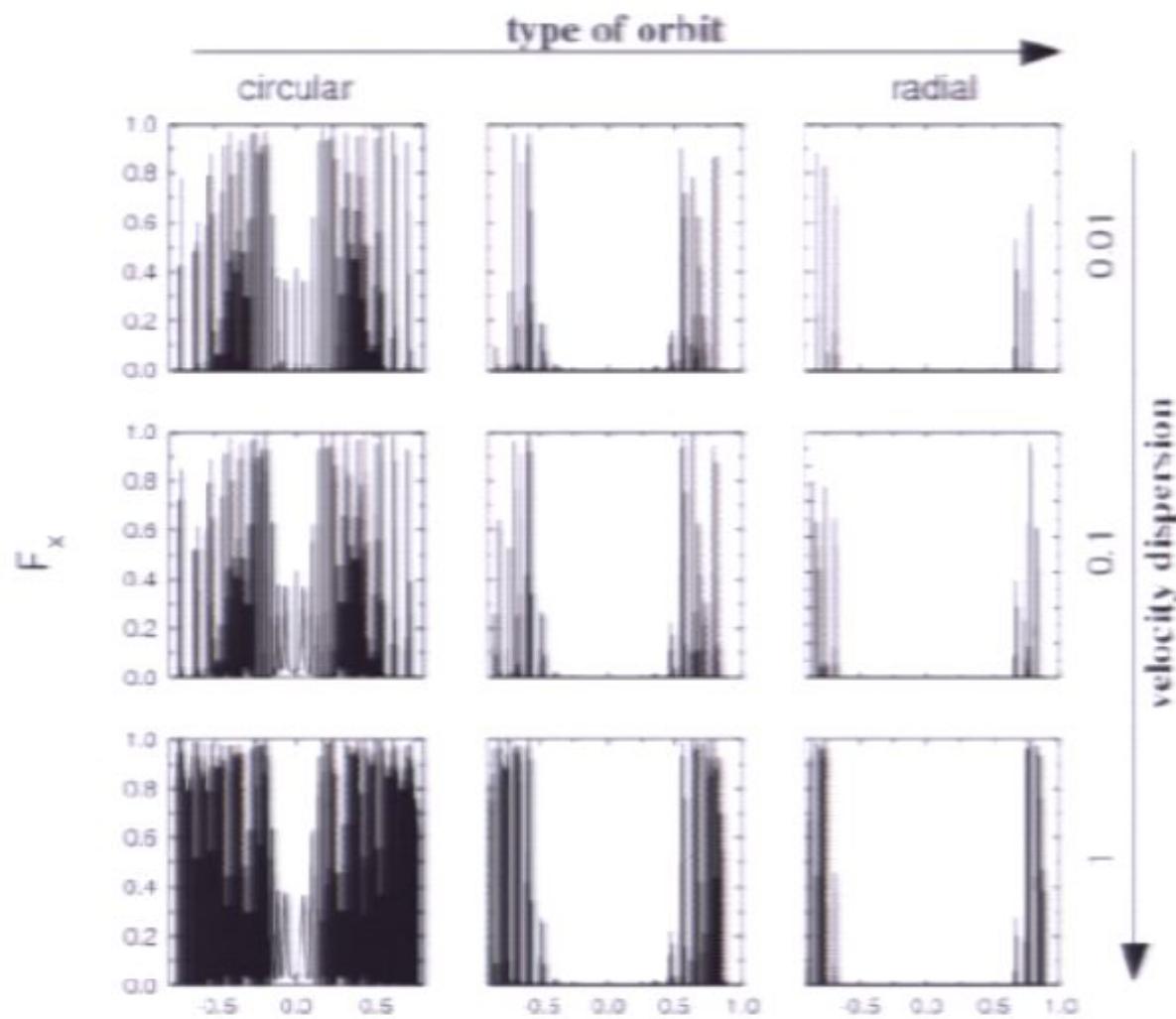
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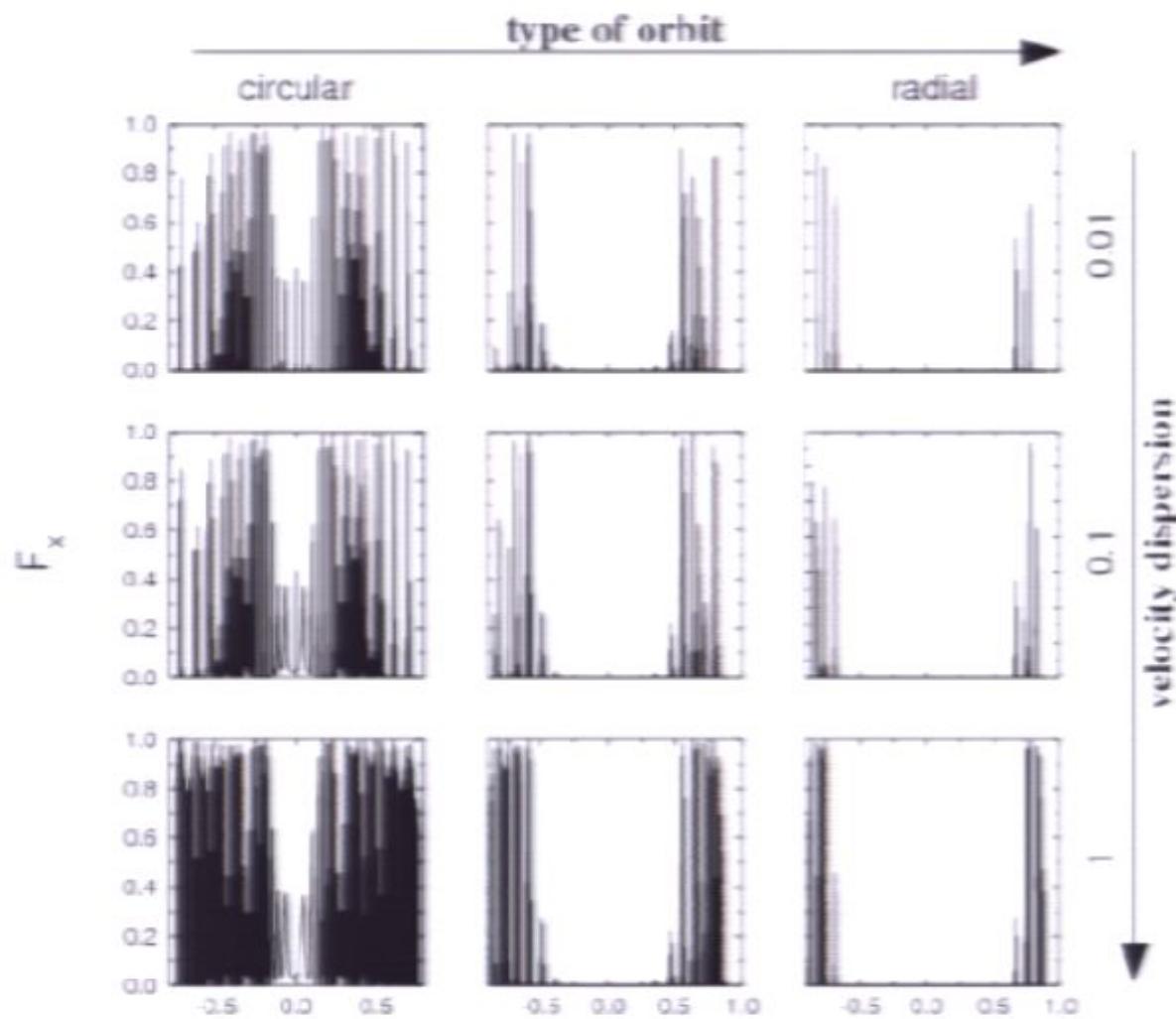
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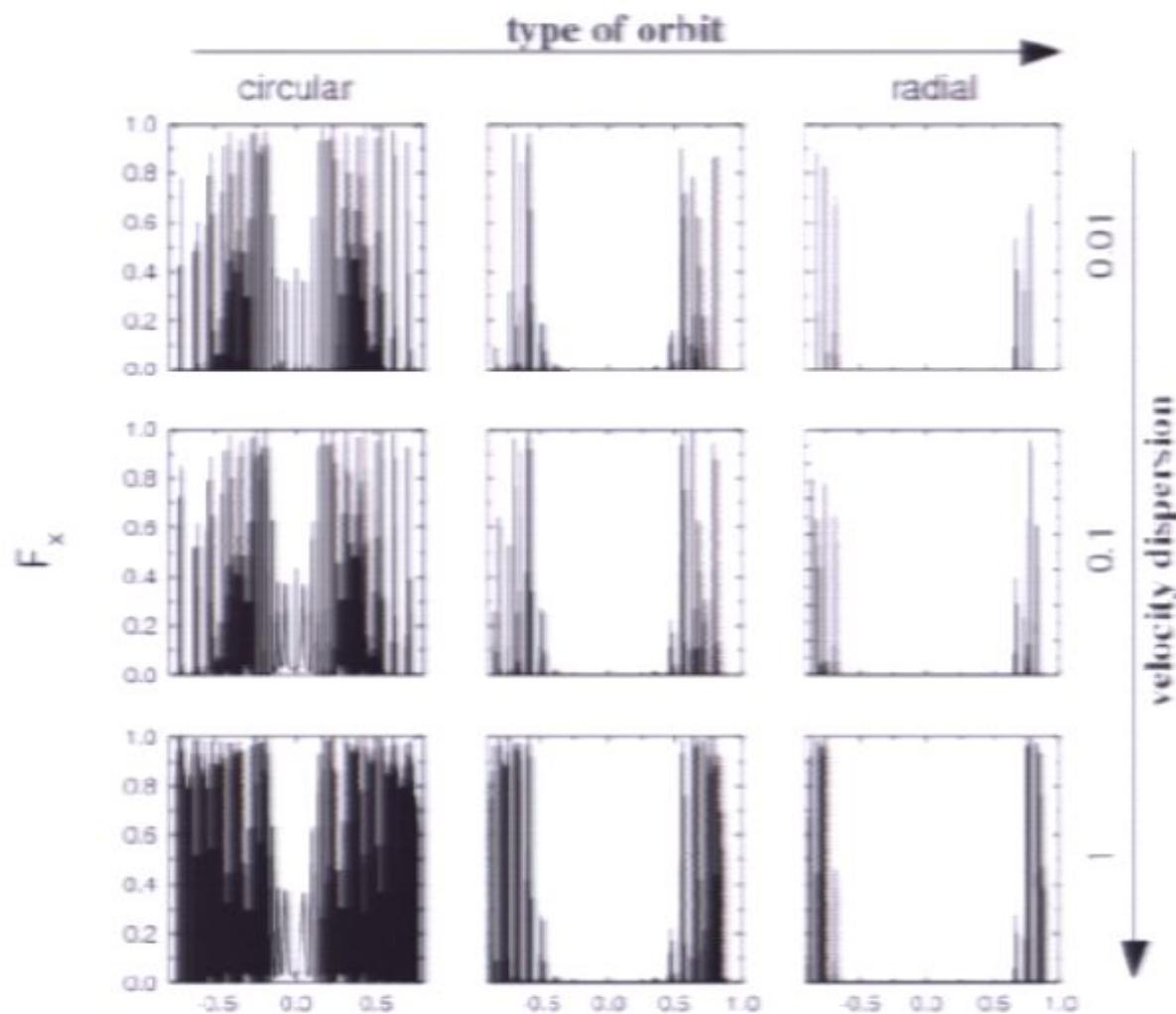
Fantin et al. explore different orbits and effect of initial internal dispersion of the satellite



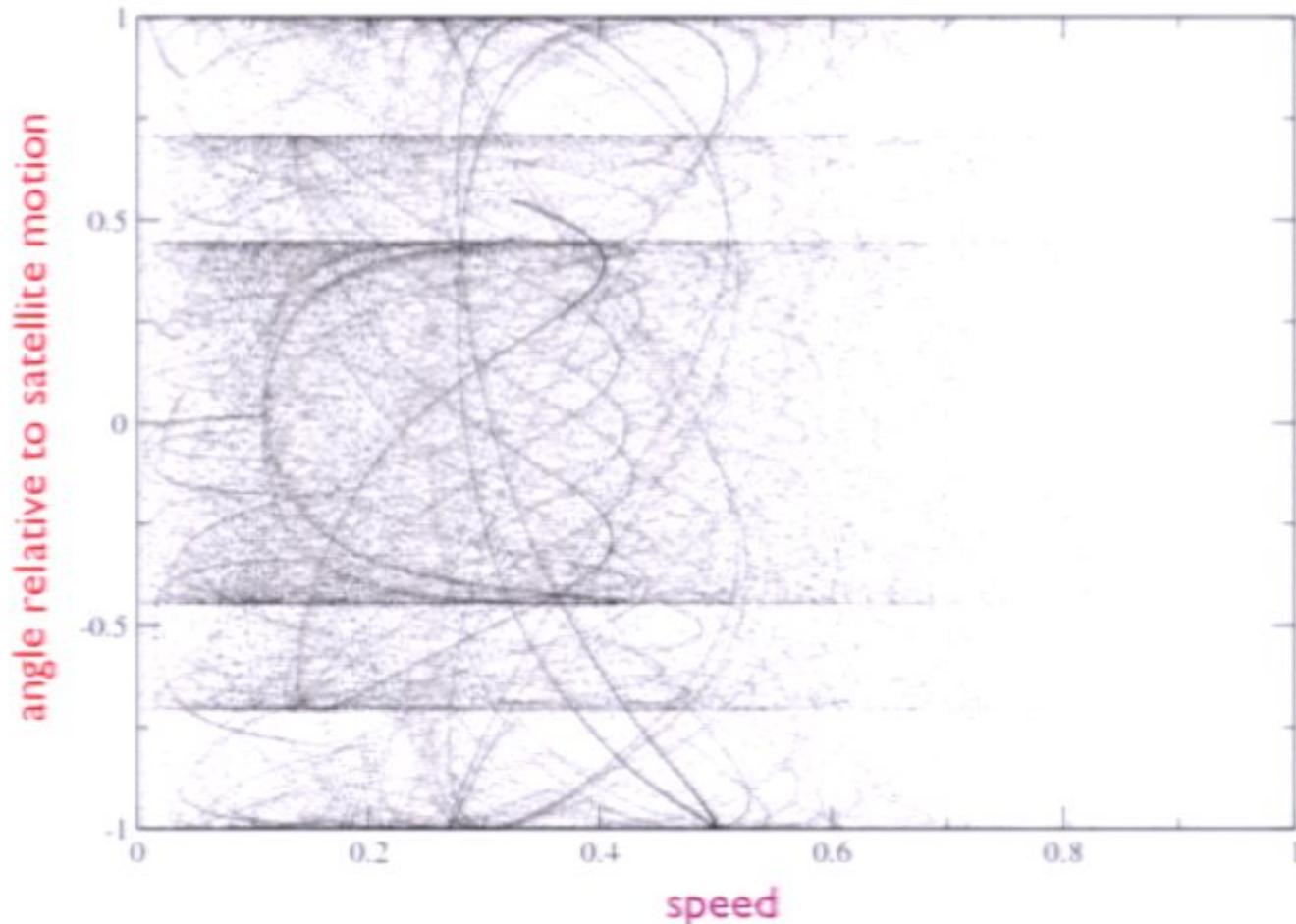
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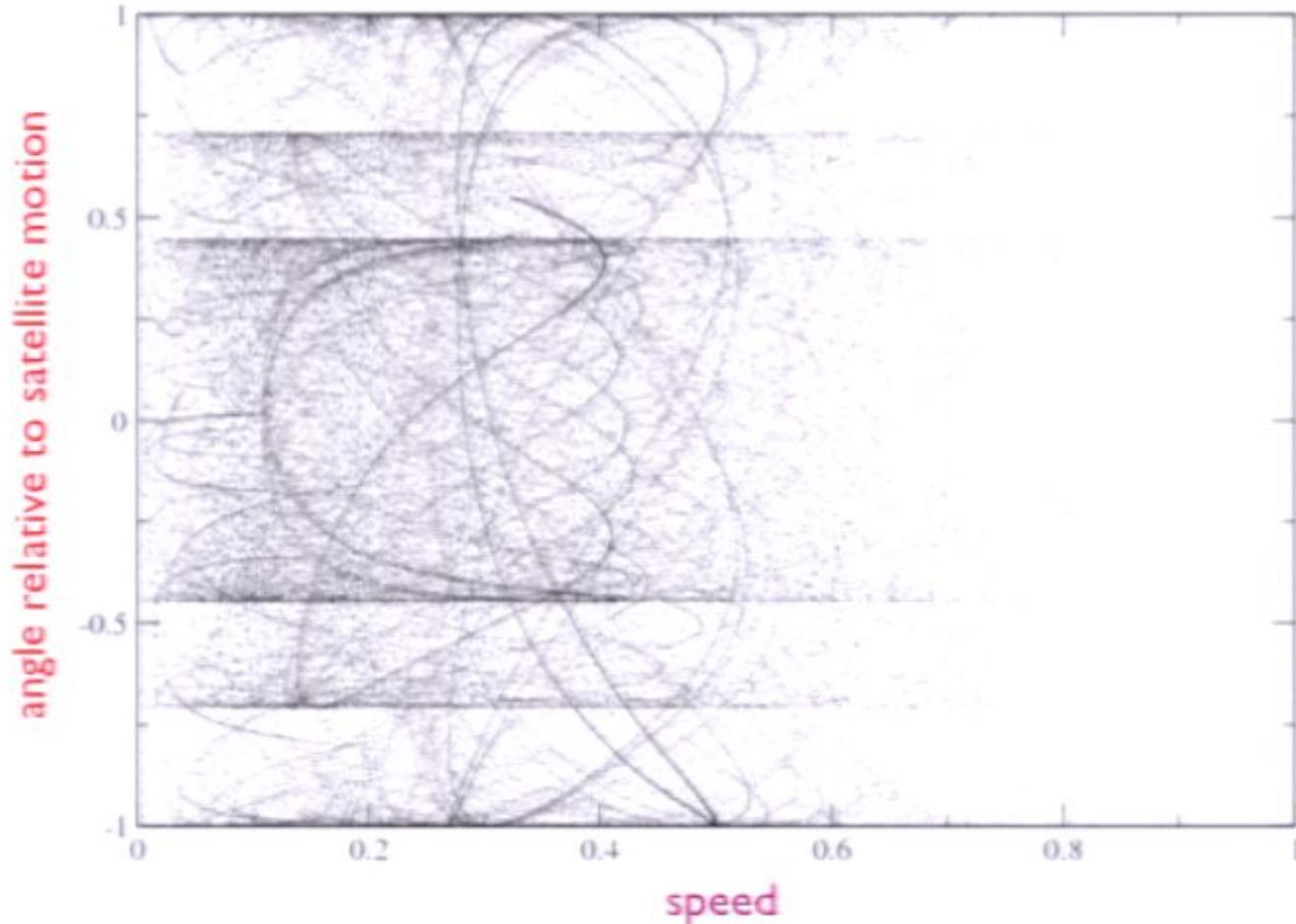


velocity-space distribution in a terrestrial detector due to three subhalos (Fantin et al. 2007)



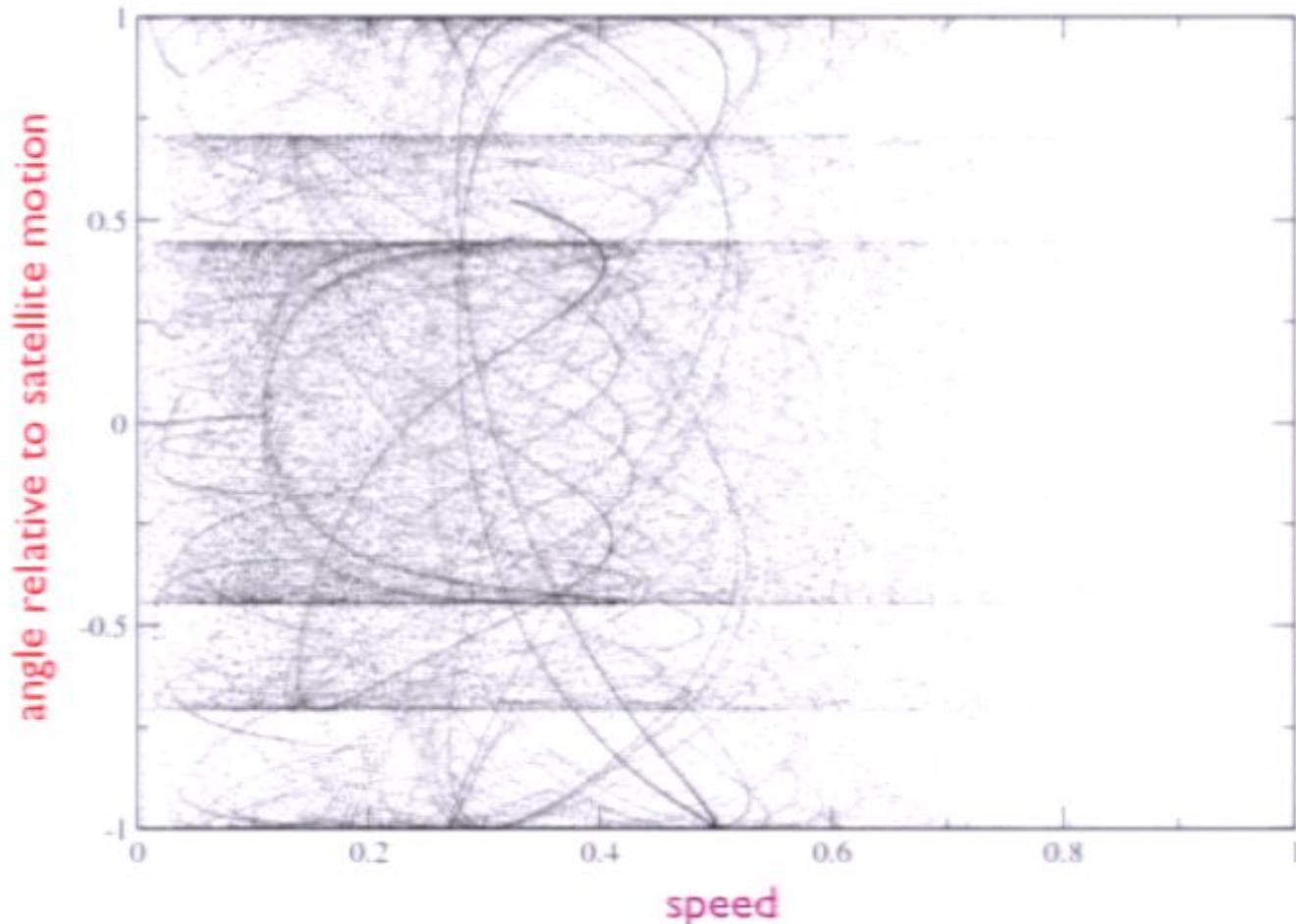
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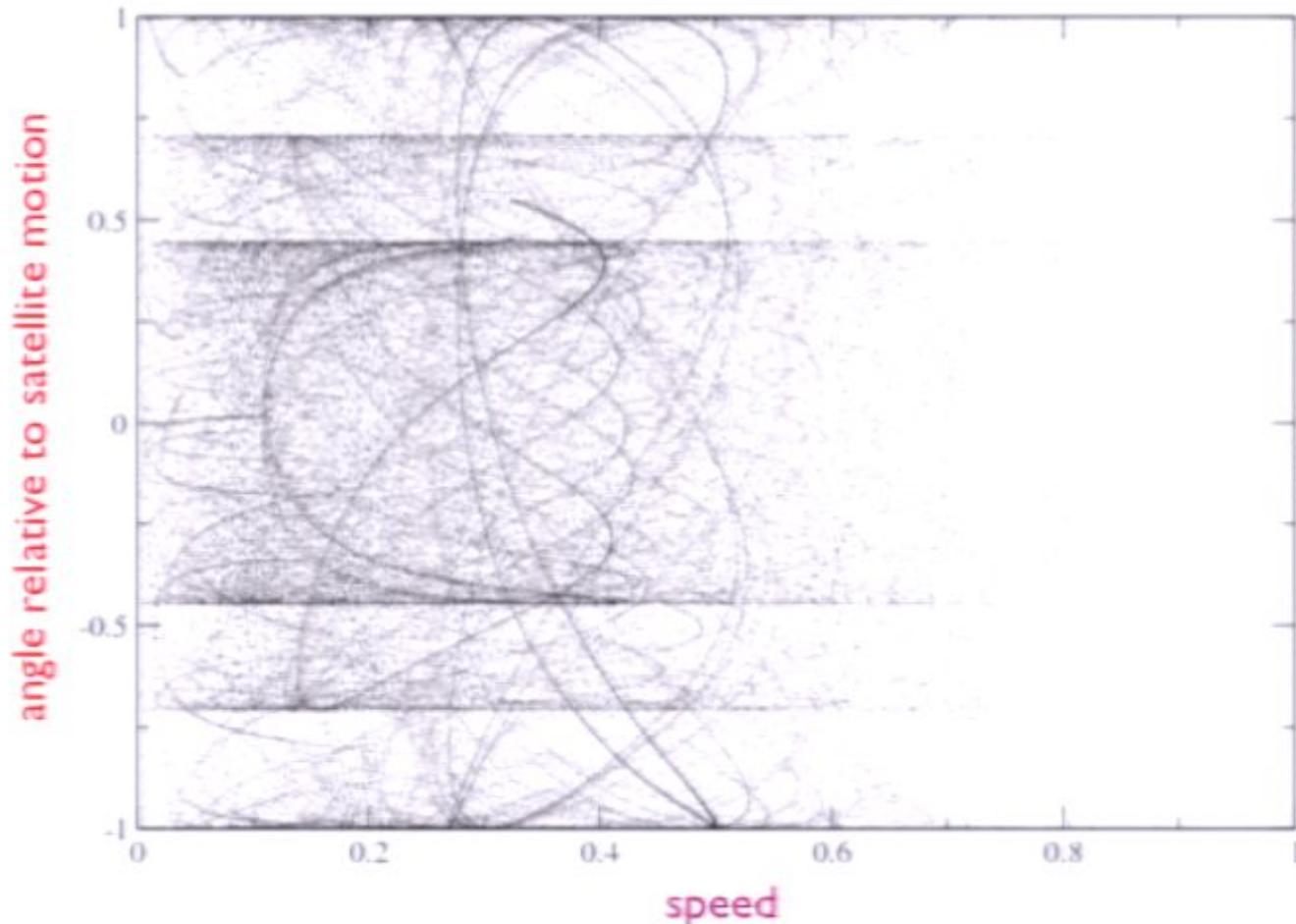
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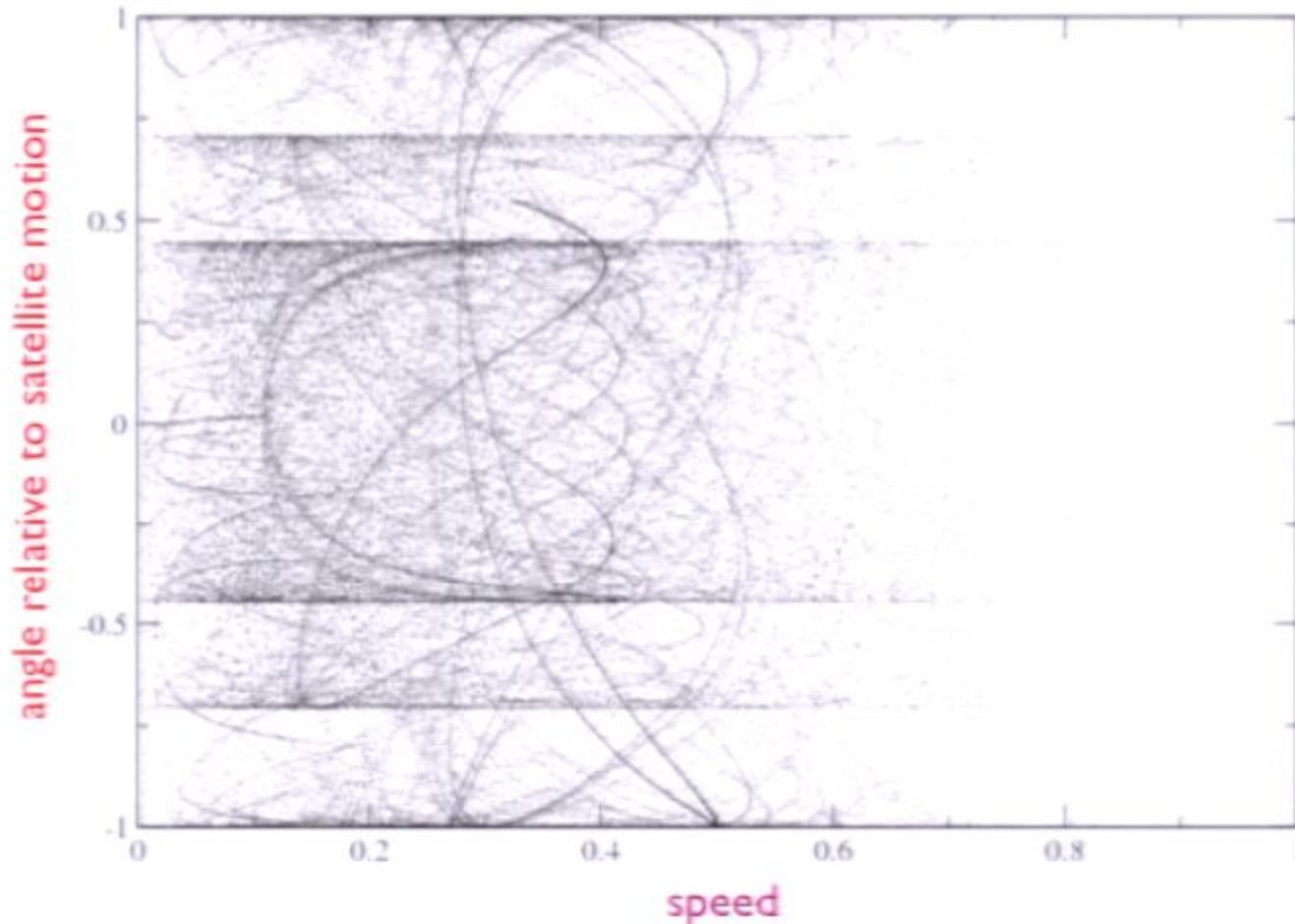
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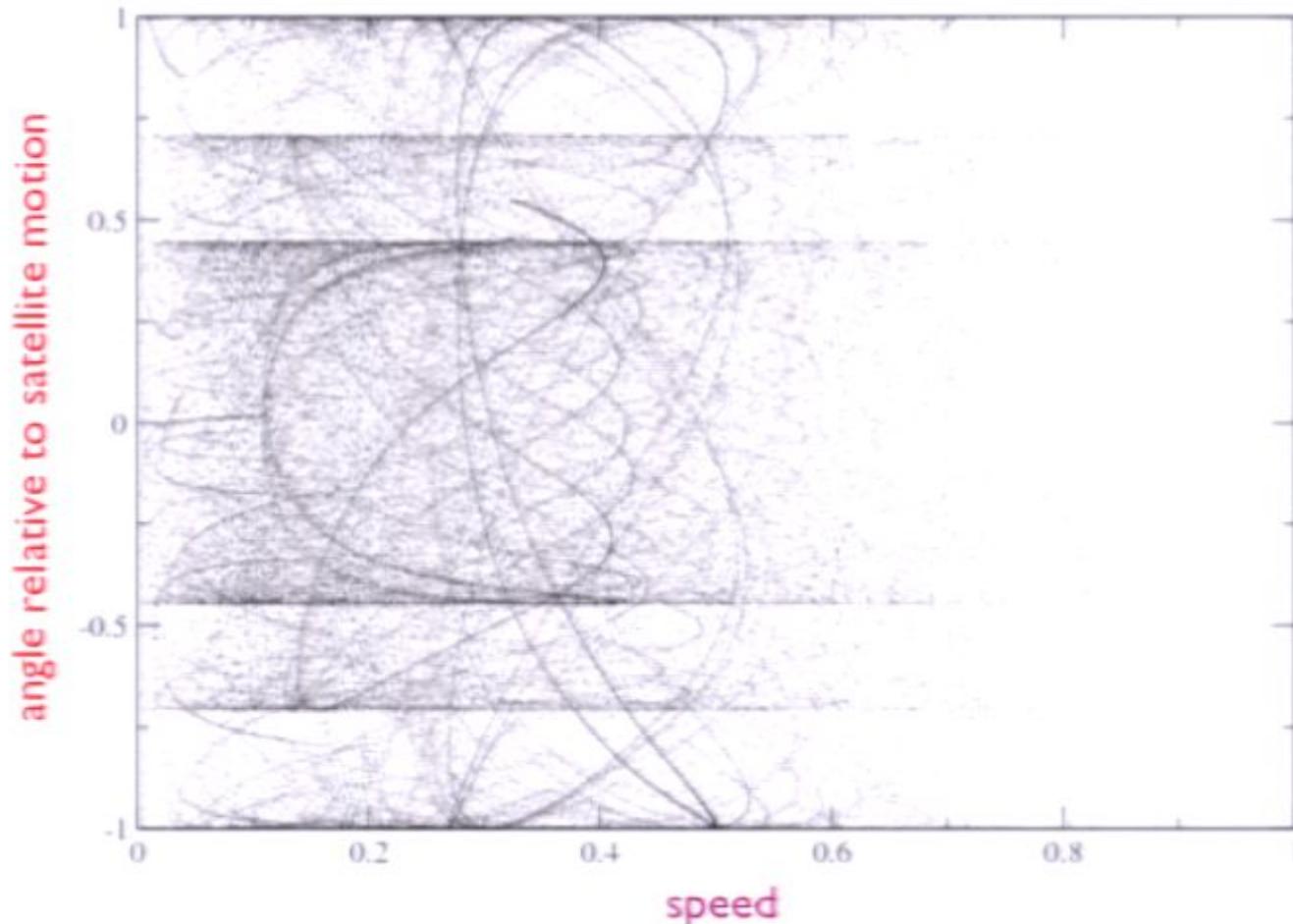
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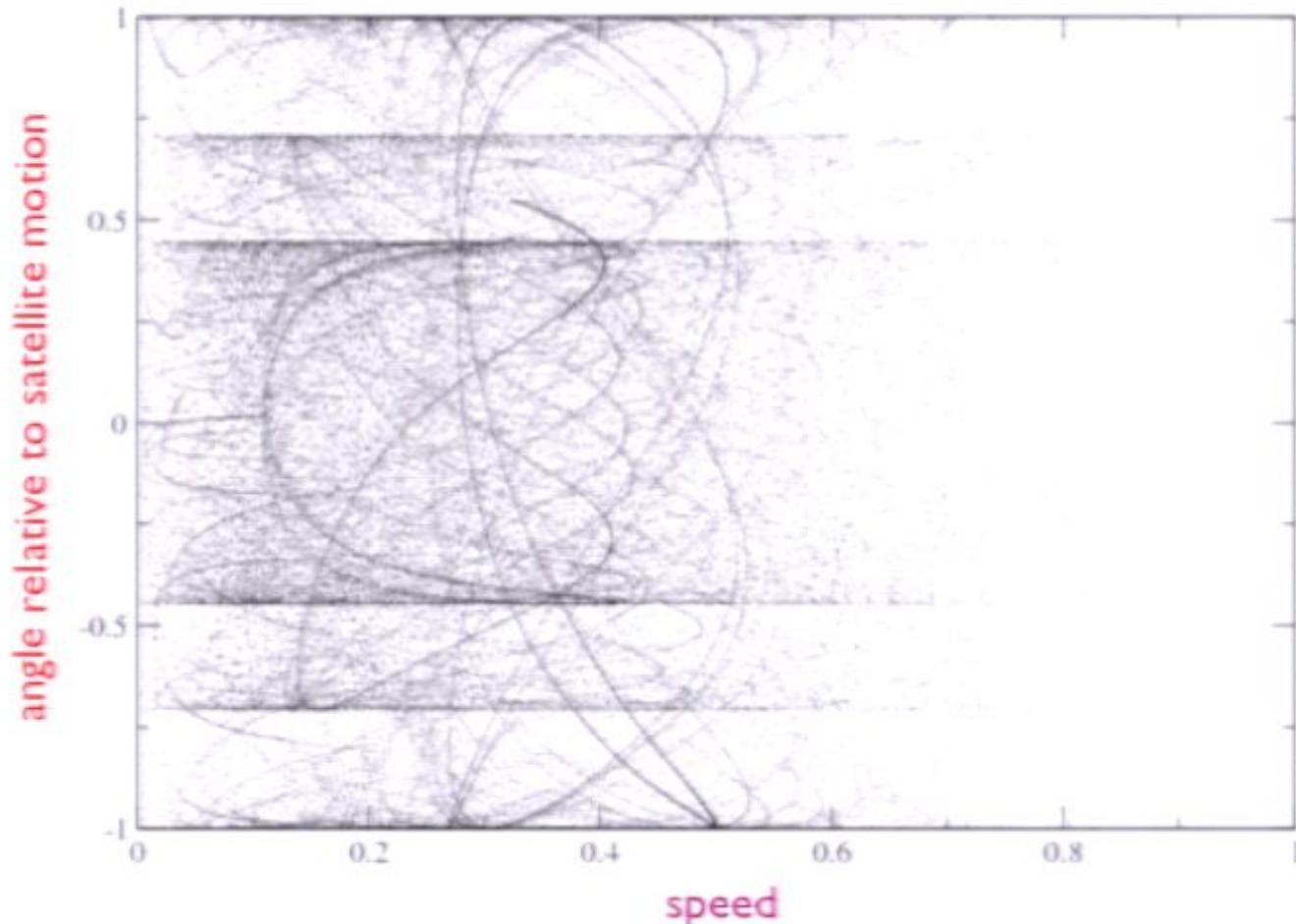
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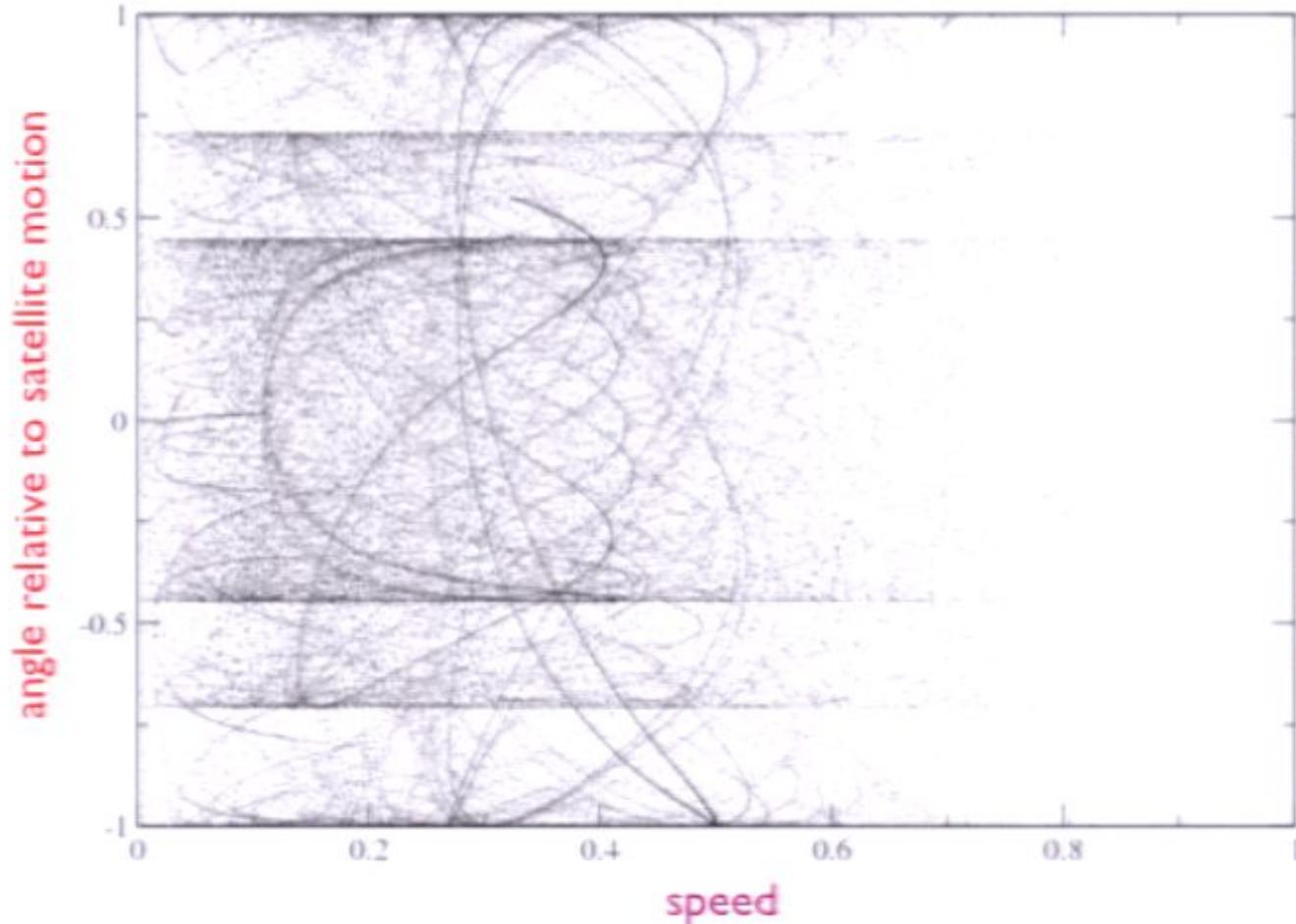
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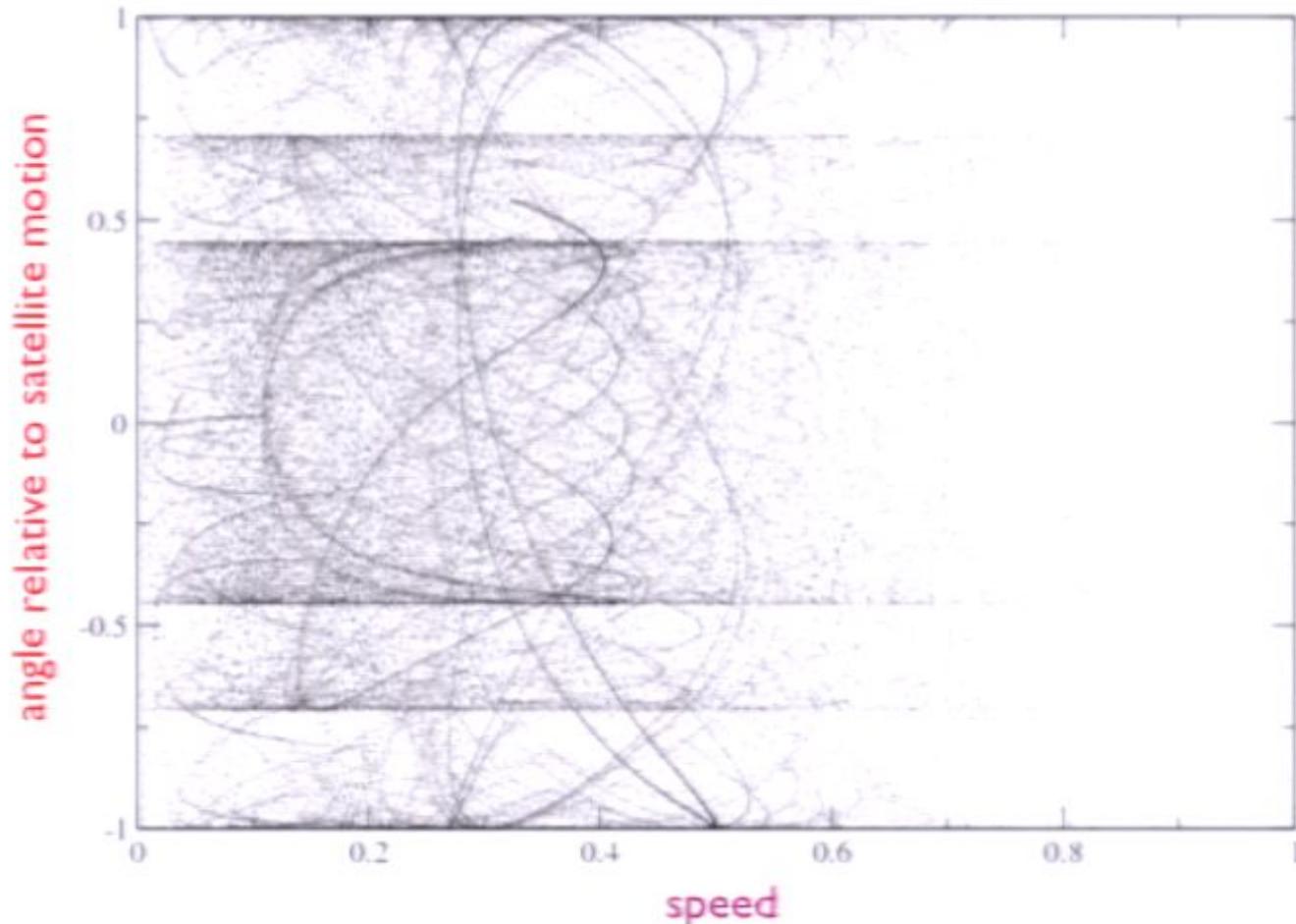
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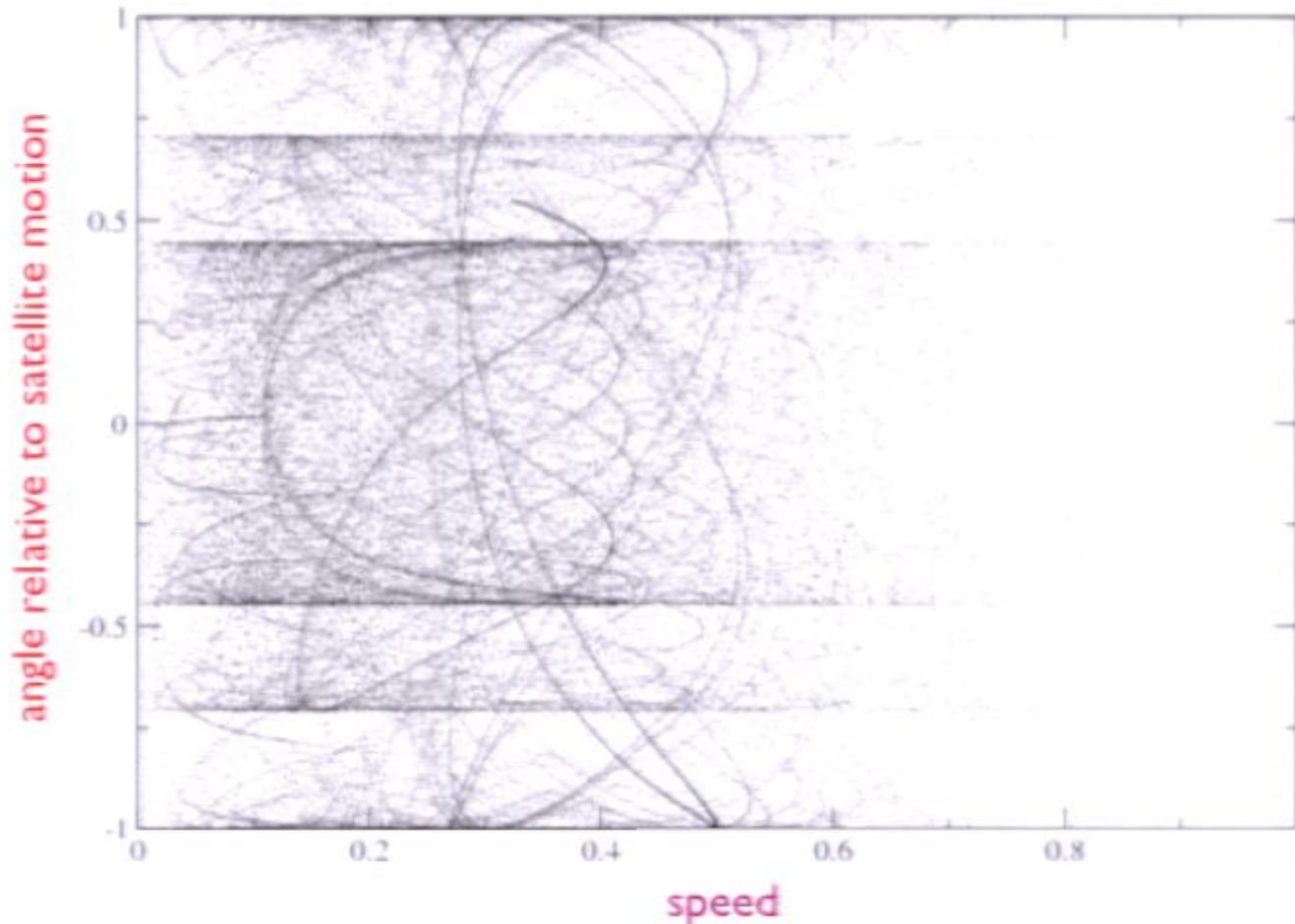
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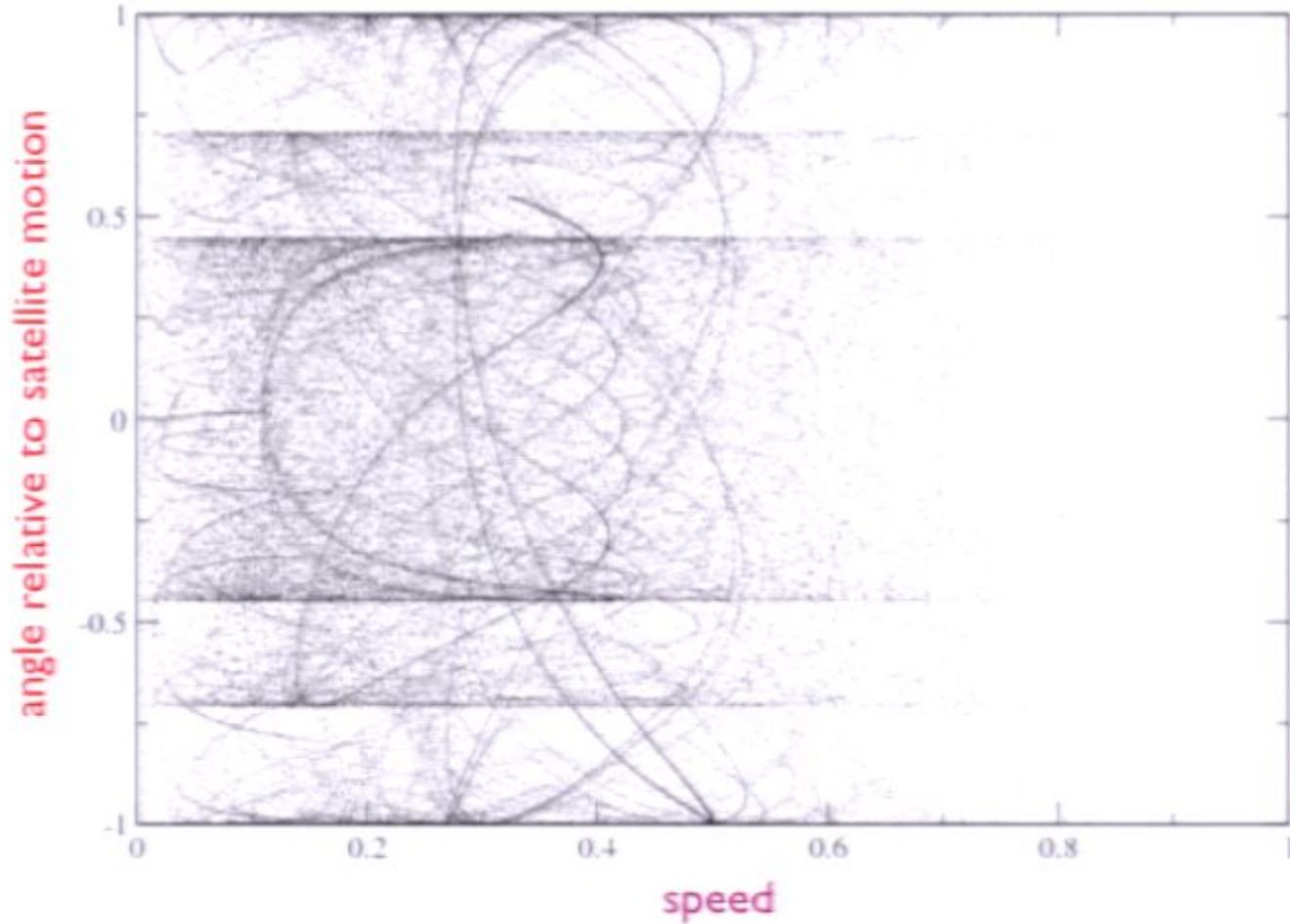
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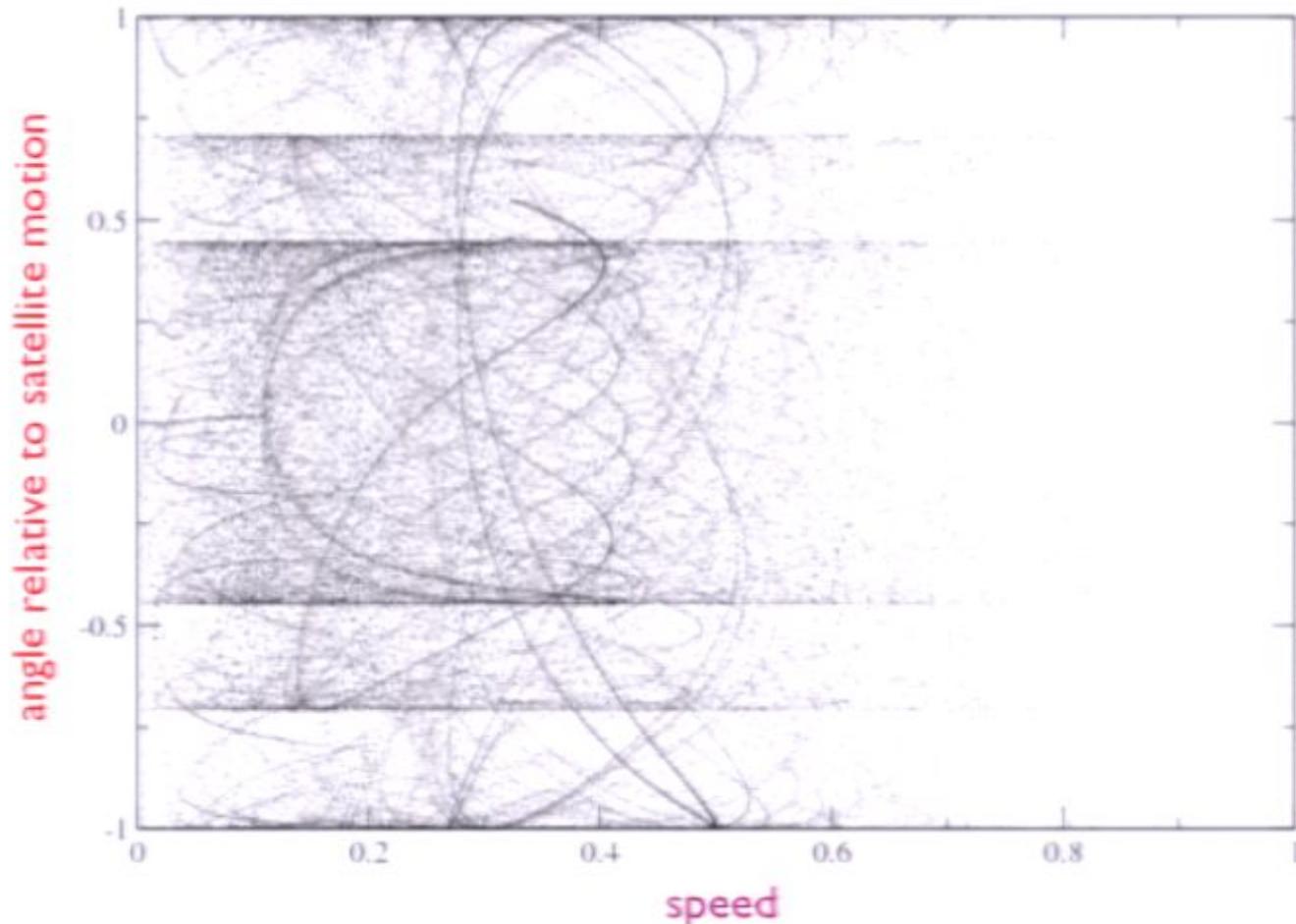
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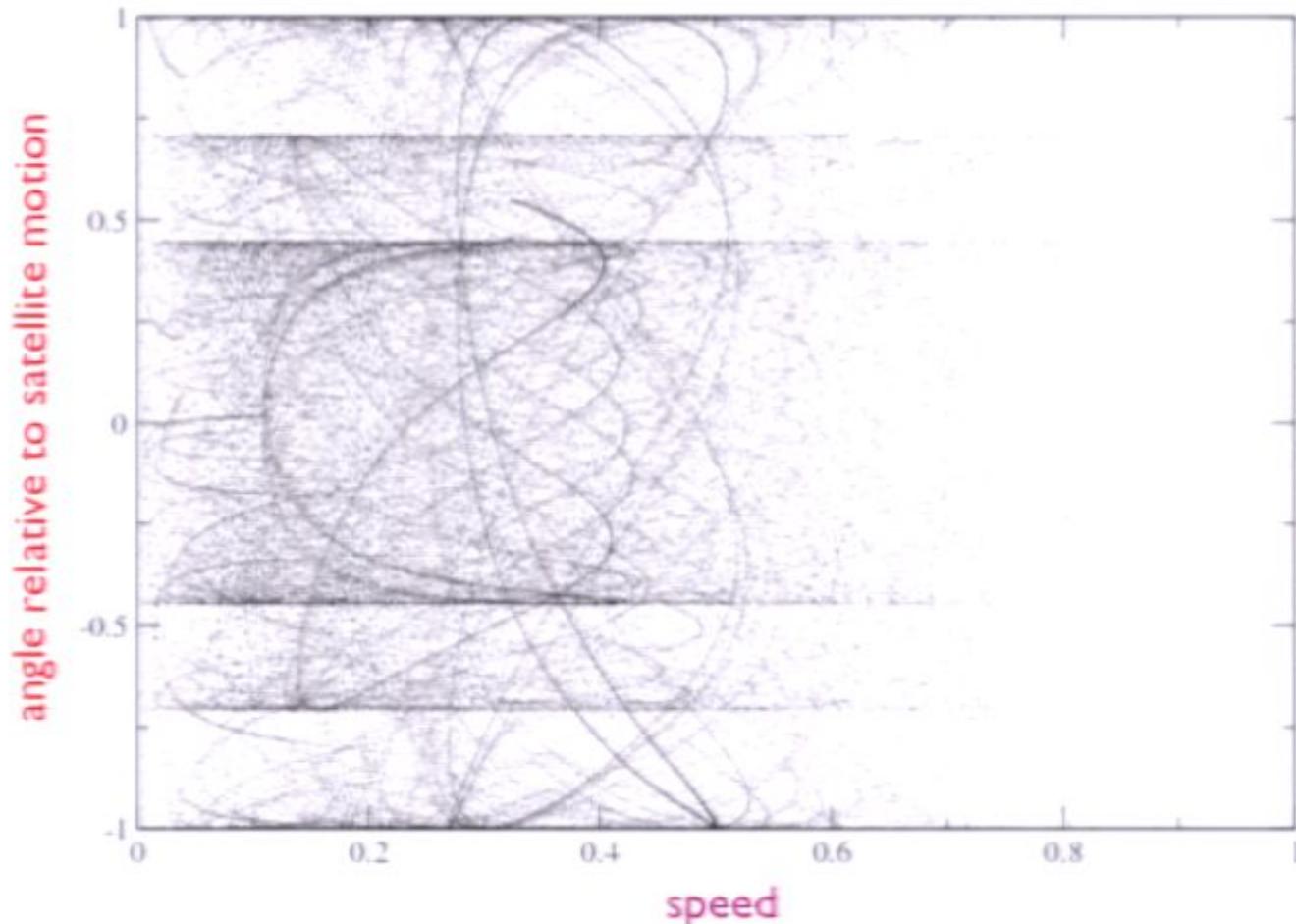
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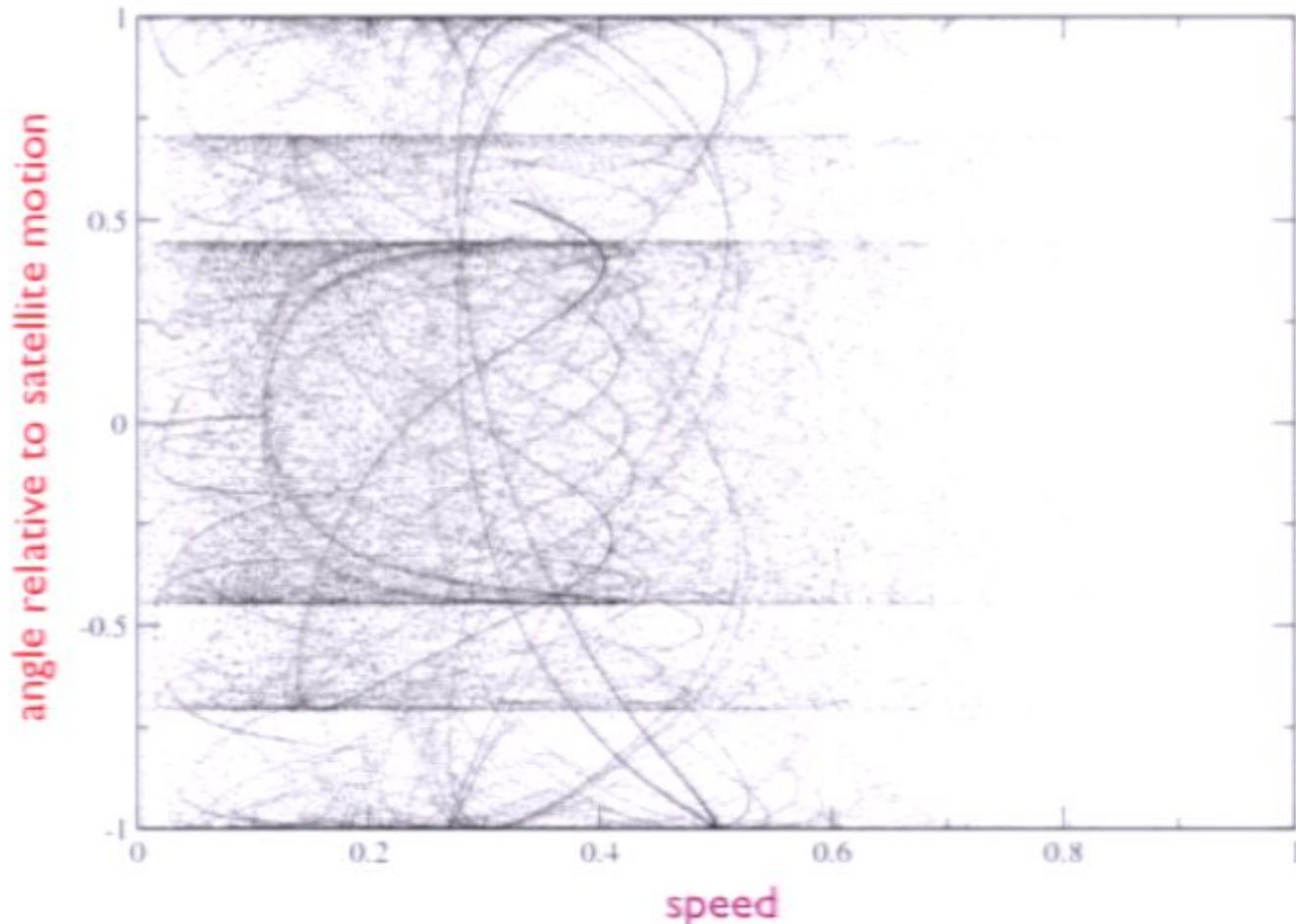
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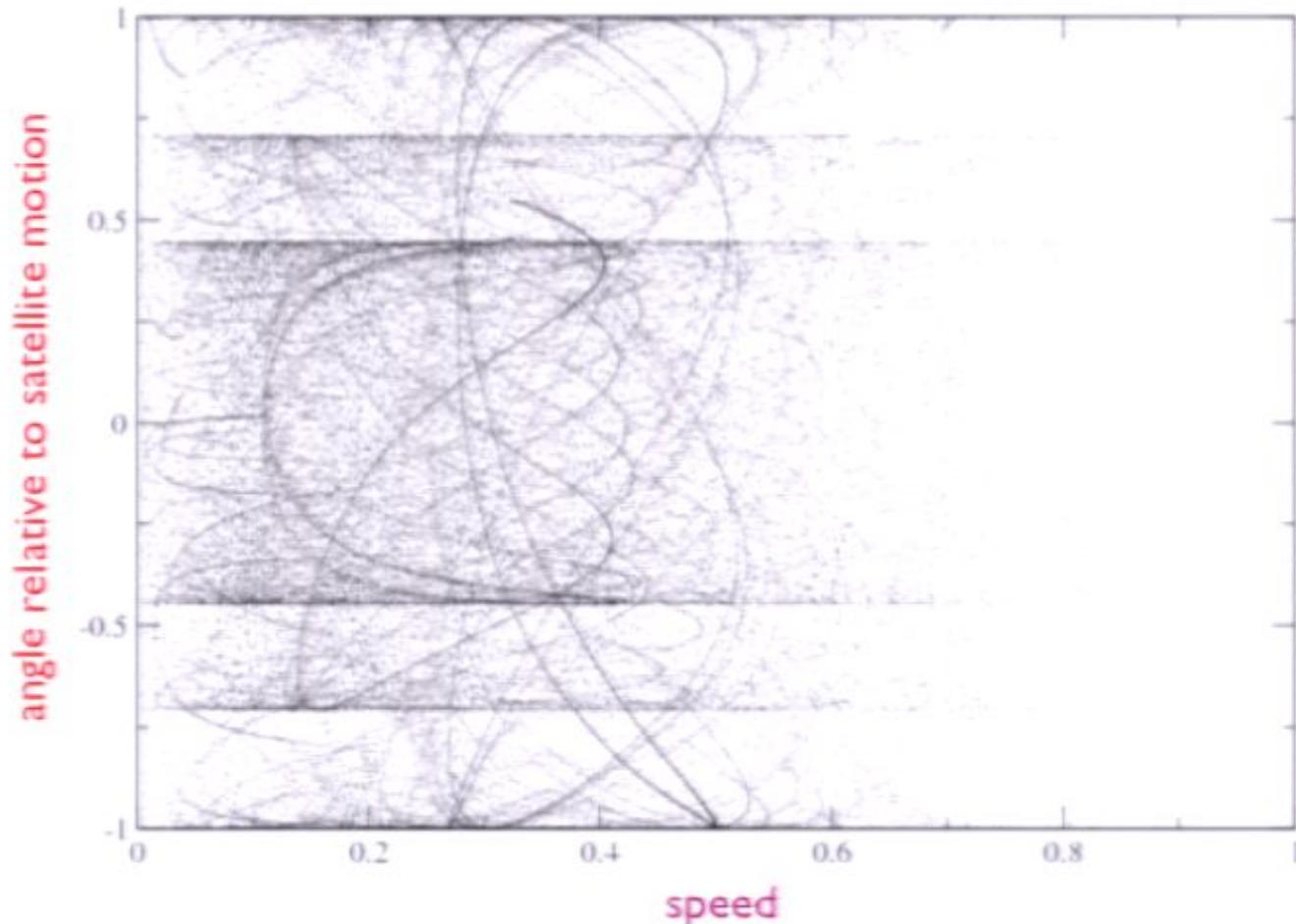
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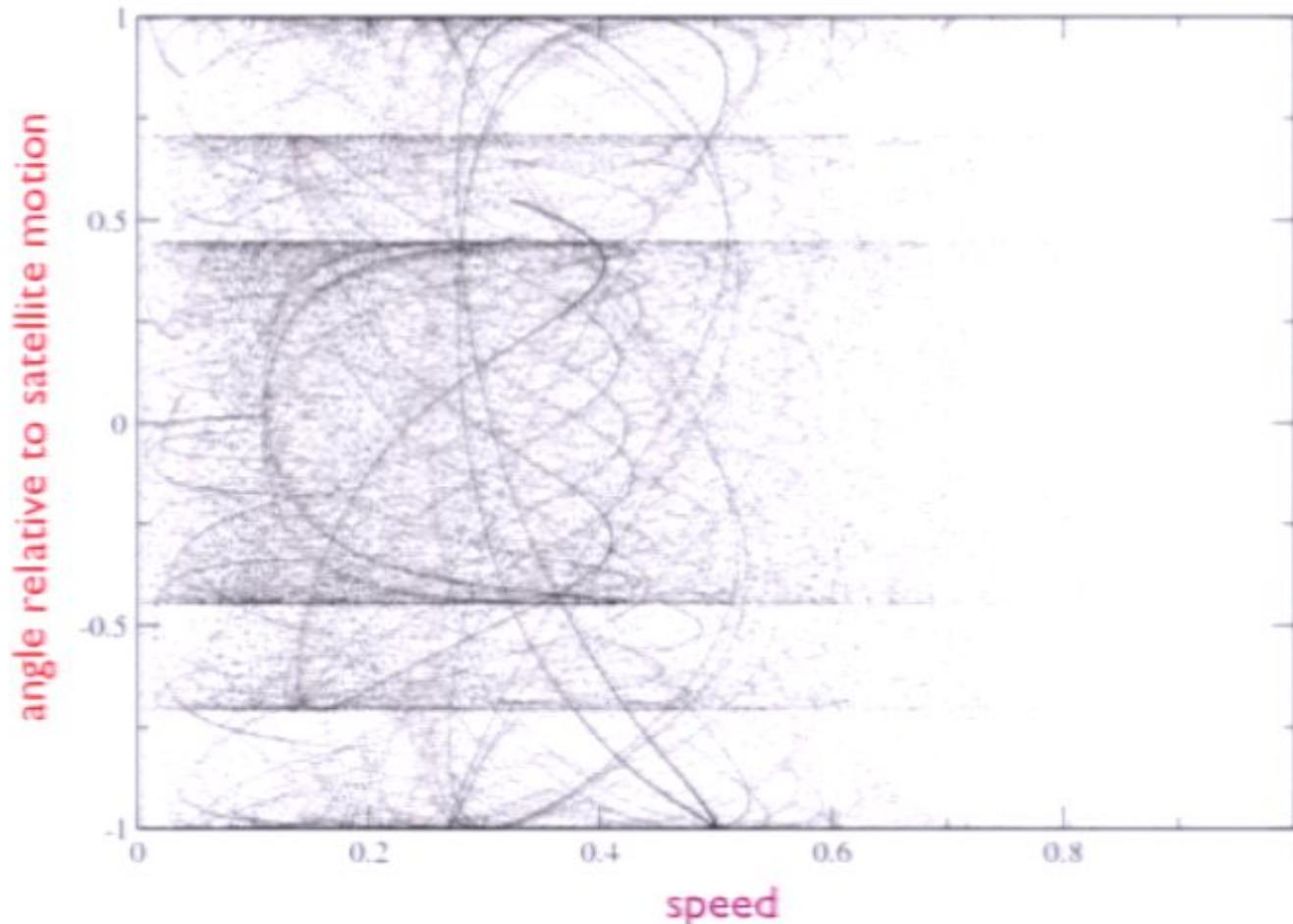
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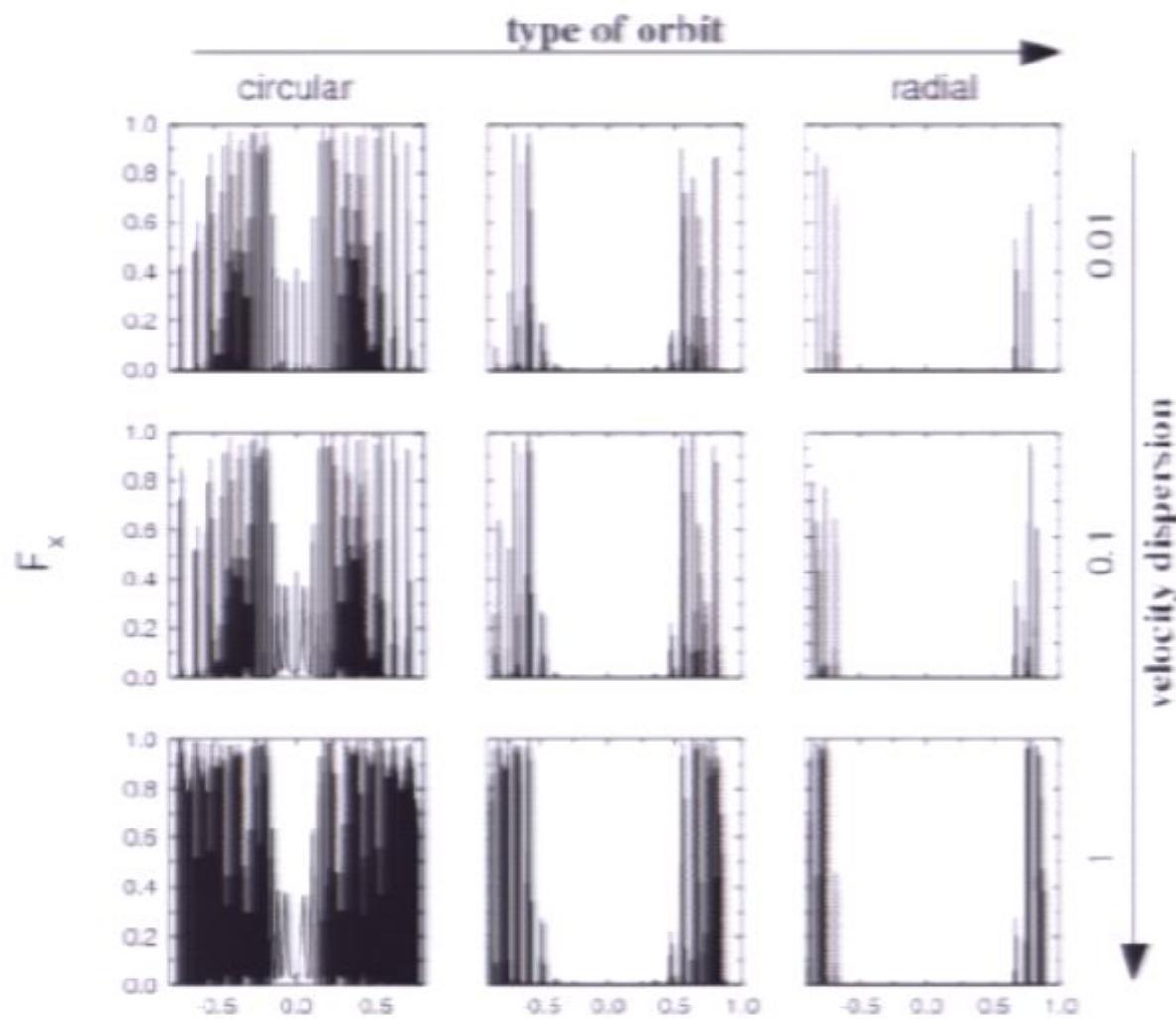
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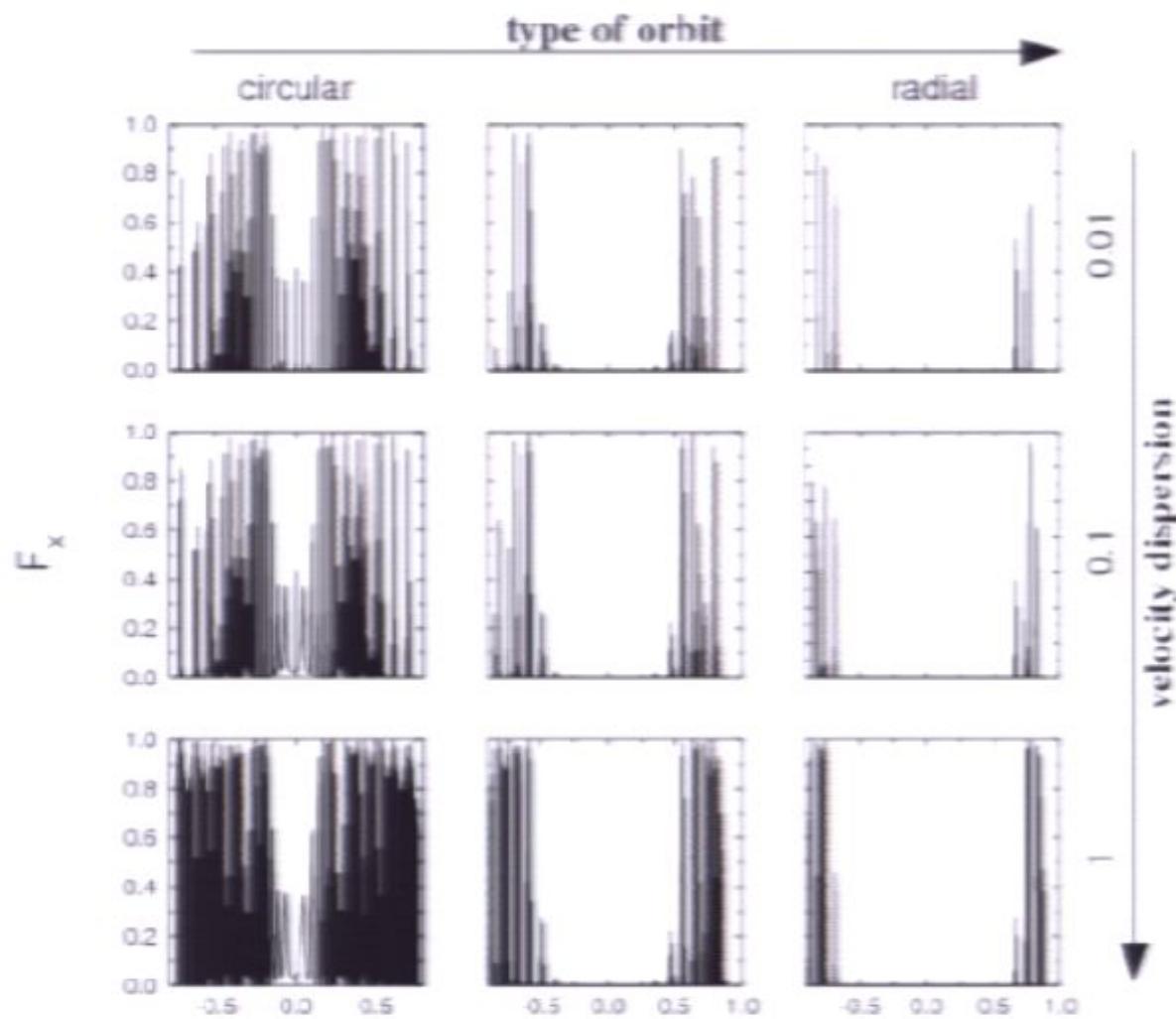


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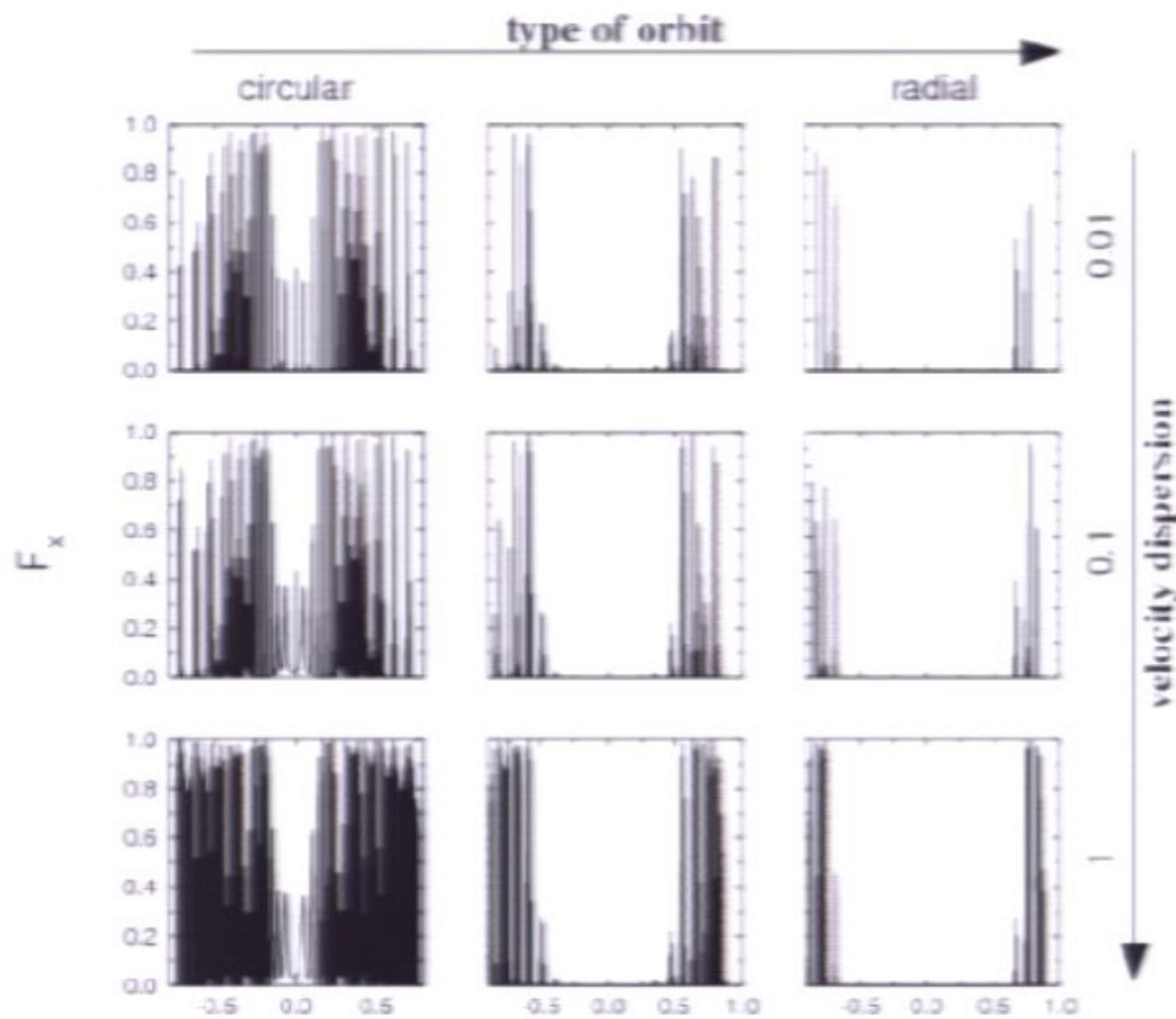
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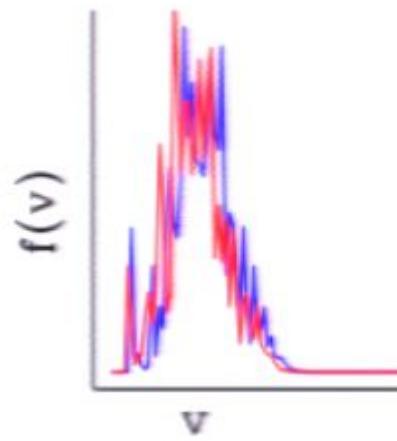
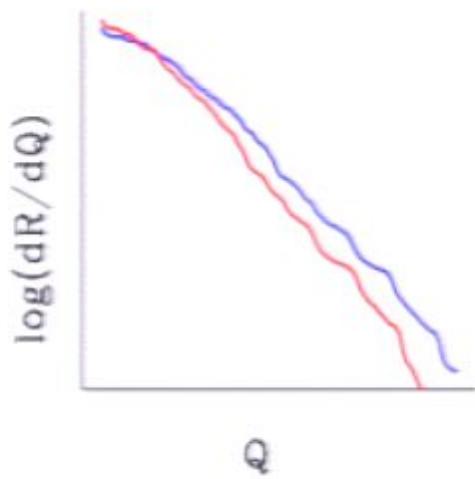
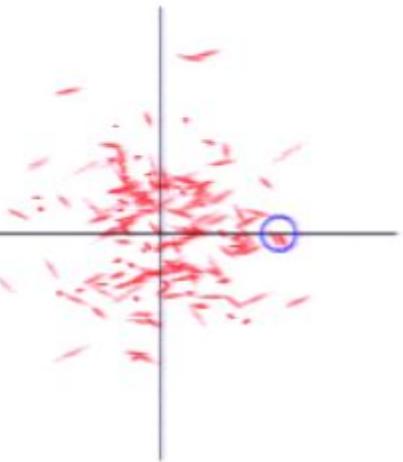


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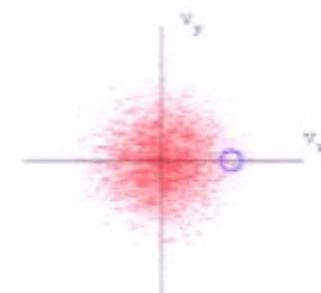




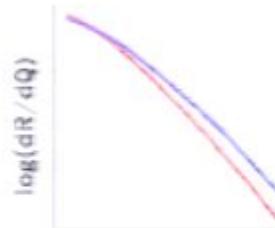
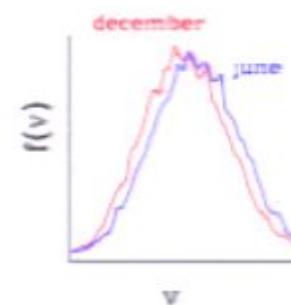
Possible phenomenological model for local $f(v)$

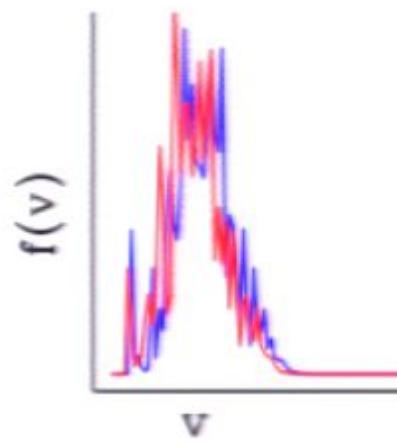
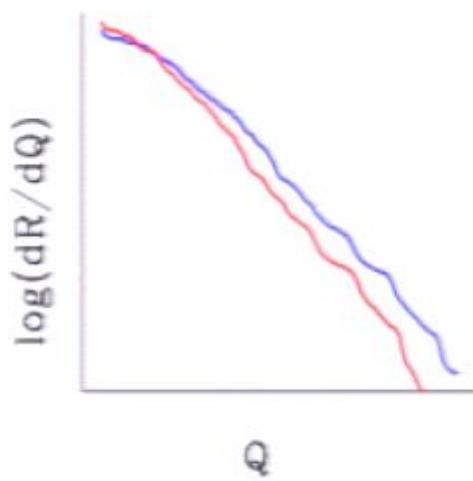
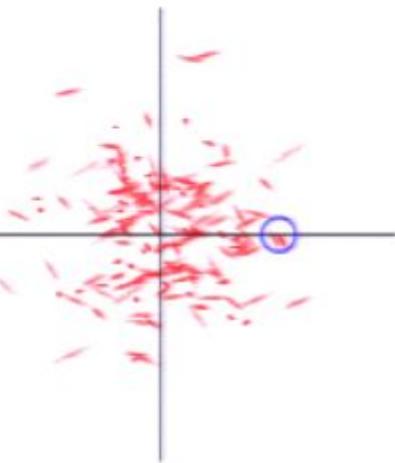
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standard Maxwellian

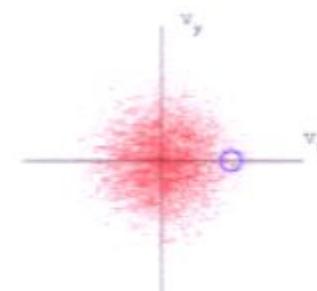




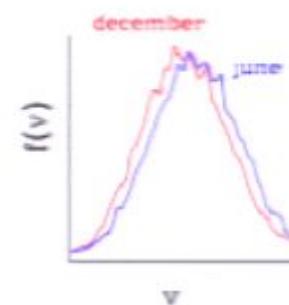
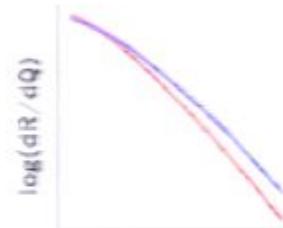
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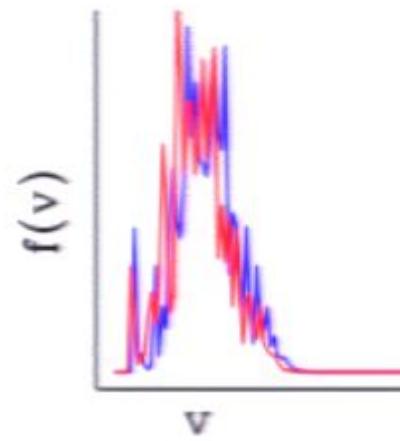
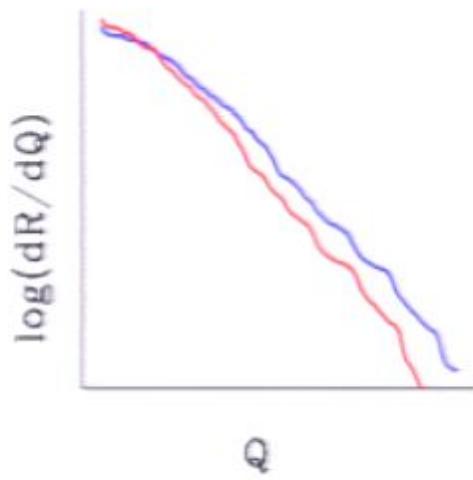
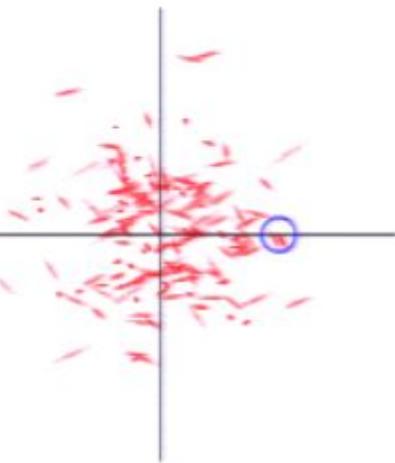
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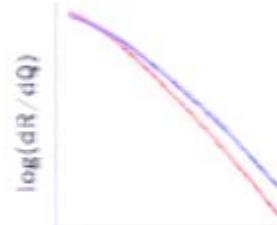
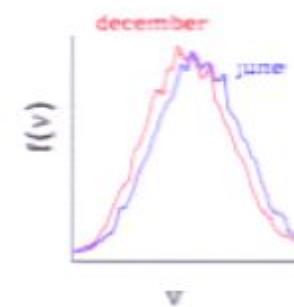
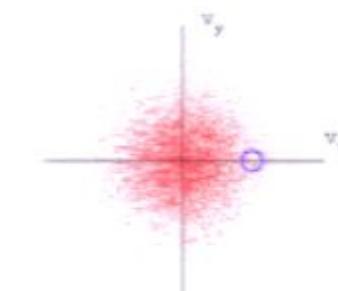


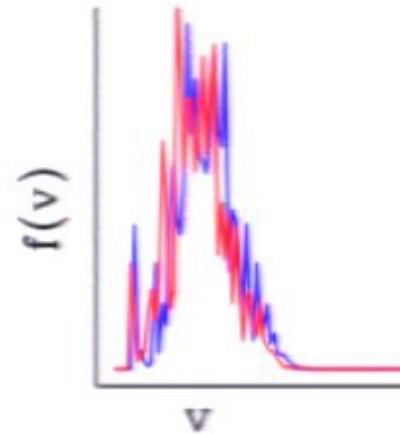
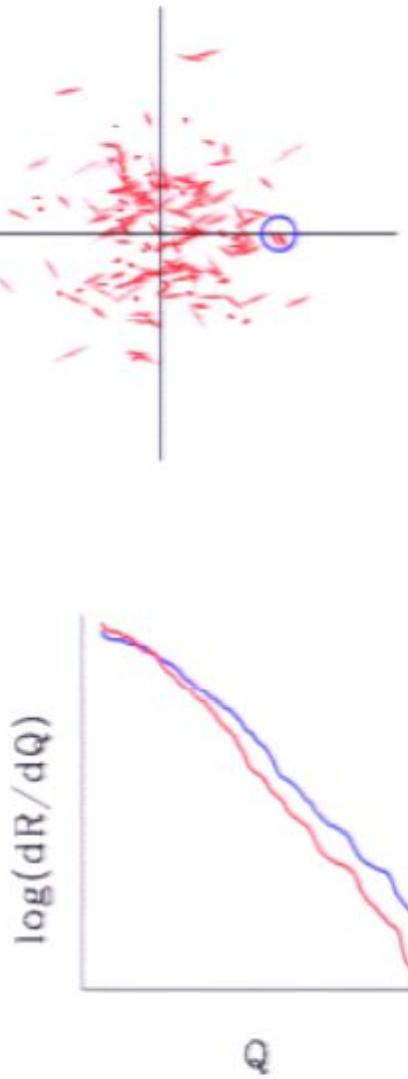


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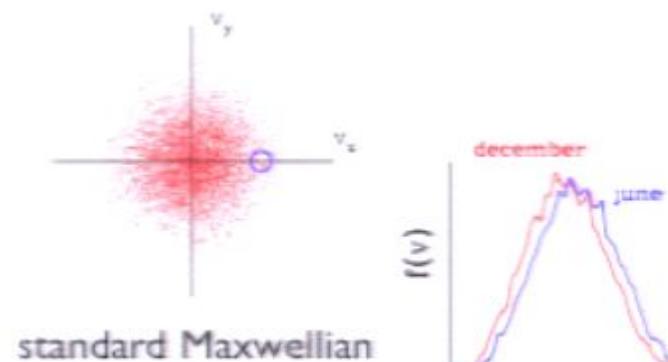




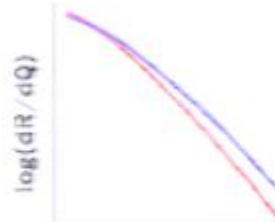
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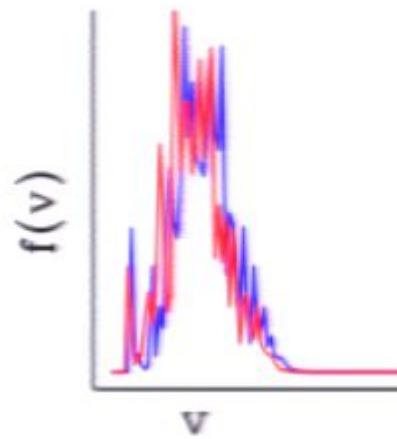
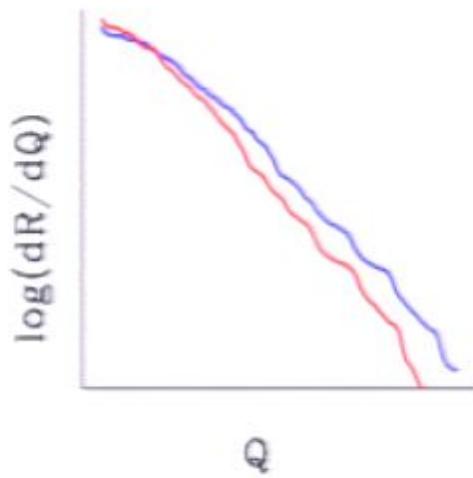
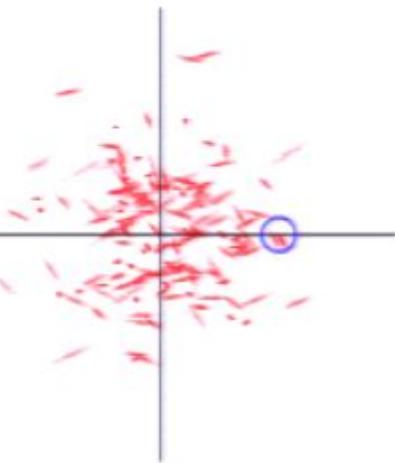
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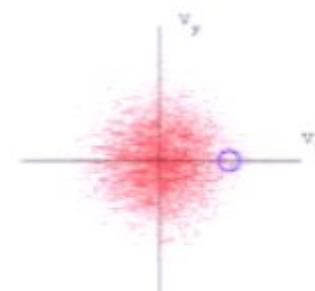




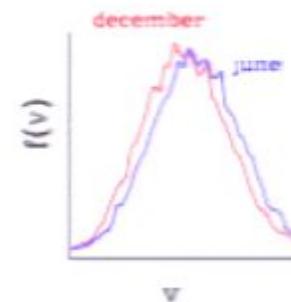
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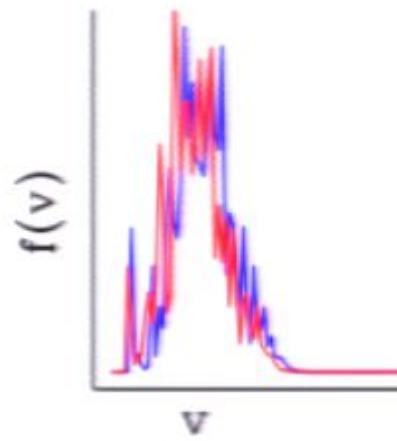
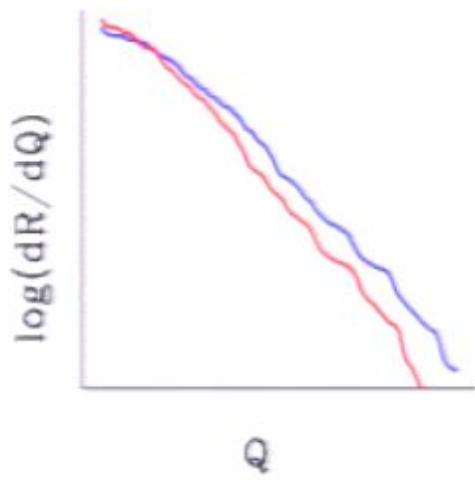
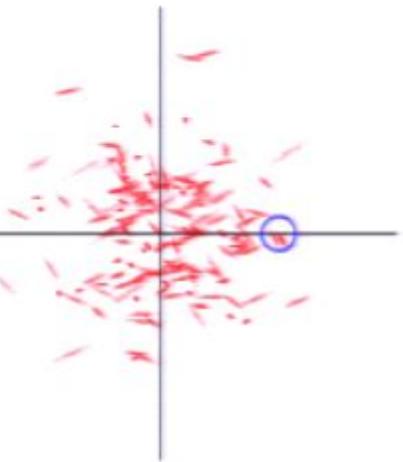
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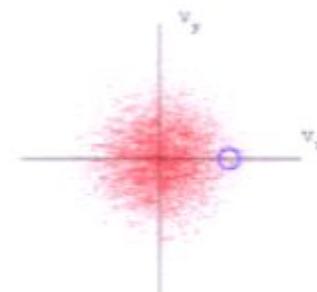




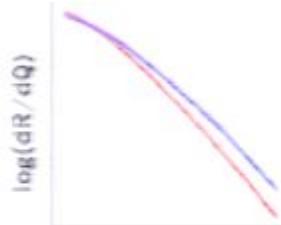
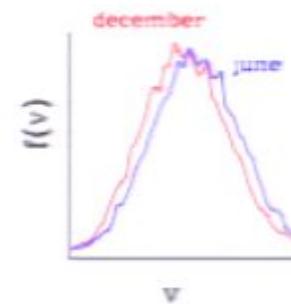
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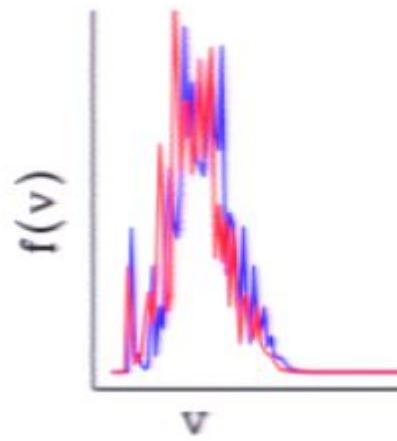
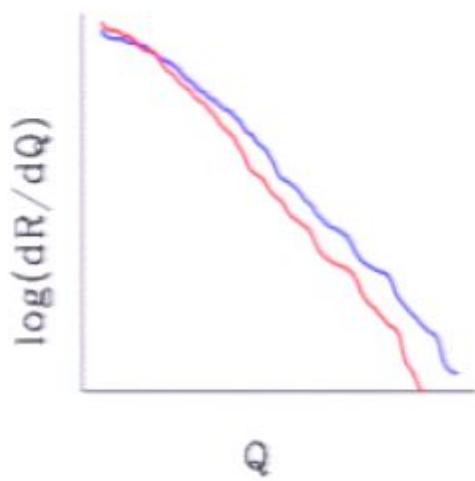
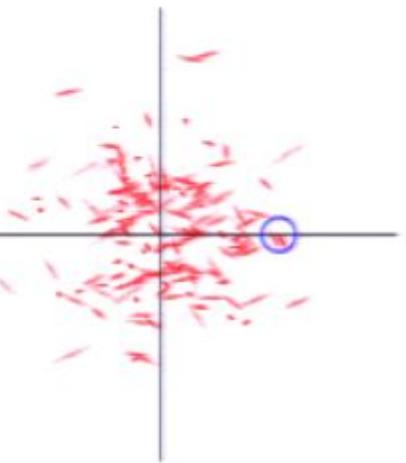
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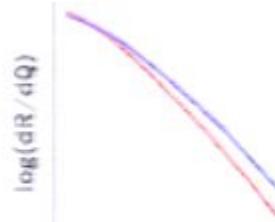
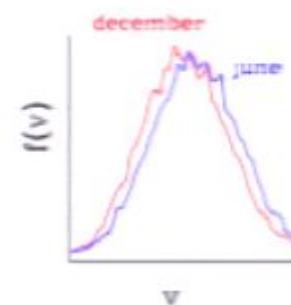
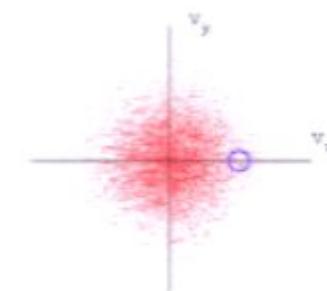


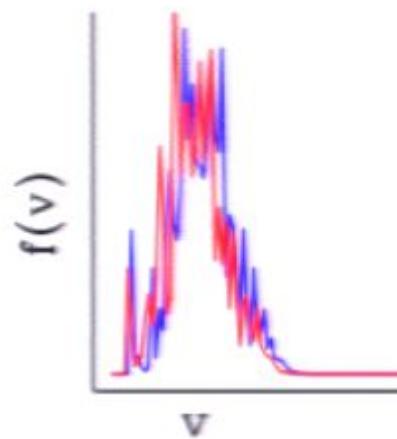
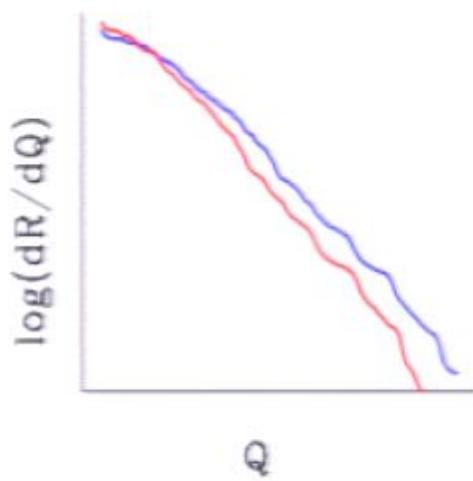
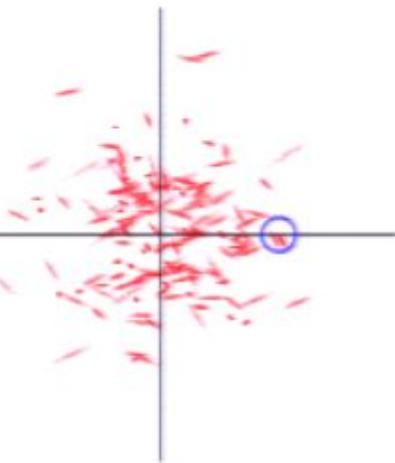


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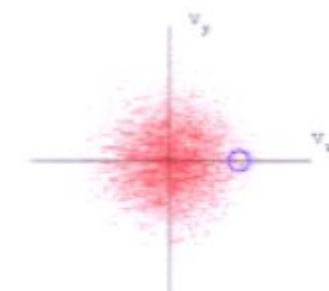




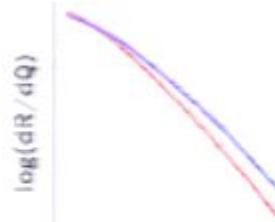
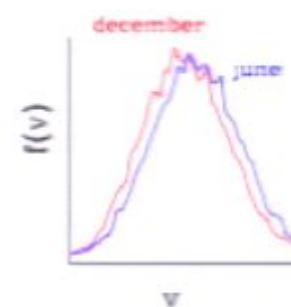
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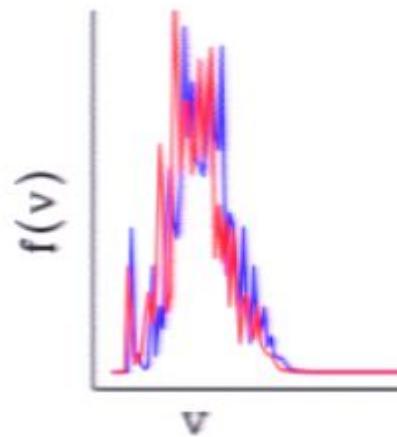
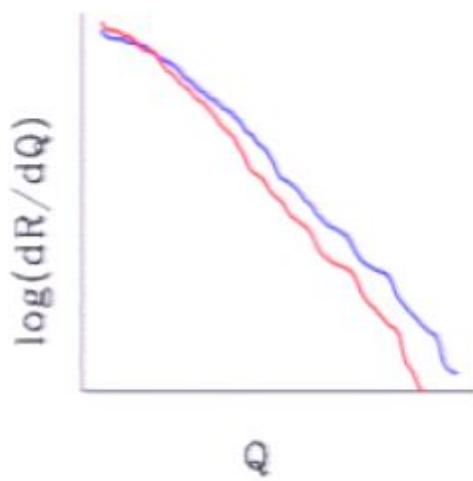
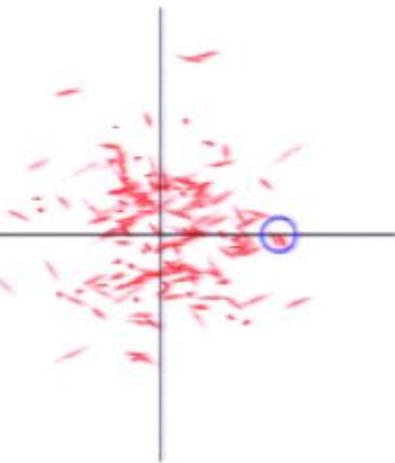
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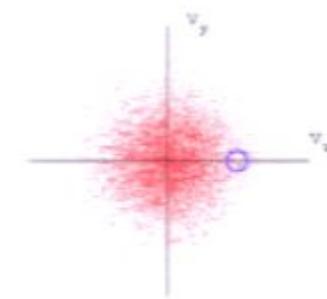




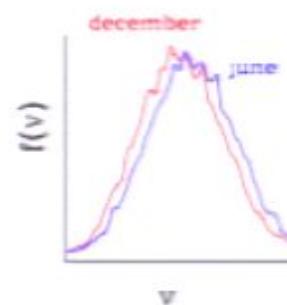
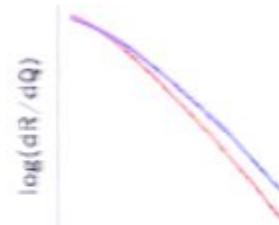
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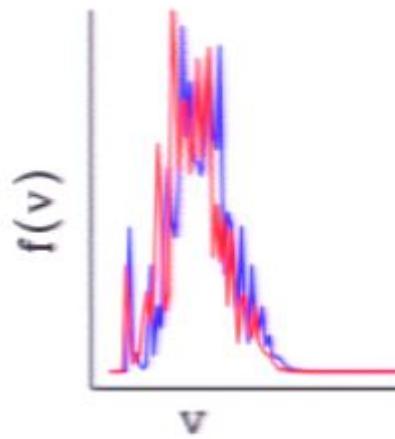
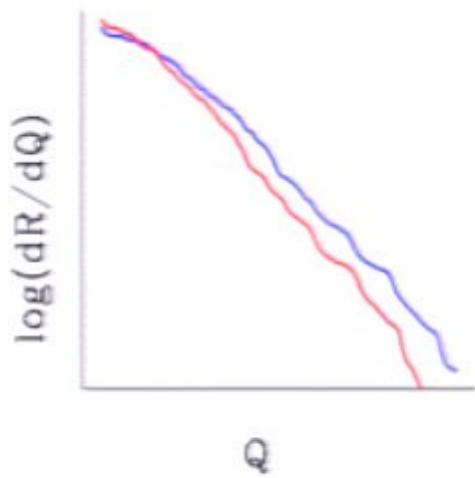
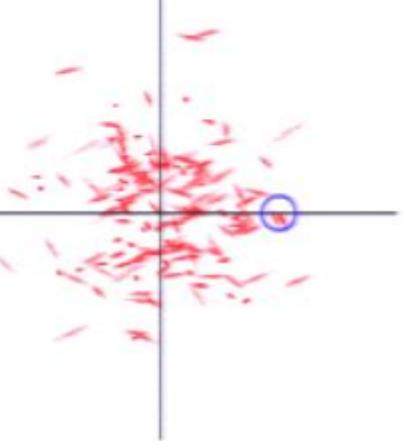
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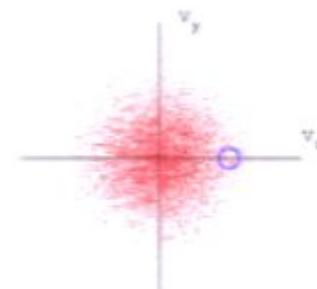




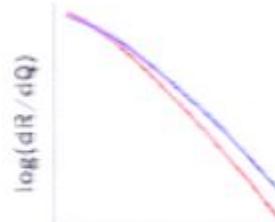
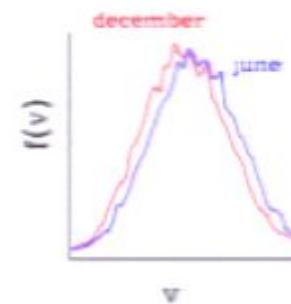
Possible phenomenological model for local $f(v)$

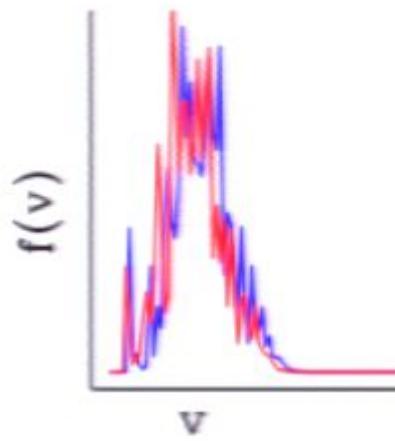
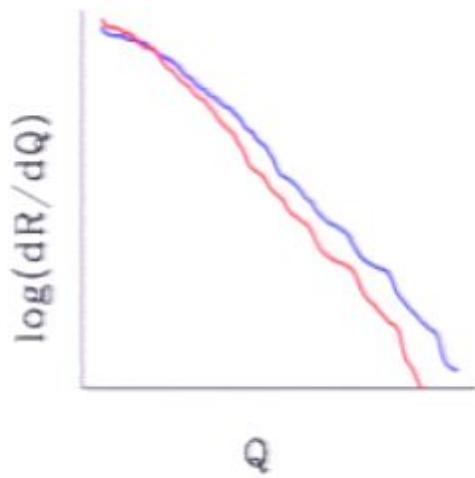
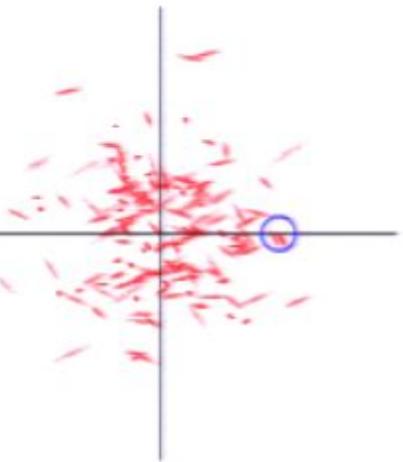
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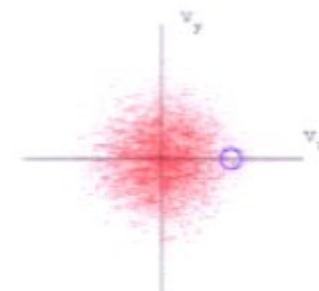




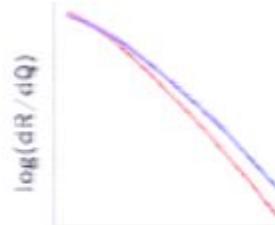
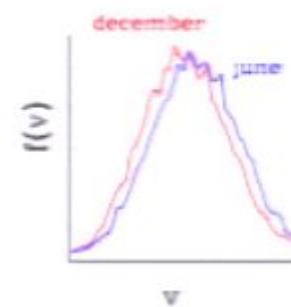
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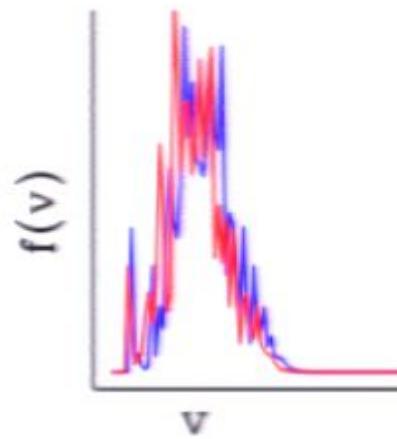
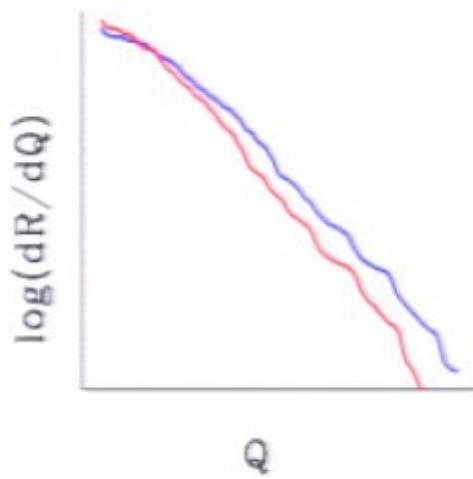
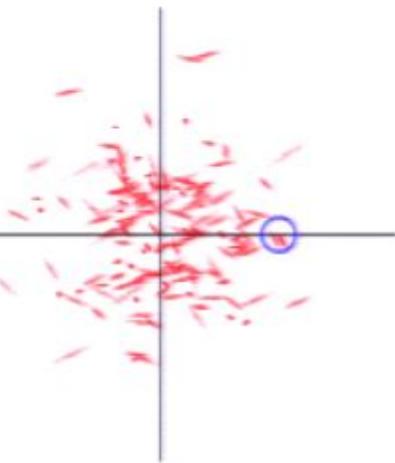
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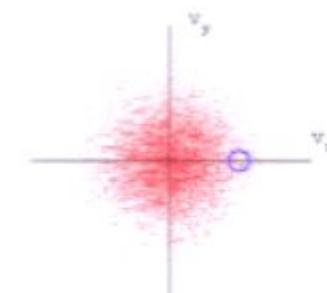




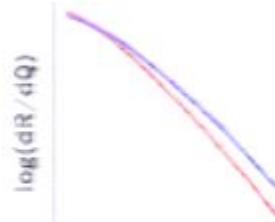
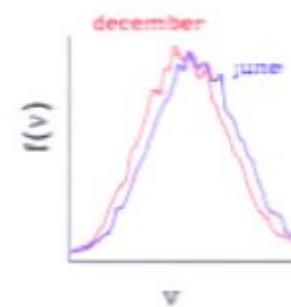
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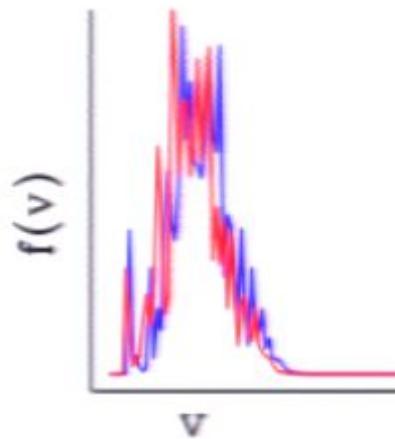
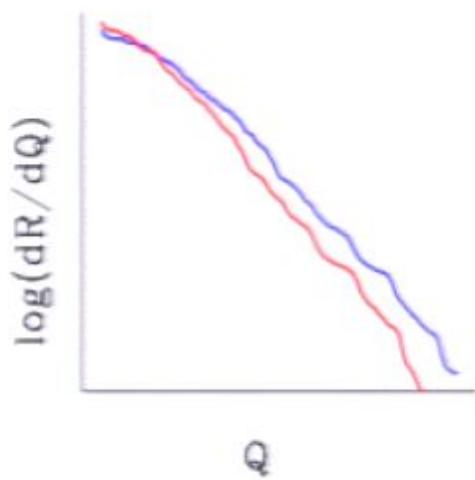
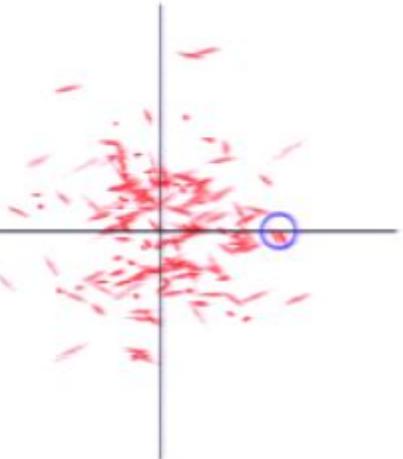
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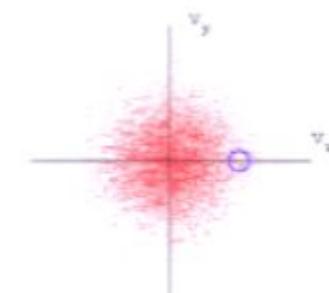




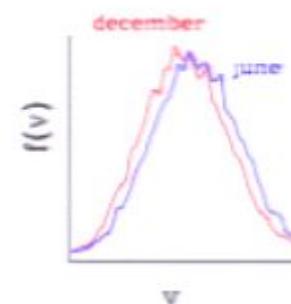
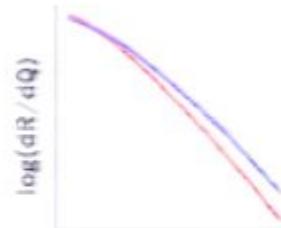
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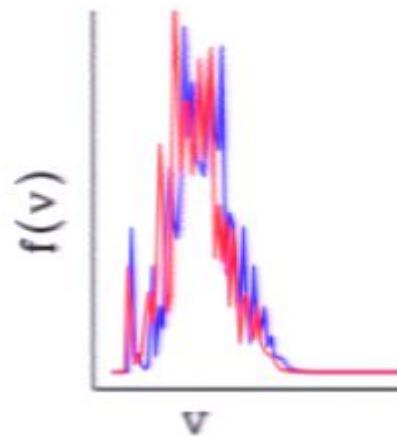
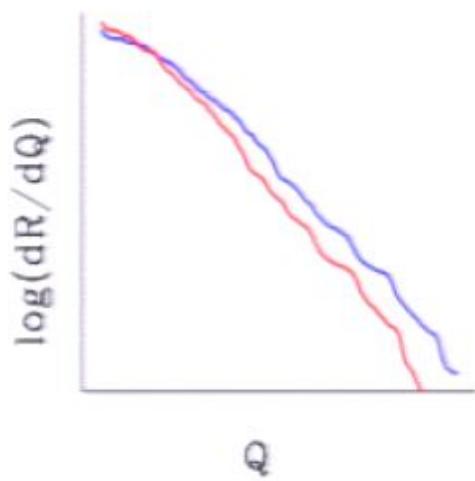
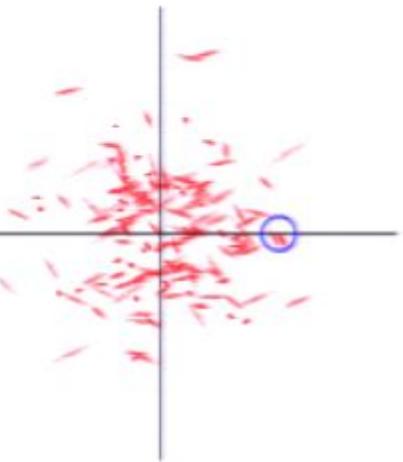
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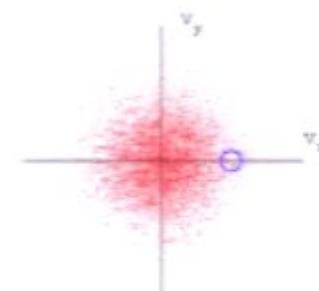




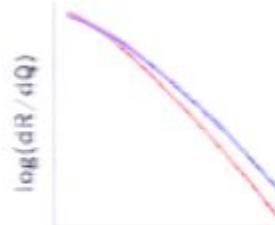
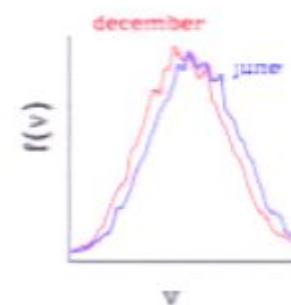
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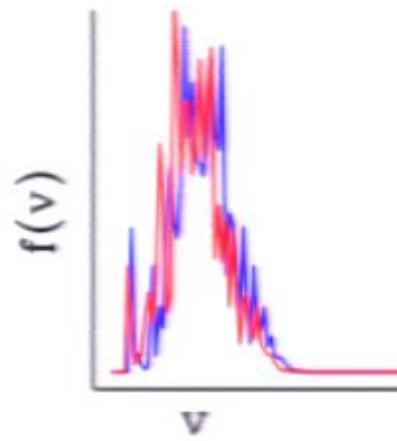
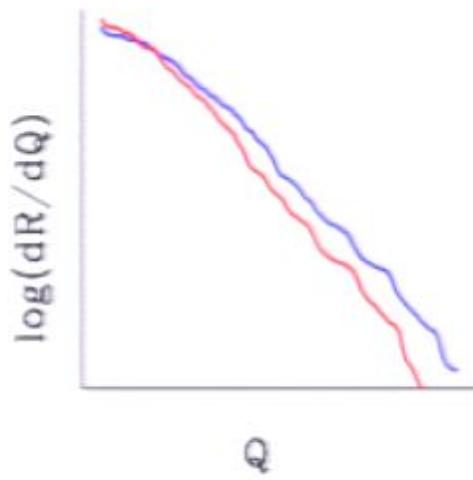
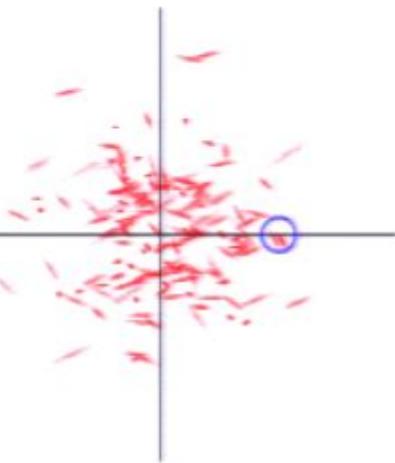
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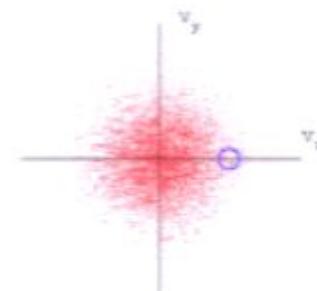




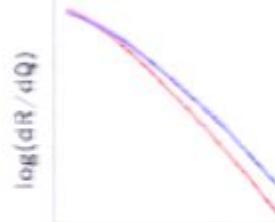
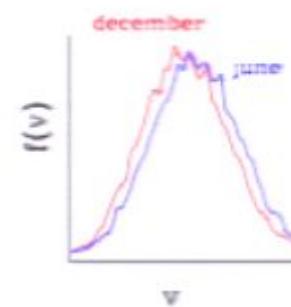
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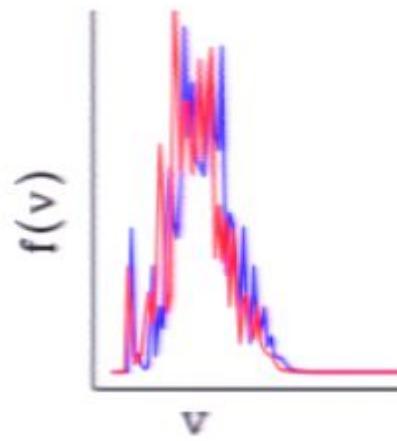
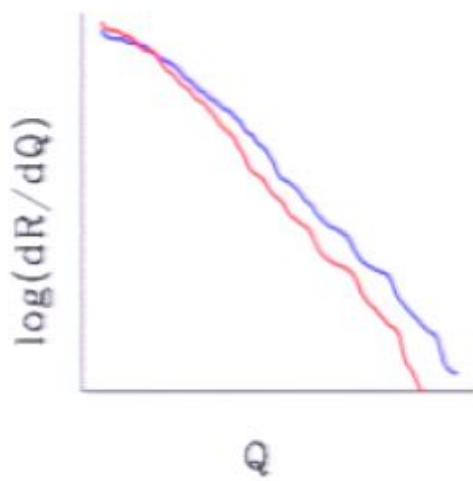
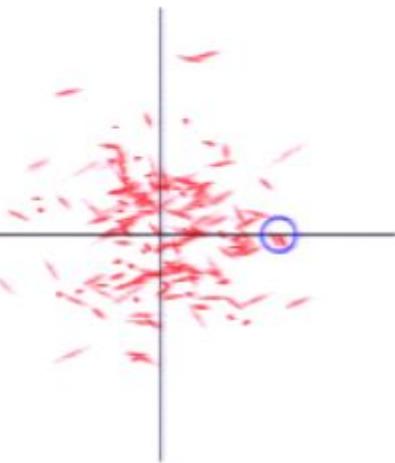
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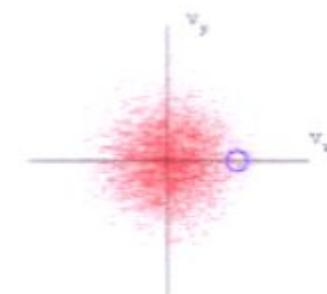




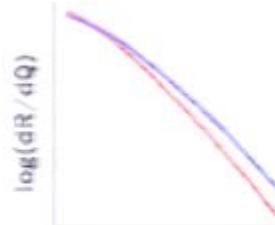
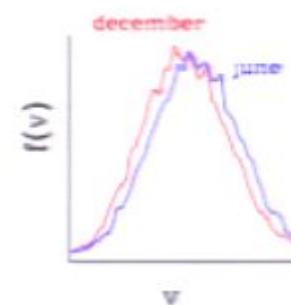
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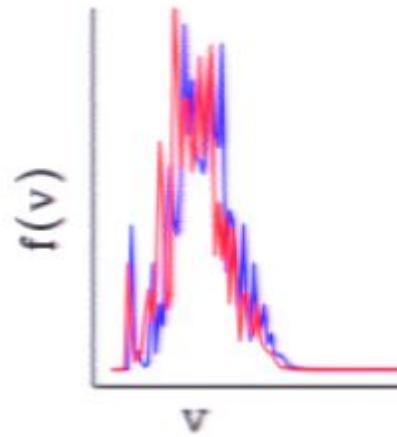
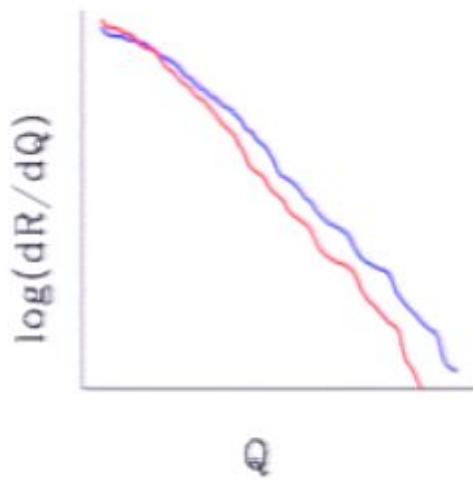
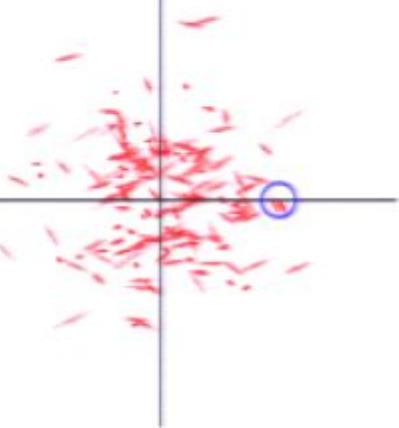
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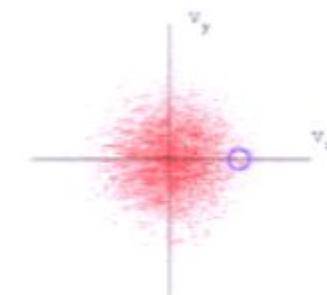




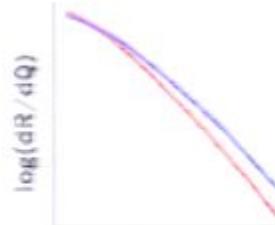
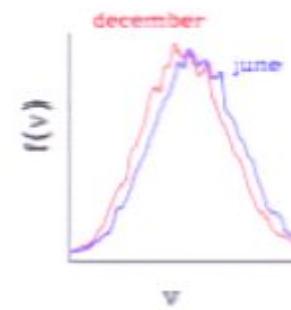
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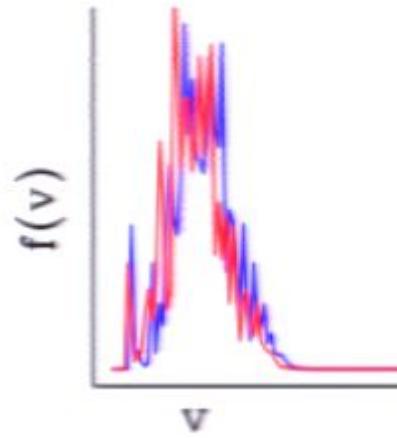
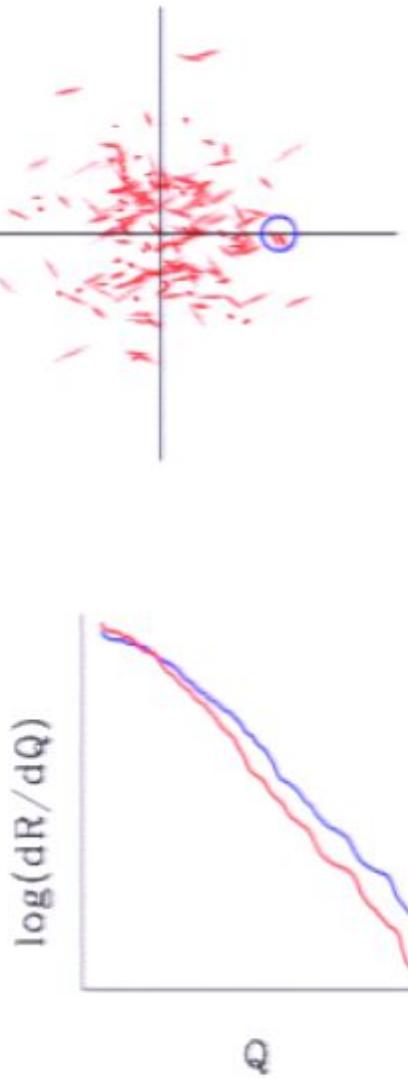
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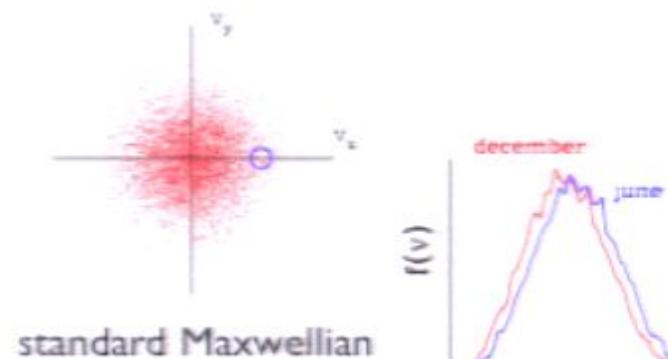




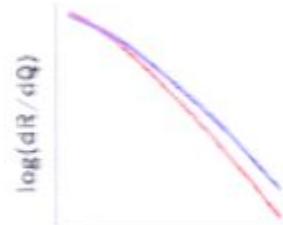
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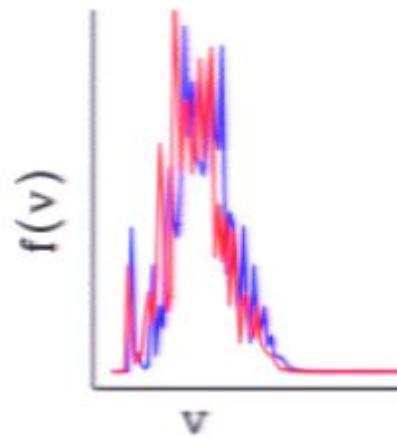
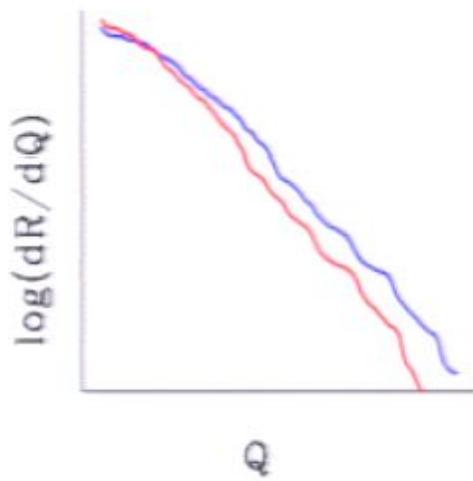
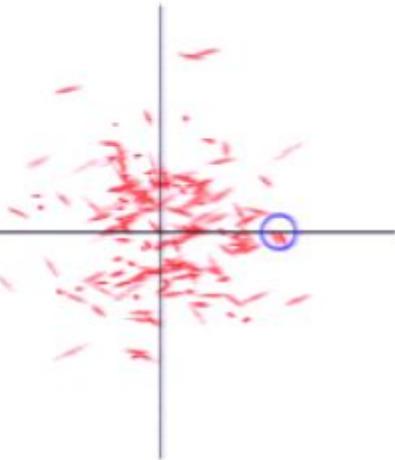
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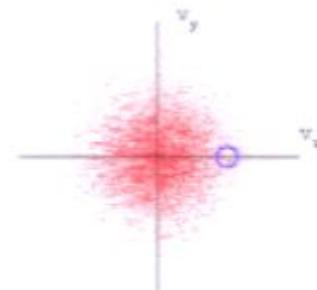




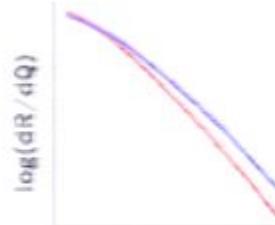
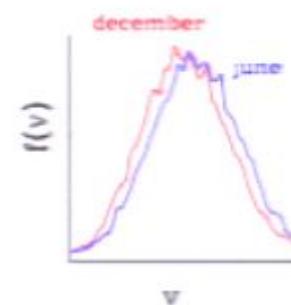
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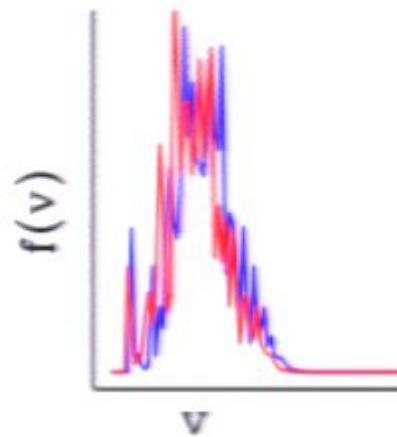
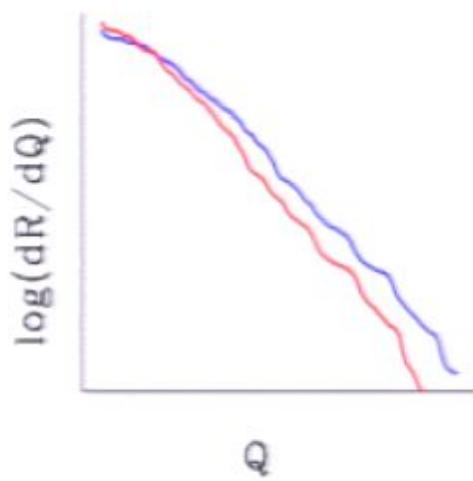
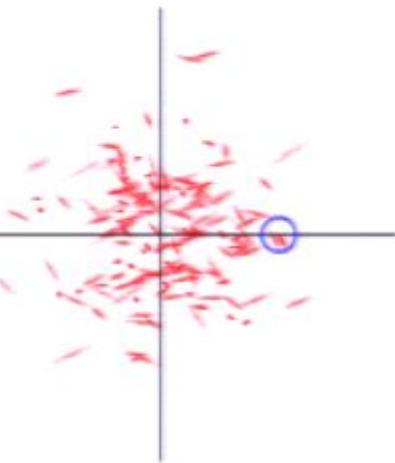
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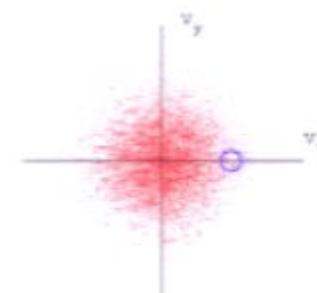




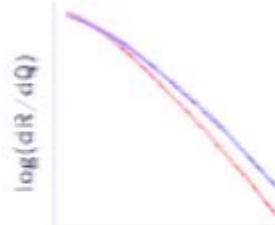
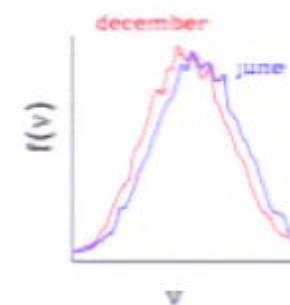
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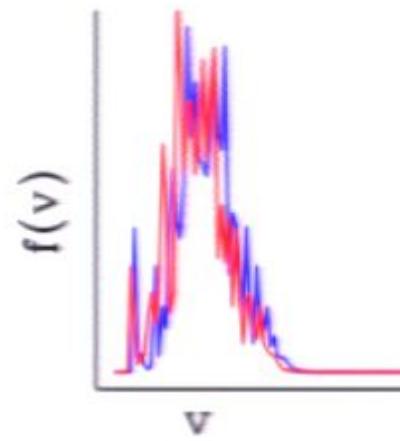
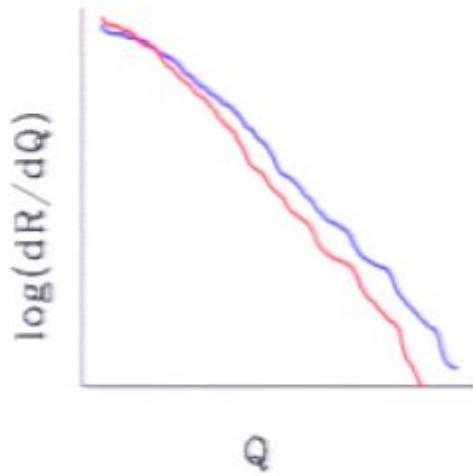
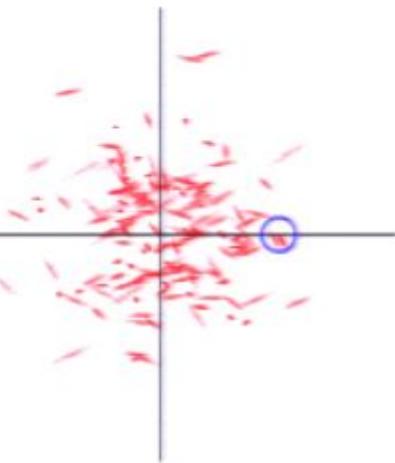
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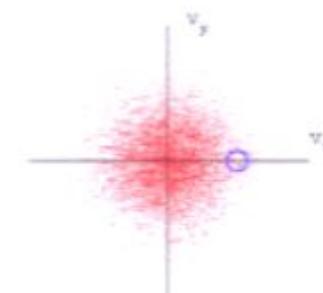




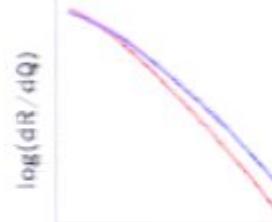
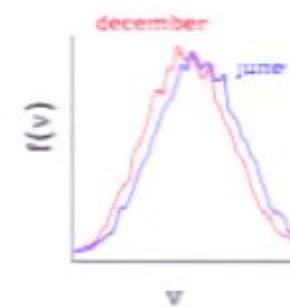
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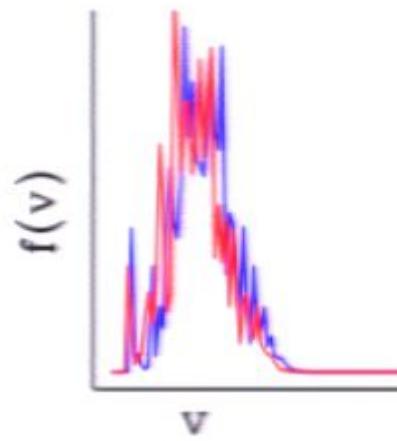
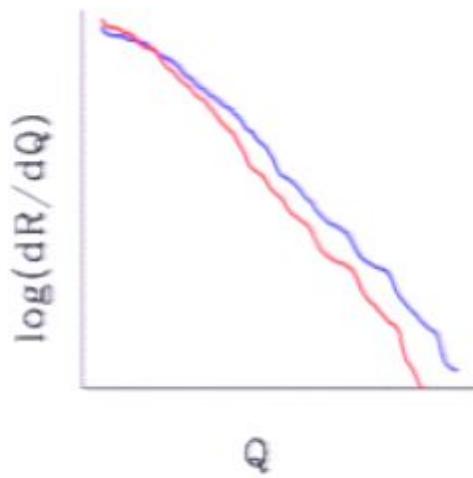
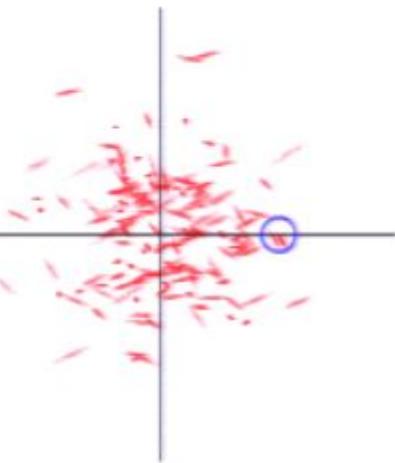
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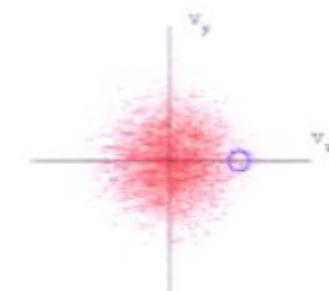




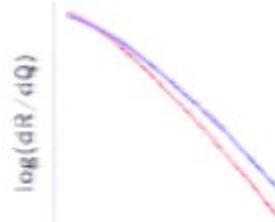
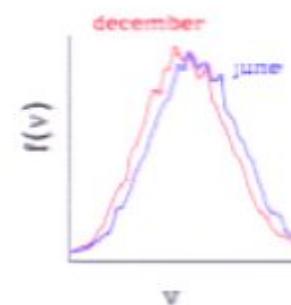
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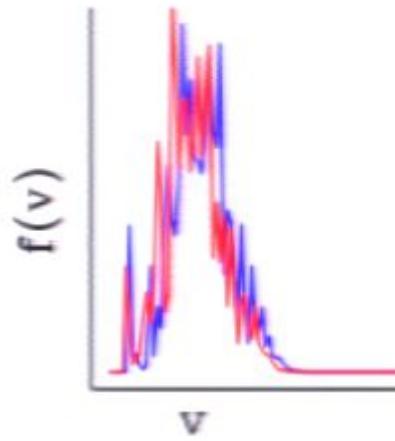
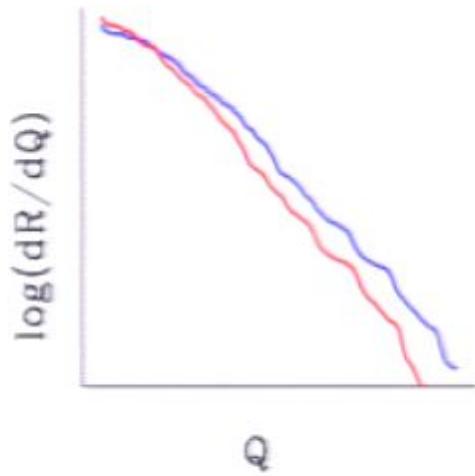
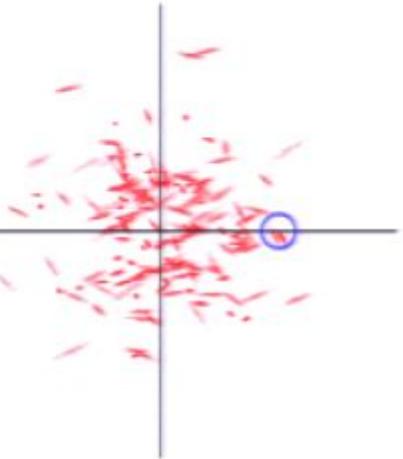
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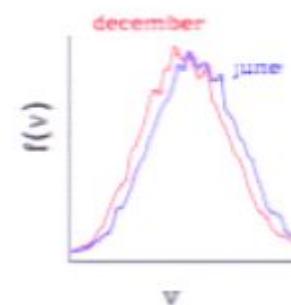
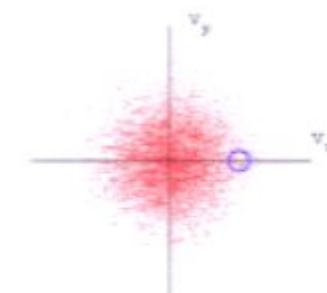


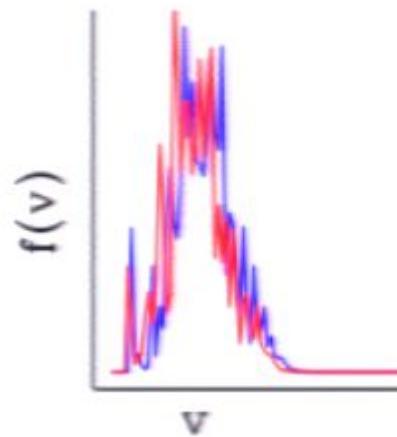
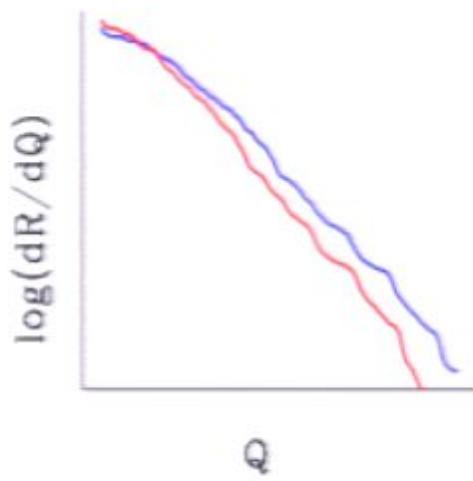
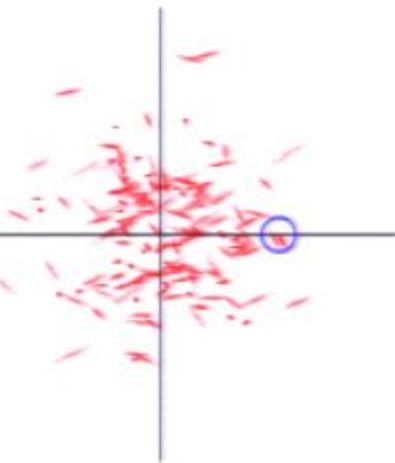


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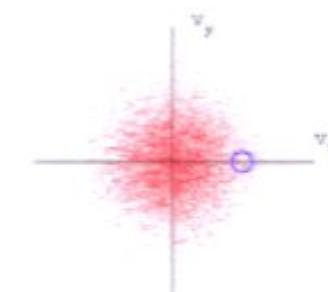




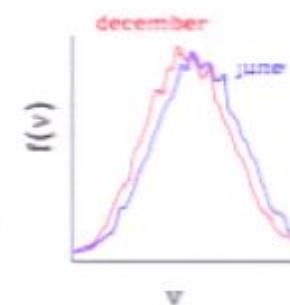
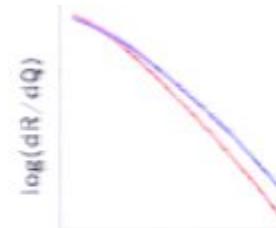
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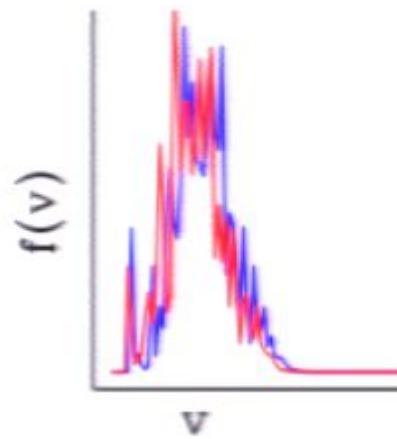
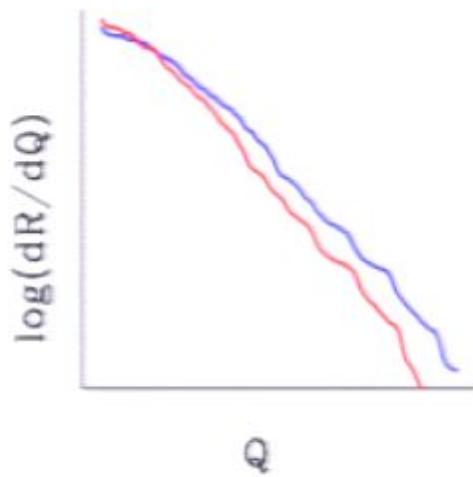
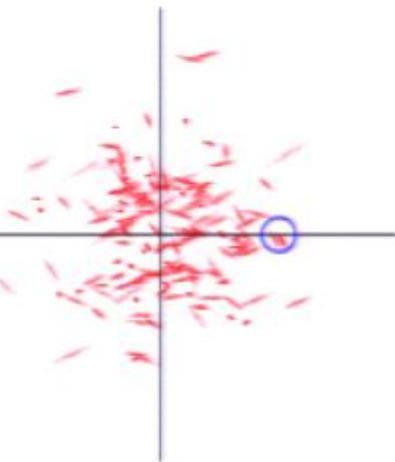
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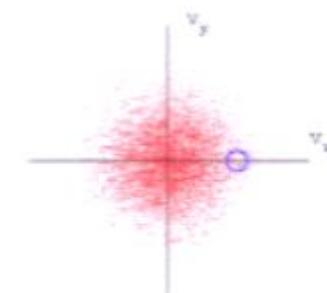




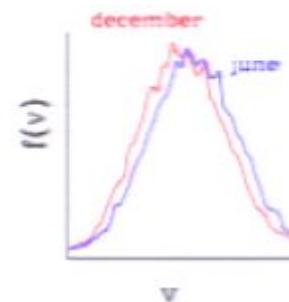
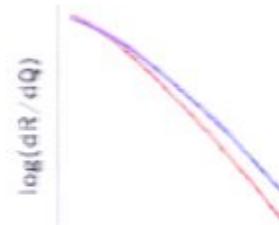
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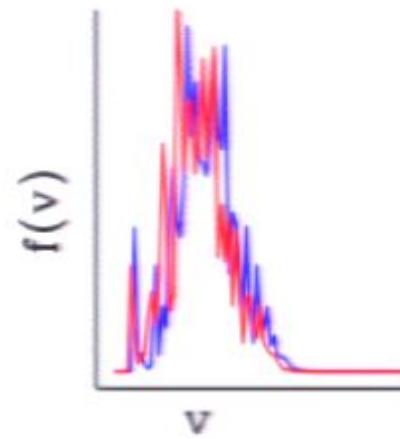
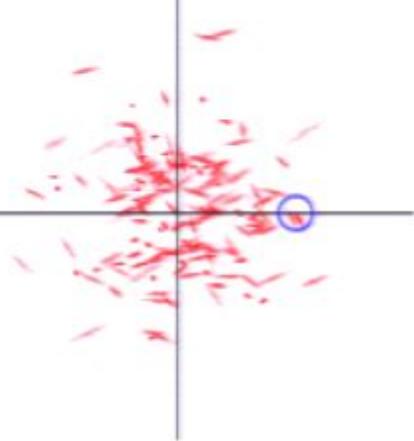
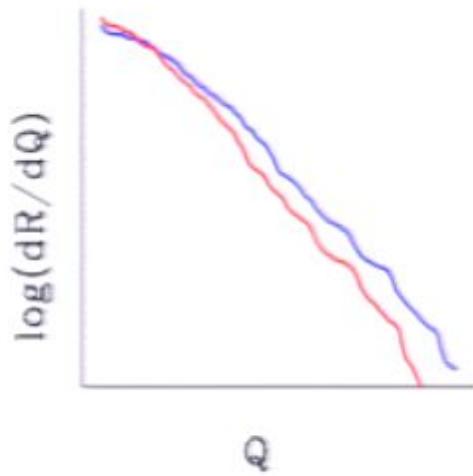
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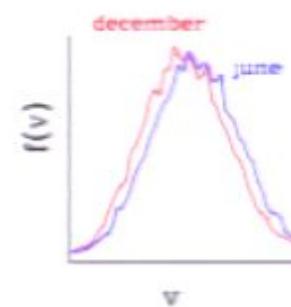
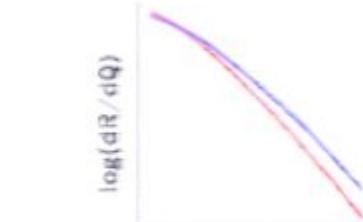
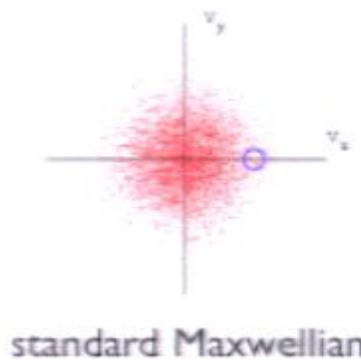


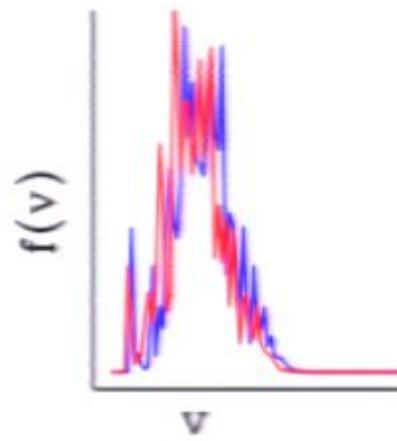
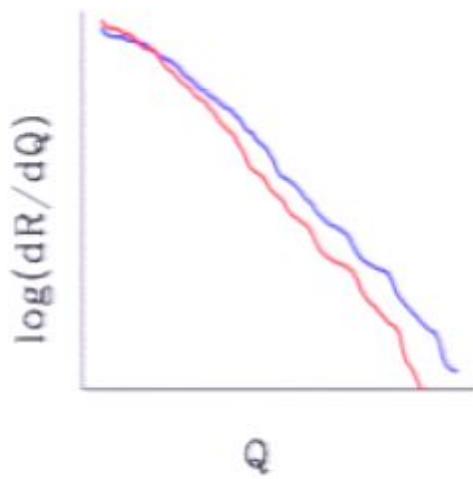
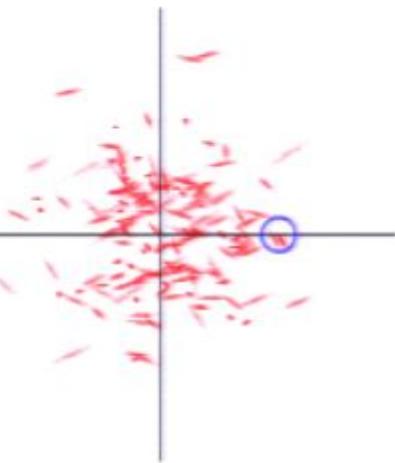


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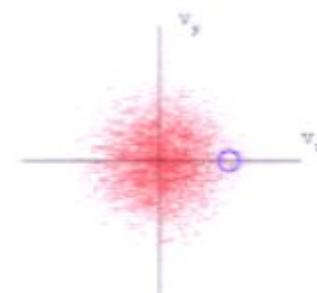




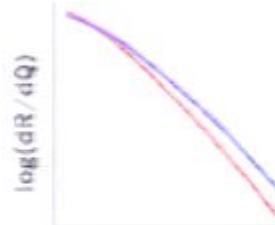
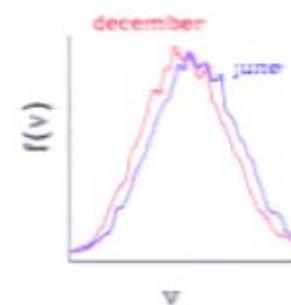
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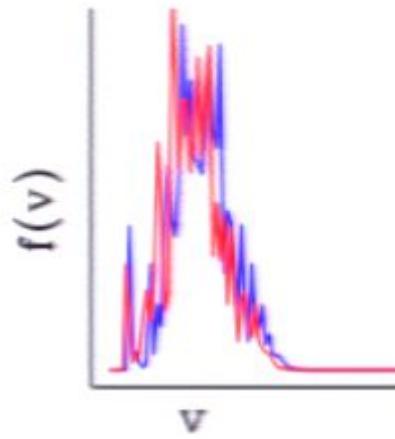
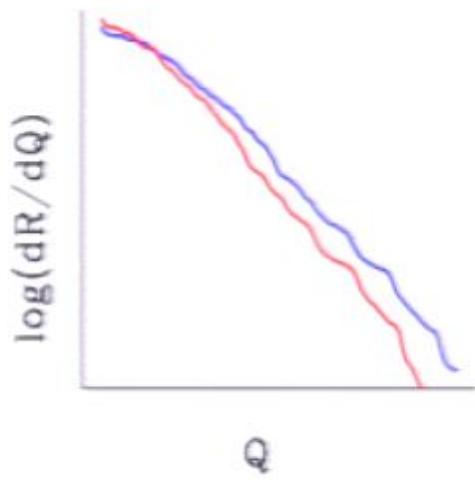
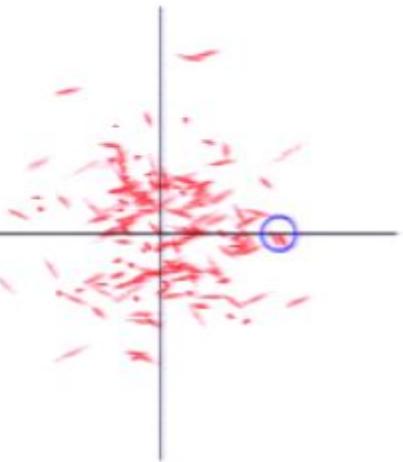
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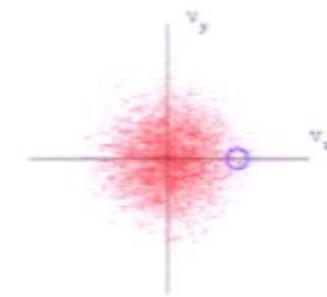




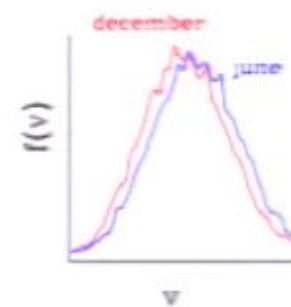
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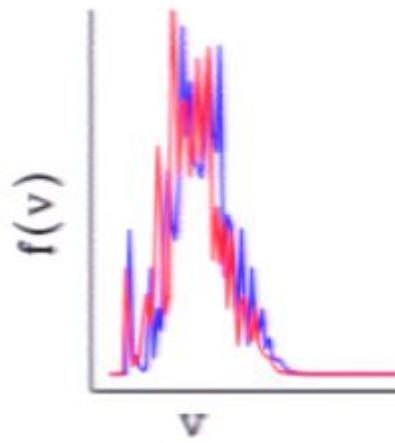
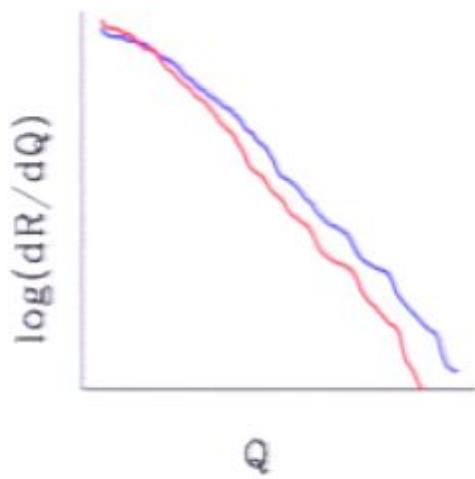
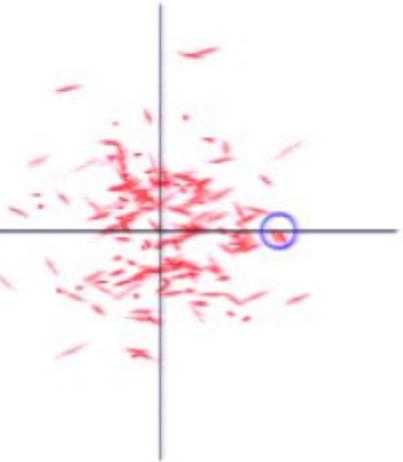
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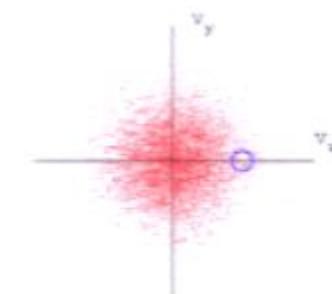




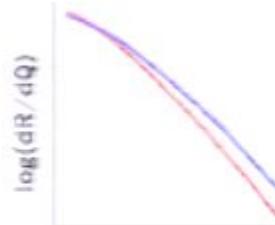
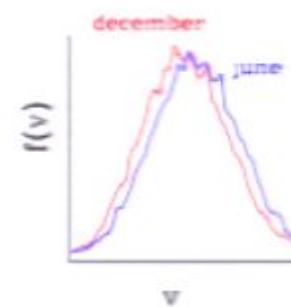
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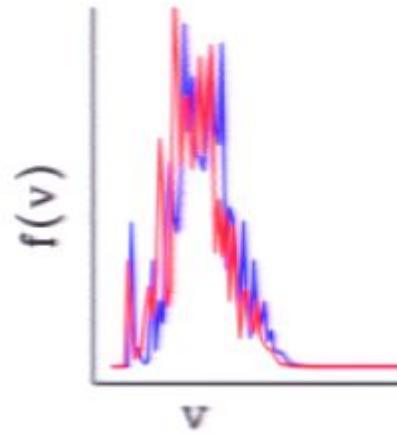
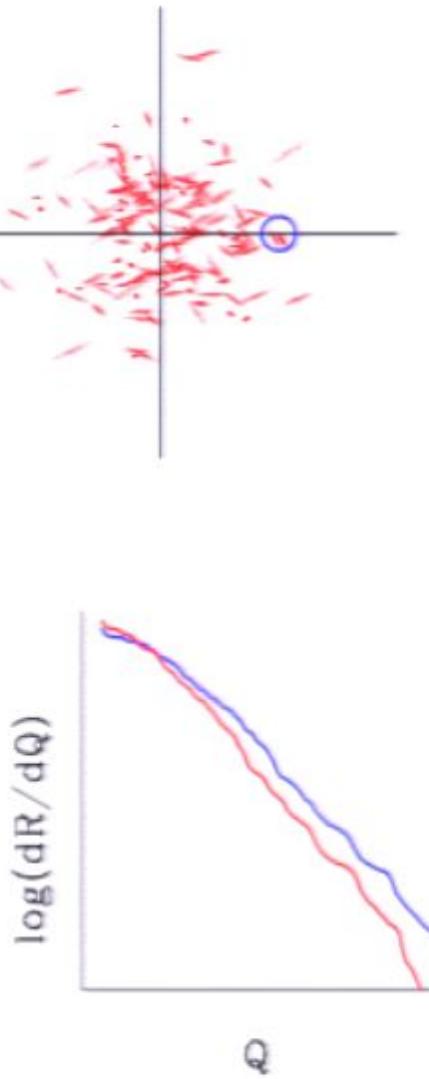
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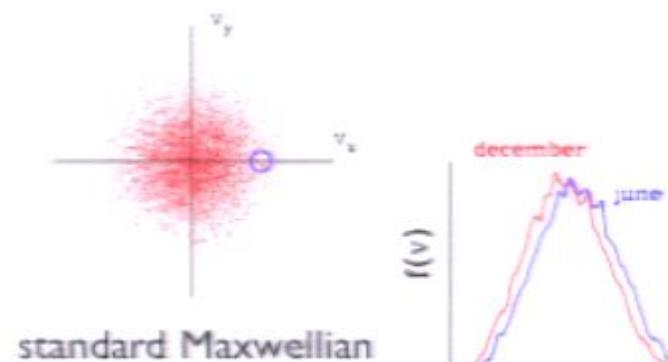


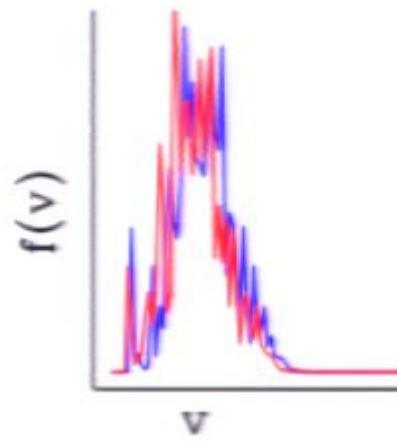
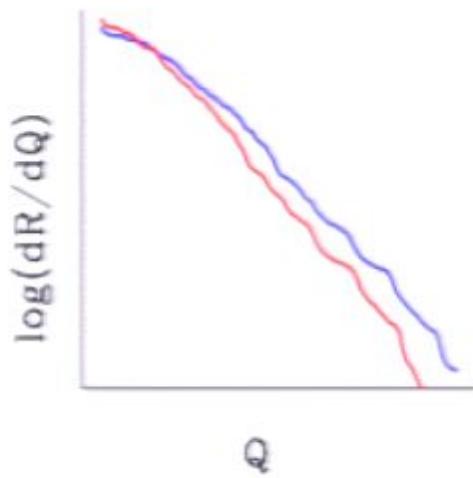
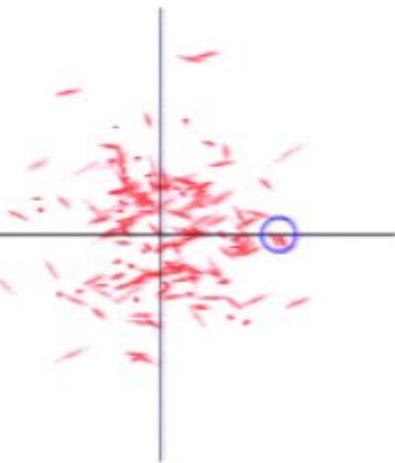


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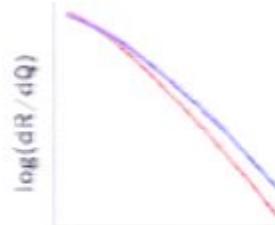
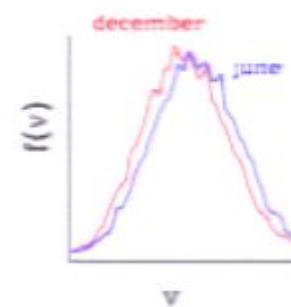
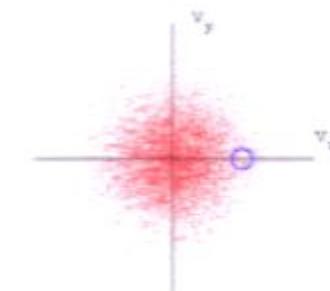


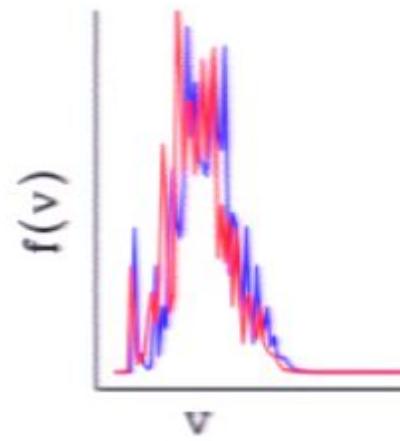
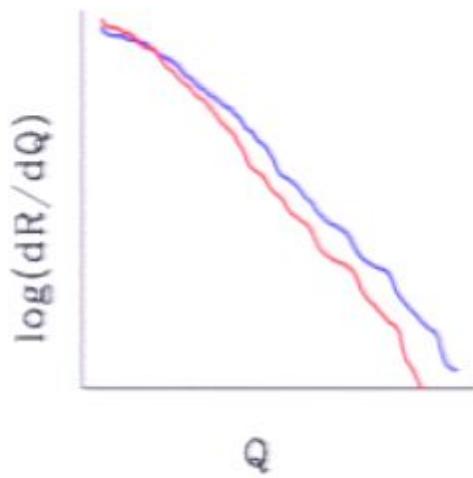
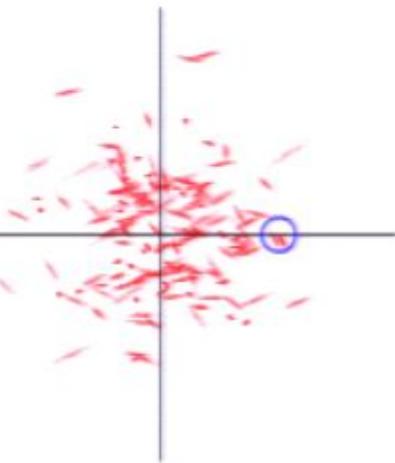


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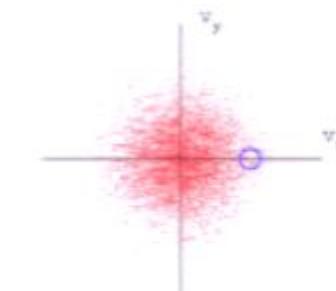




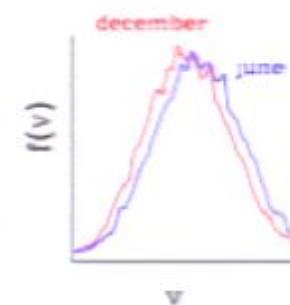
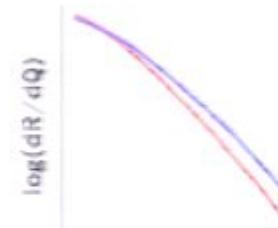
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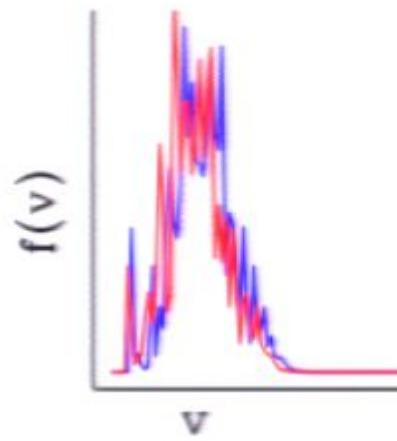
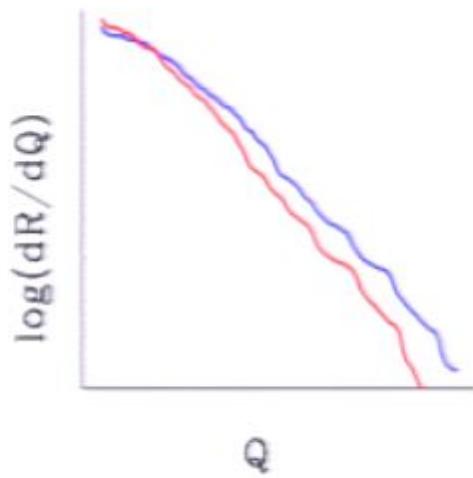
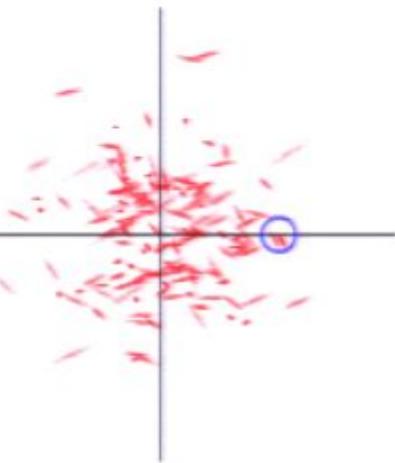
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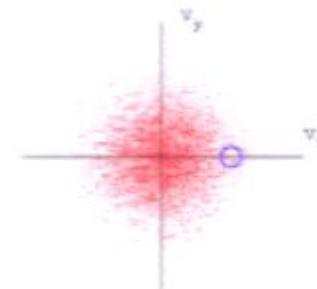




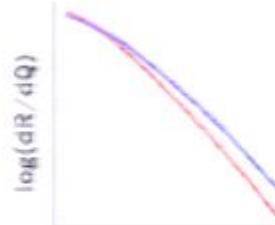
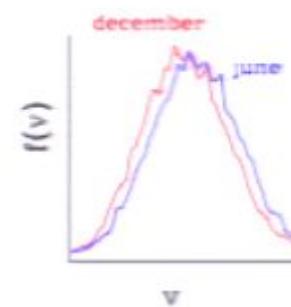
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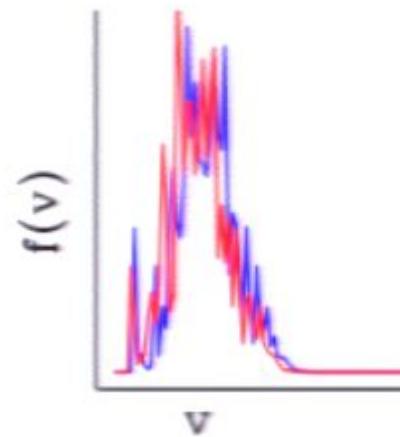
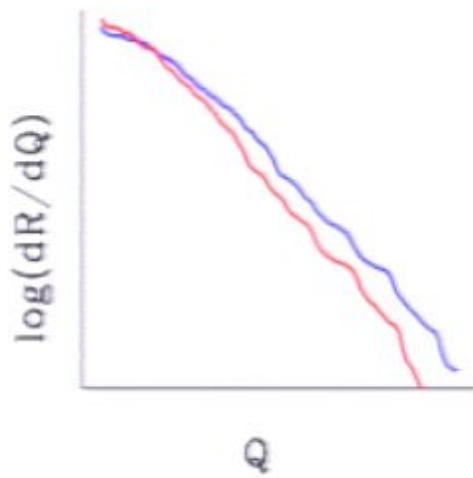
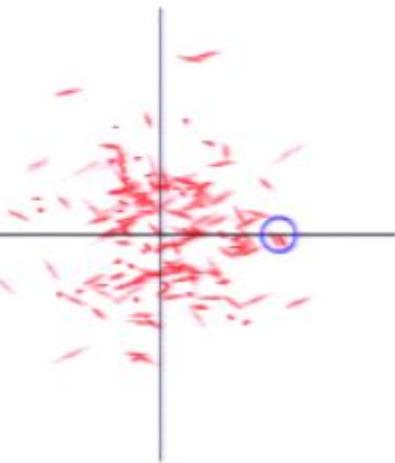
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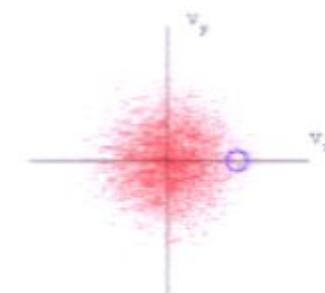




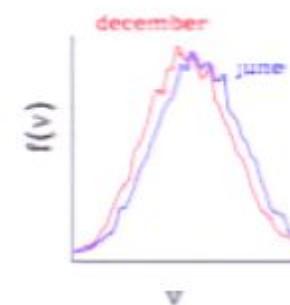
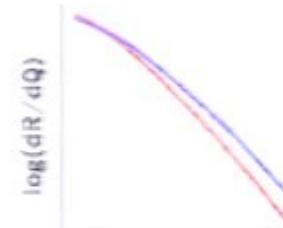
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Identification of halos, subhalos, and tidal streams is a time-honored endeavor in numerical cosmology

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hunting for peaks above the Maxwellian Sea
(Elahi, LMW, Thacker 2011)

Model the distribution of dark matter as a smooth background (the parent halo DF) and contributions from substructure (streams and subhalos)

$$f(\mathbf{x}, \mathbf{v}) = f_h(\mathbf{x}, \mathbf{v}) + \sum_s f_s(\mathbf{x}, \mathbf{v})$$

In a small region of the parent halo, DF is the product of density and an anisotropic Maxwellian-like velocity distribution

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$$f_h(\mathbf{x}, \mathbf{v}) = \rho_h(\mathbf{x}) e^{-v_i \sigma^{ij} v_j / 2\sigma_h^2}$$

STructure Finder hunting for peaks above the Maxwellian Sea (Elahi, LMW, Thacker 2011)

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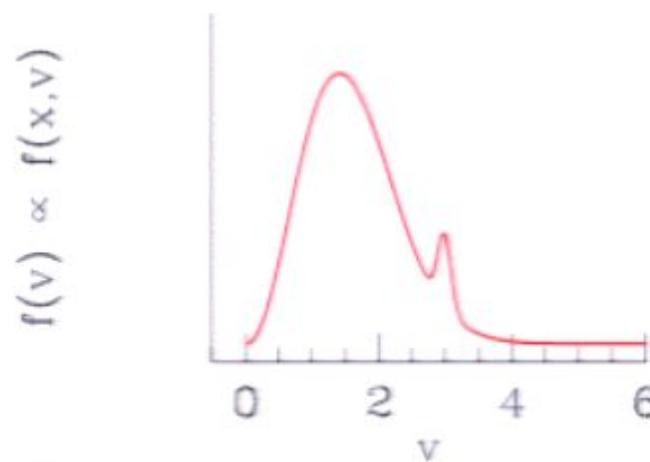
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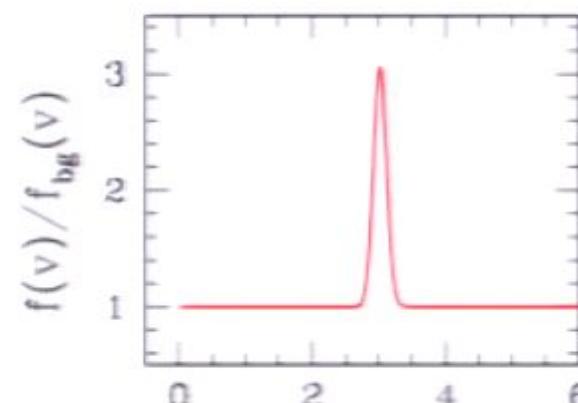
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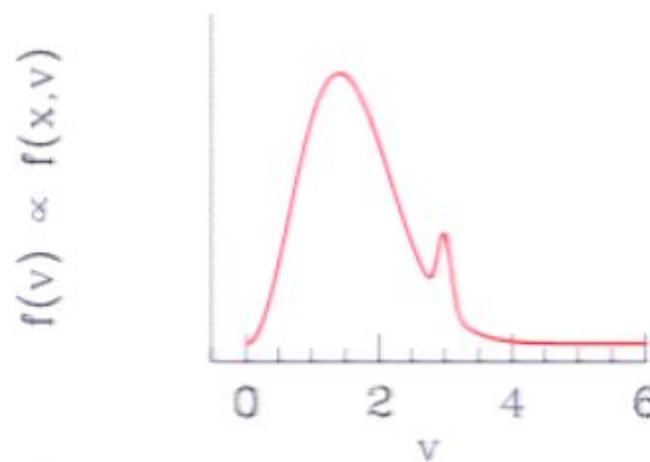


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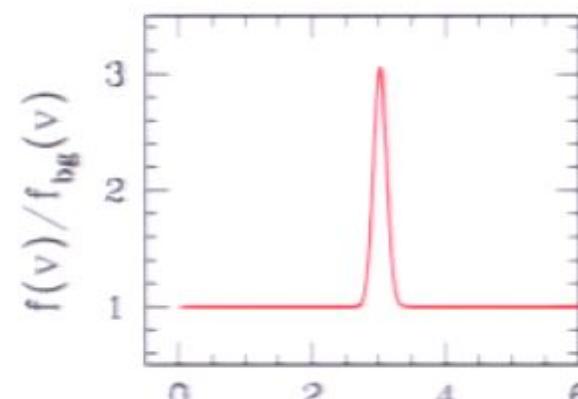


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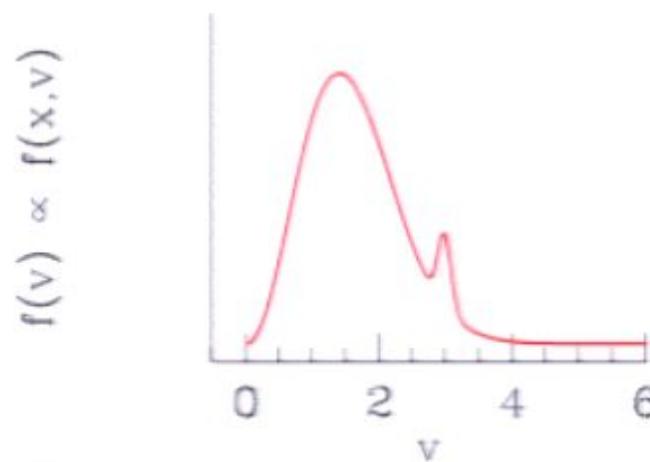


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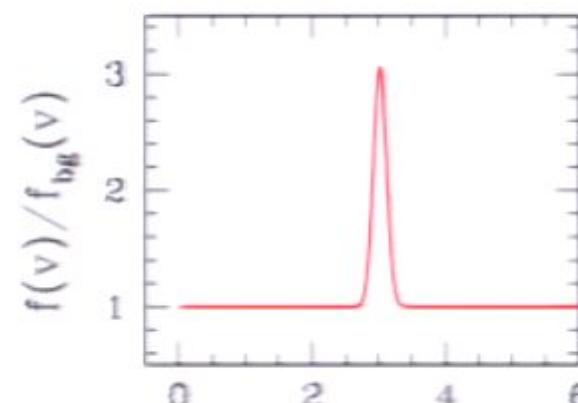


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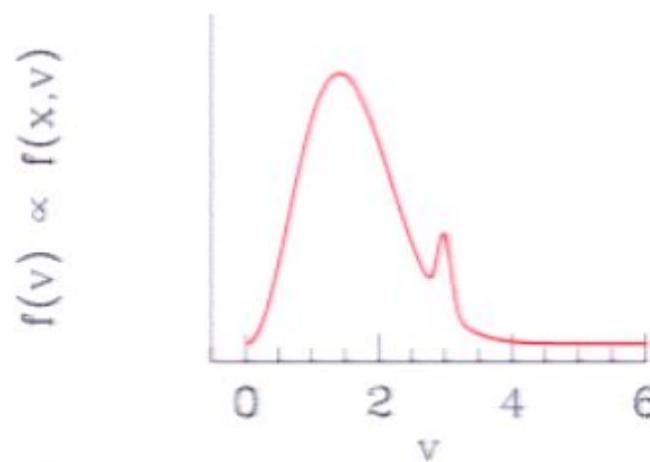


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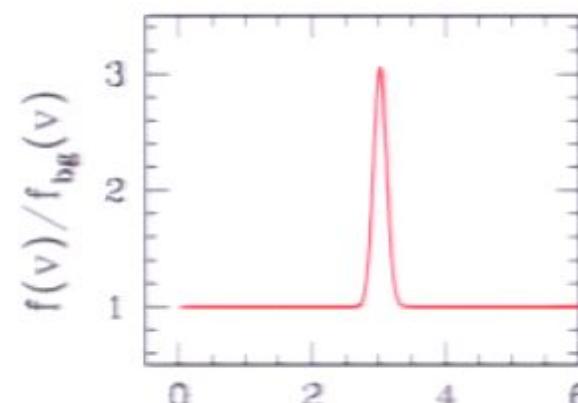


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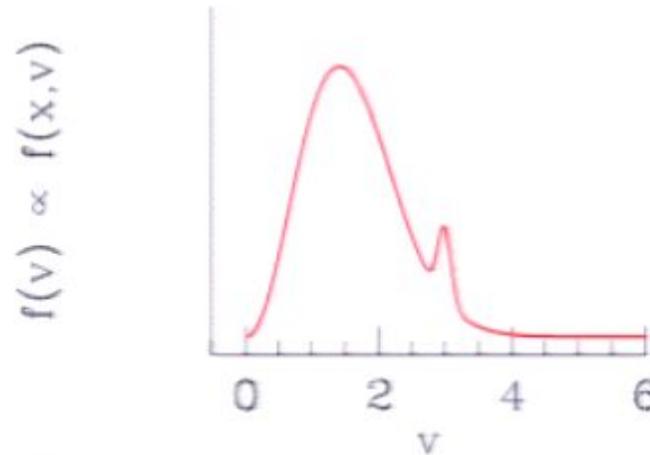


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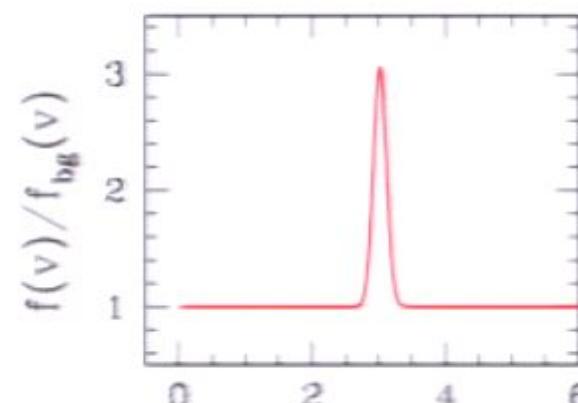


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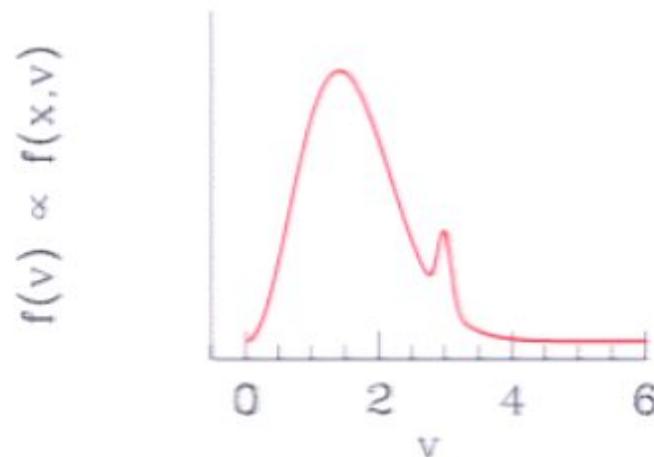


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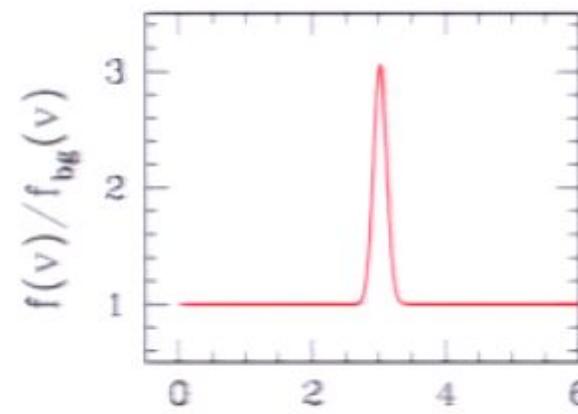


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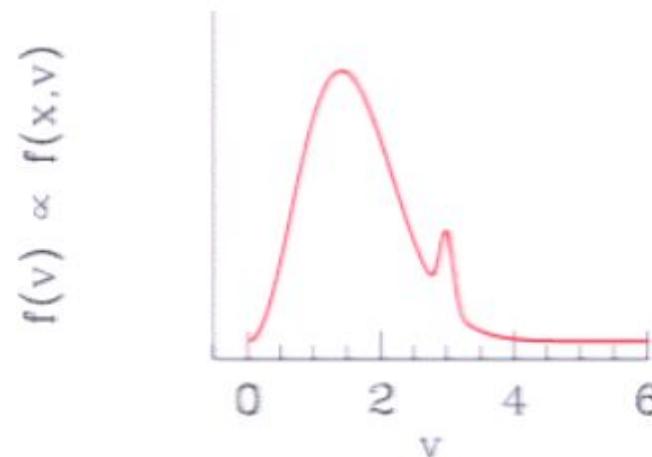


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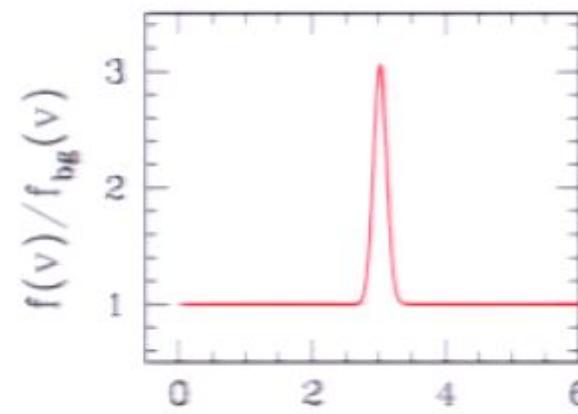


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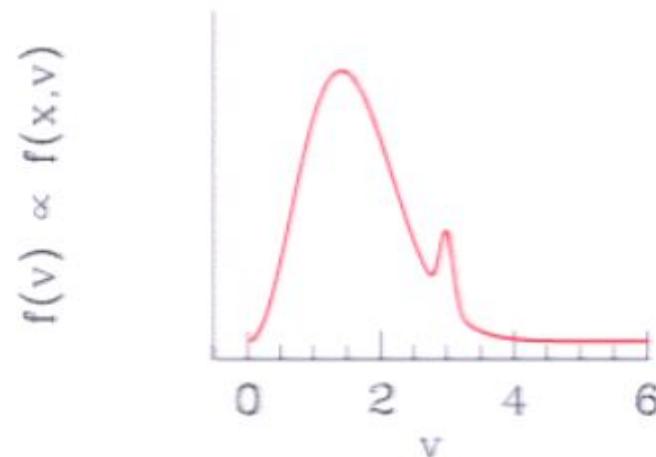


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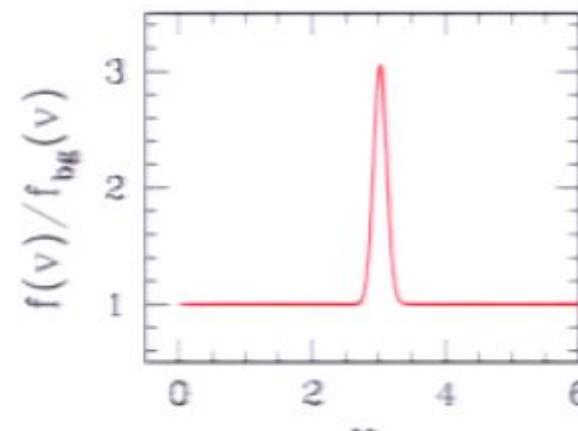


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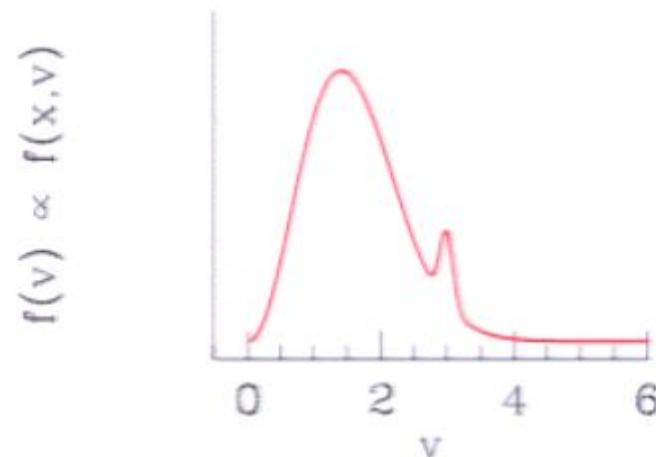


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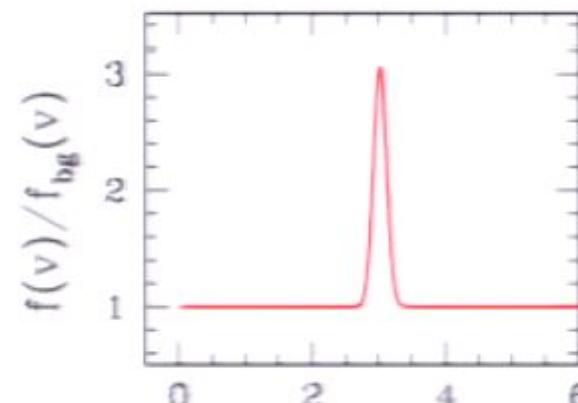


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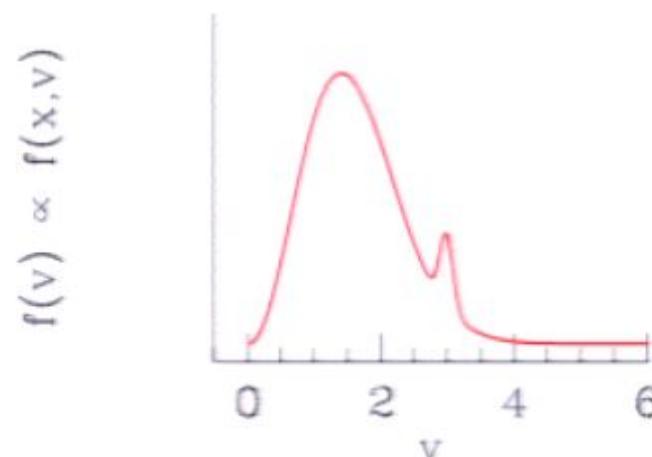


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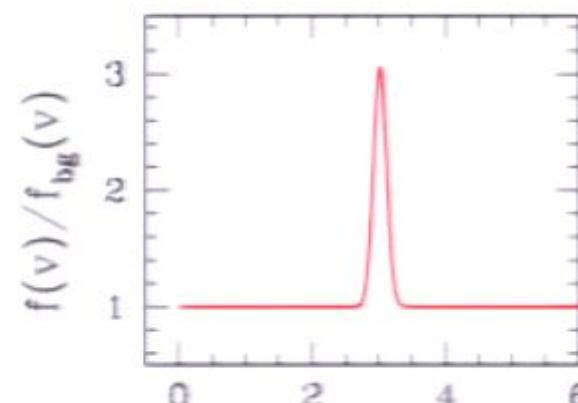


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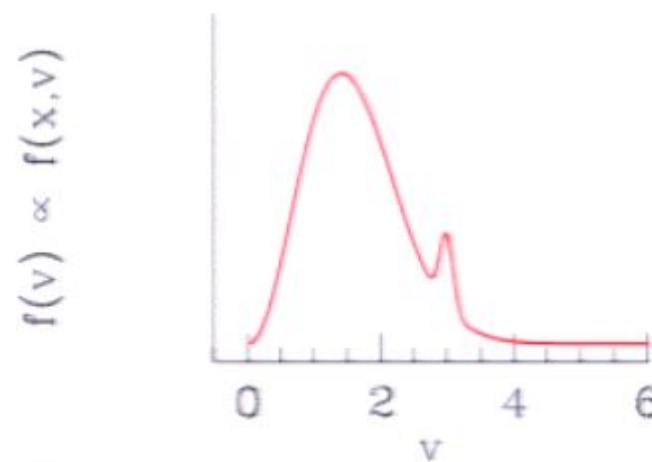


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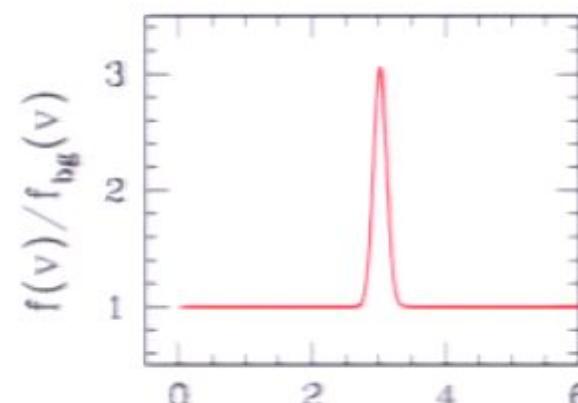


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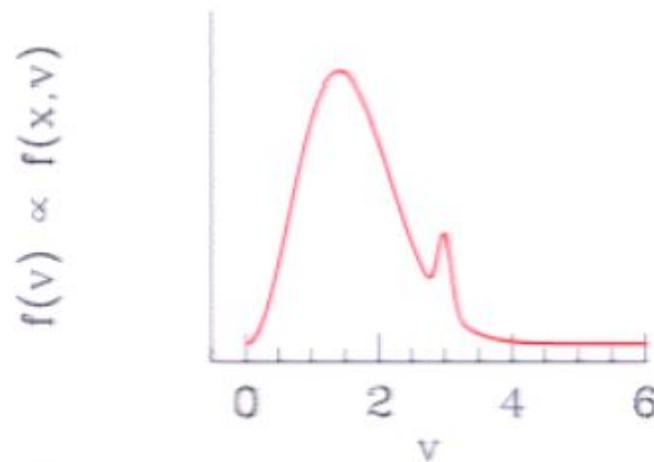


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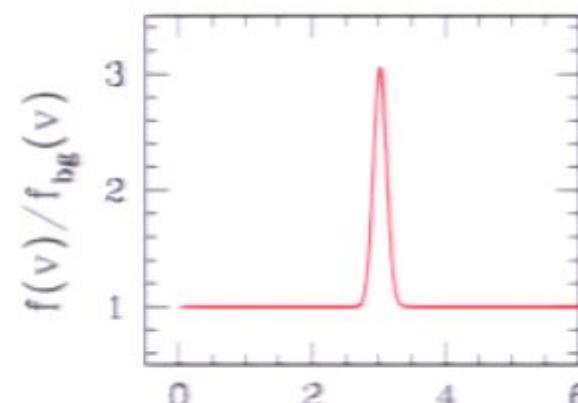


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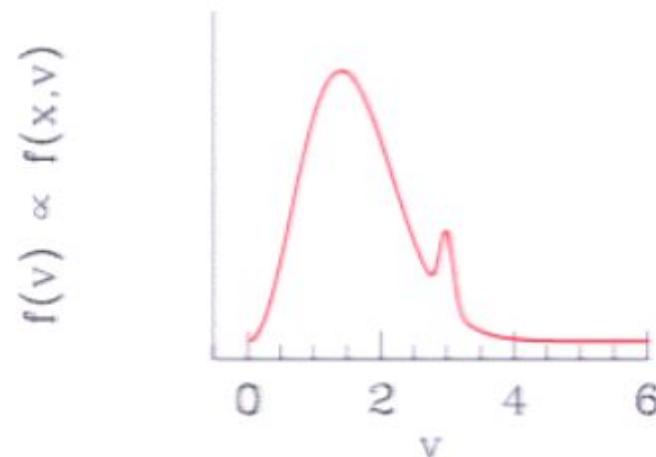


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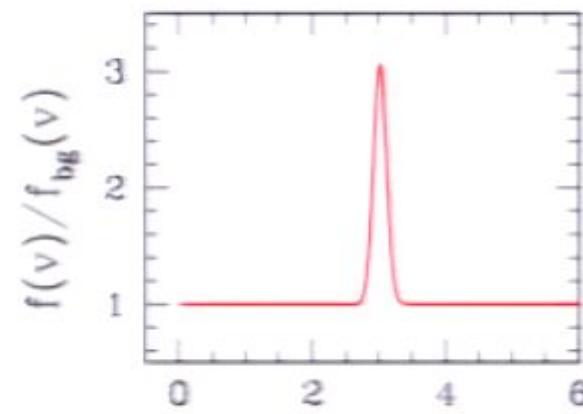


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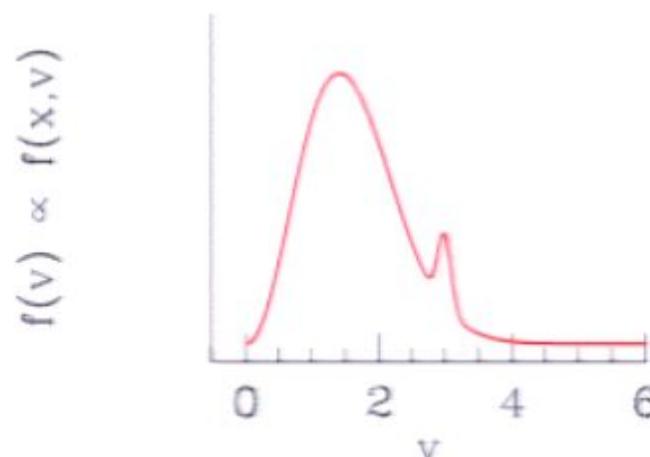


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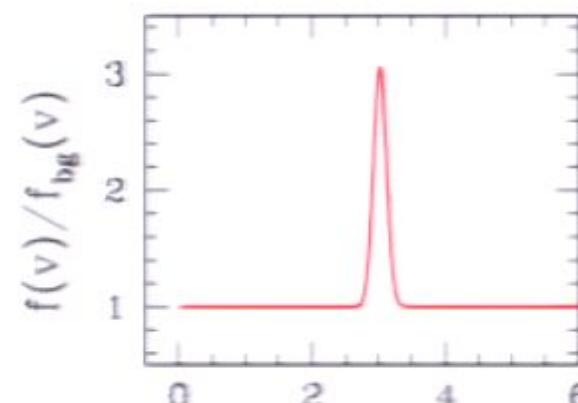


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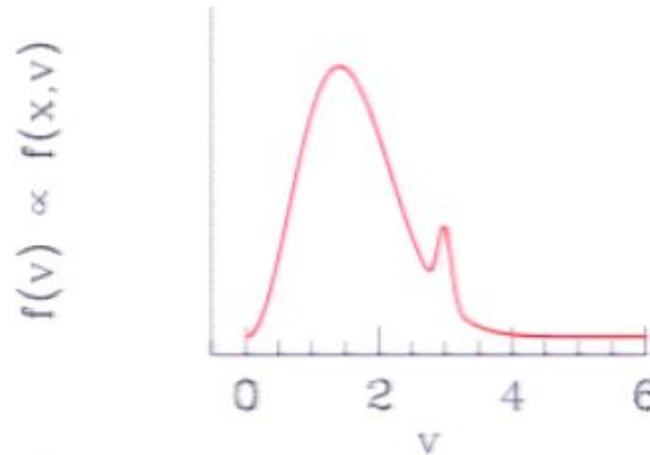


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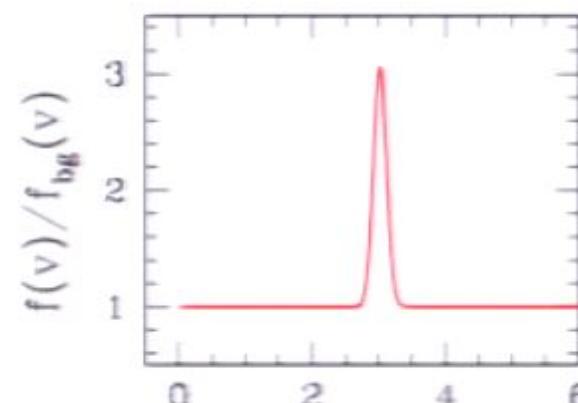


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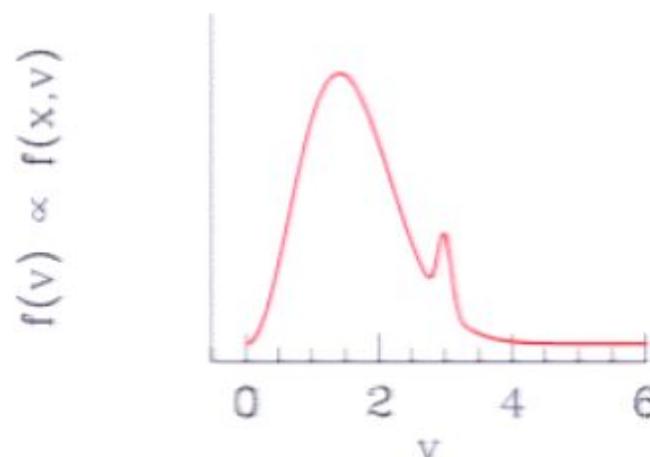


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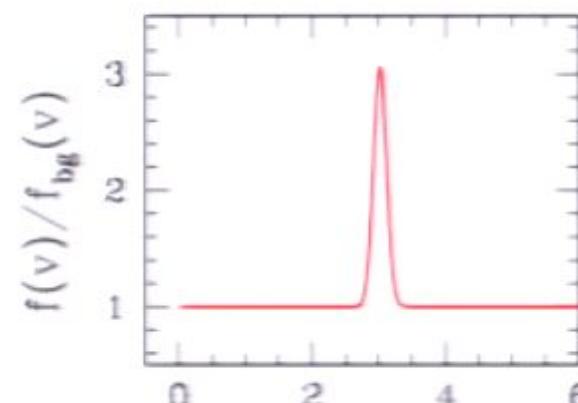


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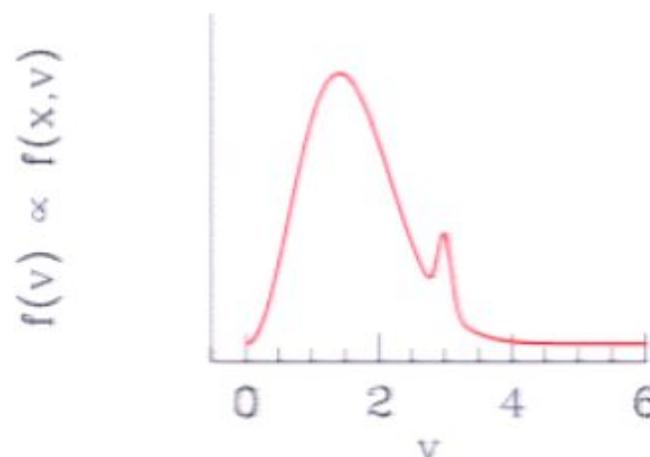


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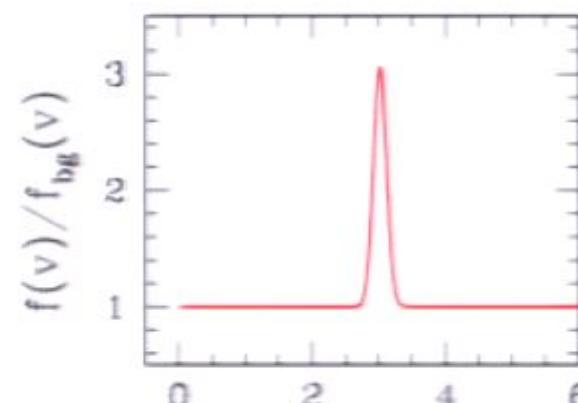


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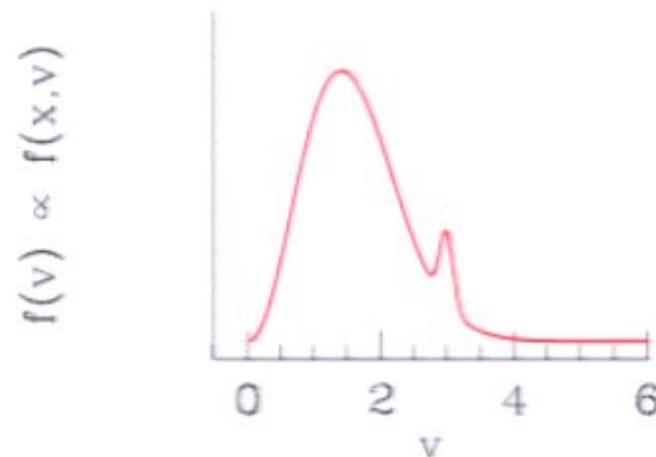


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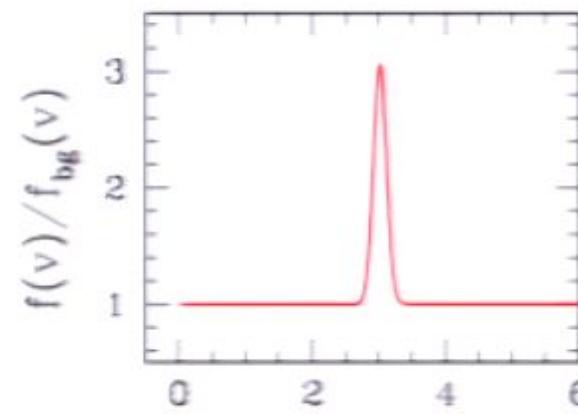


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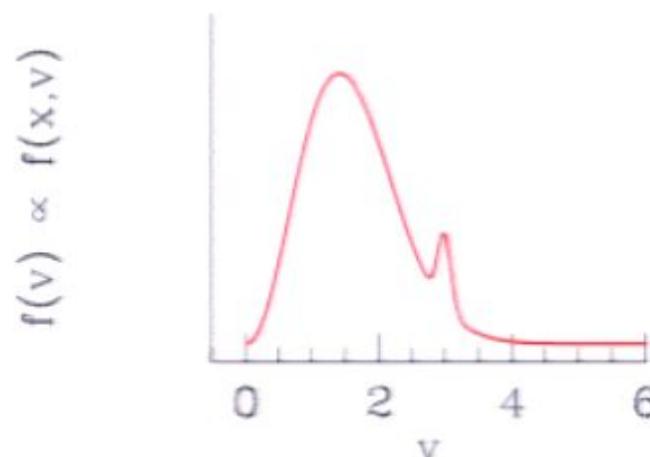


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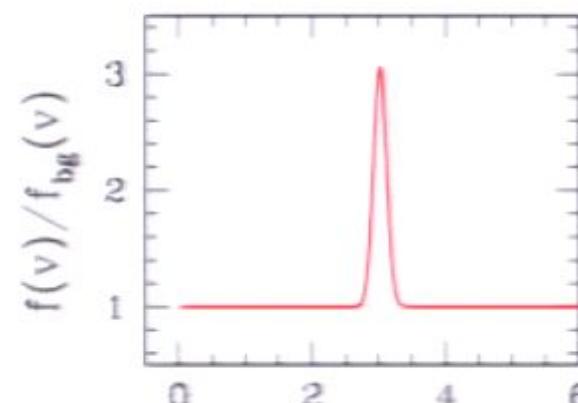


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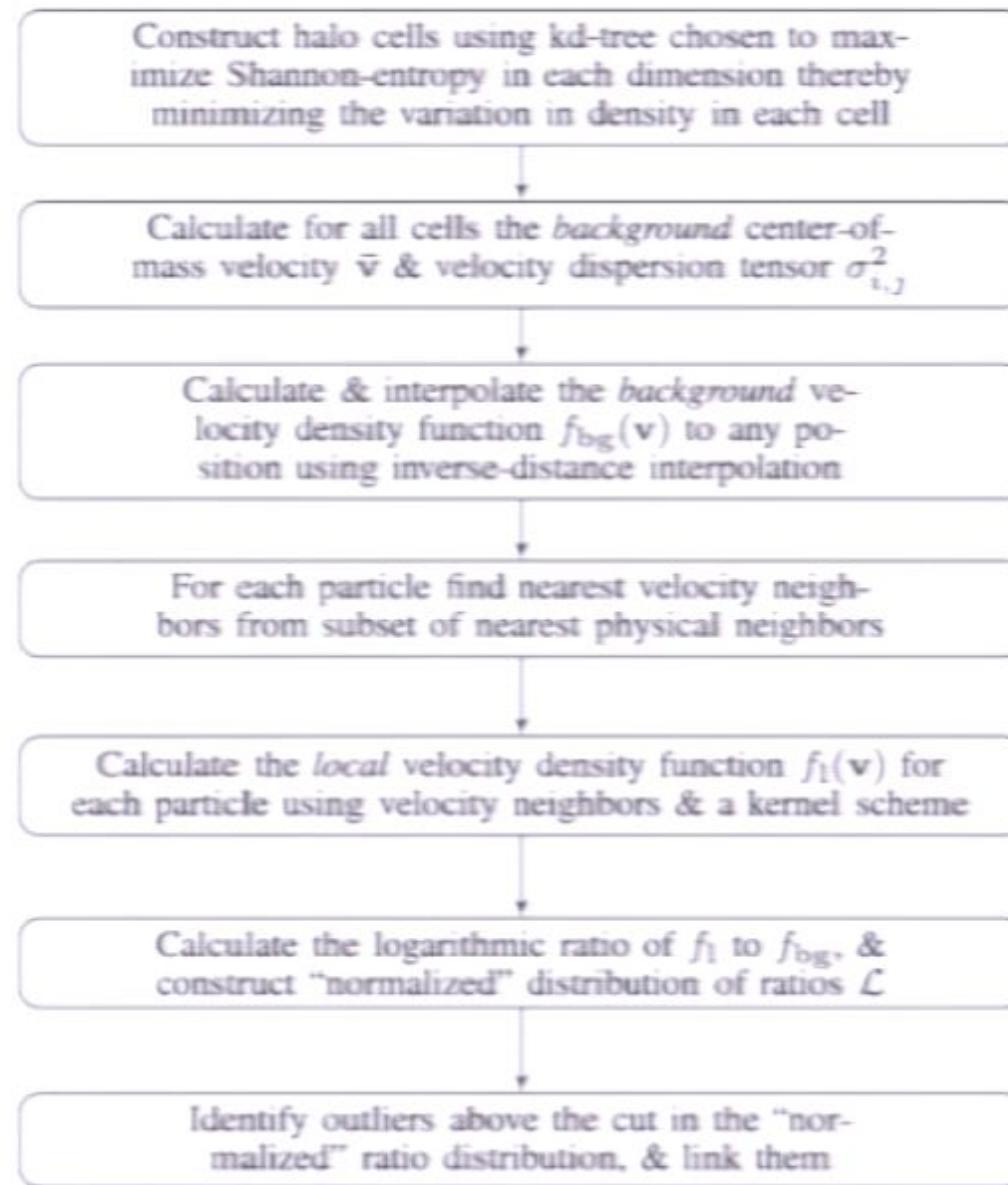


Figure 1. Flow chart of STF algorithm.

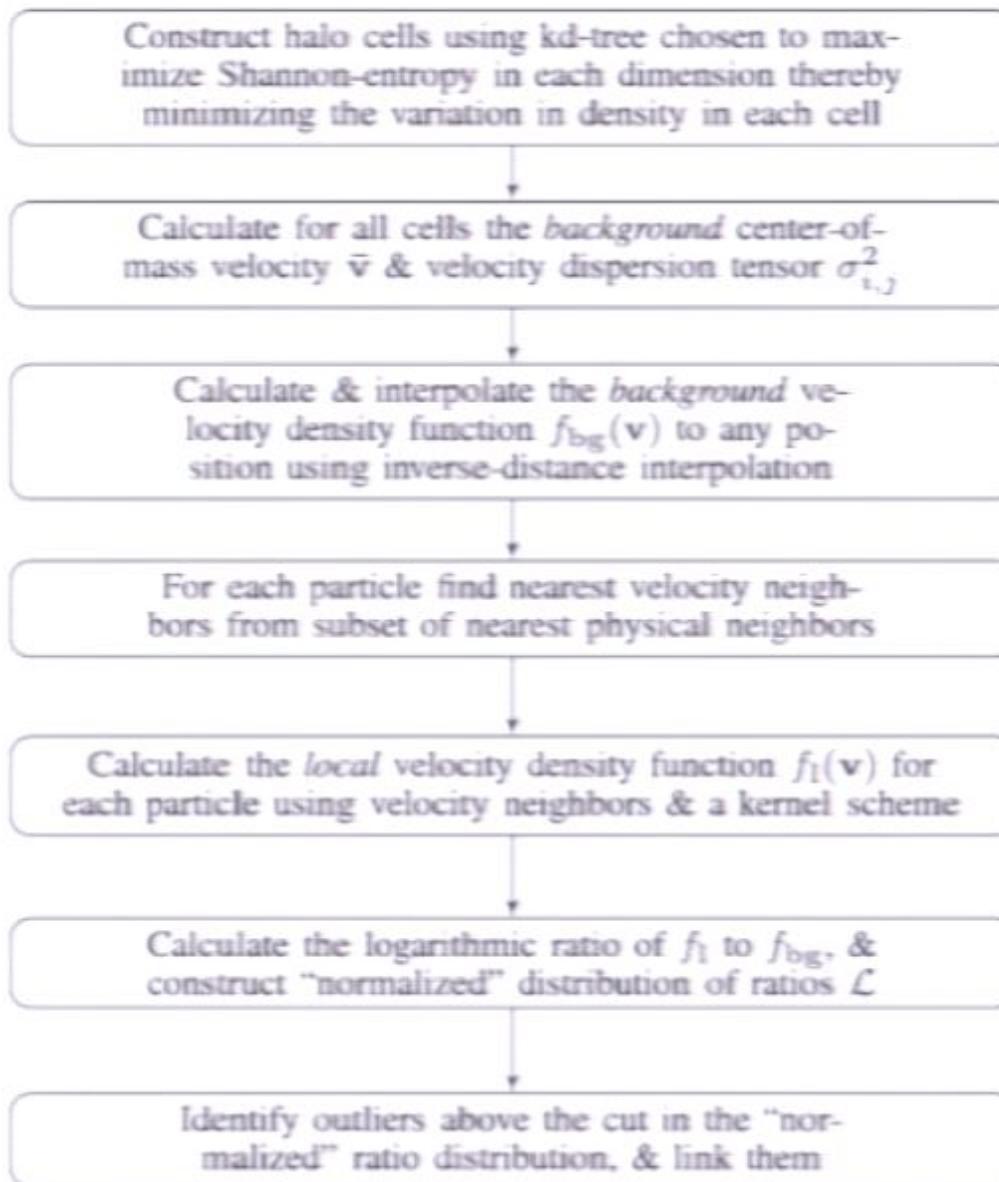


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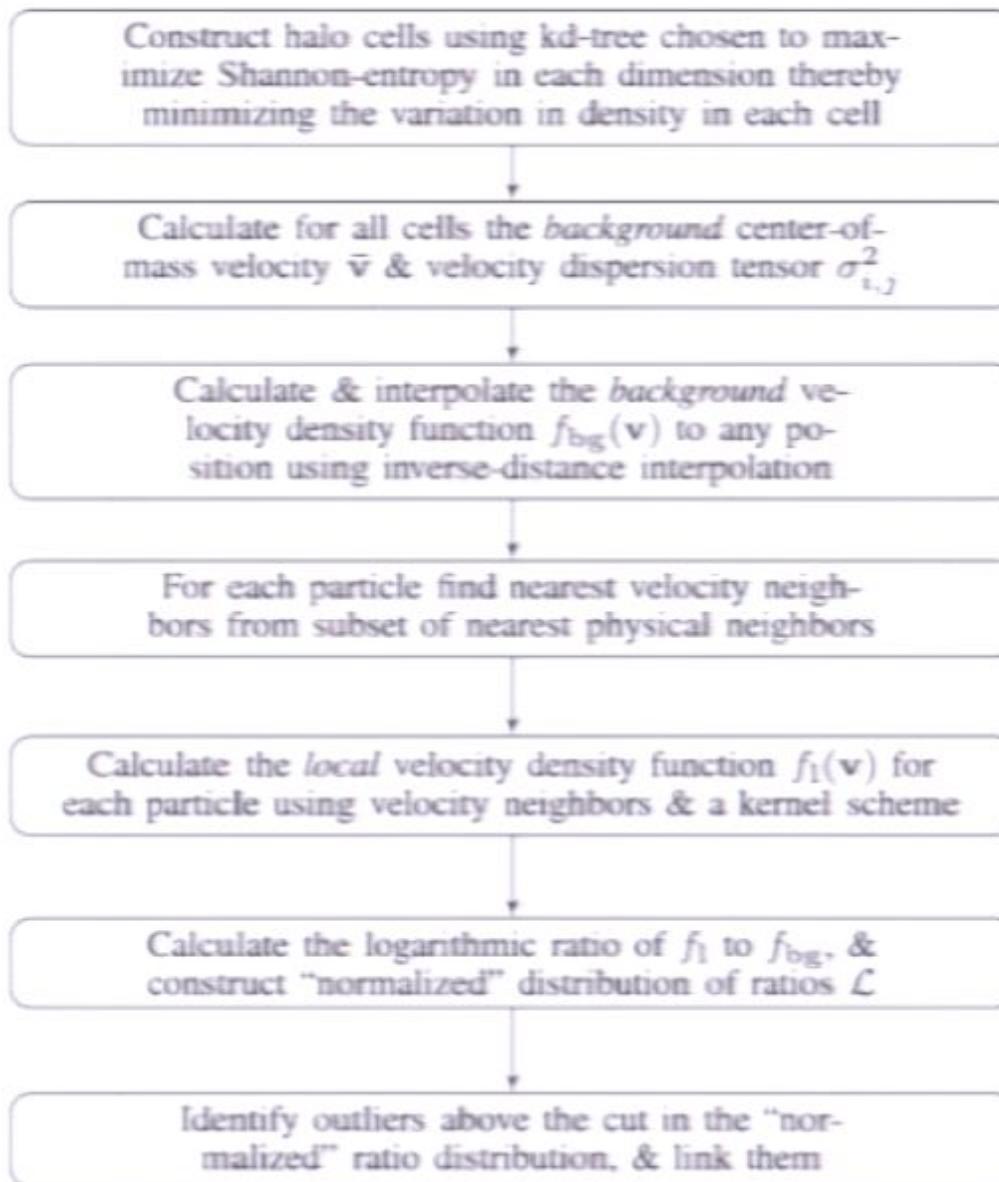


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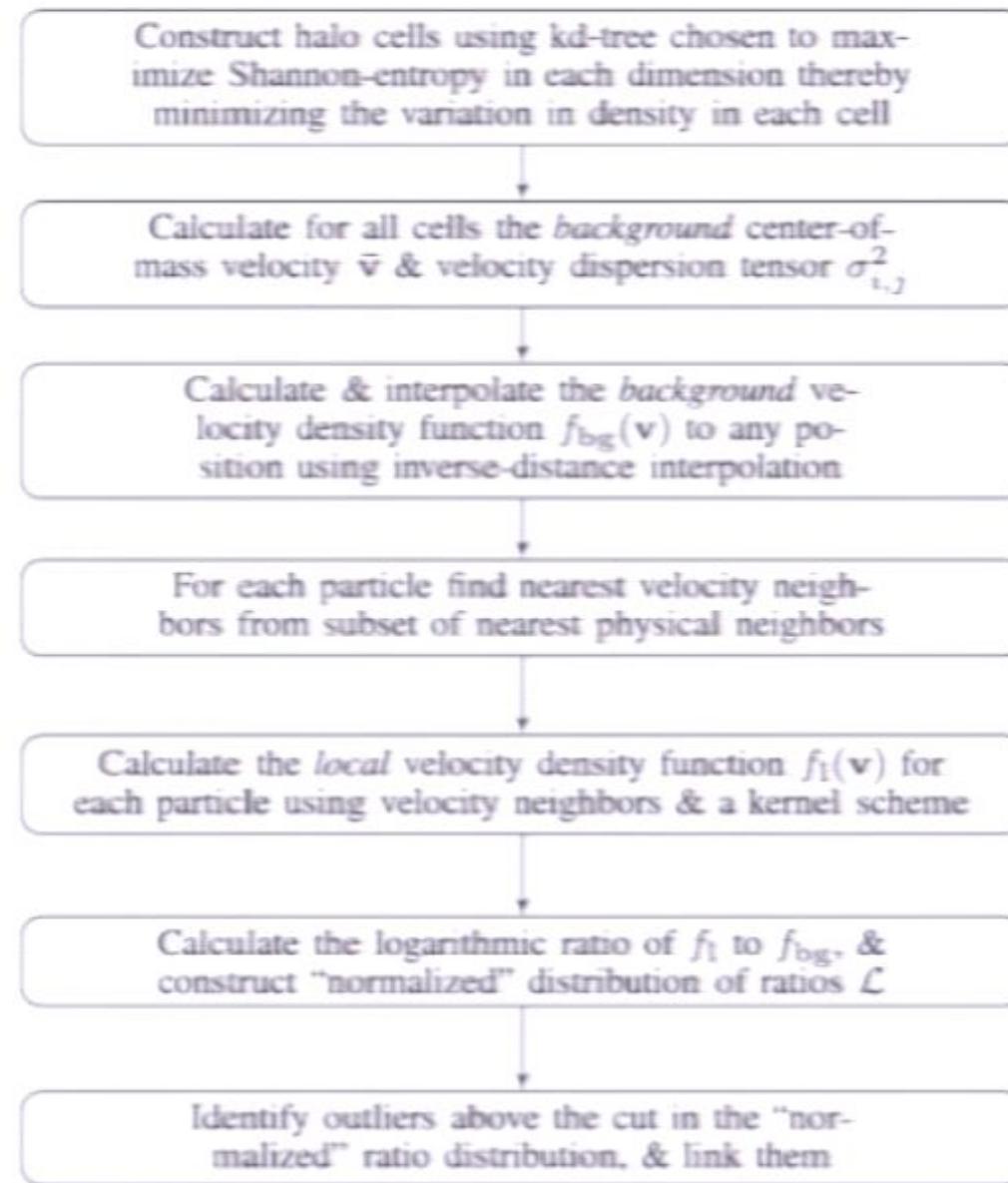


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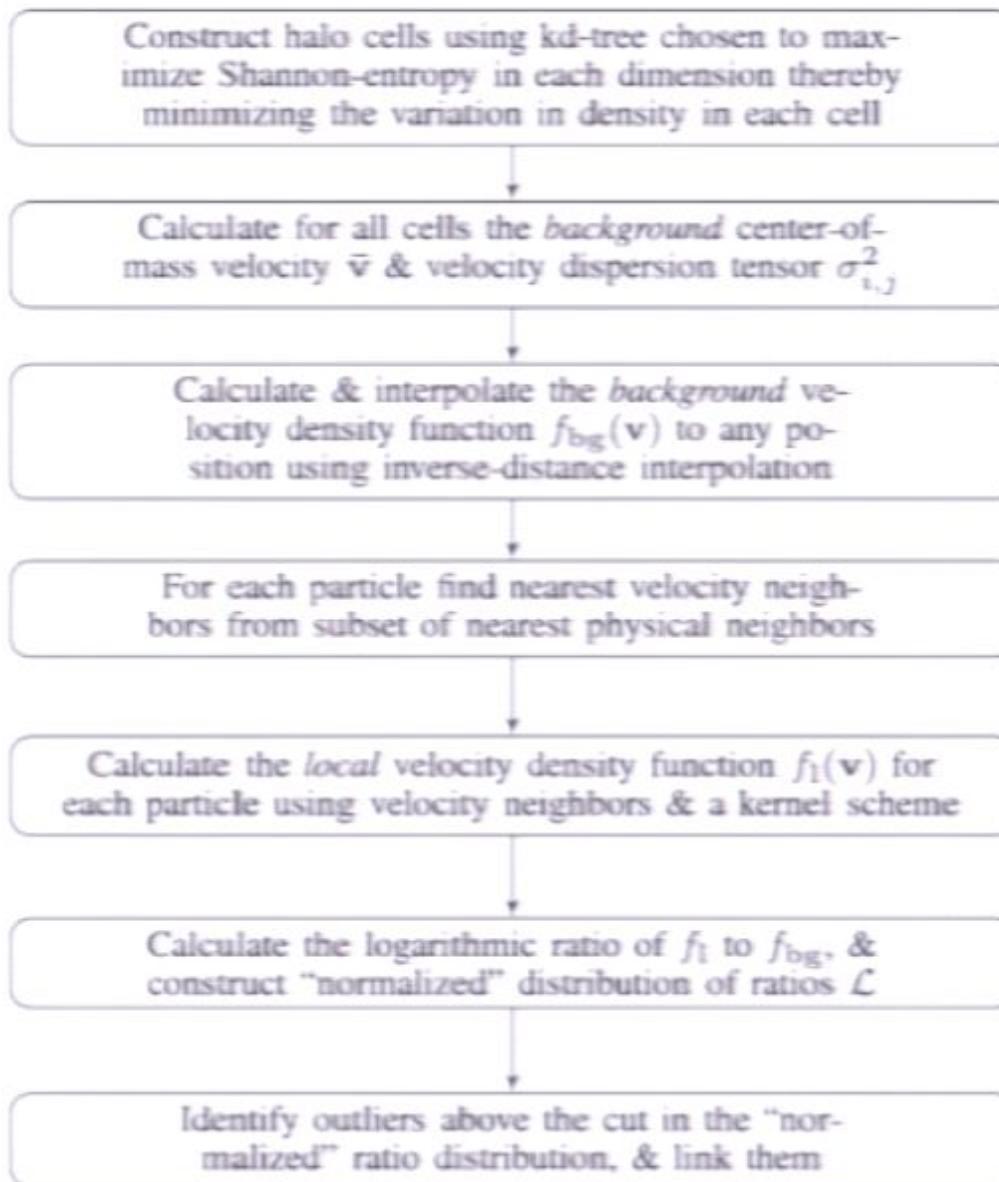


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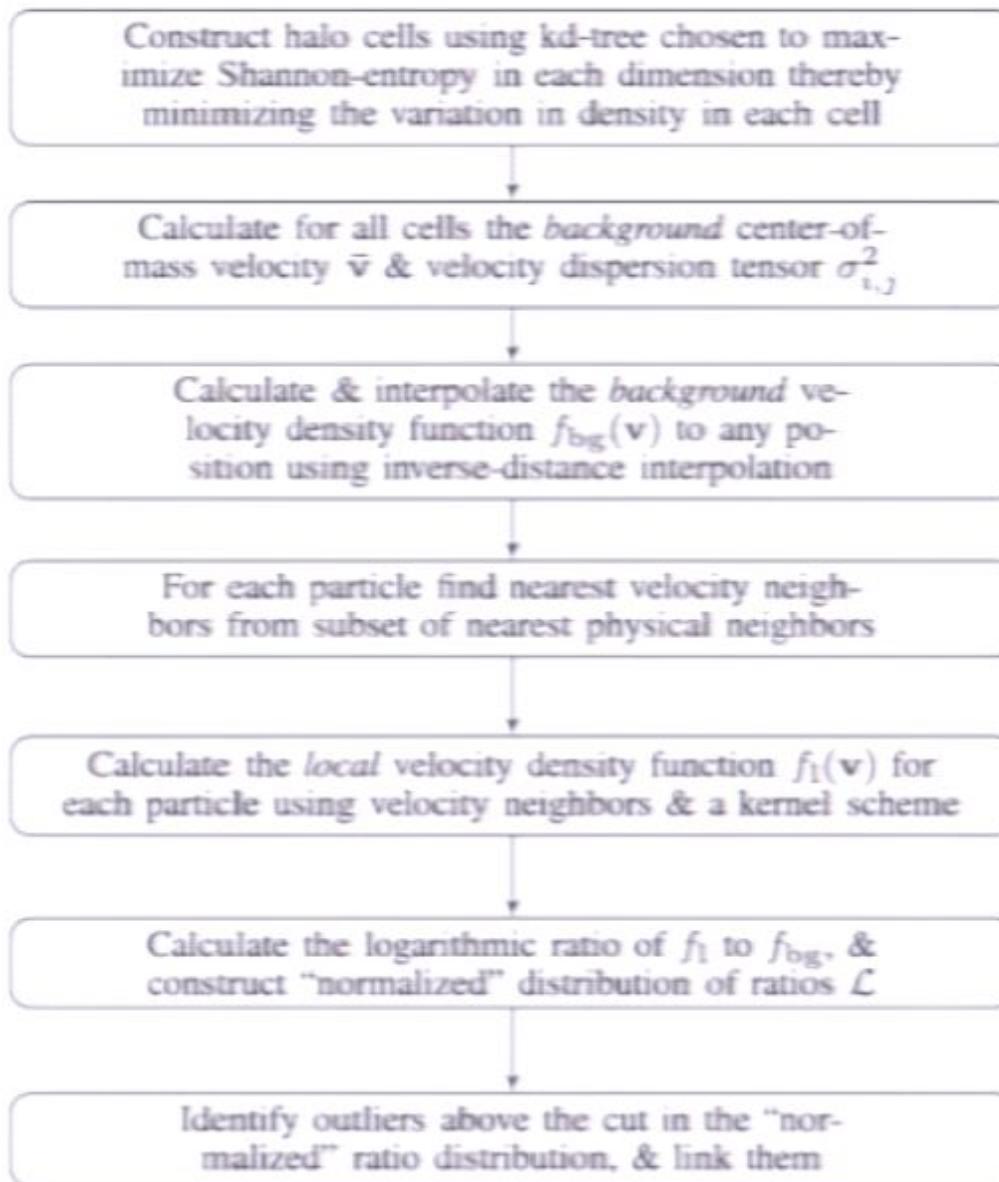


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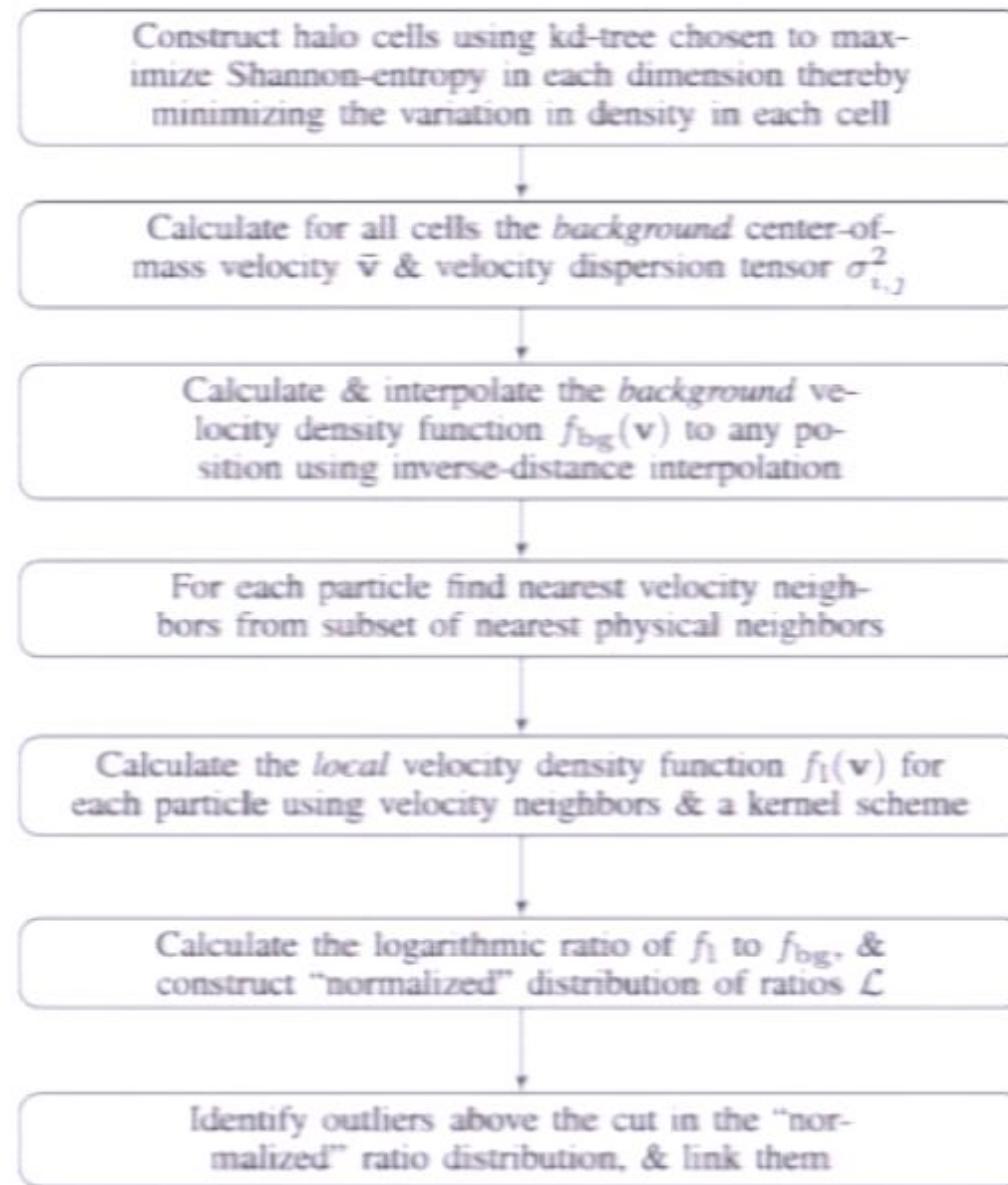


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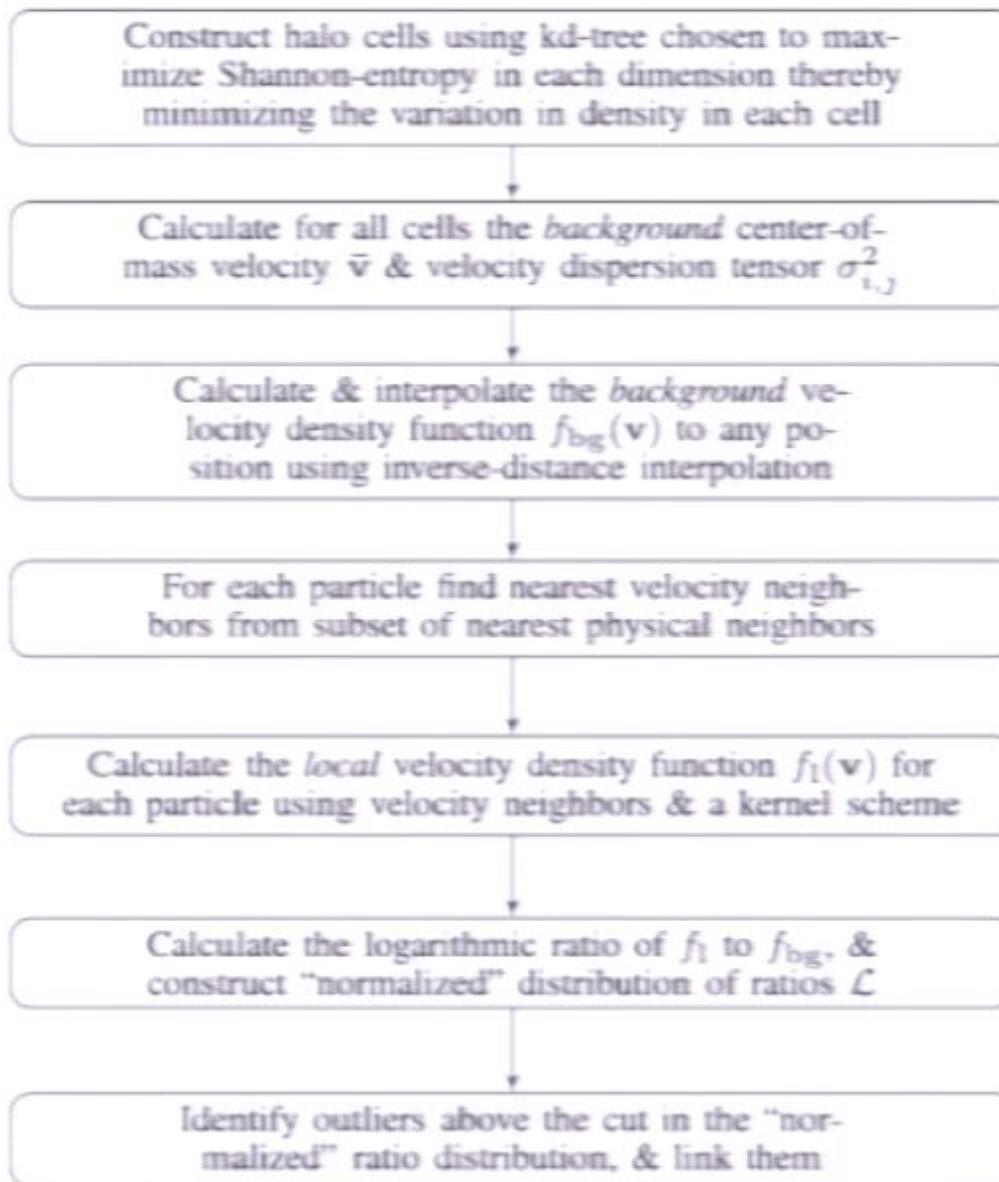


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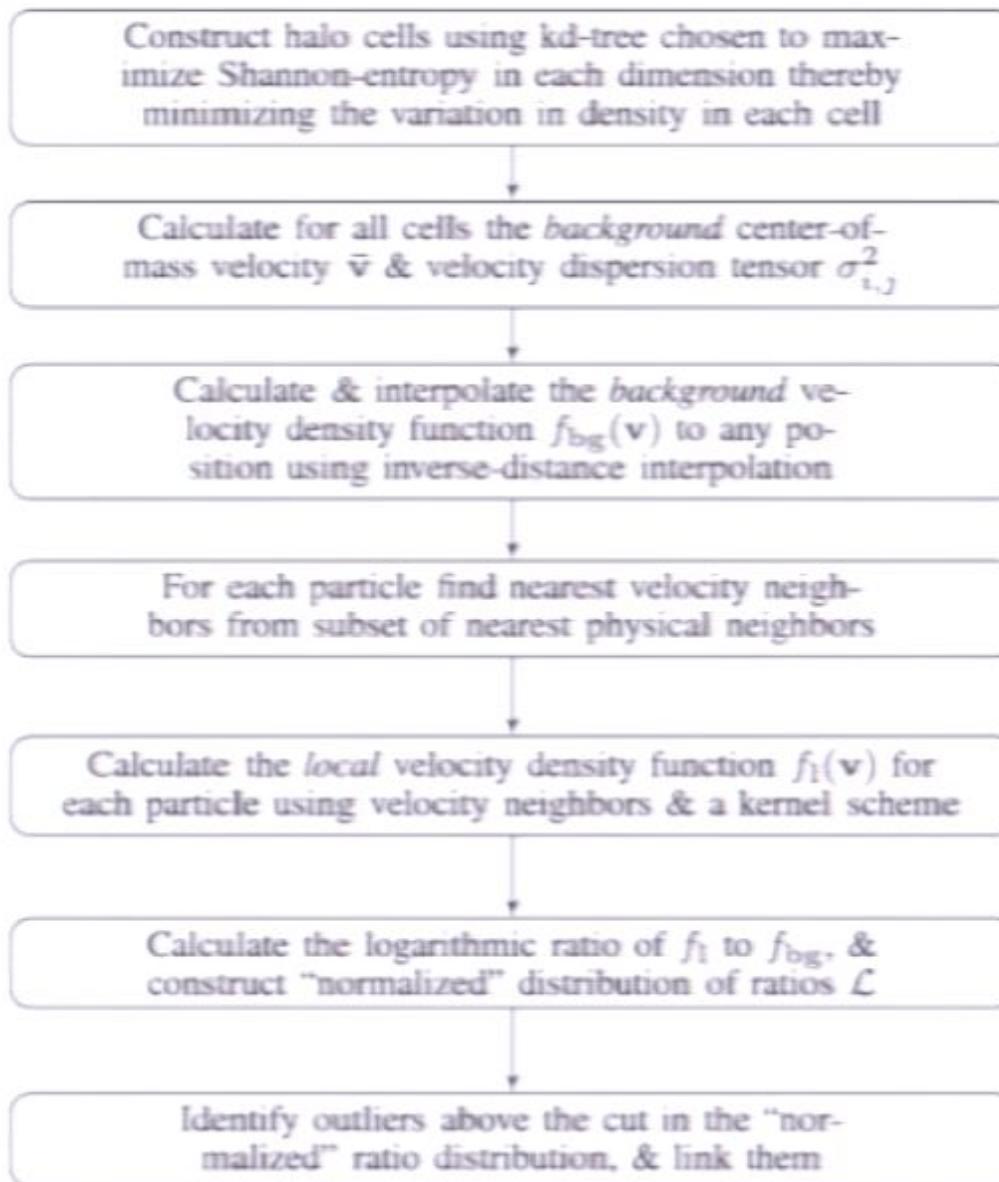


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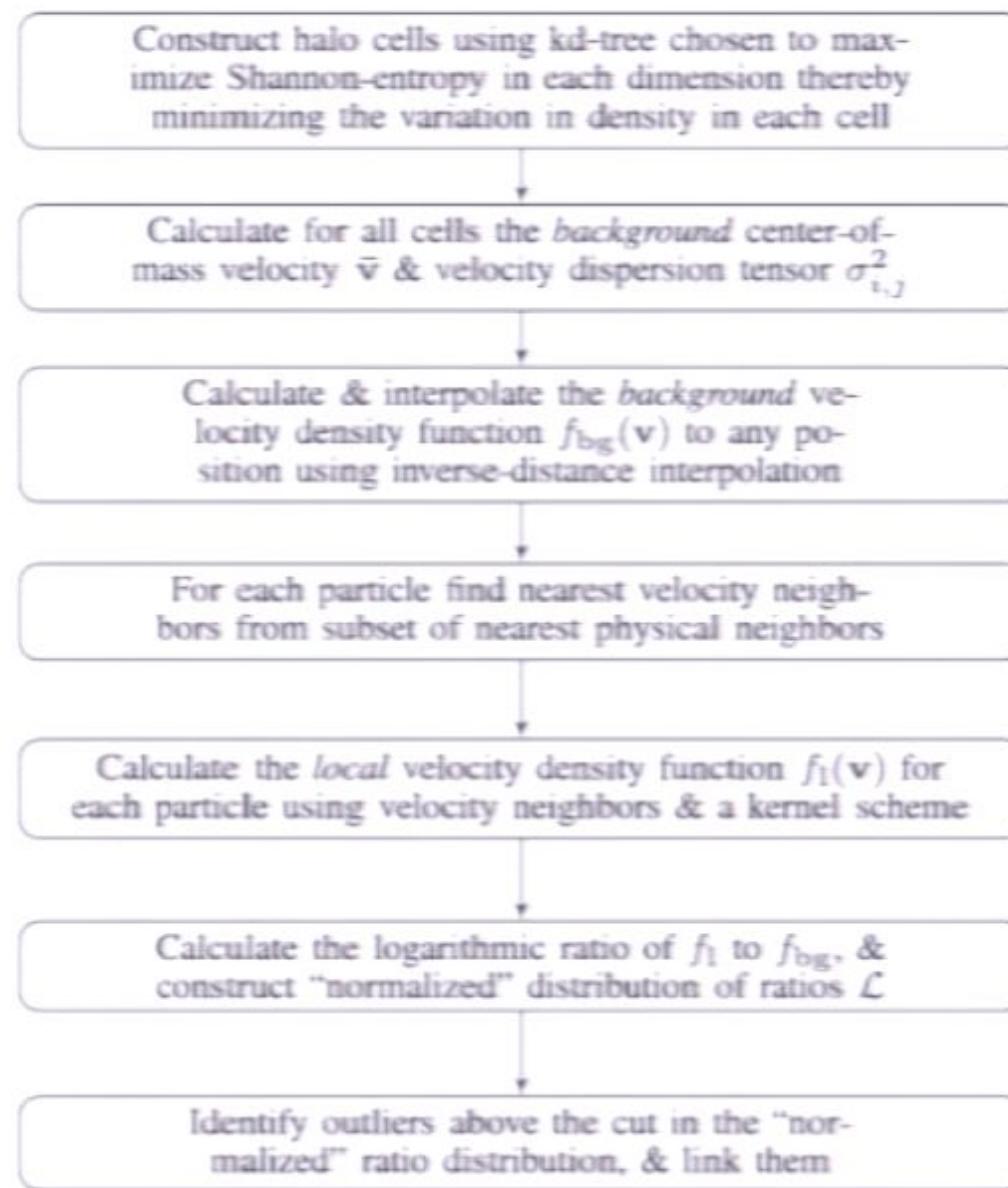


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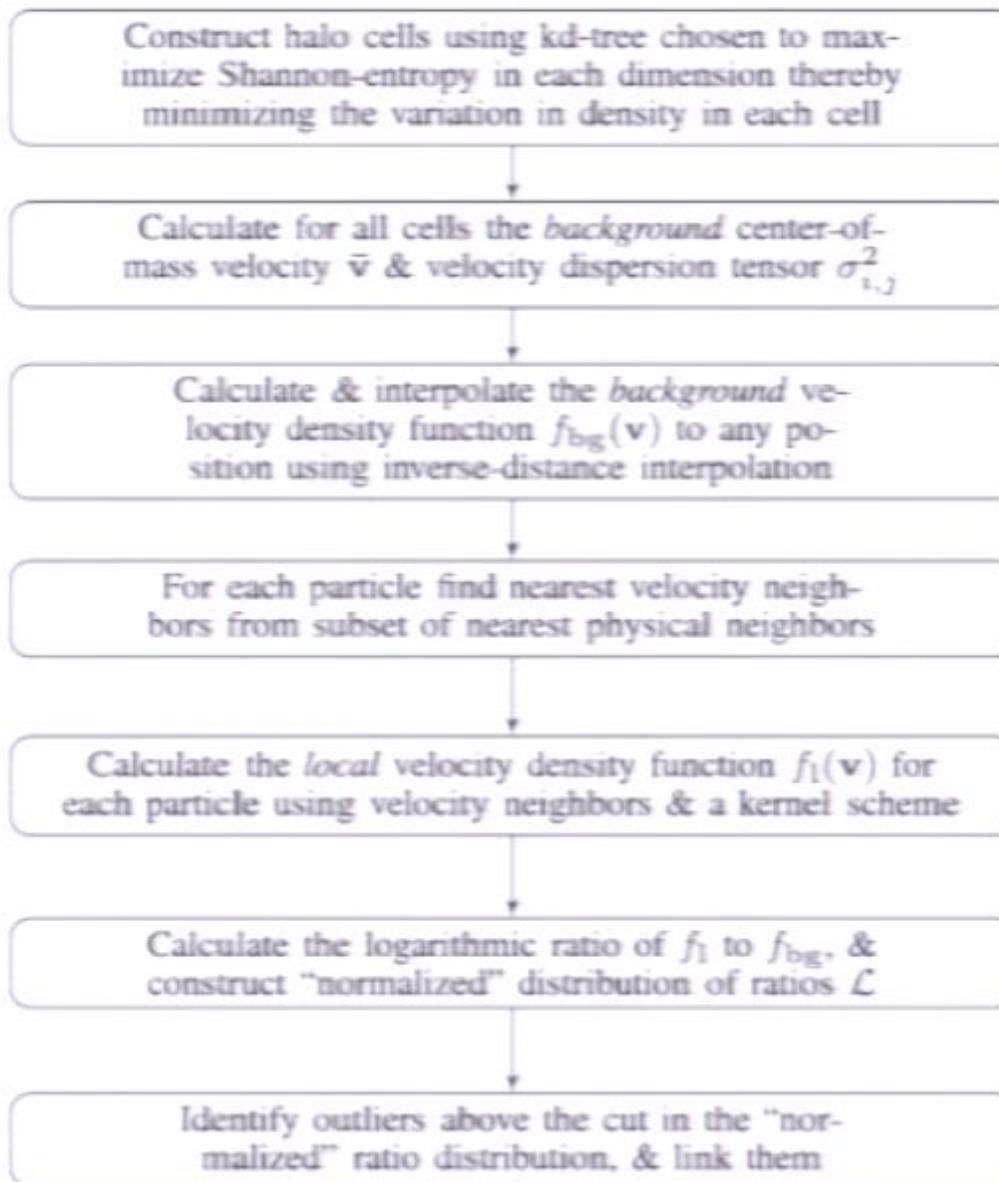


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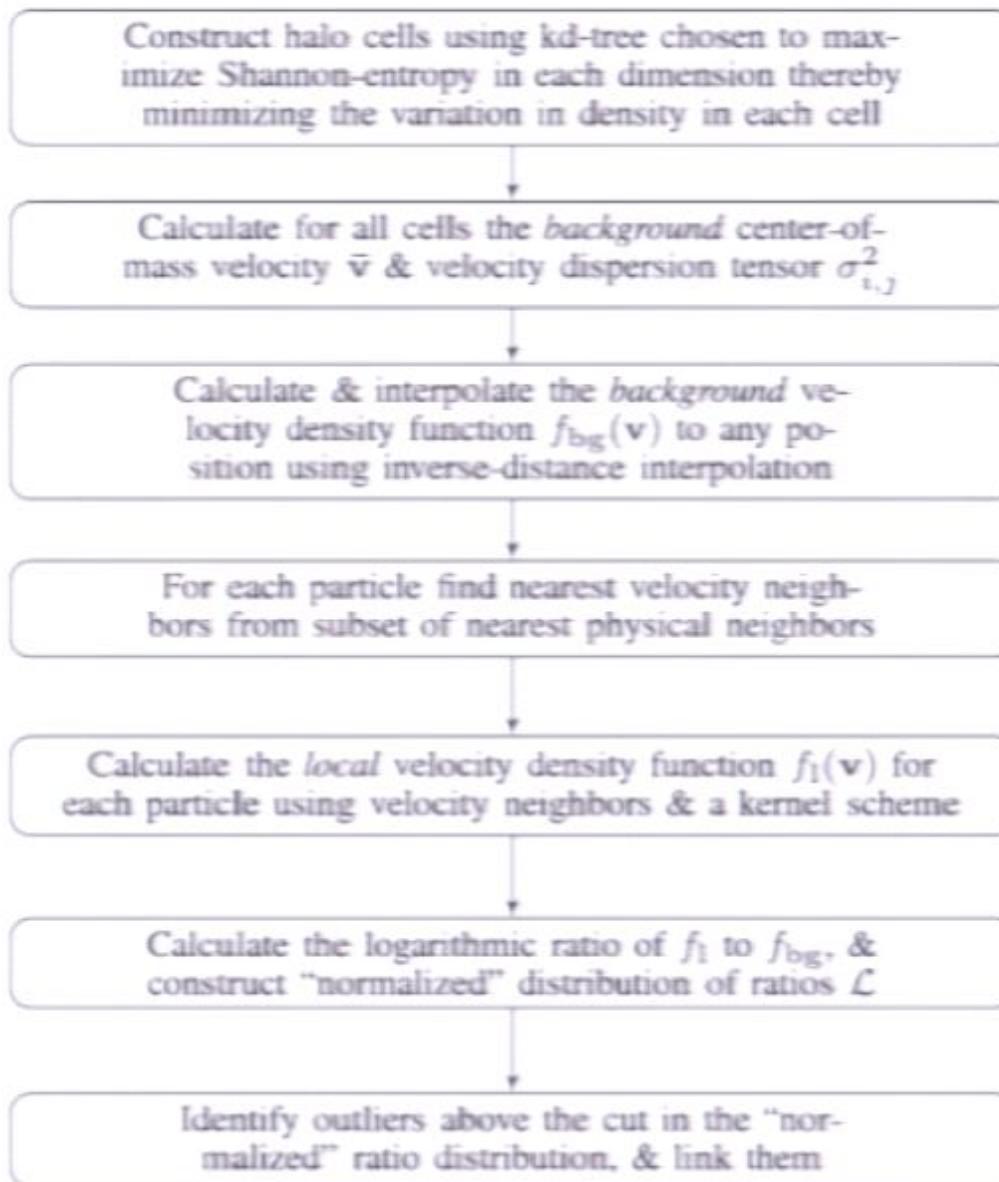


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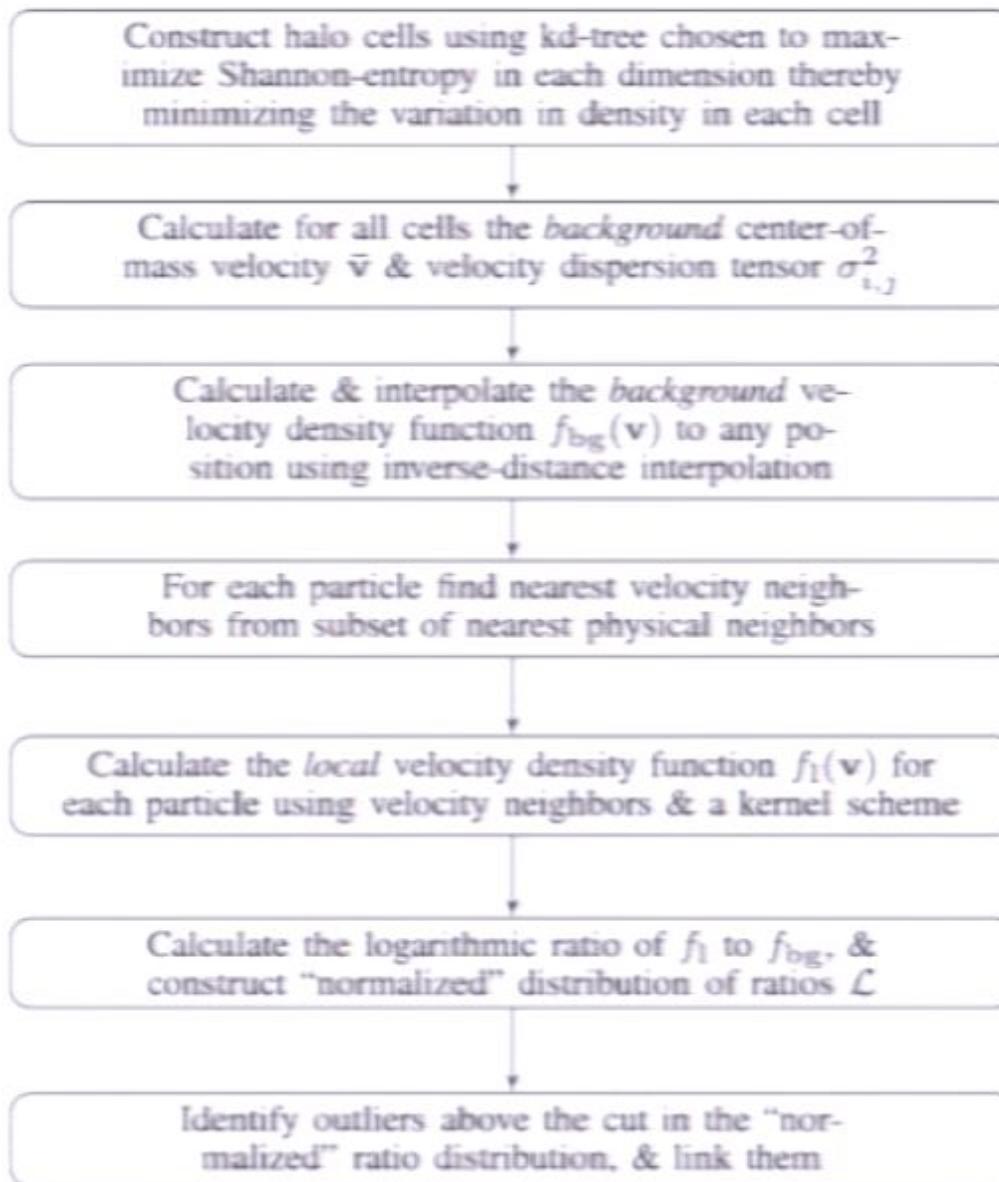


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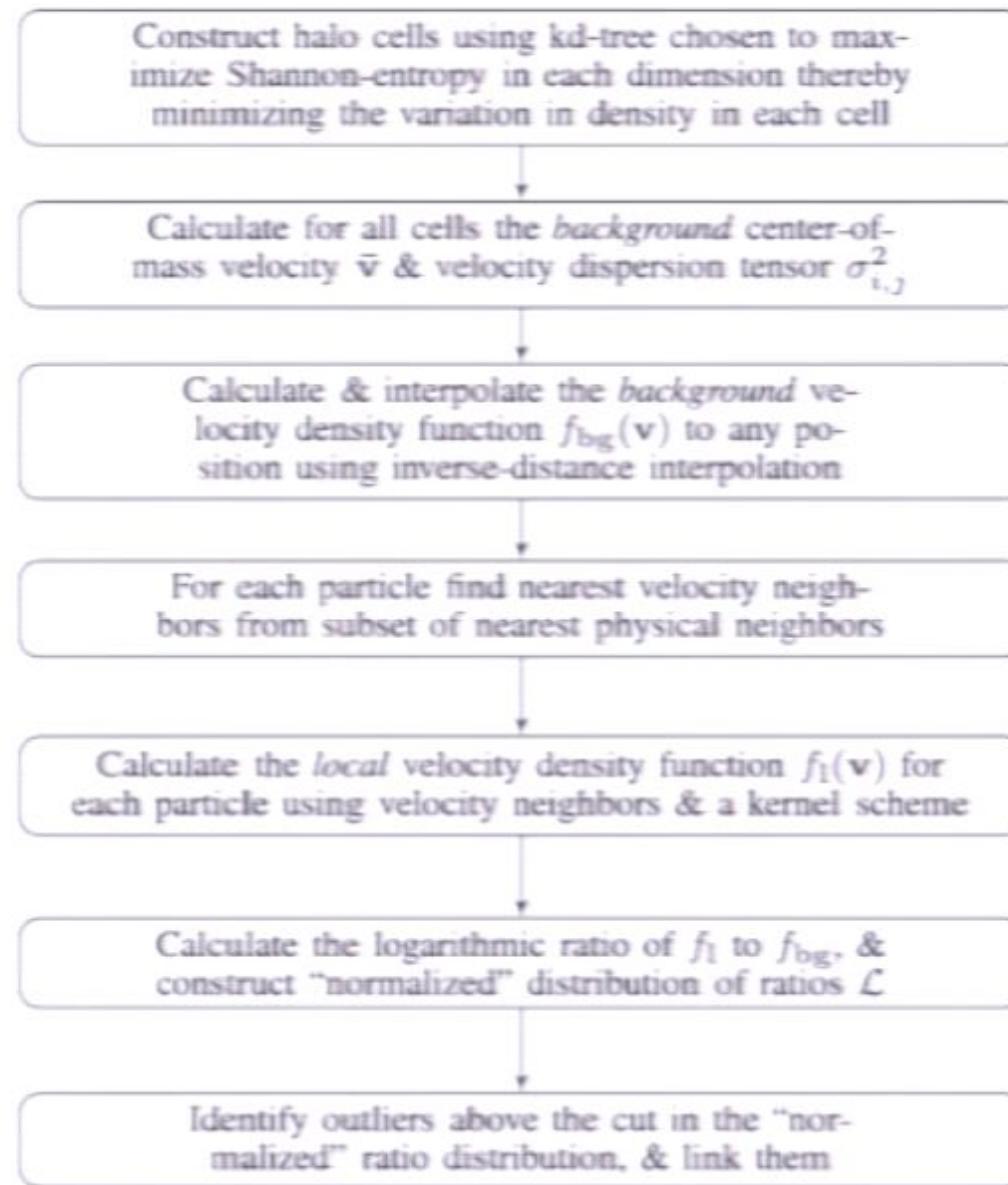


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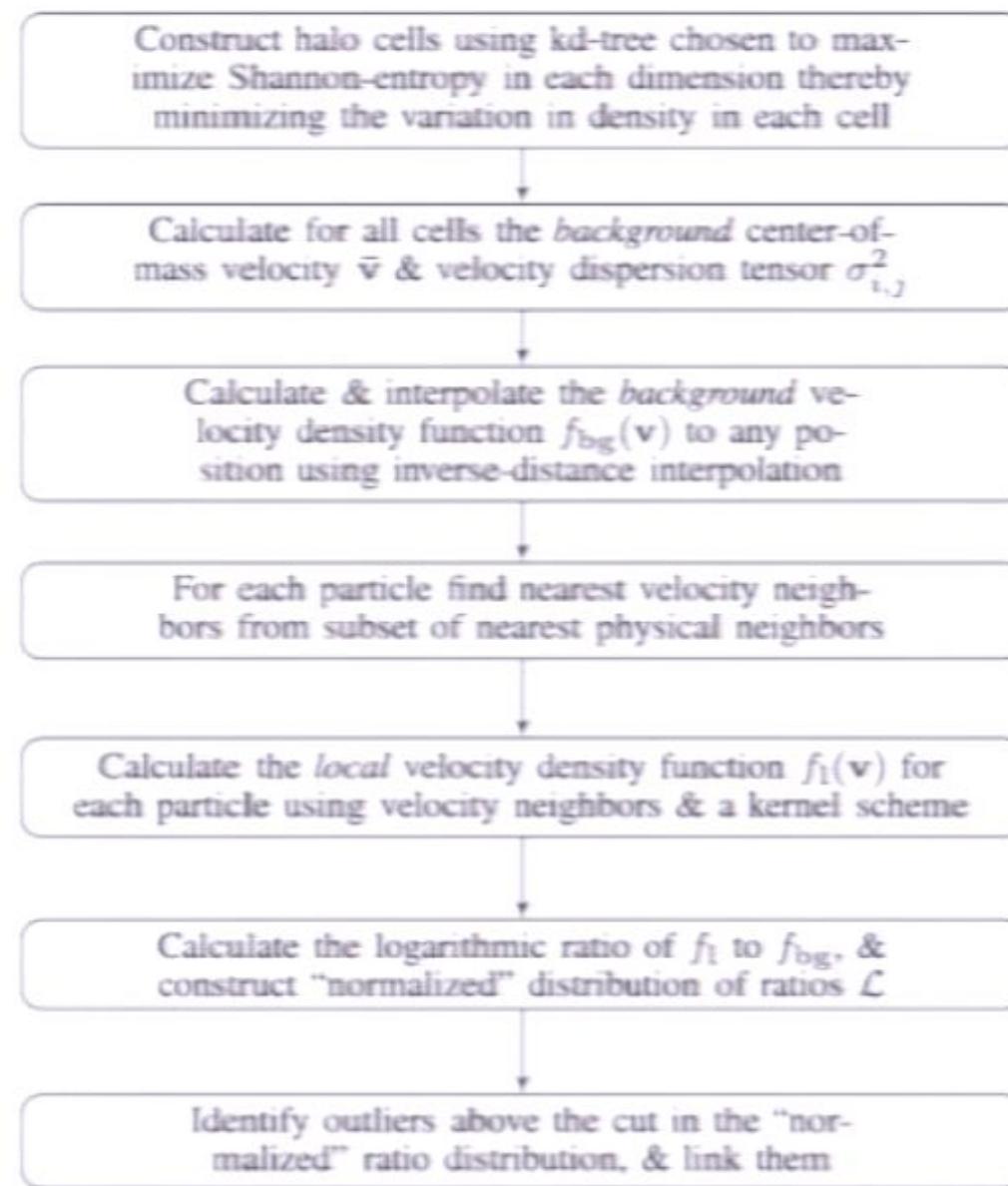


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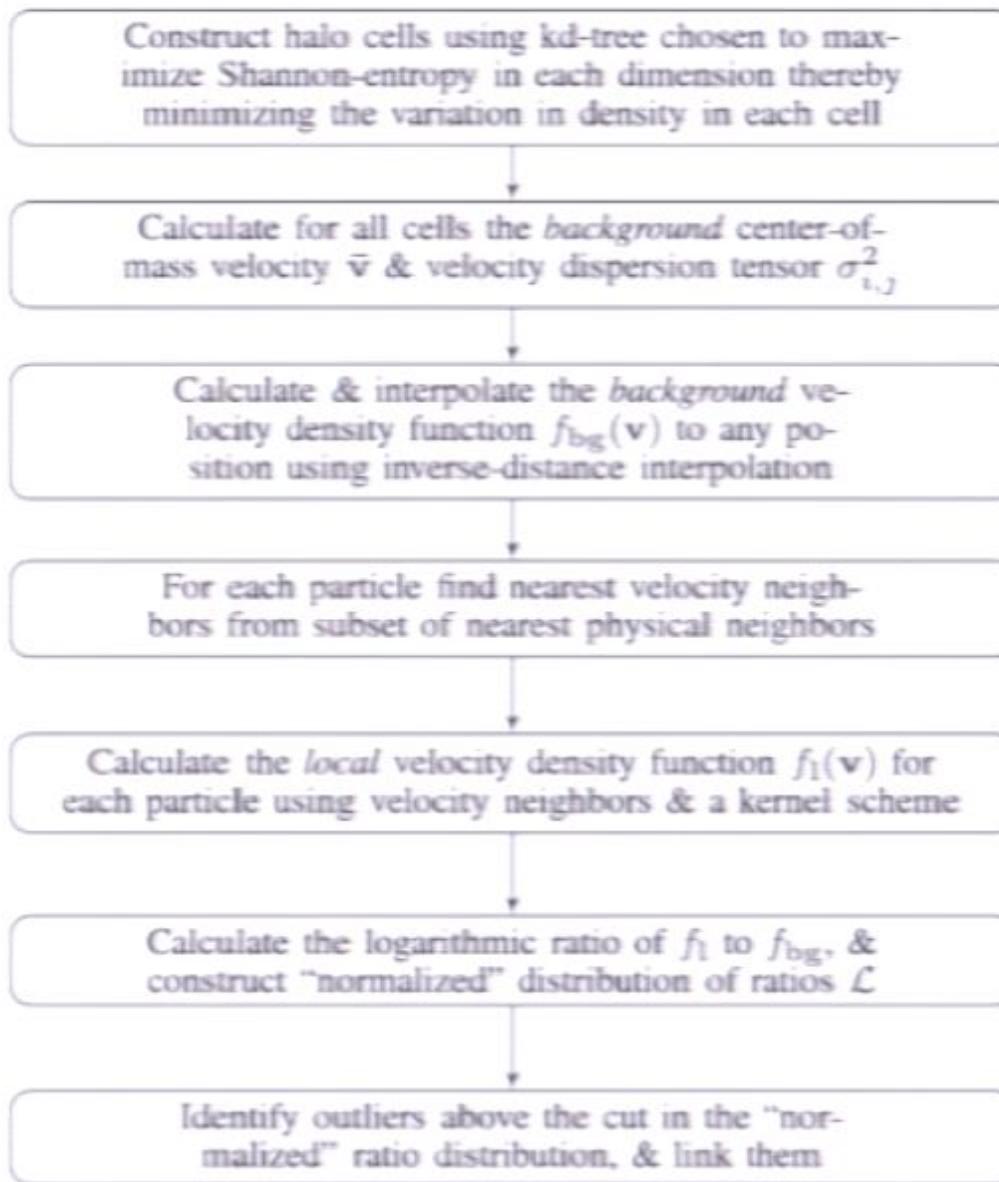


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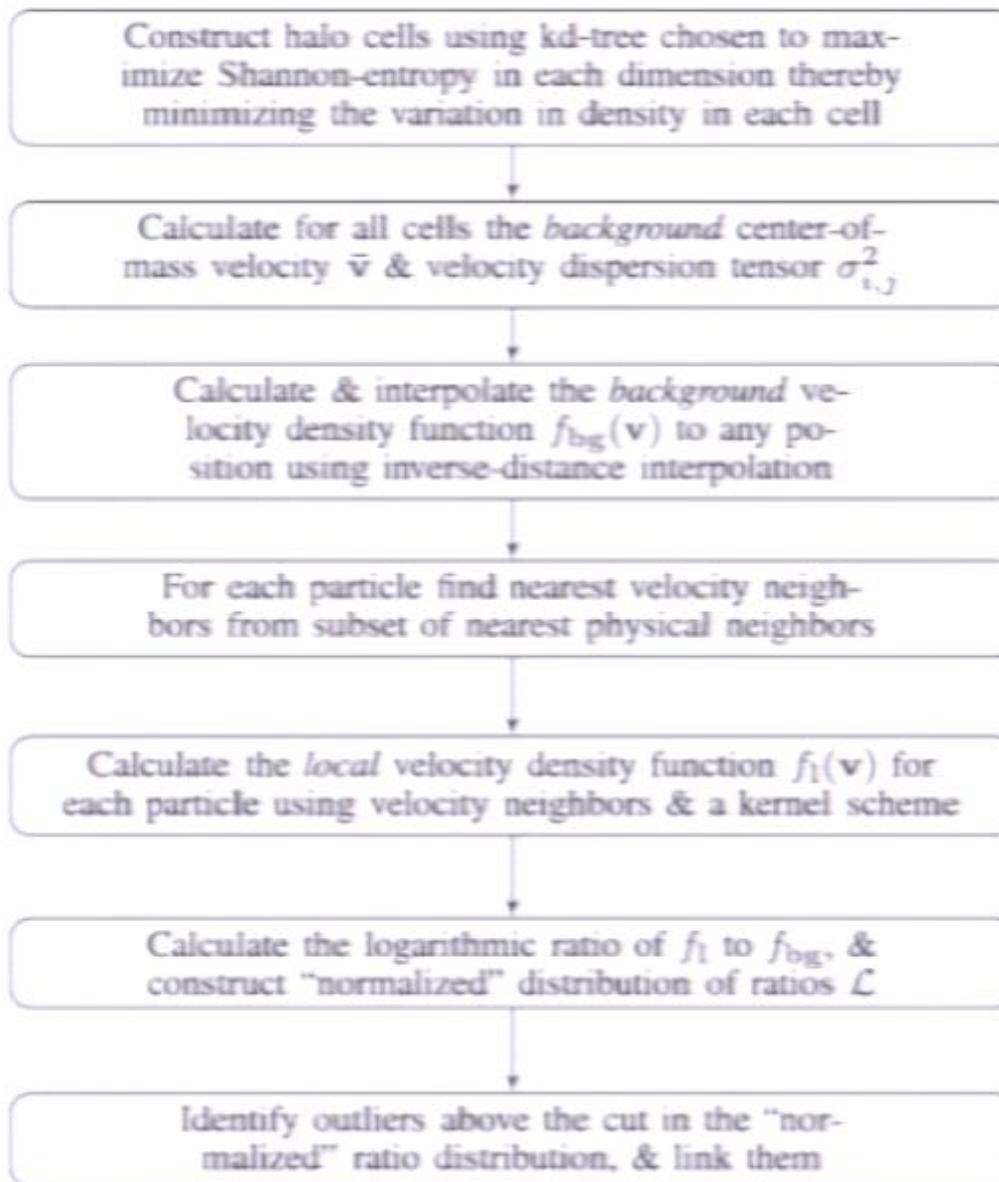


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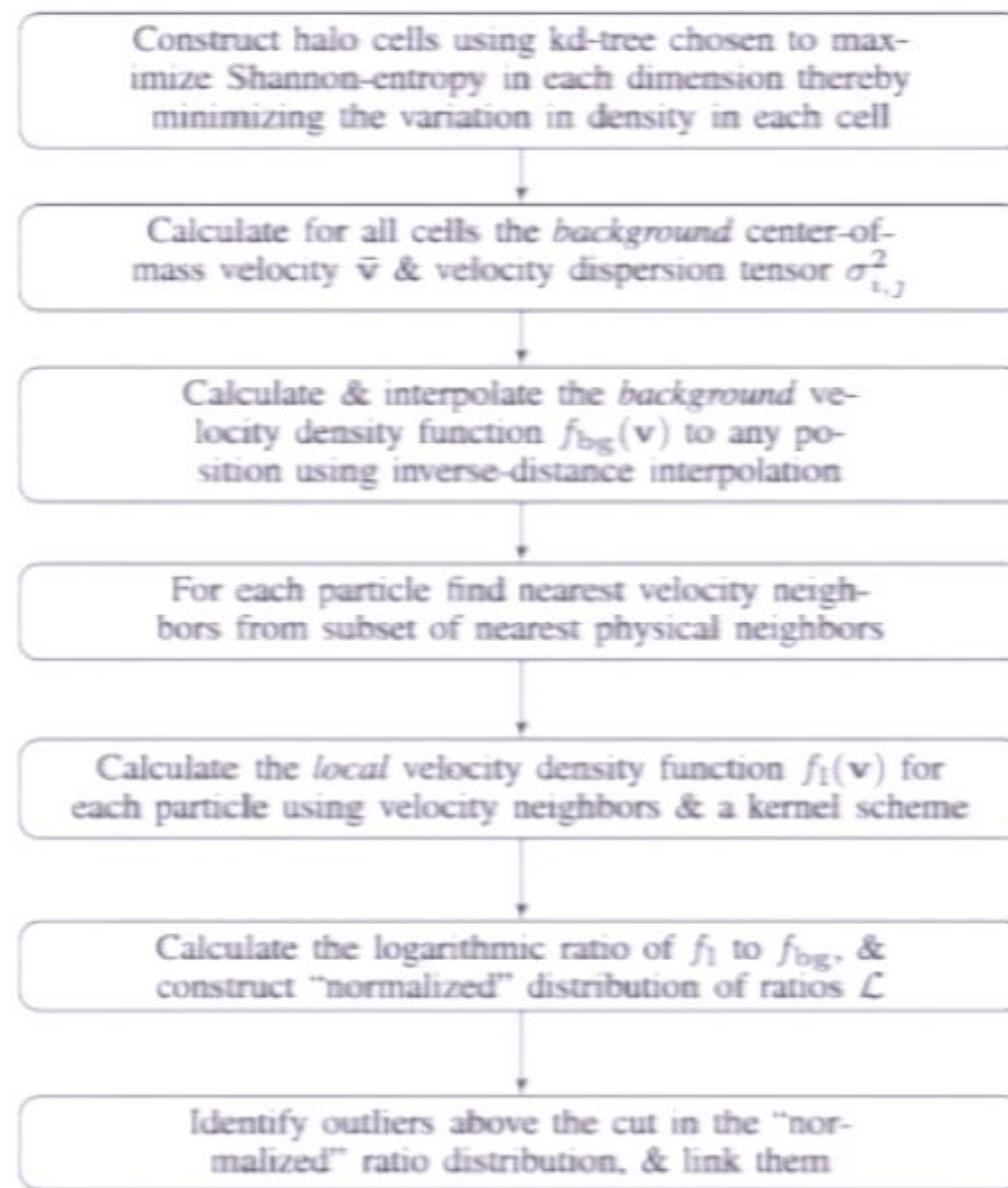


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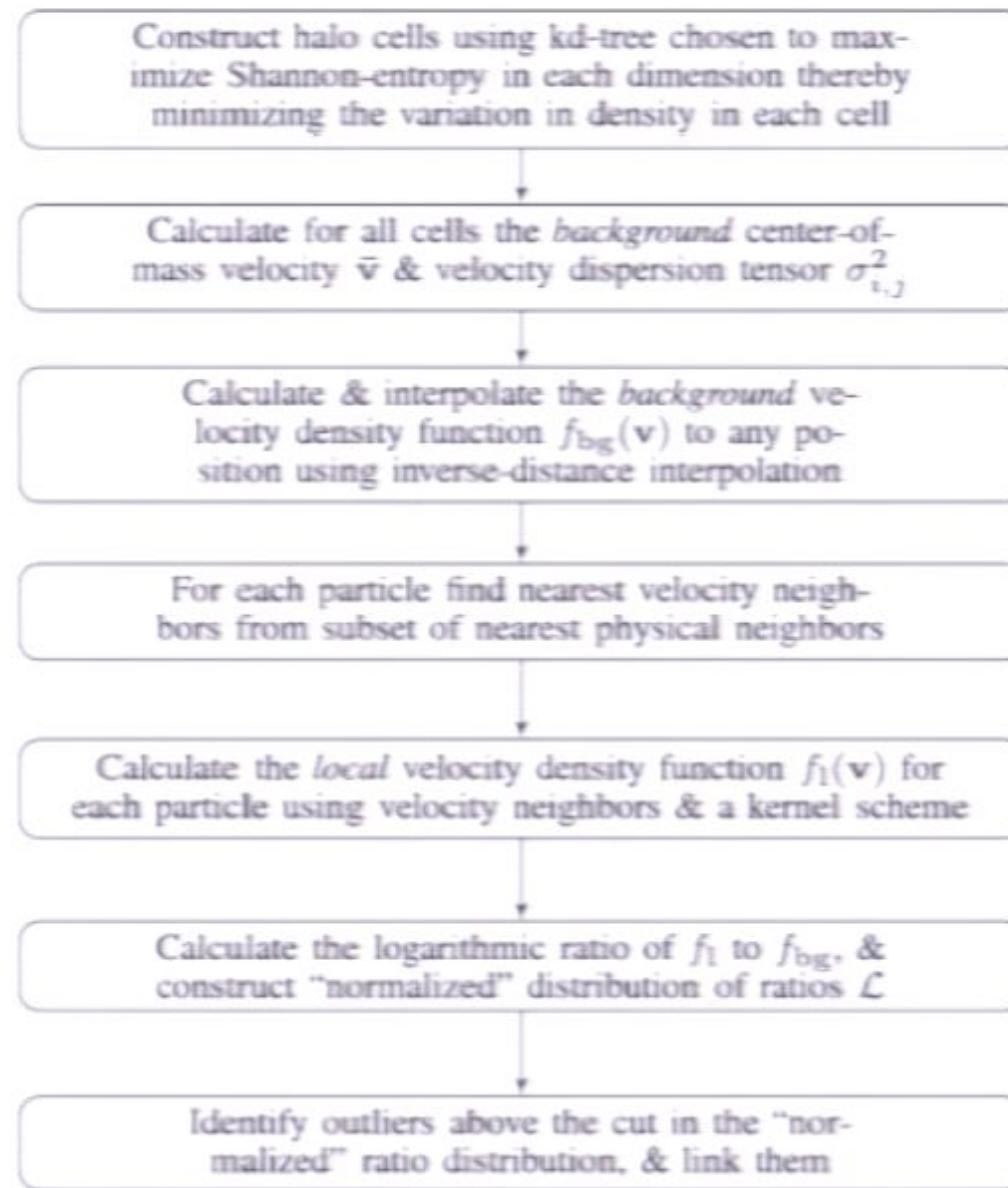


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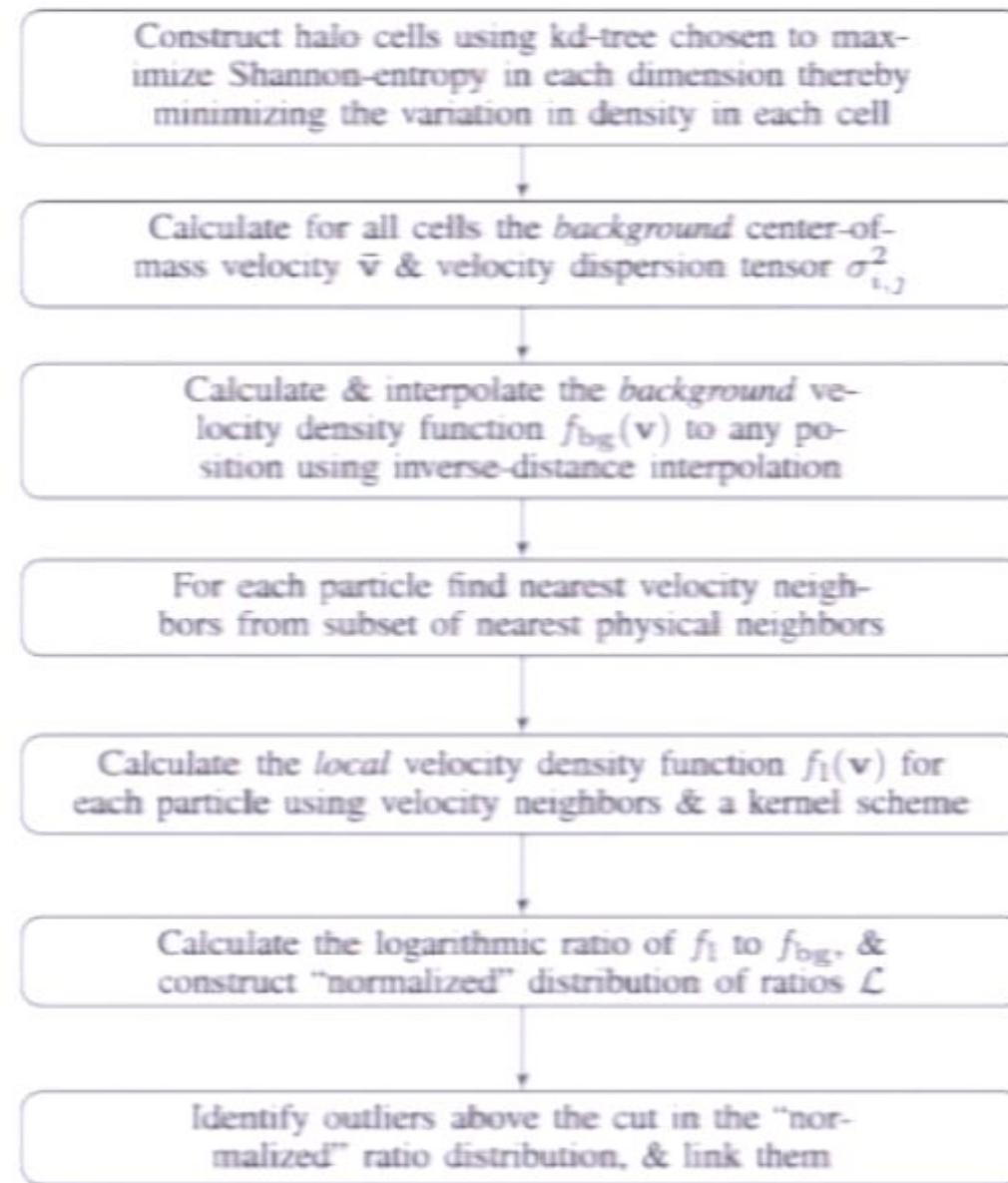


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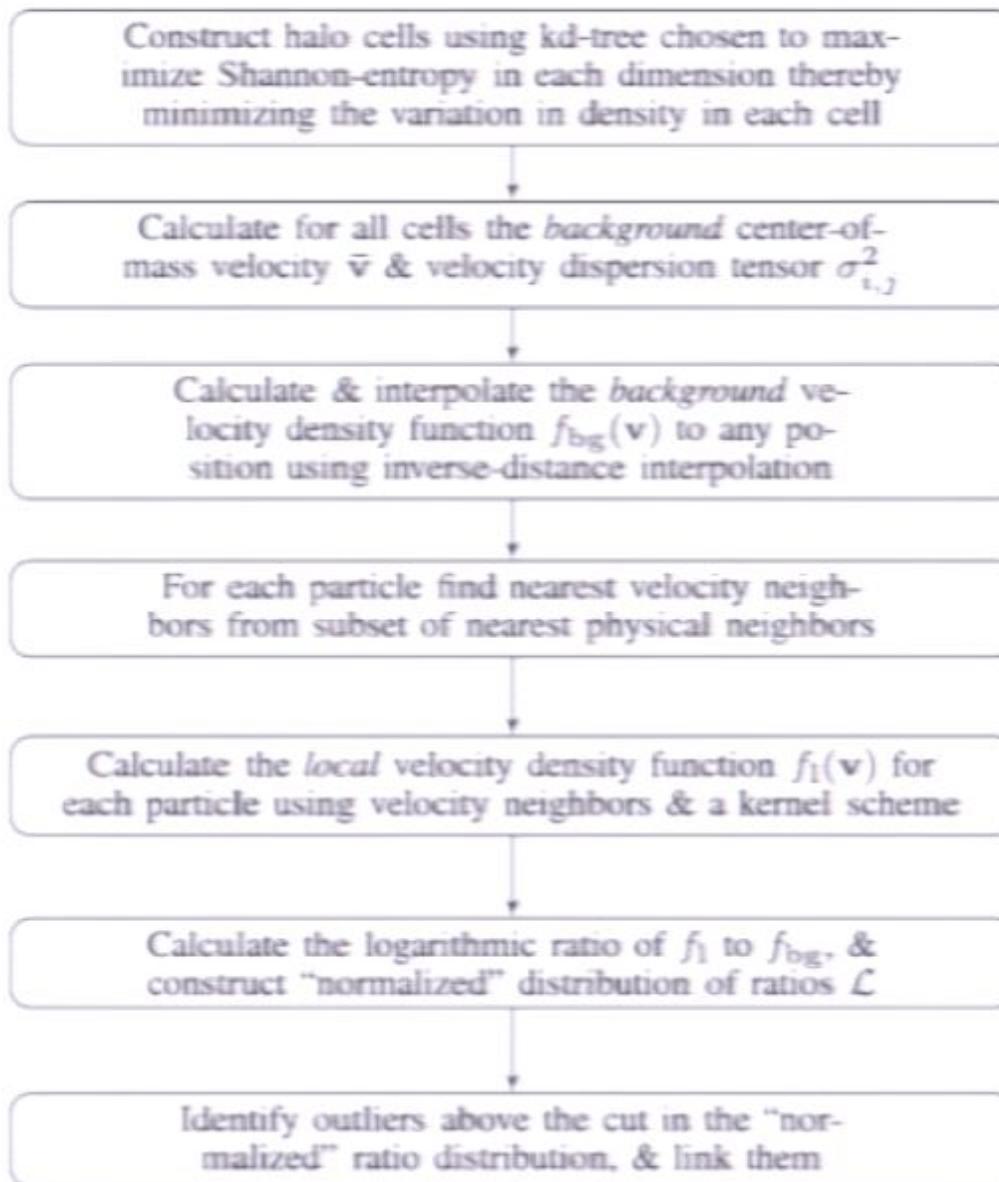


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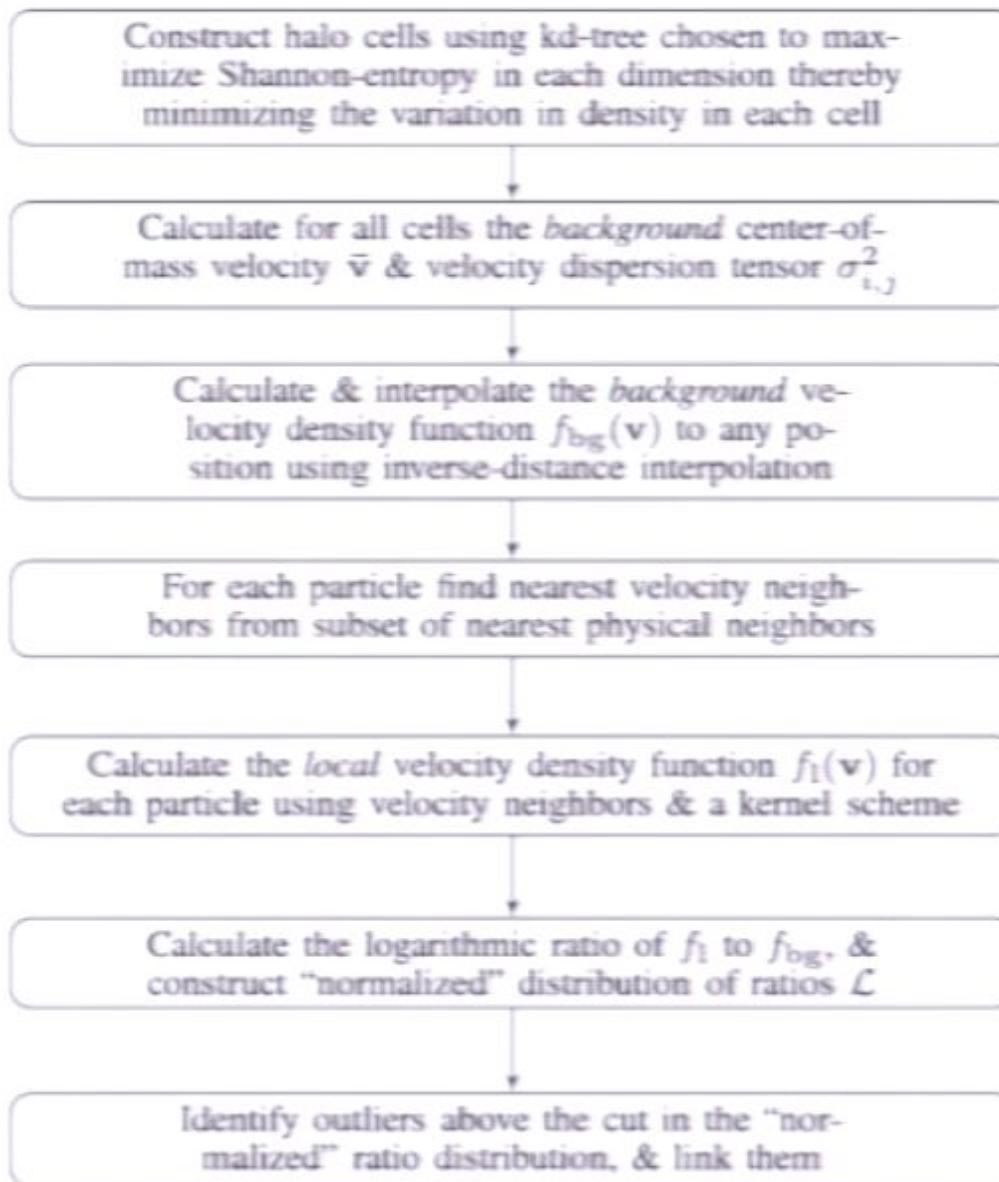


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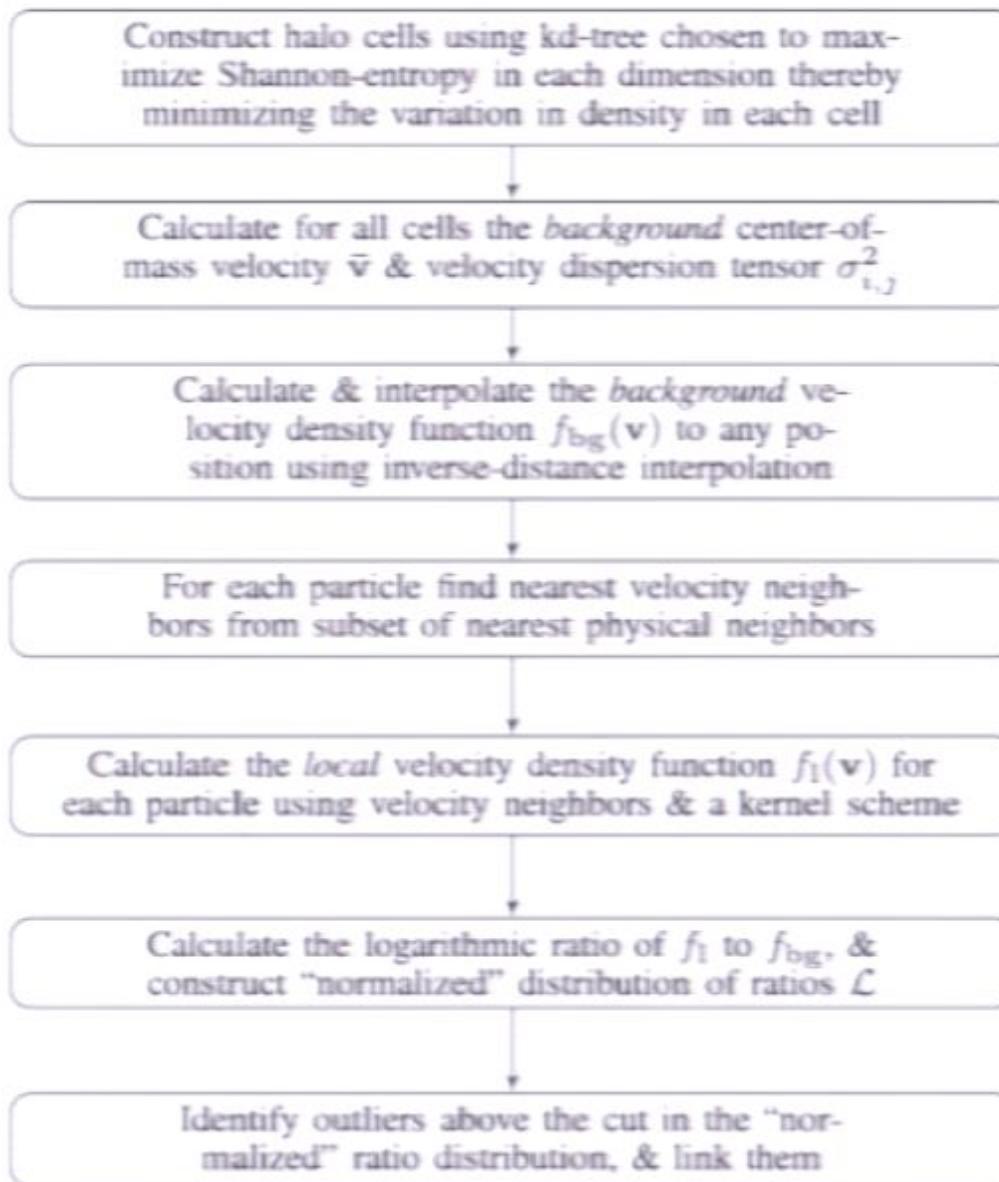


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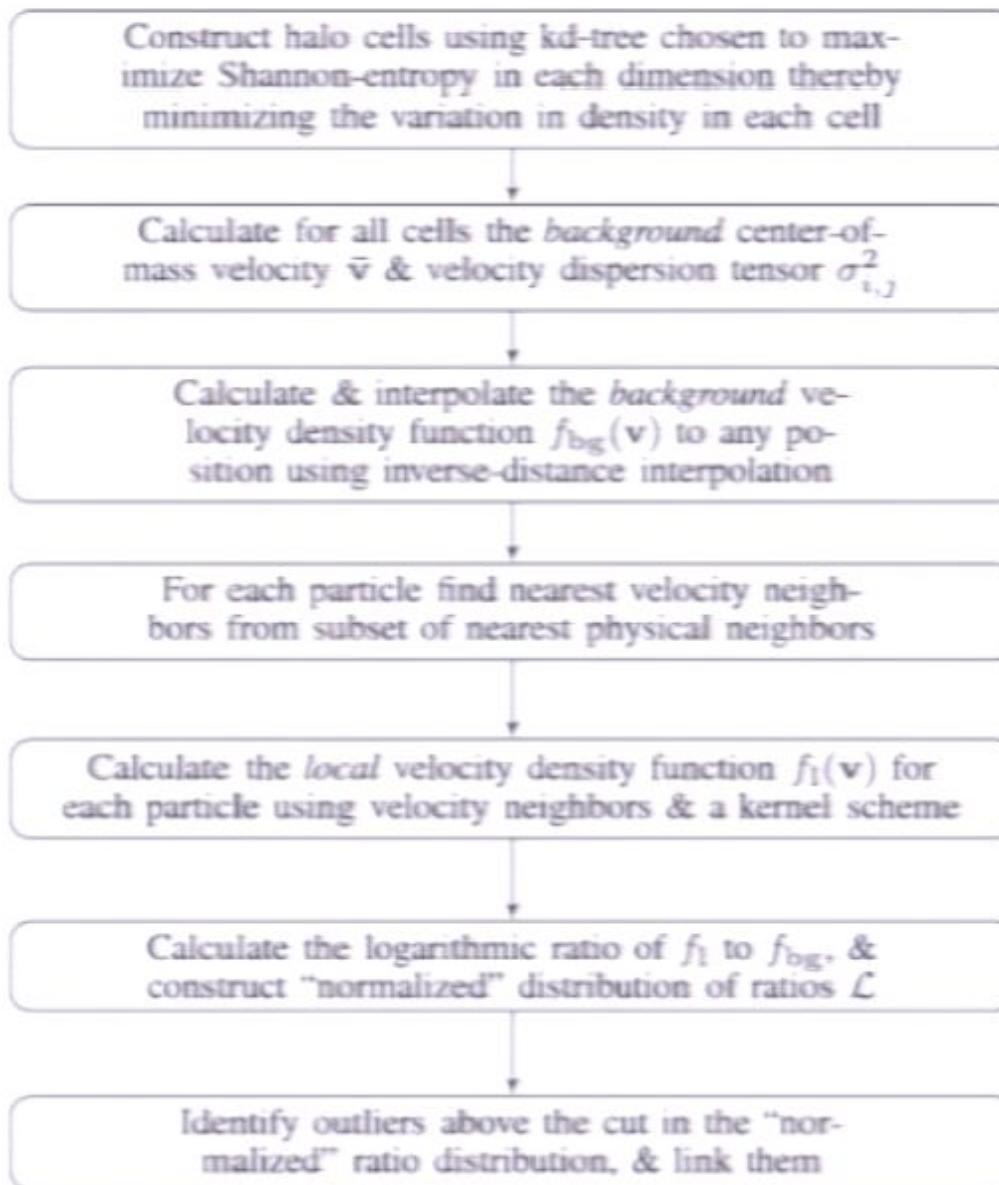


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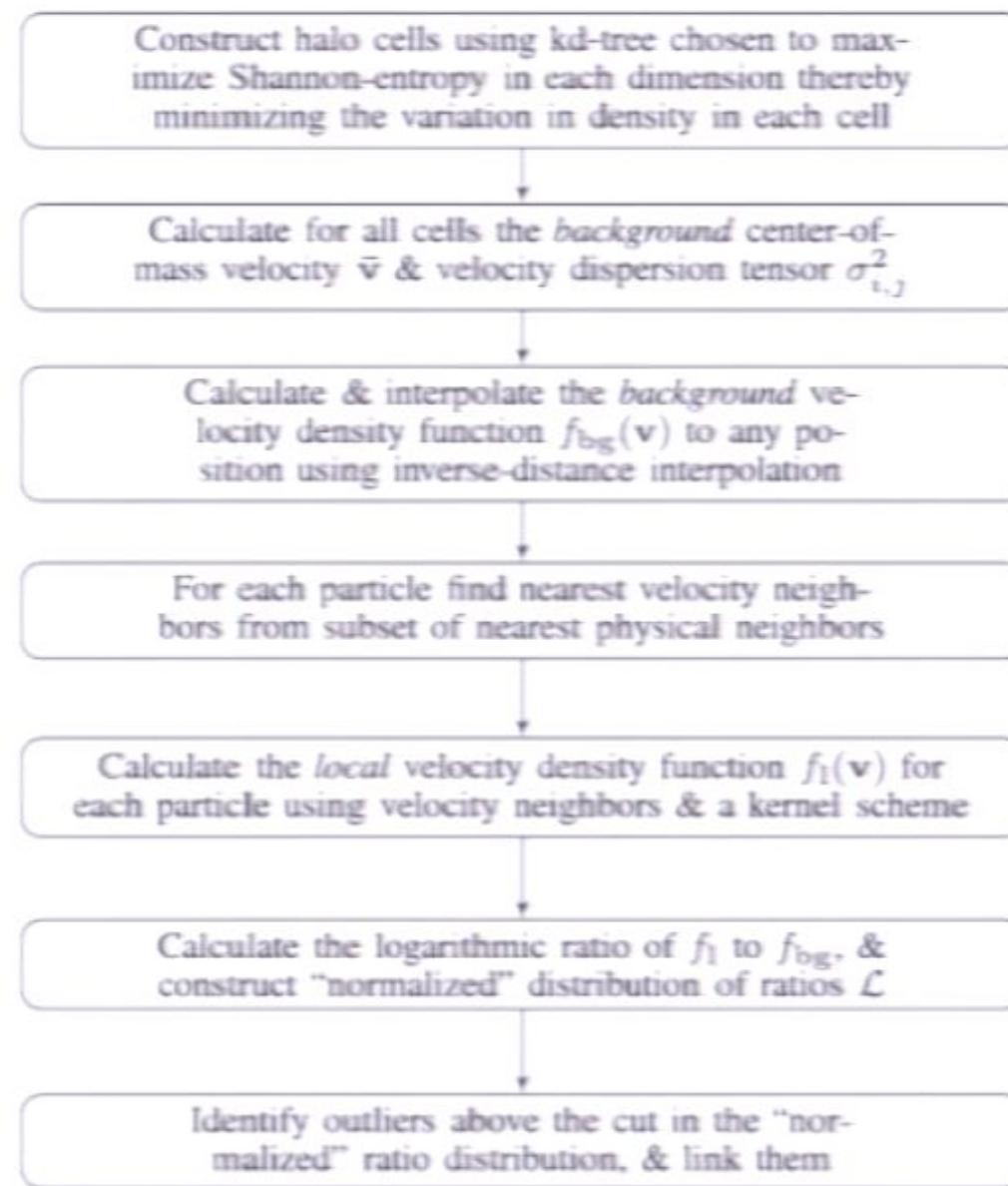


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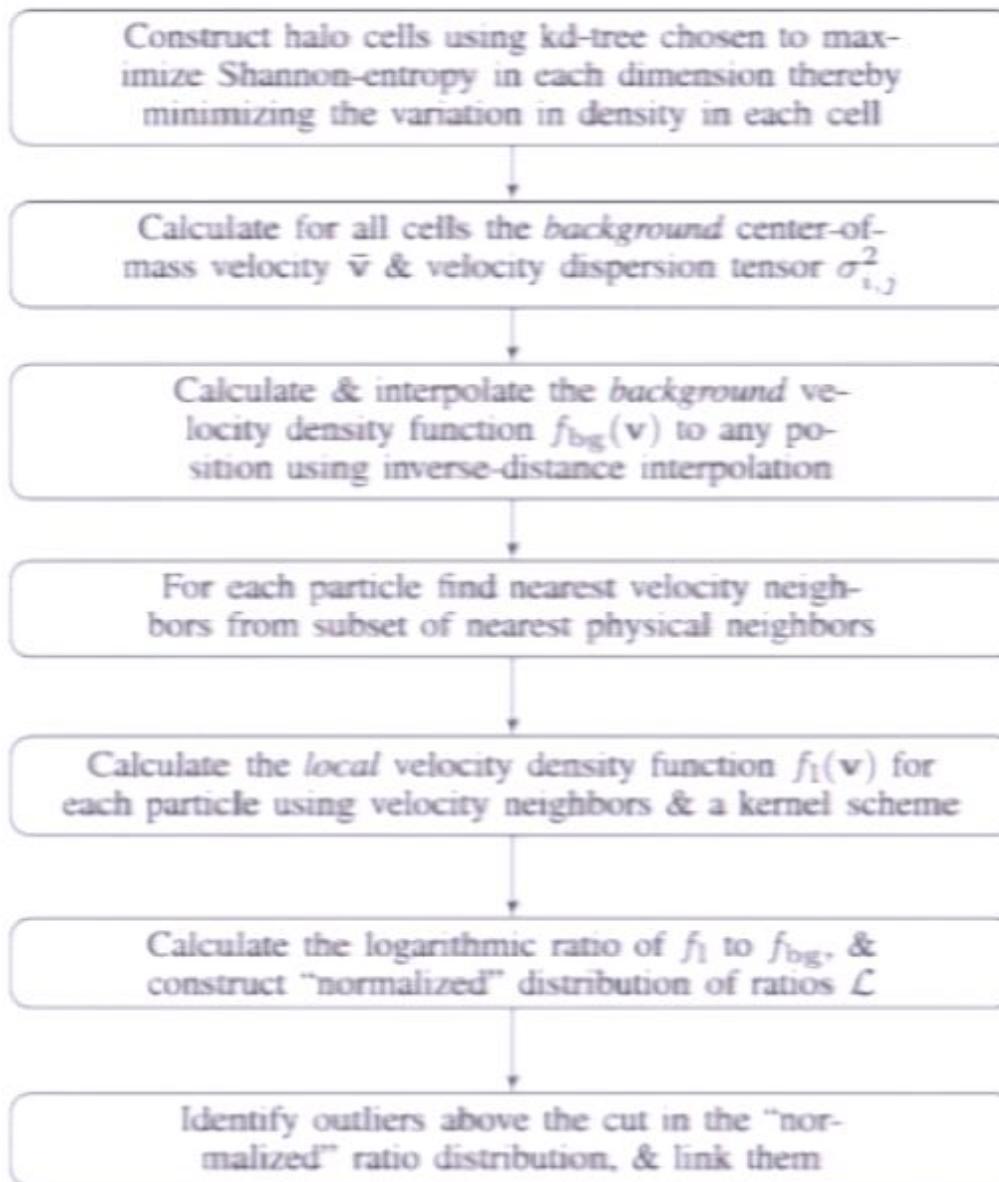


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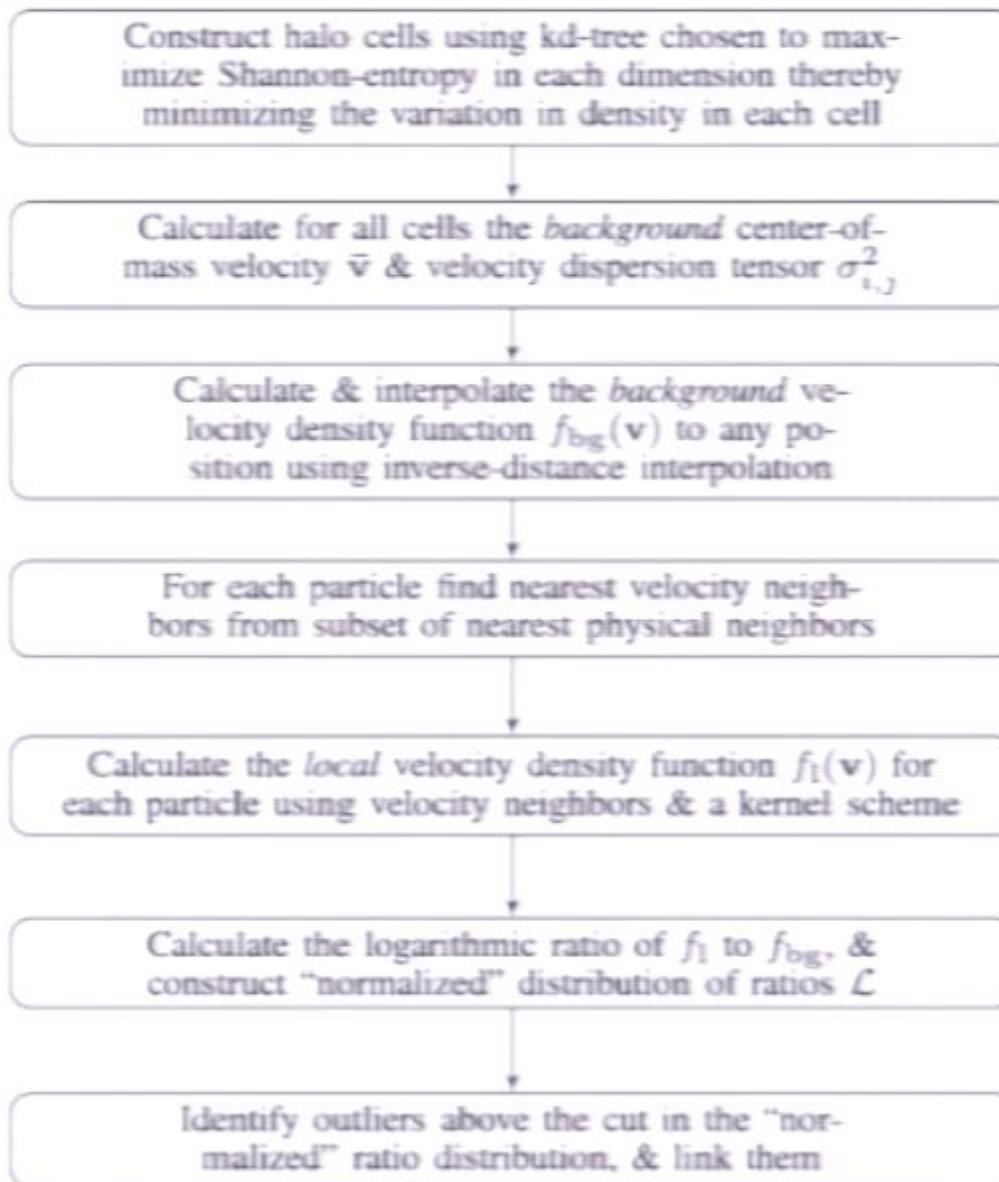


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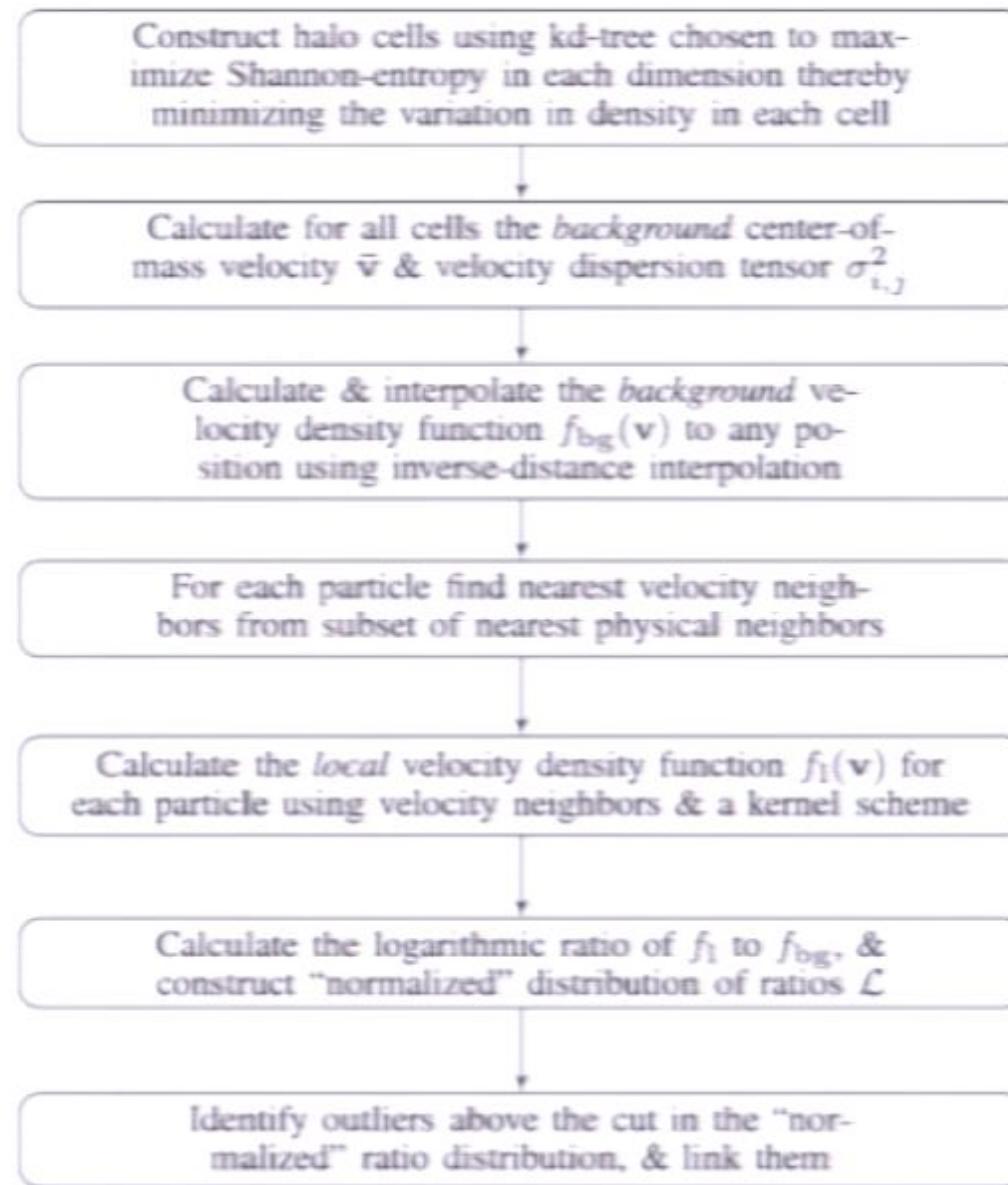


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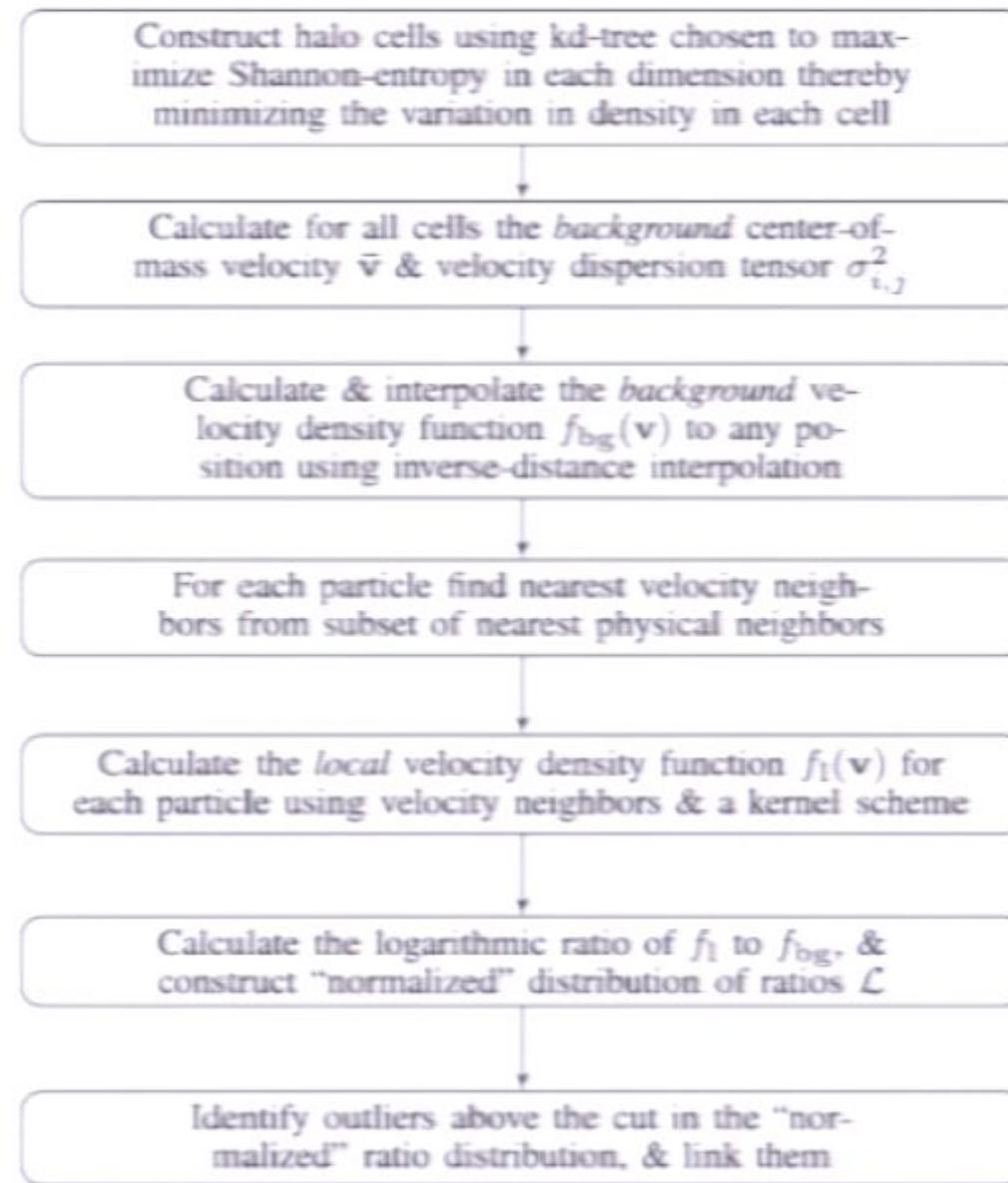


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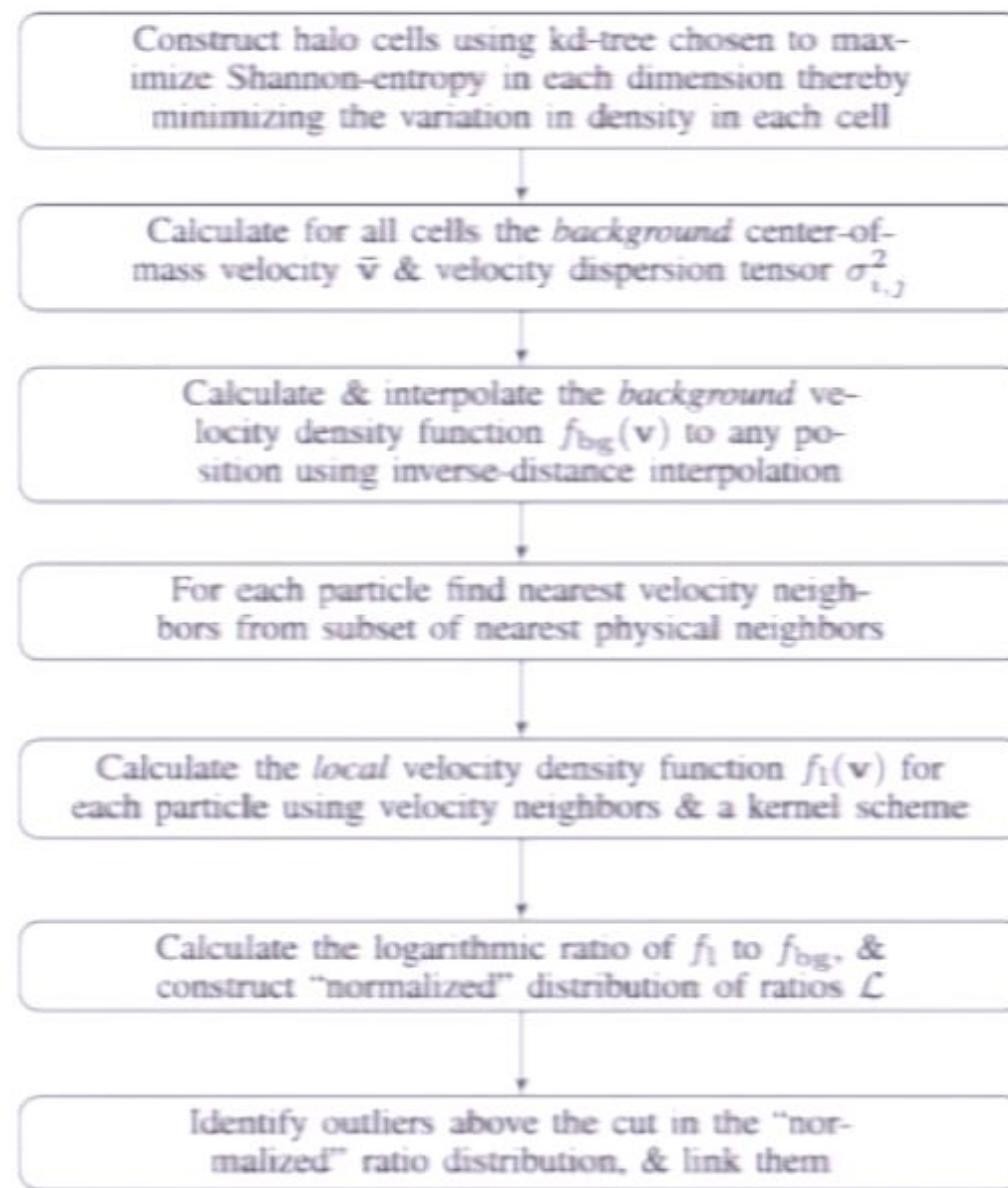


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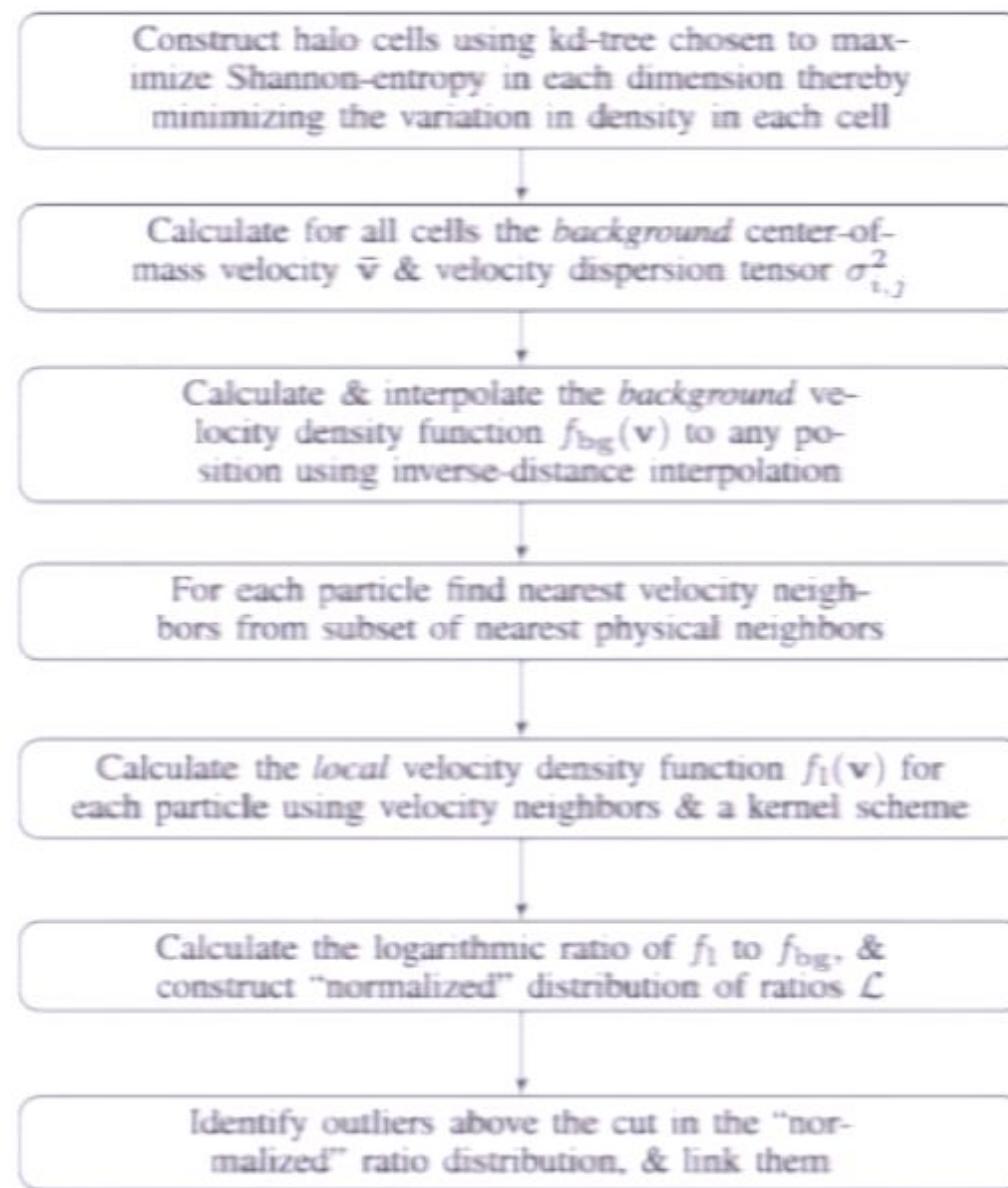


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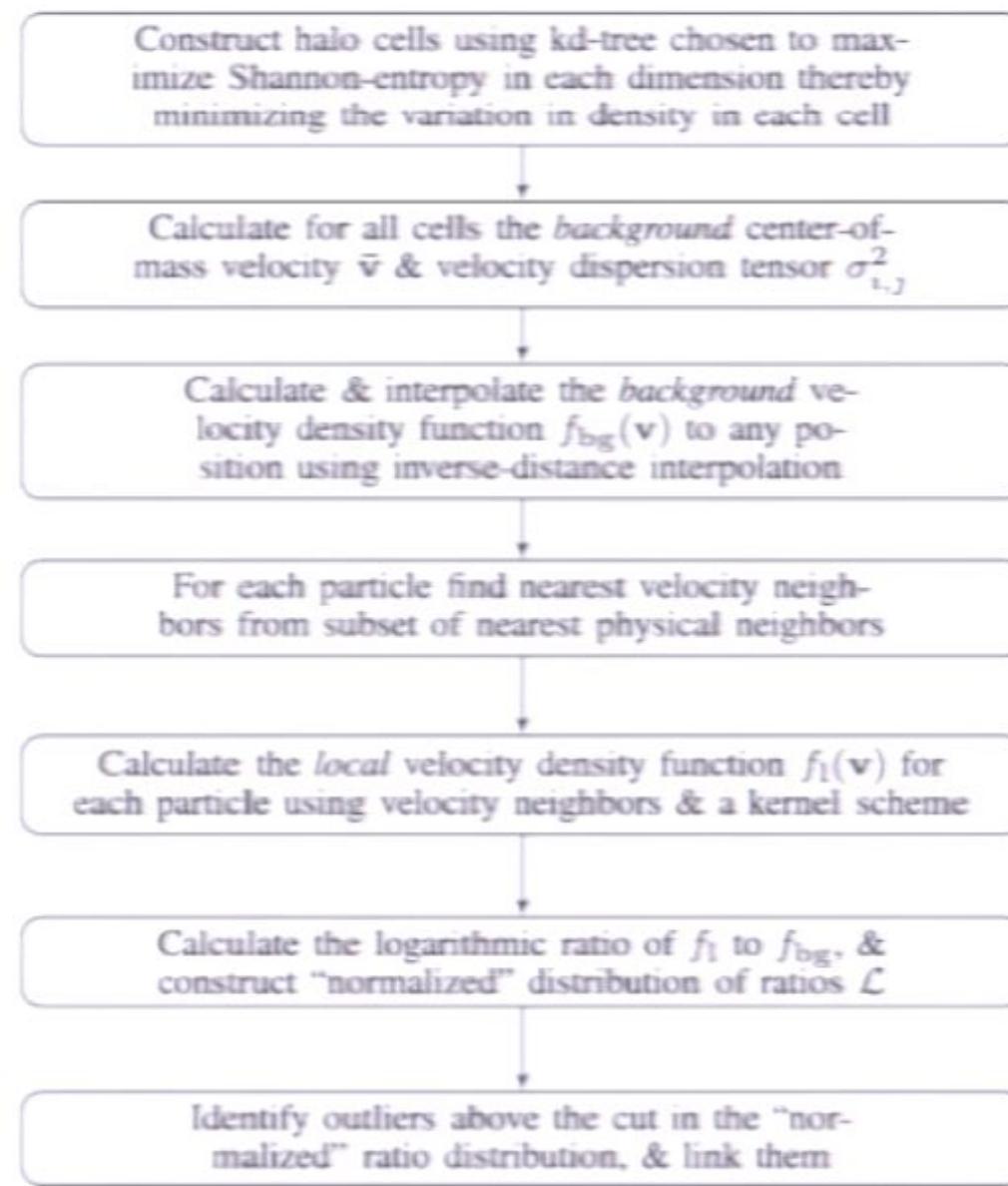


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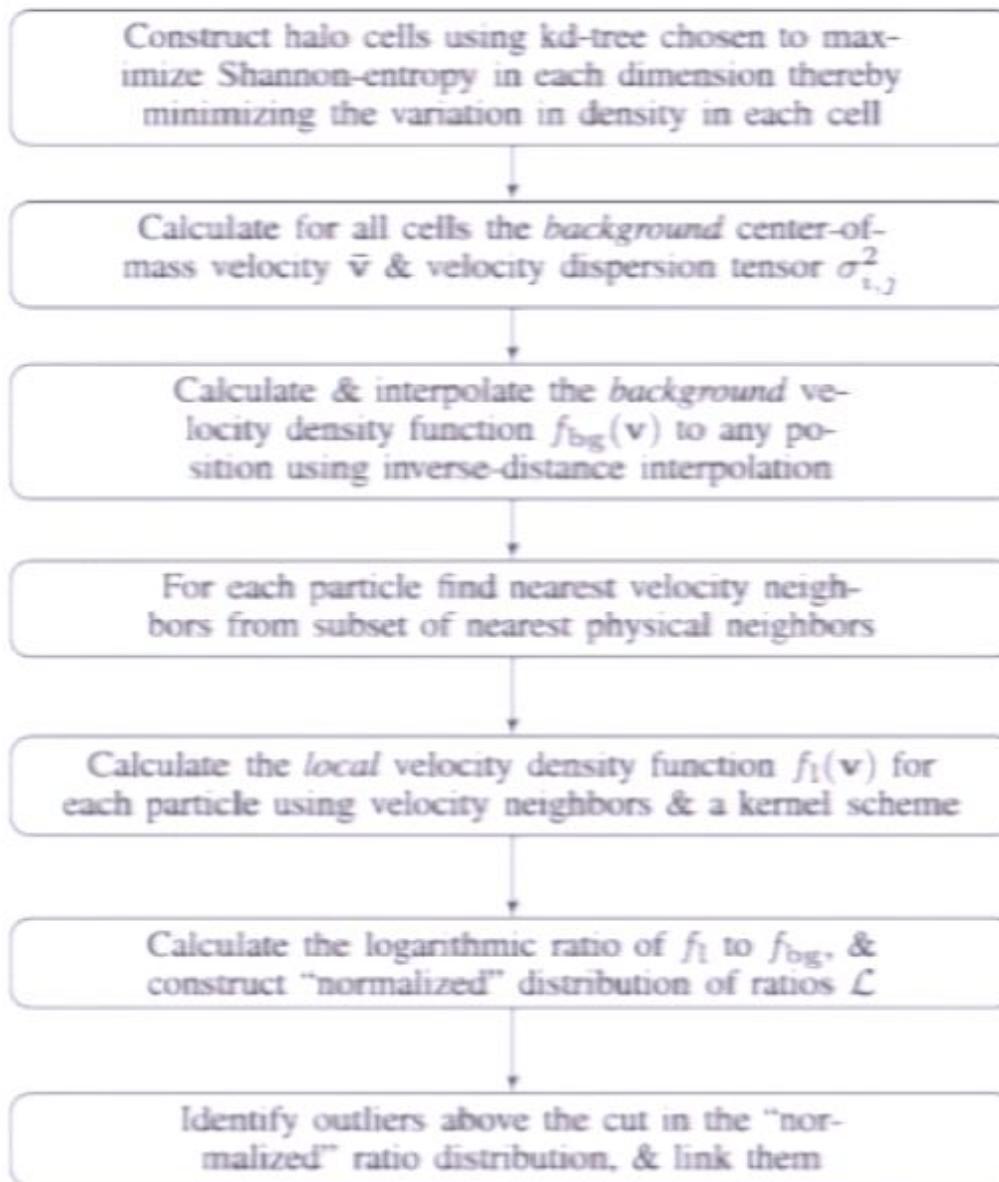


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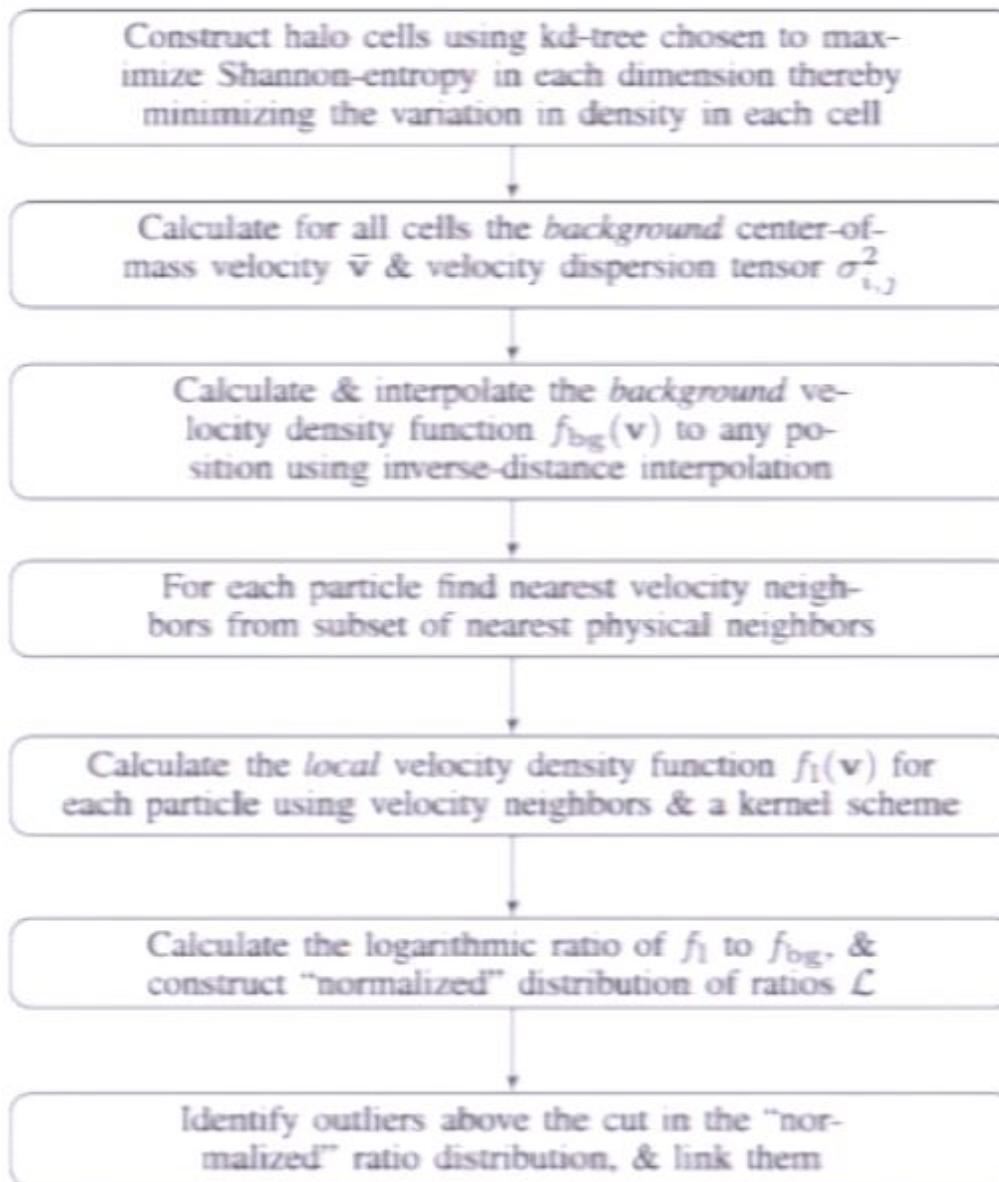


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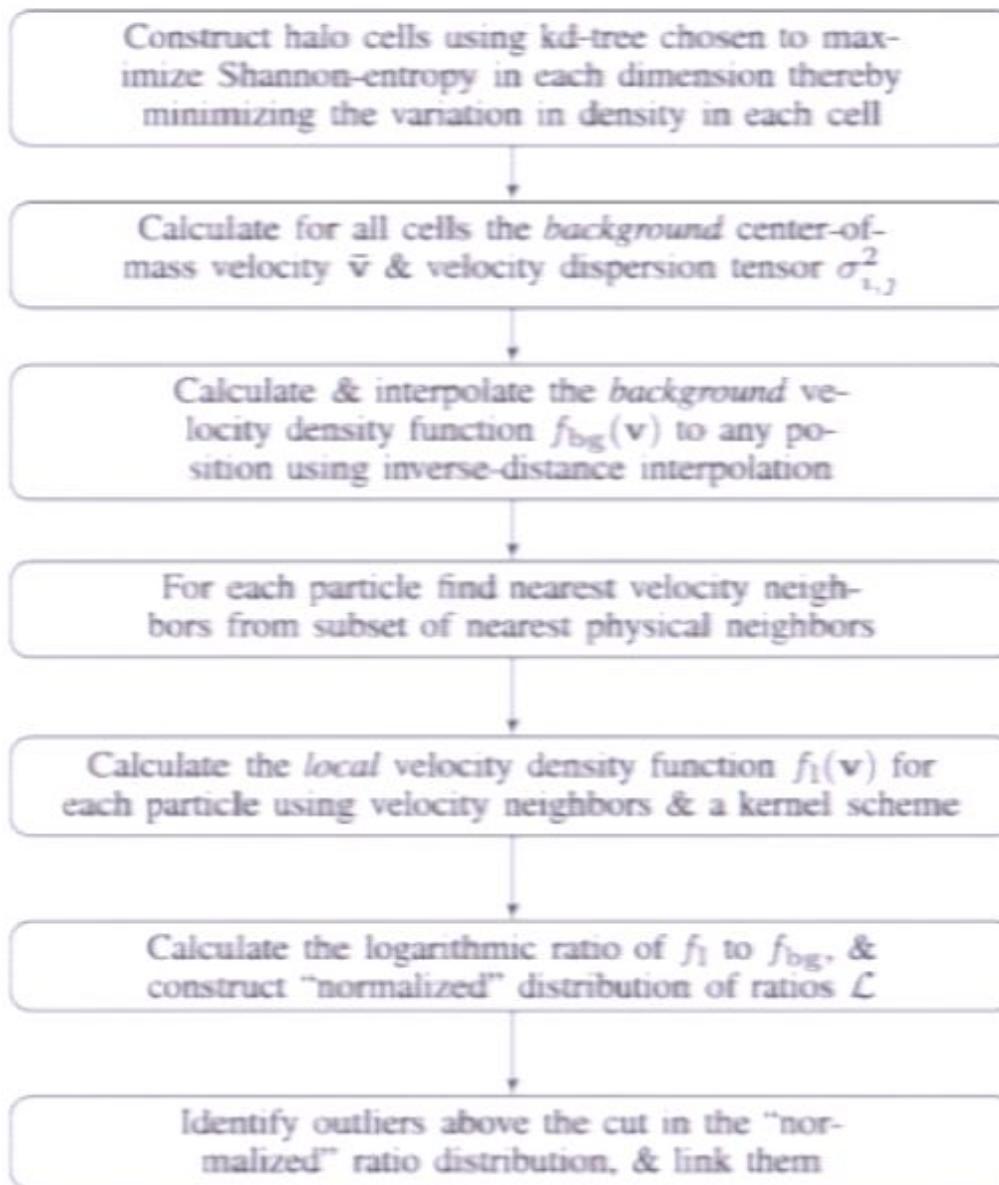


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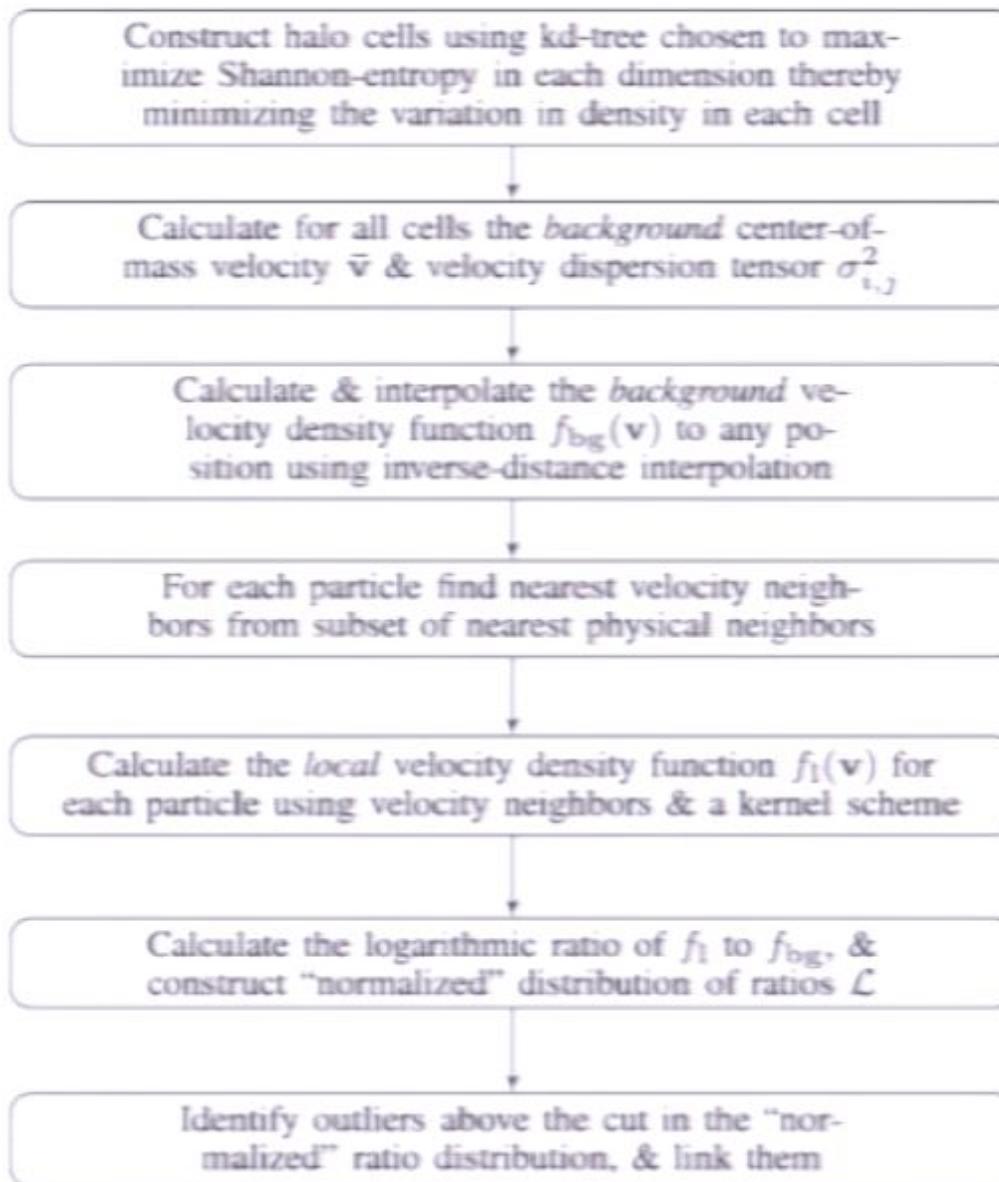


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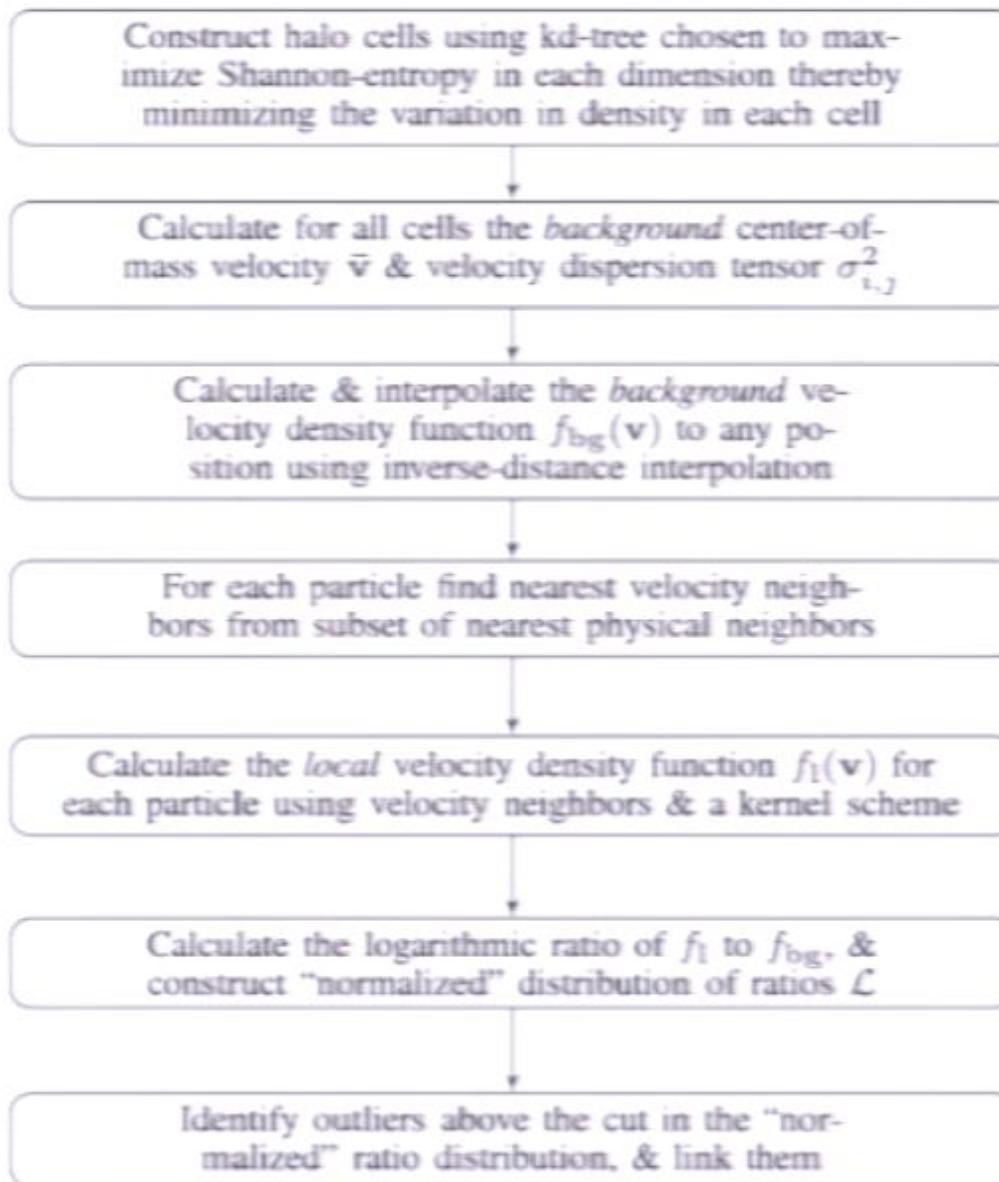


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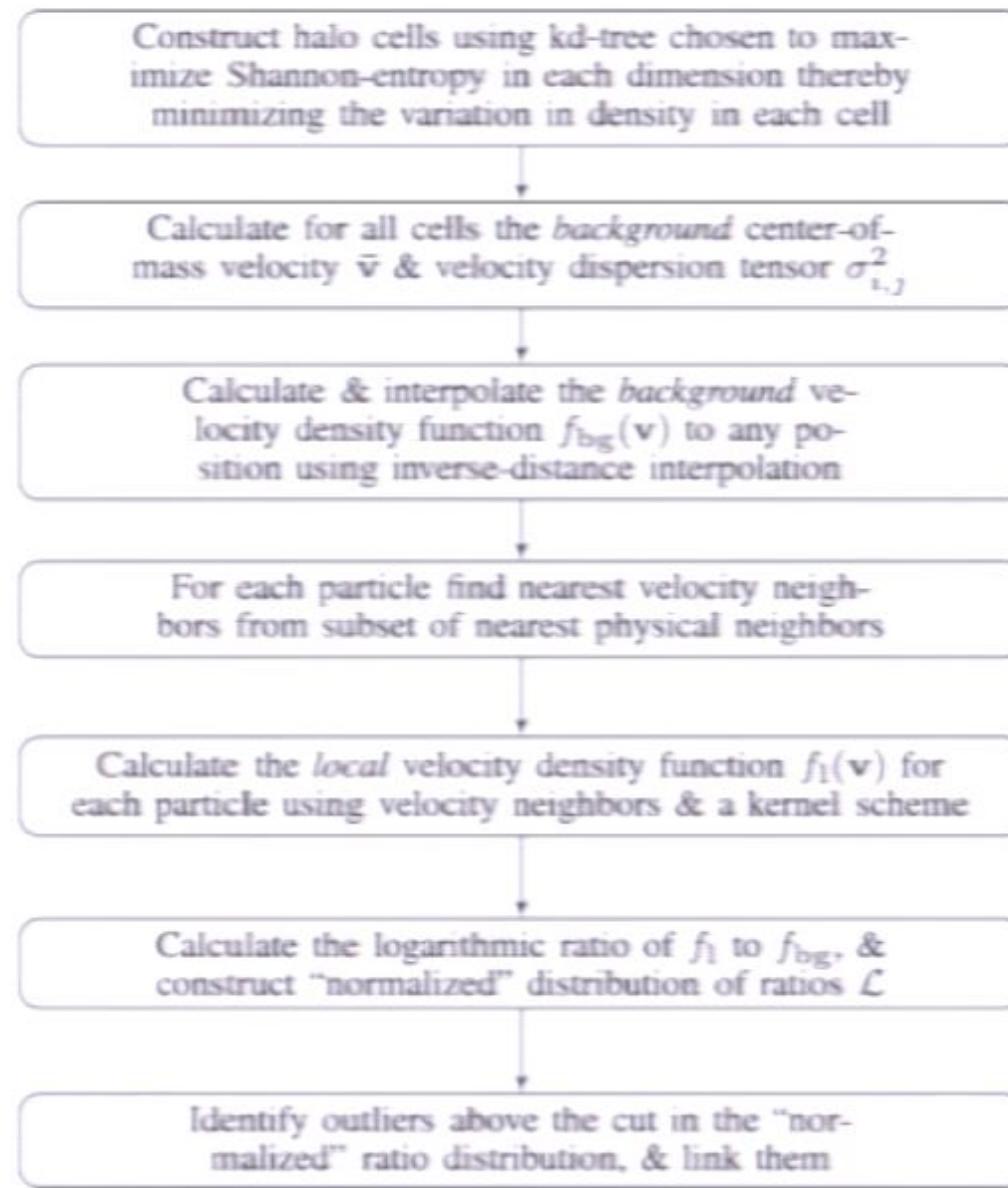


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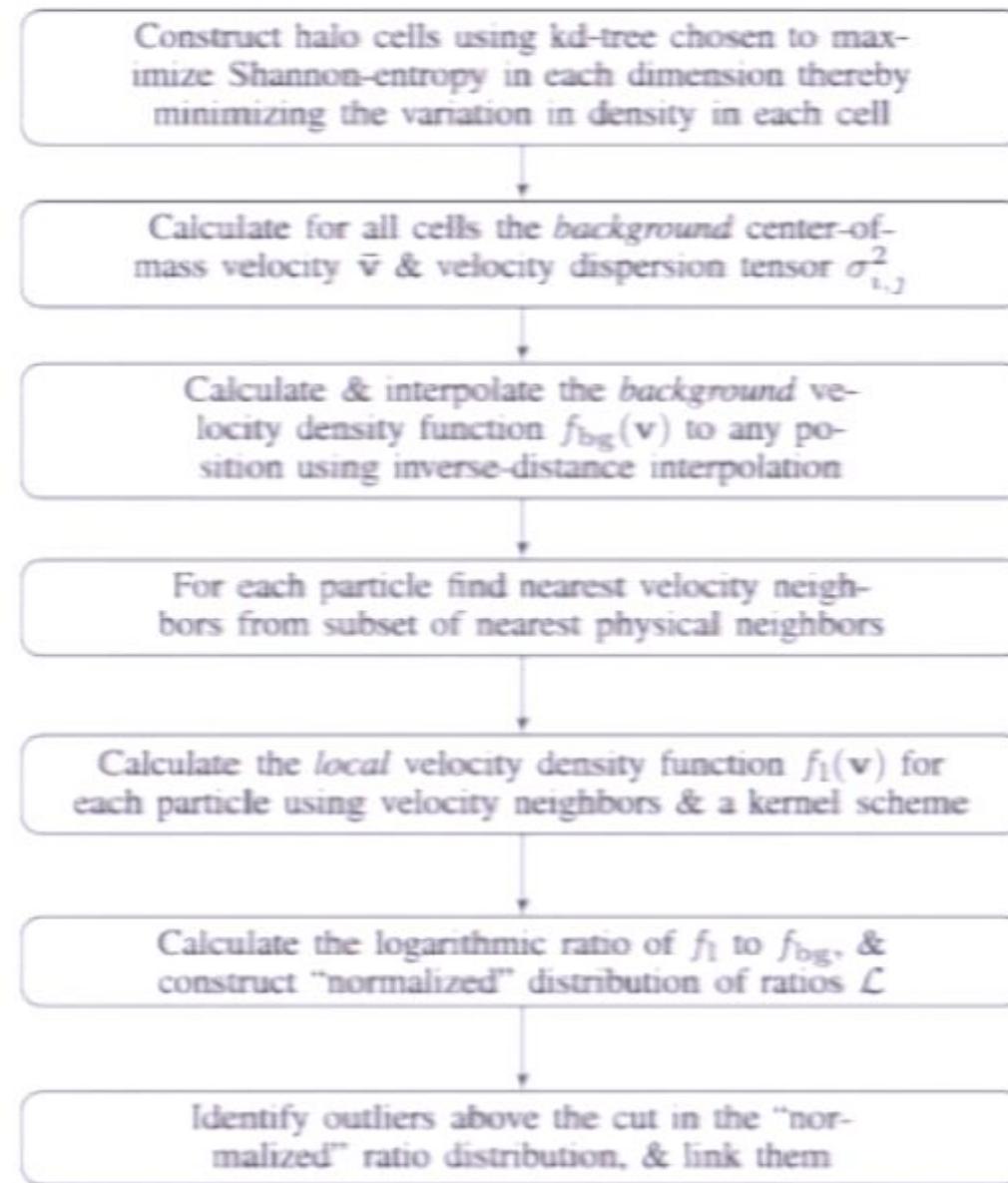


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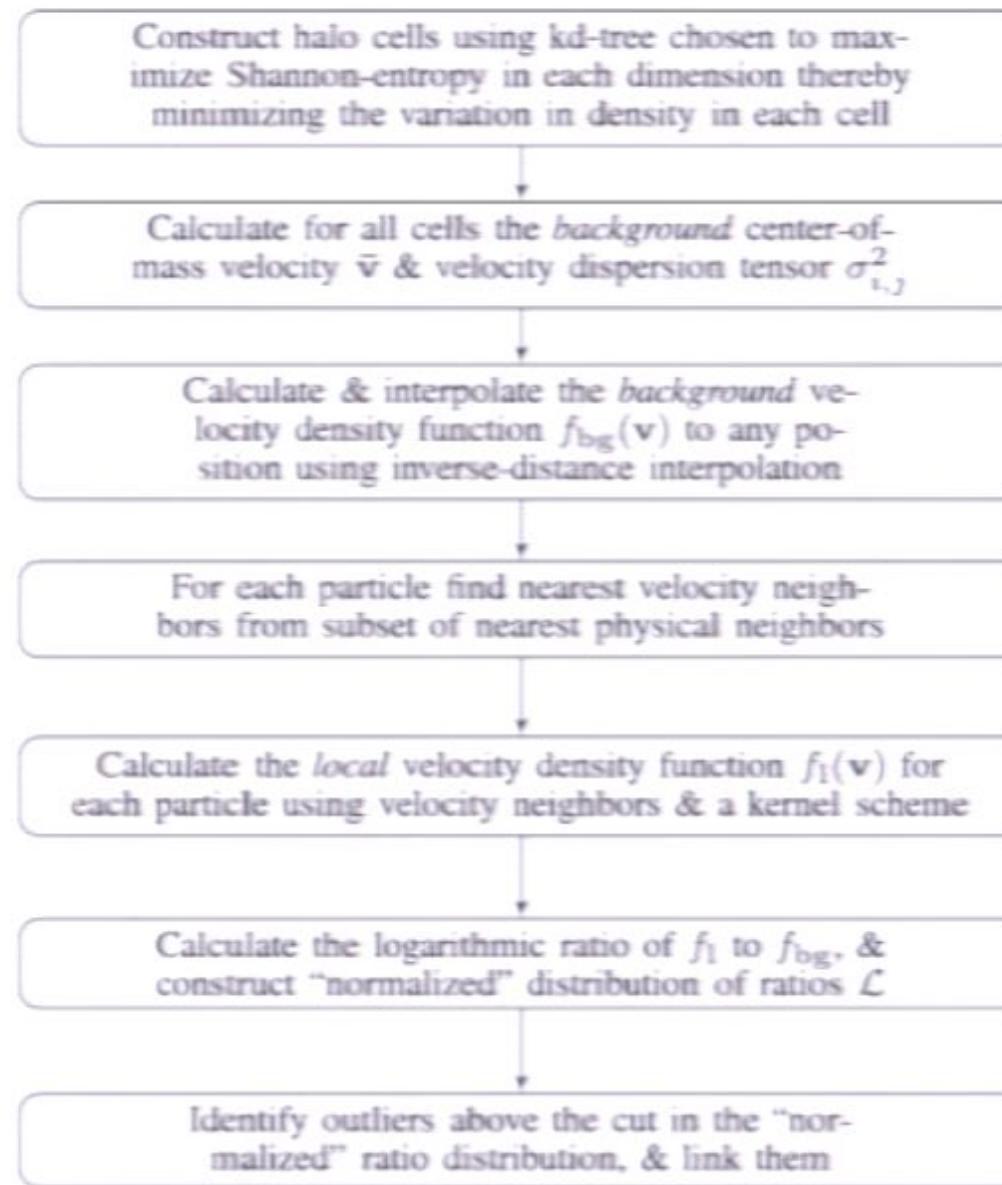


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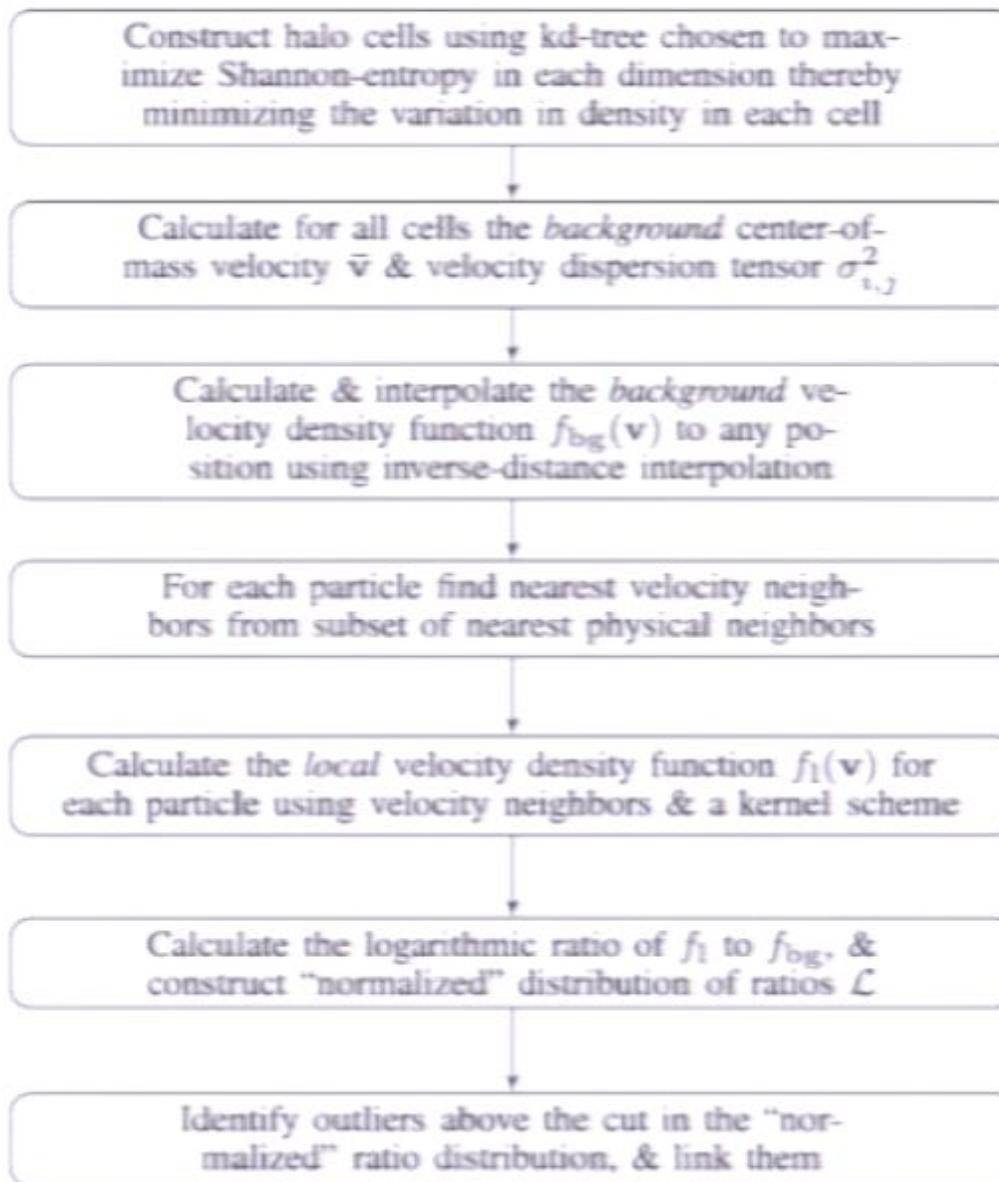


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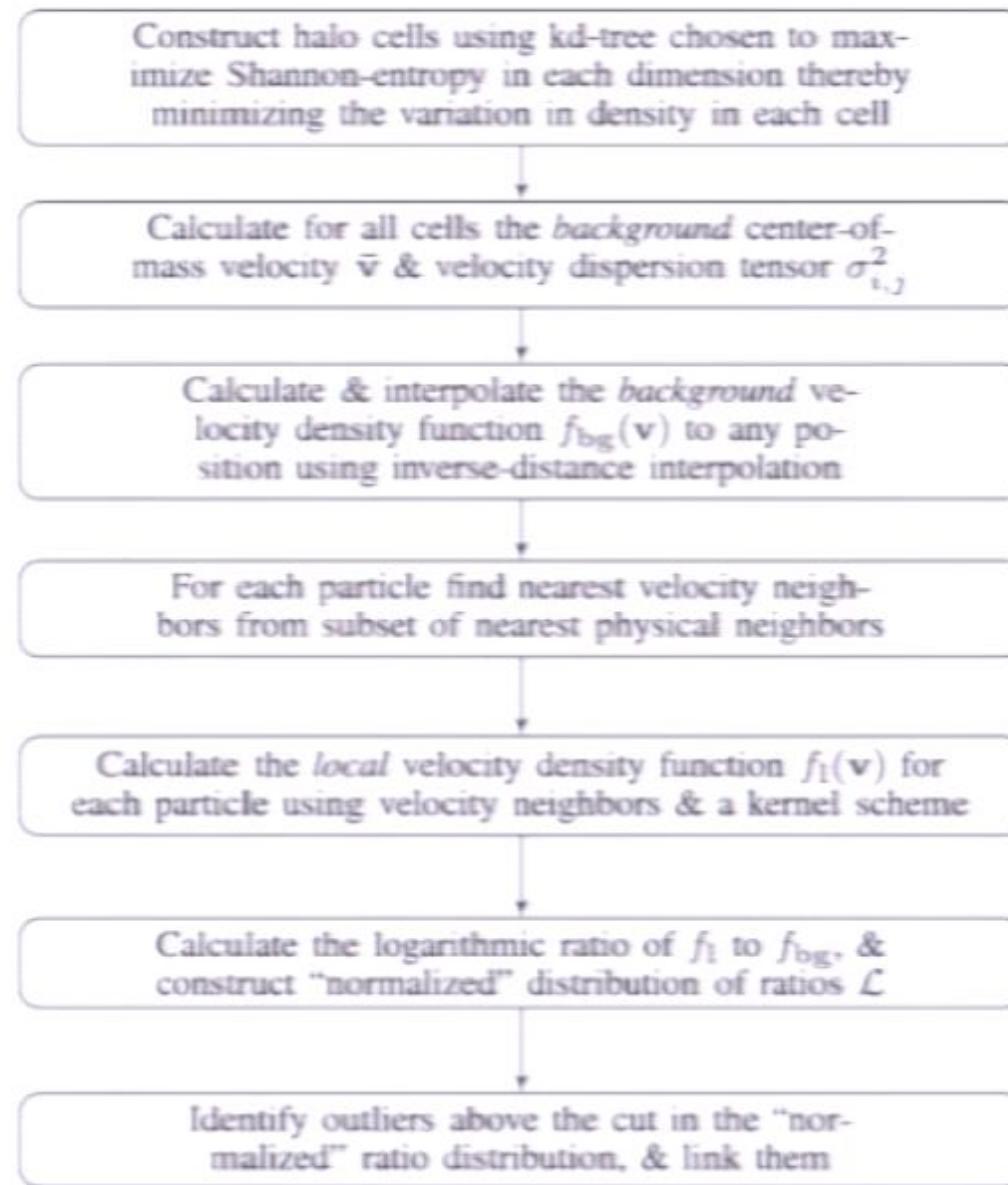


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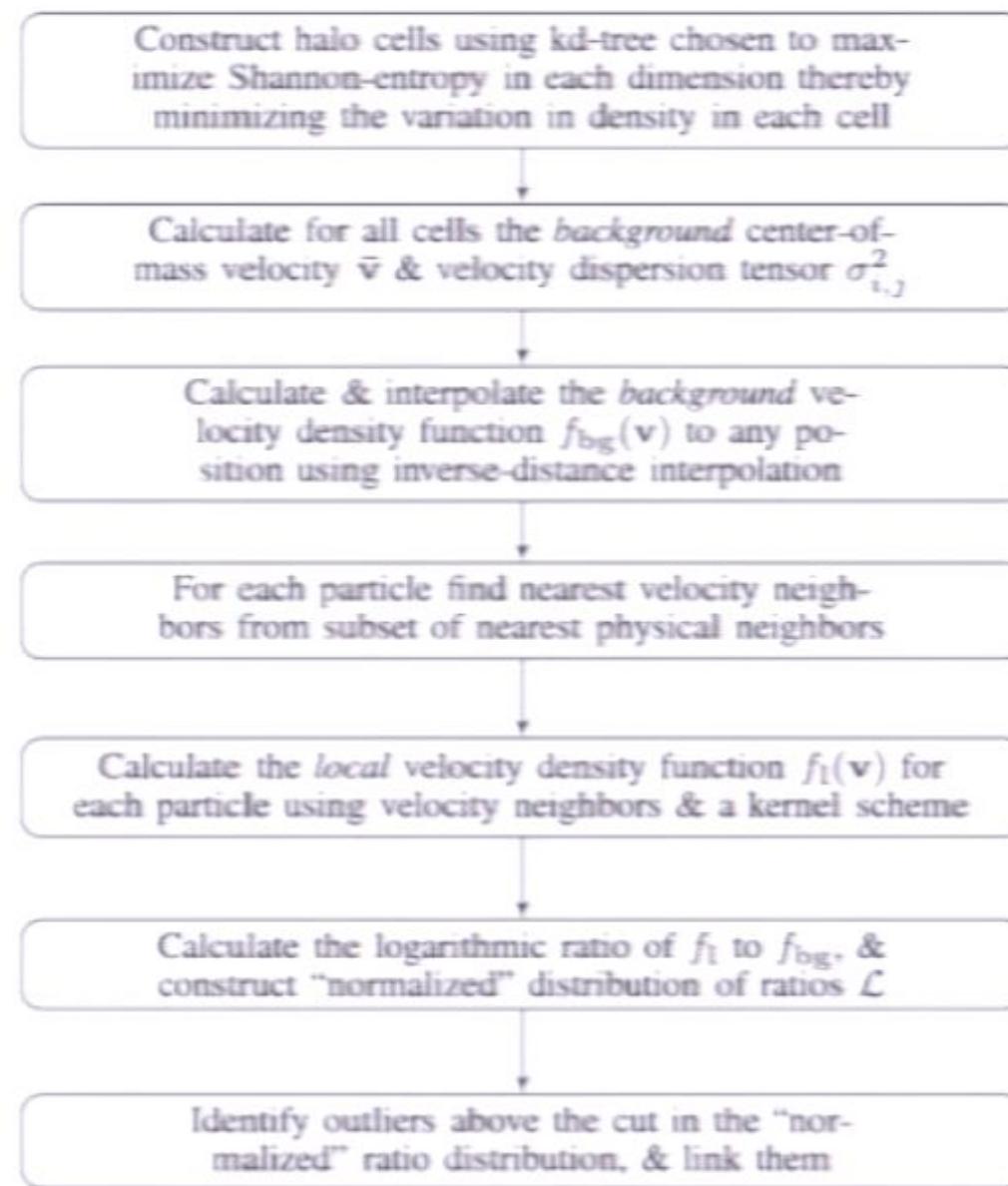


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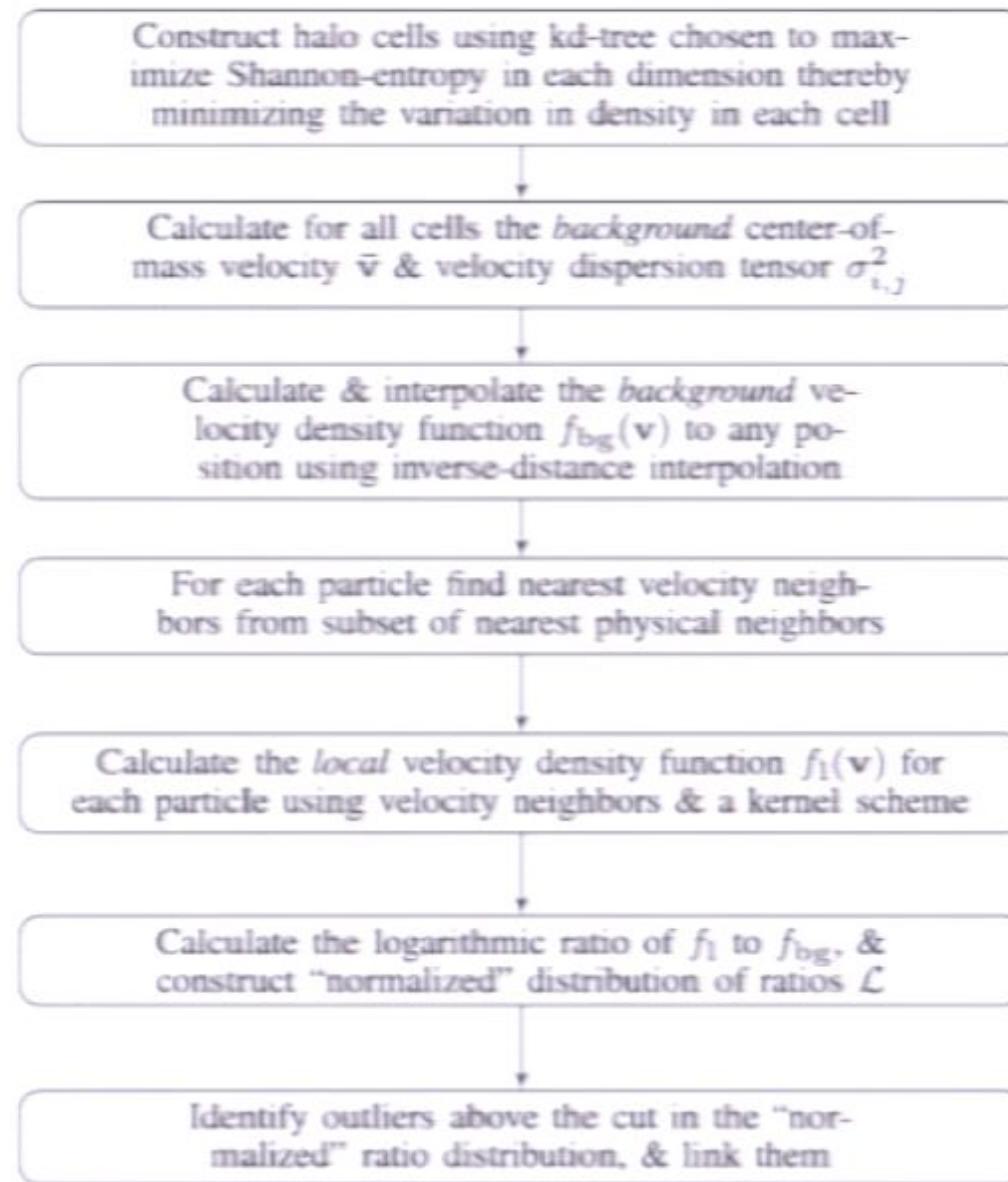


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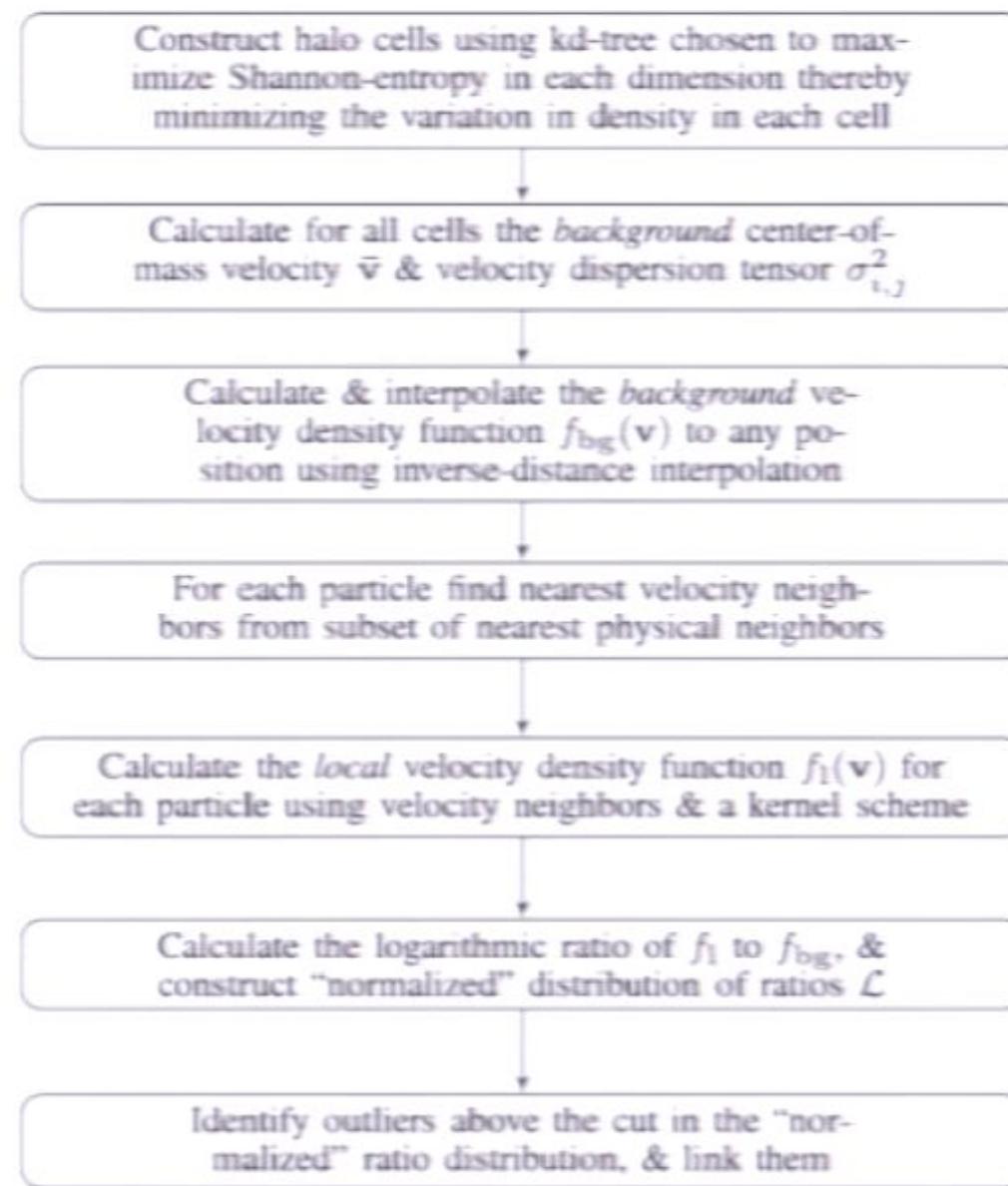


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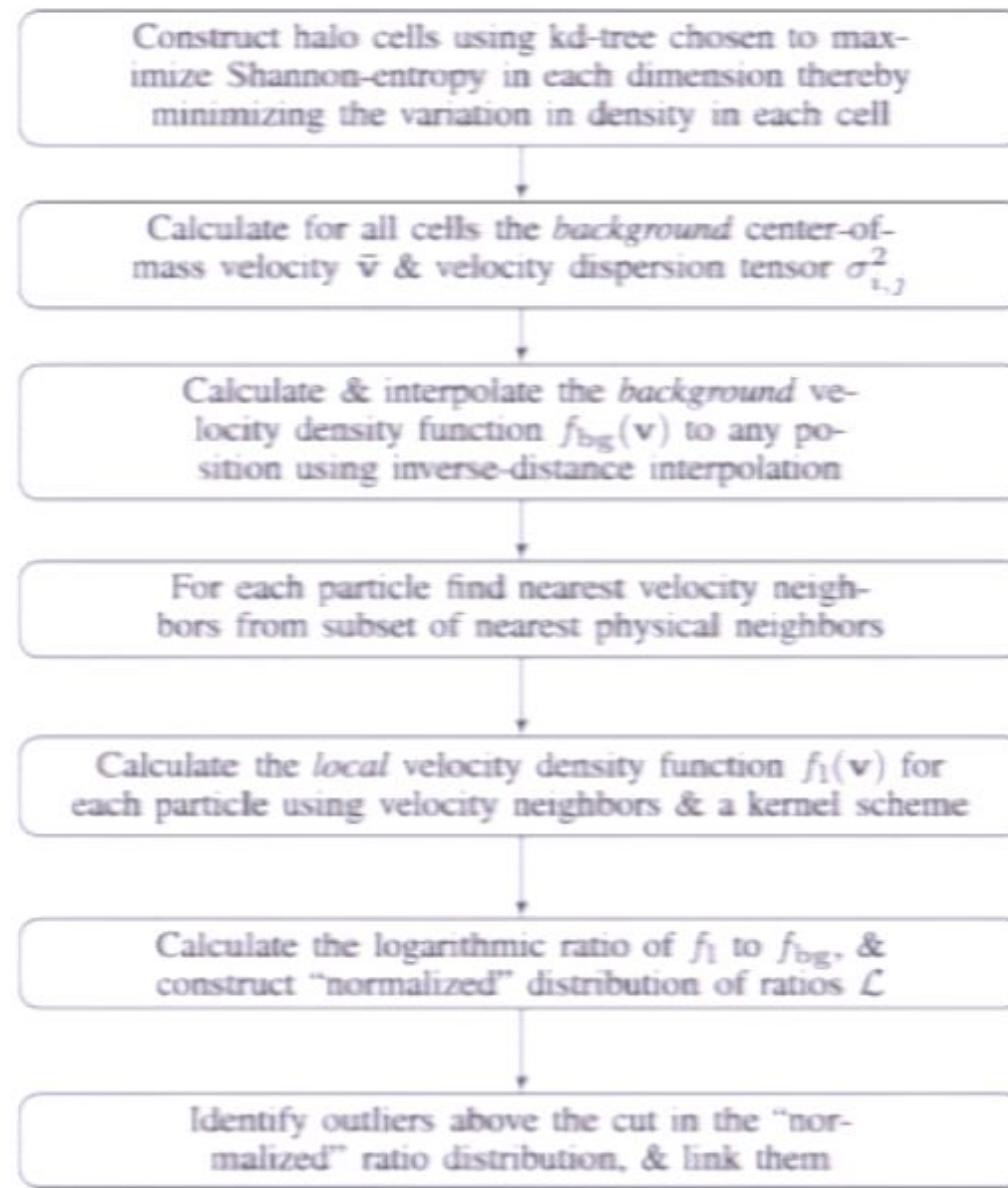


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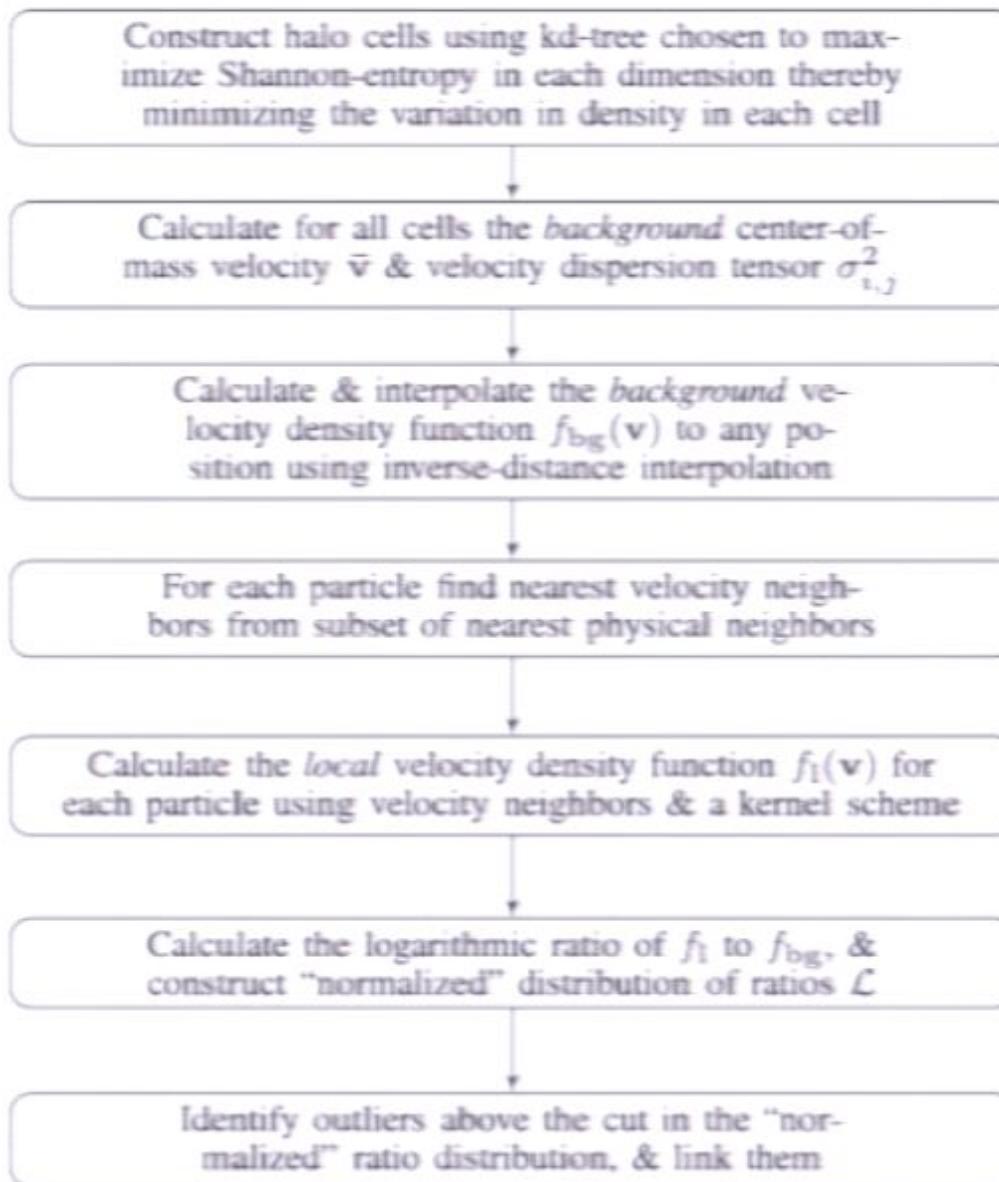


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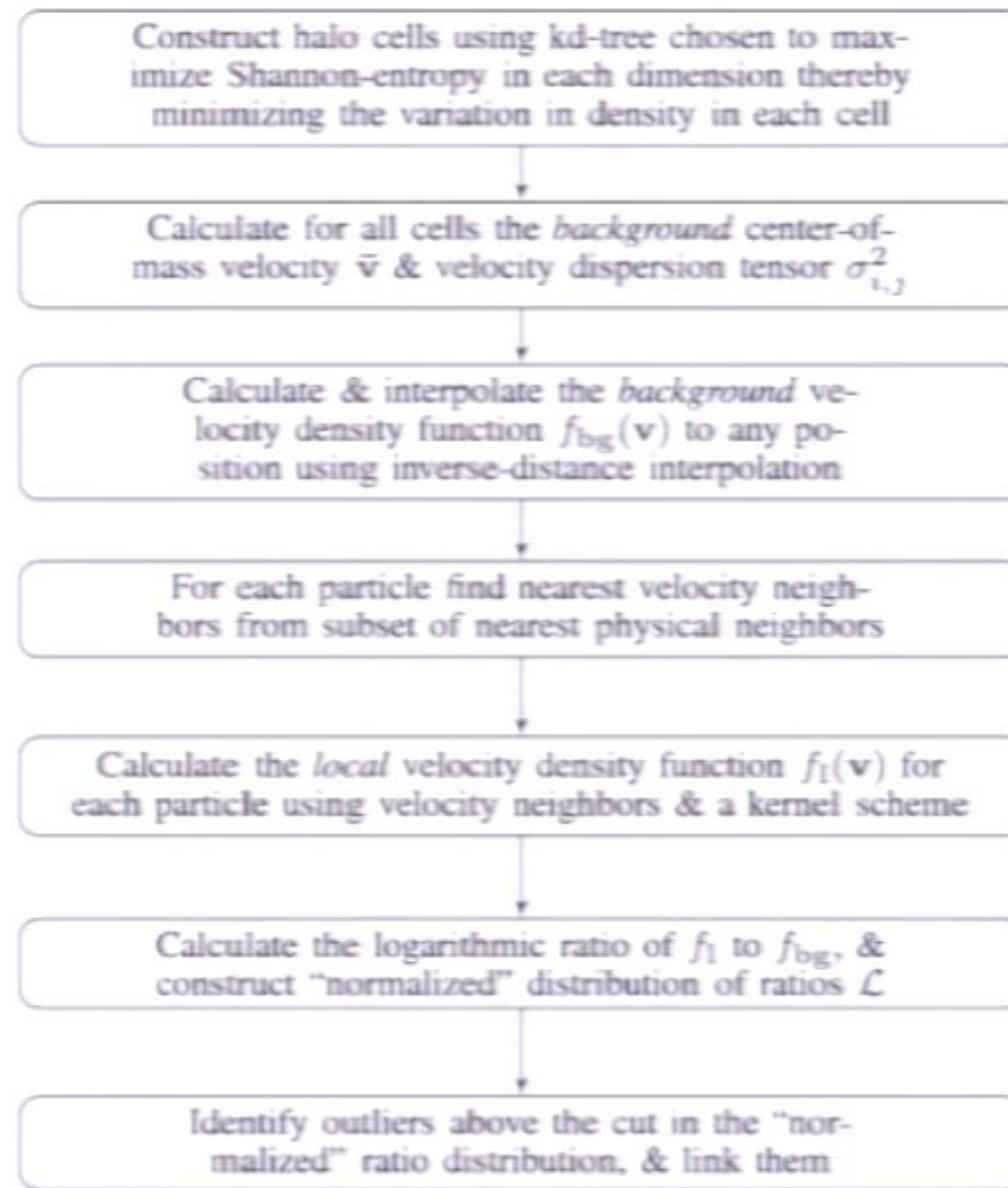


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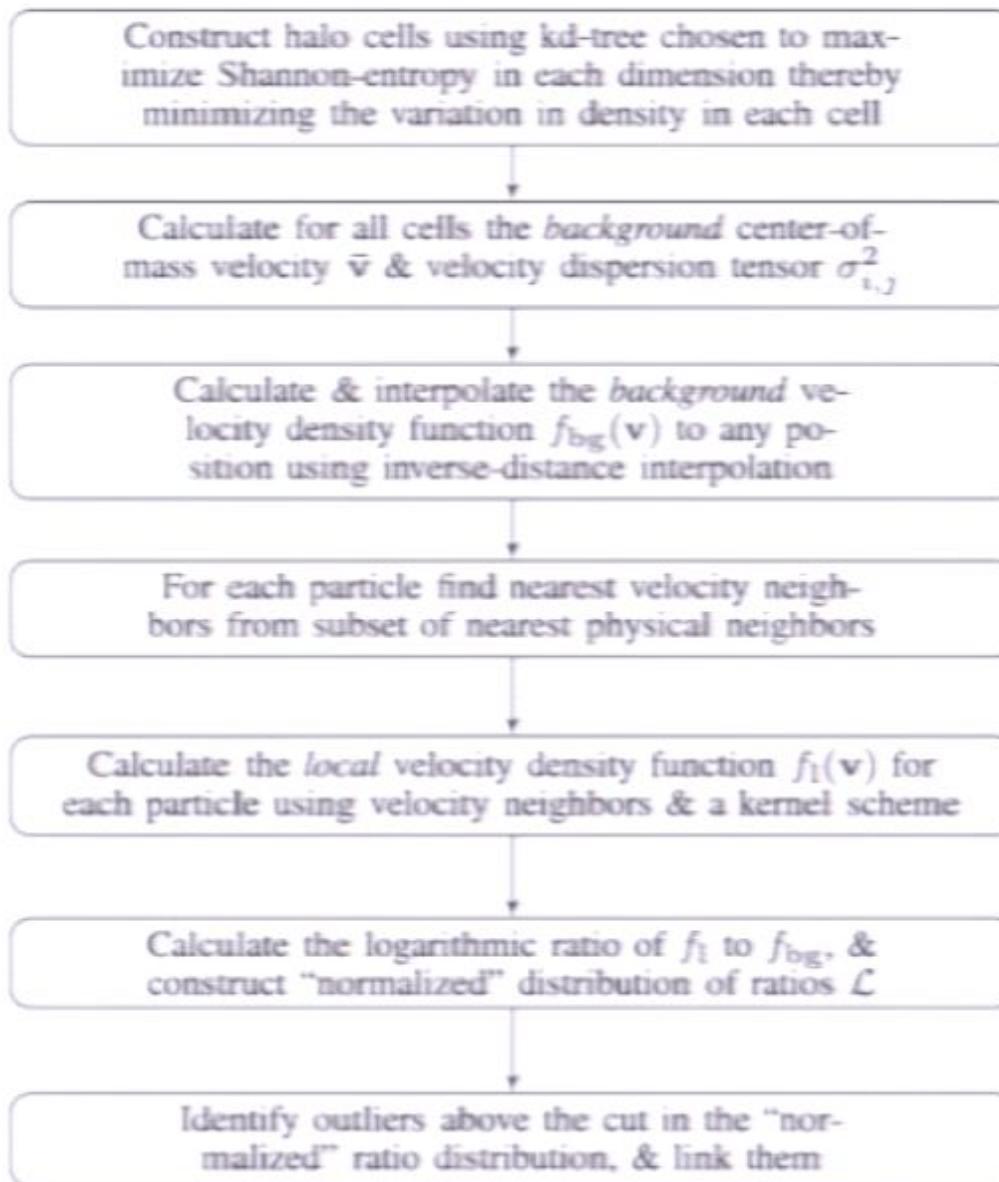


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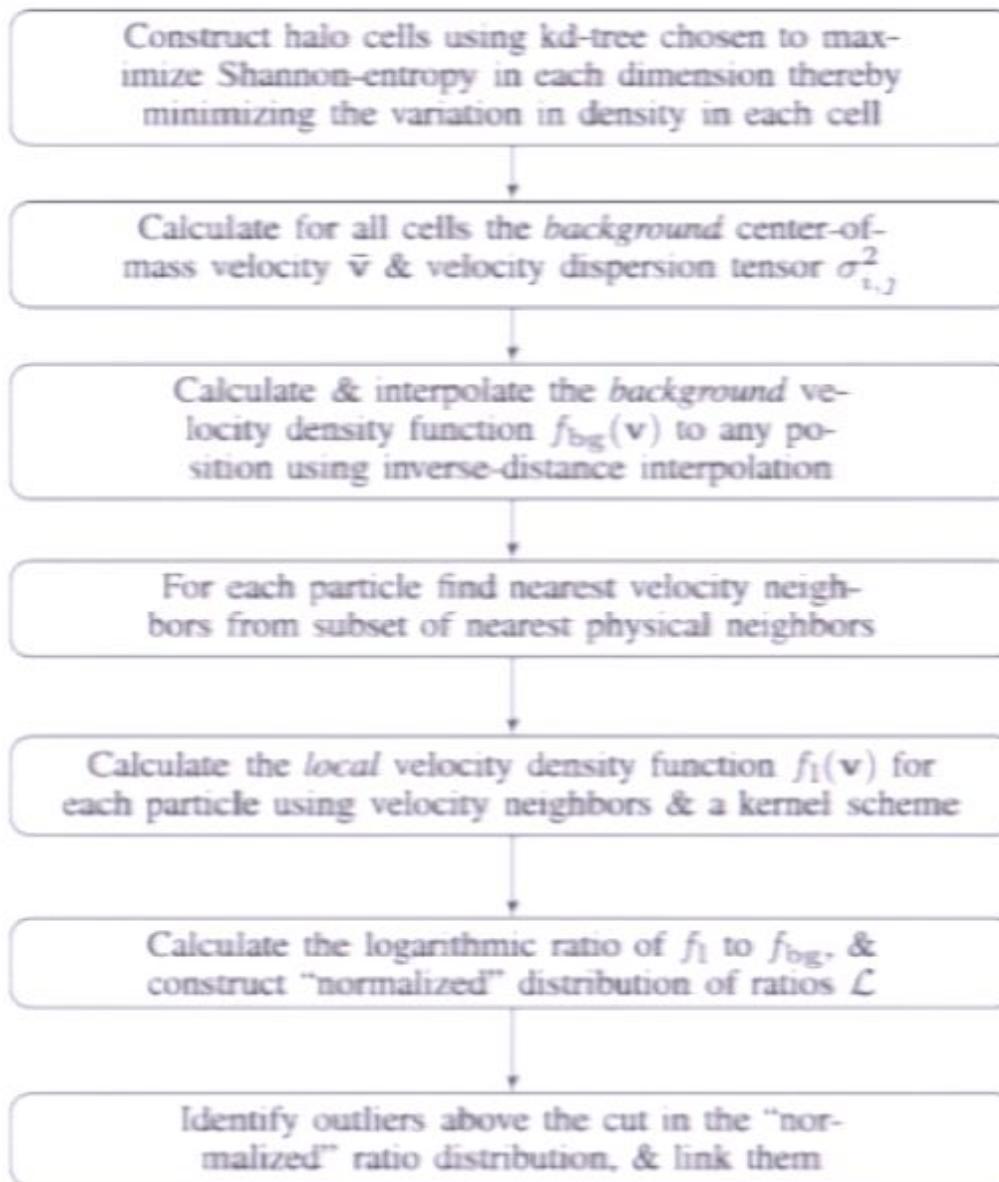


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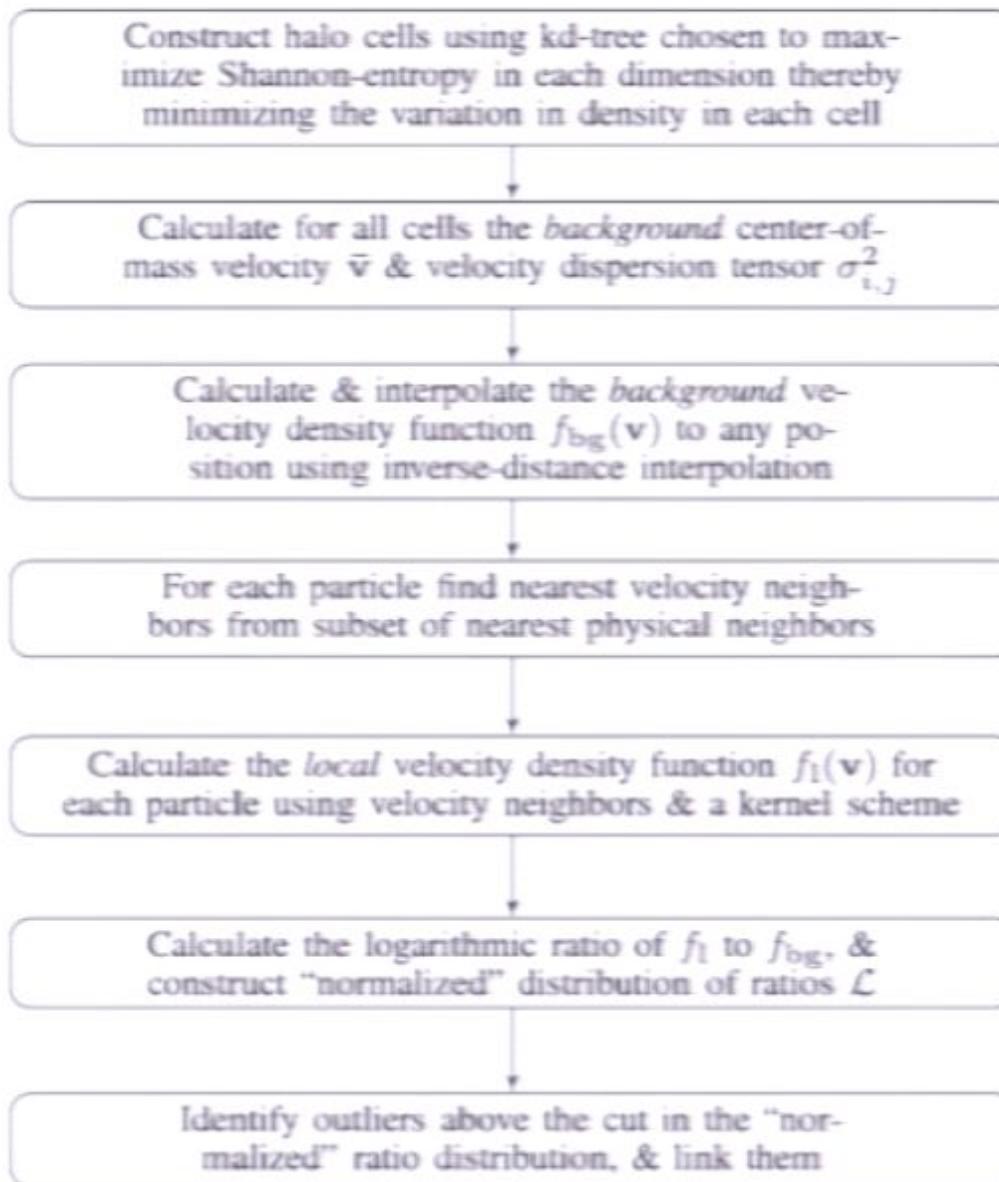
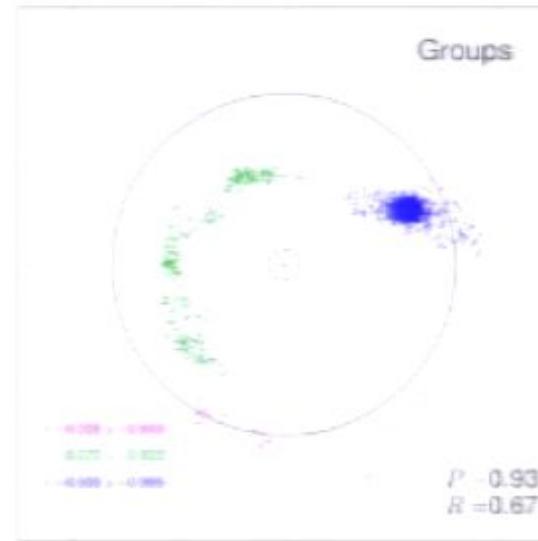
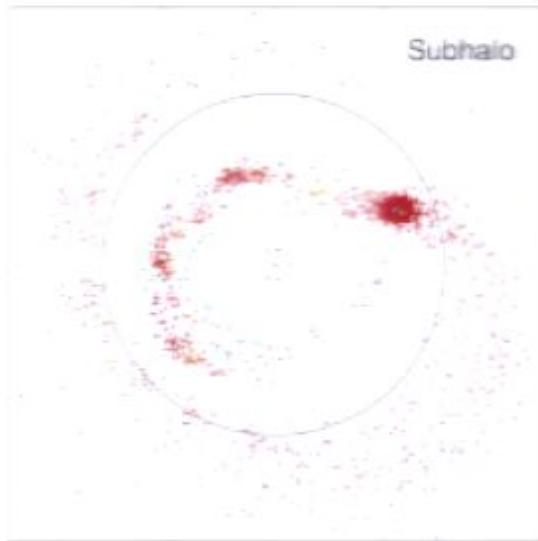
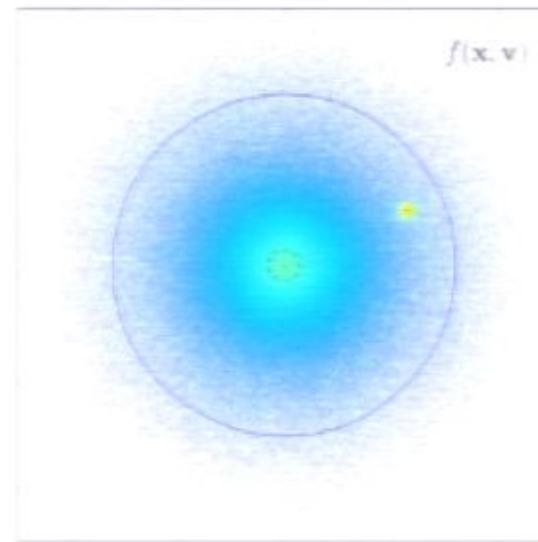
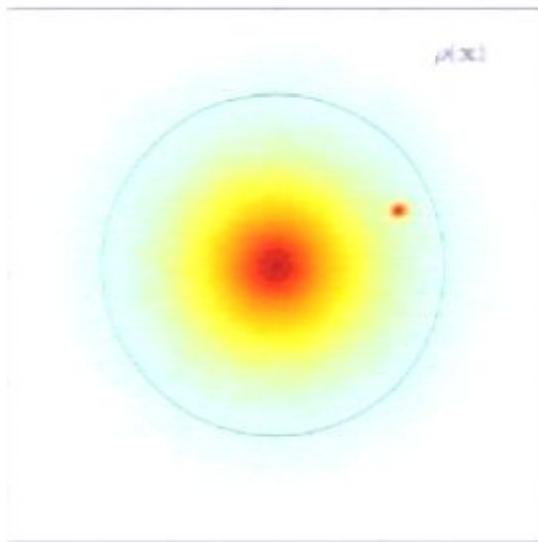
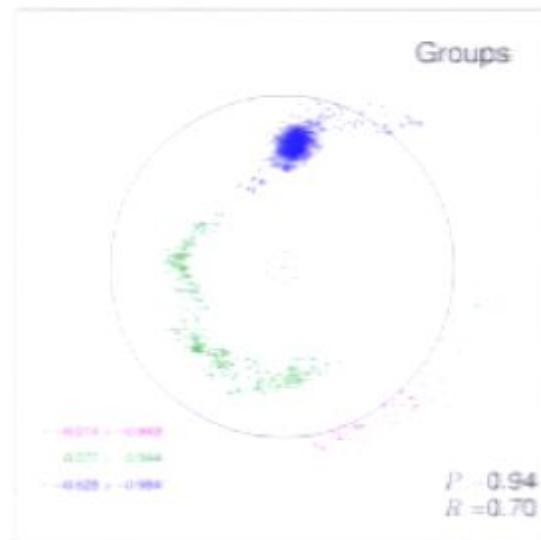
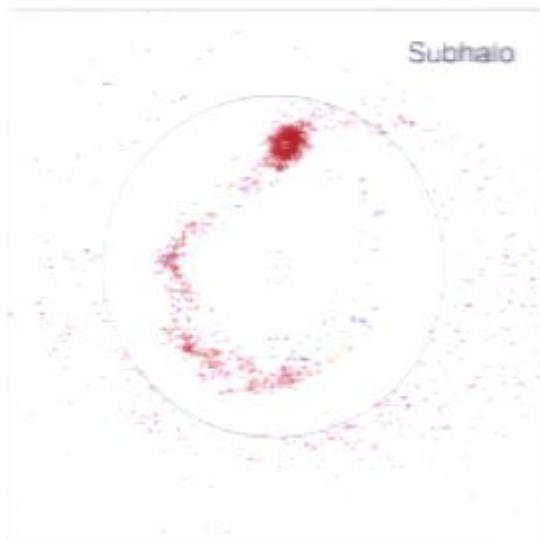
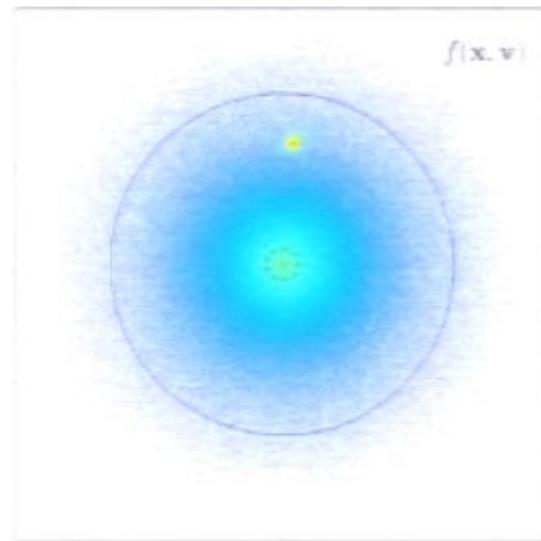
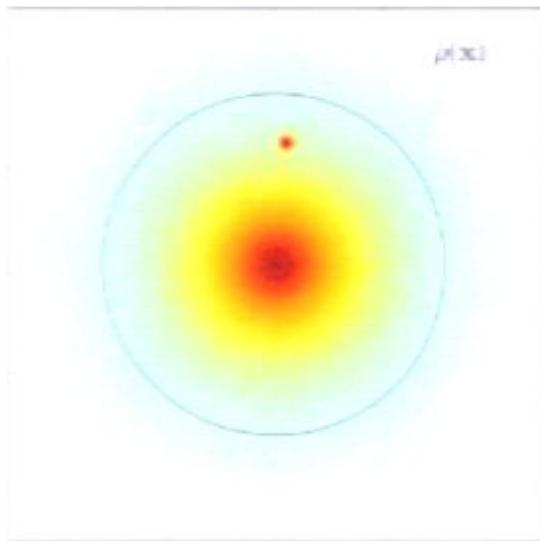
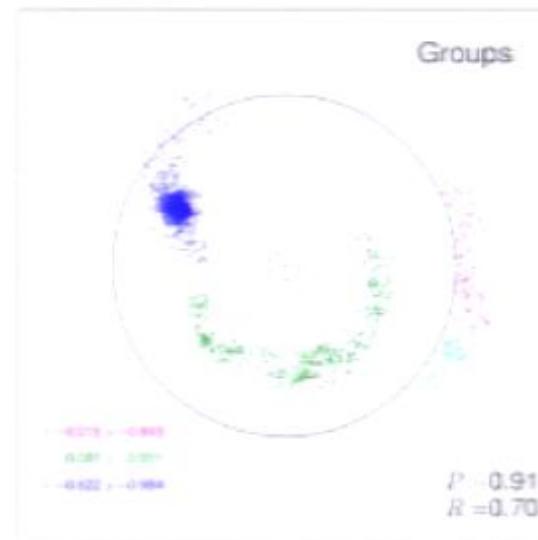
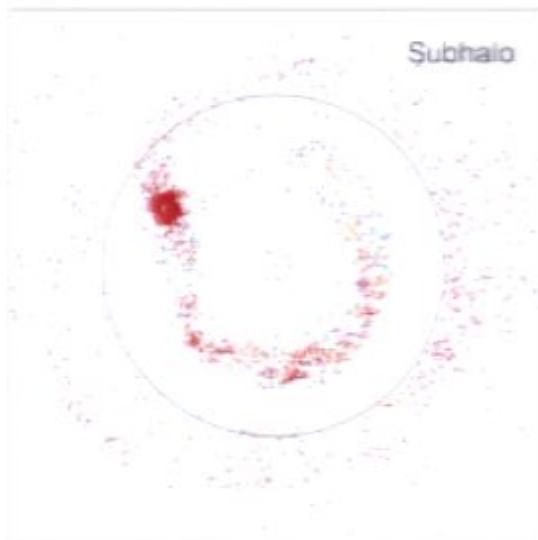
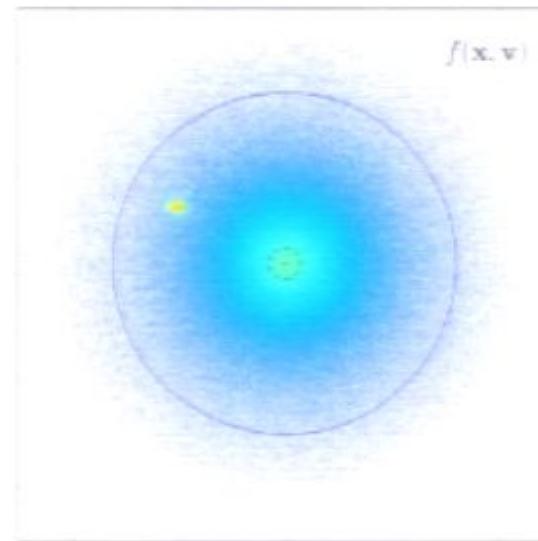
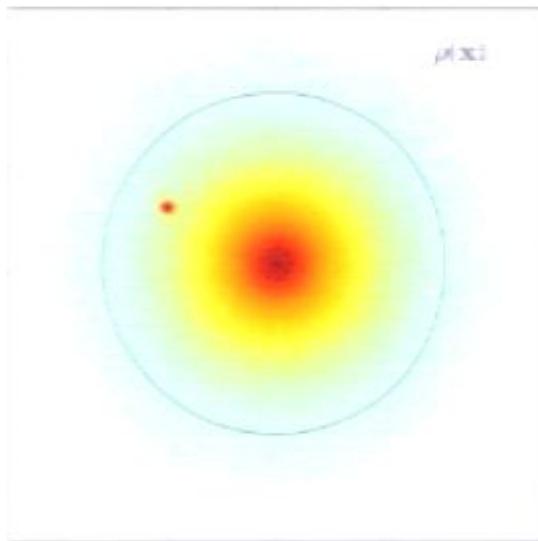
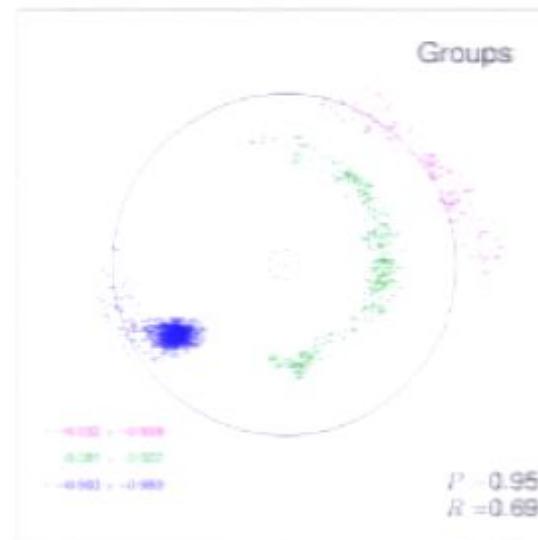
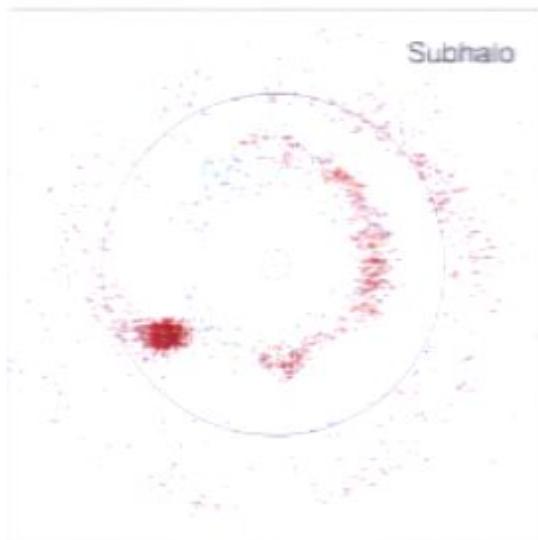
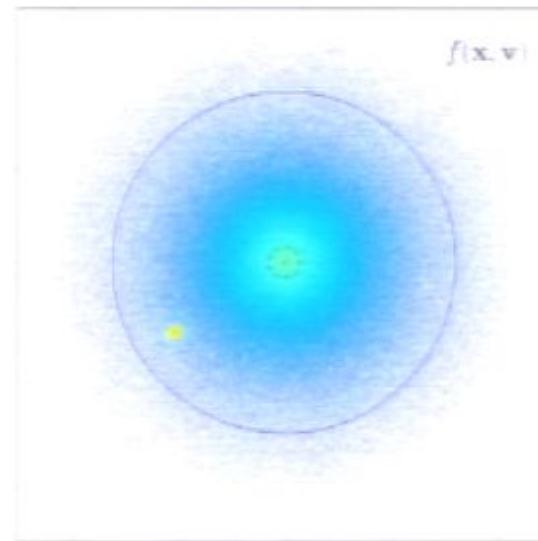
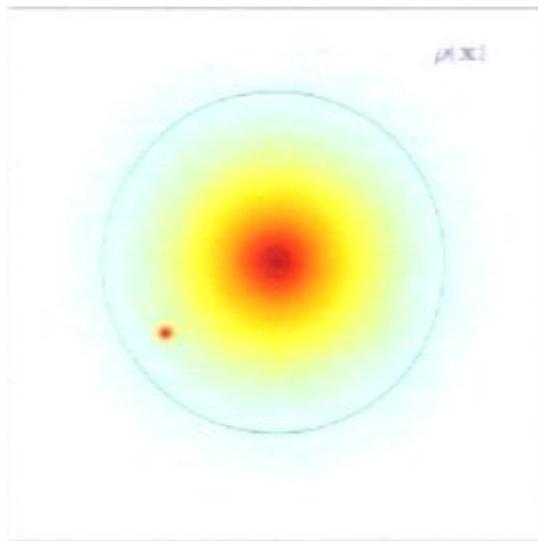


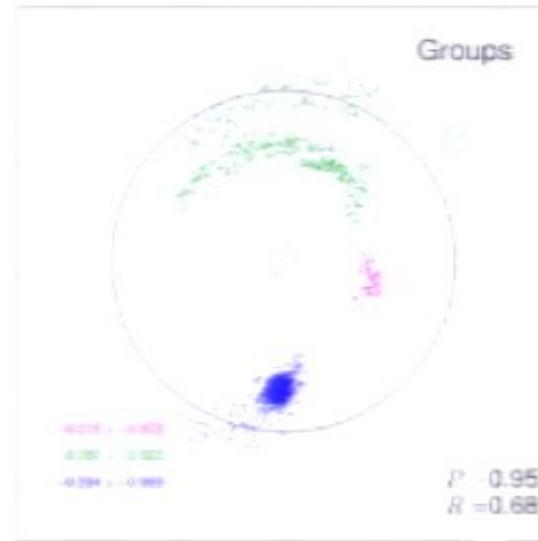
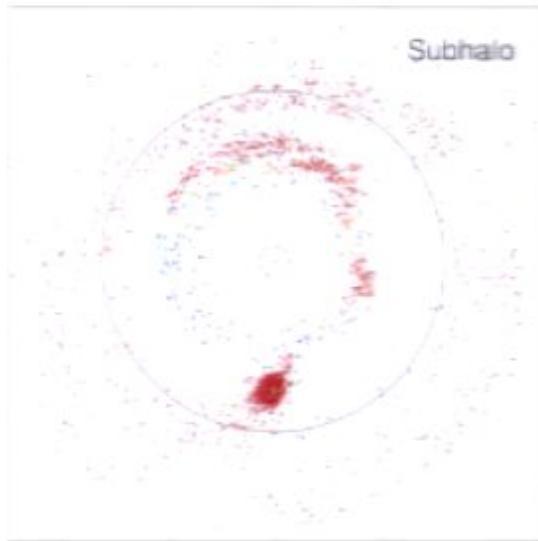
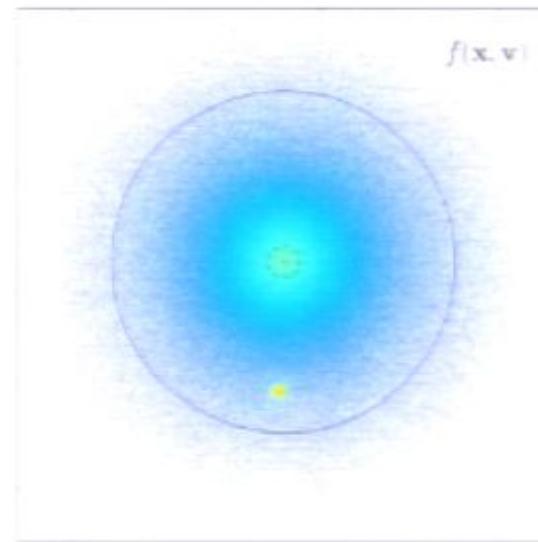
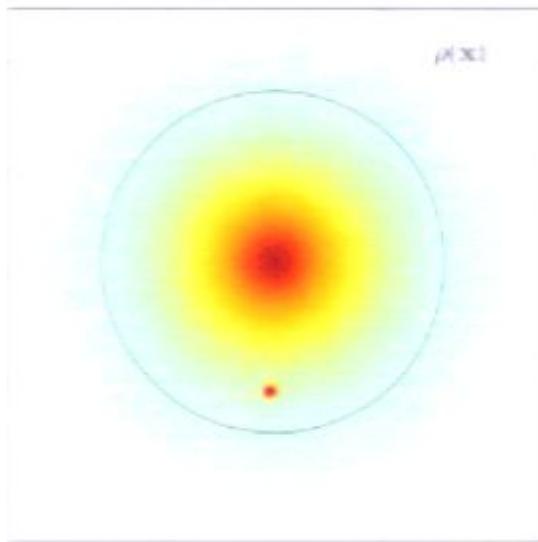
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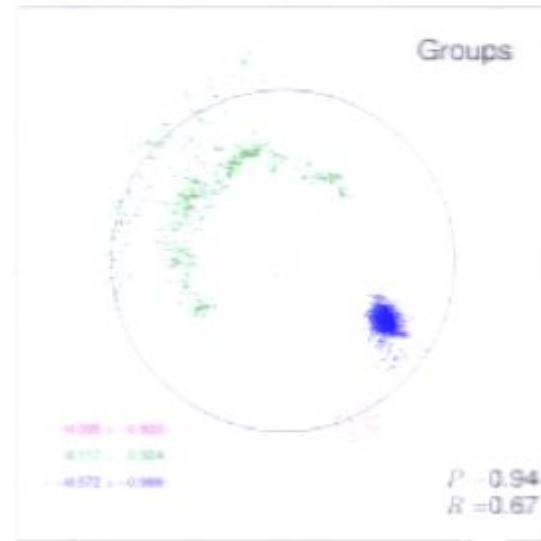
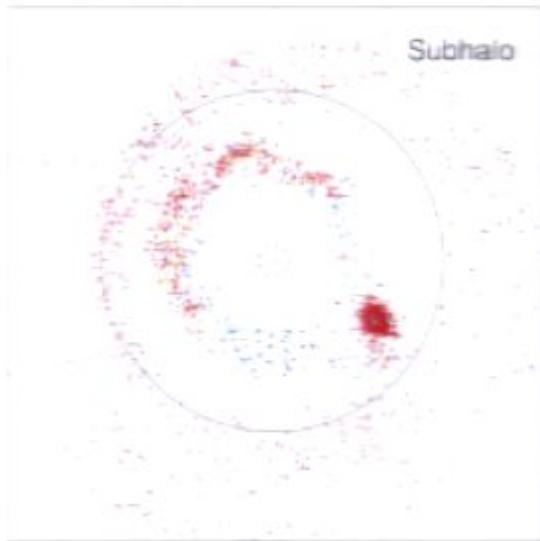
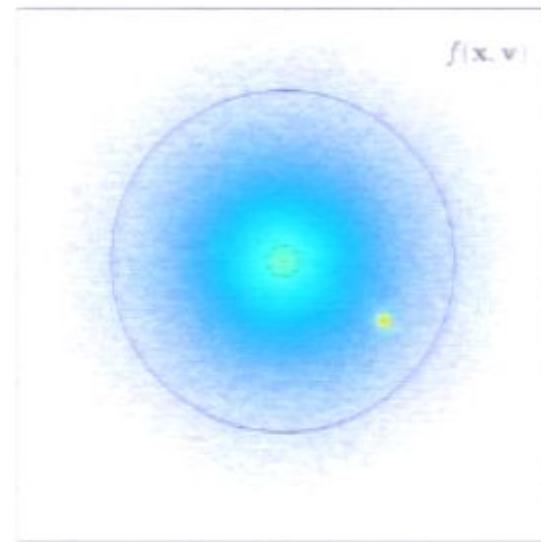
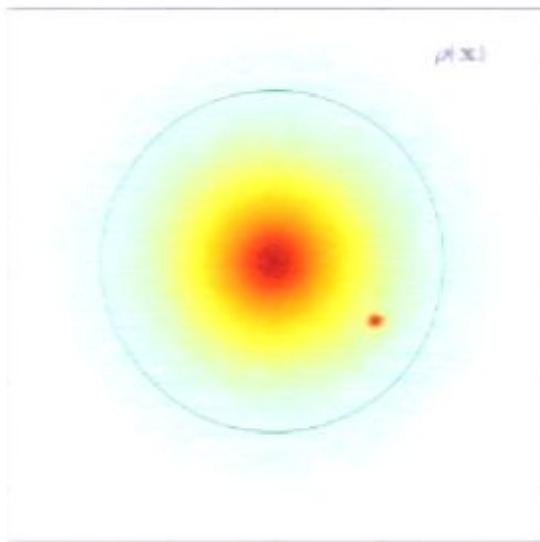


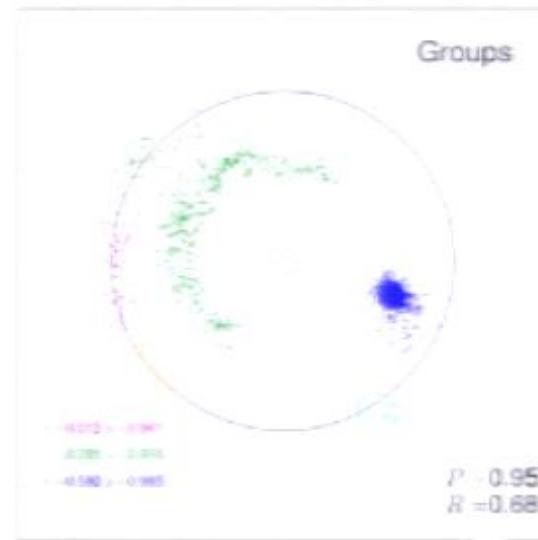
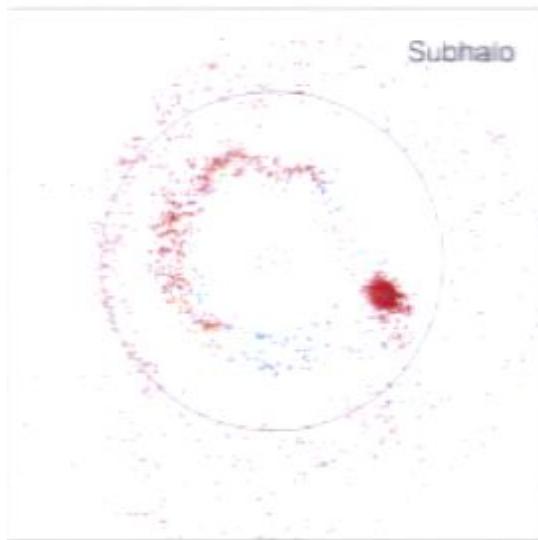
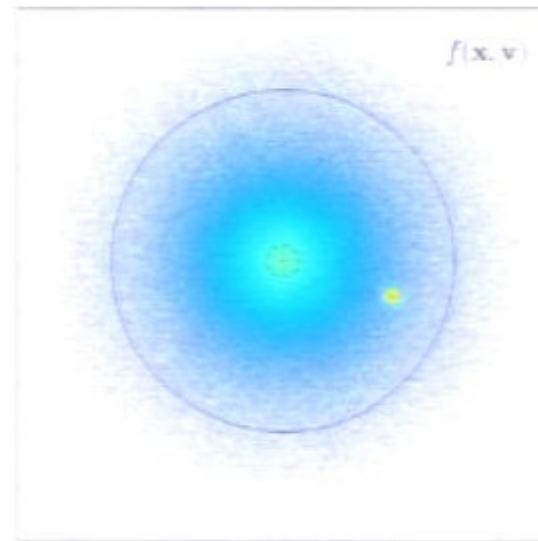
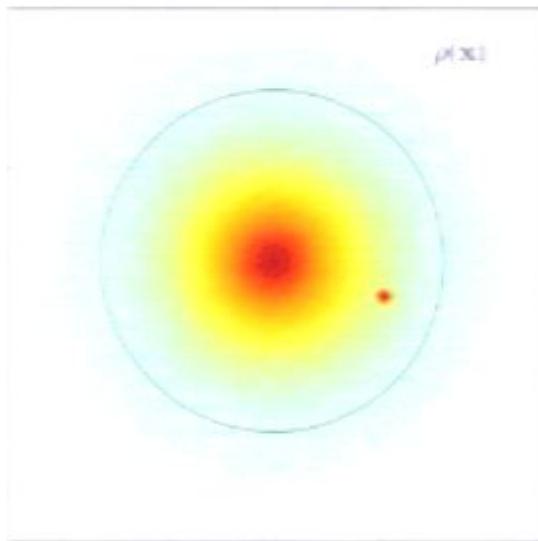


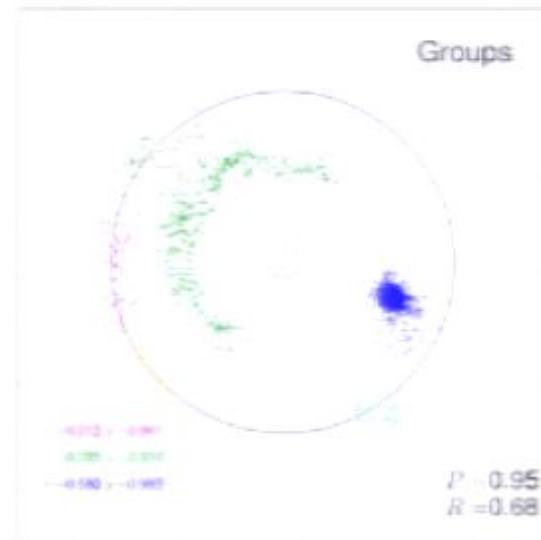
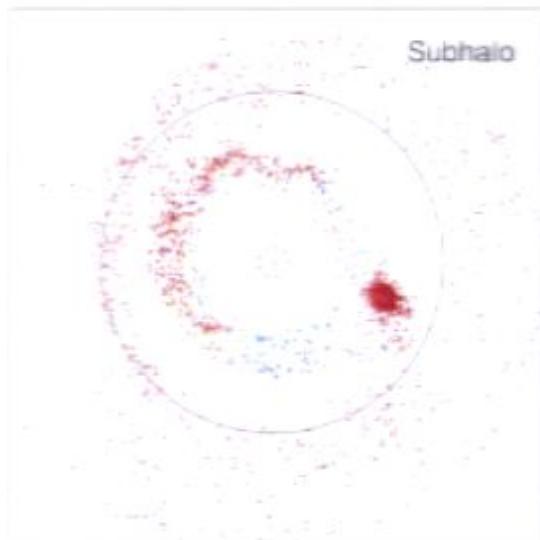
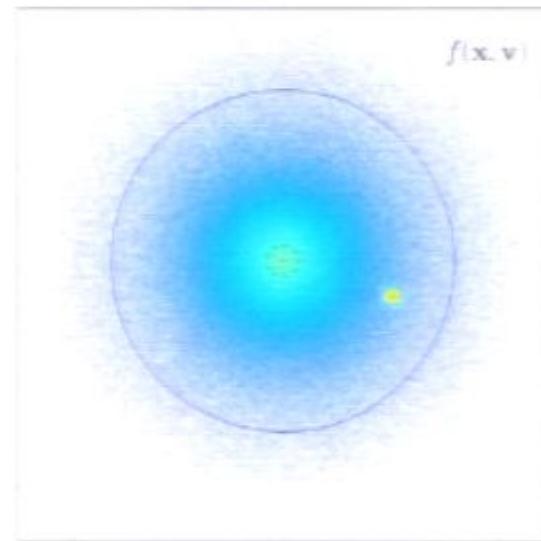
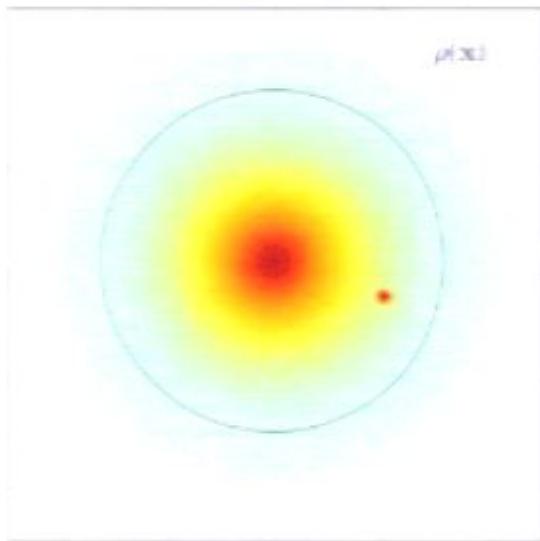


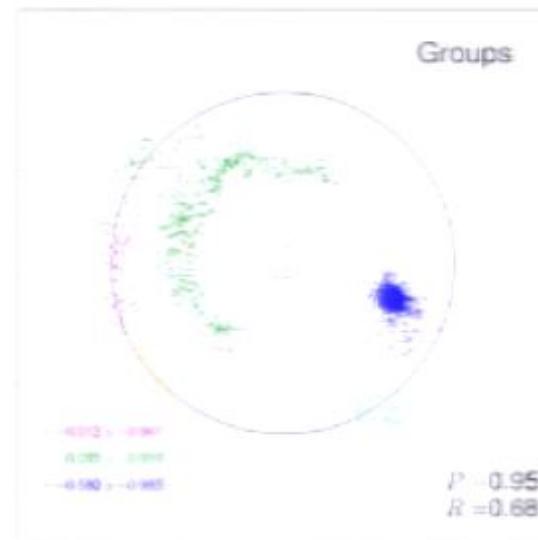
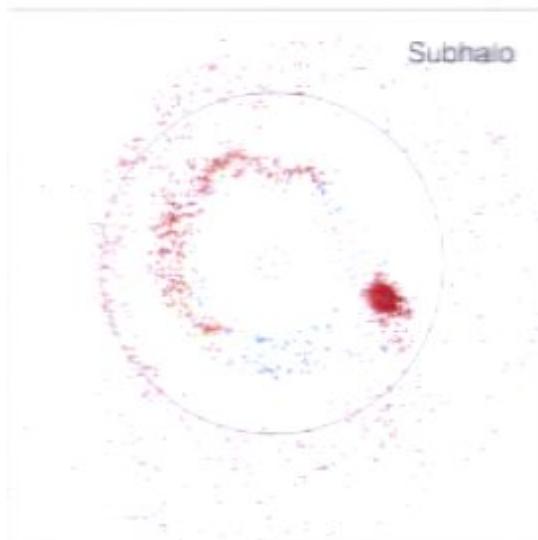
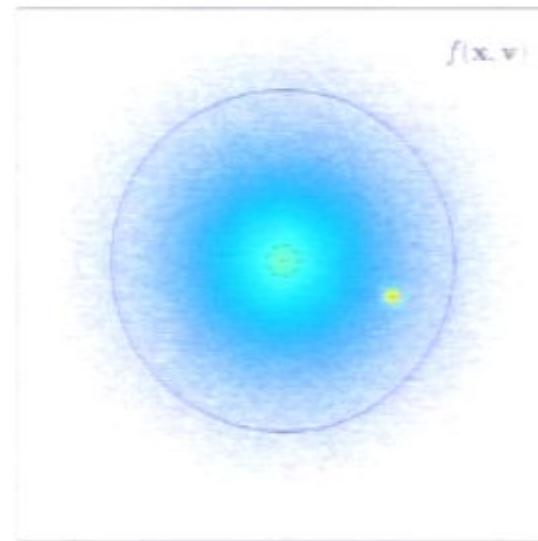
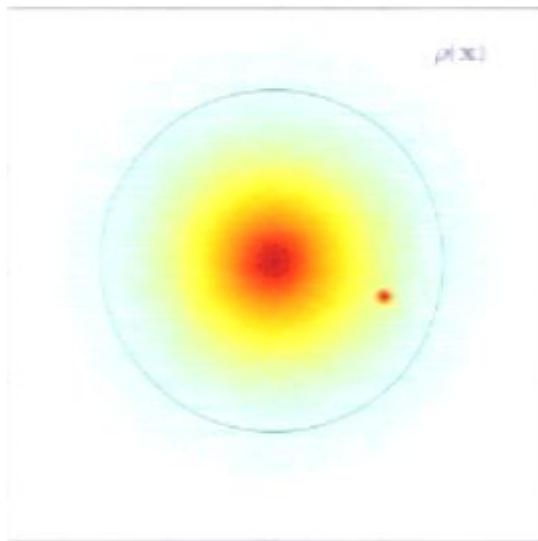


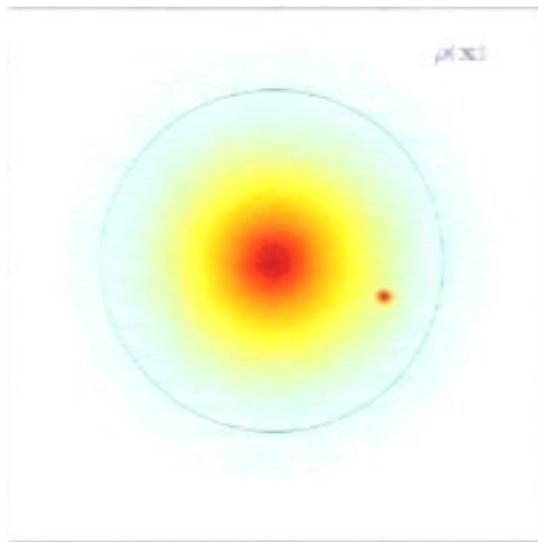




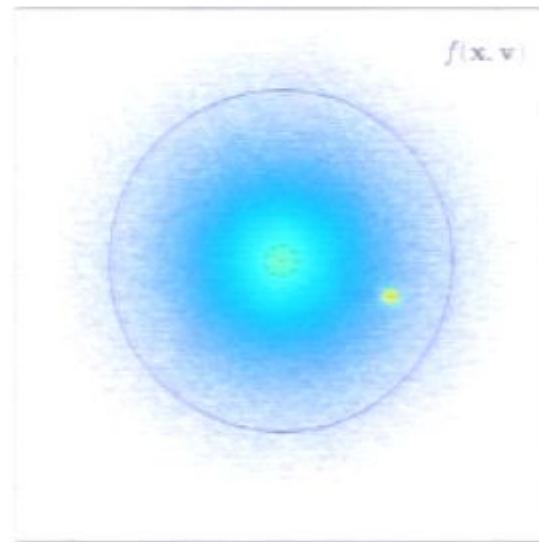




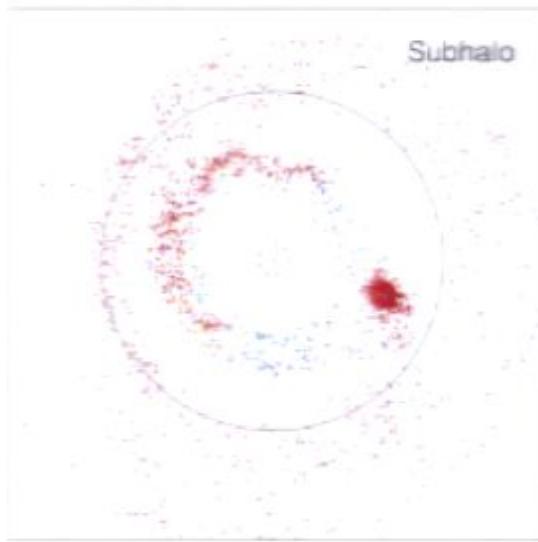




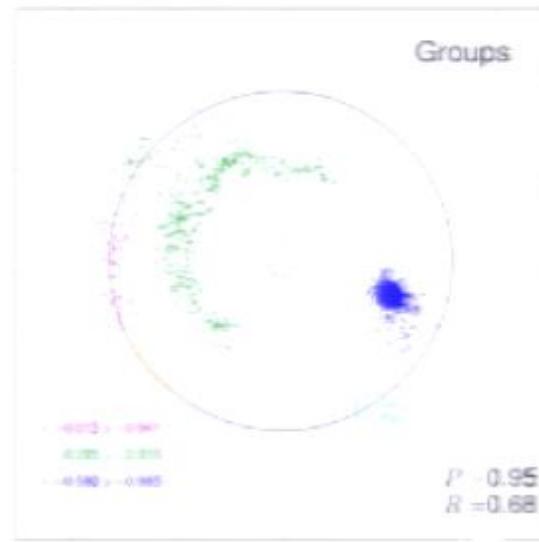
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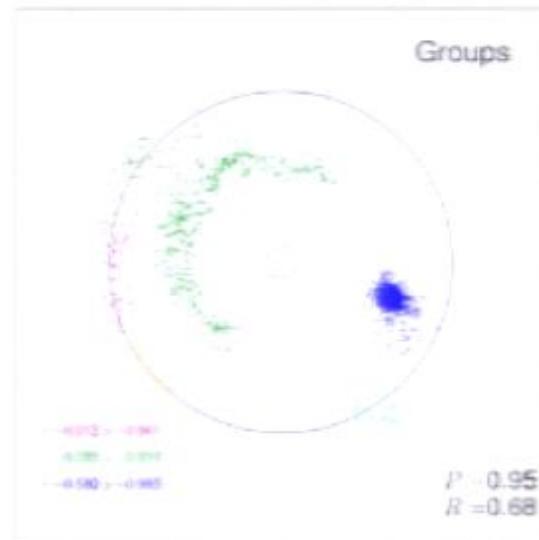
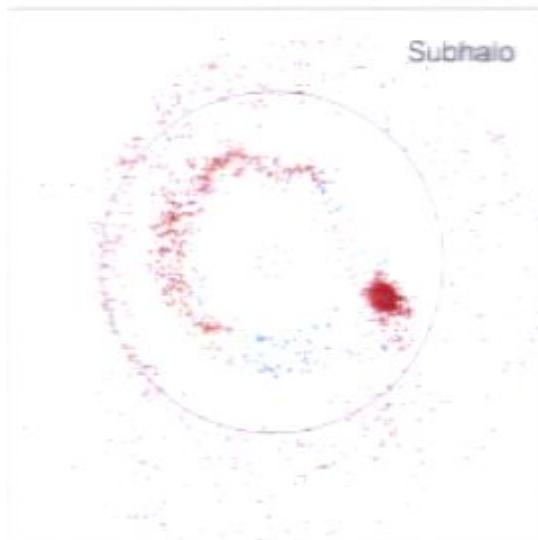
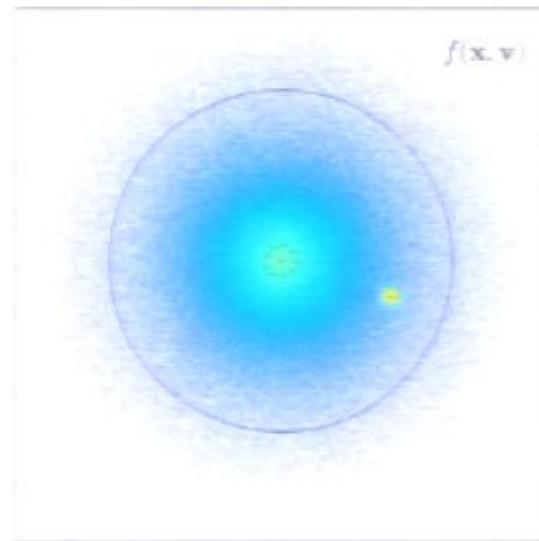
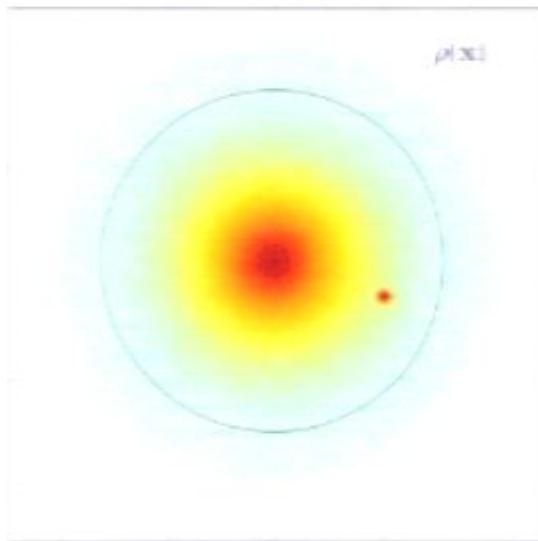
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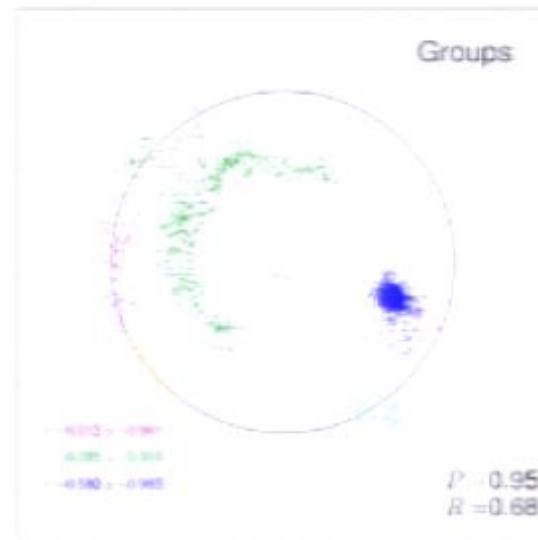
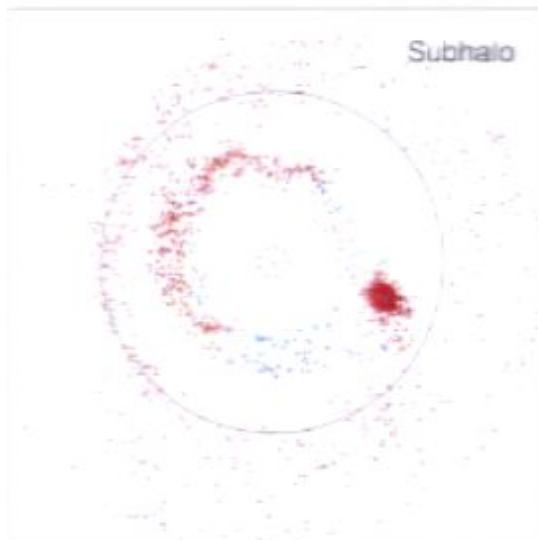
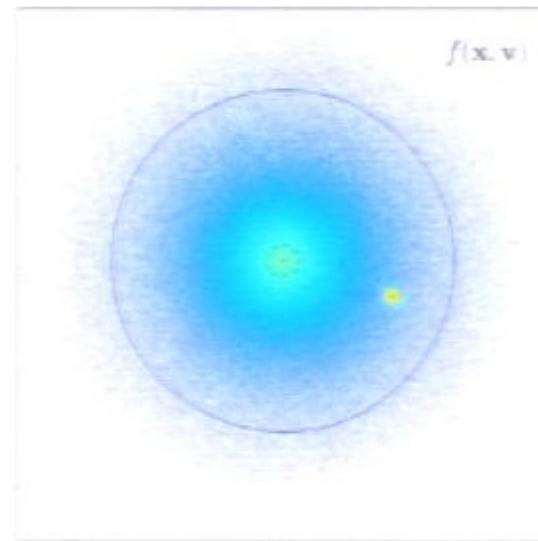
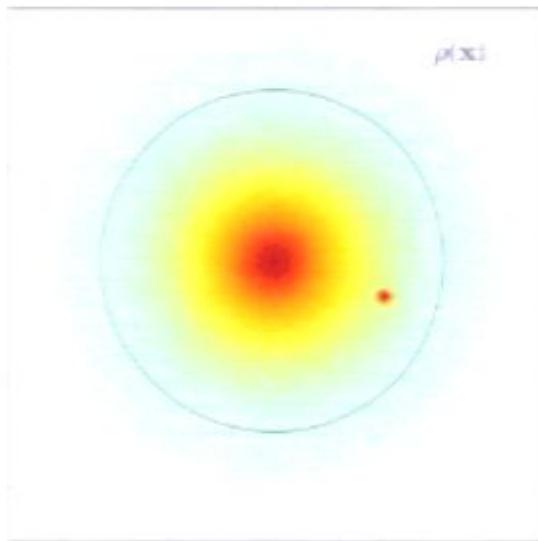


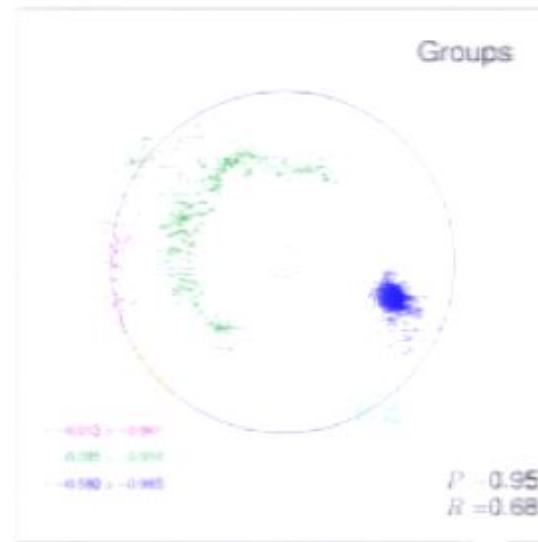
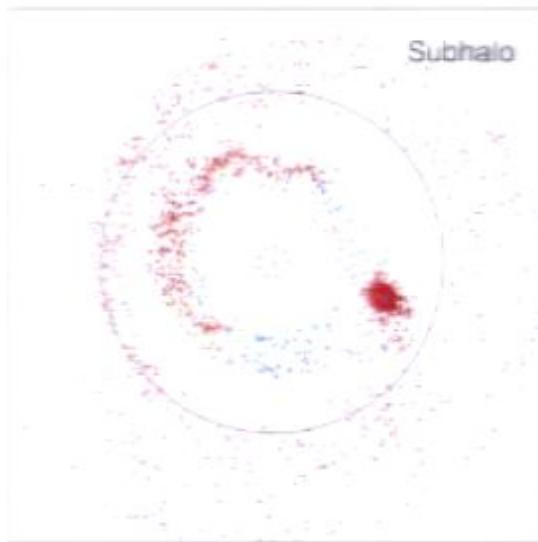
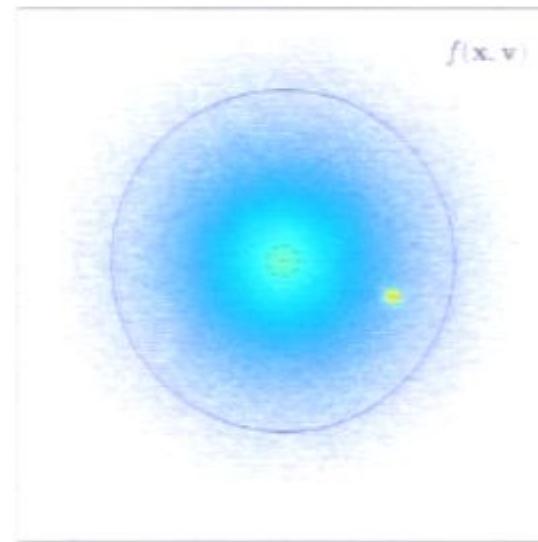
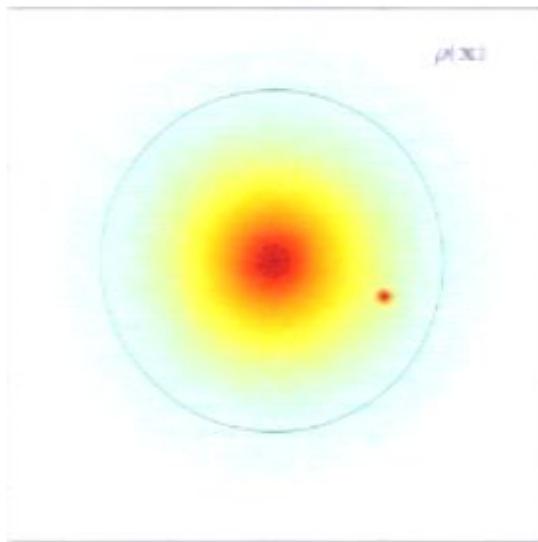
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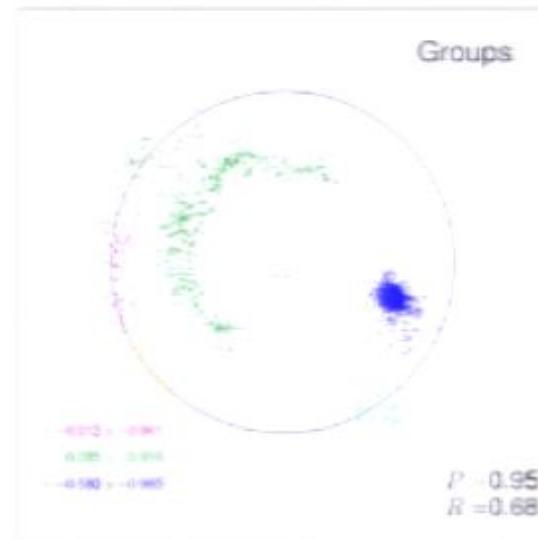
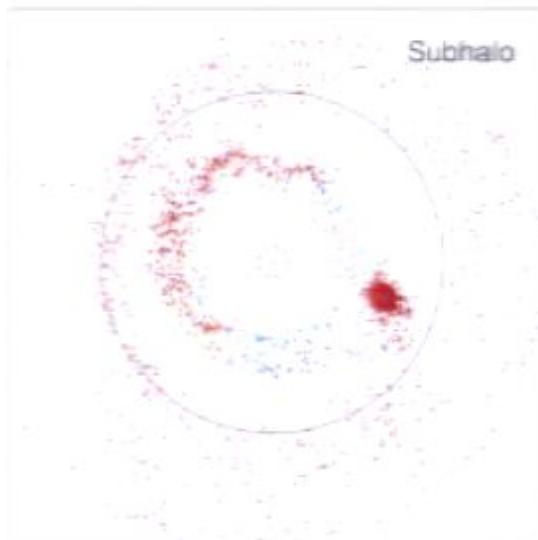
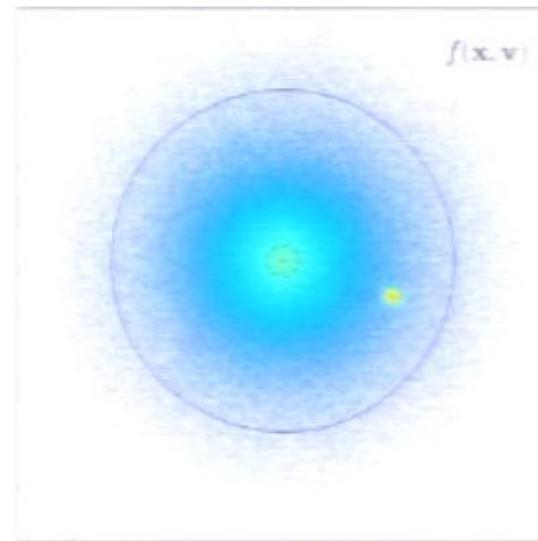
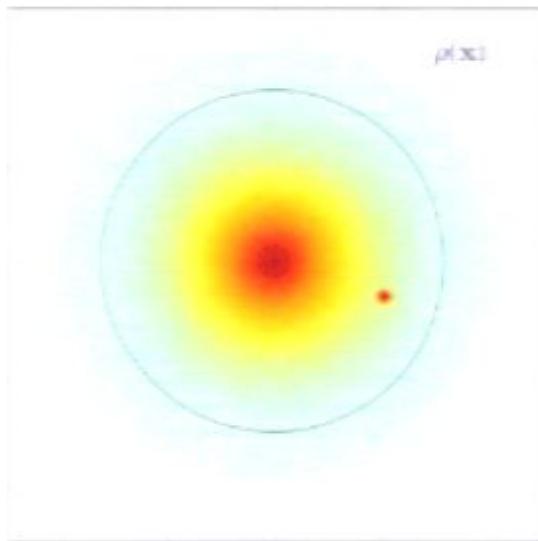


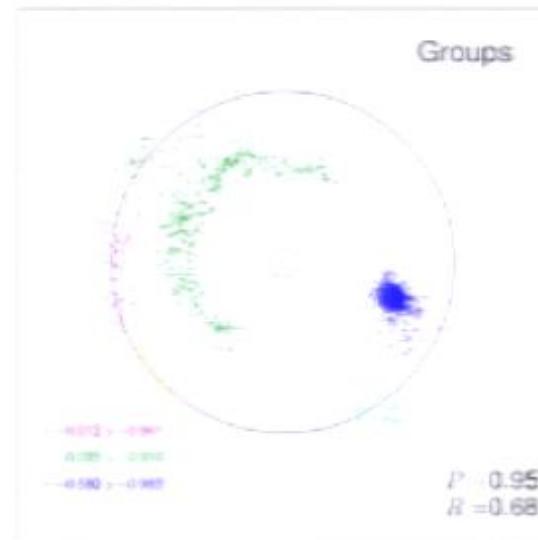
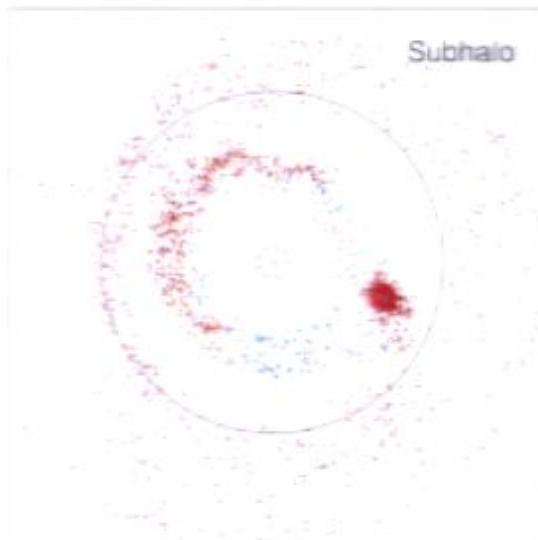
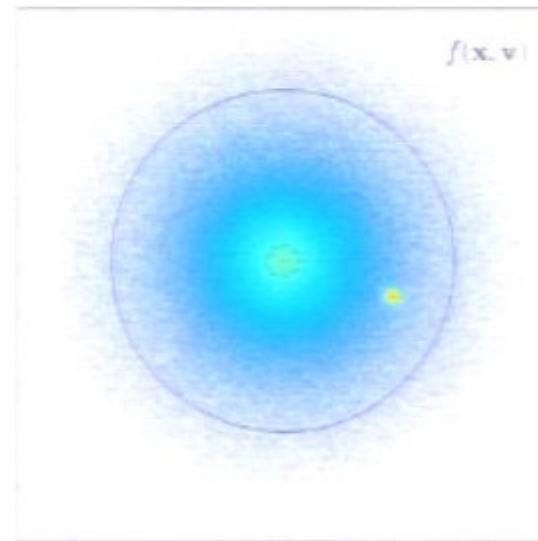
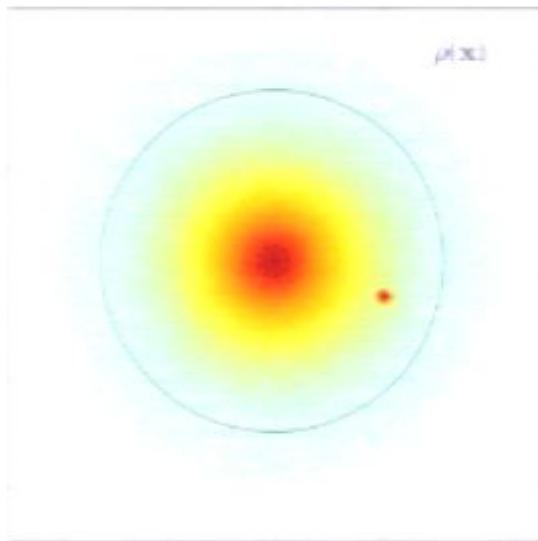
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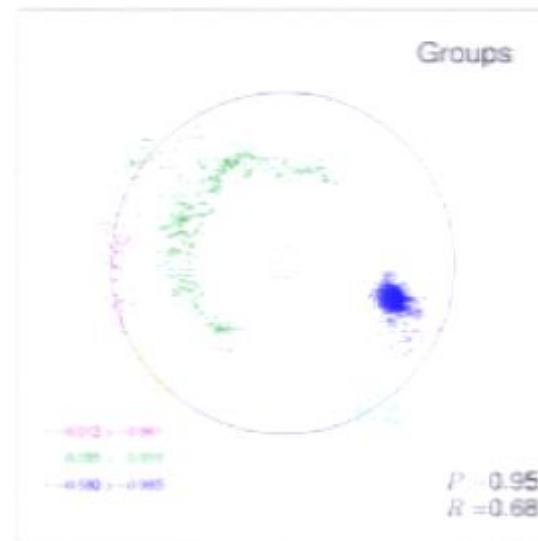
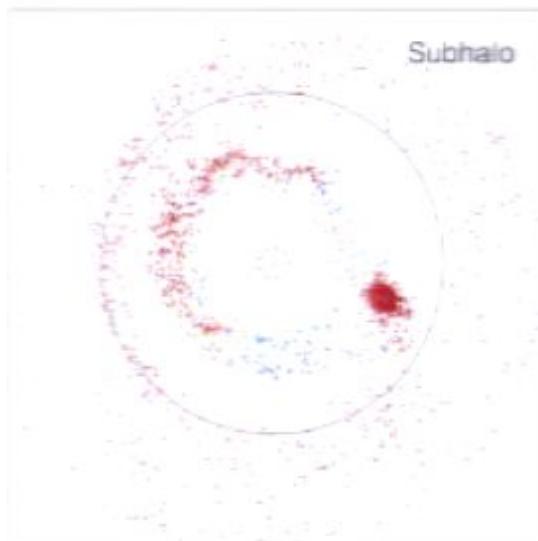
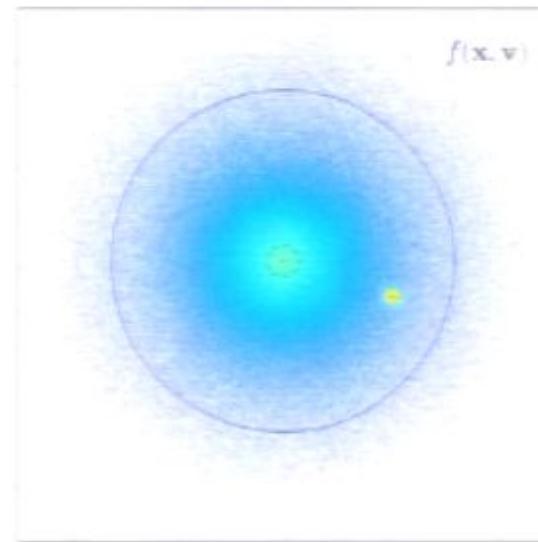
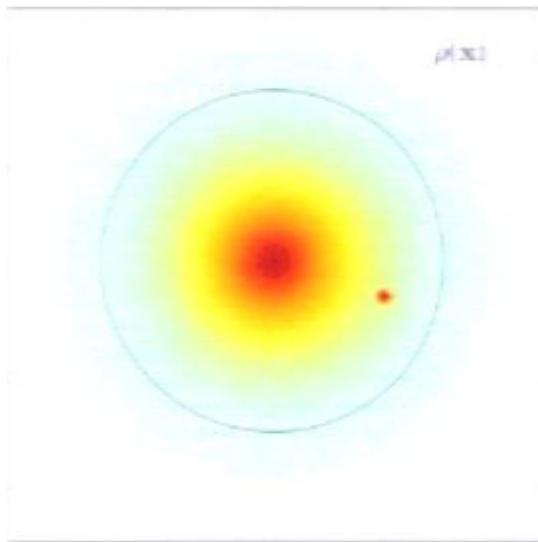


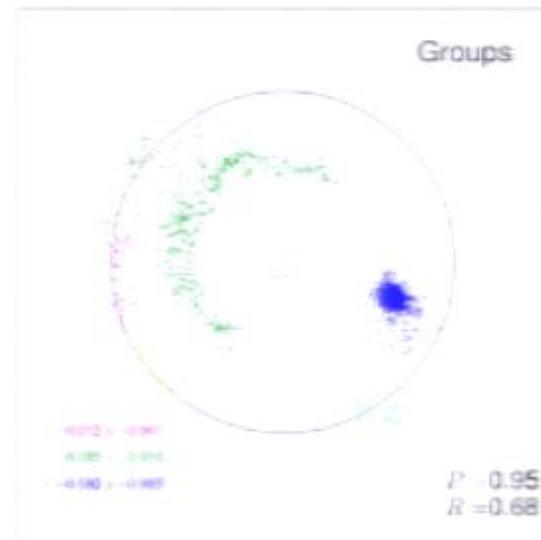
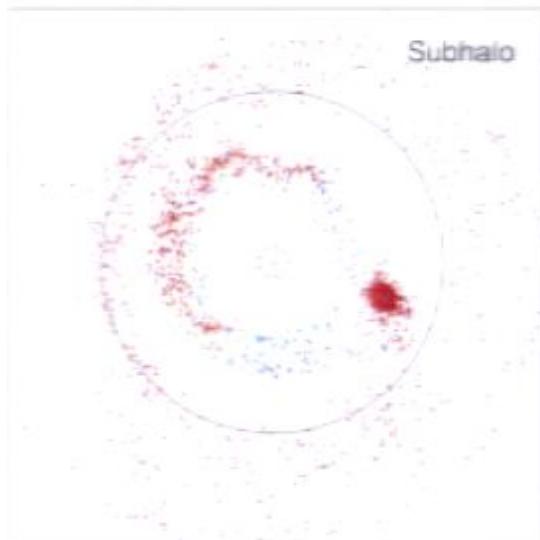
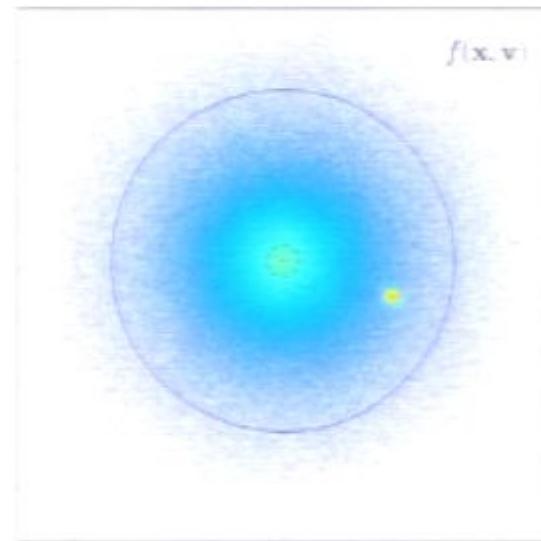
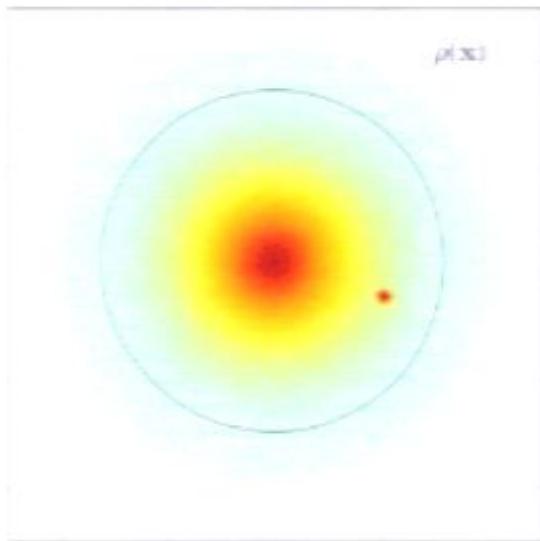


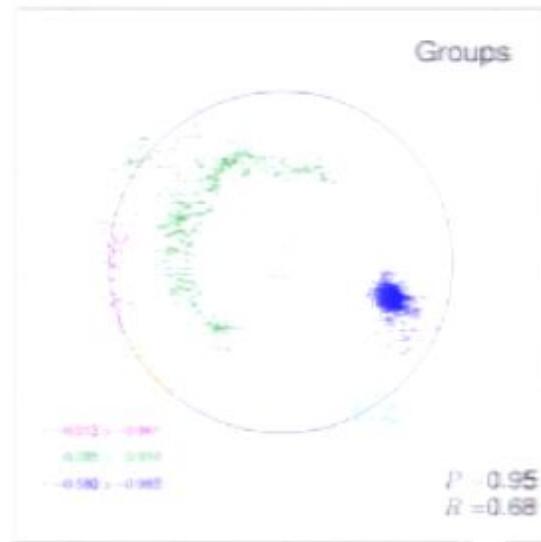
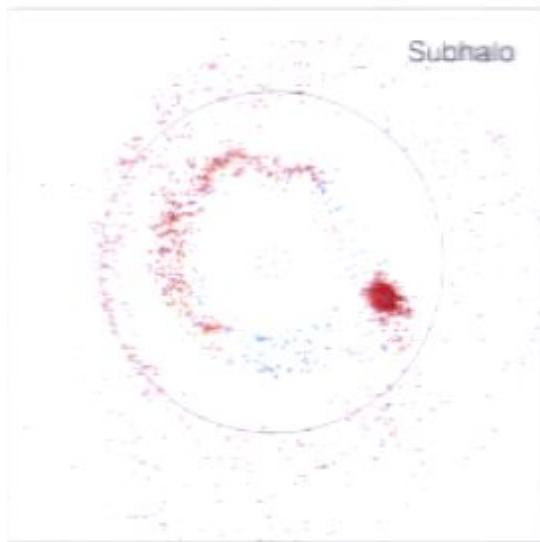
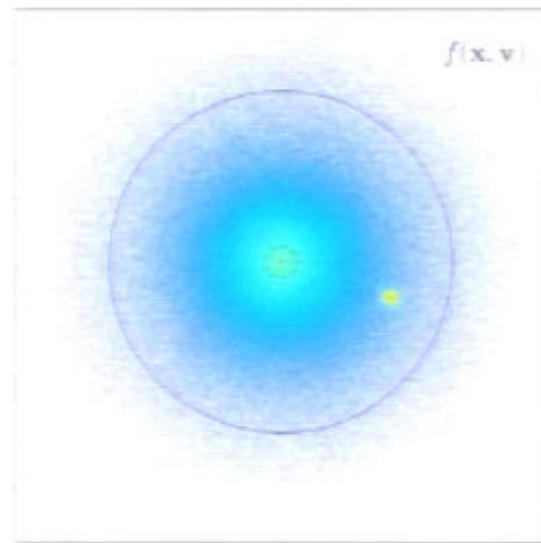
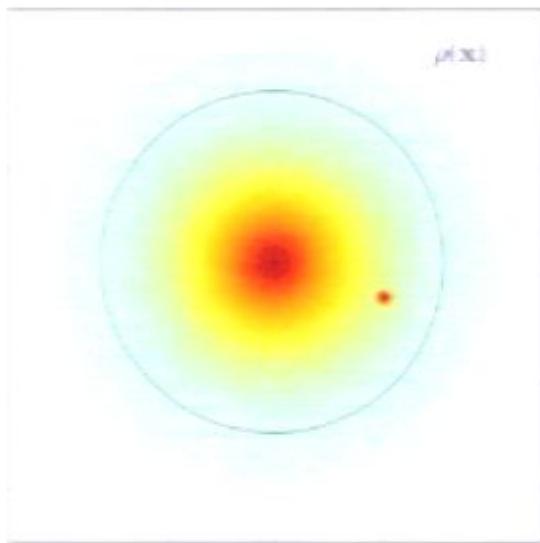


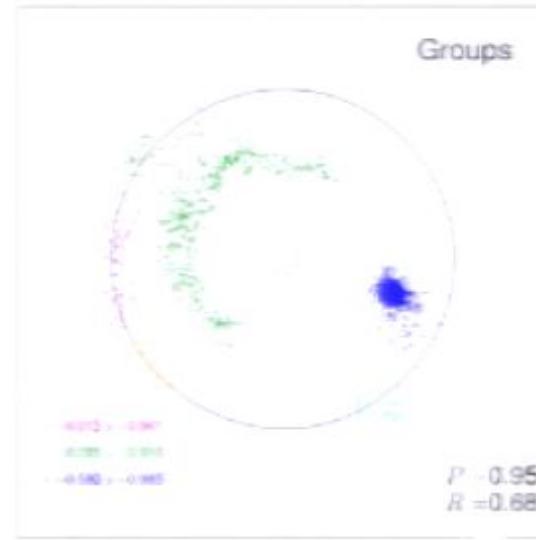
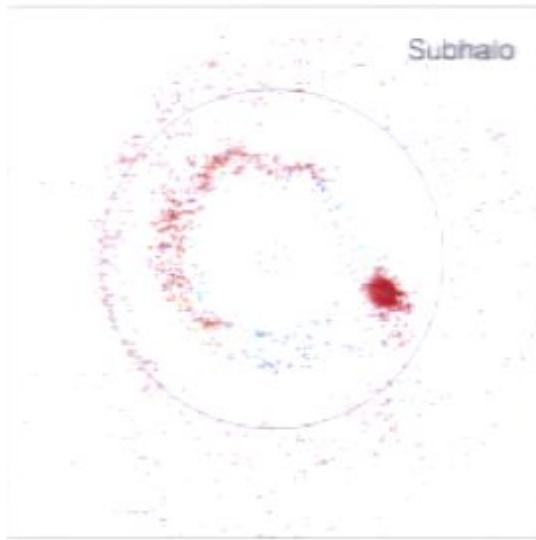
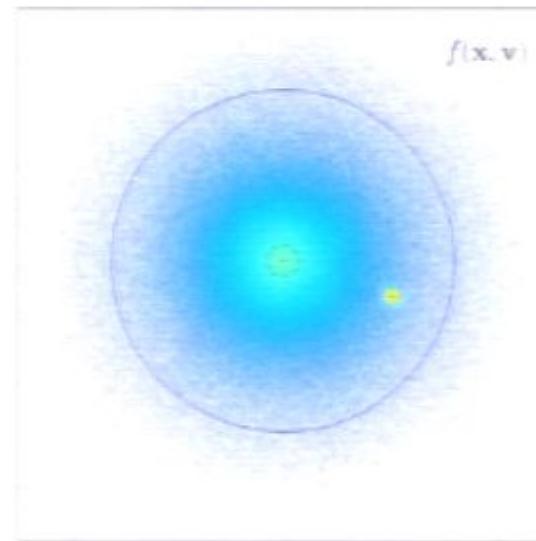
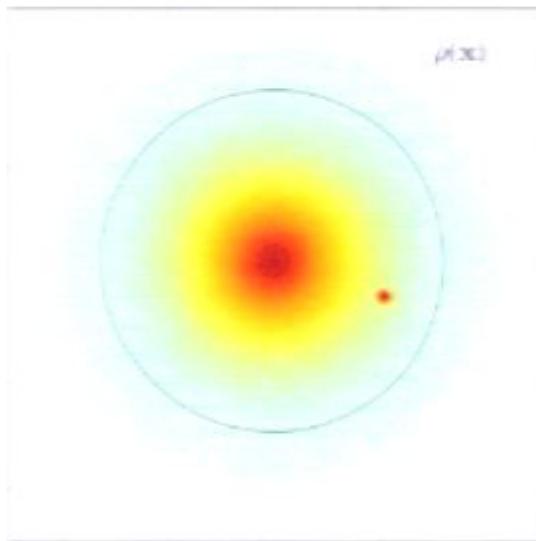


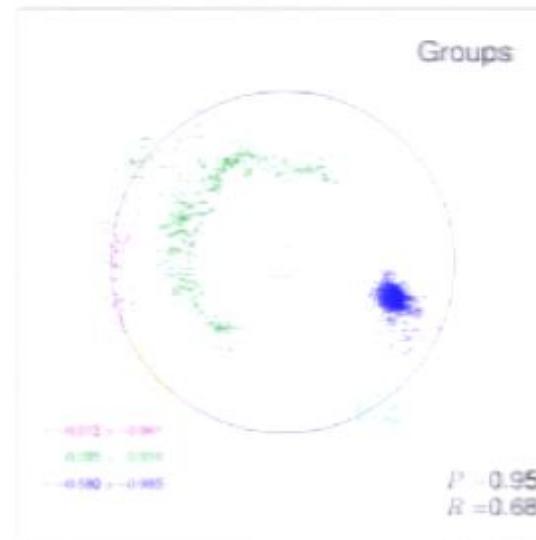
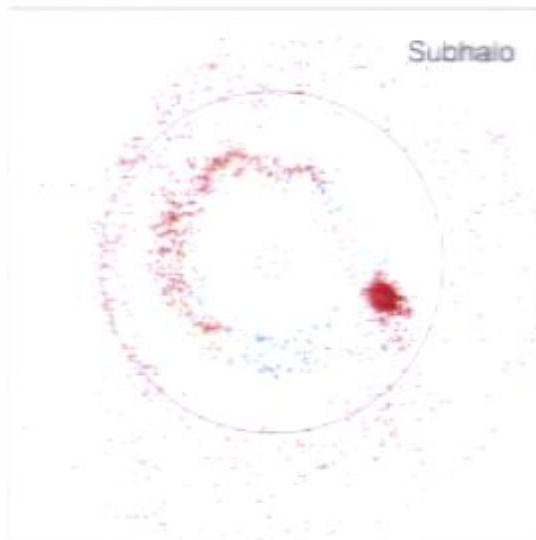
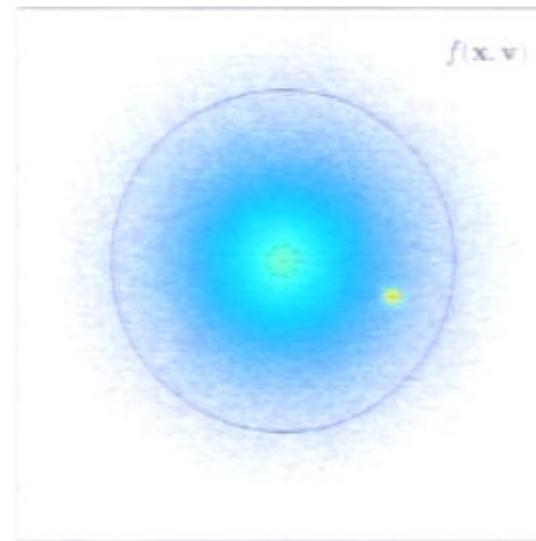
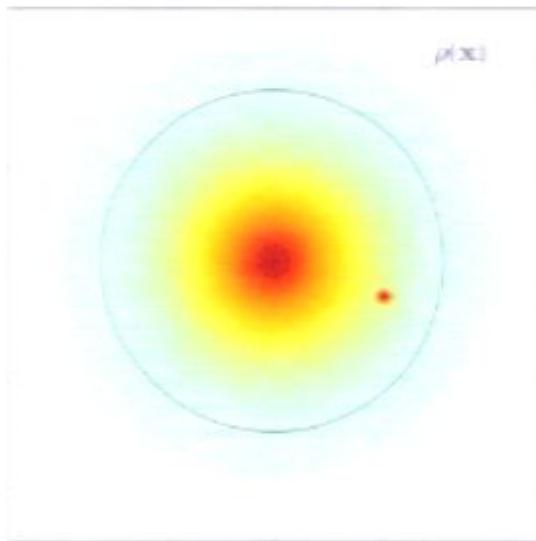


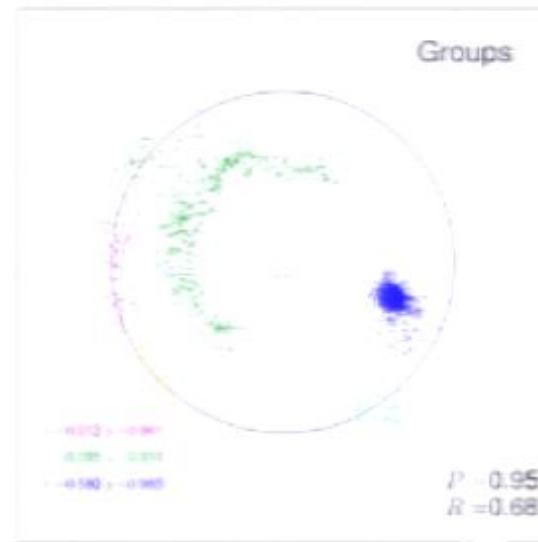
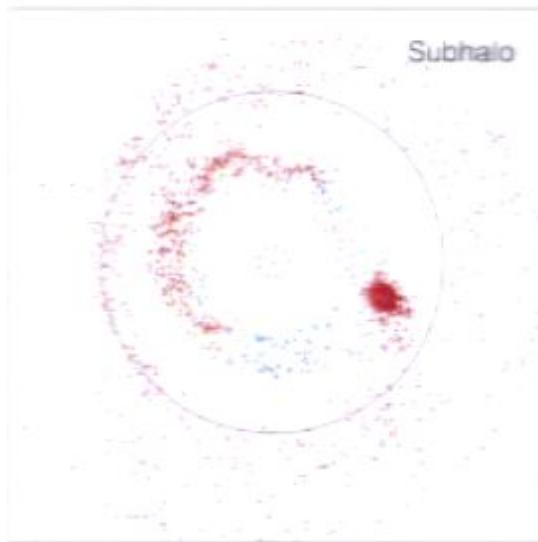
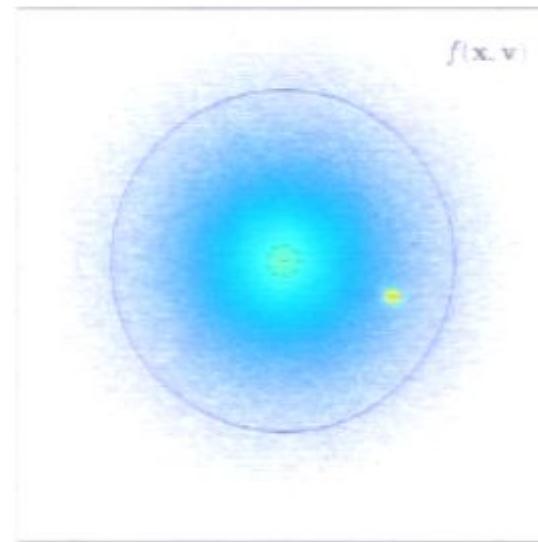
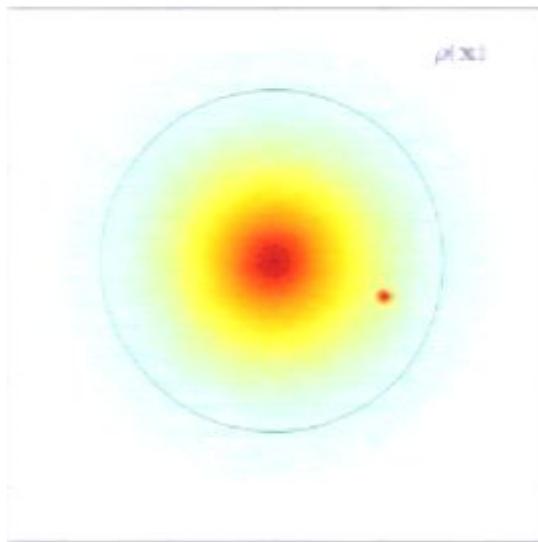


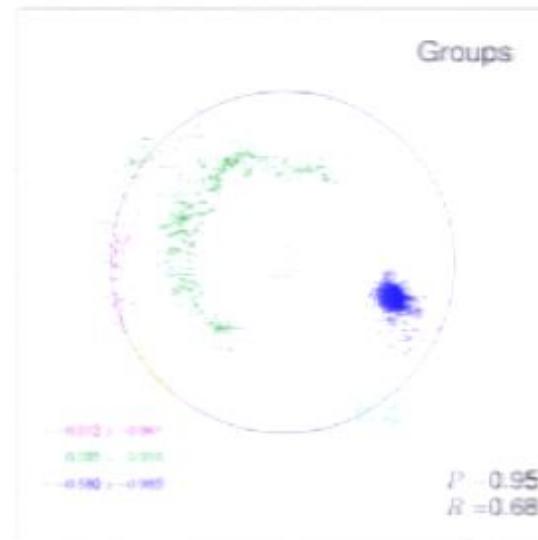
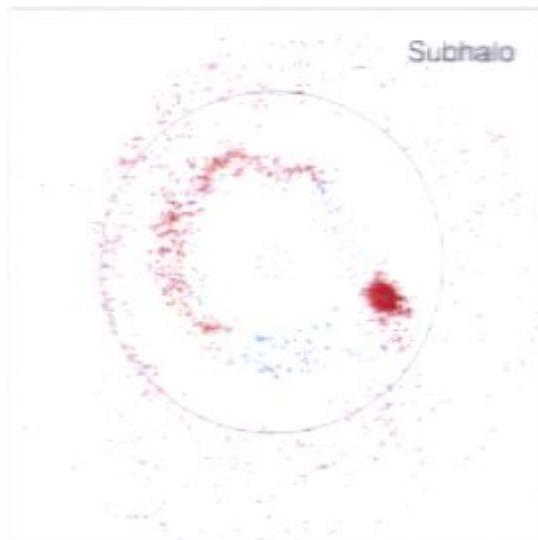
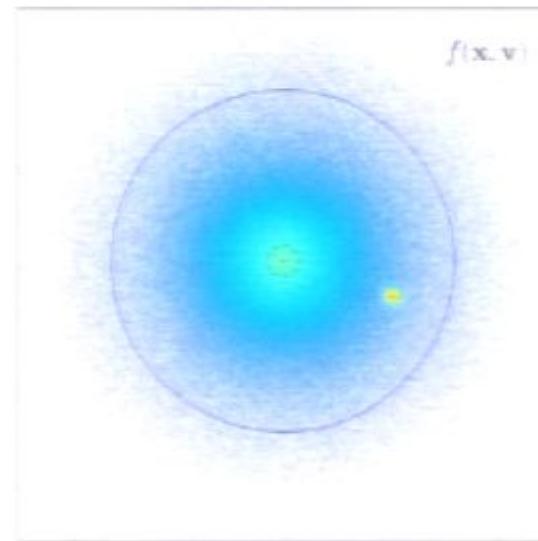
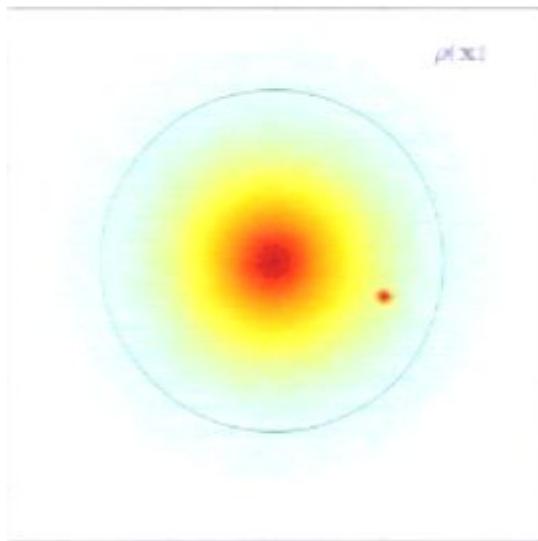


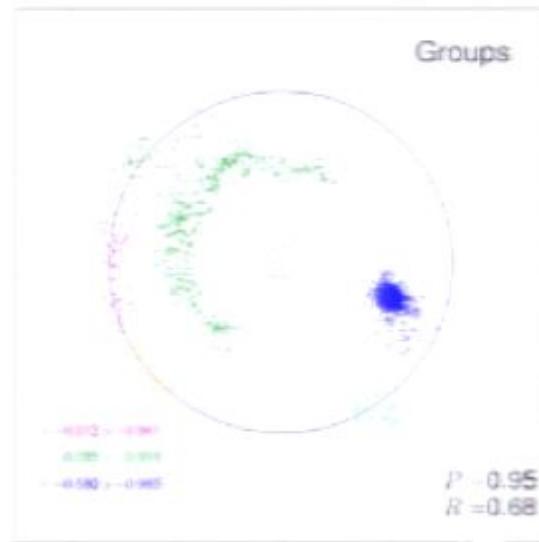
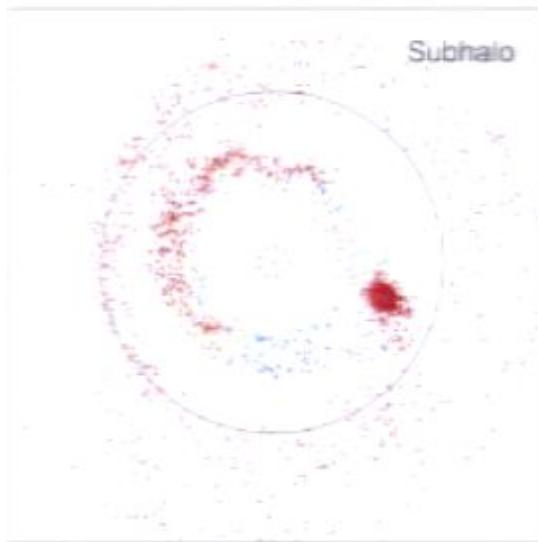
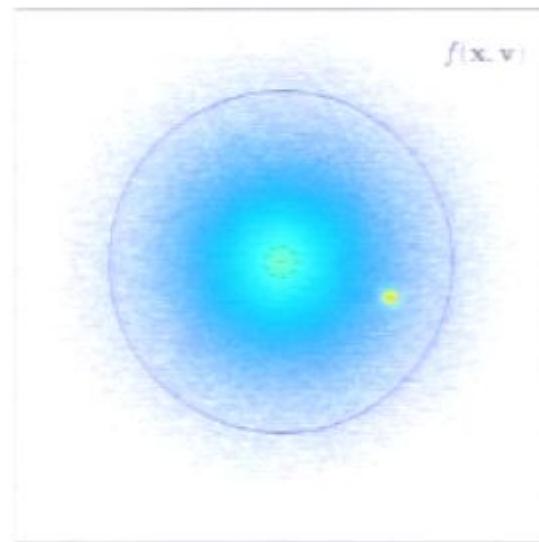
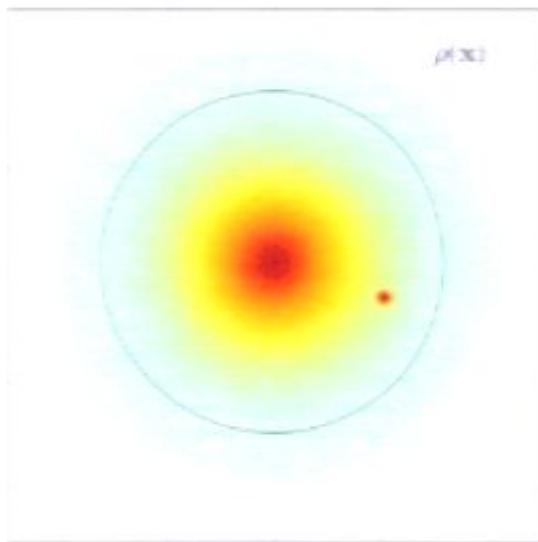


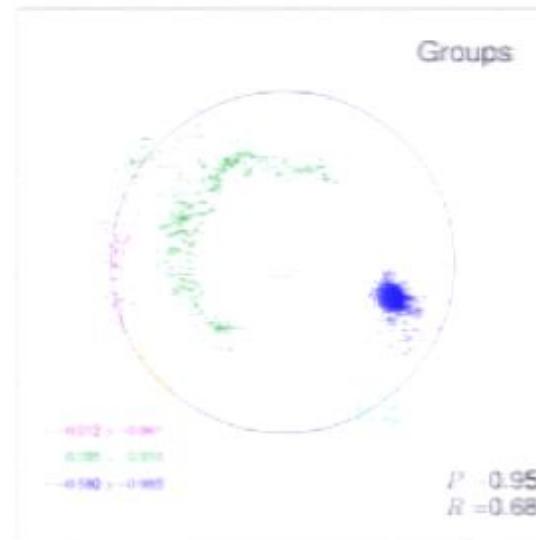
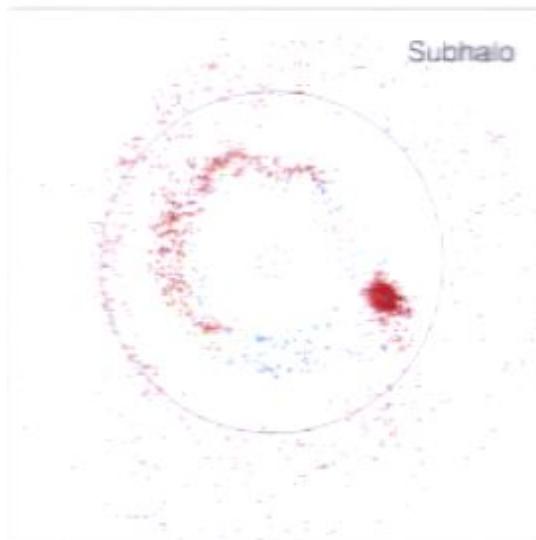
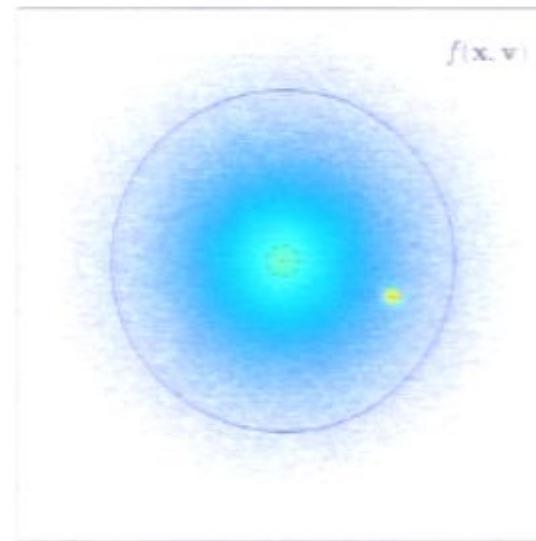
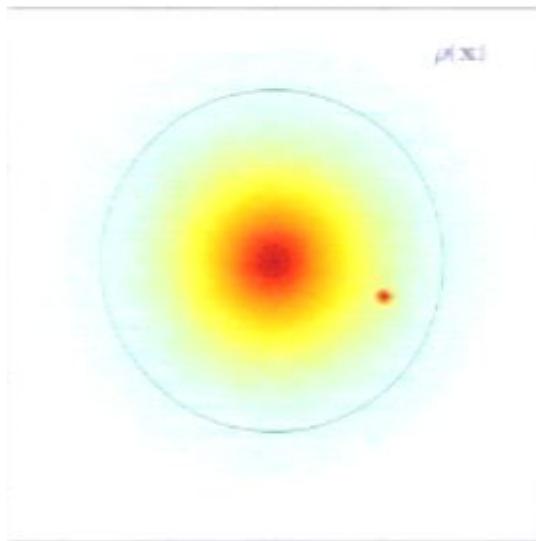


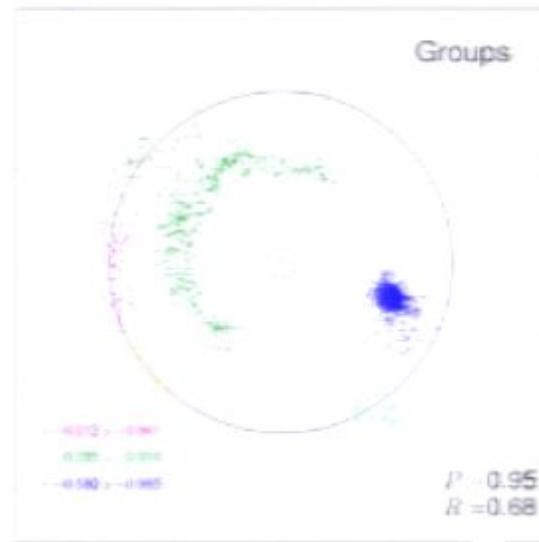
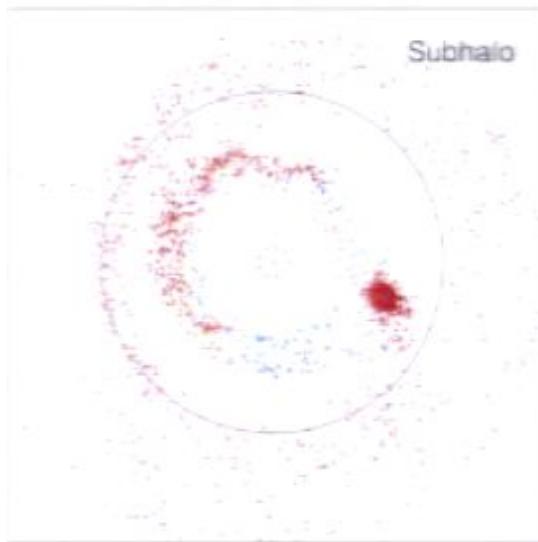
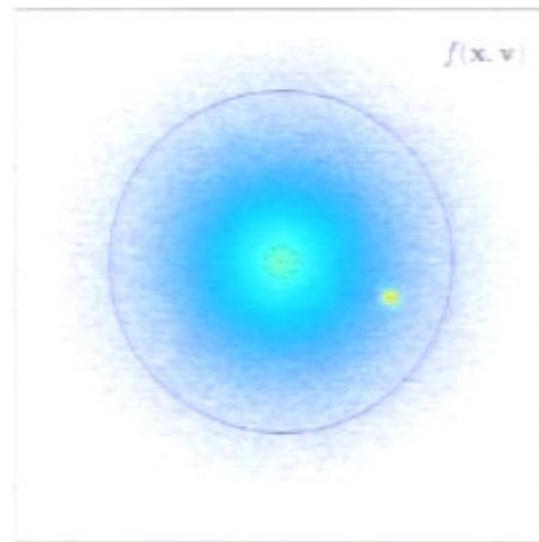
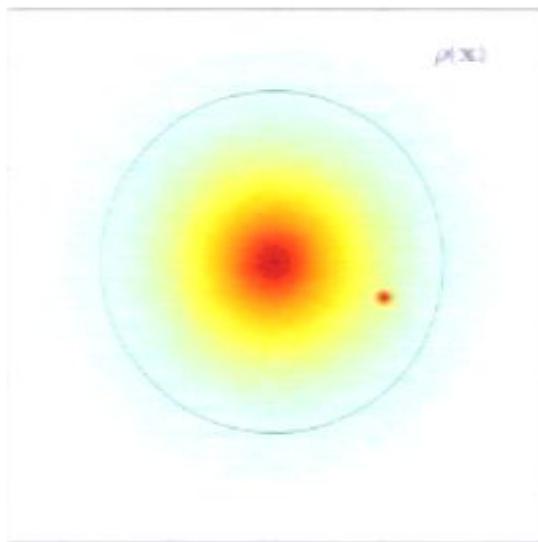


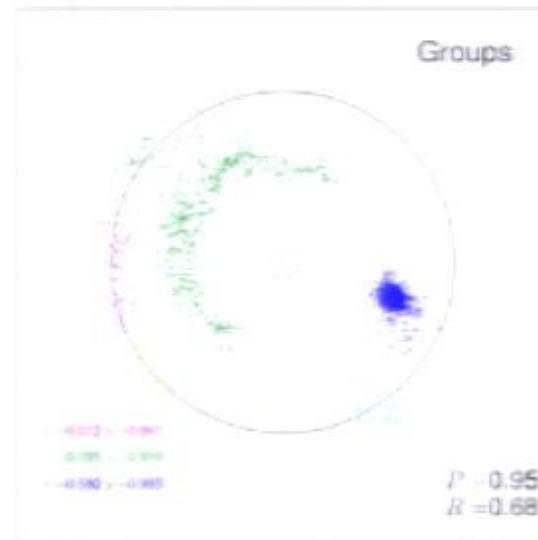
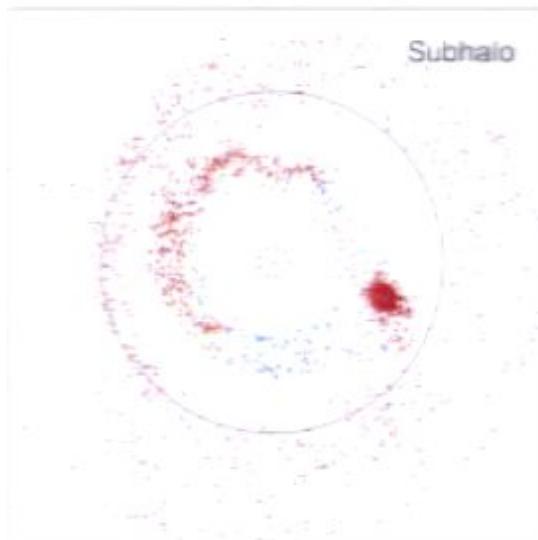
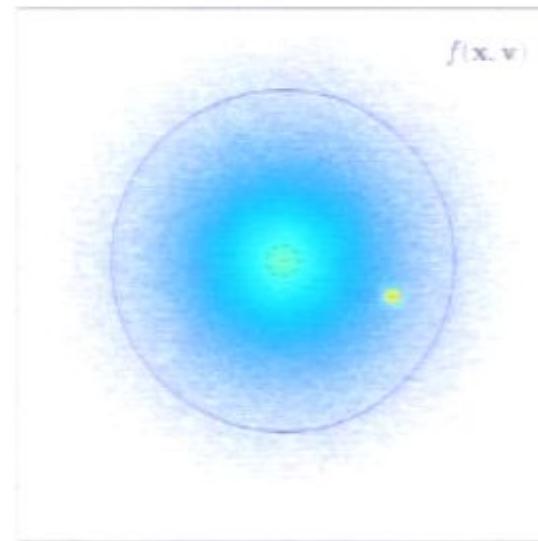
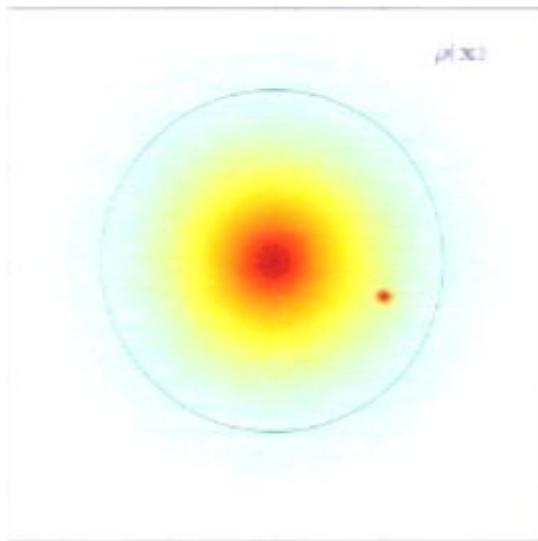


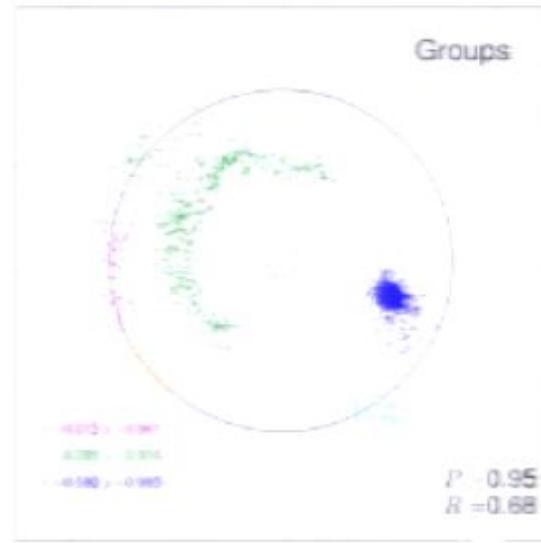
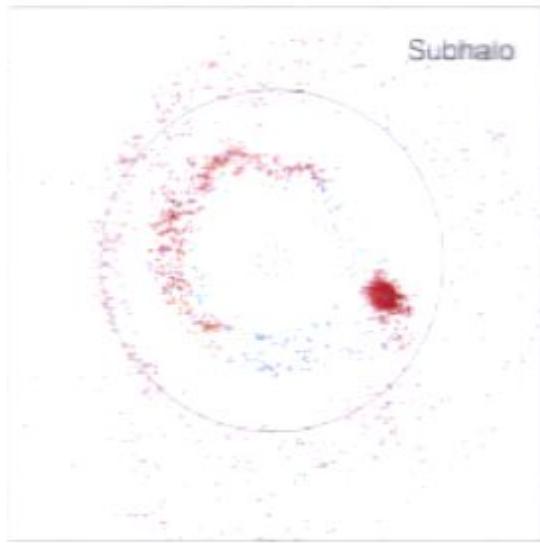
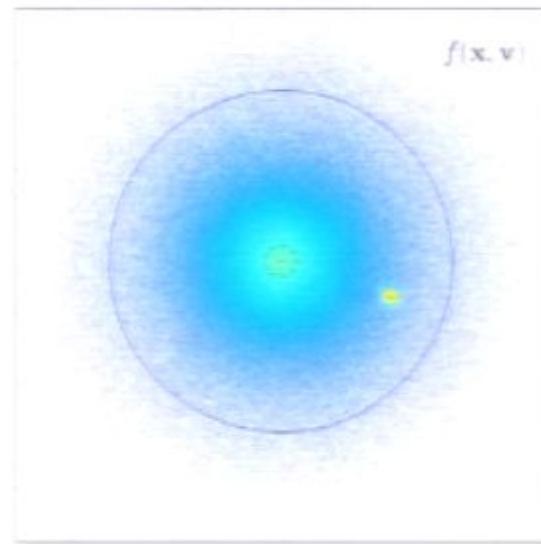
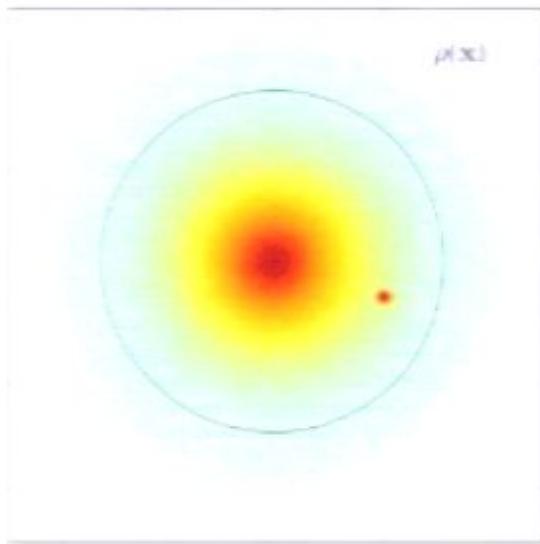


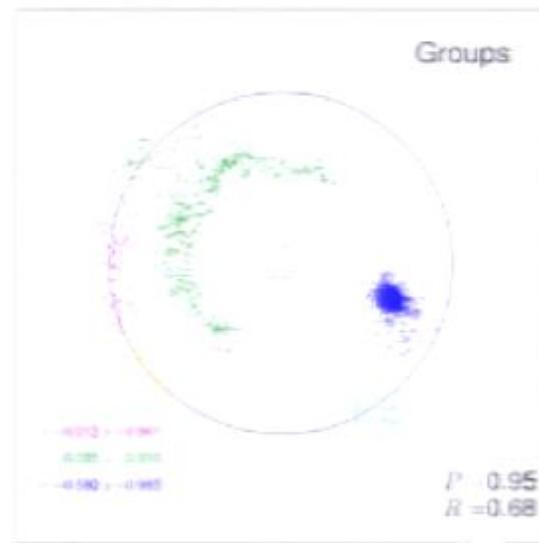
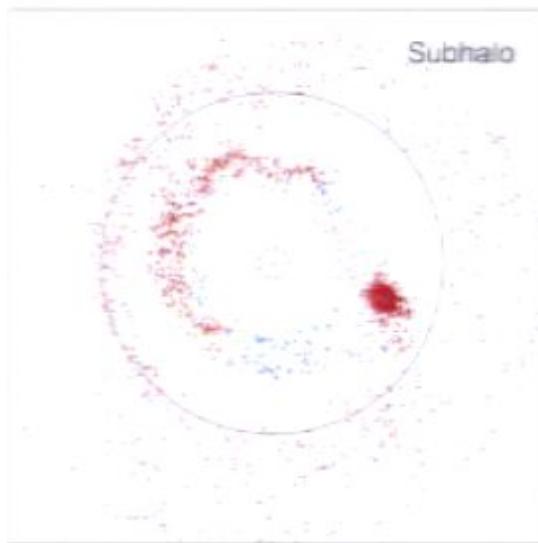
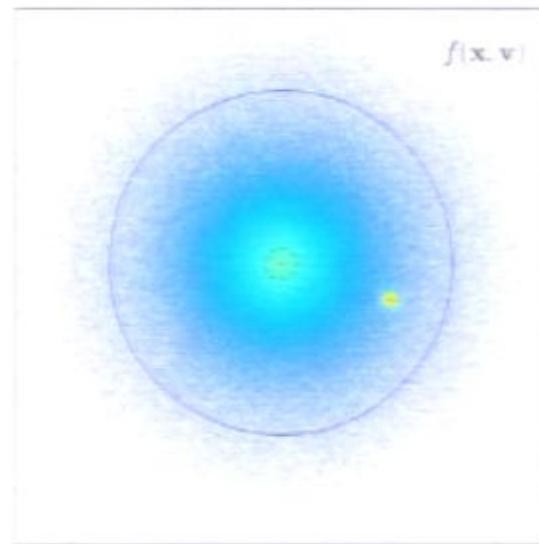
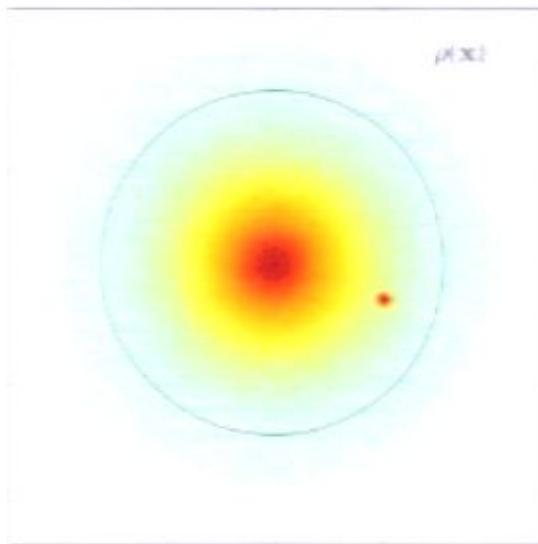


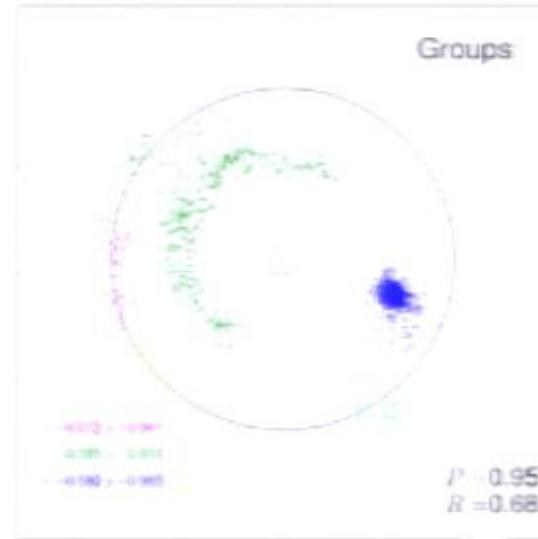
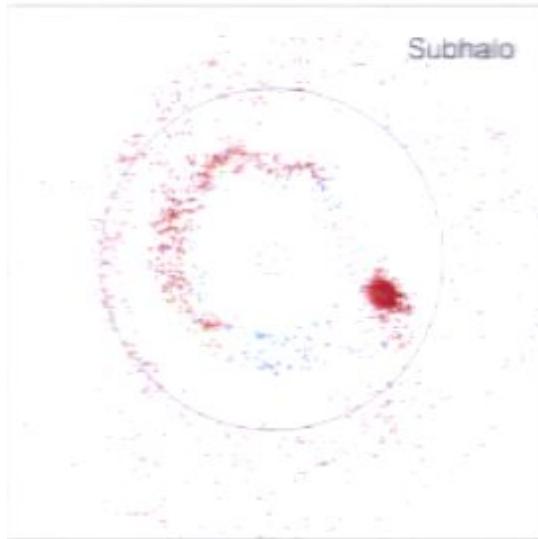
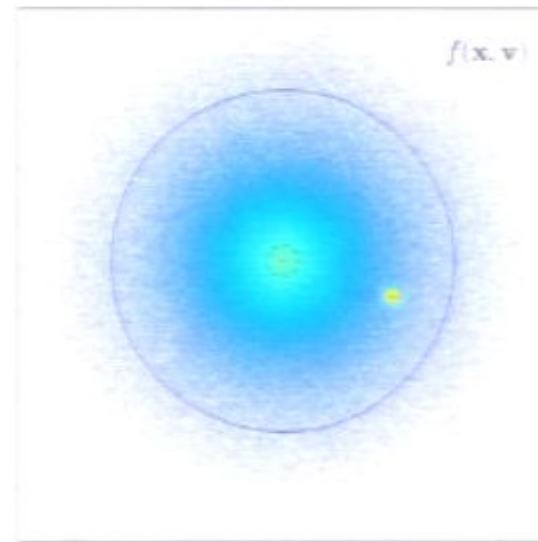
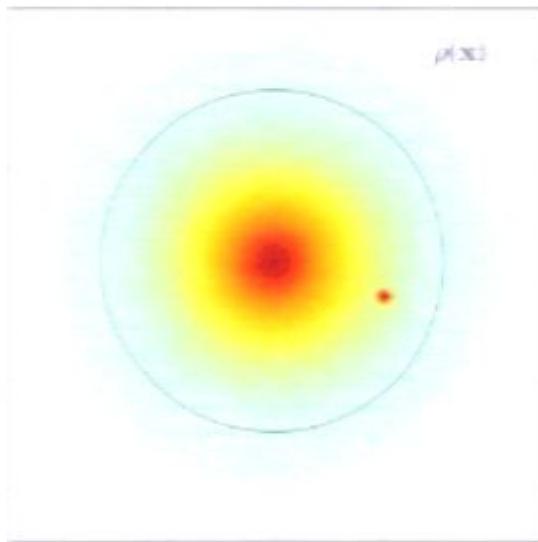


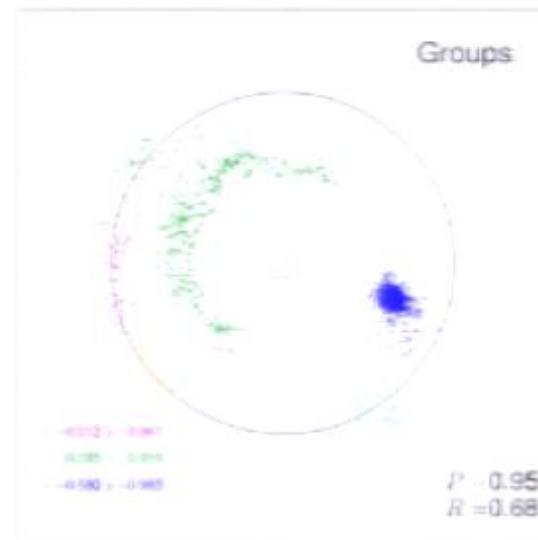
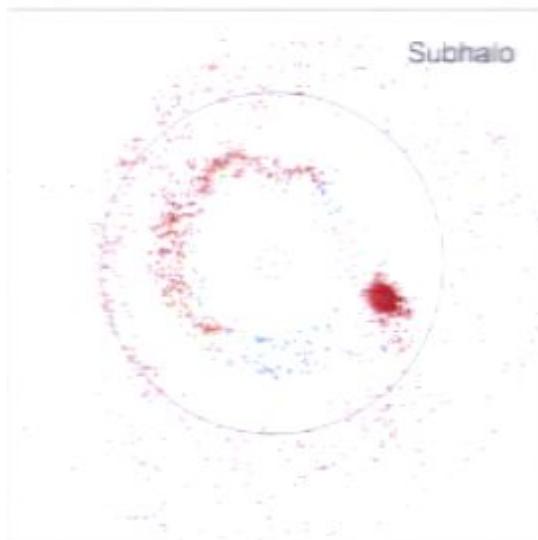
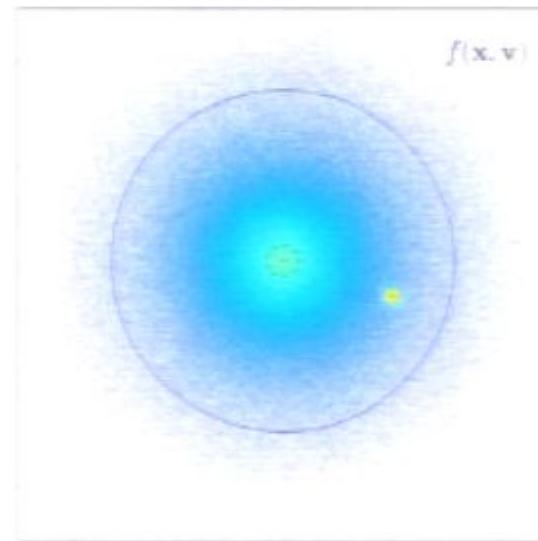
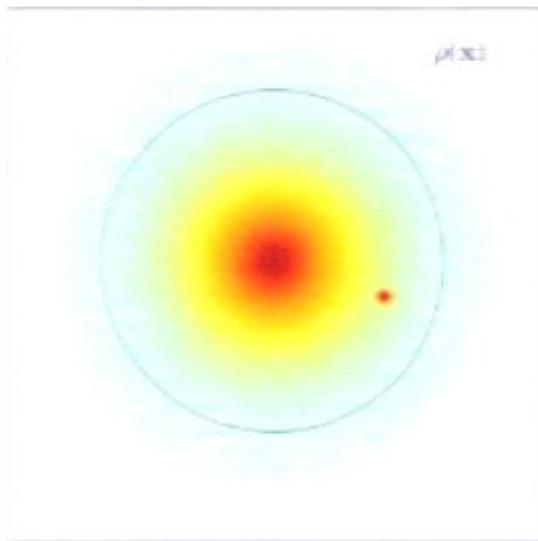


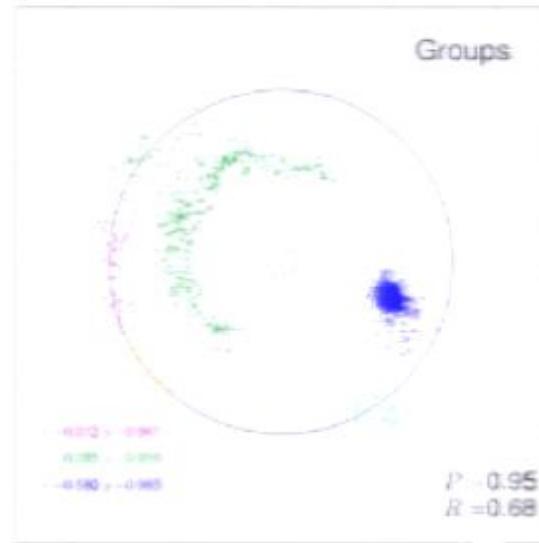
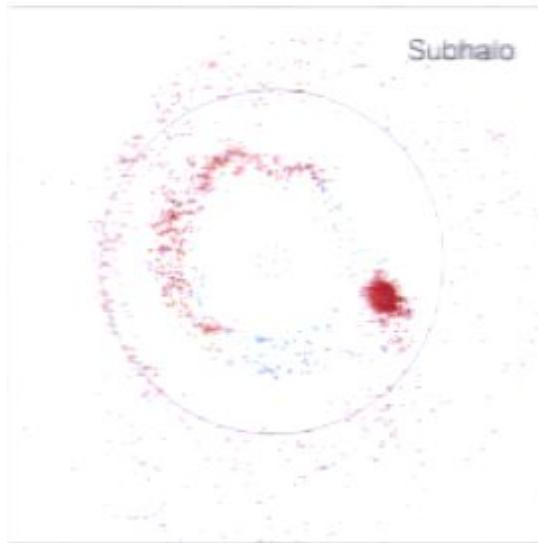
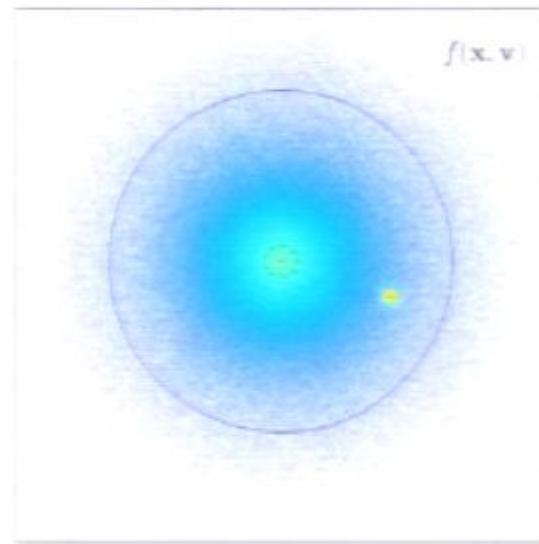
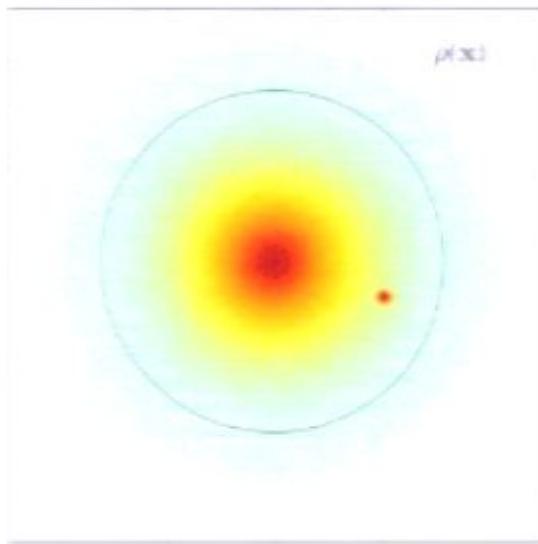


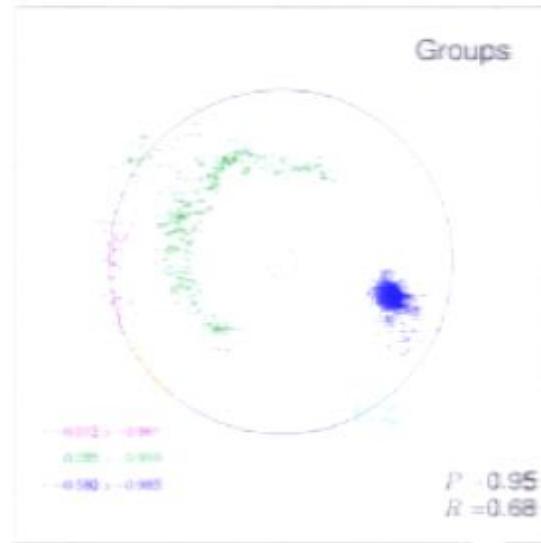
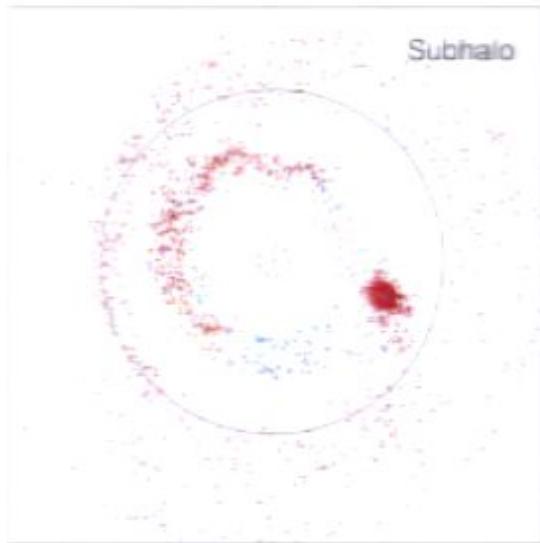
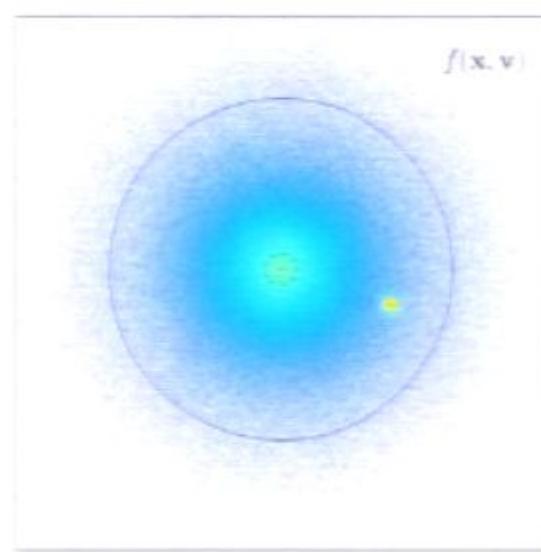
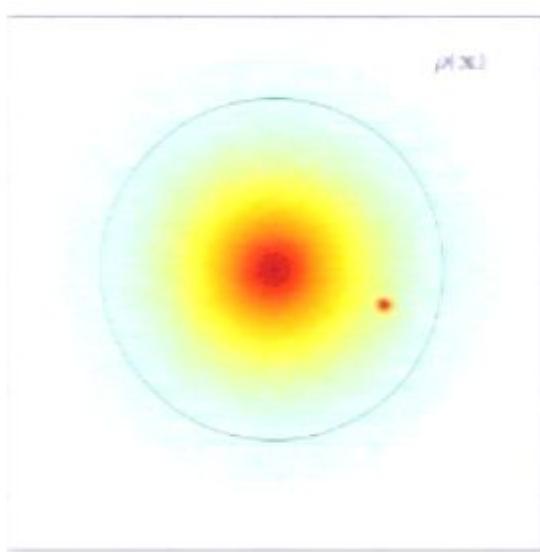


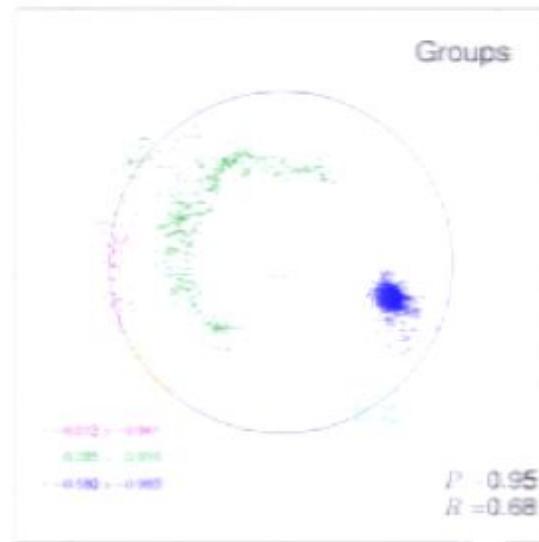
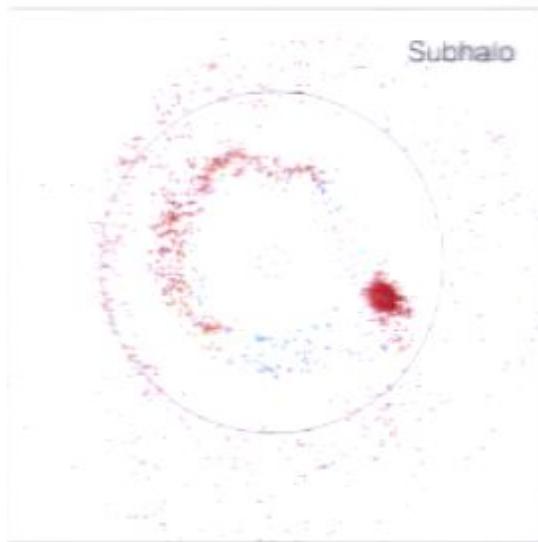
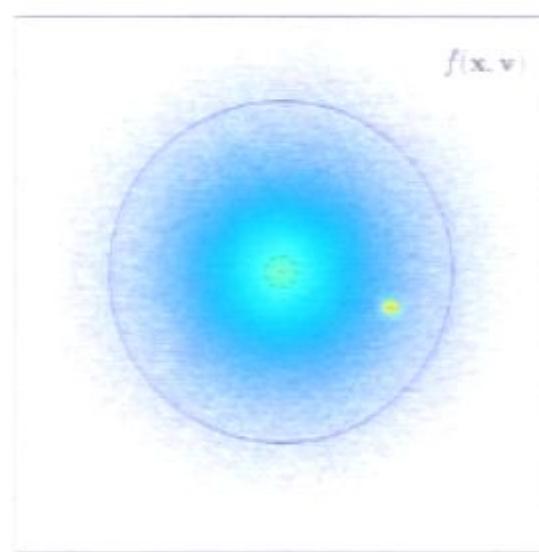
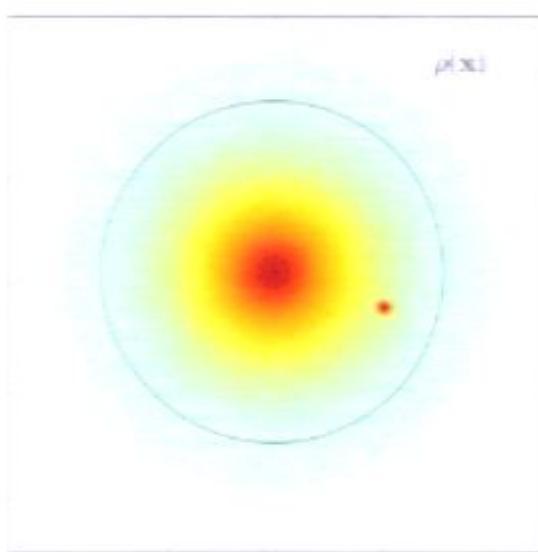


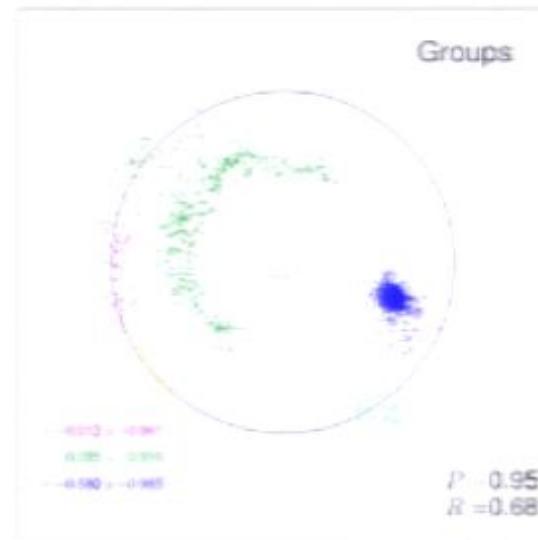
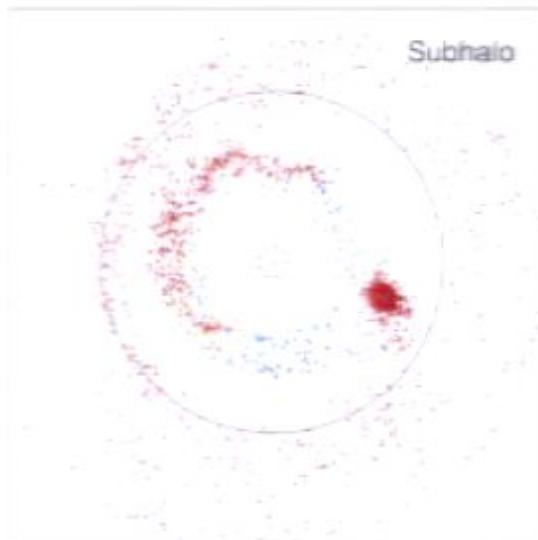
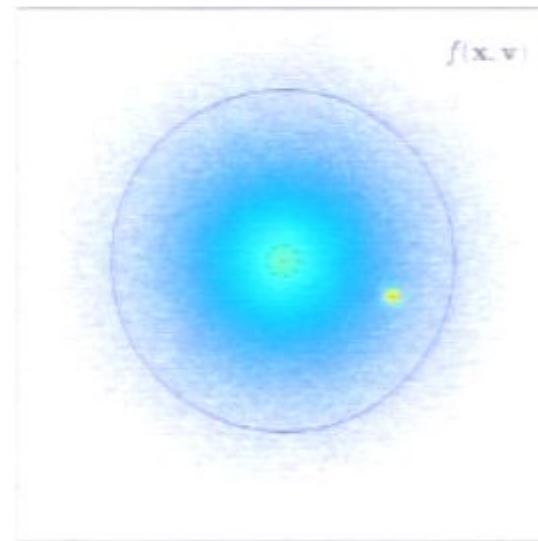
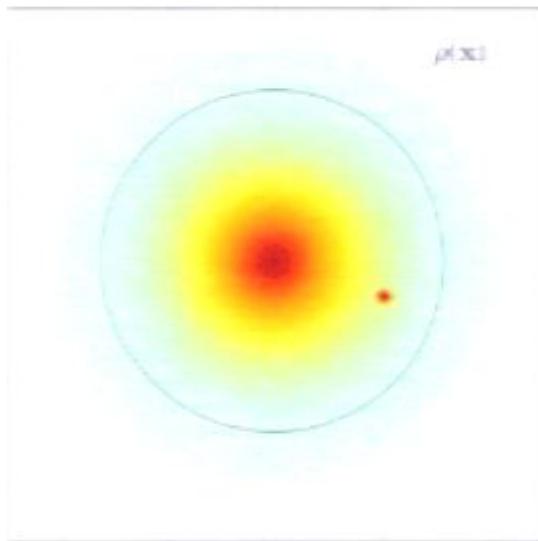


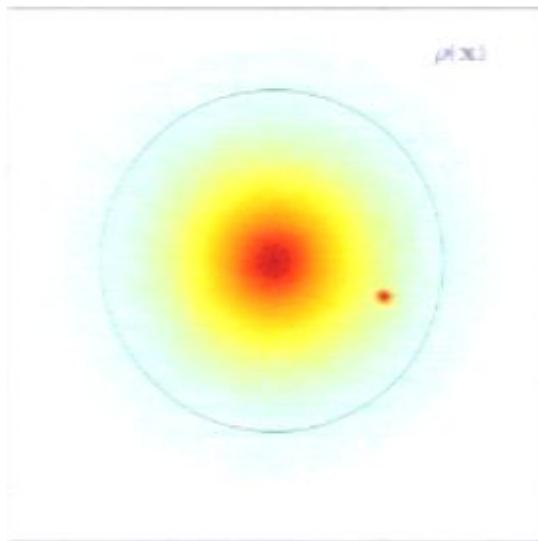




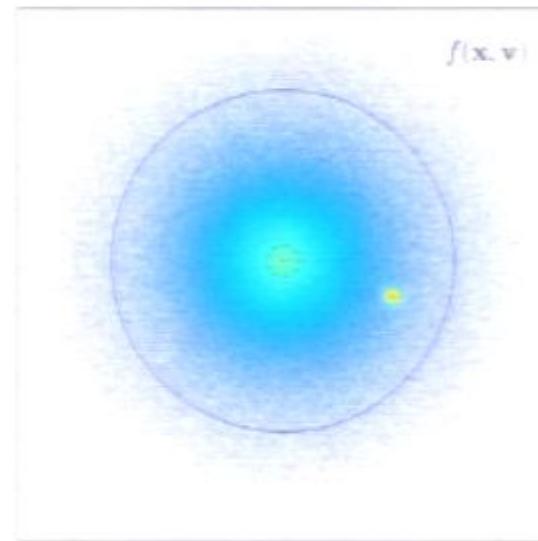




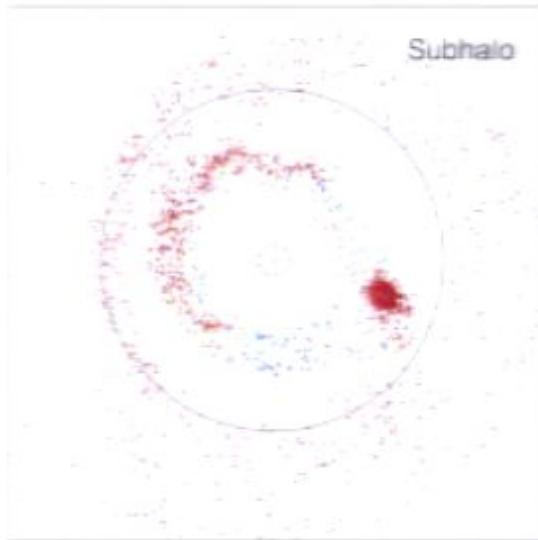




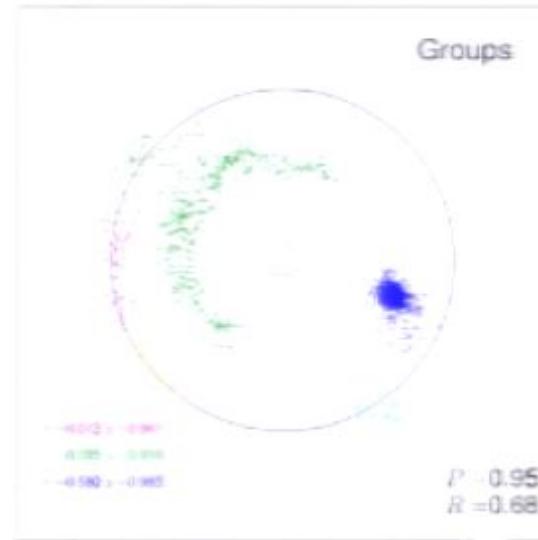
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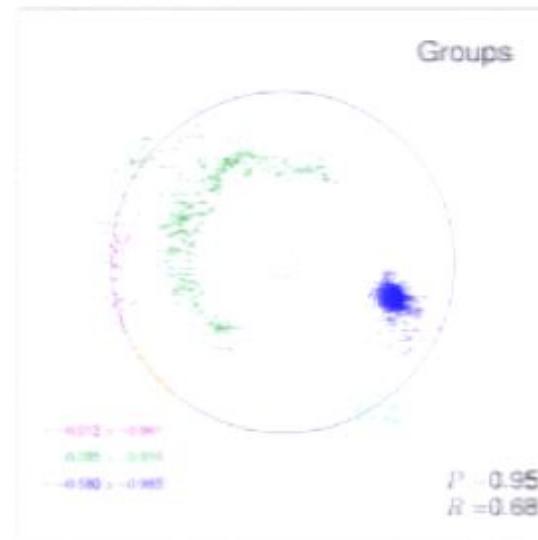
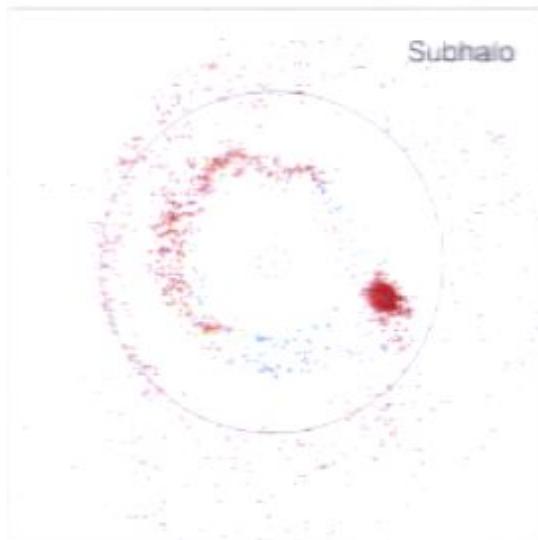
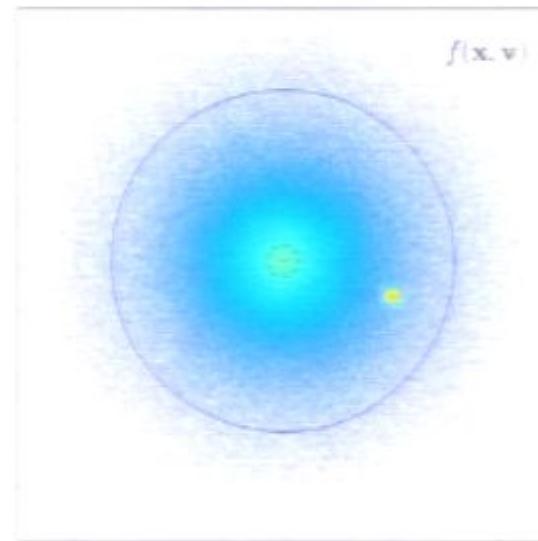
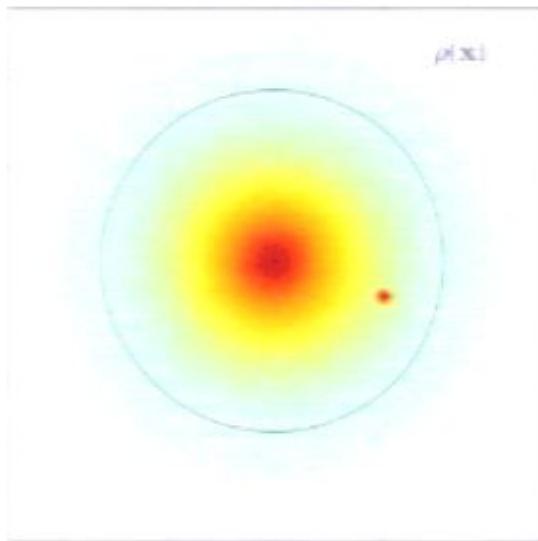
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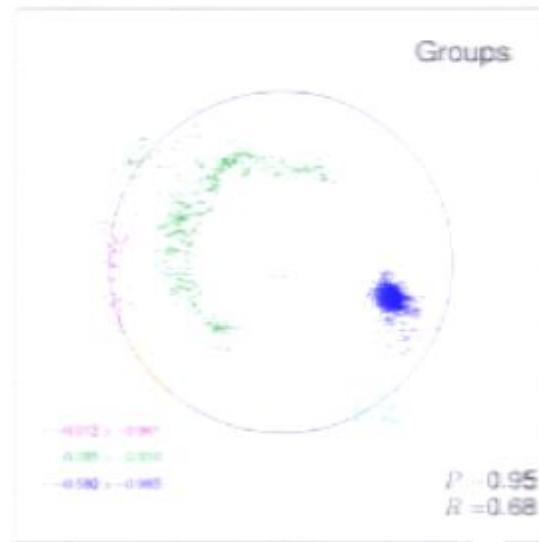
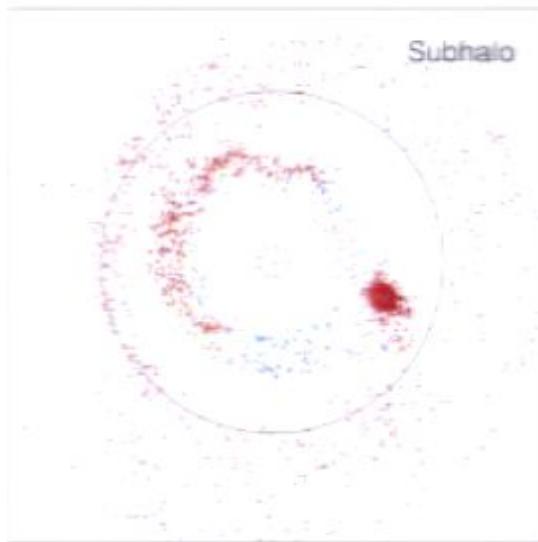
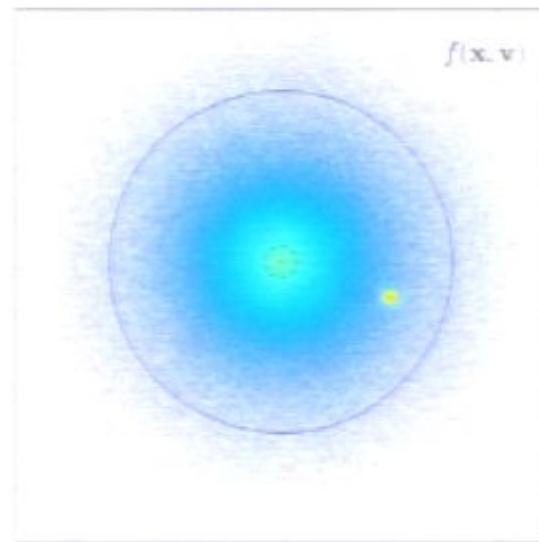
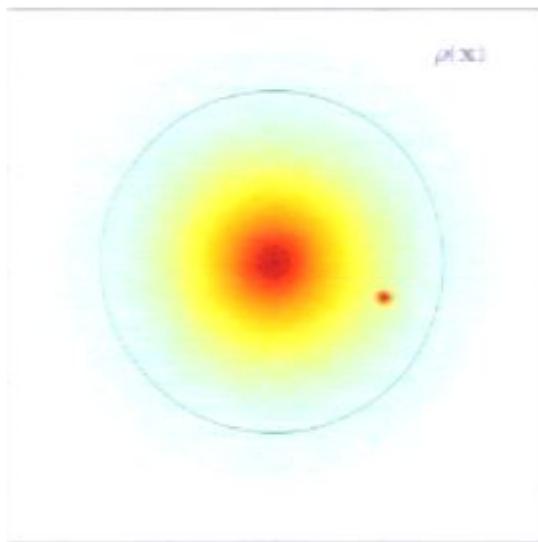


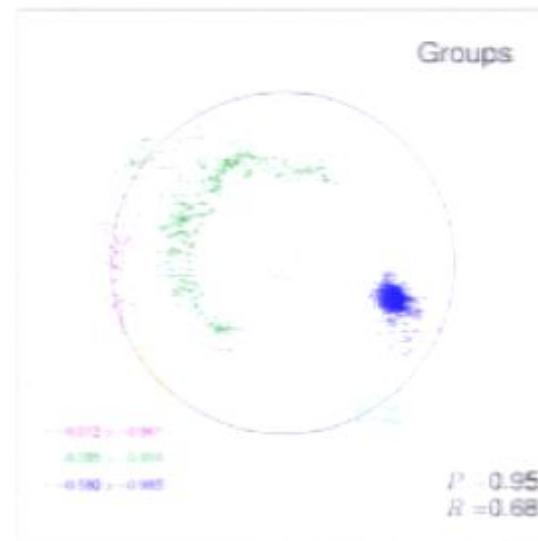
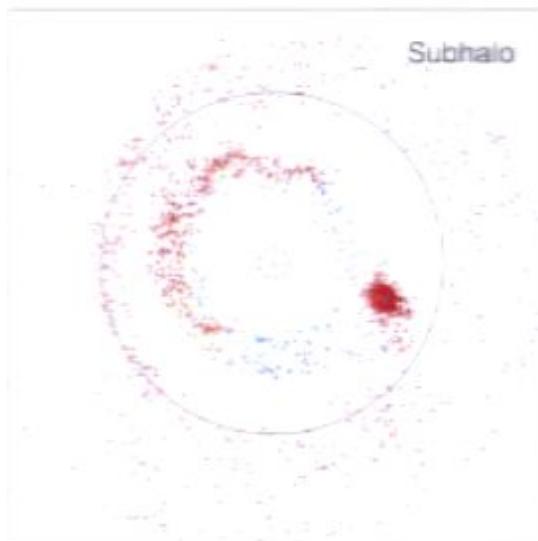
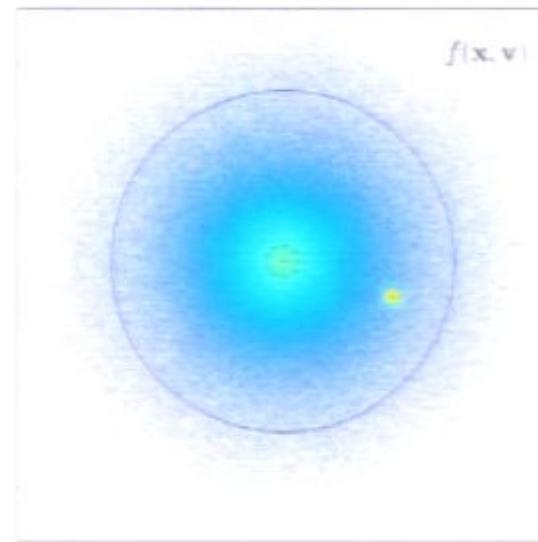
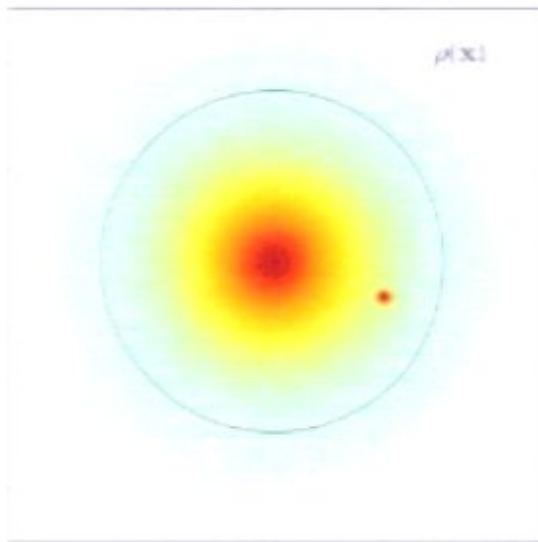
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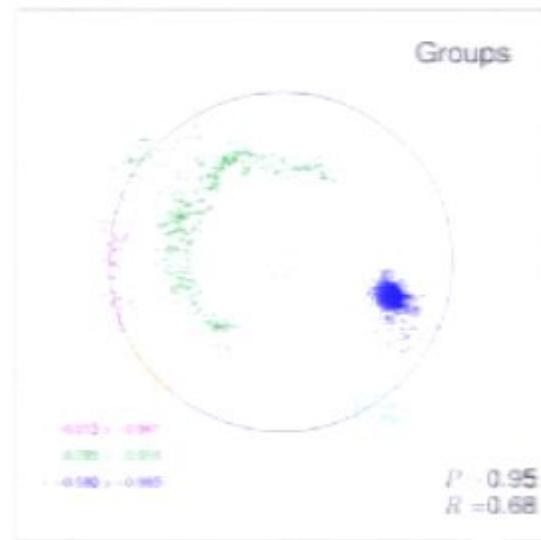
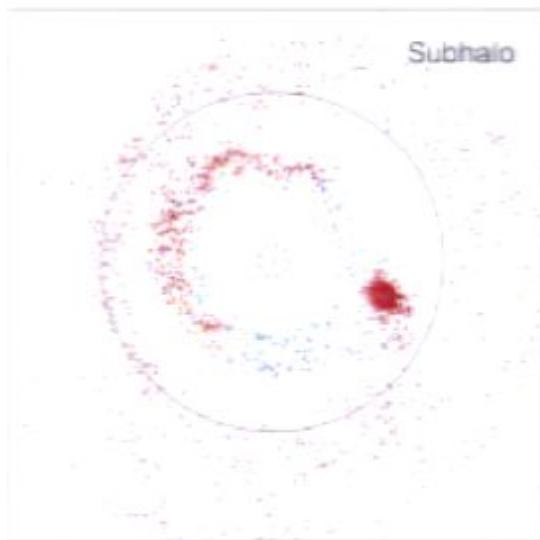
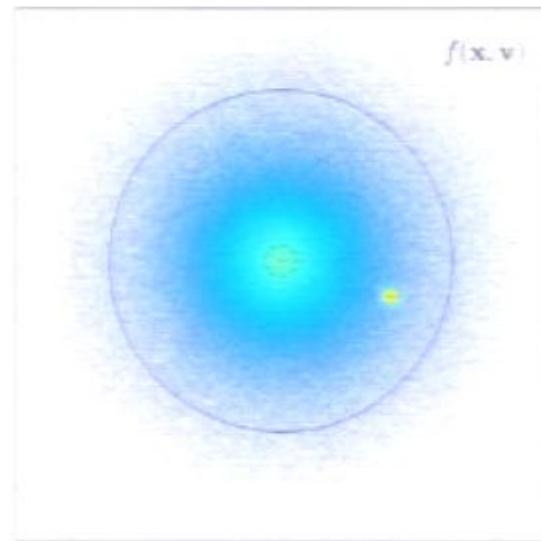
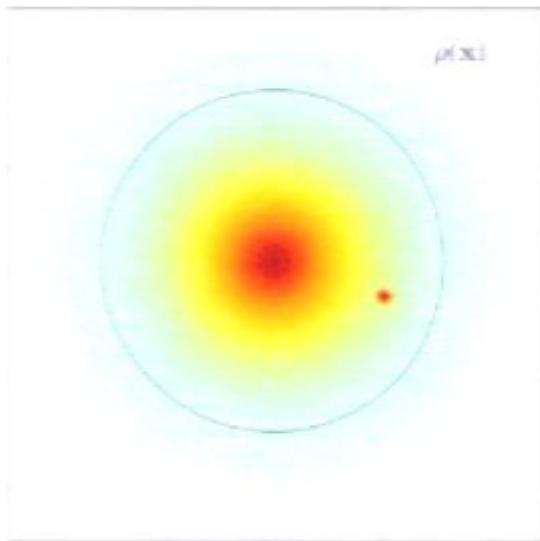


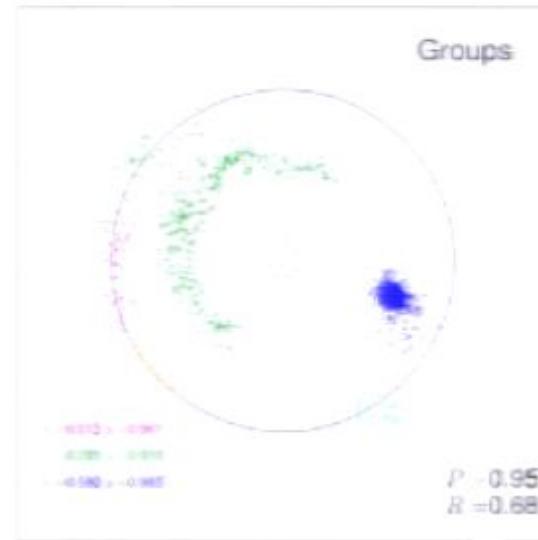
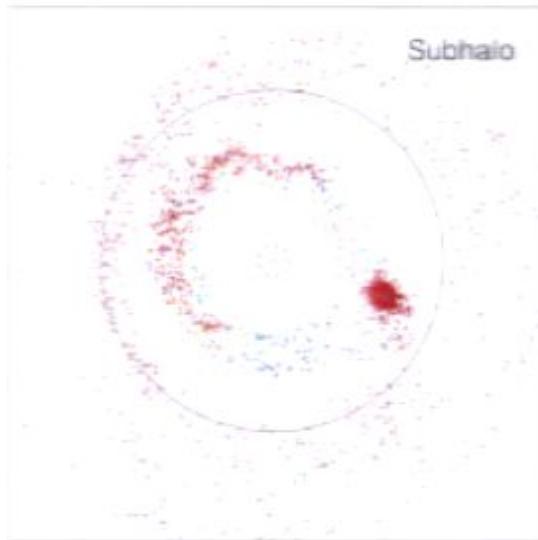
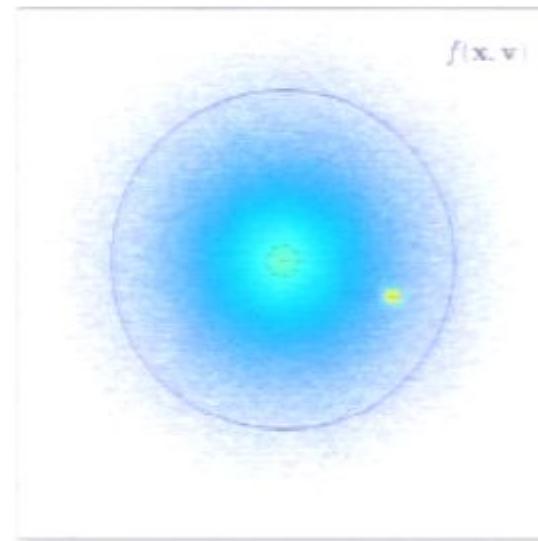
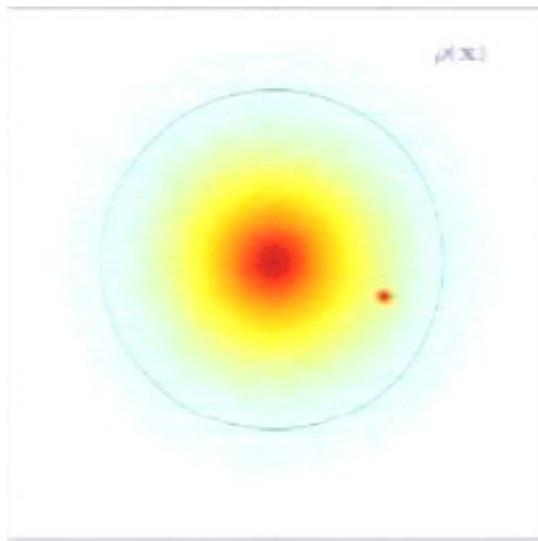
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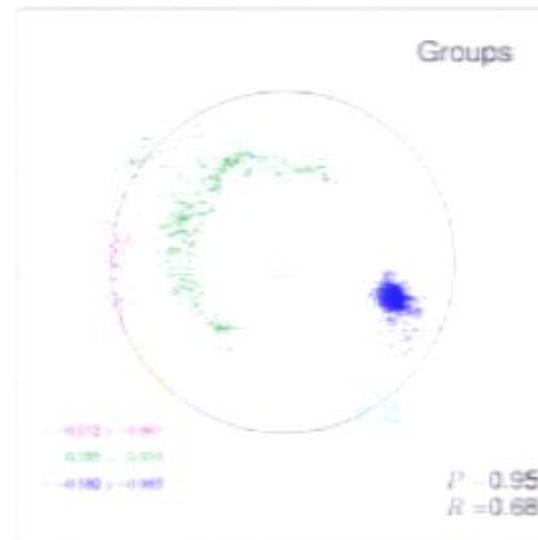
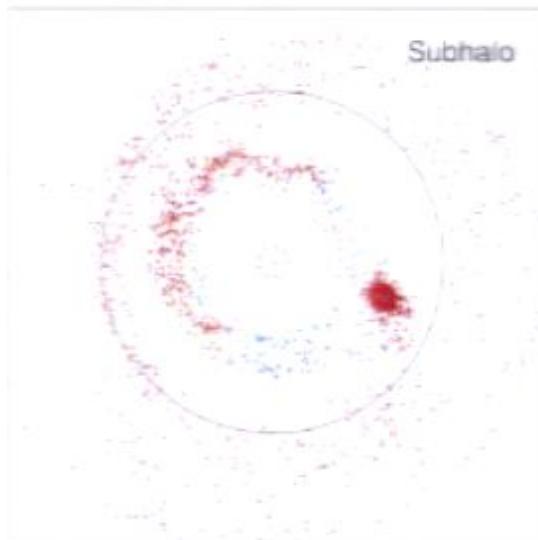
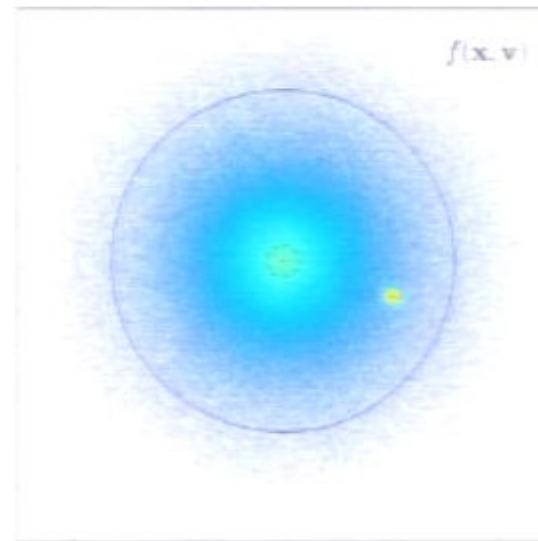
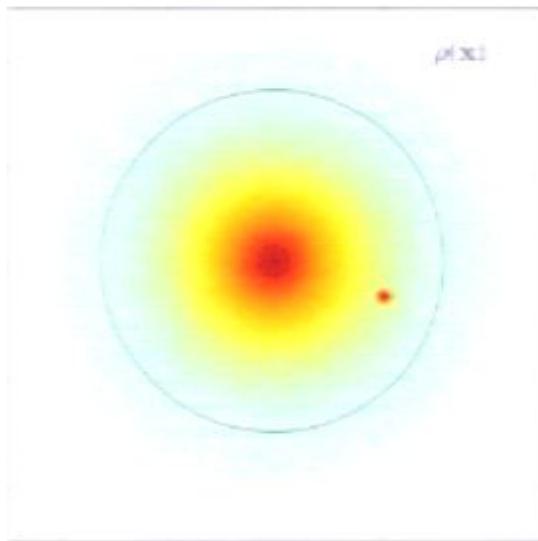


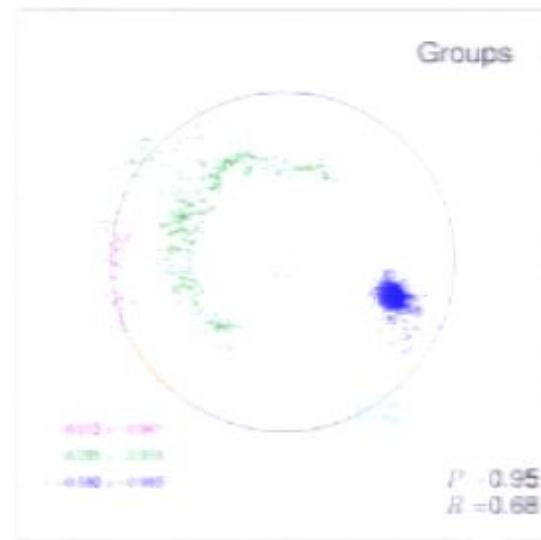
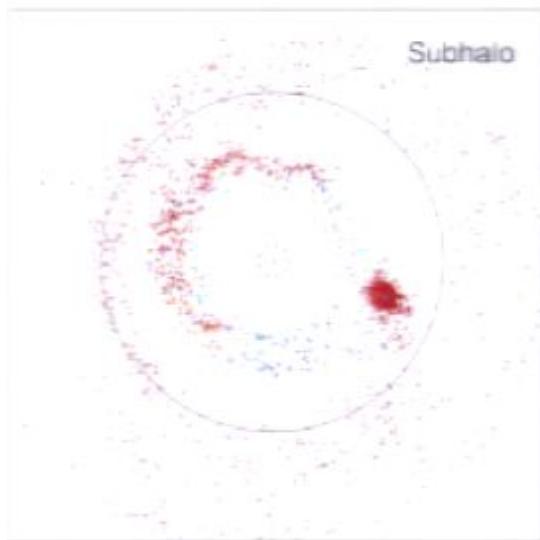
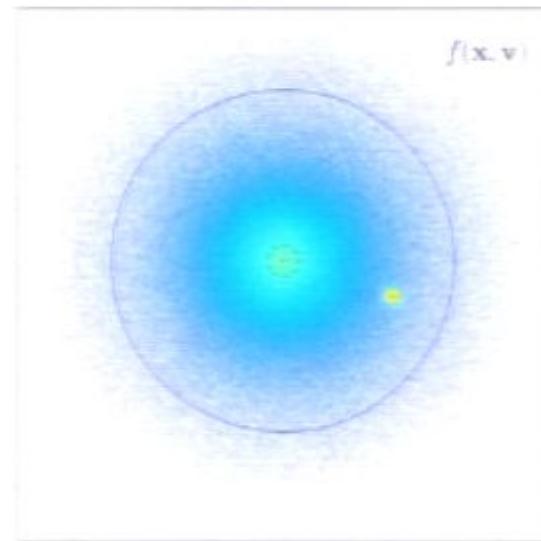
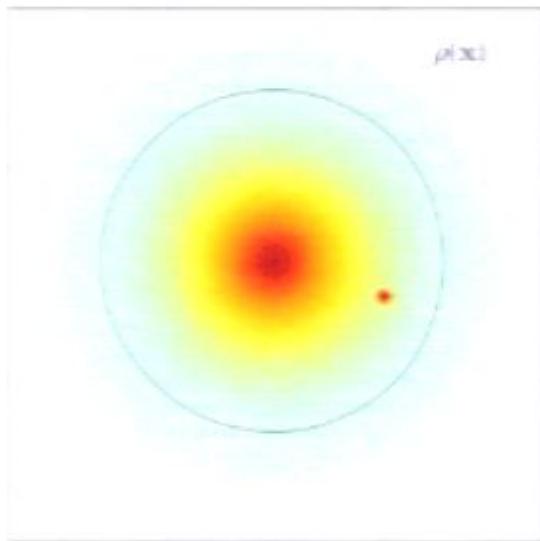


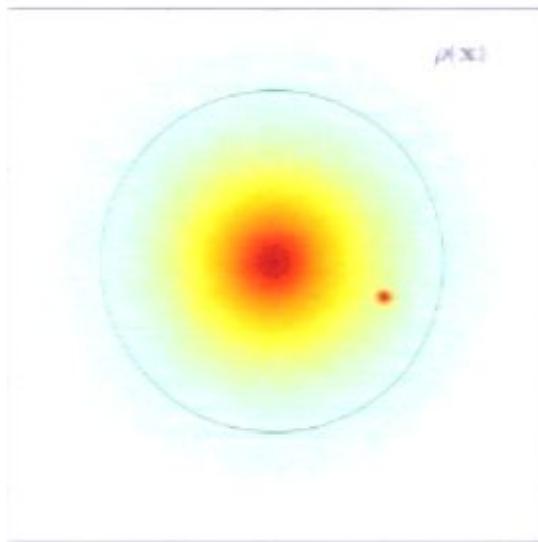




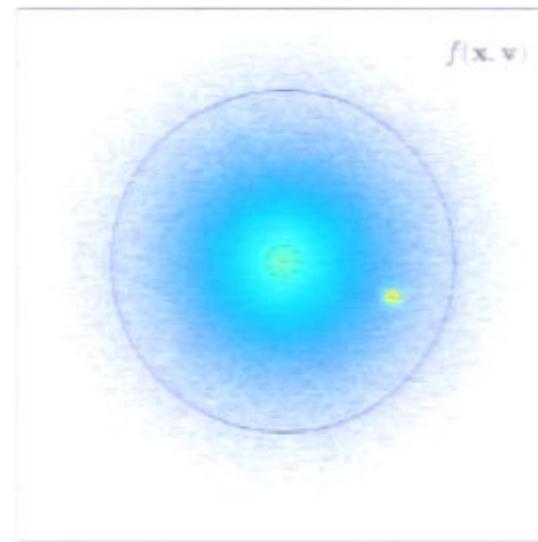




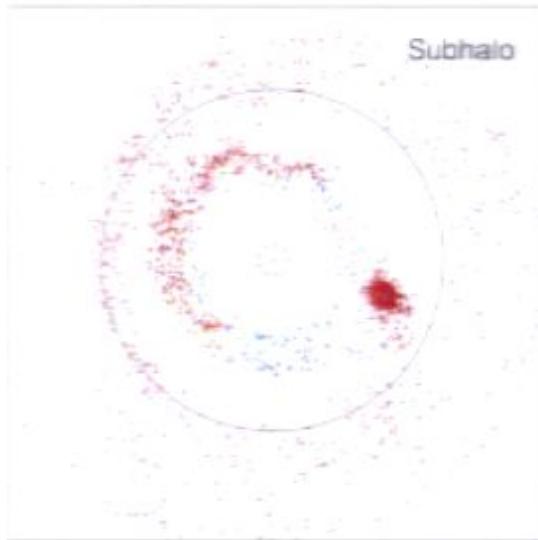




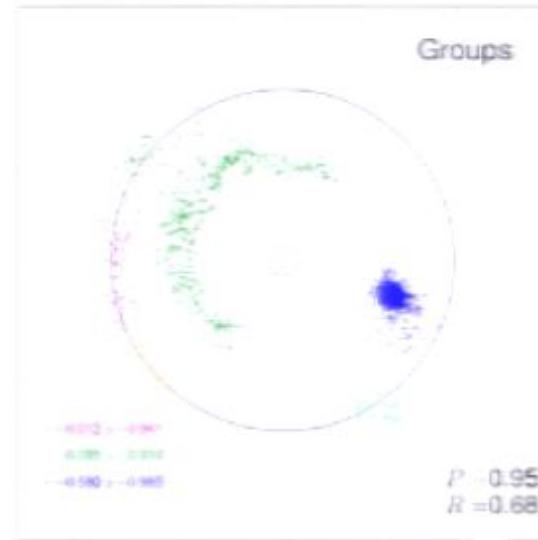
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$f(\mathbf{x}, \mathbf{v})$



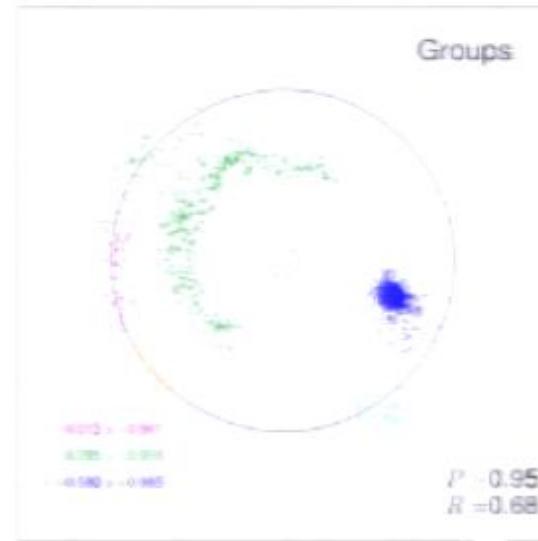
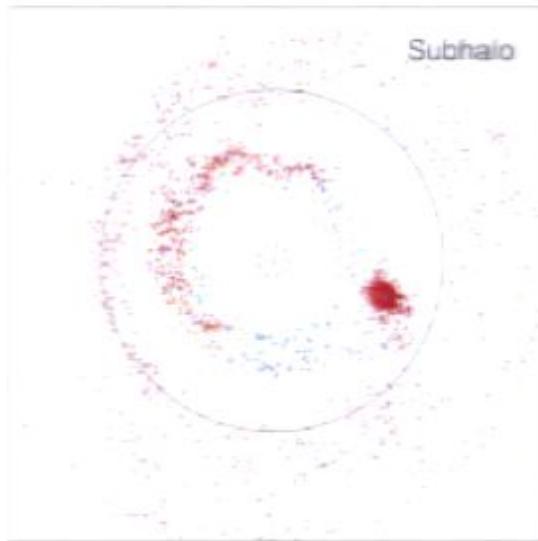
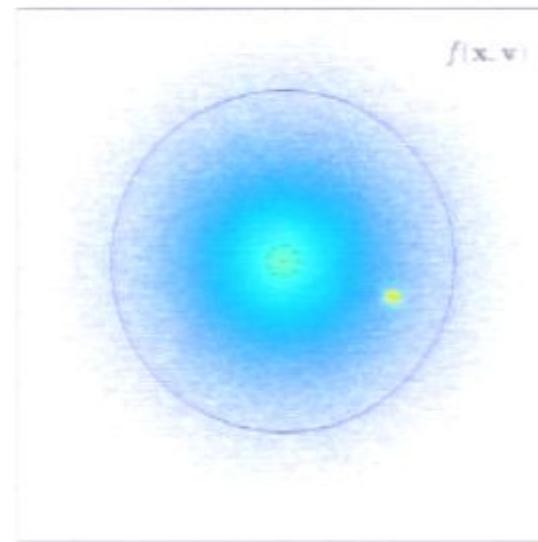
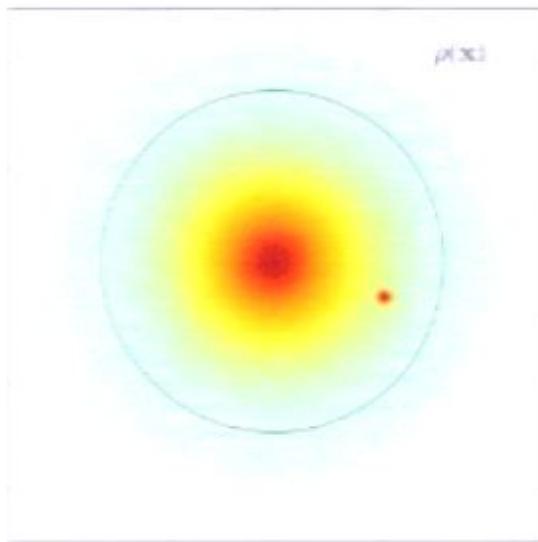
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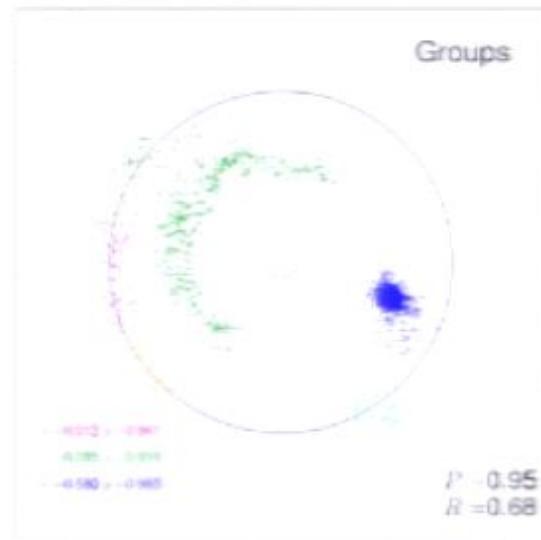
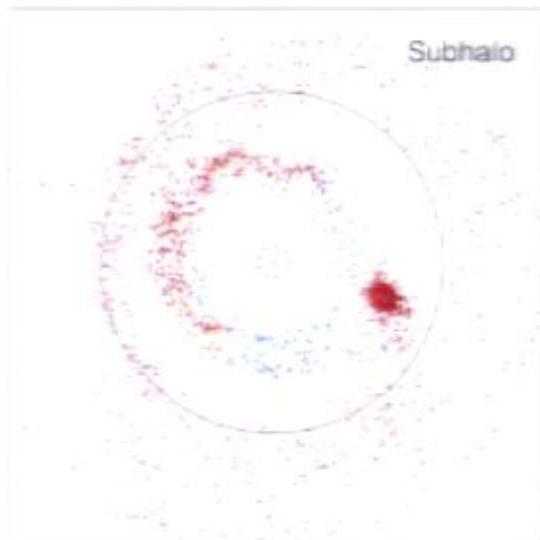
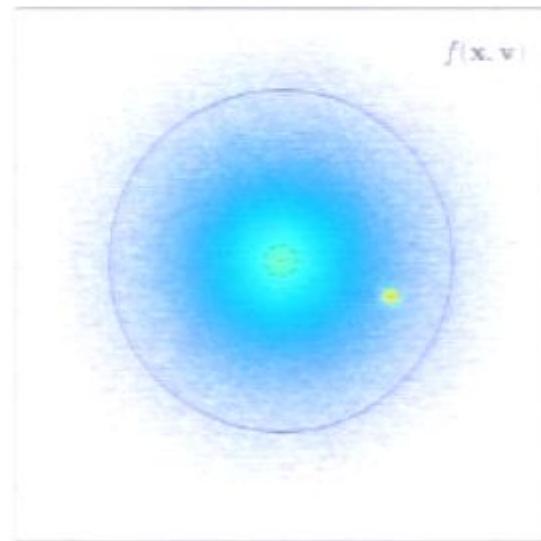
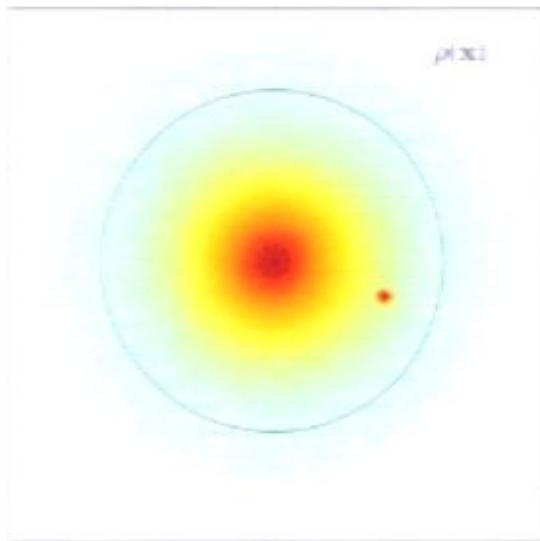


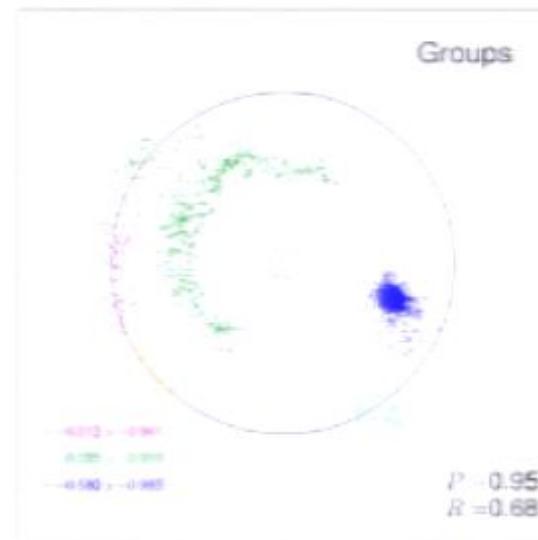
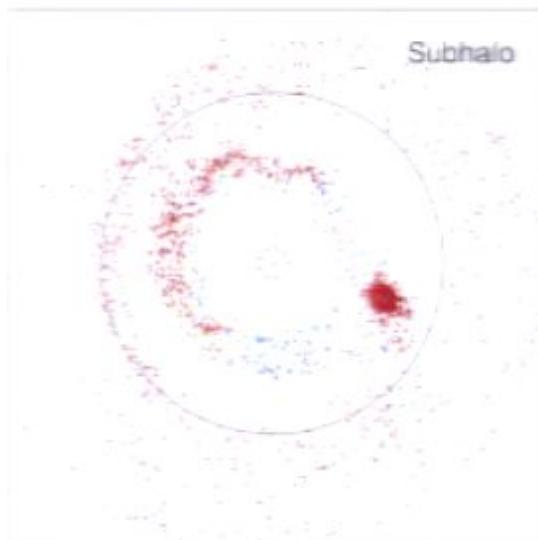
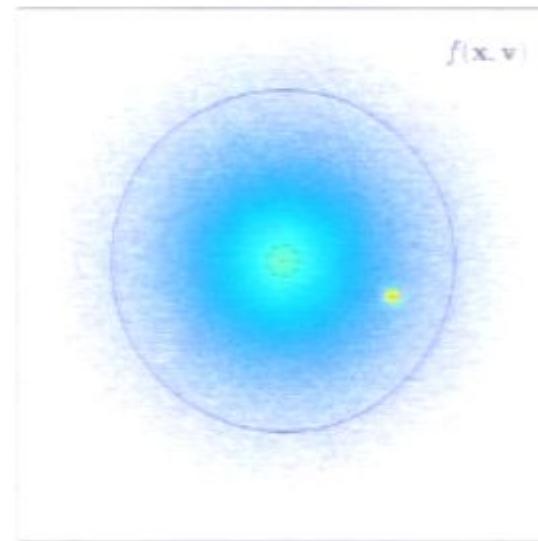
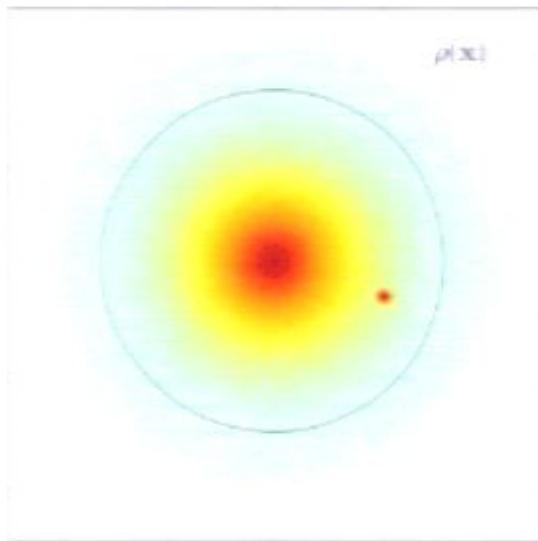
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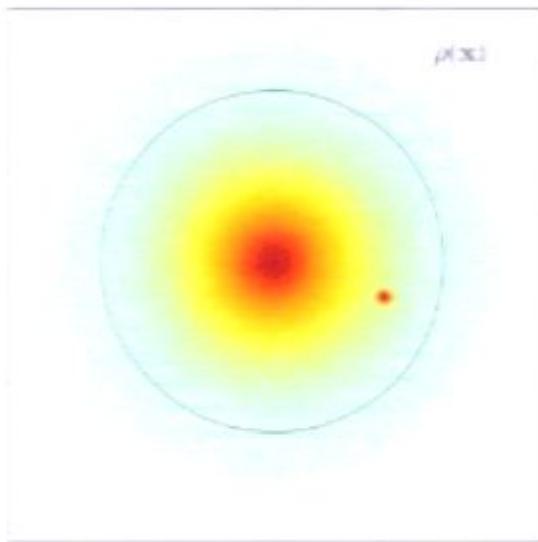
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$P = 0.95$
 $R = 0.68$

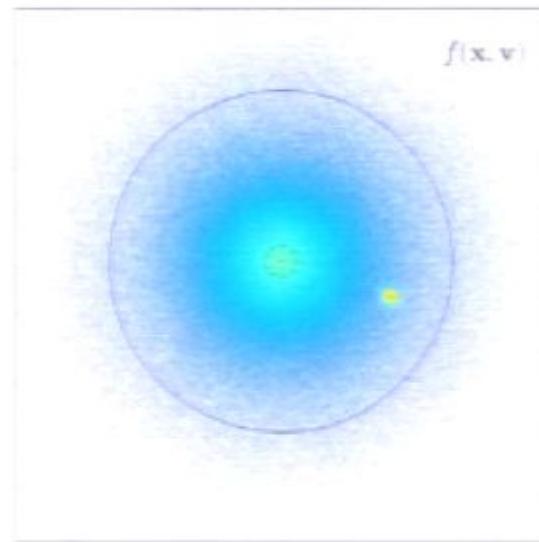




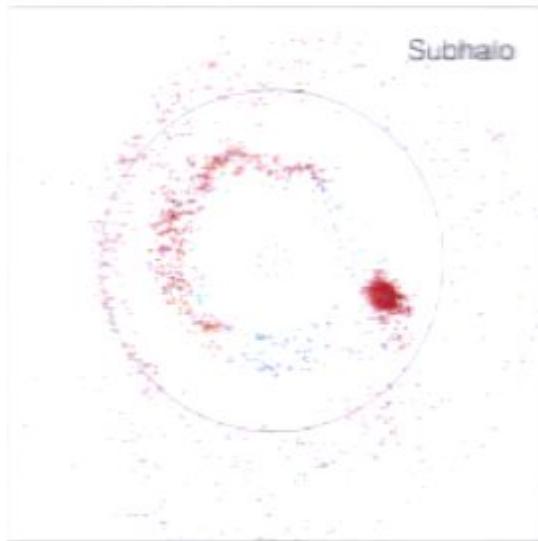




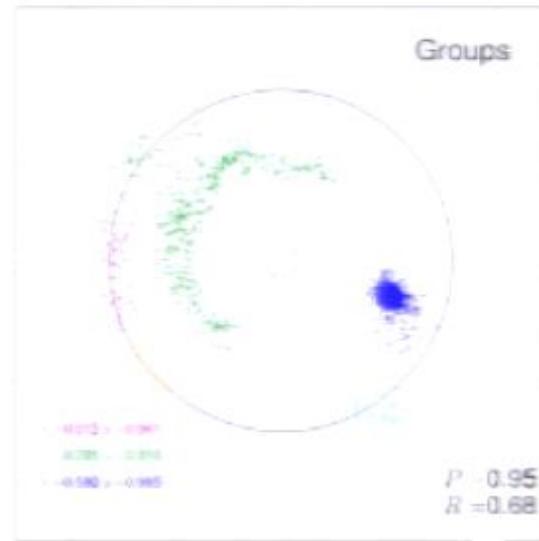
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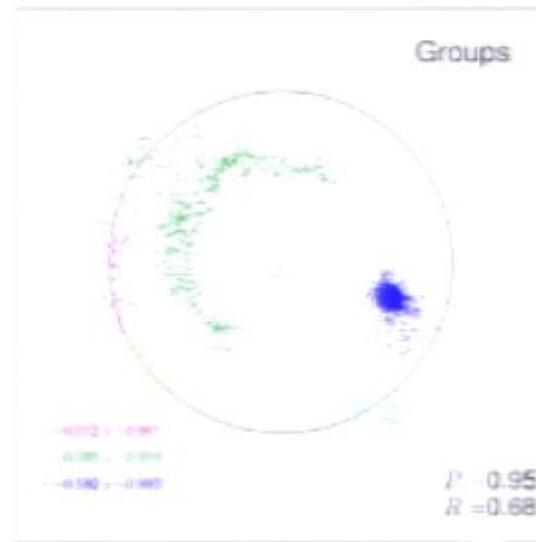
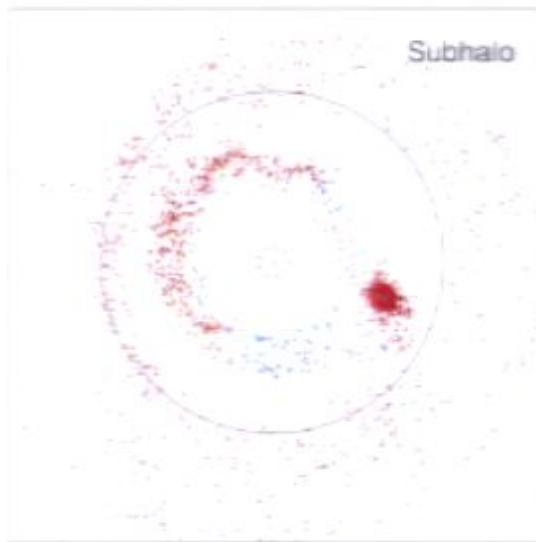
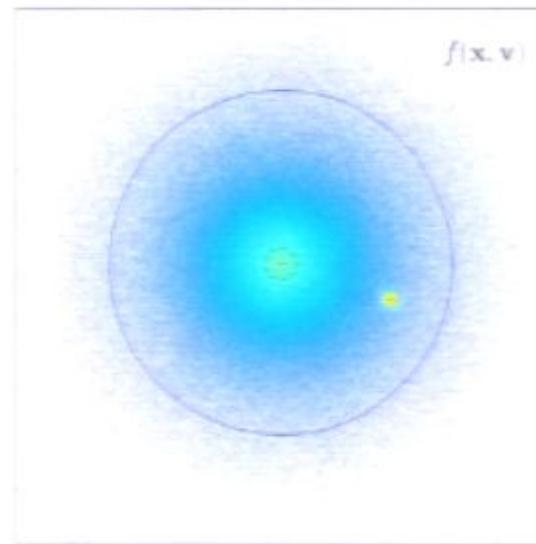
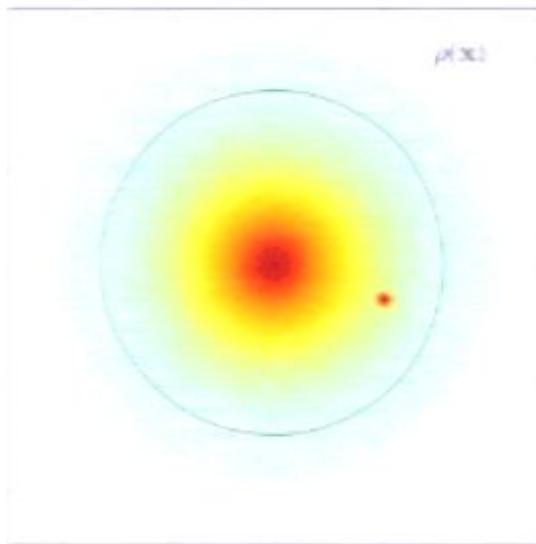


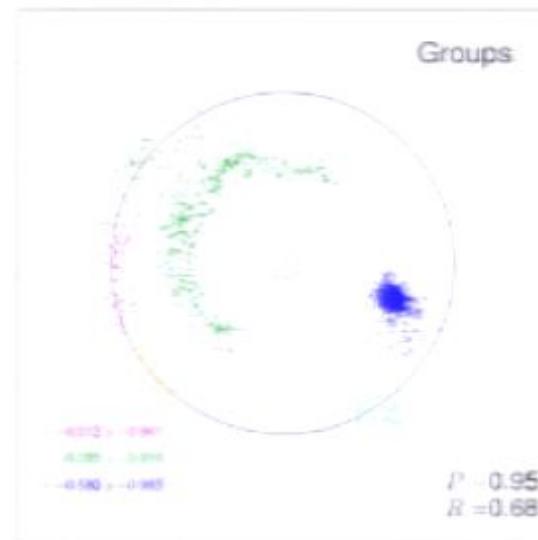
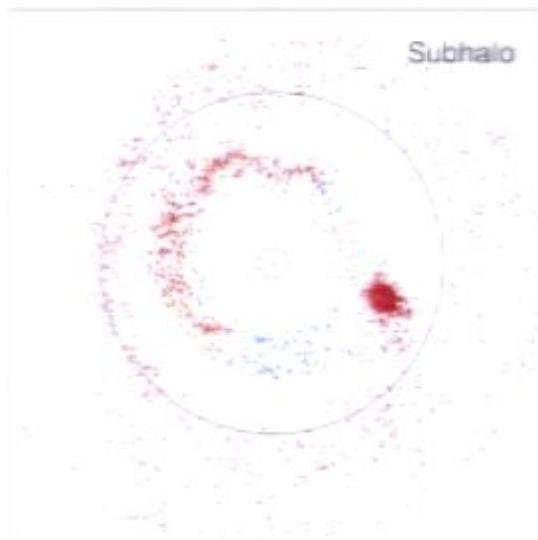
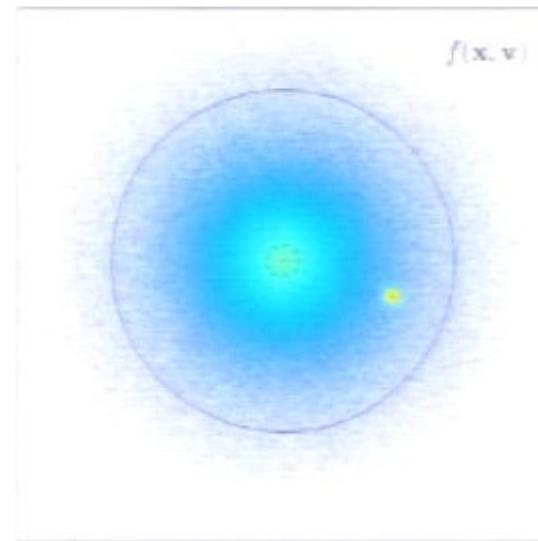
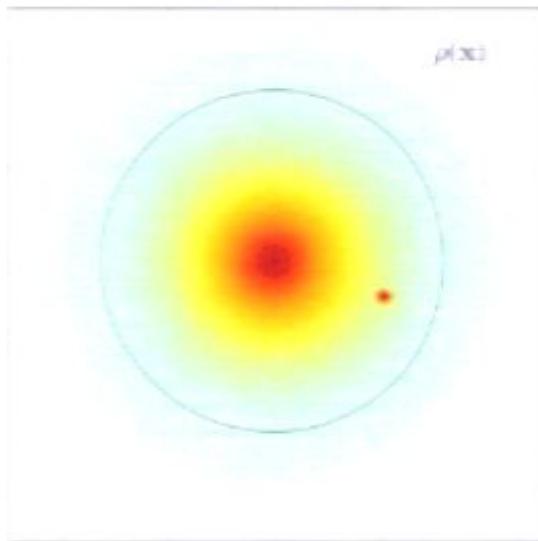
$f(\mathbf{x}, \mathbf{v})$

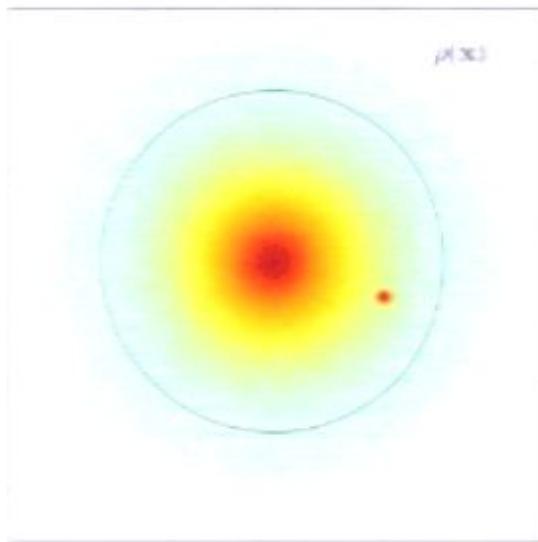


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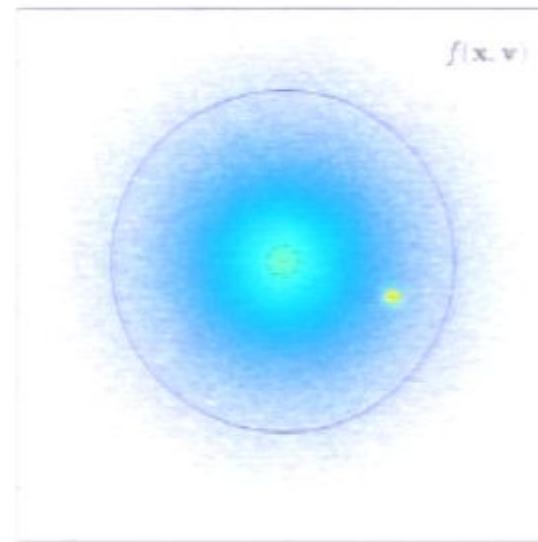




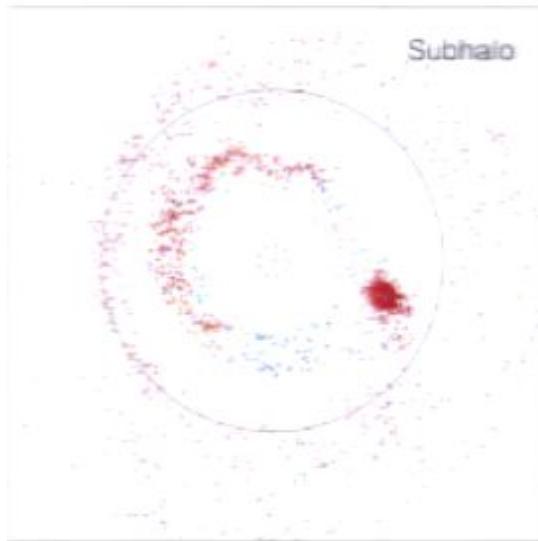




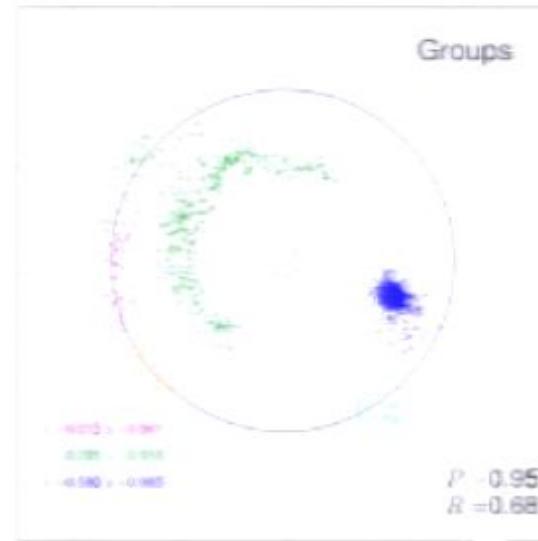
$\rho(\mathbf{x})$



$f(\mathbf{x}, \mathbf{v})$

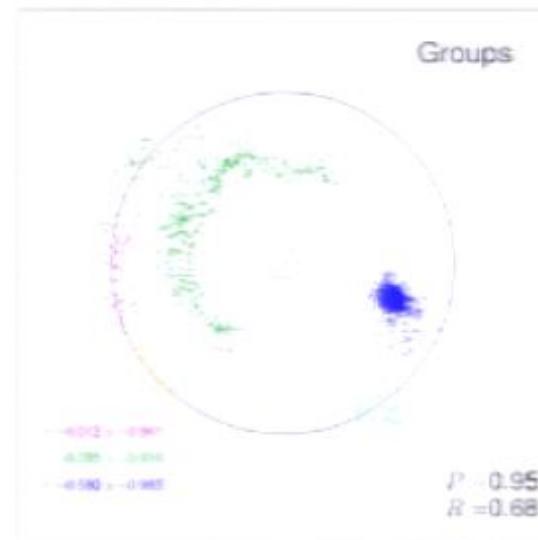
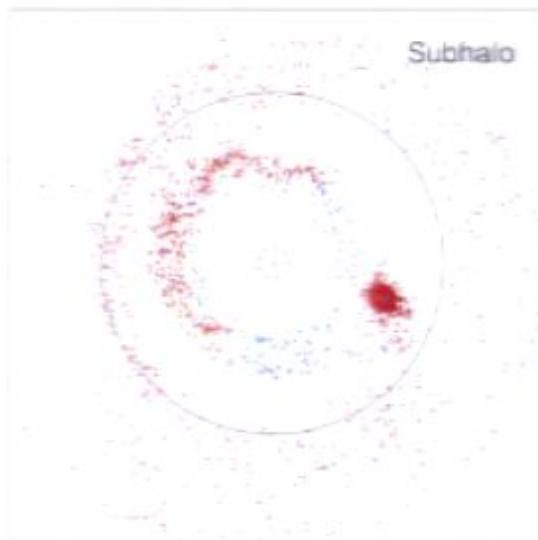
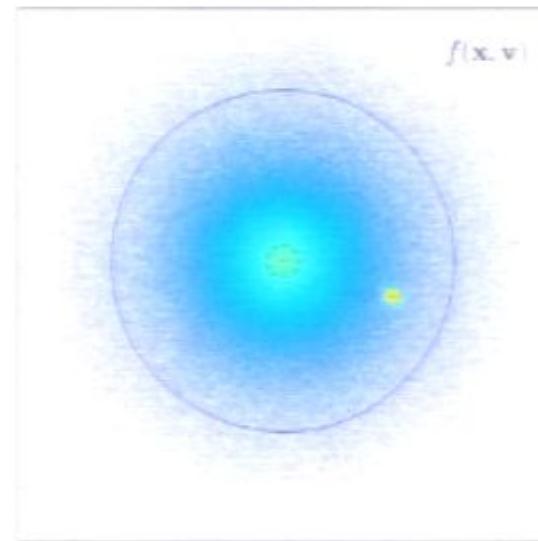
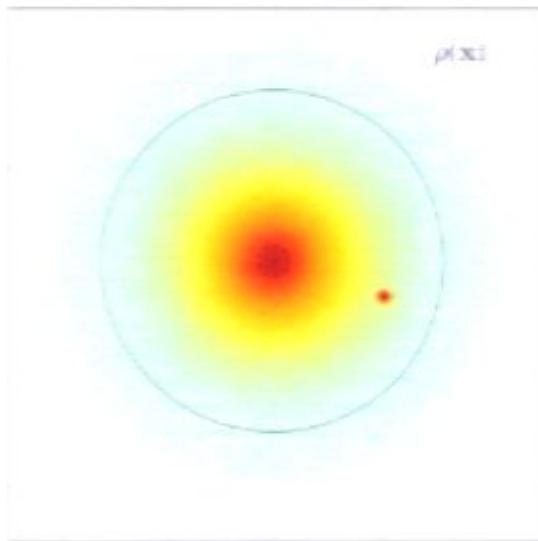


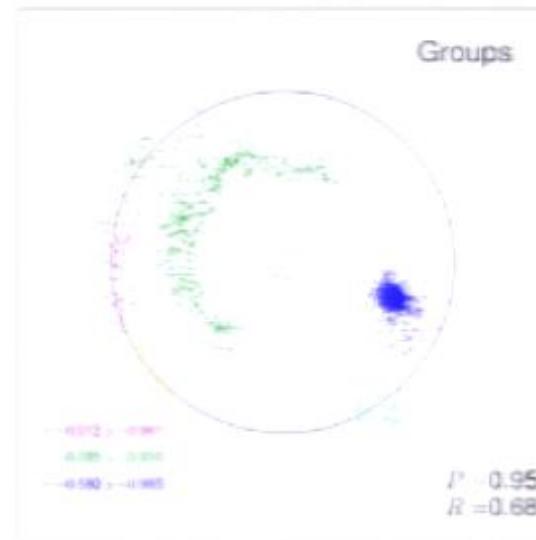
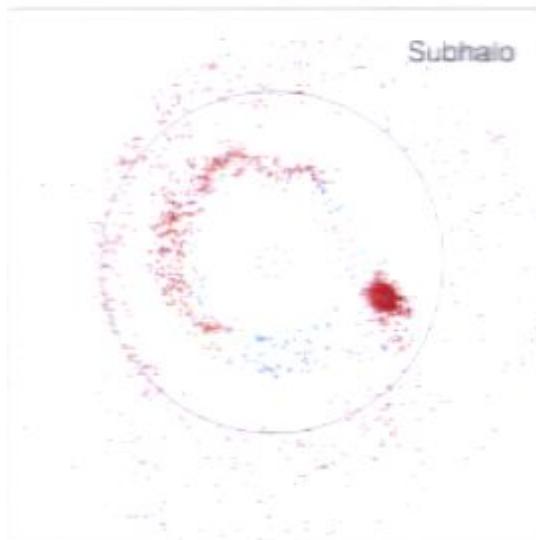
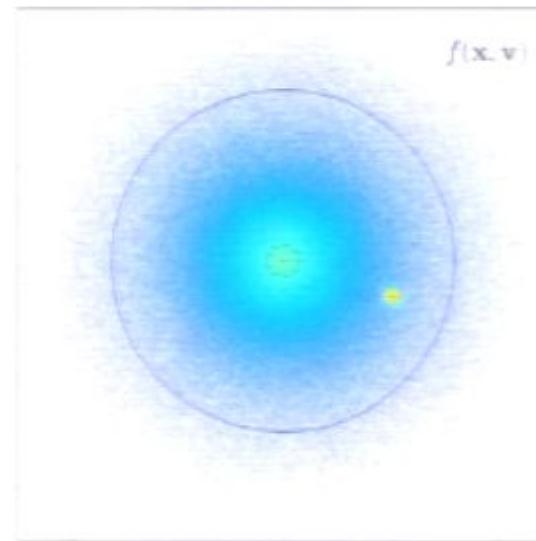
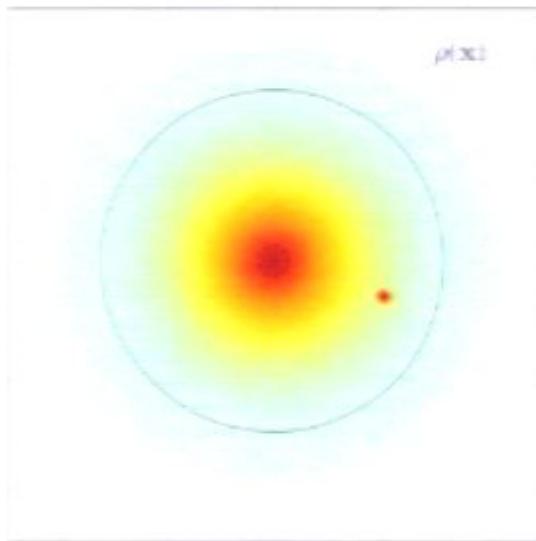
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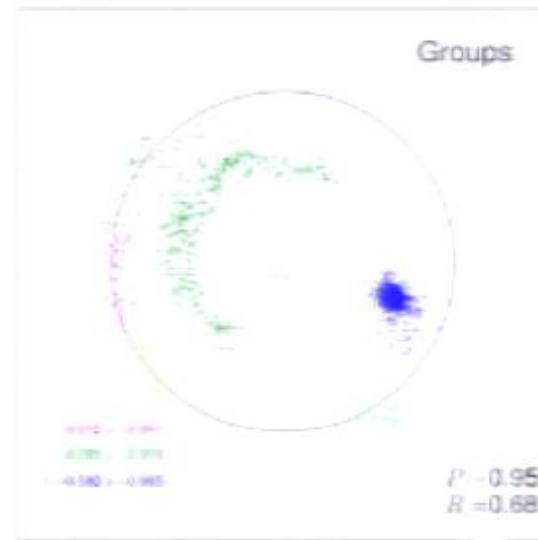
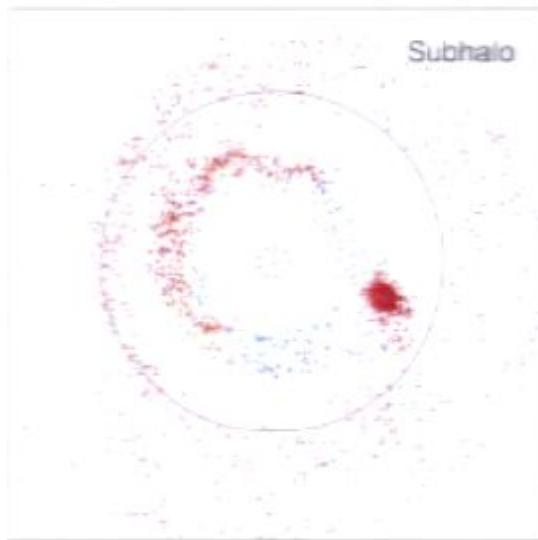
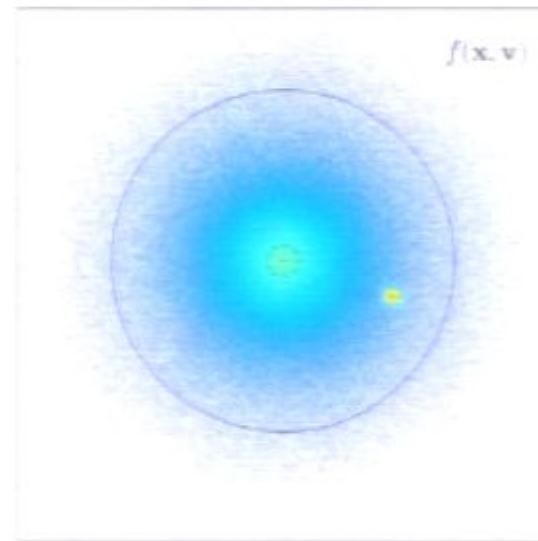
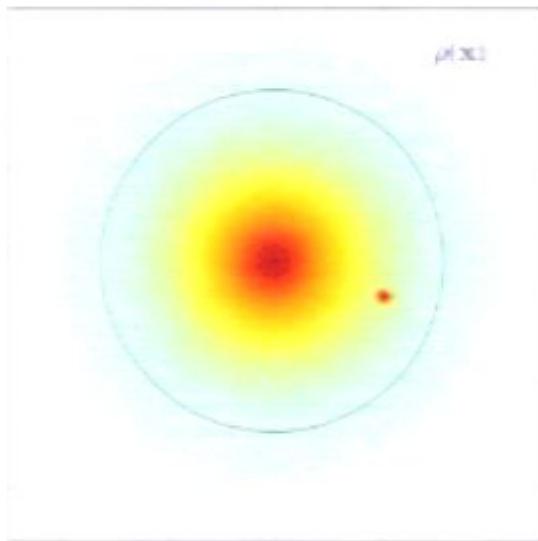


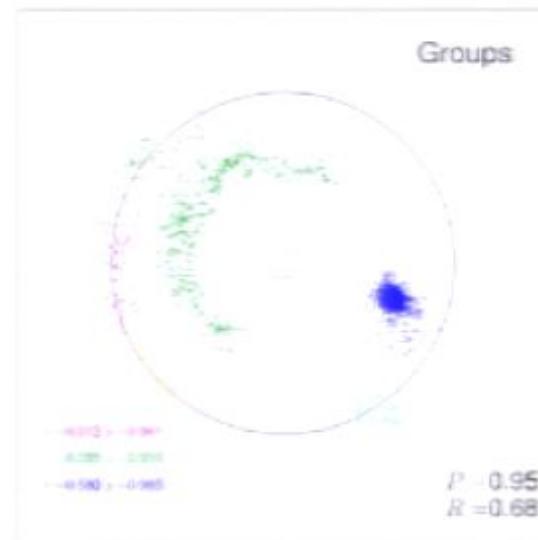
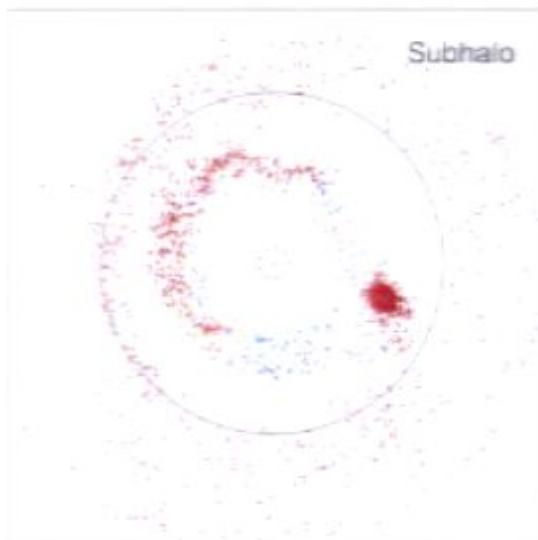
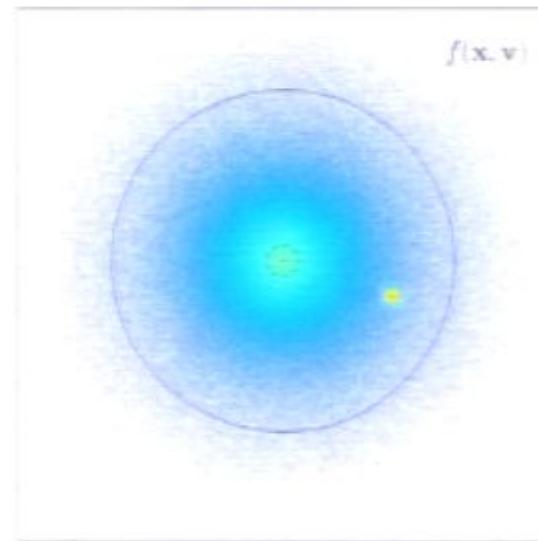
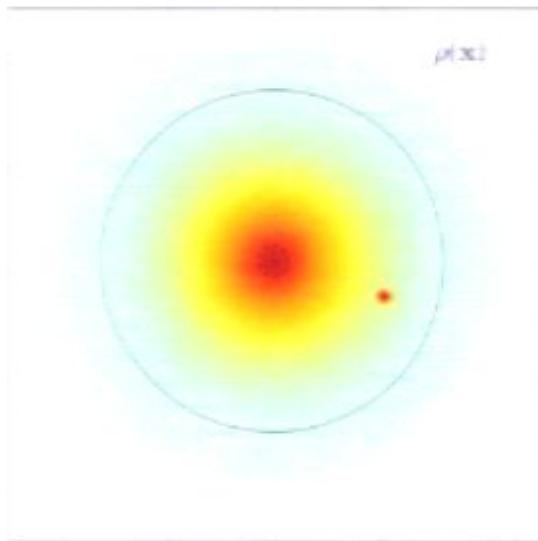
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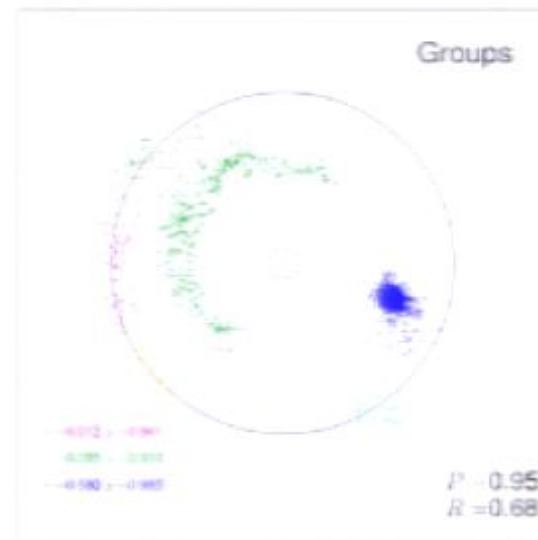
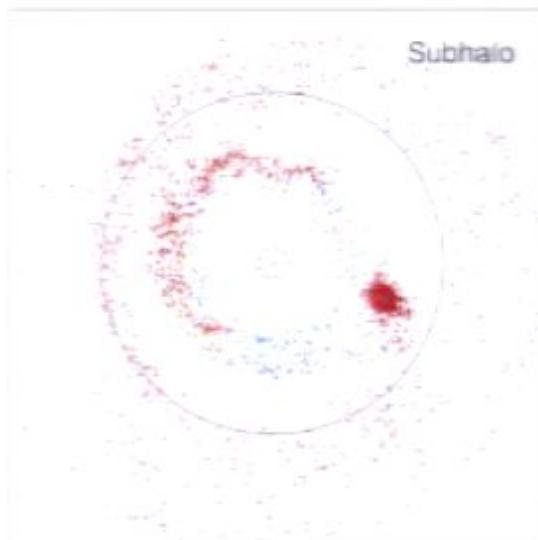
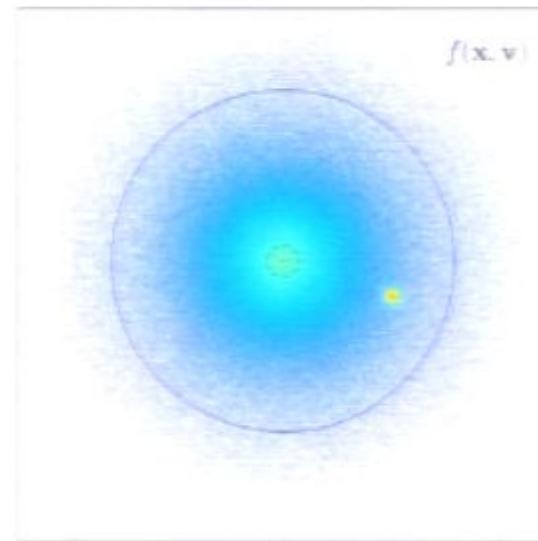
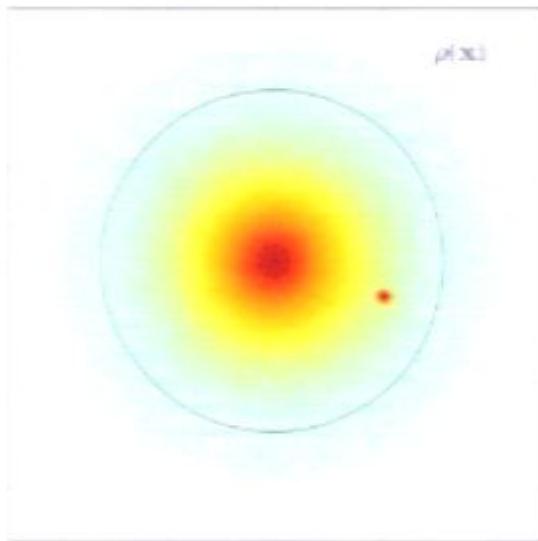
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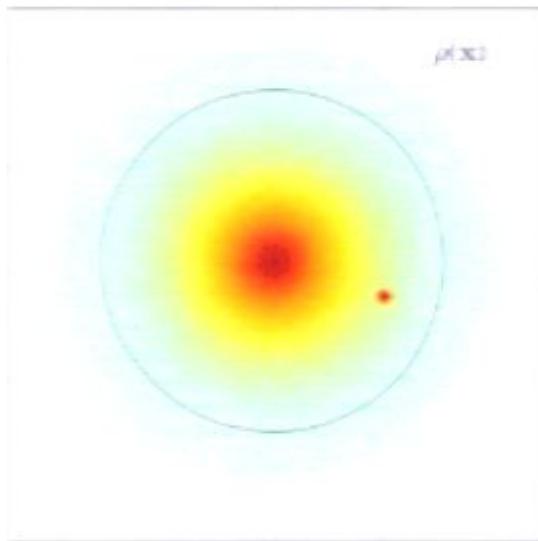




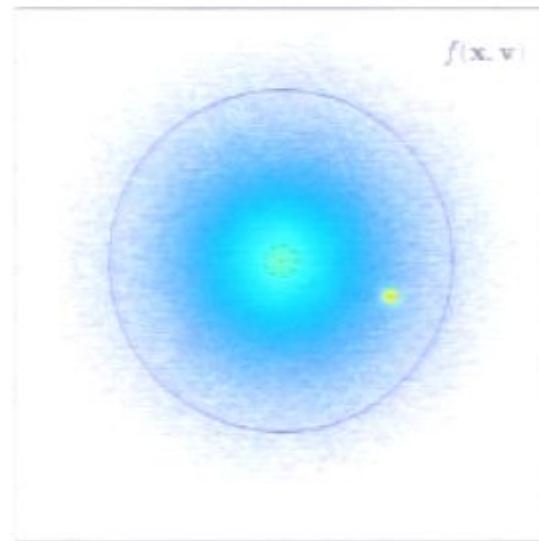




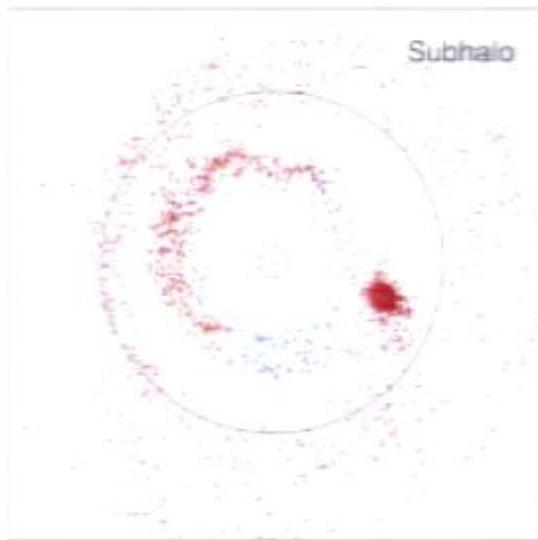




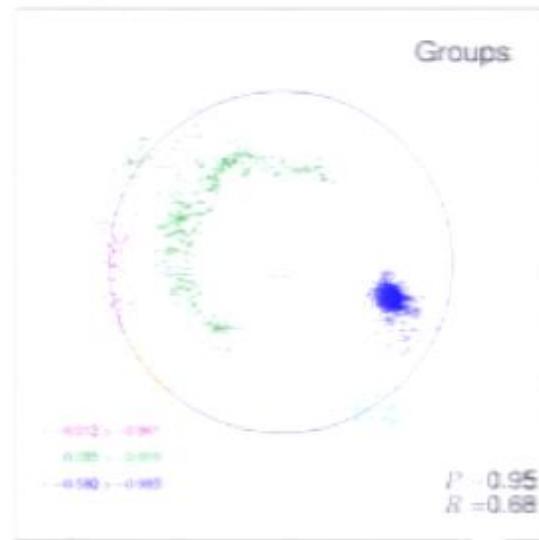
$\rho(\mathbf{x})$



$f(\mathbf{x}, \mathbf{v})$

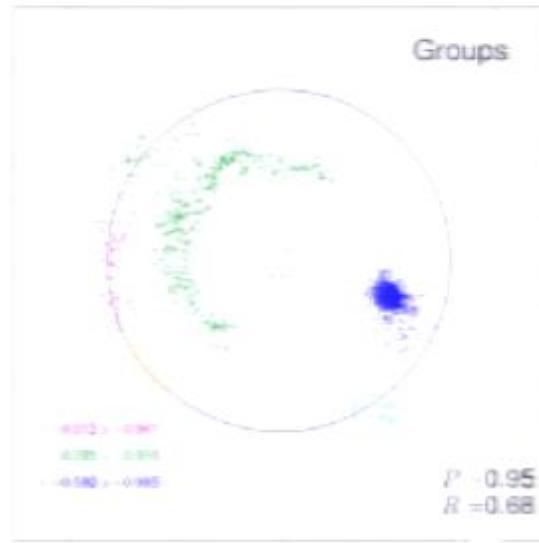
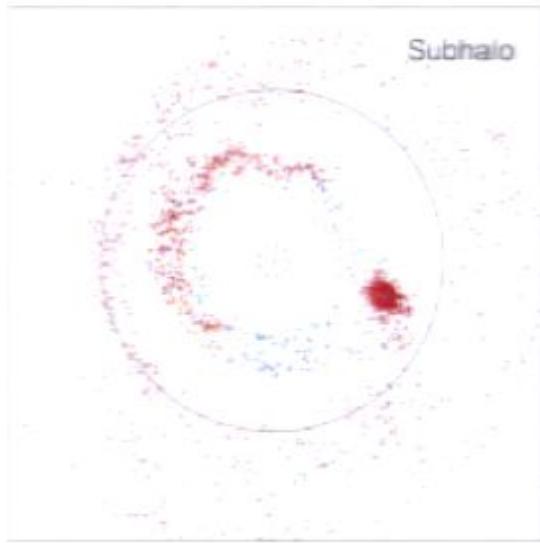
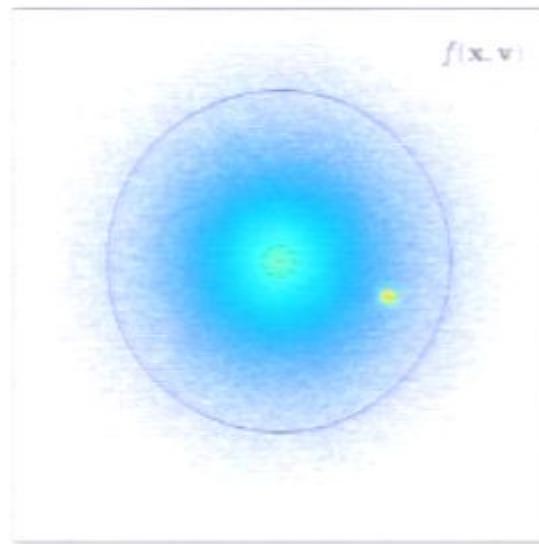
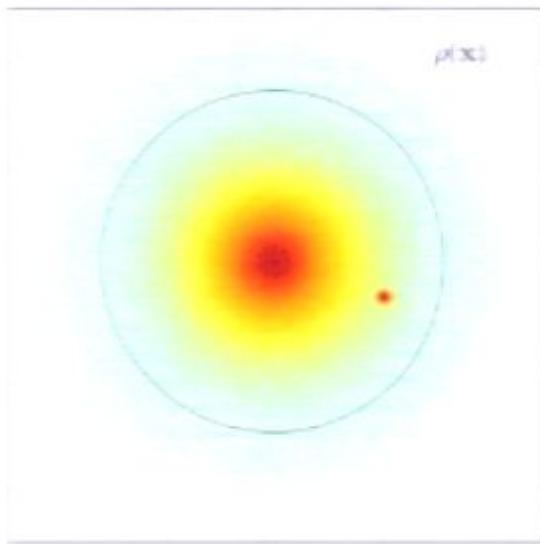


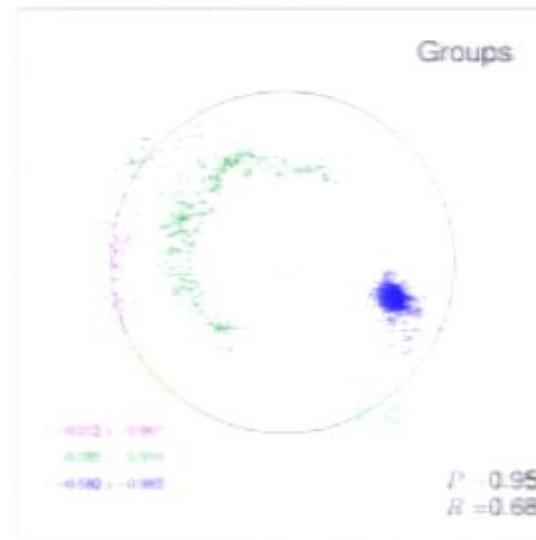
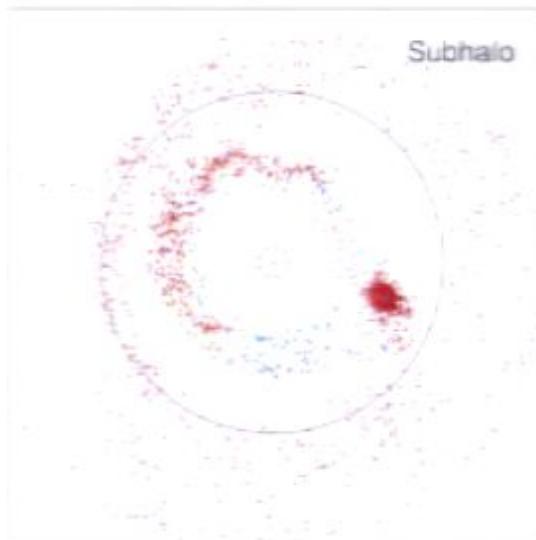
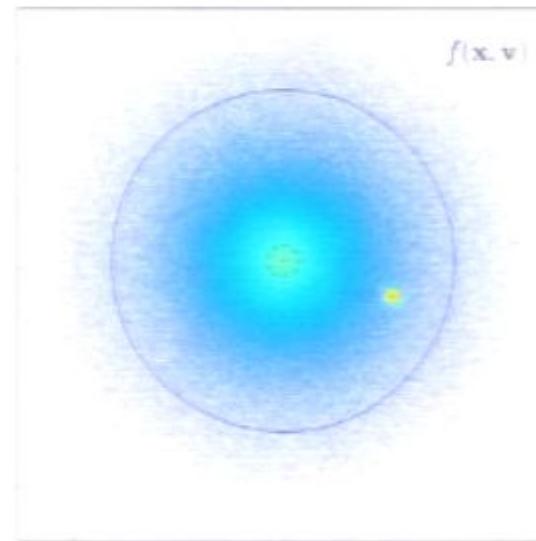
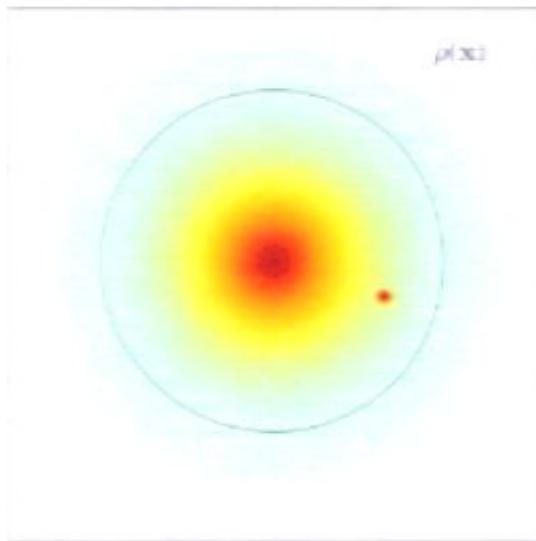
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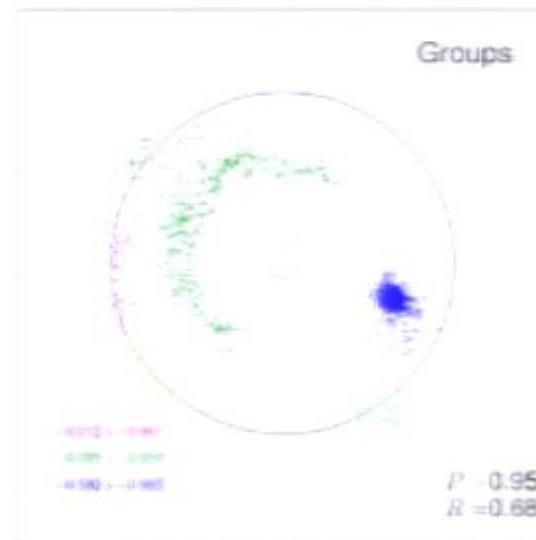
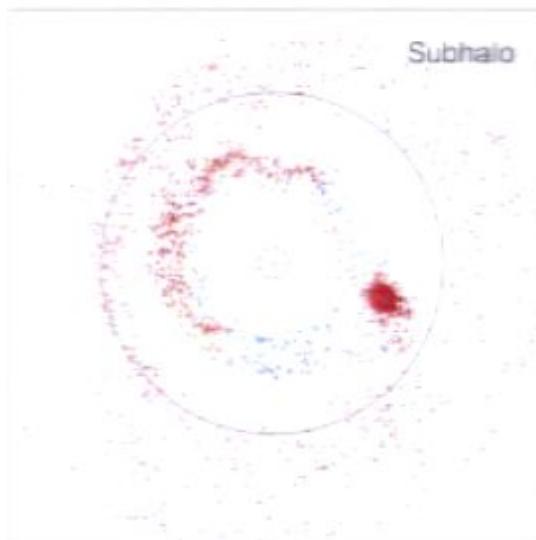
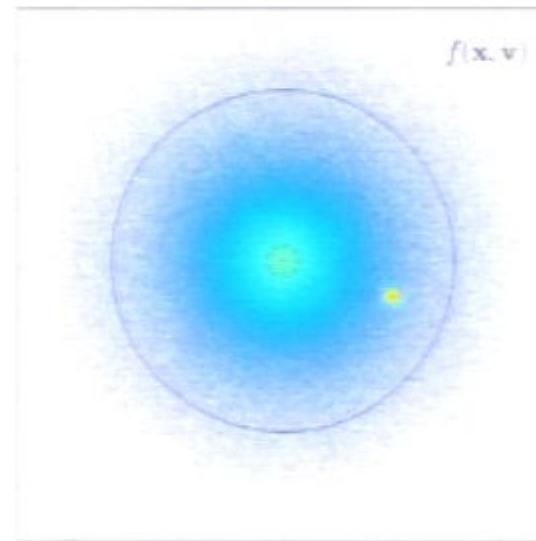
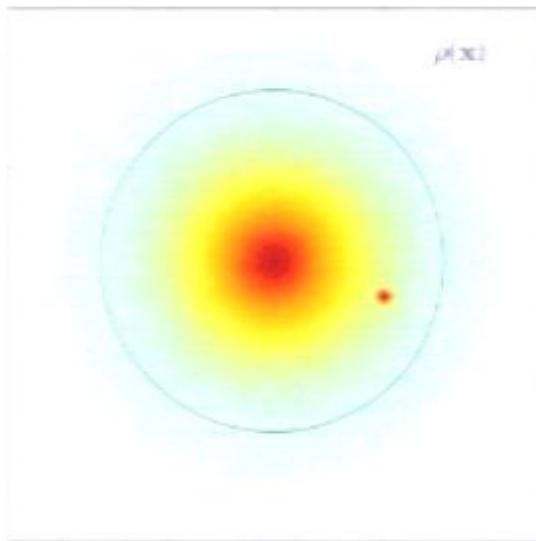


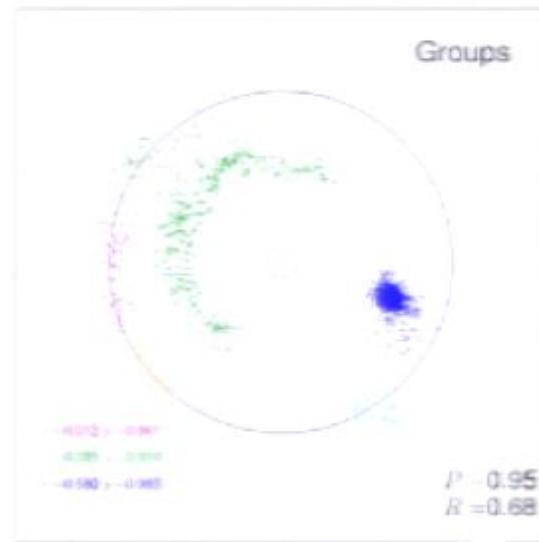
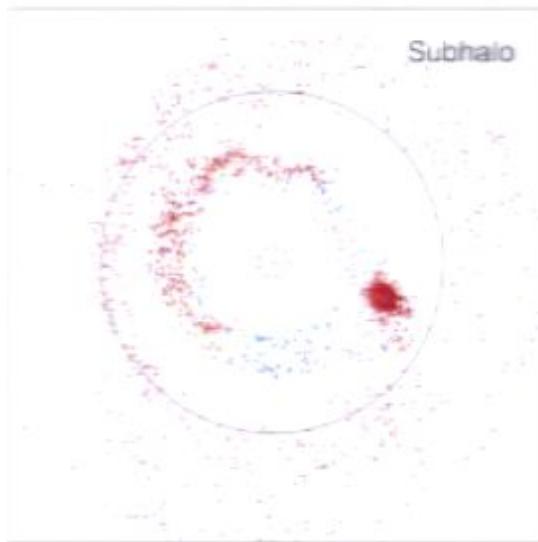
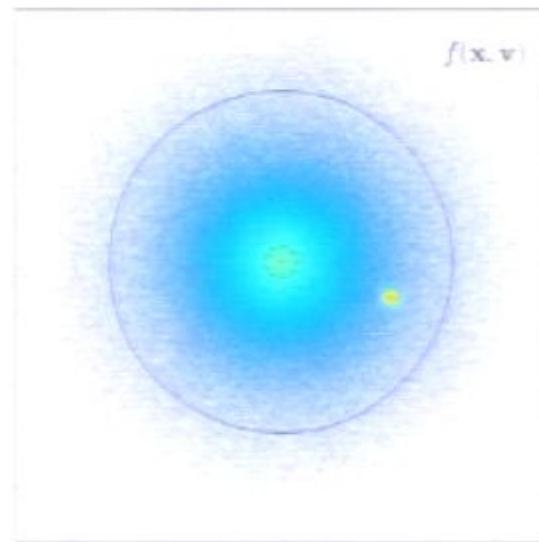
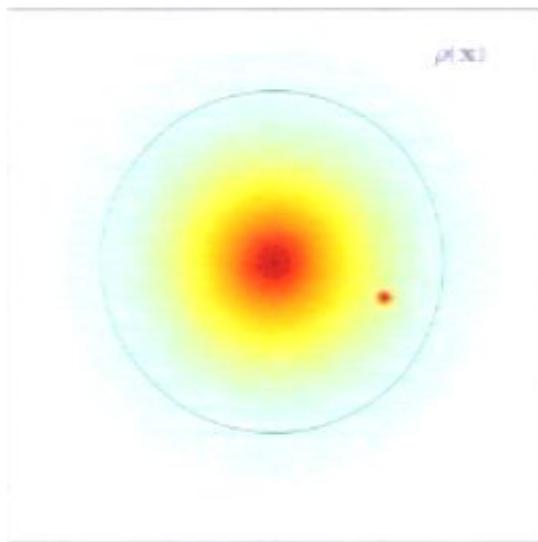
Groups

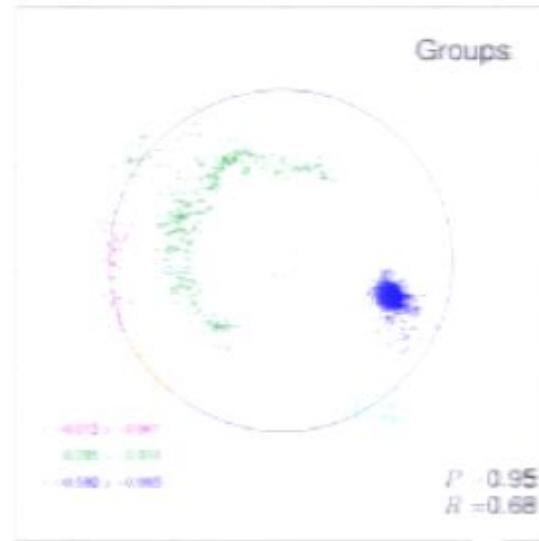
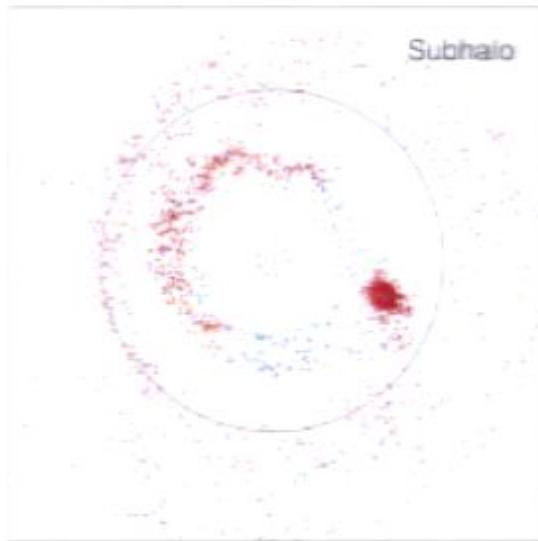
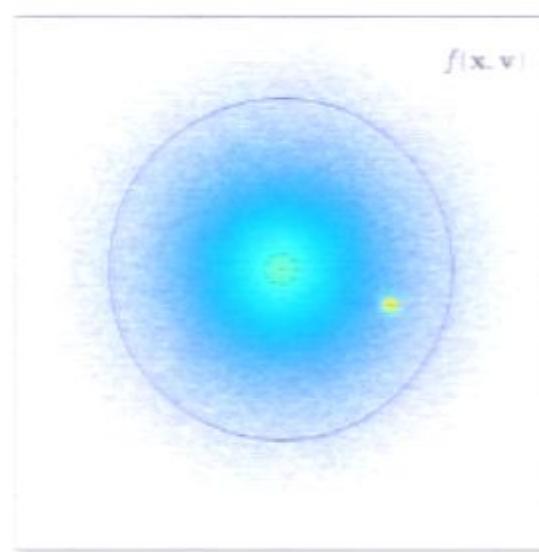
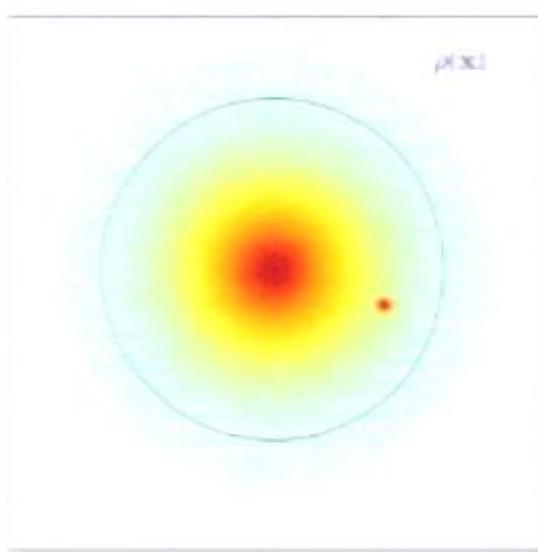
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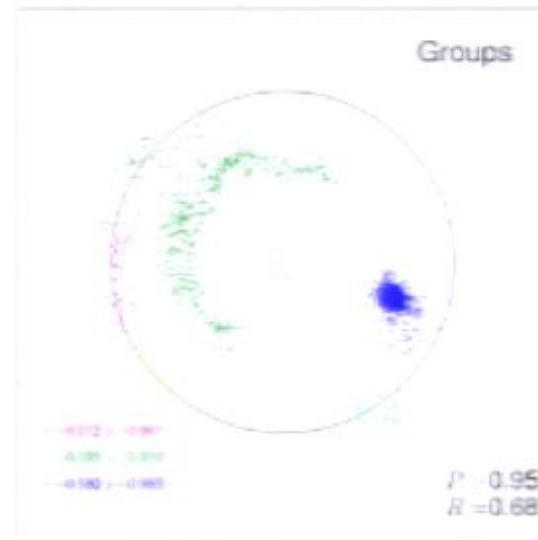
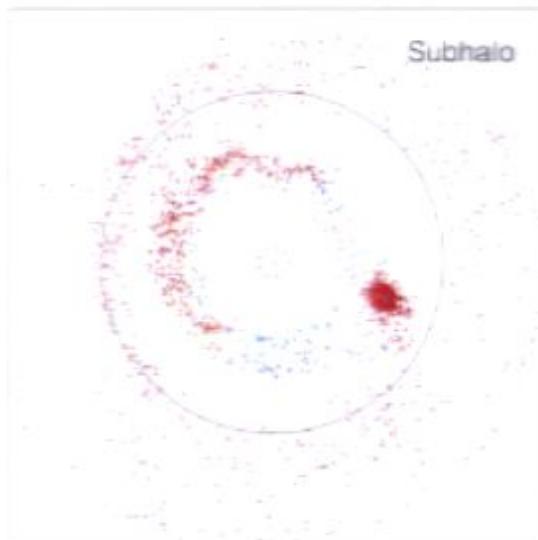
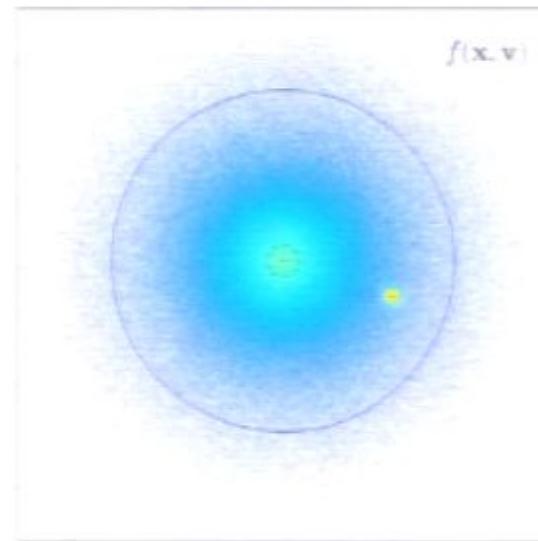
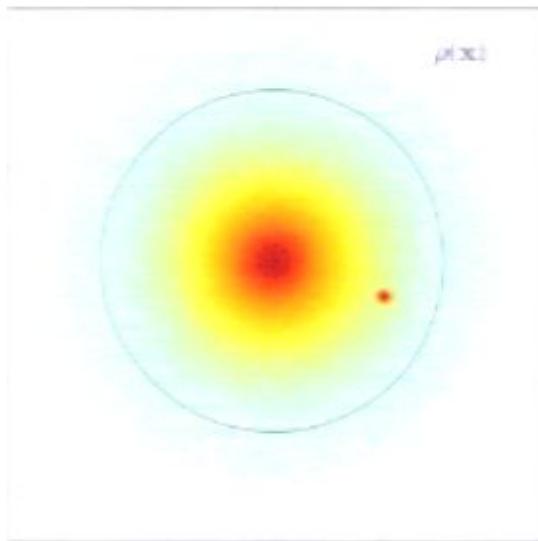


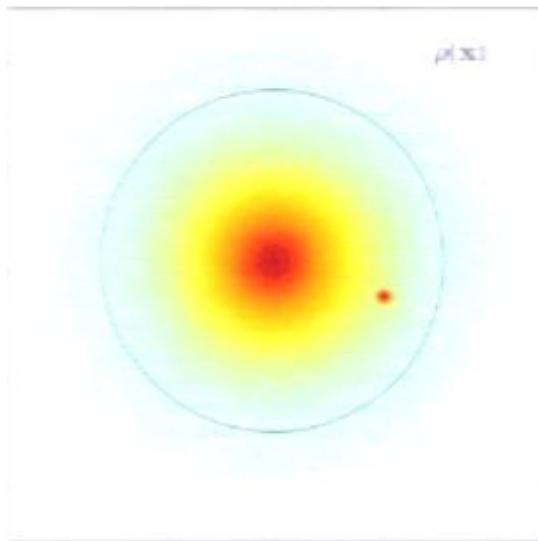




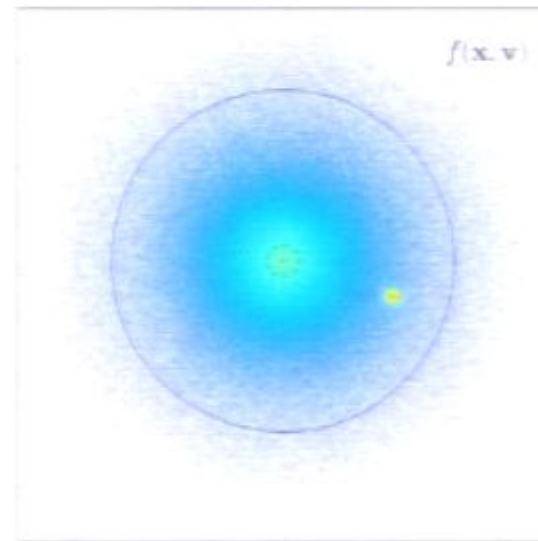




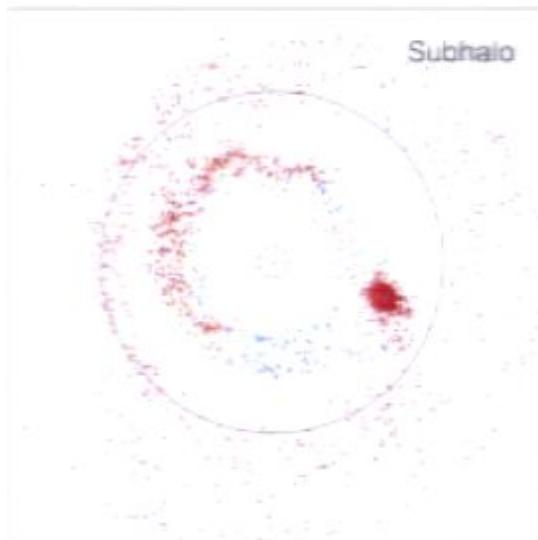




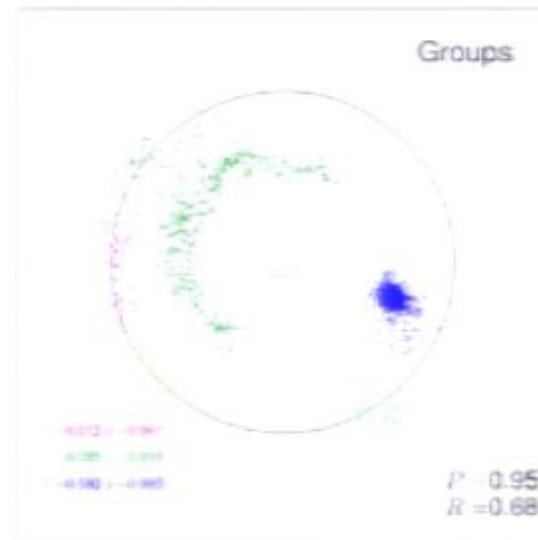
$\rho(\mathbf{x})$



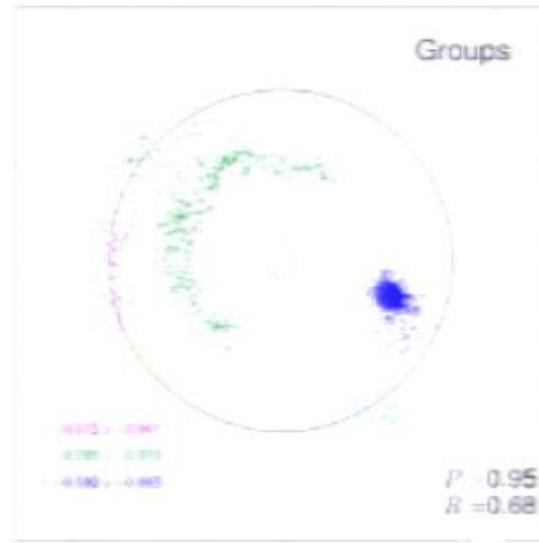
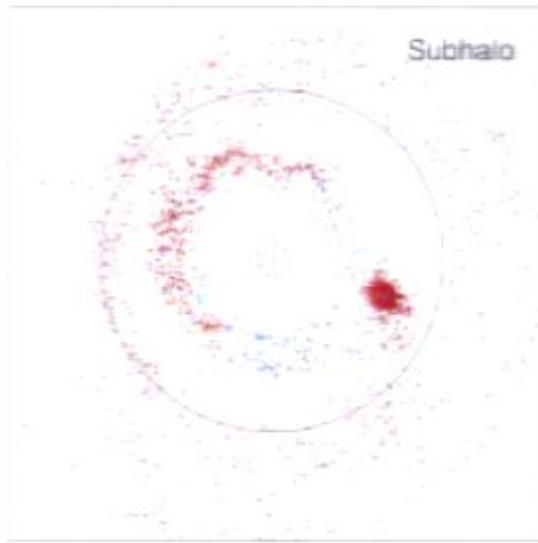
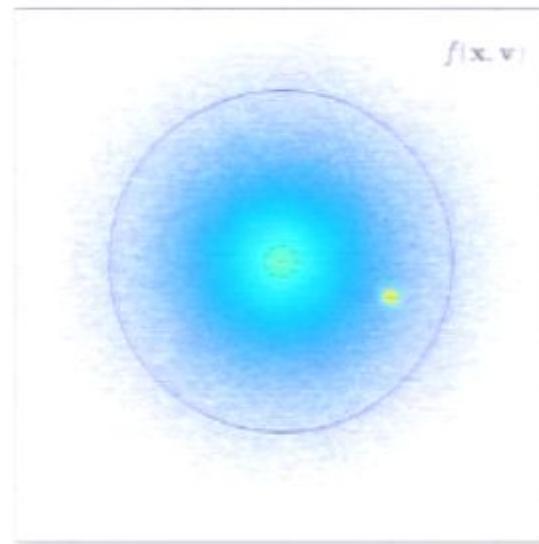
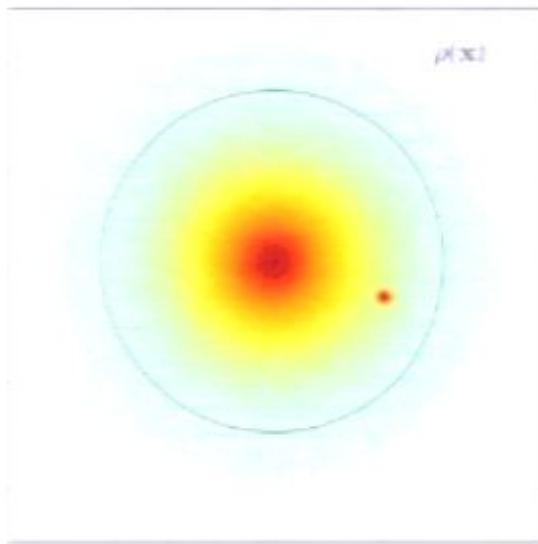
$f(\mathbf{x}, \mathbf{v})$

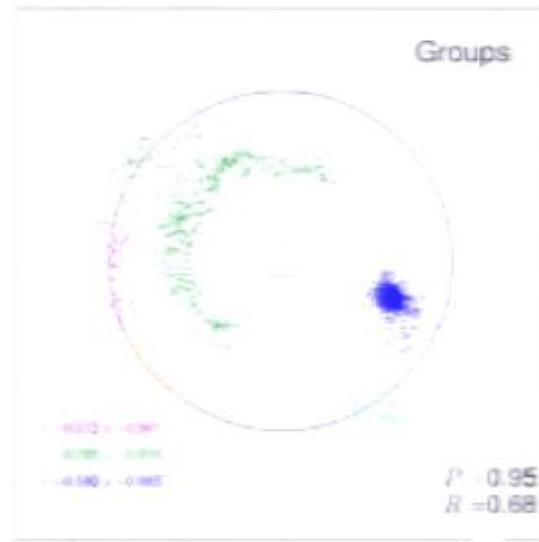
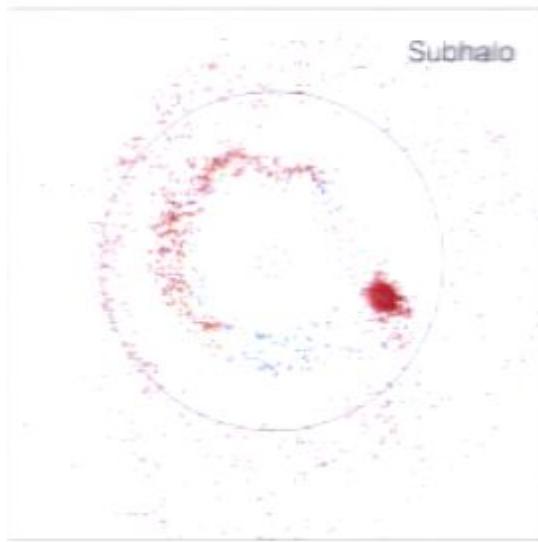
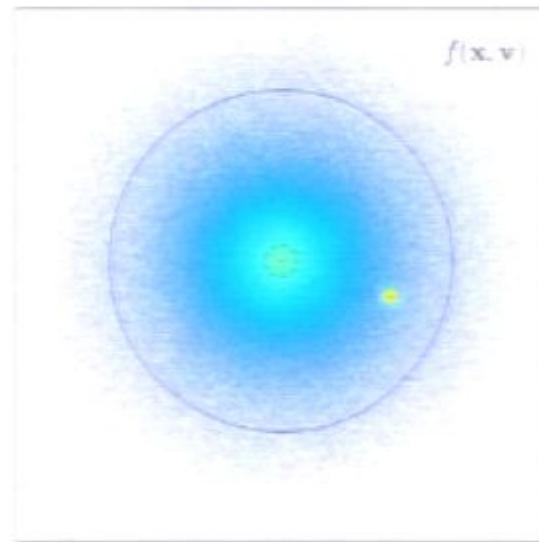
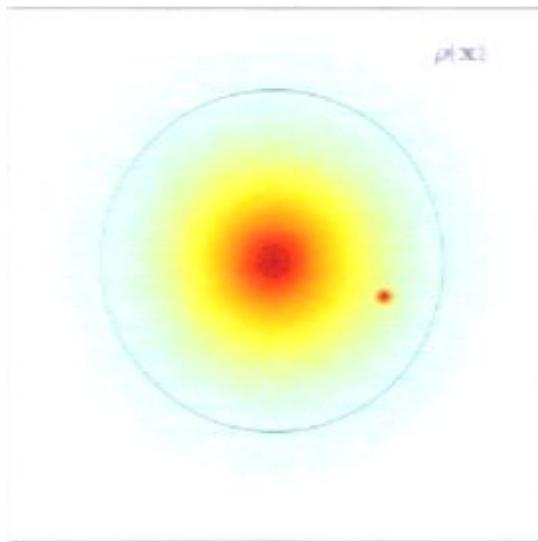


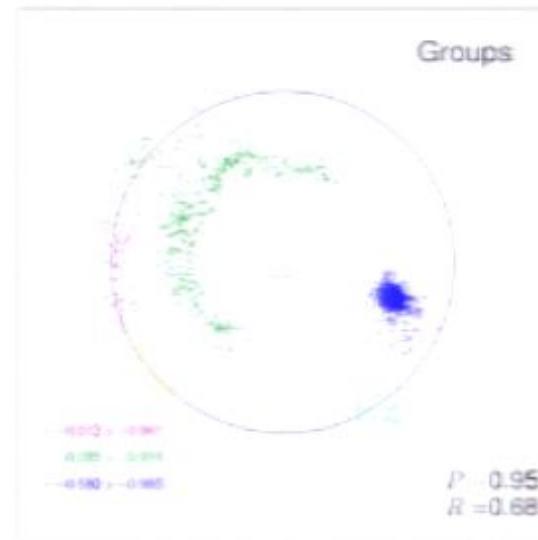
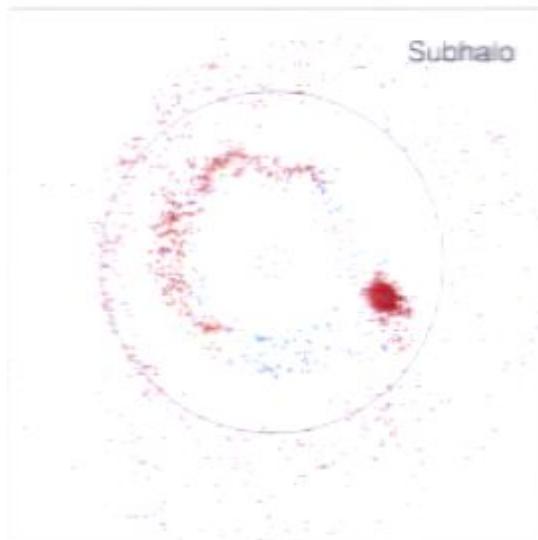
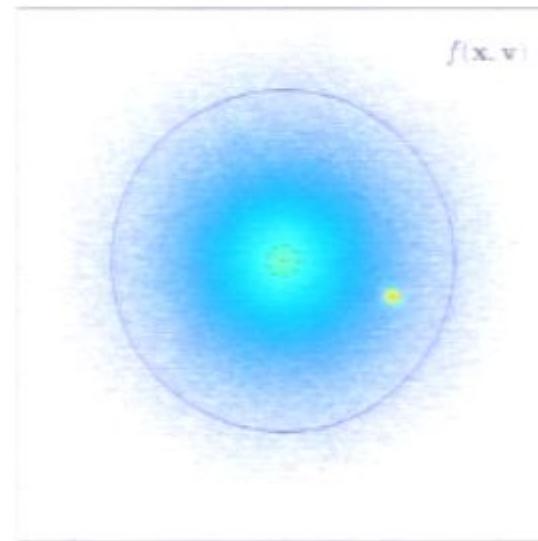
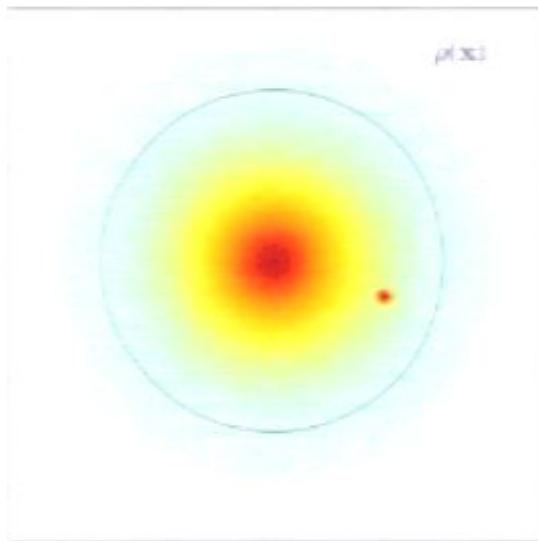
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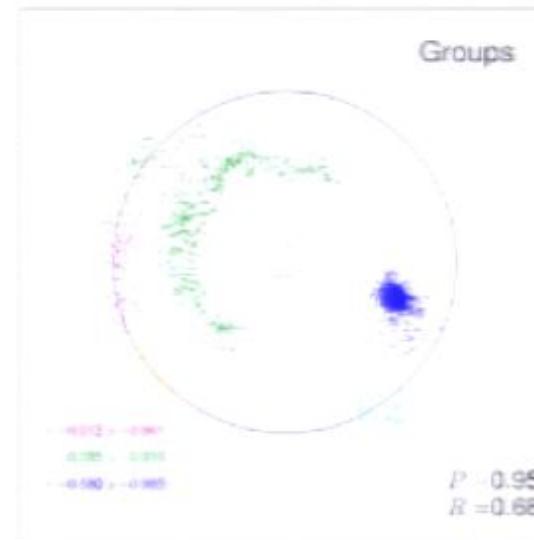
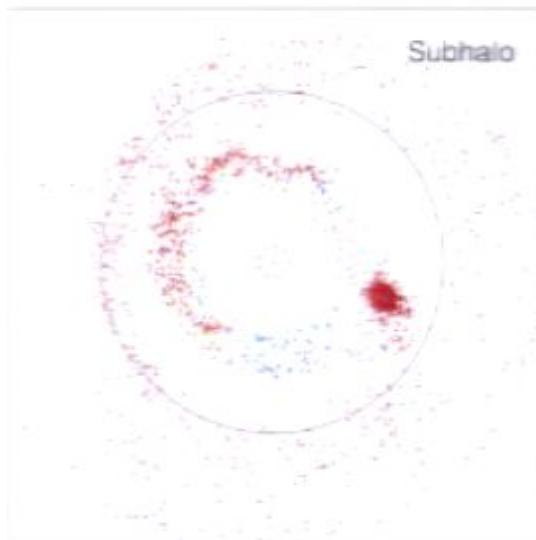
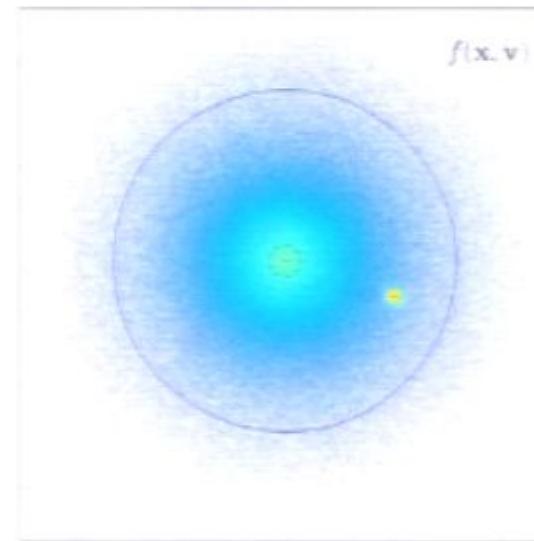
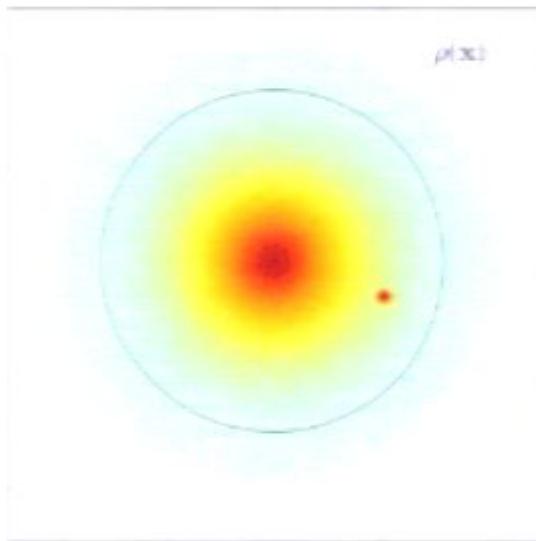


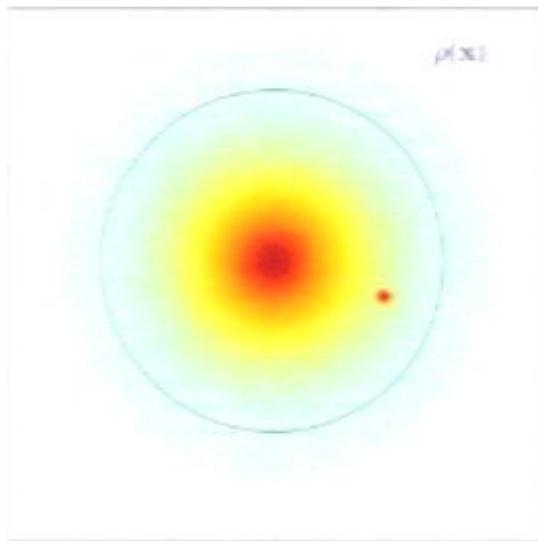
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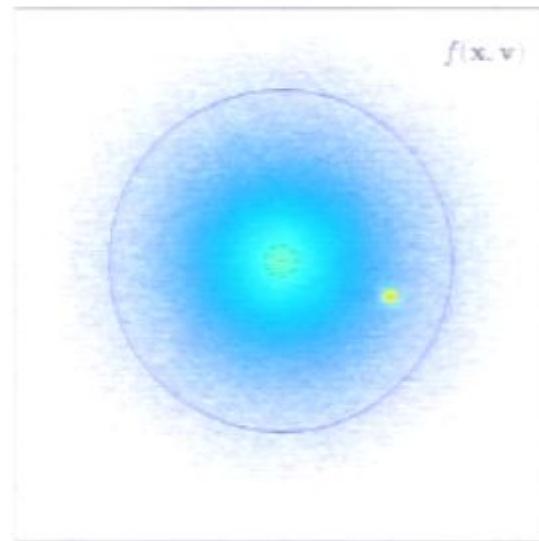




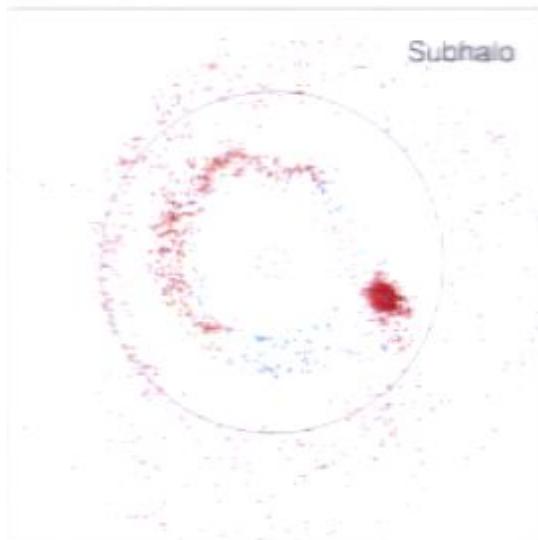




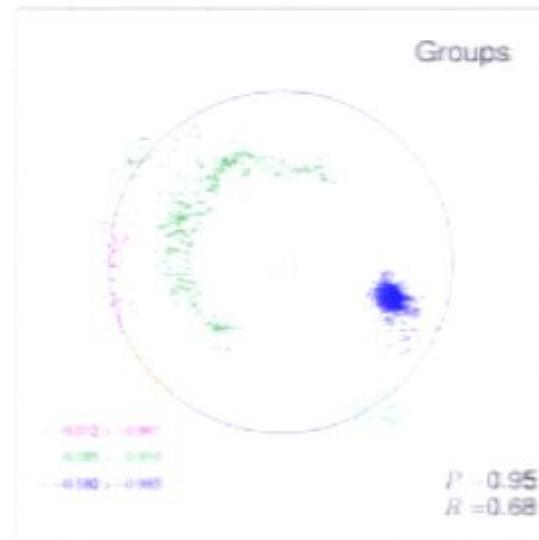
$\rho(\mathbf{x})$



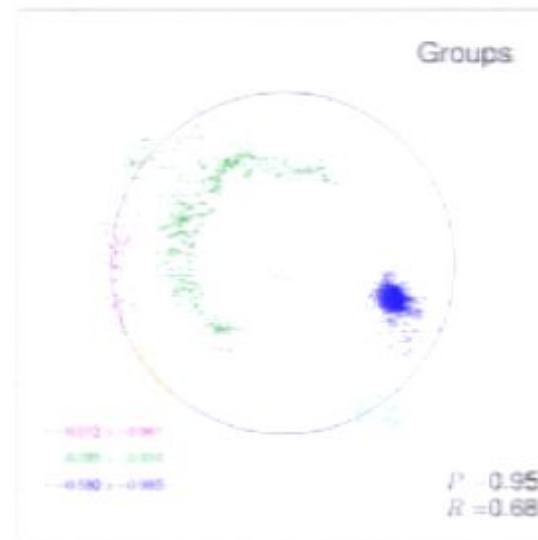
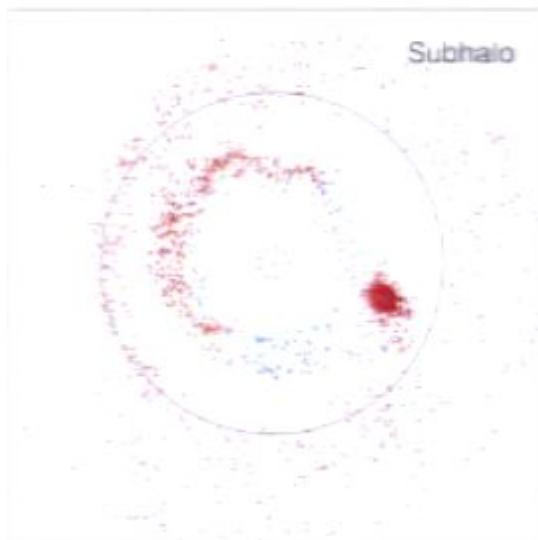
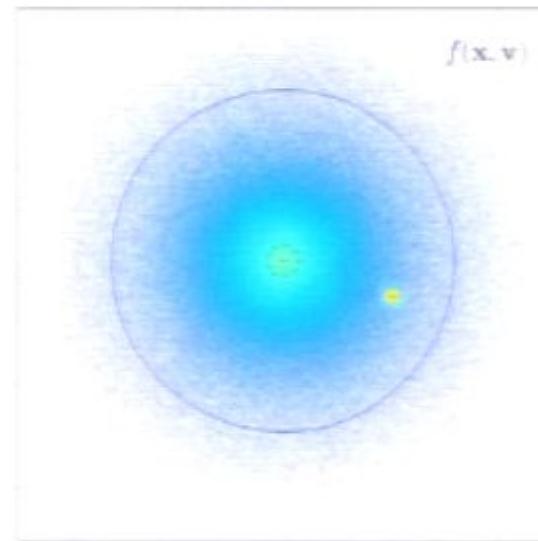
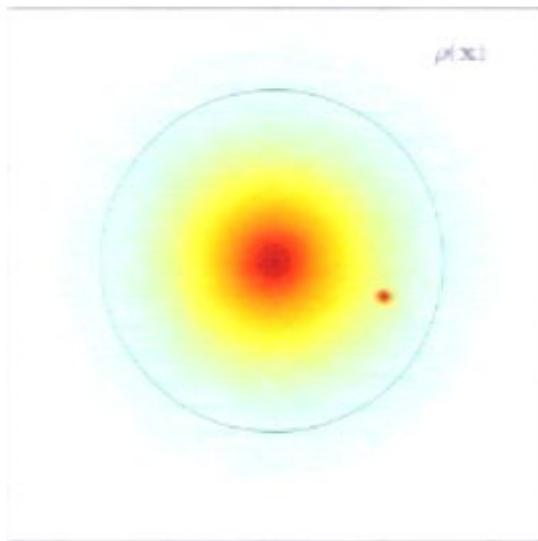
$f(\mathbf{x}, \mathbf{v})$

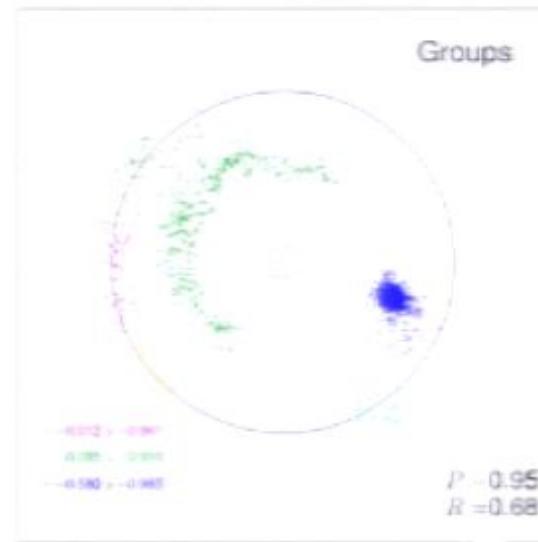
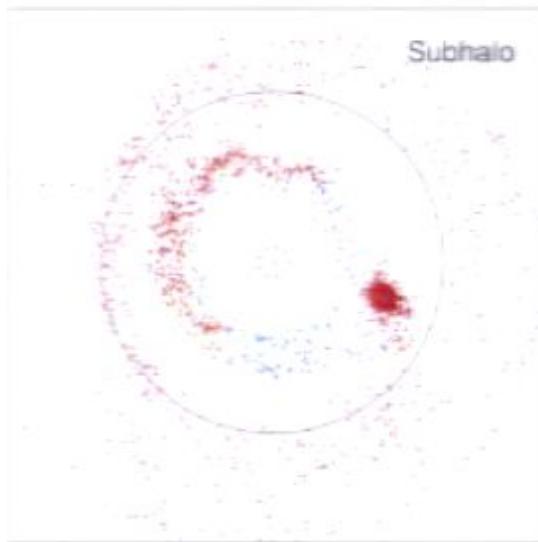
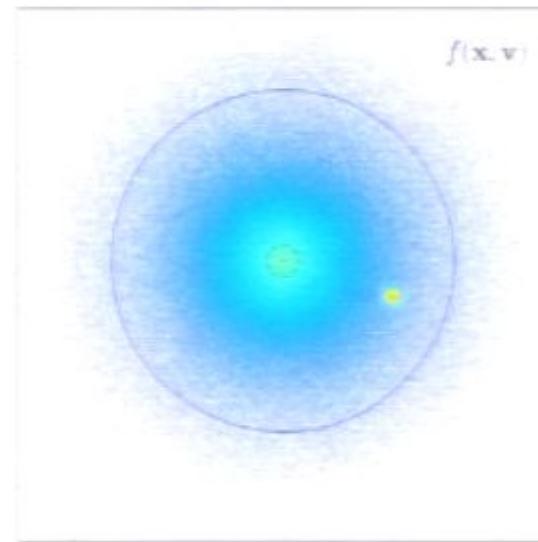
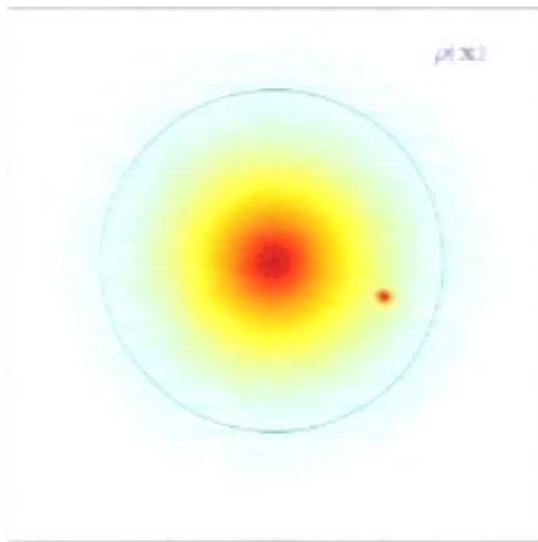


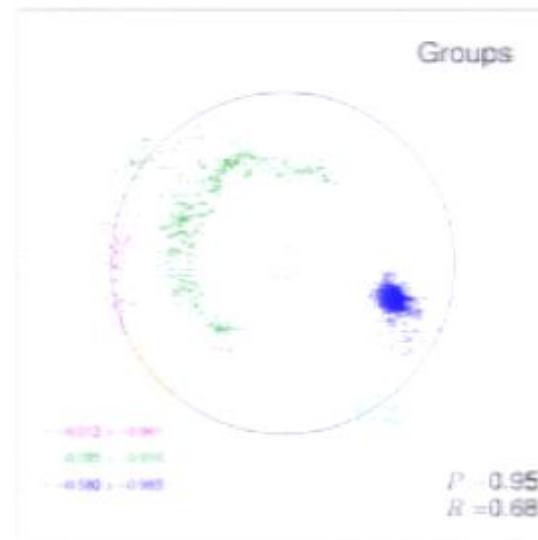
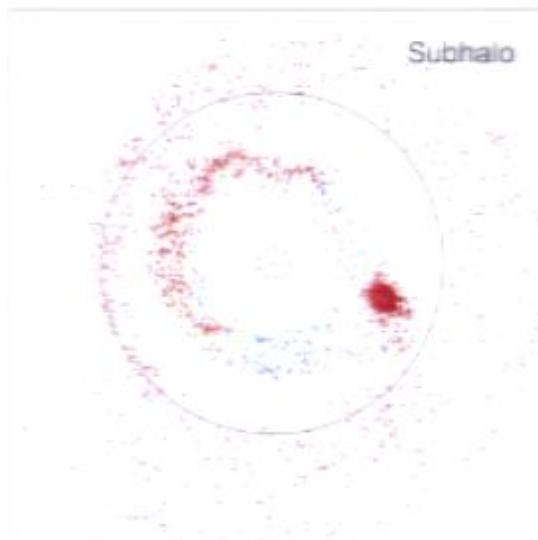
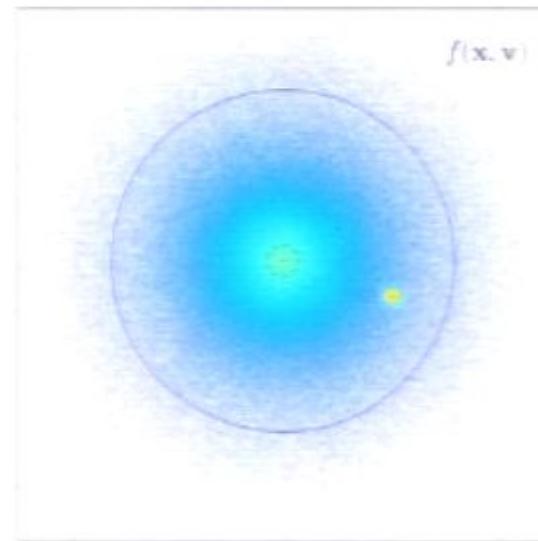
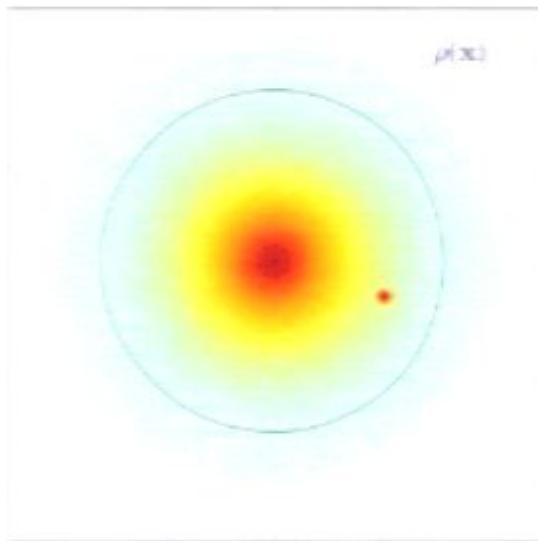
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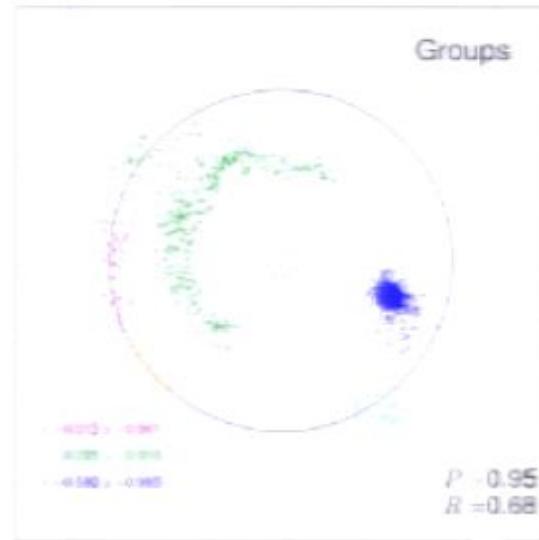
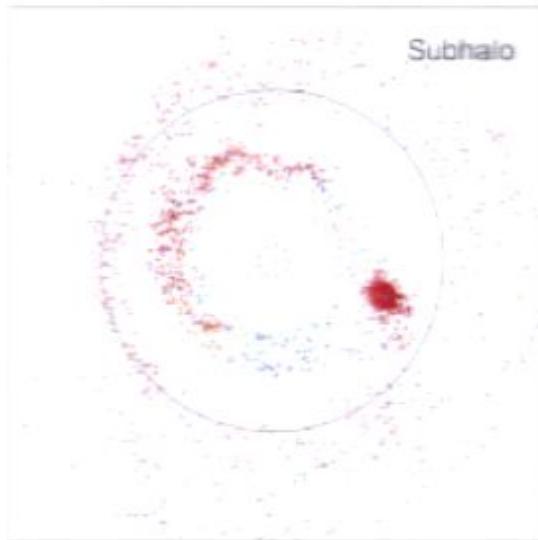
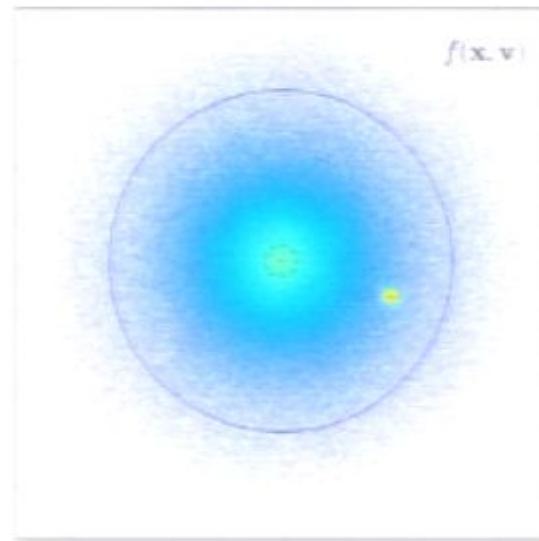
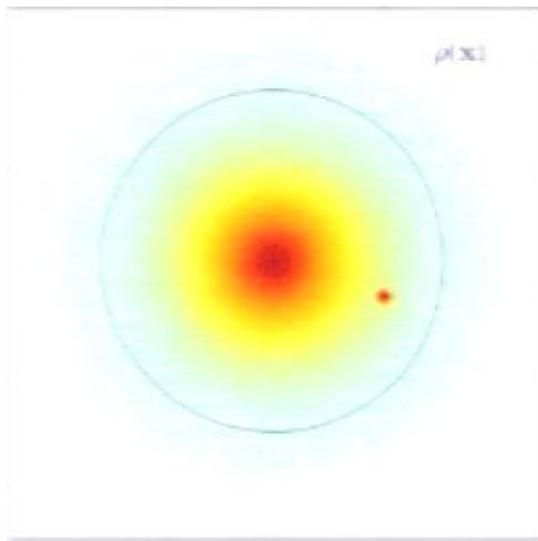


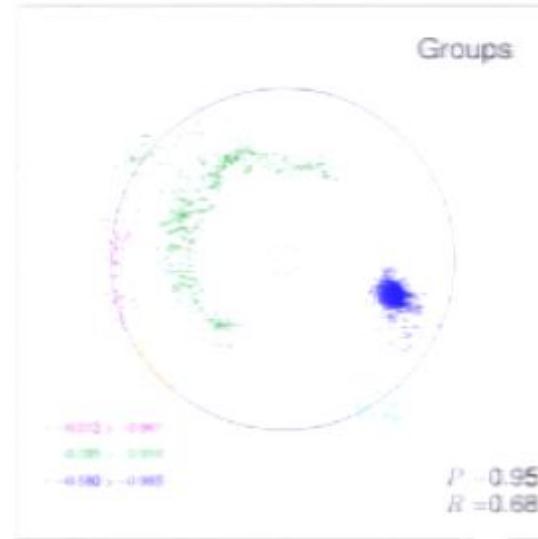
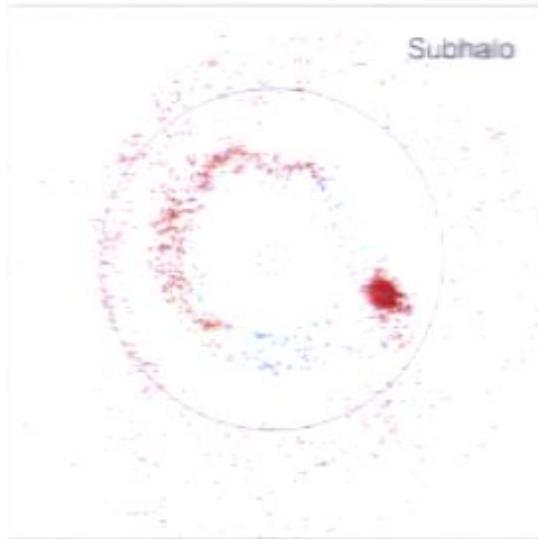
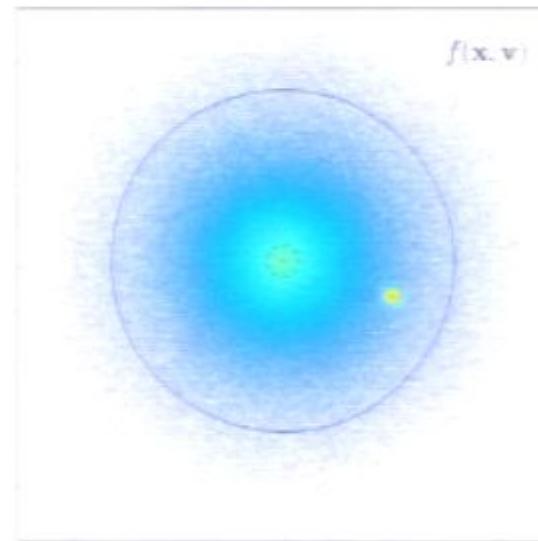
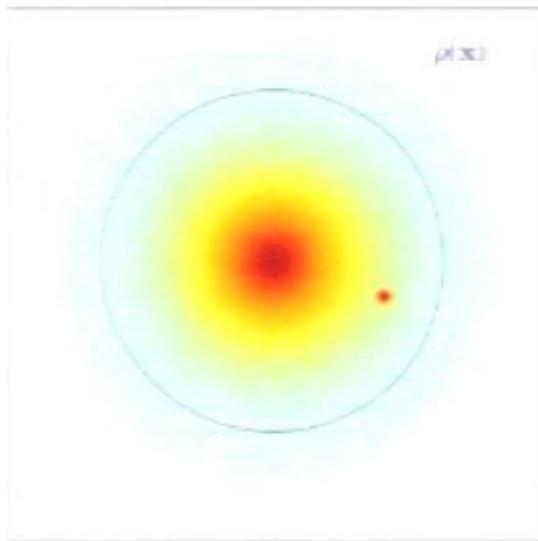
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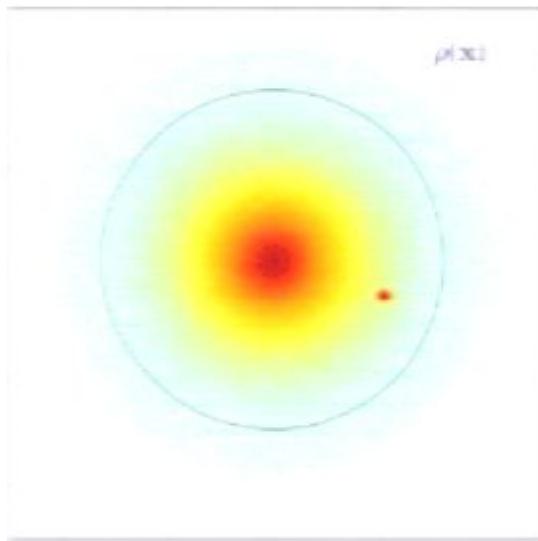




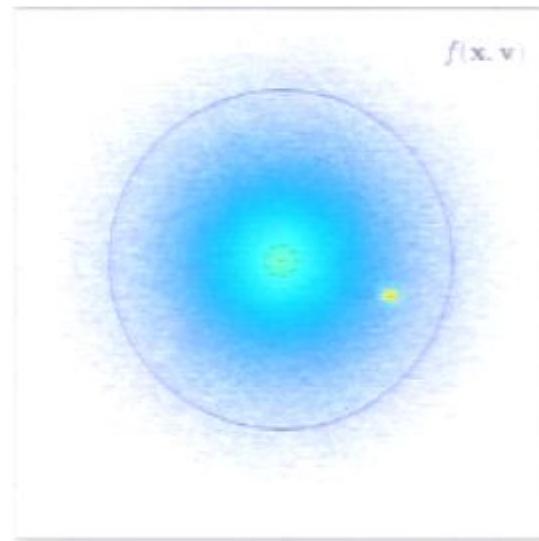




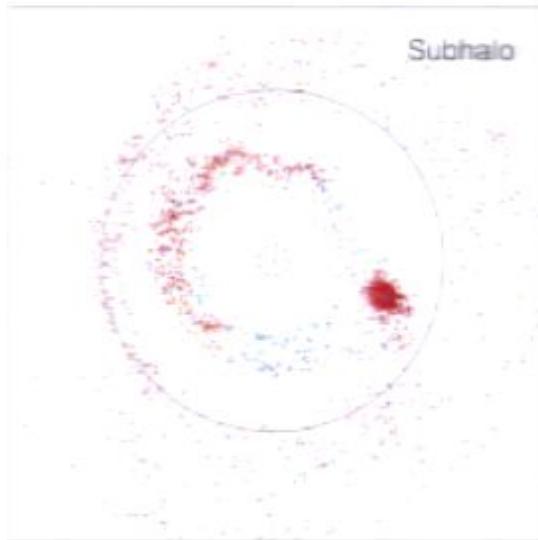




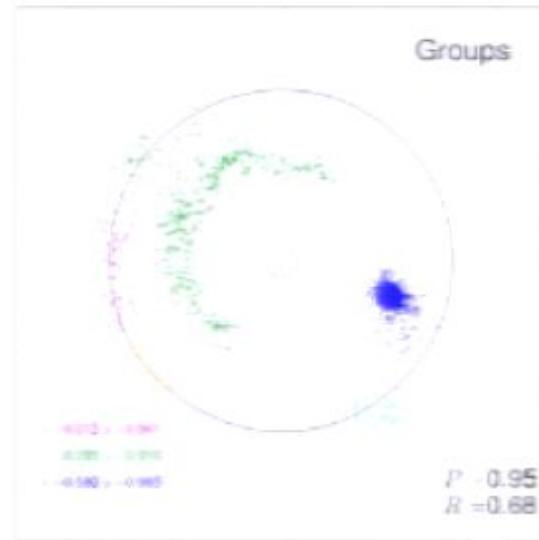
$\rho(30)$



$f(\mathbf{x}, \mathbf{v})$



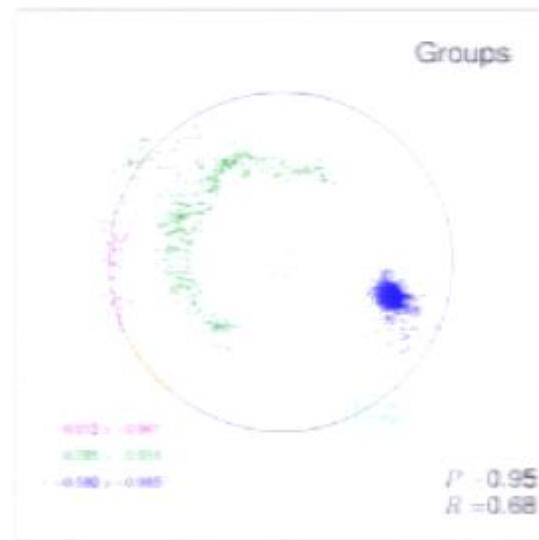
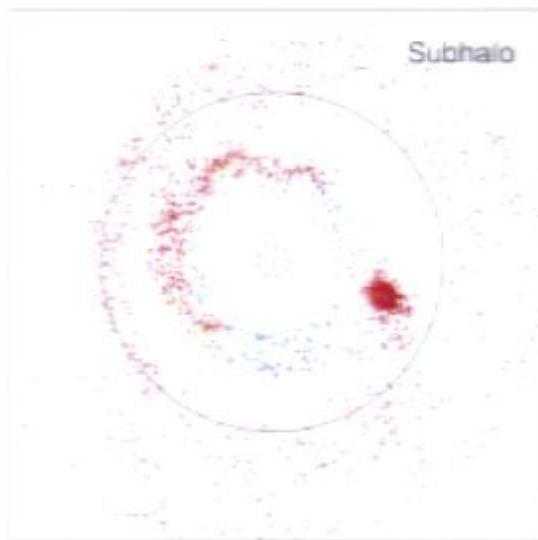
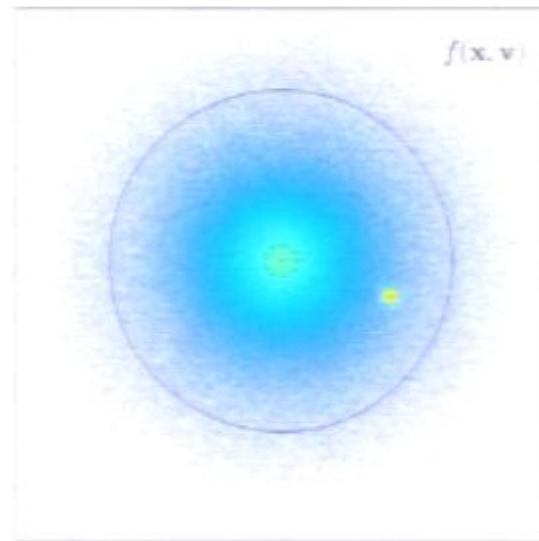
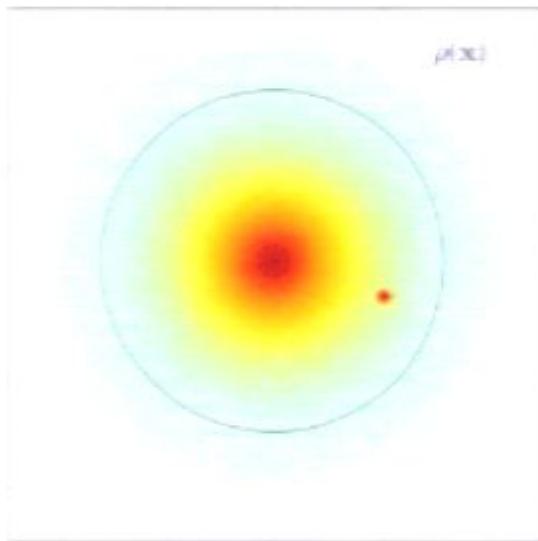
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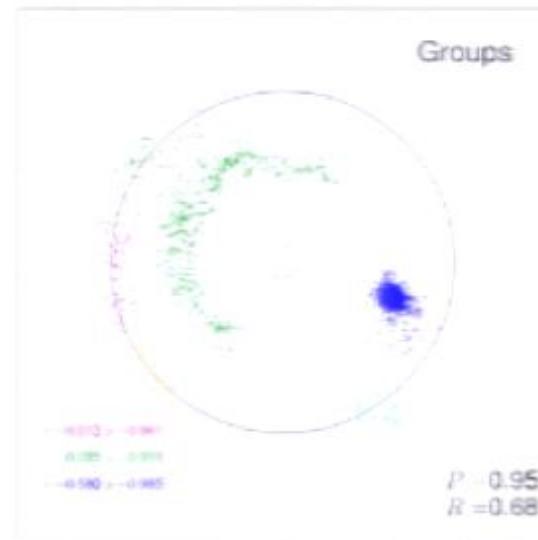
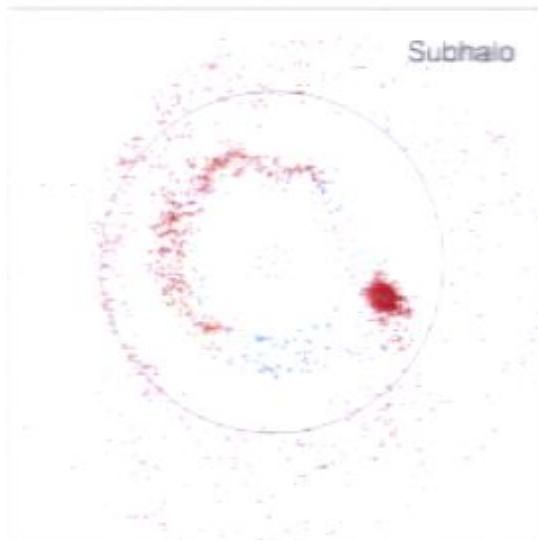
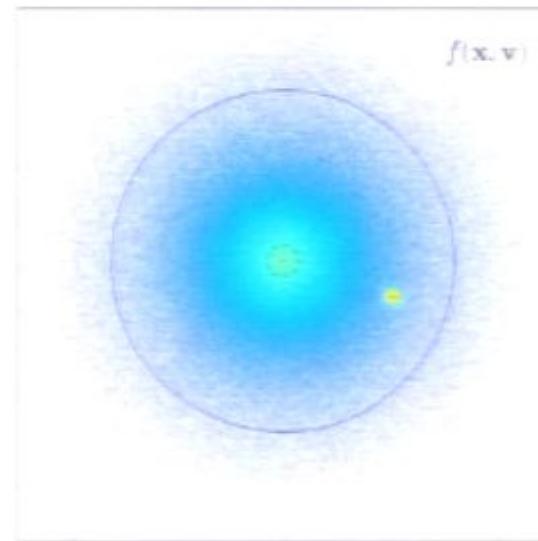
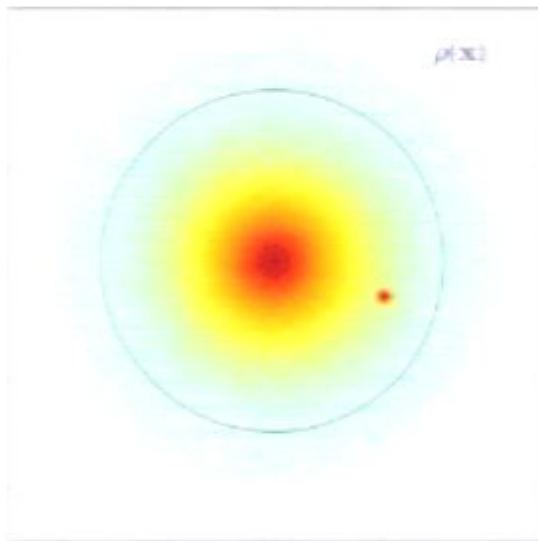


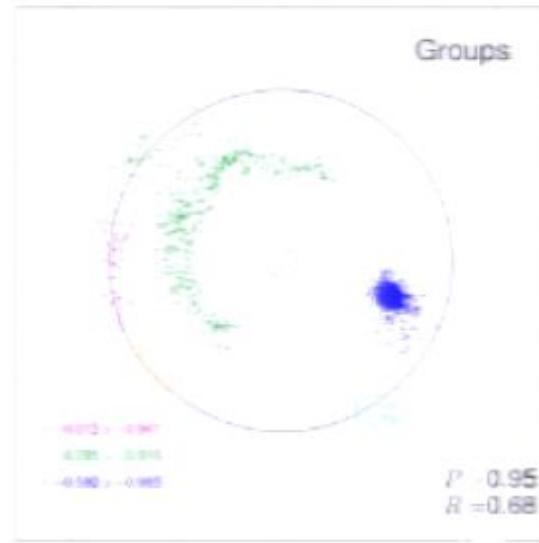
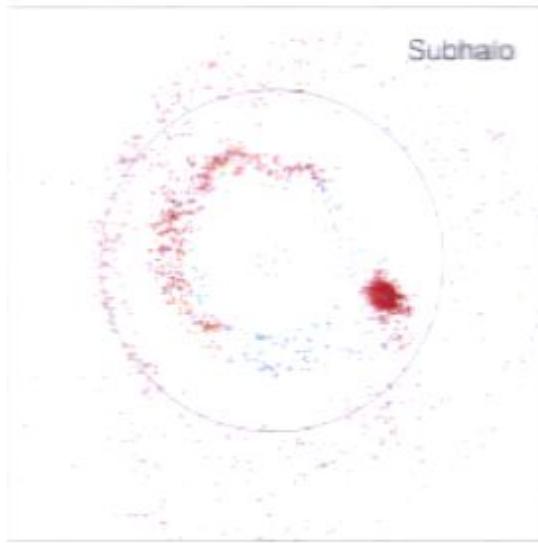
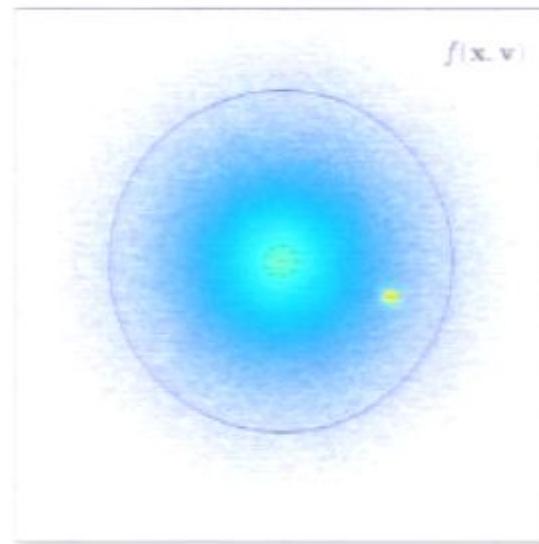
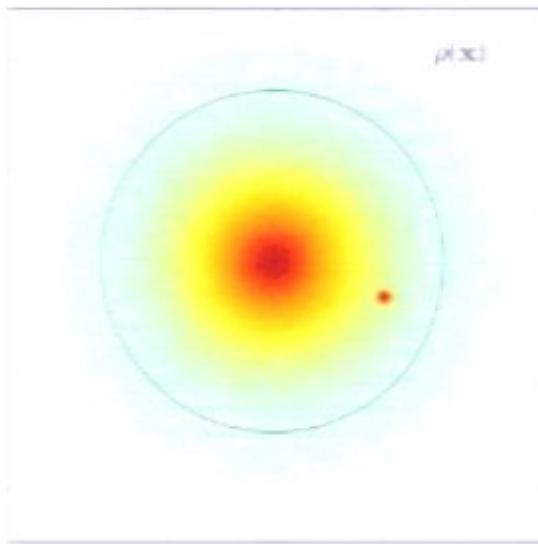
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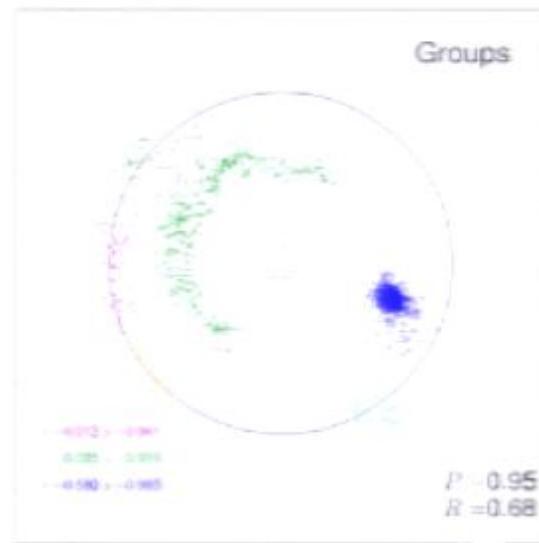
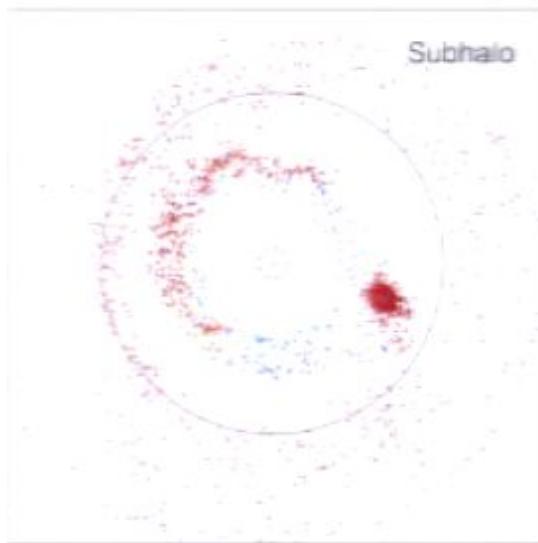
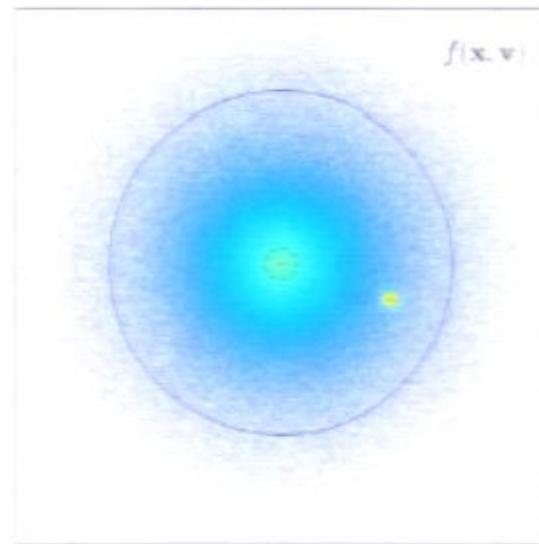
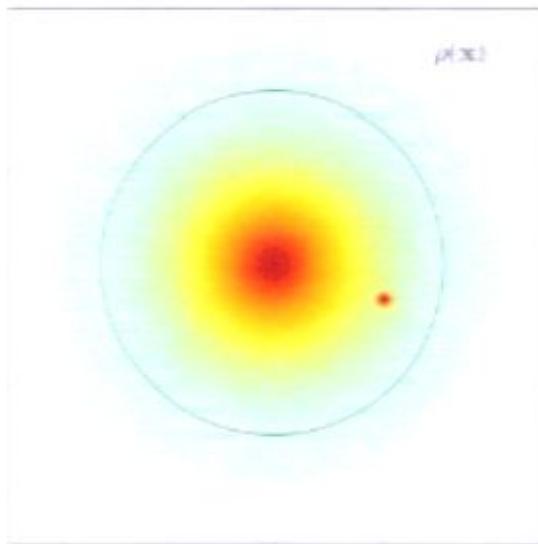
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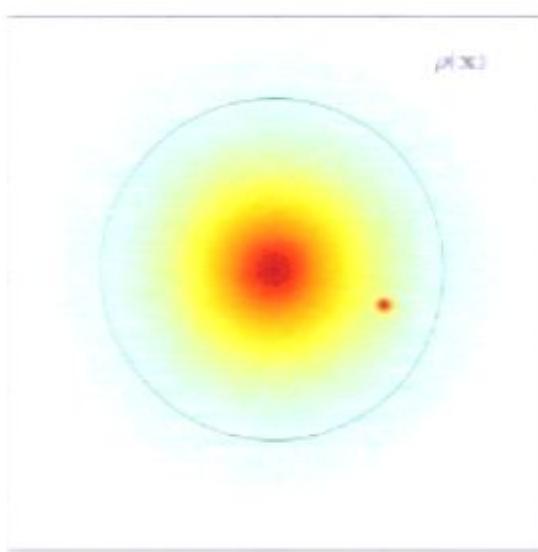
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 $R = 0.68$



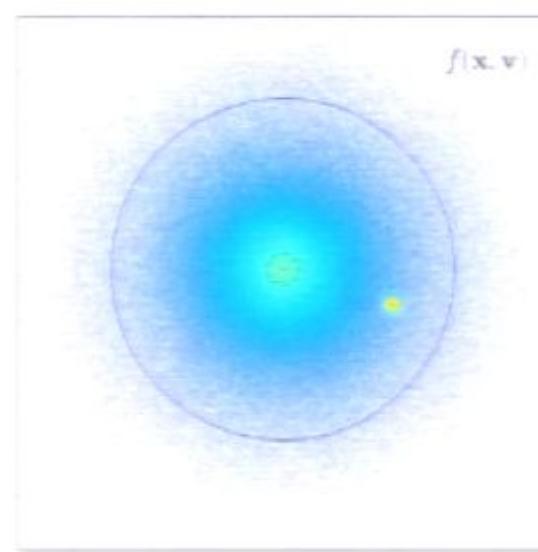




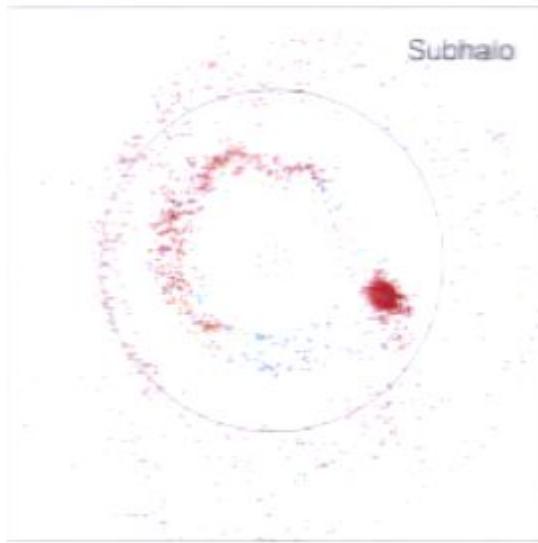




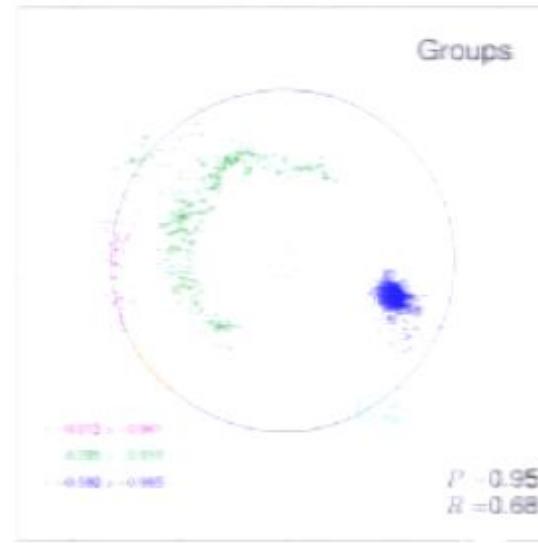
$\rho(\mathbf{x})$



$f(\mathbf{x}, \mathbf{v})$

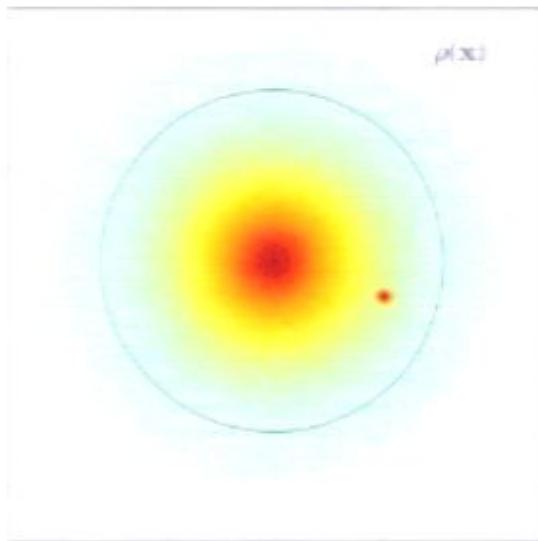


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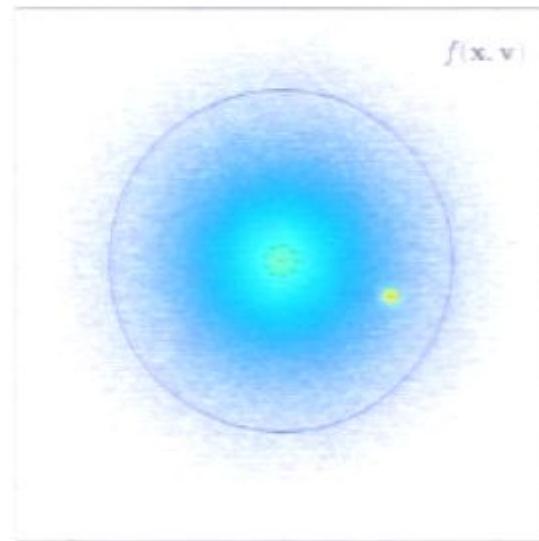


Groups

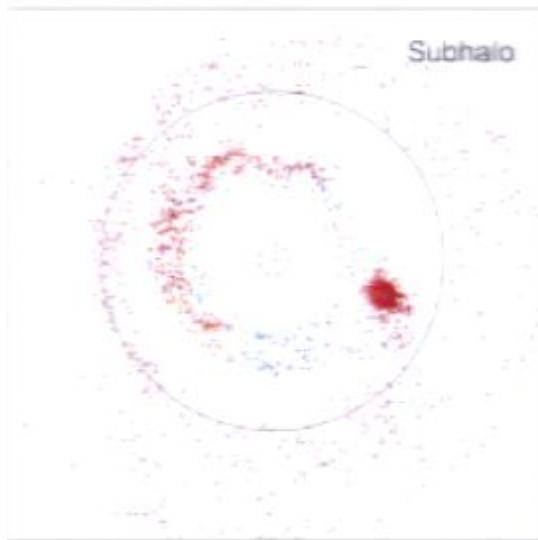
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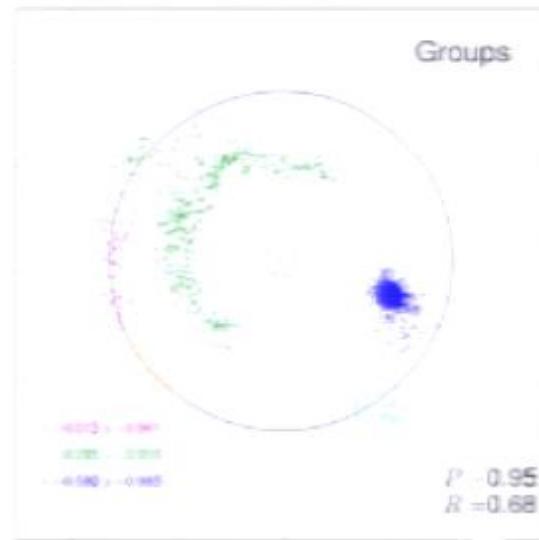
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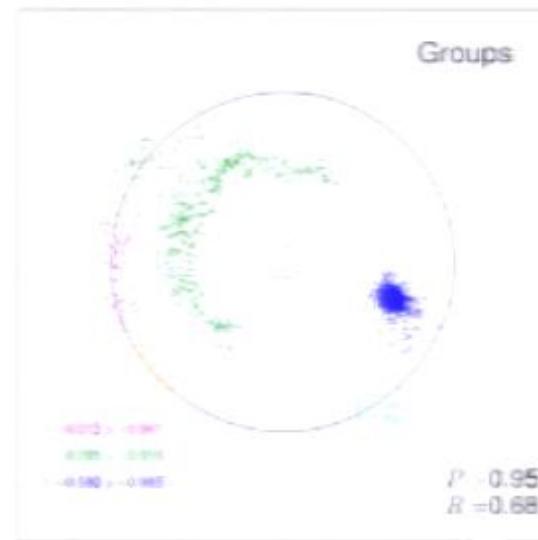
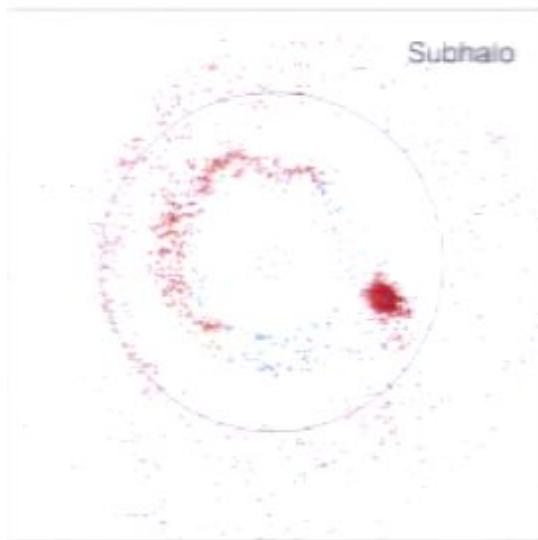
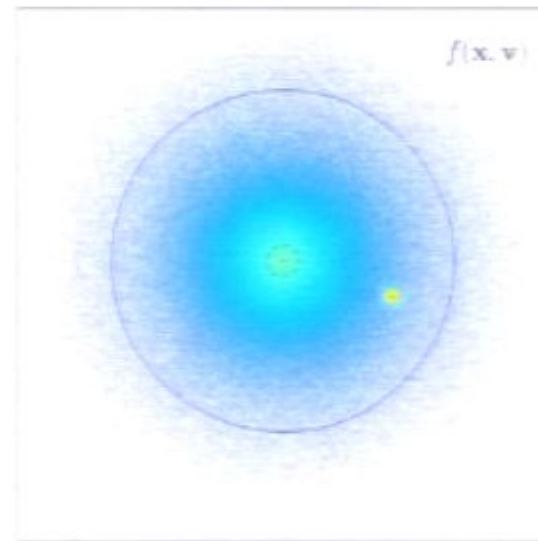
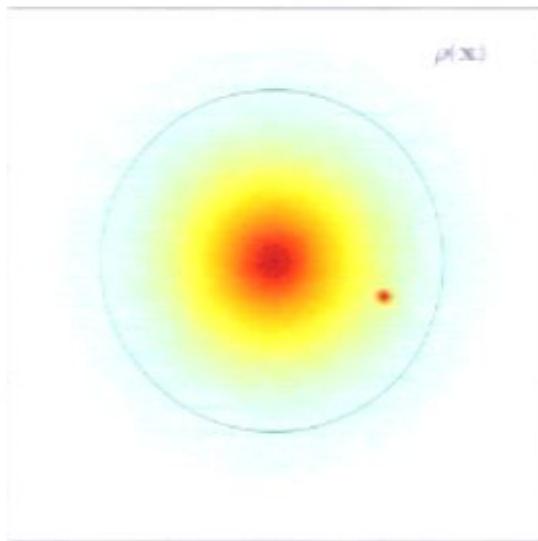
$f(\mathbf{x}, \mathbf{v})$

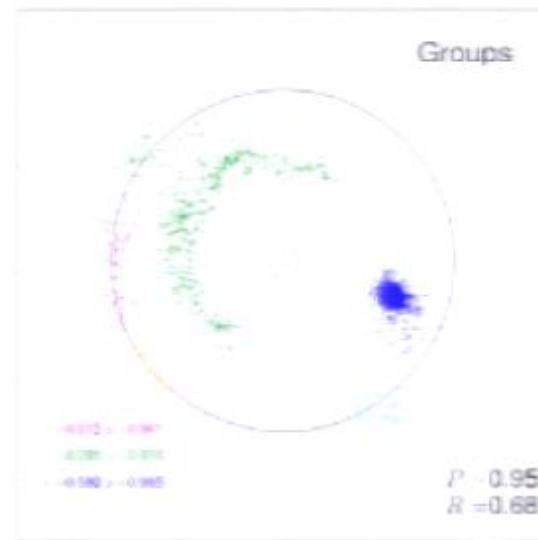
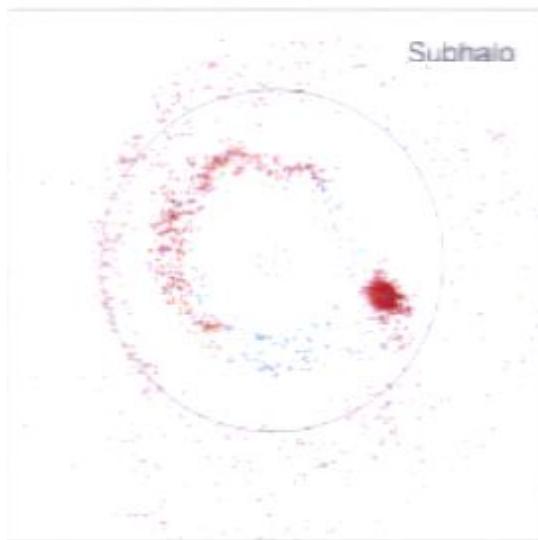
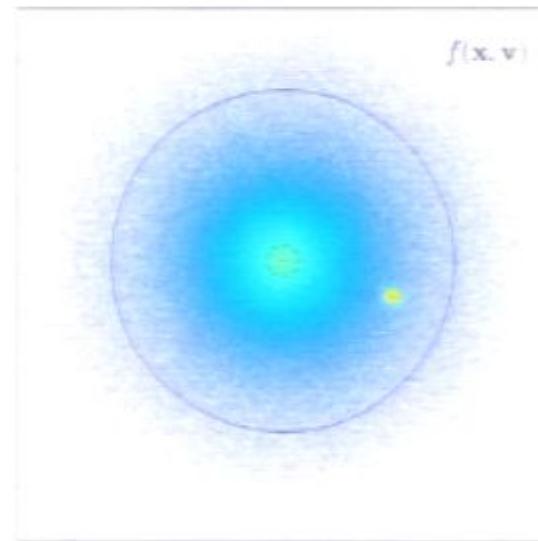
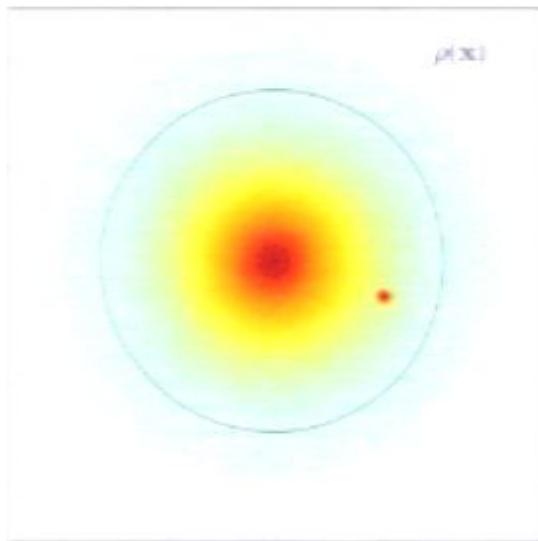


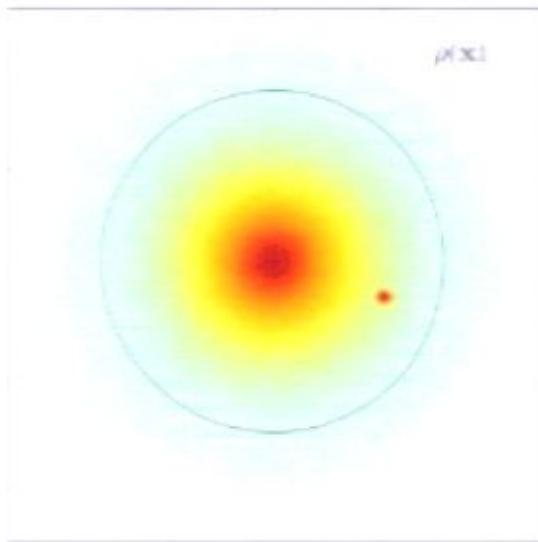
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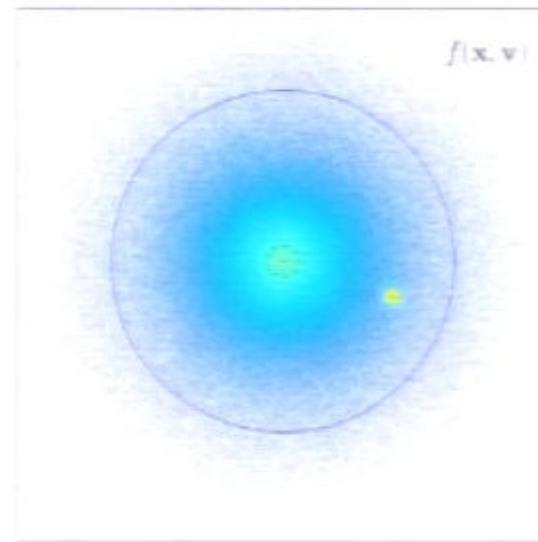
Groups



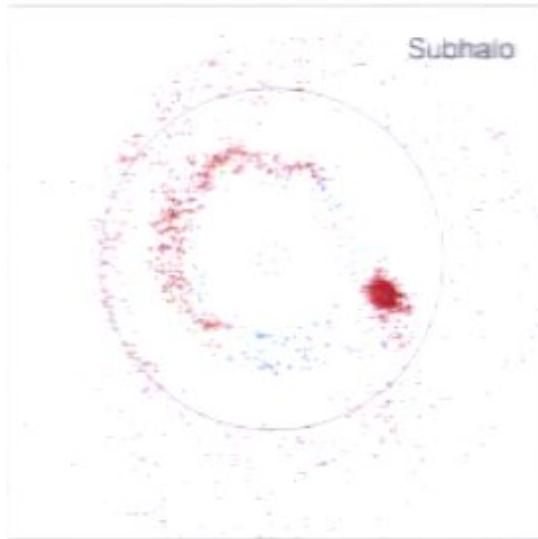




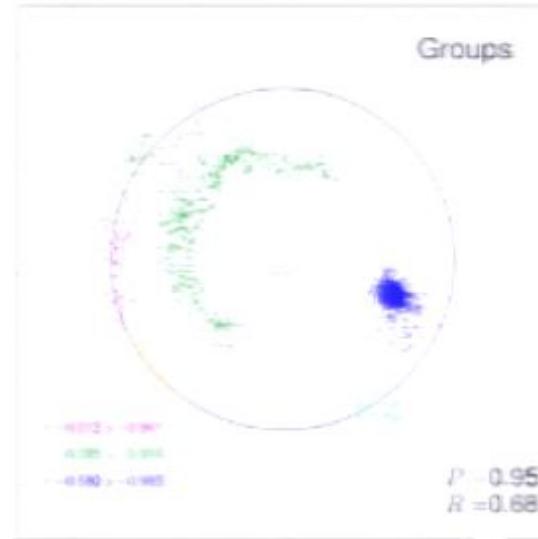
$\rho(\mathbf{x})$



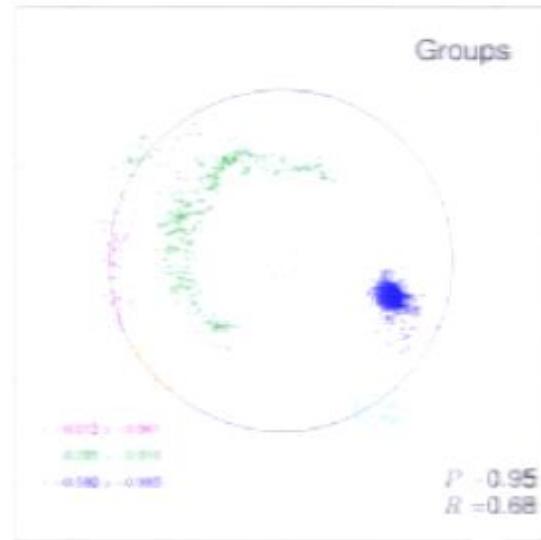
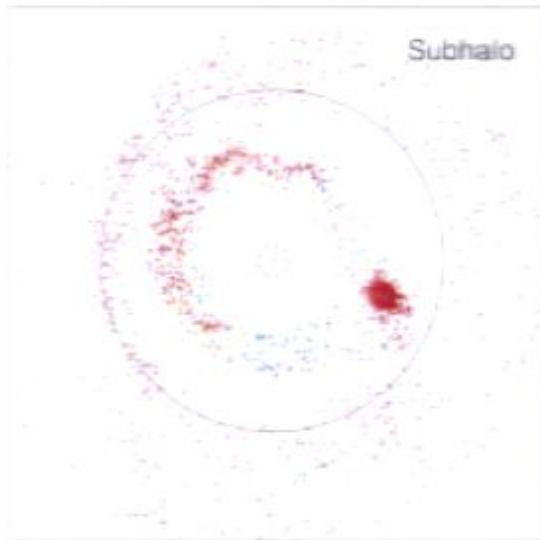
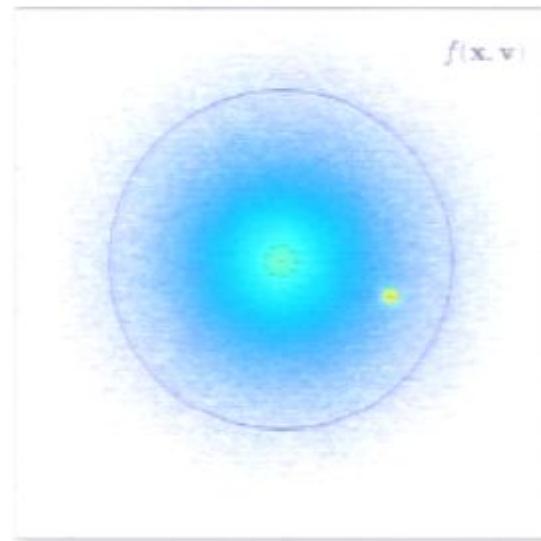
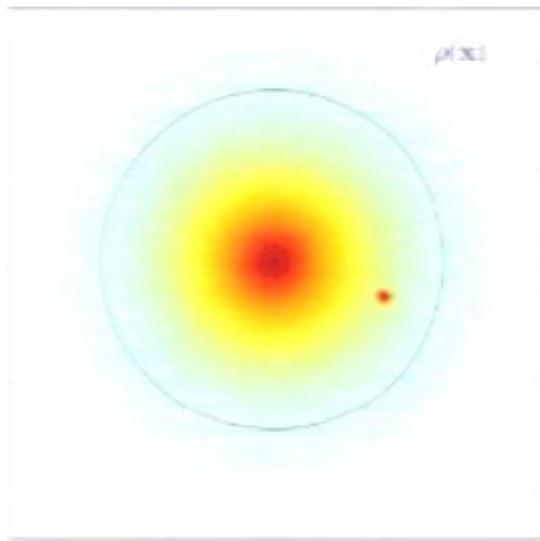
$f(\mathbf{x}, \mathbf{v})$

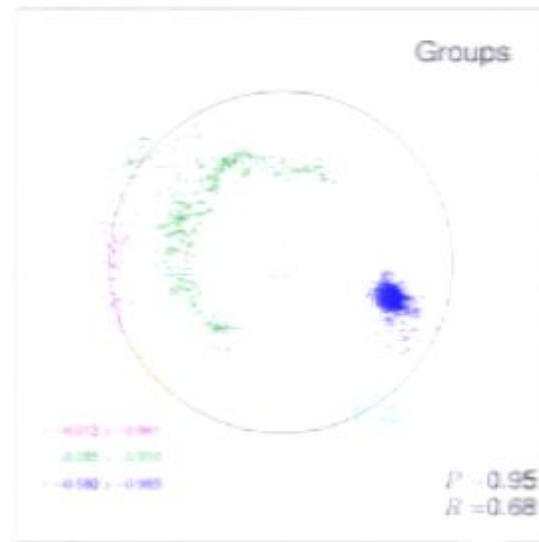
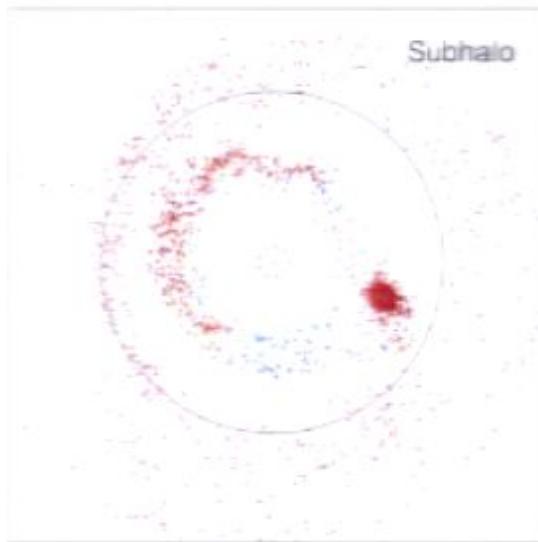
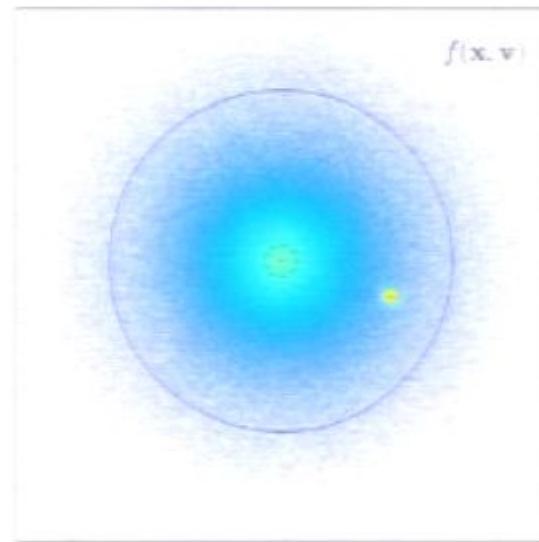
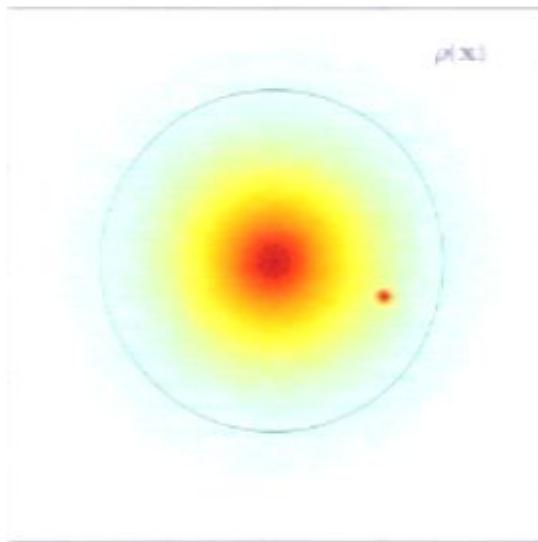


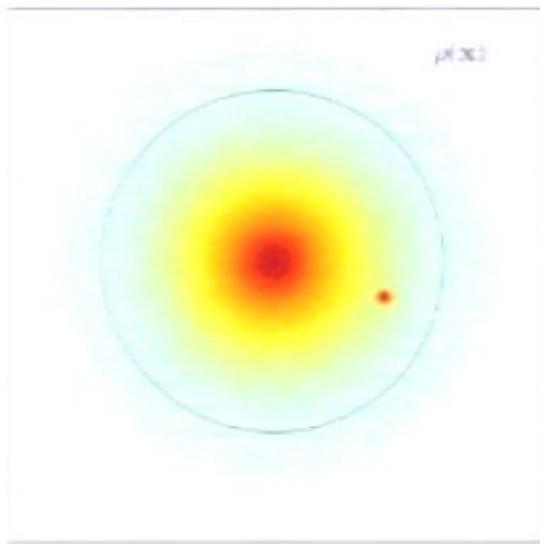
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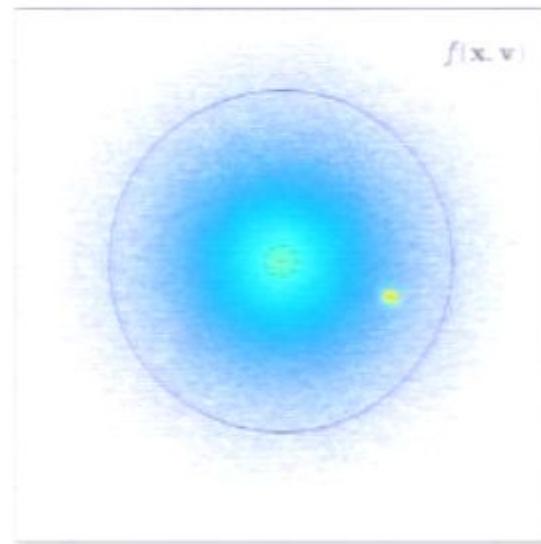
Groups



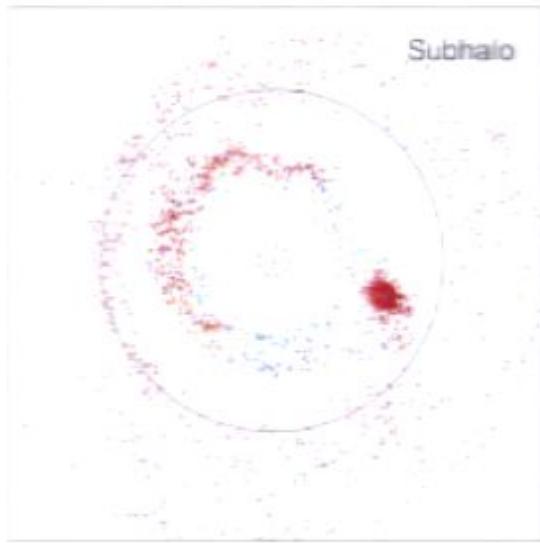




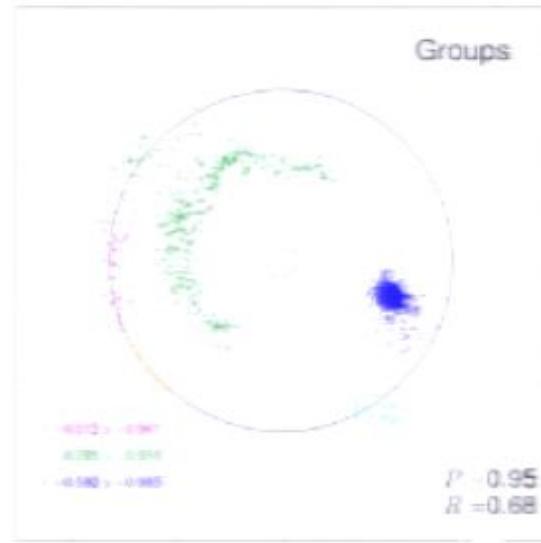
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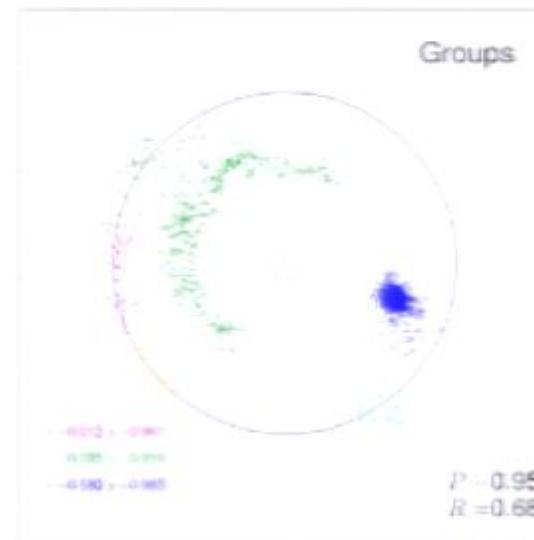
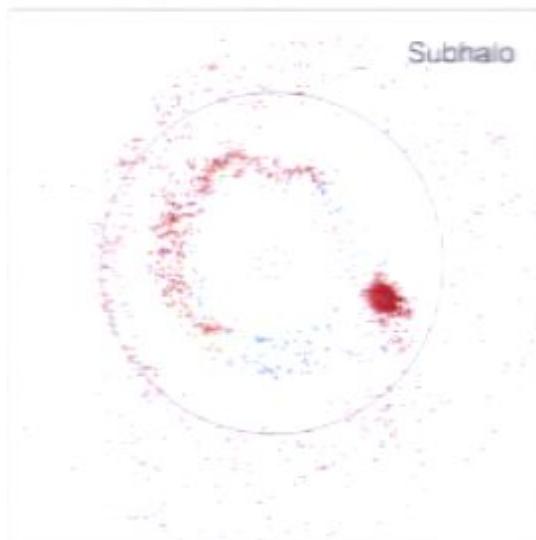
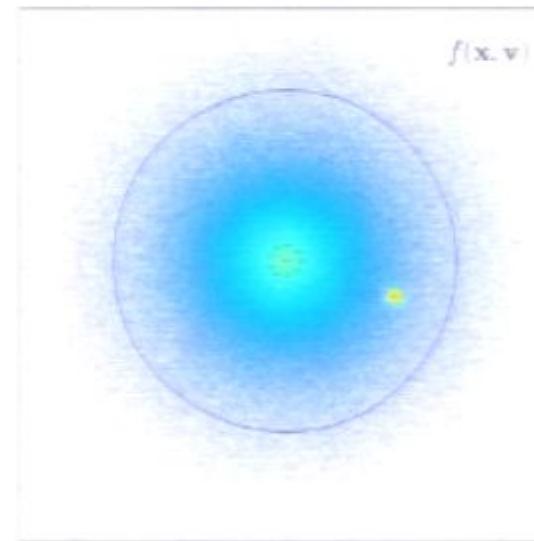
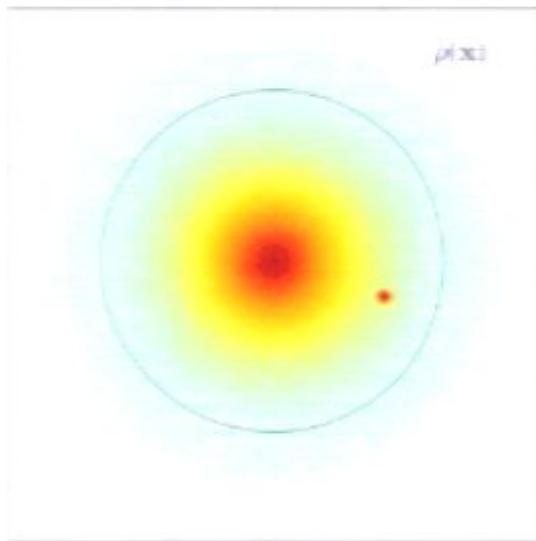
$f(\mathbf{x}, \mathbf{v})$

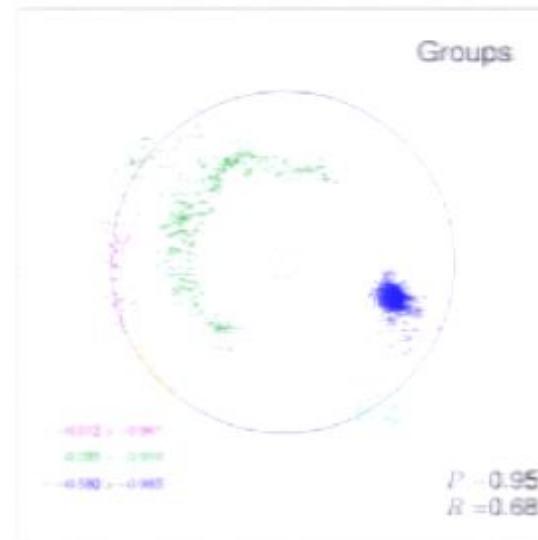
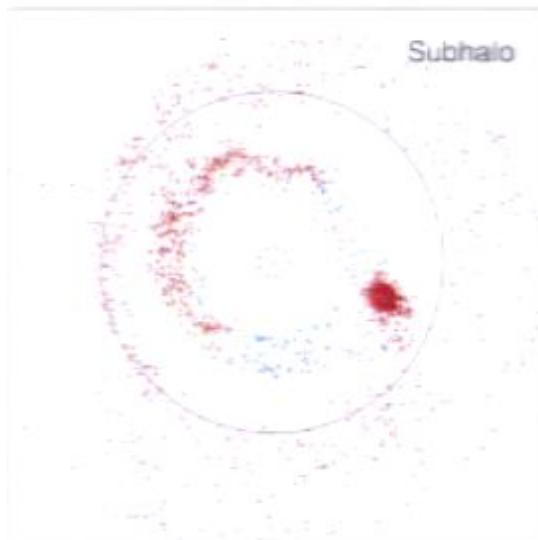
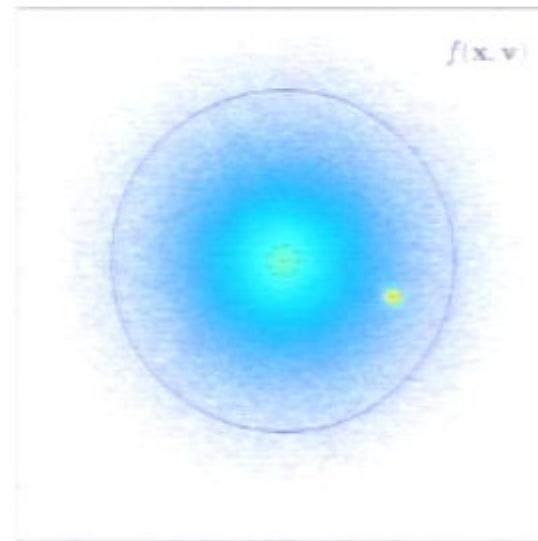
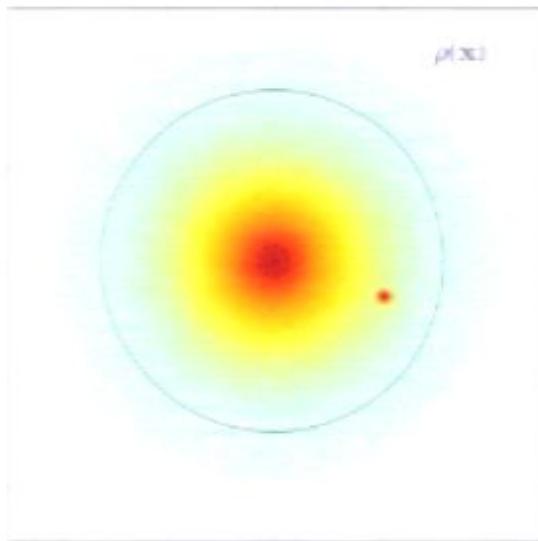


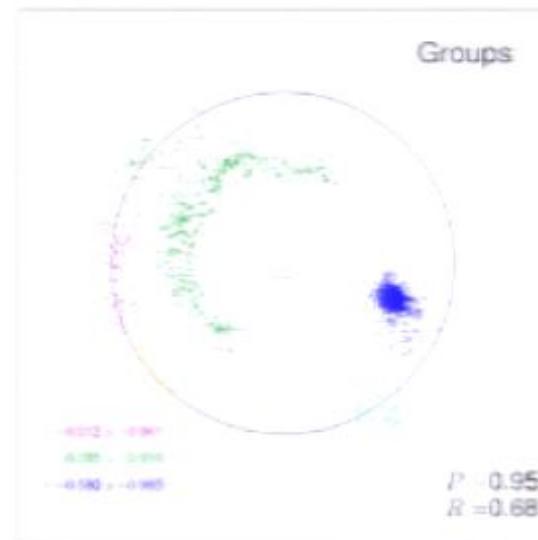
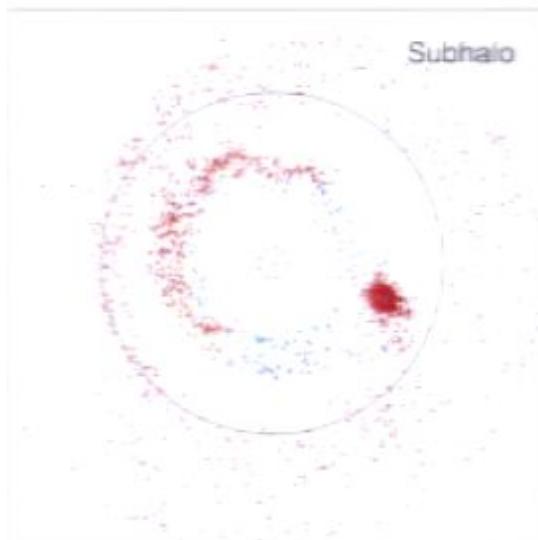
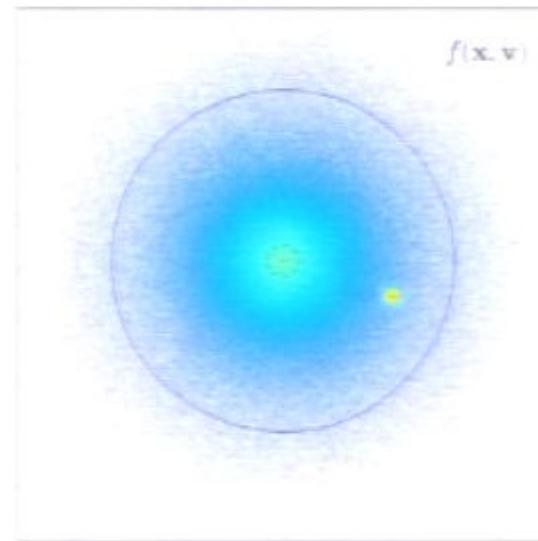
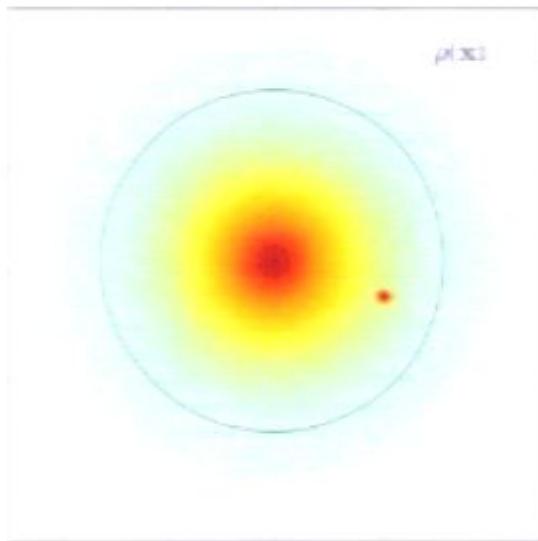
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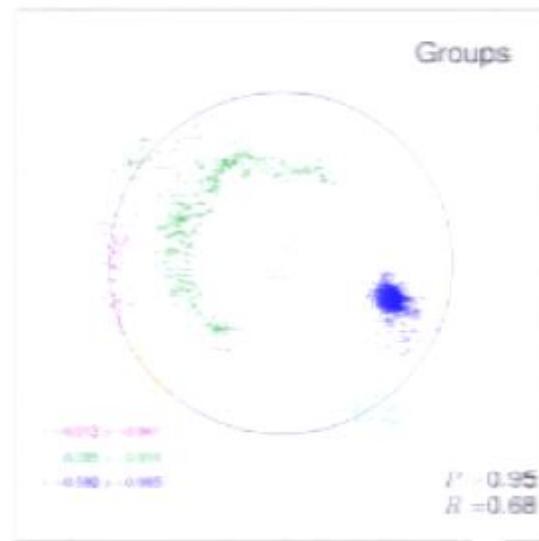
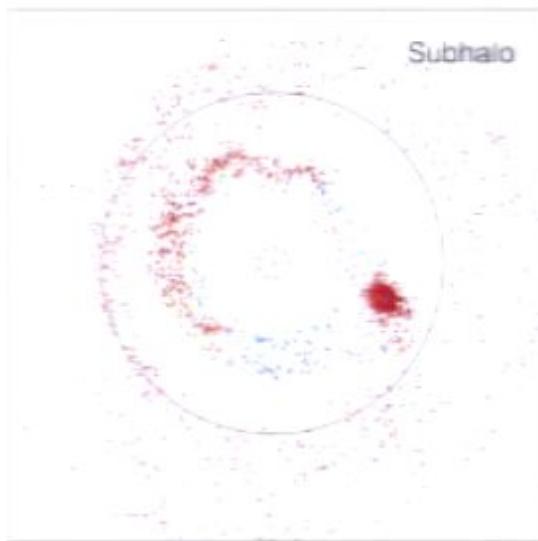
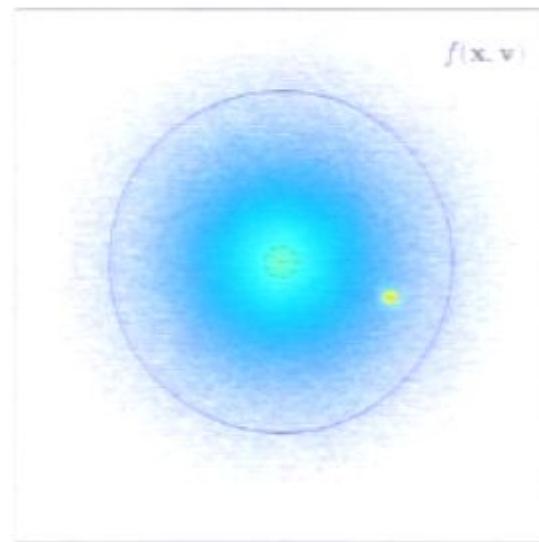
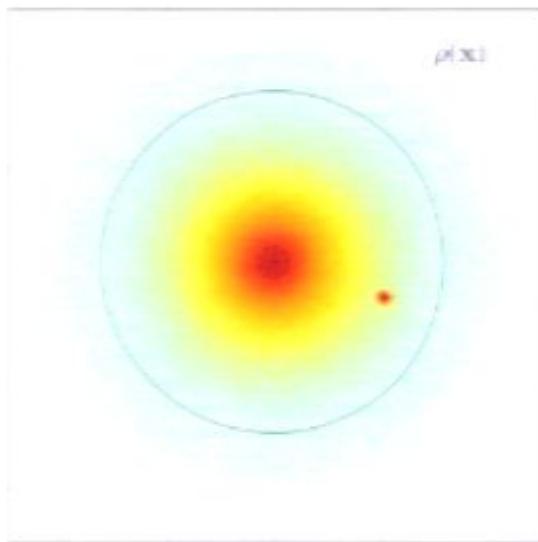


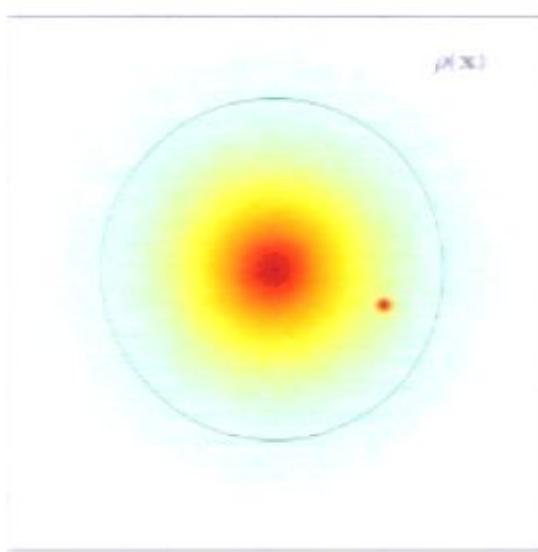
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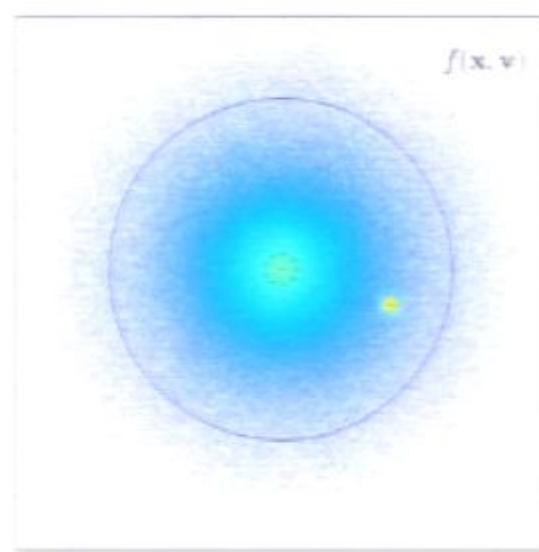




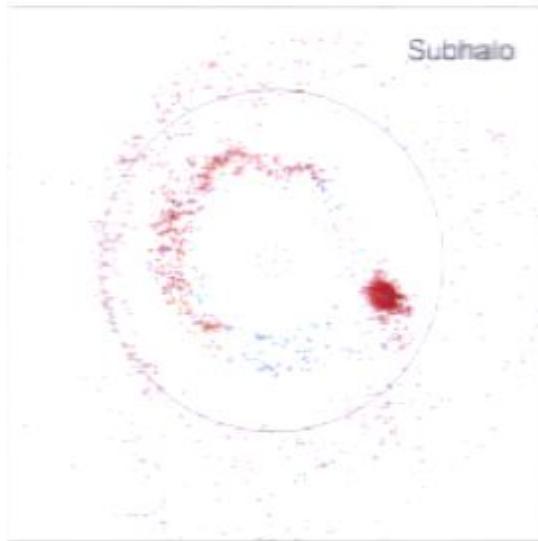




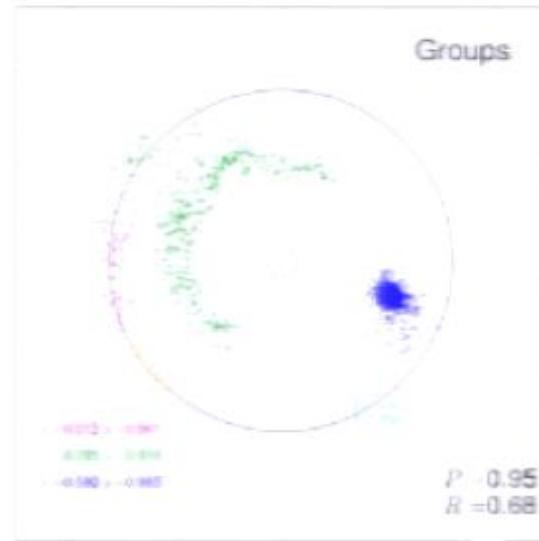
$\rho(\mathbf{x})$



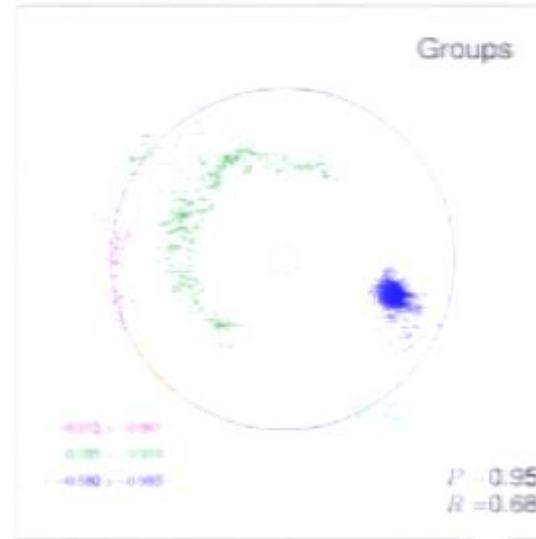
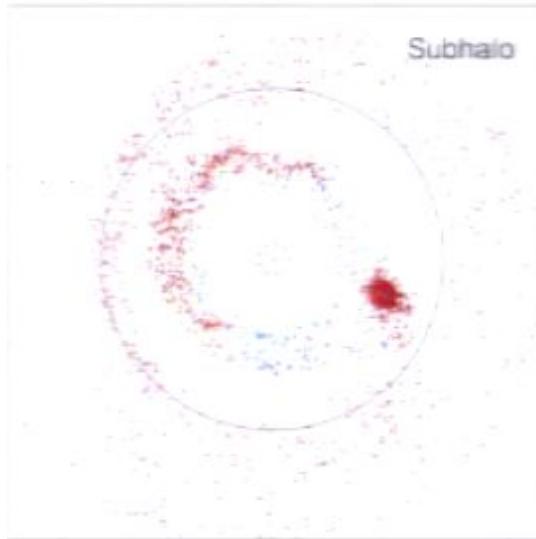
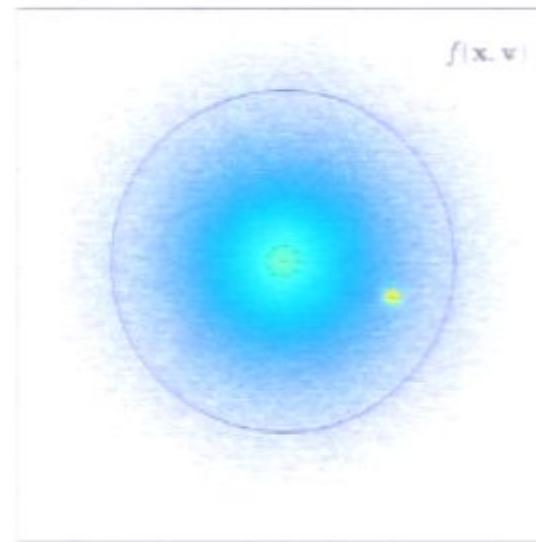
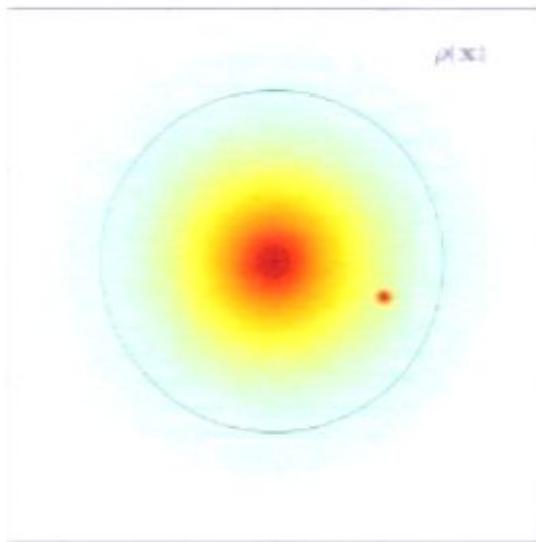
$f(\mathbf{x}, \mathbf{v})$

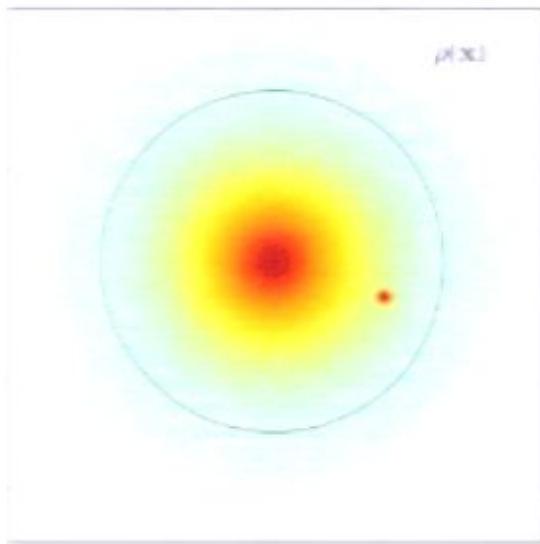


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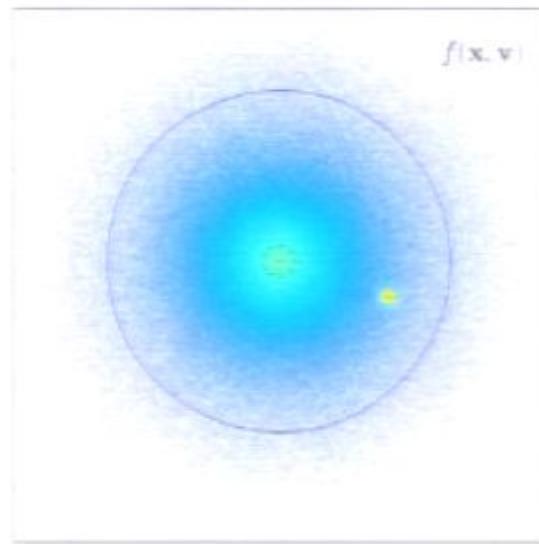


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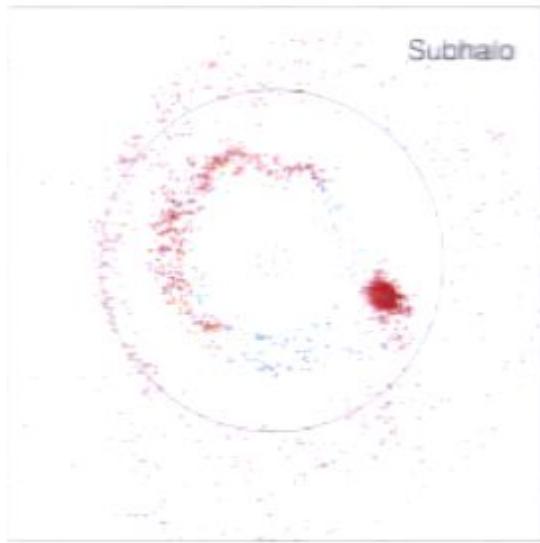




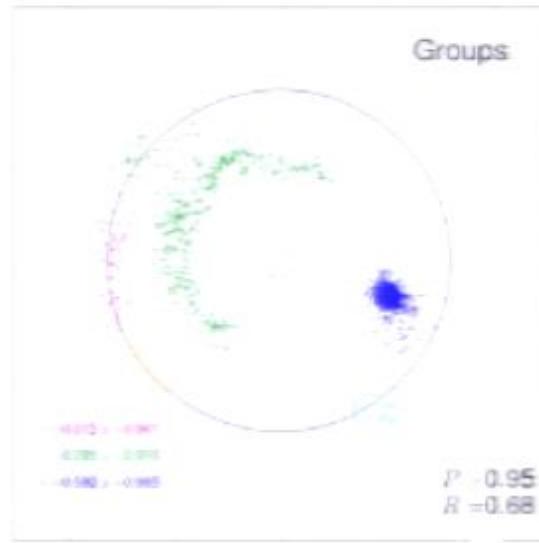
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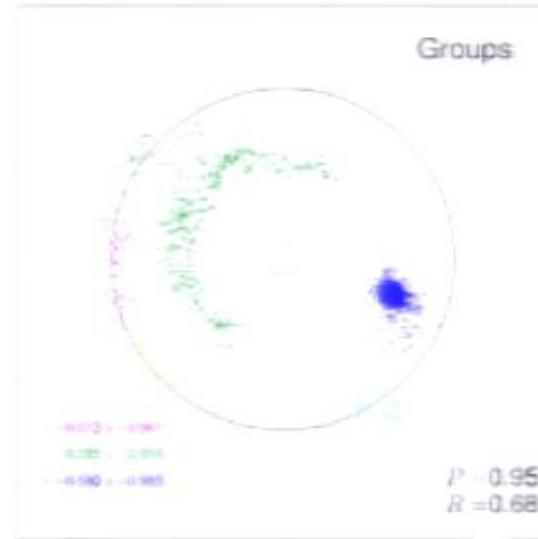
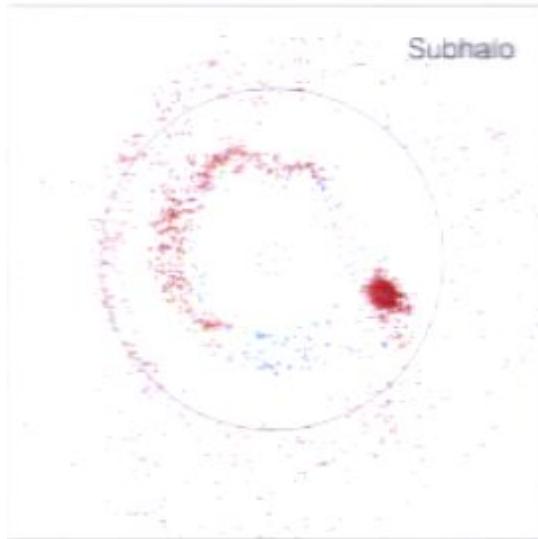
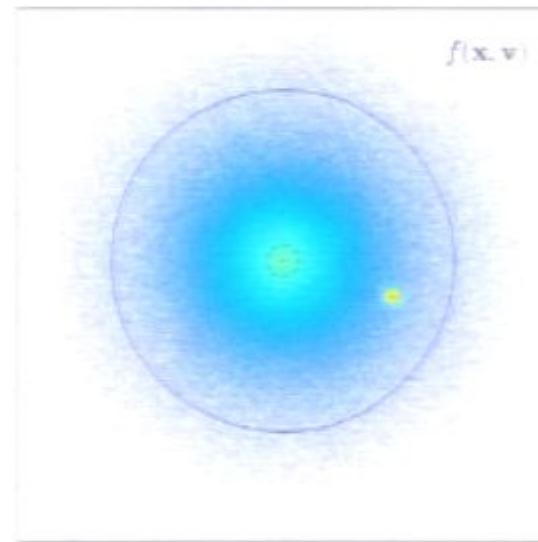
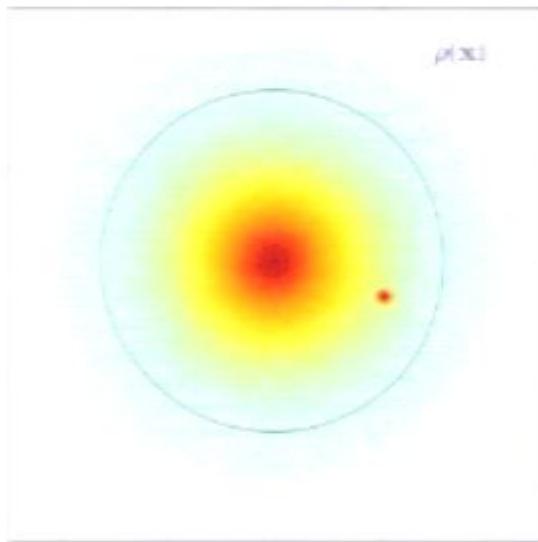


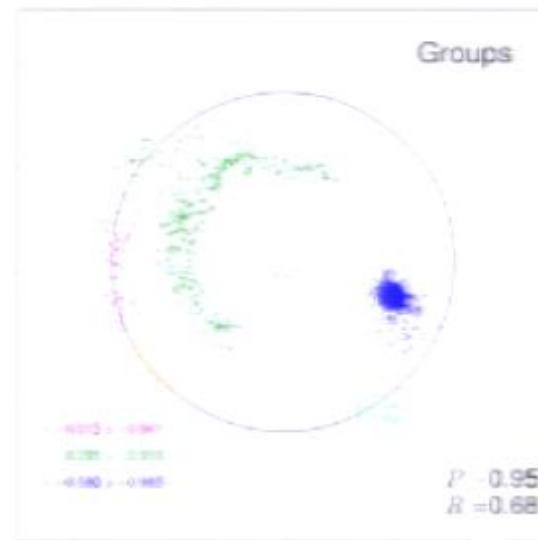
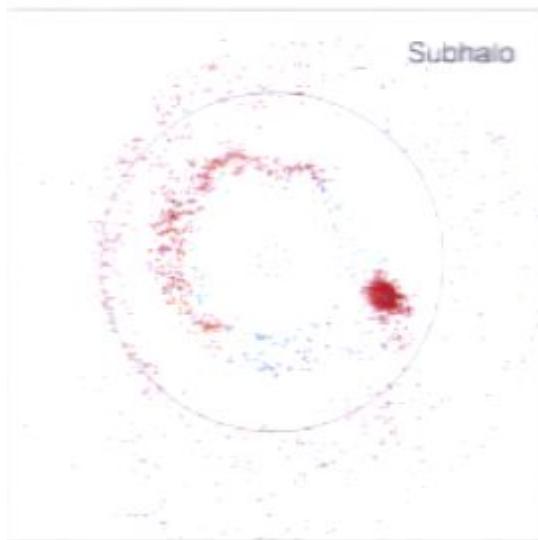
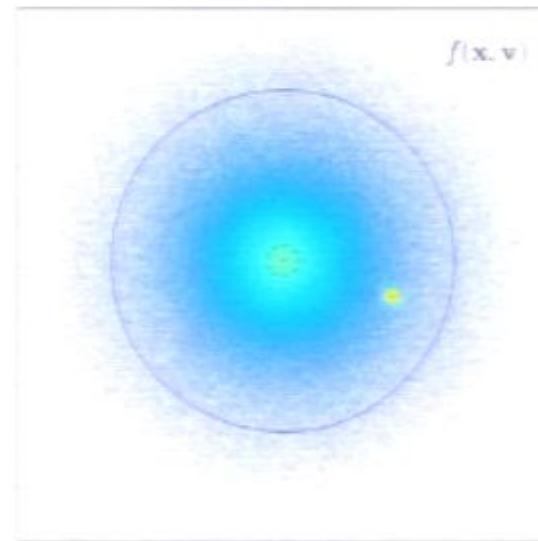
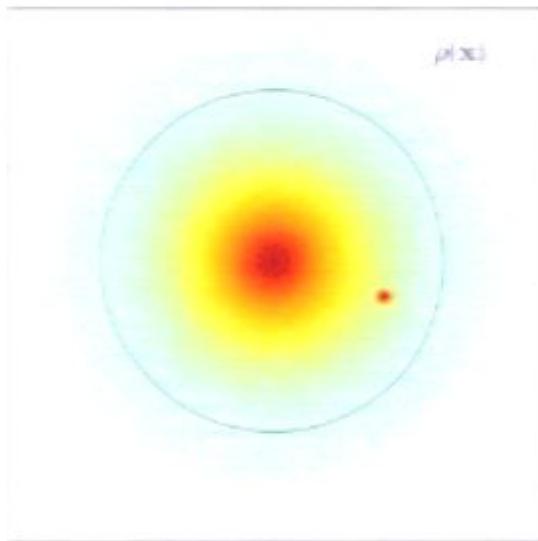
$f(\mathbf{x}, \mathbf{v})$

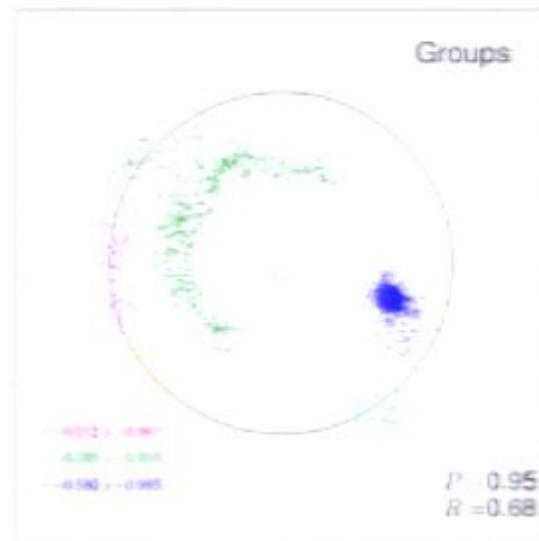
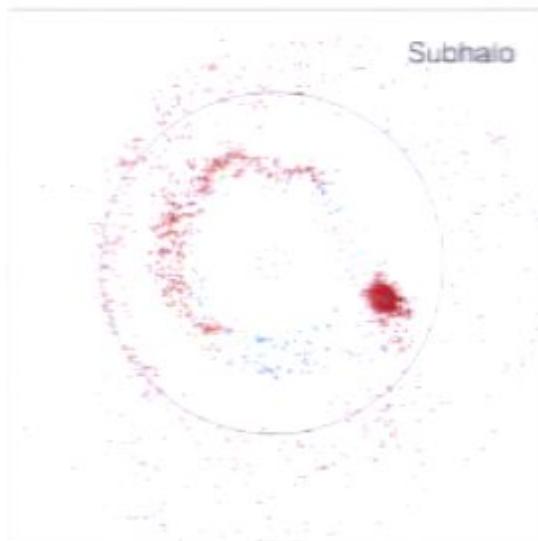
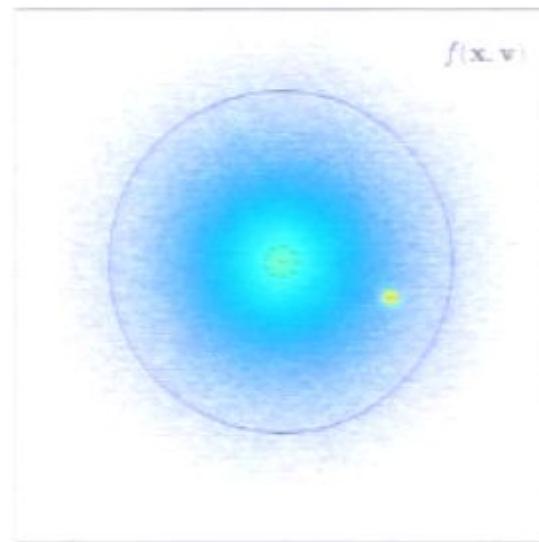
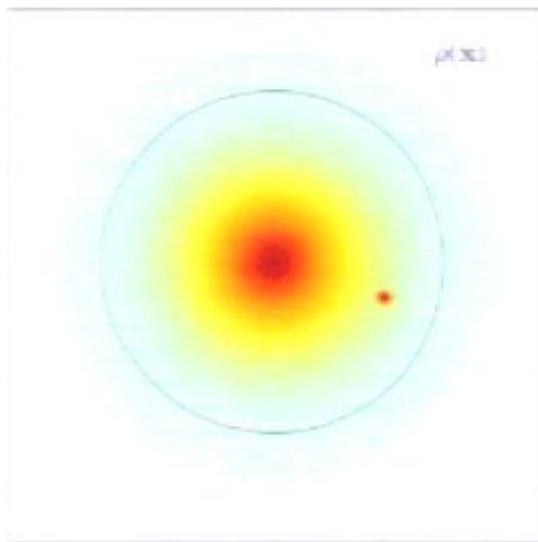


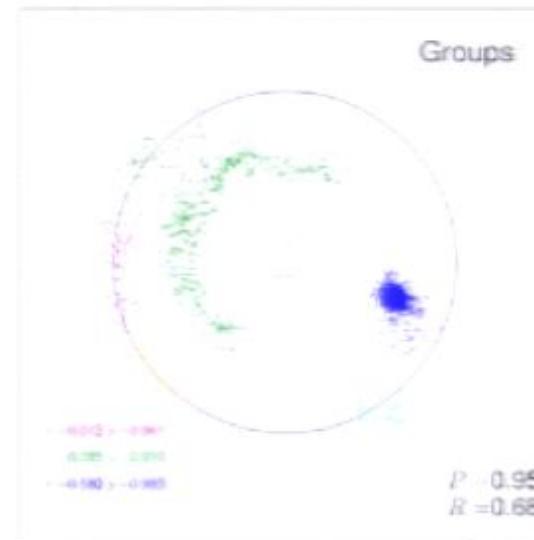
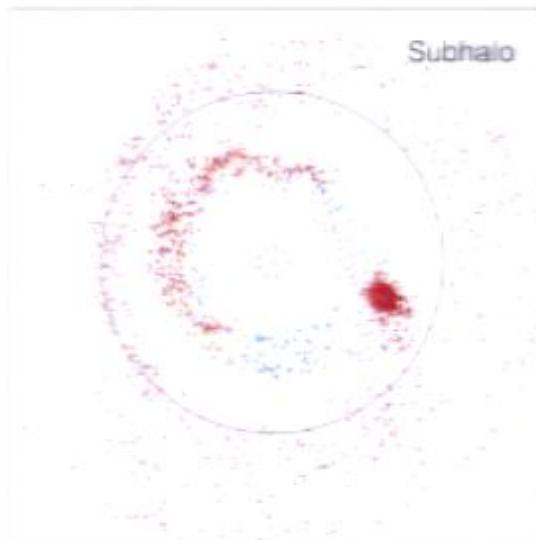
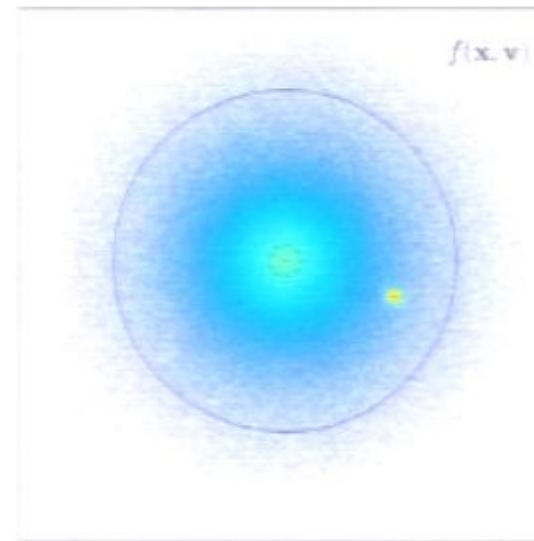
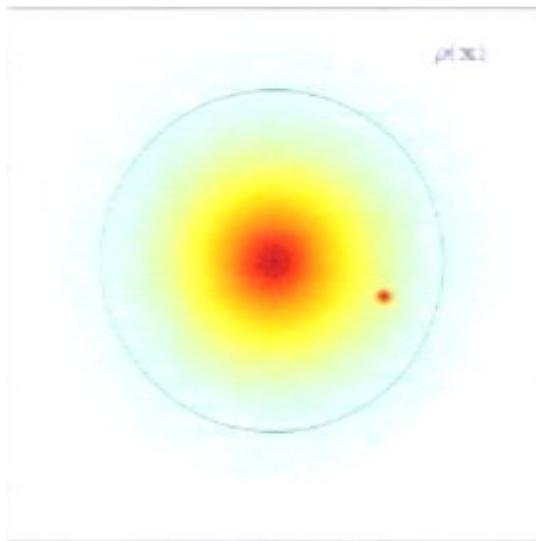
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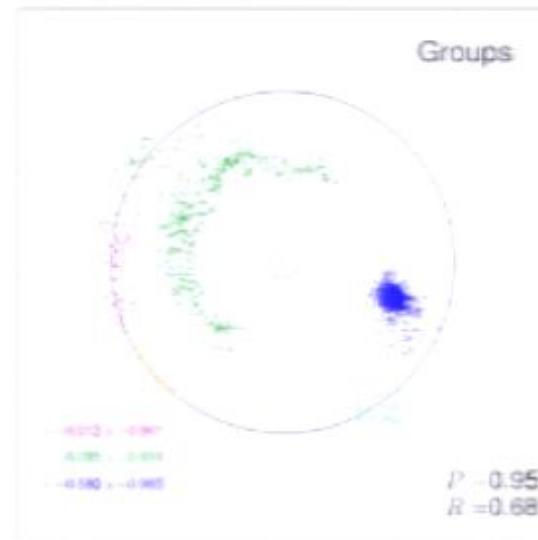
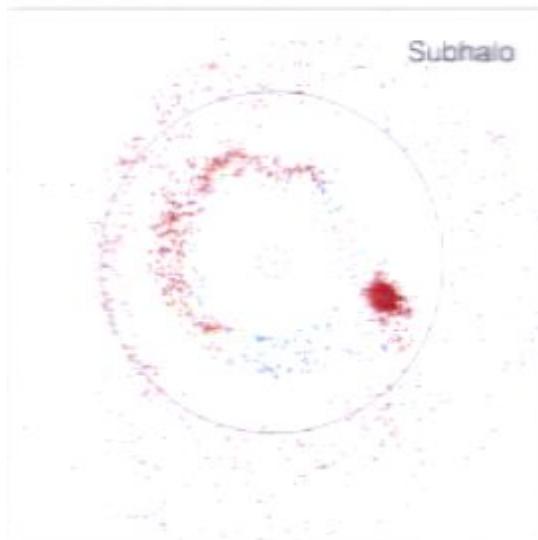
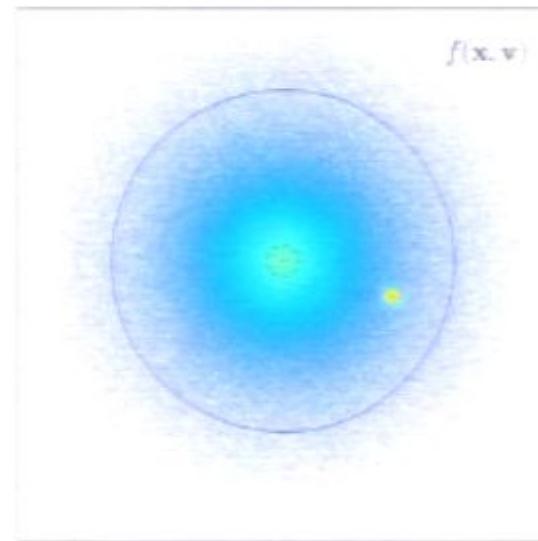
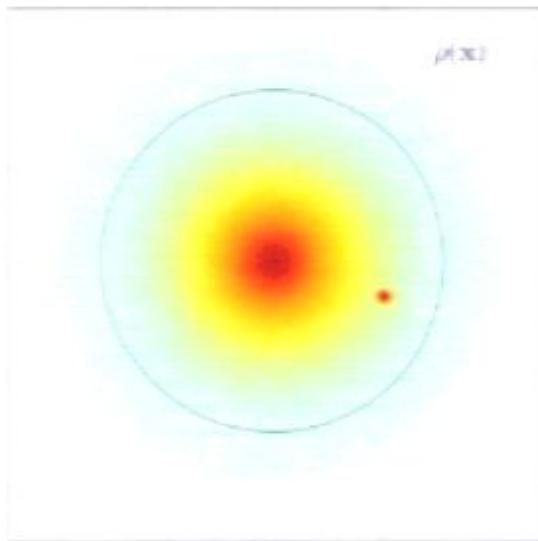


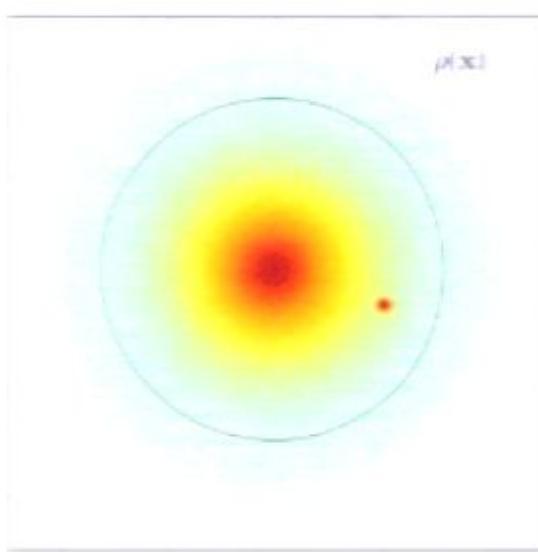




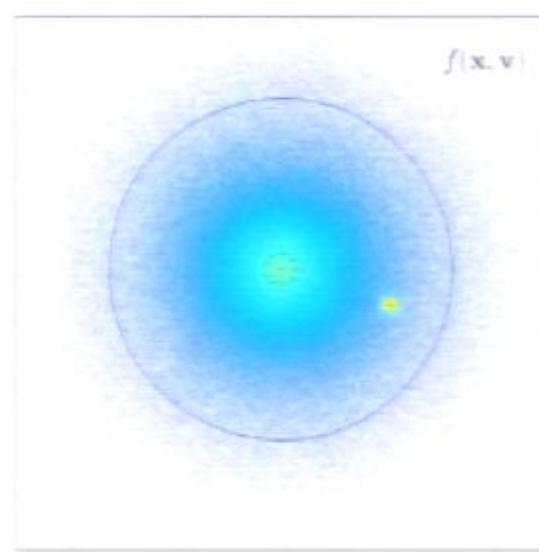




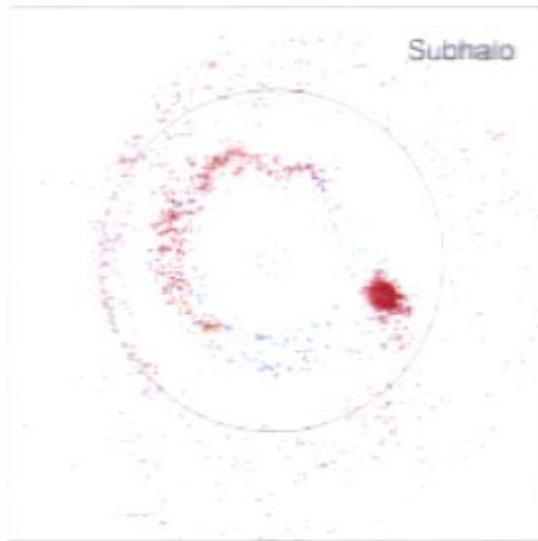




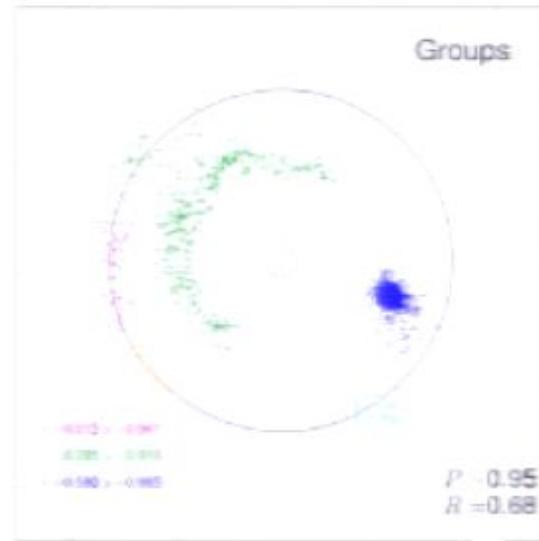
$\rho(\mathbf{x})$



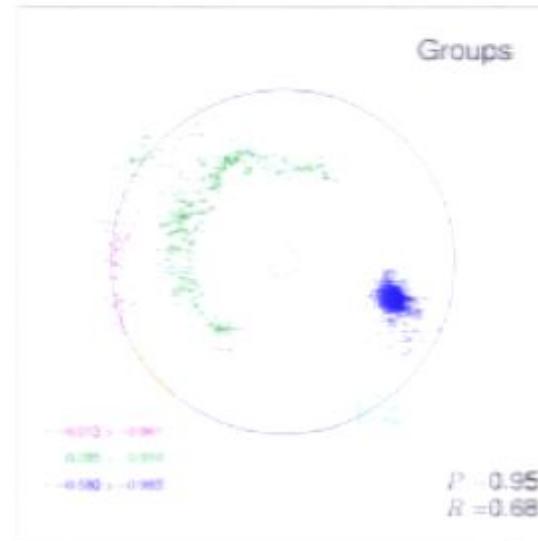
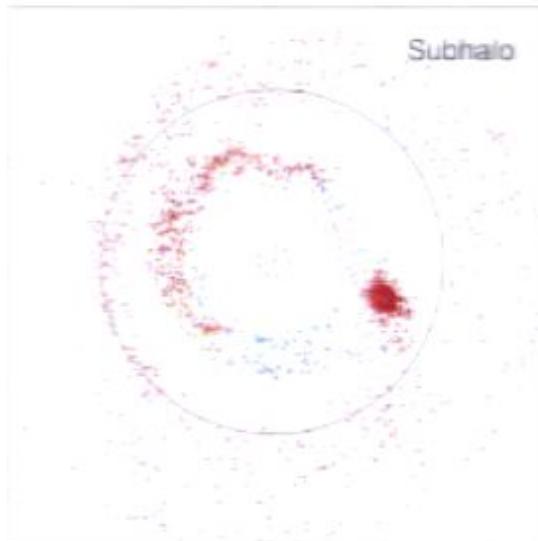
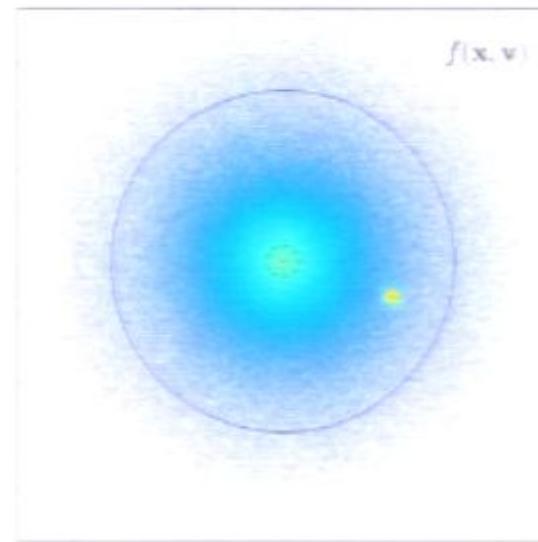
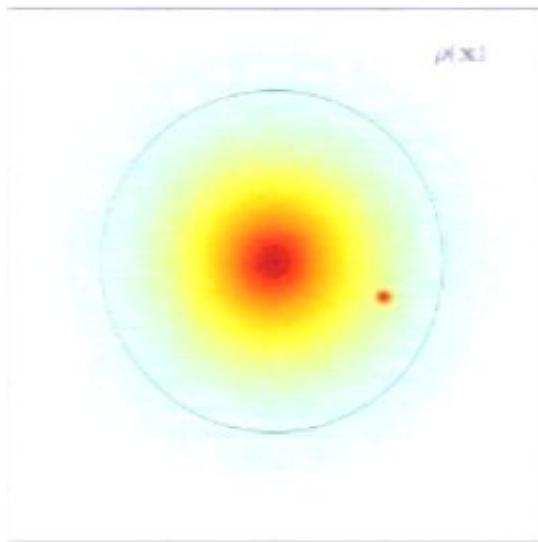
$f(\mathbf{x}, \mathbf{v})$

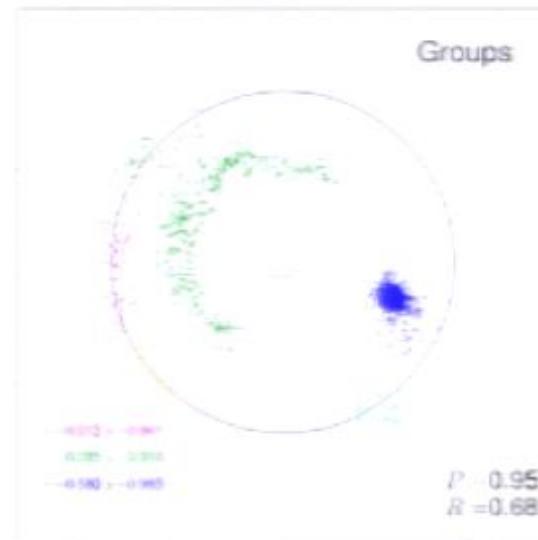
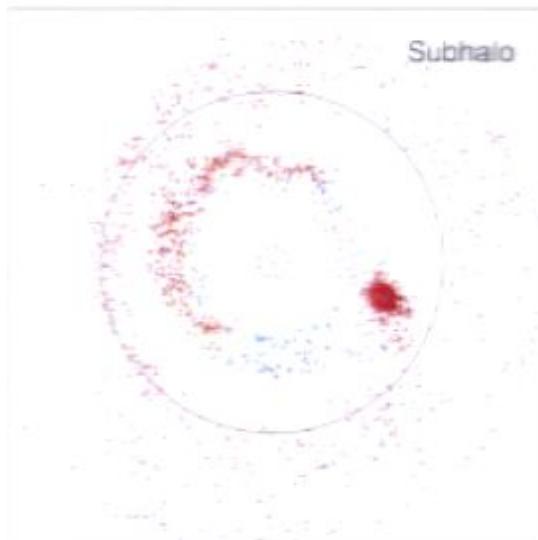
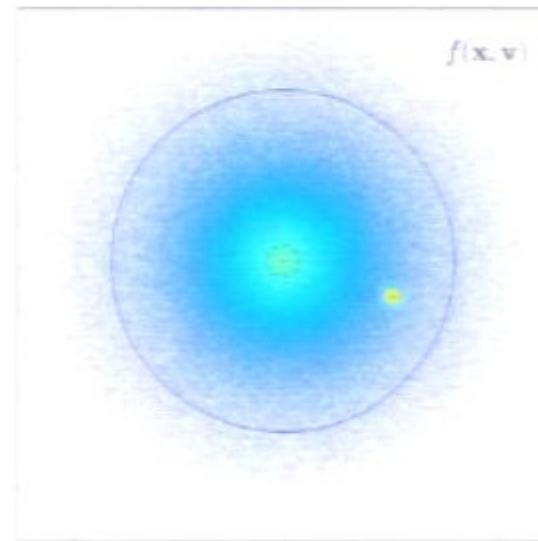
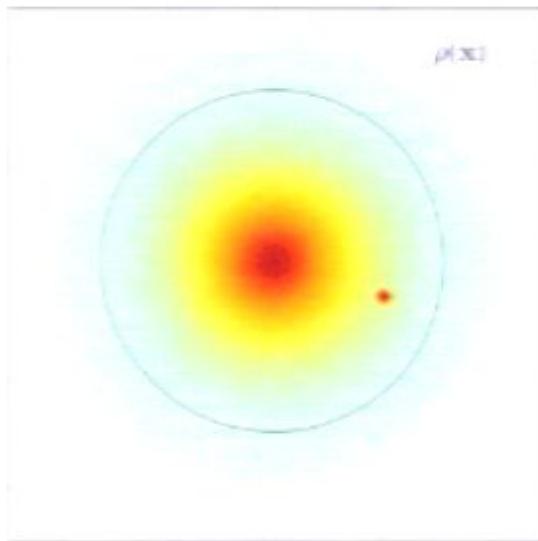


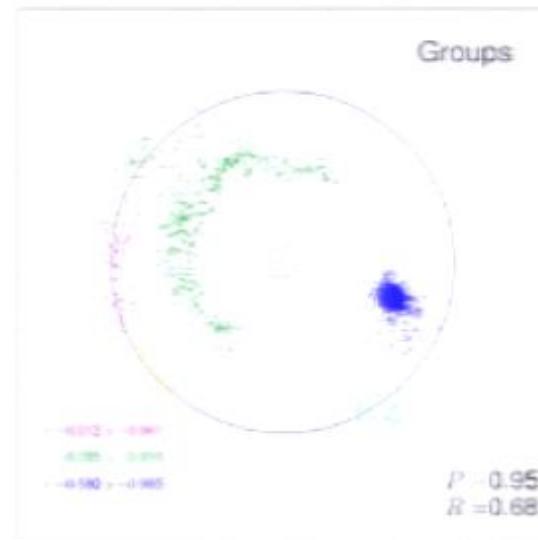
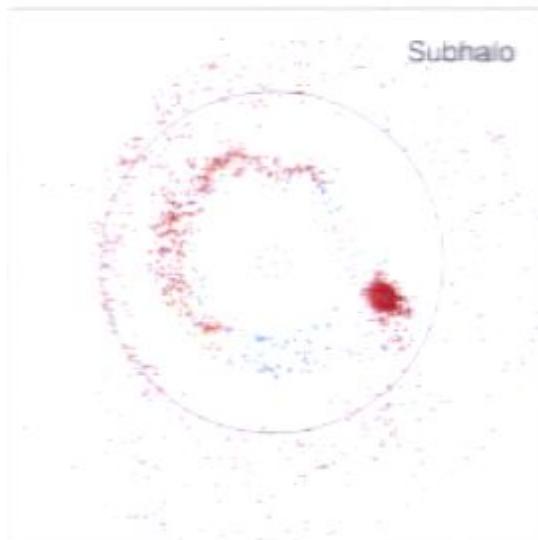
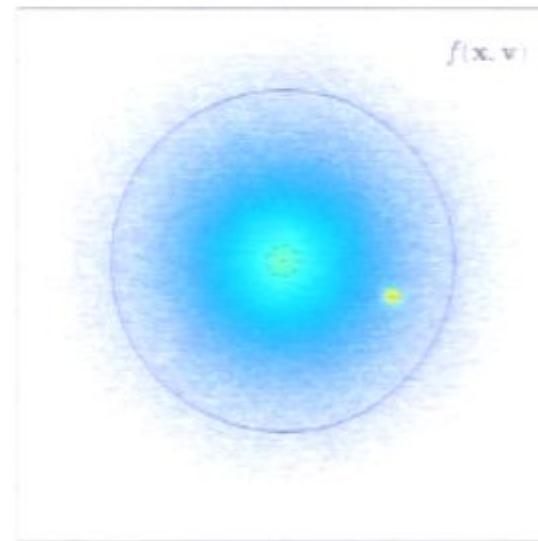
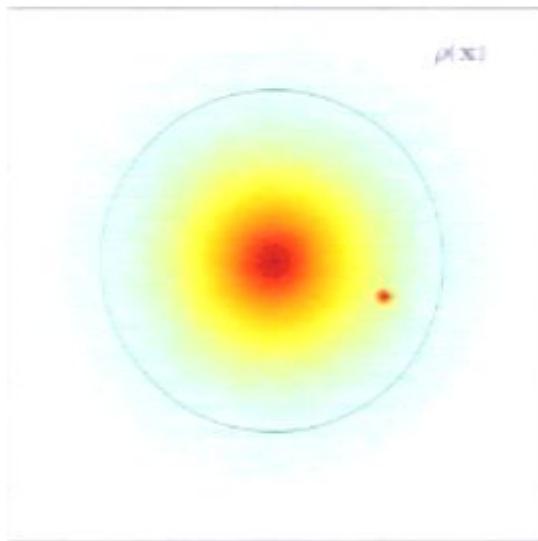
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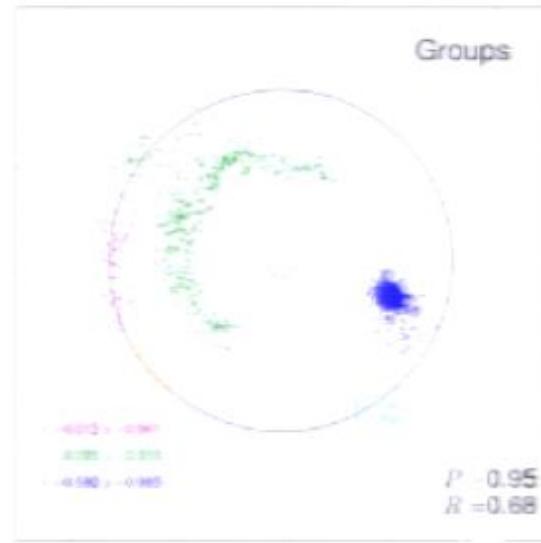
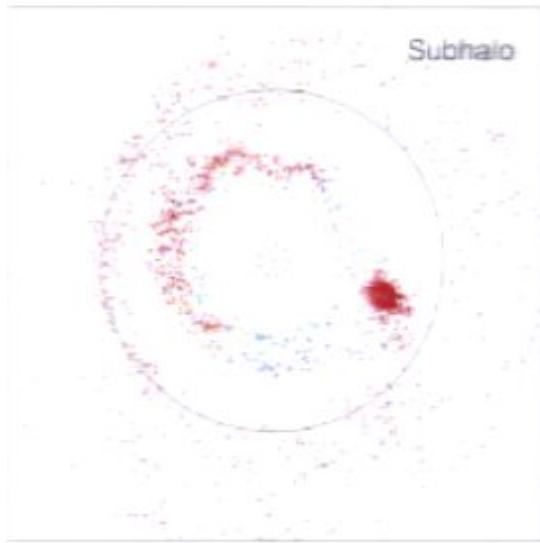
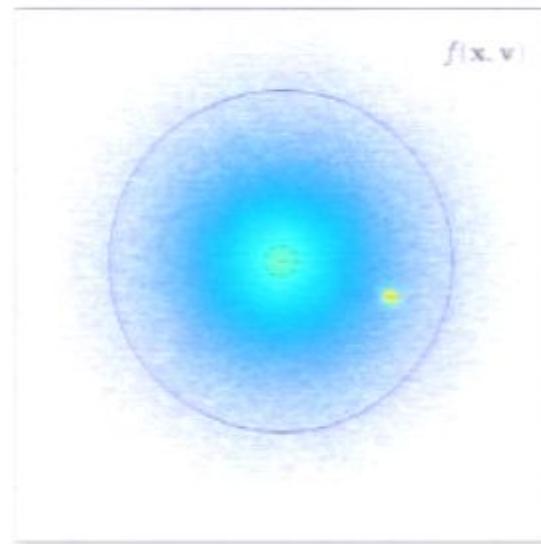
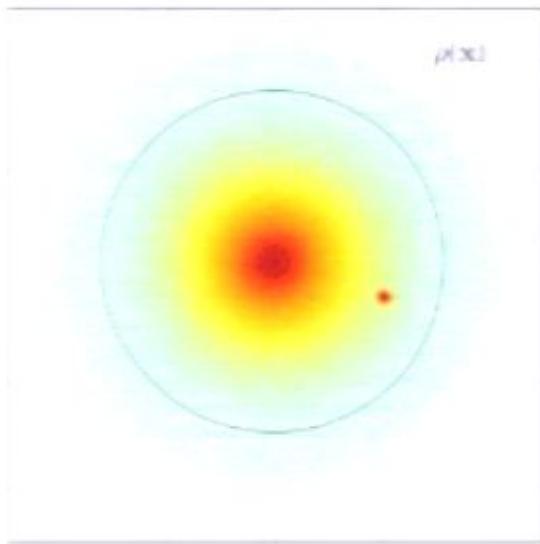


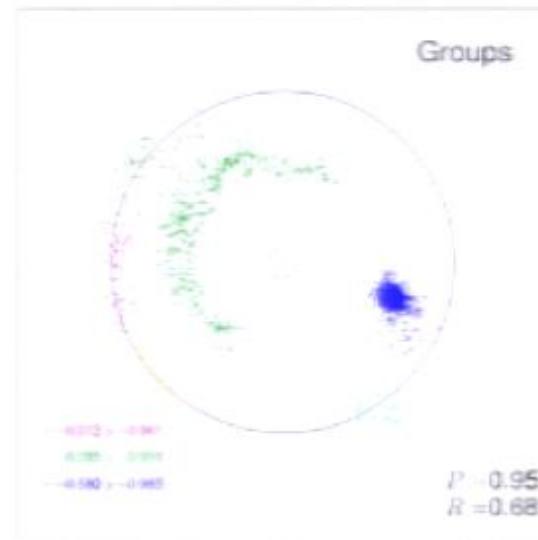
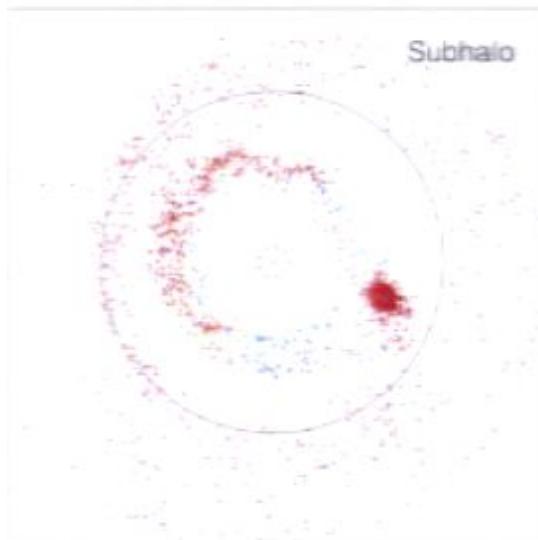
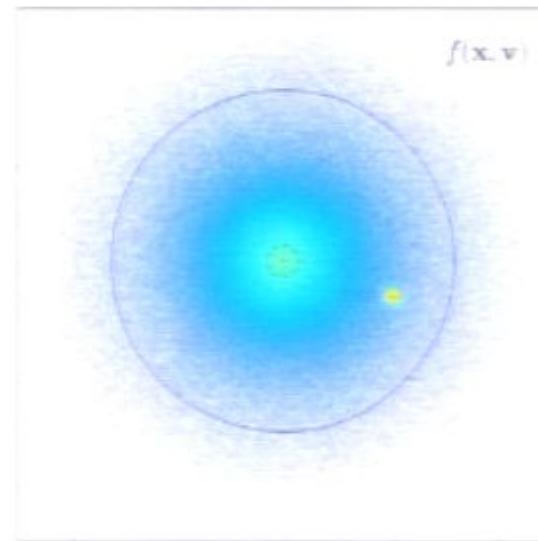
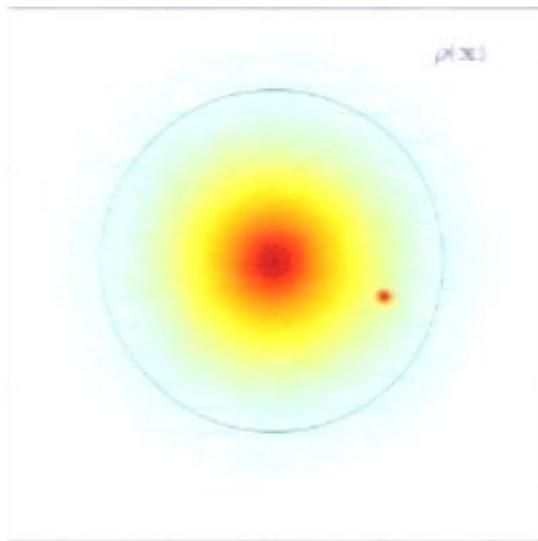
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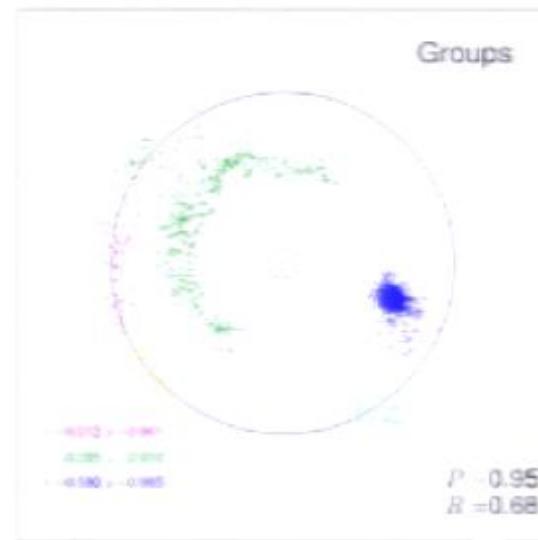
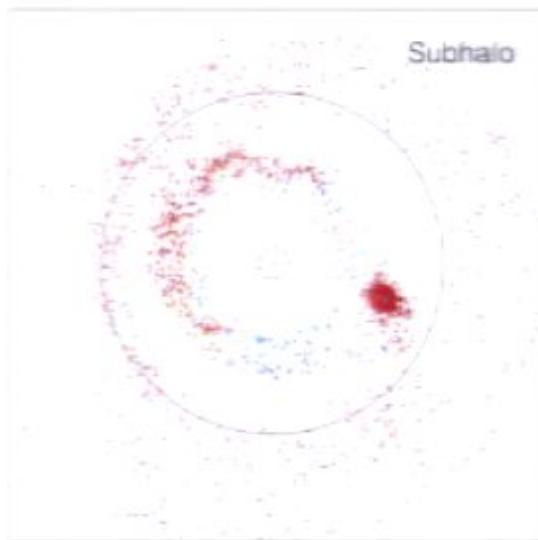
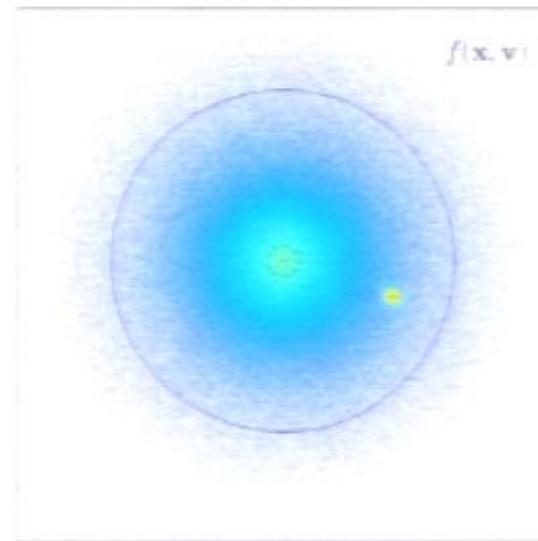
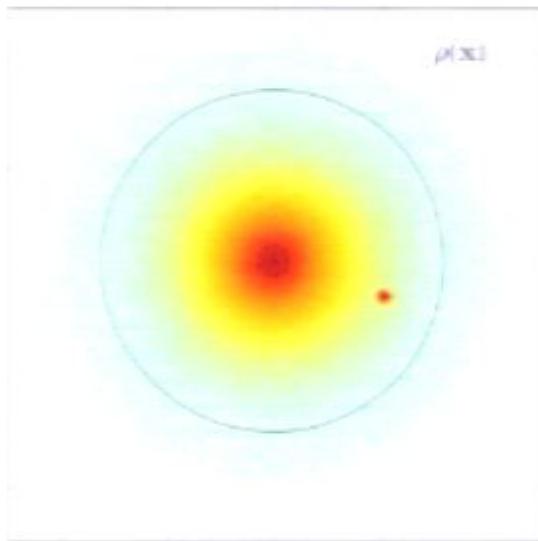


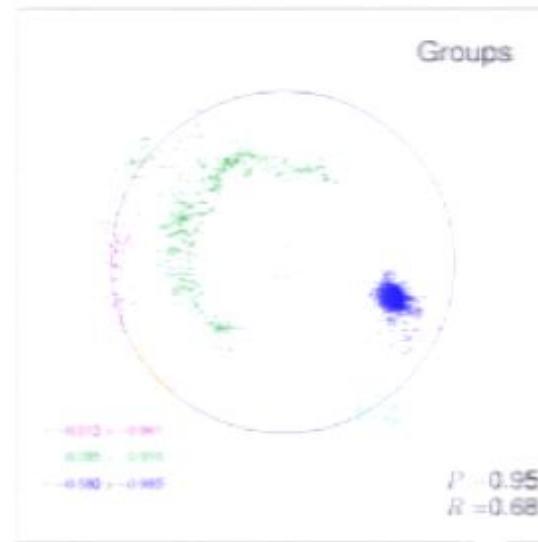
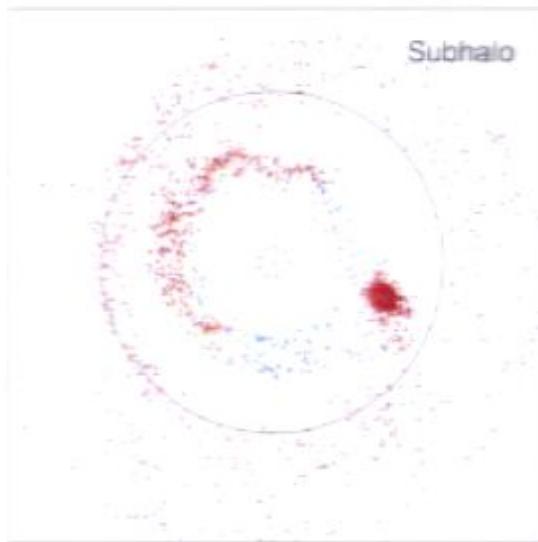
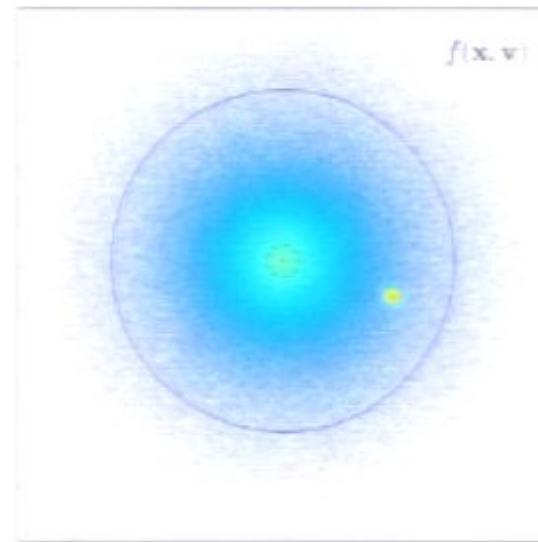
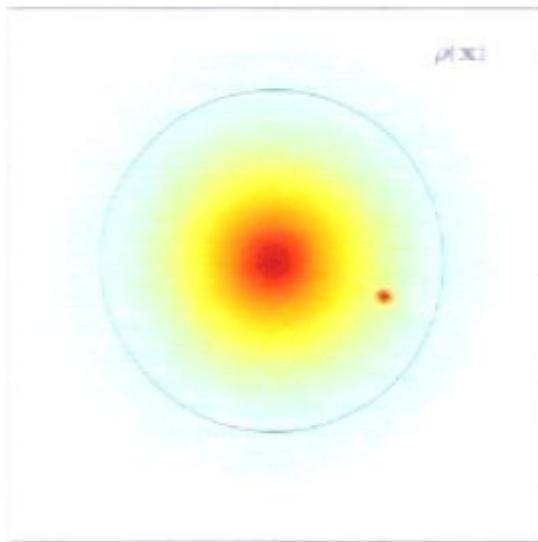


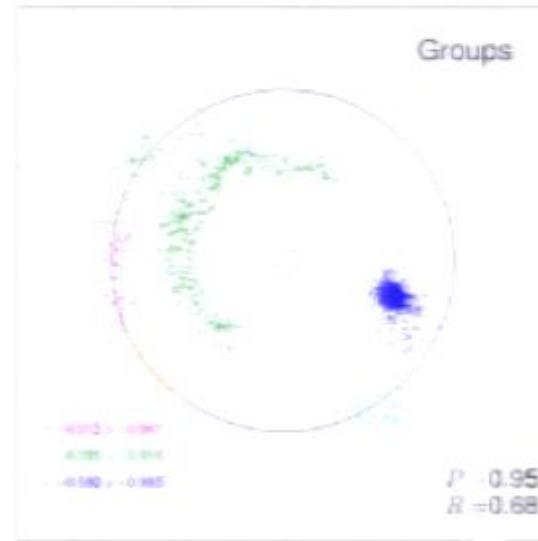
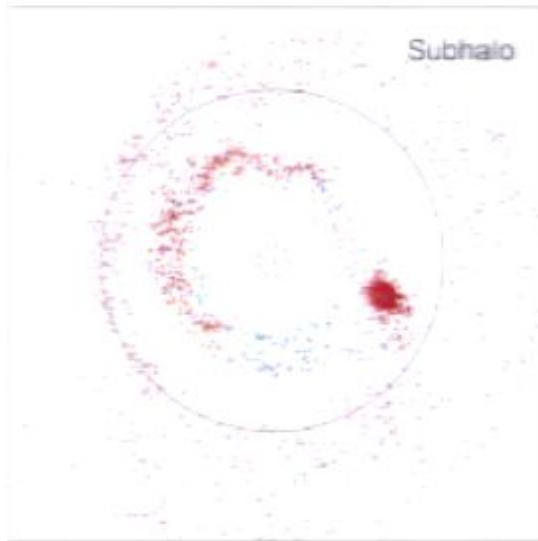
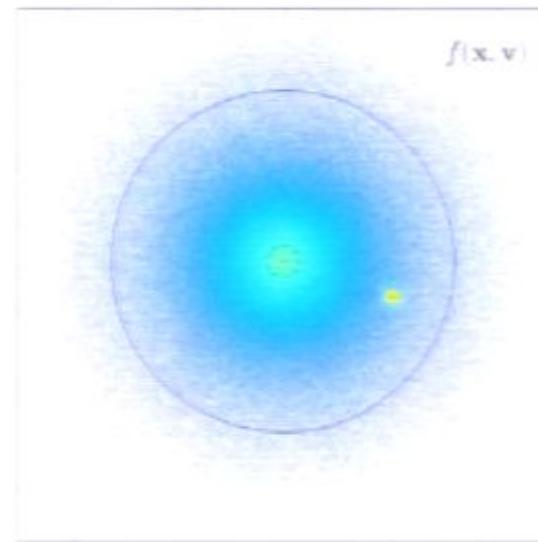
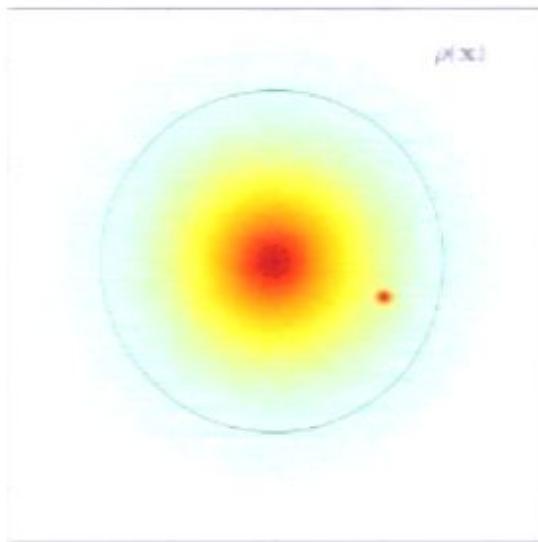


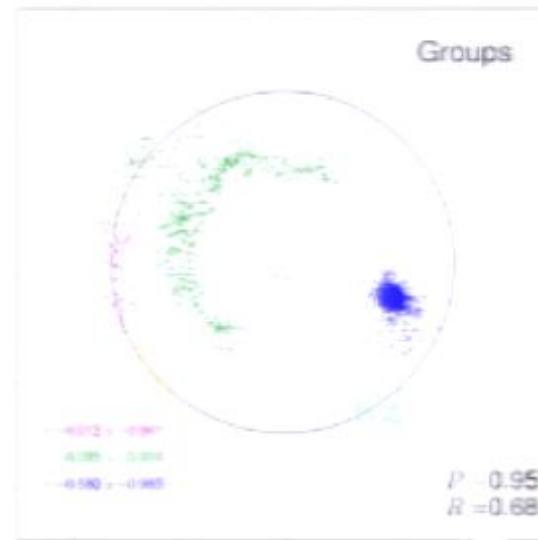
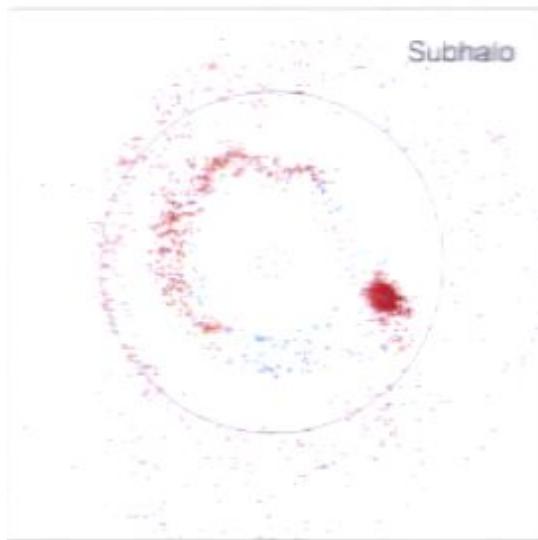
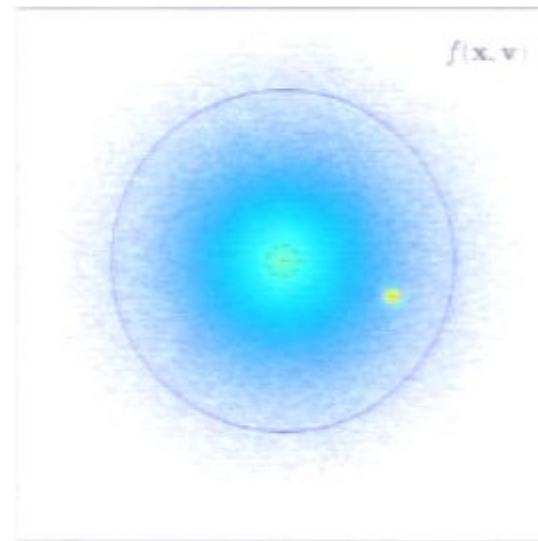
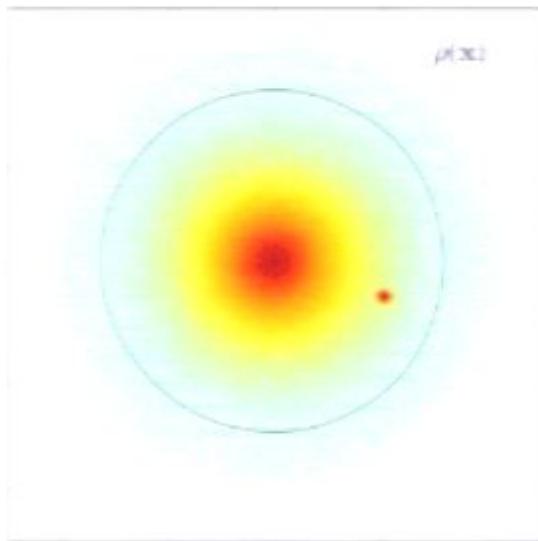


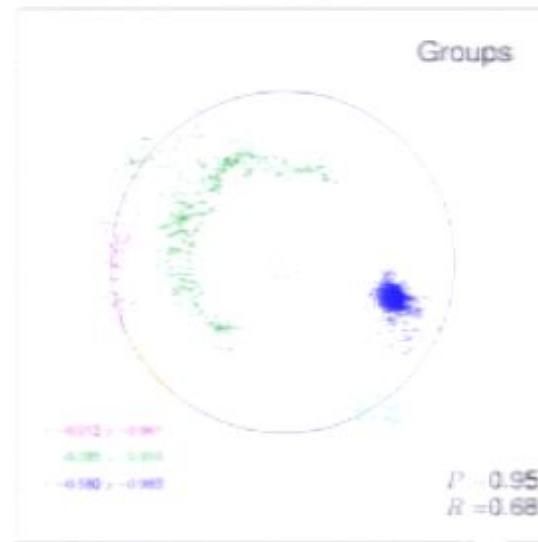
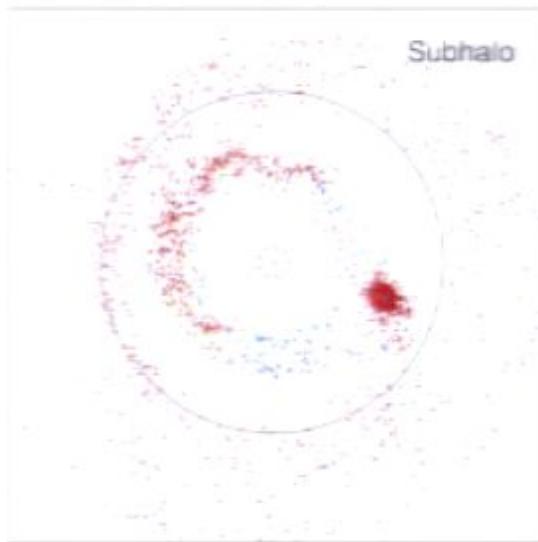
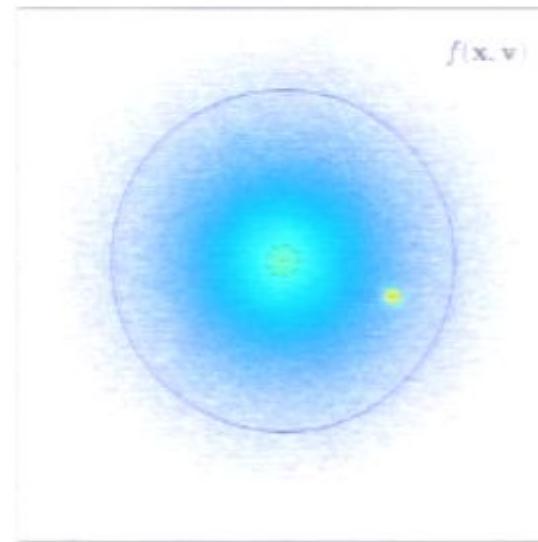
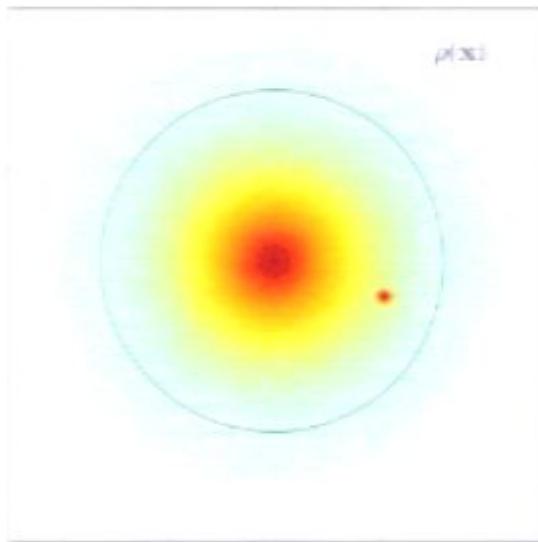


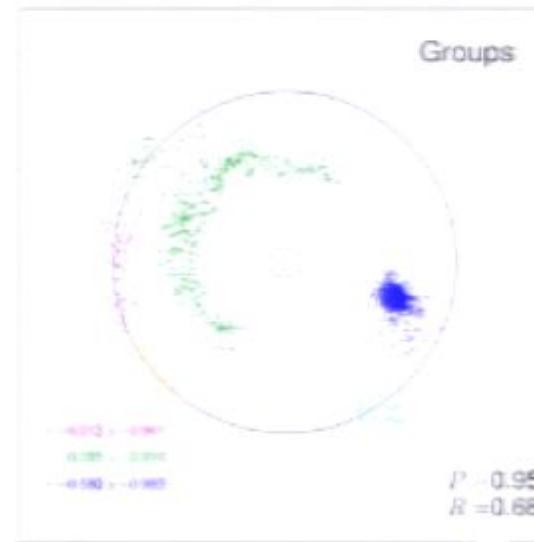
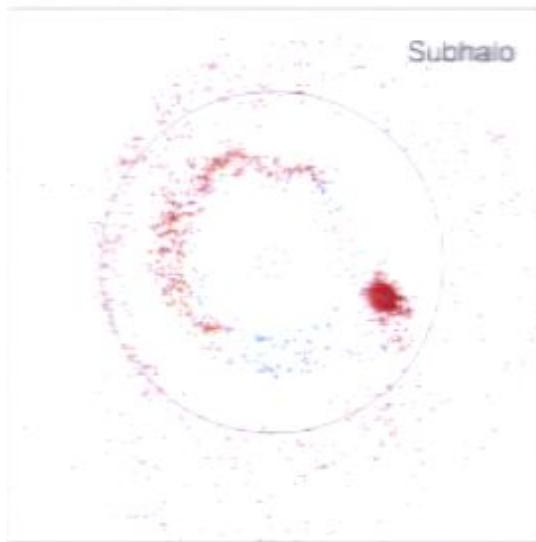
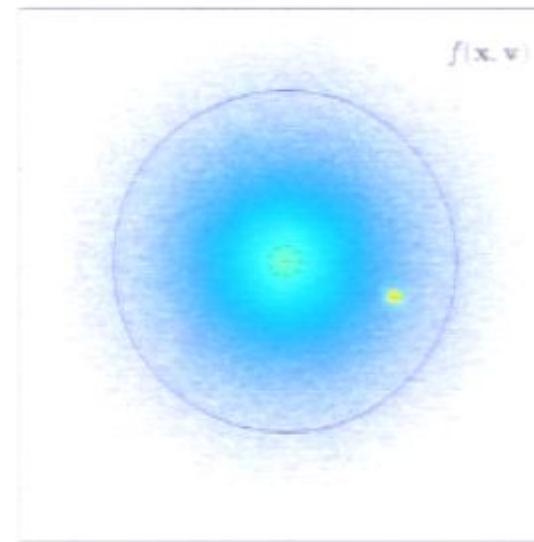
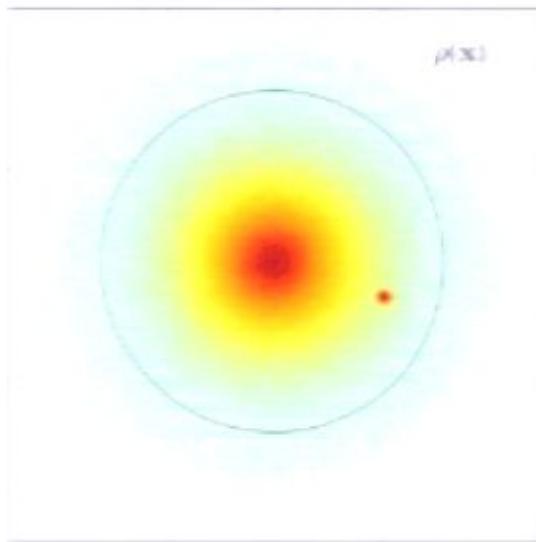


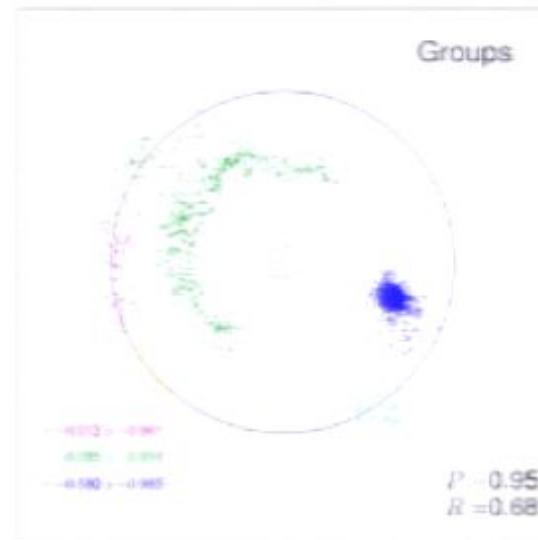
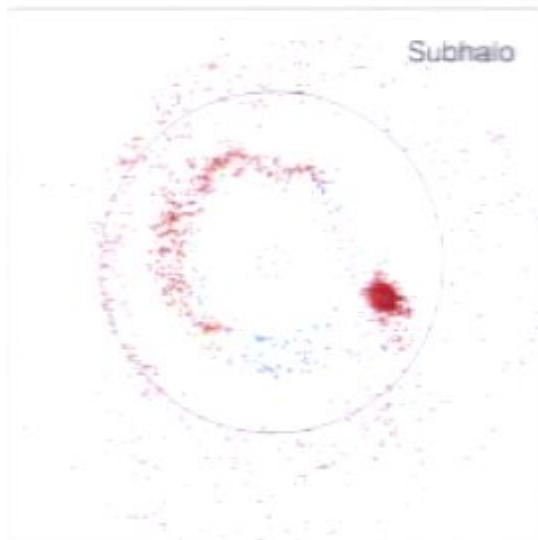
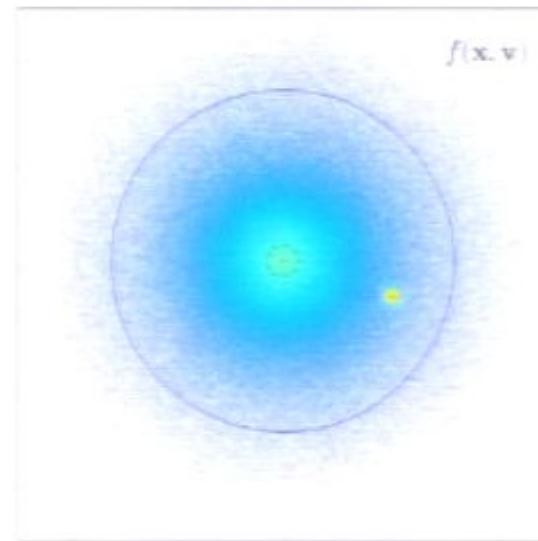
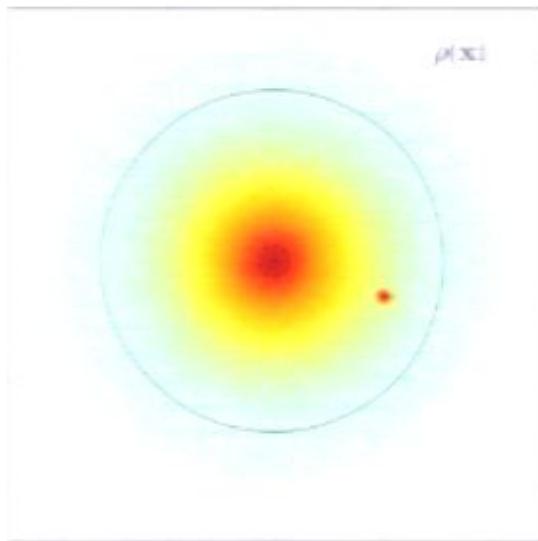


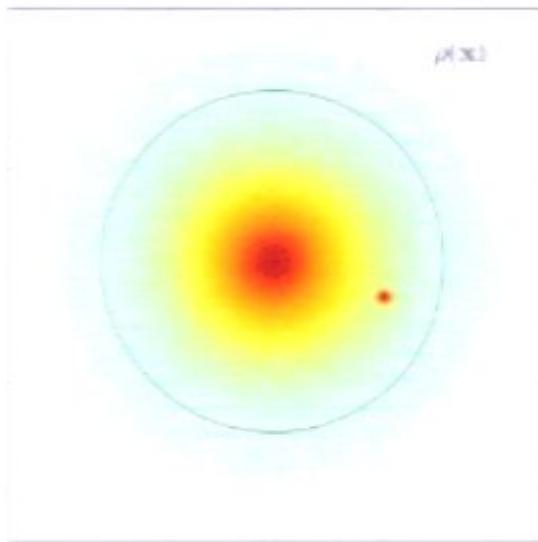




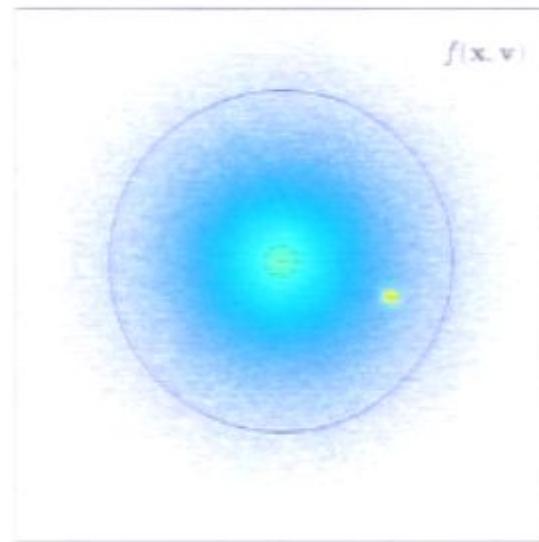




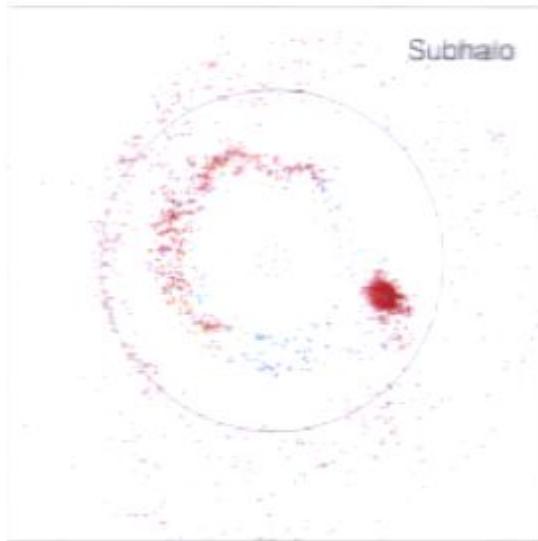




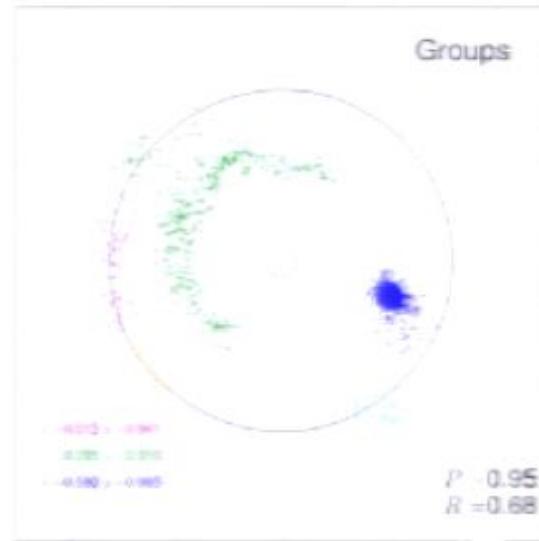
$\rho(\mathbf{x})$



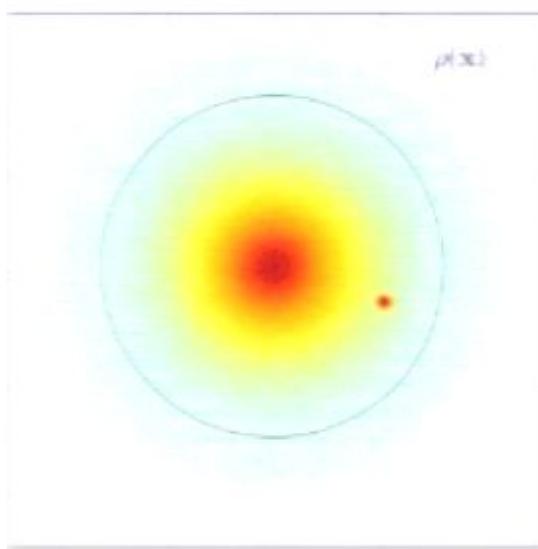
$f(\mathbf{x}, \mathbf{v})$



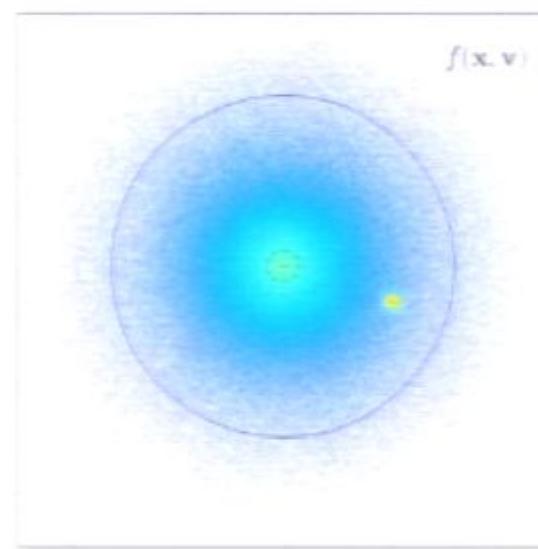
Subhalo



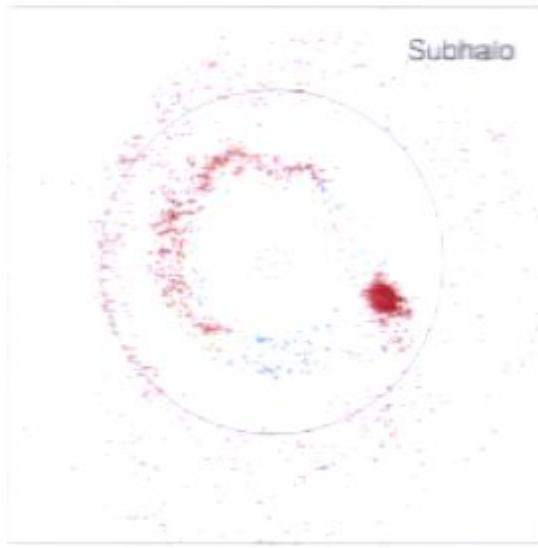
Groups



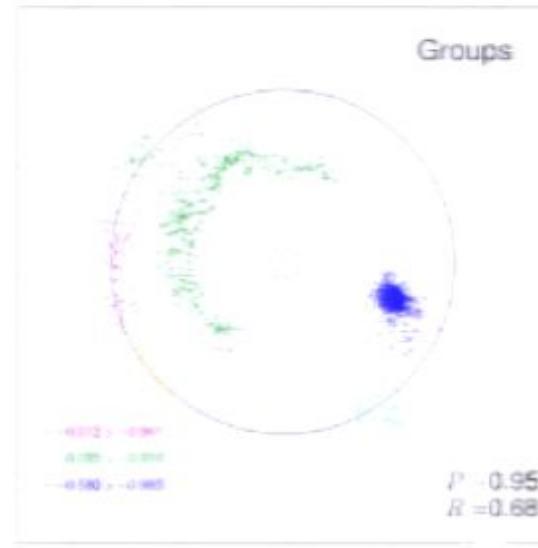
$\rho(\mathbf{x}_i)$



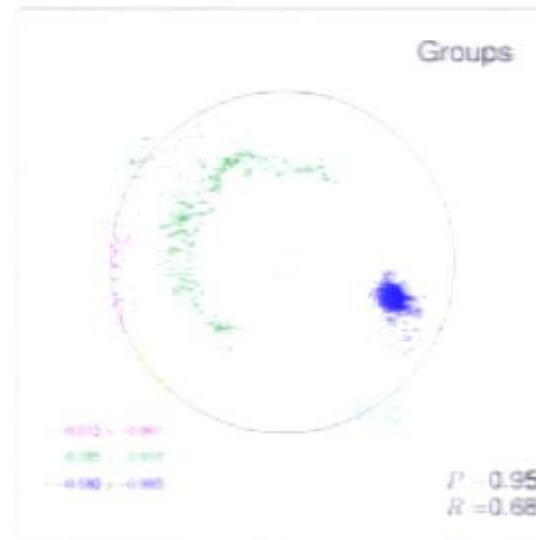
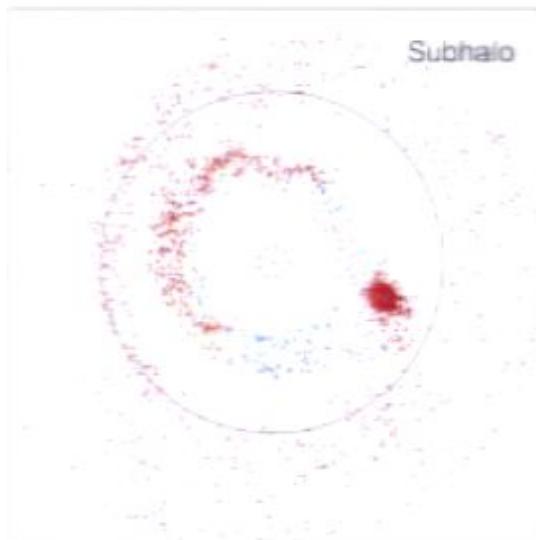
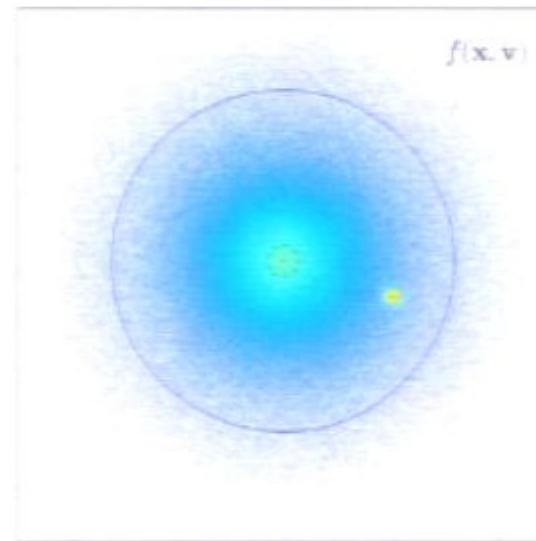
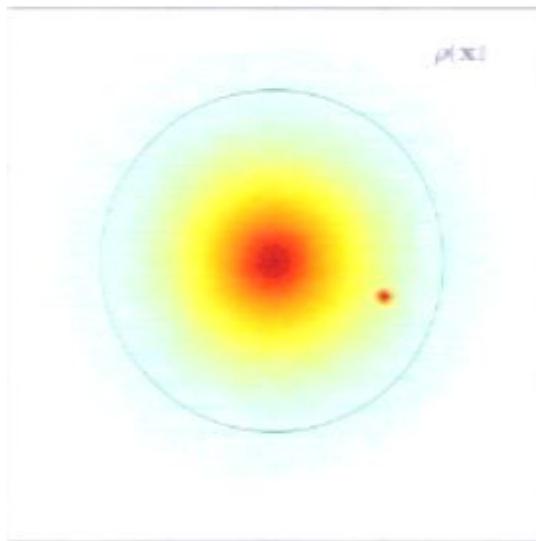
$f(\mathbf{x}, \mathbf{v})$

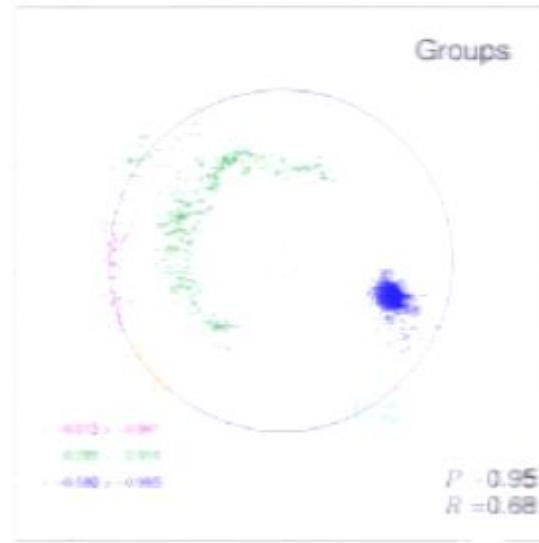
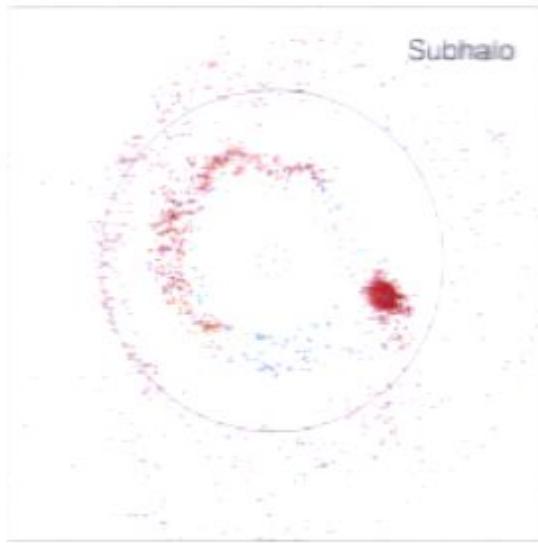
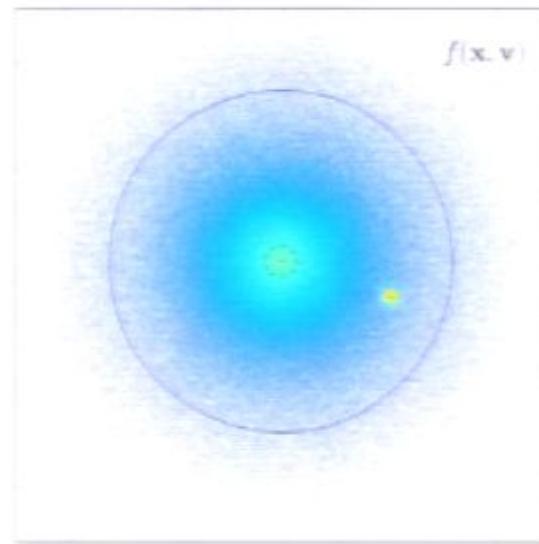
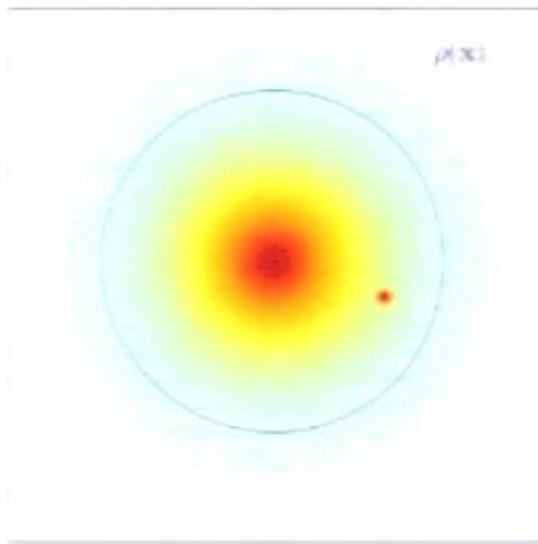


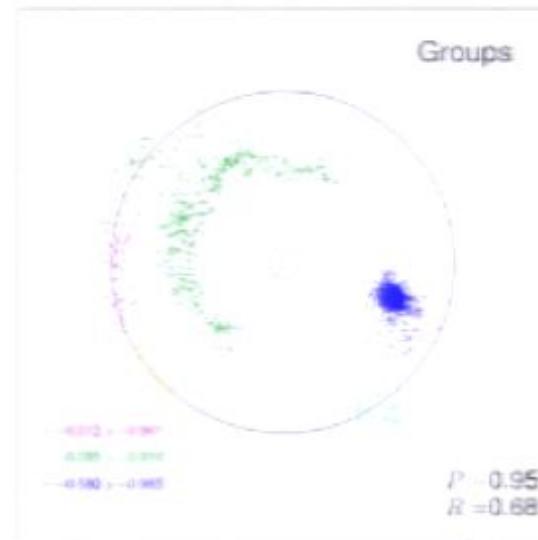
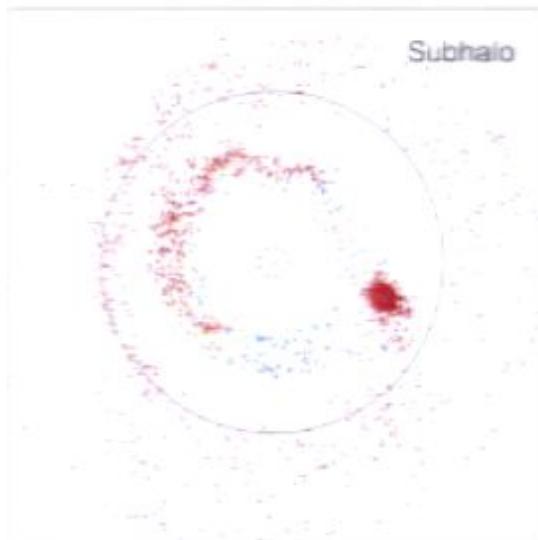
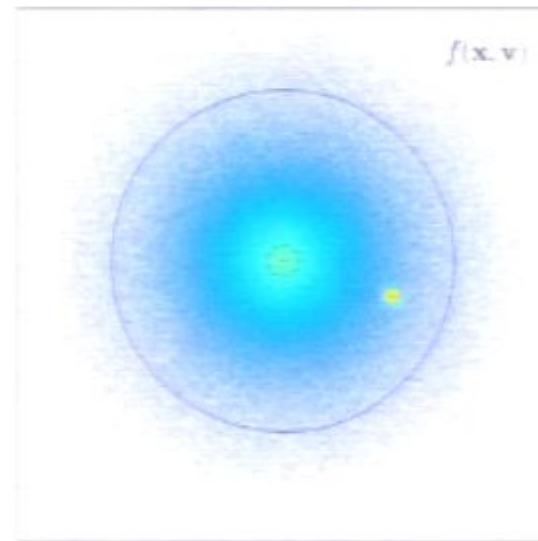
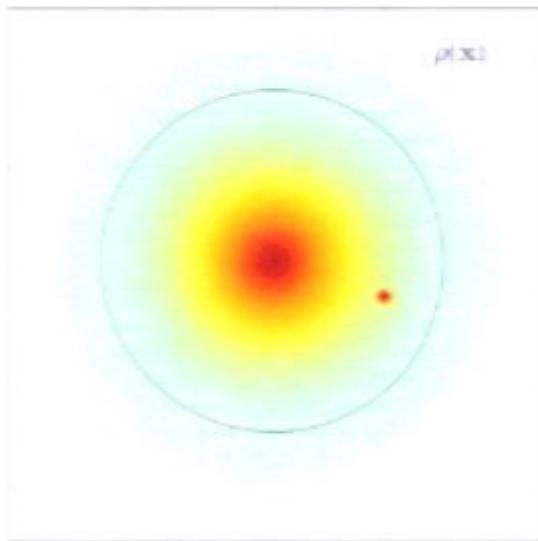
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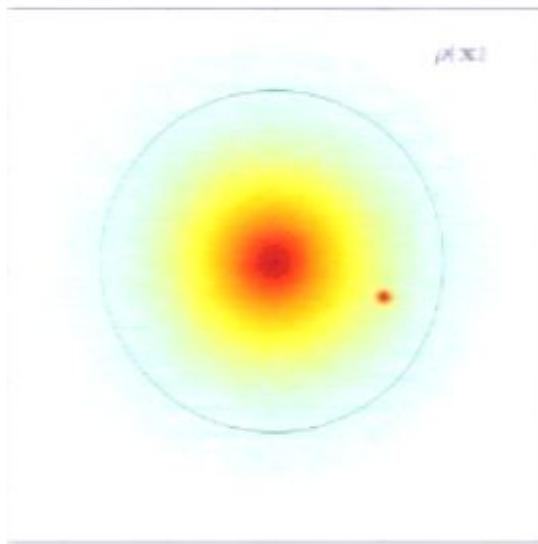


Groups

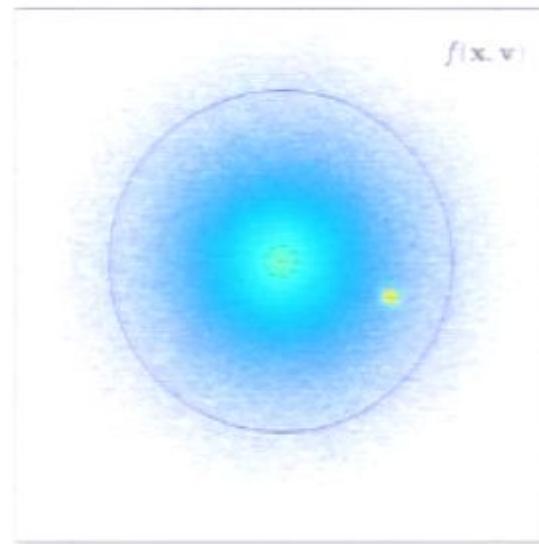




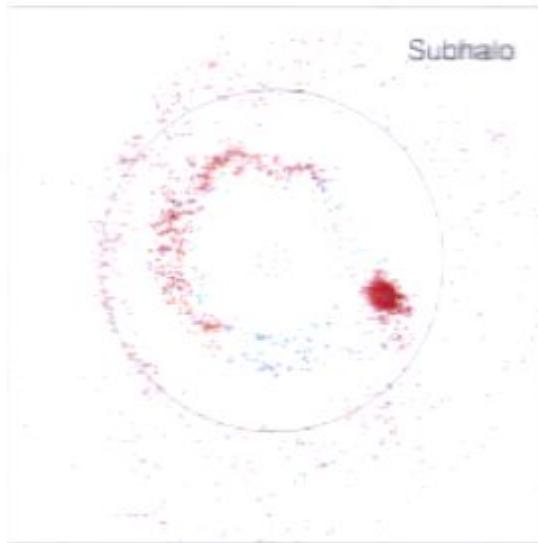




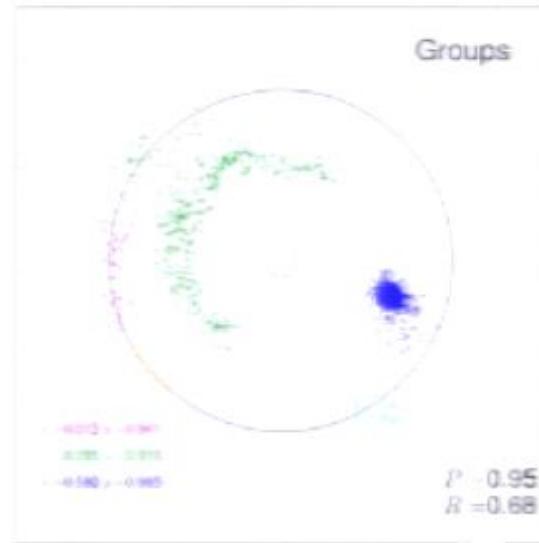
$\rho(\mathbf{x})$



$f(\mathbf{x}, \mathbf{v})$



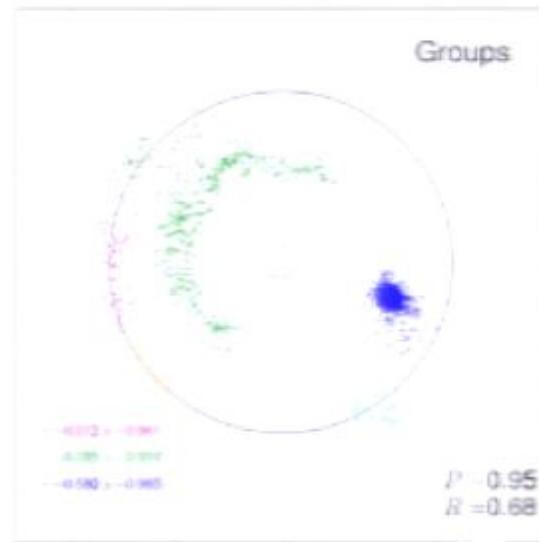
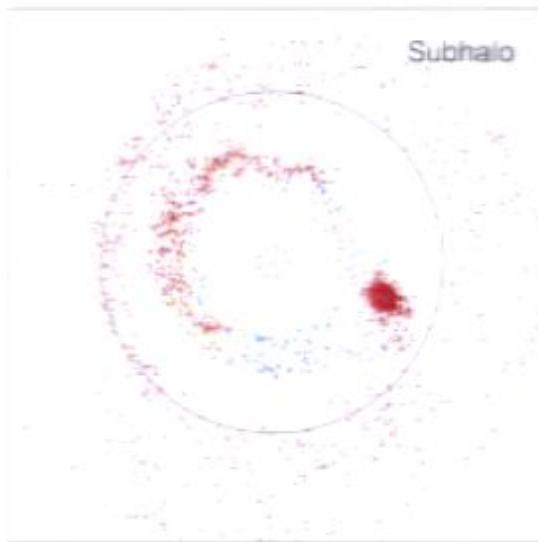
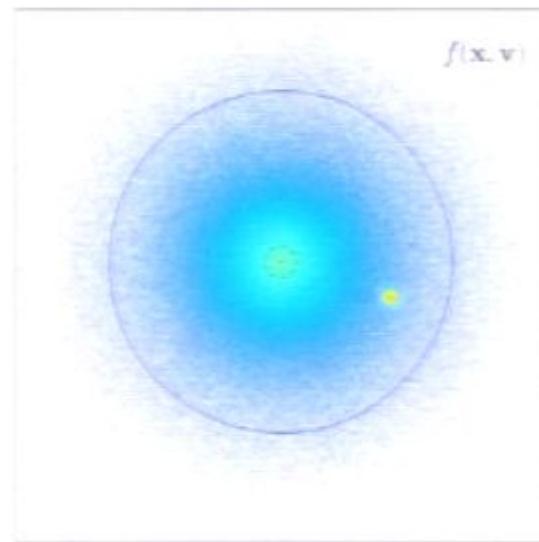
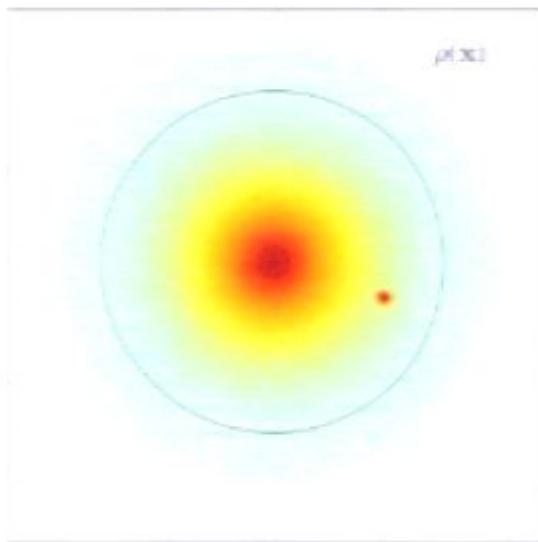
Subhalo

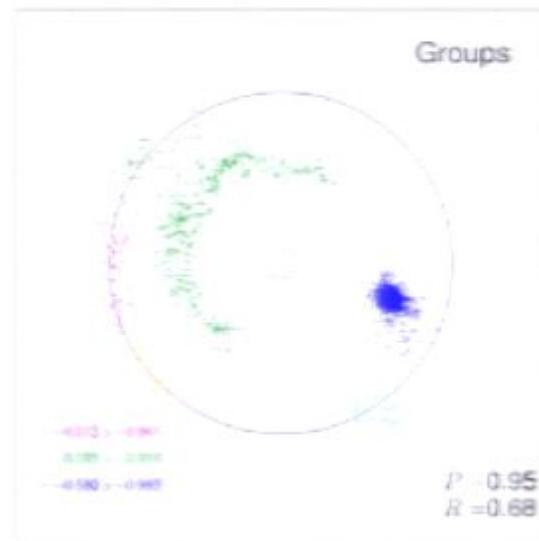
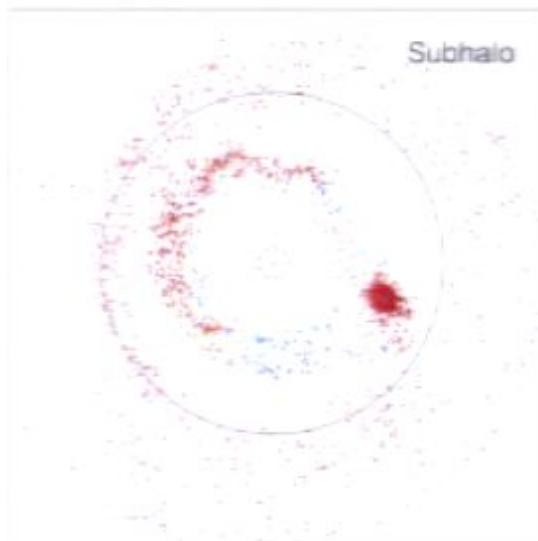
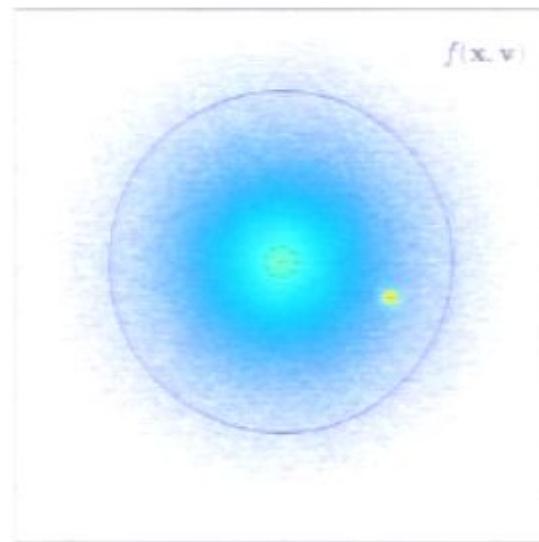
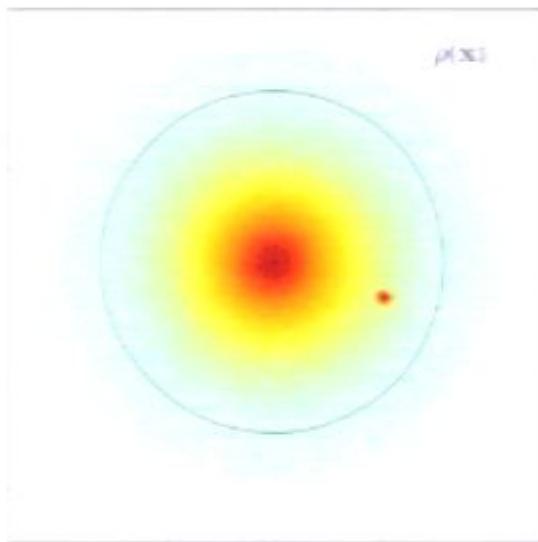


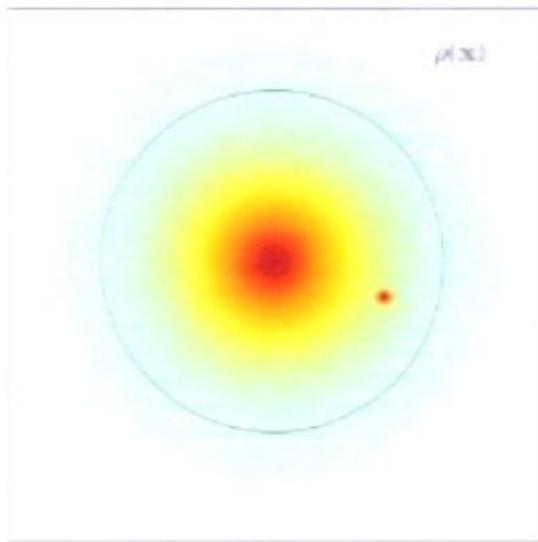
Groups

-0.012 < -0.007
0.005 < 0.010
-0.040 < -0.035

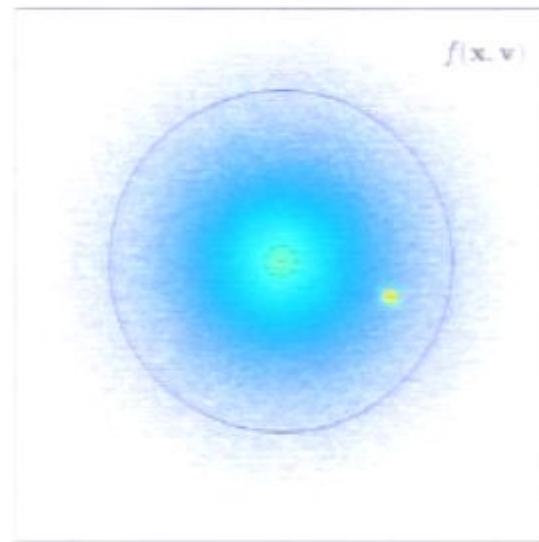
$P = 0.95$
 $R = 0.68$



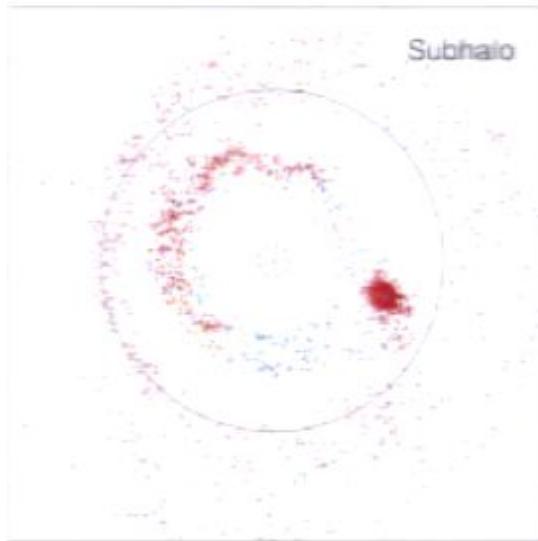




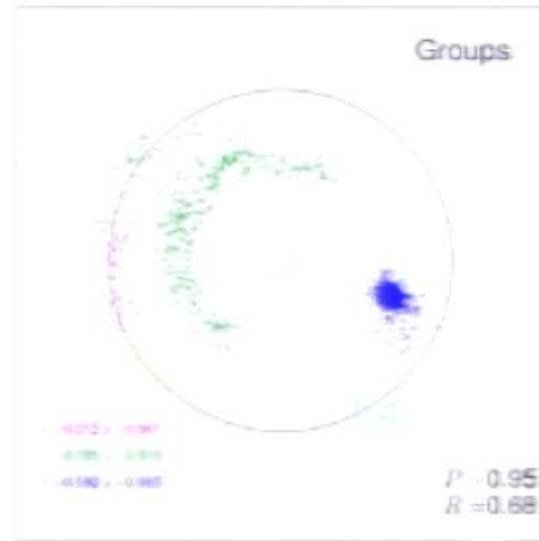
$\rho(\mathbf{x})$



$f(\mathbf{x}, \mathbf{v})$



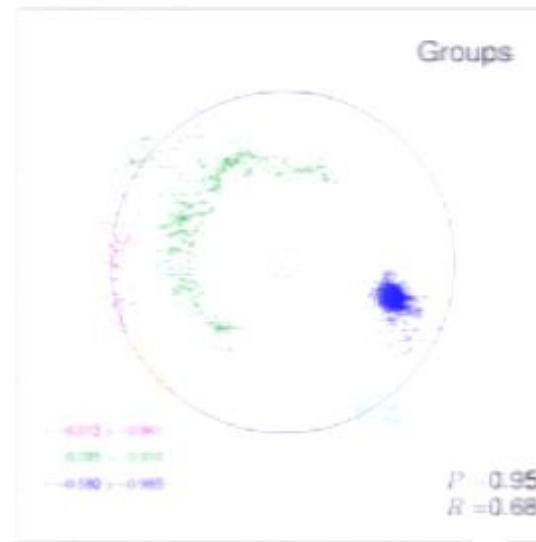
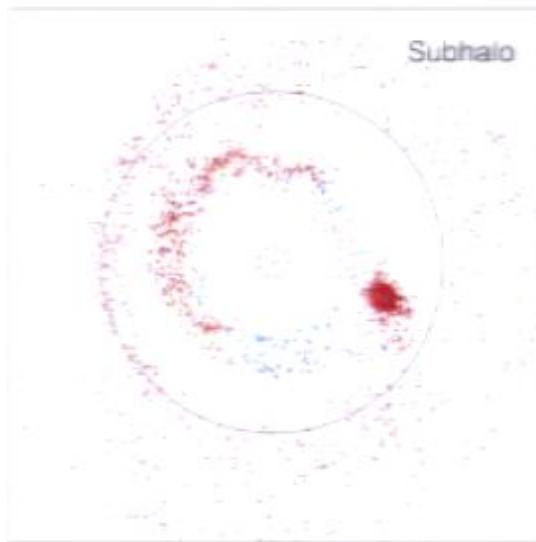
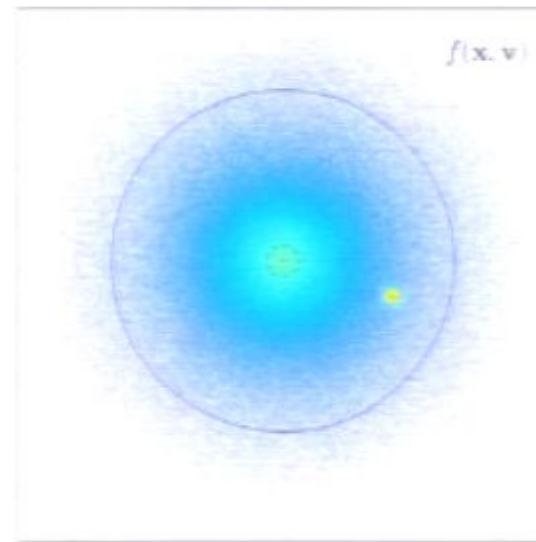
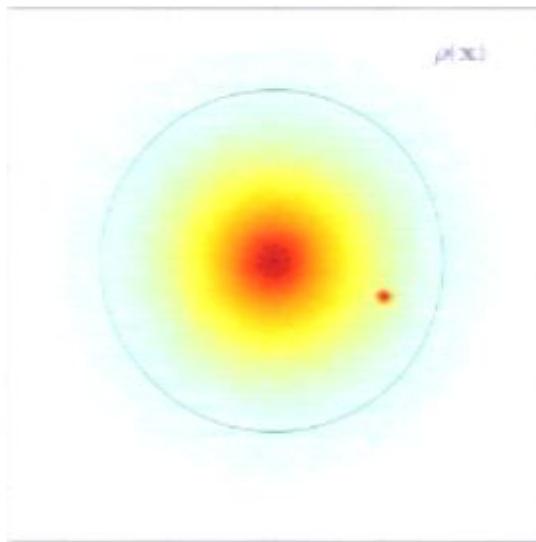
Subhalo

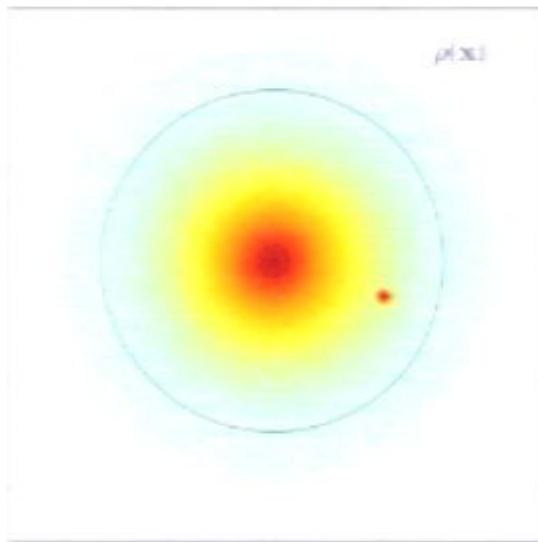


Groups

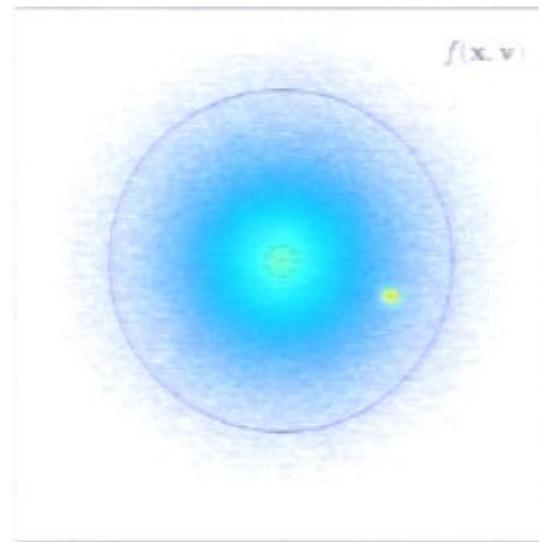
4012 > 0.997
0.999 > 0.999
0.999 > 0.999

$P = 0.95$
 $R = 0.68$

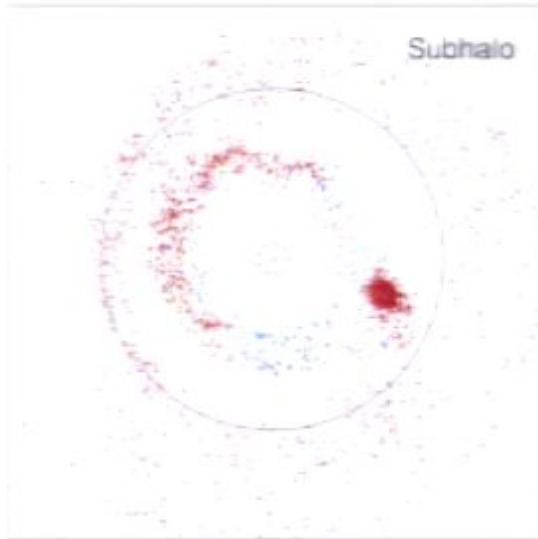




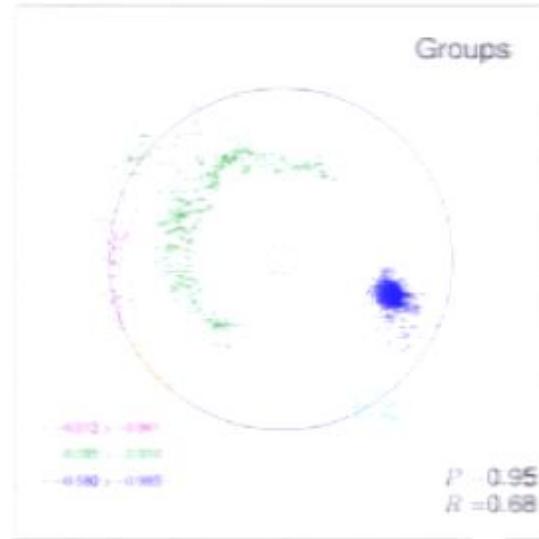
$\rho(\mathbf{x}, t)$



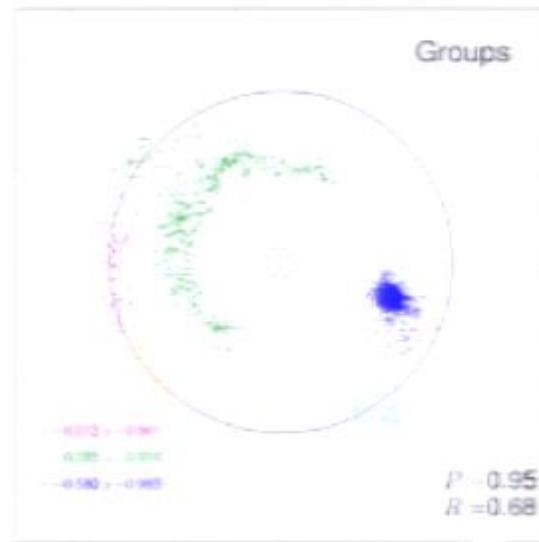
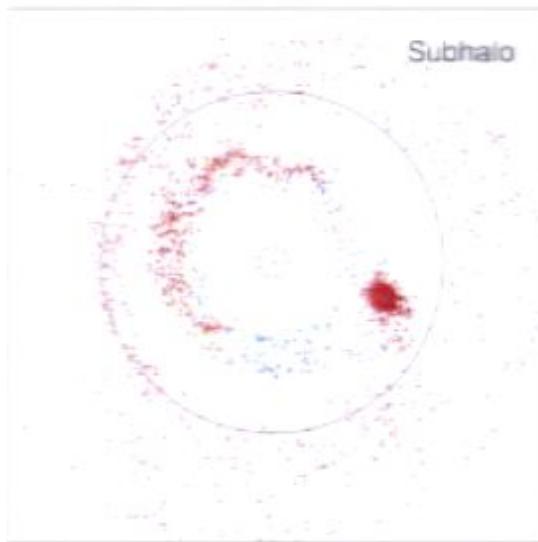
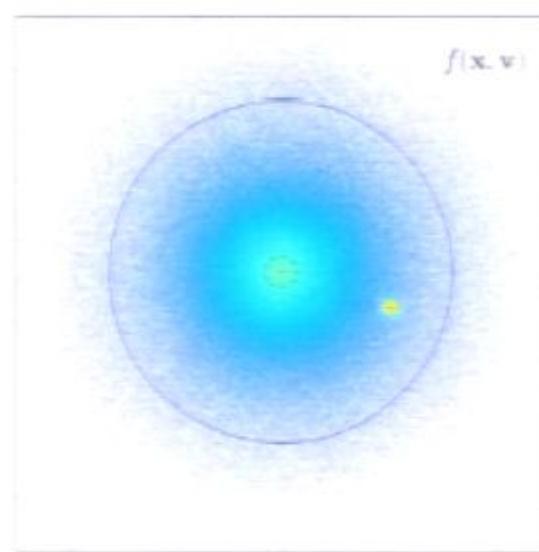
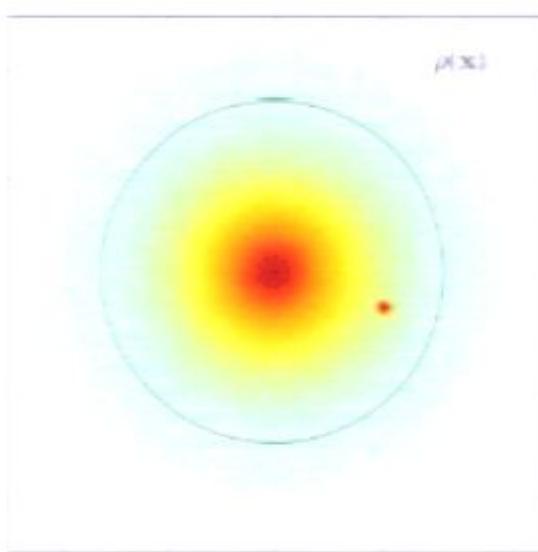
$f(\mathbf{x}, \mathbf{v})$

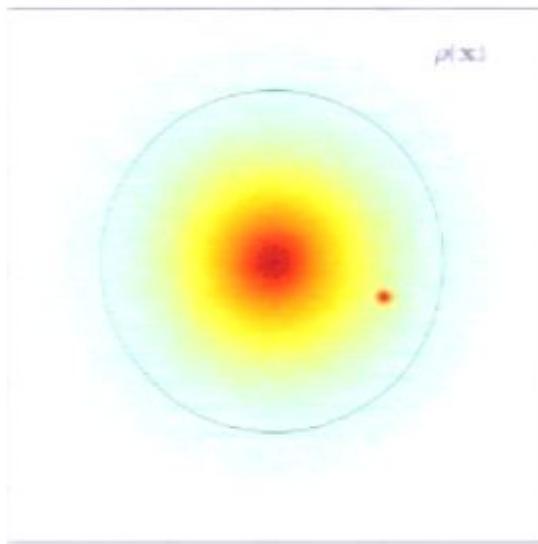


Subhalo

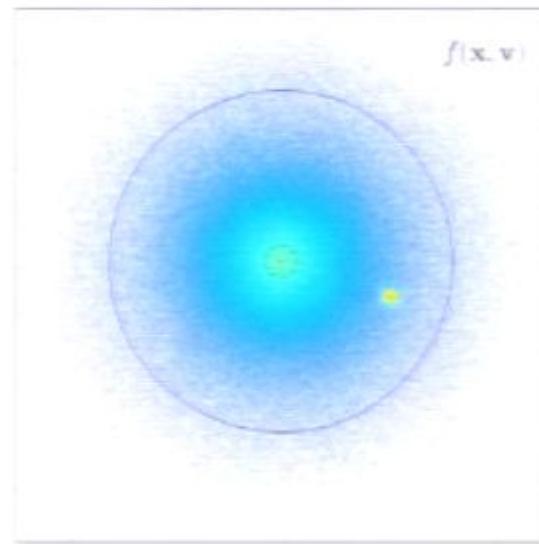


Groups

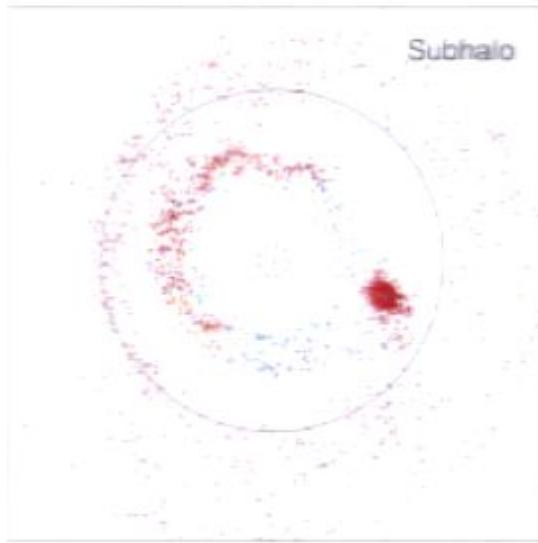




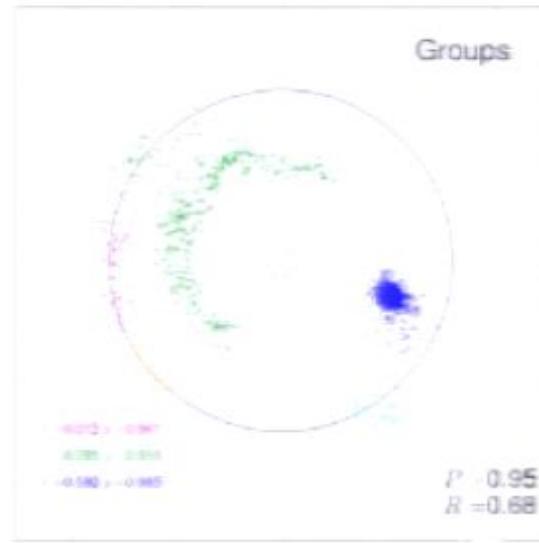
$\rho(\mathbf{x})$



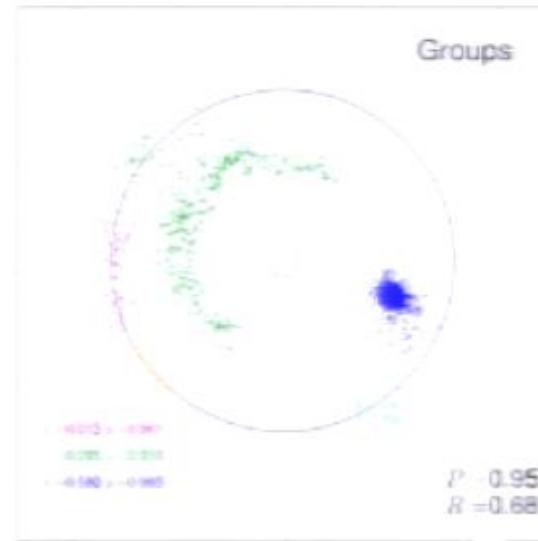
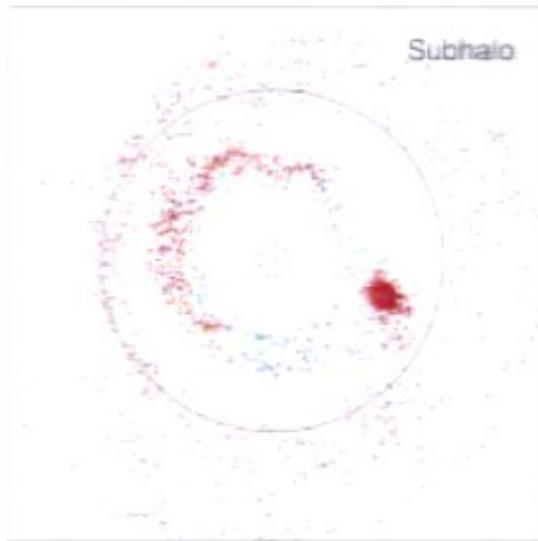
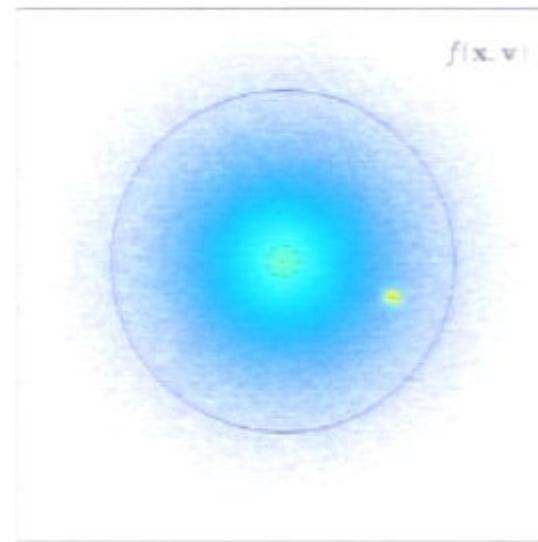
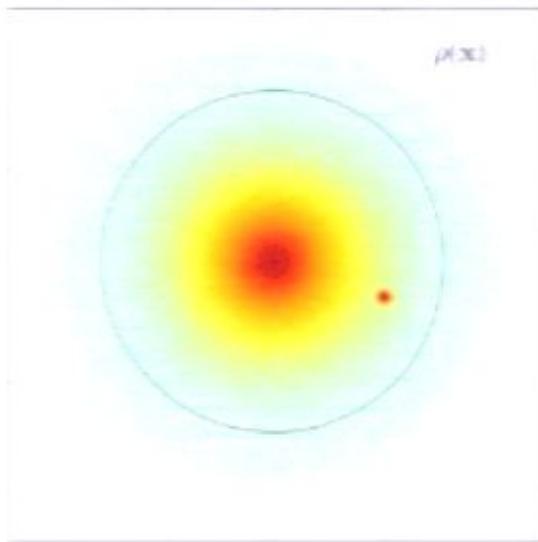
$f(\mathbf{x}, \mathbf{v})$

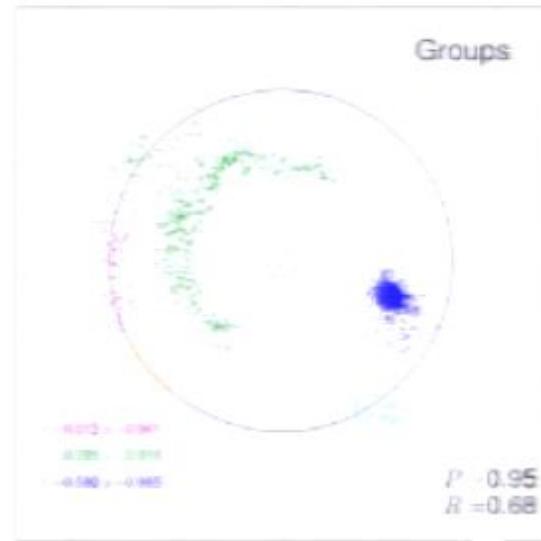
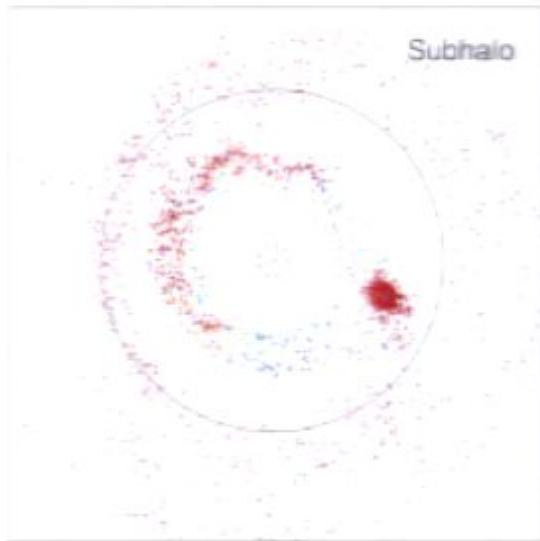
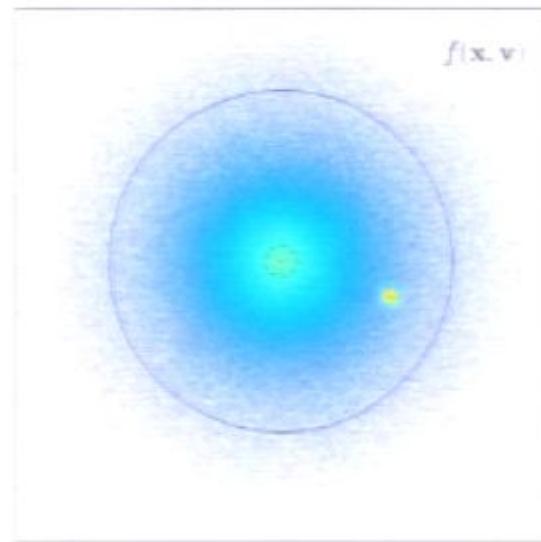
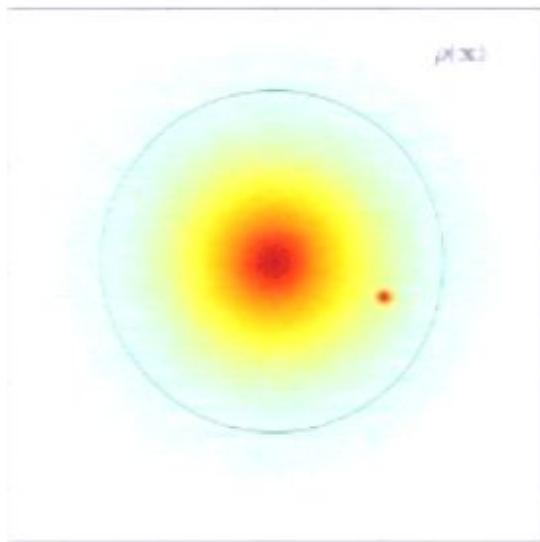


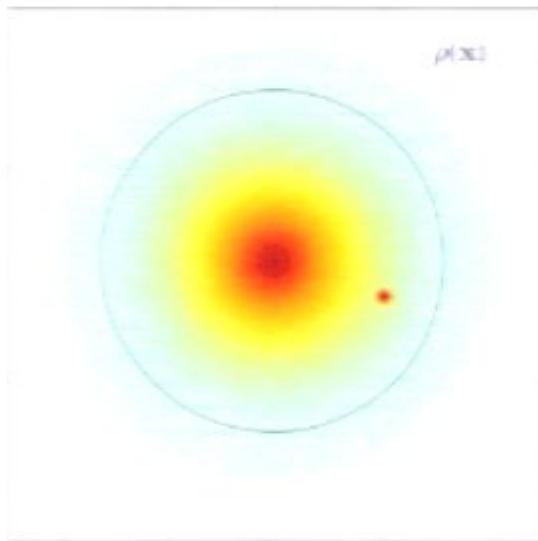
Subhalo



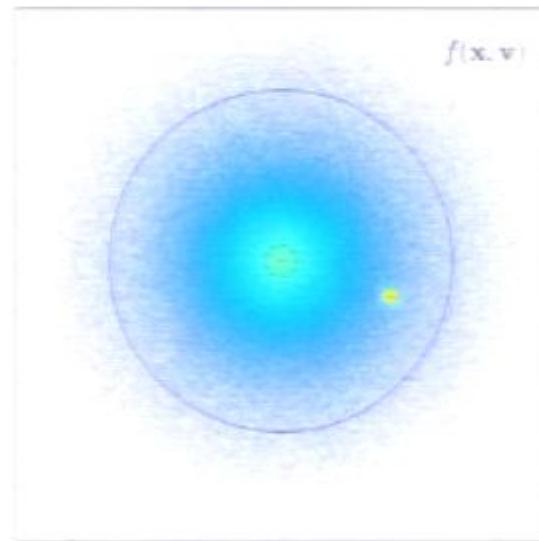
Groups



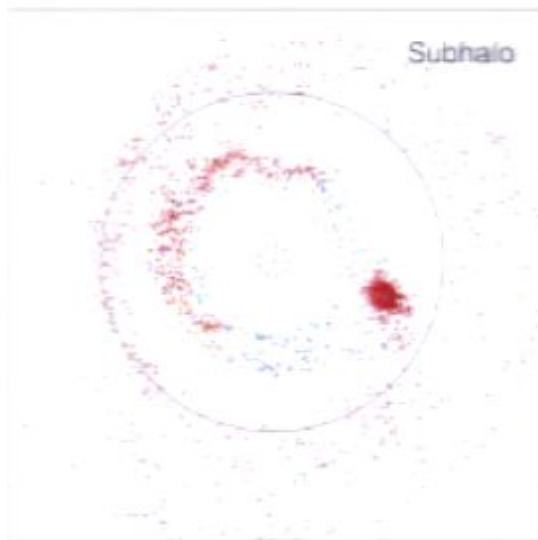




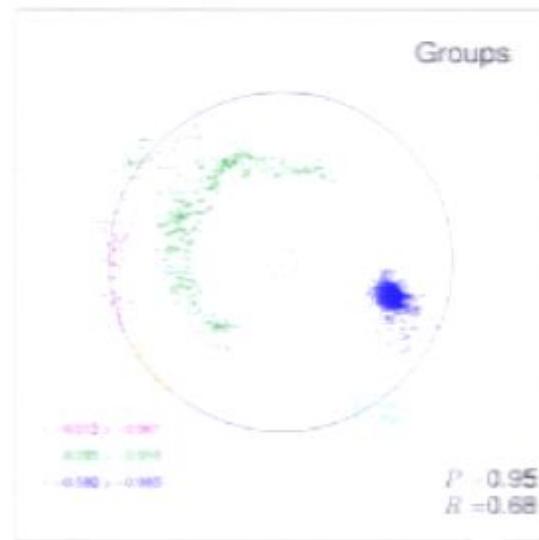
$\rho(26)$



$f(x, v)$

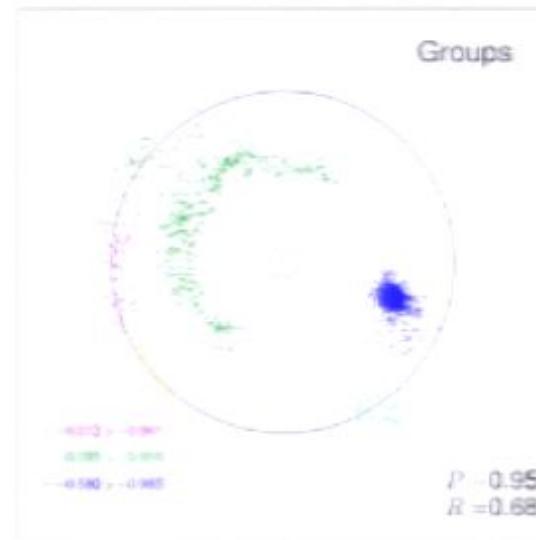
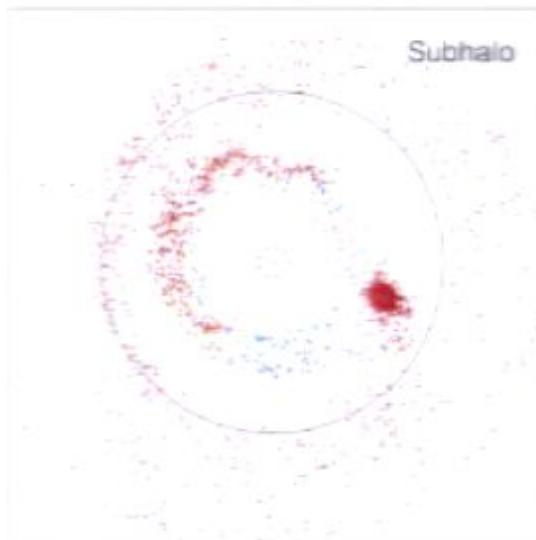
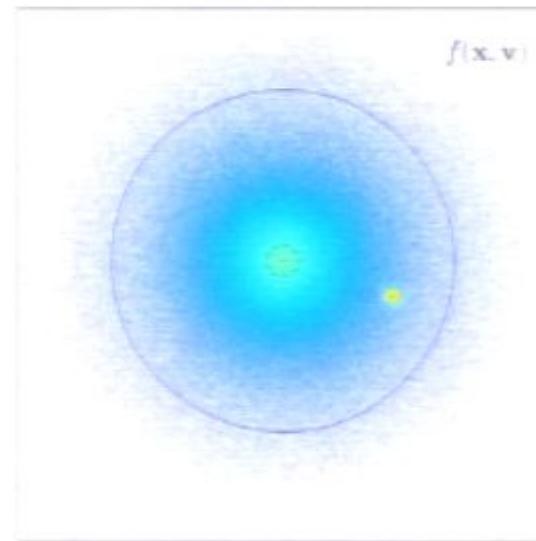
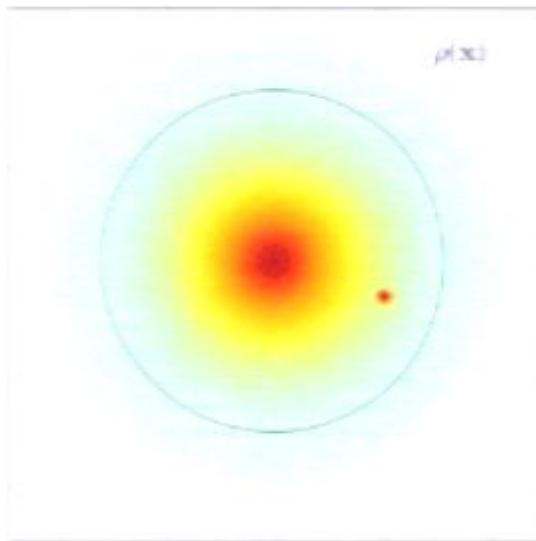


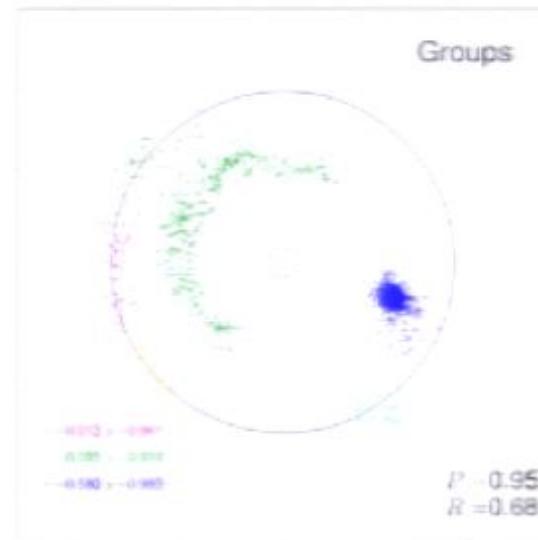
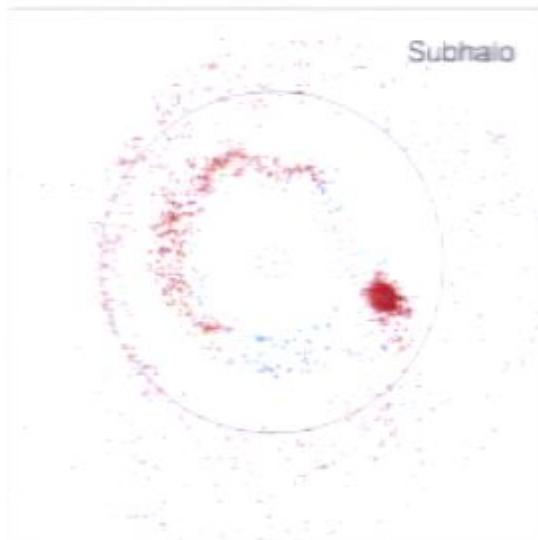
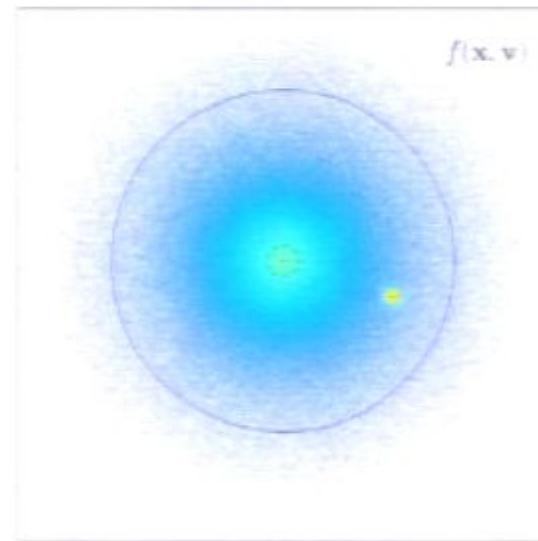
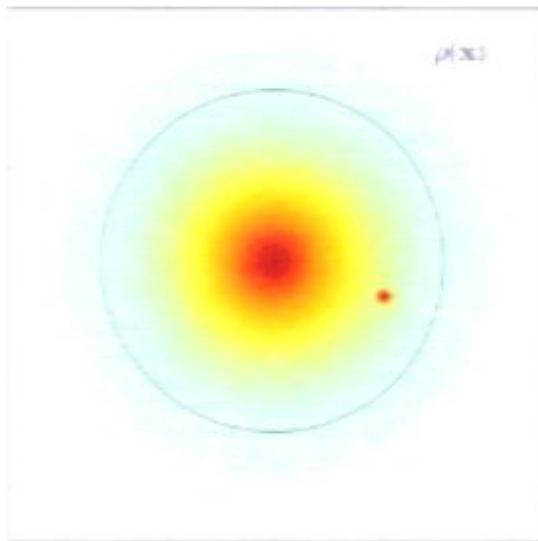
Subhalo

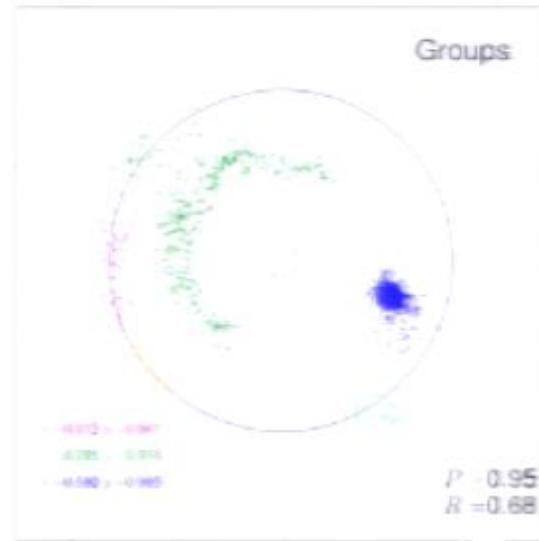
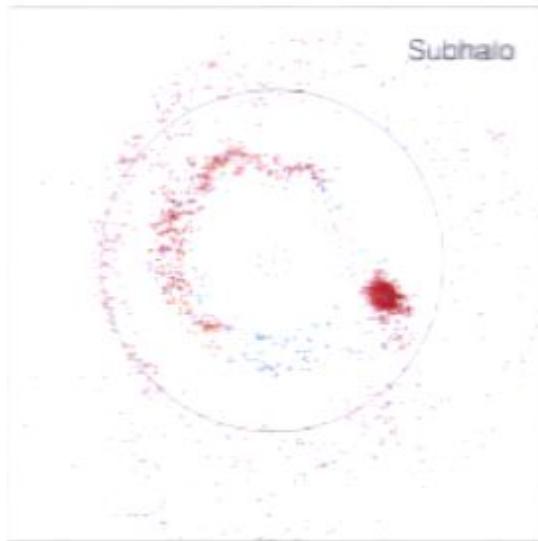
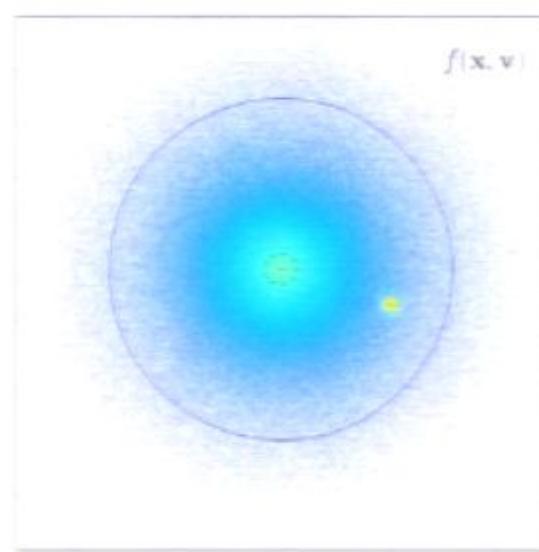
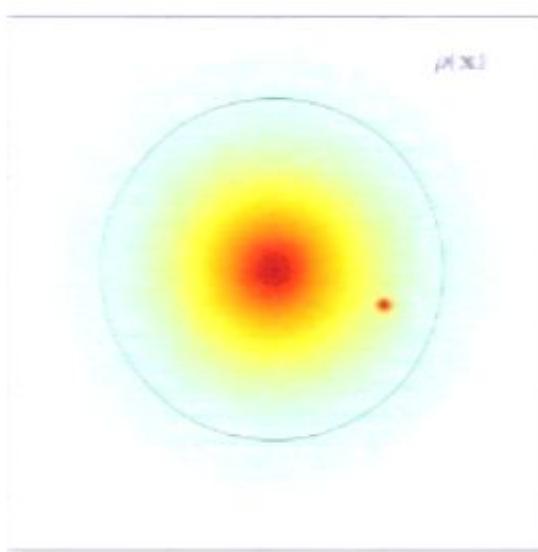


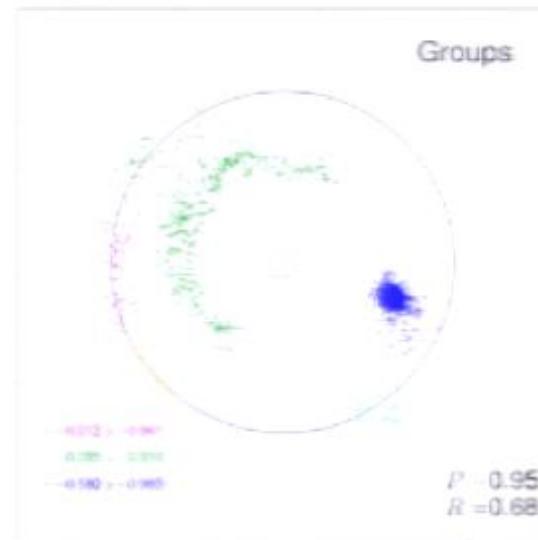
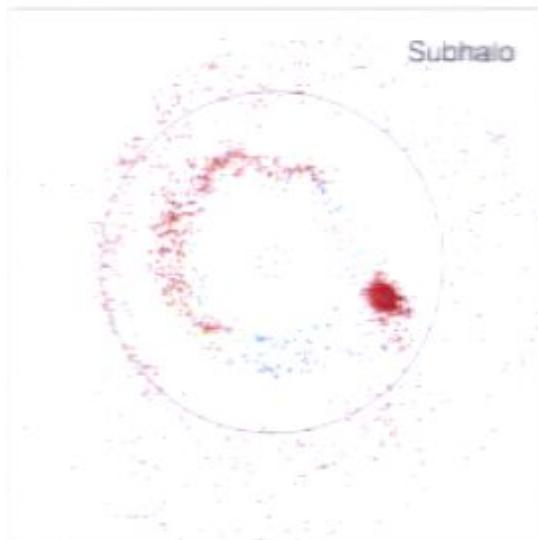
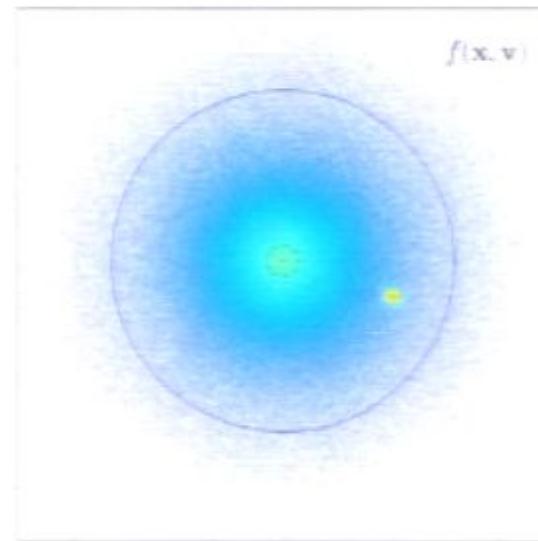
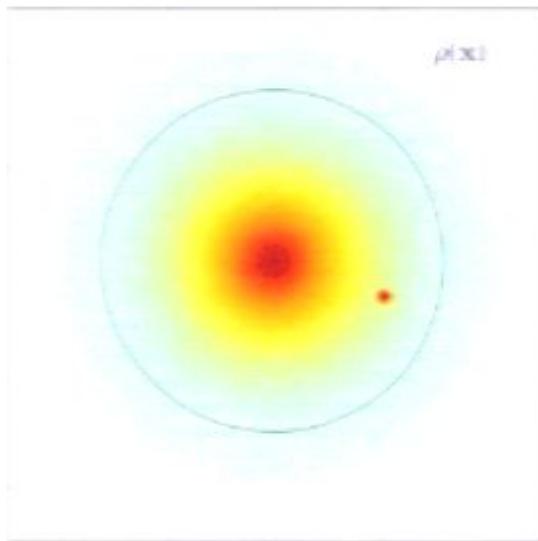
Groups

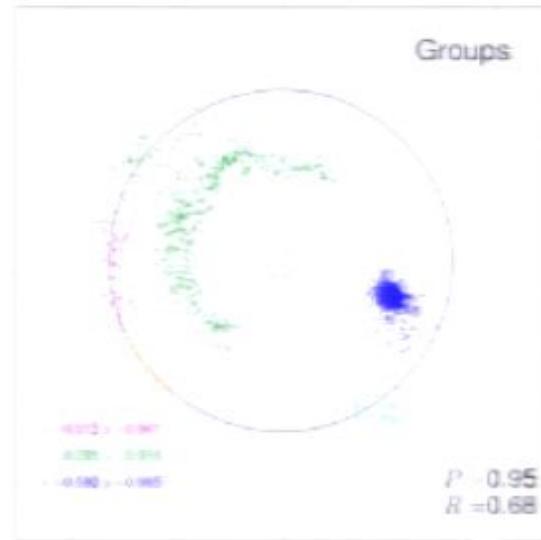
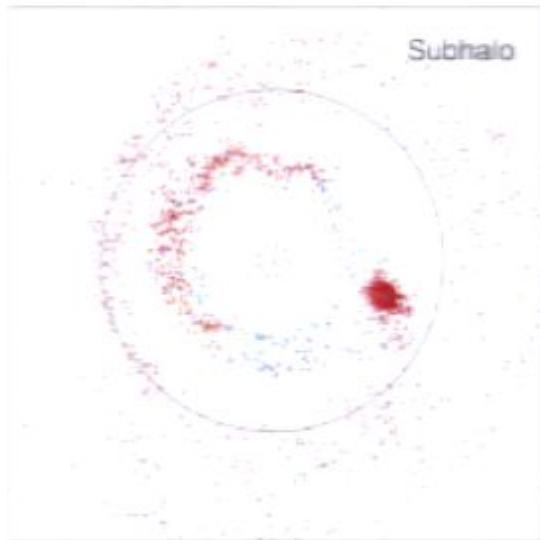
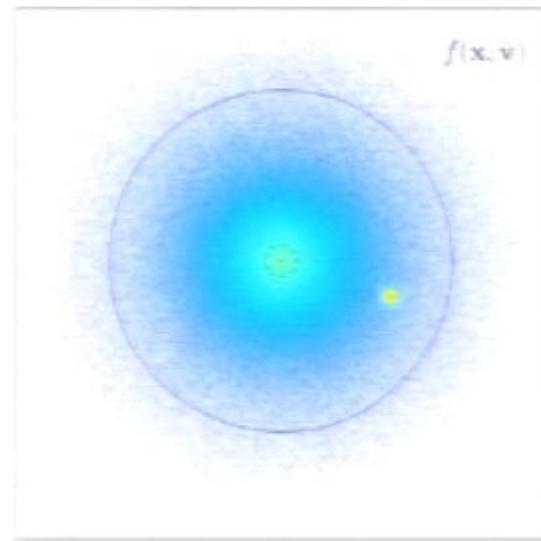
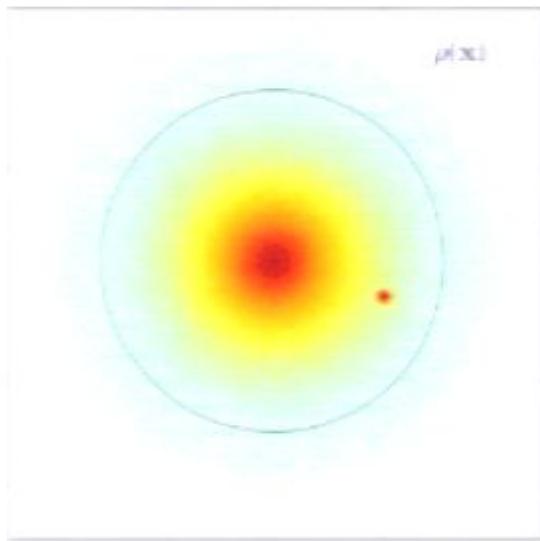
$P = 0.95$
 $R = 0.68$

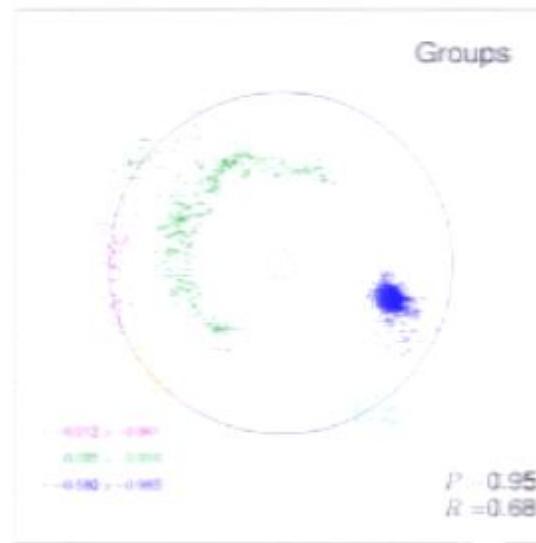
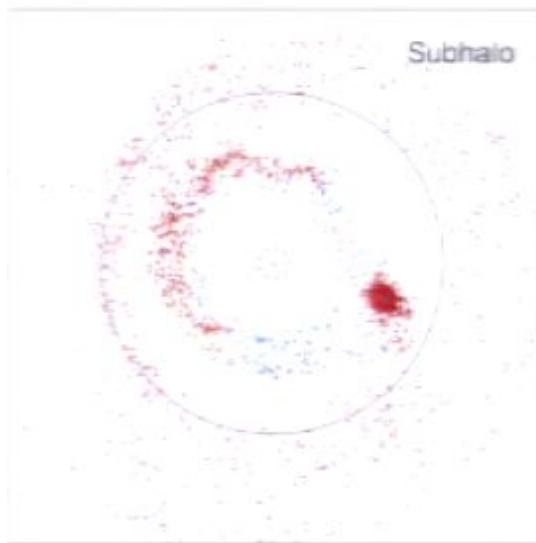
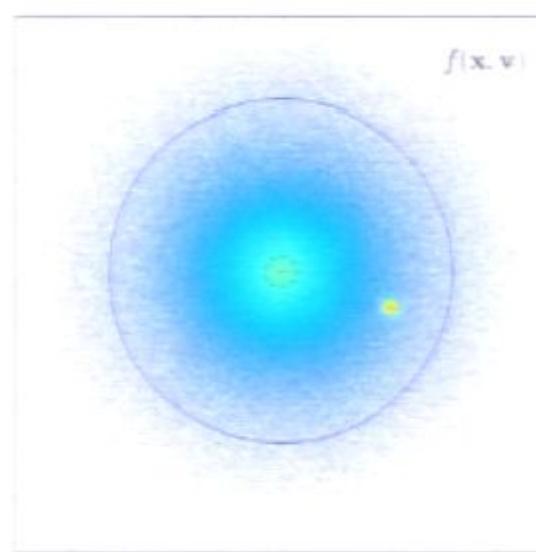
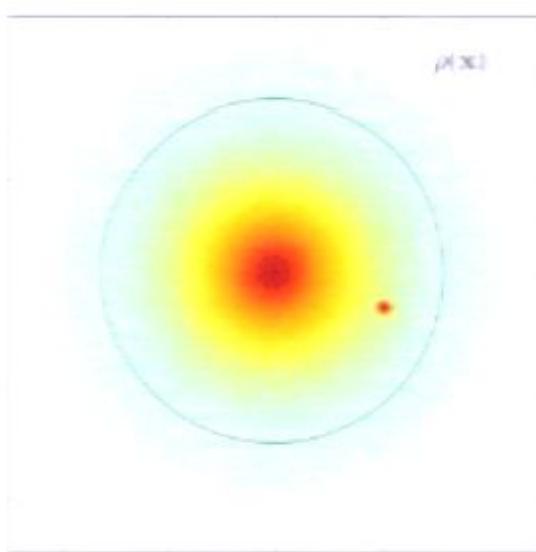


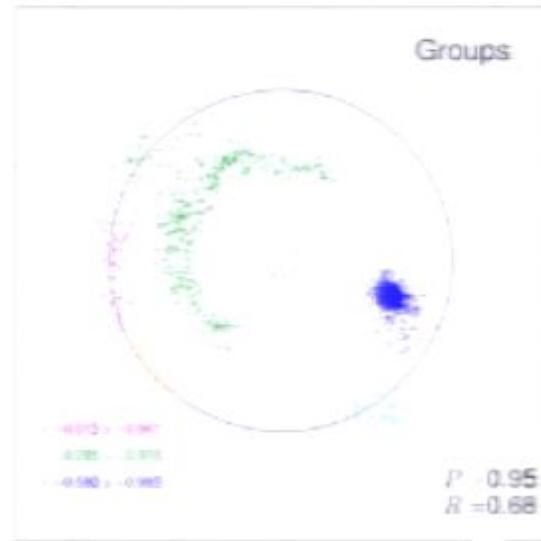
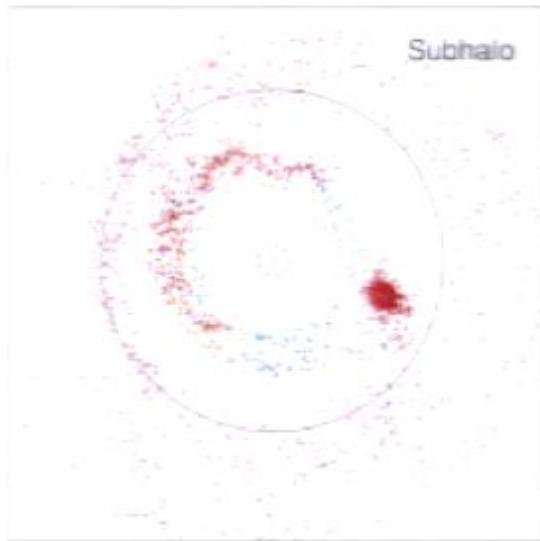
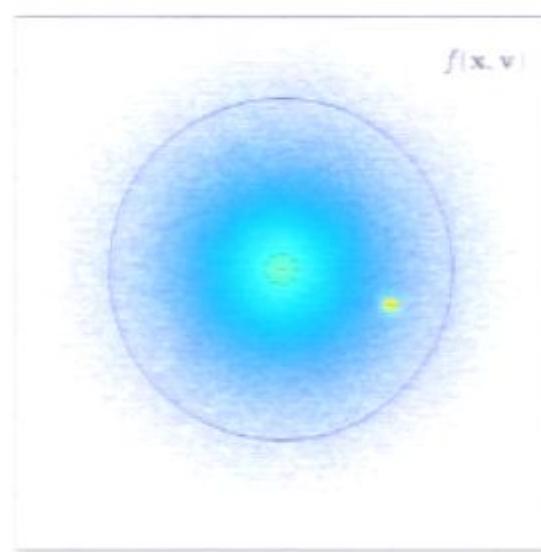
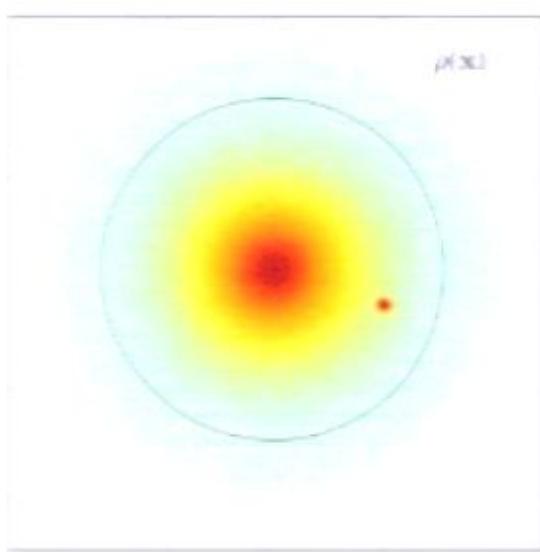


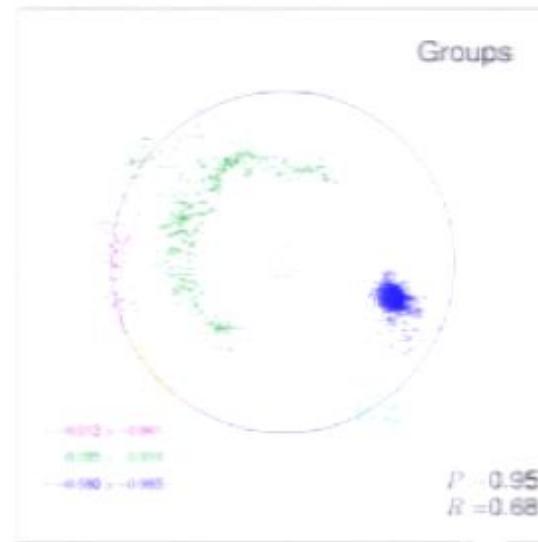
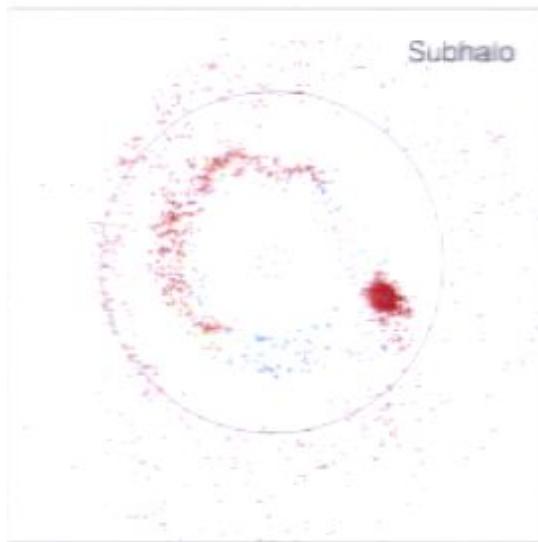
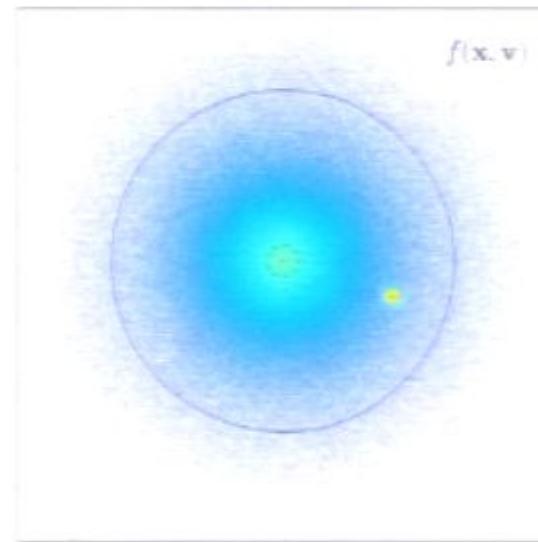
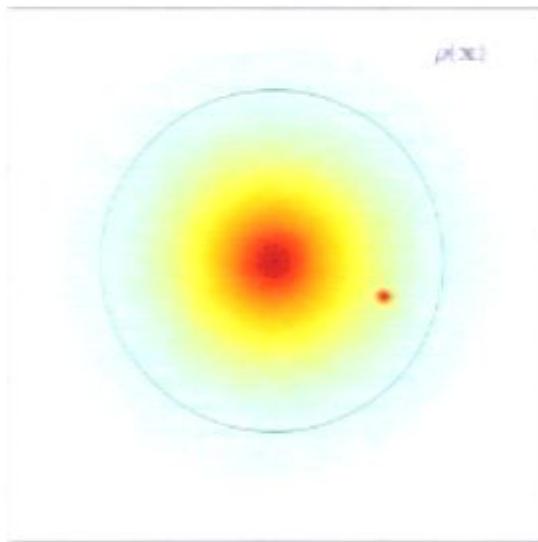


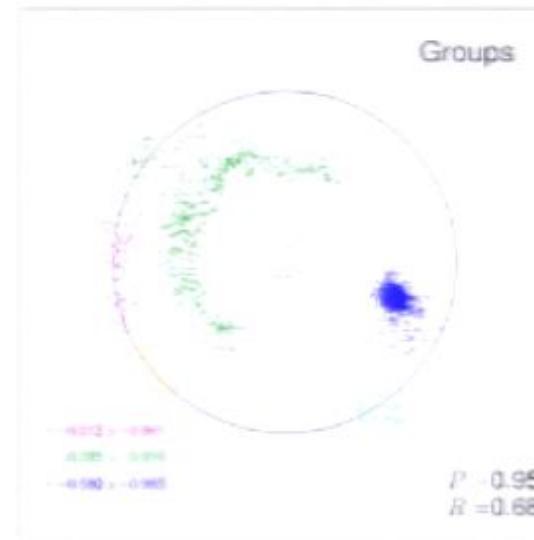
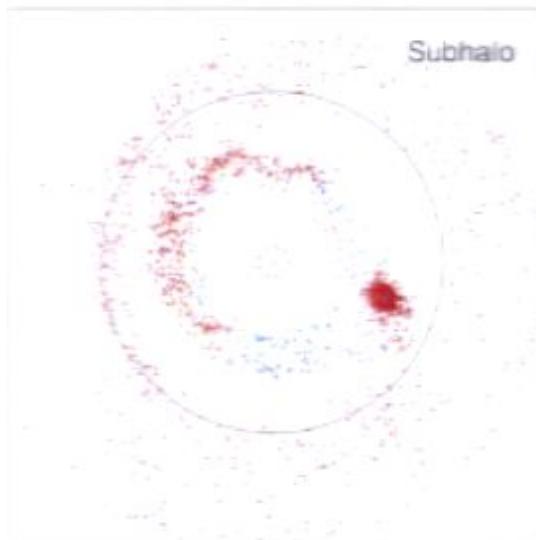
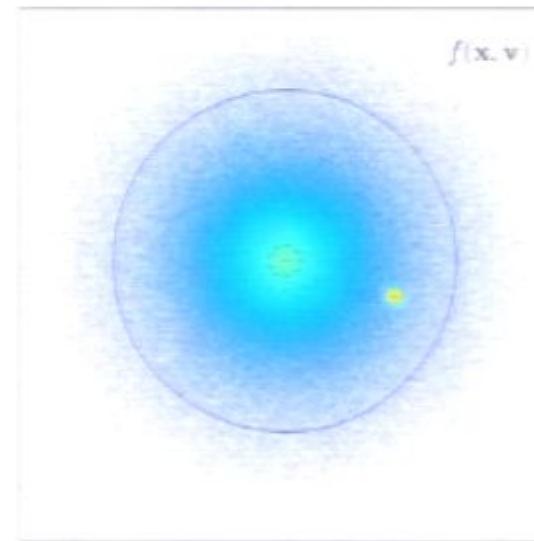
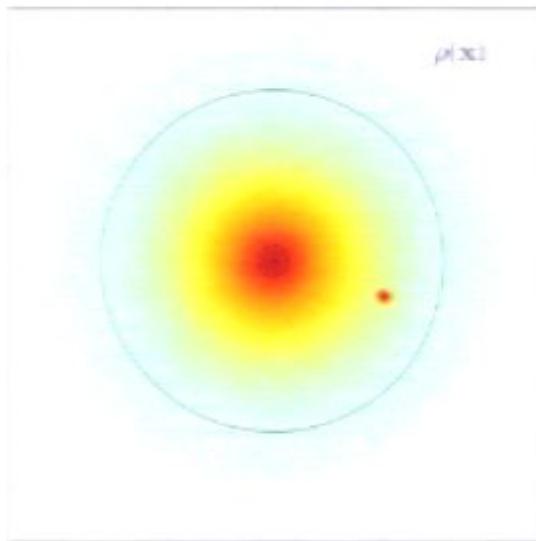


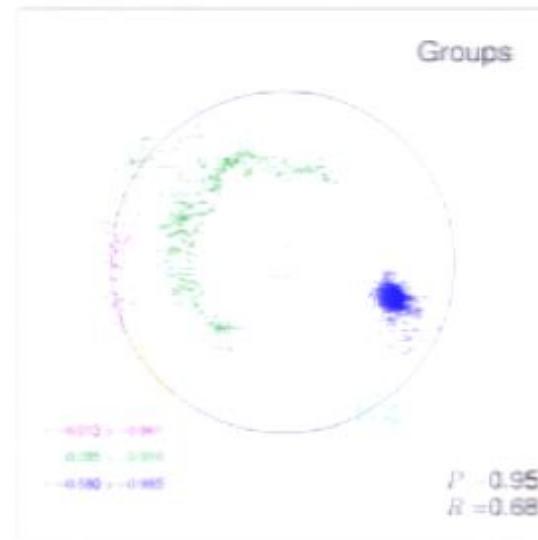
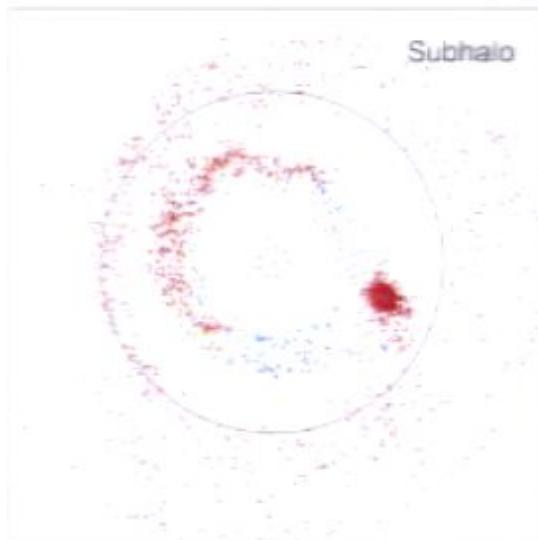
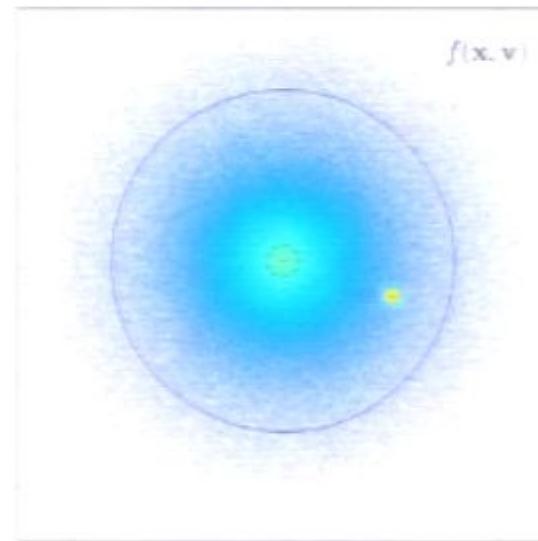
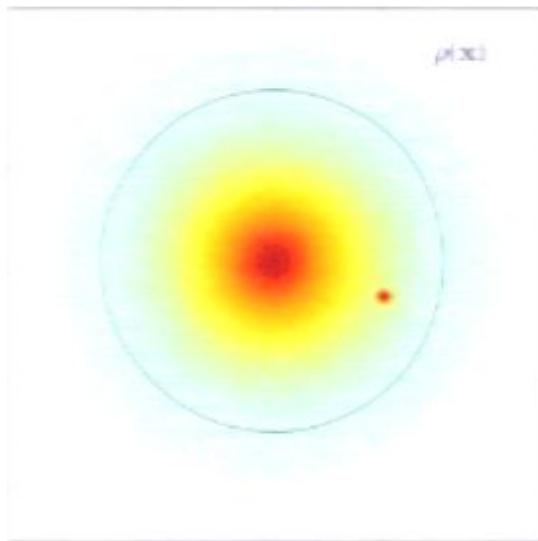


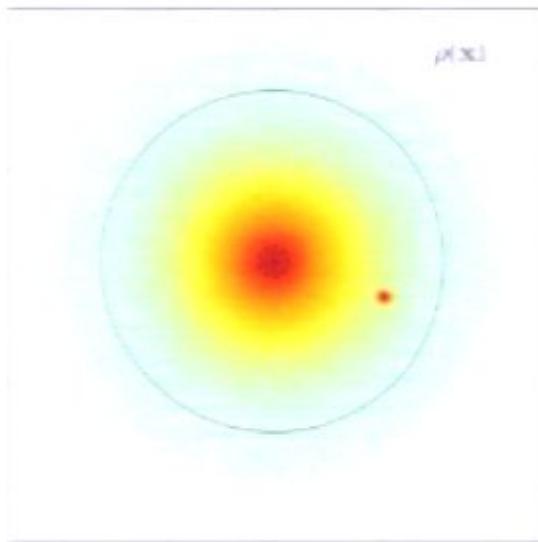




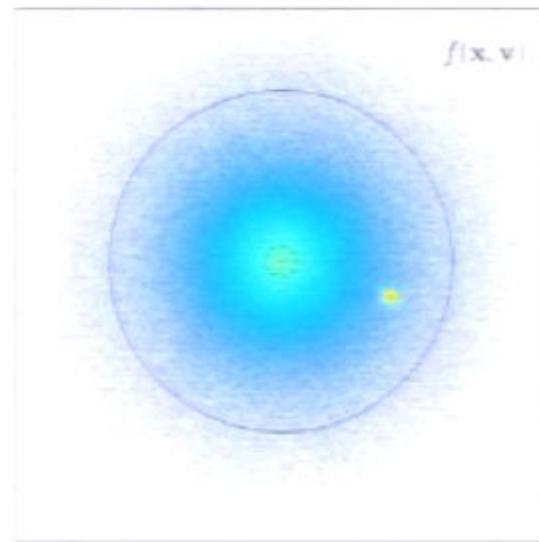




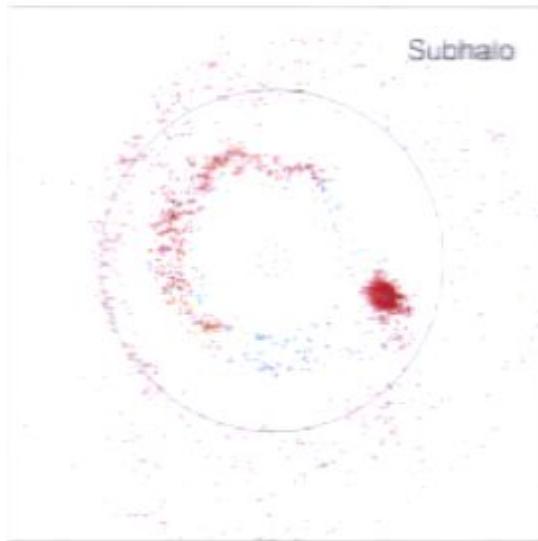




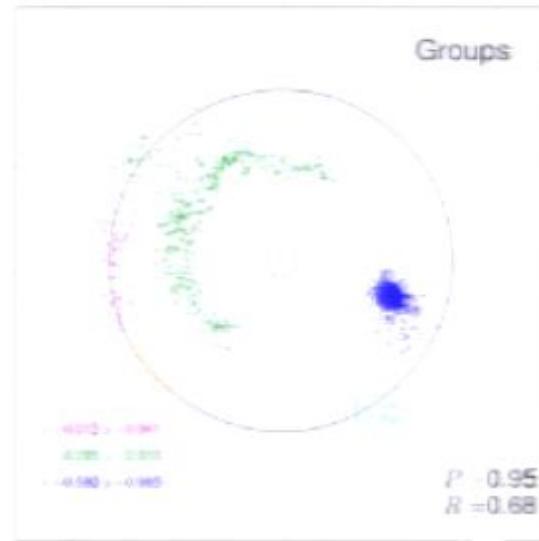
$\rho(\mathbf{x})$

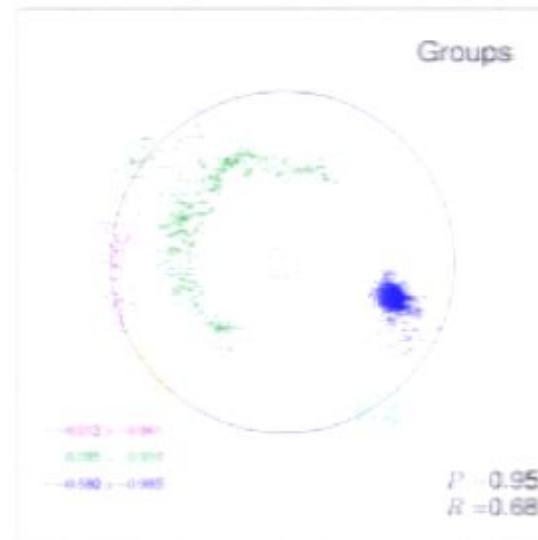
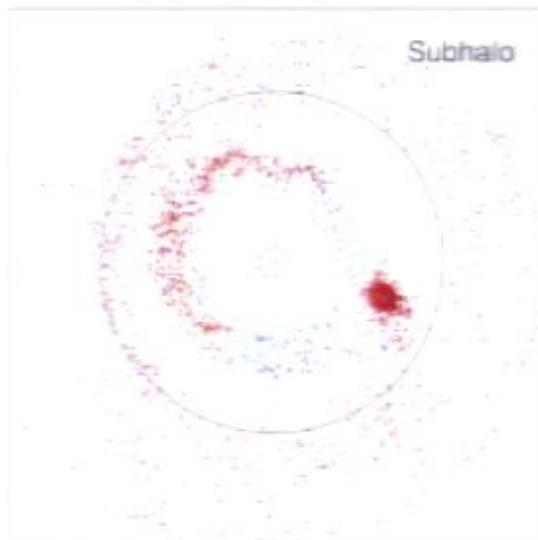
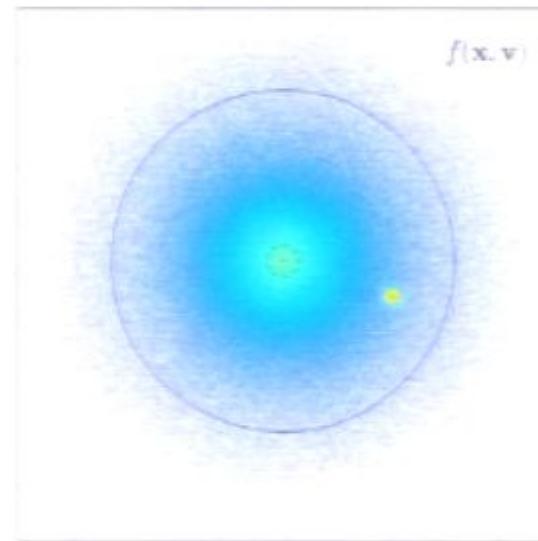
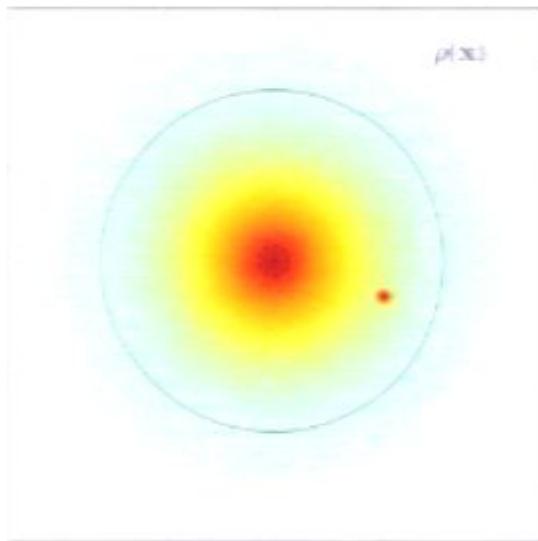


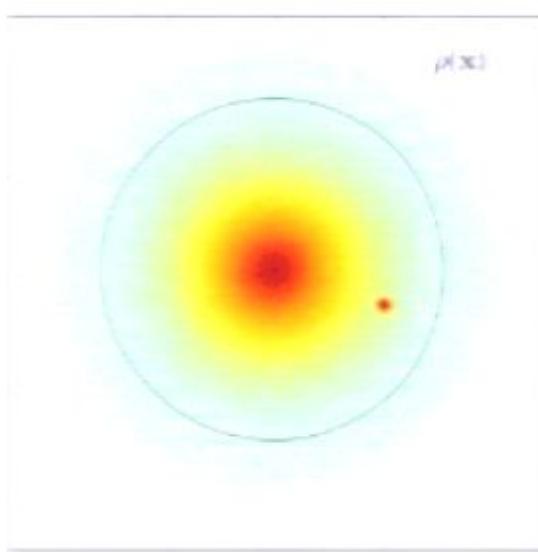
$f(\mathbf{x}, \mathbf{v})$



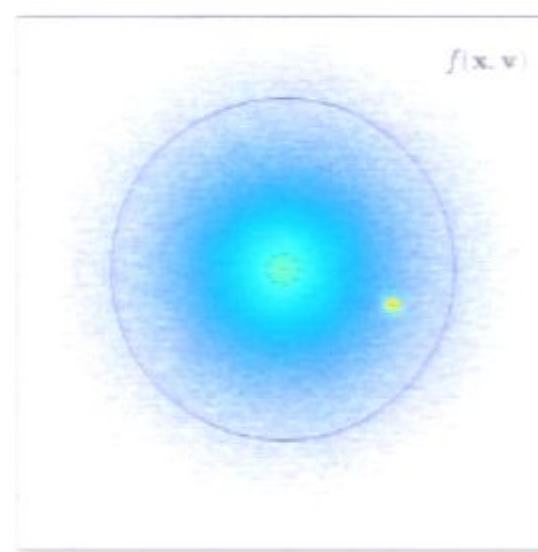
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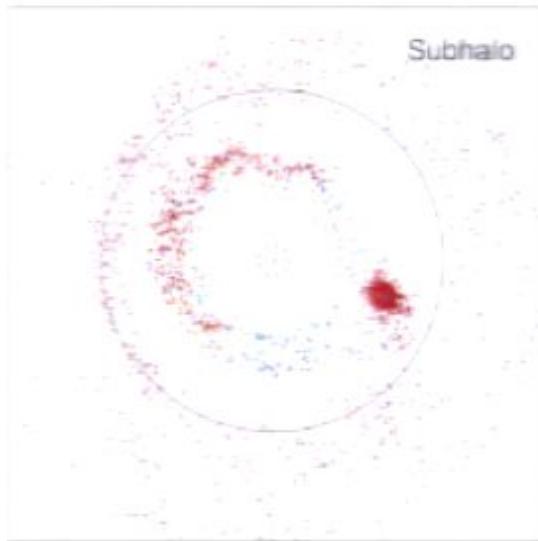




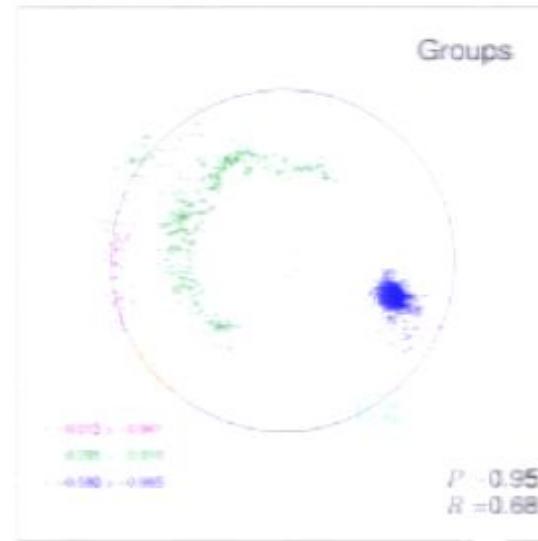
$\rho(\mathbf{x})$



$f(\mathbf{x}, \mathbf{v})$

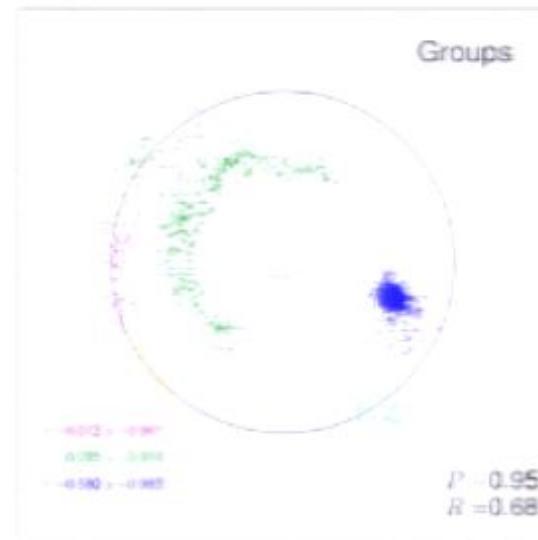
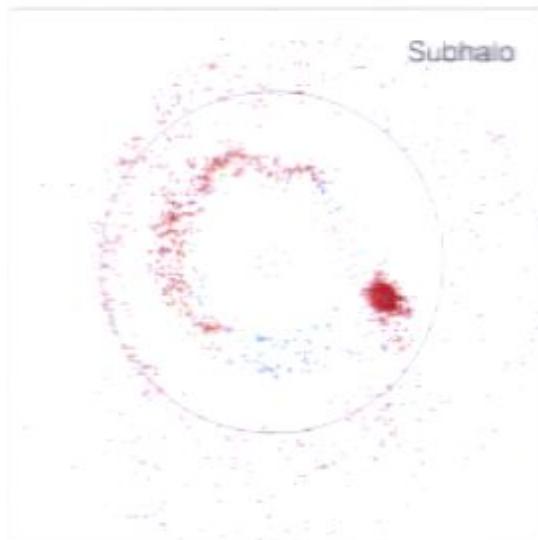
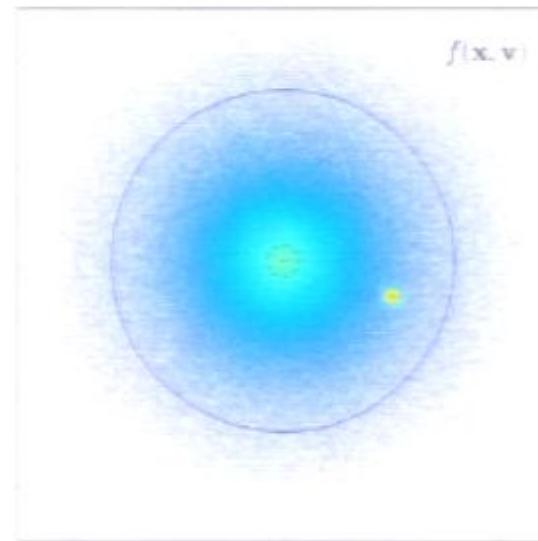
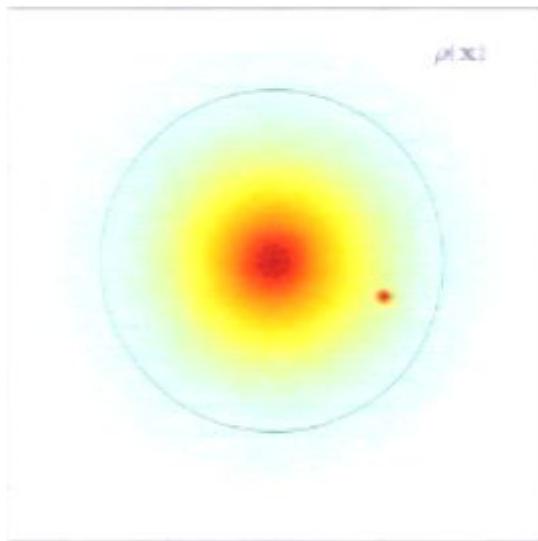


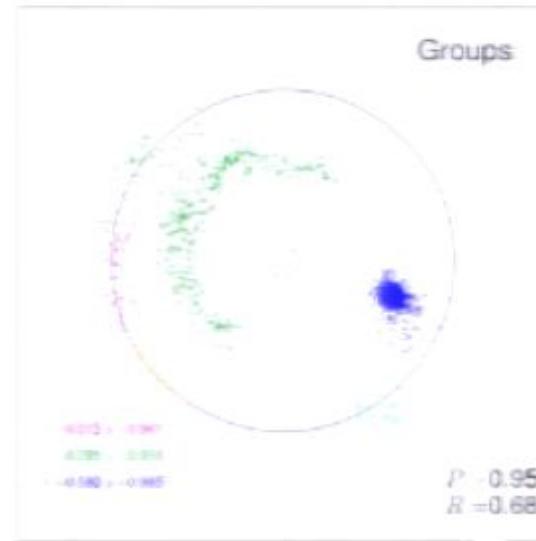
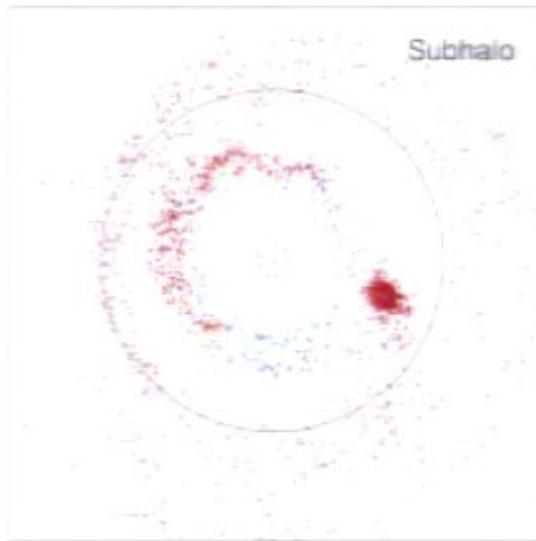
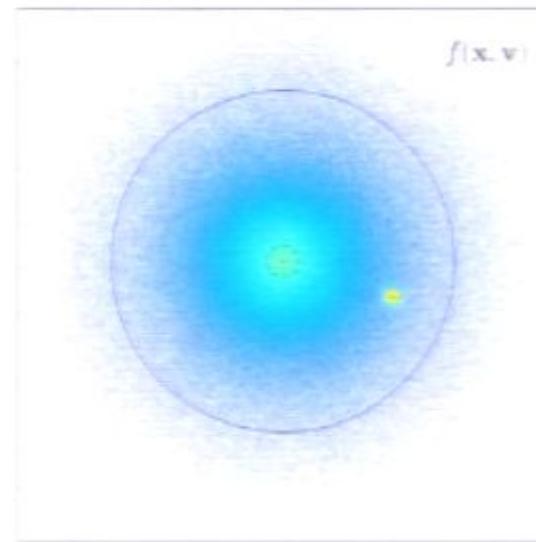
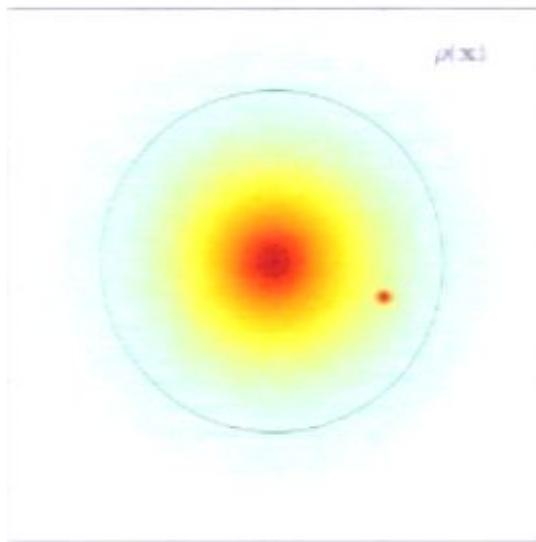
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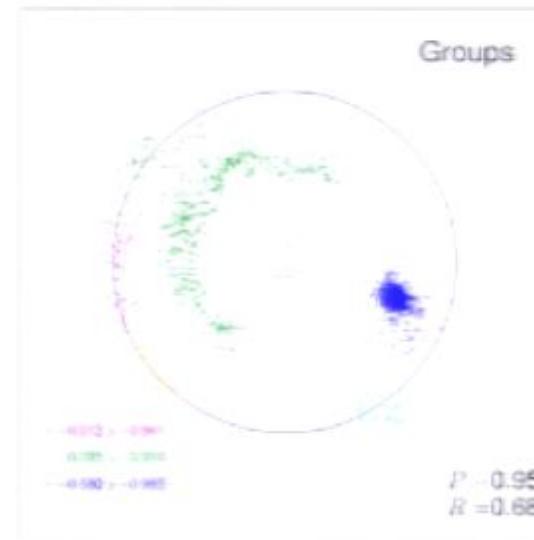
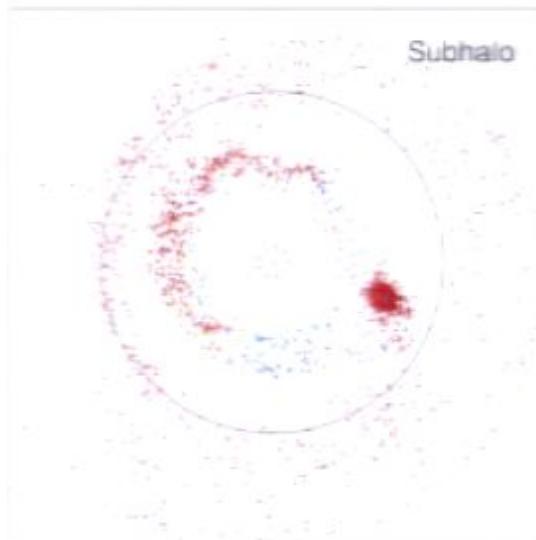
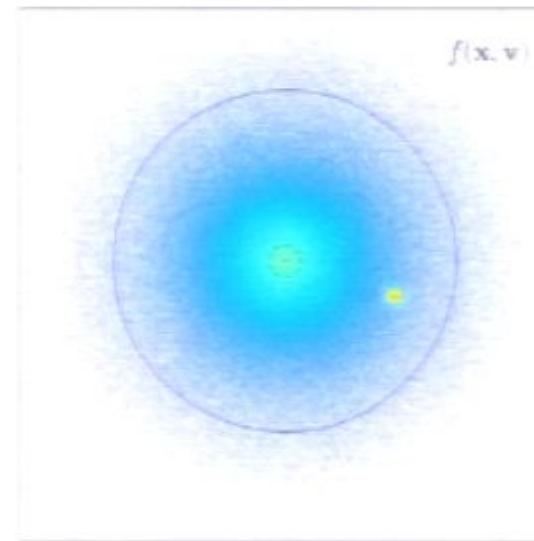
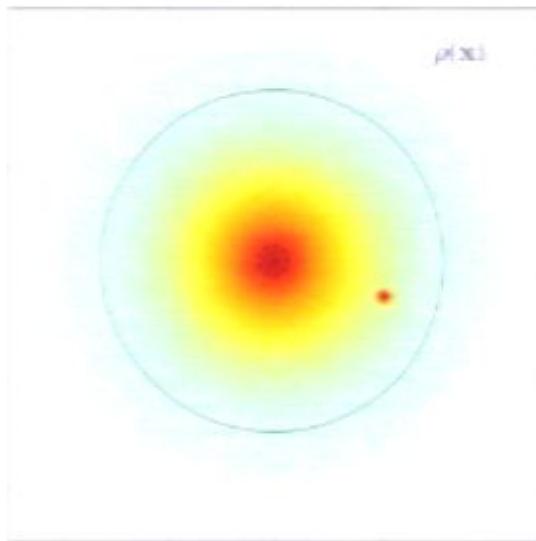


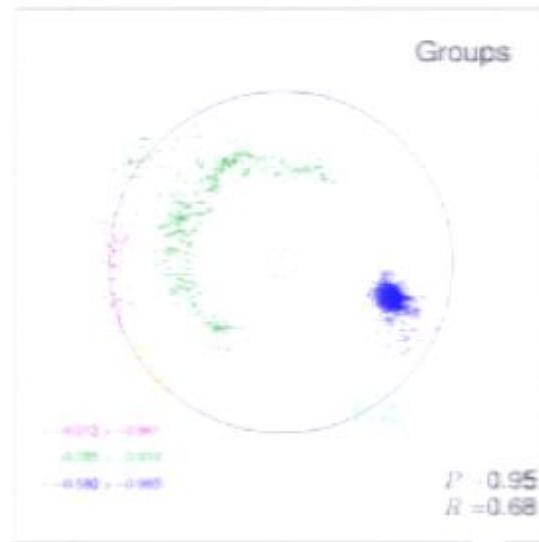
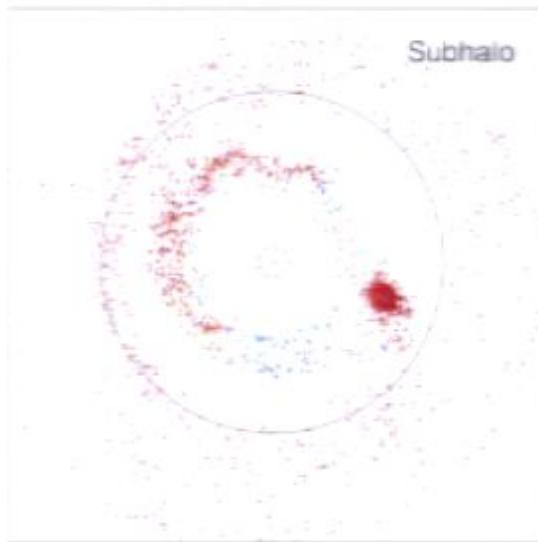
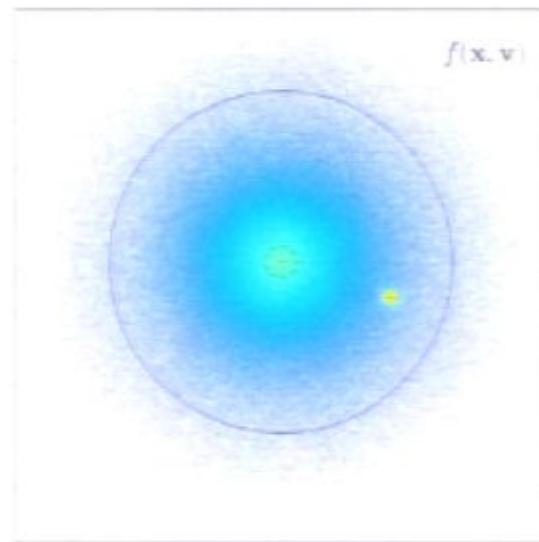
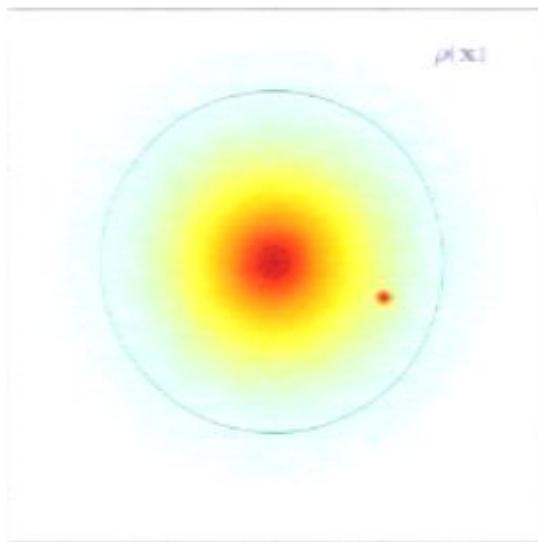
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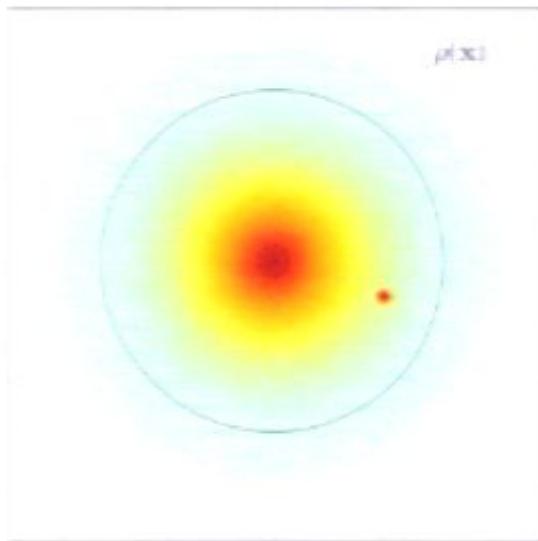
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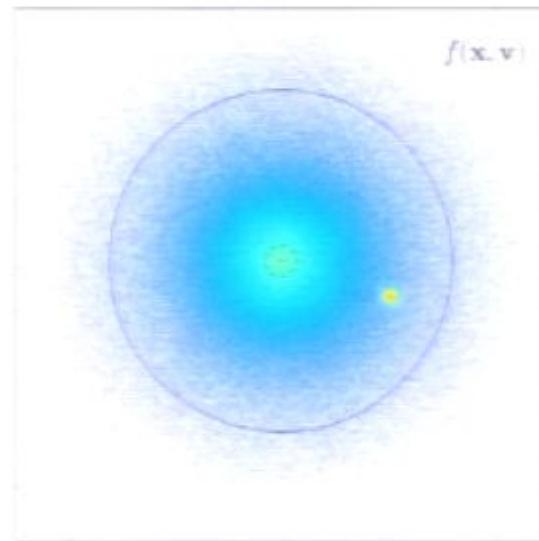




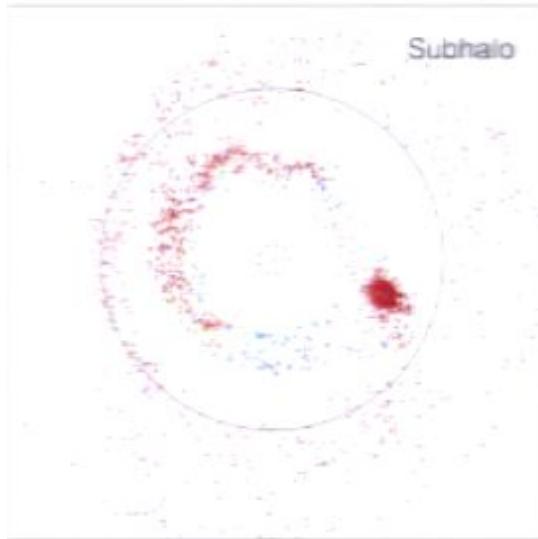




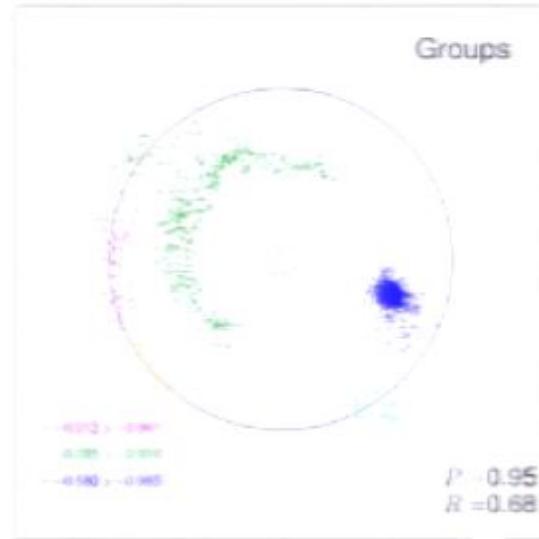
$\rho(\mathbf{x})$



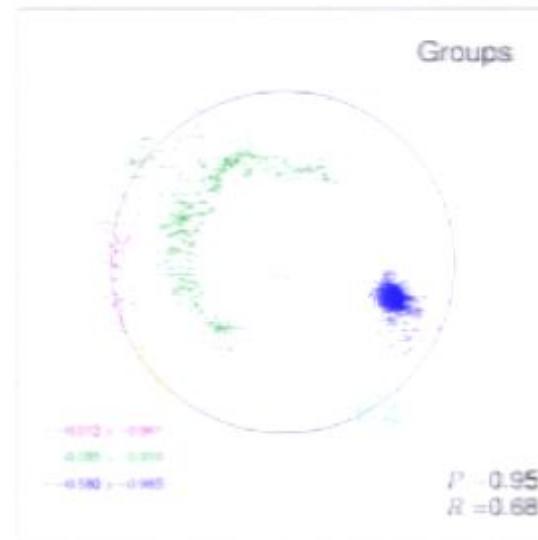
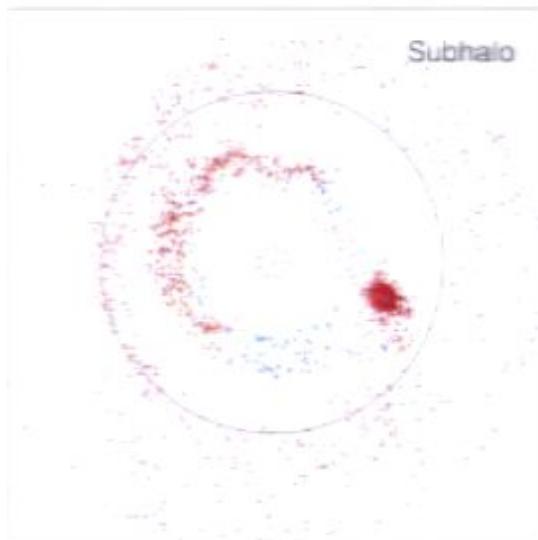
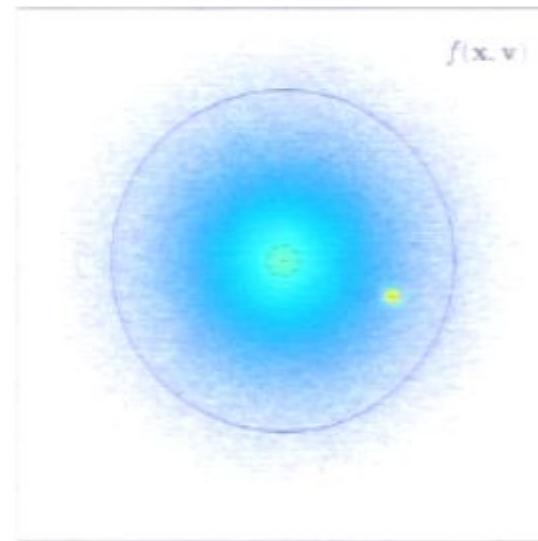
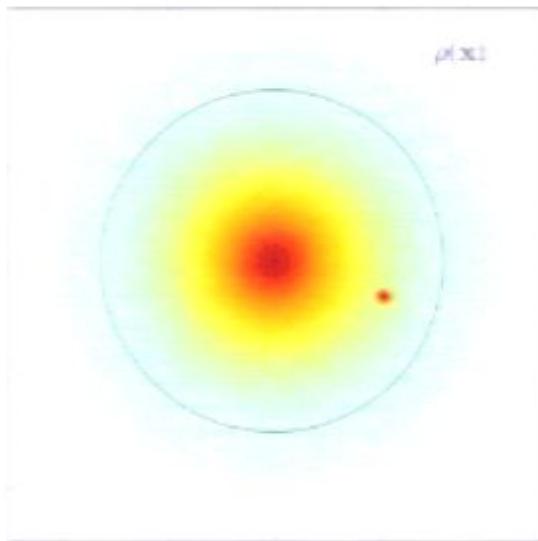
$f(\mathbf{x}, \mathbf{v})$

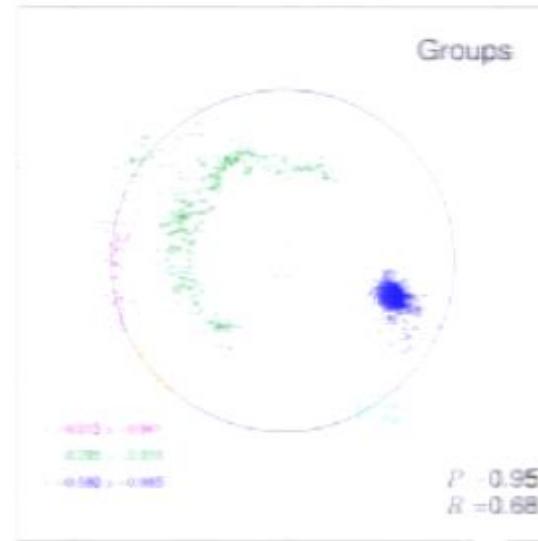
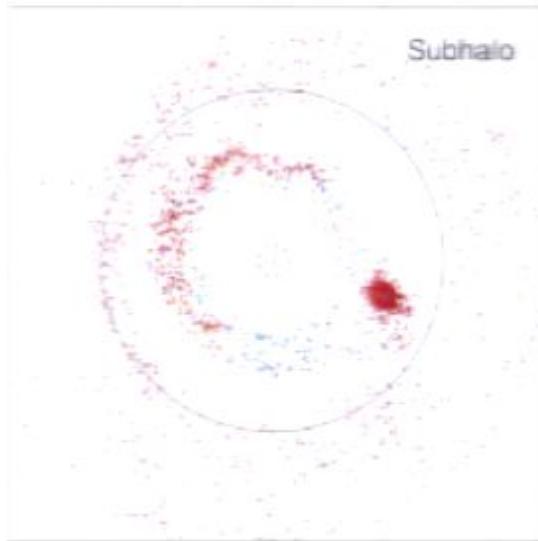
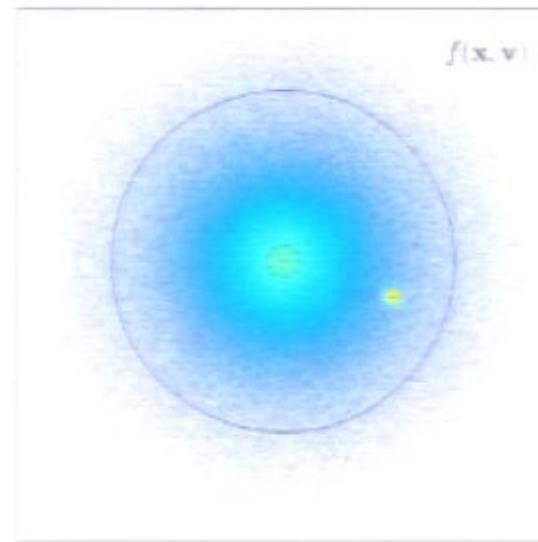
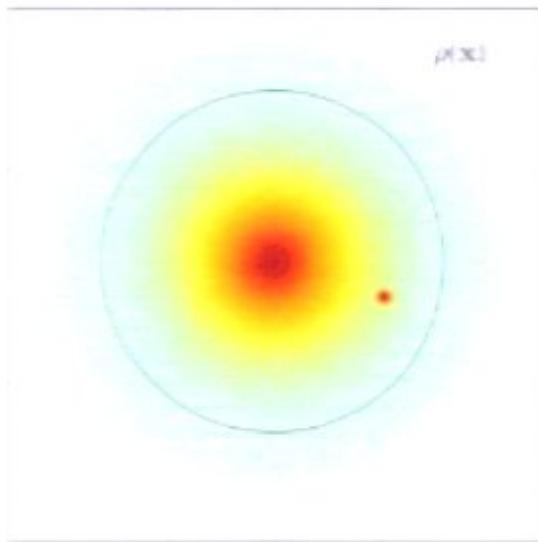


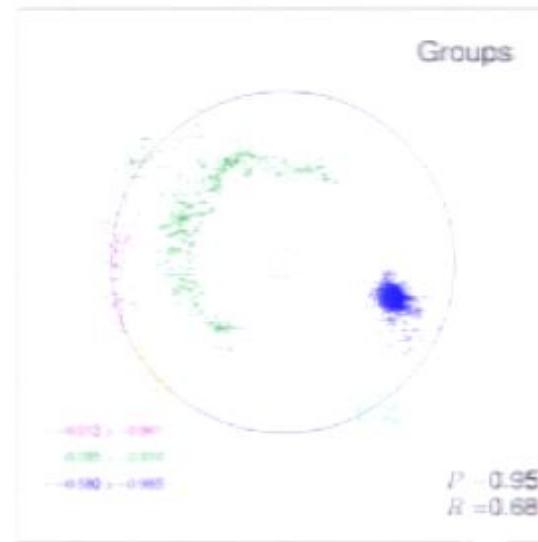
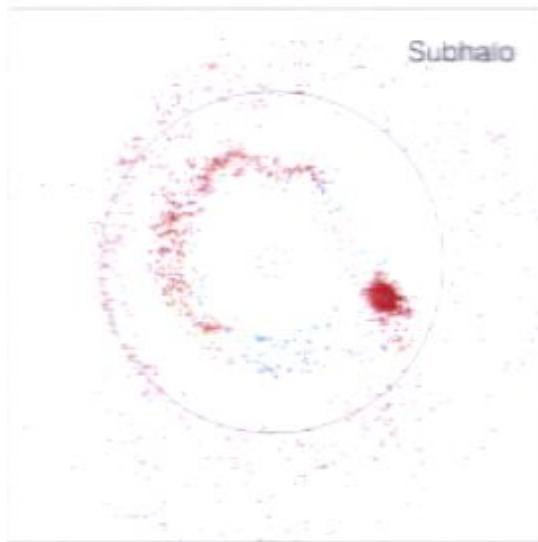
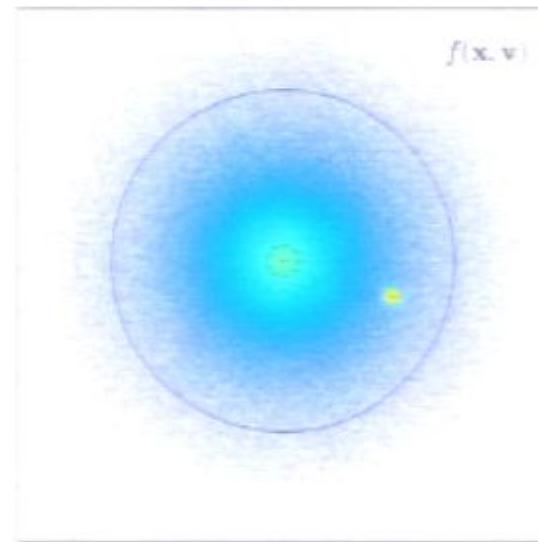
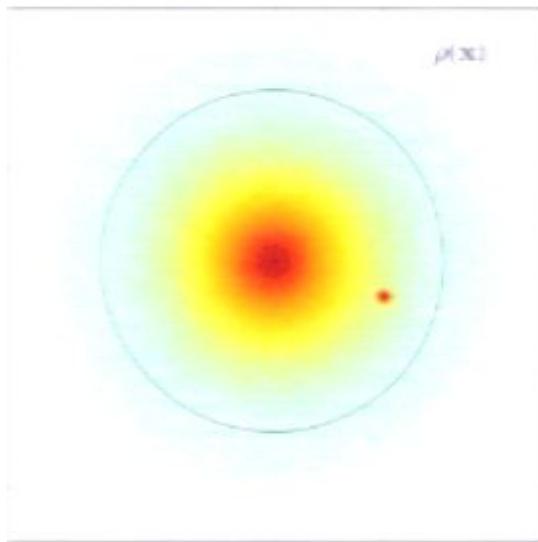
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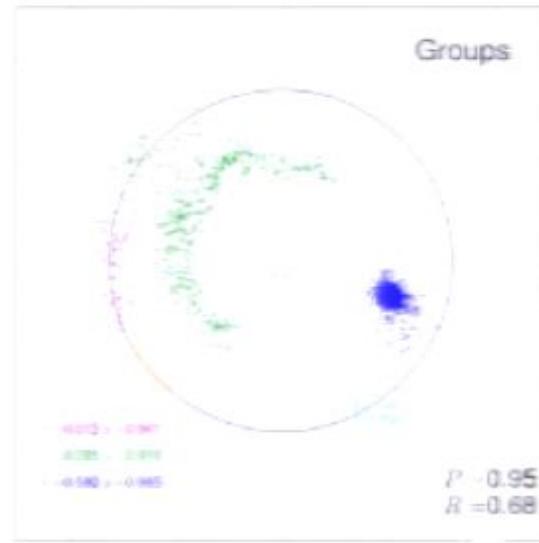
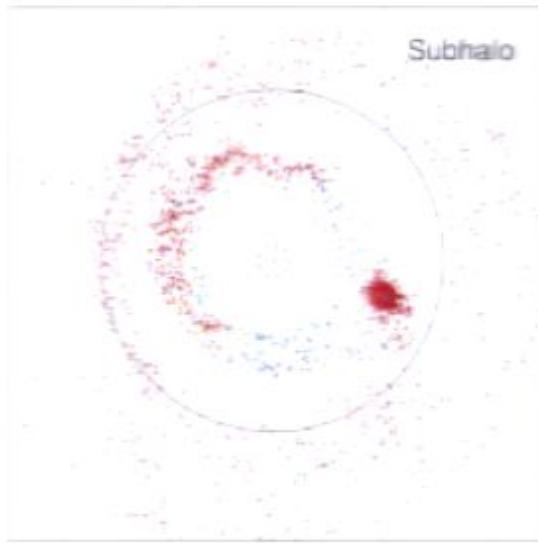
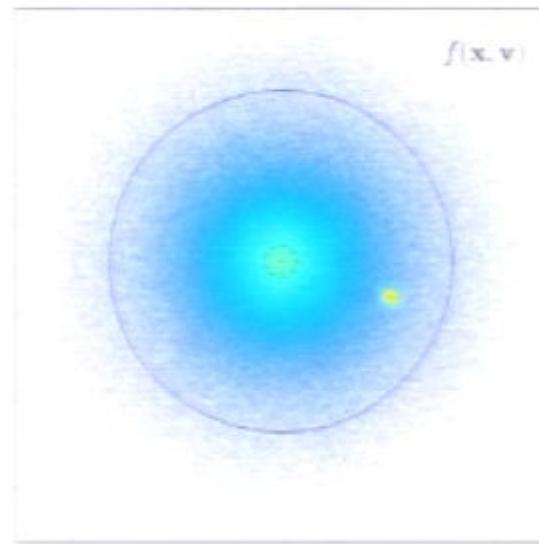
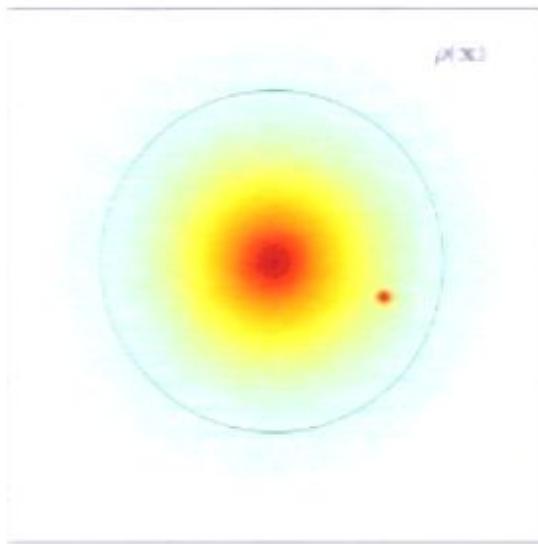


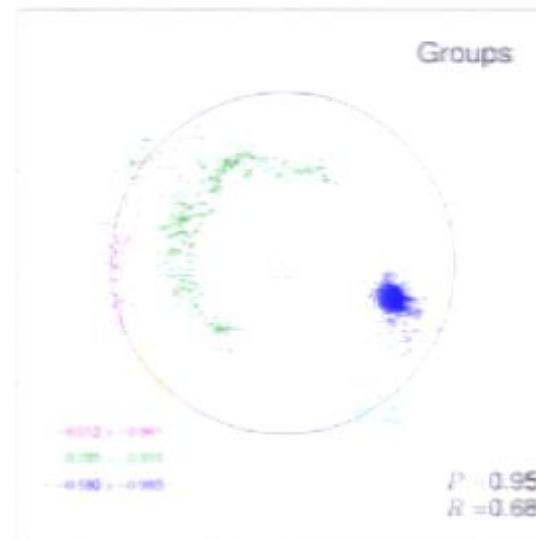
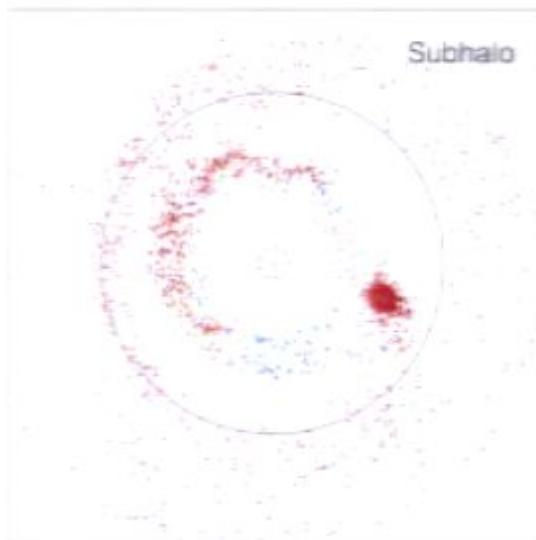
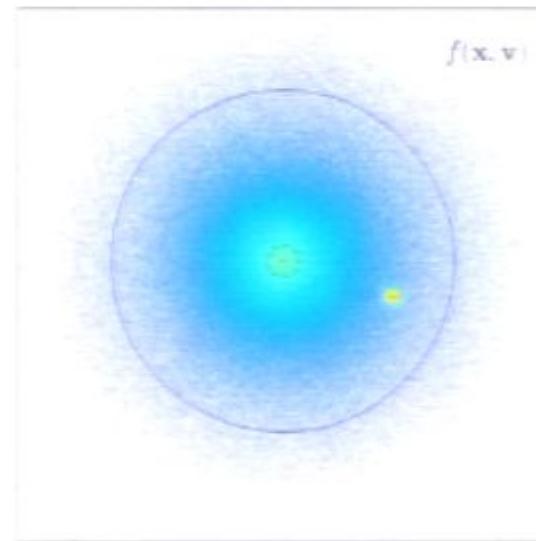
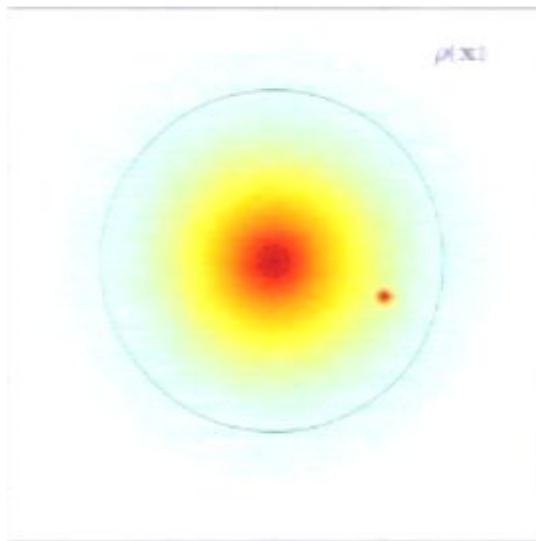
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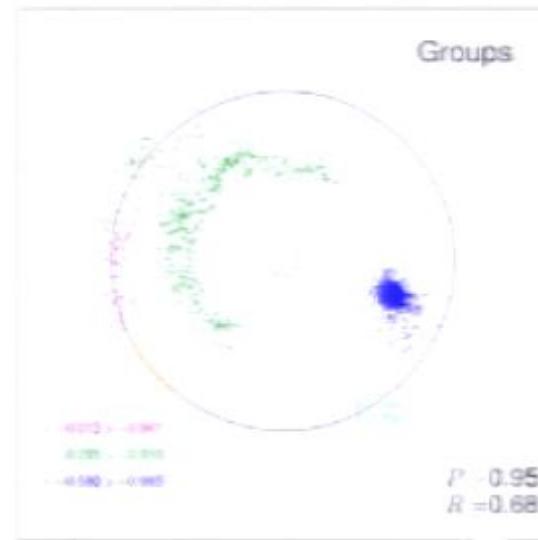
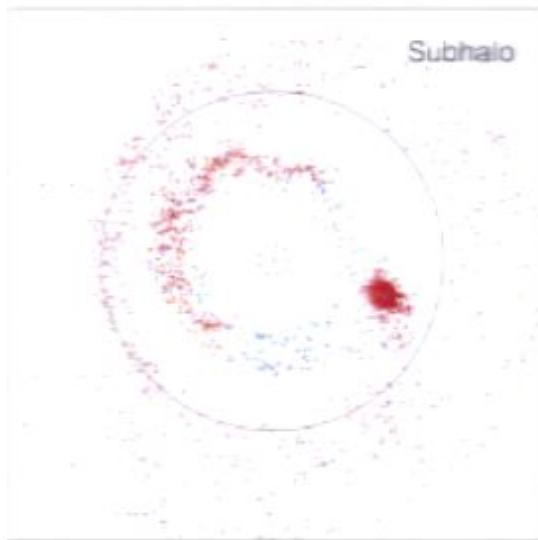
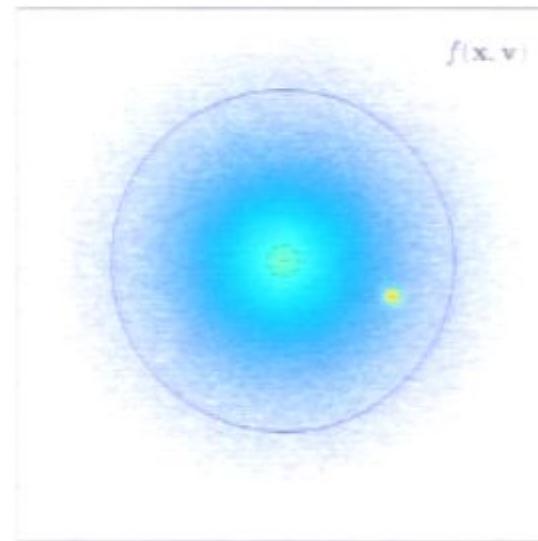
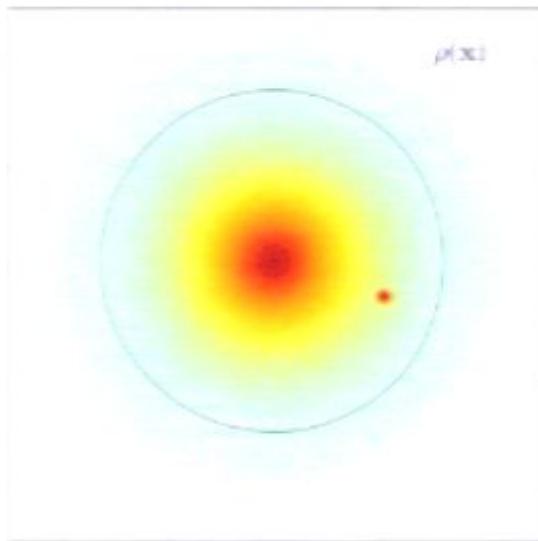


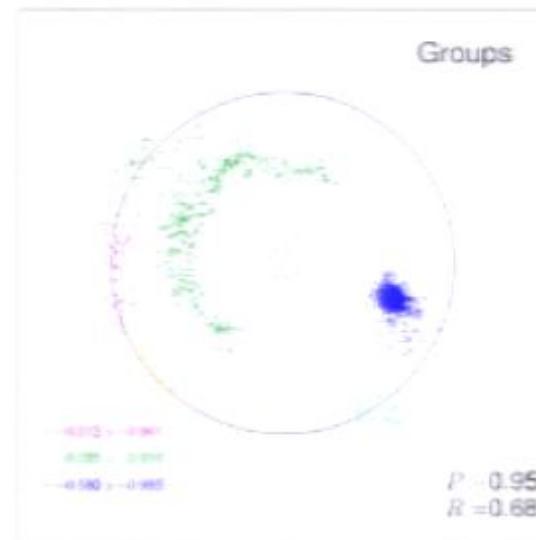
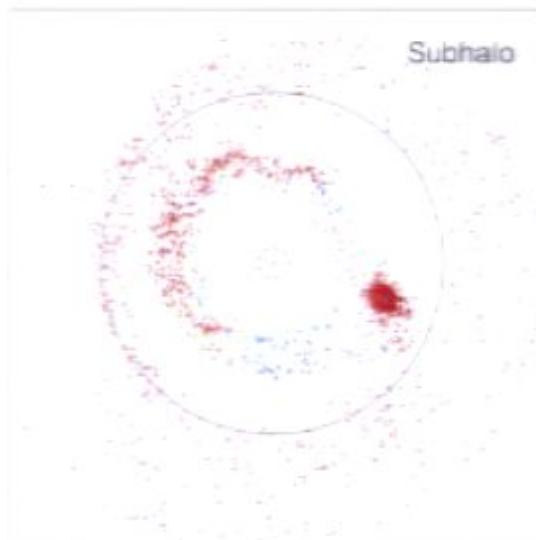
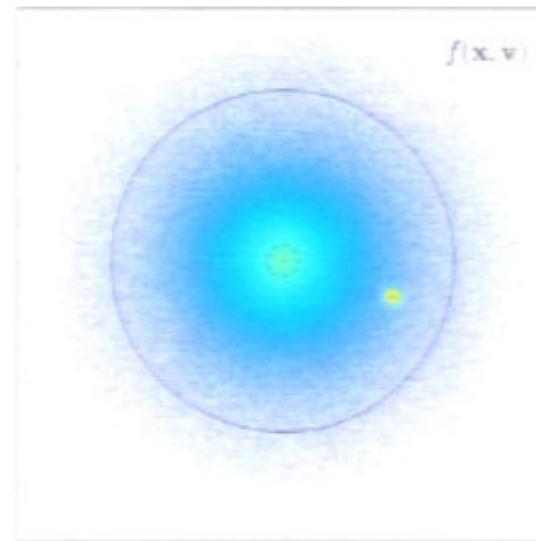
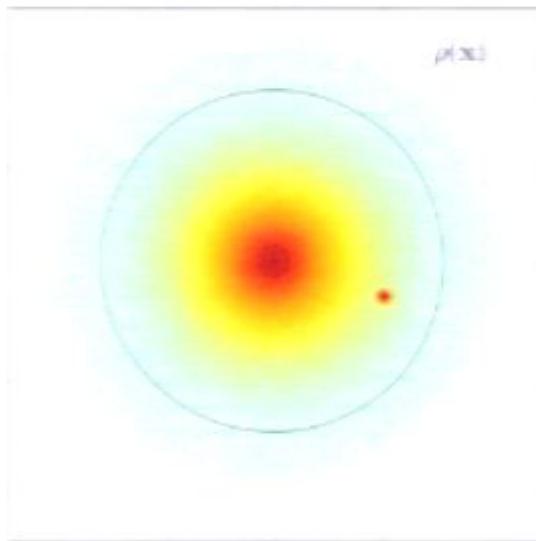


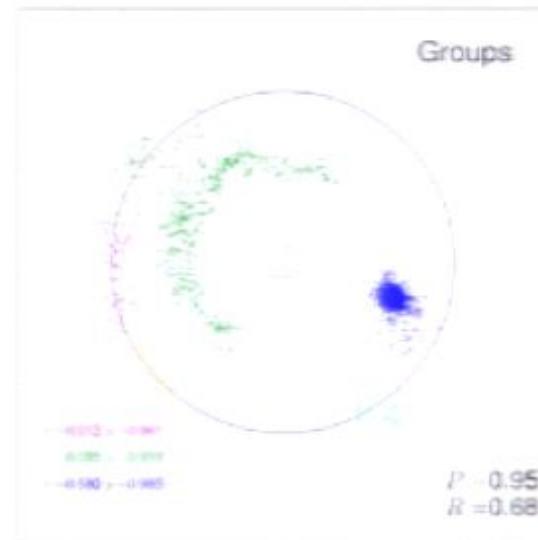
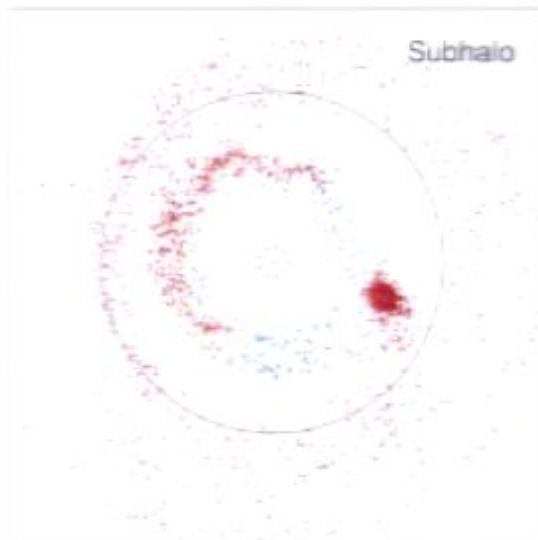
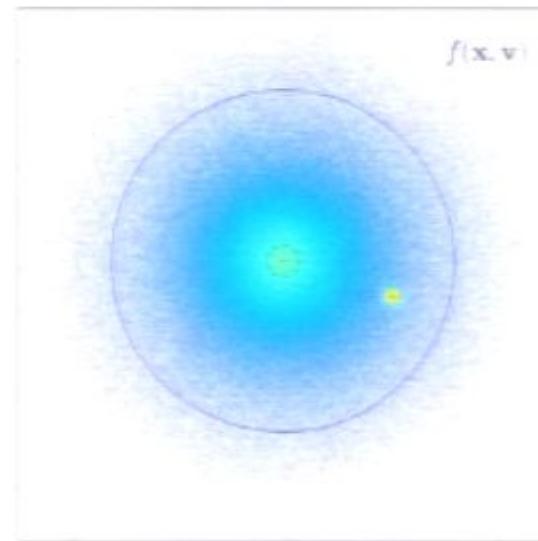
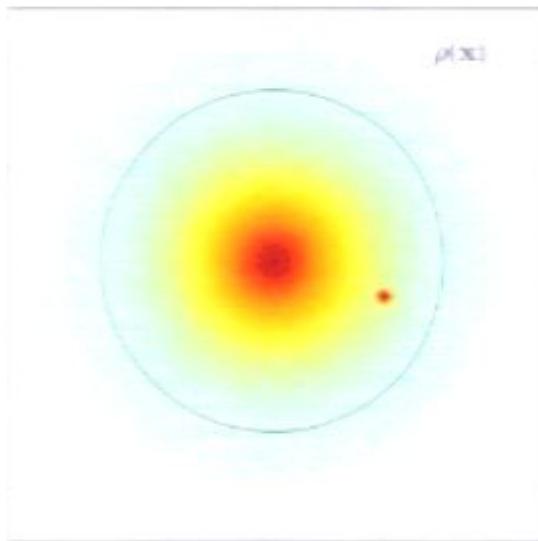


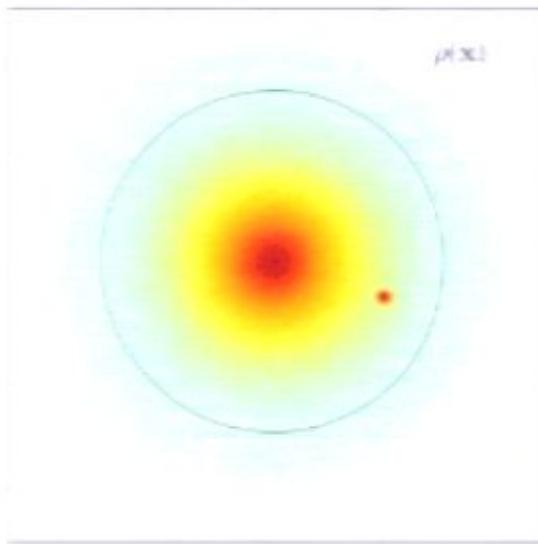




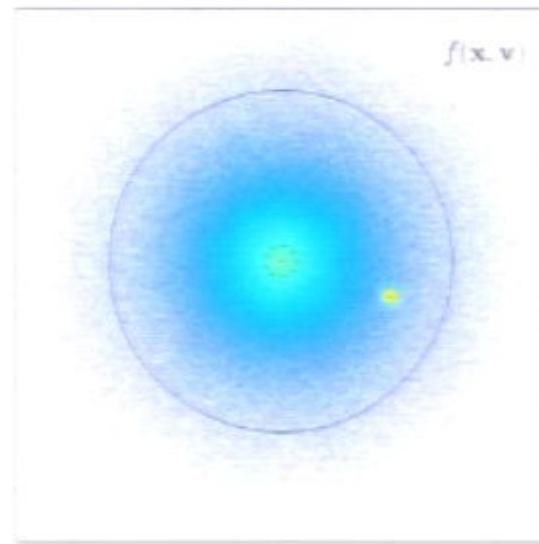




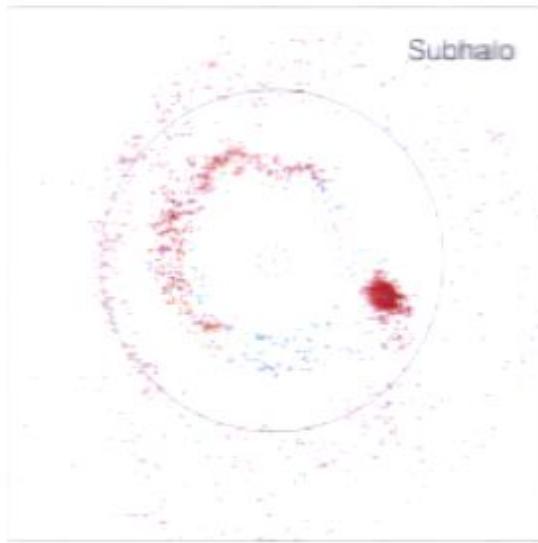




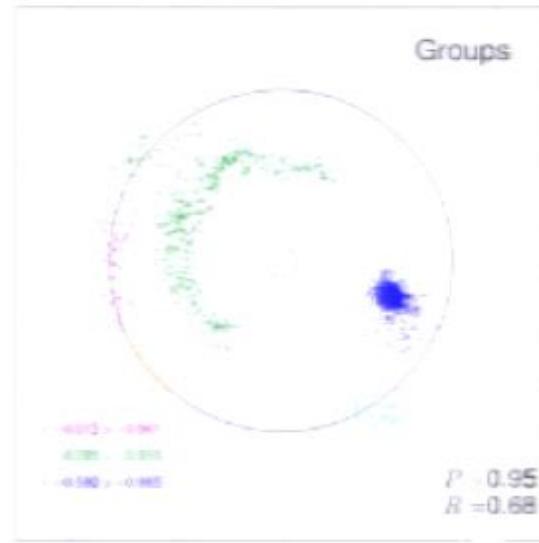
$\rho(\mathbf{x}_i)$



$f(\mathbf{x}, \mathbf{v})$



Subhalo



Groups

