

Title: A CoGeNT Modulation Analysis

Date: Sep 23, 2011 11:20 AM

URL: <http://pirsa.org/11090084>

Abstract: We analyze the recently released CoGeNT data with a focus on their time-dependent properties. Using various statistical techniques, we confirm the presence of modulation in the data, and find a significant component at high ($E_{\text{ee}} \gtrsim 1.5 \text{ keVee}$) energies. We find that standard elastic WIMPs in a Maxwellian halo do not provide a good description of the modulation. We consider the possibility of non-standard halos, using halo independent techniques, and find a good agreement with the DAMA modulation for QNa

A CoGeNT modulation analysis

Patrick Fox



with Joachim Kopp, Mariangela Lisanti
and Neal Weiner
(arXiv:1107.0717)

or related works see:

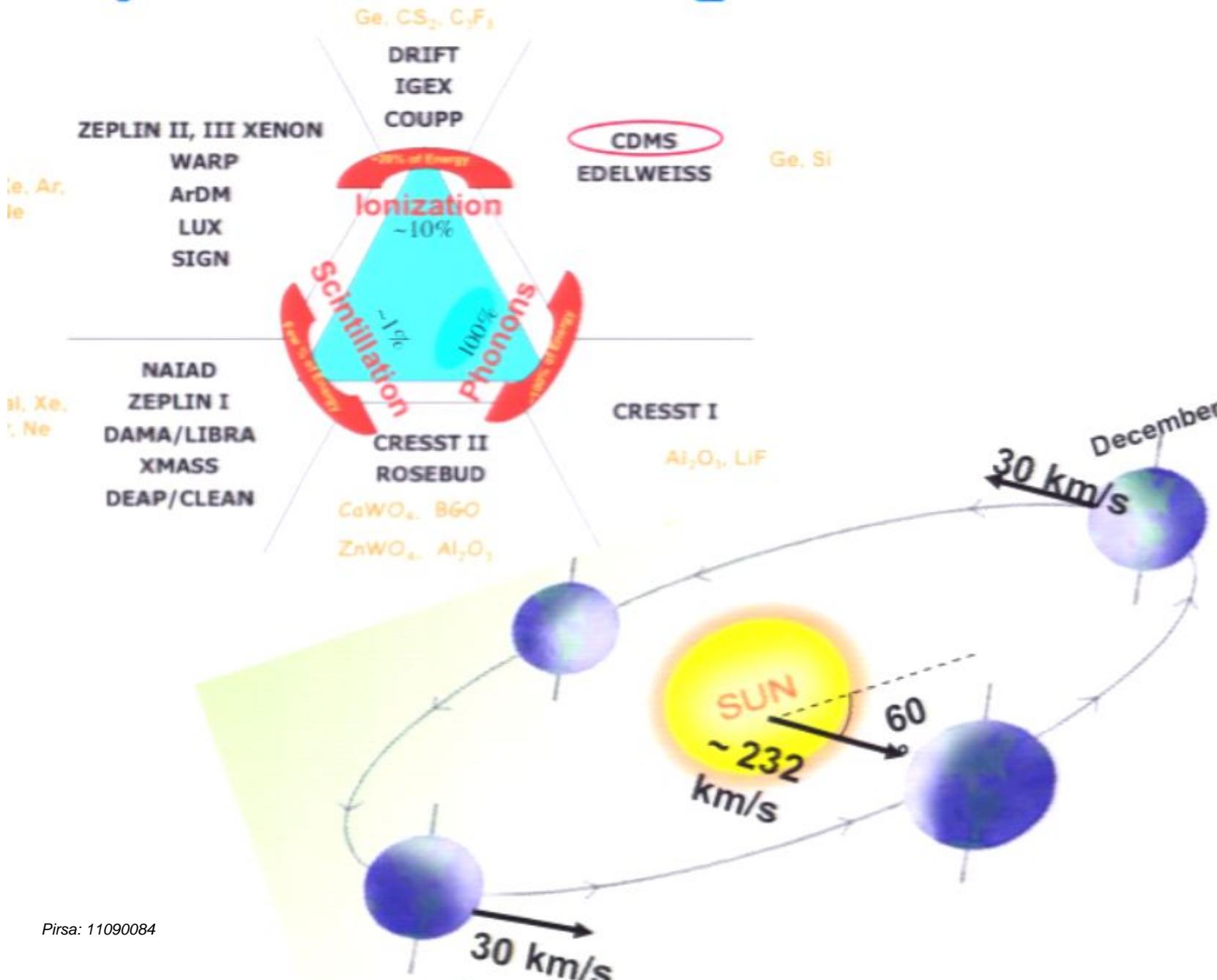
Dollar, Hall, Hooper and McKinsey (1007.1005)

Hooper and Kelso (1106.1066)

Zupan and Schwetz (1106.6241)

Farina, Pappadopulo, Strumia and Volansky (1107.0715)

Ways to find a DM signal

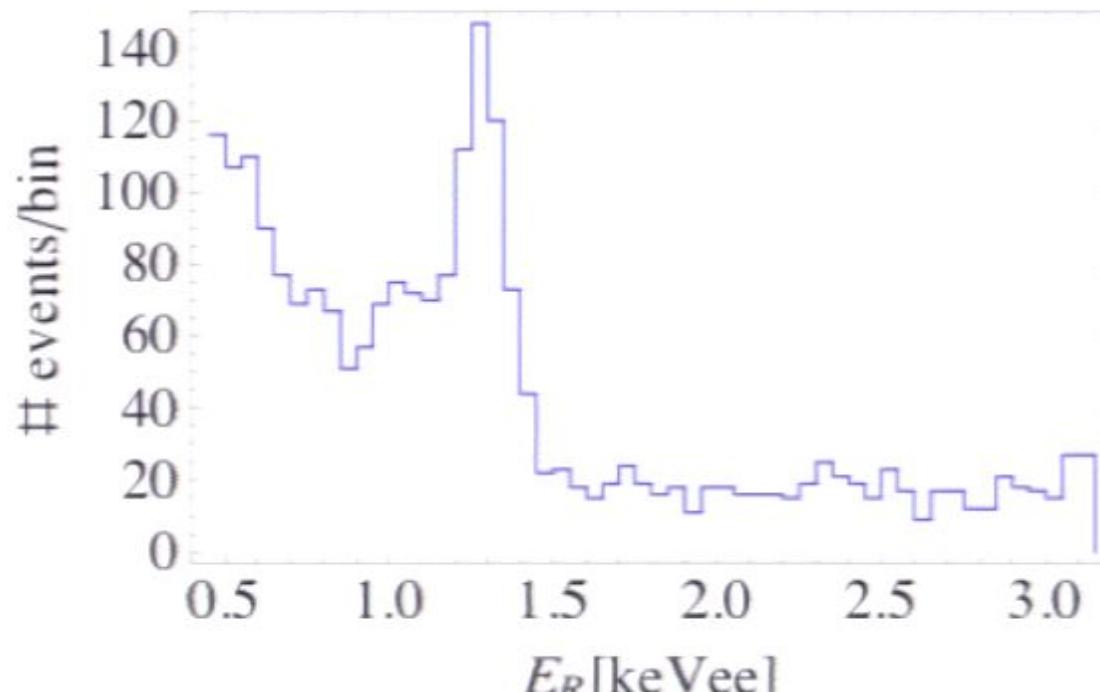


CoGeNT

- 330 g Ge detector, very low threshold
- Data taking [Dec. 4, 2009-Soudan fire]
• runs 458 days, 442 live days

Collaboration kindly provided time-stamped data

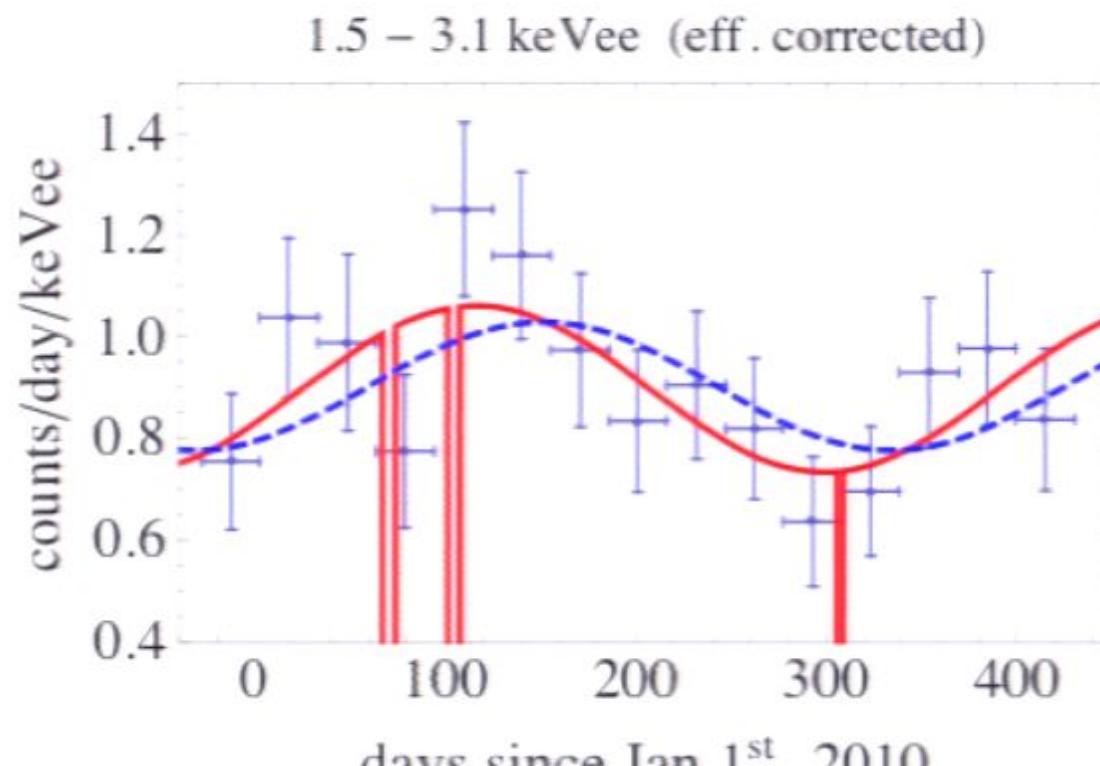
t [days]	E [keVee]
0.419763	0.576587
0.587532	0.362542
0.677681	1.11571
0.840586	1.35503
0.961931	0.414923
1.016340	2.22568
1.028094	1.82074
1.029199	2.67502
1.358892	0.872435
1.834660	0.329431
1.892096	0.366899
1.919907	1.30808
2.395985	1.05583
2.408826	1.04175
2.492828	0.6463
2.500264	0.443454
2.503603	1.04213
2.612478	0.563312
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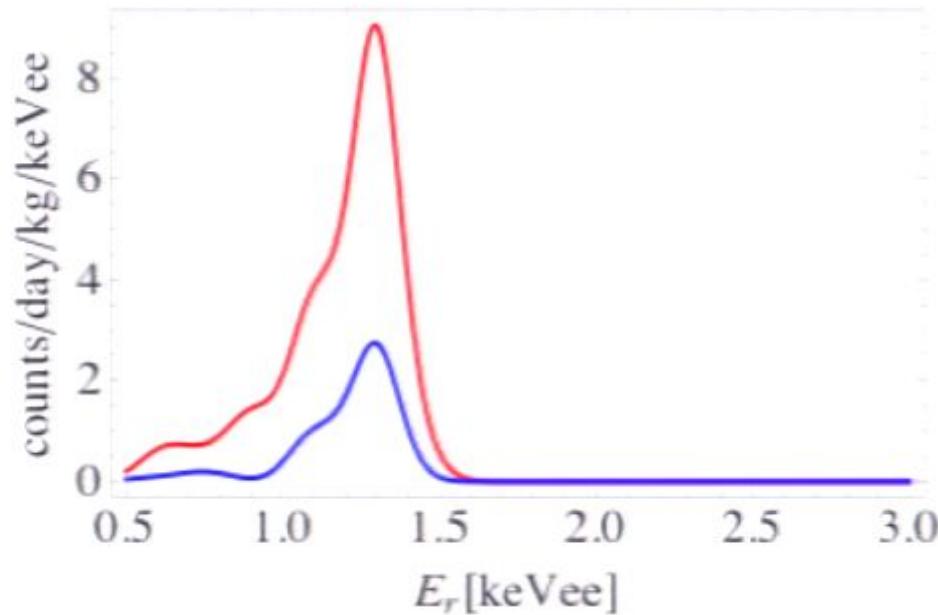
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Statistical techniques

After taking into account cosmogenic backgrounds



Fit what remains to $R(t) = A_0(1 + A_1 \cos(\omega(t - t_0))$

A_0 contains DM piece and constant background

SHM predicts $t_0 = 152$ days

Statistical techniques

Binned in energy and time (least squares)

Binned in energy unbinned in time (max likelihood)

Completely unbinned (max likelihood)

$$(t) = \left[0.33 \text{ kg} \times \Delta E \bar{f}_{\text{eff}}(E) \left(A_0 (1 + A_1 \cos(\omega(t - t_0))) \right) + \int_{E_{\text{low}}}^{E_{\text{high}}} f_{\text{cosmo}}(E, t) f_{\text{eff}}(E) \right] f_{\text{gaps}}(t)$$

$$2 \log L(A_0, A_1, \omega, t_0) = 2 \sum_i \phi(t_i) - 2 \int_{t_{\text{start}}}^{t_{\text{end}}} dt \phi(t)$$

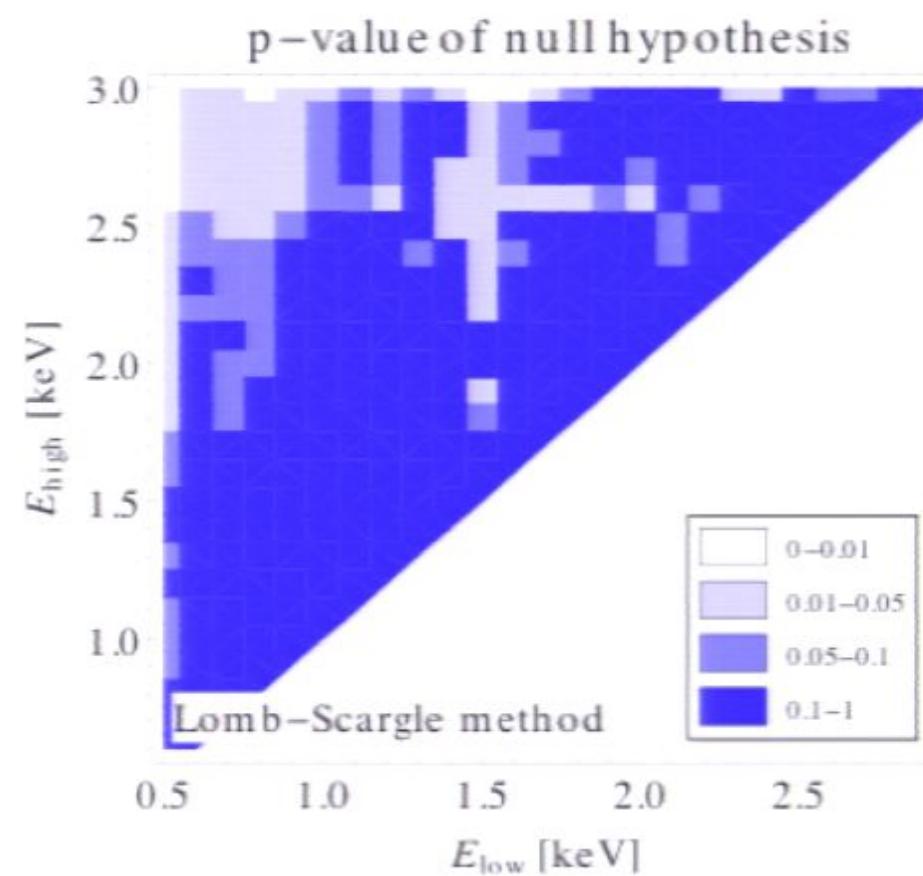
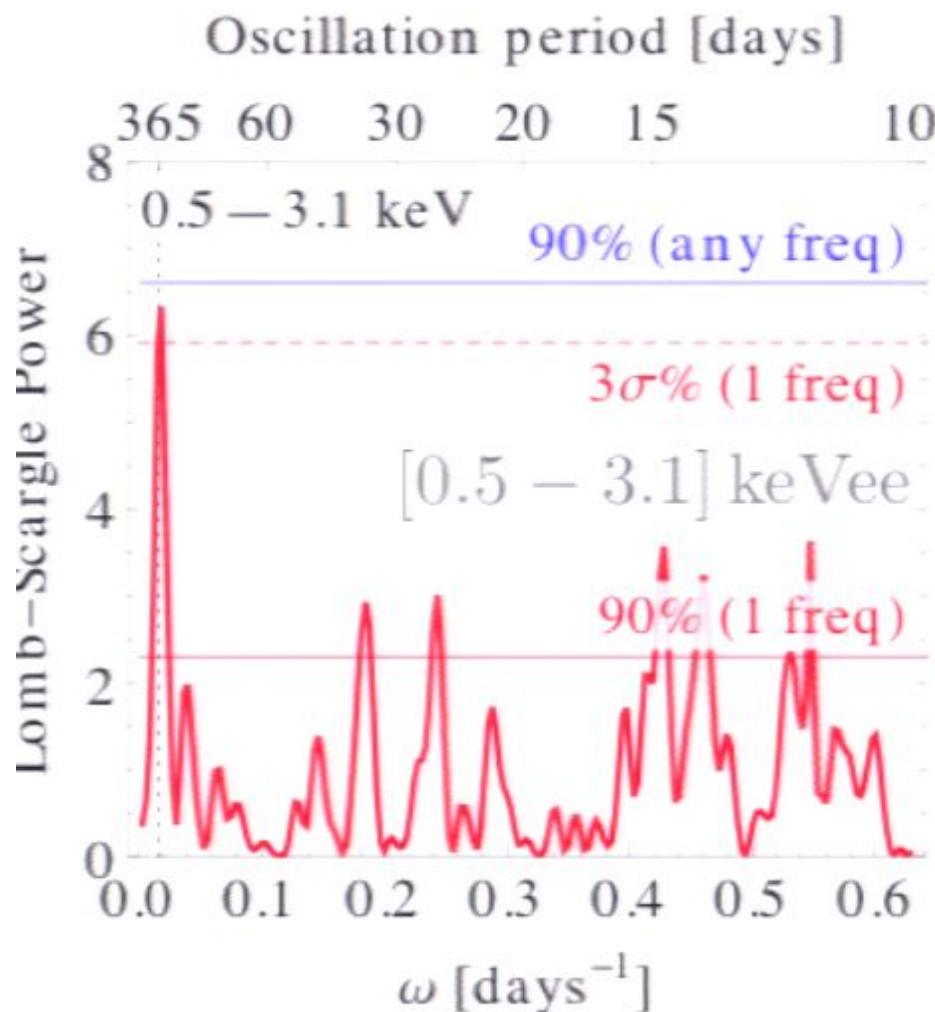
Weighted Lomb-Scargle periodogram

$$P(f) = \frac{1}{2\sigma^2} \left(\frac{\left[\sum_{i=1}^N W_i (y(t_i) - \bar{y}) \cos \omega(t_i - \tau) \right]^2}{\sum_{i=1}^N W_i \cos^2 \omega(t_i - \tau)} + \frac{\left[\sum_{i=1}^N W_i (y(t_i) - \bar{y}) \sin \omega(t_i - \tau) \right]^2}{\sum_{i=1}^N W_i \sin^2 \omega(t_i - \tau)} \right)$$

What's in the data?

Modulation frequency

(Lomb-Scargle)



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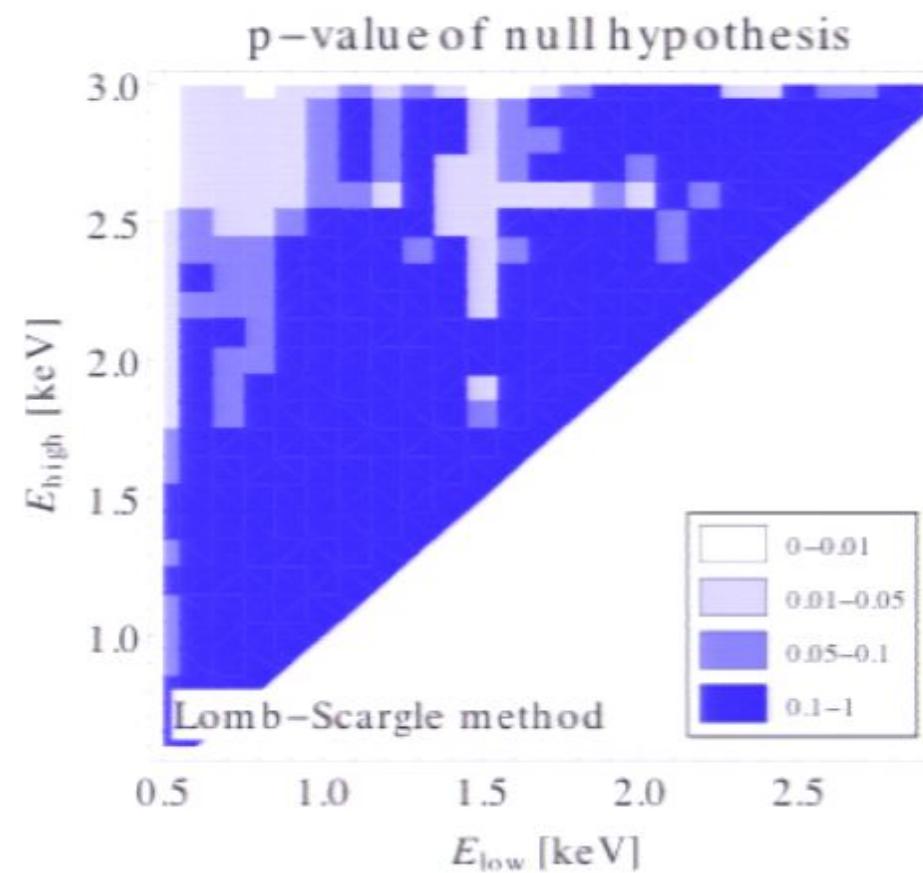
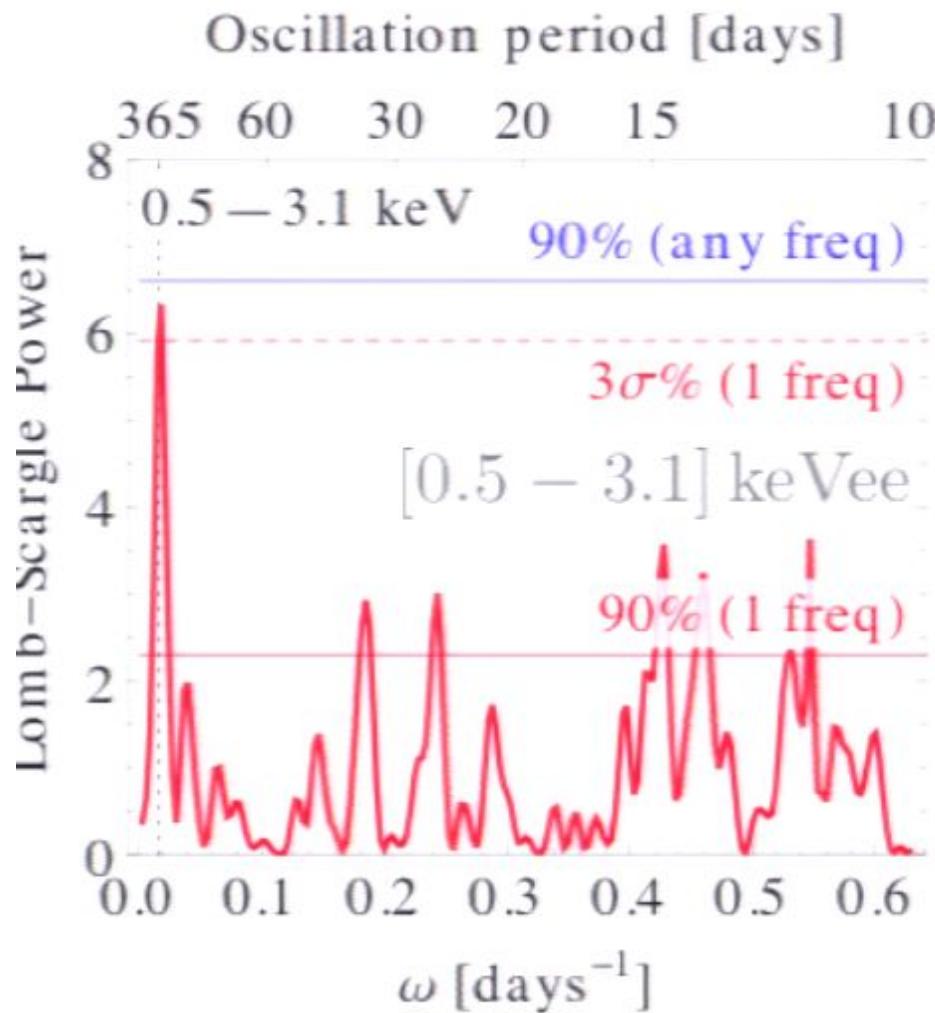
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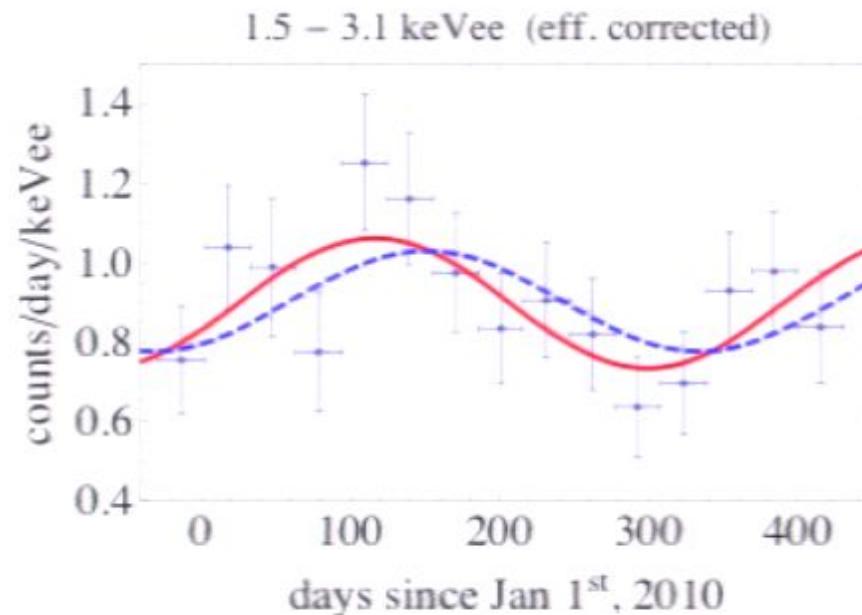
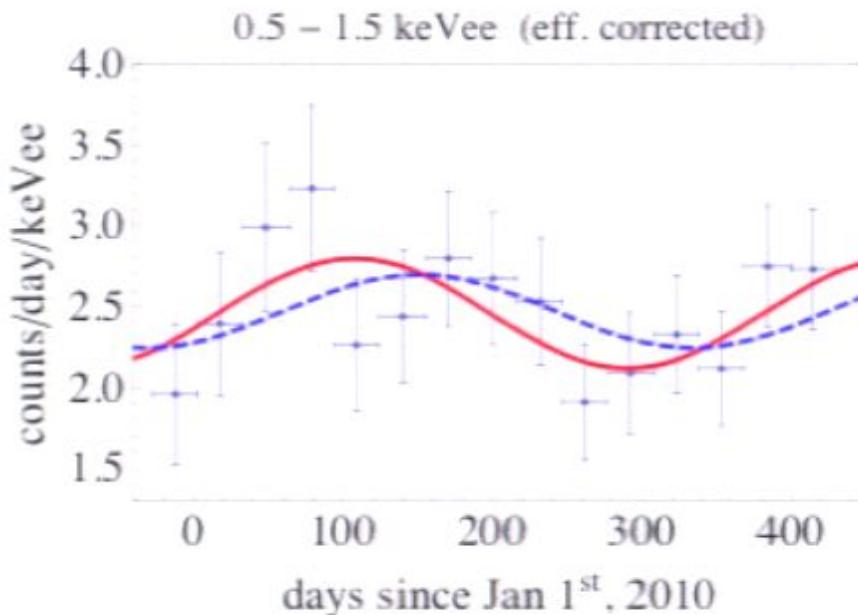
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Modulation frequency

(Lomb-Scargle)



Modulation amplitude (Binned)



Free phase

$$A_0 = 7.4 \text{ events/day/kg/keVee}$$

$$A_1 = 0.14 \quad t_0 = 107 \text{ days}$$

$$\Delta\chi^2 = 4.7$$

$$A_0 = 2.7 \text{ events/day/kg/keVee}$$

$$A_1 = 0.18 \quad t_0 = 116 \text{ days}$$

$$\Delta\chi^2 = 8.2$$

MB phase

$t_0 = 152$ days

$$A_0 = 7.5 \text{ events/day/kg/keVee}$$

$$A_1 = 0.09$$

$$\Delta\chi^2 = 2.3$$

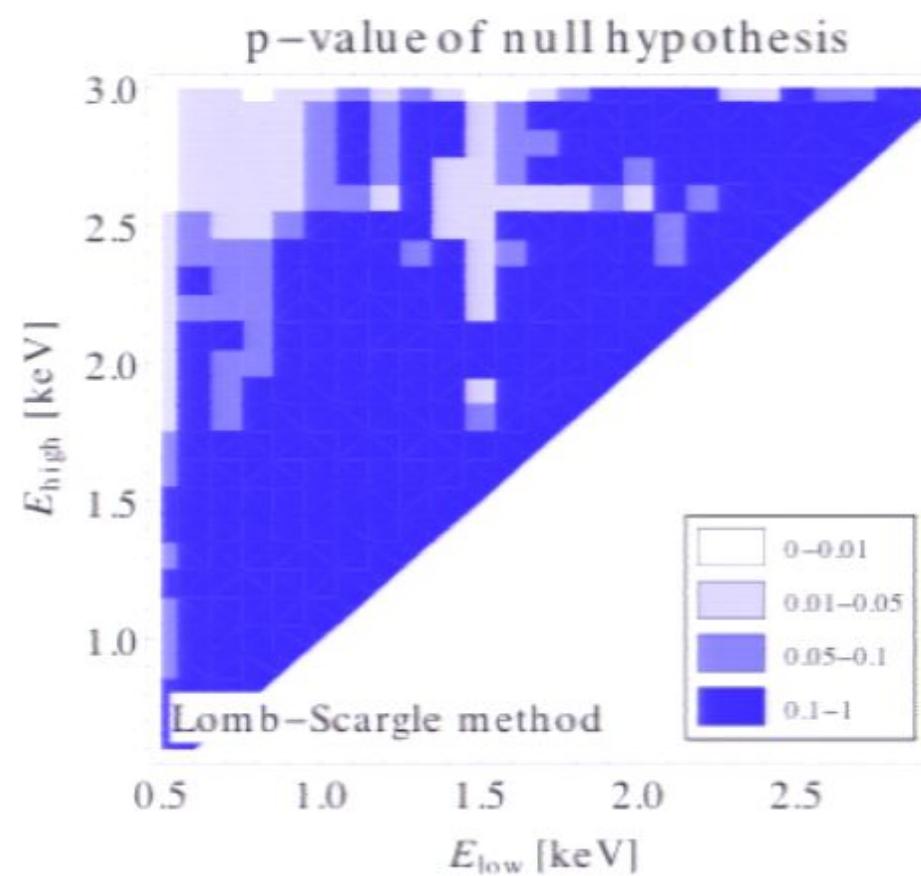
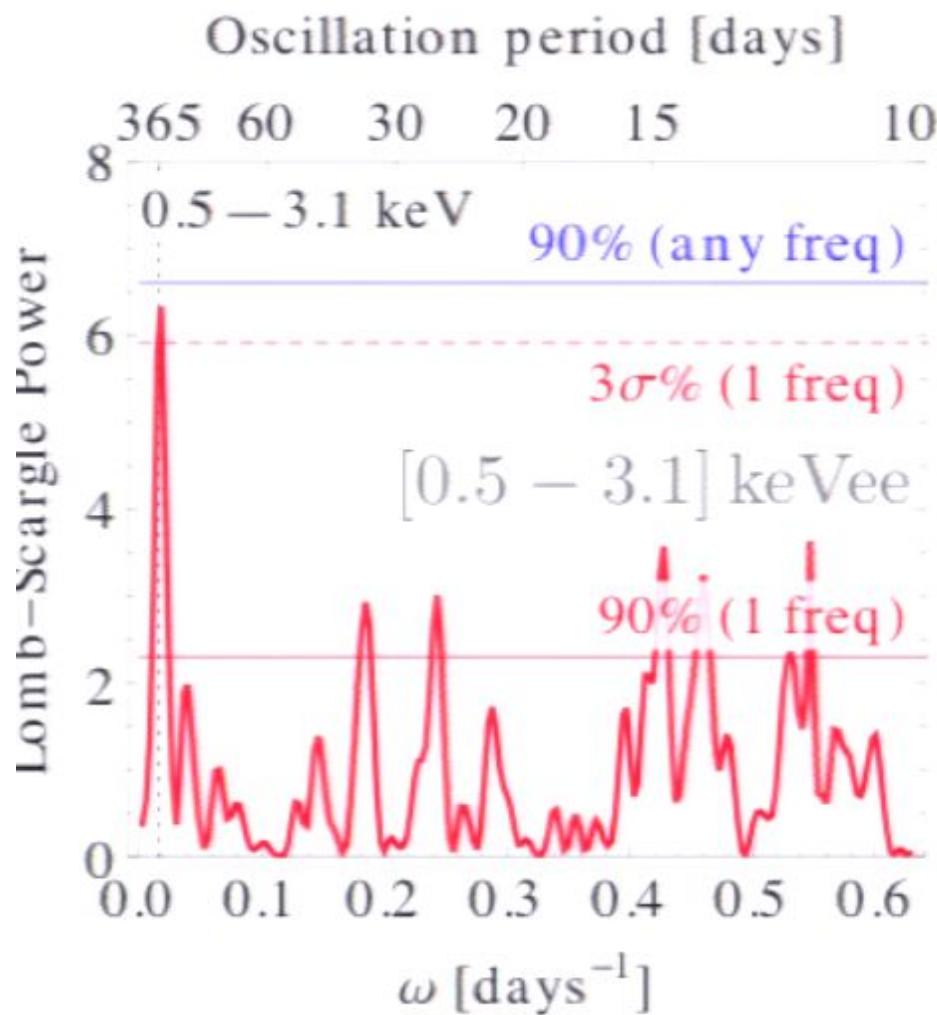
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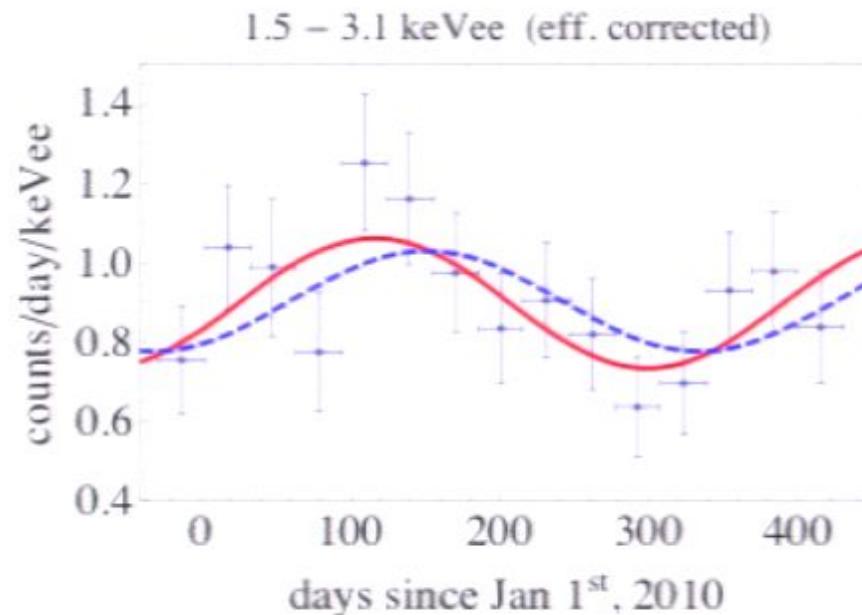
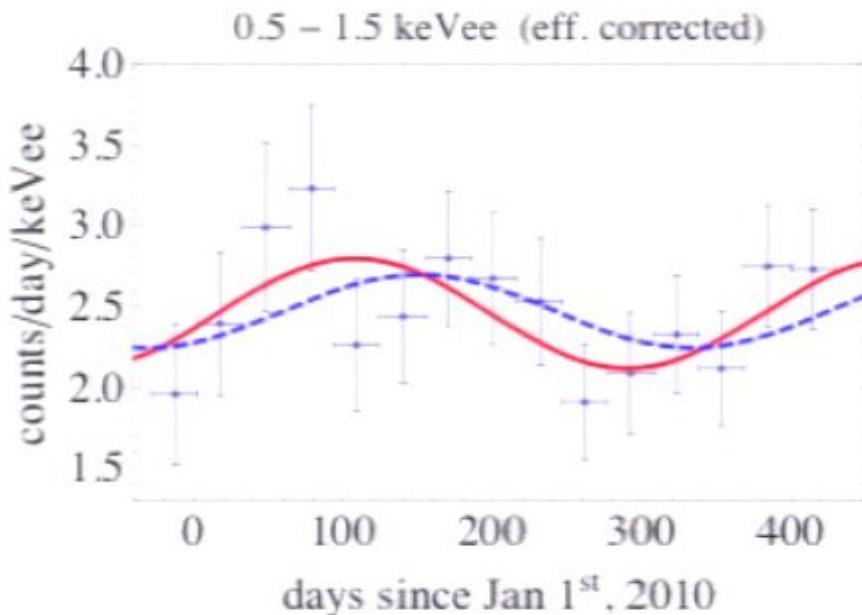
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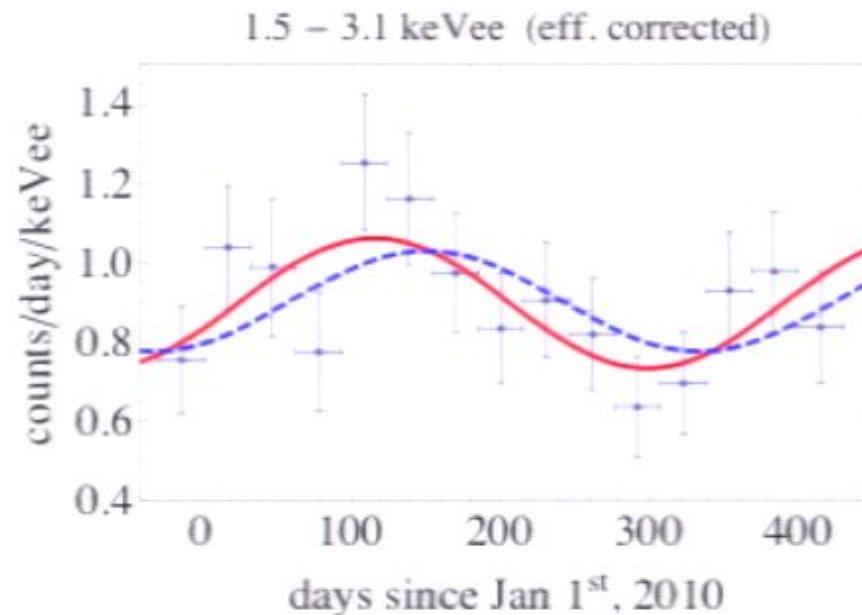
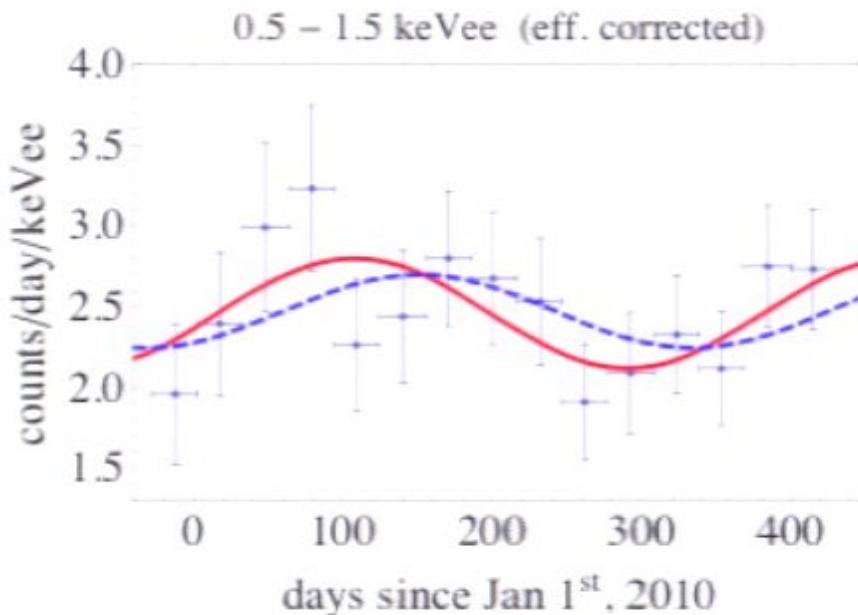
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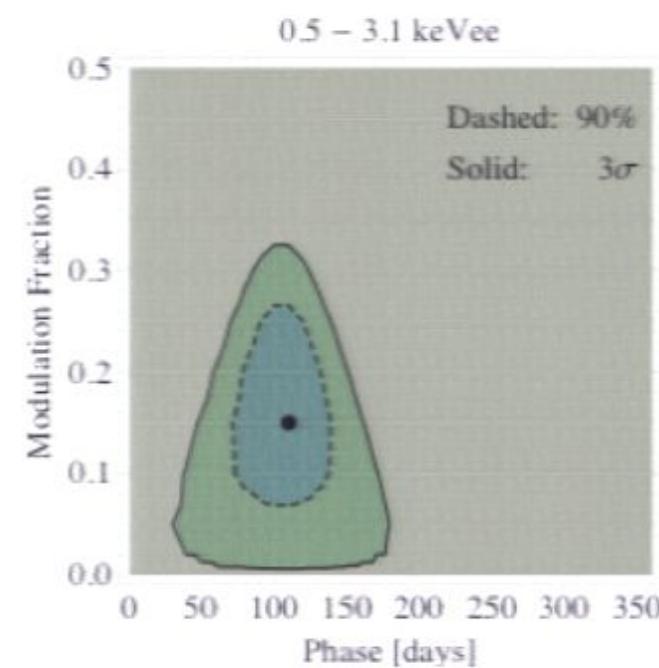
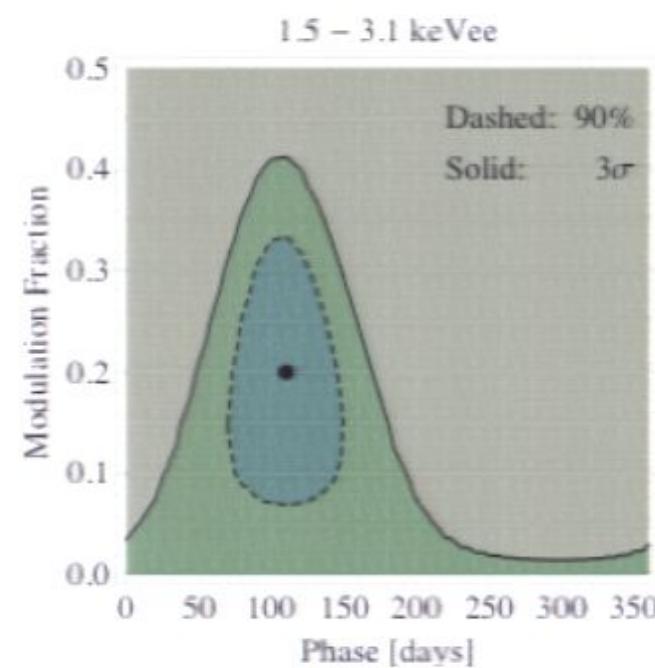
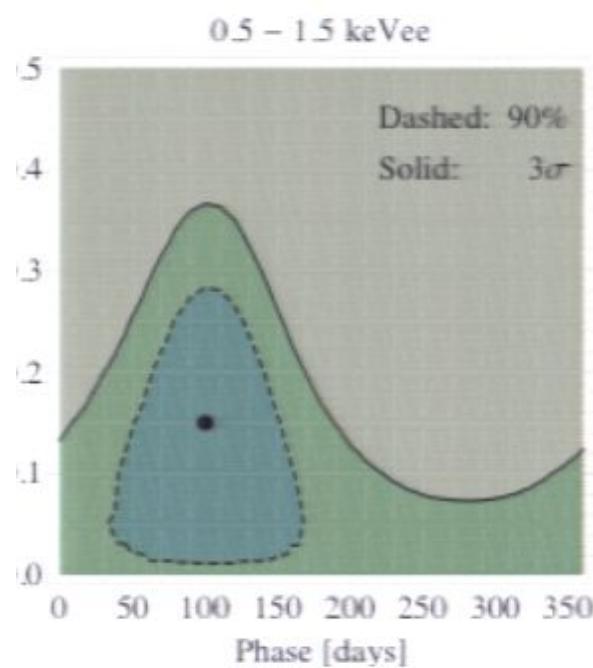
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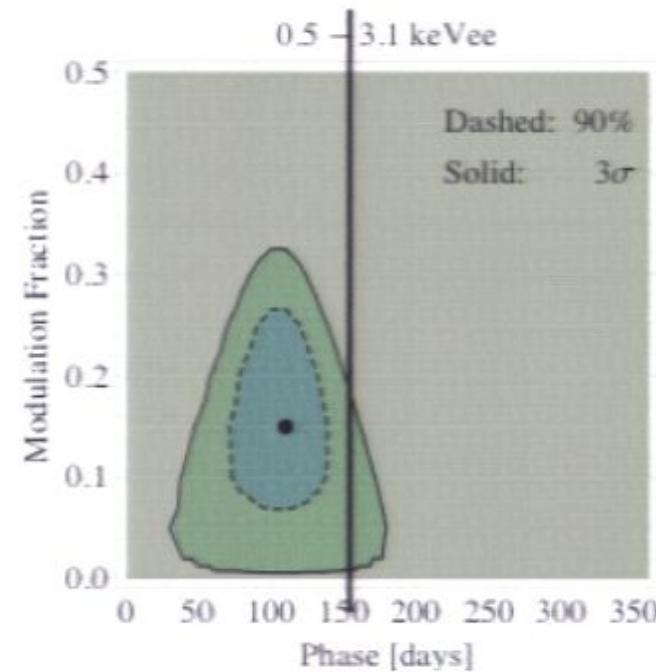
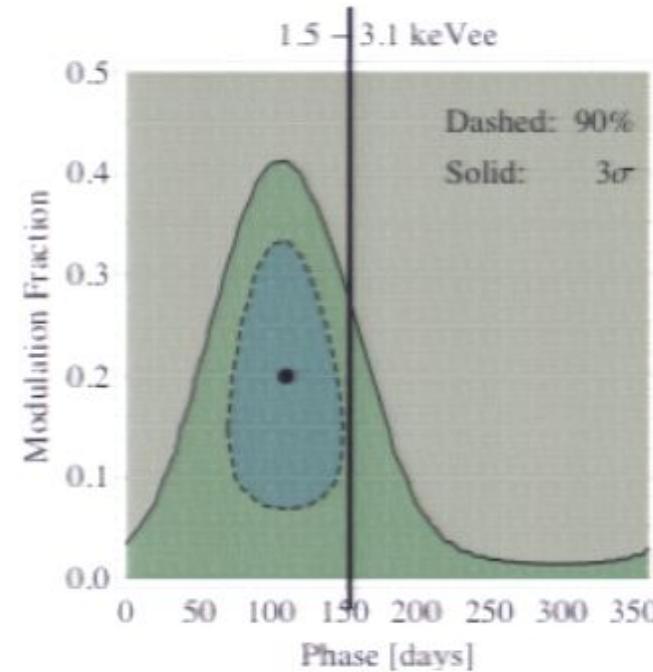
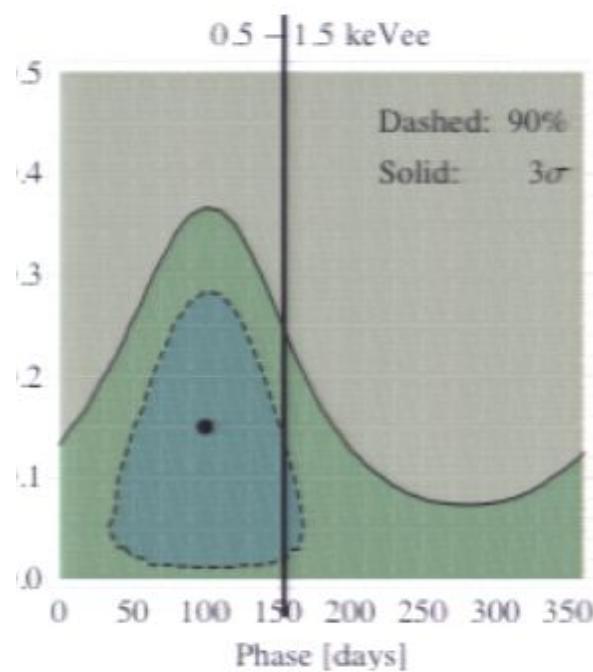
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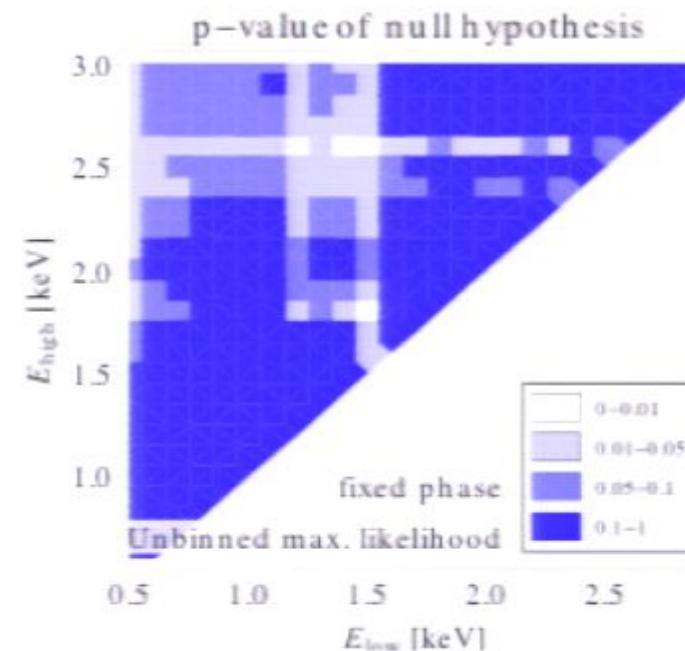
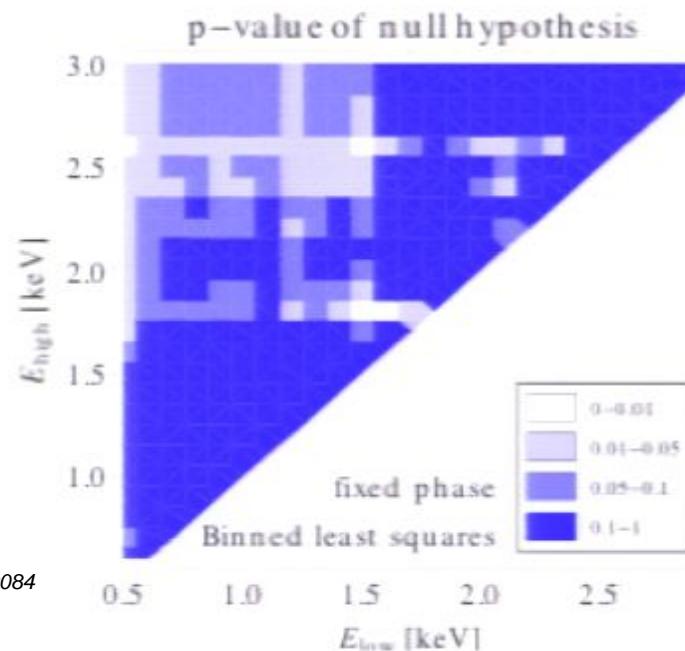
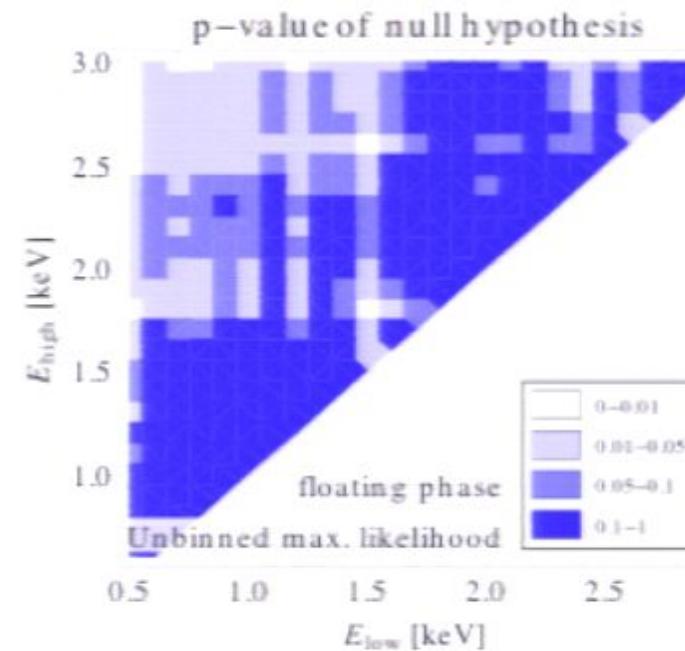
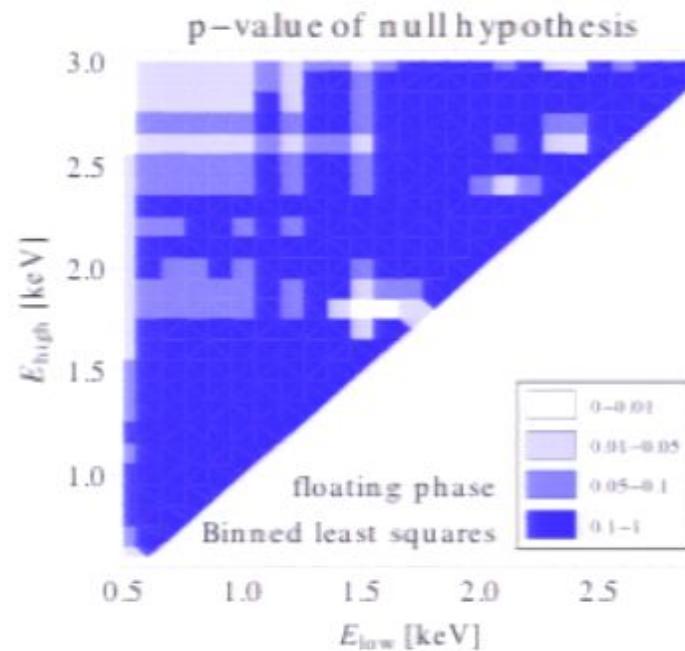
Phase and amplitude (Maximum likelihood)



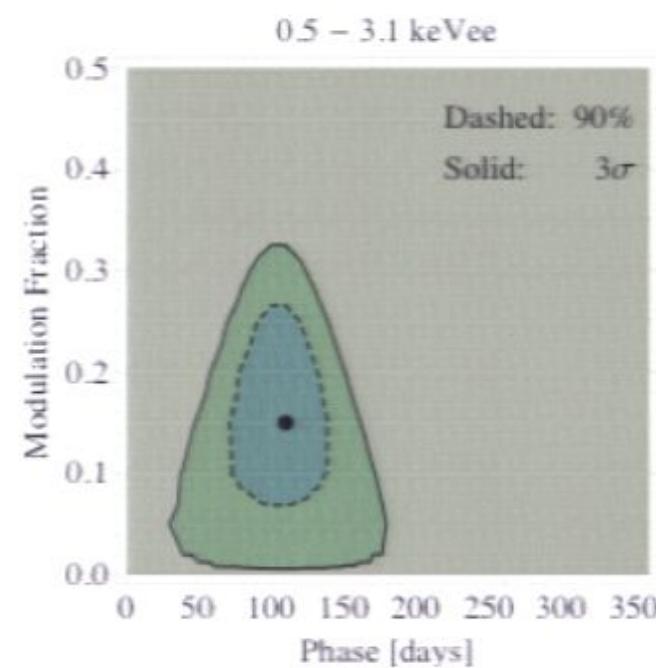
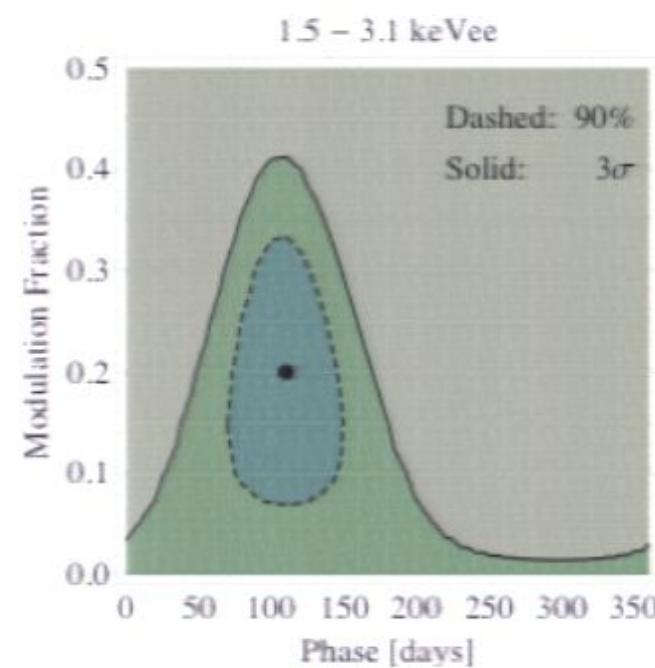
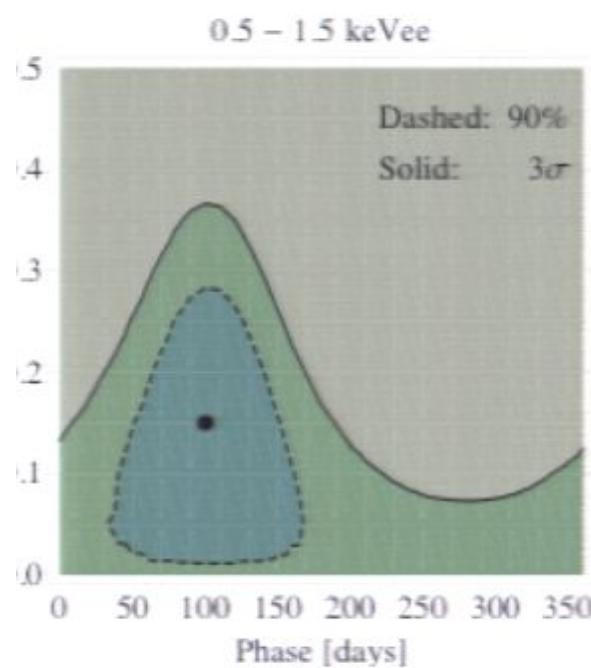
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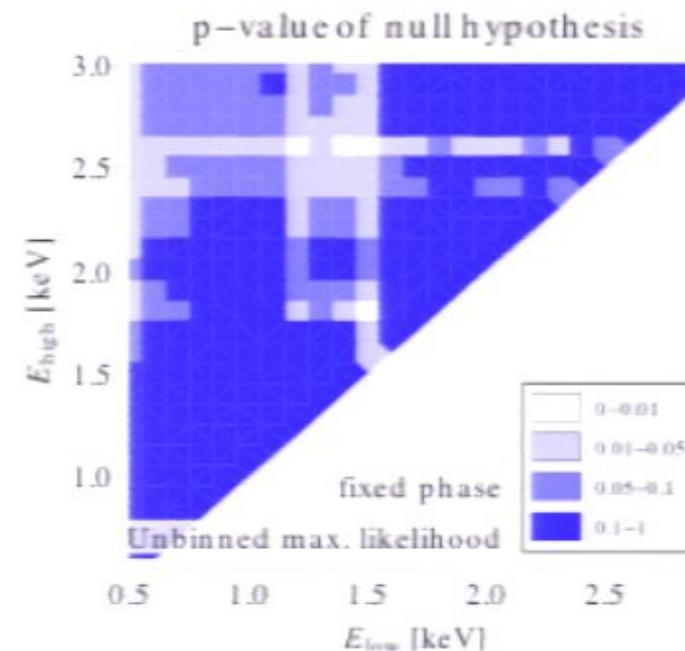
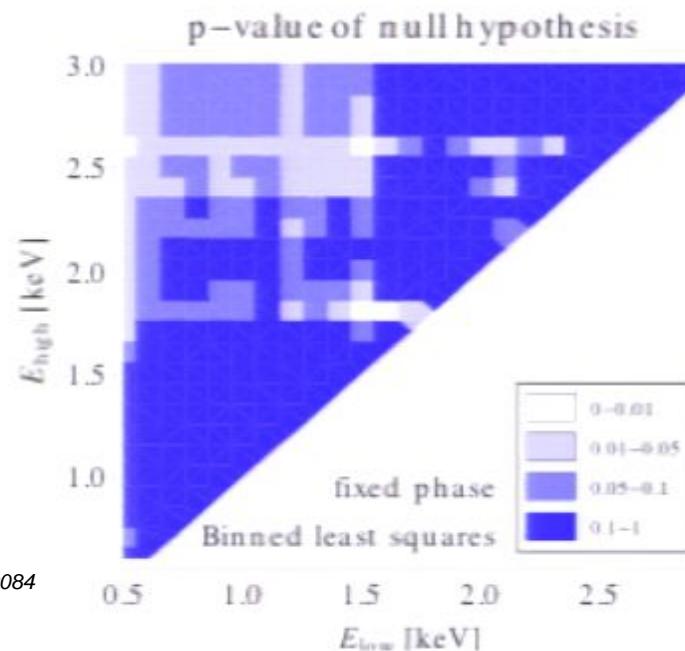
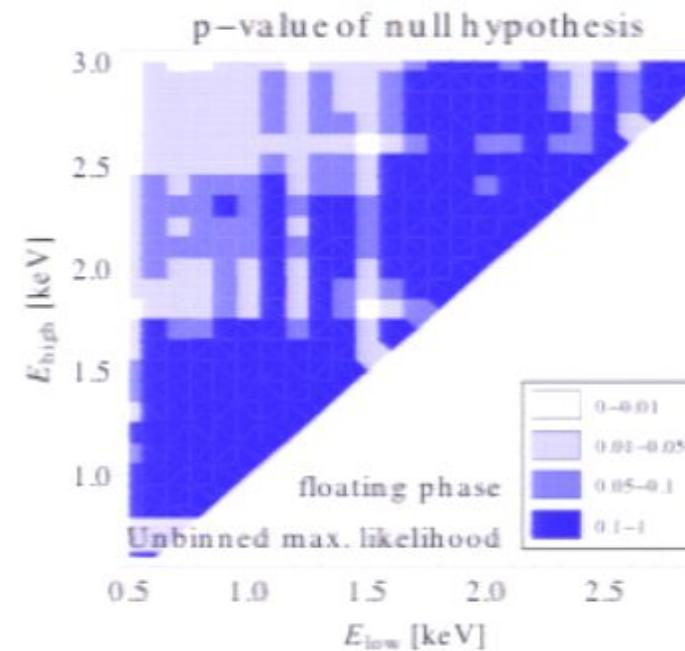
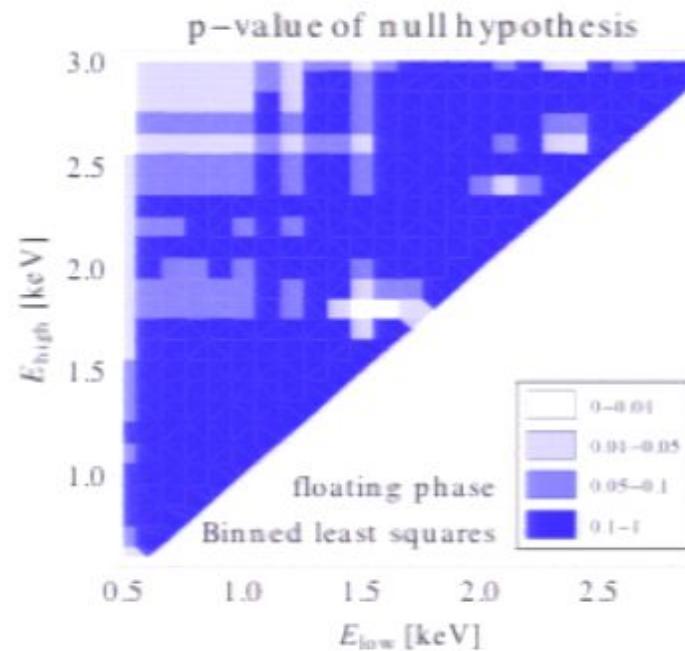
p-values



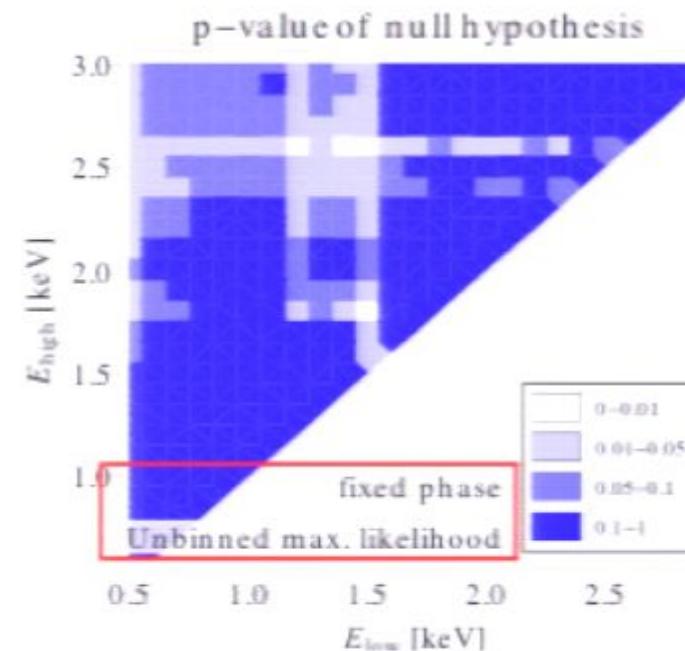
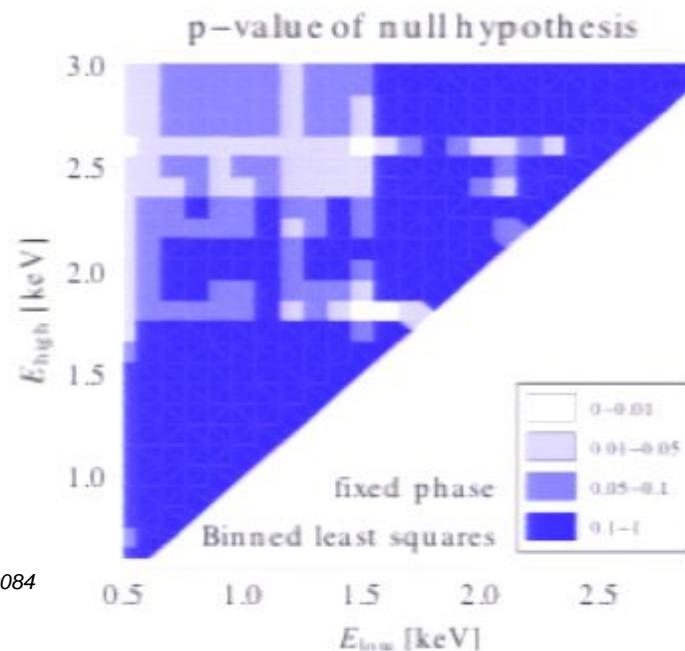
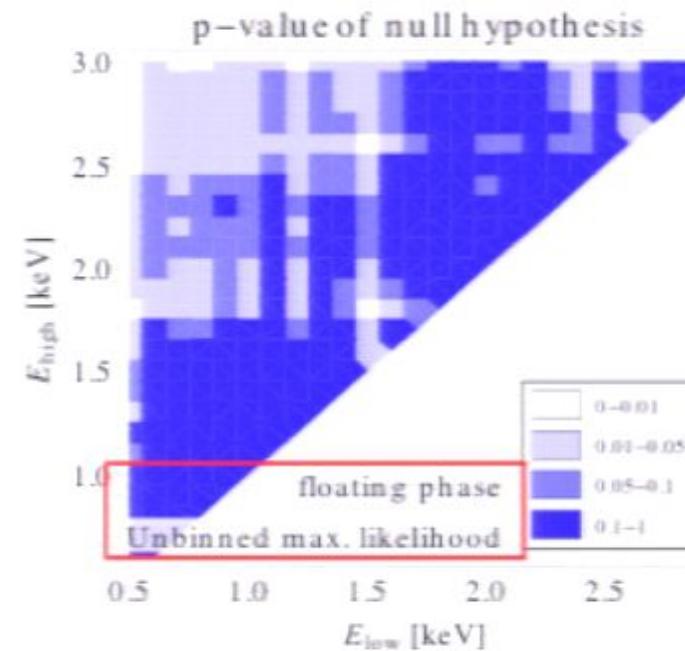
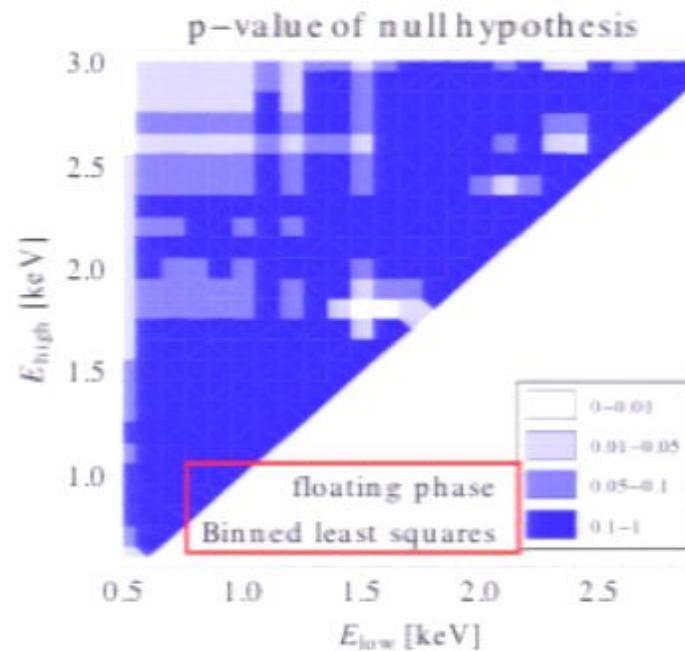
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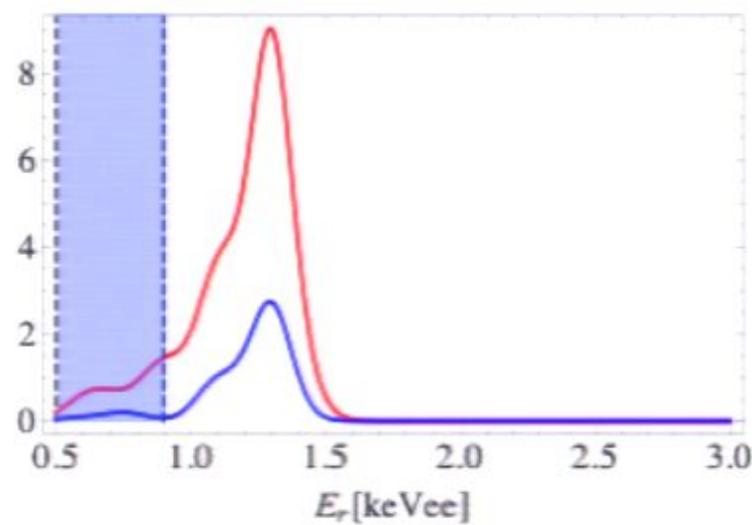
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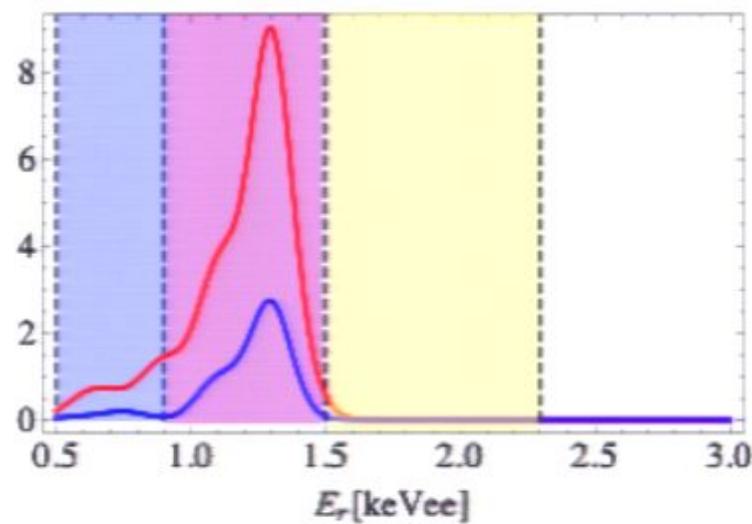
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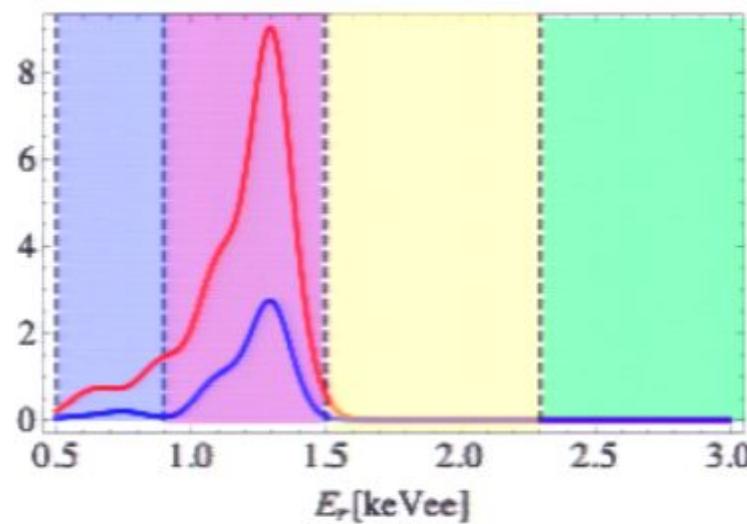
Spectra



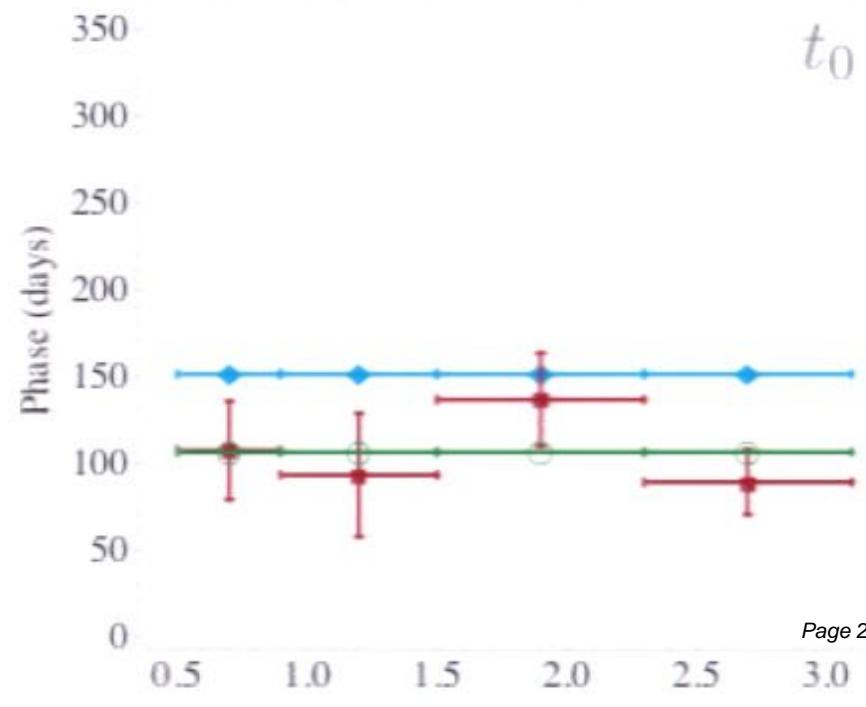
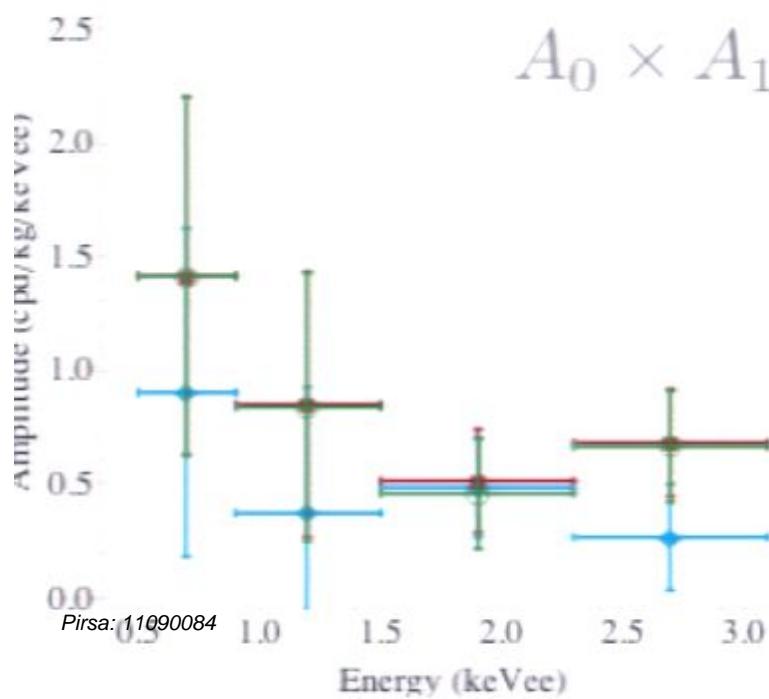
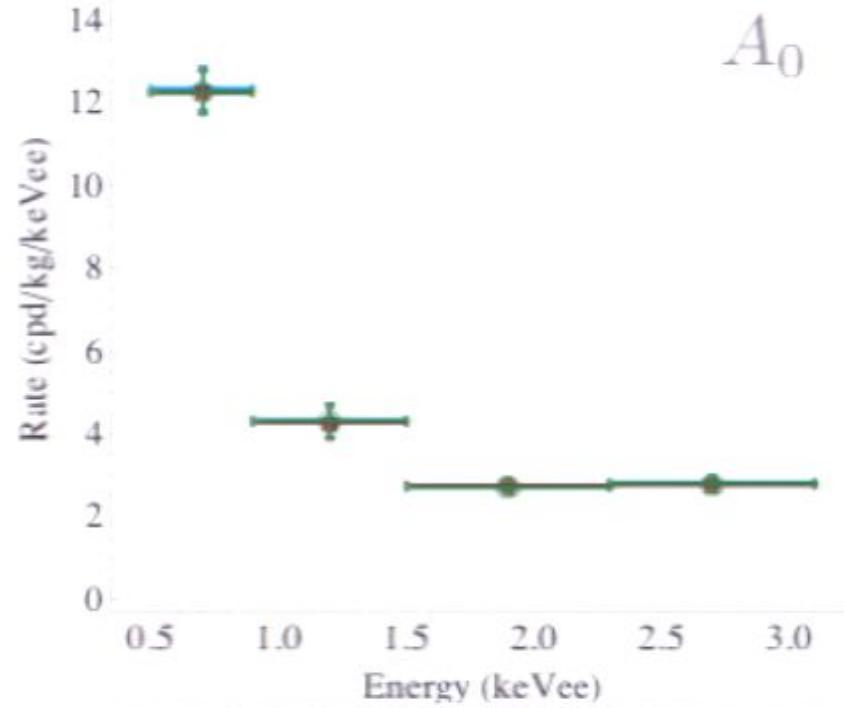
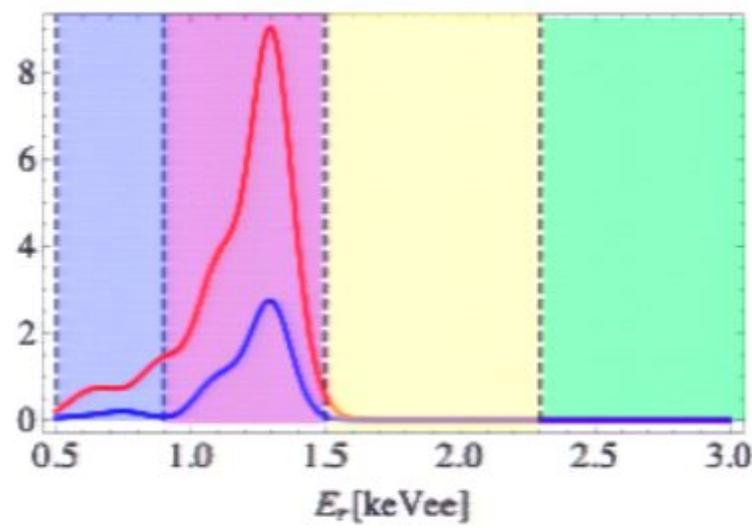
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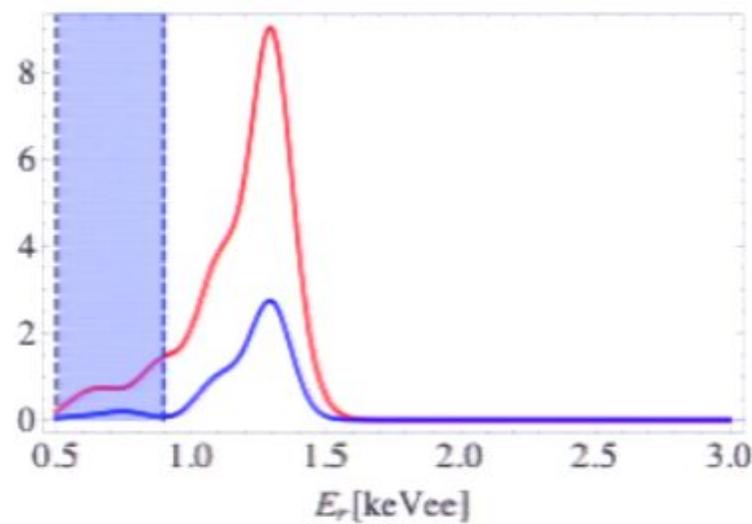
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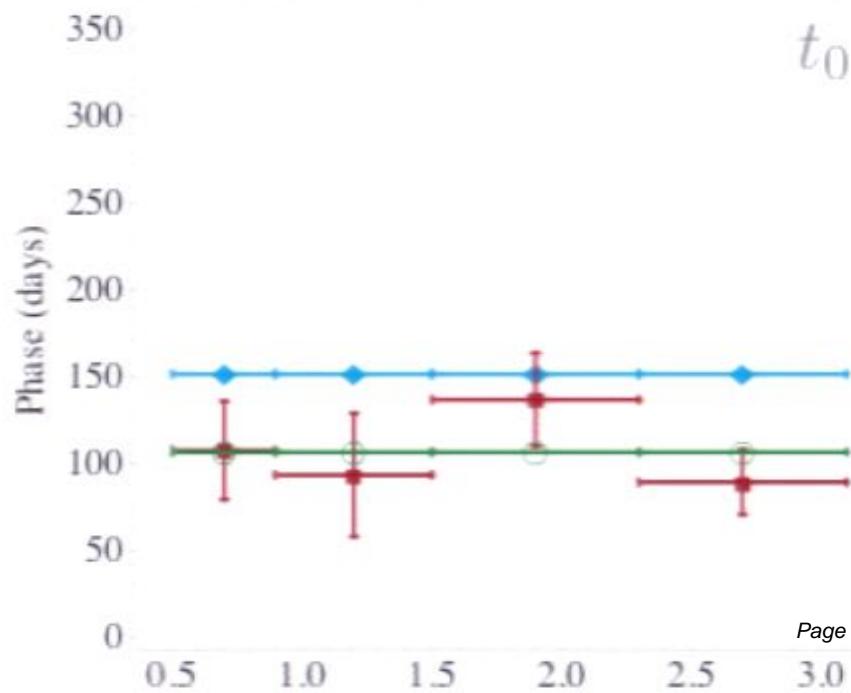
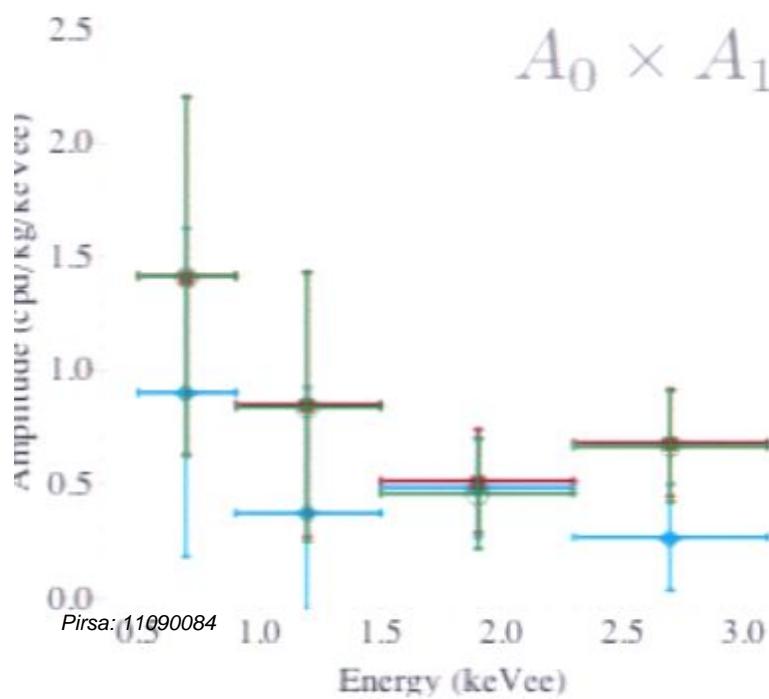
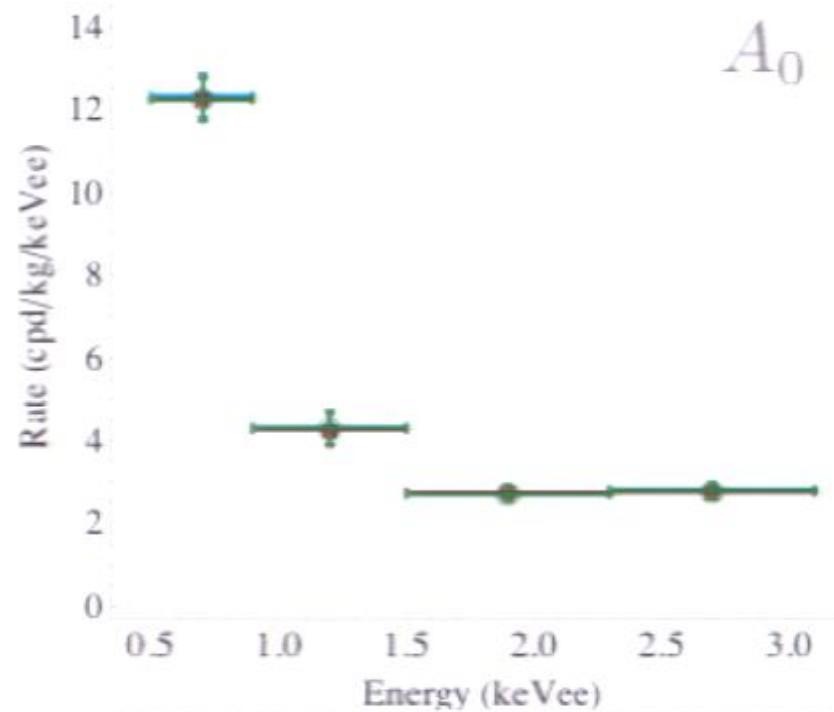
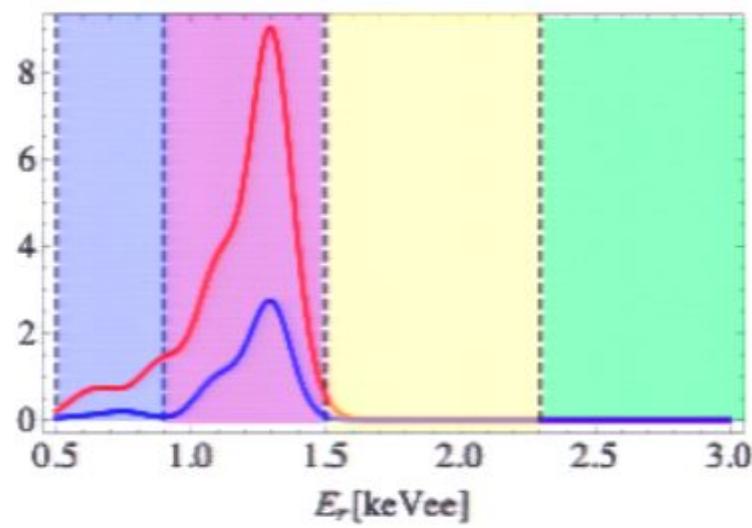
Data Summary

There is modulation in the data 2-3 sigma
Most significant at high recoil energy, ~20% modulation
Best fit phase ~106 days
Does not appear to be consistent with vanilla WIMP (MB
alo etc)

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**What about
Maxwell
Boltzmann?**

DM fits

$$f(v) \propto (e^{-v^2/v_0^2} - e^{-v_{esc}^2/v_0^2}) \Theta(v_{esc} - v)$$

$$v_{\text{esc}} = 550 \text{ km/s} \quad v_0 = 220 \text{ km/s}$$

Scenario	Spect+Mod	Spect only	Mod only ($c_i \geq 0$)	Mod only	iDM
					Mod only
d.o.f.	72	[n/a]	47	[n/a]	68
$\sigma/10^{-41} \text{ cm}^2$	13.8	[8.9]	10.1	[8.2]	6.0
m_χ/GeV	7.2	[8.1]	7.7	[8.2]	10.0
δ/keV					24
$c_i/(\text{cpd/kg/keVee})$	2.5	[2.5]	2.5	[2.6]	$\begin{pmatrix} 1.5 \\ 0 \\ 1.5 \\ 2.3 \\ 2.3 \end{pmatrix} \quad \begin{pmatrix} -5.9 \\ -4.6 \\ -1.0 \\ 1.1 \\ 1.7 \end{pmatrix} \quad \begin{pmatrix} 8.7 \\ 0.4 \\ 0.4 \\ 1.2 \\ 1.5 \end{pmatrix}$
χ^2	57.3	[n/a]	50.8	[n/a]	53.7
					51.4
					51.3

DM fits

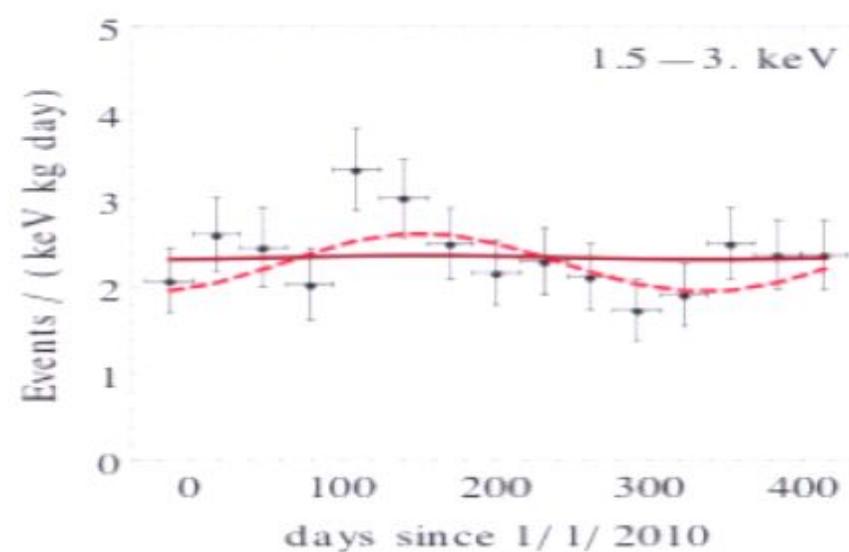
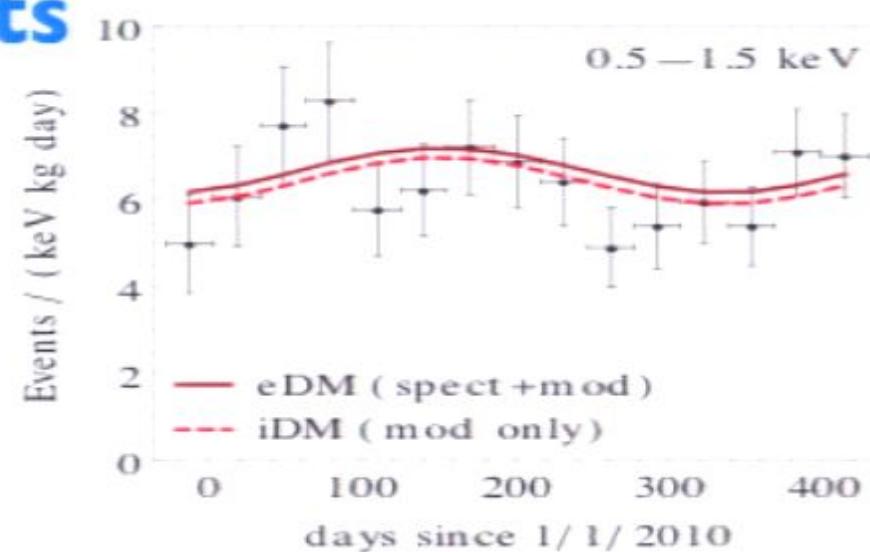
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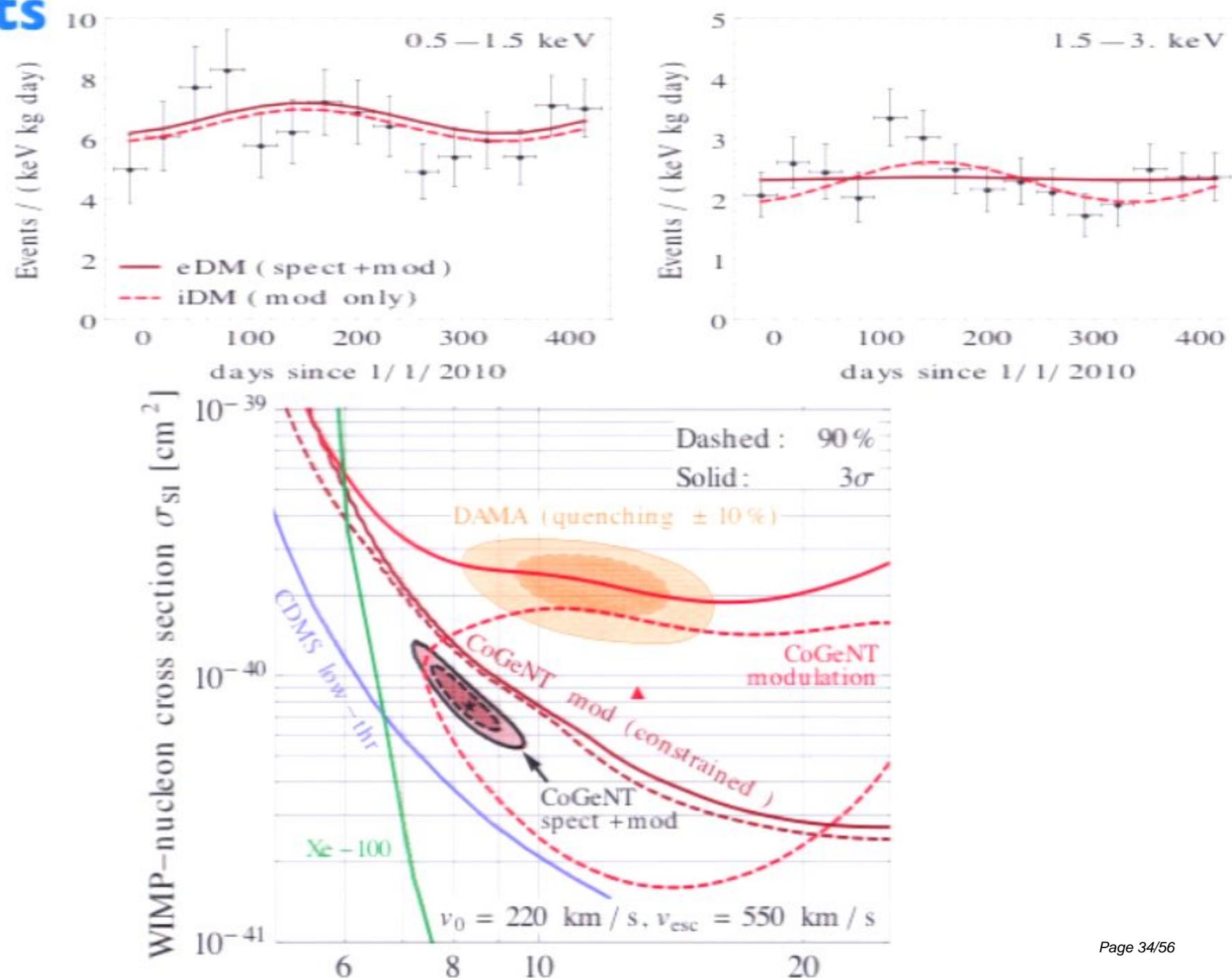
Null: $\chi^2 = 58.2$ for 70 degrees of freedom

Scenario	Spect+Mod	Spect only	Mod only ($c_i \geq 0$)	Mod only	iDM
					Mod only
d.o.f.	72	[n/a]	47	[n/a]	68
$\sigma/10^{-41} \text{ cm}^2$	13.8	[8.9]	10.1	[8.2]	6.0
m_χ/GeV	7.2	[8.1]	7.7	[8.2]	10.0
δ/keV					12.0
$c_i/(\text{cpd/kg/keVee})$	2.5	[2.5]	2.5	[2.6]	$\begin{pmatrix} 1.5 \\ 0 \\ 1.5 \\ 2.3 \\ 2.3 \end{pmatrix} \quad \begin{pmatrix} -5.9 \\ -4.6 \\ -1.0 \\ 1.1 \\ 1.7 \end{pmatrix} \quad \begin{pmatrix} 8.7 \\ 0.4 \\ 0.4 \\ 1.2 \\ 1.5 \end{pmatrix}$
χ^2	57.3	[n/a]	50.8	[n/a]	53.7
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					51.3

DM fits

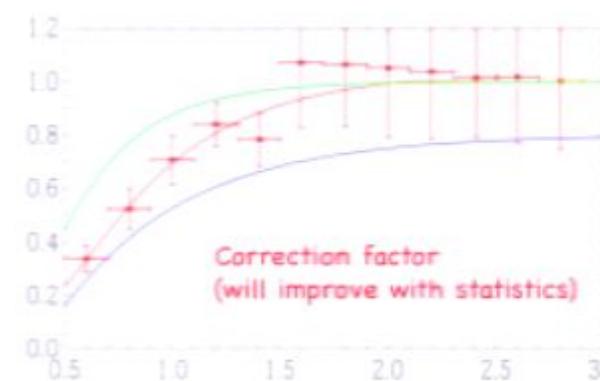


DM fits

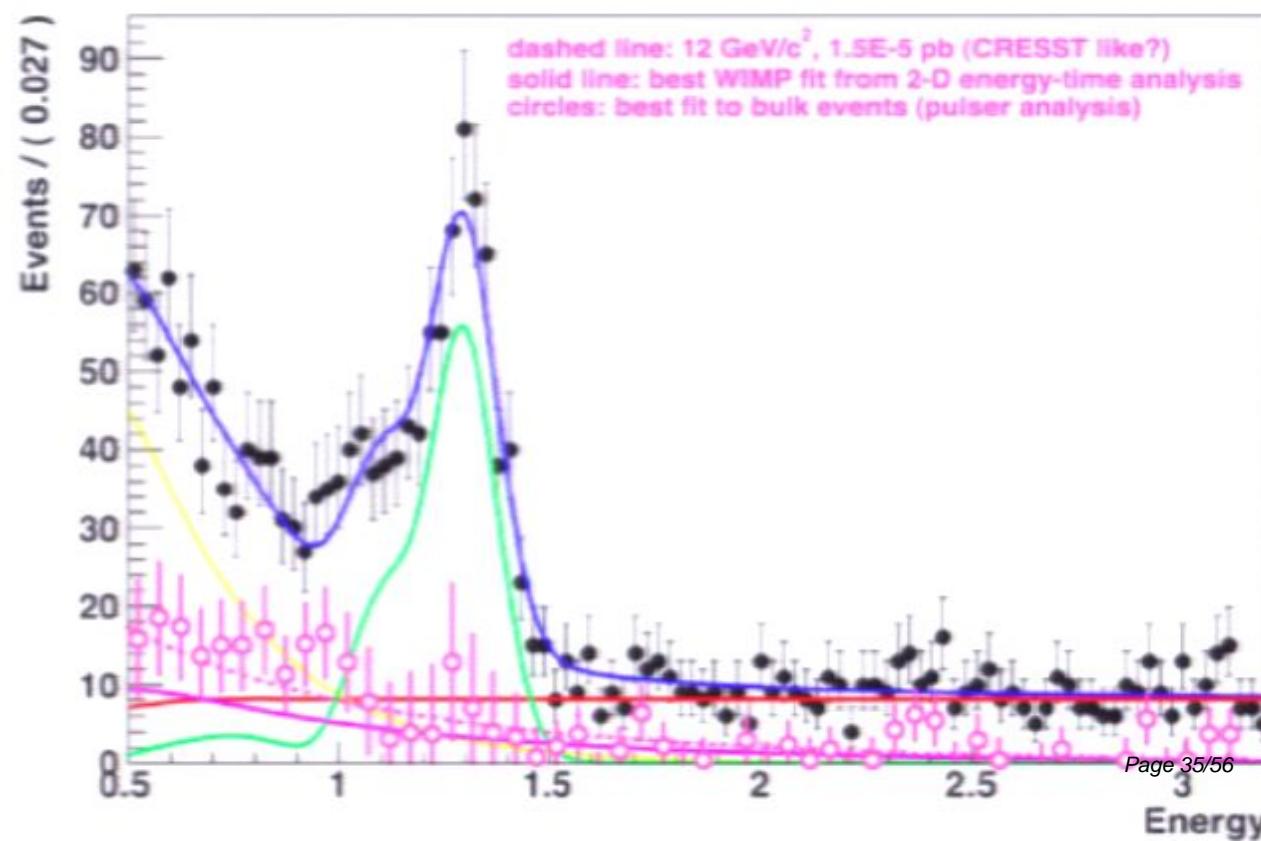


CoGeNT updates

Makes “tension” with MB even larger



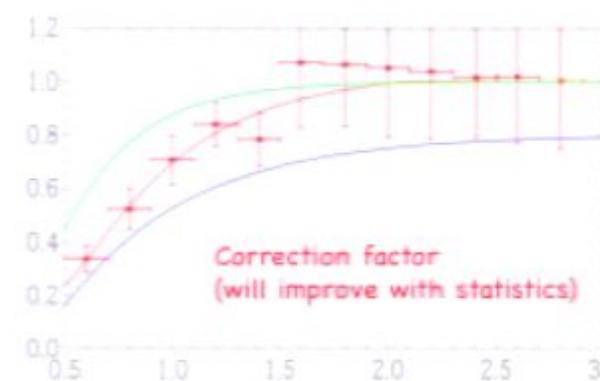
Data projected on energy PRELIMINARY (work in progress)



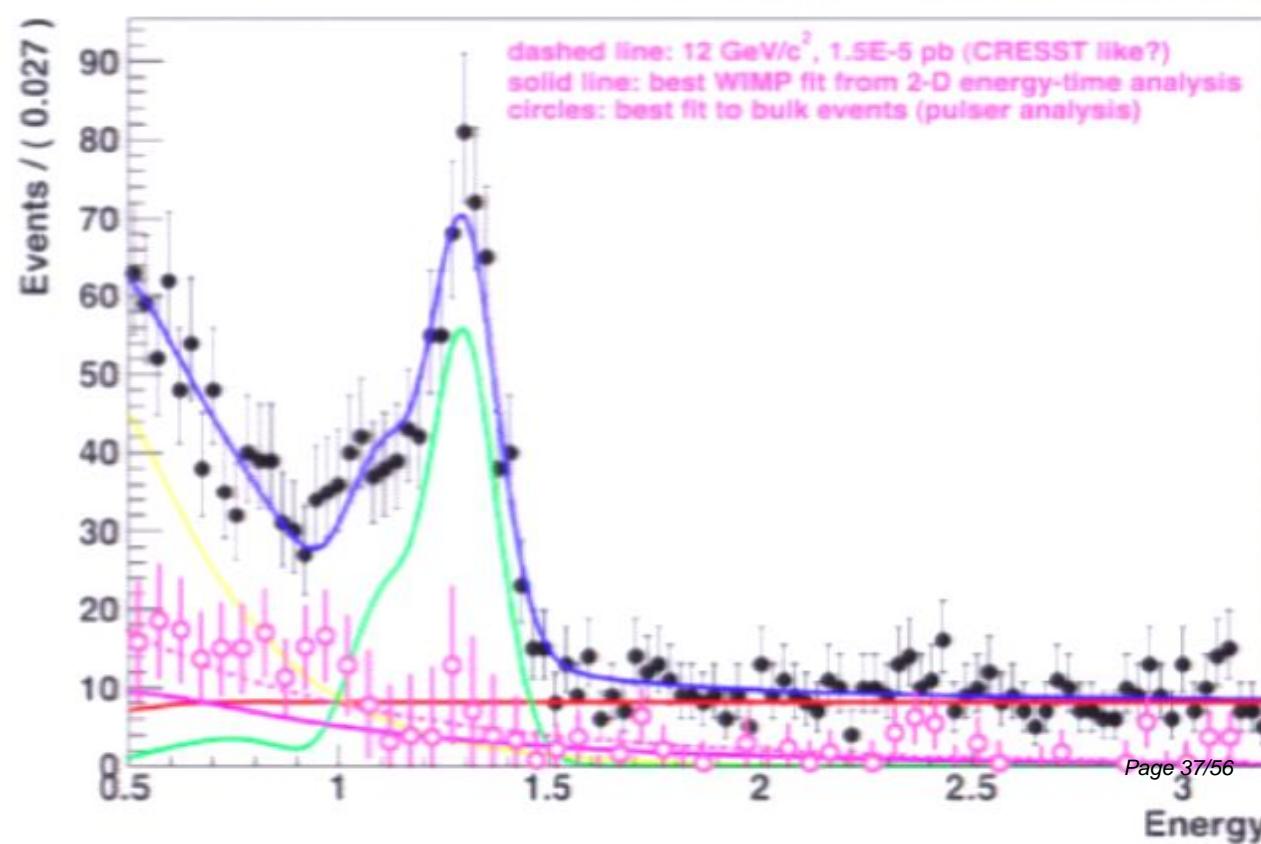
What about other halo models?

CoGeNT updates

Makes “tension” with MB even larger



Data projected on energy PRELIMINARY (work in progress)



What about other halo models?

Halo model independence

$$\frac{dR}{dE_R} = \frac{N_T M_T \rho}{2m_\chi \mu^2} \int_{v_{min}}^{v_{max}} d^3\vec{v} \frac{f(\vec{v}, \vec{v}_E)}{v} \sigma(E_R)$$

$\overbrace{\hspace{10em}}$
 $g(v)$

$$\frac{dR}{dE_R} = \frac{N_T M_T}{2\mu^2} \frac{\rho \sigma}{m_\chi} g(v)$$

Target Target
specific independent

$$v_{min} = \sqrt{\frac{M_T E_R}{2\mu^2}}$$

Recoil energy uniquely determines
minimum DM velocity

Halo model independence

$$N_T = \kappa N_A m_p / M_T$$

Solve for $g(v)$

$$g(v_{min}) = \frac{2m_\chi\mu^2}{N_A\kappa m_p \rho \sigma(E_R)} \frac{dR_1}{dE_1}$$

$$\frac{dR_1}{dE_1} \iff g(v_{min}) \iff \frac{dR_2}{dE_2}$$

The master formula (SI):

$$C_T^{(i)} = \kappa^{(i)} \left(f_p Z^{(i)} + f_n (A^{(i)} - Z^{(i)}) \right)^2$$

$$\frac{dR_2}{dE_R}(E_2) = \frac{C_T^{(2)}}{C_T^{(1)}} \frac{F_2^2(E_2)}{F_1^2 \left(\frac{\mu_1^2 M_T^{(2)}}{\mu_2^2 M_T^{(1)}} E_2 \right)} \frac{dR_1}{dE_R} \left(\frac{\mu_1^2 M_T^{(2)}}{\mu_2^2 M_T^{(1)}} E_2 \right)$$

Halo model independence

$$\frac{dR}{dE_R} = \frac{N_T M_T \rho}{2m_\chi \mu^2} \int_{v_{min}}^{v_{max}} d^3\vec{v} \frac{f(\vec{v}, \vec{v}_E)}{v} \sigma(E_R)$$

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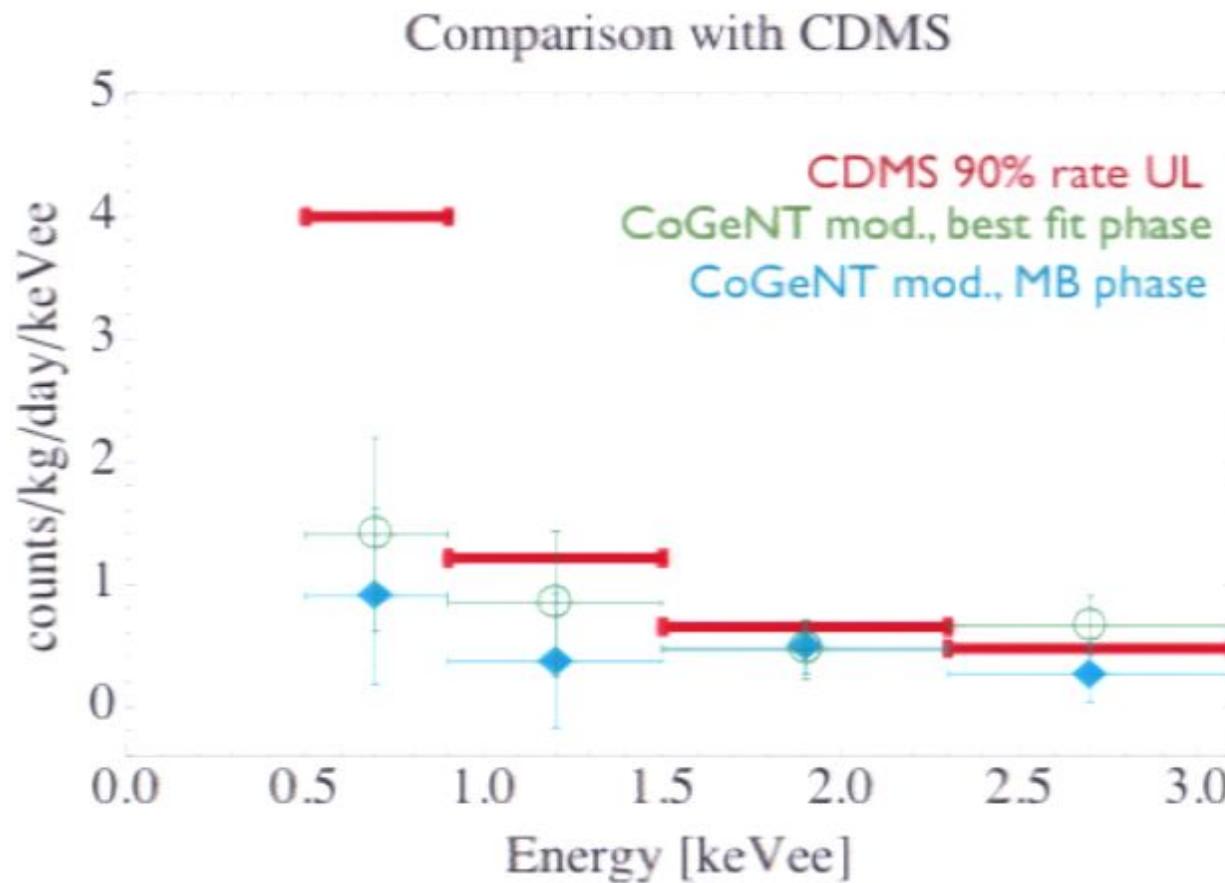
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CDMS (Ge) and CoGeNT

(easy one since both Ge)



Modulation should be large in CDMS

CRESST (O) and CoGeNT

CoGeNT CRESST (O)

$[0.5, 1.5] \rightarrow [5.8 - 15.6]$

Expect ~ 11 modulation events,
they see ~ 30

$[1.5, 3.1] \rightarrow [15.6, 29.8]$

$(m_\chi = 7 \text{ GeV})$

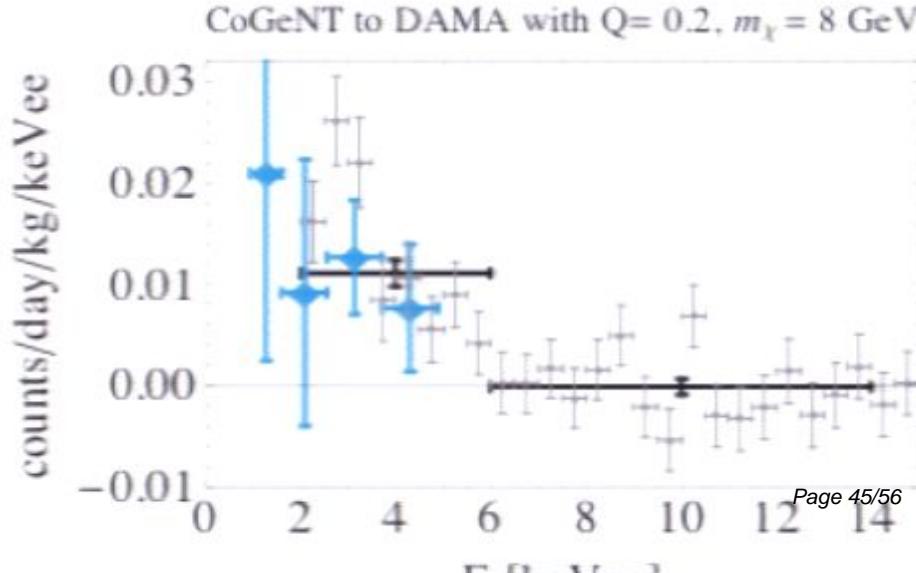
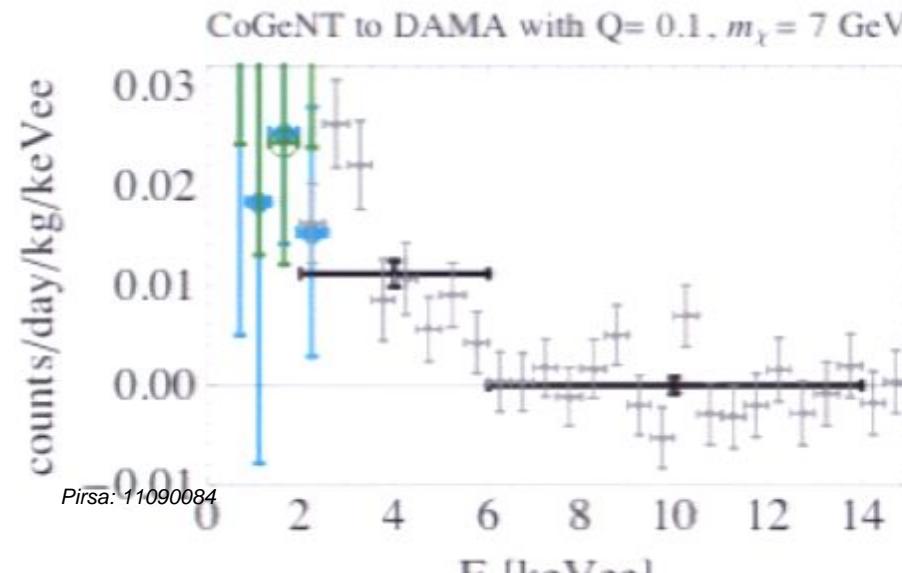
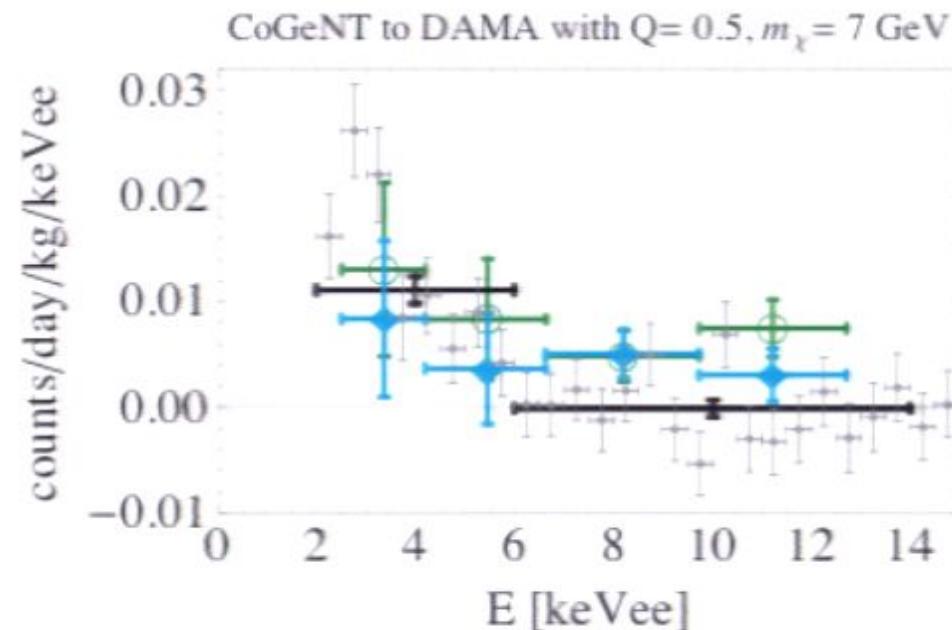
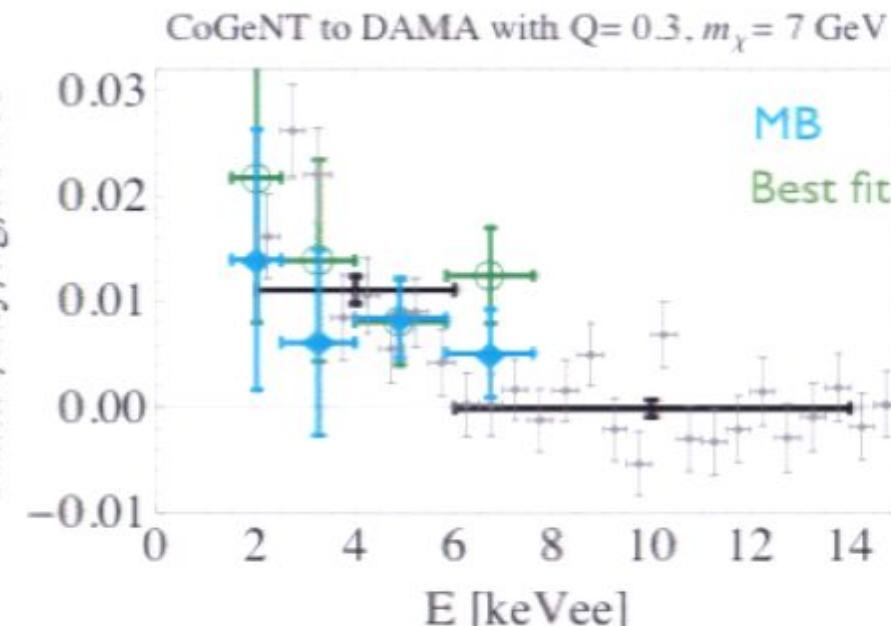
Xenon100 and CoGeNT

All about L_{eff} : $1.5 \text{ keVee} \rightarrow 3.7 \text{ keVnr}$

$$L_{eff} = 0.07 \Rightarrow (\sim 15, \sim 30)$$

Need $L_{eff} \lesssim 0.04$ to get event rate down

DAMA and CoGeNT



Conclusions-CoGeNT

If CoGeNT modulation is from DM it is not a MB halo
Should compare to other results in halo-independent way

Conclusions-global

Conclusions-CoGeNT

If CoGeNT modulation is from DM it is not a MB halo
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Conclusions-CoGeNT

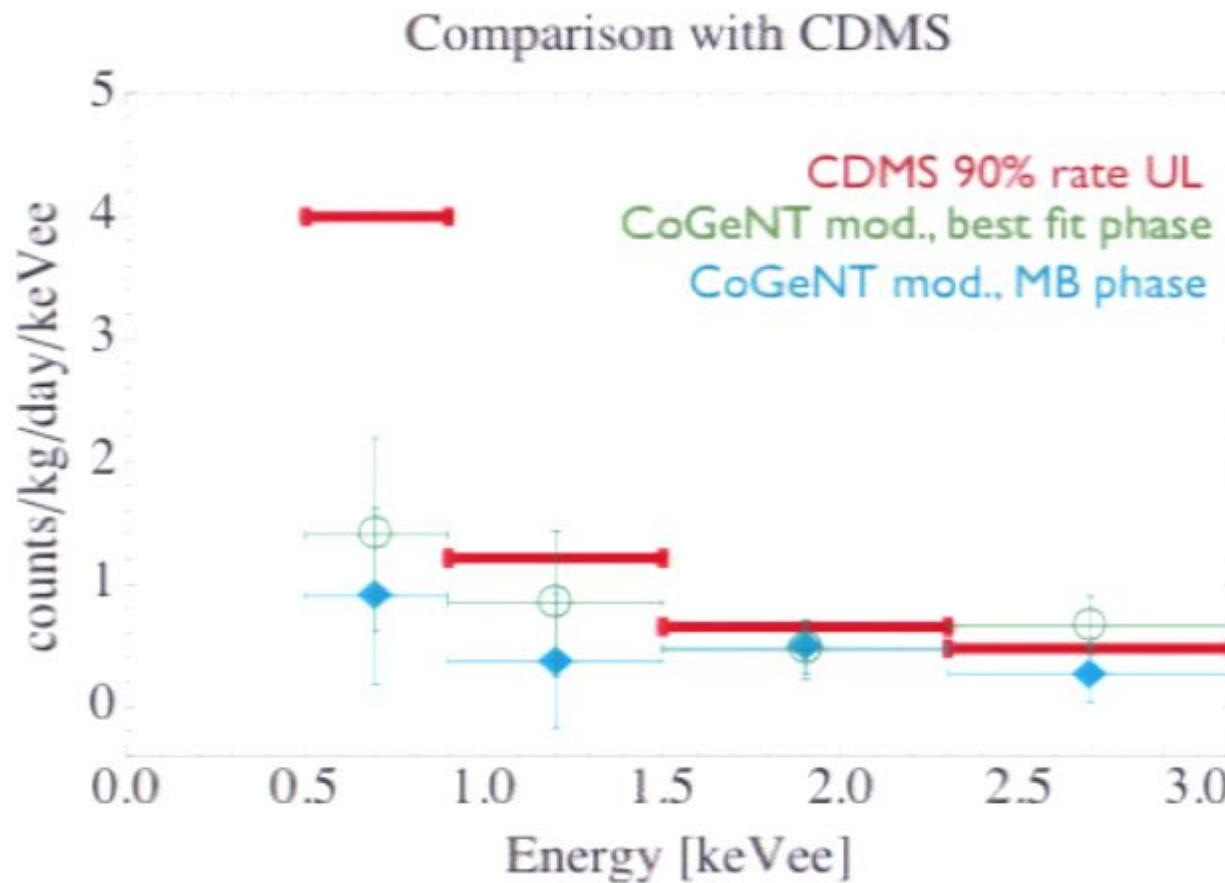
If CoGeNT modulation is from DM it is not a MB halo
Should compare to other results in halo-independent way

Conclusions-global



CDMS (Ge) and CoGeNT

(easy one since both Ge)



Modulation should be large in CDMS

DM fits

$$f(v) \propto (e^{-v^2/v_0^2} - e^{-v_{esc}^2/v_0^2}) \Theta(v_{esc} - v)$$

$$v_{\text{esc}} = 550 \text{ km/s} \quad v_0 = 220 \text{ km/s}$$

Scenario	Spect+Mod	Spect only	Mod only ($c_i \geq 0$)	Mod only	iDM
					Mod only
d.o.f.	72	[n/a]	47	[n/a]	68
$\sigma/10^{-41} \text{ cm}^2$	13.8	[8.9]	10.1	[8.2]	6.0
m_χ/GeV	7.2	[8.1]	7.7	[8.2]	10.0
δ/keV					24
$c_i/(\text{cpd/kg/keVee})$	2.5	[2.5]	2.5	[2.6]	$\begin{pmatrix} 1.5 \\ 0 \\ 1.5 \\ 2.3 \\ 2.3 \end{pmatrix} \quad \begin{pmatrix} -5.9 \\ -4.6 \\ -1.0 \\ 1.1 \\ 1.7 \end{pmatrix} \quad \begin{pmatrix} 8.7 \\ 0.4 \\ 0.4 \\ 1.2 \\ 1.5 \end{pmatrix}$
χ^2	57.3	[n/a]	50.8	[n/a]	53.7
					51.4
					51.3

DM fits

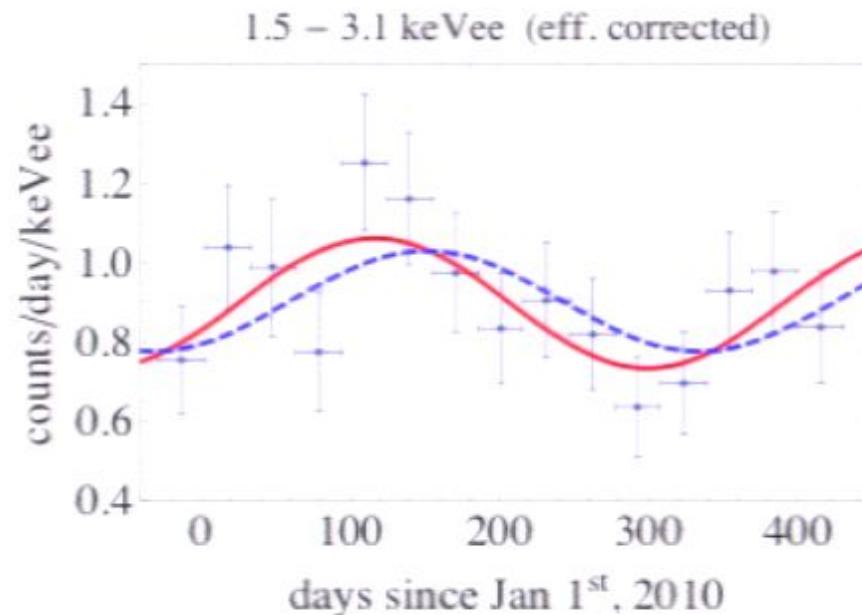
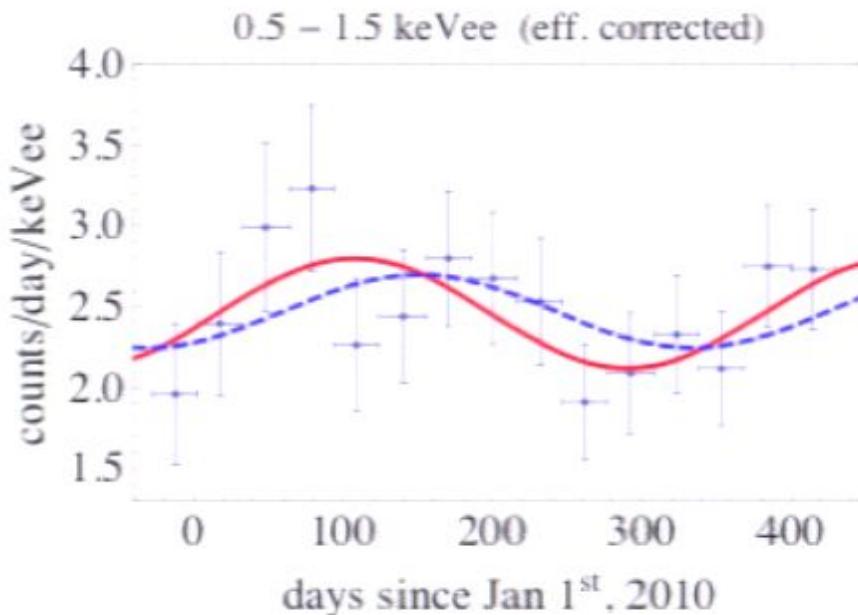
$$f(v) \propto (e^{-v^2/v_0^2} - e^{-v_{esc}^2/v_0^2}) \Theta(v_{esc} - v)$$

$$v_{\text{esc}} = 550 \text{ km/s} \quad v_0 = 220 \text{ km/s}$$

Null: $\chi^2 = 58.2$ for 70 degrees of freedom

Scenario	Spect+Mod	Spect only	Mod only ($c_i \geq 0$)	Mod only	iDM
					Mod only
d.o.f.	72	[n/a]	47	[n/a]	68
$\sigma/10^{-41} \text{ cm}^2$	13.8	[8.9]	10.1	[8.2]	6.0
m_χ/GeV	7.2	[8.1]	7.7	[8.2]	10.0
δ/keV					12.0
$c_i/(\text{cpd/kg/keVee})$	2.5	[2.5]	2.5	[2.6]	$\begin{pmatrix} 1.5 \\ 0 \\ 1.5 \\ 2.3 \\ 2.3 \end{pmatrix} \quad \begin{pmatrix} -5.9 \\ -4.6 \\ -1.0 \\ 1.1 \\ 1.7 \end{pmatrix} \quad \begin{pmatrix} 8.7 \\ 0.4 \\ 0.4 \\ 1.2 \\ 1.5 \end{pmatrix}$
χ^2	57.3	[n/a]	50.8	[n/a]	53.7
					51.4
					51.3

Modulation amplitude (Binned)



Free phase

$$A_0 = 7.4 \text{ events/day/kg/keVee}$$

$$A_1 = 0.14 \quad t_0 = 107 \text{ days}$$

$$\Delta\chi^2 = 4.7$$

$$A_0 = 2.7 \text{ events/day/kg/keVee}$$

$$A_1 = 0.18 \quad t_0 = 116 \text{ days}$$

$$\Delta\chi^2 = 8.2$$

MB phase

$$A_0 = 7.5 \text{ events/day/kg/keVee}$$

$$A_1 = 0.09$$

$$\Delta\chi^2 = 2.3$$

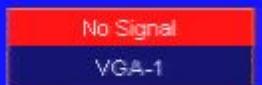
$$A_0 = 2.7 \text{ events/day/kg/keVee}$$

$$A_1 = 0.14$$

$$\Delta\chi^2 = 5.2$$

No Signal

VGA-1





No Signal
VGA-1