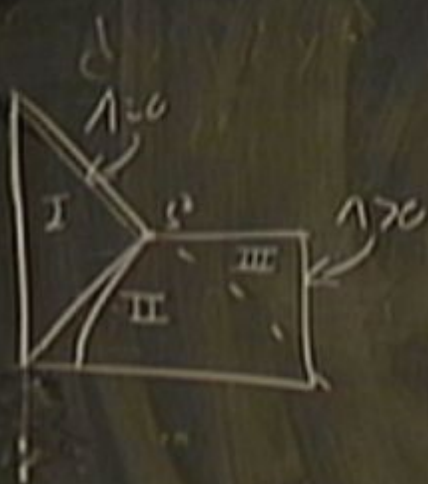


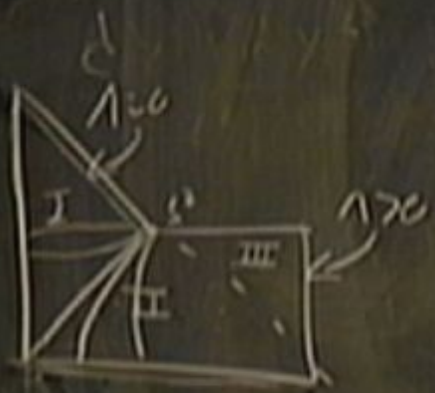
Title: Liouville theory and Holographic Cosmology

Date: Jun 23, 2011 04:10 PM

URL: <http://pirsa.org/11060050>

Abstract: I will explain how Liouville theory with complex values of its parameters arises naturally in speculative holographic cosmologies. We will encounter Liouville theory of both the ``spacelike" and ``timelike" variety. I will then use this as motivation to present some new results on the analytic continuation of Liouville theory recently obtained with Maltz and Witten.





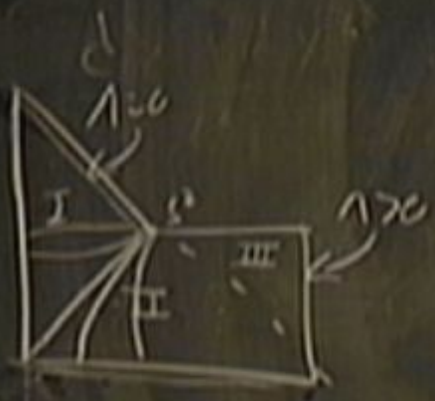
Time like

dS_3

$$\lim_{\epsilon \rightarrow 0} \Psi_{\text{NAW}} \left(\frac{\vec{x}}{\epsilon}, \vec{\phi} \right) \beta = \text{CFT}$$

= some $-\text{Scal} + \dots$
 (M, \mathbb{R})

$$c = \frac{\lambda_{\text{AdS}}}{\ell_p}$$



Time like

dS_3

$$\lim_{\epsilon \rightarrow 0} \Psi_{\text{NAW}} \left(\frac{\vec{x}}{\epsilon}, \vec{q} \right) \beta = \text{CFT}$$

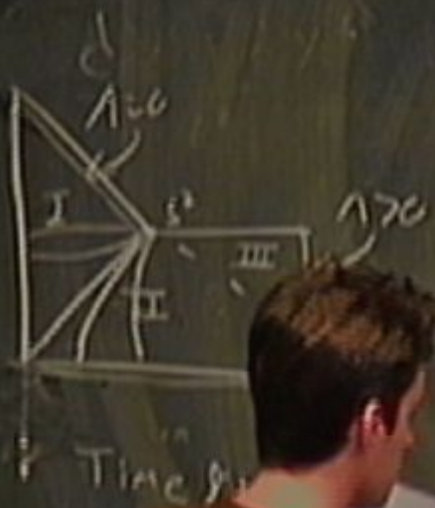
= some $e^{-S_{\text{CFT}}(\vec{x}, \vec{q})}$

$$C_{\text{NS}} = \frac{\lambda_{\text{NS}}}{G}$$

$$\lambda_{\text{NS}} = i \lambda_{\text{S}}$$

$$C_{\text{S}} = \frac{i \lambda_{\text{S}}}{G}$$

$$\Psi_{\text{NAW}} \sim e^{-i \int \frac{1}{\epsilon} \text{AdS} - \frac{\lambda_{\text{NS}}}{G} \mathbb{I} + \dots}$$



dS_3

$$\lim_{\epsilon \rightarrow 0} \Psi_{NAW} \left(\frac{\vec{x}}{\epsilon}, \vec{q} \right) \stackrel{\text{def}}{=} \text{CET}$$

= some $e^{-\text{Scat}} \text{ of } \dots$

$$C_{\text{ret}} = \frac{\lambda_{\text{NAW}}}{G}$$

$$\lambda_{\text{NAW}} = i \lambda_{\text{US}}$$

$$C_{\text{adv}} = \frac{i \lambda_{\text{US}}}{G}$$

$$\Psi_{\text{ret}} = e^{-i \int \frac{1}{\epsilon} \int d^3x \sqrt{g} - \frac{\lambda_{\text{US}}}{G} \int \dots}$$

$$T_{\mu\nu} = \int d^3\sigma d^3\sigma' \sqrt{-g} \sqrt{-g'}, R(\sigma) R(\sigma') \zeta_\gamma(\sigma, \sigma')$$

$$\int d^4x d^4y \epsilon(x, y) \psi^{\rightarrow} \psi$$

$$T_{00} = \int d^3\sigma d^3\sigma' \sqrt{-g} \sqrt{-g'} R(\sigma) R(\sigma') \zeta(\sigma, \sigma')$$

$$\Delta = \frac{1}{2} + \frac{1}{2} \sqrt{1 - 4\eta^2}$$

$$\int d^4\tilde{x} d^4\tilde{y} f(\tilde{x}, \tilde{y}) \psi^{\rightarrow} \psi^{\leftarrow}$$

$$C_{TC} = 0 + O(\eta)$$

$$S_1 = \frac{1}{4\pi} \int d^3x \left[(\dot{\phi})^2 + q \vec{R} \phi + 4\pi e^{2\phi} \right]$$

$$\langle e^{2\omega_1 \phi} \dots e^{2\omega_n \phi} \rangle_{S^2} = \int D\phi e^{2\omega_1 \phi} \dots e^{2\omega_n \phi} e^{-S_1}$$

$$S_1 = \frac{1}{4\pi} \int d^3x \left[(\dot{\phi})^2 + \vec{\nabla} \phi \cdot \vec{\nabla} \phi + 4m^2 \phi^2 \right]$$

$$\langle e^{i\int d^3x \phi(x) J(x)} \rangle = \int \mathcal{D}\phi e^{i\int d^3x \left[\mathcal{L}(\phi) + \phi J \right]}$$



$$S_1 = \frac{1}{4\pi} \int d^3x \left[(\dot{\phi})^2 + Q \tilde{R} \phi + 4\pi n e^{2\alpha\phi} \right]$$

$$\left\langle e^{2\alpha_1\phi} \dots e^{2\alpha_n\phi} \right\rangle_{S^2} = \int D\phi e^{2\alpha_1\phi} \dots e^{2\alpha_n\phi} e^{-S_1}$$

$$\sum \alpha_i > Q$$

$$\sum \alpha_i < Q$$

$$\delta\left(\int d^2\sigma \epsilon^{\alpha\beta} - A\right)$$

d

$$S_1 = \frac{1}{4\pi} \int d^3x \left[(\partial_t \phi)^2 + Q \vec{\nabla} \phi + \mu^2 \phi^2 \right]$$

$$\langle e^{2\alpha_1 \phi} \dots e^{2\alpha_n \phi} \rangle_S = \int D\phi e^{2\alpha_1 \phi} \dots e^{2\alpha_n \phi} e^{-S_1}$$



$$\sum \alpha_i > Q$$

$$\sum \alpha_i < Q$$

$$\delta \left(\int d^3x \phi - A \right)$$

$$|C| = \sum ds$$

$$Q^2 \approx |C|$$

$$\sum \alpha = Q$$

$$\alpha \approx \frac{D}{Q}$$

$$\sum \alpha^2 \approx Q^2 \approx \sum ds$$