

Title: Dark matter models with uniquely spin-dependent detection

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URL: <http://pirsa.org/11050012>

Abstract: It is often assumed that the first evidence for direct dark matter detection will come from experiments probing spin-independent interactions, because of higher sensitivities due to coherence effects. We explore the possibility of models that would be invisible in such experiments, but detectable via spin-dependent interactions. The existence of much larger (or only) spin-dependent tree-level interactions is not sufficient, due to potential spin-independent subdominant or loop-induced interactions. We find that most models with detectable spin-dependent interactions would also generate detectable spin-independent interactions. Models in which a light pseudoscalar acts as the mediator seem to uniquely evade this conclusion. We present a viable dark matter model generating such an interaction.

DM with only spin-dependent detection possibilities

Zoltan Ligeti

Perimeter Institute, May 6, 2011

- Introduction, direct detection experiments
- Spin-dependent vs. spin-independent [Freytsis and ZL, 1012.5317]
- An unusual DM search in $B \rightarrow K^{(*)} \ell^+ \ell^-$ [Freytsis, ZL, Thaler, 0911.5355]
- Conclusions

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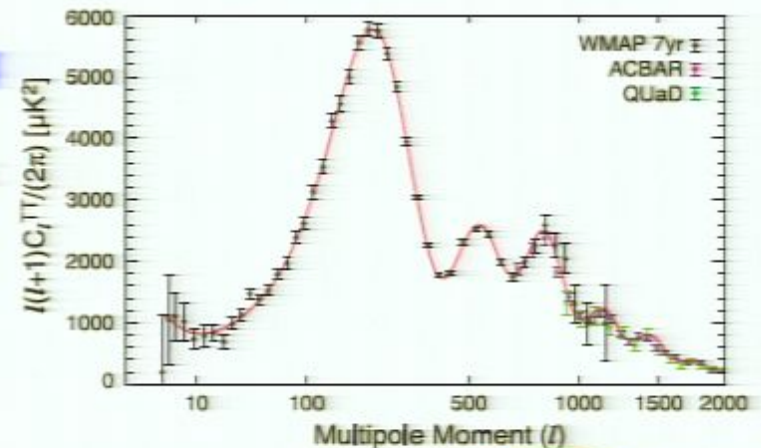
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What is dark matter?

- Homogeneous, isotropic, spatially flat, expanding

Basic paradigm is well-established, looking for corrections:

- Seek deviations from $w = -1$
- Order 10–20 % corrections allowed

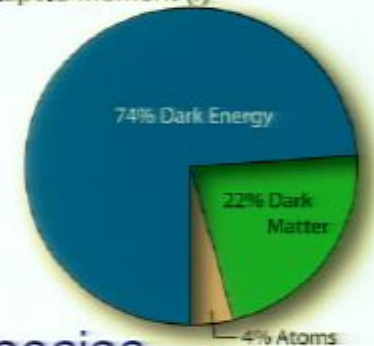


- Dark matter: rotation curves, gravitational lensing, cosmology

- We know that DM is not a particle we know:

Know: non-baryonic, long lived, neutral, abundance

Don't know: interactions, mass, quantum numbers, one/many species



- Maybe thermal relic of early universe: weakly interacting massive particle (WIMP)

If so, WIMP mass has to be around the TeV scale — LHC may directly produce it



Direct DM detection



- All evidence is gravitational — no unambiguous direct detection signals yet
Steady increase in experimental sensitivity, and will continue for a while
- The focus is often on spin-independent (SI) detection experiments, due to coherence effects giving an A^2 enhancement and (much) larger nuclear cross sections
- Uncertainties in nuclear matrix elements is substantial — I'll not talk about them
Recent LQCD results imply a few times smaller SI cross section than often used [e.g., DarkSUSY]
- Experiments capable of detecting SI cross section down to the irreducible atmospheric neutrino background, 10^{-48} cm^2 , seem possible in the foreseeable future

[Note (since I can't remember): $10^{-36} \text{ cm}^2 = 1 \text{ pb}$]



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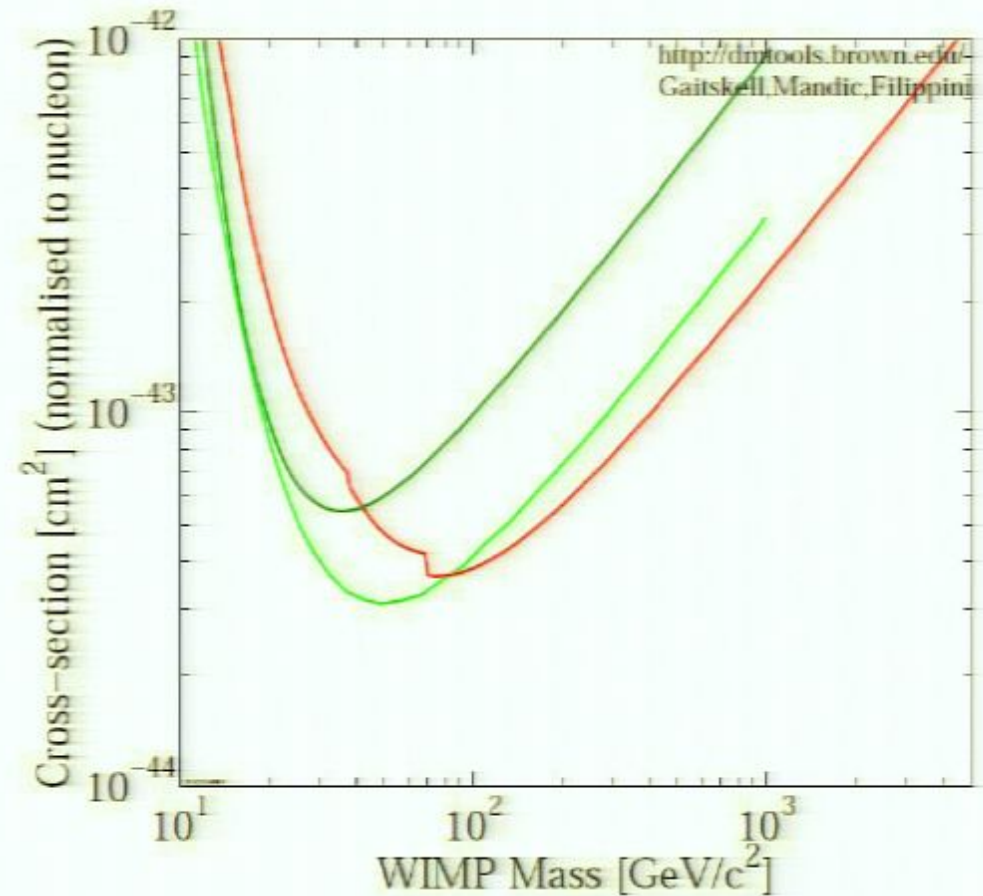


Current bounds — SI



- Cross sections normalized to per-nucleon value for comparison between experiments

Many other experiments not shown



DATA listed top to bottom on plot
— XENON10 2007, measured L_{eff} from Xe cube
— CDMS: Soudan 2004–2009 Ge
— XENON100 SI 161 kg days



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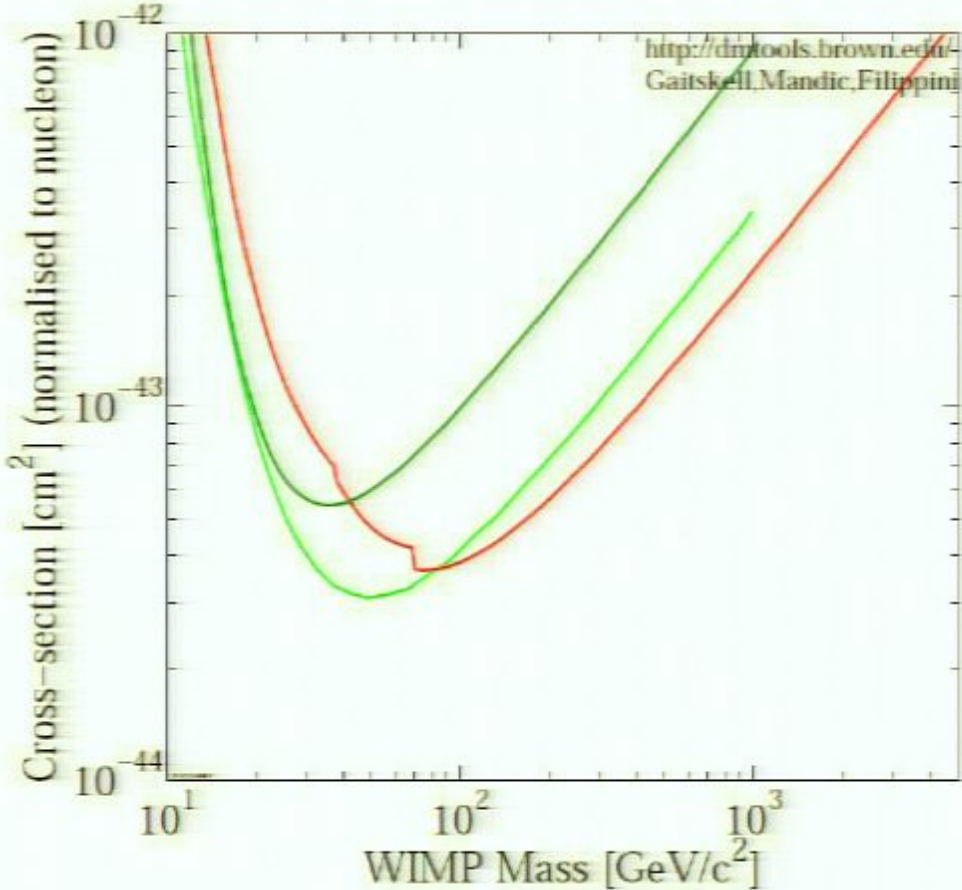


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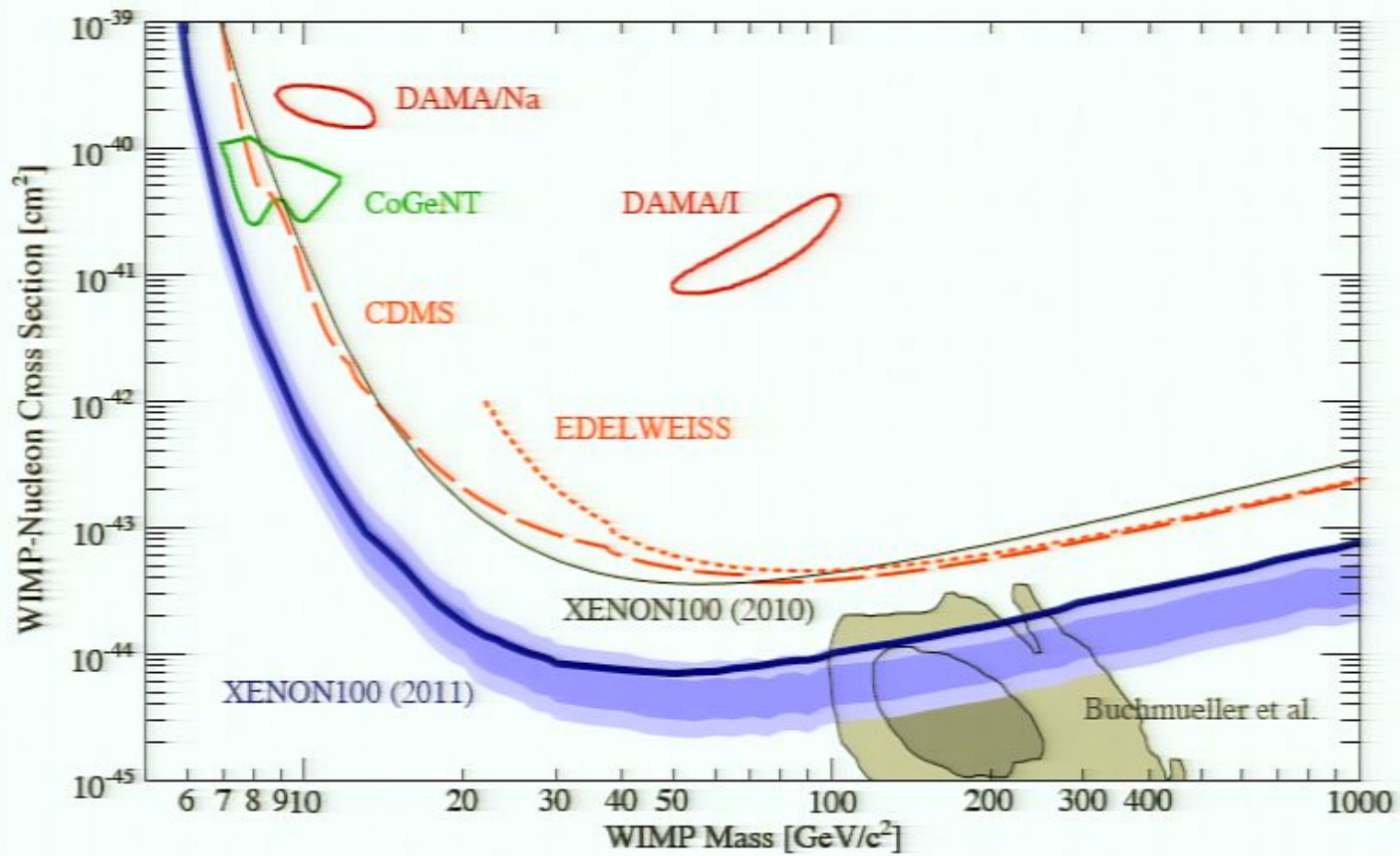


<http://dmtools.brown.edu/>
Gaitskell, Mandic, Filipini

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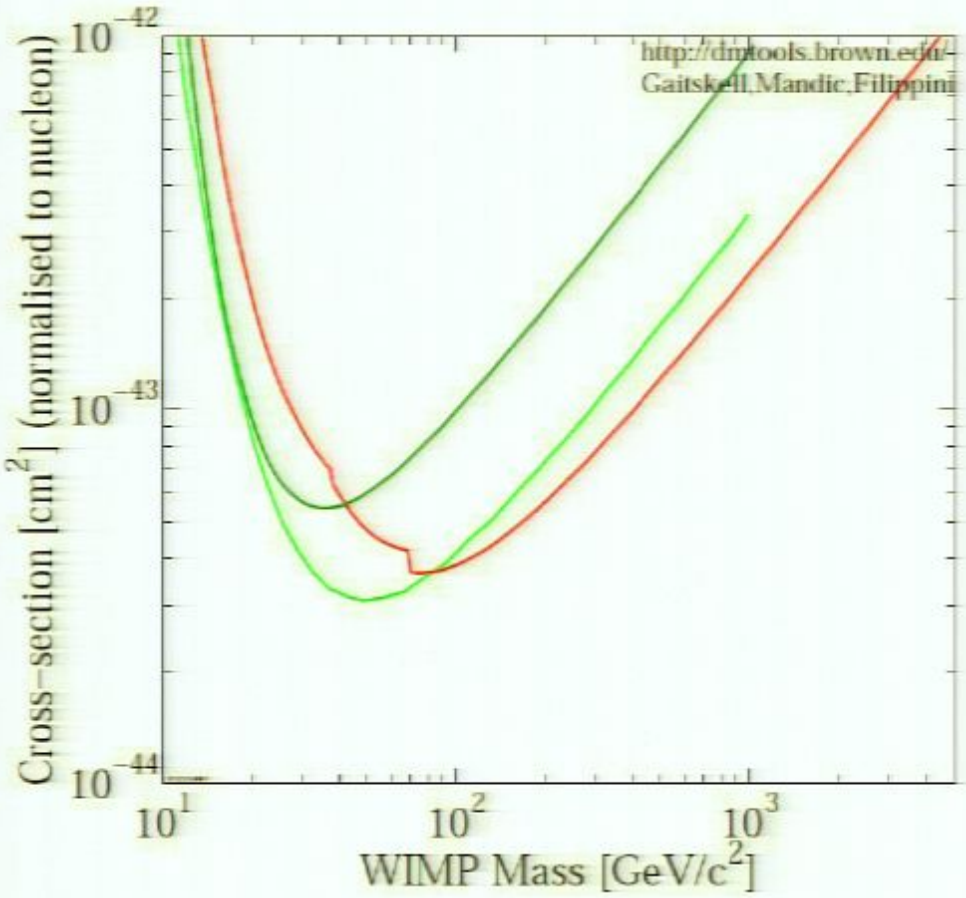
XENON 100, just last month



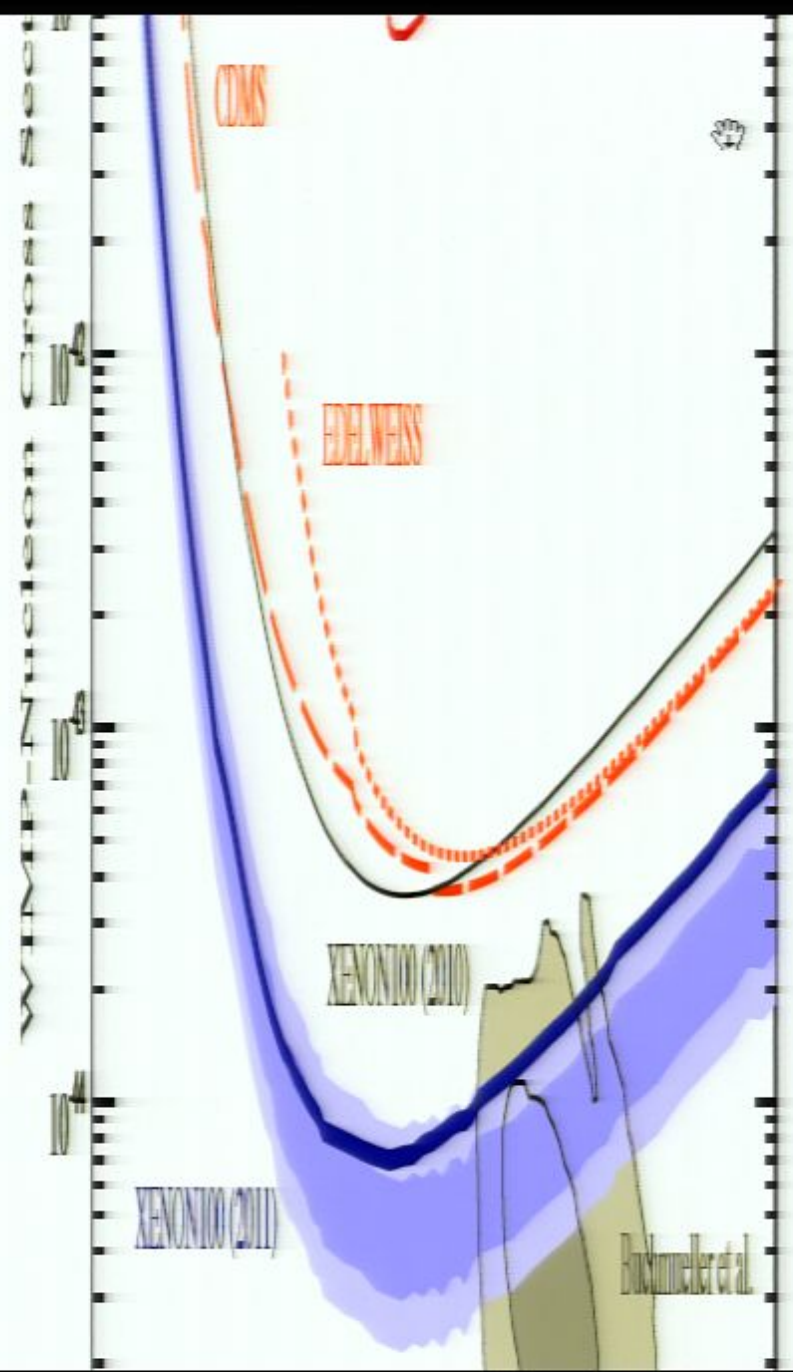
[arXiv:1104.2549]



The comparison



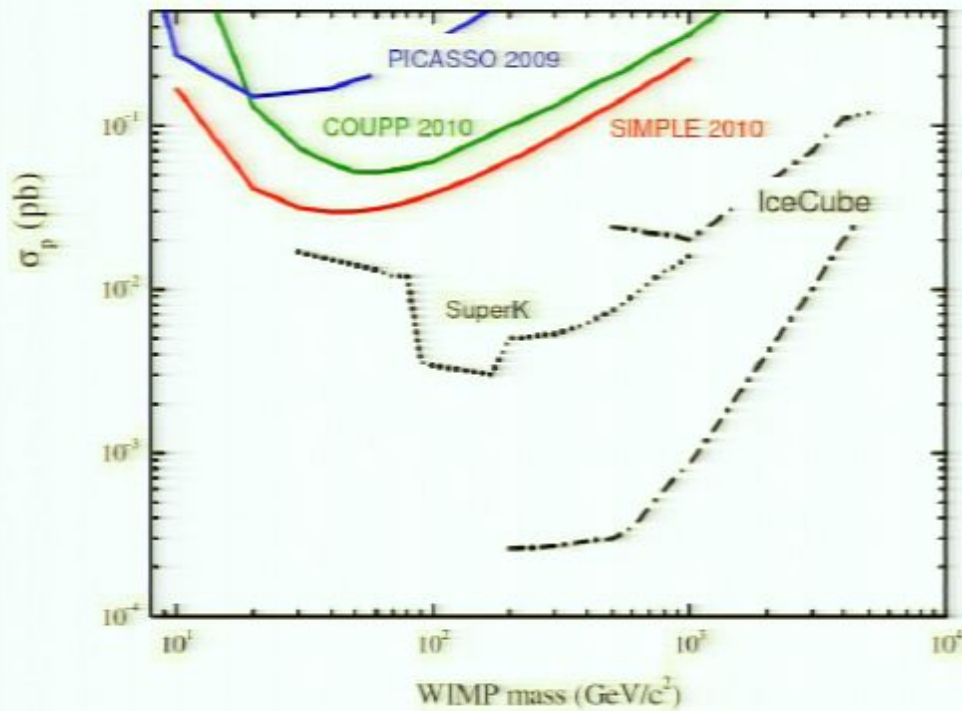
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Current bounds — SD

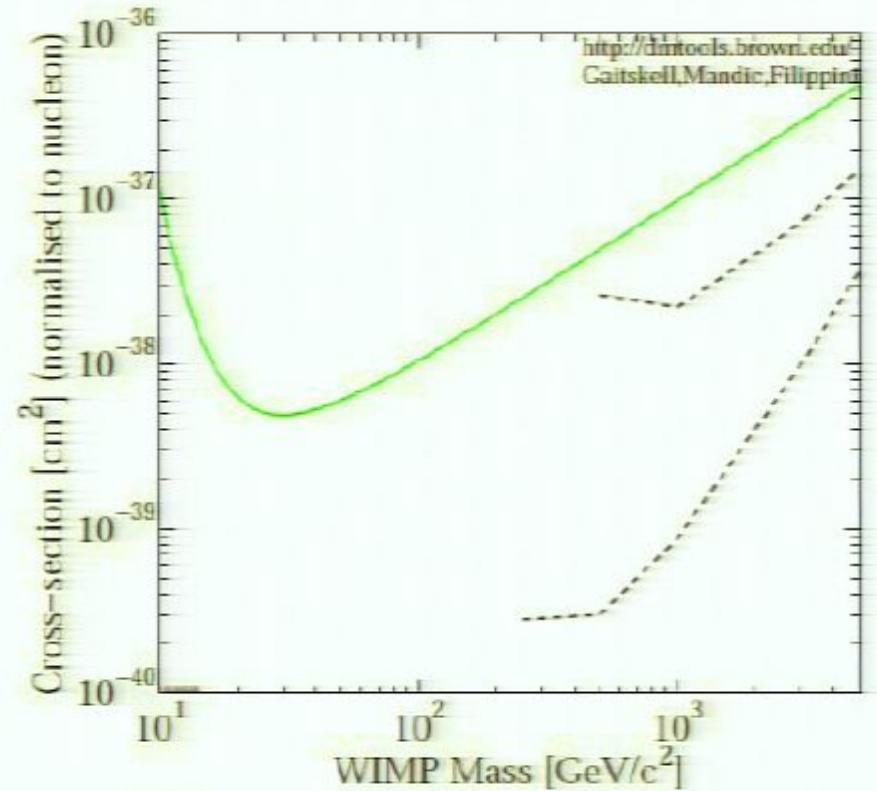


DM–proton cross section



[Girard et al., 1101.1885]

DM–neutron cross section

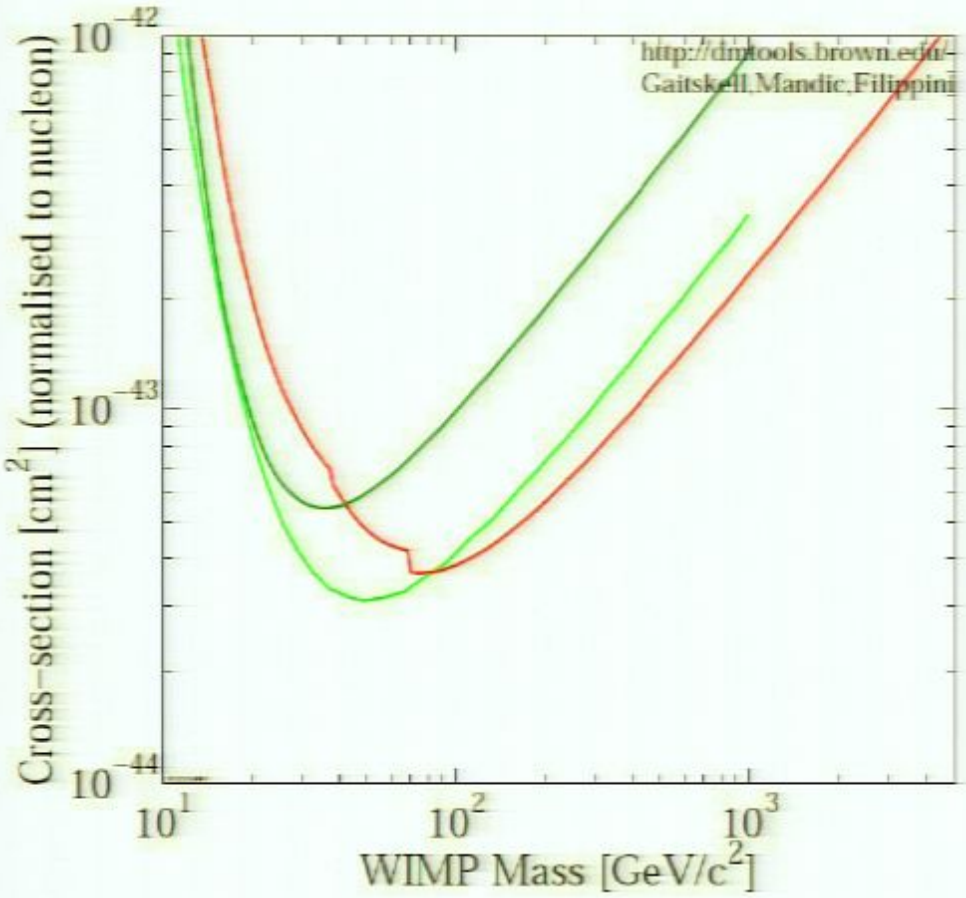


DATA listed top to bottom on plot:
 IceCube 2009 indirect SD–proton (assuming annihilation to $b\text{-}\bar{b}$)
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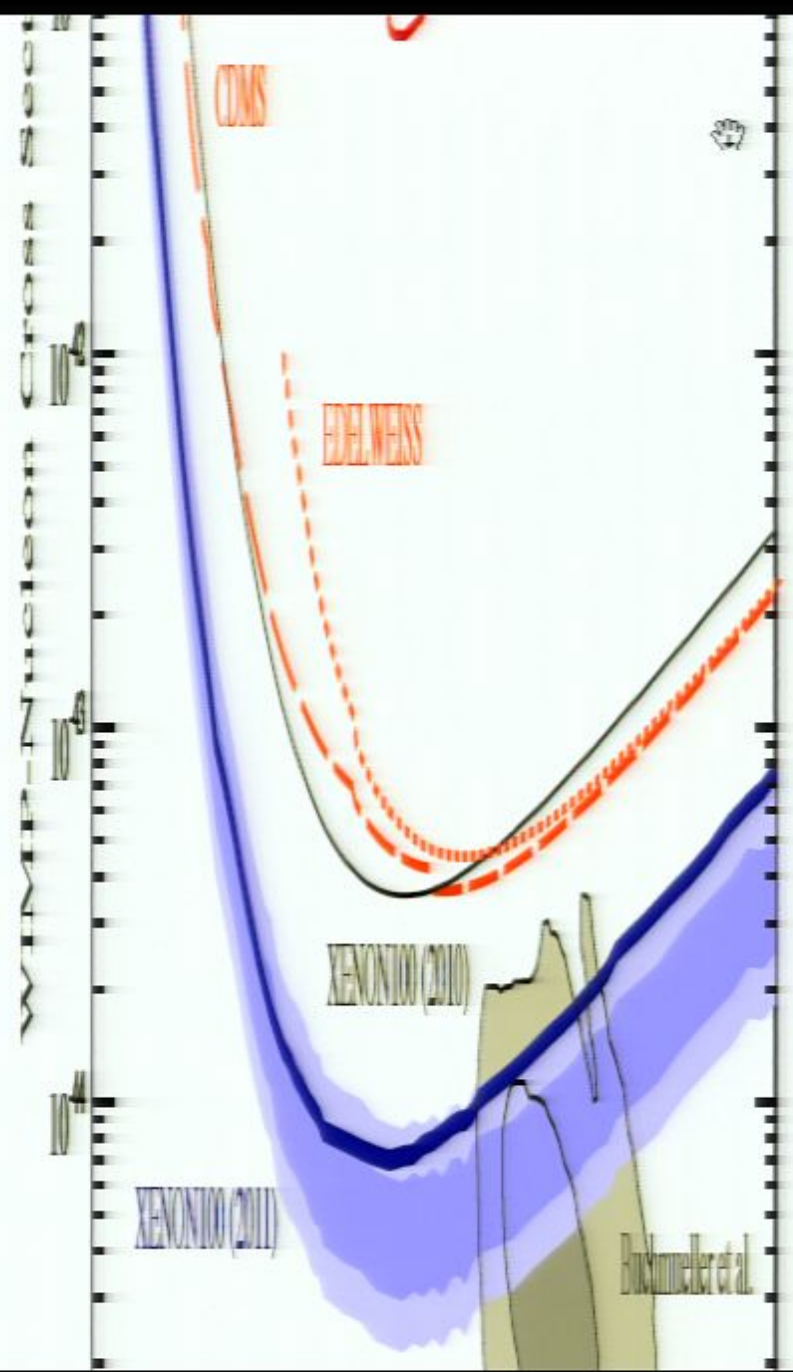
SD bounds are 5–6 orders of magnitude weaker than SI



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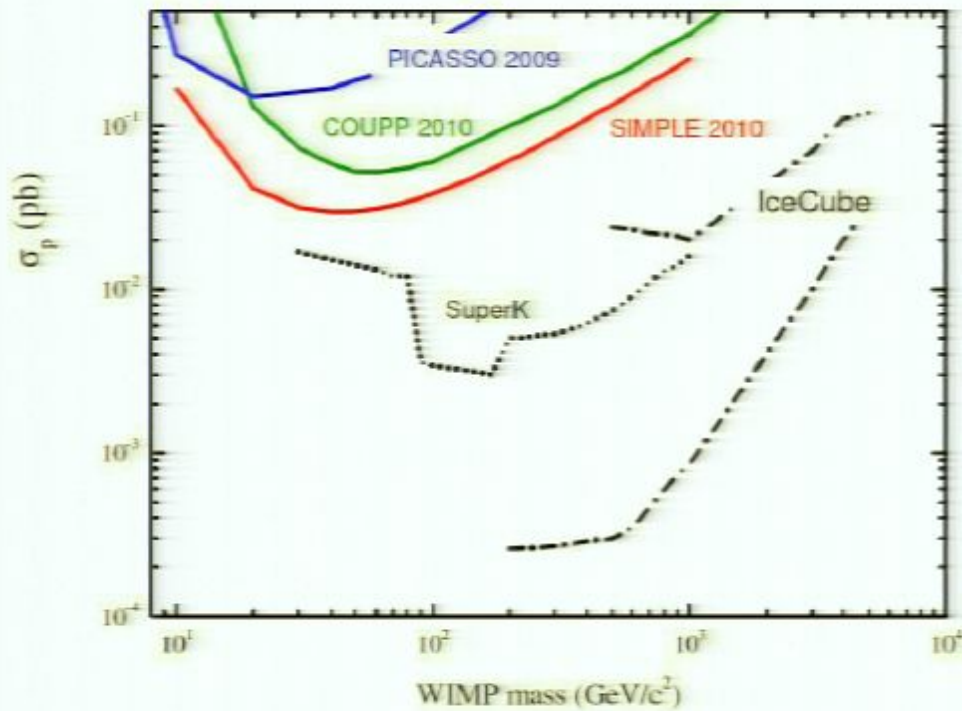
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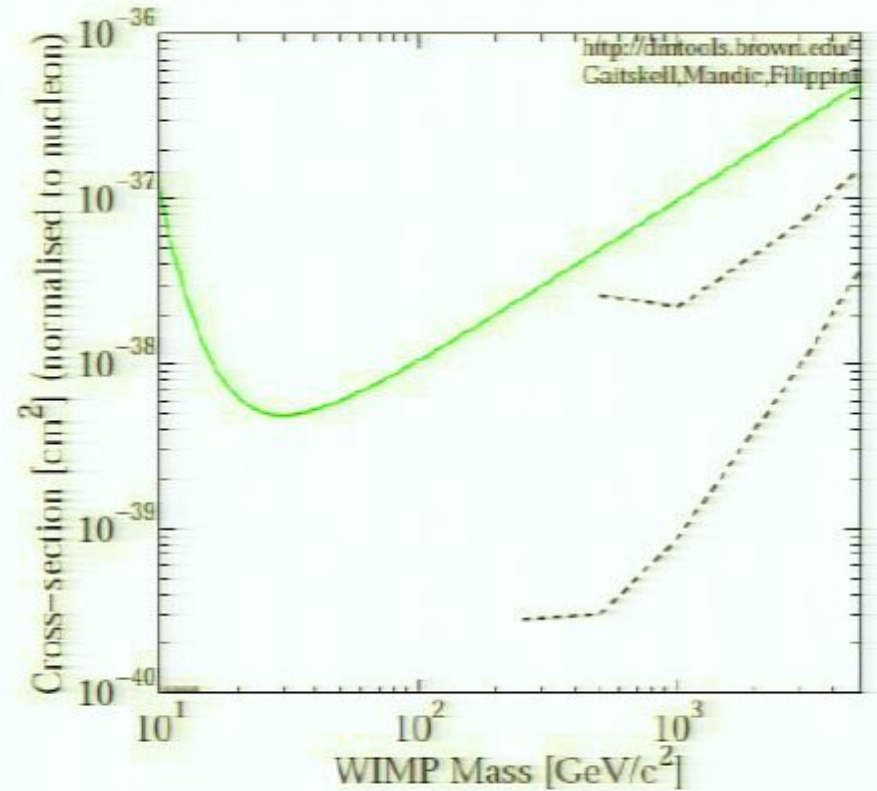


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Future sensitivity

Near future:

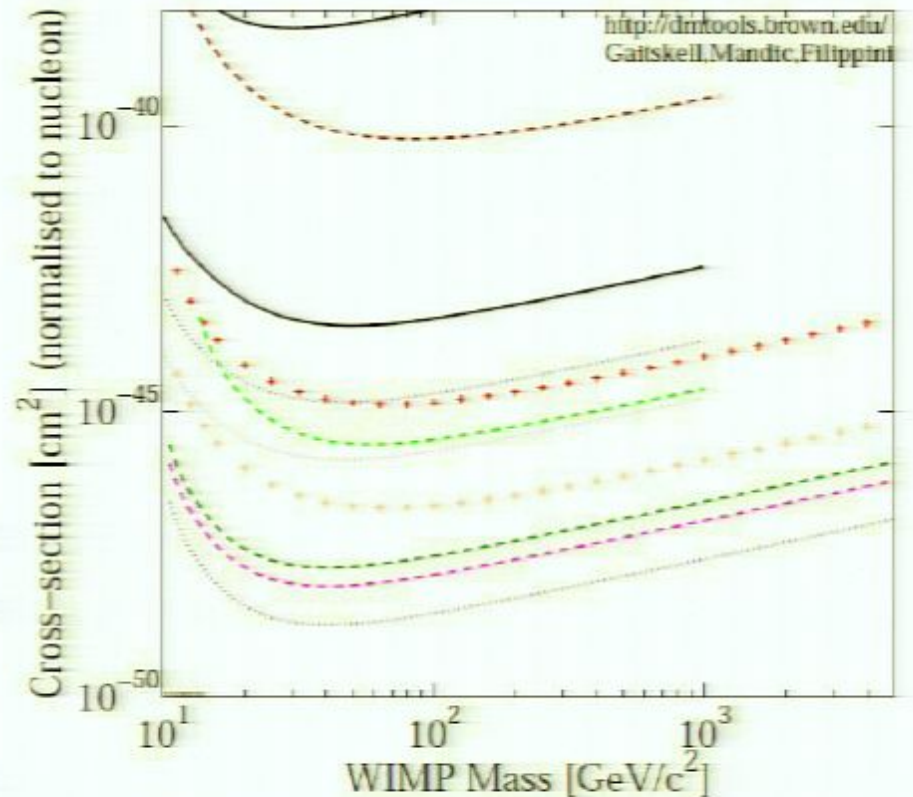
XENON 100 will improve a factor of few

SD experiments (not shown) to $\sim 10^{-40}$

Next decade (or two):

SI experiments will maintain (maybe even increase) their greater reach, down to the irreducible neutrino background

SD sensitivity (COUPP-500) $< 10^{-42} \text{ cm}^2$



Why bother with spin-dependent?



- Reasons to think about spin-dependent detection appearing on its own:

- The optimist:**

Could the SD dark matter cross section tell us something interesting?

Significant correlation between SI & SD cross sections is claimed in most standard frameworks (MSSM, UED, little Higgs) — observing the contrary would tell us we are on to something unexpected [Bertrone *et al.*, 0705.2502; Belanger *et al.*, 0810.1362; Cohen *et al.*, 1001.3408]

- The pessimist:**

What if no sign of SI dark matter interactions down to the neutrino background?

Do we give up on direct detection? Could there still be something to detect?

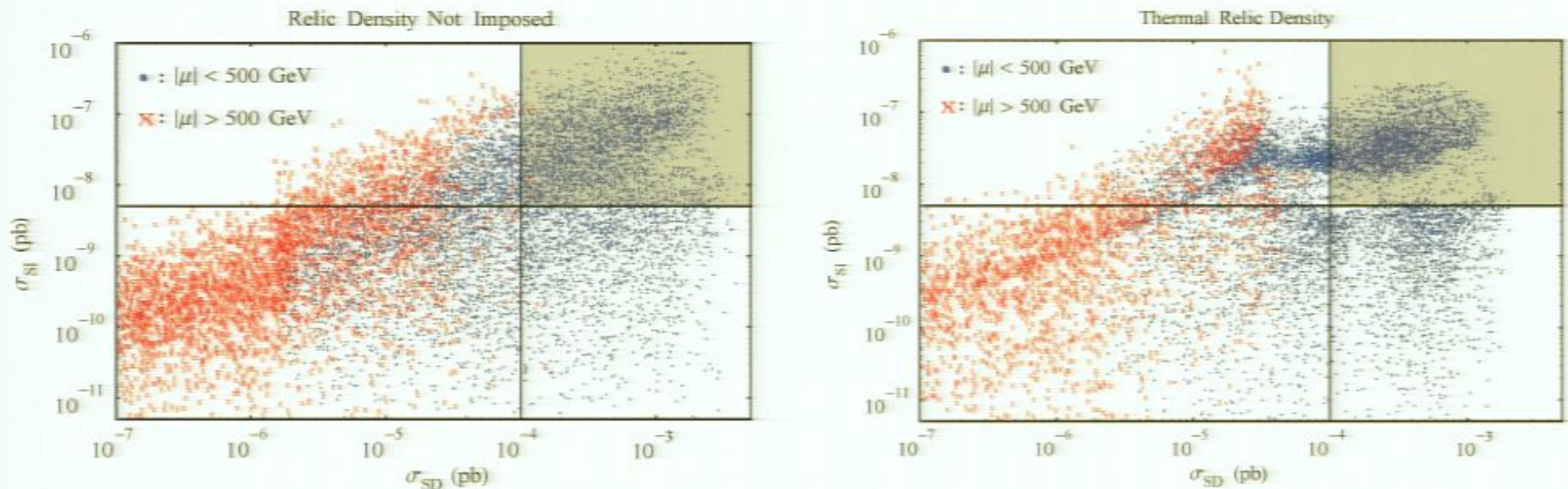
- Can spin-dependent win?**



E.g.: SUSY parameter space



- Knowing m_Z and m_h in the MSSM implies correlation between SI and SD signals



[Cohen, Phalen, Pierce, 1001.3408]

- In most of the parameter space $\sigma_{SI} > 10^{-5} \sigma_{SD}$
- Can select parameter values where huge cancellations occur (highly non-generic)



The challenge of uniquely SD detection



- If the tree-level interactions are dominantly (exclusively) SD, they may still be detectable by other means

Since the sensitivity to SI cross sections is higher by several orders of magnitude (and will stay like that), a signal will only be seen in SD experiments if there are:

- No detectable **kinematically suppressed** SI interactions at tree level
- No detectable **loop-induced** SI contributions

Either of these could be just as easy to detect as the “dominant” SD interaction

- **Operators:** which ones lead to SD scattering in the non-relativistic limit?
- **Mediators:** are there models w/ enhanced SD operators, while **subleading effects** are absent or highly suppressed?

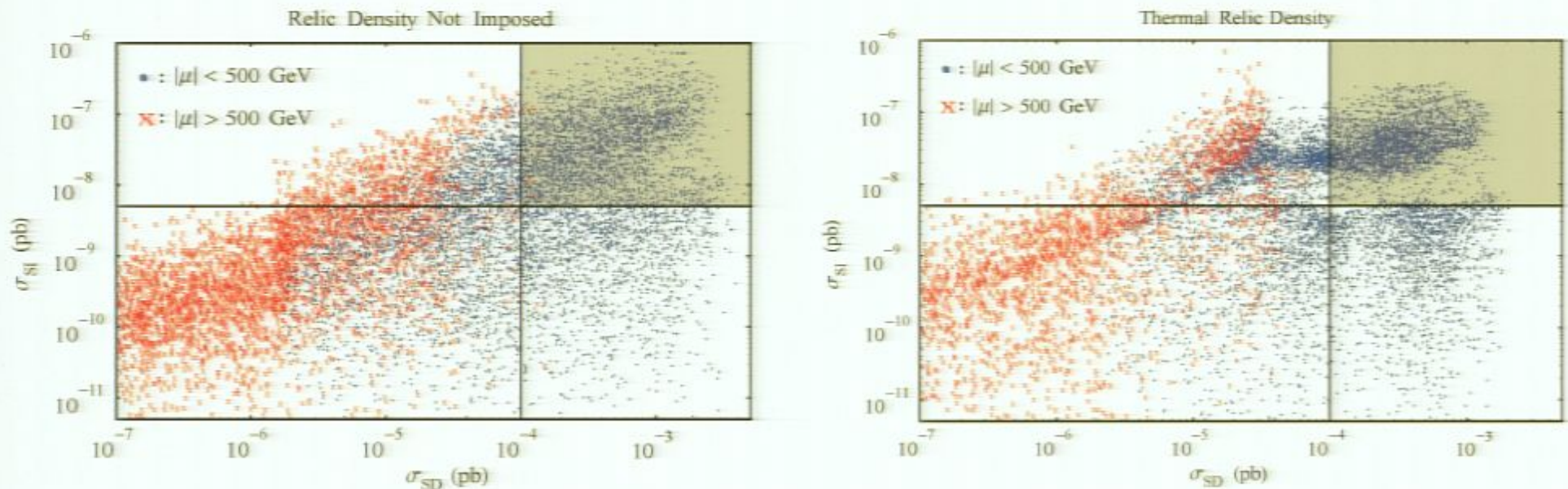
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Operator analysis (1)



Scalar DM:

Operator	SI / SD	Suppression
$\mathcal{O}_1^s = \phi^2 \bar{q}q$	SI	—
$\mathcal{O}_2^s = \phi^2 \bar{q}\gamma^5 q$	SD	q^2
$\mathcal{O}_3^s = \phi^\dagger \partial^\mu \phi \bar{q}\gamma_\mu q$	SI	—
$\mathcal{O}_4^s = \phi^\dagger \partial^\mu \phi \bar{q}\gamma_\mu \gamma^5 q$	SD	v^2

\mathcal{O}_3^s and \mathcal{O}_4^s are only present for complex scalar DM



Operator analysis (2)



• Fermion DM:

	Operator	SI / SD	Suppression
$\mathcal{O}_1^f =$	$\bar{\chi}\chi \bar{q}q$	SI	—
$\mathcal{O}_2^f =$	$\bar{\chi}i\gamma^5\chi \bar{q}q$	SI	q^2
$\mathcal{O}_3^f =$	$\bar{\chi}\chi \bar{q}i\gamma^5q$	SD	q^2
$\mathcal{O}_4^f =$	$\bar{\chi}\gamma^5\chi \bar{q}\gamma^5q$	SD	q^4
$\mathcal{O}_5^f =$	$\bar{\chi}\gamma^\mu\chi \bar{q}\gamma_\mu q$	SI	—
$\mathcal{O}_6^f =$	$\bar{\chi}\gamma^\mu\gamma^5\chi \bar{q}\gamma_\mu q$	SI SD	v^2 q^2
$\mathcal{O}_7^f =$	$\bar{\chi}\gamma^\mu\chi \bar{q}\gamma_\mu\gamma^5q$	SD	v^2 or q^2
$\mathcal{O}_8^f =$	$\bar{\chi}\gamma^\mu\gamma^5\chi \bar{q}\gamma_\mu\gamma^5q$	SD	—
$\mathcal{O}_9^f =$	$\bar{\chi}\sigma^{\mu\nu}\chi \bar{q}\sigma_{\mu\nu}q$	SD	—
$\mathcal{O}_{10}^f =$	$\bar{\chi}i\sigma^{\mu\nu}\gamma^5\chi \bar{q}\sigma_{\mu\nu}q$	SI	q^2

If the DM is Majorana fermion, \mathcal{O}_5^f , \mathcal{O}_7^f , \mathcal{O}_9^f , \mathcal{O}_{10}^f vanish identically (odd under C)



Operator analysis (3)



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$\mathcal{O}_7^v =$	$\epsilon_{\mu\nu\rho\sigma} B^\mu \partial^\nu B^\rho \bar{q}\gamma^\sigma q$	SI	v^2
		SD	q^2
$\mathcal{O}_8^v =$	$\epsilon_{\mu\nu\rho\sigma} B^\mu \partial^\nu B^\rho \bar{q}\gamma^\sigma \gamma^5 q$	SD	—

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Viabile operators



- Traditionally, disregard kinematically suppressed operators — this leaves for SD:

$$\mathcal{O}_8^f = \bar{\chi} \gamma_\mu \gamma^5 \chi \bar{q} \gamma^\mu \gamma^5 q$$

$$\mathcal{O}_9^f = \bar{\chi} \sigma_{\mu\nu} \chi \bar{q} \sigma^{\mu\nu} q$$

$$\mathcal{O}_8^v = \epsilon_{\mu\nu\rho\sigma} B^\mu \partial^\nu B^\rho \bar{q} \gamma^\sigma \gamma^5 q$$

- A kinematically suppressed operator can be competitive if the mediator scale is low and unsuppressed operators are absent [Chang et al., 0908.3192]

Both features are present in the case of spontaneous global symmetry breaking

- Need to consider suppressed operators with pseudoscalar currents:

$$\mathcal{O}_2^s = \phi^2 \bar{q} \gamma^5 q$$

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Operators from mediators



- Assume: operators from integrating out mediators w/ renormalizable interactions

What interactions yield the SD operators and nothing else?

This was studied for heavy mediators

[Agrawal *et al.*, 1003.5905]

Include the possibility of light mediators



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Fermion DM with SD scattering



Mediator	Process	Comments
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X		Majorana + chiral coupling
ϕ		Majorana + chiral coupling



Boson DM with SD scattering



Mediator	Process	Comments
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a		
Q		Real + chiral coupling



Fermion DM with SD scattering



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Kinematically suppressed contributions



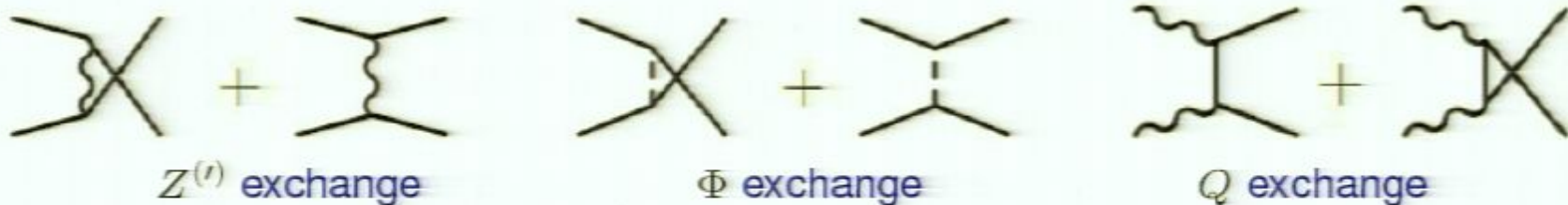
- Usually only the leading contribution from each operator is retained

Kinematic suppressions: $v^2 \sim 10^{-6}$

$$q^2/m_p m_\chi \sim 10^{-6} \text{ for } q^2 \sim (100 \text{ MeV})^2$$

This is comparable to (SD sensitivity) / (SI sensitivity) — if suppressed SI contributions exist, they should lead to detectable effects at roughly the same time

- Typically get v^2 suppressed SI operators:



- Pseudoscalar a exchange — no additional contributions generated



Loop-induced subleading effects

- Consider, e.g., Z exchange at tree level:

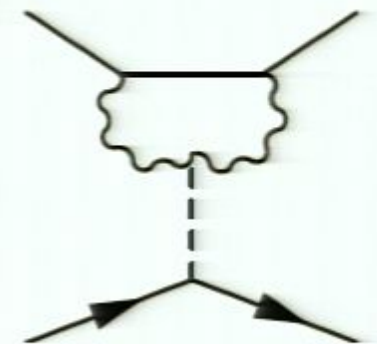
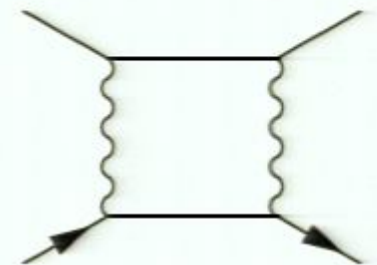
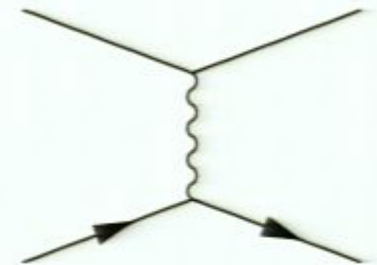
$$\frac{g_2^2}{2 \cos^2 \theta_W} T_3^q Q \frac{1}{m_Z^2} \bar{\chi} \gamma^\mu \gamma^5 \chi \bar{q} \gamma_\mu \gamma^5 q$$

- Loop effects induce the SI operator:

$$\frac{1}{4\pi} \frac{g_2^4 Q^2}{\cos^4 \theta_W m_Z} \left[\frac{(T_3^q)^2}{2m_Z^2} + \frac{1}{4m_h^2} \right] m_q \bar{\chi} \chi \bar{q} q \quad (m_q \ll m_Z \ll m_\chi)$$

For $Q \sim \mathcal{O}(1)$, the SI cross section is 6–7 orders of magnitude smaller than the SD one

However, if the DM is in an $SU(2)$ n-plet, the cross section $\sim n^2$



Kinematically suppressed contributions



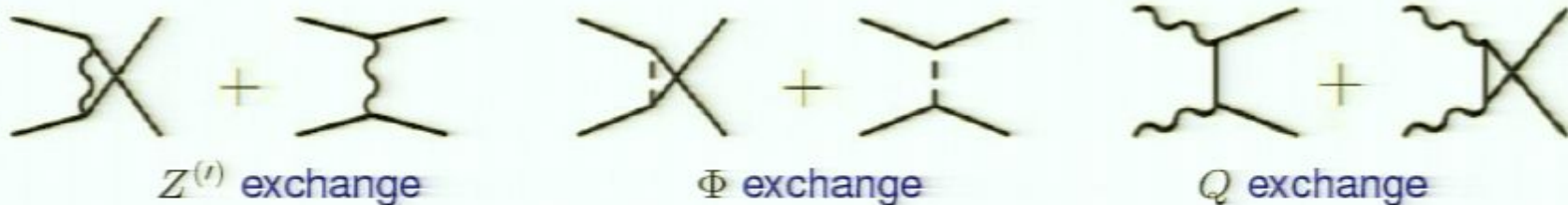
- Usually only the leading contribution from each operator is retained

Kinematic suppressions: $v^2 \sim 10^{-6}$

$$q^2/m_p m_\chi \sim 10^{-6} \text{ for } q^2 \sim (100 \text{ MeV})^2$$

This is comparable to (SD sensitivity) / (SI sensitivity) — if suppressed SI contributions exist, they should lead to detectable effects at roughly the same time

- Typically get v^2 suppressed SI operators:



- Pseudoscalar a exchange — no additional contributions generated



Loop-induced subleading effects

- Consider, e.g., Z exchange at tree level:

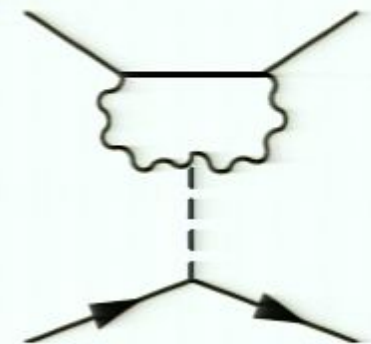
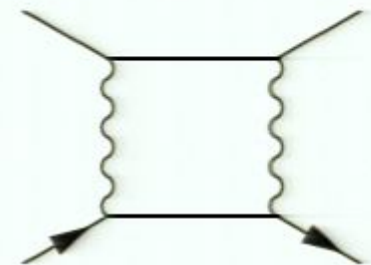
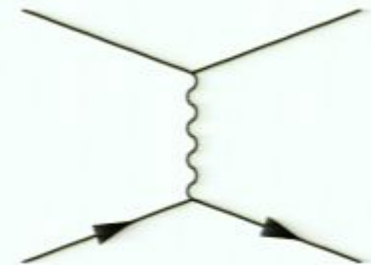
$$\frac{g_2^2}{2 \cos^2 \theta_W} T_3^q Q \frac{1}{m_Z^2} \bar{\chi} \gamma^\mu \gamma^5 \chi \bar{q} \gamma_\mu \gamma^5 q$$

- Loop effects induce the SI operator:

$$\frac{1}{4\pi} \frac{g_2^4 Q^2}{\cos^4 \theta_W m_Z} \left[\frac{(T_3^q)^2}{2m_Z^2} + \frac{1}{4m_h^2} \right] m_q \bar{\chi} \chi \bar{q} q \quad (m_q \ll m_Z \ll m_\chi)$$

For $Q \sim \mathcal{O}(1)$, the SI cross section is 6–7 orders of magnitude smaller than the SD one

However, if the DM is in an $SU(2)$ n-plet, the cross section $\sim n^2$



Loop effects: pseudoscalar mediator



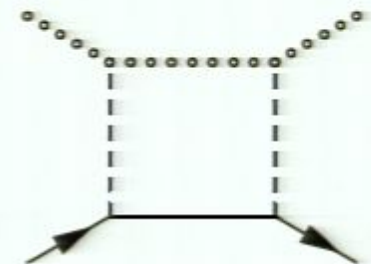
- Tree-level exchange:

$$\frac{1}{m_a^2} \xi y_q m_\phi \phi^\dagger \phi \bar{q} i \gamma^5 q$$



- Loop effects induce the SI operator:

$$\frac{1}{(4\pi)^2} \frac{\xi^2 y_q^2}{m_\phi^2} C_S \phi^\dagger \partial^\mu \phi \bar{q} \gamma_\mu q$$



Box diagram can be evaluated in terms of Passarino-Veltman integrals — for $m_a/m_\phi = 0.01$, $C_S \sim 80$

[involves $\ln^2(m_a/m_\phi)$ and $\ln(m_a/m_\phi)$ dependence]

- Worked out other cases as well...



Subleading effects — pseudoscalar mediator

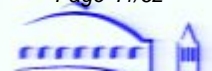
- Unlike Z case, we expect these loops have no detectable contribution:
 - The loop contribution is suppressed by an additional factor of m_a^2/m_χ^2
 - As Goldstone bosons would be expected to couple proportional to quark mass, loop effects should go as $(\text{Yukawa})^2$

A light pseudoscalar mediator is the only interaction providing a detectable SD cross section that would not be seen in a SI experiment roughly the same time

- So far, considered exclusively the role of mediators in direct detection
- Can this mechanism be implemented in a viable model?

The “axion portal” provides an example

[Nomura, Thaler, 0810.5397]



The axion portal



- A scalar field charged under a new global $U(1)_X$, spontaneously breaks to

$$S = \left(f_a + \frac{s}{\sqrt{2}} \right) \exp \left(\frac{ia}{\sqrt{2}f_a} \right)$$

New fermion coupled via $\xi S \chi \chi^c + \text{h.c.}$, then generate fermion mass $m_\chi = \xi f_a$

Let the SM (w/ 2HDM) be charged under $U(1)_X$ so the only coupling to S is via $\mathcal{L} = \lambda S^n H_u H_d + \text{h.c.}$, and $U(1)_X$ is a Peccei-Quinn symmetry

- The new fermions stay in thermal equilibrium though

$$\langle \sigma v \rangle_{\chi \chi^c \rightarrow sa} = \frac{m_\chi^2}{64\pi f_a^4} \left(1 - \frac{m_s^2}{4m_\chi^2} \right) + \mathcal{O}(v^4)$$

- If all scales $\sim \text{TeV}$, $\langle \sigma v \rangle \sim 3 \times 10^{-26} \text{ cm}^3/\text{s}$, so χ has right relic density to be DM



Axion portal and direct detection



- Two mediators present, s and a — Is the SI cross section due to s negligible?

$$\sigma_{\text{SI}}^{\chi N} \approx (2 \times 10^{-42} \text{ cm}^2) \xi^2 \epsilon^2 \left(\frac{100 \text{ GeV}}{m_s} \right)^4 \quad (\epsilon \sim v_{\text{ew}} f_a)$$

In original scenario $m_s \sim 10 \text{ GeV}$ for Sommerfeld enhancement, which was in tension with current bounds — removing this requirement, it is natural to have $m_s \sim f_a$; then the SI cross section is below the neutrino background

- Is the SD cross section large enough?

$$\sigma_{\text{SD}}^{\chi p} \approx (2 \times 10^{-37} \text{ cm}^2) \xi^2 \sin^2 \theta \frac{q_{\text{ref}}^2}{4m_\chi^2} \left(\frac{1 \text{ GeV}}{m_a} \right)^4$$

with $\tan \theta = n \sin 2\beta [v_{\text{ew}} / (2f_a)]$, $q_{\text{ref}}^2 = (100 \text{ MeV})^2$, and $m_a = \mathcal{O}(\text{few } 100 \text{ MeV})$, the cross section is $\text{few} \times 10^{-40} \text{ cm}^2$, near current bounds

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Axion portal and B decays

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Axion portal and B decays

Dark sectors — the motivation two years ago

- Observations of cosmic ray excesses led to a flurry of DM model building
Standard WIMPs unable to fit the data (lack of antiprotons, hard lepton spectrum)
- Idea: DM annihilates to SM through light bosons [Pospelov, Ritz, Voloshin; Arkani-Hamed *et al.*]

$$\chi\chi \rightarrow \phi^{(*)}\phi^{(*)}, \quad \phi \rightarrow \ell^+\ell^-, \pi^+\pi^-, \dots$$

“Dark bosons” couple to leptons with $\alpha_X = \lambda_X^2/(4\pi)$, lots of different constraints depending on mass and coupling

- Most popular scenario: ϕ^μ couples to $\bar{\psi}\gamma_\mu\psi$ and mixes with γ (“dark photons”)



The axion portal in $B \rightarrow K^{(*)} \ell^+ \ell^-$?



- The new particle could also be a scalar with axion-like couplings [Nomura, Thaler, 0810.5397]

$$\mathcal{L}_{\text{int}} = \frac{\lambda}{f_a} (\bar{\psi} \gamma^\mu \gamma_5 \psi) \partial_\mu a \rightarrow \frac{\lambda m_\psi}{f_a} (\bar{\psi} \gamma_5 \psi) a$$

The most interesting part of parameter space is thought to be:

$$m_K - m_\pi < m_a \lesssim 800 \text{ MeV}, \quad f_a \sim (1 - 3) \text{ TeV}$$

- Coupling to fermions $\propto m_\psi$, so FCNC $b \rightarrow sa$ loops are enhanced by m_t

With only \mathcal{L}_{int} , divergent loops \Rightarrow need to embed in a renormalizable theory

- A simple explicit model: Peccei-Quinn symmetric NMSSM (2HDM + a singlet)
(SUSY part not directly relevant for us, more general DFSZ-axion)

- At one loop: $\mathcal{M}(b \rightarrow sa) \propto \mathcal{M}(b \rightarrow sA^0)_{2\text{HDM}}$ (from tW , tH , tHW penguins)



The 2HDM calculation



[Hall and Wise, NPB 187 (1981) 397]

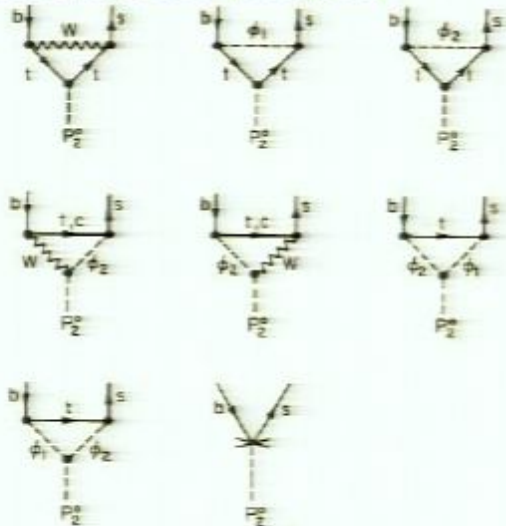
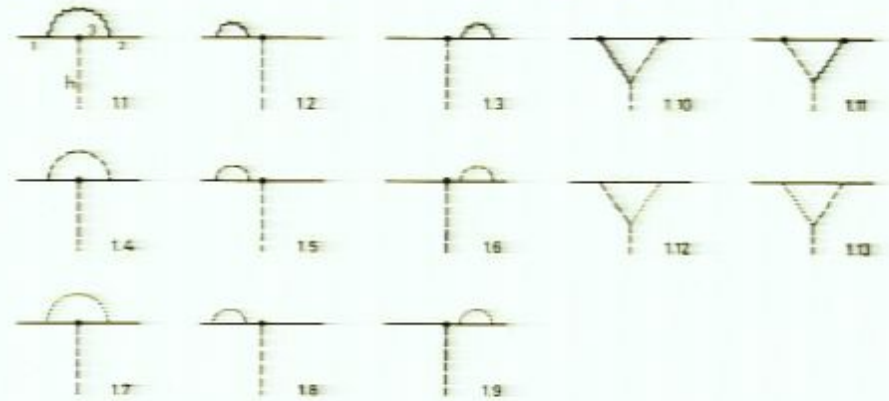


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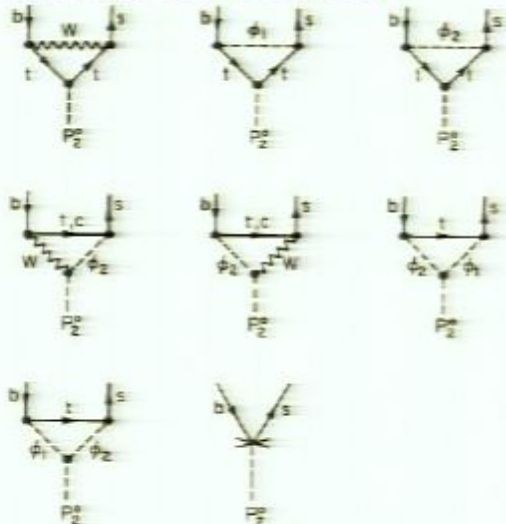
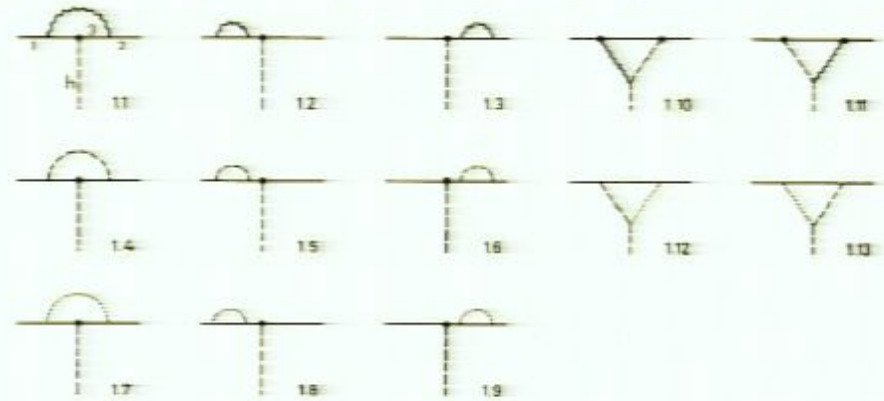


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The current data



- Considering the combined BaBar / Belle rate measurements and the spectra...

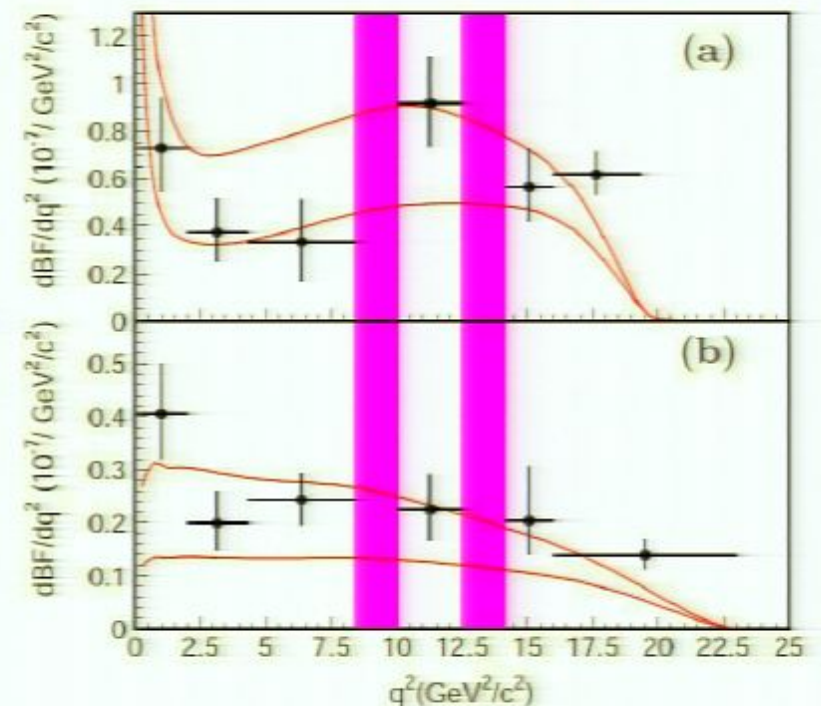
we used: $\mathcal{B}(B \rightarrow Ka) \times \mathcal{B}(a \rightarrow \mu^+ \mu^-) < 10^{-7}$

[at a high, but who-knows-what CL...]

For this physics $K\ell^+\ell^-$ may be better than $K^*\ell^+\ell^-$, since no O_7 (photon penguin) enhancement at small q^2 in K mode

Can improve independent of form factor uncertainties

[Wei *et al.*, Belle Collaboration, PRL 103 (2009) 171801]



- BaBar and Belle should be able to set a significantly better bound
- LHCb should be able to improve it substantially (more than an order of magnitude)

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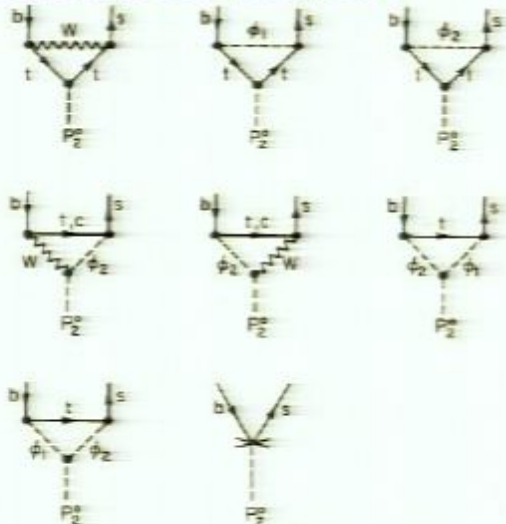
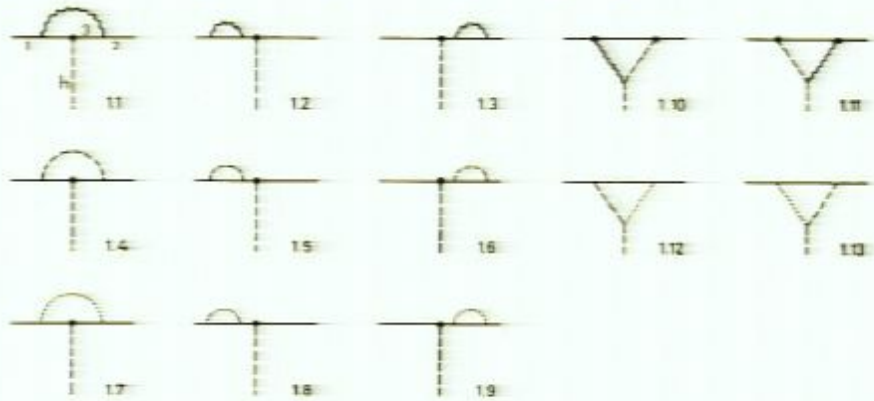


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$$\begin{aligned}
 A_1(m) = & \frac{-3M_W^4}{(M_W^2 - m^2)(M_W^2 - M_H^2)} \left(1 + \frac{M_H^2 - m^2}{M_W^2 - m^2} \right) \\
 & \times \ln \frac{m^2}{M_W^2} + \frac{M_H^2}{(M_H^2 - m^2)} \left(\frac{6M_W^2}{M_W^2 - M_H^2} + \frac{M_H^2}{M_H^2 - m^2} \right) \\
 & \times \ln \frac{m^2}{M_W^2} + 2 + \frac{3M_W^2}{M_W^2 - m^2} + \frac{M_H^2}{M_H^2 - m^2}, \quad (9)
 \end{aligned}$$

- Results disagree, neither knew about other
- Many papers cited both, none commented on disagreement... so we computed it all...



The current data



- Considering the combined BaBar / Belle rate measurements and the spectra...

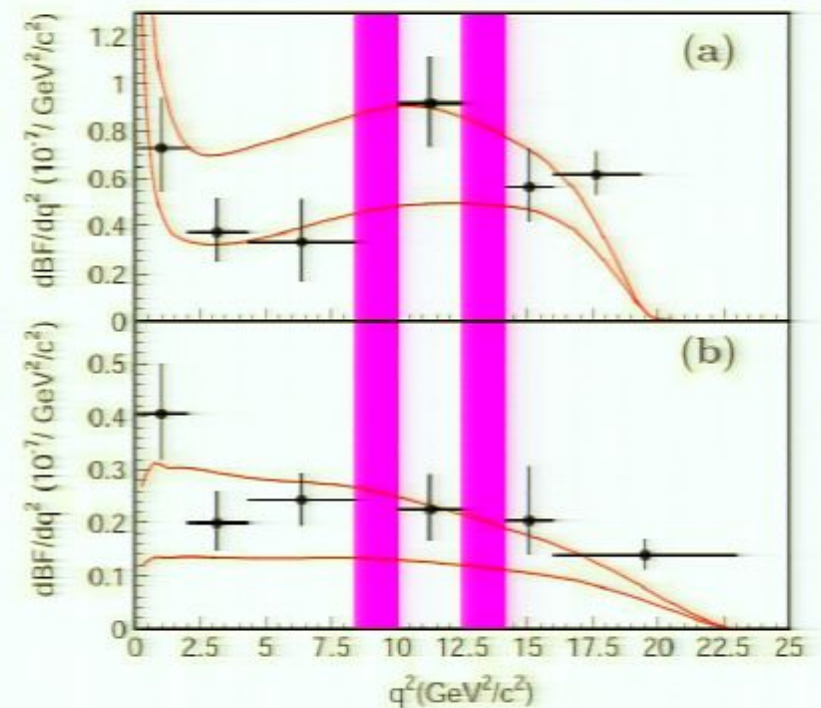
we used: $\mathcal{B}(B \rightarrow Ka) \times \mathcal{B}(a \rightarrow \mu^+ \mu^-) < 10^{-7}$

[at a high, but who-knows-what CL...]

For this physics $K\ell^+\ell^-$ may be better than $K^*\ell^+\ell^-$, since no O_7 (photon penguin) enhancement at small q^2 in K mode

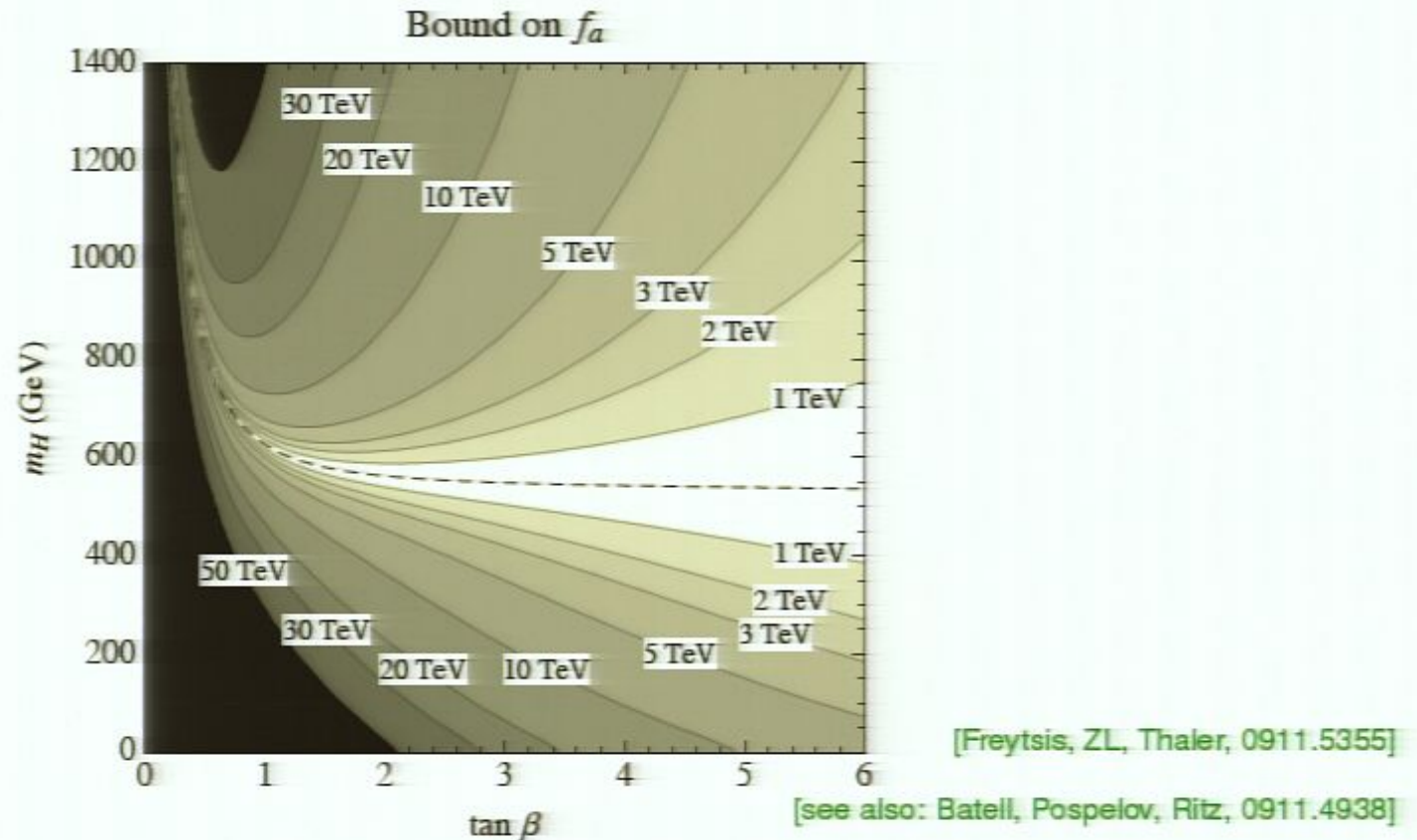
Can improve independent of form factor uncertainties

[Wei *et al.*, Belle Collaboration, PRL 103 (2009) 171801]



- BaBar and Belle should be able to set a significantly better bound
- LHCb should be able to improve it substantially (more than an order of magnitude)

The bound from $B \rightarrow K\ell^+\ell^-$



- Cancellation in a narrow region near the dashed line (between $\cot \beta$ and $\cot^3 \beta$ terms)
- In most of the parameter space this is the best bound (then $\Upsilon(3S) \rightarrow \gamma A^0$)
[BaBar, 0902.2176]



Conclusions

Summary



- Most DM models with dominantly spin-dependent interactions will not be (much) harder to see in spin-independent experiments through subleading interactions
- Unique exception seems to be cases where the mediators are light pseudoscalars
- A viable dark matter model already exists with such a mechanism
- A significant part of its parameter space is best probed in B and Υ decays
- Do other (not much more complex) models with SD interaction dominance exist?



The current data



- Considering the combined BaBar / Belle rate measurements and the spectra...

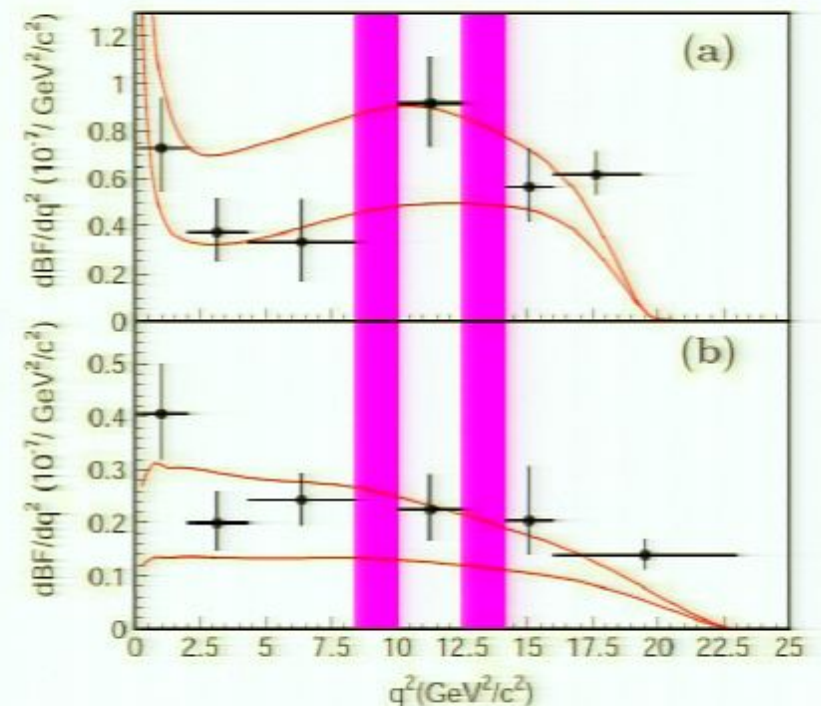
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