

Title: Explorations in Numerical Relativity - Lecture 5

Date: Apr 08, 2011 11:30 AM

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Abstract:

PHY 387NSPHERICAL SYMMETRY (SR)

- IN $\Rightarrow$ SPACETIME NEED MATTER FOR DYNAMICS
(BIRKHOFF'S THM, UNIQUENESS of SCHWARZSCHILD
SOL^N AS SOL^N of $G_{ab} = 0$)
- WILL RESTRICT MATTER CONTENT TO SINGLE, MASSLESS,
MINIMALLY-COUPLED SCALAR FIELD, ϕ
 - GOOD MODEL PROBLEM FOR STUDYING STRONG-
FIELD, RADIATIVE SIT'S - INCLUDING BLACK
HOLE FORMATION
 - EXHIBITS INTERESTING PHYSICAL BEHAVIOUR $\rightarrow$
CRITICAL PHENOMENA aka BLACK HOLE PIPESTEM
PHENOMENA

$$^{(3)}R + K^2 - K^i_T K^T_i = 16\pi S$$

$$D_b K^{bT} - D^i K = 8\pi J^i$$

$$2_t g_{i,T} - \mathcal{L}_\beta g_{i,T} = -2\alpha K_{i,T}$$

$$2_t K_{i,T} - \mathcal{L}_\beta K_{i,T} = -D^j D_{i,j} \alpha + \alpha [R_{i,T} + K K_{i,T} +$$

$$8\pi \left[\frac{1}{2} g_{i,T} (S - S) - S \right]$$

PH7 387NSPHERICAL SYMMETRY (SR)

- IN 3D SPACETIME NEED MATTER FOR DYNAMICS
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PHENOMENA

AS EINK & EKG (EINSTEIN-TIMLESS KLEIN-GORDON)

LAGRANGIAN DENSITY FOR EINK SYSTEM

$$\begin{aligned} \mathcal{L} &= \mathcal{L}_{\text{grav}} + \mathcal{L}_\phi \\ &= \sqrt{-g} \left(R - \frac{1}{2} \nabla_a \phi \nabla^a \phi \right) \end{aligned}$$

"CONSTRAINT" E.O.M

$$G_{ab} = 8\pi T_{ab} = 8\pi \left(\nabla_a \phi \nabla_b \phi - \frac{1}{2} g_{ab} \nabla_c \phi \nabla^c \phi \right)$$

$$\square \phi \equiv \nabla^a \nabla_a \phi = 0$$

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$$\zeta_{ab} = g_{\pi\pi} T_{ab} = g_{\pi\pi} (\nabla_a \phi \nabla_b \phi - \frac{1}{2} g_{ab} \nabla_c \phi \nabla^c \phi)$$

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3.1 FORM of SPACETIME METRIC IN SS

• COORDINATES (t, r, θ, ϕ) ADAPTED TO S.S.

METRIC ON UNIT 2-SPHERE $d\Omega^2 = d\theta^2 + \sin^2 \theta d\phi^2$

• THEN MOST GENERAL 3-METRIC IS

$$h_{ij} = \text{diag}(a^2(r, t), r^2 b^2(r, t), r^2 b^2 \sin^2 \theta) \quad (1)$$

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THE LAPSE FUNCTION IS $\alpha(r,t)$, AND THE SHIFT VECTOR $\beta^i(r,t)$ HAS ONLY A RADIAL COMPONENT, $\beta(r,t)$

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$$\beta^i = (\beta, 0, 0) \quad (2)$$

$$\beta_i = \gamma_{ij} \beta^j = (a^2 \beta, 0, 0) \quad (3)$$

• THE MOST GENERAL 4-METRIC IS THEN



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THE MOST GENERAL 4-METRIC IS THEN

$$\begin{aligned} ds^2 &= (-\alpha^2 + \beta^i \beta_i) dt^2 + 2\beta_i dt dx^i + \gamma_{ij} dx^i dx^j \\ &= (-\alpha^2 + a^2 \beta^2) dt^2 + 2a^2 \beta dt dr + a^2 dr^2 + r^2 b^2 d\phi^2 \quad (4) \end{aligned}$$

CORRESPONDING EXTRINSIC CURVATURE TENSOR, K^i_j ,
LIKE γ_{ij} , HAS ONLY TWO INDEPENDENT COMPONENTS

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EASY TO SHOW EQUALITY

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$$\partial_t K_{ij} - \mathcal{L}_\beta K_{ij} = -D^i D_j \alpha + \alpha [R_{ij} + K K_{ij} +$$

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EASY TO SHOW EQUALITY
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$$K_{ij} = (\partial_\alpha)^{-1} (-\partial_t \gamma_{ij} + D_i \gamma_{j\alpha} + D_j \gamma_{i\alpha})$$

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SPHERICAL SYMMETRY

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* SO, HAVE REDUCED TOTAL # OF GRAVITATIONAL KIN. !

DYN VOLS FROM 16 TO 6, AND, OF COURSE, THESE
VOLS ARE FUNCTIONS ONLY OF (r,t) RATHER THAN

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EASY TO SHOW EQUALITY
FROM

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SPHERICAL SYMMETRY

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 (x, y, z, t)

EINSTEIN EQUATIONS

(1) CONSTRAINTS

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 (x, y, z, t) EINSTEIN EQUATIONS(1) CONSTRAINTS

$$R - K^i_i; K^i_i - K^2 = 16\pi\mathcal{D} \quad (6)$$

$$D_j K^i_j - D_i K = 8\pi j_i \quad (7)$$

(NOTE INDEX SHIFT RELATIVE TO PREVIOUS FORM)

WHERE

$$\mathcal{D} = n_\mu n_\nu T^{\mu\nu} \quad (8)$$

$$j_i = \gamma_{ik} j^k = -n_\mu T^\mu_i \quad (9)$$

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(NOTE INDEX SHIFT RELATIVE TO PREVIOUS TEST)

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RECALL: $n_\mu = (-\alpha, 0, 0, 0)$ (10)

(:) EVOLUTION EQUATIONS $(\cdot \equiv \frac{\partial}{\partial t} \equiv \partial_t)$

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$$\dot{\gamma}_{ij} = -2\alpha \gamma_{ik} K^k_j + \beta^k \partial_k \gamma_{ij} + \gamma_{ik} \partial_j \beta^k + \gamma_{kj} \partial_i \beta^k \quad (11)$$

$$\dot{K}^i_j = \beta^k \partial_k K^i_j - \partial_k \beta^i K^k_j + \partial_j \beta^k K^i_k - D^i D_j \alpha$$

$$+ \alpha (R^i_j + K K^i_j + 4\pi(S-\rho)\delta^i_j - 8\pi S^i_j) \quad (12)$$

WHERE $\rho = n_\mu n_\nu T^{\mu\nu}$ (8)

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$$\text{WHERE } S_{ij} = T_{ij} \quad (13) \quad S^i_j = \gamma^{ik} S_{kj} \quad (14) \quad S = S^i_i \quad (15)$$

NEED CHRISTOFFEL SYMBOLS Γ^i_{jk} , RICCI COMPONENTS R^i_j AND RICCI SCALAR R ASSOCIATED WITH γ_{ij} (1). USING STANDARD FORMULAE FIND FOLLOWING NON-VANISHING Γ^i_{jk} ($' \equiv \frac{\partial}{\partial r} \equiv \partial_r$)

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 ($' \equiv \frac{\partial}{\partial r} \equiv \partial_r$)

$$\Gamma^r_{rr} = \frac{a'}{a}$$

$$\Gamma^r_{\theta\theta} = -\frac{(r^2 b^2)'}{2a^2}$$

$$\Gamma^r_{\phi\phi} = \sin^2\theta \Gamma^r_{\theta\theta}$$

$$\Gamma^\theta_{r\theta} = \Gamma^\theta_{\theta r} = \frac{(r^2 b^2)'}{2(r^2 b^2)}$$

$$\Gamma^\theta_{\phi\phi} = -\sin\theta \cos\theta \quad (16a-g)$$

$$\Gamma^\phi_{r\phi} = \Gamma^\phi_{\phi r} = \Gamma^\theta_{\theta r}$$

$$\Gamma^\phi_{\theta\phi} = \Gamma^\phi_{\phi\theta} = \cot\theta$$

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$$2(r^2 b^2)$$

$$\Gamma^d_{rd} = \Gamma^d_{dr} = \Gamma^e_{er} \quad \Gamma^d_{ed} = \Gamma^d_{de} = \cot \theta$$

FROM THESE WE COMPUTE Ricci-tensor R^i_j

$$R^r_r = -\frac{2}{arb} \left(\frac{(rb)'}{a} \right)' \quad (17a)$$

$$R^e_e = R^d_d = \frac{1}{a(rb)^2} \left(a - \left(\frac{rb}{a} (rb)' \right)' \right) \quad (17b)$$

AND FINALLY, THE SCALAR CURVATURE, R , IS

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$$= -\frac{2}{ab} \left(\left(\frac{(rb)'}{a} \right)' + \frac{1}{rb} \left(\left(\frac{rb}{a} (rh)' \right)' - a \right) \right) \quad (1E)$$

IN THE EVALUATION EQUATION FOR K^i_j , WE NEED TO
EVALUATE $D^i D_j \alpha$:

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$$\begin{aligned} D^i D_i \alpha &= \gamma^{ik} D_k D_i \alpha = \gamma^{ik} D_k \partial_i \alpha \\ &= \gamma^{ik} (\partial_k \partial_i \alpha - \Gamma^m_{ik} \partial_m \alpha) \end{aligned}$$

USING RESULTS FROM ABOVE, WE FIND

$$D^r D_r \alpha = \frac{1}{a} \left(\frac{\alpha'}{a} \right)' \quad (19a)$$

$$D^\theta D_\theta \alpha = D^\phi D_\phi \alpha = \frac{\alpha' (r b)'}{r b} \quad (19b)$$

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$$D^\theta D_\theta \alpha = D^\phi D_\phi \alpha = \frac{\alpha'(rb)'}{a^2 r b} \quad (19b)$$

$$D^0 D_0 \alpha = D^d D_d \alpha = \frac{\alpha'(rb)'}{a^2 rb} \quad (196)$$

• ALSO NEED STRESS-TENSOR "COMPONENTS" J_i
 j_i, S'_i

• IN SPIRIT OF HAMILTONIAN APPROACH, IT IS CONVENIENT TO INTRODUCE AUXILIARY FUNCTIONS

$$\Phi(r, t) \equiv \phi'(r, t) = \partial_r \phi(r, t) \quad (20)$$

$$\pi(r, t) \equiv \frac{a}{\alpha} (\dot{\phi} - \beta \phi') \quad (21)$$

VIEW Φ, π AS "REAL" DYNAMICAL VLS FOR

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VIEW Φ, π AS "REAL" DYNAMICAL VLS FOR SCALAR FIELD; NOTE: FOR MASSLESS FIELD, VALUE OF ϕ IS MEANINGLESS ($\phi \rightarrow \phi + \text{const}$ STILL SATISFIES $\square \phi = 0$) ALL "ACTION" IS IN GRADIENTS OF ϕ (I.E. IN Φ AND π)

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ALSO NOTE THAT WE HAVE (C/F (A))

$$g_{tt} = -\alpha^2 + a^2 \beta^2$$

$$g_{tr} = g_{rt} = a^2 \beta$$

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ALSO NOTE THAT WE HAVE (c/f (a))

$$g_{tt} = -\alpha^2 + a^2 \beta^2 \quad g_{tr} = g_{rt} = a^2 \beta \quad (22a-c)$$

$$g_{rr} = a^2 \quad g_{\theta\theta} = r^2 b^2 \quad g_{\phi\phi} = r^2 b^2 \sin^2 \theta$$

AND

$$g^{tt} = -\alpha^{-2} \quad g^{tr} = g^{rt} = \beta \alpha^{-2} \quad (23a-c)$$

$$g^{rr} = a^{-2} - \beta^2 \alpha^{-2} \quad g^{\theta\theta} = (rb)^{-2} \quad g^{\phi\phi} = (rb \sin \theta)^{-2}$$

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AND

$$g^{tt} = -\alpha^{-2}$$

$$g^{tr} = g^{rt} = \beta \alpha^{-2} \quad (23a-e)$$

$$g^{rr} = \alpha^{-2} - \beta^2 \alpha^{-2}$$

$$g^{\theta\theta} = (r\alpha)^{-2}$$

$$g^{\phi\phi} = (r\alpha \sin\theta)^{-2}$$

THEN WE FIND

$$\nabla_\mu \Phi = \partial_\mu \Phi = g_{\mu\nu} \partial^\nu \Phi = -\frac{1}{\alpha a} \quad (24)$$

$$(\nabla^\mu \Phi)(\nabla_\mu \Phi) = \partial'^\mu \Phi \partial_{,\mu} \Phi = \frac{\Phi^2 - \Pi^2}{a^2} \quad (25)$$

AND, AGAIN RECALLING THAT $n_\mu = (-\alpha, 0, 0, 0)$, WE FIND

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THEN WE FIND

$$\nabla^+ \phi \equiv \phi'^t = g^{tt} \partial_t \phi + g^{tr} \partial_r \phi = -\frac{\pi}{\alpha a} \quad (24)$$

$$(\nabla^\mu \phi)(\nabla_\mu \phi) = \phi'^\mu \phi_{,\mu} = \frac{\Phi^2 - \pi^2}{a^2} \quad (25)$$

AND, AGAIN RECALLING THAT $n_\mu = (-\alpha, 0, 0, 0)$, WE FIND

$$\begin{aligned} \rho = n_\mu n_\nu T^{\mu\nu} &= \alpha^2 T^{tt} = \alpha^2 \left(\phi'^t \phi'^t - \frac{1}{2} g^{tt} \phi'^\mu \phi_{,\mu} \right) \\ &= \frac{\Phi^2 + \pi^2}{2a^2} \quad (26) \end{aligned}$$

$$(\nabla^\mu \phi)(\nabla_\mu \phi) = \dot{\phi}^{\prime\mu} \dot{\phi}_{,\mu} = \frac{\dot{\Phi}^2 - \Pi^2}{a^2} \quad (25)$$

AND, AGAIN RECALLING THAT $n_\mu = (-\alpha, 0, 0, 0)$, WE FIND

$$\begin{aligned} \rho = n_\mu n_\nu T^{\mu\nu} &= \alpha^2 T^{tt} = \alpha^2 \left(\dot{\phi}^{\prime t} \dot{\phi}^{\prime t} - \frac{1}{2} g^{tt} \dot{\phi}^{\prime\mu} \dot{\phi}_{,\mu} \right) \\ &= \frac{\dot{\Phi}^2 + \Pi^2}{2a^2} \quad (26) \end{aligned}$$

$$j_i = (j_i, 0, 0)$$

$$j_i = -n_\mu T^\mu_i = \alpha T^0_i = \alpha \dot{\phi}^{\prime t} \dot{\phi}_{,i} = -\frac{\dot{\Phi} \Pi}{a} \quad (27)$$

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$$\begin{aligned} \rho &= n_\mu n_\nu T^{\mu\nu} = \alpha^2 T^{tt} = \alpha^2 \left(\dot{\phi}^{t+} \dot{\phi}^{t+} - \frac{1}{2} g^{++} \dot{\phi}^{t+} \dot{\phi}_{t+} \right) \\ &= \frac{\vec{E}^2 + \pi^2}{2a^2} \quad (26) \end{aligned}$$

$$j_i = (j_i, 0, 0)$$

$$j_r = -n_\mu T^{\mu}_i = \alpha T^0_i = \alpha \dot{\phi}^{t+} \dot{\phi}_{i+} = -\frac{\vec{E} \cdot \pi}{a} \quad (27)$$

$$= \frac{\bar{g}^2 + \pi^2}{2a^2} \quad (26)$$

$$j_i = (j_r, 0, 0)$$

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PHY 387H SPHERICAL SYMMETRY ④

FINALLY, WE HAVE THE SPATIAL STRESS COMPONENTS

$$S^i_j = \gamma^{ik} S_{kj} = \gamma^{ik} T_{kj}$$

2a

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PHY 382M SPHERICAL SYMMETRY

④

FINALLY, WE HAVE THE SPATIAL STRESS COMPONENTS

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FROM WHICH WE FIND

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FROM WHICH WE FIND

$$S^r_r = \rho = \frac{\Phi^2 + \pi^2}{2a^2} \quad (28)$$

$$S^e_e = S^d_d = \frac{\pi^2 - \Phi^2}{2a^2} \quad (29)$$

$$S - \rho = 2S^e_e = \frac{\pi^2 - \Phi^2}{a^2} \quad (30)$$

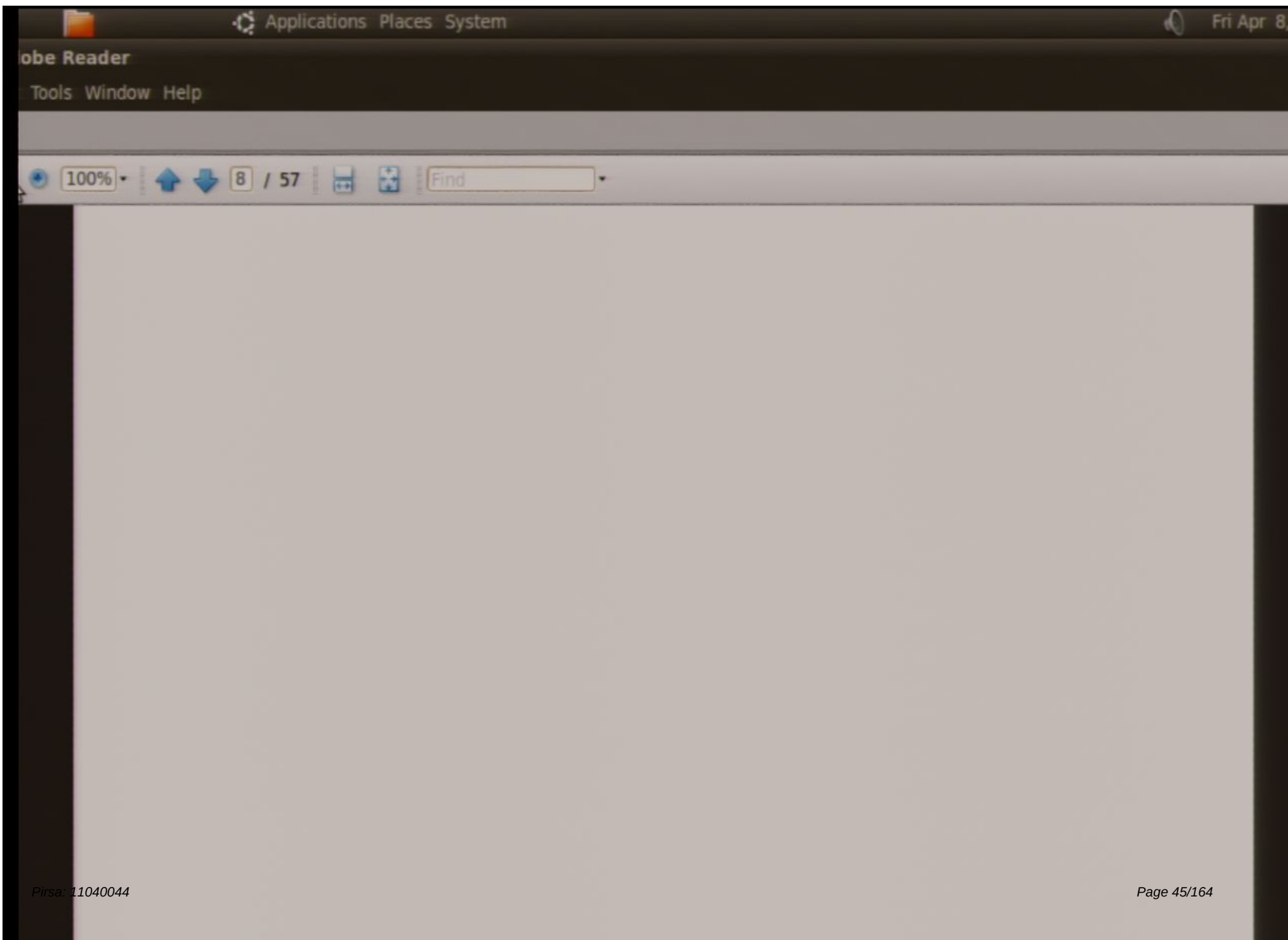
$$\frac{1}{a^2}$$

• WE CAN NOW ASSEMBLE THE ABOVE RESULTS TO REDUCE THE SPHERICALLY-SYMMETRIC SPECIALIZATION OF THE GENERAL 3+1 EQUATIONS (6), (7), (11) ; (12)

A) HAMILTONIAN CONSTRAINT

$$R - K^i_j K^j_i + K^2 = 16\pi\rho$$

$$\begin{aligned} -K^i_j K^j_i + K^2 &= -(K^r_r{}^2 + 2K^e_e{}^2) + (K^r_r + 2K^e_e)^2 \\ &= 4K^r_r K^e_e + 2K^e_e{}^2 \end{aligned}$$



$$= K_r' + 2 \Gamma_{re} (K_r' - K_e^0)$$

$$D_r K = (K_r' + 2 K_e^0)'$$

THEN WE HAVE FROM (2) AND (27)

$$K_e^0' + \frac{(rb)'}{rb} (K_e^0 - K_r') = 4\pi \frac{\Phi \Pi}{a} \quad (32)$$

c) EVOLUTION EQUATIONS FOR γ_{ij} (a, b)

* FOLLOW DIRECTLY FROM (11), RECALL, CAN BE VIEWED AS DEFⁿ OF K^i ;

$$\dot{a} = -\alpha a K_r' + (a \beta)'$$

$$K^i; K^j; + K^2 = 16\pi \rho$$

$$\begin{aligned} -K^i; K^j; + K^2 &= -(K_r^r{}^2 + 2K_e^e{}^2) + (K_r^r + 2K_e^e)^2 \\ &= 4K_r^r K_e^e + 2K_e^e{}^2 \end{aligned}$$

$$R + 4K_r^r K_e^e + 2K_e^e{}^2 = 8\pi \frac{\bar{\Phi}^2 + \Pi^2}{a^2} \quad (31)$$

SEE (18)

B) MOMENTUM CONSTRAINT (ONLY r-COMPONENT IS NON-TRIVIAL)

• FIRST NOTE THAT

$$D_i K^i{}_r = \partial_i K^i{}_r + \Gamma^i{}_{mi} K^m{}_r - \Gamma^m{}_{ri} K^i{}_m$$

A) HAMILTONIAN CONSTRAINT

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$$D_r K = (K^r_r + 2K^e_e)'$$

THEN WE HAVE FROM (2) AND (27)

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$$\frac{1}{a^2} \left(\frac{1}{2} \frac{d}{dt} \left(\frac{1}{2} \frac{d}{dt} \right) \right) = \frac{1}{2} \frac{d}{dt} \left(\frac{1}{2} \frac{d}{dt} \right) + \frac{1}{2} \frac{d}{dt} \left(\frac{1}{2} \frac{d}{dt} \right) \quad (31)$$

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C) EVOLUTION EQUATIONS FOR γ_{ij} (a, b)

rb

a

C) EVOLUTION EQUATIONS FOR γ_{ij} (a, b)

- FOLLOW DIRECTLY FROM (11), RECALL, CAN BE VIEWED AS DEFⁿ OF K^i ;

$$\dot{a} = -\alpha a K^r_r + (a\beta)' \quad (33)$$

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D) EVOLUTION EQUATIONS FOR K^i (K^r_r, K^e_e)

- FOLLOW DIRECTLY FROM (12) AND OTHER RESULTS ABOVE

$$\dot{K}^r_r = 2K^r_r' - 1(\alpha')' + \alpha(-2 \int_r (rb)')' \cdot K^r_r = \frac{3}{8\pi} \frac{\Phi^2}{r^2}$$

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$$\dot{K}^r_r = \beta K^r_r' - \frac{1}{a} \left(\frac{\alpha'}{a} \right)' + \alpha \left(-\frac{2}{rab} \left(\frac{(rb)'}{a} \right)' + K K^r_r - 5\pi \frac{\Phi^2}{a^2} \right)$$

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• FOLLOW DIRECTLY FROM (12) AND OTHER RESULTS ABOVE

$$\dot{K}^r_r = \beta K^r_r' - \frac{1}{a} \left(\frac{\alpha'}{a} \right)' + \alpha \left(-\frac{2}{r a b} \left(\frac{(r b)'}{a} \right)' + K K^r_r - 5 \pi \frac{\dot{\Phi}^2}{a^2} \right) \quad (35)$$

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SPHERICAL SYMMETRY

⑤

$$\dot{K}^{\theta}_{\theta} = \beta K^{\theta}_{\theta}' + \frac{\alpha}{(r b)^2} - \frac{1}{a (r b)^2} \left(\frac{\alpha r b (r b)'}{a} \right)' + \alpha K K^{\theta}_{\theta} \quad (36)$$

• (31)-(36) ARE THE COMPLETE 3+1 EQUATIONS FOR THE

GEOMETRIC VARIABLES (NOTE: WE HAVE NOT YET

PHY 387M SPHERICAL SYMMETRY ①

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PHI 3E7NSPHERICAL SYMMETRY

⑤

$$\dot{K}_0^0 = \beta K_0^0 + \frac{x}{(rh)^2} - \frac{1}{a(rh)^2} \left(\frac{drh}{a} (rh)' \right)' + \alpha K K_0^0 \quad (36)$$

← (31)-(36) ARE THE COMPLETE 3+1 EQUATIONS FOR THE GEOMETRIC VARIABLES (NOTE: WE HAVE SAID NOTHING YET RE COORDINATE CHOICES, I.E. RE SPECIFICATIONS OF α AND β)

MASSLESS KLEIN GORDON EQUATION

WAVE E.O.M. FOR $\bar{\phi}$ AND π

WHEED E.O.P. FOR $\bar{\Phi}$ AND π

* RECALL DEFⁿ OF π , (23)

$$\pi \equiv \frac{\alpha}{\lambda} (\dot{\phi} - \beta \phi')$$

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* BUT $\dot{\phi}' = \dot{\bar{\Phi}}$, so

$$\dot{\bar{\Phi}} = \left(\beta \bar{\Phi} + \frac{\alpha}{\lambda} \pi \right)' \quad (33)$$

* TO FIND π EQU., RECALL THAT

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©

CHARACTERISTIC ANALYSIS of THE SCALAR FIELD

REF: COURANT : HILBERT "METHODS of MATHEMATICAL PHYSICS",
VOL II, ch 5

* EQNS (37)-(38) ARE A 1ST-ORDER, QUASI-LINEAR SYS
FOR OUR RADIATION FIELD. DEFINING

$$u \equiv (\mathbf{E}, \pi)^T$$

WE CAN WRITE

$$\rightarrow \partial_t (\sqrt{-g} g^{tt} \partial_t \phi) + \partial_r (\sqrt{-g} g^{rr} \partial_r \phi)$$

$$= \partial_t (\alpha a r^2 b^2 (-\alpha^{-2} \dot{\phi} + \beta \alpha^{-2} \phi'))$$

$$+ \partial_r (\alpha a r^2 b^2 ((\alpha^{-2} - \beta^2 \alpha^{-2}) \phi' + \beta \alpha^{-2} \dot{\phi}))$$

$$= 0$$

$$\rightarrow (r^2 b^2 \pi) = (r^2 b^2 (\beta \pi + \frac{\alpha}{a} \dot{\phi}))'$$

(WRITE AS $(r^2 b^2) \dot{\pi} + (r^2 \dot{b}^2) \pi$ AND USE
EQUATION (39) FOR $\dot{b}$

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B) MOMENTUM CONSTRAINT (only r-component is non-trivial)

First note that

$$\begin{aligned} D_i K^i_r &= \partial_i K^i_r + \Gamma^i_{mi} K^m_r - \Gamma^m_{ri} K^i_m \\ &= K^r_r' + 2\Gamma^e_{re} (K^r_r - K^e_e) \end{aligned}$$

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Then we have from (2) and (2')

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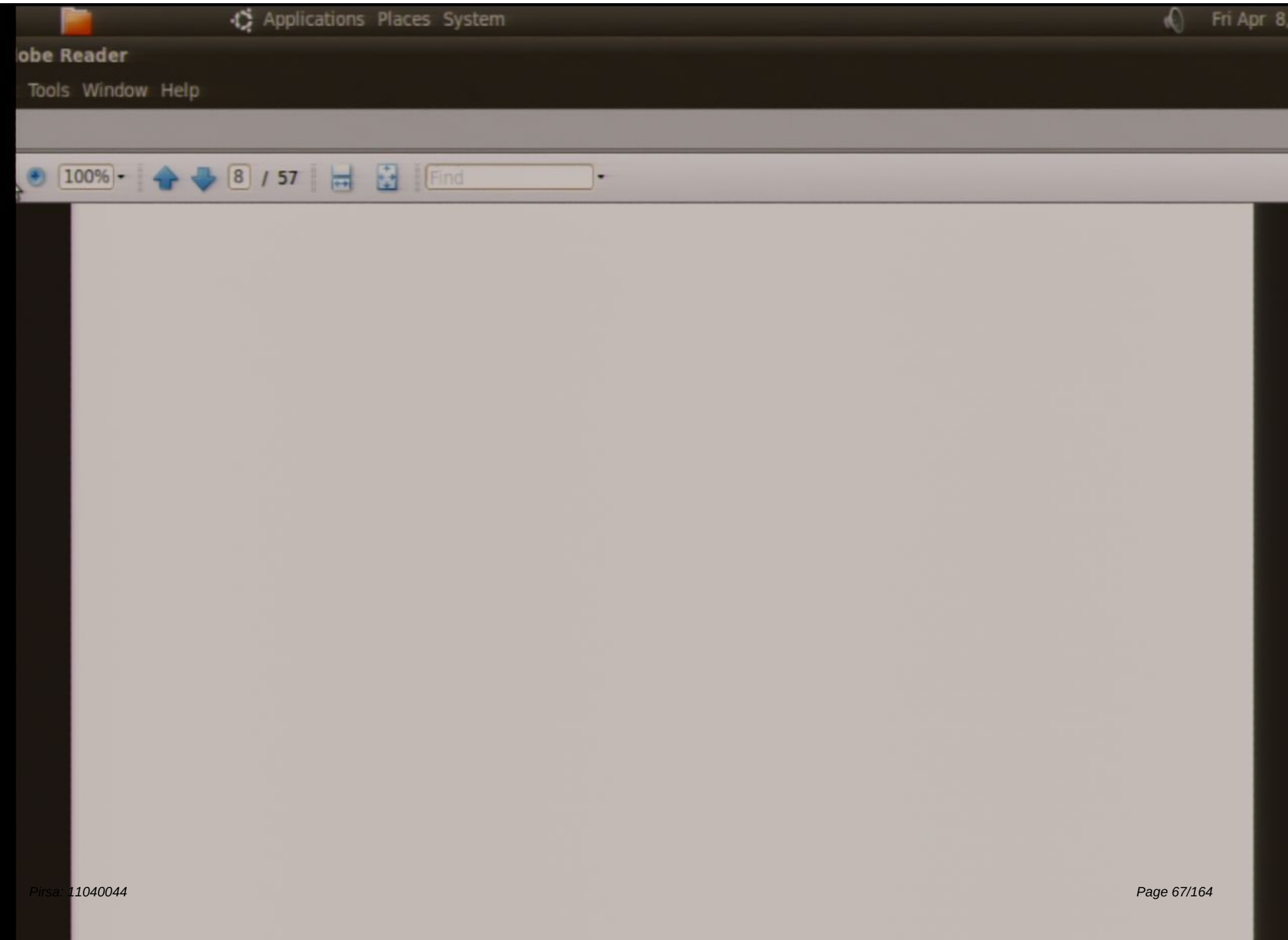
$$\dot{a} = -\alpha a K^r_r + (a\beta)' \quad (33)$$

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$$\dot{K}^r_r = \beta K^r_r' - \frac{1}{\alpha} (\alpha')' + \alpha \left(-\frac{2}{r} \left(\frac{(rb)'}{r} \right)' + K K^r_r - 5\pi \frac{\phi^2}{r^3} \right)$$





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• But $\dot{\phi}' = \dot{\Phi}$, so

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• TO FIND π EQU., RECALL THAT

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GEOMETRIC VARIABLES (NOTE: WE HAVE SAID NOTHING
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PHYS 387NSPHERICAL SYMMETRY

⑤

$$\dot{K}_G^G = \beta K_G^{G'} + \frac{\chi}{(rh)^2} - \frac{1}{a(rh)^2} \left(\frac{\alpha rh (rh)'}{a} \right)' + \alpha K K_G^G \quad (36)$$

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MASSLESS KLEIN GORDON! EQUATION!

WHEED E.O.M. FOR $\bar{\Phi}$ AND Π

* RECALL DEF OF Π (24)



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Find

YET RE COORDINATE CHOICES, I.E. RE SPECIFICATIONS
OF α AND β)

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• TO FIND π EGN, RECALL THAT

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$$\rightarrow \phi = \frac{\alpha}{a} \pi + \beta \bar{\phi} = \frac{\alpha}{a} \pi + \beta \bar{\phi}$$

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$$\pi + \beta \bar{\Phi}$$

$$^{(3)} R + K^2 - K^i{}_T K^T{}_i = 16\pi S$$

$$D_b K^{ij} - D^i K = 8\pi J^i$$

$$\partial_t g_{ij} - \mathcal{L}_\beta g_{ij} = -2\alpha K_{ij}$$

$$\partial_t K_{ij} - \mathcal{L}_\beta K_{ij} = -D_i D_j \alpha + \alpha [R_{ij} + K K_{ij} +$$

$$8\pi \left[\frac{1}{2} g_{ij} (S - S) - S \right]$$

• TO FIND Π EQU, RECALL THAT

$$\square \phi = \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} g^{\mu\nu} \partial_\nu \phi)$$

$$\square \phi = 0 \rightarrow \partial_\mu (\sqrt{-g} g^{\mu\nu} \partial_\nu \phi) = 0$$

$$\rightarrow \partial_t (\sqrt{-g} g^{tt} \partial_t \phi) + \partial_r (\sqrt{-g} g^{rr} \partial_r \phi)$$

$$= \partial_t (\alpha a r^2 b^2 (-\alpha^{-2} \dot{\phi} + \beta \alpha^{-2} \phi'))$$

$$+ \partial_r (\alpha a r^2 b^2 ((\alpha^{-2} - \beta^2 \alpha^{-2}) \phi' + \beta \alpha^{-2} \dot{\phi}))$$

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$$= 0$$

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$$\sqrt{-g} \rightarrow (\dots)$$

$$\Box \phi = 0 \rightarrow \partial_\mu (\sqrt{-g} g^{\mu\nu} \partial_\nu \phi) = 0$$

$$\rightarrow \partial_t (\sqrt{-g} g^{tt} \partial_t \phi) + \partial_r (\sqrt{-g} g^{rr} \partial_r \phi)$$

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$$= 0$$

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$$= 0$$

$$\rightarrow (r^2 b^2 \pi) = (r^2 b^2 (\beta \pi + \frac{\alpha}{a} \dot{\phi}))'$$

$$\rightarrow \text{WRITE AS } (r^2 b^2) \dot{\pi} + (r^2 b^2) \pi \text{ AND USE}$$

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= 0

$$\rightarrow (r^2 b^2 \pi)' = (r^2 b^2 (\beta \pi + \frac{\alpha}{a} \underline{\pi}))'$$

(WRITE AS $(r^2 b^2) \dot{\pi} + (r^2 b^2) \pi$ AND USE
EVALUATION EQUATION (39) FOR $\dot{b}$

$$\frac{1}{r^2 b^2} \left(\frac{r^2 b^2}{r^2 b^2} \left(\beta \pi + \frac{\alpha}{a} \underline{\pi} \right) \right)$$

(38)

$$+ 2 \left(\alpha K^0 e - \beta \frac{(r b)'}{r b} \right) \pi$$

WRITE AS $(r^2 b^2) \ddot{\pi} + (r^2 \dot{b}^2) \dot{\pi}$ AND USE
EQUATION (39) FOR $\dot{b}$

$$\ddot{\pi} = \frac{1}{r^2 b^2} \left(r^2 b^2 \left(\beta \pi + \frac{\alpha}{a} \underline{\Phi} \right) \right)' \quad (52)$$

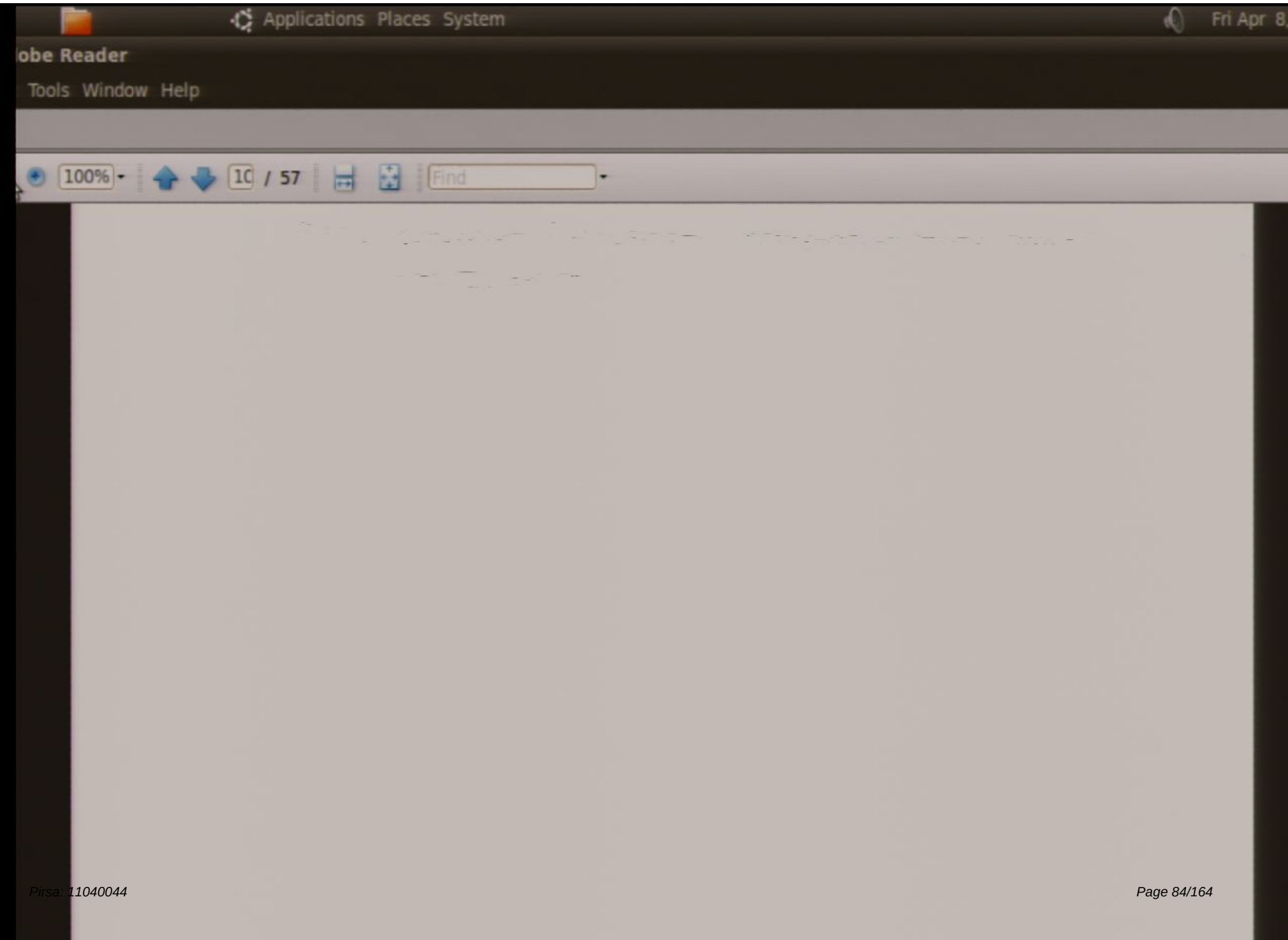
$$+ 2 \left(\alpha K^{\circ} e - \beta \frac{(r b)'}{r b} \right) \dot{\pi}$$

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REF: COUPANT: HILBERT "METHODS OF MATH. PHYS",
VOL II, ch 5

- EQNS (37)-(38) ARE A 1ST-ORDER, CURI-LINEAR SYS
FOR OUR RADIATION FIELD. DEFINING

$$u = (\underline{E}, \pi)^T$$

WE CAN WRITE

WE CAN WRITE

$$u_t + A u_x = B \quad (39)$$

$$A = - \begin{pmatrix} B & x/c \\ 1 & \end{pmatrix} \quad (40)$$

PHY 3821 SPHERICAL SYMMETRY (6)

CHARACTERISTIC ANALYSIS of THE SCALAR FIELD

REF: COURANT : HILBERT "METHODS of MATH. PHYS",
VOL II, ch 5

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$$u \equiv (\Phi, \Pi)^T$$

WE CAN WRITE

PHY 3821 SPHERICAL SYMMETRY

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CHARACTERISTIC ANALYSIS of THE SCALAR FIELD

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$$u = (\mathbf{E}, \pi)^T$$

WE CAN WRITE



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VOL II, ch 5

- EQU'S (37)-(38) ARE A 1ST-ORDER, QUASI-LINEAR SYS FOR OUR RADIATION FIELD. DEFINING

$$u \equiv (\mathbf{E}, \pi)^T$$

WE CAN WRITE

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VOL II, cel 5

- EENS (37)-(38) ARE A 1ST-ORDER, CURI-LINEAR SYS FOR OUR RADIATION FIELD. DEFINING

$$u = (\mathbf{E}, \pi)^T$$

WE CAN WRITE

$$u_t + A u_x = B \quad (39)$$

$$A = - \begin{pmatrix} \beta & x/a \\ x/a & \beta \end{pmatrix} \quad (40)$$

AND B IS ANOTHER MATRIX WHICH DOES NOT INVOLVE DERIVATIVES OF u .

$$A = - \begin{pmatrix} B & \chi/a \\ \chi/a & B \end{pmatrix} \quad (40)$$

AND B IS ANOTHER MATRIX WHICH DOES NOT INVOLVE DERIVATIVES OF ψ .

- THE CHARACTERISTIC DIRECTIONS $\tau \equiv dr/dt$ OF (39)-(40) ARE GIVEN BY

$$|A - \tau I| = 0$$

→

$$\tau = -B \pm \frac{\chi}{a} \quad (41)$$

THESE ARE THE "LOCAL SIGNAL SPEEDS" FOR THE SCALAR

DERIVATIVES of ψ .

- THE CHARACTERISTIC DIRECTIONS $\tau \equiv dr/dt$ of (39)-(40) ARE GIVEN BY

$$|A - \tau| = 0$$

→

$$\tau = -\rho \pm \frac{\alpha}{a} \quad (41)$$

THESE ARE THE "LOCAL SIGNAL SEEDS" FOR THE SCALAR FIELD

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→

$$\dot{r} = -\dot{\phi} \pm \frac{\alpha}{a}$$

(41)

THESE ARE THE "LOCAL SIGNAL SPEEDS" FOR THE SCALAR FIELD

• MASSLESS SCALAR FIELD - WEAK (NON-SELF GRAVITATING)
SIGNALS TRAVEL ALONG NULL GEODESICS → ALTERNATE
DERIVATION OF (41)

$$ds^2 = -\alpha^2 dt^2 + a^2 (dr + \dot{\phi} dt)^2 = 0$$

REGULARITY / LOCAL FLATNESS AT $r=0$

(MASSLESS SCALAR FIELD = WEAK CURVE SELF GRAVITATION)

SIGNALS TRAVEL ALONG NULL GEODESICS $\rightarrow$ ALTERNATE
DERIVATION of (4.1)

$$ds^2 = -\alpha^2 dt^2 + a^2(dr + \beta dt)^2 = 0$$

REGULARITY / LOCAL FLATNESS AT $r=0$

• OUR CAUCHY PROBLEM FOR THE EINSTEIN MODEL IS TO BE
SOLVED ON

$$t \geq 0, \quad r \geq 0$$

• BOUNDARY CONDITIONS AS $r \rightarrow \infty$ WILL FOLLOW FROM ASYM-
PTOTIC FLATNESS, NO INCOMING RADIATION; $r=0$ NOT
A REAL BOUNDARY, BUT COMPUTATIONALLY (I.E. WHEN FINITE

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- OUR CAUCHY PROBLEM FOR THE EINSTEIN MODEL IS TO BE SOLVED ON

$$t \geq 0, r \geq 0$$

- BOUNDARY CONDITIONS AS $r \rightarrow \infty$ WILL FOLLOW FROM ASYMPTOTIC FLATNESS, NO INCOMING RADIATION; $r=0$ NOT A REAL BOUNDARY, BUT COMPUTATIONALLY (I.E. WHEN FINITE DIFFERENCING) IS EFFECTIVELY ONE

- GET CONDITIONS AT $r=0$ BY DEMANDING THAT SCALAR GRAV. FIELDS BE REGULAR, AND THAT THE S.T. BE LOCALLY FLAT THERE

TRICKY ISSUE IN GENERAL UNLESS WE MAKE ASSUMPTIONS

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DERIVATION of (4.1)

$$ds^2 = -\alpha^2 dt^2 + a^2(dr + j dt)^2 = 0$$

REGULARITY / LOCAL FLATNESS at $r=0$

- OUR CAUCHY PROBLEM FOR THE EMKG MODEL IS TO BE SOLVED ON

$$t \geq 0, \quad r \geq 0$$

- BOUNDARY CONDITIONS AS $r \rightarrow \infty$ WILL FOLLOW FROM ASYMPTOTIC FLATNESS, NO INCOMING RADIATION; $r=0$ IS NOT A REAL BOUNDARY, BUT COMPUTATIONALLY (I.E. WHEN FINITE DIFFERENCING) IS EFFECTIVELY ONE

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DERIVATION of (4.1)

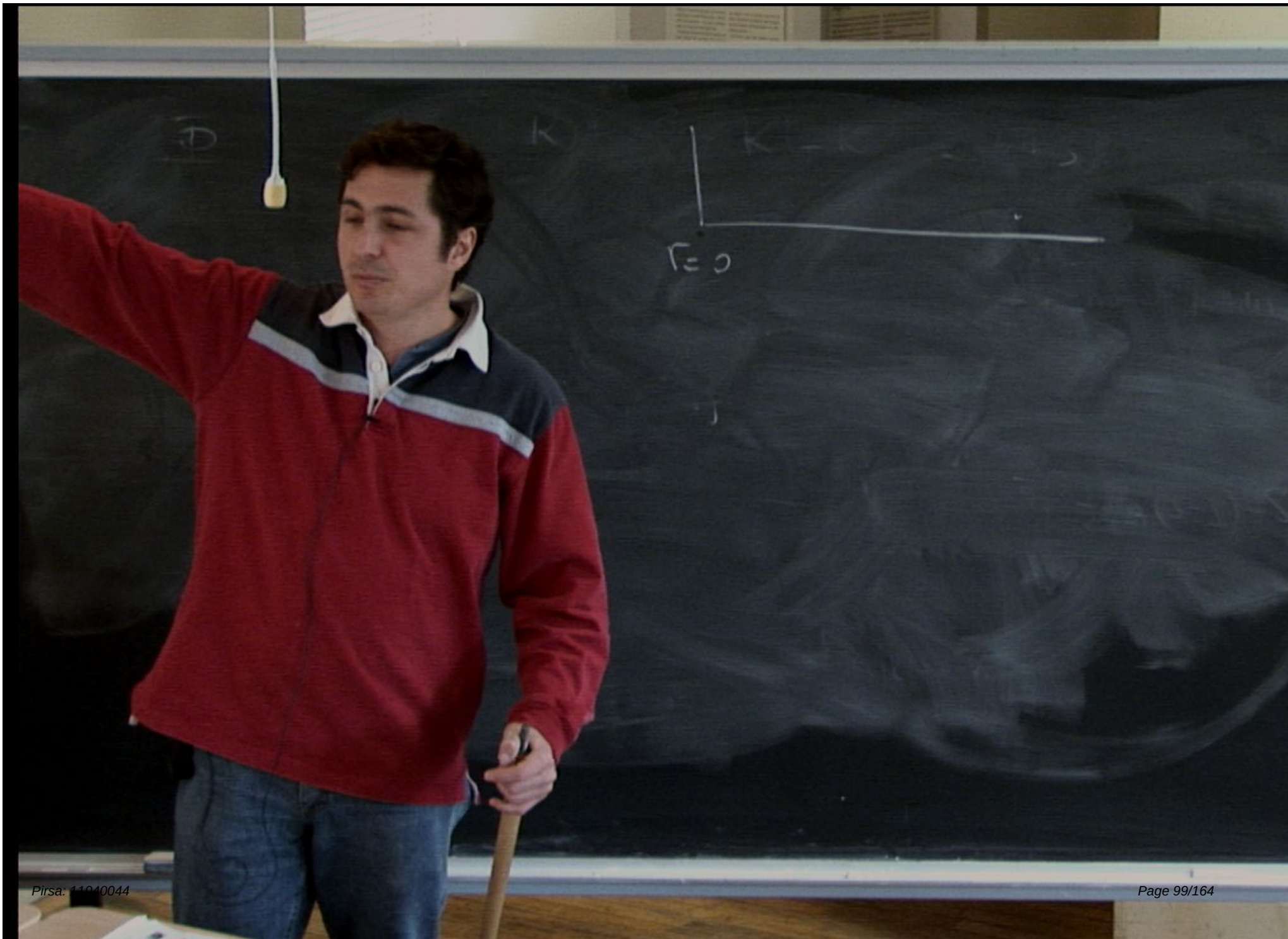
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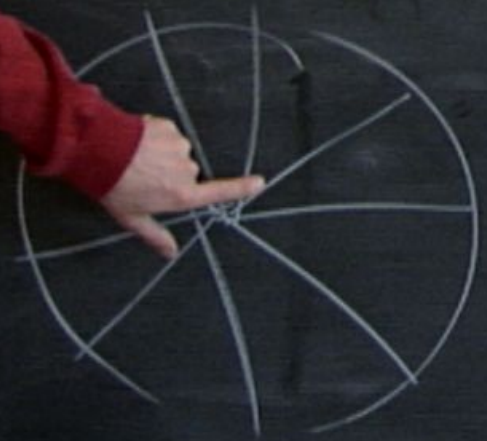
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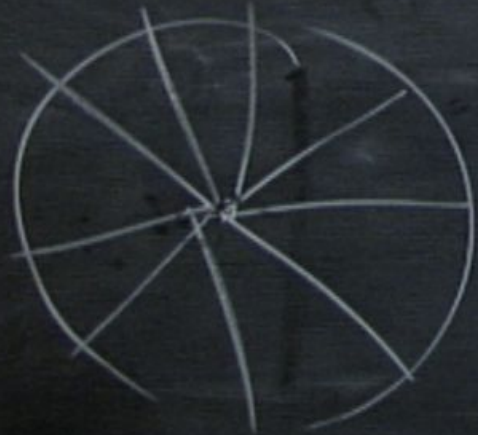
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SOLVED ON

$$t \geq 0, r \geq 0$$

• BOUNDARY CONDITIONS AS $r \rightarrow \infty$ WILL FOLLOW FROM ASYMPTOTIC FLATNESS, NO INCOMING RADIATION; $r=0$ NOT A REAL BOUNDARY, BUT COMputationALLY (I.E. WHEN FINITE DIFFERENCING) IS EFFECTIVELY ONE

• GET CONDITIONS AT $r=0$ BY DEMANDING THAT SCALAR
GRAV. FIELDS BE REGULAR, AND THAT THE S.T. BE
LOCALLY FLAT THERE

• TRICKY ISSUE IN GENERAL UNLESS WE MAKE ASSUMPTIONS ABOUT SLICING, SPATIAL COORDINATES (SEE BAROOSH; PIRAN, PHYSICS REPORTS 96 (1983) 205-250); WE WILL

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ASYMPTOTIC FLATNESS, NO INCOMING RADIATION; $r=0$ NOT A REAL BOUNDARY, BUT CONDITIONALLY (I.E. WHEN FINITE DIFFERENCING) IS EFFECTIVELY ONE

- GET CONDITIONS AT $r=0$ BY DEMANDING THAT SCALAR GRAM. FIELDS BE REGULAR, AND THAT THE S.T. BE LOCALLY FLAT THERE
- TRICKY ISSUE IN GENERAL UNLESS WE MAKE ASSUMPTIONS ABOUT SLICING, SPATIAL COORDINATES (SEE BAROEN; PIRAN, PHYSICS REPORTS 96 (1983) 205-250); WE WILL ASSUME EVERYWHERE SMOOTH SLICINGS AND SPATIAL COORDS

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(*) SCALAR FIELD

$$\nabla \phi = \phi' \vec{r}$$

NOT DEFINED AT $r=0$ UNLESS

COMPONENTS CAN BE EXPANDED IN NON-NEG. POWERS
OF x, y, z :

$$x = r \sin \theta \cos \phi \quad y = r \sin \theta \sin \phi \quad z = r \cos \theta$$

• APPLYING TO OUR CASE (WITH ASSUMPTION OF STRONG
COORDS) FIND

$$g(a, t) = g_0(t) + r^2 g_2(t) + \dots$$

WHERE g IS ANY OF a, b, K^r OR K^e

$$\rightarrow a'(a, t) = b'(a, t) = K^r'(a, t) = K^e'(a, t) = 0$$

(43a-d)

• THESE CONDITIONS WILL BE CONSISTENT WITH EVOL.

$$x = r \sin \theta \cos \phi \quad y = r \sin \theta \sin \phi \quad z = r \cos \theta$$

• APPLYING TO OUR CASE (WITH ASSUMPTION OF STRAIGHT COORDS) FIND

$$g(r, t) = g_0(t) + r^2 g_2(t) + \dots$$

WHERE g IS ANY OF a, b, K^r_r OR K^e_e

$$\rightarrow a'(r, t) = b'(r, t) = K^r_r'(r, t) = K^e_e'(r, t) = 0 \quad (43a-d)$$

• THESE CONDITIONS WILL BE CONSISTENT WITH EVOL. EQUIS ONLY IF

WHERE g IS ANY OF a, b, K^r OR K^e

$$\rightarrow a'(0,t) = b'(0,t) = K^r'(0,t) = K^e'(0,t) = 0 \quad (43a-d)$$

THESE CONDITIONS WILL BE CONSISTENT WITH EVOL. EQUIS ONLY IF

$$\alpha'(0,t) = \beta(0,t) = 0 \quad (44)$$

$$\beta(0,t) = r\beta_2(t) + c(r^3)$$

(SMOOTHNESS OF COORDINATES)

LOCAL FLATNESS: CONSIDER A TRANSPORT OF AN ARBITRARY VECTOR ABOUT CIRCUMFERENCE

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$$J(a, t) = r B_2(t) + o(r^3)$$

(SMOOTHNESS OF COORDINATES)

LOCAL FLATNESS: CONSIDER A TRANSPORT OF
AN ARBITRARY VECTOR AROUND CLOSED LOOP ENCLOSED

$r \rightarrow 0$ on $\theta = \frac{\pi}{2}$ (EQUATORIAL PLANE); DEMAND THAT
THERE BE NO NET ROTATION IN LIMIT LOOP SHRUNK TO
POINT $\Rightarrow$ $\lim_{r \rightarrow 0}$ PROPER CIRCUM / PROPER RADIUS = 2π
AND

$$\nabla^\mu(rb) \nabla_\mu(rb) = 1$$

USING PREVIOUS RESULTS INCLUDING ABOVE REG. CONDITIONS

LOCAL FLATNESS: CONSIDER A TRANSPORT OF
AN ARBITRARY VECTOR ABOUT CLOSED LOOP ENCLOSING

$r=0$ on $\Theta = \frac{\pi}{2}$ (EQUATORIAL PLANE); DEMAND THAT
THERE BE NO NET ROTATION IN LIMIT LOOP SHRUNK TO
POINT $\Rightarrow$ $\lim_{r \rightarrow 0}$ PROPER CIRCUM / PROPER RADIUS $= 2\pi$
FIND

$$\nabla^\mu(rb) \nabla_\mu(rb) = -1$$

USING PREVIOUS RESULTS INCLUDING ABOVE REG. CONDITIONS
FIND

$$a(0,t) = b(0,t)$$

(45)

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$r \rightarrow 0$ on $\theta = \frac{\pi}{2}$ (EQUATORIAL PLANE); DEMAND THAT
 THERE BE NO NET ROTATION IN LIMIT LOOP SHRUNK TO
 POINT $\Rightarrow$ $\lim_{r \rightarrow 0}$ PROPER CIRCUM / PROPER RADIUS $= 2\pi$
 FIND

$$\nabla^2(rb) \nabla_\mu(rb) = 1$$

USING PREVIOUS RESULTS INCLUDING ABOVE REG. CONDITIONS
 FIND

$$a(a,t) = b(a,t) \quad (45)$$

THIS + $\vec{a}, \vec{b}$ EGMC + DEC. THEN IT PLN

$$K'_r(0,1) = K^e_e(0,1) \quad (AG)$$

NOTE: REGULARITY CONDITIONS MUCH MORE INVOLVED
FOR AXISYMMETRY; MAINTAINING REGULARITY IN 4D.
SIMULATIONS ALSO MUCH MORE PROBLEMATIC THAN S.S.

TRAPPED SURFACES / APPARENT HORIZONS

WANT TO STUDY BLACK HOLE FORMATION; BH CHARACTERIZED
BY AN EVENT HORIZON WHICH CAN ONLY BE DETERMINED
ONCE COMPLETE S.T. HAS BEEN CONSTRUCTED

WANT TO STUDY BLACK HOLE FORMATION; BH CHARACTERIZED BY AN EVENT HORIZON WHICH CAN ONLY BE DETERMINED ONCE COMPLETE S.T. HAS BEEN CONSTRUCTED

USEFUL TO BE ABLE TO COMPUTE "INSTANTANEOUS" APPRX. (I.E. ON ANY HYPERSURFACE $\Sigma(t)$) TO E.H. - DIVIDED BY APPARENT HORIZON $\equiv$ OUTERMOST MAXIMALLY TRAPPED SURFACE

PHI ZERN

SPHERICAL SYMMETRY

(E)

TRAPPED SURFACE: 2-SURFACE WITH TOPOLOGY S^2 SUCH THAT DIVERGENCE OF OUTGOING NULL

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USING PREVIOUS RESULTS INCLUDING ABOVE REG. CONDITIONS
FIND

$$a(0,t) = b(a,t) \quad (45)$$

THIS + $\ddot{a}, \dot{b}$ ELEM + DER. THEN IMPLY

$$K'_r(0,t) = K^e_e(0,t) \quad (46)$$

NOTE: REGULARITY CONDITIONS MUCH MORE INVOLVED
FOR AXISYMMETRY; MAINTAINING REGULARITY IN AXI.
SIMULATIONS ALSO MUCH MORE PROBLEMATIC THAN S.S.

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(SMOOTHNESS OF COORDINATES)

§ LOCAL FLATNESS: CONSIDER A TRANSPORT OF
AN ARBITRARY VECTOR ABOUT CLOSED LOOP ENCLOSING

$r=0$ on $\Theta = \frac{\pi}{2}$ (EQUATORIAL PLANE); DEMAND THAT
THERE BE NO NET ROTATION IN LIMIT LOOP SHRUNK TO
POINT $\Rightarrow$ $\lim_{r \rightarrow 0}$ PROPER CIRCUM / PROPER RADIUS = 2π
FIND

$$\nabla^{\mu}(rb) \nabla_{\mu}(rb) = -1$$

USING PREVIOUS RESULTS INCLUDING ABOVE REG. CONDITIONS

FIND

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LOCAL FLATNESS: CONSIDER A TRANSPORT OF
AN ARBITRARY VECTOR ABOUT CLOSED LOOP ENCLOSING

$r=0$ on $\theta = \frac{\pi}{2}$ (EQUATORIAL PLANE); DEMAND THAT
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POINT $\Rightarrow$ PROPER CIRCLE / PROPER RADIUS $= 2\pi$
AND

$$\nabla^2(rb) \nabla_\mu(rb) = 1$$

USING PREVIOUS RESULTS INCLUDING ABOVE REG. CONDITIONS
FIND

$$a(0,t) = b(a,t)$$

(45)

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USING PREVIOUS RESULTS INCLUDING ABOVE REG. CONDITIONS
FIND

$$a(0,t) = b(0,t) \quad (45)$$

THIS + $\ddot{a}, \ddot{b}$ B.C.'S + REG. THEN IMPLY

$$K_r^r(0,t) = K_e^e(0,t) \quad (46)$$

NOTE: REGULARITY CONDITIONS MUCH MORE INVOLVED
FOR ELASTICITY; MAINTAINING REGULARITY IN ANI.
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TRAPPED SURFACES / APPARENT HORIZONS

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 BY AN EVENT HORIZON WHICH CAN ONLY BE DETERMINED
 ONCE COMPLETE S.T. HAS BEEN CONSTRUCTED

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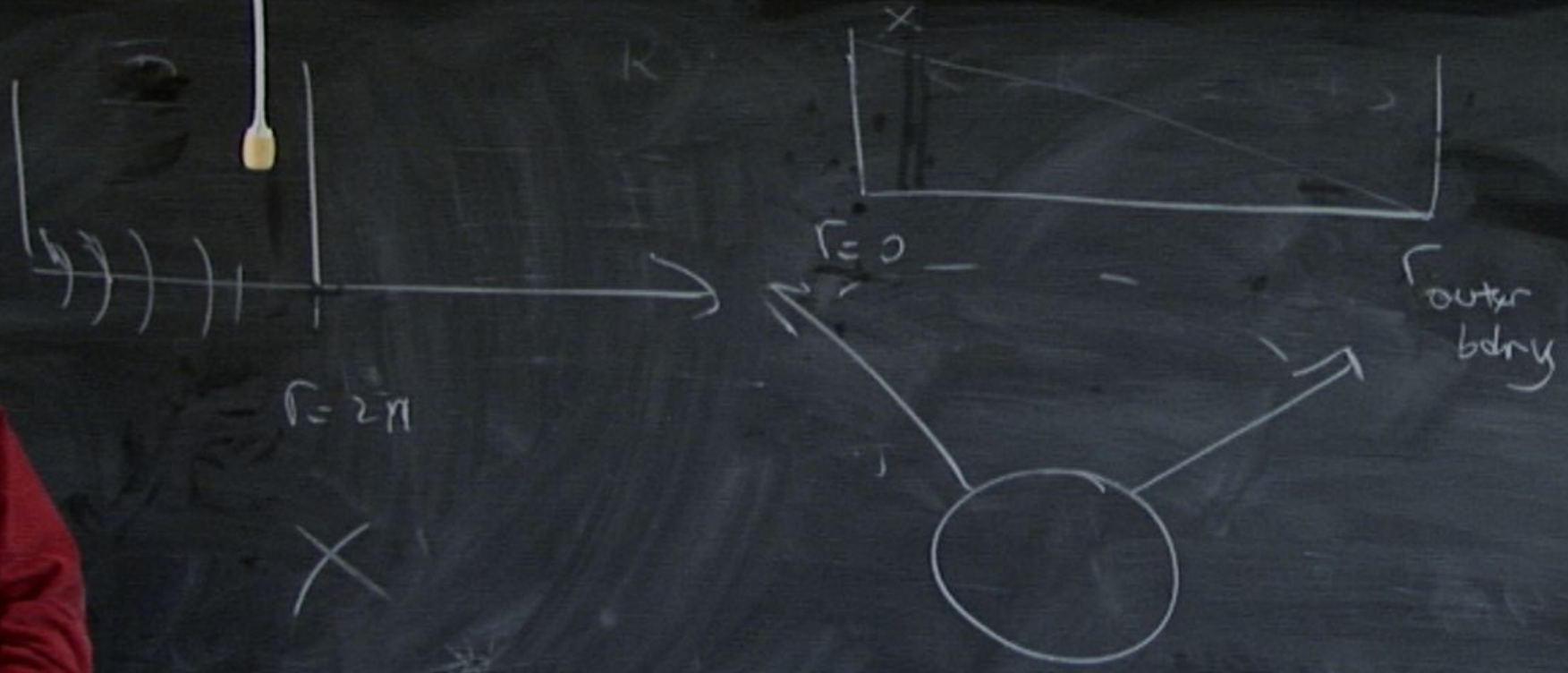
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MARGINALLY TRAPPED SURFACE

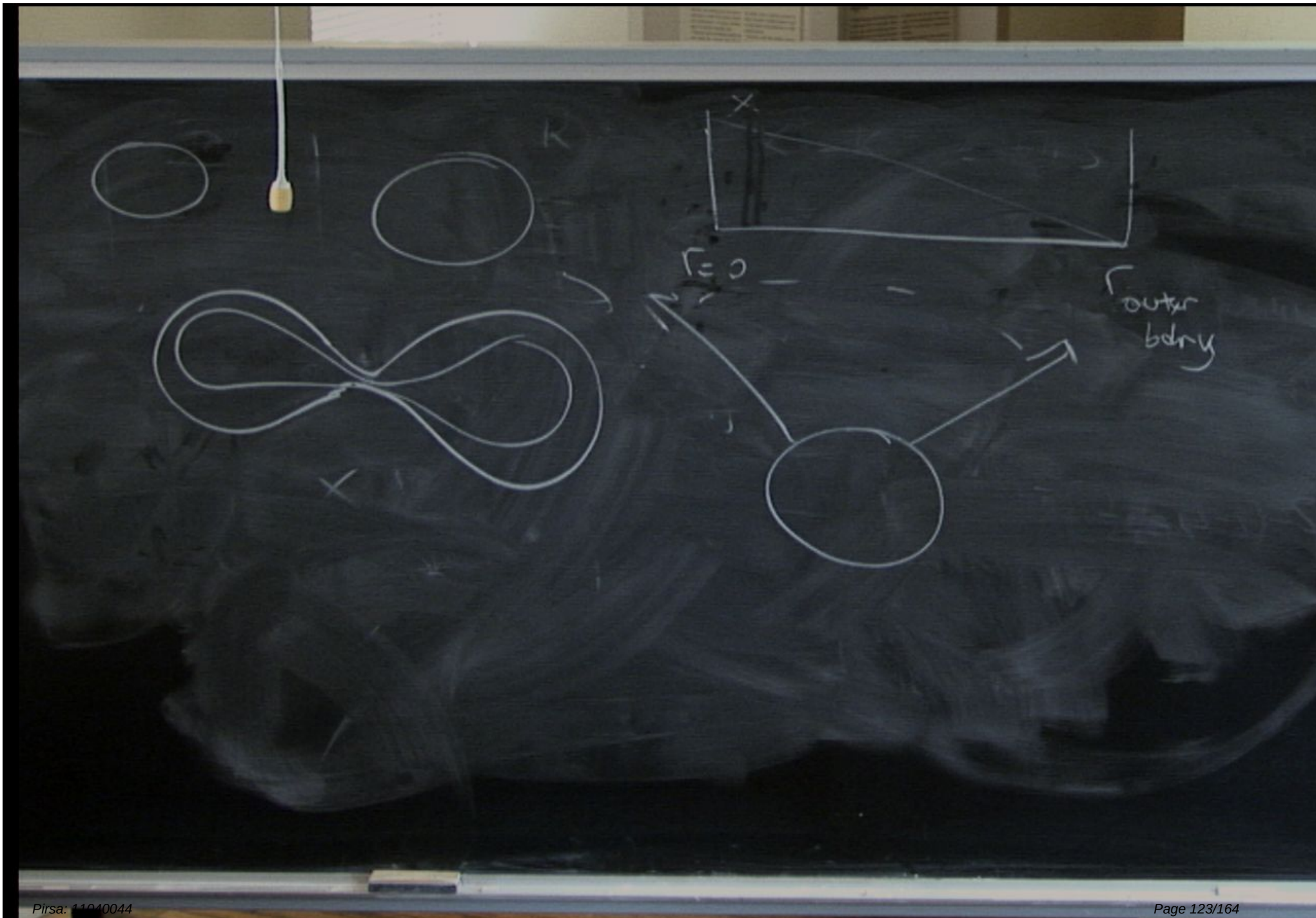


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MAXIMALLY TRAPPED SURFACE



• WILL TEND TO USE TERMS $\Delta H, T_3$; ITS INTER-
CHANGEABLY, BUT SHOULD BE AWARE OF DISTINCTIONS

(MARGINALLY) TRAPPED SURFACE EGM (ΔH EGM)

• CONSIDER A 2-SURFACE WITH OUTGOING NULL
TANGENT u^a WHICH IS MARGINALLY TRAPPED,
THEN

$$\nabla_a u^a = 0$$

NOW, CAN WRITE u^a AS

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now, can write u^a as

$$u^a = s^a + n^a$$

↳ UNIT FUTURE-DIRECTED TIMELIKE NORMAL TO Σ
 ↳ UNIT OUTWARDS-POINTING SPACELIKE NORMAL TO 2-SURF.

IN A 3+1 DECOMPOSITION, METRIC h_{ab} , h^{ab} IS INDUCED
 ON THE 2-SURFACE BY PROJECTION

$$h^{ab} = \gamma^{ab} - s^a s^b = g^{ab} + n^a n^b - s^a s^b$$

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(MARGINALLY) TRAPPED SURFACE EGM (ΔH EGM)

- CONSIDER A 2-SURFACE WITH OUTGOING NULL TANGENT u^a WHICH IS MARGINALLY TRAPPED, THEN

$$\nabla_a u^a = 0$$

NOW, CAN WRITE u^a AS

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CHANGEABLY, BUT SHOULD BE AWARE OF DISTINCTIONS

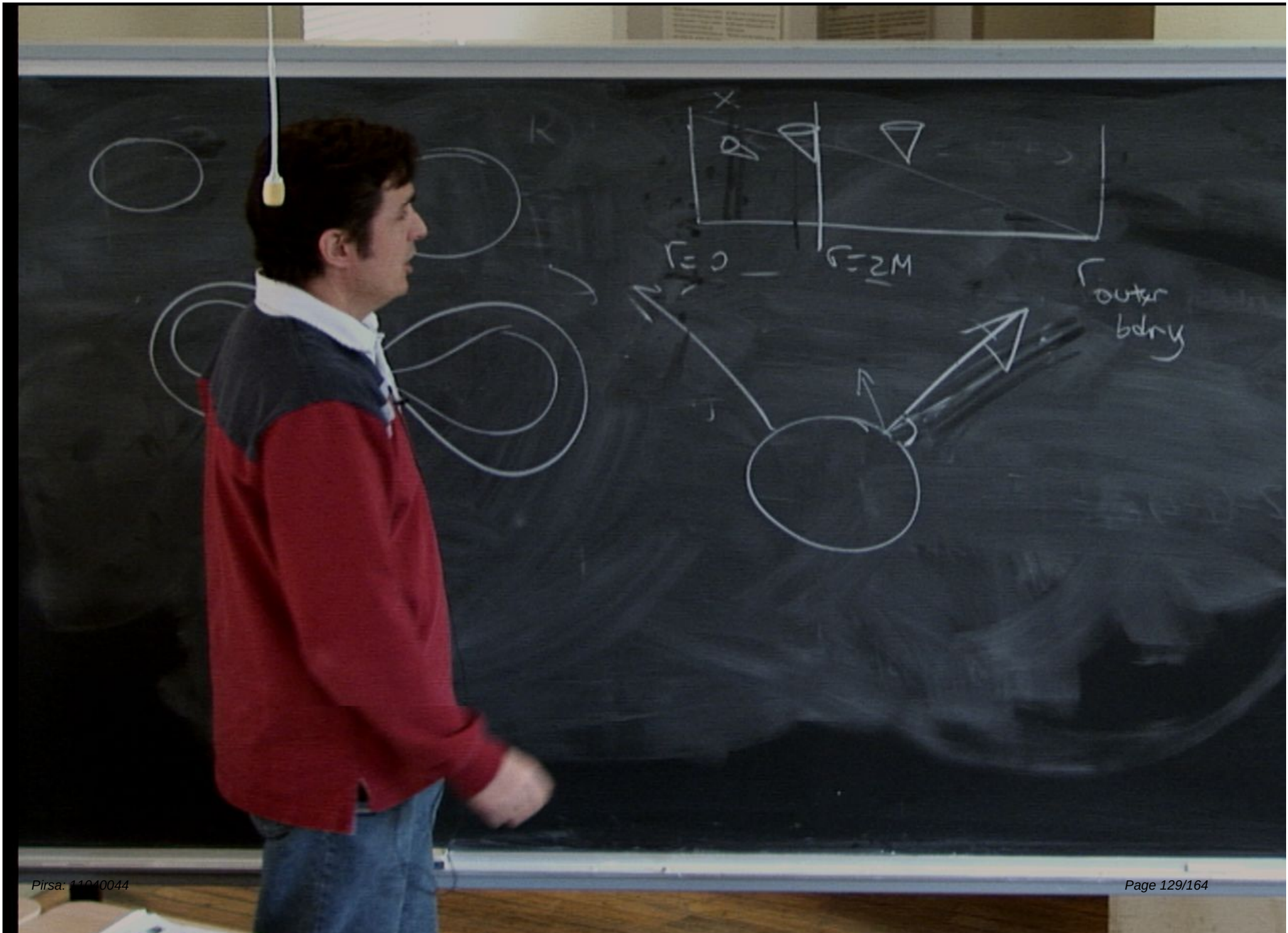
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 $\hookrightarrow$ UNIT OUTWARDS-POINTING SPACELIKE NORMAL TO 2-SURF.

IN $\neq$ TO 3+1 DECOMPOSITION, METRIC h_{ab} , h^{ab} IS INDUCED ON THE 2-SURFACE BY PROJECTION

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THEN

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NOW, CAN WRITE u^a AS

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IN $3+1$ DECOMPOSITION, METRIC h_{ab} , h^{ab} IS INDUCED
ON THE 2-SURFACE BY PROJECTION



$$u_a = 0$$



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$$\nabla_a u^a = 0$$

now, can write u^a as

$$u^a = s^a + n^a$$

$\hookrightarrow$ UNIT FUTURE-DIRECTED TIMELIKE NORMAL TO Σ
 $\hookrightarrow$ UNIT OUTWARDS-POINTING SPACELIKE NORMAL TO 2-SURF.

IN $3+1$ DECOMPOSITION, METRIC h_{ab} , h^{ab} IS INDUCED
ON THE 2-SURFACE BY PROJECTION

UNIT FUTURE-DIRECTED TIMELIKE NORMAL TO Σ
 UNIT OUTWARDS-POINTING SPACELIKE NORMAL TO 2-SURF.

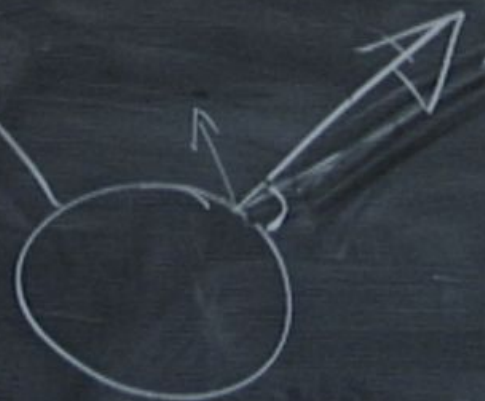
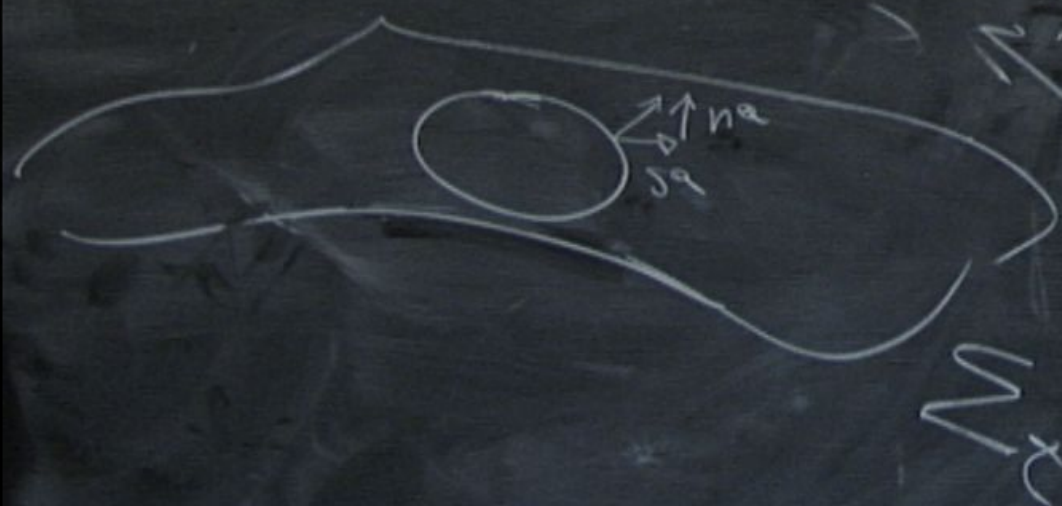
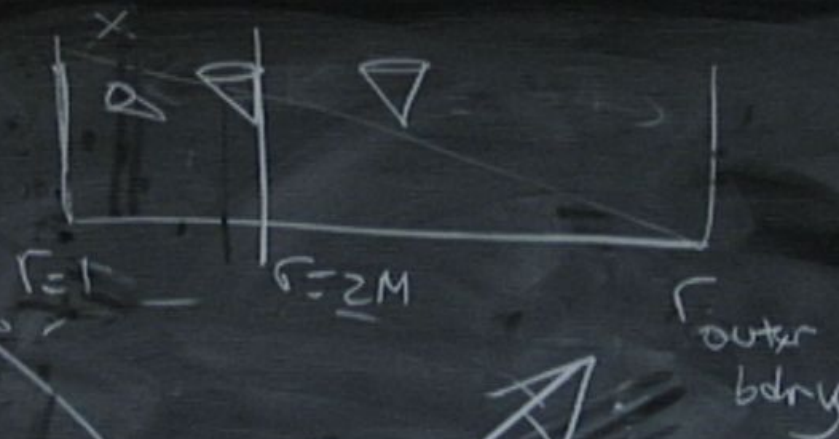
IN $3+1$ DECOMPOSITION, METRIC h_{ab} , h^{ab} IS INDUCED ON THE 2-SURFACE BY PROJECTION

$$h^{ab} = \gamma^{ab} - s^a s^b = g^{ab} + n^a n^b - s^a s^b$$

ONE SHOW (EXERCISE) THAT $\nabla_a u^a$ IS A "2-VECTOR";
 I.E. IS INTRINSIC TO 2-SURFACE; I.E. DOES NOT DEPEND ON
 HOW 2-SURFACE IS EMBEDDED IN Σ , THEN

$$\begin{aligned} \nabla_a u^a &= g^{ab} \nabla_a u_b = h^{ab} \nabla_a u_b \\ &= h^{ab} \nabla_a (s_b + n_b) \end{aligned}$$

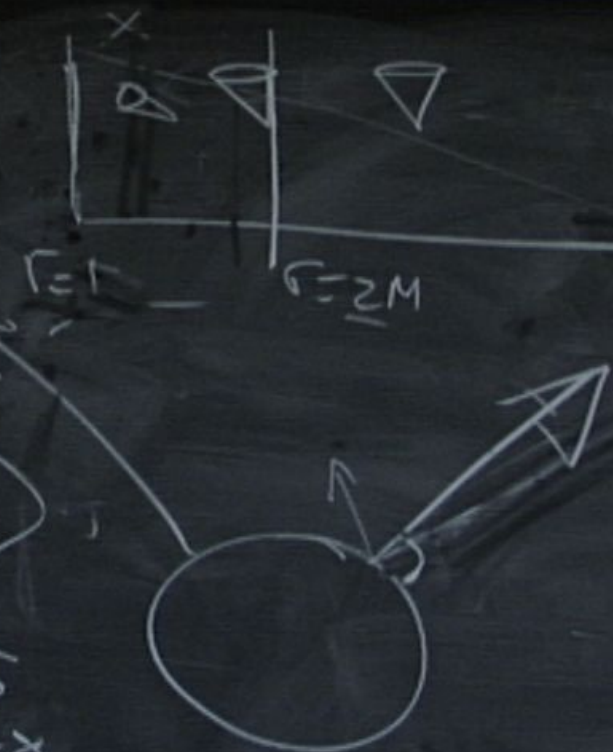
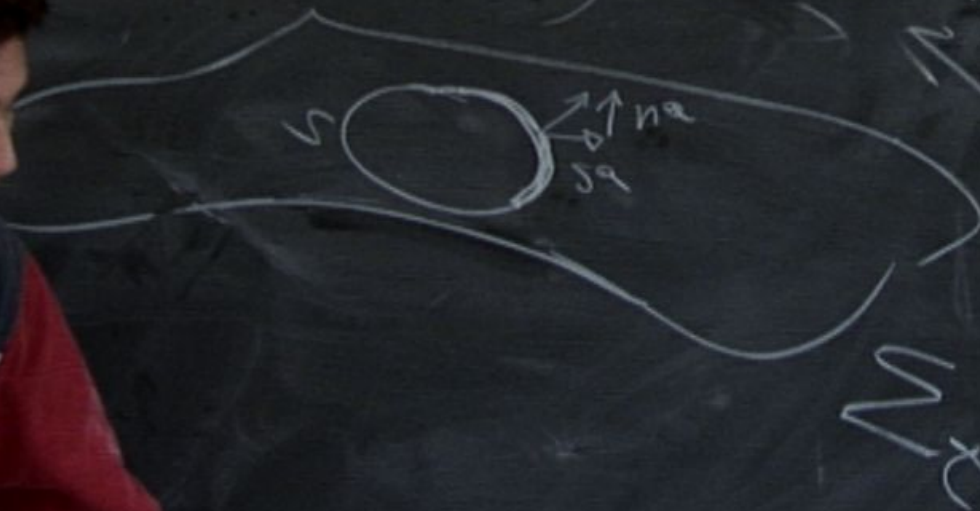
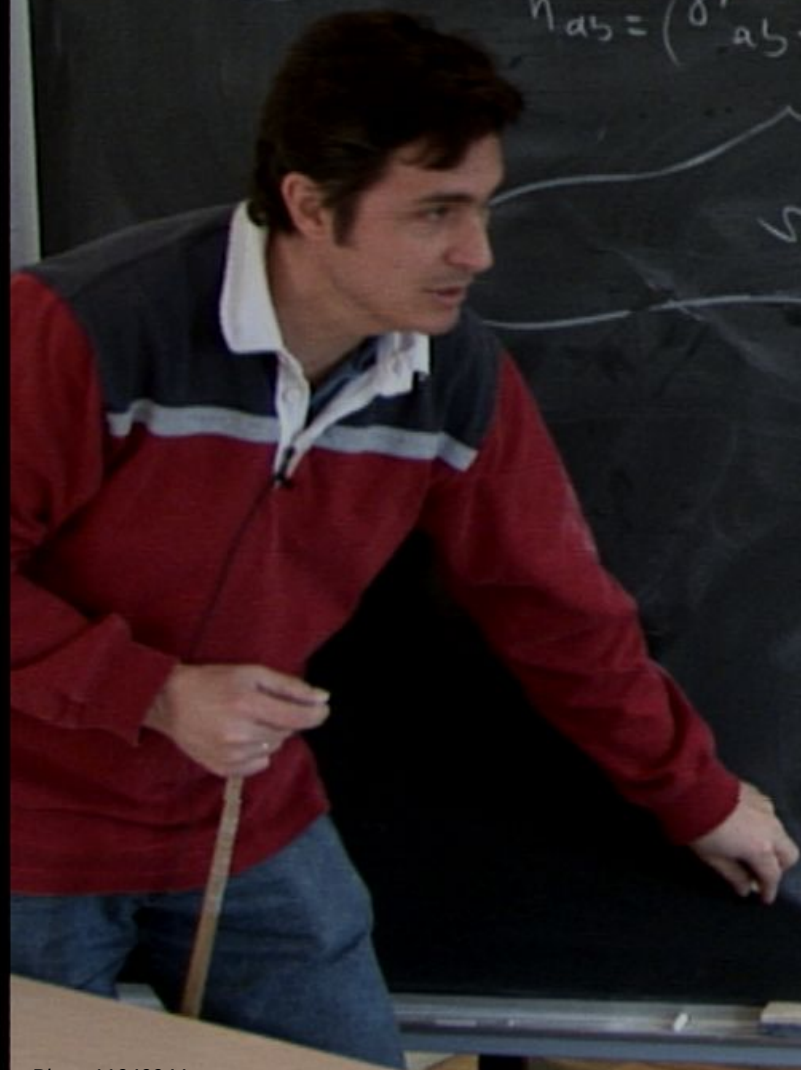
$$\delta_{ab} = \frac{1}{a_5} = g_{ab} + n_a n_b$$



$$4D \rightarrow \Sigma \quad \delta_{ab} = \frac{1}{a_5} = g_{ab} + h_a n_b$$

$$\Sigma \rightarrow S$$

$$h_{ab} = (\delta_{ab} - S_a S_b)$$





100%



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PHY 327MSPHERICAL SYMMETRY

①

THUS, OUR TRAPPED SURFACE (ΔH) EQUATION IS

$$D^a s_a - K + s^a s^b K_{ab} = 0$$

(47)

AND ARGUING AS WE DID FOR THE 3+1 EQNS WE
GET A VALID COMPONENT PART OF THIS EQUATION
BY TAKING $a \rightarrow i, b \rightarrow j$

$$D^i s_i - K + s^i s^j K_{ij} = 0$$

(48)

SPECIALIZING NOW TO SPHERICAL SYMMETRY

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NOW, RECALL EVOL. EQN (34) FOR b

$$\dot{b} = -\alpha b K^0 + \beta (rb)'$$

$$\rightarrow K^0 = -(\alpha b)^{-1} (\dot{b} - \beta (rb)')$$

$$rK^0 = -(\alpha b)^{-1} ((rb) - \beta (rb)')$$

SO (49) CAN BE REWRITTEN AS

$$(rb) + \left(\frac{\alpha}{a} - \beta \right) (rb)' = 0 \quad (50)$$

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NOW, RECALL EVOL. EQN (34) FOR b

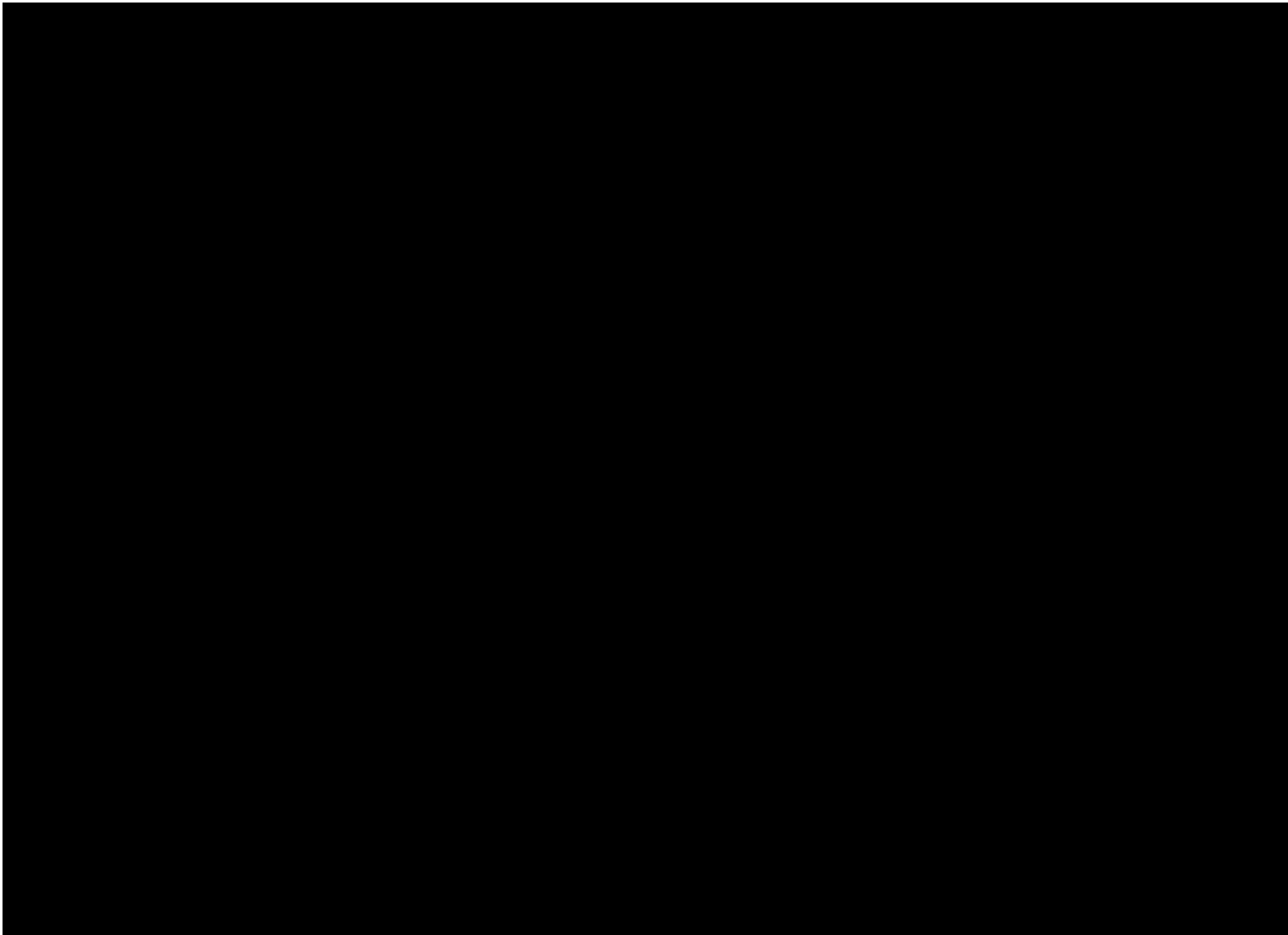
$$\dot{b} = -\alpha b |K^0| + \frac{\beta}{r} (rb)'$$

$$\rightarrow |K^0| = -(\alpha b)^{-1} \left(\dot{b} - \frac{\beta}{r} (rb)' \right)$$

$$r|K^0| = -(\alpha b)^{-1} \left((rb)' - \beta (rb)' \right)$$

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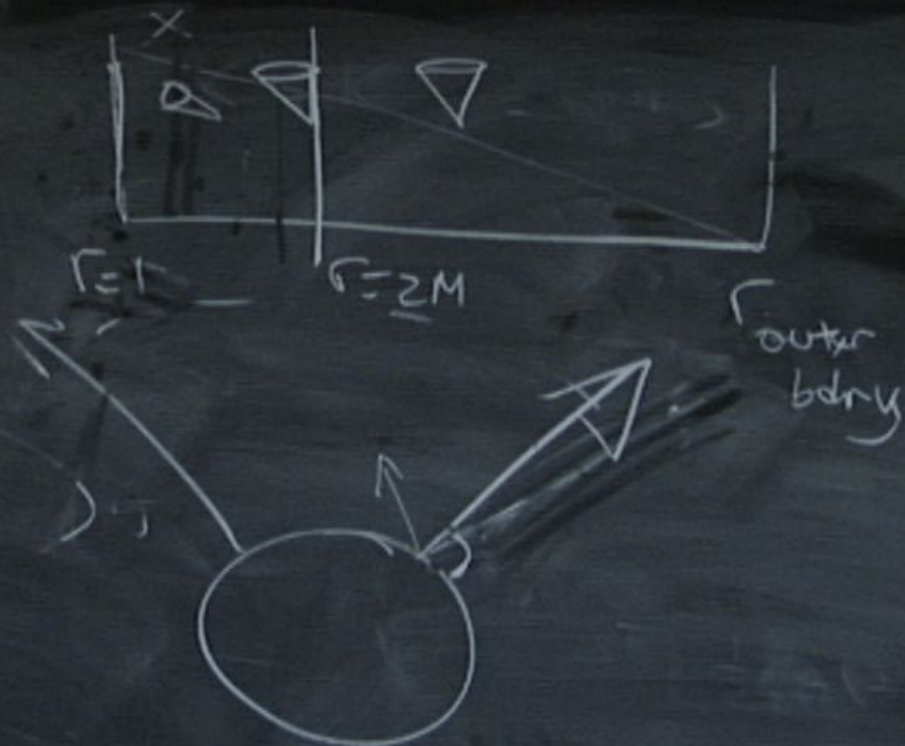
Stability Analysis

- One of the most frustrating/fascinating features of FD solutions of time dependent problems: discrete solutions often “blow up”—e.g. floating-point overflows are generated at some point in the evolution
- ‘Blow-ups’ can sometimes be caused by legitimate (!) “bugs”—i.e. an incorrect implementation—at other times it is simply the *nature of the FD scheme* which causes problems.
- Are thus lead to consider the *stability* of solutions of difference equations
- Again consider the 1-d wave equation, $u_{tt} = u_{xx}$
- Note that it is a *linear, non-dispersive* wave equation
- Thus the “size” of the solution does *not* change with time:

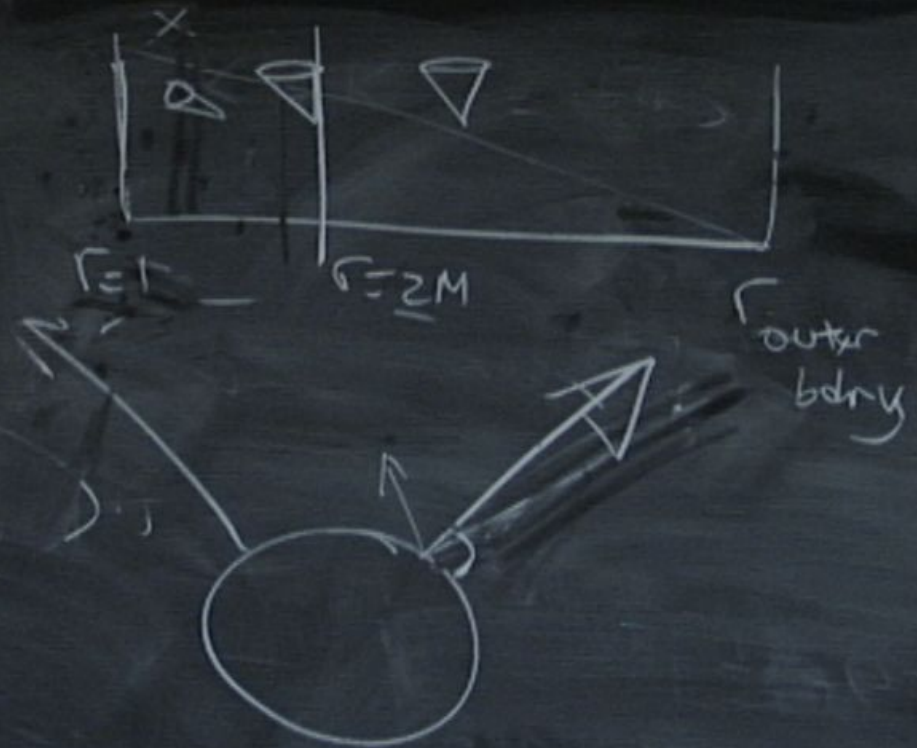
$$\|u(x, t)\| \sim \|u(x, 0)\|, \quad (95)$$

where $\|\cdot\|$ is an suitable norm, such as the L_2 norm:

$\| U(\epsilon) \|$



$$\|u(t)\| \sim \|u(t=0)\|$$

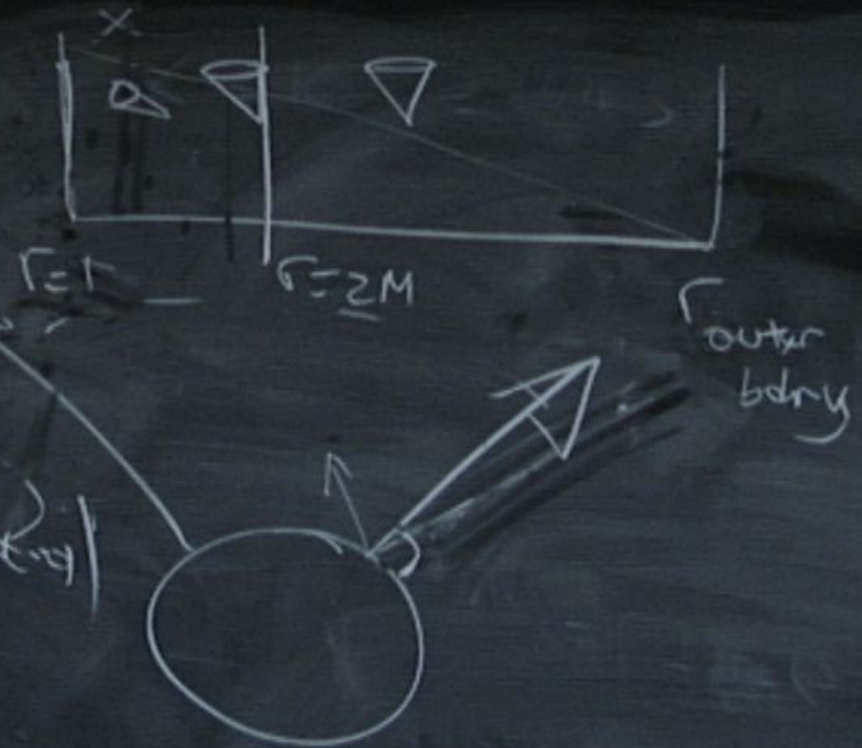


$$\|u(t)\| \sim \|u(t=0)\|$$

1) Unique soln

$$u_1(t=0), u_2(t=0)$$

$$|u_2(t) - u_1(t)| \sim |u_2(t=0) - u_1(t=0)|$$

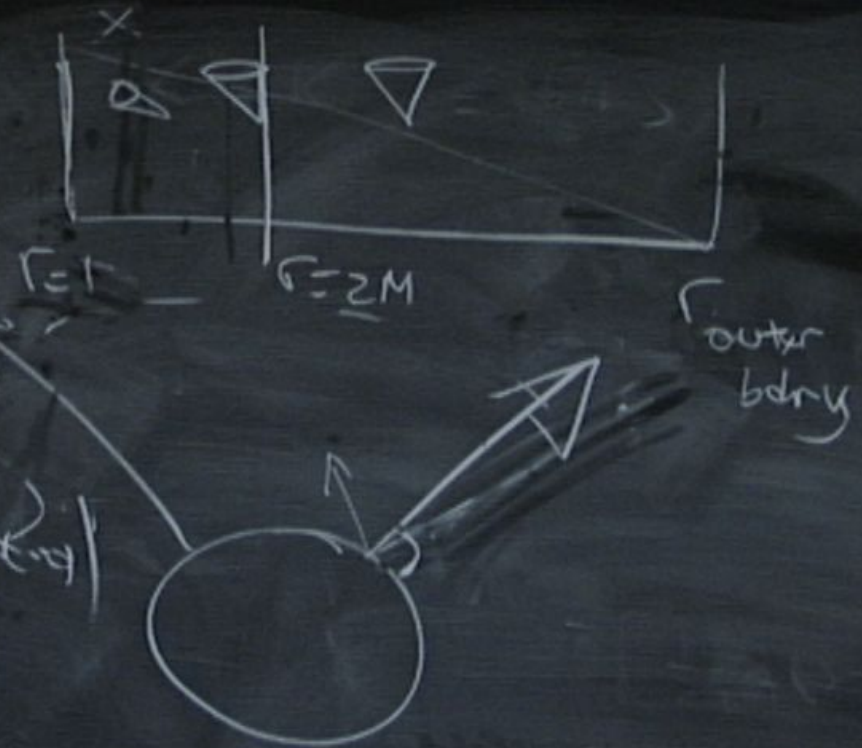


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Stability Analysis

- Will use the property captured by (95) as working definition of stability.
- In particular, if you believe (95) is true for the wave equation, then you believe the wave equation is stable.
- Fundamentally, if FDA approximation *converges*, then expect the same behaviour for the difference solution:

$$\|u_j^n\| \sim \|u_j^0\|. \quad (97)$$



- FD solution constructed by *iterating in time*, generating

$$u_j^0, u_j^1, u_j^2, u_j^3, u_j^4, \dots$$

in succession, using the FD equation

$$u_j^{n+1} = 2u_j^n - u_j^{n-1} + \lambda^2 \left(u_{j+1}^n - 2u_j^n + u_{j-1}^n \right).$$

Stability Analysis

- Not guaranteed that (97) holds for all values of $\lambda \equiv \Delta t / \Delta x$.
- For certain λ , have

$$\|u_j^n\| \gg \|u_j^0\|,$$

and for those λ , $\|u^n\|$ diverges from u , even (especially!) as $h \rightarrow 0$ —that is, the difference scheme is *unstable*.



- For many wave problems (including all linear problems), given that a FD scheme is *consistent* (i.e. so that $\hat{\tau} \rightarrow 0$ as $h \rightarrow 0$), *stability is the necessary and sufficient condition for convergence* (Lax's theorem).

Heuristic Stability Analysis

- Write general time-dependent FDA in the form

$$\mathbf{u}^{n+1} = \mathbf{G}[\mathbf{u}^n], \quad (98)$$

- $\mathbf{G}$ is some *update operator* (linear in our example problem)
- $\mathbf{u}$ is a column vector containing sufficient unknowns to write the problem in first-order-in-time form.
- Example: introduce new, auxiliary set of unknowns, v_j^n , defined by

$$v_j^n = u_j^{n-1},$$

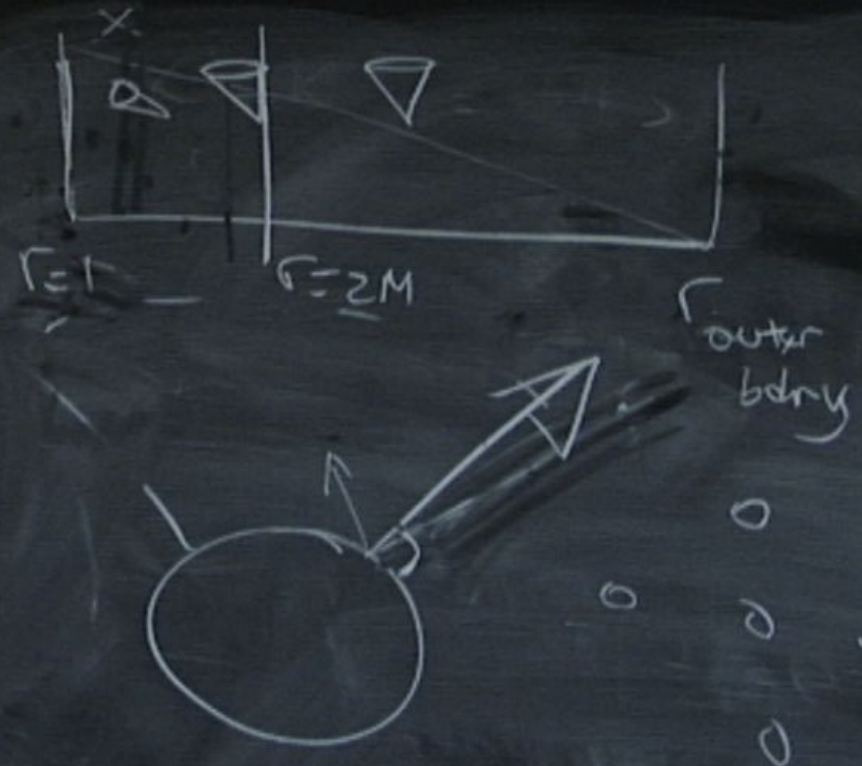
then can rewrite differenced-wave-equation (16) as

$$u_j^{n+1} = 2u_j^n - v_j^n + \lambda^2 (u_{j+1}^n - 2u_j^n + u_{j-1}^n), \quad (99)$$

$$v_j^{n+1} = u_j^n, \quad (100)$$

$$u_t = u_x$$

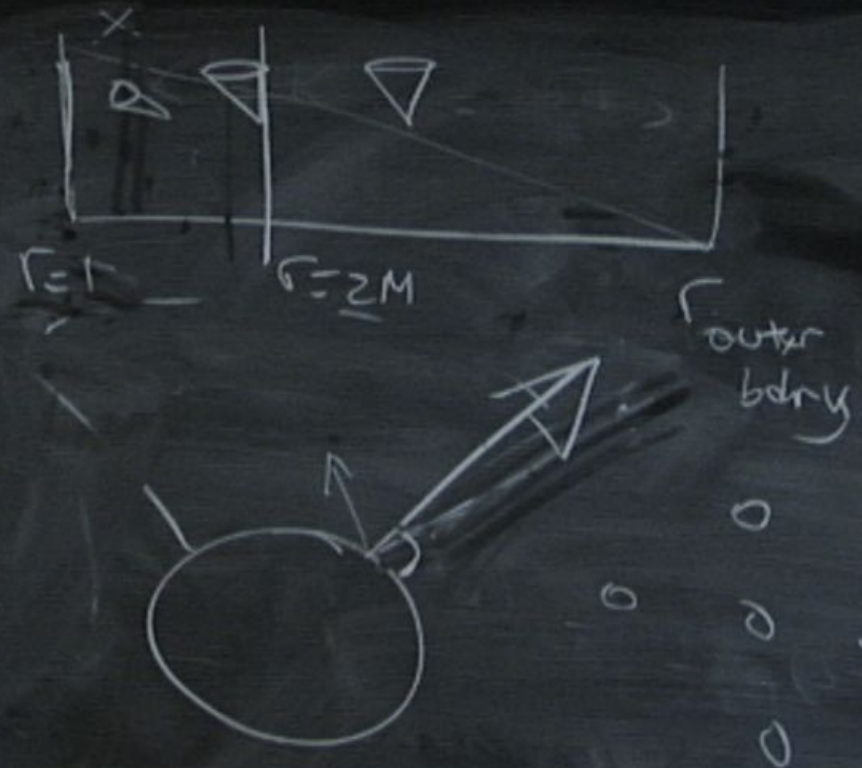
$$\frac{u_{j+1}^{n+1} - u_j^n}{\Delta t} = \frac{u_{j+1}^n - u_{j-1}^n}{2\Delta t}$$



$$u_{,t} = u_{,x}$$

$$\frac{u_J^{n+1} - u_J^n}{\Delta t} = \frac{u_{J+1}^n - u_{J-1}^n}{2\Delta t}$$

$$u_J^{n+1} = \left\{ u_J^n, u_{J+1}^n, u_{J-1}^n \right\}$$



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Heuristic Stability Analysis

- If difference scheme is to be stable, must have

$$|\mu_k| \leq 1 \quad k = 1, 2, \dots, J \quad (104)$$

(Note: μ_k will be complex in general, so $|\mu|$ denotes the complex modulus, $|\mu| \equiv \sqrt{\mu\mu^*}$.)

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$$u_{j,t} = \lambda u_{j,x}$$

$$\frac{u_j^n}{\Delta t} = \frac{u_{j+1}^n - u_{j-1}^n}{2\Delta x}$$

$$\Delta t = \mu \Delta x$$

$$J = \left\{ u_j^n, u_{j+1}^n, u_{j-1}^n \right\}$$

$$u_{i,t} = u_{i,t-1} + \frac{u_{i,t-1}^n - u_{i,t-1}^{n-1}}{\Delta t}$$

$$u_{i,t}^{n+1}$$

$$J =$$

$$u_{i,t} = \lambda u_{i,x}$$

$$\Delta t = \mu \Delta x \sim u = \lambda \frac{\Delta t}{\Delta x} u$$

$$u_{j,t} = u_{j,x}$$

$$\frac{u_{j,t}^{n+1} - u_{j,t}^n}{\Delta t} = u_{j,t}^n$$

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$$j-1)$$

$$u, t = u, x$$

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Von-Neumann (Fourier) Stability Analysis (Summary)

- Von-Neumann (VN) stability analysis based on the ideas sketched above
- Assumes that difference equation is linear with constant coefficients, periodic boundary conditions boundary conditions are periodic
- Can then use Fourier analysis: difference operators in real-space variable $x \longrightarrow$ algebraic operations in Fourier-space variable k
- VN applied to wave-equation example shows that must have

$$\lambda \equiv \frac{\Delta t}{\Delta x} \leq 1,$$

for stability of scheme (16).

- Condition is often called the CFL condition—after Courant, Friedrichs and Lewy who derived it in 1928

- This type of instability has “physical” interpretation, often summarized by the statement: *the numerical solution of the wave equation is unstable if the time step is too large*



$$u_{1,t} = \lambda u_{1,x}$$

$$\Delta t = \mu \Delta x \quad \sim u = \lambda \mu u$$



$$u_{1,t} = \lambda u_{1,x}$$

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