

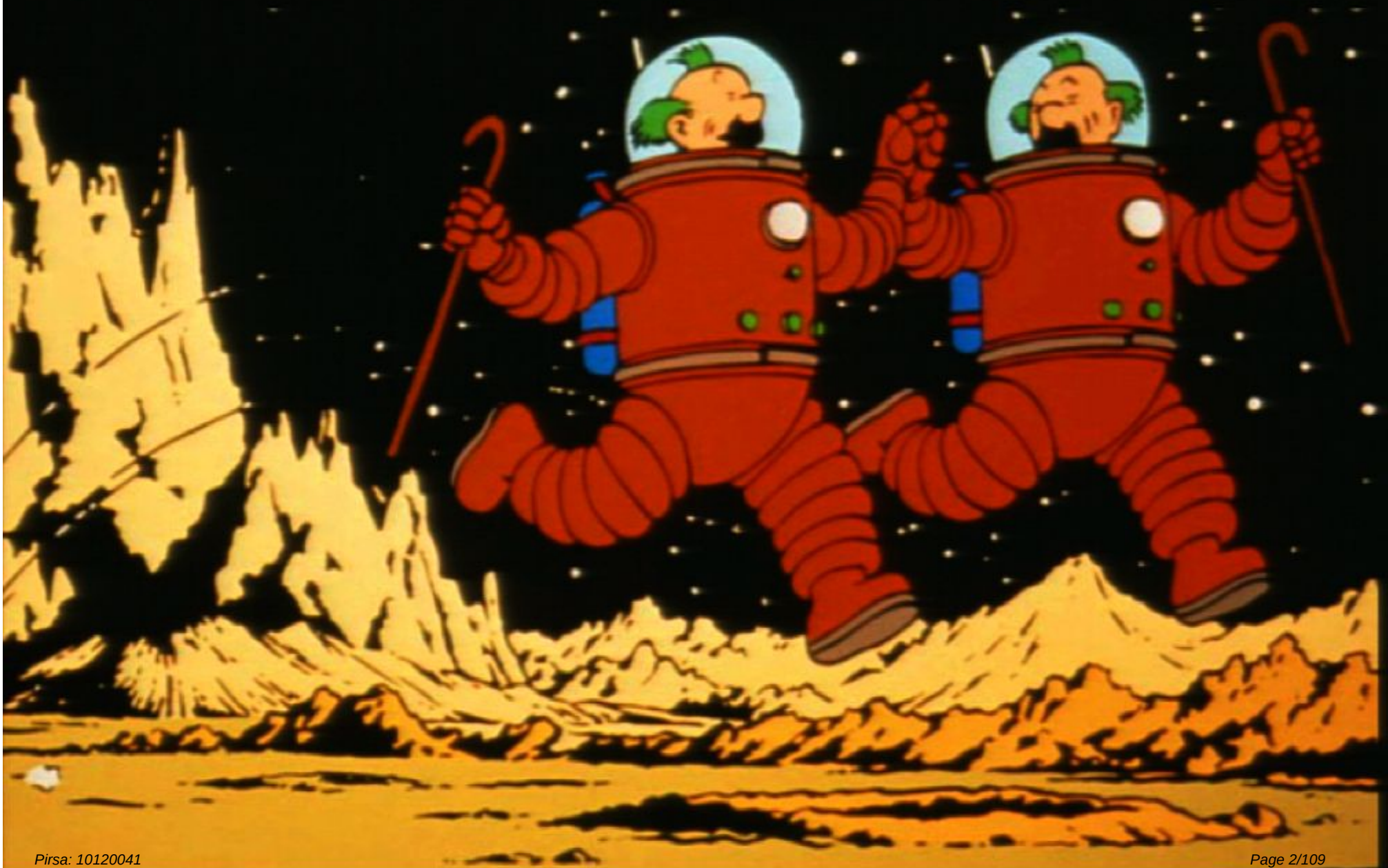
Title: The Fastest Decay in the Landscape

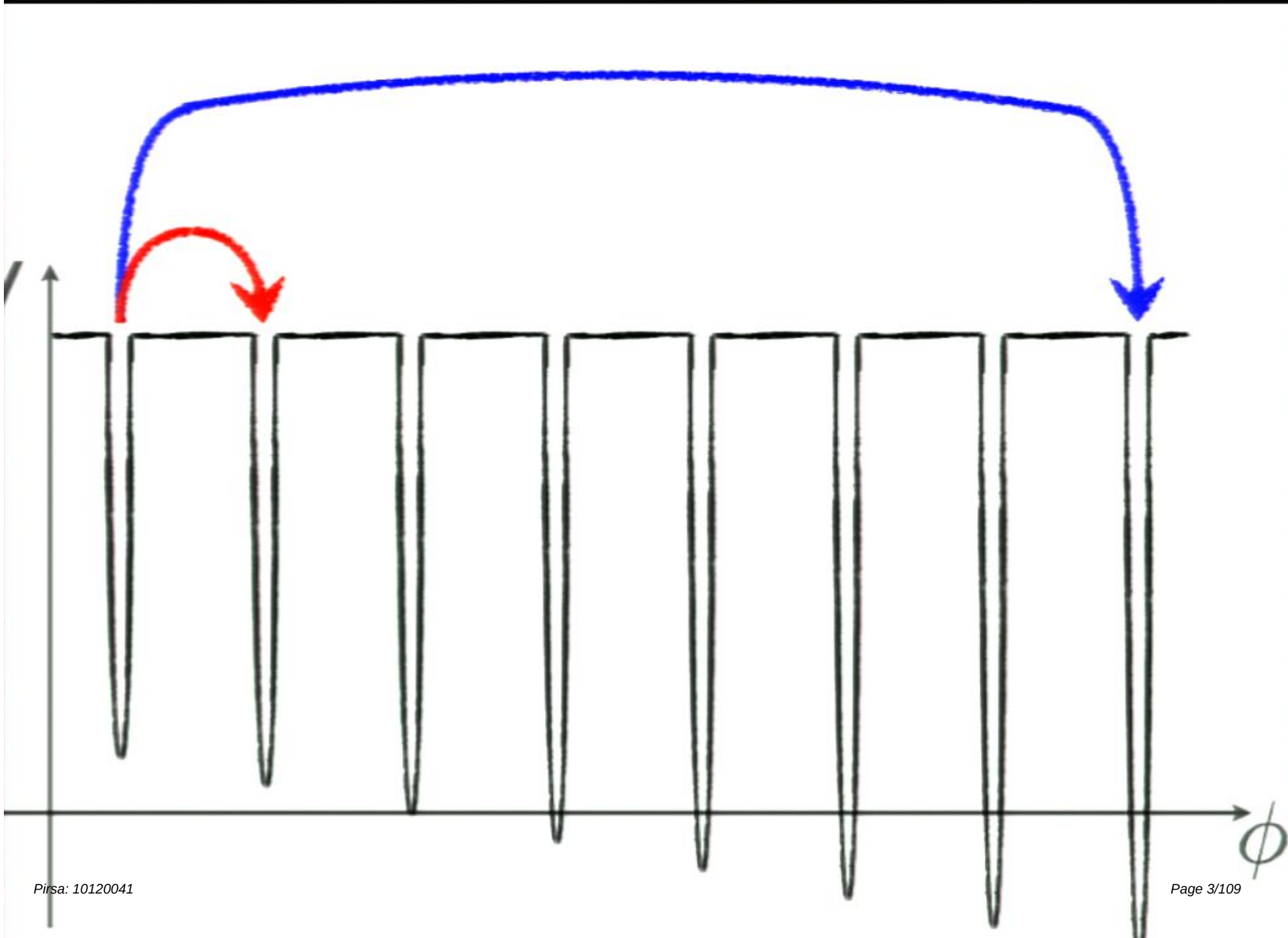
Date: Dec 09, 2010 01:00 PM

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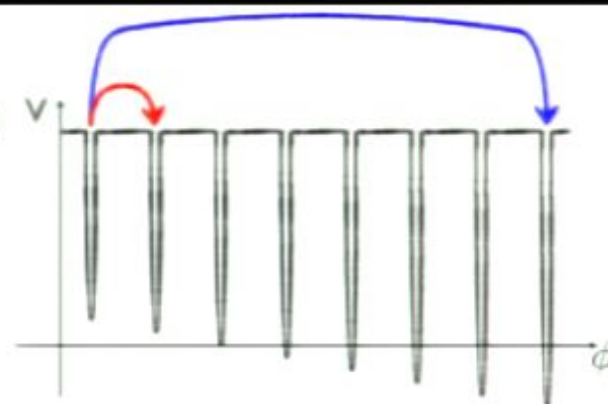
Abstract: Theories with extra dimensions naturally give rise to a large landscape of vacua stabilized by flux. I will show that the fastest decay is a giant leap to a wildly distant minimum, in which many different fluxes discharge at once. Indeed, the fastest decay is frequently the giantest leap of all, where all the fluxes discharge at once, which destabilizes the extra dimensions and begets a bubble of nothing. Finally, I will discuss how these giant leaps are mediated by the nucleation of "monkey branes" that wrap the extra dimensions.

The Fastest Decay in the Landscape





This potential gives small steps



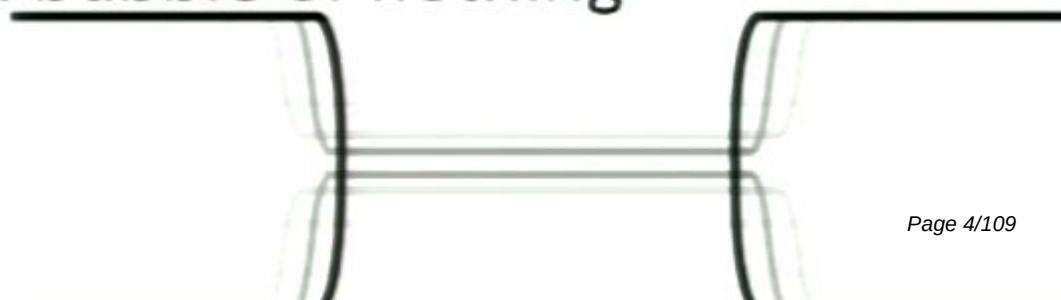
Flux compactifications with many fluxes give giant leaps

Monkey branes

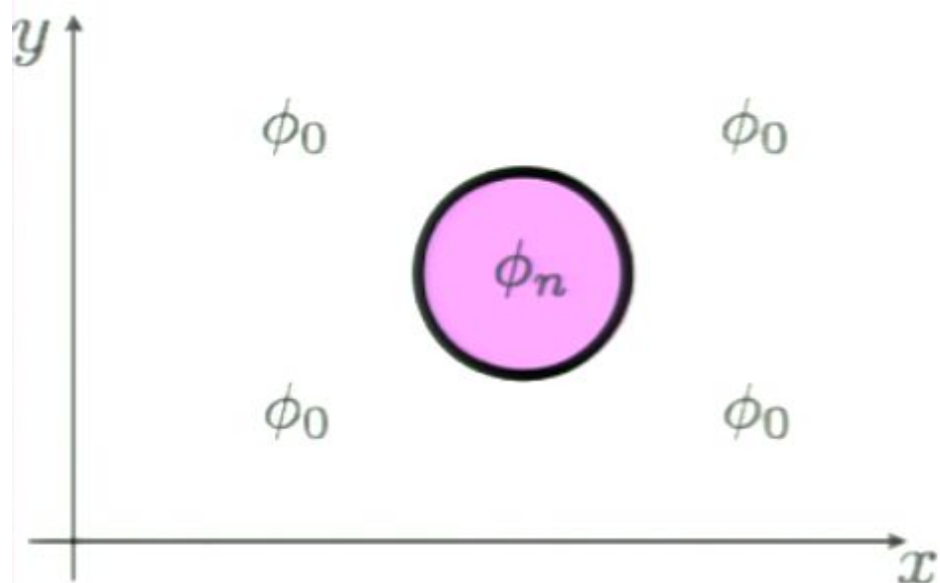
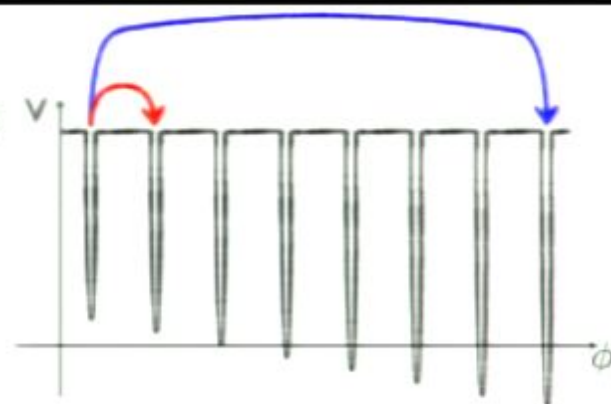


Adding back-reaction

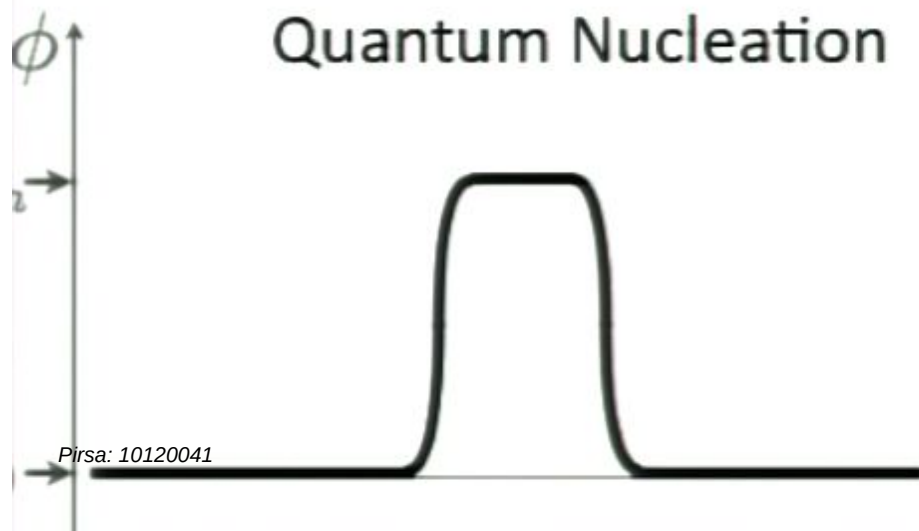
The giantest leap of all is a bubble of nothing



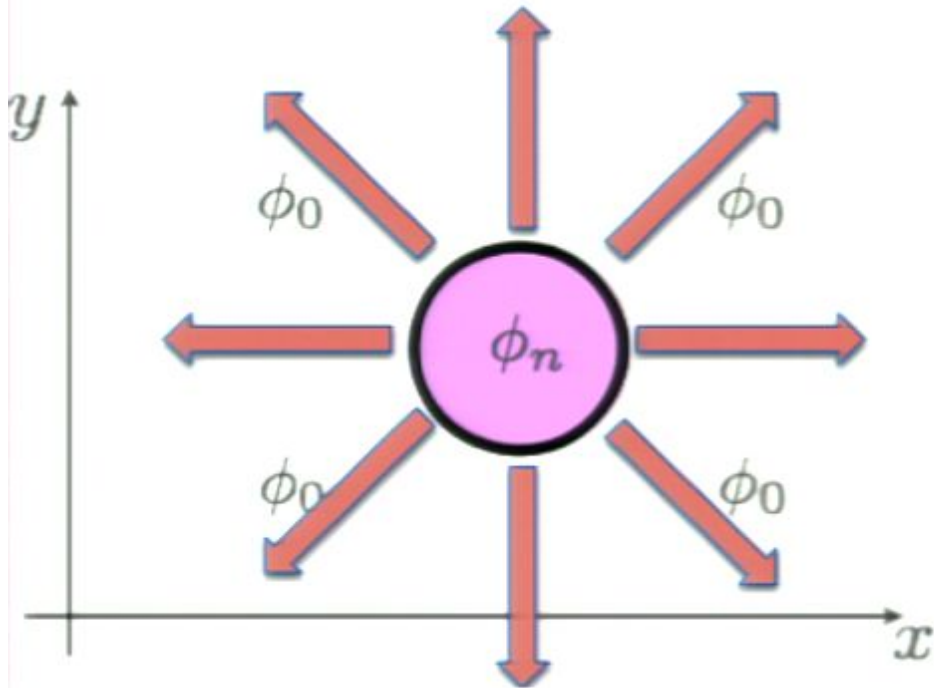
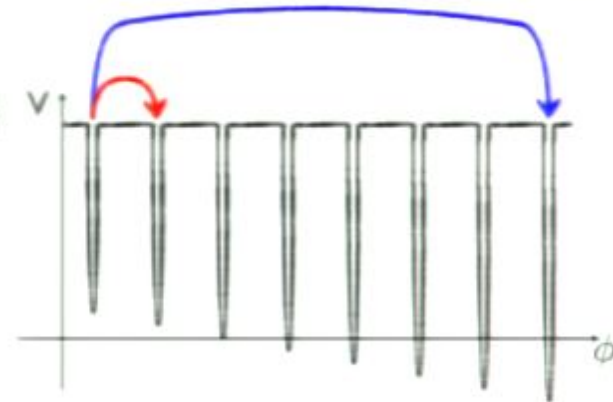
This potential gives small steps



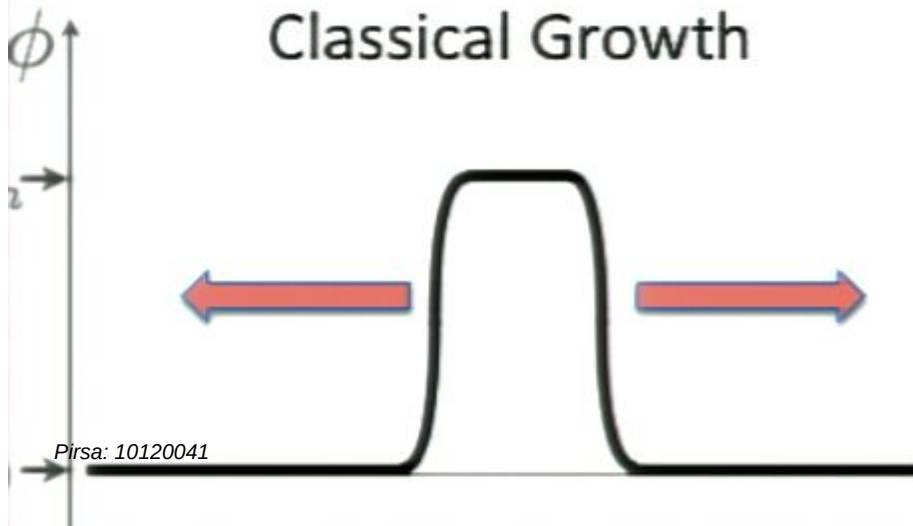
Quantum Nucleation



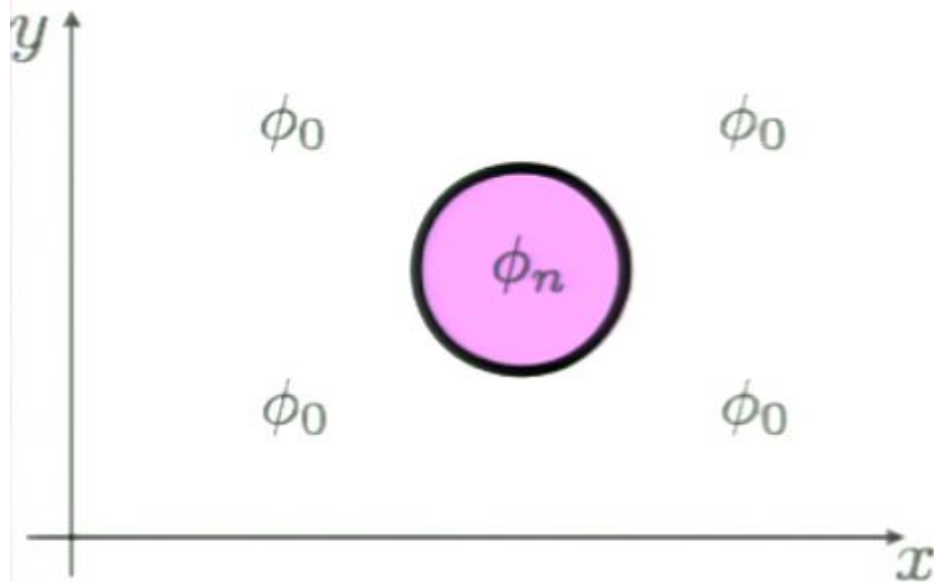
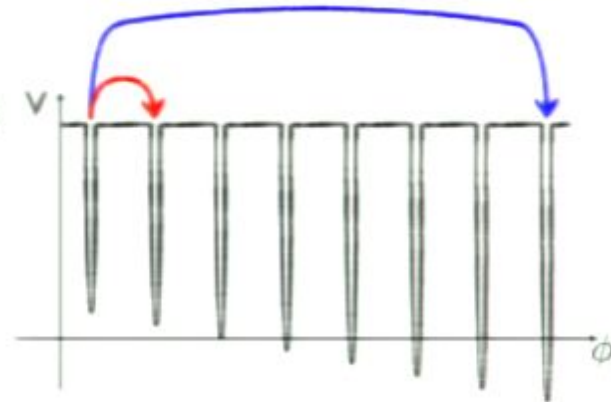
This potential gives small steps



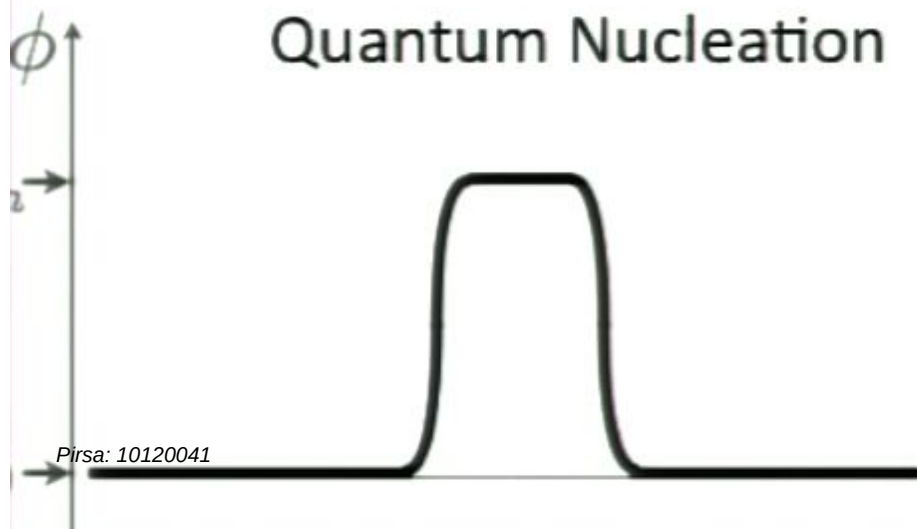
Classical Growth



This potential gives small steps



- thin wall
- flat spacetime



$$\Gamma_n \sim \exp\left[-\frac{\#}{\hbar} \frac{\sigma_n^4}{\epsilon_n^3}\right]$$

Coleman (1977)

$$\sigma_n \equiv \int_{\phi_0}^{\phi_n} d\phi \sqrt{2[V(\phi) - V(\phi_0)]}$$

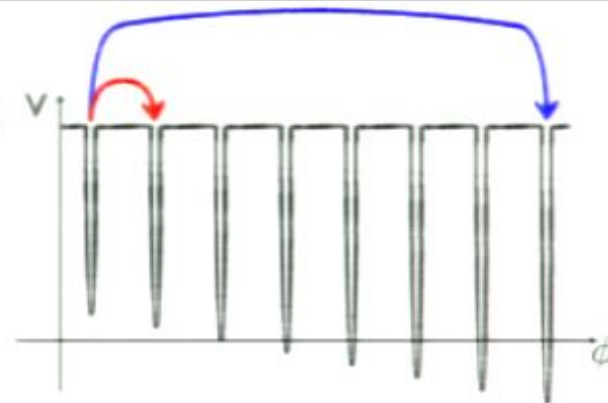
$$\epsilon_n \equiv V(\phi_0) - V(\phi_n)$$

This potential gives small steps

$$\sigma_n = n\sigma_1$$

$$\epsilon_n = n\epsilon_1$$

$$\frac{\sigma_n^4}{\epsilon_n^3} = n \frac{\sigma_1^4}{\epsilon_1^3}$$



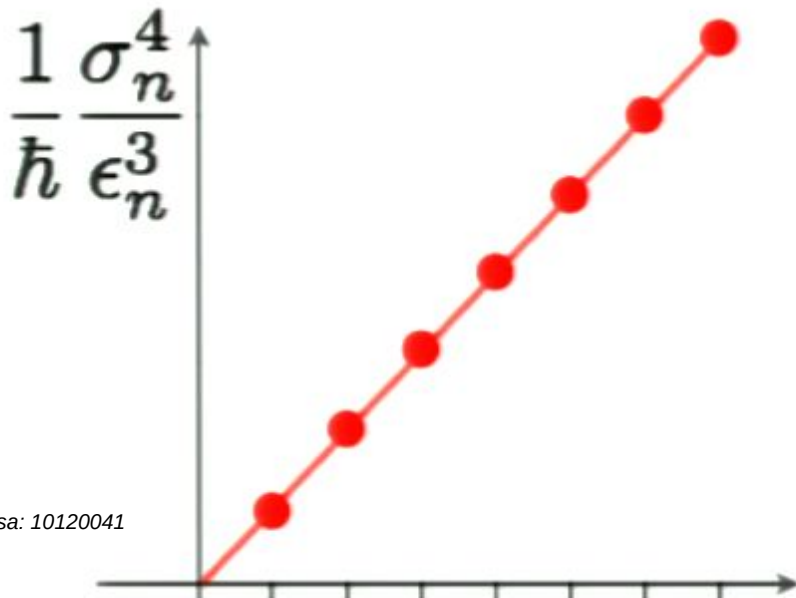
- thin wall
- flat spacetime

$$\Gamma_n \sim \exp\left[-\frac{\#}{\hbar} \frac{\sigma_n^4}{\epsilon_n^3}\right]$$

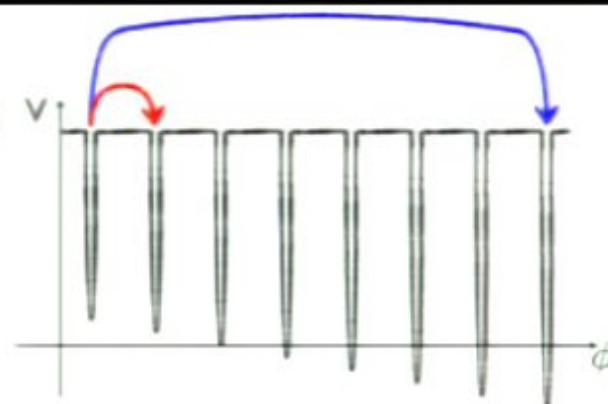
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This potential gives small steps



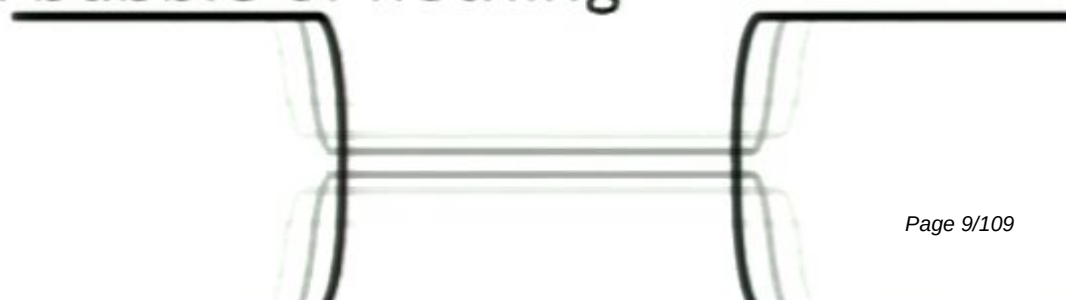
Flux compactifications with many fluxes give giant leaps

Monkey branes



Adding back-reaction

The giantest leap of all is a bubble of nothing



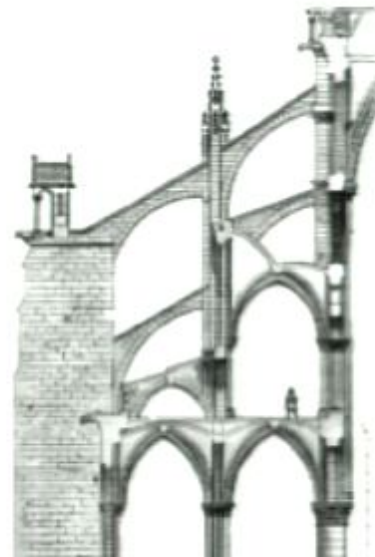
Flux Landscapes

Extra dimensions and fluxes give rise to a landscape

Flux shifts cosmological constant

$$\Lambda_{\text{eff}} = \Lambda_0 + \frac{1}{2}F^2$$

Flux stabilizes extra dimensions



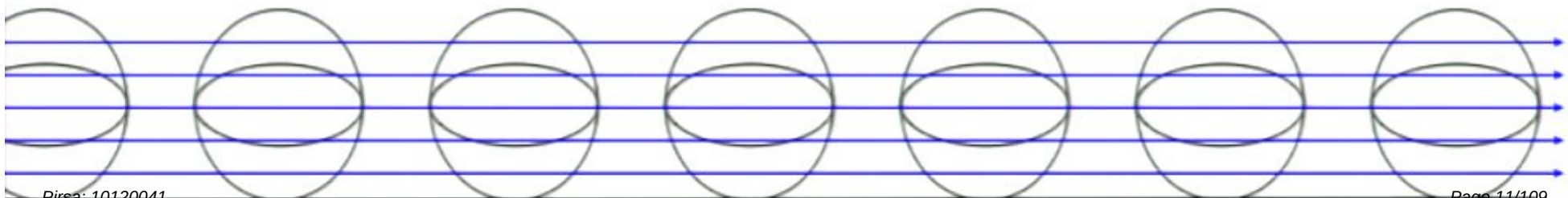
Consider flux tunneling in 1+1+2 dimensions

space time sphere



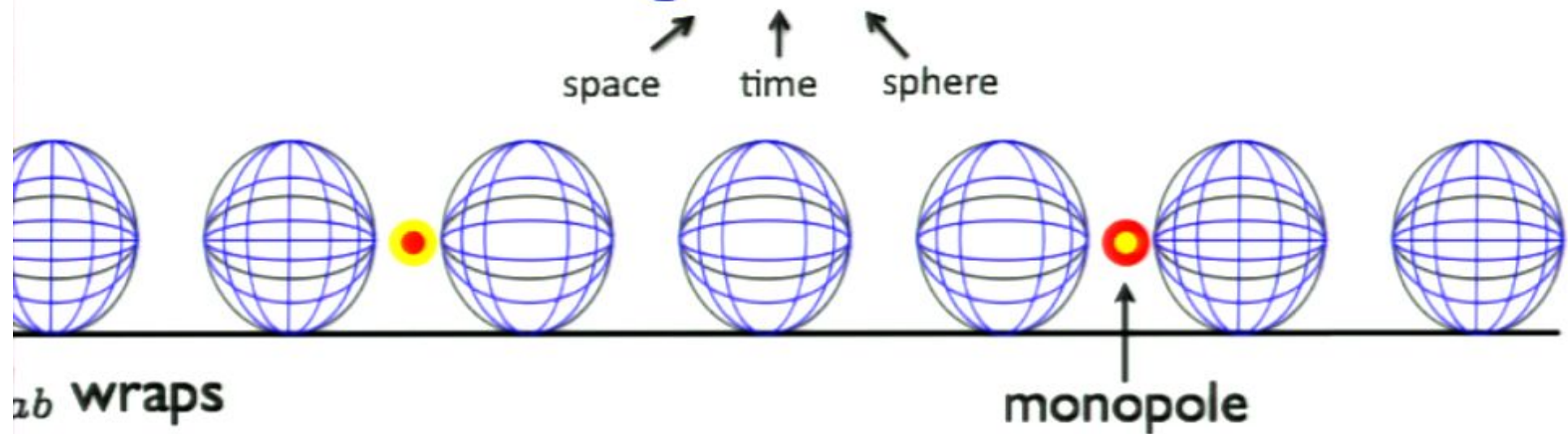
z wraps

1 units of flux N units of flux N units of flux

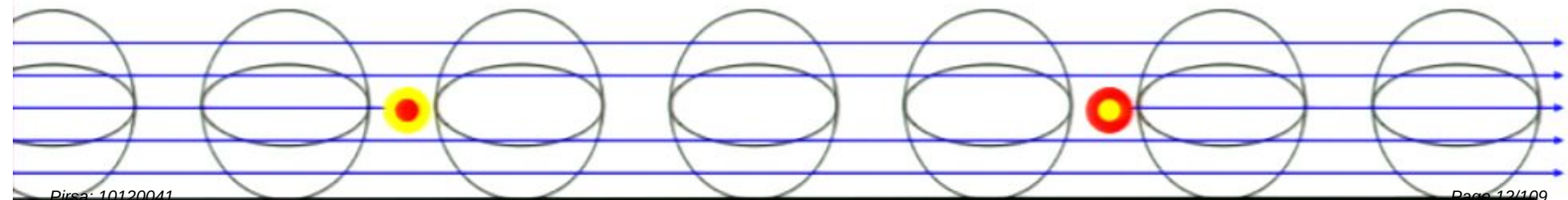


$\vec{z}$ threads

Consider flux tunneling in 1+1+2 dimensions



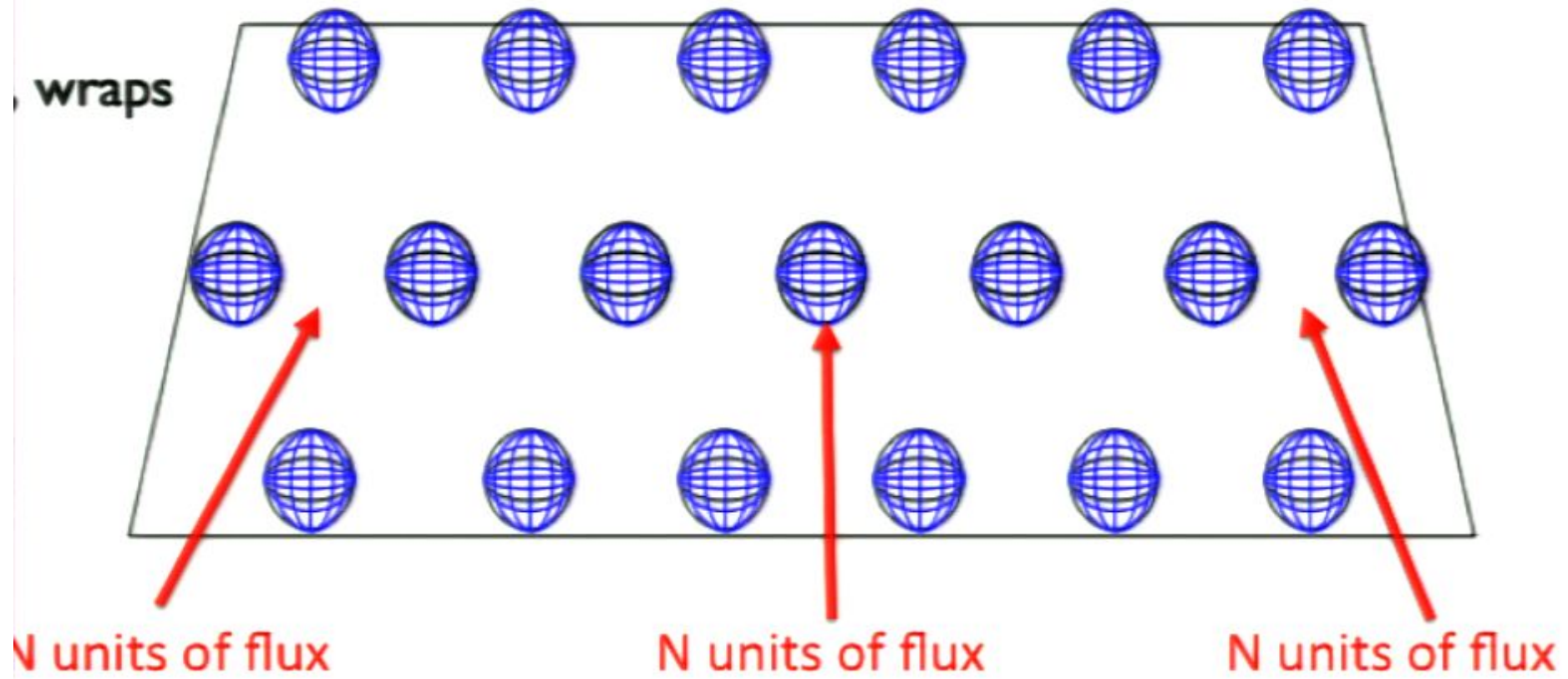
1 units of flux N-1 units of flux N units of flux



z threads

Consider flux tunneling in 3+1+2 dimensions

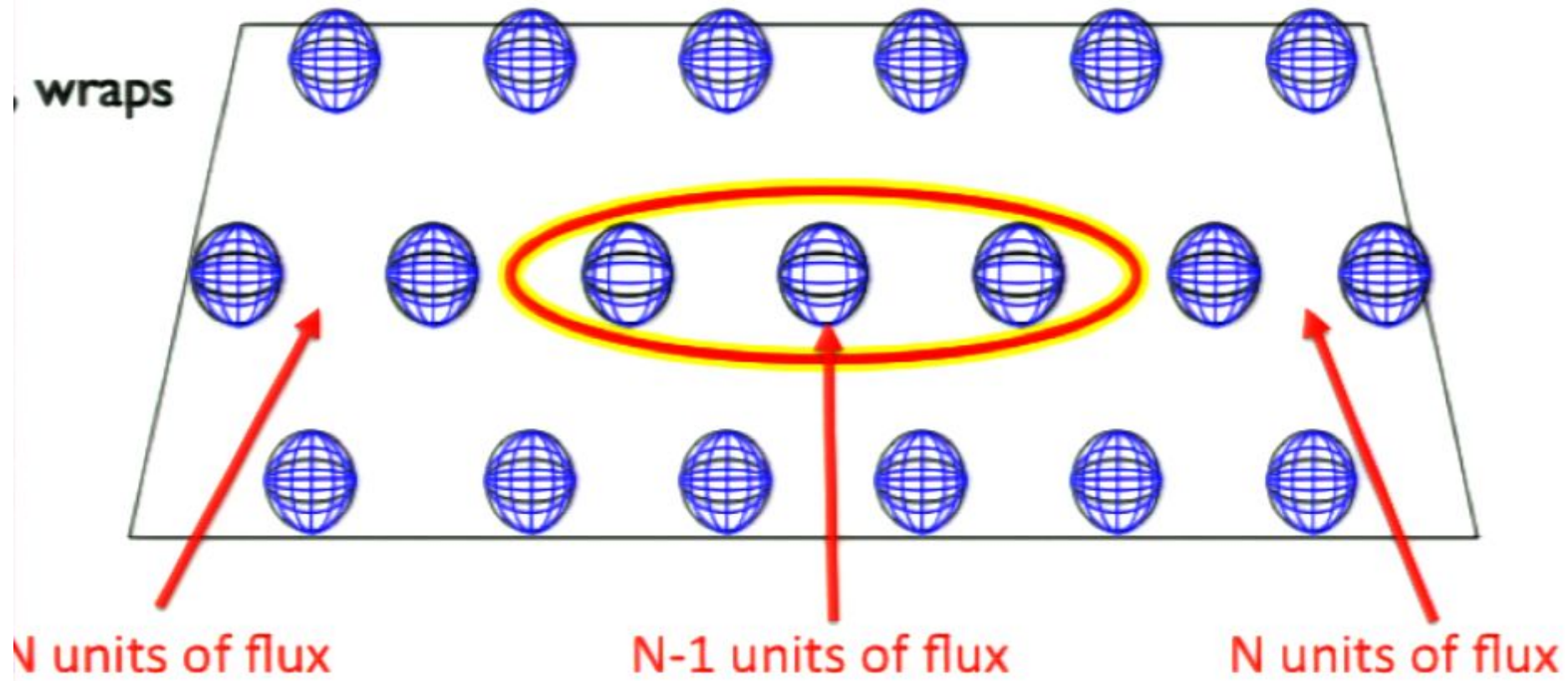
space time sphere



Consider flux tunneling in 3+1+2 dimensions

space time sphere

Nucleate an extremal black 2-branes

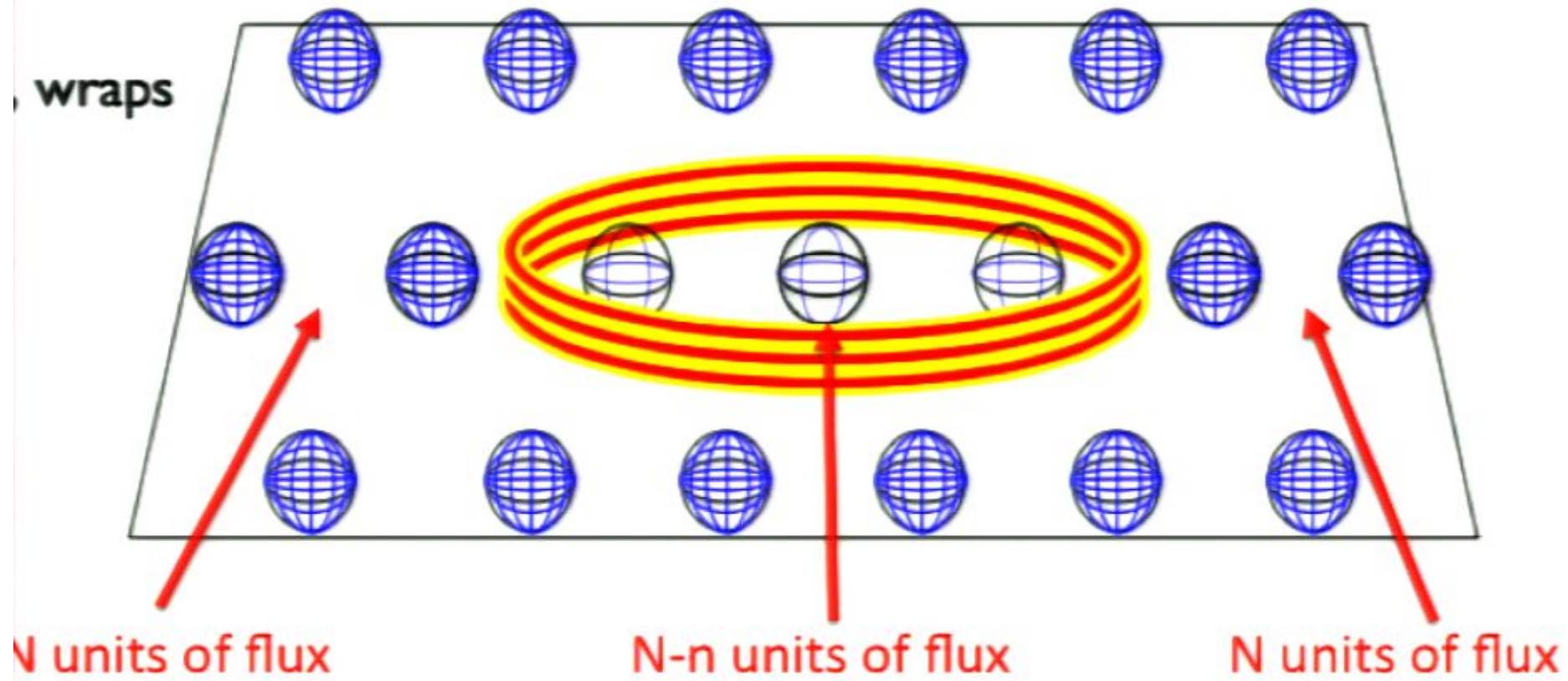


Consider flux tunneling in 3+1+2 dimensions

Giant Leaps?

space time sphere

Nucleate a stack of n 2-branes



What's the tension of the brane?

as light as possible, given its charge



A diagram showing a single yellow circle with a red center. Two horizontal arrows originate from the circle: one pointing left towards the expression $\frac{GM^2}{r}$ and one pointing right towards the expression $\frac{Q^2}{r}$.

$$\frac{GM^2}{r} \quad \leftarrow \quad \bullet \quad \rightarrow \quad \frac{Q^2}{r}$$



A diagram showing two yellow circles with red centers placed side-by-side.

$$M = Q$$

What's the tension of the brane?

as light as possible, given its charge



$$\frac{GM^2}{r} \leftarrow \text{yellow circle with red center} \rightarrow \frac{Q^2}{r}$$



$$M = Q$$





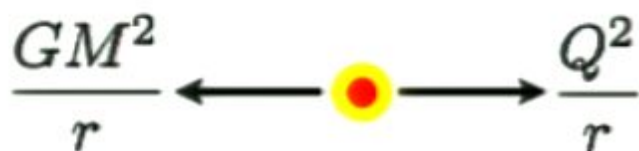
$$\frac{GM^2}{r} \leftarrow \text{yellow circle with red center} \rightarrow \frac{Q_1^2}{r} + \frac{Q_2^2}{r}$$



$$M = \sqrt{Q_1^2 + Q_2^2}$$

What's the tension of the brane?

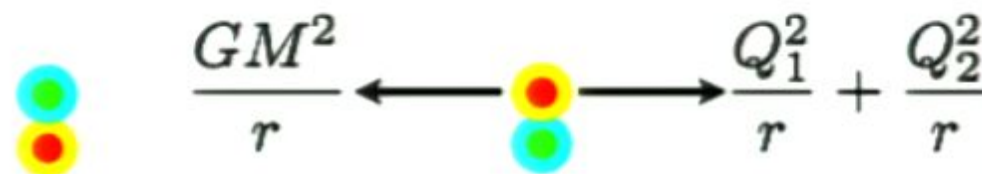
as light as possible, given its charge



$$\frac{GM^2}{r} \leftarrow \text{yellow circle} \rightarrow \frac{Q^2}{r}$$



$$M = Q$$



$$\frac{GM^2}{r} \leftarrow \text{yellow circle} \rightarrow \frac{Q_1^2}{r} + \frac{Q_2^2}{r}$$



$$M = \sqrt{Q_1^2 + Q_2^2}$$

$$T \sim \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{1/2}$$

Large landscape: many different types of fluxes

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2$$

$$T \sim \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{1/2}$$

Large landscape: many different types of fluxes

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2$$

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Two decay directions:

MONOFLUX decay

MULTIFLUX decay

large landscape: many different types of fluxes

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2$$

$$T \sim \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{1/2}$$

two decay directions:

MONOFLUX decay

1. $T_n \sim n$

MULTIFLUX decay

1. $T_n \sim \sqrt{n}$

large landscape: many different types of fluxes

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2$$

$$T \sim \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{1/2}$$

two decay directions:

MONOFLUX decay

I. $T_n \sim n$

II. $\Delta(F^2) \sim (2N - n)n$



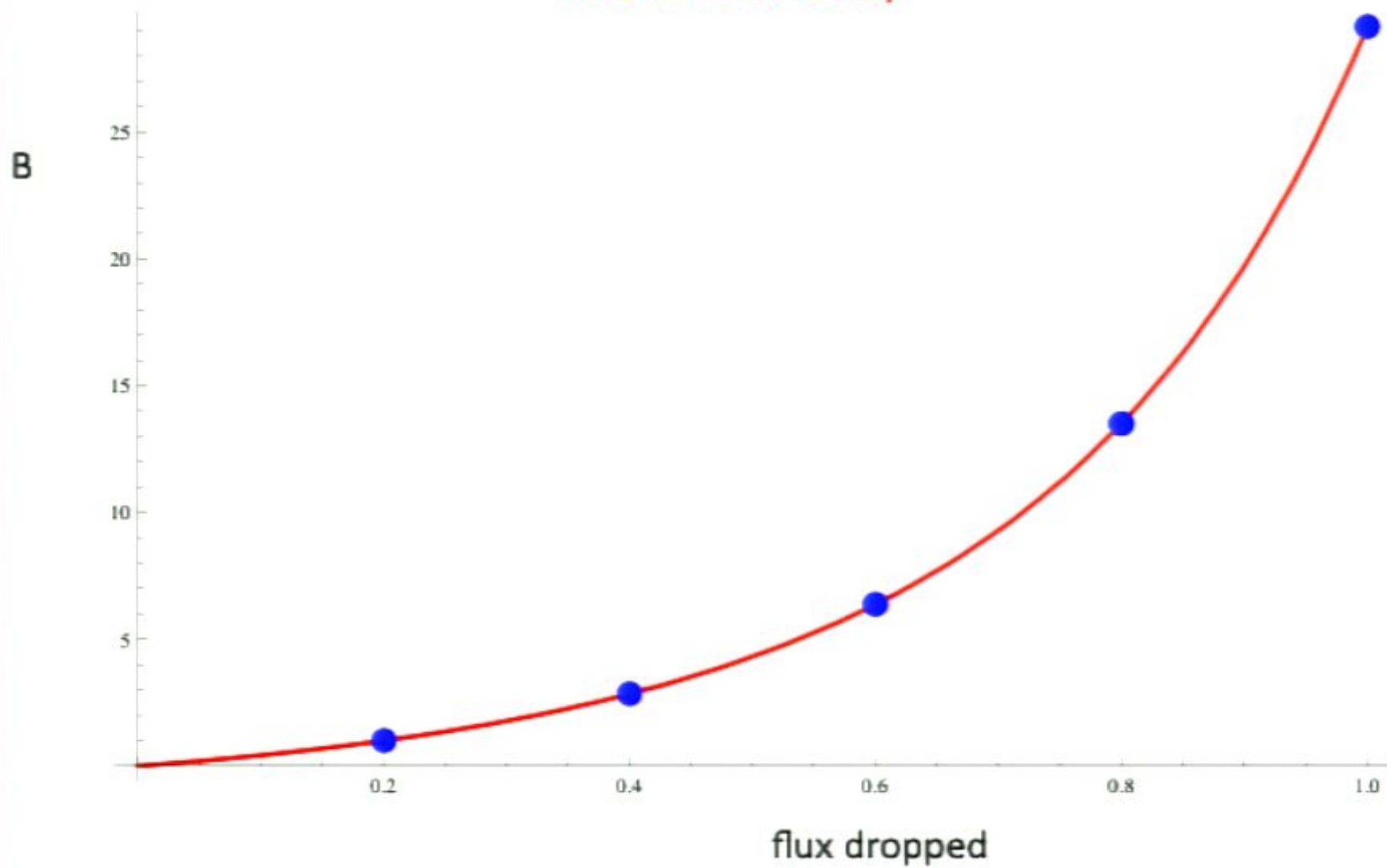
MULTIFLUX decay

I. $T_n \sim \sqrt{n}$

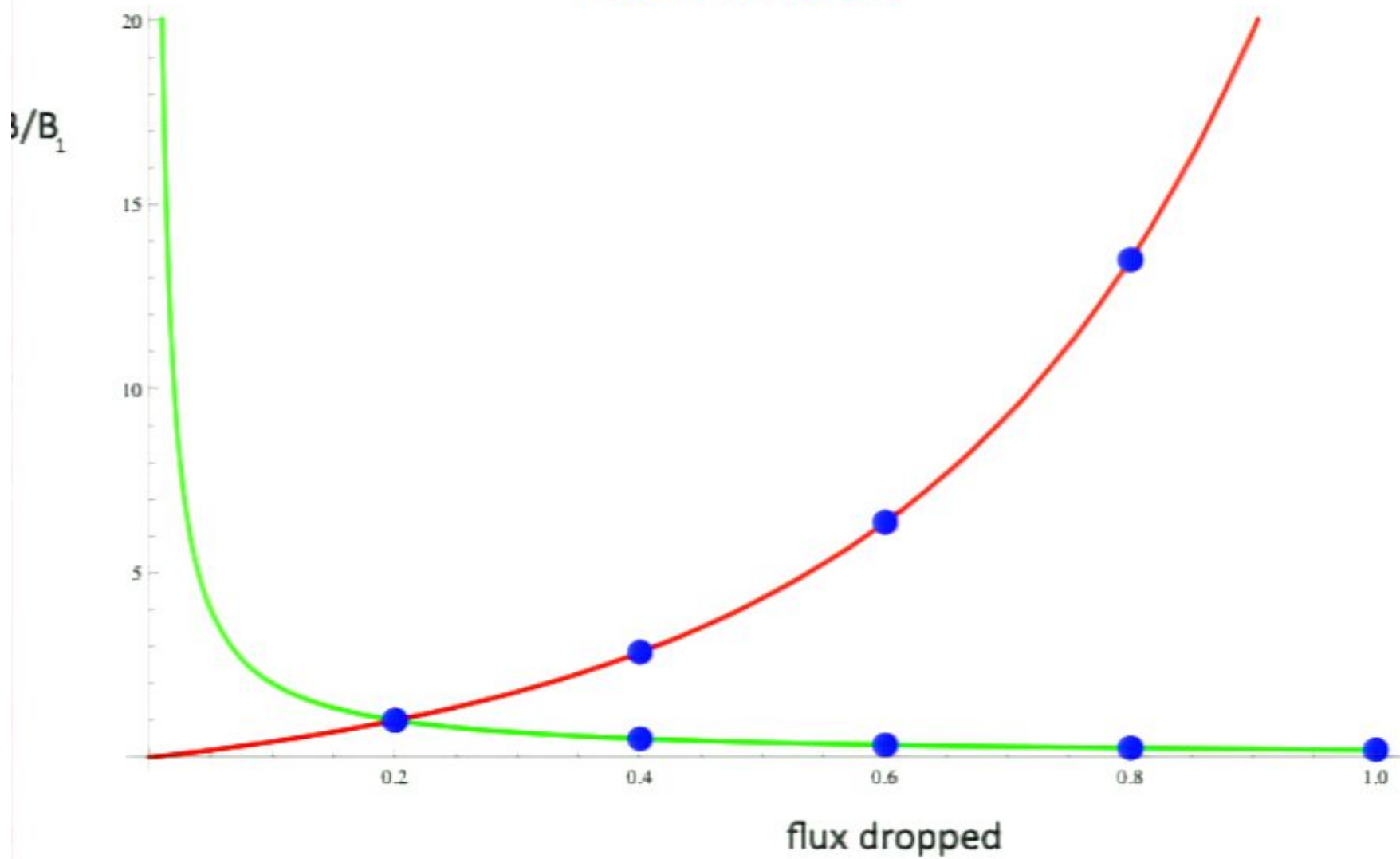
II. $\Delta(F^2) \sim n$



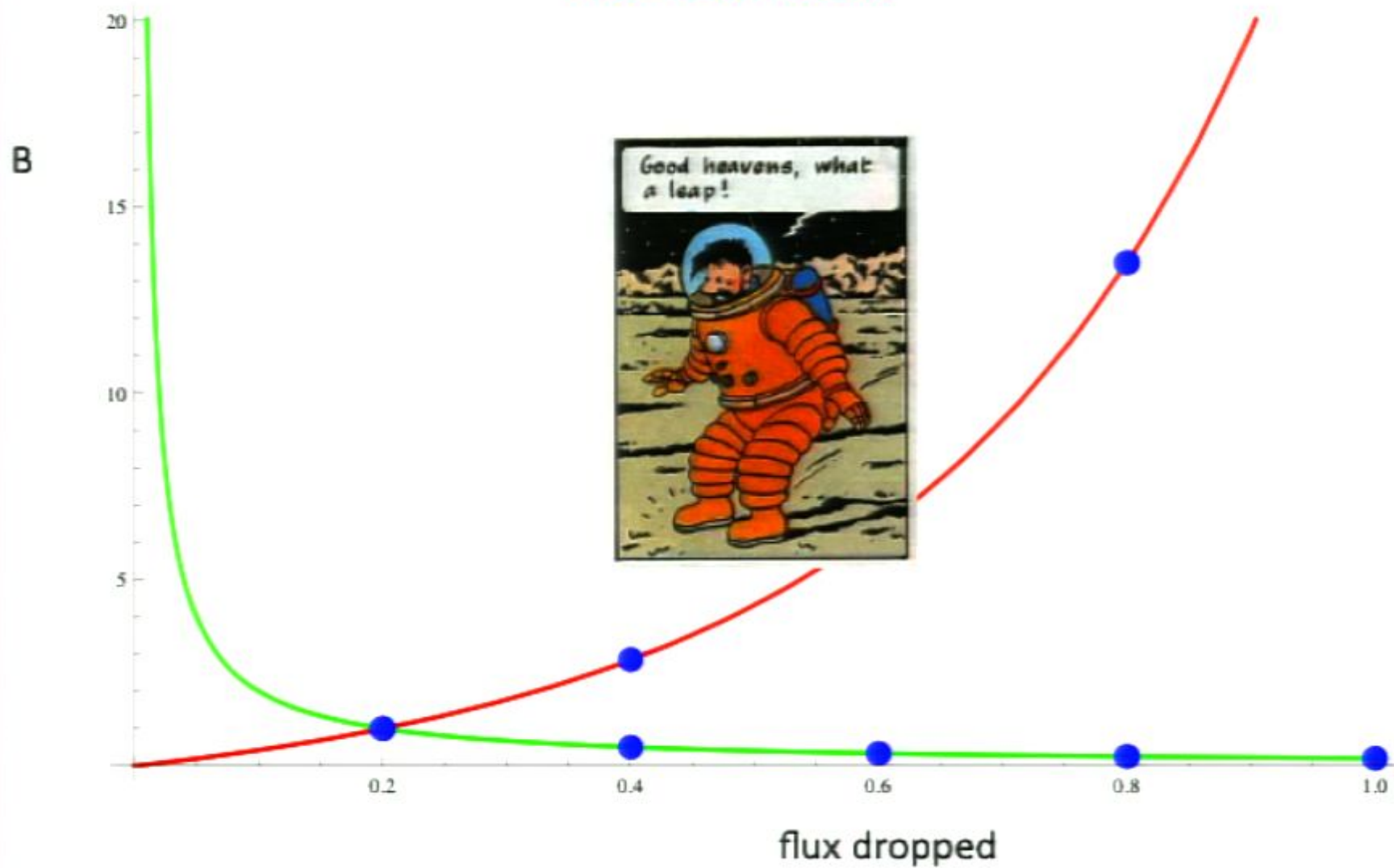
MONOFLUX decay



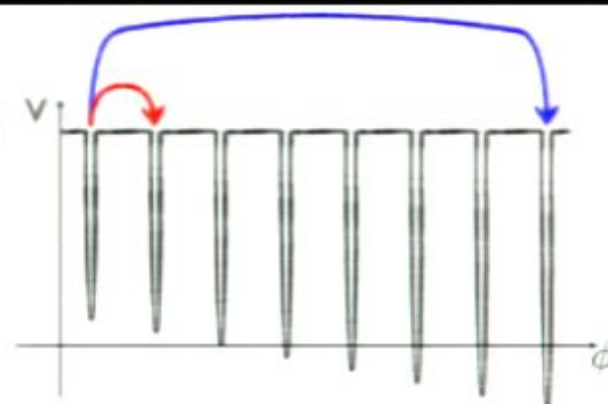
MULTIFLUX decay



MULTIFLUX decay



This potential gives small steps



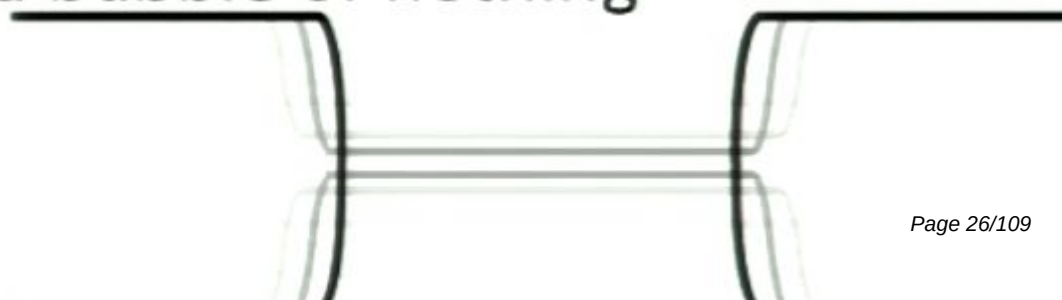
Flux compactifications with many fluxes give giant leaps

Monkey branes

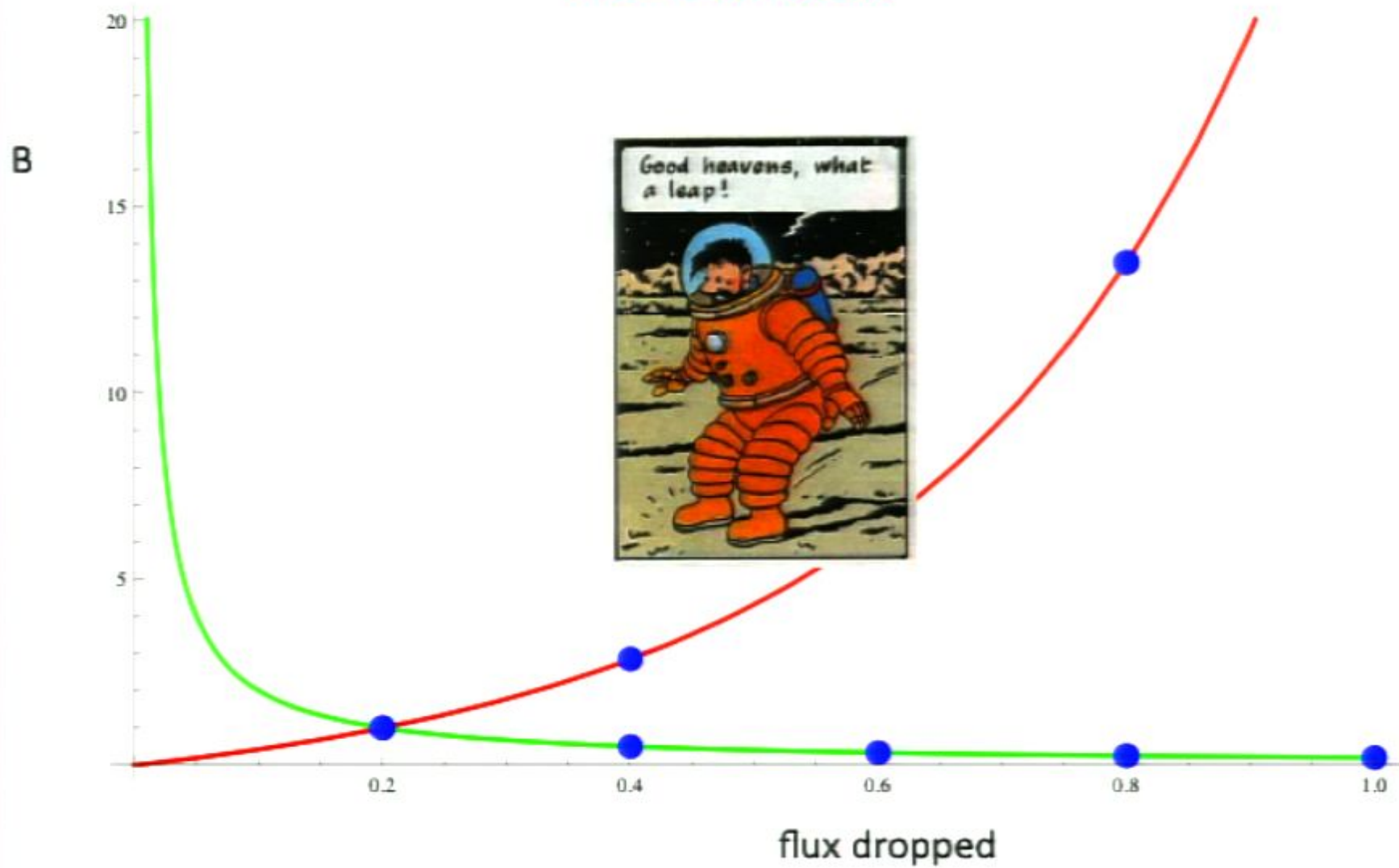


Adding back-reaction

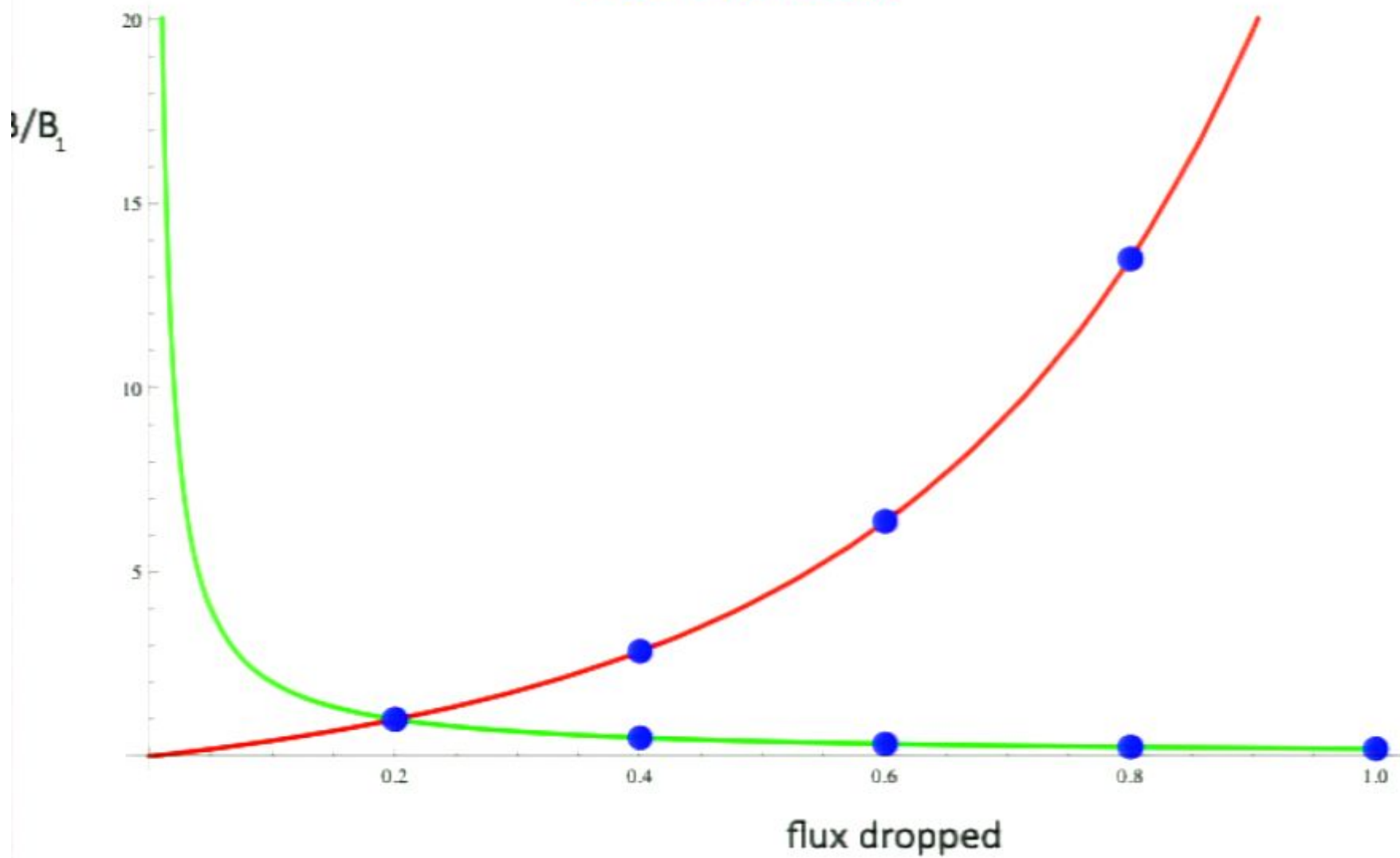
The giantest leap of all is a bubble of nothing



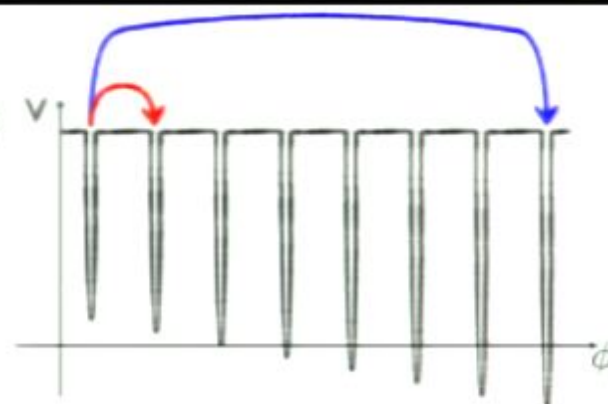
MULTIFLUX decay



MULTIFLUX decay



This potential gives small steps



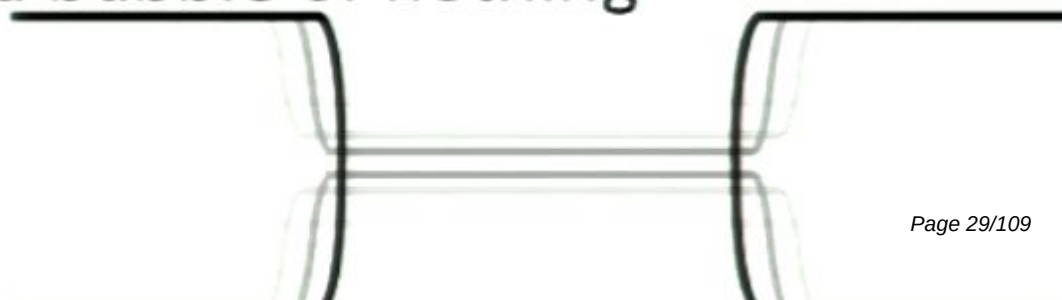
Flux compactifications with many fluxes give giant leaps

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The giantest leap of all is a bubble of nothing



Lifting to Higher Dimensions

A single higher dimensional flux that wraps many distinct cycles in the internal manifold.



Many fluxes

Bousso-Polchinski model

To what extent does the story stay true?

(fix the radion)

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2 ?$$

$$T \sim \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{1/2} ?$$

Lifting to Higher Dimensions

q -form $\mathbf{F}$

flat m -torus T^m

$$\mathbf{F} = \sum_{b=1}^{\mathfrak{N}} F_b dw_{i_1} \wedge \cdots \wedge dw_{i_q}$$

$\mathbf{F}$ must point ALL legs down the extra dimensions

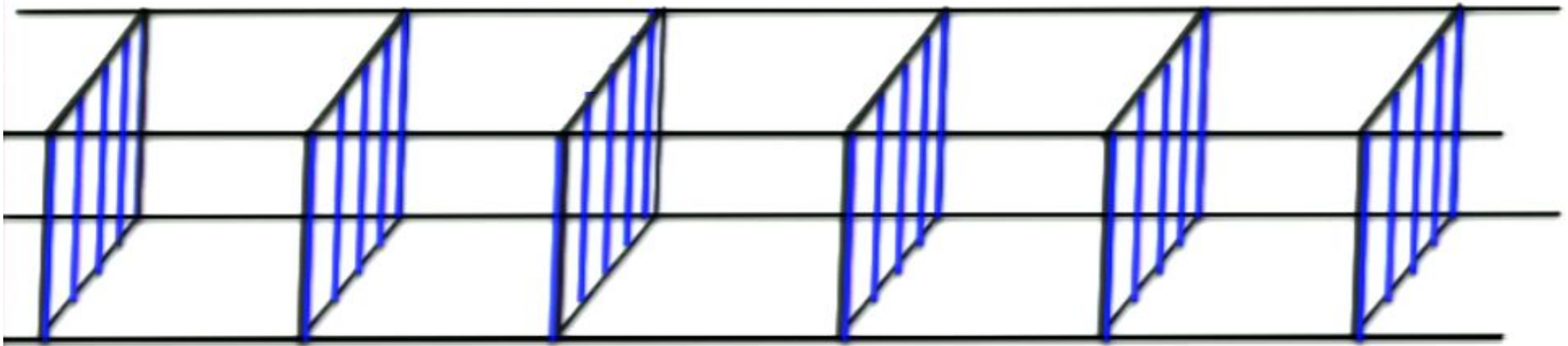
	t	r	θ	ψ	w_1	$\dots$	w_q	$\dots$	w_m
$\mathbf{F}$					x	x	x		
$\star\mathbf{F}$	x	x	x	x				x	x

$$\mathfrak{N} = \binom{m}{q}$$

$$\frac{1}{2}\mathbf{F}^2 \rightarrow \sum_{b=1}^{\mathfrak{N}} F_b^2$$

Consider flux tunneling in 1+1+2 dimensions

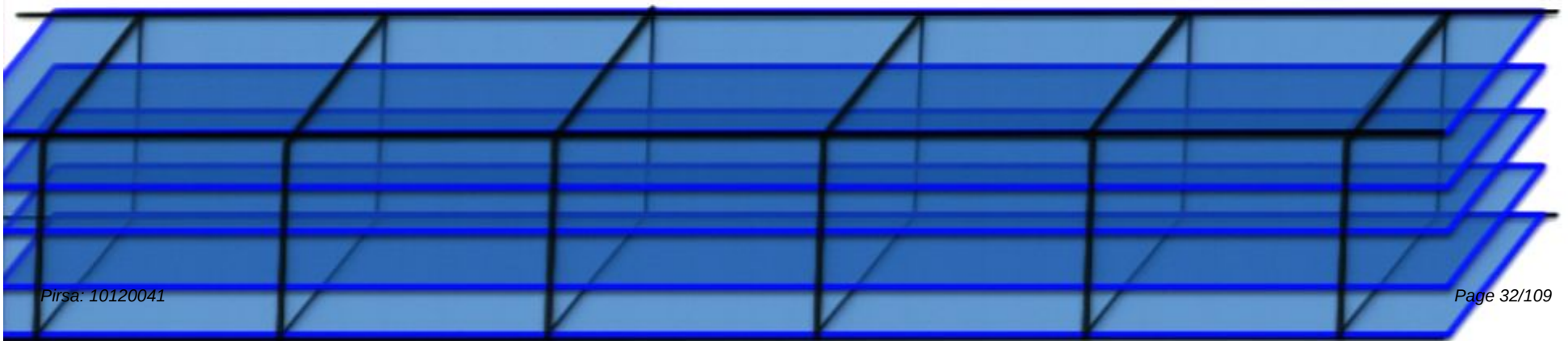
space time sphere



N units of flux

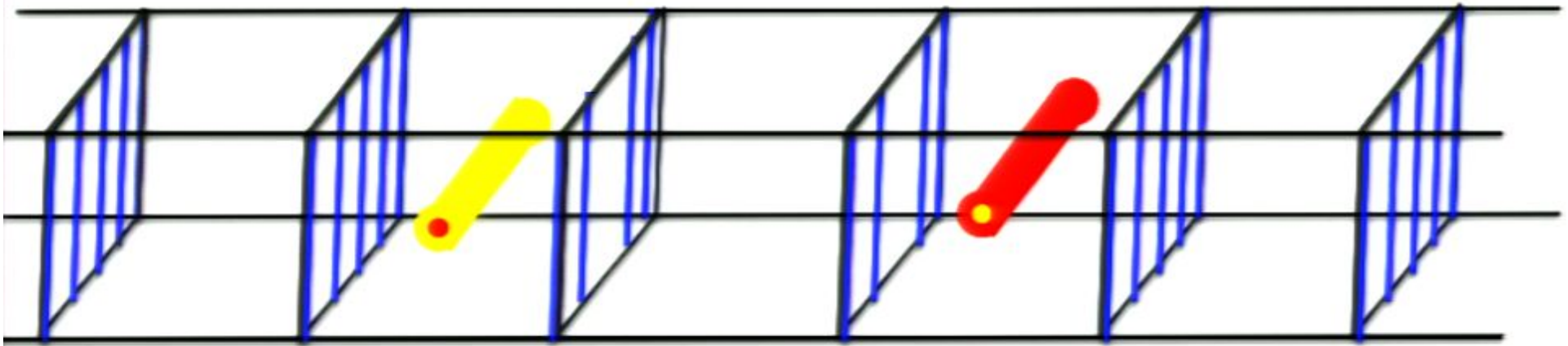
N units of flux

N units of flux

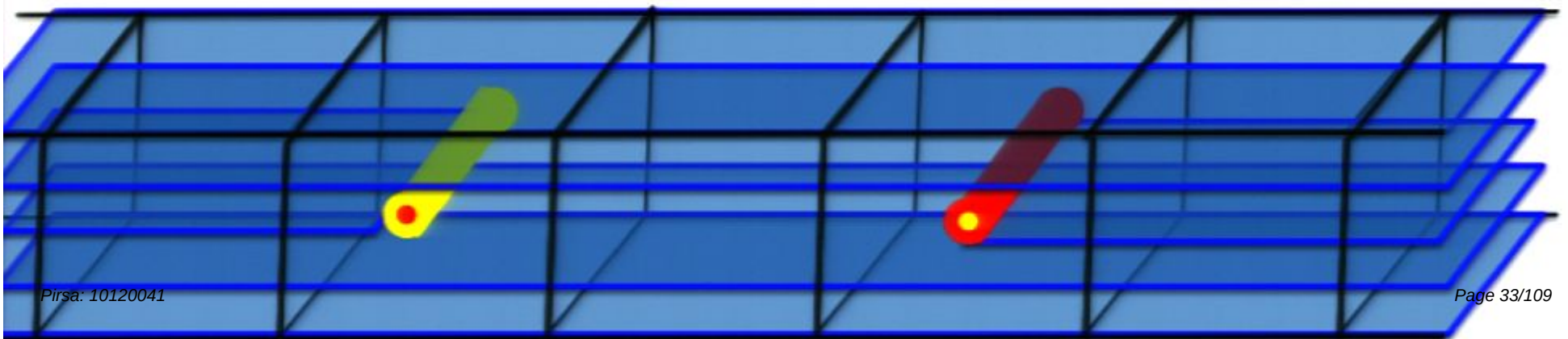


Consider flux tunneling in 1+1+2 dimensions

space time sphere



N units of flux N-1 units of flux N units of flux



Lifting to Higher Dimensions

q -form $\mathbf{F}$

flat m -torus T^m

$$\mathbf{F} = \sum_{b=1}^n F_b dw_{i_1} \wedge \cdots \wedge dw_{i_q}$$

$\mathbf{F}$ must point ALL legs down the extra dimensions

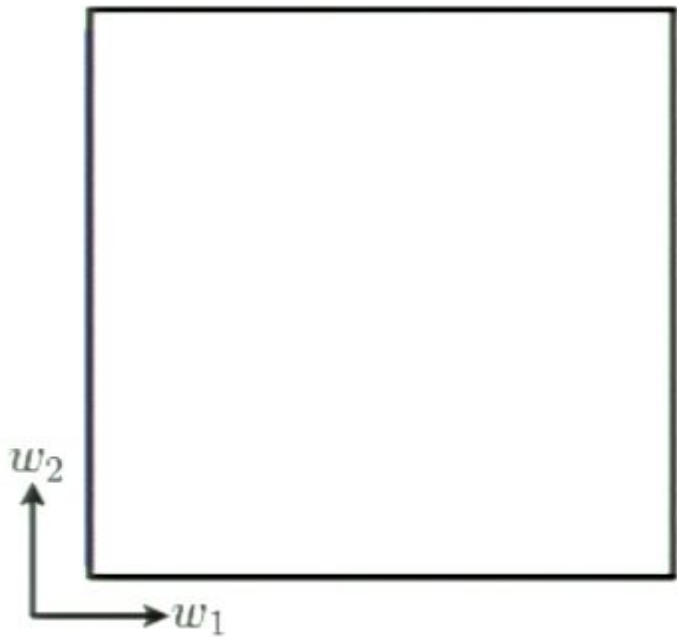
	t	r	θ	ψ	w_1	$\dots$	w_q	$\dots$	w_m
$\mathbf{F}$					x	x	x		
$\star\mathbf{F}$	x	x	x	x				x	x
brane	x		x	x				x	x

Brane makes a bubble in the extended directions
wraps the dual cycle

Let's see how this works in some specific cases

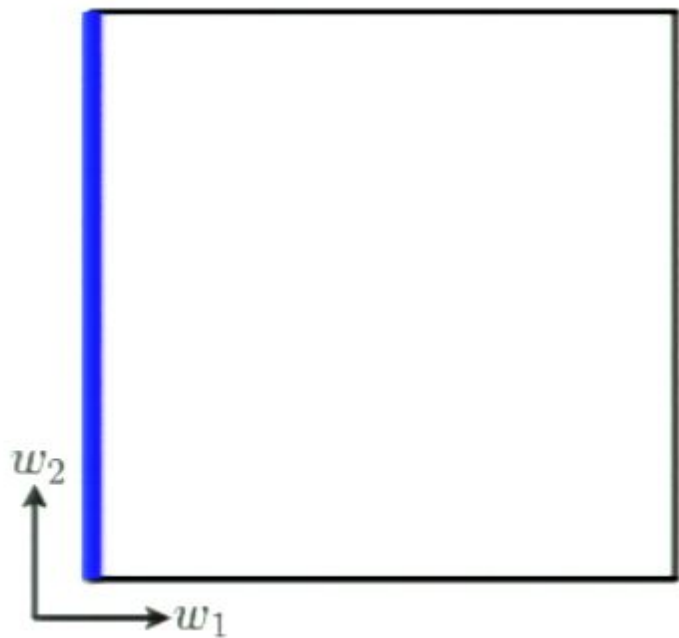
Lifting to Higher Dimensions

1-form $\mathbf{F}$
flat 2-torus T^2 $\mathfrak{N} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$



Lifting to Higher Dimensions

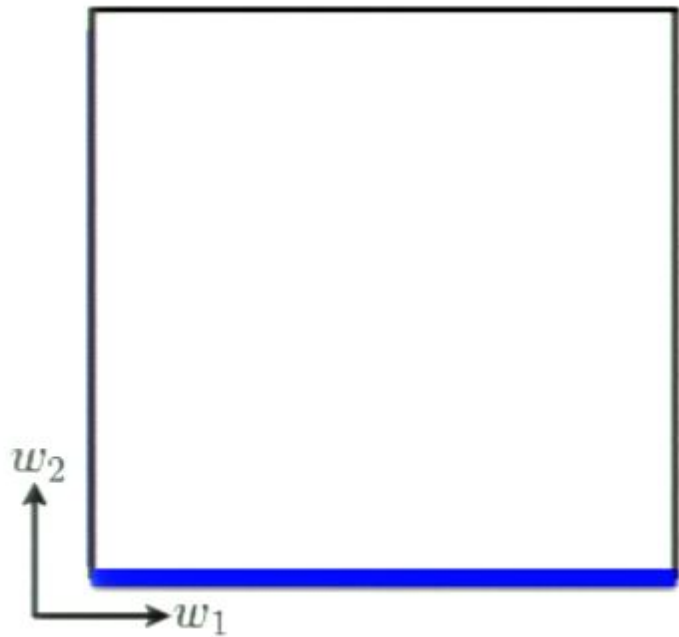
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$$\Delta \mathbf{F} = dw_1$$

Lifting to Higher Dimensions

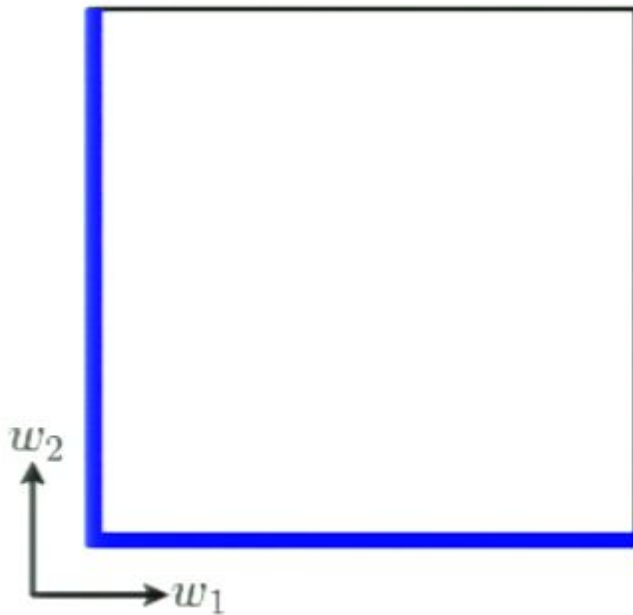
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$$\Delta \mathbf{F} = dw_2$$

Lifting to Higher Dimensions

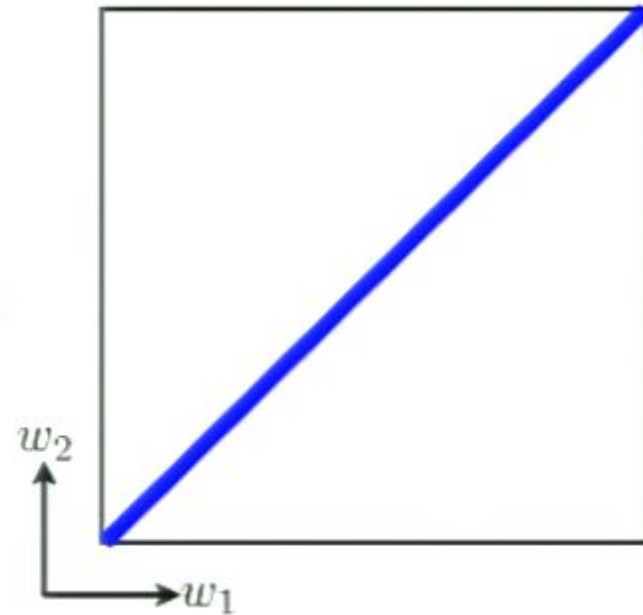
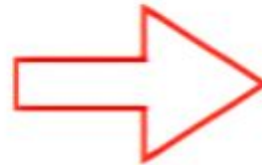
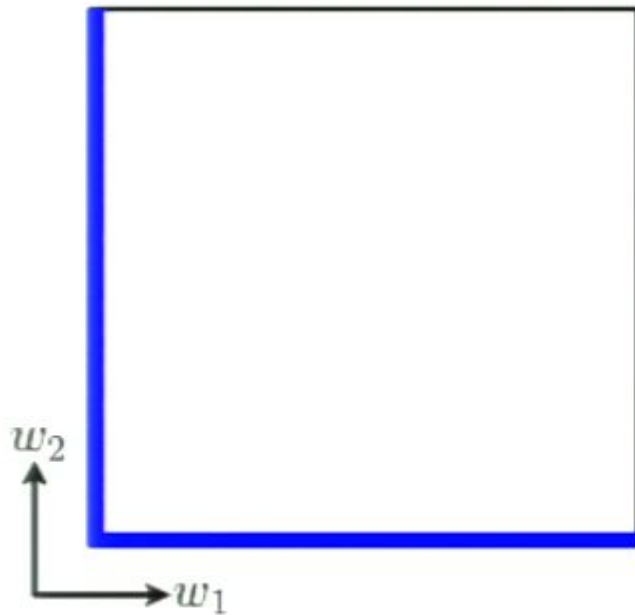
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flat 2-torus T^2 $\mathfrak{N} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$



Drop one unit of both $T \sim n?$

Lifting to Higher Dimensions

1-form $\mathbf{F}$
 flat 2-torus T^2 $\mathfrak{N} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$



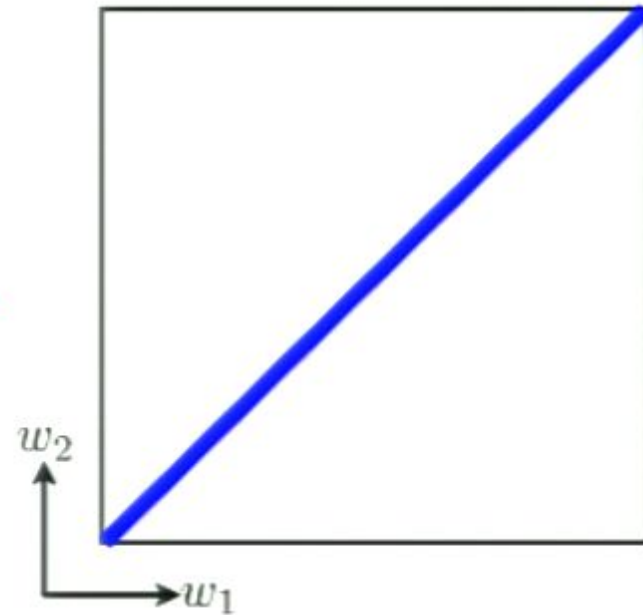
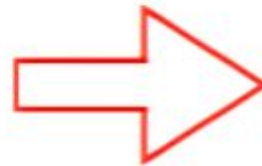
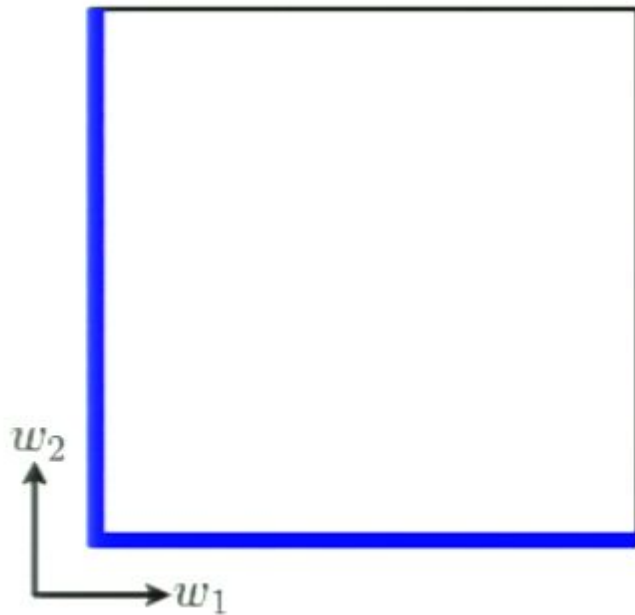
Drop one unit of both

~~$T \sim n$~~ ?

$T \sim \sqrt{n}$

Lifting to Higher Dimensions

1-form $\mathbf{F}$
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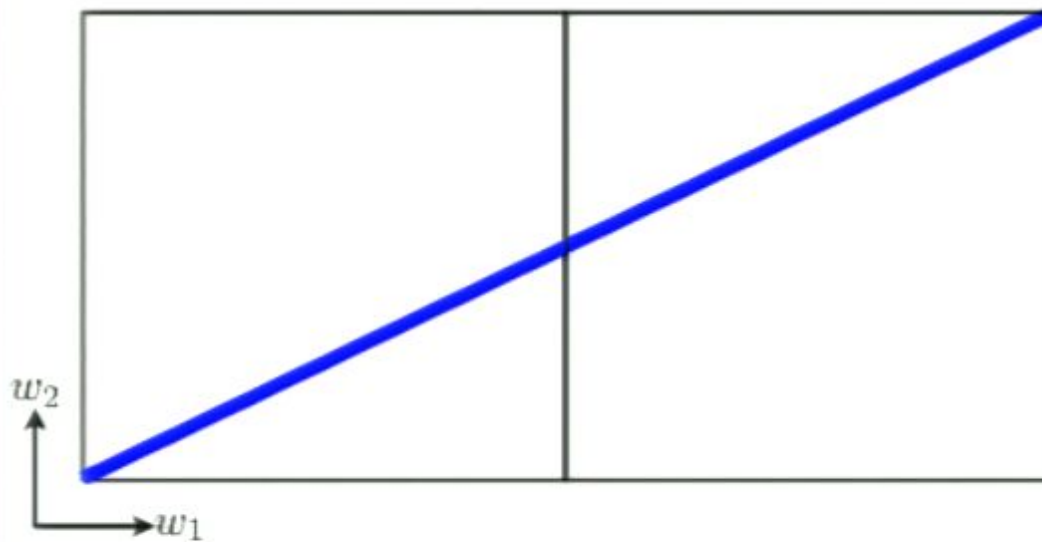
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~~$T \sim n$~~

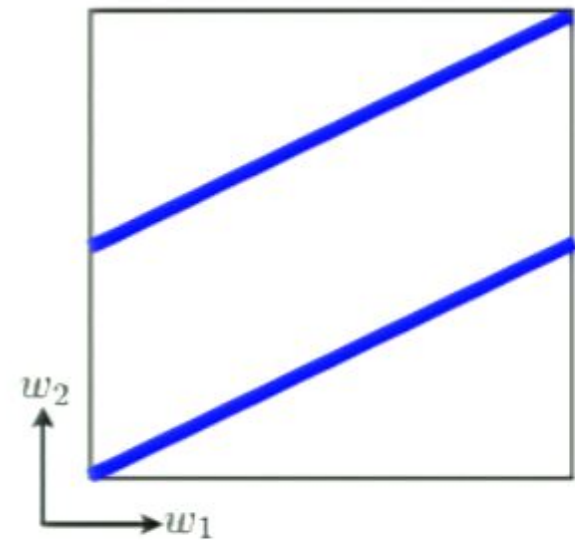
$T \sim \sqrt{n}$

Lifting to Higher Dimensions

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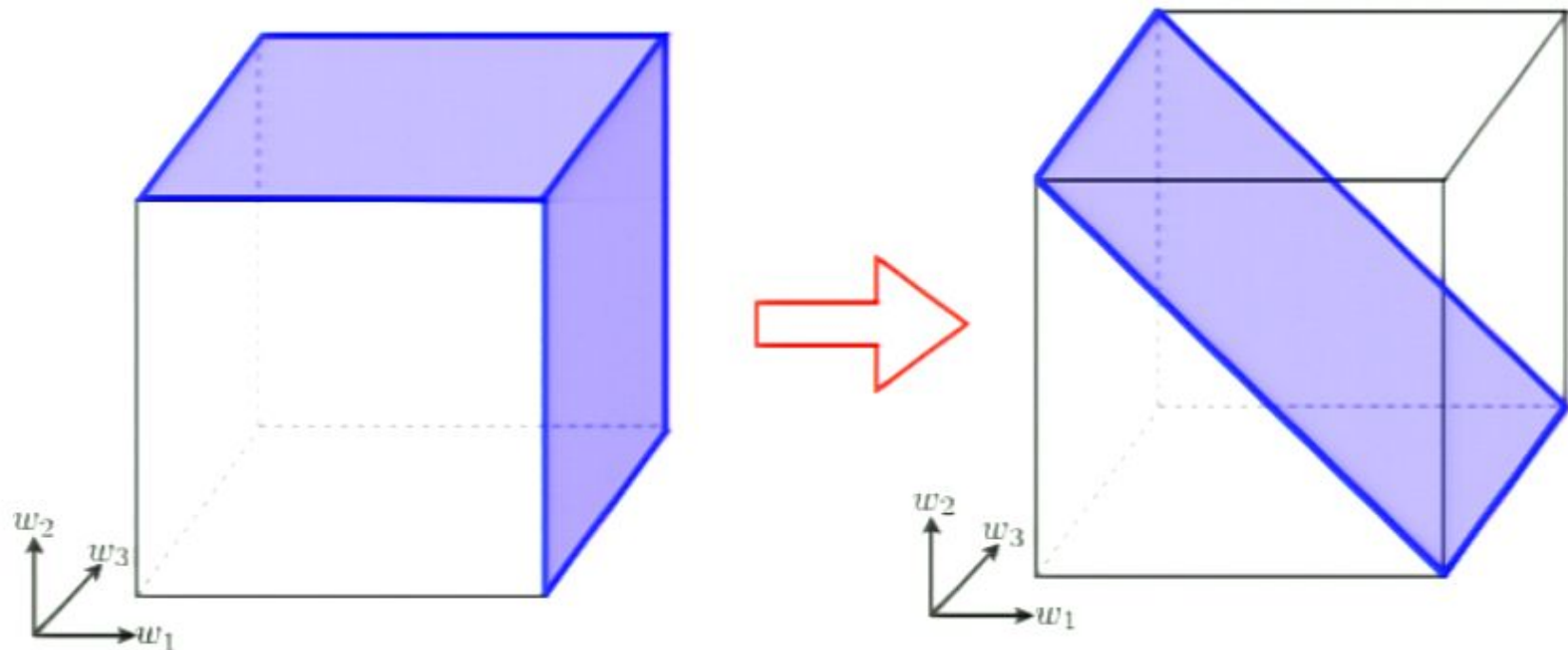
=



$$T \sim \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{1/2} = \text{Pythagoras' Theorem}$$

Lifting to Higher Dimensions

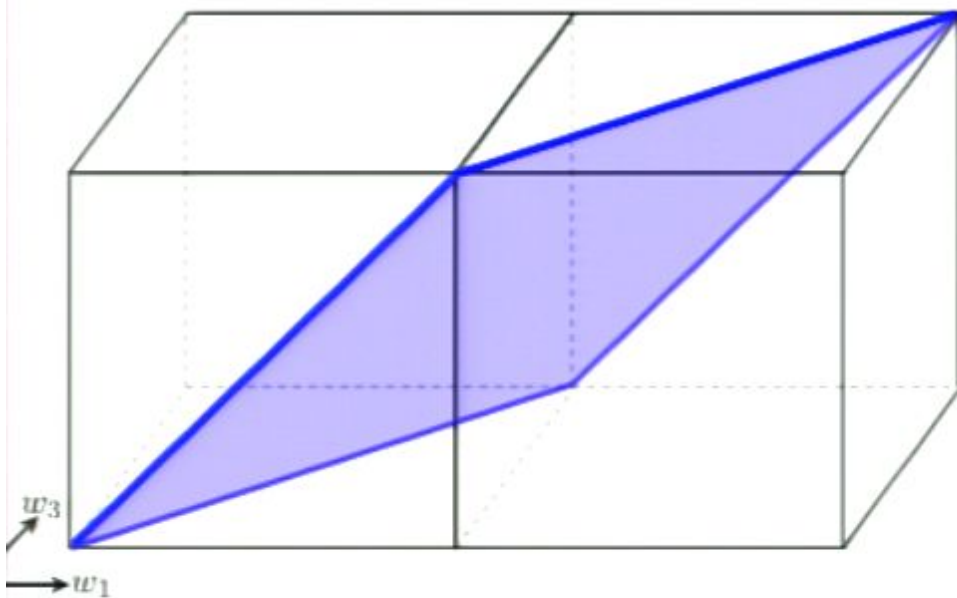
1-form $\mathbf{F}$
flat 3-torus T^3 $\mathfrak{N} = \begin{pmatrix} 3 \\ 1 \end{pmatrix} = 3$



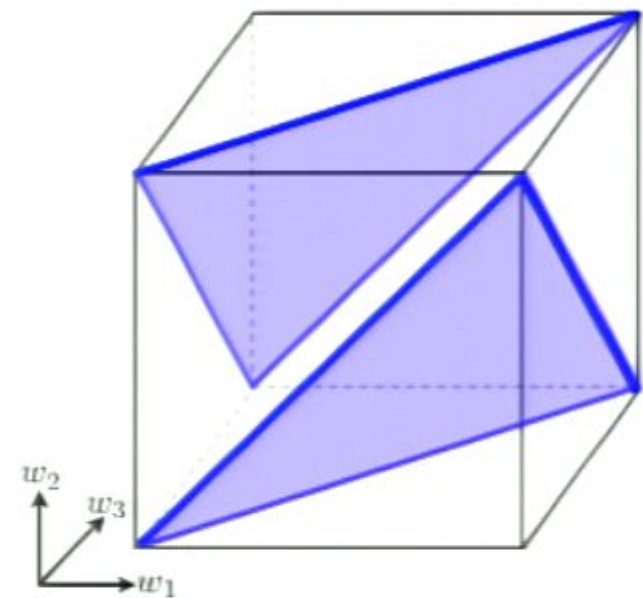
Still true

Lifting to Higher Dimensions

1-form $\mathbf{F}$
flat 3-torus T^3 $\mathfrak{N} = \begin{pmatrix} 3 \\ 1 \end{pmatrix} = 3$



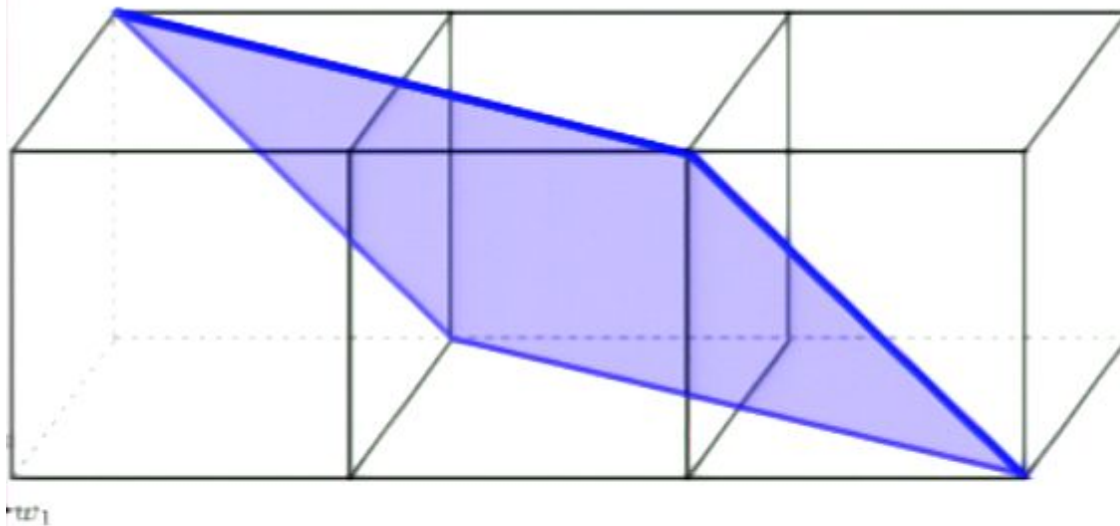
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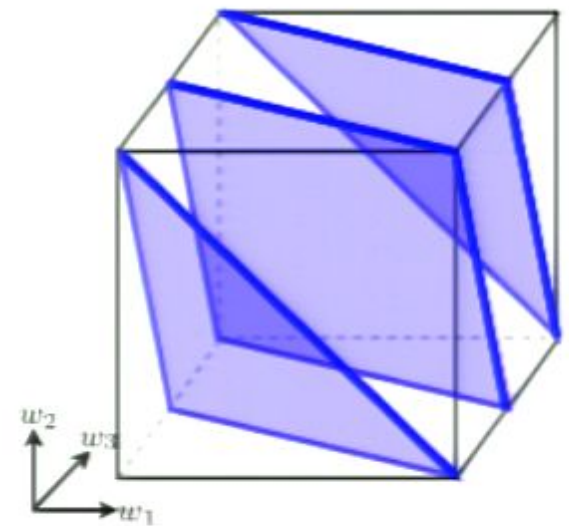
Still true

Lifting to Higher Dimensions

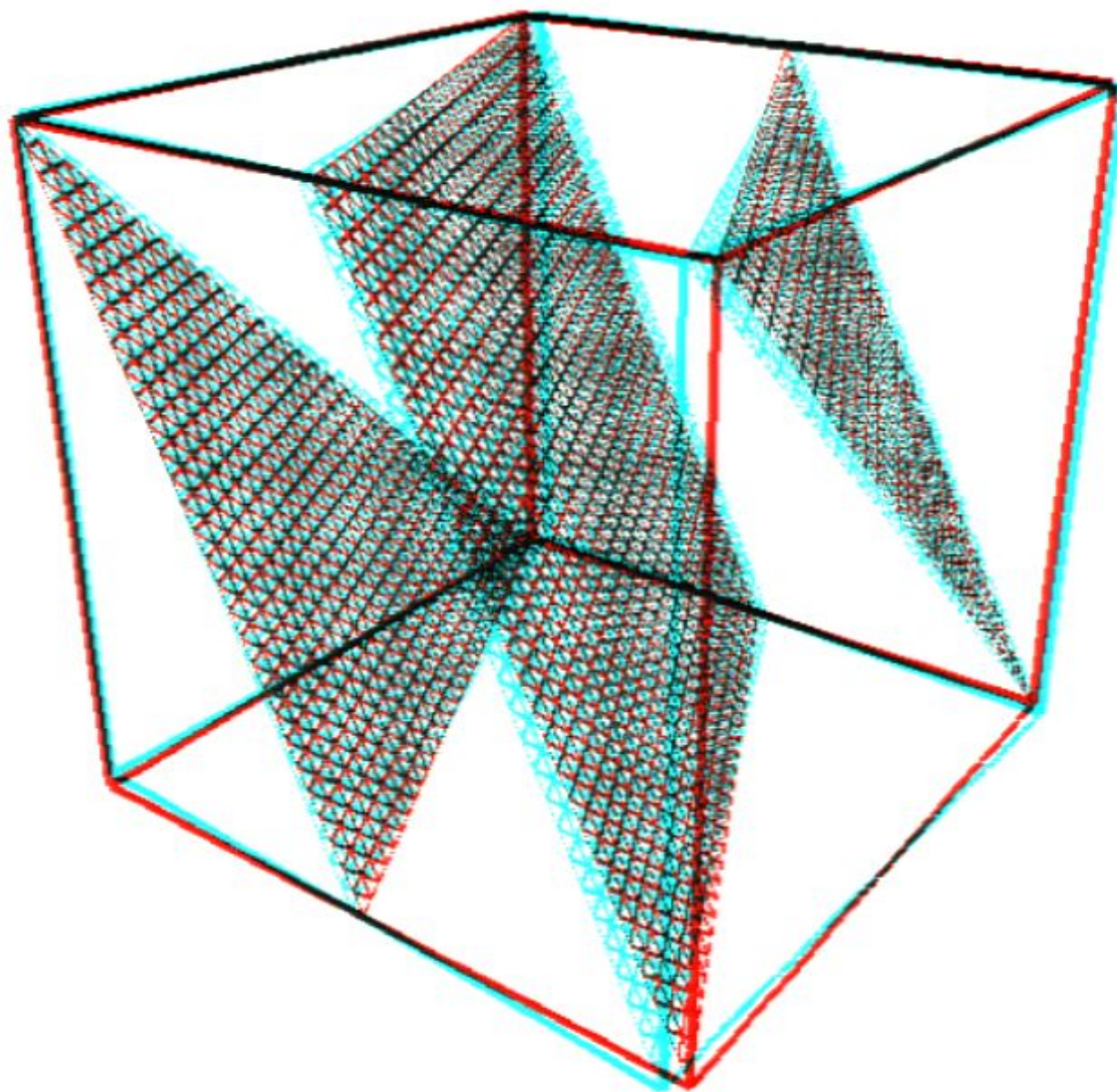
1-form $\mathbf{F}$
 flat 3-torus T^3 $\mathfrak{N} = \begin{pmatrix} 3 \\ 1 \end{pmatrix} = 3$



=



Still true



Lifting to Higher Dimensions

2-form $\mathbf{F}$

flat 4-torus T^4

$$\mathfrak{N} = \binom{4}{2} = 6$$

$$\Delta \mathbf{F} = dw_1 \wedge dw_2 + dw_3 \wedge dw_4$$

Lifting to Higher Dimensions

2-form $\mathbf{F}$

flat 4-torus T^4

$$\mathfrak{N} = \binom{4}{2} = 6$$

$$\Delta \mathbf{F} = dw_1 \wedge dw_2 + dw_3 \wedge dw_4$$

No longer true!

Requires 2 branes

$$T \sim n$$

Lifting to Higher Dimensions

The general case

$$dF = \star j$$

$$\Delta F = \star g Y_1 \wedge \cdots \wedge Y_q$$

This tells us when a single, flat brane can be found.

DECOMPOSABILITY

When it can be, $T \sim \sqrt{n}$.

Otherwise worse

Lifting to Higher Dimensions

2-form $\mathbf{F}$

flat 4-torus T^3

$$\mathfrak{N} = \binom{4}{2} = 6$$

$$\Delta \mathbf{F} = dw_1 \wedge dw_2 + dw_3 \wedge dw_4$$

No longer true!

Requires 2 branes

$$T \sim n$$

Lifting to Higher Dimensions

The general case

$$dF = \star j$$

$$\Delta F = \star g Y_1 \wedge \cdots \wedge Y_q$$

This tells us when a single, flat brane can be found.

DECOMPOSABILITY

When it can be, $T \sim \sqrt{n}$.

Otherwise worse

Lifting to Higher Dimensions

The general case

Γ is intermediate

$$\sim \sqrt{n} \quad T \sim n$$

	w_1	w_2	w_3	w_4
1	x	x		
2	x		x	
3		x	x	
4			x	x
5		x		x
6	x			x

Ideal configuration is CURVED

Lifting to Higher Dimensions

The general case

Γ is intermediate

$$\sim \sqrt{n} \quad T \sim n$$

Clustering

	w_1	w_2	w_3	w_4
1	x	x		
2	x		x	
3		x	x	
4			x	x
5		x		x
6	x			x

Ideal configuration is CURVED

breaks **rotation** symmetry of landscape

$$\mathbf{F} = dw_1 \wedge dw_2 + dw_2 \wedge dw_3$$



$$\mathbf{F} = dw_1 \wedge dw_2 + dw_3 \wedge dw_4$$



breaks **reflection** symmetry of landscape

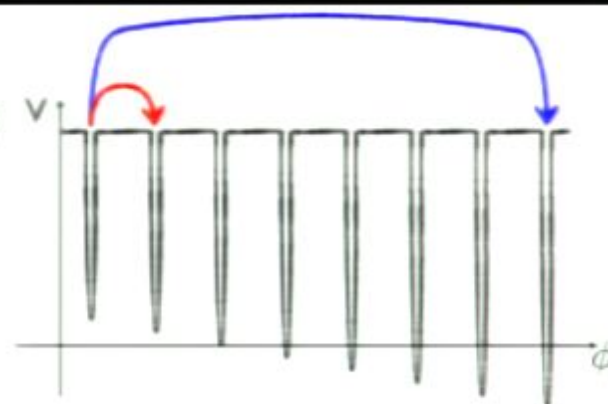
$$\mathbf{F} = dw_1 \wedge dw_3 + dw_1 \wedge dw_4 + dw_2 \wedge dw_3 + dw_2 \wedge dw_4$$



$$\mathbf{F} = dw_1 \wedge dw_3 + dw_1 \wedge dw_4 - dw_2 \wedge dw_3 + dw_2 \wedge dw_4$$



This potential gives small steps



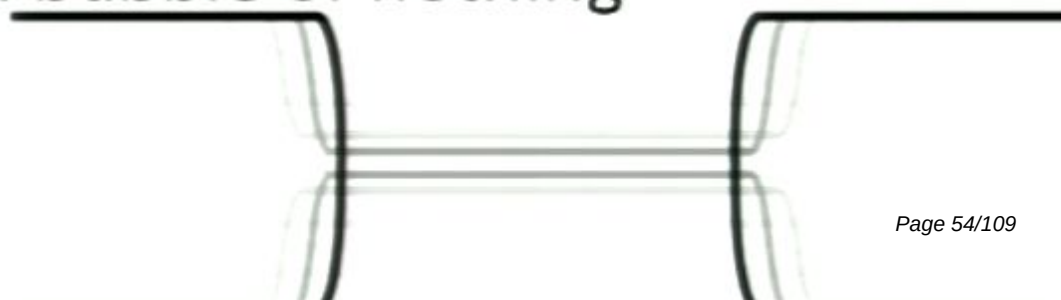
Flux compactifications with many fluxes give giant leaps

Monkey branes



Adding back-reaction

The giantest leap of all is a bubble of nothing

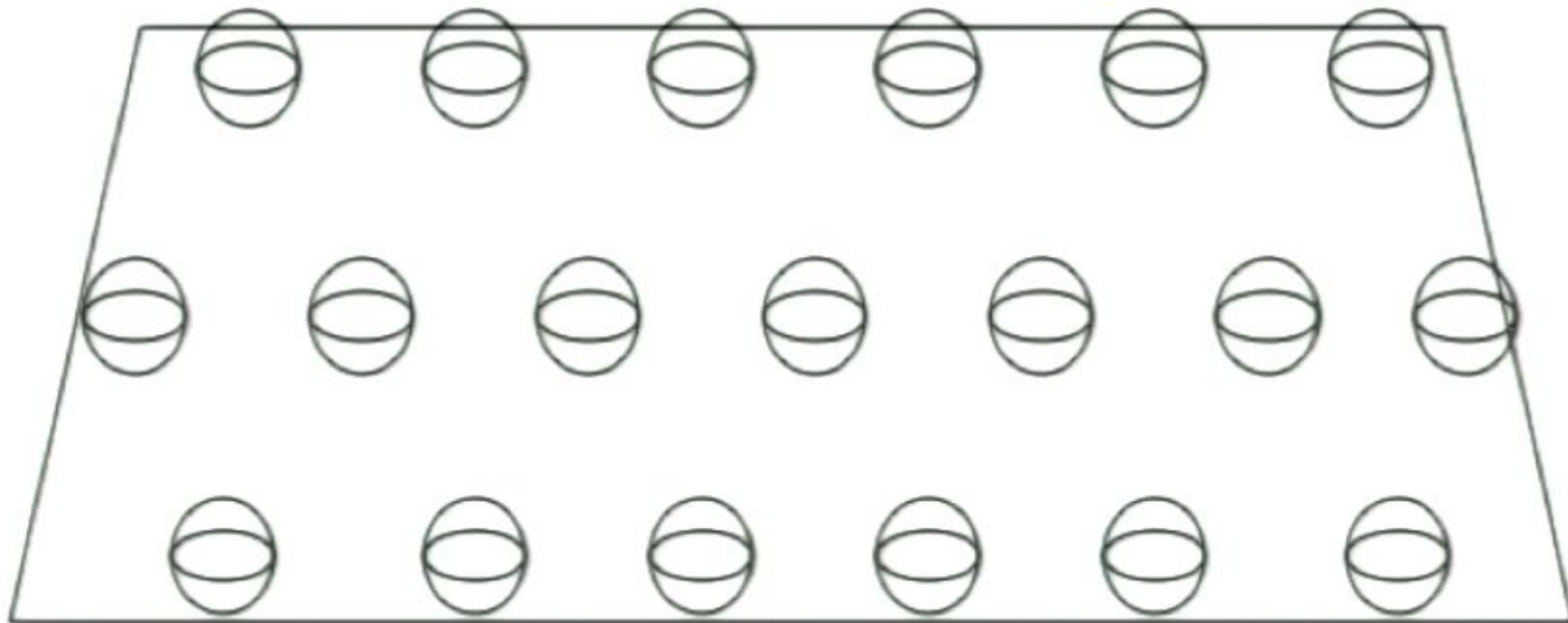


simplest possible model with **stabilized XD**s that supports **Minkowski and de Sitter** 4D slices is **6D Einstein-Maxwell**

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu + R^2 d\Omega_2^2$$

4D dS, Min, or AdS

2D sphere



three ingredients:

- spatial **curvature** of 2-spheres

- 2-form F_{ab} **flux** wrapping 2-spheres

i.e. Maxwell
magnetic flux

$$S_{\text{EM}} = \int d^6x \sqrt{-G} \left(\frac{1}{2} \mathcal{R}^{(6)} - \frac{1}{4} F_{AB} F^{AB} - \Lambda_6 \right)$$

Freund Rubin (1980)

effective 4D theory:

$$ds^2 = e^{-\psi(x)/M} g_{\mu\nu} dx^\mu dx^\nu + e^{\psi(x)/M} R^2 d\Omega_2^2$$

with different two-form fluxes, each with field strength

$$\mathbf{F}_i = \frac{g_i N_i}{4\pi} \sin \theta d\theta \wedge d\phi,$$

$$F^2 = \sum_{i=1}^n g_i^2 N_i^2$$

$$S = \int d^4x \sqrt{-g} \left(\frac{1}{2} M^2 \mathcal{R}^{(4)} - \frac{1}{2} \partial_\mu \psi \partial^\mu \psi - V(\psi) \right)$$

$$V(\psi) = 4\pi \left(\frac{1}{2} \frac{F^2}{4\pi M^2} e^{-3\psi/M} - e^{-2\psi/M} + R^2 \Lambda_6 e^{-\psi/M} \right)$$

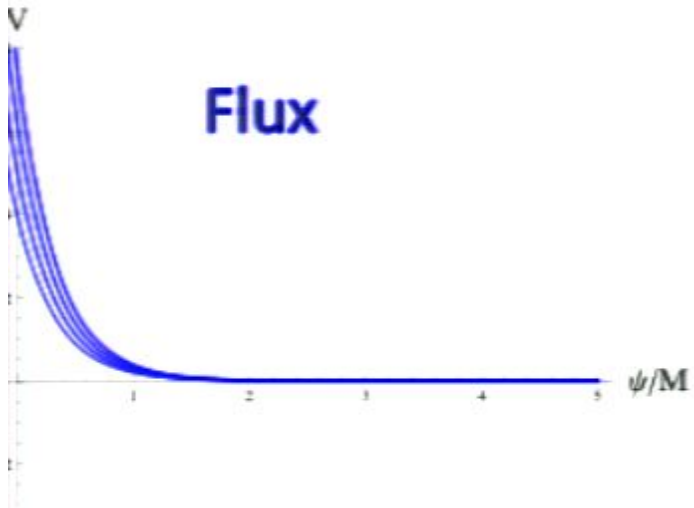
canonically
normalized
RADION

FLUX
'repulsive'

CURVATURE
'attractive'
medium-range

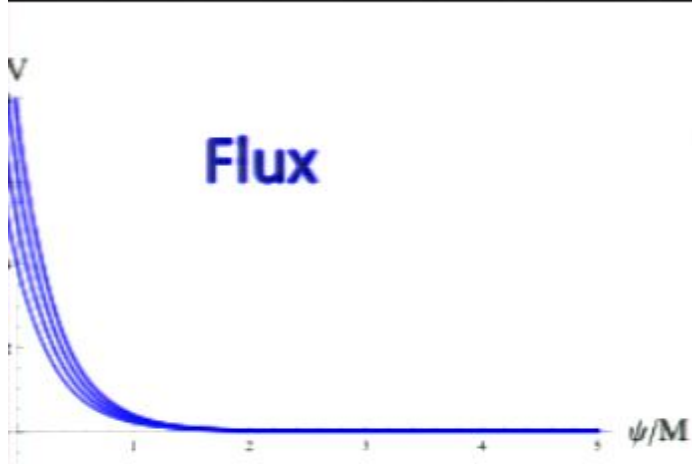
C.C.
'repulsive'
long-range

Flux

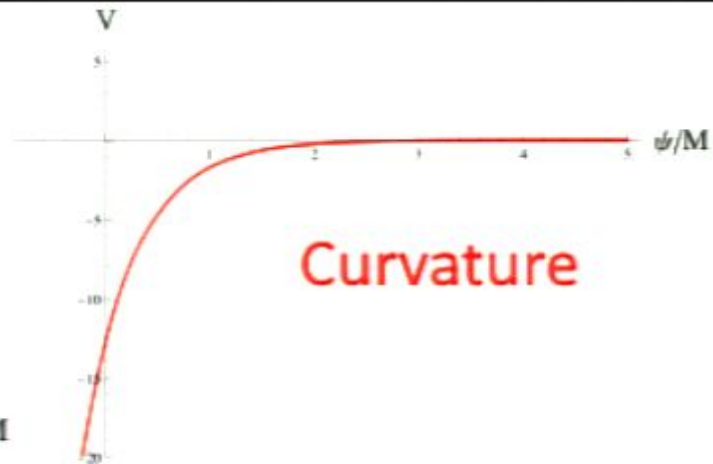


$$V(\psi) = 4\pi \left(\frac{1}{2} \frac{F^2}{M^2} e^{-3\psi/M} - e^{-2\psi/M} + R^2 \Lambda_0 e^{-\psi/M} \right)$$

Flux

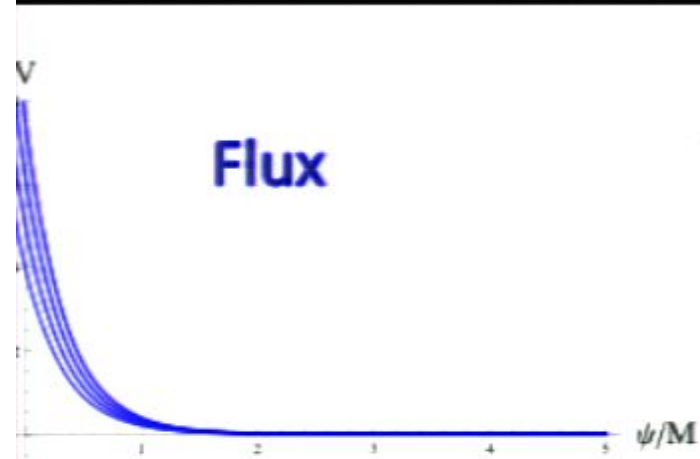


Curvature

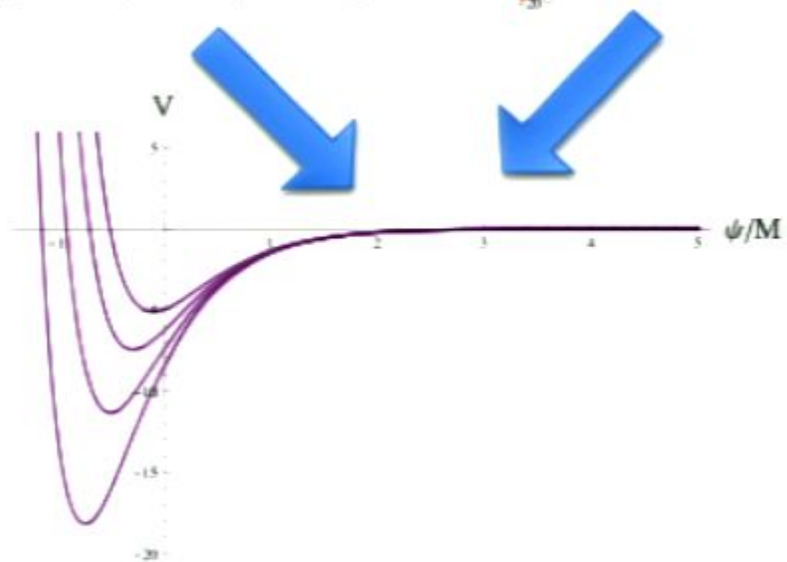
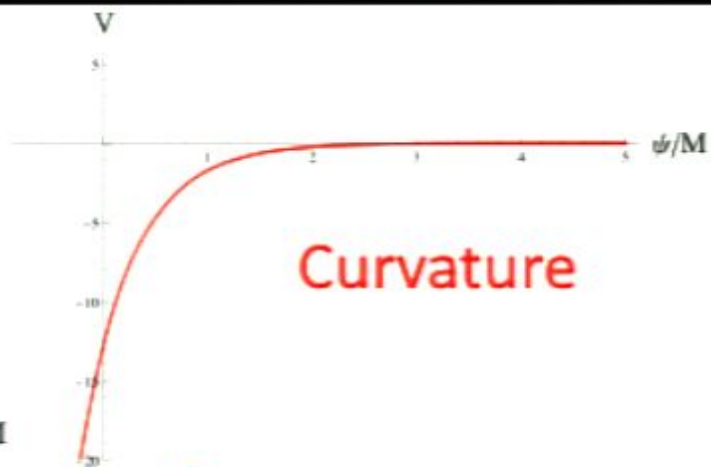


$$V(\psi) = 4\pi \left(\frac{1}{2} \frac{F^2}{e^{-3\psi/M}} - e^{-2\psi/M} + R^2 \Lambda_0 e^{-\psi/M} \right)$$

Flux

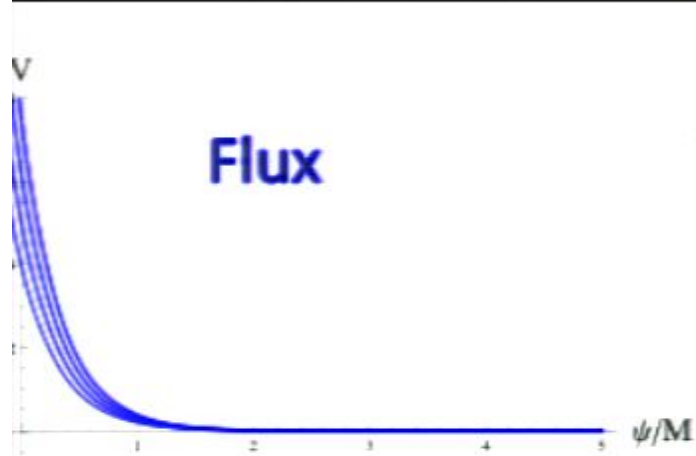


Curvature

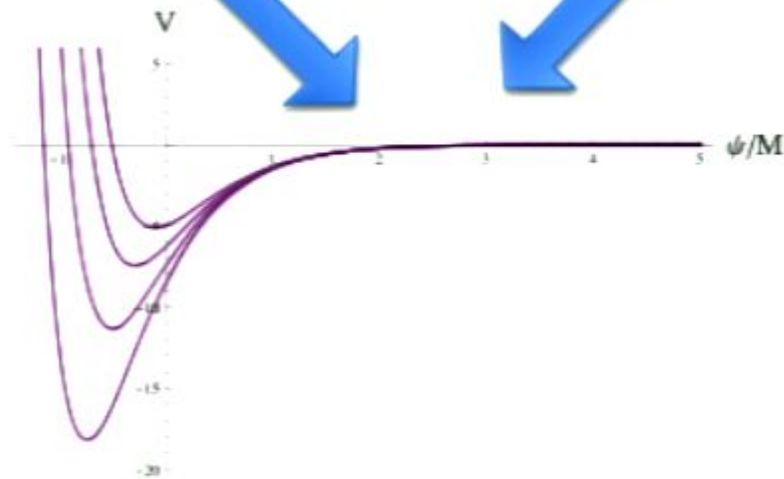
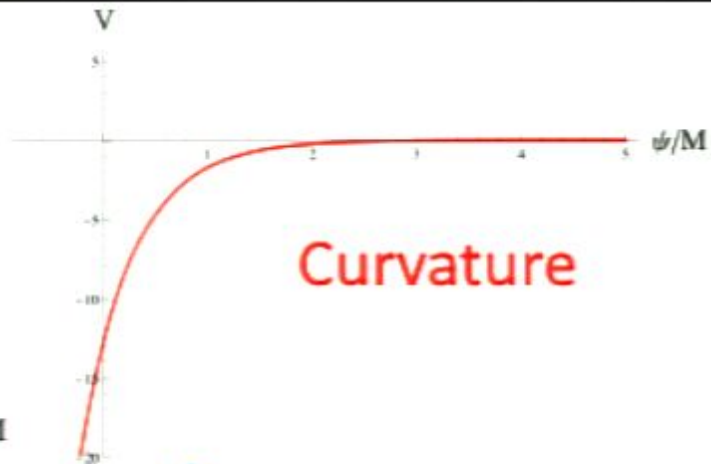


$$V(\psi) = 4\pi \left(\frac{1}{2} \frac{F^2}{M} e^{-3\psi/M} - e^{-2\psi/M} + R^2 \Lambda_0 e^{-\psi/M} \right)$$

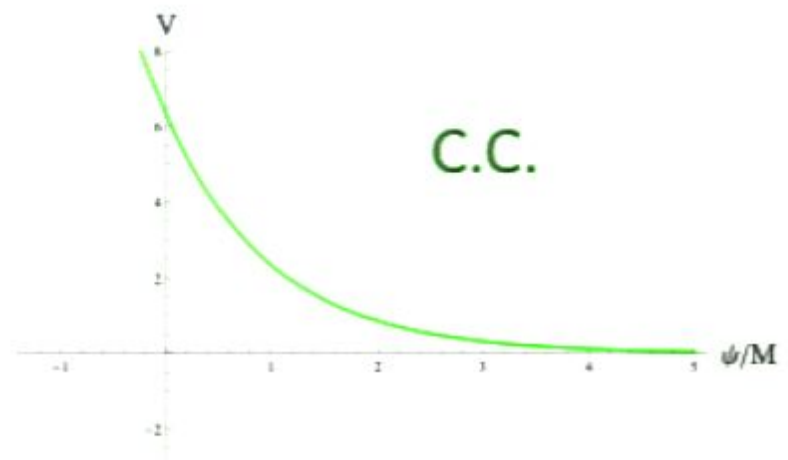
Flux



Curvature

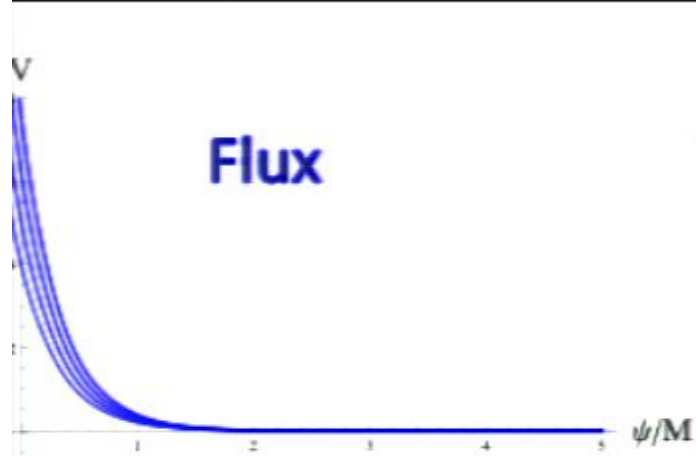


C.C.

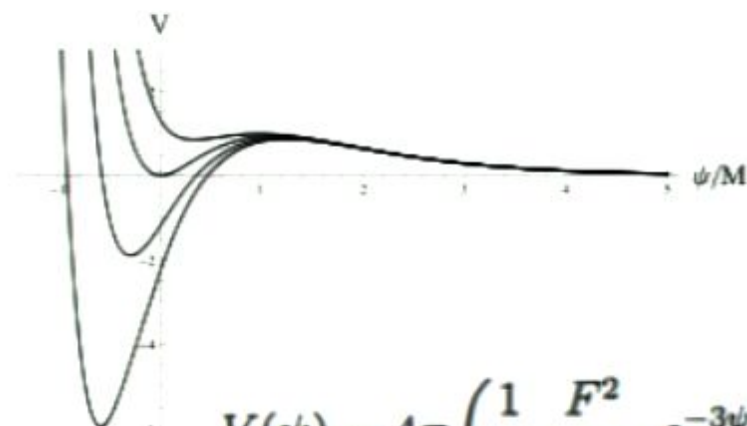
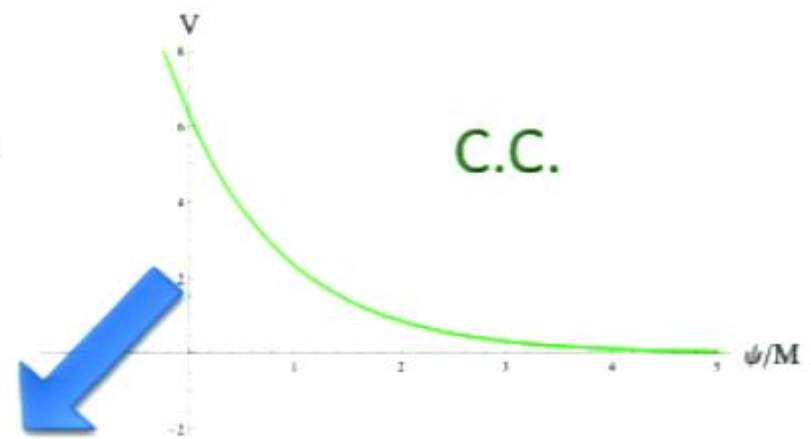
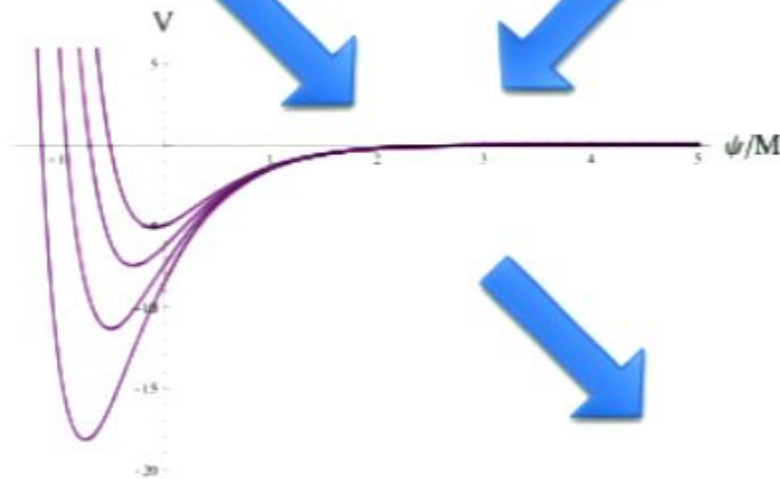
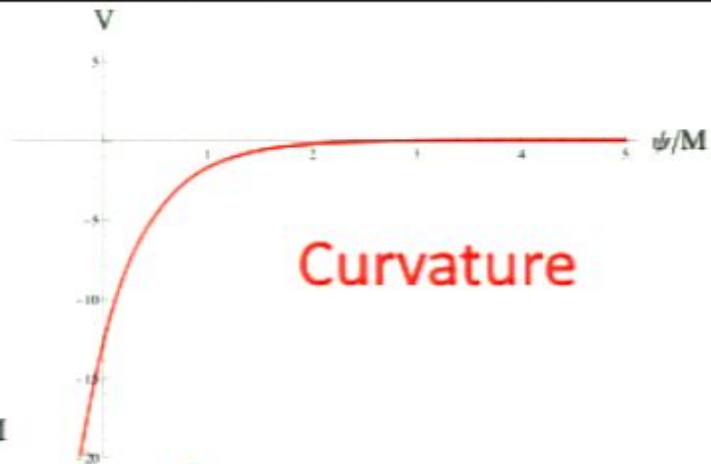


$$V(\psi) = 4\pi \left(\frac{1}{2} \frac{F^2}{M^2} e^{-3\psi/M} - e^{-2\psi/M} + R^2 \Lambda_0 e^{-\psi/M} \right)$$

Flux

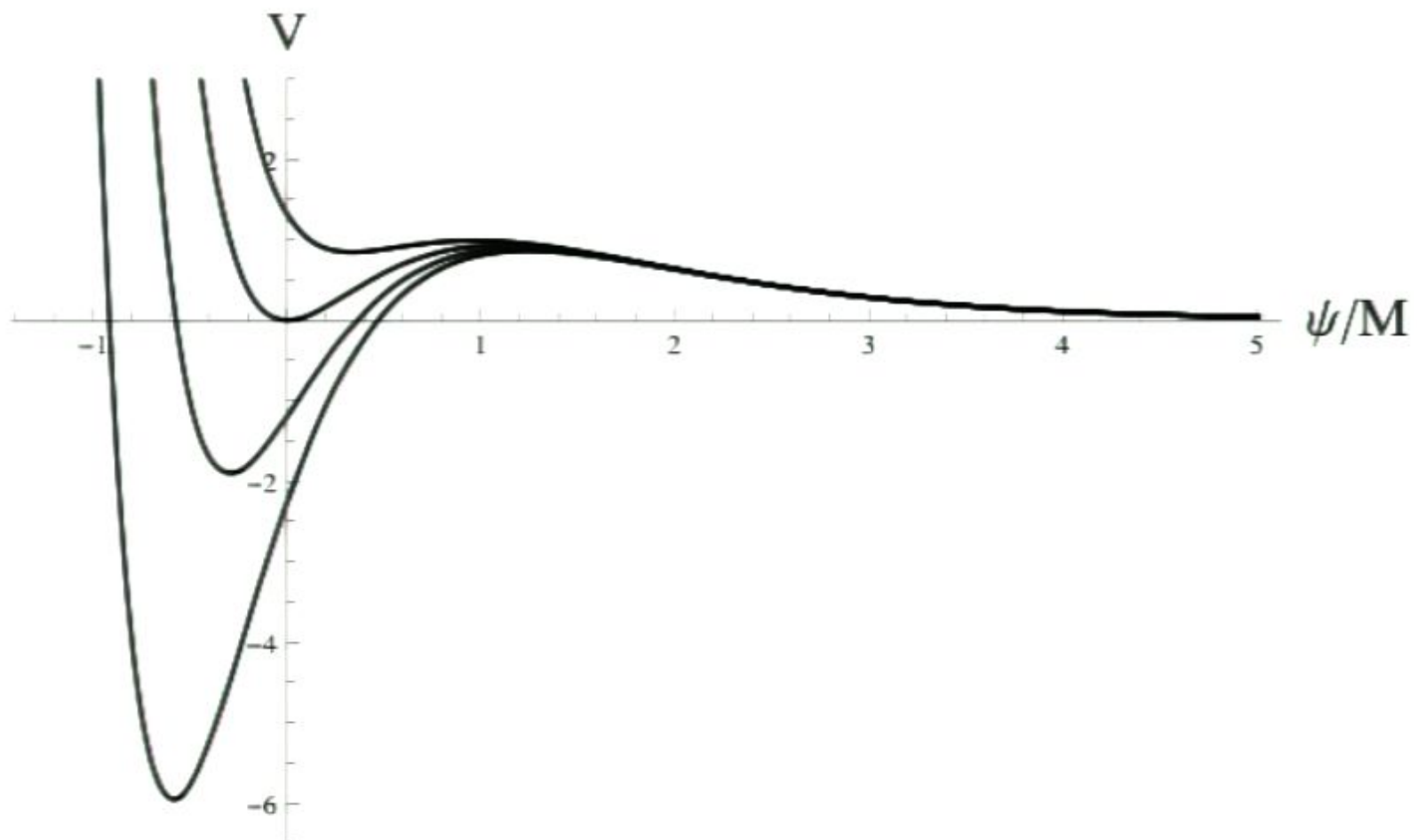


Curvature

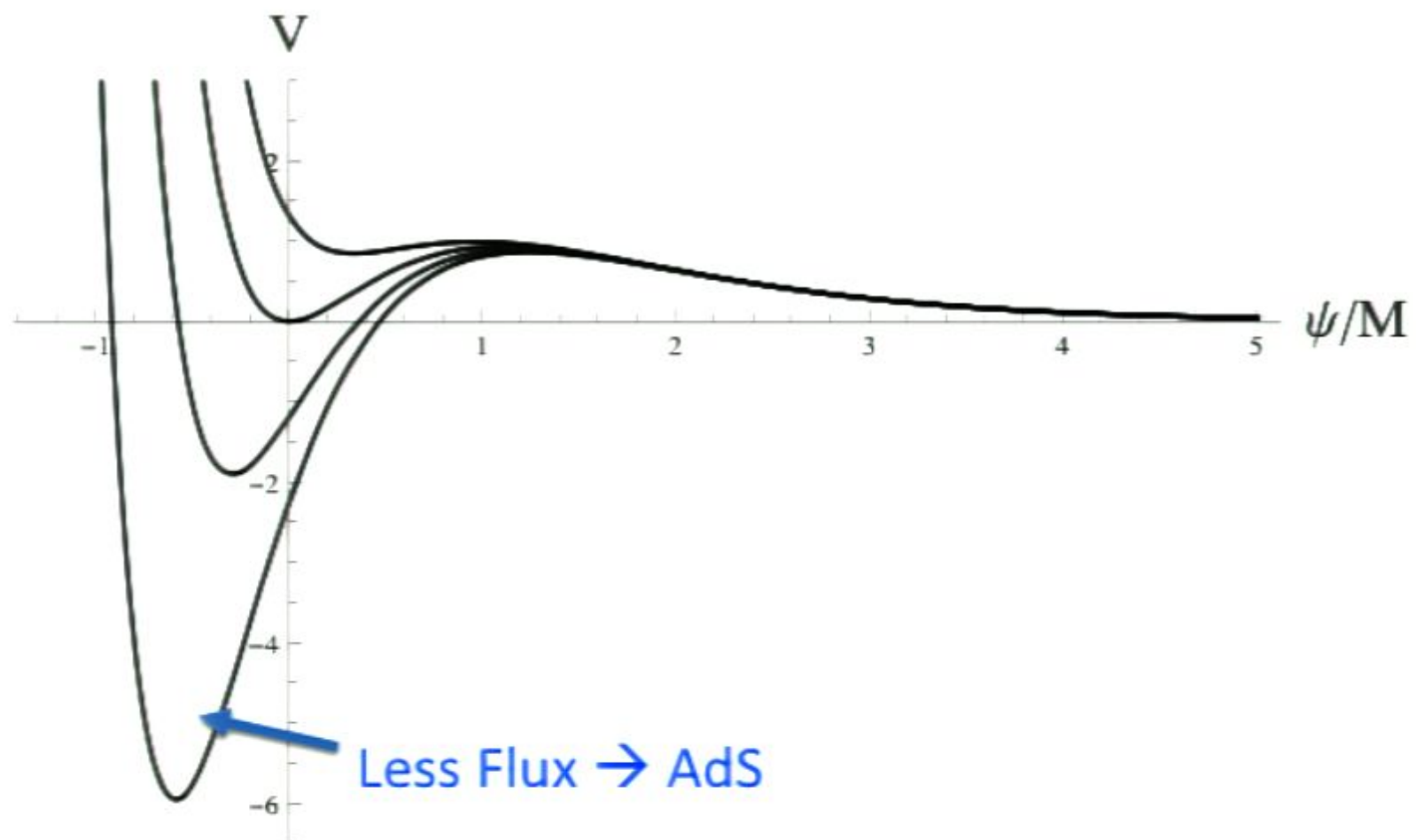


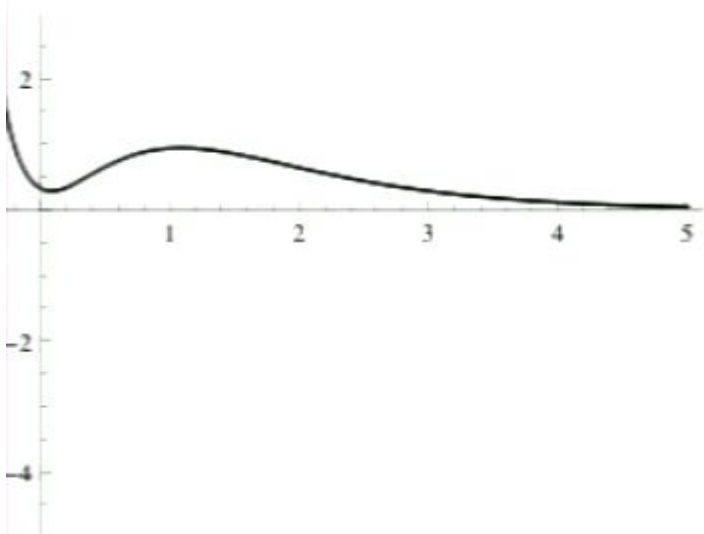
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A landscape of (discrete) vacua

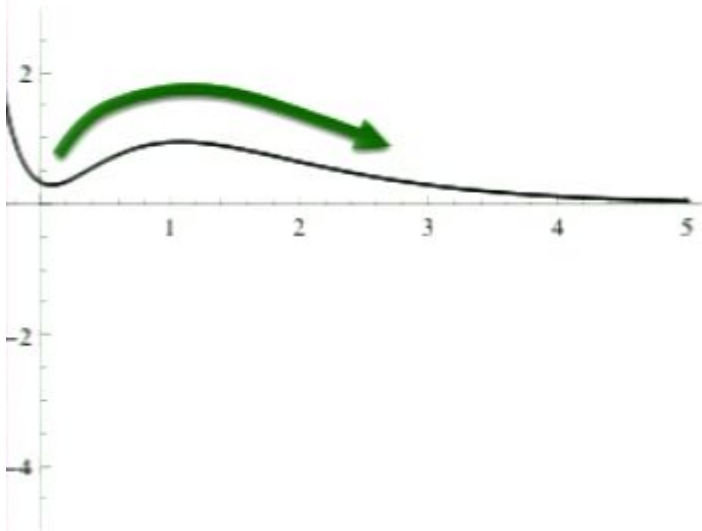


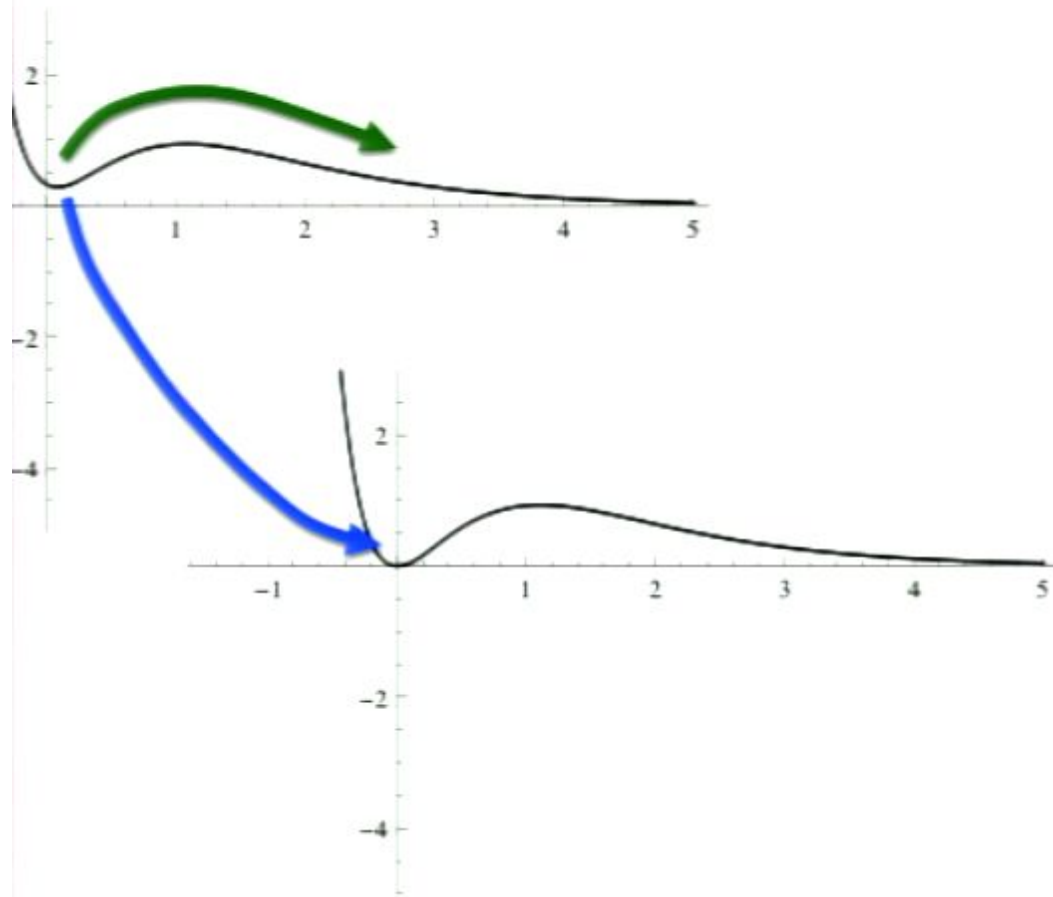
A landscape of (discrete) vacua





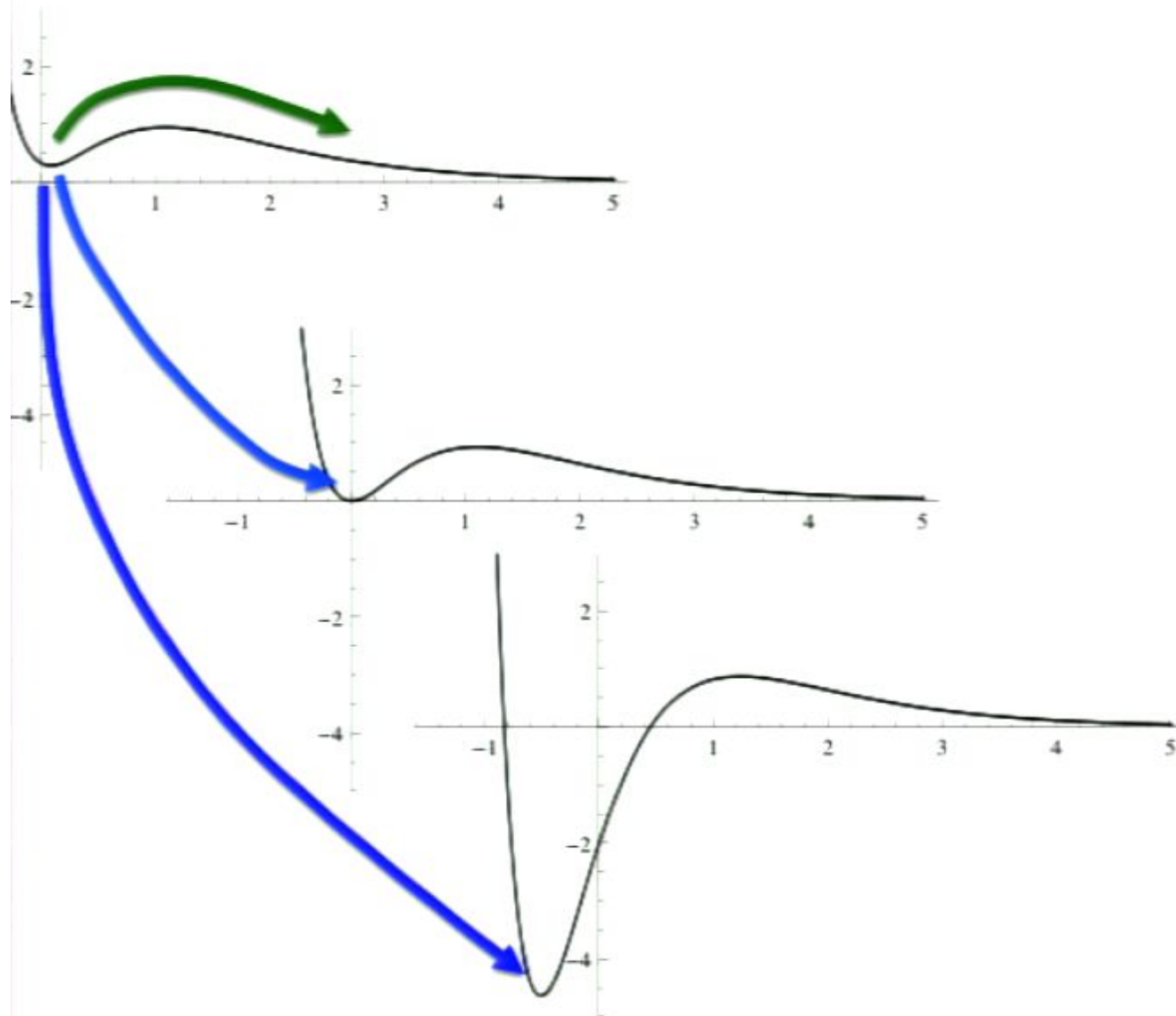
Decompactify





Decompactify

Drop one unit
(small step)



Decompactify

Drop one unit
(small step)

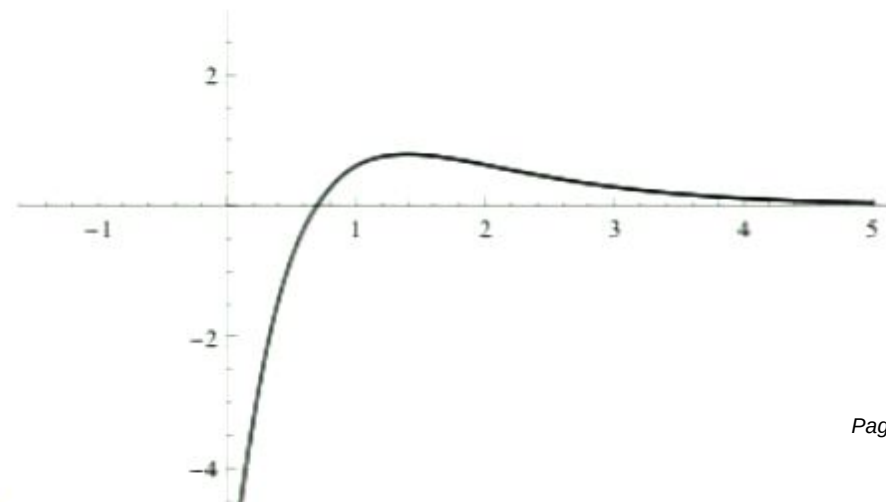
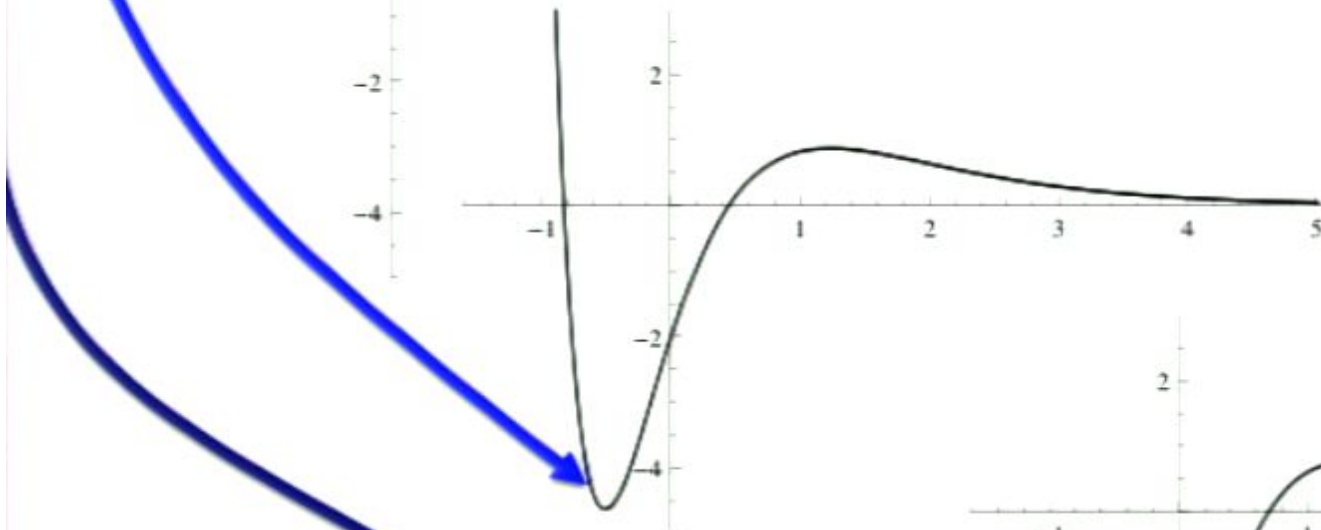
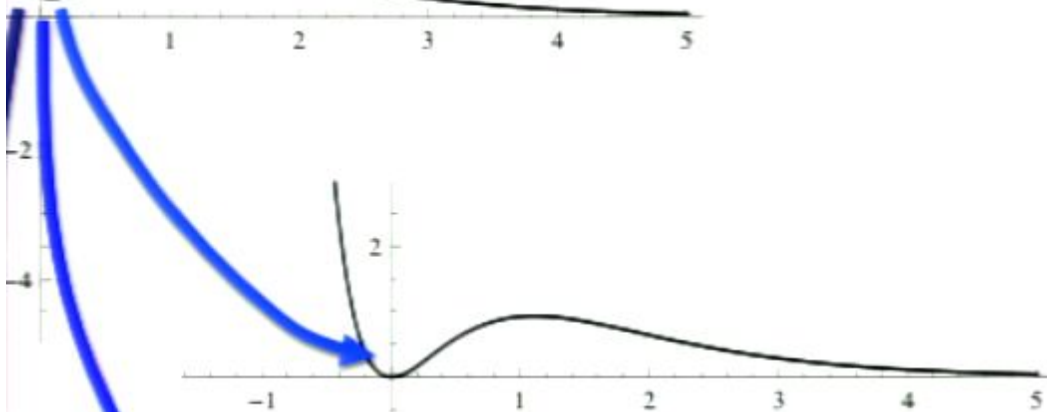
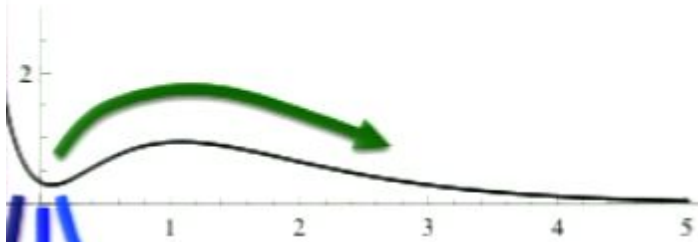
Drop many units
(giant leap)

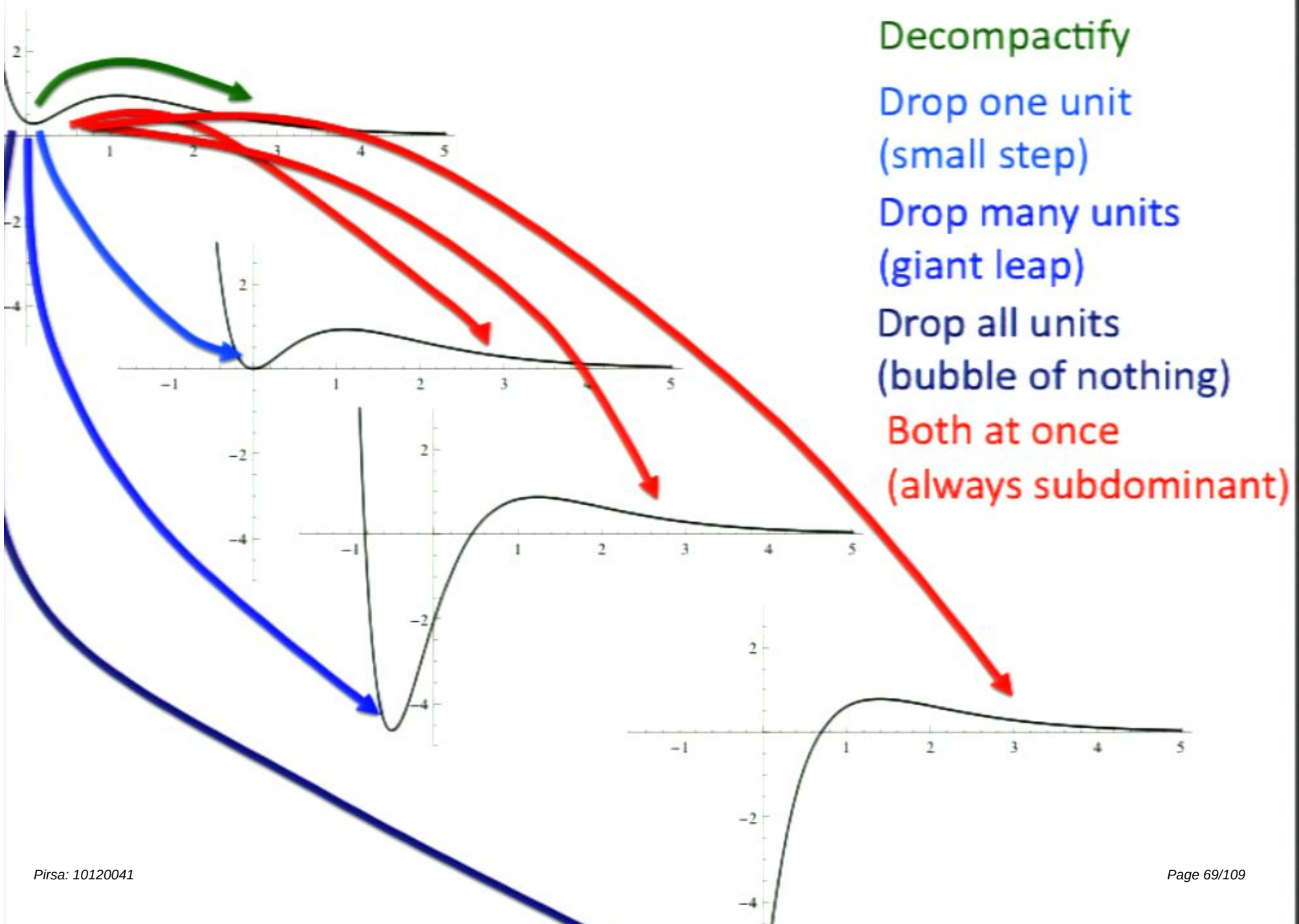
Decompactify

Drop one unit
(small step)

Drop many units
(giant leap)

Drop all units
(bubble of nothing)





Decompactify

Drop one unit
(small step)

Drop many units
(giant leap)

Drop all units
(bubble of nothing)

Both at once
(always subdominant)

2 reasons back-reaction helps giant leaps

Radion mediates attractive force

$$S_{\text{brane tension}} = -T \int_{\Sigma} \sqrt{-\gamma} d^3\xi \qquad ds^2 = e^{-\psi(x)/M_4} g_{\mu\nu} dx^\mu dx^\nu + e^{\psi(x)/M_4} R^2 d\Omega_2^2$$

$$S_{\text{brane tension}} = -T e^{-3\psi(\xi)/2M_4} \times (\text{surface area})$$

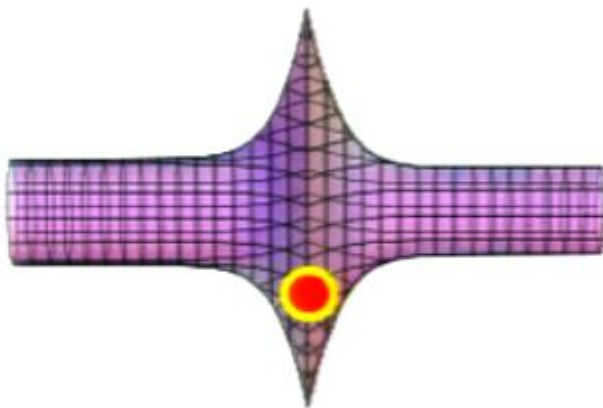
2 reasons back-reaction helps giant leaps

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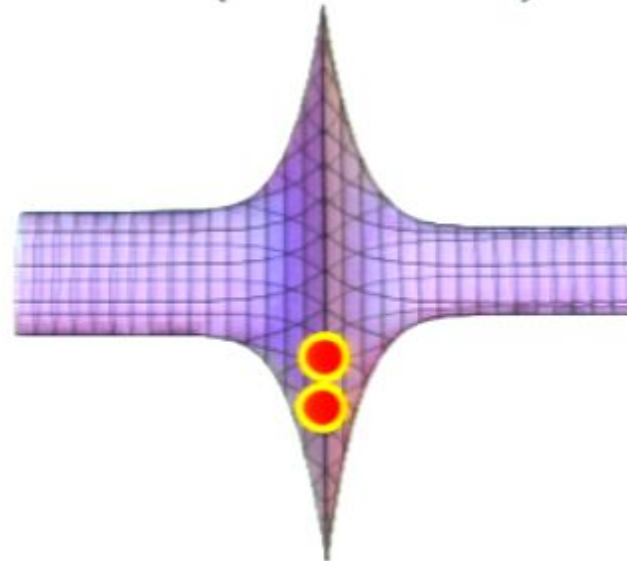
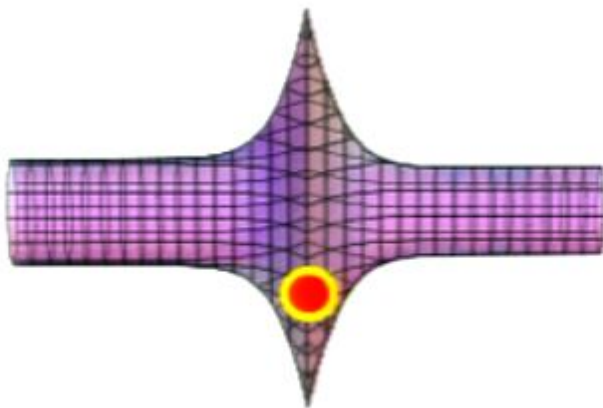
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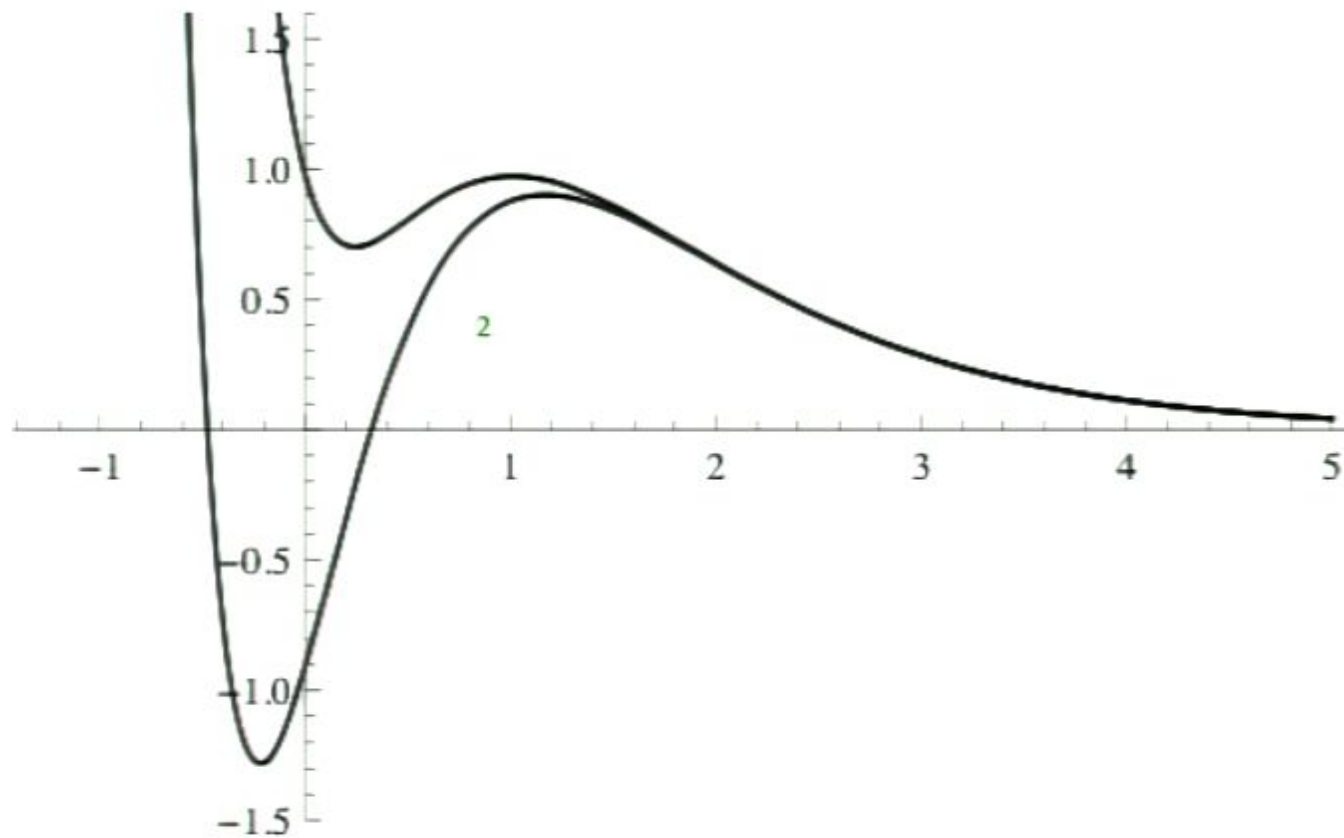
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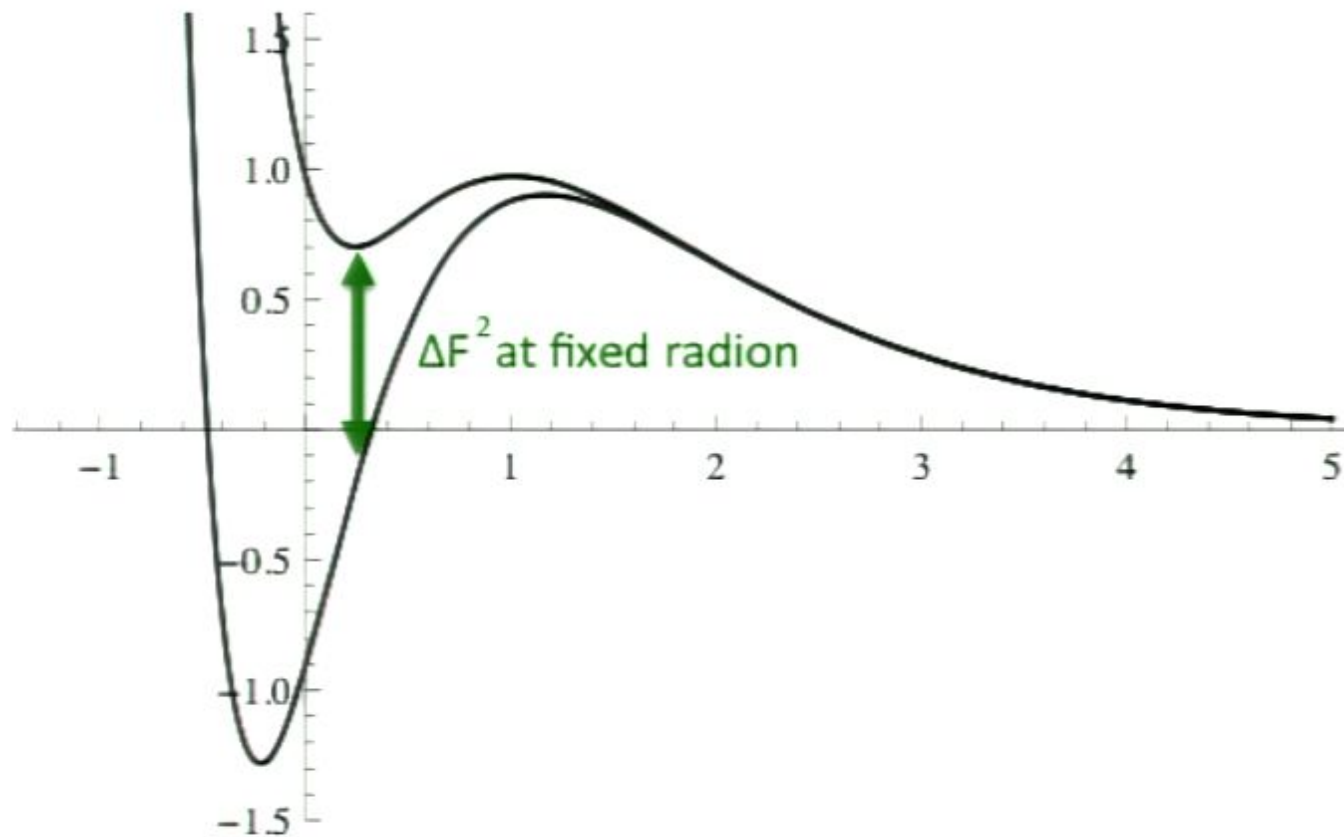
2 reasons back-reaction helps giant leaps

Radion can readjust for extra energy



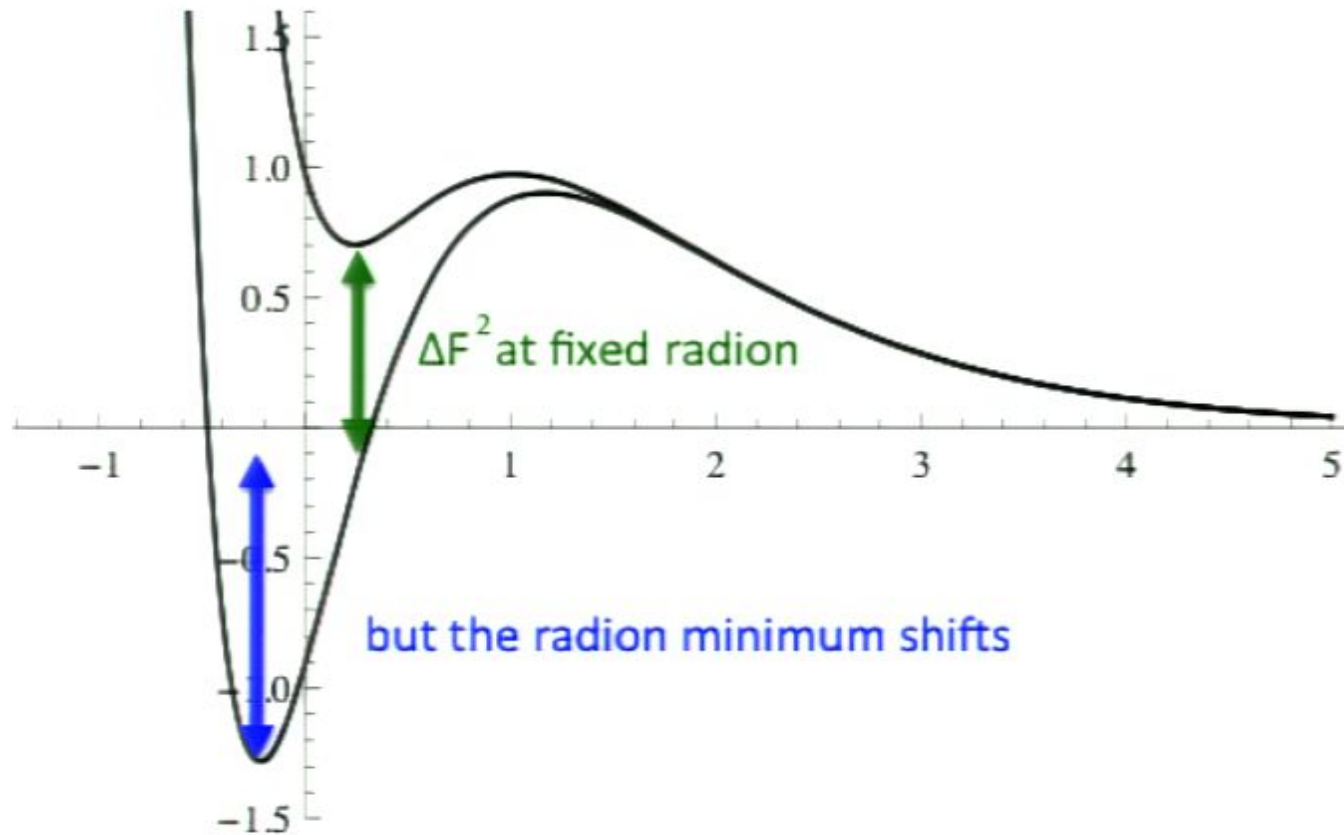
2 reasons back-reaction helps giant leaps

Radiation can readjust for extra energy



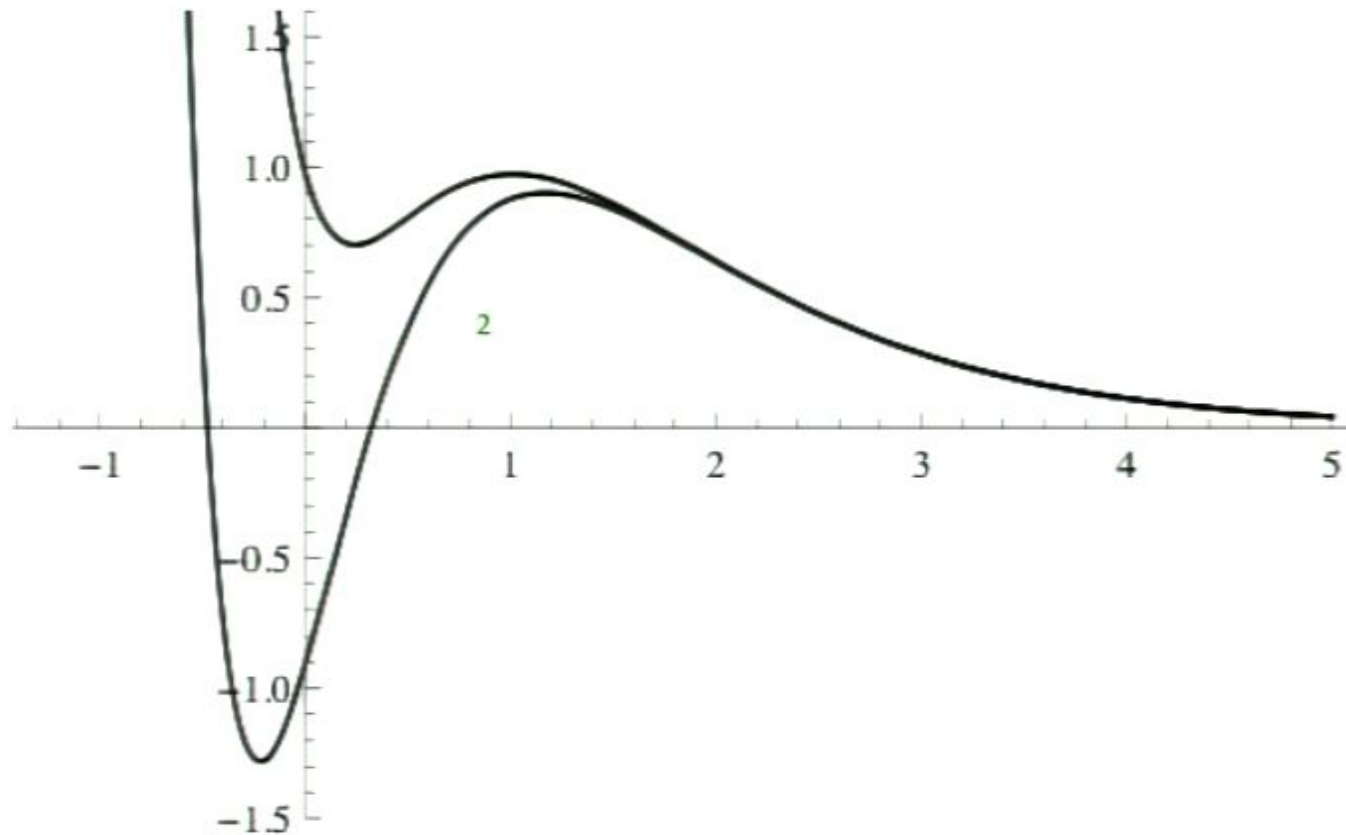
2 reasons back-reaction helps giant leaps

Radiation can readjust for extra energy



2 reasons back-reaction helps giant leaps

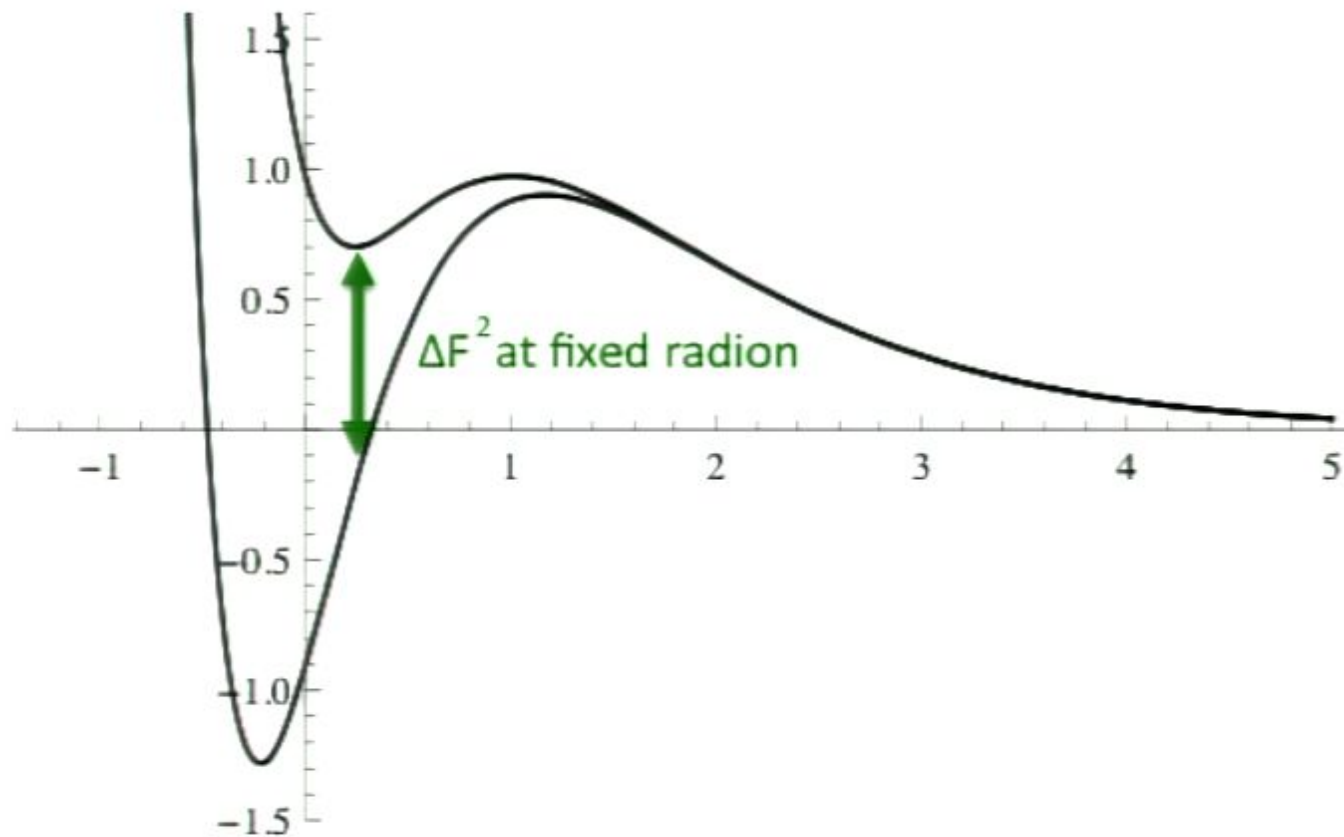
Radion can readjust for extra energy



1 reason it hurts giant leaps:

2 reasons back-reaction helps giant leaps

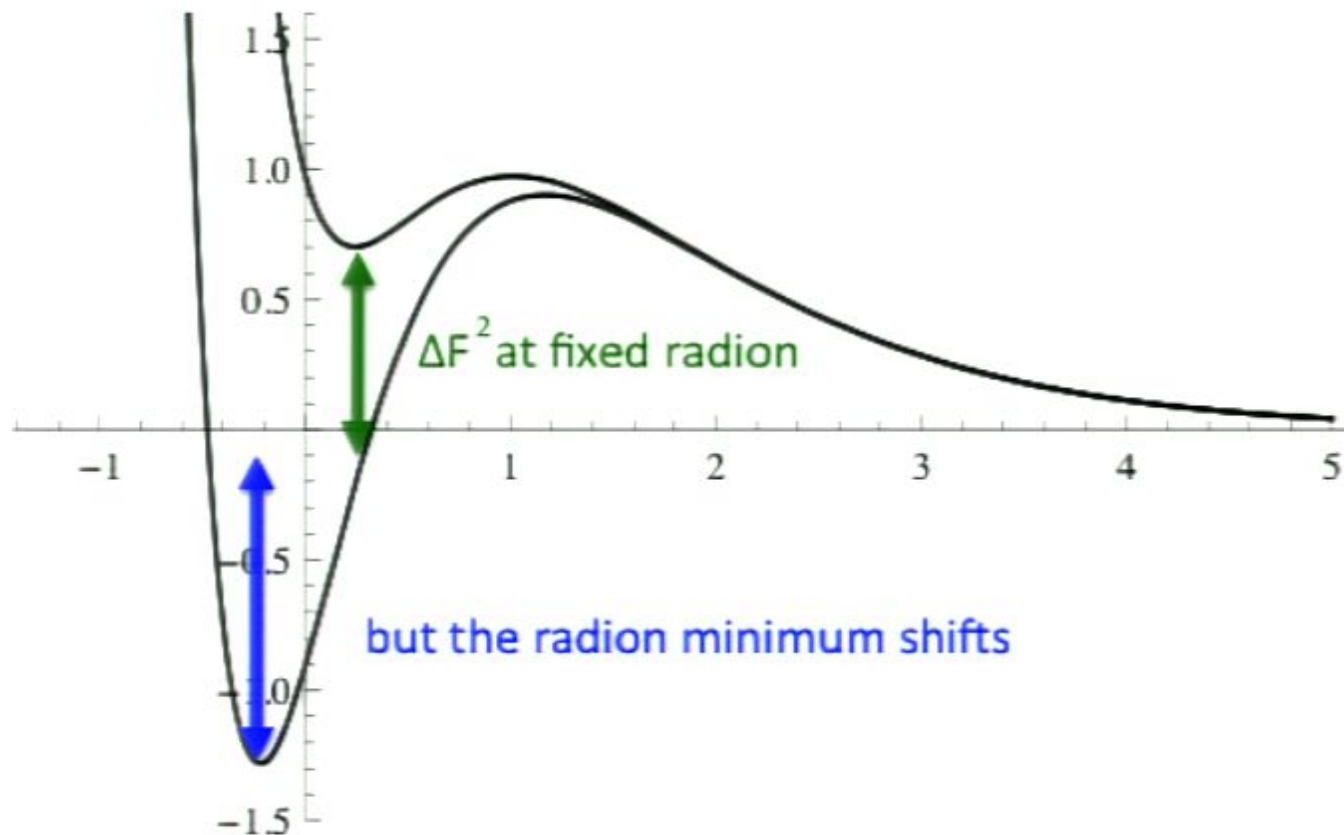
Radiation can readjust for extra energy



1 reason it hurts giant leaps:

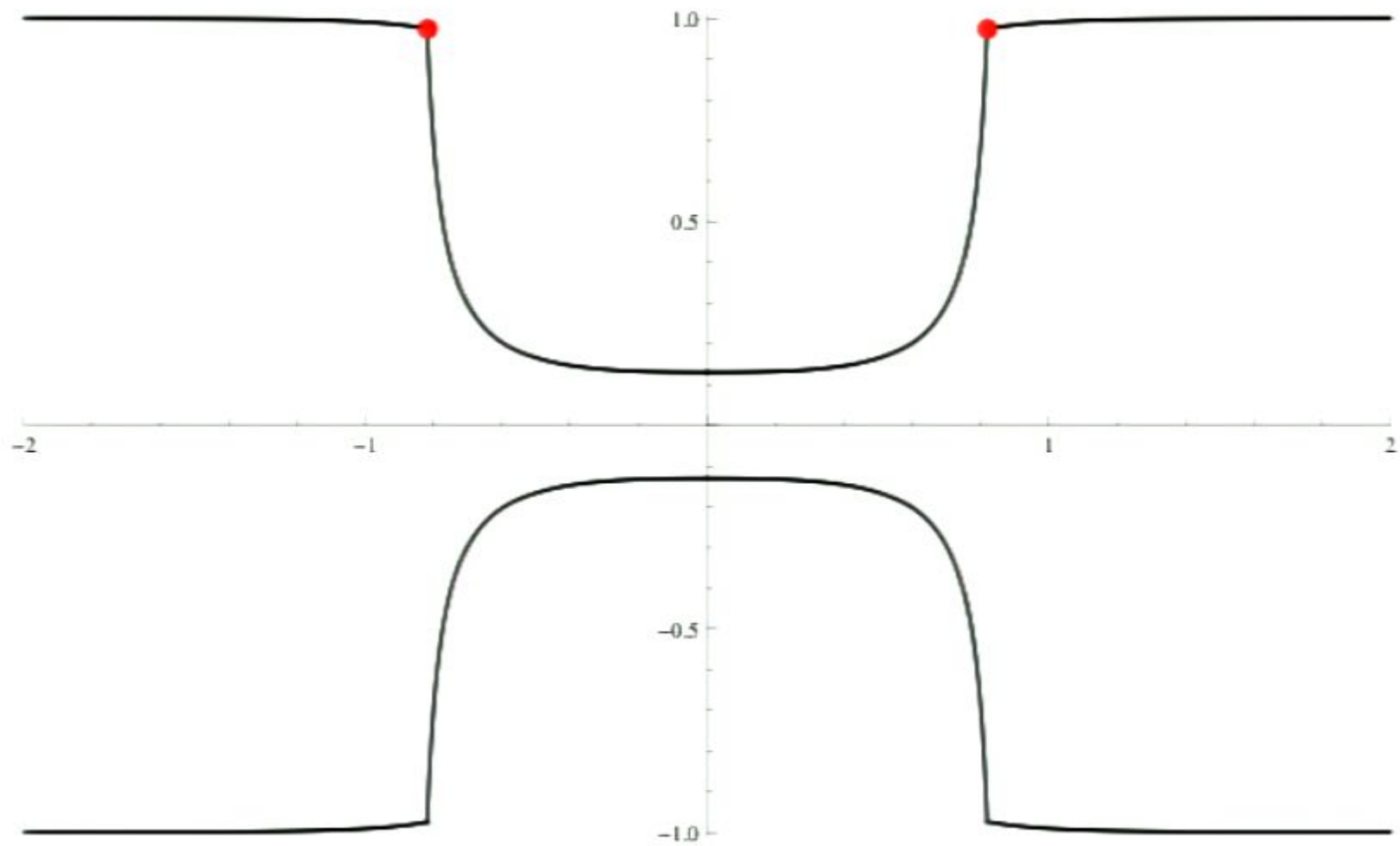
2 reasons back-reaction helps giant leaps

Radion can readjust for extra energy



1 reason it hurts giant leaps:

Sample instanton profile



Let's look at two extreme examples

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2$$

$$T = \frac{2}{\sqrt{3}} \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{\frac{1}{2}}$$

MONOFLUX

A single type of flux and many units of it

$$\mathfrak{N} = 1 \quad N \gg 1 \quad F^2 = g^2 N^2$$

MULTIFLUX

Many different types of flux, each with a single unit

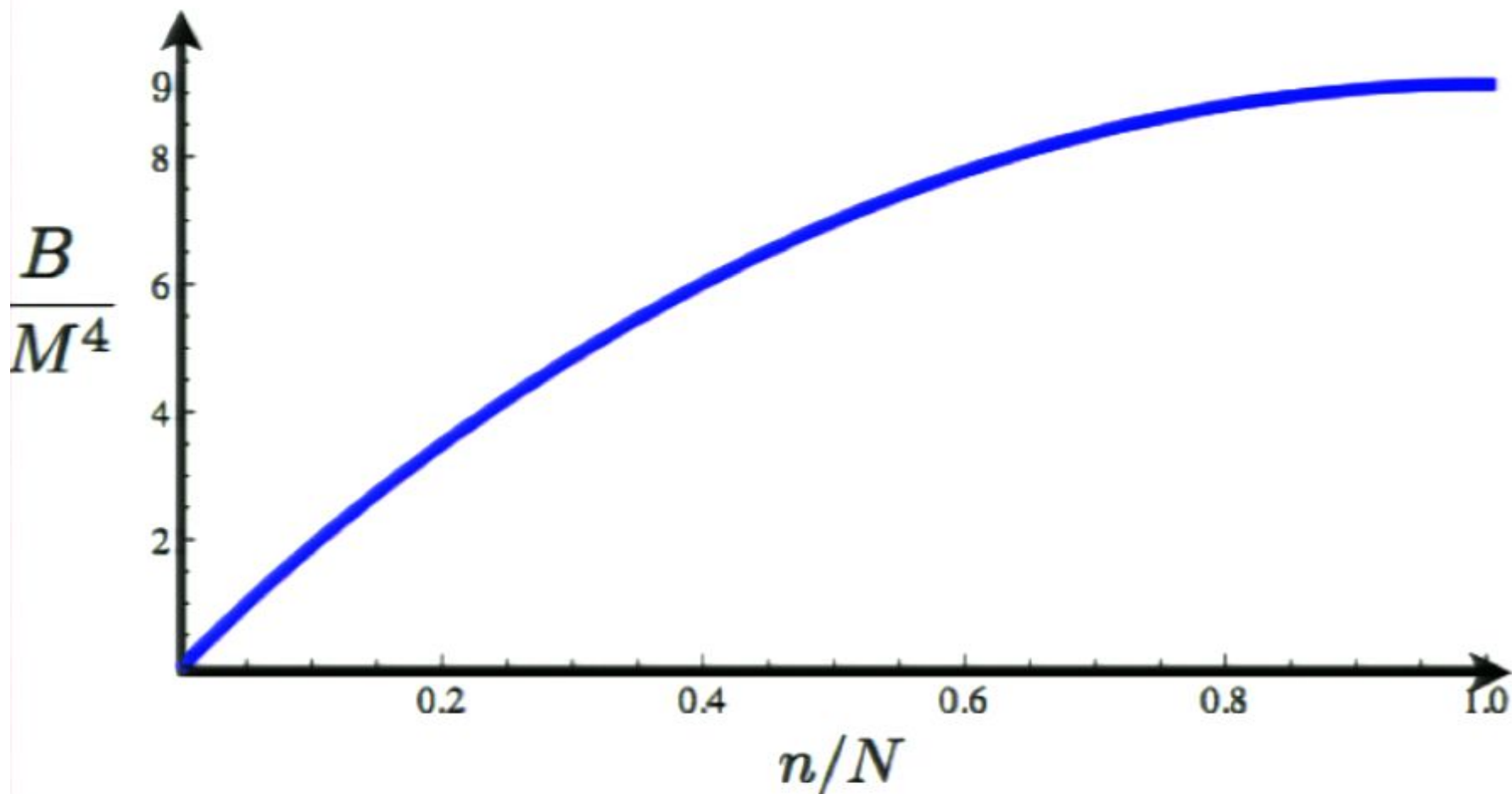
$$\mathfrak{N} \gg 1 \quad N_i = 1 \quad F^2 = g^2 \mathfrak{N}$$

Decay Rates

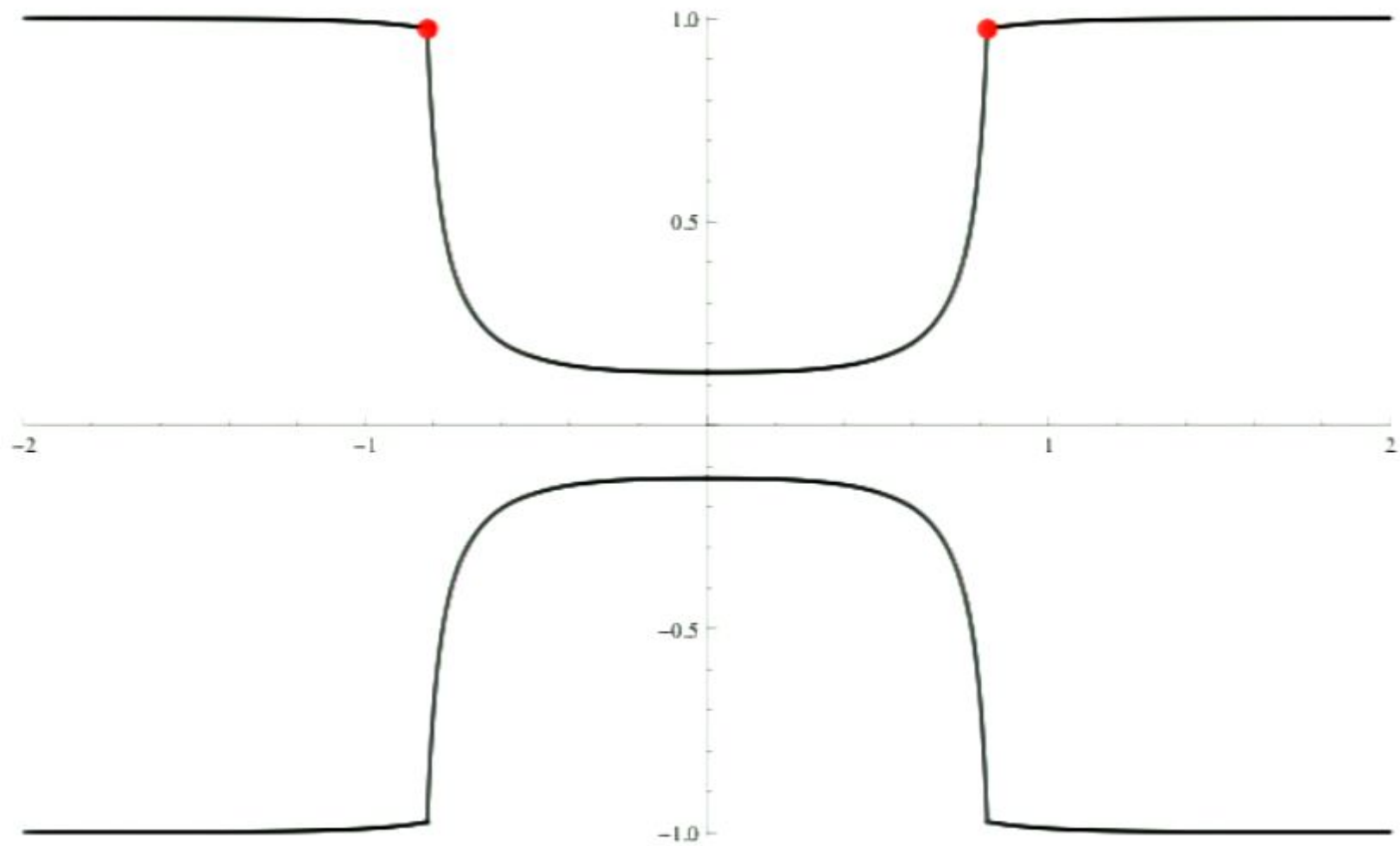
$$\Gamma \sim e^{-B/\hbar}, \quad B = S_E(\text{instanton}) - S_E(\text{false vacuum})$$

MONOFLUX

From $V=0$



Sample instanton profile



Let's look at two extreme examples

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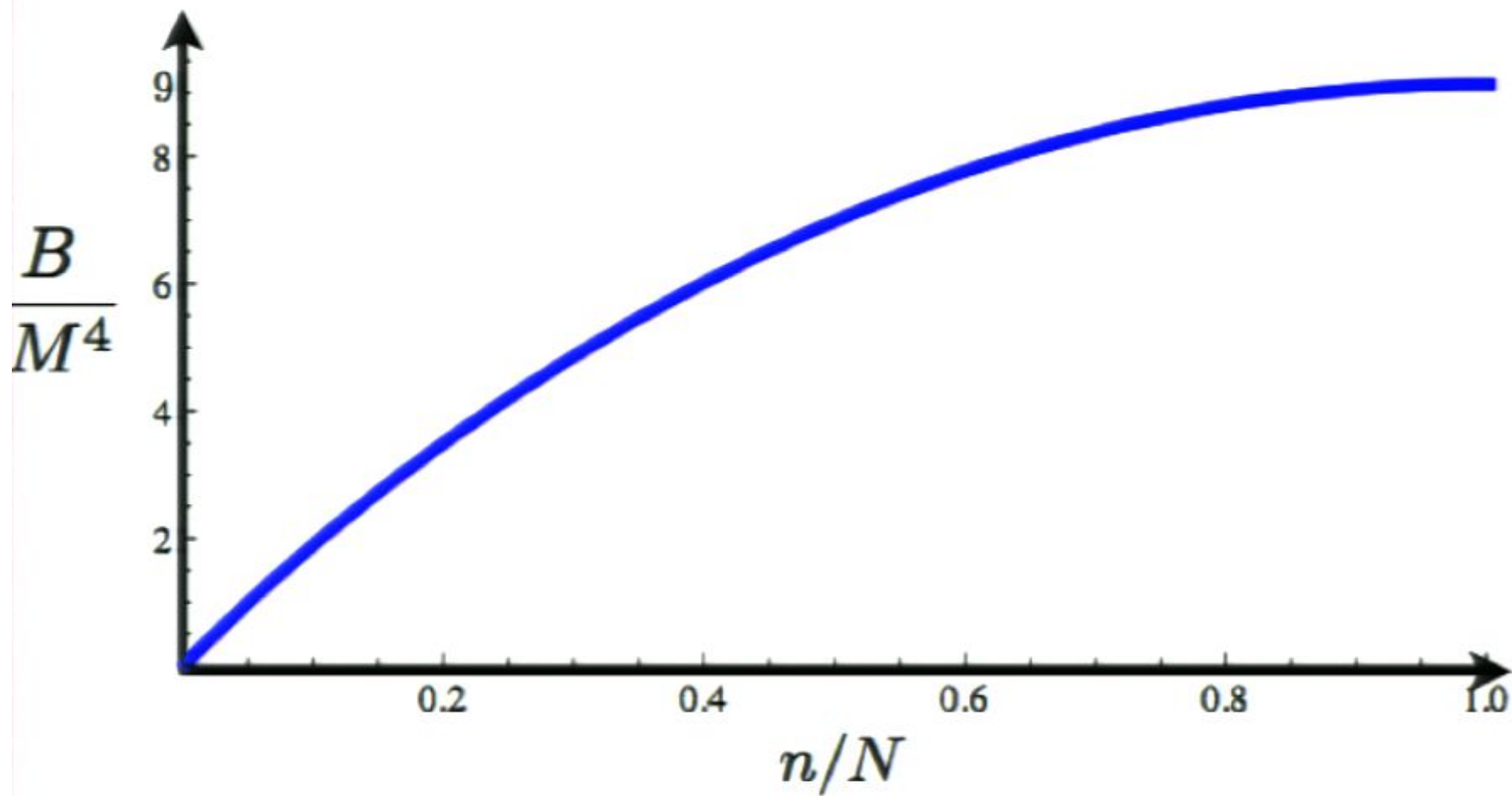
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MONOFLUX

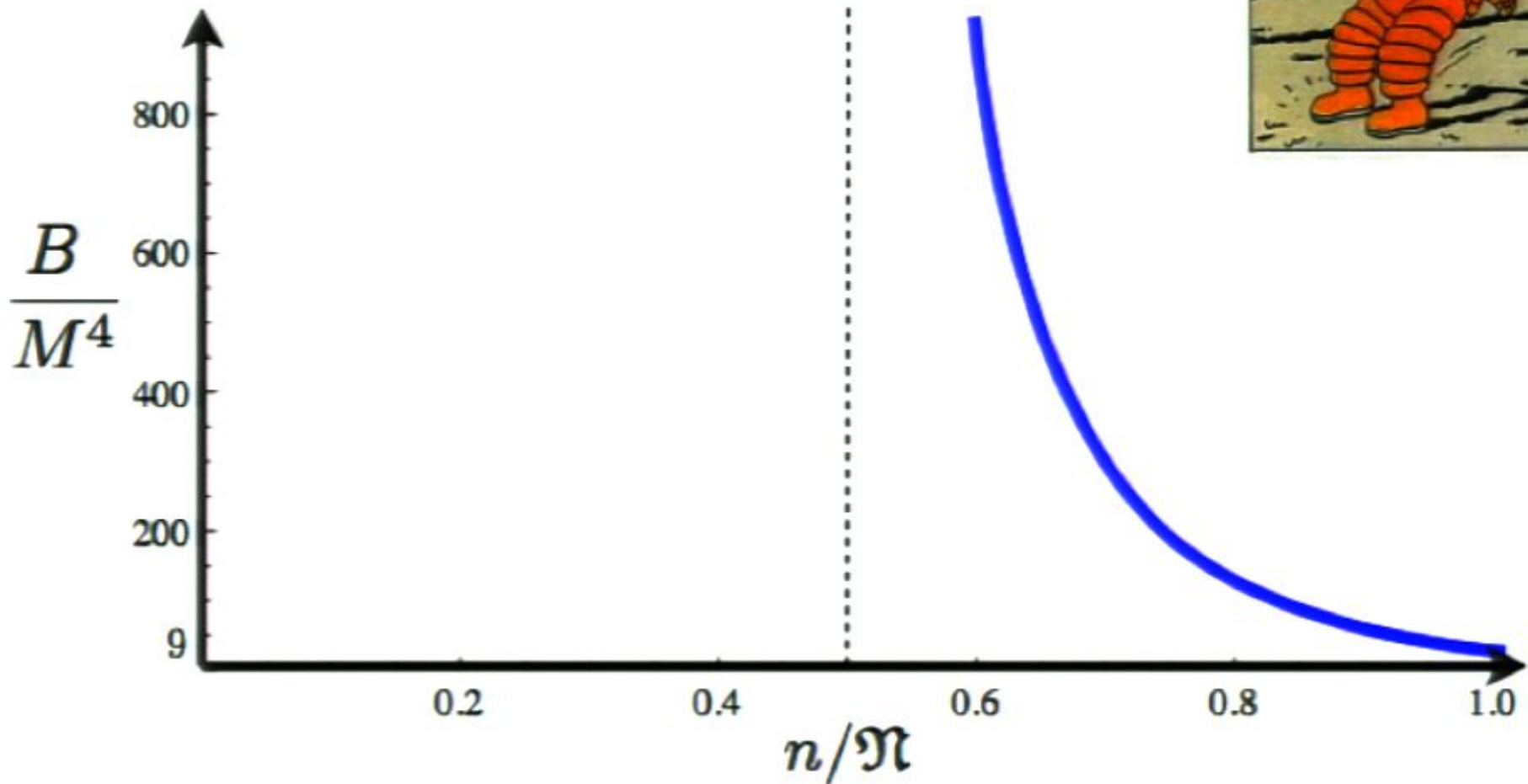
From $V=0$



Decay Rates

MULTIFLUX

From $V=0$



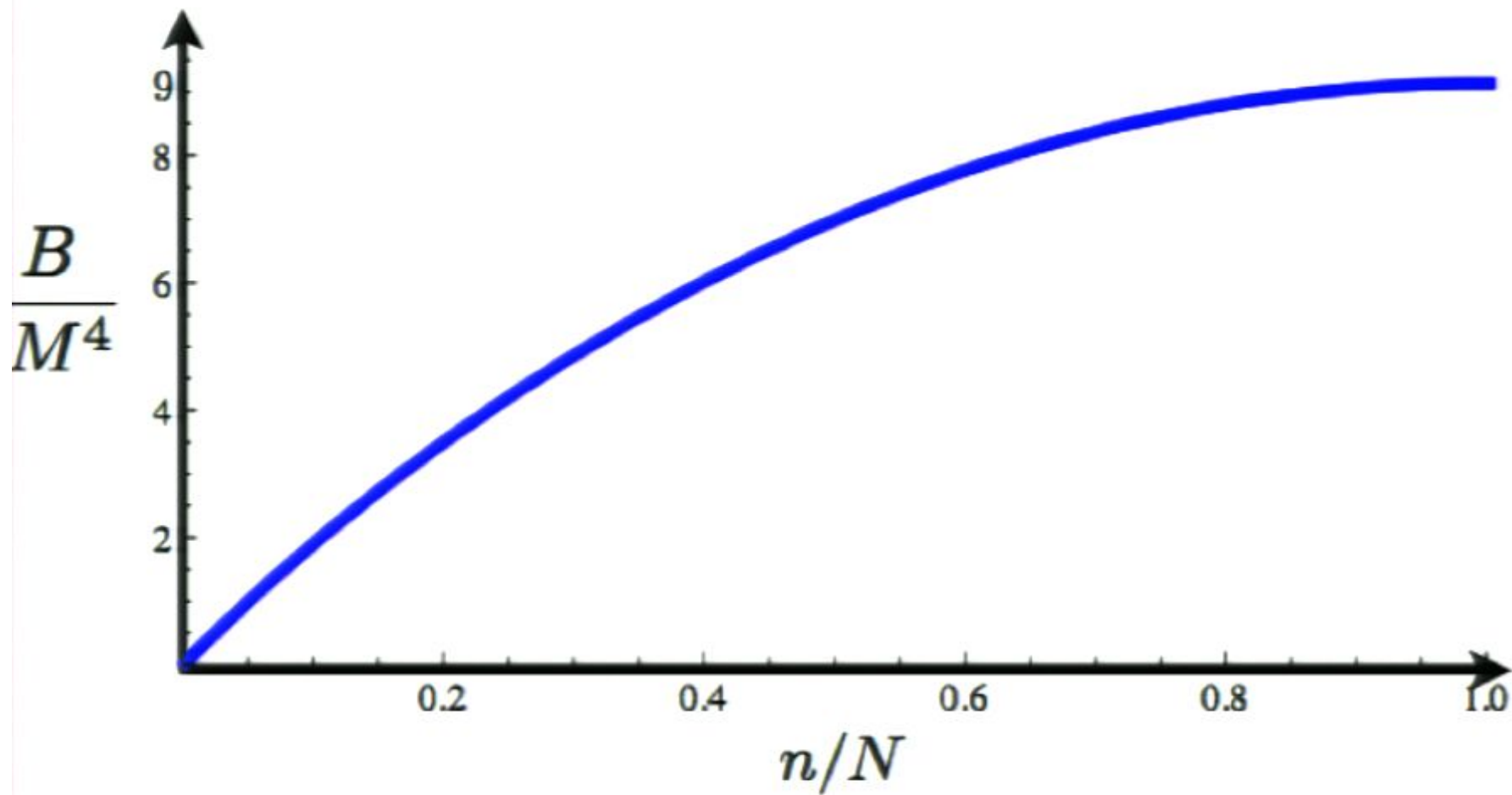
GIANT LEAPS WIN

Decay Rates

$$\Gamma \sim e^{-B/\hbar}, \quad B = S_E(\text{instanton}) - S_E(\text{false vacuum})$$

MONOFLUX

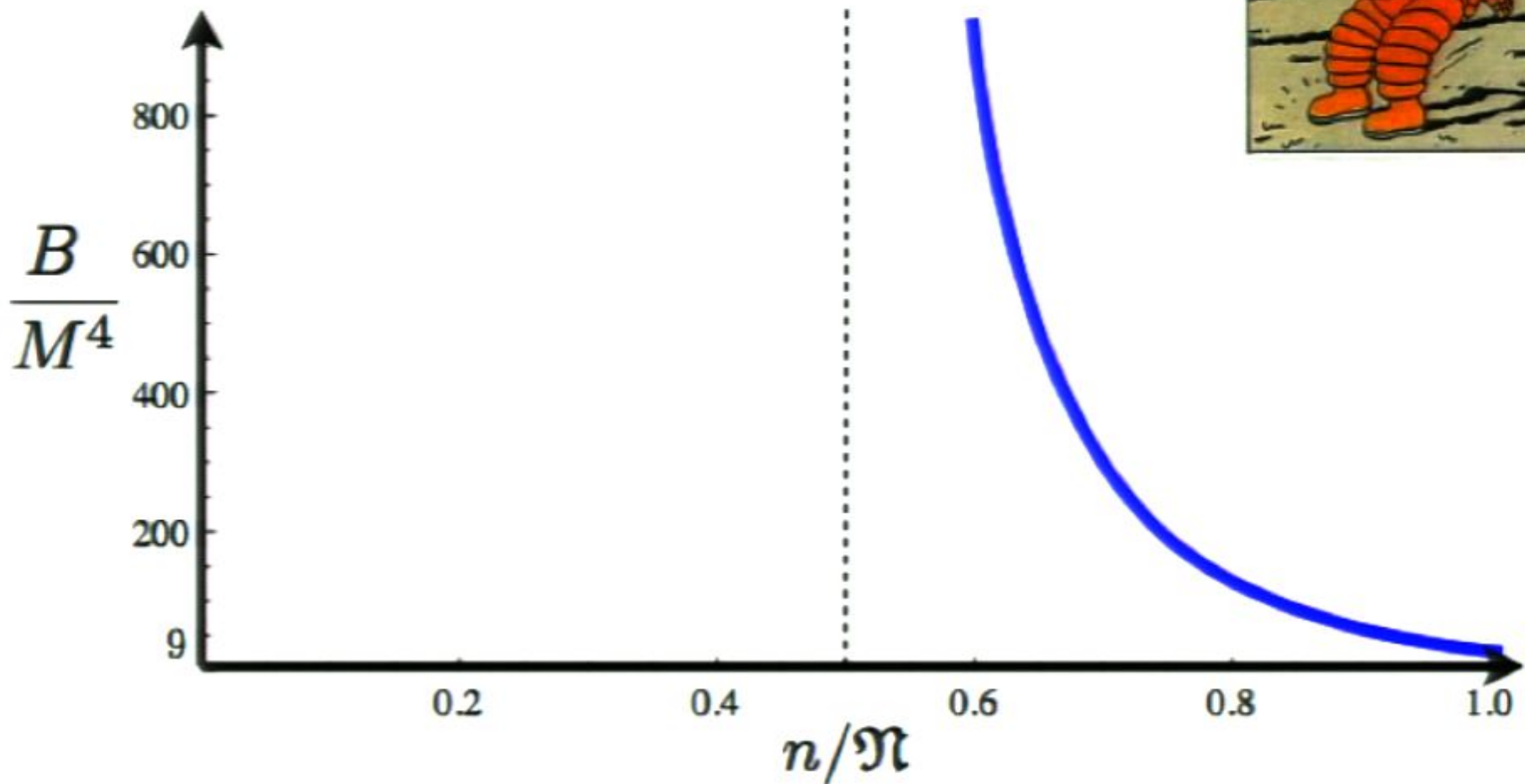
From $V=0$



Decay Rates

MULTIFLUX

From $V=0$

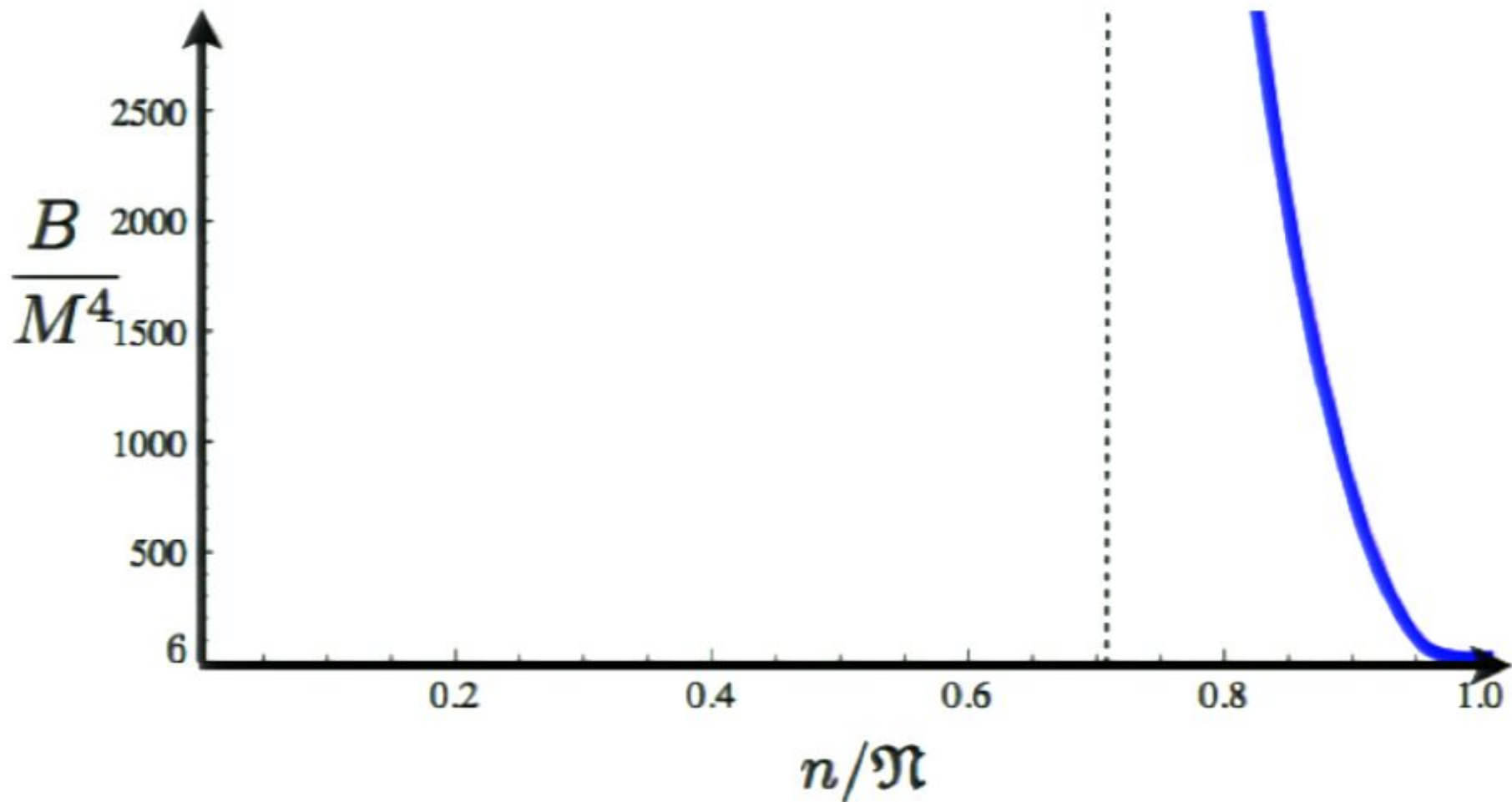


GIANT LEAPS WIN

Decay Rates

MULTIFLUX

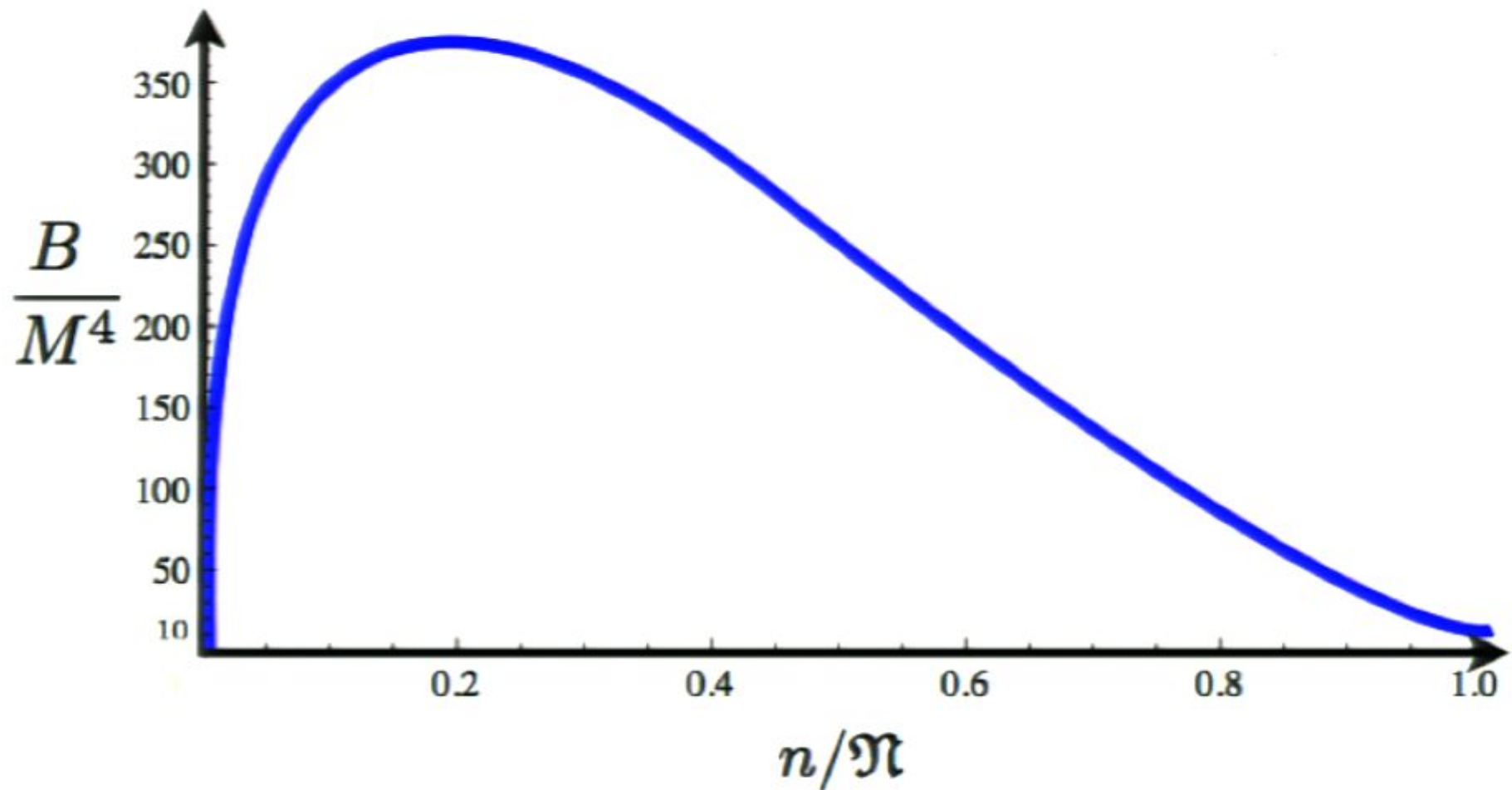
From $V < 0$

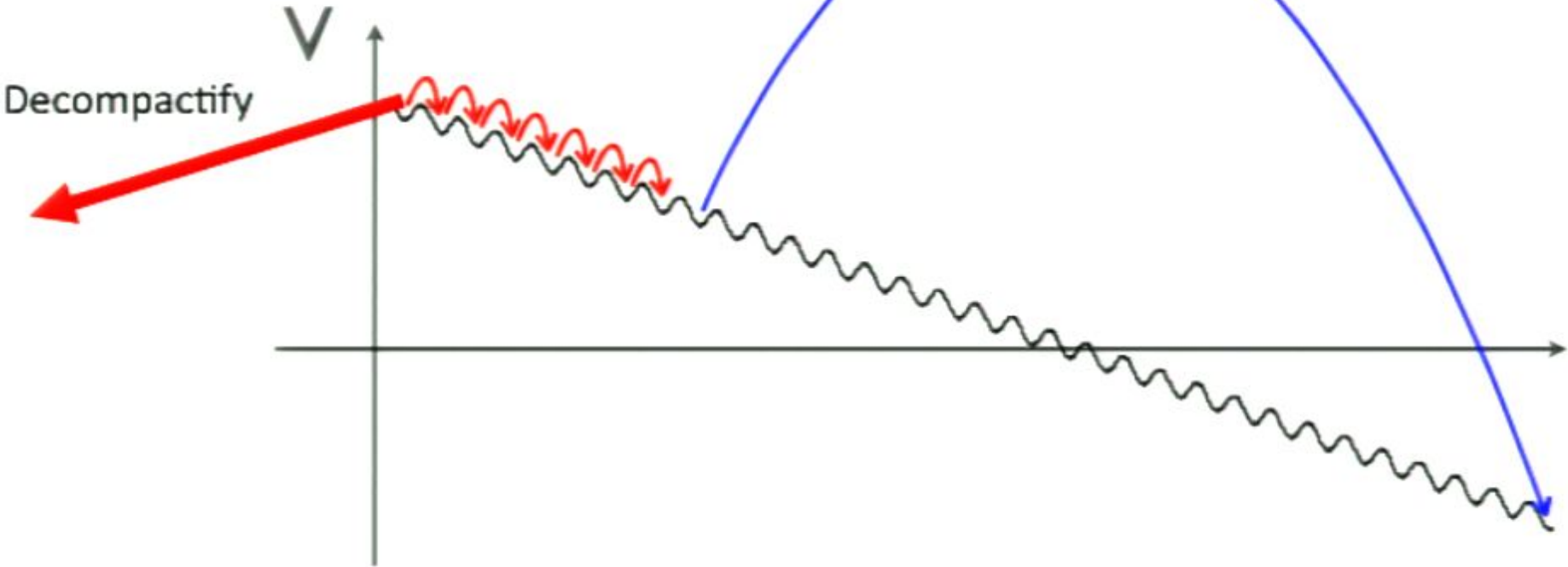


Decay Rates

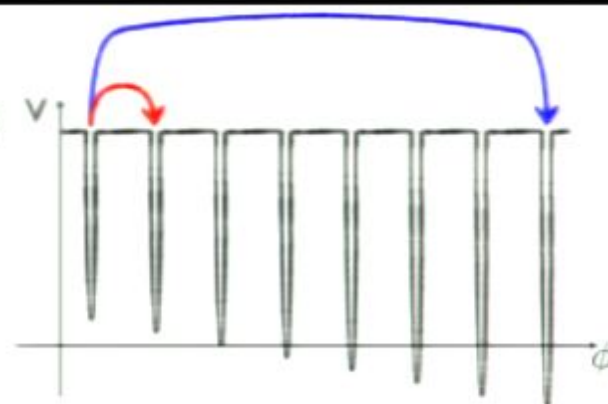
MULTIFLUX

From $V > 0$





This potential gives small steps



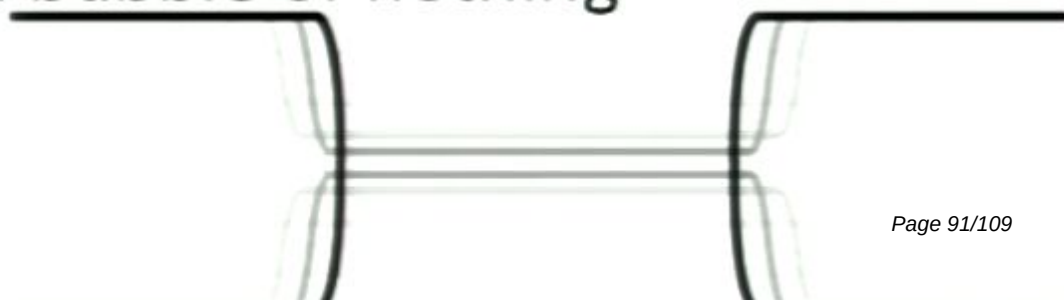
Flux compactifications with many fluxes give giant leaps

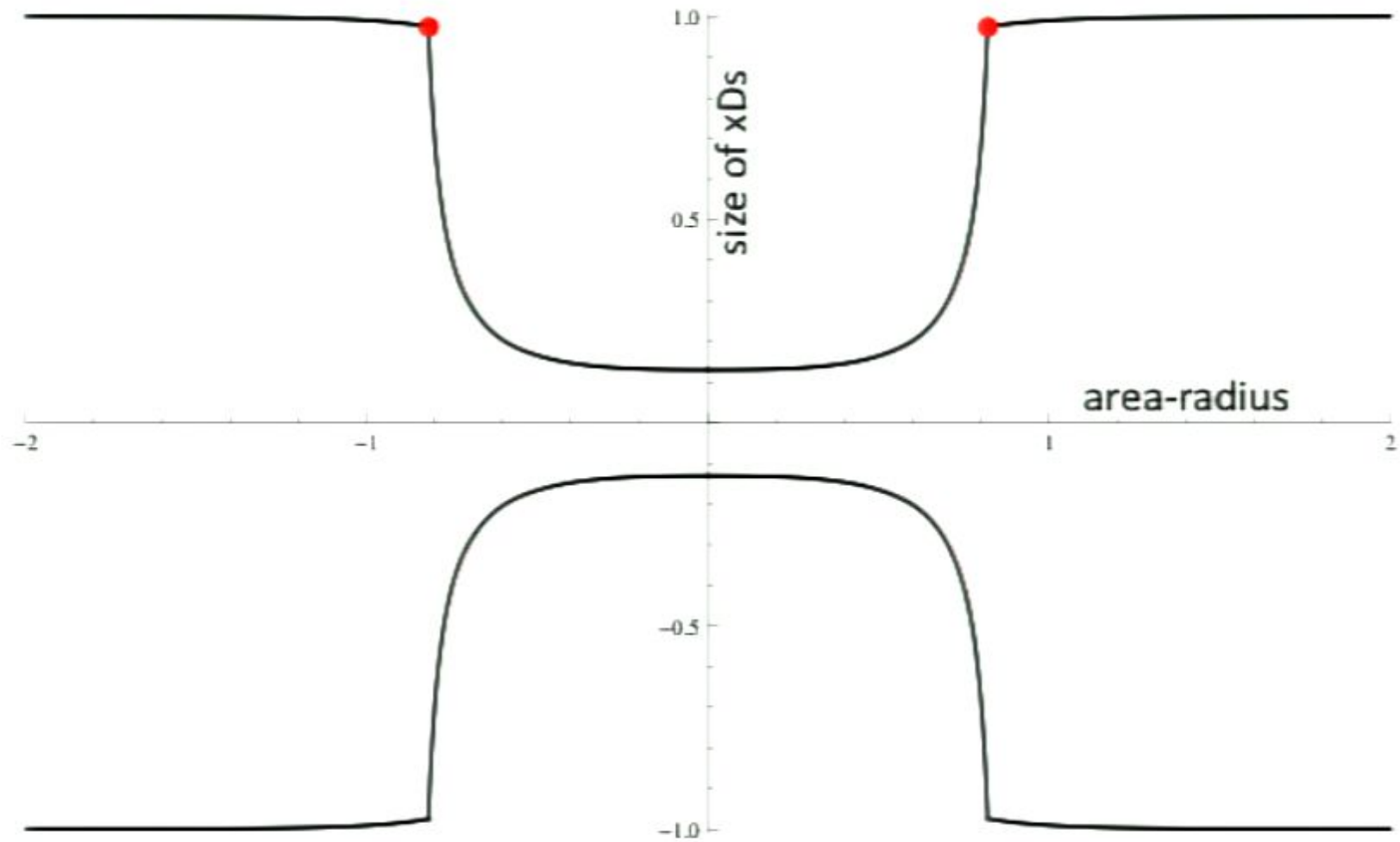
Monkey branes

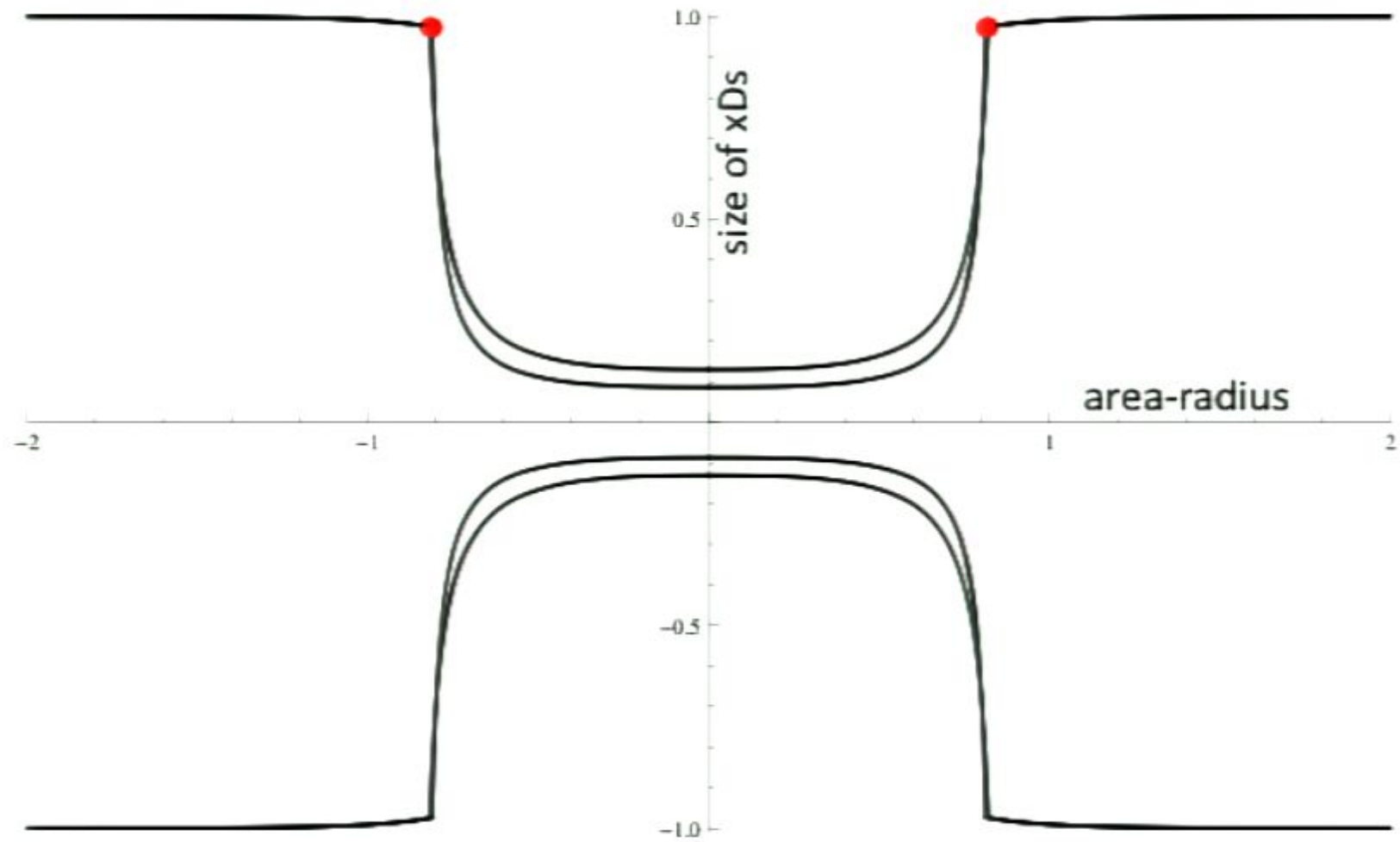


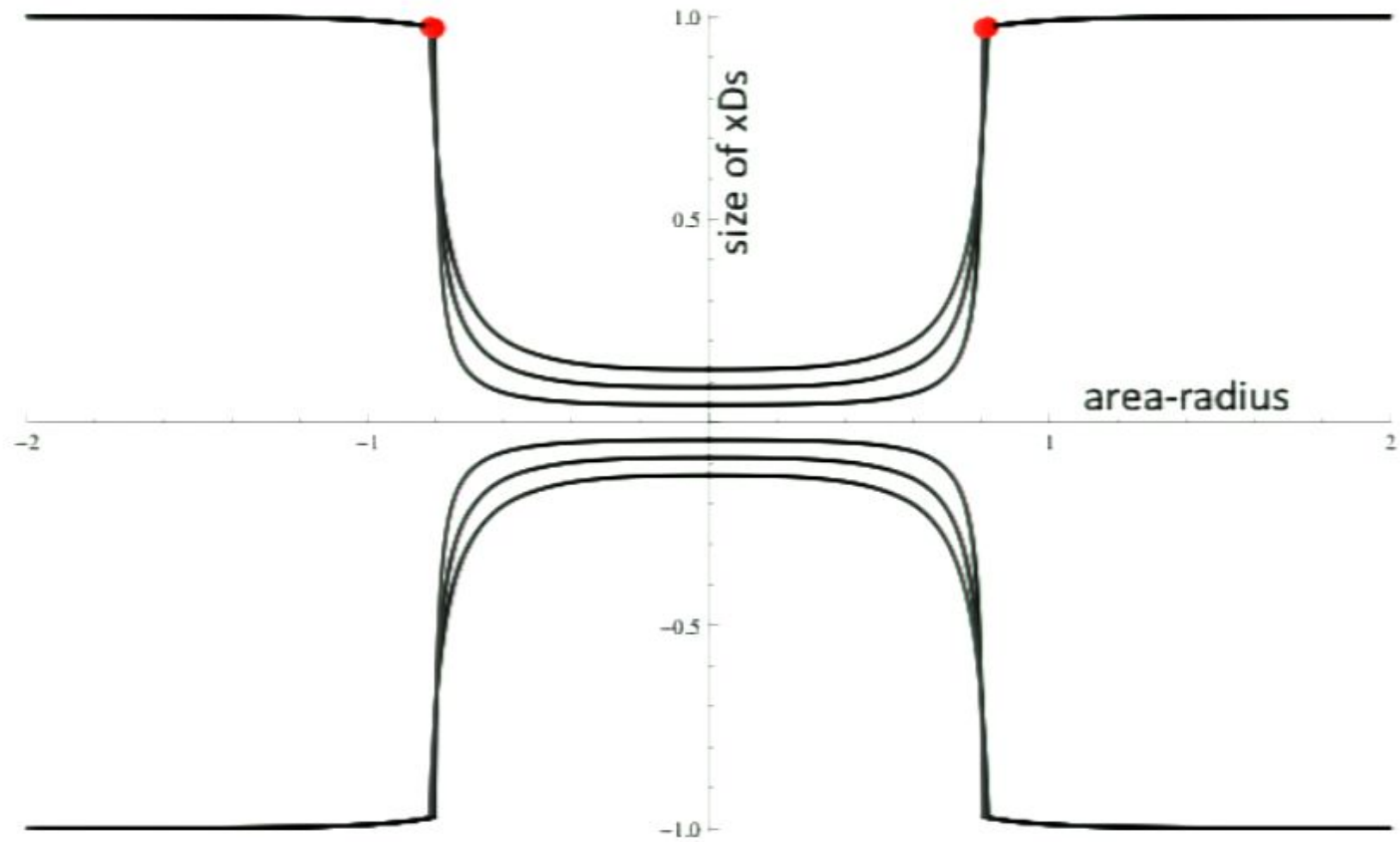
Adding back-reaction

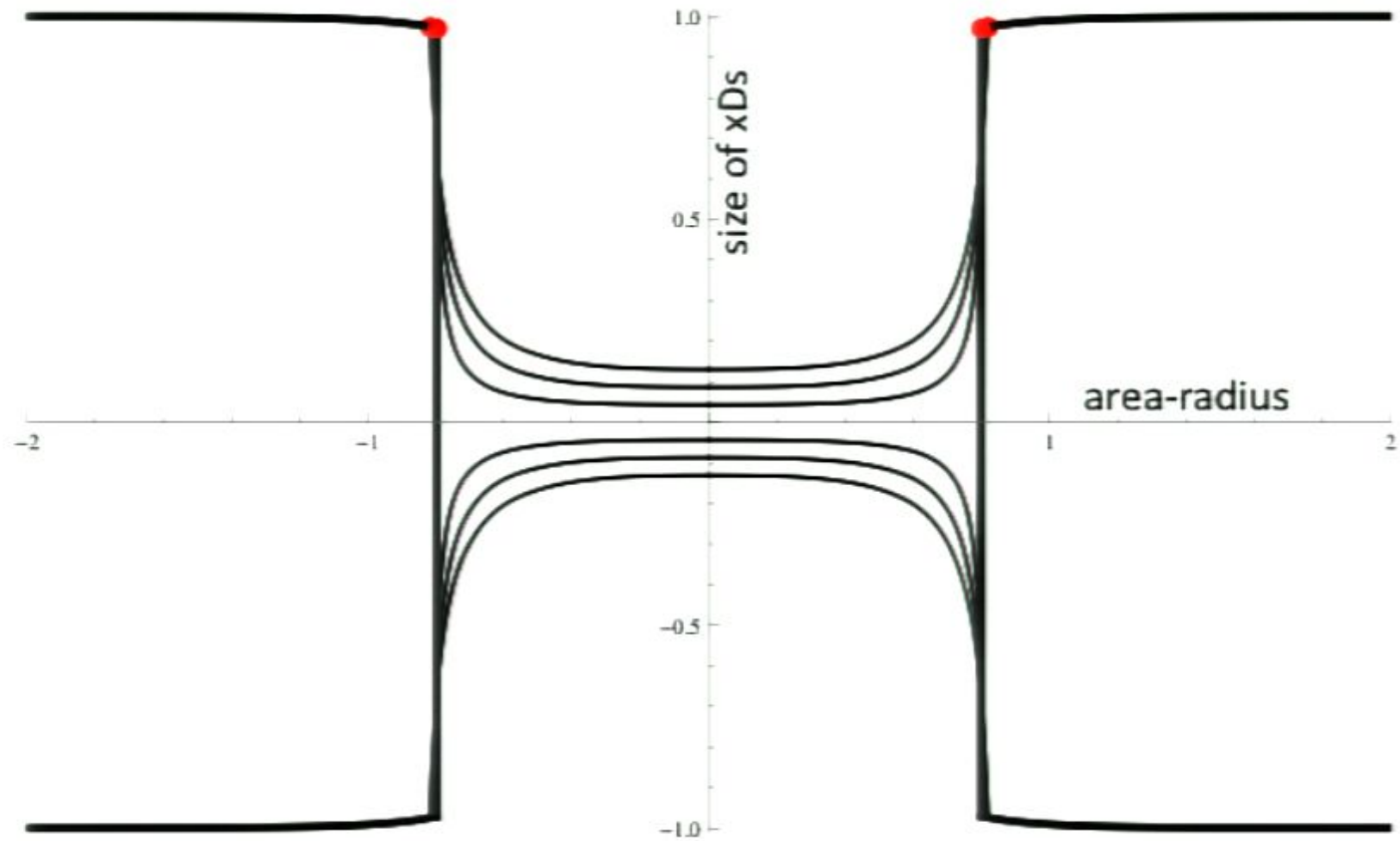
The giantest leap of all is a bubble of nothing









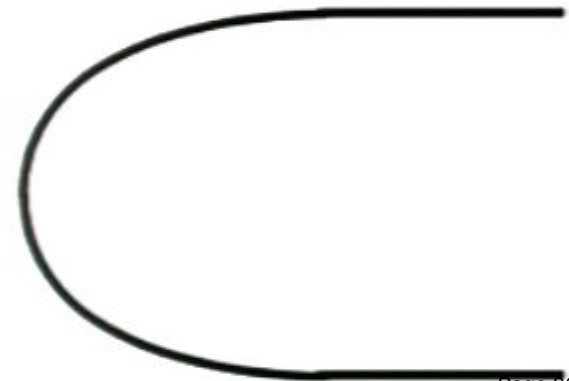
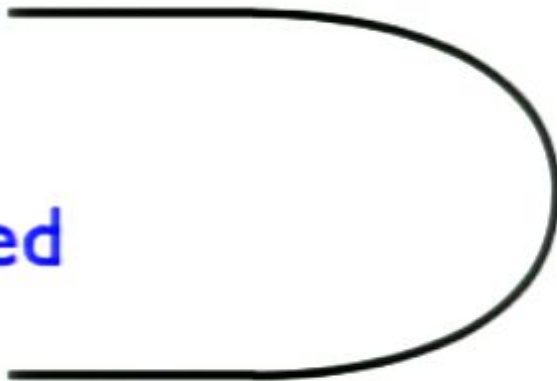


ubble of Nothing

6d
stabilized



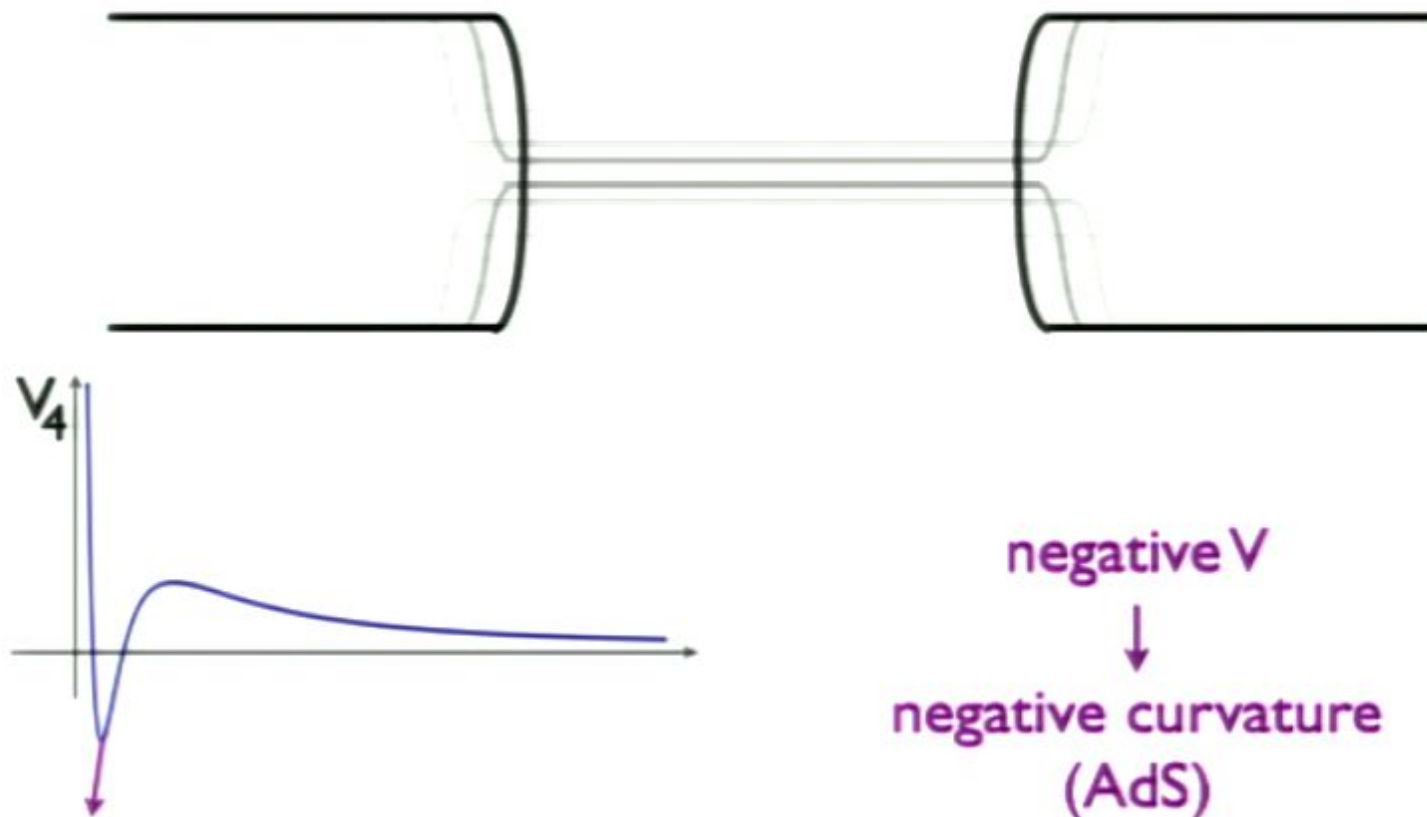
5d
unstabilized



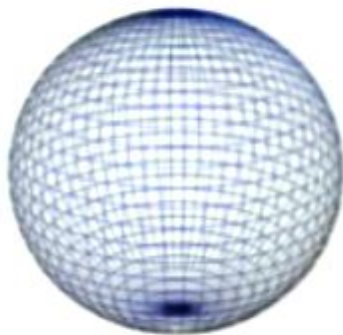


In the limit that **ALL** flux discharged:

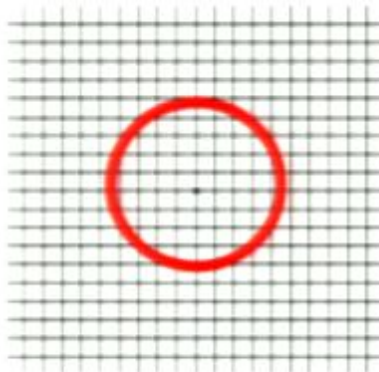
- I. Extra dimensions shrink to zero size
- II. What about the **3-volume** of a slice through the bubble?
like an infinitely thin pancake?



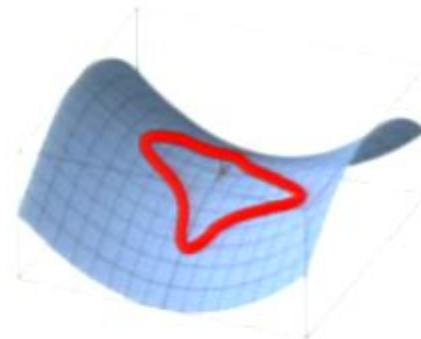
II. What about the **3-volume** of a slice through the bubble?
like an infinitely thin pancake?



$$\text{Area} > \pi r^2$$

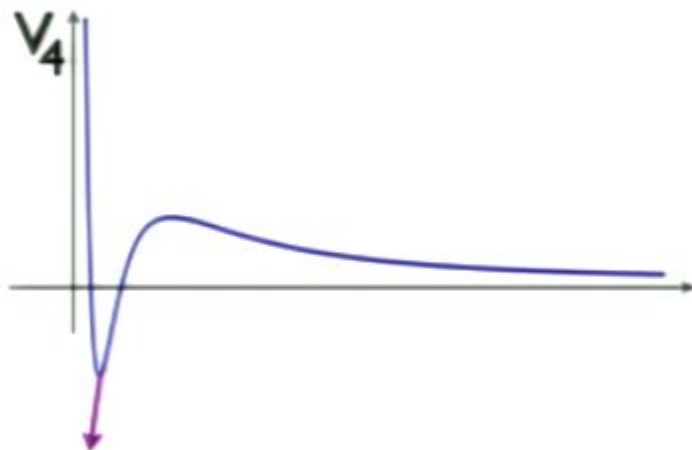


$$\text{Area} = \pi r^2$$



$$\text{Area} < \pi r^2$$

$$\text{Area} \sim r l_{\text{curv}}$$



negative V



negative curvature
(AdS)

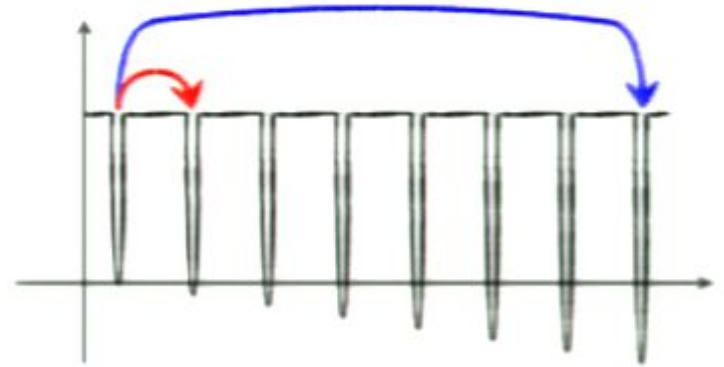
II. What about the **3-volume** of a slice through the bubble?
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$$\text{Volume} \sim \text{Area} \times l_{\text{curv}}$$
$$l_{\text{curv}} \rightarrow 0$$

II. What about the **3-volume** of a slice through the bubble?
like an infinitely thin pancake?

This potential gives small steps



Flux compactifications with many fluxes give giant leaps
Monkey branes

Back-reaction on the compactification

The giantest leap of all is a bubble of nothing

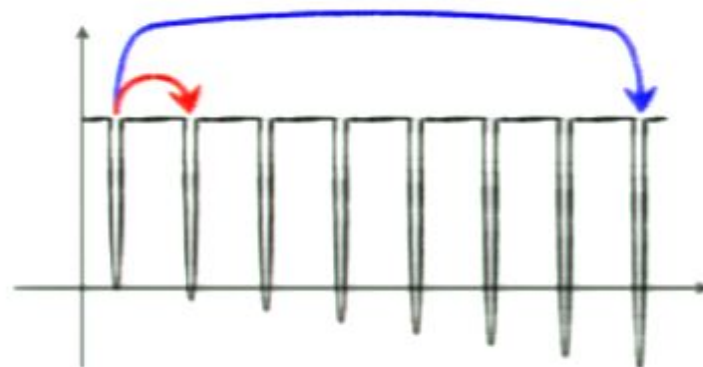




$$\text{Volume} \sim \text{Area} \times l_{\text{curv}}$$
$$l_{\text{curv}} \rightarrow 0$$

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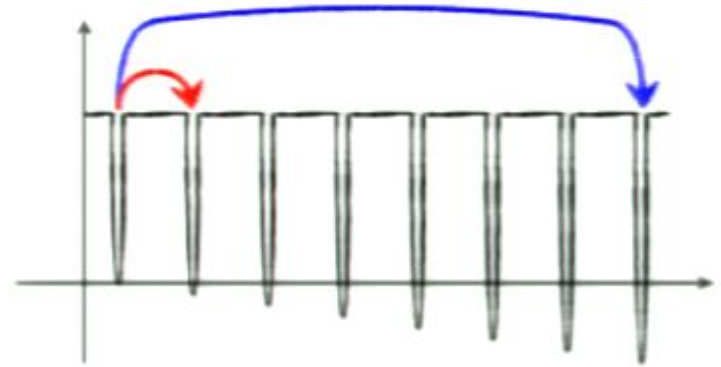
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Back-reaction on the compactification

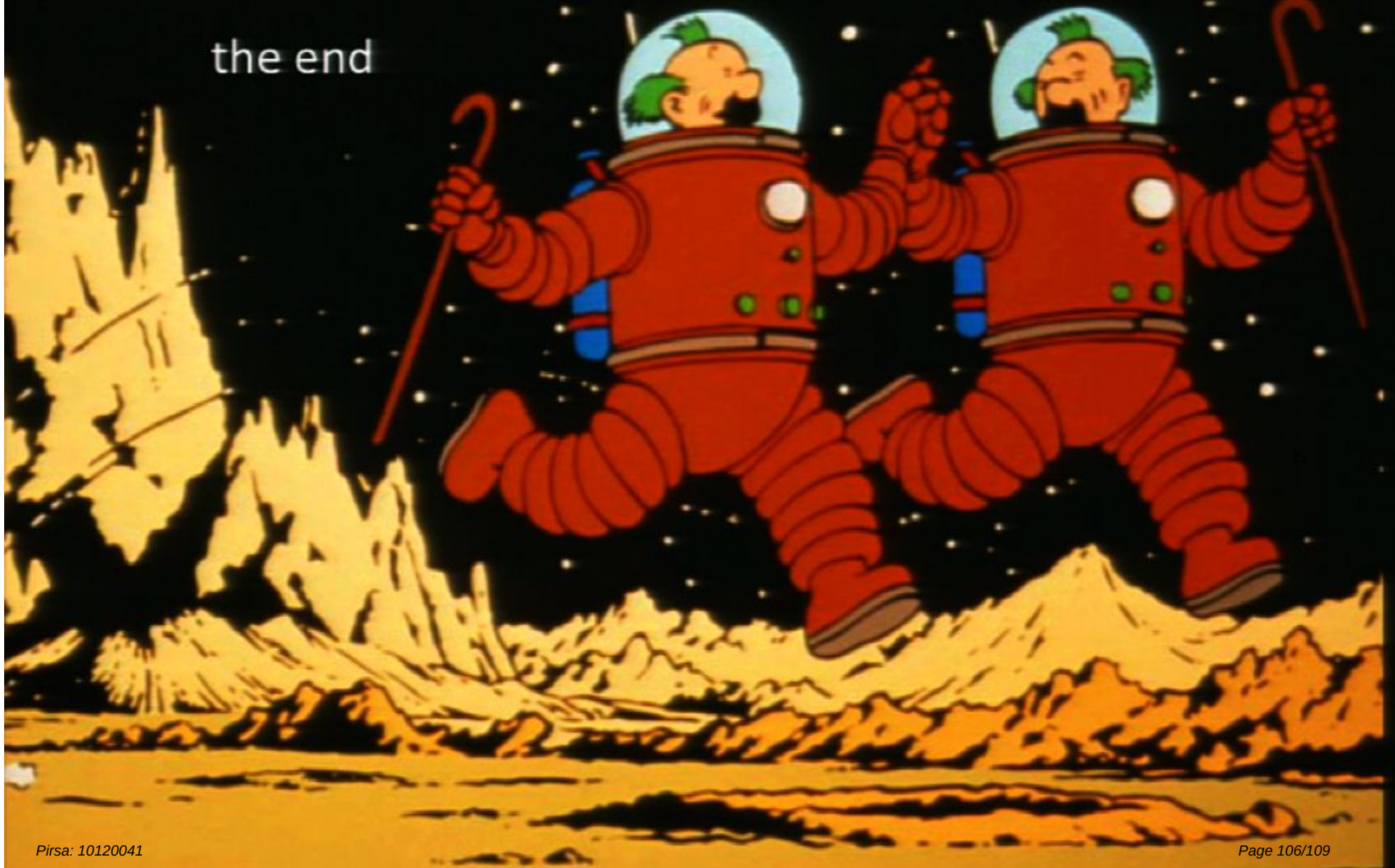
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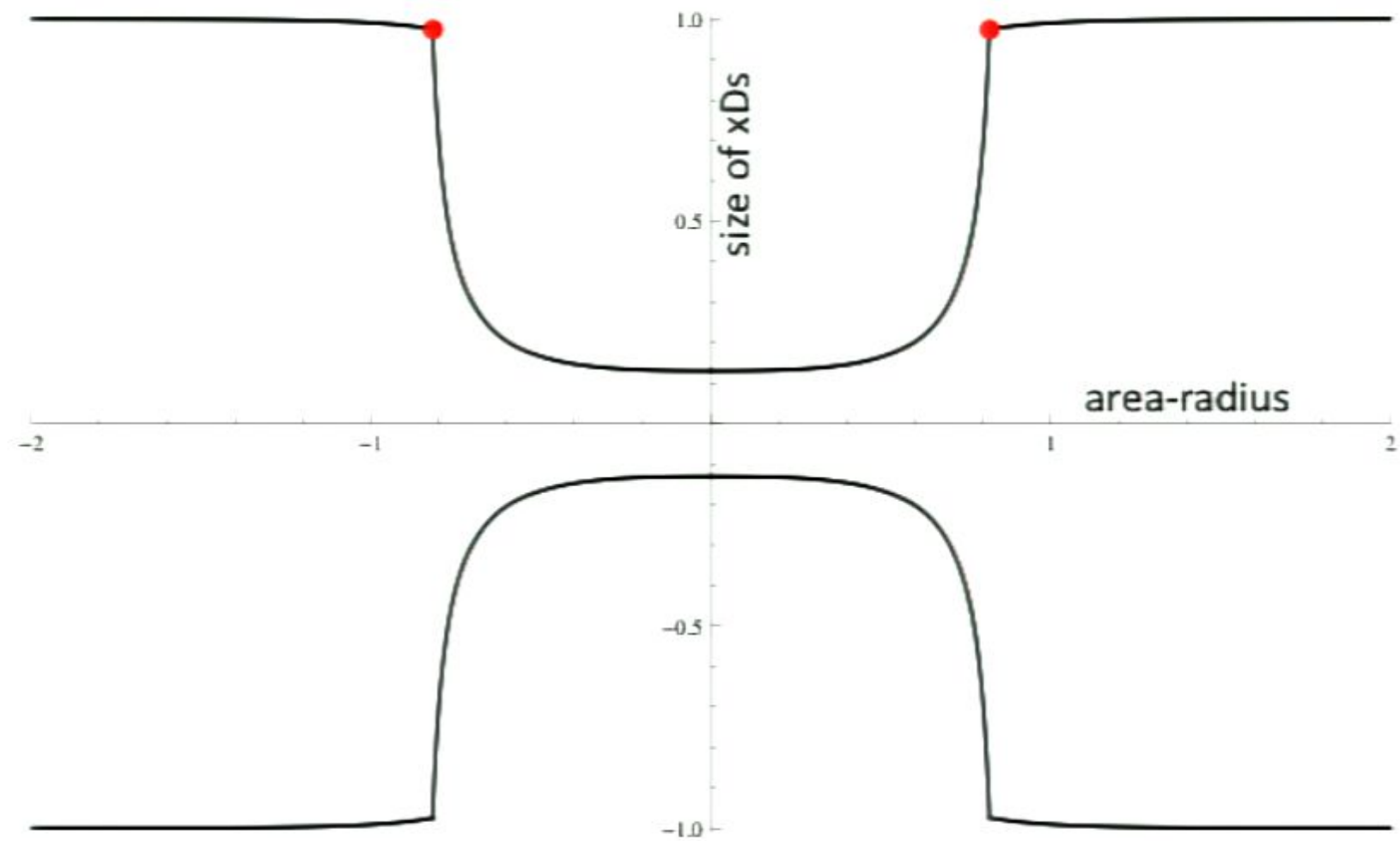


Summary

- Giant leaps beat small steps in landscapes
- The bubble of nothing can be the fastest decay
- It is realized smoothly as the limit of flux tunneling
- When the many fluxes come from higher dimensional model, still expect enhancement. Richer behavior.
- If we live in a multflux landscape, then we draw two conclusions. We got here by an exponentially subdominant decay, and we will leave here by a bubble of nothing.

the end

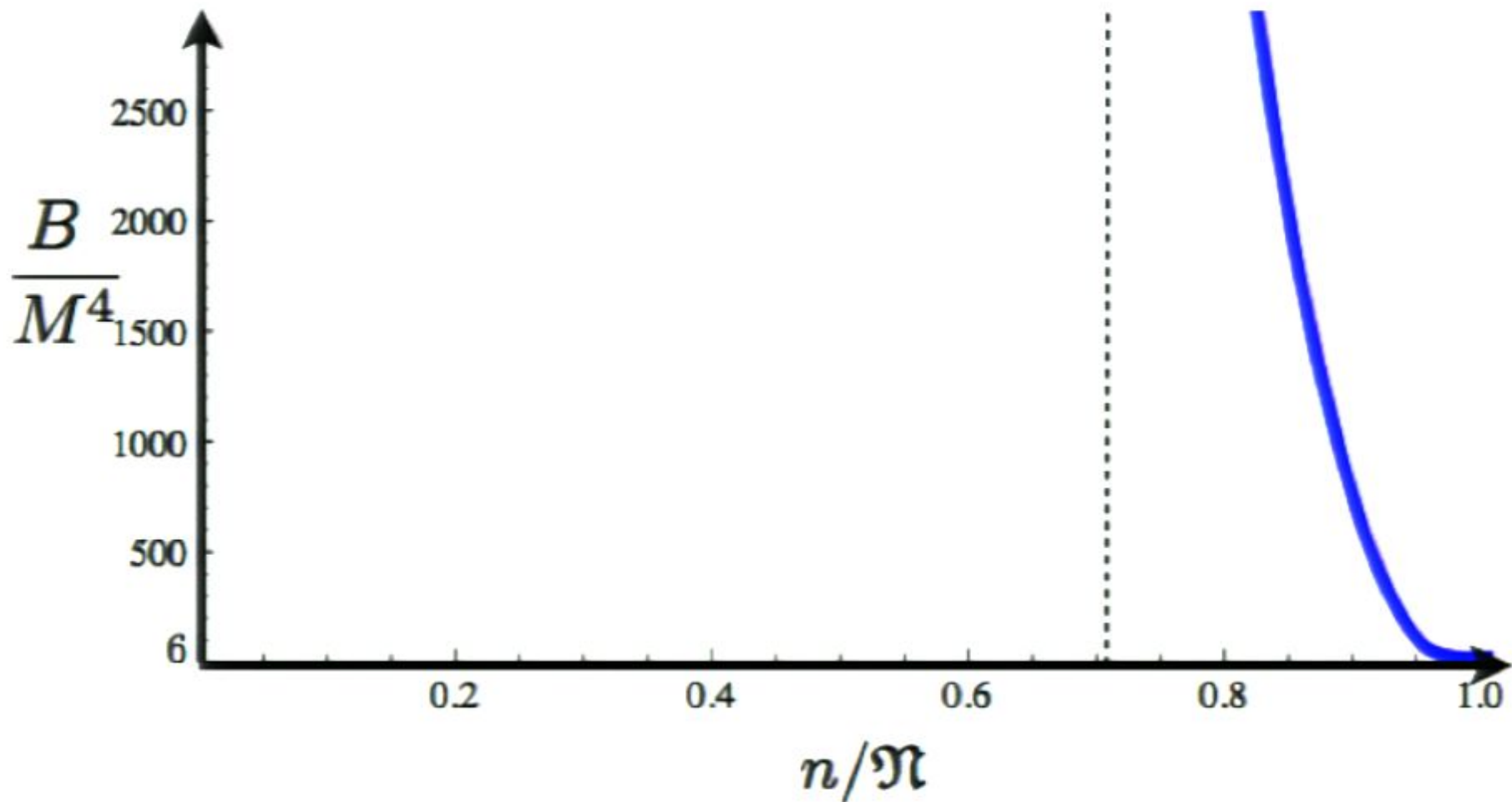




Decay Rates

MULTIFLUX

From $V < 0$



Decay Rates

$$\Gamma \sim e^{-B/\hbar}, \quad B = S_E(\text{instanton}) - S_E(\text{false vacuum})$$

MONOFLUX

From $V=0$

