Title: The Physics of Black Hole Interiors: The Most Extreme Physics in the Universe

Date: Nov 10, 2010 02:00 PM

URL: http://pirsa.org/10110059

Abstract: This talk will describe the best current understanding of the interior structure of astronomically realistic black holes.

A common misconception is that matter falling into a black hole simply falls to a central singularity, and that's that.

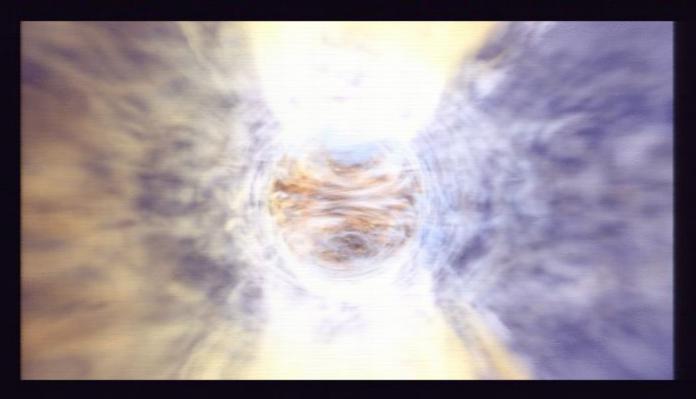
Reality is much more interesting. Rotating black holes have not only outer horizons, but also inner horizons. Penrose (1968) first pointed out that an infaller falling through the inner horizon would see the outside Universe infinitely blueshifted, and he speculated that this would destabilize the inner horizon.

The expectation was supported by linear perturbation theory, but it was not until 1990 that Poisson & Disson &

What does Nature do with this remarkable accelerator?

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The Physics of Black Hole Interiors: The Most Extreme Physics in the Universe



More extreme than the Big Bang? Yes.

Schwarzschild black hole

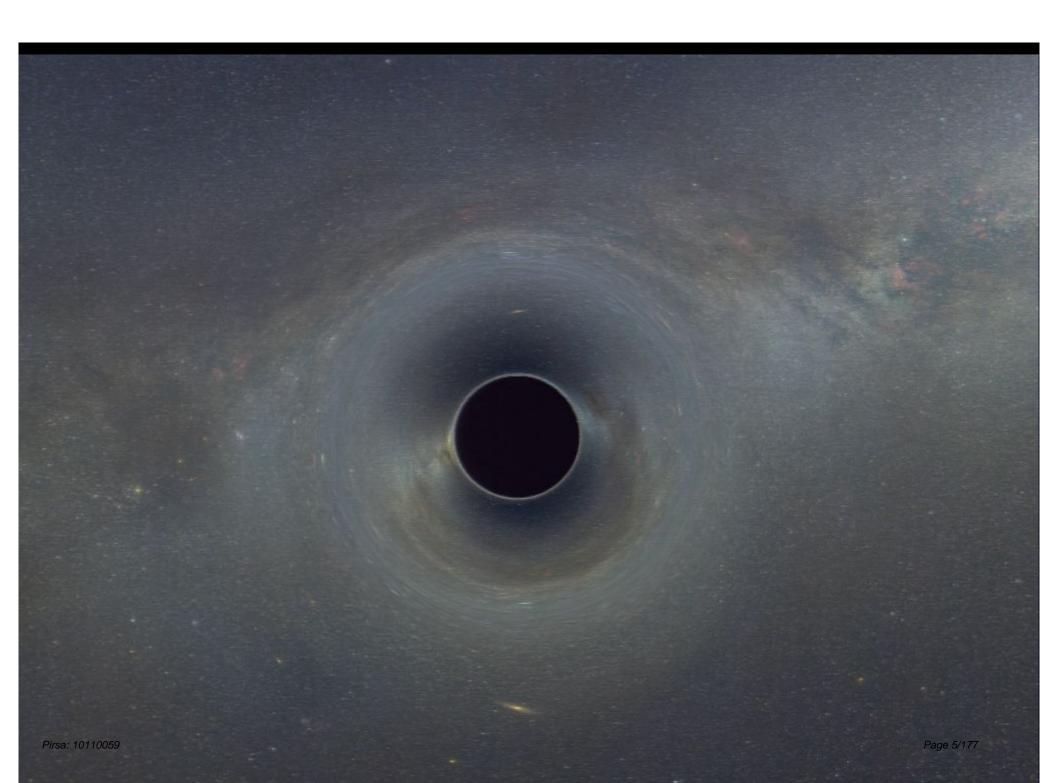
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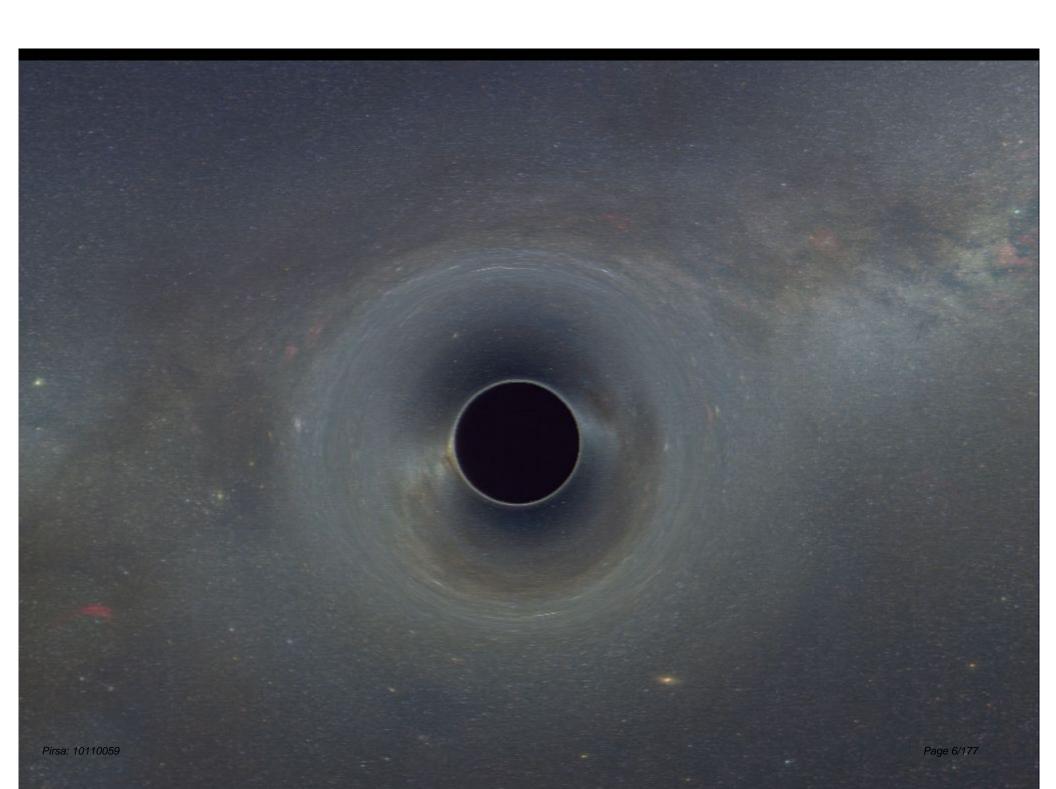


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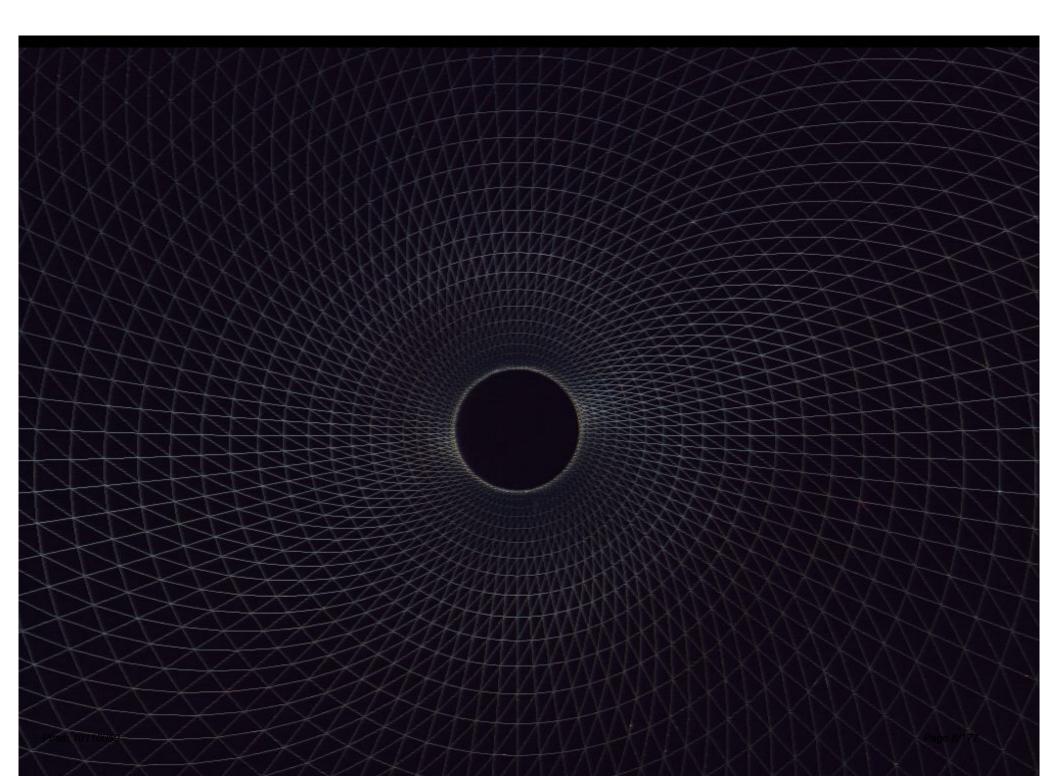
Schwarzschild black hole







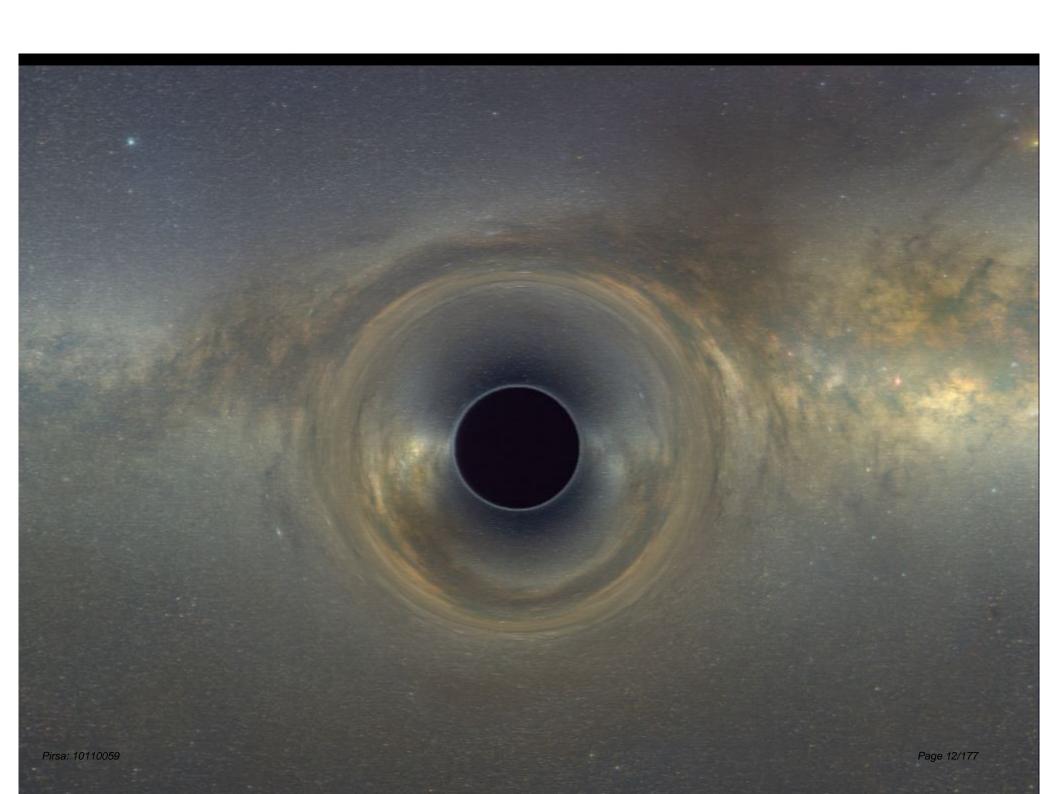


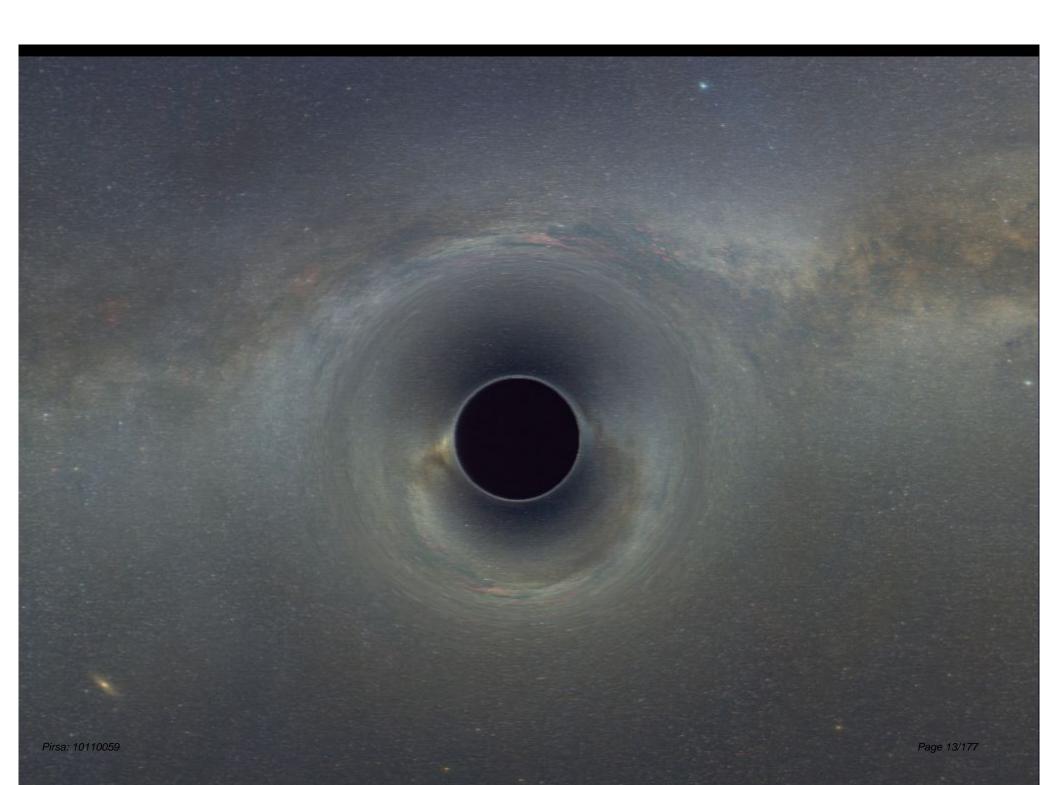


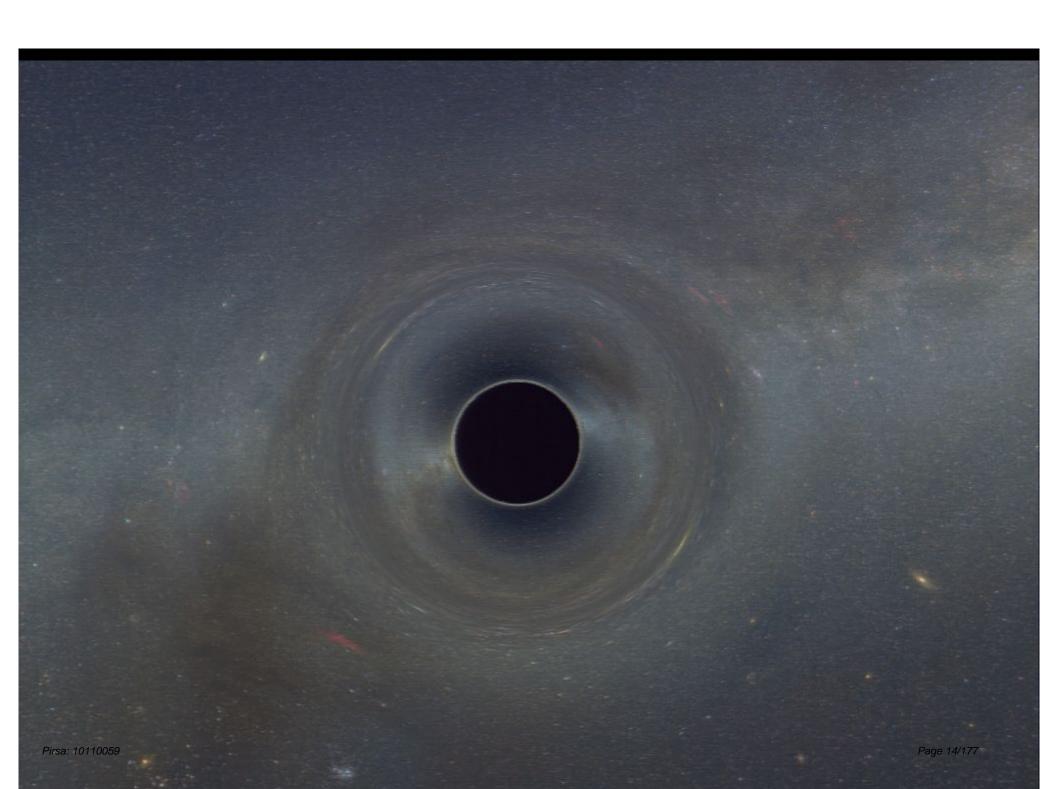


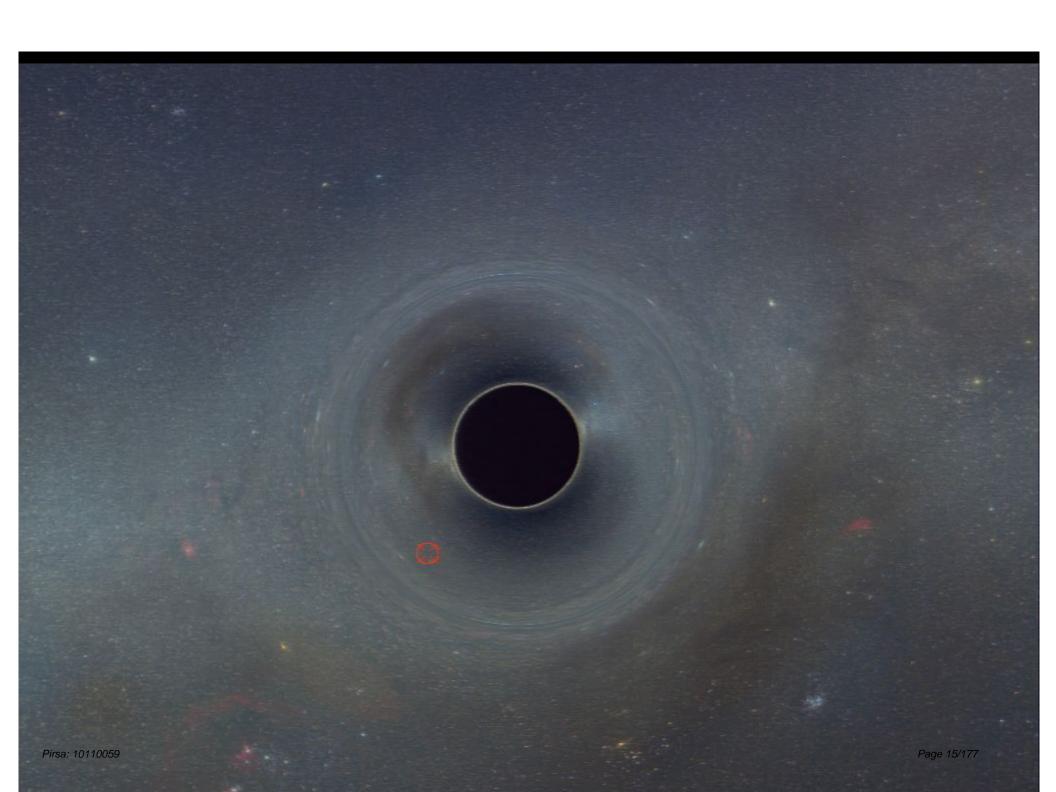


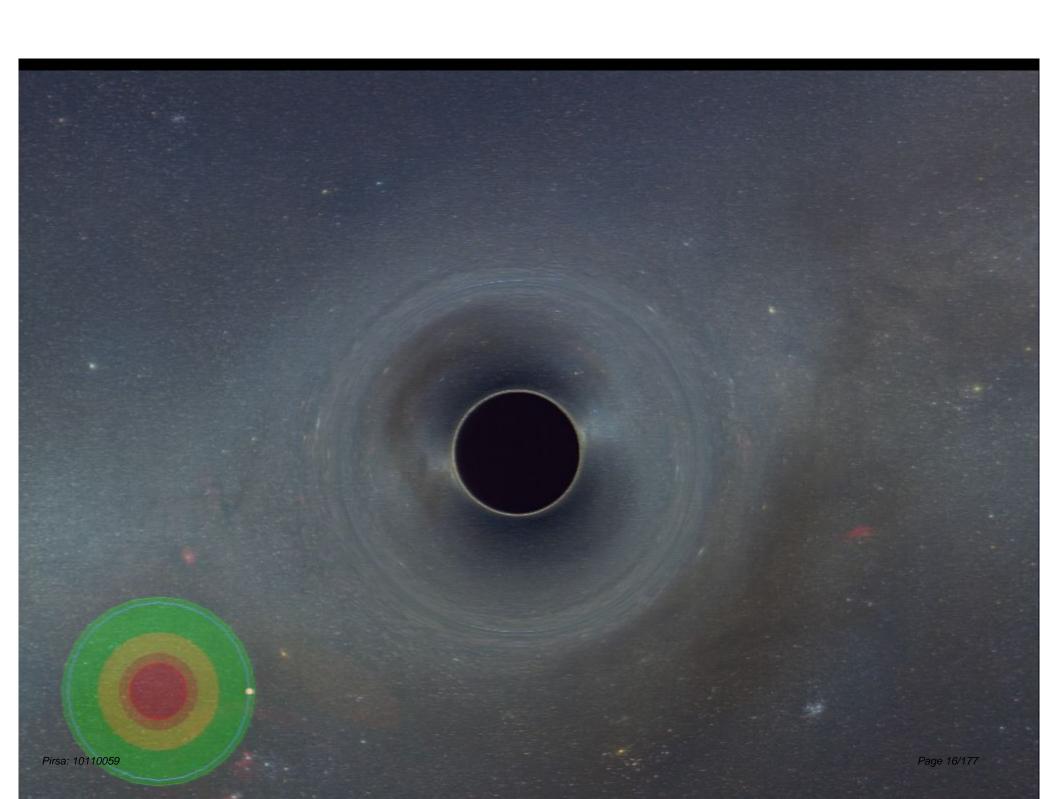


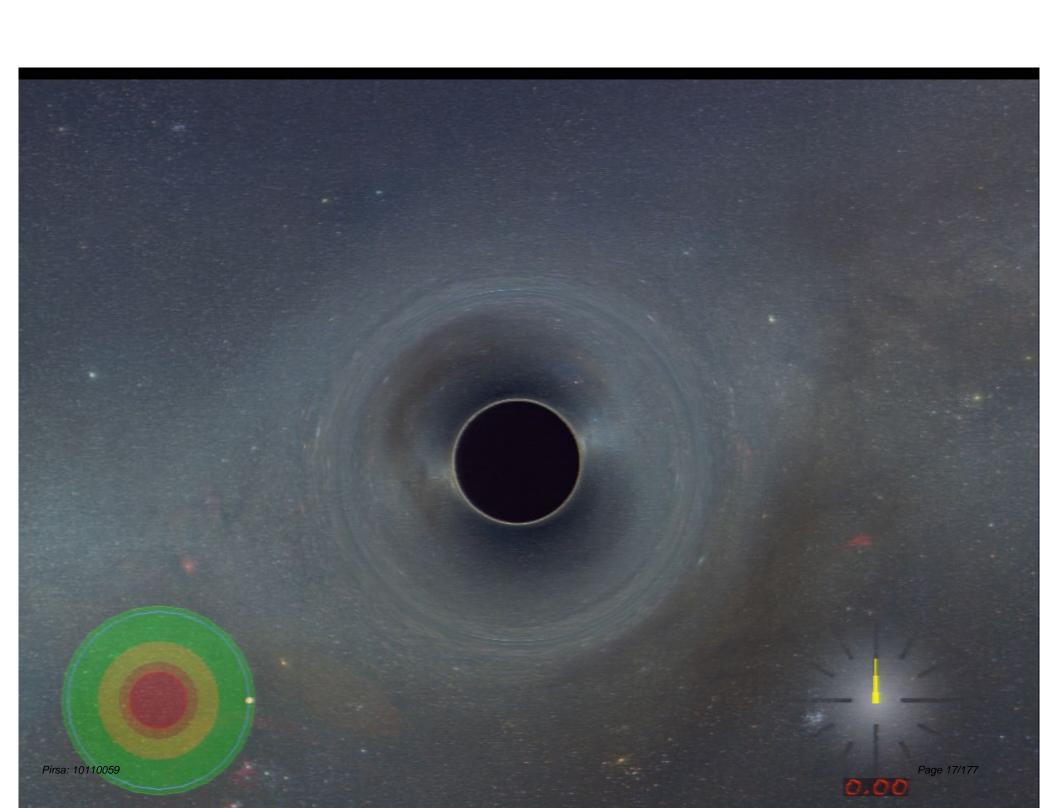


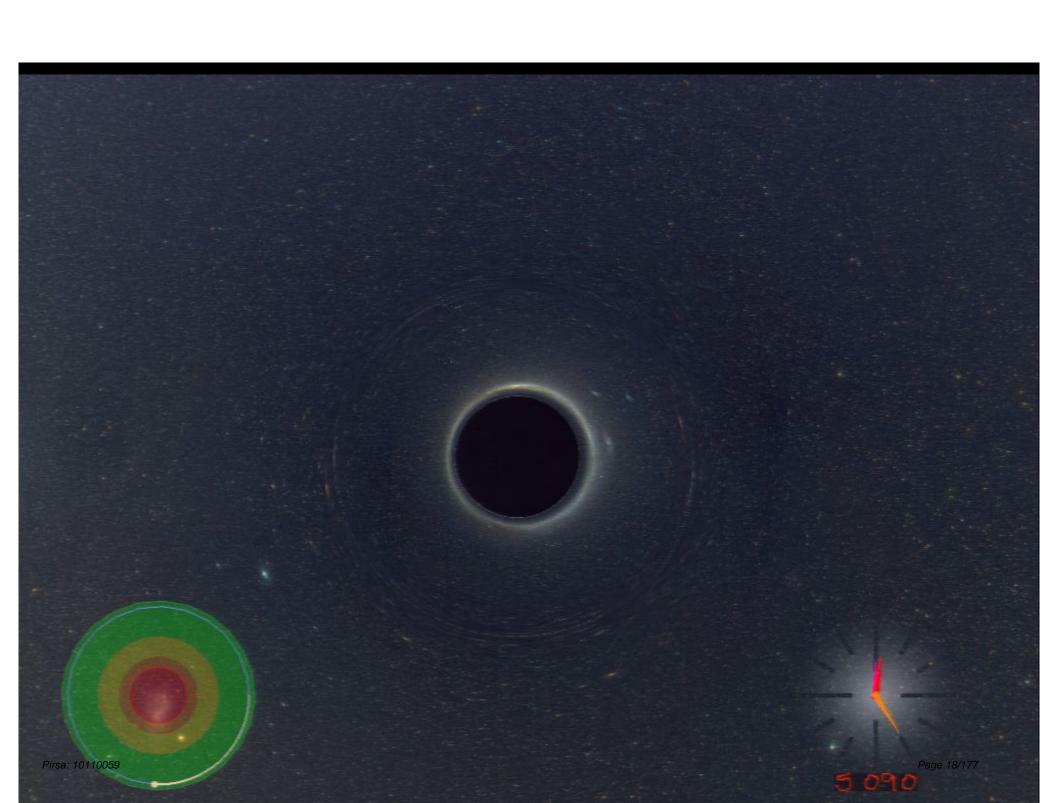


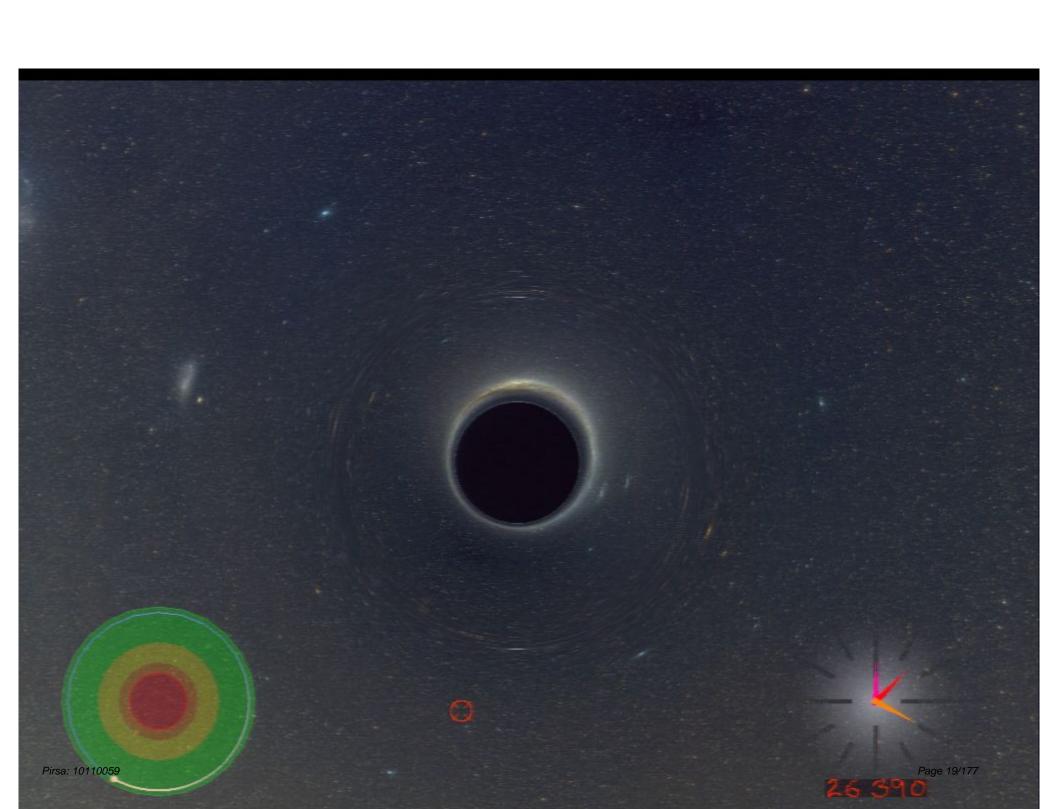


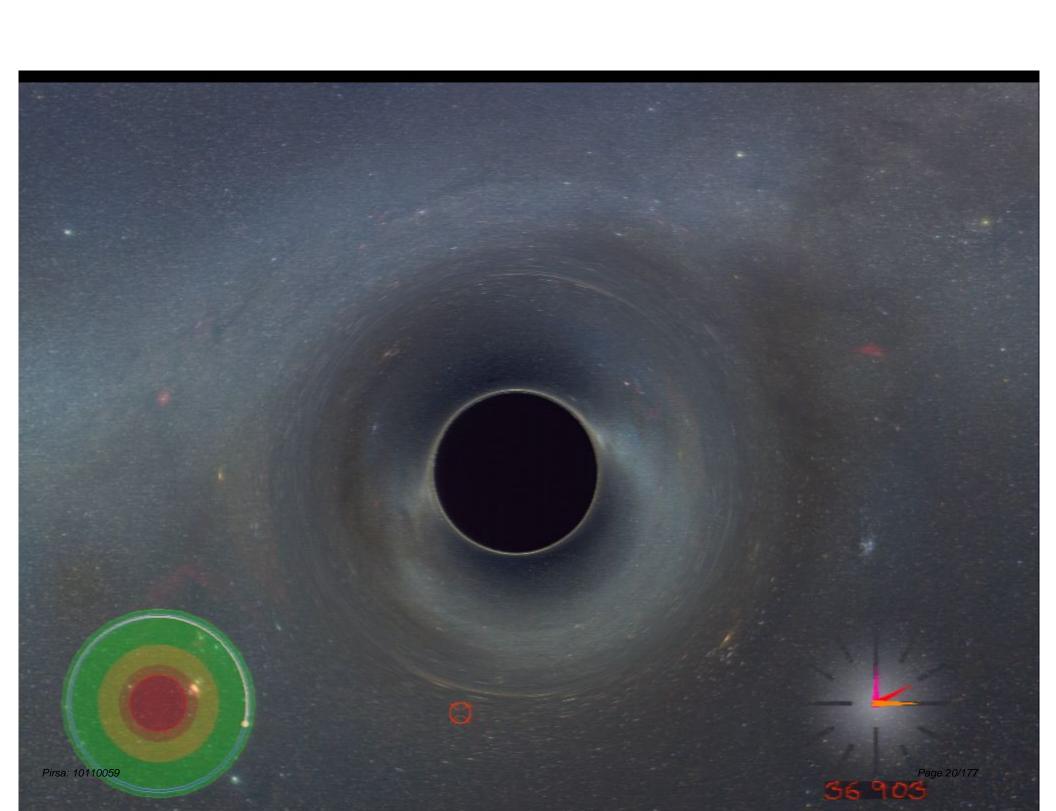


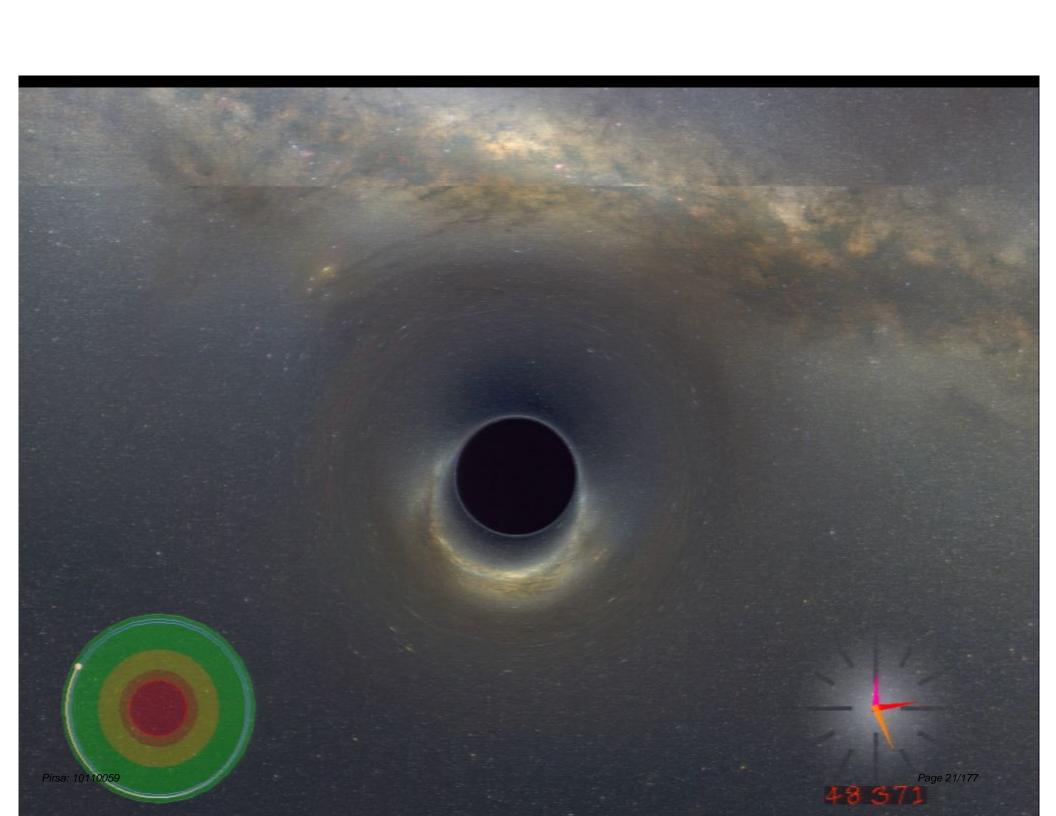


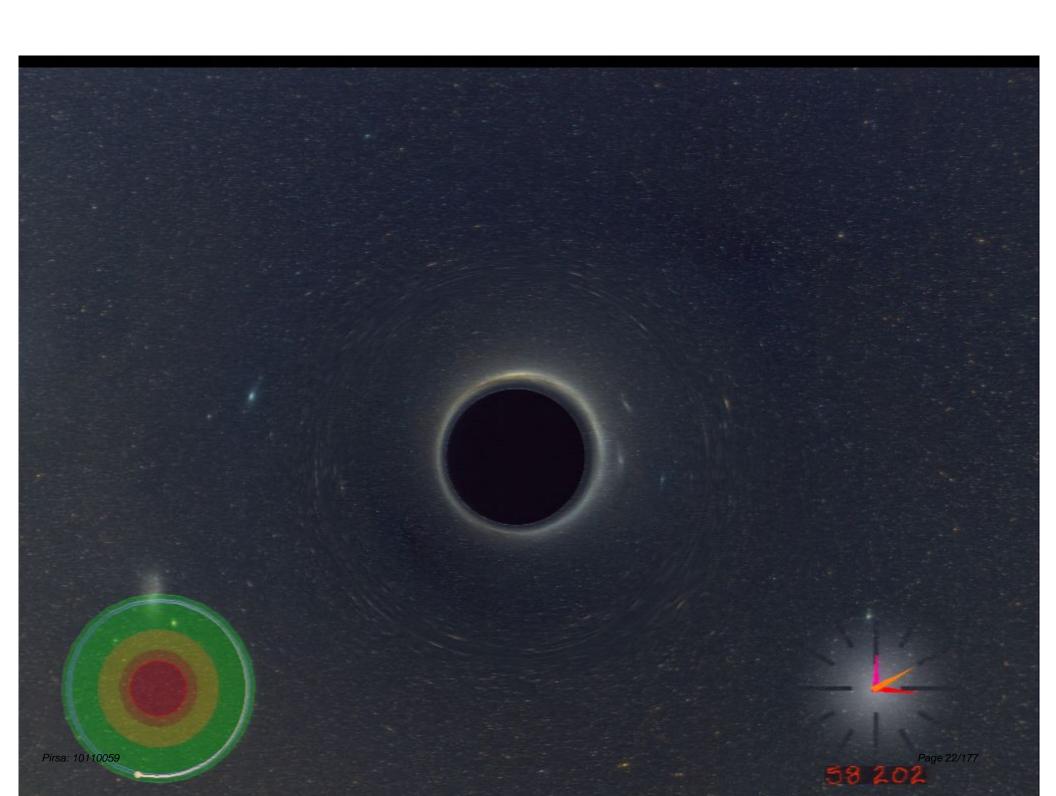


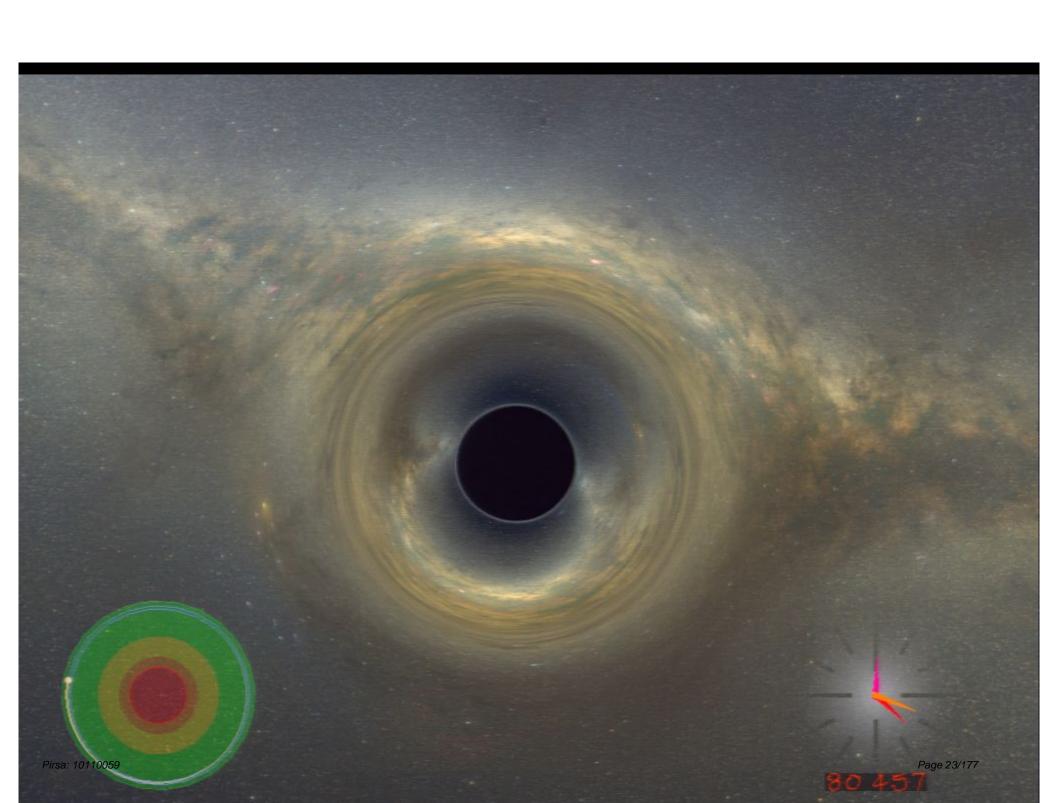


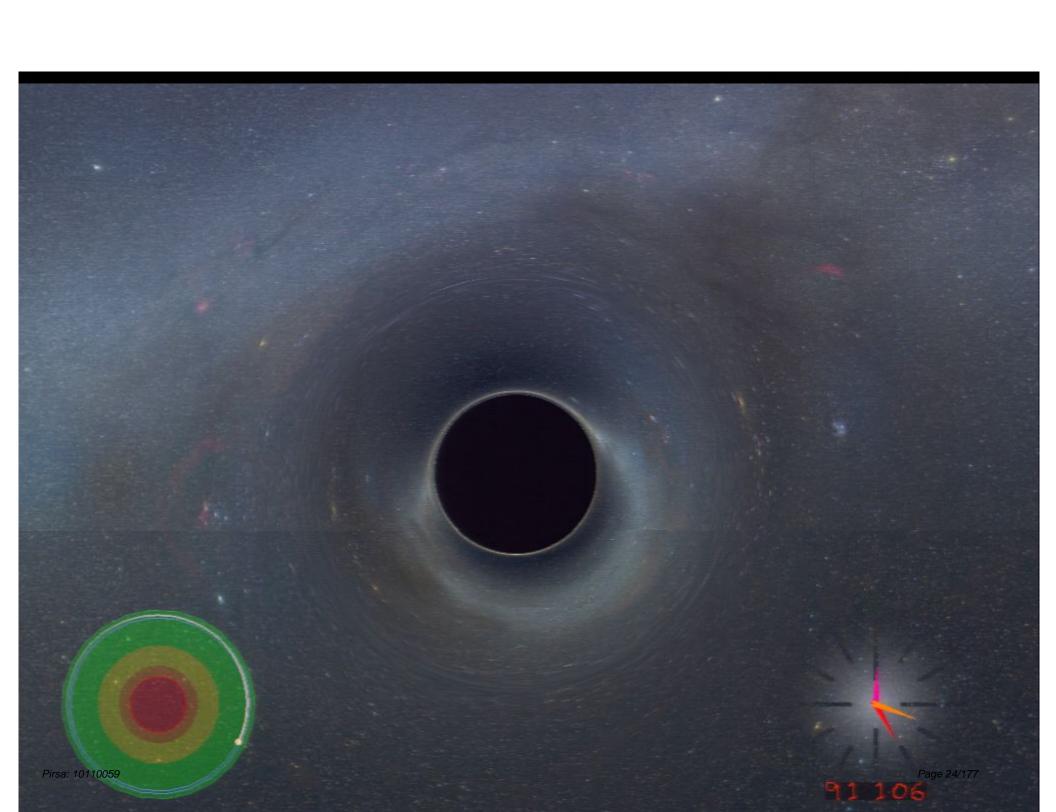


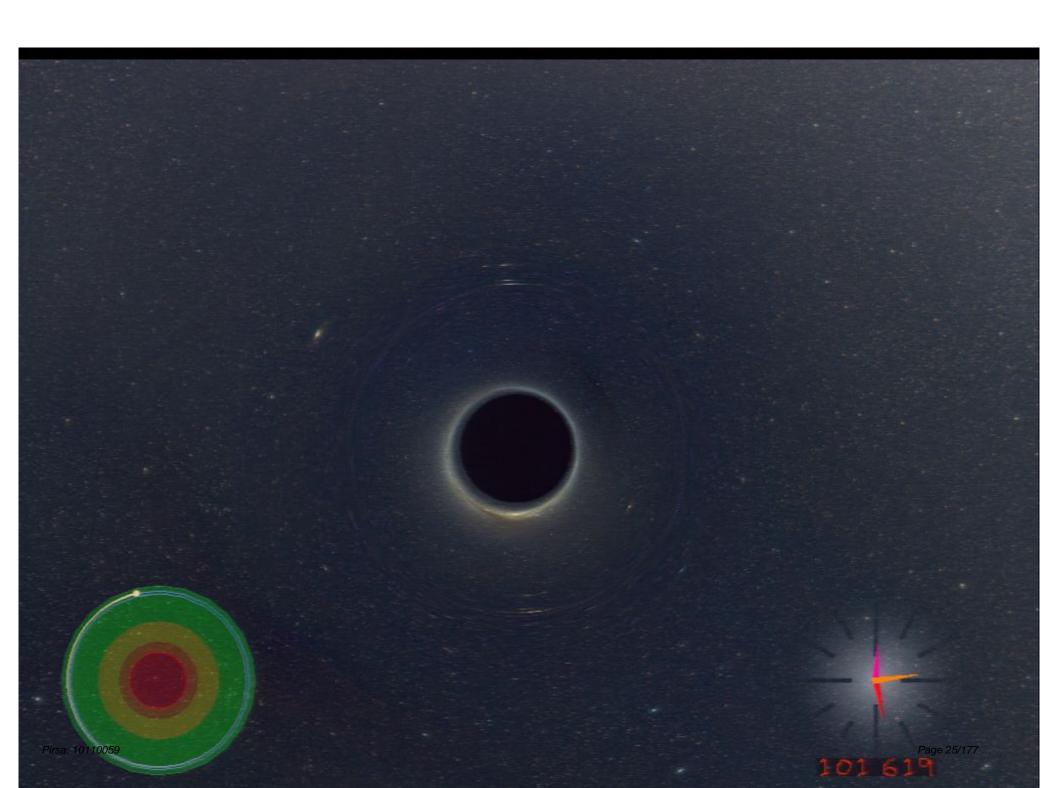


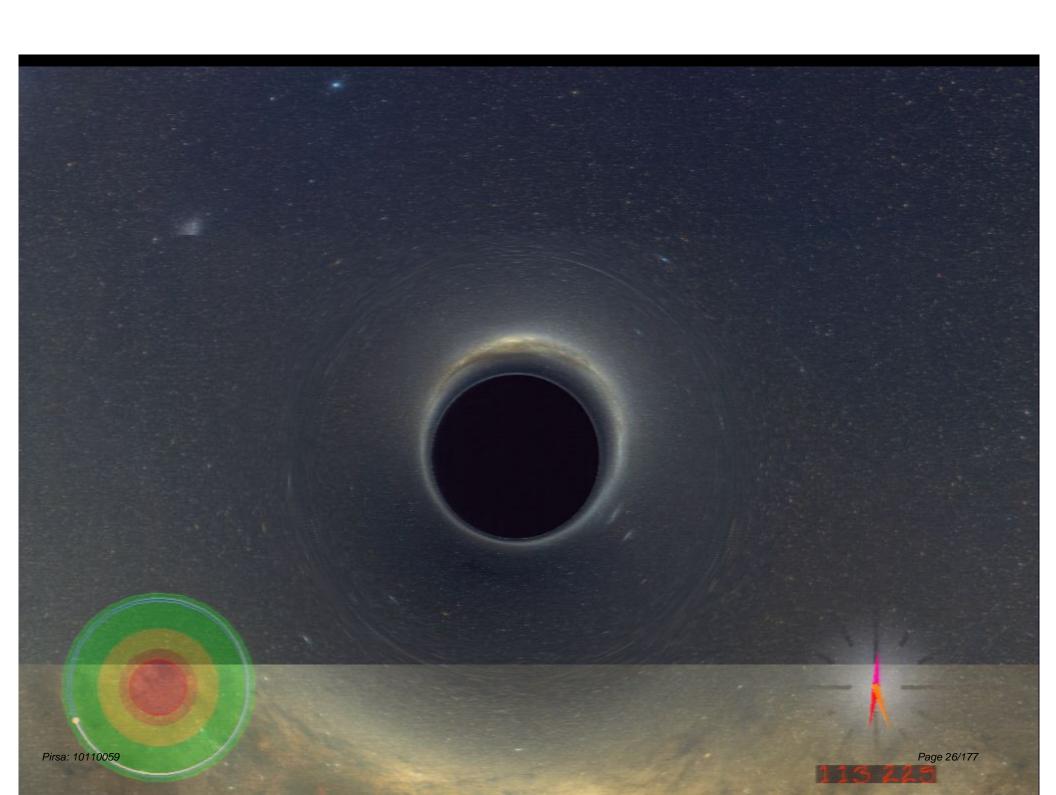


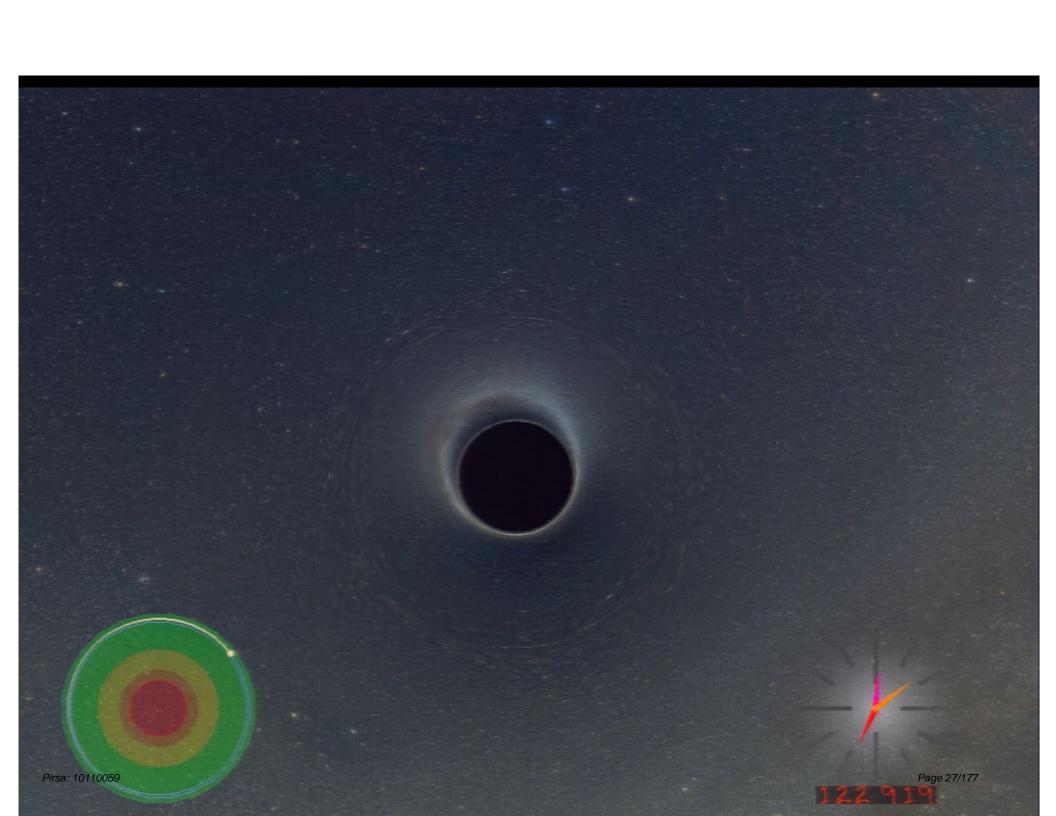


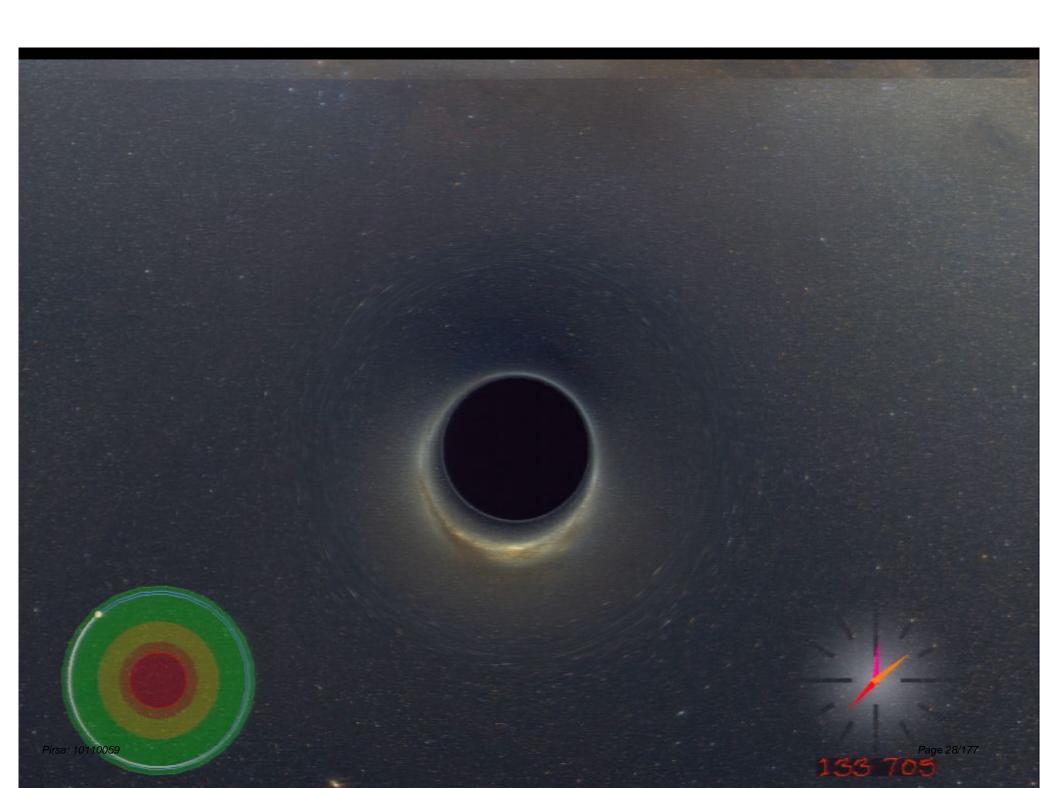


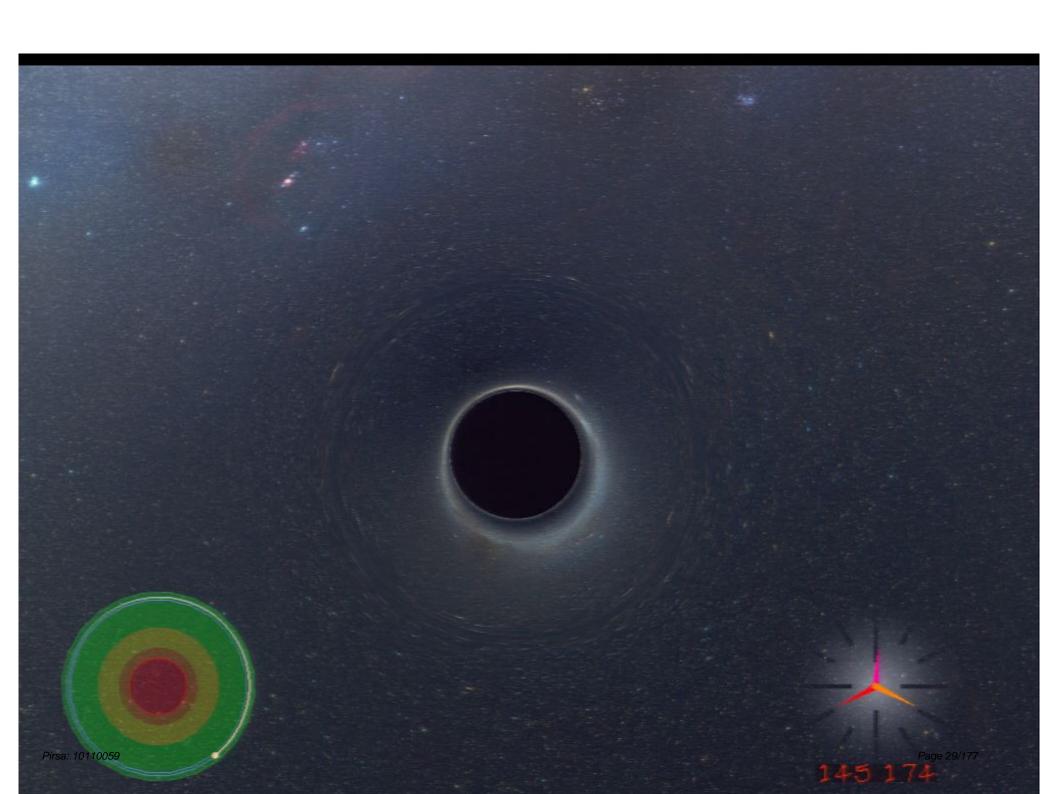


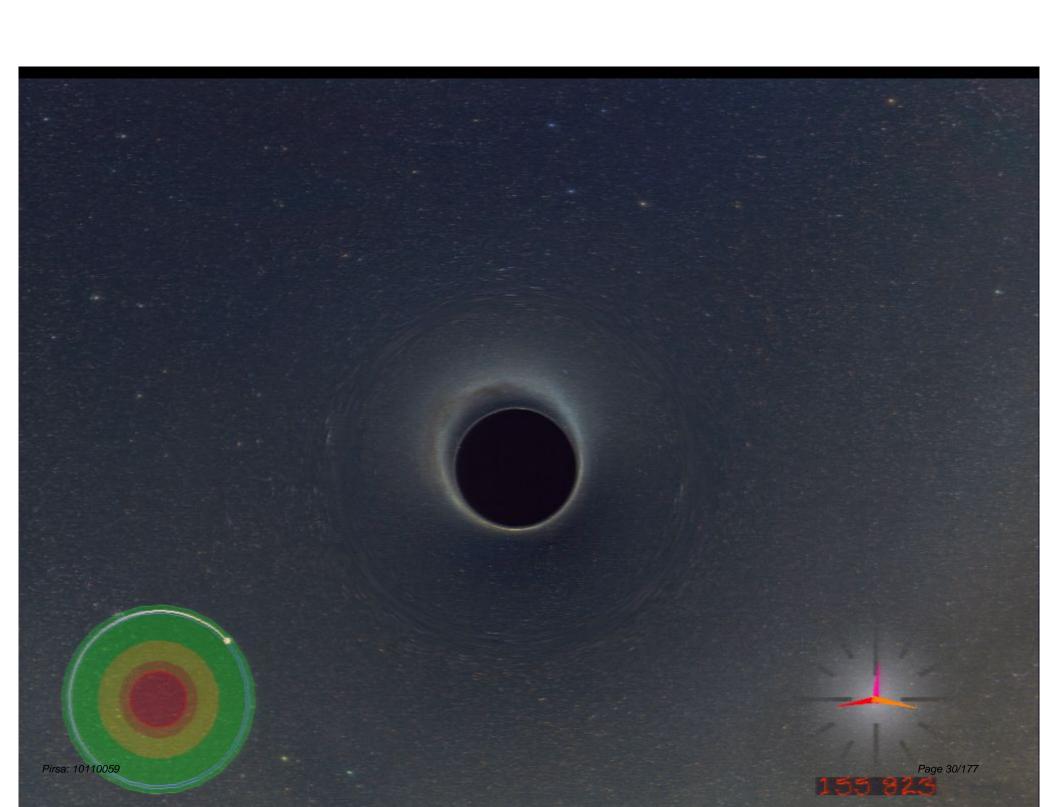


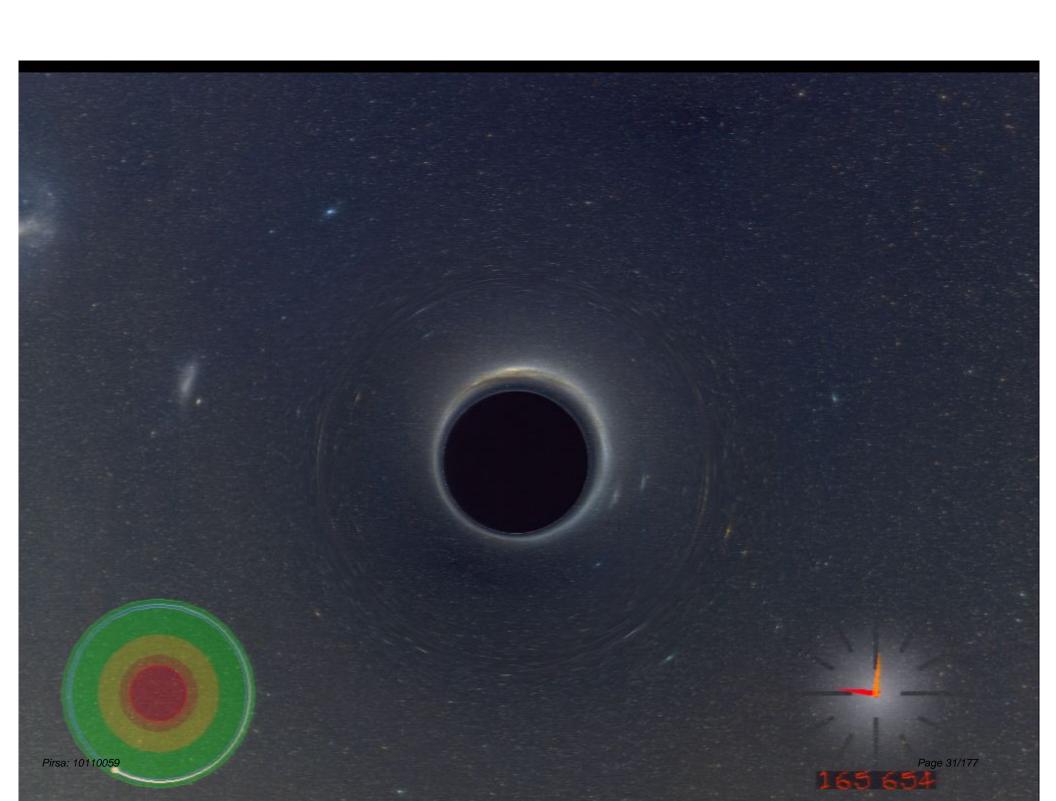


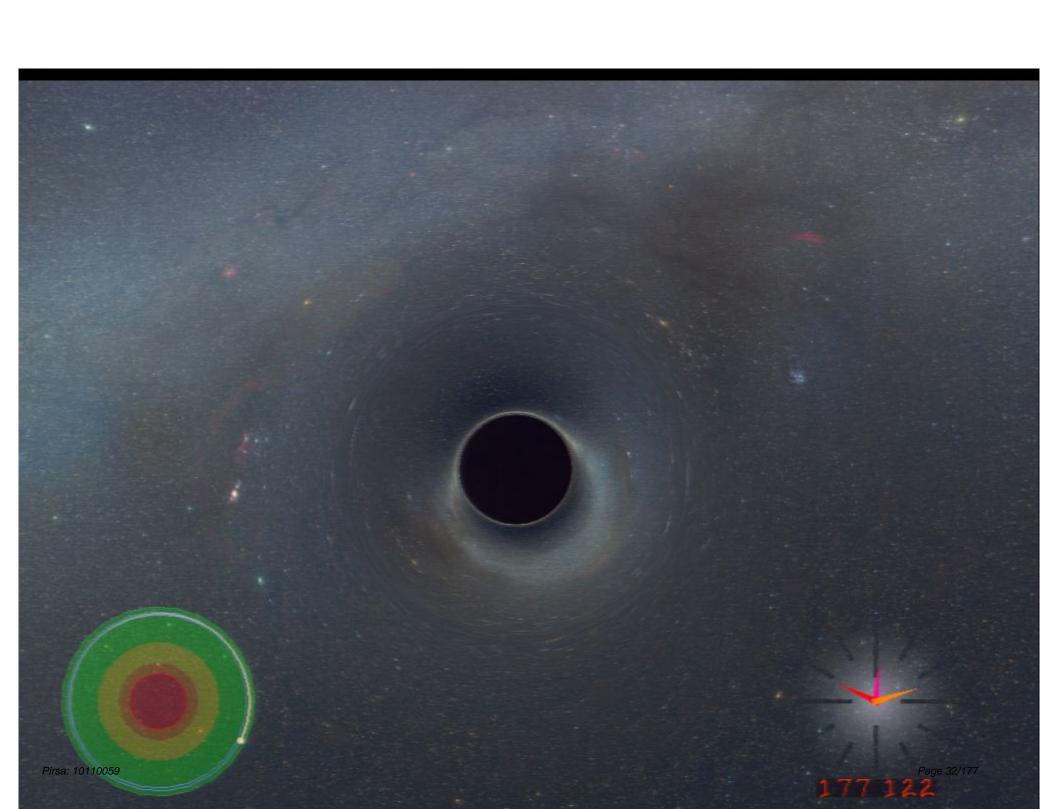


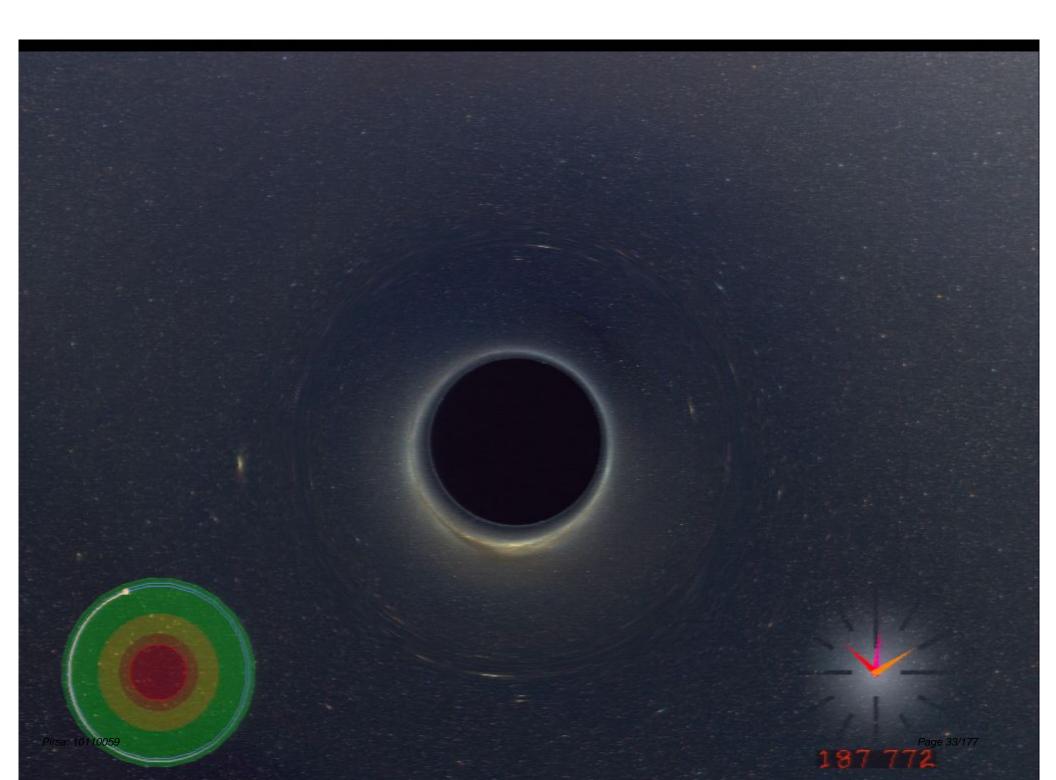


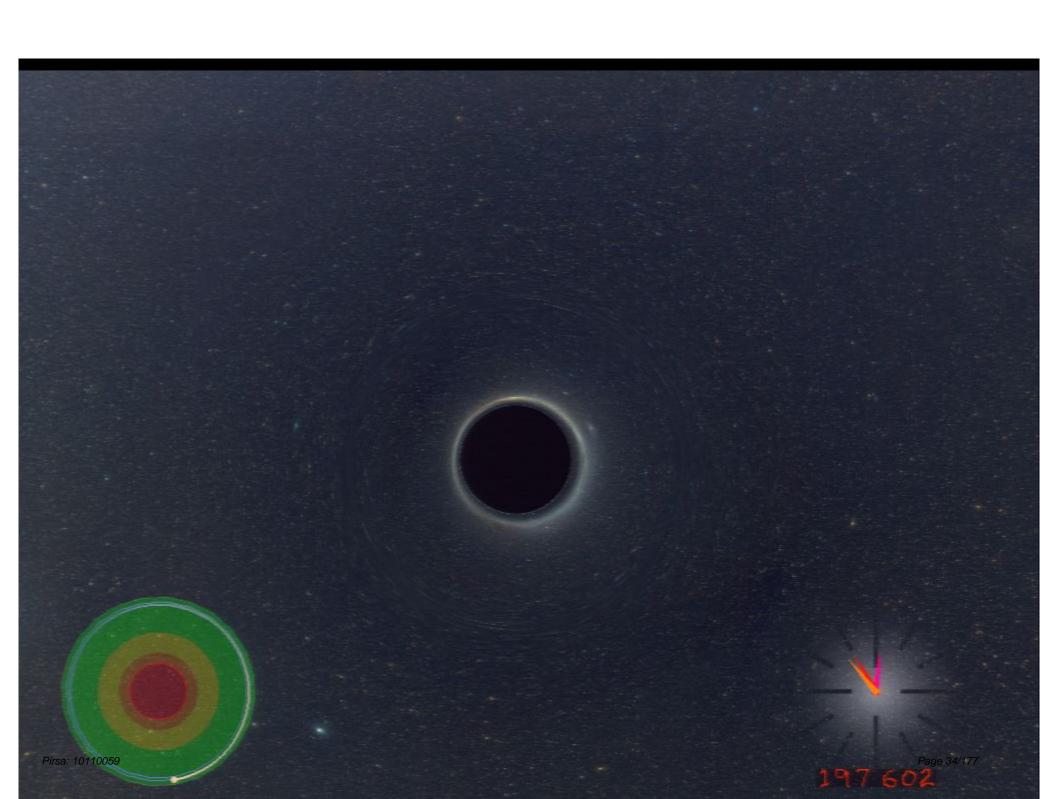


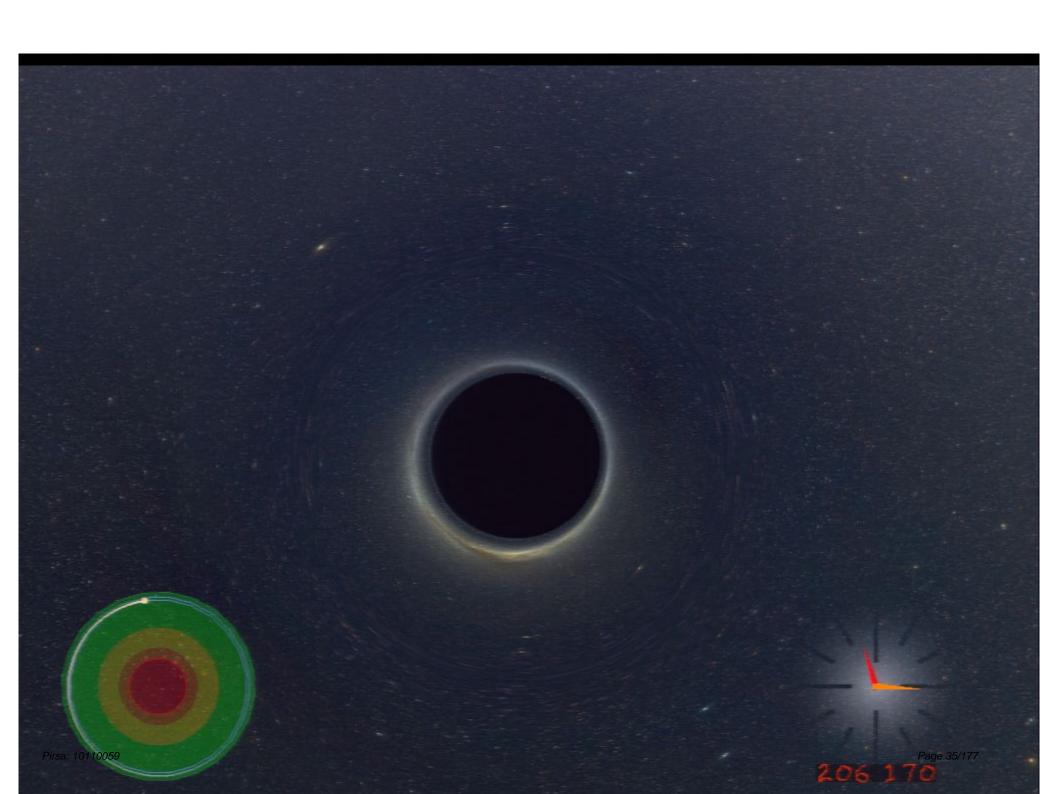


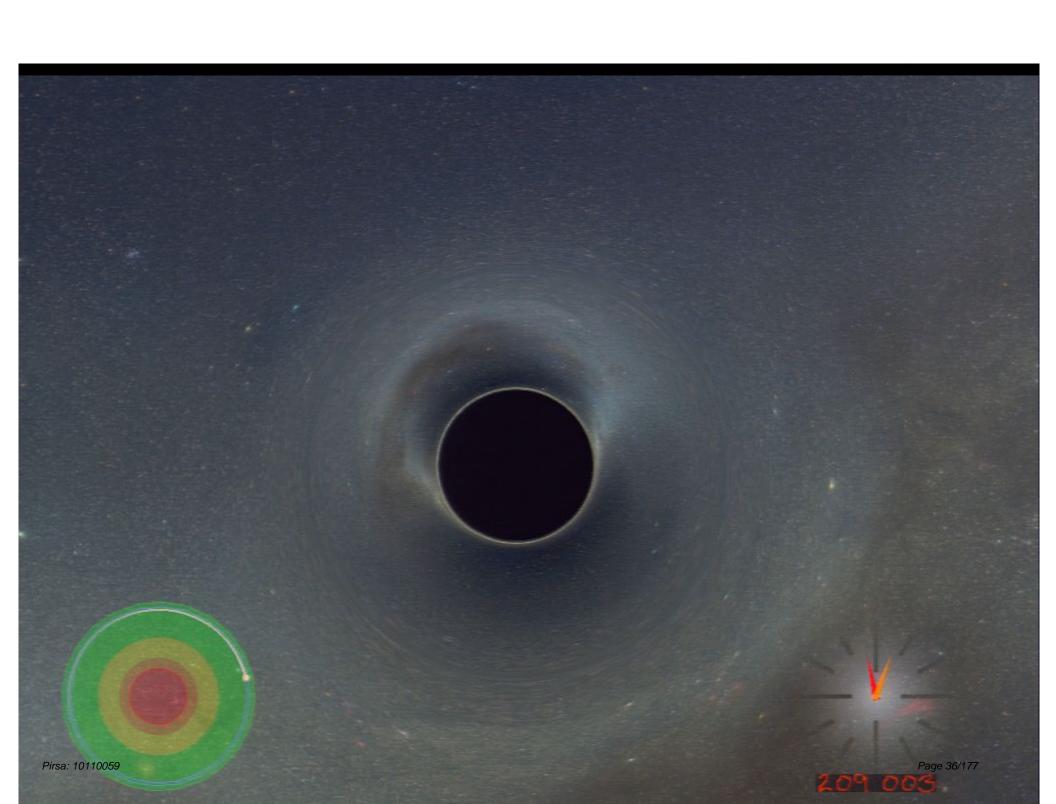


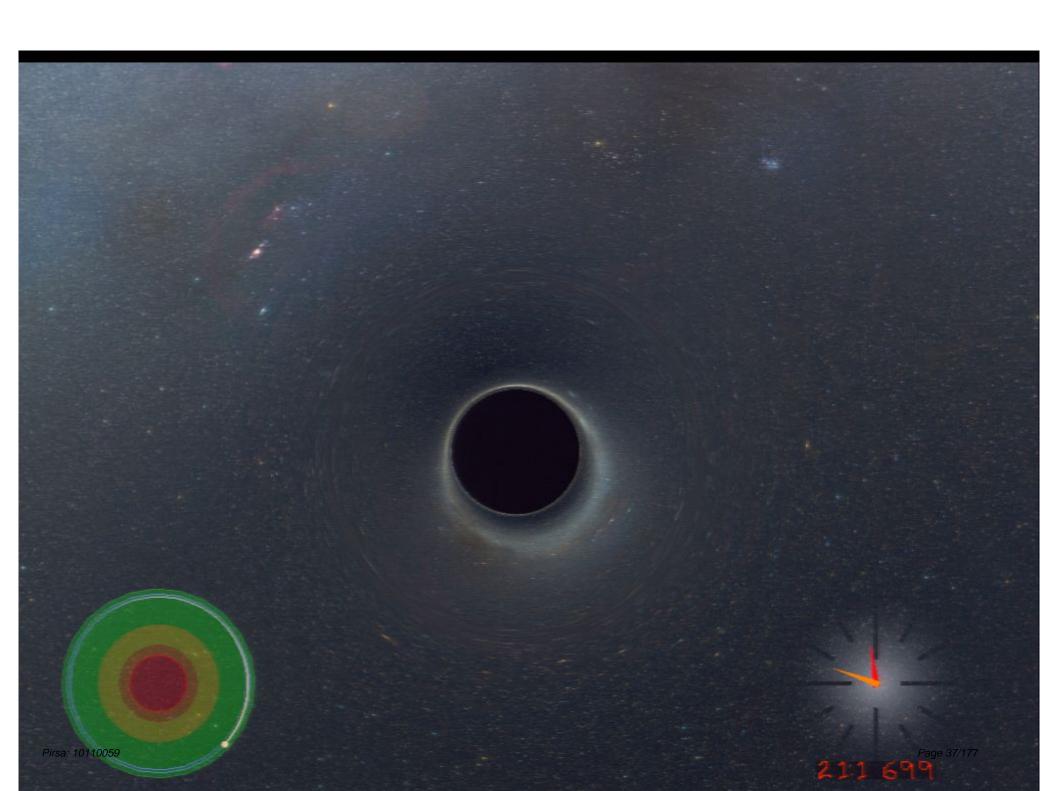


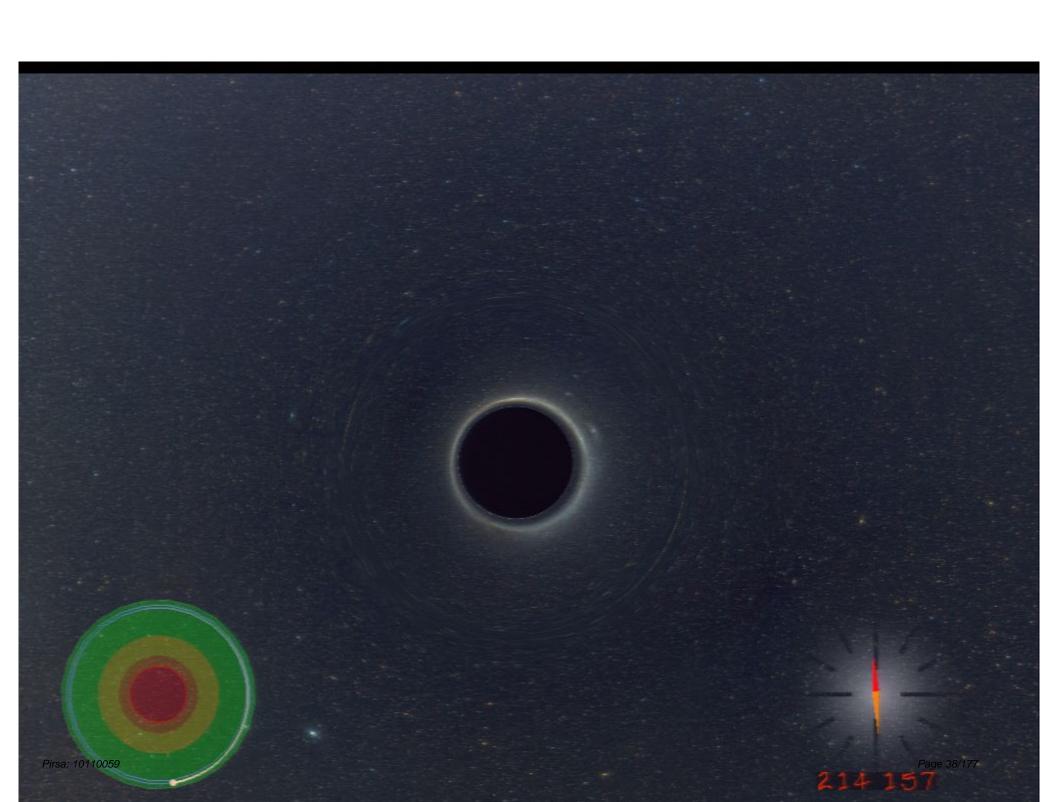


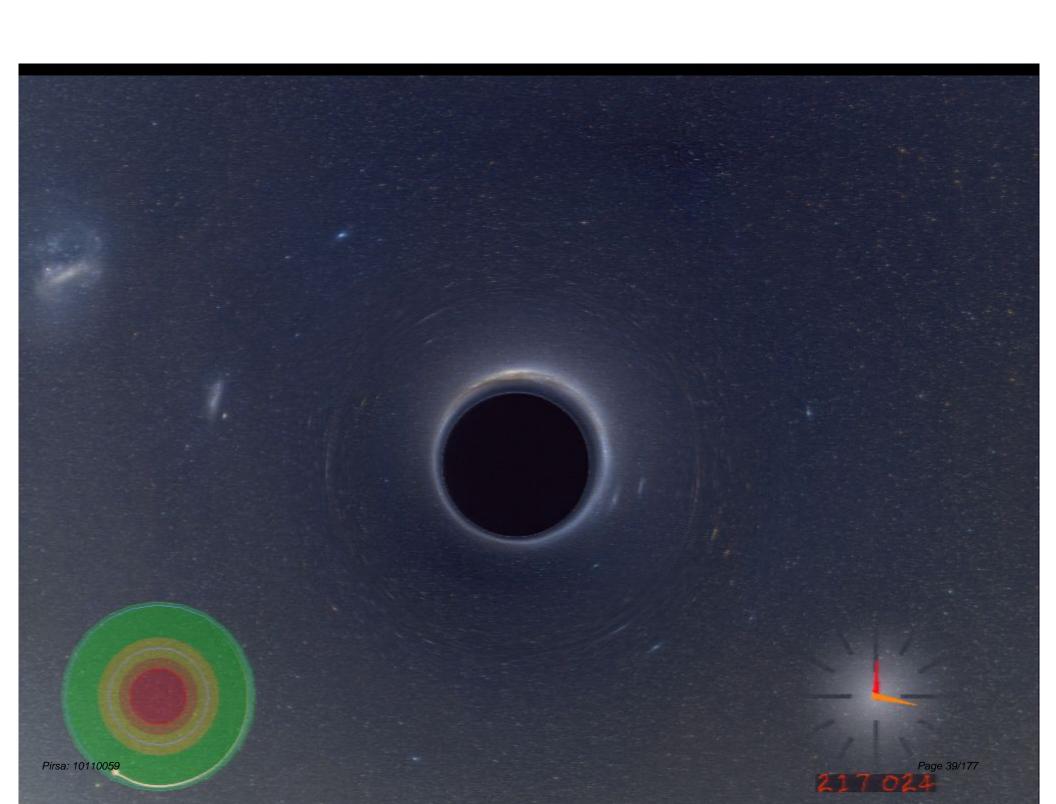


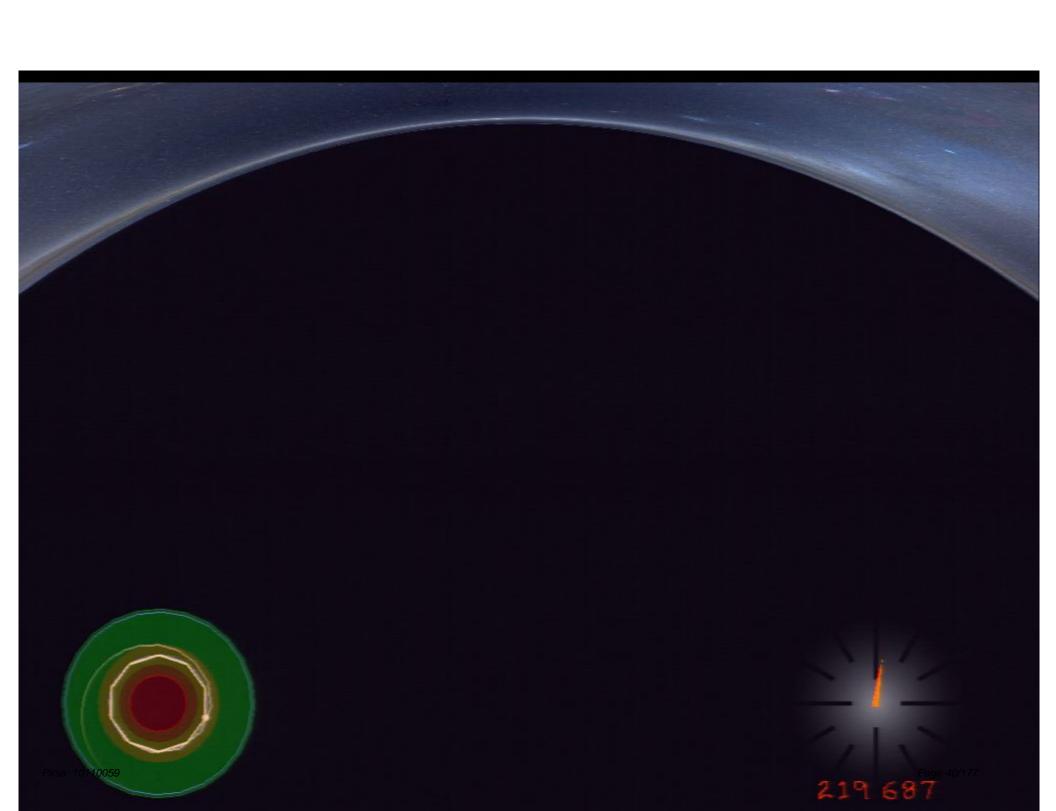


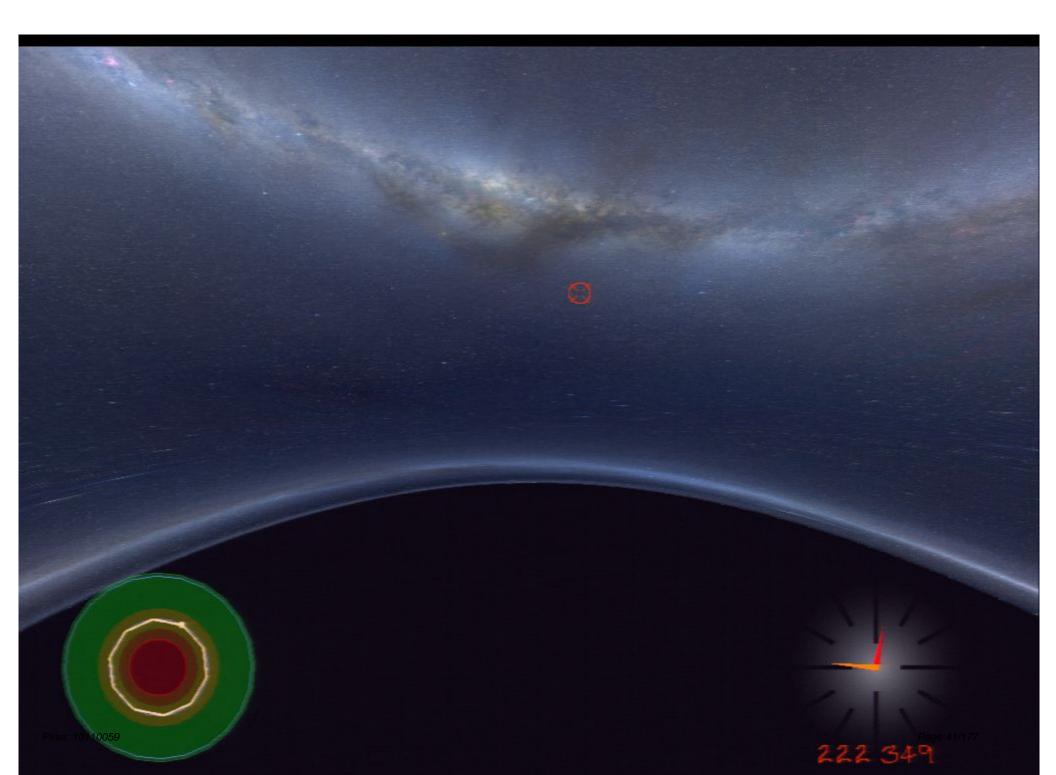


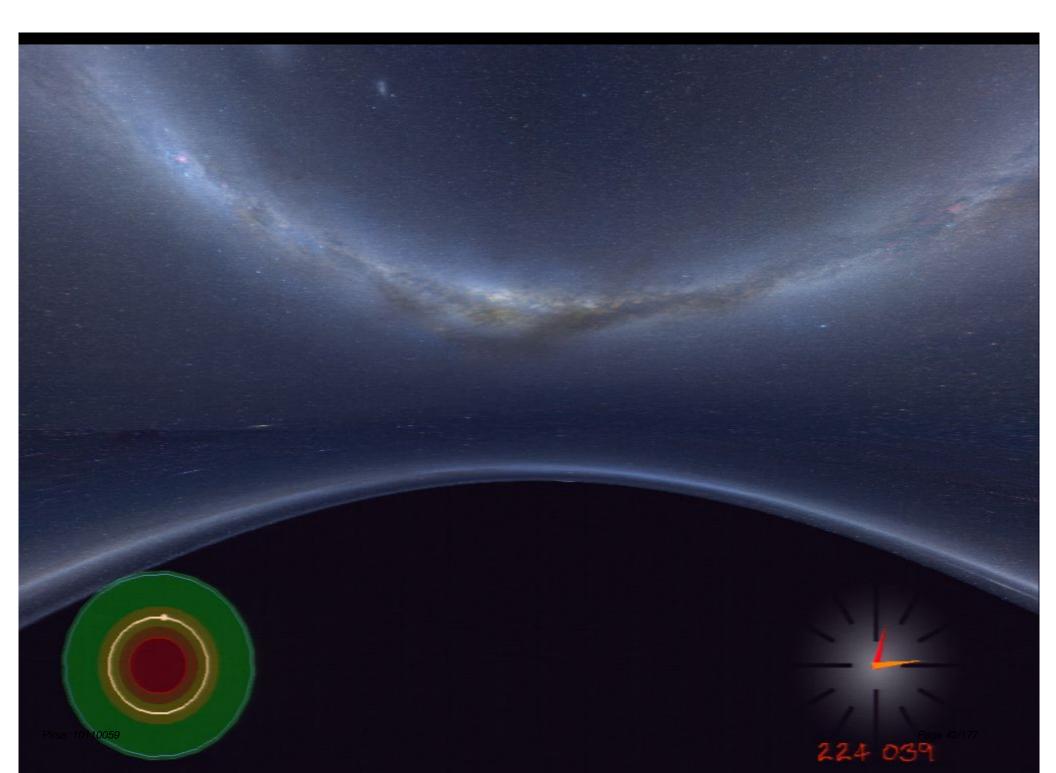


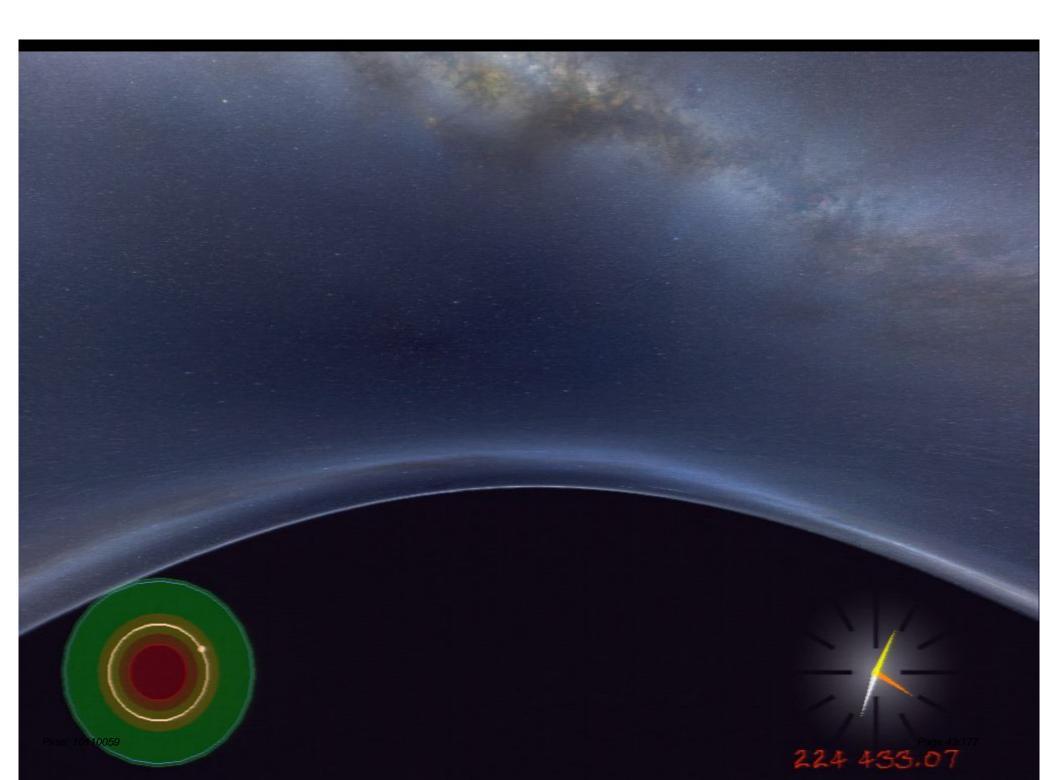


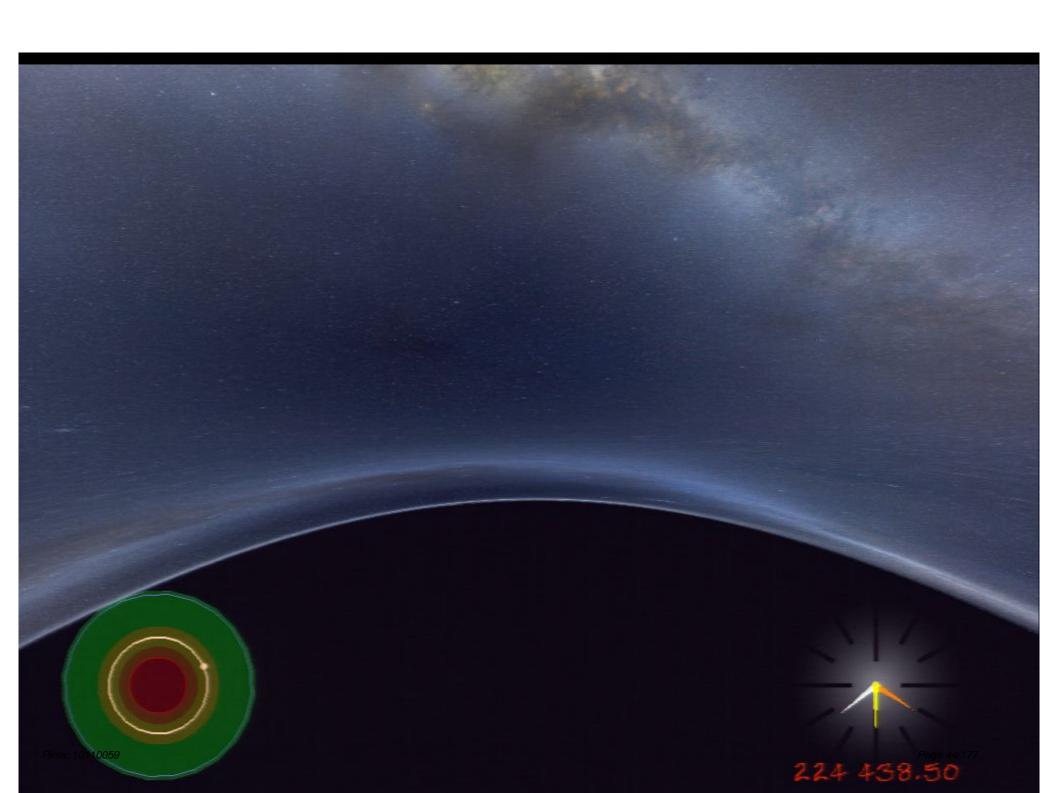


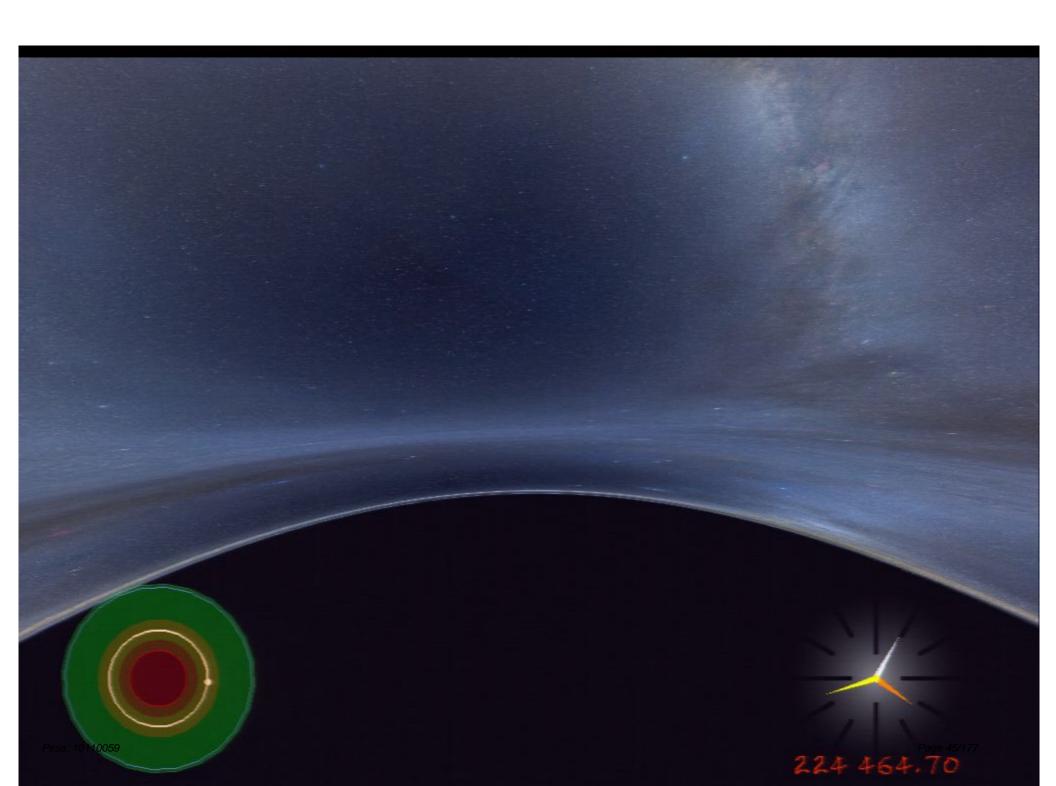


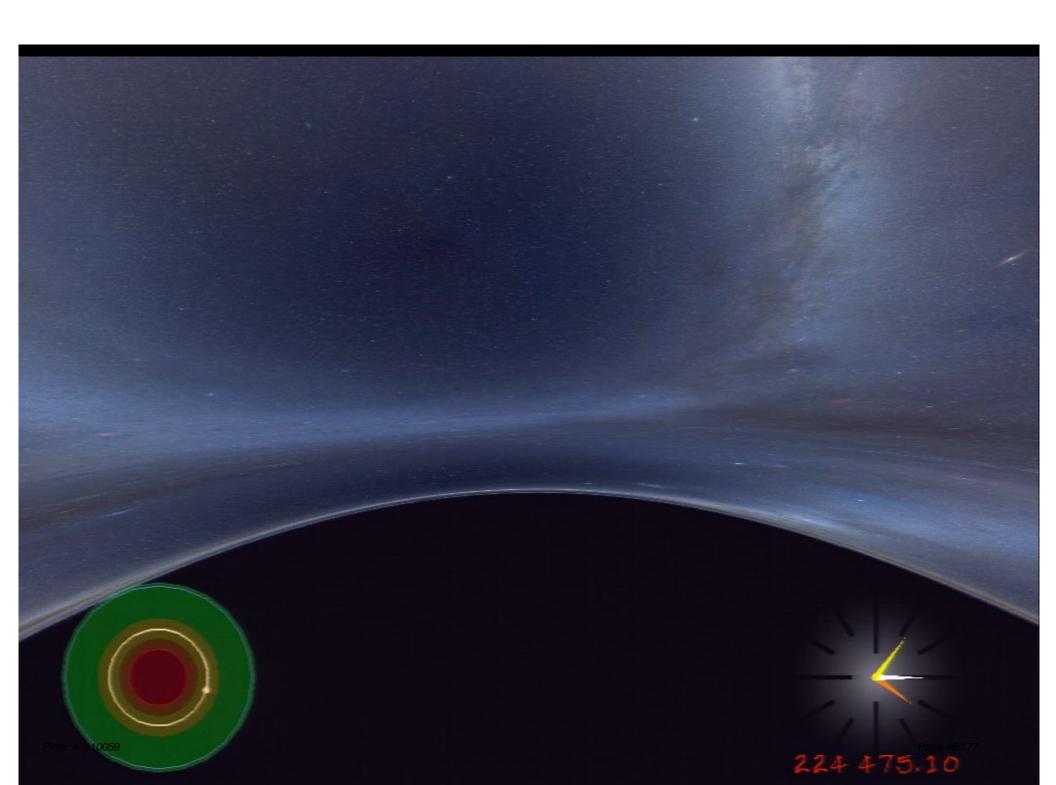


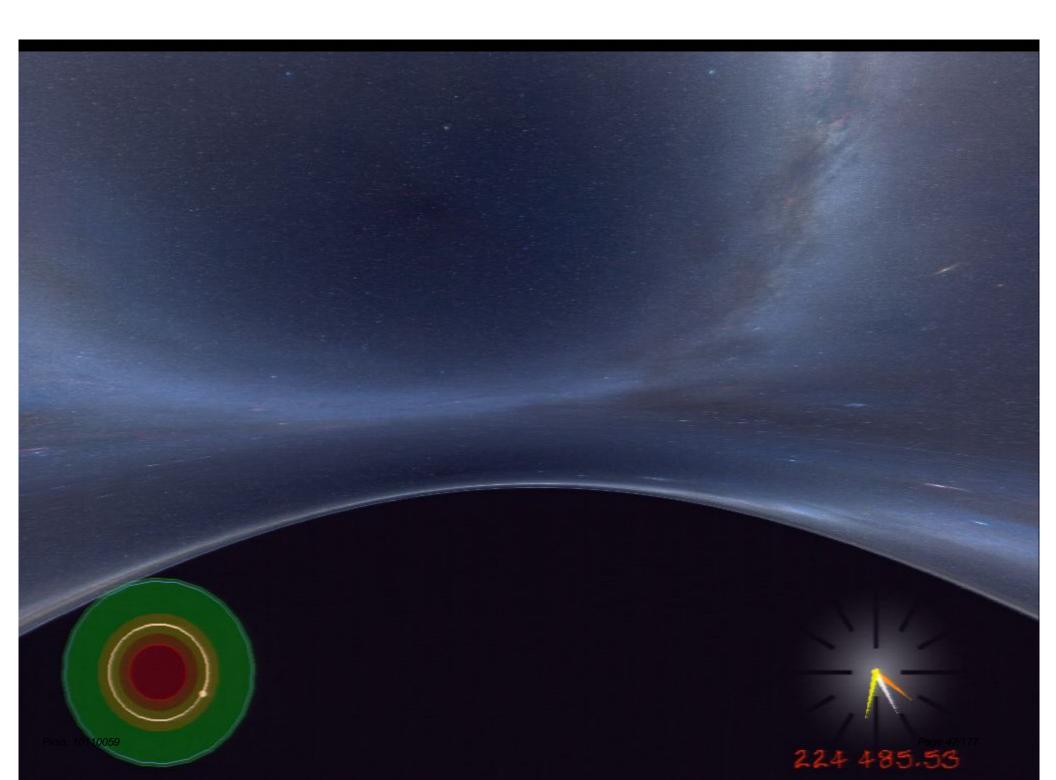


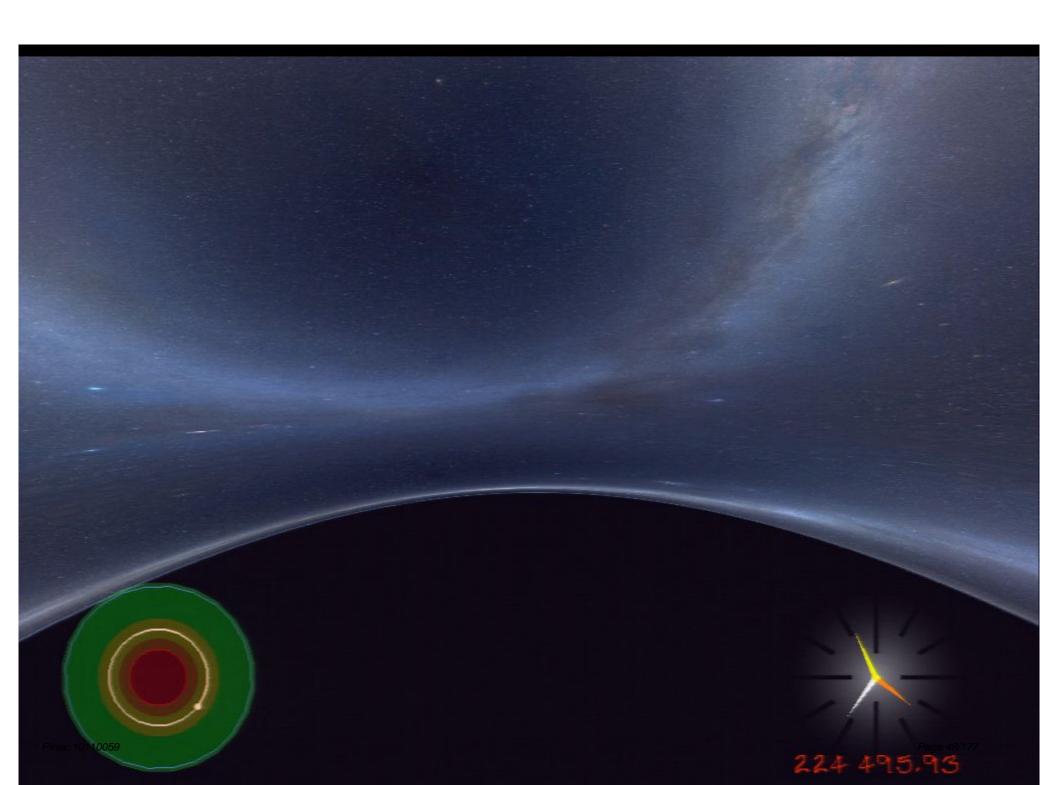


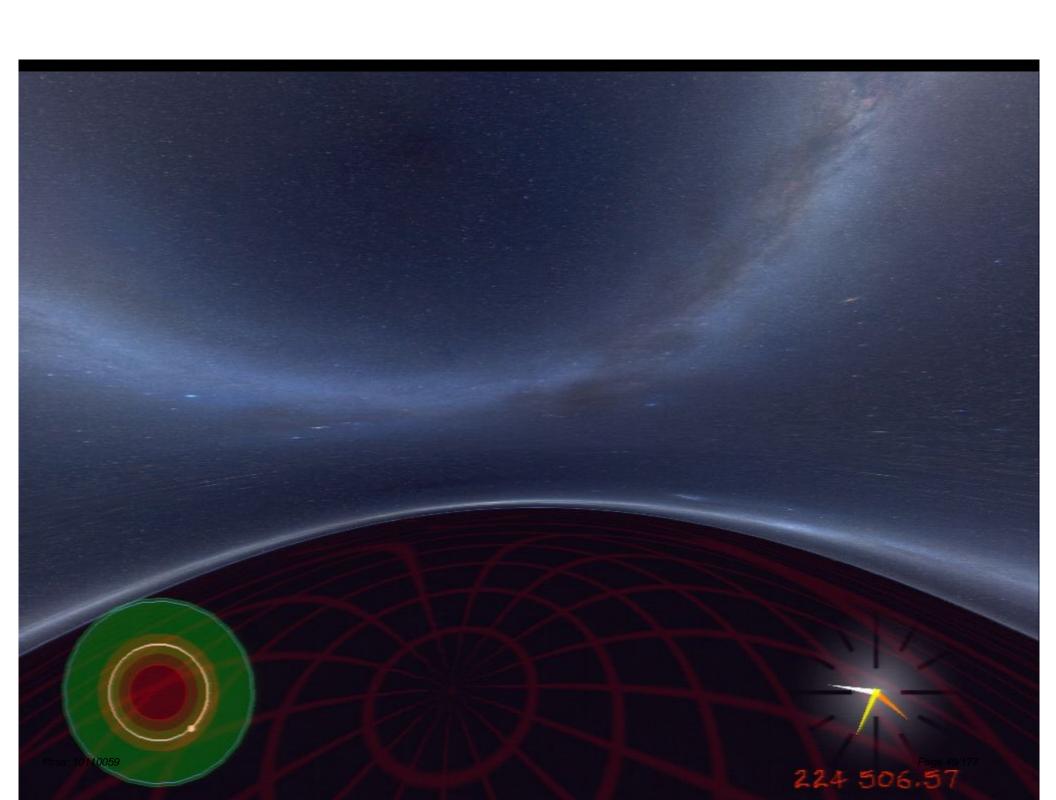


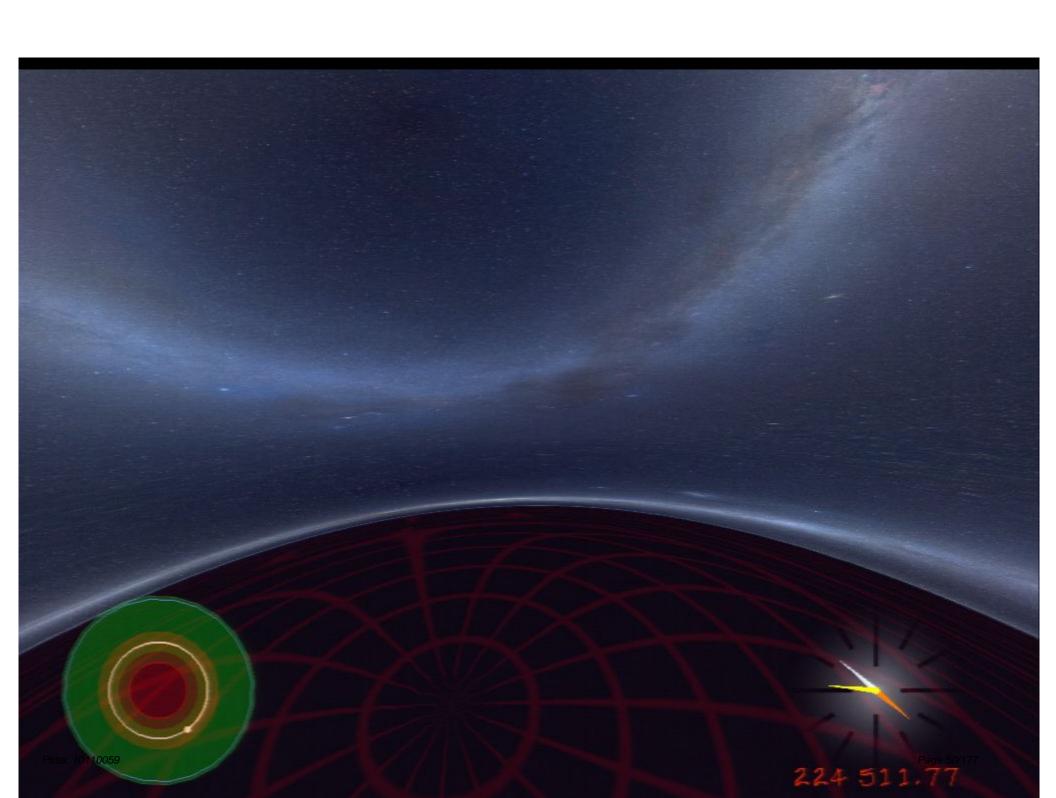


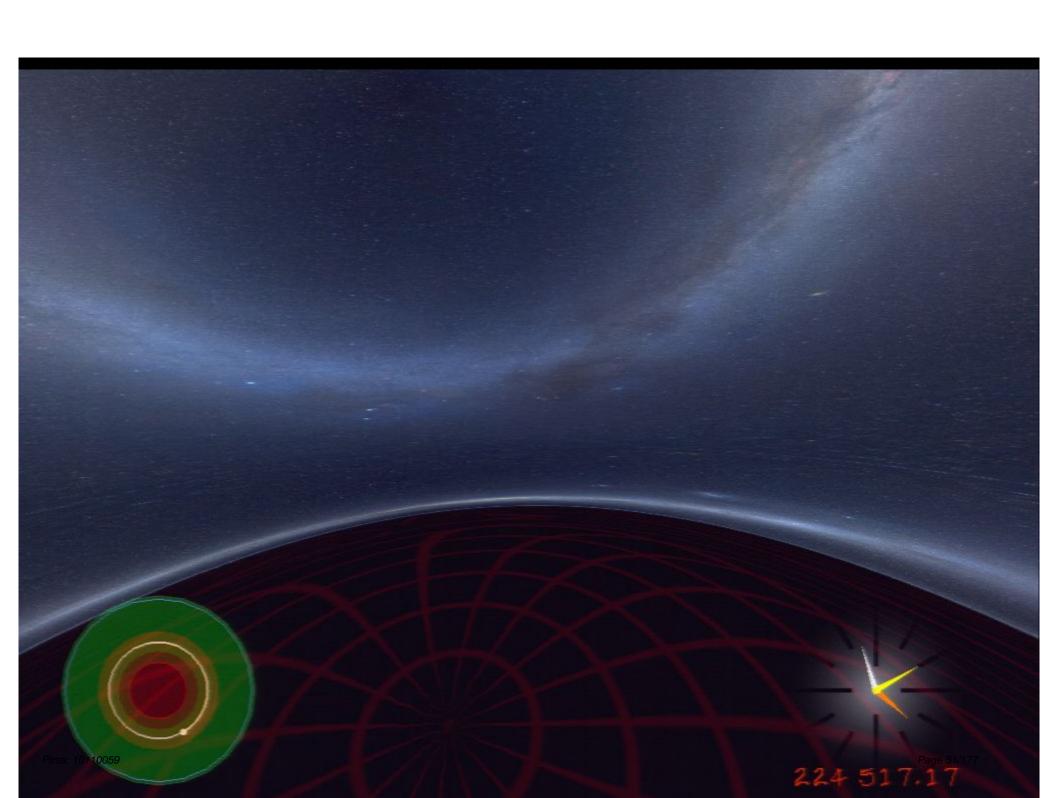


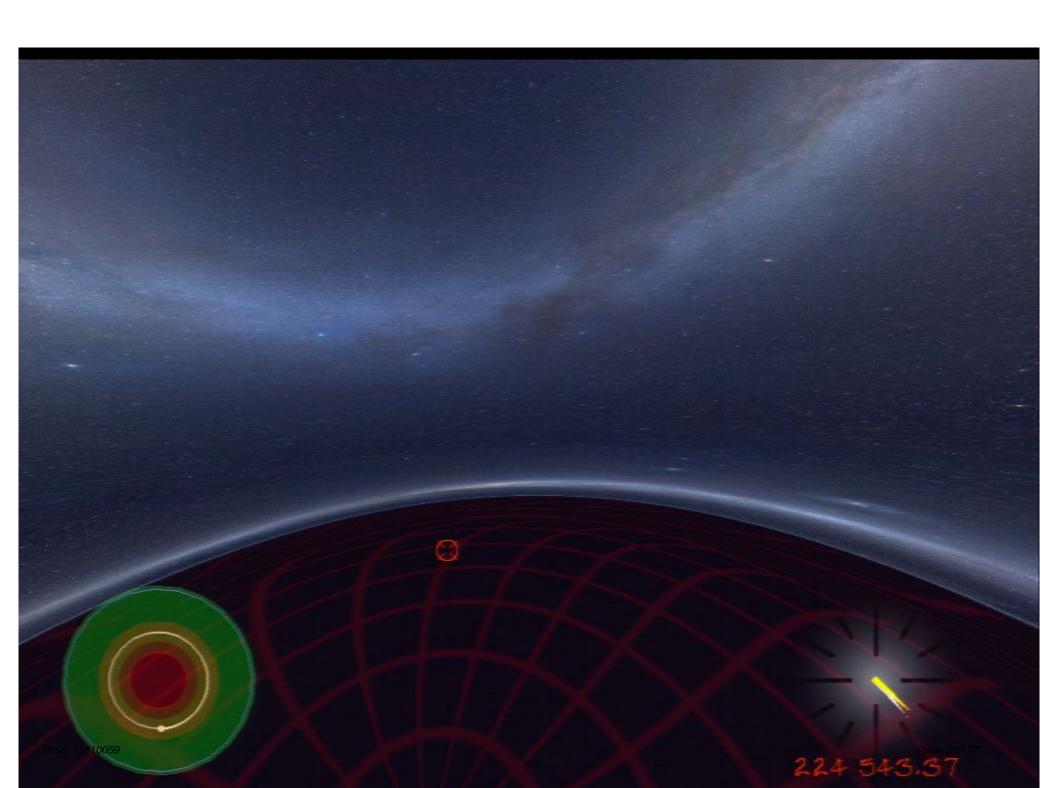


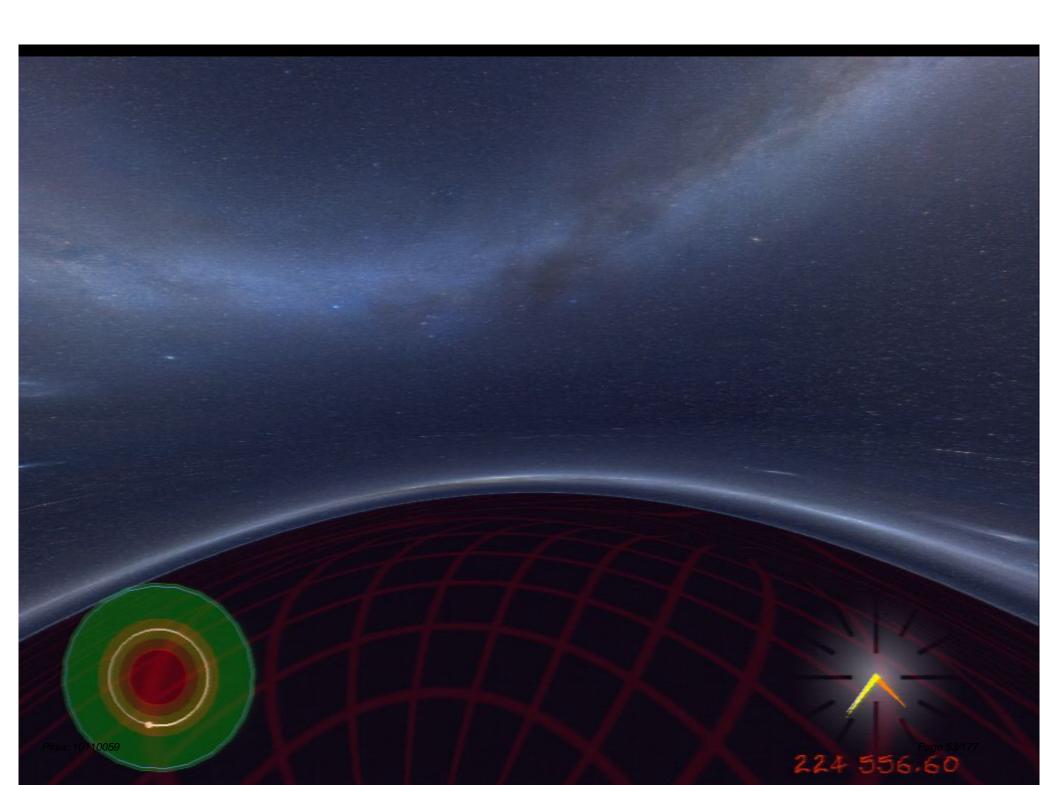


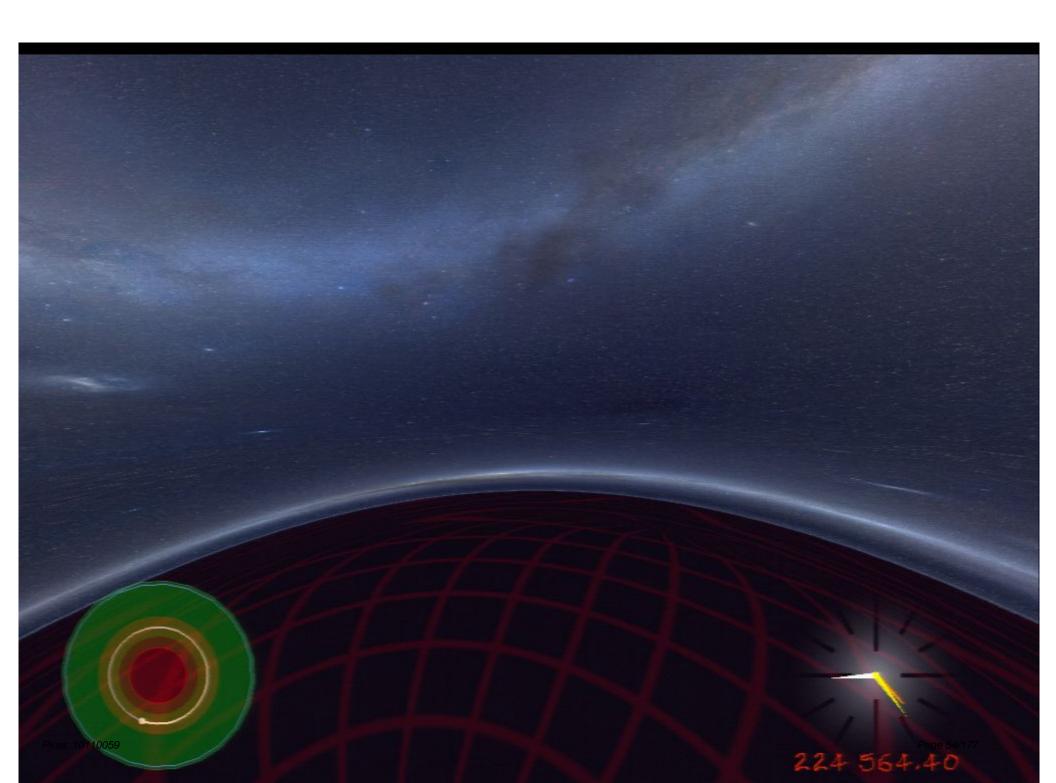


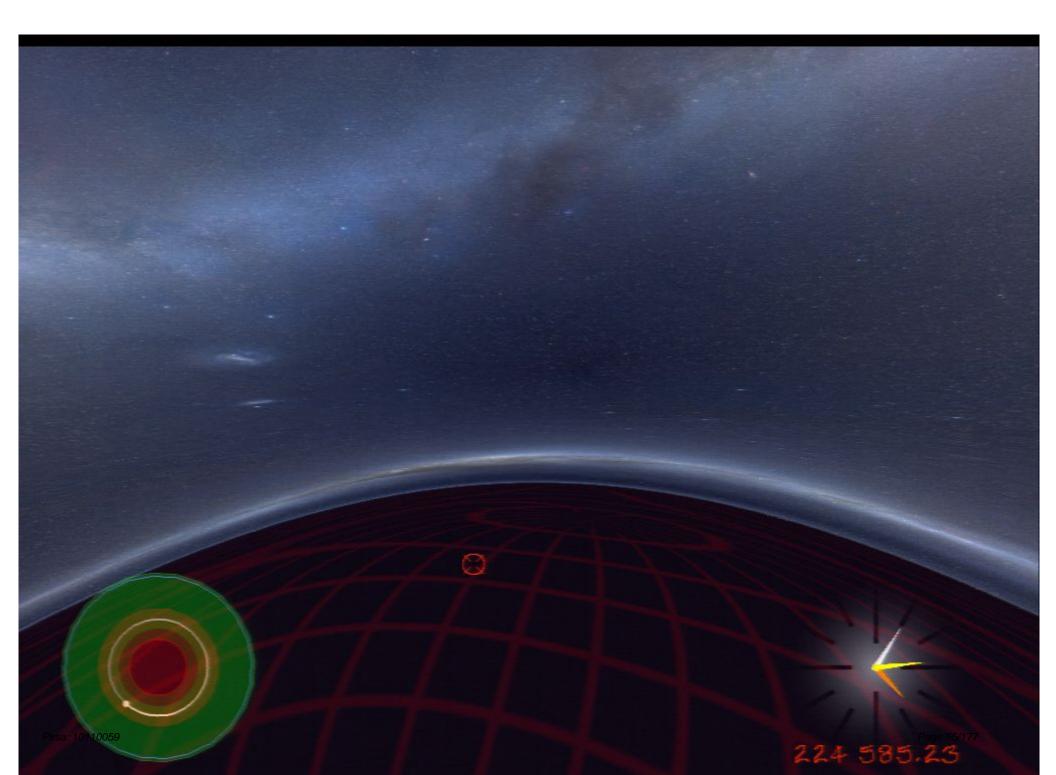


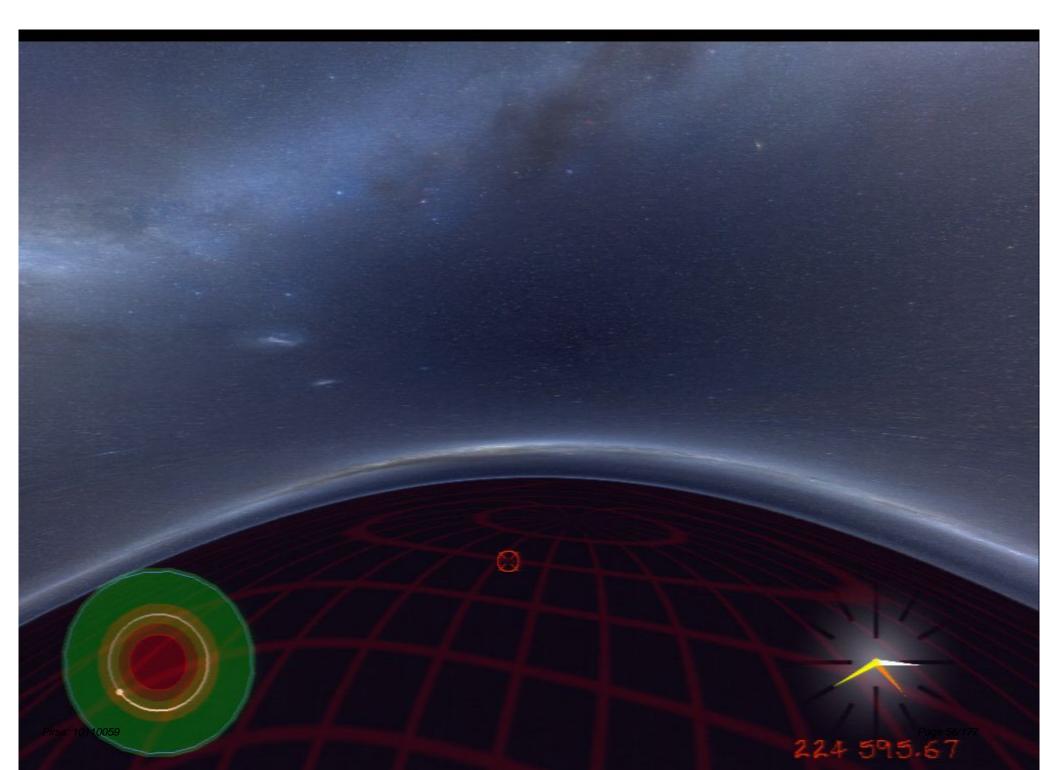


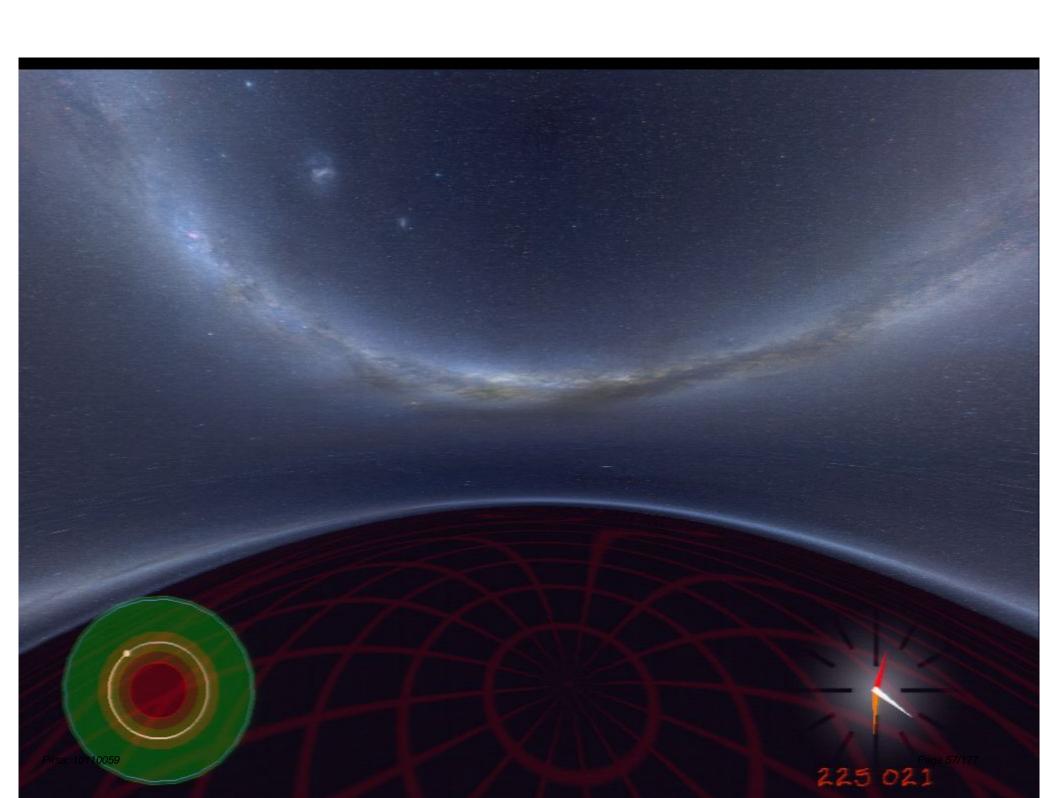


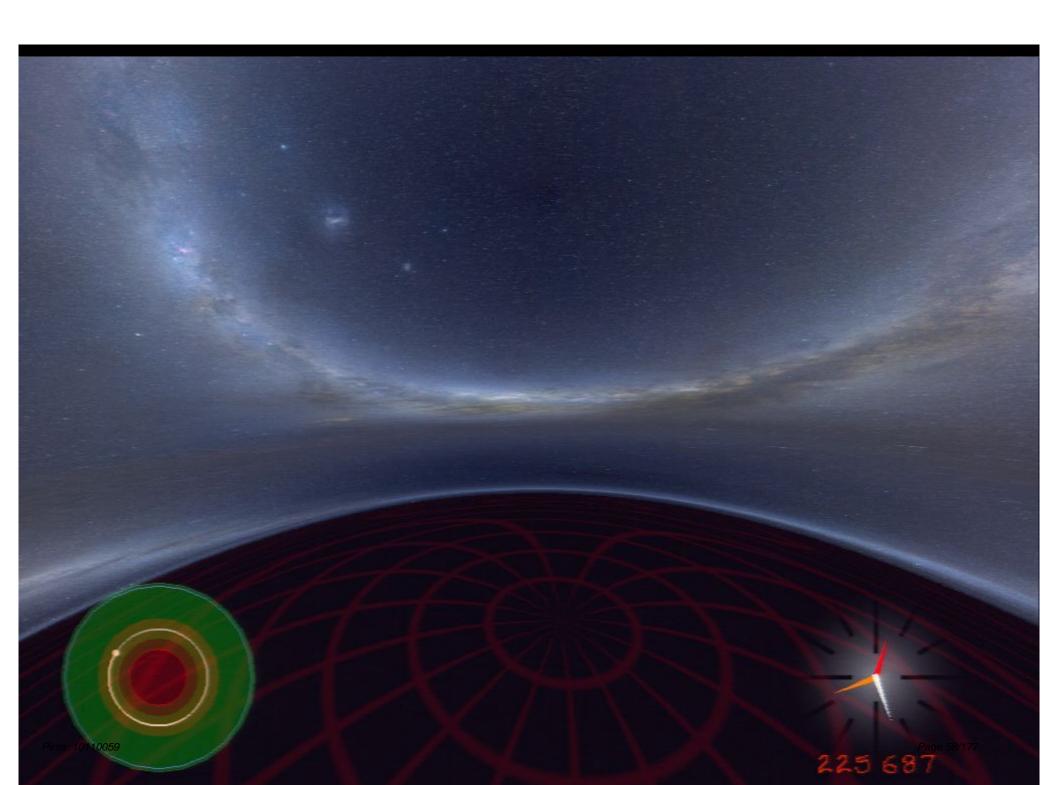


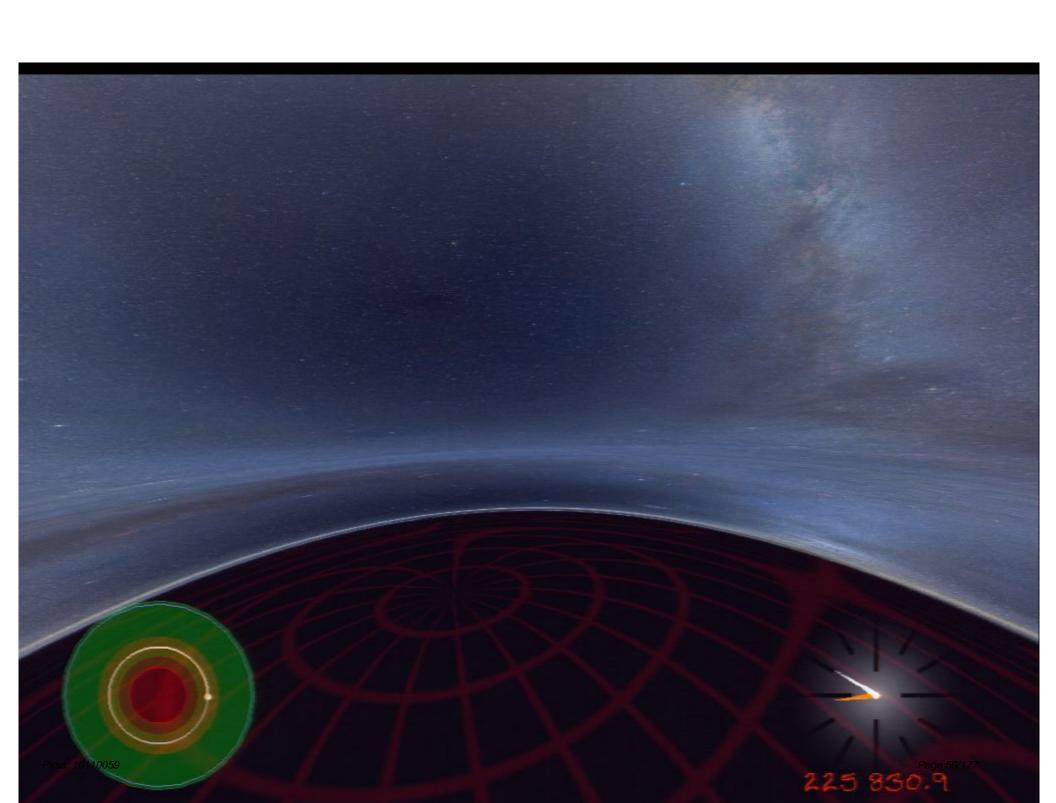


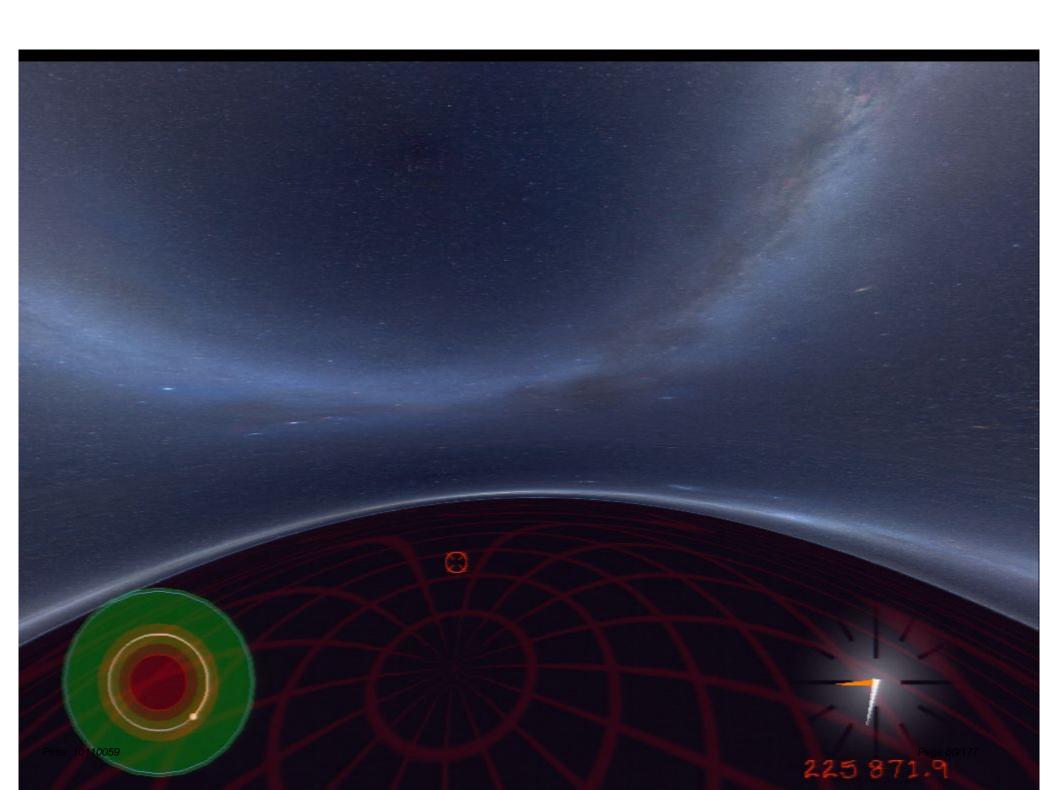


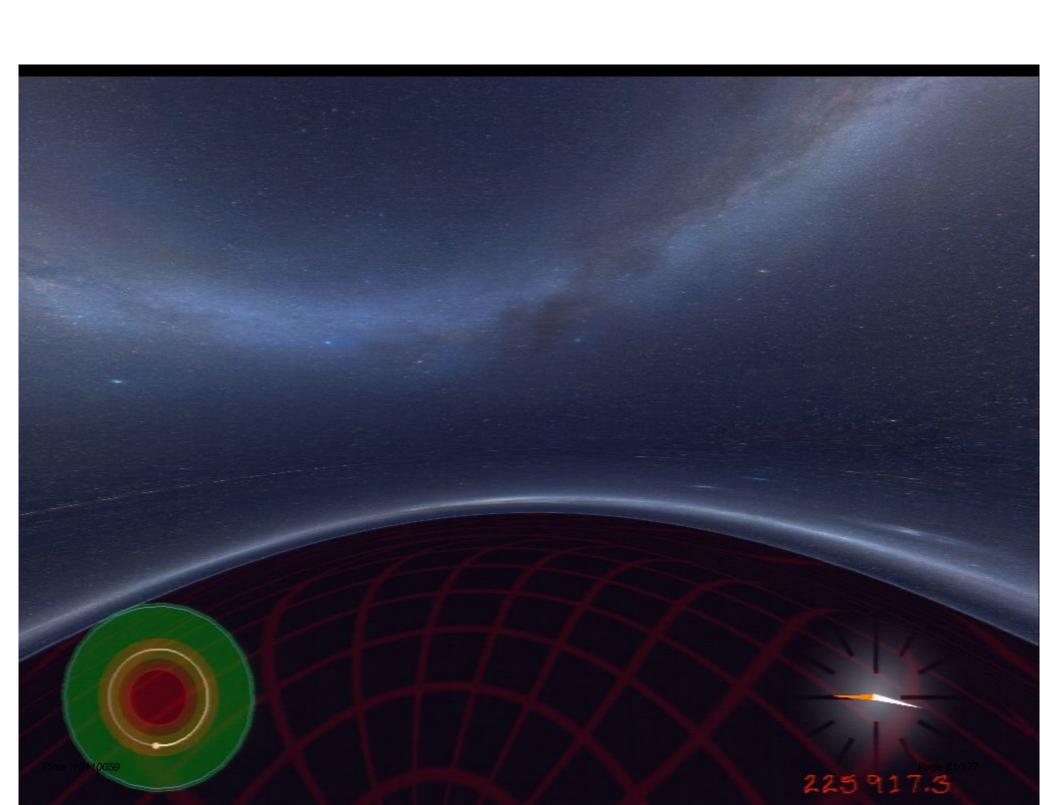


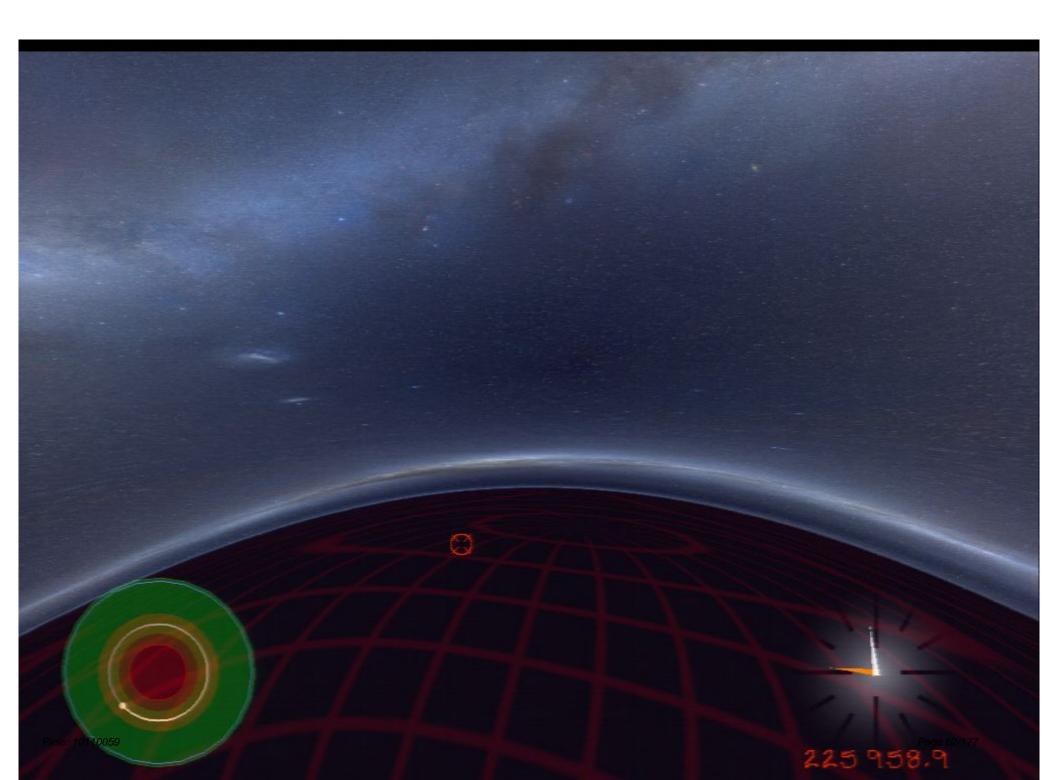


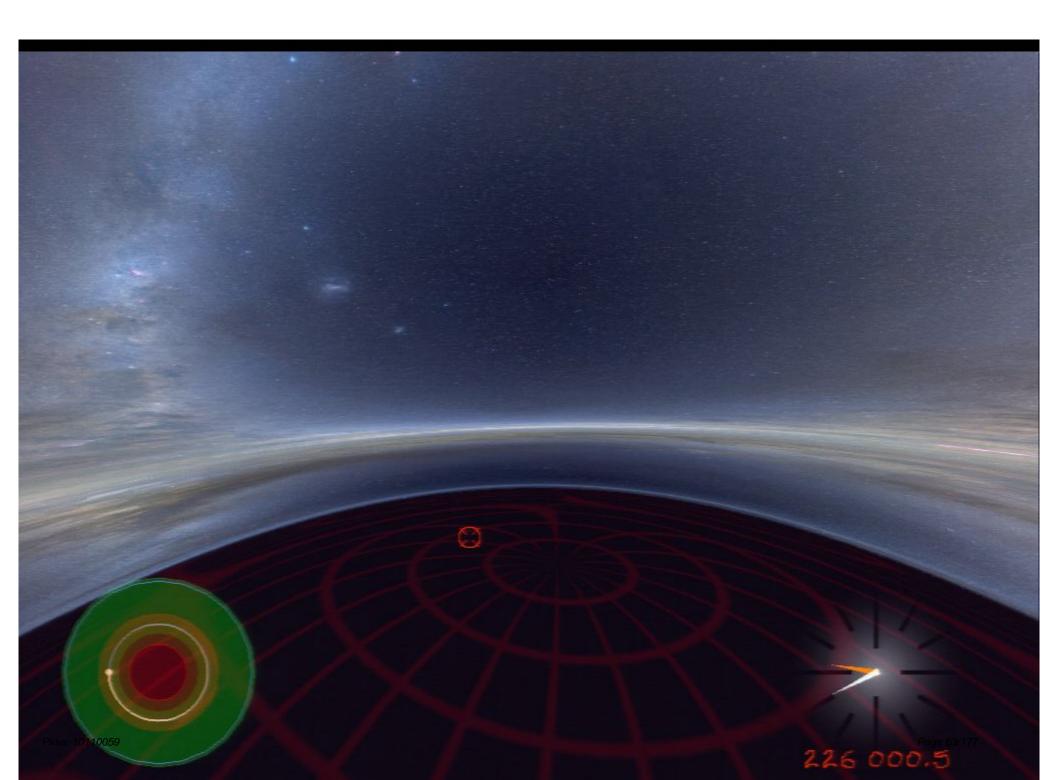


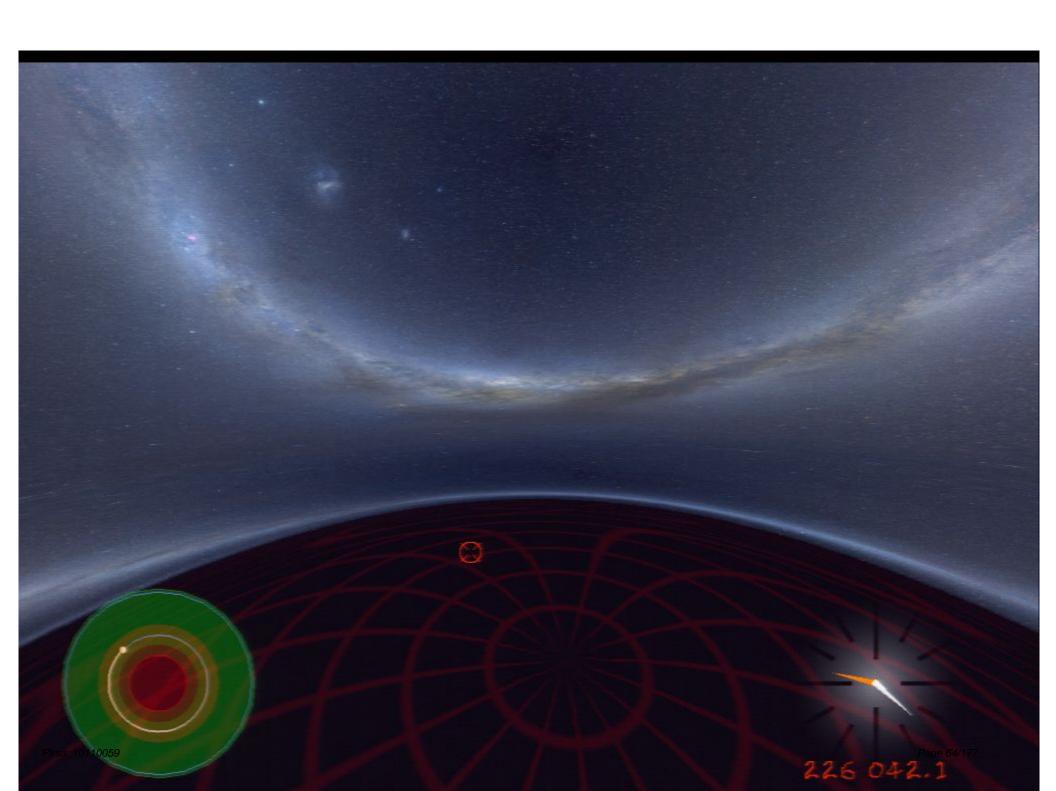


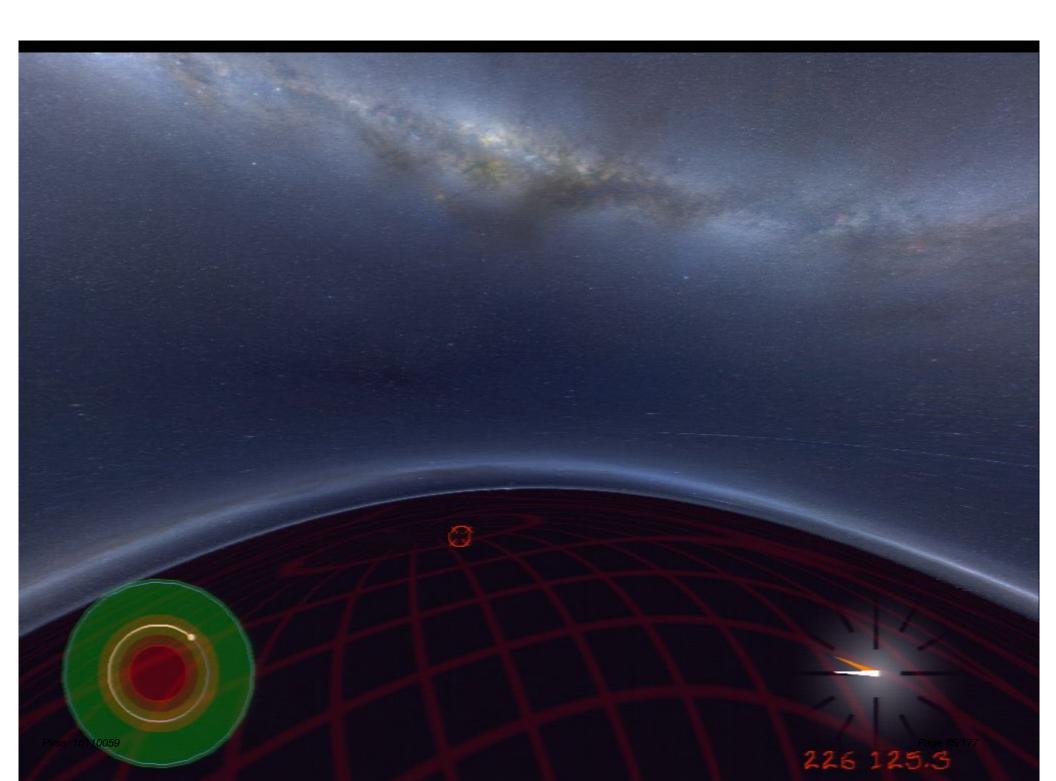


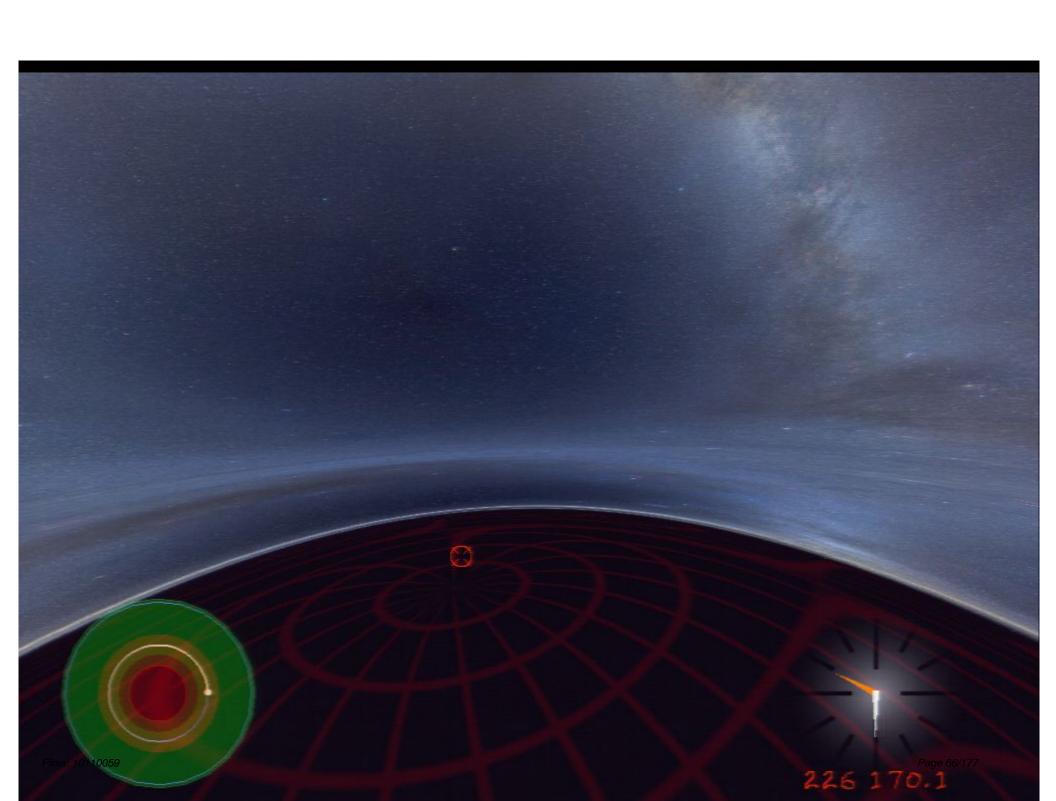


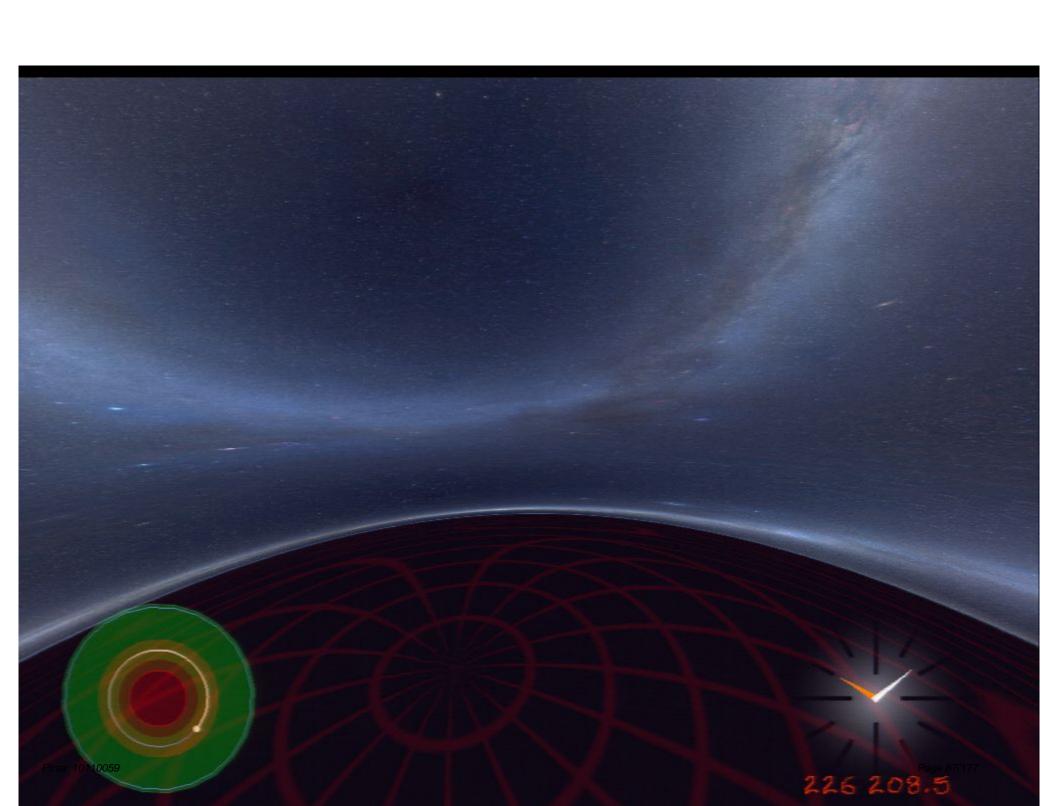


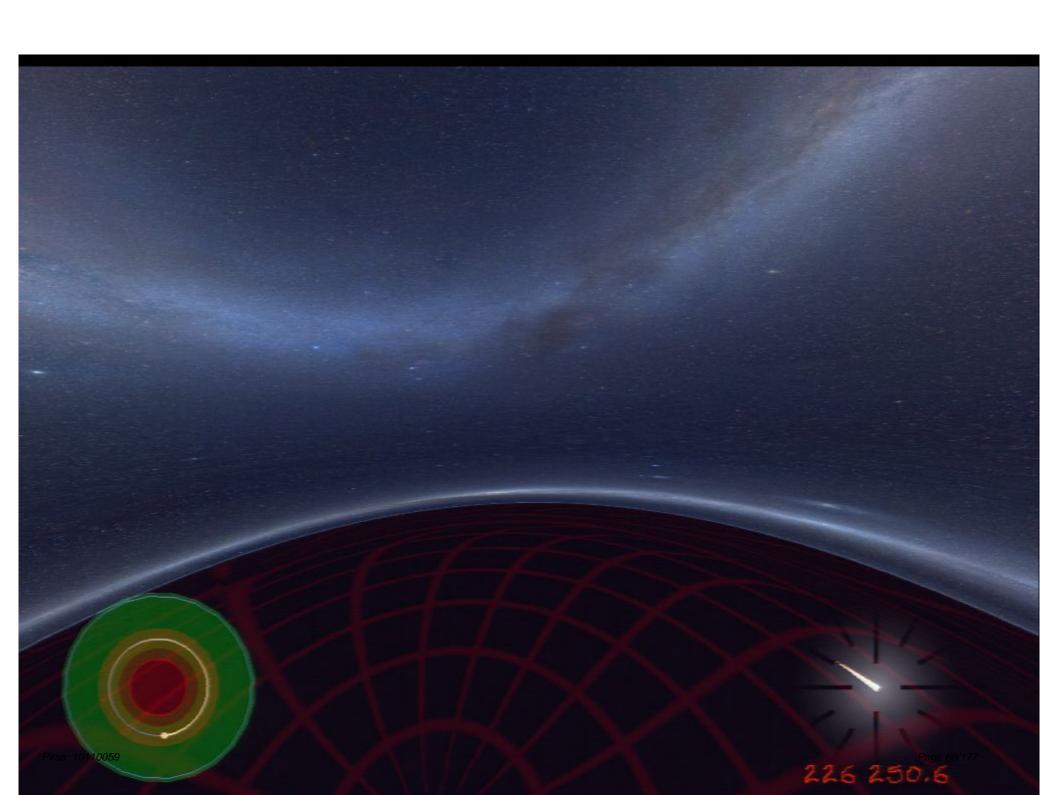


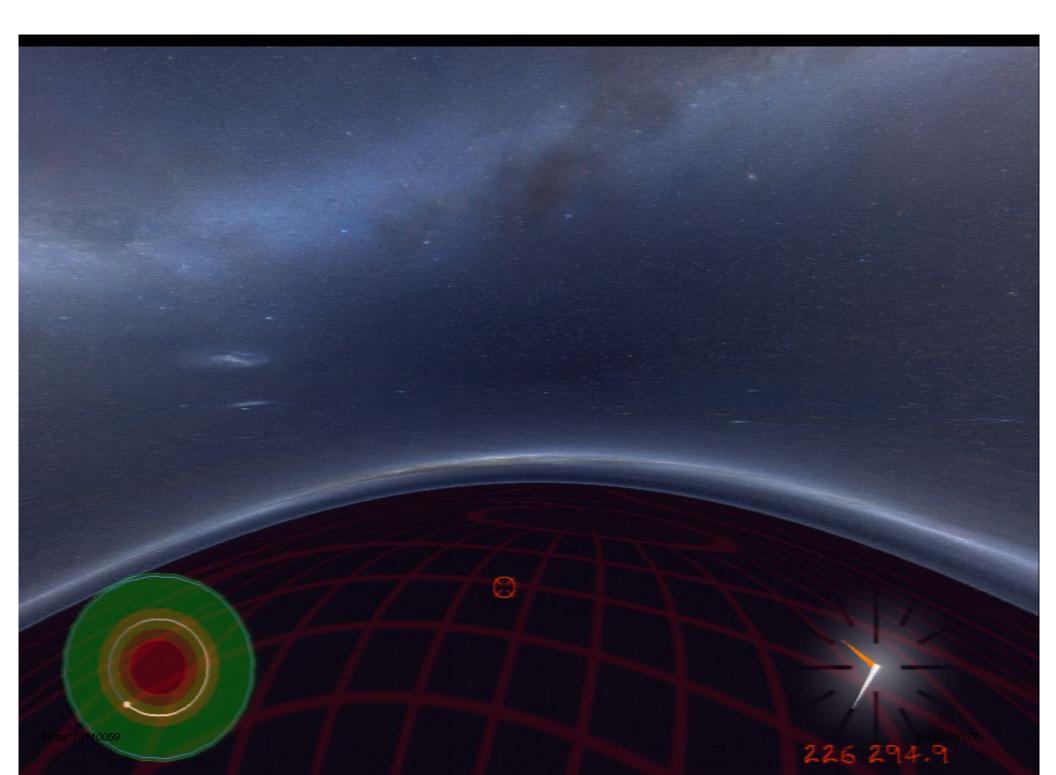


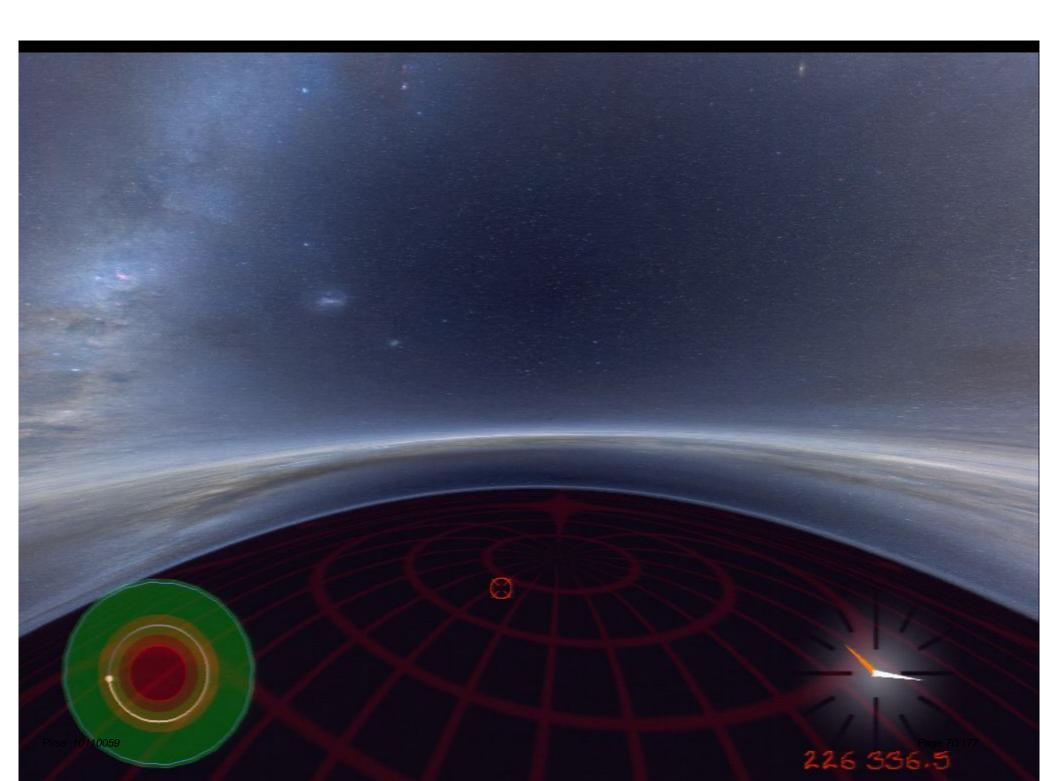


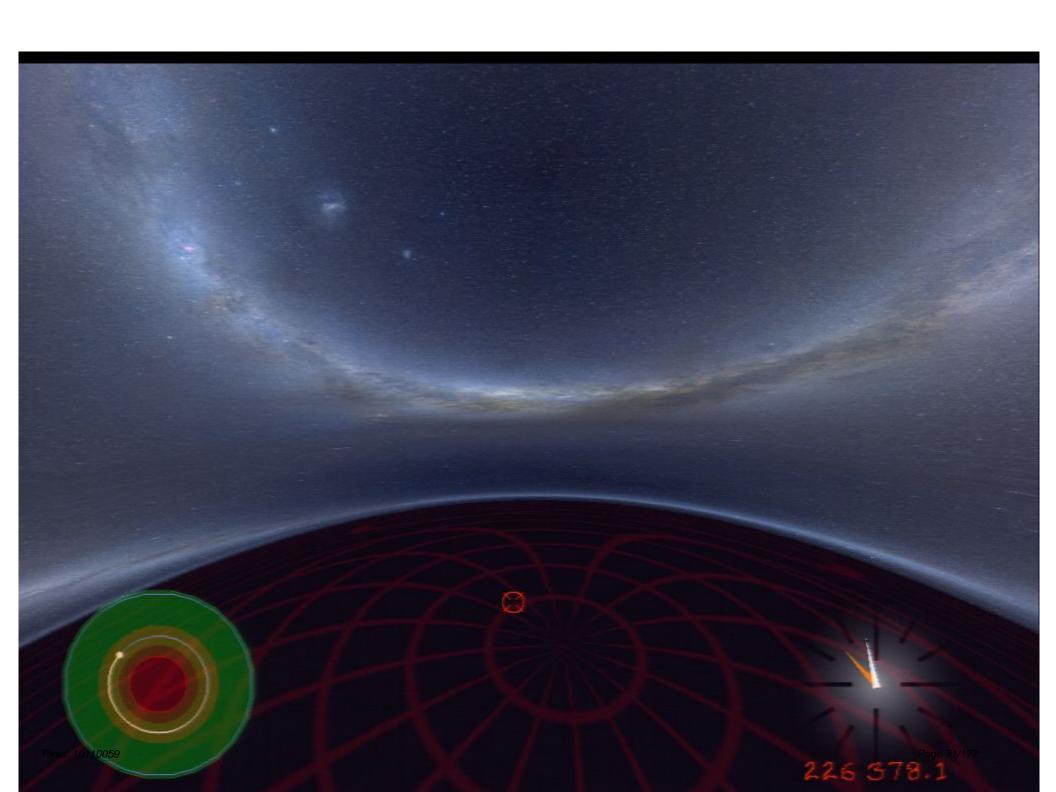


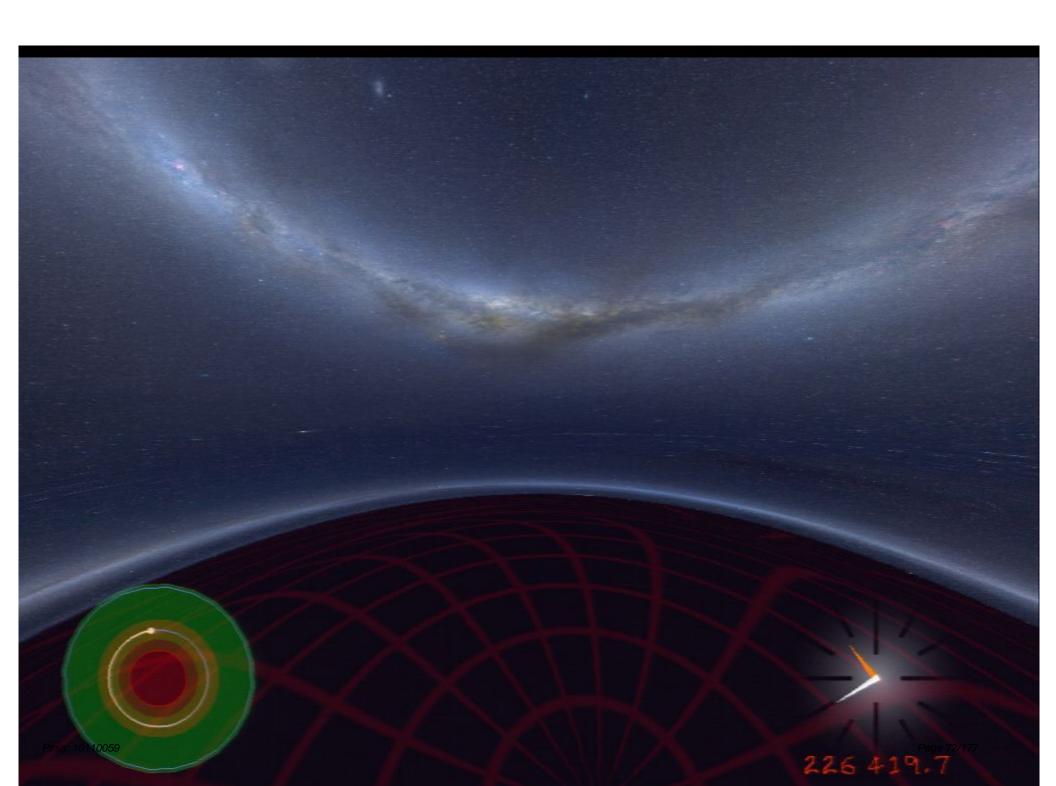


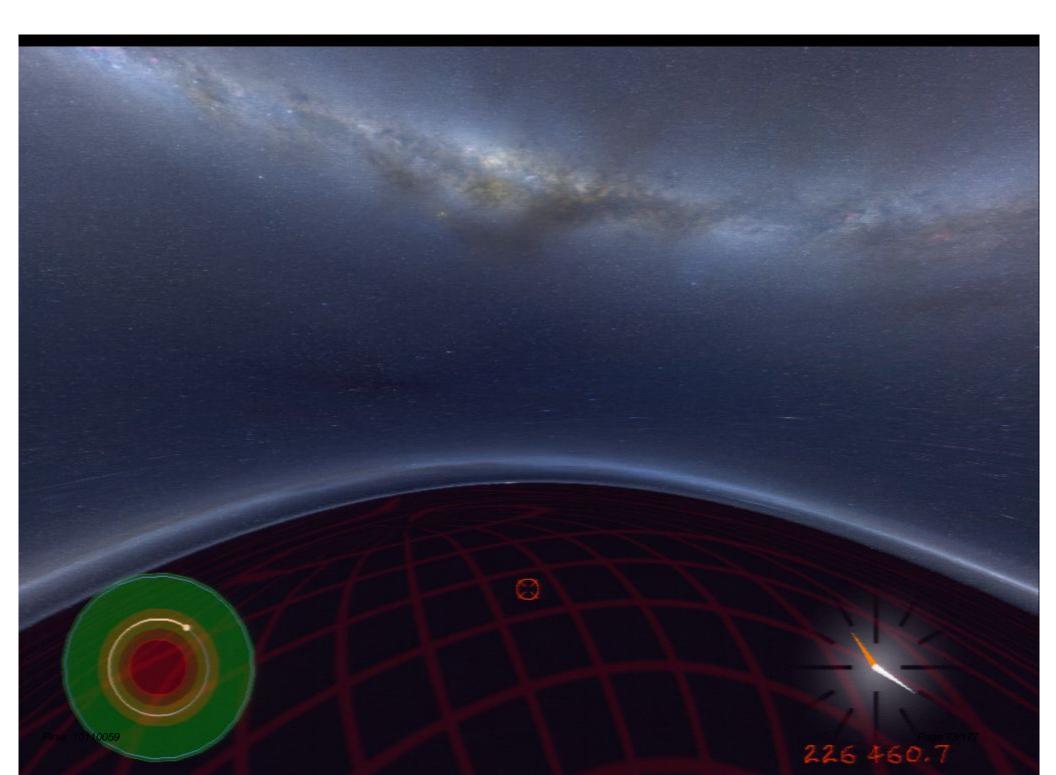


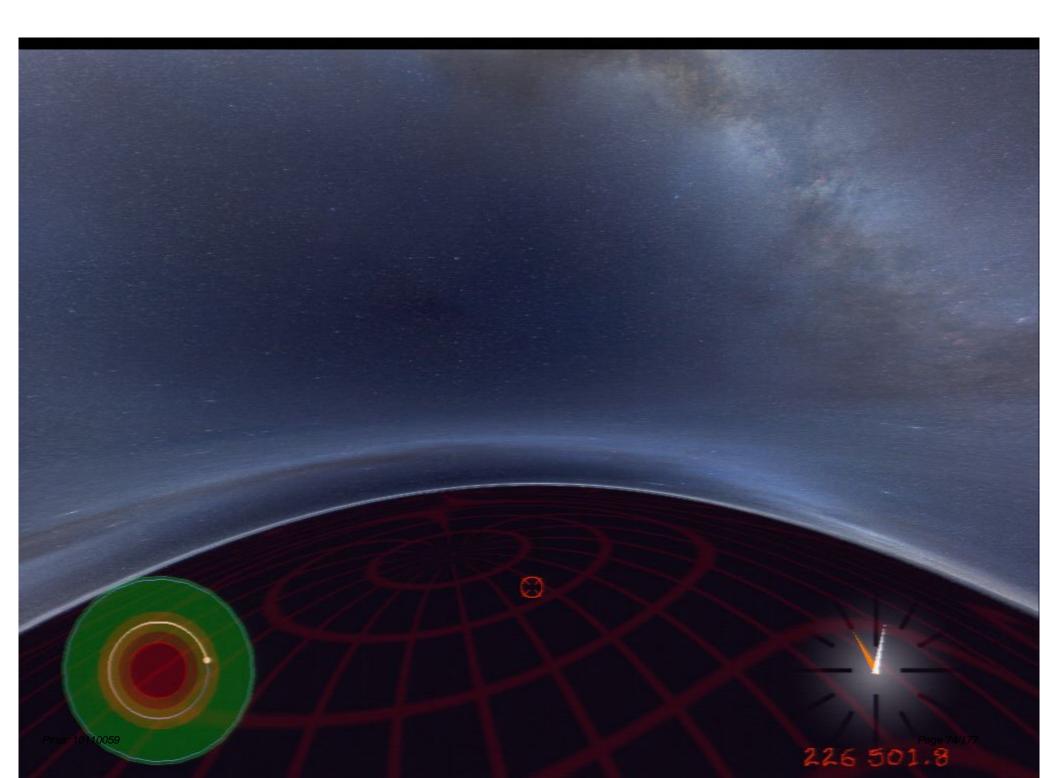


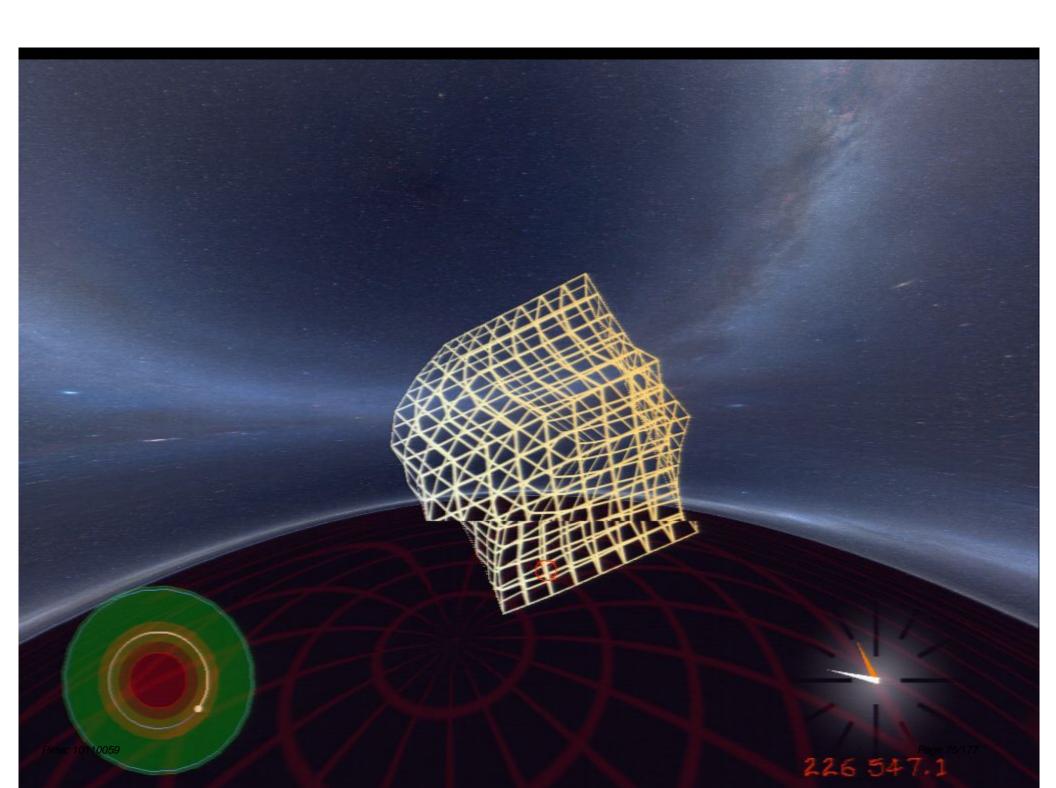


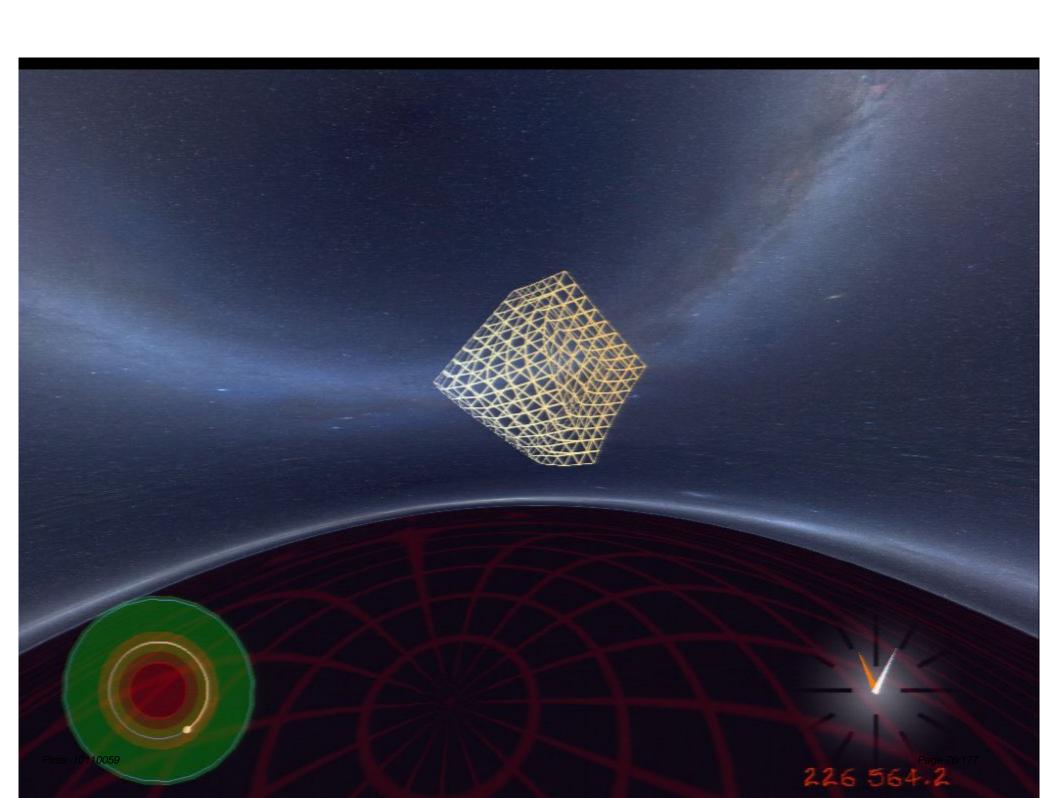


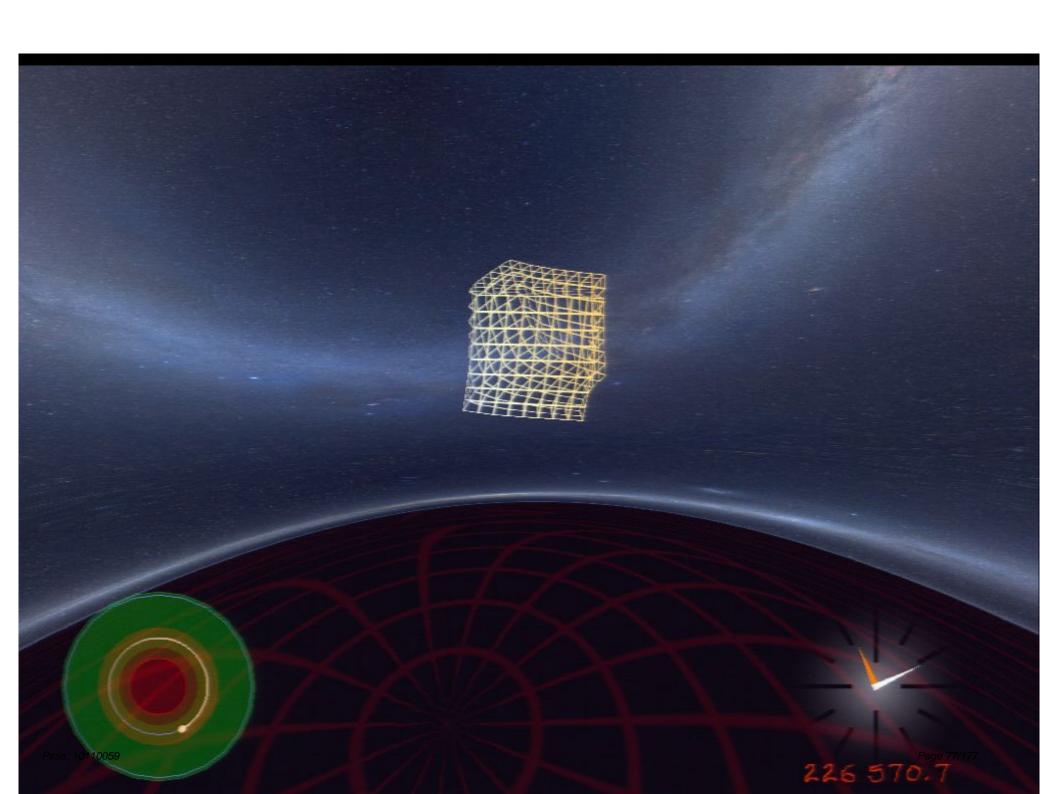


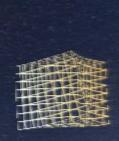






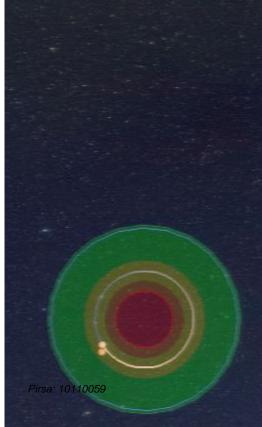




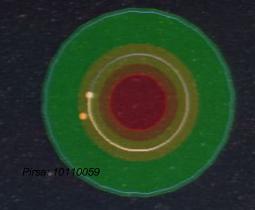


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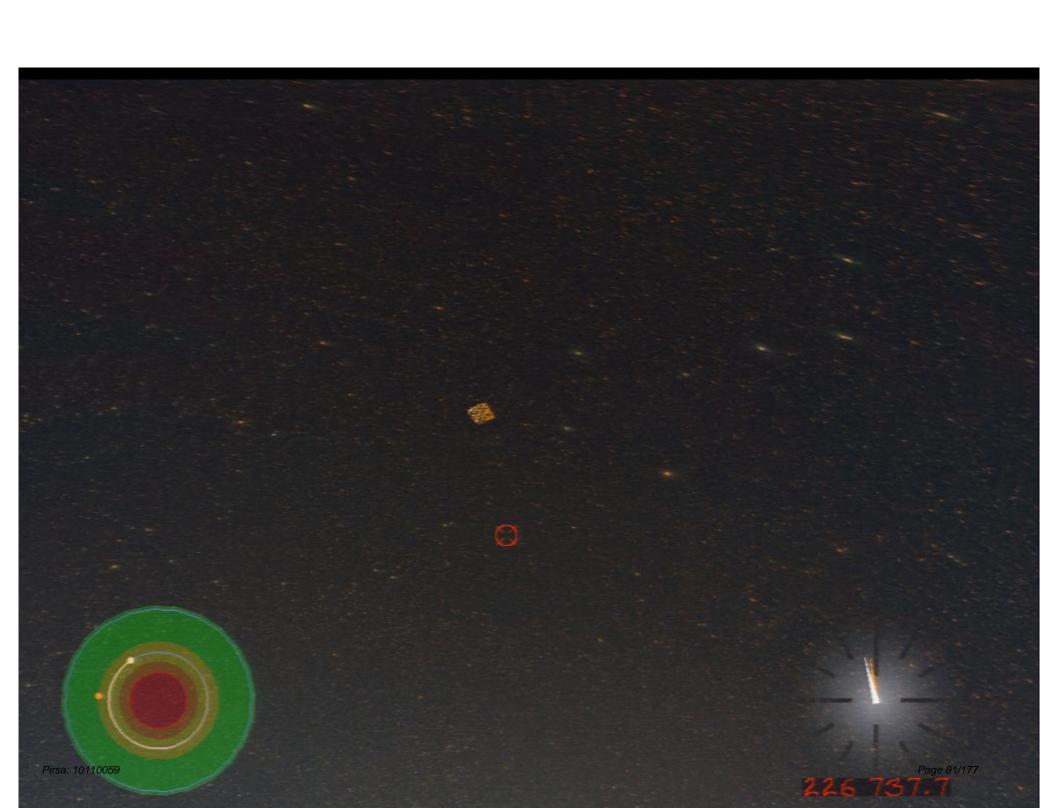


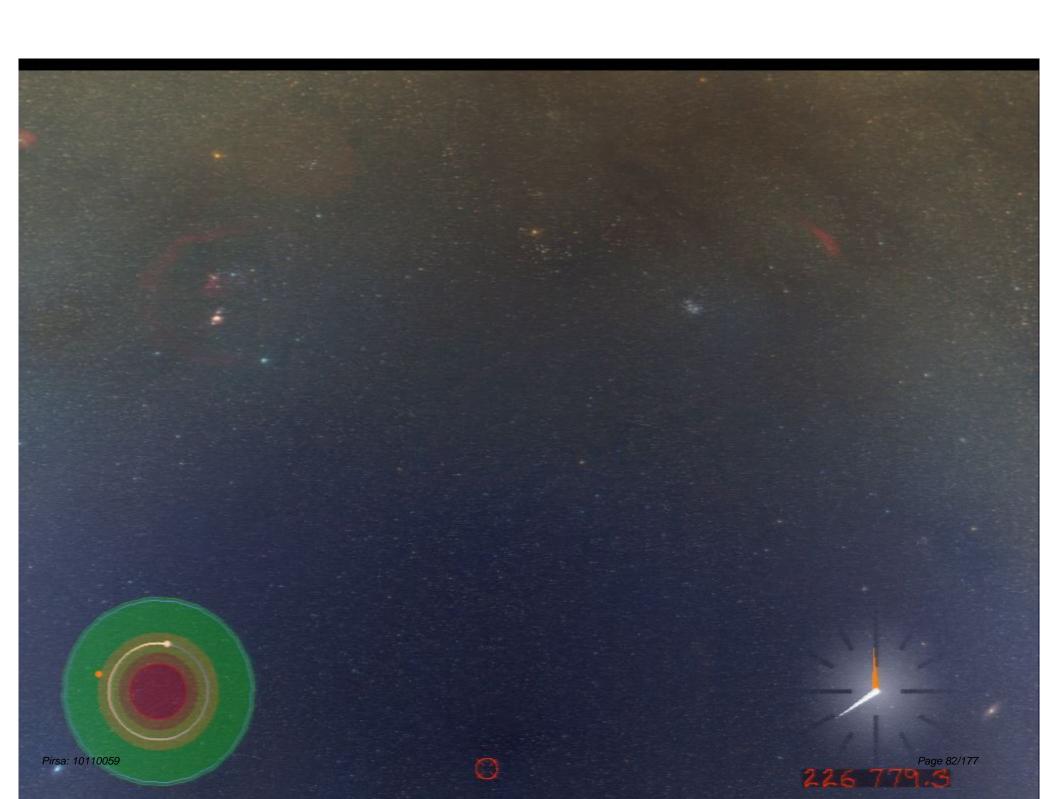
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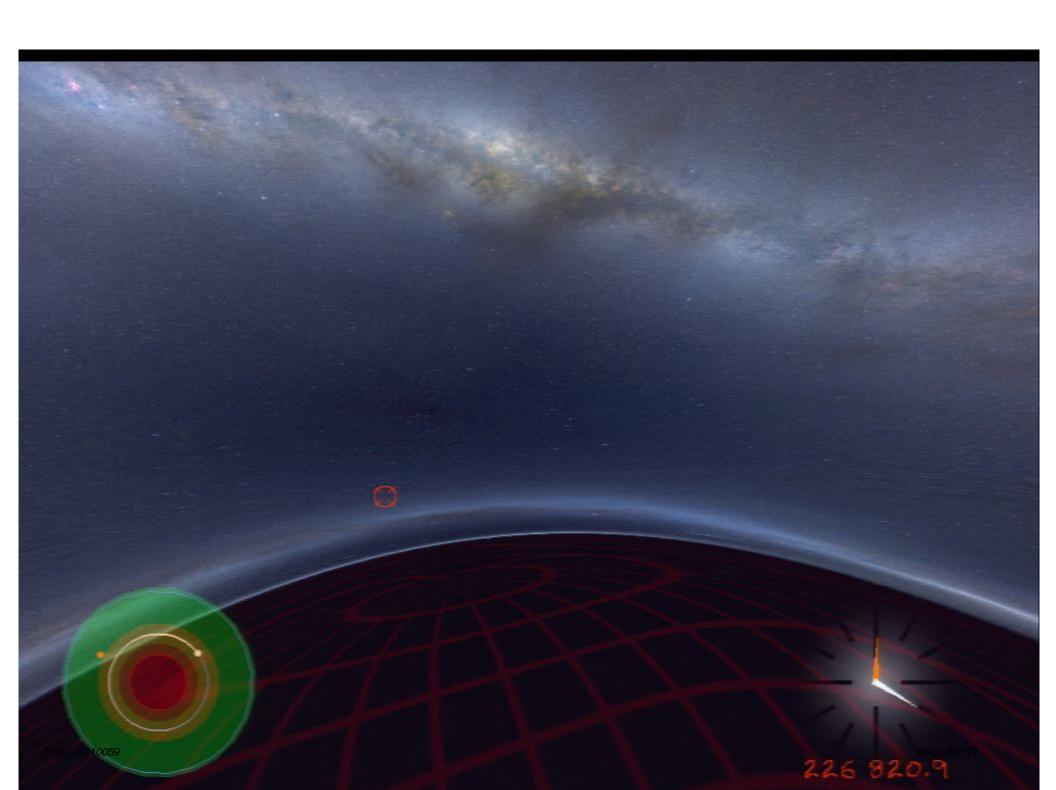


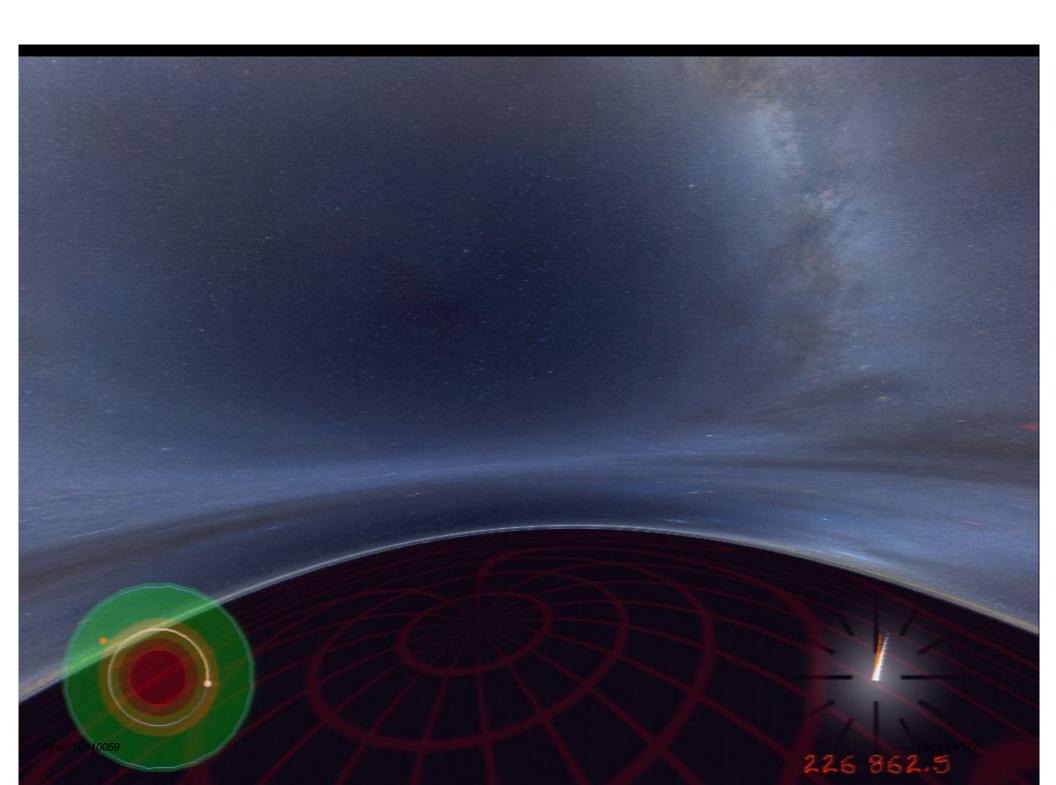


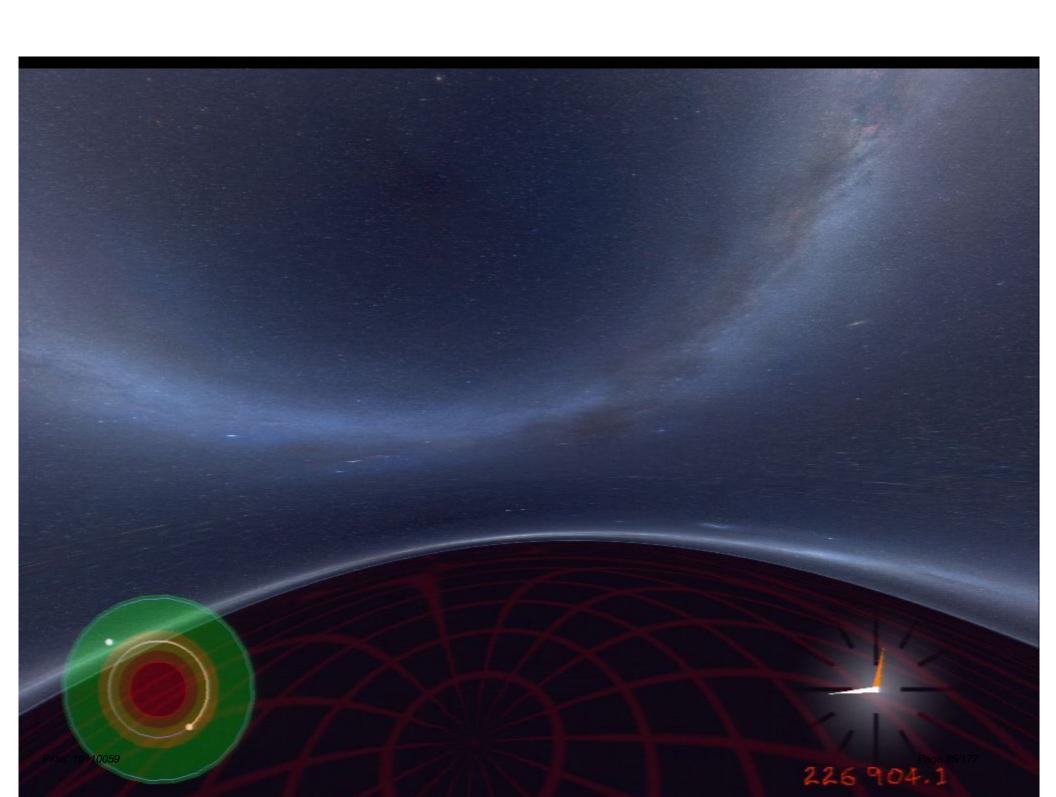
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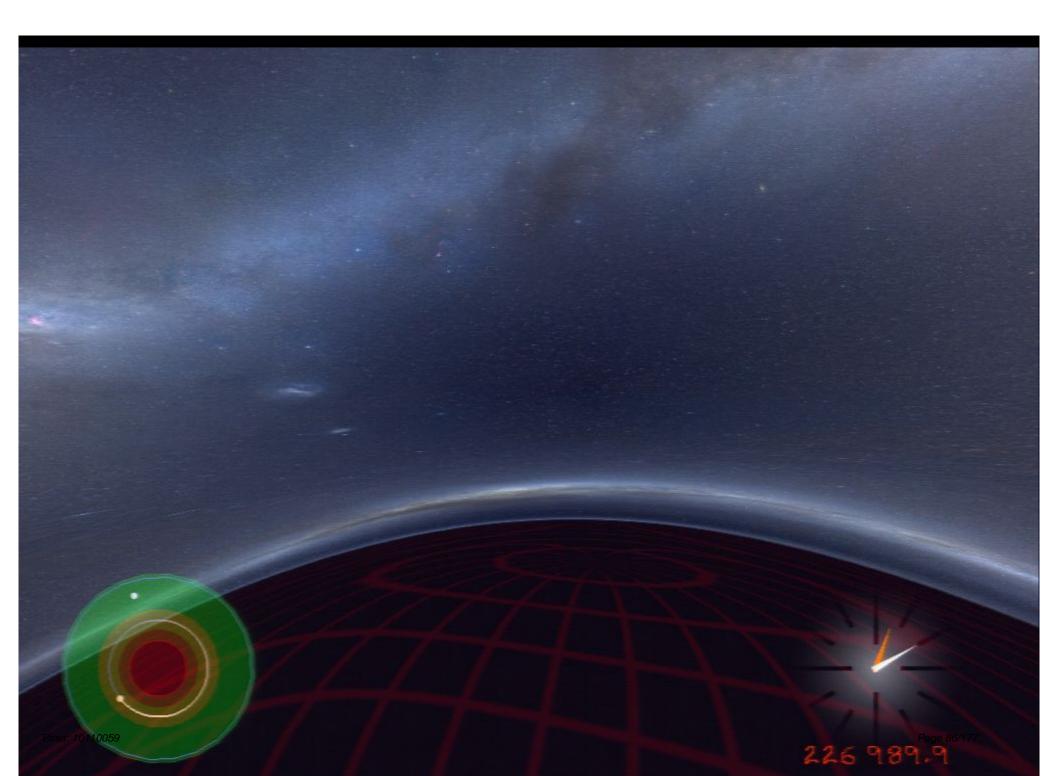


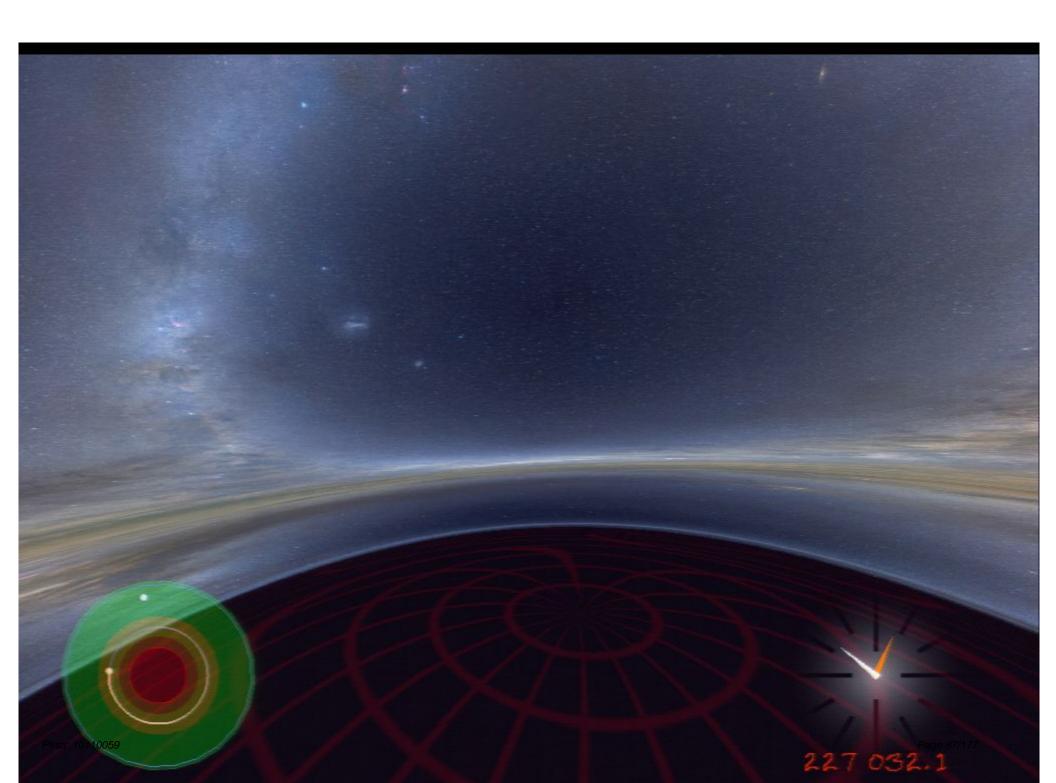


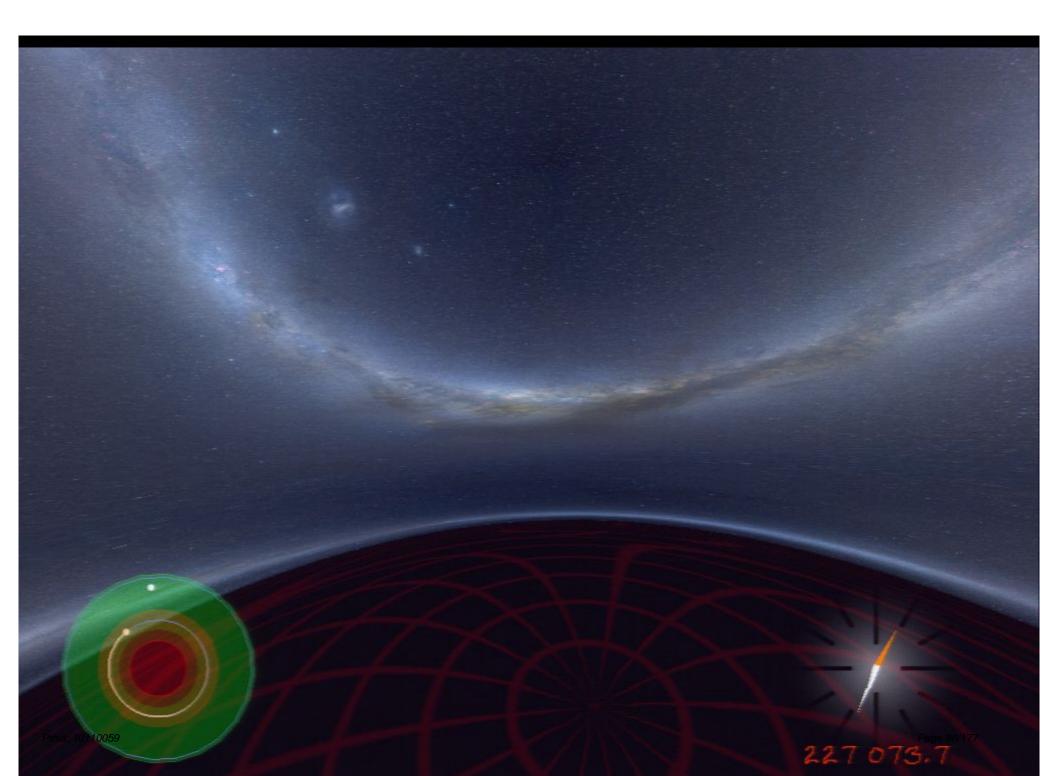


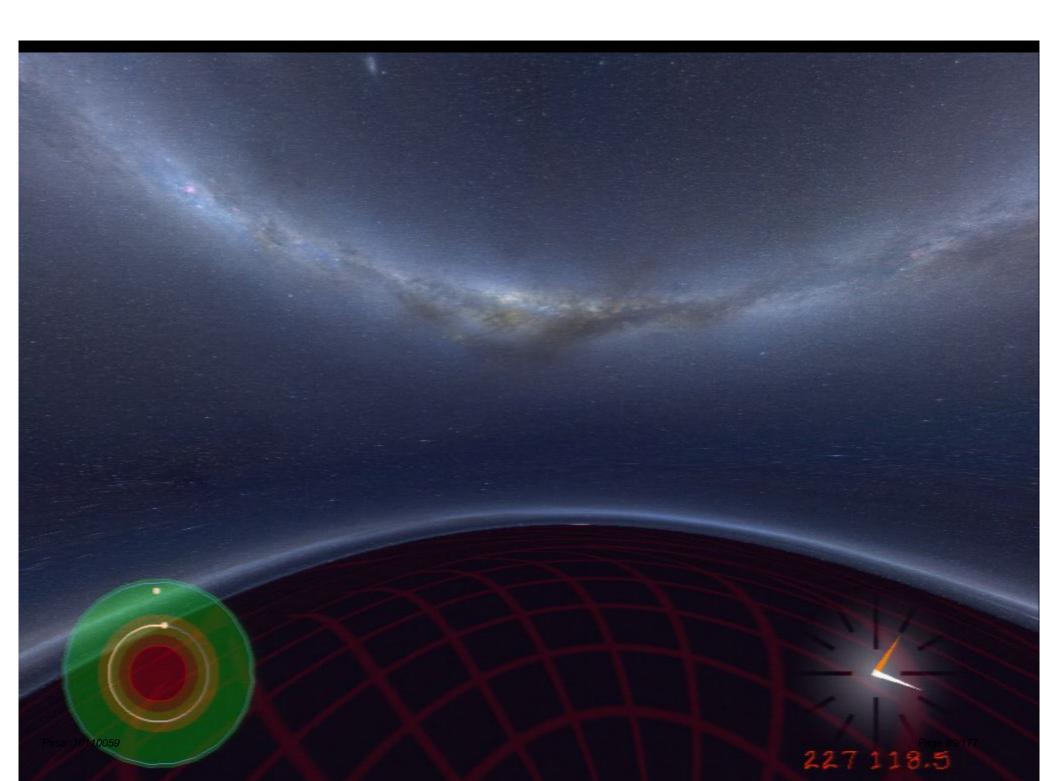


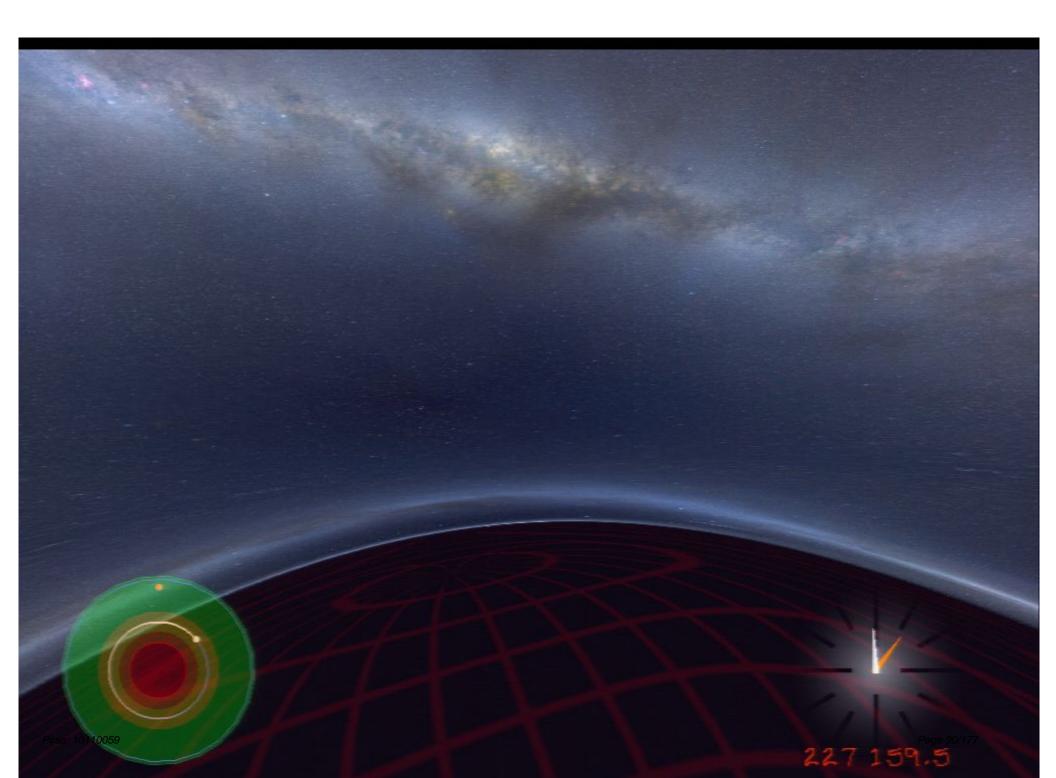


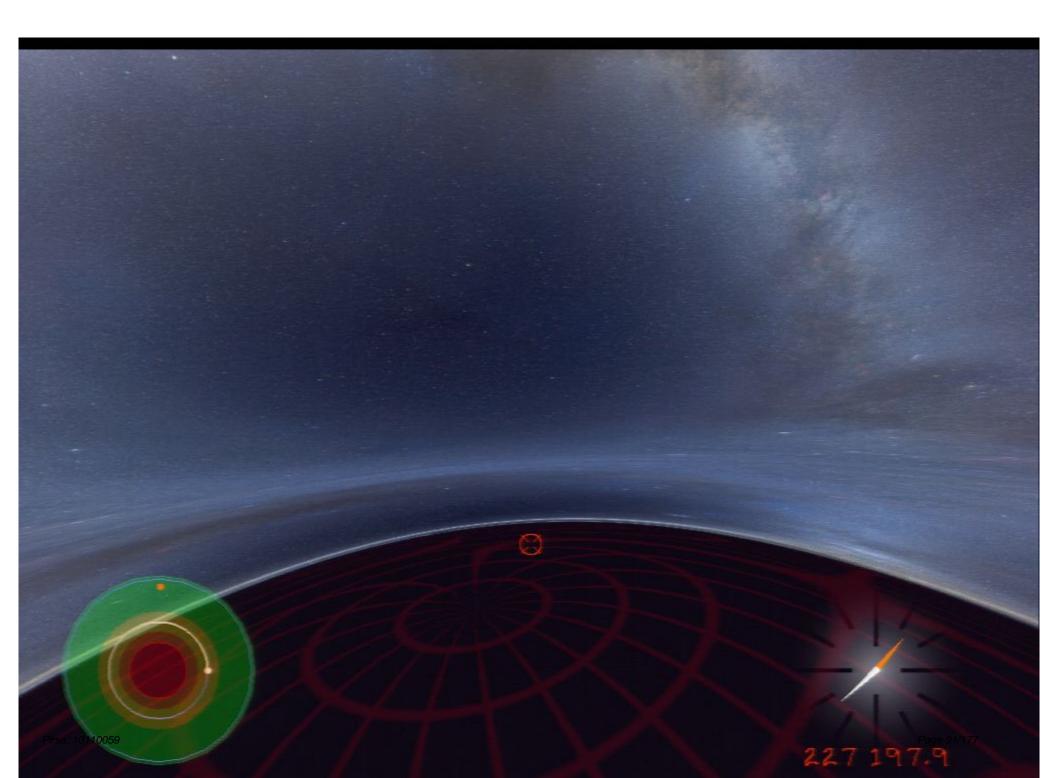


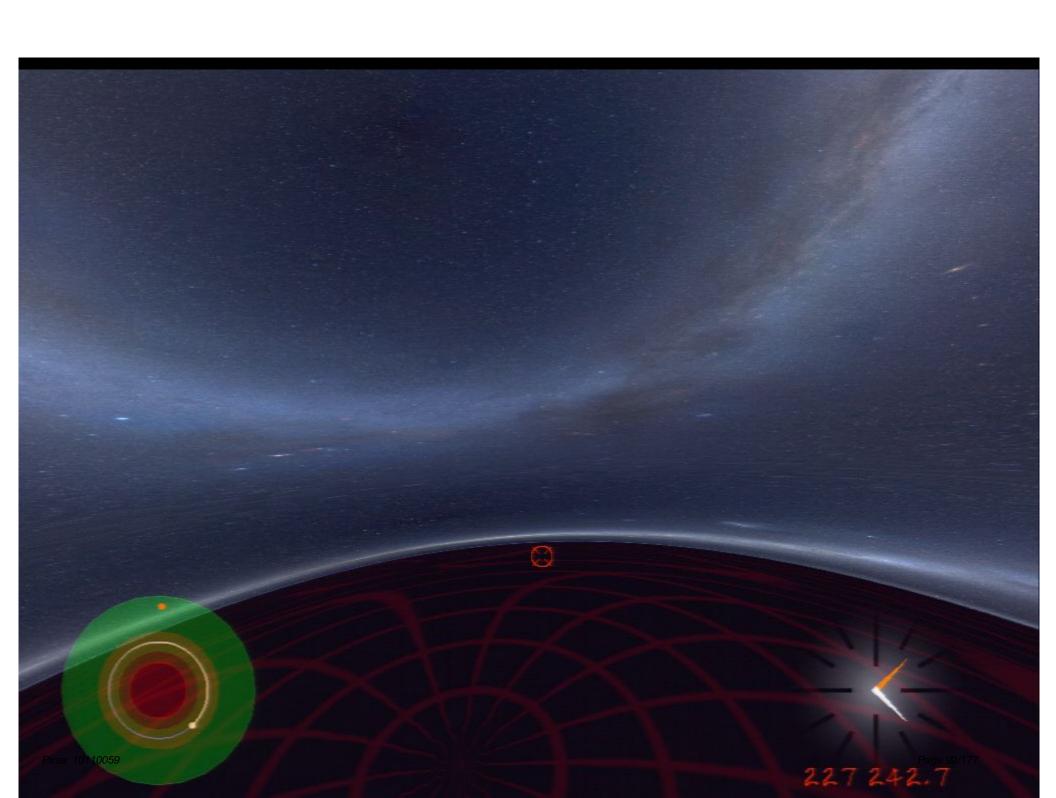


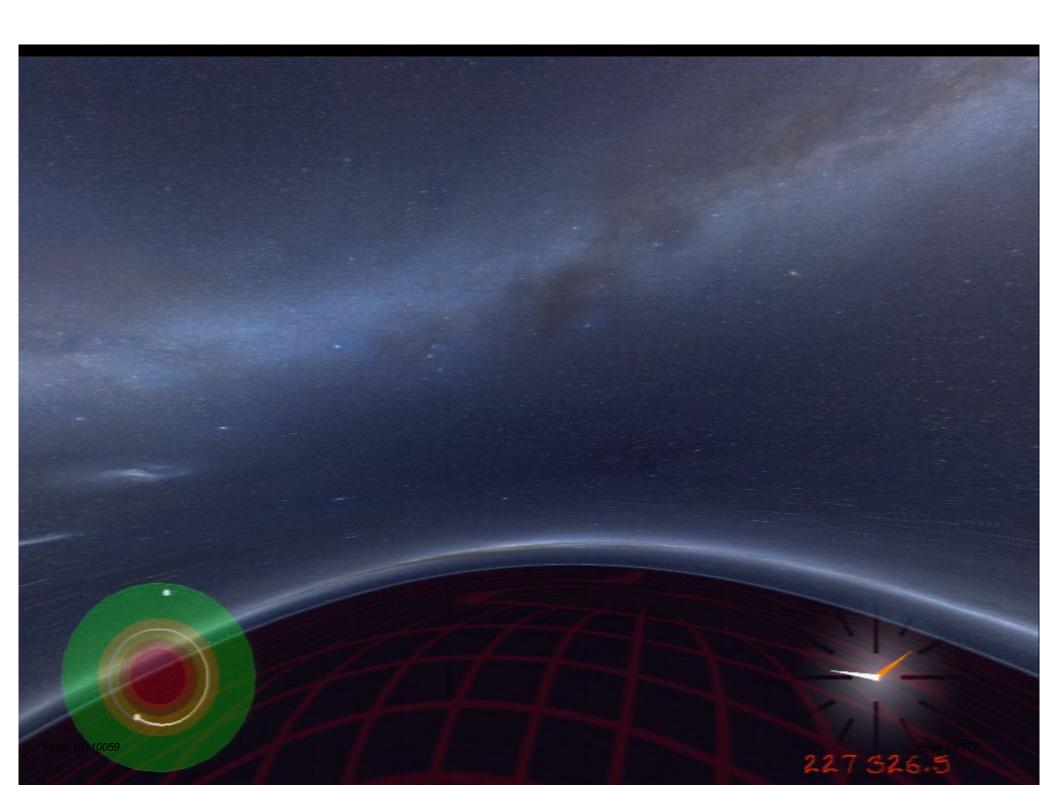


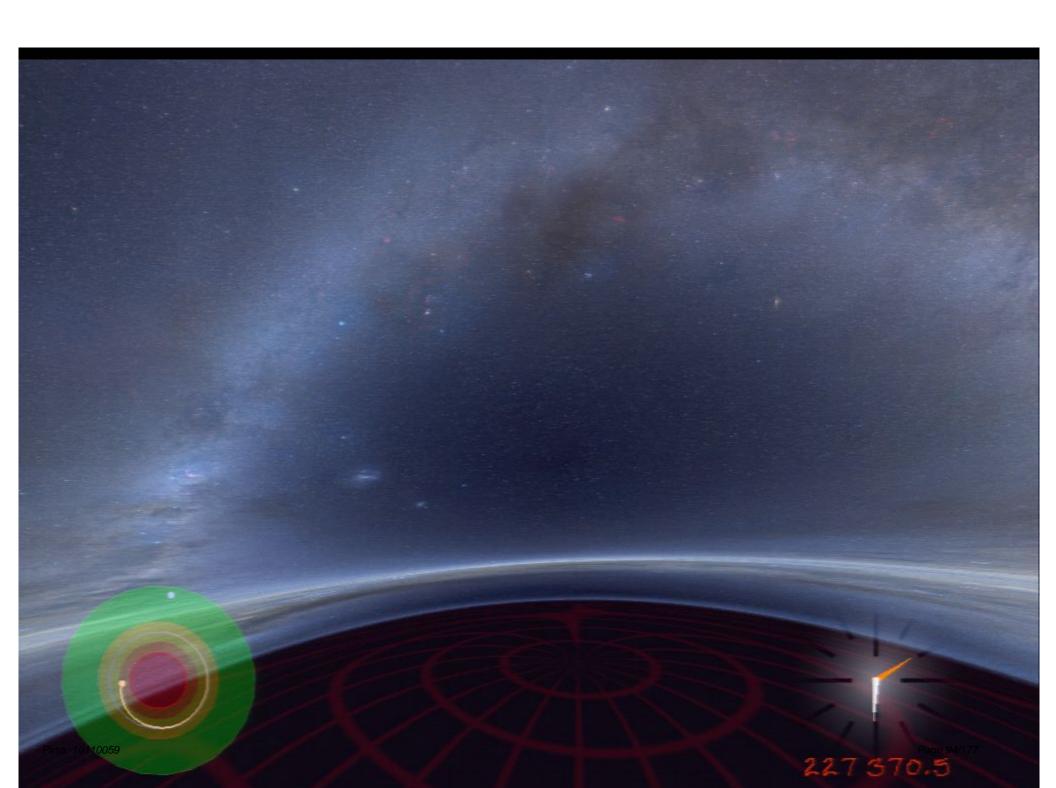


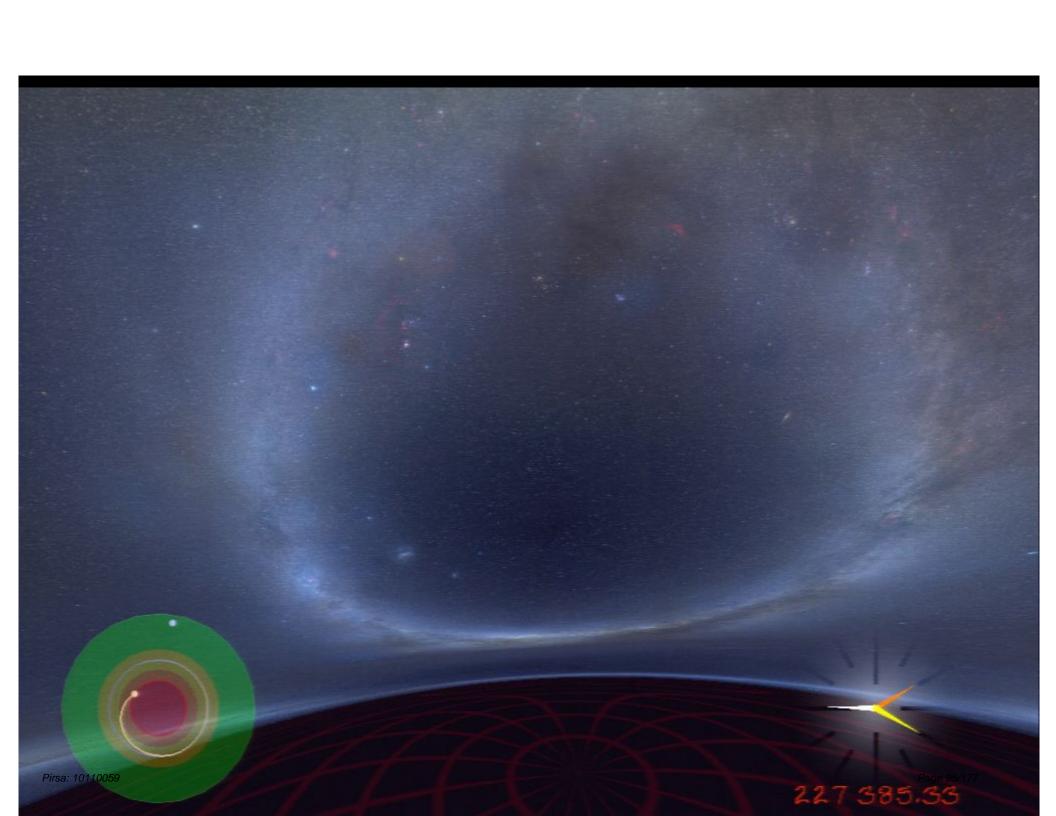


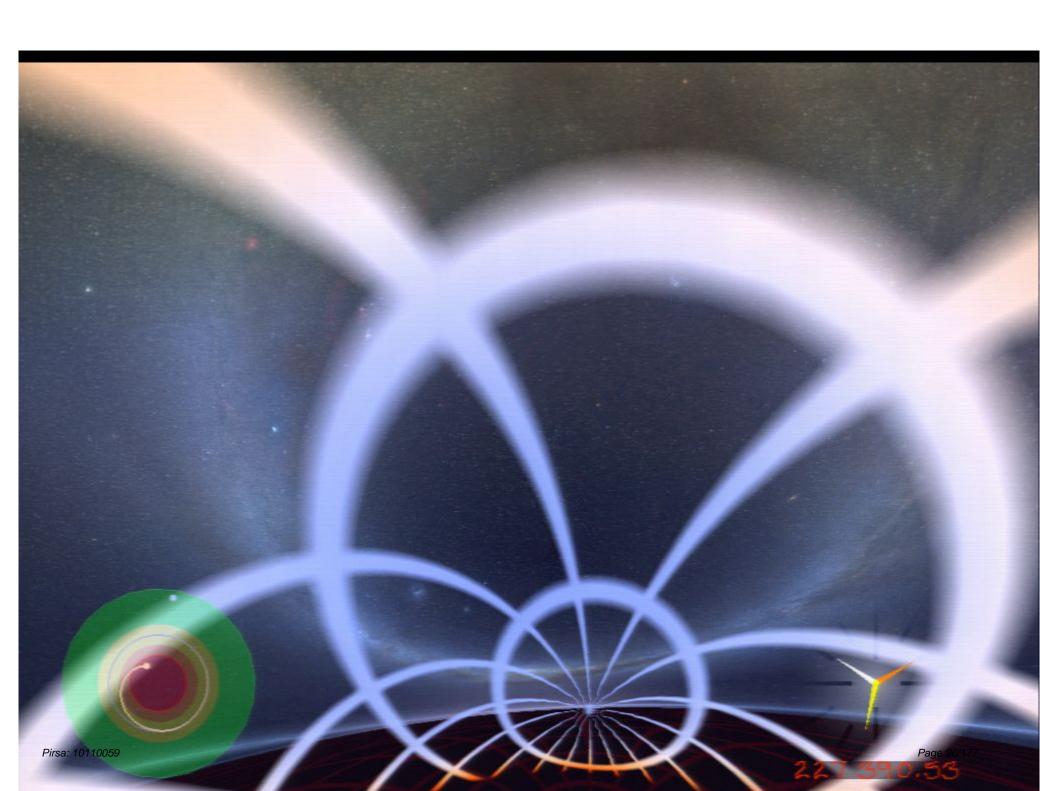








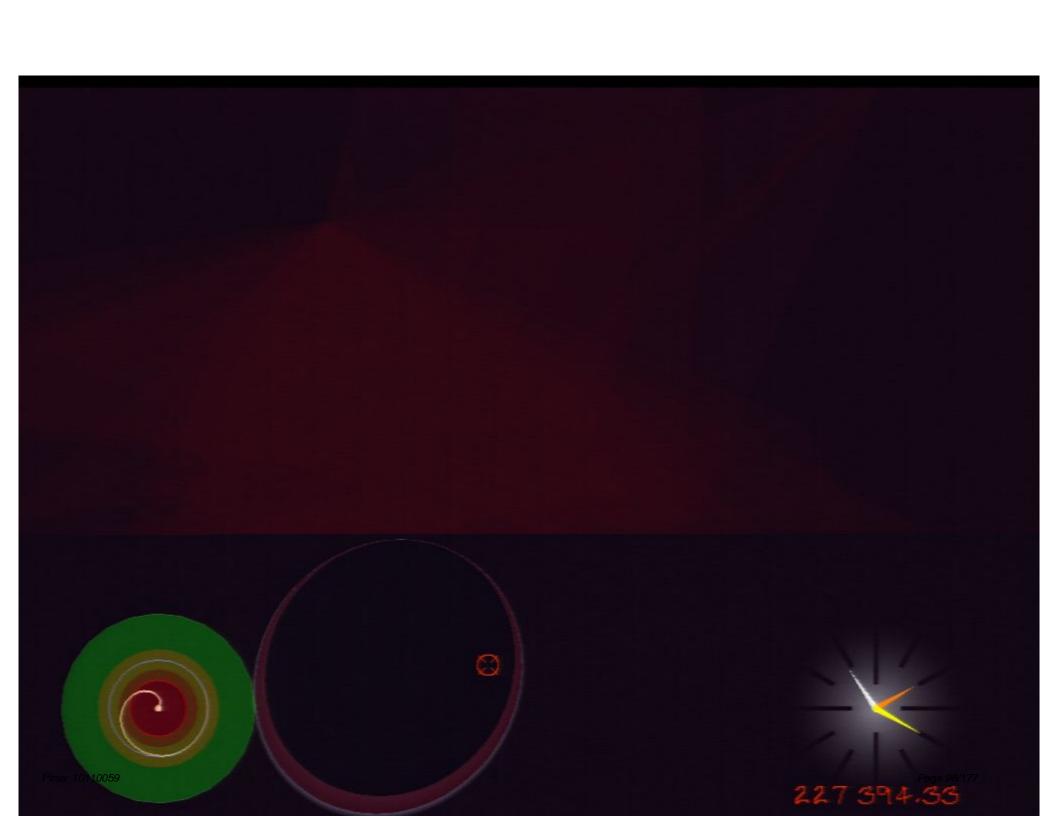


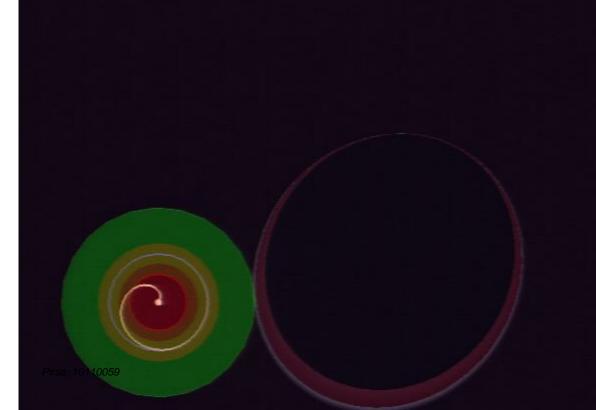




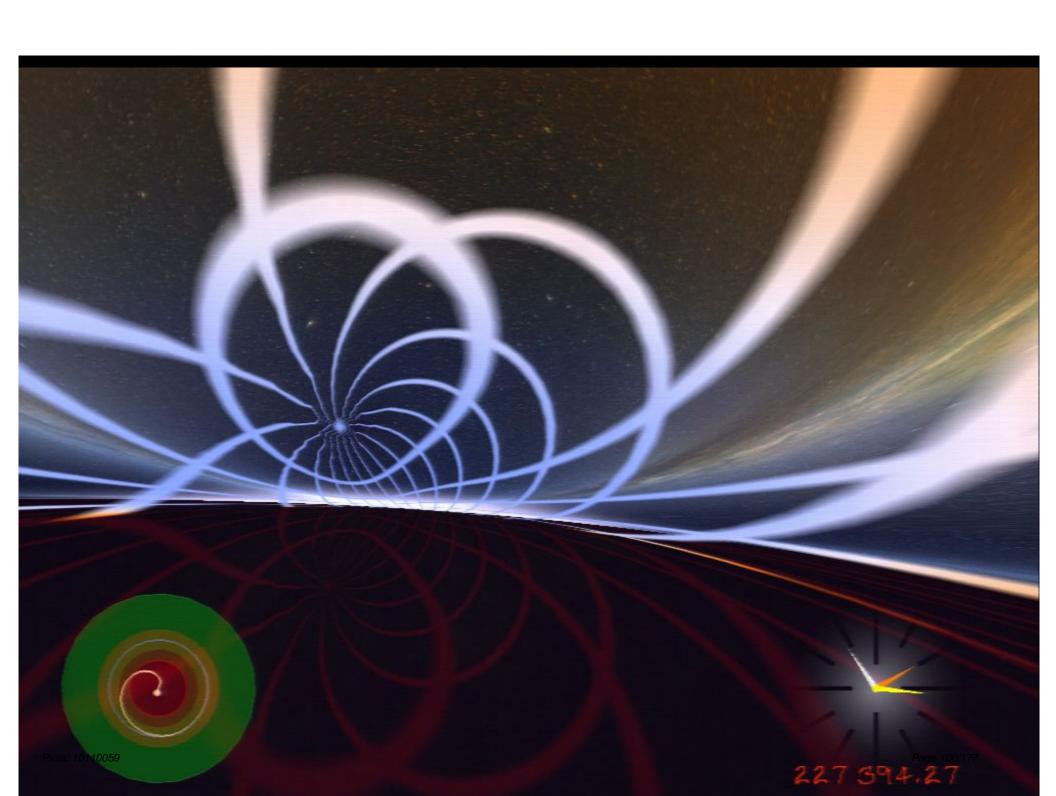


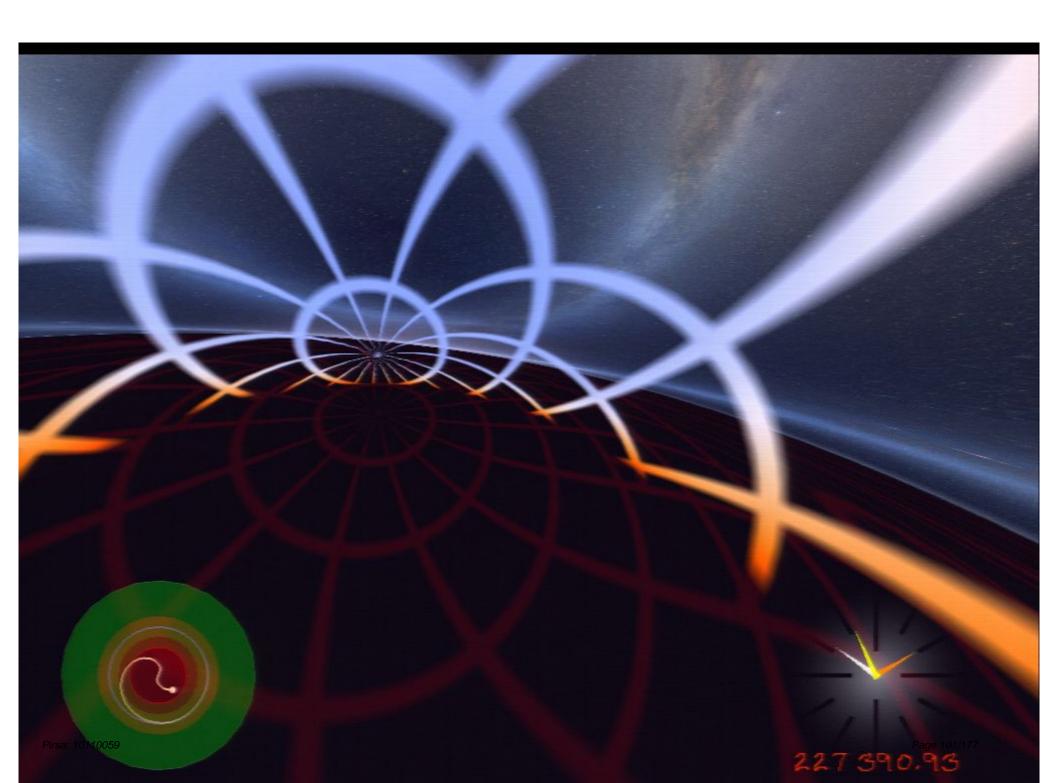
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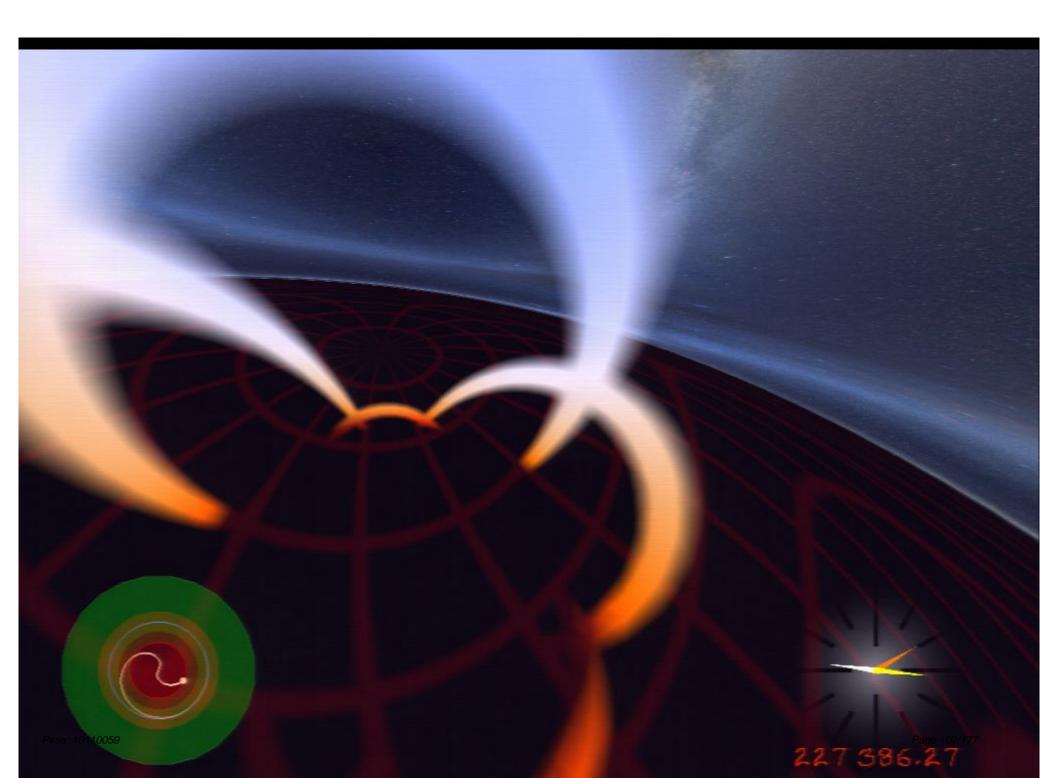


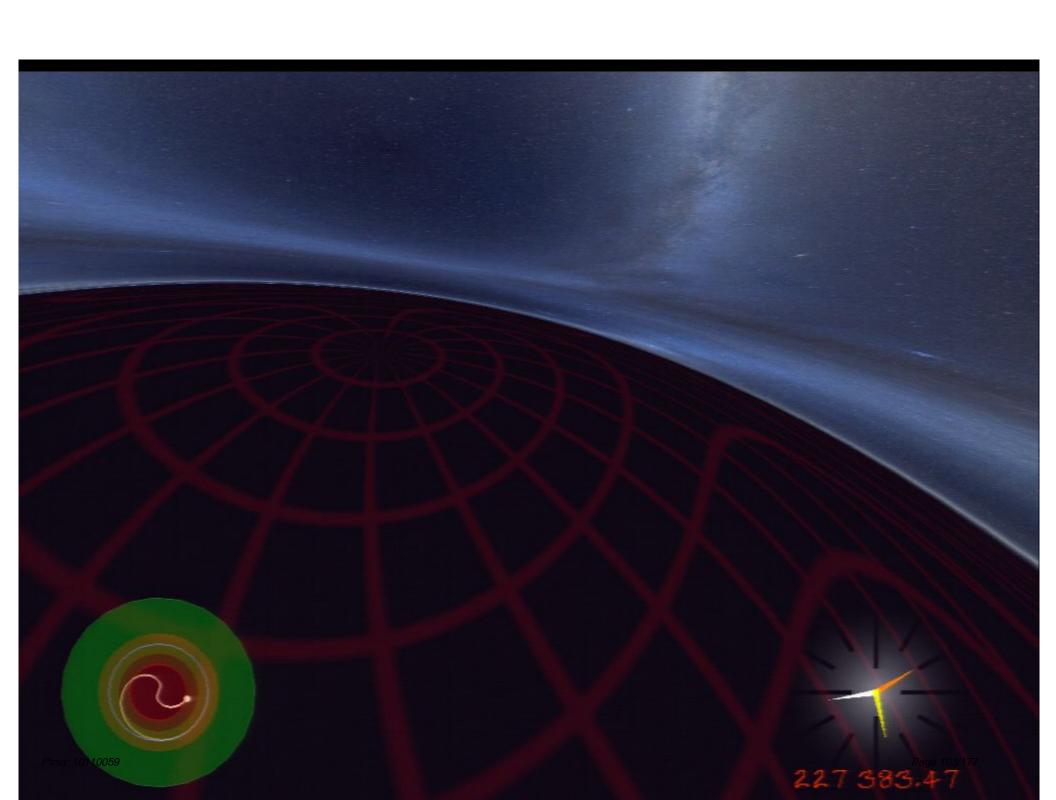


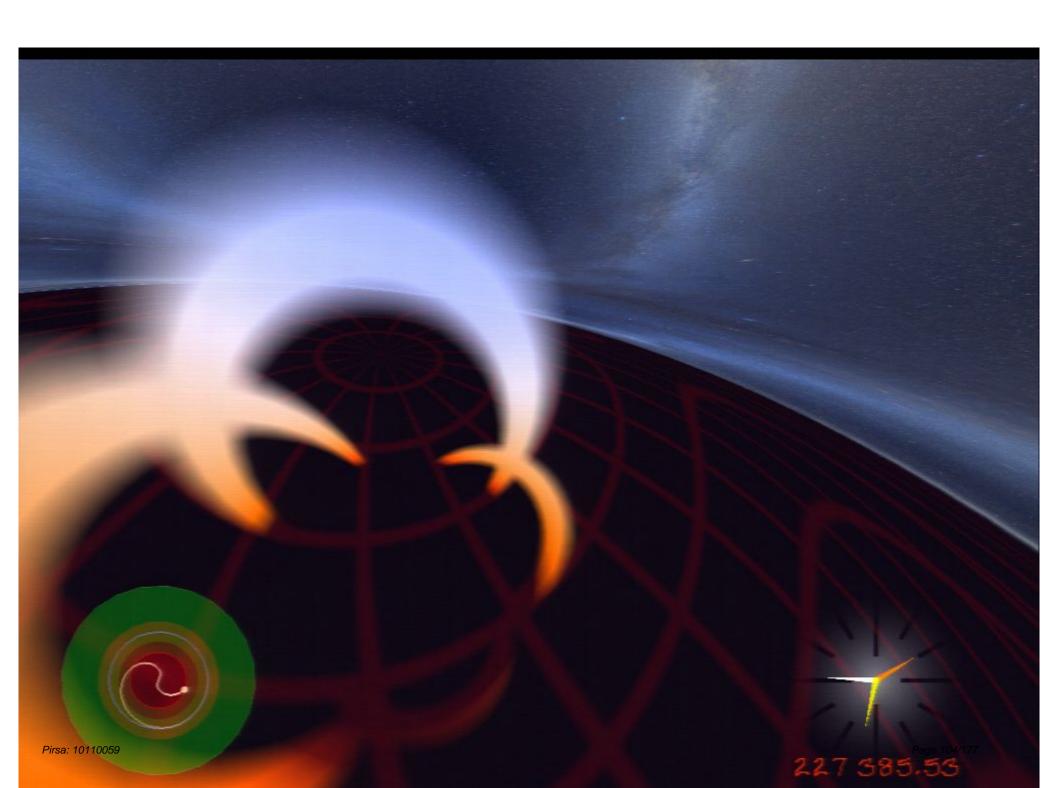


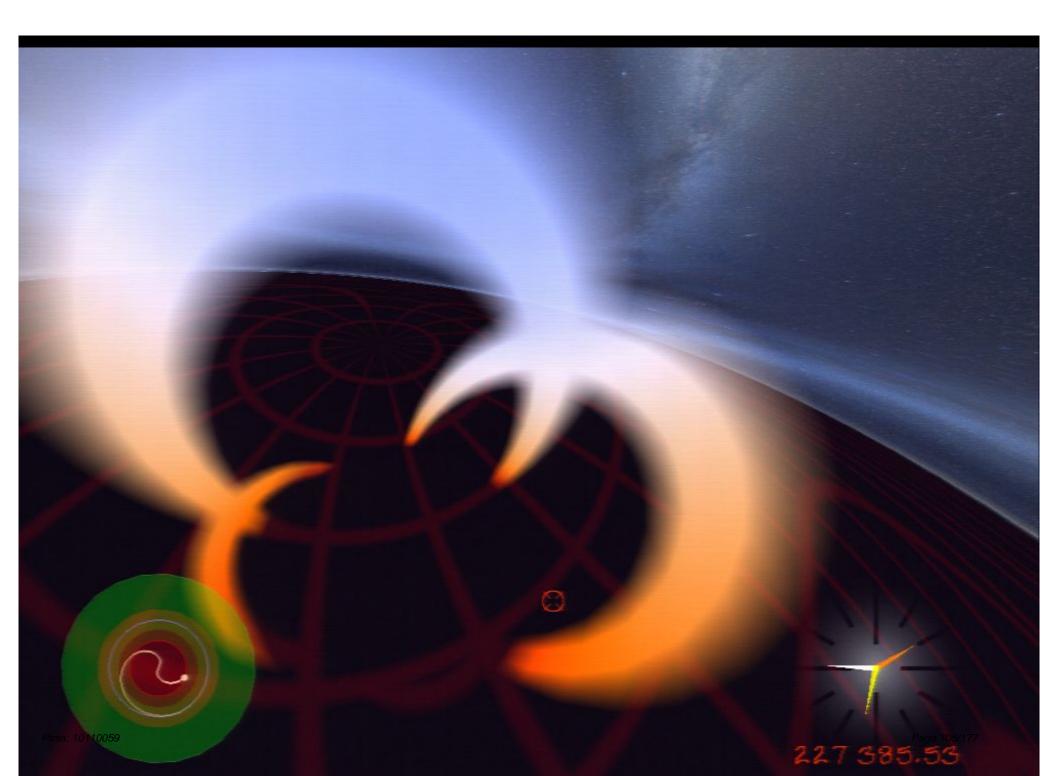


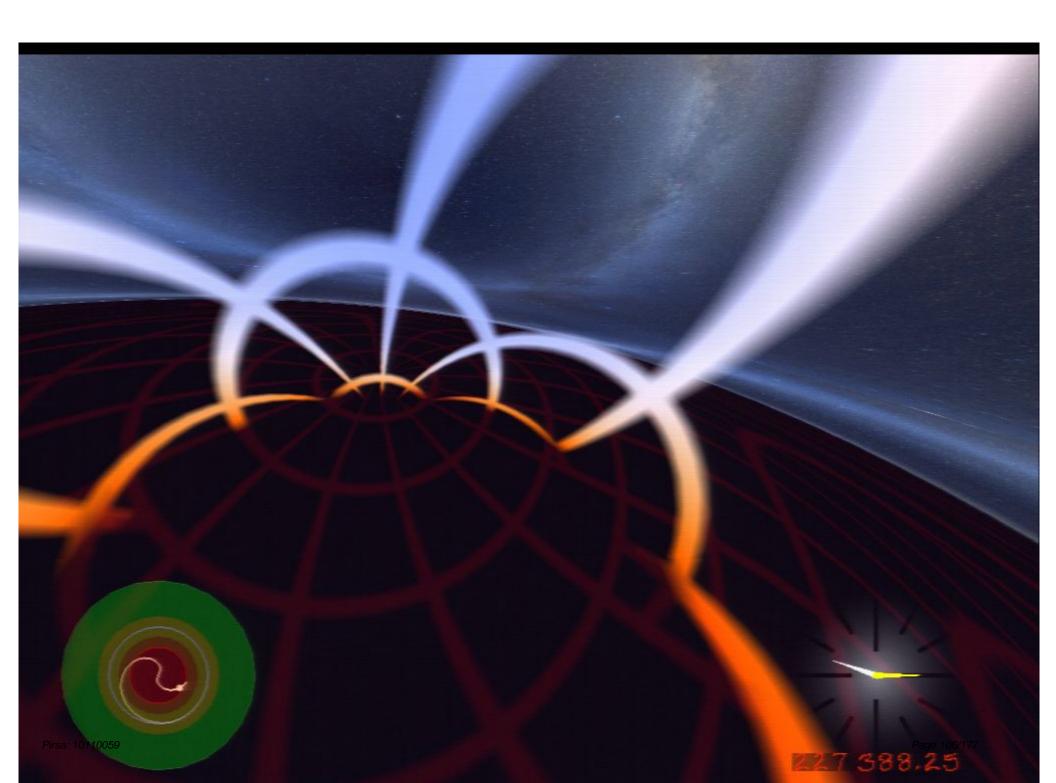


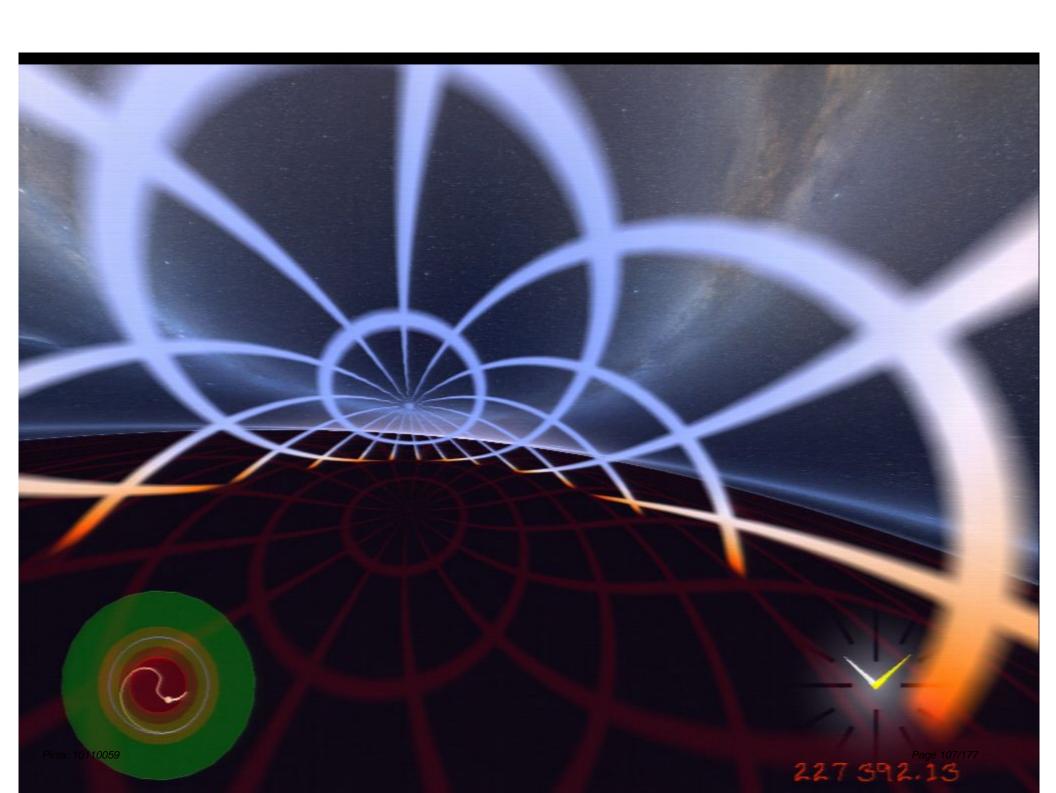


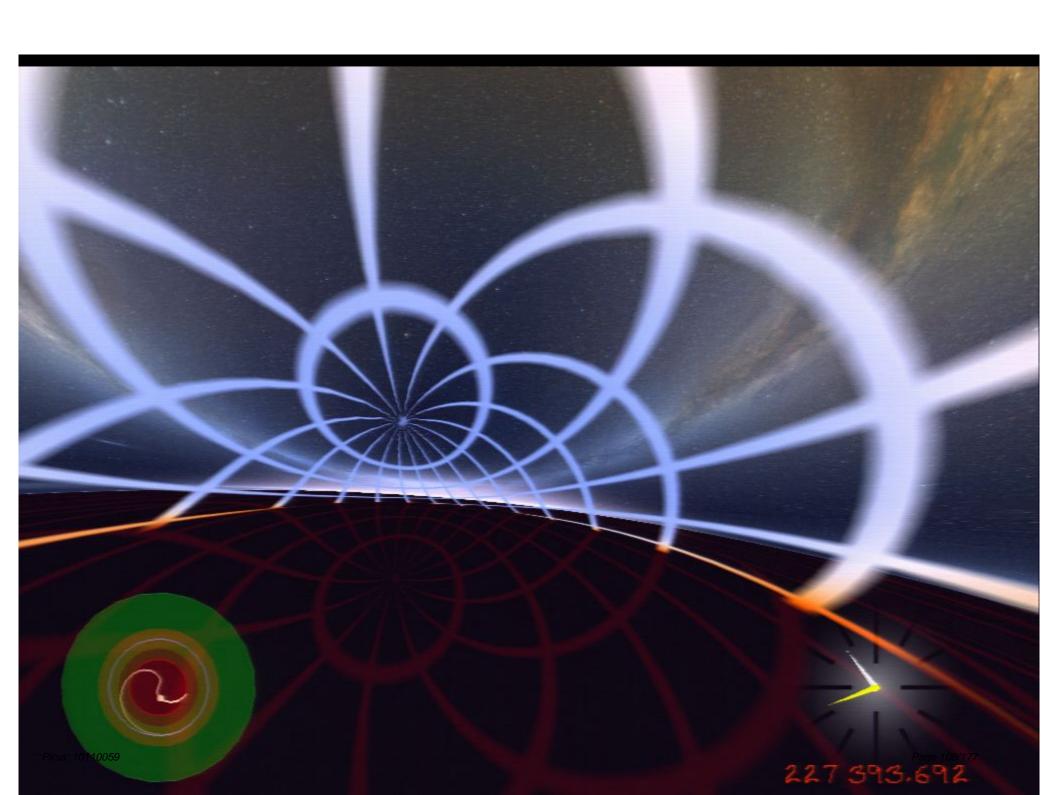


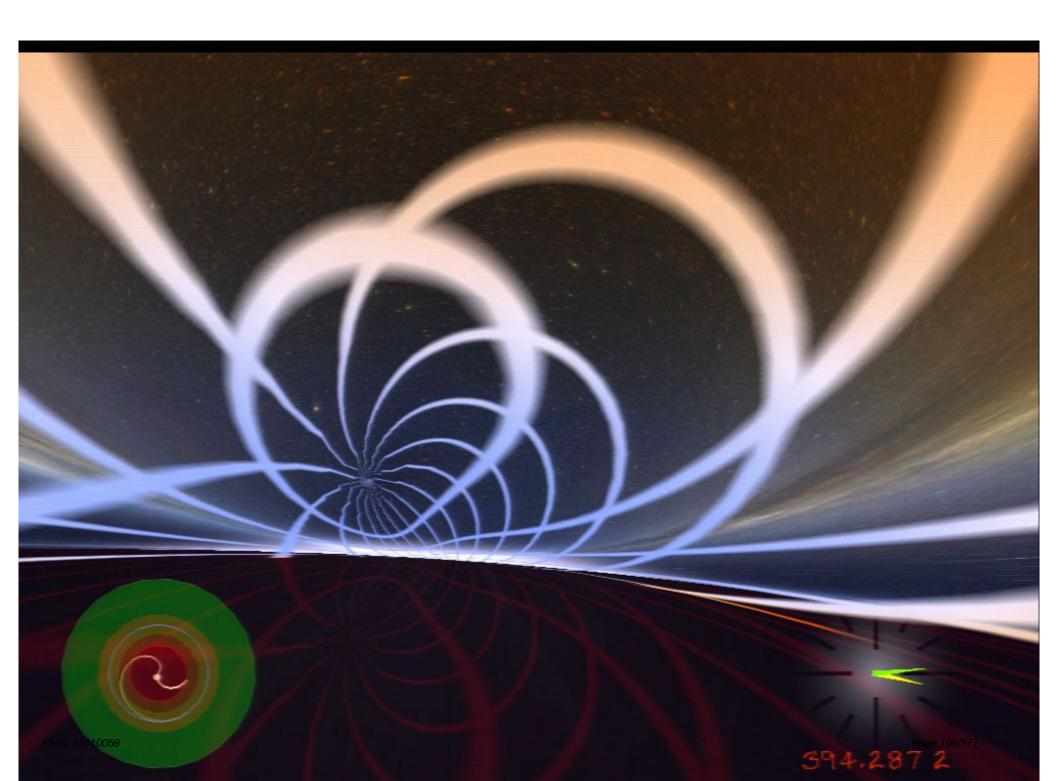


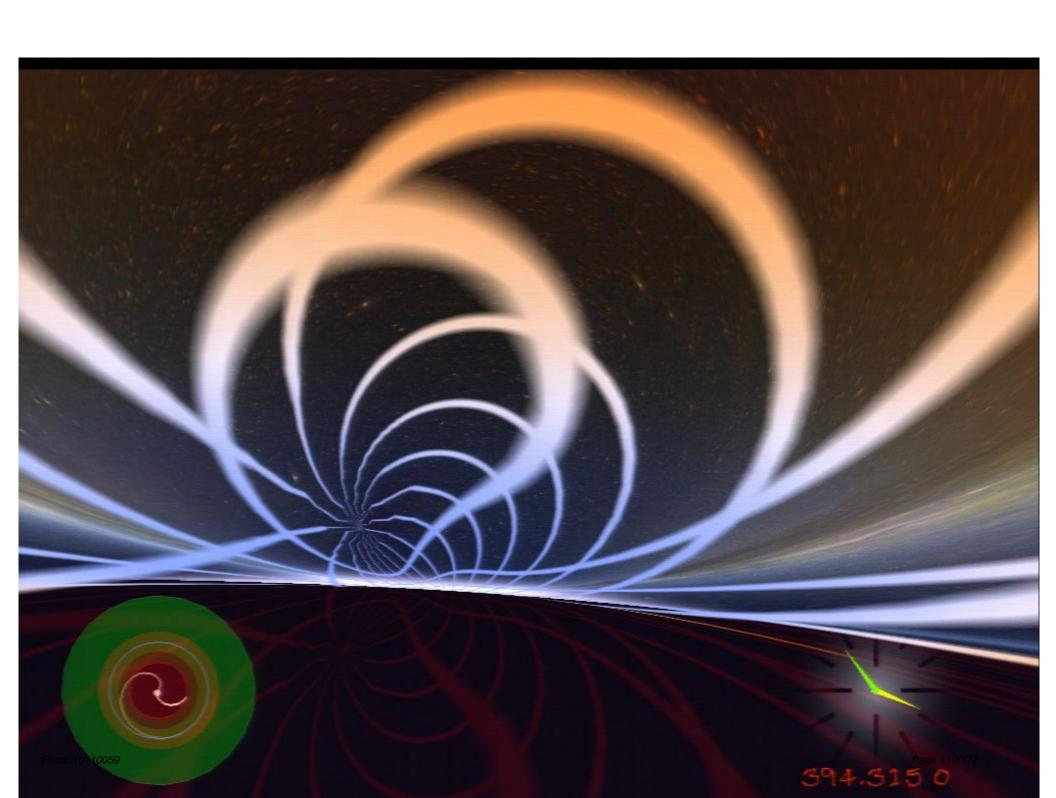


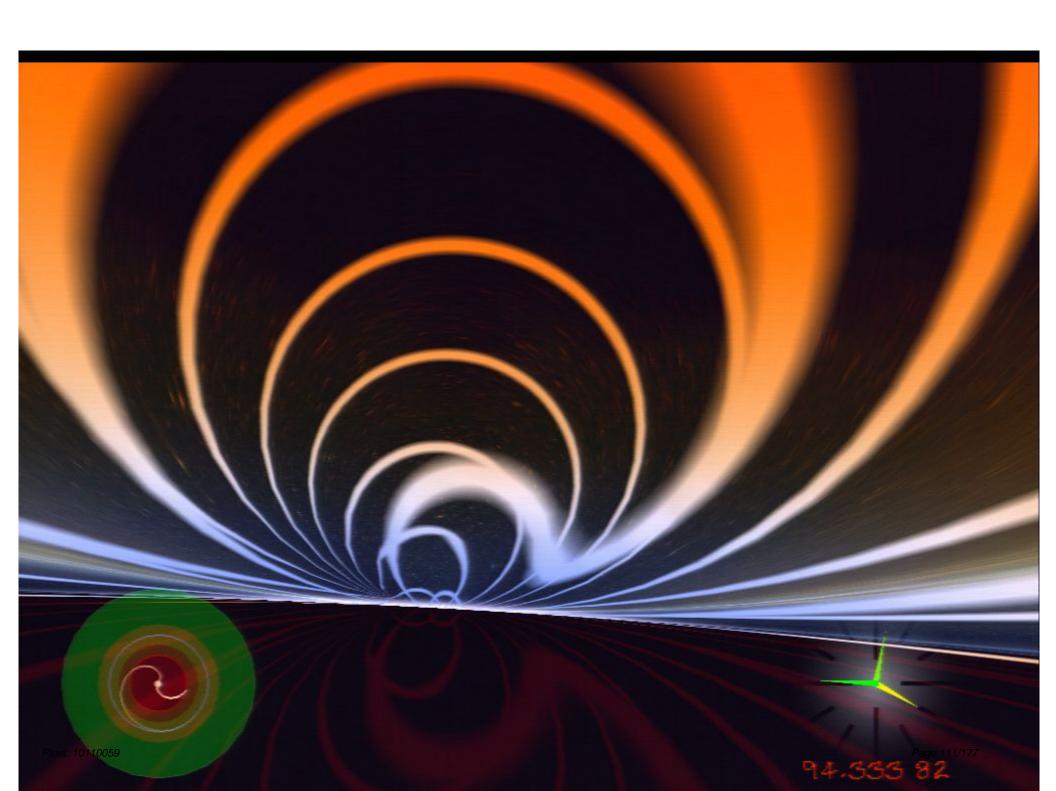


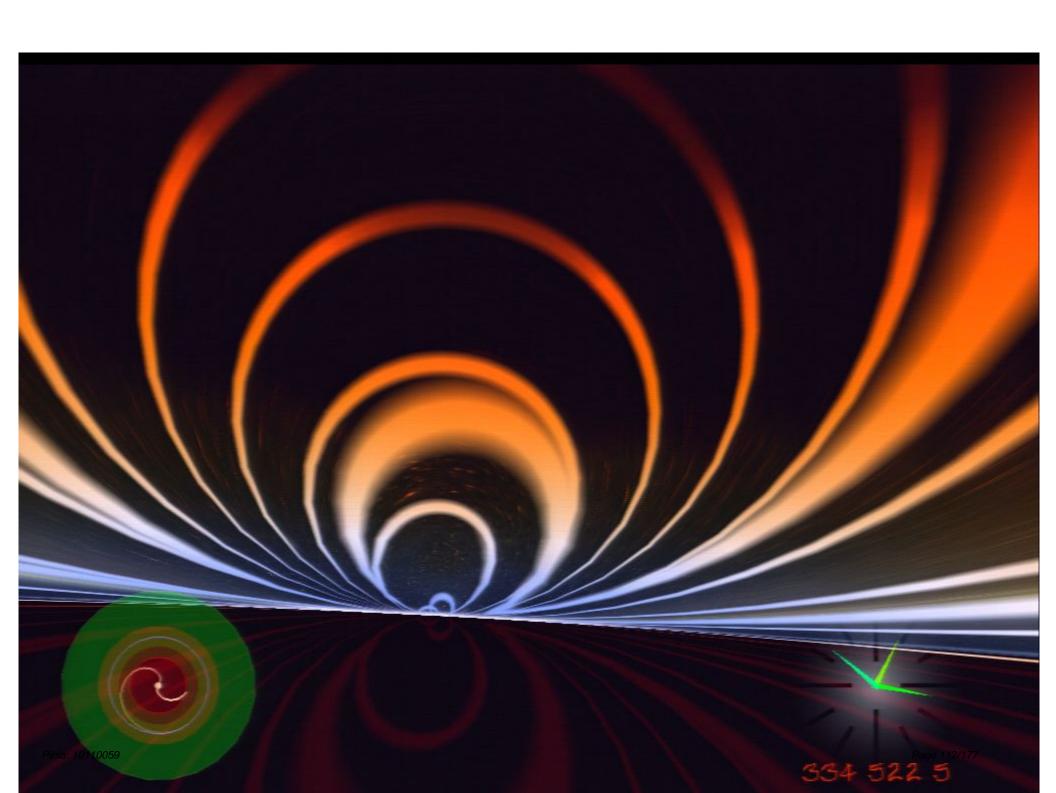


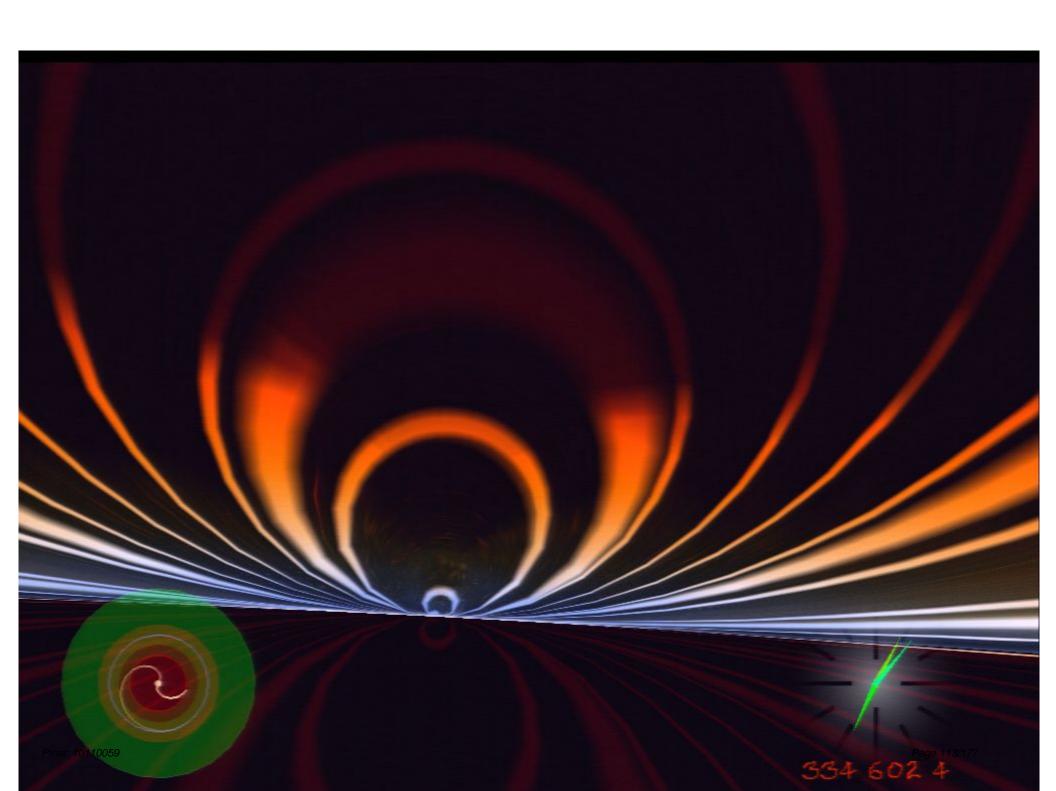


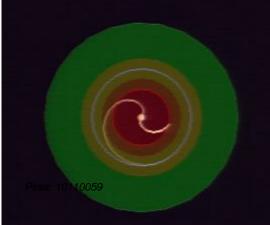




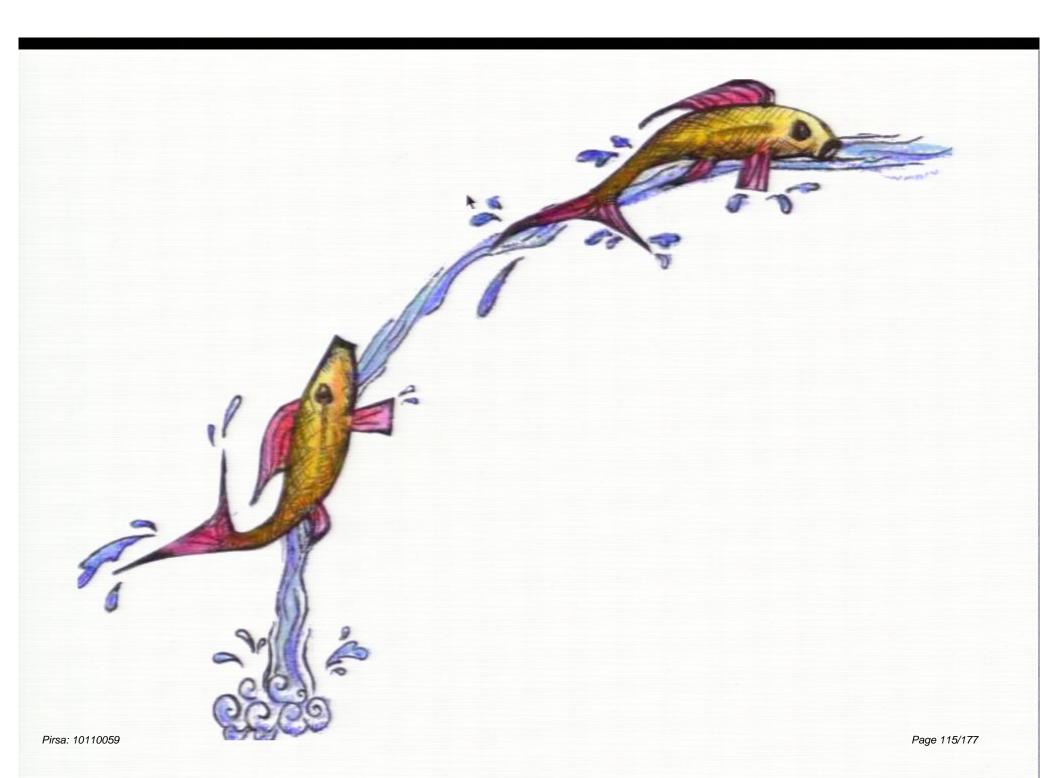




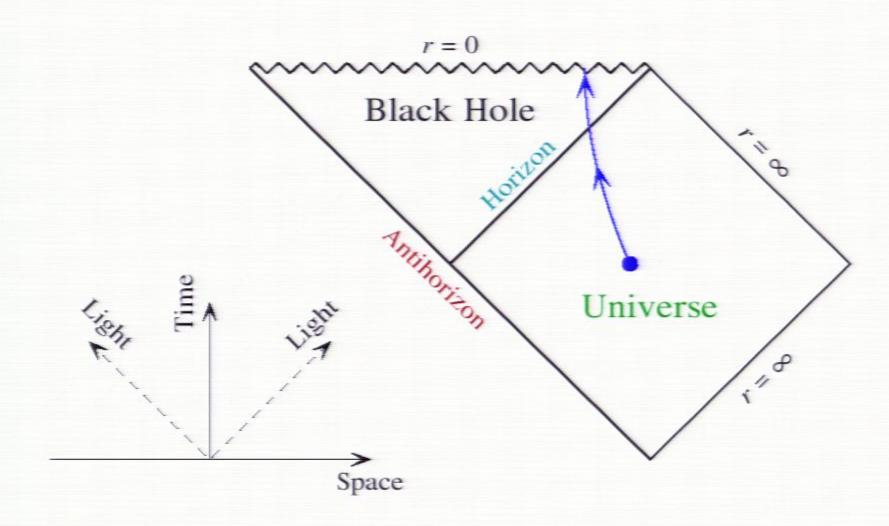






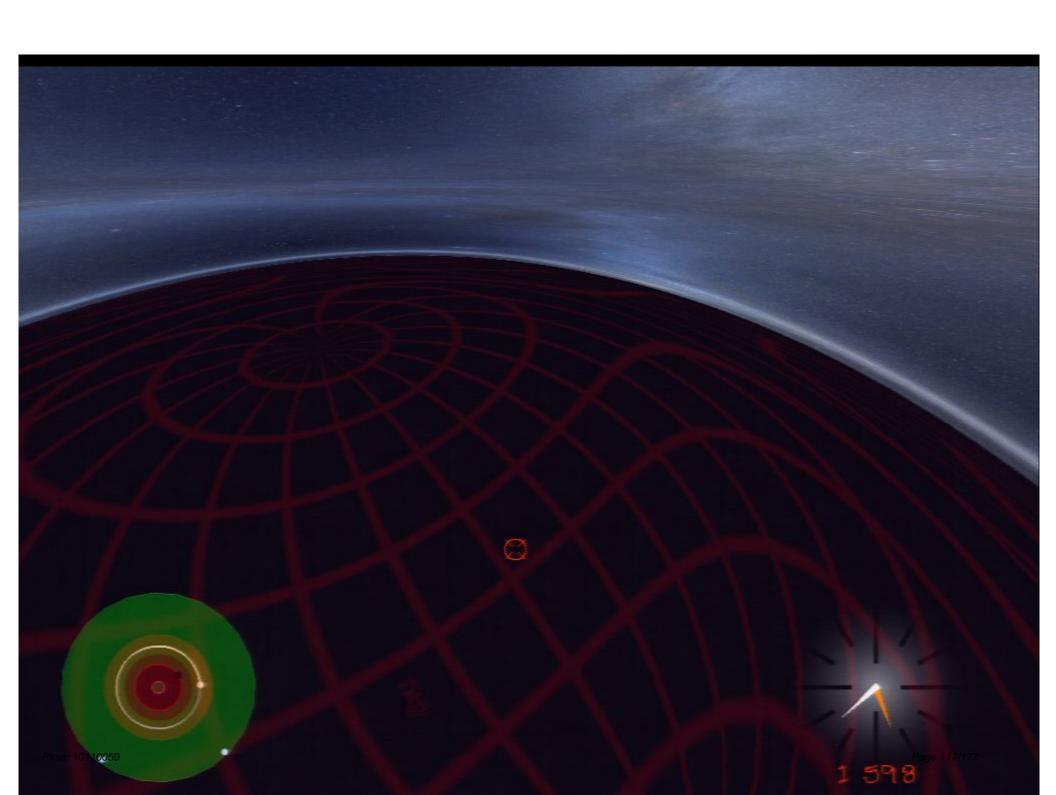


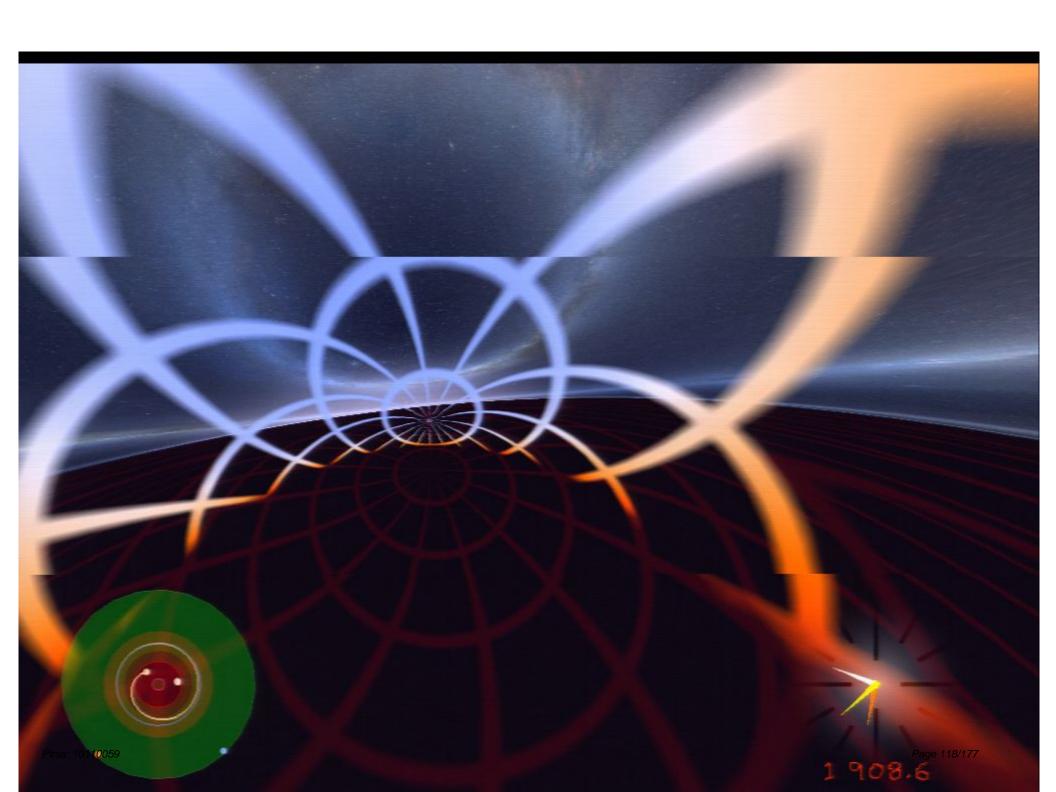
Penrose diagram of a Schwarzschild Black Hole



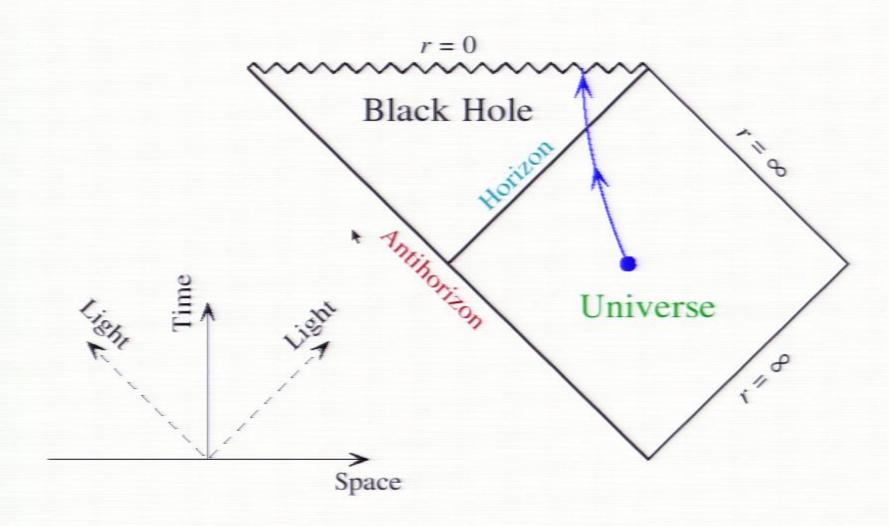
No hair theorem

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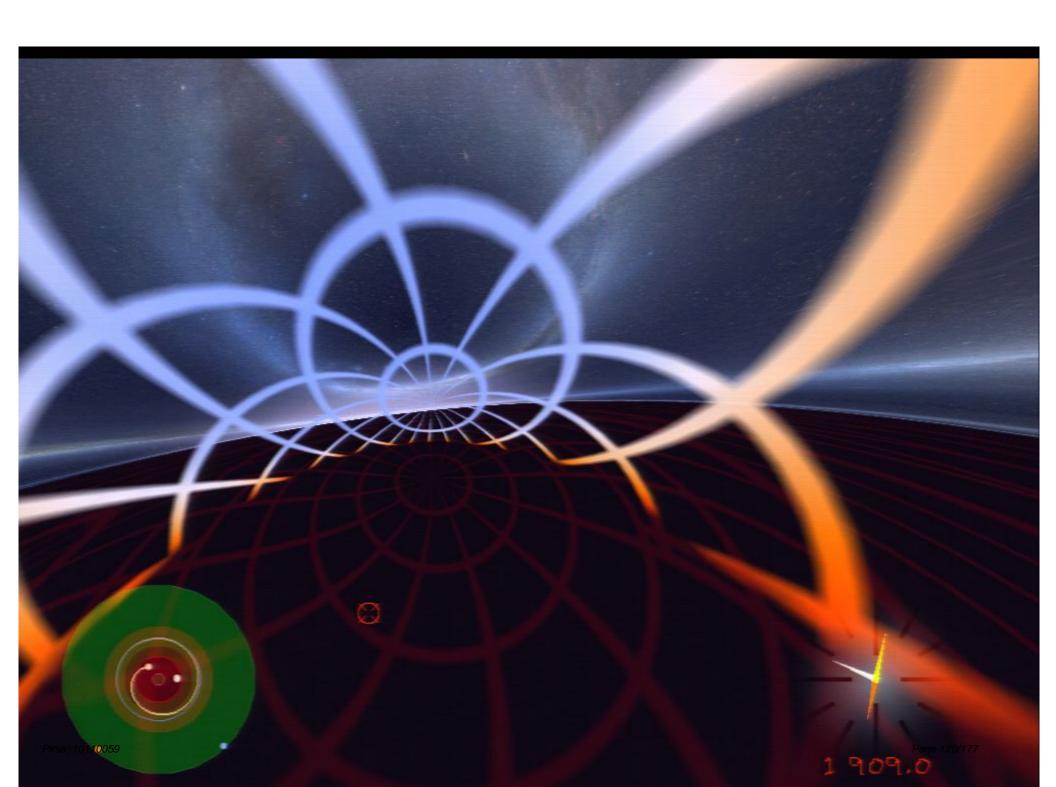


Penrose diagram of a Schwarzschild Black Hole

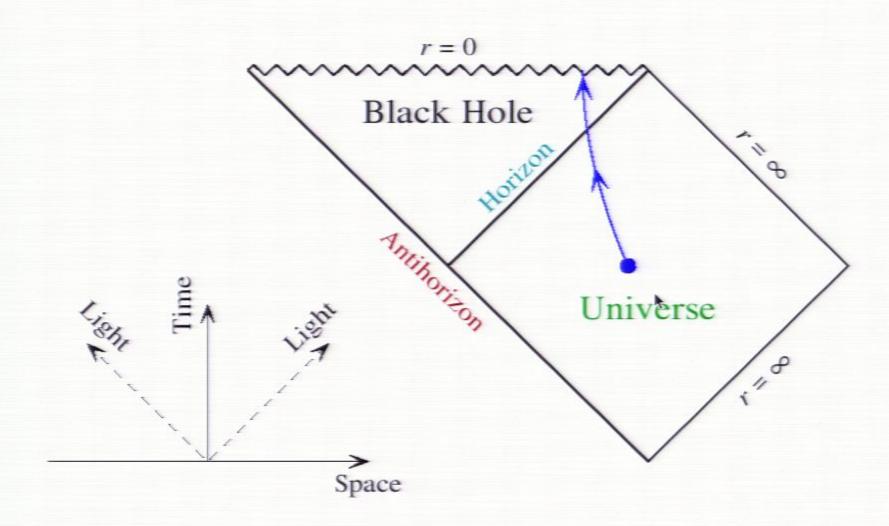


No hair theorem

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Penrose diagram of a Schwarzschild Black Hole



No hair theorem

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To hair theorem

Outside its horizon, an isolated black hole rapidly (on few light crossing imes) evolves to a state characterized by just 3 properties:

- Mass
- Electric charge
- Spin

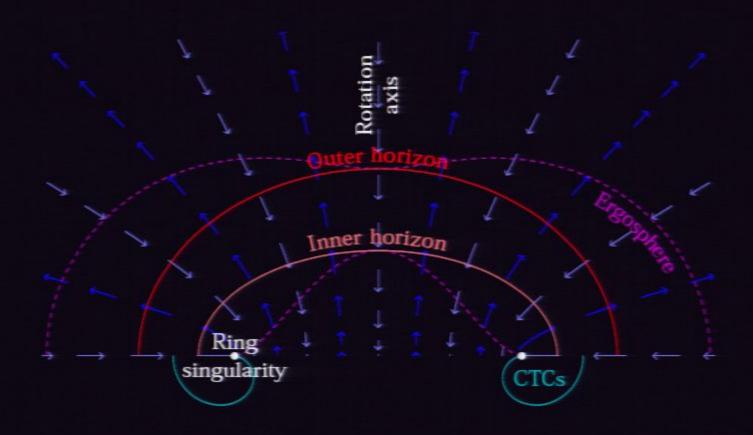
t loses its "hair" by gravitational radiation.

■ Kerr (rotating) black hole

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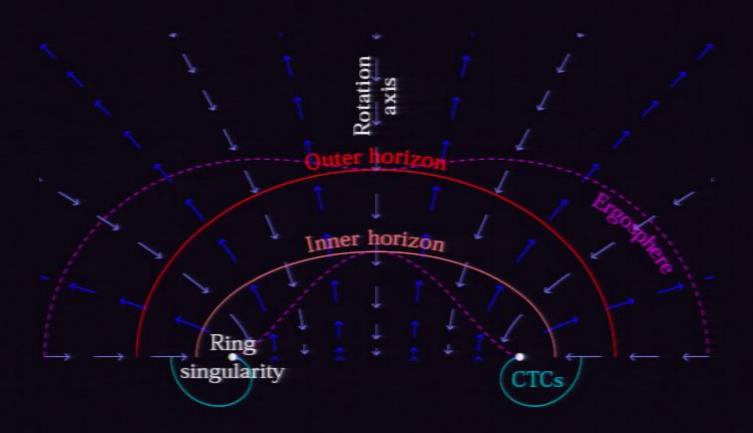
Kerr (rotating) black hole



Reissner-Nordström (charged) black hole

B

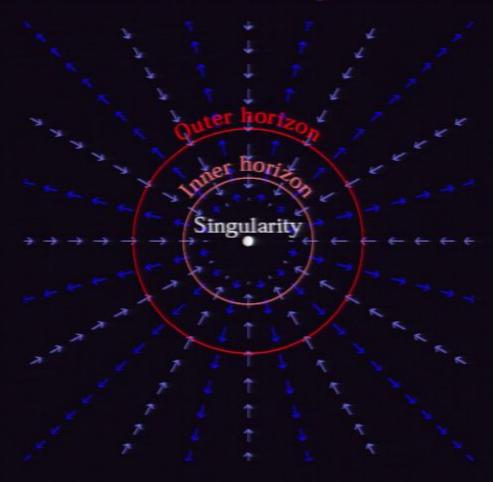
Kerr (rotating) black hole



Reissner-Nordström (charged) black hole

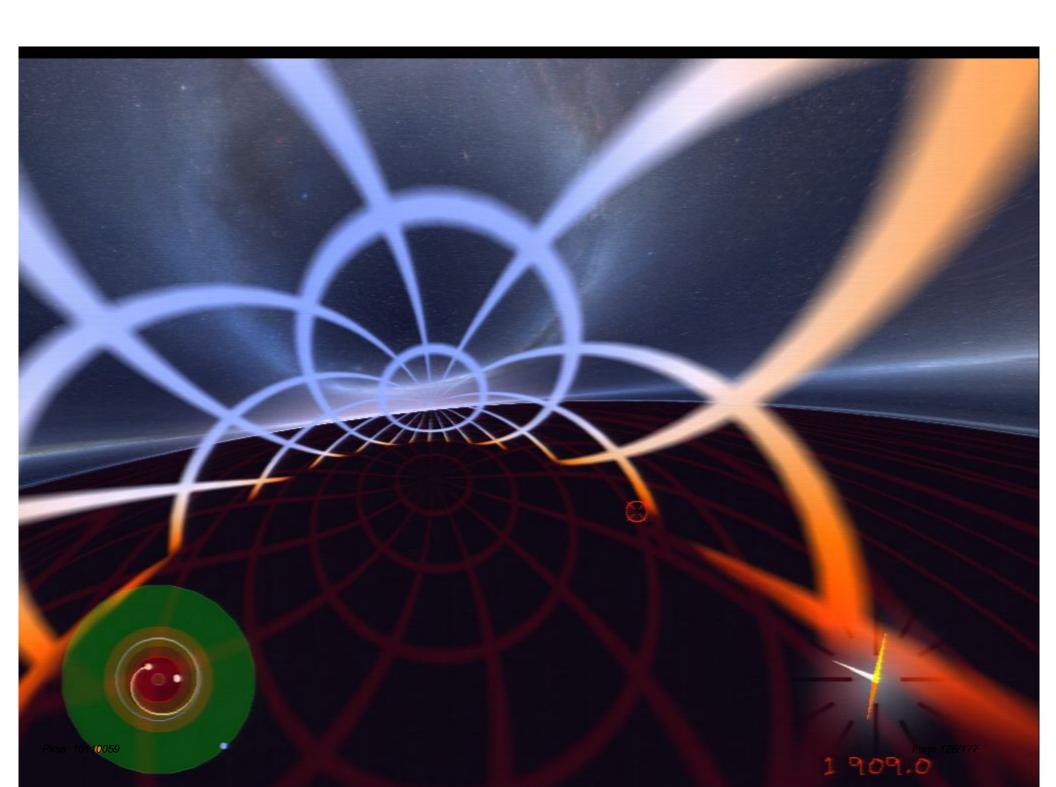
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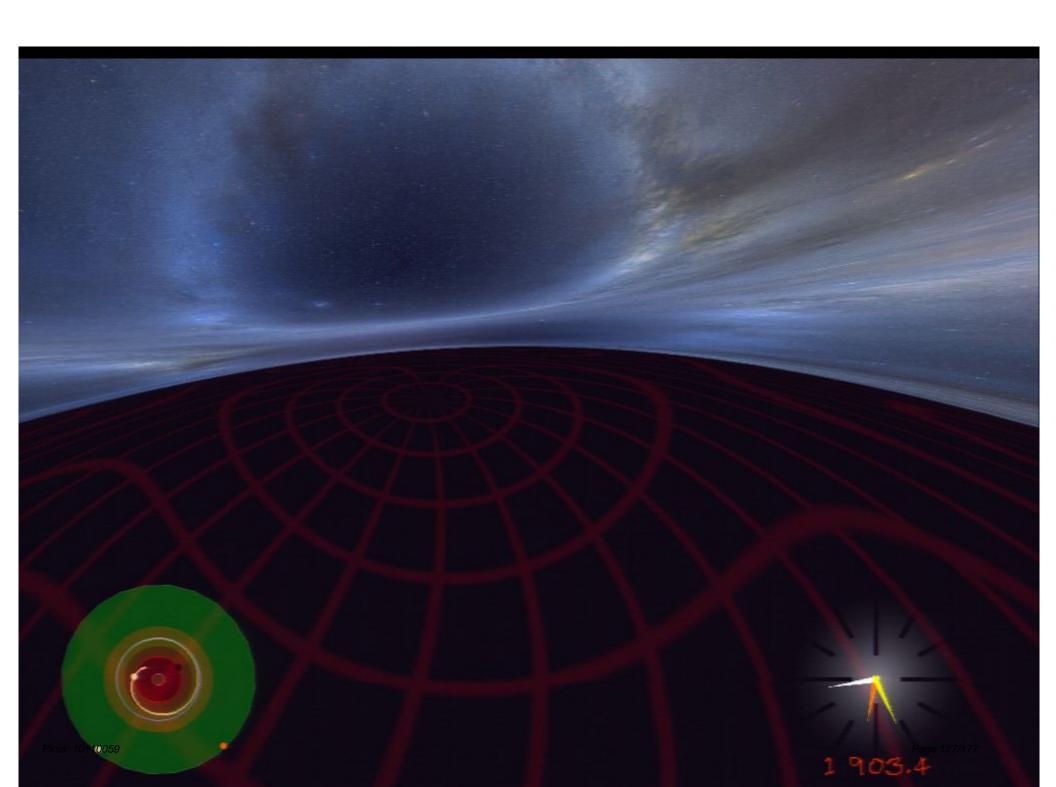
Reissner-Nordström (charged) black hole

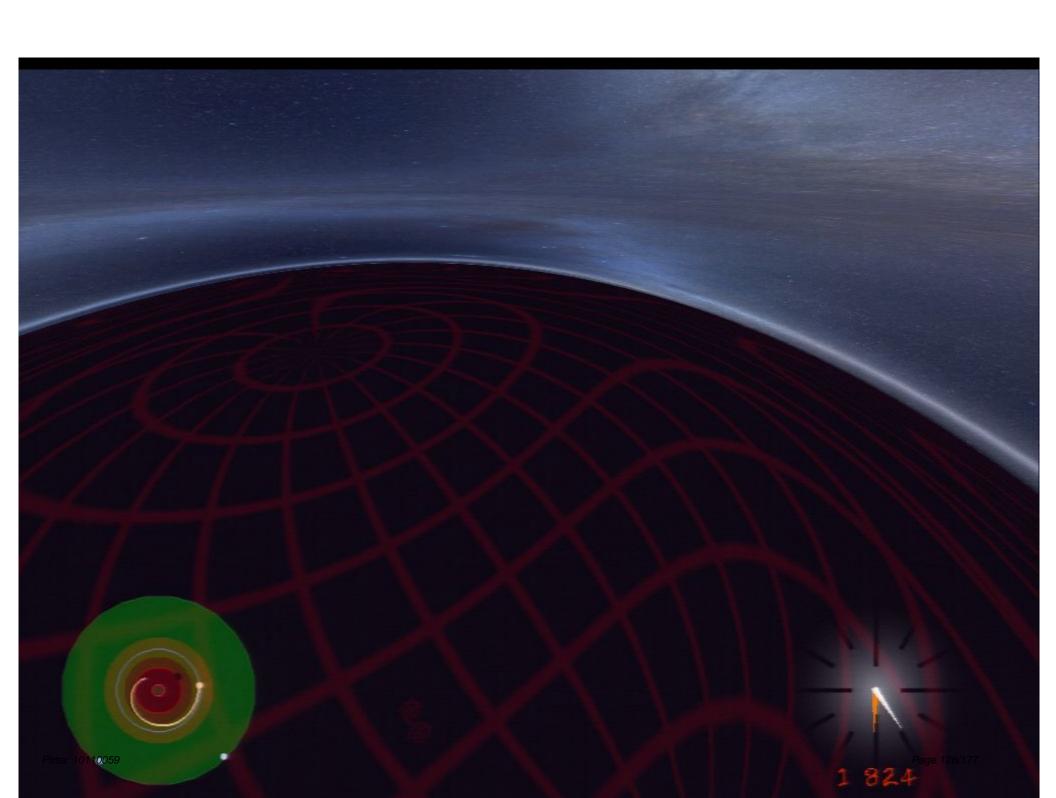


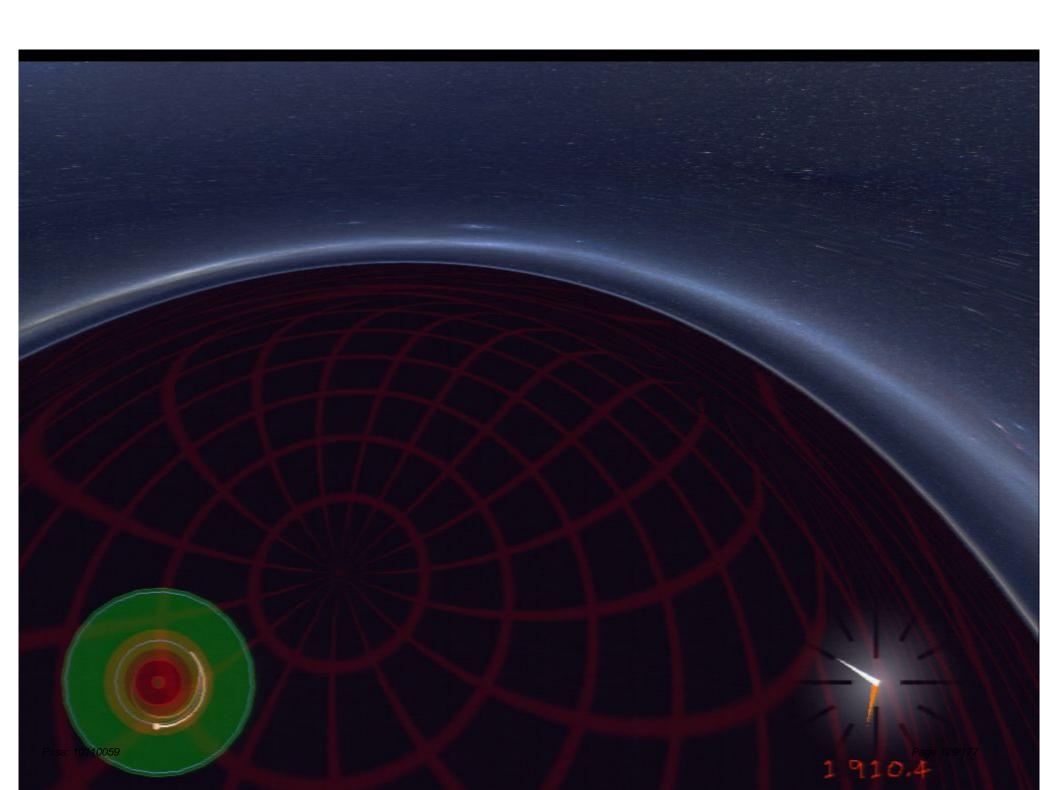
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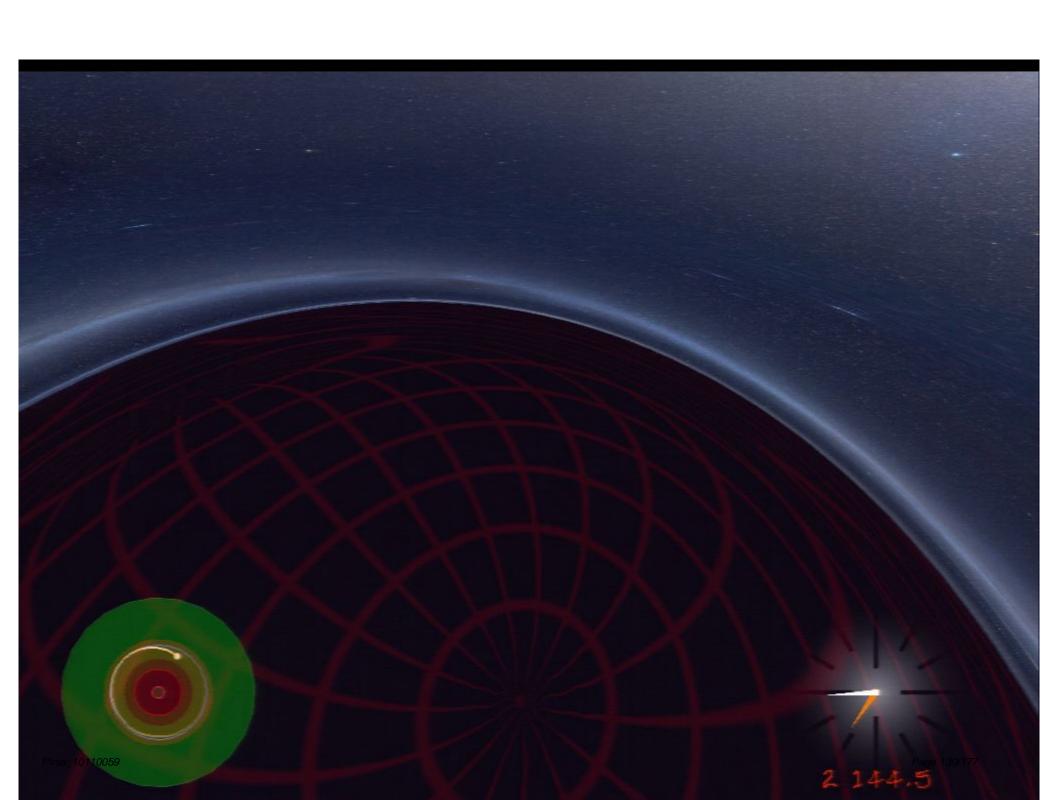
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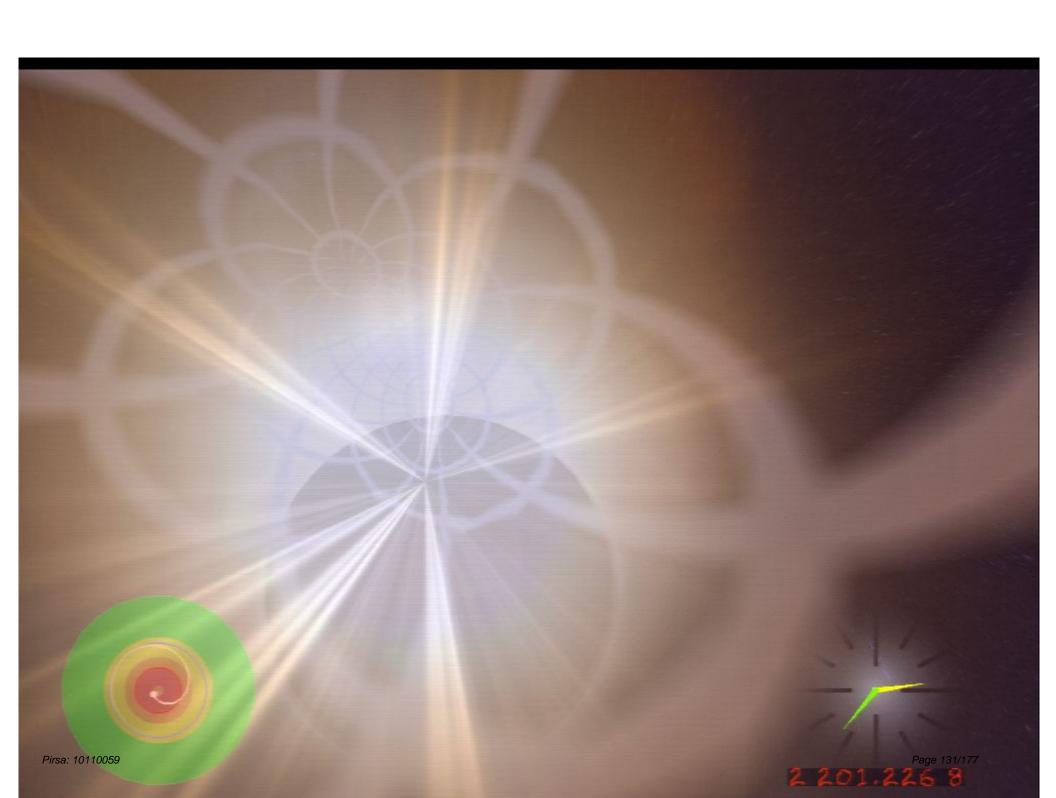


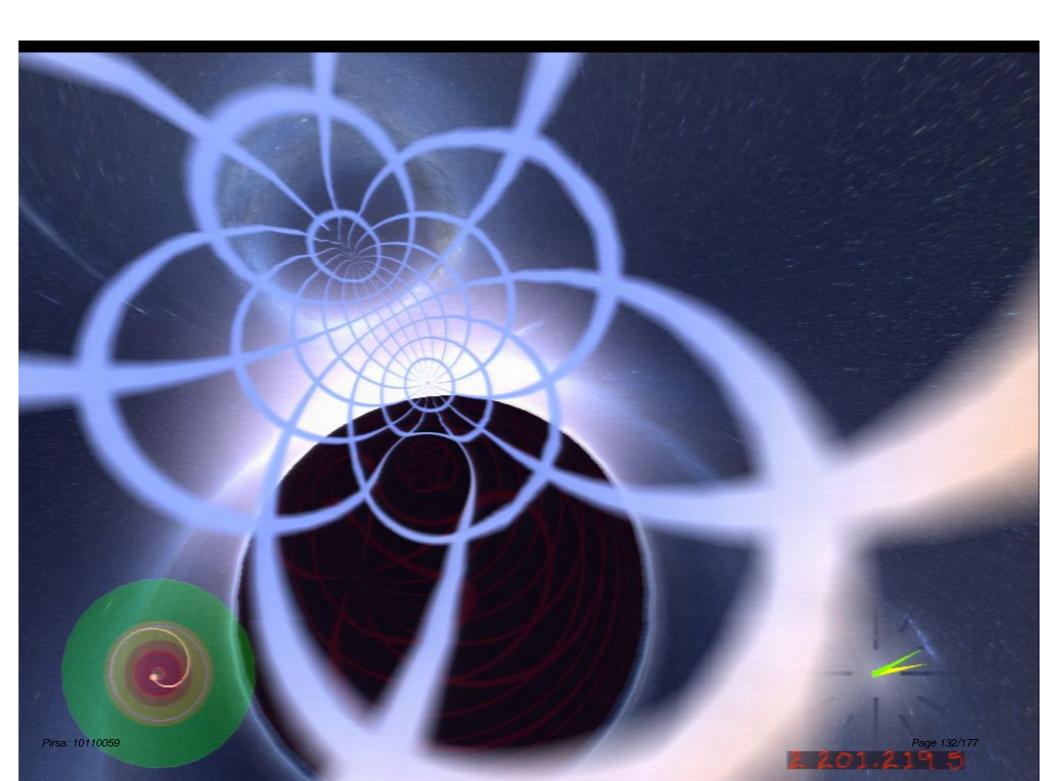


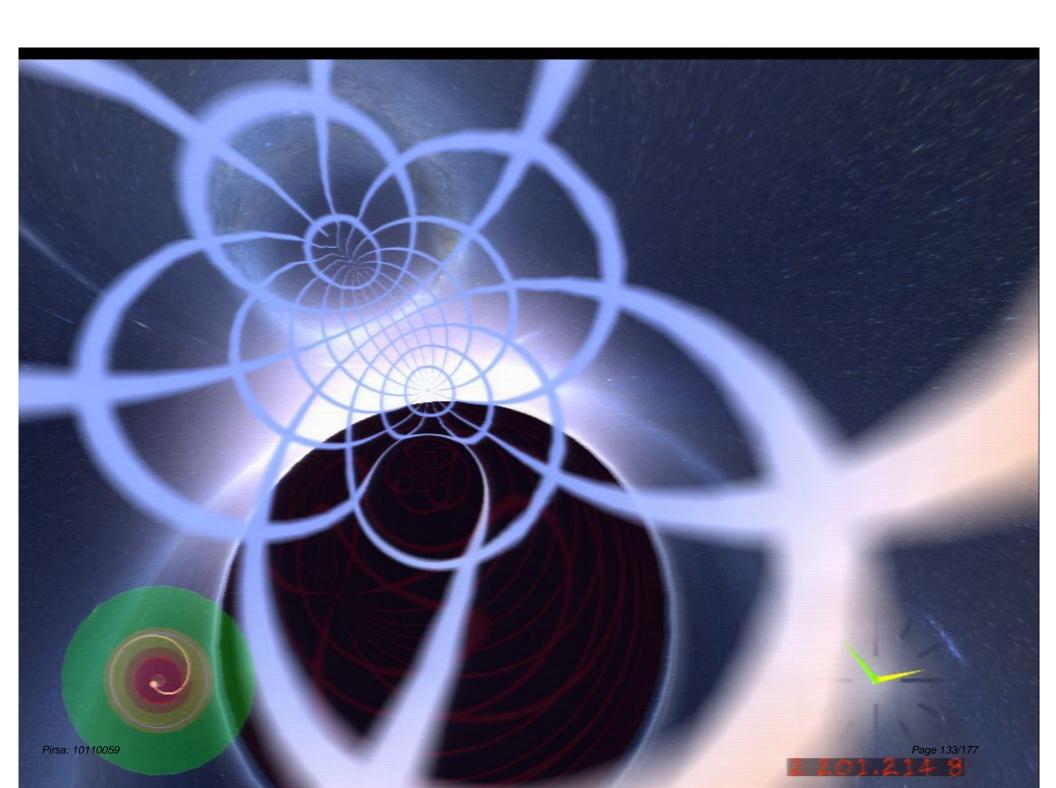


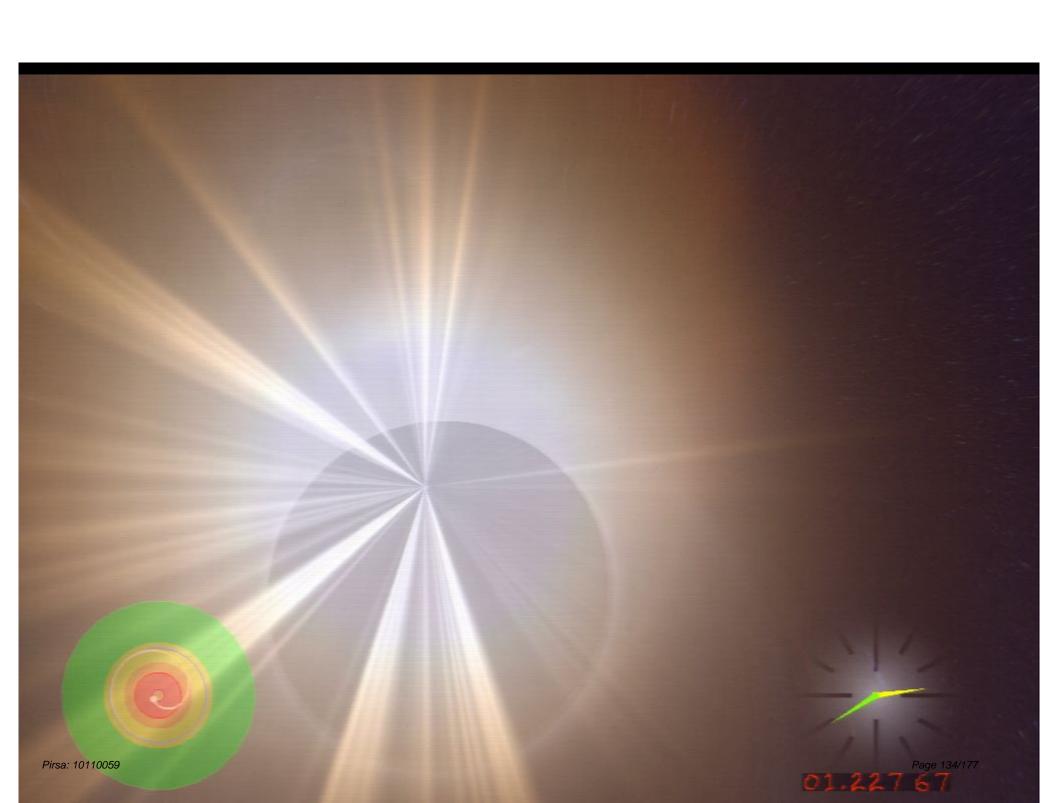




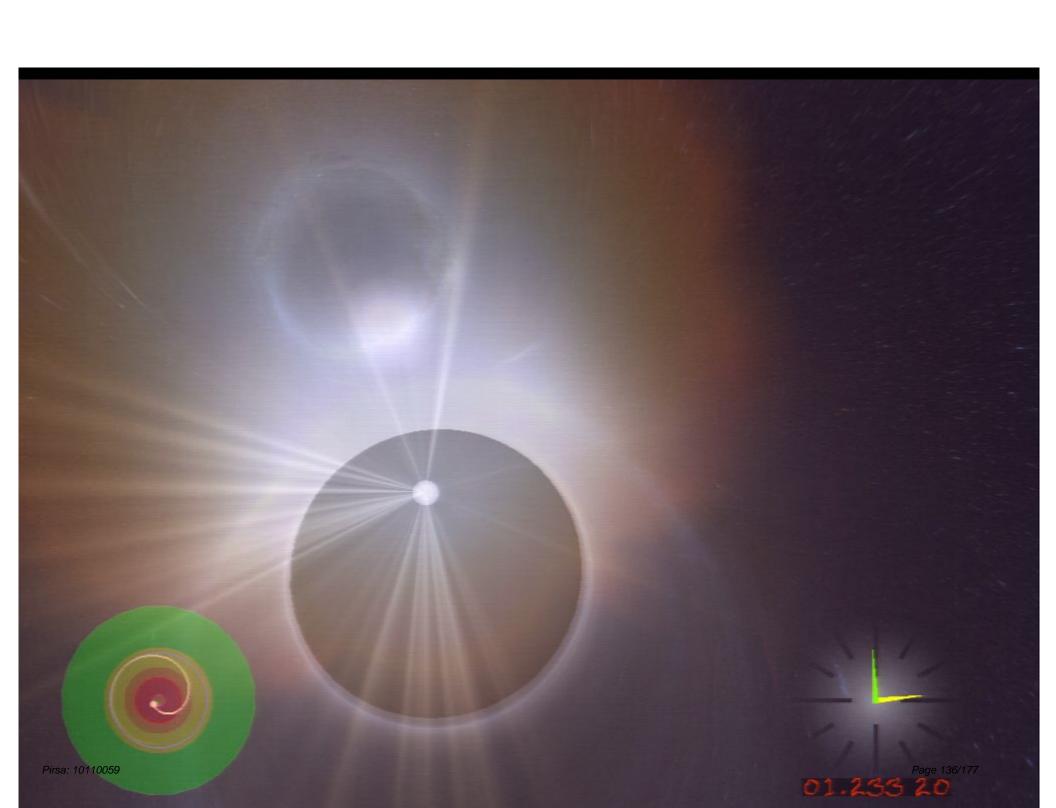


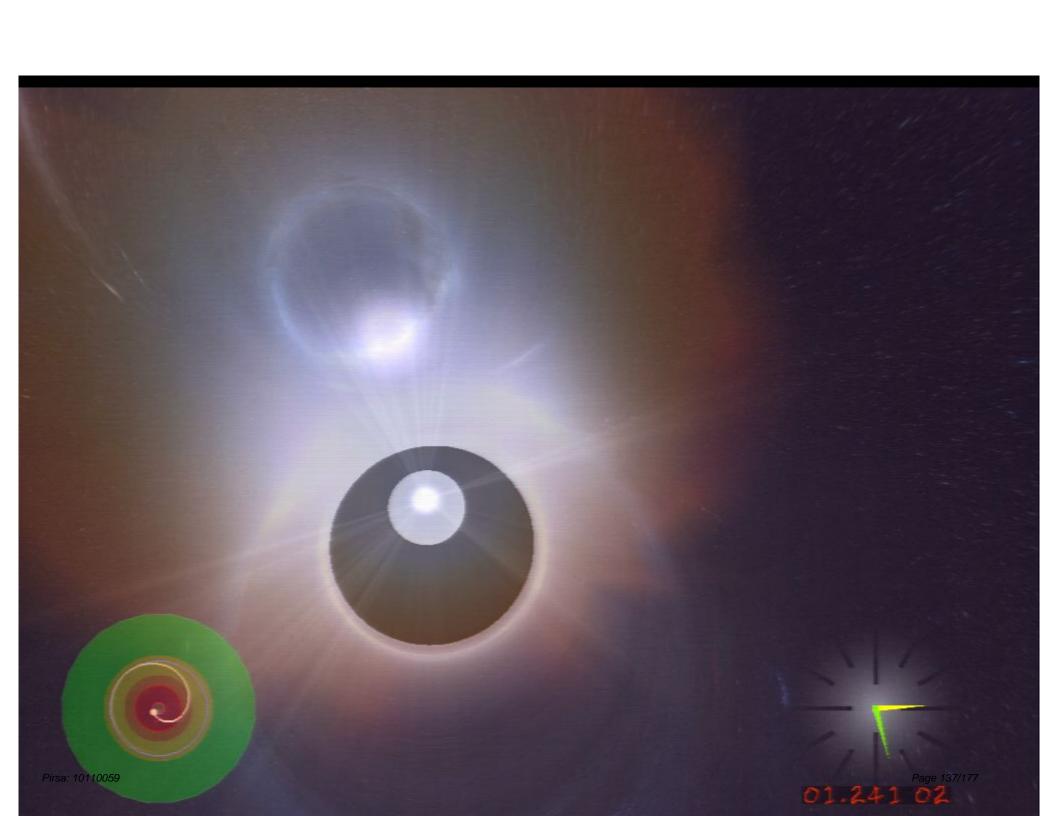


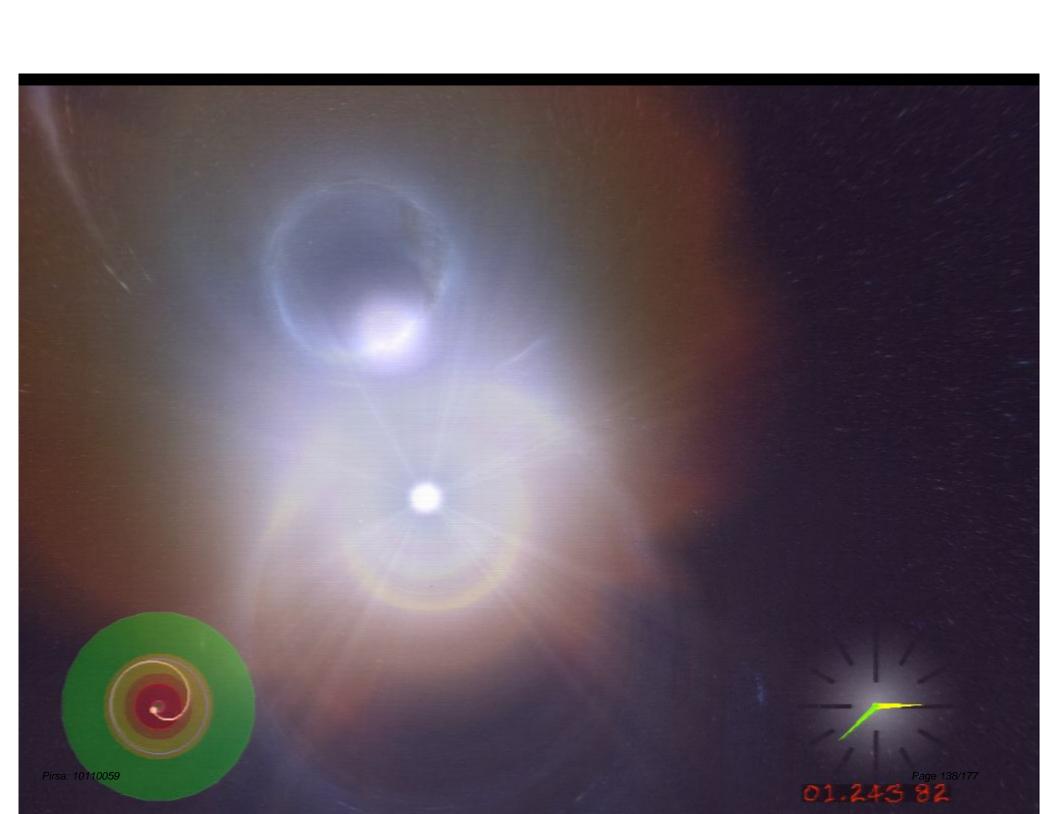


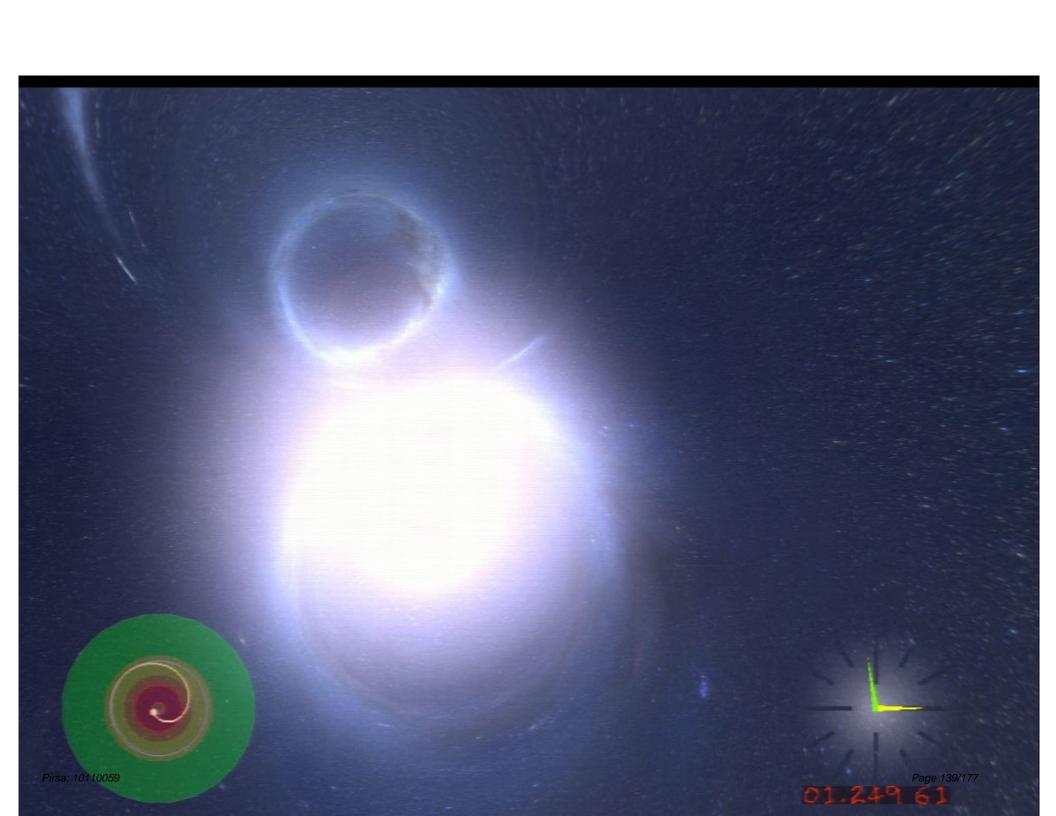




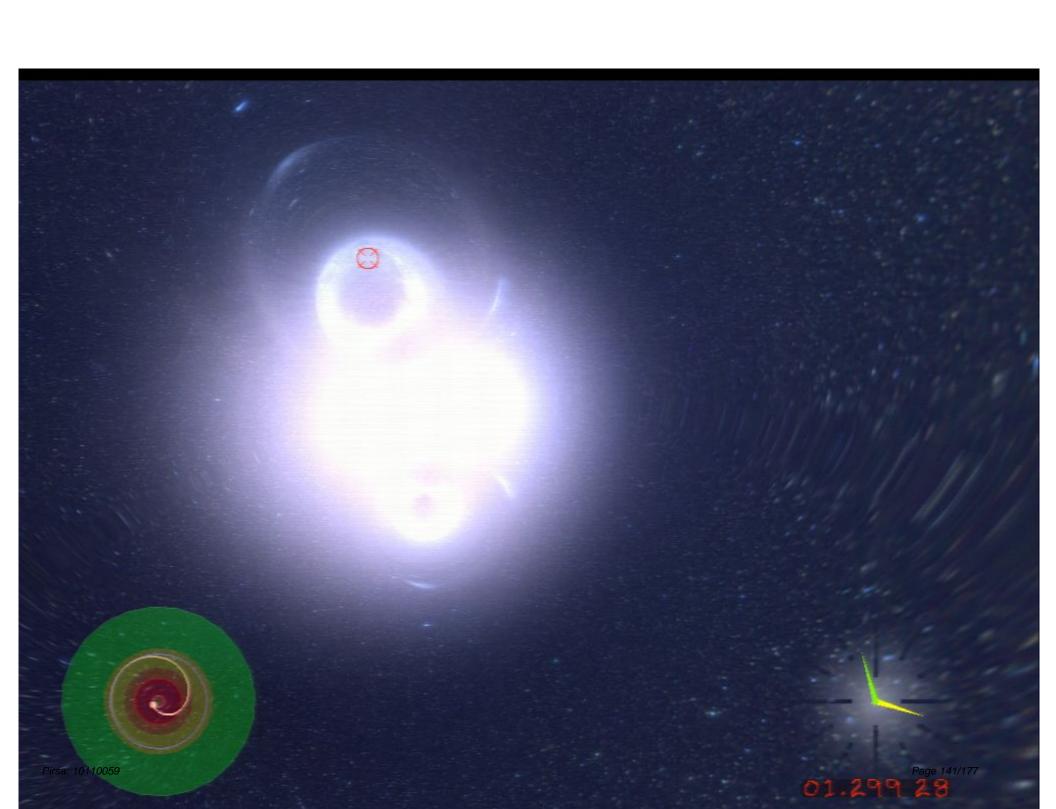


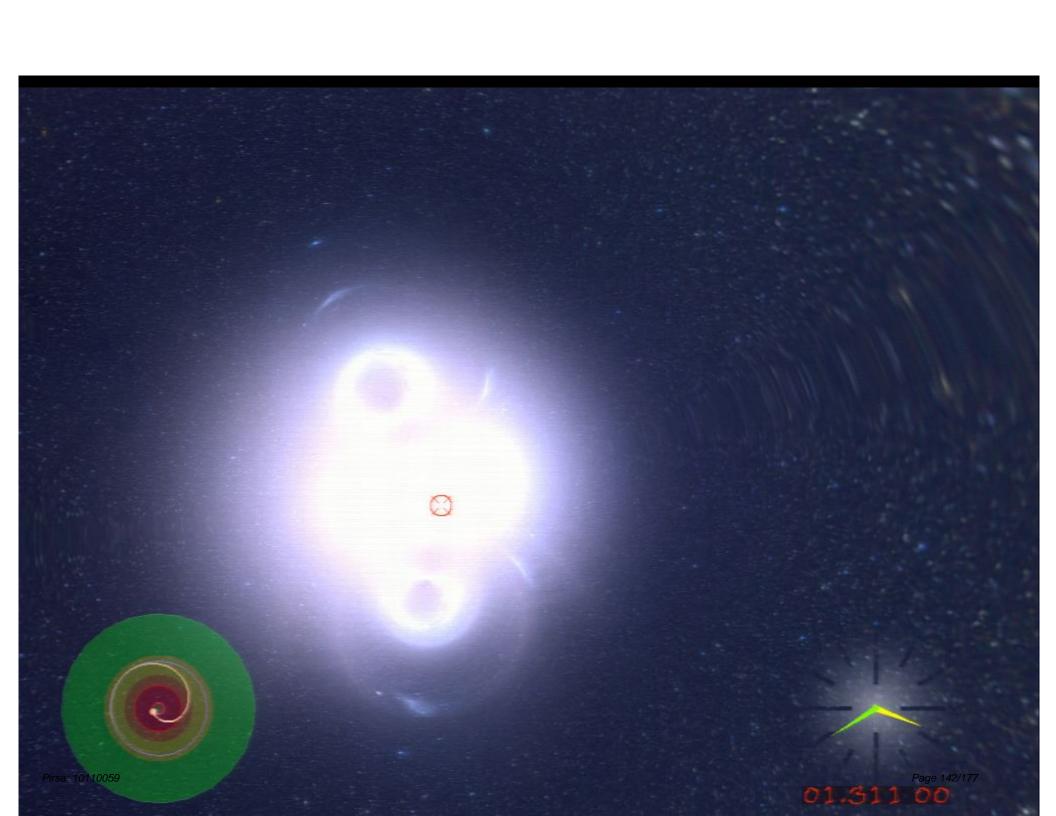


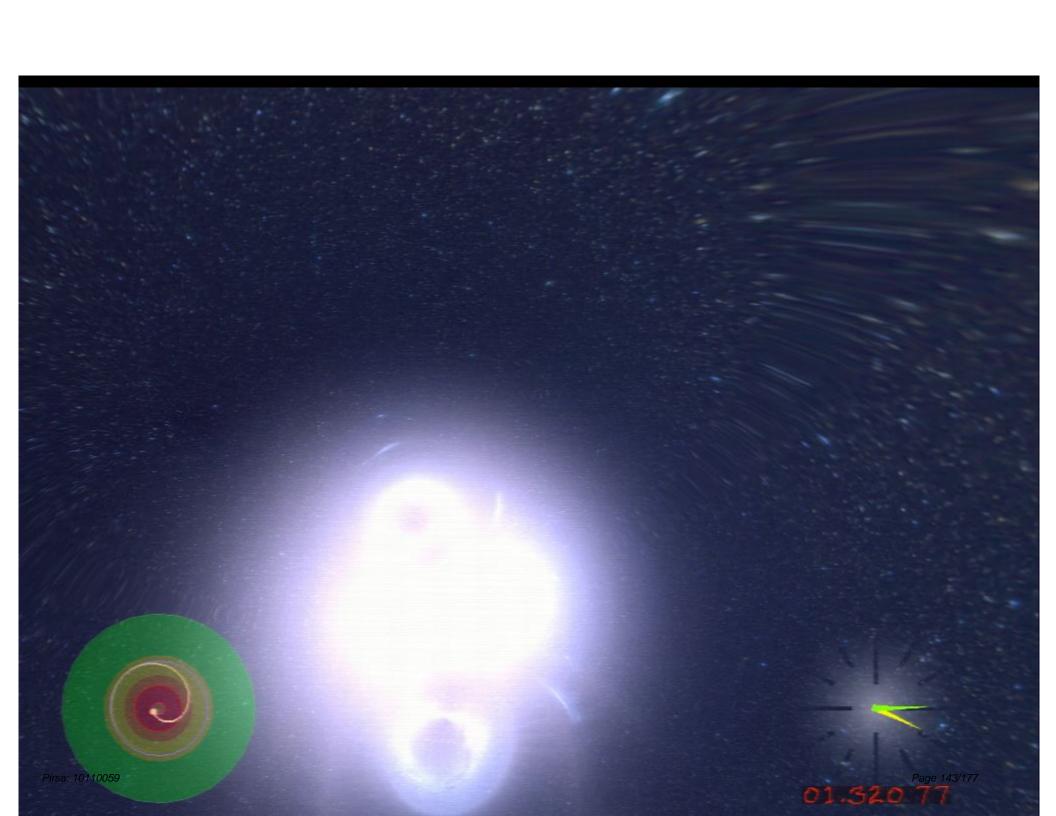








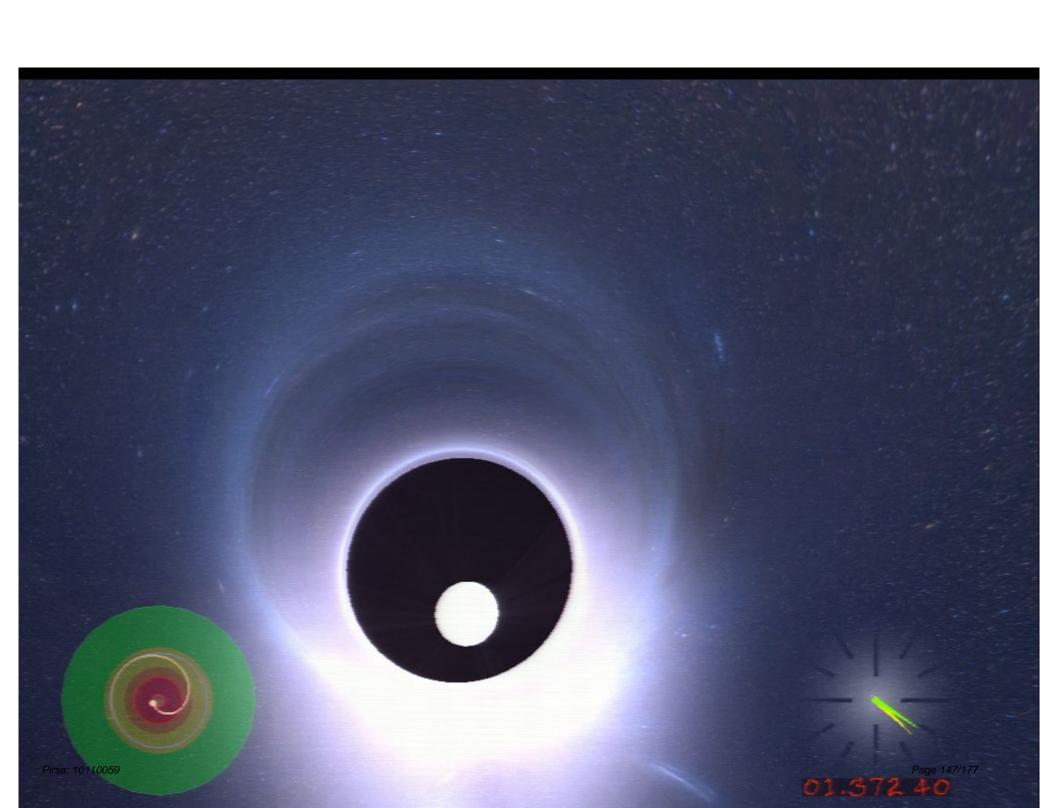


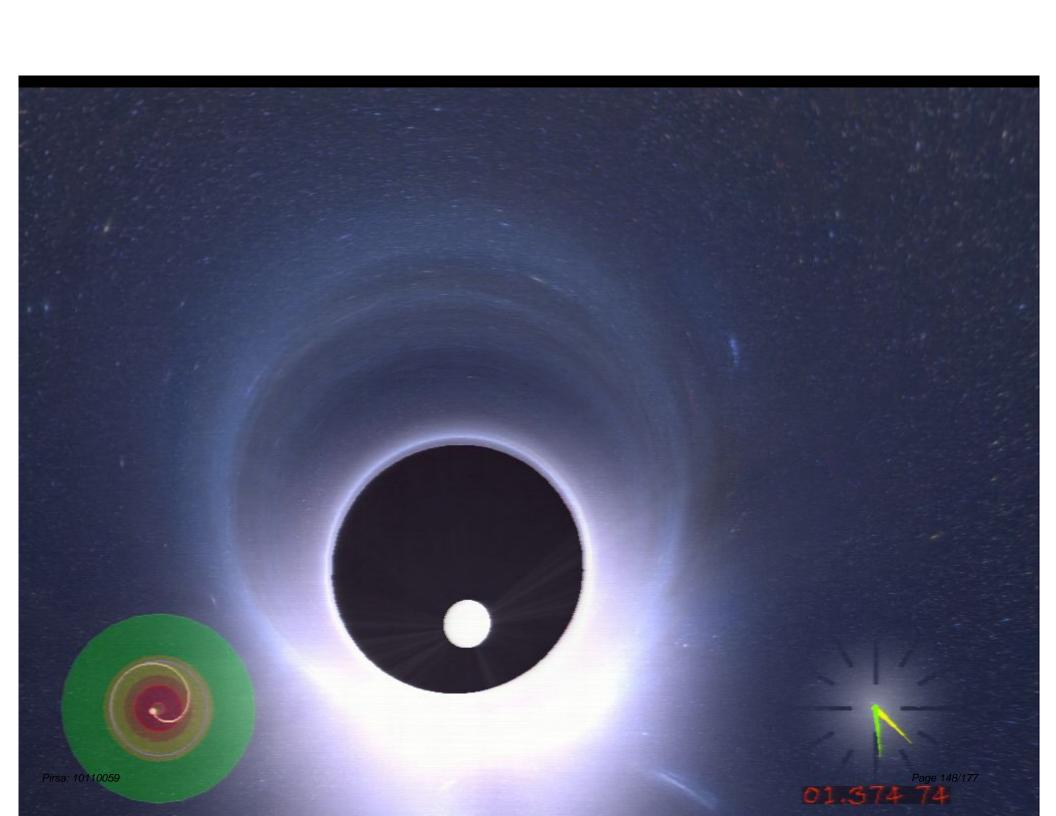


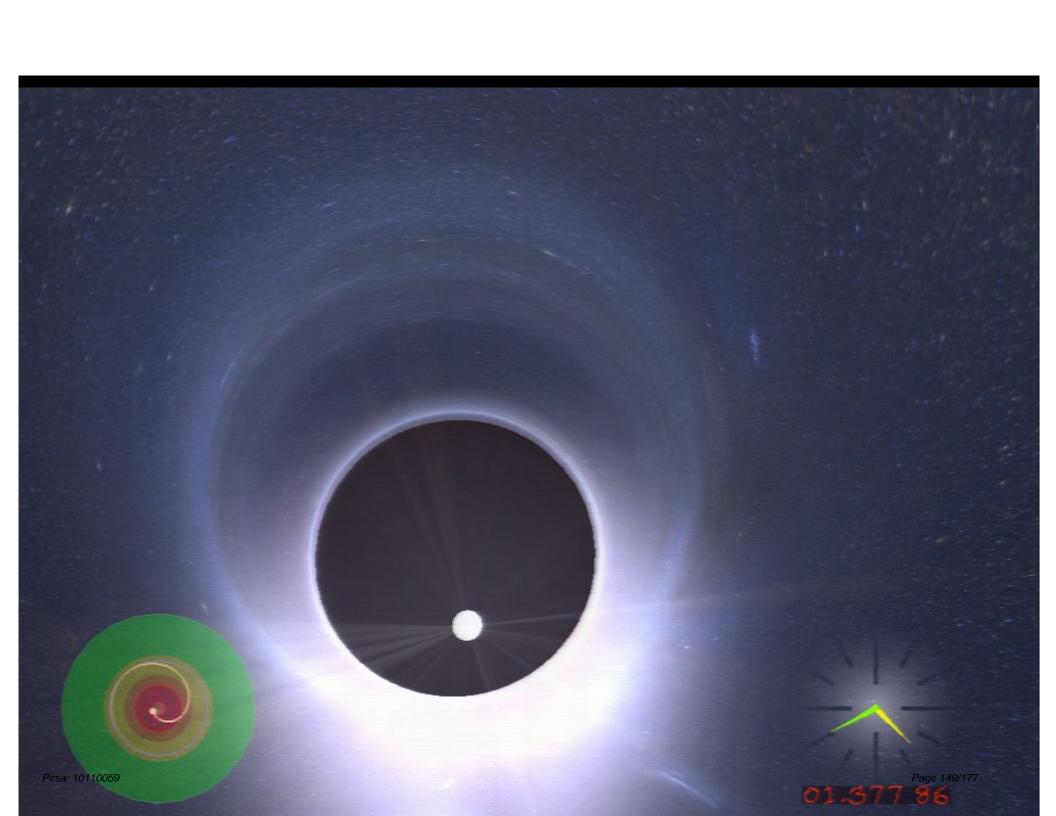


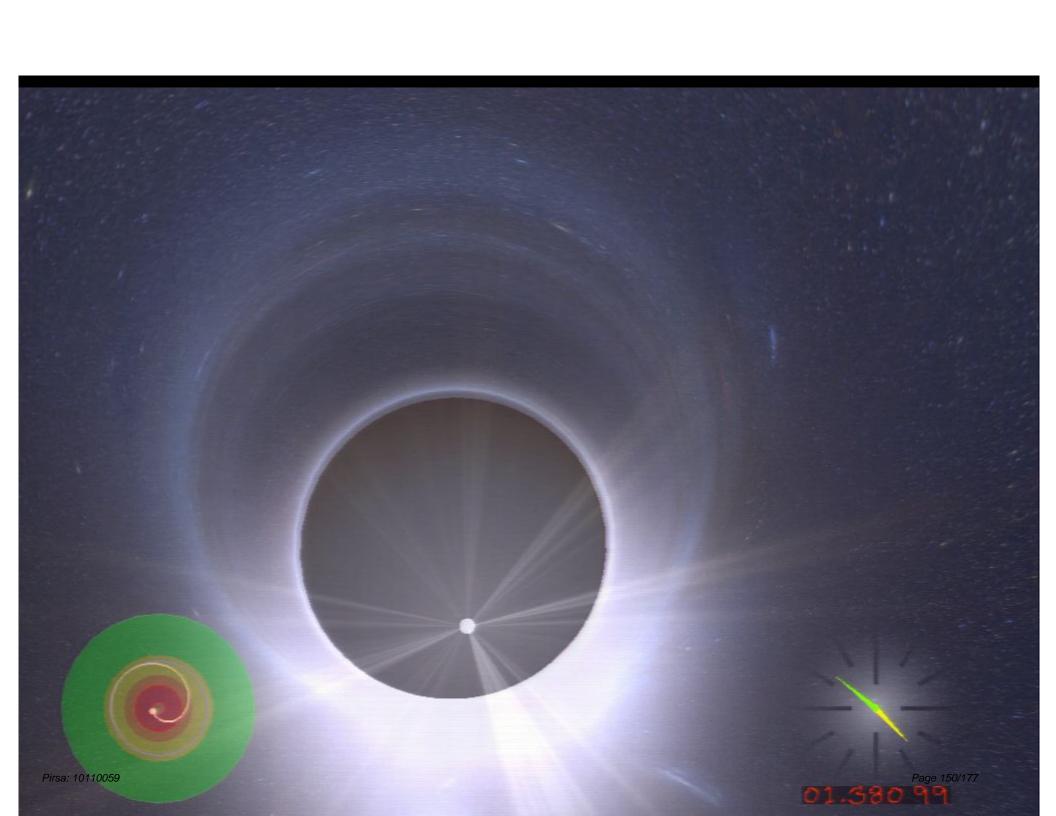


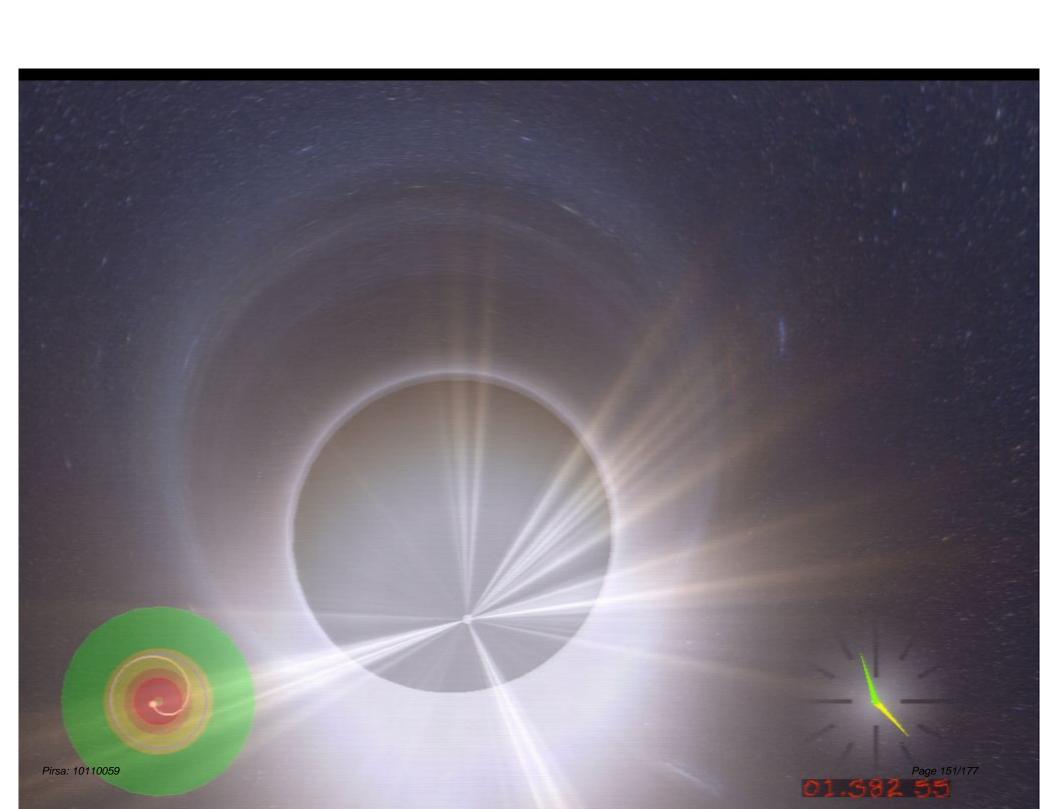


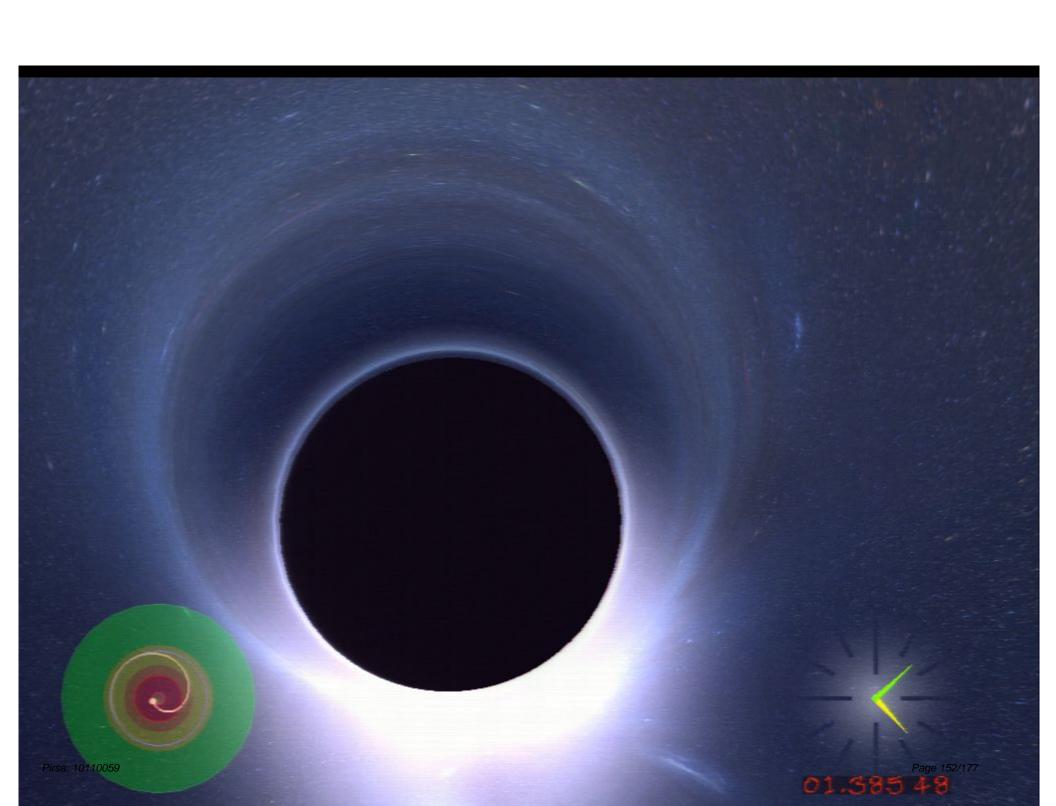


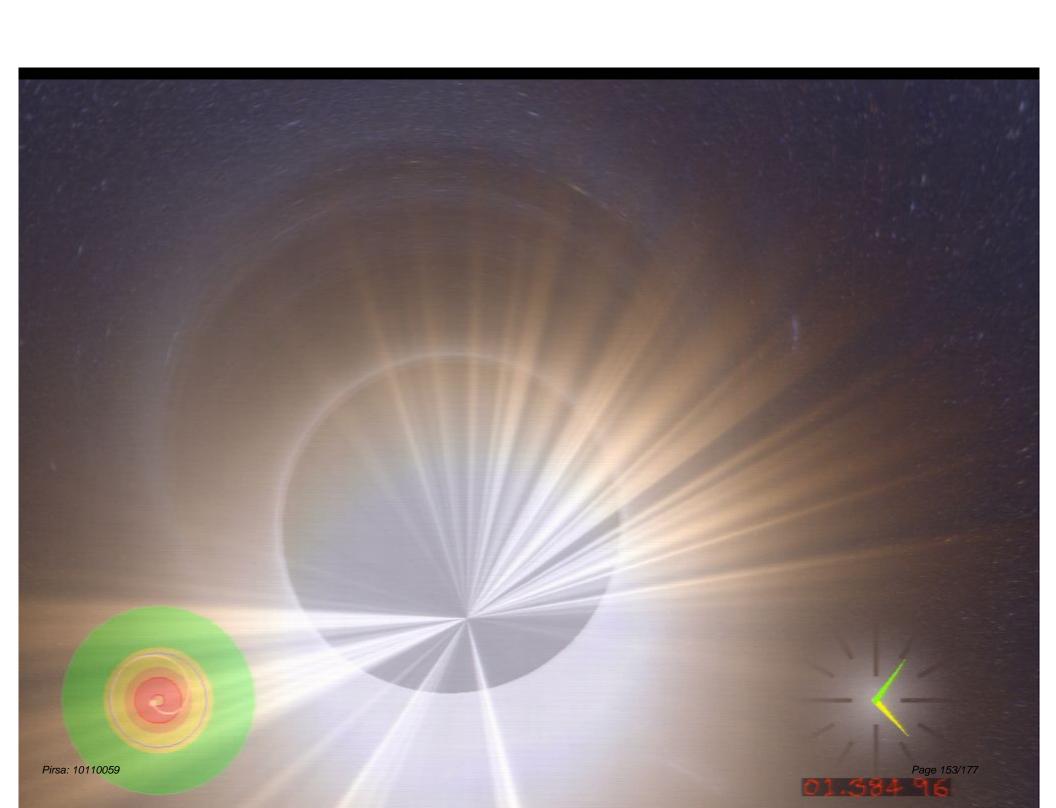


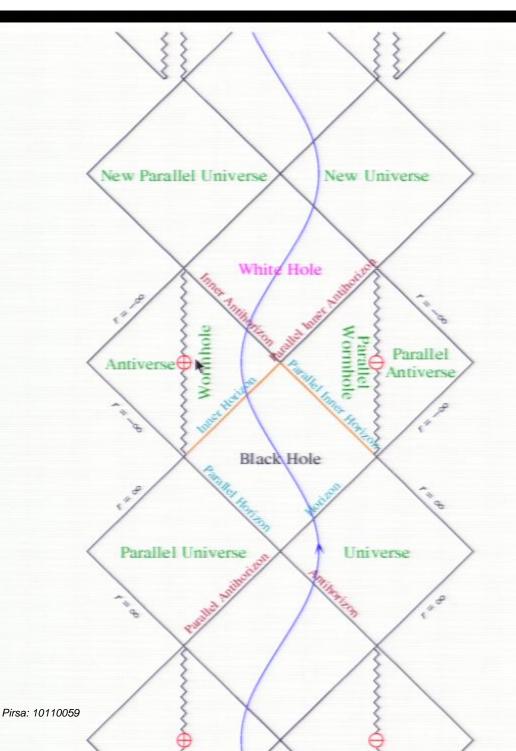






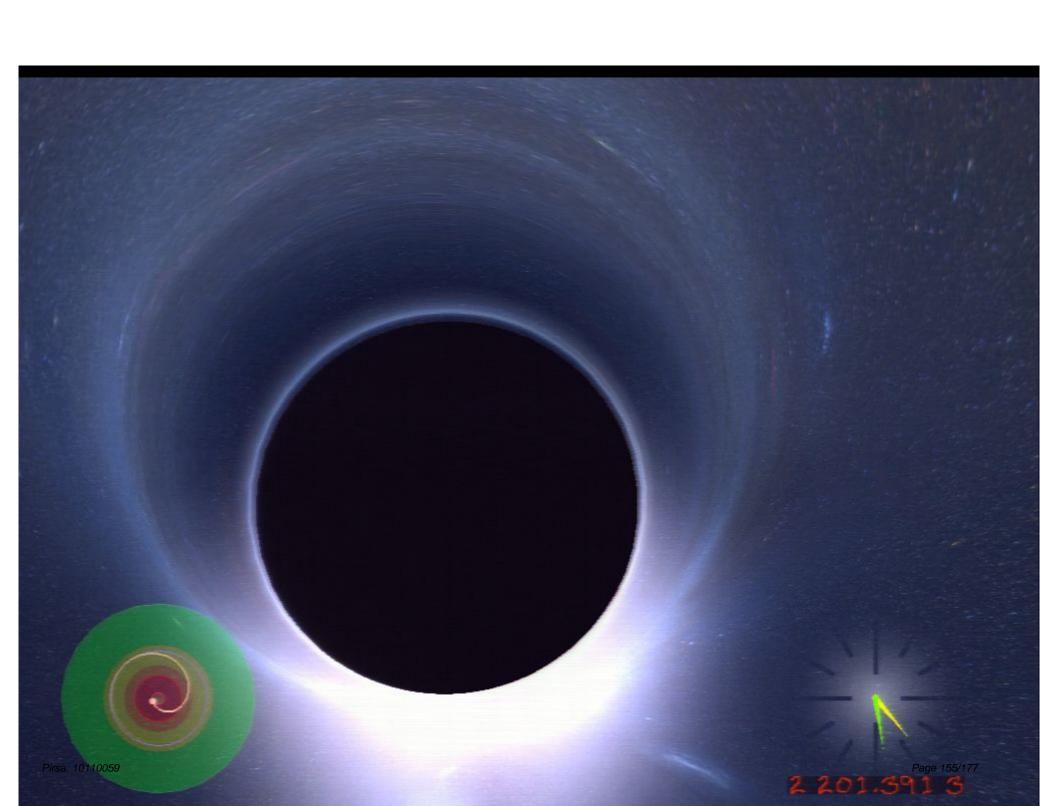


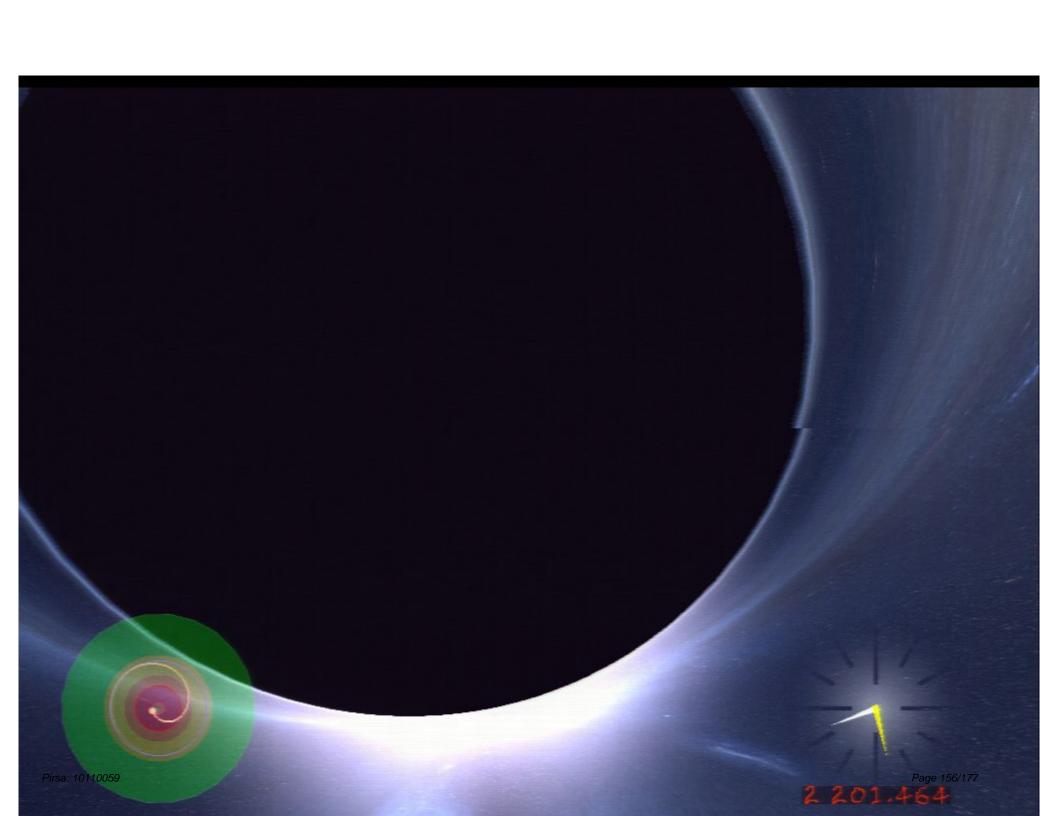


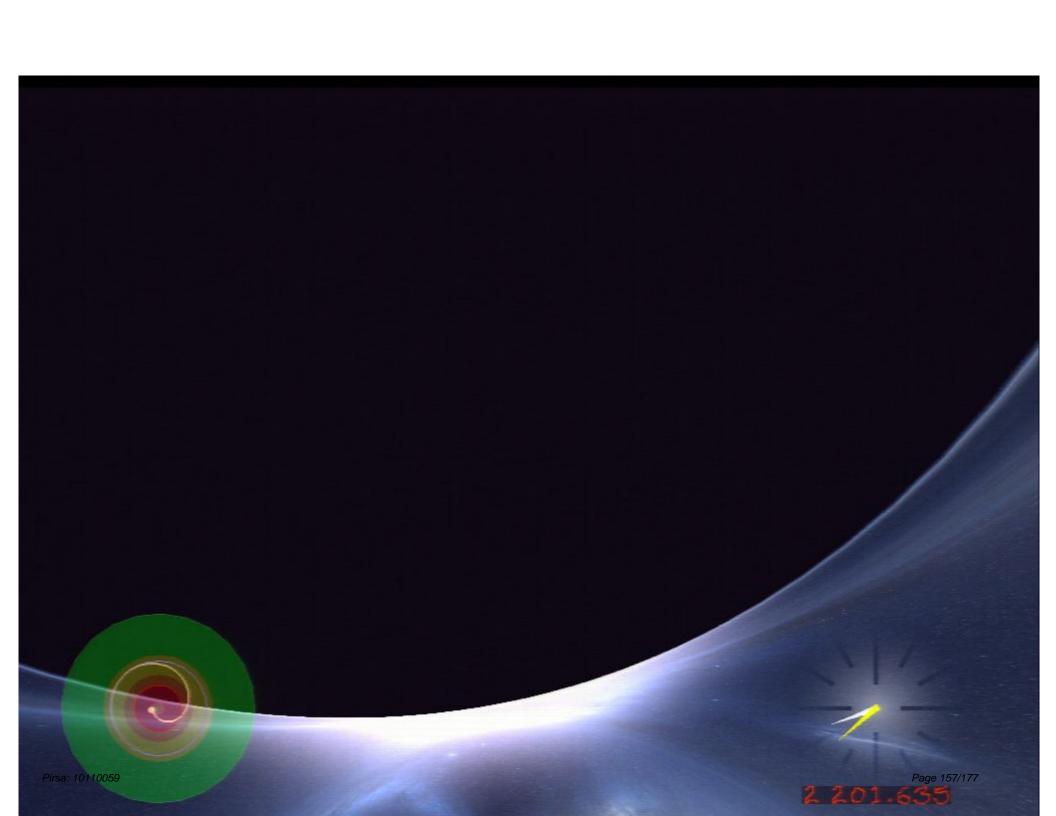


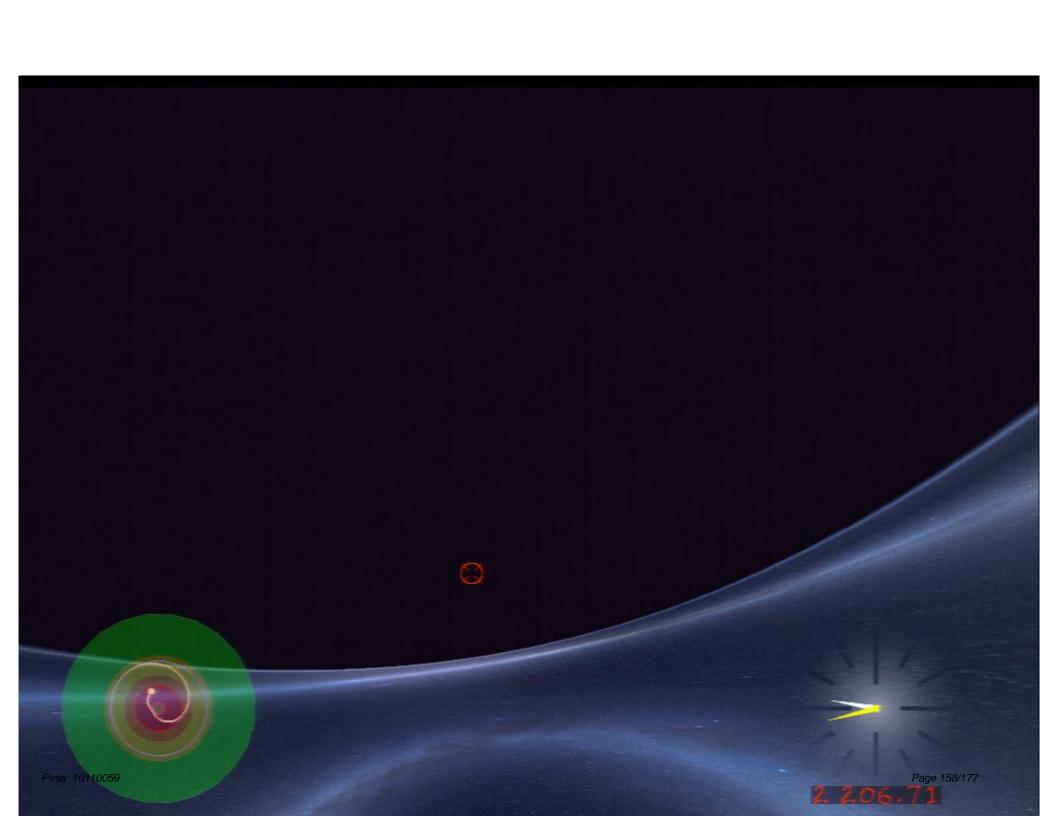
Penrose diagram of a Kerr-Newman (rotating) black hole

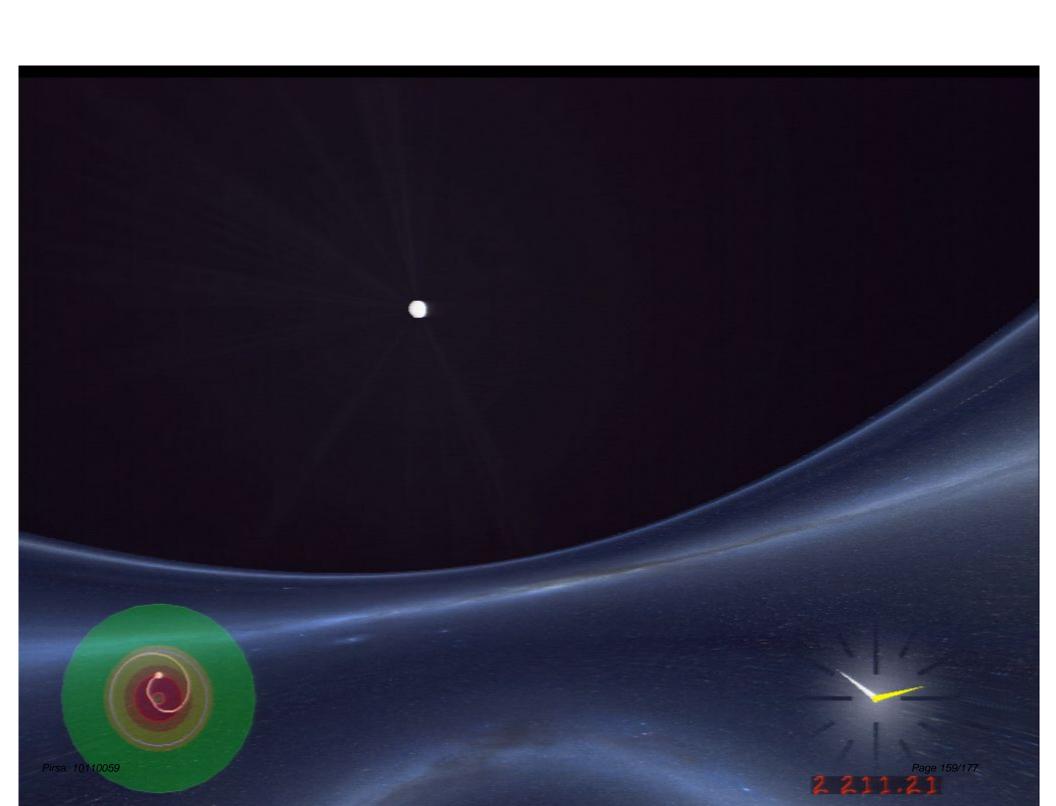
The Penrose diagram of a black hole Reissner-Nordström (charged) is similar.

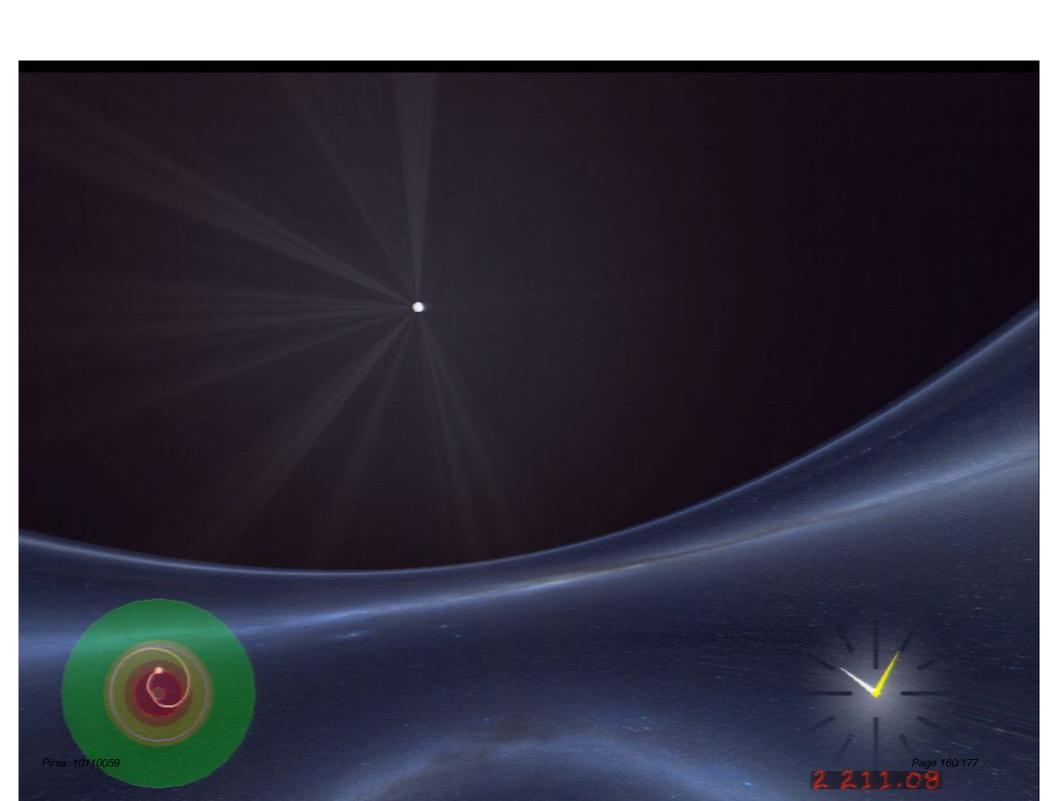


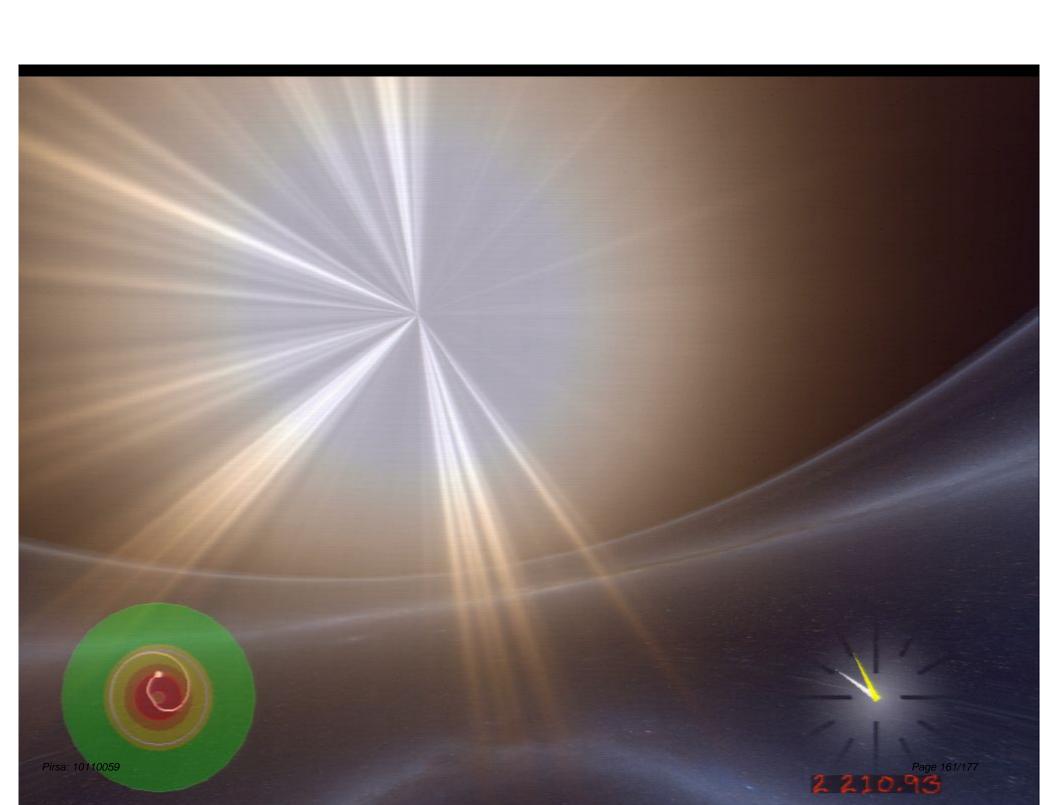


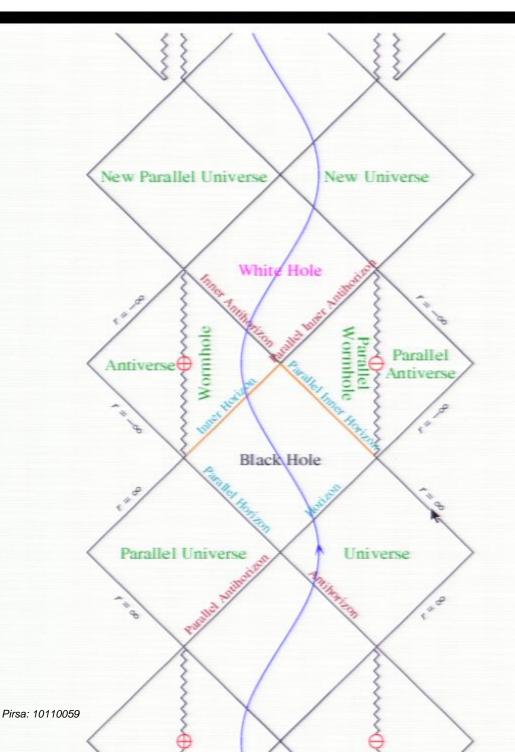












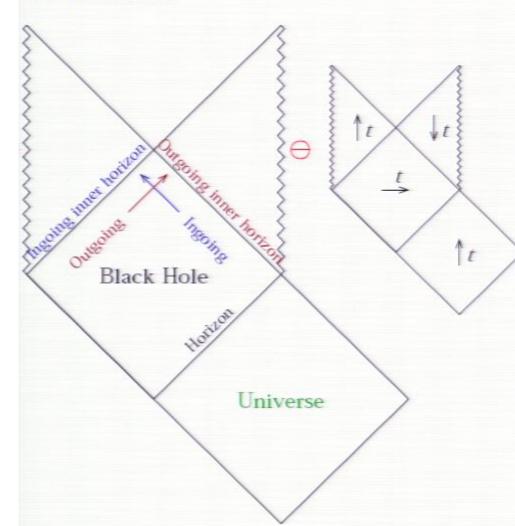
Penrose diagram of a Kerr-Newman (rotating) black hole

The Penrose diagram of a black hole Reissner-Nordström (charged) is similar.

Wormhole waterfall



ass inflation



Penrose (1968) first pointed out that a person crossing an (outgoing) inner horizon would see the outside universe infinitely blueshifted. Poisson & Israel (1990) clarified the full non-linear nature of the instability, which they called "mass inflation".

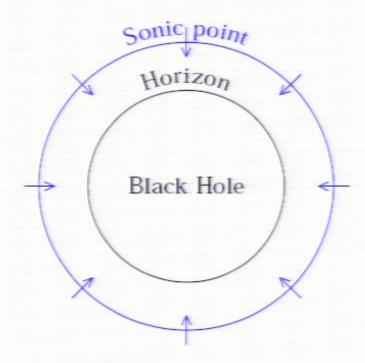
Mass inflation is caused by relativistic counterstreaming between ingoing and outgoing streams near the inner horizon. To drop through the inner horizon, the streams must exceed the speed of light relative to each other, which they cannot do.

The counter-streaming pressure generates a gravitational force, which accelerates the streams faster through each other.

Self-similar model

elf-similar model

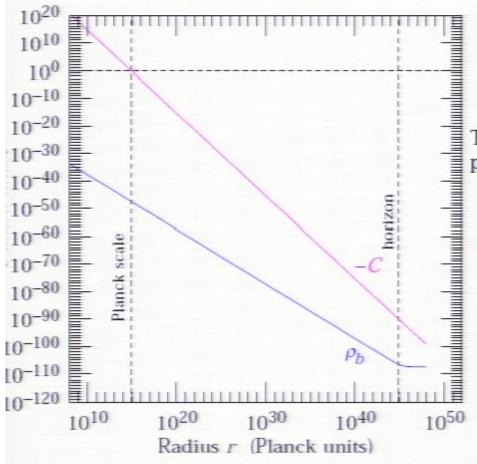
lopt fully non-linear general relativistic model in which black hole accretes self-similarly with $M \propto t$.



Uncharged black hole

.

ncharged black hole

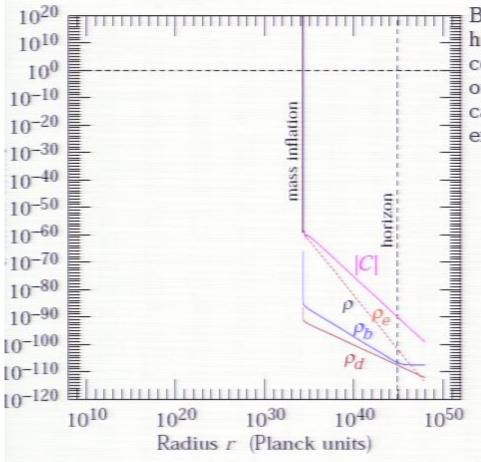


The familiar "Schwarzschild" case: the plasma plunges to a central curvature-dominated singul

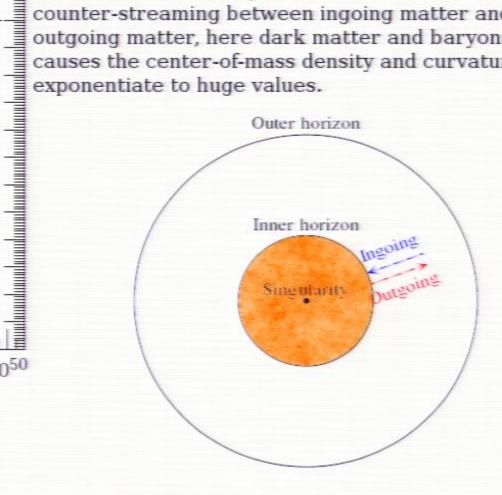
Black hole mass $M_{\bullet} = 4 \times 10^6 \,\mathrm{M}_{\odot}$ Accretion rate dM_{\bullet}/dt $= 4 \times 10^6 \,\mathrm{M}_{\odot}$ / Age of Universe $\approx 20 \,\mathrm{sec} / 4 \times 10^{17} \,\mathrm{sec} \approx 10^{-16}$ plunges to a central curvature-dominated singularity.

Mass inflation by counter-streaming baryons and dark matter

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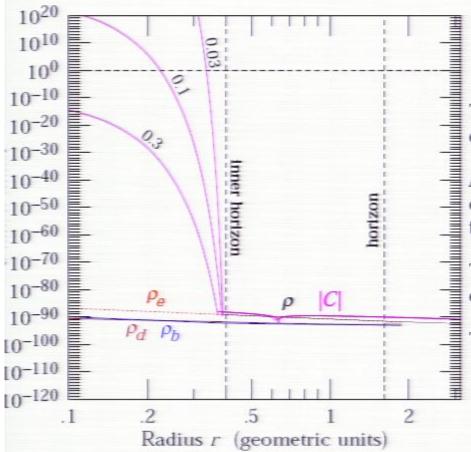


But if the black hole is charged, so it has an inner horizon like a rotating black hole, then relativistic counter-streaming between in a counter-stream i counter-streaming between ingoing matter and outgoing matter, here dark matter and baryons, causes the center-of-mass density and curvature to



Lower accretion rates yield faster inflation

ower accretion rates yield faster inflation



*

The smaller the accretion rate, the faster inflation exponentiates.

Almost all the energy is in the counter-streaming energy. The baryon and dark matter streams themselves remain relatively cold.

The center-of-mass energy exceeds the Planck energy, but the entropy remains low.

The label on each curve is the accretion rate dM_{\bullet}/dt .

Sagittarius A* parameters

he Milky Way Black Hole

Mass = 4 million solar masses

Radius = GM/c^2 = 6 million km = 10 solar radii

Natural timescale = GM/c^3 = 20 seconds

Dimensionless accretion rate =
$$\frac{20 \text{ seconds}}{\text{Age of Universe}} = 10^{-16}$$

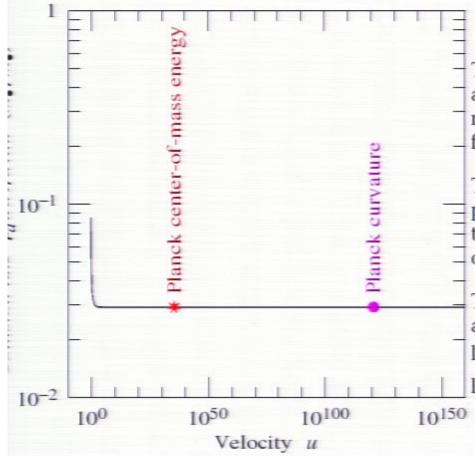
Inflationary scale length = 10^{-16} × Radius = 10^{-3} millimeters

 \Rightarrow In 1 millimeter, the streaming velocity, density, curvature, at the inner horizon exponentiate by e^{1000} , exceeding Planck.

f quantum gravity did not intervene, he streaming velocity would exponentiate to $e^{10^{16}}$ efore inflation ends and streams collapse to singularity at zero radius.

The black hole particle accelerator

he black hole particle accelerator



The black hole acts like a particle accelerator. The acceleration is exponential: particles undergo equal numbers of collisions in equal decades of Lorentz factor.

The collider is powered by its own gravity. The pressure of the counter-streaming generates gravity that accelerates the streams faster through each other.

The smaller the accretion rate, the faster the acceleration, and correspondingly the smaller the luminosity. For $dM_{\bullet}/dt=10^{-16}$,

luminosity = 10^{35} fb⁻¹ per year.

Baby universes?

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can the black hole particle accelerator make baby universes?

- Easily enough energy, and oodles of luminosity.
- Cooling problem. Our Universe began with a low entropy, vacuumdominated equation of state, over at least several Planck lengths.
 Smashing things together in a particle accelerator tends to produce lots of entropy.
- Puture work. The prevailing idea is that super-Planckian collisions generically produce mini-black holes. This would (we expect) predict a burst of mini-black hole creation, followed by hastened collapse.

<u> History</u>

B

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lack hole history

- 315 Schwarzschild discovers the solution for a spherical black hole.
- 363 47 years later, Kerr discovers the exterior solution for a rotating black hole.
- 110 47 years later, "No satisfactory interior solutions are known"
 Stephani et al. Exact Solutions to Einstein's Field Equations, 2nd edition.
- Rotating inflationary black hole solutions

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Rotating inflationary black hole solutions

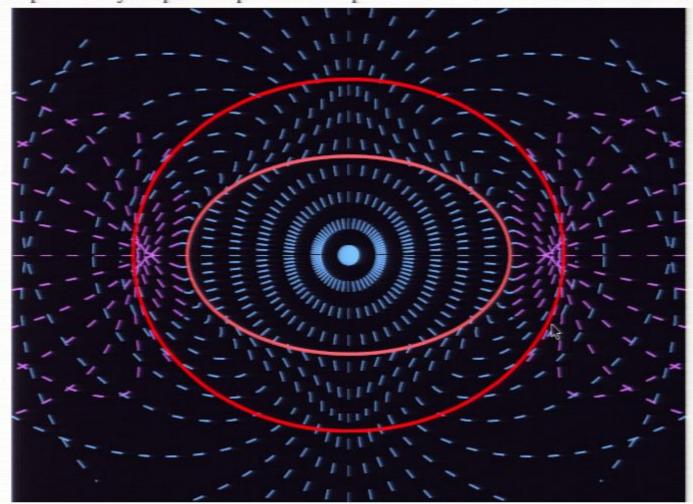
otating inflationary black hole solutions amilton & Polhemus (2010) arXiv:1010.1269,71,72

Rotating black hole accretes collisionless fluid.

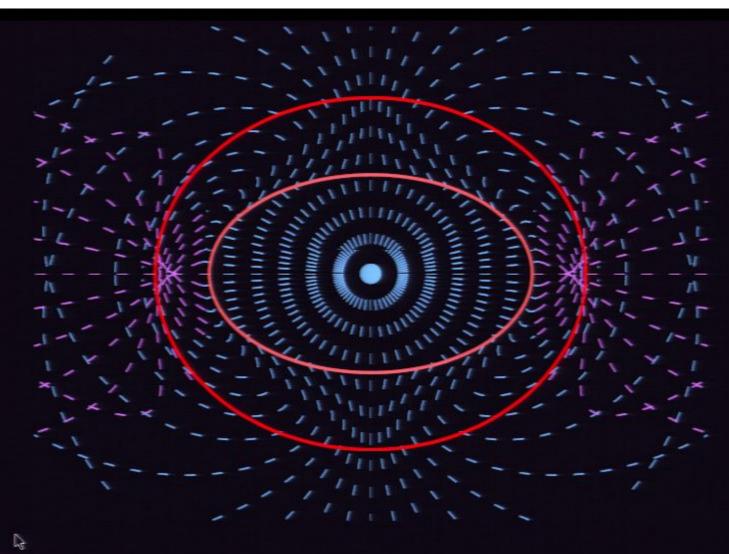
Hyper-relativistic counter-streaming drives inflation just above inner horizon.

Solutions are exact, conformally stationary, axisymmetric, separable.

Separability requires special flow pattern incident on inner horizon.



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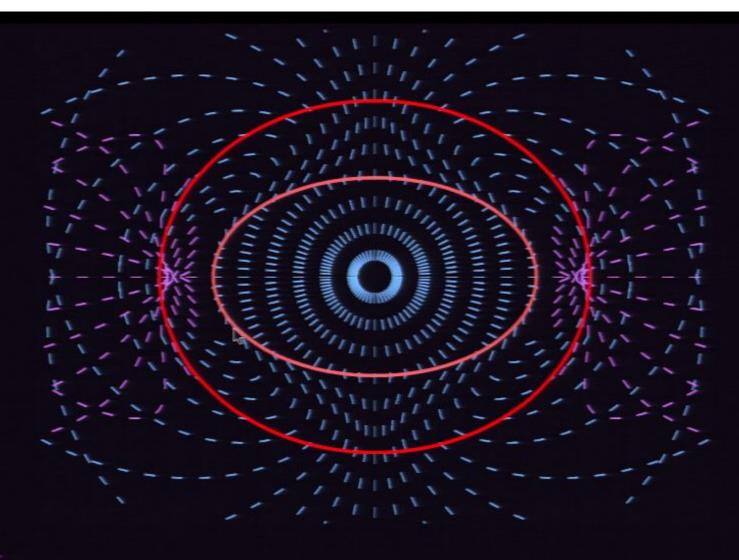


onclusions

- A black hole is a place where space falls faster than light.
- The inflationary instability at the inner horizon acts like a particle accelerator of extraordinary power.
- There exist exact solutions for rotating black holes that undergo inflation at the inner horizon.

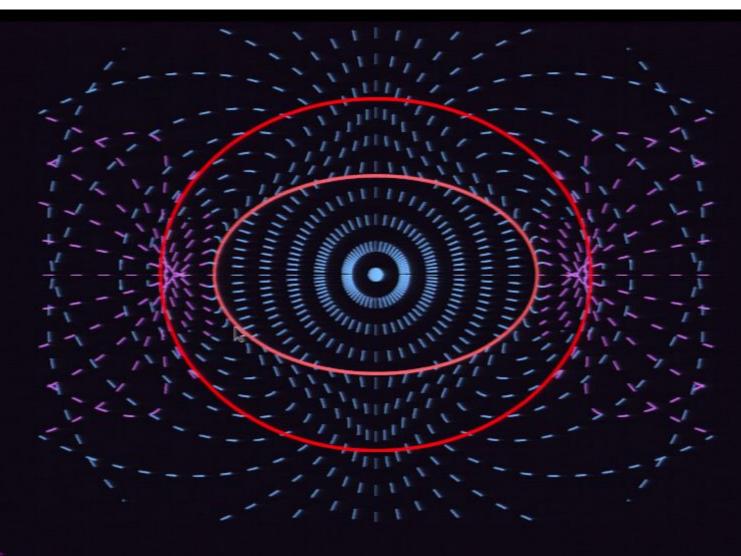
Why does Nature do what she does inside black holes?

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