

Title: On Hydrodynamics of holographic p-wave Superfluids

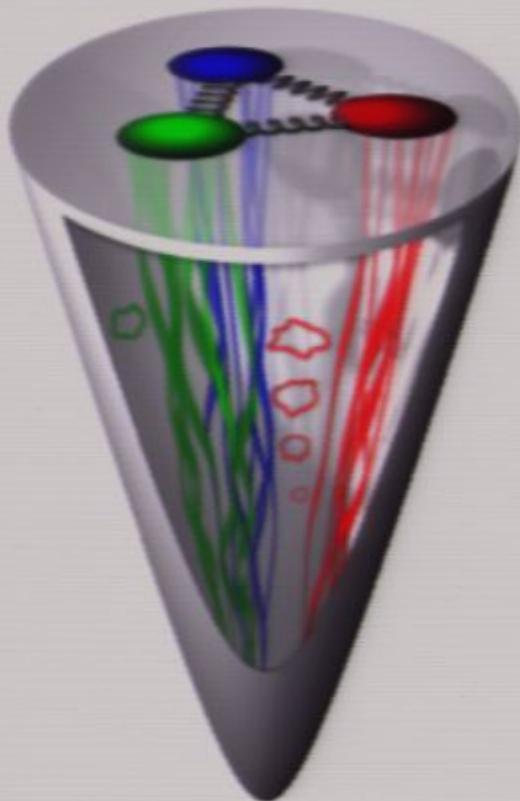
Date: Oct 19, 2010 11:00 AM

URL: <http://pirsa.org/10100083>

Abstract: We discuss holographic duals of strongly interacting gauge theories which show properties of p-wave superfluids which in addition to an Abelian symmetry also break the spatial rotational symmetry. The gravity duals of these superfluid states are black hole solutions with a vector hair which we construct in a non-Abelian Einstein-Yang-Mills theory and in the D3/D7 brane setup. The latter allows us to identify the dual field theory explicitly. After we constructed the vector hair state we study the conductivity and shear viscosity which is non-universal due to the breaking of the rotational symmetry.

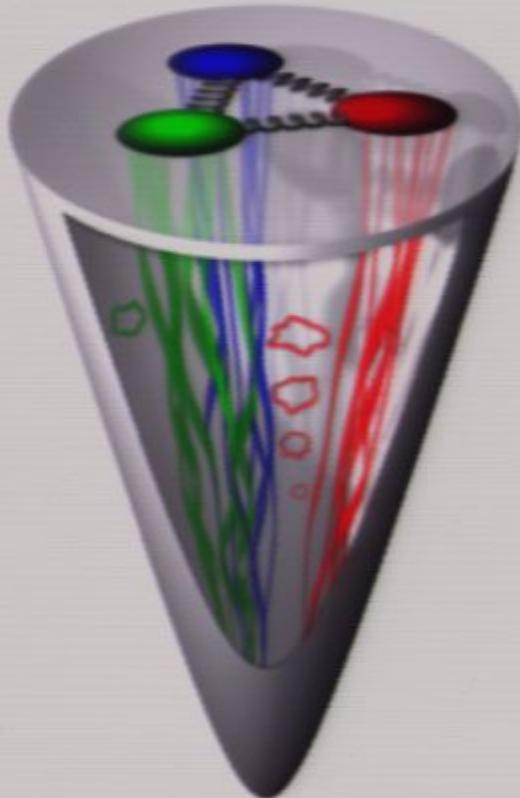
Why is Gauge/Gravity duality useful for Hydrodynamics?

- Allows for simple calculation of real-time correlators [Son, Starinets;...]

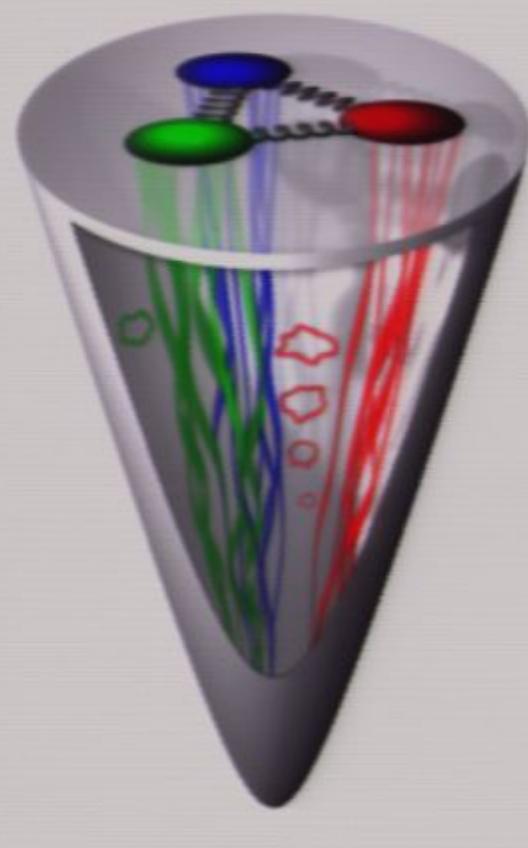


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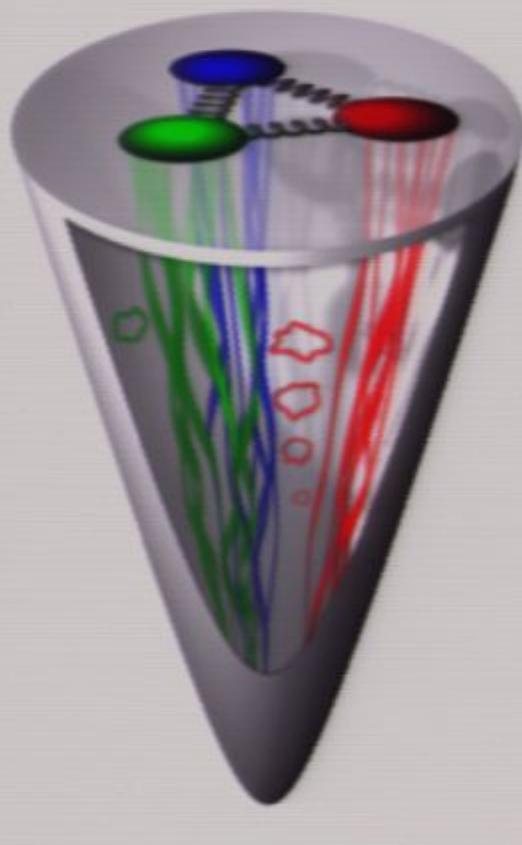


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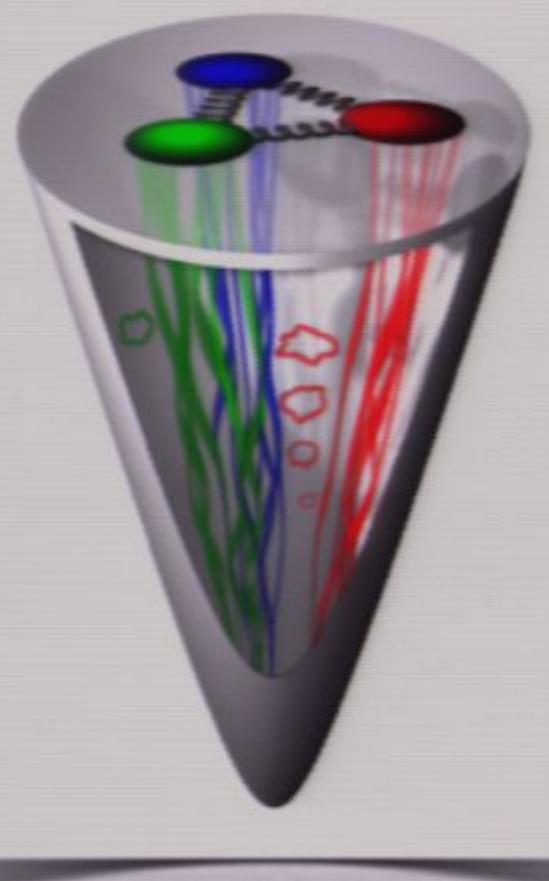
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- Low-lying poles determine hydrodynamic modes and their dispersion relations
- Unambiguous construction of hydrodynamic description possible
- Discovery of new transport coefficients, e.g. in systems with anomaly [Erdmenger et al.; Son, Son]

Why Hydrodynamics of p-wave Superfluids?

- Superfluids in general interesting since new hydrodynamic mode (Goldstone bosons) present, e.g. second and fourth sound [Herzog, Pufu; Yarom; Herzog, Yarom]

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- In p-wave Superfluids rotational symmetry also spontaneously broken \Rightarrow more hydrodynamic modes
- Rotational symmetry is often spontaneously broken in condense matter systems, e.g. liquid crystals, d-wave superconductors

Outline

(1) Motivation (over)

(2) Construct holographic p-wave Superfluid as black
hole with vector hair

[Ammon, Erdmenger, Grass, P.K., O'Bannon]

(3) Embedding into String Theory [Ammon, Erdmenger, Kaminski, P.K.]

(4) Towards the hydrodynamic description

- i. Classification and Decoupling of fluctuations
- ii. Conductivity and Shear Viscosity

(5) Summery and Outlook

Gauge/Gravity Duality

- Type IIB SUGRA on $AdS_5 \times X_5$ is dual to Conformal Field Theory at large N_c and large λ in the sense

$$Z_{\text{SUGRA}} [\phi(x, r)|_{r \rightarrow r_{\text{bdy}}} = \phi_0(x)] = \left\langle e^{\int d^4x \phi_0(x) \mathcal{O}(x)} \right\rangle$$

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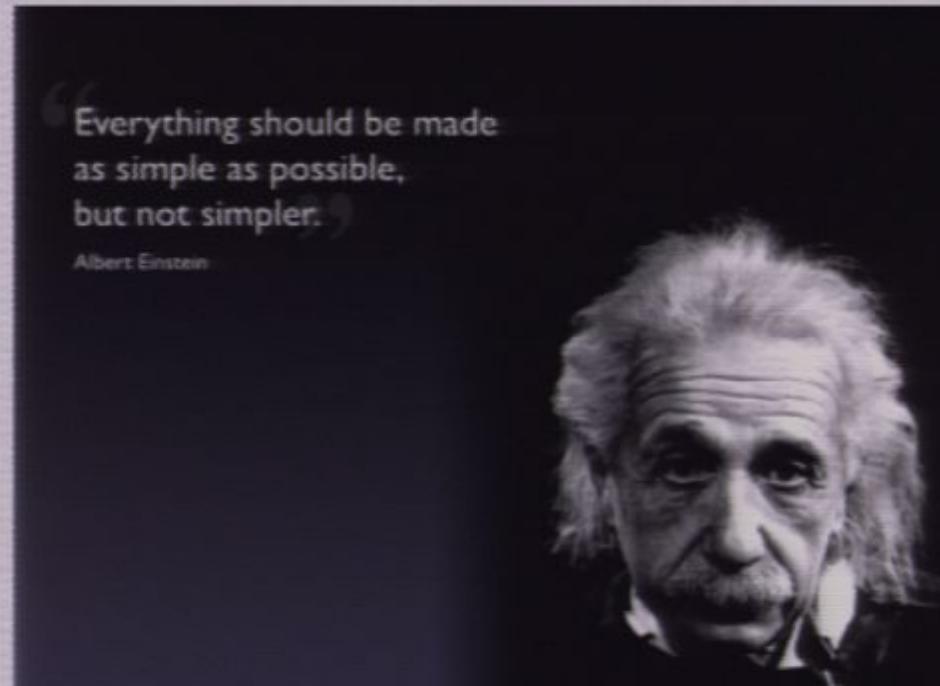
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- Black holes correspond to thermal field theories
- Gauge fields A_μ^a are dual to global currents J_a^μ especially vevs A_t^3 induce finite chemical potentials μ_I (source) and finite densities $\langle J_3^t \rangle$.

Gravity model

- Einstein-Yang-Mills theory with $SU(2)$ gauge group

$$S = \int d^5x \sqrt{-g} \left[\frac{1}{2\kappa_5^2} (R - \Lambda) - \frac{1}{4\hat{g}^2} F_{\mu\nu}^a F^{a\mu\nu} \right] \quad \alpha = \frac{\kappa_5}{\hat{g}}$$



“Everything should be made
as simple as possible,
but not simpler.”

Albert Einstein

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Take $\langle J_1^x \rangle$ dual to A_x^1 (only the vev)

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- Superfluid condensate in addition to chemical potential: !
Take $\langle J_1^x \rangle$ dual to A_x^1 (only the vev)
- $\langle J_1^x \rangle$ spontaneously breaks $U(1)_3$ down to \mathbb{Z}_2 and $SO(3)$ down to $SO(2) \Rightarrow$ p-wave superfluid

Field Theory Interpretation of α

- Holographic calculations of Weyl anomaly:
 $1/\kappa_5^2 \propto c$, c : number of degrees of freedom
- Correlators of $SU(2)$ currents proportional to $1/\hat{g}^2$
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- Intuitively,

$$\alpha^2 = \frac{\kappa_5^2}{\hat{g}^2} \propto \frac{\# \text{ charged degrees of freedom}}{\# \text{ total degrees of freedom}}$$

Behavior at finite chemical potential

- Reissner-Nordström black hole

$$ds^2 = -N(r)dt^2 + \frac{1}{N(r)}dr^2 + r^2d\vec{x}^2$$

$$N(r) = r^2 - \frac{2m_0}{r^2} + \frac{2\alpha^2q^2}{3r^4}$$

$$A = \left(\mu - \frac{q}{r}\right)\tau^3 dt$$

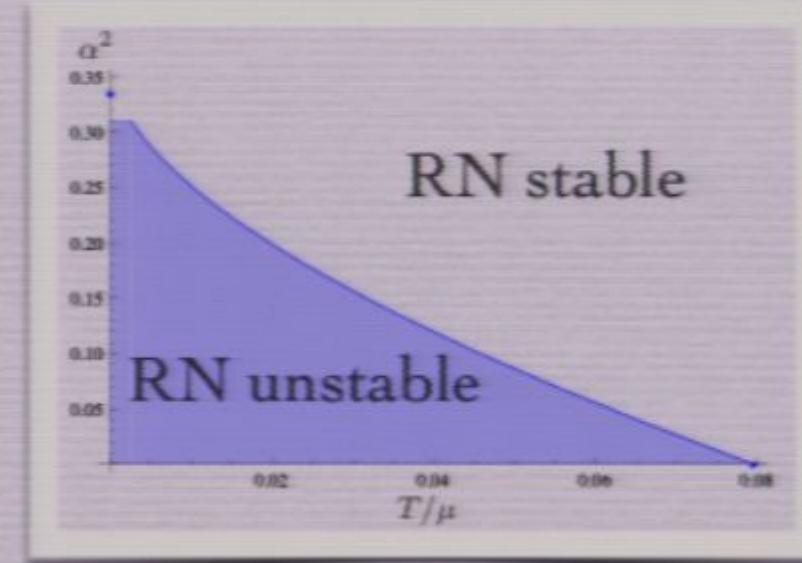
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- At low temperature the back hole is unstable against fluctuations in $A_x^1 \Rightarrow$ Condensation

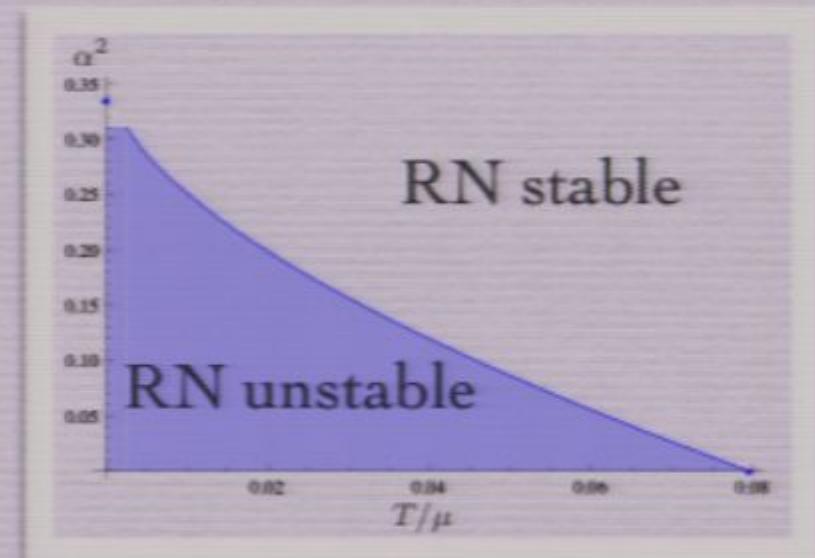
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Condensation Process

Gubser

- Sketchy action for A_x^1 :

$$S \sim \partial_\mu A_x^1 \partial^\mu A_x^1 + \underbrace{2g^{tt} g^{xx} (A_t^3)^2}_{=m_{\text{eff}}} (A_x^1)^2$$

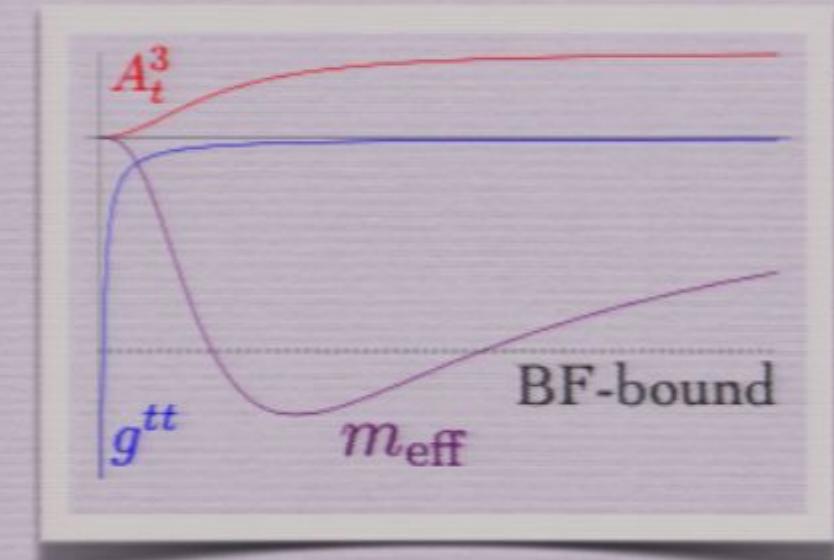
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 m_{eff} can be lower than BF-bound
⇒ Instability
⇒ Condensation ⇒ vector hair



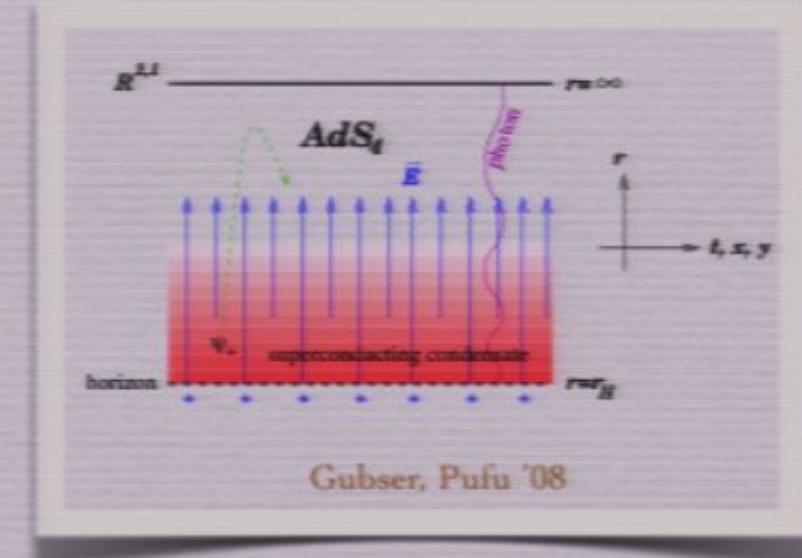
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Gubser, Pufu '08

Solutions in the broken phase

- We numerically solve the Einstein-Yang-Mills equations for the ansatz

$$\begin{aligned} ds^2 = & -N(r)\sigma(r)^2 dt^2 + \frac{1}{N(r)} dr^2 \\ & + r^2 f(r)^{-4} dx^2 + r^2 f(r)^2 (dy^2 + dz^2) \\ A = & \phi(r)\tau^3 dt + w(r)\tau^1 dx \quad N(r) = r^2 - \frac{2m(r)}{r^2} \end{aligned}$$

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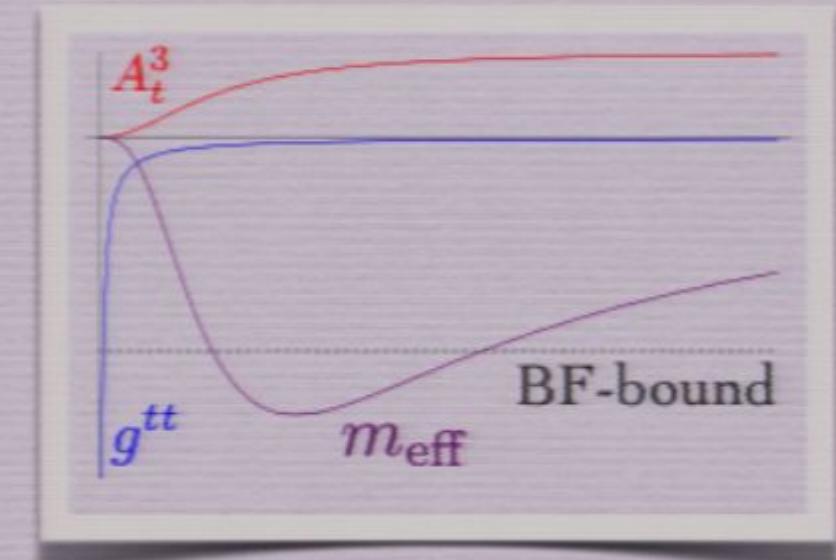
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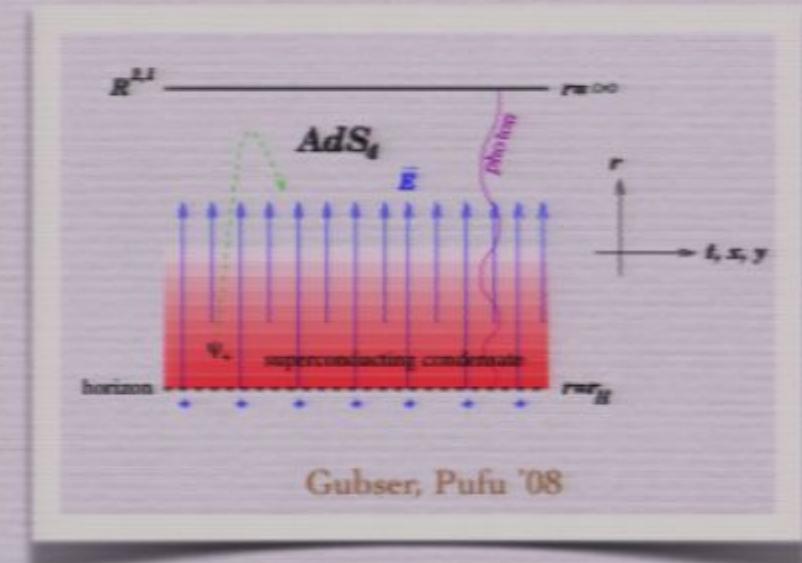
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- Hair is stabilized by the equilibrium of electric and gravitational force in AdS space.



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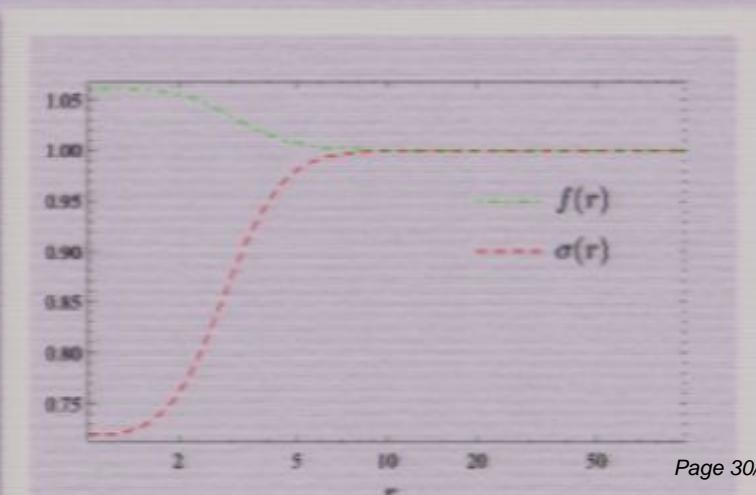
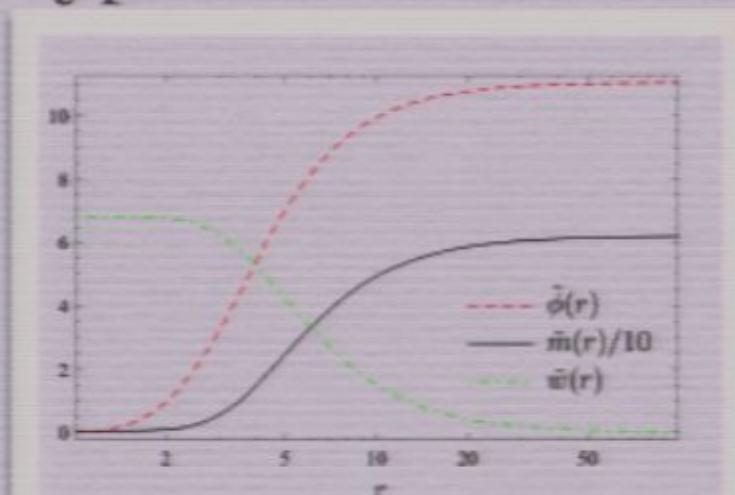
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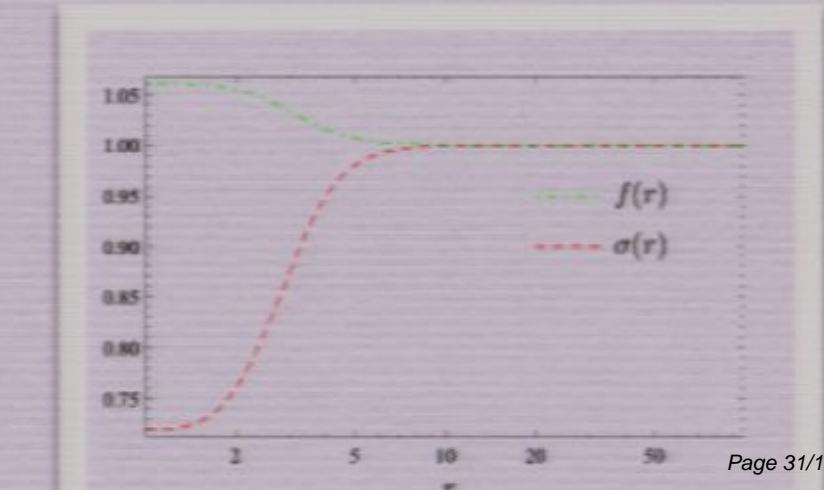
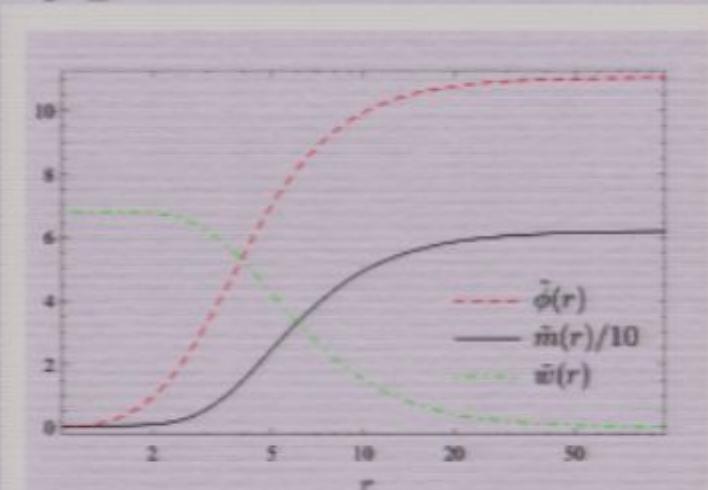
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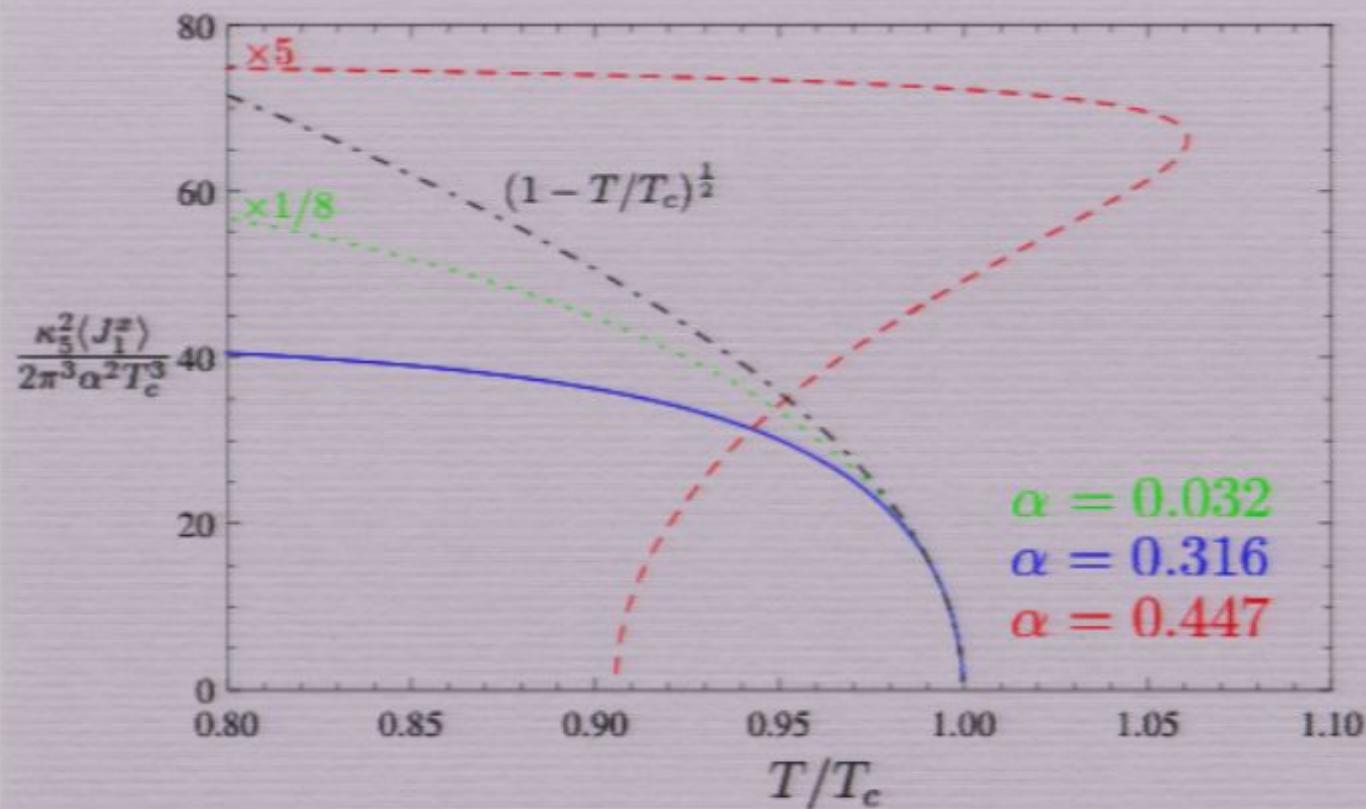
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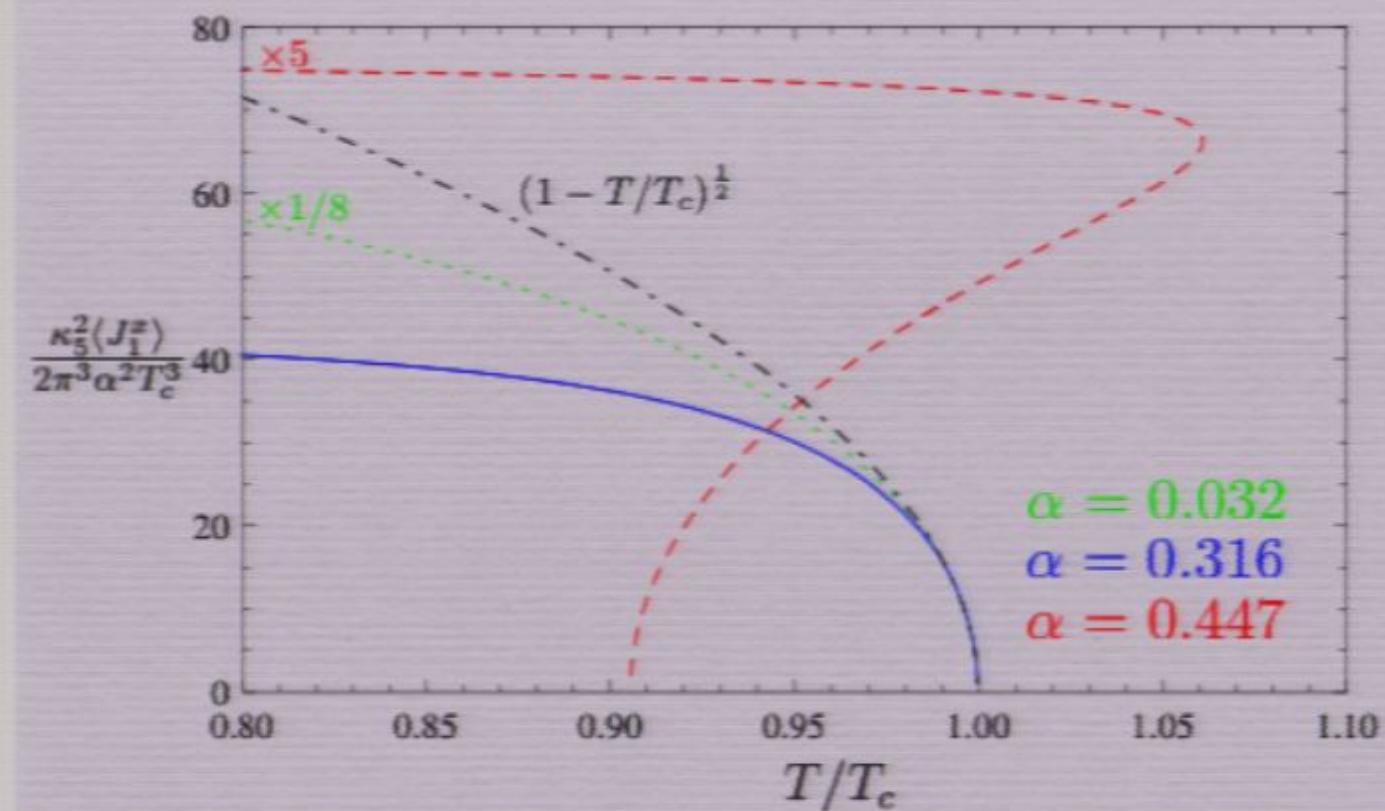
Phase transition

- Order parameter $\langle J_1^x \rangle$ determined by boundary behavior of A_x^1 .



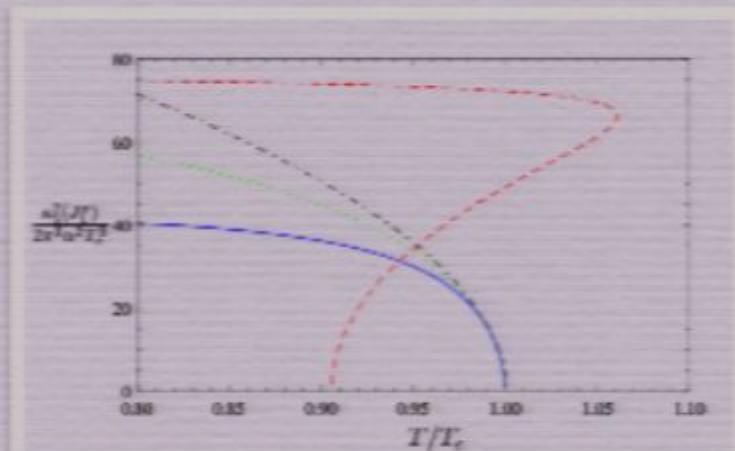
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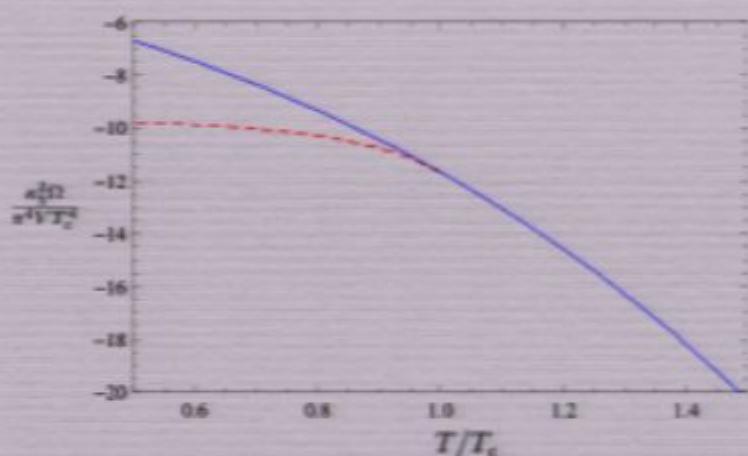
Phase transition

- Order parameter $\langle J_1^x \rangle$ determined by boundary behavior of A_x^1 .
- For $\alpha < \alpha_c$ order parameter increases monotonically \Rightarrow 2nd order transition (mean field)
- For large $\alpha > \alpha_c$ order parameter becomes multivalued \Rightarrow 1st order transition



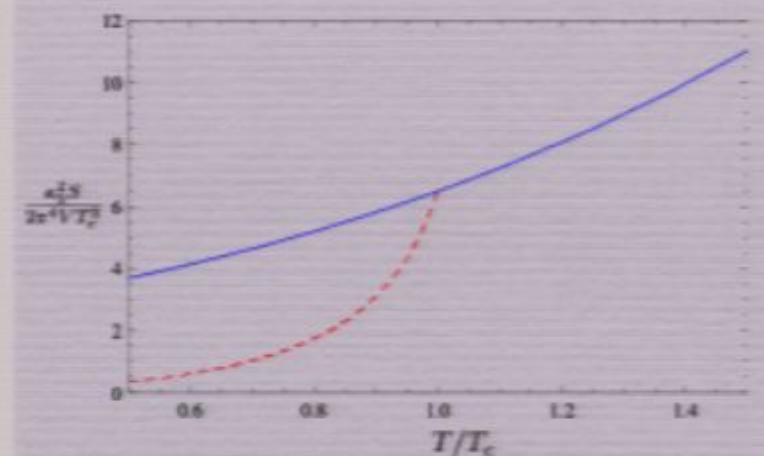
Thermodynamics

$$\alpha < \alpha_c$$



Grand potential

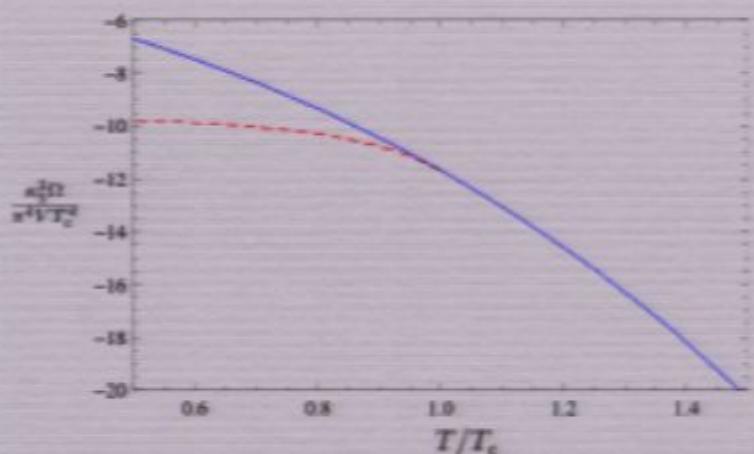
normal phase
superfluid phase



Entropy

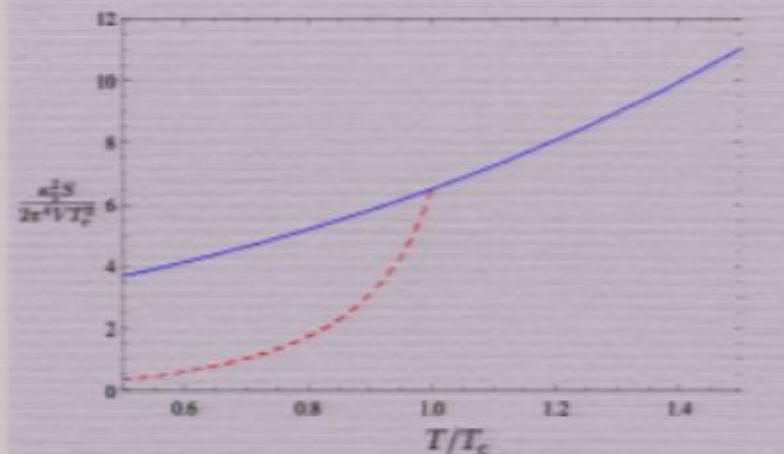
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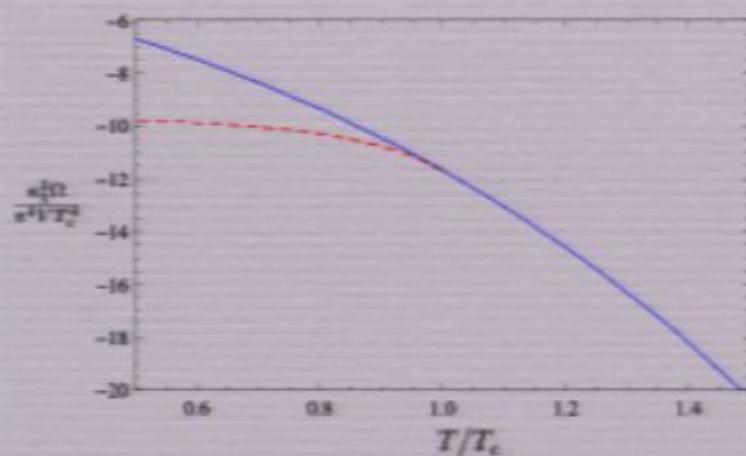
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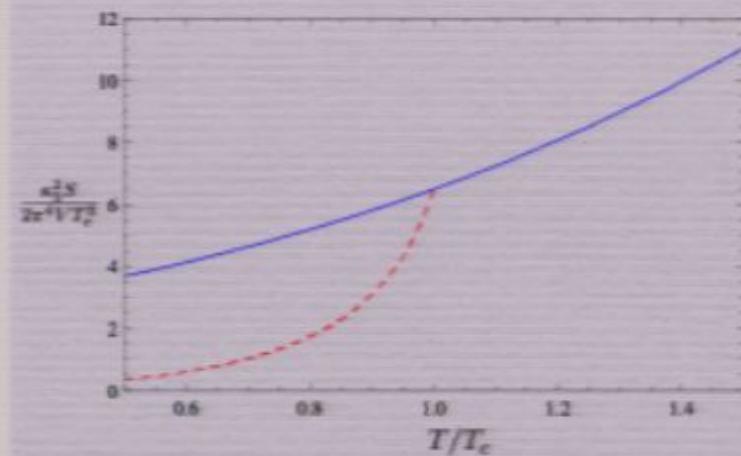
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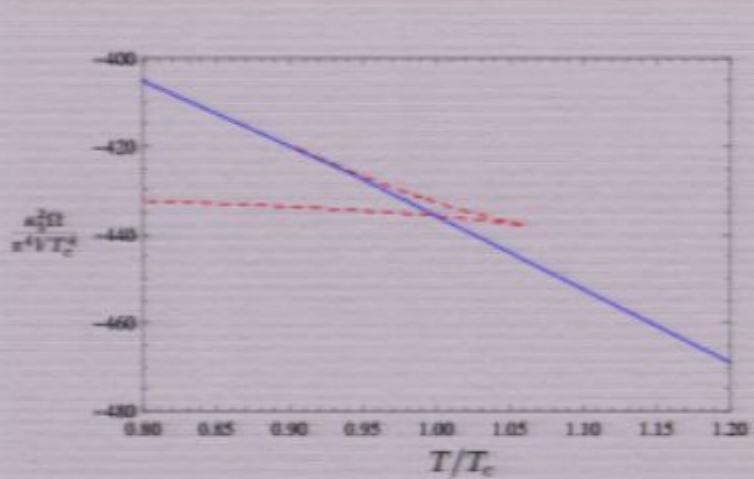
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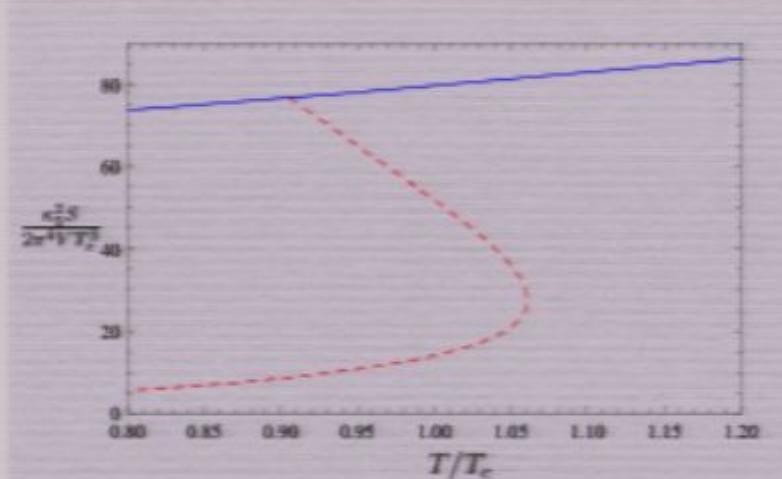
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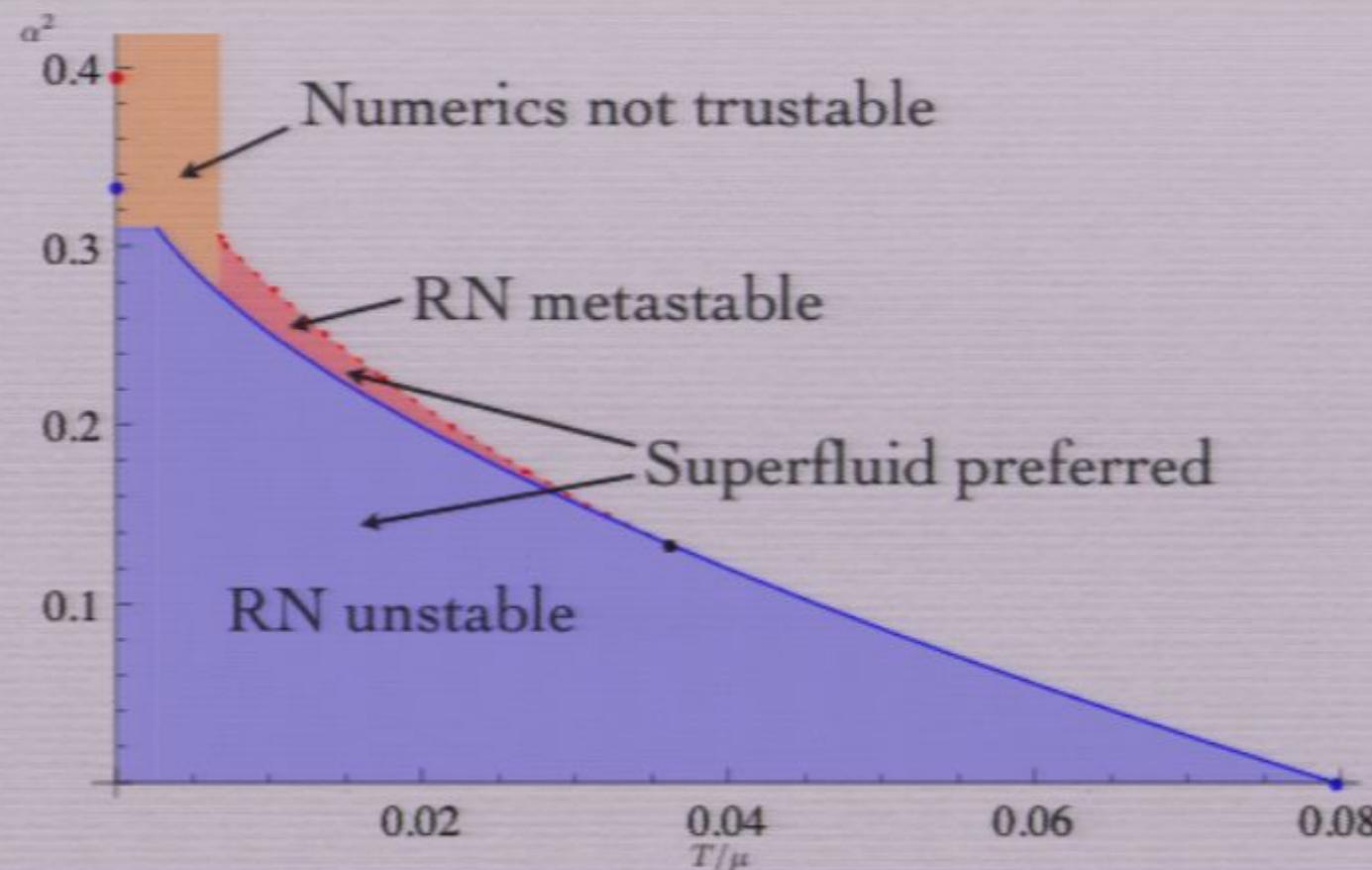
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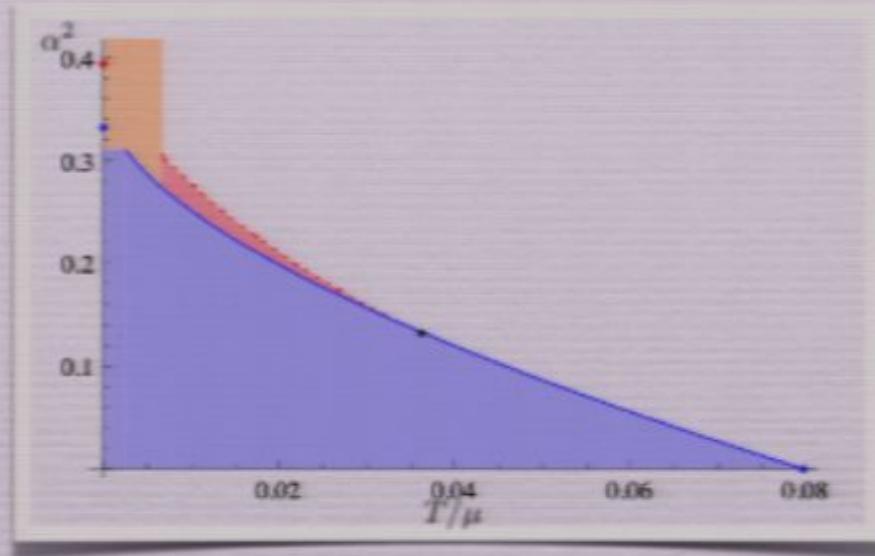


Entropy

“Phase diagram”



“Phase diagram”



- ~ Superfluid thermodynamically preferred in red and blue region
- ~ Phase transition second order for $\alpha < \alpha_c$ and first order for $\alpha > \alpha_c$ with $\alpha_c = 0.365$
- ~ For $\alpha > 0.628$ no superfluid phase available

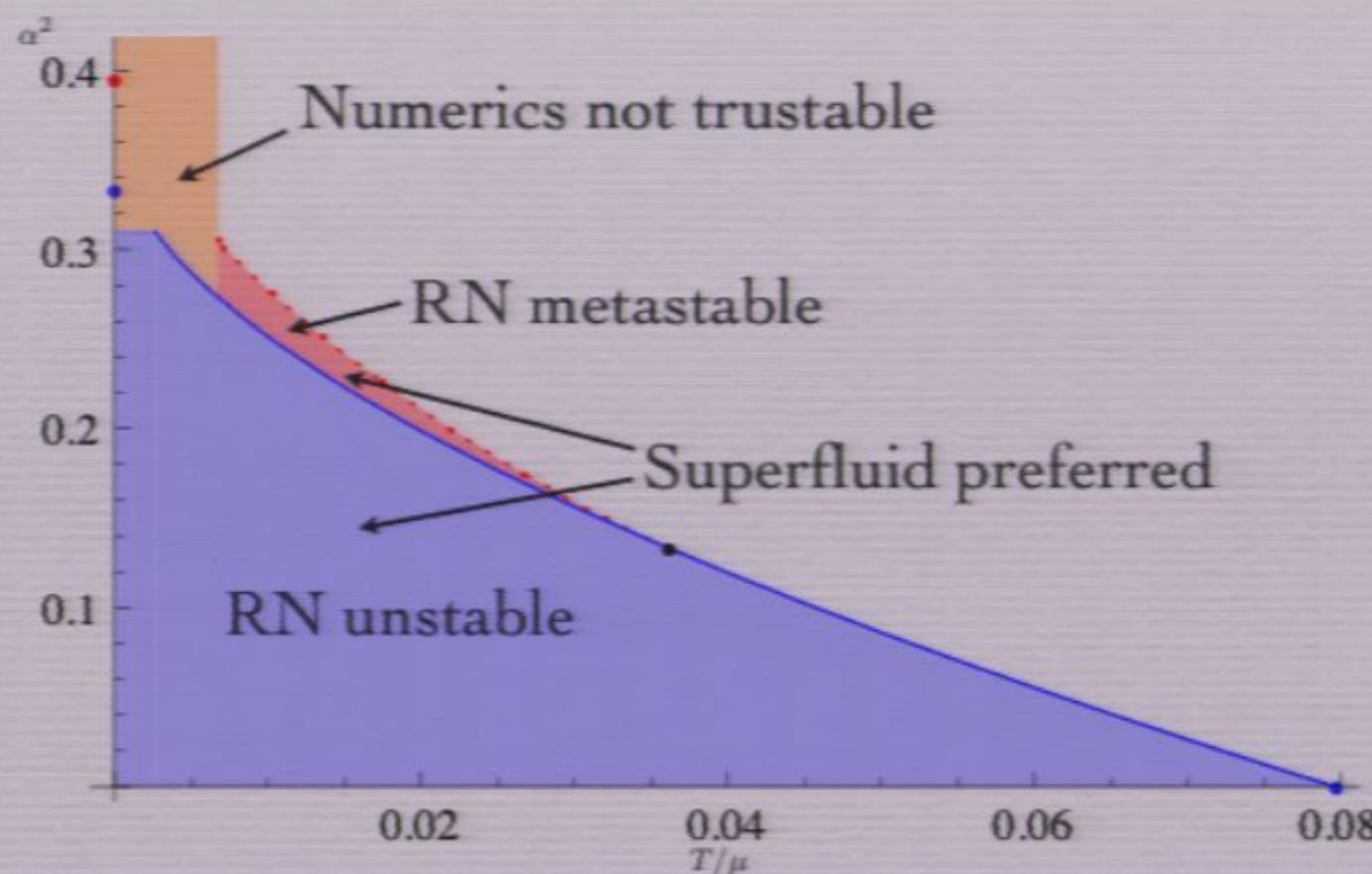
Embedding in String Theory



ZOMBIE STRING THEORISTS
A User's Guide to the Universe

- Gravity model can be embedded into D3/D7 brane setup
- D3/D7 brane setup dual to $\mathcal{N} = 4$ $SU(N_c)$ SYM coupled to $\mathcal{N} = 2$ hypermultiplets
- Coincident D7-branes provide a non-Abelian gauge field \Rightarrow Global flavor symmetry

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Embedding in String Theory non-Abelian DBI action

- Need non-Abelian DBI action (best guess, correct up to $(\alpha')^4$)

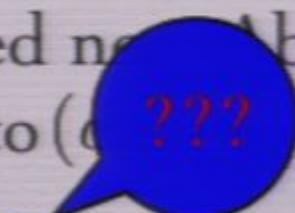
$$S_{\text{DBI}} = T_{D7} \text{Str} \int d^8 \xi \sqrt{\det Q} \sqrt{\det \left(P_{ab} \left[E_{MN} + E_{Mi} (Q^{-1} - \delta)^{ij} E_{jN} \right] + 2\pi\alpha' F_{ab} \right)}$$

$$E_{MN} = g_{MN} + B_{MN} \quad Q_j^i = \delta_j^i + i2\pi\alpha' [\Phi^i, \Phi^k] E_{kj} \quad \begin{matrix} M, N = 0, \dots, 9 \\ a, b = 0, \dots, 7 \\ i, j = 8, 9 \end{matrix}$$

Embedding in String Theory

non-Abelian DBI action

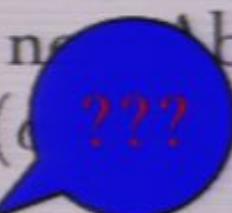
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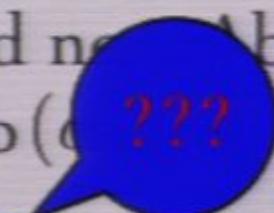
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Erdmenger, Kamiski, PK, Rust 0807.2663

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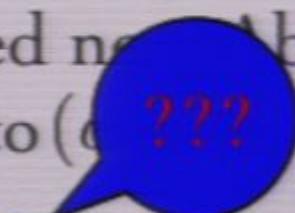
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- Need non-Abelian DBI action (best guess, correct up to α')


$$S_{\text{DBI}} = T_D \text{Str} \int d^8\xi \sqrt{\det Q} \sqrt{\det \left(P_{ab} \left[E_{MN} + E_{Mi} (Q^{-1})^{ij} E_{jN} \right] + 2\pi\alpha' F_{ab} \right)}$$

$$E_{MN} = g_{MN} + B_{MN} \quad Q_j^i = \delta_j^i + i2\pi\alpha' [\Phi^i, \Phi^k] E_{kj} \quad \begin{matrix} M, N = 0, \dots, 9 \\ a, b = 0, \dots, 7 \\ i, j = 8, 9 \end{matrix}$$

- Use symmetries of our setup

Erdmenger, Kamiski, PK, Rust 0807.2663

$$\Phi^9 = 0 \Rightarrow [\Phi^i, \Phi^j] = 0 \Rightarrow Q_j^i = \delta_j^i$$

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Looks like
Abelian action

String Theory Embedding Evaluation of Str

- Str prescription only correct up to fourth order
- We use 2 different approaches:
 - 1) expand action to fourth order
 - +: include maximal number of terms we can trust
 - : approximation breaks down in superfluid phase
 - 2) adapt Str prescription: set $(\tau^i)^2 = 1$ inside Str
 - +: can handle **all** terms
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- give the same physics!!!*

Embedding in String Theory Comparison to Gravity Model

- D7-branes are probes in D3-brane background.
Background determined by Type IIB SUGRA

$$S_{\text{IIB}} \supset N_c^2 \int d^5x \sqrt{-g} R$$

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- Action is similar to EYM action $\alpha \propto \sqrt{N_f/N_c}$ here $= 0$
with back-reaction dilaton have to be considered, too

Hydrodynamics



Fantom XP

- Effective Theory for slowly varying perturbations about the equilibrium $\omega l_{\text{mfp}} \ll 1, ql_{\text{mfp}} \ll 1$
- EOMs: conservation laws for conserved quantities
$$\nabla_\mu T^{\mu\nu} = 0, \quad \nabla_\mu J^\mu = 0$$
- Constitutive equations:
Dependence on dynamical fields, e.g. velocity u^μ
- Defines transport coefficients, e.g. viscosity
Determinable by microscopic theory

Fluctuations about the black hole with vector hair

- Consider fluctuations in metric and gauge field

$$\hat{g}_{\mu\nu}(t, x, y, r) = g_{\mu\nu}(r) + h_{\mu\nu}(r)e^{-i\omega t + iq_x x + iq_y y}$$

$$\hat{A}_\mu^a(t, x, y, r) = A_\mu^a(r) + a_\mu^a(r)e^{-i\omega t + iq_x x + iq_y y}$$

- Gauge fixing

$$\hat{g}_{\mu r} = \frac{\delta_{\mu r}}{N(r)} \quad \Rightarrow \quad h_{\mu r} = 0$$

$$\hat{A}_r^a = 0 \quad \Rightarrow \quad a_r^a = 0$$

Classification of Fluctuations

- Finite transverse momentum $q_y \neq 0$
 - ⇒ Rotational symmetry broken down to \mathbb{Z}_2
 - ⇒ all fluctuations couple 10+12 dynamical and 5+3 constraints ⇒ 14 physical modes

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- Finite longitudinal momentum $q_x \neq 0$ or no momentum ⇒ $SO(2)$ Rotational symmetry left
 - Helicity 2: $h_{yz}, h_{yy} - h_{zz}$ 2 physical
 - Helicity 1: $h_{ty}, h_{xy}, h_{yr}; a_y^a$ 2x4 physical
 - Helicity 0: $h_{tt}, h_{tx}, h_{xx}, h_{yy} + h_{zz}, h_{tr}, h_{xr}, h_{rr}; a_t^a, a_x^a, a_r^a$ 4 physical
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Kovtun, Son, Starinets;
Buchel, Liu; Liu, Iqbal,...

- Effective action for $\phi = h_z^y$ (minimal coupled scalar)

$$S = -\frac{1}{2\kappa^2} \int d^5x \sqrt{-g} \frac{1}{2} (\nabla\phi)^2$$

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Helicity 1 at zero momentum

- The 4 physical modes decouple into two blocks (1+3)

$$h'_{xy} = 2h_{xy} \left(\frac{1}{r} + \frac{f'}{f} \right) - 2\alpha^2 \phi' a_y^3$$

1st block:

$$a_y^{3\nu} + a_y^{3\nu} \left(\frac{1}{r} - \frac{2f'}{f} + \frac{N'}{N} + \frac{\sigma'}{\sigma} \right) + a_y^3 \left(-\frac{f^4 w^2}{r^2 N} + \frac{\omega^2}{N^2 \sigma^2} - \frac{2\alpha^2 \phi'^2}{N \sigma^2} \right) = 0$$

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2nd block:

$$\begin{aligned} a_y^{1\nu} = & \frac{2i\omega a_y^2 \phi}{N^2 \sigma^2} + a_y^1 \left(-\frac{\omega^2}{N^2 \sigma^2} - \frac{\phi^2}{N^2 \sigma^2} \right) + \frac{f^4 h'_{xy} w'}{r^2} + h_{xy} \left(-\frac{2f^4 w'}{r^3} - \frac{2f^3 f' w'}{r^2} \right) \\ & + a_y^{1\nu} \left(-\frac{1}{r} + \frac{2f'}{f} - \frac{N'}{N} - \frac{\sigma'}{\sigma} \right) \\ a_y^{2\nu} = & -\frac{2i\omega a_y^1 \phi}{N^2 \sigma^2} + \frac{i\omega f^4 h_{xy} w \phi}{r^2 N^2 \sigma^2} + a_y^2 \left(\frac{f^4 w^2}{r^2 N} - \frac{\omega^2}{N^2 \sigma^2} - \frac{\phi^2}{N^2 \sigma^2} \right) + a_y^{2\nu} \left(-\frac{1}{r} + \frac{2f'}{f} - \frac{N'}{N} - \frac{\sigma'}{\sigma} \right) \end{aligned}$$

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$$a_y^{2\mu} = -\frac{2i\omega a_y^1 \phi}{N^2 \sigma^2} + \frac{i\omega f^4 h_{xy} w \phi}{r^2 N^2 \sigma^2} + a_y^2 \left(\frac{f^4}{r^2 N} - \frac{N'}{f} - \frac{\sigma'}{N} \right)$$

conductivity

- 1st block $\Rightarrow G^R = \begin{pmatrix} \langle J_3^y J_3^y \rangle & \langle J_3^y T_{ty} \rangle \\ \langle T_{ty} J_3^y \rangle & \langle T_{ty} T_{ty} \rangle \end{pmatrix}$

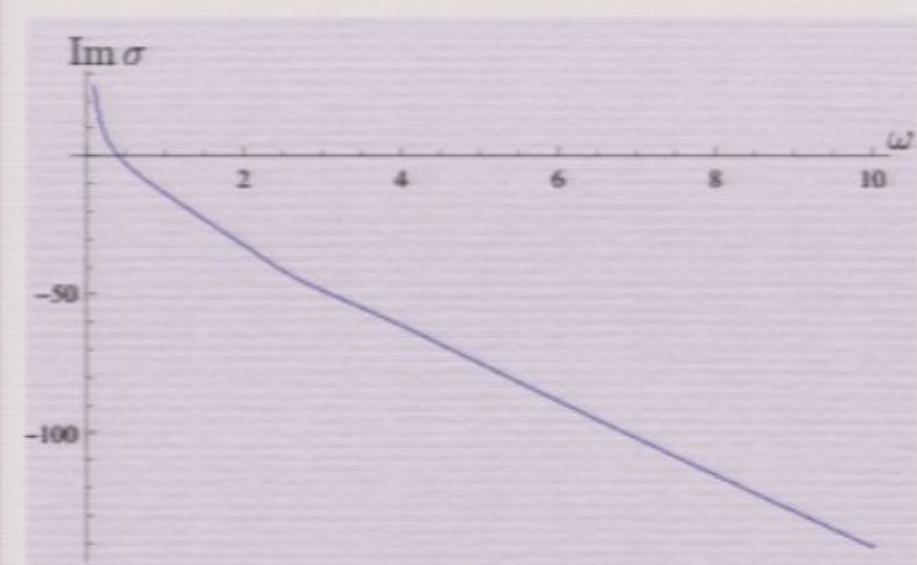
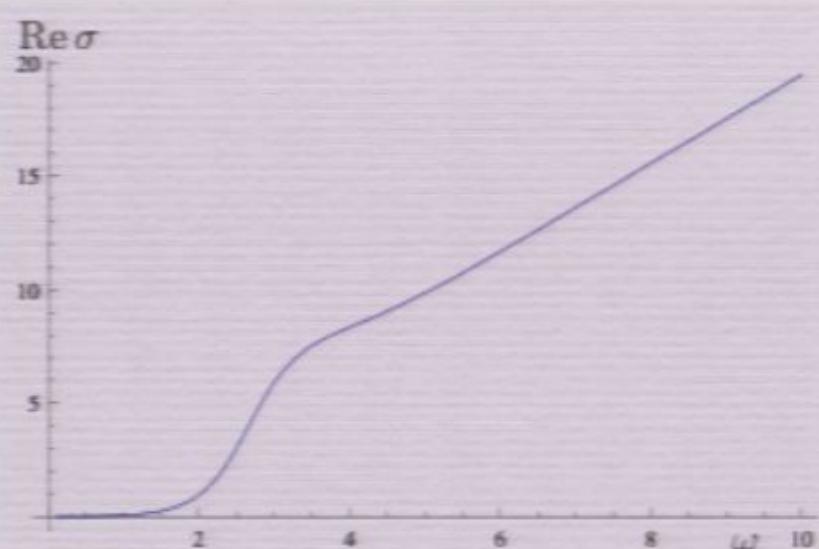
shear

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viscosity

Electric Conductivity

- Kubo relation: $\sigma(\omega) = \frac{i}{\omega} \langle J_3^y J_3^y \rangle$



Viscosity in general fluids

- Viscosity refers to dissipation due to internal motion

$$T_{\text{diss}}^{ij} = -\eta^{ijkl} (\partial_k u_l + \partial_l u_k) \quad [\text{Landau, Lifshitz}]$$

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- In general 21 independent components

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$$T_{\text{diss}}^{ij} = -\eta^{ijkl} (\partial_k u_l + \partial_l u_k) \quad [\text{Landau, Lifshitz}]$$

- η^{ijkl} is a rank four tensor with symmetries

$$\eta^{ijkl} = \eta^{jikl} = \eta^{ijlk} = \eta^{klij}$$

- In general 21 independent components
- In isotropic fluid only 2 independent components

$$\eta_{xxxx} = \eta_{yyyy} = \eta_{zzzz} = \xi + \frac{4\eta}{3}$$

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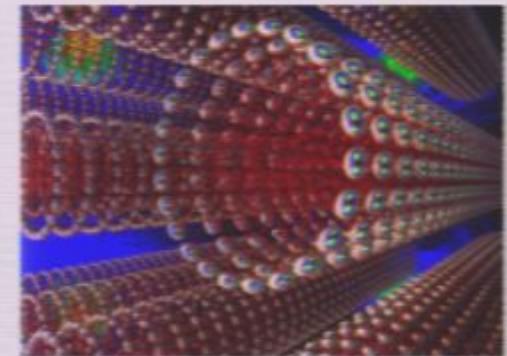
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Viscosity in transversely isotropic fluids



ISIS Facility

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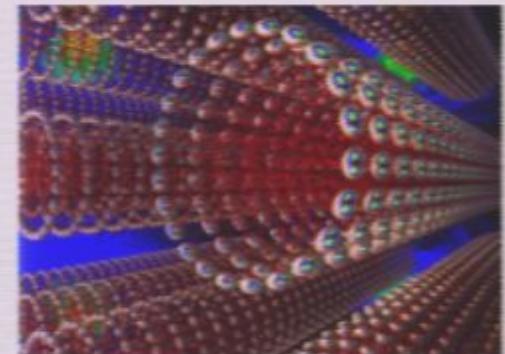
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Shear Viscosity from Gravity

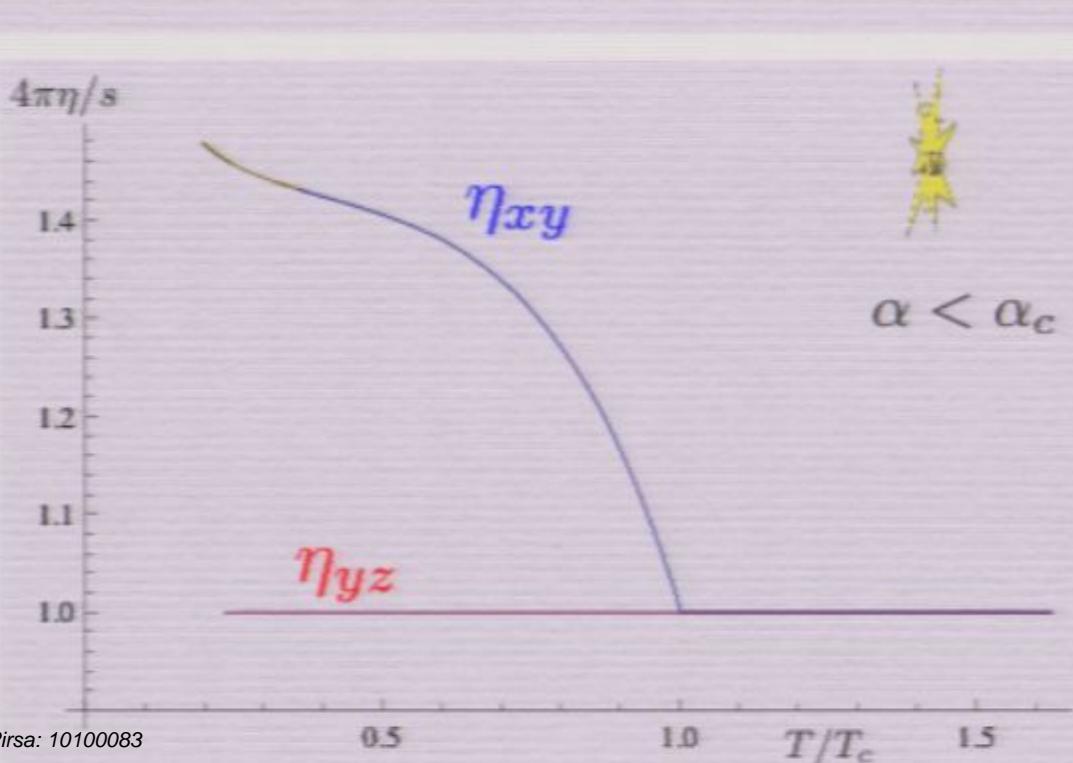
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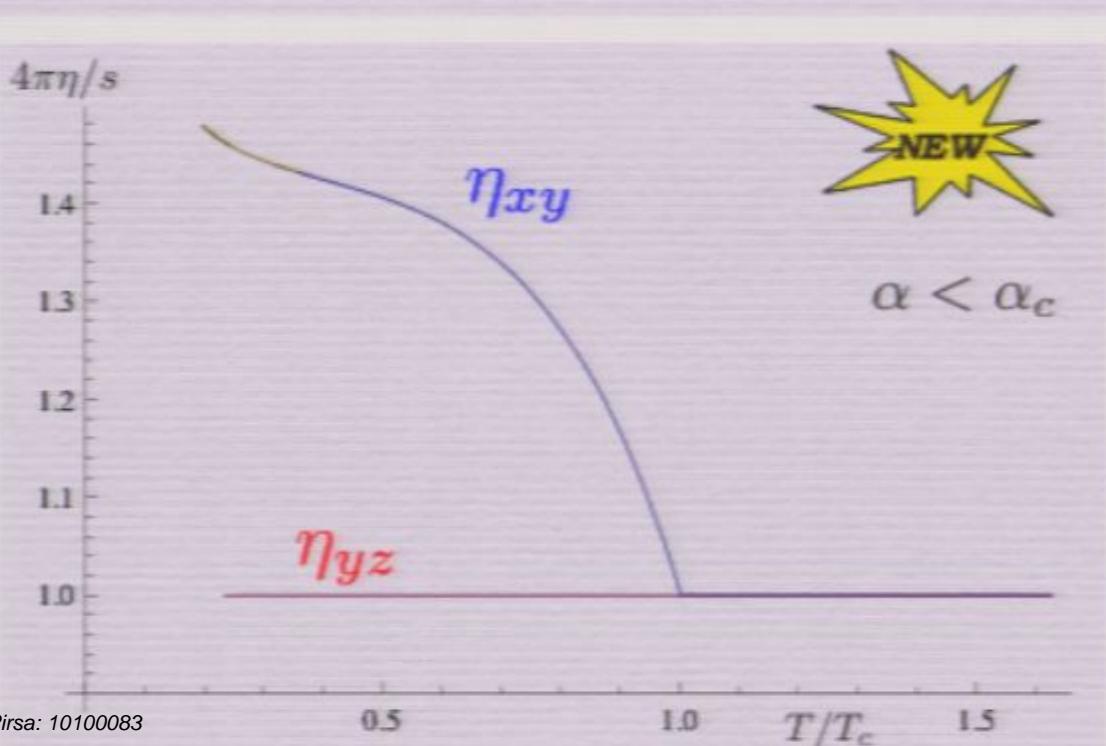


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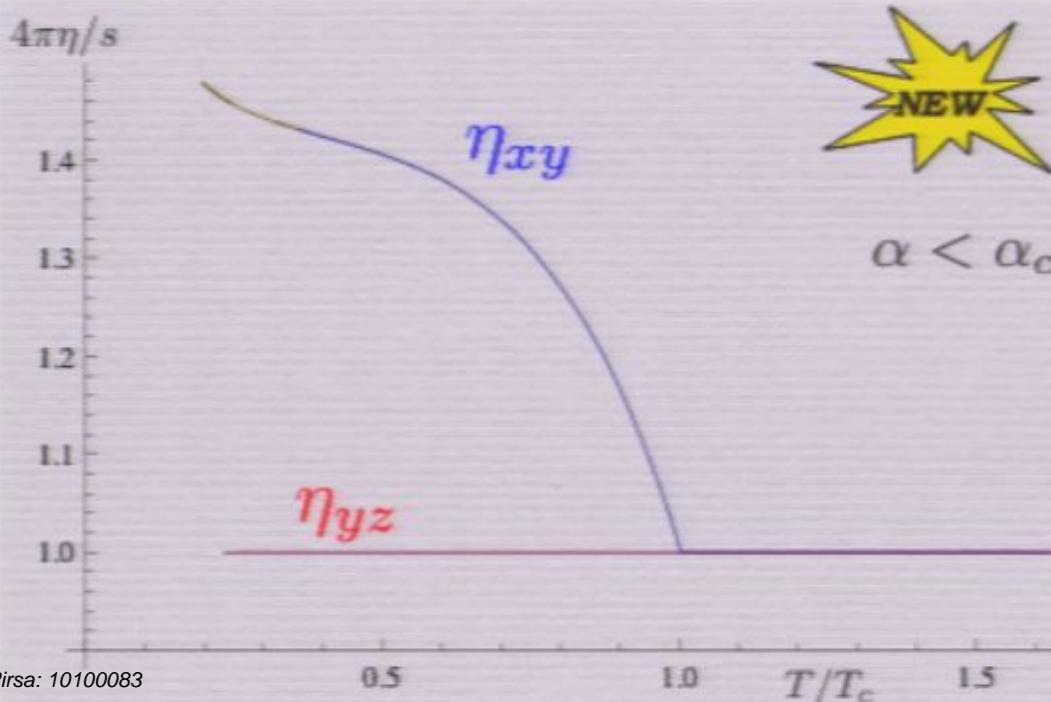


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Helicity 1 at zero momentum

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1st block:

$$h'_{xy} = 2h_{xy} \left(\frac{1}{r} + \frac{f'}{f} \right) - 2\alpha^2 \phi' a_y^3$$

$$a_y^{3\mu} + a_y^{3\nu} \left(\frac{1}{r} - \frac{2f'}{f} + \frac{N'}{N} + \frac{\sigma'}{\sigma} \right) + a_y^3 \left(-\frac{f^4 w^2}{r^2 N} + \frac{\omega^2}{N^2 \sigma^2} - \frac{2\alpha^2 \phi'^2}{N \sigma^2} \right) = 0$$

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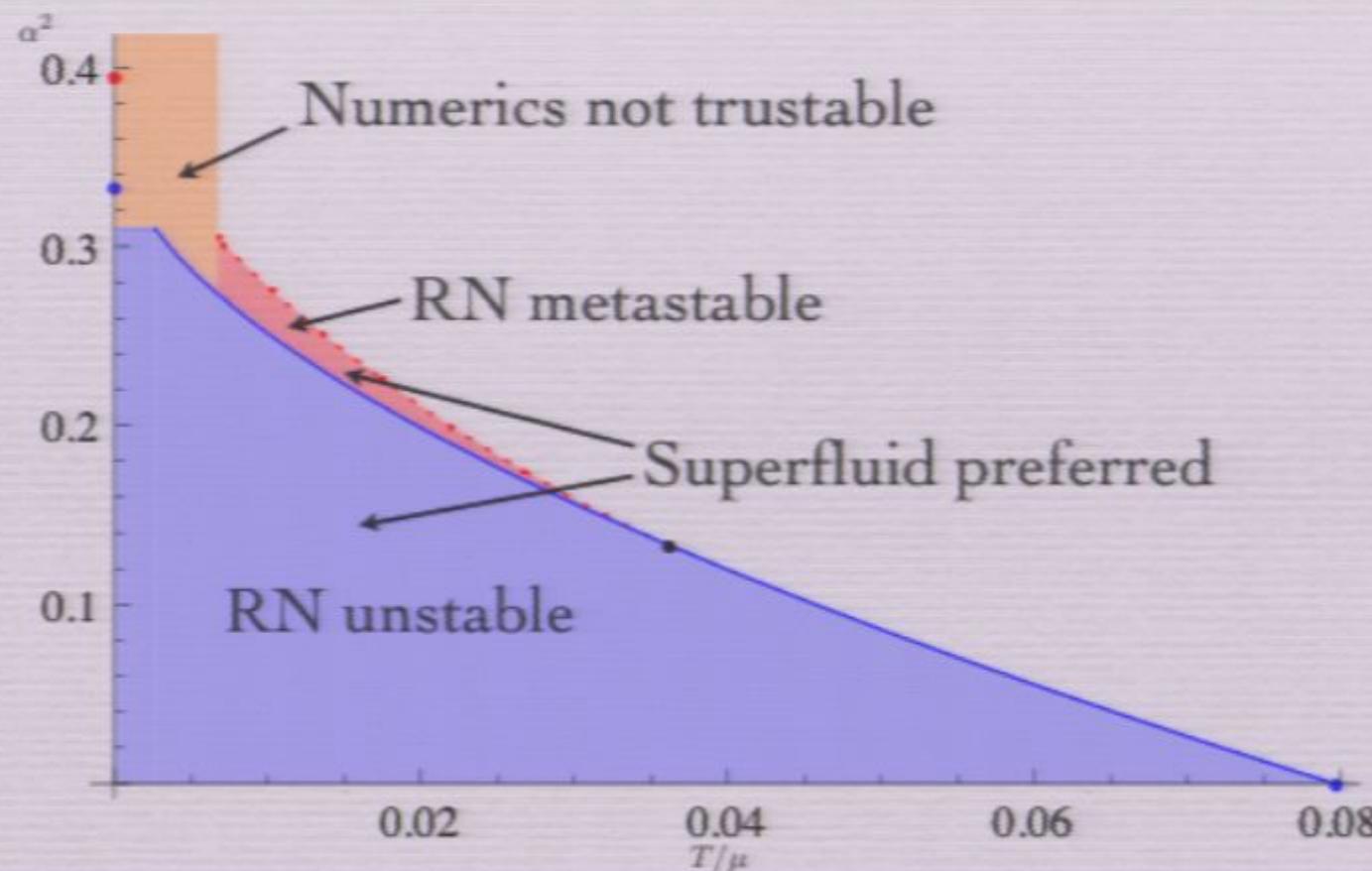
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Embedding in String Theory Comparison to Gravity Model

- D7-branes are probes in D3-brane background.
Background determined by Type IIB SUGRA

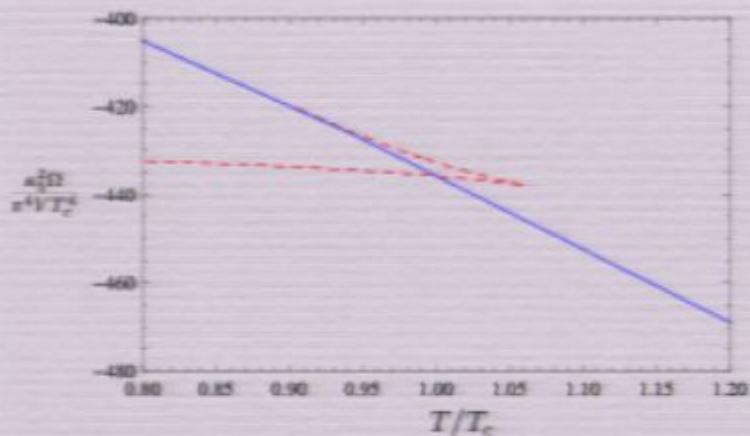
$$S_{\text{IIB}} \supset N_c^2 \int d^5x \sqrt{-g} R$$

“Phase diagram”



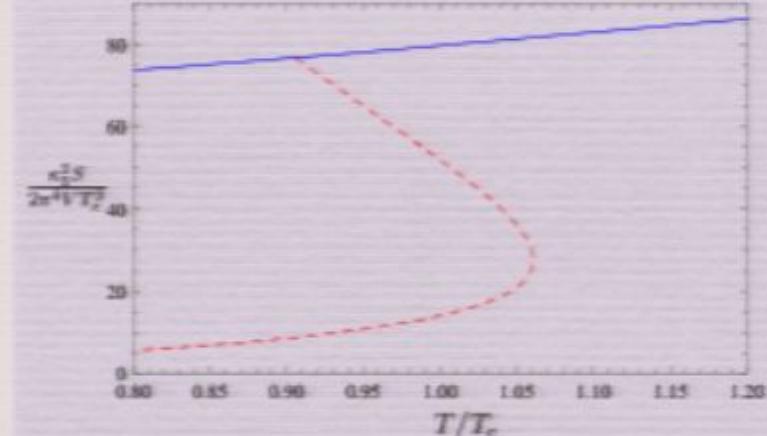
Thermodynamics

$$\alpha > \alpha_c$$



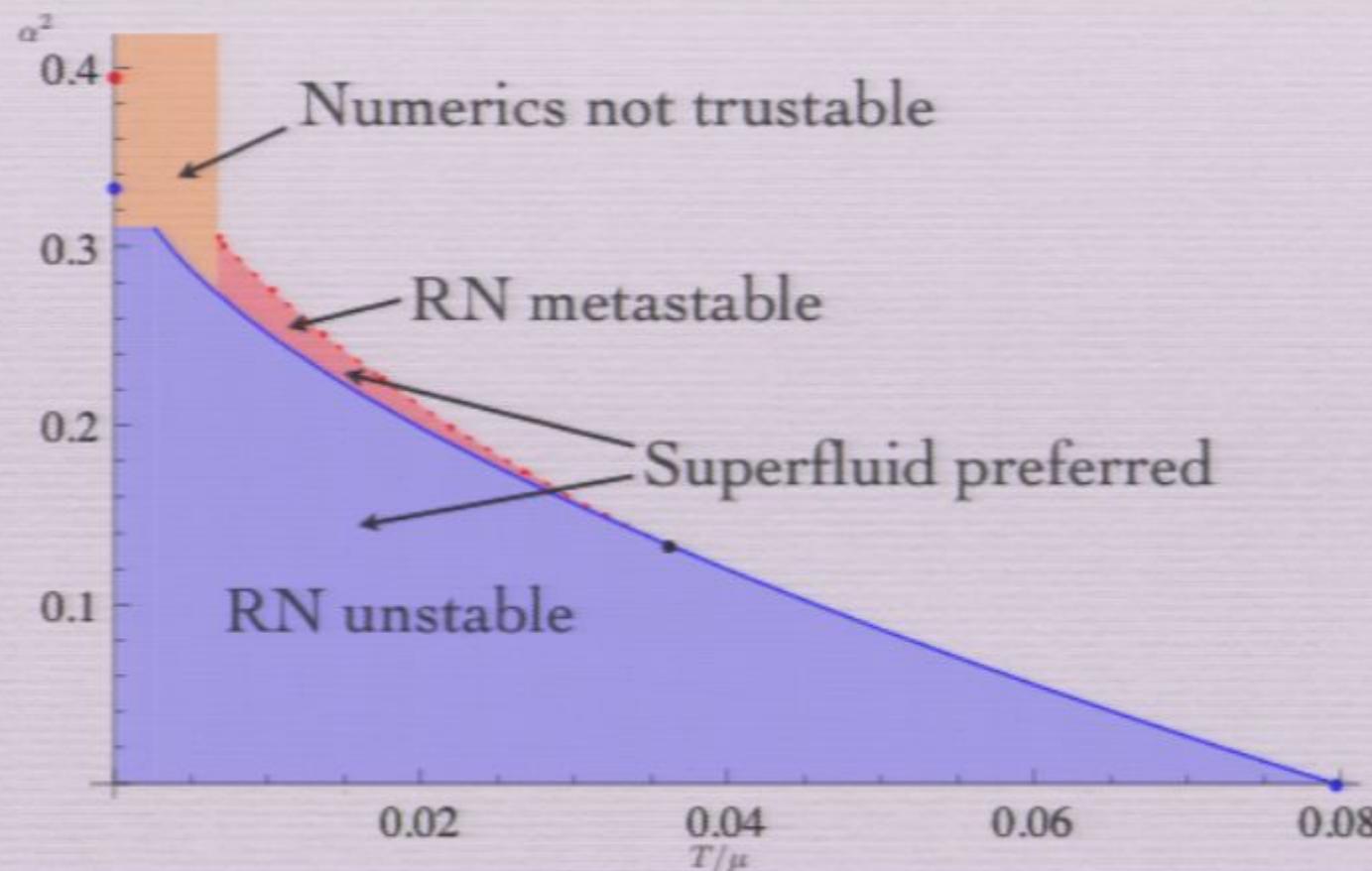
Grand potential

normal phase
superfluid phase



Entropy

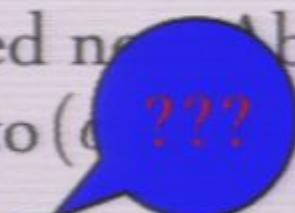
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Embedding in String Theory

non-Abelian DBI action

- Need non-Abelian DBI action (best guess, correct up to α')


$$S_{\text{DBI}} = T_D \text{Str} \int d^8\xi \sqrt{\det Q} \sqrt{\det \left(P_{ab} \left[E_{MN} + E_{M\bar{i}} (Q^{-1} - \delta)^{ij} E_{jN} \right] + 2\pi\alpha' F_{ab} \right)}$$

$$E_{MN} = g_{MN} + B_{MN} \quad Q_j^i = \delta_j^i + i2\pi\alpha' [\Phi^i, \Phi^k] E_{kj} \quad \begin{matrix} M, N = 0, \dots, 9 \\ a, b = 0, \dots, 7 \\ i, j = 8, 9 \end{matrix}$$

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Classification of Fluctuations

- Finite transverse momentum $q_y \neq 0$
 - ⇒ Rotational symmetry broken down to \mathbb{Z}_2
 - ⇒ all fluctuations coupled to 12 dynamical and 5+3 constraints
Complicated! No results yet! ⇒ 14 physical modes

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 - ⇒ all fluctuations ~~complicated!~~ **No results yet!** 12 dynamical and 5+3 constants \rightarrow 14 physical modes
- Finite longitudinal momentum $q_x \neq 0$ or no momentum ⇒ **trivial!** rotational symmetry left
 - Helicity 2: $h_{yz}, h_{yy} - h_{zz}$ **new!** 2 physical
 - Helicity 1: $h_{ty}, h_{xy}, h_{yr}; a_y^a$ 2x4 physical
 - Helicity 0: $h_{tt}, h_{tx}, h_{xx}, h_{yy} + h_{zz}, h_{tr}, h_{xr}, h_{rr}; a_t^a, a_x^a, a_r^a$ 4 physical
 - total: 14 physical

Helicity 2 (review)

Kovtun, Son, Starinets;
Buchel, Liu; Liu, Iqbal,...

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- Viscosity is given by

$$\eta = \lim_{\omega \rightarrow 0} \lim_{r \rightarrow \infty} \frac{\Pi}{i\omega\phi} \quad \Pi = -\frac{1}{2\kappa^2} \sqrt{-g} g^{rr} \partial_r \phi$$

- For $\omega \ll 1$ EoMs trivial

$$\partial_r \Pi = 0 + \mathcal{O}(\omega^2 \phi) \quad \partial_r(\omega \phi) = 0 + \mathcal{O}(\omega \Pi)$$

- Ingoing boundary condition at horizon

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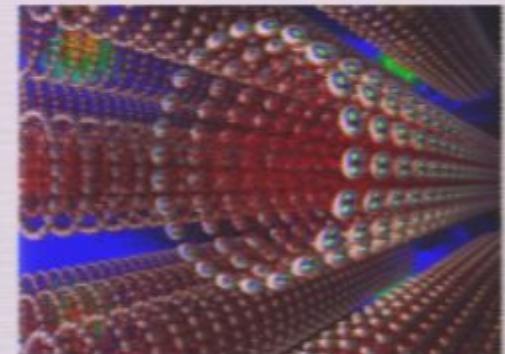
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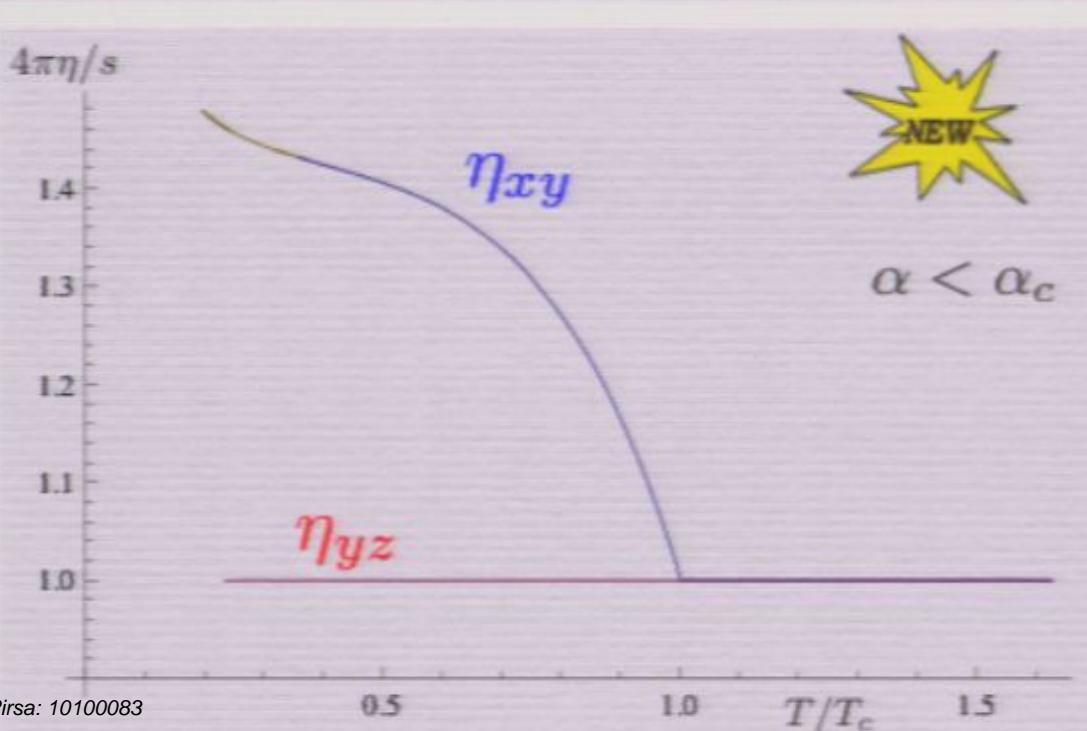
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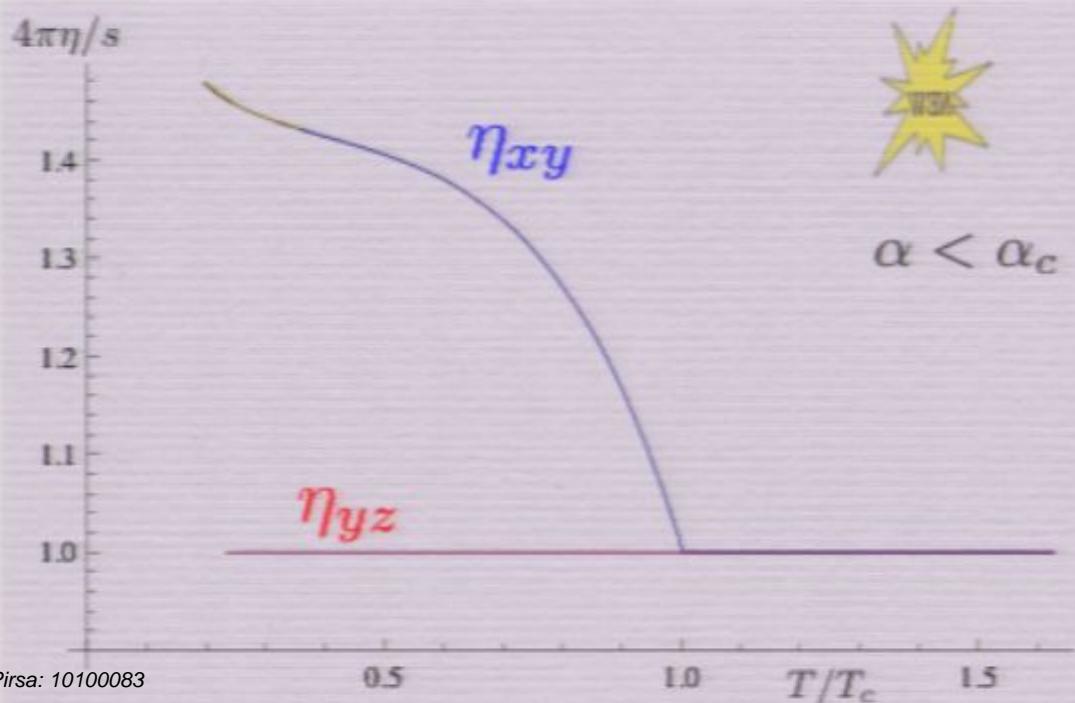
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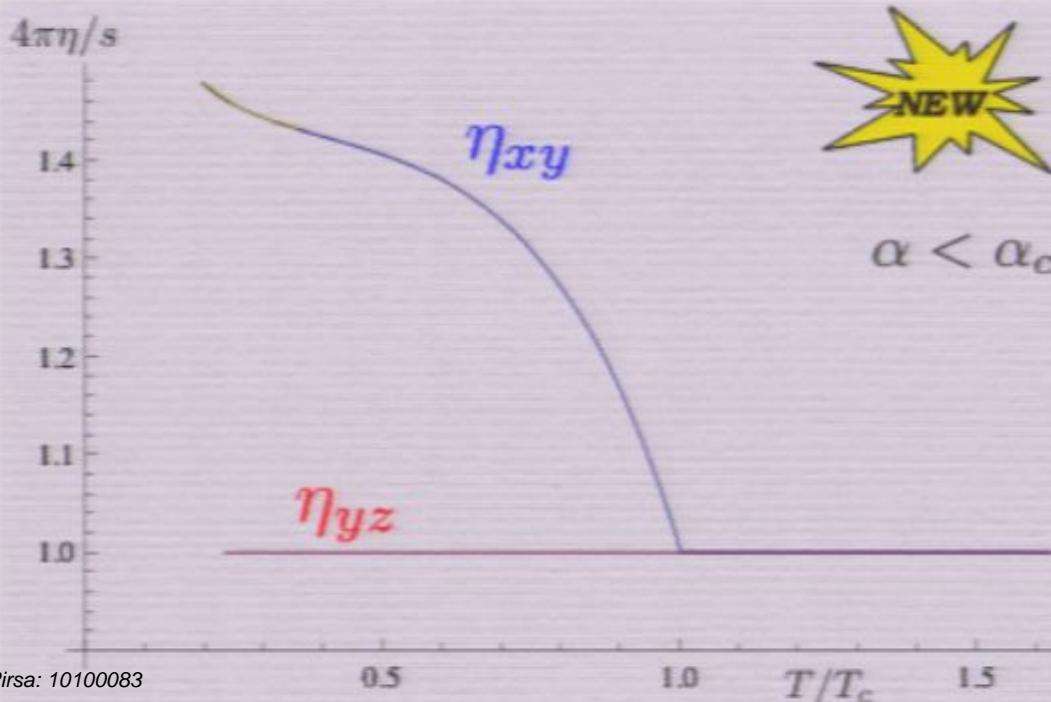


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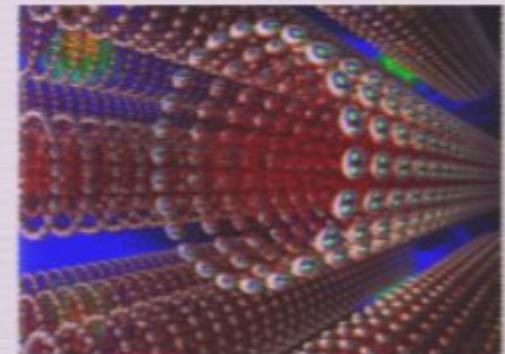
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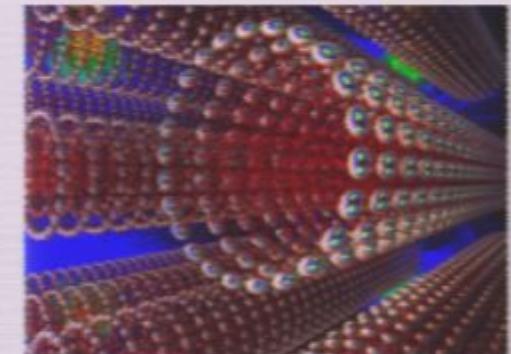
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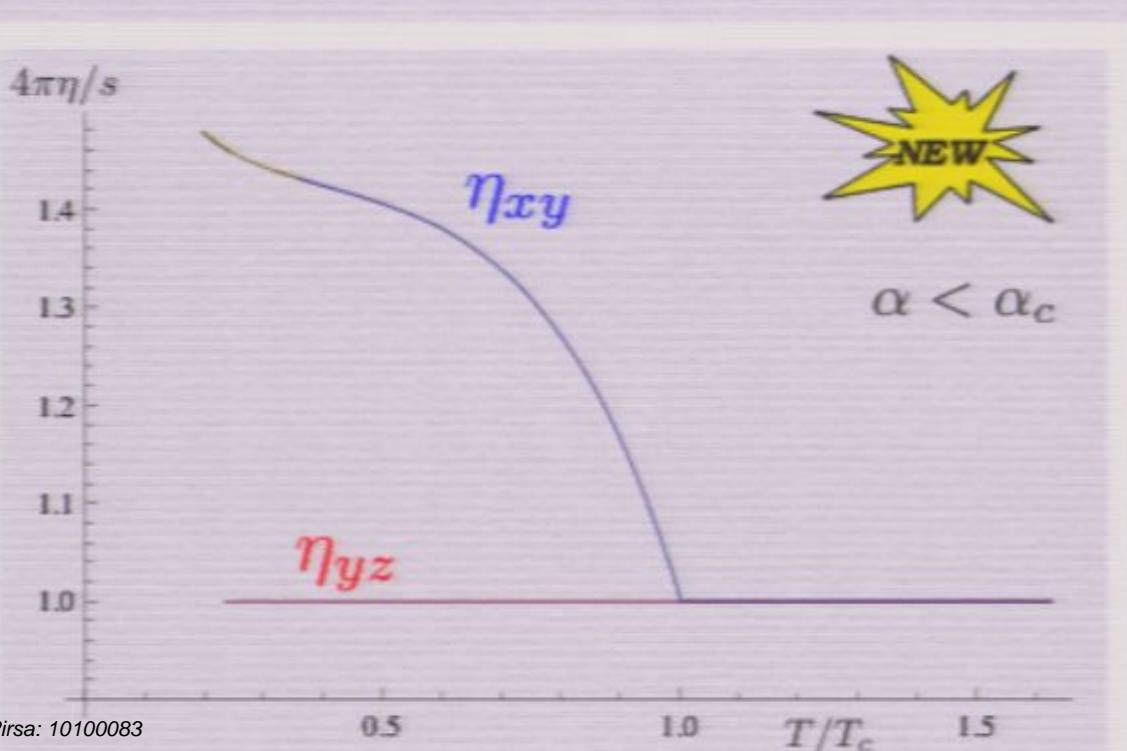
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Thank you!