

Title: Strong Gravity

Date: May 06, 2010 01:00 PM

URL: <http://pirsa.org/10050000>

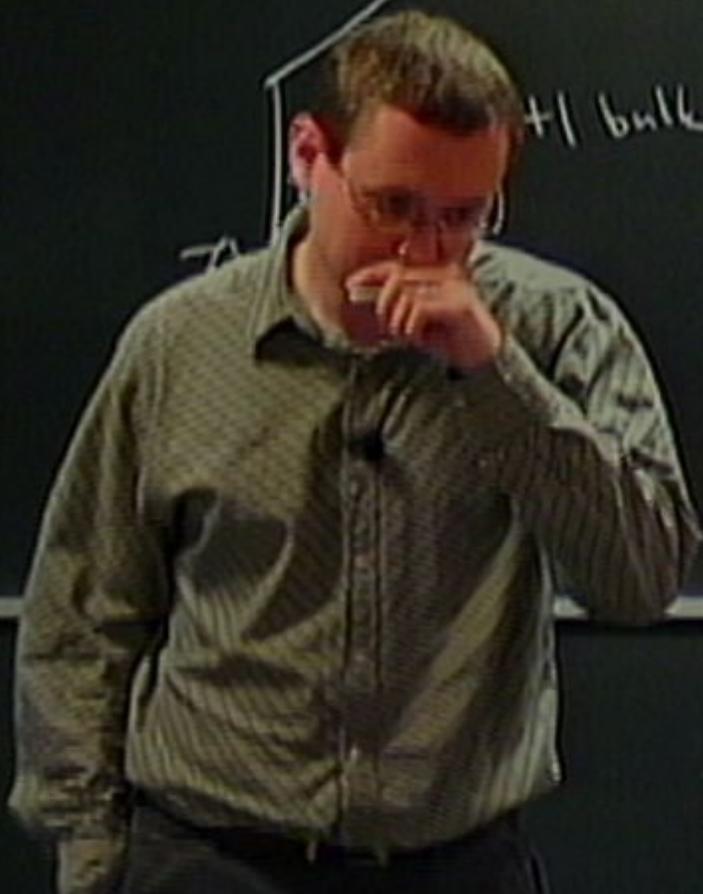
Abstract: TBA

Kaus & HSR 0901.4236

The RS2 mode

Bulk: 5d FR, -ve Λ

+1 bulk



The RS2 mode

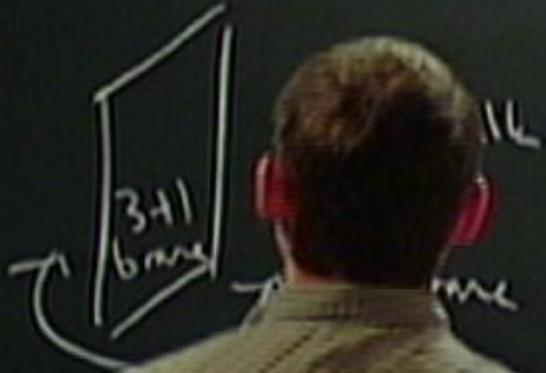


++1 bulk

Bulk: 5d FR, -ve Λ

AdS₅ solⁿ, radius $\frac{1}{\alpha}$ $\Lambda \sim -\frac{1}{\alpha^2}$

The RS2 model



Bulk: 5d FR, -ve Λ

AdS₅ solⁿ, radius l $\Lambda \sim -\frac{1}{2l^2}$

action: $S = \int d^4x \sqrt{-g} (-\sigma + \mathcal{L}_{matter})$

The RS2 model



Bulk: 5d GR, -ve Λ

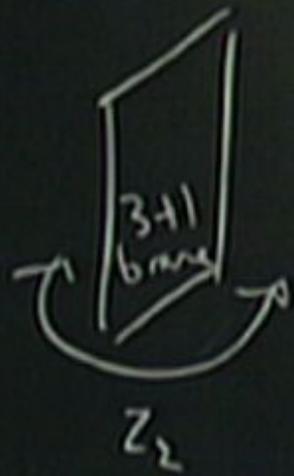
AdS₅ solⁿ, radius $\frac{1}{\alpha}$ $\Lambda \sim -\frac{1}{2L^2}$

brane: $S = \int d^4x \sqrt{-h} (-\sigma + \mathcal{L}_{matter})$

grav. backreaction via Israel eq

$K_{ablm} \sim \sigma_{ab} + \mathcal{L}_{matter}$

The RS2 model



4+1 bulk

Bulk: 5d GR, -ve Λ

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grav. backreaction via Israel eq

$K_{ablm} \sim \sigma_{ab} + \mathcal{L}_{matter}$

$$d \cdot \frac{\sum_{ij} dx^i dx^j + dz^k}{(z + z_0)^2} .$$

brane @ $z = 0$.

$$ds^2 = \frac{\eta_{ij} dx^i dx^j + dz^2}{(z+z_0)^2}$$

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$$ds^2 = \frac{\eta_{ij} dx^i dx^j + dz^2}{(z+z_0)^2} \quad \text{brane @ } z=0.$$

$$\sigma = \frac{3}{4\pi G_5 \ell} \quad (\text{RS value for } \sigma)$$

$$ds^2 = \frac{\eta_{ij} dx^i dx^j + dz^2}{(12l + z_0)^2} \quad \text{brane @ } z=0.$$

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Part theory: for $r \gg l$ gravity of brane obeys inverse square law

$$ds^2 = \frac{\eta_{ij} dx^i dx^j + dz^2}{(l^2 + z^2)^2} \quad \text{brane @ } z=0.$$

$$\sigma = \frac{3}{4\pi G_5 l} \quad (\text{RS value for } \sigma)$$

Part theory: for $r \gg l$ gravity of brane obeys inverse square law with $G_4 = G_5/l$.

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Part theory: for $r \gg l$ gravity of brane obeys inverse square law with $G_4 = G_5/l$.

- ∞ 5th dimension!

gravity - for $l \gg \lambda$ gravity on brane obeys inverse square law with $G_4 = G_5/l$.

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What about nonlinear 4d GR?

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RS black holes

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RS black holes



Many attempts to find solⁿ (exact or numerical)

for $l \gg \lambda$ gravity on brane obeys inverse square law with $G_4 = G_5/l$.

- ∞ 5th dimension!

What about nonlinear 4d GR?

RS black holes



Many attempts to find solⁿ (exact or numerical) describing RS2 BH with $G_4 M \gg l$.

AdS/CFT \Rightarrow S^2 is equivalent to 4d off. theory
Gibbs matter + large N , strongly coupled CFT

AdS/CFT \Rightarrow RS2 is equivalent to 4d eff. theory
GR + matter + large N , strongly coupled CFT

AdS/CFT \Rightarrow RSL is equivalent to 4d eff. theory

GR + matter + large N , strongly coupled CFT
with cut-off $\sim l^{-1}$.

4d effective Einstein eq: $R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = 8\pi G_N \langle T_{\mu\nu}^{CF} \rangle$

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responsible for deviation from GR.

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RS2 black holes are quantum corrected BHs from 4d perspective

Sphere

5th dimension

RS2 black holes are quantum corrected BHs from 4d perspective.

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A brane BH would Hawking radiate into $\mathcal{E}T \Rightarrow$ time-dependent!

Tanaka '02.

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lower dimensions: 2+1 brane in 3+1 bulk. Tanaka '02.

\exists exact brane BH solⁿs

AdS/CFT \Rightarrow RSL is equivalent to 4d eff. theory

GR + matter + large N , strongly coupled CFT
with cut-off $\sim l^{-1}$.

4d effective Einstein eq: $R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \left[\overset{8\pi G}{\cancel{T}}^{\text{matter}} + \langle T_{\mu\nu}^{\text{CFT}} \rangle \right]$
responsible for deviation from GR.

RS2 black holes are quantum corrected BHs from 4d perspective.

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New idea:

Many attempts to find solⁿ (exact or numerical)

describing RS2 BH with $G_{4M} \gg \lambda$.

lower dimensions: 2+1 brane in 3+1 bulk
 \exists exact brane BH solⁿs

New idea: look at extreme brane BH's $T=0$ \Rightarrow no Hawking radⁿ
 \Rightarrow Future-integr^l solⁿs.

pts to find solⁿ (exact or numerical)

BH with $G_{4M} \gg l$.

Exact base BH solⁿs

New idea: look at extreme base BH's $T=0$ \Rightarrow no Hawking radⁿ
 \Rightarrow finite-integ solⁿs.

What about non-linear 4d GR?

RS black holes



Many attempts to find solⁿ (exact or numerical)
describing RS2 BH with GEM \gg L.

e.g. Maxwell field on brane: look for analogue of extreme RN.
bulk fields depend on 2 coords

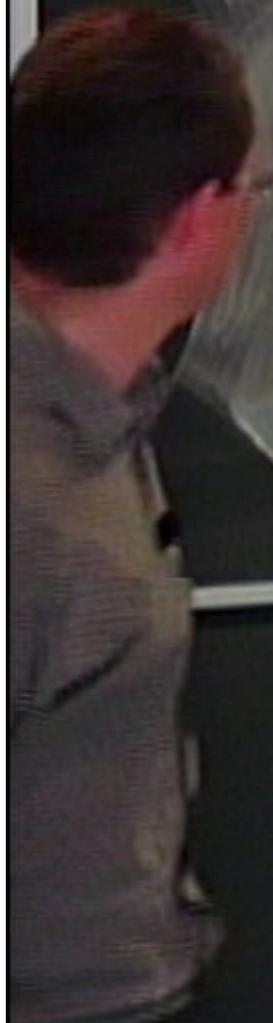
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Instead: determine near-horizon geometry of extreme charged brane (B.H.).

e.g. extreme RN \rightarrow AdS₂ x S⁵



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From n.h. geom. can determine $S(Q)$
radius Q (charge of BH).

$$4\pi G_N : S_6 = \frac{\pi Q^2}{4\pi}$$

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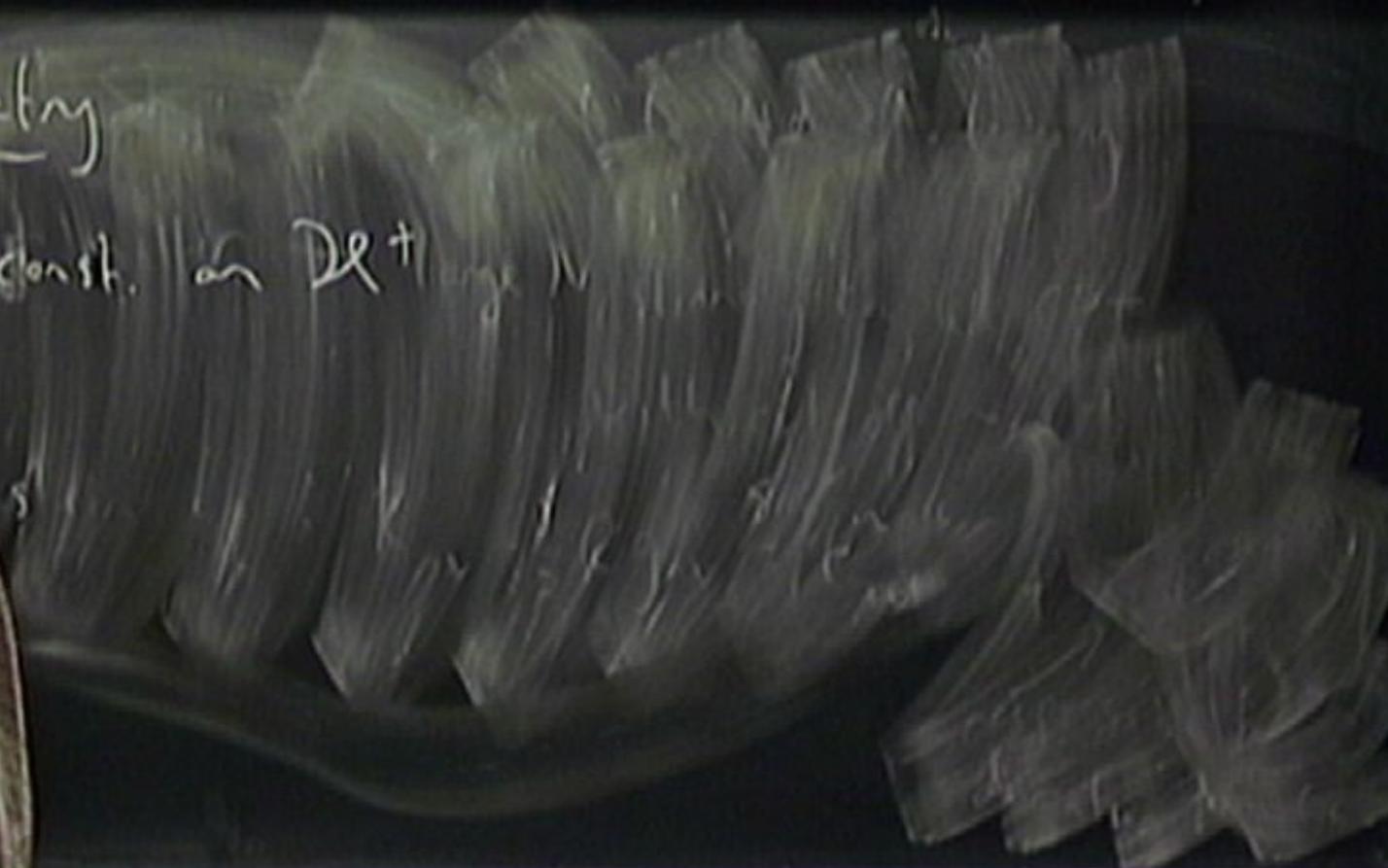
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New-L geometry

O4L $R = \text{const.}$ on $D \times \mathbb{R}^4$



New-Horizon geometry

- $K = \text{const.}$ on $\mathbb{R}^+ \times \mathbb{N}$
 $\Rightarrow K = 0$ everywhere if $K < 0$ on brane.

Non-constant geometry

- $\Rightarrow K = \text{const. on } \mathbb{R}^+$
- $\Rightarrow K = 0$ everywhere if $K < 0$ on some.
- $\Rightarrow \exists$ non-constant limits.

New-horizon geometry

$O(4)_h \Rightarrow K = \text{const. on } \mathbb{R}^{1,3}$

$\Rightarrow K = 0$ everywhere if $K < 0$ on brane.

$\Rightarrow \exists$ new-horizon limit.

New-horizon geometry

$\Rightarrow K = \text{const. on } \mathcal{D}^+$

$\Rightarrow K = 0$ everywhere if $K = 0$ on \mathcal{B} .

$\Rightarrow \exists$ new-horizon limit.

dim, Lucietti & HSR '07: any static new-horizon geometry form

$0 \text{th} \Rightarrow K = \text{const. on } \mathcal{X}^1$

$K=0$ everywhere if $K=0$ on brane .

\exists new-horizon limit.

Kundt metric \in HSR '07: any static new-horizon geometry

$$ds^2 = A(x)^2 d\Sigma_2^2$$

$\Rightarrow K=0$ everywhere if $K=0$ on brane.

$\Rightarrow \exists$ near-limit.

Kunduri, Lucietti & HSR
any static new-horizon geometry
has form $ds^2 = -dt^2 + g_{ab}(x) dx^a dx^b$.

$\Rightarrow R=0$ everywhere if $k=0$ on brane.

$\Rightarrow \exists$ near-horizon limit.

Kunduri, Lucietti & HSR '07: any static near-horizon geometry has form $ds^2 = A(x)^2 d\Sigma_{n-2}^2 + g_{ab}(x) dx^a dx^b$.

$\Rightarrow K=0$ everywhere if $K=0$ on brane.

$\Rightarrow \exists$ Newtonian limit.

Kunduri, Lucie

07: any static newtonian geometry

has form

$$= A(x)^2 d\Sigma^2 + g_{ab}(x) dx^a dx^b$$

\nearrow
describes
R^{1,3}

$\Rightarrow K=0$ everywhere if $k=0$ on brane.

$\Rightarrow \exists$ near-horizon limit.

Kundin
has

(HSR '07): any static near-horizon geometry

$$ds^2 = A(x)^2 d\Sigma^2 + g_{\mu\nu}(x) dx^\mu dx^\nu.$$

\nearrow dS_2 $k=1$
 $\mathbb{R}^{1,1}$ $k=0$ adS_2 $k=-1$.

$\Rightarrow R=0$ everywhere if $k=0$ on brane.

$\Rightarrow \exists$ new-horizon limit.

Kunduri, Lucietti & HSR '07: any static new-horizon geometry has form

$$ds^2 = A(x)^2 d\Sigma_k^2 + g_{\mu\nu}(x) dx^\mu dx^\nu.$$

$\begin{matrix} \nearrow \\ \text{deS}_2 \\ \mathbb{R}^{1,1} \end{matrix}$ $\begin{matrix} k=1 \\ k=0 \end{matrix}$ $\text{adS}_2 \quad k=-1.$

case $k=1$
 R^2 $k=0$ adS_2 $k=-1$.

e.g. Maxwell field on brane: look for analogue of extreme RN.

bulk solⁿ: depend on 2 coords (numerics?)

Instead: determine near-horizon geometry of extreme charged brane (B.H.).

HSR 0901.4236

$U(1) \times U(1)$ $k=1$
 $U(1) \times U(1)$ $k=0$ adS_2 $k=-1$.

Assume sph. sym on brane S^2 acting S^2

$$ds^2 = A(x)^2 d\xi^2 + \dots$$

Kruskal & HSR

why?

adS_2 $k=1$
 R^2 $k=0$ adS_2 $k=-1$.

Assume sph. sym on brane $SO(3)$ acting S^2

$$ds^2 = A(x)^2 d\xi_k^2 + dr^2 + R(r)^2 d\Omega^2$$

Kruskal & HSR 0901.4236

why?

adS_2 $k=1$
 R^2 $k=0$ adS_1 $k=-1$.

A time sph. sym on brane $SO(3)$ acting S^2

$$A(x^3) d\xi_k^2 + \underbrace{d\rho^2 + R(\rho)^2 d\Omega^2}_{\text{spatial geom. of } \mathbb{R}^3}$$

& HSR 0901.4236

why?

adS_2 $k=1$ adS_1 $k=-1$

Assume sph sym on brane $SO(3)$ acting S^2

$$ds^2 = A(x) \left[dt^2 + R(x)^2 d\Omega^2 \right]$$

spatial geom. of $2t$

Kaus 01.4236

wi

$U(1) \times U(1)$ $k=1$ $k=0$ adS_2 $k=-1$.

Assume sym on brane $SO(3)$ acting S^2

$$ds^2 = \dots + \underbrace{d\rho^2 + R(\rho)^2 d\Omega^2}_{\text{spatial gen. of } 2\text{d}}$$

ρ to vanish somewhere wlog @ $\rho=0$.

0901.4236

$U(1) \times U(1)$ $k=1$
 $SO(2,1)$ $k=0$ adS_2 $k=-1$.

Assume sph. sym on brane $SO(3)$ acting S^2

$$ds^2 = A(x^3) d\xi_k^2 + \underbrace{dp^2 + R(p)^2 d\Omega^2}_{\text{spatial gen. of } \mathbb{R}^{1,1}}$$

Indeed $R(p)$ to vanish somewhere wlog @ $p=0$.

Kruskal & HSR 0901.4236

Why?

$\frac{d\epsilon^2}{dt^2} \quad \left. \begin{matrix} k=1 \\ k=0 \end{matrix} \right\} \text{adS}_k \quad k=-1.$

Ass. $\text{SO}(3,1)$ sym on brane $\text{SO}(3,1)$ acting S^2

$$ds^2 = d\epsilon_k^2 + \underbrace{d\rho^2 + R(\rho)^2 d\Omega^2}_{\text{spatial gen. of } \mathbb{R}^{2,1}}$$

$R(\rho)$ to vanish somewhere wlog @ $\rho=0$.

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?

case $k=1$
 $R=1$ $k=0$ adS_2 $k=-1$.

Assume sph. sym on brane $SO(3)$ acting S^2

$$ds^2 = A(\rho)^2 d\xi_k^2 + \underbrace{d\rho^2 + R(\rho)^2 d\Omega^2}_{\text{spatial gen. of } \mathbb{R}^{2,1}}$$

Indeed $R(\rho)$ to vanish somewhere wlog @ $\rho=0$.

Kruskal & HSR 0901.4236

Why?

$\frac{d\epsilon^2}{R^2} \quad \frac{k=1}{k=0} \quad \text{adS}_2 \quad k=-1.$

Assume sph. sym on brane $SO(3)$ acting S^2

$$= A(\rho)^3 d\xi_k^2 + \underbrace{d\rho^2 + R(\rho)^2 d\Omega^2}_{\text{spatial geom. of } \mathbb{R}^3}$$

need $R(\rho)$ to vanish somewhere wlog @ $\rho=0$.

& HSR 0901.4236

why?

case $k=1$
 $R=1$ $k=0$ adS_2 $k=-1$.

Assume sph. sym on brane $SO(3)$ acting S^2

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need $R(\rho)$ to vanish somewhere wlog @ $\rho=0$.
nothness $\Rightarrow R(\rho) = \rho + \mathcal{O}(\rho^3)$

& HSR 0901.4236

why?

case $k=1$
 $\mathbb{R}^{1,1}$ $k=0$ adS_2 $k=-1$.

Assume sol. sym on brane $SO(3)$ acting S^2

$$ds^2 = A^2 dt^2 + \underbrace{dp^2 + R(p)^2 d\Omega^2}_{\text{spatial geom. of } \Sigma_t}$$

to vanish somewhere wlog @ $p=0$.

$$R(p) = p + O(p^3), \quad A(p) = A_0 + O(p^2)$$

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adS_2 $k=1$
 R^2 $k=0$ adS_1 $k=-1$.

Assume sph. sym on brane $SO(3)$ acting S^2

$$ds^2 = A(\rho)^2 d\xi_k^2 + \underbrace{d\rho^2 + R(\rho)^2 d\Omega^2}_{\text{spatial gen. of } \mathbb{R}^3}$$

Indeed $R(\rho)$ to vanish somewhere wlog @ $\rho=0$.
Smoothness $\Rightarrow R(\rho) = \rho + \mathcal{O}(\rho^3)$, $A(\rho) = A_0 + \mathcal{O}(\rho^2)$

Kruskal & HSR 0901.4236

Why?

Einstein eq \Rightarrow ODE's for $R(\rho), A(\rho)$

Einstein eq \Rightarrow ODE's for $R(\rho), A(\rho)$ solⁿ unique given k, A_0

Assume brane @ $\rho = \rho \Rightarrow ds_4^2 = L_1^2 d\Sigma_4^2 + L_2^2 d\Omega^2$

$$\begin{array}{ccc} \parallel & & \parallel \\ A(\rho_0)^2 & & R(\rho_0)^2 \end{array}$$

Assume Maxwell field on brane is

$$\star_4 F = Q d\Omega \quad \text{charge } Q.$$

Israel eqs:

Einstein eq \Rightarrow ODE's for $R(\rho), A(\rho)$ solⁿ unique given k, A_0 .

Assume brane @ $\rho = \rho_0 \Rightarrow ds_{41}^2 = L_1^2 d\Sigma_4^2 + L_2^2 d\Omega^2$

$$\begin{matrix} \parallel & & \parallel \\ A(\rho_0)^2 & & R(\rho_0)^2 \end{matrix}$$

Assume Maxwell field on brane is $*F = Q d\Omega$ charge Q .

Israel eqs: must keep region $\rho \leq \rho_0$.

Assume Maxwell field on base is $*F = Q d\Omega$. charge Q .

Israel eqs: must keep region $\rho \leq \rho_0$.

Israel + "Hamiltonian constraint" $\Rightarrow \frac{1}{L_1^2} + \frac{1}{L_2^2} = \frac{-2L^2 Q^2}{2L_2^2}$

From this can determine $S(Q)$ radius Q (charge of BH).

$S_6 = \frac{\pi Q^2}{4r}$

Assume Maxwell field on base is $*F = Q d\Omega$. $R(p_0)^2$. charge Q .

Israel eqs: must keep region $p \leq p_0$.

Israel + "Hamiltonian constraint" $\Rightarrow \frac{k}{L_1^2} + \frac{1}{L_2^2} = \frac{-k^2 Q^2}{2L_2^2} \Rightarrow k = -1 \Rightarrow AdS_2 \times S^2$.

n.h. geom can determine $S(Q)$ radius Q (charge of BH).

4d GR: $S_6 = \frac{\pi Q^2}{4}$

Israel eqs: must keep region $p \leq p_0$ ⁴

Israel + "Hamiltonian constraint" $\Rightarrow \frac{k}{L_1^2} + \frac{1}{L_2^2} = \frac{-l^2 Q^2}{2L_2^2} \Rightarrow k = -1$ ✓
 $\Rightarrow \text{AdS}_5 \times S^5$ ✓

but $L_1^2 < L_2^2$ (4d, 19R: $L_1 = L_2 = Q$)

Israel eqs: must keep region $\rho \leq \rho_0$ + charge Q .
Israel + "Hamiltonian constraint" $\Rightarrow \frac{k}{L_1^2} + \frac{1}{L_2^2} = \frac{-l^2 Q^2}{2L_2^2} \Rightarrow k = -1$
 $\Rightarrow \text{AdS}_5 \times S^5$ ✓

but $L_1 < L_2$ (4d, 19d: $L_1 = L_2 = Q$)

Solutions

- (i) determine bulk -1 parameter A_0
- (ii) Israel determine A_0, ρ_0 uniquely to Q .

Special case

$$: A_0 = \frac{\lambda}{2} \rightarrow A = \frac{\lambda}{2}$$

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Why?

$$ds^2 = A(r)^2 dt^2 + dr^2 + R(r)^2 d\Omega^2.$$

Special case

$$: A_0 = \frac{\rho}{2} \rightarrow A = \frac{\rho}{2} \quad R = \frac{\rho}{\sqrt{2}} \sinh \frac{\sqrt{2}\rho}{\lambda}$$

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special solⁿ

$$: A_0 = \frac{\ell}{2} \rightarrow A = \frac{\ell}{2} \quad R = \frac{\ell}{\sqrt{2}} \sinh \frac{\sqrt{2} \rho}{\ell}$$

bulk solⁿ is $AdS_2 \times H^3$.

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why?

special solⁿ

$$: A_0 = \frac{\ell}{2} \rightarrow A = \frac{\ell}{2} \quad R = \frac{\ell}{\sqrt{2}} \sinh \frac{\sqrt{2} \rho}{\ell}$$

bulk solⁿ is $AdS_2 \times H^3$.

Israel eq: $\sinh \frac{\sqrt{2} \rho_0}{\ell} = 1$

Kruskal & HSR 0901.4236

Why?

Special solⁿ

$$A_0 = \frac{\ell}{2} \rightarrow A = \frac{\ell}{2} \quad R = \frac{\ell}{\sqrt{2}} \sinh \frac{\sqrt{2} \rho}{\ell}$$

solⁿ is $AdS_2 \times H^3$.

$$\sinh \frac{\sqrt{2} \rho_0}{\ell} = 1 \quad Q = \frac{\ell}{\sqrt{2}}$$

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Special solⁿ

$$A_0 = \frac{\ell}{2} \rightarrow A = \frac{\ell}{2} \quad R = \frac{\ell}{\sqrt{2}} \sinh \frac{\sqrt{2} p}{\ell}$$

bulk solⁿ is $AdS_2 \times H^3$.

Israel eq: $\frac{\ell}{\sqrt{2}} \sinh \frac{\sqrt{2} p_0}{\ell} = 1 \quad Q = \frac{\ell}{\sqrt{2}}$

$$\frac{\ell}{\sqrt{2}} Q \quad L_2 = Q.$$

Si... vanish somewhere wlog @ $p=0$.
 $R(p) = p + O(p^3), \quad A(p) = A_0 + O(p^2)$

K... 09... 236

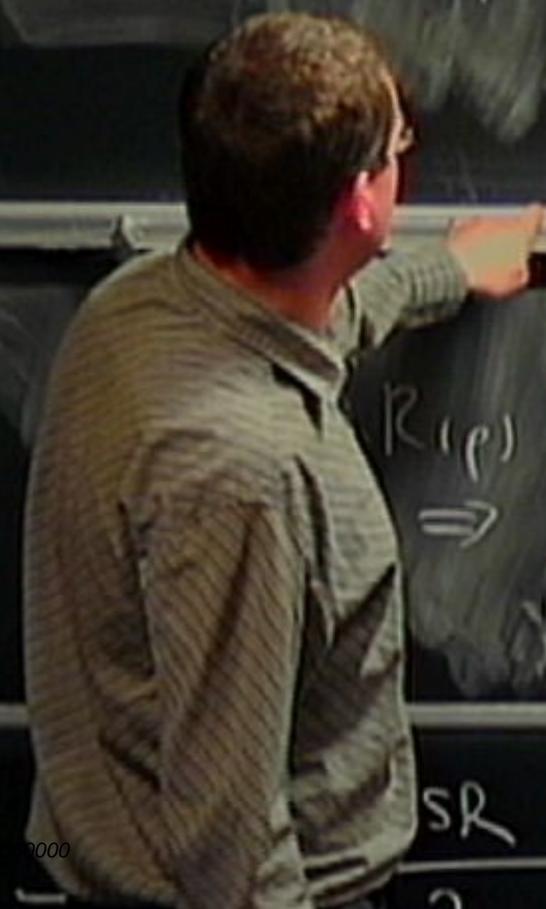
$$A_0 = \frac{1}{2} \quad A = 2 \quad K = \frac{1}{\sqrt{2}} \sinh \frac{\sqrt{2}}{2} r$$

bulk solⁿ is $AdS_2 \times H^3$.

Israel eq: $\sinh \frac{\sqrt{2} p_0}{2} = 1 \quad Q = \frac{p}{\sqrt{2}}$

$$L_1 = \frac{Q}{\sqrt{2}} \quad L_2 = Q$$

Small brane BH



$\{ \dots \}$
spatial gen. of $2+1$

$R(p)$ to vanish somewhere wlog @ $p=0$.
 $\Rightarrow R(p) = p + O(p^3), \quad A(p) = A_0 + O(p^2)$

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$$A_0 = 2 \quad A_1 = 2 \quad K = \frac{1}{\sqrt{2}} \sinh \frac{\sqrt{2}}{2} r$$

bulk solⁿ is $AdS_2 \times H^3$.

Israel eq: $\sinh \frac{\sqrt{2} p_0}{2} = 1 \quad Q = \frac{p}{\sqrt{2}}$

$$L_1 = \frac{Q}{\sqrt{2}} \quad L_2 = Q$$

Small brane BH

$$A_0 = 2 \Rightarrow$$

$\{ \dots \}$
Spatial gen. of $2+1$

instead $R(p)$ to vanish somewhere

Smoothness $\Rightarrow R(p) = p + O(p^3)$

@ $p=0$
 $A_0 + O(p^2)$

Kaus & HSR 0901

$$A_0 = 2 \quad \lambda = 2 \quad K = \frac{1}{\sqrt{2}} \sinh \frac{\sqrt{2} r}{\lambda}$$

bulk solⁿ is $AdS_2 \times H^3$.

Israel eq: $\sinh \frac{\sqrt{2} p_0}{\lambda} = 1 \quad Q = \frac{p}{\sqrt{2}}$

$L_1 = \frac{Q}{\sqrt{2}} \quad L_2 = Q$

Small brane BH

$A_0 = 2 \Rightarrow$ bulk solⁿ is AdS_5

is dead
Smooth

spatial gen. of $2+1$
ish somewhere w/ log @ $p=0$.
 $= p + O(p^3), \quad A(p) = A_0 + O(p^2)$

0901.4236

Israel eq $\sinh \frac{2r_0}{\lambda} = 1$ $Q = \frac{r_0}{\sqrt{2}}$

$L_1 = \frac{Q}{\sqrt{2}}$ $L_2 = Q$

Small brane (BH)

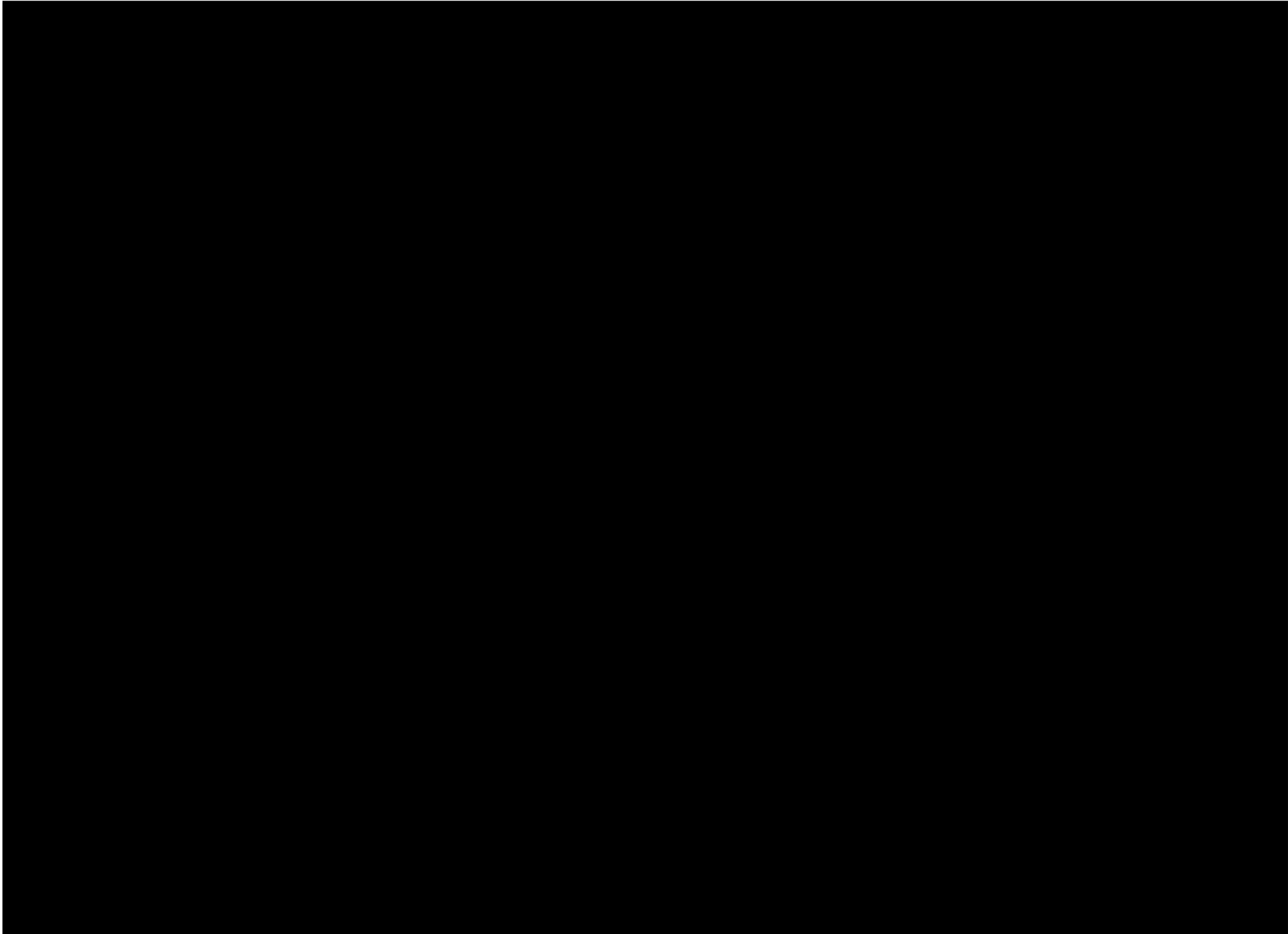
$A_0 = 2 \Rightarrow$ bulk solⁿ is AdS_5 , no solⁿ to Israel...

Assume sph. sym on S^2 $SO(3)$ acting S^2

$ds^2 = A(p)^2 d\Sigma_k^2 + (p)^2 d\Omega^2$

Instead $R(p)$ to \dots gen. of $2d$.
 Smoothness $\Rightarrow R(p) \dots$ here wlog @ $p=0$.
 $A(p) = A_0 + O(p^2)$

Kaus & HSR 09



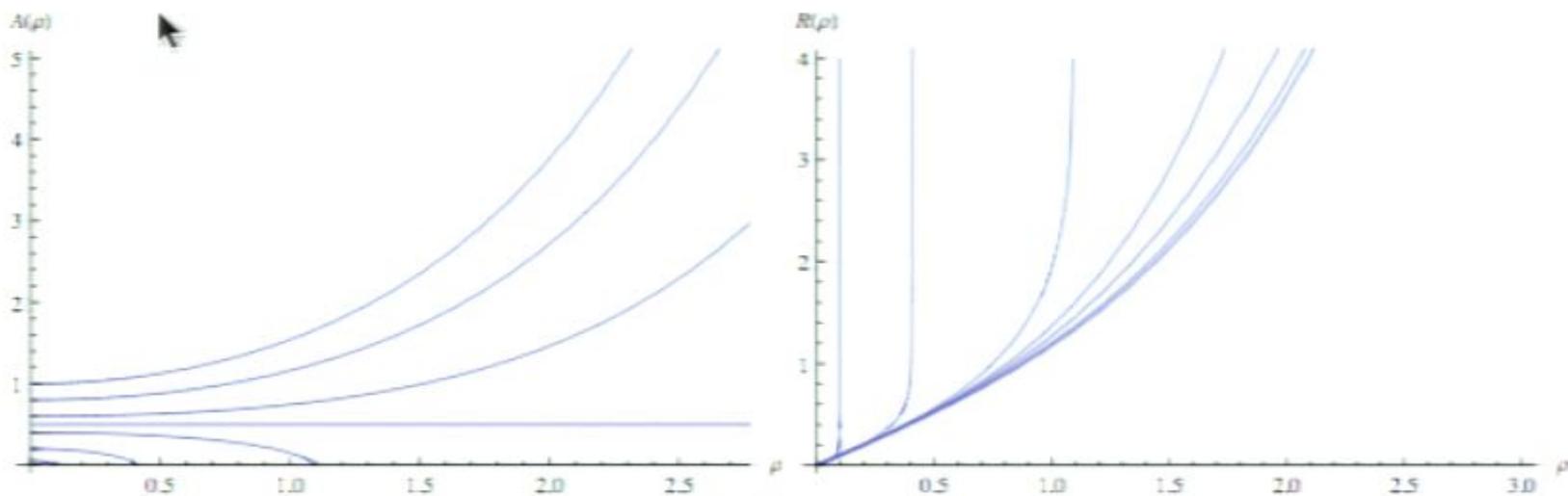


Figure 1: Bulk solutions for $A_0 = 0.05, 0.2, 0.4, 0.5, 0.6, 0.8, 1.0$ (from bottom to top on left plot, from left to right on right plot, units $\ell = 1$).

which is just AdS_5 written in coordinates adapted to a foliation by $AdS_2 \times S^2$ hypersurfaces. For the AdS_5 solution, the Israel equation cannot be satisfied. For the $AdS_2 \times H^3$ solution we find that the Israel equations are satisfied if

$$\sinh\left(\frac{\sqrt{2}\rho_0}{\ell}\right) = 1, \quad Q = \frac{\ell}{\sqrt{2}}. \quad (2.20)$$

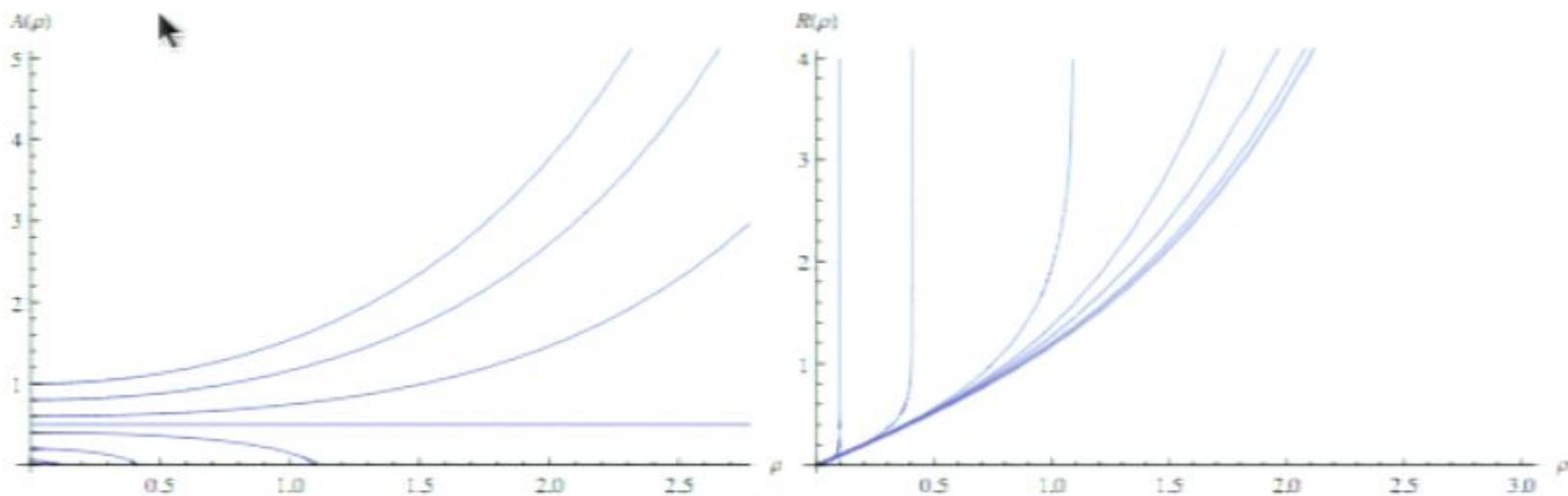


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Solⁿ to Israel eqs exists for $0 < A_0 < 1$.

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ρ_0, Q monotonic in A_0 , $\rightarrow \infty$ as $A_0 \rightarrow l$.

Solⁿ to Israel eqs exists for $0 < A_0 < l$.

($p_0 < p_1 = \text{critical sing}$)

p_0, Q monotonic in A_0 , $\rightarrow \infty$ as $A_0 \rightarrow l$.

Treat Q as indep variable.

Einstein eq \Rightarrow ODE's for $R(\rho), A(\rho)$ solⁿ unique given k

Assume brane @ $\rho = \rho_0 \Rightarrow ds_4^2 = L_1^2 d\Sigma_2^2 + L_2^2 d\Omega^2$
 \parallel $A(\rho_0)^2$ \parallel $R(\rho_0)^2$

Assume Maxwell field on brane is $\star F = Q d\Omega$ charge

Israel eqs: must keep region $\rho \leq \rho_0$

Israel + "Hamiltonian constraint" $\Rightarrow \frac{k}{L_1^2} + \frac{1}{L_2^2} = -\frac{l^2 Q^2}{2L_2^2} \Rightarrow k = \dots \Rightarrow \text{AdS}_5$

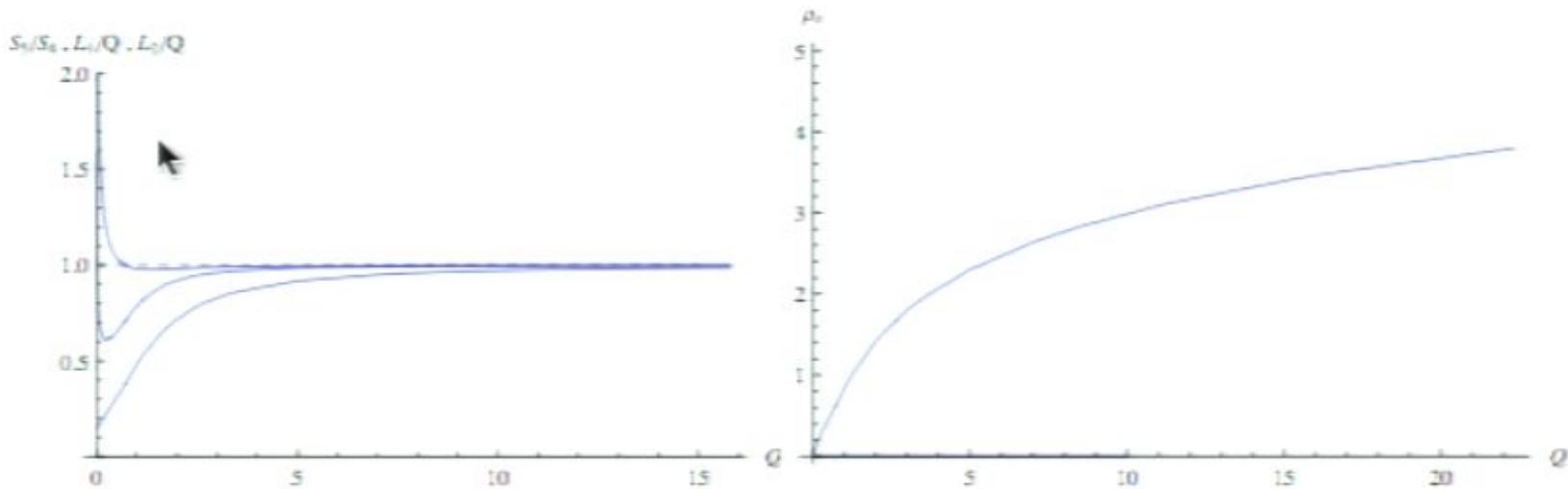


Figure 2: Left: L_2/Q (top), L_1/Q (middle) and S_5/S_4 (bottom). Note that the first two curves diverge at small Q . Right: ρ_0 , the proper length of the horizon transverse to the brane. (Units $\ell = 1$.)

monotonically increasing functions of A_0 which vanish as $A_0 \rightarrow 0$ and diverge as $A_0 \rightarrow \ell$. Physically, it is more interesting to take Q , rather than A_0 as the dependent variable, and we shall do so henceforth.

Figure 2 shows how L_1 , L_2 and ρ_0 depend on Q . L_1/Q and L_2/Q both approach 1 for large Q/ℓ . Hence the induced geometry on the brane agrees with the prediction of 4d GR for large black holes. ρ_0 grows as $\ell \log(Q/\ell) \approx \ell \log(L_2/\ell)$ for large Q , in agreement with general

$$ds^2 = A(r)^2 dt^2 + dr^2 + R(r)^2 d\Omega^2.$$

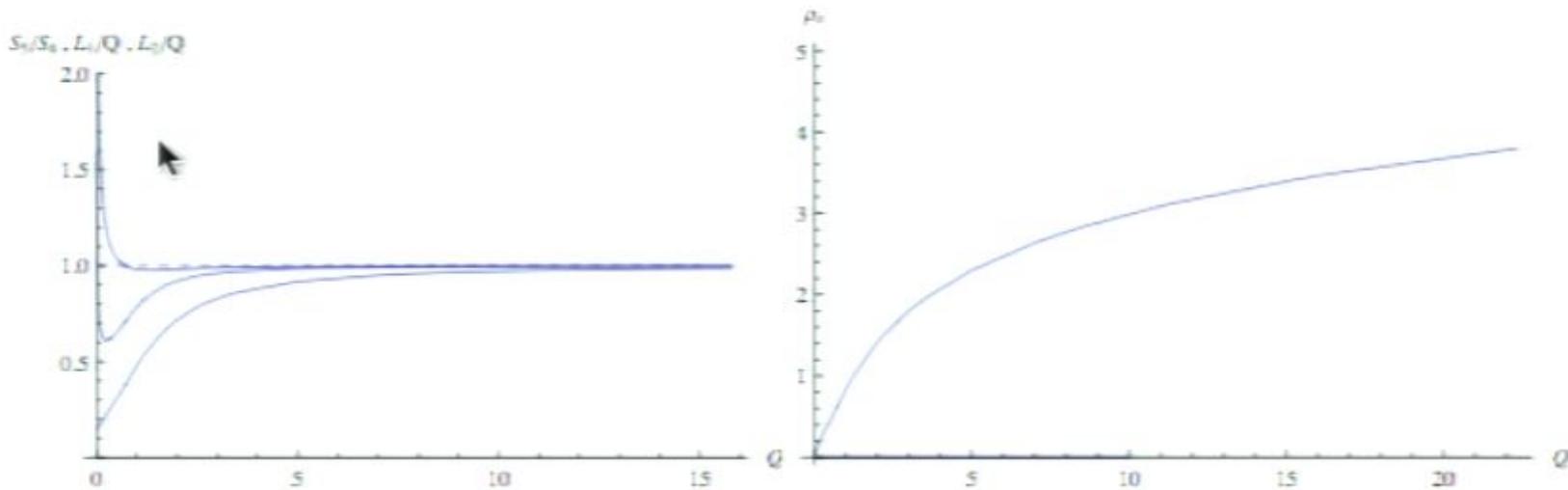


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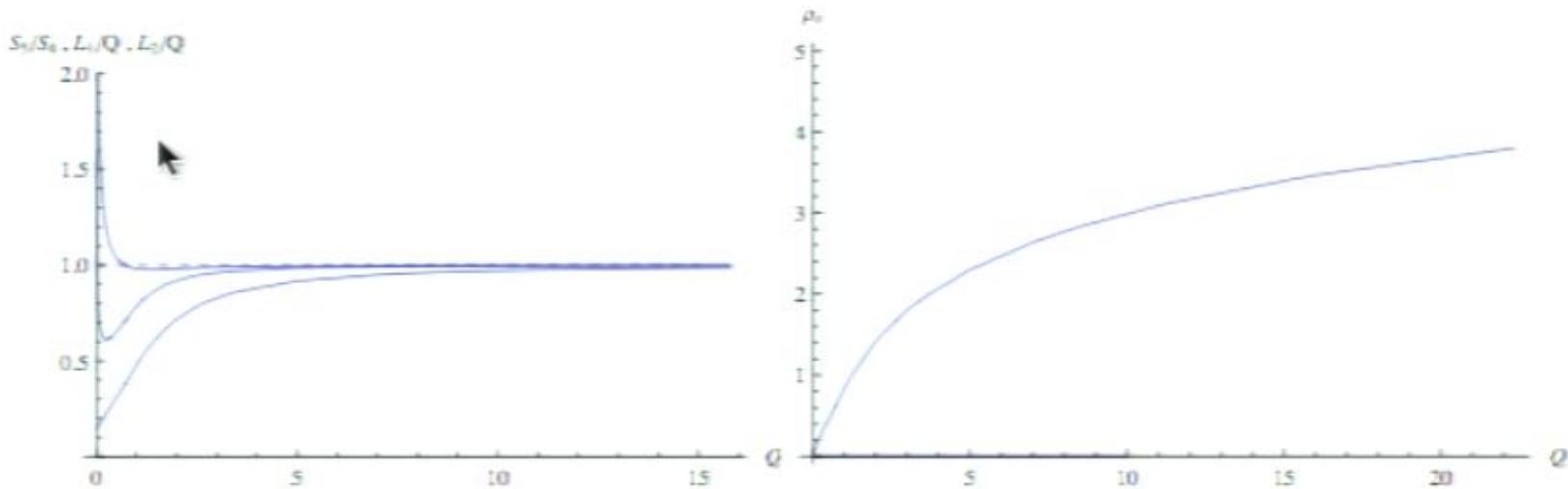


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3H entropy $S_S = \frac{2\pi}{3} \int_0^{\rho_s} K(\rho)^2 d\rho$

3H entropy $S_5 = \frac{2\pi}{\epsilon_5} \int_0^R K(r)^2 dr$

$$S_4 = \frac{\pi Q^2}{\epsilon_4} = \frac{\pi Q^2 l}{\epsilon_5}$$

Navigation and search controls: Previous, Next, Zoom (minus/plus), Move, Text Select, Sidebar, Search.

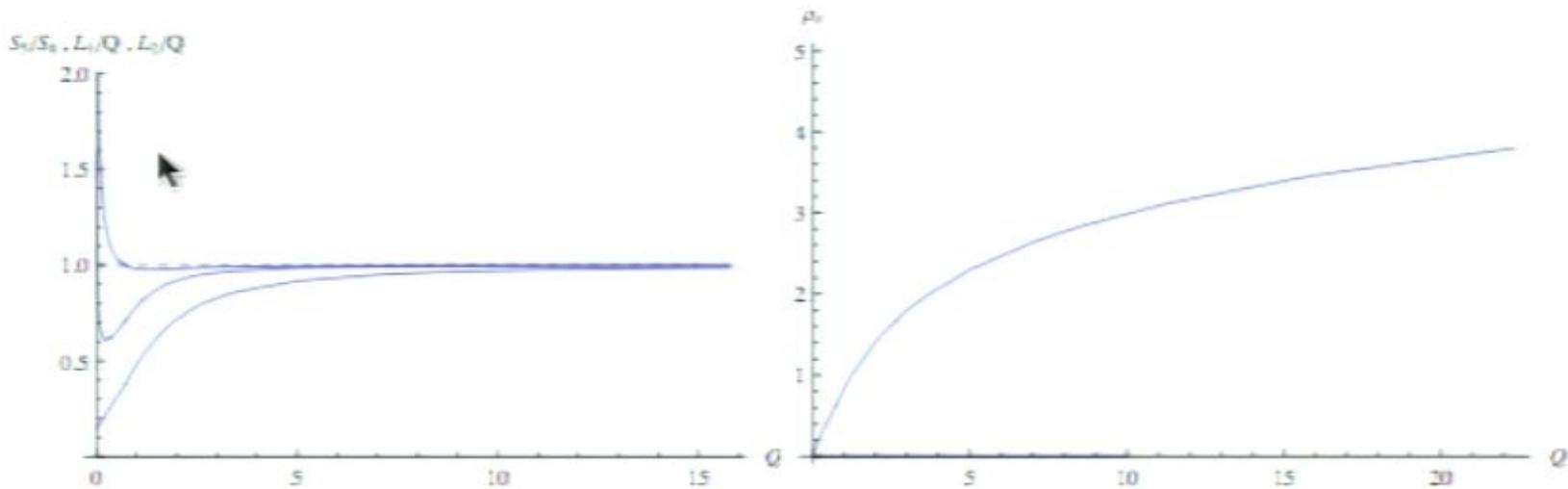


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Analytic results for large BH.

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$$L_1^2 = Q^2 - \frac{3\lambda^2}{4} + \dots$$

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Analytic results for large BH: $\left(\frac{Q}{\lambda} \gg 1\right)$

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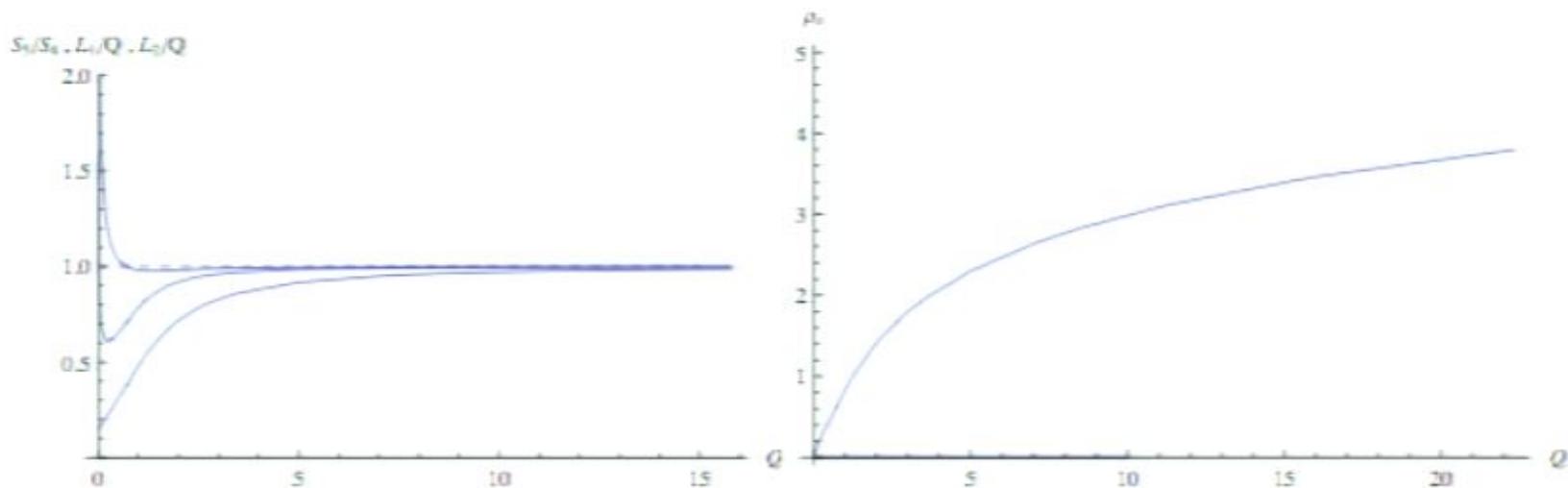


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$$L_1^2 = Q^2 - \frac{3\lambda^2}{4} + \dots$$

$$L_2^2 = Q^2 - \frac{\lambda^2}{4} + \dots$$

$$S_5 = \frac{\pi Q^2 \ell}{\zeta_5} - \frac{\pi \lambda^3}{\zeta_5} \log \frac{Q}{\lambda} + \dots$$

$$- \frac{3\zeta_4 N^2}{\zeta_5}$$

$$S_5 = \frac{\pi Q^2 \ell}{\zeta_5} - \frac{\pi \ell^3}{\zeta_5} \left(\log \frac{Q}{\ell} + \dots \right)$$

$$L_1^2 = Q^2 - \frac{3\zeta_9 N^2}{2\pi} + \dots, \quad L_2^2 = Q^2 - \frac{\zeta_9 N^2}{2\pi}, \quad S_3 = S_4 - 2N^2 \log \frac{Q}{\ell}$$

$$S_4 = \frac{\pi Q^2 \ell}{\zeta_4}$$

$$S_5 = \frac{\pi Q^2 l}{\zeta_5} - \frac{\pi l^3}{\zeta_5} \log \frac{Q}{l} + \dots$$

$$L_1^2 = Q^2 - \frac{3\zeta_4 N^2}{2\pi} + \dots, \quad L_2^2 = Q^2 - \frac{\zeta_4 N^2}{2\pi} + \dots, \quad S_5 = S_4 - 2N^2 \log \frac{Q}{l}$$

$$S_4 = \frac{\pi Q^2}{\zeta_4} = \frac{\pi Q^2 l}{\zeta_5}$$

$$L_1^2 = Q^2 - \frac{3\zeta_4 N^2}{2\pi} + \dots, \quad L_2^2 = Q^2 - \frac{\zeta_4 N^2}{2\pi} + \dots$$

$$S_5 = S_4 - 2N^2 \log \frac{Q}{l}$$

3H entropy $S_5 = \frac{2\pi}{\zeta_5} \int_0^Q R(r)^2 dr$

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Ignore brane

$$A_0 > \left\{ \begin{array}{l} \dots \end{array} \right.$$

Navigation controls: Next, Zoom, Move, Text, Select, Sidebar, Search

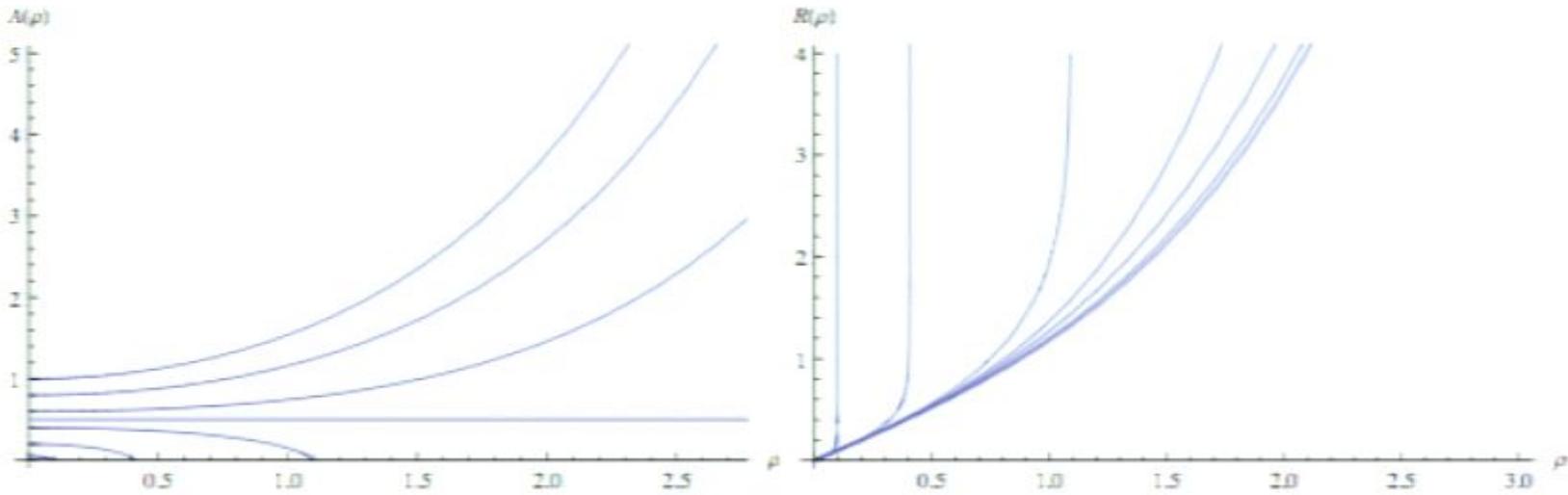


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$$\sinh\left(\frac{\sqrt{2}\rho_0}{\ell}\right) = 1, \quad Q = \frac{\ell}{\sqrt{2}}. \tag{2.20}$$

We then have

$$L_1^2 = Q^2 - \frac{3\zeta_4 N^2}{2\pi} + \dots, \quad L_2^2 = Q^2 - \frac{\zeta_4 N^2}{2\pi} + \dots$$

$$S_5 = S_4 - 2N^2 \log \frac{Q}{l_p}$$

3H entropy $S_5 = \frac{2\pi}{\zeta_5} \int_0^{\rho} R(\rho)^2 d\rho$

$$S_4 = \frac{\pi Q^2}{\zeta_4} = \frac{\pi Q^2 l}{\zeta_5}$$

ignore brane

$A_0 > l_p$ bulk asymp locally adS_5 , conformal boundary $adS_2 \times S^2$

$$L_1^2 = Q^2 - \frac{3\zeta_4 N^2}{2\pi} + \dots, \quad L_2^2 = Q^2 - \frac{\zeta_4 N^2}{2\pi} + \dots$$

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$$S_4 = \frac{\pi Q^2}{\zeta_4} = \frac{\pi Q^2 l}{\zeta_5}$$

ignore brane

$A_0 > l$ bulk asymp locally adS_5 , conformal boundary

$adS_2 \times S^2$ (sim $adS_3 \times S^2$, $adS_2 \times S^1 \times S^1 \dots$)

$A_0 = 1 \Rightarrow$ bulk solⁿ is AdS_5 , no solⁿ to Israel.

$A_0 = 2$ sph. sym on brane $SO(3)$ acting S^2

$$A(p)^3 d\Sigma_k^2 + \underbrace{dp^2 + R(p)^2 d\Omega^2}_{\text{spatial geom. of } \mathcal{R}^4}$$

lead $R(p)$ to vanish somewhere wlog @ $p=0$.
mass $\Rightarrow R(p) = p + O(p^3)$, $A(p) = A_0 + O(p^2)$

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why?

3H entropy $S_5 = \frac{2\pi}{G_5} \int_0^{\rho} R(\rho)^2 d\rho$

$$S_4 = \frac{\pi Q^2}{G_4} = \frac{\pi Q^2 \ell}{G_5}$$

ignore brane

$A_0 > \frac{1}{2}$ bulk asymp locally adS_5 , conformal boundary

$\mathbb{R} \times S^1 \times S^1$ $adS_2 \times S^2$ (sim) $adS_3 \times S^2$, $adS_2 \times S^2 \times S^1 \dots$

3H entropy $S_5 = \frac{2\pi}{G_5} \int_0^{\rho} R(\rho)^2 d\rho$

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$A_0 > \frac{1}{2}$

bulk asymp locally adS_5 , conformal boundary

$adS_3 \times S^2$

(sim $\rightarrow adS_3 \times S^2, adS_2 \times S^1 \times S^1 \dots$)

$\mathbb{R} \times S^1 \times S^1$