Title: Quantum Spin Simulations (PHYS 7380) - Lecture 5

Date: Apr 09, 2010 11:00 AM

URL: http://pirsa.org/10040047

Abstract:

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More general finite-size scaling hypothesis

has been justified using the renormalization-group theory

$$Q(t, L) = L^{\sigma} f(\xi/L),$$

Jsing $\xi \sim |t|^{-1/\nu} \rightarrow$

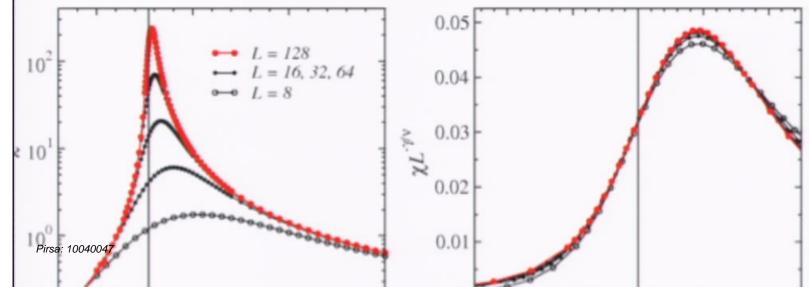
$$Q(t, L) = L^{\sigma} g(tL^{1/\nu})$$

From this we must be able to reproduce infinite-size form:

$$Q(t, L \to \infty) \sim |t|^{-\kappa}$$

which is the case if $g(x) \sim x^{-\kappa}$ and $\sigma = \kappa/
u$

Test: susceptibility of 2D Ising model (Monte Carlo)



$$T_c = 2/\ln(1+\sqrt{2})$$

 $\nu = 1, \gamma = 7/4$

Normally: adjust Tc and exponents so that the wata

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Nonte Carlo methods - based on random numbers Stanislav Ulam's terminology

- his uncle frequented the Casino in Monte Carlo

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Random (pseudo random) number generator on the computer Less glamorous than roulette tables or cards, but faster...

>109 random numbers per second

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Monte Carlo simulations in statistical physics

normally refers to importance sampling of configurations (e.g., spins)

he Metropolis algorithm

Metropolis, Rusenbluth, Rosenbluth, Teller, and Teller, Phys. Rev. 1953]

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he Metropolis algorithm

Metropolis, Rusenbluth, Rosenbluth, Teller, and Teller, Phys. Rev. 1953]

Generate a series of configurations (Markov chain); $C_1 \rightarrow C_2 \rightarrow C_3 \rightarrow C_4 \rightarrow ...$

C_{n+1} obtained by modifying (updating) C_n

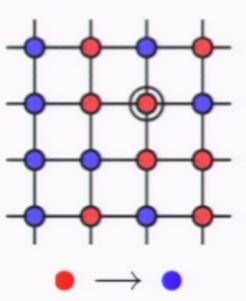
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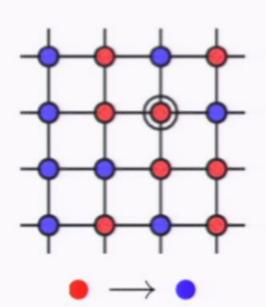
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$$\frac{P_{\text{change}}(A \to B)}{P_{\text{change}}(B \to A)} = \frac{W(B)}{W(A)} \qquad W(A) = e^{-E(A)/T}$$



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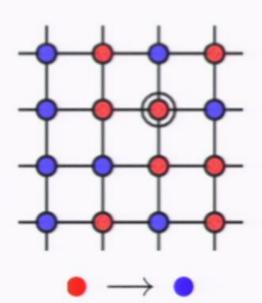
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Starting from any configuration, such a stochastic process eads to configurations distributed according to W

- the process has to be ergodic
 - any configuration reachable in principle
- it takes some time to reach equilibrium



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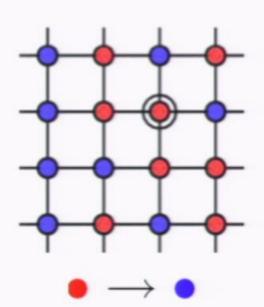
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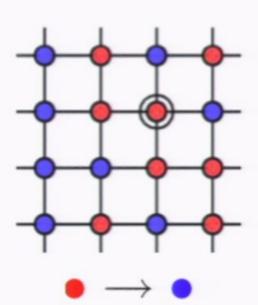
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Metropolis algorithm for the Ising model. For each update perform:

- select a spin i at random; consider flipping it σ_i → -σ_i
- compute the ratio $R=W(\sigma_1,...-\sigma_i,...,\sigma_N)/W(\sigma_1,...\sigma_i,...,\sigma_N)$
 - for this we need only the spins neighboring i
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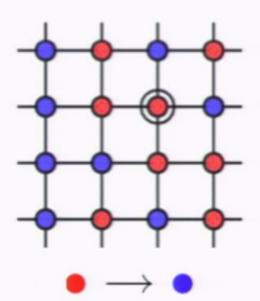
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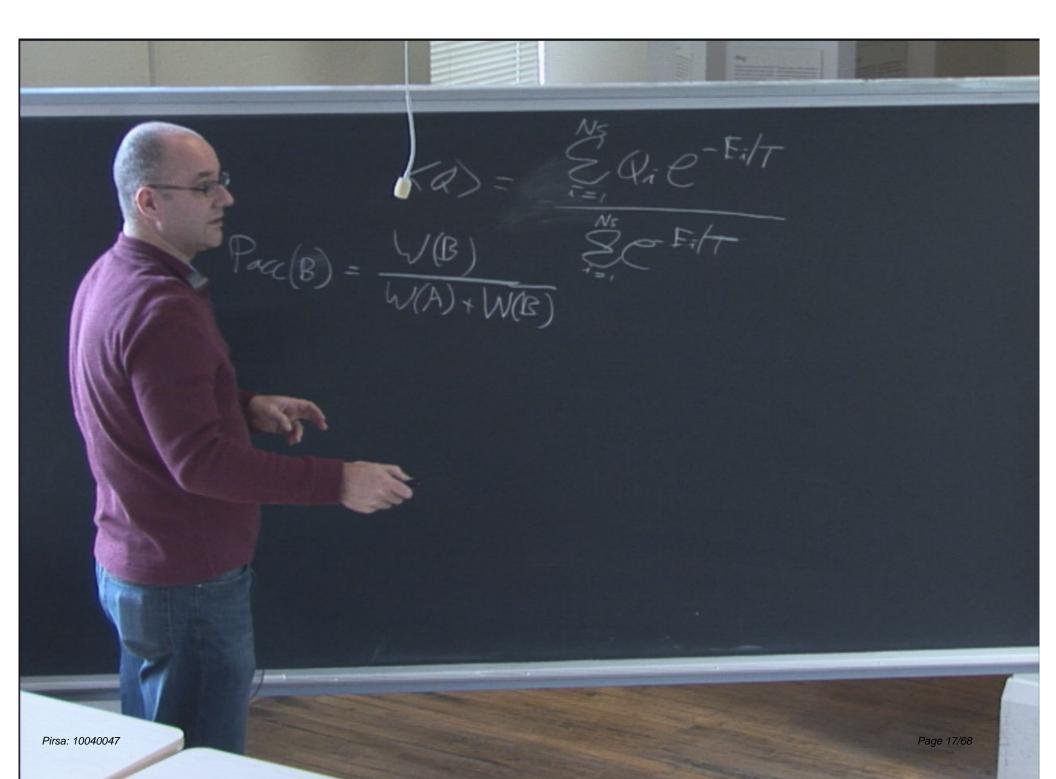


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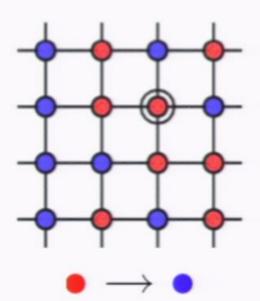
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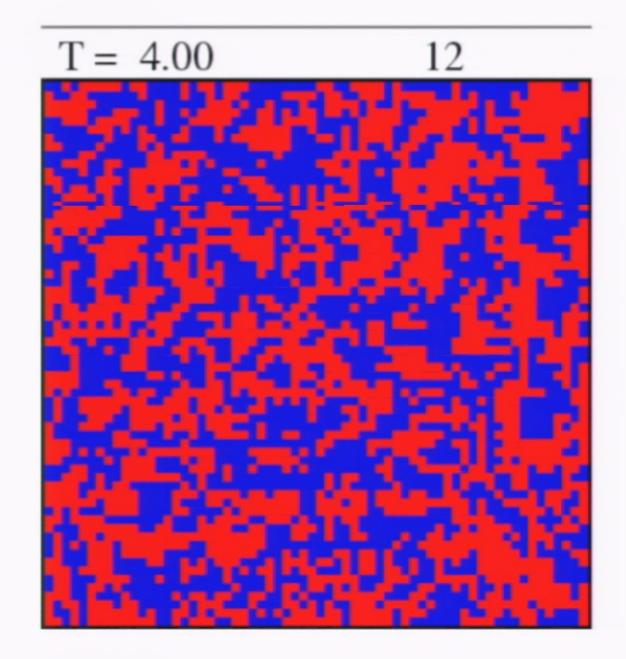
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=0 simulations

28×128 lattice N=16384)

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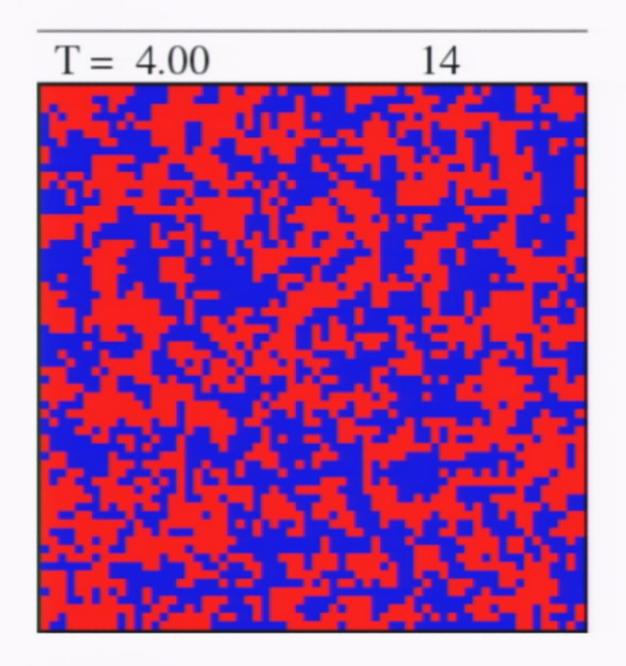


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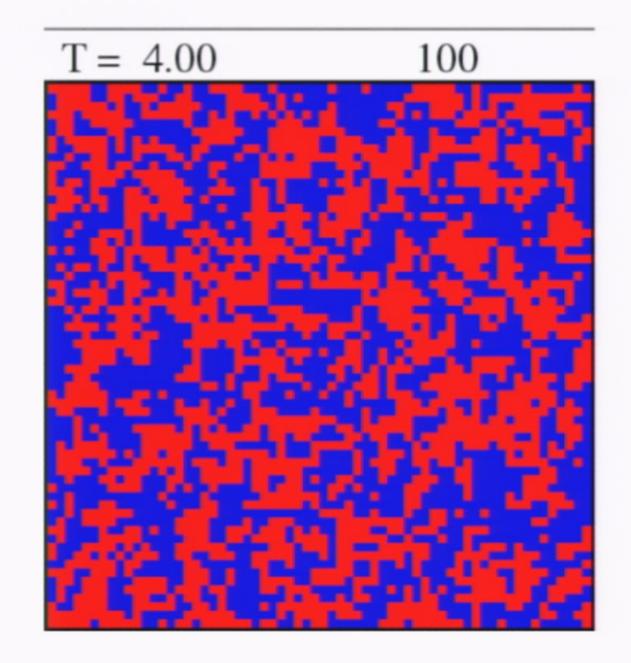


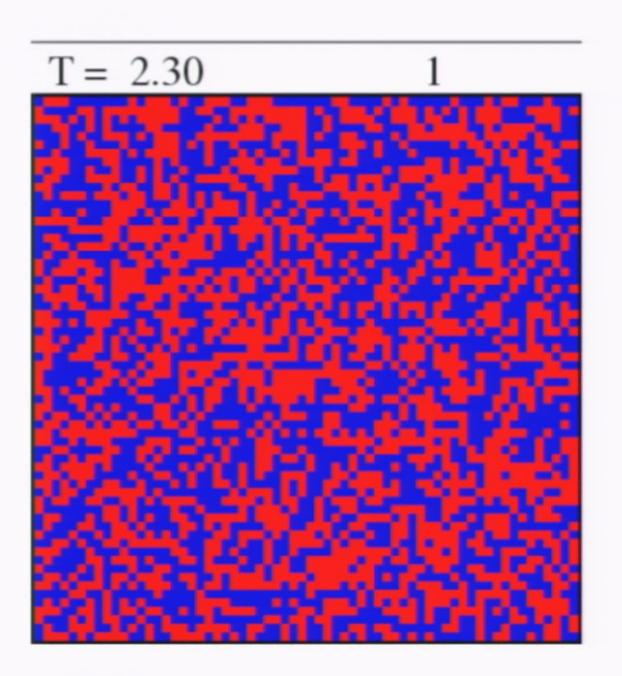
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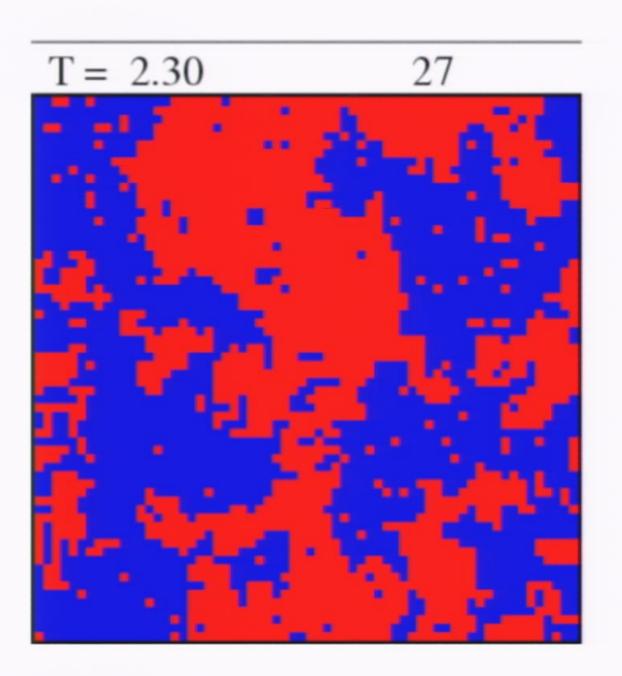
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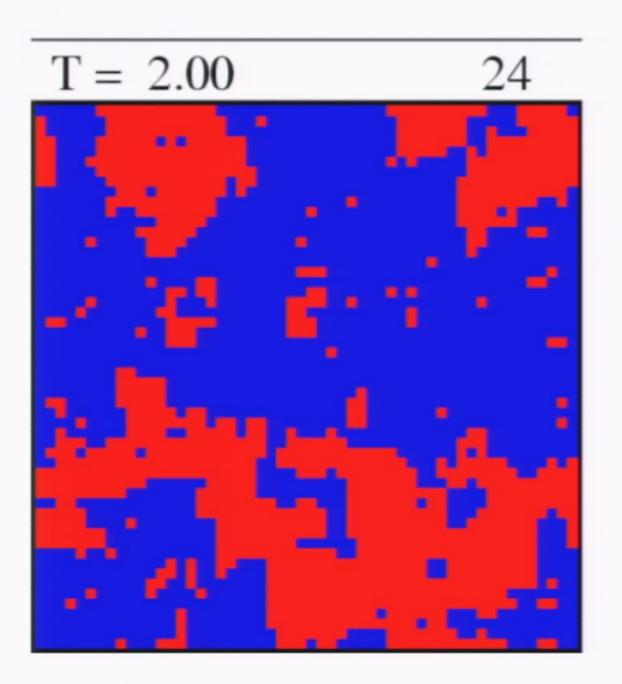
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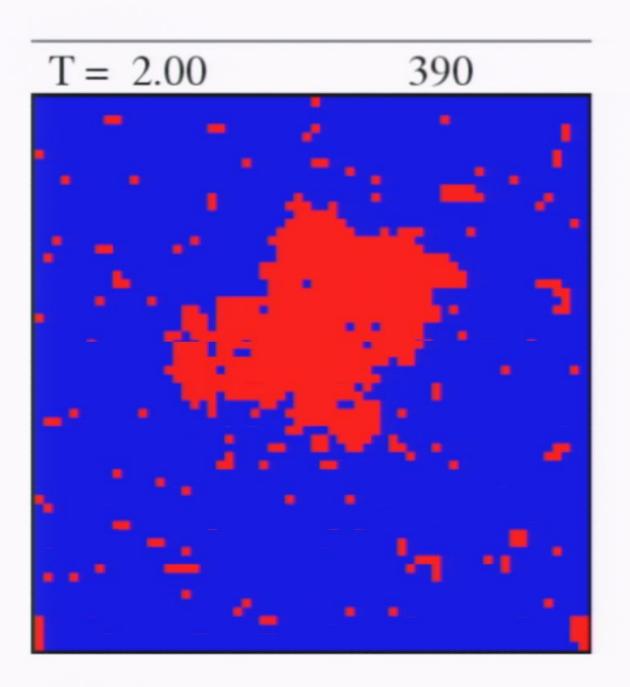
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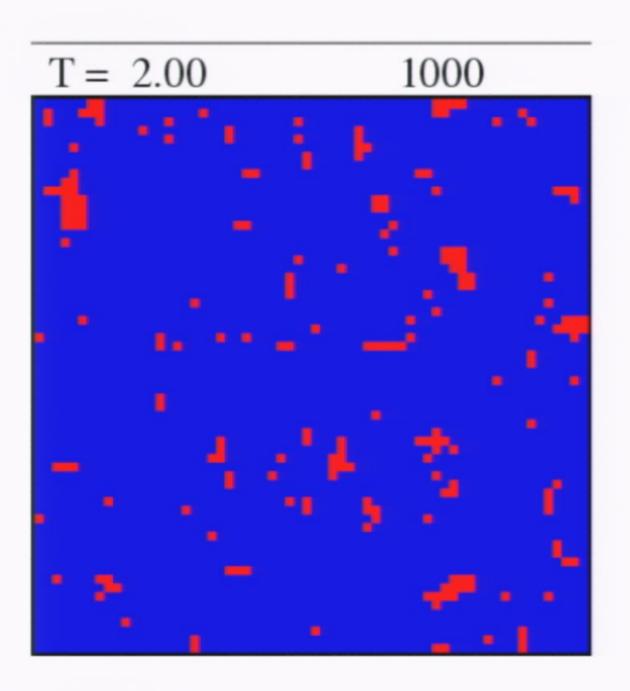










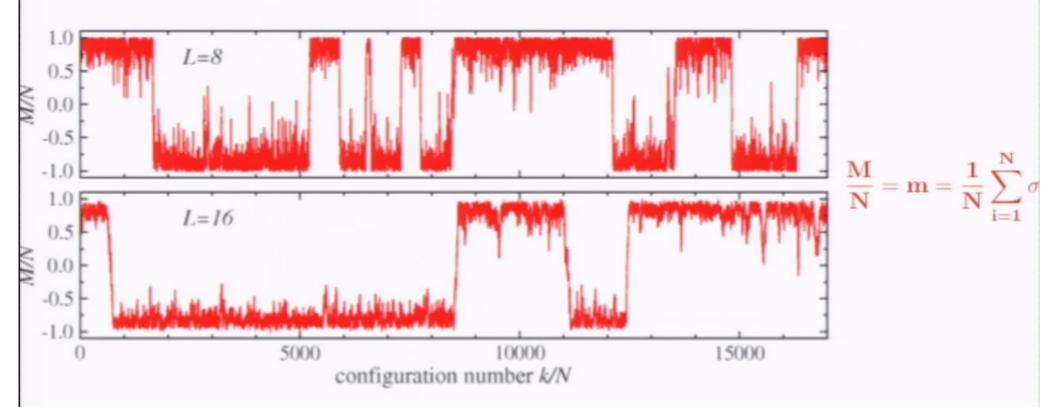


A magnetized state, <m>≠0, breaks a symmetry (E invariant under all $\sigma_i \rightarrow -\sigma_i$) strictly, mathematically we must have <m>=0 symmetry breaking (phase transition) can take place when N→∞ how can we understand the symmetry breaking for N large but finite?

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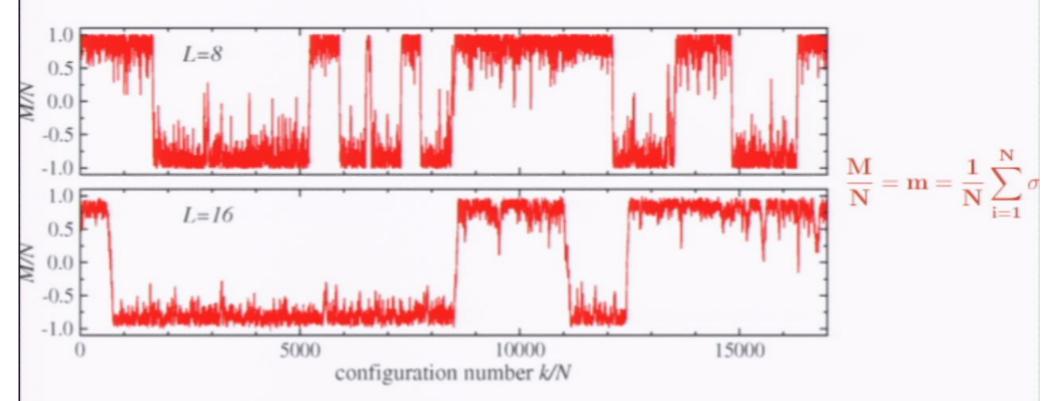
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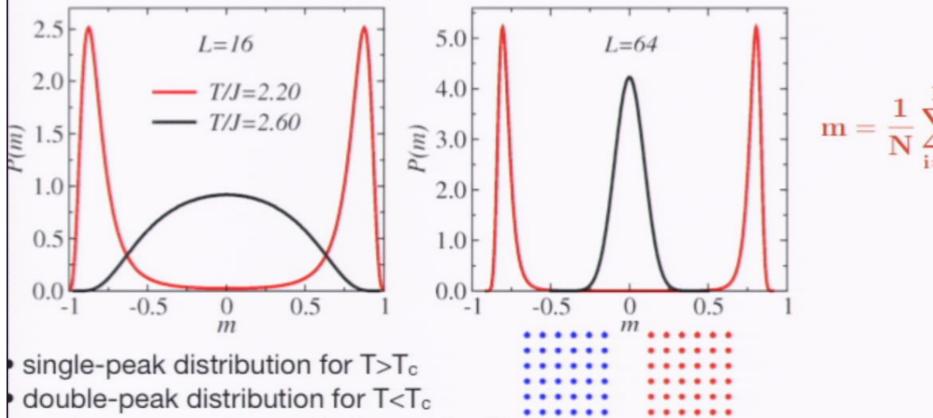


here is a characteristic "reversal" time between m>0 and m<0 configurations
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reversal time diverges for N→∞

probability distrubution (histogram) of m during the simulation

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probability distrubution (histogram) of m during the simulation



$$\mathbf{m} = \frac{1}{N} \sum_{i=1}^{N} \sigma$$

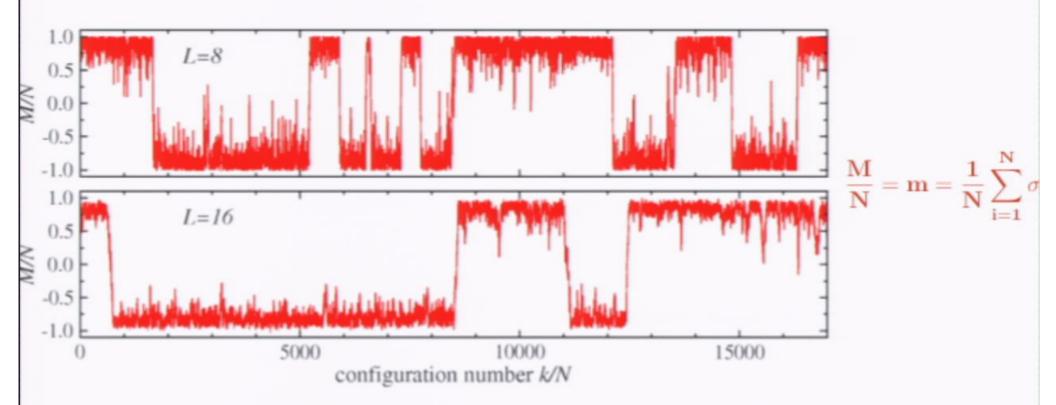
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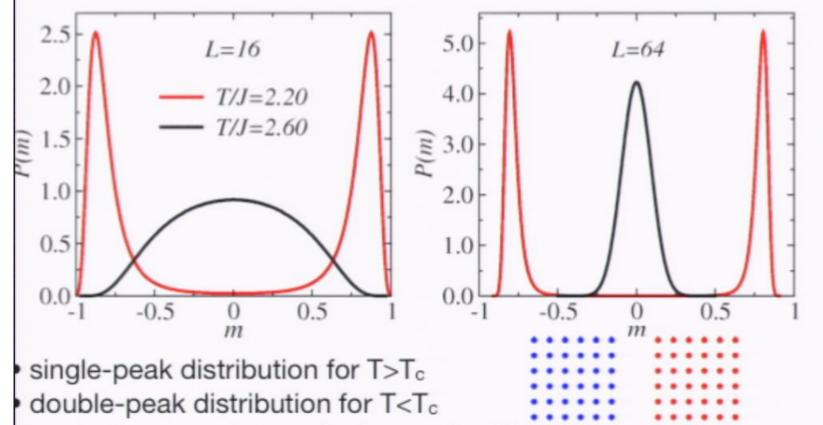
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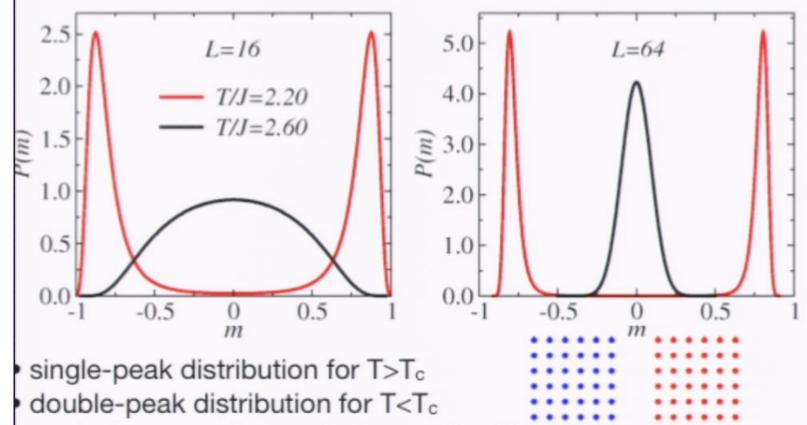
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Why this peak structure? balance between large number of m≈0 configurations with high energy small mumber of months at high T internal energy at low. The strong dominates at hight T internal energy at low. The strong dominates at hight T internal energy at low.

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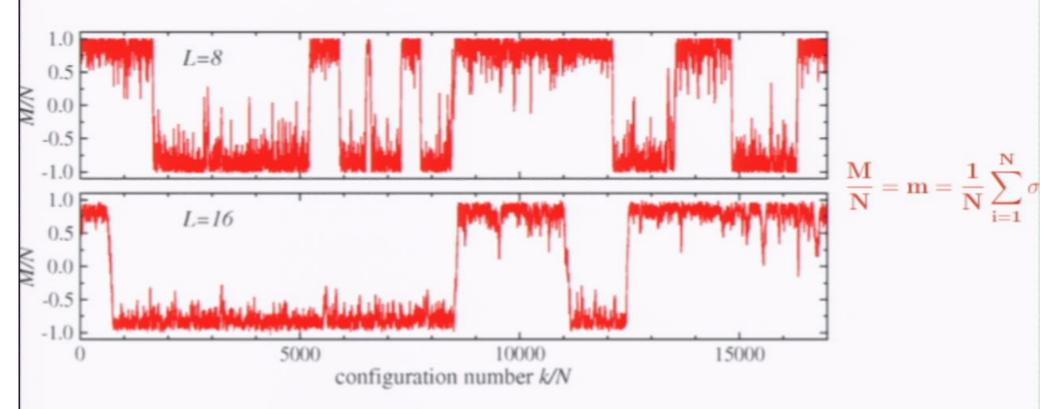
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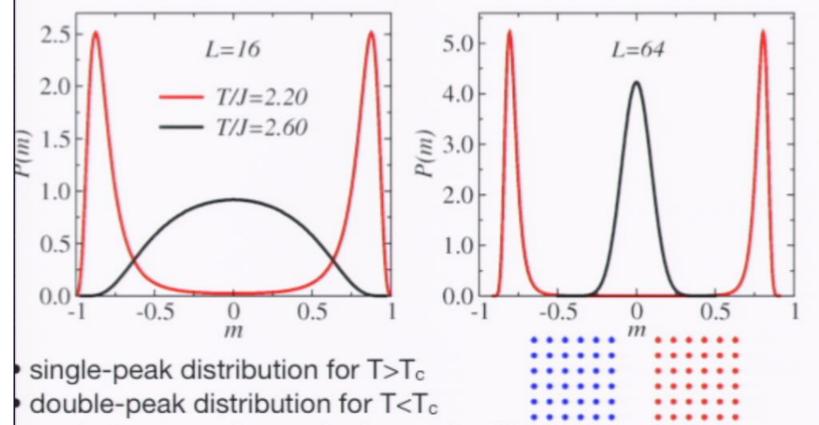
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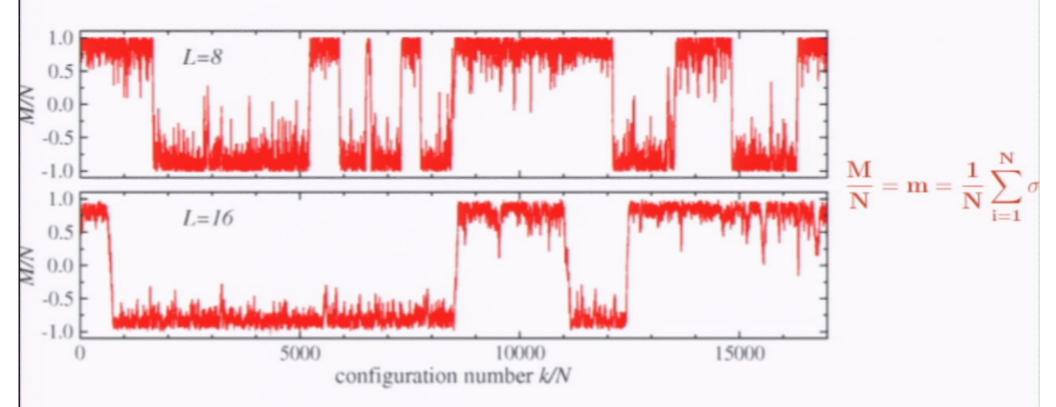
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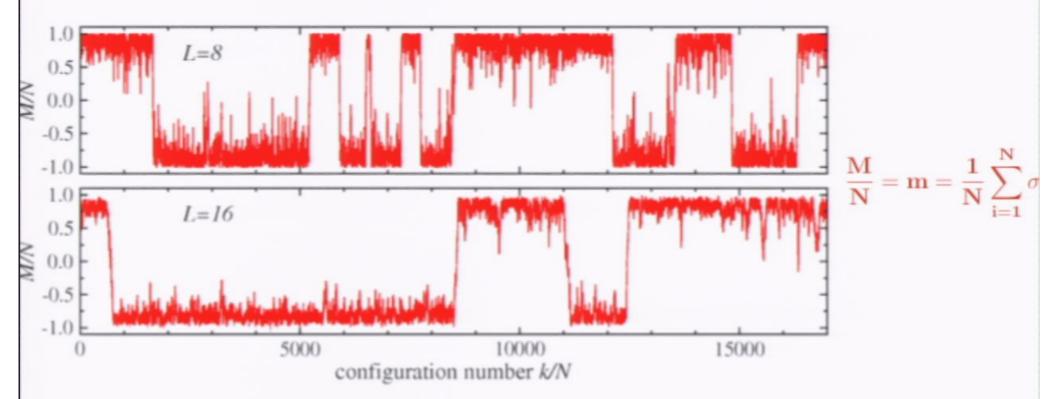
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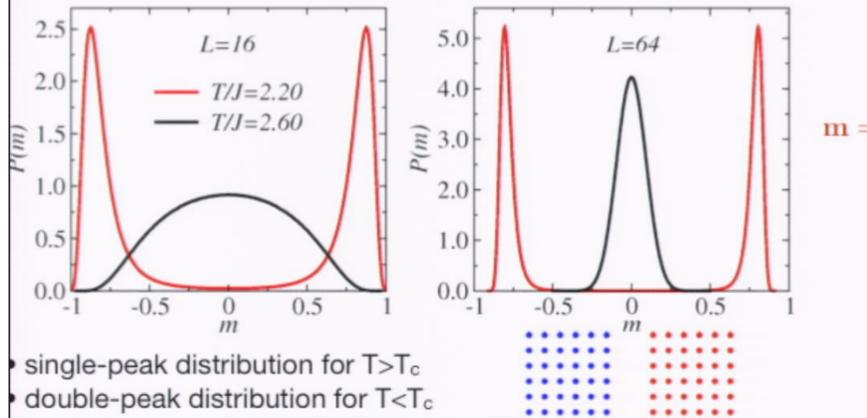
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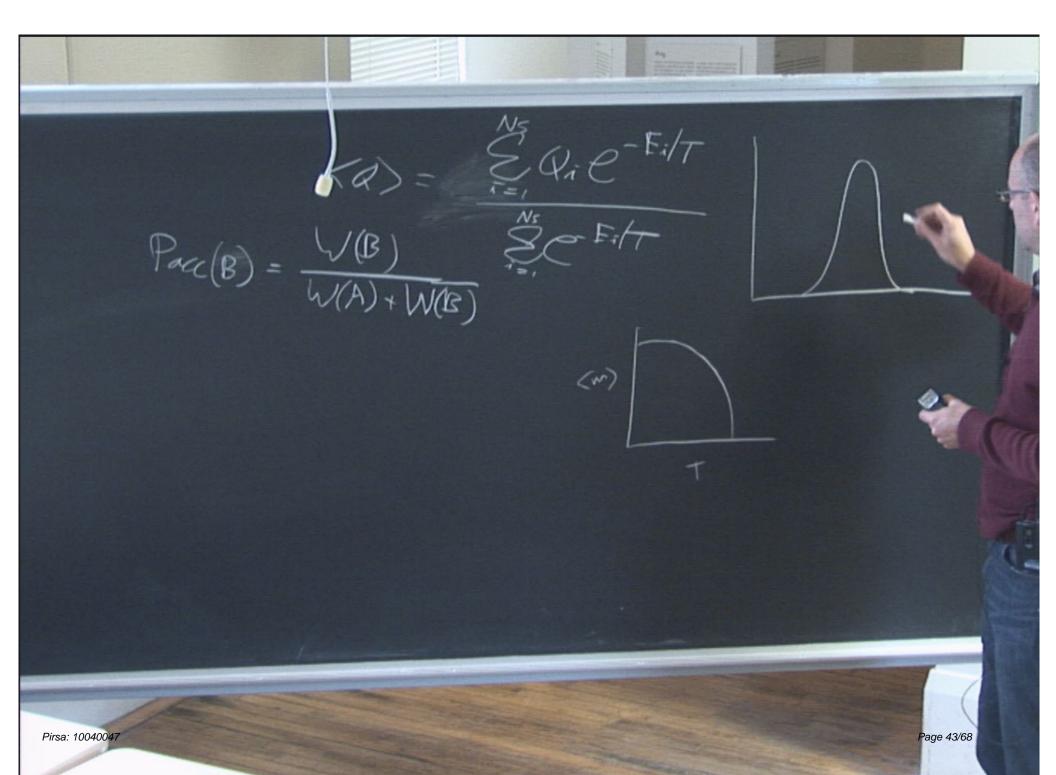
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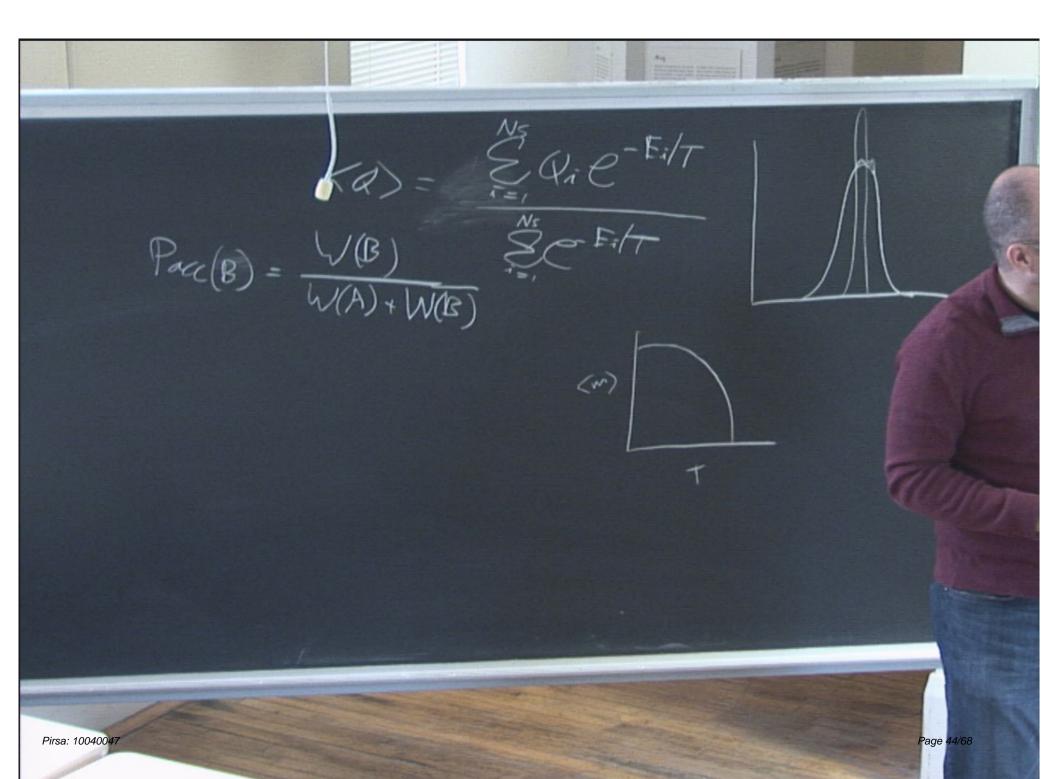
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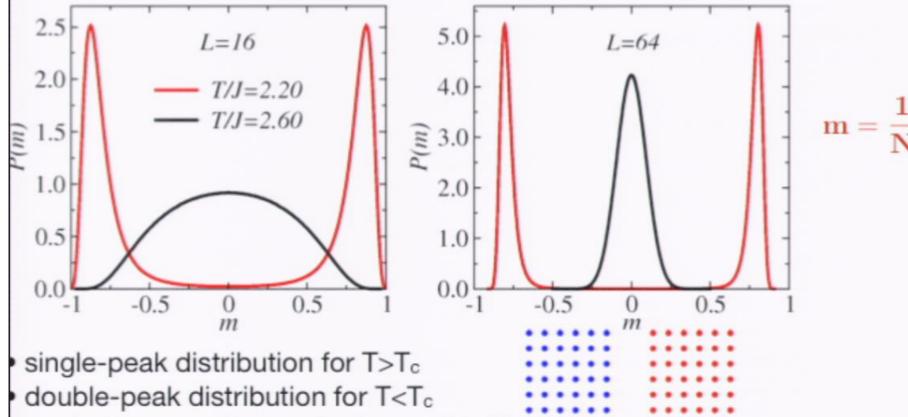


(m)

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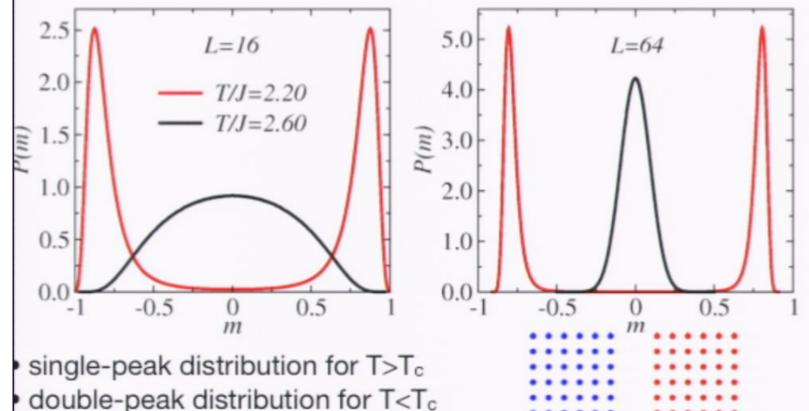
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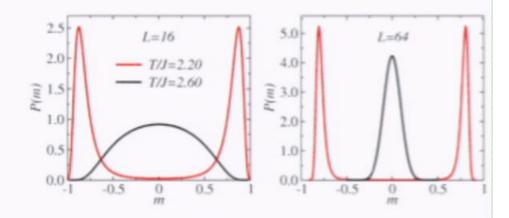
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- double-peak distribution for T<Tc
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Vhy this peak structure? balance between large number of m≈0 configurations with high energy SPirsa: 2004-0047 umber of |m|≈1 configuration with low energy entropy dominates at hight T internal energy at low T



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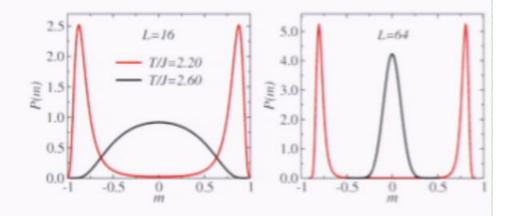
Consider the dimensionless ratio

$$R_2 = \frac{\langle m^4 \rangle}{\langle m^2 \rangle^2}$$

We can compute R₂ exactly for N→∞

for T<T_c: P(m)→δ(m-m*)+δ(m+m*)
 m*=|peak m-value|

$$R_2 \rightarrow 1$$



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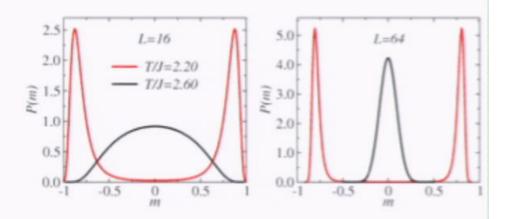
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R₂→3 (properties of Gaussian integrals)

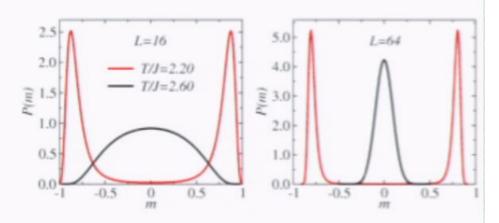
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The **Binder cumulant** is defined as (n-component order parameter; n=1 for Ising)

$$U_2 = \frac{3}{2} \left(\frac{n+1}{3} - \frac{n}{3} R_2 \right) \to \begin{cases} 1, & T < T_c \\ 0, & T > T_c \end{cases}$$

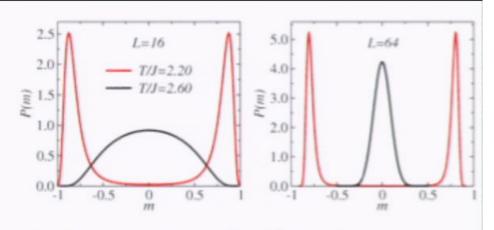
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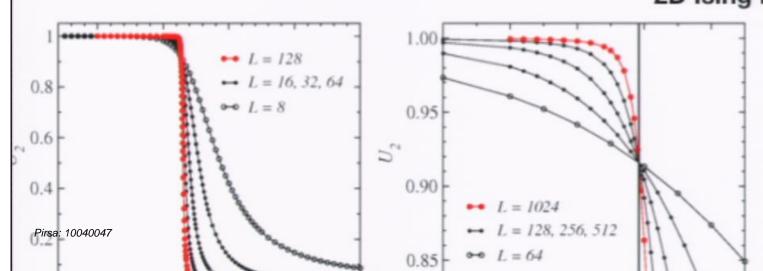
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2D Ising model; MC results



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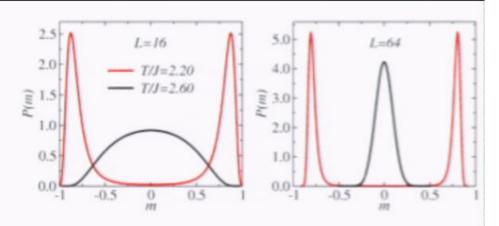
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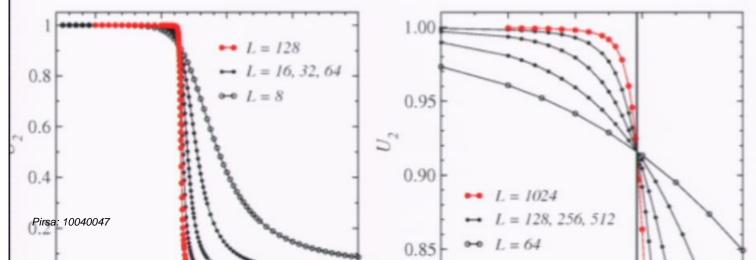
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2D Ising model; MC results



Curves for different L normally cross each other close to T_c

Extrapolate crossing for sizes L and 2L to infinite size

• converges faster than

Definition: Monte Carlo sweep = N spin-flip attempts

a natural unit of simulation "time"

"measure" observables after every (or every n) sweep

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Definition: Monte Carlo sweep = N spin-flip attempts

- a natural unit of simulation "time"
- "measure" observables after every (or every n) sweep
- Boltzmann probability accounted for at sampling stage →

$$\bar{Q} = \frac{1}{N_s} \sum_{i=1}^{N_s} Q_i, \quad N_s = \text{number of samples}$$

s the estimate for the true expectation value;

$$\bar{Q} \to \langle Q \rangle, \quad (N_s \to \infty)$$

Definition: Monte Carlo sweep = N spin-flip attempts

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- "measure" observables after every (or every n) sweep
- Boltzmann probability accounted for at sampling stage →

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f M is sufficiently large (>> autocorrelation time) the average and error are

characterization of how measurements become statistically independent

$$A_Q(t) = \frac{\langle Q(i+t)Q(i)\rangle - \langle Q\rangle^2}{\langle Q^2\rangle - \langle Q\rangle^2}, \quad (\to e^{-t/\Theta}, \ t \to \infty)$$

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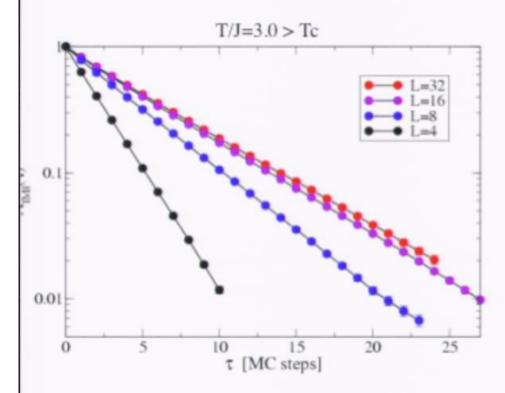
he autocorrelation time Θ grows as $T \rightarrow T_c$ (diverges for $N \rightarrow \infty$, $T \rightarrow T_c$)

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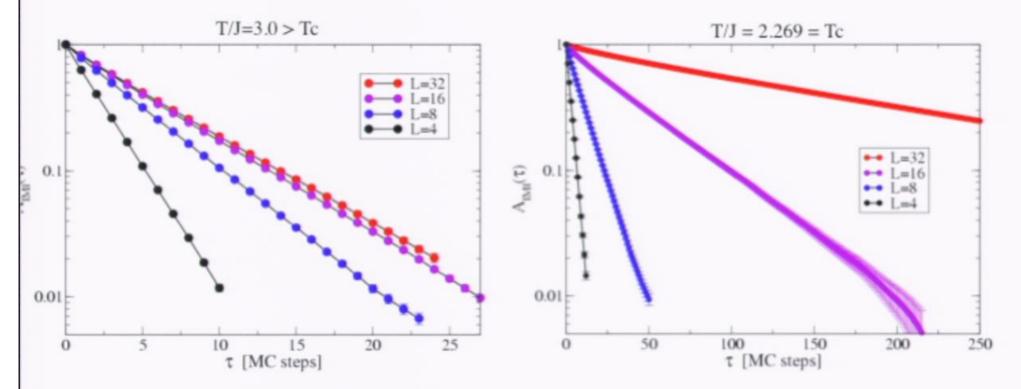


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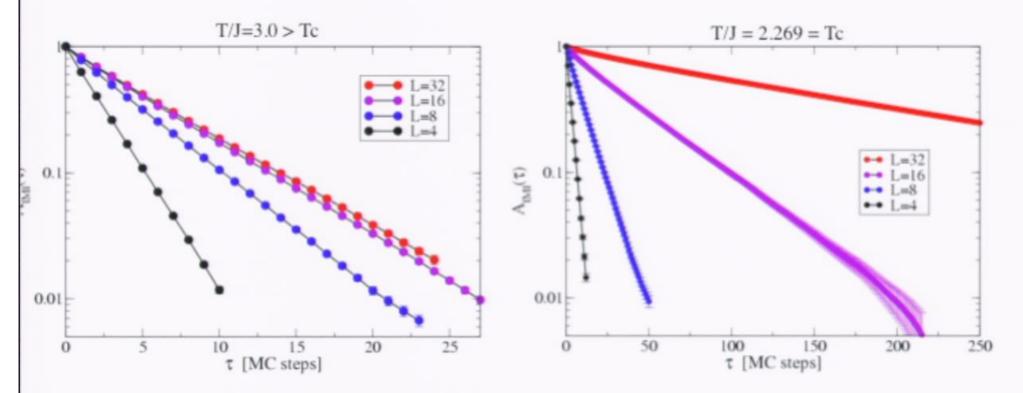


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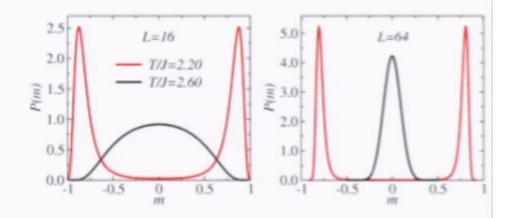
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This problem can be largely overcome by using cluster algorithms for standard Ising, XY, Heisenberg,...

but not in all cases, e.g., in the presence of external fields, frustrated systems,...

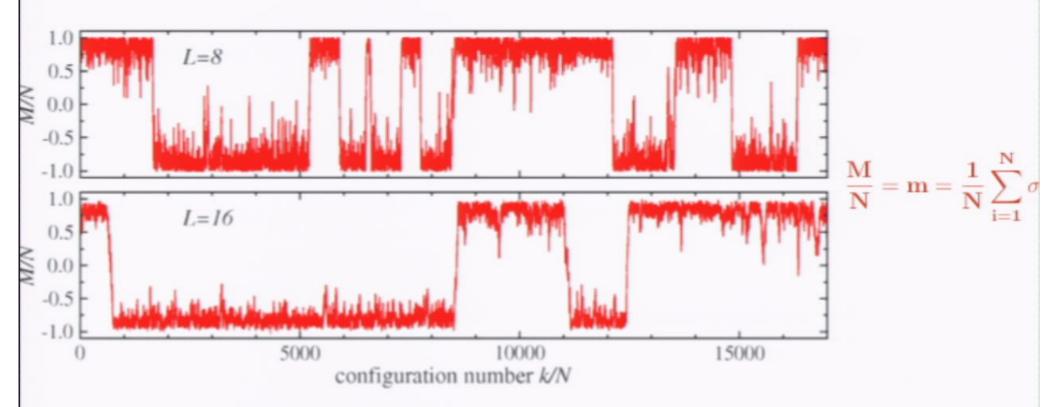


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symmetry breaking (magnetic phase transition) for h=0

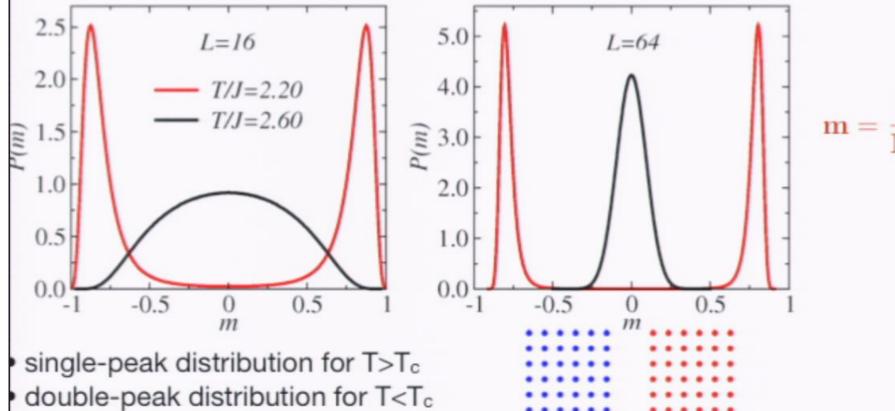
A magnetized state, <m>≠0, breaks a symmetry (E invariant under all $\sigma_i \rightarrow -\sigma_i$) strictly, mathematically we must have <m>=0 symmetry breaking (phase transition) can take place when N→∞ how can we understand the symmetry breaking for N large but finite?

ime series of simulation data; magnetization vs simulation "time" for T<Tc



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probability distrubution (histogram) of m during the simulation



$$\mathbf{m} = \frac{1}{N} \sum_{i=1}^{N} \sigma_i$$

- peaks become sharper for increasing N
- no probability to fluctuate between m<0 and m>0 peaks for N→∞
- have to go through low-probability m≈0 configurations

probability distrubution (histogram) of m during the simulation

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