

Title: Holographic non-Fermi Liquids

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Abstract: We describe a class of non-Fermi liquid systems, using the AdS/CFT correspondence. The Fermi surfaces are studied by computing the response functions of fermionic operators. The scaling behavior near the Fermi surfaces is determined by conformal dimensions in an emergent IR CFT. The low-energy excitations near the Fermi momenta are not Landau quasiparticles. When the operator is marginal in the IR CFT, the full spectral function is precisely of the 'marginal Fermi liquid' form, introduced as a phenomenological model of the 'strange metal' phase of high temperature superconductors.

# Non-Fermi liquids from holography

David Vegh  
Simons Center for Geometry and Physics



Hong Liu, John McGreevy, DV [arXiv:0903.2477](#)

Thomas Faulkner, Hong Liu, John McGreevy, DV [arXiv:0907.2694](#)

Thomas Faulkner, Gary Horowitz, John McGreevy, Matthew Roberts, DV [arXiv:0911.3402](#)

(see also: Sung-Sik Lee, [arXiv:0809.3402](#) ; Cubrovic, Zaanen, Schalm, [arXiv:0904.1933](#))

## (Non-)Fermi liquids

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AdS/CFT background

Review of AdS/CFT

$AdS_4$  – BH geometry

Spinor Green's function

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Fermi surfaces

Near-horizon  $AdS_2$  and emergent IR CFT

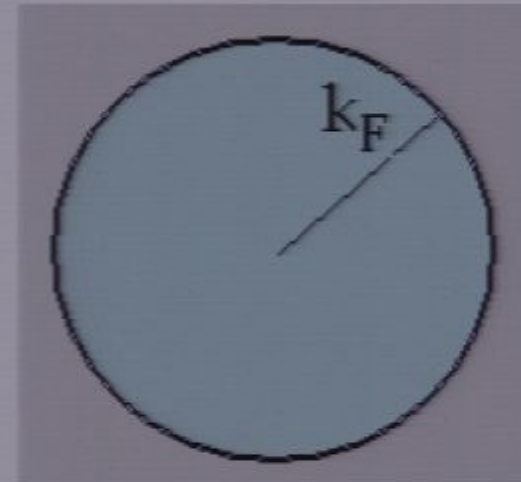
Small  $\omega$  expansion from matching

# Fermions at finite density

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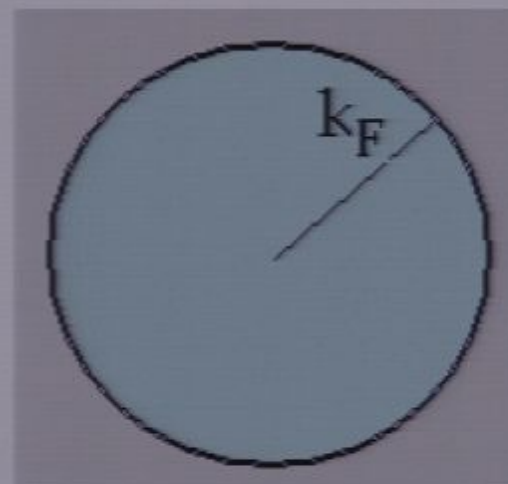
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- ▶ stable RG fixed point [Polchinski, Shankar]  
(modulo BCS instability) [Benfatto-Gallivotti]

- ▶ weakly interacting *quasiparticles*  
→ thermodynamics, transport properties

- ▶ appear as *poles* in the single-particle Green's function:

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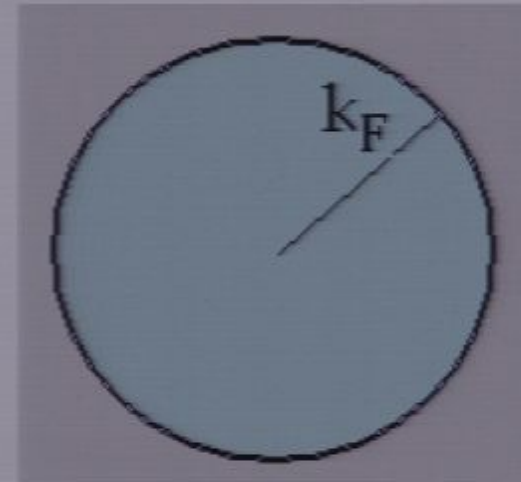
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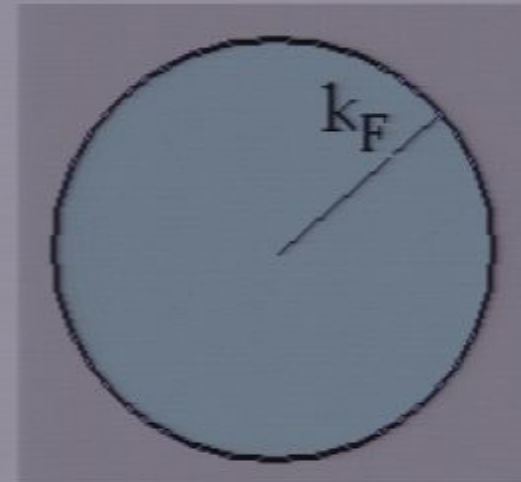
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$$G_R(\omega, \vec{k}) = \frac{Z}{\omega - v_F k_\perp + i\Gamma} + \dots, \quad k_\perp \equiv |\vec{k}| - k_F \quad \Gamma \sim \omega_*^2$$



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$$A(\omega, \vec{k}) \equiv \text{Im } G_R(\omega, \vec{k}) \xrightarrow{k_\perp \rightarrow 0} Z \delta(\omega - v_F k_\perp) \quad \text{with } Z \text{ finite}$$

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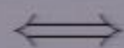
### *Organizing principle for non-Fermi liquids?*

# The AdS/CFT correspondence

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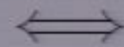


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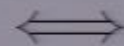
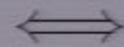
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black hole in AdS

finite chemical potential

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electrically charged black hole

## *AdS<sub>4</sub> – BH geometry*

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Relativistic CFT with gravity dual and conserved  $U(1)$  global symmetry

$$S = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} \left( \mathcal{R} + \frac{6}{R^2} - \frac{2\kappa^2}{g_F^2} F_{\mu\nu} F^{\mu\nu} + \dots \right)$$

Charged black hole solution

$$ds^2 = \frac{r^2}{R^2} (-f(r) dt^2 + d\vec{x}^2) + R^2 \frac{dr^2}{f(r)r^2}$$

$$f(r) = 1 + \frac{3}{r^4} - \frac{4}{r^3} \quad A = \mu \left( 1 - \frac{1}{r} \right) dt$$

where  $\mu$  = chemical potential, horizon at  $r = 1$ .



# ARPES

$$1 + \frac{Q^2}{\epsilon} - \frac{Q^2}{\epsilon}$$

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$$1 + \frac{Q^2}{4} - \frac{Q^2 + 1}{2}$$

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Introduce  $\psi$  spinor field in the  $AdS$  —  $BH$  background

$$S_{probe} = \int d^4x \sqrt{-g} (\bar{\psi}(\not{D} - m)\psi + interactions)$$

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- ▶ Solve the Dirac equation for the bulk spinor in  $AdS - BH$
- ▶ Impose infalling boundary conditions at the horizon
- ▶ Expand the solution at the boundary

$$\psi = (-gg^{rr})^{-1/4} e^{-i\omega t + ikx} \Psi \quad \Phi_\alpha = \frac{1}{2}(1 - (-1)^\alpha \Gamma^r \Gamma^t \Gamma^x) \Psi$$

$$\Phi_\alpha \stackrel{r \rightarrow \infty}{\approx} a_\alpha r^m \begin{pmatrix} 0 \\ 1 \end{pmatrix} + b_\alpha r^{-m} \begin{pmatrix} 1 \\ 0 \end{pmatrix} \quad G_\alpha(\omega, k) = \frac{b_\alpha}{a_\alpha} \quad \alpha = 1, 2$$

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**Holographic non-Fermi liquids**  
Superconducting phase

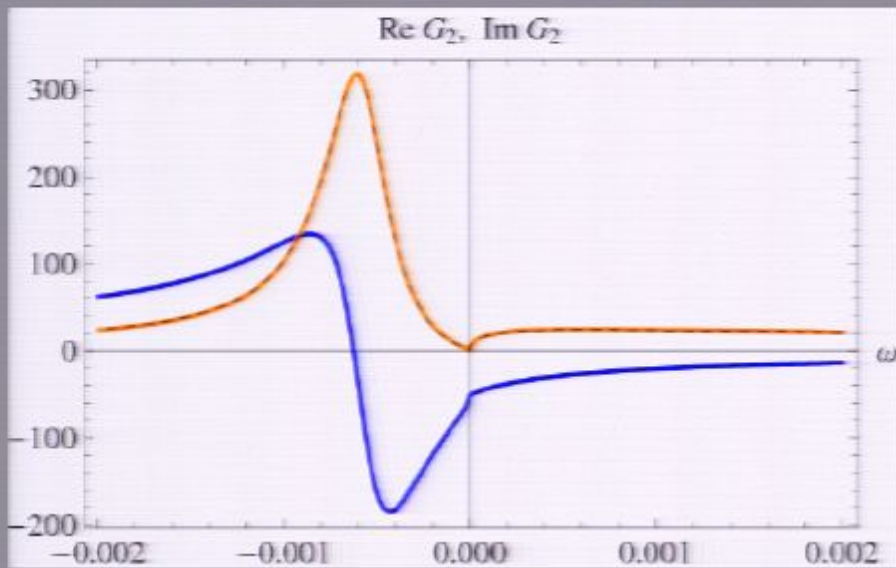
**Fermi surfaces**  
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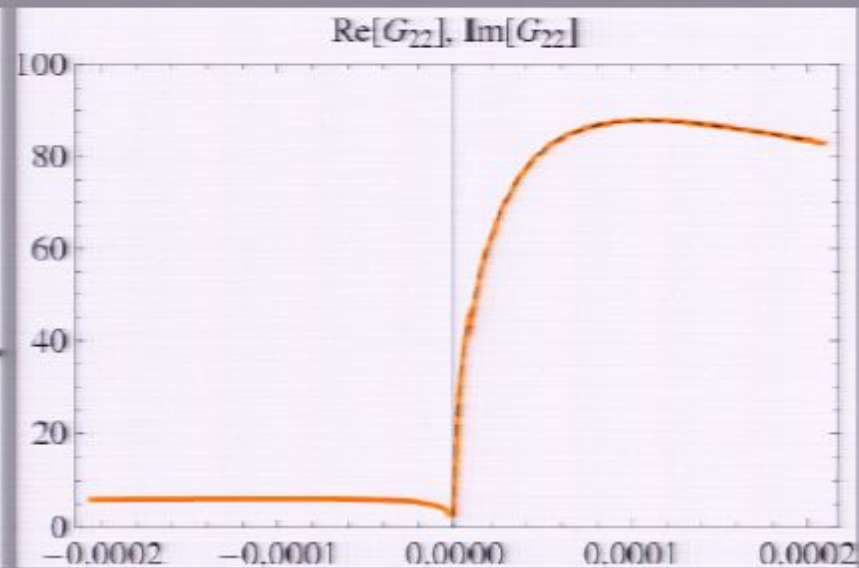


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At  $q = 1, \Delta = 3/2$  the numerical computation gives



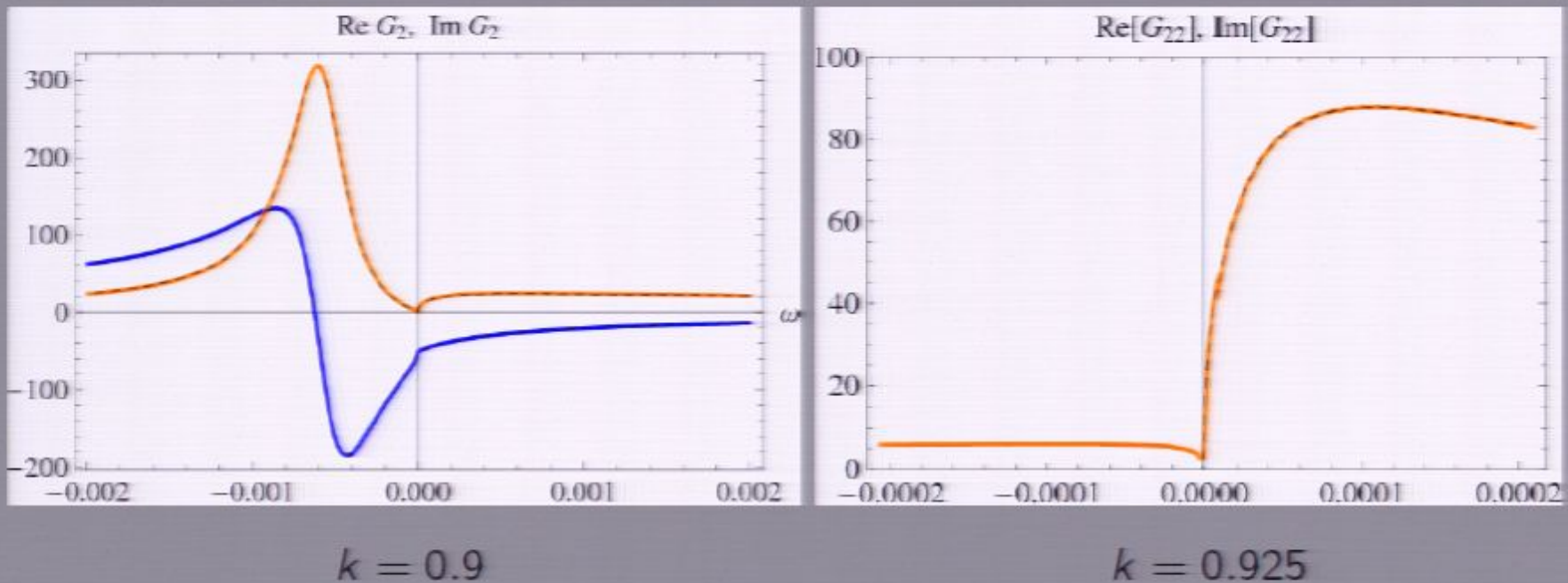
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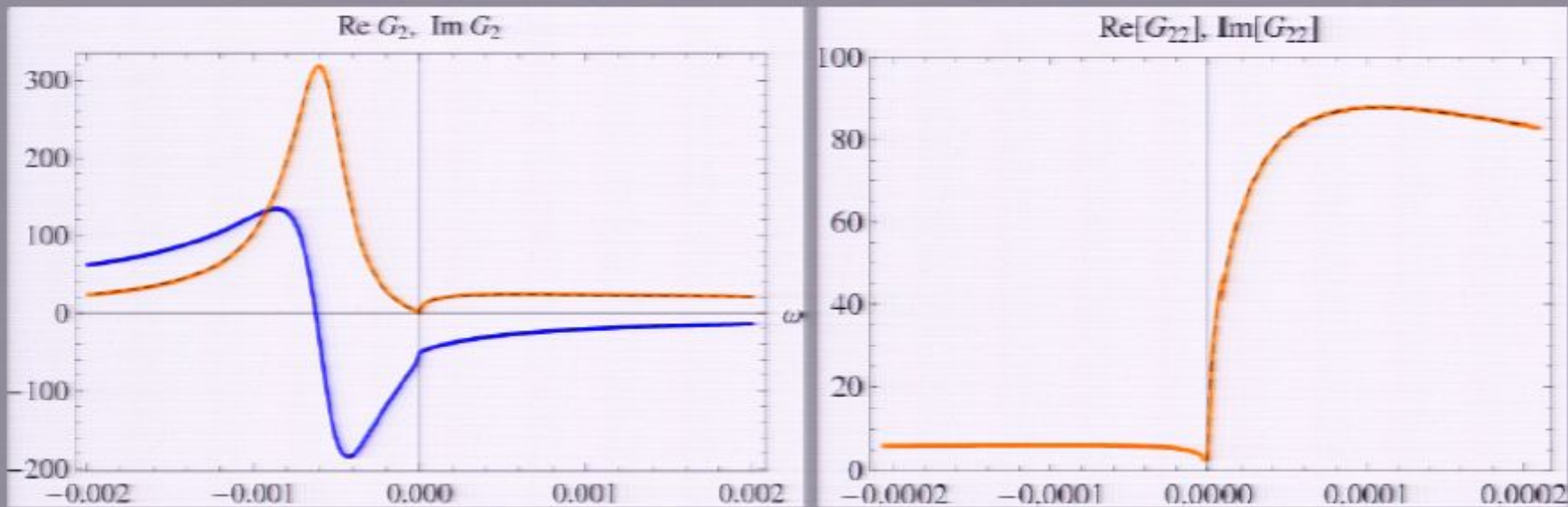
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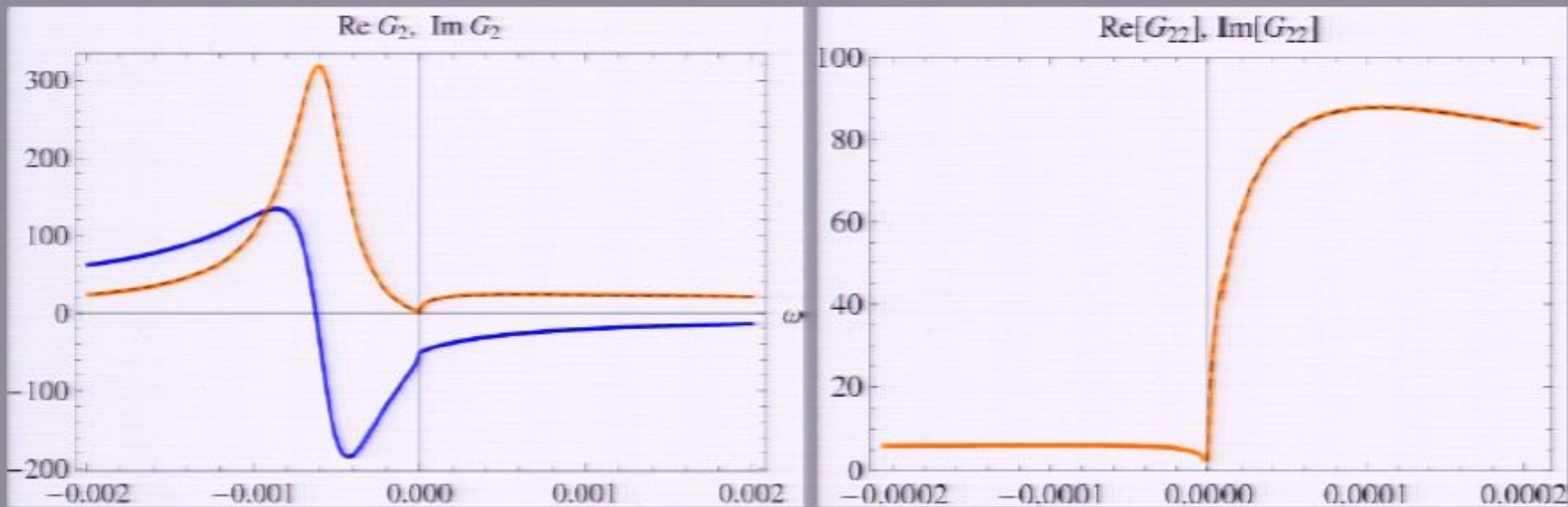
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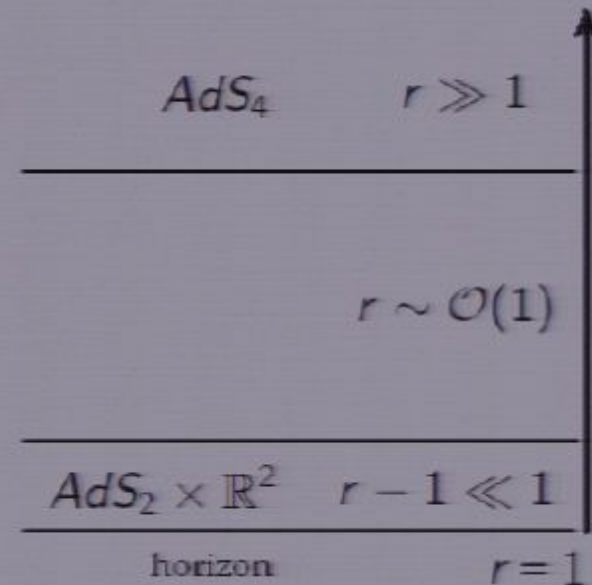
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- ▶ non-Fermi liquid!

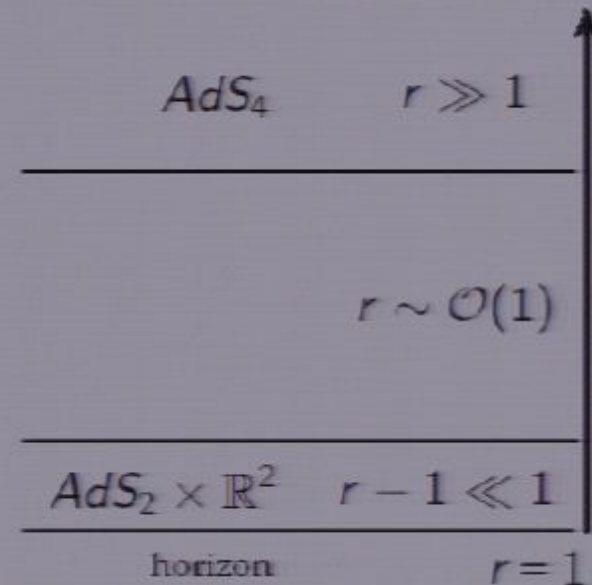
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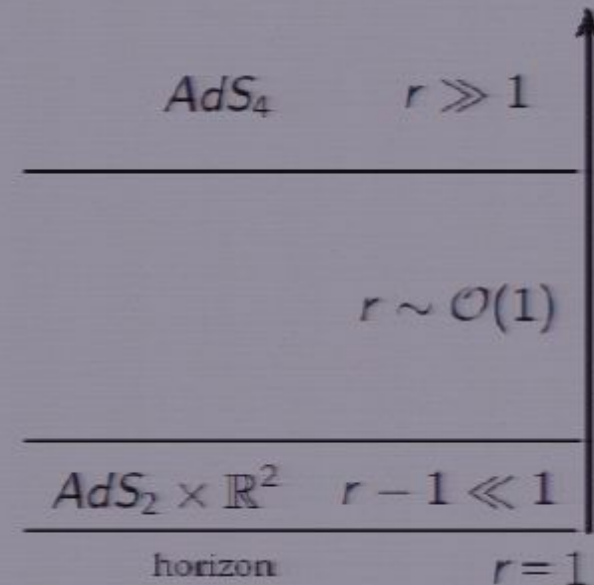


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► *Emergent IR CFT*

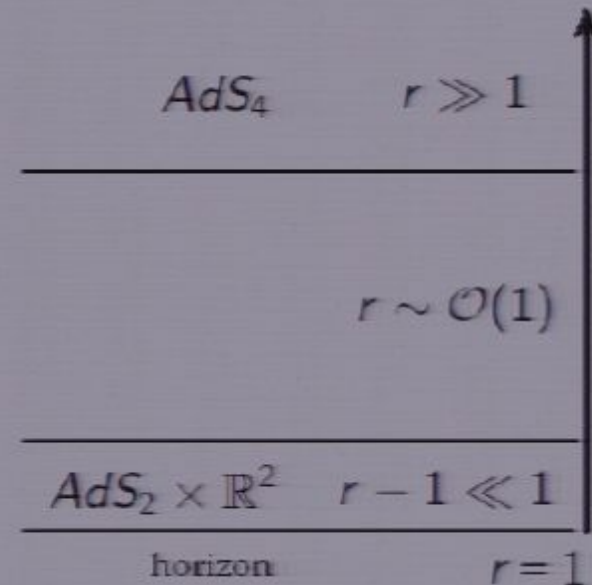
Gravity in  $AdS_2 \Leftrightarrow (0+1)$ -d CFT



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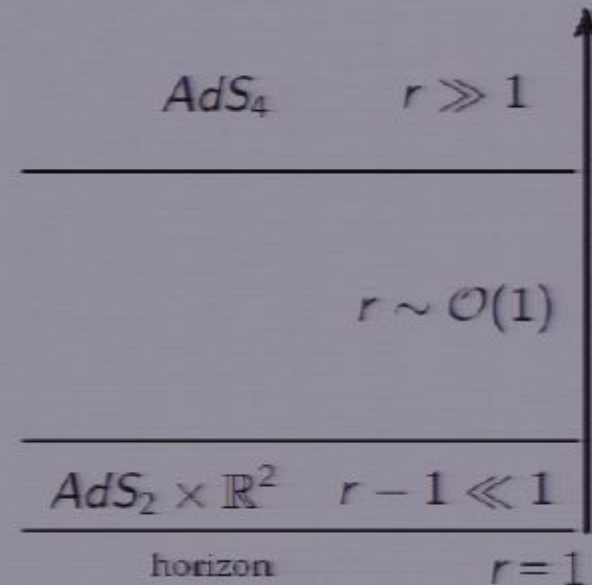




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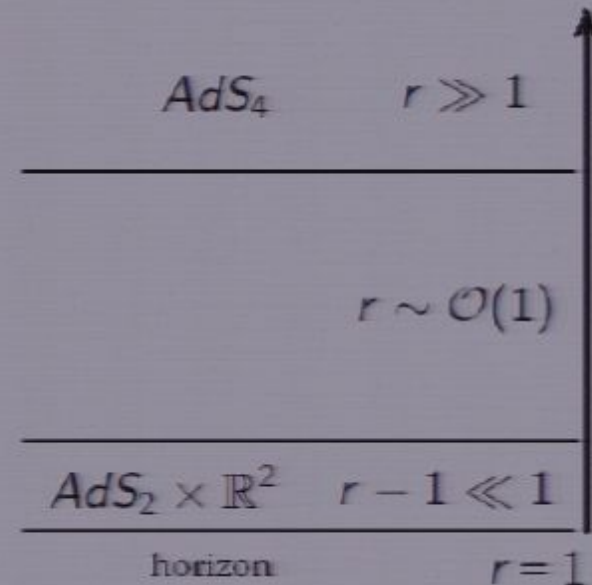
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▶ Conformal dimensions in IR

$$\delta_{\vec{k}} = \frac{1}{2} + \nu_{\vec{k}} \quad \nu_{\vec{k}} = \frac{1}{\sqrt{6}} \sqrt{(\Delta - \frac{3}{2})^2 + k^2 - \frac{q^2}{2}} \quad \mathcal{G}_{\vec{k}}(\omega) = c(k)\omega^{2\nu_{\vec{k}}}$$



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$\mathcal{G}_k(\omega) \in \mathbb{C}$  retarded correlator for  $\mathcal{O}_{\vec{k}}$  in IR CFT



$$G = c(k) \omega^{2D}$$



ARPES

$$1 + \frac{G^2}{2} + 1$$

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$$G_R(\omega, k) = \frac{b_+^{(0)} + \omega b_+^{(1)} + O(\omega^2) + \mathcal{G}_k(\omega) \left( b_-^{(0)} + \omega b_-^{(1)} + O(\omega^2) \right)}{a_+^{(0)} + \omega a_+^{(1)} + O(\omega^2) + \mathcal{G}_k(\omega) \left( a_-^{(0)} + \omega a_-^{(1)} + O(\omega^2) \right)}$$

$\mathcal{G}_k(\omega) \in \mathbb{C}$  retarded correlator for  $\mathcal{O}_{\vec{k}}$  in IR CFT

$a_{\pm}^{(0)}, a_{\pm}^{(1)}, b_{\pm}^{(0)}, b_{\pm}^{(1)} \in \mathbb{R}$  are  $k$ -dependent functions (from UV region)



# ARPES

$$x \approx \frac{\gamma}{\omega} x$$

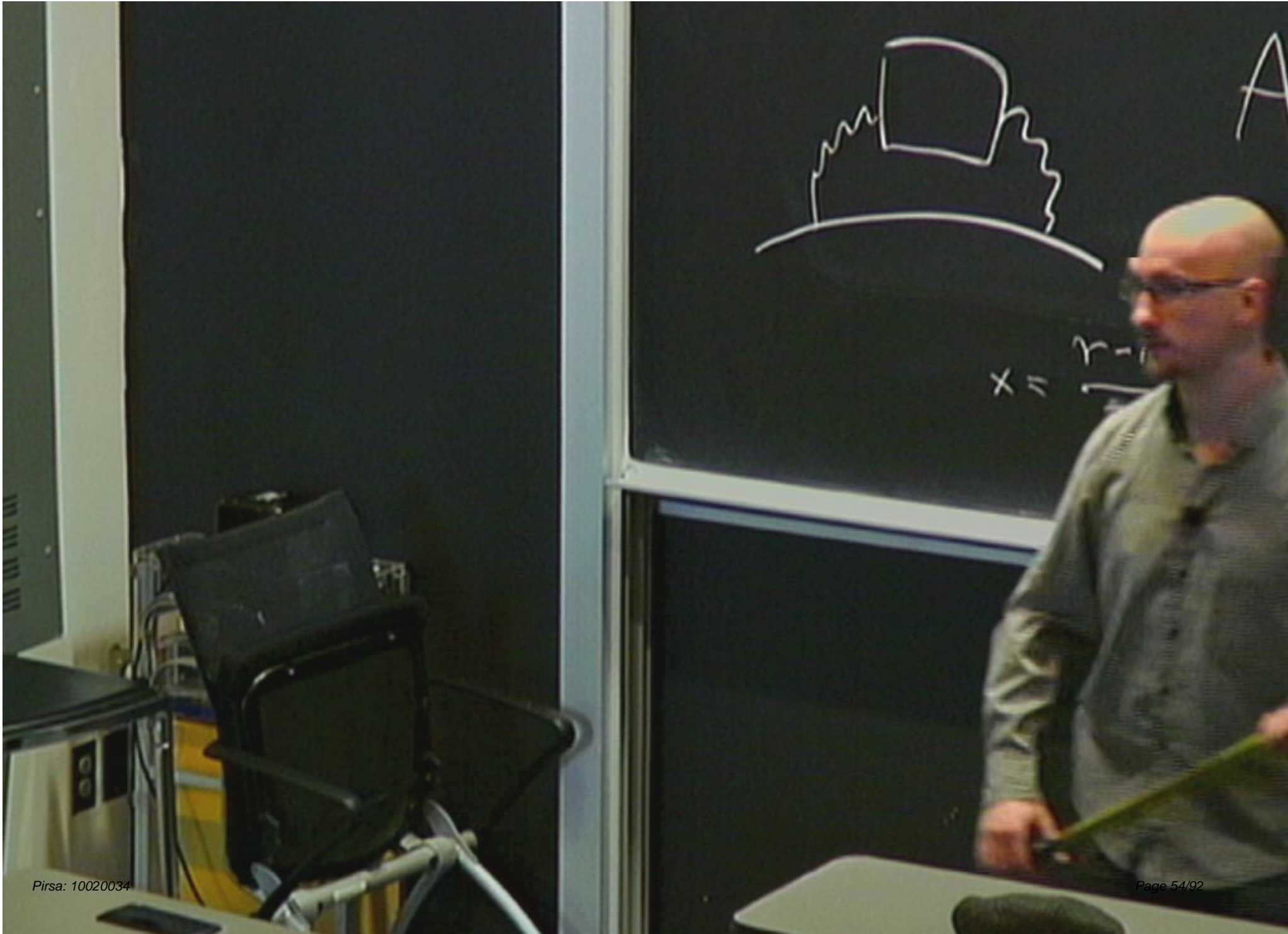
$$\frac{2}{\omega^2} - \frac{Q^2 + 1}{\omega^2}$$

# ARPES



$$x = \frac{\gamma - \gamma_*}{\omega}$$

$$1 + \frac{Q^2}{2} - \frac{Q^2 + 1}{2}$$



## Small $\omega$ expansion

- ▶ To understand scaling around FS, study low- $\omega$  behavior of correlators.
- ▶ For extremal black hole, this is not straightforward, because  $\omega$ -dependent terms in the Dirac equation always become singular near the horizon.

### Strategy

- ▶ Separate spacetime into UV region and the  $AdS_2 \times \mathbb{R}^2$  IR region.
- ▶ Perform small  $\omega$  expansions separately.
- ▶ Match them at the overlapping region.

### Result

$$G_R(\omega, k) = \frac{b_+^{(0)} + \omega b_+^{(1)} + O(\omega^2) + \mathcal{G}_k(\omega) \left( b_-^{(0)} + \omega b_-^{(1)} + O(\omega^2) \right)}{a_+^{(0)} + \omega a_+^{(1)} + O(\omega^2) + \mathcal{G}_k(\omega) \left( a_-^{(0)} + \omega a_-^{(1)} + O(\omega^2) \right)}$$

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## Fermi surfaces

$$G_R(\omega, k) = \frac{b_+^{(0)} + \omega b_+^{(1)} + O(\omega^2) + \mathcal{G}_k(\omega) \left( b_-^{(0)} + \omega b_-^{(1)} + O(\omega^2) \right)}{a_+^{(0)} + \omega a_+^{(1)} + O(\omega^2) + \mathcal{G}_k(\omega) \left( a_-^{(0)} + \omega a_-^{(1)} + O(\omega^2) \right)}$$



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Suppose that for some  $k_F$ :  $a_+^{(0)}(k_F) = 0$

Then, at small  $\omega, k_\perp$  we have

$$G_R(k, \omega) = \frac{h_1}{k_\perp - \frac{1}{v_F}\omega - h_2 \mathcal{G}_{k_F}(\omega)} + \dots$$

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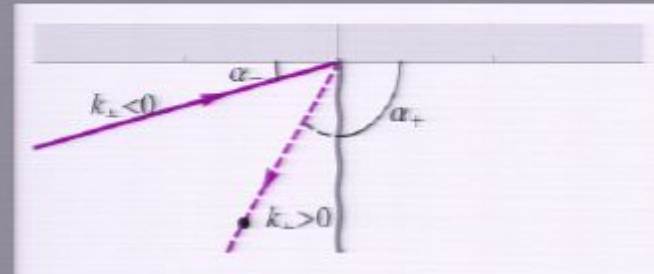
$$\mathcal{G}_k(\omega) = c(k)\omega^{2\nu_k} \quad \nu_{\vec{k}} = \frac{1}{\sqrt{6}} \sqrt{m^2 + k^2 - \frac{q^2}{2}} \quad h_1, h_2, v_F \in \mathbb{R}$$

This is the quasiparticle peak we saw earlier.

# Fermi surfaces: singular and non-singular

Suppose  $\nu_k < \frac{1}{2}$

$$G_R(k, \omega) = \frac{h_1}{k_{\perp} - h_2 G_{k_F}(\omega)} + \dots$$

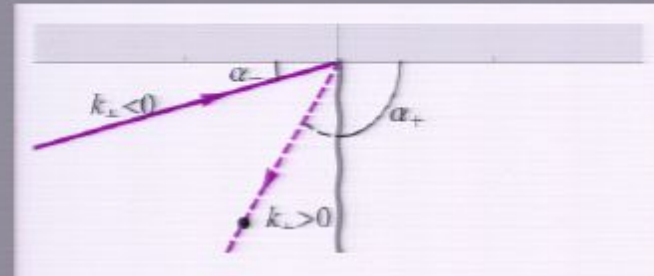


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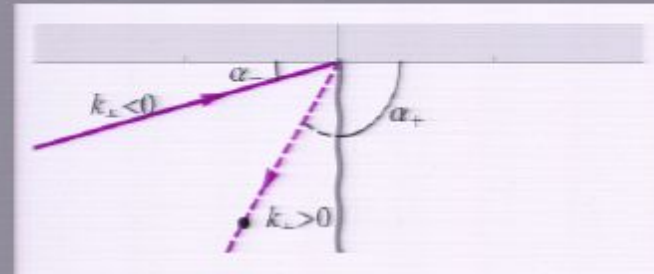
$$\omega_*(k) \sim k_\perp^z \quad z = \frac{1}{2\nu_{k_F}} > 1 \quad \frac{\Gamma(k)}{\omega_*(k)} = \text{const} \quad Z \sim k_\perp^{\frac{1-2\nu_{k_F}}{2\nu_{k_F}}} \rightarrow 0, k_\perp \rightarrow 0$$



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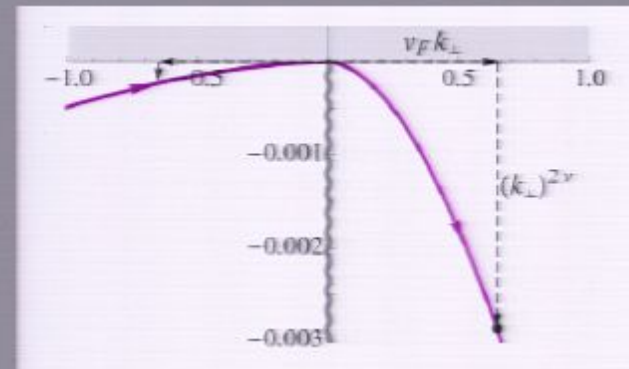
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$$v \sim \sqrt{m^2 + k^2} - \frac{q^2}{2}$$

$$G = c(k) \omega^{2\nu}$$



# ARPES

$$1 + \frac{Q^2}{2^4} - \frac{Q^2 + 1}{\pi^5}$$

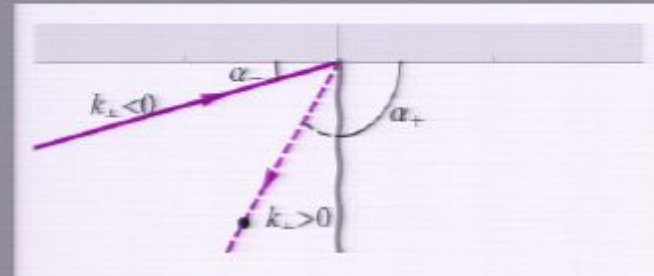
x =

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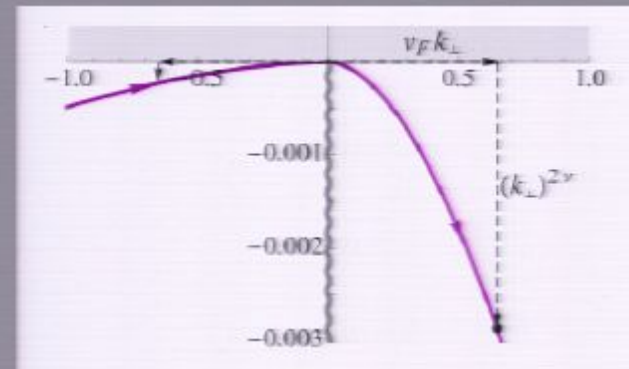
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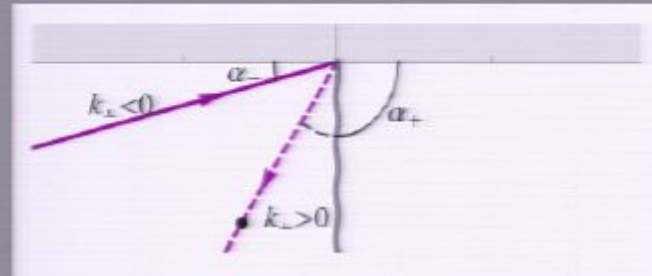
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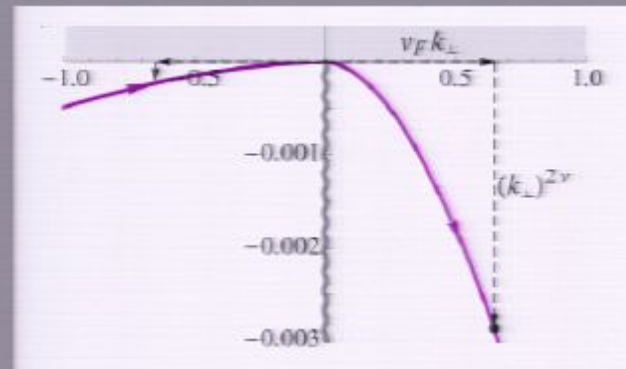
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$$\omega_*(k) \sim v_F k_\perp \quad \frac{\Gamma(k)}{\omega_*(k)} = k_\perp^{2\nu_{k_F}-1} \rightarrow 0 \quad Z = h_1 v_F$$



## Fermi surfaces: Marginal Fermi liquid

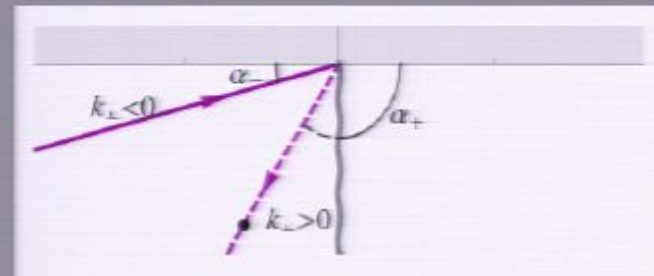
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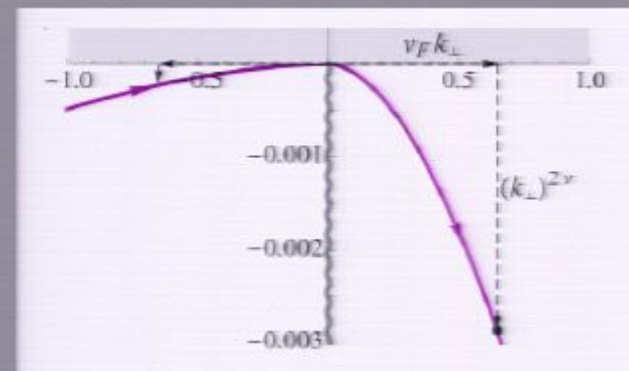
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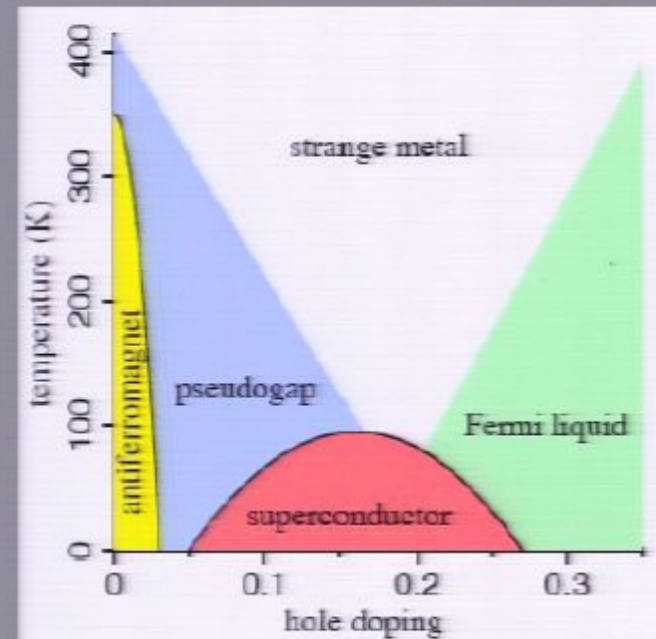
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$$G_R(k, \omega) = \frac{h_1}{k_{\perp} + \tilde{c}_1 \omega \log \omega + c_1 \omega} + \dots$$

where  $\tilde{c}_1 \in \mathbb{C}$ ,  $c_1 \in \mathbb{R}$ .

This is the Marginal Fermi liquid  
Green's function [Varma, 1989]



(Non-)Fermi liquids  
AdS/CFT background  
**Holographic non-Fermi liquids**  
Superconducting phase

Fermi surfaces  
Near-horizon  $AdS_2$  and emergent IR CFT  
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$$\sqrt{v^2 + k^2} - \frac{\omega}{2}$$

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$\omega$



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$$1 + \frac{Q^2}{2^4} - \frac{Q^2}{r}$$

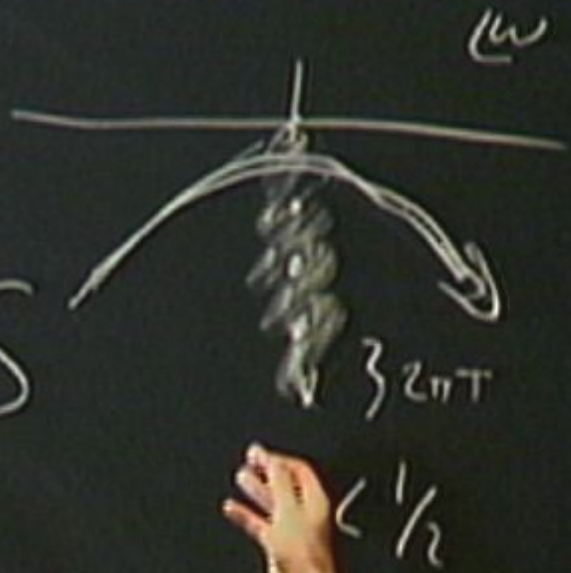
$$x = \frac{r - r_x}{r}$$



$$\sqrt{m^2 + k^2} - \alpha/2$$

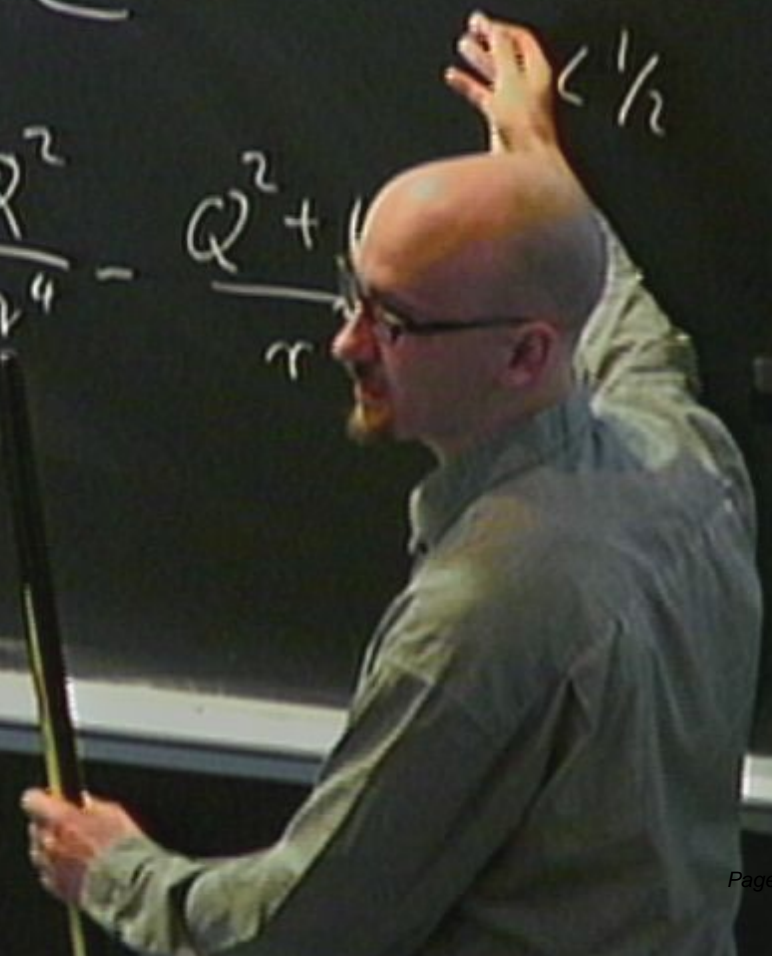
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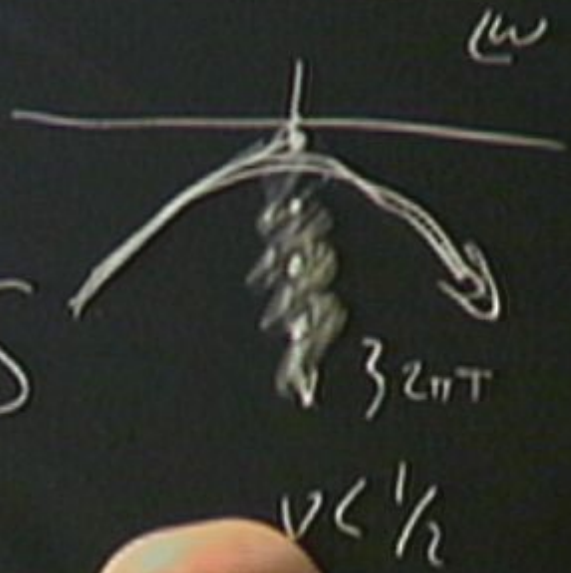
$$x = \frac{\gamma - \gamma_k}{\Gamma}$$

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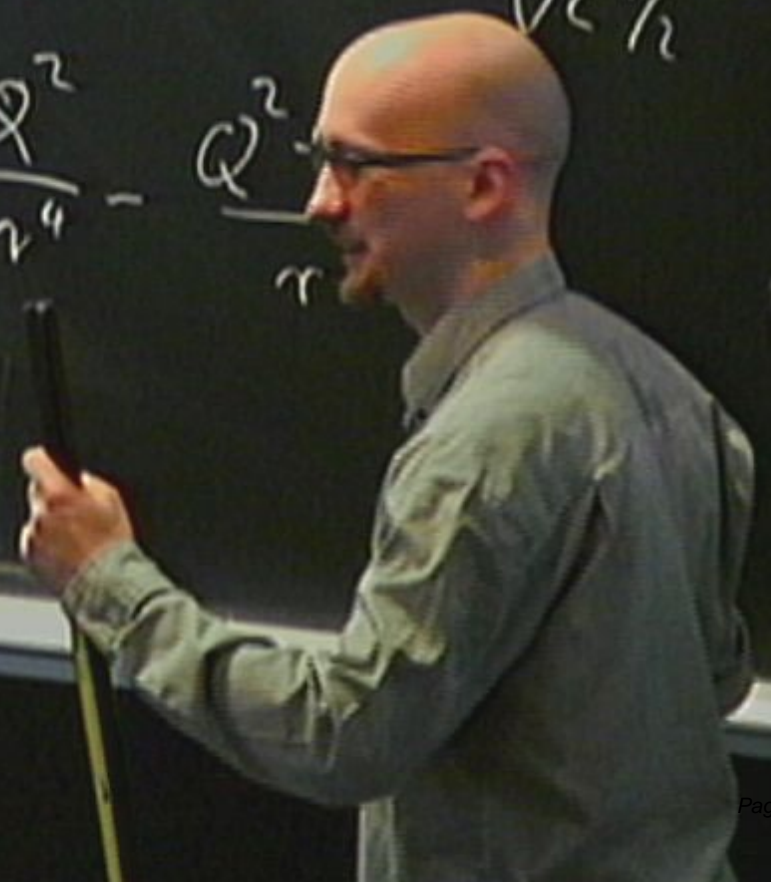


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$$x = \frac{r - r_*}{\hbar}$$

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## The gap

$$\kappa^2 \mathcal{L} = R + \frac{6}{L^2} - \frac{1}{4}(dA)^2 - |(\nabla - iq_\varphi A)\varphi|^2 - m_\varphi^2 |\varphi|^2$$

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$$ds^2 = -g(r)e^{-\chi(r)} dt^2 + \frac{dr^2}{g(r)} + r^2(dx^2 + dy^2)$$

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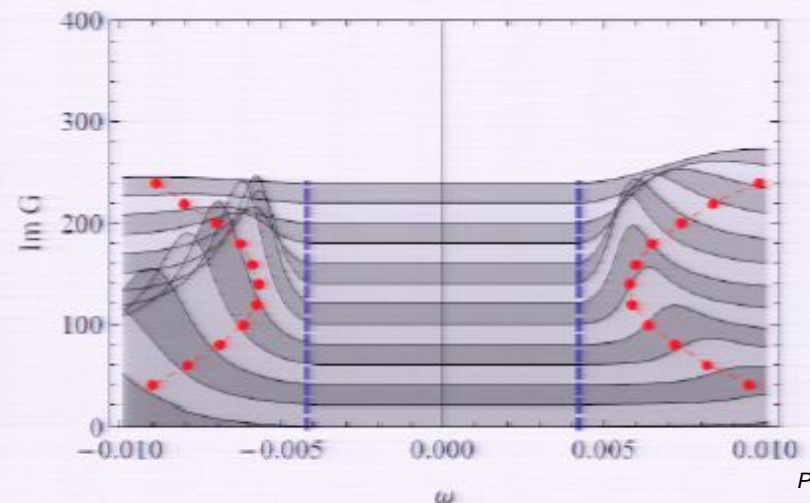
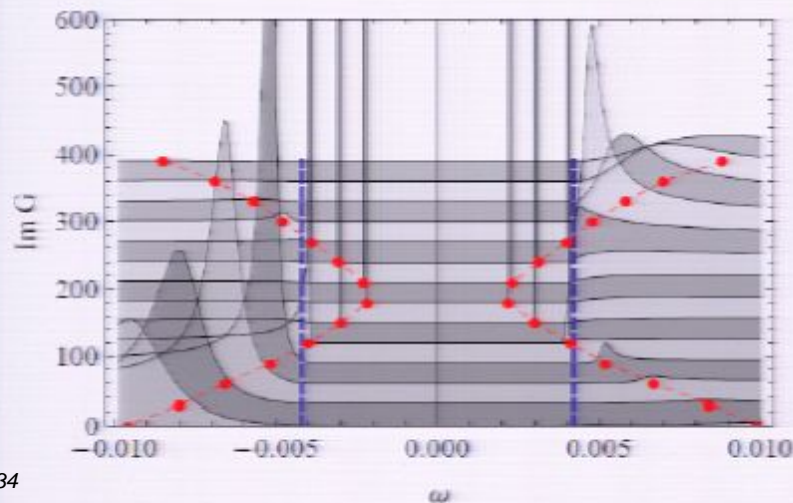
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