

Title: Effective Field Theory Approach to the Post-Newtonian Physics

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Abstract:

# *Effective Field Theory approach to the Post-Newtonian physics*

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**Hebrew University, Jerusalem**

# Outline

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## ***1. Einstein-Infeld-Hoffman Lagrangian (Improved derivation)***

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**2. Unrealistic scalar toy model**

- *redundant operators*
- *classical RG flow*

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**2. Unrealistic scalar toy model**

- *redundant operators*
- *classical RG flow*

**3. Finite size effects (work in progress)**

**4. Results and open questions**

# *Post-Newtonian approximation*

# Post-Newtonian approximation





# Post-Newtonian approximation



$$r_s \ll R$$

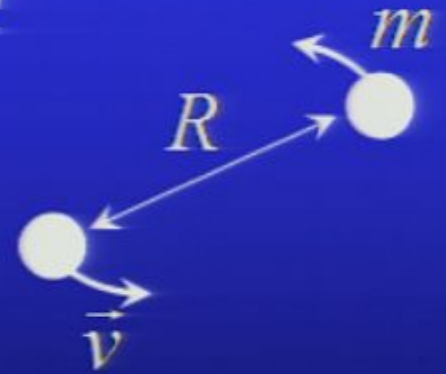
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# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

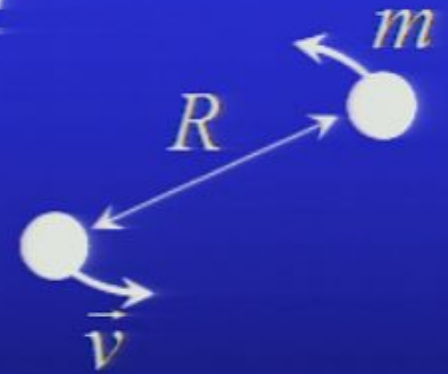


# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

(W. Goldberger and I. Rothstein)

Virial theorem:  $\vec{v}^2 \sim \frac{Gm}{R} \sim \frac{r_s}{R} = 1PN$

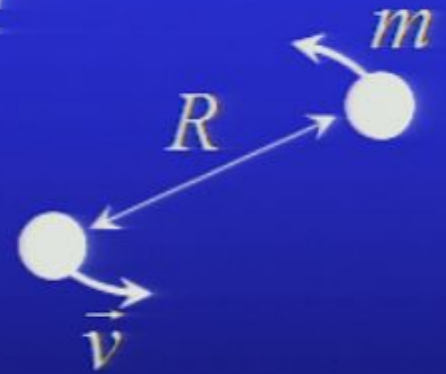


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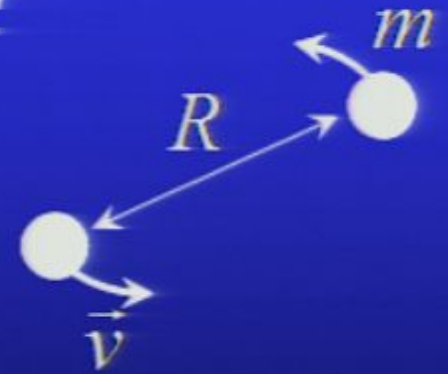
Typical lengths:

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Typical lengths:

Corresponding fields:

short length  $r_s$

strong

potential length  $R$

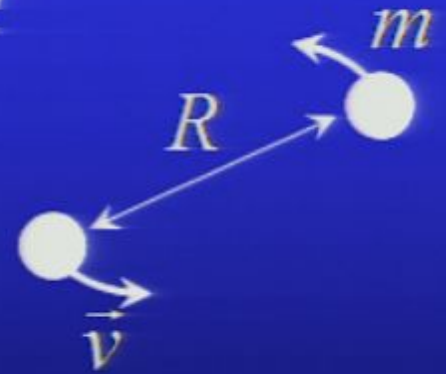
radiation length  $\frac{R}{v}$

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Typical lengths:

Corresponding fields:

short length  $r_s$

strong field  $g_{\mu\nu}$

potential length  $R$

potential field  $H_{\mu\nu}$

radiation length  $\frac{R}{v}$

radiation field  $h_{\mu\nu}$

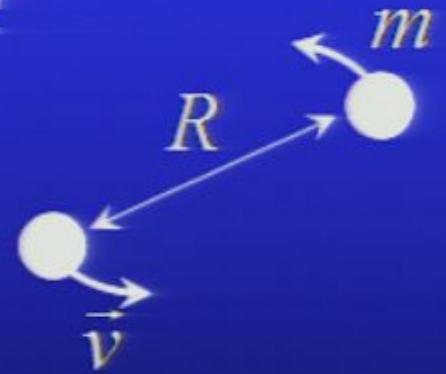
$$\frac{\partial H_{\mu\nu}}{\partial x^i} \sim \frac{1}{R} H_{\mu\nu}, \quad \frac{\partial H_{\mu\nu}}{\partial t} \sim \frac{v}{R} H_{\mu\nu}$$

$$\frac{\partial h_{\mu\nu}}{\partial x^i} \sim \frac{v}{R} h_{\mu\nu}, \quad \frac{\partial h_{\mu\nu}}{\partial t} \sim \frac{v}{R} h_{\mu\nu}$$

# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

## Step 1





# Post-Newtonian approximation

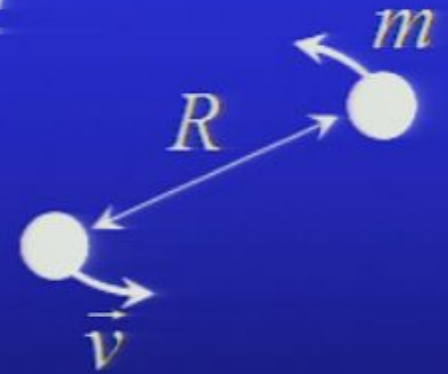
Einstein-Infeld-Hoffman Lagrangian derivation:

## Step 1

"Integrate out  $r_s$ "

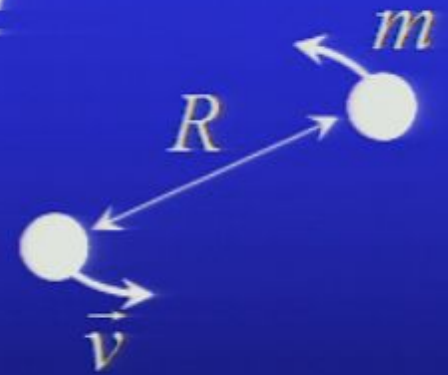
$$S = -\frac{1}{16\pi G} \int \sqrt{-\bar{g}} d^4x R[\bar{g}] - \sum_{a=1}^2 m_a \int d\tau_a + \dots$$

where  $\bar{g}_{\mu\nu} = \eta_{\mu\nu} + H_{\mu\nu} + h_{\mu\nu}$  and  $d\tau_a^2 = \bar{g}_{\mu\nu} dx_a^\mu dx_a^\nu$



# Post-Newtonian approximation

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## Step 2

Apply temporal dimensional reduction ansatz:

$$ds^2 = e^{2\phi} (dt - A_i dx^i)^2 - e^{-2\phi} \gamma_{ij} dx^i dx^j$$

This defines a change of variables  $\bar{g}_{\mu\nu} \rightarrow (\gamma_{ij}, A_i, \phi)$

# Post-Newtonian approximation

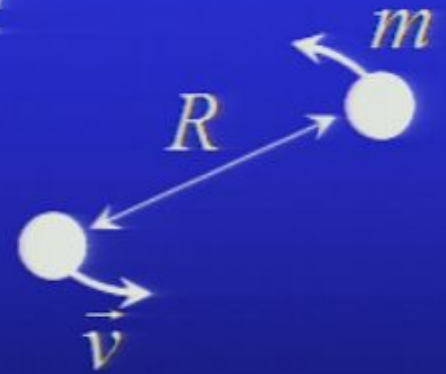
Einstein-Infeld-Hoffman Lagrangian derivation:

As a result we get

$$S_{pp} = -\sum_{a=1}^2 m_a \int dt e^{\phi} \sqrt{(1 - \vec{A} \cdot \vec{v}_a)^2 - e^{-4\phi} \gamma_{ij} v_a^i v_a^j} + \dots$$

$$S_{bulk} = -\frac{1}{16\pi G} \int \sqrt{\gamma} d^4x \left[ R[\gamma] + 2(\nabla\phi)^2 - \frac{1}{4} e^{4\phi} F_{ij} F^{ij} \right] + \frac{1}{8\pi G} \int d^4x \dot{\phi}^2 + \dots$$

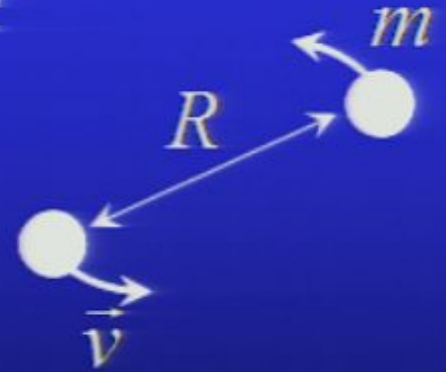
where  $F_{ij} = \partial_i A_j - \partial_j A_i$  is the gravito-magnetic field strength



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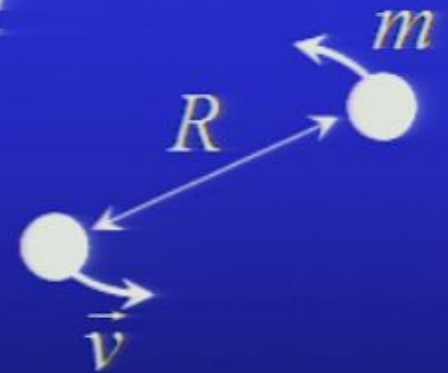
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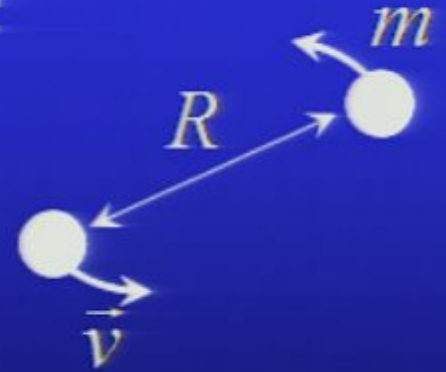
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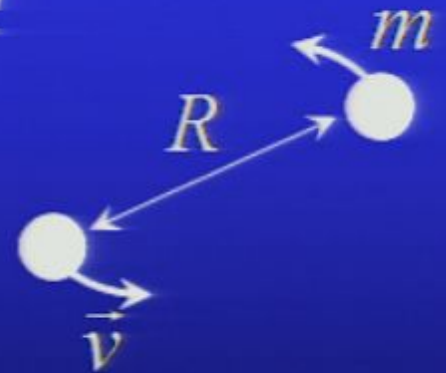
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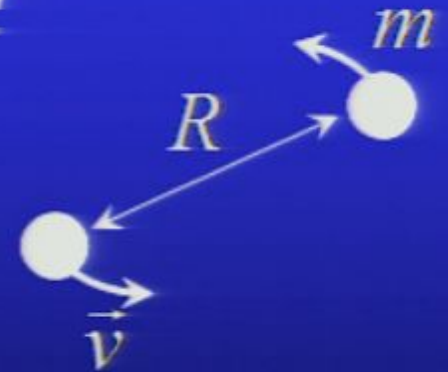


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*Exact bulk action within  
the stationary approximation*

*Time derivative term  
treated as perturbation*

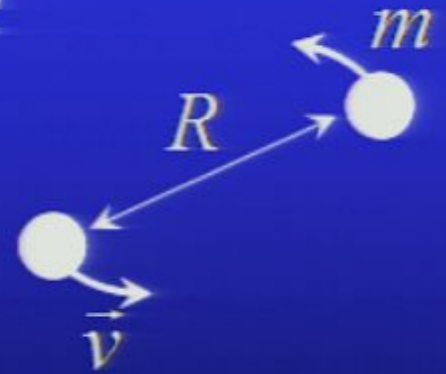
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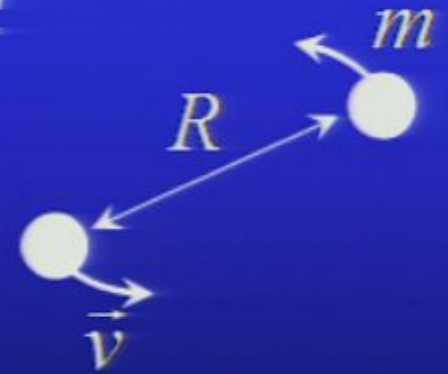
Step 3



# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

## Step 3



*Decompose all the fields into potential and radiation components*

$$\left( \gamma_{ij}, A_i, \phi \right) \rightarrow \left( \gamma_{ij} + \bar{\gamma}_{ij}, A_i + \bar{a}_i, \phi + \bar{\phi} \right)$$

*and integrate out the potential modes in order to obtain an action for two interacting massive particles which depends on their velocities as well as their locations*

# Post-Newtonian approximation

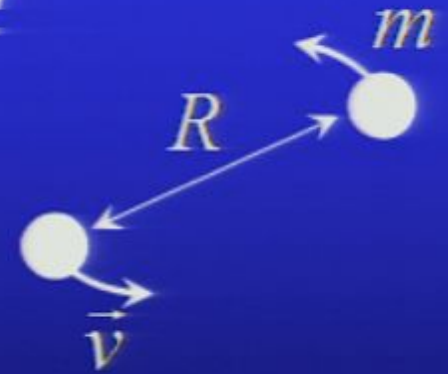
Einstein-Infeld-Hoffman Lagrangian derivation:

$$S_{bulk} = -\frac{1}{8\pi G} \int d^4x \left[ (\nabla\phi)^2 - \frac{1}{8} F_{ij} F^{ij} - \dot{\phi}^2 \right] + \dots$$

# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

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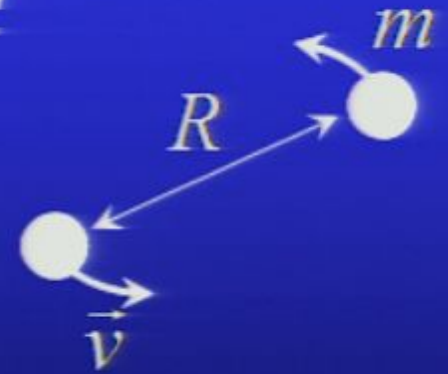
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$$S_{pp} = -\sum_{a=1}^2 m_a \int dt \left( 1 - \frac{\vec{v}_a^2}{2} - \frac{(\vec{v}_a^2)^2}{8} + \phi - \vec{A} \cdot \vec{v}_a + \frac{3}{2} \vec{v}_a^2 \phi + \frac{\phi^2}{2} \dots \right)$$

*Integrating out the potential modes leads to*

# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

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Integrating out the potential modes leads to

$$S_{eff}(\vec{x}_a, \vec{v}_a) = \sum_{a=1}^2 \int dt \frac{m_a \vec{v}_a^2}{2} + \text{---} + \sum_{a=1}^2 \int dt \frac{m_a (\vec{v}_a^2)^2}{8} + \text{---} + \text{---}$$



# Post-Newtonian approximation


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$$S_{bulk} = -\frac{1}{8\pi G} \int d^4x \left[ (\nabla\phi)^2 - \frac{1}{8} F_{ij} F^{ij} - \dot{\phi}^2 \right] + \dots$$

$$\frac{\phi}{\phi} = 4\pi G \delta(t_1 - t_2) \int \frac{d^3k}{(2\pi)^3} \frac{e^{i\vec{k}(\vec{x}_1 - \vec{x}_2)}}{k^2}$$


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Einstein-Infeld-Hoffman Lagrangian derivation:

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# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

$$S_{eff}(\vec{x}_a, \vec{v}_a) = \sum_{a=1}^2 \int dt \frac{m_a \vec{v}_a^2}{2} + \left[ \text{Diagram: two vertical lines connected by a horizontal line} \right] + \sum_{a=1}^2 \int dt \frac{m_a (\vec{v}_a^2)^2}{8} + \left[ \text{Diagram: two vertical lines connected by a dashed horizontal line} \right] + \left[ \text{Diagram: two vertical lines connected by a triangle} \right] + \left[ \text{Diagram: two vertical lines connected by a horizontal line with a circle containing an 'x' in the middle} \right] + \left[ \text{Diagram: two vertical lines connected by a horizontal line with 'v^2' above it} \right]$$

# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

$$L_{\text{EIH}}(\vec{x}_a, \vec{v}_a) = \frac{1}{8} \sum_{a=1}^2 m_a (\vec{v}_a^2)^2 + v^2 \left[ \text{---} + \text{---} \otimes \text{---} + \text{---} \text{---} + \text{---} \right]$$

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$$\frac{3Gm_1m_2}{2r} (\vec{v}_1^2 + \vec{v}_2^2)$$
  

$$\frac{Gm_1m_2}{2r} \vec{v}_{1\perp} \cdot \vec{v}_{2\perp}$$
  

$$-4 \frac{Gm_1m_2}{r} \vec{v}_1 \cdot \vec{v}_2$$
  

$$\frac{G^2 m_1 m_2 (m_1 + m_2)}{2r^2}$$

# Post-Newtonian approximation

Einstein-Infeld-Hoffman Lagrangian derivation:

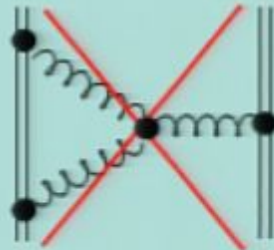
$$\mathcal{L}_{EIH}(\vec{x}_a, \vec{v}_a) = \frac{1}{8} \sum_{a=1}^2 m_a (\vec{v}_a^2)^2 + \frac{Gm_1m_2}{2r} \left[ 3(\vec{v}_1^2 + \vec{v}_2^2) - 8\vec{v}_1 \cdot \vec{v}_2 + \vec{v}_{1\perp} \cdot \vec{v}_{2\perp} \right] - \frac{G^2 m_1 m_2 (m_1 + m_2)}{2r^2}$$

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**Improvement:** The pay-off of dimensional reduction  $\bar{g}_{\mu\nu} \rightarrow (\gamma_{\bar{\mu}\bar{\nu}}, A_{\bar{\mu}\bar{\nu}}, \phi)$  is that the triple vertex diagram gets eliminated and there is no need to cope with a loop diagram containing voluminous 3-graviton vertex.





$$(\delta\psi)^{\leftarrow} - \frac{\lambda}{4!} \psi^{\leftarrow} - \rho\psi$$

$$\rho = \frac{1}{\pi^{3/2} r_s^3} \exp\left[-\left(\frac{r}{r_s}\right)^2\right]$$

$$= (\nabla\psi)^2 - \frac{\lambda}{4!} \psi^4 - g\psi$$

$$g = \int \frac{d^3x}{(2\pi)^3} \int_0^\infty r_s^3 \exp\left[-\left(\frac{r_s}{\tau_s}\right)^2\right]$$

$$\psi = (\psi)^k - \frac{\lambda}{4!} \psi^4 - g \psi$$

$$g = \frac{1}{\pi^{3/2} r_s^3} \exp\left[-\left(\frac{r}{r_s}\right)^2\right]$$

$$\mathcal{L} = (\nabla\varphi)^2 - \frac{\lambda}{4!} \varphi^4 - g\varphi$$

$$g = \frac{1}{\pi^{3/2} r_s^3} \exp\left[-\left(\frac{r}{r_s}\right)^2\right]$$

$$\mathcal{L} = (\partial\varphi)^2 - \frac{\lambda}{4!} \varphi^4 - g\varphi \quad g = \frac{g_1}{\pi^{3/2} r_s^3} \exp\left[-\left(\frac{r}{r_s}\right)^2\right]$$

$$\mathcal{L}_{\text{EFT}} = (\partial\varphi)^2 - \frac{\lambda}{4!} \varphi^4 - g_0 \int d\tau \varphi + c \int \square\varphi d\tau$$

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$$\varphi \rightarrow \varphi + \frac{1}{3} \delta(\mathbf{x} - \mathbf{x}(t))$$

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$$\psi \rightarrow \psi + \frac{1}{3} \delta(x - x(t))$$

$$\int (\partial\psi)^2 = \int \psi \square\psi \approx \frac{1}{3} \int \square\psi d\tau$$



$$= (\hbar\psi)^2 - \frac{\lambda}{4!} \psi^4 - g\psi \quad g = \frac{\bar{g}}{\pi^{\frac{3}{2}} r_s^3} \exp\left[-\left(\frac{r}{r_s}\right)^2\right]$$

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$$\int (\hbar\psi)^2 = \int \psi \square \psi \supset \frac{1}{3} \int \square \psi dt$$

$$(\psi + \delta) \square (\psi + \delta)$$

$$\mathcal{L} = (\partial\varphi)^2 - \frac{\lambda}{4!} \varphi^4 - g\varphi \quad g = \frac{g_0}{\pi^{\frac{3}{2}} r_s^3} \exp\left[-\left(\frac{r}{r_s}\right)^2\right]$$

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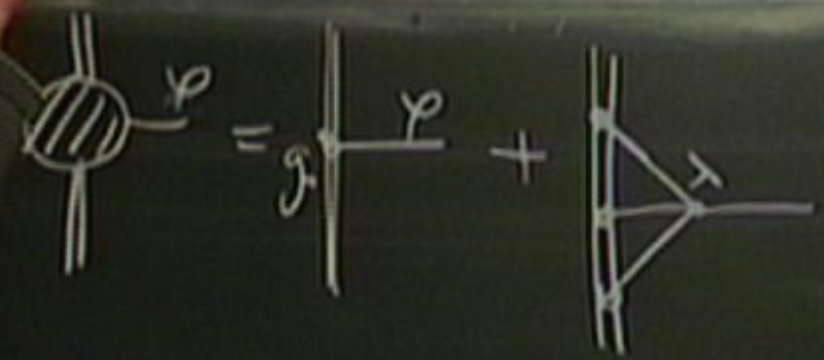
$$\varphi \rightarrow \varphi + \frac{1}{3} \delta(x - x_0(t)) \quad m \int d\tau + c_2 \int R d\tau + c_3 \int R_{\mu\nu} \gamma^{\mu\nu} d\tau$$

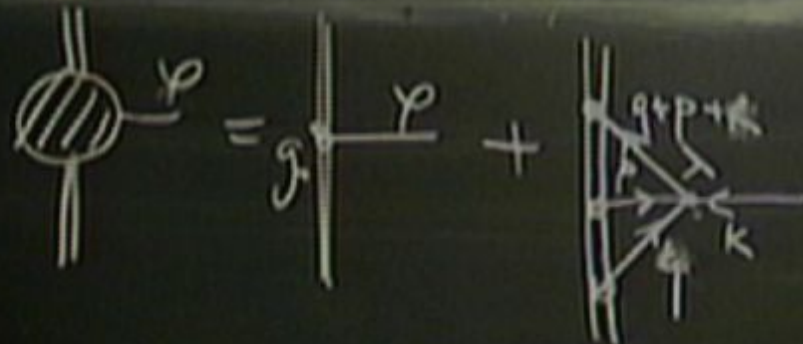
$$\int (\partial\varphi)^2 = \int \varphi \square \varphi \supset \frac{1}{3} \int \square \varphi d\tau$$

$$(p+\pi) \square (p+\delta b)$$

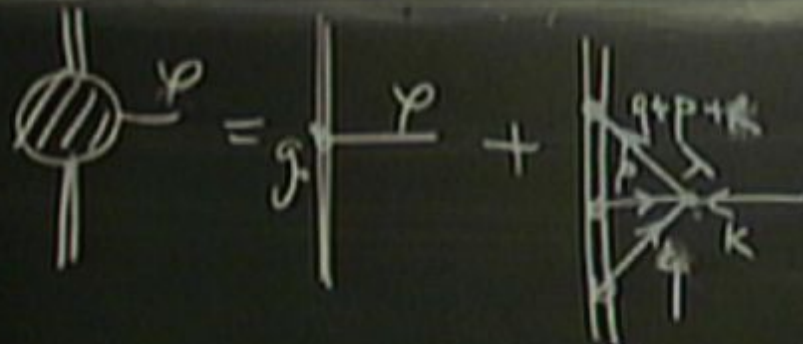
$(p+\sqrt{p}) \square (p+\sqrt{p})$

$\sim R_{10} \times \dots$





$$\frac{g_0^3 \lambda}{6} \int \frac{d^3 k}{(2\pi)^3} \phi(k) \int \frac{d^3 p d^3 q}{(2\pi)^6} \frac{1}{p^2 q^2 (q+p+k)^2}$$



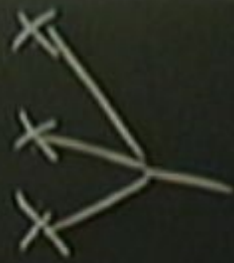
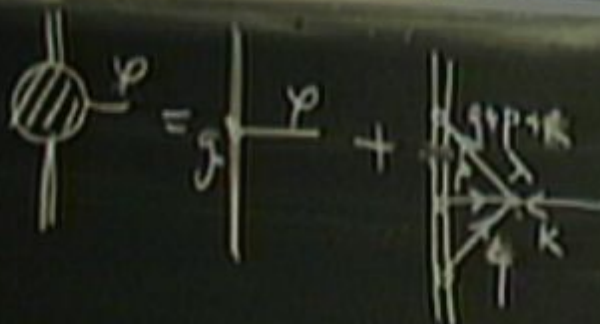
$$\frac{g_0^3 \lambda}{6} \int \frac{d^3 k}{(2\pi)^3} \phi(k) \int \frac{d^3 p d^3 q}{(2\pi)^6} \frac{1}{p^2 q^2 (q+p+k)^2}$$

$$\delta S_{ct} = - \frac{g_0^3 \lambda}{(2\pi)^6} \left( \frac{1}{\epsilon} + 3 - \gamma + \ln(4\pi) \right)$$

$$g_0 = g L^{\epsilon} \left[ 1 + \frac{g^2 \lambda}{2(4\pi)^2} \left( \frac{1}{\epsilon} + 3 - \gamma + \ln 4\pi \right) \right]$$

$$\epsilon = 4 - d$$

$$g_0 = g L^{\epsilon} \quad 0 = \frac{dg}{d \log L} \Rightarrow \beta = - \frac{g^3 \lambda}{96 \pi^2}$$



$$\frac{g_0^3 \lambda}{6} \underbrace{\int \frac{d^3 k}{(2\pi)^3} \epsilon(k) \int \frac{d^3 p d^3 q}{(2\pi)^6}}_{\left( \frac{1}{\epsilon} + 3 - \gamma + \ln \frac{\mu}{m} \right)} \frac{1}{p^2 q^2 (q+p+k)^2}$$

$$\mathcal{L}_{\text{ct}} = - \frac{g_0^3 \lambda}{(2\pi)^6} \left( \frac{1}{\epsilon} + 3 - \gamma + \ln \frac{\mu}{m} \right)$$

$$g_0 = g L^{-\epsilon} \left[ 1 + \frac{g^2 \lambda}{12(4\pi)^2} \left( \frac{1}{\epsilon} + 3 - 8 + \ln 4\pi \right) \right]$$

$$\epsilon = 4 - d$$

$$g_0 = g L^{-\epsilon} \quad 0 = \frac{dg_0}{d \log L} \Rightarrow \beta = -\frac{g^3 \lambda}{96\pi^2}$$

$$\mathcal{L}_{\text{EFT}} = g \int \frac{d^3 L}{(2\pi)^3} \phi(K)$$

$$+ \frac{\lambda g^3}{12(4\pi)^2} \int \frac{d^3 K}{(2\pi)^3} \phi(K) \ln[(KL)^2]$$



$$g_0 = g L^{-\epsilon} \left[ 1 + \frac{g^2 \lambda}{2(4\pi)^2} \left( \frac{1}{\epsilon} + 3 - \gamma + \ln 4\pi \right) \right]$$

$$\epsilon = 4 - d$$

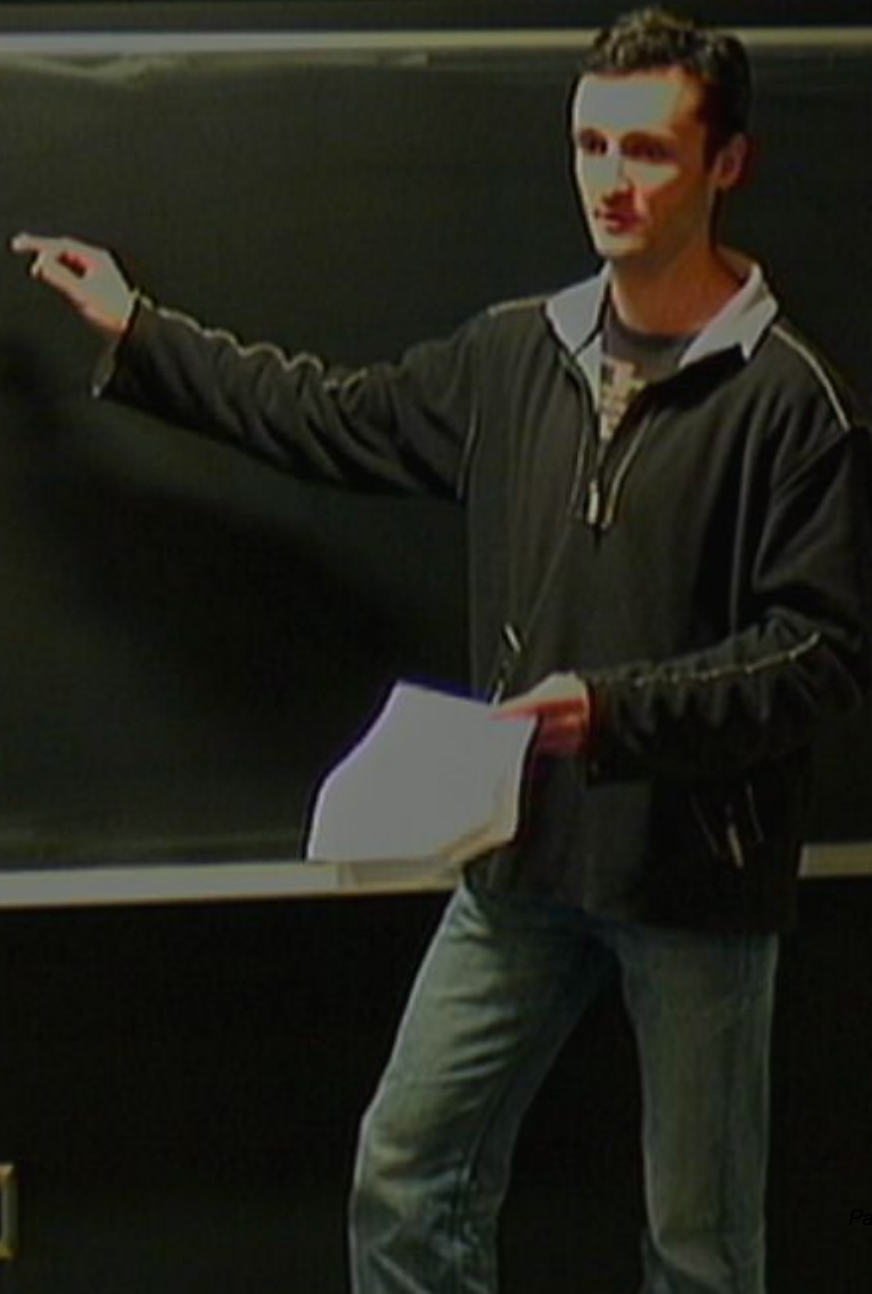
$$g_0 = g L^{-\epsilon} \Rightarrow \beta = -\frac{g^3 \lambda}{96\pi^2}$$

$$\mathcal{L}_{\text{EFT}} = g \int \frac{d^3 k}{(2\pi)^3} \phi(k)$$

$$+ \frac{\lambda g^2}{12(4\pi)^2} \int \frac{d^3 k}{(2\pi)^3} \phi(k) \ln[(kL)^2]$$

$$\phi_{\text{EFT}} = -\frac{g}{4\pi r} + \frac{g^2 \lambda}{6(4\pi)^3} \frac{\ln(r/L)}{r}$$

$$\phi_{\text{full}} = -\frac{\bar{g}}{4\pi r} + \frac{\bar{g}^2 \lambda}{6(\text{km})^3} \frac{\ln(r/r_s)}{r}$$



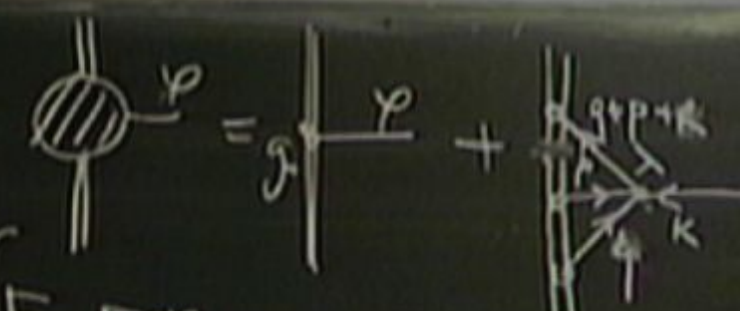
$$\phi_{\text{full}} = -\frac{\bar{g}}{4\pi r} + \frac{\bar{g}^2 \lambda}{6(4\pi)^3} \frac{\ln(r/r_s)}{r}$$

$$g(L) = \bar{g} + \frac{\bar{g}^2 \lambda}{6(4\pi)^3} \ln \frac{r_s}{L}$$

$$j(\omega) = -j\omega P \rightarrow \int \omega P d\tau$$

$$(p + \gamma p) / D(\omega)$$

$$j(\omega) d\tau + c \int R_{\mu\nu} \gamma_{\mu\nu}$$



$$c \int E_{\alpha\beta} E^{\alpha\beta} d\tau$$

$$E_{\alpha\beta} = R_{\alpha\gamma\beta\delta} \dot{x}^\gamma \dot{x}^\delta$$

$$\frac{g_0^3 \lambda}{6} \int \frac{d^3 k}{(2\pi)^3} \phi(k) \int \frac{d^3 p d^3 q}{(2\pi)^6} \frac{1}{p^2 q^2 (q+p+k)^2}$$

$$= -\frac{g_0^3 \lambda}{4\pi^2 (4\pi)^2} \left( \frac{1}{\epsilon} + 3 - \gamma + \ln(4\pi) \right)$$