

Title: Cosmology - Review (PHYS 621) - Lecture 11

Date: Dec 14, 2009 10:00 AM

URL: <http://pirsa.org/09120093>

Abstract:

Plan

- * Features of E, F eqns
Stress Energy Tensor
- * Horizon Problem
and its solution
- * Inflation

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- * Features of E, F eqns
Stress Energy Tensor
- * Horizon Problem
and its solution
- * Inflation
- * Cosmological Constt
& dark energy
- * Novel ideas on inflation
Bubbles

Plan

- * Features of E, F eqns
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Einstein Field Equations.

Einstein Field Equations

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu}$$

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$\rho \rightarrow$ energy density

$P \rightarrow$ pressure

$$T_{\mu\nu} = (\rho + P) u_{\mu} u_{\nu} - P g_{\mu\nu}$$

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$u_{\mu} \rightarrow$ 4-velocity of the fluid

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Raychaudhuri Eqn.

$$\frac{\dot{a}}{a} = -\frac{4\pi G}{3} (\rho + 3P)$$

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density

re

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velocity of the fluid

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Conservation law.

$$\rho + 3 \frac{\dot{\rho}}{\rho} (\rho + P) = 0$$

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velocity of the fluid

$$T_{\mu\nu ; \mu} = 0$$

Raychaudhuri Eqn.

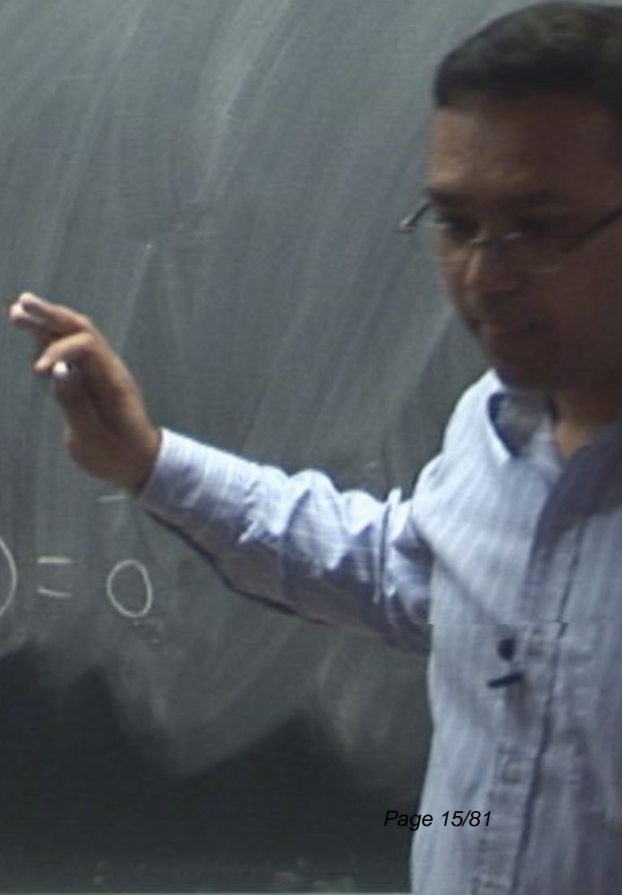
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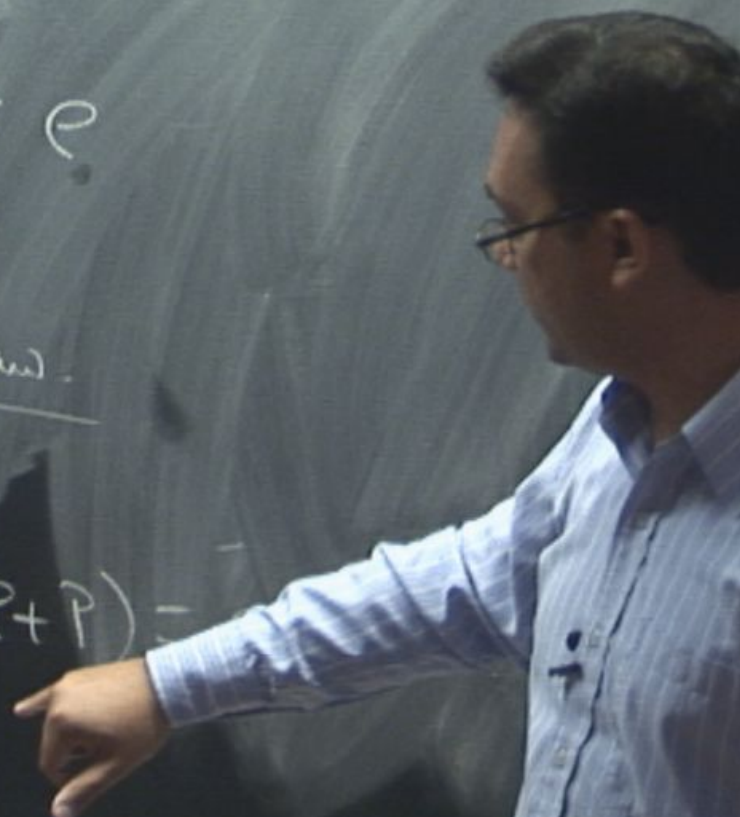
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ions.

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density

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$$dv = P g_{\mu\nu}$$

velocity of the fluid

Raychaudhuri. Eqn.

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Friedman Eqn

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Conservation law.

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$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu}$$

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$$w = \frac{P}{\rho}$$

Raychaudhuri

$$\frac{\ddot{a}}{a} =$$

Friedmann

$$\frac{\dot{a}}{a}$$

Conservation

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Conservation law

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Uni. Eqn.

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Adiabatic Expansion

$$dE = -pdV$$

Plan

* Features
Stress

Horizon
and its

Infla

Cosmo



Uni. Eqn.

$$-\frac{4\pi G}{3}(\rho + 3P)$$

man Eqn

$$\frac{z}{\lambda^2} = \frac{8\pi G}{3} \rho$$

conservation law

$$\dot{\rho} + 3\frac{\dot{a}}{a}(\rho + P) = 0$$

Adiabatic Expansion

$$dE = -PdV \quad ; V = a^3$$

E

Plan

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Adiabatic Expansion

$$dE = -pdV \quad ; V = a^3$$

$$E = \rho V$$

$$\begin{aligned} d(\rho a^3) &= -p d(a^3) \\ &= -3pa^2 da \end{aligned}$$

$$\frac{d}{da}$$

Plan

- * Features Stress

- * Horizon and its

- * Inflation

- * Cosmology & dark energy

- * Novel ideas: Bubble universes

Uni. Eqn.

$$-\frac{4\pi G}{3}(\rho + 3P)$$

man Eqn.

$$\frac{\pi G}{3} \rho$$

law.

$$(\rho + P) = 0$$

Adiabatic Expansion

$$dE = -pdV \quad ; V = a^3$$

$$E = \rho V$$

$$\begin{aligned} d(\rho a^3) &= -\rho d(a^3) \\ &= -3\rho a^2 da \end{aligned}$$

$$\frac{d}{da}(\rho a^3) = -3\rho a^2$$

$$\frac{d}{dt}(\rho a^3) = -3a^2 \dot{a} \rho$$

Plan

- * Features Stress
- * Horizon and its
- * Inflation
- * Cosmology & dark
- * Novel Bubble

Uni. Eqn.

$$-\frac{4\pi G}{3}(\rho + 3P)$$

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$$\frac{\dot{a}^2}{a^2} = \frac{8\pi G}{3}\rho$$

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Plan

* Features
Stress

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Plan

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Stress

* Horizon
and its

* Infla

* Cosmo
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ions.

$$R = 8\pi G T_{\mu\nu}$$

density

$$u_\nu = P g_{\mu\nu}$$

velocity of the fluid

$$T_{\mu\nu ; \mu} = 0$$

Dust.

$$P = 0 ; \rho > 0$$

Adiabatic

$$dE =$$

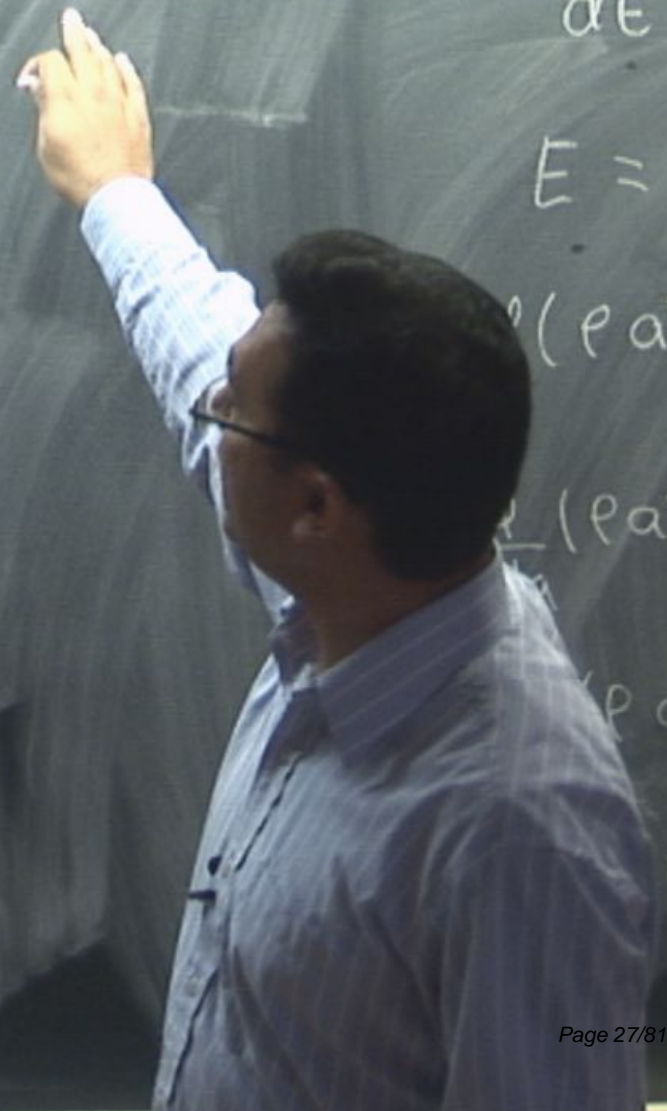
$$E = \rho V$$

$$(\rho a^3) =$$

$$(\rho a^3) =$$

$$(\rho a^3) =$$

$$= -3$$



ions.

$$\rho = 8\pi G T_{\mu\nu}$$

density

ne

$$u_\nu = P g_{\mu\nu}$$

locity of the fluid

Dust.

$$P = 0 ; \rho > 0$$

$$\therefore w = 0$$

$$T_{\mu\nu ; \mu} = 0$$

Adiabatic

$$dE =$$

$$E = \rho V$$



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$$\rho = 8\pi G T_{\mu\nu}$$

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velocity of the fluid

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Radiation

Adiabatic

$$dE =$$

$$E = \rho V$$

$$d(\rho a^3) =$$

$$\frac{d}{da} (\rho a^3) =$$

$$\frac{d}{dt} (\rho a^3) =$$

$$\dot{\rho} = -3$$

ions.

$$\rho = 8\pi G T_{\mu\nu}$$

density

$$P g_{\mu\nu}$$

ity of the fluid

$$T_{\mu\nu ; \mu} = 0$$

Dust.

$$P = 0 ; \rho > 0$$

$$\therefore w = 0$$

Radiation

$$P = \frac{1}{3} \rho$$

Adiabatic

$$dE =$$

$$E = \rho V$$

$$d(\rho a^3) =$$

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Dust.

$$p = 0 ; \rho > 0$$

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Radiation

$$p = \frac{1}{3} \rho$$

$$\therefore w = \frac{1}{3}$$

$$\rho \propto a^{-3(1+w)}$$

ρV

$$d(\rho a^3) =$$

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$T_{\mu\nu}$

fluid

$$T_{\mu\nu} = 0$$

Dust.

$$p = 0 ; \rho > 0$$

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Radiation

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$$\frac{d}{dt} a^3 =$$

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$$\therefore \omega = 0$$

Radiation

$$p = \frac{1}{3} \rho$$

$$\therefore \omega = \frac{1}{3}$$

$$T_{\mu\nu} = 0$$

$$\rho \propto a^{-3(1+\omega)}$$

$$\longrightarrow \rho \propto a^{-3}$$

$$\longrightarrow \rho \propto a^{-4}$$

Dust.

$$p = 0; \quad \rho > 0$$

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Radiation

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$$T_{\mu\nu}; \quad \mu = 0$$

$$\rho \propto a^{-3(1+w)}$$

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$$P = 0; \ell > 0$$

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$$T_{\mu\nu} = 0$$

$$P \propto a^{-3}(1+\omega)$$

$$\longrightarrow P \propto a^{-3} \propto (1+z)^3$$

$$P \propto a^{-4} \propto (1+z)^4$$

$$\text{If } a \rightarrow 0 \quad P_{\text{dust}} \rightarrow \infty$$

$$P_{\text{rad}} \rightarrow \infty$$

Dust.

$$P = 0 ; \ell > 0$$

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Radiation

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$$P_{\text{rad}} \rightarrow \infty$$

$$T_{\mu\nu} = (\partial_\mu \phi)(\partial_\nu \phi) - g_{\mu\nu} \left[\frac{1}{2} (\partial_\alpha \phi)(\partial^\alpha \phi) - V(\phi) \right]$$

$$T_{\mu\nu} = (\rho + P) u_\mu u_\nu - P g_{\mu\nu}$$

$$u_\mu = \delta^0_\mu$$

Dust.

$$P = 0 ; \rho$$

$$\therefore \omega = 0$$

Radiation

$$P = \frac{1}{3} \rho$$

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$$T_{\mu\nu} = (\rho + P) \delta^0_\mu \delta^0_\nu - P g_{\mu\nu}$$

$$\rho_\phi = \frac{1}{2} \dot{\phi}^2 + V$$

$$\dot{\phi} \neq 0$$

$$\frac{\partial \phi}{\partial x^a} = 0$$

Dust.

$$P = 0 ; \rho$$

$$\therefore \omega = 0$$

Radiation

$$P = \frac{1}{3} \rho$$

$$\therefore \omega = \frac{1}{3}$$

$$T_{\mu\nu ; \mu} = 0$$

$$T_{\mu\nu} = (\partial_\mu \phi)(\partial_\nu \phi) - g_{\mu\nu} \left[\frac{1}{2} (\partial_\gamma \phi)(\partial^\gamma \phi) - V(\phi) \right]$$

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$$\rho_\phi = \frac{1}{2} \dot{\phi}^2 + V(\phi)$$

$$P_\phi = \frac{1}{2} \dot{\phi}^2 - V(\phi)$$

$$\dot{\phi} \neq 0$$

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Dust.

$$P = 0 ; \rho >$$

$$\therefore \omega = 0$$

Radiation

$$P = \frac{1}{3} \rho$$

$$\therefore \omega = \frac{1}{3}$$

$$T_{\mu\nu ; \mu} = 0$$

$$T_{\mu\nu} = (\partial_\mu \phi)(\partial_\nu \phi) - g_{\mu\nu} \left[\frac{1}{2} (\partial_\rho \phi)(\partial^\rho \phi) - V(\phi) \right]$$

Dust

$$p = 0 ; e \rightarrow$$

-0

$$T_{\mu\nu} = (e+p) u_\mu u_\nu - p g_{\mu\nu}$$

$$u_\mu = \delta^0_\mu$$

$$\dot{\phi} \neq 0$$

$$\frac{\partial \phi}{\partial x^a} = 0$$

$$T_{\mu\nu} = (e+p) \delta^0_\mu \delta^0_\nu - p g_{\mu\nu}$$

$$p_\phi = \frac{1}{2} \dot{\phi}^2 + V(\phi)$$

$$p_\psi = \frac{1}{2} \dot{\phi}^2 - V(\phi)$$

$$T_{\mu\nu ; \mu} = 0$$

$$b) -V(\phi)$$

$$\omega\phi = \frac{1}{2}\dot{\phi}^2 - V$$

$$\frac{1}{2}\dot{\phi}^2 + V$$

$$\dot{\phi} \neq 0$$
$$\frac{\partial \mathcal{L}}{\partial x^a} = 0$$

$$T_{\mu\nu} \delta_{\mu\nu} = 0$$

$$\alpha(1+z)$$

$$(1+z)$$

→ a

→ ∞

b) $-V(\phi)$

$$\omega\phi = \frac{\frac{1}{2}\dot{\phi}^2 - V}{\frac{1}{2}\dot{\phi}^2 + V}$$

$$\dot{\phi} \neq 0$$
$$\frac{\partial \phi}{\partial x^a} = 0$$

$$\text{If } V=0, \omega = 1$$

$$T_{\mu\nu} = 0$$

$$\alpha(1+z)$$

$$(1+z)$$

(ϕ)

$$\omega\phi = \frac{\frac{1}{2}\phi^2 - V}{\frac{1}{2}\phi^2 + V}$$

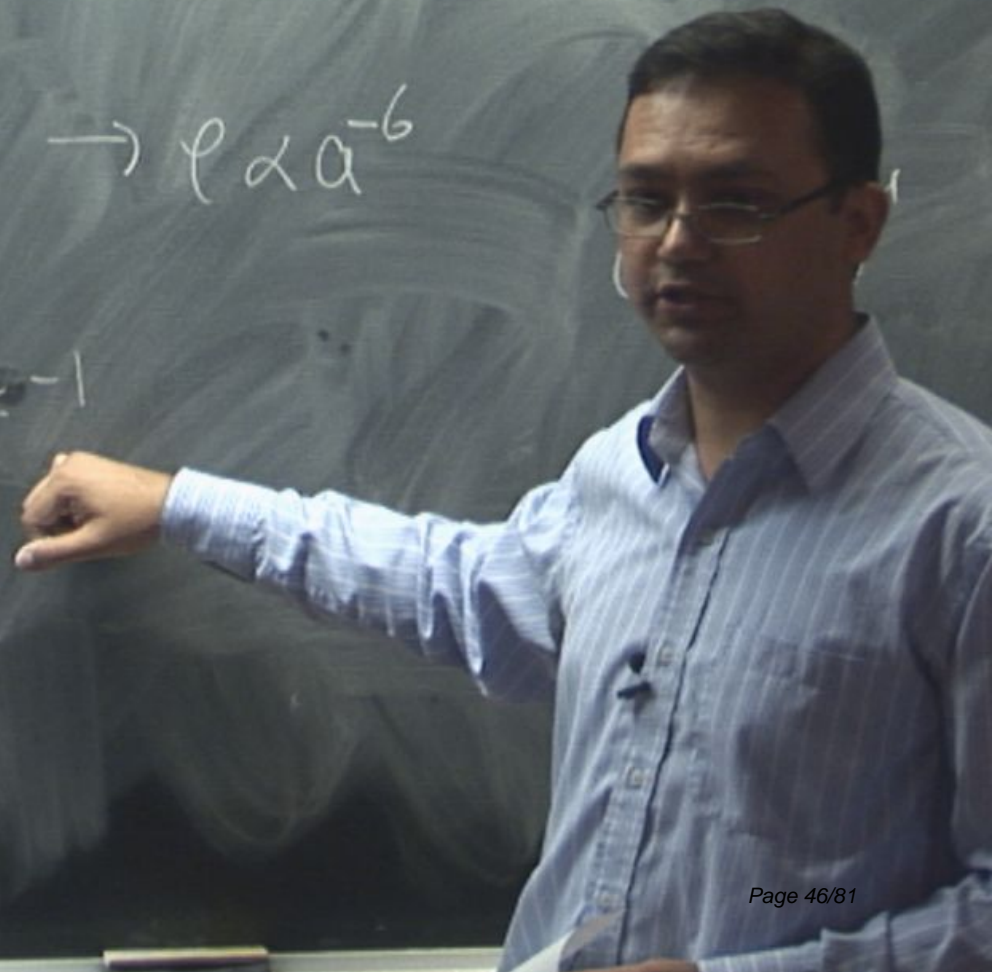
$$\rho \propto a^{-3(1+w)}$$

$$\propto (1+z)^3$$

If $V=0$, $\omega = 1 \rightarrow \rho \propto a^{-6}$

If $\phi^2 \ll V$, $\omega \approx -1$

$T_{\mu\nu} = 0$



(ϕ)

$$\omega\phi = \frac{\frac{1}{2}\phi^2 - V}{\frac{1}{2}\phi^2 + V}$$

$$\rho \propto \omega^3 (1+\omega)$$

$$\propto (1+z)^3$$

If $V=0$, $\omega = 1 \rightarrow \rho \propto a^{-6}$

$$(1+z)^4$$

If $\phi^2 \ll V$, $\omega \approx -1$

$T_{\mu\nu} = 0$

$\rightarrow \infty$

$\rightarrow \infty$

(ϕ)

$$\omega\phi = \frac{\frac{1}{2}\phi^2 - V}{\frac{1}{2}\phi^2 + V}$$

If $V=0$, $\omega = 1 \rightarrow \rho \propto a^{-6}$

If $\phi^2 \ll V$, $\omega \approx -1$

$T_{\mu\nu} = 0$

$$\rho \propto a^{-3}(1+w)$$

$$\propto (1+z)^3$$

$$\rho_{\text{dust}} \propto a^{-3}$$

$$\rho_{\text{rad}} \propto a^{-4}$$

$$(1+z)^5$$

(ϕ)

$$\omega\phi = \frac{\frac{1}{2}\phi^2 - V}{\frac{1}{2}\phi^2 + V}$$

If $V=0$, $\omega = 1 \rightarrow \rho \propto a^{-6}$

If $\phi^2 \ll V$, $\omega \approx -1$

$T_{\mu\nu} = 0$

$$\rho \propto a^{-3}(1+w)$$

$$\propto (1+z)^3$$

$$\rho_{\text{dust}} \propto a^{-3}$$

$$\rho_{\text{rad}} \propto a^{-4}$$

$$(1+z)^4$$

$\rightarrow \infty$

$\rightarrow \infty$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu}$$

$$\omega\phi = \frac{1}{2}\dot{\phi}^2$$

$$\frac{1}{2}\phi^2$$

If $V=0$

If $\phi \ll V$

$$T_{\mu\nu ; \mu} = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

$$\omega\phi = \frac{1}{2} \dot{\phi}^2$$

$$\frac{1}{2} \phi^2$$

If $V=0$,

If $\phi \ll V$

$$T_{\mu\nu ; \mu} = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

$$T_{\mu\nu}^{(0)}$$

$$\omega\phi = \frac{1}{2} \dot{\phi}^2$$

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If $V=0$

If $\phi \ll V$

$$T_{\mu\nu ; \mu} = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

$$T_{\mu\nu}^{(\Lambda)}$$

$$T_{\mu\nu}^{(\Lambda)} \equiv \frac{\Lambda}{8\pi G} g_{\mu\nu}$$

$$\omega\phi = \frac{1}{2} \dot{\phi}^2$$

$$\text{If } V=0$$

$$\text{If } \phi \ll V$$

$$T_{\mu\nu ; \mu} = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

$$\underbrace{\hspace{10em}}_{T_{\mu\nu}^{(\Lambda)}}$$

$$T_{\mu\nu}^{(\Lambda)} \equiv \frac{\Lambda}{8\pi G} g_{\mu\nu}$$

$$T_{\mu\nu} = (p + \rho) \delta_{\mu}^0 \delta_{\nu}^0 - p g_{\mu\nu}$$

$$p = -\rho \quad ; \quad \omega = -1$$

$$T_{\mu\nu} = -p g_{\mu\nu}$$

$$\omega \phi = \frac{1}{2} \dot{\phi}^2$$

$$\frac{1}{2} \phi^2$$

If $V=0$

If $\phi \ll V$

$$T_{\mu\nu ; \mu} = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

$$\underbrace{T_{\mu\nu}}^{(\Lambda)}$$

$$\boxed{T_{\mu\nu}^{(\Lambda)} \equiv \frac{\Lambda}{8\pi G} g_{\mu\nu}}$$

$$T_{\mu\nu} = (p + \rho) \delta_{\mu}^0 \delta_{\nu}^0 - p g_{\mu\nu}$$

$$p = -\rho \quad ; \quad \omega = -1$$

$$T_{\mu\nu} = -p g_{\mu\nu}$$

$$\omega \phi = \frac{1}{2} \dot{\phi}^2$$

$$\frac{1}{2} \phi^2$$

If $V=0$

If $\phi \ll V$

$$T_{\mu\nu ; \mu} = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

$$\underbrace{\hspace{10em}}_{T_{\mu\nu}^{(\Lambda)}}$$

$$T_{\mu\nu}^{(\Lambda)} \equiv \frac{\Lambda}{8\pi G} g_{\mu\nu}$$

$$T_{\mu\nu} = (p + \rho) \delta_{\mu}^0 \delta_{\nu}^0 - p g_{\mu\nu}$$

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$$\rho_{\Lambda} = \frac{\Lambda}{8\pi G}$$

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If $V=0$,

If $\dot{\phi} \ll V$

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$$\underbrace{\hspace{10em}}_{T_{\mu\nu}^{(\Lambda)}}$$

$$T_{\mu\nu}^{(\Lambda)} \equiv \frac{\Lambda}{8\pi G} g_{\mu\nu}$$

$$T_{\mu\nu} = (p + P) \delta_{\mu}^0 \delta_{\nu}^0 - P g_{\mu\nu}$$

$$P = -p \quad ; \quad \omega = -1$$

$$T_{\mu\nu} = -P g_{\mu\nu} = p g_{\mu\nu}$$

$$P_{\Lambda} = \frac{\Lambda}{8\pi G}$$

$$\omega \phi = \frac{1}{2} \dot{\phi}^2$$

$$\frac{1}{2} \phi^2$$

If $V=0$,

If $\phi \ll V$

$$T_{\mu\nu ; \mu} = 0$$

Energy Conditions.

$$v_i \mu = 0$$

Energy Conditions.

$$T_{\mu\nu} = (e + p) u_\mu u_\nu - p g_{\mu\nu}.$$

i) Strong Energy

$$u^\mu u^\nu T_{\mu\nu} = 0$$

Energy Conditions.

$$T_{\mu\nu} = (\rho + p) u_\mu u_\nu - p g_{\mu\nu}.$$

i) Strong Energy Condition:

$$u^\mu u^\nu T_{\mu\nu} \geq 0$$

Energy Conditions.

$$T_{\mu\nu} = (\rho + p) u_\mu u_\nu - p g_{\mu\nu}.$$

i) Strong Energy Condition:

\forall timelike t^μ

$$T_{\mu\nu} t^\mu t^\nu \geq \frac{1}{2} T^\lambda{}_\lambda t^\sigma t_\sigma$$

$$R_{\mu\nu} t^\mu t^\nu \geq 0$$

$$u_i = 0$$

Energy Conditions.

$$T_{\mu\nu} = (\rho + p) u_\mu u_\nu - p g_{\mu\nu}.$$

i) Strong Energy Condition:

\forall timelike t^μ

$$T_{\mu\nu} t^\mu t^\nu \geq \frac{1}{2} T^\lambda{}_\lambda t^\sigma t_\sigma$$

$$R_{\mu\nu} t^\mu t^\nu \geq 0$$

$$\Rightarrow \rho + p \geq 0, \rho + 3p \geq 0$$

Energy Conditions.

$$T_{\mu\nu} = (\rho + p) u_\mu u_\nu - p g_{\mu\nu}.$$

i) Strong Energy Condition:

\forall timelike t^μ

$$T_{\mu\nu} t^\mu t^\nu \geq \frac{1}{2} T^\lambda{}_\lambda t^\sigma t_\sigma$$

$$R_{\mu\nu} t^\mu t^\nu \geq 0$$

$$\Rightarrow \boxed{\rho + p \geq 0, \rho + 3p \geq 0}$$

$u_i = 0$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

* Dominant Energy Condition:

Energy

$$T_{\mu\nu} = (\rho + p) u_{\mu} u_{\nu} + p g_{\mu\nu}$$

i) Strong Energy

\forall timelike

$$T_{\mu\nu} t^{\mu} t^{\nu} \geq 0$$

$$\Rightarrow \rho + p \geq 0$$

$$T_{\mu\nu} i^{\mu} i^{\nu} = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

* Dominant Energy Condition:

$$\forall t^\mu; T_{\mu\nu} t^\mu t^\nu \geq 0 \quad \& \quad T_{\mu\nu} T^\nu{}_\lambda t^\mu t^\lambda \leq 0$$

$$\boxed{\rho + p \geq 0, \rho - p \geq 0}$$

- Energy

$$T_{\mu\nu} = (\rho + p) u_\mu u_\nu + p g_{\mu\nu}$$

i) Strong Energy

\forall timelike

$$T_{\mu\nu} t^\mu t^\nu \geq 0$$

$$\Rightarrow \boxed{\rho + p \geq 0}$$

$$T_{\mu\nu} i^\mu = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

* Dominant Energy Condition:

$$\forall t^\mu; T_{\mu\nu} t^\mu t^\nu \geq 0 \text{ \& } T_{\mu\nu} T^\mu{}_\lambda t^\lambda t^\nu \leq 0$$

$$\boxed{\rho + p \geq 0, \rho - p \geq 0}$$

* Weak energy condition

$$T_{\mu\nu} t^\mu t^\nu \geq 0 \Rightarrow \boxed{\rho \geq 0, \rho + p \geq 0}$$

* Null Energy condition

$$T_{\mu\nu} \ell^\mu \ell^\nu \geq 0 \quad \forall \ell_\mu \rightarrow \boxed{\rho + p \geq 0}$$

Energy

$$T_{\mu\nu} = (\rho, p)$$

i) Strong

\forall timelike

$$T_{\mu\nu} t^\mu t^\nu$$

$$\Rightarrow \boxed{\rho + p}$$

$$T_{\mu\nu} \ell^\mu \ell^\nu = 0$$

$$\boxed{\rho + p \geq 0}$$

Conditions.

$$+P) u_{\mu} u_{\nu} - P g_{\mu\nu}.$$

Energy Condition:

for t^{μ}

$$\geq \frac{1}{2} T^{\lambda}_{\lambda} t^{\sigma} t_{\sigma} \quad | \quad R_{\mu\nu} t^{\mu} t^{\nu} > 0$$

$$P \geq 0, \quad P + 3P \geq 0$$

Conditions.

$$+P) u_{\mu} u_{\nu} - P g_{\mu\nu}.$$

Energy Condition:

for t^{μ}

$$\geq \frac{1}{2} T^{\lambda}_{\lambda} t^{\sigma} t_{\sigma} \quad | \quad R_{\mu\nu} t^{\mu} t^{\nu} > 0$$

$$P \geq 0, \quad P + 3P \geq 0$$

Dust:

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} (\rho + 3P)$$

Conditions.

$$+P) u_{\mu} u_{\nu} - P g_{\mu\nu}.$$

Energy Condition:

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If SEC is satisfied then $\ddot{a} < 0$

Conditions.

$$+P) u_{\mu} u_{\nu} - P g_{\mu\nu}.$$

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$$P \geq 0, \quad P + 3P \geq 0$$

Dust:

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} (\rho + 3P)$$

If SEC is satisfied then $\ddot{a} < 0$

$$\omega = 0, \quad P = 0, \quad \rho > 0$$
$$\rho + 3P > 0$$

Conditions.

$$+P) u_{\mu} u_{\nu} - P g_{\mu\nu}.$$

Energy Condition:

for t^{μ}

$$\geq \frac{1}{2} T^{\mu\nu} t^{\sigma} t_{\sigma} \quad | \quad R_{\mu\nu} t^{\mu} t^{\nu} > 0$$

$$P \geq 0, \quad P + 3P \geq 0$$

Dust:

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} (\rho + 3P)$$

If SEC is satisfied then $\ddot{a} < 0$

$$w = 0, \quad P = 0, \quad \rho > 0$$

$$\rho + 3P > 0$$

$$w = \frac{1}{3}, \quad \rho + 3P = \rho + \rho = 2\rho > 0$$

$$w = -1; \quad \rho + 3P = \rho - 3\rho = -2\rho < 0$$

$$\frac{\ddot{a}}{a} > 0$$

Conditions.

$$+P) u_{\mu} u_{\nu} - P g_{\mu\nu}.$$

Energy Condition:

for t^{μ}

$$\geq \frac{1}{2} T^{\mu\nu} t^{\sigma} t_{\sigma}$$

$$P \geq 0, P + 3P \geq 0$$

Dust:

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} (\rho + 3P)$$

If SEC is satisfied then $\ddot{a} < 0$

$$w = 0, P = 0, \rho > 0$$

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$$w = \frac{1}{3}, \rho + 3P = \rho + \rho = 2\rho > 0$$

$$w = -1, \rho + 3P = \rho - 3\rho = -2\rho < 0$$

$$\frac{\ddot{a}}{a} > 0$$

Conditions. $\rho \propto a^{-3(1+w)}$

$$+P) u_\mu u_\nu - P g_{\mu\nu}.$$

Energy Condition:

for t^μ

$$\geq \frac{1}{2} T^{\mu\nu} t^\sigma t_\sigma \rho_{\mu\nu}$$

$$\rho \geq 0, \rho + 3P \geq 0$$

Dust:

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} (\rho + 3P)$$

If SEC is satisfied then $\ddot{a} < 0$

$$w = 0, \rho = 0, \rho > 0$$

$$\rho + 3P > 0$$

$$w = \frac{1}{3}, \rho + 3P = \rho + \rho = 2\rho > 0$$

$$w = -1; \rho + 3P = \rho - 3\rho = -2\rho < 0$$

$$\frac{\ddot{a}}{a} > 0$$

Energy Condition

like t^μ

$$t^\nu \geq \frac{1}{2} T^\lambda_{\lambda\sigma} t^\sigma \quad | \quad R_{\mu\nu} t^\mu t^\nu > 0$$

$$+P \geq 0, \quad \rho + 3P \geq 0$$

$$\rho = \frac{E}{V}, \quad p = \frac{h\nu}{V}$$

SEC

$$\omega = 0$$

$$\rho + 3P$$

$$\omega = \frac{1}{3}$$

$$\omega = -1$$

$$t^\mu t^\mu \leq 0$$

i) Strong Energy Condition:

\forall timelike t^μ

$$T_{\mu\nu} t^\mu t^\nu \geq \frac{1}{2} T^\lambda{}_\lambda t^\sigma t_\sigma \quad | \quad R_{\mu\nu} t^\mu t^\nu$$

$$\Rightarrow \rho + p \geq 0, \quad \rho + 3p \geq 0$$

$$p \geq 0$$

$$T_{\mu\nu} i_\mu = 0$$

$$\forall l_\mu \rightarrow \rho + p \geq 0$$

$$\dot{\rho} + 3H(\rho + p) = 0 \quad \rho = E$$

Strong Energy Condition:

\forall timelike t^μ

$$T_{\mu\nu} t^\mu t^\nu \geq \frac{1}{2} T^\lambda{}_\lambda t^\sigma t_\sigma \quad | \quad R_{\mu\nu} t^\mu t^\nu > 0$$

$$\Rightarrow \rho + P \geq 0, \quad \rho + 3P \geq 0$$

$$\frac{d\rho}{dt} = -2\frac{\dot{a}}{a}(1+w)\rho$$

$$\dot{\rho} + 3H(\rho + P) = 0 \quad \left| \quad \rho = \frac{E}{V}, \quad p = \frac{h\nu}{V} \right.$$

$$\rho + P \geq 0$$

Strong Energy Condition:

timelike t^μ

$$T_{\mu\nu} t^\mu t^\nu \geq \frac{1}{2} T^\lambda{}_\lambda t^\sigma t_\sigma \quad | \quad R_{\mu\nu} t^\mu t^\nu \geq 0$$

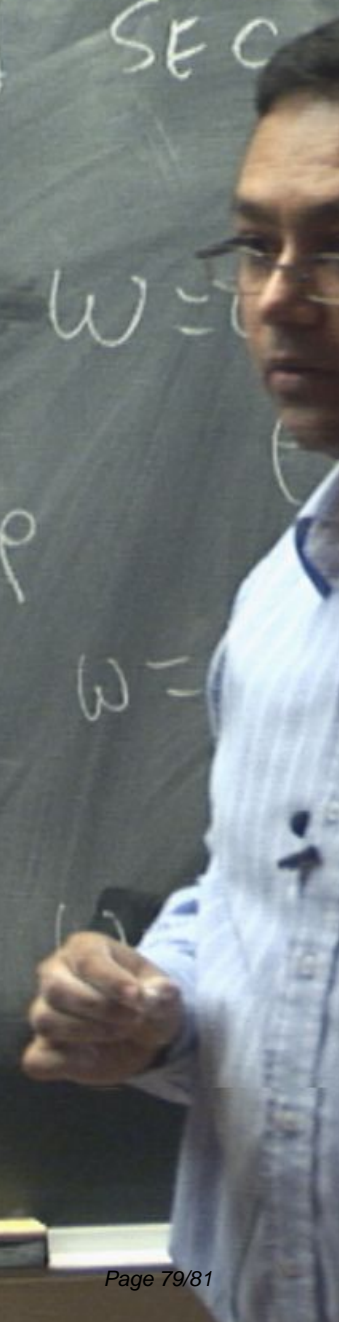
$$\frac{d\rho}{dt} = 3 \frac{da}{a} (1+w) \rho \quad \text{SEC}$$

$$\rho + P \geq 0, \quad \rho + 3P \geq 0$$

$$\frac{d\rho}{dt} = 2 \frac{da}{a} (1+w) \rho \quad \text{w} =$$

$$\dot{\rho} + 3H(\rho + P) = 0 \quad \rho = \frac{E}{V}, \quad \rho = \frac{h\nu}{V}$$

$$\rho + P \geq 0$$



Energy

$$T_{\mu\nu} = (\rho + P) u_\mu u_\nu - P g_{\mu\nu} \quad \rho \propto \dot{a}^2$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3}(\rho + P)$$

i) Strong Energy Condition:

$$\frac{d\rho}{\rho} = 3 \frac{da}{a} \quad (\text{if } w=0)$$

SEC is satisfied

\forall timelike t^μ

$$T_{\mu\nu} t^\mu t^\nu \geq \frac{1}{2} T^\lambda{}_\lambda, \quad t^\sigma t_\sigma > 0 \quad | \quad R_{\mu\nu} t^\mu t^\nu > 0$$

$$\Rightarrow \boxed{\rho + P \geq 0, \quad \rho + 3P \geq 0}$$

$$\frac{d\rho}{dt} = 2 \frac{d\rho}{da}$$

$u=0$

$$\rightarrow \boxed{\rho + P \geq 0}$$

$$\dot{\rho} + 3H(\rho + P) = 0 \quad \left| \quad \rho = \frac{E}{V}, \quad P = \frac{h\nu}{V}$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} + \Lambda g_{\mu\nu}$$

* Dominant Energy Condition:

$$\forall t^\mu; T_{\mu\nu} t^\mu t^\nu \geq 0 \text{ \& } T_{\mu\nu} T^\mu{}_\lambda t^\lambda t^\nu \leq 0$$

$$\boxed{\rho + p \geq 0, \rho - p \geq 0}$$

* Weak energy condition

$$T_{\mu\nu} t^\mu t^\nu \geq 0 \Rightarrow \boxed{\rho \geq 0, \rho + p \geq 0}$$

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$$T_{\mu\nu} \ell^\mu \ell^\nu \geq 0 \quad \forall \ell_\mu \rightarrow$$

- Energy Con

$$T_{\mu\nu} = (\rho + p)$$

i) Strong Energy

\forall timelike

$$T_{\mu\nu} t^\mu t^\nu$$

$$\Rightarrow \boxed{\rho + p}$$

$$T_{\mu\nu} i_\mu = 0$$