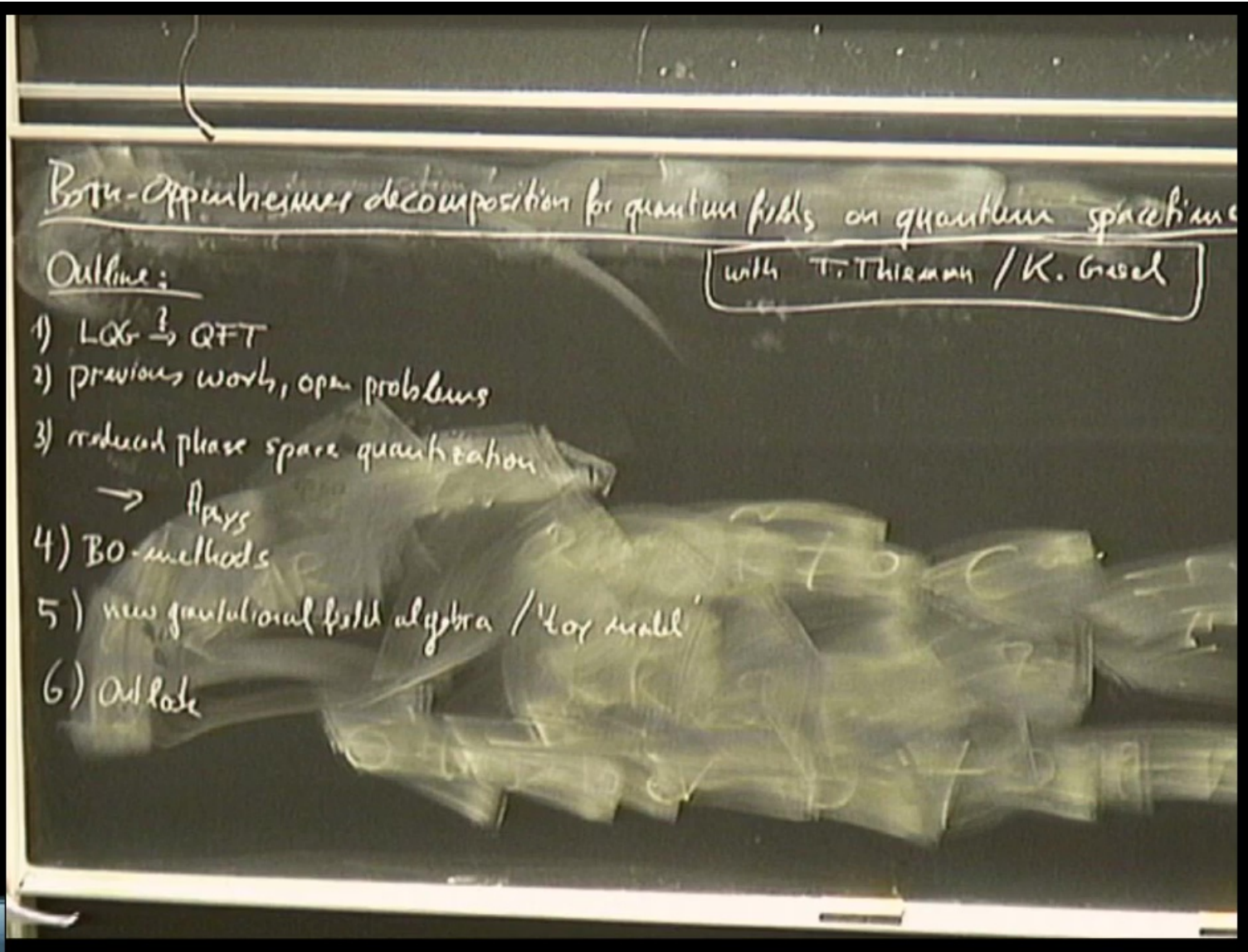
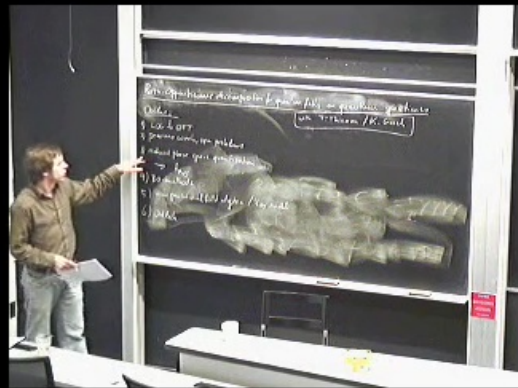


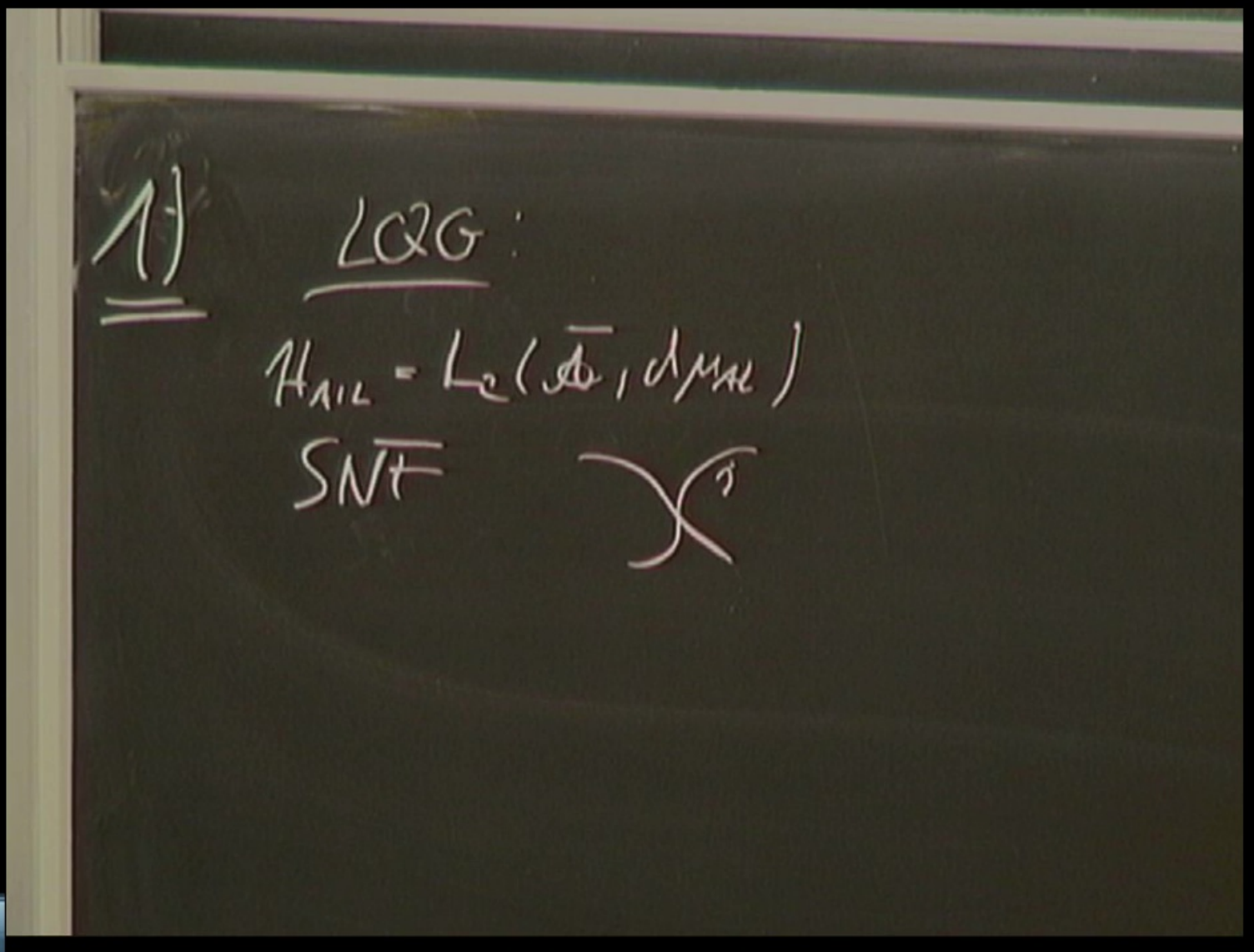
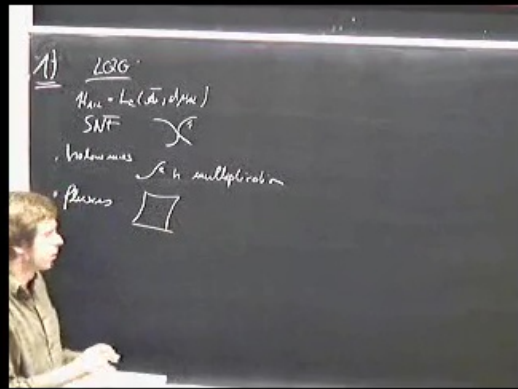
Title: Born--Oppenheimer approximation for quantum fields on quantum spacetimes

Date: Nov 25, 2009 04:00 PM

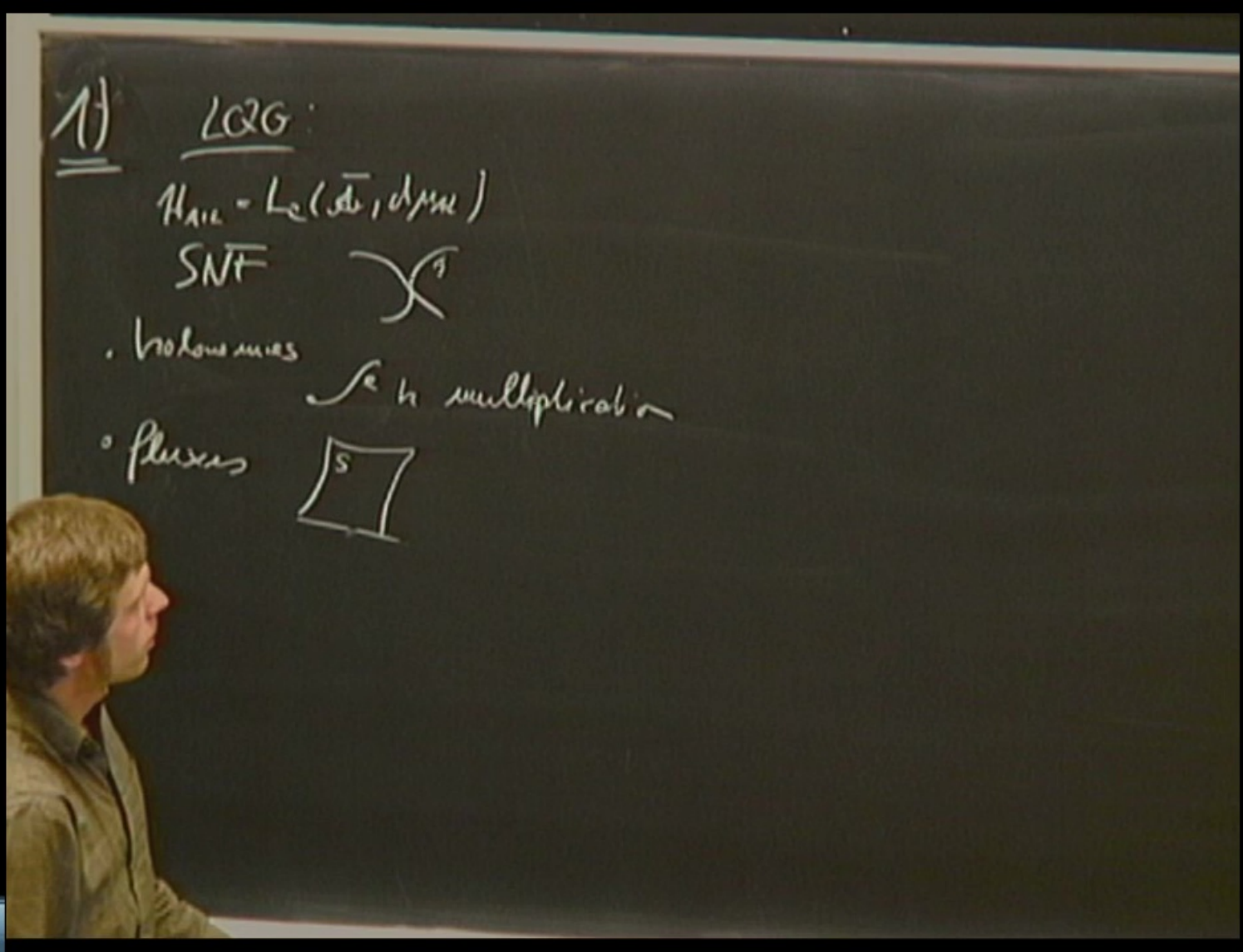
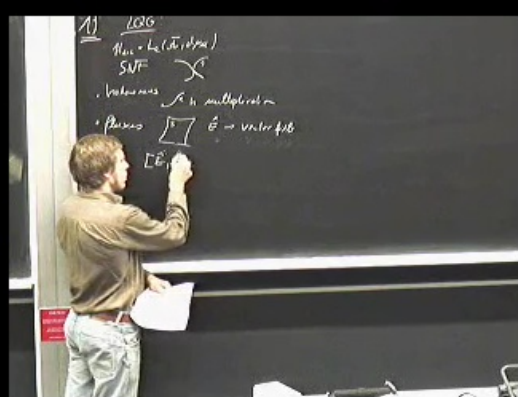
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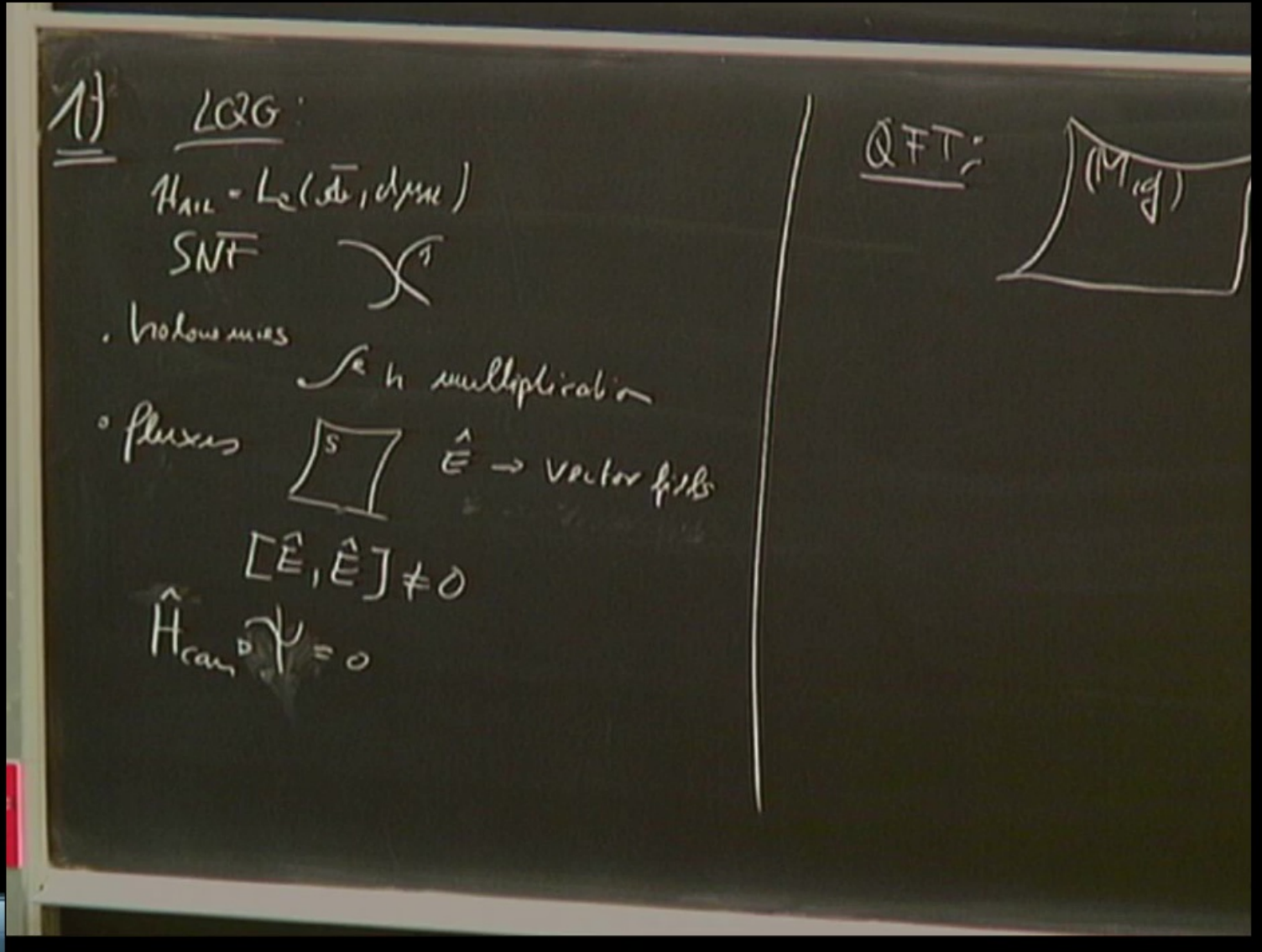
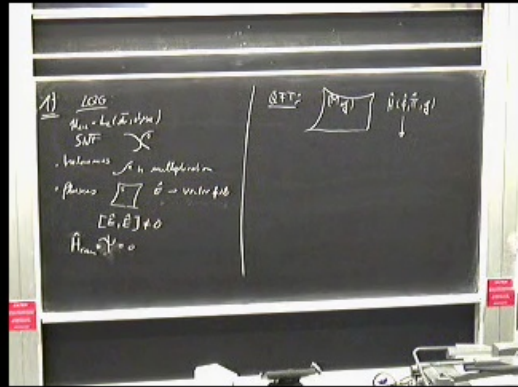
Abstract: The relation between loop quantum gravity (LQG) and ordinary quantum field theory (QFT) on a fixed background spacetime still bears many obstacles. When looking at LQG and ordinary QFT from a mathematical perspective it turns out that the two frameworks are rather different: Although LQG is a true continuum theory its Hilbert space is defined in terms of certain embedded graphs which are labeled by irreducible representations of  $SU(2)$ . The natural arena for ordinary QFT, on the other hand, is a Fock space which strongly uses the metric properties of the underlying continuum spacetime. In this talk I will review this issue and show how one can use Born--Oppenheimer methods to further progress towards an understanding of (matter) quantum field theories from first principles.

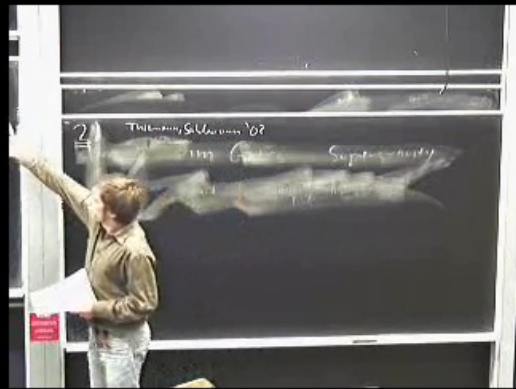












1) LQG  
 $\mathcal{H}_{\text{kin}} = L_2(\mathcal{S}^1, \mathfrak{su}(2))$   
 SNF  $\chi^1$

- holonomies  $\int_{\gamma} A$  multiplication
- fluxes  $\int_{\Sigma} \hat{E} \rightarrow$  vector fib

$[\hat{E}, \hat{E}] \neq 0$   
 $\hat{H}_{\text{can}} \psi = 0$

QFT:  $(M, g)$

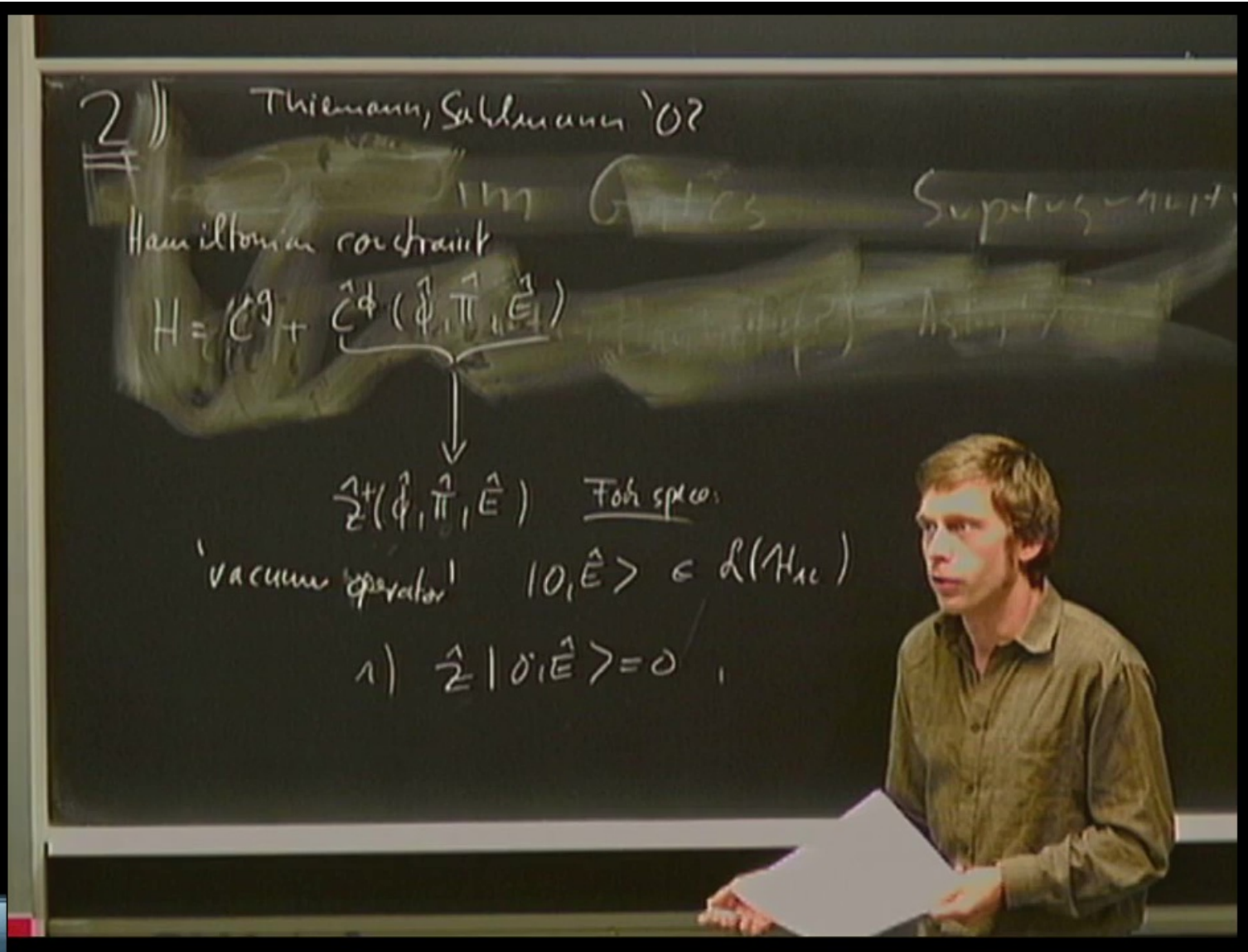
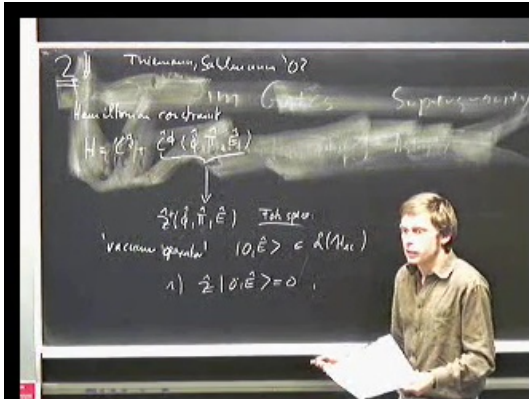
$\hat{H}(\hat{\pi}, \hat{\pi}, g) \downarrow \hat{H}_g$   
 $\hat{Z}^+(i\hbar, g), \hat{Z}$   
 $|\mathcal{O}\rangle_g$

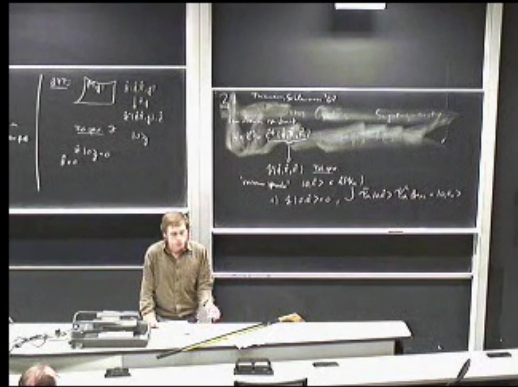
Fock space:  $\mathcal{F}$

$\hat{H} | \mathcal{O} \rangle_g = 0$   
 $\hat{H} | + 0$









2) Thiemann, Seibener '07

im Gauge Supergravity

Hamiltonian constraint

$$H = \mathcal{L}_g + \mathcal{L}_d(\hat{q}, \hat{\pi}, \hat{E})$$

↓

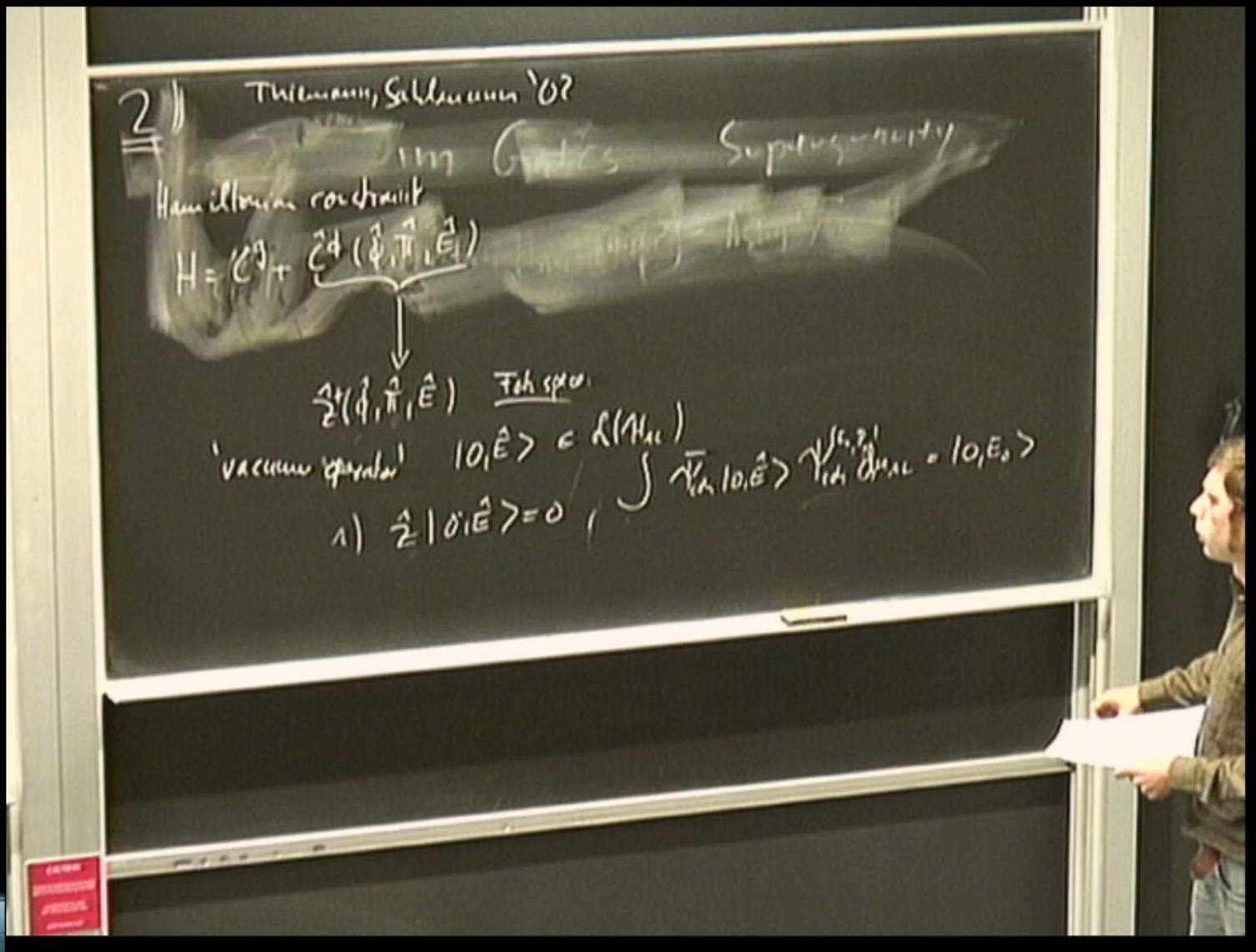
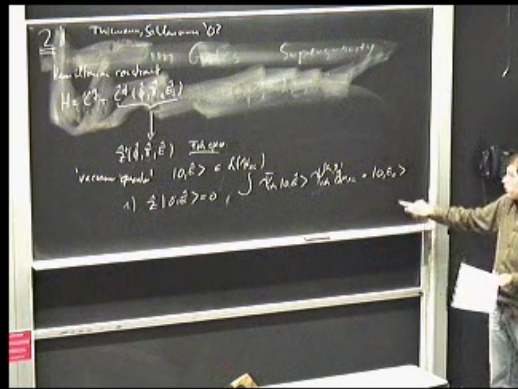
$\hat{H}(\hat{q}, \hat{\pi}, \hat{E})$  Fock space

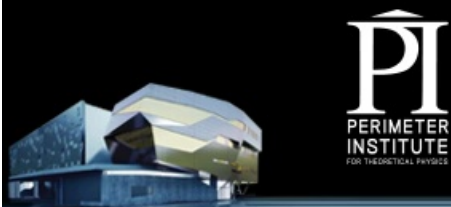
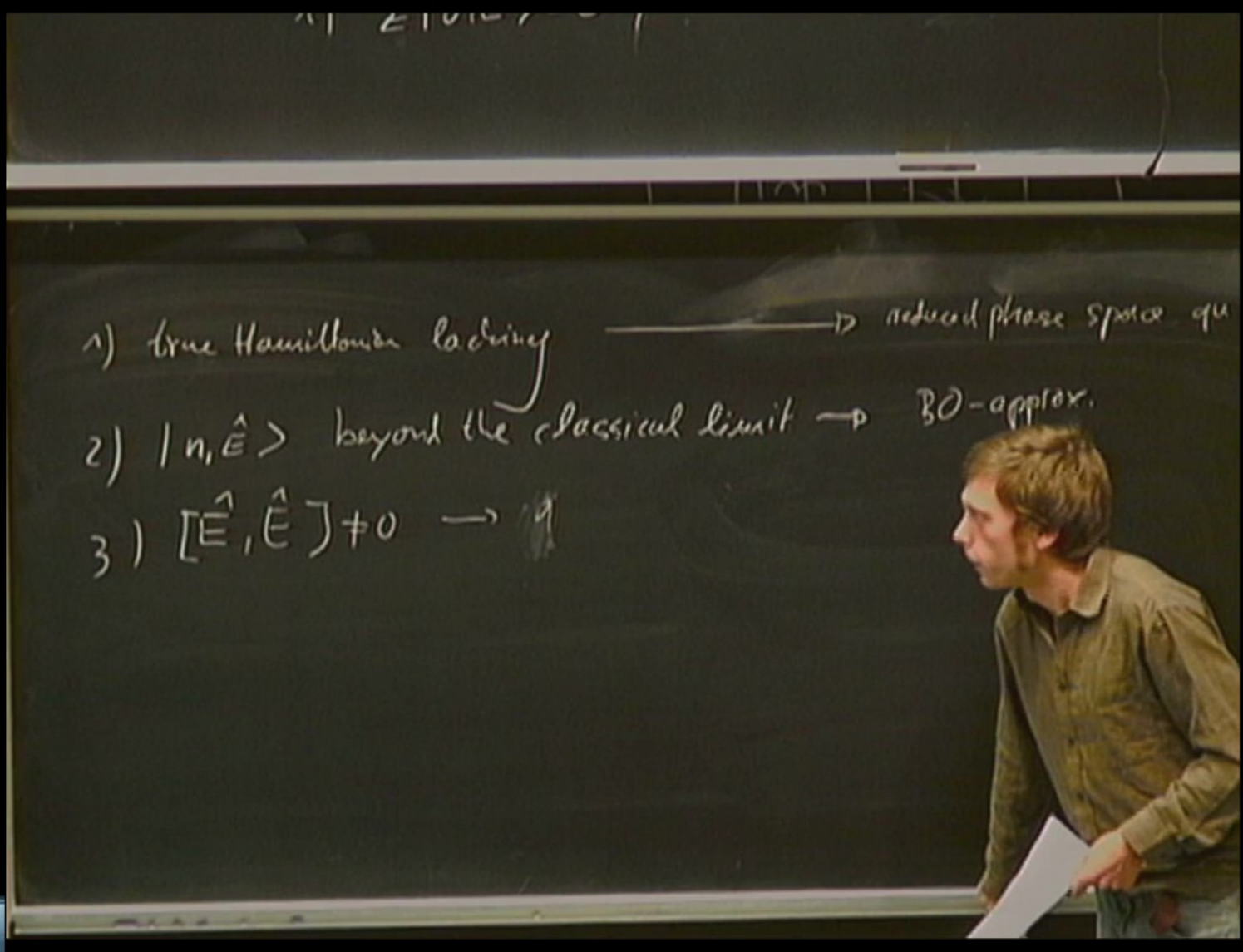
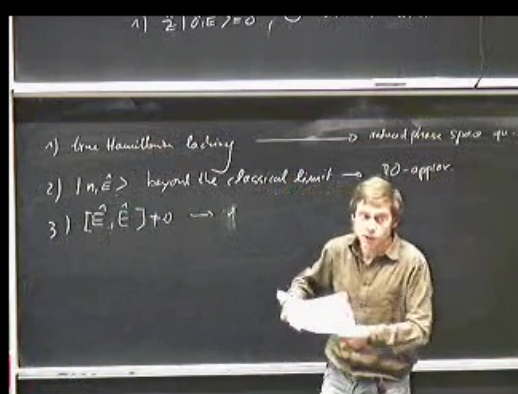
'vacuum operator'  $|0, \hat{E}\rangle \in \mathcal{L}(\mathcal{H}_{\mathcal{L}})$

1)  $\hat{H}|0, \hat{E}\rangle = 0$ ,  $\int \bar{\Psi}_m |0, \hat{E}\rangle \Psi_{ch}^m d\mu_{\mathcal{L}} = |0, E_0\rangle$













So  $H \neq 0$

Dirac Program

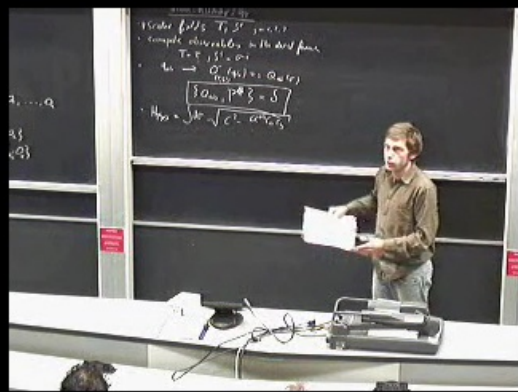
- i) quantize all dof's
- ii)  $H_{\text{red}} \Psi = 0$
- iii) dynamics?

RPA

- i) solve constraints  $\rightarrow$  Dirac Observables  $a_1, \dots, a_n$
- $\{a_i, a_j\} = ?$
- ii) representation of  $\{a_i, a_j\}$
- iii) dynamics  $\frac{d}{dt} a_i = \{H_{\text{red}}, a_i\}$







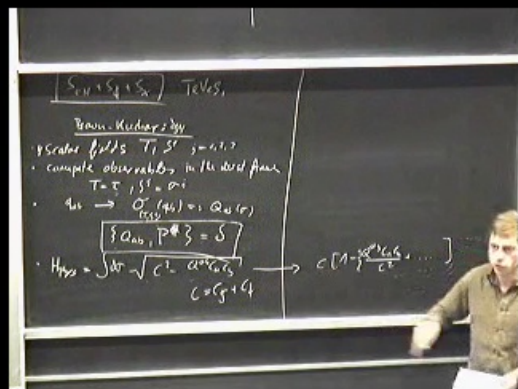
Brown-Kuchar: 231

- 4 scalar fields  $T, S^i, j=1,2,3$
- compute observables in the dust frame  
 $T = \tau, S^i = \sigma^i$
- $q_{ab} \rightarrow \tilde{\sigma}_{(T, S^i)}(q_b) =: Q_{ab}(\sigma)$

$\{Q_{ab}, P^{\mu}\} = \delta$

- $H_{\text{phys}} = \int d\sigma \sqrt{c^2 - \alpha^{ab} c_a c_b}$





$S_{EH} + S_f + S_x$  Teves,

Brown-Kuchar: 234

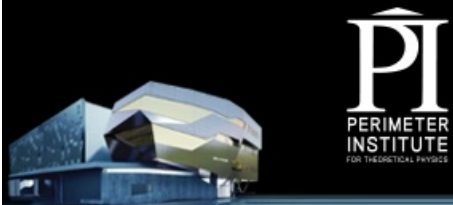
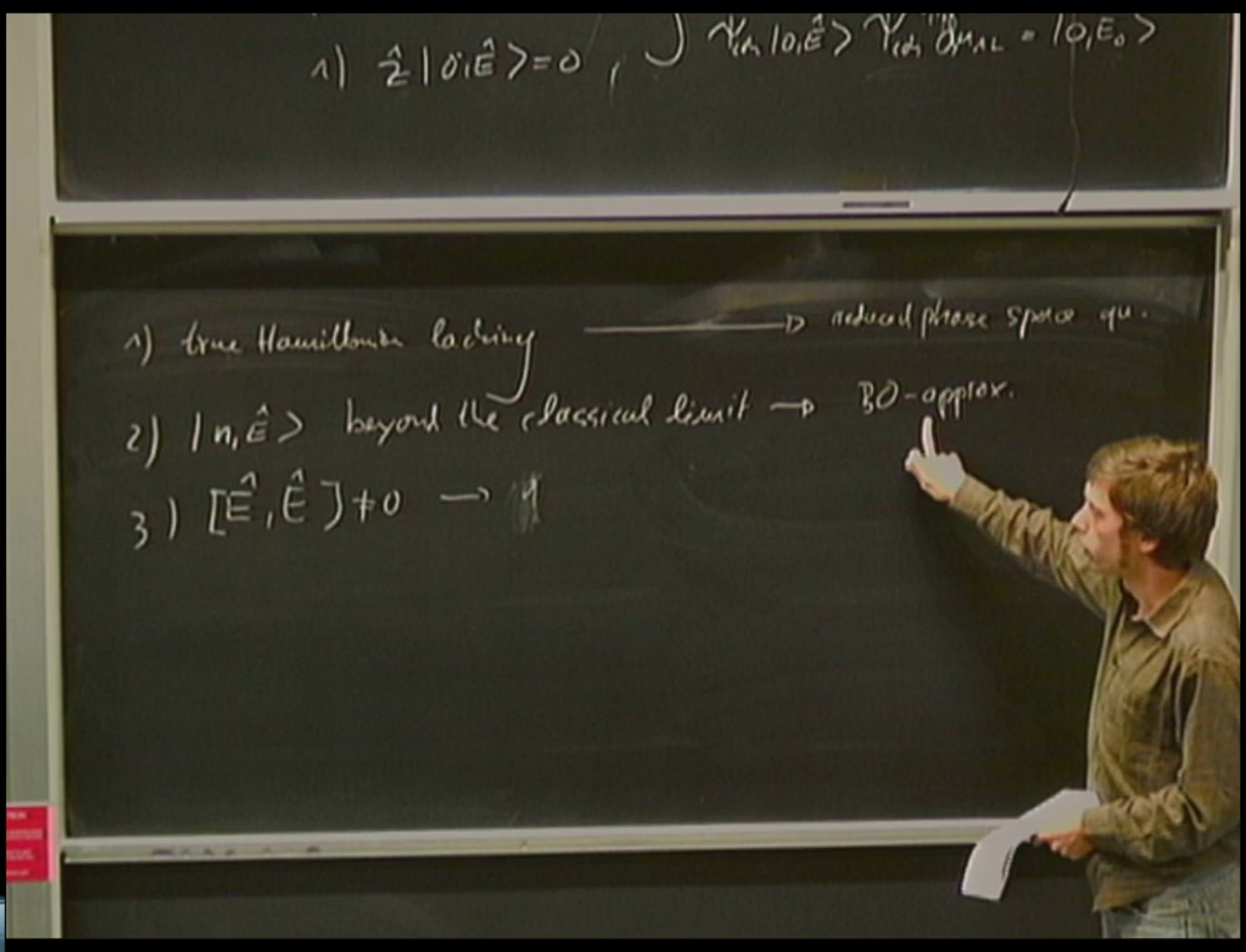
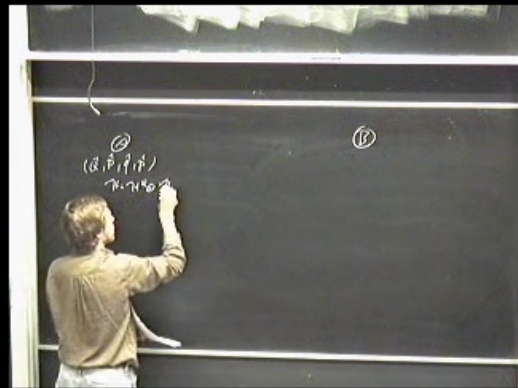
- 4 scalar fields  $T_i, S^i, j=1,2,3$
- compute observables in the dust frame  
 $T = \tau, S^i = \sigma^i$
- $q_{ab} \rightarrow \sigma_{(T, S^i)}^{(ab)} =: Q_{ab}(\sigma)$

$\{Q_{ab}, P^{\mu}\} = \mathcal{S}$

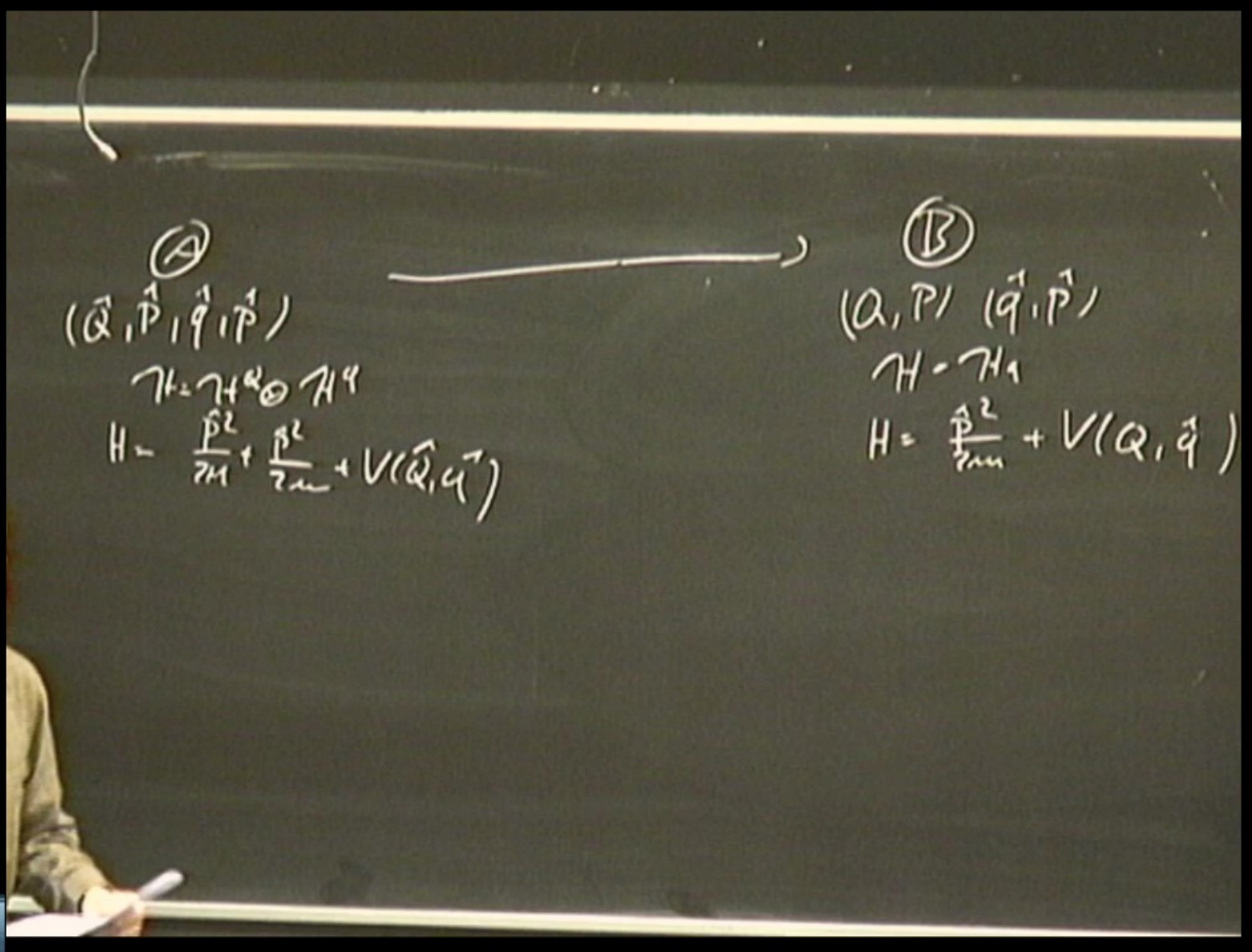
- $H_{phys} = \int dV \sqrt{c^2 - Q^{ab}c_{ab}}$   $\rightarrow c \left[ 1 - \frac{1}{2} \frac{Q^{ab}c_{ab}}{c^2} + \dots \right]$   
 $c = c_g + c_f$













Ⓐ

$$(\vec{Q}, \vec{P}, \vec{q}, \vec{p})$$

$$\mathcal{H} = \mathcal{H}^Q \oplus \mathcal{H}^q$$

$$H = \frac{\vec{P}^2}{2M} + \frac{\vec{p}^2}{2m} + V(\vec{Q}, \vec{q})$$

Ⓑ

$$(Q, P) (q, p)$$

$$\mathcal{H} = \mathcal{H}_1$$

$$\hat{H}_q = \frac{\vec{p}^2}{2m} + V(Q, \vec{q})$$

---

$$[\hat{H}_q - \lambda(Q)] \chi_q(q) = 0$$





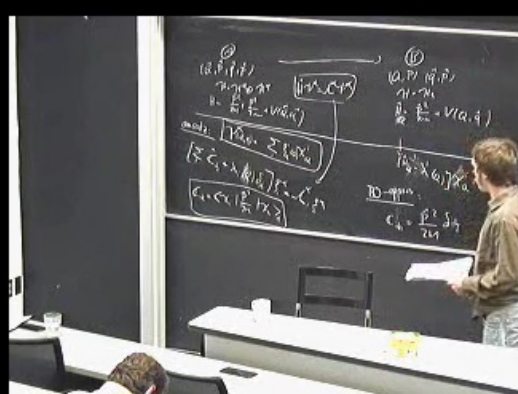
(A)  $(Q, P, \dot{Q}, \dot{P})$   
 $\mathcal{H} = \mathcal{H}^0 + \mathcal{H}^1$   
 $\mathcal{H} = \frac{p^2}{2m} + V(Q, q)$   
 ansatz:  $\Psi(Q, q) = \sum_k c_k \chi_k$   
 $[\sum_k \hat{c}_k + \lambda_i \delta_{ik}] \delta_{ik} = \sum_k \hat{c}_k \delta_{ik}$   
 $c_k = \langle \chi_k | \frac{p^2}{2m} | \chi_k \rangle$

(A)  $(Q, P, \dot{Q}, \dot{P})$   
 $\mathcal{H} = \mathcal{H}^0 + \mathcal{H}^1$   
 $\mathcal{H} = \frac{p^2}{2m} + V(Q, q)$   
 ansatz:  $\Psi(Q, q) = \sum_k c_k \chi_k$   
 $[\sum_k \hat{c}_k + \lambda_i \delta_{ik}] \delta_{ik} = \sum_k \hat{c}_k \delta_{ik}$   
 $c_k = \langle \chi_k | \frac{p^2}{2m} | \chi_k \rangle$

(B)  $(Q, P, \dot{Q}, \dot{P})$   
 $\mathcal{H} = \mathcal{H}_1$   
 $\hat{H}_Q = \frac{p^2}{2m} + V(Q, q)$   
 $[\hat{H}_Q - \lambda(Q)] \chi_k^i(Q) = 0$







Ⓐ

$(Q, P, \dot{q}, \dot{P})$   
 $\mathcal{H} = \mathcal{H}^0 + \mathcal{H}^1$   
 $\mathcal{H} = \frac{\vec{p}^2}{2m} + \frac{\vec{p}^2}{2M} + V(Q, q)$

ansatz:  $\psi(Q, q) = \sum_k \psi_k(Q) \chi_k(q)$

$[\sum_k \hat{c}_k + \lambda_i \delta_{ik}] \psi_k = \epsilon_k \psi_k$

$c_{ik} = \langle \chi_i | \frac{\vec{p}^2}{2M} | \chi_k \rangle$

Ⓑ

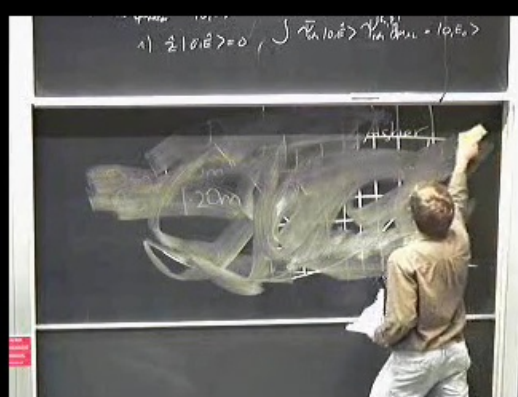
$(Q, P) (\dot{q}, \dot{P})$   
 $\mathcal{H} = \mathcal{H}_1$   
 $\hat{H}_Q = \frac{\vec{p}^2}{2m} + V(Q, q)$

$[\hat{H}_Q - \epsilon_k(Q)] \chi_k(q) = 0$

BD-approx.

$c_{ih} = \frac{\vec{p}^2}{2M}$





$C = C_g + C_d$

$$\left[ \frac{\hat{p}^2}{2M} + \lambda_i(\hat{Q}) \right] \xi_n^i(\alpha) = \Lambda_n \xi_n^i(\alpha)$$

↑  
effectively taking into account  
the presence of  $\psi$

- i) 2 separated energy scales
- ii)  $\hat{Q}$  multiplication operator





