

Title: Workshop Summary and Discussion

Date: Nov 10, 2009 03:00 PM

URL: <http://pirsa.org/09110066>

Abstract:

projectable  
vs

non projectable

detailed balance  
or  
not?

expanding around  
flat or curved?

Strong coupling or weak?



{ projectable  
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flat or curved?

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propagating/not?



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detailed balance  
or  
not?

Add  $\left(\frac{\partial M^2}{\partial N}\right)$  or not?

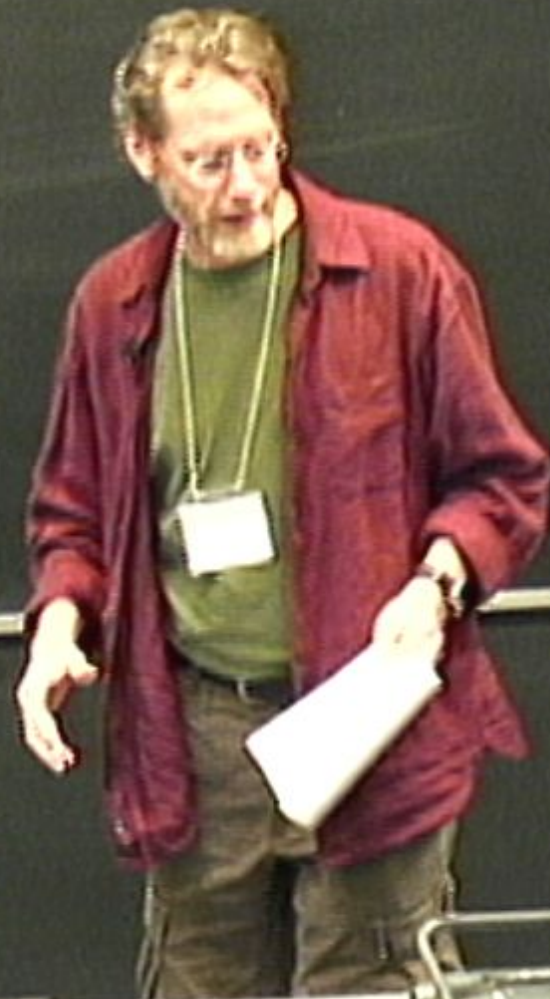
expanding around  
flat or curved?

Strong coupling or weak?

propagating/not?

Cosmology - primordial fluct. w/o inflation

Lifshitz  
Robert B: "matter bounce"





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Robert B: "matter bounce" - scale inv. by amplif.  
of longer wavelengths  
"by accident" scale inv.  
if matter dominated

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↳ scale inv. from  $\Phi \rightarrow b^{-\frac{(2-\gamma)}{2}} \Phi$

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Stability at bkgnds

→ projectable at flat space unstable scalar mode

## Stability at bkgnds

→ projectable at flat space unstable scalar mode.

→ add  $\left(\frac{\partial_i N}{N}\right)^2$  term. (Dregu).



$$N = N(t) \rightarrow u_n = \frac{\nabla_n t}{\sqrt{\partial_n \partial_n t}} \quad \text{unit.}$$



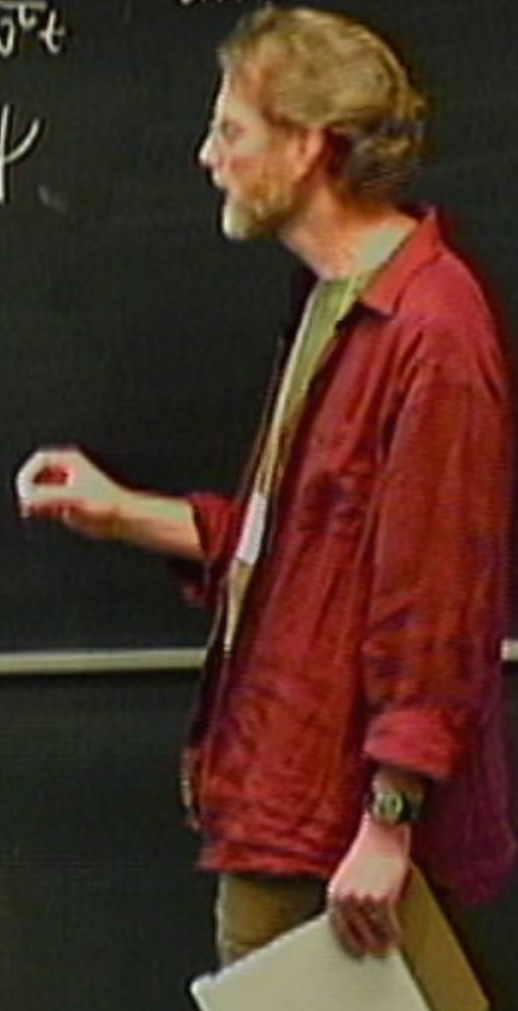
$$N = N(t) \rightarrow U_n = \frac{\nabla_n t}{\sqrt{g_{ij} \dot{x}^i \dot{x}^j}}$$

unit.

geodesics

$$= \nabla_n \Psi$$

$\uparrow U_n$   $t = \text{const.}$



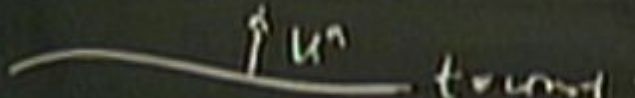
CAUTION  
UNIVERSITY OF  
TORONTO



$$N = N(t) \rightarrow U_a = \frac{\nabla_a t}{\sqrt{g_{ab} \dot{x}^b \dot{x}^a}} \quad \text{unit.}$$

geodesics

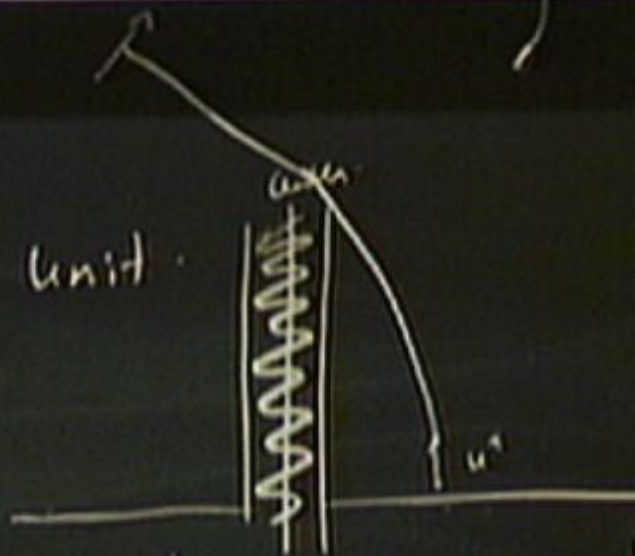
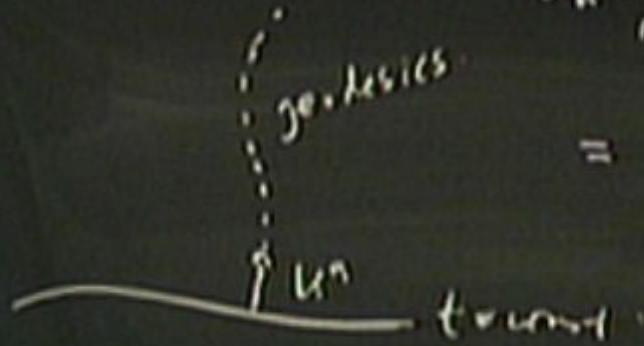
$$= \nabla_a \Psi$$



$$U^b \nabla_b U_a = U^b \nabla_a U_b \\ = \frac{1}{2} \nabla_a (U^b U_b)$$

$$N = N(t) \rightarrow U_a = \frac{\nabla_a t}{\sqrt{g_{ab} \dot{x}^b \dot{x}^a}}$$

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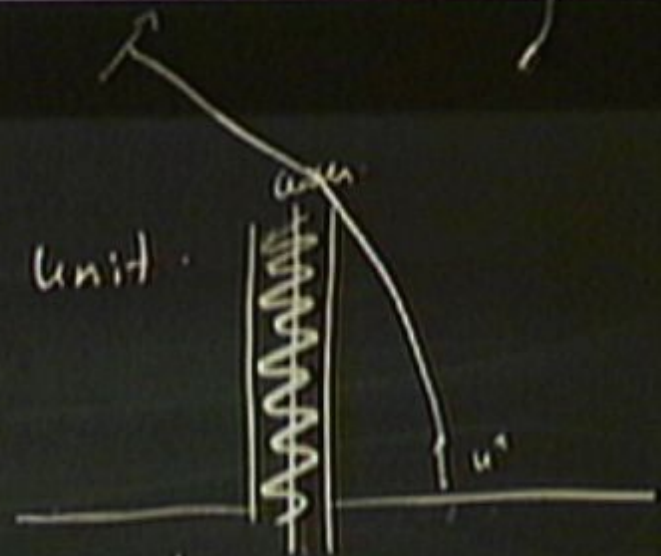
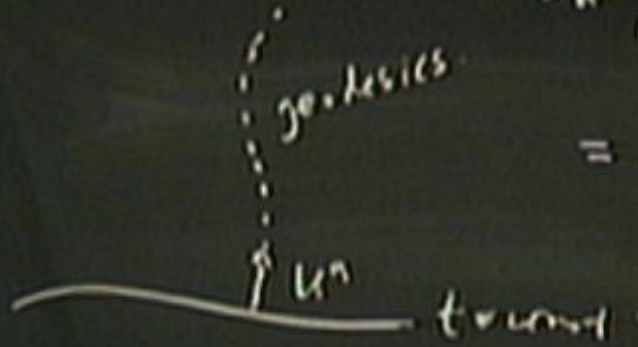
$$\rightarrow \frac{1}{2} \nabla_a (U^b U_b) = 0$$





$$N = N(t) \rightarrow U_a = \frac{\nabla_a t}{\sqrt{g_{ab} \dot{x}^a \dot{x}^b}}$$

$$= \nabla_a \Psi$$



$$U^b \nabla_b U_a = U^b \nabla_a U_b$$

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Niaysh:  $(1-\lambda)K^2$  term





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claim:  $K$  fixed by b.c.'s to be const. on preferred foliation.

$$K = \nabla_\mu u^\mu \quad 2 \nabla_\mu u^\mu \varphi - \varphi^2 \quad \delta\varphi \Rightarrow \varphi = \nabla_\mu u^\mu = K.$$

$$- 2 u^\mu \nabla_\mu \varphi - \varphi^2$$

$$- 2 \frac{\nabla_\mu \varphi}{\sqrt{g_{\mu\nu}}} \nabla^\mu \varphi - \varphi^2$$

Niaysh:  $(1-\lambda)K^2$  term

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$$K = \nabla_n u^m \quad 2 \nabla_n u^m \varphi - \varphi^2 \quad \delta\varphi \Rightarrow \varphi = \nabla_n u^m = K$$

$$- 2 u^m \nabla_n \varphi - \varphi^2$$

$$- 2 \frac{\nabla_n \varphi}{\sqrt{g_{mn} \delta\varphi}} \nabla_n \varphi - \varphi^2 \xRightarrow{\delta\psi_{qm}} \varphi = \varphi(\psi)$$



Naively:  $(1-\lambda)k^2$  term

claim:  $k$  fixed by b.c.'s to be const. on prefixed  $\mathcal{M}$  action

$$k = \nabla_\mu u^\mu$$

$$2 \nabla_\mu u^\mu \varphi - \varphi^2$$

$$\delta q \Rightarrow \varphi = \nabla_\mu u^\mu = k.$$

$$- 2 (u^\mu \nabla_\mu \varphi - \varphi^2)$$

$$\varphi = \varphi(\psi, \partial^\nu \psi \partial_\alpha \psi)$$

$$- 2 \frac{\nabla_\mu \varphi}{\sqrt{g_{\mu\nu}}} \nabla^\mu \varphi - \varphi^2$$

$$\delta \psi_{\mu\nu} + b.c. \Rightarrow \varphi = \varphi(\psi)$$

Stability -- projectable case



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Scalar constraint dropped, preferred time

→ true hamiltonian      no "problem of time".

$$E_{ADM} = \int$$

asymptotically flat case



$$E_{ADM} = \mathcal{E} = \int_{\Sigma} \{ (\mathcal{G}_{ik} - T_{ik}) + j_k \} \lambda \Sigma^k$$

asymptotically flat case

$$E_{ADM} = \mathcal{E} = \int_{\Sigma} \{ (\Theta_{ik} - T_{ik}) + j_k \} \underline{d^3\Sigma^k}$$

asymptotically flat case

$\int_{\Sigma} j_k \underline{d^3\Sigma^k} \geq 0$   
 can be arranged  
 (Witten)



$$E_{ADM} = \mathcal{E} = \int_{\Sigma} \left\{ (G_{\mu\nu} - T_{\mu\nu}) + j_{\mu} \right\} \underline{d}\Sigma^{\mu}$$

asymptotically flat case

$$= \int_{\Sigma} \mathcal{H}_{\mu\nu} d\Sigma^{\mu\nu}$$

$\int_{\Sigma} j_{\mu} d\Sigma^{\mu} \geq 0$   
can be arranged

(Witten)

$$E_{ADM} = \mathcal{E} = \int_{\Sigma} \left\{ \underbrace{(G_{\mu\nu} - T_{\mu\nu})}_{T_{\mu\nu}^{HL}} + j_{\mu} \right\} \underbrace{d^3x}_{n^{\mu} d^3x}$$

asymptotically flat case

$$= \int_{\Sigma} g^{HL} \omega_{\mu} \omega_{\nu}$$

$$j_{\mu} n^{\mu} \geq 0$$

can be arranged

$$\int_{\Sigma} g^{HL} d^3x$$

(Witten)



$$E_{ADM} = \oint_{\Sigma} \left\{ (G_{in} - T_{in}) + \underline{j}_n \right\} \underline{d}\Sigma^n$$

asymptotically flat case

$$= \oint_{\Sigma} S^{HL} \mu_{HL}$$

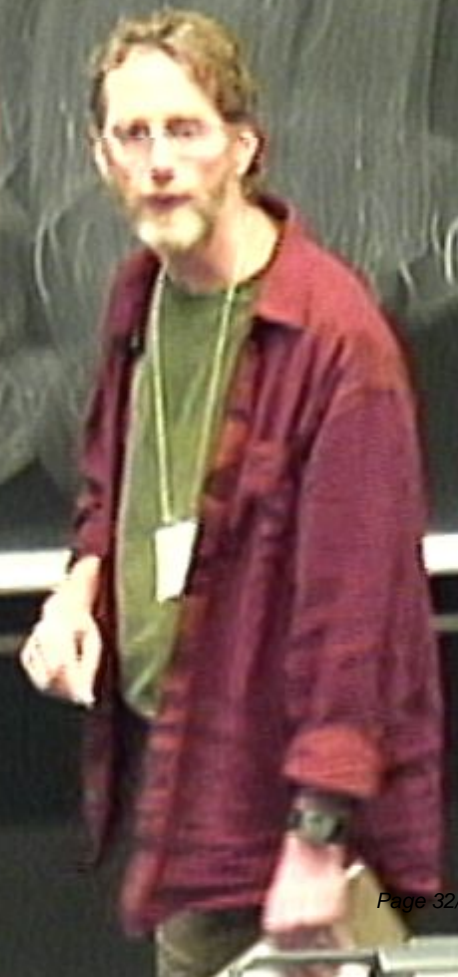
$\oint_{\Sigma} j_n \geq 0$   
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$$\int_{\Sigma} S^{HL} d^3x = 0$$



A particles slower  
B particles faster





A partides slowen

B partides fosten

