

Title: Hydrodynamic long-time tails from AdS/CFT

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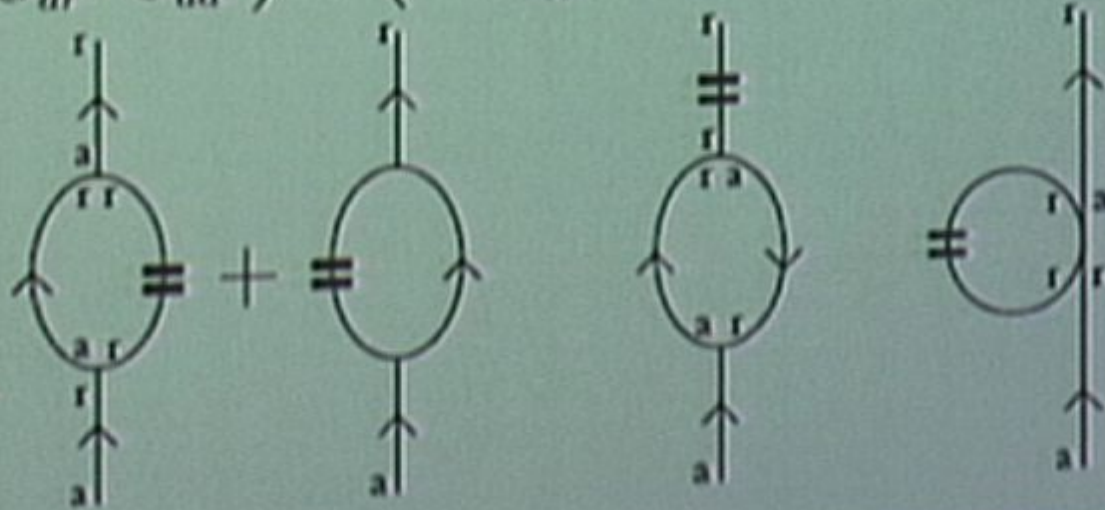
Abstract: For generic field theories at finite temperature, a power-law falloff of correlation functions of conserved currents at long times is a prediction of non-linear hydrodynamics. We demonstrate, through a one-loop computation in Einstein gravity in Anti de Sitter space, that this effect is reproduced by the dynamics of black hole horizons. The result is in agreement with the gauge-gravity correspondence.

Diagrams

- Not every diagram is important at long times
- In Schwinger-Keldysh formalism in ra -basis

$$\phi_r = \frac{1}{2}(\phi_1 + \phi_2), \quad \phi_a = \phi_1 - \phi_2.$$

$$G = \begin{pmatrix} G_{rr} & G_{ra} \\ G_{ar} & G_{aa} \end{pmatrix} = \begin{pmatrix} -i(G_R - G_A) \left(\frac{1}{2} + n_B(\omega)\right) & -iG_R \\ -iG_A & 0 \end{pmatrix}$$

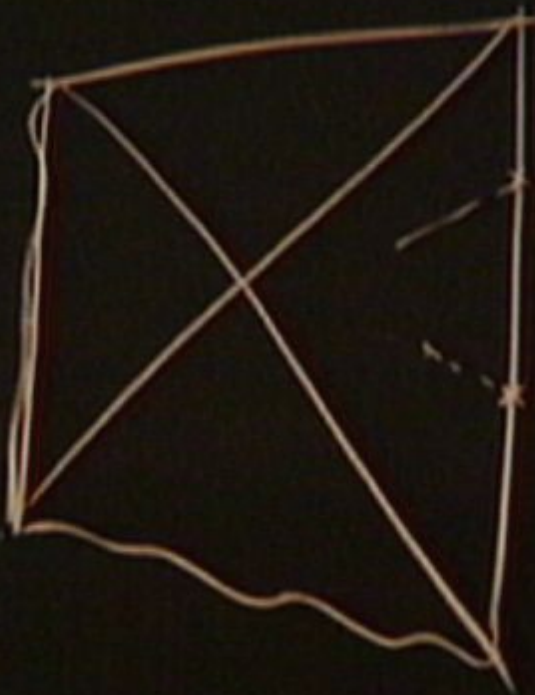


AdS-Schw

$V=4$

$S \sim \mathbb{Z}_2$

$J^M = P_{2M}$



AdS-Schw



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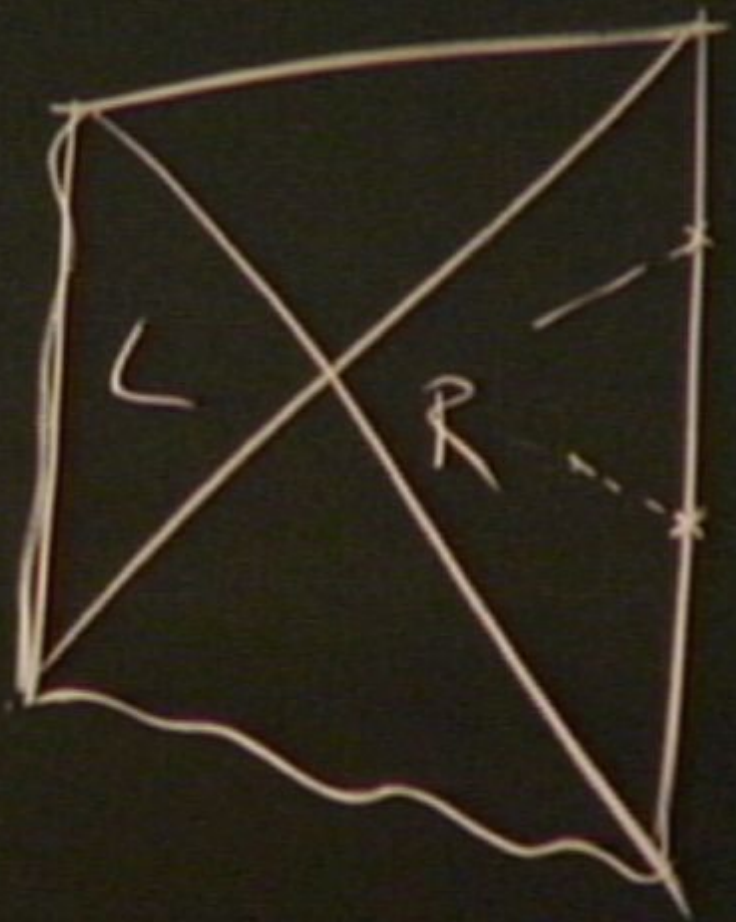
P_{21}

AdS-Schw

$N=4$

$S \sim \mathbb{Z}_2$

$\bar{J}^M = P_{\mu}^M$

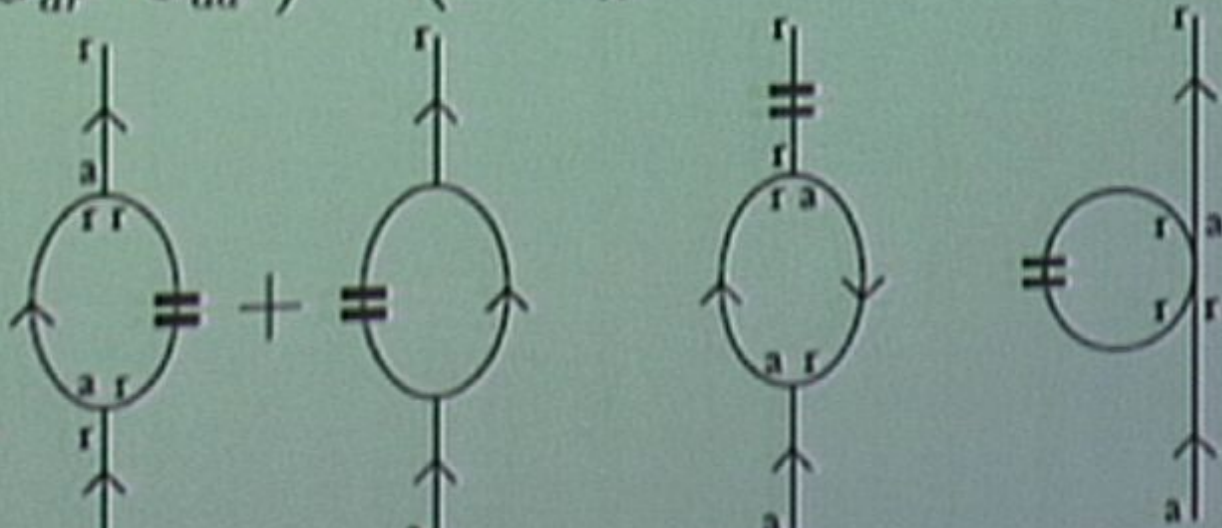


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The origin of the boundary KMS conditions

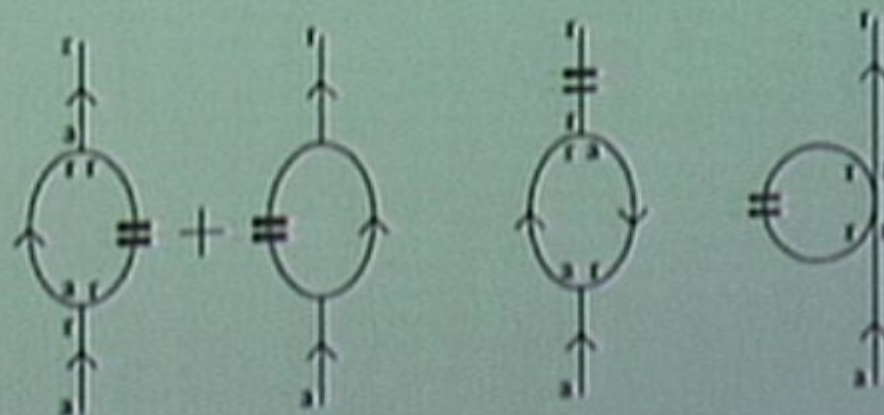
- ✦ In QFT in curved space limit of QG: BH is in Thermal equilibrium with the environment to all orders in perturbation theory [Gibbons-Perry '76]

$$G_{12}(r, r'; p) = -in_B(p)(G_R(r, r'; p) - G_A(r, r'; p)), \quad \text{etc.}$$

- ✦ Fluctuations around H-H state are thermal to all orders in perturbation theory
- ✦ The **KMS** conditions are obeyed in the bulk
- ✦ So in an AdS/CFT setting, the boundary **KMS** conditions are inherited from the bulk

Dominant fields in the hydrodynamic regime: Generalities

- Both perturbations in the loop: long-lived and diffusive
- Charge fluctuations dominate over current perturbations $A_t/A_z \sim k/\omega$
- h_{tx} dominates over h_{zx} : momentum fluctuation dominates over stress
- Spin-two fluctuations are suppressed by viscous effects



Gravity Vertices j^x

- Third variation of the Yang-Mills action

$$V_{A-A-g} = \delta_g \delta_A \delta_A S_{YM} = - \int d^{d+1}x \delta A_x \partial_M \left[\delta_g \delta_A \left(\frac{\sqrt{-g} F^{xM}}{g_{d+1}^2} \right) \right]$$

$$= \int d^{d+1}x \frac{\delta A_x}{g_{d+1}^2} \partial_M \left[\sqrt{-g} (\delta F^{\mu M} h^x{}_\mu + \delta F^{x\mu} h^M{}_\mu - \frac{1}{2} \delta F^{xM} h^\sigma{}_\sigma) \right].$$

- Terms in the second sum: k -suppressed or zero by gauge choice

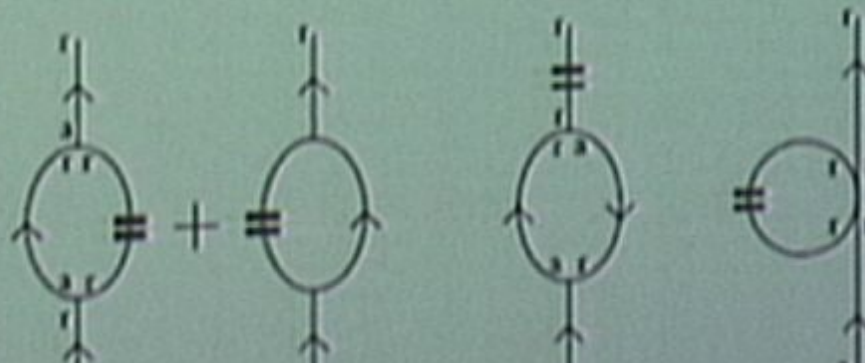
$$V_{A-A-g} = \int d^{d+1}x \delta A_x \frac{\sqrt{-g}}{g_{d+1}^2} \left[\delta F^{\mu r} \partial_r h^x{}_\mu - \frac{1}{2} \delta F^{xr} \partial_r h^\sigma{}_\sigma \right]$$

- The last term mixes a sound mode and diffusive mode: no long-time tail here

$$V_{A-A-g} = \int d^{d+1}x \delta A_x \frac{\sqrt{-g} \delta F^{tr}}{g_{d+1}^2} \partial_r h^x{}_t.$$

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$$\omega = \omega(k) \rightarrow \cdot$$

$$k \rightarrow \cdot$$

$$N = 4$$

$$S \sim N_c^2$$

$$T \sim \left(k_{\mu\nu} - \frac{1}{2} k h_{\mu\nu} \right) \sim \pi_{\mu\nu}^h = p_{\mu\nu}^h$$

Gravity Vertices: j^x

- ↳ Rewriting in terms of the bulk local currents

$$\begin{aligned} V_{A-A-g} &= \int d^{d-1}x \delta j^t(\tau) \delta t^t_x(\tau) \int \frac{dr 16\pi G g^{xx}}{\sqrt{-g} |g^{tt}|} + \mathcal{O}(1/\tau) \\ &= \int d^{d-1}x \frac{\delta j^t(\tau) \delta t^t_x(\tau)}{\epsilon + p}, \end{aligned}$$

- ↳ Gauge-gauge-gauge vertices are also sub-leading :
 $\delta A^x \partial_r \delta A^t$ and $\delta A^t \partial_r \delta A^x$ involve current fluctuation

$$V \sim g(r) \\ A \cdot A - g$$

$$N = 4$$

$$S \sim \frac{2}{c^2}$$

ADS-SC



$$T^{M \nu}$$

$$(T_{M \nu} - \frac{1}{2} K h_{M \nu}) \sim \pi_{M \nu} = P_{M \nu}$$

Gravity Vertices: t^{xy}

- Third variation of the Einstein-Hilbert action

$$\begin{aligned}
 V_{g-g-g} &= \delta_g \delta_g \delta_g S_{EH} = \frac{1}{16\pi G} \delta^3 \int d^{d+1}x \sqrt{-g} (R + d(d-1)/\ell^2) \\
 &= -\frac{1}{8\pi G} \int d^{d+1}x h_{xy} \delta^2 \sqrt{-g} g^{x\mu} g^{y\nu} \left[R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R - \frac{d(d-1)}{2\ell^2} g_{\mu\nu} \right].
 \end{aligned}$$

- Two derivative action

$$V_{g-g-g} = -\frac{1}{8\pi G} \int d^{d+1}x \sqrt{-g} h^{xy} \delta^2 \left[2K_{\sigma x} K^{\sigma}_y - K^{\sigma}_{\sigma} K_{xy} - \frac{1}{2} \partial_r^2 g_{xy} \right]$$

- No long-lived spin-2:

$$V_{g-g-g} = \frac{1}{16\pi G} \int d^{d+1}x \sqrt{-g} h^{xy} |g_{tt}| \delta K^t_x \delta K^t_y$$

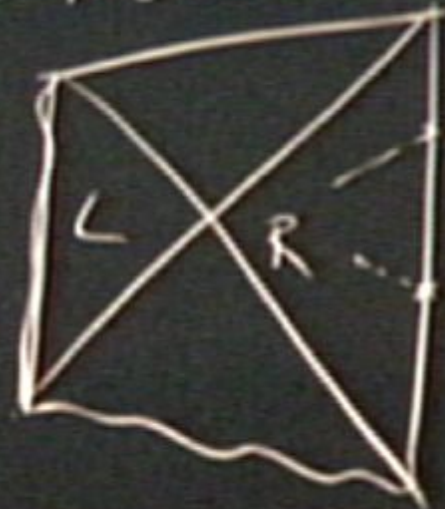
$$V \sim g(x) \int \dots$$

$$N = 4$$

$$S \sim \dots$$

$$T \sim (r_{m+1} - \frac{1}{2} k h_m) \sim \dots$$

ADS-Schw.



Diagrams

At one loop



(a)



(b)



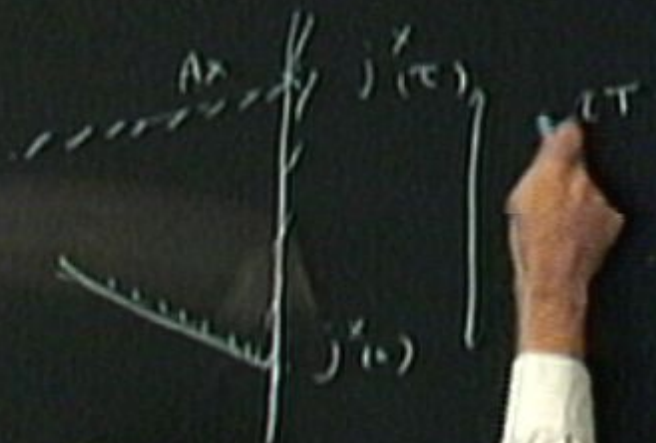
(c)

Feynman diagrams contributing to (a) current correlator (b) stress tensor correlator. Four-point vertices (c) will be found to have negligible effects. Wavy lines are bulk Yang-Mills fields and double lines are gravitons.

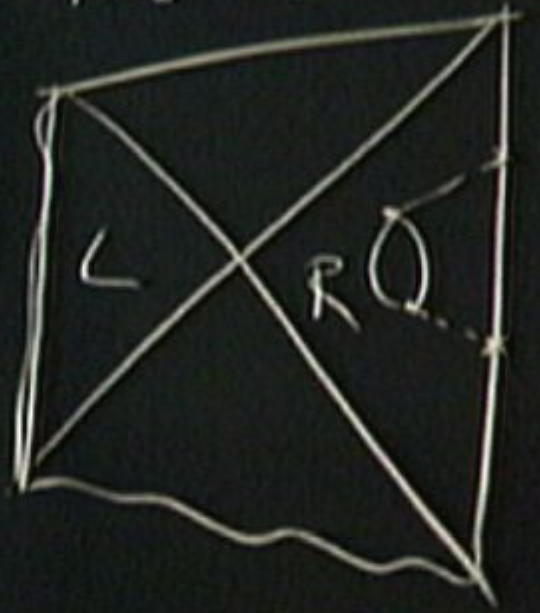
$V \sim g(r)$
 $A \cdot A - g$

$T \sim$
 $M \cdot V$

horizon



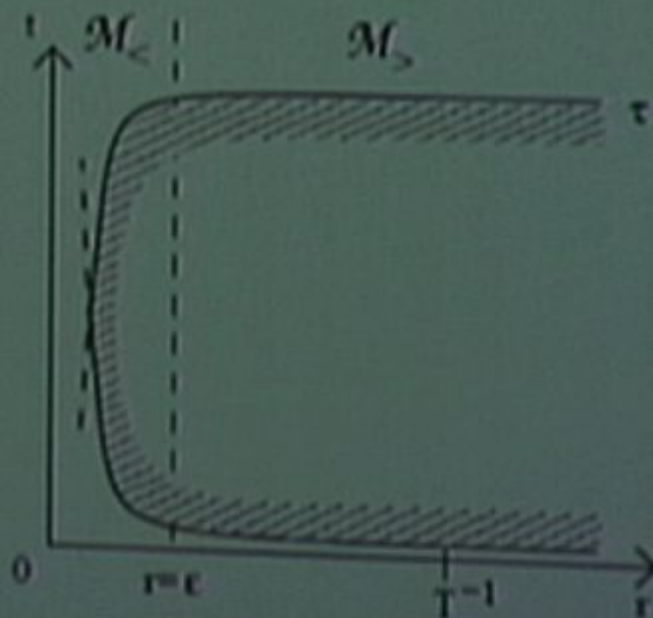
AIS-Schw



External Wavefunction

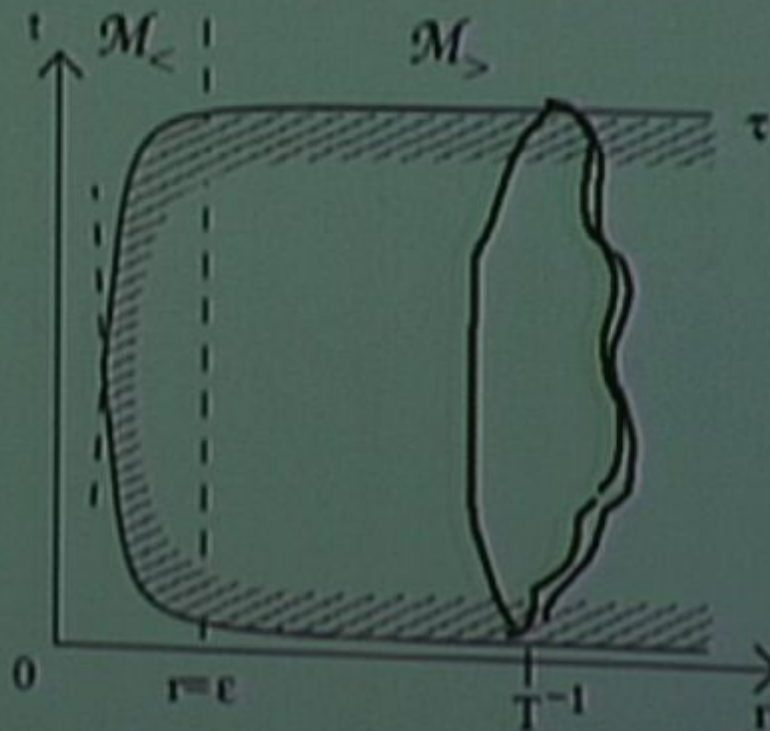
- A_x relaxes microscopically. A_x dissipate on time scale T^{-1}
- The field profile falls into the horizon along a null ray

$$A_x^{\text{ret}}(r, t) \rightarrow A_x^{\text{ret}}(t - r_*(r))$$



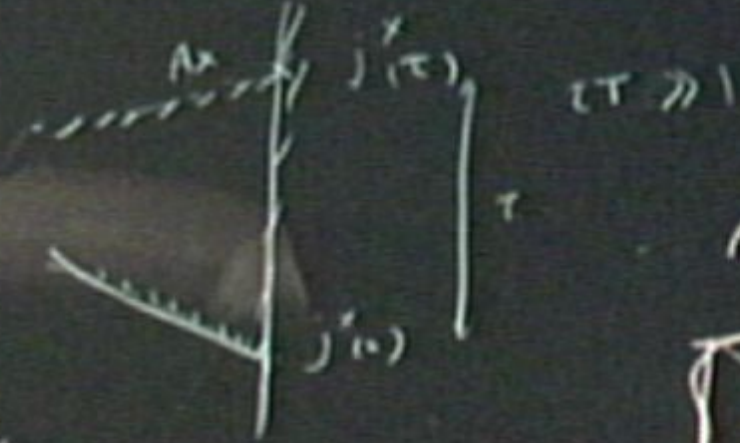
Emergence of Hydrodynamics from Anti de Sitter Space

→ $r-t$ plane : the causal diamond: insertions at 0 and τ



Emergence of non-linear Hydro from Anti de Sitter Space

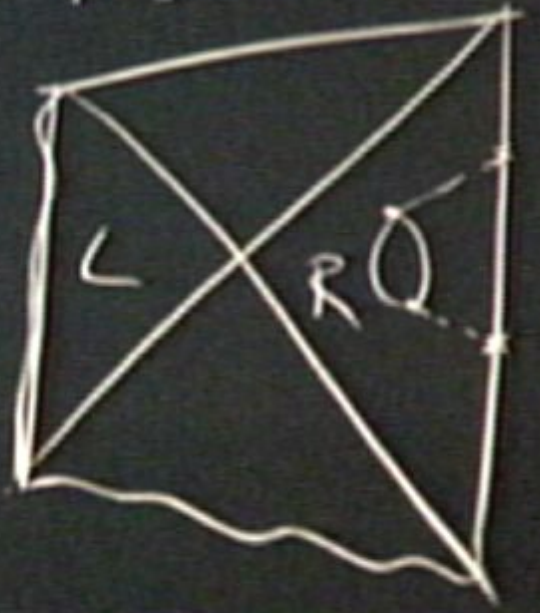
- Tip of the half-almond $r_0(\tau) \approx \frac{\sqrt{2}}{2\pi T} e^{-\pi T \tau}$.
- The shaded region $T^{-1} \ll t$
- Not all of the half-almond contributes: the external disturbances relax on microscopic scale T^{-1}
- Vertices near the tip would correspond to a scale of non-locality in the boundary of the order $\sim \tau$
- In contrast, in hydro the vertices are instantaneous on the microscopic time scale T^{-1}
- A minimal choice for ϵ is $\epsilon(\tau) \sim \frac{1}{T} e^{-\pi T \tau / 2}$.
- Separation in $\mathcal{M}_>$ is $\sim \tau$.
- The Membrane picture for hydrodynamic is lost from the point of view of the boundary observer




$V \sim g(x)$
 $AA-g$
 x
 horizon
 $N = -4$

$S \sim \frac{1}{2} \pi$

AdS-Schw



$$T \sim \left(\frac{1}{2} \pi - \frac{1}{2} \pi \right) \sim \frac{1}{2} \pi$$



Bulk-bulk correlators

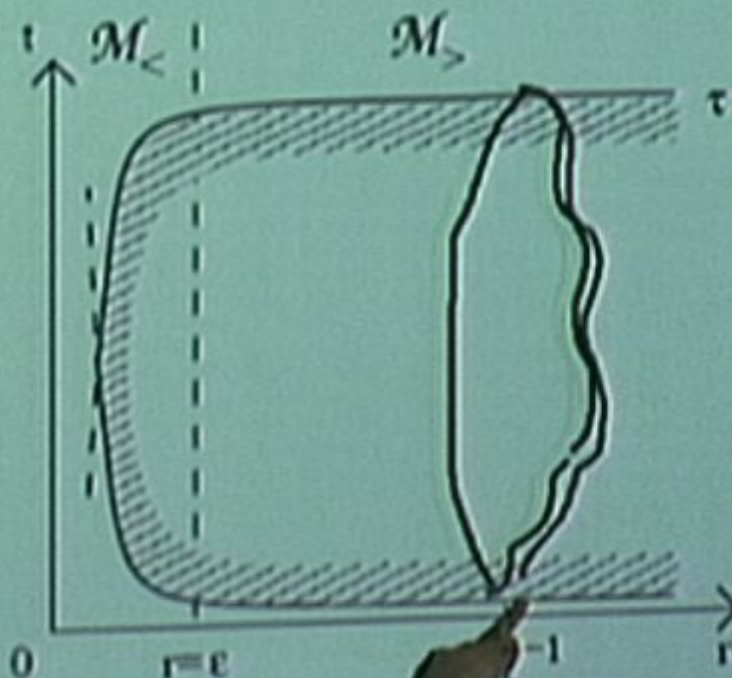
- The full bulk-bulk propagator

$$\begin{aligned} G_{\text{R}}^{tt}(r, r'; p) + \frac{\delta(r-r')}{C(r)g^{zz}} &= i\sigma k^2 \frac{G(r)F(r')\theta(r-r') + F(r)G(r')\theta(r'-r)}{2(\omega + i\frac{\sigma}{\Xi}k^2)} \\ &= \frac{i\sigma k^2}{\omega + i\frac{\sigma}{\Xi}k^2} + \mathcal{O}(\omega \log \frac{1}{r}, k^2). \end{aligned}$$

- It loses its radial dependence up to corrections which are only large in $\mathcal{M}_<$

Emergence of Hydrodynamics from Anti de Sitter Space

→ $r-t$ plane : the causal diamond: insertions at 0 and τ





Bulk Variable

- The bulk computation is best organized in momentum variables rather than in field space.

$$t^{\mu}_{\nu} = 2 \frac{\delta S_{\text{bulk}}}{\delta \gamma_{\mu\nu}} = \frac{\sqrt{-\gamma}}{8\pi G_N} (\delta^{\mu}_{\nu} K - K^{\mu}_{\nu}),$$

$$j^{\mu} = \frac{\delta S_{\text{bulk}}}{\delta A_{\mu}} = \frac{\sqrt{-\gamma}}{g_{d+1}^2} F^{\mu r},$$

- Consider them as extending in the bulk: local density of the currents

$$t^{\mu}_{\nu} = t^{\mu}_{\nu}(r), \quad j^{\mu} = j^{\mu}(r)$$

Radial Flow equation for the Hydrodynamic Variable

- Bulk Hamiltonian (Einstein and Yang-Mills): the radial evolution of the momenta. It can be organized written as a derivative expansion


$$\partial_r(\delta t^\mu{}_\nu) = \frac{1}{8\pi G} \delta(\sqrt{-g}^{(d)} R^\mu{}_\nu) = 0 + \mathcal{O}(\omega^2, q^2) \quad \mu \neq \nu,$$

$$\partial_r(\delta j^\mu) = \frac{1}{g_{d+1}^2} \delta(\partial_\nu \sqrt{-g} F^{\nu\mu}) = 0 + \mathcal{O}(\omega^2, q^2),$$

$$\frac{1}{\sqrt{g}} \partial_r(\sqrt{g} K^\mu{}_\nu) = {}^{(d)}R^\mu{}_\nu + \frac{d}{\ell^2} \delta^\mu{}_\nu.$$

- Holds up to exponentially close to the horizon due to oscillations

$$r \sim e^{-2\pi T/\omega}$$



Relation to Long-time tails in Black hole Physics

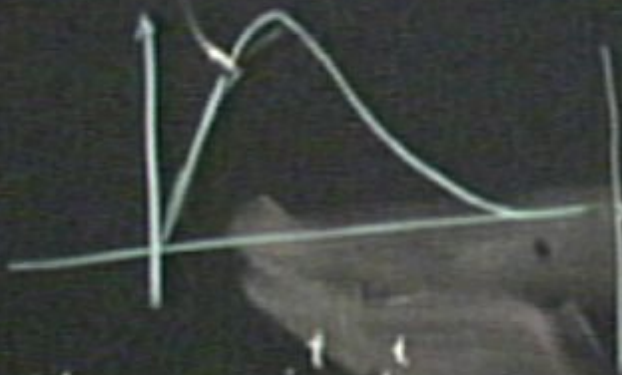
- ↳ Evolution of probe fields on black hole spacetimes
- ↳ Late-time power-law tails in flat space $\Phi \sim t^{-n}$
- ↳ In AdS space they do not exist. Dirichlet boundary condition kills it
- ↳ We have argue that they do exist even in AdS but at non-linear level
- ↳ The argument comes from an unlikely source: hydrodynamics!

$|\phi\rangle \sim \frac{1}{\sqrt{2}} (|0\rangle + |1\rangle)$



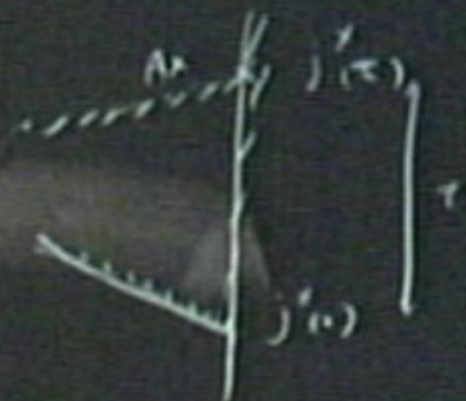
Outlook

- ⬇ Emergence of non-linear hydrodynamics from anti de Sitter Space
- ⬇ Counter-intuitive: late-time tails must have come from deep IR: fluctuations are tied to the horizon but the scale of vertices are microscopic since they are local in time
- ⬇ The role of the asymptotic region not clear. This analysis must be true in flat space as well
- ⬇ $D=3$ and below is very interesting. Work in progress
- ⬇ Mode-coupling theory of critical systems from gravity



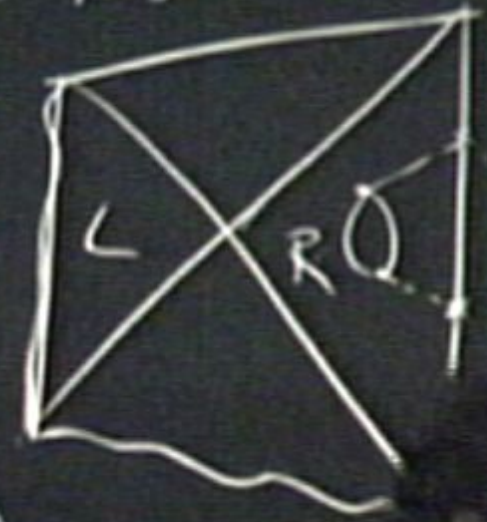
$$V_{AA-g} \sim g(r) j^t + x$$

$$N = 4$$



$(T \gg 1)$

AdS-Schw



$$|\Phi| \sim \frac{1}{r \rightarrow 0}$$

$$T^{\mu\nu} \sim (r_{\mu\nu} - \frac{1}{2} k h_{\mu\nu}) \sim j^{\mu} = p_{\mu}^{\nu}$$

horizon

Long-Time Tails in Hydrodynamic Perturbation Theory

- Hydrodynamic diagrams
- In hydrodynamics perturbation theory

$$\int d^{d-1}x \langle j^a(\tau, x) j^b(0) \rangle \simeq \int d^{d-1}x \langle \rho(\tau, x) u^a(\tau, x) \rho(0) u^b(0) \rangle$$

$$\langle \rho(\tau, x) u^a(\tau, x) \rho(0) u^b(0) \rangle = \langle \rho(\tau, x) \rho(0) \rangle \langle u^a(\tau, x) u^b(0) \rangle$$



$$\int d^{d-1}x \langle j^a(\tau, x) j^b(0) \rangle = \frac{T \Xi \delta^{ab} d - 2}{\epsilon + p} \frac{1}{d - 1} \frac{1}{[4\pi(D + \gamma_\eta) |\tau|]^{\frac{d-1}{2}}}$$

- Long-Time tails are suppressed by entropy density