

Title: Quantum Field Theory (PHYS 601) - Lecture 6

Date: Oct 05, 2009 09:00 AM

URL: <http://pirsa.org/09100111>

Abstract:

# Propagators

We could ask the following question:

# Propagators

We could ask the following question:  
disturb the field at point  $y$ . What  
is the amplitude.

$$\langle 0 | \phi(x) \phi(y) | 0 \rangle$$

$$\langle 0 | \phi(x) \phi(y) | 0 \rangle = \int \frac{d^3 p d^3 p'}{(2\pi)^6} \frac{1}{\sqrt{4E_p E_{p'}}} \langle 0 | a_p a_{p'}^\dagger | 0 \rangle e^{-i(p \cdot x - p' \cdot y)}$$

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$$= \int \frac{d^3 p}{(2\pi)^3} \frac{1}{2E_p} e^{-ip \cdot (x-y)}$$

$\equiv D(x-y)$  This is the propagator

For spacelike separation  $(x-y)^2 < 0$

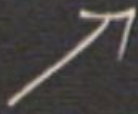
For spacelike separation  $(x-y)^2 < 0$

$$D(x-y) \sim e^{-m|x-y|} \neq 0.$$

# Feynman Propagator

The important quantity for interacting field theories is the Feynman propagator.

$$\Delta_F(x-y) = \langle 0 | T \phi(x) \phi(y) | 0 \rangle$$

time ordering 



$$\equiv \left\{ \begin{array}{ll} \langle 0 | \phi(x) \phi(y) | 0 \rangle & x^0 > y^0 \\ \langle 0 | \phi(y) \phi(x) | 0 \rangle & y^0 > x^0 \end{array} \right.$$

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Claim.  $\Delta_F(x-y) = \int \frac{d^4 p}{(2\pi)^4} \frac{i}{p^2 - m^2} e^{-i p \cdot (x-y)}$

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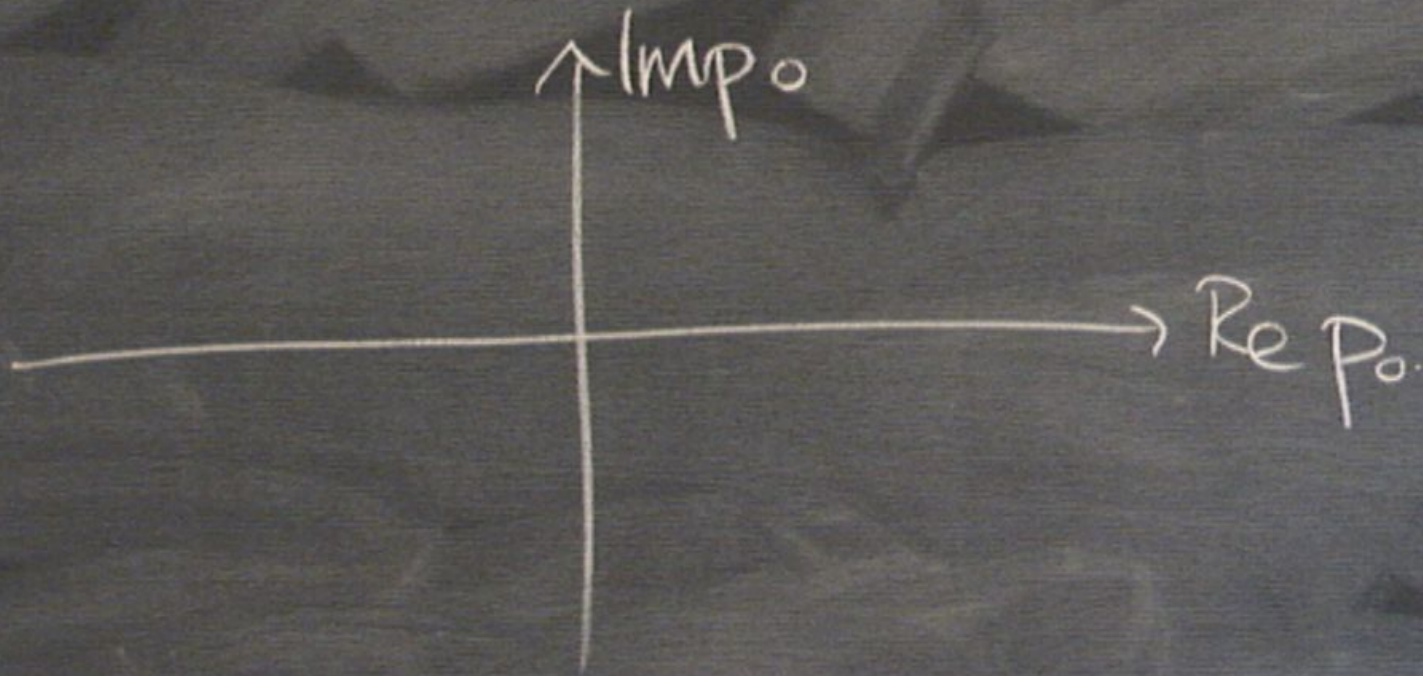
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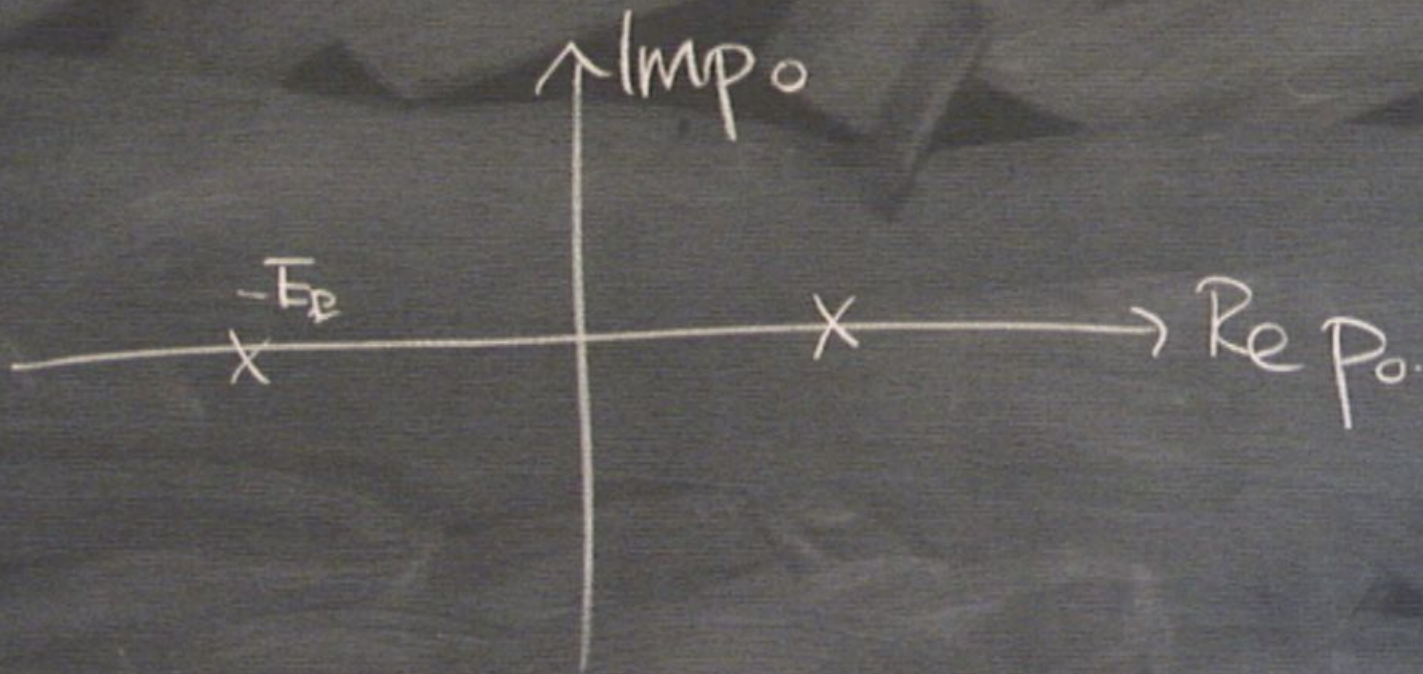
Claim:  $\Delta_F(x-y) = \int \frac{d^4 p}{(2\pi)^4} \frac{i}{p^2 - m^2} e^{-ip \cdot (x-y)}$

This isn't yet well-defined because, for  $\not{x}$ , the  $\int d^4 p_0$  has a pole when  $p_0^2 = p^2 + m^2$

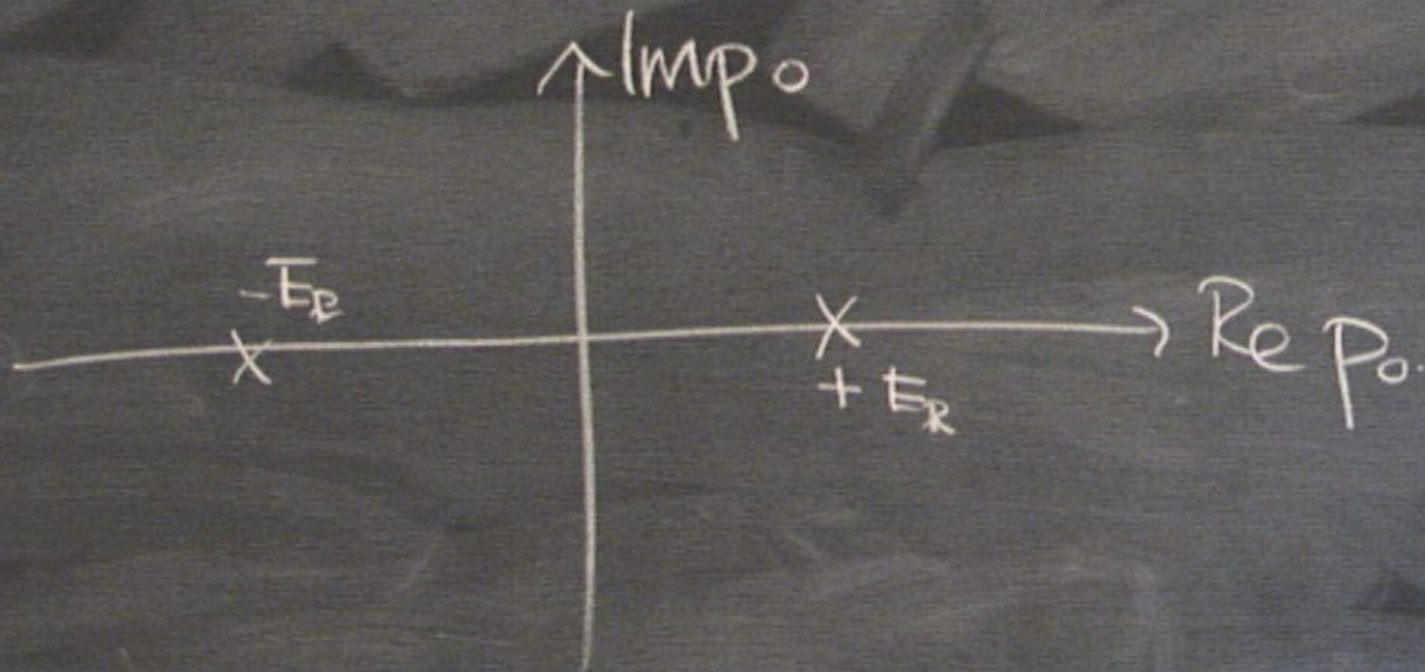
We'll define the integral by the contour  
in the complex  $p^o$  plane



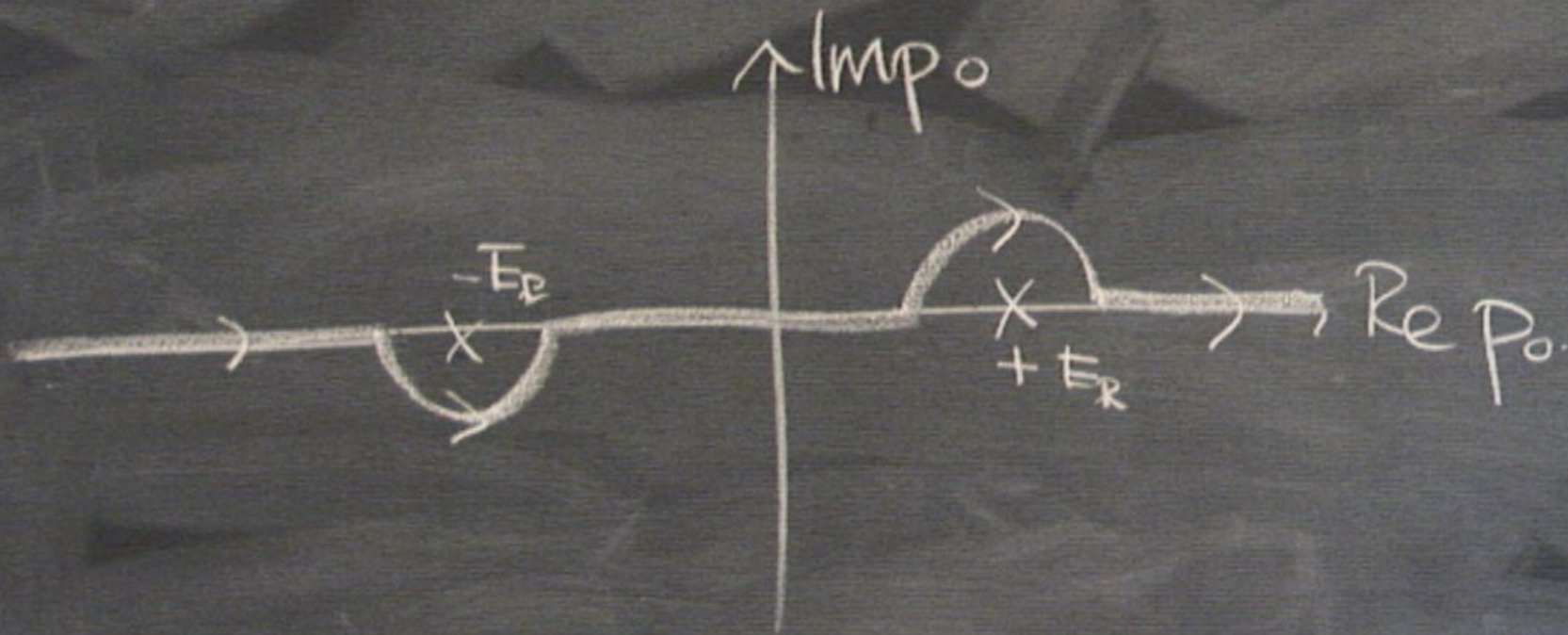
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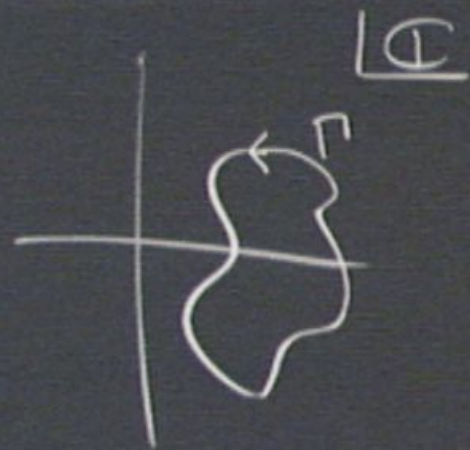
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## Residue Thm.

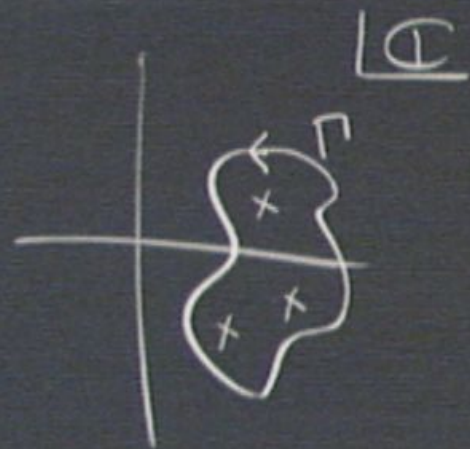
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$$\oint_{\Gamma} f(z) = 2\pi i \sum \text{Residues}$$



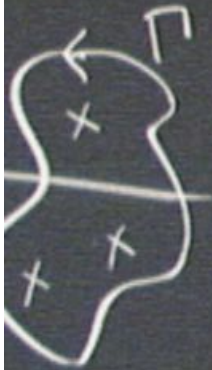
## Residue Thm:

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$$f(z) =$$

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$\mathbb{C}$



Residue Thm: something

$f(z)$  is ~~holomorphic~~ analytic. Let  $\Gamma$  be a closed anticlockwise contour in  $\mathbb{C}$ .

Near  $z_0$

$$(z-z_0)^2 + (z-z_0) + \text{const}$$

$$\oint_{\Gamma} f(z) = 2\pi i \sum \text{Residues}$$

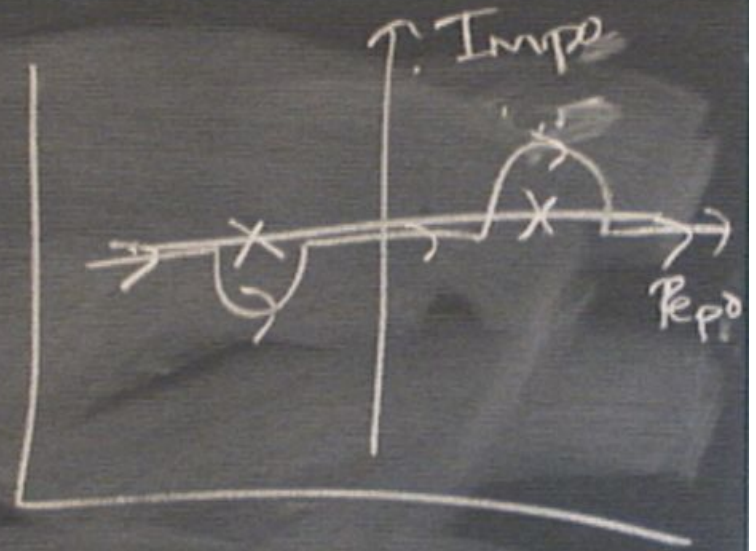
$$+ \frac{\text{Res}}{(z-z_0)} + \frac{\#}{(z-z_0)^2} + \dots$$

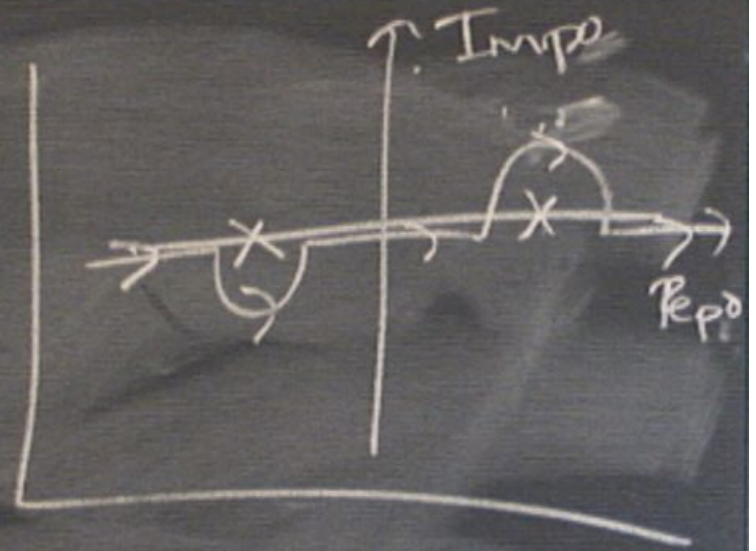
Proof:

$$\frac{1}{p^2 - m^2} = \frac{1}{(p^0)^2 - E_p^2}$$

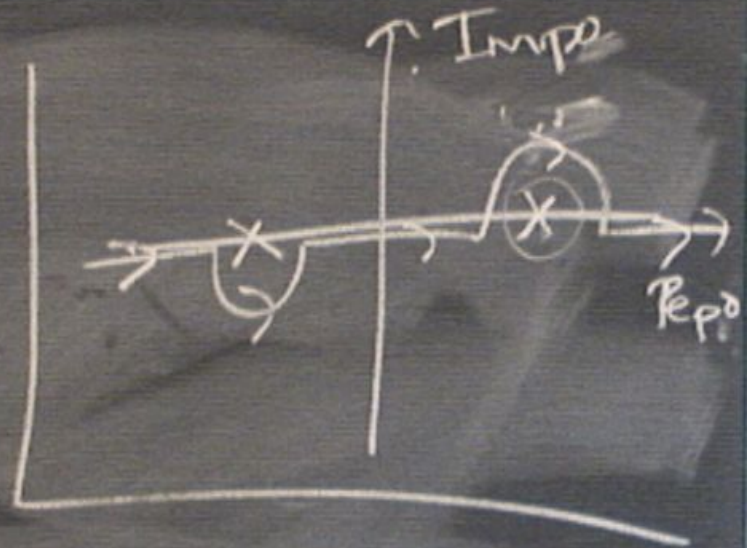
$$= \frac{1}{(p^0 - E_p)(p^0 + E_p)}$$

Residue at  $p^0 = +E_p$  is  $+\frac{1}{2E_p}$





en  $x^0 \rightarrow y^0$ , we close the  
 star in lower-half plane



$$p_0 \rightarrow -i\infty$$

$$\Rightarrow e^{-ip^0(x^0 - y^0)}$$

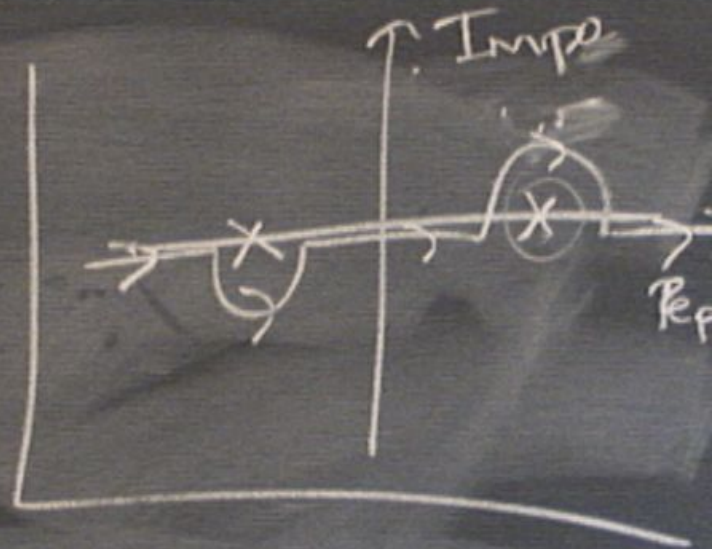
$$\rightarrow e^{-\infty} = 0$$

$\Rightarrow \int dp^0$  picks up residue at  $\mathbb{F}_{\mathbb{R}}$

$$\int \frac{d^4 p}{(2\pi)^4} \frac{i}{p^2 - m^2} e^{-ip \cdot (x-y)}$$



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Residue is  $+1/2E_p$

$$\int \frac{d^4 p}{(2\pi)^4} \frac{i}{p^2 - m^2} e^{-ip \cdot x}$$

$$\Rightarrow \Delta(x-y) = \int \frac{d^3 p}{(2\pi)^4} \frac{- (2\pi i)}{2E_p} \cdot i e^{-i E_p (x^0 - y^0)} \times e^{i p \cdot (x - y)}$$

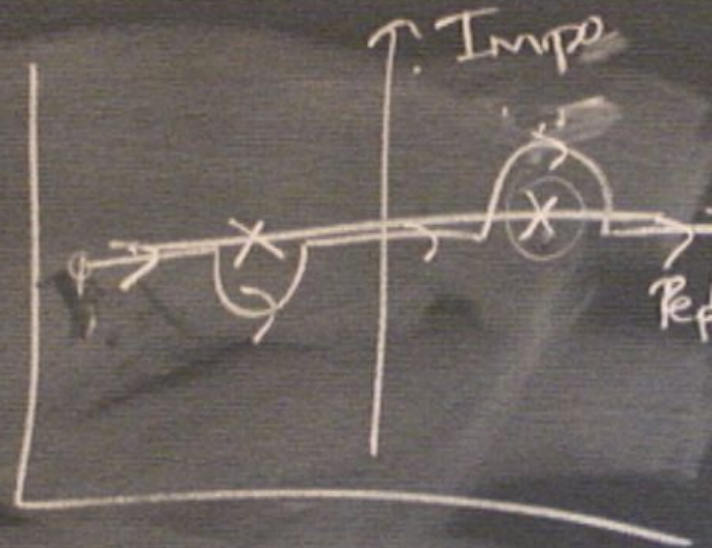
because clockwise contour

Residue

$$= \int \frac{d^3 p}{(2\pi)^3} \frac{1}{2E_p} e^{-i p \cdot (x - y)}$$

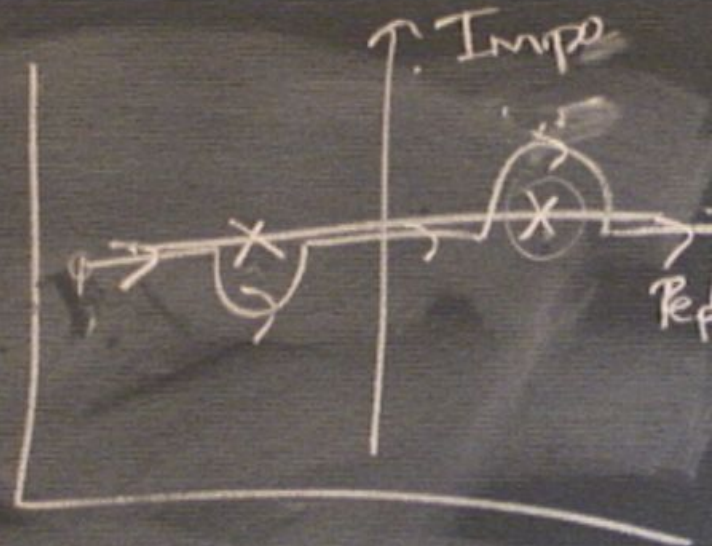
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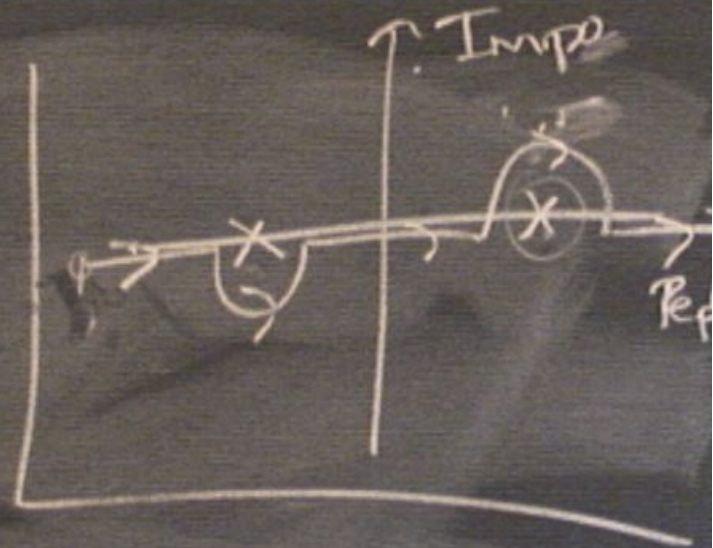
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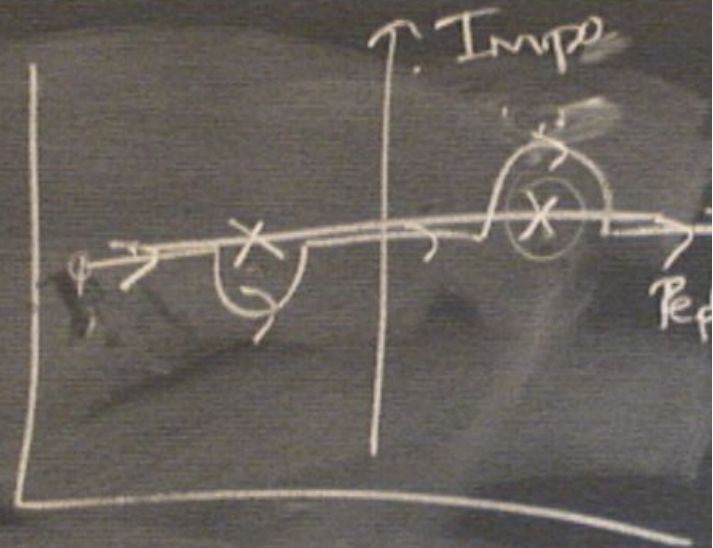
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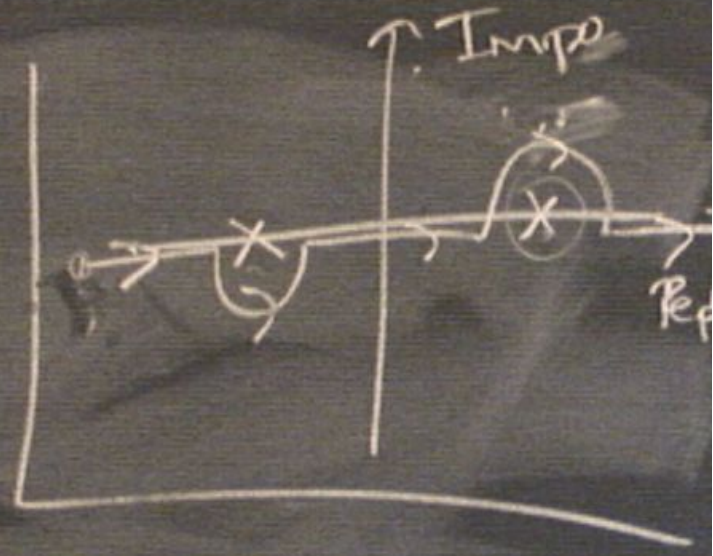
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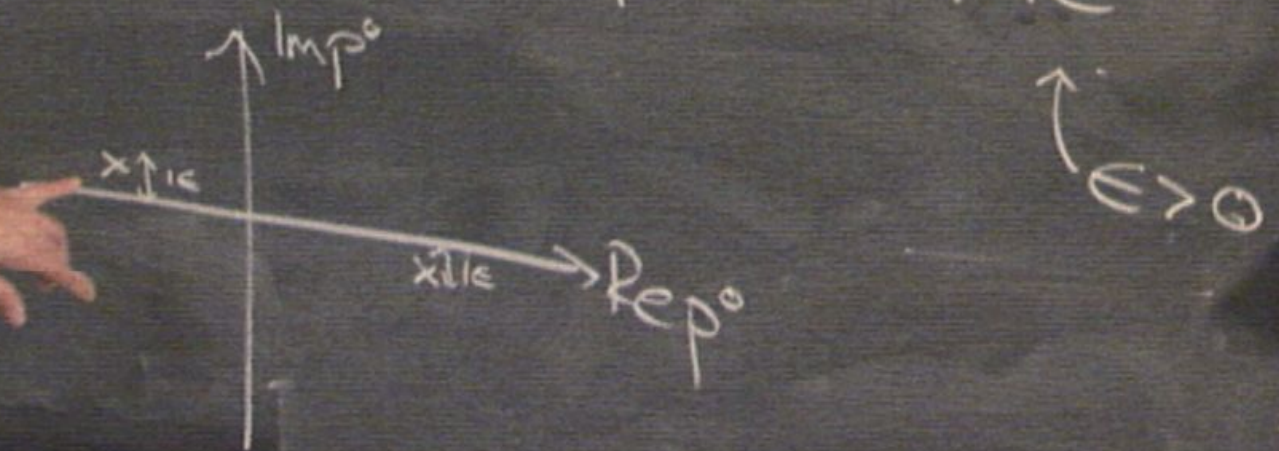
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Instead of specifying the contour,  
we may instead write

$$\Delta_F(x-y) = \int \frac{d^4 p}{(2\pi)^4} \frac{i e^{-i p \cdot (x-y)}}{p^2 - m^2 + i\epsilon}$$





## Another Avatar

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$$(\partial_t^2 - \nabla^2 + m^2) \Delta_F(x-y) = \int \frac{d^4 p}{(2\pi)^4} \frac{i}{p^2 - m^2} (-p^2 + m^2) e^{-ip(x-y)}$$

$$= -i \int \frac{d^4 p}{(2\pi)^4} e^{-i p \cdot (x-y)}$$

$$= -i \delta^{(4)}(x-y)$$

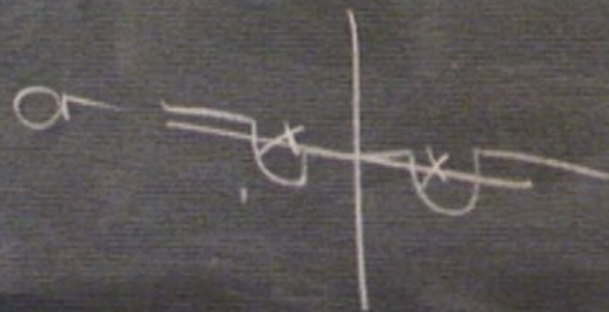
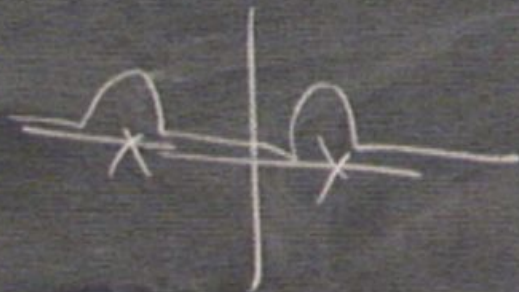
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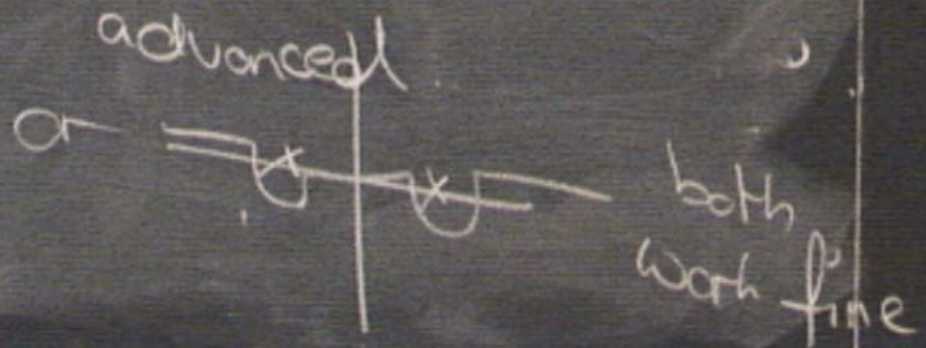
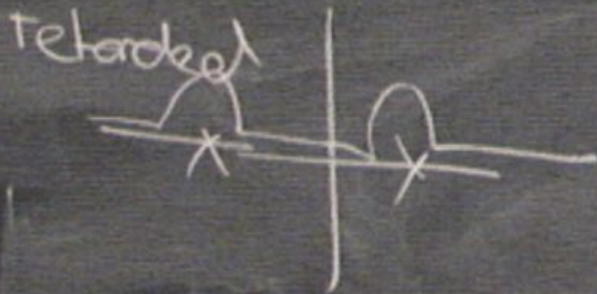


both work fine

$$= -i \int \frac{d^4 p}{(2\pi)^4} e^{-ip \cdot (x-y)}$$

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Consider the Lagrangian

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 - \sum_{n=3}^{\infty} \frac{\lambda_n}{n!} \phi^n$$

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$\lambda_n$  are called coupling constants.

Some dimensional analysis:

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$$[S] = 0$$

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↑  
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$$[\phi] = 1 \Rightarrow [\psi] = 1$$

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$$[k] = 0$$

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↑  
[ ] = 4

$$[\theta] = 1 \Rightarrow [\phi] = 1 \text{ and } [M] = 1$$

and  $[\lambda_n] = 4 - n$

what does " $\lambda_n$  small" mean?

$[\lambda_3] = 1 \implies$  dimensionless  
parameter is  $(\lambda_3 / E)$

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where  $E$  is energy scale of typical configuration

or process that we're interested in

• small perturbation of  $E \gg \lambda_3$

• large perturbation of  $E \ll \lambda_3$

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For relativistic theories,  $E > M \Rightarrow$  small perturbation of  $\ll M$

2)  $\frac{\lambda_4 \phi^4}{4!}$

$[\lambda_4] = 0 \Rightarrow \lambda_4 \ll 1$  is small perturbations

These perturbations are called marginal.

3)

on  $\lambda_5 \in M$

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$$3) \quad \frac{\lambda_n \phi^n}{n!} : [\lambda_n] < 0 \Rightarrow (\lambda E^{4-n}) \text{ is dimensionless}$$

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