Title: Quantum Field Theory (PHYS 601) - Lecture 2

Date: Sep 29, 2009 09:00 AM

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Abstract:

Recall : $\quad \phi_a(\hat{x}, t).$

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Lorentz Transformations: We want to construct relativistic field theories, with time and space placed on equal footing

The theory should be invariant under Lorentz transformations:

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$$\Lambda^\mu{}_\sigma = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \cos\theta & -\sin\theta & 0 \\ 0 & \sin\theta & \cos\theta & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \quad \text{rotates by } \theta \text{ about } z\text{-axis}$$

$$\text{or} \quad \Lambda^{\mu}_{\ \sigma} = \begin{pmatrix} \gamma & -\gamma r & 0 & 0 \\ -\gamma r & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

is a boost by v along the x-axis, with $\gamma = \dfrac{1}{\sqrt{1-v^2}}$

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The Lorentz transformations have a representation on fields. For a scalar field, this is

$$\phi(x) \rightarrow \phi'(x) = \phi(\Lambda^{-1}x)$$

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This is an <u>active</u> transformation. A theory is said to be Lorentz invariant whenever $\phi(x)$ is a sol^{n} $\Rightarrow \phi(\Lambda^{-1}x)$ is also a solution. We can ensure that this is true by requiring that S is invariant.

○ Exercise: check this for

$$S = \int d^4x \; \frac{1}{2} \partial_\mu \phi \, \partial^\mu \phi - \frac{1}{2} m^2 \phi^2$$

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An example of a Lorentz transformation on a vector field, $A_\mu(x)$.

$$A_\mu(x) \rightarrow A'_\mu(x) = \Lambda_\mu{}^\nu A_\nu(\Lambda^{-1} x)$$

Symmetries & Noether's Theorem

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Noether's Thm: Every continuous symmetry of the Lagrangian gives rise to a conserved current $j^\mu(x)$, such that the equations of motion imply

$$\boxed{\partial_\mu j^\mu(x) = 0}$$

i.e. $\dfrac{\partial}{\partial t} j^0 + \vec{\nabla} \cdot \vec{j} = 0$

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<u>note</u>. A conserved current $\Rightarrow$ a conserved <u>charge</u>

$$Q = \int_{\mathbb{R}^3} d^3x \, j^0$$

which follows from

$$\dot{Q} = \int_{\mathbb{R}^3} \frac{\partial j^0}{\partial t} \, d^3x$$

$$= -\int_{\mathbb{R}^3} \nabla \cdot \vec{j} \, d^3x$$

$$= 0 \quad \text{assuming } \vec{j} \to 0$$
$$\text{as } x \to \infty.$$

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But the existence of a current is stronger — it tells us that "charge" is conserved **locally**. Consider a fixed volume V

$$Q_V = \int_V j^0 \, d^3\underset{\sim}{x}$$

$$\dot{Q}_V = -\int_V \nabla \cdot j \, d^3\underset{\sim}{x} = -\int_A j \cdot d\underset{\sim}{S}$$

Proof of Noether's thm

We'll work infinitesimally

$$\phi_a(x) \rightarrow \phi'_a(x) = \phi_a(x) + \epsilon \Delta \phi_a$$

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This is a <u>symmetry</u> if $\delta \mathcal{L} = 0$.

Consider making an arbitrary transformation

$$\phi_a \rightarrow \phi_a + \delta\phi_a$$

$$\delta\mathcal{L} = \frac{\partial\mathcal{L}}{\partial\phi_a}\,\delta\phi_a + \frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi_a)}\,\delta(\partial_\mu\phi_a)$$

$$= \int \frac{\partial\mathcal{L}}{\partial\phi_a}$$

Consider making an arbitrary transformation.

$$\phi_a \rightarrow \phi_a + \delta\phi_a$$

$$\delta\mathcal{L} = \frac{\partial\mathcal{L}}{\partial\phi_a}\delta\phi_a + \frac{\partial\mathcal{L}}{\partial(\partial_i\phi_a)}\delta(\partial_i\phi_a)$$

$$= \left[\frac{\partial\mathcal{L}}{\partial\phi_a} - \partial_i\frac{\partial\mathcal{L}}{\partial(\partial_i\phi_a)}\right]\delta\phi_a$$

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$$\phi_a \longrightarrow \phi_a + \delta\phi_a$$

$$\delta \mathcal{L} = \frac{\partial \mathcal{L}}{\partial \phi_a} \delta\phi_a + \frac{\partial \mathcal{L}}{\partial (\partial_r \phi_a)} \delta(\partial_r \phi_a)$$

$$= \left[\frac{\partial \mathcal{L}}{\partial \phi_a} - \partial_r \frac{\partial \mathcal{L}}{\partial(\partial_r \phi_a)} \right] \delta\phi_a + \partial_r \left(\frac{\partial \mathcal{L}}{\partial(\partial_r \phi_a)} \delta\phi_a \right)$$

When the equations of motion hold.

$$\delta \mathcal{L} = \partial_r \left(\frac{\partial \mathcal{L}}{\partial (\partial_r \phi_a)} \, \delta \phi_a \right)$$

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When the equations of motion hold

$$\delta \mathcal{L} = \partial_r \left(\frac{\partial \mathcal{L}}{\partial(\partial_r \phi_a)} \delta \phi_a \right)$$

Now suppose that $\delta \phi_a = \epsilon \Delta \phi_a$, a
symmetry, then $\delta \mathcal{L} = 0$

$$\Rightarrow \partial_r j^\mu = 0 \quad \text{where} \quad \boxed{j^\mu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi_a)} \Delta \phi_a}$$

slight Generalization

We can weaken our requirement of a symmetry to

$$\delta \mathcal{L} = \partial_\mu F^\mu \qquad \text{for some } F^\mu.$$

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$$\Rightarrow \quad j^\mu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi_a)} \Delta\phi_a - F^\mu \qquad \text{is conserved}$$

Examples

1) Space + Time + Energy + Momentum

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Consider $x^\mu \rightarrow x^\mu + \epsilon^\mu$

$\Rightarrow$ $\phi_a(x) \rightarrow \phi_a(x) + \epsilon^\nu \partial_\nu \phi_a(x)$

If the Lagrangian has no explicit x dependence

$$\mathcal{L}(x) \rightarrow \mathcal{L}(x) + \epsilon^\nu \partial_\nu \mathcal{L}$$

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If the Lagrangian has no explicit x dependence

$$\mathcal{L}(x) \rightarrow \mathcal{L}(x) + \epsilon^\nu \partial_\nu \mathcal{L}(x)$$

$\Rightarrow$ we get four conserved currents

Examples

$$(j^\mu)_\nu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi_a)} \partial_\nu \phi_a - \delta^\mu_\nu \mathcal{L}$$

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$$\equiv T^\mu{}_\nu , \text{ is called the}$$

This obeys $\partial_\mu T^\mu{}_\nu = 0$

energy-momentum tensor

Four conserved quantities

$$E = \int d^3x \, T^{00}$$ is total energy of field

$$P^i = \int d^3x \, T^{0i}$$ is total momentum of field

An example:

$$\mathcal{L} = \frac{1}{2}\partial_\mu\phi\,\partial^\mu\phi - \frac{1}{2}m^2\phi^2$$

$$\Rightarrow T^{\mu\nu} = \partial^\mu\phi\,\partial^\nu\phi - \eta^{\mu\nu}\mathcal{L}$$

An example:

$$\mathcal{L} = \frac{1}{2}\partial_\mu\phi\,\partial^\mu\phi - \frac{1}{2}m^2\phi^2$$

$$\Rightarrow T^{\mu\nu} = \partial^\mu\phi\,\partial^\nu\phi - \eta^{\mu\nu}\mathcal{L}$$

$$\Rightarrow E = \int d^3x \;\frac{1}{2}\dot\phi^2 + \frac{1}{2}(\nabla\phi)^2 + \frac{1}{2}m^2\phi^2$$

$$P^i = \int d^3x \;\dot\phi\,(\partial^i\phi)$$