

Title: PSI - Relativity (PHYS 604) - 17

Date: Sep 25, 2009 10:30 AM

URL: <http://pirsa.org/09090076>

Abstract:

Newton

Einstein

vacuum

$$\vec{\nabla}^2 \Phi = 0$$

$$R_{\alpha\beta} = 0$$

$$R = g^{\alpha\beta} R_{\alpha\beta}$$

with
matter

$$\vec{\nabla}^2 \Phi = 4\pi G \rho$$

$$R_{\alpha\beta} + \lambda g_{\alpha\beta} R = \kappa T_{\alpha\beta}$$

$$\lambda = ?$$

$$\kappa = ?$$

↑
Some symmetric
tensor

Newton

Einstein

vacuum

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$$R = g^{\alpha\beta} R_{\alpha\beta}$$

with
matter

$$\vec{\nabla}^2 \Phi = 4\pi G \rho$$

$$R_{\alpha\beta} + \lambda g_{\alpha\beta} R = K T_{\alpha\beta}$$

$$\lambda = ?$$

$$K = ?$$

↑
some symmetric
tensor

"dust" - cloud of non-interacting particles

→ mass: m

→ u -velocity describes flow: $\underline{u}(x)$

"proper" particle density: $N_0(x)$
for observer with $u^*(x)$

"proper" energy/mass density: $\rho = m N_0(x)$

$$T_{\alpha\beta} = \rho u_\alpha u_\beta$$

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in rest frame: $u_\alpha = (-1, 0, 0, 0)$

$$T_{\alpha\beta} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{bmatrix}$$

$$T_{\alpha\beta} = \rho u_\alpha u_\beta$$

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→ combined with Newtonian limit,
this says, $\lambda \neq 0$

$$T_{\alpha\beta} = \rho u_\alpha u_\beta$$

- in analogy to charge conservation
(eg $\partial_\alpha j_\alpha = 0$), we can write
an eq that says energy is
locally conserved:

black hole

$$T_{\alpha\beta} = \rho u_\alpha u_\beta$$

- in analogy to charge conservation
 (eg $\partial_\alpha j^\alpha = 0$), we can write
 an eq that says energy is
 locally conserved:

$$\begin{aligned}\rho' &= \gamma^2 \rho \\ &= \frac{\rho}{1-v^2}\end{aligned}$$

$$T_{\alpha\beta} = \rho u_\alpha u_\beta$$

- in analogy to charge conservation
 (eg $\partial_\alpha j^\alpha = 0$), we can write
 an eq that says energy is
 locally conserved:

recall $\rho' = \gamma^2 \rho$ $\frac{\partial \rho'}{\partial t}$
 $= \frac{\rho}{1-v^2}$

$$T_{\alpha\beta} = \rho u_\alpha u_\beta$$

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recall

$$\rho' = \gamma^2 \rho$$

$$= \frac{\rho}{1-v^2}$$

$$\frac{\partial \rho'}{\partial t} + \vec{\partial} \cdot (\rho' \vec{v})$$

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recall $\rho' = \gamma^2 \rho$
 $= \frac{\rho}{1-v^2/c^2}$

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 an eq that says energy is
 locally conserved:

recall $\rho' = \gamma^2 \rho$
 $= \frac{\rho}{1-v^2/c^2}$

$$\frac{\partial \rho'}{\partial t} + \vec{\partial} \cdot (\rho' \vec{v}) = 0$$

can be expressed in terms of $T_{\mu\nu}$

$$\partial_\mu (\rho u^\mu u^\nu)$$

↓

can be expressed in terms of T_{AB}

$$\partial_\mu (\rho u^\mu u^\nu) = 0$$

can be expressed in terms of $T_{\alpha\beta}$

$$\partial_\alpha (\rho u^\alpha u^\alpha) = 0 = \partial_\alpha T^{\alpha\alpha}$$

in flat space (or local inertial frame)

can be expressed in terms of $T_{\alpha\beta}$

$$\partial_\alpha (\rho u^\alpha u^\alpha) = 0 = \partial_\alpha T^{\alpha\alpha}$$

in flat space (or local inertial frame)

→ generalises to curved space

$$\nabla_\alpha T^{\alpha\alpha} = 0$$

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in flat space (or local inertial frame)

→ generalises to curved space

$$\nabla_\alpha T^{\alpha\alpha} = 0$$

→ in fact there is a vector's worth
of conservation eq's

$$\nabla_\alpha T^{\beta\alpha} = 0$$

$$T_{\alpha\beta} = \rho u_\alpha u_\beta$$

→ spatial components

$$\partial_\alpha (\rho u^i u^\alpha) = 0$$

$$\Rightarrow \frac{\partial}{\partial t} (\rho' v^i) + \frac{\partial}{\partial x_j} (\rho' v^i v^j) = 0$$

$$\rho' = \gamma^2 \rho$$

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$$T_{\alpha\beta} = \rho u_\alpha u_\beta$$

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$$\frac{\partial \rho'}{\partial t}$$

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$$= \frac{\rho}{1-v^2}$$

Navier-Stokes
eq

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recall $\rho' = \gamma^2 \rho$

$$= \frac{\rho}{1-v^2}$$

Navier-Stokes
eq

$$T^{\alpha\beta} = \rho u^{\alpha} u^{\beta}$$

$$T^{\alpha\beta} = \rho u^\alpha u^\beta = \begin{bmatrix} \rho & \rho v^i \\ \rho v^i & \rho v^i v^j \end{bmatrix}$$

$$T^{++} =$$

$$T^{\alpha\beta} = \rho u^\alpha u^\beta = \begin{bmatrix} \rho' & \rho' v^i \\ \rho' v^i & \rho' v^i v^j \end{bmatrix}$$

$$T^{++} = \text{energy density}$$

$$T^{+i} = T^{i+} = \text{momentum density}$$

$$T^{\alpha\beta} = \rho u^\alpha u^\beta = \begin{bmatrix} \rho' & \rho' v^i \\ \rho' v^i & \rho' v^i v^j \end{bmatrix}$$

$$T^{++} = \text{energy density}$$

$$T^{+i} = T^{i+} = \text{momentum density} = \text{flux of energy density}$$

$$0 \quad T^{ij} = T^{ji} =$$

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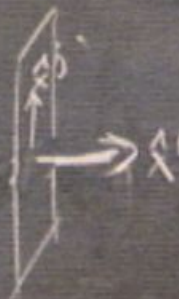
$$T^{++} = \text{energy density}$$

$$T^{+i} = T^{i+} = \text{momentum density} = \text{flux of energy density}$$

$$T^{ij} = T^{ji} = \text{flux of momentum density} \\ = \text{"stress"}$$

$$T^{\alpha\beta} = \rho u^\alpha u^\beta = \begin{bmatrix} \rho' & \rho' v^i \\ \rho' v^i & \rho' v^i v^j \end{bmatrix}$$

Stress



$T^{jk} \sim$ j th comp
of force
on surface
element with
normal $\hat{x}^k$

T^{++} = energy density

$T^{+i} = T^{i+}$ = momentum density = flux of energy density

$T^{ij} = T^{ji}$ = flux of momentum density
= "stress"

$$-x_\beta = \rho u^\alpha u_\beta = \begin{bmatrix} \rho' & \rho' v^i \\ \rho' v^j & \rho' v^i v^j \end{bmatrix}$$

Stress



$T^{jk} \sim j$ th comp
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Stress



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Stress



$T^{jk} \sim$ jth comp. of force on surface element with normal $\hat{n}^k$

$$-T^{++} = \text{energy density}$$

$$T^{+i} = T^{i+} = \text{momentum density} = \text{flux of energy density}$$

$$T^{ij} = T^{ji} = \text{flux of momentum density} = \text{"stress"}$$

$T_{\alpha\beta}$ = stress-energy
tensor

or
energy-momentum
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$$\nabla_\alpha T^{\alpha\beta} = 0$$

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← applies in
general

$T_{\alpha\beta}$ = stress-energy
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← applies in
general

→ reflects the conservation
of energy + momentum

Why energy conservation?

$$\bar{E} =$$

" why energy conservation?

$$E = \int d^3x T^{++}$$

$$\frac{\partial E}{\partial t} = \int d^3x \frac{\partial T^{++}}{\partial t}$$

" why energy conservation?

$$E = \int d^3x T^{++}$$

$$\frac{\partial E}{\partial t} = \int d^3x \frac{\partial T^{++}}{\partial t}$$

$$= \int d^3x \left(- \frac{\partial}{\partial x^i} T^{i+} \right)$$

$$= \oint d^2x n_i T^{i+}$$

Why energy conservation?

$$E = \int d^3x T^{++}$$

$$\frac{\partial E}{\partial t} = \int d^3x \frac{\partial T^{++}}{\partial t}$$

$$= \int d^3x \left(- \frac{\partial}{\partial x^i} T^{i+} \right)$$

$$= \oint d^2x n_i T^{i+} = 0$$

flux of
energy
escaping
 $\rightarrow \infty$

recall

$$\nabla^{\mu} (R_{\mu\nu} + \lambda g_{\mu\nu} R) = \cancel{K \nabla^{\mu} \lambda} \nabla_{\mu} \lambda$$

recall

$$\nabla^\alpha (R_{\alpha\beta} + \lambda g_{\alpha\beta} R) = \cancel{K \nabla^\alpha \lambda g_{\alpha\beta}}$$

$$\nabla^\alpha R_{\alpha\beta} = -\lambda \nabla_\beta R$$

recall

$$\nabla^\alpha (R_{\alpha\beta} + \lambda g_{\alpha\beta} R) = \cancel{\nabla^\alpha R_{\alpha\beta}} + \lambda \nabla^\alpha R_{\alpha\beta}$$

$$\nabla^\alpha R_{\alpha\beta} = -\lambda \nabla_\beta R$$

→ need this to be true always

recall

$$\nabla^\alpha (R_{\alpha\beta} + \lambda g_{\alpha\beta} R) = \cancel{K \nabla^\alpha T_{\alpha\beta}}$$

$$\nabla^\alpha R_{\alpha\beta} = -\lambda \nabla_\beta R$$

→ need this to be true always

from Bianchi identity, we know

this is satisfied if $\lambda = -\frac{1}{2}$

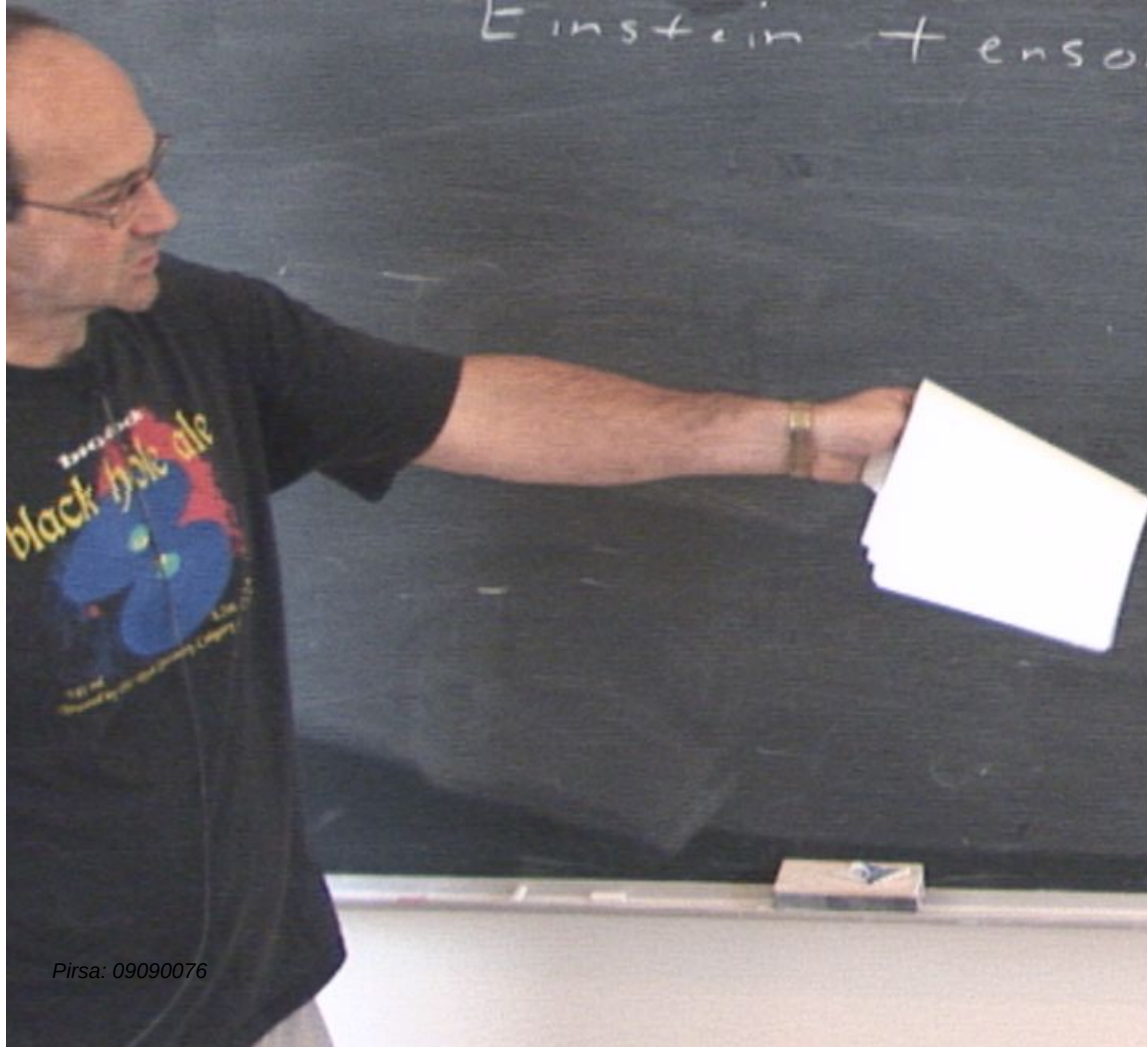
$$R_{\alpha\beta} - \frac{1}{2} g_{\alpha\beta} R = \kappa T_{\alpha\beta}$$



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Einstein tensor, $G_{\alpha\beta}$



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Einstein tensor, $G_{\alpha\beta}$

$$\rightarrow \nabla^\alpha G_{\alpha\beta} = 0$$

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Einstein tensor, $G_{\alpha\beta}$

$$\rightarrow \nabla^\alpha G_{\alpha\beta} = 0$$

to fix κ :

consider Newtonian limit with
the dust stress tensor

$$R_{\alpha\beta} - \frac{1}{2} g_{\alpha\beta} R = \kappa T_{\alpha\beta}$$



Einstein tensor, $G_{\alpha\beta}$

$$\rightarrow \nabla^\alpha G_{\alpha\beta} = 0$$

to fix κ :

- consider Newtonian limit with
dust stress tensor

→ compare to $\nabla^2 \Phi = 4\pi G \rho$

$$K = \frac{8\pi G}{c^4}$$

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Action Principle

→ classical mechanics: $\int dt L(\dot{x}, x)$

→ compare to $\nabla^2 \Phi = 4\pi G \rho$

$$\boxed{K = \frac{8\pi G}{c^4}}$$

Action Principle

→ classical mechanics $\int dt L(\dot{x}^a, x^a)$

→ compare to $\nabla^2 \Phi = 4\pi G \rho$

$$K = \frac{8\pi G}{c^4}$$

Action Principle

→ classical mechanics: $\int dt L(\dot{x}^a, x^a)$

→ fields: $\int L(\dot{\phi}, \partial_i \phi, \phi)$

→ compare to $\nabla^2 \Phi = 4\pi G \rho$

$$K = \frac{8\pi G}{c^4}$$

Action Principle

→ classical mechanics $\int dt L(\dot{x}^a, x^a)$

→ fields $\int dt d^3x L(\dot{\phi}, \partial_i \phi, \phi)$ (eg $\int dt d^3x (-\frac{1}{4} F_{\mu\nu} F^{\mu\nu})$)

→ compare to $\nabla^2 \Phi = 4\pi G \rho$

$$K = \frac{8\pi G}{c^4}$$

Action Principle

→ classical mechanics $\int dt L(\dot{x}^0, x^1)$

→ fields $\int dt d^3x L(\dot{\phi}, \partial_i \phi, \phi)$ (eg $\int dt d^3x (-\frac{1}{4} F_{\mu\nu} F^{\mu\nu})$)

gravity? field $\rightarrow g_{\mu\nu}(x)$

gravity? field $\rightarrow g_{\alpha\beta}(x)$

$$\int d\tau \dots$$

gravity? field $\rightarrow g_{\alpha\beta}(x)$

$\int d^4x \sqrt{-g}$ "scalar" (

gravity? field $\rightarrow g_{\alpha\beta}(x)$

$\int d(\text{"proper volume"})$ "scalar" (

gravity? field $\rightarrow g_{\alpha\beta}(x)$

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consider diagonal metric
de

gravity? field $\rightarrow g_{\alpha\beta}(x)$

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consider diagonal metric

$$ds^2 = \underbrace{g_{tt}}_{\text{negative}} dt^2 + g_{xx} dx^2 + \dots$$



$$dV_{\text{proper}} = \sqrt{-g_{tt} g_{xx} g_{yy} g_{zz}} dt dx dy dz$$

gravity? field $\rightarrow g_{\alpha\beta}(x)$

$$\int d(\text{"proper volume"}) \quad \text{"scalar"} ($$

$$\hookrightarrow \sqrt{-g} d^4x$$

consider diagonal metric $\hookrightarrow \det g_{\alpha\beta}$

$$ds^2 = \underbrace{g_{tt}}_{\text{negative}} dt^2 + g_{xx} dx^2 + \dots$$


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gravity? field $\rightarrow g_{\alpha\beta}(x)$

$$\int d(\text{"proper volume"}) \quad \text{"scalar"}$$

general result

$$\sqrt{-g} d^4x$$

$\leftarrow \det g_{\alpha\beta}$

consider diagonal metric

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gravity? field $\rightarrow g_{\alpha\beta}(x)$

$\int d(\text{"proper volume"})$ "scalar" $(g, \partial g)$

general result

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$\leftarrow \det g_{\alpha\beta}$

consider diagonal metric

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gravity? field $\rightarrow g_{\alpha\beta}(x)$

$$\int d(\text{"proper volume"}) \quad \text{"scalar"}(g, \partial g)$$

general result $\rightarrow \sqrt{-g} d^4x \rightarrow \text{none (constant)}$

consider diagonal metric $\leftarrow \det g_{\alpha\beta}$

$$ds^2 = \underbrace{g_{tt}}_{\text{negative}} dt^2 + g_{xx} dx^2 + \dots$$


$$dV_{\text{proper}} = \sqrt{-g_{tt} g_{xx} g_{yy} g_{zz}} dt dx dy dz$$

gravity? field $\rightarrow g_{\alpha\beta}(x)$

$$\int d(\text{"proper volume"}) \quad \text{"scalar"}(g, \partial g)$$

general result

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$\rightarrow$ none (constant)

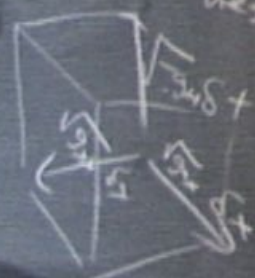
consider diagonal metric

$$\det g_{\alpha\beta}$$

"scalar" $(g, \partial g, \partial^2 g)$

$$ds^2 = g_{tt} dt^2 + g_{xx} dx^2 + \dots$$

negative



$$dV_{\text{proper}} = \sqrt{-g_{tt} g_{xx} g_{yy} g_{zz}} dt dx dy dz$$

metric

gravity? field $\rightarrow g_{\alpha\beta}(x)$

$$\int d(\text{"proper volume"}) \quad \text{"scalar"}(g, \partial g)$$

general result

$$\sqrt{-g} d^4x$$

$\rightarrow$ none (constant)

consider diagonal metric

$$\det g_{\alpha\beta}$$

"scalar" $(g, \partial g, \partial \partial g)$

$$ds^2 = \underbrace{g_{tt}}_{\text{negative}} dt^2 + g_{xx} dx^2 + \dots$$

$\rightarrow$ one

$$R = g^{\alpha\beta} R_{\alpha\beta}$$



$$dV_{\text{proper}} = \sqrt{-g_{tt} g_{xx} g_{yy} g_{zz}} dt dx dy dz$$

metric

variations

$$S \Gamma = \int d^4x \quad \delta g_{\mu\nu} (EOM)^{\mu\nu}$$

instead vary inverse metric

$$- S \Gamma = \int d^4x \quad \delta g^{\mu\nu} (\widetilde{EOM})_{\mu\nu}$$

variations

$$S \mathcal{I} = \int d^4x \quad \delta g_{\mu\nu} (EOM)^{\mu\nu}$$

instead vary inverse metric

$$- S \mathcal{I} = \int d^4x \quad \delta g^{\mu\nu} (\widetilde{EOM})_{\mu\nu}$$

Note

$$g^{\mu\nu} g_{\nu\alpha} = \delta^\mu_\alpha$$

$$\delta g^{\mu\nu} g_{\nu\alpha} + g^{\mu\nu} \delta g_{\nu\alpha} = 0$$

$$\delta g^{\alpha\beta} = - g^{\alpha\mu} g^{\beta\nu} \delta g_{\mu\nu}$$

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$$I = \int d^4x \sqrt{-g} g^{\mu\nu} R_{\mu\nu}$$

$$\delta g^{\alpha\beta} = - g^{\alpha\mu} g^{\beta\nu} \delta g_{\mu\nu}$$

$$I = \int d^4x \sqrt{-g} g^{\mu\nu} R_{\mu\nu}$$

$$\delta I = \delta I_1 + \delta I_2 + \delta I_3$$

$\nearrow$ $\nearrow$ $\nearrow$
 $\sqrt{-g}$ $g^{\mu\nu}$ $R_{\mu\nu}$

$$\delta I_2 = \int d^4x$$

$$\delta g^{\alpha\beta} = - g^{\alpha\mu} g^{\beta\nu} \delta g_{\mu\nu}$$

$$I = \int d^4x \sqrt{-g} g^{\mu\nu} R_{\mu\nu}$$

$$\delta I = \delta I_1 + \delta I_2 + \delta I_3$$

$\nearrow$ $\nearrow$ $\nearrow$
 $\sqrt{-g}$ $g^{\mu\nu}$ $R_{\mu\nu}$

$$\delta I_2 = \int d^4x \sqrt{-g} \delta g^{\mu\nu} R_{\mu\nu}$$

$$\delta Z_1 = ?$$

$$\delta \det(-g_{\mu\nu}) =$$

$$\delta Z_1 = ?$$

$$\begin{aligned} \delta \det(g_{\mu\nu}) &= \delta \left(\frac{1}{\det(g^{\mu\nu})} \right) \\ &= - \frac{1}{\det(g^{\mu\nu} + \delta g^{\mu\nu})} \\ &= - \frac{1}{\det(g^{\mu\nu}) \det(\delta^\mu_\nu + g_{\mu\sigma} \delta g^{\sigma\nu})} \end{aligned}$$

$$\delta Z_1 = ?$$

$$\delta \det(g_{\mu\nu}) = \delta \left(\frac{1}{\det(g^{\mu\nu})} \right)$$

$$= - \frac{1}{\det(g^{\mu\nu} + \delta g^{\mu\nu})}$$

$$= - \frac{1}{\det(g^{\mu\nu}) \det(\delta^\mu_\nu + g_{\mu\sigma} \delta g^{\sigma\nu})}$$

$$\delta Z_1 = ?$$

$$\begin{aligned} \delta g &= g^{\mu\nu} \delta g_{\mu\nu} \\ &= -g^{\mu\nu} \delta g_{\mu\nu} \end{aligned}$$

$$\delta Z_1 = ?$$

$$\begin{aligned}\delta g &= -g \cdot g^{uv} \delta g_{uv} \\ &= -g \cdot g^{uv} \delta g^{uv}\end{aligned}$$

$$\delta Z_1 = ?$$

$$\begin{aligned} \delta g &= g^{\mu\nu} \delta g_{\mu\nu} \\ &= -g^{\mu\nu} \delta g_{\mu\nu} \end{aligned}$$

$$S(-\theta)$$

$$\delta Z_1 = ?$$

$$\delta g = g \cdot g^{uv} \delta g_{uv}$$

$$= -g \cdot g^{uv} \delta g^{uv}$$

$$\delta \sqrt{-g} = +\frac{1}{2} \sqrt{-g} \delta g$$

$$= -\frac{1}{2} \sqrt{-g} g_{uv} \delta g^{uv} \quad (\text{answer})$$

$$\delta \mathcal{I}_1 = ?$$

$$\begin{aligned} \delta g &= g_{\mu\nu} \delta g^{\mu\nu} \\ &= -g_{\mu\nu} \delta g^{\mu\nu} \end{aligned}$$

$$\delta \sqrt{-g} = +\frac{1}{2} \sqrt{-g} \delta g$$

$$= -\frac{1}{2} \sqrt{-g} g_{\mu\nu} \delta g^{\mu\nu}$$

$$\delta \mathcal{I}_1 = \int d^4x \sqrt{-g} \left(-\frac{1}{2} g_{\mu\nu} \delta g^{\mu\nu} \right) R$$

$$S I_3 = 11$$

$$R_{\mu\nu} = R^{\lambda}{}_{\mu\lambda\nu} = \partial_{\mu}\Gamma_{\nu} - \Gamma_{\mu}\Gamma_{\nu}$$

$$R_{\mu\nu}$$

$$\delta \bar{I}_3 = 1$$

$$R_{\mu\nu} = R^{\lambda}_{\mu\lambda\nu} = \partial \Gamma + \Gamma \Gamma$$

$$\delta R_{\mu\nu} = \delta R^{\lambda}_{\mu\lambda\nu} = \partial \delta \Gamma + \Gamma \delta \Gamma$$

$$\delta \Gamma_3 = 11$$

$$R_{\mu\nu} = R^{\lambda}_{\mu\lambda\nu} = \partial \Gamma + \Gamma \Gamma$$

$$\delta R_{\mu\nu} = \delta R^{\lambda}_{\mu\lambda\nu} = \partial \delta \Gamma + \Gamma \delta \Gamma$$

↑
tensor

$$= \nabla_{\lambda} (\delta \Gamma^{\lambda}_{\nu\mu}) - \nabla_{\nu} (\delta \Gamma^{\lambda}_{\lambda\mu})$$

$$\delta \bar{L}_3 = 11$$

$$R_{\mu\nu} = R^{\lambda}_{\mu\lambda\nu} = \partial\Gamma + \Gamma\Gamma$$

$$\delta R_{\mu\nu} = \delta R^{\lambda}_{\mu\lambda\nu} = \partial \delta\Gamma + \Gamma \delta\Gamma$$

↑
tensor

$$g^{\mu\nu} \delta R_{\mu\nu} = \nabla_{\lambda} (\delta\Gamma^{\lambda}_{\mu\nu}) - \nabla_{\nu} (\delta\Gamma^{\lambda}_{\lambda\mu}) - \nabla^{\lambda} (\delta\Gamma^{\mu}_{\lambda\nu}) + \nabla^{\mu} (\delta\Gamma^{\lambda}_{\lambda\nu})$$

$$\delta \bar{L}_3 = 0$$

$$R_{\mu\nu} = R^{\lambda}{}_{\mu\lambda\nu} = \partial^{\lambda} \Gamma_{\mu\lambda\nu} - \partial_{\nu} \Gamma_{\mu\lambda}{}^{\lambda}$$

$$\delta R_{\mu\nu} = \delta R^{\lambda}{}_{\mu\lambda\nu} = \partial^{\lambda} \delta \Gamma_{\mu\lambda\nu} - \partial_{\nu} \delta \Gamma_{\mu\lambda}{}^{\lambda}$$

↑ tensor

$$\begin{aligned} g^{\mu\nu} \delta R_{\mu\nu} &= \nabla_{\lambda} (\delta \Gamma^{\lambda}{}_{\mu\nu}) - \nabla_{\nu} (\delta \Gamma^{\lambda}{}_{\lambda\mu}) \\ &= \nabla^{\lambda} (\delta \Gamma_{\lambda\mu}{}^{\mu}) - \nabla^{\mu} (\delta \Gamma_{\lambda\mu}{}^{\lambda}) \\ &= \nabla^{\mu} (\delta S_{\mu}) \end{aligned}$$

$$\delta \bar{L}_3 = 11$$

$$R_{\mu\nu} = R^{\lambda}{}_{\mu\lambda\nu} = \partial\Gamma + \Gamma\Gamma$$

$$\delta R_{\mu\nu} = \delta R^{\lambda}{}_{\mu\lambda\nu} = \partial \delta\Gamma + \Gamma \delta\Gamma$$

↑
tensor

$$\begin{aligned} g^{\mu\nu} \delta R_{\mu\nu} &= \nabla_{\lambda} (\delta\Gamma^{\lambda}{}_{\nu\mu}) - \nabla_{\nu} (\delta\Gamma^{\lambda}{}_{\lambda\mu}) \\ &= \nabla^{\lambda} (\delta\Gamma^{\lambda}{}_{\mu}) - \nabla^{\mu} (\delta\Gamma^{\lambda}{}_{\lambda}) \\ &= \nabla^{\mu} (\delta S_{\mu}) \end{aligned}$$

$$\delta Z_3 = \int d^4x \sqrt{-g} \nabla^\mu (S w_\mu)$$

$$= \int d^4x \sqrt{-g} n^\mu S w_\mu$$

$$\delta Z_3 = \int d^4x \sqrt{-g} \nabla^\mu (S w_\mu)$$

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↑ use divergence theorem

$$\delta \mathcal{I}_3 = \int d^4x \sqrt{-g} \nabla^\mu (S w_\mu)$$

$$= \int d^4x \sqrt{-g} n^\mu S w_\mu$$

, $\uparrow$ use divergence theorem

$$\delta \mathcal{I} = \delta \mathcal{I}_1 + \delta \mathcal{I}_2 + \cancel{\delta \mathcal{I}_3}^{\text{boundary term}}$$

$$= \int d^4x \sqrt{-g} \delta g^{\mu\nu} (R_{\mu\nu})$$

$$\delta \mathcal{I}_3 = \int d^4x \sqrt{-g} \nabla^\mu (S w_\mu)$$

$$= \int d^4x \sqrt{-g} n^\mu S w_\mu$$

use divergence theorem

$$\delta \mathcal{I} = \delta \mathcal{I}_1 + \delta \mathcal{I}_2 + \cancel{\delta \mathcal{I}_3} \quad \text{boundary term}$$

$$= \int d^4x \sqrt{-g} \delta g^{\mu\nu} \left(R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right)$$

$$\delta I_3 = \int d^4x \sqrt{-g} \nabla^\mu (S w_\mu)$$

$$= \int d^4x \sqrt{-g} n^\mu S w_\mu$$

↑ use divergence theorem

$$S I = S I_1 + S I_2 + \cancel{S I_3}^{\text{boundary term}}$$

$$= \int d^4x \sqrt{-g} \underbrace{\delta g^{\mu\nu} \left(R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right)}_{G_{\mu\nu}}$$

EOM:

$$G_{\mu\nu} = 0$$

in absence of any
matter

$$\begin{aligned} g^{\mu\nu} \delta R_{\mu\nu} &= \nabla_\lambda (\delta \Gamma^\lambda_{\mu\nu}) - \nabla_\nu (\delta \Gamma^\lambda_{\lambda\mu}) \\ &= \nabla^\lambda (\delta \Gamma_{\lambda\mu}{}^\mu) - \nabla^\mu (\delta \Gamma_{\lambda\mu}{}^\lambda) \\ &= \nabla^\mu (\delta w_\mu) \end{aligned}$$

$$\delta I_3 = \int d^4x \sqrt{-g} \nabla^\mu (\delta w_\mu)$$

$$= \int d^4x \sqrt{-g} n^\mu \delta w_\mu$$

use divergence theorem

$$\delta I = \delta I_1 + \delta I_2 + \cancel{\delta I_3}^{\text{boundary term}}$$

$$= \int d^4x \sqrt{-g} \delta g^{\mu\nu} \left(\underbrace{R_{\mu\nu}}_{G_{\mu\nu}} - \frac{1}{2} g_{\mu\nu} R \right)$$

Aside: $(\nabla^\nu \xi^\mu + \nabla^\mu \xi^\nu)$

$$\delta I_3 = \int d^4x \sqrt{-g} \nabla^\mu (\delta w_\mu)$$

$$= \int d^4x \sqrt{-g} n^\mu \delta w_\mu$$

↑ use divergence theorem

$$\delta I = \delta I_1 + \delta I_2 + \cancel{\delta I_3}^{\text{boundary term}}$$

$$= \int d^4x \sqrt{-g} \delta g^{\mu\nu} \underbrace{\left(R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right)}_{G_{\mu\nu}}$$

Aside:

$$(\nabla^\nu \xi^\mu + \nabla^\mu \xi^\nu)$$

$$\text{EOM: } G_{\mu\nu} = 0$$

in absence of any
matter

$$\text{if } \delta g^{\mu\nu} = \nabla^\mu \xi^\nu + \nabla^\nu \xi^\mu \quad \leftarrow \text{infinitesimal coord change}$$

$$\delta I = 0$$

EOM: $G_{\mu\nu} = 0$

in absence of any matter

if $\delta g^{\mu\nu} = \nabla^\mu \xi^\nu + \nabla^\nu \xi^\mu$ ← infinitesimal coord change

$\delta I = 0 \rightarrow \nabla^\mu G_{\mu\nu} = 0$

$$\delta I_3 = \int d^4x \sqrt{-g} \nabla^\mu (\delta w_\mu)$$

$$= \int d^4x \sqrt{-g} n^\mu \delta w_\mu$$

↑ use divergence theorem

$$\delta I = \delta I_1 + \delta I_2 + \cancel{\delta I_3} \quad \text{boundary term}$$

$$= \int d^4x \sqrt{-g} \delta g^{\mu\nu} \left(\underbrace{R_{\mu\nu}}_{G_{\mu\nu}} - \frac{1}{2} g_{\mu\nu} R \right)$$

Aside: $(\nabla^\nu \xi^\mu + \nabla^\mu \xi^\nu)$