

Title: Spacetime can be simultaneously discrete and continuous, in the same way that information can.

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Abstract: TBA

**Spacetime could be simultaneously discrete *and* continuous,
in the same way that information can.**

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Pisa: 09090005



Page 2/27

Spacetime at short distances ?

A gedanken experiment (Heisenberg):

Assume we try to resolve a distance more and more precisely

=> increasing momentum uncertainty

momentum gravitates and thus curves space

=> increasing curvature uncertainty

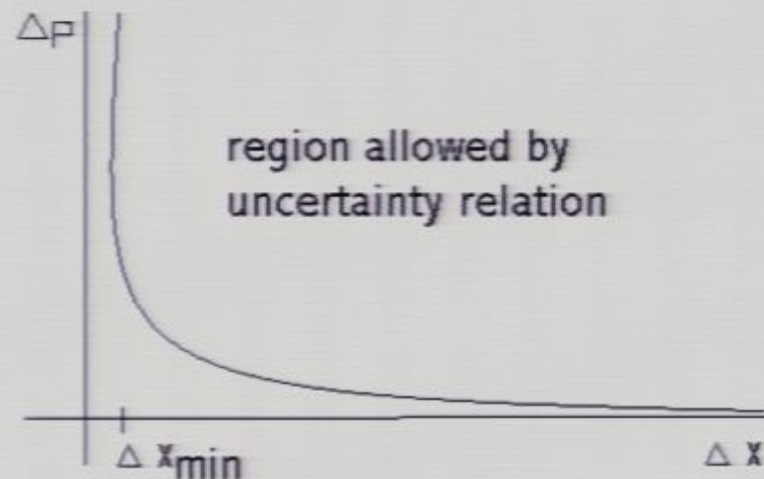
=> increasing distance uncertainties

=> a limit to how precisely distances can be resolved

=> Estimate of “natural ultraviolet cutoff”: Planck length, i.e., 10^{-35}m .

Spacetime at the Planck scale ?

General quantum gravity studies (incl. string theory) suggest e.g.:



There could be an underlying structure, below the UV cutoff scale:

- strings, foams, causal sets, emergent, etc...
- several key approaches are being developed and pursued here at PI.
- so far, **no experimental access** and no consensus.

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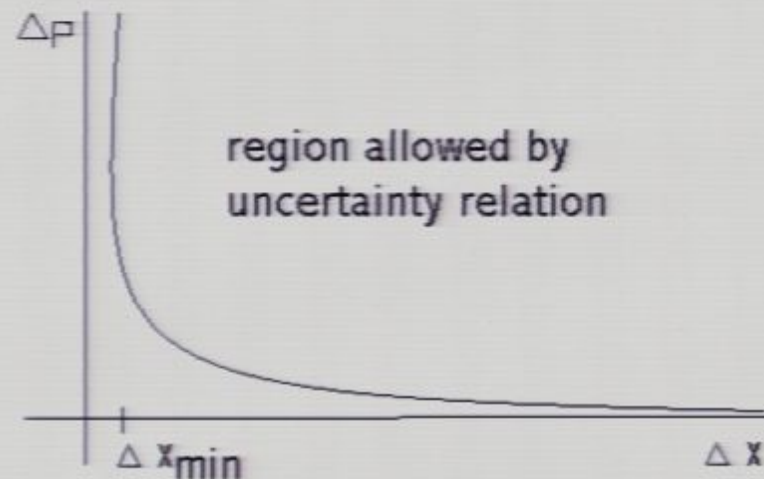
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Spacetime at the Planck scale ?

Problem hard since QFT and GR provide conflicting indications!

General relativity:

- Fields naturally described as living on a differentiable spacetime manifold.

Quantum field theory:

- Fields are quantizable generally only if spacetime described as discrete.

Recall: Information can be both! Let us try this idea:

Spacetime could be both discrete *and* continuous, in the same way that information can.

Discrete vs. continuous in information theory

Information can be:

- continuous (music, speech, etc):



- discrete (letters, digits, etc):

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Unified in 1949 by Shannon, through: Sampling theory

Applications are ubiquitous in all of:

- analog / digital conversion
- communication engineering
- signal processing and analysis
- scientific data taking



The basic Shannon sampling theorem

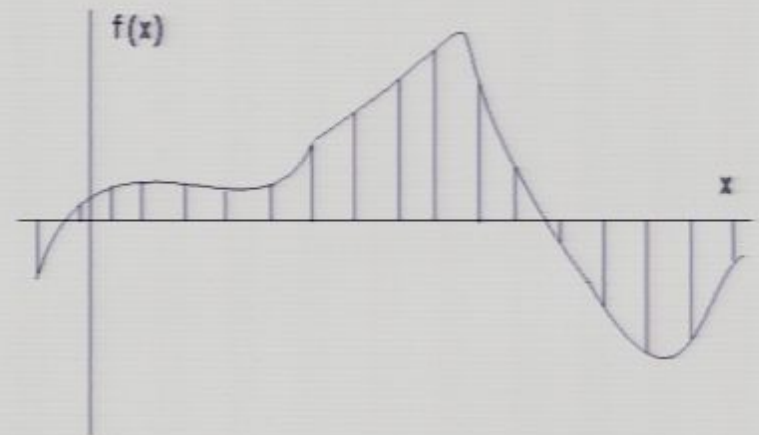
- Assume f is “bandlimited”, i.e:

$$f(x) = \int_{-\omega_{\max}}^{\omega_{\max}} \tilde{f}(\omega) e^{-2\pi i \omega x} d\omega$$

- Take samples of $f(x)$ with spacing:

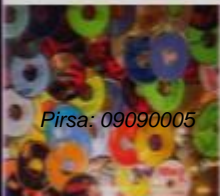
$$x_{n+1} - x_n = (2\omega_{\max})^{-1}$$

- Then, **exact** reconstruction is possible:



samples

$$f(x) = \sum_n f(x_n) \frac{\sin[2\pi(x-x_n)\omega_{\max}]}{\pi(x-x_n)\omega_{\max}}$$



No magic required

Traditional proof:

- uses Fourier theory
- but Fourier theory would not be useful on curved spaces



New method of proof:

1) Notice: Trivial for any N -dimensional function space

- Consider a function space specified by N given basis functions.
 - Any function, f , is determined by its N coefficients in that basis.
 - The values of f at N points yield N equations to determine the N coefficients.
- => we obtain f in a basis, thus we know f completely, i.e., everywhere.

2) For suitable function spaces, can take the limit $N \rightarrow \text{infinity}$.

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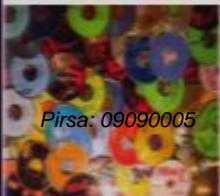
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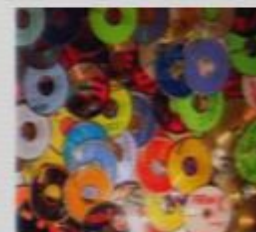
Fun with bandlimited calculus

- Differential operators are also finite difference operators.
- Differential equations are also finite difference equations.
- Integrals are also series:

$$\int_{-\infty}^{\infty} f(x)^* g(x) dx = \frac{1}{2\omega_{\max}} \sum_{n=-\infty}^{\infty} f(x_n)^* g(x_n)$$

Notice:

Useful also as a summation tool for series
(traditionally used, e.g., in analytic number theory)



Sampling theory for physical fields ?

Let us ask: (for now only in the euclidean signature case)

Could there be a natural UV cutoff in QFT such that fields possess a suitable “finite spatial bandwidth” ?

Consequences, if yes:

- **Fields, actions and equations of motion possess equivalent representations:**
 - **on a spacetime manifold**
(shows preservation of external symmetries incl. Killing vector fields)
 - **on any lattice of sufficiently dense spacing**
(showing UV finiteness)



“Bandlimited” physical fields

Generalize “bandlimitation”:

- Cut off spectrum of the Laplacian at Λ , e.g., at the Planck scale.
- Thus, the space of fields, \mathbf{F} , that is summed over in QFT path integral is spanned only by the eigenfunctions with eigenvalues $\lambda_i < \Lambda$.
- Cutoff is covariant since Laplacian is scalar.

Observation: On any finite-volume region of space, \mathbf{F} is finite dimensional.

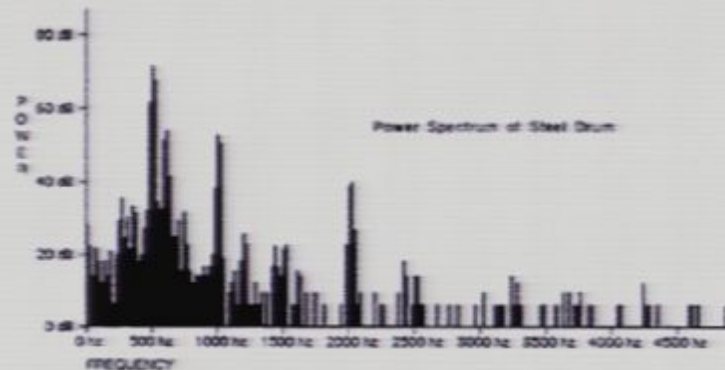
⇒ Thus, we have a sampling theorem for any finite region of space !

Q: Does (# samples/Volume) stay finite as (Volume \rightarrow infinity) ?

A: Yes: Can be shown via Weyl’s asymptotic formula of spectral geometry.

Recall: Spectral Geometry

- M. Kac 1966: Can one hear the shape of a drum ?



- Spectral geometry: Basically yes!

except in a few subtle cases (but with UV cutoff a coarser equivalence will suffice)

Shape of Riemannian manifold \cong Spectrum of Laplacian

Sampling theory of spacetime itself ?

Idea: Nontrivial shape \Leftrightarrow nontrivial distance relations \Leftrightarrow nontrivial 2-point correlators

Method: to sample and reconstruct a piece of a euclidean signature spacetime:
(see [arxiv:0908.3061](https://arxiv.org/abs/0908.3061))

- At **N** points, sample propagator matrix elements, e.g., $\langle x_a | 1/\Delta | x_b \rangle$
- From matrix elements, $\langle x_a | 1/\Delta | x_b \rangle$, calculate lowest eigenvalues of Δ .
- Use spectral geometry to reconstruct the piece of spacetime, up to UV cutoff.
- **A piece of (euclidean signature) spacetime with this UV cutoff**
= an equivalence class of mflds with same spectrum up to the UV cutoff.
(a spacetime is an equivalence class of manifolds differing only by sub-Planckian ripples).

Consistent: Sub-Planckian ripples in space cannot be resolved as fields are bandlimited.



Representation-theoretic view

- **In general relativity:**

- A choice of coordinates is merely a choice of representation of an underlying Riemannian manifold.

- **If there is also this UV cutoff, one has, further:**

- Even the choice of Riemannian manifold is merely a choice of representation of an underlying spacetime with UV cutoff.
(the spacetime is specified by a spectrum up to Λ).
- One can represent a spacetime as any manifold whose Laplacian's spectrum is as specified, up to Λ .



Dynamics of spacetime ?

- Consider a simple action:

$$\begin{aligned} S_{matter} &= \int d^n x \sqrt{|g|} \frac{1}{2} \phi(x) (\Delta + m^2) \phi(x) \\ &= \sum_{i=1}^N \frac{1}{2} \phi_i (\lambda_i + m^2) \phi_i \\ &= \text{Tr} \left(\frac{1}{2} (\Delta + m^2) | \phi \rangle \langle \phi | \right) \end{aligned}$$

- In contains the degrees of freedom of the field and the spacetime, except for N.
- Simplest possible action for N?** $S_{Size} = \alpha N = \alpha \text{Tr}(1)$



Tr(1) = Einstein action

A result of spectral geometry (Gilkey 1975, but anticipated via Sakharov mechanism 1969):

$$N = \frac{1}{16\pi} \int d^4x \sqrt{|g|} \left\{ \frac{\Lambda^2}{2} + \frac{\Lambda}{6} R + O(R^2, \Lambda^{-1}) \right\}$$

=> Einstein action emerges when choosing $\alpha = 6\Lambda$.

(Note: an analogous result for the Dirac operator has been used in NC geometry by Rovelli, Landi, Connes et al)

Here, obtain new interpretation of role of cosmological constant and curvature:

Without curvature, the density of samples (i.e. of degrees of freedom) is:

$$\frac{N}{V} = \frac{\Lambda^2}{32\pi} \quad (\text{i.e., around the Planck density})$$

=> Curvature can be viewed as local perturbation of density of degrees of freedom.



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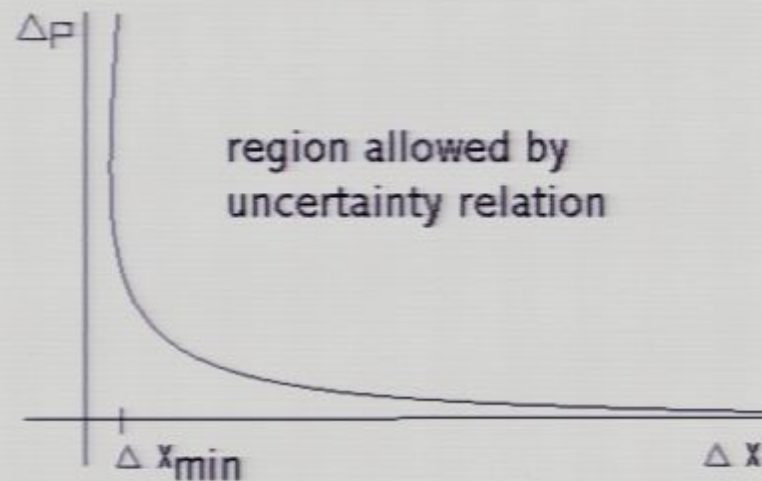
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Any theoretical evidence ?

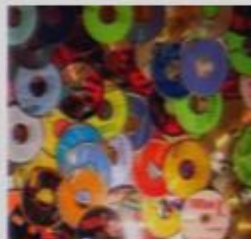
- Recall quantum gravity studies:

=> there may exist a minimum length uncertainty Δx_{\min} :



- Functional analysis shows:

Functions (wave functions or fields etc) in the domain of any operator X with lower uncertainty bound necessarily possess the sampling property.



Experimental evidence?



- **Applications to inflationary cosmology:**

- Sampling theoretic UV cutoff is easily applied to FRW spacetimes.
- Multiple groups have worked out predictions for CMB.

=> **Predicted experimental signatures:**

- Characteristic, $O(10^{-5})$ modulations in CMB power spectrum.

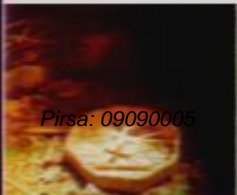
However:

Modulations could have many causes, e.g., suitable inflaton potential.

- Characteristic deviation from scalar/tensor consistency relation in B-polarization data.

However:

B-polarization is being sought but is very hard to measure.



Summary

Spacetime can be discrete & continuous, in same way information can.

- combines mathematical languages of QT and GR
- expresses minimum length uncertainty principle
- **Fields, actions and equations of motion possess equivalent representations**
 - on a spacetime manifold
 - on any lattice of sufficiently dense spacing
- **Describes gravity via eigenvalues of the Laplacian:**
 - representational degrees of freedom are modded out
 - Einstein action is automatically induced

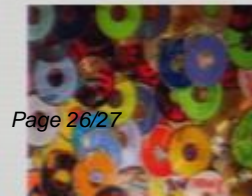
Outlook:

- **develop Lorentz covariant framework further** (see e.g. PRL 2004, new soon).
- **study spacetime as a channel with fundamental quantum noise** (see arxiv:0908.3144)



If not a fundamental property of nature?

- Any new sampling theoretic methods could be useful in communication engineering and signal processing (2 patents)
- Covariant sampling theory yields tools that should be useful in various quantum gravity theories that describe substructure (as in number theory):
 - For any discrete theory can define an equivalent continuum theory
 - Use, e.g., to express finite difference equations as differential equations.
 - Use to define and stabilize dimension of lattices.
 - Replace continuum limit with reconciling discrete and continuum formulations.



Lorentzian signature sampling

- Hard because Lorentzian spectral geometry less developed.
- Results on fully covariant sampling theory exist for:
 - Minkowski space
[AK, Phys.Rev.Lett., 92, 221301 (2004)]
 - The FRW cases of de Sitter space, and power law inflation
[work with R.T. Martin, soon to be posted]
- Key feature of Lorentz covariant sampling:
 - Spatial modes obey temporal sampling theorem, and vice versa
 - This way, Lorentz contraction and time dilatation ensure covariance.