

Title: The Beauty and Basics of BAO

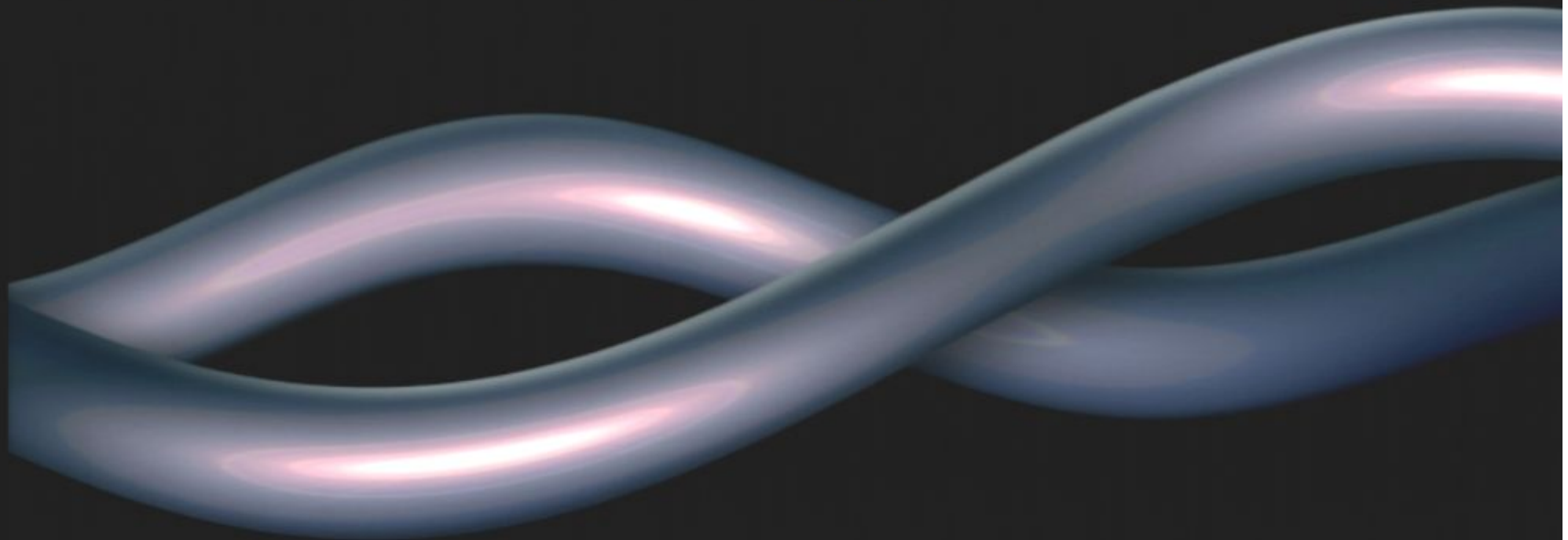
Date: Jul 31, 2009 11:00 AM

URL: <http://pirsa.org/09070039>

Abstract: The Baryon Acoustic Oscillations (BAO) are the latest weapon in the quest for precision cosmology and dark energy. Many presentations on BAO are complicated and unclear and I will therefore present BAO with particular emphasis on trying to give the simplest theoretical description, both at the linear and nonlinear level, and will describe some of the observational challenges to measuring BAO.

# Baryon Acoustic Oscillations

Bruce Bassett (SAAO/UCT)



# Overview

# Overview

- Statistical Standard Rulers



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- Statistical Standard Rulers
- Origins, uses & measurements of BAO

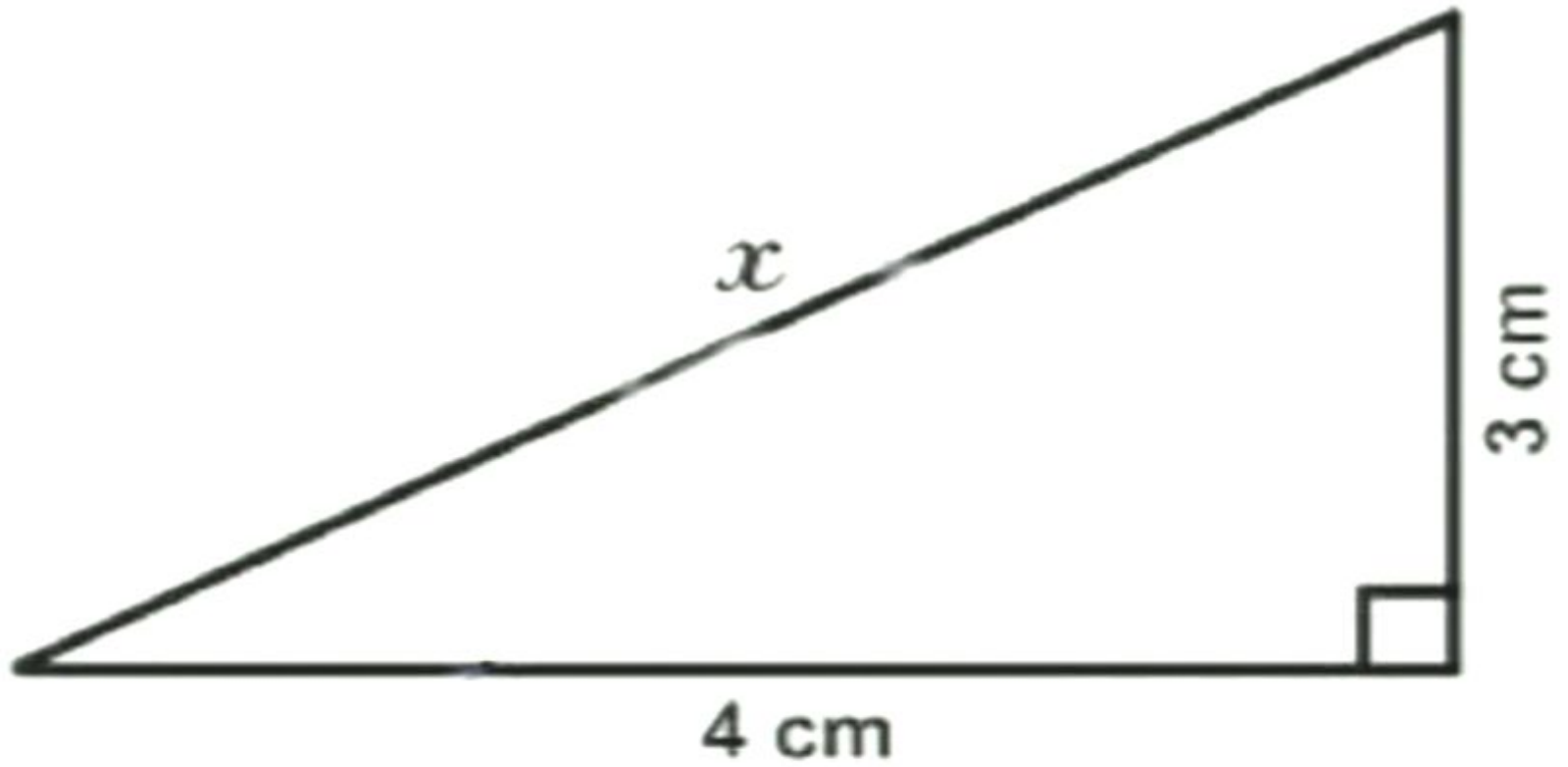
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- Statistical Standard Rulers
- Origins, uses & measurements of BAO
- Targets for BAO surveys

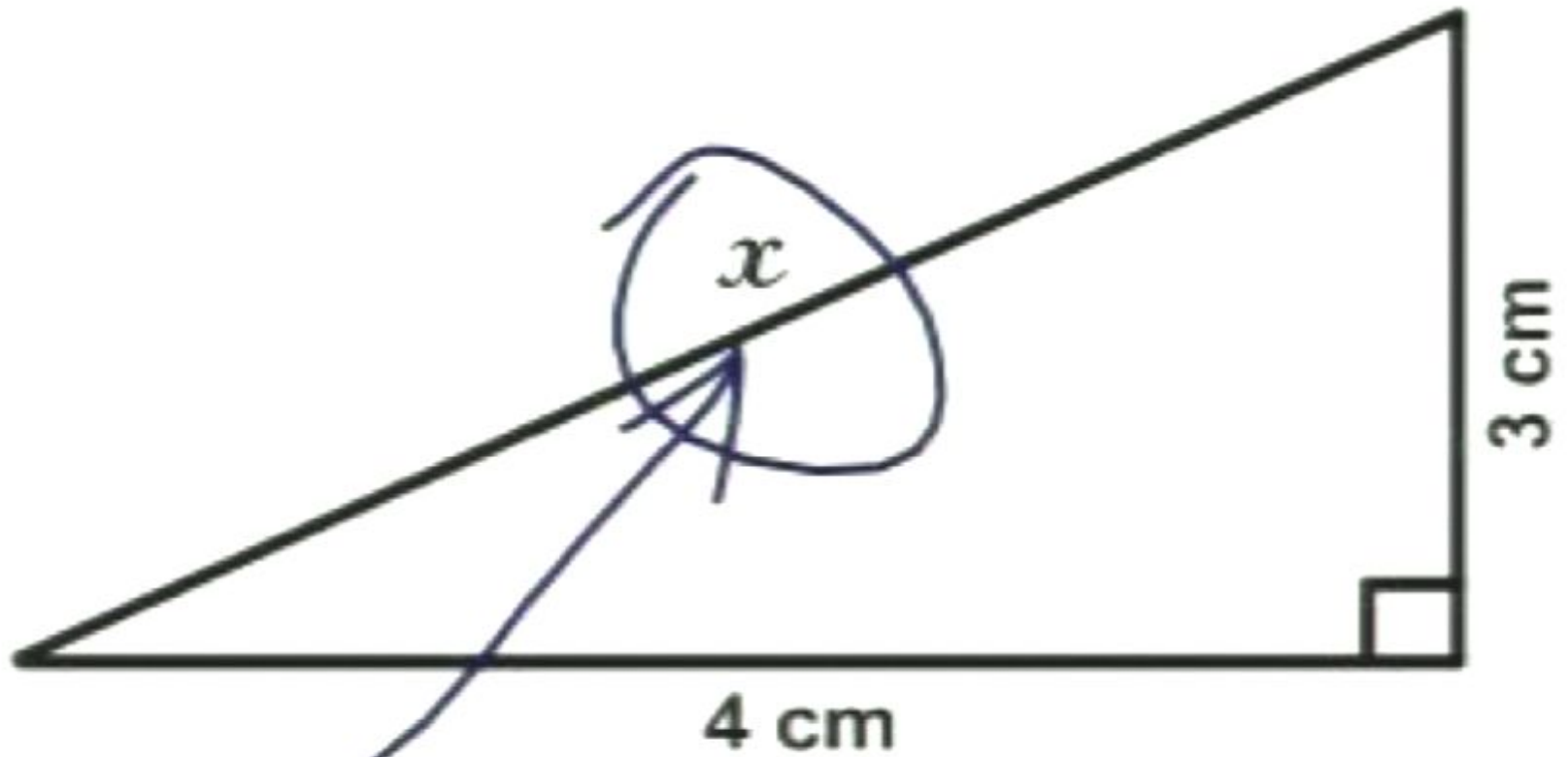
# Overview

- Statistical Standard Rulers
- Origins, uses & measurements of BAO
- Targets for BAO surveys
- Complications
  - Nonlinearities
  - Photometric vs spectroscopic surveys

**3. Find  $x$ .**



3. Find  $x$ .



Here it is

# The Importance of Communication in Science

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## Diabetes Care

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### ARTICLES

## A mathematical model for the determination of total area under glucose tolerance and other metabolic curves

MM Tai

Obesity Research Center, St. Luke's-Roosevelt Hospital Center, New York, New York.



OBJECTIVE--To develop a mathematical model for the determination of total areas under curves from various metabolic studies. RESEARCH DESIGN AND METHODS--In Tai's Model, the total area under a curve is computed by dividing the area under the curve between two designated values on the X-axis (abscissas) into small segments (rectangles and triangles) whose areas can be accurately calculated from their respective geometrical formulas. The total sum of these individual areas thus represents the total area under the curve. Validity of the model is established by comparing total areas obtained from this model to these same areas obtained from graphic method (less than +/- 0.4%). Other formulas widely applied by researchers under- or overestimated total area under a metabolic curve by a great margin. RESULTS--Tai's model proves to be able to 1) determine total area under a curve with precision; 2) calculate area with varied shapes that may or may not intercept on one or both X/Y axes; 3) estimate total area under a curve plotted against varied time intervals (abscissas), whereas other formulas only allow the same time interval; and 4) compare total areas of metabolic curves produced by different studies. CONCLUSIONS--The Tai model allows flexibility in experimental conditions, which means, in the case of the glucose-response curve, samples can be taken with differing time intervals and total area under the curve can still be determined with precision.

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# (Statistical) Standard Rulers



# (Statistical) Standard Rulers



# “Standard” is a Crucial Ingredient

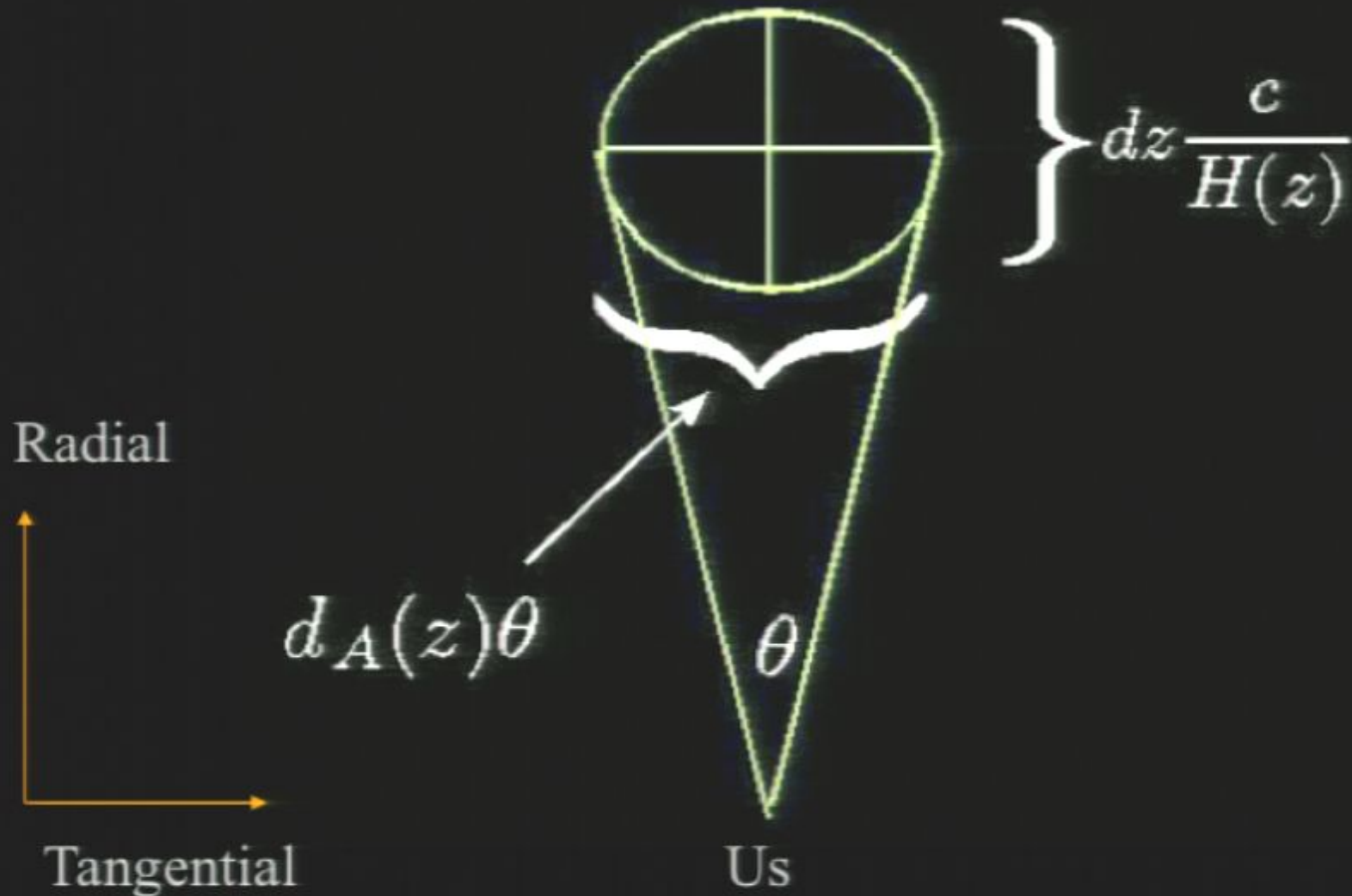
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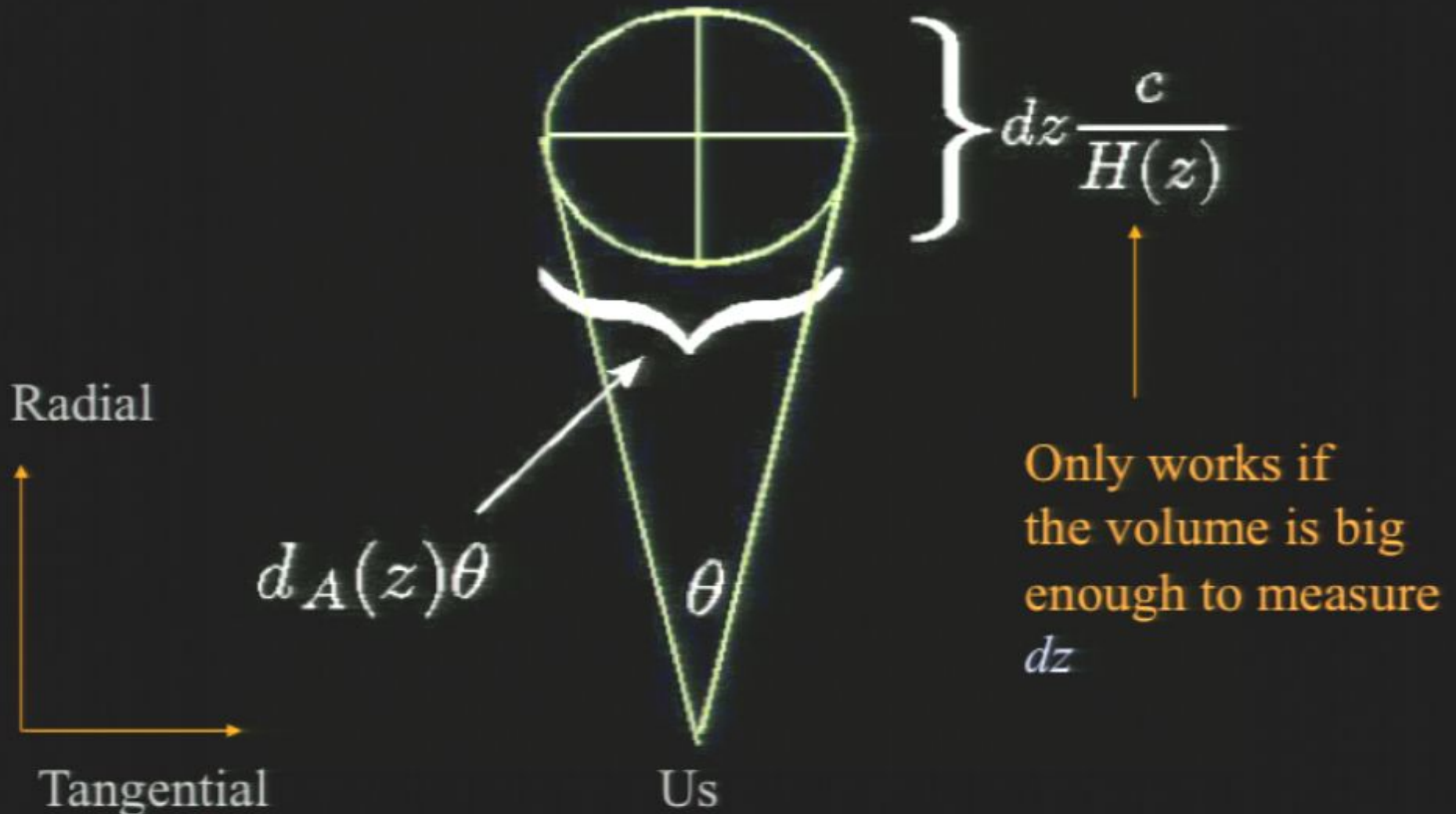


# The Beauty of Standard Volumes



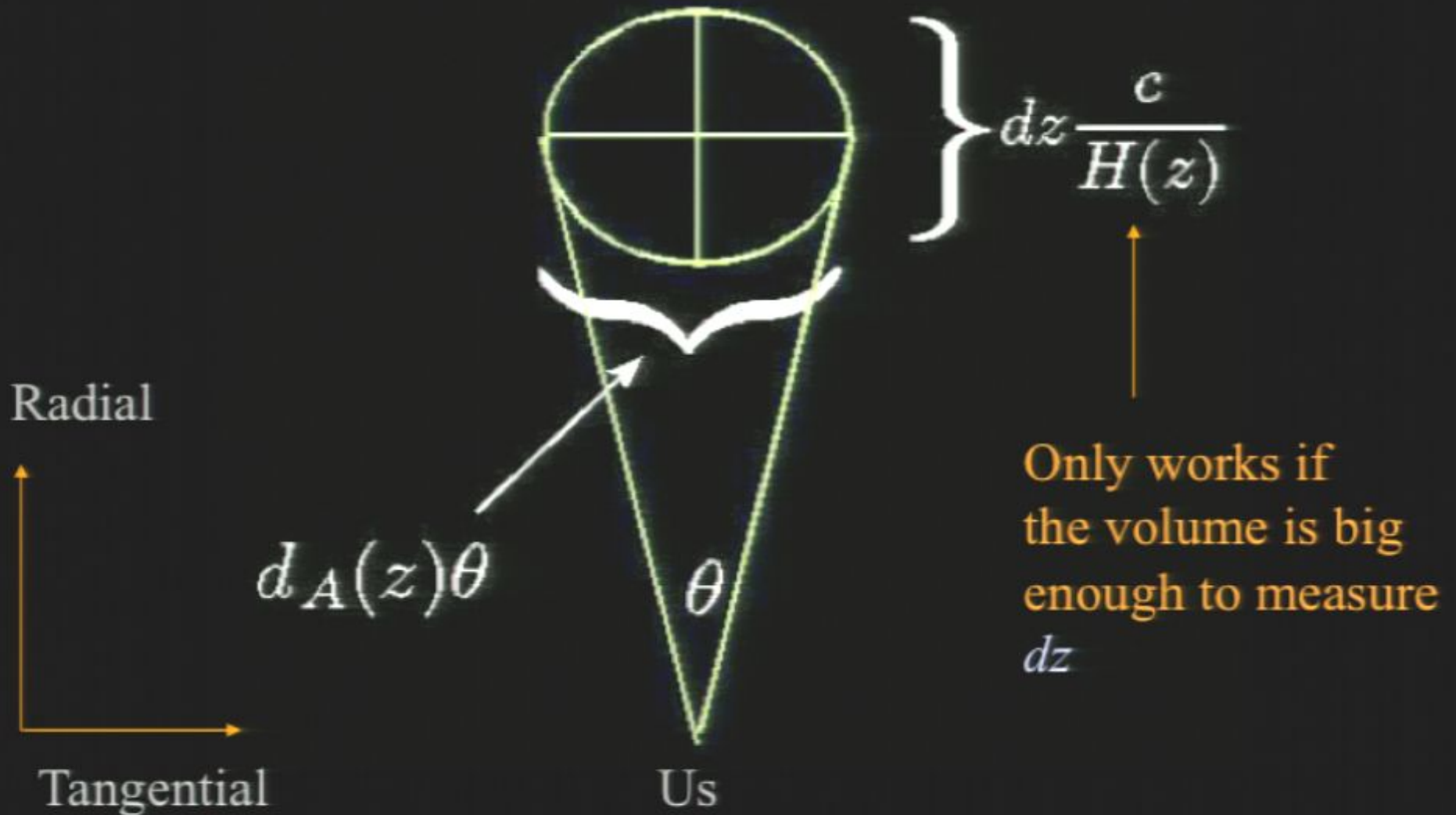


# The Beauty of Standard Volumes



# BAO and AP

# The Beauty of Standard Volumes



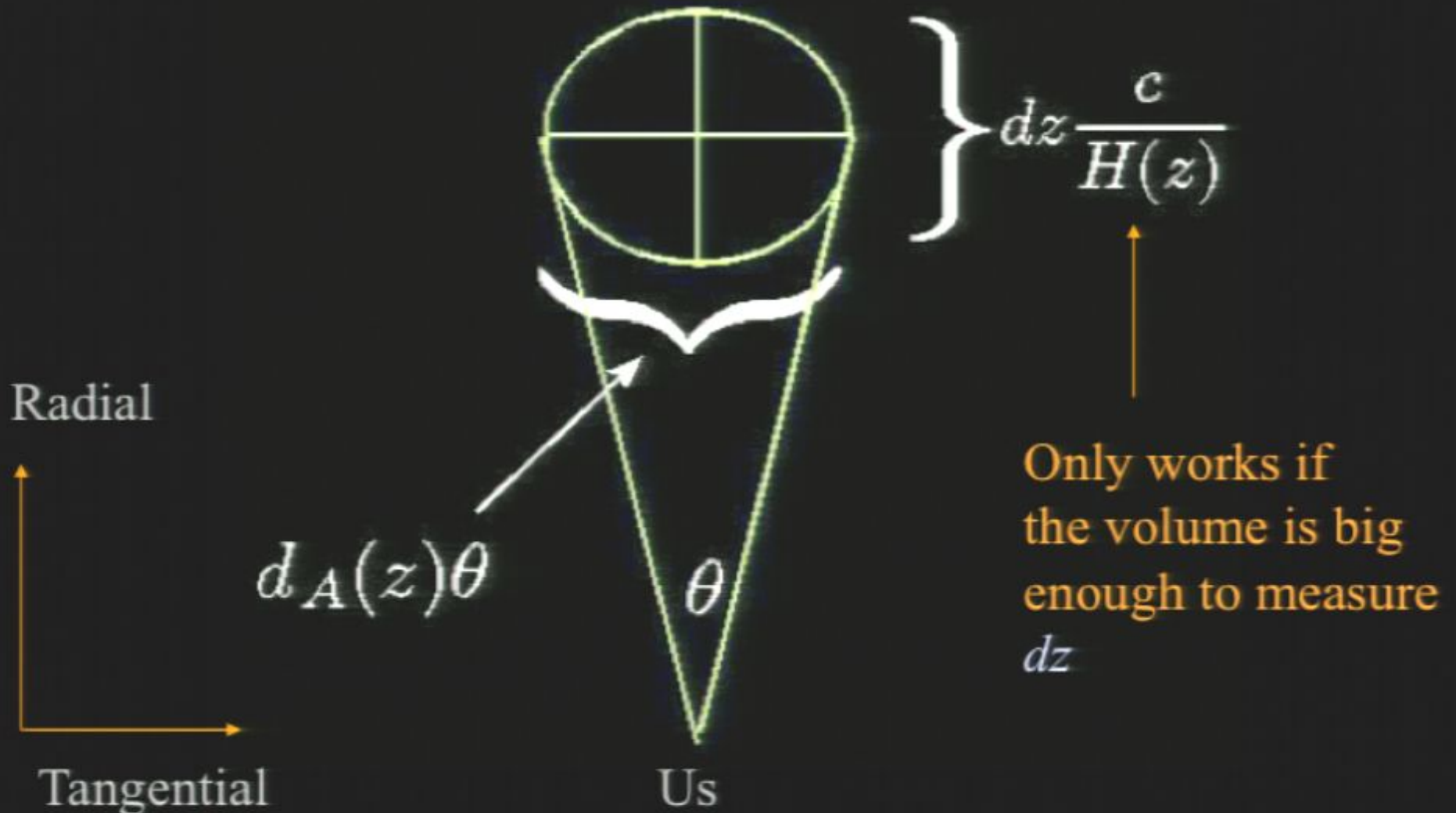
# BAO and AP

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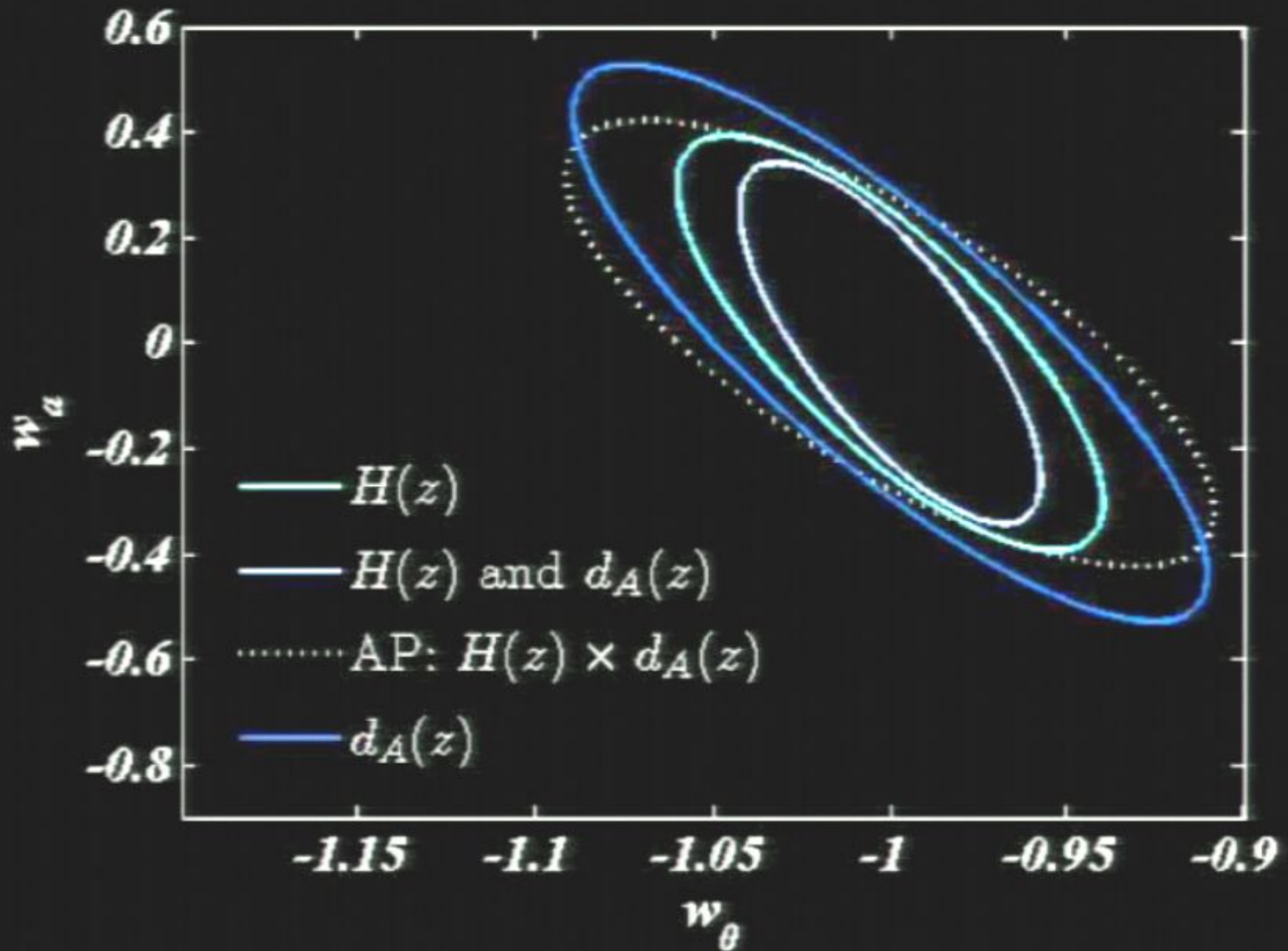
- If we *only* know the volume is spherical, then we constrain the product  $H(z) d_A(z)$  (the Alcock-Paczynski (AP) test)
- If we also know the *absolute* sizes in both radial and tangential directions (like with BAO), we constrain both  $d_A(z)$  and  $H(z)$  separately.

# BAO and AP

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- If we also know the *absolute* sizes in both radial and tangential directions (like with BAO), we constrain both  $d_A(z)$  and  $H(z)$  separately.
- BAO thus provide an *absolute* AP test.

$$d_A(z) = \frac{c}{H_0 \sqrt{-\Omega_k} (1+z)} \sin \left( \sqrt{-\Omega_k} \int_0^z \frac{H_0}{H(z')} dz' \right)$$

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# Testing the Copernican Principle

- Test isotropy of gravitational collapse via comparison of BAO with other probes (e.g. SNIa)
- Compare curvature at different redshifts using

$$\Omega_k = \frac{[H(z)D'(z)]^2 - 1}{[H_0 D(z)]^2}$$

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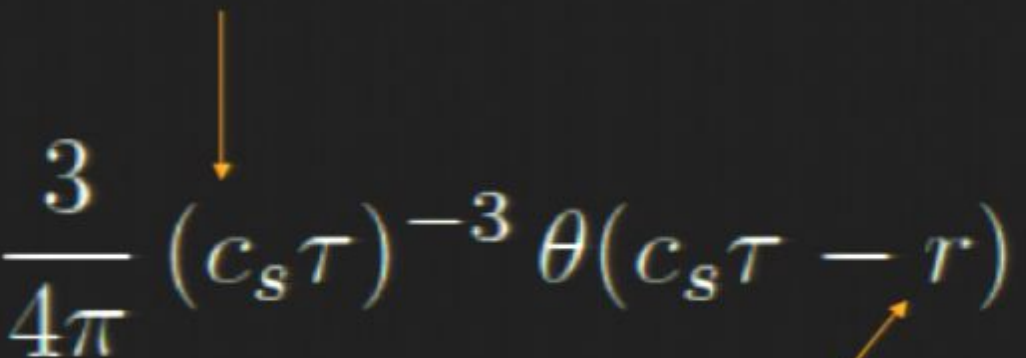
$$\Omega_k = \frac{[H(z)D'(z)]^2 - 1}{[H_0 D(z)]^2}$$



What are BAO and why do we think they are standard rulers?

# 3D Green's Function for $\Phi$

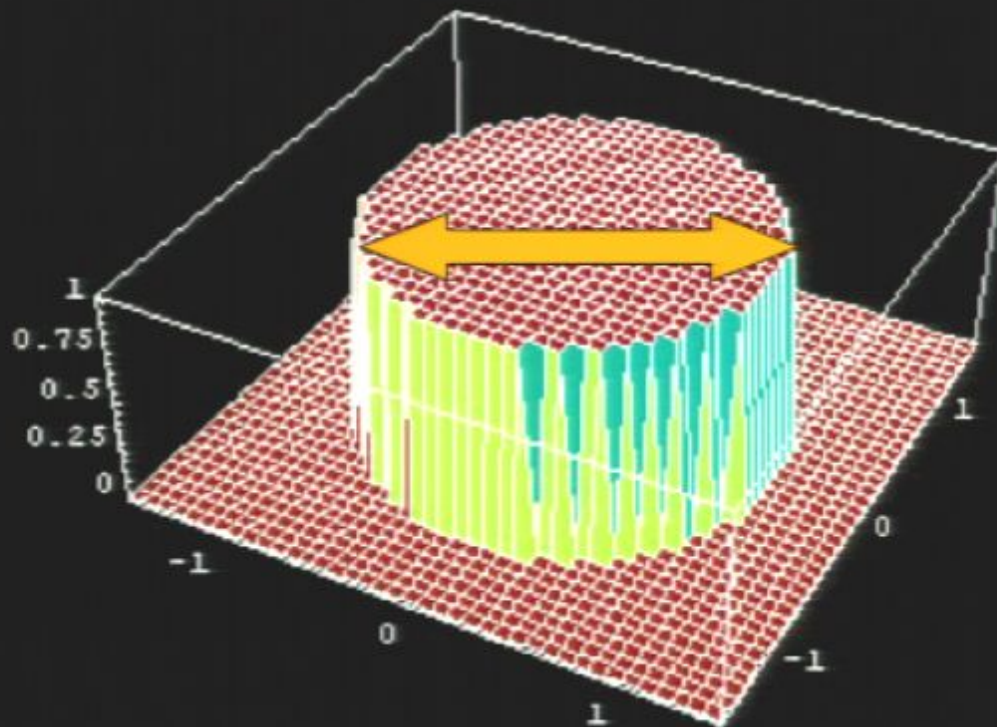
Speed of Sound

$$\phi^{(3)}(r, \tau) = \frac{3}{4\pi} (c_s \tau)^{-3} \theta(c_s \tau - r)$$
The diagram shows the equation  $\phi^{(3)}(r, \tau) = \frac{3}{4\pi} (c_s \tau)^{-3} \theta(c_s \tau - r)$ . An orange arrow points from the text 'Speed of Sound' to the term  $c_s$  in the denominator. Another orange arrow points from the text 'Sound Horizon' to the term  $r$  in the argument of the Heaviside step function  $\theta(c_s \tau - r)$ .

Sound Horizon

*Bashinsky and Bertschinger, 2001*

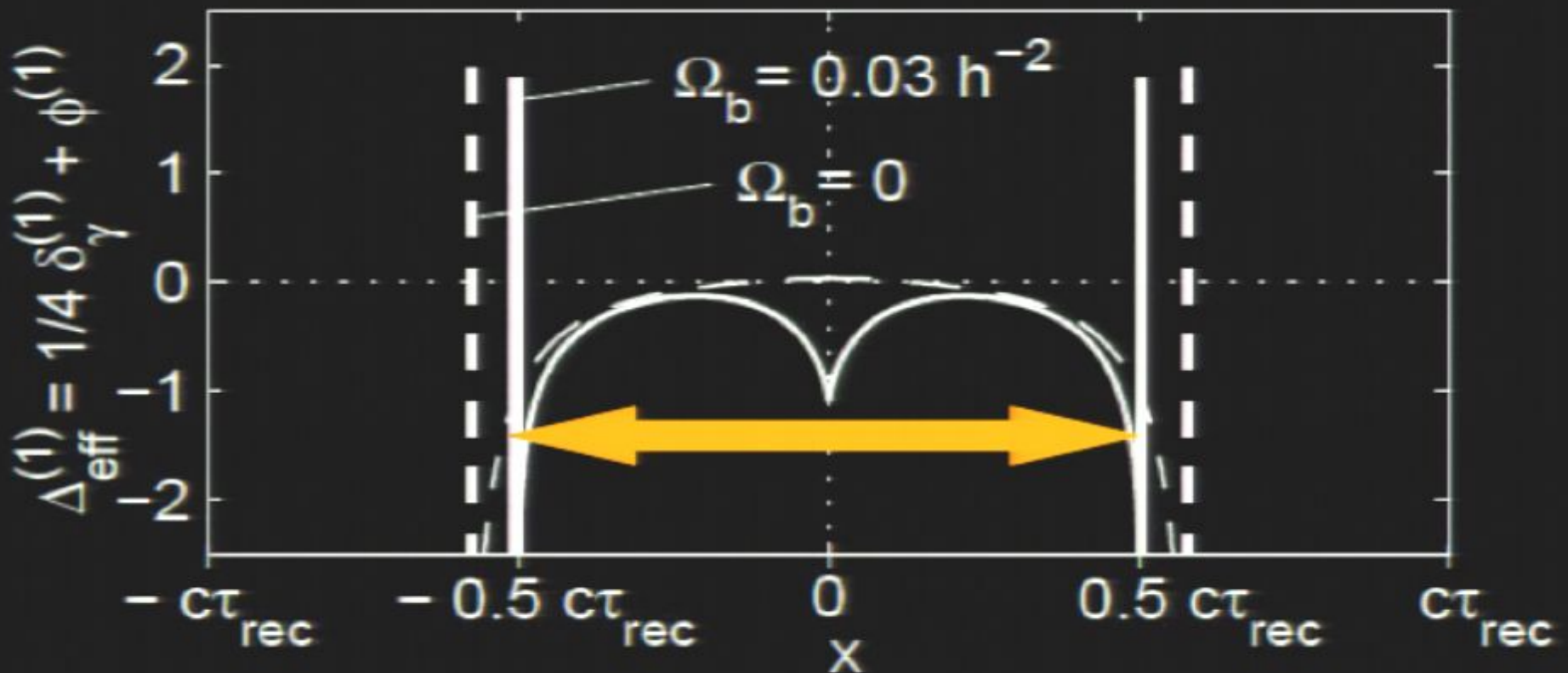
$$\phi^{(3)}(r, \tau) = \frac{3}{4\pi} (c_s \tau)^{-3} \theta(c_s \tau - r)$$



Light travels faster than sound.  
This is why some people appear  
bright until you hear them speak.

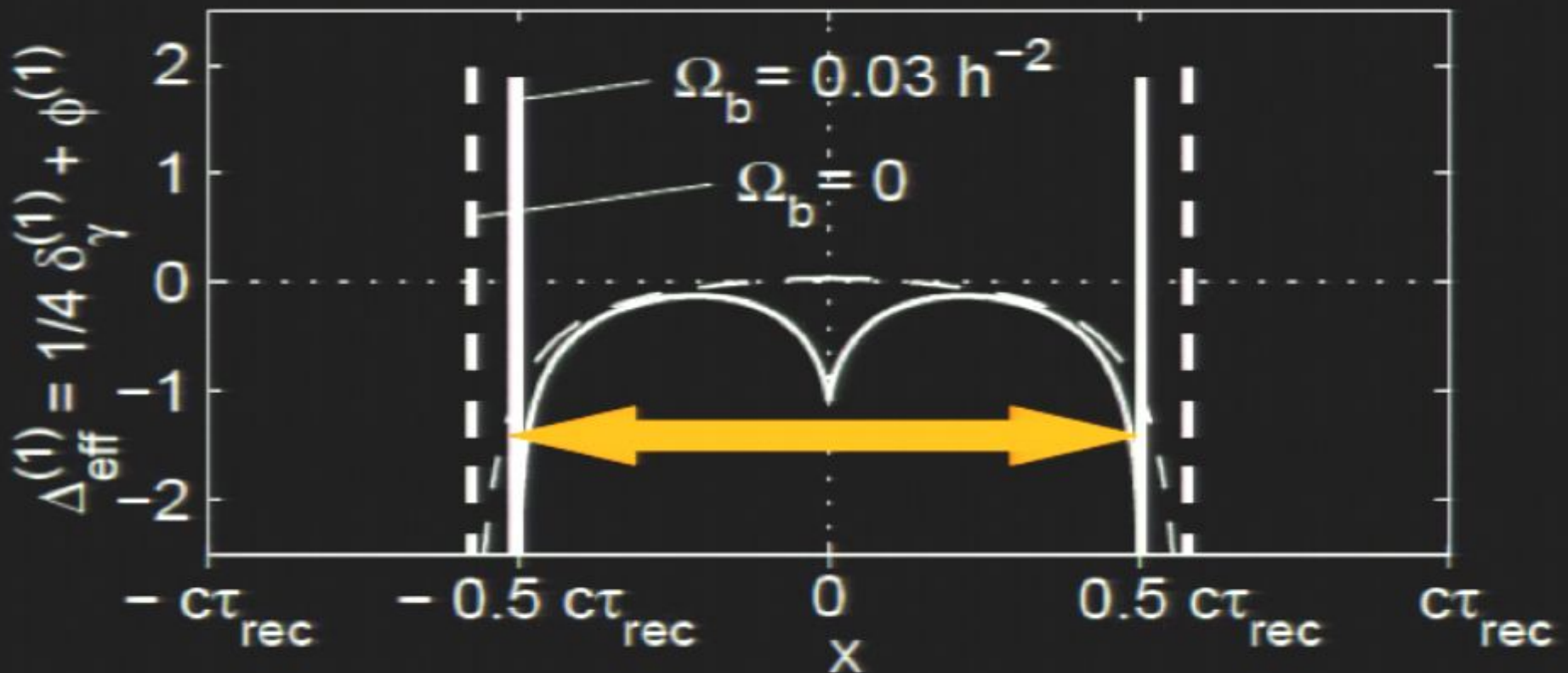
.... only partially true before decoupling

# Real Space Transfer Function





# Real Space Transfer Function

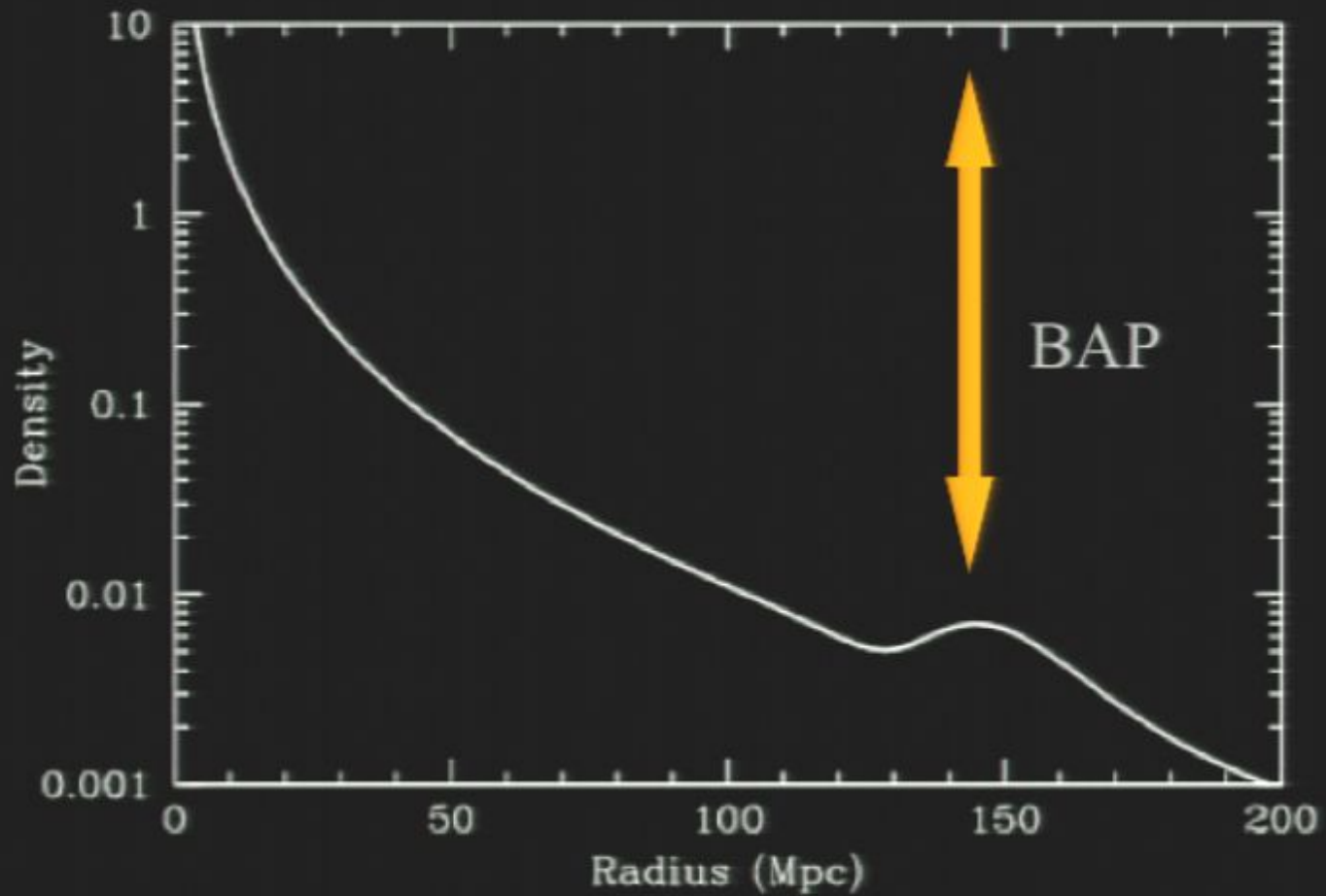
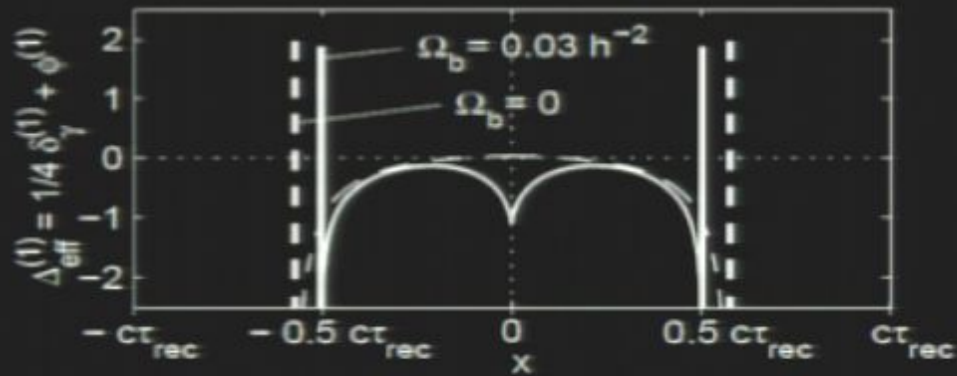


$$c_s^2 = (1/3) / [1 + (3\rho_b)/(4\rho_\gamma)]$$

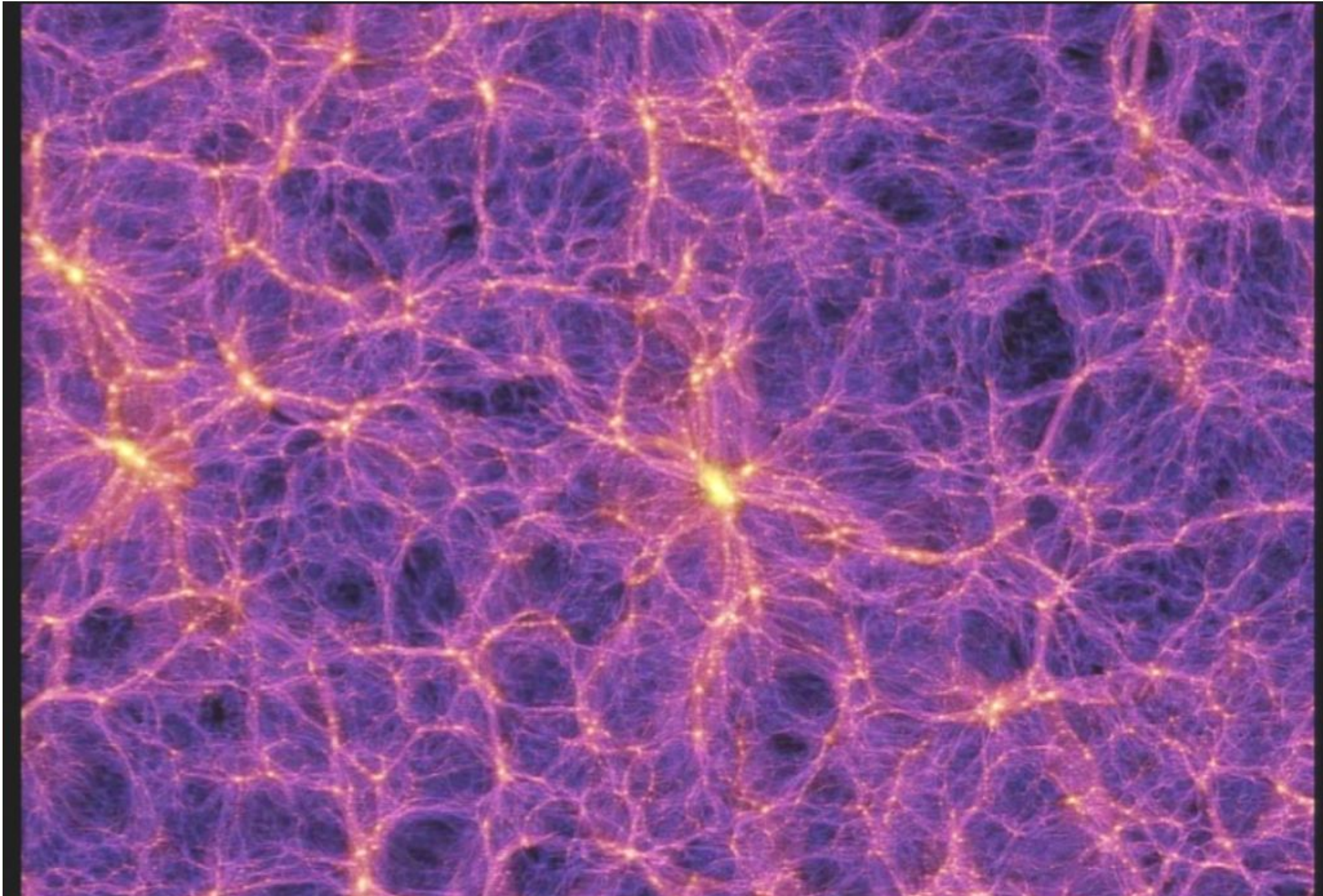
*Bashinsky and Bertschinger, 2*

It is the same standard ruler  
that suggests the cosmos is flat











# 2-point Correlation function

$$\xi(r)$$



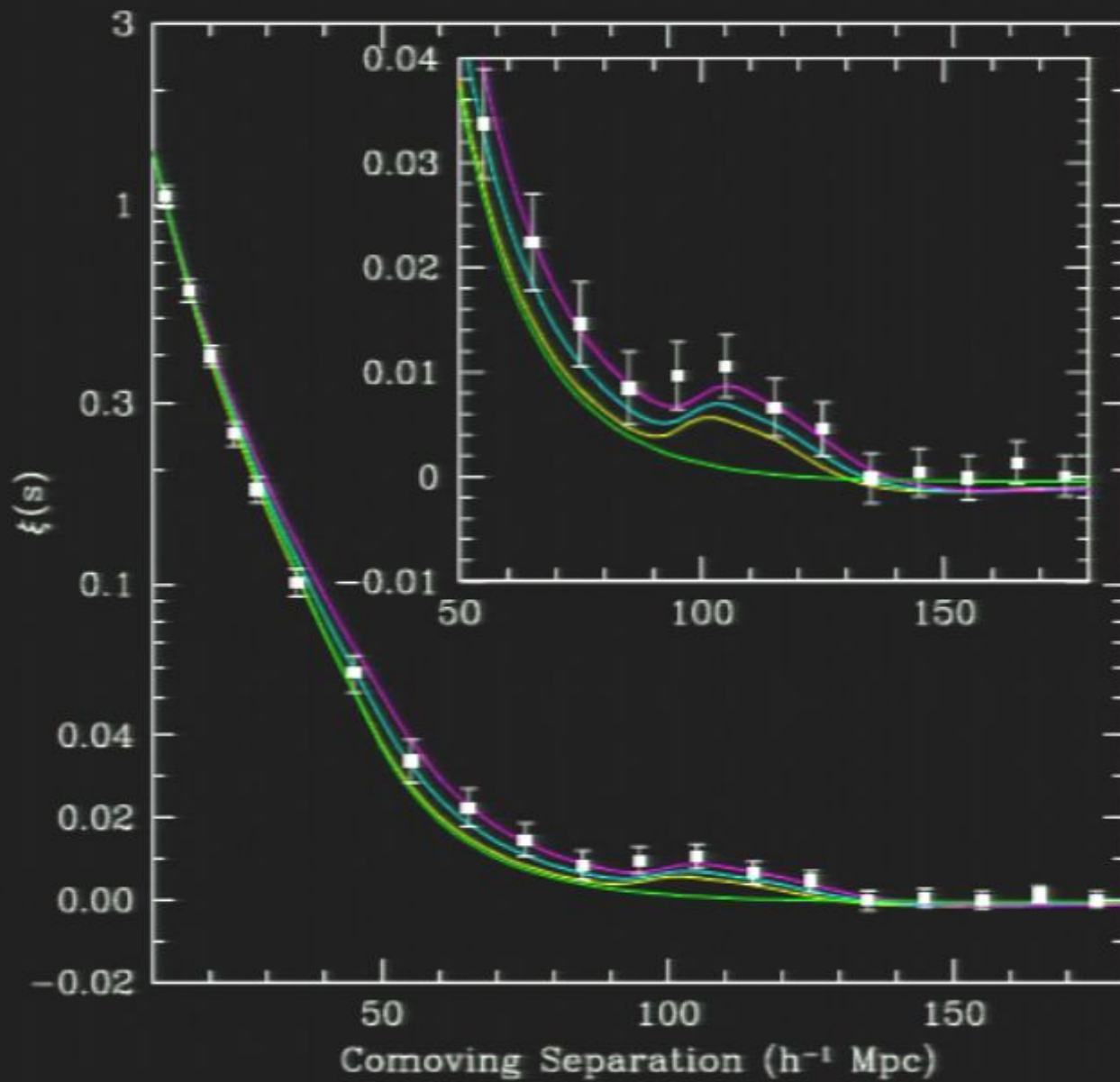


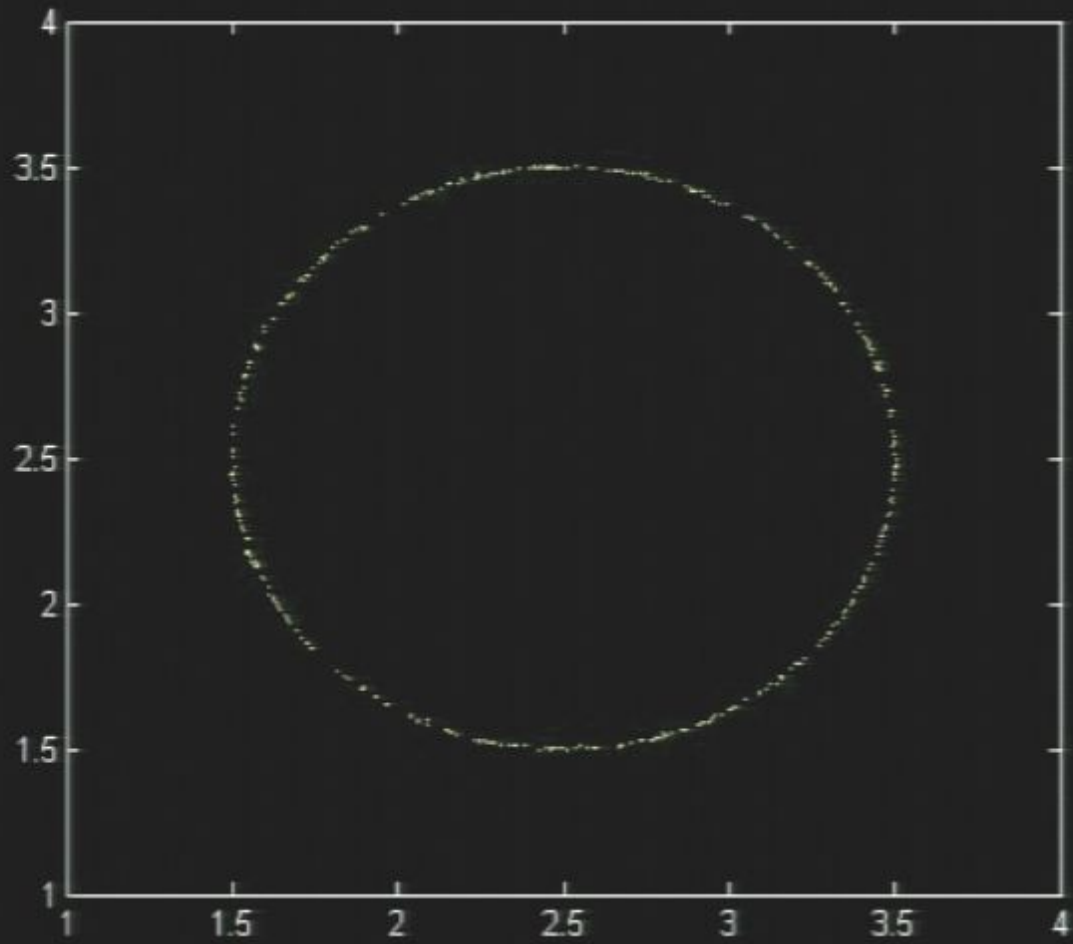
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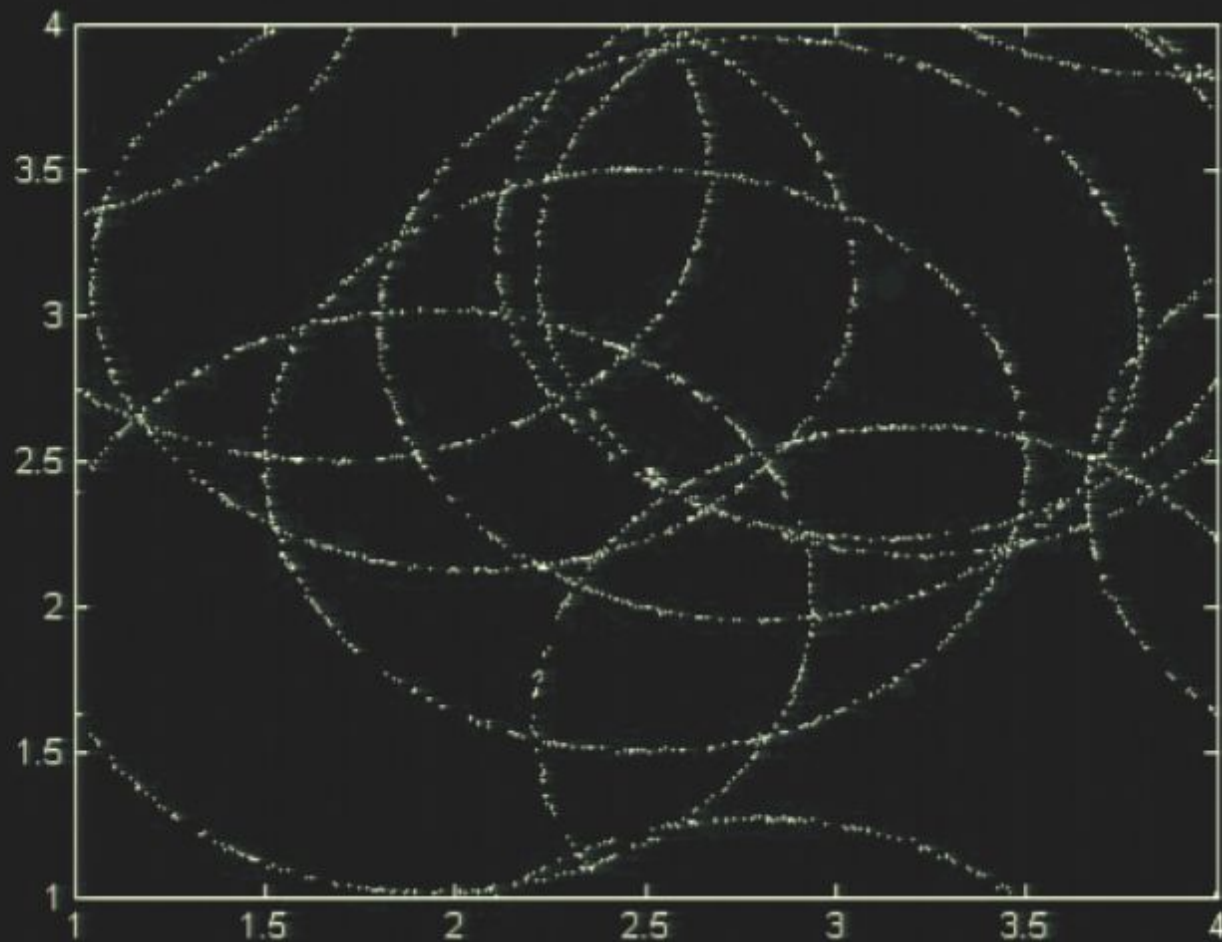
$$\xi(r)$$



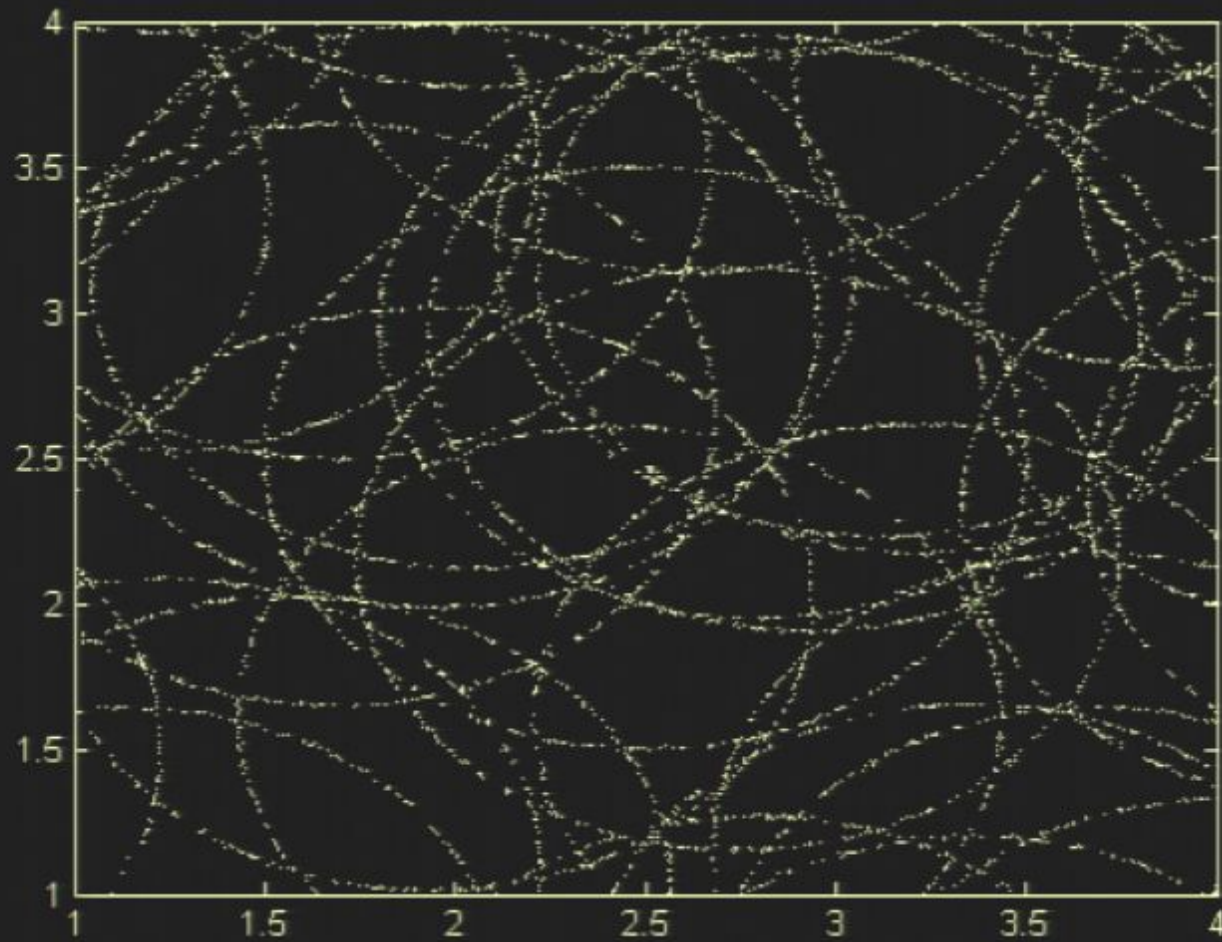
We expect to see an increase  
in  $\xi(r)$  at  $r \sim 100 \text{ Mpc}$

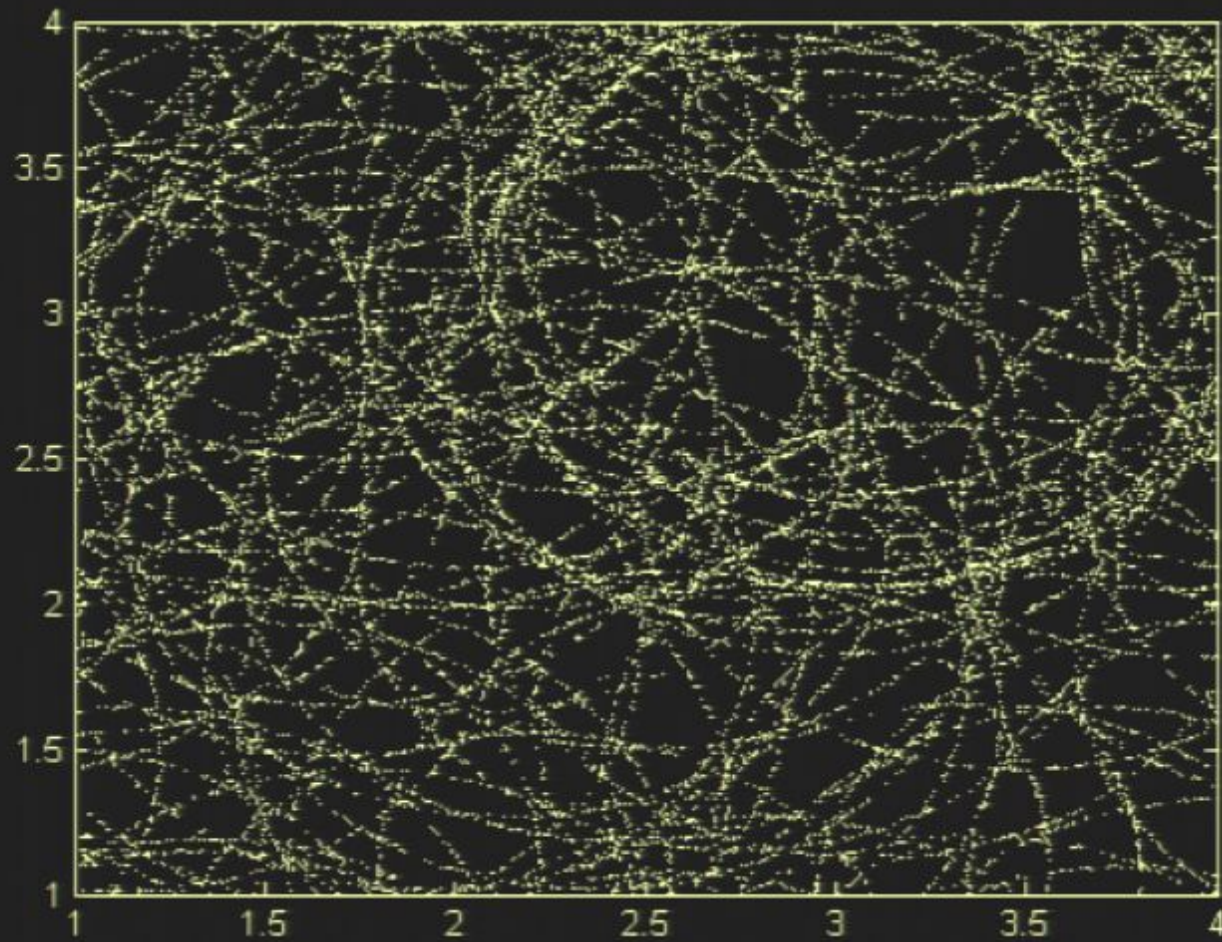




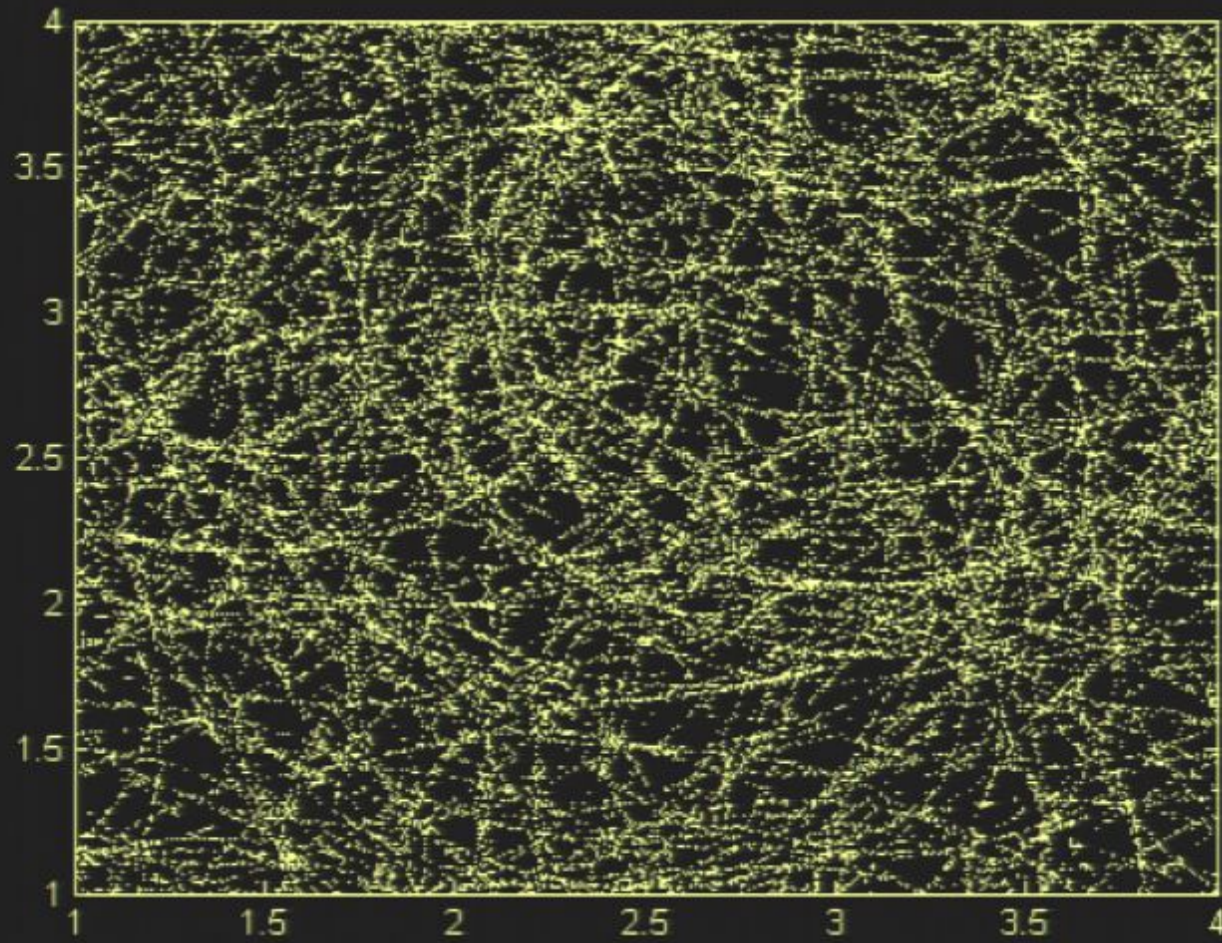




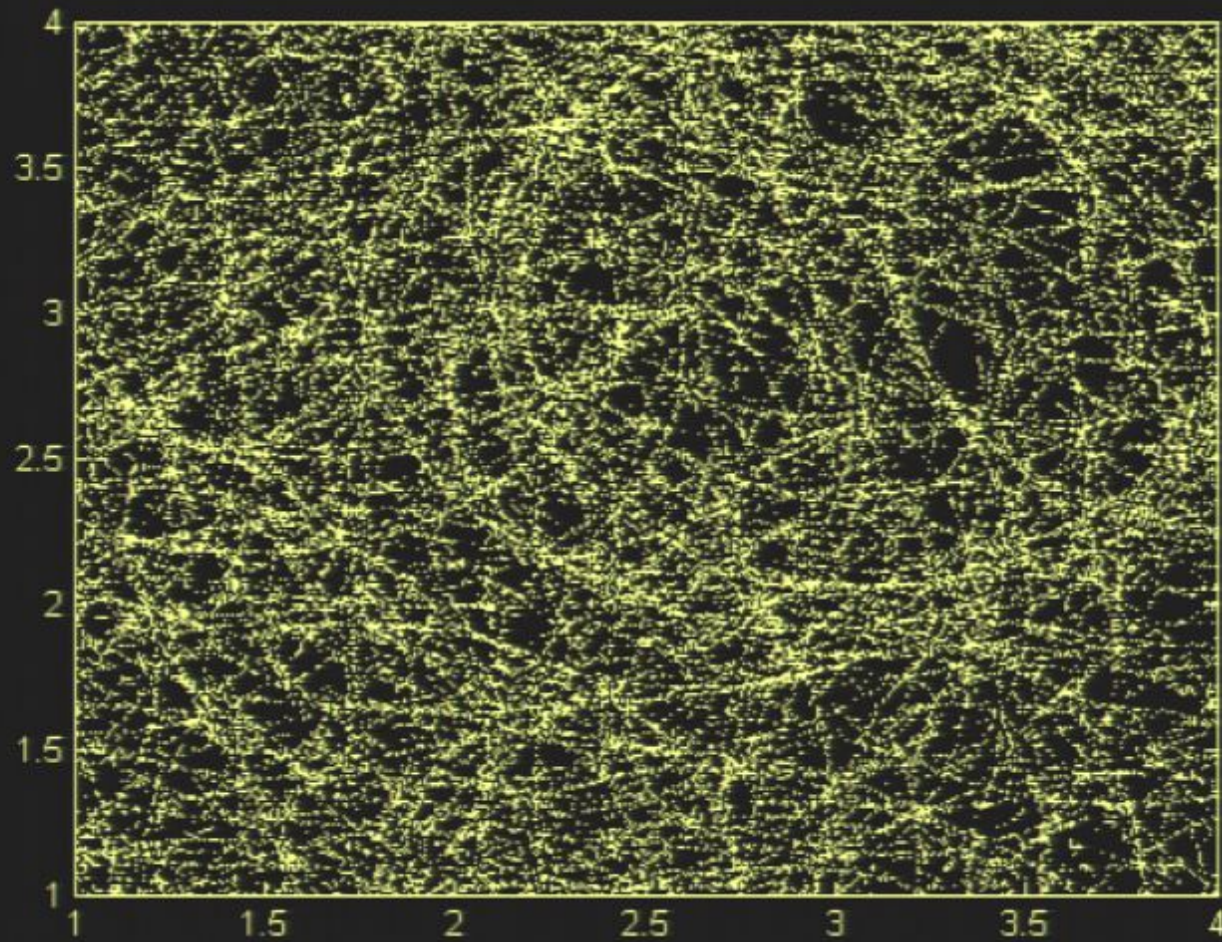




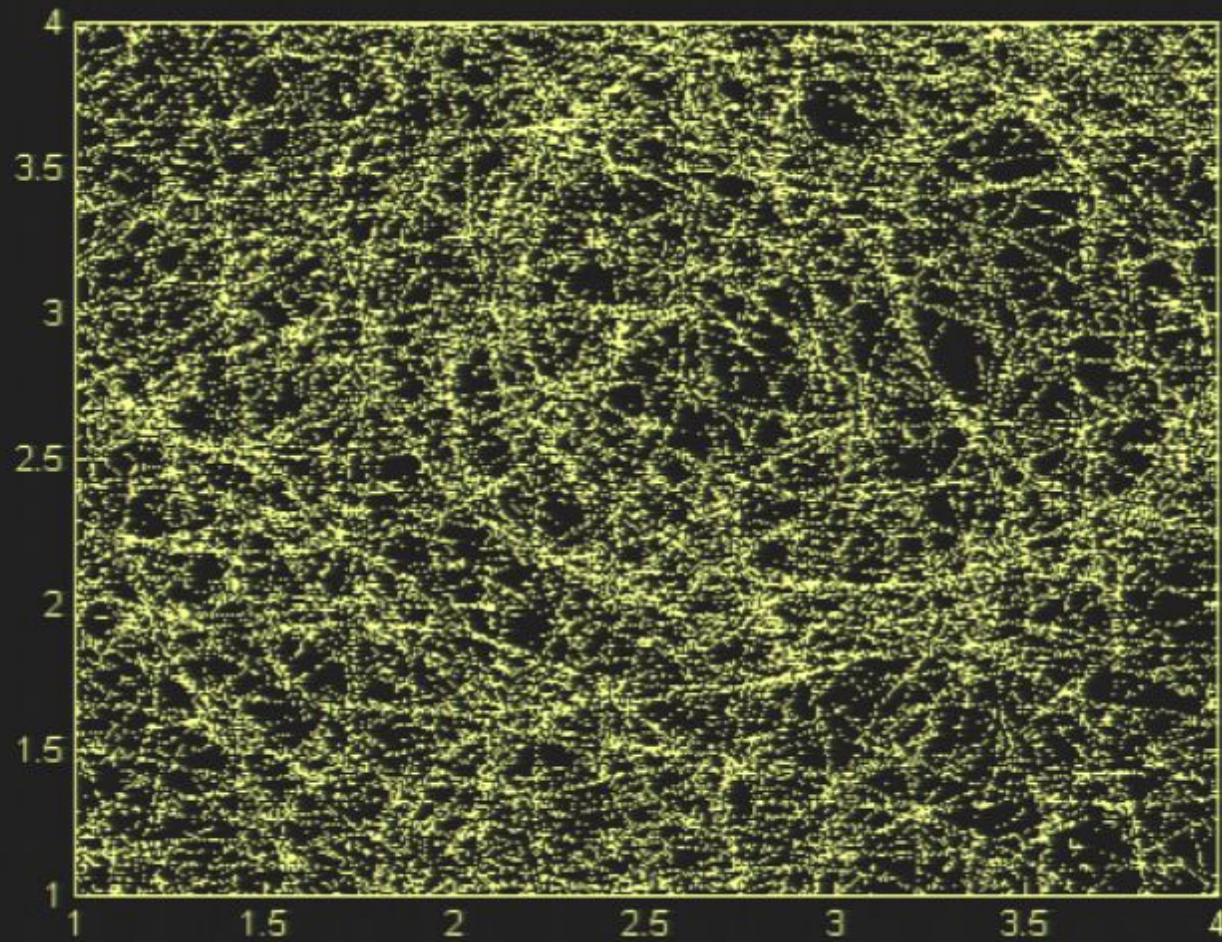






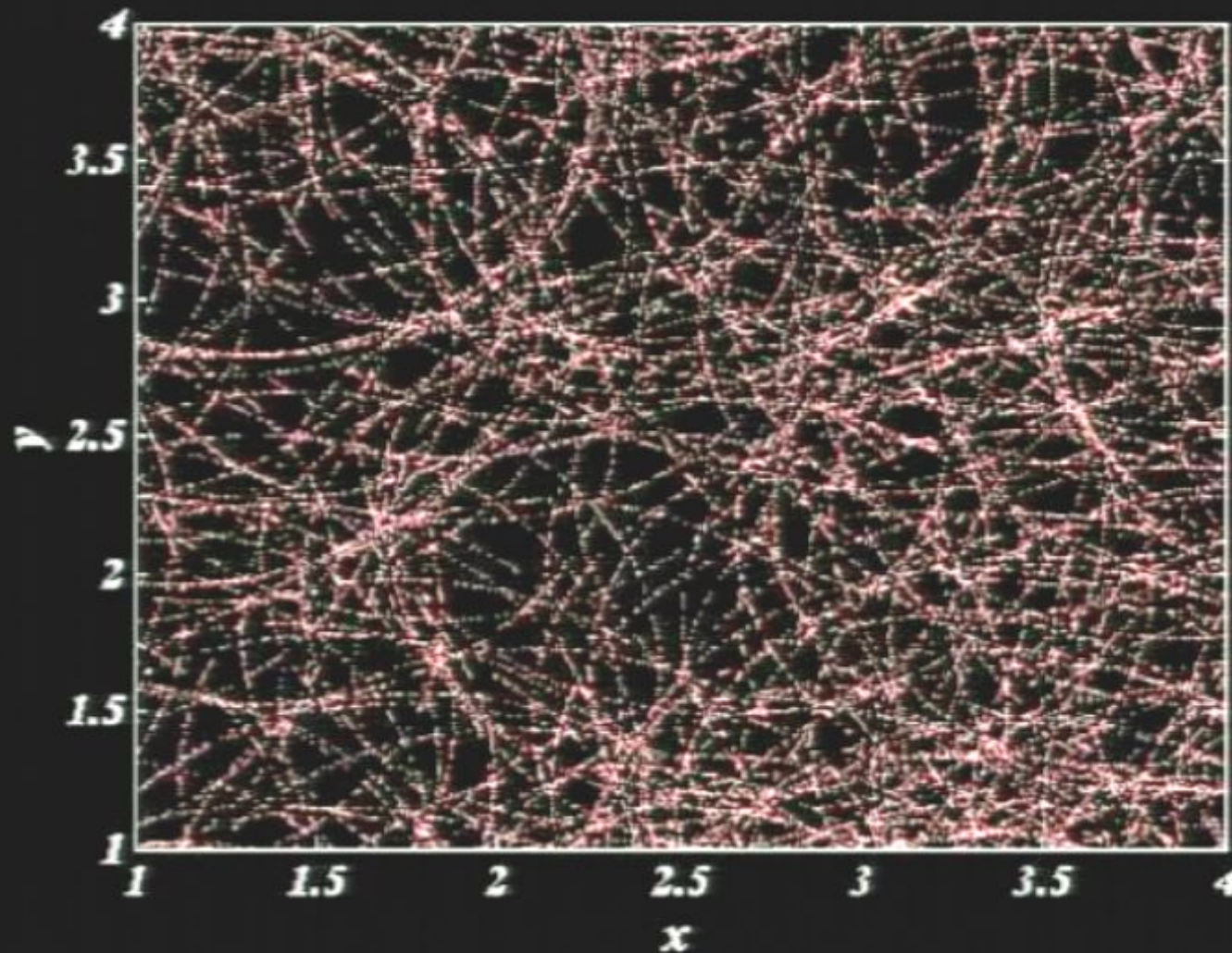




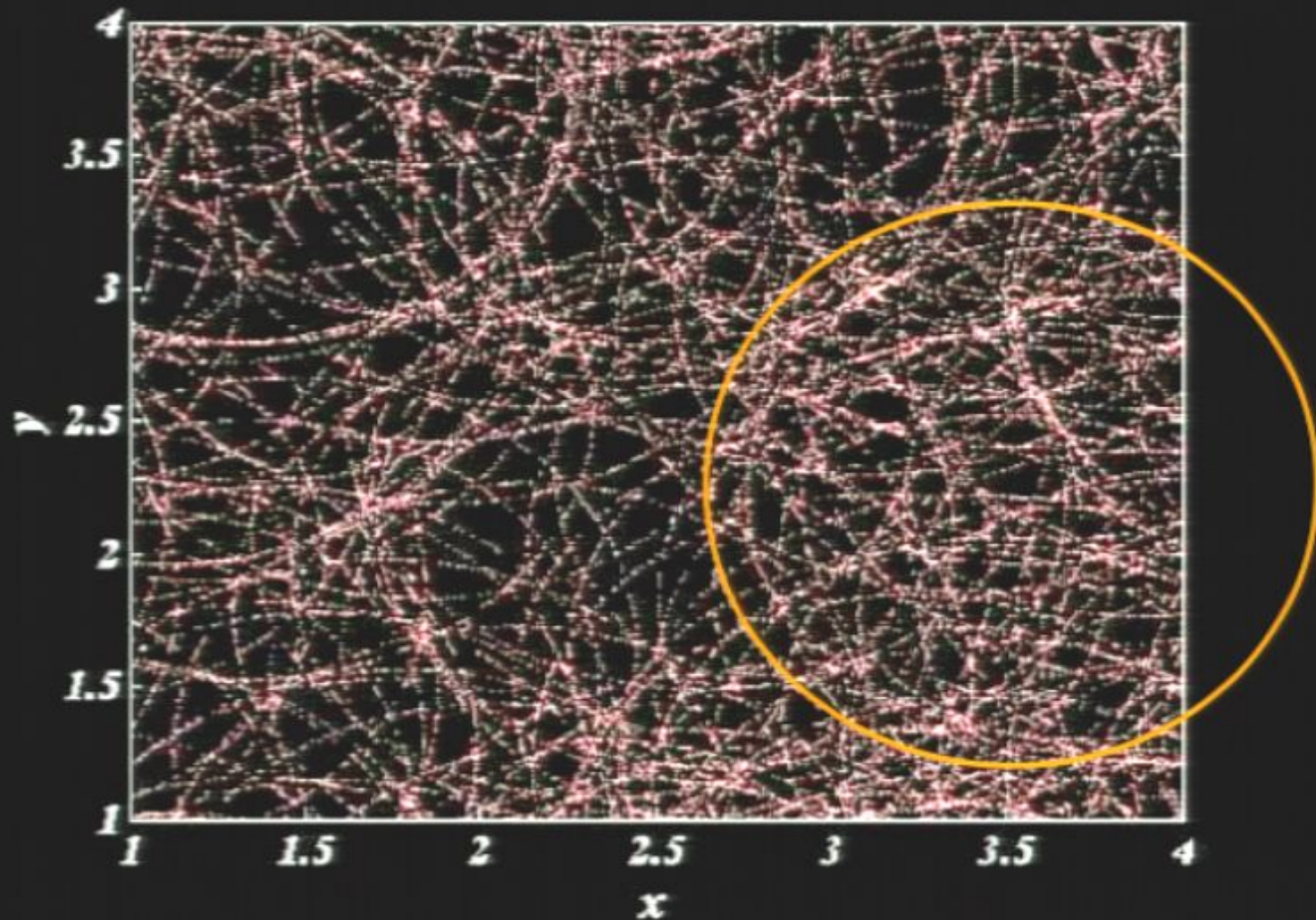




# Statistical Standard Rulers

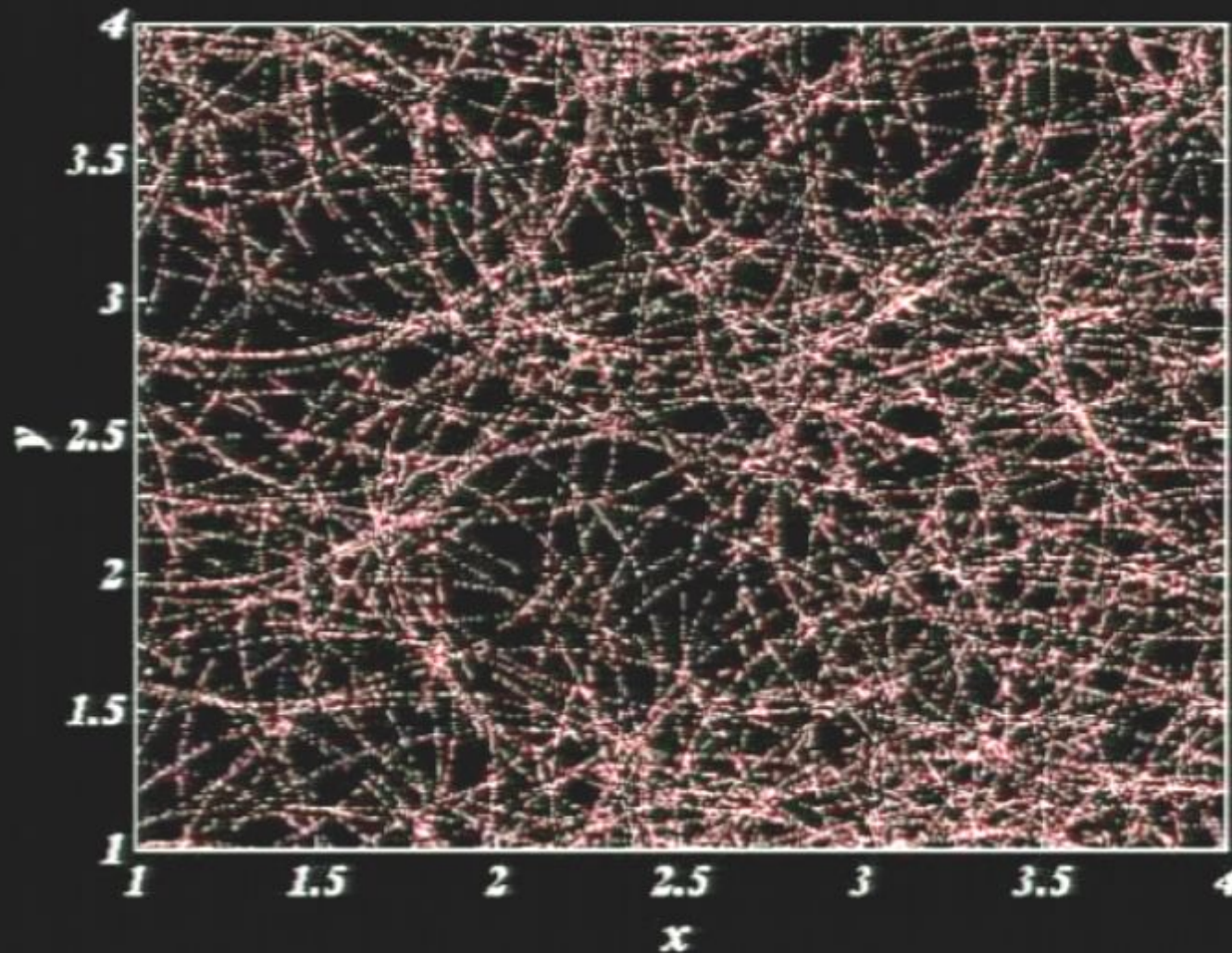


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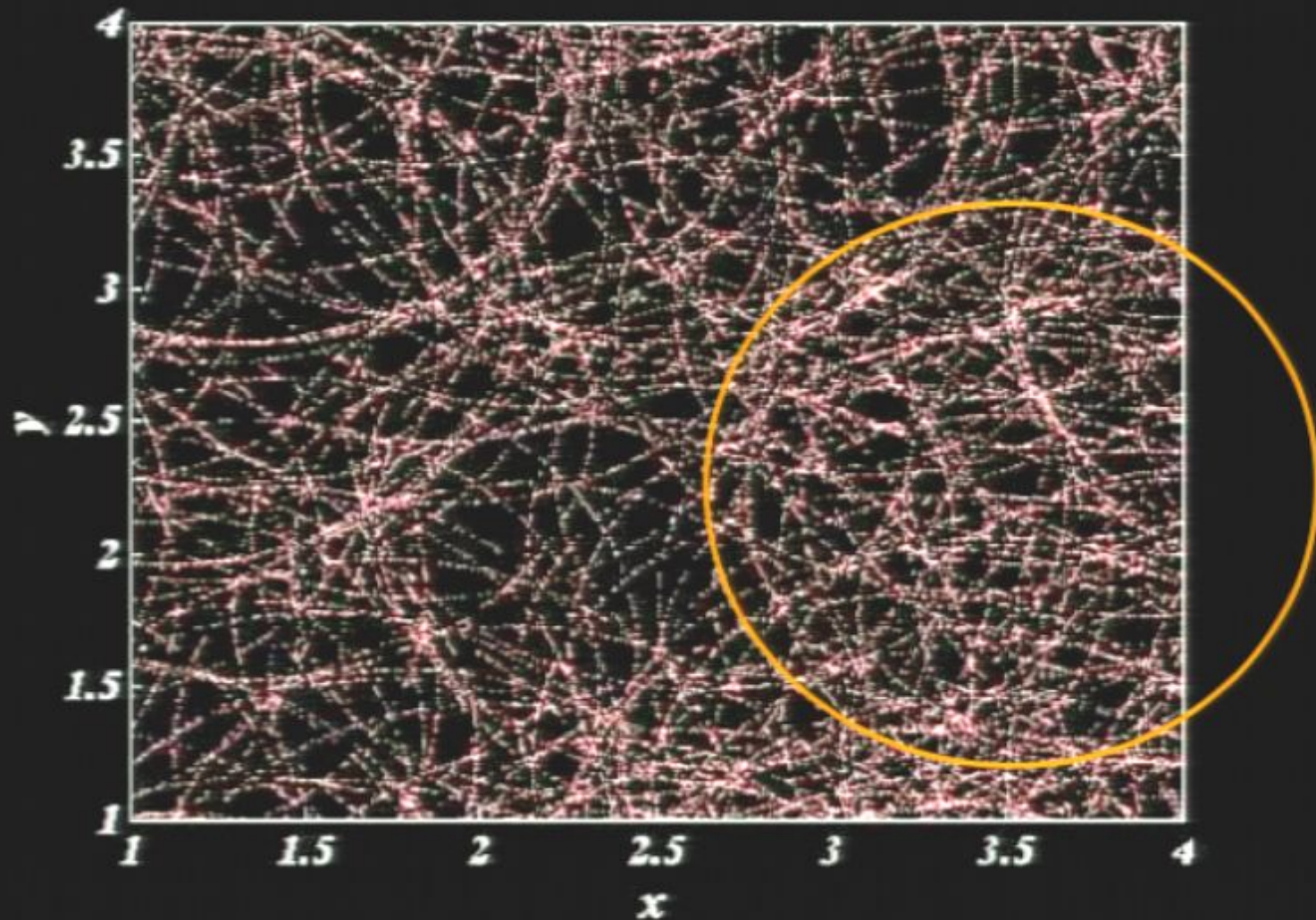




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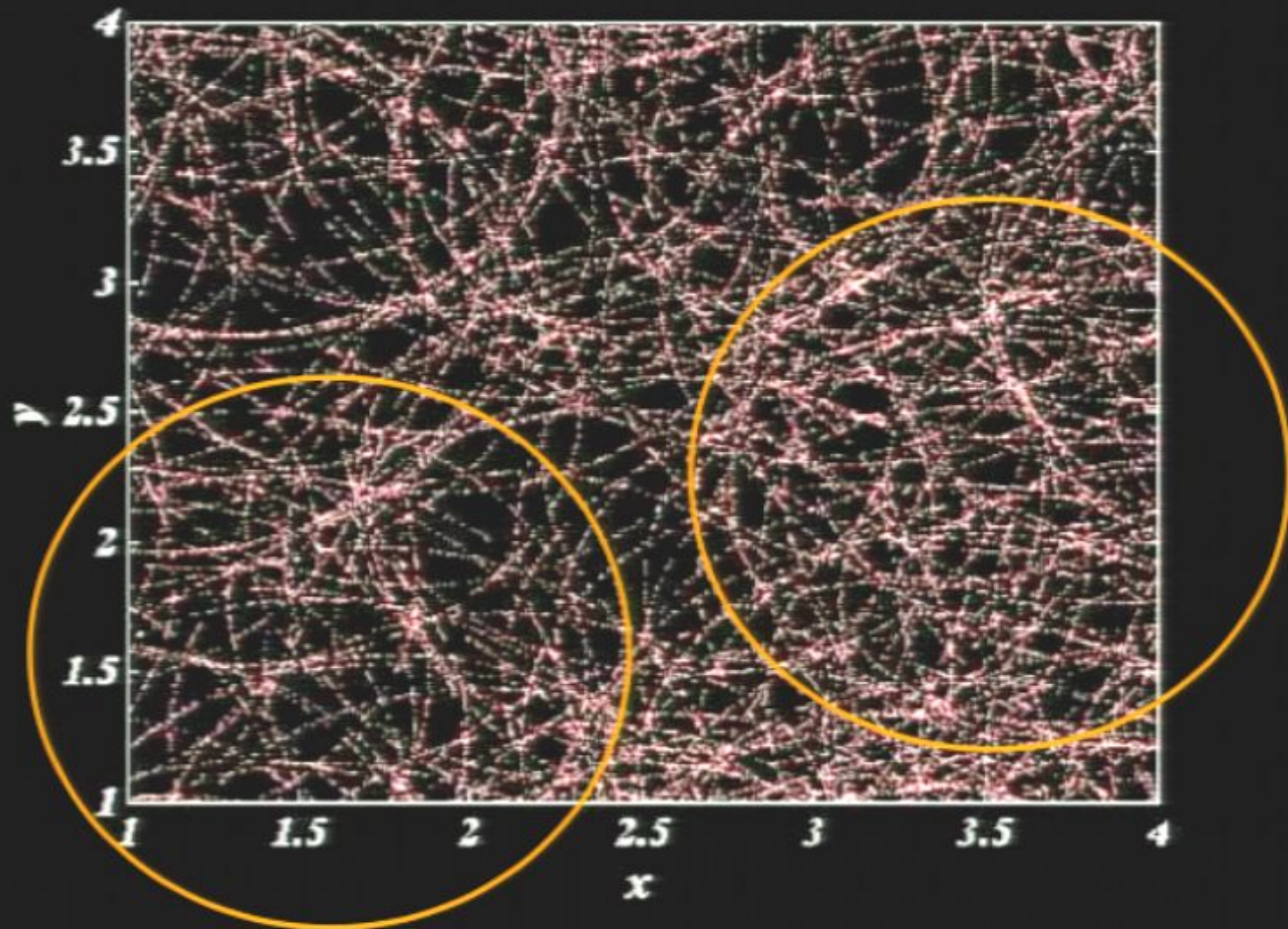


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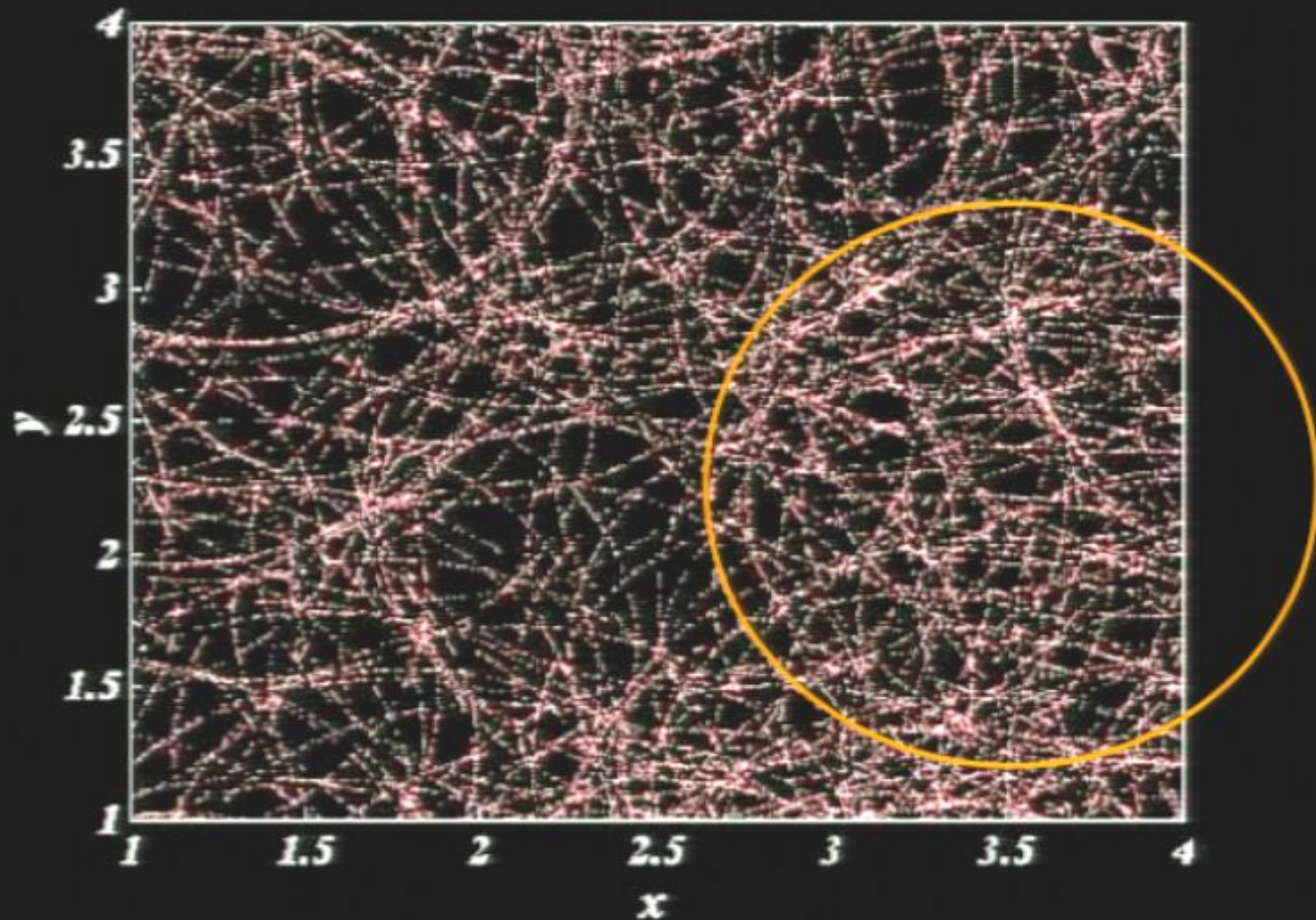


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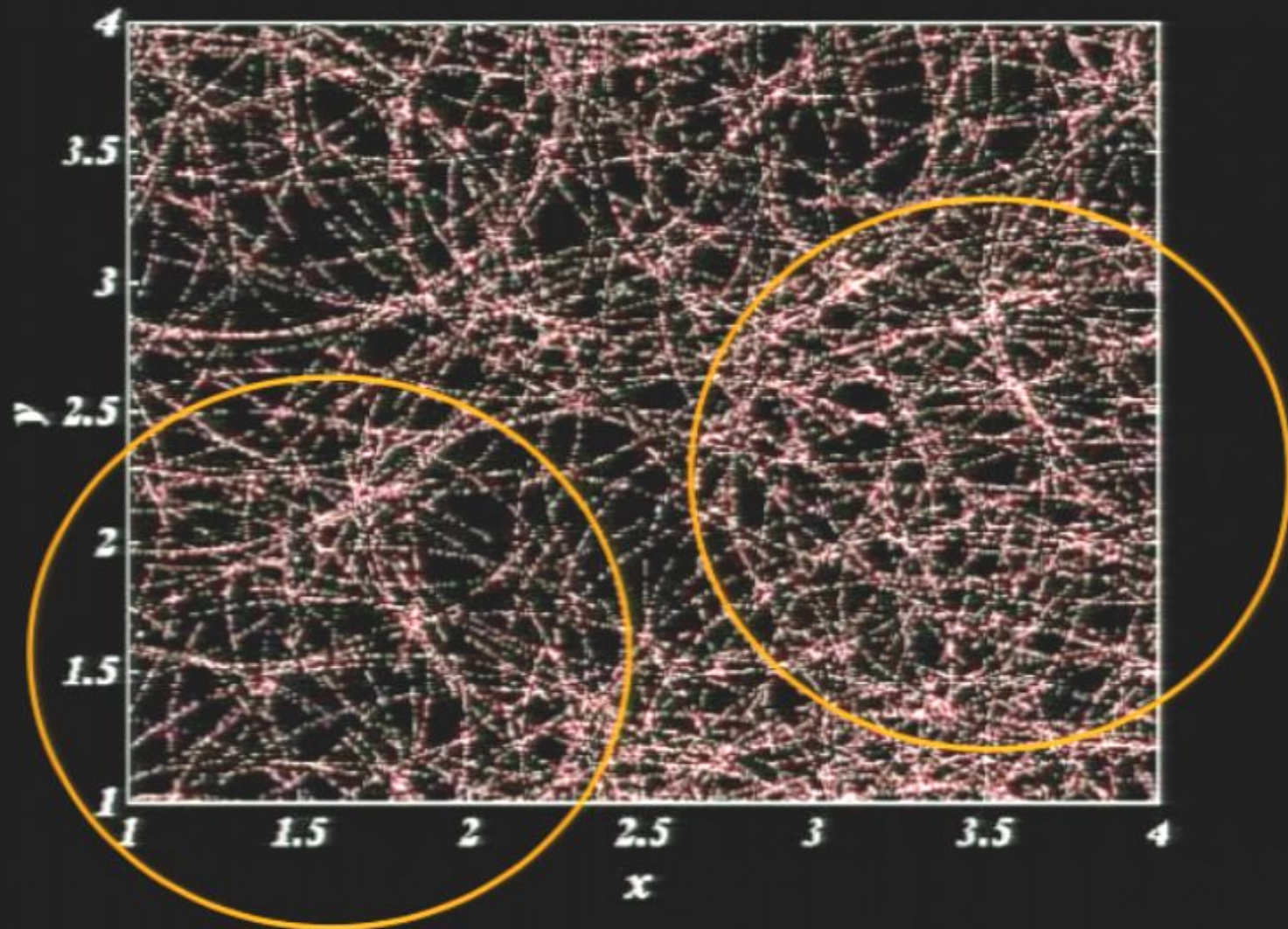


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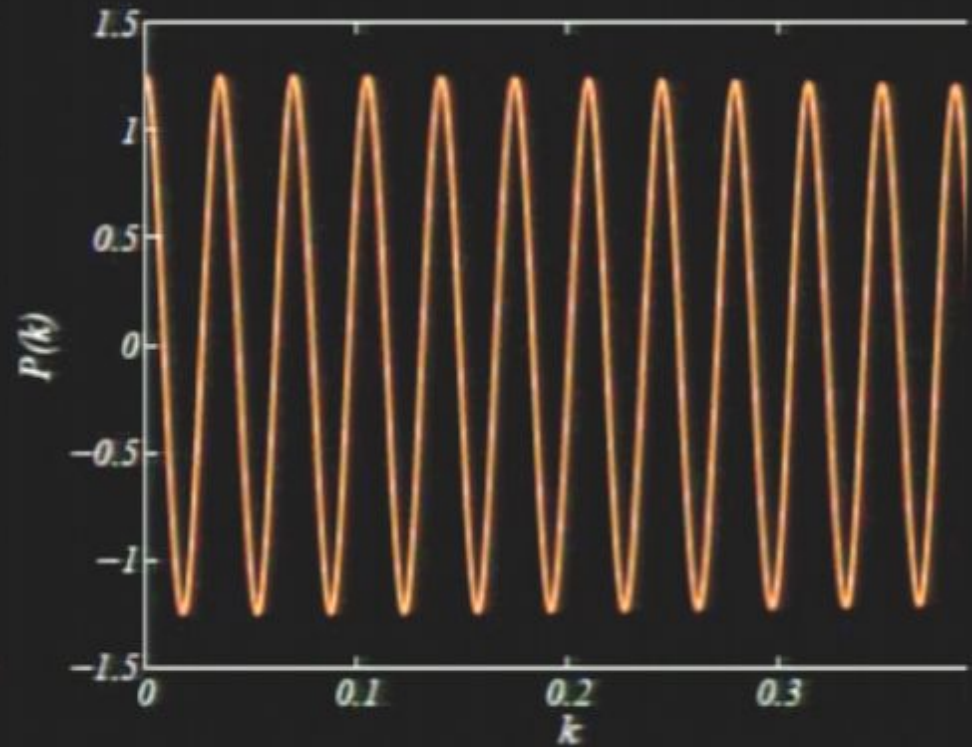
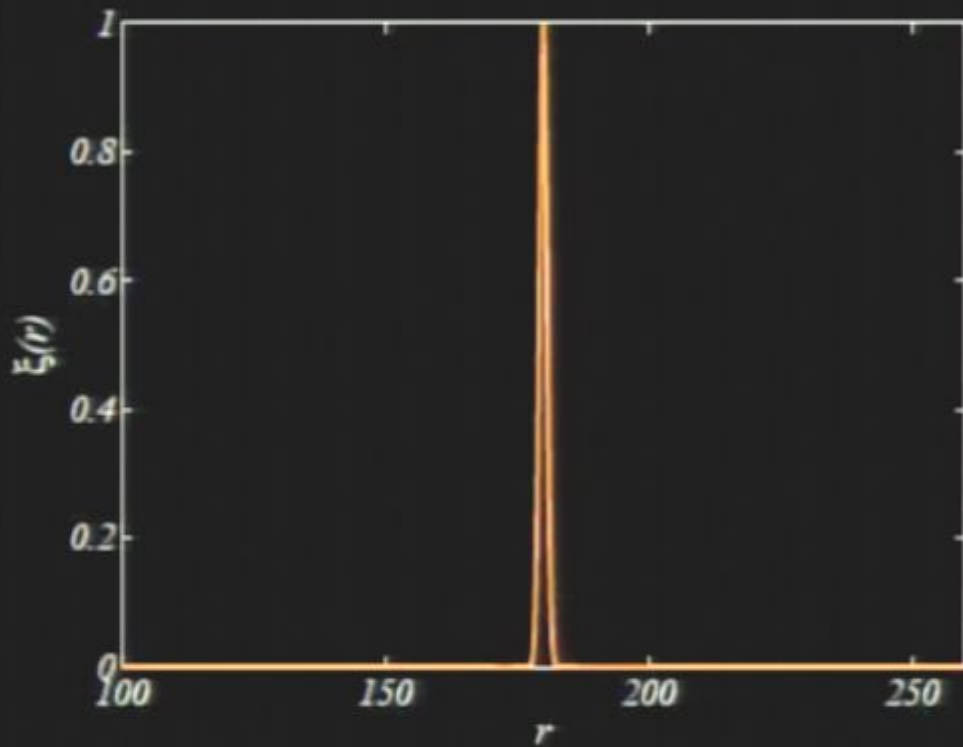




# Statistical Standard Rulers



# Where are the oscillations in BAO?



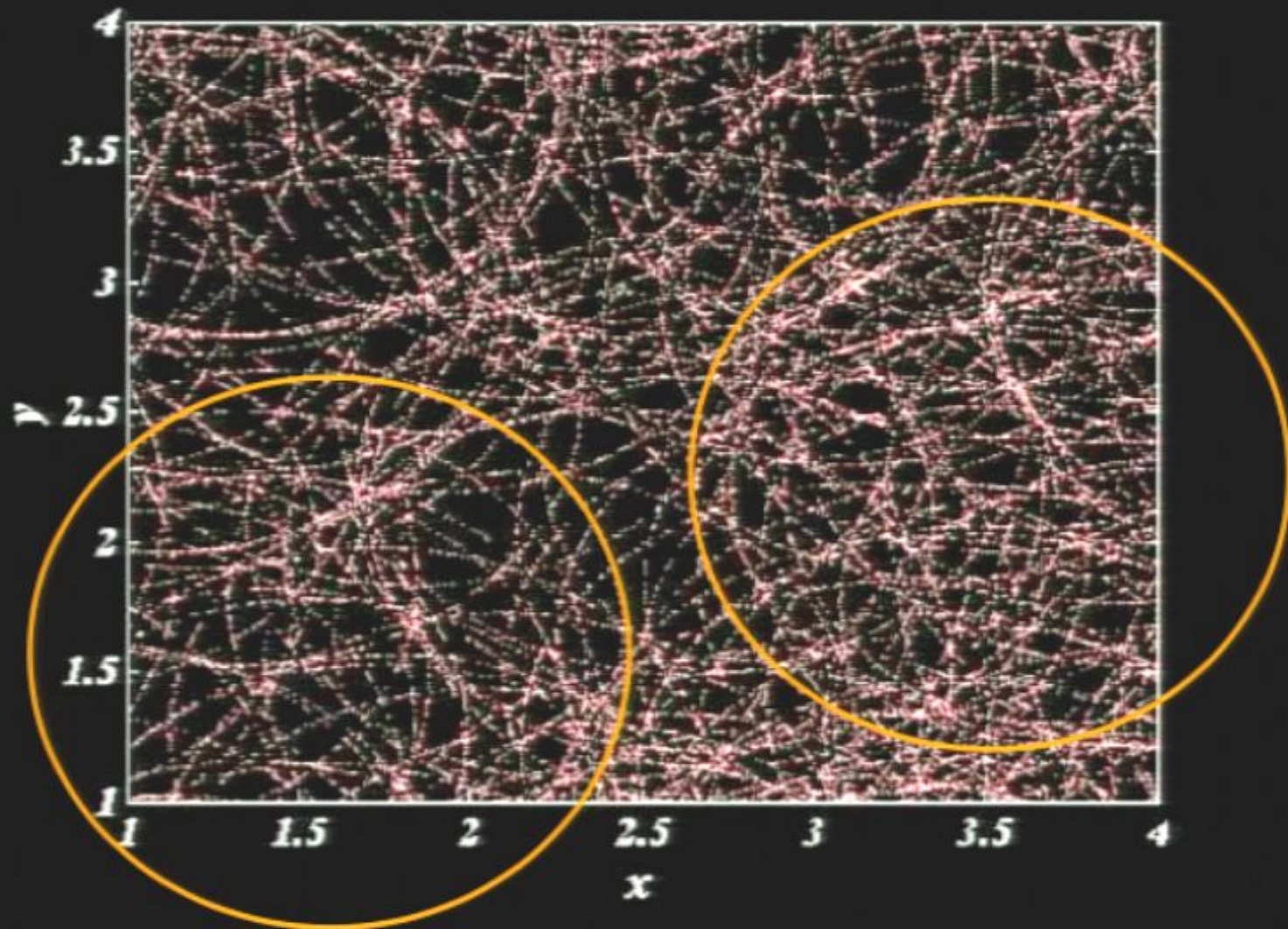
$\xi(r)$



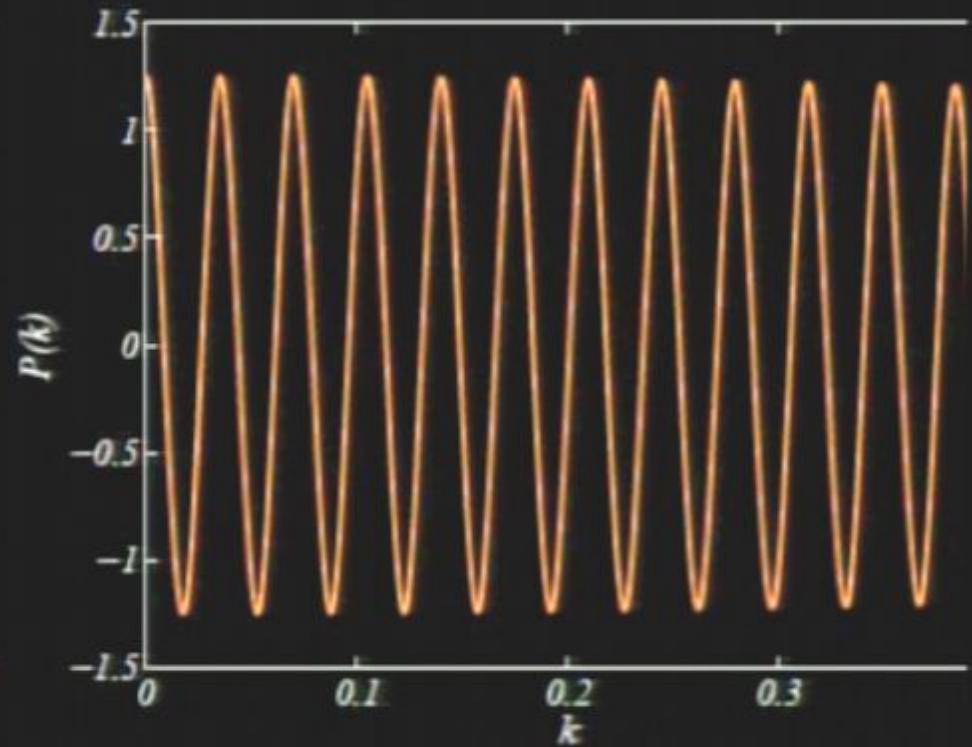
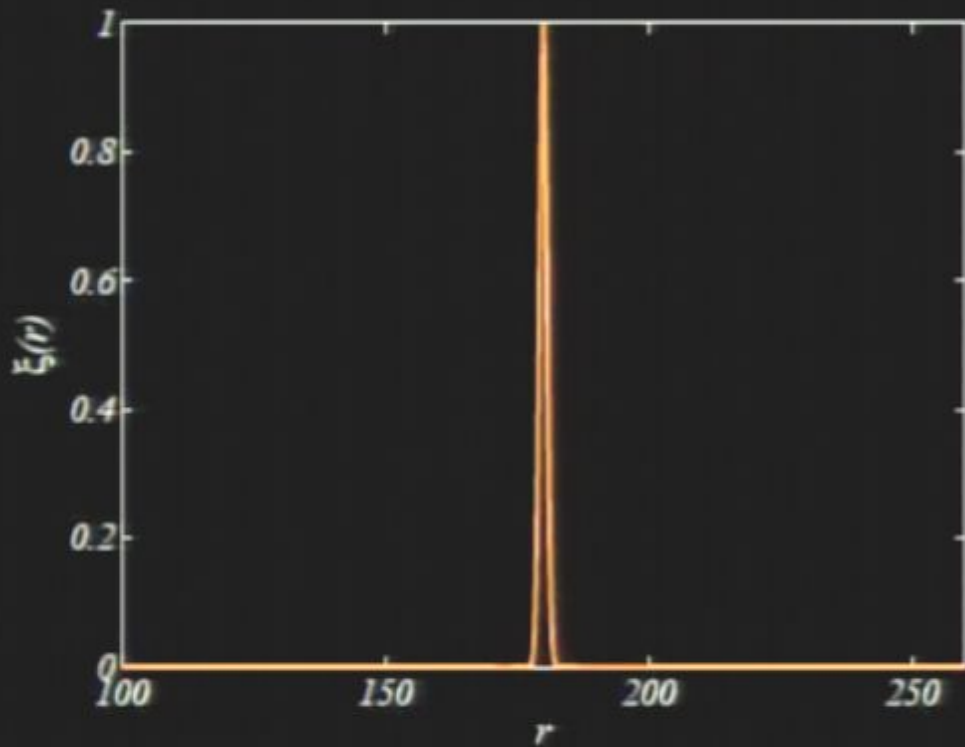
$P(k)$



# Statistical Standard Rulers



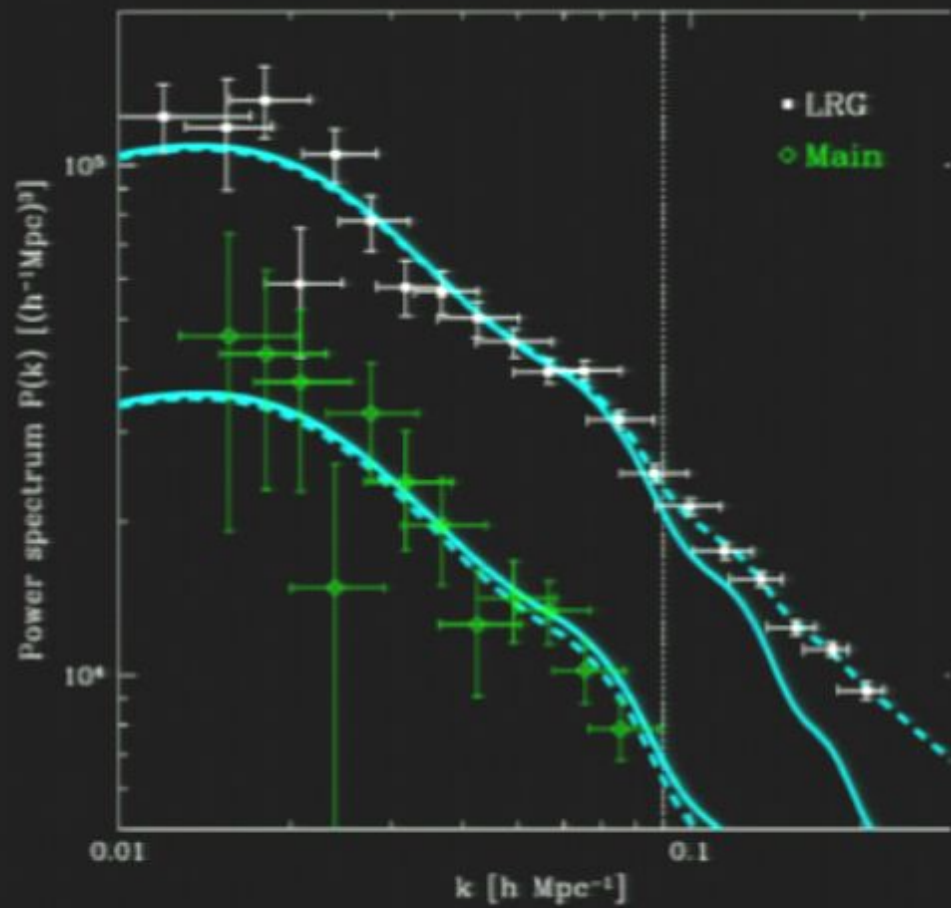
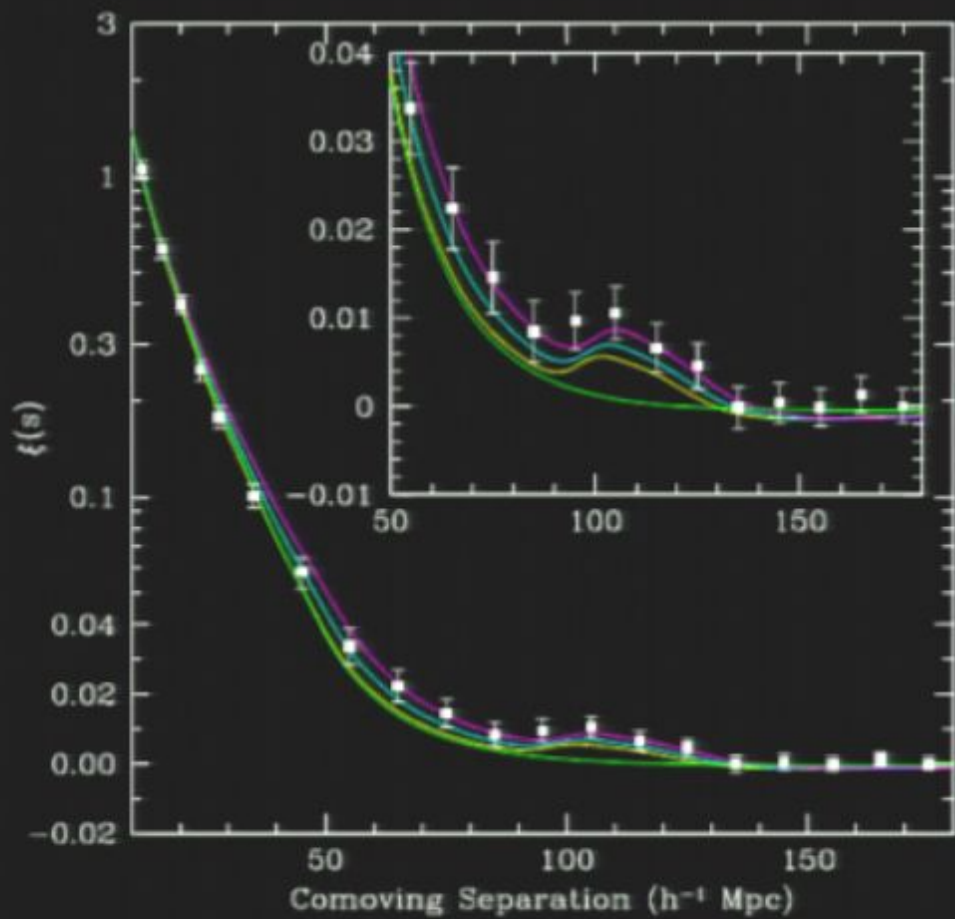
# Where are the oscillations in BAO?



$\xi(r)$



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$\xi(r)$



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- $n \sim 10^{-4} h^3 \text{ Mpc}^{-3}$
- Volume =  $20 \text{ Mpc} \times 100^2 \text{ Mpc}^2 \times 5 \sim 10^6 h^{-3} \text{ Mpc}^3$

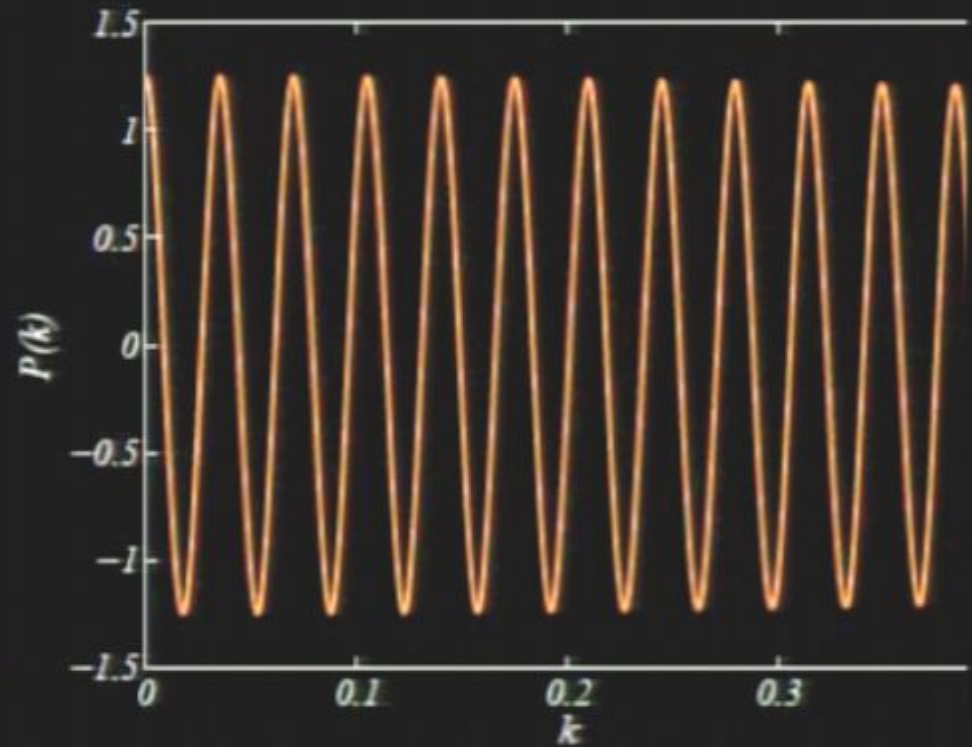
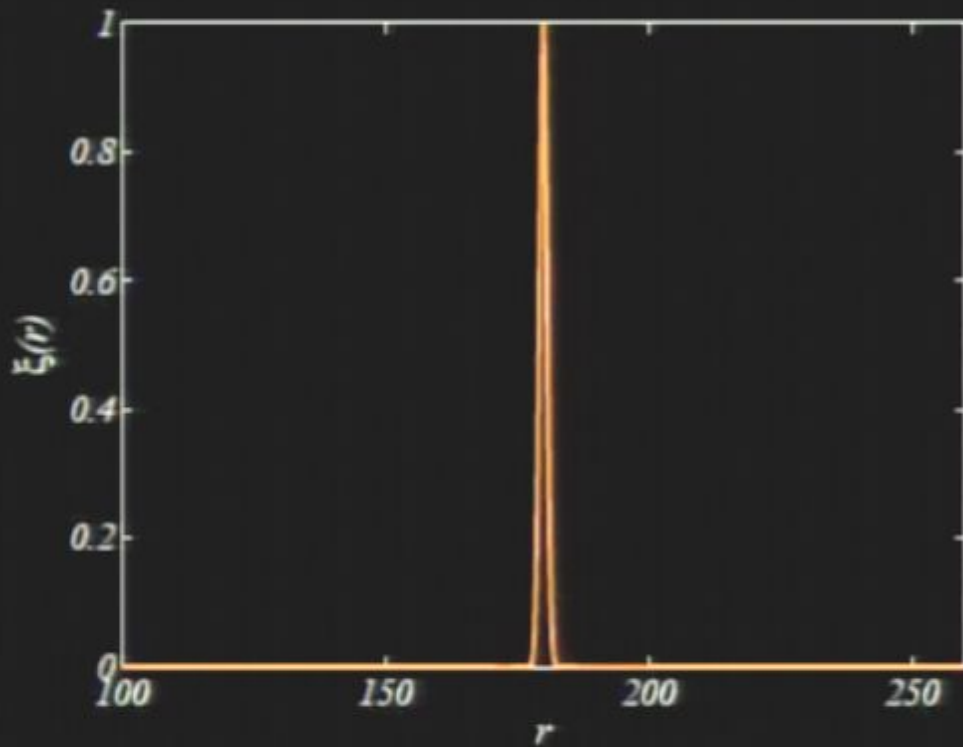
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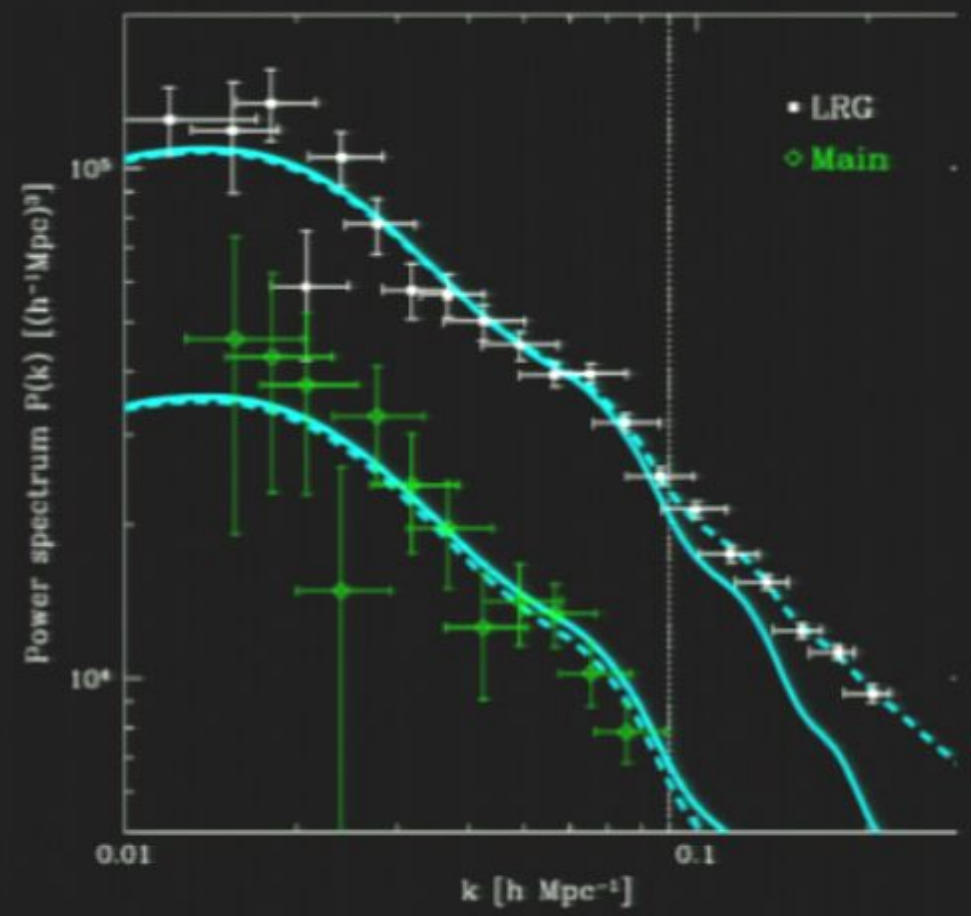
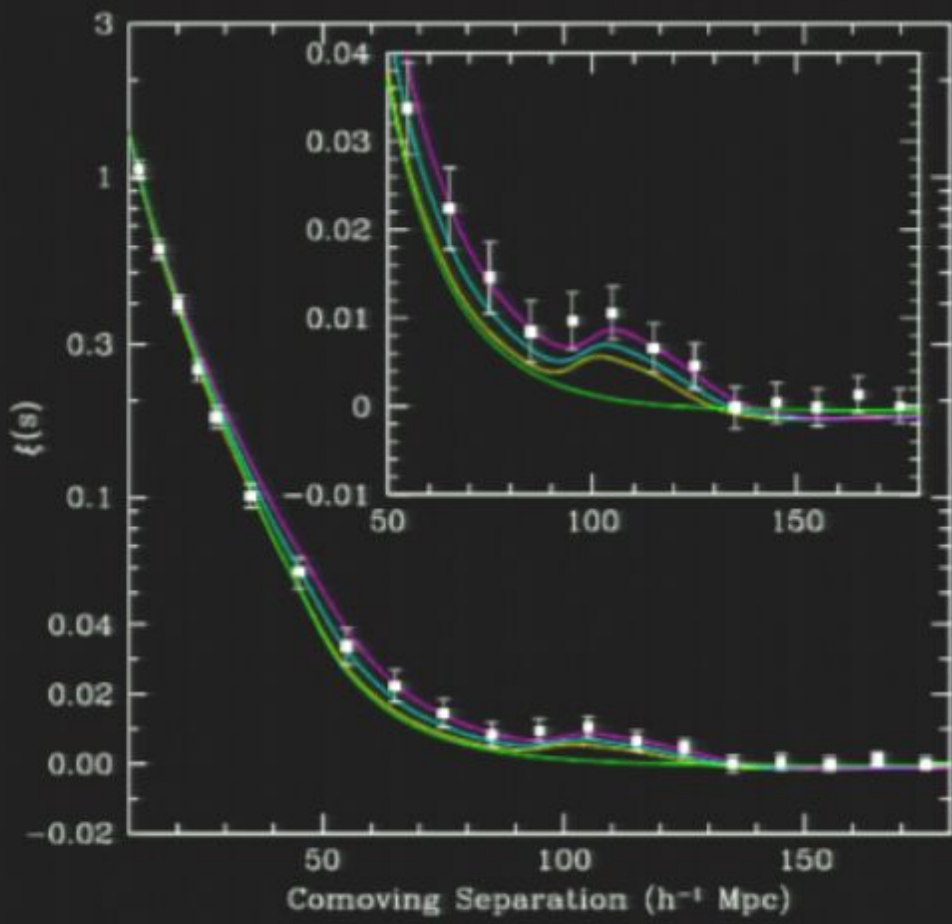


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$P(k)$





$\xi(r)$



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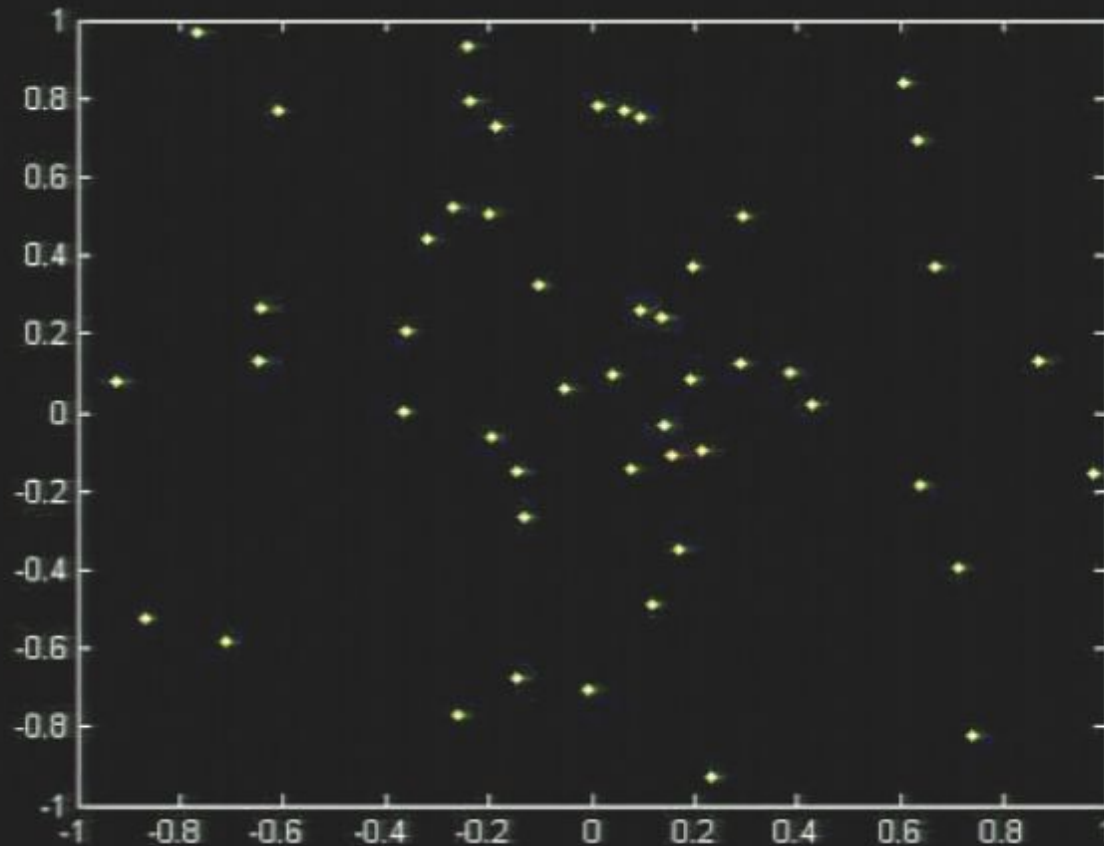
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# Power Spectrum Errors

$$\frac{\delta P}{P} = \underbrace{\frac{1}{\sqrt{m}}}_{\text{}} \left( 1 + \underbrace{\frac{1}{nP}}_{\text{}} \right)$$

# Shot Noise



Fixed Volume

50 points



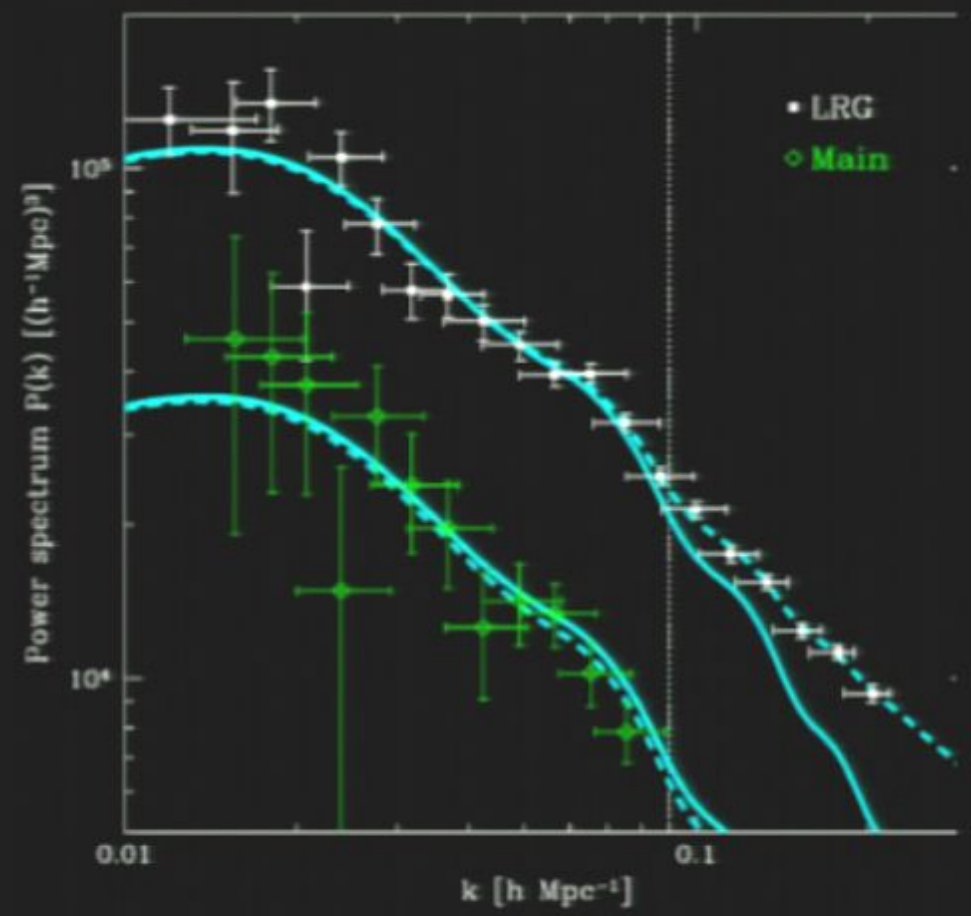
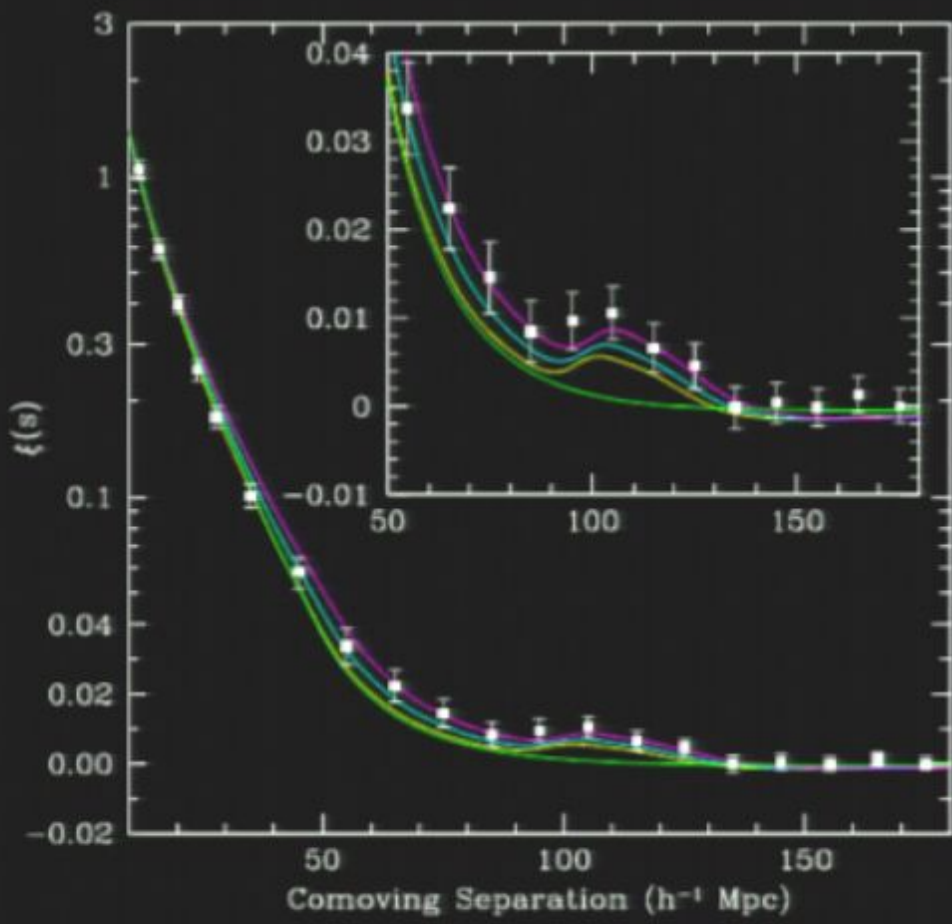
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*m = number of Fourier modes measured in the survey*  
*n = mean galaxy number density in the survey*

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- Need lots of volume and large numbers of galaxies...so we can see the extra galaxy on average!



$\xi(r)$



$P(k)$

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**Cosmic Variance**      **Shot Noise**



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**Cosmic Variance**      **Shot Noise**

# BAP Detection Is Hard!

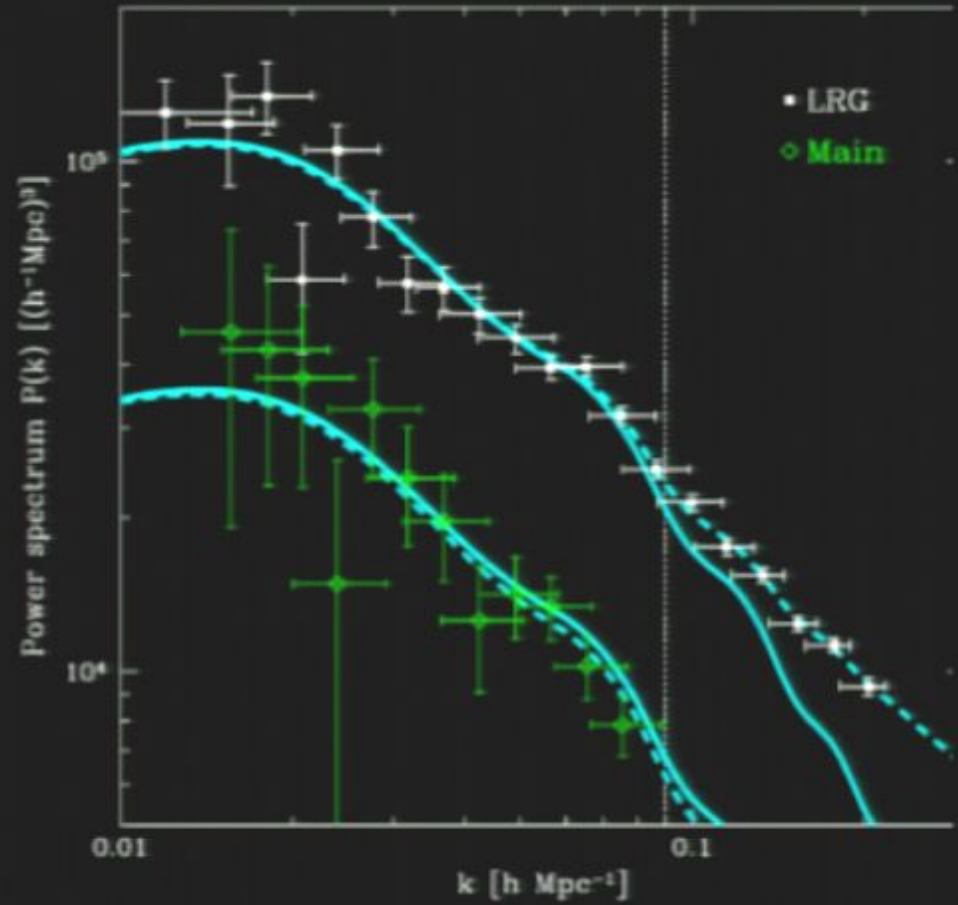
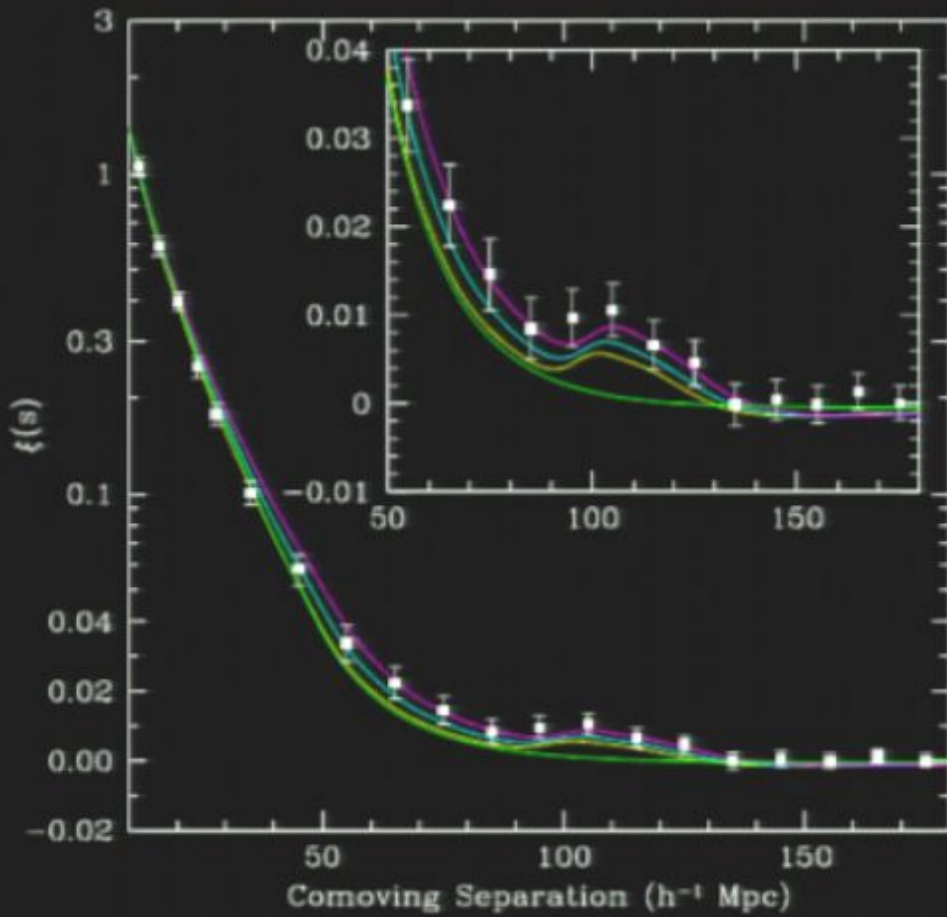
- Consider a typical LRG. How many LRGs do we expect in a BAP shell around it?
- $n \sim 10^{-4} h^3 \text{ Mpc}^{-3}$
- Volume =  $20 \text{ Mpc} \times 100^2 \text{ Mpc}^2 \times 5 \sim 10^6 h^{-3} \text{ Mpc}^3$   
→ 100 LRGs from purely uniform distribution
- BAP in  $\xi(r)$  gives  $\sim 1\%$  excess probability
- So gives a *single extra LRG* on average.
- Need lots of volume and large numbers of galaxies...so we can see the extra galaxy on average!



# BAP Detection Is Hard!

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# BAP Detection Is Hard!



$\xi(r)$



$P(k)$

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# Power Spectrum Errors

$$\frac{\delta P}{P} = \frac{1}{\sqrt{m}} \left( 1 + \frac{1}{nP} \right)$$

  
**Cosmic Variance**      **Shot Noise**

*m = number of Fourier modes measured in the survey*

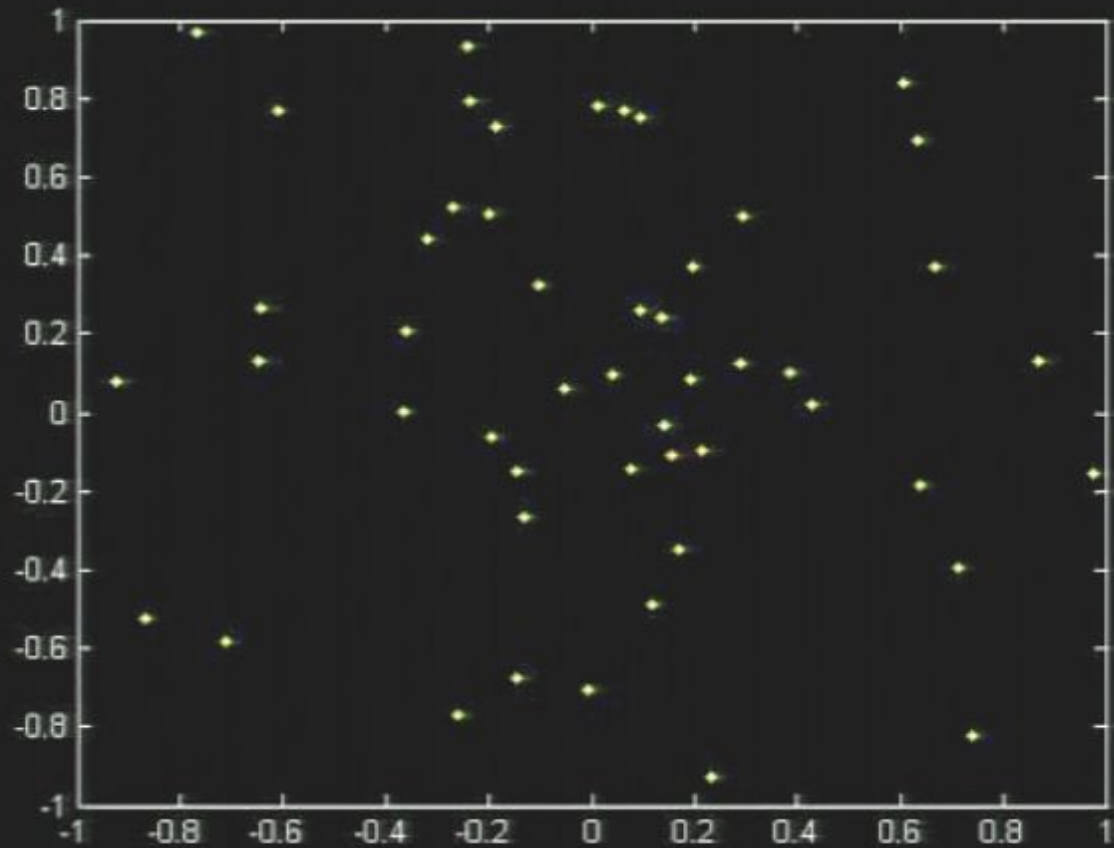
# Power Spectrum Errors

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*m* = number of Fourier modes measured in the survey  
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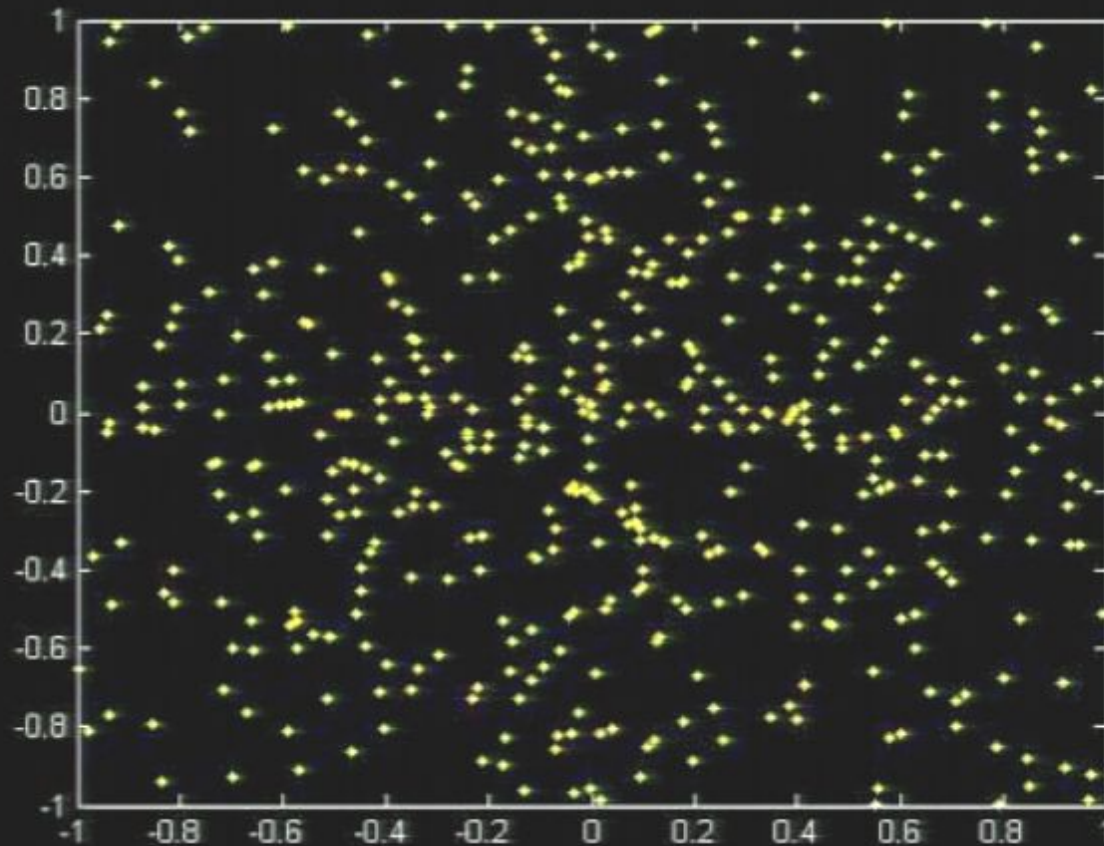
# Shot Noise



Fixed Volume

50 points

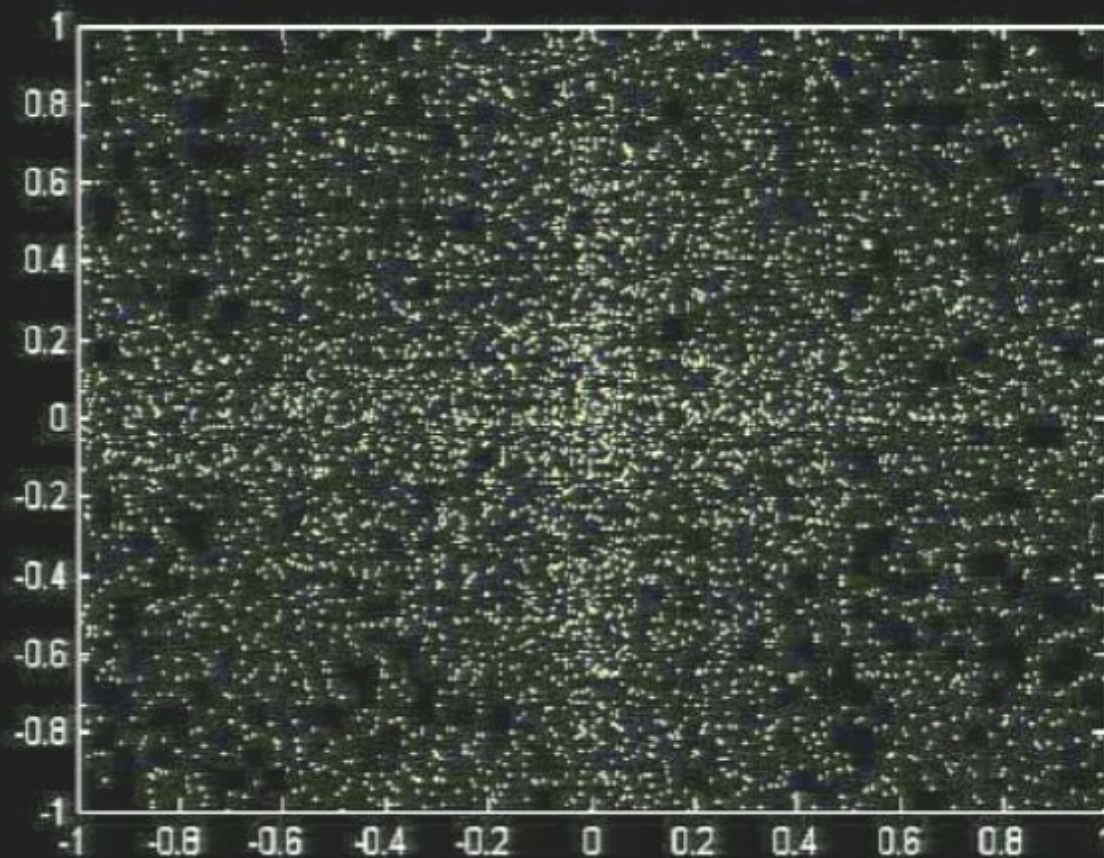
# Shot Noise



Fixed Volume

500 points

# Shot Noise

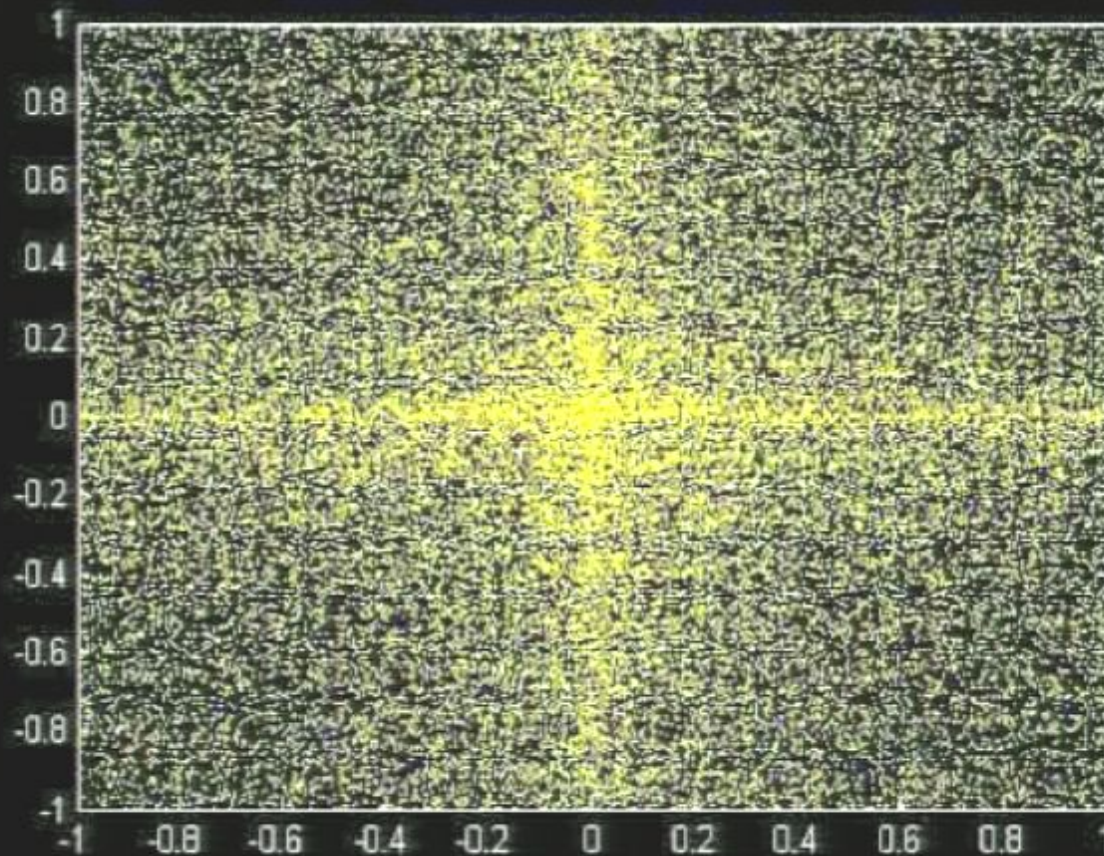


Fixed Volume

5000 points



# Shot Noise

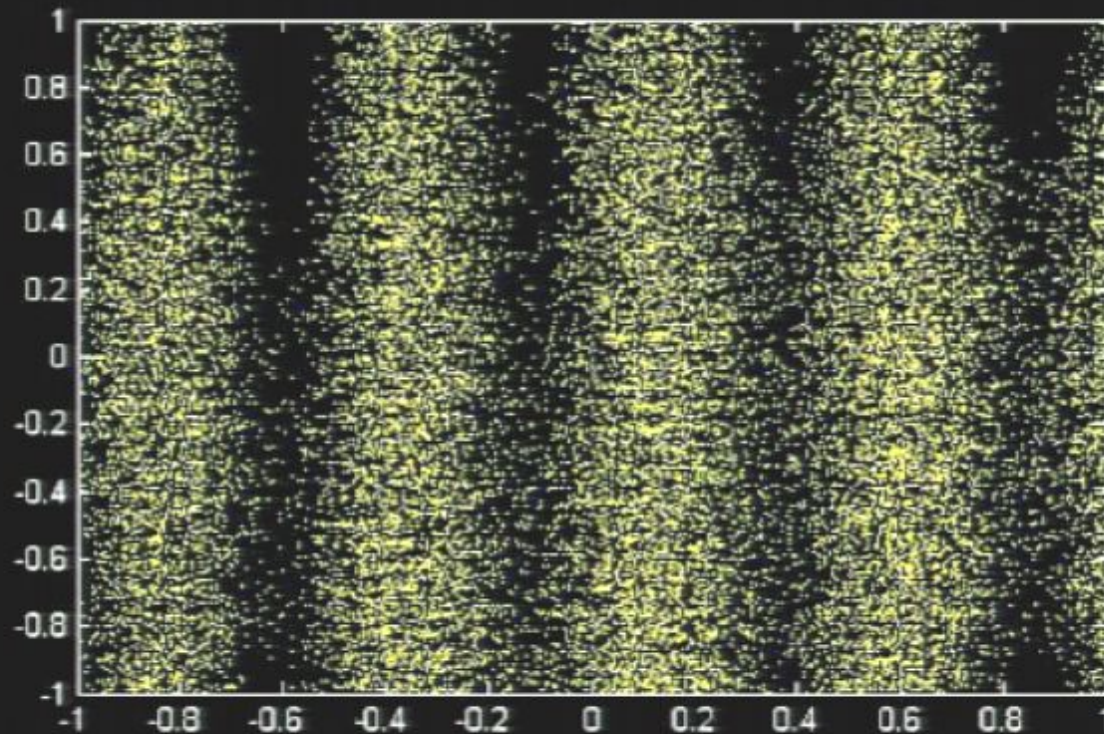


Fixed Volume

50k points



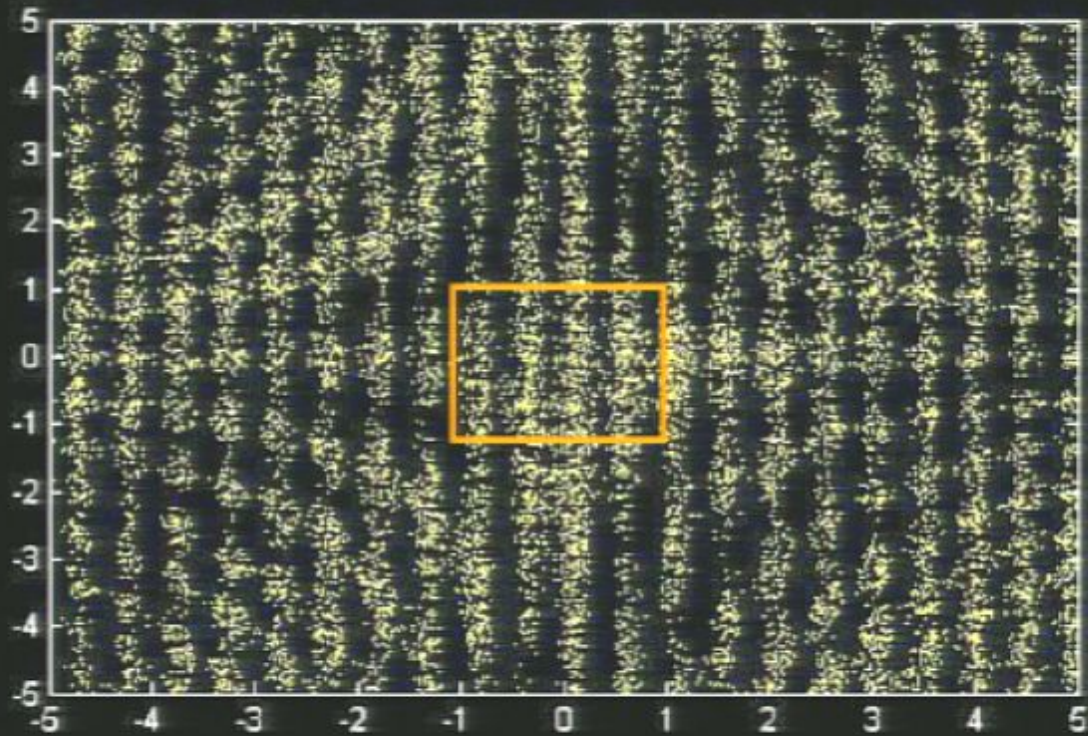
# Cosmic Variance



1 x Volume

Fixed 50k points

# Cosmic Variance

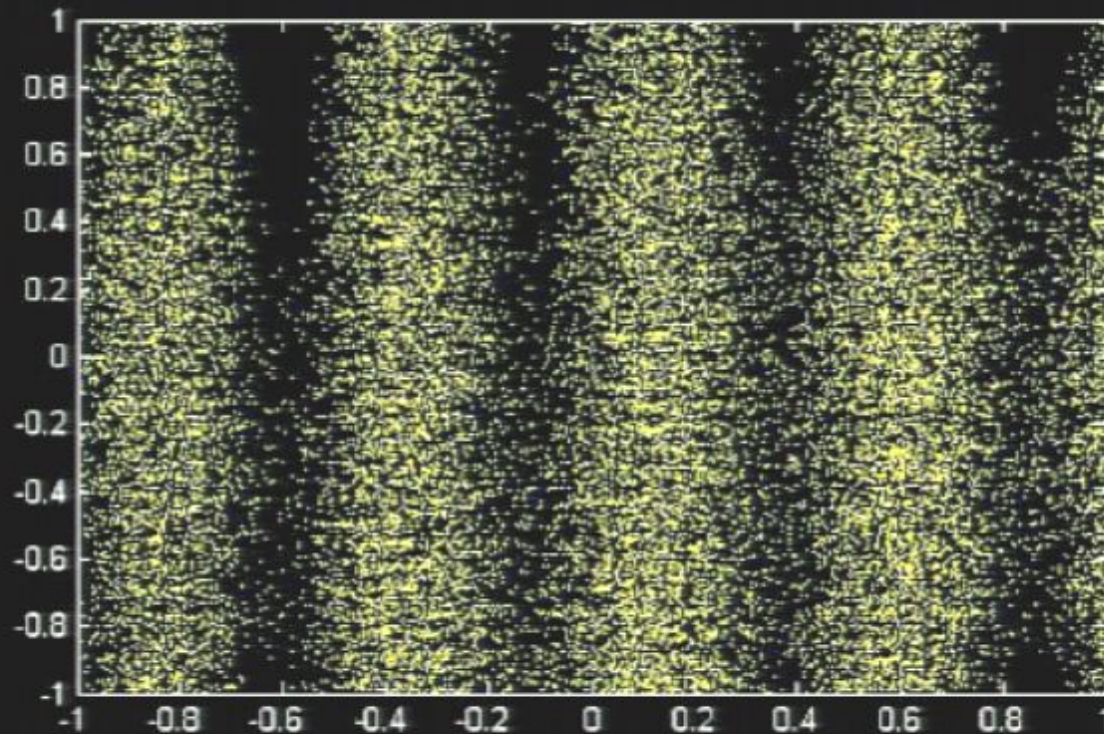


25 x Volume

Fixed 50k points



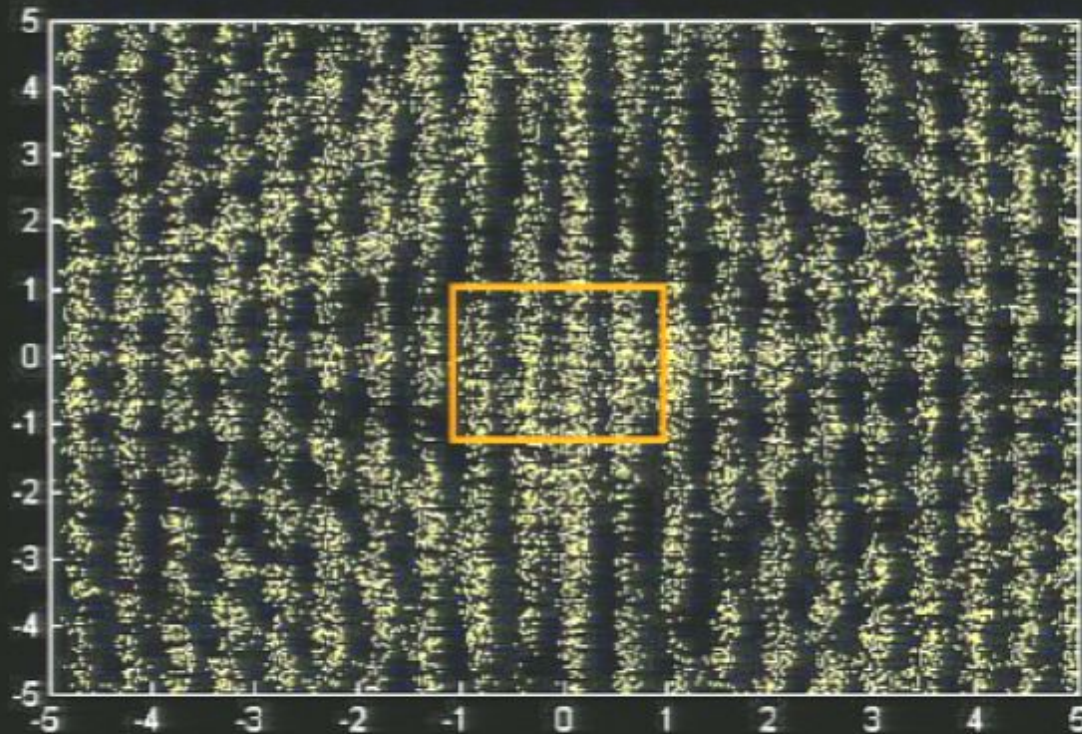
# Cosmic Variance



1 x Volume

Fixed 50k points

# Cosmic Variance

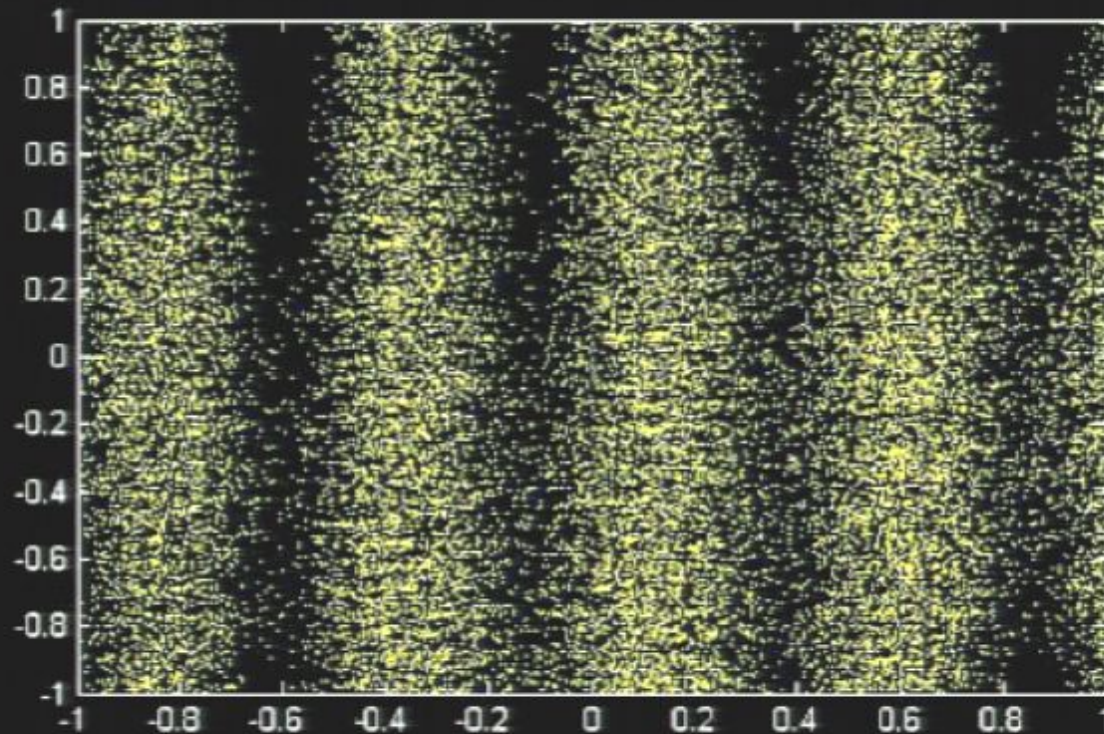


25 x Volume

Fixed 50k points



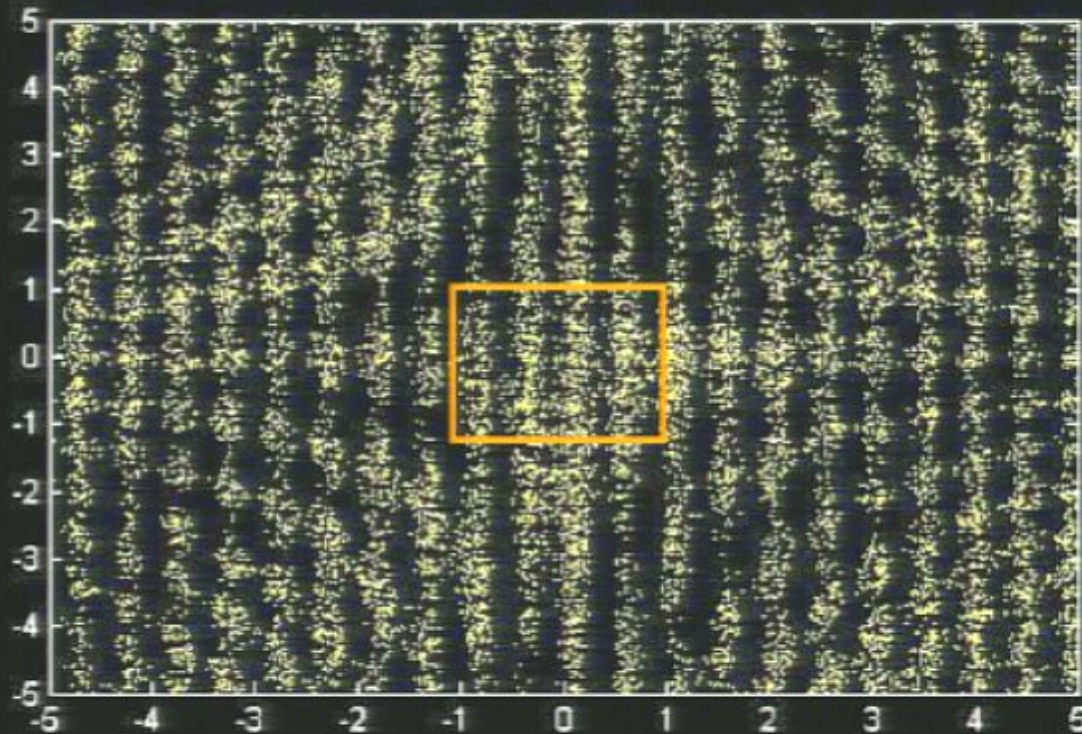
# Cosmic Variance



1 x Volume

Fixed 50k points

# Cosmic Variance

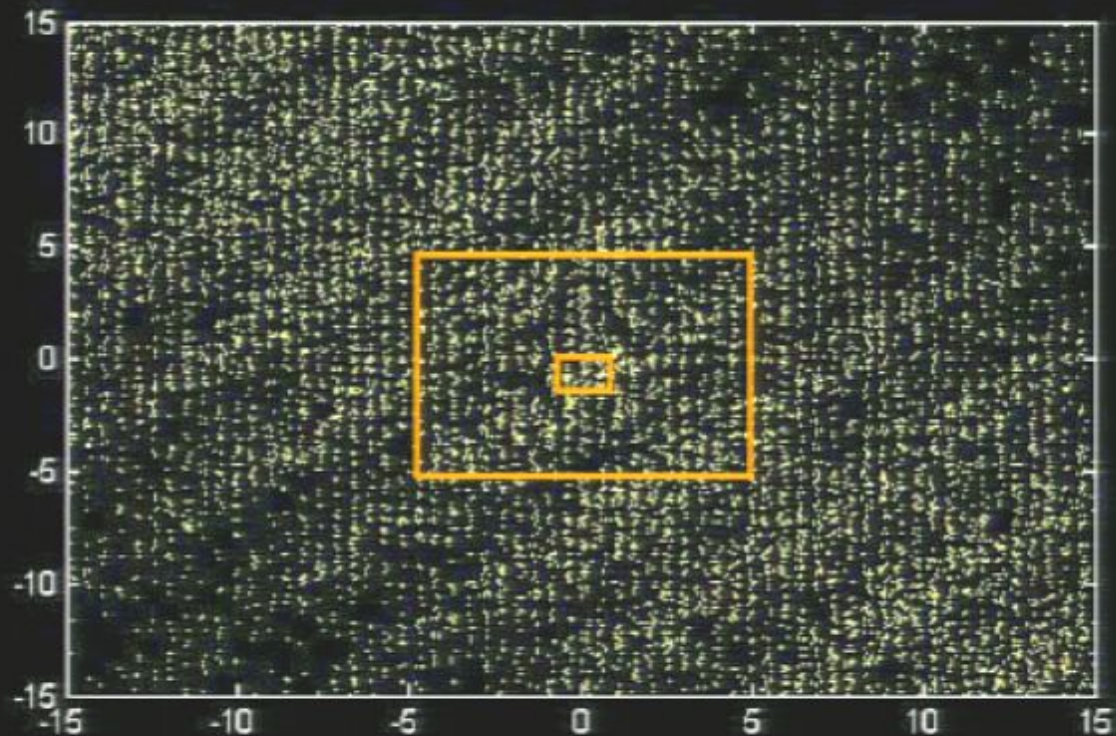


25 x Volume

Fixed 50k points



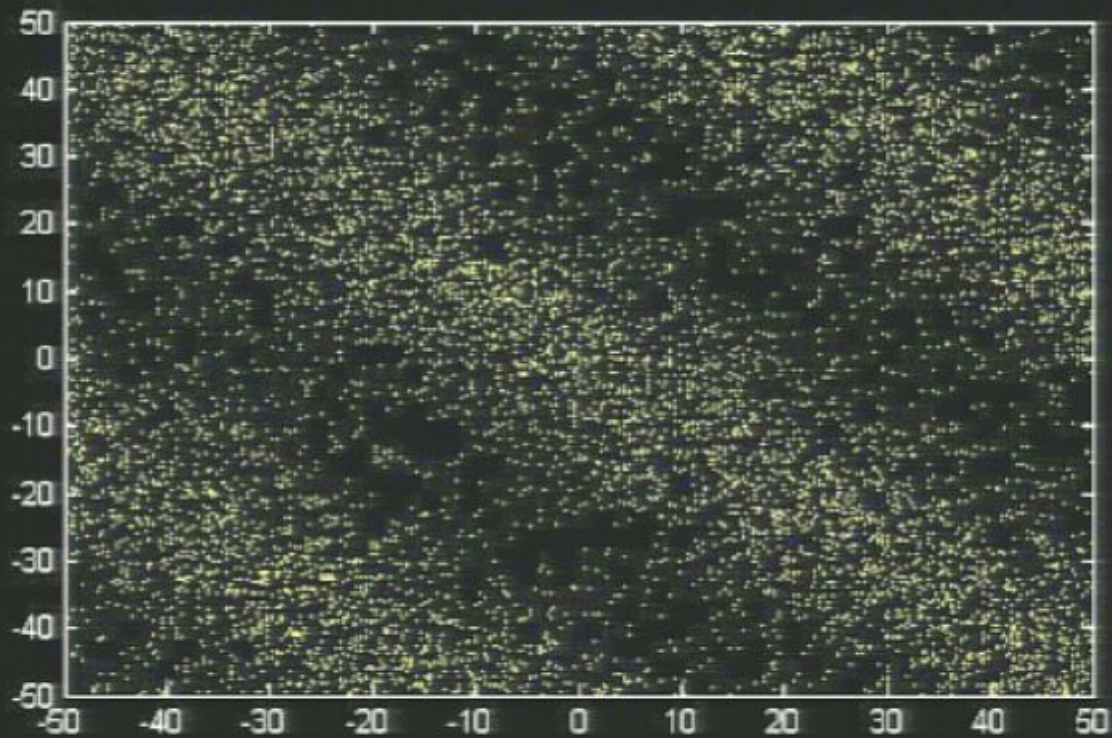
# Cosmic Variance



225 x Volume

Fixed 50k points

# Cosmic Variance

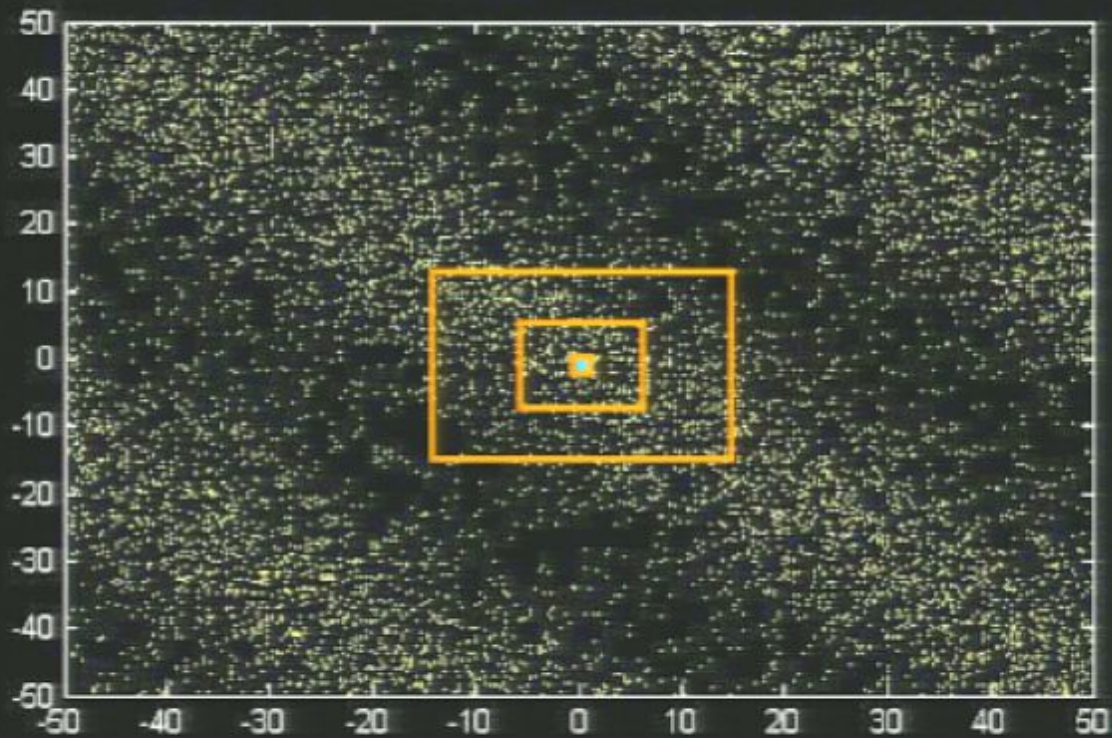


2200 x Volume

Fixed 50k points



# Cosmic Variance



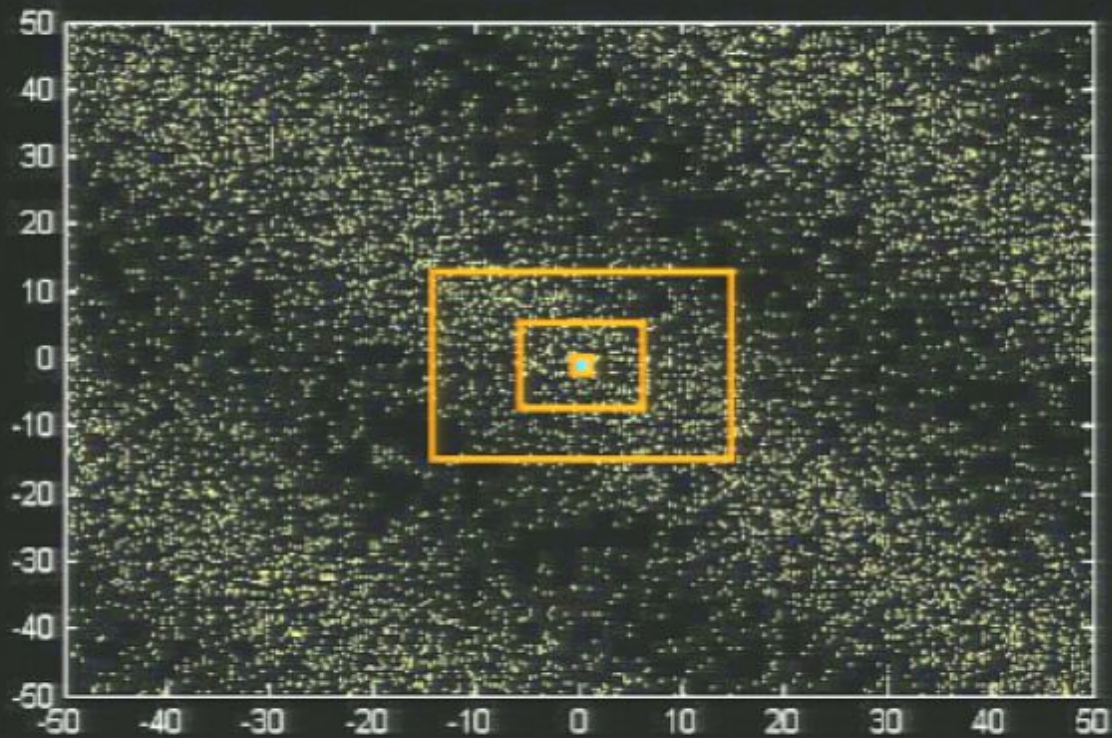
2200 x Volume

Fixed 50k points

# Optimal Survey Design

- What is the optimal sampling of a page of text?
- Depends what information you want.
- E.g. what language it is written in...
- E.g. what page layout is being used...
- Trade-off between the two...
- $nP \sim 1-5$  is optimal

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2200 x Volume

Fixed 50k points



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r of point  
r of target  
, leading t

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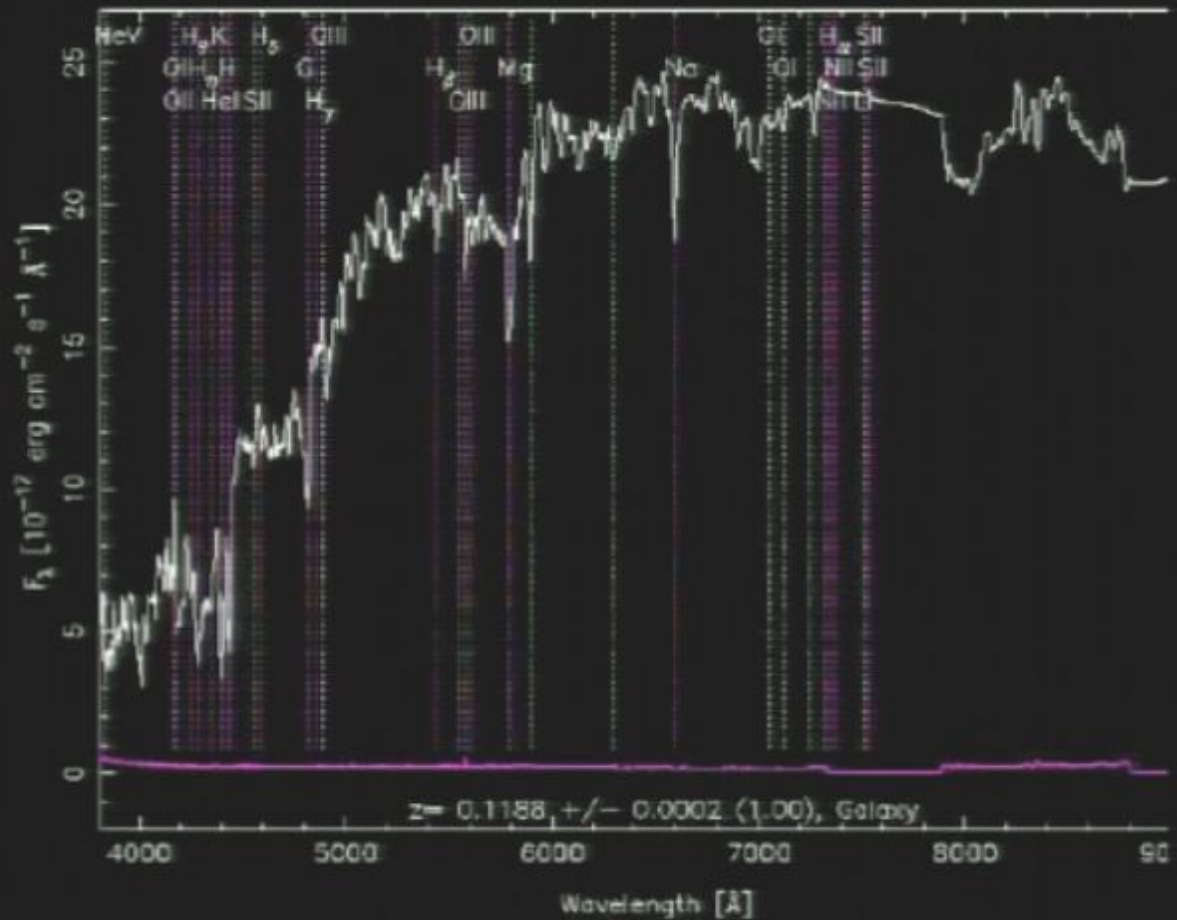
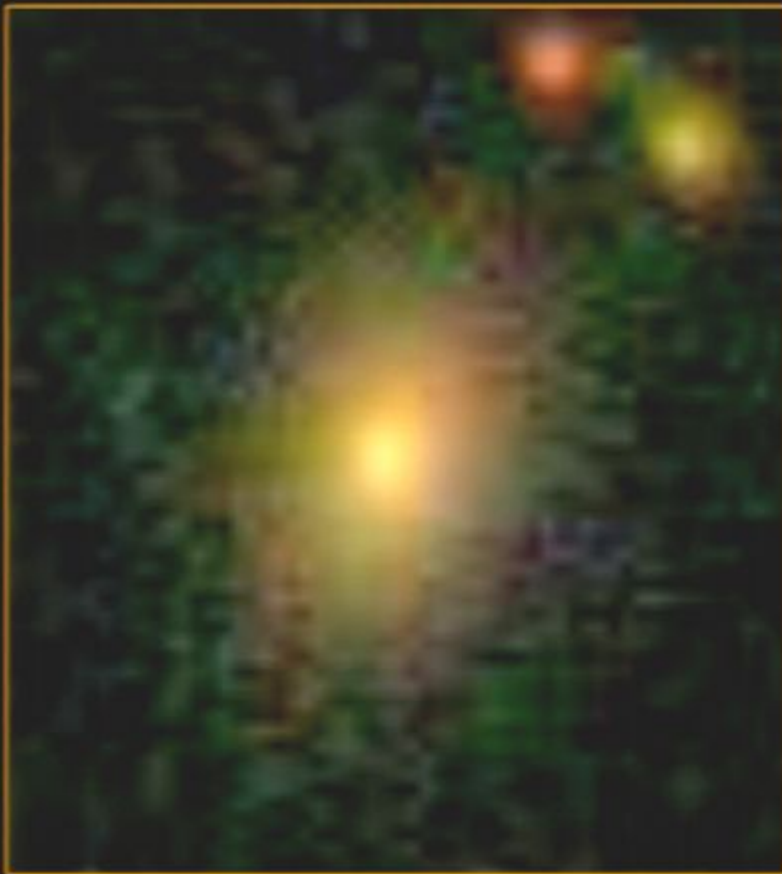


# BAO Targets

What targets should one use for measuring the BAO? Options are:

- Luminous Red Galaxies (LRGs)
- Star-forming galaxies (blue)
- Neutral Hydrogen (HI)
- Lyman alpha forest

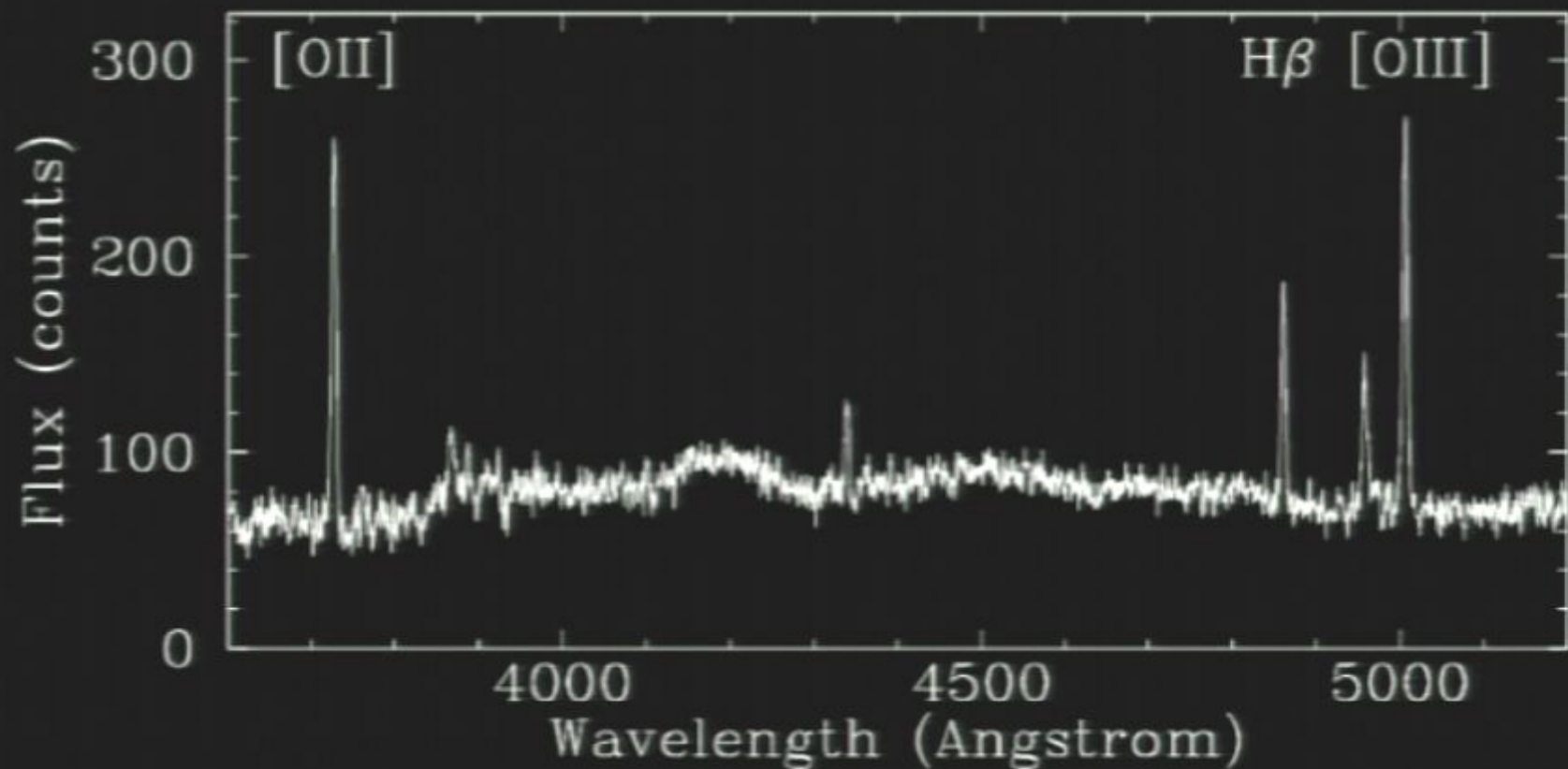
# LRGs



- Used by SDSS, BOSS and LAMOST

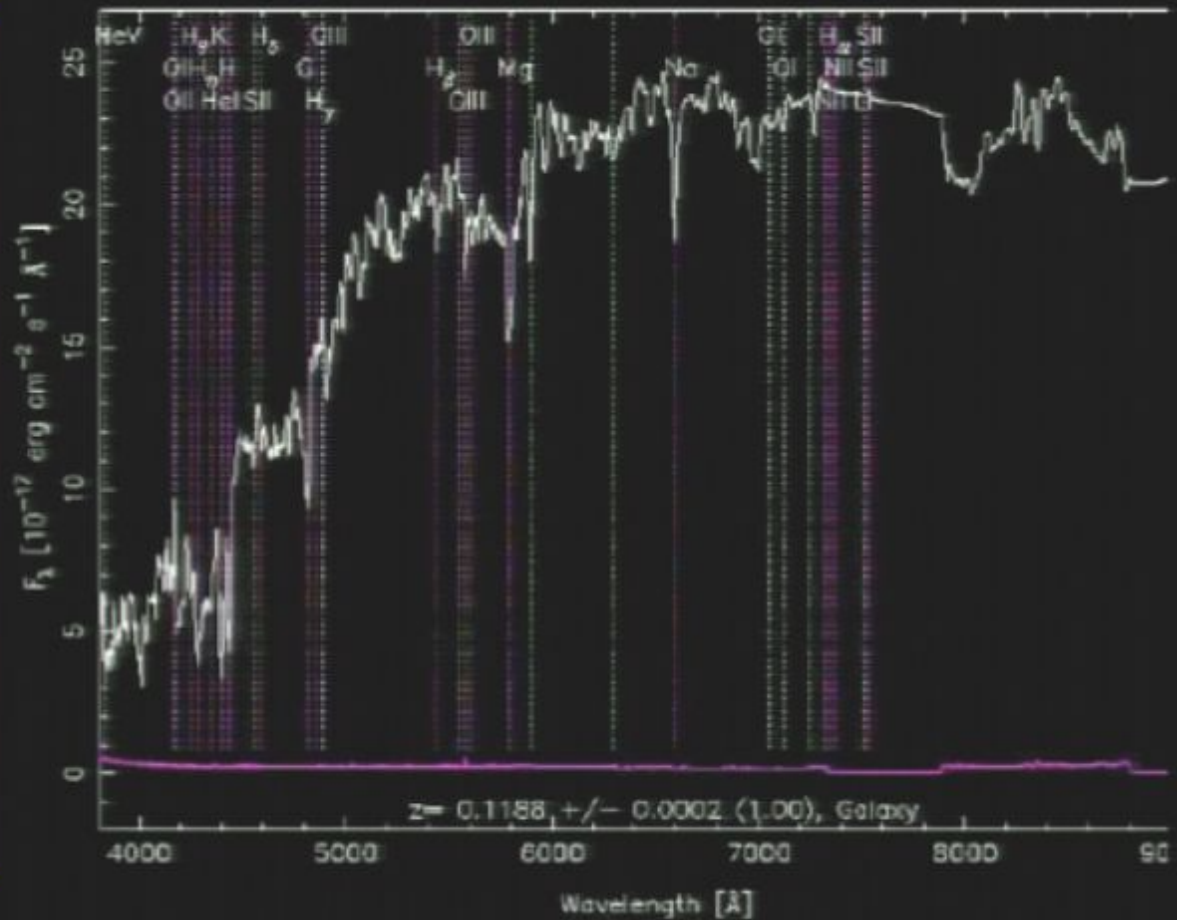
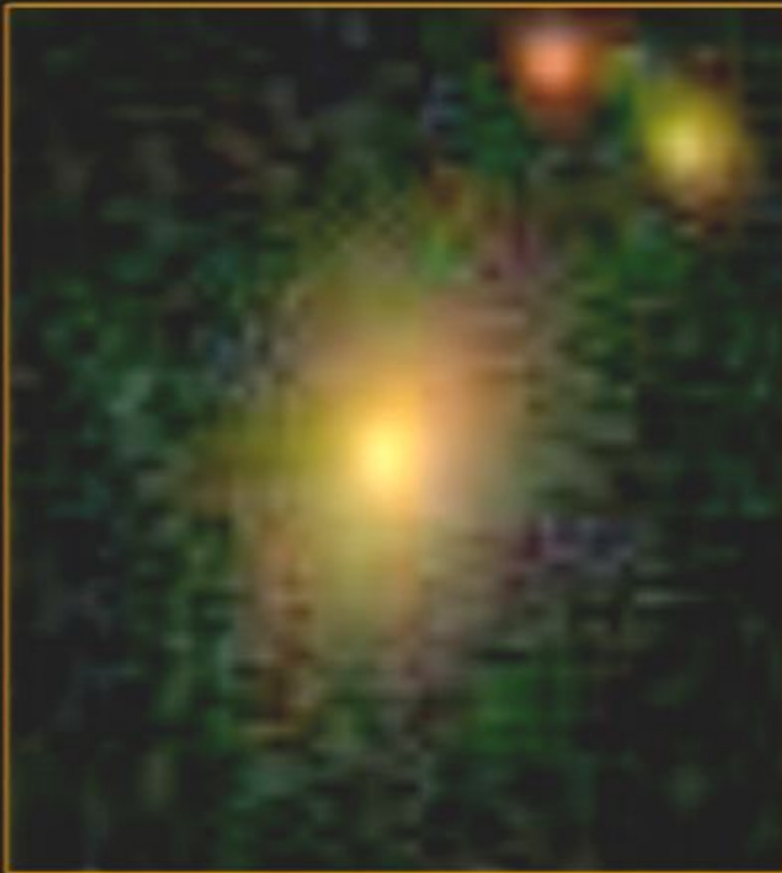


# Blue, star-forming galaxies



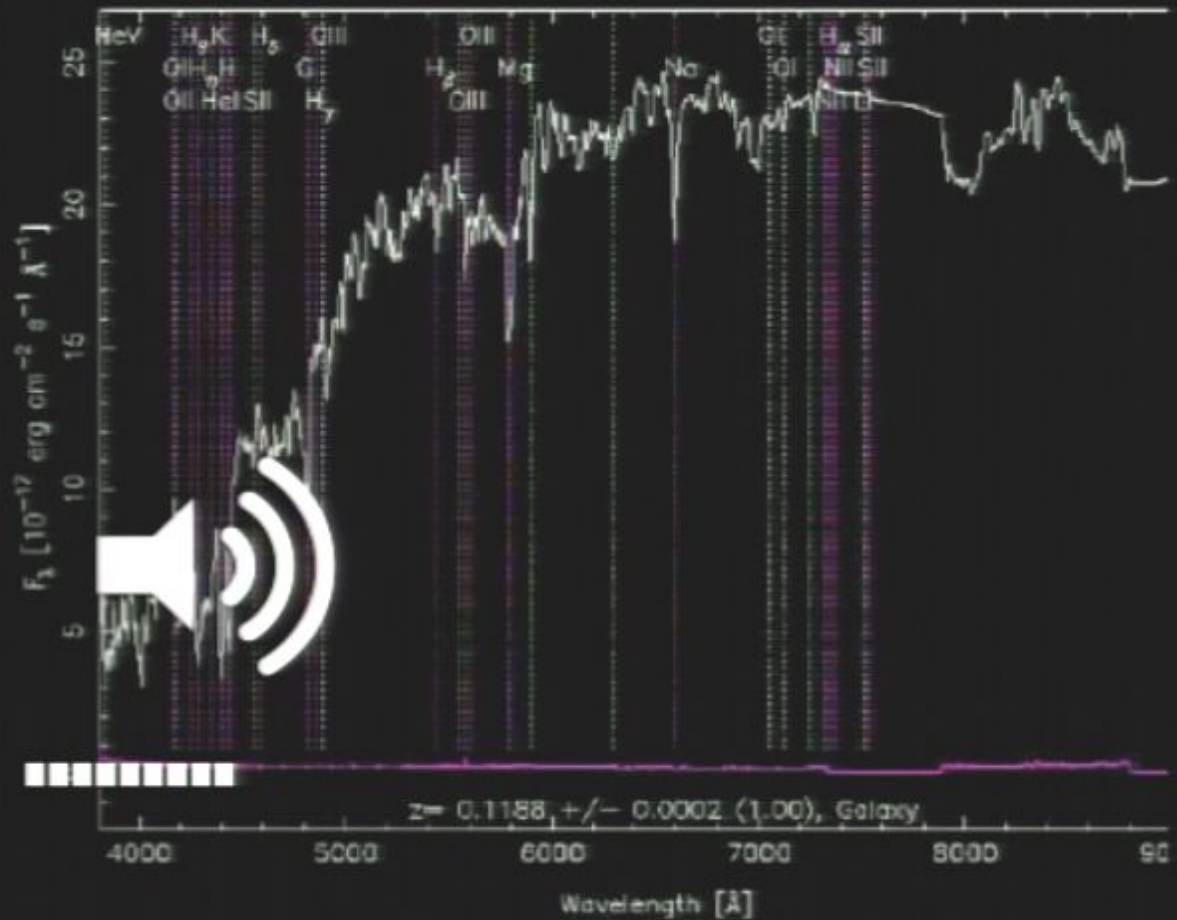
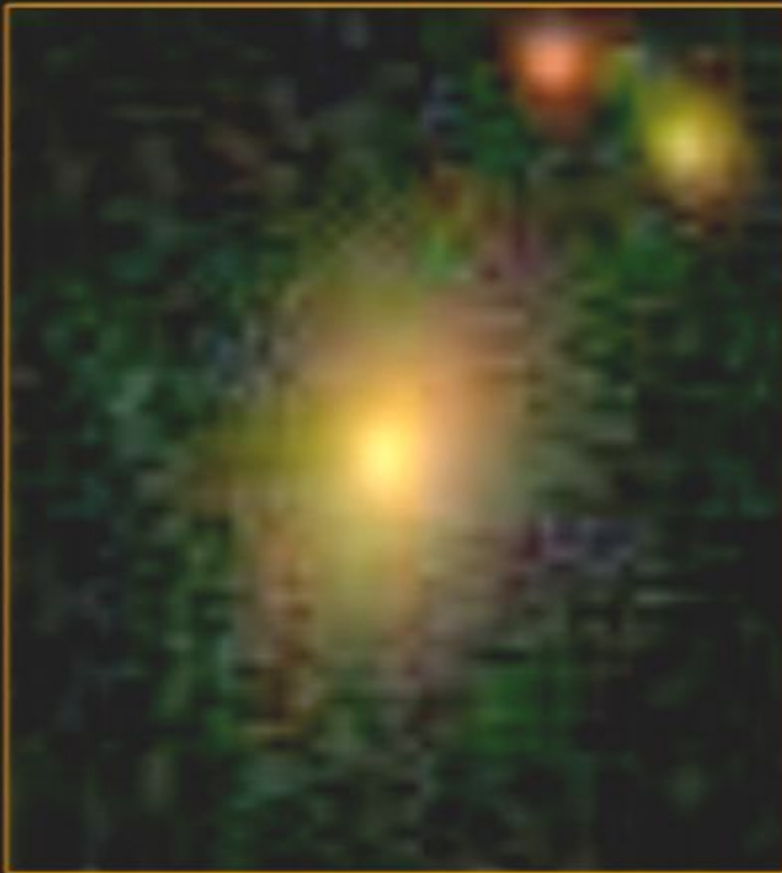
- Used by WiggleZ, preferred for 8m class

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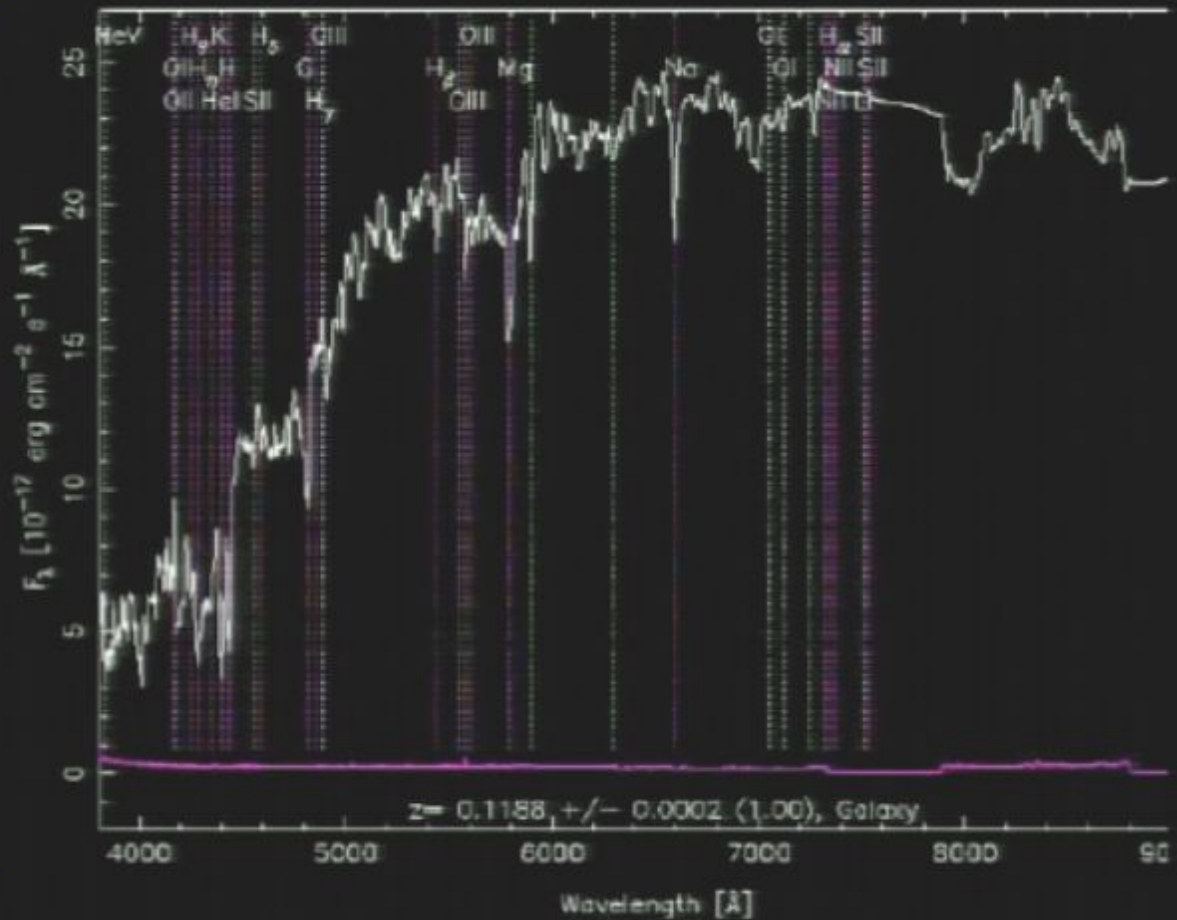
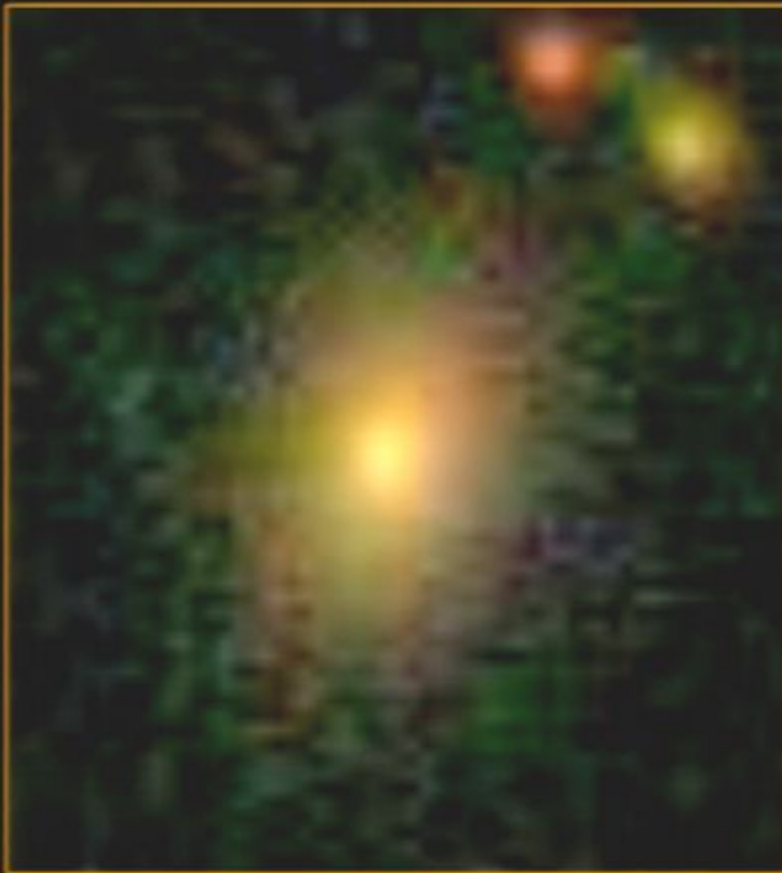
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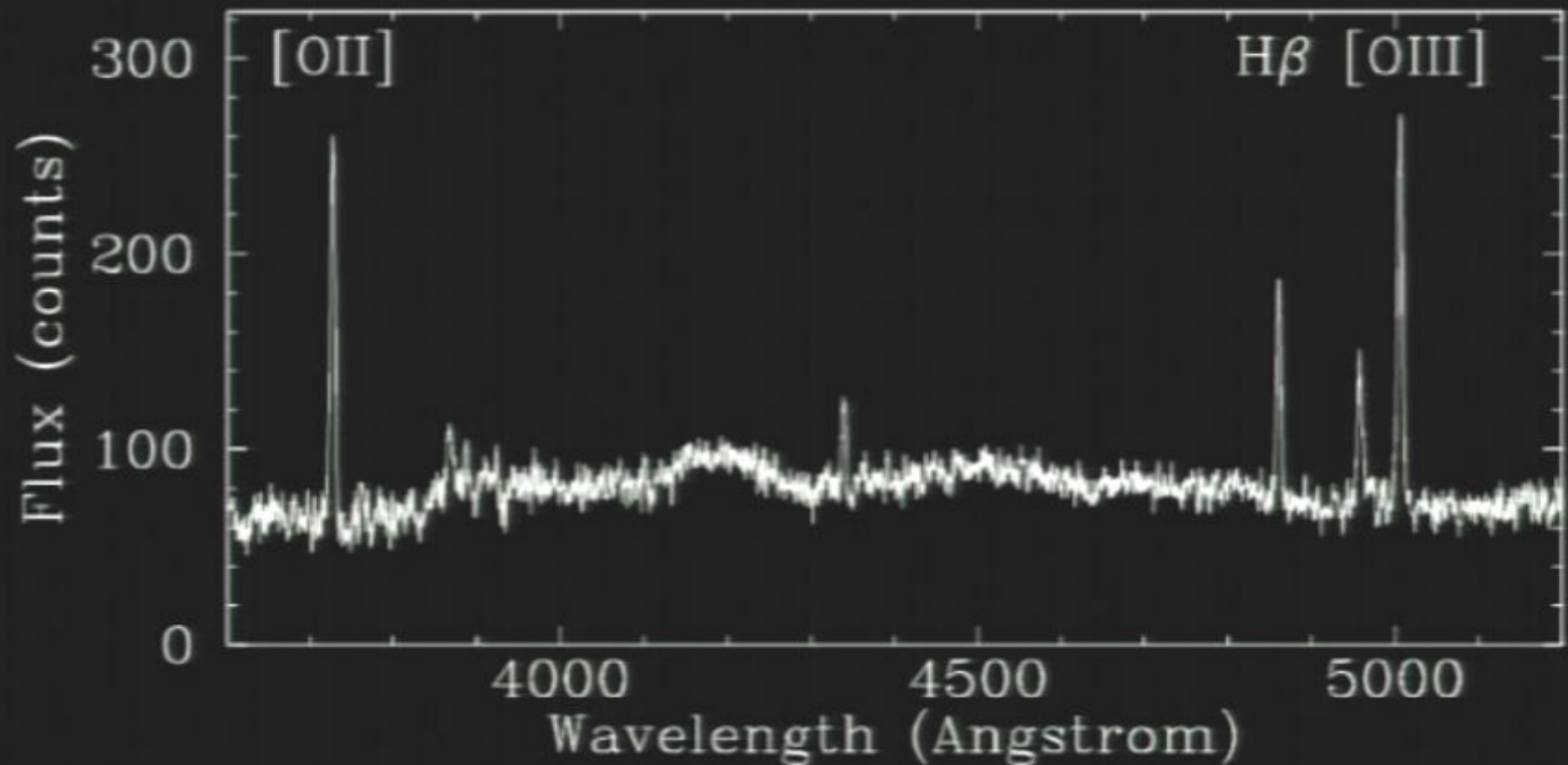
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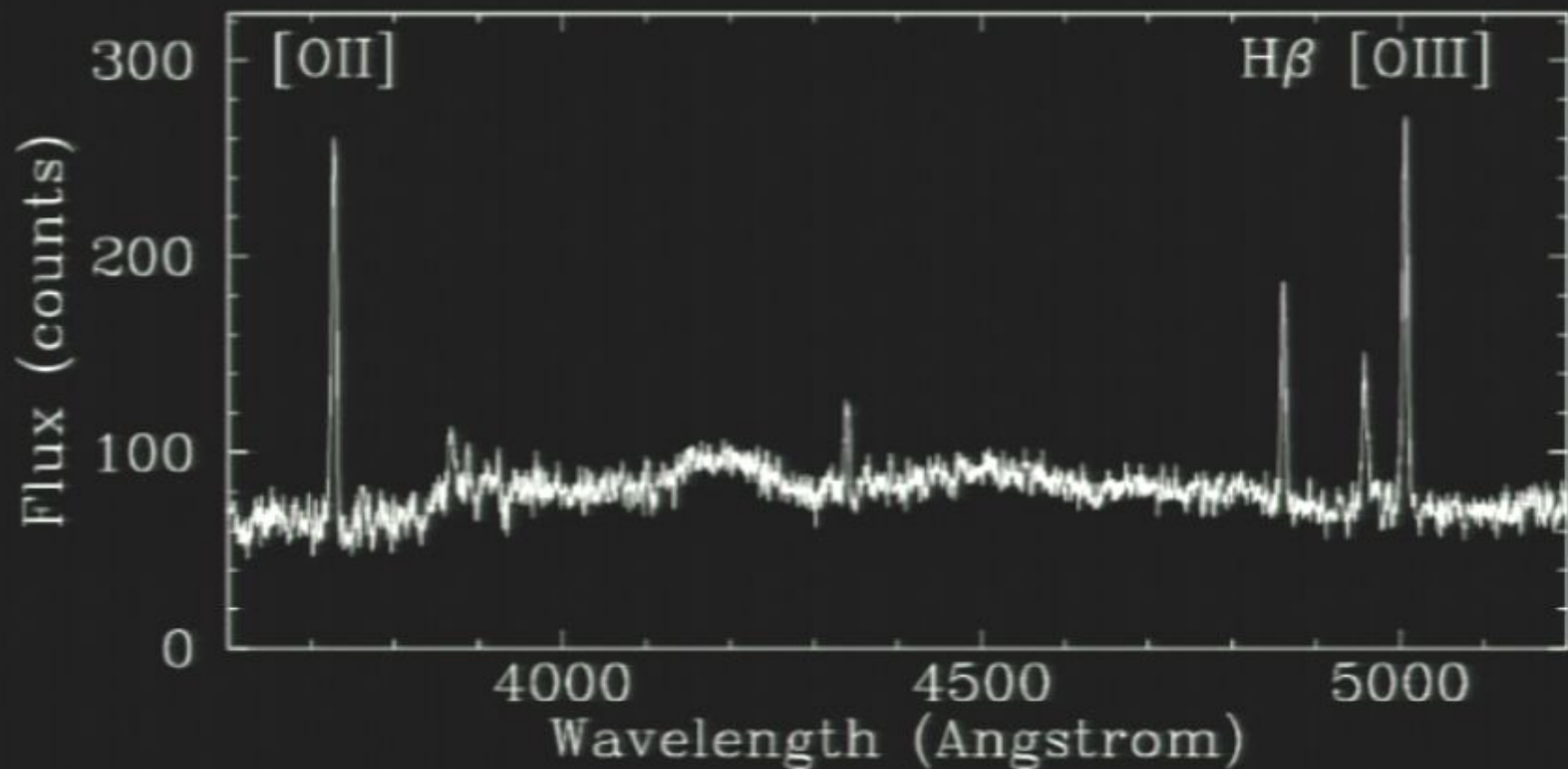


# Blue, star-forming galaxies



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# Neutral Hydrogen (HI) and Lyman Alpha Forest

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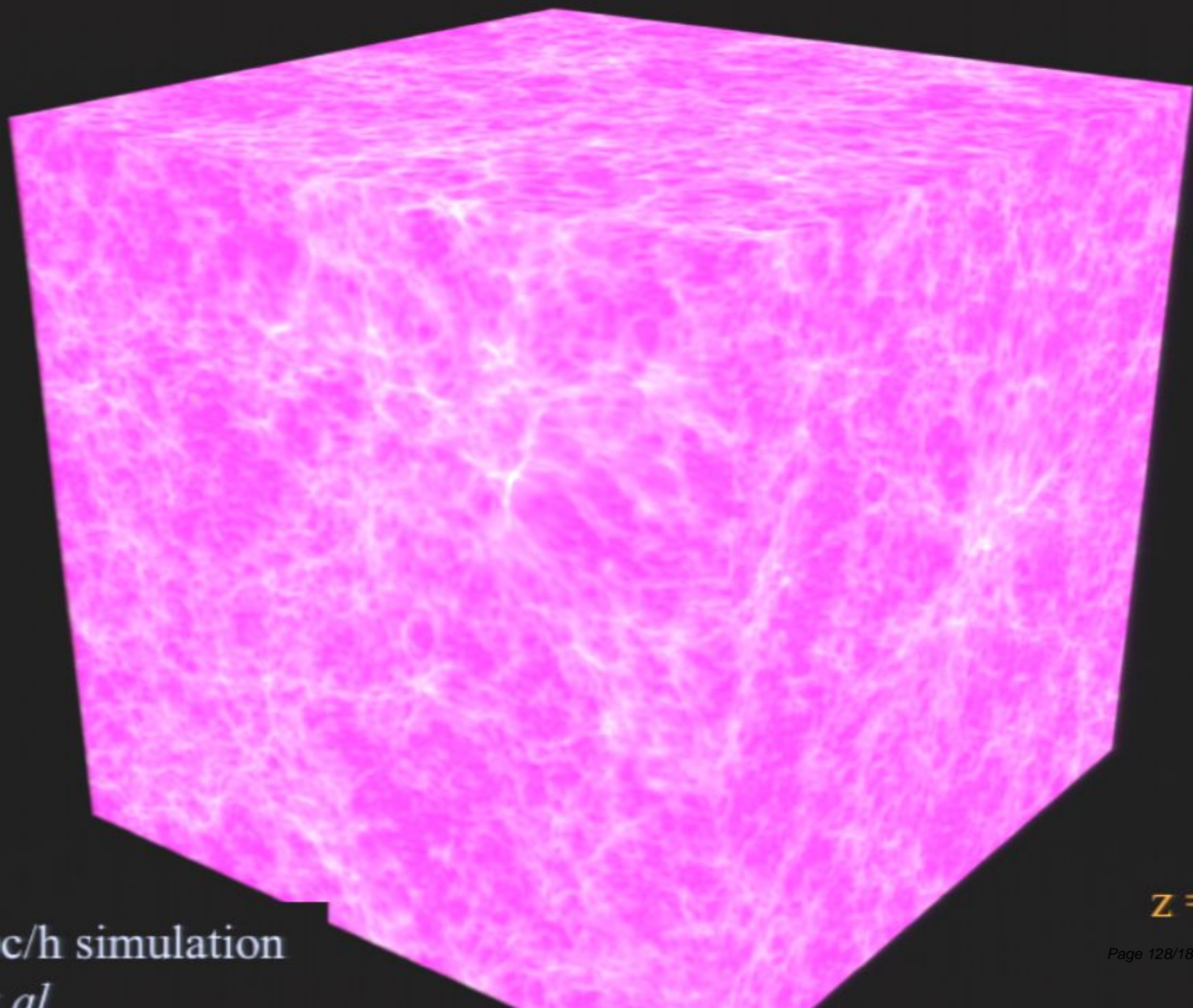


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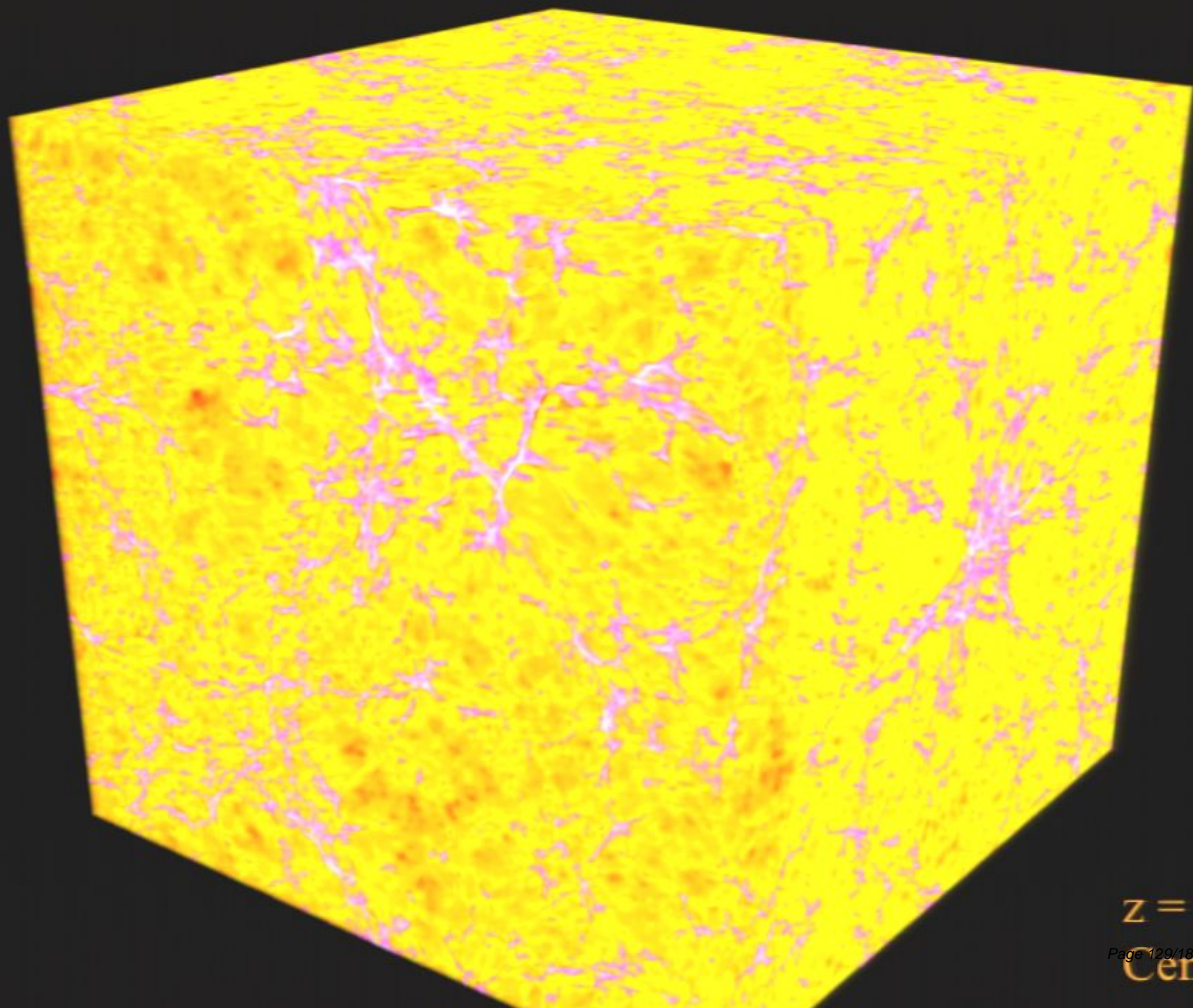
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- Neutral hydrogen absorbs at the Lyman Alpha frequency ( $\sim 126\text{nm}$ ). Works in the optical for  $2 < z < 6$ .
- Both of these can be used to map neutral hydrogen when the nonlinear scale was very small.



25 Mpc/h simulation  
Cen *et al*

$z = 11$

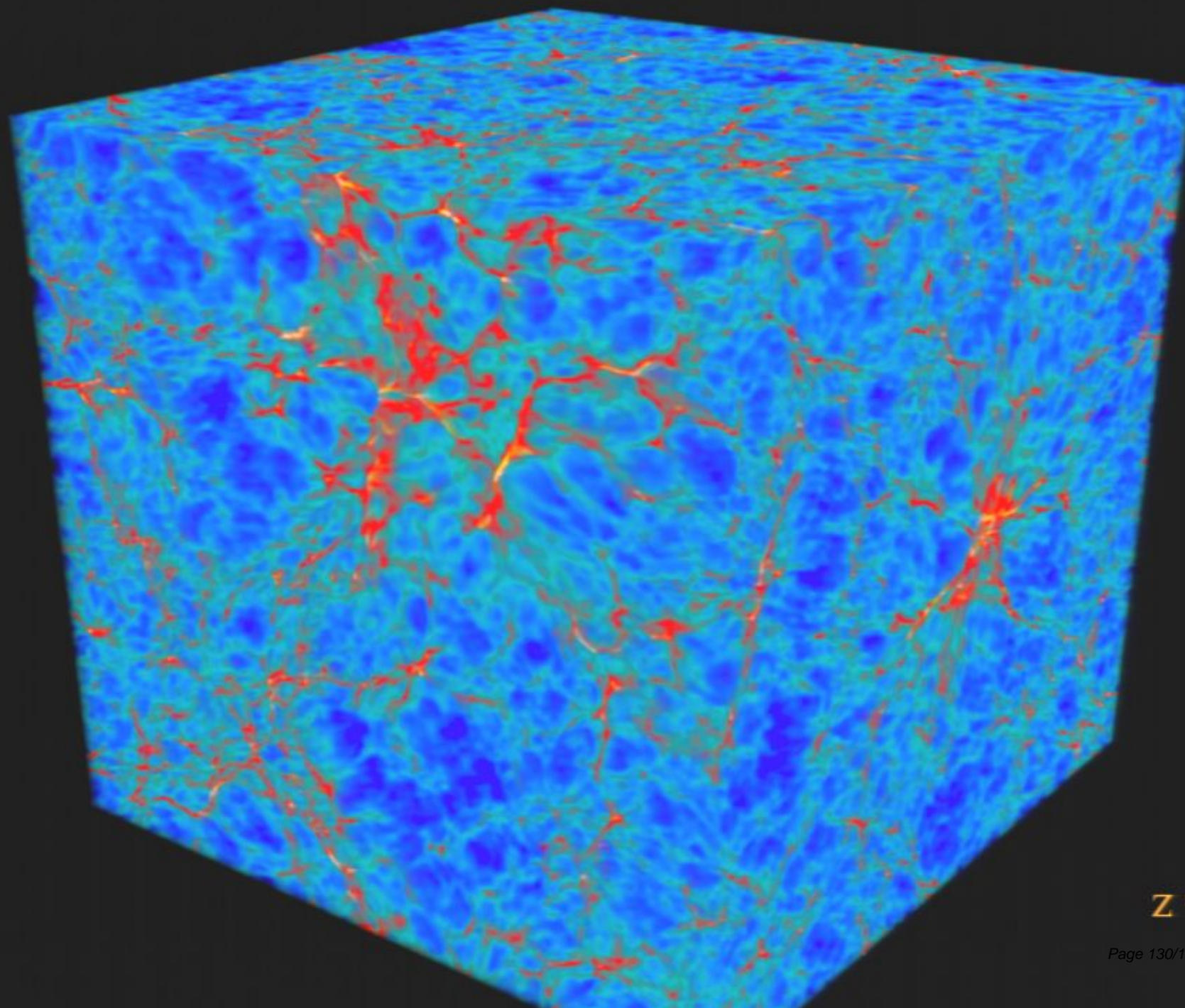




$z = 8.55$

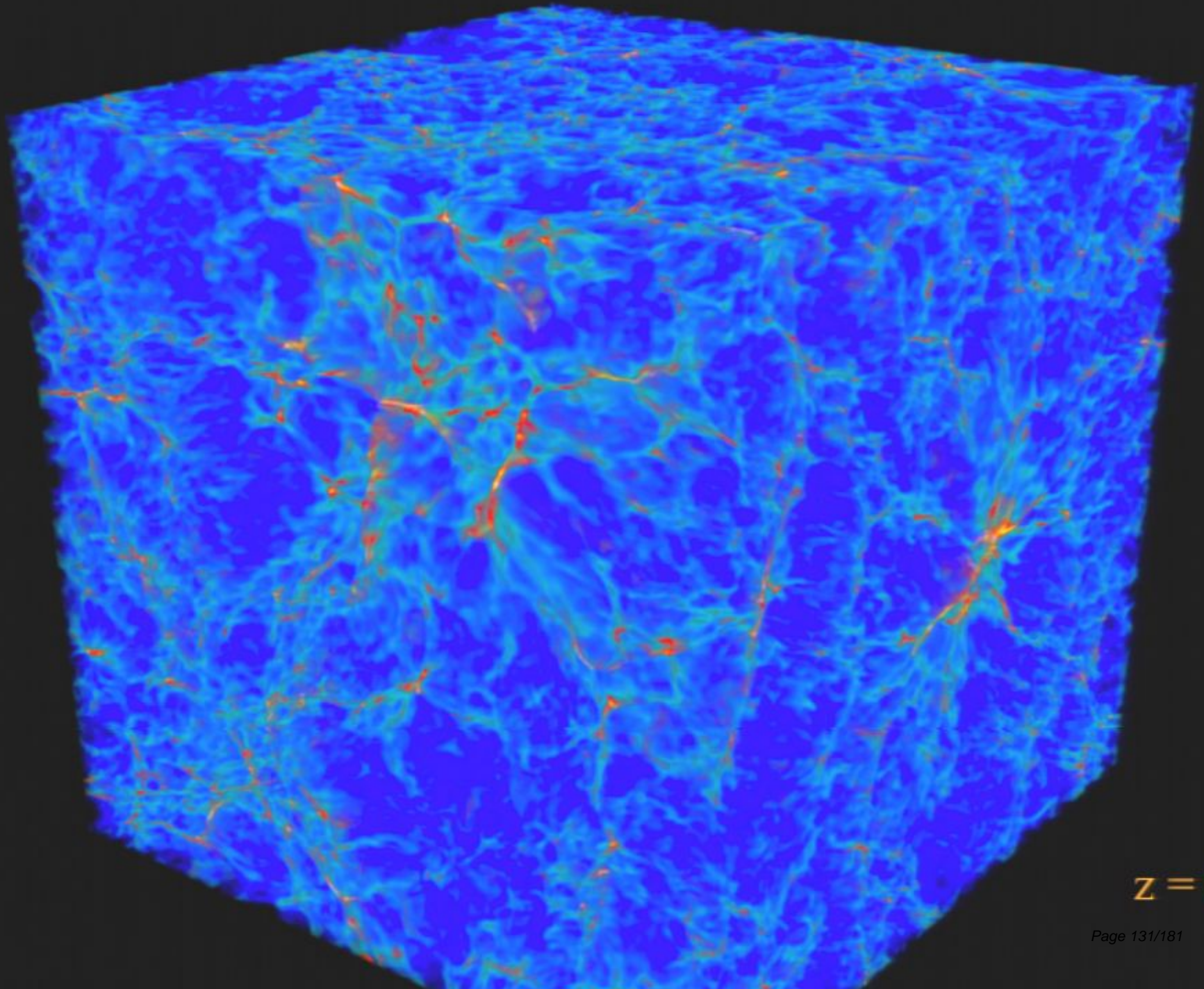
Cen





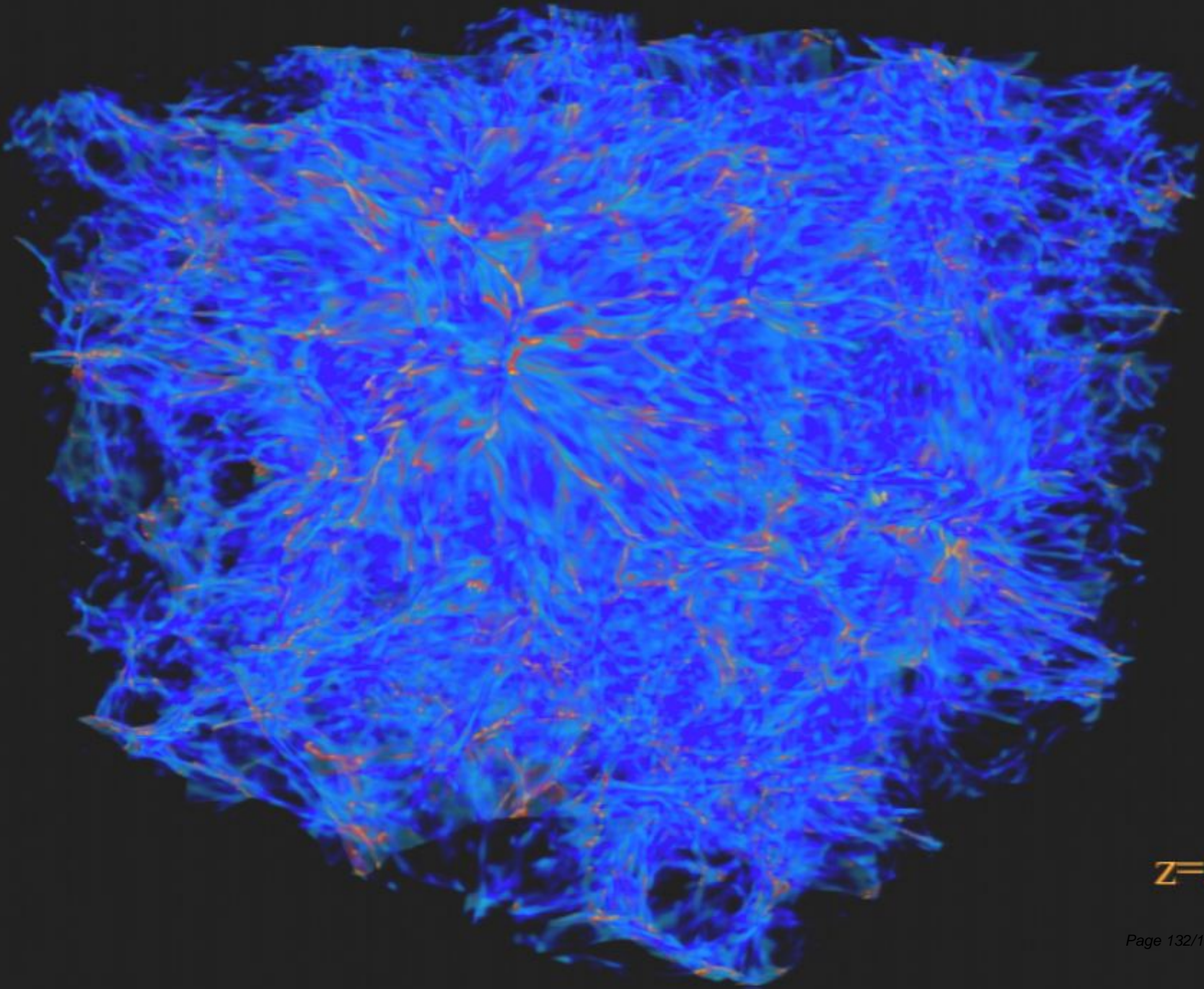
$z = 8.8$





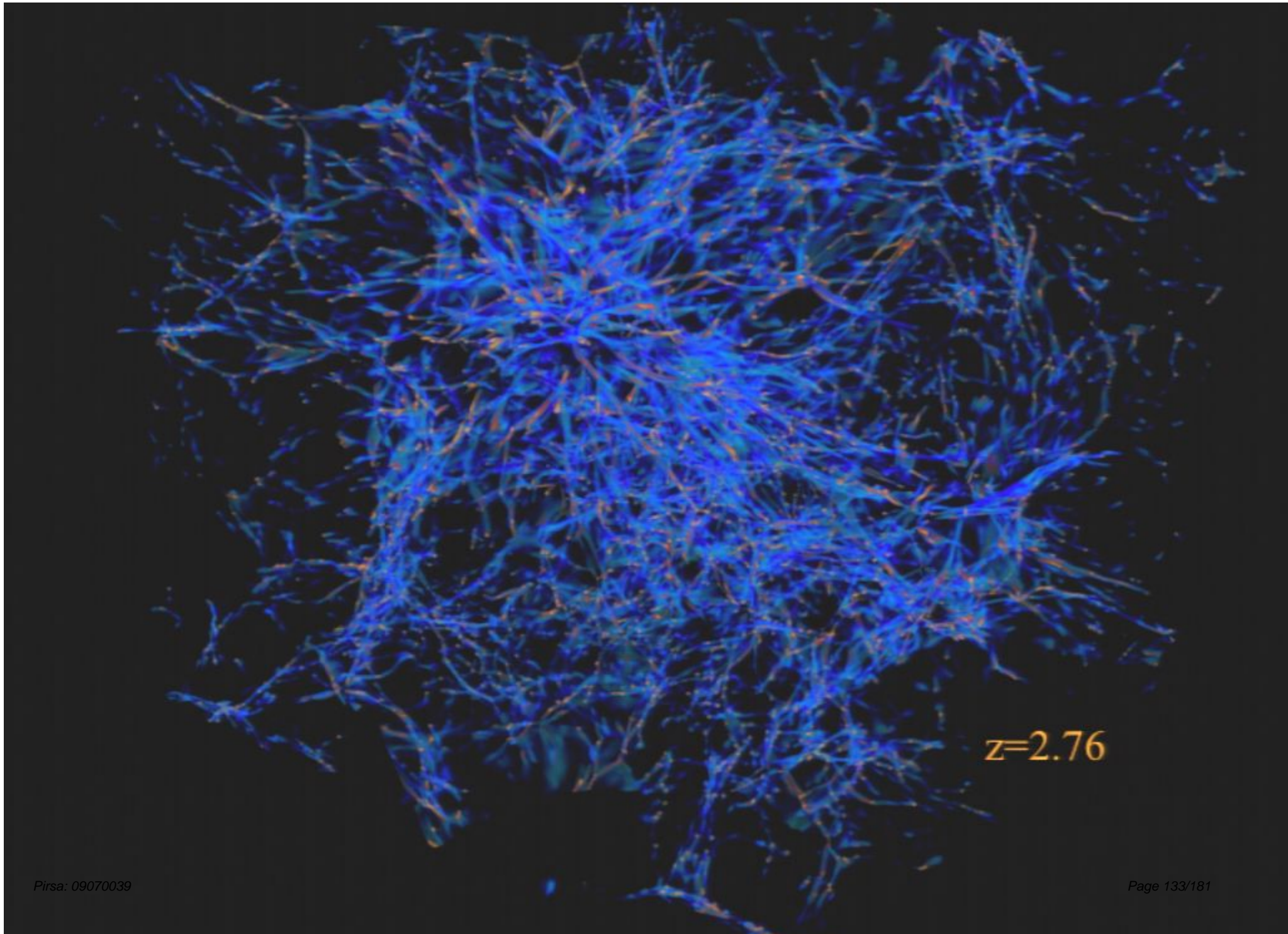
$z = 6.5$





$z=3.5$

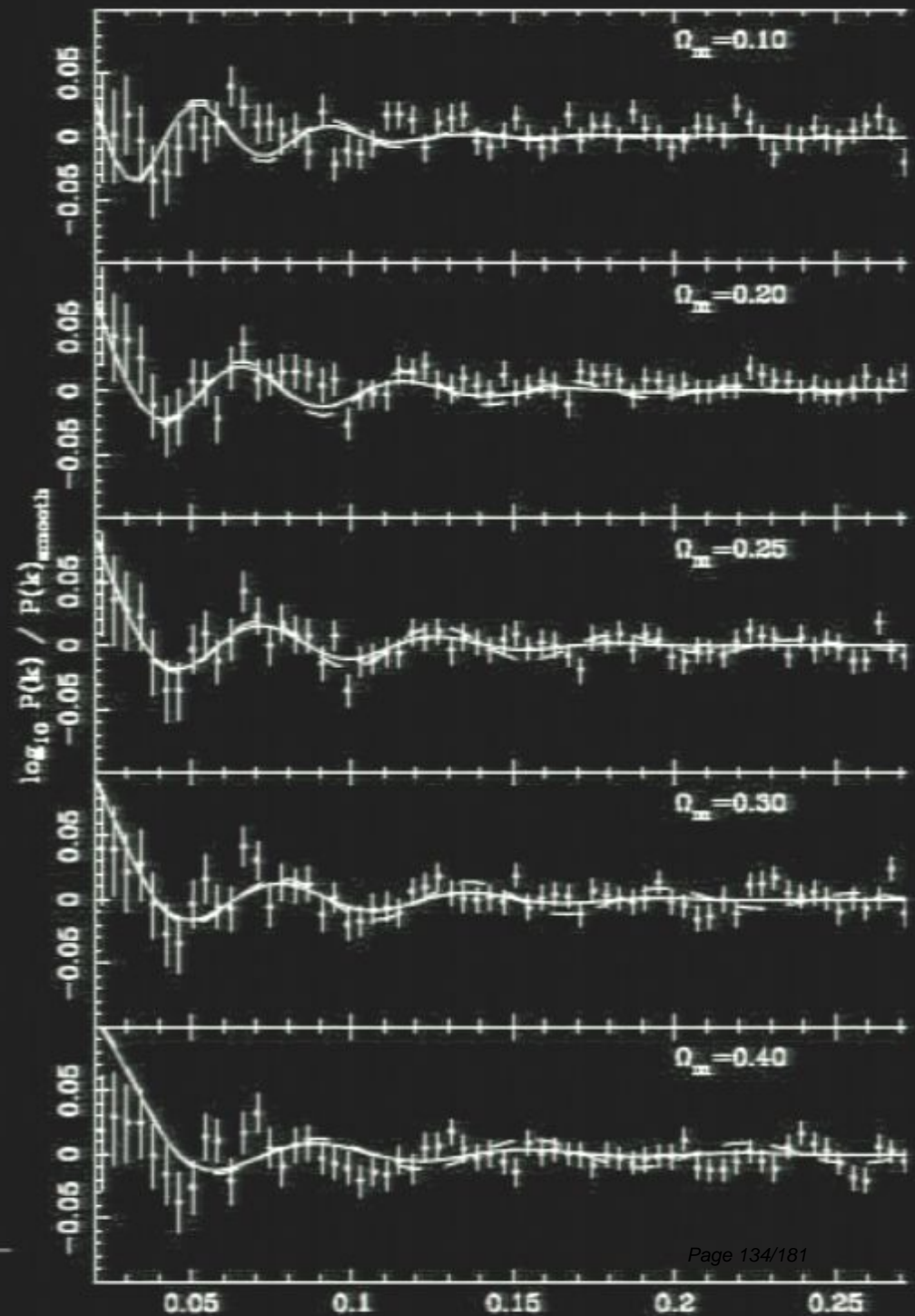




$z=2.76$

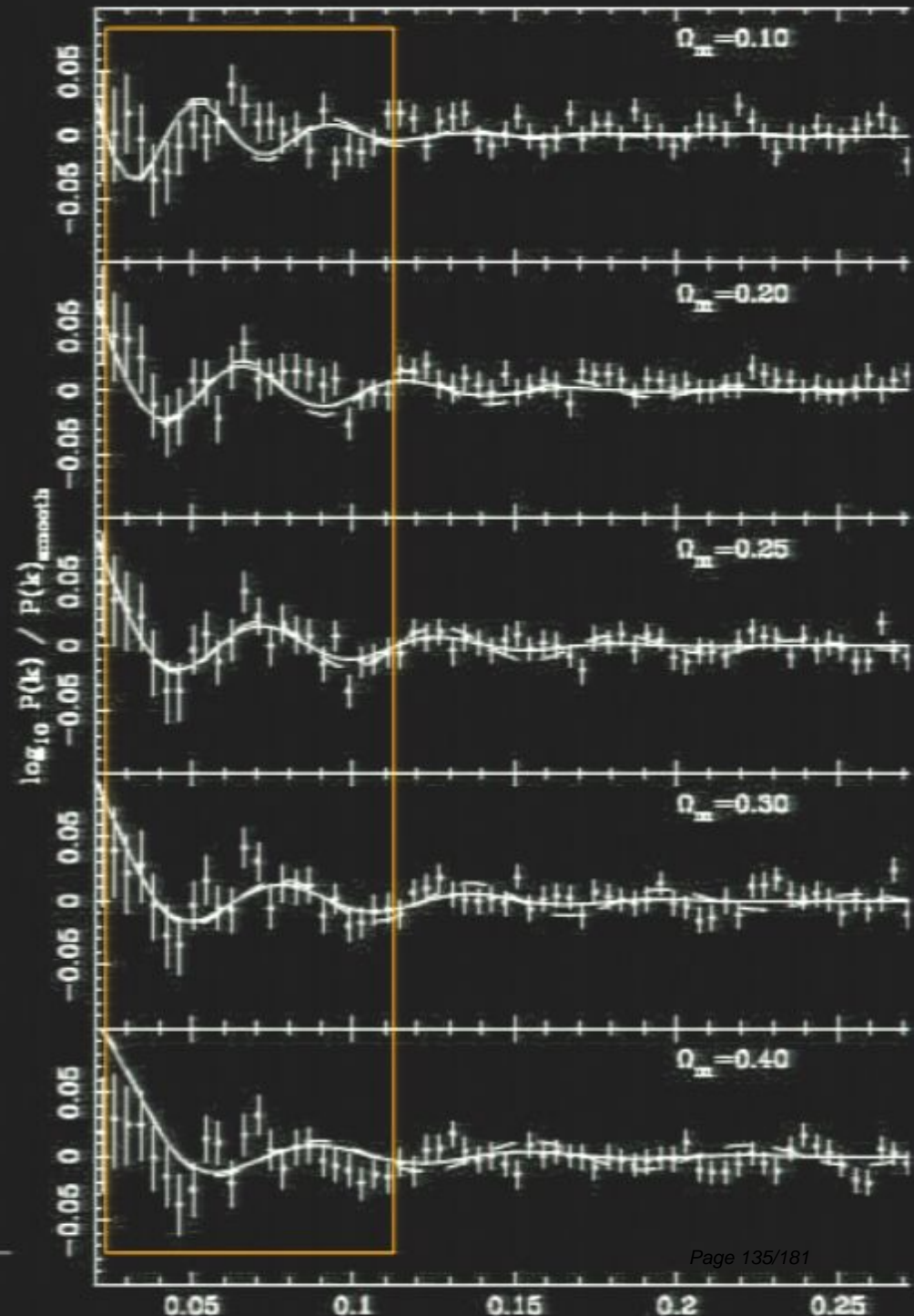


# Current Data



# Current Data

- Just the BAO wiggles
- Produced for SDSS + 2df galaxies combined

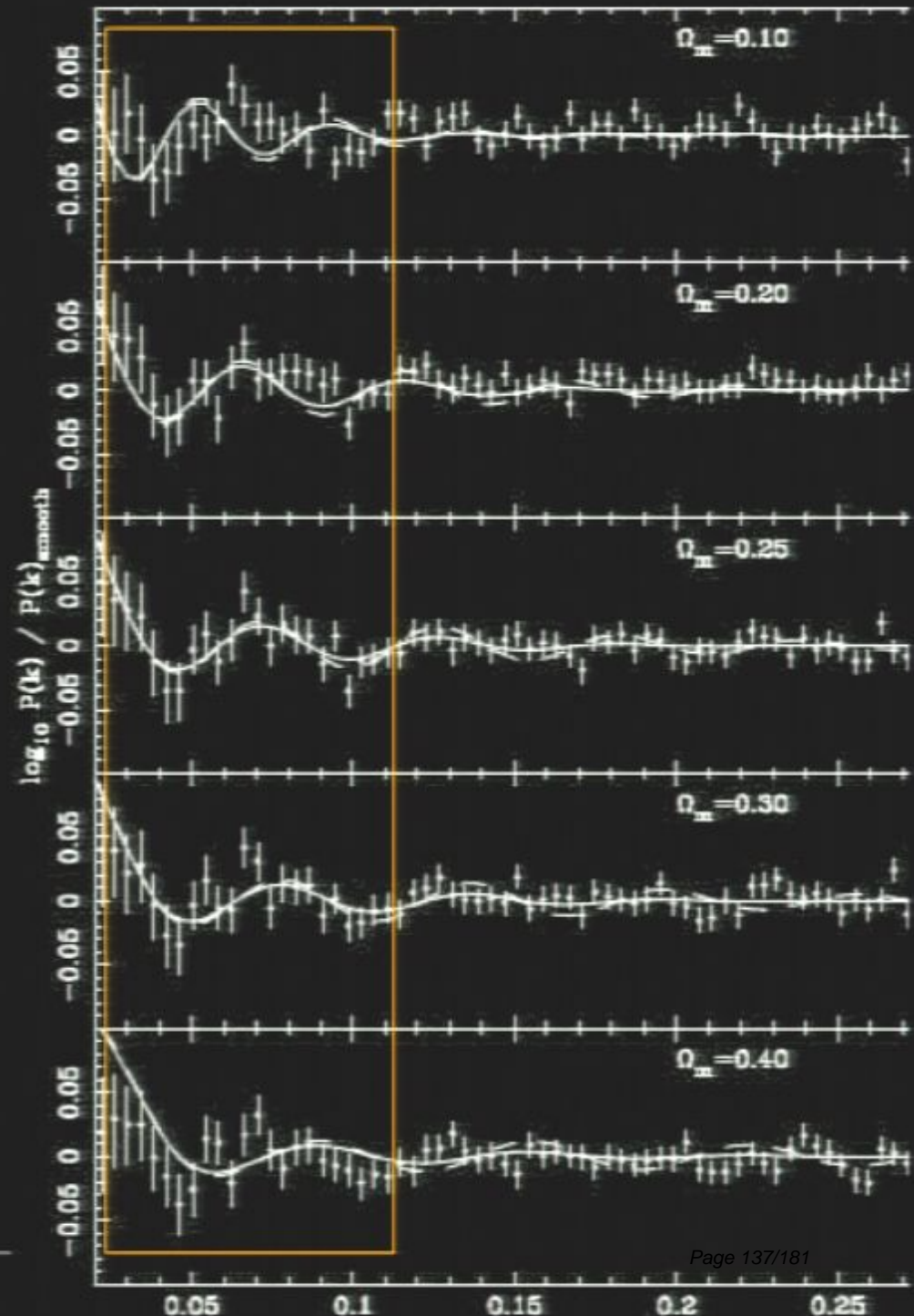


# BAO

have provided the *first* ever  
cosmic distance measure based  
on *linear* physics

# Current Data

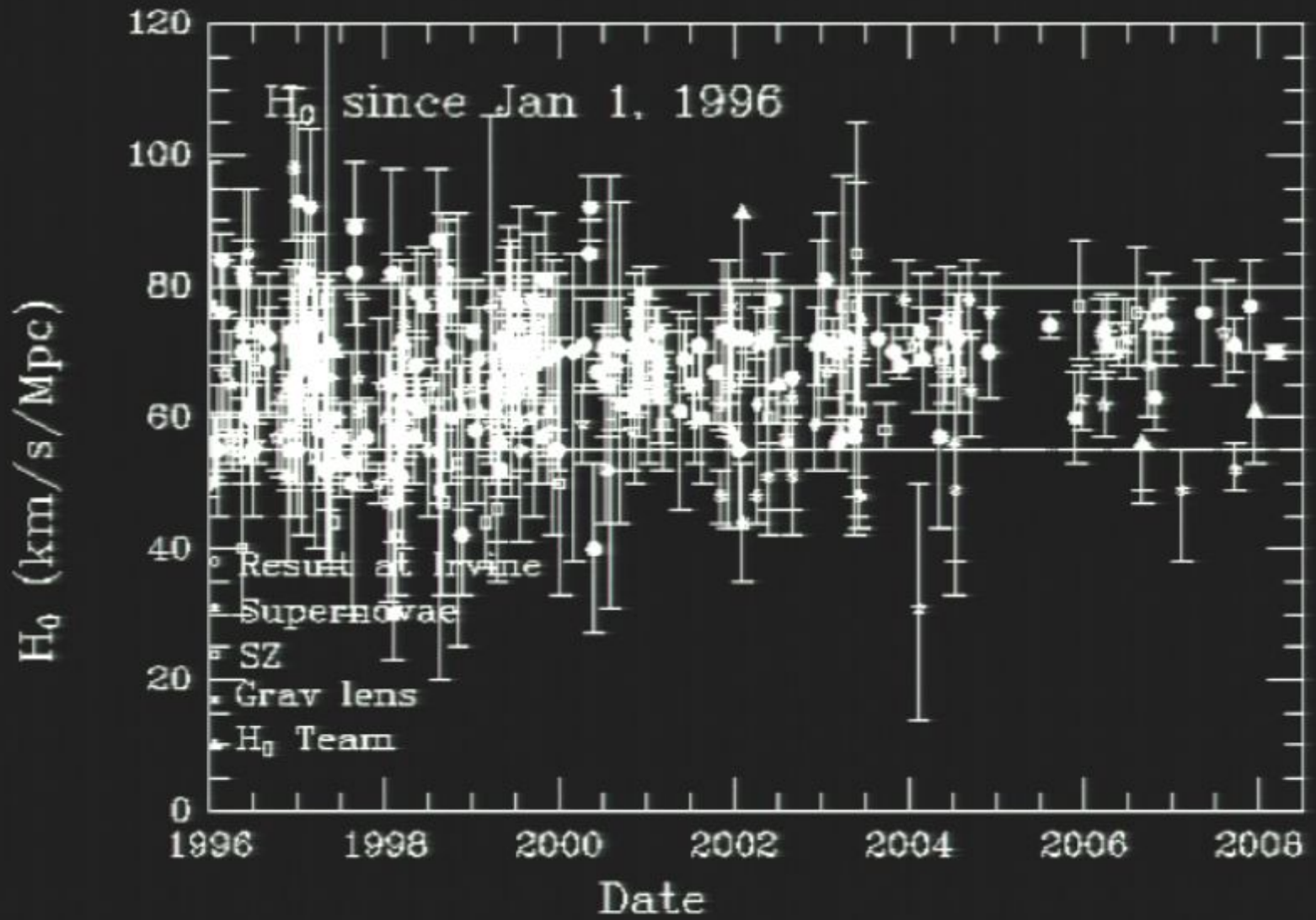
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# BAO

have provided the *first* ever  
cosmic distance measure based  
on *linear* physics



*J. Huchra*

# Complications

A man falls from a window and on the way down someone asks... ``*how's it going*''?

The man replies: "*So far so good!*"

# Complications

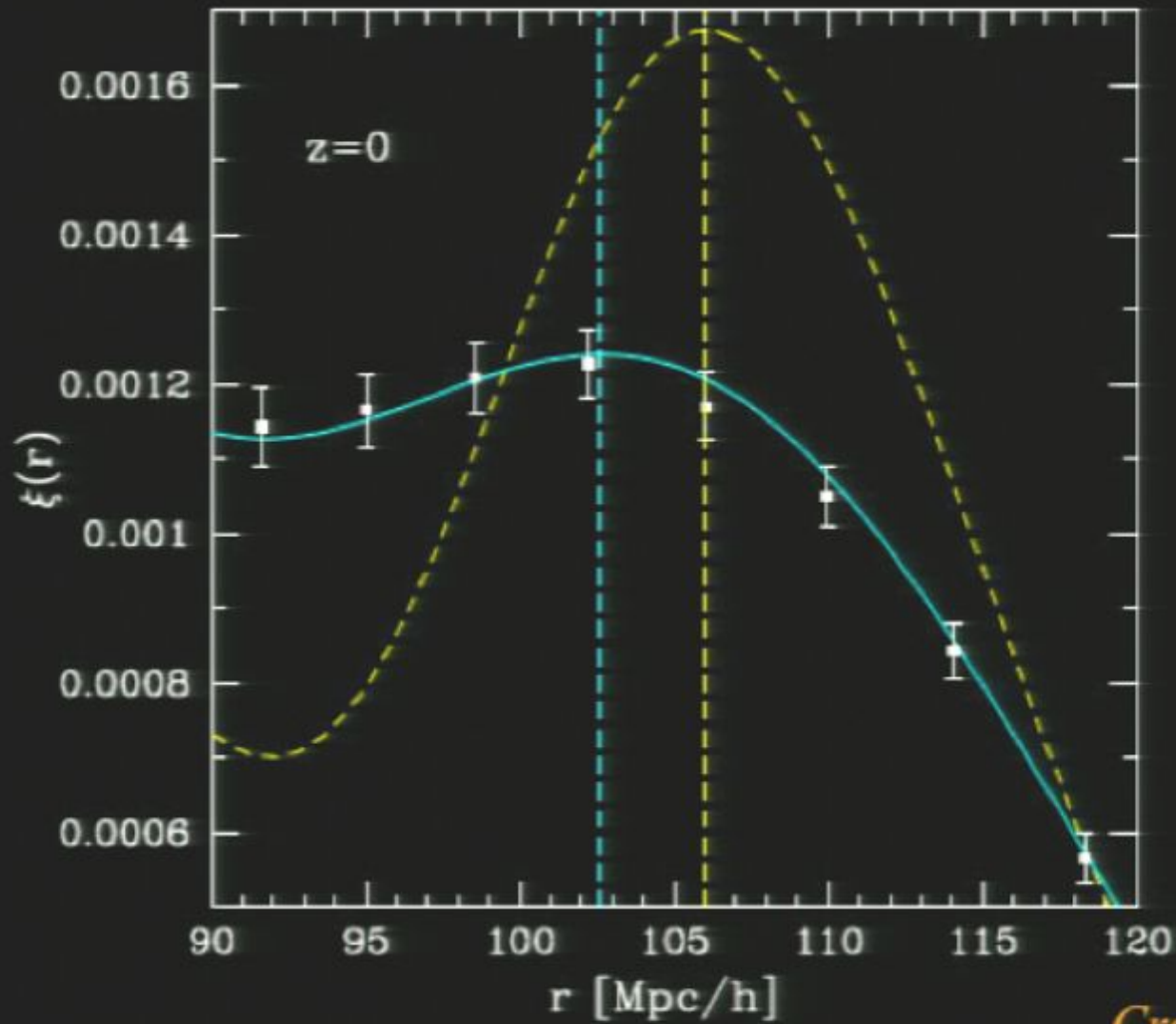
A man falls from a window and on the way down someone asks... ``*how's it going*''?

The man replies: "*So far so good!*"

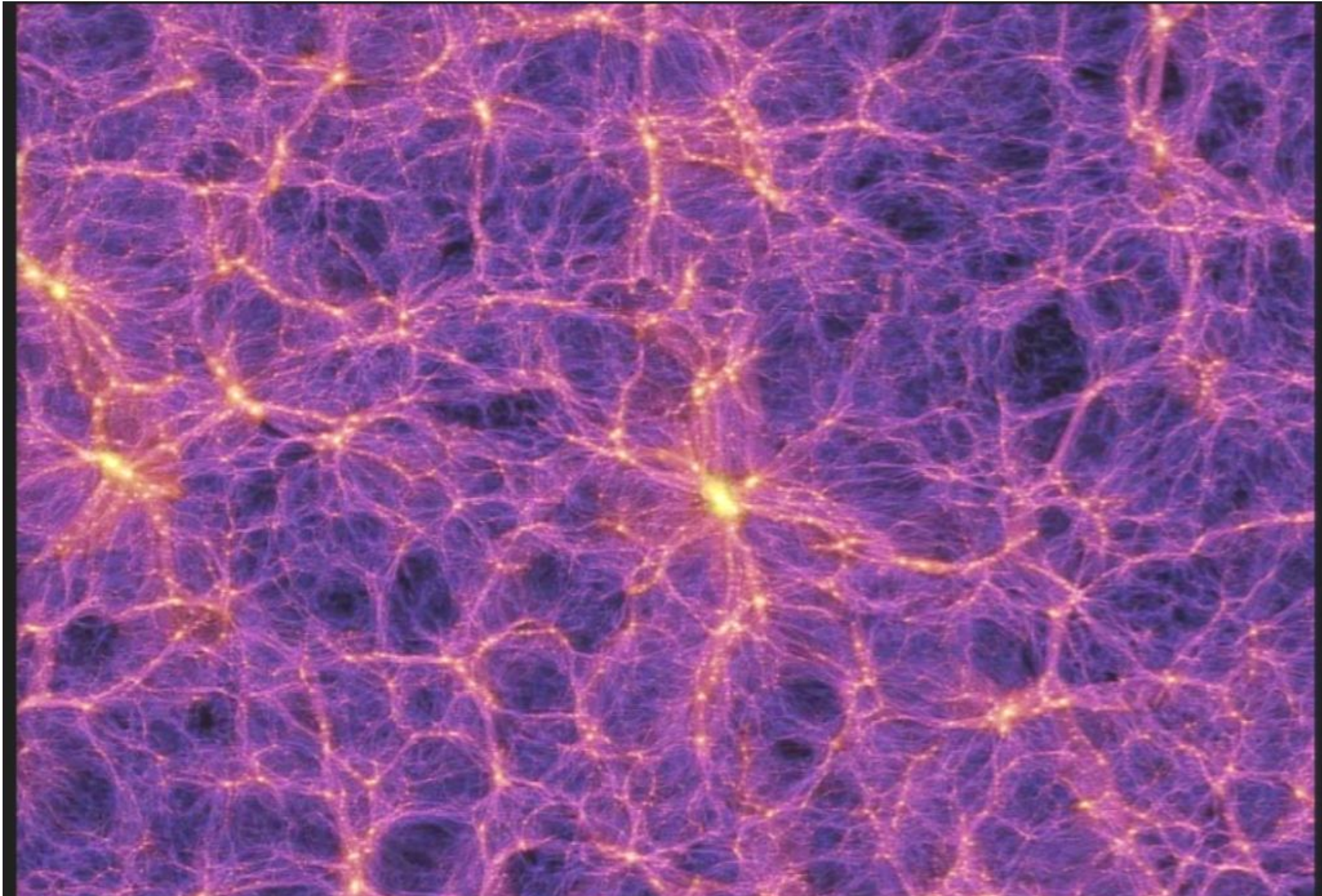
Cosmological Methods are similar – they have a free-fall zone and a hard, systematic “floor” where things get tough



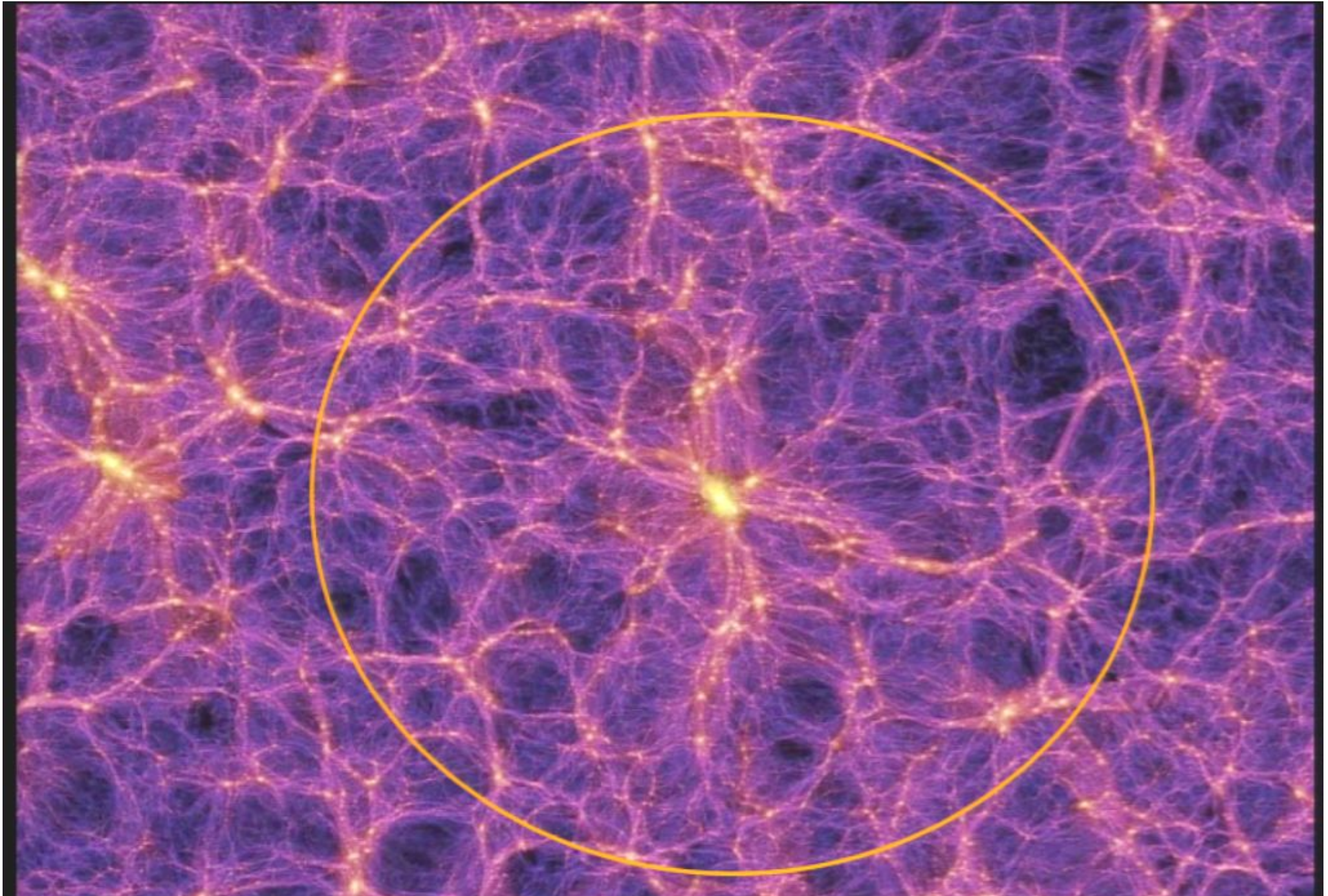
# BAO and Nonlinearity



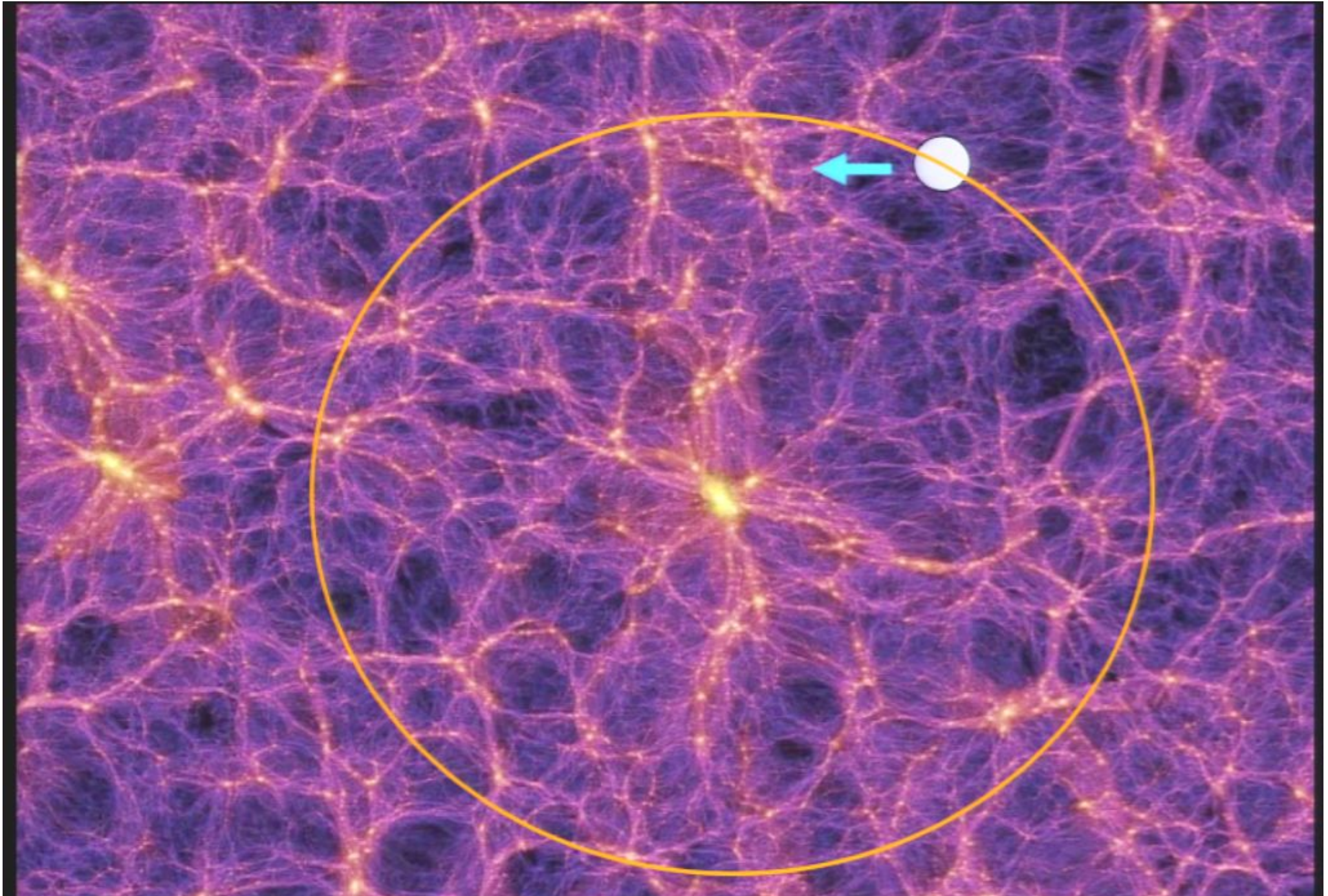
*Crocce et al*



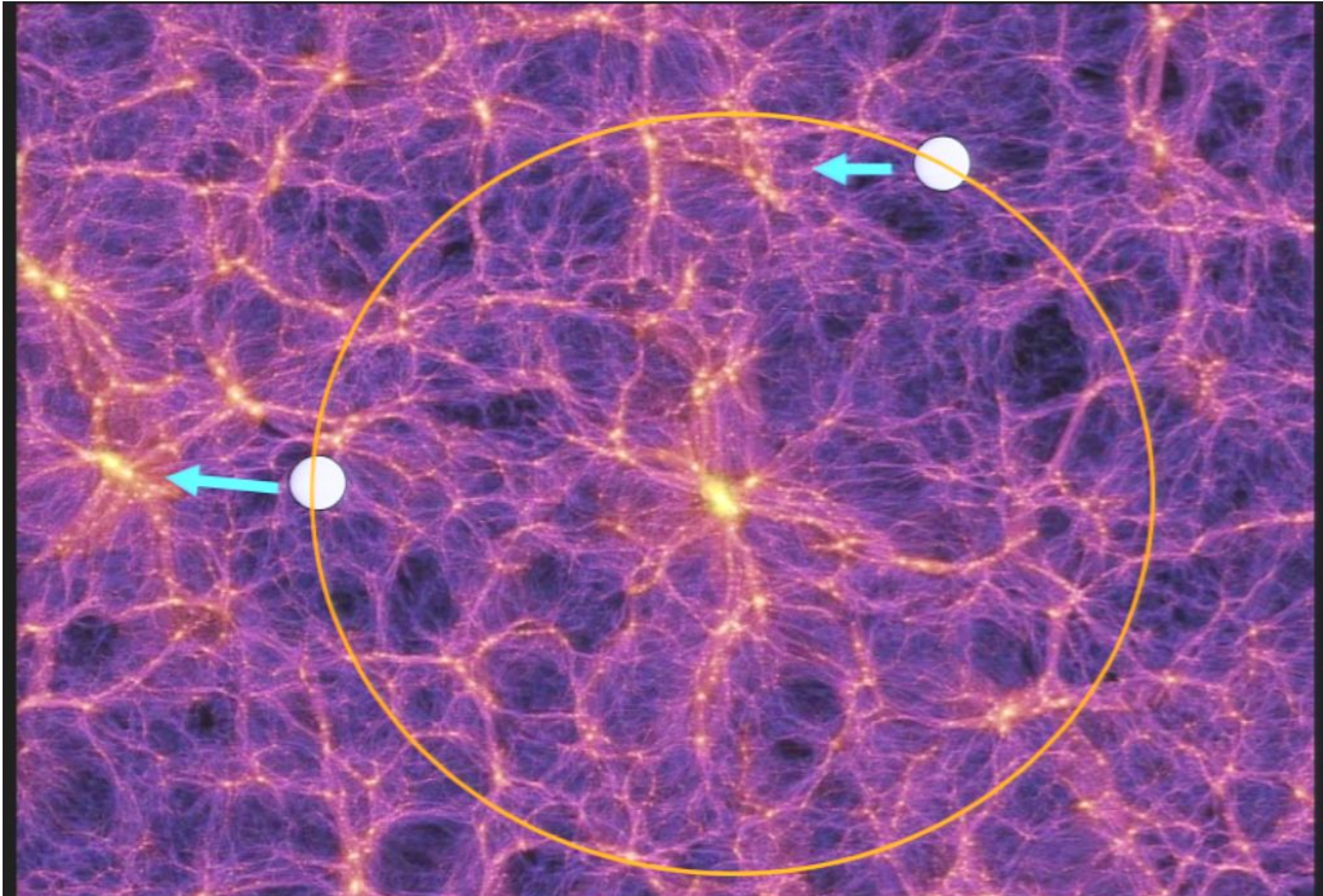




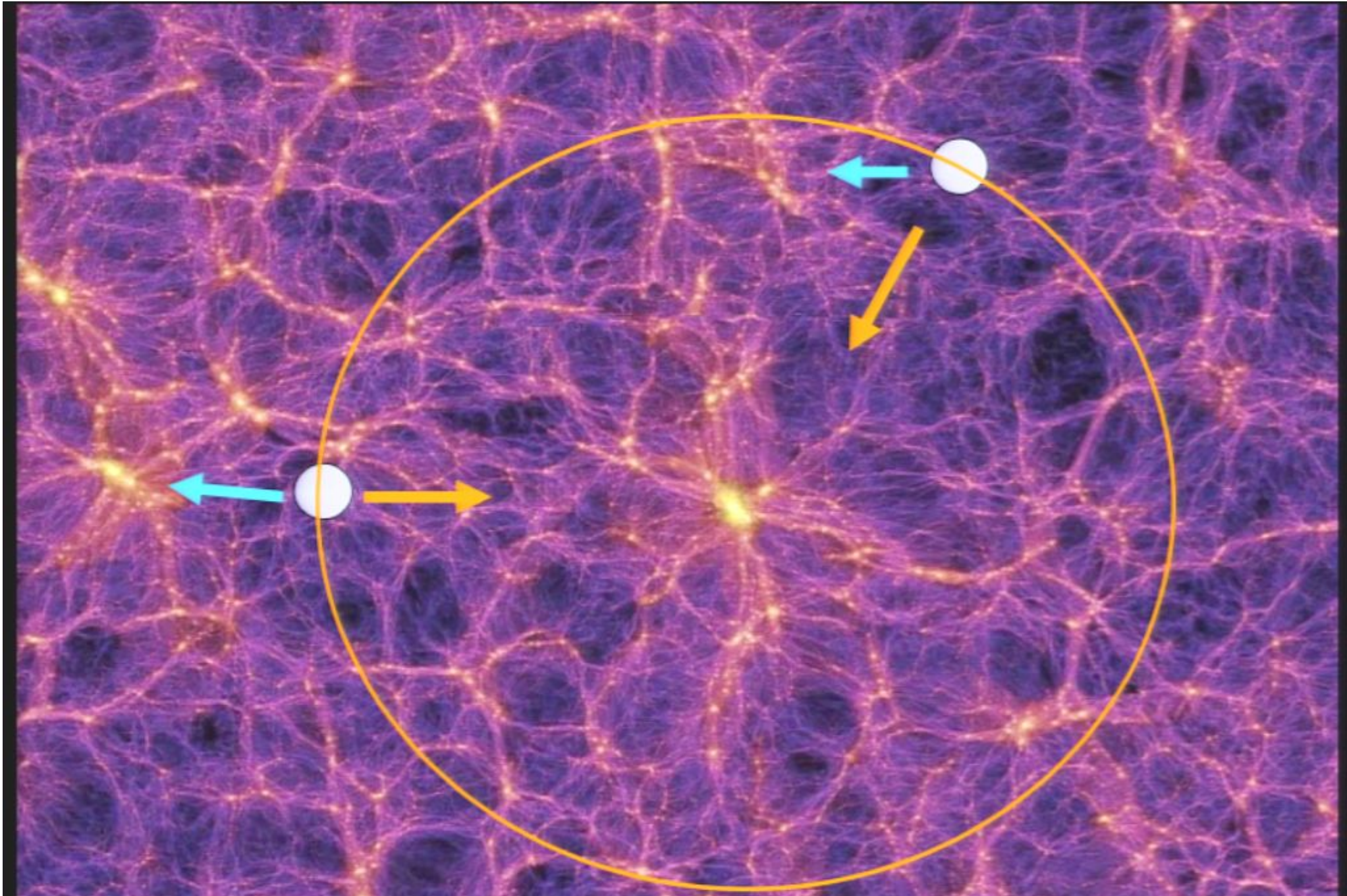




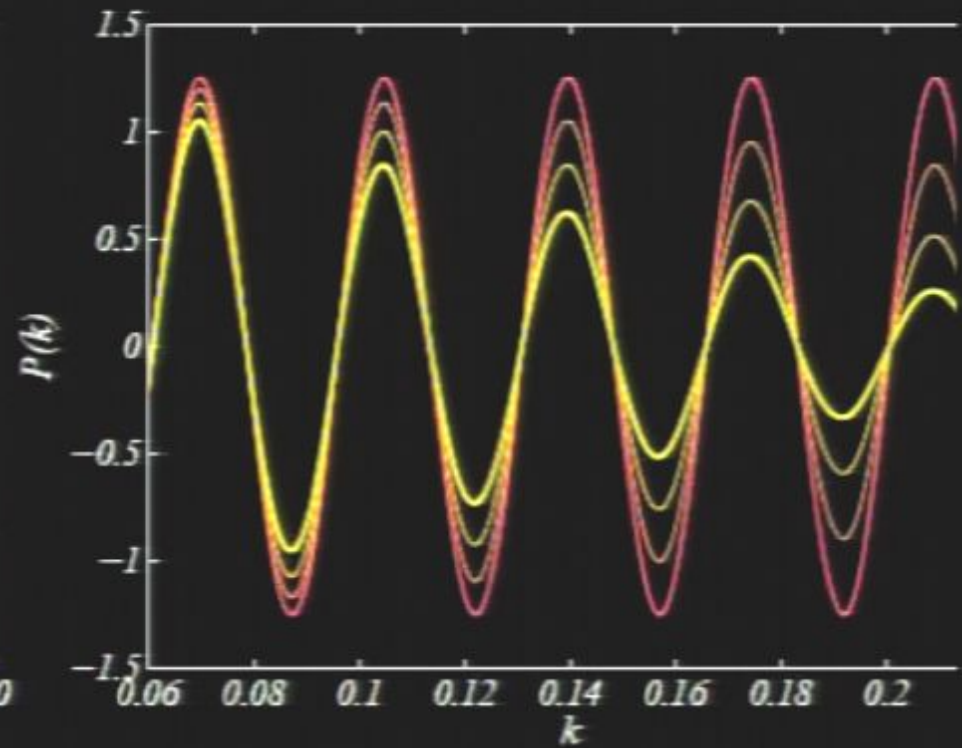
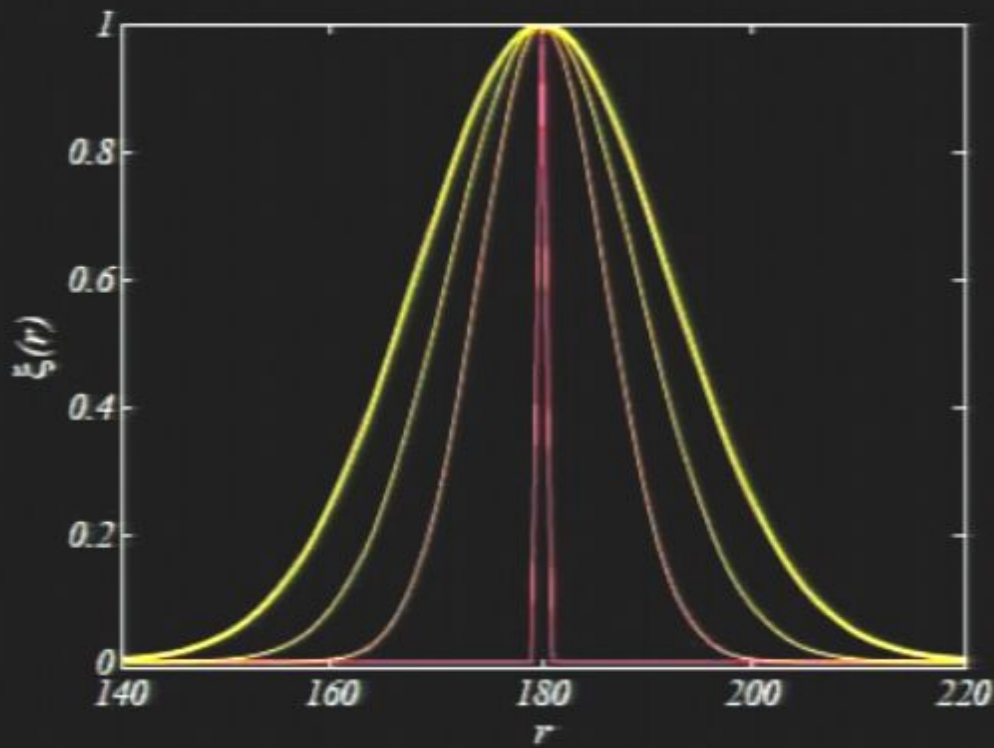




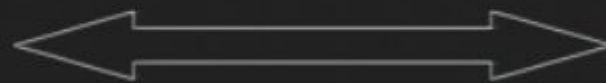




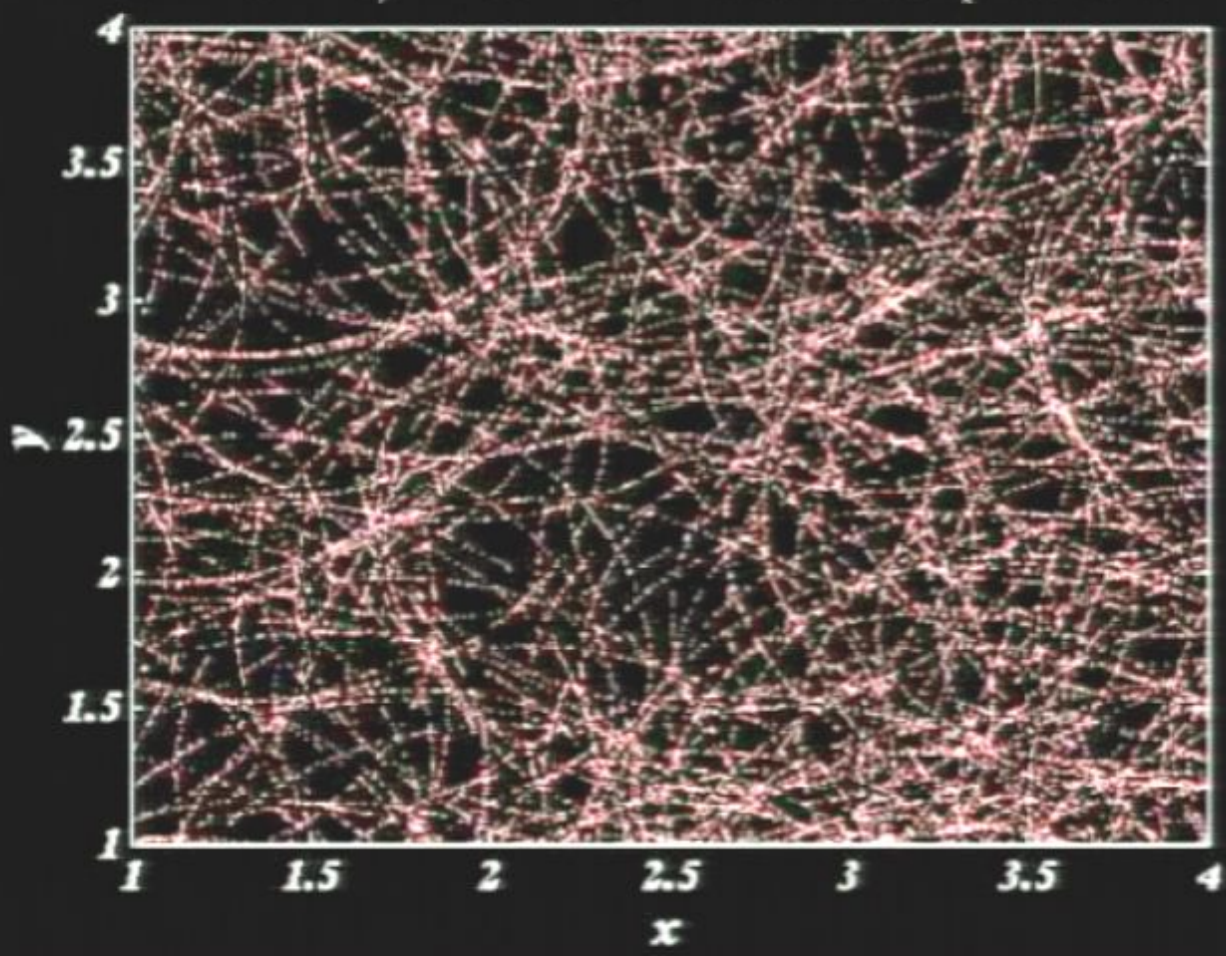
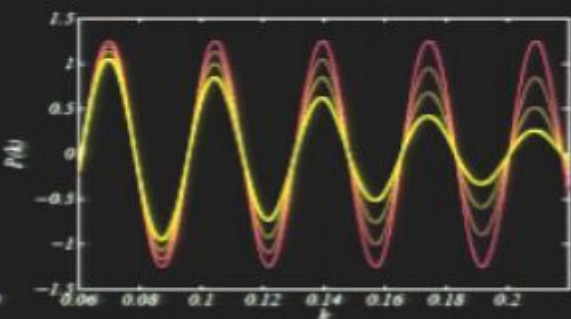
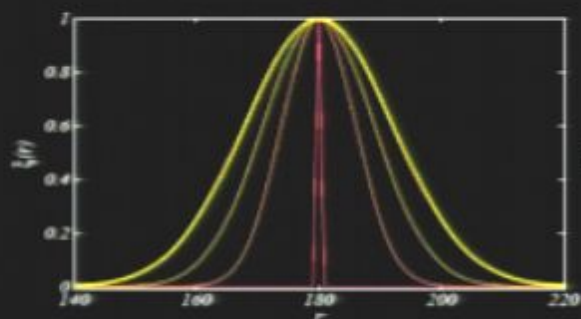




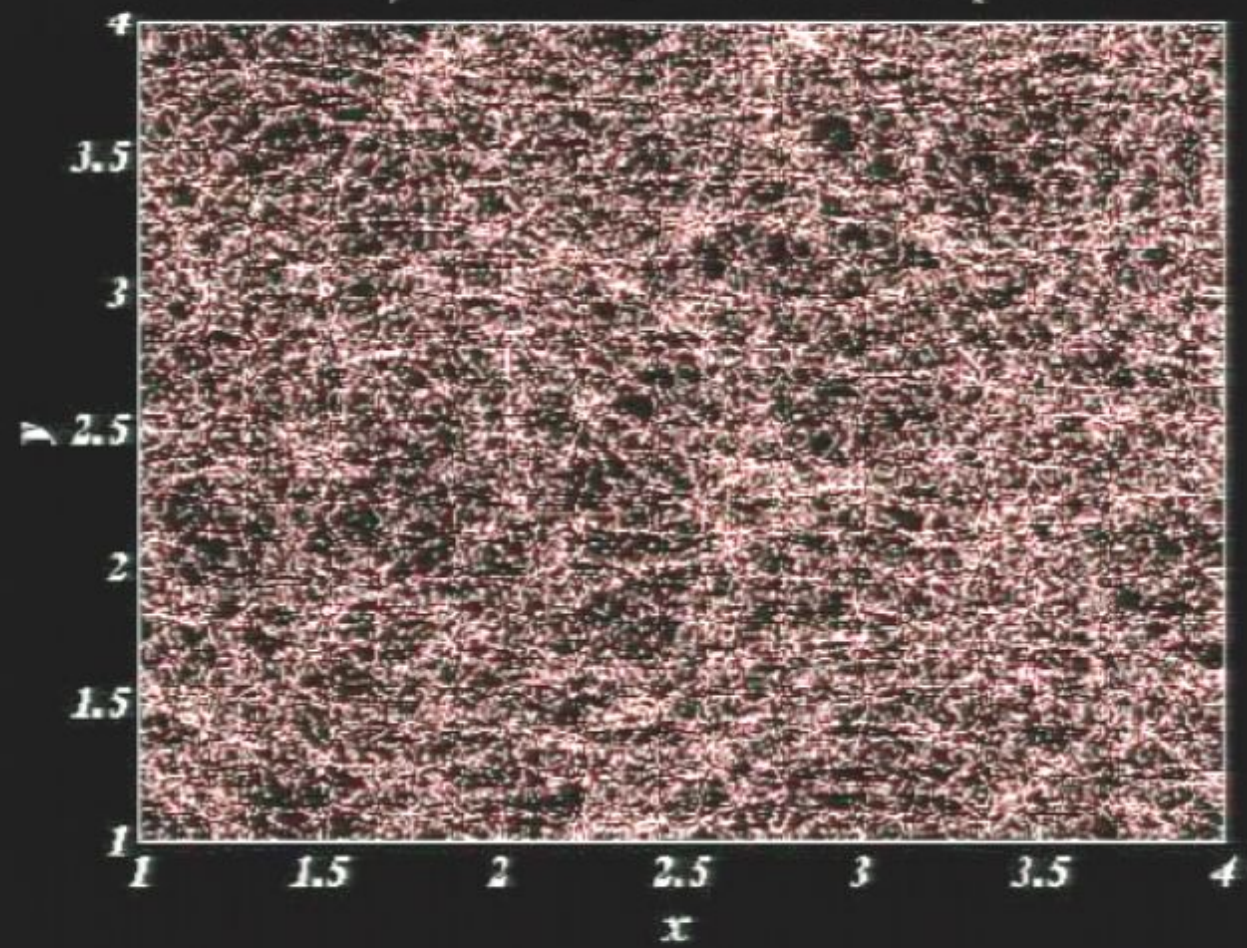
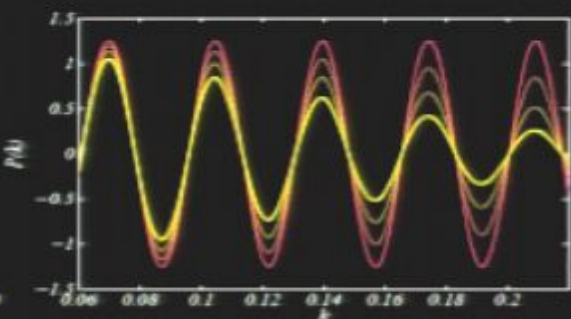
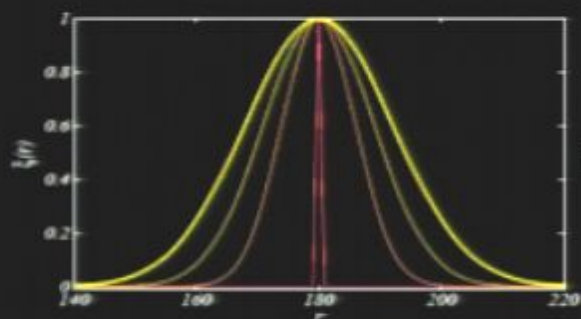
$\xi(r)$

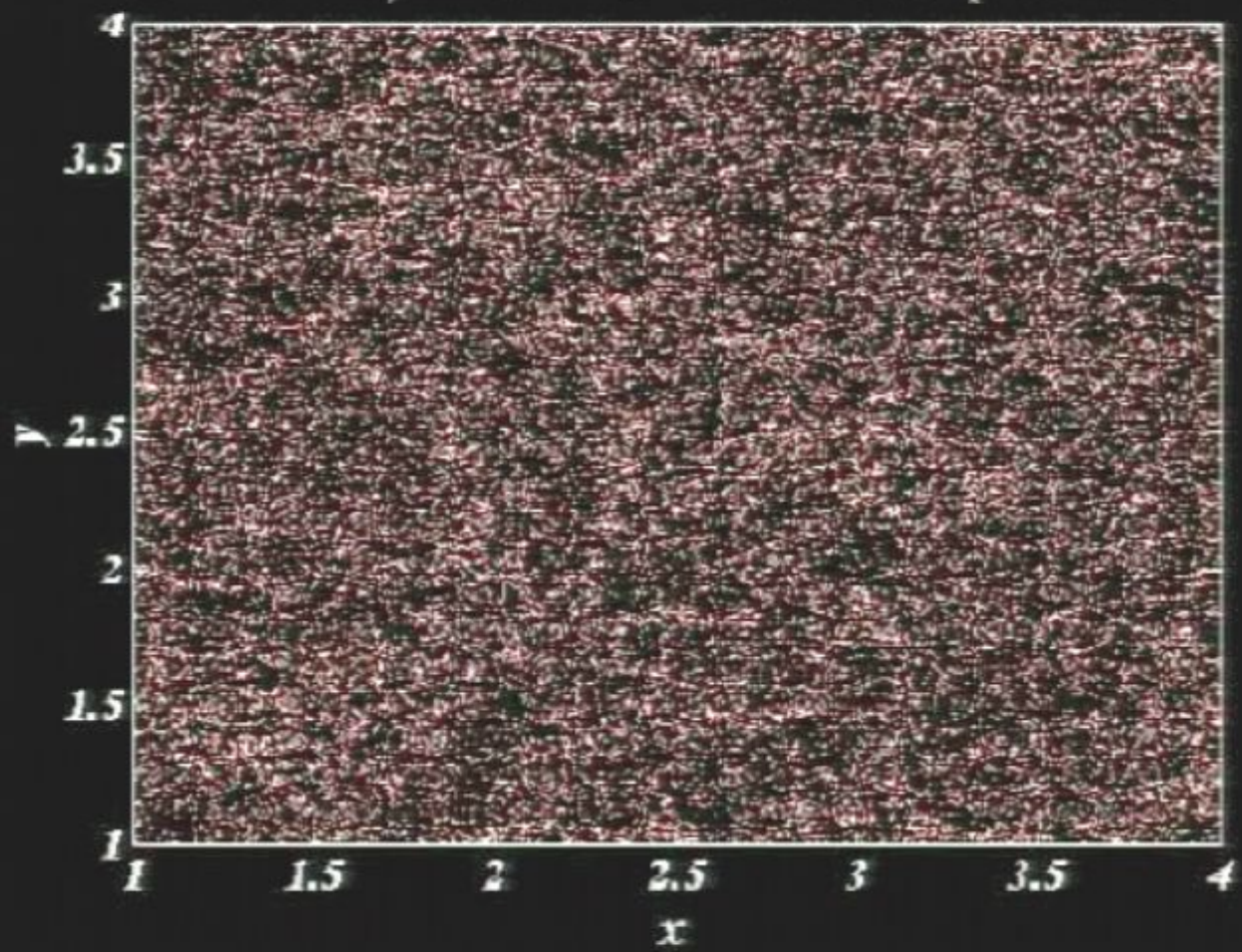
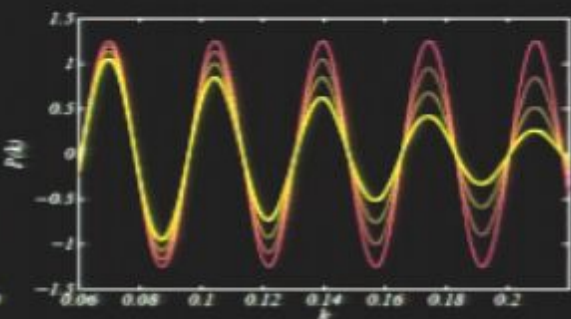
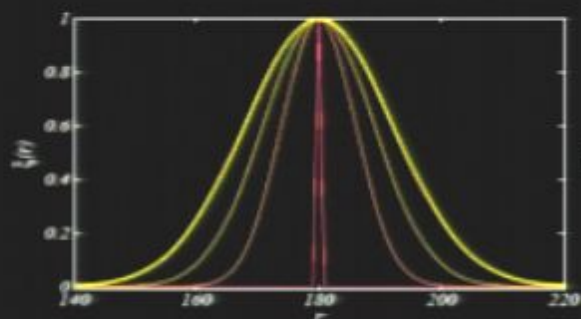


$P(k)$











# Calibrating Nonlinearity

- Although it is (weakly) nonlinear, gravity is still quite clean (compare turbulent MHD in strong gravitational field)
- Good reasons to believe the nonlinear effects can be calibrated at the 1% level for BAO/BAP
- Still needs more theoretical and numerical study
- Only true if background is FLRW ... what do we do in LTB?



# Photometric BAO Surveys

It is tempting to want to just do an imaging survey instead of taking spectra

Photometric BAO surveys primarily sacrifice  $H(z)$  information and need much larger volume to compete.





Skin colour: Earthy Copper

Background: none

Default input: See & Eisenstein 2005

**Parameters**

H\_0    O\_m    O\_k    w\_0    w\_a

**Base parameters**

70	0.3	0	-1	0
----	-----	---	----	---

**Prior matrix**

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

Use Prior

**Observable**

H    d\_A    G

**Derivative Type**

Analytical   Analytical   Numerical

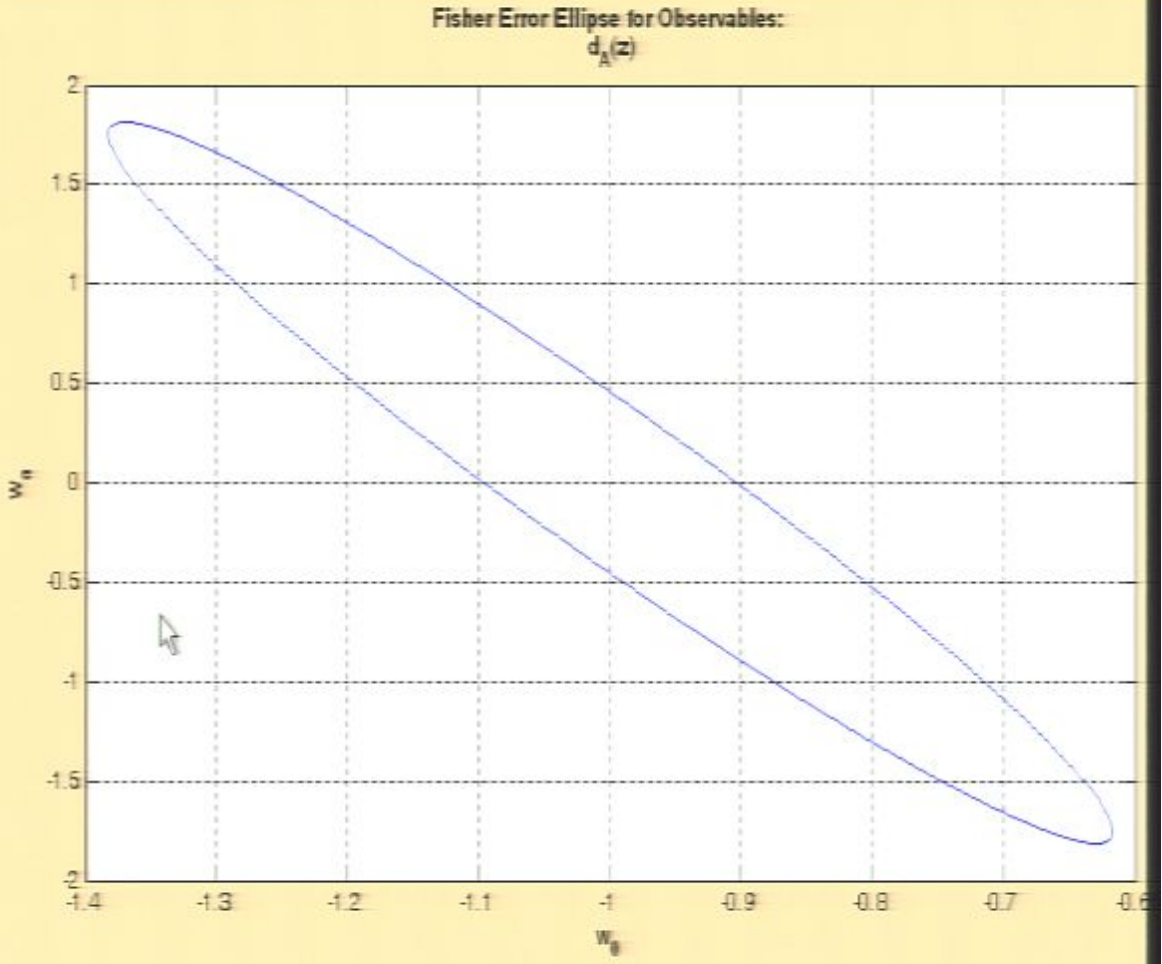
**Data Redshifts**

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

Normalise Growth at z = 0

**Fractional Errors on Observables**

0.098	0.0519	
0.0519	0.049	
0.0359	0.022	
0.0294	0.023	
0.0253	0.0209	
0.0148	0.0119	
	0.0022	



Hold on   Line Style: -    xlim: [-1.6 -0.6]   Clear

Area Fill   Sigma Level: 1 sigma   Edit axis labels

Line Color    Grid   over plot    ylim: [-1 1]   Saving Features

Run

Figure of Merit =

DEF (1/Area 2-sigma)

Local Area Connection 2  
A network cable is unplugged.

Page 155/181

Reset Form







Skin colour: **Earthy Copper**

Background: **none**

Default Input: **Seo & Eisenstein 2003**

**Parameters**

H\_0  O...  O\_k  w\_0  w\_a

**Base parameters**

70	0.3	0	-1	0
----	-----	---	----	---

**Prior matrix**

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

Use Prior

**Observable**

H  d\_A  G

**Derivative Type**

Analyti... Analyti... Numerical

**Data Redshifts**

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

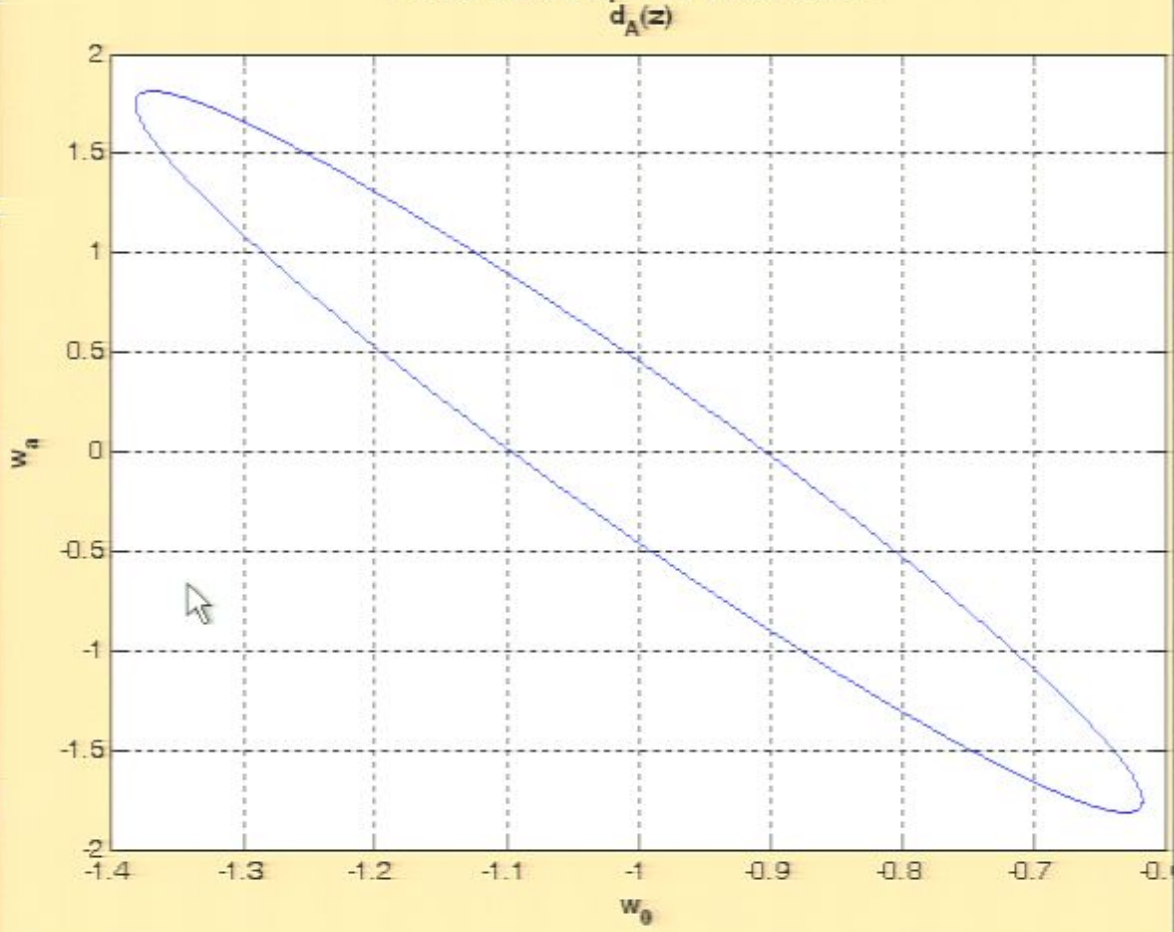
Normalise Growth at z = 0

**Fractional Errors on Observables**

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Pirsa: 09070039

**Fisher Error Ellipse for Observables:**



Hold on  Area Fill  Line Color

Line Style: **-** Sigma Level: **1 sigma**  Grid

xlim: **[-1.5 -0.5]**  ylim: **[-1 1]**

Buttons: Clear, Edit axis labels, Saving Featur...

**Run**

Figure of

**Local Area Connection 2**

A network cable is unplugged.

DETF (1/Area 2-...)

Page 157/181

Reset FoM



Skin colour: **Earthy Copper**

Background: **none**

Default Input: **Seo & Eisenstein 2003**

**Parameters**

H\_0  O...  O\_k  w\_0  w\_a

**Base parameters**

70	0.3	0	-1	0
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**Prior matrix**

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0	10000	0	0	0
0	0	10000	0	0
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0	0	0	0	0

**Observable**

H  d\_A  G

**Derivative Type**

Analyti... Analyti... Numerical

**Data Redshifts**

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

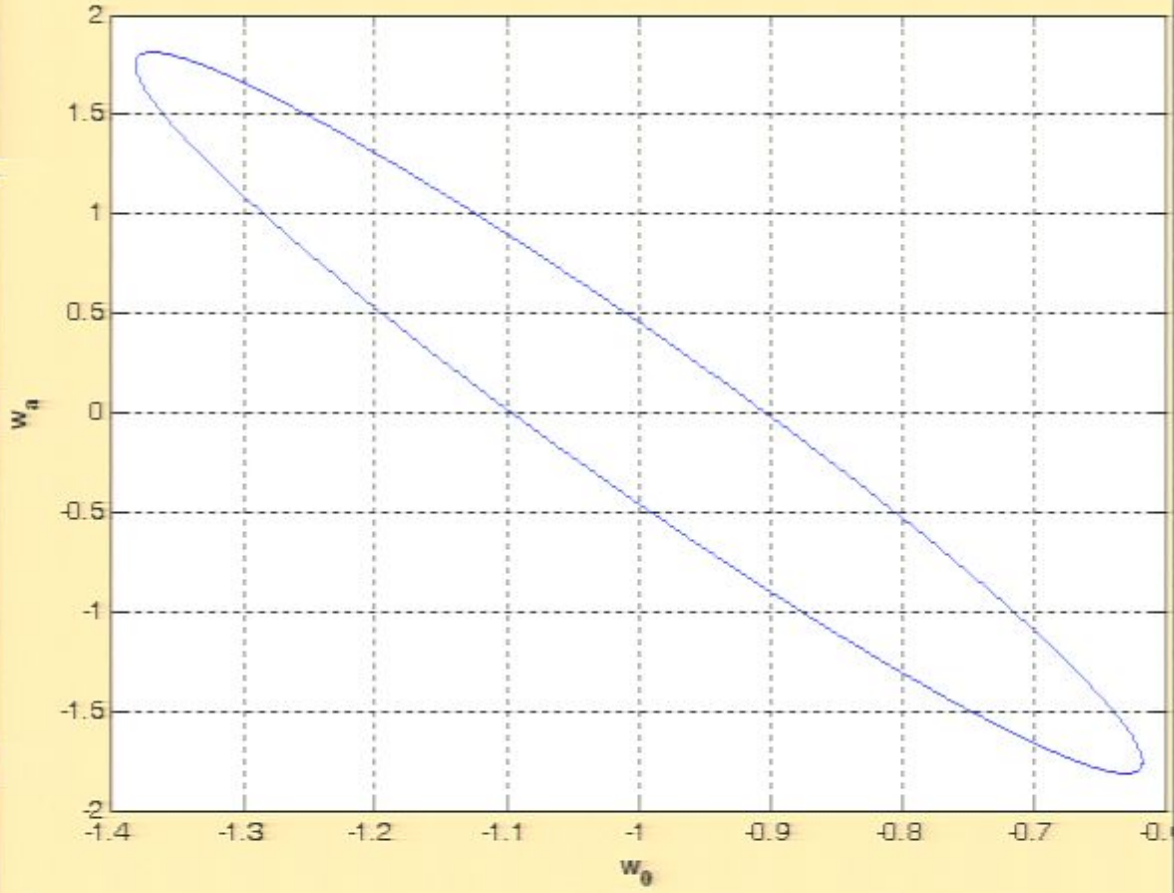
Normalise Growth at z = 0

**Fractional Errors on Observables**

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Fisher Error Ellipse for Observables:

$$d_A(z)$$



Hold on  Area Fill  Line Color

Line Style: **-** Sigma Level: **1 sigma**  Grid

xlim: [-1.5 -0.5]  ylim: [-1 1]

Buttons: Clear, Edit axis labels, Saving Featur...

**Run**

Figure of Merit = **1.70**

DETF (1/Area 2-...  
Page 158/181  
Reset FoM



### Choose Fill Color

Basic colors:

Custom colors:

Define Custom Colors >>

OK Cancel

view Insert Tools Desktop Window Help

Earthy Copper

none

Seo & Eisenstein 2003

O\_k  w\_0  w\_a

0	-1	0
0	0	0
0	0	0
0000	0	0
0	0	0
0	0	0

Observable

Derivative Type

Data Redshifts

Fractional Errors on Observables

Pirsa: 09070039

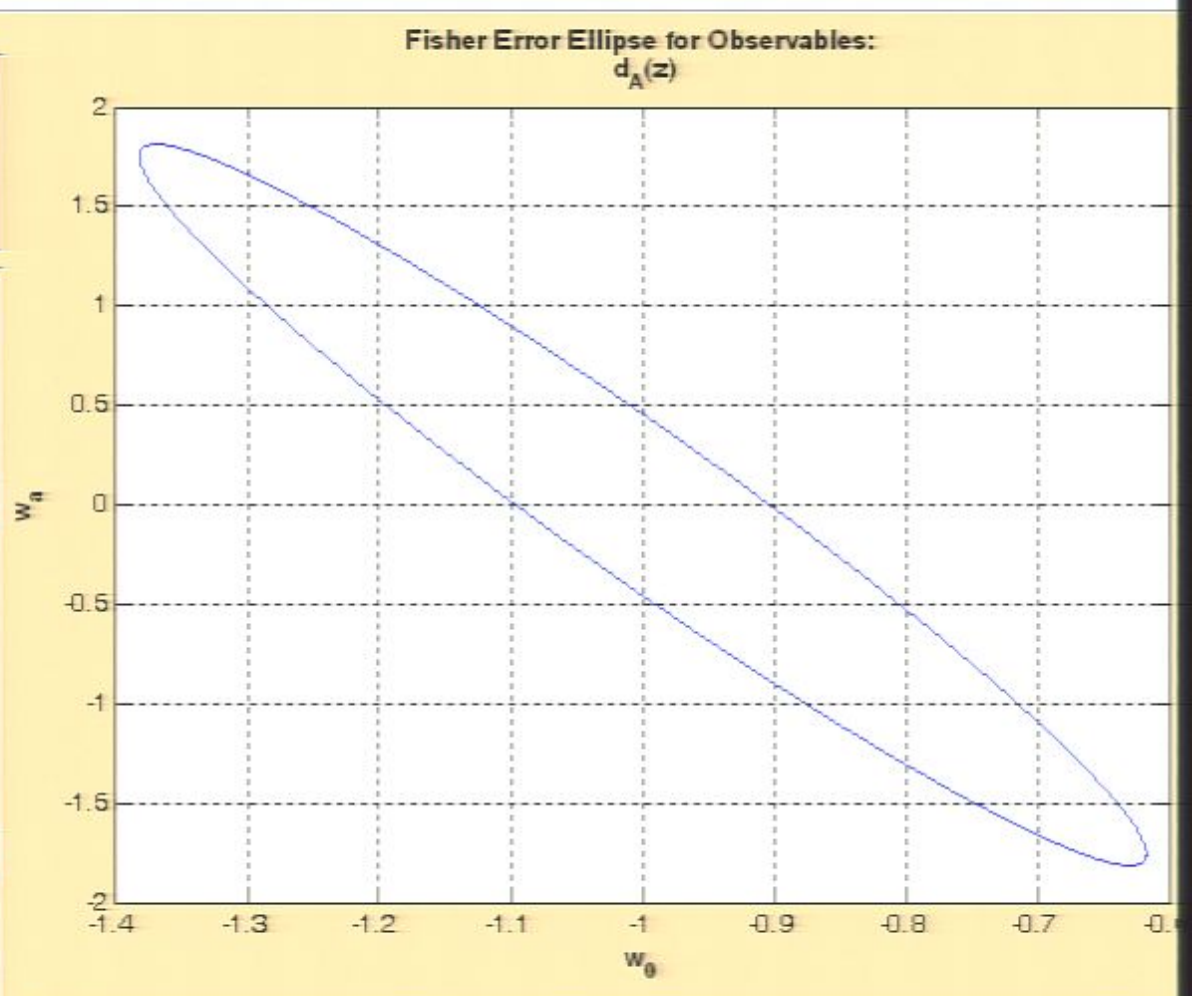
H  d\_A  G

Analyti... Analyti... Numerical

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	



Hold on  xlim [-1.5 -0.5]

Area Fill  ylim [-1 1]

Line Color  Grid over plot

Line Style: -

Sigma Level: 1 sigma

Clear Edit axis labels Saving Featur...

Run Figure of Merit = 1.70

DETF (1/Area 2-...

Page 159/181 ResetFoM



Skin colour: **Earthy Copper**

Background: **none**

Default Input: **Seo & Eisenstein 2003**

**Parameters**

H\_0  O...  O\_k  w\_0  w\_a

**Base parameters**

70	0.8	0	-1	0
----	-----	---	----	---

**Prior matrix**

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

**Observable**

H  d\_A  G

**Derivative Type**

Analyti... Analyti... Numerical

**Data Redshifts**

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

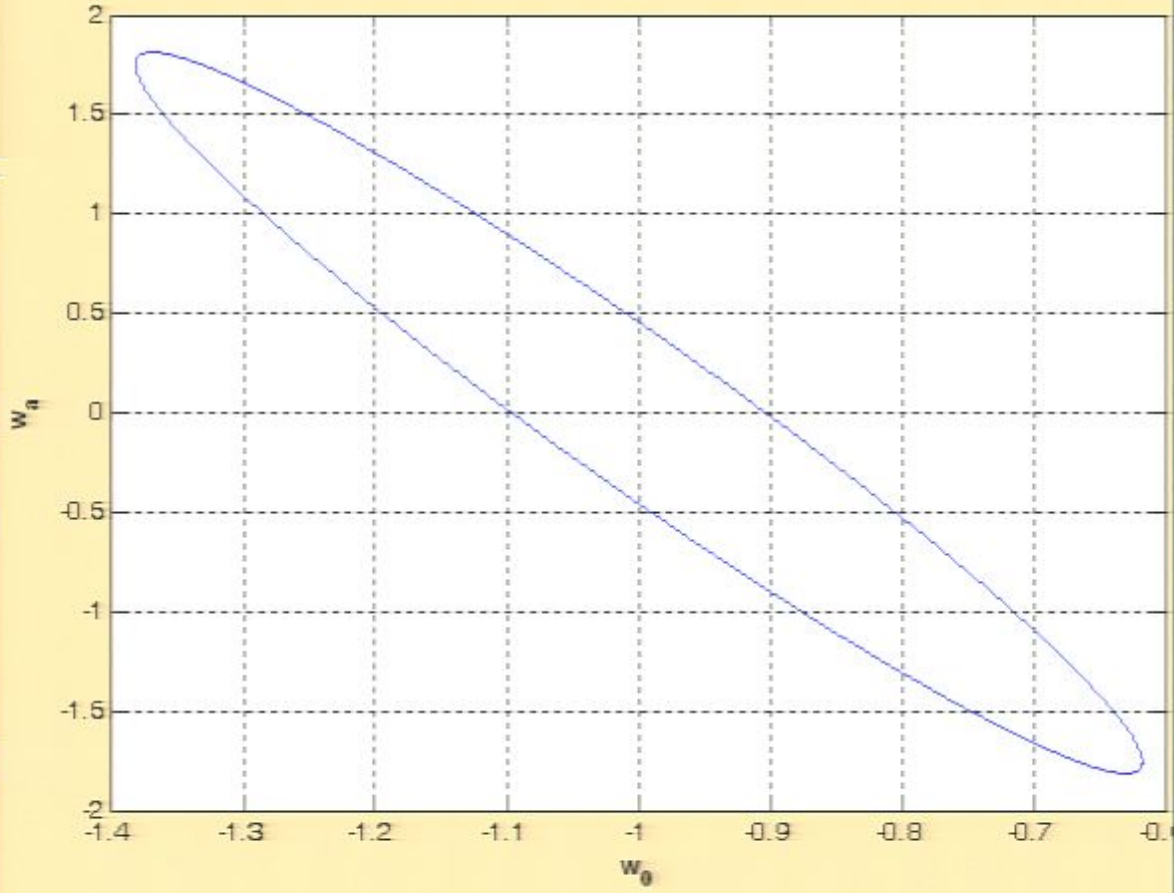
**Fractional Errors on Observables**

Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.028	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Fisher Error Ellipse for Observables:

$d_A(z)$



Hold on    Line Style: **-**     xlim: [-1.5-0.5]    Clear

Area Fill    Sigma Level: **1 sigma**     ylim: [-1 1]    Edit axis labels

Line Color     Grid: **over plot**    Saving Featur...

**Run**

Figure of Merit = **1.70**

DETF (1/Area 2-...  
Page 160/181  
Reset FoM





Skin colour:

Background:

Default Input:

**Parameters**

H\_0  O...  O\_k  w\_0  w\_a

**Base parameters**

70	0.3	0	-1	0
----	-----	---	----	---

**Prior matrix**

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

Use Prior

**Observable**

H  d\_A  G

**Derivative Type**

Analyti... Analyti... Numerical

**Data Redshifts**

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

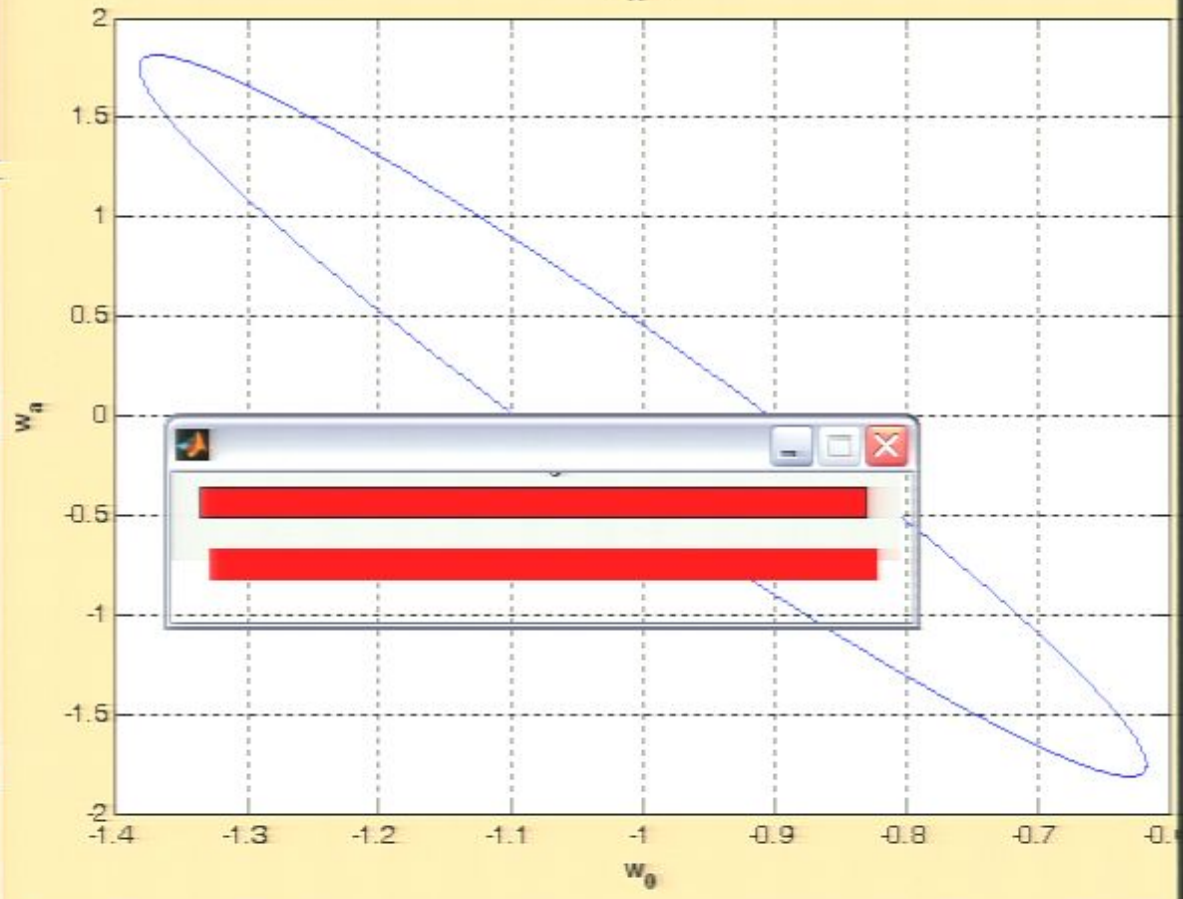
Normalise Growth at z =

**Fractional Errors on Observables**

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Pirsa: 09070039

Fisher Error Ellipse for Observables:  
 $d_A(z)$



Hold on  Line Color  Area Fill  Grid  xlim [-1.5 -0.5]  ylim [-1 1]



Figure of Merit =

DETF (1/Area 2-...





Skin colour:

Background:

Default Input:

Parameters

H\_0  O...  O\_k  w\_0  w\_a

Base parameters

70	0.3	0	-1	0
----	-----	---	----	---

Prior matrix

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

Use Prior

Observable

H  d\_A  G

Derivative Type

Analyti... Analyti... Numerical

Data Redshifts

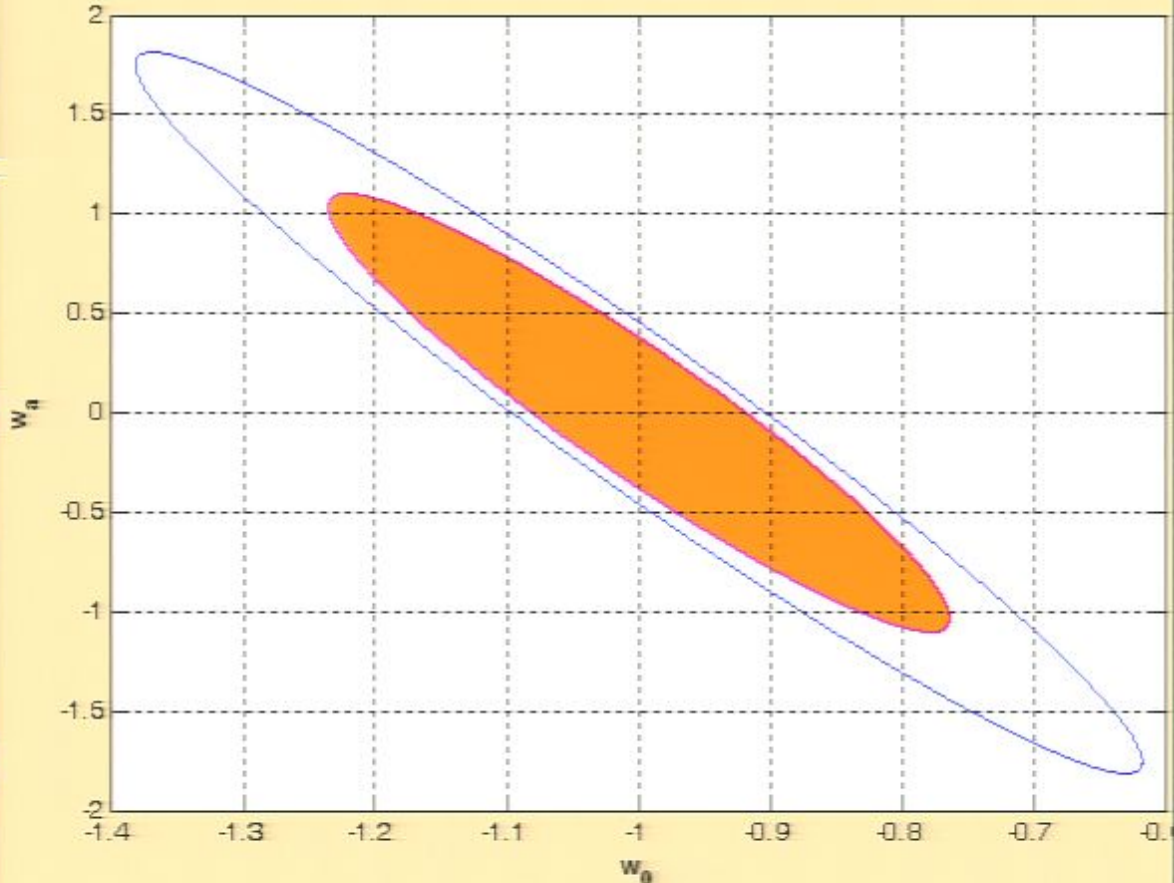
0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

Normalise Growth at z =

Fractional Errors on Observables

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Fisher Error Ellipse for Observables:  $H(z), d_A(z)$



Hold on  xlim [-1.5 -0.5]

Area Fill  ylim [-1 1]

Line Color  Grid over plot

Line Style:  Sigma Level:



Figure of Merit =



Skin colour: **Earthy Copper**

Background: **none**

Default Input: **Seo & Eisenstein 2003**

**Parameters**

H\_0  O...  O\_k  w\_0  w\_a

**Base parameters**

70 0.3 0 -1 0

**Prior matrix**

```

10000 0 0 0 0
0 10000 0 0 0
0 0 10000 0 0
0 0 0 0 0
0 0 0 0 0
    
```

Use Prior

**Observable**

H  d\_A  G

**Derivative Type**

Analyti... Analyti... Numerical

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

**Data Redshifts**

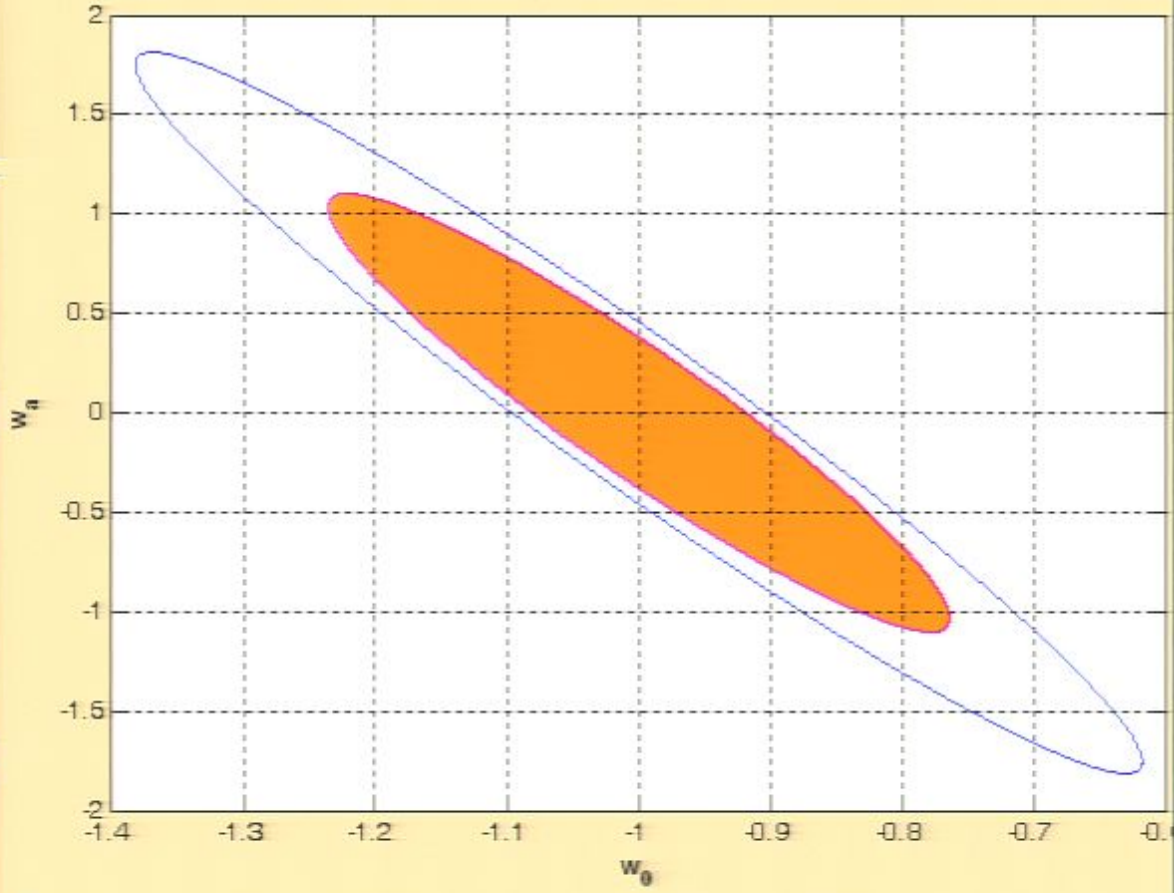
Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

**Fractional Errors on Observables**

Pirsa: 09070039

**Fisher Error Ellipse for Observables:**  
H(z), d\_A(z)



Hold on  xlim [-1.5 -0.5]

Area Fill  Line Color

Line Style: **-** Sigma Level: **1 sigma**  Grid **over plot**

Buttons: Clear, Edit axis labels, Saving Featur...

**Run**

Figure of Merit =

3.30  
1.70  
0

DETF (1/Area 2-...  
ResetFoM





Skin colour:

Background:

Default Input:

**Parameters**

H\_0  O...  O\_k  w\_0  w\_a

**Base parameters**

70 0.3 0 -1 0

**Prior matrix**

```

10000 0 0 0 0
0 10000 0 0 0
0 0 10000 0 0
0 0 0 0 0
0 0 0 0 0
    
```

Use Prior

**Observable**

H  d\_A  G

**Derivative Type**

Analyti... Analyti... Numerical

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

**Data Redshifts**

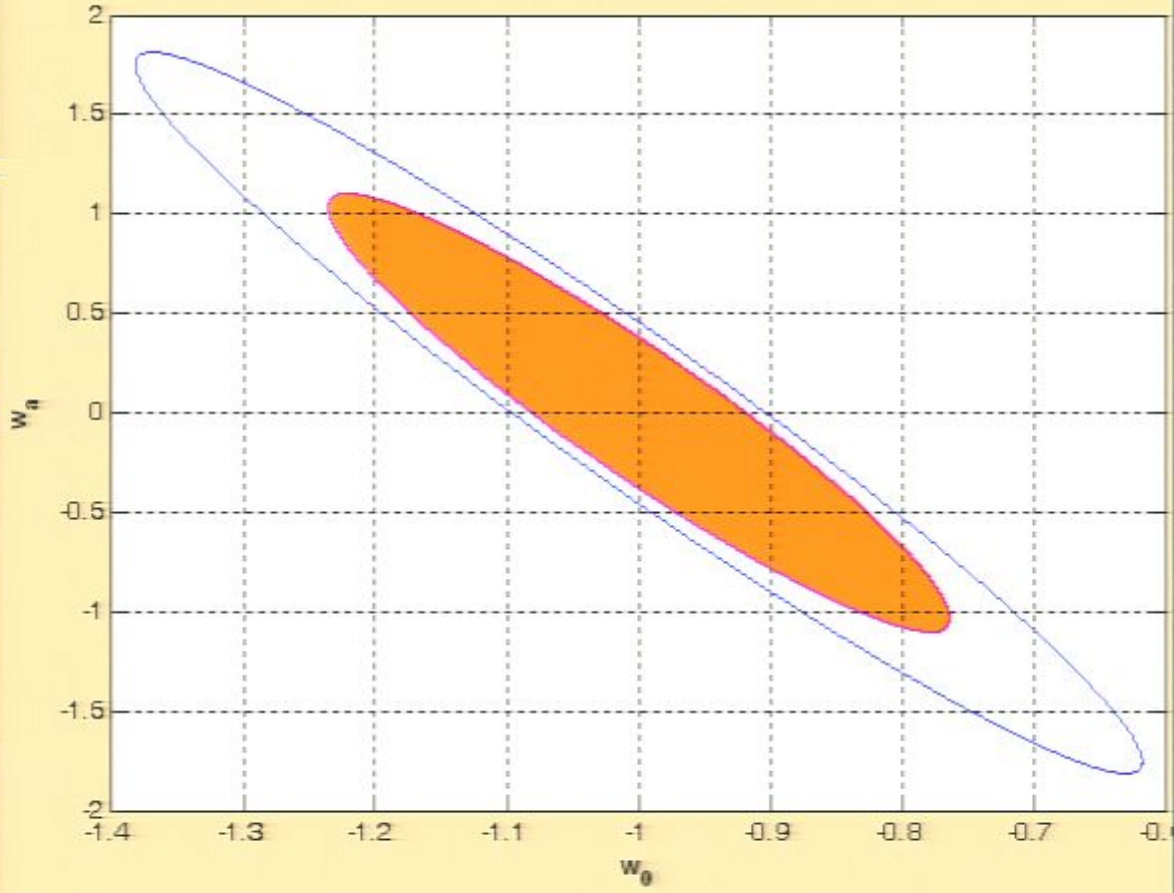
Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

**Fractional Errors on Observables**

Pirsa: 09070039

Fisher Error Ellipse for Observables:  
H(z), d\_A(z)



Hold on  xlim [-1.5 -0.5]  ylim [-1 1]

Area Fill  Line Color  Grid  over plot

Line Style:  Sigma Level: 1 sigma

Figure of Merit = 3.30

- DETF (1/Area 2-sigma)
- 1/Area 1-sigma
- Area 1-sigma
- Trace(cov)
- sum(cov^2)
- DETF (1/Area 2-sigma)
- ResetFoM





Skin colour: **Earthy Copper**

Background: **none**

Default Input: **Seo & Eisenstein 2003**

**Parameters**

H\_0  O...  O\_k  w\_0  w\_a

**Base parameters**

70	0.3	0	-1	0
----	-----	---	----	---

**Prior matrix**

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

**Observable**

H  d\_A  G

**Derivative Type**

Analyti... Analyti... Numerical

**Data Redshifts**

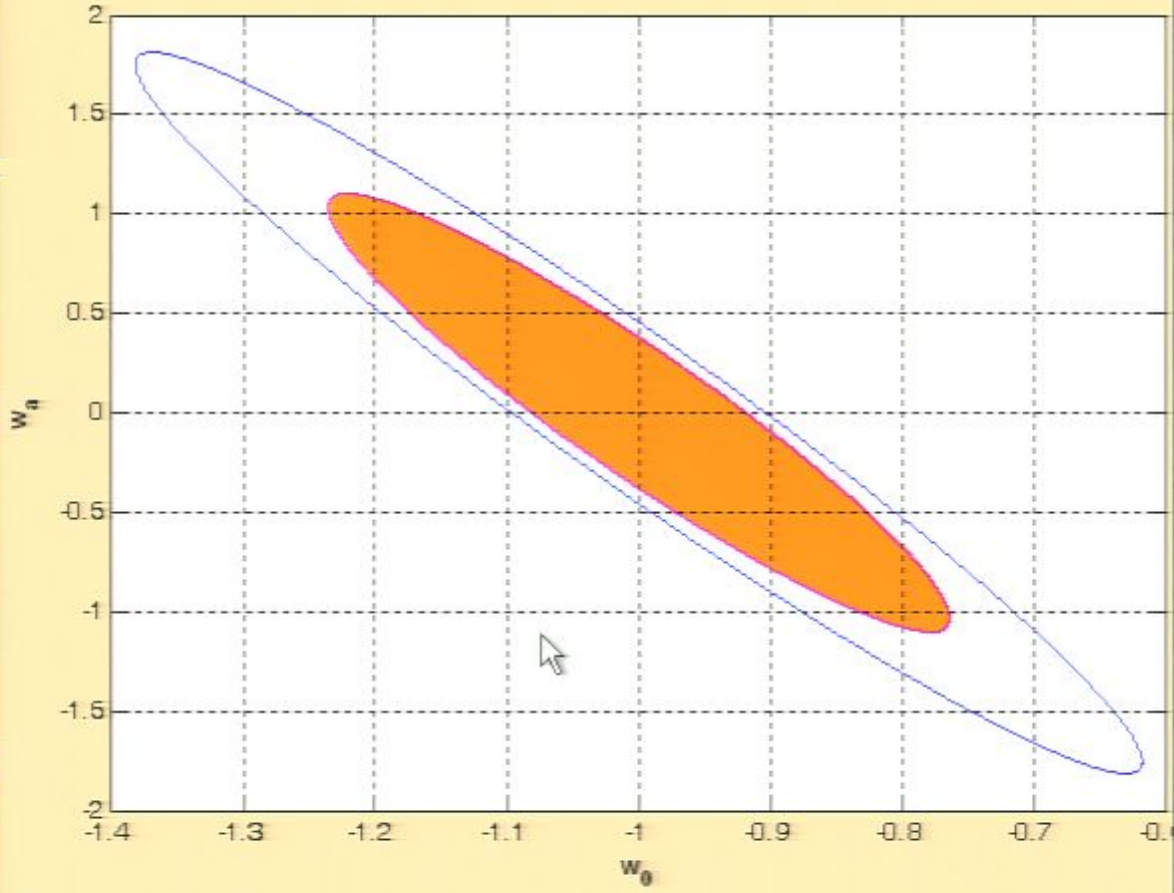
0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

**Fractional Errors on Observables**

Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

**Fisher Error Ellipse for Observables:**  
H(z), d\_A(z)



Hold on    Line Style: **-**     xlim: [-1.5 -0.5]    Clear

Area Fill    Sigma Level: **1 sigma**     ylim: [-1 1]    Edit axis labels

Line Color     Grid: **over plot**    Saving Featur...

**Run**

Figure of Merit = **3.30**

DETF (1/Area 2-...)  
Page 165/181  
Reset FoM



Skin colour: Earthy Copper

Background: none

Default Input: none

- none
- WMAP
- Millenium Simulation
- The Matrix
- Load from file

Parameters

Base parameters

<input type="checkbox"/> H <sub>0</sub>	<input type="checkbox"/> Ω <sub>b</sub>	<input type="checkbox"/> Ω <sub>c</sub>	<input type="checkbox"/> Ω <sub>m</sub>	<input type="checkbox"/> Ω <sub>Λ</sub>
70	0.3	0	-1	0

Prior matrix

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

Use Prior

Observable

H  d<sub>A</sub>  G

Derivative Type

Analyti... Analyti... Numerical

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

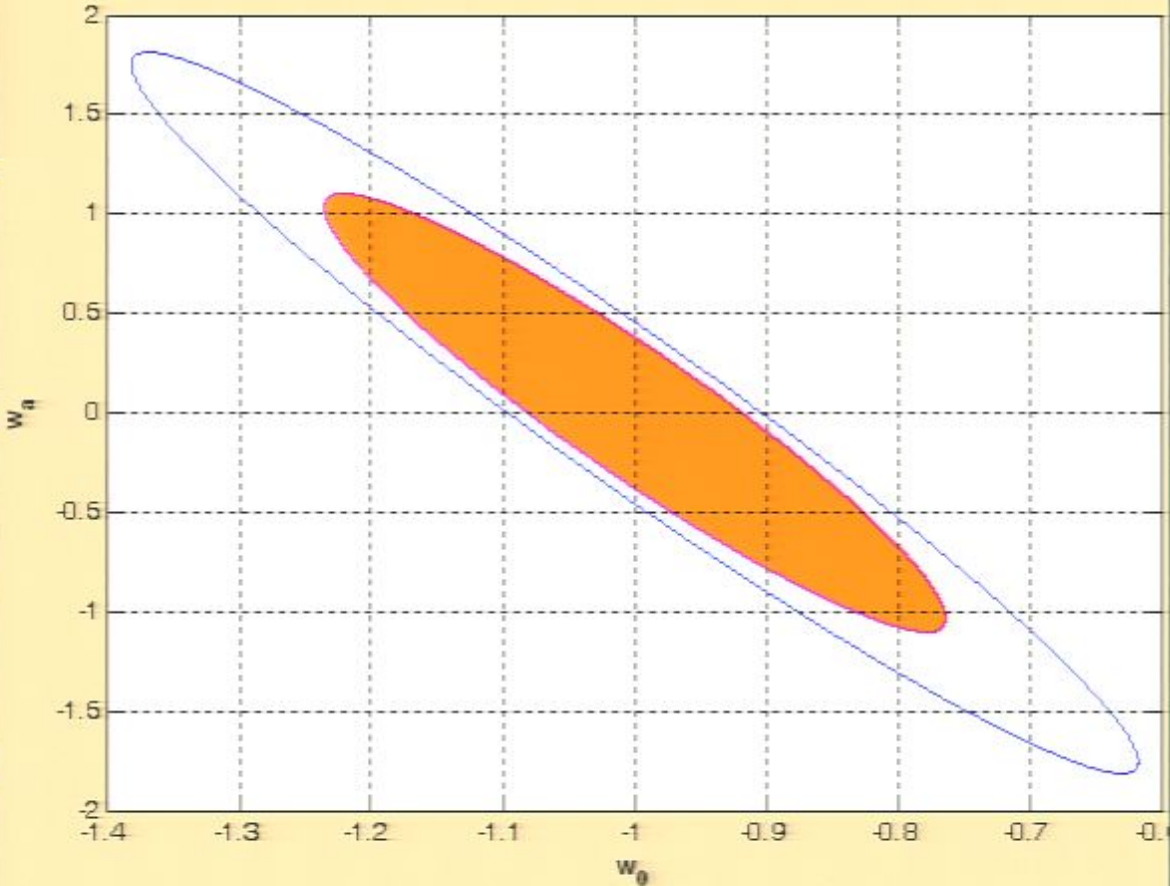
Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Fractional Errors on Observables

Pirsa: 09070039

Fisher Error Ellipse for Observables: H(z), d<sub>A</sub>(z)



Hold on    Line Style:      xlim: [-1.5 -0.5]    Clear  
 Area Fill    Sigma Level: 1 sigma     ylim: [-1 1]    Edit axis labels  
 Line Color     Grid: over plot    Saving Featur...

Run

Figure of Merit = 3.30

DETF (1/Area 2-...  
 Page 166/181  
 Reset FoM





**Fisher 4cast**

Skin colour: Earthy Copper

Background: WMAP

Default Input: Seo & Eisenstein 2003

**Parameters**

Base parameters:  H<sub>0</sub>  O...  O<sub>k</sub>  w<sub>0</sub>  w<sub>a</sub>

70	0.8	0	-1	0
----	-----	---	----	---

Prior matrix:

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

Use Prior

Observable:  H  d<sub>A</sub>  G

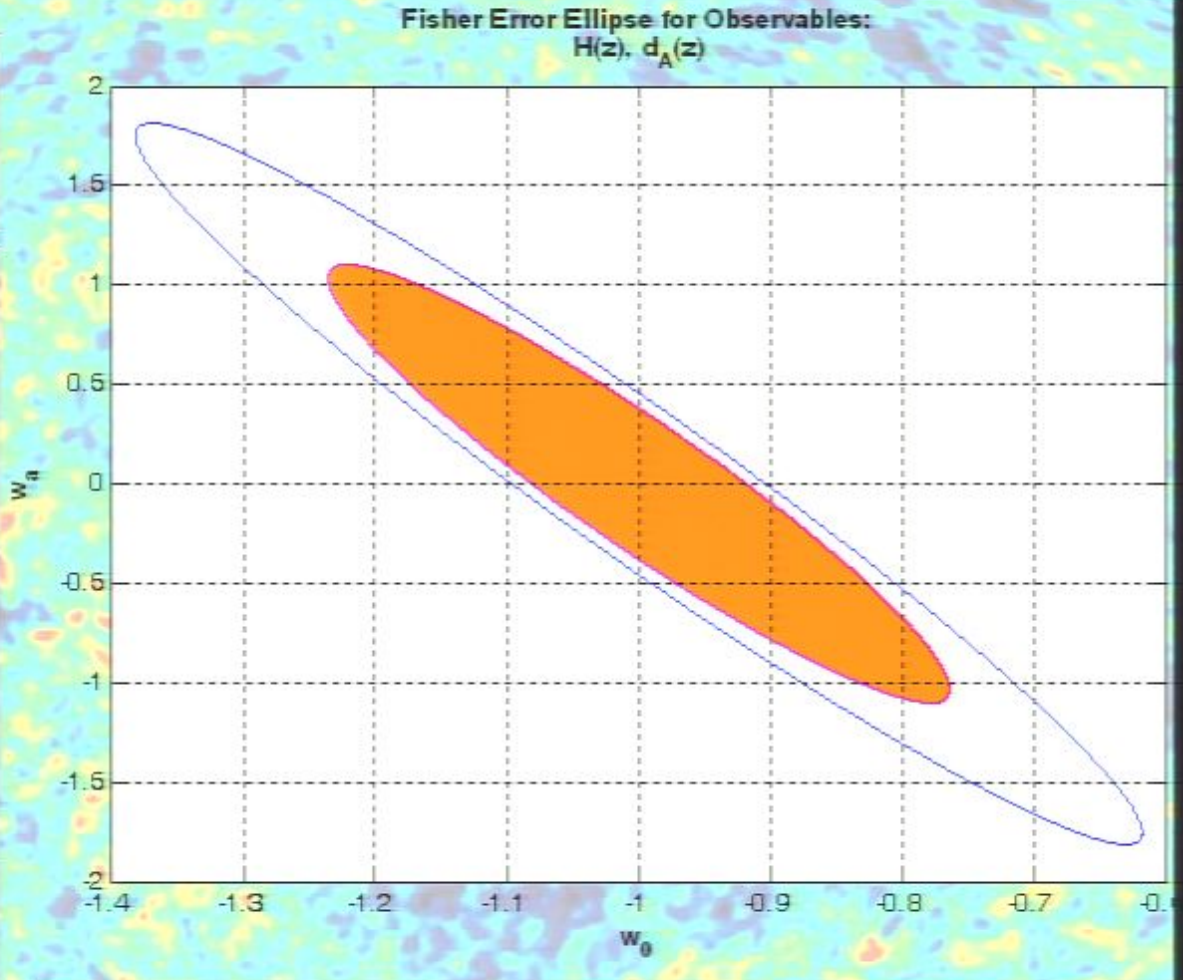
Derivative Type: Analyti... Analyti... Numerical

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Fractional Errors on Observables



Hold on

Area Fill

Line Color

Line Style:

Sigma Level: 1 sigma

Grid

over plot

xlim: [-1.5 -0.5]

ylim: [-1 1]

Clear

Edit axis labels

Saving Featur...

**Run** Figure of Merit = 3.30

DETF (1/Area 2-...

Page 167/181

Reset FoM





Skin colour: Earthy Copper

Background: WMAP

Default Input: Seo & Eisenstein 2003

- Parameters
- Base parameters
- Prior matrix
- Use Prior
- Observable
- Derivative Type
- Data Redshifts
- Fractional Errors on Observables

H\_0    O...    O\_k    w\_0    w\_a

70	0.8	0	-1	0
----	-----	---	----	---

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

H    d\_A    G

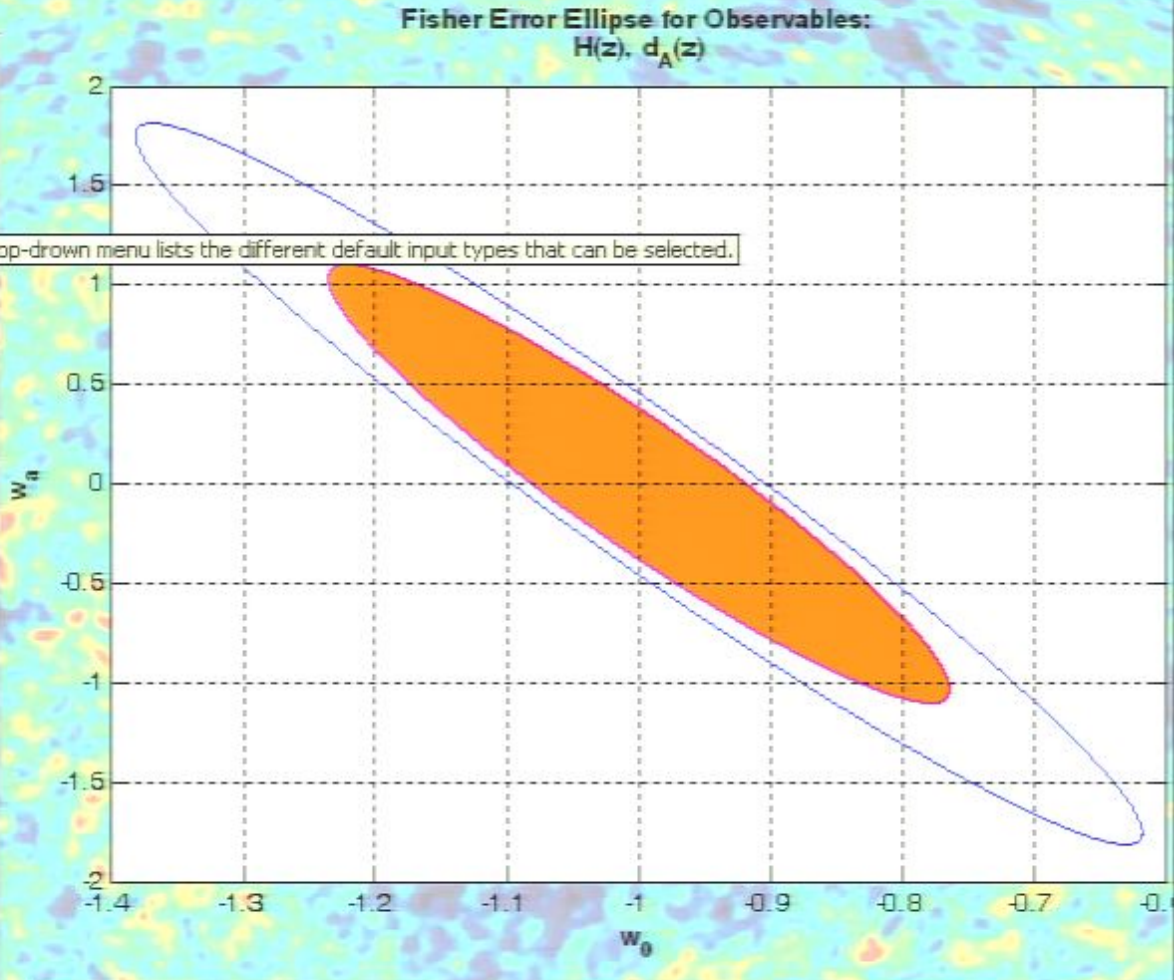
Analyti...   Analyti...   Numerical

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

This drop-down menu lists the different default input types that can be selected.



Hold on   Line Style: -    xlim: [-1.5 -0.5]   Clear

Area Fill   Sigma Level: 1 sigma    ylim: [-1 1]   Edit axis labels

Line Color    Grid   over plot   Saving Featur...

**Run**   Figure of Merit = 3.30   DETF (1/Area 2-...   **ResetFoM**



**Fisher 4cast**

Skin colour: Earthy Copper

Background: WMAP

Default Input: Seo & Eisenstein 2003

**Parameters**

Base parameters:  H\_0  O...  O\_k  w\_0  w\_a

70	0.8	0	-1	0
----	-----	---	----	---

Prior matrix:

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

Use Prior

Observable:  H  d\_A  G

Derivative Type: Analyti... Analyti... Numerical

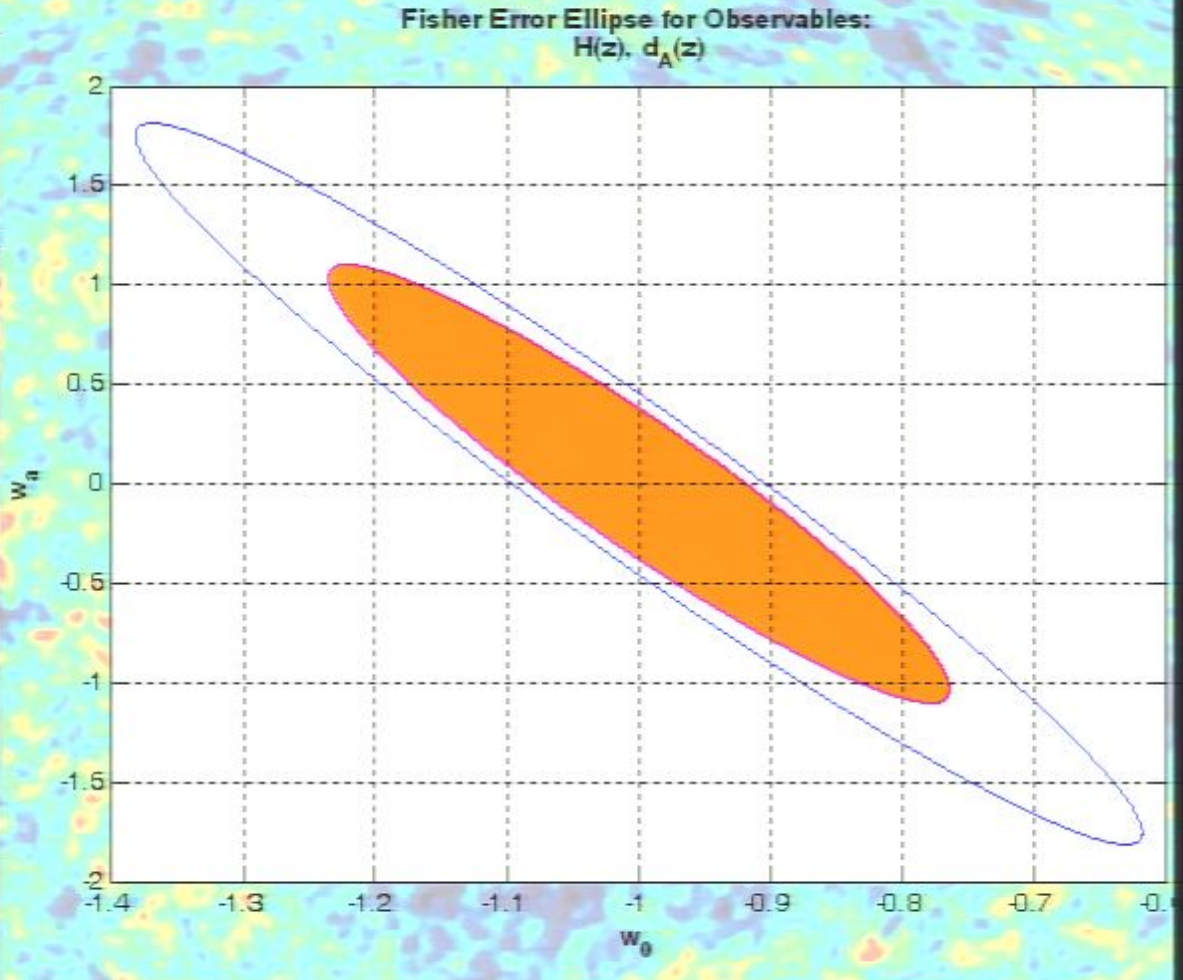
Data Redshifts:

0.3	0.3	
0.6	0.6	
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Fractional Errors on Observables



Hold on    Line Style: -     xlim: [-1.5 -0.5]    Clear

Area Fill    Sigma Level: 1 sigma     ylim: [-1 1]    Edit axis labels

Line Color     Grid: over plot    Saving Featur...

**Run**    Figure of Merit = 3.30    DETF (1/Area 2-...    Page 169/181    Reset FoM





Skin colour: Earthy Copper  
 Background: WMAP  
 Default Input: Seo & Eisenstein 2003

**Parameters**

Base parameters

Prior matrix

Use Prior

Observable

Derivative Type

Data Redshifts

Fractional Errors on Observables

H\_0    O...    O\_k    w\_0    w\_a

70	0.8	0	-1	0
----	-----	---	----	---

10000	0	0	0	0
0	10000	0	0	0
0	0	10000	0	0
0	0	0	0	0
0	0	0	0	0

H    d\_A    G

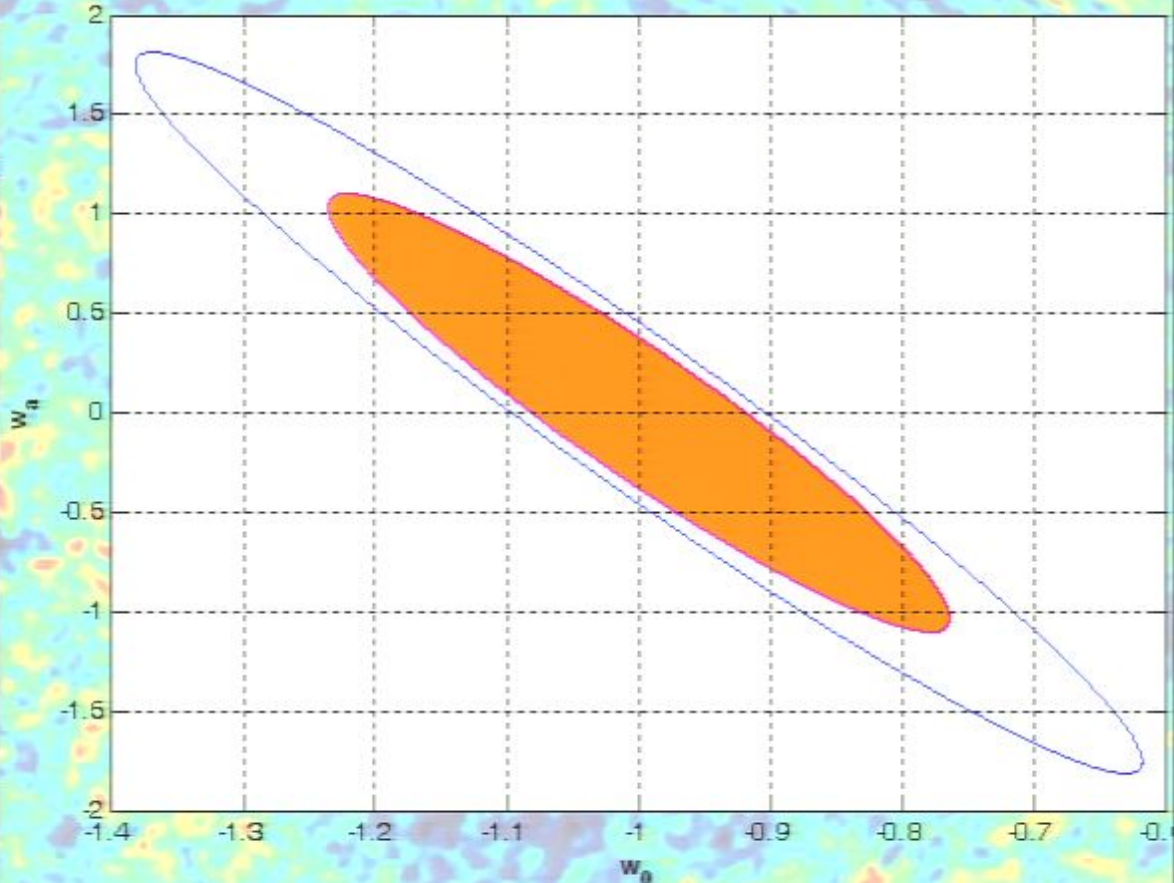
Analytical   Analytical   Numerical

0.3	0.8	
0.6		
0.8	0.8	
1	1	
1.2	1.2	
3	3	
	1000	

Normalise Growth at z = 0

0.058	0.0519	
0.0519	0.043	
0.0369	0.0322	
0.0284	0.023	
0.0253	0.0203	
0.0148	0.0119	
	0.0022	

Fisher Error Ellipse for Observables:  
 $H(z), d_A(z)$



Hold on   Line Style: -    xlim: [-1.5 -0.5]   Clear

Area Fill   Sigma Level: 1 sigma    ylim: [-1 1]   Edit axis labels

Line Color    Grid   over plot   Saving Featur...

**Run**   **Figure of Merit = 3.30**   DETF (1/Area 2-...   **Page 170/181**   ResetFoM



```
1  * -----
2  * Copyright (C) 2008-2009
3  * Bruce Bassett Yabebai Fantaye  Renee Hiozek  Jacques Kotze
4  *
5  *
6  *
7  * This file is part of Fisher4Cast.
8  *
9  * Fisher4Cast is free software: you can redistribute it and/or modify
10 * it under the terms of the Berkeley Software Distribution (BSD) license.
11 *
12 * Fisher4Cast is distributed in the hope that it will be useful,
13 * but WITHOUT ANY WARRANTY; without even the implied warranty of
14 * MERCHANTABILITY or FITNESS FOR A PARTICULAR PURPOSE.  See the
15 * BSD license for more details.
16 * -----
17 *
18 * FM_GUI sets a Gui for the Fisher matrix tacklebox.
19 *
20 * It excutes the callbacks which are
21 * USERIN- the input for the GUI. During the first
22 * initialisation of this function, USERIN is zero
23 * which opens the new GUI window. All buttons and controls follow a specific
24 * format as outlined in the Matlab Help files, and indentation is key in
25 * this code. The GUI also contains floating help to aid the user.
26 * For more information on GUI's, please see the Matlab documentation.
27
28
29 function FM_GUI(USERIN)
30
31 - global axis_spec tacklebox
32
33
34 - if nargin==0
```


```

1 function isoFoM_movie
2 % movie of iso FoM surface - either slice or isosurface
3 % BB 1 June 09
4
5 - close all
6 %load v_30_new.mat; old data
7 - load new_detf_fom_vol_out.mat % new DETF FoM data
8 %v = v_30_new;
9 - v = fom_vol;
10 - H=linspace(0.1,5,30); da=linspace(0.1,5,30); G=linspace(0.1,5,30);
11 - [x y z] = meshgrid(H,da,G);
12
13 - xmin = 0.1;
14 - xmax = 5;
15 - ymin = 0.1;
16 - ymax = 5;
17 - zmin = 0.1;
18 - zmax = 5;
19 - minv = min(min(min(v)));
20 - maxv = max(max(max(v)));
21
22 - figure(2)
23 - whitebg('white') % colour of background
24 %*****
25 - n = 75; % number of threshold steps
26 - step = (maxv-minv)./n; % union step in threshold
27 - thresh = minv:step:maxv; % threshold vector
28 %*****
29
30 - slice_flag = 1; % 1 if slice animation, 0 if union
31
32 - nslice = 50;
33
34 %*****

```

```
1 function isoFoM_movie
2 % movie of iso FoM surface - either slice or isosurface
3 % BB 1 June 09
4
5 - close all
6 %load v_30_new.mat; old data
7 - load new_detf_fom_vol_out.mat % new DETF FoM data
8 %v = v_30_new;
9 - v = fom_vol;
10 - H=linspace(0.1,5,30); da=linspace(0.1,5,30); G=linspace(0.1,5,30);
11 - [x y z] = meshgrid(H,da,G);
12
13 - xmin = 0.1;
14 - xmax = 5;
15 - ymin = 0.1;
16 - ymax = 5;
17 - zmin = 0.1;
18 - zmax = 5;
19 - minv = min(min(min(v)));
20 - maxv = max(max(max(v)));
21
22 - figure(2)
23 - whitebg('white') % colour
24 %*****
25 - n = 75; % number of threshold steps
26 - step = (maxv-minv)./n; % union step in threshold
27 - thresh = minv:step:maxv; % threshold vector
28 %*****
29
30 - slice_flag = 1; % 1 if slice animation, 0 if union
31
32 - nslice = 50;
33
34 %*****
```

**MATLAB Editor**

 File D:\Bruce\tex\Fisher Tackle\Papers\isoFoM\_movie.m is not found in the current directory or on the MATLAB path.

To run this file, select one of the following:

- Change MATLAB current directory
- Add directory to the top of the MATLAB path
- Add directory to the bottom of the MATLAB path

OK Cancel





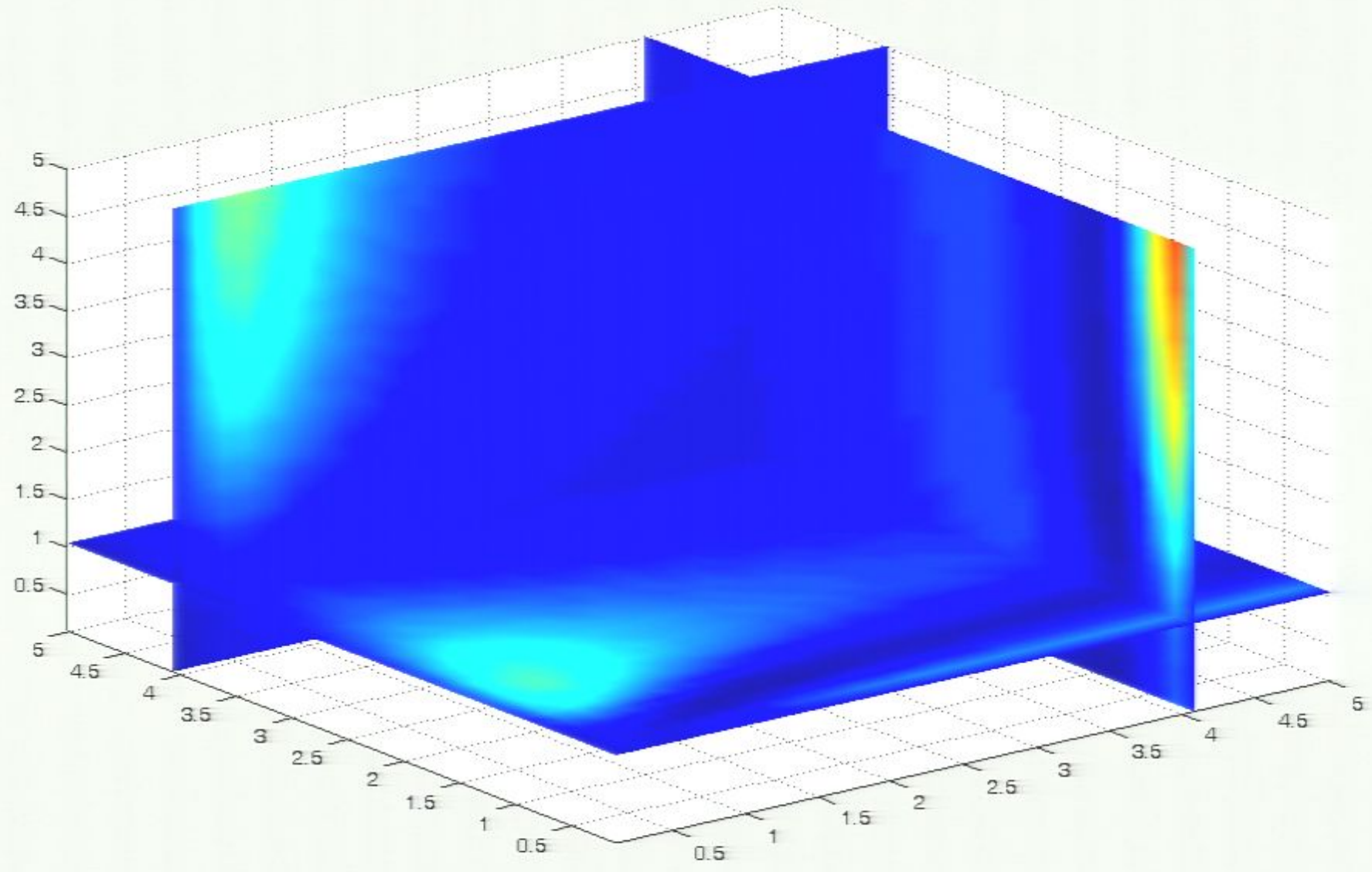
```

1 function isoFoM_movie
2 % movie of iso FoM surface - either slice o
3 % BB 1 June 09
4
5 - close all
6 %load v_30_new.mat; old data
7 - load new_detf_fom_vol_out.mat % new DETF F
8 %v = v_30_new;
9 - v = fom_vol;
10 - H=linspace(0.1,5,30); da=linspace(0.1,5,30)
11 - [x y z] = meshgrid(H,da,G);
12
13 - xmin = 0.1;
14 - xmax = 5;
15 - ymin = 0.1;
16 - ymax = 5;
17 - zmin = 0.1;
18 - zmax = 5;
19 - minv = min(min(min(v)));
20 - maxv = max(max(max(v)));
21
22 - figure(2)
23 - whitebg('white') % colour of background
24 %*****
25 - n = 75; % number of threshold steps
26 - step = (maxv-minv)./n; % union step in threshold
27 - thresh = minv:step:maxv; % threshold vector
28 %*****
29
30 - slice_flag = 1; % 1 if slice animation, 0 if union
31
32 - nslice = 50;
33
34 %*****

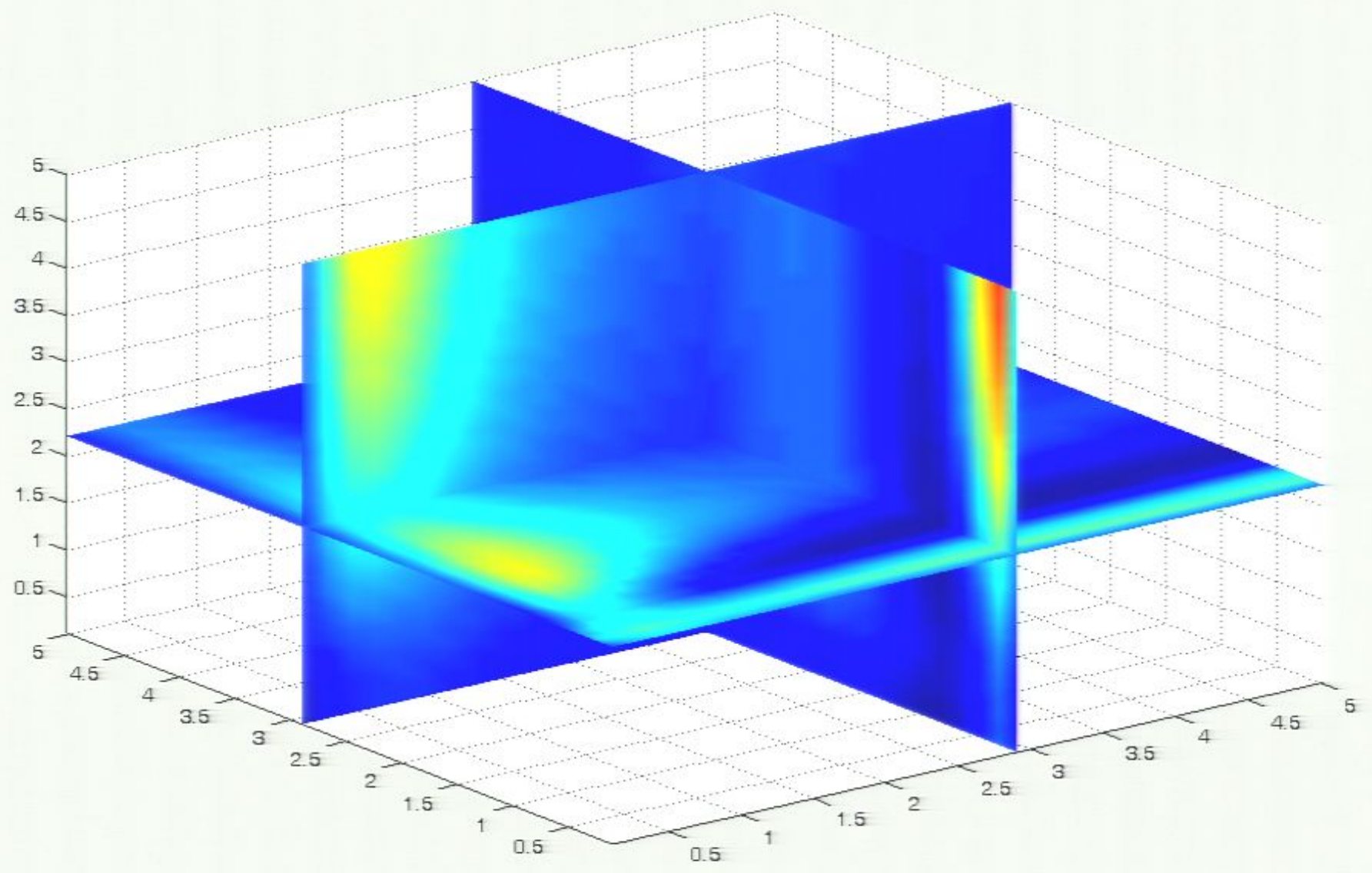
```

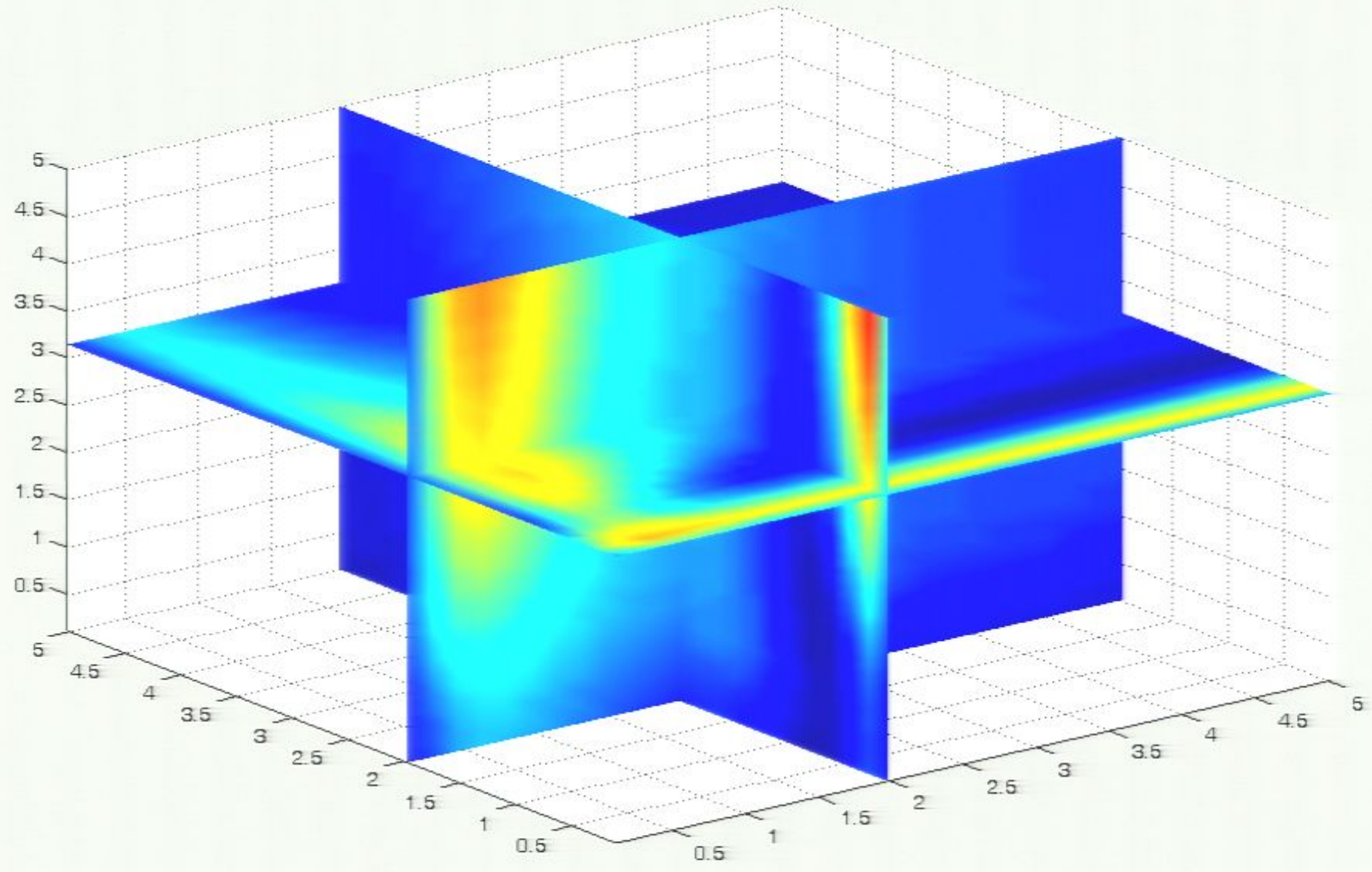
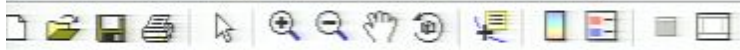


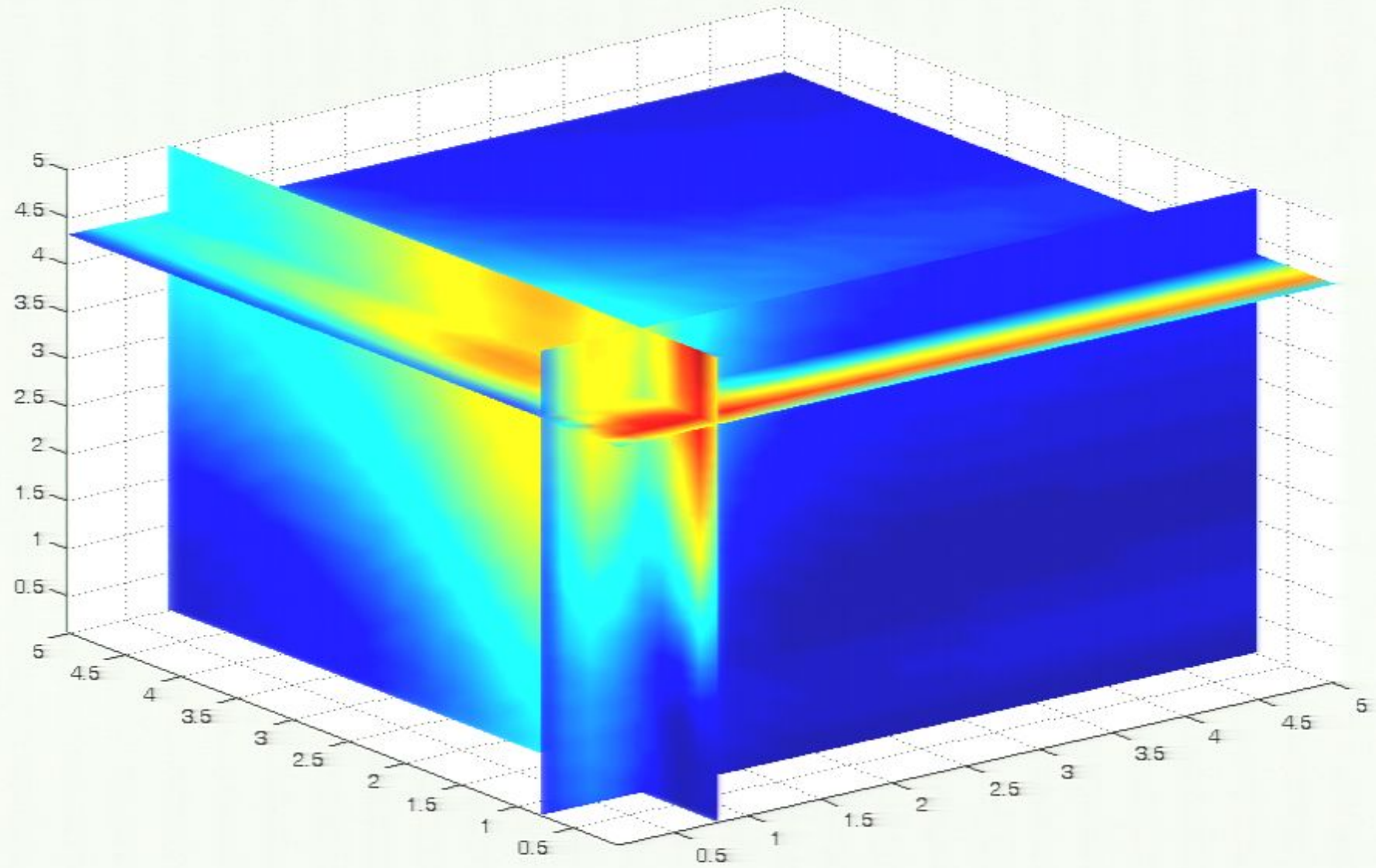
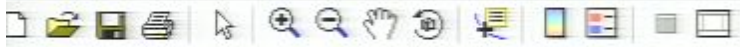




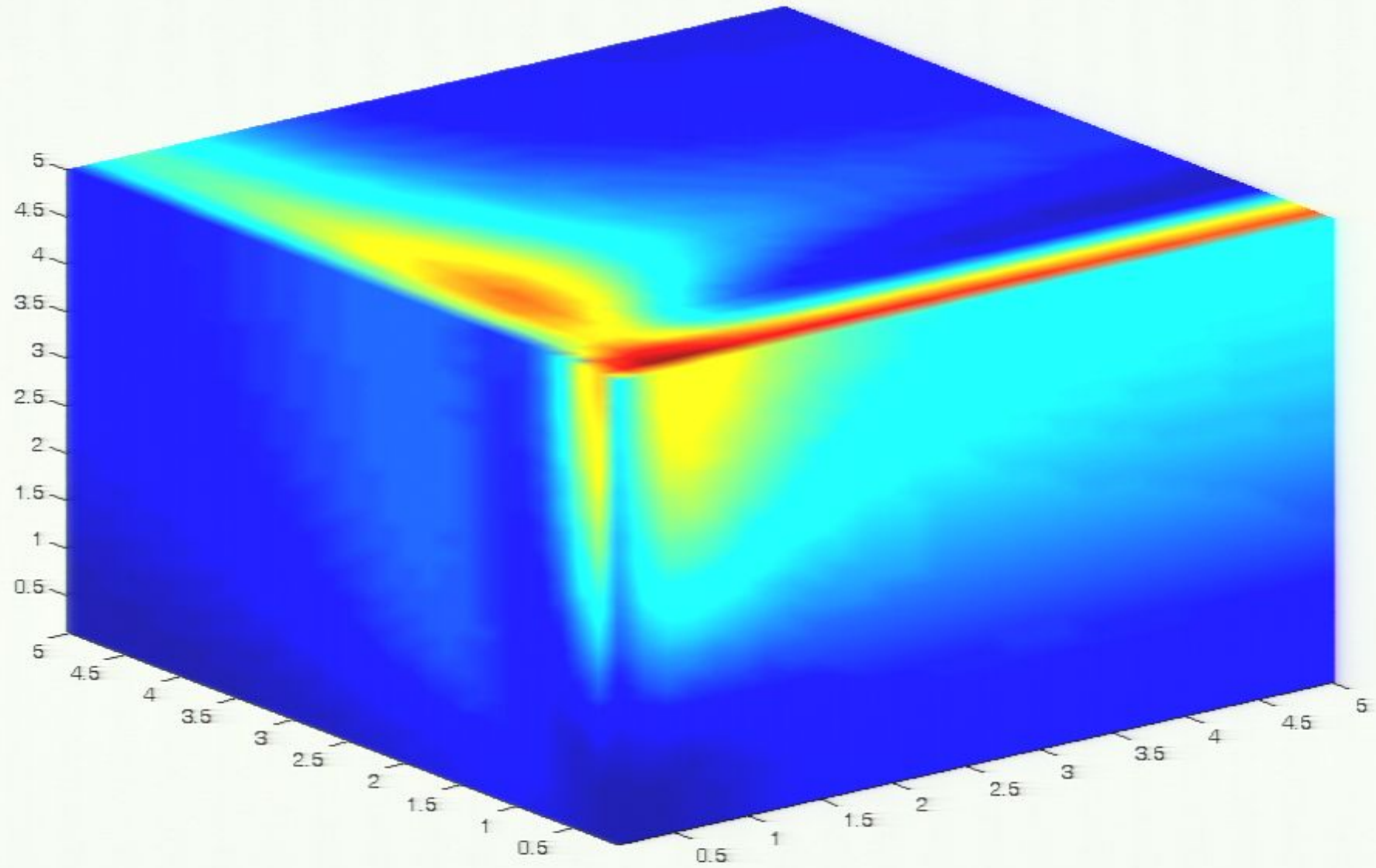












```
1 function isoFoM_movie
2 % movie of iso FoM surface - either slice or isosurface
3 % BB 1 June 09
4
5 - close all
6 %load v_30_new.mat; old data
7 - load new_detf_fom_vol_out.mat % new DETF FoM data
8 %v = v_30_new;
9 - v = fom_vol;
10 - H=linspace(0.1,5,30); da=linspace(0.1,5,30); G=linspace(0.1,5,30);
11 - [x y z] = meshgrid(H,da,G);
12
13 - xmin = 0.1;
14 - xmax = 5;
15 - ymin = 0.1;
16 - ymax = 5;
17 - zmin = 0.1;
18 - zmax = 5;
19 - minv = min(min(min(v)));
20 - maxv = max(max(max(v)));
21
22 - figure(2)
23 - whitebg('white') % colour of background
24 %*****
25 - n = 75; % number of threshold steps
26 - step = (maxv-minv)./n; % union step in threshold
27 - thresh = minv:step:maxv; % threshold vector
28 %*****
29
30 - slice_flag = 1; % 1 if slice animation, 0 if union
31
32 - nslice = 50;
33
34 %*****
```