

Title: Ekpyrotic/Cyclic Cosmology - Lecture 3

Date: Jun 29, 2009 10:00 AM

URL: <http://pirsa.org/09060080>

Abstract:

III Cosm. Pert. : Entropic mechanism + Non-Gaussianity

2 scalars

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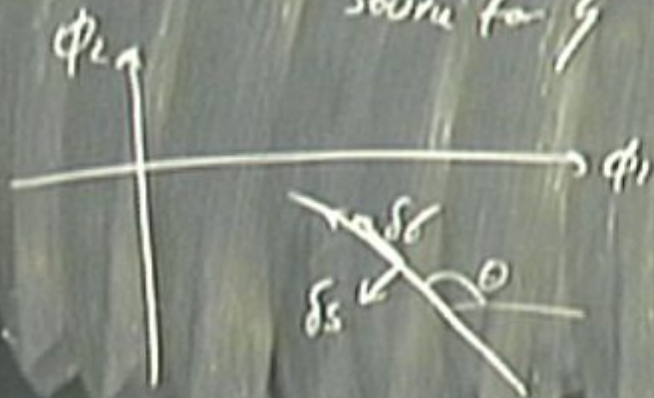
2 scalars

relative pert $\equiv$ entropy pert

2 scalars

relative part $\equiv$ entropy part ; gauge inv.

same for ϕ



III Cosm. Pert.: Entropic mechanism + Non-Gaussianity

2 scalars

relative pert $\equiv$ entropic pert, gauge inv. $\dot{\sigma} = \dot{\phi}_1 \cos \theta + \dot{\phi}_2 \sin \theta$
 source for ζ $\cos \theta = \frac{\dot{\phi}_1}{\sigma}$



III Cosm. Pert.: Entropic mechanism + Non-Gaussianity

2 scalars

relative pert $\equiv$ entropy pert, gauge inv.

source for

$$\dot{\sigma}^L = \dot{\phi}_1^L + \dot{\phi}_2$$

$$\dot{\sigma} = \cos \theta \cdot \dot{\phi}_1 + \sin \theta \cdot \dot{\phi}_2$$

$$\cos \theta = \frac{\dot{\phi}_1}{\sigma}$$

$$\delta \sigma = \cos \theta \delta \phi_1$$

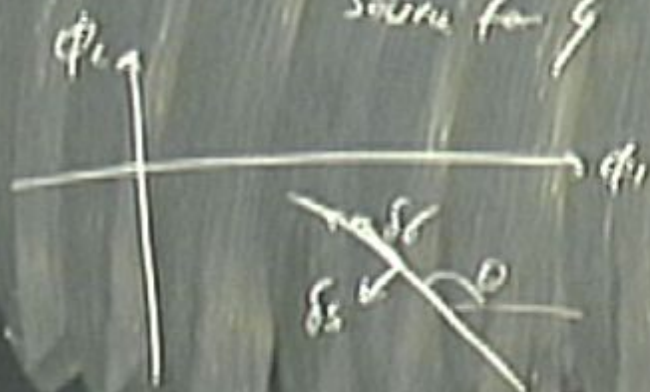


σ

III Cosm. Pert.: Entropic mechanism + Non-Gaussianity

2 scalars

relative part $\equiv$ entropy part, gauge inv. $\vec{\phi} = |\cos\theta \cdot \phi_1 + \sin\theta \cdot \phi_2|$
 same for ζ $\cos\theta = \frac{\phi_1}{\sigma}$

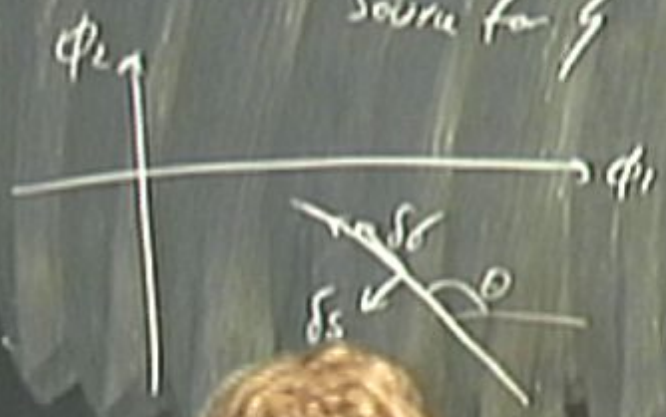


$$\delta\phi = \cos\theta \delta\phi_1 + \sin\theta \delta\phi_2$$

$$\delta\zeta = \cos\theta \delta\phi_2 - \sin\theta \delta\phi_1$$

adiabatic mode
 entropic mode

relative part $\equiv$ entropy part, gauge inv. $\phi = \cos \theta \cdot \phi_1 + \sin \theta \cdot \phi_2$
 source for ξ $\cos \theta = \frac{\dot{\phi}_1}{\sigma}$



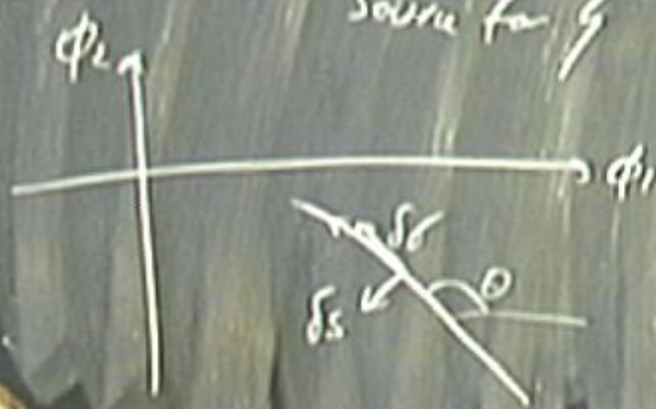
$$\delta\phi = \cos \theta \delta\phi_1 + \sin \theta \delta\phi_2 \quad \text{adiabatic mode}$$

$$\delta s = \sin \theta \delta\phi_1 - \cos \theta \delta\phi_2 \quad \text{entropic mode}$$

$$\delta\phi_1 \rightarrow \delta\phi_1 + \angle_\xi \phi_1 = \delta\phi_1 = \xi^\mu \partial_\mu \phi_1$$



relative part $\equiv$ entropy part ; gauge inv. $\phi = \cos \theta \cdot \phi_1 + \sin \theta \cdot \phi_2$
 source for ξ $\cos \theta = \frac{\dot{\phi}_1}{\sigma}$



$$\delta\phi = \cos \theta \delta\phi_1 + \sin \theta \delta\phi_2 \quad \text{adiabatic mode}$$

$$\delta s = \sin \theta \delta\phi_1 - \cos \theta \delta\phi_2 \quad \text{entropic mode}$$

$$\delta\phi_1 \rightarrow \delta\phi_1 + \angle_\xi \phi_1 = \delta\phi_1 + \int^\mu \partial_\mu \phi_1$$

$$\downarrow$$

$$\int^\mu \dot{\phi}_1 \cdot \phi_1(t)$$

eqns of motion

$$\delta\ddot{s} + 3H\delta\dot{s} + \left(\frac{k^2}{a^2} + \underline{V_{,ss}} + 3\dot{\Theta}^2\right)\delta s = \frac{\partial}{\partial\phi} \frac{4k^2}{a^2} \Phi$$

$$\zeta = -\frac{2H}{\dot{\phi}} \dot{\Theta} \delta s$$

↳ Newtonian pert.

eqns of motion

$$\delta\ddot{s} + 3H\delta\dot{s} + \left(\frac{k^2}{a^2} + \underline{V_{,ss}} + 3\dot{\Theta}^2\right)\delta s = -\frac{\dot{\Theta}}{\dot{\phi}} \frac{4k^2}{a^2} \Phi$$

↳ Newtonian pert.

$$\ddot{\zeta} = -\frac{2H}{\dot{\phi}} \dot{\Theta} \delta s$$

$\zeta(\delta\phi, \delta s, \text{rot}, \mu)$

eqns of motion

$$\left\{ \delta \ddot{s} + 3H \delta \dot{s} + \left(\frac{k^2}{a^2} + \underline{V_{,ss}} + 3\dot{\theta}^2 \right) \delta s = - \frac{\partial}{\partial s} \frac{4k^2}{a^2} \Phi \right.$$

$$\left\{ \ddot{\zeta} = - \frac{2H}{\dot{\zeta}} \dot{\theta} \delta s \quad (\text{large-scale}) \right.$$

$\zeta \text{ (} \delta \psi, \delta \phi, \text{rot} \text{)}_{\text{rot}}$

↳ Newtonian pert.

background

$$V_{el} = -V_0 e^{-\sqrt{2}\epsilon^6} [1 + \epsilon s^2 + \dots]$$

~~XXXX~~

background

$$V_{\text{eff}} = -V_0 e^{-\sqrt{2\epsilon} \phi} [1 + \epsilon s^2 + \dots]$$

ekpyrotic pot. for ϕ

tachyonic pot. for s

$$V_{\text{eff}} = -V_0 e^{\epsilon s^2} \left[1 + \epsilon s^2 + \dots \right]$$

ekpyrotic pol. for ϵ

tachyonic pol. for s

scaling sol

$$a \sim (-t)^{1/\epsilon} \quad \sigma = -\sqrt{\frac{2}{\epsilon}} \ln(-\sqrt{\epsilon} t) \quad s=0$$

scaling sol

$$a \sim (t-t_c)^{\alpha} \quad \phi = -\sqrt{2} \ln(-t/t_c) \quad \gamma = 0$$

$\theta = 0$ along ridge is potential

$$V_{el} = -V_0 e$$

$$[1 + \epsilon s^2 + \dots]$$

ekpyrotic pot. for σ

tachyonic pot. for s

$$V_{ss} = 2\epsilon V(s=0) < 0$$

scaling sol

$$a \sim (-t)^{1/\epsilon}$$

$$\sigma = -\sqrt{\frac{2}{\epsilon}} \ln(-\sqrt{\epsilon} V_0 t)$$

$$s=0$$

$$\dot{\theta} = 0$$

along ridge in potential

$$V_{100} = 22 \text{ V}$$
$$\text{for } s=0 \quad V_{100} = V_{ss}$$

$$V_{100} = 22 \text{ V}$$

for $s=0$ $V_{100} = V_{155}$
 $V_{15} = 0$ $V_{10} \neq 0$

$$V_{100} = 2\varepsilon V$$

$$\text{for } s=0 \quad V_{100} = V_{155}$$

$$V_{15} = 0 \quad V_{10} \neq 0$$

generation of pert

$$\ddot{\delta s} + 3H\dot{\delta s} + \left(\frac{k^2}{q^2} + V_{,11}\right)\delta s = 0$$

$$V_{,\sigma\sigma} = 2\varepsilon V$$

$$\text{for } s=0 \quad V_{,\sigma\sigma} = V_{,ss}$$

$$V_{,s} = 0 \quad V_{,\sigma} \neq 0$$

generation of pert

$$\ddot{\delta s} + 3H\dot{\delta s} + \left(\frac{k^2}{a^2} + V_{,ss}\right)\delta s = 0$$

$$\text{conf time } dt = a d\tau \quad \text{if } \delta S = a \delta s$$

$$\delta S'' + \left(k^2 - \frac{a''}{a} + a^2 V_{,ss}\right)\delta S = 0$$

$$V_{,\sigma\sigma} = 2\varepsilon V$$

$$\text{for } s=0 \quad V_{,\sigma\sigma} = V_{,ss}$$

$$V_{,ss} = 0 \quad V_{,\sigma\sigma} \neq 0$$

generation of pert

$$\delta\ddot{s} + 3H\delta\dot{s} + \left(\frac{k^2}{a^2} + V_{,ss}\right)\delta s = 0$$

$$\text{conf time } dt = a d\tau$$

$$\delta\ddot{s} + \left(k^2 - \frac{\ddot{a}}{a} + a^2 V_{,ss}\right)\delta s = 0$$

$$\frac{\ddot{a}}{a} = a^2 H^2 (2 - \varepsilon)$$

$$\text{exp in } \frac{1}{\varepsilon}$$

$$V_{,ss} = 2\varepsilon V$$

$$\text{for } s=0 \quad V_{,ss} = V_{,ss}$$

$$V_{,s} = 0 \quad V_{,ss} \neq 0$$

generation of pert

$$\delta \ddot{s} + 3H \delta \dot{s} + \left(\frac{k^2}{a^2} + V_{,ss} \right) \delta s = 0$$

$$\text{conf time } dt = a d\tau \quad \text{def } \delta S = a \delta s$$

$$\delta S'' + \left(k^2 - \frac{\ddot{a}}{a} + a^4 V_{,ss} \right) \delta S = 0$$

$$\frac{\ddot{a}}{a} = a^4 H^4 (2 - \varepsilon)$$

exp in $\frac{a^2}{\varepsilon}$

$$dN = d \ln a$$

$$\frac{d\varepsilon}{dH}$$

$$\text{diff } \epsilon = \frac{\dot{\sigma}^2}{2H^4} \quad \ddot{\sigma} + 3H\dot{\sigma} + V_{,\sigma} = 0 \quad V_{1ss} = V_{,ss}$$

$$\frac{V_{1ss}}{H^4} = -2\epsilon^4 + 6\epsilon^3 + \frac{5}{2}\epsilon_{,N} + \dots$$

$$\text{diff } \epsilon = \frac{\dot{\sigma}^2}{2H^4}$$

$$\ddot{\sigma} + 3H\dot{\sigma} + V_{,\sigma} = 0$$

$$V_{1ss} = V_{,\sigma\sigma}$$

$$\frac{V_{1ss}}{H^4} = -2\epsilon^2 + 6\epsilon + \frac{5}{2}\epsilon_{,\nu} + \dots$$

$$\alpha H = \frac{1}{\epsilon\tau} \left(1 + \frac{1}{\epsilon} + \frac{\epsilon_{,\nu}}{\epsilon^2} \right)$$

$$\text{diff } \epsilon = \frac{\dot{\sigma}}{2H^2} \quad \ddot{\sigma} + 3H\dot{\sigma} + V_{,\sigma} = 0 \quad V_{,ss} = V_{,\sigma\sigma}$$

$$\frac{V_{,ss}}{H^2} = -2\epsilon^2 + 6\epsilon' + \frac{5}{2}\epsilon_{,\mathcal{N}} + \dots$$

$$\alpha H = \frac{1}{\epsilon\tau} \left(1 + \frac{1}{\epsilon} + \frac{\epsilon_{,\mathcal{U}}}{\epsilon^2} \right)$$

$$\frac{\alpha''}{\alpha} - n^2 V_{,ss} = \frac{2}{\tau^2} \left(1 - \frac{3}{2\epsilon} + \frac{3}{4} \frac{\epsilon_{,\mathcal{N}}}{\epsilon^2} \right)$$

$$\frac{a''}{a} - a^2 V_{,rr} = \frac{2}{r^2} \left(1 - \frac{3}{2\varepsilon} + \frac{3}{4} \frac{\varepsilon_{,rr}}{\varepsilon^2} \right)$$

solved in terms of, Horndeski's

Minkowski vacuum as $t \rightarrow -\infty$

$$\delta S = \frac{\sqrt{-\pi\epsilon}}{2} H_\nu^{(1)}(-k\tau) \quad \nu = \frac{3}{2} \left(1 - \frac{2}{3\epsilon} + \frac{\epsilon_H}{3\epsilon^2} \right)$$

Minkowski vacuum $\alpha, \tau \rightarrow -\infty$

$$\delta S = \frac{\sqrt{-\pi\epsilon}}{2} H_{\nu}^{(1)}(-k\tau) \quad \nu = \frac{3}{2} \left(1 - \frac{2}{3\epsilon} + \frac{\epsilon\nu}{3\epsilon^2} \right)$$

late time $\tau \rightarrow 0$

$$\delta S = \frac{1}{\sqrt{2}(-\tau)k^{\nu}}$$

Minkowski vacuum as $t \rightarrow -\infty$

$$\delta S = \frac{\sqrt{-\pi\epsilon}}{2} H_\nu^{(1)}(-k\tau) \quad \nu = \frac{3}{2} \left(1 - \frac{2}{3\epsilon} + \frac{\epsilon\nu}{3\epsilon^2} \right)$$

$$V = -\frac{1}{\epsilon t^2}$$

late time $\tau \rightarrow 0$

$$\delta S = \frac{1}{\sqrt{2}(-\tau)k^\nu}$$

at end of chaotic phase

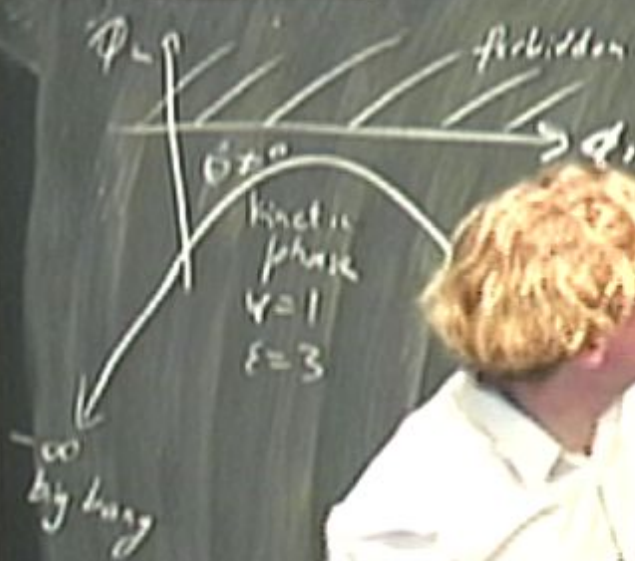
$$\delta S \approx \frac{| \epsilon V_{\text{chaotic}} |^{1/2}}{k^\nu}$$

Conversion

ϕ_L

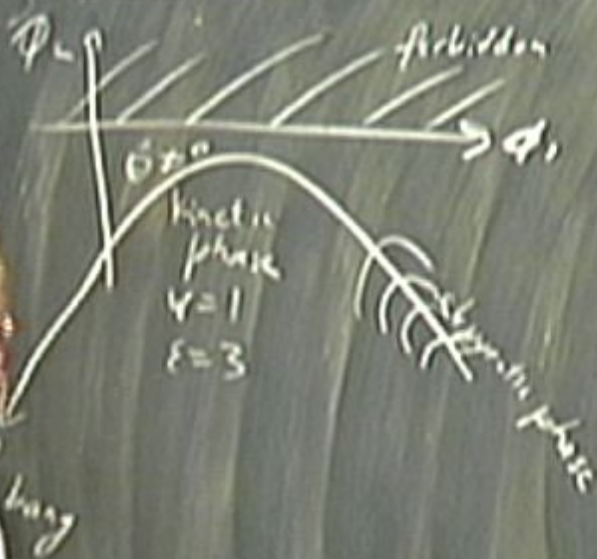
ϕ_1

Conversion



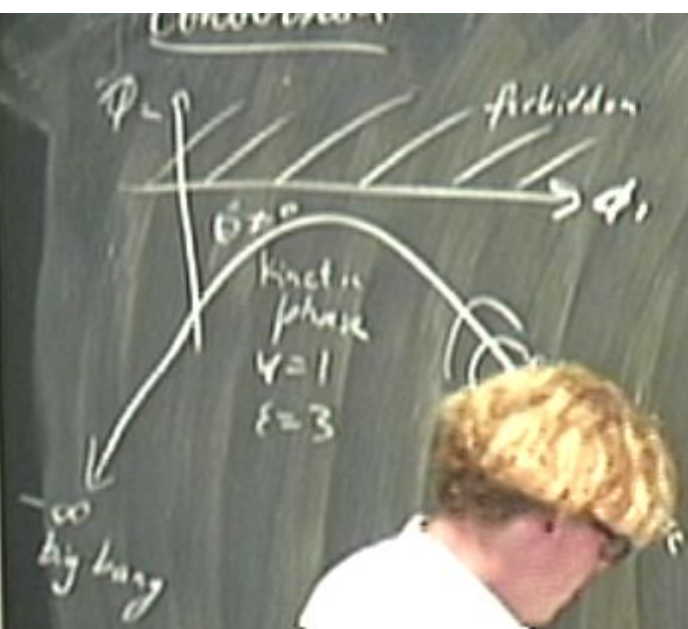
$$\dot{\zeta} = -\frac{2H}{\sigma} \dot{\phi} \mathcal{S}$$

Conversion



$$\dot{\zeta} = - \underbrace{\frac{2H}{\sigma}}_{\text{no } k\text{-dependence}} \dot{\sigma} \delta S$$

$$3H^2 = \frac{1}{2} \dot{\sigma}^2 \quad \frac{H}{\dot{\sigma}} = -\frac{1}{\sqrt{6}}$$



$$\dot{\zeta} = - \underbrace{\frac{2H}{\dot{\sigma}} \dot{\sigma}}_{\text{no } k\text{-dependence}} \dot{S} S$$

$$3H^2 = \frac{1}{2} \dot{\sigma}^2 \quad \frac{H}{\dot{\sigma}} = -\frac{1}{\sqrt{6}}$$

approximate $\dot{S} S$ const during conversion

phase

too big long

$$\int \dot{\theta} = \Delta \theta \sim O(1) \text{ rad}$$

$$\zeta = \int \frac{2}{\pi} \Delta \theta \, \delta s$$

ζ [rad], [deg], [mrad]

astro-ph/0005131



too big bang

phase

along horizon

$$\int \dot{\theta} = \Delta\theta \sim O(1) \text{ rad}$$

$$\zeta = \int \frac{2}{16} \Delta\theta \delta S$$

$$\zeta \approx \delta S_{\text{cl-end}} \quad (k \ll 1)$$



astro-ph/0003131

$$\langle \epsilon^2 \rangle = \int \frac{d^3k}{(2\pi)^3} \frac{|\epsilon V_{el-mel}|}{k^{2\nu}}$$

$$= \int \frac{dk}{k} \frac{\epsilon |V_{el-mel}|}{2\pi^2} k^{n_s-1}$$

$$n_s =$$

$$\begin{aligned}
 \langle \epsilon^2 \rangle &= \int \frac{d^3k}{(2\pi)^3} \frac{|\epsilon V_{el-rod}|}{k^{2\nu}} \\
 &= \int \frac{dk}{k} \frac{\epsilon |V_{el-rod}|}{2\pi^2} k^{n_s-1} \\
 n_s &= 1 + \frac{2}{\epsilon} - \frac{\epsilon_{1d}}{\epsilon^2}
 \end{aligned}$$

$$\langle \eta^2 \rangle = \int \frac{d^3k}{(2\pi)^3} \frac{|\varepsilon V_{el-ax}|}{k^{2\nu}}$$

$$= \int \frac{dk}{k} \frac{\varepsilon |V_{el-ax}|}{2\pi^2} k^{n_s-1}$$

$$n_s = 1 + \frac{2}{\varepsilon} - \frac{\varepsilon_{UV}}{\varepsilon^2}$$

$$0.92 < n_s < 1.02$$



2) ekpyrotic conditions

$$\frac{H}{\sigma} \sim \frac{1}{\sqrt{c}}$$

$$\zeta \approx \frac{1}{\sqrt{c}} \delta_3$$



2) ekpyrotic conditions

$$\frac{H}{\delta} \sim \frac{1}{\sqrt{c}}$$

$$\zeta \approx \frac{1}{\sqrt{c}} \delta \zeta$$

GK - don't expect to see any

Non-gaussianity

Non-gaussianity

linear eqns $\rightarrow$ quadratic interaction — prob. $\sim e^{-S}$ Gaussian dist.

quadratic

Non-gaussianity

linear eqn → quadratic interaction — prob. $\sim e^{-S}$ Gaussian dist.
quadratic cubic
cubic quadratic kurtosis

Non-gaussianity

linear eqn → quadratic interaction
quadratic cubic
cubic quadratic

$$prob \sim e^{-5}$$

Gaussian dist.
Gaussian
Gaussian



$$r\zeta = \underbrace{\zeta}_{\text{gauge}} + \frac{3}{5} f m \zeta^2 + \frac{3}{25} g m \zeta^3$$



SALES
REPRESENTATIVE
JAMES
JAMES

$$r \zeta = \underbrace{\zeta_\mu}_{\text{gallia}} + \frac{3}{5} f_m \zeta_\mu^2 + \frac{3}{25} g_m \zeta_\mu^3$$

$$f_m = 40 \pm 40 \text{ (20)}$$

0-10-10-10
0-10-10-10
0-10-10-10
0-10-10-10

$$r \cdot \zeta = \zeta_{\mu} + \frac{3}{5} f_{\mu} \zeta_{\mu}^2 + \frac{3}{25} g_{\mu} \zeta_{\mu}^3$$

ζ_{μ} geolin

$$f_{\mu} = 60 \pm 40 \text{ (20)}$$

$$V_{\text{eff}} = -V_0 e^{-12\pi\sigma} \left[1 + \epsilon s^2 + \frac{K_2}{s!} \epsilon^2 s^2 + \frac{K_3}{s!} \epsilon^3 s^3 + \dots \right]$$



$f_m = 40 \pm 40$ (20)

$$V_{el} = -V_0 e^{-i2\pi\sigma} \left[1 + \epsilon s^2 + \frac{K_1}{s!} \epsilon^2 s^2 + \frac{K_2}{s!} \epsilon^2 s^2 + \dots \right]$$

kinetic conversion

during elyptosis

$$\delta s = \delta s_{el}$$

$$\frac{1}{t}$$

f_{geom} $f_M = 40 \pm 40$ (20)

$$V_{\text{rel}} = -V_0 e^{-\sqrt{2} \epsilon \sigma} \left[1 + \epsilon s^2 + \frac{K_1}{s!} \epsilon^2 s^2 + \frac{K_1}{4!} \epsilon^2 s^4 + \dots \right]$$

kinetic conversion

during ebbings

$$\delta s = \delta s_L + \frac{K_1 \sqrt{\epsilon}}{8} \delta s_L^2 + \epsilon \left(\frac{K_1}{60} + \frac{K_1^2}{80} - \frac{2}{5} \right) \delta s_L^3$$

$\frac{1}{t}$

$$V_{cl} = -V_0 c \left(1 + \epsilon S + \frac{\epsilon^2}{2} S^2 + \dots \right)$$

kinetic conversion

during algebraic

$$\delta S = \delta S_L + \frac{K_1 \sqrt{\epsilon}}{8} \delta S_L^2 + \epsilon \left(\frac{K_1}{80} + \frac{K_1^2}{80} - \frac{2}{\epsilon} \right) \delta S_L^3$$

$$f_{\text{res}} \sim \frac{3}{2} K_1 \sqrt{\epsilon}$$

dynamic phase

$$\delta S \approx \frac{|i V_{cl, \text{res}}|}{k^2}$$

$$V_{cl} = -V_0 \epsilon \left(1 + \epsilon S + \frac{1}{3!} \epsilon^2 S^2 + \frac{1}{5!} \epsilon^4 S^4 + \dots \right)$$

kinetic conversion

during adiabatic

$$S_S = S_{S_L} + \frac{K_1 \sqrt{\epsilon}}{8} S_{S_L}^2 + \epsilon \left(\frac{K_1}{80} + \frac{K_1^2}{80} - \frac{2}{5} \right) S_{S_L}^3$$

$$\begin{cases} f_{ac} \sim \frac{3}{2} K_1 \sqrt{\epsilon} \\ g_{ac} \sim 100 \epsilon \left(\frac{K_1}{80} + \frac{K_1^2}{80} - \frac{2}{5} \right) \end{cases}$$

at end of adiabatic phase

$$S_S \approx \frac{|V_{cl, end}|^2}{K^2}$$

$$V_{\text{eff}} = -V_0 \epsilon$$

kinetic conversion

during ekyronis

$$\delta S = \delta S_L + \frac{K_1 \sqrt{\epsilon}}{8} \delta S_L^2 + \epsilon \left(\frac{K_1}{80} + \frac{K_1^2}{80} - \frac{2}{5} \right) \delta S_L^3$$

$$\begin{cases} f_{\text{me}} \sim \frac{1}{2} K_1 \sqrt{\epsilon} & \sim O(10) \\ g_{\text{me}} \sim 100 \epsilon \left(\frac{K_1}{80} + \frac{K_1^2}{80} - \frac{2}{5} \right) \sim O(1000) \end{cases} \quad \epsilon \sim 50$$

at end of ekyronis phase

$$\delta S \approx \frac{|V_{\text{eff-end}}|^{1/2}}{k^2}$$

during egyptic $\delta S = \delta S_L + \frac{K_1 \sqrt{\epsilon}}{8} \delta S_L^2 + \epsilon \left(\frac{K_1}{80} + \frac{K_1^2}{80} - \frac{2}{5} \right) \delta S_L^3$

$\begin{cases} f_{12} \sim \frac{3}{2} K_1 \sqrt{\epsilon} \\ g_{12} \sim 100 \epsilon \left(\frac{K_1}{80} + \frac{K_1^2}{80} - \frac{2}{5} \right) \sim \alpha(-1000) \end{cases}$ $\sim 0(10)$ $\epsilon \sim 50$

late time $\tau \rightarrow \infty$ $\delta S = \frac{1}{\sqrt{1-\tau} L_h}$

at end of egyptic phase

$\delta S \approx \frac{|E V_{\text{end}}|}{k^2}$