

Title: Dilution of the Cosmological Constant: Higher Codimension Branes and Higher Curvature Terms

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Abstract: TBA

Dilution of the Cosmological Constant

Higher Codimension Branes and Higher Curvature Terms

Alberto Iglesias

University of California, Davis

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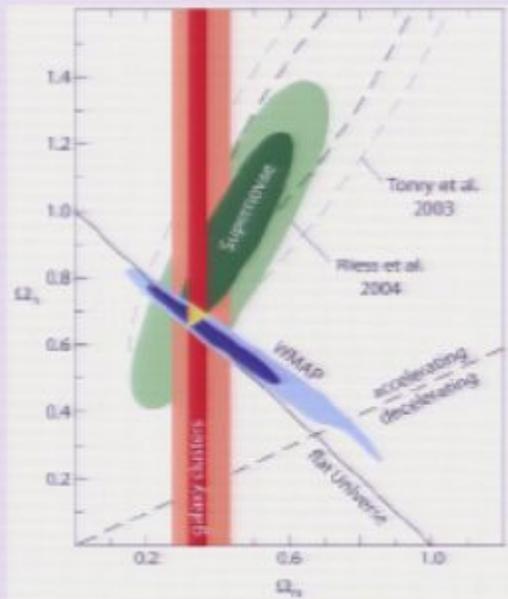
Outline

- Cosmological Constant Problem
- Infinite Volume Extra Dimensions
Setup: higher curvature terms \rightarrow Gauss-Bonnet
- Codimension 2 case
- Higher Codimension
Singularities and smoothing
- Codimension 3 case
- Conclusions

Cosmological Constant Problem

$$\mathcal{L} = M_{pl}^2 \int d^4x \sqrt{-g} (R - 2\Lambda) + \mathcal{L}_{SM}$$

Observations



$$M_{pl}^2 \Lambda_{obs} \sim (10^{-3} eV)^4$$

QFT with cutoff

$$\langle \mathcal{E} \rangle = \int_0^{M_{cutoff}} k^2 dk \frac{\omega_k}{2}$$

$$M_{pl}^2 \Lambda_{QFT} \sim M_{cutoff}^4$$

Goal

- Since $\Lambda_{obs} \ll \Lambda_{QFT}$
- Find a way to make geometry of our $3 + 1$ spacetime *insensitive* to Λ_{QFT} → embed in extra dimensions
- Use $\langle \mathcal{E} \rangle$ to deform extra dimensions
 - Codimension 2: create a deficit angle
 - Higher codimension: create curvature
- In particular: find compatibility of **Minkowski $3 + 1$** for a wide range of values of Λ

Setup

Brane of codimension d in bulk of dimension D

$$S = \hat{M}^{D-4} \int_{\Sigma_{D-d}} (\hat{R} - \hat{\Lambda}) + M^{D-2} \int_{\mathcal{M}_D} [R + \xi(R^2 - 4R_{MN}^2 + R_{MNR S}^2)]$$

Codimension 2

Codimension 2

$$S = - \int_{\Sigma_{D-2}} f + M^{D-2} \int_{\mathcal{M}_D} [R + \xi(R^2 - 4R_{MN}^2 + R_{MNRs}^2)]$$

Ansatz:

$$ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu + e^{2\omega} \delta_{ij} dx^i dx^j$$

EoM:

$$\partial^j \partial_{i\omega} = -\frac{f}{2M^{D-2}} \delta^{(2)}(x^i)$$

$$\omega = -\frac{f}{8\pi M^{D-2}} \log\left(\frac{\rho^2}{a^2}\right)$$

$$ds_2^2 = dr^2 + e^{-2\beta} r^2 d\phi^2$$

$$r \sim \rho^{1-\theta/2\pi}$$

$$e^{-\beta} = 1 - \frac{f}{4\pi M^{D-2}}$$

$$\theta = 2\pi[1 - e^{-\beta}] = \frac{f}{2M^{D-2}}$$

$$0 < f < 4\pi M^{D-2}$$

Linearized Gravity

- At linear level EoM coincide with those for

$$S = \hat{M}^{D-4} \int_{\Sigma_{D-2}} (\hat{R} - \hat{\Lambda}) + M^{D-2} \int_{\mathcal{M}_D} R + S_{int}$$

$$\hat{M}^{D-4} = 2\xi f, \quad \hat{\Lambda} = 1/2\xi$$

- Some of the diffs are broken (tension)

$$\delta h_{MN} = \nabla_M \xi_N + \nabla_N \xi_M$$

$$0 = \partial_i \xi_\mu |_{r=0}$$

$$0 = \partial^j [\xi_j \delta^{(2)}(x^j)]$$

Solution to Linearized EoM

$p^2 \neq 0$

$$H_{\mu\nu} = \chi = 0 \text{ for } p \neq 0$$

$$H_{\mu\nu}(p, p \neq 0) = \frac{\hat{M}^{4-D}}{p^2} \left[T_{\mu\nu}(p) - \frac{1}{D-4} \eta_{\mu\nu} T(p) + \frac{2}{D-4} \frac{p_\mu p_\nu}{p^2} T(p) \right]$$

$$\chi(p, p \neq 0) = \frac{2}{D-4} \frac{\hat{M}^{4-D}}{p^2} T(p)$$

$p^2 = 0$

$$\chi(p) = \frac{M^{2-D}}{2\pi(D-2)} T(p) \log \left(\frac{\rho^2 + \epsilon^2}{b^2} \right)$$

$$H_{\mu\nu}(p) = -\frac{M^{2-D}}{4\pi} \left[T_{\mu\nu}(p) - \frac{1}{D-2} \eta_{\mu\nu} T(p) + \hat{M}^{D-4} p_\mu p_\nu \chi(0) \right] \log \left(\frac{\rho^2 + \epsilon^2}{b^2} \right)$$

Higher Codimension

Singularities and smoothing

Tension \rightarrow Background singularities

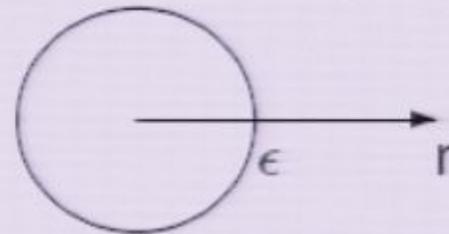
- Origin
- Finite distance (codim 3 and higher)

$(D - d)$ worldvolume (codimension d)

Replace brane worldvolume $R^{D-d-1,1}$ by $R^{D-d-1,1} \times S_\epsilon^{d-1}$

$R^{D-d-1,1}$
•

$R^{D-d-1,1} \times S_\epsilon^{d-1}$



$$S = \tilde{M}^{D-3} \int_{R^{D-d-1,1} \times S_\epsilon^{d-1}} (\tilde{R} - \tilde{\Lambda}) + M^{D-2} \int_{\mathcal{M}_D} [R + \xi(GB)]$$

$$S = \tilde{M}^{D-3} \int_{\Sigma_{D-1}} (\tilde{R} - \tilde{\Lambda}) + M^{D-2} \int_{\mathcal{M}_D} [R + \xi(R^2 - 4R_{MN}^2 + R_{MNRs}^2)]$$

Ansatz:

$$ds^2 = e^{2A(r)} \eta_{\mu\nu} dx^\mu dx^\nu + e^{2B(r)} (dr^2 + r^2 d\Omega_{d-1}^2)$$

↓

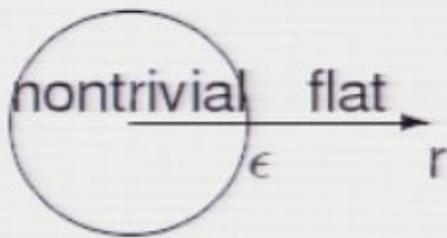
$$R^{D-d-1,1}$$

- EoM have no third or fourth derivatives
- Specialize to $d = 3$: simplification of EoM (GB)
- Want to find solutions with $\tilde{\Lambda}$ not determined by the other parameters

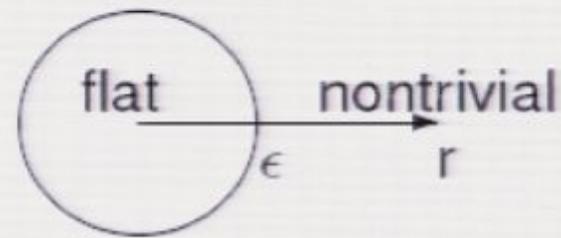
Structure of the EoM

$$\begin{aligned} (rr) \quad & \{A', B, B', B''\} = 0 \\ (\text{others}) \quad & \{A', A'', B, B', B''\} = \delta(r - \epsilon) \end{aligned}$$

- Flat Solution: A, B constant
- Nontrivial solutions
- Impose junction conditions at $r = \epsilon$



Interior Solutions



Exterior Solutions

Studied cases

No Einstein-Hilbert bulk term

- String in 5D
- Membrane in 6D
- 3-brane in 7D

EH plus GB bulk term

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String in 5D

$$S = \tilde{M}^2 \int_{R^{1,1} \times S^2} (\tilde{R} - \tilde{\Lambda}) + M^3 \int_{\mathcal{M}_3} \xi (R^2 - 4R_{MN}^2 + R_{MNRS}^2)$$

$$(A')^2 \left[3B'^2 + 6\frac{B'}{r} + \frac{2}{r^2} \right] = 0$$

$$(A'' + A'^2 - A'B')(B'^2 + 2\frac{B'}{r}) + 2A'(B' + \frac{1}{r})(B'' + \frac{B'}{r}) = \frac{e^{3B}}{8\xi} (\tilde{\Lambda} - 2\lambda)L\delta(r - \epsilon)$$

$$A'^2(B'' + \frac{B'}{r}) + 2(A'' + A'^2 - A'B')A'(B' + \frac{1}{r}) = \frac{\tilde{\Lambda}}{8\xi} e^{3B} L\delta(r - \epsilon)$$

Solutions

- A, B constant
- $A = \log(r_A/r)^{\alpha-1}, \quad B = \log(r_B/r)^\alpha, \quad \alpha_\pm = 1 \pm 1/\sqrt{3}$

- Singularities

$$ds^2 = (r_A/r)^{2(\alpha-1)} \eta_{\mu\nu} dx^\mu dx^\nu + (r_B/r)^{2\alpha} (dr^2 + r^2 d\Omega_2^2)$$

$$R \sim r^{2(\alpha-1)} = r^{\pm 2/\sqrt{3}} \quad \text{for } \alpha_{\pm}$$

$$\sigma \sim r^{\mp 2/\sqrt{3}} + \text{finite} \quad (\text{radial geodesics})$$

Interior sol: Choose α_+ Exterior sol: Choose α_-
 Geod. Complete, $R \rightarrow 0$ as $r \rightarrow 0$ (idem $r \rightarrow \infty$)

- Junction conditions

$$\pm \frac{\alpha(\alpha-1)(\alpha-2)}{\epsilon} = -\frac{\kappa-2}{8\xi} L(r_B/\epsilon)^\alpha$$

$$\pm \frac{(\alpha-1)^3}{\epsilon} = -\frac{\kappa}{8\xi} L(r_B/\epsilon)^\alpha$$

$$\xi = \frac{\sqrt{3}L\epsilon}{4(\alpha-1)} (r_B/\epsilon)^\alpha, \quad \tilde{\Lambda} = \frac{8L^2}{\xi^2} \quad \text{Fixed!}$$

String in 5D

$$S = \tilde{M}^2 \int_{R^{1,1} \times S^2} (\tilde{R} - \tilde{\Lambda}) + M^3 \int_{\mathcal{M}_3} \xi (R^2 - 4R_{MN}^2 + R_{MNRS}^2)$$

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3-brane in 7D with EH + GB bulk terms

Define

$$V = rA' , \quad U = rB' , \quad z = \xi^{-1} r^2 e^{2B}$$

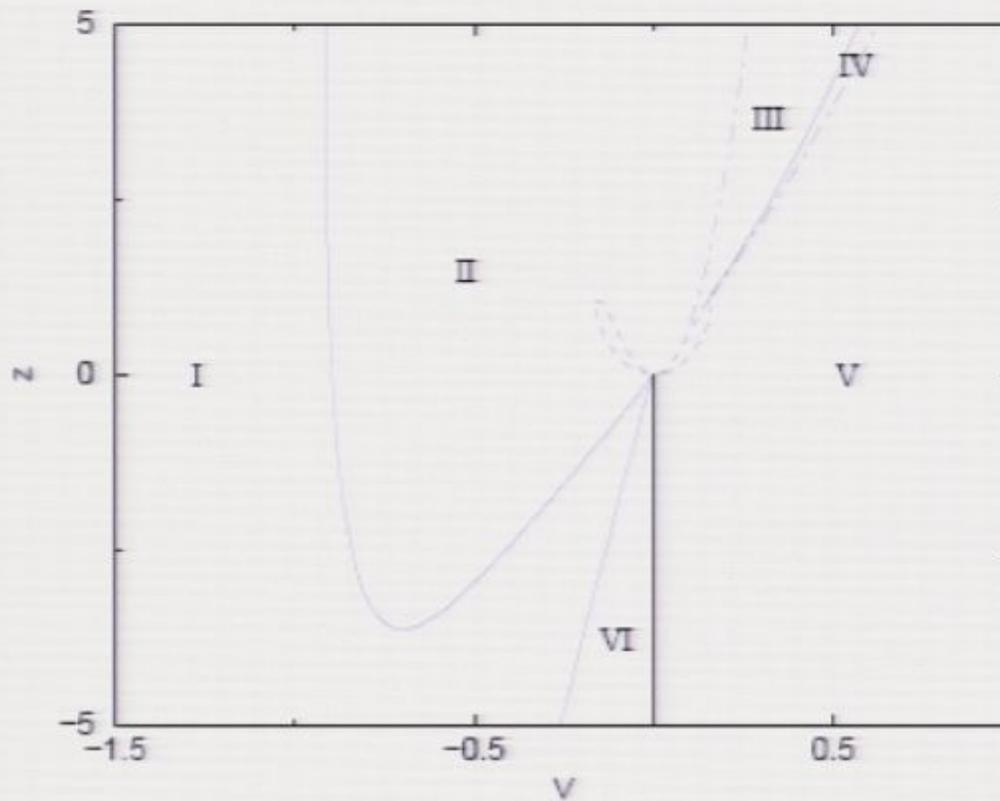
$$(rr) \quad 0 = U^2(72V^2 - z) - 8UV(z - 12V^2) - z(6V^2 - 1) - 12V^2(2 - V^2)$$

$$U \equiv U_{\pm}(V, z)$$

$$(rr - \alpha\beta) \quad 2zU \frac{dV}{dz} = \frac{g_1}{g_2} , \quad g_{1,2} \text{ pol in } V, z$$

- Impose matching conditions
- Study flow of solutions in (V, z) plane
- Approximate equations in regions of possible divergences
- Identify coordinate/true singularities

3-brane in 7D EH+GB bulk terms: + case



I,II,V,VI $\rightarrow \tilde{\Lambda} < 0$
III,IV $\rightarrow \tilde{\Lambda} > 0$

3-brane in 7D with EH + GB bulk terms

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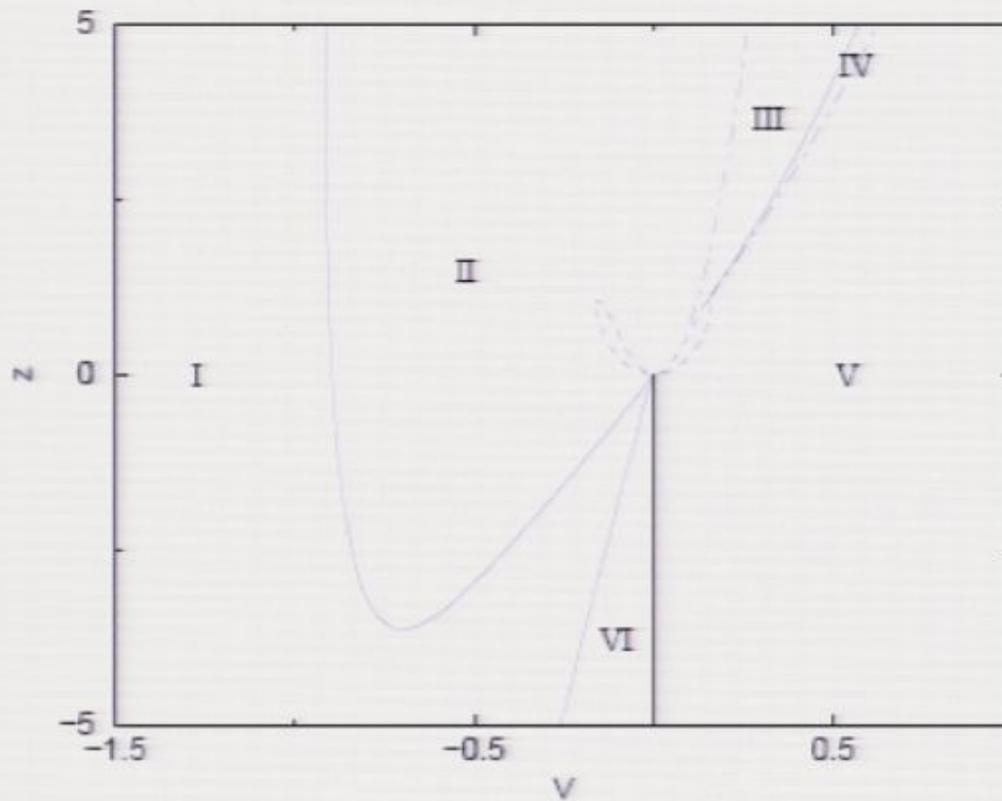
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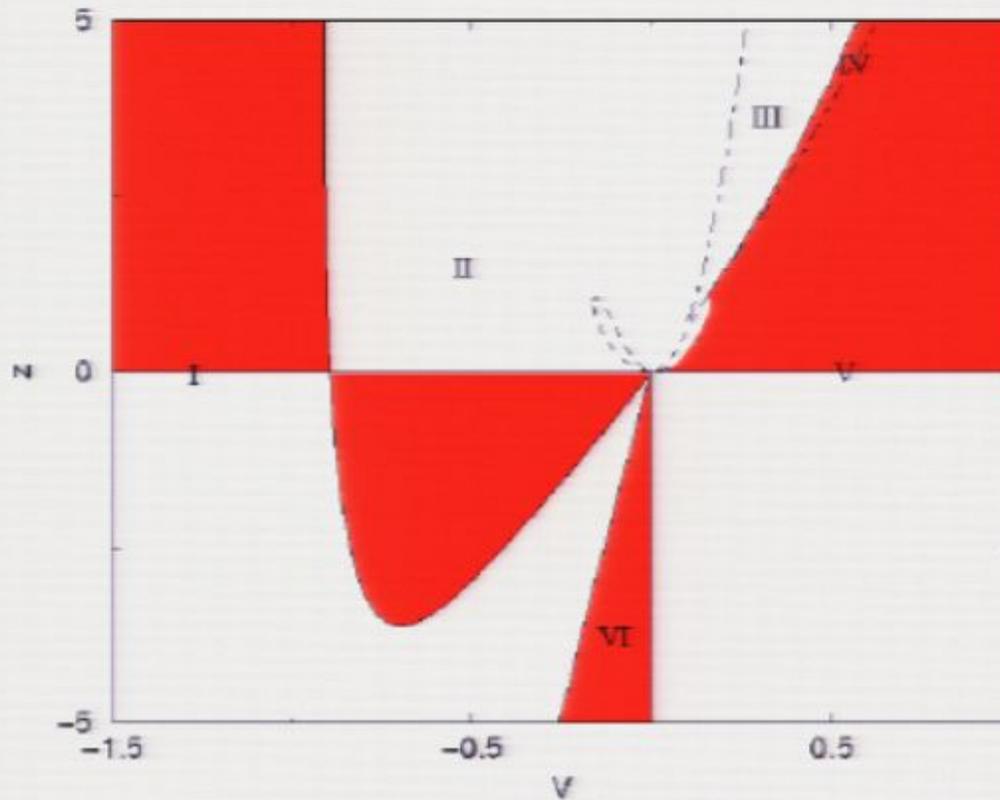
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3-brane in 7D EH+GB bulk terms: + case



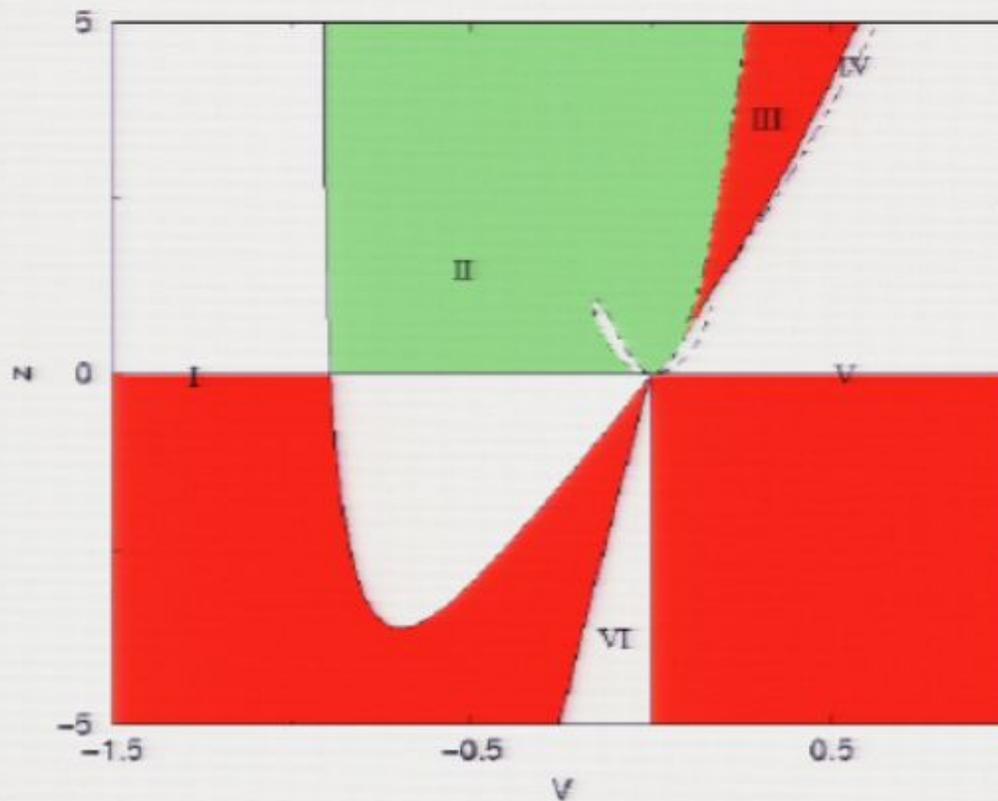
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3-brane in 7D EH+GB bulk terms: + case



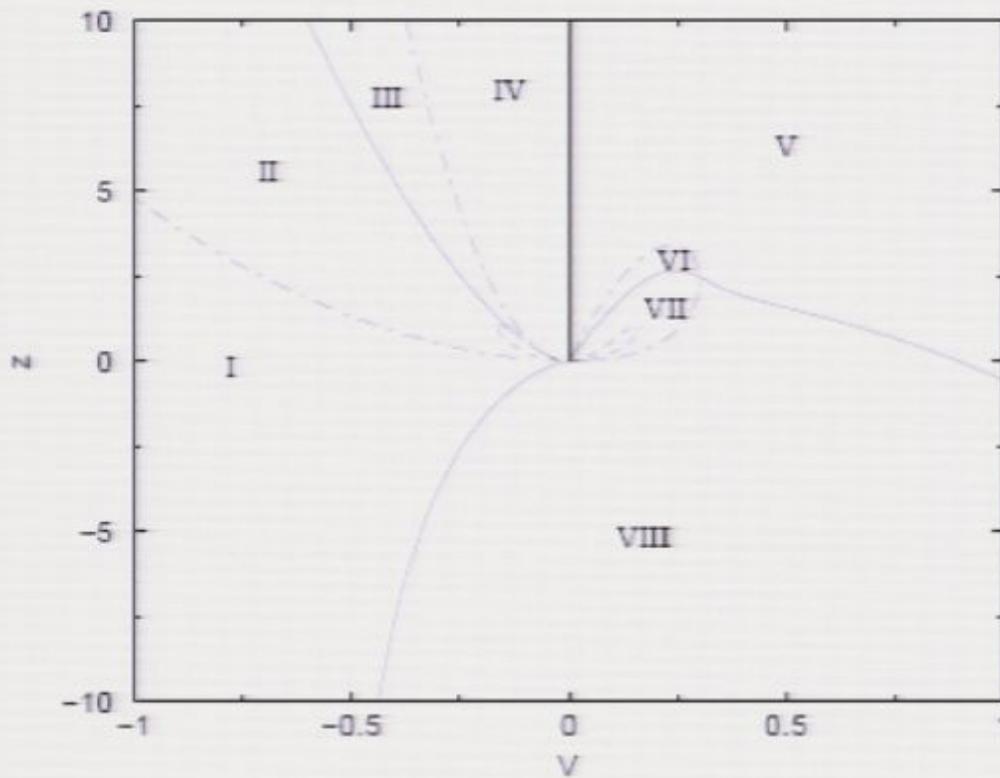
Exterior solutions
all singular

3-brane in 7D EH+GB bulk terms: + case



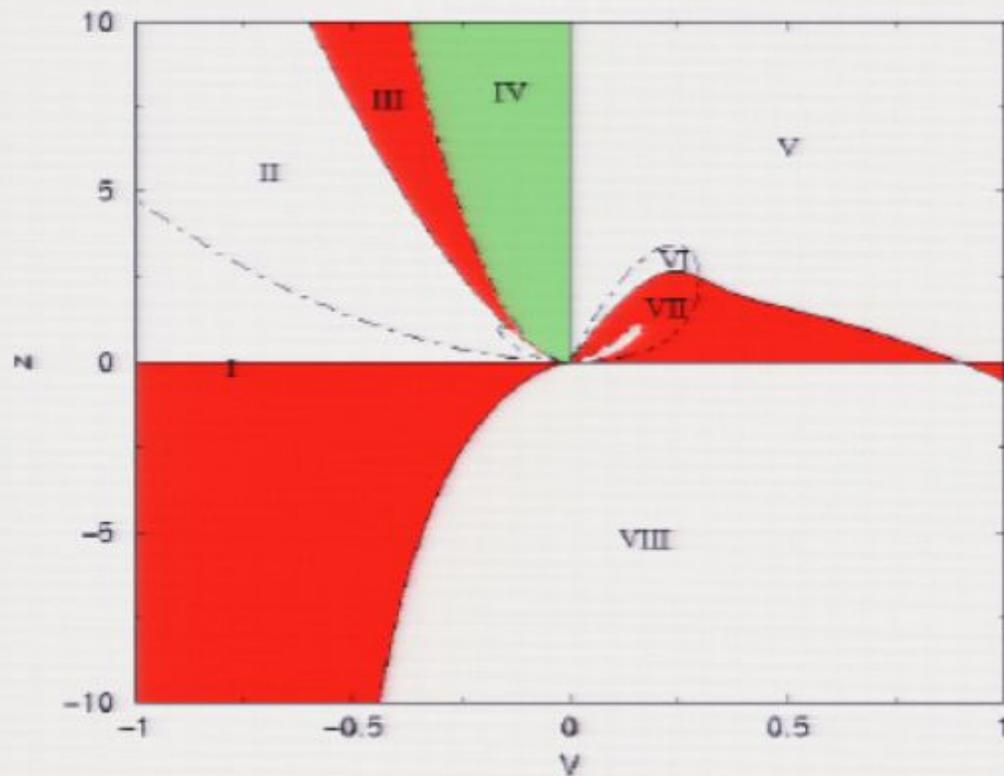
Interior solutions
Region II
 $A \sim \text{const}, B \sim -2 \log r/r_0$
coordinate singularity
 $\tilde{\Lambda} < 0$

3-brane in 7D EH+GB bulk terms: - case



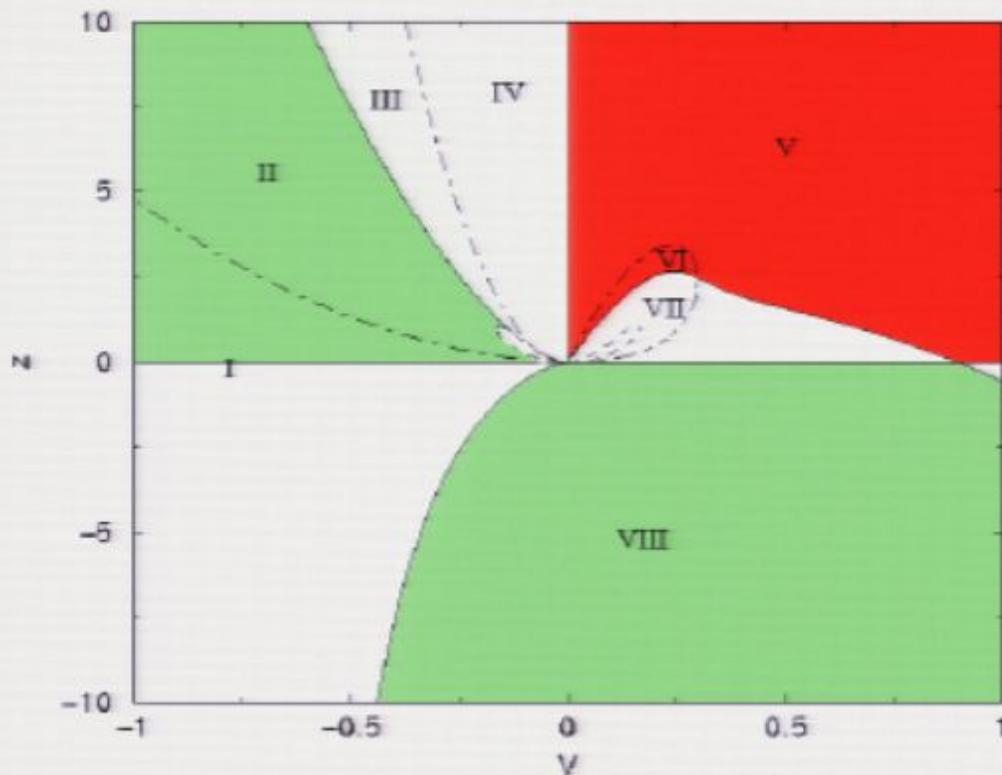
I, III, V, VII $\rightarrow \tilde{\Lambda} > 0$
II, IV, VI, VIII $\rightarrow \tilde{\Lambda} < 0$

3-brane in 7D EH+GB bulk terms: - case



Exterior solutions
Region IV
 $A, B \sim \text{const}$
 $\tilde{\Lambda} < 0$

3-brane in 7D EH+GB bulk terms: - case



Interior solutions

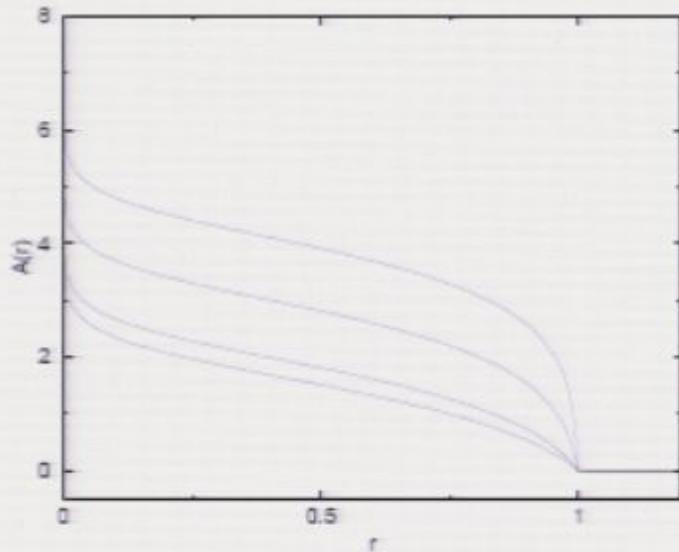
Region VIII

$A \sim \text{const}, B \sim -2 \log r/r_0$
coordinate singularity
 $\tilde{\Lambda} < 0$

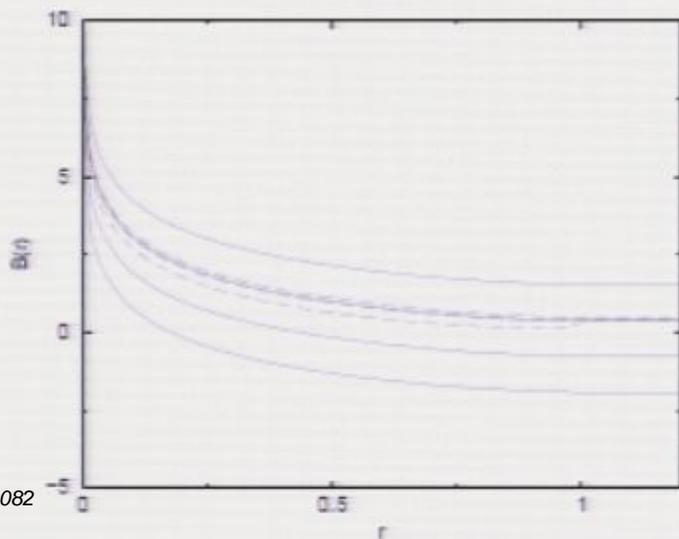
Regions I, II

$A \sim -\log(\log r/r_0),$
 $e^{-B} \sim r \log r/r_0$
coordinate singularity
 $\tilde{\Lambda} > 0$ in I

3-brane in 7D EH+GB bulk terms: - case interior solutions $\tilde{\Lambda} > 0$



$$\tilde{M}^4 / M^2 \xi = 10^8, 10^6, 10^4, 1$$



$$\tilde{\Lambda} = .1, 1, 10, 100$$

Conclusions

- Found smooth interior infinite volume 3-brane solutions with flat worldvolume for a wide range of positive (as well as negative) values of the brane tension that could account for zero point energy of SM fields
- Stability should be further analyzed
- Viability of 4D gravity in this setups needs to be carefully addressed
- These setups provide a fruitful arena to study the phenomenon of dilution of the CC and might hint on a possible way towards a solution to the CC Problem