Title: The Stellar Components of Nearby Galaxy Clusters

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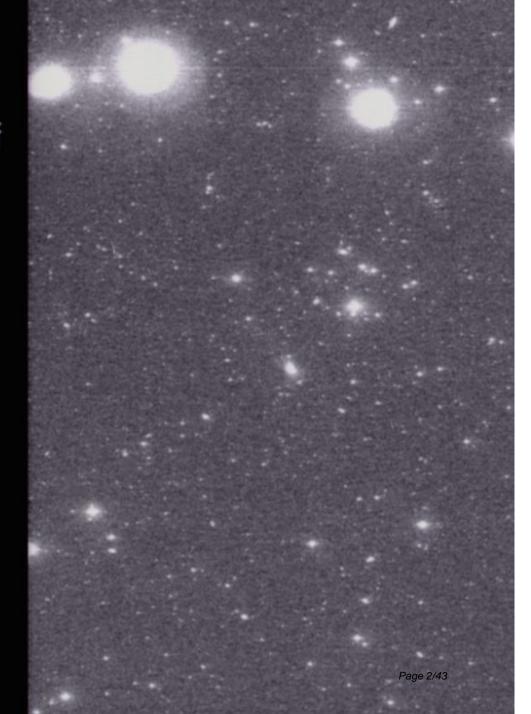
Abstract: A detailed understanding of galaxy clusters is essential in limiting the potential for systematic effects in the use of galaxy clusters for cosmological measurements. I will synthesize our current work on understanding the stellar components of galaxy clusters, including the intergalactic stellar populations and the structural relationship between stellar populations in galaxies and clusters. Ultimately, the characterization of the relationship between dark matter halos and the stellar populations within them will play a key role in unraveling what clusters tell us about the Universe.

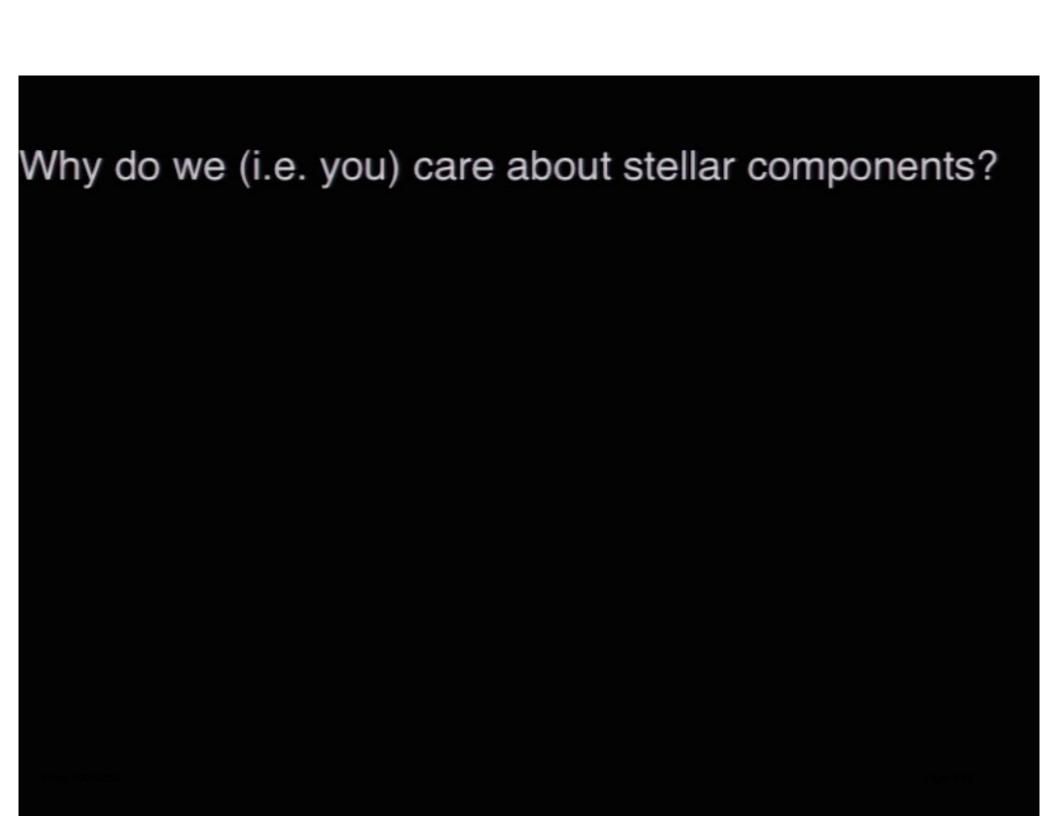
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The Stellar Components of Nearby Galaxy Clusters

Dennis Zaritsky

University of Arizona





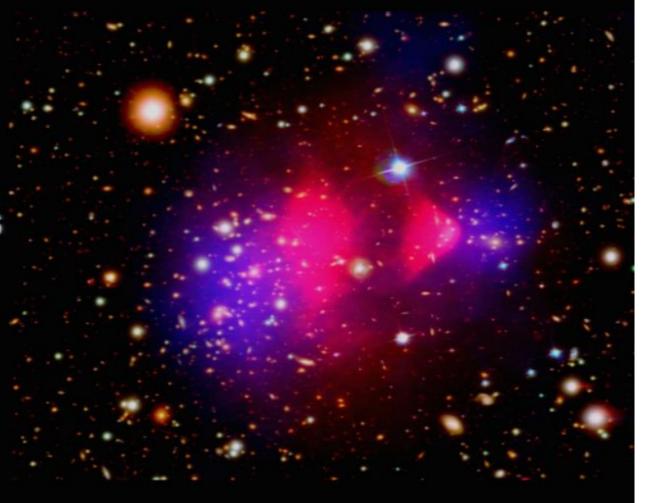
Why do we care about stellar components?

- what's happening



1E0657-56 (the "bullet cluster")

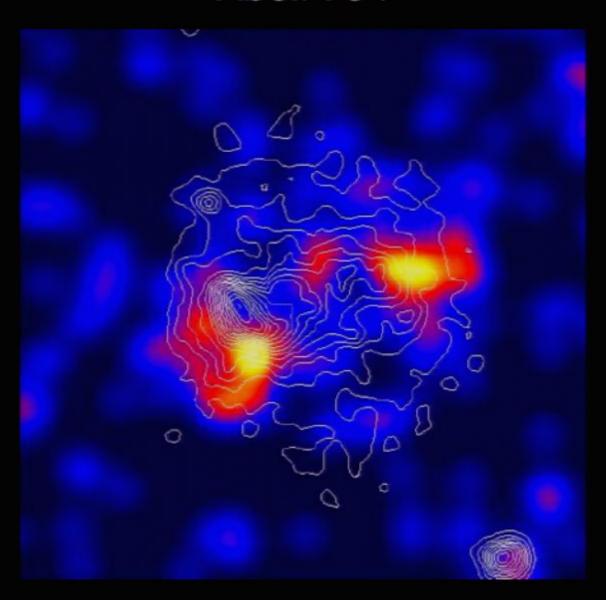
pink = plasma blue = lensing mass image = stars



Clowe, Bradac, Gonzalez, Markevitch, Randall, Jones, & Zaritsky 2006

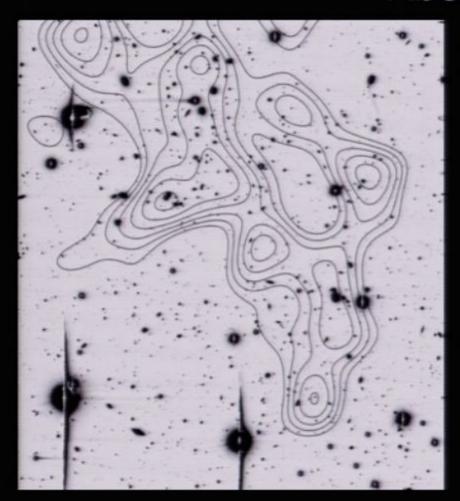
Abell 754

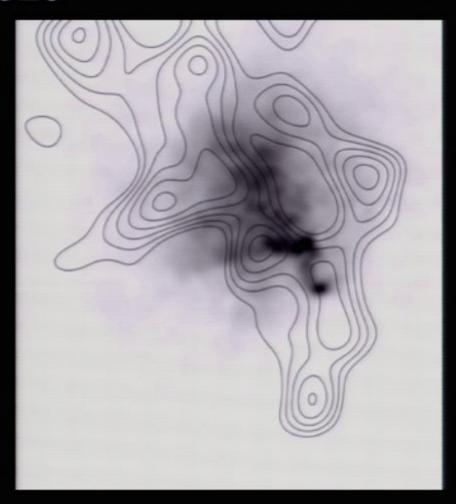
contours = plasma colors = galaxies



Zabludoff & Zaritsky 1995

Abell 520

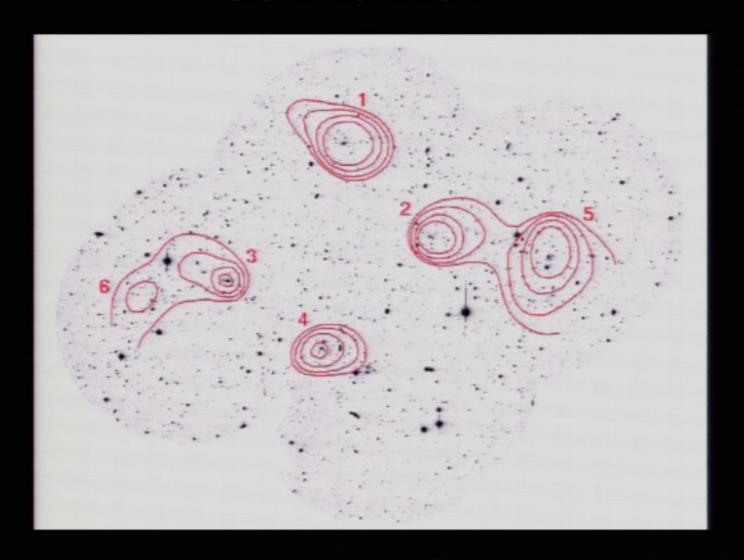




galaxies plasma contours = lensing mass

Clowe et al. in prep.

SG1120-1202



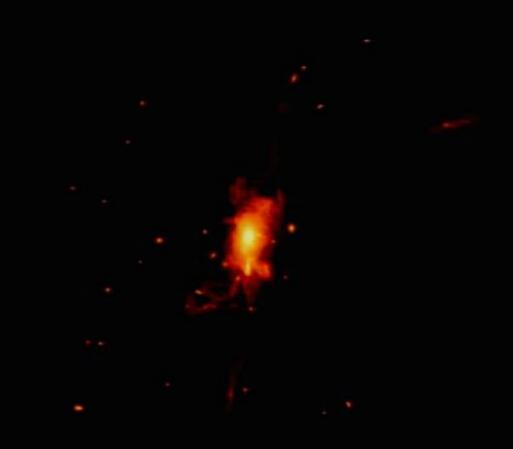
contours = plasma

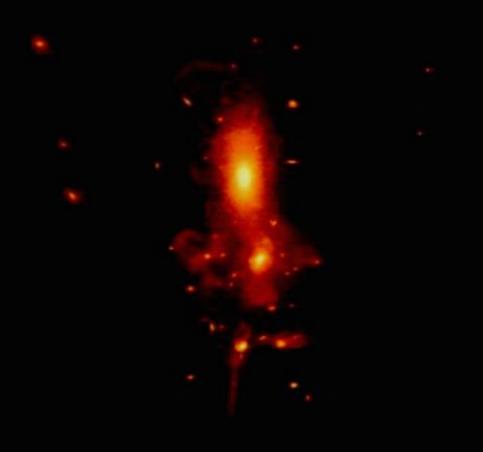
see Gonzalez, Tran, Conbere, & Zaritsky 2005

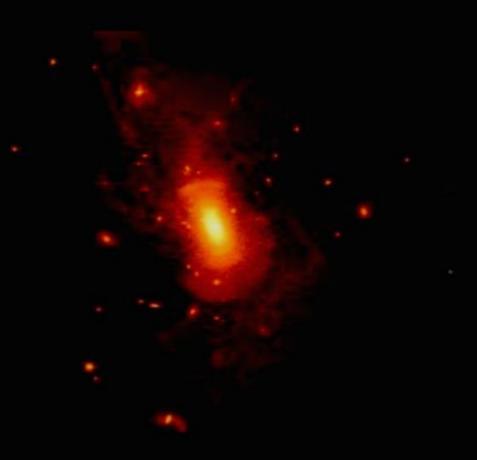
Why do we care about stellar components?

- what's happening
- what happened

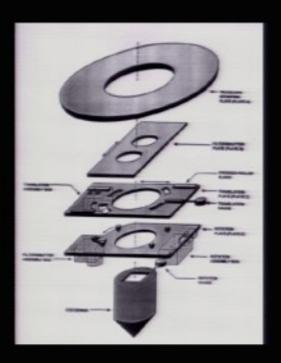






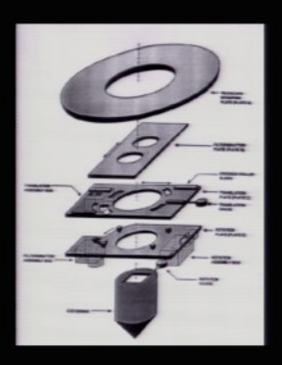




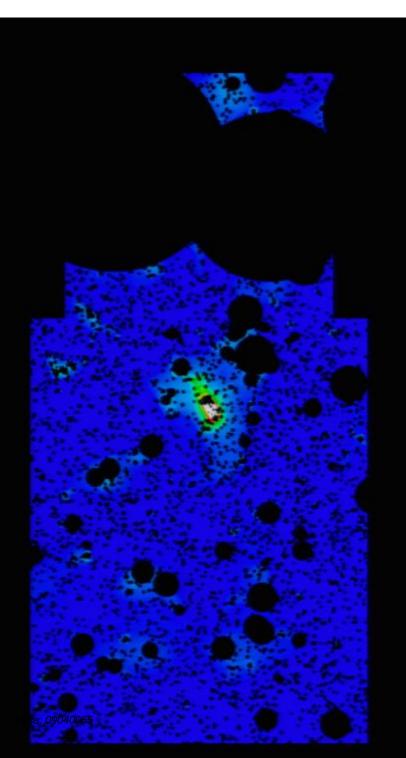


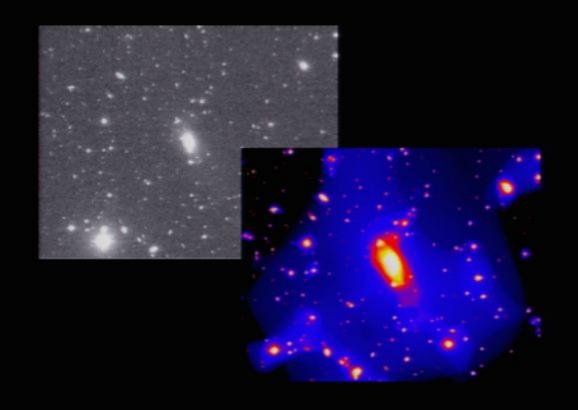








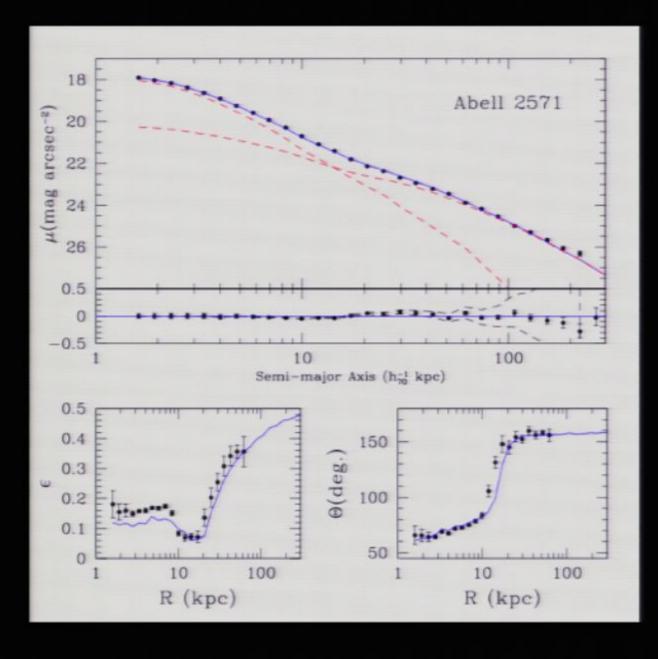


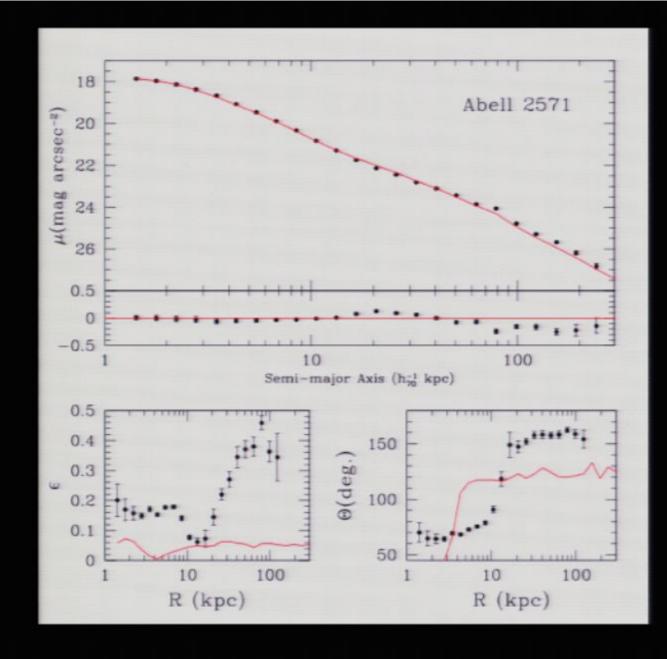


Gonzalez, Zabludoff,
Zaritsky & Dalcanton 2000

Gonzalez, Zabludoff, & Zaritsky 2005

"typical" case

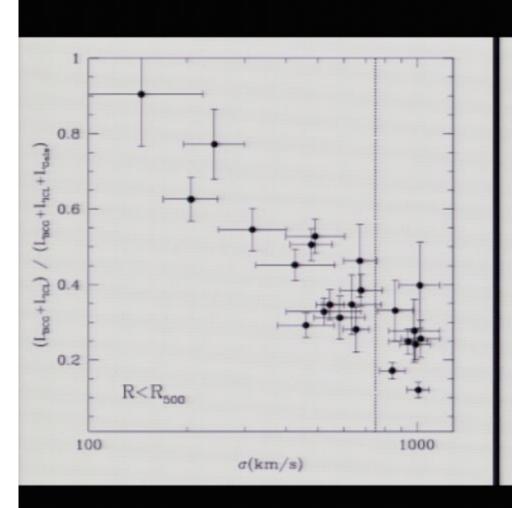


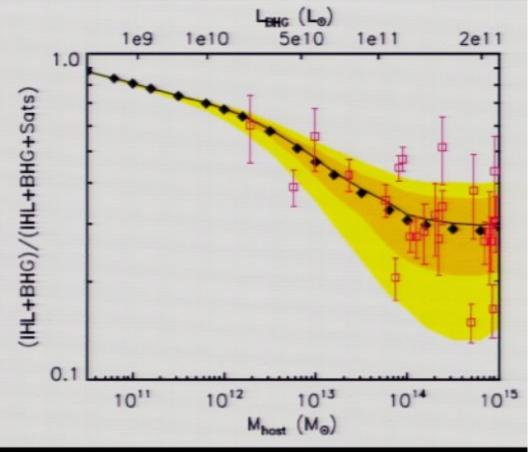


single Sersic $(r^{1/n})$ component

BCG-ICL alignment angle

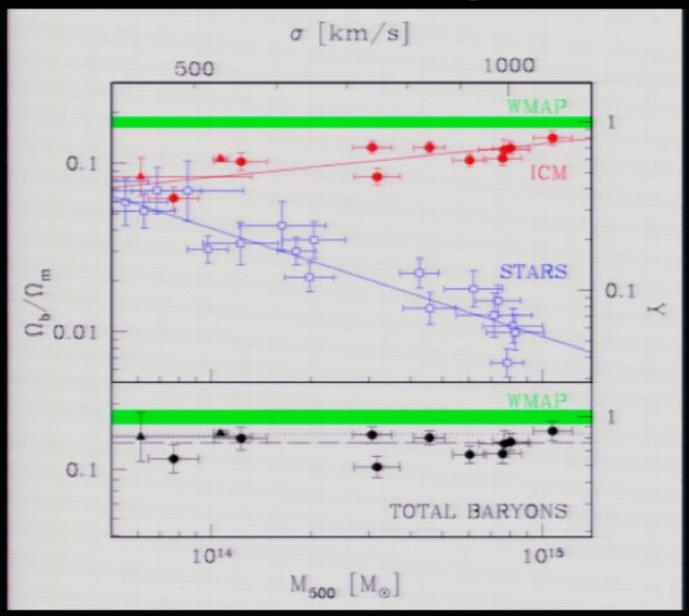






Why do we care about stellar components?

- what's happening
- what happened
- who participated

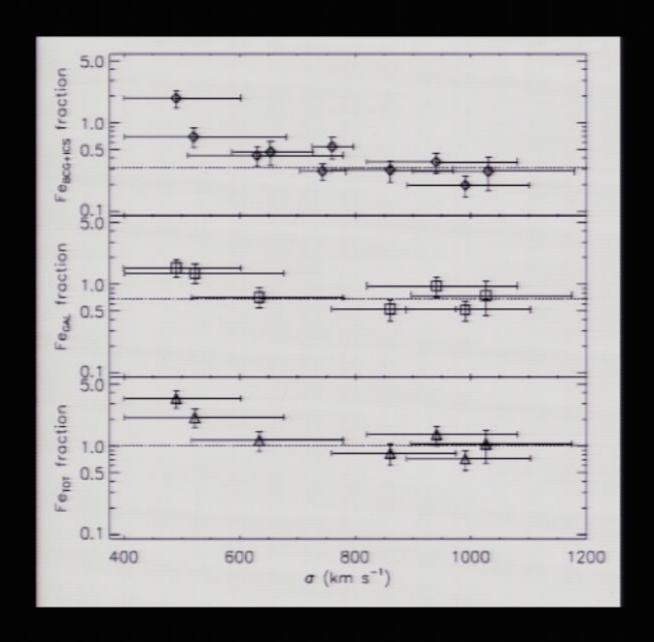


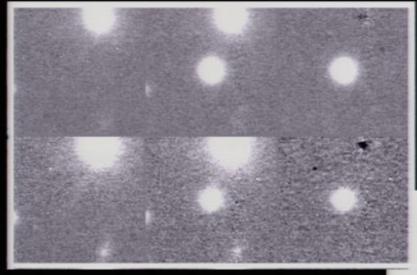
Gonzalez, Zaritsky & Zabludoff 2007

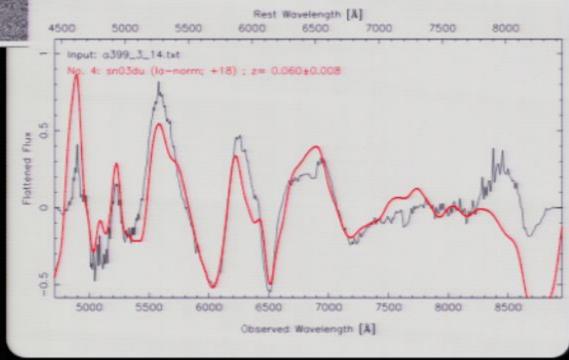
BCG+ICS

Galaxies (85% metal loss)

Total







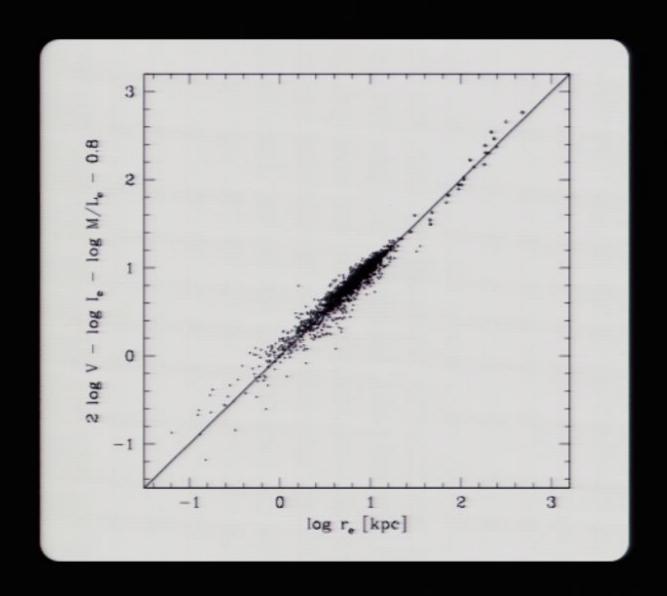
preliminary rate

= $0.039^{+0.025}_{-0.023} \text{ SN}/100 \text{ yr}/10^{10} M_{\odot}$

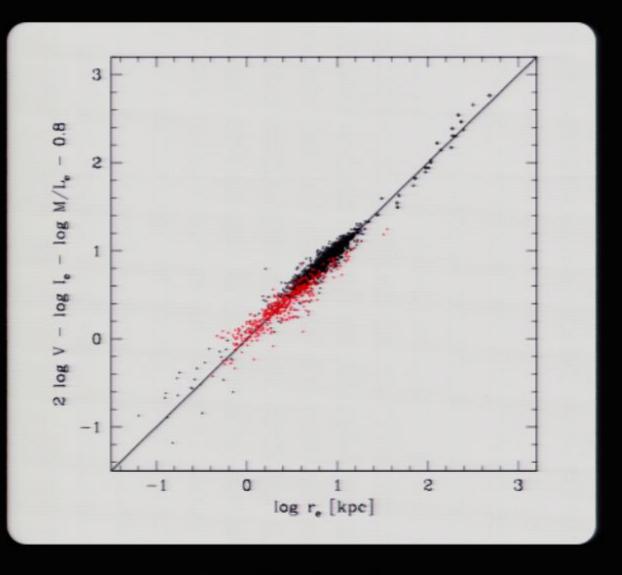
Sand et al. in prep

Why do we care about stellar components?

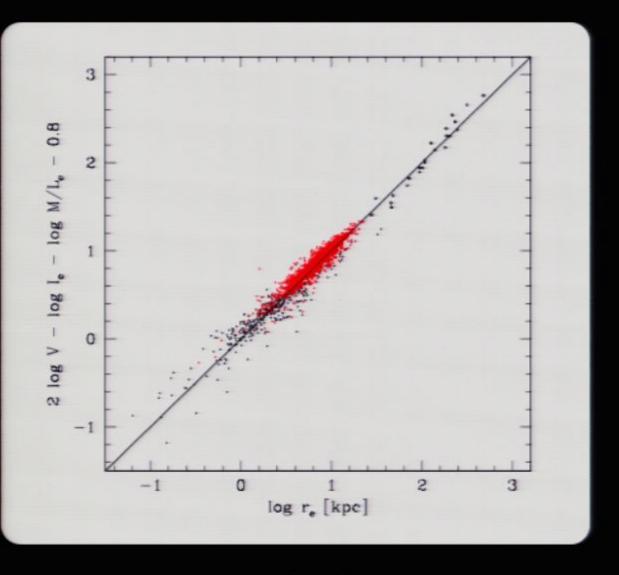
- what's happening
- what happened
- who participated
- how did this happen?



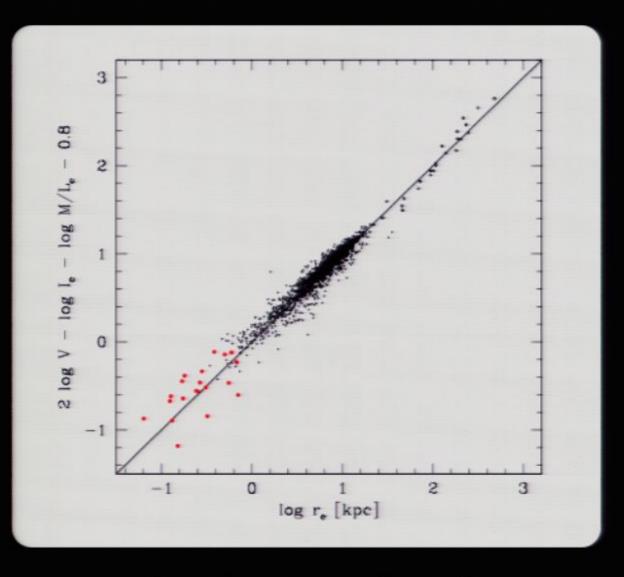
Zaritsky, Gonzalez, & Zabludoff 2006 Zaritsky, Zabludoff, & Gonzalez 2008



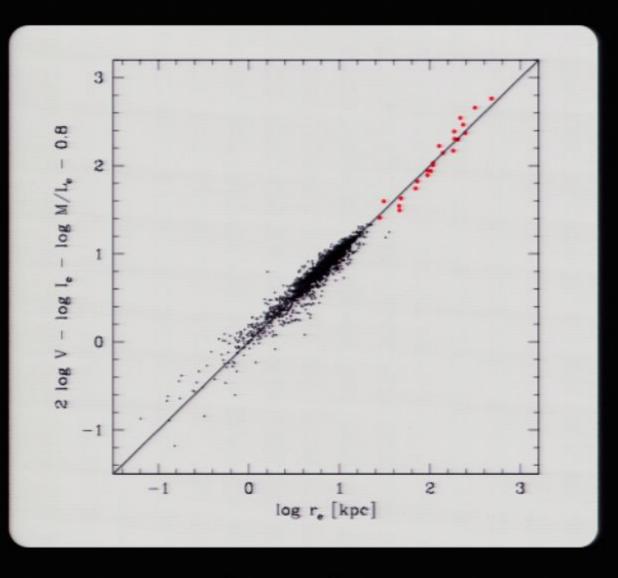
ellipticals



spirals



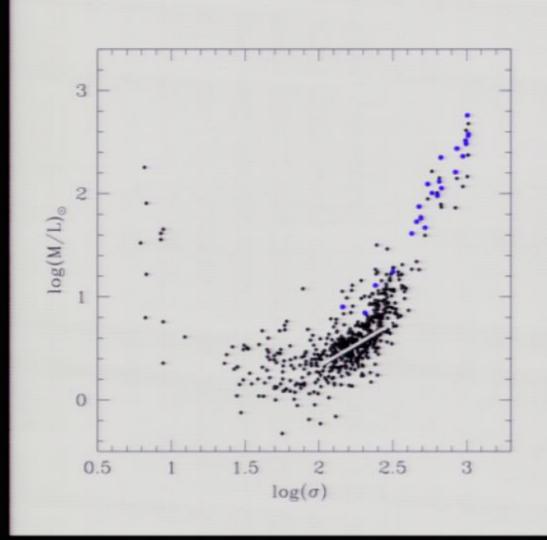
local group dwarfs

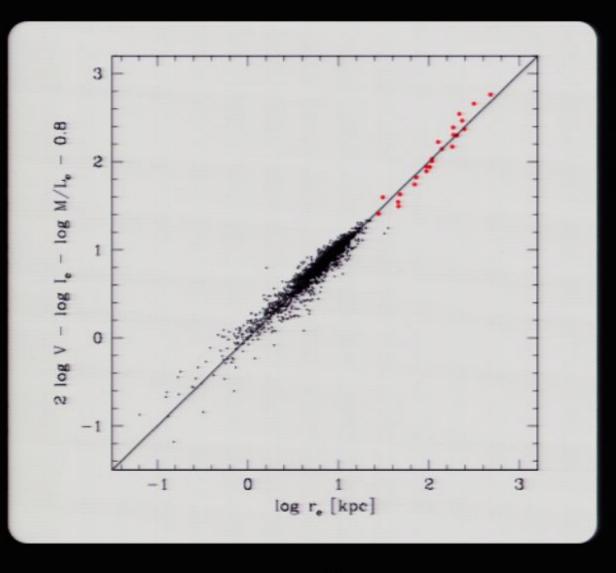


galaxy clusters

black => fitting formula

blue => stars & x-ray determined masses

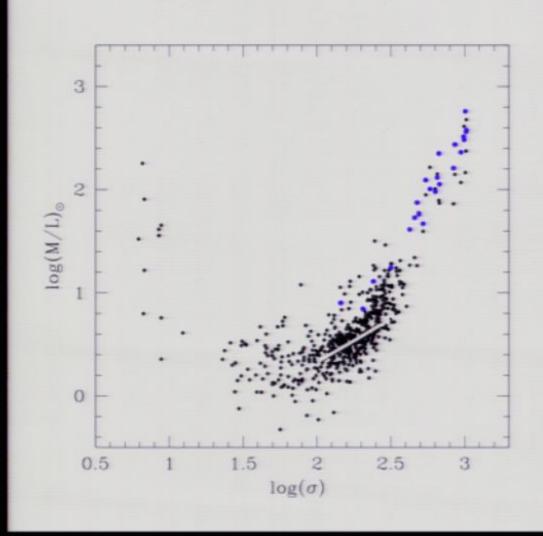




galaxy clusters

black => fitting formula

blue => stars & x-ray determined masses



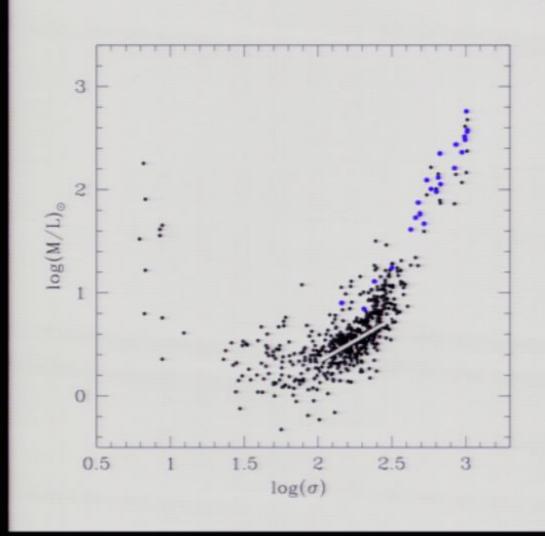
SUMMARY

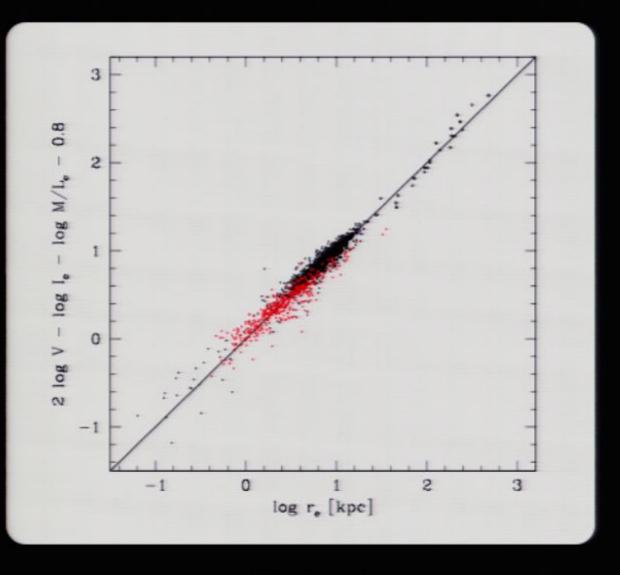
- what's happening (mergers/dark matter)
- what happened (ICL as a witness to the past)
- who participated (baryon & metal accounting)
- how did this happen?

(constraints from a single scaling relation -path to z & sigma?)

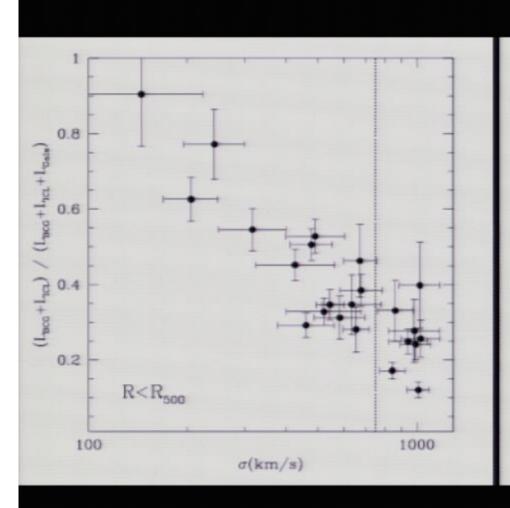
black => fitting formula

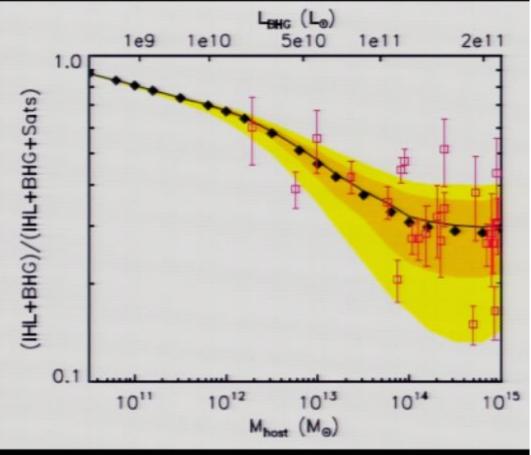
blue => stars & x-ray determined masses

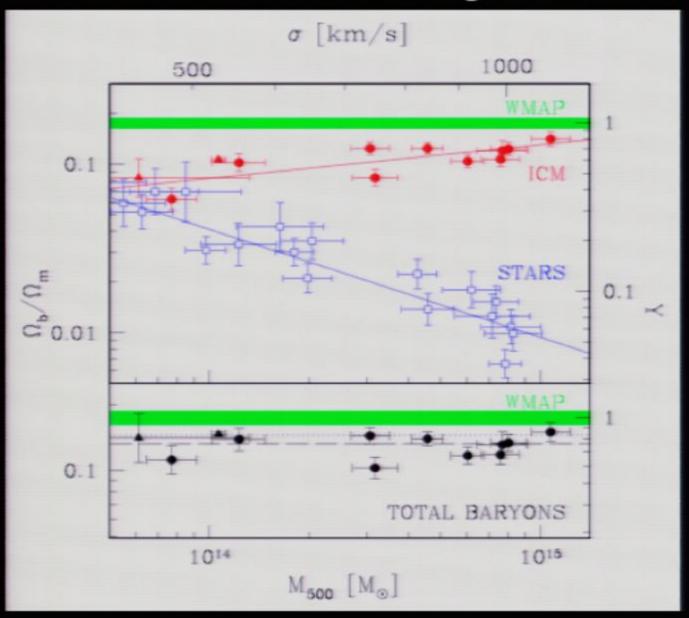




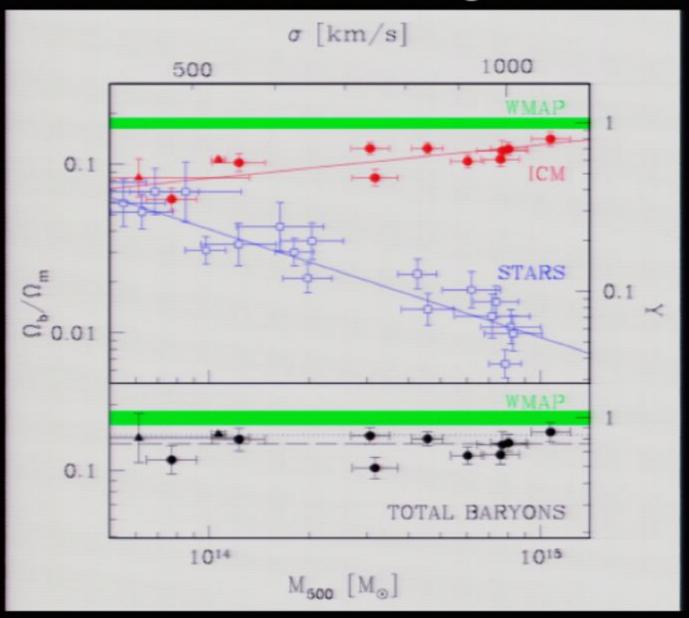
ellipticals



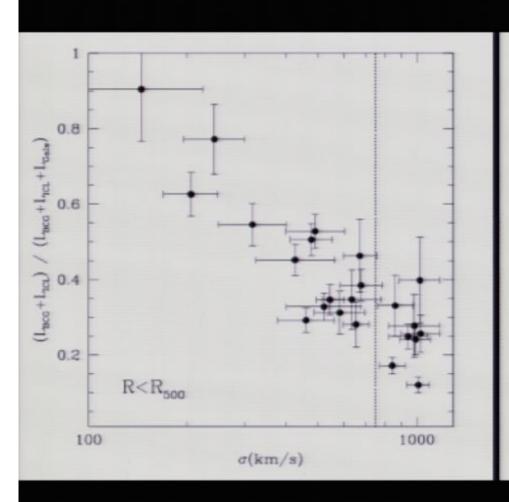


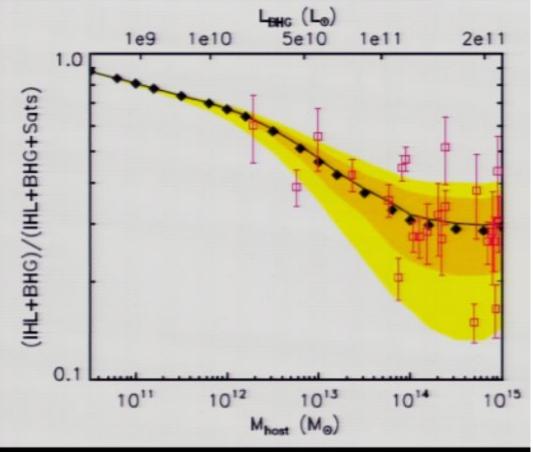


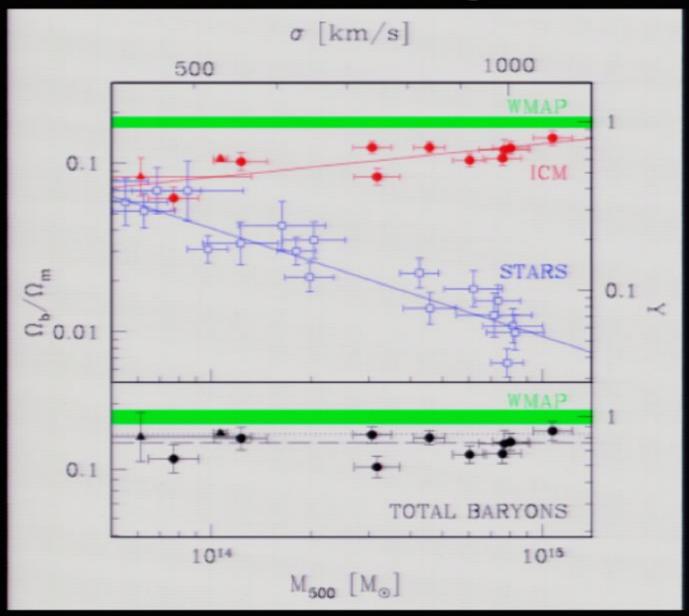
Gonzalez, Zaritsky & Zabludoff 2007



Gonzalez, Zaritsky & Zabludoff 2007







Gonzalez, Zaritsky & Zabludoff 2007