

Title: Asymptotic Expansion of the N=4 Dyon Degeneracy

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Abstract: We study various aspects of power suppressed as well as exponentially suppressed corrections in the asymptotic expansion of the degeneracy of quarter BPS dyons in N=4 supersymmetric string theories. In particular we explicitly calculate the power suppressed corrections up to second order and the first exponentially suppressed corrections. We also propose a macroscopic origin of the exponentially suppressed corrections using the quantum entropy function formalism. This suggests a universal pattern of exponentially suppressed corrections to all four dimensional extremal black hole entropies in string theory.

Macroscopic Side

- We will consider quarter BPS dyonic Black Holes in the Heterotic theory.
- Restricting to Supergravity approximation, we can find the leading entropy carried by these Black Holes .

Microscopic Side

- By duality, we can also regard the states associated with this Black Holes as states of some particular quarter BPS D-brane configuration in the type IIB theory.
- Considering the dynamics of various fields in the D-brane configuration, the complete degeneracy function has been evaluated (J. David, A. Sen; NB, D. Jatkar, A. Sen).

Degeneracy Formula

- The microscopic degeneracy is,

$$d(\vec{Q}, \vec{P}) = (-1)^{Q \cdot P + 1} A \int_{\mathcal{C}} d\rho d\sigma d\nu \frac{e^{-\pi i(\rho Q^2 + \sigma P^2 + 2\nu Q \cdot P)}}{\Phi(\rho, \sigma, \nu)}$$

- Contour \mathcal{C} is a three real dimensional subspace of the complex dimensional space labeled by (ρ, σ, ν) .

- For $N = 1$ theory, the function $\Phi(\rho, \sigma, \nu)$ is a modular form of weight 10.
- The analogous modular forms are also known for many other models .

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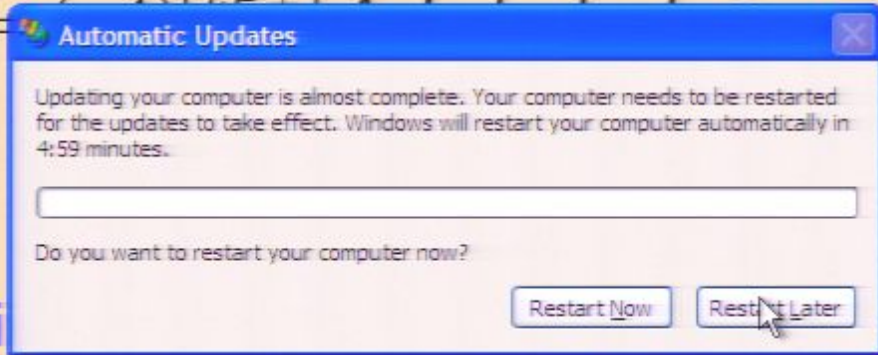
Degeneracy Formula

- The microscopic degeneracy is,

$$d(\vec{Q}, \vec{P}) = \frac{1}{(2\pi)^{3N}} \int d^3v \int d^3\sigma \int d^3\rho e^{-\pi i(\rho Q^2 + \sigma P^2 + 2vQ \cdot P)} \Phi(\rho, \sigma, v)$$

- Contour \mathcal{C} is in the N -dimensional space of the complex dimensional space labeled by (ρ, σ, v) .

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Asymptotic Expansion

A .

- For a given set of charges, there are single centered and multi centered Black Hole solutions.
- We are interested in single centered Black Hole entropy.
- We organize the integral such that the result can pick up the contribution from single centered Black holes. This is done by choosing the integration contour \mathcal{C} in a specific way.
- In particular, we need to set the asymptotic values of the moduli fields equal to their attractor values.

B .

- We have to do three integrals, over (ρ, σ, v) . For this, we need the pole of the integrand.
- The function $\Phi(\rho, \sigma, v)$ has a second order zero at,

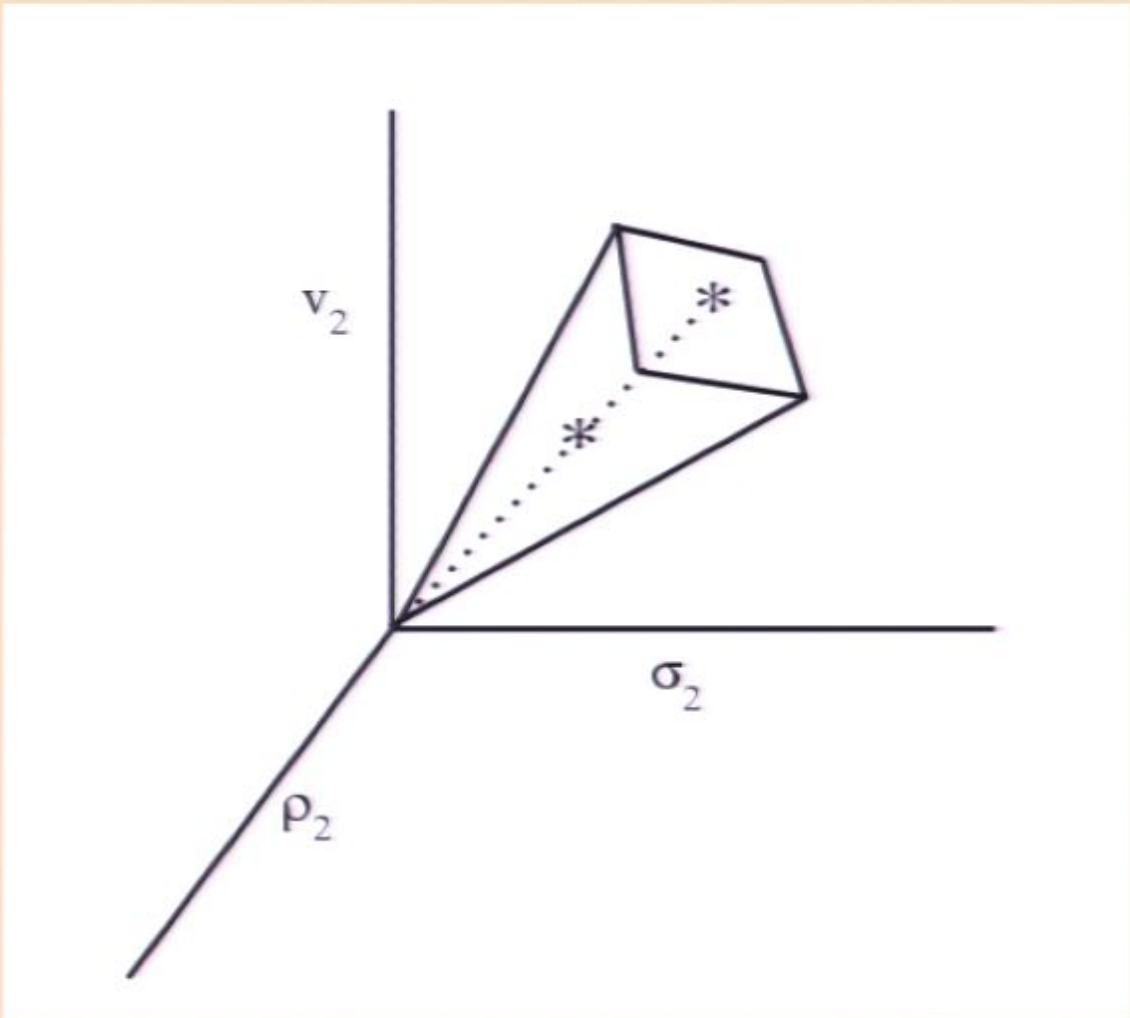
$$n_2(\sigma\rho - v^2) + jv + n_1\sigma - m_1\rho + m_2 = 0$$

for

$$m_1, n_1, m_2, n_2 \in \mathbb{Z}, j \in 2\mathbb{Z} + 1, \quad m_1 n_1 + m_2 n_2 + \frac{j^2}{4} = \frac{1}{4}$$

- We consider cases with $n_2 \geq 1$.

Pole structure



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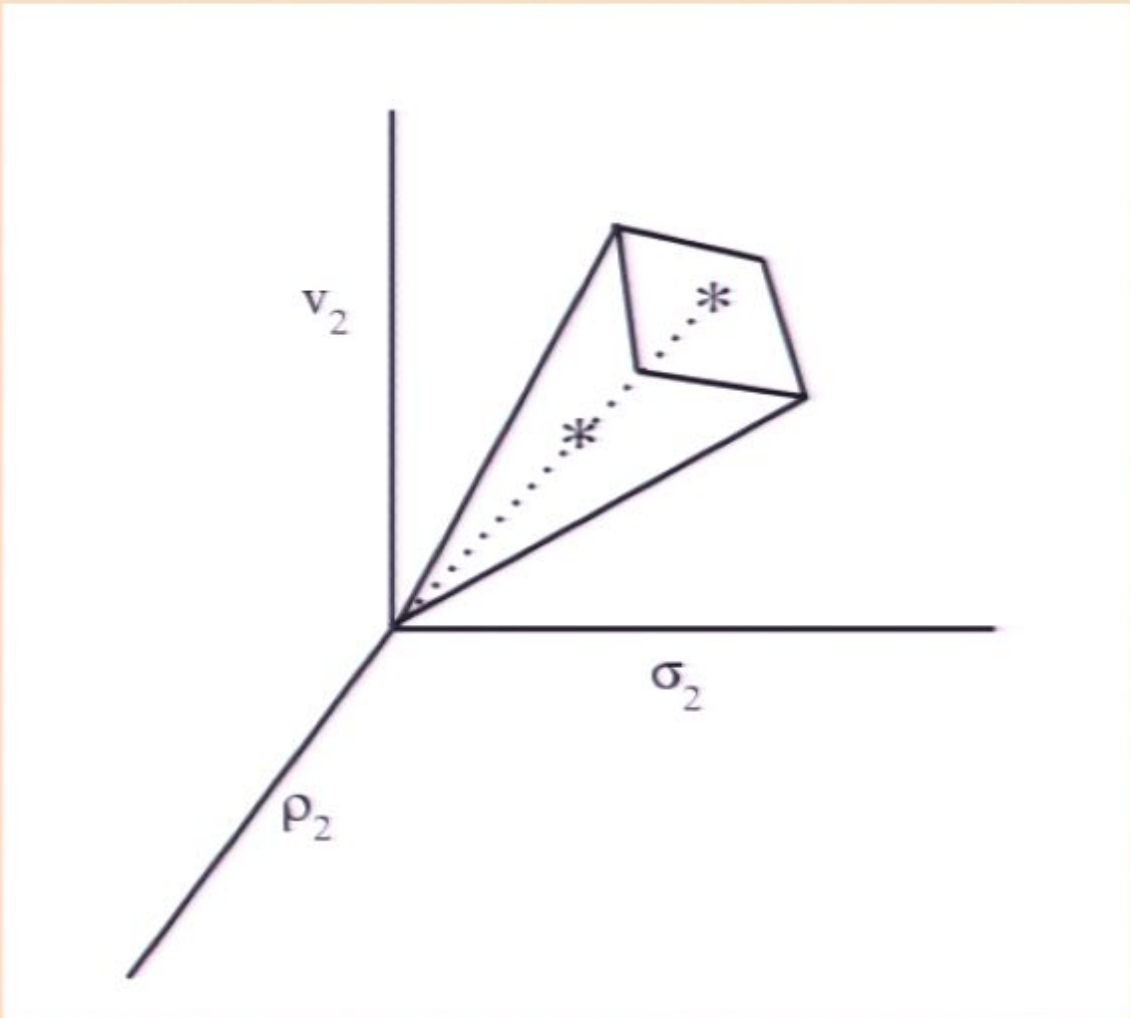
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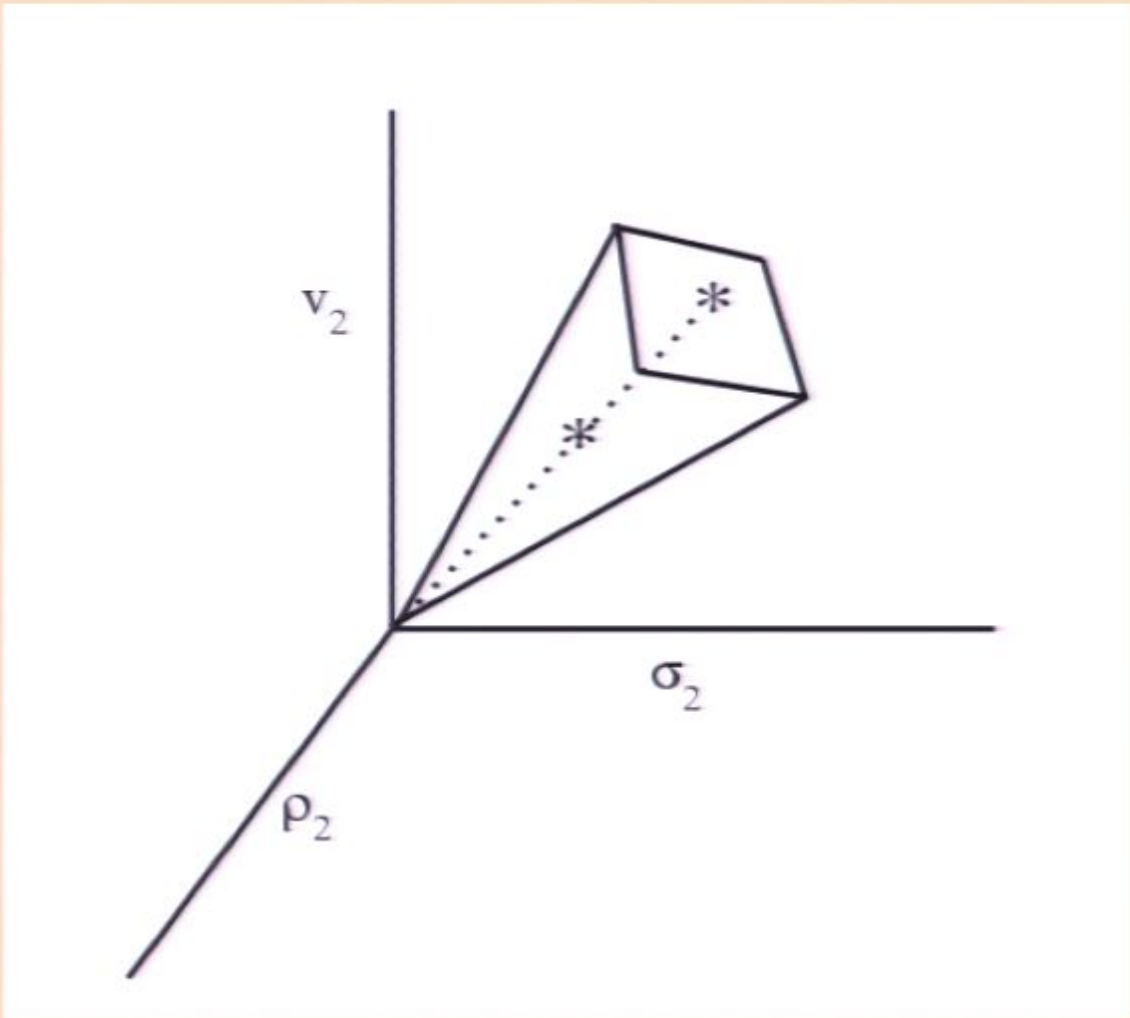
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B.1

- In saddle point approximation of the integral, the degeneracy is,

$$d(\vec{Q}, \vec{P}) = e^{\pi \sqrt{Q^2 P^2 - (Q \cdot P)^2} / n_2}$$

- For $n_2 = 1$, we get the maximum contribution to the degeneracy.
- For $n_2 \geq 2$, the degeneracy is exponentially suppressed compared to the leading one. Hence, to compute exponentially suppressed contribution, we need to look at these sub-leading poles.

B.2

- For the leading pole, one integral can be done by residue method. The other two integrals are done by saddle point analysis.
- The v integral is done by residue method. Near the pole, the function Φ behaves as.

$$\Phi(\rho, \sigma, v) \rightarrow v^2 g(\rho)g(\sigma) + O(v^4)$$

- The (ρ, σ) integral takes the form,

$$e^{S_{stat}(\vec{Q}, \vec{P})} = d(\vec{Q}, \vec{P}) \sim \int \frac{d^2\tau}{\tau_2^2} e^{F(\vec{\tau})}$$

where, $\rho = \tau_1 + i\tau_2$ and $\sigma = -\tau_1 + i\tau_2$

- The function $F(\tau)$ can be easily computed after doing the v integral.

- This can be regarded as a zero dimensional field theory with fields (τ_1, τ_2) with action $F(\vec{\tau}) - 2 \ln \tau_2$.
- The result for statistical entropy S_{stat} can be obtained by computing the possible diagrams of this field theory up to any desired order in charges.
- This will produce all sub-leading correction which are power suppressed in charges.

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Up to $1/\text{charge}^2$ corrections

Statistical Entropy Function

- Leading result:

$$\Gamma_0(\vec{\tau}_B) = -\frac{\pi}{2\tau_{B2}} |Q - \tau_B P|^2$$

- The charge^0 correction:

$$\Gamma_1(\vec{\tau}_B) = \ln g(\tau_B) + \ln g(-\bar{\tau}_B) + (k + 2) \ln(2\tau_{B2})$$

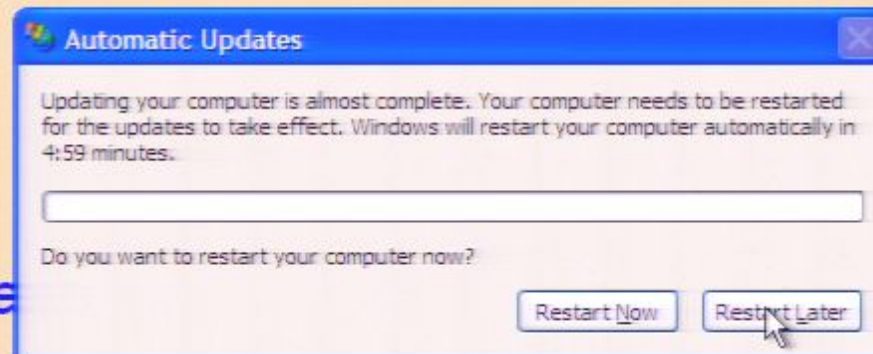
The $1/\text{charge}^2$ correction :

$$\Gamma_2(\vec{\tau}_B) = \ln d_2(\vec{Q}, \vec{P}) = -\frac{\tau_{2B}}{\pi |Q - \tau_B P|^2} (k + 2)$$

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- The charge^0 correction:

$$S^{(1)} = -\ln g(\tau_{(0)}) - \ln g(-\bar{\tau}_{(0)}) - (k + 2) \ln(2\tau_{(0)}\bar{\tau}_{(0)})$$

- The $1/\text{charge}^2$ correction :

$$S^{(2)} = \frac{2 + k}{2\pi \sqrt{Q^2 P^2 - (Q \cdot P)^2}} + \frac{4\tau_{(0)}^3 \bar{\tau}_{(0)}^2}{\pi |Q - \tau_{(0)} P|^2}$$

$$\left[\left(\frac{g'(\tau_{(0)})}{g(\tau_{(0)})} + \frac{k + 2}{\tau_{(0)} - \bar{\tau}_{(0)}} \right) \left(\frac{g'(-\bar{\tau}_{(0)})}{g(-\bar{\tau}_{(0)})} + \frac{k + 2}{\tau_{(0)} - \bar{\tau}_{(0)}} \right) \right]$$

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BMPV Black Holes

Microscopic Configuration

- It involves a D1-D5 system carrying certain momenta and angular momenta in flat space-time.

Results [NB]

- We derive the higher derivative corrections to classical entropy.
- the results are valid even out of Farey tail limit.
- The leading walls of marginal stability is absent in this case.

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Exponential Suppressed Correction

- The zeros of the function Φ ,

$$n_2(\sigma\rho - v^2) + jv + n_1\sigma - m_1\rho + m_2 = 0$$

- We get such corrections from the sub-leading poles,
 $n_2 \geq 2$.

- For this we define

$$\Omega = \begin{bmatrix} \rho & v \\ v & \sigma \end{bmatrix}$$

- We look for a symplectic transformation of the form:

$$\begin{bmatrix} \hat{\rho} & \hat{V} \\ \hat{V} & \hat{\sigma} \end{bmatrix} \equiv \hat{\Omega} = (A \Omega + B)(C \Omega + D)^{-1},$$

such that

$$\hat{V} = \frac{n_2(\sigma\rho - v^2) + jv + n_1\sigma - m_1\rho + m_2}{\det(C\Omega + D)}$$

- The behavior of Φ_k near the zero is,

$$\Phi_k(\rho, \sigma, v) = -\{\det(C \Omega + D)\}^{-k} 4\pi^2 \hat{V}^2 g(\hat{\rho}) g(\hat{\sigma}) + O(v^4)$$

Comparison

Q^2, P^2	$Q.P$	$d(Q, P)$	S_{stat}	$S_{stat}^{(0)}$	$+S_{stat}^{(1)}$	$+S_{stat}^{(2)}$	Δd
2	0	5×10^4	10.82	6.28	10.62	11.58	34.6
4	0	3×10^7	17.31	12.57	16.90	17.38	480.6
6	0	1×10^{10}	23.51	18.85	23.19	23.51	18573
6	3	9×10^8	20.64	16.32	20.41	20.77	0
6	-3	2×10^9	21.78	16.32	20.41	20.77	-

Comparison

Q^2, P^2	$Q.P$	Generic Host Process for Win32 Services				(1) $stat$	$+S_{stat}^{(2)}$	Δd
2	0	<p>Generic Host Process for Win32 Services has encountered a problem and needs to close. We are sorry for the inconvenience.</p> <p>If you were in the middle of something, the information you were working on might be lost.</p> <p>Please tell Microsoft about this problem.</p> <p>We have created an error report that you can send to help us improve Generic Host Process for Win32 Services. We will treat this report as confidential and anonymous.</p> <p>To see what data this error report contains, click here.</p> <p><input type="button" value="Send Error Report"/> <input type="button" value="Don't Send"/></p>				62	11.58	34.6
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Wireless Network Connection is now connected
 Connected to: GuestPass
 Signal Strength: Excellent

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Automatic Updates [X]

Updating your computer is almost complete. Your computer needs to be restarted for the updates to take effect. Windows will restart your computer automatically in 4:59 minutes.

Do you want to restart your computer now?

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Automatic Updates

Updating your computer is almost complete. Your computer needs to be restarted for the updates to take effect. Windows will restart your computer automatically in 4:19 minutes.

Do you want to restart your computer now?

Restart Now Restart Later

Unable to connect to preferred wireless network

Windows could not connect to any of your preferred wireless networks. Windows will keep trying to connect. To see a list of all networks, including others you can connect to, click this message.

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